

A study of a phenomena deviated from random feature in EAS energy around 100TeV with new Kinki Array

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Abstract

An EAS observation system was rebuilt in July 1998 at Kinki University which consists of seven $0.25m^2$ plastic scintillators with fast PMTs. The system works well to detect EASs energy around 100TeV at sea level. In the observation of more than 10^3 hours, the detector records some phenomenae deviated from average feature in triggering rate, which means each series of the events has extremely small expected frequency in short term data acquisition period. For number of total events this set of consecutive events with upper limit of significant level of 3.6×10^{-2} was found. Although this is not strong evidence deviating from random Poisson distribution, but it has a value of 5σ from average rate.

The new Kinki Array is reported briefly and this phenomena is also reported in detail.

Key words: HE EAS, non-random component, consecutive EAS

1 Introduction

There have been published many reports on excess of event from expected feature in poisson random process [1, 2, 3]. One of them reported an existing of burst of 32 EASs in PeV energy within a 5 minutes time interval. They claimed the probability of such phenomena was 10^{-35} for assumption of each inde-

pendent EASs[1]. And more other experiment was reported suggesting a possibility of an existing of non-random component for time separation $< 40s$ in EAS above $10^{14}eV$ in the region of specific celestial coordinate [3].

2 Experimental Set Up

Since 1989 we have been searched non-random phenomenae in extensive air showers(EAS) events

from 1989 using compact detector with CAMAC electronic circuit. It had four $0.25m^2$ and four $1m^2$

plastic scintillation detectors. We have been published some reports elsewhere [4]. It had been worked till the end of 1996 when some parts of electric circuit were moved to another location.

We reestablished a new observation system in July 1998 at Kinki University (135.530E, 34.649N) at sea level. The system is consisting of seven $0.25m^2$ plastic scintillation detectors with PMT as the fast triggering detectors at vertice of two equilateral triangles and its centre as shown in Fig.-1. The length of sides are about $20.5m$ and $10.3m$ for large triangle and for small triangle respectively, the area is about $182.6m^2$. A block diagram of the electronic circuit is shown in Fig.-2. The data acquisition is performed

by *NIM – CAMAC* electronics which is controlled by using *CAMAC* computer system. It also has a *CAMAC* based GPS (Global Positioning System) module to record arrival time of the EAS in accuracy of micro-second. The trigger scheme is establish to detect the zenith angle of an EAS up to 45 degrees. It has been worked from the beginning of July 1998, although there were deficit of observations for several weeks owing to circuit trouble by a typhoon in summer, up to the present, more than 20k EAS events were recorded. In this report, about 16k events are analysed for searching non-random component in the data of EASs.

3 Data Analysis

There are many basic data analysing to check the performance of the detector system. For 16k events of EASs with 1228.15 hours data acquisition time, the average triggering rate is 4.61 ± 0.75 (min/ev). Frequencies of specific arrival time intervals are shown in Fig.-3(a) to (d). The triggering rates deriveing from fit to these distributions in the figure are consistent with above value within their errors. These values are consistent with previous observation system which had almost the same area with threshold energy around $10^{14}eV$. Some other triggering values for the different time intervals are calculated and shown in Table-1. In the process of data analysis, some EAS events arriving successively and we evaluate statistical quantities using statistics for these series of events. Extracted events exceed more than three σ (standard deviation:SD) are searched by scanning of all data and listed in Table-2. Table-2(a) is the number of events and their σ within 1 hour and Table-2(b) is that of within 2 hours. There is worth notice of exceeding four σ from the average number of triggering rate for 28 events within 1 hour and exceeding five σ for 52 events within 2 hours are found. Fig.-4 shows the time sequence of the events for the time of 13.00 to 16.00 on 19th July 1998. The beginning of time bin is shifted to find the maximum number of events in a certain time intervals for Fig.-4. There is a position having the maximum counts for 2 hours time interval which contains 53 events. To calculate the probability for this 53 events within 2 hours, i.e. $(53 - 26.03)/5.06 = 5.33\sigma$, it corresponds to the probability of 2.9×10^{-7} which gives expected events of 4.6×10^{-3} to the number of

total events from the accumulative normal distribution. These events are payed attention to analyse in detail for the candidate of deviating events expeted from average poisson feature.

