Open Research Online



The Open University's repository of research publications and other research outputs

Hunting for Life on Mars by Studying Life on Earth

Journal Item

How to cite:

Macey, Michael; Grand, Ann; Fox-Powell, Mark; Ramkissoon, Nisha; Cousins, Claire and Olsson-Francis, Karen (2021). Hunting for Life on Mars by Studying Life on Earth. Frontiers for Young Minds, 9, article no. 665929.

For guidance on citations see FAQs.

 \bigodot 2021 The Authors



https://creativecommons.org/licenses/by/4.0/

Version: Version of Record

Link(s) to article on publisher's website: http://dx.doi.org/doi:10.3389/frym.2021.665529

Copyright and Moral Rights for the articles on this site are retained by the individual authors and/or other copyright owners. For more information on Open Research Online's data <u>policy</u> on reuse of materials please consult the policies page.

oro.open.ac.uk



NEW DISCOVERY Published: October 18, 2021

Hunting for Life on Mars by **Studying Life on Earth**

Authors



Young Reviewers

Faydh Mohammed 🏻 🏫 Ginny



Articles

Participate

Sections

About

water-rich, like Earth. Scientists believe that, as water was lost on Mars, the remaining water would have become increasingly salty and rich in sulphur. By researching environments on Earth with a similar chemistry, we can develop hypotheses about whether life could have existed on Mars. In our research, we studied Colour Peak, a sulphur-rich, salty spring system in the Canadian High Arctic. This environment contains bacteria that use sulphur to get their energy. By studying water and sediment samples from this region, we confirmed that bacteria that used sulphur as an energy source could potentially have lived on Mars. These results can help scientists who are analysing data from Mars missions to identify evidence of former life on Mars.

THE WATER OF MARS

Collections

All living things need water. Since 1971, nine missions have successfully landed on Mars, and there are currently 14 satellites orbiting the planet. None of these missions have found any evidence of water or life on the surface today. However, those missions have observed structures (such as river basins) and waterbearing rocks and minerals (such as clays) on Mars' surface, which may be evidence for the presence of water [1]. Scientists believe that, billions of years ago, there was water on Mars.

Mars lost its water due to the partial loss of its atmosphere, but thanks to data from missions, scientists have an idea of the environment on early Mars. Scientists believe that, as the water

sulphide. This is the chemical that gives the smell to rotten eggs! We also know that the key elements required for life on Earth (carbon, hydrogen, nitrogen, oxygen, phosphorous and sulphur; abbreviated as <u>CHNOPS</u>) have all been identified in martian rocks and soils.

If life ever existed on Mars, it would have existed in or near bodies of water, and these potential martian organisms might have left behind evidence of their existence. Before we send missions to Mars to search for evidence of life, we need to have an idea of what kind of evidence we should look for. This preparation is important, because sending missions to Mars is incredibly complicated and expensive. Scientists start their preparation by studying environments on Earth that have similar water chemistries to that of early Mars. No environment on Earth is a perfect fit, but there are many that are similar enough for us to develop our understanding of martian water and whether life could have ever existed in it.

One such environment is the Colour Peak spring system on Axel Heiberg Island in the High Arctic of Canada (Figure 1), where the water is very salty and sulphur-rich. The average air temperature at Colour Peak is -20° C. In winter, it can go as low as -40° C, but the springs on the island never freeze because of their high salt content (in the same way that salt is spread on roads to stop ice from forming) [2, 3]. We collected sediment and water samples from Colour Peak and stored them at a low temperature until they reached a laboratory in the United Kingdom.

