

1 **Intensity of the Earth's magnetic field: evidence for a Mid-Paleozoic dipole low**

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16

17 **Abstract**

18 The Mesozoic Dipole Low (MDL) is a period, covering at least ~80 million years, of low dipole
19 moment that ended at the start of the Cretaceous Normal Superchron. Recent studies of
20 Devonian age Siberian localities identified similarly low field values a few tens of million years
21 prior to the Permo-Carboniferous Reverse Superchron (PCRS). To constrain the length and
22 timing of this potential new dipole low, this study presents new paleointensity estimates from
23 Strathmore (~411-416 Ma) and Kinghorn (~332 Ma) lava flows, UK. Both localities have been
24 studied for paleomagnetic poles (Q values of 6-7) and the sites were assessed for their suitability

25 for paleointensity from paleodirections, rock magnetic analysis, and microscopy. Thermal- and
26 microwave-IZZI protocol experiments were used to determine site mean paleointensity
27 estimates of $\sim 3\text{-}51 \mu\text{T}$ ($6\text{-}98 \text{ZAm}^2$) and $4\text{-}11 \mu\text{T}$ ($9\text{-}27 \text{ZAm}^2$) from the Strathmore and Kinghorn
28 localities, respectively. These, and all the sites from 200-500 Ma from the (updated) PINT15
29 database, were assessed using the Qualitative Paleointensity criteria (Q_{PI}). The procurement of
30 reliable ($Q_{PI} \geq 5$), weak paleointensity estimates from this and other studies indicates a period of
31 low dipole moment (median field strength of 17ZAm^2) for ~ 80 Myrs, from 332-416 Ma. This
32 “Mid-Paleozoic Dipole Low (MPDL)” bears a number of similarities to the MDL, including the
33 substantial increase in field strength near the onset of the PCRS. The MPDL also adds support to
34 inverse relationship between reversal frequency and field strength and a possible ~ 200 -million-
35 year cycle in paleomagnetic behavior relating to mantle convection.

36 **Significance Statement**

37 Variations in past geomagnetic field strength are important indicators of variation in deep-Earth
38 processes over hundreds of millions of years. Most other geophysical methods only provide a
39 snapshot of the Earth’s recent interior and deep Earth materials are poorly represented in the
40 geological record. New measurements from Scotland (Northern UK), in addition to the existing
41 datasets, show the field was relatively weak (less than half the strength of the long-term
42 average field) for tens of millions of years between 332-416 Ma. The similarities between this
43 and a later period of low field strength provide further evidence for a ~ 200 -million-year cycle
44 linked to deep Earth processes.

45

46 **Introduction**

47 The evolution of the Earth’s deep interior, although critical to our understanding of the planet’s
48 history, is poorly constrained. Reconstructions of convection patterns, such as the configuration
49 and volume of subduction through time (1, 2) and the occurrence of mantle plumes (often
50 inferred from the occurrence of Large Igneous Provinces) (3), as well as the stability of the Large
51 Low-Shear Wave Velocity Provinces (4), are generally ill defined prior to ~ 300 Ma. This lack of

52 constraint is primarily due to poor preservation of materials that formed deep within the Earth
53 and because most geophysical techniques (i.e. seismic tomography, gravity inversion, etc.) can
54 only constrain geologically recent, deep Earth processes. Comparatively, the paleomagnetic
55 record, and paleointensity in particular, has the potential to serve as a key indicator of early
56 deep Earth processes, such as the initiation of the geodynamo (5) and inner core nucleation (6,
57 7). During the Phanerozoic, superchrons (periods of tens of millions of years without magnetic
58 polarity reversals) are suggested to be linked to changes in Earth's deep interior (8, 9). Three
59 superchrons have been identified (10) and correspond with peaks in field strength (11). The
60 superchrons alternate with suspected periods of frequently-reversing weak dipole moments. If
61 confirmed, this pattern would suggest at the existence of a ~200-million-year (Myr) cycle in
62 paleomagnetic behavior, likely resulting from deep Earth processes (8), which alternates
63 between superchrons and these periods of low magnetic dipole moments, such as the Mesozoic
64 Dipole Low (MDL). First proposed by Prevot et al. (12), the MDL is a period of low dipole
65 moment suggested to have lasted for at least the ~80 Myrs preceding the Cretaceous Normal
66 Superchron (CNS: 84-126 Ma). The MDL has since been confirmed by subsequent studies (13,
67 14), which have potentially placed its origin near the end of the Permo-Carboniferous (Kiaman)
68 Reversed Superchron (PCRS; 267-315 Ma) (15). It has also been suggested that the MDL is
69 actually confined to ~150-170 Ma, while the rest of the "MDL" is biased towards low field values
70 due to rock magnetic effects (16). Recent research from Siberian sites (17, 18) found a similar,
71 persistent low dipole strength magnetic field during the Devonian (359 – 419 Ma), lasting ~50
72 Myrs, ~35 Myrs prior to the start of the PCRS. However, it is still unclear if the behavior of the
73 magnetic field during the Devonian and Early Carboniferous (~100 Myrs pre-PCRS) is comparable
74 to that of the MDL, as there is very little available data, with only five studies from this age in
75 the Paleointensity Database (PINT15) (19).

76 To quantify the length of this potential dipole low, two localities from the east coast of Scotland,
77 UK, were selected from this time period to augment the previously published studies (Fig. 1).
78 The first of these, lava flows from the Strathmore region (411-416 Ma) (20), were initially
79 studied comprehensively by Sallomy and Piper (21), who found paleodirections consistent with
80 this early Devonian age. A follow-up paleointensity study by Kono (22), based on a subset of

81 these sites, gave a mean virtual dipole moment (VDM) of $\sim 35 \text{ ZAm}^2$, which is substantially lower
82 than the present-day field strength ($\sim 80 \text{ ZAm}^2$). However, developments in the field of
83 paleointensity in the last 40 years mean that the reliability of this study is now uncertain, as it
84 was done prior to the development of modern day paleointensity techniques and selection
85 criteria. No checks for alteration or multi-domain behavior were included, and recent studies
86 have shown that the perpendicular protocol that was used therein can give artificially low
87 paleointensities (23). The original paleodirectional study has also since been superseded by
88 Torsvik (24). This updated study argued that the high degree of scatter and the presence of sites
89 with 'transitional' directions in the original study, several of which were used for paleointensity,
90 was likely due to bias introduced by the demagnetization techniques used and local tectonic
91 effects.

92 The second locality, lava flows from the beaches along Kinghorn and Burntisland, Scotland (332
93 ± 5.6) (25), has not been studied for paleointensity previously. A paleodirectional study carried
94 out on these lavas by Torsvik (26) found primary directions that were used to determine an Early
95 Carboniferous pole. The new experimental results are presented herein alongside a detailed
96 meta-analysis of published datasets dated to 200–500 Ma using paleointensity quality criteria
97 (Q_{PI}) (27). The outcome supports a new key feature of long-term geomagnetic behavior: the
98 mid-Paleozoic dipole low (MPDL).

99 **Materials and Methods**

100 Detailed geological backgrounds of the Strathmore and Kinghorn localities, along with a
101 description of the sampling techniques used, are provided in *Geological Background and*
102 *Sampling (SI)* and the location of the sites sampled are shown in Fig. 1. The suitability of these
103 sites for paleointensity analysis was first determined using paleodirectional and rock magnetic
104 analysis. The majority of sites were initially stepwise thermally demagnetized (see *Methods:*
105 *Paleodirections (SI)* for details) to determine if the samples carried a stable magnetic
106 remanence. Selection criteria applied to the individual paleodirections obtained were an
107 anchored, maximum angular deviation ($MAD_{ANC} \leq 10^\circ$) and the angle between the anchored and
108 unanchored directions ($\alpha \leq 10^\circ$). Additionally, site directions required some degree of clustering

109 ($k \geq 15$), before being compared with those from the previous studies (24, 26) to determine if the
110 magnetization was of the correct age. Rock magnetic analysis was also performed on all sites,
111 including hysteresis, isothermal remanent magnetization (IRM), back-field, and thermomagnetic
112 (Curie) measurements (see *Methods: Rock Magnetism (SI)* for details). These measurements, in
113 conjunction with scanning electron microscopy (SEM) analysis, were used to determine if the
114 remanence carriers were consistent with the sites carrying a primary Thermal Remanent
115 Magnetization (TRM). Sites that passed all the criteria were then used for paleointensity
116 analysis. Both microwave and thermal Thellier-type paleointensity experiments were performed
117 using the IZZI protocol (28) with partial TRM (pTRM) checks (29). Full details of the experimental
118 procedure used, including selection criteria applied, is provided in *Methods: Paleointensity (SI)*.

119 **Results**

120 **Strathmore paleointensity.** A detailed outline of the results of the paleodirectional and rock
121 magnetism analysis carried out on this locality is described under *Strathmore Results (SI)*. Based
122 on these results, of the eleven Strathmore sites that gave acceptable paleodirections, six were
123 determined to be suitable for paleointensity experimentation. The five sites that produced
124 thermomagnetic curves consistent with (titano-)maghemite (CB2-CB4, TH1 and SN3) were
125 excluded from further analysis because (titano-)maghemite is unlikely to be the original
126 (primary) mineralogy of these lavas (see *Strathmore Results; Rock Magnetism (SI)*; Fig. S3a).
127 Comparatively, the sites measured for paleointensity were deemed to likely retain a primary
128 TRM. Sites CB1 and SN1 produced thermomagnetic curves that indicated the presence of two
129 magnetic minerals, magnetite and hematite (Fig. S3b). The unblocking temperature range of the
130 ChRM supports that the primary remanence is carried by both minerals (see Fig. S1bi and
131 *Strathmore Results; Rock Magnetism* for details). The remaining sites (WB2-WB5) produced
132 thermomagnetic curves consistent with magnetite (Fig. S3d) and gave representative SEM
133 microscopy that showed no evidence of low temperature oxidation (Fig. S3g). Site WB2 was
134 combined with the hematite-bearing (Fig. S3c) baked sediment (WB1) into a single
135 paleointensity site (WB1/2) because the similar characteristic remanent magnetisation (ChRM)
136 directions for the two sites (see *Strathmore Results; Paleodirections (SI)* for details) suggests that

137 WB1 was completely remagnetized by the overlying lava (WB2), which means they would have
138 acquired their TRM in the same field.

139 From the Strathmore locality, 35 out of 82 paleointensity measurements passed selection criteria
140 (given in Table S1; pass rate of 43%) across the six sites. All but one site gave low site-mean
141 paleointensity estimates (3.1-19.7 μT), corresponding to VDMs between 5.6 and 46.2 ZAm^2 (Table
142 1), while site CB1 gave a singularly high site mean estimate of 50.9 μT (98.0 ZAm^2). The majority
143 of accepted Strathmore estimates are from thermal experiments because the microwave
144 demagnetization mechanism was largely unsuitable for hematite bearing sites (CB1, SN1, WB1).
145 The other estimates were split approximately evenly between the two techniques. The pass rate
146 for microwave experiments was 48% vs. 36% for thermal results for the sites that used both
147 methods. Characteristic Arai plots (Fig. 2) show the range of behavior exhibited from the six sites
148 and the different techniques used. The hematite-bearing sites (Fig. 2a-b) showed minimal
149 demagnetization at temperatures below 300°C and linear Arai plots across the temperature
150 ranges for magnetite or both magnetite and hematite. From Wormit Bay, sites WB2-WB3 (Fig. 2c-
151 d) behave similarly, as do WB4-WB5 (Fig. 2e-f), likely because of the similarities in grain size, based
152 on the hysteresis properties of the sites (Fig S3i). Sites WB2-WB3 lie close to the MD range but
153 produce near-linear orthogonal vector and Arai plots, whereas sites WB4-WB5 exhibit some zig-
154 zagging of the corresponding orthogonal plots and the main occurrence of prominently two-
155 sloped Arai plots; however, as seen in the orthogonal plots from the corresponding thermal
156 demagnetization experiments (Fig. S1di), the samples display more than one directional
157 component, and the temperature range of the ChRM corresponds to the selected component
158 from the thermal paleointensity experiments. Accepted measurements are considered
159 unaffected by anisotropy, based on measurement gamma values (γ), and non-linear TRM effects,
160 as both the ancient and applied fields used are relatively low ($\leq 60 \mu\text{T}$). The Wormit Bay sites are
161 also unlikely to have been affected by cooling rate based on the samples' grain size distribution
162 (non-single domain (SD) grains (30); Fig S2i) and any variation between results from techniques
163 with different cooling rates (thermal vs. microwave) or mineralogy with different grain size
164 distributions (magnetite vs. hematite) being inconsistent with the expected effects from cooling

165 rate differences (see *Paleointensity Reliability Assessment; Strathmore Q_{PI} scoring*, the ACN
166 criterion, for further details).

167

168 **Kinghorn paleointensity.** A detailed outline of the results of the paleodirectional and rock
169 magnetism analysis carried out on this locality is described under *Kinghorn Results (SI)*. All the
170 Kinghorn sites that passed the paleodirectional selection criteria (six out of nine sites) were
171 deemed suitable for paleointensity, along with two additional sites (KHA-KHB; Fig. S2b), for a total
172 of eight paleointensity sites. Some sites had curves that were consistent with relatively Ti-rich
173 titanomagnetite. (KH1 and KHA-KHB; Fig. S3f), while the rest (KH2, KH4 and KH7-KH10; KH8/9 is a
174 single site) were consistent with magnetite or low-Ti titanomagnetite (Fig. S3e). Representative
175 SEM microscopy showed primary igneous textures (i.e. coarse exsolution structures) and no
176 evidence of low temperature oxidation (Fig. S3h), which is consistent with the sites carrying a
177 primary TRM.

178

179 Of the Kinghorn locality samples, 53 out of 143 measurements passed selection criteria (given in
180 Table S2; pass rate of 37%). All sites produced very low site mean estimates (3.7-10.9 μT ; Table
181 1), corresponding to exceptionally low VDM estimates (9.6-27.0 ZAm^2). The majority of the
182 accepted measurements were made using the microwave system, as it had a much higher
183 success rate (54% success rate vs. the 9% success rate for thermal experiments). This may be
184 because the relatively Ti-rich titanomagnetite is prone to altering more in the thermal than the
185 microwave experiments due to reduced bulk heating of the samples in the latter (31). The
186 appearance of the Arai plots varies, with some sites producing near linear plots with minimal
187 overprints (Fig. 2a, f), whereas others exhibit varying degrees of two-slope behavior (Fig 2b-e).
188 Like sites WB4-WB5, the two-slope Arai plots all show a corresponding change in direction,
189 which is also indicated in the thermal demagnetization orthogonal plots (Fig. S1ei-fi). These plots
190 suggest that the steep, low-temperature slope is likely an overprint rather than due to non-ideal
191 behavior. This interpretation is also supported by a lack of zig-zagging when using IZZI protocol
192 and the low degrees of curvature (which was evaluated using curvature, $|k'|$). There is no clear
193 correlation between the appearance of the Arai plots and either the titanium content of the
194 titanomagnetite or the apparent grain size (Fig. S3j). Accepted measurements are considered

195 unaffected by cooling rate, due to the samples' non-SD grain size (Fig S2j) and the lavas being
196 "fast-cooled" (27), and non-linear TRM effects (for the same reason as the Strathmore sites).
197 Anisotropy checks on the Kinghorn samples, performed due to relatively high γ values, show
198 that some of the Kinghorn sites are slightly anisotropic, but the anisotropic corrections required
199 (which vary depend on the applied field direction) are negligible and average out to give nearly
200 identical results to the uncorrected site means (see *Paleointensity Reliability Assessment*;
201 *Kinghorn Q_{PI} scoring*, the ACN criterion, for further details and quantification).

202 **Discussion**

203 **Reliability of 200-500 Ma sites.** To further assess the reliability of these new site-mean
204 paleointensity estimates, and to provide a framework for comparing them to others from the
205 Paleozoic (~252-541 Ma), all sites were evaluated using Q_{PI} criteria. Biggin and Paterson (32)
206 proposed these nine criteria to acknowledge and mitigate the potential biases that affect the
207 interpretation of paleointensity data and are applied in a similar way to Q criteria for
208 paleomagnetic poles (33). Sites that have published information addressing a criterion pass
209 (score a 1), and, if not, they fail (score a 0). The Q_{PI} score for the site is the sum of the individual
210 criterion scores. Detailed descriptions of these nine Q_{PI} criteria and how they have been
211 assessed for this study are included in *Paleointensity Reliability Assessment; Q_{PI} criteria (SI)* and
212 explanations for how the Kinghorn and Strathmore sites were scored is provided in
213 *Paleointensity Reliability Assessment; Strathmore scoring* and *Kinghorn scoring (SI)* respectively.
214 A summary of the individual criteria scores for each site is provided in Table 1 and in Dataset S3.
215 The new Strathmore sites in this study received Q_{PI} scores ranging from 5 to 8 (median: 6.5,
216 mean: 6.3), out of a possible 9. The new Kinghorn sites similarly received Q_{PI} scores ranging from
217 6 to 8 (with a slightly higher median of 7 and mean of 6.8). STAT and LITH are the Q_{PI} criteria
218 that the majority of sites failed. The failure to meet STAT is largely because the sites gave low
219 paleointensity estimates, which are less likely to pass STAT due to its definition (see
220 *Paleointensity Reliability Assessment; Strathmore scoring for details (SI)*), and LITH because only
221 one site had a suitable contact lithology that could be sampled. MD and TRM failing for some of
222 the Strathmore sites is why its Q_{PI} scores are slightly lower than those of the Kinghorn sites.
223 However, the sites are not considered unreliable as their failure to pass these criteria need not

224 require that they have been remagnetized and/or affected by MD behavior. We note that our
225 results are also very similar to those obtained from Siberian rocks of a similar age (7-39 ZAm²
226 with a single outlier site of 98 ZAm²) (17).

227 Integration of these new estimates with the existing Paleozoic dataset first requires
228 determination of what published data are sufficiently robust for meta-analysis. All the data in
229 the PINT15 database (19) from 200-500 Ma were checked against their corresponding study to
230 fix any errors. Ages were recalculated where possible (e.g. stratigraphic ages were revised to be
231 consistent with the most recent timescale (ICS2020/v1), isotopic ages were replaced where
232 superseding ages are known, etc.). The biggest reassessment of site ages comes from the
233 apparently Middle-Late Carboniferous from Uzbekistan (34, 35). The relative ages between sites
234 and the single inclination sign across multiple sections, with 13-40 sites per section, indicate that
235 these sites could only come from the part of the Carboniferous during the PCRS (i.e. Moscovian-
236 Gzhelian; 298.9-315.2 Ma). Sites from 5 studies published since the last PINT15 update were
237 also added (references listed in Dataset S4). Q_{PI} criteria were applied to all the sites, based on
238 the published information from the corresponding studies. This time period covers both the
239 PCRS and the surrounding time periods, which allows paleomagnetic field strengths during the
240 superchron to be compared with those from a reversing field. This time period also
241 complements two other Q_{PI} studies that assessed the PINT15 database for 500-3500 Ma (6) and
242 65-200 Ma (16).

