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1 Tectonic rotations in Central Chile: new insights on the southern limit of 2 the Maipo Orocline

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15 Abstract

16 The Maipo Transition Zone or Maipo Orocline represents a change in strike of the
17 topography from N-S between 30° and 33°S to an NNE-SSW trend to the south. The southern
18 flank of the Maipo Orocline can be traced along strike to around 38°S. There the Andean
19 margin exposes another orogen bending, the Arauco Orocline. Although the tectonic
20 evolution of this segment of the Andean chain is relatively well constrained, the limit between
21 both oroclinal is unclear. In this paper, we show the results of a paleomagnetic study carried
22 out in three plutons of Upper Triassic age located along the Coastal Cordillera of central Chile
23 between 35° to 37°S. Paleomagnetic analysis from Constitución Granite and Cobquecura
24 Pluton shows an increase of clockwise vertical-axis rotations from north to south consistent
25 with previous data, whereas results obtained in the Hualpén Stock, located to the south, shows
26 minor counter-clockwise rotations. This change in the rotation pattern would reflect the
27 Maipo and Arauco Oroclines limit. The similar amounts of tectonic rotations recorded by
28 Mesozoic and Miocene rocks indicate that the whole area behaved as a single block and
29 constrains the maximum possible age for the occurrence of tectonic rotations to Miocene.

30 Keywords: Maipo Orocline, Paleomagnetism, Tectonic rotation, Triassic Plutons

32 1. Introduction

33 The Andean Cordillera extends for more than 8000 km on South America's western
34 margin from the Caribbean Sea at its north end to the Scotia Ridge at the south end. Although
35 this mountain range is a continuous element, its geological history involves a complex record

36 of accretions, collisions and two important periods of subduction-related magmatism, the
37 Carboniferous to Permian Gondwana Orogeny (e.g., Mpodozis and Ramos 1989; Bahlburg
38 and Hervé 1997) and the Jurassic to Recent Andean Cycle, with its ongoing subduction along
39 the Andean trench (e.g., Coira et al. 1982; Mpodozis and Ramos 1989). This generates a
40 complicated segmentation where tectonics, magmatism and sedimentation processes change
41 over time and space.

42 Three large pronounced curvatures of tectonic trends are present along the topographic
43 front of the southern Central Andes, including important changes in the rotations pattern and
44 from north to south these are the Vallenar, Maipo and Arauco oroclines (Arriagada et al.
45 2009, 2013; Ferrando et al. 2014).

46 The orocline concept was formulated by Carey (1955) for an orogenic system which
47 has been flexed in plan view to an elbow shape in a subsequent deformation phase. While
48 having a genetic connotation, this term is often incorrectly used in the literature as a
49 geometric description for any orogenic curvature regardless of the actual kinematics of the
50 curvature formation. One of the mechanisms most common involved in the formation of an
51 orocline, proposed by Isacks (1988), is the presences of differential shortening along the belt
52 causing tectonic rotations.

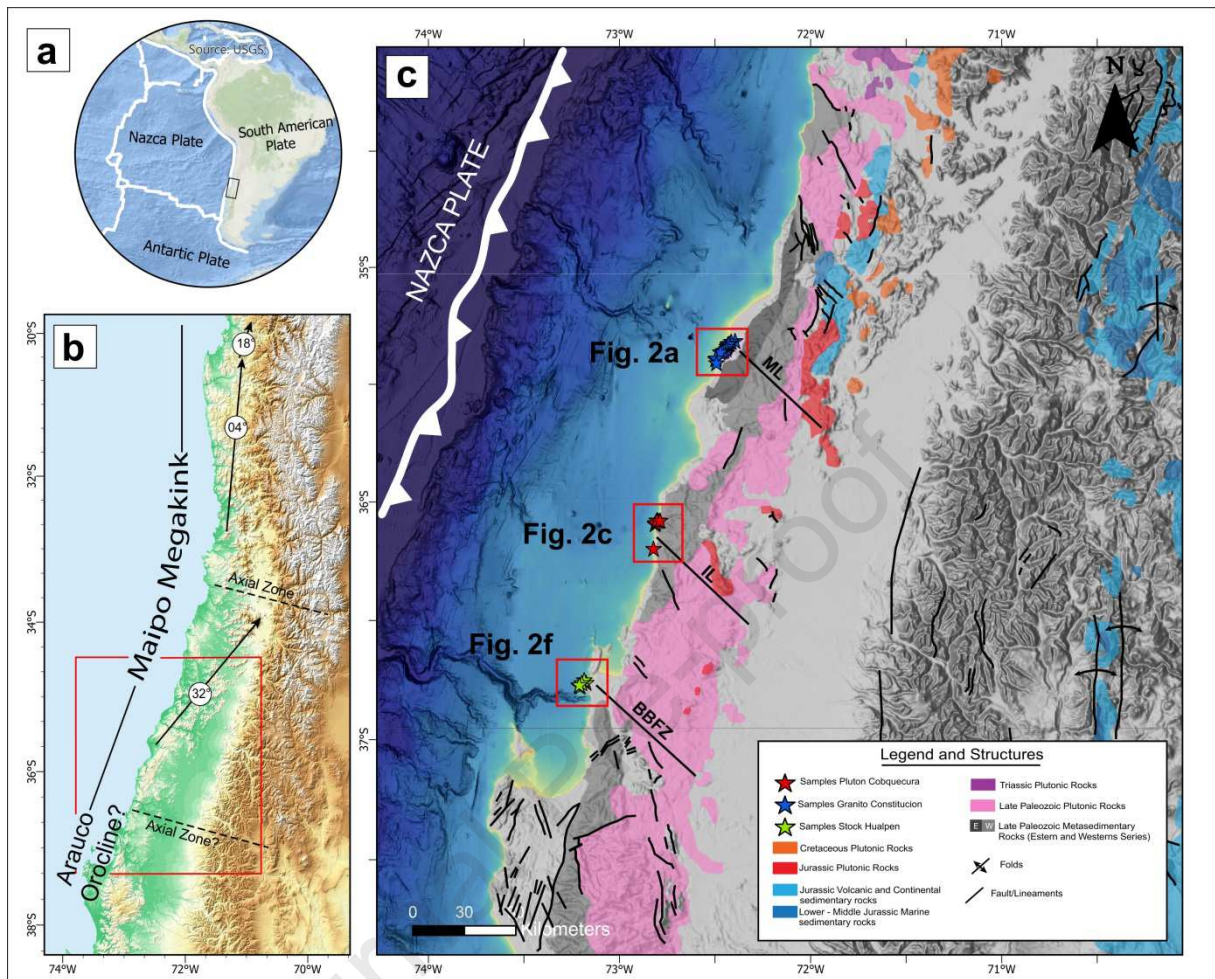
53 The Maipo Transition Zone (Yáñez et al. 2002) or Maipo Orocline (Farías et al. 2008)
54 represents a change in strike in the topography, geologic units and major structures. The trend
55 of major structures changes from N-S between 30° and 33°S in the north to a NNE-SSW trend
56 to the south. While the northern segment of the curvature corresponds with the ongoing
57 subduction of the Pampean flat-slab segment, which converges nearly horizontally beneath
58 the South American lithosphere, the southern segment coincides with the normal subduction
59 segment developed to the south of 33°S (Isacks 1988; Ramos et al. 2002).

60 Previous paleomagnetic analysis from Jurassic to Late Cretaceous rocks in the
61 Chilean-Pampean flat-slab segment shows small to non-significant clockwise vertical-axis
62 rotations (Beck et al. 1986, 1990; Creixell et al. 2006). In contrast, towards the south on the
63 normal subduction segment, paleomagnetic results obtained in Late Jurassic to Neogene rocks
64 show systematical clockwise rotations of up to 40° (Arriagada et al. 2013). The southern flank
65 of the Maipo Orocline can be traced along strike to around 38°S. There, the Andean margin
66 exposes another orogen bending, the Arauco “Orocline” (Fig 1; Melnick et al. 2009).

67 However, currently, no detailed paleomagnetic studies confirm the occurrence of tectonic
68 rotations and the limit between both oroclinal segments is still unknown.