The actual observed events within 1 hour and 2 hours are shown in Fig.-5(a), (b) and are compared with the expected events from poisson probability in Table-3. Furthermore, change of σ values as a function of the time binning is investigated by shift the beginning time. They are shown in Fig-6(a) and (b). There are two peaks for the events within 2 hours. The first peak corresponds to the maximum 5.33σ . It is easy to imagine of event arrival time sequece in Fig.-4 roughly from above distributions.

The following statistical methods are adopted to the series of 53 events within 2 hours.

(A) the Chebyshev's inequality expression
The Chebyshev's inequality gives an upper limit of the probability. The probability for observing any specific number of counts given the Poisson probability with a mean of x_0 and standard deviation σ is calculated by the Chebyshev's theorem which gives an upper limit of the probability for the frequency of exceed $k\sigma$ as following:

$$Prob(|x - x_0| > k\sigma) \leq 1/k^2.$$

In this case,

$$x_0 = m = 4.61 = \sigma^2$$

i.e. $\sigma = \sqrt{m} = 5.06, k\sigma = 53 - 26.03$ events, there-

fore, $k = 5.33$

$$Prob(|x - x_0| > 52) = (1/5.33)^2 = 0.036.$$

This value corresponds to the significance level of 3.6%.

(B) cumulative Poisson probability:

The cumulative Poisson probability function is defined by

$$P_x(< k) = 1 - P(k, x)$$

where $P_x(< k)$ is the probability that the number of Poisson random events occurring will be between 0 and $k - 1$ inclusive, x is the mean number, $P(k, x)$ is the incomplete gamma function [5].

The calculated values of the cumulative Poisson probability with the real data are given in Table-3(a) and (b).

4 Summary

The upper limit of significant level of $\alpha = 3.6 \times 10^{-2}$, that is, the confidence level of 96% is obtained to accept an hypothesis for 52 events within 2 hours deviating from random Poisson process. From cumulative Poisson probability, the number of expected events 1.28×10^{-3} is obtained for the total events. Although this value is not strong evidence deviating from random phenomena. However, if we assume the normal distribution for the errors of triggering rate, the probability of order of 10^{-7} might be obtained. For validity of the candidate in detail, there needs

some other statistical methods and also needs to improve statistics of the data.

There is a topic of some gamma-ray bursts (GRBs) coming from a soft gamma-ray repeater (SGR), Magnetar SGR1900+14, on 27th August 1998 [6]. We tried to find a possibility of event excess from random distribution but no excess was found in the data at the time. We are keeping to watch the EAS events due to GRBs from the Magnetars.

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6.51 ± 2.47	evs./30min.
13.01 ± 3.55	evs./hr.
26.03 ± 5.06	evs./2hrs.

Table-1: EAS events triggering rates.

(a)

date	time	events	$\sigma/30min.$
1998/07/11	00:21:56-01:51:56	15	3.33
1998/07/11	00:51:56-02:21:56	16	3.72
1998/07/11	01:21:56-02:51:56	16	3.72
1998/07/17	01:51:56-03:21:56	15	3.33
1998/07/19	02:21:56-03:51:56	18	4.50
1998/07/24	02:51:56-04:21:56	15	3.33
1998/08/01	03:21:56-04:51:56	15	3.33
1998/08/03	03:51:56-05:21:56	15	3.33
1998/08/29	04:21:56-05:51:56	15	3.33
1998/08/31	04:51:56-06:21:56	16	3.72
1998/09/04	05:21:56-06:51:56	15	3.33
1998/09/15	05:51:56-07:21:56	15	3.33