243 The revised PINT15 data for 200-500 Ma, including Q_{PI} scoring, is included in Dataset S3. A
244 workflow for the scores provided is outlined in Dataset S4, and the age distribution, coverage,
245 and reliability of the revised 200-500 Ma PINT15 data is illustrated in Fig. 4. Given that most
246 studies from this time period were published before Q_{PI} criteria existed, their Q_{PI} scores tend to
247 be lower because there is insufficient information published to confirm that a potential issue
248 has been addressed, rather than it being clear that that issue has affected the estimates. Only
249 sites with Q_{PI} scores of 0 will be excluded entirely; these either have no published information to
250 support the reliability of the site means or they have been confirmed to be unreliable. All the
251 highest scoring sites ($Q_{PI} \geq 5$) are found in the time periods immediately before (16 sites) and
252 after (26 sites) the PCRS, which comprises the data from this study along with recently published

253 studies (17, 18). While there are numerous sites with PCRS ages, 144 of the 195 sites (74%)
254 covering this period come from just four studies: (34, 36–38). The Q_{PI} scores for these are low
255 because these publications use outdated paleointensity methods (i.e. calculating the ratio of
256 total NRM/total TRM) (39) and include very little supporting information.

257 **Paleozoic Field Variation.** Based on the pattern of field strength variation from these new site
258 means and the existing PINT15 dataset, weighted by Q_{PI} score (Fig.4), a relatively long period of
259 low dipole moment presents itself in the period preceding the PCRS, followed by a substantial
260 increase in field strength during the superchron relative to periods of reversing field. To evaluate
261 whether this variation is similar to that observed during the Mesozoic, an analysis was
262 performed, following the methodology of previous studies (6, 16), by comparing the field
263 strength distribution of different periods, filtered by Q_{PI} scoring. The combined dataset was
264 grouped into 3 bins using the superchron as an anchor: PRE (315-416 Ma), PCRS (267-315 Ma),
265 and POST (200-267 Ma). Kulakov et al. (16) was able to identify periods of distinct dipole
266 moment during the MDL based on reversal frequency. However, the reversal record prior to the
267 PCRS is too sparse to apply the same technique (Fig. 4), so the PRE bin was not divided further.
268 Its maximum age bound was set to 416 Ma to avoid the Ordovician Reversed Superchron (ORS;
269 461-480 Ma) (40) and because there are no estimates with $Q_{PI} \geq 1$ between 416-461 Ma. There is
270 no analysis of the ORS, or the period before it, as between 461-500 Ma there are only 3
271 available estimates. The age distribution of the POST bin is also substantially skewed (skewness
272 = -6.14 at $Q_{PI} \geq 3$; see Table S2) as there are only 13 site-mean results between 200-250 Ma and
273 an abundance of data around ~250 Ma. This peak in the data is almost entirely the result of a
274 large number of studies from the Siberian Traps (see Anwar et al. (15) for details); however, the
275 paucity of data between 200-250 Ma means it cannot be connected to the Early bin from
276 Kulakov et al. (16) and should be considered independent of it.

277 Q_{PI} filtering was applied based on the total Q_{PI} score for the site, up to a $Q_{PI} \geq 5$ (because the
278 PCRS bin does not include any sites that have a total score greater than five). Fig. 5a-e illustrates
279 the distribution of site VDMs in these bins for different Q_{PI} minima, while details of the data
280 included in these bins are included in Table S2. However, not all the Q_{PI} criteria are considered to
281 be of equal weight. As discussed in Kulakov et al. (16), AGE, STAT, TRM, MD and ALT are

282 considered the priority criteria; however, it is not possible to make all these criteria mandatory
283 and still have sufficiently large datasets for statistical analysis. AGE and ALT are considered
284 essential criteria to retain (16), but the way the TRM criterion is judged (see *Paleointensity*
285 *Reliability Assessment; Q_{PI} criteria (SI)*) means that a site can carry a TRM but not successfully
286 meet the criterion. Too few PCRS sites pass TRM (11 sites) for further analysis to be statistically
287 meaningful (especially when requiring AGE and ALT to pass as well). The inverse problem exists
288 for the STAT criterion: very few sites pass the criterion for the PRE (18 sites) and POST (16 sites)
289 bins, probably because of the weak site mean value, which is discussed in *Paleointensity*
290 *Reliability Assessment; Kinghorn Scoring (SI)*. However, requiring the inclusion of the MD
291 criterion still retained a statistically rigorous number of sites across all three age bins (171 sites
292 passed for AGE, ALT and MD). The inclusion of the MD criterion is also particularly important
293 because MD effects on paleointensity experiments have been cited as a primary cause for the
294 apparent long-lasting weak strength of the paleomagnetic field, compared to the present day
295 field value (41). The distribution of VDMs that pass all these criteria (AGE+ALT+MD) is shown in
296 Fig. 5f, and details of the data included in these bins are included in Table S2. Finally, to check
297 that the bin statistics are not being biased by oversampling of the field at certain time periods
298 (i.e. 250 Ma), the subsets of sites with total Q_{PI} scores ≥ 1 and ≥ 3 , were binned in 10 Ma and 5
299 Ma intervals before averaging (details in Table S2).

300 A series of Kolmogorov-Smirnov (K-S) tests were run to compare the bins pair-wise (Table S2).
301 The tests reject the null hypothesis that the datasets come from the same distribution at the 1%
302 significance level up to $Q_{PI} \geq 5$ (the PCRS and POST bins have too few data to be significant if a
303 higher Q_{PI} cut-off is applied; see Table S2). All three time periods, therefore, have distinct VDM
304 distributions. Fig. 5(a-f) indicates that the dipole moment during the PCRS is substantially higher
305 than the surrounding time periods, regardless of Q_{PI} filtering, although its average strength is
306 unclear. The median values for the PCRS range from 39–94 ZAm^2 depending on the Q_{PI} filters
307 and binning used. Although the median CNS values (48-59 ZAm^2), based on similar analysis (16),
308 fit within this range, the large range in average field strength during the PCRS points to issues
309 with the current PCRS dataset. The difference emerges from the greater number of low
310 estimates recorded for the CNS than the PCRS. The higher median values are generally from

311 subsets with lower levels of reliability (low Q_{PI} scores/number of bins) because the four studies
312 that make up the bulk of the PCRS are older studies that give generally high field values and
313 have Q_{PI} scores ≤ 3 . There are a greater number of, and more recent, studies available for the
314 CNS, so the median values for it are more likely to represent reliable estimates of field strength
315 during a superchron. Thus, while the difference in average field strength between the PCRS and
316 the rest of the Paleozoic is likely to remain, further studies are needed to evaluate the average
317 strength of the field during the PCRS and if the potential differences between the average field
318 strength during the two superchrons is valid or due to data bias.

319 Both the PRE and POST bins gave consistently low values across all Q_{PI} filters (Fig. 5), including
320 AGE + ALT + MD, suggesting that MD behavior is not responsible for the appearance of these
321 periods of low field strength. Comparatively, the AGE + ALT + MD filter provided one of the
322 lowest median values for the PCRS (48 ZAm^2) suggesting MD behavior could instead be biasing
323 the PCRS values high (if taken from the lower temperature slope). The aging of TRM has also
324 been considered as a potential cause of long-term field strength underestimation. While this
325 process is less well understood than MD effects, experiments have shown that specimens
326 affected by aging tend to show more pronounced MD-like behavior (42) so they should also be
327 excluded by the MD criterion. Dividing the POST bin into 5 and 10 Ma intervals gives marginally
328 higher field values. This result suggests that the low average field strength of the POST bin and,
329 potentially, the distinct VDM distribution from the PRE bin (despite their similar median values),
330 could be due to under-sampling of the average field behavior during this time bracket. As
331 demonstrated by the highly skewed age of the bins and the low N values for 5 and 10 Ma sub-
332 bins, almost all the sites from this time period were emplaced over $\sim 800,000$ years (the Siberian
333 Trap sites) (43). The low N values mean that, while they may still be underestimating the long-
334 term strength of the field, they are starting to approach the strength of the field following
335 another superchron, the CNS (Late period; 36 ZAm^2 at $Q_{PI} \geq 3$) (16).

336 The strength of the field prior to the PCRS is the most reliable period herein due to the
337 consistency between the median values for the PRE bin, regardless of filtering and age binning.
338 The new paleointensity data presented in this study are also in close agreement with previous
339 studies, with median field estimates for the PRE bin ranging from 12.2-18.9 ZAm^2 , based on the

340 same filtering of Q_{PI} scores as used in Table S2. This range of median values for the PRE bin
341 remains the same if sites from the Kinghorn locality or the Strathmore locality, or both, are
342 excluded. The consistency between these new estimates and the legacy weak-field data, despite
343 differences in reliability between the new sites and those from the older studies, is encouraging
344 because it indicates the field strength during this time period is not being underestimated. At
345 $\sim 17 \text{ ZAm}^2$ ($Q_{PI} \geq 3$), the PRE bin is considerably lower than the average for the Phanerozoic, which
346 is estimated to be 42 ZAm^2 (13) to 50 ZAm^2 (6) and is probably the typical strength of the field
347 outside of the Bruhnes (44); i.e., the period since the last geomagnetic field reversal (the last
348 $\sim 770 \text{ ka}$). The PRE bin values are closest in strength to the Jurassic Hyperactivity period (JHAP;
349 $155\text{-}171 \text{ Ma}$; 26 ZAm^2 for all data points, 35 ZAm^2 at $Q_{PI} \geq 3$ (45), which had an average reversal
350 frequency of ~ 11 reversals/Myr. In comparison, the other periods of reversing field during the
351 Mesozoic (Early, Mid, Late) had median field strengths of $36\text{-}48 \text{ ZAm}^2$, with reversal frequencies
352 of $1\text{-}3$ reversals/Myr. The low average field strength may be partly due to recent studies (with
353 higher Q_{PI} scores) tending to sample periods of very high reversal frequency, like the JHAP. This
354 tendency is difficult to constrain because the magnetostratigraphic records before the PCRS are
355 generally too sparse to provide a reliable record of reversal frequency. A recent
356 magnetostratigraphic study from the Canning Basin (46) suggests reversal frequencies of a
357 minimum of $2\text{-}5$ reversals/Myr around the same time that the Viluy sites cooled ($\sim 360 \text{ Ma}$) (18).
358 In addition, an evaluation of reversal frequency $\sim 8\text{-}14$ Myrs before the PCRS (47) suggests
359 reversal frequencies of ~ 12 reversals/Myr, very similar to the high JHAP values, occurring ~ 5 Myr
360 after the Kinghorn lavas (this study) erupted. These studies produced notably weak site mean
361 values in the range of $4\text{-}27 \text{ ZAm}^2$. A greater non-dipole component to the field, relative to the
362 present day, is also indicated for some of the Siberian sites, such as the Minusa (17) and the
363 aforementioned Viluy Traps (18), although the new sites herein provide no indication of a
364 greater non-dipole component, despite comparable timing and field strength estimates.

365 **A Mid-Paleozoic Dipole Low and its implications for deep mantle variation.** Despite some
366 potential biases of the dataset, as discussed in the previous sections, the evaluation of the
367 period of low dipole moment leading up to the PCRS provided by this study suggests it is a
368 significant and distinct feature of the Paleozoic paleomagnetic record (see K-S tests in Table S2).

369 The proposed term for this feature, the mid-Paleozoic dipole low (MPDL), is based on its
370 similarities to the MDL. These similarities include the weak field (discussed in previous section)
371 and the ~80 Myr duration, from ~416-332 Ma. In both cases, however, gaps in the record make
372 it difficult to confirm if the period of low dipole moment extends back further in time (18). The
373 average field strength has been shown to vary throughout the MDL (16), and there is also
374 evidence for a difference (see Fig. 4) between the relatively strong early part of the MPDL in the
375 interval 390-416 Ma (median of 36 ZAm²), based on the Minusa (17) and Strathmore (this study)
376 sites, and the rest of the MPDL (317-390 Ma; median of 14 ZAm²). Unlike the MDL, the low field
377 strength is difficult to relate directly to reversal frequency due to the paucity of
378 magnetostratigraphic records at this time. Finally, there is a clear increase in field strength
379 around the onset of the PCRS, estimated to be 3-4 times the strength of the pre-PCRS field if the
380 average for the PCRS is reliable. We point out, however, that a gap in the dataset exists from
381 ~20 Myrs prior to the PCRS to an unknown point in the superchron (assumed to be within the
382 Carboniferous part of the PCRS; 299-315 Ma). The large age uncertainties associated with the
383 early PCRS sites prevent a clear determination of this transition.

384 The newly assessed paleointensity record provided in this study gives an improved indication of
385 patterns in Phanerozoic paleomagnetic field behavior across 10–100 Myr timescales (8) back to
386 ~415 Ma. The similarities observed herein between the MPDL and the MDL prior to their
387 respective superchrons provide more evidence for the proposed inverse relationship between
388 field strength and reversal frequency. There are insufficient site-mean paleointensity estimates
389 prior to (3 site means) and during (0 site means) the ORS to test this theory further back in time.
390 There have been several mechanisms proposed for this variation in field behavior relating to
391 mantle plumes (9, 48), subduction (2) and True Polar Wander (8). The extension of our reliable
392 Phanerozoic paleointensity record will assist future studies in linking these processes to
393 paleomagnetic evolution.

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401 **References**

- 402 1. D. G. van der Meer, D. J. J. van Hinsbergen, W. Spakman, Atlas of the underworld: Slab
403 remnants in the mantle, their sinking history, and a new outlook on lower mantle
404 viscosity. *Tectonophysics* **723**, 309–448 (2018).
- 405 2. M. W. Hounslow, M. Domeier, A. J. Biggin, Subduction flux modulates the geomagnetic
406 polarity reversal rate. *Tectonophysics* **742–743**, 34–49 (2018).
- 407 3. R. E. Ernst, K. L. Buchan, “Large mafic magmatic events through time and links to mantle-
408 plume heads” in *Mantle Plumes: Their Identification through Time*, R. E. Ernst, K. L.
409 Buchan, Eds. (Geological Society of America, 2001), pp. 483–575.
- 410 4. T. H. Torsvik, K. Burke, B. Steinberger, S. J. Webb, L. D. Ashwal, Diamonds sampled by
411 plumes from the core-mantle boundary. *Nature* **466**, 352–355 (2010).
- 412 5. J. A. Tarduno, R. D. Cottrell, W. J. Davis, F. Nimmo, R. K. Bono, A Hadean to Paleoproterozoic
413 geodynamo recorded by single zircon crystals. *Science* (80-.). **349**, 521–524 (2015).
- 414 6. A. J. Biggin, *et al.*, Palaeomagnetic field intensity variations suggest Mesoproterozoic
415 inner-core nucleation. *Nature* **526**, 245–248 (2015).
- 416 7. R. K. Bono, J. A. Tarduno, F. Nimmo, R. D. Cottrell, Young inner core inferred from
417 Ediacaran ultra-low geomagnetic field intensity. *Nat. Geosci.* **12**, 143–147 (2019).
- 418 8. A. J. Biggin, *et al.*, Possible links between long-term geomagnetic variations and whole-
419 mantle convection processes. *Nat. Geosci.* **5**, 526–533 (2012).
- 420 9. P. Olson, H. Amit, Mantle superplumes induce geomagnetic superchrons. *Front. Earth Sci.*
421 **3**, 1–11 (2015).
- 422 10. V. Pavlov, Y. Gallet, A third superchron during the Early Paleozoic. *Episodes* **28**, 78–84
423 (2005).
- 424 11. L. Tauxe, H. Staudigel, Strength of the geomagnetic field in the Cretaceous Normal
425 Superchron: New data from submarine basaltic glass of the Troodos Ophiolite.

- 426 *Geochemistry, Geophys. Geosystems* **5** (2004).
- 427 12. M. Prévot, M. E.-M. Derder, M. McWilliams, J. Thompson, Intensity of the Earth's
428 magnetic field: Evidence for a Mesozoic dipole low. *Earth Planet. Sci. Lett.* **97**, 129–139
429 (1990).
- 430 13. L. Tauxe, J. S. Gee, M. B. Steiner, H. Staudigel, Paleointensity results from the Jurassic:
431 New constraints from submarine basaltic glasses of ODP Site 801C. *Geochemistry,*
432 *Geophys. Geosystems* **14** (2013).
- 433 14. D. N. Thomas, A. J. Biggin, Does the mesozoic dipole low really exist? *Eos Trans. AGU* **84**,
434 97–104 (2003).
- 435 15. T. Anwar, L. Hawkins, V. A. Kravchinsky, A. J. Biggin, V. E. Pavlov, Microwave
436 paleointensities indicate a low paleomagnetic dipole moment at the Permo-Triassic
437 boundary. *Phys. Earth Planet. Inter.* **260**, 62–73 (2016).
- 438 16. E. V Kulakov, *et al.*, Analysis of an updated paleointensity database (QPI-PINT) for 65-200
439 Ma: Implications for the long-term history of dipole moment through the Mesozoic. *J.*
440 *Geophys. Res. Solid Earth* **124**, 9999–10022 (2019).
- 441 17. V. V. Shcherbakova, *et al.*, Was the Devonian geomagnetic field dipolar or multipolar?
442 Palaeointensity studies of Devonian igneous rocks from the Minusa Basin (Siberia) and
443 the Kola Peninsula dykes, Russia. *Geophys. J. Int.* **209**, 1265–1286 (2017).
- 444 18. L. M. A. Hawkins, *et al.*, An exceptionally weak Devonian geomagnetic field recorded by
445 the Viluy Traps, Siberia. *Earth Planet. Sci. Lett.* **506**, 134–145 (2019).
- 446 19. A. J. Biggin, G. H. M. A. Strik, C. G. Langereis, The intensity of the geomagnetic field in the
447 late-Archaeon: New measurements and an analysis of the updated IAGA palaeointensity
448 database. *Earth, Planets Sp.* **61**, 9–22 (2009).
- 449 20. M. F. Thirlwall, Geochronology of Late Caledonian magmatism in northern Britain. *J.*
450 *Geol. Soc. London.* **145**, 951–967 (1988).
- 451 21. J. T. Sallomy, J. D. A. Piper, Palaeomagnetic Studies in the British Caledonides - IV Lower
452 Devonian Lavas of the Strathmore Region. *Geophys. Journal, R. Astron. Soc.* **34**, 47–68
453 (1973).
- 454 22. M. Kono, Palaeomagnetism and palaeointensity studies of Scottish Devonian volcanic
455 rocks. *Geophys. J. Int.* **56**, 385–396 (1979).