69 In Triassic to Jurassic times, the Chilean active continental margin experienced
70 significant changes in the tectonic and magmatic regimes due to the transition from the
71 Gondwana to the Andean subduction process. In the segment between 34° and 42°S, the Late
72 Paleozoic-Early Mesozoic magmatic activity is typical for a magmatic arc experiencing
73 abundant crustal recycling, with the formation of large granitic batholiths and decreasing
74 magma volumes with time toward the Late Triassic (Martin et al. 1999; Lucassen et al. 2004).
75 Early Mesozoic intrusions in the western part of the Cordillera de la Costa comprise, from
76 north to south, La Estrella Granite, Pichilemu Granite, Constitución Granite, Cobquecura
77 Pluton, and the Hualpén Stock. In this contribution, we report new paleomagnetic data of the
78 Late Triassic Constitución, Cobquecura and Hualpen plutons, which crop out along the
79 Cordillera de la Costa between 33 ° - 37 ° S. They are post-orogenic intrusions with respect to
80 the Gondwanic orogenic cycle and their study allows us not only to investigate the southern
81 limit of the Maipo Orocline but also to work out the possibility of tectonic rotations related to
82 the Gondwana Breakup.

83



84

85 **Figure 1: (A) Tectonics plates; the black box shows the study area. (B) Paleomagnetic rotations domains showing the**
 86 **Maipo and Arauco Oroclines (Arriagada et al 2009). (C) Geological map, the red boxes indicates the sampling areas. BBFZ:**
 87 **Bio-Bio Fault Zone(Rehak et al. 2008); ML: Maule Lineament and IL: Itata Lineament (Willner et al. 2005).**

88 2. Geological Setting

89 During the transit from the Gondwanan to the Andean subduction, the active
 90 continental margin experienced significant changes in the tectonic and magmatic regimes.
 91 The igneous activity in the Late Paleozoic between 34° and 42°S shows typical features of
 92 magmatic arc with abundant crustal recycling, with the formation of large granitic batholiths
 93 (Berg and Breikreuz 1983; Berg et al. 1983), that intrude previously metamorphosed
 94 sedimentary units, leading to the formation of the Late Paleozoic Paired Metamorphic belt of
 95 Central and Southern Chile (Aguirre et al. 1972; Hervé et al. 1988).

96 The units that form this paired metamorphic complex are subdivided in the Western
 97 and the Eastern Series (Aguirre et al., 1972 and Fig. 1C). The Eastern Series consists of meta-
 98 greywacke and metapelite and is almost devoid of metabasites. It shows metamorphic

99 gradients from very low-grade rocks to amphibolite – granulite gneisses, related to the
100 proximity of Late Paleozoic calcalkaline granitoids (Hervé et al. 1988; Beck et al. 1991;
101 Martin et al. 1999). Based on petrological, structural, and age characteristics, the Eastern
102 Series is interpreted as resulting from subduction-related accretion and deformation of former
103 passive margin sediments, followed by intrusion of arc granitoids (Hervé et al. 1988; Willner
104 et al. 2005; Glodny et al. 2006). The Western Series represents the accretionary prism and is
105 characterized by meta-greywackes interlayered with lenses of oceanic crust. The latter consist
106 of mainly green-schist, rare blue-schist and associated Fe-Mn-rich meta-sediments, minor
107 meta-chert and serpentinite. All these rocks are pervasively deformed and characterized by a
108 second flat-lying transposition foliation. The Western Series is interpreted as a
109 paleoaccretionary complex built by basal accretion and the Eastern Series shows all
110 characteristics of frontally accreted sediments (Glodny et al. 2005; Richter et al. 2007).

111 Subsequently, after the definitive amalgamation of Pangea Supercontinent in late
112 Permian, the previously formed Orogen collapsed as result of the drastic diminish in plate
113 convergence, leading to the formation of elongated NNW-SSE extensional basins and large
114 volumes of mostly anorogenic silic magmatism of the Choiyoi province (Uliana and Biddle
115 1988; Mpodozis and Ramos 1989; Mpodozis and Kay 1990; Suarez and Bell 1992)

116 First evidence of post Gondwana magmatism in Southern Central Chile is a series of
117 isolated Upper Triassic high silica granitoids that crops out in Coastal Cordillera (Vásquez et
118 al., 2011). Their compositions are mainly monzogranitic (Hervé et al. 1988), except for the
119 Cobquecura Pluton, which also comprises gabbroic rocks (Hervé et al. 1988; Vásquez et al.
120 2009). These plutonic rocks crop out on both the western and the eastern flanks of the
121 Cordillera de la Costa (Hervé et al. 1988; Vásquez and Franz 2008). This granitoid intrudes
122 exclusively Late Paleozoic magmatic and metamorphic rocks of the previous tectonic cycle,
123 and are closely associated with small Upper Triassic sedimentary basins with an incipient
124 northwest orientation (Charrier 1979; Hervé et al. 1988).

125 The Early Mesozoic Constitución, Cobquecura and Hualpén plutons were all
126 emplaced at shallow crustal levels, display narrow contact aureoles, lack indications for
127 deformation or metamorphism, are post-kinematic with respect to the main foliation of their
128 metamorphic host rocks, and are exposed discontinuously in only small areas (Vásquez et al.
129 2011). The coincidence of the Triassic plutonic bodies with NW–SE trending lineaments have
130 led Vásquez et al. (2011) to propose oblique subduction conditions generating extension-
131 related NW–SE structures, which in turn facilitated the ascent of magmas followed by an

132 episode of slab steepening at around 210–197 Ma (Fig. 1C). Extensional conditions was
133 maintained for most of Jurassic and Early Cretaceous time, permitting the dominance of
134 extensional tectonic conditions on the continental margin, intense magmatic activity along the
135 arc, and abundant sedimentation in the backarc basin. These basins were inverted during the
136 Late Cretaceous compressive deformation and created a major regional unconformity
137 (Charrier et al. 2009).

138 The Constitución Granite (35.36°S, 72.42°W; Fig. 2a and b) is a homogeneous, K-
139 feldspar-porphyric biotite-monzogranite (Vásquez et al. 2011). It is intruded into metapelites
140 and metabasites of the Late Paleozoic metamorphic basement. The Constitución Granite has
141 been dated as 208 ± 6 Ma; 202 ± 6 Ma by K-Ar in biotite by Gana and Hervé (1983). Older
142 ages of 224 ± 1 Ma were obtained using Pb evaporation on zircon by Willner et al. (2005).

143 The Cobquecura Pluton is located to the south of Constitución Granite (~ 36°10'S,
144 72°47'W; Fig. 2c and d), intruded into the meta-sedimentary metamorphic basement. The
145 pluton, compared with other intrusions of the area, shows a bi-modal composition with
146 gabbroic and granitic end-members, and less abundant hybrid rocks consisting of tonalites
147 evidencing mixing (resorbed cores of plagioclase) and mingling (composite dikes, mafic
148 pillow-shaped enclaves and net-vein complexes) processes. This rock additionally contains a
149 few metamorphic enclaves (Vásquez and Franz 2008). Abundant pillows of mafic
150 composition in granitic magmas are typical for the mingling of coeval felsic and mafic
151 magmas (Fig. 2d). Granites contain fayalitic olivine, Fe-rich clinopyroxene, amphibole,
152 biotite, and ilmenite, whereas gabbros are mostly composed of Mg-rich pyroxenes,
153 amphiboles and minor forsteritic olivine and biotite (Vásquez et al. 2011). The age of the
154 Cobquecura magmatism is fairly well constrained. A K–Ar age of 211±4 Ma on ferroedenitic
155 amphibole was obtained for a fayalite granodiorite, a K–Ar Mg-hornblende age of 213±5 Ma
156 for a gabbro (Vásquez et al. 2005) and 209 ± 2 Ma; 203 ± 2 Ma (sample 04-22) by U-Pb
157 zircon (Vásquez et al. 2011). Field observations, namely the evidence for magma mingling,
158 show that the mafic and felsic magmas of the Cobquecura Pluton were strictly
159 contemporaneous. This inference is backed up by major element correlations for fayalite
160 granitoids and gabbro, suggesting a co-magmatic evolution for all Cobquecura igneous rocks
161 (Vásquez and Franz 2008).

