

Exploration of co-sputtered Ta₂O₅–ZrO₂ thin films for gravitational-wave detectors

M Abernathy^{1,2}, A Amato³, A Ananyeva⁴, S Angelova⁵,
B Baloukas⁶, R Bassiri⁷, G Billingsley⁴, R Birney^{5,8},
G Cagnoli³, M Canepa^{9,10}, M Coulon¹¹, J Degallaix¹¹,
A Di Michele¹², M A Fazio¹³, M M Fejer⁷, D Forest¹¹,
C Gier⁵, M Granata^{11,*}, A M Gretarsson¹⁴,
E M Gretarsson¹⁵, E Gustafson⁴, E J Hough¹⁶, M Irving¹⁷,
É Lalande¹⁸, C Lévesque¹⁸, A W Lussier¹⁸, A Markosyan⁷,
I W Martin¹⁶, L Martinu⁶, B Maynard¹⁷, C S Menoni^{13,19},
C Michel¹¹, P G Murray¹⁶, C Osthelder⁴, S Penn¹⁷,
L Pinard¹¹, K Prasai⁷, S Reid⁵, R Robie¹⁶, S Rowan¹⁶,
B Sassolas¹¹, F Schiettekatte¹⁸, R Shink¹⁸, S Tait¹⁶,
J Teillon¹¹, G Vajente^{4,*}, M Ward¹⁸ and L Yang¹⁹

¹ Department of Physics, American University, Washington, DC, United States of America

² Johns Hopkins University Applied Physics Laboratory, Laurel, MD, United States of America

³ Université de Lyon, Université Claude Bernard Lyon 1, CNRS, Institut Lumière Matière, F-69622 Villeurbanne, France

⁴ LIGO Laboratory, California Institute of Technology, Pasadena, CA, United States of America

⁵ SUPA, Department of Biomedical Engineering, University of Strathclyde, Glasgow, United Kingdom

⁶ École Polytechnique de Montréal, Montréal, Québec, Canada

⁷ Edward L Ginzton Laboratory, Stanford University, Stanford, CA 94305, United States of America

⁸ SUPA, Institute of Thin Films, Sensors and Imaging, University of the West of Scotland, Paisley, United Kingdom

⁹ OPTMATLAB, Dipartimento di Fisica, Università di Genova, Via Dodecaneso 33, 16146 Genova, Italy

¹⁰ INFN, Sezione di Genova, Via Dodecaneso 33, 16146 Genova, Italy

¹¹ Laboratoire des Matériaux Avancés, Institut de Physique des 2 Infinis de Lyon, CNRS/IN2P3, Université de Lyon, Université Claude Bernard Lyon 1, F-69622 Villeurbanne, France

¹² Dipartimento di Fisica e Geologia, Università degli Studi di Perugia, Via Pascoli, 06123 Perugia, Italy

*Authors to whom any correspondence should be addressed.



Original content from this work may be used under the terms of the [Creative Commons Attribution 4.0 licence](#). Any further distribution of this work must maintain attribution to the author(s) and the title of the work, journal citation and DOI.

¹³ Department of Electrical and Computer Engineering, Colorado State University, Fort Collins, CO, United States of America

¹⁴ Embry-Riddle Aeronautical University, 3700 Willow Creek Rd., Prescott, AZ 86301, United States of America

¹⁵ Wyant College of Optical Sciences, University of Arizona, 1630 E University Blvd., Tucson, AZ 85721, United States of America

¹⁶ SUPA, Institute for Gravitational Research, University of Glasgow, Glasgow, United Kingdom

¹⁷ Department of Physics, Hobart and William Smith Colleges, Geneva, NY, United States of America

¹⁸ Département de physique, Université de Montréal, Montréal, Québec, Canada

¹⁹ Department of Chemistry, Colorado State University, Fort Collins, CO, United States of America

E-mail: m.granata@lma.in2p3.fr and vajente@caltech.edu

Received 29 March 2021, revised 29 June 2021

Accepted for publication 5 August 2021

Published 13 September 2021



Abstract

We report on the development and extensive characterization of co-sputtered tantalum–zirconia (Ta_2O_5 - ZrO_2) thin films, with the goal to decrease coating Brownian noise in present and future gravitational-wave detectors. We tested a variety of sputtering processes of different energies and deposition rates, and we considered the effect of different values of cation ratio $\eta = Zr/(Zr + Ta)$ and of post-deposition heat treatment temperature T_a on the optical and mechanical properties of the films. Co-sputtered zirconia proved to be an efficient way to frustrate crystallization in tantalum thin films, allowing for a substantial increase of the maximum annealing temperature and hence for a decrease of coating mechanical loss φ_c . The lowest average coating loss was observed for an ion-beam sputtered sample with $\eta = 0.485 \pm 0.004$ annealed at $800\text{ }^\circ\text{C}$, yielding $\overline{\varphi}_c = 1.8 \times 10^{-4}$ rad. All coating samples showed cracks after annealing. Although in principle our measurements are sensitive to such defects, we found no evidence that our results were affected. The issue could be solved, at least for ion-beam sputtered coatings, by decreasing heating and cooling rates down to $7\text{ }^\circ\text{C h}^{-1}$. While we observed as little optical absorption as in the coatings of current gravitational-wave interferometers (0.5 parts per million), further development will be needed to decrease light scattering and avoid the formation of defects upon annealing.

Keywords: thin films, thermal noise, gravitational-wave detectors

(Some figures may appear in colour only in the online journal)

1. Introduction

The detection of gravitational-wave (GW) signals from astrophysical systems [9] has ushered in a new era in astronomy [10, 11], allowing the study of previously invisible systems. The current generation of GW interferometric detectors are sensitive enough to allow a confident candidate event about once per week [13, 14]. Increasing the sensitivity of the detectors is

important not only to increase the detection rate but also to improve the signal to noise ratio and sky localization of events, allowing for a more precise probing of the properties of nuclear matter in neutron stars [12] and of the fundamental physics of gravitation and general relativity [8].

Ironically, the sensitivity of current detectors to signals from very large and massive astrophysical systems is limited by random noise in microscopic systems, that is, fluctuations resulting from the internal mechanical dissipation in the material constituting the mirrors used as test masses [42]. In the most sensitive region of the detection band of Advanced LIGO [7] and Advanced Virgo [15], between 50 and 300 Hz, the sensitivity is limited by an equal contribution of quantum noise [26] and thermal noise, the latter being dominated by coating Brownian noise [72]. Indeed, all GW detectors currently operating are laser interferometers with km-scale optical resonators [56], where the test masses used to probe the space time metric fluctuations are high-quality fused silica mirrors with high-reflection (HR) dielectric coatings on the front surfaces [27, 46, 84]. Of all the sources of thermal noise in the mirrors, the Brownian motion of particles in the coating layers is the limiting factor and hence needs to be reduced for future sensitivity upgrades [40, 48, 50].

The dielectric coatings used in the current generation of GW detectors are Bragg reflectors [75] made of alternate ion-beam sputtered (IBS) thin layers of silica (SiO_2) and tantalum–titania ($\text{Ta}_2\text{O}_5\text{--TiO}_2$) of optimized thickness [58]. Silica is the low refractive index material ($n = 1.45$ at 1064 nm), and the tantalum–titania is the high refractive index material ($n = 2.07$ at 1064 nm) [32]. These materials and their synthesis by various methods have been extensively investigated in the context of optical materials in general and interferometric filters for photonic, lasers, and astronomy applications in particular [75], including for example their mechanical properties [3, 32, 38], showing that their performance is intimately related to parameters such as their internal stress and growth parameters [32].

The detection of GW signals is performed by monitoring the change in the relative distance between the mirrors, measured as a change in the phase of the laser beams [78]. Therefore, any fluctuation in the thickness or in the optical properties of the coating materials acts like a spurious displacement noise. Thermo-refractive and thermo-elastic fluctuations are not relevant [28], and the dominant source of phase noise is due to the Brownian motion of all the layers. The fluctuation dissipation theorem [1] links Brownian noise to the rate of energy dissipation in the coating material due to internal friction. The amount of friction is measured by the material loss angle φ , that is the fraction of energy lost per cycle when the material is driven by a sinusoidal external force. In the coating materials used so far, the worst offender is tantalum–titania, with measured values $(2.3\text{--}3.4) \times 10^{-4}$ [32, 82]. The silica layers have one order of magnitude lower loss angle, $(2.3\text{--}4.0) \times 10^{-5}$ [32, 44].

The goal for the GW detector upgrades in the next 2–5 years is to roughly double the sensitivity to GW inspiral signals [40], which translates to an improvement of the same factor of the sensitivity in the 100 Hz region. In Advanced LIGO, for instance, this sets a target of $6.6 \times 10^{-21} \text{ m Hz}^{-1/2}$ at 100 Hz in terms of coating Brownian noise, to be compared with the current estimated level of $1.3 \times 10^{-20} \text{ m Hz}^{-1/2}$ [23]. The Brownian noise amplitude spectral density scales with the square root of the total loss angle of the HR coating [36]. Therefore the upgrade goal translates into improving the high refractive index material loss angle by at least of a factor 4. Marginal or no improvement is necessary for the low refractive index material. For this reason, most of the coating research in the Virgo and LIGO collaborations focused on finding an alternative material to replace the tantalum–titania layers [33, 84], while keeping the same outstanding optical properties (absorption, scattering, etc) of the HR coatings. The main line of research explores the co-sputtering of other oxides with tantalum, following the success of

co-sputtered tantalum–titania coatings [37] which allowed a reduction of the coating loss angle by about 25% [32] with respect to tantalum.

Beginning in 2012, one of the authors (SP) characterized a series of zirconia-based coatings. That investigation was motivated by zirconia (ZrO_2) being a high-index material ($n \geq 2$ at 1064 nm, like tantalum) possessing a similar molecular structure and physical parameters to silica. However, unlike silica, when treated at high temperatures, zirconia will crystallize unless stabilized. Tewg *et al* [67] showed that the addition of zirconia to tantalum thin films allowed for a higher maximum heating temperature, before the onset of crystallization. Post-deposition heat treatment, referred to as *annealing* hereafter, is a standard procedure to decrease the internal stress, the optical absorption and the mechanical loss of coatings [32]. As a general empirical rule, the higher the annealing peak temperature, the lower the resulting coating mechanical loss. In the work of Tewg *et al*, the annealing had a duration of 10 min, while typical annealing times for HR coatings used in GW detectors are of the order of 10 h [32]. Indeed, higher annealing temperatures (up to 900 °C) and duration would be beneficial also for the silica layers, decreasing their mechanical loss [17, 31] as a result of structural relaxation, leading to HR coatings with overall lower loss angle and hence lower thermal noise.