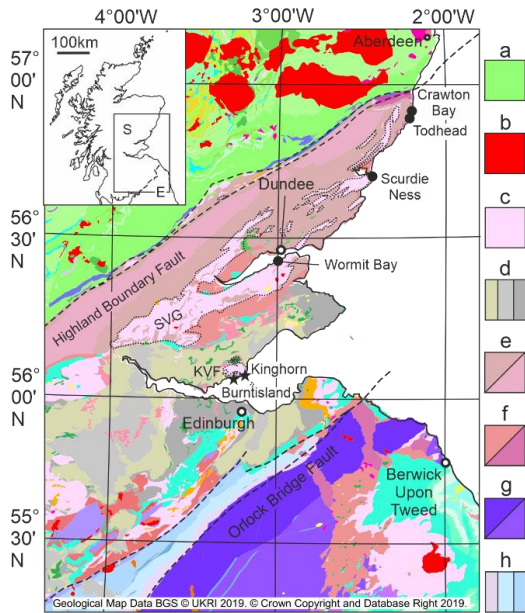
- 456 23. J. M. Grappone, A. J. Biggin, M. J. Hill, Solving the mystery of the 1960 Hawaiian lava
457 flow: implications for estimating Earth's magnetic field. *Geophys. J. Int.* **218**, 1796–1806
458 (2019).
- 459 24. T. H. Torsvik, Magnetic properties of the Lower Old Red Sandstone lavas in the Midland
460 Valley, Scotland; palaeomagnetic and tectonic considerations. *Phys. Earth Planet. Inter.*
461 **39**, 194–207 (1985).
- 462 25. S. Brindley, E. Spinner, Palynological assemblages from Lower Carboniferous deposits,
463 Burntisland district, Fife, Scotland. *Proc. Yorksh. Geol. Soc.* **47**, 215–231 (1989).
- 464 26. T. H. Torsvik, O. Lyse, G. Atterås, B. J. Bluck, Palaeozoic palaeomagnetic results from
465 Scotland and their bearing on the British apparent polar wander path. *Phys. Earth Planet.*
466 *Inter.* **55**, 93–105 (1989).
- 467 27. A. J. Biggin, G. A. Paterson, A new set of qualitative reliability criteria to aid inferences on
468 palaeomagnetic dipole moment variations through geological time. *Front. Earth Sci.* **2**, 1–
469 9 (2014).
- 470 28. Y. Yu, L. Tauxe, Testing the IZZI protocol of geomagnetic field intensity determination.
471 *Geochemistry, Geophys. Geosystems* **6** (2005).
- 472 29. R. S. Coe, S. Grommé, E. A. Mankinen, Geomagnetic paleointensities from radiocarbon-
473 dated lava flows on Hawaii and the question of the Pacific nondipole low. *J. Geophys.*
474 *Res.* **83**, 1740–1756 (1978).
- 475 30. A. J. Biggin, S. Badejo, A. R. Muxworthy, M. J. Dekkers, The effect of cooling rate on the
476 intensity of thermoremanent magnetization (TRM) acquired by assemblages of pseudo-
477 single domain, multidomain and interacting single-domain grains. *Geophys. J. Int.* **193**,
478 1239–1249 (2013).
- 479 31. N. Suttie, J. Shaw, M. J. Hill, Direct demonstration of microwave demagnetization of a
480 whole rock sample with minimal heating. *Earth Planet. Sci. Lett.* **292**, 357–362 (2010).
- 481 32. G. A. Paterson, L. Tauxe, A. J. Biggin, R. Shaar, L. C. Jonestrask, On improving the selection
482 of Thellier-type paleointensity data. *Geochemistry, Geophys. Geosystems* **15**, 1180–1192
483 (2014).
- 484 33. R. Van der Voo, The reliability of paleomagnetic data. *Tectonophysics* **184**, 1–9 (1990).
- 485 34. G. M. Solodovnikov, Paleostrength of the geomagnetic field in Middle-Late Carboniferous.

- 486 *Izv. - Phys. Solid Earth* **28**, 327–331 (1992).
- 487 35. G. M. Solodovnikov, Geomagnetic intensity in the Middle Carboniferous of Uzbekistan.
488 *Izv. - Russ. Acad. Sci. Phys. Solid Earth* **28**, 511–515 (1993).
- 489 36. W. E. Senanayake, M. W. McElhinny, A palaeointensity method for use with highly
490 oxidised basalts, and application to some Permian volcanics. *J. Geophys.* **52**, 85–96
491 (1983).
- 492 37. A. S. Bolshakov, G. M. Solodovnikov, Y. K. Vinogradov, Paleointensity of the Geomagnetic
493 Field in the Early Permian. *Izv. - Phys. Solid Earth*, 70–78 (1989).
- 494 38. G. M. Solodovnikov, Paleointensity of the geomagnetic field in the Lower Permian. *Izv. -*
495 *Phys. Solid Earth* **28**, 718–722 (1992).
- 496 39. J. G. Koenigsberger, Natural residual magnetism of eruptive rocks. *Terr. Magn. Atmos.*
497 *Electr.* **43**, 299 (1938).
- 498 40. J. G. Ogg, G. Ogg, F. M. Gradstein, *A concise geologic time scale 2016. [electronic book]*
499 (2016).
- 500 41. A. V. Smirnov, E. V. Kulakov, M. S. Foucher, K. E. Bristol, Intrinsic paleointensity bias and
501 the long-term history of the geodynamo. *Sci. Adv.* **3** (2017).
- 502 42. R. Shaar, L. Tauxe, Instability of thermoremanence and the problem of estimating the
503 ancient geomagnetic field strength from non-single-domain recorders. *Proc. Natl. Acad.*
504 *Sci. U. S. A.* **112**, 11187–11192 (2015).
- 505 43. S. D. Burgess, S. A. Bowring, High-precision geochronology confirms voluminous
506 magmatism before, during, and after Earth's most severe extinction. *Sci. Adv.* **1** (2015).
- 507 44. G. Cromwell, L. Tauxe, S. A. Halldörsson, New paleointensity results from rapidly cooled
508 Icelandic lavas: Implications for Arctic geomagnetic field strength. *J. Geophys. Res. Solid*
509 *Earth* **120**, 2913–2934 (2015).
- 510 45. M. Tominaga, W. W. Sager, M. A. Tivey, S.-M. Lee, Deep-tow magnetic anomaly study of
511 the Pacific Jurassic Quiet Zone and implications for the geomagnetic polarity reversal
512 timescale and geomagnetic field behavior. *J. Geophys. Res. Solid Earth* **113**, B07110
513 (2008).
- 514 46. J. Hansma, *et al.*, Late Devonian carbonate magnetostratigraphy from the Oscar and
515 Horse Spring Ranges, Lennard Shelf, Canning Basin, Western Australia. *Earth Planet. Sci.*

- 516 *Lett.* **409**, 232–242 (2015).
- 517 47. M. W. Hounslow, Geomagnetic reversal rates following Palaeozoic superchrons have a
518 fast restart mechanism. *Nat. Commun.* **7** (2016).
- 519 48. H. Amit, P. Olson, Lower mantle superplume growth excites geomagnetic reversals. *Earth*
520 *Planet. Sci. Lett.* **414**, 68–76 (2015).
- 521 49. G. A. Paterson, A simple test for the presence of multidomain behavior during
522 paleointensity experiments. *J. Geophys. Res. Solid Earth* **116** (2011).
- 523 50. E. Thébault, *et al.*, International Geomagnetic Reference Field: the 12th generation.
524 *Earth, Planets Sp.* **67**, 79 (2015).
- 525

526 **Figures and Tables**

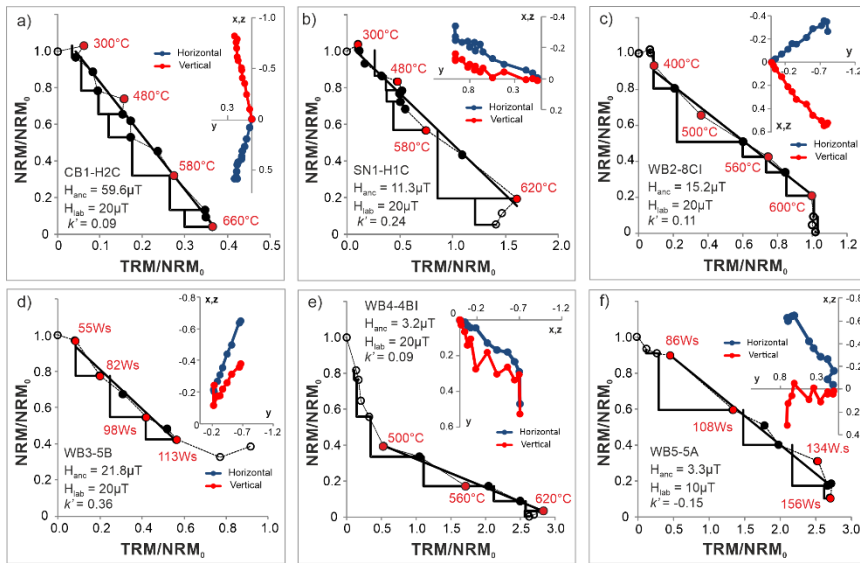
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530 **Figure 1.** Geological map showing the Strathmore (black circles) and Kinghorn (black stars)
531 localities sampled for paleodirection and paleointensity sites in this study. The geological units
532 come from the 1:50000 solid geology map from the British Geological Survey (BGS) ©UKRI
533 2019, accessed via Edina Digimap, and generalized descriptions are listed in the legend on the
534 right. Key cities are highlighted as white circles and key faults as dashed lines. The Strathmore
535 Group Volcanic (SGV) units and the Kinghorn Volcanic Formation (KVF) are highlighted with
536 dotted outlines. The location of the geological map is outlined in the inset map in the top left
537 corner of the Northern UK. The metamorphic and igneous geological units are a) Neoproterozoic
538 metamorphics, b) Silurian-Early Devonian felsic intrusions, c) Silurian–Devonian mafic extrusives.
539 The remaining units are clastic sedimentary rocks from d) Visean to Westphalian, e) Arbuthnott-
540 Garvock/Strathmore Groups, f) other Devonian, g) Llandovery-Wenlock and h) Caradoc to
541 Ashgill.

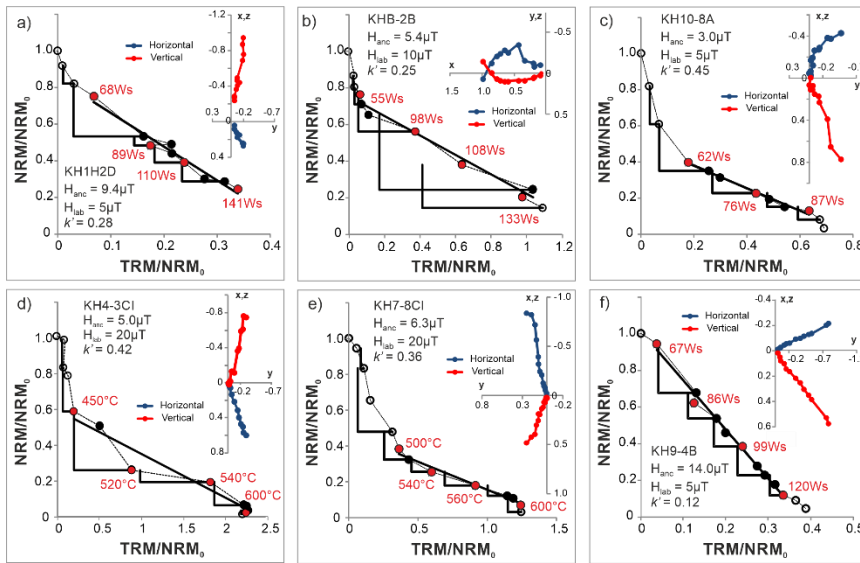


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543

544 **Figure 2.** Representative Arai plots from the six Strathmore sites a) CB1, b) SN1, c) WB(1/2), d)
 545 WB3, e) WB4 and f) WB5, illustrating the different Arai plot behaviors observed. All the
 546 measurements were done using microwave and thermal Thellier-type experiments using the IZZI
 547 protocol, with the thermal plots showing the highlighted temperature steps (°C) and the
 548 microwave plots showing the highlighted power steps (W.s, Watts per second). The thick black
 549 lines connecting measurement steps are the pTRM checks. The k' value is the curvature of the
 550 selected component on the Arai plot (49). The corresponding orthogonal plots are inset in the
 551 top right corners of the Arai plots. Plots (a-b) are examples from the sites where there are only
 552 thermal measurements as the components come from both the magnetite and hematite
 553 temperature ranges (Fig. S2a). The remaining plots (c-f) come from the magnetite (Fig. S2b) only
 554 samples and show microwave and thermal examples from sites that have similar hysteresis
 555 properties (Fig. S2e), WB2-WB3 (c-d) and WB4-WB5 (e-f).

556



557

558

559 **Figure 3.** Representative Arai plots from six of the Kinghorn sites a) KH1, b) KHB, c) KH10, d) KH4,

560 e) KH7 and f) KH8/9, illustrating the different Arai plot behaviors observed. All the

561 measurements were done using microwave and thermal Thellier-type experiments using the IZZI

562 protocol, with the thermal plots showing the highlighted temperature steps (°C) and the

563 microwave plots showing the highlighted power steps (W.s, Watts per second). The thick black

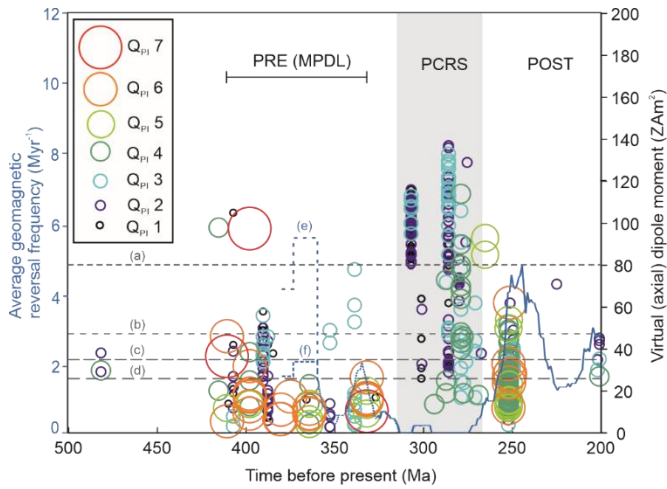
564 lines connecting measurement steps are the pTRM checks. The k' value is the curvature of the

565 selected component on the Arai plot (49). The corresponding orthogonal plots are inset in the

566 top right corners of the Arai plots. All the Arai plots represent (low-Ti titano-)magnetite apart

567 from a) KH1 and b) KHB.

568

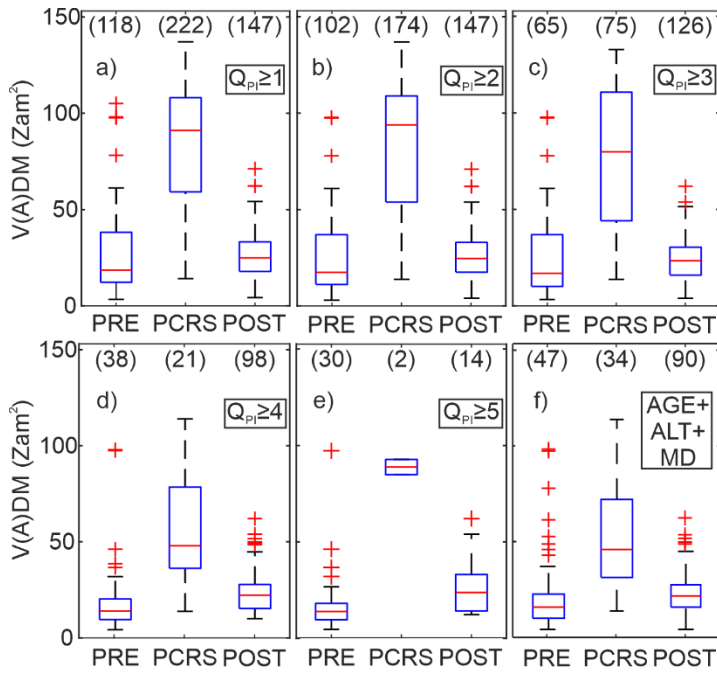


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570

571 **Figure 4.** The age distribution of all the V(A)DM values with $Q_{PI} > 0$ between 200–500 Ma. A
 572 summary of the Q_{PI} scores applied to each of the studies from this period are outlined in
 573 Datasets S3 and S4. The size and the color of the circles representing the V(A)DM values
 574 corresponds to the Q_{PI} scoring as outlined in the key. The PRE-, PCRS, and POST- section refer to
 575 the same age bins used for the Kolmogorov-Smirnov tests in Figure 5 and Table S2. The dashed
 576 lines represent a) the present day field strength (50), b) CNS at $Q_{PI} \geq 3$ (16), c) JHAP $Q_{PI} \geq 3$ (16),
 577 d) JHAP $Q_{PI} \geq 0$ (16), e) the maximum possible and f) the minimum possible reversal frequency
 578 from the Canning basin magnetostratigraphy (46).

579



580

581

582 **Figure 5.** Boxplots showing the V(A)DM distribution of the PRE, PCRS and POST bins (same as in
 583 Table S2). The boxplots are filtered based on the total Q_{pi} scores applied to the sites, between
 584 $Q_{pi} \geq 1$ to $Q_{pi} \geq 5$ (a-e) and for sites that pass the AGE, ALT and MD criteria (f). The numbers over
 585 the boxes display the number of sites in each of the age bins. On each box, the central mark is
 586 the median, the edges of the box are the 25th and 75th percentiles, the dashed lines extend to
 587 the most extreme data points not considered outliers, and outliers are plotted individually (+).

588

589 **Table 1.** Summary of paleointensity results and Q_{PI} scores for all the Strathmore and Kinghorn
 590 sites.

Site	Strathmore						Kinghorn							
	CB1	SN1	WB1/2	WB3	WB4	WB5	KH2	KH1	KHA	KHB	KH10	KH4	KH7	KH8/9
Paleointensity results:														
n_{INT}	9	11	22	12	15	13	20	28	12	21	17	10	12	23
N_{INT}	5	5	11	7	5	2	6	12	3	9	4	6	4	9
N_T	5	5	8	3	2	1	1	1	-	1	-	2	1	-
N_{MW}	-	-	3	4	3	1	5	11	3	8	4	4	3	9
mean (μT)	50.9	12.6	16.8	19.7	3.1	6.3	6.6	6.1	3.7	6.7	5.2	5.3	10.9	8.6
s.d. (μT)	15.9	3.6	5.8	6.4	0.3	4.3	3.0	1.6	0.5	2.5	1.1	0.7	4.3	4.0
s.d./ mean (%)	31	29	34	32	10	68	45	26	13	38	22	13	40	47
VDM (ZAm^2)	98.0	20.2	36.6	46.2	5.6	12.0	16.4	15.7	9.3	16.8	13.4	10.4	26.6	20.2
Q_{PI} scores:														
AGE	1	1	1	1	1	1	1	1	1	1	1	1	1	1
STAT	0	0	0	0	1	0	0	0	0	0	0	1	0	0
TRM	0	0	1	1	1	1	1	1	1	1	1	1	1	1
ALT	1	1	1	1	1	1	1	1	1	1	1	1	1	1
MD	1	1	1	1	0	0	1	1	1	1	1	1	1	1
ACN	1	1	1	1	1	1	1	1	1	1	1	1	1	1
TECH	0	0	1	1	1	1	1	1	0	1	0	1	1	0
LITH	0	0	1	0	0	0	0	0	0	0	0	0	0	0
MAG	1	1	1	1	1	1	1	1	1	1	1	1	1	1
Q_{PI}	5	5	8	7	7	6	7	7	6	7	6	8	7	6

591

592 n_{INT} : number of samples measured; N_{INT} : number of measurements that passed selection

593 criteria; N_T : number of accepted measurements from thermal IZZI; N_{MW} : number of

594 measurements from microwave IZZI; s.d.: standard deviation; VDM: virtual dipole moment. The

595 nine Q_{PI} are described in full in Biggin and Paterson (27), 1 is a pass and 0 is a fail to meet the
596 qualitative criteria, and Q_{PI} is total score of all these criteria. Site longitude, latitude and the
597 corresponding site directions are available in Table S2.