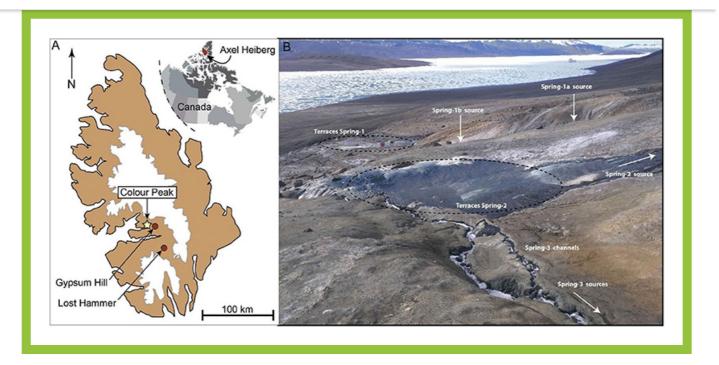


Figure 1 - (A) Map of Axel Heiberg Island, with land shown in brown and icecaps shown in white.

Colour Peak is marked by a star and two other salty springs are marked by red dots. **(B)** Photograph of Colour Peak with labels showing the sources of its springs. The map was produced by modifying an image from Google Maps (Map Data @2020 Google) using the program Illustrator Creative Cloud version 21.0.2. These images are reproduced from Macey et al., [3] under creative commons licence 4.0 (creativecommons.org/licenses/by/4.0/).

We analysed the chemistry of the water from Colour Peak springs using a technique called <u>inductively coupled plasma optical</u> <u>emission spectroscopy (ICP-OES)</u>. ICP-OES works by heating water to create plasma (a very hot, energised gas) and detecting the specific colours of light given off by each element in the water. Using geological data from the surface of Mars that was collected by the NASA Curiosity rover, we used a computer programme to model the chemical composition of the ancient water on Mars [4]. The composition of the predicted martian water was incredibly similar to the composition of the water from Colour Peak. Th means the bacteria living in the Colour Peak springs are survive

About

Articles

Collections

Participate

Sections

About

WHAT LIVES IN COLOUR PEAK WATER?

We identified the bacteria that live in the Colour Peak springs by looking at their deoxyribonucleic acid (DNA). The well-known double helix of DNA is the genetic code that instructs cells to produce many kinds of proteins and regulates how those proteins are used in cellular functions. There is one key gene that lets us identify the bacteria that live in an environment: the <u>16S rRNA gene</u>. This gene, which exists in a related but slightly different form in all types of bacteria, performs a key role in bacterial replication and growth. By studying this gene in an environment, we can identify the various kinds of bacteria that live there.

However, before we could study the gene, we first had to extract (remove) the DNA from the bacteria in our samples. To do this, we had to overcome two major problems. First, the standard way to extract DNA from bacteria is to use chemicals to burst the cells open, leaving only the DNA behind. The high amounts of salt in the Colour Peak water caused some of these chemicals to react in strange ways, meaning that we could not collect any DNA. Second, because the Colour Peak water is cold and salty, the DNA from bacteria that died and burst thousands of years ago was preserved, and we cannot tell the DNA from long-dead bacteria apart from that of living bacteria [5].

To overcome the first problem, we washed the samples with ultrapure water to wash away the salt from the bacteria. To solve the second problem, we studied ribonucleic acid (RNA) as well as DNA. RNA is a single-stranded copy of the DNA that a cell rea produce proteins. RNA breaks down incredibly quickly, even i

alive when the samples were taken. Using this approach, our research revealed that most bacteria living in Colour Peak belong to a group called <u>sulphur-oxidising bacteria</u> (SOB) [<u>3</u>].

WHAT DO THESE RESULTS MEAN FOR MARS?

SOB use sulphur as a source of energy. Given the large amount of sulphur in Colour Peak water, it makes sense that this is a successful way to live in this environment. The water of ancient Mars would also have contained lots of sulphur, meaning that the main energy source for SOB would have been available. Many SOB obtain the carbon they need by capturing it from carbon dioxide in the air, like plants do—this is called <u>autotrophy</u>. Autotrophy would also be a good strategy on Mars, due to the large amount of carbon dioxide in the atmosphere [95,000 parts per million (ppm) compared to 415 ppm on Earth]. Some SOB can also grow without oxygen (which is very low on Mars), and can instead use nitrogen-based chemicals, called nitrates, which have been detected in the rocks of Mars (Figure 2) [6].

