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1 **Exhumation and structural evolution of the high-elevation Malcante Range, Eastern Cordillera,**  
2 **NW Argentina**

3

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**29 Abstract**

30 As an integral part of the Eastern Cordillera, the fault-bounded Malcante mountain range (up to 5,100  
31 m) in the NW Argentine Andes (ca. 25°S) is located in the transition between the arid Puna Plateau to  
32 the west and the humid broken foreland to the east. At this latitude, the topographic gradient of the  
33 eastern Andean margin forms an efficient orographic barrier that causes pronounced east–west rainfall  
34 and surface-process gradients. In this setting, the Malcante Range is an important, yet poorly studied  
35 structural high formed during the Cenozoic topographic growth of the Central Andes. In this study, we  
36 combine (a) detailed field observations, (b) a two-dimensional structural reconstruction, (c) apatite  
37 fission track and (U-Th-Sm)/He thermochronology of bedrock samples from a vertical transect across  
38 the western flank of the Malcante Range, and (d) inverse thermal modelling using QTQt software with  
39 the aim of deciphering the exhumation history of this mountain range. Field data indicate the presence  
40 of an angular unconformity between Cenozoic foreland deposits and older sedimentary strata,  
41 suggesting an initial episode of deformation during the middle-late Eocene, while our thermal model  
42 constrains the onset of exhumation at ~10 Ma. We suggest that exhumation was related to the  
43 unroofing of the easily erodible sedimentary cover, which prevented significant initial surface uplift.  
44 This may have changed as more resilient bedrock was exposed at ~5 Ma according to the thermal  
45 model, promoting rapid rock uplift. In combination with published data, our thermochronology allows  
46 us to speculate on the existence of a zone of deformation concentrated in the area of the present-day  
47 Pasha (24.5°S), Malcante (25°S), and Agua de Castilla (25.4°S) mountain ranges by ca. 10 Ma.

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**50 Keywords**

51 Thermochronology, apatite fission track, apatite (U–Th)/He, mountain building, Eastern Cordillera,  
52 NW Argentina

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**54 1. Introduction**

55 The orogenic wedge model predicts that a pulse of rock uplift causes an instantaneous erosional  
56 response, and explains the conditions in which deformation propagates in a forward sequence at the

57 orogen scale (e.g., Davis et al., 1983; Dahlen, 1990). In a contractile setting, significant fault  
58 displacement can strongly influence the distribution of rainfall, erosion, and vegetation cover (Lenters  
59 and Cook, 1997; Bookhagen and Strecker, 2012; Burbank et al., 2012; Pingel et al., 2020). In  
60 basement-involved provinces, deformation is accommodated by a combination of reactivation of pre-  
61 existing heterogeneities and the generation of new faults (Burtman, 1975; Sibson, 1995; Holdsworth et  
62 al., 2001; Buiter and Pfiffner, 2003).

63 The Cenozoic topographic growth of the Central Andes is linked to the eastward propagation of  
64 deformation, which is a consequence of the subduction of the oceanic Nazca plate beneath the South  
65 American plate (e.g., Barazangi and Isacks, 1976; Jordan et al., 1983). The Andes in NW Argentina at  
66 ca. 25°S latitude have been divided into 3 major, roughly N-S-oriented morphotectonic provinces  
67 (Fig. 1A). From west to east these are the Puna (high elevation, low relief, arid climate), the Eastern  
68 Cordillera (high elevation, high relief, semi-arid climate), and the Santa Bárbara System-broken  
69 foreland (distributed range uplifts, humid climate) (Turner, 1979; Allmendinger et al., 1983; Kley and  
70 Monaldi, 2002;). The Eastern Cordillera forms an efficient orographic barrier for humid air masses  
71 from the Atlantic and the Amazon, which causes strong across-strike climatic and vegetational  
72 gradients with high rainfall amounts and dense vegetation (Yungas, Fig. 1B) along its eastern flanks  
73 and semi-arid to arid conditions in the Andean hinterland that is also reflected in the generally  
74 decreasing surface process rates from east to west (Fig. 1A).

75 The orogen-scale propagation of Cenozoic deformation formed new faults and caused the  
76 reactivation of previously-formed structures between the Puna and the Santa Bárbara System (e.g.,  
77 Allmendinger et al., 1983, Hongn et al., 2007; Carrapa and DeCelles, 2008; Payrola et al., 2012;  
78 Pearson et al., 2013; Reiners et al., 2015). Numerous studies have focused on whether this  
79 deformation has propagated in- or out-of-sequence (e.g., Pearson et al., 2013; Zhou et al., 2017;  
80 Payrola et al., 2020), and to what extent fault reactivation is promoted by basement heterogeneities  
81 (Hongn et al., 2010; Pearson et al., 2013; Payrola et al., 2020; Zapata et al., 2020). The main  
82 precursors of out-of-sequence deformation are: 1) fault-properties (e.g., fault orientation, fluid  
83 pressure, friction coefficient); 2) shortening rate; 3) lithology and rock rheology; and 4) the state of  
84 lithostatic stress caused by tectonic loading or erosional unloading (e.g., Beaumont et al., 1992;

85 Willett, 1999; Hilley et al., 2005; Pingel et al., 2013; Ballato et al., 2019). As this deformation  
86 propagates eastward, the growth of mountain ranges formed orographic barriers that focussed erosion  
87 and caused a W-E gradient in rainfall (e.g., Pingel et al., 2014; Pingel et al., 2019).

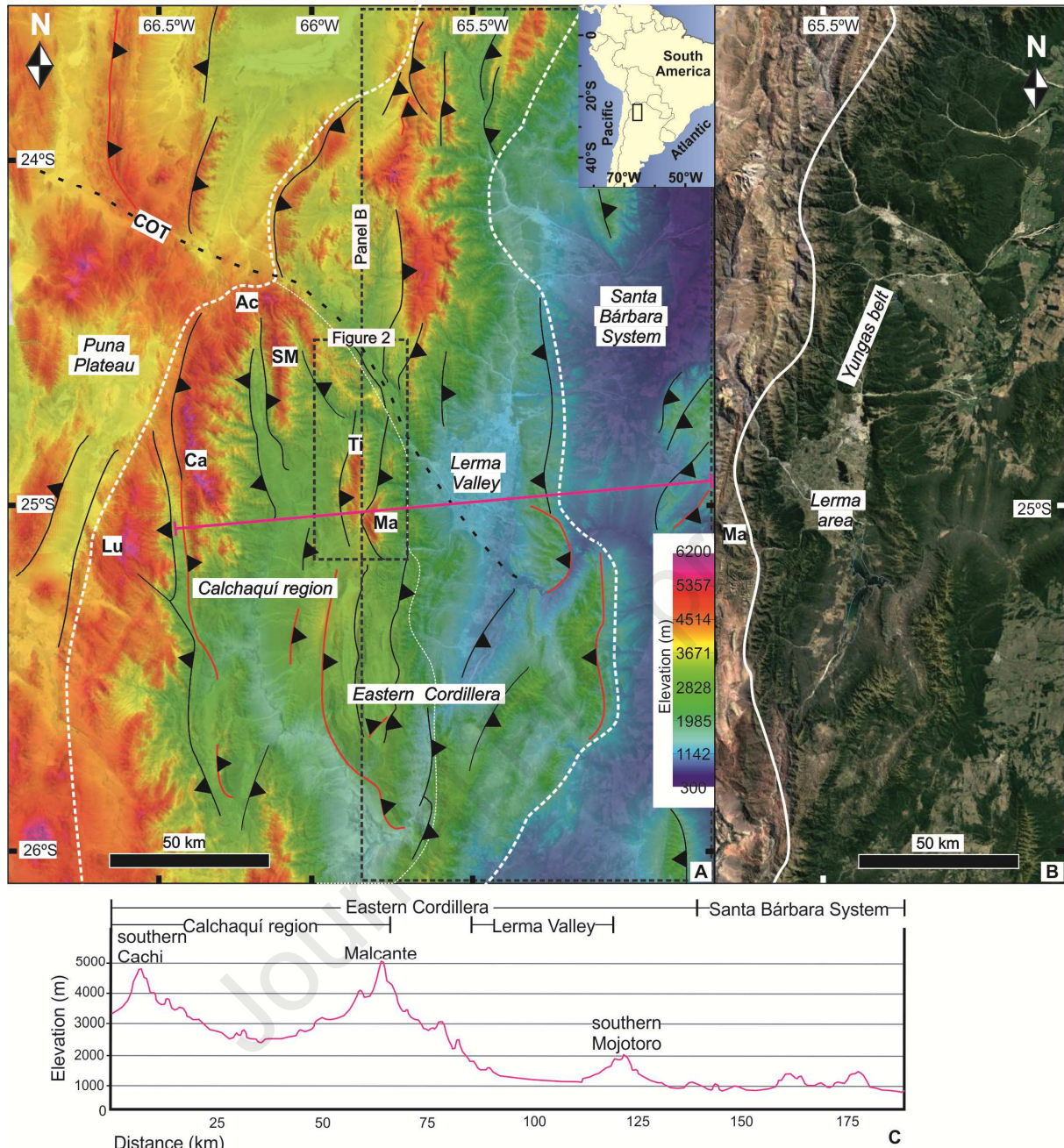
88 Since the Paleogene, deformation in the present-day Eastern Cordillera and parts of the eastern  
89 Puna Plateau has been localized by the reactivation of Cretaceous rift NNW-SSE-trending structures  
90 that root in the Neoproterozoic-Lower Cambrian basement (Fig. 1A). Under a compressional stress  
91 regime, reactivation of Cretaceous normal faults with a thin sedimentary cover (<5,000 m) leads to  
92 high-angle basement reverse faults, and hence, thick-skinned deformation (Grier et al., 1991; Hongn  
93 and Seggiaro, 2001; Carrera and Muñoz, 2013). The first documented compressional event occurred  
94 in the middle-late Eocene, marked by local angular unconformities, growth-strata, deposition of  
95 coarsening upward alluvial successions, and AHe and AFT cooling ages along the Puna-Eastern  
96 Cordillera margin (Hongn et al 2007; Payrola et al 2009; Pearson et al., 2012). Oligocene to early  
97 Miocene compressional events are well-documented in the Puna-Eastern Cordillera margin as well as  
98 in the interior of the Eastern Cordillera by local unconformities, growth-strata, seismites, and AFT  
99 ages (e.g., Deeken et al., 2006; Payrola et al., 2020; Espinoza et al., 2020). Middle-late Miocene to  
100 Pliocene compressional events created high topography in the interior of the Eastern Cordillera  
101 according to structural and cooling ages data (Carrera and Muñoz, 2008; Hain et al., 2011; Carrapa et  
102 al., 2011; Pearson et al., 2013). Finally, Quaternary compressional events modified the landscape and  
103 produced large alluvial fan conglomerates (Bookhagen and Strecker, 2012; McCarthy et al., 2019).  
104 However, the ages, rates and mechanisms of particularly each range uplift that led to the present-day  
105 topography in the area remain unresolved.