598 **Supplementary Information Text**

599 **Geological Background and Sampling**

600 **Strathmore Locality.** The older of the northern UK localities is part of the Lower “Old Red
601 Sandstone” suite of the Strathmore region, from the northern Midland Valley in Scotland (Fig.
602 1). This succession is represented by interbedded fluvial conglomerates and sandstones,
603 punctuated by calc-alkaline volcanism, and was deposited as part of three successive, graben
604 bound sedimentary basins: the (a) Stonehaven (Stonehaven Group), (b) Crawton (Dunnotar-
605 Crawton Group), and (c) Strathmore (Arbuthnott–Garvock and Strathmore Groups) basins (1).
606 The magmatism originates from the subducting Laurentian plate under the Iapetus suture (2)
607 and was greatest during the deposition of the Dunnotar-Crawton Group to the Lower
608 Arbuthnott–Garvock Group. The basaltic lava flows of the Crawton Volcanic Formation (VF;
609 Crawton Group) and the Montrose and Ochil Hill VF’s (Arbuthnott Group) provided the material
610 for this study and for the paleomagnetic pole (3).

611 Published Rb-Sr age dates exist from units correlated to the Crawton and Ochil Hills VFs, but
612 there is no flow-level age data available. The Lintrathen Tuff Member of the Crawton VF has
613 been correlated to the Glenbervie Porphyry member north of the highland boundary fault,
614 which has been dated using Rb-Sr (4) and recalculated to 415.5 ± 5.8 Ma (5). For the Ochil Hills
615 Volcanic Formation, Rb-Sr age dating was done on rhyolite from the base of the Wormit Bay
616 section (4) and recalculated to 410.8 ± 5.6 Ma (5). This age agrees closely with muscovite
617 assemblages from the sedimentary rocks of the Wormit Bay section to the Lockhavian (419.2
618 ± 3.2 Ma to 410.8 ± 2.8 (6)). No isotope age date exists for the Montrose Volcanic Formation,
619 although it has been correlated with ignimbrite situated 120 m stratigraphically above the top of
620 Crawton lavas (1) and lies below the Ochil lavas.

621 Paleodirections from the Strathmore region were originally measured by Sallomy and Piper (7)
622 to determine the Devonian pole for Britain, which was later superseded by the paleomagnetic
623 pole determined by Torsvik (3). Torsvik (3) found that the complex spread of the directions of
624 the original study could be largely explained by the demagnetization techniques used, with peak
625 alternating field (AF) values of ~ 30 -50 mT (considered low by modern standards), and local

626 tectonics. Torsvik (3) also reported two groups of directions. The Group 1 directions are similar
627 to those from Sallomy and Piper (7) and are considered primary, based on the nearly antiparallel
628 normal and reversed directions and passing fold tests and conglomerate tests on samples from
629 below flows 1 and 4 from the Crawton VF (Crawton Bay). Furthermore, samples taken from the
630 conglomerate near the contact with flow 1 (within 0-0.5 m) gave similar directions to the
631 overlying flow, suggesting they might pass a baked contact test. The Group 2 directions were
632 only observed at 3 sites (including two from Crawton Bay); however, despite coming from the
633 highest temperature component, these near-horizontal directions were interpreted to be due to
634 remagnetization (see *Strathmore results: Rock Magnetism* for further discussion).

635 The paleomagnetic pole, determined only from the Group 1 directions, has a quality (Q) factor
636 of seven (8, 9), which evaluates whether seven qualitative criteria that affect the reliability of
637 the paleomagnetic pole were addressed in the study. However, the data indicate that the
638 locality mean directions (Table III; Torsvik (3)) were calculated with all the specimen-level
639 measurements having the same weighting, which is inconsistent with modern methods for
640 determining the locality mean directions. To compare the locality mean directions from Torsvik
641 (3) to those from this study, first the site means were calculated from the Group 1 directions
642 (Table II; Torsvik (3)), before then averaging these to get the locality mean. There is a $<2.5^\circ$
643 difference between these revised normal, reverse and combined mean directions and the
644 original Torsvik mean directions (although the k values are no longer artificially inflated by the
645 high N values). The updated directions are included in this summary table (Table S1), along with
646 the site mean results from this study. A common true mean direction (CTMD) test was used (as
647 a reversal test) to check that the distributions of site mean directions were antiparallel came
648 back positive at classification C (Angle = 7° , Critical Angle = 17° (10–12)). In addition to the
649 paleodirectional work, there has also been a paleointensity study (13), which used some of the
650 same samples as the original paleodirectional study (7), including some of those considered
651 ‘transitional.’ However, questions regarding the reliability of the results (outlined in *Motivation*)
652 indicated that revisiting this locality was worthwhile.

653 Sampling was done with the intent to closely mimic the sites locations used for the
654 paleomagnetic pole (3), but precise matches were not always possible because of the limited

655 information published about the site locations. Of the locations sampled (Fig. 1), the Crawton
656 Bay and Todhead flows (Crawton VF) are the same as those used for the published pole, the
657 Scurdie Ness sites (Montrose VF) are from roughly the same location, while the Wormit Bay
658 (Ochil Hills VF) sites (14) are likely new sites, though they seem to be located close to the
659 Tayside site, based on the map from Torsvik (3). Samples were collected from five lava flows
660 from the Crawton VF (Crawton Bay and Todhead; CB1-CB4 and TH1), three from the Montrose
661 VF (Scurdie Ness; SN1-SN3) and four from the Ochil Hills VF (Wormit Bay; WB2-WB5) for a total
662 of twelve potential sites. All were visually distinct lava flows, apart from the Scurdie Ness lava
663 flows (the boundaries were not clearly visible in the field), where sites were taken substantial
664 distances (i.e. >100 m) from each other. Samples were also collected from a baked sediment
665 (WB1; underlying the WB2) for a baked contact test and from clasts within the conglomerate at
666 Todhead (TH2; underlying TH1) for a conglomerate test. Most of these samples were collected
667 as oriented drill cores, while the rest were collected as oriented hand samples, such as CB3-CB4,
668 SN2, WB1. Attempts were made at collecting sun compass readings, although weather
669 conditions largely prevented it; the few available readings were largely consistent with the
670 magnetic readings, so all the samples were oriented using magnetic readings for consistency.
671 Tilt corrections were applied from readings from sedimentary units (WB1 and a sandy layer
672 within TH2) or from clear flow boundaries.

673 **Kinghorn Locality.** The Kinghorn Volcanic Formation, part of the Visean volcanic sequences of
674 the Midland Valley, near Kinghorn, Scotland (Fig. 1), provided the second sample set. This
675 formation comprises a thick sequence (~485 m) of lava flows interspersed with minor thin
676 intercalations of sedimentary and volcanoclastic layers (15). These lava flows are predominantly
677 olivine basalts, which approach picrite compositions in some areas (16). Thirty flows have been
678 mapped in the region, dipping moderately (20-30 degrees) to the NE, and ranging in thickness
679 from 2.5-12 m. Magmatism in this region is thought to be the result of lithospheric extension
680 caused by the Variscan front to the south, which also led to rifting and the development of fault-
681 bound basins.

682 Stratigraphically, the age of the Kinghorn lavas are well constrained by correlations to the Sandy
683 Craig and Pathead Formations of the Strathclyde Group, whose strata are found above and

684 below, as well as interbedded with the Kinghorn Volcanic Formation (17). The micropore
685 assemblages in these correlated sedimentary rocks constrain the age of the Kinghorn lavas to
686 the Asbian-Brigantian (~337.5-326.4 Ma (17)). Isotopic age dating has been largely unsuccessful,
687 generally underestimating ages, possibly due to Argon loss (18). The closest isotopic age date to
688 the Asbian-Brigantian comes from a K-Ar age date of 338 ± 4 Ma from a sample collected
689 between Burntisland and Kinghorn (19) but, as the paleontology appears more robust, the mean
690 stratigraphic age is used (332.0 ± 5.6 Ma) in this study.

691 The Kinghorn locality has only one prior paleomagnetic study; its paleomagnetic pole, which was
692 determined by Torsvik et al. (20), has a Q factor of 6 (8, 21). This study has a lower Q score than
693 the Strathmore paleomagnetic pole because field tests were not performed. However, the
694 directions from the Carboniferous lavas agree with the high temperature component from the
695 underlying Old Red Sandstone (ORS) units, which is due to remagnetization of the ORS based on
696 the negative fold and conglomerate tests and it being a local feature (the primary direction is
697 found elsewhere in the ORS). As with the Strathmore locality, the aim was to sample as close to
698 the original sites as possible, but as with the Strathmore sites, identical sampling was difficult
699 due to the limited published information. Samples were collected from eleven distinct lava flows
700 from the Kinghorn locality; nine sites (KH1-KH10; KH8 and KH9 were later recognized to be a
701 single flow/site (KH8/9)) were collected on the first trip and two from a later trip (KHA-KHB). For
702 the Kinghorn (KH) sites, one site was collected on Burntisland (KH2), and the rest were collected
703 between Pettycur Harbour and Kinghorn Harbour (KH1, KHA-KHB, KH3-KH10). The majority of
704 samples were collected as oriented drill-cores, while the rest were collected as oriented hand
705 samples (KH1-KH2, KHA-KHB, and the odd sample from other sites) that were then drilled in the
706 laboratory. Most core samples were oriented using sun compass readings, with a few oriented
707 using magnetic readings if they were hand samples or sun readings could not be taken (the two
708 readings were generally consistent with each other). Tilt corrections were applied from readings
709 taken from clear flow boundaries and sedimentary interbeds.

710 **Methods**

711 **Paleodirections.** All the paleodirections were obtained using stepwise thermal demagnetization.
712 About half of these measurements (most of the samples from KH3-KH8/9, TH1-TH2 and WB2-
713 WB5) were made using full 2.5cm cores, heated in the 24-sample Magnetic Measurements
714 Thermal Demagnetizer (MMTD24) and measured using an Agico JR6 Spinner magnetometer.
715 The rest came from half-height 2.5cm cores that were heated in a Super Cooled MMTD(SC) and
716 measured on a RAPID 2G SQUID magnetometer (22). Samples were progressively heated to a
717 maximum temperature between 580-680 °C, when the remaining magnetic intensity of the
718 sample had decreased to <10 % of the natural remanent magnetization (NRM). The high
719 temperature magnetic components, interpreted as the Characteristic Remanent Magnetization
720 (ChRM), were selected based on the orthogonal plots and calculated using principle component
721 analysis (PCA (23)). If the mean angular dispersion (MAD°) and the angle between the anchored
722 and unanchored directions (α°) of the individual directions were $\geq 15^\circ$, these directions were not
723 included in the site mean analysis. Sites with dispersed, non-clustering paleodirections ($k < 15$)
724 were also excluded from further analysis (see Table S1). These site means are not intended to
725 supersede previous paleomagnetic studies (as many of the site N values are fairly low) but
726 rather to determine if the sites can reliably be used for paleointensity.

727 **Rock Magnetism.** To broadly determine these sites' magnetic mineralogy, rock magnetic
728 analysis was performed on representative specimens from each site. Hysteresis loops,
729 isothermal remanent magnetization (IRM) and back-field IRM curves, and thermomagnetic
730 (Curie) curves were run in air on crushed specimens on a Magnetic Measurements Variable Field
731 Translation Balance (MMVFTB). The thermomagnetic curves are used for categorizing the types
732 of magnetic minerals present in the samples based on the Curie temperatures (T_c) indicated,
733 whether the minerals appear to alter and, if so, over what temperature range. Determining the
734 magnetic mineralogy helps identify whether the sample is likely to carry a primary TRM.
735 However, the magnetization is often dominated by the larger of the population of magnetic
736 grains in the sample, which are unlikely to be the primary remanence carriers and so the
737 irreversibility of these curves is not a good indicator for how low-temperature alteration of the
738 NRM might affect the paleointensity experiments (see *Methods: Paleointensity (SI)* for details on

739 the alteration checks used instead). The hysteresis and IRM measurements are used as
740 indicators of the bulk grain size distribution.

741 Scanning electron microscopy (SEM) analysis was also performed on representative sites with
742 accepted paleointensity results using a Hitachi Table-top Microscope TM3000. Back-scattered
743 electron (BSE) images of representative thin sections were used to confirm the presence of
744 igneous textures, consistent with the samples carrying a primary thermoremanent
745 magnetization. This analysis included looking for textures that are consistent with the rock
746 magnetic results, such as exsolution lamellae, euhedral vs. skeletal structures, etc., and for any
747 cracking of magnetite grains that could be consistent with a volume reduction of the grains due
748 maghematization, which would result in a thermochemical remanent magnetization (TCRM)
749 that is not suitable for paleointensity experimentation. Energy-dispersive X-ray (EDX) analysis
750 was used, in conjunction with the BSE images, to assist in the identification of the magnetic
751 mineralogy.

752 **Paleointensity.** Sites were deemed suitable for paleointensity analysis if they a) passed the
753 paleodirectional selection criteria, b) had directions consistent with previous studies, and c)
754 produced rock magnetic results that were not inconsistent with a primary TRM. Two
755 paleointensity techniques were applied to specimens from these sites: thermal and microwave
756 Thellier-style experiments using the “IZZI” protocol (24), starting with a zero-field (Z) step.
757 Partial TRM (pTRM) checks (25) were used to check that the remanence carriers did not alter
758 due to heating during the experiment. Thermal experiments used oriented 2.5 cm diameter core
759 specimens (~1 cm in height), heated in air in the super cooled MMTDSC and then measured on
760 the RAPID 2G SQUID magnetometer. A field of 20 μ T was applied along the core’s Z-axis for the
761 in-field (I) steps. An alternating field (AF) step of 5 mT was applied before measuring each
762 heating step, as an ‘AF cleanse’, in order to reduce any potential non-single domain (SD) effects,
763 such as pTRM tails (26), on the paleointensity estimate (27, 28). Temperature steps were
764 determined from the behavior of sister specimens from thermal demagnetization and the rock
765 magnetic data. An initial temperature of 300 °C was selected to avoid the low temperature
766 overprints observed in the paleodirection data. Steps of 20–50 °C were used to a maximum of
767 600–660 °C, depending on the magnetic mineralogy.

768 Microwave experiments were performed on unoriented 5 mm diameter cores that were both
769 (de)magnetized and measured in air using the “Tristan” 14 GHz microwave SQUID
770 magnetometer system at the University of Liverpool (29). Each microwave specimen is run
771 individually, so a field of 3–20 μT , calibrated using other paleointensity estimates from each
772 respective site, was applied, at an angle of 45–90 ° to the NRM. The selection criteria used in
773 this study (provided in Datasets S1 and S2) are comparable to those used in recent studies of a
774 similar age from Siberia (30, 31), but with a stricter FRAC cut-off used for this new study. The
775 FRAC value (≥ 0.35) is still lower than recommended (32); however, these suggested criteria are
776 based on single component Arai plots. A larger FRAC value is used here because, on average, the
777 overprint on these samples represents a smaller part of the NRM than the other studies and it
778 reduces the misfit of the site mean data, while enough estimates are accepted to still be
779 statistically rigorous.

780 **Strathmore Results**

781 **Paleodirections.** Out of the 100 samples from the 12 sites that were thermally demagnetized
782 (this does not include WB1 and TH2, which were used for field tests), 90 passed the
783 paleodirection selection criteria. All the sites had $N \geq 3$ directions and representative orthogonal
784 plots (Fig. S1ai-di), with corresponding intensity curves (Fig. S1aii-dii), from accepted
785 measurements show that the NRM is made up of potentially three distinct components. The
786 data indicate a low temperature component between 0-100 °C (likely a viscous remanent
787 magnetization; VRM), a mid-temperature component, and the high-temperature component,
788 which is considered to be the ChRM, from around ~ 340 -480 °C and to 600-680 °C. The different
789 unblocking temperature ranges for the different sites reflect the different magnetic
790 mineralogies observed (see *Strathmore Results: Rock Magnetism* for further details). All the
791 measurement data is included in the MagIC database (<https://earthref.org/MagIC/17067>).

792 For these 12 sites, all but one of the site-mean directions (SN2) were reasonably well-clustered
793 with $k > 15$ (Table S1) and covered both the normal and reverse field states (Fig. S2a). Both the
794 normal and reverse directions from this study agree fairly well with the revised mean directions
795 from Torsvik (3), with angular differences of 7.8° and 10.7° respectively and substantial overlap

796 of their respective α_{95} circles. From the Crawton Bay flows, only CB1 has a reversed direction,
797 which suggests that it is the same as flow 4 from the Torsvik study (3), with flows 1, 2, and 3
798 being CB3, CB4, and CB2, respectively, based on their relative positions. Interestingly, sites CB3-
799 CB4 lack a shallow, high-temperature component (Fig. S1ai) consistent with the Group 2
800 directions from Torsvik (3), despite showing them in the original study; however, this
801 remagnetization was only recorded in some of their samples. A conglomerate test was
802 performed on a new site, the conglomerate overlying TH1 (site TH2), which showed that the
803 directions are uniform with weak support (Fig. S2b; $P(H_A|R) = 0.71$ (33)). The random nature of
804 these directions suggests that the conglomerate has not been remagnetized, but the degree of
805 confidence is limited by the low number of samples used ($n = 8$). A baked contact test was also
806 performed, using a CTMD test between the directions from the baked sediment at Wormit Bay
807 (WB1) and the overlying lava flow (WB2), which was positive at classification C (Angle = 7° ,
808 Critical Angle = 10° (10–12)).

809 **Rock Magnetism.** The magnetic mineralogy of the Strathmore sites can be divided into four
810 types; the first of which produced rock magnetic results consistent with 'Type B' magnetic
811 mineralogy from the Torsvik study (3). These sites (CB2-CB4, TH1 and SN3) gave
812 thermomagnetic curves that are largely reversible up to 300°C but then show a $\sim 30\text{-}40\%$
813 decrease in magnetization between $300\text{-}600^\circ\text{C}$ and provide Curie temperatures (T_c) of $\sim 470\text{-}$
814 540°C (Fig. S3a), apart from TH1, which produced a T_c of $\sim 620\text{-}640^\circ\text{C}$. These results are
815 consistent with the presence of (titano-)maghemite, which converts into (titano-)hematite when
816 heated above 300°C . The presence of (titano-)maghemite suggests that these sites underwent
817 chemical alteration after the lavas were emplaced (maghemite is not an original, or primary,
818 mineral in lavas) and may carry a TCRM (12). Also, as (titano-)maghemite is only metastable
819 (12), it is likely that some of it may have inverted into hematite at some point during the history
820 of the sites, which may explain the high-temperature, near-horizontal directional components
821 categorized as the Group 2 directions (3).