This paper focuses on collaborative work, carried out by several groups within the LIGO and Virgo Collaborations to produce and characterize co-sputtered tantalum–zirconia thin films, for possible use as high refractive index layers for the new sets of mirrors of the Advanced LIGO + and Advanced Virgo + detector upgrades. We considered the effect of adding various amounts of zirconia, and measured the optical and mechanical properties of the films as functions of their chemical composition and annealing temperature. As coating mechanical loss depends in general on the nature of the growth technique, we tested a variety of sputtering processes of different energies and deposition rates. A detailed description of the coating samples is given in section 2. Different techniques and methodologies have been used to characterize the optical and mechanical properties of the coating samples, described with their results in sections 3 and 4. In sections 4 and 5 we show that the main effect of co-sputtering zirconia with tantalum is a significant increase of the maximum annealing temperature T_a , hence a reduction of coating mechanical loss. The main issues we encountered are cracking of all annealed single-layer samples and bubble-like defects in annealed IBS HR coatings. Later we found that cracking could be avoided by drastically slowing down the annealing heating and cooling rates of the samples. Bubble-like defects, however, remain a pending issue to date. In section 6 we present the results of Raman spectroscopy measurements, and discuss the possibility of using Raman peaks as a probe of the state of relaxation or a predictor of the loss angle in a material. Finally, we discuss the results and their implications in section 7.

2. Deposition methods and film composition

The co-sputtered tantalum–zirconia coating samples used for this work were produced by one industrial manufacturer and several academic research groups, using different growth techniques:

- Ion-beam sputtering (IBS), at MLD Technologies (MLD) and Laboratoire des Matériaux Avancés (LMA)
- Electron-cyclotron-resonance ion-beam deposition (ECR-IBD), at University of Strathclyde (UoS)
- Reactive biased-target deposition (RBTD), at Colorado State University (CSU), Fort Collins

Table 1. Measured atomic cation ratio $\eta = \text{Zr}/(\text{Zr} + \text{Ta})$, analysis method used to determine η and density ρ of co-sputtered Ta_2O_5 – ZrO_2 thin films. RF-MS samples had Ta deposited by RF-MS and Zr deposited by DC-MS, HiPIMS samples had Ta deposited by HiPIMS and Zr deposited by RF-MS. Density values are plotted and discussed in section 3.5, where correlation between oxygen concentration, density and refractive index are considered.

Sample	Process	η (at.%)	Method	ρ (g cm ⁻³)
MLD2014	IBS	48.5 ± 0.4	EDX	
MLD2018	IBS	50.2 ± 0.6	EDX	
LMA	IBS	19 ± 1	EDX	6.9 ± 0.1
UoS	ECR-IBD	34.0 ± 0.5	EDX	
UMP 551	RF-MS	41 ± 3	RBS	6.1 ± 0.2
UMP 554	RF-MS	43 ± 3	RBS	5.9 ± 0.2
UMP 658	HiPIMS	24 ± 3	RBS	7.0 ± 0.2
UMP 659	HiPIMS	24 ± 2	RBS	7.1 ± 0.2
UMP 542	HiPIMS	7 ± 2	RBS	7.3 ± 0.2
UMP 664	HiPIMS	7 ± 2	RBS	7.1 ± 0.2
UMP 678	HiPIMS	47 ± 2	RBS	6.6 ± 0.2
UMP 680	HiPIMS	47 ± 2	RBS	6.5 ± 0.2
CSU II	RBTD	23 ± 1	XPS	7.1^{a}
CSU III	RBTD	54 ± 3	XPS	7.1^{a}

^aEstimated from component values and concentration.

- High-power impulse (HiPI) and radio-frequency (RF) magnetron sputtering (MS), at Université de Montréal and Polytechnique Montréal (UMP).

Acronyms of growth techniques and groups are used throughout this paper to identify the samples. The measured atomic cation ratio $\eta = \text{Zr}/(\text{Zr} + \text{Ta})$ and density ρ of all samples considered in this work are summarized in table 1 and plotted and discussed in section 3.5. Obtained values of η depended on the different coating facilities used, in terms of their equipment, configuration, accessible growth parameters, and availability. Considering all our sample sets, we eventually managed to reproduce almost every cation ratio already tested by Tewg *et al* in their study [67]. Details on the different growth processes, sample specifications and characterizations follow below.

After coating deposition, all coating samples were annealed in air at increasing peak temperature T_a by steps of 50 or 100 °C, starting from 300 °C, until they showed signs of crystallization. The annealing time for each soaking temperature T_a was 10 h, except for ECR-IBD samples for which it was 5 h.

2.1. IBS

Ion-beam-sputtered films were produced by MLD Technologies and by the LMA research group. Details are described below.

2.1.1. MLD Technologies. The provider is a manufacturer of high performance and specialty optical coatings located in Mountain View, CA, with coating facilities in Eugene, OR. The company deposited the coatings using standard IBS, where a beam of Ar ions is directed at a metal target and the sputtered metal atoms are oxidized in a low density O₂ atmosphere before striking the substrate. For all deposition runs, MLD coated: (i) thin super-polished disk substrates of Corning 7980 fused silica that were 76.2 mm in diameter and 2.5 mm thick; (ii) a ~100 μm thick fused-silica cantilever welded to a 10 mm × 10 mm × 3 mm

clamping block; (iii) and a number of $35\text{ mm} \times 5\text{ mm} \times \sim 65\text{ }\mu\text{m}$ silicon cantilevers with a $10\text{ mm} \times 5\text{ mm} \times \sim 550\text{ }\mu\text{m}$ clamping block. As is standard practice, each run included fused-silica witness samples with 25.4 mm diameter that were used to test the coating composition and thickness.

MLD conducted 2 runs, whose samples are referenced as MLD2014 and MLD2018, of identical specifications: the design goal was a single 500 nm thick tantalum-zirconia coating with atomic cation ratio $\eta = \text{Zr}/(\text{Zr} + \text{Ta}) \sim 36\%$, as specified by Tewg *et al* [67] for achieving amorphous coatings stable to peak annealing temperatures of 800 °C, where the ratio was intended to denote atomic number density. MLD performed post-deposition energy-dispersive x-ray spectroscopy (EDX) measurements that reported cation ratios in agreement with that goal. However, while Tewg *et al* reported atomic number densities, MLD reported ratios were later learned to be molecular number densities. Moreover, a reanalysis of the EDX data showed that the η was $48.5 \pm 0.4\%$ for MLD2014 and $50.2 \pm 0.6\%$ for MLD2018. Thus the actual molecular number densities were $\text{ZrO}_2 = 65.3 \pm 0.4\%$ and $\text{Ta}_2\text{O}_5 = 34.7 \pm 0.4\%$ for MLD2014 and $\text{ZrO}_2 = 66.8 \pm 0.6\%$ and $\text{Ta}_2\text{O}_5 = 33.2 \pm 0.6\%$ for MLD2018. MLD also reported coating thicknesses of 590 nm for MLD2014 and 483 nm for MLD2018.

After coating deposition, to test the effect of annealing on coating mechanical loss, the fused-silica disk and the silicon cantilevers were progressively annealed with steps of increasingly high temperatures, ranging from 300 to 800°C.

2.1.2. LMA. Co-sputtered Ta_2O_5 – ZrO_2 coatings were produced with the custom-developed so-called *Grand Coater* (with a volume of 10 m^3) used to coat the mirrors of the Advanced LIGO, Advanced Virgo and KAGRA detectors [27, 32, 46]. Prior to deposition, the base pressure was less than 10^{-7} mbar . Ar gas was fed into the sputtering source while O₂ gas was fed directly into the vacuum chamber, for a total working pressure of $2 \times 10^{-4}\text{ mbar}$. Energy and current of the sputtering Ar ions were of the order of 1 keV and 0.1 A, respectively. During deposition, the co-sputtered Ta_2O_5 and ZrO_2 particles impinged on substrates heated up to about 120 °C.

Single layers with a thickness between 500 and 650 nm have been grown on various substrates for different purposes: (i) on fused-silica witness samples (25.4 mm in diameter, 6 mm thick) for transmission spectrophotometry; (ii) on silicon wafers (75 mm in diameter, 0.5 mm thick) for reflection spectroscopic ellipsometry and EDX analyses; (iii) on a Corning 7980 fused-silica disk-shaped mechanical resonator of 75 mm in diameter and 1 mm thick for measuring coating mechanical loss, Young modulus, Poisson ratio and density.

Additionally, Ta_2O_5 – ZrO_2 /SiO₂ HR coatings with the same design of those of Advanced LIGO and Advanced Virgo most reflective mirrors (end test masses, ETMs) [32] have been deposited on micro-polished fused-silica witness samples for optical absorption and scattering measurements. By design, those ETM coatings have a transmission of 5 parts per million (ppm) at 1064 nm, determined by 38 layers of optimized thickness [58], for a total thickness of 5.9 μm [32].

Prior to coating deposition, to release the internal stress due to manufacturing and induce relaxation, the fused-silica disk was annealed in air at 900 °C for 10 h. After deposition, all coating samples were annealed in air for 10 h: optical samples were annealed at 500 °C, the disk underwent consecutive annealing treatments of increasing peak temperature (500, 600 and 700 °C) and the mechanical loss of its coating was measured after each annealing step.

We used an analytical balance to measure the mass of the disk before and after the coating deposition, as well as after the 500 °C annealing. To estimate the surface area coated in the deposition process, we measured its diameter with a Vernier caliper. As the coating thickness

was known with high accuracy from the optical characterization (see section 3.1.2), the coating density $\rho = 6.9 \pm 0.1 \text{ g cm}^{-3}$ was straightforwardly estimated as the mass-to-volume ratio.

We used a Zeiss LEO 1525 field-emission scanning electron microscope (SEM) and a Bruker Quantax system equipped with a Peltier-cooled XFlash 410 M silicon drift detector to analyze the surface and elemental composition of the single-layer coatings. The SEM beam was set to 15 keV for the survey. We performed multiple EDX analyses on different coating sample spots and with different magnifications (from $100\times$ to $5000\times$). Semi-quantitative (standardless) results were based on a peak-to-background evaluation method of atomic number, self-absorption and fluorescence effects (P/B-ZAF correction) and a series fit deconvolution model provided by the Bruker Esprit 1.9 software. Using the self-calibrating P/B-ZAF standard-based analysis, no system calibration had to be performed. According to our EDX analyses, our thin films had an average atomic cation ratio $\eta = \text{Zr}/(\text{Zr} + \text{Ta}) = 19 \pm 1\%$. We also found that they incorporated 2.9 ± 0.6 at.% of Ar.