(b)

date	time	events	$\sigma/hr.$
1998/07/11	01:21:56-02:21:56	26	3.60
1998/07/16	13:55:31-14:55:31	24	3.04
1998/07/19	13:34:01-14:34:01	30	4.79
1998/07/19	14:55:31-15:55:31	24	3.04
1998/07/24	21:29:37-22:29:37	2	-3.05
1998/08/04	00:09:31-01:09:31	2	-3.05
1998/08/04	02:09:31-03:09:31	2	-3.05
1998/09/02	21:17:08-22:17:08	2	-3.05
1998/09/16	18:16:28-19:16:28	24	3.04
1998/09/16	20:16:28-21:16:28	24	3.04
1998/09/16	22:16:28-23:16:28	30	4.71

(c)

date	time	events	$\sigma/2hrs.$
1998/07/13	23:21:56-01:21:56	9	-3.33
1998/07/14	23:21:56-01:21:56	10	-3.14
1998/07/19	13:34:01-15:34:01	53	5.33
1998/07/25	22:29:37-00:29:37	10	-3.14
1998/07/29	00:29:37-02:29:37	8	-3.53
1998/07/30	21:09:31-23:09:31	9	-3.33
1998/07/31	21:09:31-23:09:31	10	-3.14
1998/07/31	23:09:31-01:09:31	7	-3.72
1998/08/03	01:09:31-03:09:31	10	-3.14
1998/08/03	13:09:31-15:09:31	43	3.32
1998/09/12	23:50:20-01:50:20	9	-3.33
1998/09/13	05:50:20-07:50:20	42	3.13
1998/09/14	13:50:20-17:50:20	42	3.13
1998/09/16	17:16:28-19:16:28	44	3.52
1998/09/16	19:16:28-21:16:28	44	3.52
1998/09/16	21:16:28-23:16:28	44	3.52
1998/09/19	23:16:28-01:16:28	9	-3.33

Table-2(a): List of excess events greater than 3σ (a) within 30 min, (b) within 1 hour, and (c) within 2 hours time interval.

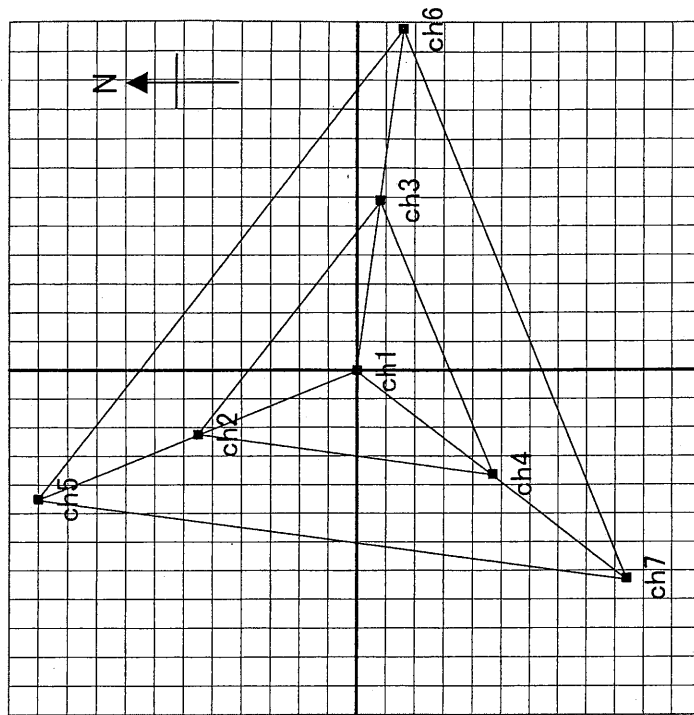


Fig.-1: The new Kinki Array was rebuilt. The arrangement of seven plastic scintillation detectors are shown. A lattice length is 1m.