822 The other three magnetic mineralogy types produced thermomagnetic curves that generally
823 showed greater reversibility ($\leq 20\%$ decrease in magnetization when heated to $600\text{-}700^\circ\text{C}$) than
824 the Type B curves. The curves from sites CB1 and SN1 show two apparent values of T_c (Fig. S3b),

825 the first at ~580 °C, indicative of magnetite, and the second at ~680 °C, indicative of hematite.
826 This curve is consistent with the Type C curves from Torsvik (3). Hematite can be a primary
827 magnetic mineral in igneous rocks but is also produced by alteration (12). As the
828 paleodirectional component is from a temperature range that is consistent with both the
829 magnetite and hematite temperature ranges (340-680 °C; Fig. S1bi), both are likely primary
830 minerals. These sites plot in the upper part of the bulk domain stability plot in Fig. S3i, trending
831 above and perpendicularly away from the bulk domain stability (BDS) line (34), which is
832 consistent with the presence of hematite. There is no SEM analysis available for samples of this
833 rock magnetic type.

834 The thermomagnetic curve (Fig. S3c) and hysteresis properties (Fig S3i) for the baked sediment
835 (WB1) are similar to those from CB1 and SN1 because WB1 has a T_c consistent with hematite
836 (~680 °C) but not magnetite. The hematite component likely carries a primary TRM since its
837 direction is consistent with the overlying lava. The last mineralogy type (sites WB2–WB5) had
838 associated thermomagnetic curves with values of T_c from 520-580 °C (Fig. S3d), which
839 corresponds to the (titano-)magnetite Type A curves from Torsvik (3). SEM analysis of the
840 samples showed grains of low-Ti (titano-)magnetite next to ilmenite, both on the order of tens
841 of μm in length, with no finer exsolution structure apparent (Fig. S3g). All these samples plot
842 near but just above the 'BDS trend' line (Fig. S3i).

843 **Kinghorn Results**

844 **Paleodirections.** Out of the 57 samples, from 9 sites sampled, that were thermally
845 demagnetized for paleodirections (KH8/9 were combined into a single site and KHA-KHB were
846 not included in the paleodirectional analysis), 47 passed the paleodirection selection criteria.
847 Three sites failed to produce consistent directions (KH3, KH5-KH6), because of both individual
848 directions failing to pass the selection criteria and the site directions not clustering, which may
849 relate to the magnetic mineralogy of the samples (see the *Kinghorn; Rock Magnetism* section).
850 Representative orthogonal plots (Fig. S1ei-fi), and corresponding intensity curves (Fig. S1eii-fii),
851 from accepted measurements indicate a possible VRM between 0-100°C and a small mid-
852 temperature overprint (the intensity of these overprints combined is <25% of the NRM), before

853 the ChRM, which generally starts at 300–480 °C and finishes between 540–600 °C, depending on
854 the titanium content of the sites (see *Kinghorn; Rock Magnetism* section for further details).

855 Of the accepted paleodirectional sites from this study, the majority of the lavas sampled here
856 are normal in polarity (four of the six; Table S1 and Fig. S2c), unlike Torsvik (20), which sampled
857 more reversed polarity lavas to the north of those herein. The mean normal polarity direction
858 from this study agrees well with that from Torsvik, with $<3^\circ$ angular difference. The mean
859 reverse direction from this study is very poorly constrained ($\alpha_{95} >100^\circ$) because the
860 paleodirectional sites (KH1–KH2) also only have $N = 2$, while the other two reversed sites (KHA–
861 KHB) were not included in the paleodirectional analysis because they only have a single oriented
862 hand sample each. However, the directions from both the paleodirectional sites (especially KH2)
863 and from oriented paleointensity measurements (which both lie within the α_{95} circle of KH1) are
864 also sufficiently similar, given the low N (n) values, to the mean reversed direction from Torsvik
865 (20) to be considered for paleointensity analysis (Fig. S3c; there are no individual site directions
866 from Torsvik (20) to compare against, which might have agreed better). All the measurement
867 data are included in the MagIC database (<https://earthref.org/MagIC/17067>).

868 **Rock Magnetism.** The majority of the Kinghorn sites that gave acceptable paleodirections (KH2,
869 KH4 and KH7–KH8/9) produced thermomagnetic curves similar to the Type A curves from Torsvik
870 (3), with a T_C range from ~ 540 – 580 °C, indicative of low-Ti (titano-)magnetite (Fig. S3e). Sites
871 KH1, KH10 and KHA–KHB, as well as all the sites that failed to give acceptable paleodirections,
872 gave substantially lower T_C 's in the range of 370–480 °C (the lowest values came from the sites
873 that failed to pass paleodirection selection criteria; KH3 and KH5–KH6). These curves showed an
874 increase in magnetization when heated above ~ 500 °C (Fig. S3f), which is indicative of the
875 formation of magnetite from exsolution of the titanium-rich titanomagnetite upon heating.
876 SEM analysis indicates the presence of very coarse exsolution structures between the
877 titanomagnetite and ilmenite phases (Fig. S3h), with the titanomagnetite forming around the
878 edge of large, skeletal grains of ilmenite (hundreds of μm in length). All the sites plot close to
879 the BDS trendline, with moderate M_{rs}/M_s and H_{cr}/H_c values, apart from KHB, which has the
880 lowest T_C (indicating moderate-Ti titanomagnetite), which may have affected its BDS value (Fig.
881 S3j).

882 **Paleointensity Reliability Assessment**

883 **Q_{PI} criteria.** While selection criteria, such as those used in this study (see Datasets S1 and S2),
884 are designed to check the reliability of measurements at the specimen level, there are also a
885 number of factors that can affect the reliability of paleointensity estimates that are not
886 evaluated by these criteria. To assess these site-level factors, Biggin and Paterson (35) outlined a
887 set of nine quality criteria for paleointensity (Q_{PI}). These criteria emulate the Quality (Q) criteria
888 outlined by Van der Voo (9), which are used to assess the reliability of paleomagnetic poles, and
889 are scored the same way (a pass = 1, fail = 0 and the Q and Q_{PI} scores are the total of all the
890 criteria scores). An overview of how these nine criteria have been assessed in this study is
891 outlined here:

- 892 1. AGE: This criterion requires that the site has a reliable age constraint from either
893 radiometric or stratigraphic age dating, with errors that are less than 10% of the age.
894 Corresponding paleodirections are not required to pass this criterion because this can
895 discriminate against materials that are difficult to orient but otherwise have good age
896 constraints (i.e. some ocean drill cores, single crystals, etc.). However, where they are
897 available, the directions should not unquestionably deviate from what is expected from
898 the age of the sites (transitional sites and sites where the apparent polar wander paths
899 are not well constrained would still pass). This includes whether normal and reverse
900 directions are observed i.e. sites age dated to a superchron should not include reversals
901 and vice versa.
- 902 2. STAT: This criterion assesses the consistency between paleointensity estimates from the
903 same site as there is no geomagnetic reason for discrepancies between samples that
904 acquired their TRM in the same field. To pass STAT, the site mean must be calculated
905 from at least 5 specimens ($N \geq 5$) and the standard deviation (std.dev.) of these
906 measurements, normalized by the site mean (std.dev./mean), must be $\leq 25\%$. However,
907 because the N values for paleointensity studies tend to be statistically small, there are
908 high levels of uncertainty associated with the std.dev. and mean and so the upper 95%
909 confidence interval should be used ($\delta B_N\%$ (36)) to ensure a std.dev. $\leq 25\%$ (for $N = 5$ at
910 25% maximum, $\delta B_N\% = 15.97\%$).

- 911 3. TRM: This criterion recognizes that, while the effects of chemical alteration may be
912 identifiable from rock magnetic techniques, microscopy is well suited to identify clear
913 occurrences of low-temperature alteration that may otherwise be missed. In the
914 majority of cases, the study must present evidence for primary igneous textures that are
915 consistent with TRM acquisition and no apparent alteration to pass this criterion. There
916 are some exceptions to this, such as when the remanent magnetization is unlikely to
917 have been acquired by any means other than thermally (i.e. where the magnetization is
918 held by inclusions in silicates (37)) or there is other compelling evidence for thermal
919 acquisition (i.e. a baked contact with consistent paleodirections as the igneous body).
- 920 4. ALT: Alteration of the magnetic carriers can present a major issue for the reliability of
921 paleointensity estimates as they are calculated assuming that the magnetic recorders
922 have remained unchanged between ChRM and (lab induced) TRM acquisition and
923 almost all conventional paleointensity techniques include at least one heating step (38).
924 To pass the ALT criterion, measurements should include some form of alteration check.
925 This is generally pTRM checks for Thellier-type experiments (25), while other methods,
926 such as the Shaw method (39), pass by comparing ARM acquisition before and after
927 heating. Ideally, some selection criteria relating to these checks should be applied to the
928 measurements; however, the type of criteria should not be taken into consideration,
929 unless it has shown to be completely ineffective at detecting alteration. The main
930 exception to this is if the paleointensity experiment includes no heating step (i.e.
931 calibrated pseudo-Thellier (40)), it will automatically pass ALT.
- 932 5. MD: It is well established that the behavior of multi-domain (MD) grains (as well as
933 interacting grains, which have been shown to behave in a similar way (41)) can affect
934 the paleointensity estimate. For any paleointensity experiments using pTRMs (i.e.
935 Thellier-type experiments), MD behavior is expressed as tails, causing the Arai plots to
936 “sag” or “curve” below that of the linear behavior expected from SD grains (26, 42). Site
937 values from Thellier-type experiments can pass this criterion if the experiment includes
938 some protocol for checking for MD effects during the experiment. This includes pTRM
939 tail checks (43), checking for curvature of the selected component (e.g. $|k/$ or $|k'|$ (44))
940 or using the IZZI protocol (24), where pTRM tails are expressed as zig-zagging in the Arai

941 and/or orthogonal plots (the site will only pass if the selected components lack a clear
942 indication of zig-zagging). The site may also pass if the selected component is from
943 linear, single component Arai plots, which can be determined from selection criteria e.g.
944 high f or FRAC values, which indicate that the Arai plot is near single-component, and
945 low β values that suggest the component is linear (see Standard Paleointensity
946 Definitions (SPD) for details (45)). Other techniques, such as the Shaw (39) and Wilson
947 (46, 47) methods, automatically pass the MD criterion as they are considered to be
948 domain-state independent. Bulk hysteresis measurements that are consistent with the
949 “pseudo-single domain” grains (48) are insufficient to pass the MD criterion as domain
950 state effects can still be an issue (49)

951 6. ACN: This criterion is divided into three sub-criteria, reflecting three issues that can bias
952 paleointensity estimates but are not as commonly observed with paleomagnetic sites as
953 ALT and MD and are correctable. While the sub-criteria are combined into a single Q_{PI}
954 criterion, so that their weighting is kept proportional in comparison to the other criteria,
955 all the sub-criteria need to pass for the site to pass ACN (if one fails, the site fails ACN).
956 These sub-criteria are:

957 a. Anisotropy of TRM: This is not commonly an issue for paleointensity (i.e.
958 igneous) targets, but it has been shown to be an issue for some localities (50).
959 Experiments using an applied field parallel to the direction of the ChRM would
960 pass this sub-criterion automatically. Otherwise, checks that show that
961 anisotropy of TRM is not affecting the paleointensity estimates, would be
962 accepted. An acceptable check is a low γ value (51), i.e. the angle
963 between the applied field and last pTRM acquired for the selected component.
964 Anisotropy of ARM experiments could also be used to show that anisotropy of
965 TRM is not an issue for the site (52). The site can still pass if anisotropy of TRM is
966 an issue if an appropriate correction has been applied (50, 53).

967 b. Cooling rate effects: It has been shown both theoretically (54, 55) and
968 experimentally (56) that the TRM acquired by SD grains is greater when slow-
969 cooled, rather than fast-cooled. This can bias paleointensity estimates because
970 there are normally several orders of magnitude difference between the cooling

971 rates in nature and the lab. Studies suggest that this should only really be an
972 issue when the remanence carriers are SD in size (57–59) or are within a large
973 intrusion that has cooled over hundreds of thousands of years. In general, “fast-
974 cooled” igneous sites (e.g. lavas and smaller dykes and sills) pass this sub-
975 criterion as the remanence carriers are normally larger than SD grains
976 (preferably this would be supported by rock magnetic data such as bulk
977 hysteresis measurements though). It may also be possible to check for cooling
978 rate effects if the site includes measurements from different techniques with
979 substantially different cooling rates (the faster the cooling rate in the lab, the
980 higher the paleointensity estimate should be). For materials that carry SD grains
981 or slow-cooled sites (e.g. plutons and larger sills), a cooling rate correction
982 needs to be applied (54, 55) for the site to pass this sub-criterion. It is also worth
983 noting that there has been recent work that suggests that cooling rates may
984 affect more paleointensity measurements with non-SD carriers more than
985 previously considered (60). However, further work is needed to quantify how
986 this may have affected the paleomagnetic record, especially as there is
987 potentially no way to correct past studies as there was no observed correlation
988 between the cooling rate dependence and commonly measured magnetic
989 properties (i.e. bulk hysteresis).

990 c. Non-linear dependence of TRM: An assumption when calculating paleointensity
991 estimates is that the relationship between field strength and TRM acquisition is
992 linear at weak (Earth-like) field values. However, for some magnetic grain
993 assemblages, non-linear TRM acquisition can occur at sufficiently low field
994 values that it affects either NRM or lab TRM acquisition (61). If the applied field
995 during a paleointensity experiment is the same or close (within 1.5x of) the field
996 the NRM was acquired in, then the effects of non-linear TRM dependence
997 should be negligible and the site passes this sub-criterion. Sites with low field
998 values for both the applied field and the paleointensity estimates ($\leq 60 \mu\text{T}$) will
999 also pass as they should both be within the linear portion of TRM acquisition

1000 based on experimentation (61). The site may also pass if measurements using
1001 substantially different applied field values give consistent paleointensity results.

1002 7. TECH: A multitude of paleointensity techniques have been developed over time, and
1003 having accepted results from different methods can provide greater confidence in the
1004 results. To be considered different techniques, they should use significantly different
1005 styles of unblocking. For example, Thellier-type experiments, all using pTRMs but
1006 different protocols, would be considered a single technique but experiments using total
1007 TRM acquisition (46, 47) or unblocking grains using the microwave system (62) or AF
1008 (39) would be considered separate techniques. The results do not have to be consistent
1009 because this requirement is covered by the STAT criterion; however, variation between
1010 the techniques may illustrate issues relating to some of the other (sub-)criteria.
1011 Consistency between results from more than one technique can also allow sites to pass
1012 criteria that using one technique alone would not pass (e.g. sites with consistent Thellier
1013 and Wilson can pass MD because Wilson is domain state independent and ALT if there
1014 are pTRM checks for the Thellier experiment).

1015 8. LITH: The LITH criterion recognizes that where materials with different lithologies (i.e.
1016 materials with substantially different magneto-mineralogy and domain state
1017 assemblages) have acquired a TRM in the same field, this can improve the reliability of
1018 the site. This scenario is typically only met if there is a baked contact that can be
1019 sampled, or an igneous unit has chilled margins and a coarse-grained interior with
1020 substantially different mineralogies. There should also be information provided for the
1021 magnetic mineralogy of the two units, including evidence for differences in the
1022 unblocking temperatures.

1023 9. MAG: This criterion is designed to encourage the publication of the raw measurement
1024 data somewhere accessible (such as the MagIC database, which the criterion is named
1025 after, but it can be published elsewhere and pass this criterion). Whatever the
1026 format/depository chosen, the raw measurement data should allow others to be able to
1027 analyze the data for themselves. However, while this criterion has been assessed for all
1028 the 200-500 Ma sites in the PINT15 database (Dataset S3), this criterion has not been
1029 included in the total Q_{PI} scores used for long-term field strength analysis (Figures 4 and 5

1030 and Table S2) because the publication of the raw measurements is a very recent
1031 phenomenon and so this criterion would probably add unnecessary bias to the most
1032 recently published studies.

1033

1034 **Strathmore Q_{PI} Scoring.** The individual site Q_{PI} scores for the Strathmore sites can be found in
1035 Table 1, as well as Datasets S3 and S4 (the latter also includes comments on the scoring of the
1036 sites). A detailed explanation for the scoring of these sites is included below (numbering of the
1037 criteria is the same as in *Paleointensity Reliability Assessment: Q_{PI} criteria* section):

- 1038 1. AGE: The oldest (Crawton Bay/Todhead) and youngest (Wormit Bay) sites pass AGE
1039 because the sections they come from have, or have been correlated with, members that
1040 have Rb-Sr age dates (415.5 ± 5.8 Ma and 410.8 ± 5.6 Ma respectively (5)) consistent
1041 with the stratigraphic constraints (1) and errors <10% of the age. The Scurdie Ness sites
1042 pass because they are stratigraphically between the two other locations and the
1043 difference in age between the two sites (including errors) is <10% of the mean age.
- 1044 2. STAT: None of the Strathmore sites pass STAT. Most of the sites had $N_{INT} \geq 5$ but the
1045 normalized standard deviations were too high, possibly because the paleointensity
1046 estimates were mostly low (<20 μT), resulting in higher normalized std.dev. However,
1047 CB1 had a relatively high site mean value (50.9 μT) and still failed due to scattered
1048 paleointensity estimates.
- 1049 3. TRM: The Wormit Bay sites (WB1/2-WB5) all pass this criterion as the lavas all have
1050 similar mineralogy and there are representative SEM images for these sites (Fig. S3h).
1051 While WB1/2 represents two lithologies (WB1 is the baked sandstone and WB2 is the
1052 lava), and there are no SEM images for WB1, the consistency in paleodirections
1053 between the two sites indicates that the baked sandstone was overprinted by the lava.
1054 The other sites (CB1, SN1) did not pass this criterion as there is no representative
1055 microscopy for them; however, the results are still included because the consistency
1056 between the paleodirections and paleointensity estimates from the magnetite and
1057 hematite temperature ranges support that their remanence is likely a primary TRM.

1058 4. ALT: All the Strathmore sites passed as all the measurements included pTRM checks and
1059 appropriate DRAT/CDRAT selection criteria (see Dataset S1).

1060 5. MD: As the IZZI protocol was used for all the experiments and none of the sites, apart
1061 from WB4-WB5, showed any signs of zig-zagging, all the sites (other than WB4-WB5)
1062 should pass the MD criterion. Selection criteria to evaluate the curvature of the selected
1063 component was also used ($|k'| \leq 0.48$ for all accepted measurements (63)). As for sites
1064 WB4-WB5, while they fail this criterion, their low results may not be the result of MD
1065 behavior. While the zig-zagging seen in the orthogonal plot obscures it, it is clear from
1066 the thermal demagnetization experiments that there are at least two directional
1067 components in these sites (see Fig. S1di), and the selected temperature range for the
1068 paleointensity estimate (Fig. 2e) corresponds with the temperature range for the ChRM.
1069 Also, the curvature of the selected Arai plot components is still low ($|k'| < 0.31$). Finally,
1070 a simple estimation of the maximum bias of the paleointensity estimates can be
1071 calculated by selecting the first and last (before alteration) points of the Arai plot to see
1072 how much it deviates from the theoretical linear, single component plot (although this is
1073 unrealistic as the site has clearly been overprinted). The “single component estimates
1074 range from ~ 0.95 - 2.50 x the accepted estimates and, while this seems like a large
1075 deviation, in absolute terms it is $< 9 \mu\text{T}$, so the maximum paleointensity estimates would
1076 still be $< 10 \mu\text{T}$ for WB4 and $< 20 \mu\text{T}$ for WB5. This does not change whether the sites pass
1077 the MD criterion but is important for the overall consideration of the strength of the
1078 field at the time of the Strathmore lavas.