106 The topography of the Eastern Cordillera shows two main steps (Figs. 1A and C): from 1,200 m  
107 mean elevation in the eastern Lerma Valley, including 4,200 m peaks, to 2,500 m mean elevation in  
108 the western Calchaquí region, including 5,100 m peaks and minors ranges of 3,000 m in intermediate  
109 zones. This topography is partially controlled by reverse fault displacements and by a W-E gradient in  
110 rainfall and erosion (Bookhagen and Strecker, 2012; Pingel et al., 2020).

111 The Calchaqui region, located in the Eastern Cordillera morphotectonic province, comprises  
112 several high-elevation ranges in excess of 5,000-6,000 m (e.g., Cachi 6,300, Luracatao 5,900, Acay

113 5,700, San Miguel 5,700 m, Tirao 5,000, Malcante 5,200, high mountains, Figs. 1A and C). These  
114 ranges are integrated by a pre-Cenozoic basement composed by igneous (Luracatao), medium grade  
115 metamorphic (Cachi) or very low grade metamorphic rocks (to the east of Cachi Range), and are  
116 generally bounded by basement-involved, high-angle reverse faults with kilometeric-scale  
117 displacements that show a complex deformation history (Hongn and Seggiaro, 2001; Pearson et al.,  
118 2012; Payrola et al., 2020). Moreover, lower mountain ranges could have absorbed less shortening  
119 than higher mountains at the latitudinal regional scale during the Cenozoic multi-episode of fault-  
120 reactivation (e.g., Pearson et al., 2013; Payrola et al., 2020). A particularly interesting topographic  
121 feature is the more than 5,000-m high Malcante Range, located on the eastern border of the semi-arid  
122 Calchaquí region at ca. 25°S (Figs. 1A and 2). The Malcante Range forms the orographic barrier that  
123 limits the heavily vegetated Yungas belt (Fig. 1B), where strong convective storms discharge heavy  
124 rains (Hongn and Seggiaro, 2001; Salfity, 2004; Bookhagen and Strecker, 2012). This range also  
125 separates the Calchaquí Cenozoic Basin to the west, from the Lerma Valley Basin to the east (Fig. 2);  
126 the latter is located in the broken foreland of NW Argentina (Hain et al., 2011; Pearson et al., 2013).  
127 In this study, we analyze the structural geometry of the Cenozoic Malcante Range to understand the  
128 details of its formation. To achieve this, we developed a 2D kinematic structural model based on field  
129 observations and evaluate the exhumation history using low-temperature apatite (U-Th-Sm)/He and  
130 fission track thermochronology. Moreover, we use inverse thermal modelling to recognize rapid  
131 cooling episodes and interpret the timing of exhumation of the Malcante Range and compare it with  
132 the history of other ranges in the Central Andes. The results improve our understanding of the  
133 Miocene orogenic growth, which represents the main phase of deformation in the Central Andes  
134 (Jordan et al., 1983; Grier et al., 1991; Coutand et al., 2001; among others).





135  
 136 **Fig. 1. A.** Digital elevation map (JAXA/METI 2010) highlighting regional faults. Cenozoic reverse  
 137 faults are shown by black lines and Cretaceous inverted faults are shown by red lines (based on Hongn  
 138 and Seggiaro, 2001; Salfity and Monaldi, 2006). Thick dashed white lines outline morphotectonic  
 139 units (after Ramos, 1999) and thick dashed black line outlines the Calama-El Toro Lineament (COT).  
 140 Specific high mountains within the Calchaquí Region: Ca-Cachi, Lu-Luracatao, Ac-Acay, SM-San  
 141 Miguel, Ti-Tirao, Ma-Malcante. Inset shows location of the study area within South America. **B.**  
 142 Google Earth image of the eastern border of the Calchaquí region, shown by white solid line, and the  
 143 heavily vegetated Yungas green belt. Note that the generally lighter tones west of the Calchaquí  
 144 border reflect less vegetation due to more arid conditions; darker tones farther east reflect much more  
 145 abundant vegetation linked to higher mean annual precipitation.

## 146 147 2. Geological setting

148 The Calchaquí region is located in the southernmost part of the Eastern Cordillera at 65.7°W  
 149 longitude, between 24.5°S and 26.5°S latitudes (Fig. 1). Here, a complex geological history is



150 documented by Neoproterozoic-to-Lower-Cambrian basement rocks and a sedimentary cover  
151 recording Cretaceous rifting and Cenozoic foreland deposits (Fig. 2A, Hongn and Seggiaro, 2001).  
152 The Calchaquí region constitutes a series of approximately north-south oriented intermontane basins  
153 and intervening ranges related to regional faults and folds with the same strike (Salfity, 2004; Salfity  
154 and Monaldi, 2006). The regional structural style is characterized by a thick-skinned fold-and-thrust  
155 belt formed by mostly west-vergent and subordinate east-vergent faults, and the Cenozoic inversion of  
156 NW-SE-oriented Cretaceous normal faults (Hongn and Seggiaro, 2001, Carrera and Muñoz, 2013).  
157 Many of the major folds comprise basement-cored ranges with a passively-folded sedimentary cover  
158 (i.e., drape folds, e.g., Carrera et al., 2009; Payrola et al., 2012, Hernández et al., 2016).

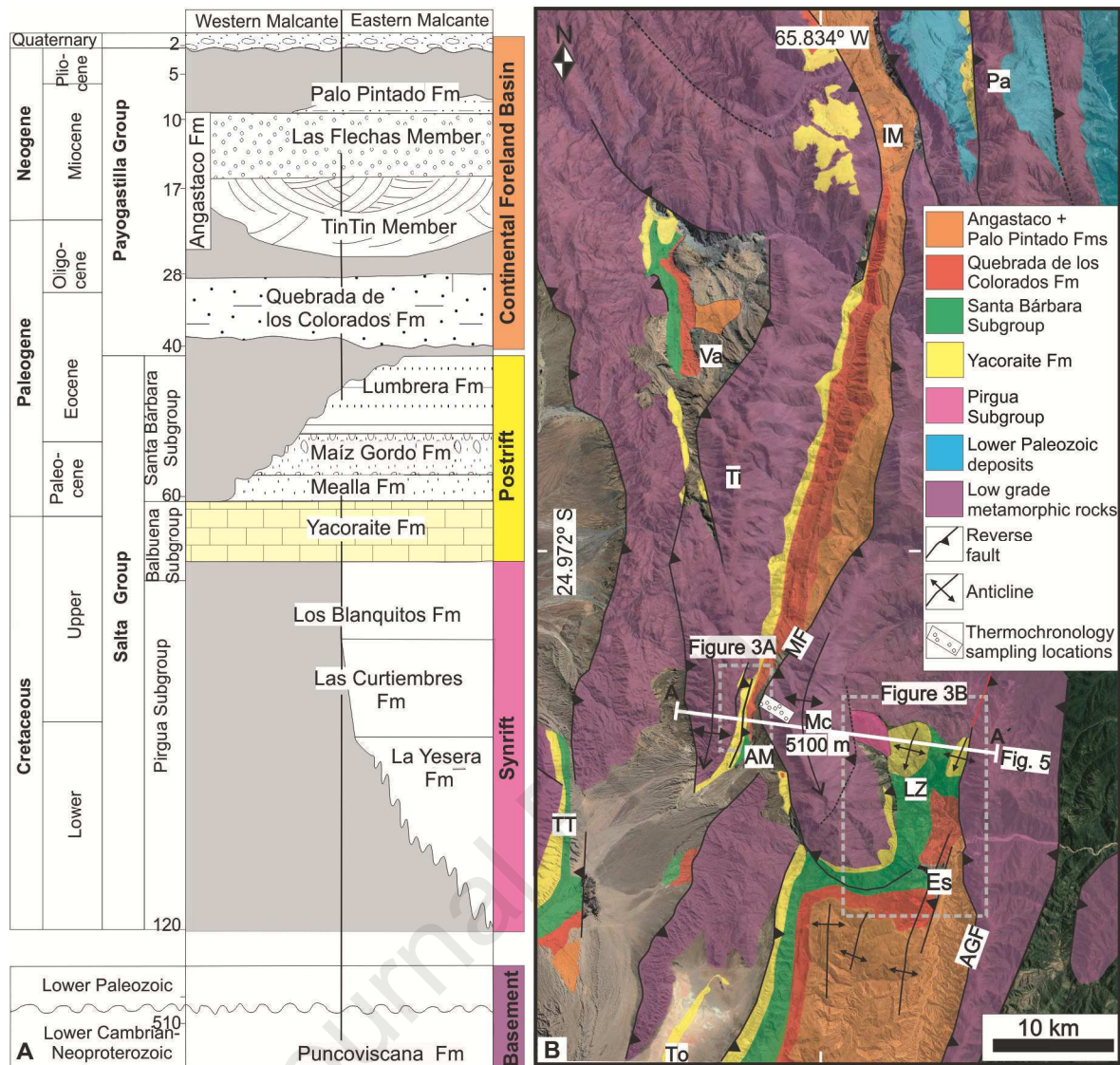
159 The regional Neoproterozoic-Lower Cambrian basement is mainly composed of low-grade  
160 metamorphic rocks of the Puncoviscana Formation (Turner, 1960; Escayola et al., 2011; Do Campo et  
161 al., 2013) with granitic intrusions in the western zone of the Cachi Range (Hongn and Seggiaro,  
162 2001). These rocks are pervasively folded in various wavelengths ranging between  $10^{-2}$  and  $10^2$  m,  
163 with a predominantly N-S-oriented axial direction concordant with the main foliations and the  
164 orientation of ductile shear zones in the region (Hongn and Seggiaro, 2001; Riller and Hongn, 2003;  
165 Hernández et al., 2016).