162 The Hualpén Stock is located at the mouth of the Biobío River, which follows a NW–
163 SE-oriented lineament. It intruded into metapelites of the Paleozoic metamorphic basement
164 and is a homogeneous two mica–monzogranite (Fig. 2 f and g; Creixell et al. 2002). There are

165 precise geochronologic data for the stock (215 ± 4 Ma, K-Ar in biotite from Hervé et al. 1988;
166 meanwhile an age by Rb-Sr mineral data $220 \pm 5 - 222 \pm 2$ Ma has been mentioned by Lucassen
167 et al. 2004, and 227 ± 7 Ma in Rb-Sr mineral data from Glodny et al. 2006).

168

169 **3. Sampling and methodology**

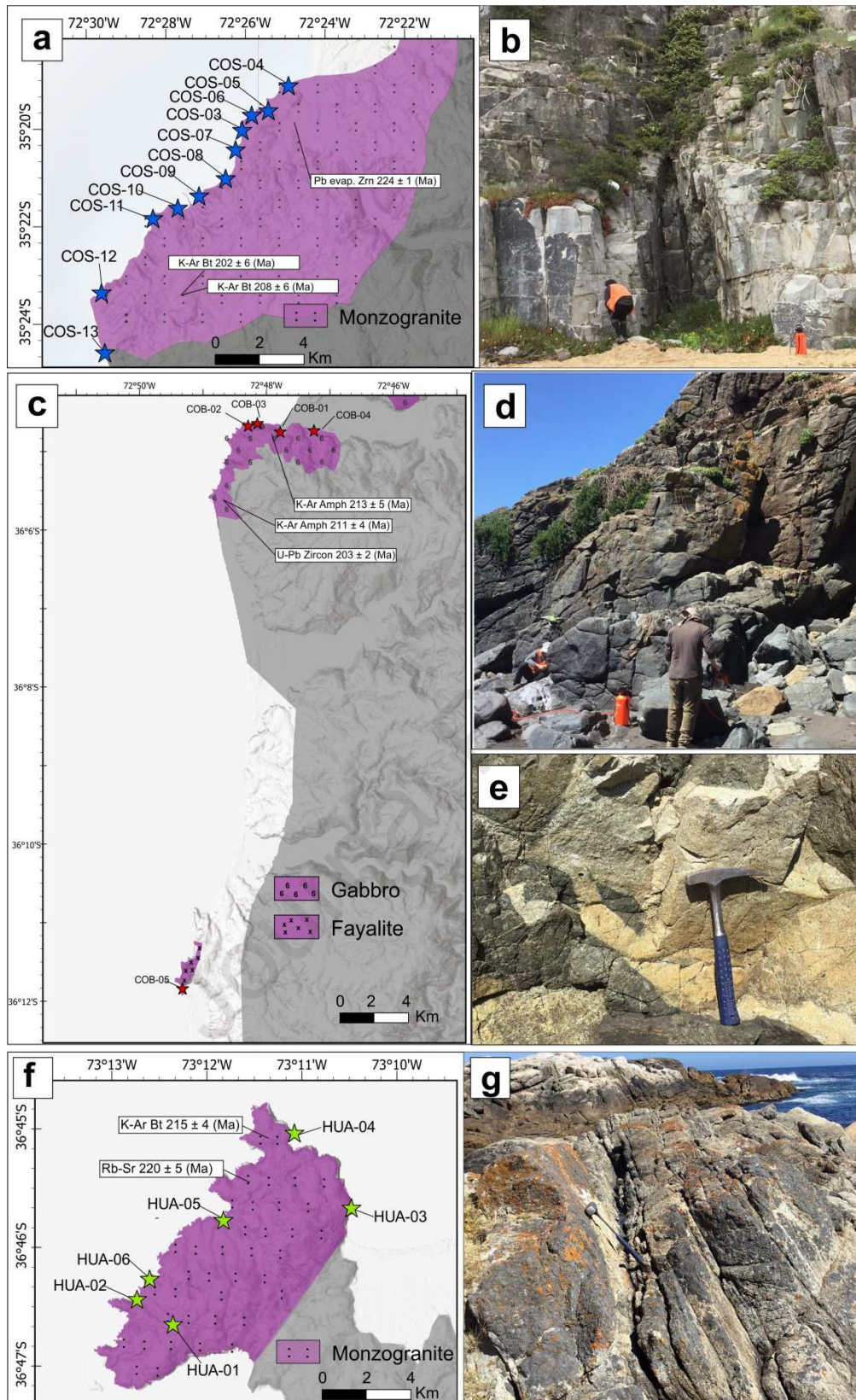
170 Paleomagnetic sampling was performed during two field seasons when more than 120
171 oriented cores were collected. The sampling was carried out with a portable gasoline-powered
172 drill system, and all samples were oriented in situ using magnetic and solar compasses. The
173 collection of oriented cores belongs to 23 stations and at least five different samples were
174 obtained from each sampling site.

175 Magnetic remanences were measured with a JR6 spinner magnetometer in the
176 Paleomagnetism Laboratory “Daniel A. Valencio” of the IGeBA (UBA-CONICET) and with
177 the “SushiBar” magnetometer (a fully automated system based on a 2G Enterprises, Inc.,
178 three-axis superconducting magnetometer) in the Geophysics Laboratory of Ludwig-
179 Maximilians-Universität (Munich). Thermal demagnetizations were carried out by an ASC
180 furnace and alternating fields (AF) demagnetizations were done with a tumbler LD-A
181 (AGICO Inc) alternating field demagnetizer. AGICO “Remasoft” software was used to
182 analyze the magnetic behaviors of the samples and determine the characteristic remanent
183 magnetizations (ChRM).

184 The ChRMs were determined, as described in Ruiz González et al. (2020), only in
185 their geographic coordinates due to the lack of paleohorizontal evidence. Characteristic
186 remanence directions (ChRM) with less than 15° of Maximum Angular Deviation (MAD)
187 calculated using at least 3 demagnetization steps were exclusively accepted for the analysis
188 (Kirschvink 1980). Samples with less than 20% of the NRM intensity after steps of 20 mT or
189 300°C of demagnetization were discarded, considering that low coercivities (H_c) or low un-
190 blocking temperatures could belong to spurious magnetic components. The paleomagnetic
191 analysis was performed using the ChRMs of the samples without averaging them per site,
192 according to Deenen et al. (2011), and each pluton was considered as belonging to the same
193 tectonic block (Fig. 1C). A cut-off angle of 40° was applied to the ChRM to filter transitional
194 and spurious data (Wilson et al., 1972). In order to validate these results, a reversal test was
195 carried out with the ChRM directions inside the cut off angle.

196 To discuss the tectonic implications of the paleomagnetic results and detect possible
197 spatial variations in the amount of tectonic rotation and flattening values with respect to stable
198 South America for each pluton, we used Demarest (1983). The references paleopoles for
199 South America for 210 Ma (Cobquecura Pluton) and 220 Ma (Constitución Granite and
200 Hualpén Stock) were taken from Torsvik et al. (2012). The mean directions for each locality
201 and the tectonic rotations and inclination errors are given in Tables 1-4.

202 Nine sites were sampled in the granite (COS3-COS12, 50 cores) and one (COS13, 5
203 cores) in the host rock (Fig. 2a). Sampling covered mainly the western sector of the pluton
204 with a distance between adjacent sites generally around hundreds of meters (660m) up to 3
205 km. The Cobquecura Pluton was sampled at 4 localities (COB1 to COB4) at the northern part
206 of the pluton and one (COB5) at the southern part near to Santa Rita (30 cores in total) (Fig.
207 2c). Also, six sites were drilled in Hualpén stock located ~79 km to the south near to
208 Concepción city (HUA-1 to HUA-6; Fig. 2f). A total of 27 independently oriented cores were
209 obtained.