1079 6. ACN: All the Strathmore sites passed ACN because all the sub-criteria were fulfilled:

1080 a. Anisotropy of TRM: All sites were considered to be unaffected by anisotropy as
1081 the locality had a mean γ value of 2.8° (range 0.3 - 6.7°).

1082 b. Cooling rate effects: All the Wormit Bay sites, other than WB1/2, pass because
1083 they represent relatively fast-cooled lava flows, whose hysteresis properties are
1084 consistent with non-SD grains (Fig. S3i, j). The thermal and microwave results
1085 from WB4 and WB5 are relatively consistent, despite the difference in cooling
1086 rate between the two techniques. There appears to be some difference
1087 between the two techniques for sites WB2 and WB3, but the microwave

1088 estimates were generally lower than the thermal estimates, in contradiction
1089 with a cooling rate effects (55).The hematite-bearing sites (CB1, SN1, WB1/2)
1090 are more likely to be affected by cooling rate because hematite grains tend to
1091 be SD in size (64). The measurements from WB1 and WB2 are consistent
1092 enough that the site has a negligible cooling rate effect. For the other two sites,
1093 reselecting estimates, where the hematite component was included, to just
1094 below magnetite's T_C produced estimates within ~15% of those selected from
1095 both components and were, in contradiction to the expected effects of cooling
1096 rate, generally higher so cooling rate is also unlikely to have affected these sites.

1097 c. Non-linear dependence of TRM: Both the paleointensity estimates and applied
1098 field values were low enough ($\leq 60 \mu\text{T}$) that the non-linearity of TRM should
1099 have no significant effect (61).

1100 7. TECH: The Wormit Bay sites (W1/2-WB5) all passed as they included both microwave
1101 and thermal experiments. The other sites (CB1 and SN1) do not pass as the microwave
1102 system was not able to demagnetize the hematite-bearing sites (this is also true of WB1
1103 but there are microwave experiments for WB2).

1104 8. LITH: Only site WB1/2 passed this criterion as it is the only site with results from two
1105 distinct mineralogy's, with distinct unblocking spectra.

1106 9. MAG: All the sites pass as the raw measurement files are included in the MagIC
1107 database (<https://earthref.org/MagIC/17067>).

1108

1109 **Kinghorn Q_{PI} Scoring.** The individual site Q_{PI} scores for the Kinghorn sites can be found in Table
1110 1, as well as Datasets S3 and S4 (comments on scoring can also be found in Dataset 4). A
1111 detailed explanation for the scoring of these sites is included below (numbering is the same as in
1112 the previous two sections, *Q_{PI} criteria and Strathmore Q_{PI} scoring*):

1113 1. AGE: The Kinghorn isotopic age dating is generally fairly poor with suspected Argon loss
1114 (18). However, the misopore ages from correlated sedimentary sections are well
1115 constrained (~337.5-326.4 Ma (17)) and are in close agreement with at least one
1116 radiometric age (338 ± 4 Ma (19)), and the difference between the ages is within 10% of
1117 the mean age.

- 1118 2. STAT: Only one site passed the STAT criterion (KH4); three of the sites had $N_{INT} \leq 5$ and
1119 the rest had std.dev values that were too high to pass STAT. Like with the Strathmore
1120 sites, this may be due to the low site mean values ($\leq 11 \mu T$), meaning that while the
1121 absolute std.dev values were low (1.6-4.0 μT for the Kinghorn sites with $N_{INT} \geq 5$ that
1122 failed STAT), the normalized std.dev. was relatively high (26-40%).
- 1123 3. TRM: Based on the rock magnetic measurements, all the Kinghorn lavas are (titano-
1124)magnetite-bearing, with varying amounts of titanium and no evidence of
1125 maghematization, so the SEM microscopy (Fig. S3h) should be representative of all the
1126 lavas, so all sites pass this criterion.
- 1127 4. ALT: All the Kinghorn sites passed as all the measurements included pTRM checks and
1128 appropriate DRAT/CDRAT selection criteria (see Dataset S2).
- 1129 5. MD: The IZZI protocol was used for all the experiments and no significant zig-zagging
1130 was noted with the Arai or orthogonal plots. Along with the checks for curvature of the
1131 selected component ($k' \leq 0.48$ for all accepted measurements (63)) there is enough
1132 evidence for all the sites to pass the MD criterion. The two-slope Arai plots also are also
1133 likely due to overprinting as the orthogonal plots from both the Arai plots (Fig. 3b-e) and
1134 those from thermal demagnetization experiments (Fig. S1ei, fi) show the selected
1135 component corresponds to the ChRM.
- 1136 6. ACN: All the Kinghorn sites passed ACN because all the sub-criteria were fulfilled:
- 1137 a. Anisotropy of TRM: Unlike for the Strathmore sites, in some cases, γ exceeded
1138 6° for some accepted Kinghorn site mean specimens. This is a cut-off value
1139 suggested for isotropic samples based on the work of Paterson (51). AARM
1140 experiments were performed on representative samples, using the same
1141 methods as described in SPD (45), to investigate this issue further. These gave p
1142 values ranging from 1.05-1.19. However, calculating the anisotropic corrections
1143 (c values), using the method described in the SPD (45) for some of the most
1144 anisotropic sites (KH1, KH4 and KH7; p values of 1.19, 1.16 and 1.12
1145 respectively) gave specimen c values of 0.91-1.10, with a median c value of 0.98,
1146 far less than the anisotropy of the specimens themselves. The c values also
1147 varied substantially, probably because the c value is dependent on the direction

1148 of the applied field, which was varied for the individual microwave
1149 measurements. The site mean paleointensities calculated from the corrected
1150 specimen measurements showed negligible offsets from the uncorrected means
1151 (corrected site means were 6.0 μT , 5.3 μT and 10.8 μT respectively KH1, KH4
1152 and KH7; the uncorrected site means are in Dataset S2). Due to the limited
1153 sample available after the paleointensity experiments, only representative
1154 specimens were analyzed. It would not have been possible to correct all the
1155 measurements consistently; however, these AARM measurements show that
1156 the effect of anisotropy on the paleointensity estimates were negligible, even
1157 for the most anisotropic of the Kinghorn sites.

1158 a. Cooling rate effects: All the sites are fast-cooled lavas, and the bulk hysteresis
1159 properties show that all the sites plot between the SD and MD regions of the
1160 (Fig S2j). Thus, it is unlikely that cooling rate would affect the paleointensity
1161 estimates. In addition, the Kinghorn sites produced a slightly lower average
1162 paleointensity estimates for the microwave experiments (6.9 μT vs 7.7 μT) than
1163 the thermal ones, in contradiction to what is expected from cooling rate effects.

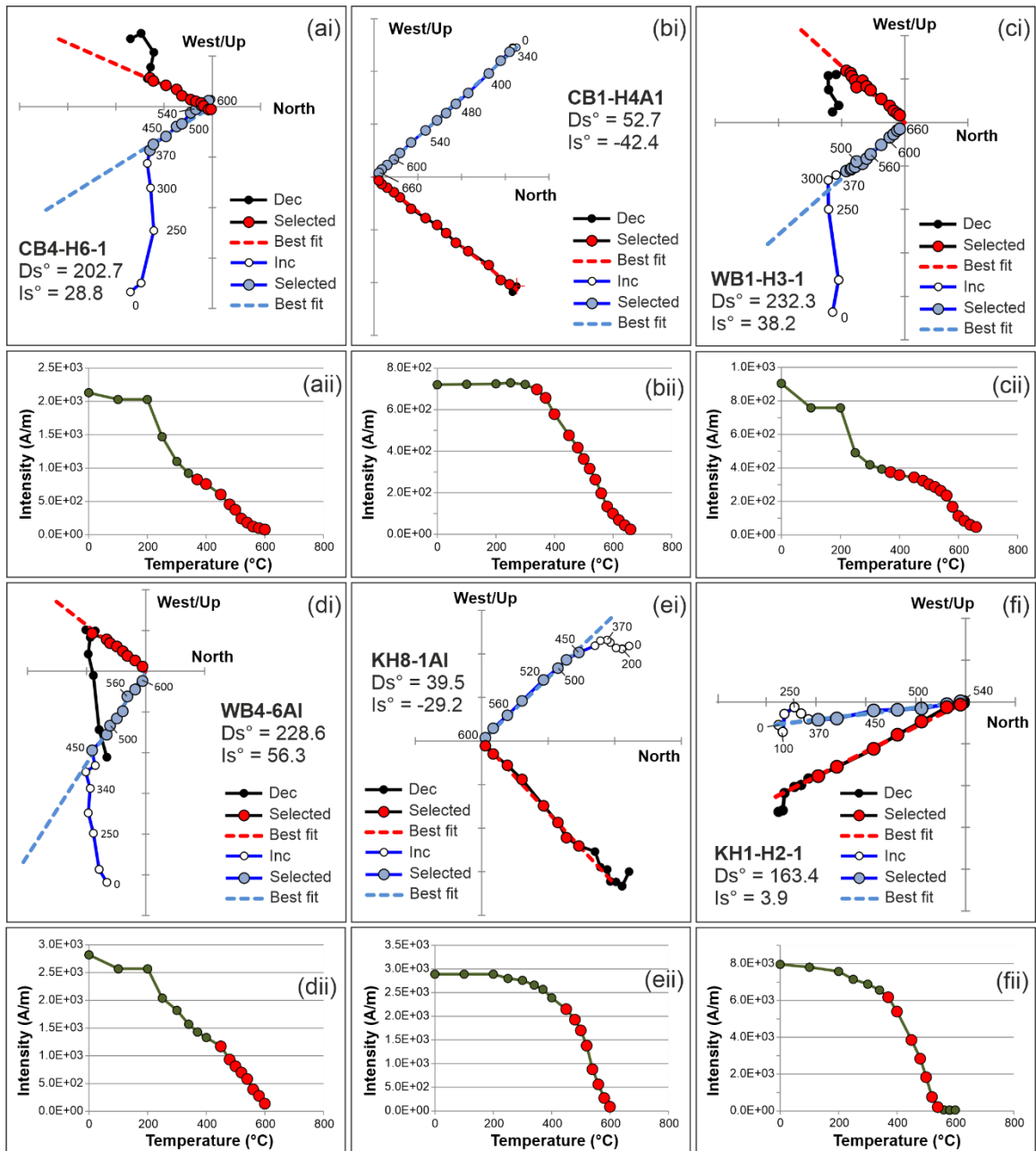
1164 b. Non-linearity of TRM: Both the paleointensity estimates and applied field values
1165 were sufficiently low ($\leq 20 \mu\text{T}$) that the non-linearity of TRM should have had no
1166 significant effect (61).

1167 7. TECH: Five out of the eight sites passed the TECH criterion as they included both
1168 successful microwave and thermal Thellier-type experiments, while the three sites that
1169 failed only had successful microwave experiments.

1170 8. LITH: None of the site pass LITH as they are all single lithology sites (lava flows).

1171 9. MAG: All the sites pass as the raw measurement files are included in the MagIC
1172 database (<https://earthref.org/MagIC/17067>).

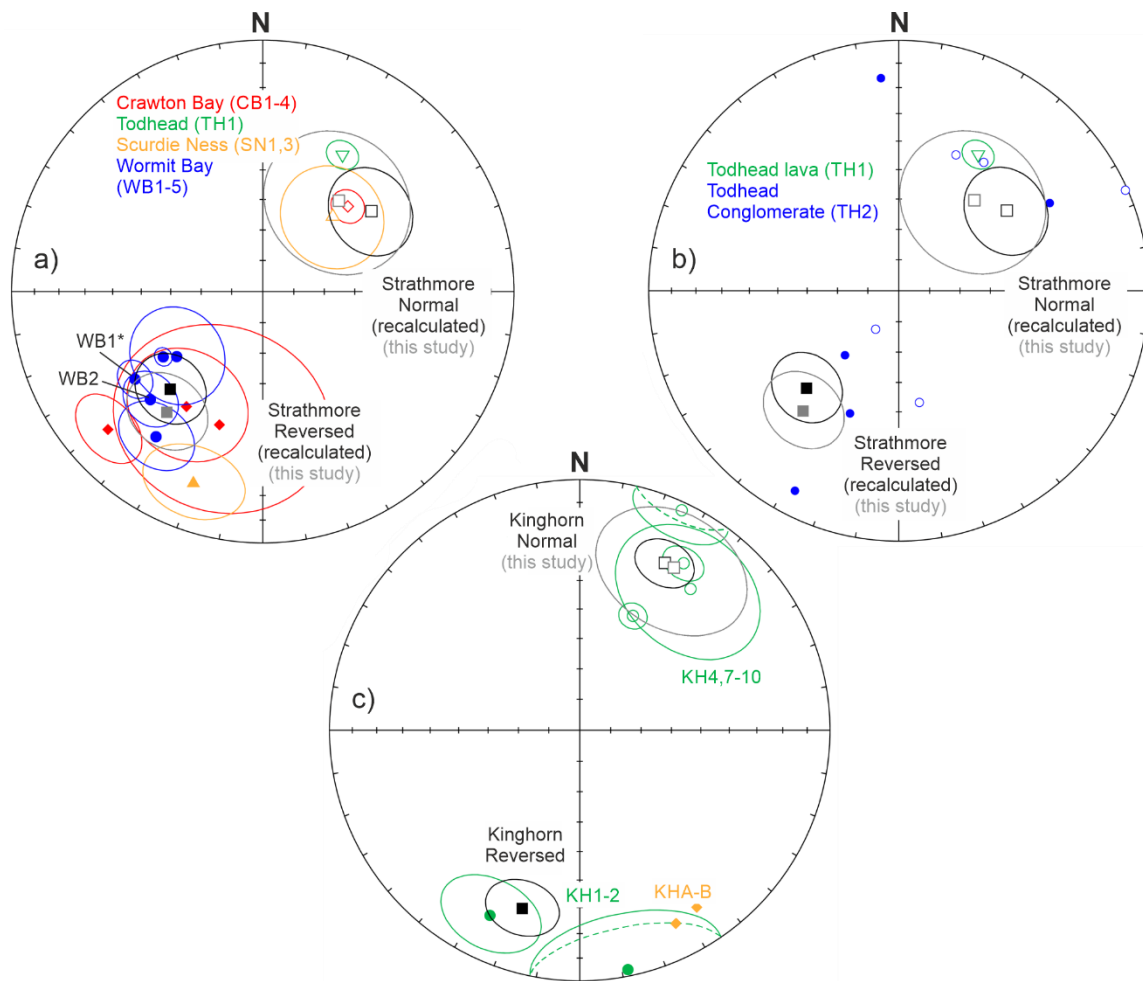
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1177 **Fig. S1.** Examples of i) orthogonal plots and the ii) corresponding intensity curves from the
 1178 thermal demagnetization experiments from Strathmore (a-d) and Kinghorn (e-f) localities. Each
 1179 of the examples represents a different magnetic mineralogy (outlined in the *Rock Magnetism*
 1180 sections from the *Strathmore Results and Kinghorn Results*) and corresponds to the example

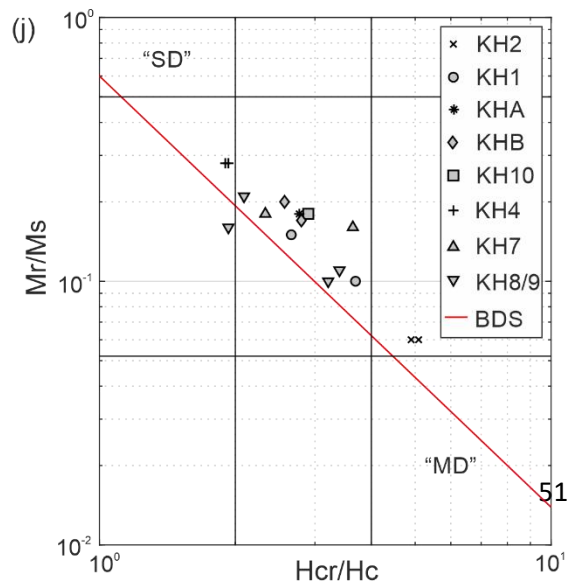
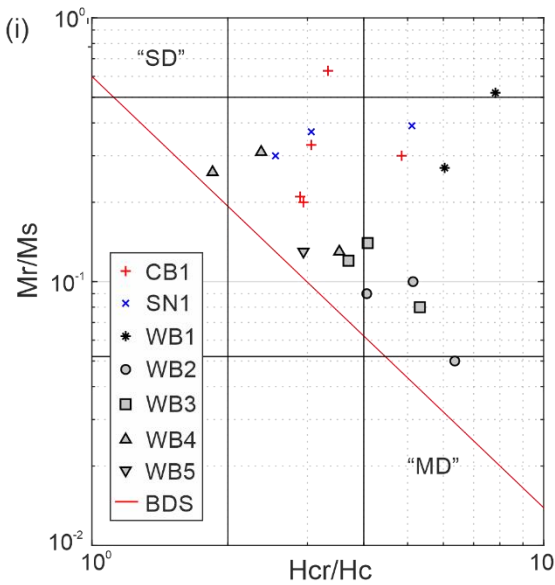
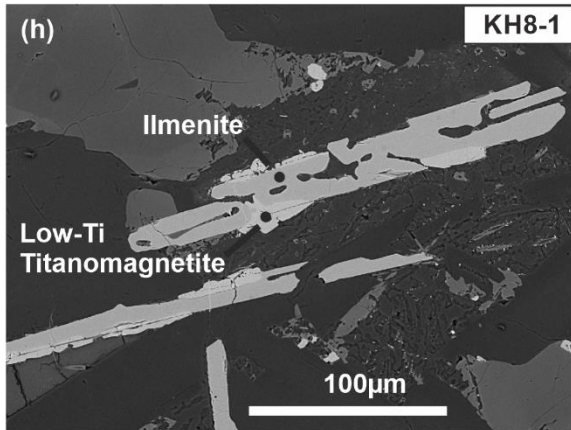
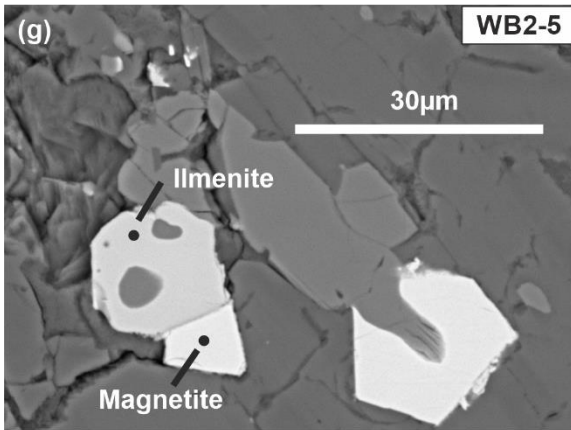
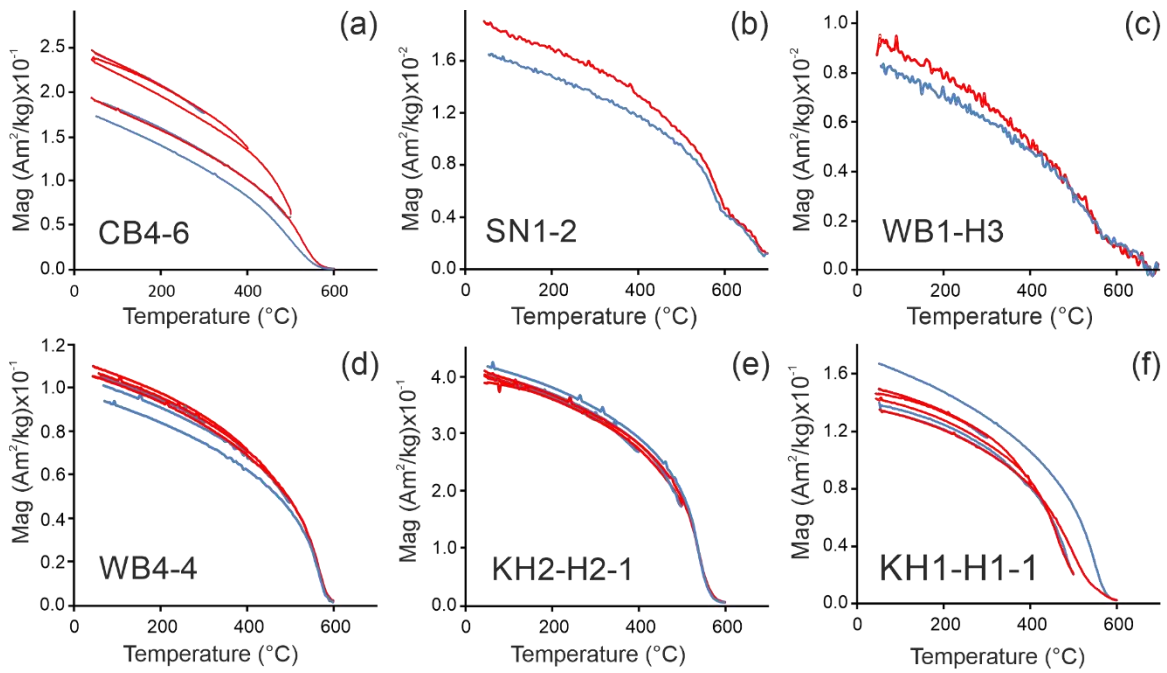
1181 thermomagnetic curves with the same letter in Fig. S3. D_s° and I_s° are the declination and
1182 inclination (in stratigraphic co-ordinates) determined from the selected component using PCA
1183 analysis (23). The numbers next to the selected points on the inclination curve are the
1184 temperature step (in °C). The maximum demagnetization temperatures of the samples are
1185 sometimes a bit higher than expected from the corresponding thermomagnetic curves (Fig. S3),
1186 which is likely due to thermochemical alteration.