2.2. ECR-IBD

Coatings were deposited using a custom-built ion beam deposition system that utilized three ECR ion sources for sputtering [20]. The ion beams are generated by injection of Ar gas into a resonance microwave cavity, with ~ 3 W injected power at 2.45 GHz. The confinement of the plasma permits the extraction of ions through a single aperture, enabling a larger range of stable extraction potentials to be explored, with low ion current. The extraction potential was 10 kV and ion currents ranged between 0.2–0.6 mA per source, resulting in a deposition rate of 0.001 nm s^{-1} . Base pressure was below 1×10^{-5} mbar prior to deposition, with pressure rising to 8×10^{-5} mbar with Ar injected through the ion sources, and increasing to 1.2×10^{-4} mbar with reactive O₂ gas fed into the chamber. The targets used during the deposition process are two circular 152.4 mm diameter targets of pure tantalum and zirconia provided by Scotech, UK, both with 99.99% purity. During the deposition process, two ion beams were directed onto the tantalum target and one beam on the zirconia target, with the aim for cation ratio $\eta = \text{Zr}/(\text{Zr} + \text{Ta})$ of $\sim 33\%$. Fused-silica witness samples and cantilevers were placed on a rotating substrate holder to ensure coating uniformity.

The cantilevers, used for coating mechanical loss measurements [55], were fabricated from 45 mm × 5 mm × (175 ± 3) μm strips of fused silica, flame welded to 10 mm × 10 mm × 3.1 mm clamping blocks [55]. They were cleaned in boiling H₂O₂ (30% w/w in H₂O) with 1 mol l⁻¹ KOH dissolved, in order to remove the majority of particulates associated with the welding process, then annealed in air for 5 h at 950 °C prior to measurement and deposition.

EDX was used to quantify the cation ratio $\eta = \text{Zr}/(\text{Zr} + \text{Ta})$, found to be $34.0 \pm 0.5\%$.

2.3. RBTD

Coatings were deposited using a LANS system by 4Wave, Inc. [63], where a low energy Ar ion plume is produced using an end-Hall ion source operated at 60 V and 5.8 A, with an Ar flow of 27 sccm. For neutralization, a hollow cathode electron source is employed at an emission current of 6.2 A with an Ar flow of 25 sccm. The Ar ions are attracted to a biased metallic target, leading to sputtering of the target material. For oxide coatings, O₂ is introduced into the chamber during the process. Substrates are placed in a sample holder facing the ion source, and rotating at 20 rpm during deposition to ensure uniformity of the film. The chamber base pressure is below 1×10^{-7} mbar, while the process pressure is about 9×10^{-4} mbar.

The LANS system has the ability to simultaneously bias up to three targets to produce mixed oxide coatings. An asymmetric, bipolar DC pulse is used to bias the targets with a negative voltage of -800 V and a positive voltage of 10 V and a period of 100 μ s. For a

Table 2. Deposition conditions for Ta_2O_5 – ZrO_2 thin films grown via RBTD. Pure Ta_2O_5 and ZrO_2 samples (I and IV, respectively) are included for reference.

Film	O_2 flow (sccm)	Target	Pulse width (μs)	Deposition rate (nm s^{-1})
CSU I	12	Ta	2	0.0212 ± 0.0001
CSU II	12	Ta–Zr	2–54	0.0294 ± 0.0001
CSU III	12	Ta–Zr	47–2	0.0217 ± 0.0001
CSU IV	12	Zr	2	0.0116 ± 0.0001

Table 3. Atomic concentrations obtained from XPS analysis at the surface of as-deposited Ta_2O_5 – ZrO_2 thin films grown via RBTD.

Film	Elemental concentration (at.%)				Cation ratio $\text{Zr}/(\text{Zr} + \text{Ta})$
	C	O	Zr	Ta	
CSU II	38 ± 1	44 ± 1	4.3 ± 0.1	14.1 ± 0.1	0.23 ± 0.01
CSU III	39 ± 1	42 ± 1	10.4 ± 0.4	8.7 ± 0.1	0.54 ± 0.03

binary oxide mixture, two targets are simultaneously biased and different mixture proportions of the materials are achieved by controlling the individual pulse duration of each target. Ta and Zr targets were employed to deposit tantalum–zirconia thin films, and the corresponding deposition conditions are presented in table 2. Deposition rates for RBTD are typically an order of magnitude lower than with a gridded IBS system, ranging from 0.0212 nm s^{-1} for tantalum to 0.0116 nm s^{-1} for zirconia. Films with a thickness between 200–300 nm were grown on: (i) 25.4 mm diameter and 6.35 mm thick UV grade fused-silica substrates; (ii) 75 mm diameter and 1 mm thick fused-silica substrates; and (iii) (111) Si wafer substrates.

X-ray photoelectron spectroscopy (XPS) measurements were performed on as-deposited coatings using a Physical Electronics PE 5800 ESCA/AES system with a monochromatic Al $\text{K}\alpha$ x-ray source. The photoelectron take-off angle was set at 45° and a charge neutralizer with a current of $10 \mu\text{A}$ was employed. The chamber base pressure was around $1 \times 10^{-9} \text{ mbar}$. High resolution scans were recorded for the main peaks of the detected elements: C 1s, O 1s, Zr 3d and Ta 4f. No Ar can be detected by XPS in this case, because the peak associated with this element is superimposed with a Ta peak. Presence of carbon is due to normal surface contamination of samples exposed to the atmosphere. In the case of insulating samples the carbon is not removed but instead used for calibration of the binding energy scale, by employing the position of the adventitious carbon peak. CasaXPS software (version 2.3.19) was used to fit the high-resolution spectra [71].

Table 3 shows the atomic concentrations obtained from XPS for the thin films as deposited. The atomic percentage cation ratio, $\eta = \text{Zr}/(\text{Zr} + \text{Ta})$, resulted in 0.23 ± 0.01 for sample II and 0.54 ± 0.03 for sample III.

Figure 1 presents typical fits of the main XPS peaks. The C 1s peak was composed mainly of adventitious carbon, and up to 5% of carbon was found to be linked to an alcohol and/or ester functionality. This further confirmed that the presence of carbon was due to surface contamination. The O 1s peak had two components: lattice O, determined as such based on the tabulated position of O 1s for zirconia and tantalum (around 530 eV), and non-lattice oxygen. In all cases, 1 to 3% of the oxygen was associated with hydroxide species based on the analysis of the C 1s peak, so the non-lattice peak could be attributed mainly to oxygen in defective sites. In the case of the Zr 3d peak, the doublet energy separation was found to be $(2.39 \pm 0.01) \text{ eV}$ and the Zr 3d_{5/2} peak position was $(182.1 \pm 0.1) \text{ eV}$ for both films. These values are in agreement with the tabulated values for zirconia [76]. For the Ta 4f doublet in both films,

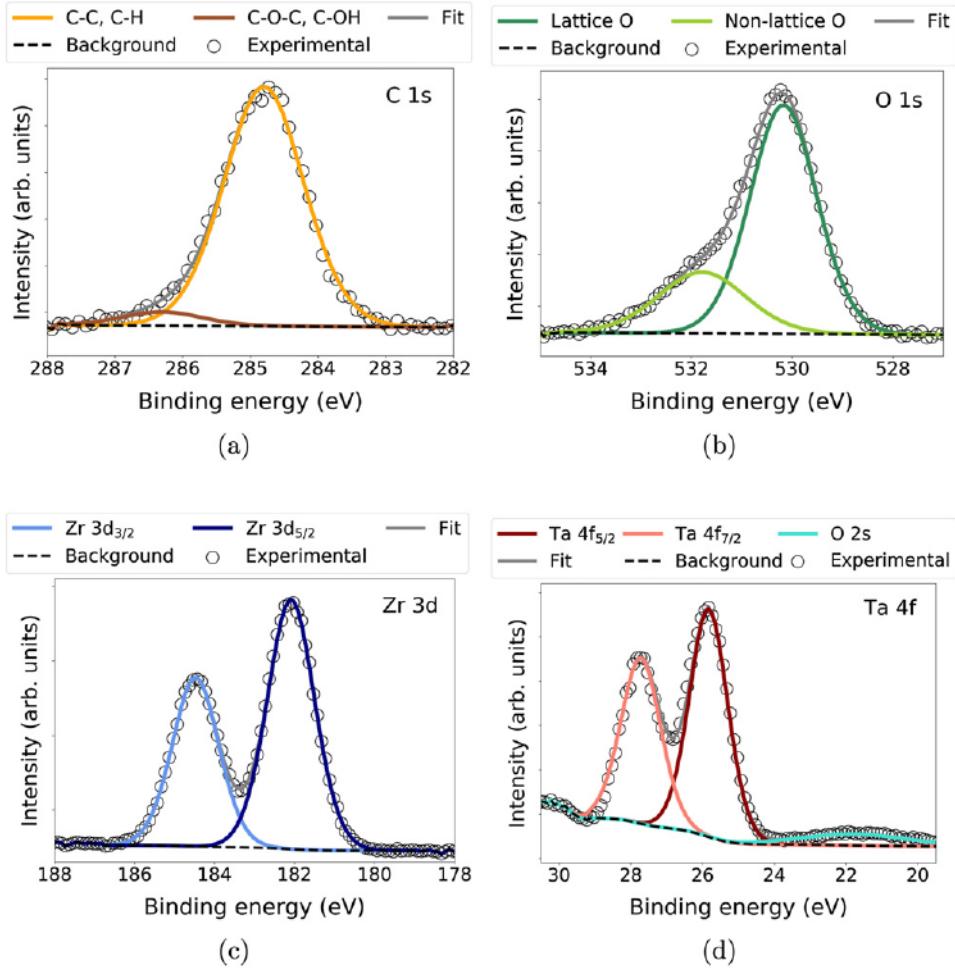


Figure 1. Typical fit of XPS peaks for an as-deposited Ta_2O_5 – ZrO_2 thin film grown via RBTD with a cation ratio of 0.54: (a) C 1s, (b) O 1s, (c) Zr 3d and (d) Ta 4f.

the energy separation was determined as (1.89 ± 0.01) eV and the $\text{Ta} 4\text{f}_{7/2}$ peak position as (26.0 ± 0.2) eV, which is in agreement with reported values for tantalum [76]. Therefore, for the evaluated zirconia-doped tantalum films as deposited, there was no noticeable modification of the chemical environment of either Ta or Zr, compared to their corresponding oxides.

2.4. HiPIMS/RF-MS

Layers of approximately 700 nm in thickness were synthesized in a Kurt J Lesker CMS-18 system, equipped with two 76 mm diameter metallic targets of Zr and Ta of 99.95% purity located ~ 25 cm beneath the substrate holder. Prior to deposition, the base pressure in the system was less than 10^{-7} mbar. Depositions were carried out on different samples: (i) the 75 mm diameter fused-silica disks were placed face down, one at a time, on a rotating substrate holder featuring a wedged hole in the middle, such that they were supported by a 1 mm ledge; (ii) silica and (iii) silicon substrates were placed next to the disks as witness samples. During deposition, the pressure was 9 mbar with a 20% O_2 :Ar gas ratio.