166 The basement is overlain by the Cretaceous to Paleogene Salta Group (Turner, 1959), which has  
167 been subdivided into the syn-rift Pirgua Subgroup and post-rift Balbuena and Santa Bárbara subgroups  
168 (Fig. 2A, reviewed in Marquillas et al., 2005). The synrift deposits consist of conglomerate, sandstone,  
169 and mudstone deposited in proximal to distal positions of alluvial fans to fluvial systems with volcanic  
170 intrusions in the lower and middle sections (Marquillas et al., 2005). Post-rift sediments are  
171 characterized by lacustrine to shallow marine carbonates of the Yacoraite Formation and mainly  
172 sandstones of fluvial origin (Marquillas et al., 2005).

173 These units are unconformably overlain by the up to 6,000-m thick Payogastilla Group (Díaz and  
174 Malizzia, 1983; Jordan and Alonso, 1987; Hongn et al., 2007), composed of an Eocene to Pleistocene  
175 assemblage of fluvial, alluvial, and eolian deposits (Fig. 2A), which document the transition from  
176 unconstrained foreland deposition to the present-day restricted intermontane basin deposition (Díaz  
177 and Malizzia, 1983; Coutand et al., 2006; del Papa et al., 2013a; Pingel et al., 2016). From bottom to

178 top, the Payogastilla Group has been subdivided into four litho-stratigraphic units (Fig. 2, Díaz and  
179 Malizzia, 1983; del Papa et al., 2013b): I) the Quebrada de los Colorados Formation (ca. 40–28 Ma,  
180 Payrola et al., 2009; DeCelles et al., 2011); II) the Angastaco Formation comprising the Tin Tin (ca.  
181 28–15 Ma, del Papa et al., 2013a; Payrola et al., 2020) and Las Flechas (ca. 15–9 Ma, Carrapa et al.,  
182 2011; del Papa et al., 2013b; Pingel et al., 2016; Aramayo et al., 2017) members; III) the Palo Pintado  
183 Formation (ca. 9–5.2 Ma, Coutand et al., 2006; Bywater-Reyes et al., 2010; Pingel et al., 2016); and  
184 IV) the San Felipe Formation (ca. 5.2 to <1.9 Ma, Bywater-Reyes et al., 2010; Pingel et al., 2016). All  
185 these units are covered by coarse-grained Quaternary alluvial fans and fluvial sediments that  
186 characterize the present-day Calchaquí region (e.g., Strecker et al., 2007; Pingel et al., 2016).

187 The Malcante Range is located on the border between the Calchaquí region and the Lerma Valley  
188 (Figs. 1, 2B) and it is a high amplitude basement anticline composed primarily of the Puncoviscana  
189 Formation, bounded on the west by the Malcante fault. To the east of the Malcante Range in the Las  
190 Zanjás area, a minor anticline-syncline-anticline fold-train is defined by the Yacoraite Formation and  
191 the strata of the Santa Bárbara Subgroup and is truncated to the east by the Agua de Castilla fault (Fig.  
192 2B).



193  
194 **Fig. 2. A.** Stratigraphy of the western flank of Malcante Range (AM) compared to the eastern flank  
195 (LZ) showing the stratigraphic position and lateral variability of stratigraphic units (based on  
196 Marquillas et al., 2005; del Papa et al., 2013b), **B.** Simplified geological map, highlighting the position  
197 of Cenozoic foreland deposits of the Eocene to Pleistocene Payogastilla Group in the Abra Malcante  
198 (AM) and Las Zanjias (LZ) areas (modified after Vergani and Starck, 1989). For location, see the  
199 white box in Figure 1. Specific areas and faults within the Calchaquí Region: TT-Tin Tin; To-Tonco;  
200 Es-Escoipe; Va-Vallecitos; IM-Ingeniero Mauri, Pa-Pascha; MF-Malcante fault; AGF-Agua de  
201 Castilla Fault. Legend code in the panel B, Fm: Formation. Lithology codes: Yacoraite Fm:  
202 limestones, Mealla Fm: sandstones, Maíz Gordo Fm: conglomeratic sandstones, Lumbrera:  
203 sandstones, Quebrada de los Colorados: sandstones and mudstones, Tin Tin Member: eolian  
204 sandstones, Las Flechas: conglomerates and sandstones, Palo Pinado: mudstones.  
205  
206

### 207 3. Methods

#### 208 3.1 Sampling and structural reconstruction

209 We carried out detailed structural and geological mapping (1:10.000) along the western and eastern  
210 flanks of the Malcante mountain range, in the Abra Malcante and Las Zanjias areas, respectively (Fig.

211 3, 4 and 5). Lithostratigraphic units were identified by regional features and their stratigraphic  
212 contacts were mapped using contrasting lithofacies assemblages, changes in stratigraphic stacking  
213 patterns, and regional structural relationships. Our new data were supplemented by published  
214 sedimentological, structural and chronological data from adjacent areas (Vergani and Starck, 1989;  
215 González Villa, 2002; Salfity and Monaldi., 2006; Payrola et al., 2020).

216 A structural cross-section was constructed (Fig. 5) using Move software 2015.2 (academic license),  
217 based on measured field structural data projected onto a W-E transect (Fig. 2). The reconstruction of  
218 the Malcante fold utilizes a fault-propagation fold mechanism, modified by trishear deformation  
219 (Erslev, 1991, Hongn and Seggiaro, 2001; Payrola et al., 2020).

### 220 *3.2 Thermochronological methods*

221 Apatite (U-Th-Sm/He) thermochronology (AHe) is based on measuring the amount of helium  
222 produced by the radiogenic decay of  $^{238}\text{U}$ ,  $^{235}\text{U}$ ,  $^{232}\text{Th}$ , and  $^{147}\text{Sm}$  (e.g., Zeitler et al., 1987; Flowers et  
223 al., 2009). Apatites quantitatively retain helium at temperatures below  $\sim 40^\circ\text{C}$ . Between ca. 40 and  
224  $80^\circ\text{C}$ , some helium is lost by diffusion; this interval is known as the PRZ-Partial Retention Zone  
225 (Wolf et al., 1998; Farley, 2000). The limits of the PRZ depend on cooling rate, grain shape and size,  
226 and the amount of alpha damage in the mineral lattice, which increases the apatite helium retentivity  
227 (e.g., Flowers et al., 2007; Gautheron et al., 2009; Brown et al., 2013). Damage effects can be  
228 identified by a correlation between effective uranium ( $e\text{U}=\text{U} + 0.235*\text{Th}$ ) and age (Flowers et al.,  
229 2007; Flowers, 2009; Gautheron et al., 2009). AHe analyses (Table 1) were carried out at Potsdam  
230 University (alphachron) and in the German Research Center for Geoscience (GFZ) (ICP-MS). Details  
231 of the analytical methods are presented in Zhou et al. (2017) and in the supplementary material.

232 Apatite fission track (AFT) thermochronology is based on the quantification of mineral lattice  
233 damage (fission tracks), which are formed by the spontaneous fission of  $^{238}\text{U}$  (e.g., Wagner et al.,  
234 1989). Tracks within the apatite mineral lattice can be partially annealed at temperatures between 60  
235 and  $\sim 120^\circ\text{C}$ , this interval is known as the Partial Annealing Zone (PAZ) (e.g., Fitzgerald et al., 1986).  
236 Resistance to annealing depends on mineral kinetics and composition. The length dimension of the  
237 intersection between the polished mineral surface and the tracks ( $D_{\text{par}}$ ) can be used as a proxy for the

238 mineral resistance to annealing (Donelick et al., 1999; Ketcham et al., 1999). AFT analyses were  
239 conducted at the Thermochronology Laboratory La.Te.Andes (Salta, Argentina, Table 2). Track  
240 measurements were performed using a binocular microscope Zeiss® AXIO Imager Z2m and the  
241 Autoscan® software TrackWorks®. Unfortunately, very few confined track lengths were identified,  
242 and thus track lengths were not measured. Ages were calculated by the external detector method  
243 (Huford and Green, 1982, 1983; Wagner and van den Haute, 1992), and the data was processed using  
244 Trackkey® Software (Dunkl, 2002). Analytical details are provided in the supplementary data.

245 Thermochronology samples were collected from bedrock (Puncoviscana Formation) along a  
246 vertical transect of the western flank of the Malcante Range (for location, see Fig. 2 and Table 1). Due  
247 to the high degree of deformation present in the Puncoviscana Formation, structural and stratigraphic  
248 markers are not clear. Seven rock samples, each composed of three to five kilograms, were collected  
249 from 4,045 to 5,100 m a.s.l., according to the location of good outcrops.

### 250 *3.3 Thermal modeling parameter and procedures*

251 Rocks located at different positions within a coherent rock body experience similar thermal  
252 histories, where the temperature at different positions depends on the thermal gradient within the rock  
253 body (Gallagher et al., 2005). Temperature variations with depth are caused by the regional  
254 geothermal gradient; this may be perturbed locally by magmatism (e.g., Murray et al., 2018; Zapata et  
255 al., 2019b). In steep mountain belts formed by contractile-driven rock uplift, rocks at higher elevation  
256 experienced relatively lower temperatures compared to rocks at lower elevations throughout the  
257 thermal history, resulting in a positive relationship between elevation and cooling age (e.g., Fitzgerald  
258 and Malusà, 2019). Therefore, modern elevation can often be used as a proxy for the paleodepth.  
259 Herein, we present a multi-sample thermal history model from samples collected at different  
260 elevations along the Malcante Range. This model presents a single form of the thermal history based  
261 on samples collected in a contiguous basement block.

262 Thermal modelling was performed using QTQt software (v. 5.7.0 Gallagher, 2018), which uses a  
263 Bayesian transdimensional Markov chain Monte Carlo statistical approach to obtain the most probable  
264 thermal histories from multiple samples analyzed with thermochronological methods (Gallagher,  
265 2012). The total number of iterations used in the sampling chain is equal to the burn-in + post-burn-in.



266 The burn-in is the number of initial iterations used to explore the model space; these initial  
267 interactions are discarded. After the burn-in stage, the post-burn-in is characterized by the number of  
268 iterations used in subsequent inferences of the thermal history. Several initial runs were performed in  
269 order to calibrate the model and select the best temporal intervals. Afterwards, model iterations were  
270 progressively increased until acceptance rates were below 0.5 and a stable plateau was observed in the  
271 likelihood chain. After 150,000 initial iterations (burn-in stage), the model was run for 350,000  
272 iterations (post-burn-in stage).