210

211 **Figure 2: (a) Geological map of Constitución Granite. Ages from Willner et al. (2005) and Gana and Hervé (1983). (b)**
 212 **Outcrop view. (c) Geological map of Cobquecura Pluton. Radiometric ages from Vásquez et al. 2005 and 2011. (d)**
 213 **Outcrop view. (e) Felsic intrusion in gabbro. (f) Geological map of Hualpén Stock. Isotopic ages from Lucassen et al.**
 214 **(2004) and Hervé et al 1988. (g) Outcrop view with clearly developed subparallel joints. Blue, red and green stars show**
 215 **sampling locations in each studied unit.**

216 4. Characteristic directions

217 In this section we analyse the paleomagnetic data obtained from the Constitución,
 218 Cobquecura and Hualpén plutons. In spite of the fact that in the three areas a characteristic
 219 remanence direction could be obtained of normal and reversed polarity and all correspond to
 220 Upper Triassic rocks, each pluton will be analyzed separately.

221

222 4.1. Constitución Granite

223 Characteristic remanence directions were obtained from most of the samples of the
 224 monzogranite (see Table 1). On the other hand, four of the five samples obtained from the
 225 host rock presented an erratic behavior during demagnetization procedures (Fig. 3f). The
 226 unique component obtained (sample COS13-3, see Table 1) shows an anomalous direction
 227 and was discarded for tectonic analysis.

228 For samples from the Constitución Granite both demagnetizing methods were efficient
 229 in isolating a stable univectorial normal or reverse polarity remanent magnetization (Fig. 3).
 230 The ChRM was recovered by thermal demagnetization in the temperature range 450-530°C,
 231 or the range 10-60 mT under AF cleaning (Fig. 3). The characteristic component was
 232 determined with MAD under 10° in most cases (see Table 1). Since both polarities were
 233 recorded, the McFadden and McElhinny (1990) reversal test was performed. An
 234 indeterminate result was obtained with a $\gamma_{\text{observed}} = 12.5^\circ$ and a $\gamma_{\text{critical}} = 28.1^\circ$.

235 The mean direction for the Constitución Granite obtained after ruling out eight ChRM
 236 considered anomalous differs from the expected Late Triassic paleomagnetic field suggesting
 237 a clockwise tectonic rotation of $13.5 \pm 11.5^\circ$ (Fig. 3 II, see Table 4).

238 **Table 1: Paleomagnetic results in geographic coordinates for the Constitución Granite.**

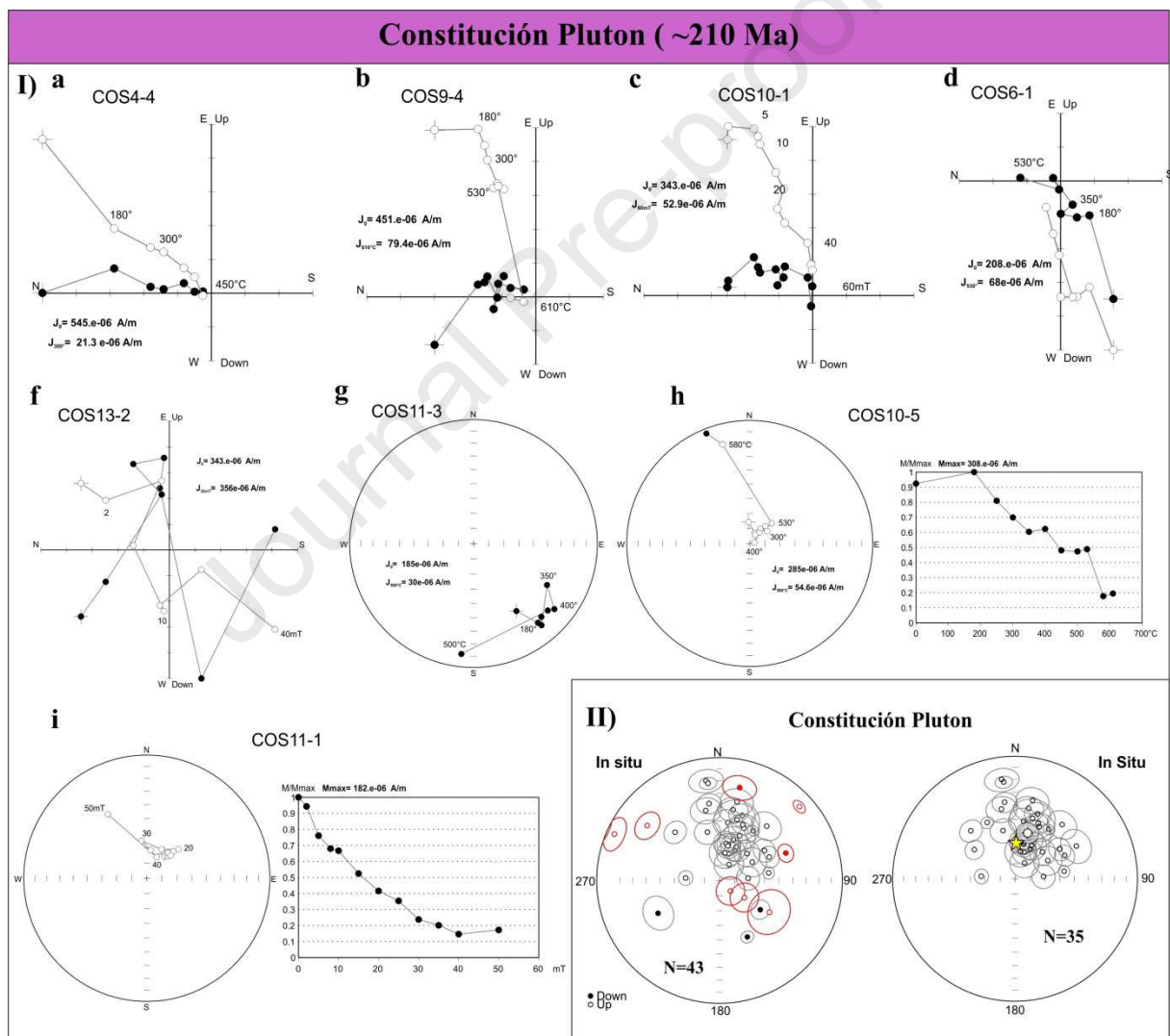
Sample	Lat.	Long.	Lithology	Dec	Inc	MAD
<i>Constitución Pluton (~210 Ma)</i>						
COS-03-1	-35.333	-72.435	Monzogranite	56.4	-53.8	7.5
COS-03-2	-35.333	-72.435	Monzogranite	----	----	----
COS-03-3	-35.333	-72.435	Monzogranite	----	----	----
COS-03-4	-35.333	-72.435	Monzogranite	----	----	----
COS-03-5	-35.333	-72.435	Monzogranite	----	----	----
COS-04-1	-35.318	-72.415	Monzogranite	11.	-41.2	9.1
COS-04-2	-35.318	-72.415	Monzogranite	352.3	-19.3	9.4
COS-04-3	-35.318	-72.415	Monzogranite	----	----	----