1187
1188

1189 **Fig. S2.** Stereographic projections of (a) the accepted site mean directions from the Strathmore
 1190 region, (b) the individual directions used for the conglomerate test for TH2 and (c) the accepted
 1191 site mean directions from the Kinghorn lava flows. The site means are shown with their α_{95}
 1192 circle. For the Strathmore site means (a), the new site directions from the different localities are
 1193 represented by different colors and symbols; red diamonds (Crawton Bay), green inverted
 1194 triangle (Todhead), orange triangles (Scurdie Ness) and blue circles (Wormit Bay). For the
 1195 conglomerate test (b), the site mean for the TH1 lavas are as in (a) with blue circles for the
 1196 individual conglomerate directions (TH2). For the Kinghorn sites (c), the new site directions are
 1197 represented by green circles and the orange diamonds represent the average direction from the
 1198 selected component from oriented paleointensity experiments for sites without separate

1199 paleointensity analysis (measurements from Grappone (65)). The black squares on both plots
1200 represent the locality mean directions used for the paleomagnetic poles from Torsvik (3) for
1201 Strathmore, recalculated to fit with modern paleodirectional analysis (see *Geological*
1202 *Background and Sampling; Strathmore Locality* for more details), and Torsvik et al. (20) for the
1203 Kinghorn. The gray squares are the recalculated normal and reverse directions from this study
1204 (the Kinghorn Reversed direction is not included as it is very poorly constrained due to a lack of
1205 sites).



1207 **Fig. S3.** Representative thermomagnetic curves (a-f), representative BSE SEM images (g-h) and
1208 (i-j) hysteresis data plotted on an H_{cr}/H_c vs. M_r/M_s or “Day” plot from the sites that provided
1209 accepted paleointensity estimates. The thermomagnetic curves show the magnetization of the
1210 sample upon heating (red curve) and cooling (blue curve). The first four thermomagnetic curves
1211 (a-d), the left SEM image (g), and the left “Day” plot (i) are representative of the Strathmore
1212 sites while the rest are representative of the Kinghorn sites. The hematite bearing sites show
1213 thermomagnetic curves (b-c) with a single heating or cooling curve to 700 °C, while the other
1214 curves (a, d-f) were heated in steps of 100 °C, from 300-600 °C to show the temperature range
1215 over which the curves are (ir)reversible, indicating the presence of alteration. For the SEM
1216 images (g-h), the key magnetic minerals are labelled, having been identified from the EDX and
1217 igneous textures. The H_{cr}/H_c vs. M_r/M_s or “Day” plots (i-j) show the hysteresis data from all the
1218 samples that provided accepted paleointensity measurements; the red bulk domain state (BDS)
1219 line comes from Paterson et al. (34) and the black lines represent the grain size boundaries from
1220 Dunlop (66) with the “single-domain” (“SD”) and “multi-domain” (“MD”) boxes highlighted.

1221

1222 **Table S1.** Summary of the accepted site and locality mean directional data from the new
 1223 Strathmore and Kinghorn sites, as well as the recalculated Strathmore sites from Torsvik (3).

1224

Site	S _{LAT} (°)	S _{LONG} (°)	N (n)	D _G (°)	I _G (°)	D _S (°)	I _S (°)	k	α ₉₅	P _{LAT}
<u>Strathmore Locality:</u>										
<i>Crawton Bay</i>										
CB1	56.9080	-2.2022	11	33.5	62.3	45.0	-50.4	71.2	5.4	-43.6
CB2	56.9077	-2.2001	7	225.6	31.1	228.3	18.4	33.7	10.5	16.8
CB3	56.9081	-2.1980	4	202.2	54.2	213.6	44.1	24.5	18.9	34.8
CB4	56.9083	-2.1992	3	184.9	50.5	198.1	43.5	15.0	33.0	31.2
<i>Todhead</i>										
TH1	56.8848	-2.2161	10	22.6	51.1	30.3	-37.2	99.7	4.9	-31.8
<i>Scurdie Ness</i>										
SN1	56.6910	-2.4424	10	44.1	16.0	50.6	-64.6	25.5	9.8	-8.2
SN2	56.6952	-2.4415	4	16.6	12.9	357.6	-55.7	2.1	88.0	-6.5
SN3	56.6810	-2.4507	8	199.0	26.7	200.0	19.8	16.7	13.9	-14.1
<i>Wormit Bay</i>										
WB1	56.4241	-2.9849	8	250.9	21.2	235.5	38.1	90.1	5.9	11.0
WB2	56.4244	-2.9847	8	244.8	24.8	227.1	38.3	37.1	9.2	13.0
WB3	56.4237	-2.9857	8	230.8	22.3	215.8	29.1	24.5	11.4	11.6
WB4	56.4207	-2.9911	9	261.3	36.2	234.8	55.7	18.4	12.3	20.1
WB5	56.4207	-2.9911	8	261.2	31.2	239.5	51.5	143.3	4.6	16.8
<i>New locality means</i>										
Normal	-	-	3	-	-	40.0	-51.1	28.5	23.5	-31.7
Reverse	-	-	8	-	-	218.6	38.4	20.0	12.7	21.6
Combined	-	-	11	-	-	218.9	41.9	21.2	10.2	24.2
<i>Redone Torsvik locality means</i>										
Normal	-	-	6	-	-	53.5	-45.3	24.3	13.9	-26.8
Reverse	-	-	9	-	-	223.5	45.3	21.3	11.4	26.8
Combined	-	-	15	-	-	227.5	45.4	23.0	8.1	26.9
<u>Kinghorn locality:</u>										
<i>Kinghorn sites</i>										
KH2	56.0586	-3.2233	2	209.6	5.8	205.9	18.8	318.6	14.0	2.9

KH1	56.0625	-3.1793	2	169.7	3.2	168.6	2.8	118.7	23.1	1.6
KHA*	56.0621	-3.1782	1 (4)	166.2	-3.5	166.0	4.8	<u>7.0</u>	37.5	2.4
KHB*	56.0622	-3.1777	1 (2)	153.6	14.9	146.5	16.3	1412.3	6.7	8.3
KH10	56.0626	-3.1749	8	24.3	8.3	24.9	-3.7	21.9	12.1	-1.8
KH3	56.0633	-3.1739	3	32.3	6.8	32.3	-7.3	<u>2.2</u>	<u>121.1</u>	3.4
KH4	56.0639	-3.1734	8	38.9	34.1	25.0	-48.3	162.2	4.4	-18.7
KH5	56.0643	-3.1733	4	257	50.4	253.6	76.3	<u>1.5</u>	<u>128.2</u>	31.1
KH6	56.0645	-3.1730	2	39.5	43.6	48.4	28.7	<u>5</u>	-	25.5
KH7	56.0669	-3.1735	7	36.2	10.0	31.9	-22.6	103.4	6.0	-5.1
KH8/9	56.0680	-3.1738	11	43.3	14.9	38.1	-29.1	49.6	6.5	-15.5
New locality means										
Normal	-	-	4	-	-	30.1	-26.0	18.0	22.3	-13.7
Reverse	-	-	2	-	-	186.7	11.4	8.4	102.1	5.7
Combined	-	-	6	-	-	22.0	-21.6	11.9	20.2	-11.2

1225

1226 S_{LAT} and S_{LONG} : site longitude and latitude; N (n): Number of samples (number of specimens); D_G

1227 and I_G : (anchored) Declination and Inclination in geographic co-ordinates; D_S and I_S : (anchored)

1228 Declination and Inclination in stratigraphic co-ordinates i.e. tectonically corrected; k and α_{95} : the

1229 precision parameter and 95% confidence limit from Fisher statistics (12); P_{LAT} : Paleolatitude. The

1230 sites that have been greyed out were not included in the new locality means. For the majority of

1231 sites, this was because the site mean directions were not clustered ($k < 15$; k values with these

1232 values are underlined, as well as high α_{95} values $> 45^\circ$). *Sites with an asterisk next to them were

1233 not included in the locality means as the site directions came from the paleointensity

1234 measurements from a single hand sample (N=1). *WB1 was included for a baked contact test*

1235 *(passed) so it should have the same direction as WB2 and was, therefore, not included in the*

1236 *locality mean.*

1237

1238 **Table S2.** Statistics for the PRE, PCRS and POST bins and the results of Kolmogorov-Smirnov test
 1239 for different levels of Q_{PI} filtering and age binning of the datasets.

Statistics	Q_{PI} filter						AGE+ ALT+ MD	Bin- ned (Q_{PI} ≥ 1 ; 10Ma)	Bin- ned (Q_{PI} ≥ 3 ; 10Ma)	Bin- ned (Q_{PI} ≥ 1 ; 5Ma)	Bin- ned (Q_{PI} ≥ 3 ; 5Ma)
	$Q_{PI} \geq 1$	$Q_{PI} \geq 2$	$Q_{PI} \geq 3$	$Q_{PI} \geq 4$	$Q_{PI} \geq 5$						
Bin statistics:											
<i>N</i> (*)	118	102	65	38	30	47	10*	9*	14*	12*	
PRE bin statistics	Age mean \pm std.dev (skewness)	378.0 ± 25.2 (-0.64)	376.7 ± 25.3 (-0.55)	371.3 ± 28.4 (-0.10)	376.4 ± 28.5 (-0.39)	374.5 ± 27.0 (-0.18)	361.8 ± 27.6 (0.56)	-	-	-	-
	VDM median/ IQR (V%)	18.4/ 26.1 (142 %)	17.5/ 25.8 (147 %)	17.0/ 26.8 (158 %)	14.1/ 10.7 (76%)	13.6/ 8.6 (63%)	15.7/ 13.2 (84%)	17.0/ 9.8 (58%)	17.9/ 30.1 (168 %)	16.5/ 7.0 (42%)	17.9/ 29.5 (165 %)
PCRS bin statistics	<i>N</i> (*)	222	174	75	21	2	34	5*	4*	7*	6*
	Age mean \pm std.dev (skewness)	290.7 ± 10.6 (-0.14)	288.7 ± 10.4 (0.12)	283.1 ± 8.4 (0.63)	279.0 ± 6.7 (-0.16)	265.4	278.9 ± 5.3 (0.52)	-	-	-	-
POST bin statistics	VDM median/ IQR (V%)	91.0/ 49.0 (54%)	94.0/ 55.0 (59%)	80.0/ 66.8 (84%)	48.0/ 42.3 (88%)	89.0/ 8.0 (9%)	45.5/ 41.0 (90%)	56.5/ 56.0 (99%)	79.8/ 62.5 (78%)	39.0/ 54.0 (138 %)	65.5/ 76.5 (117 %)
	Age mean \pm std.dev (skewness)	249.2 ± 11.0 (-4.02)	249.2 ± 11.0 (-4.02)	250.6 ± 7.8 (-6.14)	251.6 ± 5.2 (-9.73)	251.9 ± 0.3 (-1.26)	251.5 ± 5.4 (-9.31)	-	-	-	-
POST bin statistics	<i>N</i> (*)	147	147	126	98	14	90	3*	2*	3*	3*
	VDM median/ IQR (V%)	24.6/ 15.5 (63%)	24.6/ 15.5 (63%)	23.6/ 14.5 (61%)	22.3/ 12.5 (56%)	23.6/ 19.0 (81%)	21.4/ 12.6 (59%)	35.3/ 35.5 (101 %)	26.7/ 6.7 (25%)	31.9/ 27.6 (87%)	28.6/ 10.0 (35%)

Kolmogorov-Smirnov tests:

p values		PRE vs. PCRS	1.3E-35	2.5E-29	1.4E-14	2.3E-06	2.2E-02	9.0E-07	2.2E-03	2.2E-02	6.0E-03	1.9E-01
		PCRS vs. POST	4.6E-45	2.5E-40	2.2E-23	1.2E-07	1.9E-02	1.4E-09	2.2E-01	4.7E-02	5.0E-01	2.0E-01
PRE vs. POST		PRE vs. POST	4.8E-04	3.5E-04	5.4E-04	9.7E-04	5.4E-02	1.8E-03	8.5E-02	4.0E-01	3.2E-02	4.4E-01
		5%	1	1	1	1	1	1	1	1	1	0
PRE vs. PCRS	5%	1	1	1	1	0	1	1	0	1	0	
	1%	1	1	1	1	0	1	1	0	1	0	
PCR S vs. POST	5%	1	1	1	1	1	1	0	1	0	0	
	1%	1	1	1	1	0	1	0	0	0	0	
PRE vs. POST	5%	1	1	1	1	0	1	0	0	1	0	
	1%	1	1	1	1	0	1	0	0	0	0	

1240

1241 Q_{PI} : the total Qualitative Paleointensity score based on the criteria defined by Biggin and Paterson
1242 (35); AGE/ALT/MD: three of the nine Q_{PI} criteria and sites must pass all three (each must score a
1243 1) to be included in this dataset; PRE, PCRS and POST: primary age bins the VDM values from the
1244 updated PINT15 data (i.e. Dataset S3) were sorted into (from 315-416 Ma, 267-315 Ma and 200-
1245 267 Ma respectively); 10 Ma and 5 Ma, the 10 and 5 million year sub-bins the VDM values, were
1246 further divided into and the median values taken; N(*): the number of site mean estimates in
1247 the bin (*the number of sub-bins that included sites from the PINT15 database within that age
1248 range); std. dev.: standard deviation; VDM: virtual dipole moment; IQR: interquartile range; V%:
1249 IQR/median; p: asymptotic p-values from the Kolmogorov-Smirnov (k-s) test result; 5% and 1%:
1250 the significance levels the k-s tests were performed at; 1, 0 or NaN; pass, fail or insufficient data,
1251 in one or both of the bins, for the k-s test.

1252 **Dataset S1 (separate file).** Summary of all the paleointensity results from the Strathmore
1253 locality. The selection criteria cut-offs applied to the paleointensity criteria are listed at the top;
1254 n: number of selected measurement steps; FRAC: fraction of the NRM used for the best fit of the
1255 selected component; β : scatter around the best fit of the selected component; q: 'quality
1256 factor'; $|\vec{k}'|$: curvature of the selected component; MAD_{ANC}: Maximum Angular Deviation (MAD)
1257 of the (anchored) best fit direction; α : angular difference between the best-fit anchored and
1258 free-floating directions; DRAT: maximum absolute difference from a pTRM check; CDRAT:
1259 cumulative DRAT. For further details on the selection criteria, the reader is referred to the
1260 Standard Paleointensity Definitions (SPD) (45). The two methods used for determining the
1261 paleointensity estimates were microwave (MW) and thermal with an AF cleanse (Th-AF)
1262 Thellier-type measurements using the IZZI protocol with pTRM checks. The maximum and
1263 minimum steps are shown as power integrals listed in Watt seconds (Ws) for the microwave
1264 measurements and temperature (T; °C) for the thermal experiments. Experimental results that
1265 do not pass the selection criteria are in gray, and the selection criteria that failed are underlined.
1266 Pi: Paleointensity estimate; H_{lab}: Applied lab field; s.d: Standard deviation.

1267 **Dataset S2 (separate file).** Summary of all the paleointensity results from the Kinghorn locality.
1268 The selection criteria cut-offs applied to the paleointensity criteria are listed at the top; n:
1269 number of selected measurement steps; FRAC: fraction of the NRM used for the best fit of the
1270 selected component; β : scatter around the best fit of the selected component; q: 'quality
1271 factor'; $|\vec{k}'|$: curvature of the selected component; MAD_{ANC}: Maximum Angular Deviation (MAD)
1272 of the (anchored) best fit direction; α : angular difference between the best-fit anchored and
1273 free-floating directions; DRAT: maximum absolute difference from a pTRM check; CDRAT:
1274 cumulative DRAT. For further details on the selection criteria, see the Standard Paleointensity
1275 Definitions (SPD) (45). The two methods used for determining the paleointensity estimates were
1276 microwave (MW-IZZI) and thermal with an AF cleanse (Th-AF) Thellier-type measurements,
1277 using the IZZI protocol with pTRM checks. The maximum and minimum steps are shown as
1278 power integrals listed in Watt seconds (Ws) for the microwave measurements and temperature
1279 (T; °C) for the thermal experiments. Experimental results that do not pass the selection criteria

1280 are in gray, and the selection criteria that failed are underlined. Pi: Paleointensity estimate; H_{lab}:
1281 Applied lab field; s.d: Standard deviation.