273 We use the radiation damage model (RDAAM) from Flowers et al. (2009) for the AHe data and  
274 the annealing model from Ketcham et al. (2007) for the AFT data. Reproducible AHe ages were  
275 included in the model. We consider AHe ages to be reproducible when the 1 sigma interval of the  
276 distribution is less than 20% of the mean age (e.g., Flowers and Kelley, 2011; Zapata et al., 2019b).  
277 AHe single grain ages that fit the age-elevation trends and have possible eU controls were also  
278 included in the model despite not being reproducible. Such aliquots could have different closure  
279 temperatures and therefore may have significantly different ages. As a result, all aliquots were  
280 modeled because none of them was identified as an evident outlier following these criteria.

281 Geologically meaningful time-temperature constraints were included to improve the model output.  
282 To start the model with fully reset samples, we set the first constraint box above the temperature range  
283 of the PAZ at 580-520 Ma between 150 °C and 210 °C, the age of metamorphism of the Puncoviscana  
284 Formation (Hongn and Seggiaro, 2001; Aparicio González et al., 2010; Do Campo et al., 2013). To  
285 account for the prominent unconformity between the basement and the Cretaceous sediments of the  
286 Pirgua Subgroup, we set a surface temperature constraint between 80 and 60 Ma. The present-day  
287 surface temperature was set to  $10 \pm 10$  °C.

288 We present the expected thermal history model, which is the weighted average of the accepted  
289 models, the expected model predicted ages, and the 2-sigma interval of the ages predicted by the  
290 accepted models. We allowed the model to have pre-Cenozoic thermal histories ( $70 \pm 70$  Ma) because  
291 one of the AFT ages is partially reset and to allow the model to consider radiation damage effects to  
292 better constraint single grain AHe closure temperatures. The chain of likelihood, the data used as



293 constraints, and the parameters incorporated in the thermal models (Fig. 7) are presented in the  
294 supplementary data.

295

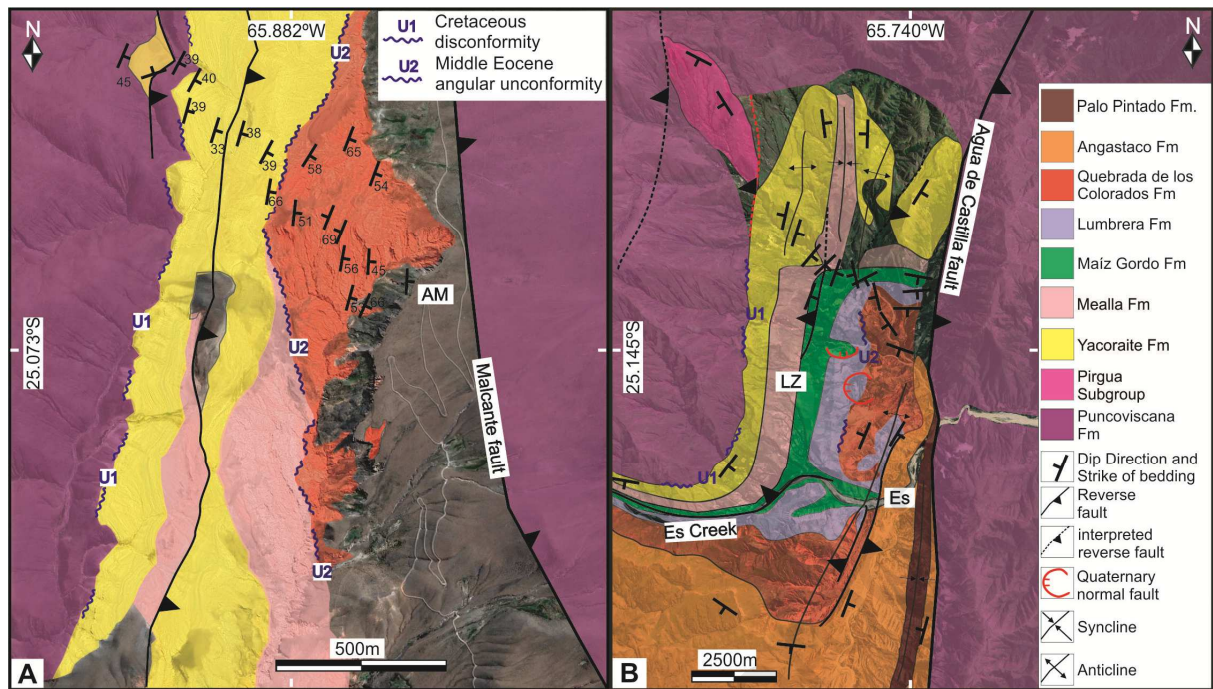
## 296 **4. Results**

### 297 *4.1 Structural inventory*

298 Field-based mapping and remote sensing analysis reveal that the flanks of the Malcante Range  
299 record complex structural relationships (unconformities, inverted strata, pinch outs, among others,  
300 Figs. 3 and 4). The field data support the interpretation that this range is an approximately north-south  
301 trending, asymmetrical basement-cored anticline bounded by the Tirao anticline and the Las Zanzas  
302 folds, to the west and east, respectively (Fig. 2). Its steeply southward plunging fold axis is well  
303 defined by the prominent limestones of the Yacoraite Formation, which were passively folded  
304 together with the underlying Puncoviscana bedrock (Fig. 4A). The fold has a width of 13,000 m, 5,000  
305 m of vertical amplitude and an axial length of 27,000 m (Fig. 2), which makes it one of the largest  
306 basement anticlines in the Calchaquí area and the Lerma Valley to the east (Figs. 1 and 2).

307 The western limb of the Malcante Range is bounded by the west-vergent Malcante reverse fault,  
308 which uplifts the Puncoviscana basement over the Quebrada de los Colorados formations (Figs. 2 and  
309 5). This fault has a curved trace to the south and continues towards the Escoipe creek with decreasing  
310 displacement (Fig. 3B).

311 Several outcrops within the footwall of the Malcante fault show a pronounced erosional  
312 unconformity between the Middle-Late Eocene Quebrada de los Colorados Formation and the  
313 Paleocene Mealla Formation (Santa Bárbara Subgroup) to the south (Abra Malcante) or between the  
314 Quebrada de los Colorados Formation and Yacoraite Formation (Balbuena Subgroup), north of the  
315 Abra-Malcante site (Figs. 3A and 4B). In addition, the Yacoraite Formation overlaps the  
316 Puncoviscana Formation in the eastern limb of the Tirao anticline. The only occurrence of Lower  
317 Cretaceous syn-rift strata of the Pirgua Subgroup is found within the eastern limb of the Malcante  
318 anticline in the Las Zanzas area (Fig. 3B).



319  
 320 **Fig. 3.** Geological maps of the western limb (AM-Abra Malcante) and the eastern limb (Las Zanzas)  
 321 of the Malcante anticline. **A.** Detailed map showing for the first time the unconformity (in red colour)  
 322 between the Quebrada de los Colorados Formation and the Mealla Formation (U1) and the Yacoraite  
 323 Formation (U2) in the Abra Malcante. **B.** Detailed map showing the complex structure of a thin belt of  
 324 Pirgua syn-rift deposits, the post-rift deposits and the foreland Payogastilla Group in the Las Zanzas  
 325 (LZ) and Escoipe (Es) areas.  
 326

327 The reconstruction of structures in the Las Zanzas area is challenging due to several reverse faults  
 328 that displace the stratigraphy and because the outcrops are often covered by young deposits (Figs. 3B,  
 329 4C and D). However, we identified an anticline-syncline-anticline fold train. Moreover, we observed  
 330 an angular unconformity that marks the contact between the Quebrada de los Colorados Formation  
 331 and the Lumbrera Formation (Fig. 3B). Minor low-angle reverse faults, with east and west vergence,  
 332 produce minor folds in the ductile levels of the Yacoraite and Maíz Gordo Formations (Fig. 4C).  
 333 Finally, at the eastern border of the study zone, the west-vergent Agua de Castilla reverse fault  
 334 uplifted the Puncoviscana Formation over the post-rift and Angastaco and Palo Pintado Formations  
 335 foreland deposits (Fig. 3B). The Malcante basement block continues to the north up to the NW-SE-  
 336 trending El Toro regional lineament and then continues to the Pascha Range. There, outside of the  
 337 study area, multiple west-vergent reverse faults uplift the Puncoviscana Formation and lower  
 338 Paleozoic strata (Fig. 2).

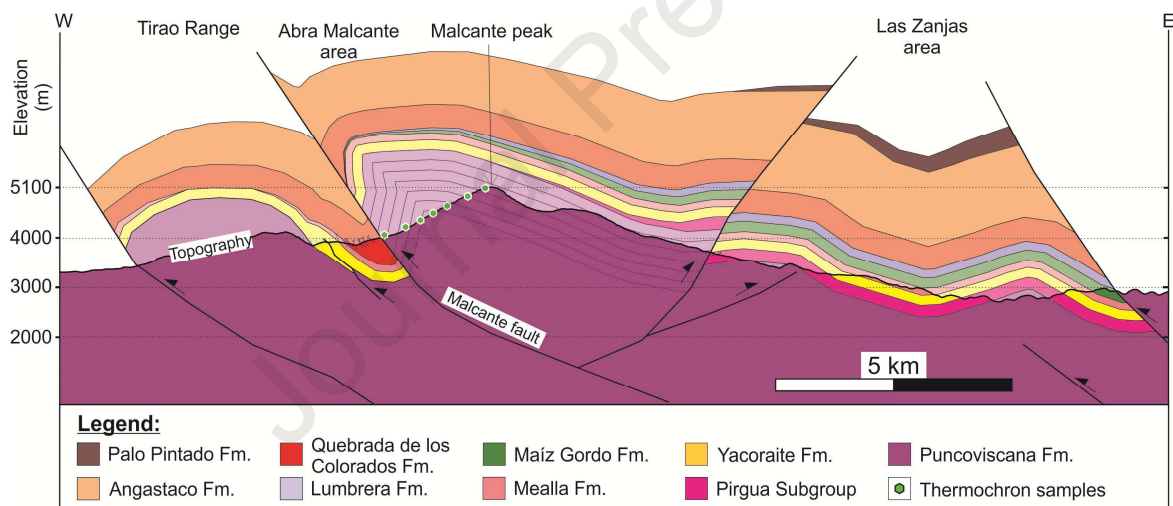




339  
 340 **Fig. 4.** A. Angular unconformity (U) between the Yacoraite Formation (Ya) and the Puncoviscana  
 341 Formation (PC) in the footwall of the Malcante fault. B. Low angle unconformity (U) between the  
 342 Quebrada de los Colorados Formation (QC) and the Mealla Formation (Me) in the Abra Malcante  
 343 area. C. Minor reverse fault that duplicates the Yacoraite Formation in the northern part of the Las  
 344 Zanjias area. D. Northeastern regional view of outcrops located between the eastern limb of the  
 345 Malcante anticline and the Agua de Castilla reverse fault in the Las Zanjias area. Abbreviations: Lu-  
 346 Lumbrera Formation, Ang-Angastaco Formation, PP-Palo Pintado Formation.