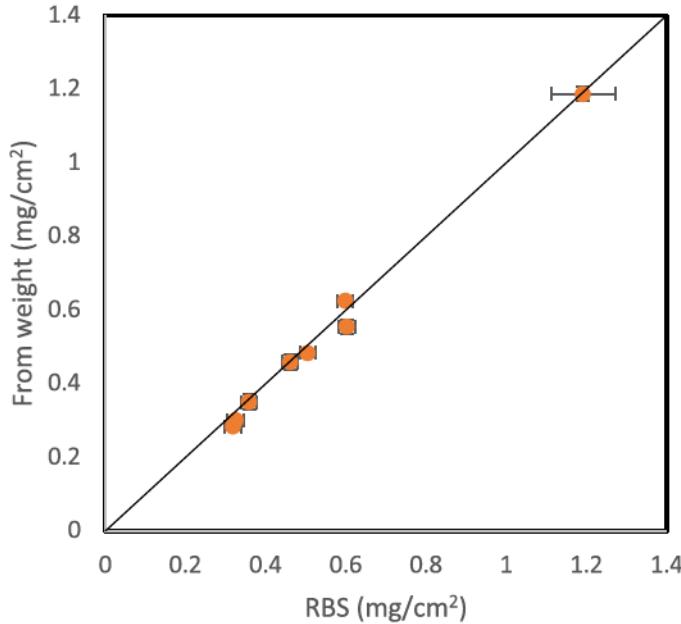


Figure 2. Areal mass density of Ta_2O_5 – ZrO_2 thin films grown via MS: values measured by sample weighing are presented as a function of those measured by RBS, for comparison. The black line shows the 1:1 relationship.

For six samples, Ta was sputtered by HiPIMS [34, 35] using a Hüttinger power supply, while the Zr was sputtered using an RF power supply (13.56 MHz) [22]. Two additional samples had their Ta deposited by RF-MS while the Zr was deposited by DC-MS, for comparison. The nominal duty cycle of the HiPIMS process was set at a voltage of 700 V, with a pulse duration of $80 \mu\text{s}$ at a frequency of 80 Hz. The nominal RF voltage applied to the Zr target varied between 200 V and 400 V, in order to achieve different doping levels. Both magnetrons were pre-sputtered for 12 min prior to deposition, in order to reach steady-state conditions before opening the main shutter. Depositions typically lasted 2 to 6 h.

Rutherford back scattering (RBS) measurements [68] were performed on the Si witness samples to measure the composition and areal atomic density of the films. Two ion beam configurations were implemented: a 2 MeV He ion beam to obtain the classical Rutherford cross section of O, and a 3.9 MeV He ion beam in order to separate the different peaks of Ta, Zr and Ar. The latter allowed for a higher precision of the concentration of these elements but with less precision on the O content, given the non-Rutherford cross-section and charging effects. The incidence angle was set at 7° and the scattering angle at 170° to maximize mass resolution. SIMNRA simulations [74] were compared to the experimental spectra in order to extract the composition and the areal atomic density of the deposited layers. Uncertainties were estimated by varying the concentrations until the simulation no longer fit the data. The resulting cation ratios $\eta = \text{Zr}/(\text{Ta} + \text{Zr})$ are reported in table 1, while the O content is discussed below. The coatings contain $\sim 1.5\%$ Ar, as expected for samples deposited in Ar ambient. Elastic recoil detection analysis in other coatings usually also reveals the presence of 1.5 to 3.0% of hydrogen. We expect this to be the case in these films as well. No concentration gradients were identified within the sensitivity of our measurements.

Table 4. Measured refractive index n at 1064, 1550 and 2000 nm of co-sputtered Ta_2O_5 – ZrO_2 thin films, and characterization method used to determine it (SP and SE indicate spectroscopic photometry and spectroscopic ellipsometry, respectively). The site of production and characterization is the same, and is indicated by the group acronym in each sample name. RF-MS samples had Ta deposited by RF-MS and Zr deposited by DC-MS, HiPIMS samples had Ta deposited by HiPIMS and Zr deposited by RF-MS. Index values are plotted and discussed in section 3.5.

Sample	Process	η (at.%)	n_{1064}	n_{1550}	n_{2000}	Method
MLD2014	IBS	48.5 ± 0.4	2.07 ± 0.01	2.06 ± 0.01		SP
MLD2018	IBS	50.2 ± 0.6	2.08 ± 0.01	2.07 ± 0.01		SP
LMA	IBS	19 ± 1	2.08 ± 0.01	2.07 ± 0.01	$2.06 \pm 0.02^{\text{a}}$	SE
UoS	ECR-IBD	34.0 ± 0.5	1.96 ± 0.01	1.95 ± 0.01	1.94 ± 0.01	SP
UMP 551	RF-MS	41 ± 3	2.01 ± 0.01	1.99 ± 0.01	1.99 ± 0.01	SE
UMP 554	RF-MS	43 ± 3	1.99 ± 0.01	1.97 ± 0.01	1.96 ± 0.01	SE
UMP 658	HiPIMS	24 ± 3	2.13 ± 0.01	2.11 ± 0.01	2.11 ± 0.01	SE
UMP 659	HiPIMS	24 ± 2	2.11 ± 0.01	2.10 ± 0.01	2.09 ± 0.01	SE
UMP 542	HiPIMS	7 ± 2	2.12 ± 0.01	2.11 ± 0.01	2.10 ± 0.01	SE
UMP 664	HiPIMS	7 ± 2	2.11 ± 0.01	2.10 ± 0.01	2.09 ± 0.01	SE
UMP 678	HiPIMS	47 ± 2	2.11 ± 0.01	2.10 ± 0.01	2.09 ± 0.01	SE
UMP 680	HiPIMS	47 ± 2	2.11 ± 0.01	2.10 ± 0.01	2.09 ± 0.01	SE
CSU II	RBDT	23 ± 1	2.07 ± 0.01	2.06 ± 0.01	$2.05 \pm 0.01^{\text{a}}$	SE
CSU III	RBDT	54 ± 3	2.10 ± 0.01	2.09 ± 0.01	$2.08 \pm 0.01^{\text{a}}$	SE

^aModel extrapolation.

The mass density of the films was calculated by the following two methods. The first one relies on the areal atomic density (at./cm^2) and composition found by RBS. Knowing the atomic mass, it was possible to compute the areal mass density (g cm^{-2}). The second method consisted in measuring the substrate mass before and after deposition using a micro-balance, and dividing it by the coated area. The latter was measured from a photograph, in order to account for the area of the uncoated region near the edges of the sample. For both methods, the areal density was divided by the thickness found by ellipsometry (see section 3.4). We note that for the density calculation based on RBS, due to the presence of thickness gradients, both measurements needed to be carried out at the same location to avoid discrepancies. Likewise, in the case of the weighing method, it was important to consider an average thickness deduced from the thickness mapping obtained by ellipsometry. Figure 2 compares the areal mass density deduced from RBS measurements and by weighing the samples, that is, without considering the thickness. We can observe that both data sets closely follow a 1:1 relationship. Given the independence of both methods, we can assume that the measured values are accurate. Densities reported in table 1 are average values between the two methods.

3. Optical properties

The optical properties of the coating samples were characterized using different techniques. Table 4 presents the measured values for the refractive index n at the wavelengths of interest for current and future GW interferometers: 1064, 1550 and 2000 nm [7, 15, 48, 50]. In some cases, the refractive index at 2000 nm was estimated by extrapolating the sample best-fit dispersion models through the infrared region, where optical absorption is expected to be negligible.

Optical absorption at 1064 nm was measured on RBDT single-layer samples and IBS multi-layer HR stacks. Light scattering at 1064 nm was measured on IBS multi-layer HR stacks.

At 1550 nm, where only ellipsometric and photometric measurements were available, extinction values were smaller than the sensitivity of the instruments ($k < 10^{-3}$). Measurements and their analyses are described below. A comparative summary is presented in section 3.5, where correlations between oxygen concentration, density and refractive index are discussed.

3.1. IBS

3.1.1. MLD. After deposition, MLD used a Cary spectrophotometer to measure the coating transmittance over the 400–1800 nm spectral range. The coating thickness and index of refraction were determined by a fit to this data.

At Embry-Riddle Aeronautical University, Prescott (Gretarsson Lab), we measured the optical thickness dependence on temperature for sample MLD2014 by monitoring the wavelength dependence of its transmission spectrum extrema as a function of temperature in the range 89°C–411°C. Details of the apparatus and method are given elsewhere [73]. Typical fits to the extrema locations are shown in figure 3(a). The measured coating optical thickness as a function of temperature, shown in figure 3(b), is linear. Note that, due to uncertainty in the physical thickness of the coating layer, optical thickness is not used here to estimate the absolute value of the index, only the change with temperature. The fractional optical thickness change as a function of temperature was found to be

$$\frac{1}{L} \frac{dL}{dT} = \alpha + \beta/n = (18.6 \pm 1) \times 10^{-6} \text{ K}^{-1}, \quad (1)$$

where α is the thermal expansion coefficient and β is the thermo-optic coefficient.

3.1.2. LMA. Refractive index and thickness of as-deposited coating samples were determined by spectroscopic ellipsometry, via coupled analysis of data sets acquired with two different J. A. Woollam Co. commercial instruments covering complementary spectral regions: a VASE for the 190–1100 nm range and a rotating-compensator M-2000 for the 245–1700 nm range. The wide range swept with both instruments allowed us to extend the analysis from ultraviolet to near infrared (0.7–6.5 eV). Consistent results were obtained in the overlapping range, within the uncertainty of the models used to fit the data. Analysis was carried out independently for VASE and M-2000 data, and eventually the information was combined in order to obtain the final results.

The optical constants were obtained by measuring the amplitude ratio Ψ and phase difference Δ of the p- and s-polarized light [64] reflected by the samples at different incidence angles $\theta = 55^\circ, 60^\circ, 65^\circ$, close to the coating Brewster angle. The optical response of the substrates (silicon wafers with a thin native oxide layer and unpolished rear surface in order to avoid backside spurious reflections) was characterized with prior dedicated measurements and analyses [18]. As in previous works on similar systems [18, 47], we analyzed the data through a three-layer model consisting of substrate, thin film and surface layer. The latter, accounting for surface roughness, was defined via a Bruggeman effective medium approximation. The resulting coating roughness was less than 1 nm, a value comparable with those measured with an atomic-force microscope on similar samples [47]. Being less than 1% of the coating thickness, the surface layer did not substantially affect the data fit, especially in the near-infrared region.

