### The FLARE Network Vicarious Cal/Val for Earth Observation Satellites

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### Lambertian vs Specular Targets

Target signal embedded in a uniform scene is elevated above the low spatial frequency background (sky path radiance, adjacency effect, stray light, etc.) and is separable

- Background becomes a bias and is subtracted out based on image data alone
- Sensor response to target radiance is integrated (DN) contained in the PSF
- Atmospheric, adjacency, multiple scattering effects reduce to transmittance only

   measured with solar
   spectrometer coincident
   with overpass



## **FLARE Nodes**

#### Alpha Node

- Arlington, SD
- SDSU Evaluation
   partnership



**FLARE** 

• Ft. Worth, TX





## **FLARE Development**

- Mobile Node future development
  - Summer '21
  - Prototype at Beta site
- Custom Campaigns
  - Manual campaigns for targeted geometry
  - Commissioning Projects - individual assets or constellations
- Planned Nodes
  - Mauna Loa (3300m)
    - Fall '21
  - Railroad Valley Playa, Tenerife – TBD
  - Atacama, Australia, Gobabeb - TBD







### **FLARE Radiometric Tower: VNIR Spectrometer & SWIR Bands**



## **Dual Traceability - Langley & Sphere Methods**



### **Small Sats and Analysis Ready Data - CEOS CARD4L**





### Maturity Matrix Radar Plot: Transparent ARD ratings



Credit: Sam Hunt et al., VH-RODA Workshop, Apr 20, 2021

## **DQR and GIQE FLARE – Feed for Maturity Matrices**



DQR = Data Quality Report

FLARE Enabled Imagery Data Quality Report Sensor: Sentinel 2B Acquisition: Jan-March 2021

FLARE

Version 1.0 | March 2021

#### 4. FLARE Operation and Sensor Acquisition Summary

4.1. Ground Station Operational Status

The LOOKS and EVALS (Table 2) utilized in this Data Quality Report were executed using the FLARE BETA node at Brock TX, USA (Lat: 32.664241', Lon: -97.961547'). All LOOKS were calibrated using the Langleyderived fitting method with absolute SI traceability to the TSIS-1 instrument. No data quality issues were identified during any of the LOOKS.

Date	Time (UTC)	Solar El. (°)	Satellite El. (°)	Signal Level	Data Ident.
04 Jan 2021	17:24:52	32.1	88.5	6	S2B_MSIL1C_20210104T171719_N0209_ R112_T14SNB_20210104T192056
23 Feb 2021	17:24:44	43.5	88.4	6	S2B_MSIL1C_20210223T171309_N0209_ R112_T14SNB_20210223T192416
05 Mar 2021	17:24:45	47.2	88.5	6	S2B_MSIL1C_20210305T171109_N0209_ R112_T14SNB_20210305T204742



Figure 2: Spectral Intensity – during overpass with instrument RSR's





#### GIQE = General Image Quality Equation

#### FLARE Enabled GIQE Report Sensor: Example Acquisition: Example

FLARE

Version 1.0 | March 2021

#### 8. GIQE 5 and NIIRS Projection Results

8.1. FLARE Enabled NIIRS Calculated for Reference Conditions

NIIRS values as calculated by GIOE are dependent on operational conditions under test. Using analysis of sensor performance as characterized by FLARE, atmospheric modeling, and specific solar and satellite positions, FLARE-Enabled GIOE can be used to project NIIRS values for any given operational condition.

Below is a table of NIIRS values calculated for the Sensor Under Test projected to example operational conditions. These conditions assume a representative atmosphere (mid-latitude summer) with varying sensor and solar angles at time of acquisition.

Table 4:	Projected	NIIRS fo	or Ref	erence	Conditi
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NIIRS Values Predicted for Sun/Sat Elevations	Sat @ 90° GSDw = 0.70m	Sat @ 60° GSDw = 0.93m	Sat @ 30° GSDw = 2.5m
Sun @ 90°	4.6	4.2	2.6
Sun @ 60°	4.6	4.2	2.6
Sun @ 30°	4.6	4.1	2.4

8.2. SDNR Calculated for Reference Conditions

The Signal to Noise coefficient (SNDR) is calculated from predicted sensor response to reference solar constant, atmosphere, and simulated surface reflectance standards.

ble	5:	Signal	to	Noise	Parameters	for	Reference	Conditions

SDNR	Sat_El = 90 °	Sat_El = 60°	Sat_El = 30°
Sol_El = 90°	42.01	39.30	15.59
Sol_El = 60°	38.59	35.70	13.93
Sol_El = 30°	25.37	23.08	8.26

#### 8.3. FLARE Enabled NIIRS Calculated for Test Conditions

For the acquisition(s) reported here, the atmosphere, solar, and sensor conditions were known. These inputs were used to calculate a predicted NIIRS rating using GIQE for the FLARE test conditions.

ble 6: NIIRS Under Test Conditions				
Calculated NIIRS	SDNR	GSD_w		

calculated NIIRS	SUNK	GSD_W	RER
4.2	37.4	0.94m	0.3771

Finally, the projected NIIRS value for the sensor under test with test pose and solar angle using the reference atmosphere is provided below.