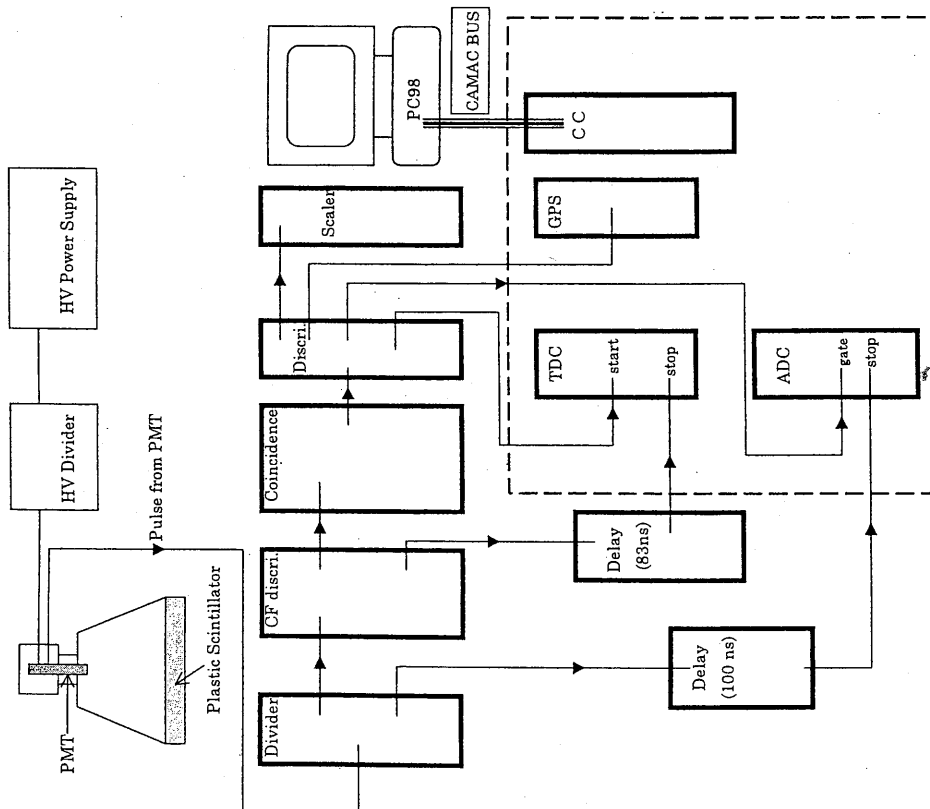


Fig.-2: The block diagram of the detection system which consists of NIM - CAMAC electronic circuit.