347  
 348 Based on field mapping, we reconstructed the possible fold geometry of the Malcante anticline,  
 349 revealing a steep frontal limb and low-angle back-limb (Fig. 5). The geometry of the sedimentary  
 350 cover allows us to project the dip domain to the basement in order to show the potential geometry of  
 351 the anticline core. The cover was drawn in Fig. 5 with thickness intervals of 200 m. According to our  
 352 structural reconstruction (Fig. 5), the Malcante reverse fault has a minimum of 1,400 m of reverse  
 353 displacement. The sedimentary succession is thicker in the Las Zanjias area than the Abra Malcante  
 354 area because to the assumption that there are unconformities over the crest of the Malcante anticline,  
 355 consistent with unconformities described in this study and the regional data (Hongn et al., 2007;  
 356 Carrera and Muñoz, 2013; Carrapa et al., 2011; Aramayo et al., 2017, Payrola et al., 2020). Details of  
 357 the parameters used for the reconstruction are presented in the supplementary data.  
 358



359  
 360 **Fig. 5.** West-east cross-section of the Malcante anticline and the Las Zanjias fold-train with the  
 361 location of the samples analyzed by AHe and AFT (thermochron samples). Reconstruction based on  
 362 field structural mapping using Move software. Erosion of the Santa Bárbara Subgroup in the crest of  
 363 the anticline is based on the unconformity in the footwall of Malcante fault, and thinning of the  
 364 foreland succession without Palo Pintado deposition over the crest of the anticline are assumed based  
 365 on the AFT and AHe ages. Black lines represent reverse faults.  
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#### 369 4.2 AHe and AFT results

370 We obtained 23 AHe single grain ages from seven bedrock samples from a vertical profile between  
 371 4,045 and 5,100 m elevation. AHe single grain ages range between ~4 and 10 Ma, with eU values  
 372 between ~1 and 84 ppm, with the exception of one with 173 ppm, and equivalent spherical radius

373 (ESR) between ~36 and 131  $\mu\text{m}$ . We obtained AFT ages from five of the samples. The four  
 374 lowermost samples (PC1.2, PC2, PC2.1, and PC3.1) pass the Chi-square test and have overlapping  
 375 ages between ~ 8 and 11.5 Ma and Dpar values between 1.8 and 2.1. Sample PC1.1, which is the  
 376 highest of the samples with AFT data, did not pass the Chi-square test. This sample exhibits two  
 377 discrete AFT grain-age populations, with central ages of 15.8 and 133.0 Ma. Analytical results for  
 378 individual AHe single grain aliquots are presented in Table 1 and sample AFT data in Table 2.

379 **Table 1:** AHe data from the Malcante anticline

Sample	H (m)	Lat ( $^{\circ}$ S)	Long ( $^{\circ}$ W)	Stratigraphic unit	Stratigraphic age (Ma)	Grain	Raw Age (Ma)	AHe Age (Ma)	$\pm 1\sigma$ (Ma)	U (ppm)	Th (ppm)	147Sm (ppm)	eU	He (nmol/g)	Ft	ESR ( $\mu\text{m}$ )	Term.	
PC1	5100	25.07458	65.85204	Puncoviscana	600-530	PC1-a	6.4	10.2	0.4	60	68	19	76	2.62	0.63	40	1	
						Fm	PC1-b	3.5	5.1	0.5	9	22	32	14	0.27	0.68	47	2
						PC1-c	4.1	6.8	0.4	27	65	12	42	0.95	0.61	38	1	
						PC1-d	4.2	6.5	0.3	29	34	41	37	0.84	0.64	42	1	
						PC1-e	4.3	7.1	0.3	53	58	114	67	1.58	0.61	39	1	
PC1.1	4900	25.07174	65.8569	Puncoviscana	600-530	PC1.1a	6.6	10.1	0.4	66	76	53	84	3.01	0.66	44	1	
						Fm	PC1.1b	5.0	8.1	0.5	29	19	60	34	0.91	0.61	39	1
						PC1.1c	4.2	6.5	2.3	2	17	8	6	0.13	0.64	41	1	
						PC1.1d	4.9	7.4	0.3	63	89	55	84	2.26	0.67	45	0	
PC1.2	4740	25.07195	65.8594	Puncoviscana	600-530	PC1.2a	3.8	6.2	1.5	7	3	36	8	0.17	0.62	39	1	
PC2	4551	25.07416	65.864	Puncoviscana	600-530	PC2-a	4.2	5.9	0.3	27	18	45	31	0.72	0.71	52	1	
						Fm	PC2-b	4.1	4.6	0.1	7	2	1	7	0.16	0.88	122	1
PC2.1	4405	25.07115	65.8656	Puncoviscana	600-530	PC2-c	3.4	4.5	0.2	19	5	10	20	0.36	0.76	62	0	
						Fm	PC2-e	3.7	5.3	2.8	1	6	1	2	0.04	0.70	49	1
						PC2.1a	3.9	6.1	0.4	41	124	25	70	1.47	0.63	41	0	
PC3	4257	25.07061	65.8694	Puncoviscana	600-530	PC3-a	3.3	5.4	1.6	6	7	2	7	0.13	0.62	39	1	
						Fm	PC3-b	5.0	7.0	0.1	30	146	14	65	1.78	0.73	55	1
						PC3-c	2.9	4.2	0.4	18	36	4	27	0.41	0.68	47	0	
						PC3-e		3.6	2.6	4	10	4	7	0.07	0.57	35	0	
PC3.1	4045	25.07572	65.8738	Puncoviscana	600-530	PC3.1a	7.0	9.2	0.5	3	13	6	6	0.24	0.76	62	0	
						Fm	PC3.1b	4.1	6.1	0.1	134	166	29	173	3.80	0.67	45	0
						PC3.1c	3.9	6.4	0.4	36	95	65	59	1.26	0.61	39	0	
						PC3.1d	3.9	6.6	1.0	19	9	32	21	0.46	0.59	37	0	

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**Notes:** Abbreviations: H, elevation; Lat, Latitude; Long, Longitude, AHe age, Corrected age, eU, effective Uranium, ESR, equivalent spherical radius, Term, Crystal terminations.

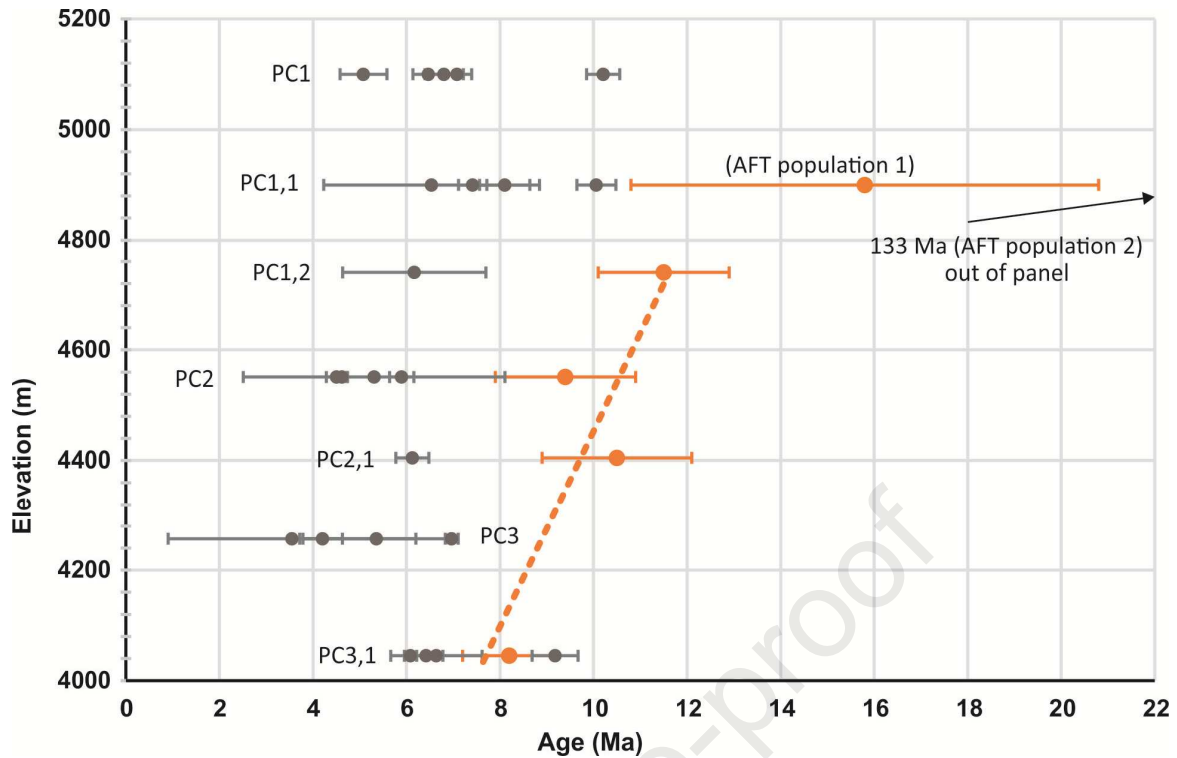
400 **Table 2:** AFT data from the Malcante anticline