COS-04-4	-35.318	-72.415	Monzogranite	12.6	-34.7	9.8
COS-04-5	-35.318	-72.415	Monzogranite	307.2	-27.0	8.5
COS-05-1	-35.326	-72.424	Monzogranite	19.1	-21.5	8.1
COS-05-2	-35.326	-72.424	Monzogranite	274.2	-67.5	4.8
COS-05-3	-35.326	-72.424	Monzogranite	----	----	----
COS-05-4	-35.326	-72.424	Monzogranite	34.1	-49.6	3.5
COS-05-5	-35.326	-72.424	Monzogranite	8.4	-50.7	9.4
COS-06-1	-35.328	-72.431	Monzogranite	241.9	42.7	10.7
COS-06-2	-35.328	-72.431	Monzogranite	41.0	-41.0	9.6
COS-06-3	-35.328	-72.431	Monzogranite	19.5	-71.3	8.
COS-06-4	-35.328	-72.431	Monzogranite	12.3	23.7	9.5
COS-06-5	-35.328	-72.431	Monzogranite	16.2	-47.4	14.2
COS-07-1	-35.340	-72.437	Monzogranite	352.2	-37.0	7.4
COS-07-2	-35.340	-72.437	Monzogranite	----	----	----
COS-07-3	-35.340	-72.437	Monzogranite	----	----	----
COS-07-4	-35.340	-72.437	Monzogranite	353.3	-21.5	4.8
COS-07-5	-35.340	-72.437	Monzogranite	----	----	----
COS-08-1	-35.350	-72.441	Monzogranite	346.0	-56.7	8.3
COS-08-2	-35.350	-72.441	Monzogranite	47.3	-12.5	3.9
COS-08-3	-35.350	-72.441	Monzogranite	293.8	-7.2	8.8
COS-08-4	-35.350	-72.441	Monzogranite	25.1	-51.2	9.9
COS-08-5	-35.350	-72.441	Monzogranite	9.9	-65.0	8.1
COS-09-1	-35.356	-72.453	Monzogranite	24.3	-57.2	10.
COS-09-2	-35.356	-72.453	Monzogranite	22.6	-49.3	6.8
COS-09-3	-35.356	-72.453	Monzogranite	19.0	-61.3	9.
COS-09-4	-35.356	-72.453	Monzogranite	11.3	-71.4	10.3
COS-09-5	-35.356	-72.453	Monzogranite	13.6	-66.4	8.3
COS-10-1	-35.360	-72.462	Monzogranite	26.2	-65.3	9.8
COS-10-2	-35.360	-72.462	Monzogranite	27.1	-67.8	6.7
COS-10-3	-35.360	-72.462	Monzogranite	134.1	-79.9	9.4
COS-10-4	-35.360	-72.462	Monzogranite	45.6	-64.4	8.8
COS-10-5	-35.360	-72.462	Monzogranite	124.5	-69.9	9.5
COS-11-1	-35.363	-72.472	Monzogranite	60.8	-65.0	9.6
COS-11-2	-35.363	-72.472	Monzogranite	317.3	-45.7	7.4
COS-11-3	-35.363	-72.472	Monzogranite	153.7	47.5	4.1
COS-11-4	-35.363	-72.472	Monzogranite	122.5	-50.1	14.4
COS-11-5	-35.363	-72.472	Monzogranite	81.3	-56.7	3.9
COS-12-1	-35.389	-72.494	Monzogranite	14.0	-65.4	7.4
COS-12-2	-35.389	-72.494	Monzogranite	85.5	-71.0	7.8
COS-12-3	-35.389	-72.494	Monzogranite	31.1	-77.8	9.7
COS-12-4	-35.389	-72.494	Monzogranite	19.2	-43.3	8.2
COS-12-5	-35.389	-72.494	Monzogranite	64.9	-53.1	6.9
COS-13-1	-35.409	-72.492	Milonite	----	----	----
COS-13-2	-35.409	-72.492	Milonite	----	----	----
COS-13-3	-35.409	-72.492	Milonite	67.6	41.20	5.6
COS-13-4	-35.409	-72.492	Milonite	----	----	----

COS-13-5	-35.409	-72.492	Milonite	---	---	---
Mean	N/n	Dec	Inc	α_{95}	k	
COS	35/55	15.4	-58.2	7.2	12	

Lat., Long.: site location (to the nearest half a minute). (n/N): number of samples involved in statistics/ total number of samples. Decl., Incl.: declination and inclination of paleomagnetic vector in geographical coordinates. α_{95} : cone of 95% confidence level around mean direction (Fisher, 1953). k: precision parameter (Fisher, 1953). Sites not considered for mean calculation are highlighted in gray. Samples with (---) indicates that ChRM could not be obtained.

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241 **Figure 3: (I) Paleomagnetic behavior observed for samples from the Constitución Granite (a-e and g-i) and the host rock**
 242 **(f). The open (full) symbols in the Zijderveld diagram correspond to the projection in the vertical (horizontal) plane; in the**
 243 **stereonet open (full) symbols indicate up-wards (downwards) inclination. (II) To the left directions of the characteristic**
 244 **remance of the specimens of Constitución Granite, samples discarded are showing in red; at right mean direction**
 245 **carried to a single polarity. Open (filled) symbols correspond to projection onto the upper (lower) hemisphere. The**
 246 **yellow star corresponds to the references direction.**

247 4.2. Cobquecura Pluton

248 Samples from the pluton exhibited stable magnetic behaviors that allowed us to define
 249 ChRM (see Table 2). These were mostly determined with $MAD < 11^\circ$. Both demagnetization
 250 methods were efficient in discriminating the magnetic components. In most cases, a unique
 251 ChRM was isolated at temperatures/ maximum alternating fields ranges of 180-580°C or 10-
 252 60mT (Fig.4I a-b and d-f). Sometimes thermal demagnetization showed an abrupt decay of
 253 the remanence towards the origin of coordinates above 500 °C (Fig. 4I-e and f). Samples from
 254 COB2 record two components, one of lower temperature (180-350°C) and other between 400-
 255 580°C (Fig.4I-c). The lower temperature component shows scattered directions and was
 256 discarded from further analysis (see Table 2). These differences in the demagnetization
 257 behaviors between different sites reflect the widely variable textural and compositional
 258 lithology of the pluton.

259 Remanences of this intrusive body record both polarities of the paleomagnetic field
 260 (Fig.4II), although positive inclinations predominate. The reversal test of McFadden and
 261 McElhinny (1990) was performed with an indeterminate result: $\gamma_{\text{observed}} = 19^\circ$ and a $\gamma_{\text{critical}} = 21.3^\circ$.
 262 The mean direction obtained for the Cobquecura Pluton differs from the expected
 263 direction (see star in Fig. 4II and Table 4) suggesting a $41.3 \pm 15^\circ$ clockwise rotation (Fig. 4II).