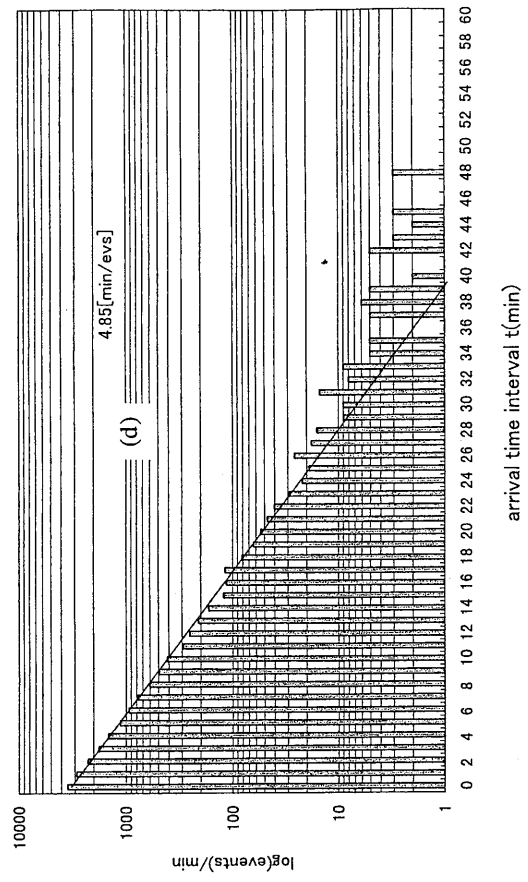
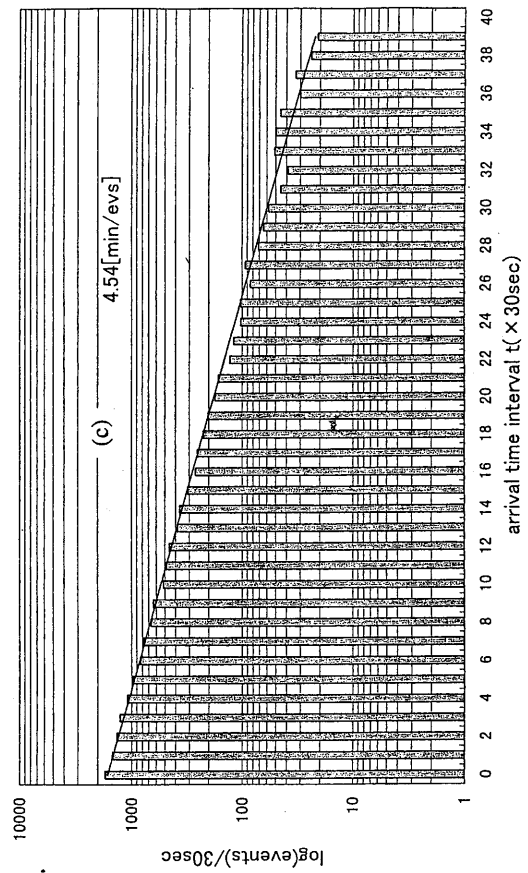
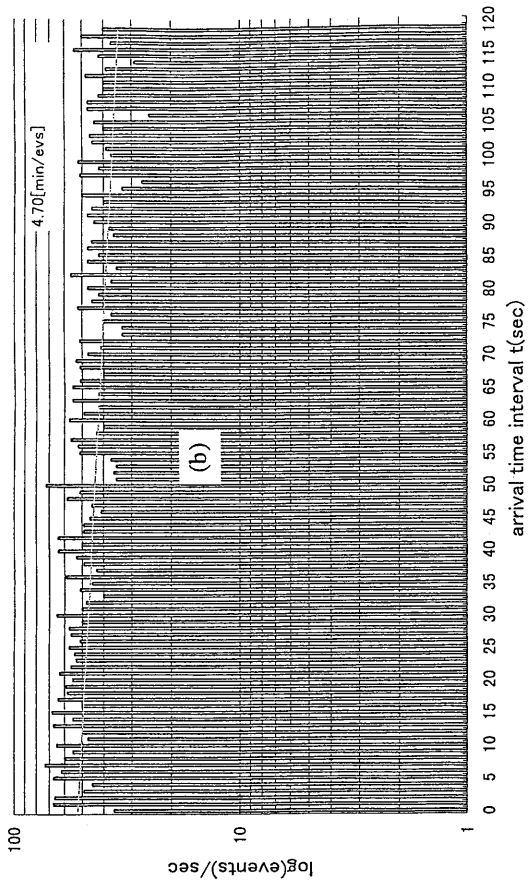
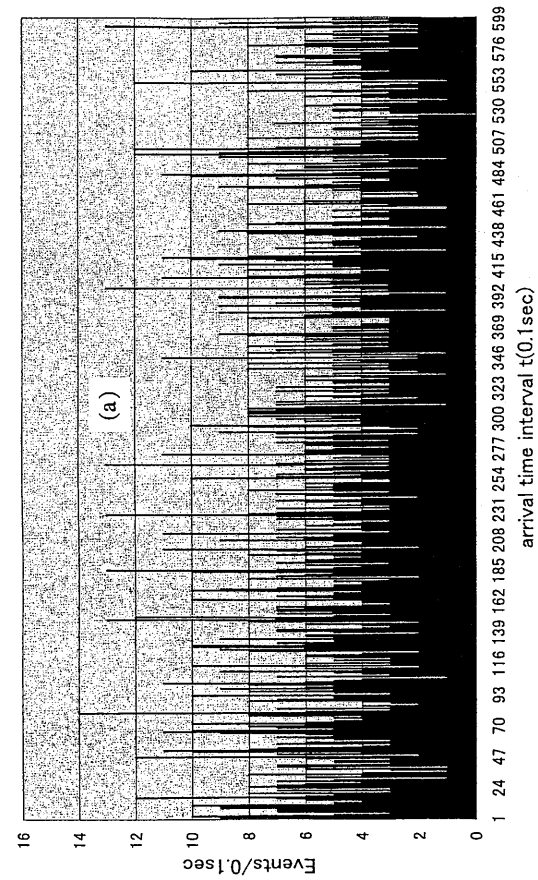


Fig.-3: The distributions of EAS arrival time interval for intervals of (a) 0.1sec, (b) 1sec, (c) 30sec, and (d) 1min.