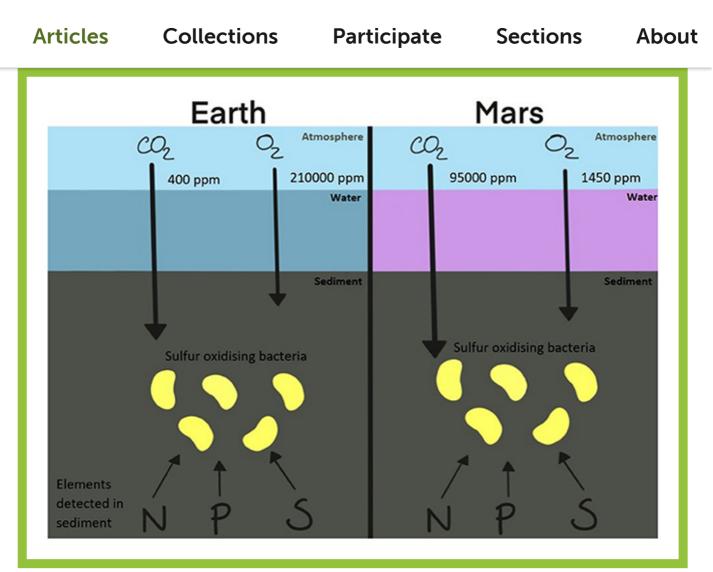


Figure 2 - Proposed sources of the life-essential elements (carbon, nitrogen, oxygen, phosphorous and sulphur - CHNOPS) for the SOB within Colour Peak springs and the water environments predicted on Mars.

The presence of the elements in the sediment and the concentrations of carbon dioxide and oxygen in the atmosphere are based on either studies of the environment of Colour Peak or data from missions to the surface of Mars [6, 7]. These images are reproduced from Macey et al., [3] under creative commons licence 4.0 (creativecommons.org/licenses/by/4.0/).

After showing that SOB can thrive in the Colour Peak water and knowing that all the key elements (CHNOPS) required for growth are present, we investigated whether the martian water we had modelled would actually support bacterial growth. This is important because, as well as needing specific elements, SOB would als need to produce enough energy to grow. If we know the

bacteria need. The equation used to calculate the energy is called the <u>Gibbs energy equation</u> and it was developed by the scientist Willard Gibbs [8]. Using the Gibbs energy equation, we can say that not only would SOB produce enough energy to grow, but they would also be one of the *only* types of bacteria that could grow in these specific waters. An outline of our work is shown in <u>Figure 3</u>.

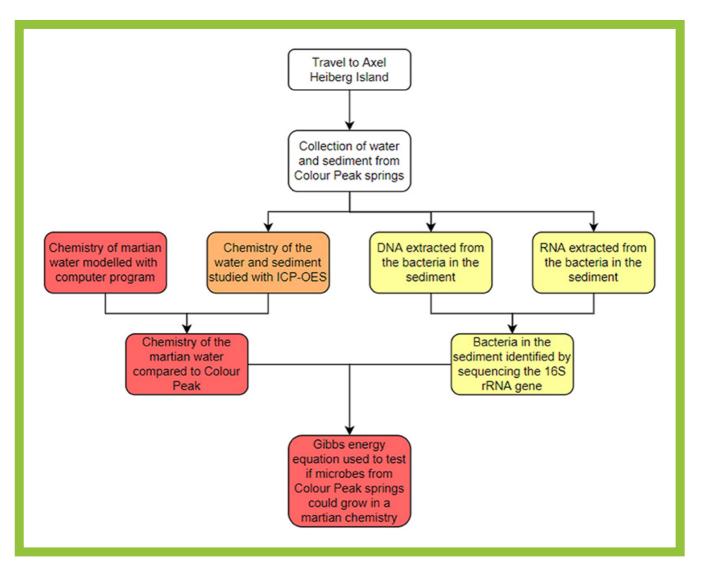


Figure 3 - The steps involved in this study.

Fieldwork steps are in white, microbiology steps are in yellow, chemistry in orange, and the Mars-focused work in red.