The dielectric functions were reconstructed by using Kramers–Kronig-consistent Tauc–Lorentz [6] and Cody–Lorentz [80] models. Both models fit the data with comparable mean squared error, and yielded consistent results. We then obtained the complex refractive index of the as-deposited thin films from the best-fit dielectric functions. The Cody–Lorentz model provided a slightly better fit in the region of the absorption edge and included also the

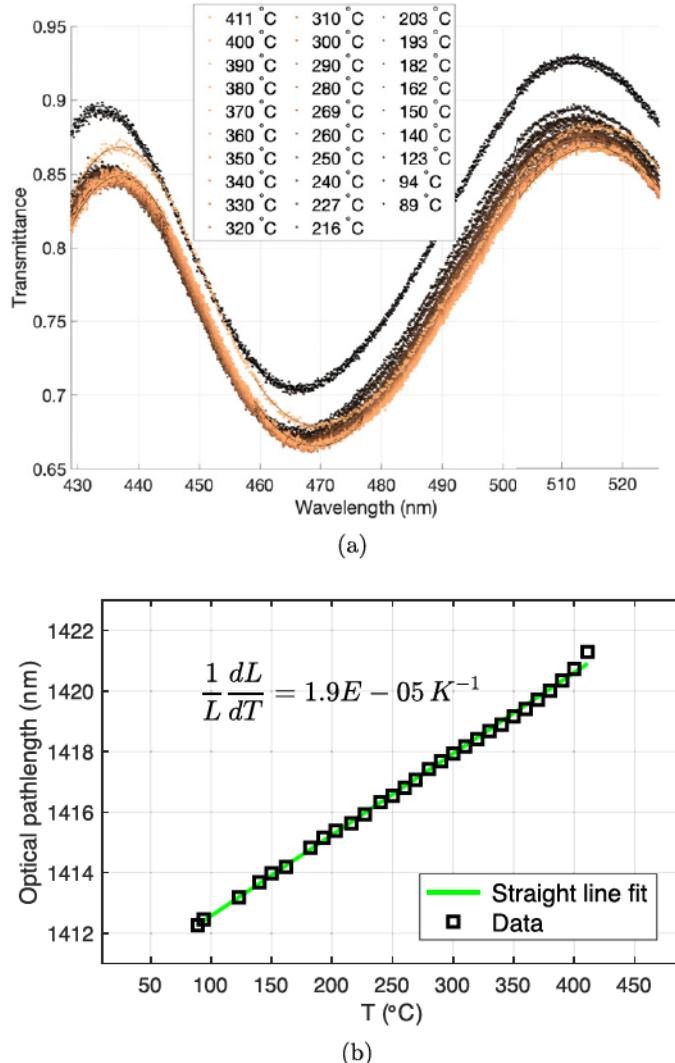


Figure 3. Characterization of optical thickness dependence on temperature of IBS $Ta_2O_5-ZrO_2$ thin film MLD2014. (a) Transmission spectrum extrema with parabolic fits to determine the extrema locations as a function of temperature. Higher temperatures shift the extrema to longer wavelengths, indicating an increasing optical thickness with temperature. (b) Optical thickness of the coating as a function of temperature.

Urbach tail absorption, which may be correlated to coating mechanical loss [19]. However, data quality for energies higher than the energy gap ($E > 4$ eV) was degraded and resulted in a large uncertainty for the Urbach energy $E_U = 130 \pm 30$ eV.

All measured spectra were fit with the same accuracy. Figure 4 shows exemplary spectra acquired at $\theta = 60^\circ$ and their best fits with a Cody–Lorentz model. Figure 5(a) shows the reconstructed coating optical constants, and figure 5(b) is a Cody plot highlighting the coating energy gap, $E_g = 4.1 \pm 0.1$ eV, which is consistent with that of pure tantalum coatings produced under identical conditions [18].

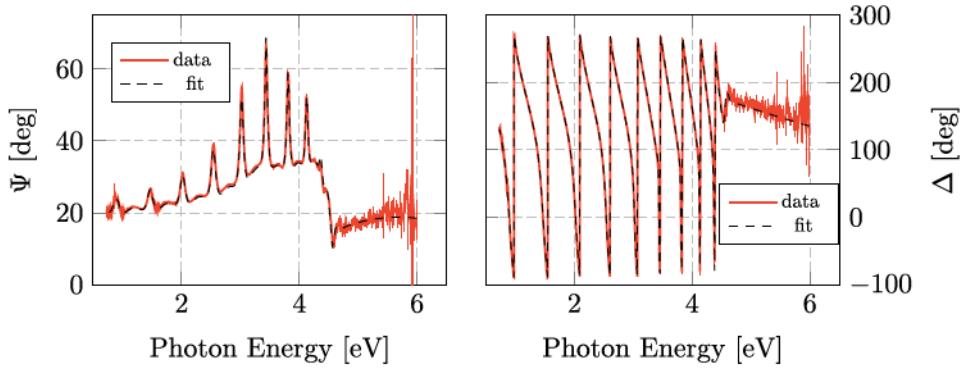


Figure 4. Measured ellipsometric spectra of an as-deposited IBS $\text{Ta}_2\text{O}_5\text{-ZrO}_2$ thin film with cation ratio $\eta = 0.19$ (LMA), acquired at an incidence angle $\theta = 60^\circ$. The best-fit Cody–Lorentz model is shown (dashed line).

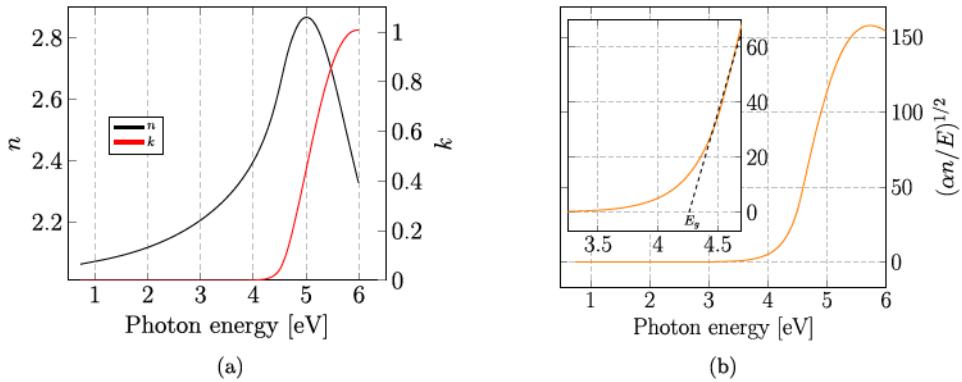


Figure 5. Measured optical properties of as-deposited IBS $\text{Ta}_2\text{O}_5\text{-ZrO}_2$ thin films with cation ratio $\eta = 0.19$ (LMA): (a) refractive index n and extinction coefficient k as function of photon energy E , the region of interest for present GW detectors is around $E = 1.2$ eV [7, 15] (corresponding to a wavelength of 1064 nm) and for future GW detectors is around $E = 0.8$ and $E = 0.6$ eV [48, 50] (corresponding to 1550 and 2000 nm, respectively); (b) Cody plot highlighting the energy gap E_g .

After annealing at 500°C in air for 10 h, single-layer coating samples were characterized in their transparency region (400–1400 nm) with a Perkin Elmer Lambda 1050 spectrophotometer, at normal incidence. The annealing reduced the refractive index in the near-infrared region by about 3%. Samples of HR ETM stacks (see section 2), also annealed at 500°C in air for 10 h, were characterized for optical absorption and light scattering with a custom-developed setup based on the photo-thermal deflection principle [21] and a commercial CASI scatterometer, respectively. We measured an optical absorption of 0.5 ppm, and 45 ppm of scattered light.

3.2. ECR-IBD

Optical-transmission spectra of as-deposited films, shown in figure 6, were measured in the 250–2100 nm range using a Varian 5000 spectrophotometer and fit to transmission simulations using SCOUT software by WTheiss. The behavior of the measured extinction coefficient,

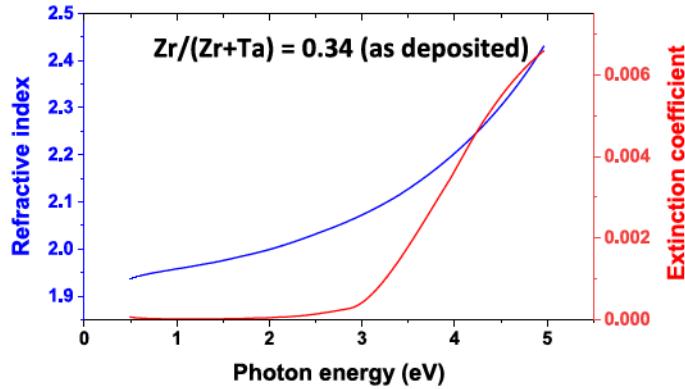


Figure 6. Refractive index n and extinction coefficient k as function of photon energy for as-deposited ECR-IBD Ta_2O_5 – ZrO_2 thin films with cation ratio $\eta = 0.34$.

particularly the short-wavelength cut-off, is likely due to the deposition parameters associated with the ECR-IBD process, e.g. the higher ion energy and slower deposition rate. At 1064 nm and 1550 nm, the refractive index n was found to be 1.96 ± 0.01 and 1.95 ± 0.01 , respectively, and the extinction coefficient $k < 10^{-3}$.

3.3. RBTD

Refractive index and extinction coefficient for the films were determined by spectroscopic ellipsometry using a Horiba UVISEL ellipsometer. Measurements were taken at an angle of 60° , for a spectral range of 0.59–6.5 eV. The films were evaluated as deposited and after annealing at the maximum temperature for which no crystallization was detected and mechanical loss minimized. A reduction of the refractive index was found after annealing, which could be due to oxygen incorporation during the process or due to formation of Ar bubbles [29]. Films previously grown by our group were found to contain 3 to 5% of Ar. When Ar bubbles are formed and coalesce upon annealing, the films usually feature a reduction in the refractive index that is accompanied by an increase in thickness. In our case, however, the thickness did not increase after annealing within the resolution of the instrument, indicating no significant presence of Ar bubbles in the coatings.

The dispersion of refractive index n and extinction coefficient k are presented in figure 7. The refractive index for both annealed films is presented in table 4. The film with the largest cation ratio featured the highest refractive index, as expected, given that the refractive index of zirconia is higher than that of tantalum.

Optical absorption at $\lambda = 1064$ nm was determined by photothermal common-path interferometry [16]. We define the absorption normalized to a quarter wave thickness as

$$a_{\text{QWT}} = \frac{a}{d} \frac{\lambda}{4n}, \quad (2)$$

with a being the absorption as measured in ppm, n the refractive index, d the coating thickness and $\lambda = 1064$ nm. a_{QWT} was 4.4 ± 0.8 ppm for the film with cation ratio $\eta = 0.54$ and 1.85 ± 0.05 ppm for the film with cation ratio $\eta = 0.23$. These results are summarized in table 5.