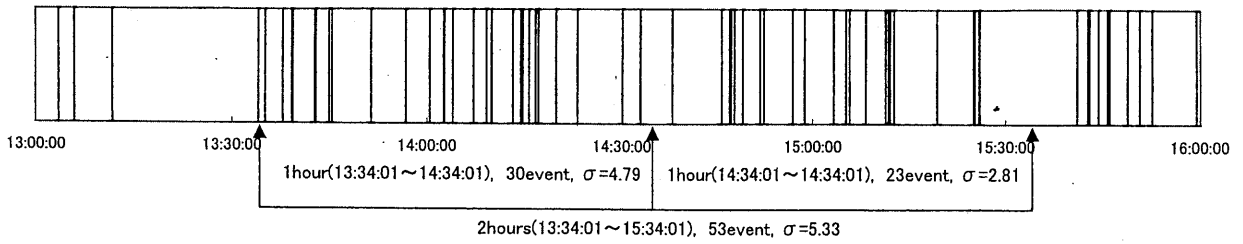


Fig.-4: The time sequence of events for exceeding 5σ within 2 hours is shown. It contains a series of 53 events in the indicated period. Also deviding to 1 hour region is shown. It include 30 events corresponding to 4.79σ These are phenomenae of the maximum consecutive events in the Table-2(b) and (c).