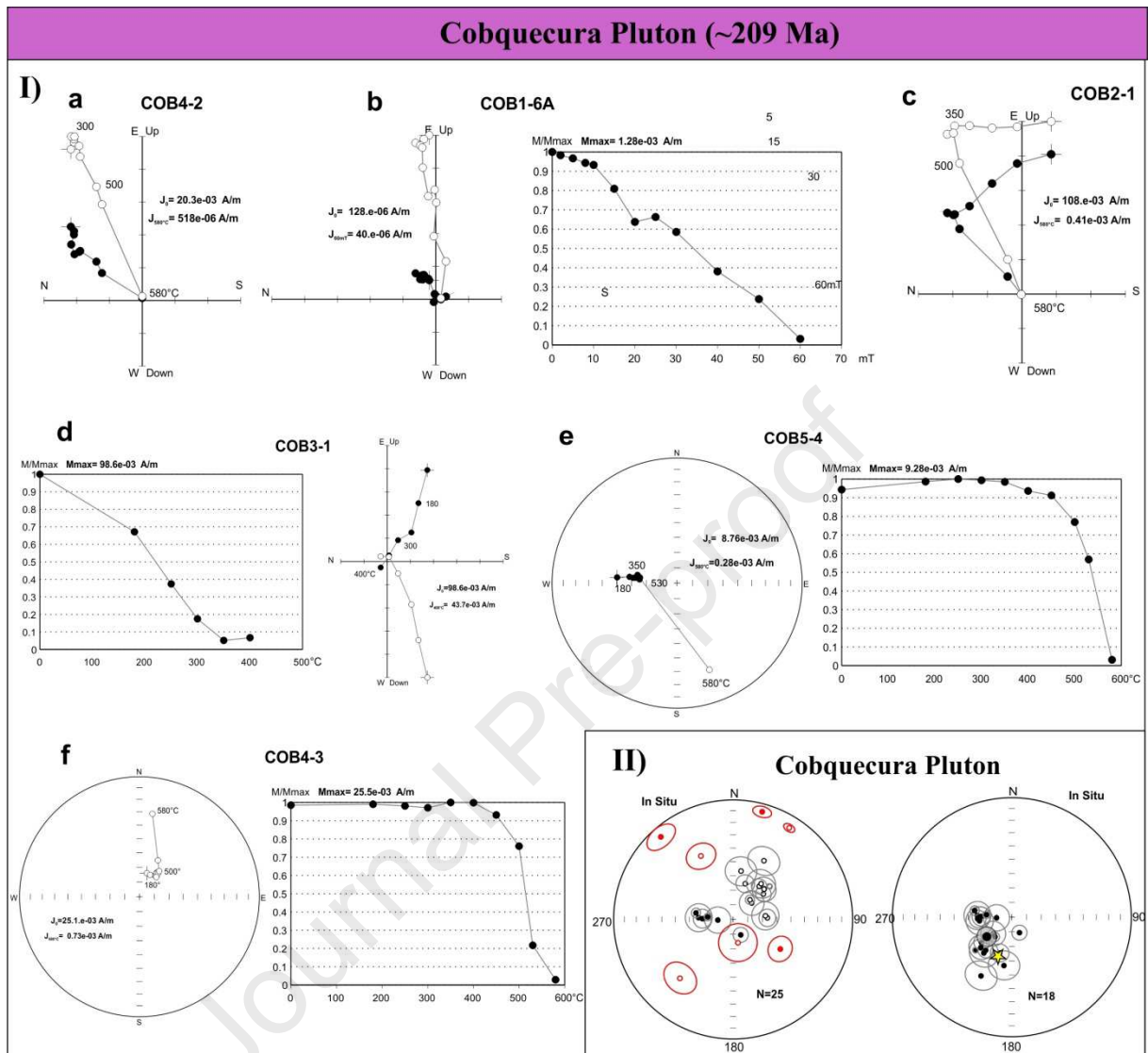
264

265 **Table 2: Paleomagnetic results from Cobquecura Pluton**

Sample	Lat.	Long.	Lithology	Dec	Inc	MAD
<i>Cobquecura Pluton (~209 Ma)</i>						
COB1-1	-36.079	-72.796	Gabbro	26.6	-12.6	8.6
COB1-2	-36.079	-72.796	Gabbro	----	----	----
COB1-3C	-36.079	-72.796	Gabbro	9.3	-56.3	10.3
COB1-4	-36.079	-72.796	Gabbro	87.3	-66.6	5.8
COB1-5	-36.079	-72.796	Gabbro	83.6	-67.8	9.4
COB1-6A	-36.079	-72.796	Gabbro	48.6	-73.1	8.4
COB2-1 ^l	-36.077	-72.802	Gabbro	139.6	1.1	2.6
COB2-1 ^h	-36.077	-72.802	Gabbro	47.8	-56.1	1.8
COB2-2C ^l	-36.077	-72.802	Gabbro	74.2	-3.7	8.6
COB2-2C ^l	-36.077	-72.802	Gabbro	337.4	4.4	4
COB2-2C ^h	-36.077	-72.802	Gabbro	333.1	-40.2	9.4
2COB-2	-36.077	-72.802	Gabbro	222.1	-34	10.8
COB2-3 ^h	-36.077	-72.802	Gabbro	49.9	-62.9	5.3
COB2-3 ^l	-36.077	-72.802	Gabbro	96.1	-37	4.2
COB2-4	-36.077	-72.802	Gabbro	31.5	-11.1	3.2

COB2-5	-36.077	-72.802	Gabbro	15.1	7.4	4.9
COB2-6	-36.077	-72.802	Gabbro	18.8	-64.2	6.2
COB3-1	-36.078	-72.805	Gabbro	122.1	51.5	8.4
COB3-2	-36.078	-72.805	Gabbro	----	----	----
COB3-3	-36.078	-72.805	Gabbro	27.7	-43.8	10.3
COB3-4	-36.078	-72.805	Gabbro	267.7	79.6	8.9
COB3-5	-36.078	-72.805	Gabbro	153.8	78.3	5.1
COB3-6B	-36.078	-72.805	Gabbro	318.8	9	8.7
COB4-1	-36.079	-72.788	Fayalite	----	----	----
COB4-2	-36.079	-72.788	Fayalite	38.3	-61.1	5.2
COB4-2A	-36.079	-72.788	Fayalite	38	-58.8	9.8
COB4-3	-36.079	-72.788	Fayalite	40.6	-72.4	3.1
COB4-4	-36.079	-72.788	Fayalite	168.3	-73.8	13
COB4-5	-36.079	-72.788	Fayalite	45.6	-60.3	8.2
COB5-1B	-36.197	-72.822	Gabbro	272.9	66.4	9.3
COB5-2	-36.197	-72.822	Gabbro	271.5	69	6.3
COB5-3	-36.197	-72.822	Gabbro	----	----	----
COB5-4	-36.197	-72.822	Gabbro	280	64.2	3.2
COB5-5	-36.197	-72.822	Gabbro	275.6	72.7	2.5
Mean	N/n	Dec	Inc	α_{95}	k	
COB	22/29	232.1	68.4	6.7	28	

Lat., Long.: site location (to the nearest half a minute). (n/N): number of samples involved in statistics/ total number of samples. Decl., Incl.: declination and inclination of paleomagnetic vector in geographical coordinates. α_{95} : cone of 95% confidence level around mean direction (Fisher, 1953). k: precision parameter (Fisher, 1953). Sites not considered for mean calculation are highlighted in gray. ^h: High temperature component. ^l: Low temperature component. Samples with (---) indicates that ChRM could not be obtained.



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268 **Figure 4: (I) Paleomagnetic behaviour observed in Cobquecura Pluton. See text for explanation. The open (full) symbols**
 269 **in the Zijderveld diagram correspond to the projection in the vertical (horizontal) plane; in the stereonet open (full)**
 270 **symbols indicate upwards (downwards) inclination.. (II) Directions of the ChRM showing in red are not involved in the**
 271 **calculation of the mean. To the right mean direction carried to reversed polarity, the star corresponds to the expected**
 272 **direction.**

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4.3. Hualpén Stock

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Characteristic remanence directions of both polarities were obtained from most of the samples of the Hualpén Stock (see Table 3, Fig.5). In most cases, a unique component decaying towards the origin could be isolated between 180-580°C or 10-60mT (Fig.5I). Sometimes an abrupt decay during thermal demagnetization was observed above 500 °C (Fig. 5I-b and e) suggesting that titanomagnetite is the main mineral carrying the remanence. A few samples from the stock record two components: one of lower temperature (180-300°C) and another between 350-580°C (Fig.5I-f). In contrast with the Cobquecura Pluton, most of

281 the remanence of lower temperature observed in the Hualpén Stock record a direction similar
 282 to the samples with a unique component (see Table 3).

283 Since both polarities were recorded, the reversal test of McFadden and McElhinny
 284 (1990) was performed. An indeterminate result was obtained with an $\gamma_{\text{observed}} = 2.6^\circ$ and a
 285 $\gamma_{\text{critical}} = 24.1^\circ$. The mean direction obtained from the stock, after applying a cut-off angle of
 286 45° to filter transitional and spurious data, differs from the expected paleomagnetic field for
 287 the Late Triassic suggesting that the stock is $16 \pm 16.3^\circ$ counter-clockwise rotated (Fig. 5II and
 288 Table 4).