This is a very exciting result, because SOB alter the chemistry their local environment as they grow. This produces evidence



bacteria are dead [9, 10]. If bacteria like SOB once grew on Mars, it is possible that the kind of evidence left behind by SOB on Earth could be found on Mars.

THE FUTURE

We have studied an environment on Earth (the salty, sulphur-rich waters of Colour Peak Springs) that is considered to be similar to environments on early Mars (4.1-3.7 billion years ago). From this work, we have been able to gain a better understanding of what kinds of life might have been able to survive on Mars. By combining microbiology, chemistry, and modelling, we showed that sulphur oxidising bacteria thrive in this Arctic spring water, and so potentially would have been able to grow on early Mars.

As we are writing this, two rovers (NASA's *Perseverance* and the Chinese Space Administration's *Tianwen-1*) have recently landed on Mars. Understanding the data that these missions obtain and determining if life did exist on Mars will take many years and will ultimately be completed by future generations of young researchers. We are excited to think that our work and methods will contribute to addressing one of the biggest questions of our time: Are we alone in the universe?

Glossary

CHONPS: \uparrow An acronym for carbon, hydrogen, oxygen, nitrogen, phosphorous, and sulphur. These are considered the key elements for life as they make up nearly 98% of all molecules in living cells.

Inductively Coupled Plasma Optical Emission Spectroscopy (ICP-OES): scientific technique that works by heating water to create plasma (a hot a



16S rRNA Gene: ↑ This gene plays a role in cell growth and replication. It is in all bacteria in a related form, and the differences in sequence can be used to identify bacteria.

Sulphur Oxidising Bacteria: \uparrow A type of bacteria that get energy from sulphur.

Autotrophy: \uparrow This is the process of making complex compounds (e.g., the building blocks that make up a cell) from simple substances (e.g., like carbon dioxide-CO₂).

Gibbs Energy Equation: \uparrow This is an equation that can be used to calculate how much energy a specific chemical reaction will produce.

CONFLICT OF INTEREST

The authors declare that the research was conducted in the absence of any commercial or financial relationships that could be construed as a potential conflict of interest.

Original Source

↑Macey, M. C., Fox-Powell, M., Ramkissoon, N. K., Stephens, B. P., Barton, T., Schwenzer, S. P., et al. 2020. The identification of sulfide oxidation as a potential metabolism driving primary production on late noachian mars. *Sci Rep.* 10:10941. doi: 10.1038/s41598-020-67815-8

References

[1] ↑ Carr, M. H., and Head, J. W. 2010. Geologic history of Mars. *Earth Planet Sci Lett.* 294:185–203. doi: 10.1016/j.epsl.2009.06.042

[2] ↑ Perreault, N. N., Andersen, D. T., Pollard, W. H., Greer, C. W., and Why[.] 2007. Characterization of the prokaryotic diversity in cold saline perennial

[3] ↑ Macey, M. C., Fox-Powell, M., Ramkissoon, N. K., Stephens, B. P., Barton, T., Schwenzer, S. P., et al. 2020. The identification of sulfide oxidation as a potential metabolism driving primary production on late noachian mars. *Sci Rep.* 10:10941. doi: 10.1038/s41598-020-67815-8

[4] ↑ Bridges, J. C., and Schwenzer, S. P. 2012. The nakhlite hydrothermal brine on Mars. *Earth Planet Sci Lett*. 359–60:117–23. doi: 10.1016/j.epsl.2012.09.044

[5] ↑ Willerslev, E., Hansen, A. J., Rønn, R., Brand, T. B., Barnes, I., Wiuf, C., et al.
2004. Long-term persistence of bacterial DNA. *Curr Biol.* 14:13–4. doi:
10.1016/j.cub.2003.12.012

[6] \uparrow Stern, J. C., Sutter, B., Freissinet, C., Navarro-González, R., McKay, C. P., Archer, P. D., et al. 2015. Evidence for indigenous nitrogen in sedimentary and aeolian deposits from the Curiosity rover investigations at Gale crater, Mars. *Proc Natl Acad Sci USA*. 112:4245–50. doi: 10.1073/pnas.1420932112