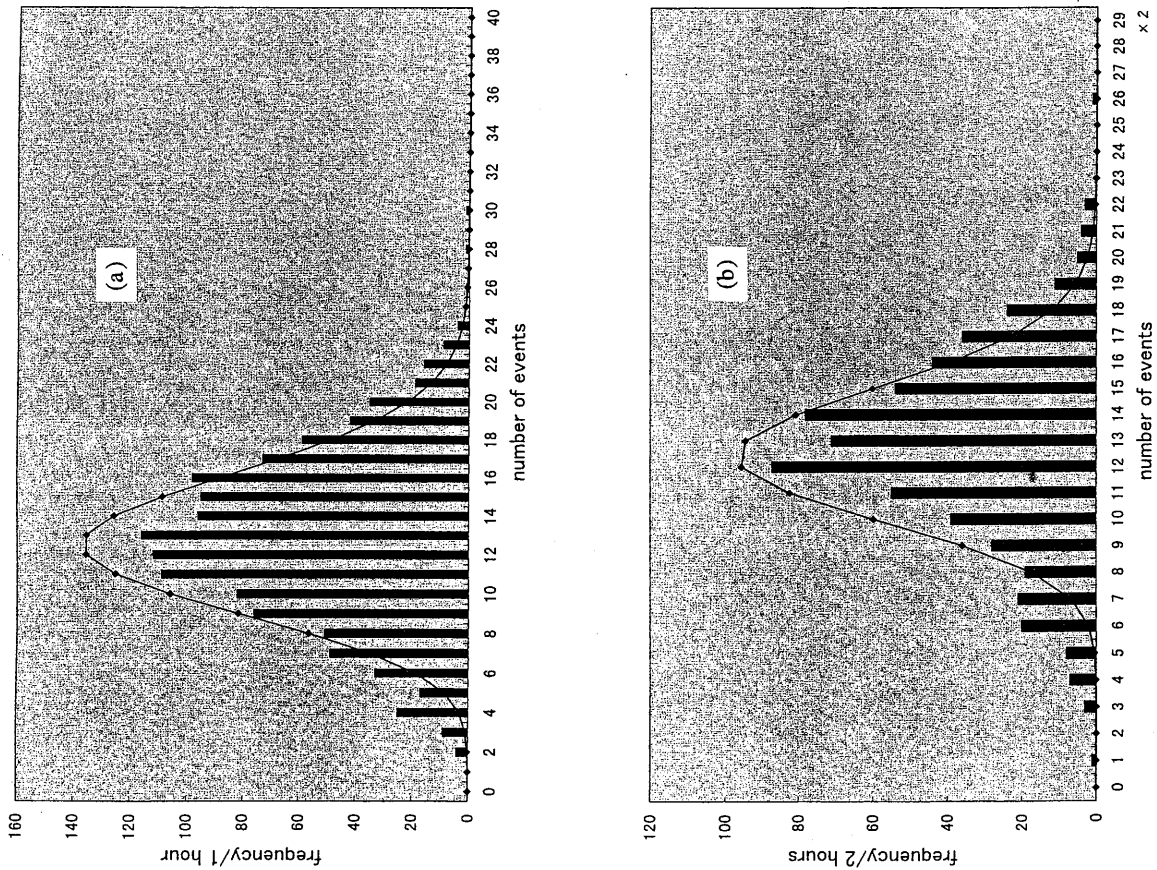


Fig.-5: The distributions of the number of events vs. frequency for (a) 1 hour time interval, and (b) 2 hours time interval. The expected events from the prediction of Poisson distribution are also drawn.