Sample	H (m)	Lat (° S)	Long (° W)	n	<i>Rho</i> D	ND	Mean Dpar ( $\mu$ m)	$\pm 1\sigma$ ( $\mu$ m)	<i>Rho</i> s	<i>N</i> s	<i>Rho</i> i	<i>N</i> i	AFT Age (Ma)	$\pm 1\sigma$ (Ma)	P ( $\chi^2$ ) (%)
PC1.1	4900	25.07174	65.8569	18	8.1	5000	1.79	0.24	1.76	102	5.34	310	133	37.3 (P1)	44
				7			1.81	0.23					15.8	5.1 (P2)	72
PC1.2	4740	25.07195	65.8594	35	7.2	5000	2.27	0.82	0.97	136	10.37	1460	11.5	1.4	49
PC2	4551	25.07416	65.864	19	7.3	5000	2.09	0.35	1.01	56	13.55	749	9.4	1.5	88
PC2.1	4405	25.07115	65.8656	27	7.4	5000	1.83	0.25	0.89	78	11.43	998	10.5	1.6	11
PC3.1	4045	25.07572	65.8738	37	8.2	5000	1.91	0.53	0.69	100	12.02	1724	8.2	1.0	21

401 **Notes:** AFT data from PC1.1 is divided into 2 populations, each of which passes the Chi-squared test.  
402 Abbreviations: H, elevation; Lat, Latitude; Lon, Longitude; n, number of analyzed crystals; *Rho*D,  
403 dosimeter density; Dpar, pit diameter of tracks; *N*s, number of spontaneous track;  $\rho_s$ , spontaneous  
404 track density ( $\times 10^5 \text{ cm}^{-2}$ ), *N*i, number of induced track;  $\rho_i$ , induced track density; P ( $\chi^2$ ), Chi-squared  
405 test. Zeta-value  $330.4 \pm 20.2$  obtained by the analyst (Guadalupe Arzadum) using IRMM 540  
406 dosimetry glasses. Sample etching conditions ( $\text{HNO}_3$ ) 5.5 N for 20' at 20°C.  
407  
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409  
410 Thermochronologic data derived from a vertical transect can be readily evaluated for changes in  
411 cooling trends. The apparent exhumation rate can be determined from a plot of AHe or AFT ages  
412 versus elevation when the samples are considered to be fully reset (e.g., Fitzgerald et al., 1986). The  
413 AFT data on figure 6 show a positive correlation between elevation and age. The young ages of the 4  
414 lower AFT samples are interpreted to reflect fully reset ages. In contrast, the 2 populations with  
415 central ages of 133.0 and 15.8 Ma from the highest sample (PC1.1) suggest that the sample has not  
416 been hotter than the AFT closure temperature during the Cenozoic. Therefore, the bottom of a fossil  
417 PAZ is placed below this sample, suggesting that there was an increase in the exhumation rate at ~11  
418 Ma.





419

420 **Fig. 6** AFT and AHe cooling age plotted vs. elevation of PC samples, showing regression line  
 421 (dashed) whose slope indicates the apparent exhumation rate. Grey dots represent AHe single grain  
 422 ages. Orange dots represent AFT ages. The exhumation rate yields  $\sim 0.2 \pm 0.1$  km/Ma from the AFT  
 423 regression line with a  $R^2 = 0.73$  (calculated without the partial reset sample PC1.1).  
 424

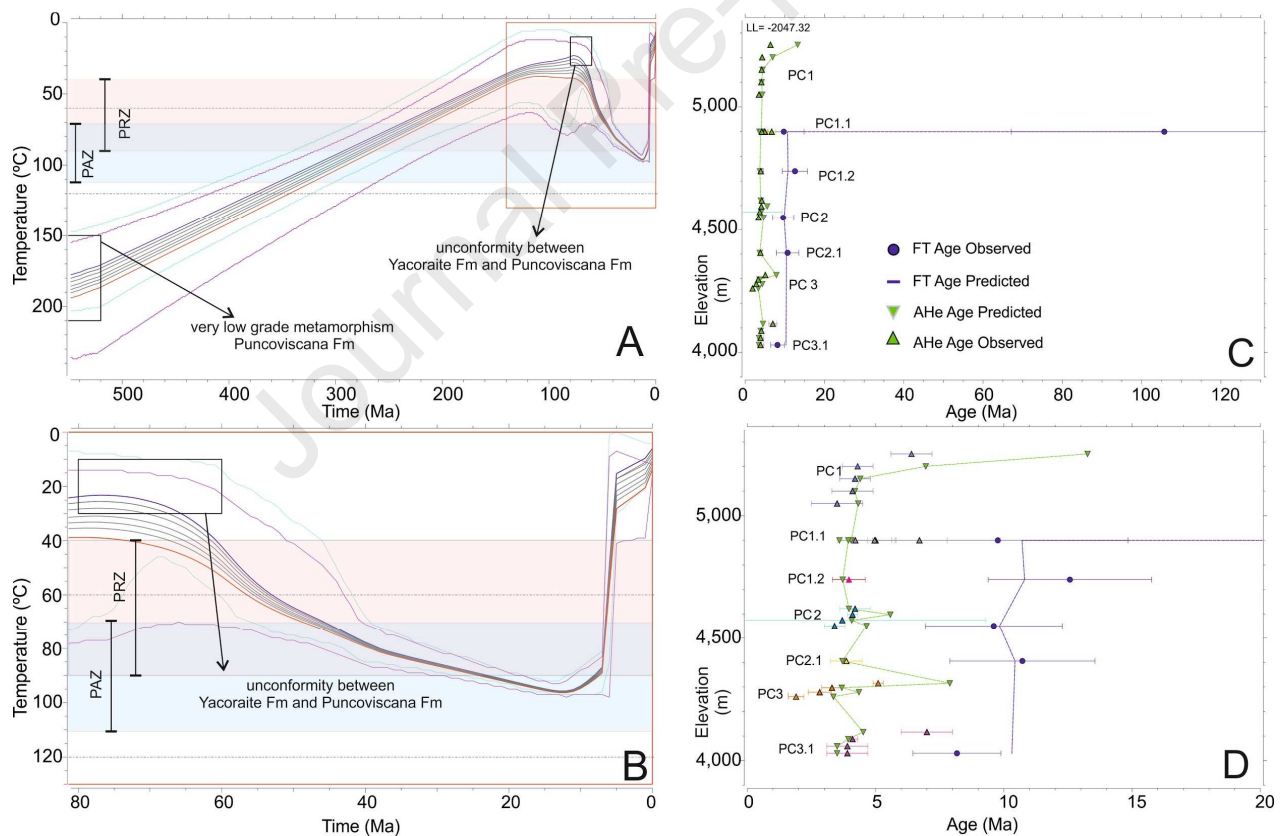
#### 425 4.3 Thermal history modeling results

426 We included all 23 AHe single grain ages in our QTQt model because they pass the criteria  
 427 presented in the methodology section. Twelve aliquots from samples PC1.1, PC2, and PC3.1 were  
 428 included because they exhibit intra-sample reproducibility. Nine aliquots from samples PC1 and PC3  
 429 were included because they exhibit a positive correlation between eU and age (Table 1). Although  
 430 samples PC1.1 and PC2.1 each had only a single aliquot AHe age, these two aliquots were included  
 431 because they overlap with the AHe ages from the samples above and below (Table 1). The PC3.1b  
 432 aliquot (in the lower sample) with high concentration of U was included because could indicate the  
 433 maximum temperature reached.

434 AFT data from five samples were included in the model. Although the two grain age populations  
 435 from sample PC1.1 show similar  $D_{par}$  values, the populations were modelled separately as  $D_{par}$  does  
 436 not account for all possible kinetic differences among the grains (e.g., Ketcham et al., 1999).  
 437 Unfortunately, we do not have information about wt% Cl from our samples. Since QTQt can resample

438 Dpar values during the iterations to improve the fit of the data (Gallagher, 2012), we have modeled  
 439 both AFT-age populations to test if the observed age dispersion can be explained by different mineral  
 440 kinetics.

441 Figure 7A shows the expected multi-sample thermal history of samples arranged according to  
 442 present-day elevations. The Max likelihood chain is shown in the supplementary data. This model  
 443 exhibits pre-80 Ma cooling, followed by Late Cretaceous to Miocene reheating up to 100°C and final  
 444 cooling between ca. 10 Ma and the present. In this model, the 23 AHe aliquots are adequately  
 445 predicted, including the two different AFT age populations of sample PC1.1. As expected, the model  
 446 predicts different Dpar values for each population. Predicted AFT ages for the lowermost sample are  
 447 at least 2 Ma older than the observed age. The model successfully predicts the remaining AFT ages  
 448 (Fig. 7C).



449

450 **Fig. 7. A.** The expected multi-sample thermal model, showing the uppermost and the lowermost  
 451 sample thermal histories plotted in blue and red, respectively. The thermal histories for all intervening  
 452 samples are drawn in grey. For the uppermost sample, the 95% credible intervals are drawn in light  
 453 blue and these reflect the uncertainty in the inferred thermal history alone. For the lowermost sample,  
 454 the 95% credible intervals are drawn in magenta and these reflect the combined uncertainty in the  
 455 inferred thermal history and also the offset parameters. Light blue and light pink shaded areas  
 456 represent the apatite partial annealing zone and apatite partial retention zone. **B.** Detail of the last 80

457 Ma from 0 and 130°C of the thermal model. **C.** Comparison of observed AHe-AFT ages versus those  
458 predicted by the thermal model, **D.** Magnified version of the C panel, highlighting the last 20 My.  
459 Legend: blue dots-fission track ages observed, blue line-fission track predicted, colours triangles-AHe  
460 ages observed, green inverted triangles-AHe ages predicted.  
461

## 462 **5. Discussion**

### 463 *5.1 Exhumation analysis of the Malcante Range*

464 According to the regional thickness of the sedimentary post-rift and foreland strata, sample PC1.1  
465 (4,900 m a.s.l.) must have been buried beneath at least 2,700 m of strata. Our structural interpretation  
466 suggests that about 500 m of basement was removed by erosion; therefore, this sample was buried by  
467 ~3,200 m. The multi-sample thermal history model shows that the upper sample experienced a  
468 maximum temperature of ~95°C at ~11 Ma. Considering a mean annual surface temperature of 10°C,  
469 this implies a paleo-geothermal gradient ~27°C/km for the late Miocene, which is higher than the  
470 estimates for the Payogastilla Group in the Angastaco Basin to the southeast of the study area ( $18 \pm 8^\circ$   
471 C/km at 16/15 Ma, Deeken et al., 2006). Moreover, if the upper sample was at 3,200 m depth at 10  
472 Ma, this would imply an average exhumation rate of  $\sim 0.3 \pm 0.5$  km/Ma, possibly faster than the  $\sim 0.2 \pm$   
473 0.1 km/Ma apparent exhumation rate predicted from the age-elevation plot (Fig. 6). Two possible  
474 explanations for these observations are that i) the amount of burial could be significantly  
475 underestimated or that ii) the rock column could have been rotated about a horizontal axis during  
476 exhumation, such that the samples now span a smaller range of elevation than when they cooled  
477 through the apatite PAZ and PRZ.