<b>abl</b> e	7: Projected	NIIRS under	Test Pos	e with Refe	rence Atmospher

Calculated NIIRS	SDNR	GSD_w	RER
4.2	37.4	0.94m	0.3771

# **FLARE Mission Quality Metrics**

IMAGE QUALITY PARAMETER	DESCRIPTION
Absolute Radiometric Performance	Imagery reported in-band radiance relative to uncertainty requirements.
Absolute Geolocation	Location error of imagery reported coordinates for FLARE signal center position relative to known values.
Multi-Spectral Registration	Inter-channel spatial band co-registration error based on evaluation of FLARE signal center position in reported bands.
Modulation Transfer Function	Nyquist MTF, other sensor resolution metrics (Point Response Function, Line Response Function, Rayleigh/Sparrow Criterion, Ground Spot Size, etc.).
National Imagery Interoperability Rating Scale	NIIRS value for provided imagery with FLARE target in-scene, derived through General Image Quality Equation v 5. Predicted NIIRS rating for sensor under alternative atmospheric conditions and solar/sensor geometries.



### Successful Engagements with Small Sat and Agency Assets





### **Current Agency Archive**



### **Post-Launch Commissioning and Trouble Shooting**



ALPHA Node – Arlington, SD USA. Imaged by Sentinel 2B July 16 2020

![](_page_12_Figure_3.jpeg)

Band-dependent halo/ringing, misregistration

### **Post-Launch Commissioning and Trouble Shooting**

![](_page_13_Picture_1.jpeg)

ALPHA Node – Arlington, SD USA. Imaged by Sentinel 2B July 16 2020

Across-track smear, thermal defocusing

## **Campaigns – Rapid Baseline Characterizations**

![](_page_14_Picture_1.jpeg)

![](_page_14_Picture_2.jpeg)

![](_page_14_Picture_3.jpeg)

![](_page_14_Picture_4.jpeg)

# **Resolution Enhancement Experiment**

Satellite Image (L1A)

![](_page_15_Figure_2.jpeg)

	ORBITAL ALTITUDE					
	500	km	450 km			
RESOLUTION METRIC	Cross (m)	Along (m)	Cross (m)	Along (m)		
Rayleigh (15.3%)	1.55 ± 0.05	$1.61 \pm 0.07$	1.37	1.39		
GSS (0.9%)	$1.23 \pm 0.04$	$1.26 \pm 0.05$	1.09	1.10		
Sparrow (0%)	$1.17 \pm 0.03$	$1.20 \pm 0.08$	1.07	1.05		

FLARE provided proof that lowering orbital altitude improved resolution capabilities and Ground Sample Distance, *before* a constellation wide maneuver

![](_page_15_Picture_5.jpeg)

## Mirror Empirical Line Method (MELM)

![](_page_16_Figure_1.jpeg)

**Different mirror in targets = multiple** 

Spectral Band	Slope: DN/Mirror	R <sup>2</sup>
Blue	17.9	0.9898
Green	25.2	0.9972
Red	22.8	0.9917
NIR	19.8	0.9965

FLARE

Results for Sept. 10, 2009 IKONOS collect.

DN/Mirror: Image po\_365282 Glass Mirror SPARC Target

![](_page_16_Figure_6.jpeg)

- MELM via Multi-LOOK events or Multi-Points in scene both provide
  - Low reflectance signal verification
  - Absolute gain
  - Linearity assessment
  - SNR

## FLARE 2021 Summer **Campaigns for UAVs** and Satellites

June 30-July 9, Hawaii Tests July 5-15, Texas Aug 9-13, South Dakota Aug 28-Sept 5 South Dakota: Landsat 8 Surface Validation

#### Contract:

Dr. Josh Hudson, FLARE Sales Manager – jhudson@labsphere.com M: 817-771-4847

![](_page_17_Picture_4.jpeg)

#### FLARE Customized events tailored to YOUR satellites

#### **Special Event! Limited Capacity! Act NOW!**

#### Complete spatial and radiometric calibration data sets for your satellite or constellation

FLARE is a revolutionary new cal/val technology that combines NIST traceable radiometry with the automated mirror tracking systems for rapid determination of important sensor metrics. Many point source targets in a single scene, all tuned to your sensor under the same atmospheric and solar conditions, yields a "one-shot" data-rich radiometric and spatial calibration event

#### What results can you expect from your FLARE campaign?

#### Spatial Performance from Point Sources

Cross-Track

FLARE Campaigns can diagnose focusing, smear, and band misregistration across a constellation -NOT detected with Lambertian targets

![](_page_17_Figure_12.jpeg)

![](_page_17_Picture_13.jpeg)

#### Video & Flyer HERE→

#### https://flare-network.com/latest-news/ © 2021 Labsphere Proprietary

## **FLARE Primary Benefits**

# New and independent method vs. classic vicarious calibration

- Radiometry & Spatial in ONE EVENT
- Fully automated synthetic target
- Small or Large target radiometry
- Impulse system response
- Band registration
- Geospatial Control Points
- Applies to all levels (L0, L1, L2, etc.)
- One-Sigma <3.5% VNIR Uncertainty

# Scales and applies to Airborne, UAV & Satellite calibration (mirror changes)

- Common radiometric calibration & harmonization for all data sources
- FLARE can be scaled from 0.01m to ≥1km GSD Pixels (OLCI, MODIS, etc.)
- Rapid identification of imaging or radiometric errors
- Speed commissioning and satellite interoperability

![](_page_18_Picture_15.jpeg)

![](_page_19_Picture_0.jpeg)

#### Better Calibration. Better Data. Better Decisions.

# **Thank you!** Questions and Comments may be directed to:

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