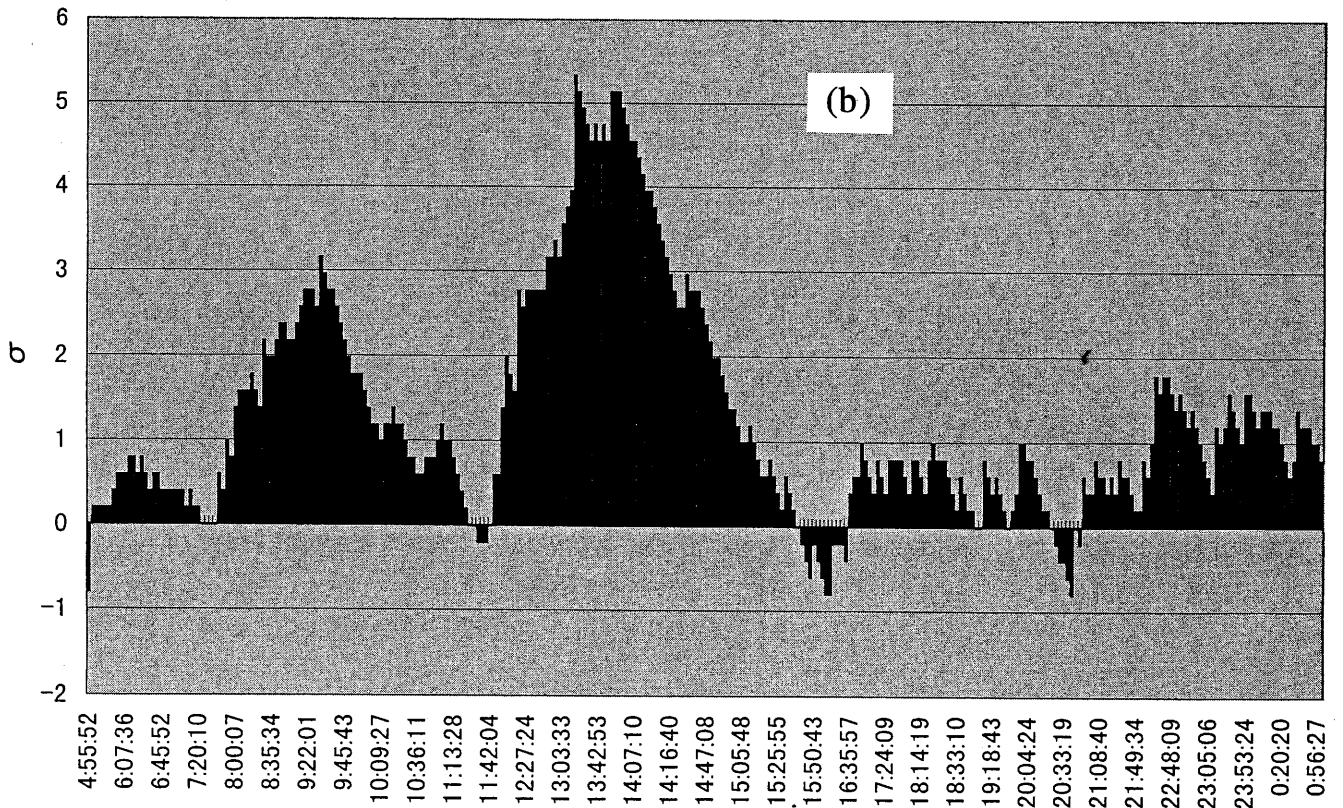
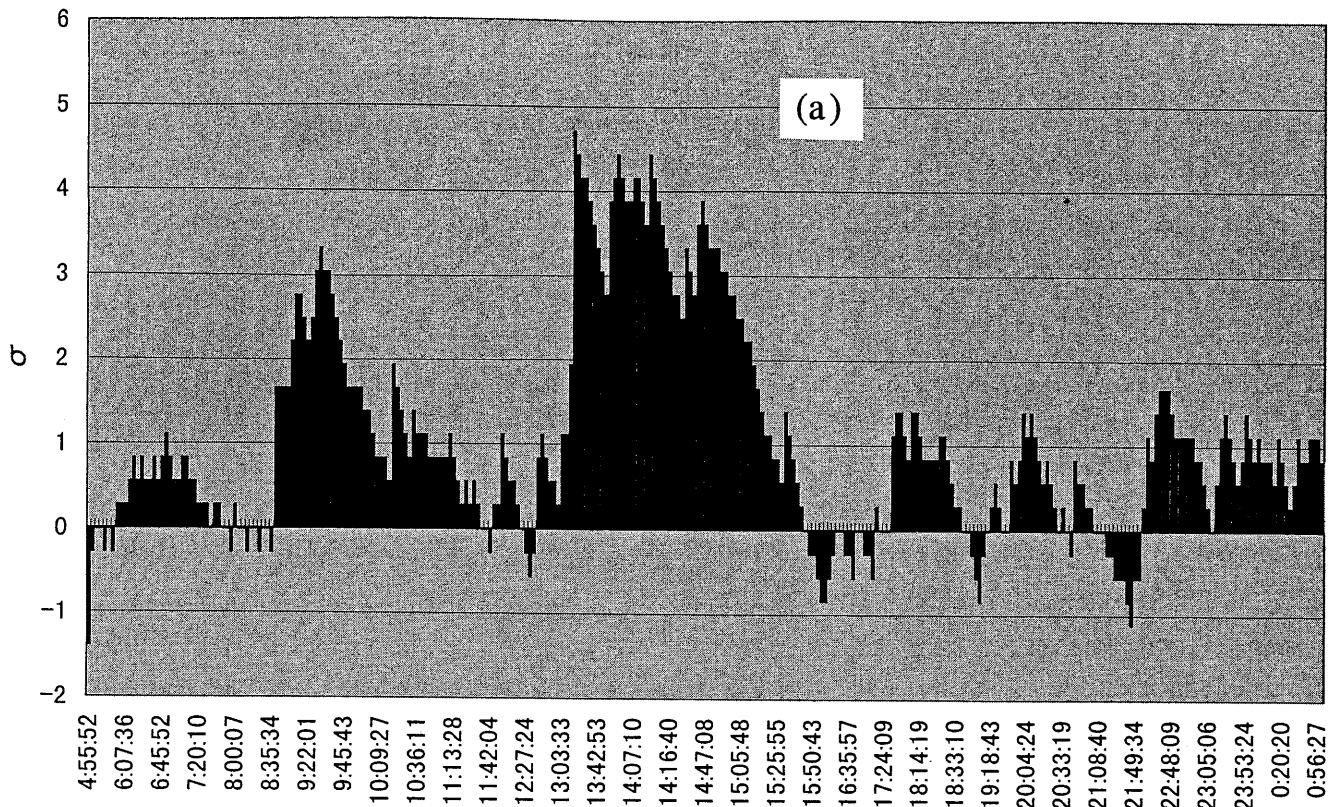


Fig.-6: The distributions of σ values as a function of the beginning time for (a) 1 hour time interval, and (b) 2 hours time interval.