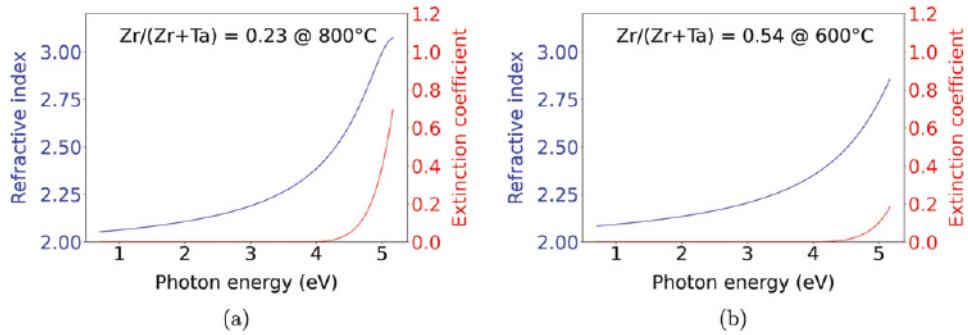


Figure 7. Refractive index n and extinction coefficient k as function of photon energy for RBTD Ta_2O_5 – ZrO_2 thin films: (a) cation ratio $\eta = 0.23$, annealed at $800\text{ }^\circ\text{C}$ for 10 h; (b) cation ratio $\eta = 0.54$, annealed at $600\text{ }^\circ\text{C}$ for 10 h.

Table 5. Absorption loss normalized to a quarter wave thickness α_{QWT} at 1064 nm for RBTD Ta_2O_5 – ZrO_2 thin films annealed at the maximum temperature T_a before crystallization (see equation (2) and section 5 for more details).

Sample	T_a ($^\circ\text{C}$)	Cation ratio η	a_{QWT} @ 1064 nm (ppm)
CSU II	800	0.23	1.85 ± 0.05
CSU III	600	0.54	4.4 ± 0.8

3.4. HiPIMS/RF-MS

Similarly to what is described in sections 3.1.2 and 3.3, spectroscopic ellipsometry was carried out between wavelengths of 210 nm and 2500 nm with a J. A. Woollam Co., Inc. RC2-XI ellipsometer. Data analysis to obtain the complex refractive index and thickness mappings was carried out using the J. A. Woollam CompleteEASE software, with the optical properties being represented by a general oscillator model consisting of a Tauc–Lorentz and Gaussian oscillators for the UV absorption. The model parameters were first fit for each sample from a series of measurements at four different incidence angles, 45° , 55° , 65° and 75° , carried out at the center of the sample. The average thickness of the samples was then established through a mapping of 48 points, measured at 70° where only the thickness and surface roughness (Bruggeman effective medium approximation) were fit while keeping the optical model constant. The measured refractive index at 1064 and 1550 nm of all HiPIMS/RF-MS samples is presented in table 4.

3.5. Comparative summary

Figure 8(a) shows the refractive index values at 1064 nm and 1550 nm reported in table 4 as a function of the cation ratio $\eta = \text{Zr}/(\text{Zr} + \text{Ta})$, for all the samples. By way of example, a comparison of dispersion laws for samples where $\eta \sim 20\%$ is presented in figure 9.

The solid line in figure 8(a) corresponds to the rule of mixtures (addition according to molar fractions, also known as Arago–Biot rule [65] for the refractive index), assuming bulk values for pure oxides. The low dependence on η is expected, since tantalum and zirconia have a similar refractive index at these wavelengths. Samples deposited by HiPIMS, IBS and RBTD feature values close to the rule of mixture. For samples where tantalum is deposited by RF-MS or when the coating is deposited by ECR-IBD, the refractive index is significantly lower.

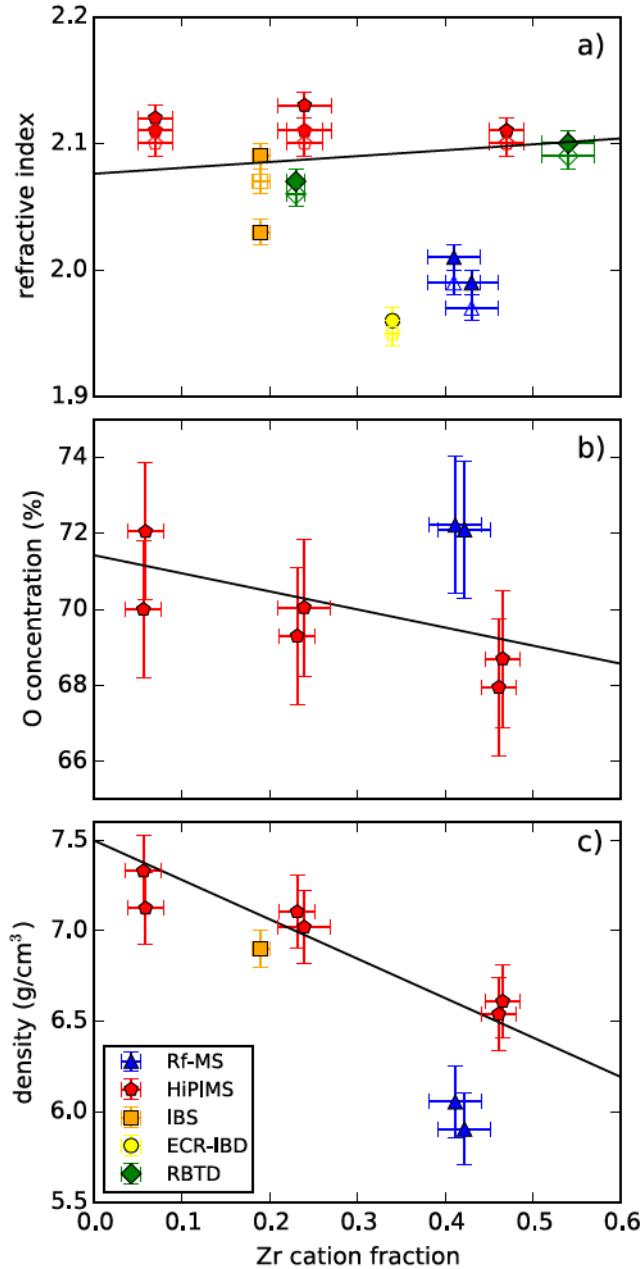


Figure 8. Correlation between oxygen concentration, density and refractive index of co-sputtered Ta_2O_5 – ZrO_2 thin films, as a function of cation ratio $\eta = \text{Zr}/(\text{Ta} + \text{Zr})$: (a) refractive index at 1064 nm (filled markers) and 1550 nm (empty markers), the black solid line is the value expected from the rule of mixtures (or Arago–Biot rule [65]) at 1064 nm when considering the refractive index found in reference [69] for bulk zirconia and reference [85] for that of bulk tantalum; (b) oxygen concentration; (c) density. The solid lines in (b) and (c) are the expected oxygen content and density, respectively, according to the rule of mixture and assuming stoichiometric tantalum and zirconia for the oxygen content, and density values for Ta_2O_5 and ZrO_2 taken from the literature [4, 55].

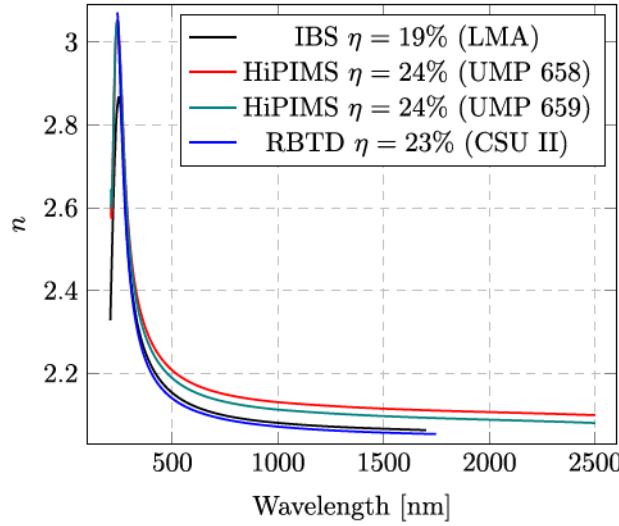


Figure 9. Comparison of refractive index n for $\text{Ta}_2\text{O}_5-\text{ZrO}_2$ thin films with cation ratio $\eta \sim 20\%$, as function of photon wavelength: as-deposited IBS sample (LMA), HiPIMS samples annealed at $700\text{ }^\circ\text{C}$ before crystallization, RBTD sample annealed at $800\text{ }^\circ\text{C}$.

In figure 8(b), we find that the oxygen content in the layers where tantalum is deposited by HiPIMS follows the rule of mixture (according to molar fractions) between stoichiometric tantalum and zirconia, while there is an excess of oxygen in the films where tantalum was deposited by RF-MS. Accordingly, as shown by figure 8(c), the density of films where tantalum is deposited by HiPIMS or IBS follow the rule of mixture considering bulk densities of tantalum and zirconia [4, 55], while those where tantalum was deposited by RF-MS show a significantly lower density. Hence, the excess of oxygen leads to films with lower density and correspondingly lower refractive index. The explanation partly lies in the fact that atoms, when deposited by HiPIMS, have a higher kinetic energy than in RF-MS, which produces more compact films with less chances to have, for example, two oxygen atoms on the same bond. It is worth noting that we did not attempt to grow stoichiometric RF-MS layers, for example by changing the $\text{O}_2:\text{Ar}$ ratio during deposition, as we wanted to compare HiPIMS and RF-MS under the same conditions.

Our results show the correlation between the oxygen concentration, the density, and the refractive index. Techniques such as HiPIMS and IBS feature atomic fluxes with higher kinetic energy, leading to higher dissociation rates of oxygen. This means that less oxygen may be used in the Ar/O_2 mixture to obtain stoichiometry. Also, the higher energy of species impinging on the growing film surface leads to coatings of higher density. Conversely, RF-MS films are generally less dense and have a lower refractive index. However, pores are frequently filled with water vapor that adds oxygen and hydrogen, leading to apparent oxygen over-stoichiometry. The former may therefore have more desirable properties for optical components. We will see in section 4 how that affects the loss angle.

4. Mechanical properties

The mechanical loss φ_c , Young modulus Y_c and Poisson ratio ν_c of the coating samples were measured using different setups, all based on the ring-down technique [83]: the mechanical resonances (modes) of a suspended coated resonator are excited using an actuator, and the

amplitude of their ring-down oscillations is monitored using an optical transducer. For a resonating mode of frequency f and ring-down time τ , the measured mechanical loss angle is $\varphi = (\pi f \tau)^{-1}$. The coating loss angle φ_c can be calculated as

$$\varphi_c = \frac{\varphi + (D - 1)\varphi_s}{D}, \quad (3)$$

where φ_s is the measured loss angle of the substrate, φ is the measured loss angle of the coated sample and D is a so-called *dilution factor*, defined as the ratio of the coating elastic energy to the total elastic energy of the coated sample.