289 **Table 3: Characteristic remanence directions from Hualpén Stock.**

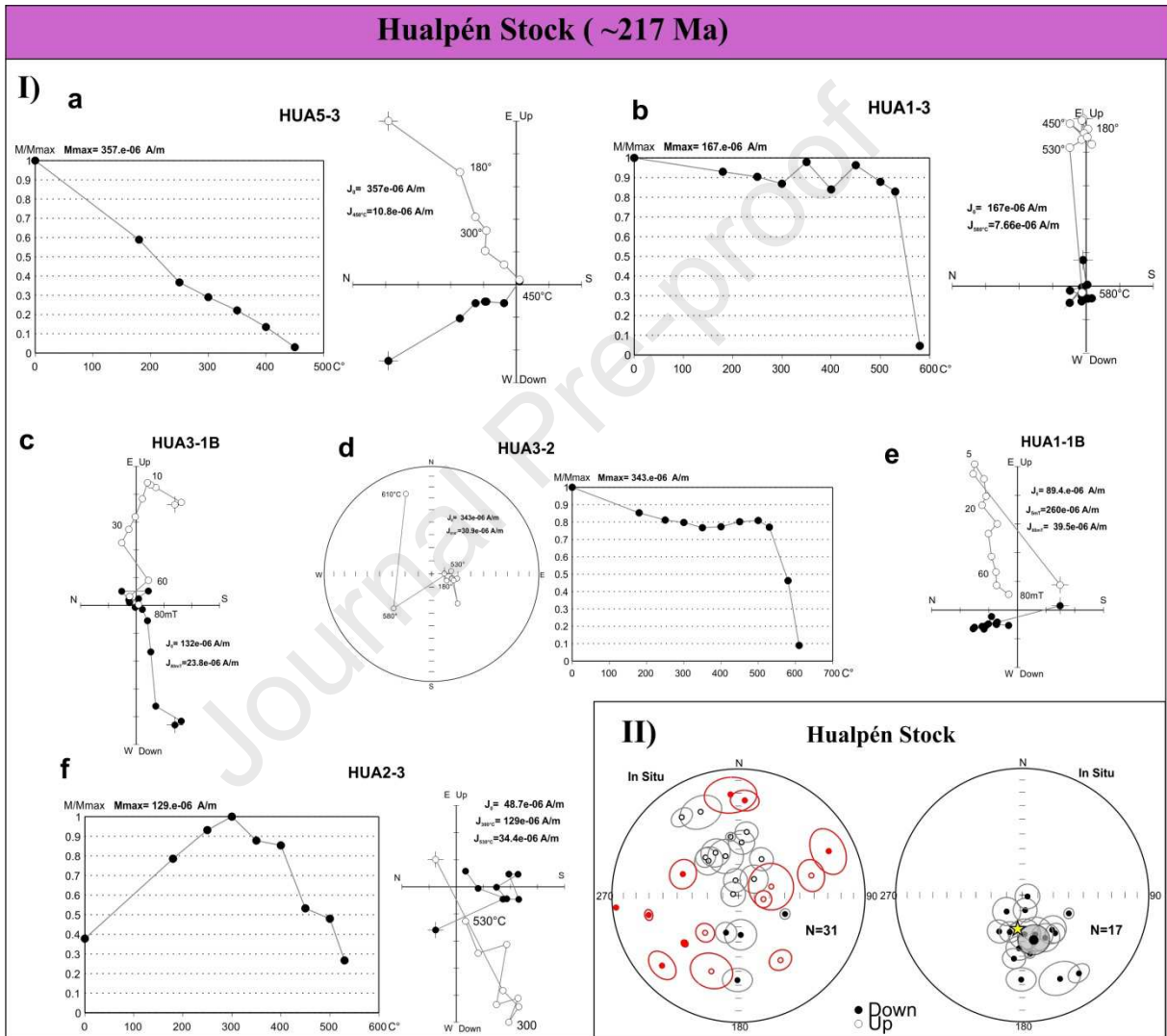
Sample	Lat.	Long.	Lithology	Dec	Inc	MAD
<i>Hualpén Stock (~217 Ma)</i>						
HUA1-1	-36.777	-73.206	Monzogranite	348.9	-80.1	7.6
HUA1-2	-36.777	-73.206	Monzogranite	3.8	25.6	7.8
HUA1-3	-36.777	-73.206	Monzogranite	285.2	-86.5	7.9
HUA2-1B	-36.774	-73.212	Monzogranite	319.2	-57.3	6.3
HUA2-2	-36.774	-73.212	Monzogranite	32.3	-62.1	7.5
HUA2-3 ^l	-36.774	-73.212	Monzogranite	319.3	-60.3	9
HUA2-3 ^h	-36.774	-73.212	Monzogranite	176.7	64.6	10
HUA2-4 ^h	-36.774	-73.212	Monzogranite	198.5	64.8	6.7
HUA2-4 ^l	-36.774	-73.212	Monzogranite	352.4	-51.3	3
HUA2-5	-36.774	-73.212	Monzogranite	180.7	33.6	8.5
HUA3-1 ^l	-36.761	-73.175	Monzogranite	264.4	3.2	0.4
HUA3-1 ^h	-36.761	-73.175	Monzogranite	41.9	57.6	6
HUA3-2	-36.761	-73.175	Monzogranite	254.6	40	9.4
HUA3-3	-36.761	-73.175	Monzogranite	278.2	73.8	5.4
HUA3-4	-36.761	-73.175	Monzogranite	175.5	-21	13.9
HUA3-5	-36.761	-73.175	Monzogranite	335.7	-28.2	11.9
HUA3-6	-36.761	-73.175	Monzogranite	326.9	39	6.8
HUA4-1	-36.751	-73.185	Monzogranite	228.4	42.6	2.2
HUA4-2	-36.751	-73.185	Monzogranite	----	----	----
HUA4-3	-36.751	-73.185	Monzogranite	257.8	28.3	3
HUA4-4	-36.751	-73.185	Monzogranite	111.5	57.5	3.3
HUA4-5	-36.751	-73.185	Monzogranite	19.7	36.6	12.4
HUA5-1	-36.763	-73.197	Monzogranite	43.4	-75.4	9
HUA5-2	-36.763	-73.197	Monzogranite	3.4	-55.3	8
HUA5-3	-36.763	-73.197	Monzogranite	331.8	-58.5	9.2
HUA5-4 ^l	-36.763	-73.197	Monzogranite	341.8	-63.2	11.2
HUA5-4 ^h	-36.763	-73.197	Monzogranite	227.5	19.2	8.6
HUA6-1	-36.771	-73.210	Monzogranite	7.4	-48	6.9
HUA6-2 ^l	-36.771	-73.210	Monzogranite	324.1	-24.6	6.3

HUA6-2 ^h	-36.771	-73.210	Monzogranite	63.8	21.6	12.8
HUA6-4	-36.771	-73.210	Monzogranite	291.1	51.1	9
HUA6-5	-36.771	-73.210	Monzogranite	254.6	68.1	14.7
Mean	N/n	Dec	Inc	α_{95}	k	
HUA	17/32	165.6	60	9.9	14	

Same as Table 2

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5. Tectonic rotations

299 For each locality, tectonic rotations and inclination anomalies were calculated
 300 according to Demarest et al. (1983) using the reference poles for 210 Ma and 220 Ma from
 301 Torsvik et al (2012) transported to South American coordinates through the translations
 302 proposed in that work (see Table 4).

303 **Table 4: Tectonic rotations**

Locality	Lat.	Long	age ref.	Dec.obs.	Inc.obs.	α_{95}	Dec.exp	Inc.exp	α_{95}	R \pm Δ R	I \pm Δ I
Constitución Granite	-35.3	-72.4	220	15.4	-58.2	7.2	1.9	-68.1	1.5	13.5 \pm 11.5	9.9 \pm 5.8
Cobquecura Pluton	-36.1	-72.8	210	232.1	68.4	6.7	190.8	66	1.5	41.3 \pm 15	-2.4 \pm 5.5
Hualpén Stock	-36.7	-73.2	220	165.6	60	9.9	181.6	68.1	1.5	-16.0 \pm 16.4	9.1 \pm 8.0

Age: estimated age of the magnetization and of the reference from Torsvik et al. (2012) used to calculate the rotation; Lat, Long: Latitude, longitude, position of the localities used in the calculation of the tectonic parameters; Dobs, Iobs: observed paleomagnetic declination and inclination; α_{95} : semi-angle of confidence; Dexp, Iexp: are the expected declination and inclination from expected pole; R($^{\circ}$), Δ R, I($^{\circ}$), Δ I: rotation and inclination and their associated errors (Demarest, 1983).