[7] ↑ Mahaffy, P. R., Webster, C. R., Atreya, S. K., Franz, H., Wong, M., Conrad, P. G., et al. 2013. Abundance and isotopic composition of gases in the Martian atmosphere from the Curiosity rover. *Science*. 341:263–6. doi: 10.1126/science.1237966

[8] ↑ McCollom, T. M. 2007. Geochemical constraints on sources of metabolic energy for chemolithoautotrophy in ultramafic-hosted deep-sea hydrothermal systems. *Astrobiology*. 7:933–50. doi: 10.1089/ast.2006.0119

[9] ↑ Battler, M. M., Osinski, G. R., and Banerjee, N. R. 2013. Mineralogy of saline perennial cold springs on axel Heiberg Island, Nunavut, Canada and implications for spring deposits on Mars. *Icarus*. 224:364–81. doi: 10.1016/j.icarus.2012.08.031

[10] \uparrow Pellerin, A., Antler, G., Holm, S. A., Findlay, A. J., Crockford, P. W., Turchyn, A. V., et al. 2019. Large sulfur isotope fractionation by bacterial sulfide oxidation. *Sci Adv*. 5:eaaw1480. doi: 10.1126/sciadv.aaw1480

Citation

Macey M, Grand A, Fox-Powell M, Ramkissoon N, Cousins C and Olsson-F K (2021) Hunting for Life on Mars by Studying Life on Earth. Front. Young Minds.



Editor

Edward Gomez

Publishing dates

Submitted: February 8, 2021; Accepted: September 23, 2021; Published online: October 18, 2021.

Copyright © 2021 Macey, Grand, Fox-Powell, Ramkissoon, Cousins and Olsson-Francis

This is an open-access article distributed under the terms of the <u>Creative</u> <u>Commons Attribution License (CC BY)</u>. The use, distribution or reproduction in other forums is permitted, provided the original author(s) and the copyright owner(s) are credited and that the original publication in this journal is cited, in accordance with accepted academic practice. No use, distribution or reproduction is permitted which does not comply with these terms.

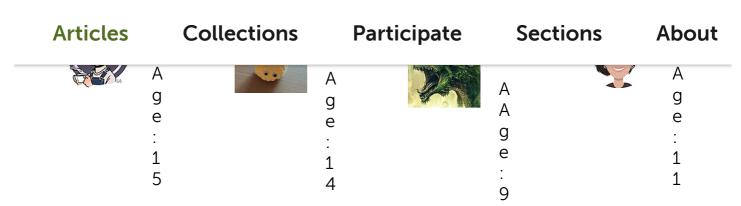
View Original Article

Related Articles



Articles	Collections	Participate	Sections	About
World:	The	Snail,	—T	he
From	Plants	the	Мо	st
the	?	Fragile	Ext	re
Botto	Micro	Coral,	me	
m of	bes!	and	An	ma
the Red	Plants get their food	the Off the	ls c The	
Sea The deep- sea brines of the Red Sea are some of the most extreme environmen ts on Earth. They have high salinity,	from the soil they grow in. But this food is only available thanks to a great diversity of	coast of Vietnam, two recent events seem to have contributed to the death of coral reefs: the rapid invasion of	toloratos	
high Authors André Antunes	Authors Aline Lacaze, Antoine Zboralski, David L. Joly	Authors Yvan Bettarel, Elyse Boudin, Sébastien Halary, Jean- Christophe Auguet, Jean Péronnin, Thierry Bouvier, Emma Rochelle- Newall, Van Ngoc Bui, Christelle Desnues	Auth Skand Elleud	de

https://kids.frontiersin.org/articles/10.3389/frym.2021.665529



View all Articles

Frontiers	Home	Blog	Contact
I I O I I I I I I		Diog	COntact

© 2007 - 2021 Frontiers Media S.A. All Rights Reserved