Coated samples were suspended under vacuum, in order to prevent systematic damping from ambient pressure. We used different suspension systems, each one specifically adapted to the different geometries of the substrates used as resonators. Because of the different suspension systems and samples, we also used different methods to estimate D , Y_c and ν_c . Specific measurements and their analyses are described in the following.

Results are summarized in table 6, where an average coating loss $\bar{\varphi}$, taken from every set of resonant modes of each sample, is used to ease comparisons. Since the lowest coating thermal noise occurs when the coating Young modulus matches that of the substrate [36], values of the parameter

$$\gamma = \bar{\varphi}_c \left(\frac{Y_c}{Y_s} + \frac{Y_s}{Y_c} \right) \quad (4)$$

are also presented, where $Y_s = 73.2$ GPa is the Young modulus of the fused silica substrate of Advanced LIGO and Advanced Virgo mirrors [7, 15].

A comparative summary is presented in figures 18 and 19 of section 4.5, where the effects of different cation ratios $\eta = \text{Zr}/(\text{Zr} + \text{Ta})$ and annealing temperatures T_a on the coating mechanical loss are discussed.

4.1. IBS

4.1.1. MLD. The coating samples used to measure coating mechanical loss, Young modulus and Poisson ratio were grown on fused-silica disks and cantilevers and on silicon cantilevers (see section 2.1.1). Samples grown on fused-silica substrates were used for measurements at ambient temperature, those on silicon cantilevers to perform cryogenic measurements.

Disks — the mechanical loss measurements were performed at Hobart and William Smith Colleges (HWS). The coated, fused silica substrates were welded to a monolithic, fused silica fiber suspension. The mechanical loss was measured for the samples' eight lowest vibrational modes, which span the frequency range from 2.5 to 18 kHz. Each mode was excited to resonance using an electrostatic comb exciter. The dissipation was measured using an ellipsometry detector to monitor the strain-induced birefringence during ring-down (see [37] for details). The loss for each mode was measured several times with both weighted and unweighted averages retained.

After each set of measurements, the samples have been annealed and remeasured. The annealing temperatures were 300, 600, 650, 700, 750 and 800 °C for MLD2014 and 400, 600 and 700 °C for MLD2018. For these experiments, the fused silica substrate is typically annealed to high temperature before being coated. The MLD2018 disk was properly annealed at 950 °C, whereas the MLD2014 disk was not annealed prior to coating. For the latter sample, the substrate loss was determined from the data using the frequency dependence of the loss as determined by finite-element modeling. The model, created using COMSOL, was used to

20

Table 6. Measured mechanical properties of co-sputtered Ta₂O₅–ZrO₂ thin films: frequency-averaged mechanical loss $\bar{\varphi}$ of samples annealed at the highest peak temperature T_a before the onset of crystallization, Young modulus Y_c , Poisson ratio ν_c , and coefficient $\gamma = \bar{\varphi}_c(Y_c/Y_s + Y_s/Y_c)$, where $Y_s = 73.2$ GPa is the Young modulus of a fused silica substrate. The characterization methods used to determine such properties are also listed, together with the site where the characterizations were performed. RF-MS samples had Ta deposited by RF-MS and Zr deposited by DC-MS, HiPIMS samples had Ta deposited by HiPIMS and Zr deposited by RF-MS. RF-MS, HiPIMS and IBS (LMA) samples have been annealed for 10 h, ECR-IBD samples for 5 h. Average coating loss values are plotted in figures 18 and 19 and discussed in section 4.5.

Sample	Process	T_a (°C)	η (at.%)	$\bar{\varphi}_c$ (10^{-4} rad)	Y_c (GPa)	ν_c	γ (10^{-4} rad)	Method
MLD2014	IBS	800	48.5 ± 0.4	4.26 ± 0.11	131 ± 2		10.0	^a
MLD2018	IBS	700	50.2 ± 0.6	1.82 ± 0.04	130 ± 2		4.3	^a
LMA	IBS	700	19 ± 1	4.41 ± 0.68	120 ± 3	0.32 ± 0.01	9.9	GeNS (LMA)
UoS	ECR IBD	750	34.0 ± 0.5	3.64 ± 0.45	176 ± 8		10.3	Clamped cantilevers (UoS)
UMP 551	RF-MS	800	41 ± 3	2.36 ± 0.85	125 ± 2	0.21 ± 0.05	5.4	GeNS (Caltech)
UMP 554	RF-MS	800	43 ± 3	2.52 ± 0.82	114 ± 2	0.36 ± 0.03	5.5	GeNS (Caltech)
UMP 658	HiPIMS	700	24 ± 3	5.59 ± 0.90	110.6 ± 1.2	0.31 ± 0.02	12.1	GeNS (Caltech)
UMP 659	HiPIMS	700	24 ± 2	5.04 ± 1.53	108 ± 5	0.35 ± 0.05	10.9	GeNS (Caltech)
UMP 542	HiPIMS	600	7 ± 2	5.69 ± 0.37	100.6 ± 1.1	0.34 ± 0.02	12.0	GeNS (Caltech)
UMP 664	HiPIMS	650	7 ± 2	4.45 ± 0.38	116.5 ± 1.3	0.33 ± 0.02	9.9	GeNS (Caltech)
UMP 678	HiPIMS	800	47 ± 2	4.38 ± 0.63	111 ± 4	0.28 ± 0.09	9.5	GeNS (Caltech)
UMP 680	HiPIMS	800	47 ± 2	4.87 ± 0.92	110 ± 4	0.28 ± 0.09	10.6	GeNS (Caltech)
CSU II	RBTD	800	23 ± 1	3.22 ± 0.53	145 ± 6	0.30 ± 0.09	8.0	GeNS (Caltech)
CSU III	RBTD	600	54 ± 3	2.67 ± 0.15	143 ± 5	0.37 ± 0.05	6.6	GeNS (Caltech)

^aDisks suspended by welded glass fibers (HWS)/clamped cantilever blades, at ambient and cryogenic temperature (IGR)/nanoindentation (Caltech).

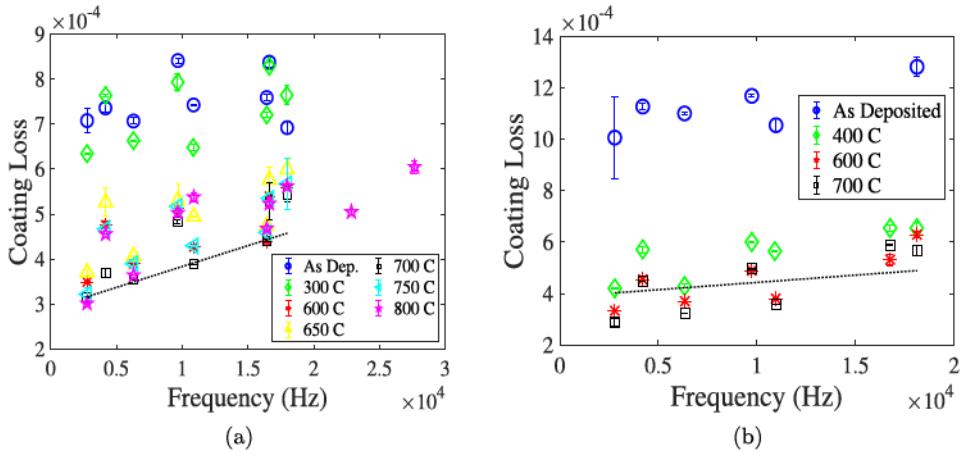


Figure 10. Coating loss as a function of frequency for IBS $\text{Ta}_2\text{O}_5\text{-ZrO}_2$ thin films, as measured on suspended disks: (a) sample MLD2014, (b) sample MLD2018. The loss is steadily reduced with increased annealing temperature T_a , until reaching a minimum near 700 °C. A weighted linear least-squares fit is shown (black line) for the data sets of the 700 °C annealing. In the MLD2014 data, panel (a), the fit line follows the low loss data due to their smaller uncertainties.

calculate the dilution factor for the total coating loss and for the decomposition of the coating loss into bulk and shear loss [86].

As shown in figure 10, the coating loss for both samples decreased with increasing annealing temperature, until reaching a minimum at 700 °C. A linear fit of the frequency-dependent coating loss after annealing at 700 °C gave:

$$\varphi_c(f) = (9.29 \pm 0.09) \times 10^{-9} \text{ Hz}^{-1} f + (2.907 \pm 0.007) \times 10^{-4} \text{ for MLD2014,}$$

$$\varphi_c(f) = (5.7 \pm 0.2) \times 10^{-9} \text{ Hz}^{-1} f + (3.87 \pm 0.02) \times 10^{-4} \text{ for MLD2018.}$$

The variation in the ratio of bulk and shear energy with mode shape, which in turn depends on frequency, introduces a scatter of the data about any function of a single coating loss angle. Instead, a decomposition of the coating loss into shear and bulk components allows a more precise fit to the data, as shown in figure 11. A fit of each data set to bulk and shear coating loss resulted in the dependence of these values with annealing temperature T_a shown in figure 12: the shear loss steadily decreased for increasing T_a , whereas the bulk loss showed no clear dependence. The substrate barrel (edge) was polished. A fit to the data showed no significant contribution from edge dependent losses [24].

The coating Young modulus was measured after annealing at two different temperatures, using a nano-indentation facility at the California Institute of Technology (Caltech). The results, $Y_c = 130 \pm 2$ GPa for MLD2018 annealed at 800 °C and $Y_c = 131 \pm 2$ GPa for MLD2014 annealed at 600 °C, assume a coating Poisson ratio of 0.298 determined by molecular dynamics models [66].

Cantilevers — the mechanical loss of the MLD2018 coating was measured also on a fused-silica cantilever at ambient temperature, and on silicon cantilevers at cryogenic temperatures. Measurements were performed at the Institute for Gravitational Research of the University of Glasgow (IGR).

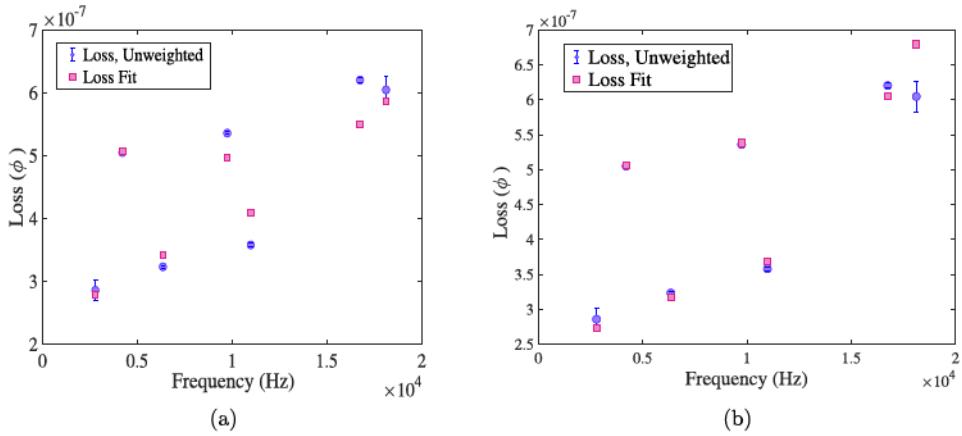


Figure 11. Comparison of analysis methods, as applied to coating loss data of IBS $\text{Ta}_2\text{O}_5\text{-ZrO}_2$ thin films (MLD) measured on suspended disks: (a) fit with single loss angle, (b) fit with bulk and shear loss angles. The need for separate analysis of bulk and shear coating loss angle is pointed out by the improved fit of panel (b).