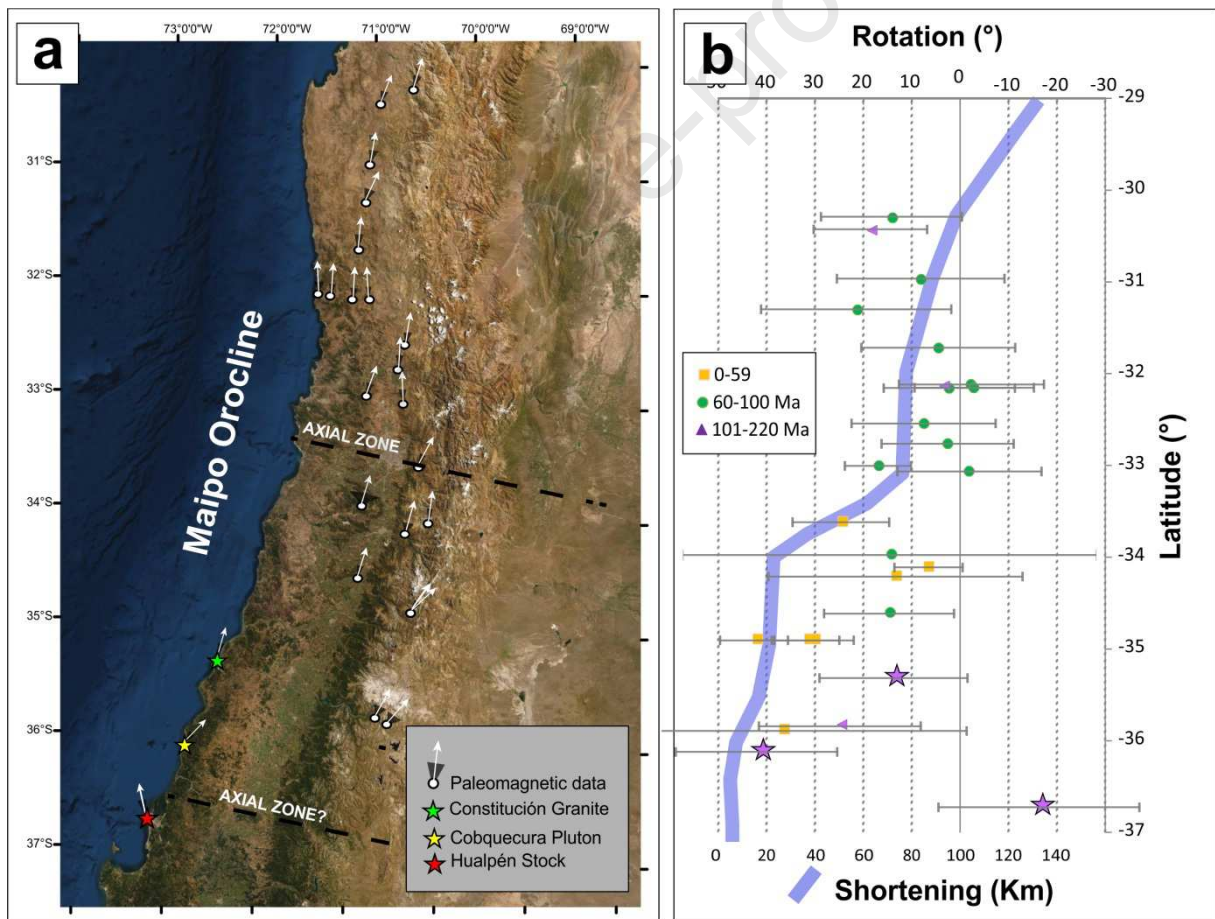
304 The data obtained indicate an increase in the magnitude of a clockwise tectonic
 305 rotation between the Constitución Granite and the Cobquecura Pluton from $\sim 13^{\circ}$ up to 41° .
 306 This is consistent with previous paleomagnetic data (Beck et al. 1986, 1990; Charrier et al.
 307 1996; Gogutchichvili et al. 2000; Parada et al. 2005; Astudillo et al. 2009; Arriagada et al.
 308 2013) (see Fig. 6). On the other hand, the Hualpén Stock recorded minor counter-clockwise
 309 tectonic rotation to no tectonic rotation. The spatial variation in the magnitude and sense of
 310 rotations shows that the Chilean Andes region between 30° and 37° S did not behave as a rigid
 311 block.

312 Previous paleomagnetic data show no significant systematic rotations in Chile between
 313 approximately 31° and 33° S (Dashwood and Taylor 2005; Peña et al. 2011; Arriagada et al.
 314 2013; Fig. 6). To the south of this last segment, a new zone of clockwise tectonic rotations is
 315 observed between approximately 33° and 36° S (see Arriagada et al. 2013 and references
 316 therein). This last area is limited to the north by the Argentinian Precordillera, where one of
 317 the orogenic maximums in Neogene horizontal shortening is observed (Fig. 6b). Horizontal
 318 shortening decreases gradually between 33° and 36° S (Giambiagi et al. 2012). Dewey and
 319 Lamb (1992) interpreted that this decrease in the horizontal shortening would be responsible
 320 for a clockwise oroclinal rotation of the trench and volcanic arc between 34° and 39° S, and
 321 the formation of the Maipo Orocline (Arriagada et al. 2009). This hypothesis is in agreement

322 with the previous paleomagnetic data (Beck et al. 1986, 1990; Charrier et al. 1996;
 323 Goguitchaichvili et al. 2000; Parada et al. 2005; Astudillo et al. 2009; Arriagada et al. 2013).

324 Japas and Re (2012) presented an alternative interpretation in which the Maipo
 325 Orocline would have been part of the Bolivian Orocline, later disconnected from northern
 326 Chile by a counter-clockwise rotation of the coastal segment between 30° and 33°S. These
 327 authors attributed this younger counter-clockwise rotation to the subduction of Juan
 328 Fernández ridge, so that the collision of the ridge would have almost entirely restored the
 329 deflection caused by the hourly rotation of the Bolivian Orocline to an orientation close to the
 330 pre-oroclinal.

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333 **Figure 6: (a) Compilation of tectonic rotations (deviation of arrows from a NS direction) with errors between 30° and**
 334 **37°S. (b) Paleomagnetic rotation colored by ages of magnetization plotted against latitude. Blue line corresponded to a**
 335 **maximum horizontal shortening according to Allmendinger et al. (1990); von Gosen (1992); Ramos et al. (1996, 2002);**
 336 **Giambiagi et al. (2012); Allmendinger and Judge (2014); Rojas Vera et al. (2014). Stars correspond to this study.**

337

338 The paleomagnetic rotations obtained in this work for the Constitución Granite and
 339 Cobquecura Pluton are similar to data reported by Arriagada et al. (2013) for a 10 Ma

340 intrusive rocks suggesting that all the area rotated as a single block during the Miocene (Fig.
341 6). Results obtained by Astudillo et al. (2009)) show that no rotations occurred since 5 Ma.
342 The large tectonic rotation observed in Cobquecura Pluton exceeds the expected tectonic
343 rotation due to horizontal shortening. This rotation can be explained as a consequence of a
344 transfer zone to accommodate the change in style and distribution of deformation between
345 Cobquecura Pluton and counter-clockwise to insignificant tectonic rotation obtained from
346 Hualpén Stock. In that case, the Maipo and Arauco Oroclines limit is located to the north of
347 the stock. However, more paleomagnetic and structural data is needed in order to corroborate
348 this hypothesis.

349

350 **6. Conclusion**

351 Paleomagnetic data obtained from the Upper Triassic Constitución Granite and
352 Cobquecura Pluton indicate an increase of clockwise tectonic rotations from north to south.
353 The similar amounts of tectonic rotations recorded by Mesozoic and Miocene rocks suggest
354 that no rotations occurred in the area during the Gondwana breakup. Considering the good
355 correspondence between the balanced section analyses with the paleomagnetic data, we
356 believe that the deflection in the trench heading and the tectonic rotations of this segment
357 respond to differential shortening north of $\sim 36^{\circ}\text{S}$. In contrast, the Hualpén granite (~ 217 Ma)
358 registers minor to insignificant counter-clockwise rotations for the southern area. This change
359 from north to south suggests that the boundary between the Maipo and Arauco oroclinal is
360 located to the north of the Hualpén granite.

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Paleomagnetic data from Upper Triassic rocks in southern Chile is presented.

Vertical-axis rotations related to Maipo Orocline formation are recorded.

Counter-clockwise rotation observed in Hualpén Stock suggests that the boundary between the Maipo and Arauco oroclines is located to the north.

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Declaration of interests

The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

The authors declare the following financial interests/personal relationships which may be considered as potential competing interests:

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