478 During the growth of the significant hanging-wall anticline, the Malcante Range experienced a  
479 transition from the erosion of structurally shallow, easily-eroded sediment to the erosion of  
480 structurally deeper, more resistant basement lithologies. The contrast in rock erodibility between the  
481 basement and the overlying sedimentary strata may have resulted in faster initial exhumation rates and  
482 reduced surface uplift during the erosion of the sedimentary cover, followed by slower exhumation  
483 rates and increased surface uplift after the basement was exposed; thus, the bulk of the surface uplift  
484 of the Malcante Range probably occurred during the final stages of the exhumation. The majority of  
485 the present 3,000 m of relief was likely formed after the removal of much of the ~2,700 m thick post-  
486 rift and foreland-basin sedimentary sequences, when the more resistant Puncoviscana Formation was

487 exposed. Similar relationships between rock erodibility, exhumation, and surface uplift have been  
488 documented in other segments of the Central Andes (Sobel and Strecker, 2003; Zapata et al., 2019b).

489 A significant rock column was exhumed between 10 and 0 Ma. Provenance studies on the Palo  
490 Pintado Formation (10-5 Ma) in the Escoipe area, to the southeast, did not reveal the exhumation of  
491 the post-rift Yacoraite Formation (Gonzalez Villa, 2002). Therefore, we suggest that the erosion  
492 products from the exhumed Malcante Range could have been transported farther east, out of the  
493 Calchaquí Basin and possibly deposited in the late Miocene-Pliocene conglomerate sequence in the  
494 adjacent Lerma Valley which contains eastward-directed paleocurrent indicators (Hain et al., 2011);  
495 this hypothesis should be corroborated in future studies.

496

## 497 *5.2 Fault reactivation history*

498 The thin sequence of syn-rift strata (Pirgua Subgroup) in the Las Zanjias area, on the eastern flank  
499 of the Malcante anticline, and the basement unconformity below the Yacoraite Formation to the west  
500 and south (Abra de Malcante and Escoipe areas, Figs. 2 and 3) indicate a westward pinch-out of the  
501 Pirgua deposits. This in turn supports the notion of a gently eastward-sloping paleotopography during  
502 the Early Cretaceous. The unconformity between the marine deposits of the Yacoraite and  
503 Puncoviscana formations implies that the area was at sea level during the Maastrichtian marine  
504 transgression (Marquillas et al., 2005).

505 From the Cretaceous to the Eocene, this region was a depocenter controlled by thermal subsidence,  
506 represented by the final post-rift deposits of the Balbuena and Santa Bárbara Subgroups (Mealla,  
507 Maíz-Gordo, and Lumbrera formations) (Marquillas et al., 2005). Subsequently, the Payogastilla  
508 Group was deposited in a foreland-basin setting (e.g., Starck and Vergani, 1989; Díaz and Malizzia,  
509 1983; Coutand et al., 2006; Payrola et al., 2009). The first stage of Malcante structural growth is  
510 defined by the low-angle unconformity between the Quebrada de los Colorados Formation and the  
511 Mealla/Yacoraite formations, highlighting a phase of middle-late Eocene deformation that caused the  
512 erosion of the 300-m-thick Santa Bárbara Subgroup from the western limb of the Malcante anticline  
513 (Fig. 3). Unfortunately, the expected thermal effect of this erosion would be ~5 to 8°C, too small to be  
514 recorded by our thermochronological data or a thermal history.

515 The middle Oligocene - middle Miocene Angastaco Formation records local angular  
516 unconformities linked to several deformation events in the Calchaquí region (Carrera and Muñoz,  
517 2008; Carrapa et al., 2012; Aramayo et al., 2017; Payrola et al., 2020; Fig. 2A). The decrease in the  
518 thickness of this unit in the western limb of the Malcante anticline (northern part) could be related to  
519 the reactivation of the Malcante fault during the middle Miocene, which is consistent with the  
520 unconformities recorded in the Tonco valley (Payrola et al., 2020) and the Quebrada de Carachi - El  
521 Toro (Pearson et al., 2013; Montero-López et al., 2017). Our modeled thermal history between 35 and  
522 ca. 10 Ma indicates a relatively long subsidence period in the region (Fig. 7), which is necessary to  
523 accumulate the foreland deposits. The thermal model constrains the beginning of the  
524 cooling/exhumation event at ~10 Ma (Fig. 7). This exhumation must be related to episodes of  
525 Malcante fault reactivation as deduced from mapped unconformities in the Malcante anticline, as well  
526 as episodes of faults reactivation recorded in the regional Miocene basin strata (Fig. 8A, Carrera and  
527 Muñoz, 2013; Pearson et al., 2013; Payrola et al., 2020).

528 This late Cenozoic cooling event is regionally consistent with the ca. 10 Ma AHe cooling ages  
529 obtained from the neighboring Tin-Tin (5.7 - 11.9 Ma), Angastaco (9.9 - 10.6 Ma), Amblayo-Agua de  
530 Castilla (8 - 12 Ma), and Sierra de los Colorados (5.7 - 9.9 Ma) sections, farther southwest (Fig. 8A;  
531 Carrapa et al., 2011; Kortyna et al., 2019; Payrola et al., 2020). In addition, thermochronological data  
532 (especially AHe ages) from the Malcante Range and other neighboring mountain ranges (Fig. 8A)  
533 (from north to south, the Pascha-Lesser (5.2 - 12.1 Ma, Pearson et al., 2013), Malcante (4.2 - 10.2 Ma,  
534 this study), Tonco-Filos del Pelado (7.4 - 17.1, Payrola et al., 2020), and Agua de Castilla ranges (8 -  
535 12 Ma, Kortyna et al., 2019)), could indicate a concentration of shortening in a relatively narrow  
536 north-south trending deformation belt. Our study area presently receives ~300 mm of mean annual  
537 precipitation (Noe et al., 2012). It is possible that the establishment of the South American low-level  
538 jet (e.g., Vera et al., 2006) at 10-7 Ma (Strecker et al., 2007; Mulch et al., 2010; Rohrman et al.,  
539 2016) could have contributed to the 7 and 5 Ma accelerated exhumation of the Malcante Range (based  
540 on the thermal model) by increasing the available precipitation in this area. A more humid setting at  
541 this time is indicated by isotopic data and fossils from the Palo Pintado Formation in other areas (  
542 Pingel et al., 2016; Rohrman et al., 2016). After 5 Ma, the exhumation of less erodible bedrock must

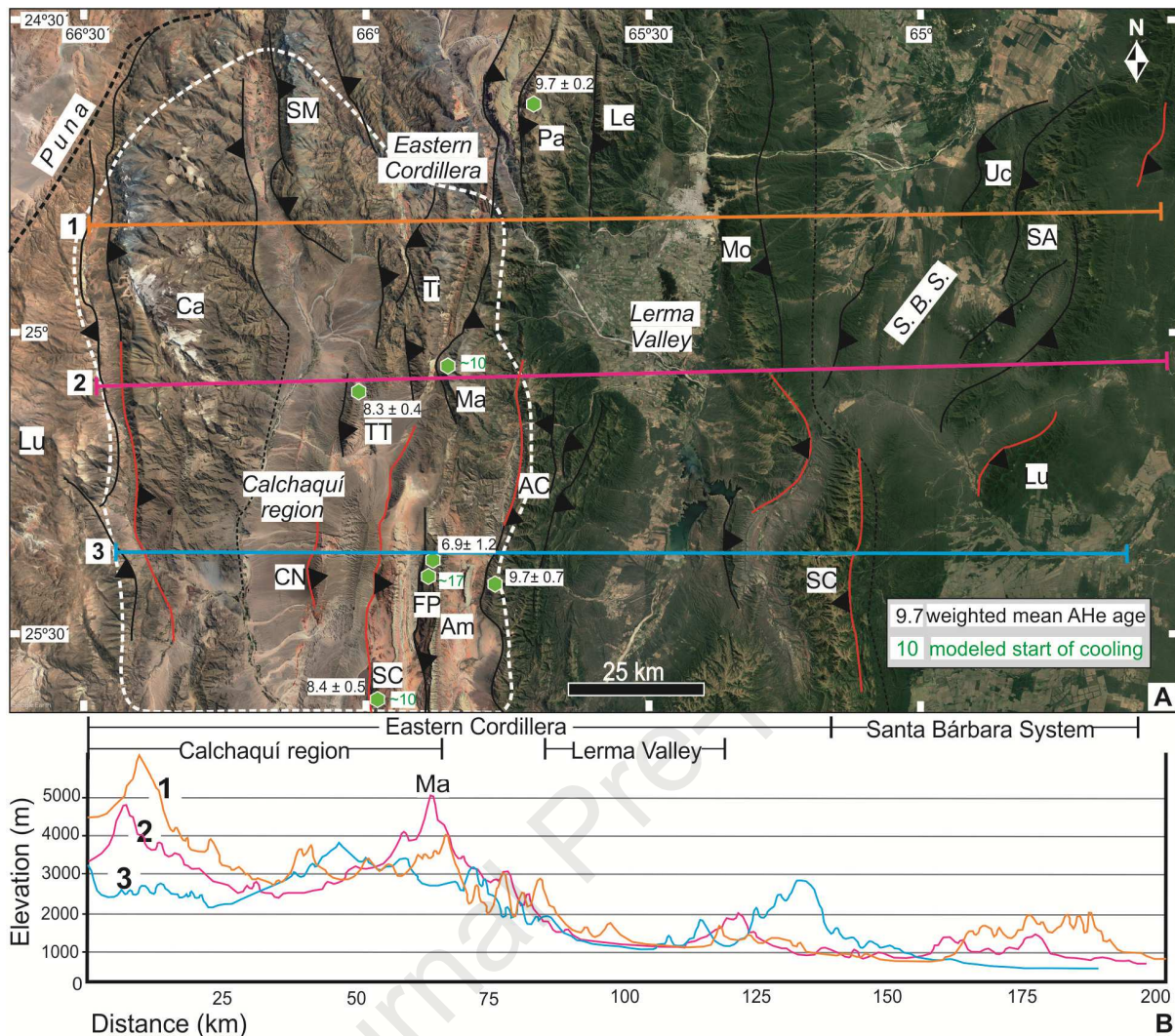
543 have facilitated the development of high topography, which is consistent with surface uplift estimates  
544 obtained from intermontane basins in this region (Rohrman et al., 2016; Pingel et al., 2016, 2020).