1282 **Dataset S3 (separate file).** All the PINT15 (67) data from between 200-500 Ma (under 'PINT
1283 data'), updated and scored for Q_{PI} ('Q_{PI} scoring'). All the headings under 'PINT data' are
1284 consistent with the headings from the PINT15 database and explained on the 'Information'
1285 sheet (available at <http://earth.liv.ac.uk/pint/>). Yellow cells represent those that have been
1286 updated/added since the last upload of the PINT15 database (excel comments on the top right
1287 cell of each block describe why the changes have been made). Sites that do not have a 'Data'
1288 number have either superseded previous sites (sites that have been superseded since the last
1289 PINT update are included at the bottom of the sheet in red) or have been added from studies
1290 published since the last update. The description for Dataset S4 gives the full reference for sites
1291 with a letter ref (a-e). The 'Q_{PI} scoring' headings are consistent with names of the nine different
1292 criteria, which are explained in detail in Biggin and Paterson (35), and the final Q_{PI} score ('Q_{PI}') is
1293 the sum of all the criteria (these are scored a 1 if they pass the criteria and 0 if they do not). The
1294 explanation for each of the site scores is given under the corresponding headings in Dataset S4.

1295 **Dataset S4 (separate file).** Summary of the Q_{PI} scoring of all the sites from the updated PINT15
1296 database (67) from between 200-500 Ma (i.e. Dataset S3). The 'Study' is given in the format of
1297 first author_publication year. The 'PINTref' number corresponds to the 'REF' in PINT15 (letters
1298 are for studies added in the Dataset S3 update) and 'N' is the number of sites included in the
1299 study. The abbreviations used in 'Method' are consistent with those on the 'PI Methods' sheet of
1300 the PINT15 database. The headings under 'Q_{PI} Values' are consistent with the nine Q_{PI} criteria
1301 which are explained in detail in Biggin and Paterson (35), and the final Q_{PI} score ('Q_{PI}') is the sum
1302 of all the criteria (these are scored a 1 if they pass the criteria and 0 if they do not). Cells with
1303 the same headings under 'Notes' provide explanations for the site scorings for the first eight
1304 criteria (MAG only passes only when the raw data is published, e.g. in the studies supplementary
1305 information, on the MagIC database, etc.). 'Other Notes' provides other information such as the
1306 revision of age dates based on newly published ages, how sites have been recalculated, why
1307 certain QPI criteria have been rescored since previously being published, etc. *a) Shcherbakova

1308 et al, 2015 (68), b) Anwar et al, 2016 (69), c) Shcherbakova et al, 2017 (31), d) Usui and Tian,
1309 2017 (70) and e) Hawkins et al., 2019 (30).

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1311 SI References

1312

- 1313 1. M. Browne, R. Smith, A. M. Aitken, "Stratigraphical framework for the Devonian (Old
1314 Red Sandstone) rocks of Scotland south of a line from Fort William to Aberdeen" (2002).
- 1315 2. N. H. Woodcock, "Early Devonian Sedimentary and Magmatic Interlude after Iapetus
1316 Closure" in *Geological History of Britain and Ireland*, Wiley Online Books., (2012), pp. 193–209.
- 1317 3. T. H. Torsvik, Magnetic properties of the Lower Old Red Sandstone lavas in the Midland
1318 Valley, Scotland; palaeomagnetic and tectonic considerations. *Phys. Earth Planet. Inter.* 39, 194–
1319 207 (1985).
- 1320 4. M. F. Thirlwall, Isotope geochemistry and origin of calc-alkaline lavas from a Caledonian
1321 continental margin volcanic arc. *J. Volcanol. Geotherm. Res.* 18, 589–631 (1983).
- 1322 5. M. F. Thirlwall, Geochronology of Late Caledonian magmatism in northern Britain. *J.*
1323 *Geol. Soc. London.* 145, 951–967 (1988).
- 1324 6. J. B. Richardson, J. H. Ford, F. Parker, Miospores, correlation and age of some Scottish
1325 Lower Old Red Sandstone sediments from the Strathmore region (Fife and Angus). *J.*
1326 *Micropalaeontology* 3, 109–124 (1984).
- 1327 7. J. T. Sallomy, J. D. A. Piper, Palaeomagnetic Studies in the British Caledonides - IV Lower
1328 Devonian Lavas of the Strathmore Region. *Geophys. Journal, R. Astron. Soc.* 34, 47–68 (1973).
- 1329 8. T. H. Torsvik, *et al.*, Phanerozoic polar wander, palaeogeography and dynamics. *Earth-*
1330 *Science Rev.* 114, 325–368 (2012).
- 1331 9. R. Van der Voo, The reliability of paleomagnetic data. *Tectonophysics* 184, 1–9 (1990).
- 1332 10. M. R. Koymans, C. G. Langereis, D. Pastor-Galán, D. J. J. van Hinsbergen,
1333 Paleomagnetism.org: An online multi-platform open source environment for paleomagnetic
1334 data analysis. *Comput. Geosci.* 93, 127–137 (2016).

- 1335 11. P. L. McFadden, M. W. McElhinny, Classification of the reversal test in
1336 palaeomagnetism. *Geophys. J. Int.* 103, 725–796 (1990).
- 1337 12. L. Tauxe, R. F. Butler, R. Van der Voo, S. K. Banerjee, *Essentials of paleomagnetism*.
1338 (University of California Press, 2010).
- 1339 13. M. Kono, N. Ueno, Paleointensity determination by a modified thellier method. *Phys.*
1340 *Earth Planet. Inter.* 13 (1977).
- 1341 14. D. Stephenson, *et al.*, *Caledonian Igneous Rocks of Great Britain* (Joint Nature
1342 Conservation Committee, 1999).
- 1343 15. D. Stephenson, S. C. Loughlin, D. Millward, C. N. Waters, I. T. Williamson, *Carboniferous*
1344 *and Permian Igneous Rocks of Great Britain North of the Variscan Front* (Joint Nature
1345 Conservation Committee, 2003).
- 1346 16. A. Geikie, *The geology of central and western Fife and Kinross* (Mem. Geol. Surv.
1347 Scotland, 1900).
- 1348 17. S. Brindley, E. Spinner, Palynological assemblages from Lower Carboniferous deposits,
1349 Burntisland district, Fife, Scotland. *Proc. Yorksh. Geol. Soc.* 47, 215–231 (1989).
- 1350 18. A. A. Monaghan, M. A. E. Browne, “Nine $^{40}\text{Ar}/^{39}\text{Ar}$ dates from Carboniferous igneous
1351 rocks of the Midland Valley of Scotland” (2010).
- 1352 19. F. J. Fitch, J. A. Miller, S. C. Williams, Isotopic ages of British Carboniferous rocks. *C.r. 6th*
1353 *Int. Congr. Carb. Strat. Geol. (Sheffield, 1967)* 2, 771–789 (1970).
- 1354 20. T. H. Torsvik, O. Lyse, G. Atterås, B. J. Bluck, Palaeozoic palaeomagnetic results from
1355 Scotland and their bearing on the British apparent polar wander path. *Phys. Earth Planet. Inter.*
1356 55, 93–105 (1989).
- 1357 21. R. Van der Voo, The reliability of paleomagnetic data. *Tectonophysics* 184, 1–9 (1990).

- 1358 22. J. L. Kirschvink, R. E. Kopp, T. D. Raub, C. T. Baumgartner, J. W. Holt, Rapid, precise, and
1359 high-sensitivity acquisition of paleomagnetic and rock-magnetic data: Development of a low-
1360 noise automatic sample changing system for superconducting rock magnetometers.
1361 *Geochemistry, Geophys. Geosystems* 9 (2008).
- 1362 23. J. L. Kirschvink, The least-squares line and plane and the analysis of palaeomagnetic
1363 data. *Geophys. Journal, R. Astron. Soc.* 62, 699–718 (1980).
- 1364 24. Y. Yu, L. Tauxe, Testing the IZZI protocol of geomagnetic field intensity determination.
1365 *Geochemistry, Geophys. Geosystems* 6 (2005).
- 1366 25. R. S. Coe, S. Grommé, E. A. Mankinen, Geomagnetic paleointensities from radiocarbon-
1367 dated lava flows on Hawaii and the question of the Pacific nondipole low. *J. Geophys. Res.* 83,
1368 1740–1756 (1978).
- 1369 26. S. Levi, The effect of magnetite particle size on paleointensity determinations of the
1370 geomagnetic field. *Phys. Earth Planet. Inter.* 13, 245–259 (1977).
- 1371 27. D. J. Dunlop, Ö. Özdemir, Beyond Néel’s theories: Thermal demagnetization of narrow-
1372 band partial thermoremanent magnetizations. *Phys. Earth Planet. Inter.* 126, 43–57 (2001).
- 1373 28. A. Biggin, M. Perrin, J. Shaw, A comparison of a quasi-perpendicular method of absolute
1374 palaeointensity determination with other thermal and microwave techniques. *Earth Planet. Sci.*
1375 *Lett.* 257, 564–581 (2007).
- 1376 29. M. J. Hill, Y. Pan, C. J. Davies, An assessment of the reliability of palaeointensity results
1377 obtained from the Cretaceous aged Suhongtu section, Inner Mongolia, China. *Phys. Earth Planet.*
1378 *Inter.* 169, 76–88 (2008).
- 1379 30. L. M. A. Hawkins, *et al.*, An exceptionally weak Devonian geomagnetic field recorded by
1380 the Viluy Traps, Siberia. *Earth Planet. Sci. Lett.* 506, 134–145 (2019).

- 1381 31. V. V. Shcherbakova, *et al.*, Was the Devonian geomagnetic field dipolar or multipolar?
1382 Palaeointensity studies of Devonian igneous rocks from the Minusa Basin (Siberia) and the Kola
1383 Peninsula dykes, Russia. *Geophys. J. Int.* 209, 1265–1286 (2017).
- 1384 32. G. A. Paterson, A. J. Biggin, E. Hodgson, M. J. Hill, Thellier-type paleointensity data from
1385 multidomain specimens. *Phys. Earth Planet. Inter.* 245, 117–133 (2015).
- 1386 33. D. Heslop, A. P. Roberts, A Bayesian Approach to the Paleomagnetic Conglomerate Test.
1387 *J. Geophys. Res. Solid Earth* 123, 1132–1142 (2018).
- 1388 34. G. A. Paterson, A. R. Muxworthy, Y. Yamamoto, Y. Pan, Bulk magnetic domain stability
1389 controls paleointensity fidelity. *Proc. Natl. Acad. Sci.* 114, 13120–13125 (2017).
- 1390 35. A. J. Biggin, G. A. Paterson, A new set of qualitative reliability criteria to aid inferences
1391 on palaeomagnetic dipole moment variations through geological time. *Front. Earth Sci.* 2, 1–9
1392 (2014).
- 1393 36. G. A. Paterson, D. Heslop, A. R. Muxworthy, Deriving confidence in paleointensity
1394 estimates. *Geochemistry, Geophys. Geosystems* 11 (2010).
- 1395 37. A. R. Muxworthy, M. E. Evans, S. J. Scourfield, J. G. King, Paleointensity results from the
1396 late-Archaean Modipe Gabbro of Botswana. *Geochemistry, Geophys. Geosystems* 14, 2198–2205
1397 (2013).
- 1398 38. L. Tauxe, T. Yamazaki, “Paleointensities” in *Treatise on Geophysics*, (2007), pp. 509–563.
- 1399 39. J. Shaw, A New Method of Determining the Magnitude of the Palaeomagnetic Field:
1400 Application to five historic lavas and five archaeological samples. *Geophys. J. R. Astron. Soc.* 39,
1401 133–141 (1974).
- 1402 40. L. V de Groot, A. J. Biggin, M. J. Dekkers, C. G. Langereis, E. Herrero-Bervera, Rapid
1403 regional perturbations to the recent global geomagnetic decay revealed by a new Hawaiian
1404 record. *Nat. Commun.* 4, 2727 (2013).

- 1405 41. E. Hodgson, J. M. Grappone, A. J. Biggin, M. J. Hill, M. J. Dekkers, Thermoremanent
1406 Behavior in Synthetic Samples Containing Natural Oxyexsolved Titanomagnetite. *Geochemistry,*
1407 *Geophys. Geosystems* 19, 1751–1766 (2018).
- 1408 42. S. Xu, D. J. Dunlop, Thellier paleointensity theory and experiments for multidomain
1409 grains. *J. Geophys. Res. Solid Earth* 109 (2004).
- 1410 43. P. Riisager, J. Riisager, Detecting multidomain magnetic grains in Thellier palaeointensity
1411 experiments. *Phys. Earth Planet. Inter.* 125 (2001).
- 1412 44. G. A. Paterson, A simple test for the presence of multidomain behavior during
1413 paleointensity experiments. *J. Geophys. Res. Solid Earth* 116 (2011).
- 1414 45. G. A. Paterson, L. Tauxe, A. J. Biggin, R. Shaar, L. C. Jonestrask, On improving the
1415 selection of Thellier-type paleointensity data. *Geochemistry, Geophys. Geosystems* 15, 1180–
1416 1192 (2014).
- 1417 46. R. L. Wilson, The Thermal Demagnetization of Natural Magnetic Moments in Rocks.
1418 *Geophys. J. R. Astron. Soc.* 5 (1961).
- 1419 47. R. L. Wilson, An Instrument for Measuring Vector Magnetization at High Temperatures.
1420 *Geophys. J. R. Astron. Soc.* 7 (1962).
- 1421 48. R. Day, M. Fuller, V. A. Schmidt, Hysteresis properties of titanomagnetites: Grain-size
1422 and compositional dependence. *Phys. Earth Planet. Inter.* 13, 260–267 (1977).
- 1423 49. A. J. Biggin, D. N. Thomas, The application of acceptance criteria to results of Thellier
1424 palaeointensity experiments performed on samples with pseudo-single-domain-like
1425 characteristics. *Phys. Earth Planet. Inter.* 138, 279–287 (2003).
- 1426 50. P. A. Selkin, J. S. Gee, L. Tauxe, W. P. Meurer, A. J. Newell, The effect of remanence
1427 anisotropy on paleointensity estimates: A case study from the Archean Stillwater Complex. *Earth*
1428 *Planet. Sci. Lett.* 183, 403–416 (2000).

- 1429 51. G. A. Paterson, The effects of anisotropic and non-linear thermoremanent
1430 magnetizations on Thellier-type paleointensity data. *Geophys. J. Int.* 193, 694–710 (2013).
- 1431 52. A. Jackson, A. R. T. Jonkers, M. R. Walker, Four centuries of geomagnetic secular
1432 variation from historical records. *Philos. Trans. R. Soc. A Math. Phys. Eng. Sci.* 358, 957–990
1433 (2000).
- 1434 53. R. J. Veitch, I. G. Hedley, J.-J. Wagner, An investigation of the intensity of the
1435 geomagnetic field during roman times using magnetically anisotropic bricks and tiles. *Arch. Sci.*
1436 *Geneve.* 37, 359–373 (1984).
- 1437 54. M. H. Dodson, E. McLelland-Brown, Magnetic blocking temperatures of single-domain
1438 grains during slow cooling. *J. Geophys. Res.* 85, 2625–2637 (1980).
- 1439 55. S. L. Halgedahl, R. Day, M. Fuller, The effect of cooling rate on the intensity of weak-field
1440 TRM in single- domain magnetite. *J. Geophys. Res.* 85, 3690–3698 (1980).
- 1441 56. F. D. Stacey, The physical theory of rock magnetism. *Adv. Phys.* 12, 45–133 (1963).
- 1442 57. Y. Yu, Importance of cooling rate dependence of thermoremanence in paleointensity
1443 determination. *J. Geophys. Res. Solid Earth* 116 (2011).
- 1444 58. A. J. Biggin, S. Badejo, A. R. Muxworthy, M. J. Dekkers, The effect of cooling rate on the
1445 intensity of thermoremanent magnetization (TRM) acquired by assemblages of pseudo-single
1446 domain, multidomain and interacting single-domain grains. *Geophys. J. Int.* 193, 1239–1249
1447 (2013).
- 1448 59. A. Ferk, *et al.*, Influence of cooling rate on thermoremanence of magnetite grains:
1449 Identifying the role of different magnetic domain states. *J. Geophys. Res. Solid Earth* 119, 1599–
1450 1606 (2014).
- 1451 60. C. N. Santos, L. Tauxe, Investigating the Accuracy, Precision, and Cooling Rate
1452 Dependence of Laboratory-Acquired Thermal Remanences During Paleointensity Experiments.
1453 *Geochemistry, Geophys. Geosystems* 20, 383–397 (2019).

- 1454 61. P. A. Selkin, J. S. Gee, L. Tauxe, Nonlinear thermoremanence acquisition and implications
1455 for paleointensity data. *Earth Planet. Sci. Lett.* 256, 81–89 (2007).
- 1456 62. M. J. Hill, J. Shaw, Palaeointensity results for historic lavas from Mt Etna using
1457 microwave demagnetization/remagnetization in a modified Thellier-type experiment. *Geophys.*
1458 *J. Int.* 139, 583–590 (1999).
- 1459 63. R. Shaar, L. Tauxe, Instability of thermoremanence and the problem of estimating the
1460 ancient geomagnetic field strength from non-single-domain recorders. *Proc. Natl. Acad. Sci. U. S.*
1461 *A.* 112, 11187–11192 (2015).
- 1462 64. G. Kletetschka, P. J. Wasilewski, Grain size limit for SD hematite. *Phys. Earth Planet.*
1463 *Inter.* 129, 173–179 (2002).
- 1464 65. J. M. Grappone, “Prototyping the next generation of versatile paleomagnetic
1465 laboratory,” University of Liverpool. (2020).
- 1466 66. D. J. Dunlop, Theory and application of the Day plot (Mrs/Ms versus Hcr/Hc) 1.
1467 Theoretical curves and tests using titanomagnetite data. *J. Geophys. Res. Solid Earth* 107 (2002).
- 1468 67. A. J. Biggin, G. H. M. A. Strik, C. G. Langereis, The intensity of the geomagnetic field in
1469 the late-Archaeon: New measurements and an analysis of the updated IAGA palaeointensity
1470 database. *Earth, Planets Sp.* 61, 9–22 (2009).
- 1471 68. V. V. Shcherbakova, G. V. Zhidkov, V. P. Shcherbakov, A. V. Latyshev, A. M. Fetisova,
1472 Verifying the mesozoic dipole low hypothesis by the Siberian trap data. *Izv. Phys. Solid Earth* 51,
1473 47–67 (2015).
- 1474 69. T. Anwar, L. Hawkins, V. A. Kravchinsky, A. J. Biggin, V. E. Pavlov, Microwave
1475 paleointensities indicate a low paleomagnetic dipole moment at the Permo-Triassic boundary.
1476 *Phys. Earth Planet. Inter.* 260, 62–73 (2016).

1477 70. Y. Usui, W. Tian, Paleomagnetic directional groups and paleointensity from the flood
1478 basalt in the Tarim large igneous province: Implications for eruption frequency. *Earth, Planets*
1479 *Sp.* 69 (2017).

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1481

1482