545 On the other hand, pre-Cenozoic basement structures can be reactivated more easily than new  
546 faults can be created (Masek and Duncan, 1998; Hilley and Strecker, 2004; Yagupsky et al., 2014).  
547 The Malcante retrovergent fault could be a pre-Andean heterogeneity that was reactivated during the  
548 Cenozoic, as suggested by the west-vergence of this fault and the absence of Cretaceous or Paleozoic  
549 units in the surrounding area. This >5,000-m high structural system could have accumulated a large  
550 amount of shortening during the Cenozoic, defining part of a zone of concentrated deformation which  
551 includes the Luracatao, Cachi, Calchaquí and Malcante fault zones (Fig. 8A), where basement with  
552 different metamorphic grades was uplifted; deeper crustal levels were exhumed in the north than in the  
553 south (Hongn et al., 2010; Payrola et al., 2012; Pearson et al., 2012; Payrola et al., 2020).

554 The 5,100-m high Malcante Range is located in an area with reduced rainfall, to the west of an  
555 orographic barrier which does not exceed 4,200 m in the heavily vegetated Yungas green belt (Fig.  
556 1B). The position of this anomalously high range (Fig. 8B) could be related to the west-east regional  
557 balance of shortening. Figure 8 depicts a southward decrease of the topography, visualized along three  
558 west-east transects. The number of significant Cenozoic thrust faults also varies between profiles. The  
559 northern transect (1) crosses the Cachi, San Miguel, and Tirao high altitude basement blocks, which  
560 have absorbed a large amount of Cenozoic fault-displacement in three well-defined zone of  
561 concentrated deformation (Hongn and Seggiaro, 2001; Payrola et al., 2012, Pearson et al., 2012 and  
562 2013). In contrast, the two southern transects depict a relative decrease in the topography, which is  
563 interpreted to reflect a decrease of fault-displacement along individual structures that is compensated  
564 by an increase in the number of regional faults in order to balance the regional crustal shortening  
565 (Salfity and Monaldi, 2006; Carrera and Muñoz, 2013; Payrola et al., 2020).

566





567  
 568 **Fig. 8. A.** Google Earth satellite image of the southern Eastern Cordillera and northern Santa Bárbara  
 569 System (S.B.S.) showing location of AHe ages (Carrapa et al., 2011; Pearson et al., 2013, Kortyna et  
 570 al., 2019; Payrola et al., 2020). Main reverse faults shown in black, Cretaceous inverted faults in red;  
 571 major mountains ranges are labelled (Lu Cumbres de Luracatao, Ca Cachi, SM San Miguel, CN Cerro  
 572 Negro, TT TinTin, Ti Tirao, Ma Malcante (this study), SC Sierra de los Colorados, FP Filos del  
 573 Pelado-Tonco, Am Amblayo, AC Agua de Castilla, Pa Pascha, Le Lesser, Mo Mojotoro, SC Crestón,  
 574 Uc Unchime, SA San Antonio, Lu Lumbrera). Dashed white lines mark the borders of the Payogastilla  
 575 Group basin, coincident with the borders of Calchaquí region. Dashed black lines show the borders  
 576 between main morphostructural units. **B.** Three topographic profiles showing the southward and  
 577 eastward decrease of altitude. Location of profiles is shown in panel A. Note the higher altitude of the  
 578 Malcante Range in profile 2 compared to the other ranges in this position along strike.  
 579

580

## 581 6. Conclusions

582 The basement anticline Malcante Range was uplifted by a retrovergent reverse fault. Based on our  
 583 structural reconstructions, this fault has a minimum Cenozoic displacement of 1,400 m. The  
 584 significant unconformity between the Quebrada de los Colorados Formation and Yacoraite Formation,

585 located on the western limb of the Malcante anticline, delineates the first episode of faulting during  
586 the middle Eocene. About 2,700 m of strata were deposited between the late Eocene and ca. 10 Ma.  
587 AFT and AHe data from a vertical profile across the Malcante Range suggest a period of rock uplift  
588 and efficient erosion that started around 11 Ma. This exhumation event was related to erosion of soft  
589 sedimentary cover during the formation of the actual Malcante orographic barrier, which separates the  
590 humid climate in the Lerma Valley from the semi-arid climate in the northern Calchaquí region,  
591 possibly during the Plio-Pleistocene. The Puncoviscana basement is now exposed in the core of the  
592 range; we suggest that exposure of this more erosionally-resistant lithology may have been responsible  
593 for significant surface uplift of the Malcante Range during the Plio-Pleistocene.

594 The anomalously high altitude of the Malcante Range compared to other ranges in similar  
595 structural positions along the eastern border of the Calchaquí region could be related to the  
596 accommodation of late Miocene-Pliocene shortening along only a few faults compared to a larger  
597 number of Cenozoic faults which uplift the basement-cored ranges to the north and south.  
598 Concentrating deformation along only a few structures could facilitate a higher amount of rock uplift  
599 in the Malcante Range. Thus, high elevation was related to the compensation of east-vergent  
600 shortening linked to the availability of pre-existing basement structures, which could have been  
601 reactivated during the Cenozoic.

602

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615

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Sample	H (m)	Lat (° S)	Long. (° W)	Stratigraphic unit	age (Ma)	Raw				II				ESR			
						(Ma)	(Ma)	(Ma)	(Ma)	(ppm)	(ppm)	(ppm)	eU	He (nmol/g)	Ft	( $\mu$ m)	Term.
PC1	5100	25.074583	5.852036	Puncoviscana Fm	600-530	PC1-a	6.4	10.2	0.4	60	68	19	76	2.62	0.63	40	1
						PC1-b	3.5	5.1	0.5	9	22	32	14	0.27	0.68	47	2
						PC1-c	4.1	6.8	0.4	27	65	12	42	0.95	0.61	38	1
						PC1-d	4.2	6.5	0.3	29	34	41	37	0.84	0.64	42	1
						PC1-e	4.3	7.1	0.3	53	58	114	67	1.58	0.61	39	1
PC1.1	4900	25.071743	65.85688	Puncoviscana Fm	600-530	PC1.1a	6.6	10.1	0.4	66	76	53	84	3.01	0.66	44	1
						PC1.1b	5.0	8.1	0.5	29	19	60	34	0.91	0.61	39	1
						PC1.1c	4.2	6.5	2.3	2	17	8	6	0.13	0.64	41	1
						PC1.1d	4.9	7.4	0.3	63	89	55	84	2.26	0.67	45	0
PC1.2	4740	25.071947	65.85936	Puncoviscana Fm	600-530	PC1.2a	3.8	6.2	1.5	7	3	36	8	0.17	0.62	39	1
PC2	4551	25.074162	65.86396	Puncoviscana Fm	600-530	PC2-a	4.2	5.9	0.3	27	18	45	31	0.72	0.71	52	1
						PC2-b	4.1	4.6	0.1	7	2	1	7	0.16	0.88	122	1
						PC2-c	3.4	4.5	0.2	19	5	10	20	0.36	0.76	62	0
						PC2-e	3.7	5.3	2.8	1	6	1	2	0.04	0.70	49	1
PC2.1	4405	25.071146	65.86561	Puncoviscana Fm	600-530	PC2.1a	3.9	6.1	0.4	41	124	25	70	1.47	0.63	41	0
PC3	4257	25.070608	65.86945	Puncoviscana Fm	600-530	PC3-a	3.3	5.4	1.6	6	7	2	7	0.13	0.62	39	1
						PC3-b	5.0	7.0	0.1	30	146	14	65	1.78	0.73	55	1
						PC3-c	2.9	4.2	0.4	18	36	4	27	0.41	0.68	47	0
						PC3-e		3.6	2.6	4	10	4	7	0.07	0.57	35	0
PC3.1	4045	25.075715	65.87382	Puncoviscana Fm	600-530	PC3.1a	7.0	9.2	0.5	3	13	6	6	0.24	0.76	62	0
						PC3.1b	4.1	6.1	0.1	134	166	29	173	3.80	0.67	45	0
						PC3.1c	3.9	6.4	0.4	36	95	65	59	1.26	0.61	39	0
						PC3.1d	3.9	6.6	1.0	19	9	32	21	0.46	0.59	37	0

Table 1 AHe data

Sample	H (m)	Lat (° S)	Long (° W)	n	RhoD	Journal Pre-proof									
						Mean ( $\mu\text{m}$ )	( $\mu\text{m}$ )	( $\mu\text{m}$ )	( $\mu\text{m}$ )	( $\mu\text{m}$ )	( $\mu\text{m}$ )	( $\mu\text{m}$ )	( $\mu\text{m}$ )	(Ma)	(Ma)
PC1.1	4900	25.071743	65.85688	18	8.1	5000	1.79	0.24	1.76	102	5.34	310	133 (P1)	37.3	44
				7			1.81	0.23						15.8 (P2)	5.1
PC1.2	4740	25.071947	65.85936	35	7.2	5000	2.27	0.82	0.97	136	10.37	1460	11.5	1.4	49
PC2	4551	25.074162	65.86396	19	7.3	5000	2.09	0.35	1.01	56	13.55	749	9.4	1.5	88
PC2.1	4405	25.071146	65.86561	27	7.4	5000	1.83	0.25	0.89	78	11.43	998	10.5	1.6	11
PC3.1	4045	25.075715	65.87382	37	8.2	5000	1.91	0.53	0.69	100	12.02	1724	8.20	1.00	21

Table 2: AFT data

**Exhumation and structural evolution of the high-elevation Malcante Range, Eastern Cordillera, NW Argentina**

**Highlights**

Calchaquí high-mountains with anomalous position.

Angular unconformity in the contact between Paleogene post-rift deposits and middle Eocene foreland deposits.

Malcante Range exhumation start at 10 Ma.

Accelerated uplifting during the Pliocene.

**Declaration of interests**

The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

The authors declare the following financial interests/personal relationships which may be considered as potential competing interests:

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