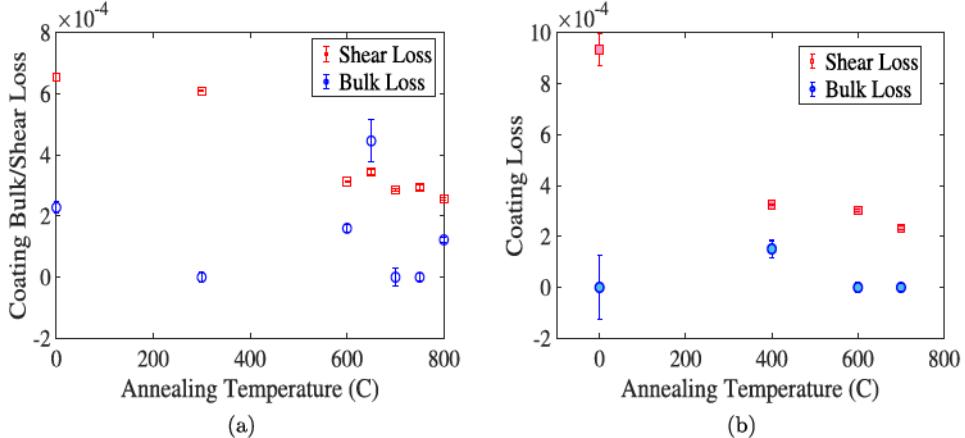


Figure 12. Bulk and shear coating loss as function of annealing temperature for IBS $\text{Ta}_2\text{O}_5\text{-ZrO}_2$ thin films, as measured on suspended disks: (a) sample MLD2014, (b) sample MLD2018. The shear loss is steadily reduced with increased annealing temperature, while the bulk loss shows no clear dependence.

The cantilevers were mounted in a clamp, under vacuum, and their bending modes were excited using an electrostatic actuator. The amplitude of the ring-down motion was monitored using an optical shadow sensor. The coating mechanical loss was then calculated using equation (3) and

$$D = \frac{3Y_c d_c}{Y_s d_s}, \quad (5)$$

where Y_c and d_c are the coating Young modulus and thickness, respectively, and Y_s and d_s are the substrate Young modulus and thickness, respectively.

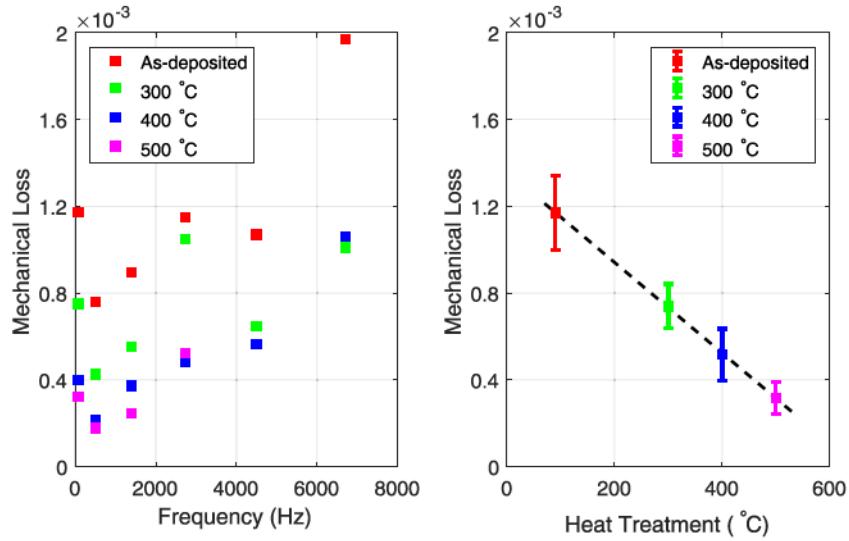


Figure 13. Coating loss of IBS Ta_2O_5 – ZrO_2 thin film MLD2018, as measured on a clamped $\sim 100 \mu\text{m}$ thick fused-silica cantilever at ambient temperature. Left: (a) the loss was measured as a function of frequency on the as-deposited sample, then re-measured after progressive annealings at 300, 400 and 500 °C; right: (b) frequency-averaged coating loss $\bar{\varphi}$ as a function of annealing temperature. The coating loss decreased after each annealing treatment.

After each set of measurements, the coated fused-silica cantilever was progressively annealed at increasing peak temperatures, up to 500 °C. Figure 13(b) shows the effect of annealing on the coating mechanical loss as a function of frequency, and that the coating mechanical loss at 1 kHz decreased by over a factor of three upon annealing, from $\bar{\varphi} = 1.1 \times 10^{-3}$ in its as-deposited state to $\bar{\varphi} = 3.2 \times 10^{-4}$ after a 500 °C annealing.

The coated silicon cantilevers were measured in a temperature-controlled, liquid-helium-cooled cryostat [77]. The coating mechanical loss was calculated using equations (3) and (5) for a large number of resonant modes, from ~ 1 to 30 kHz, as extensively detailed in [77]. Here, for sake of conciseness, we present in figure 14 only the coating loss of the third- and fourth-order bending modes, at ~ 1.4 kHz and ~ 2.7 kHz respectively.

Figure 14 shows that IBS Ta_2O_5 – ZrO_2 coating in its as-deposited state has a very small and broad loss peak of $\sim 6 \times 10^{-4}$ at ~ 30 K. This peak narrowed to 1×10^{-3} after annealing at 300 and 400 °C, and further increased to 1.2×10^{-3} after an 800 °C anneal. Below 100 K, the coating loss temperature-dependent trend of the annealed Ta_2O_5 – ZrO_2 samples is similar to that of pure IBS Ta_2O_5 films annealed at 600 °C [39]. After annealing at 800 °C, the coating loss of Ta_2O_5 – ZrO_2 does not show the large and broad loss peak observed at 80 K for IBS poly-crystalline Ta_2O_5 thin films annealed to the same temperature [39]. This is a clear indication that co-sputtering ZrO_2 with Ta_2O_5 prevented the onset of crystallization for annealing at 800 °C.

As shown in figure 13, annealing reduced the coating loss at ambient temperature. At the same time, figure 14 suggests that annealing increased the coating loss at low temperatures, which is the opposite of what has been observed on IBS poly-crystalline Ta_2O_5 – TiO_2 [41]. This is of interest for further understanding of coating loss and structure and modeling, and

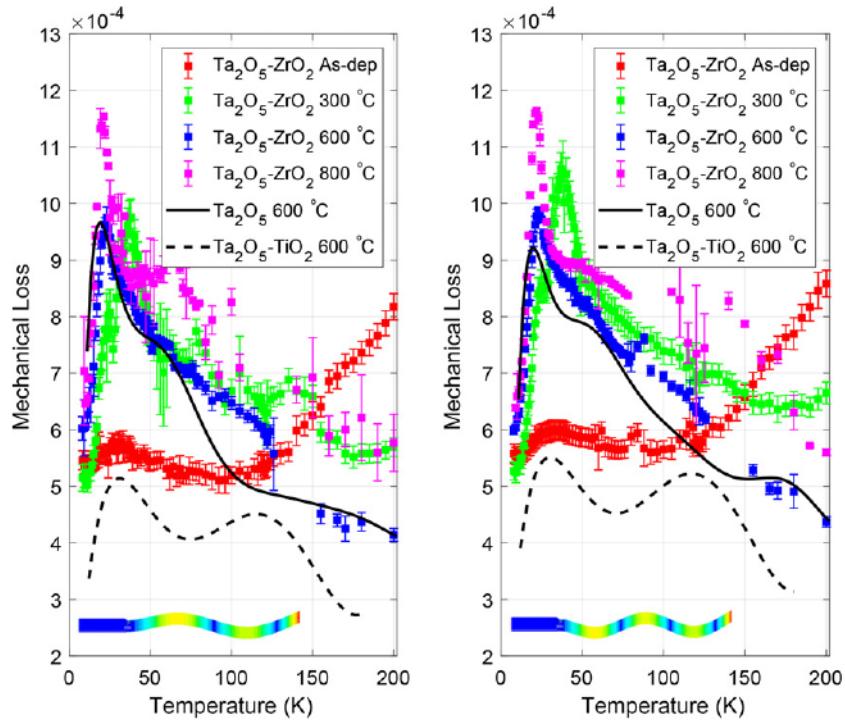


Figure 14. Cryogenic mechanical loss of IBS Ta_2O_5 - ZrO_2 coatings as a function of sample temperature, as measured on clamped $\sim 65 \mu\text{m}$ thick silicon cantilevers. Left: (a) results for the third-order bending mode at $\sim 1.4 \text{ kHz}$; right: (b) results for the fourth-order bending mode at $\sim 2.7 \text{ kHz}$. The coating loss of sample MLD2018 as deposited and after a 300°C annealing and of sample MLD2014 after 600°C and 800°C annealing are shown, and compared to data of IBS amorphous Ta_2O_5 and poly-crystalline Ta_2O_5 - TiO_2 samples annealed to 600°C [39, 41]. Finite-element representations of the measured resonant mode are visible at the bottom of each plot.

suggests that different values of the cation ratio $\eta = \text{Zr}/(\text{Zr} + \text{Ta})$ need to be investigated at low temperatures.

4.1.2. LMA. We used a gentle nodal suspension (GeNS) system [25] to measure coating mechanical loss, Young modulus and Poisson ratio via the direct measurement of the disk dilution factor [32]. As a matter of fact, the dilution factor D in equation (3) can be written as a function of the resonator mode frequency (f_0, f) and mass (m_0, m) before and after coating deposition, respectively [52]:

$$D = 1 - \left(\frac{f_0}{f} \right)^2 \frac{m_0}{m}. \quad (6)$$

The advantage of measuring the dilution factor is that the ensuing coating loss estimation is then exclusively based on measured quantities: sample loss angle, mode frequency and mass. Therefore, unlike other experimental setups based on the ring-down method, our method does not require prior knowledge of coating Young modulus and thickness. Furthermore, one

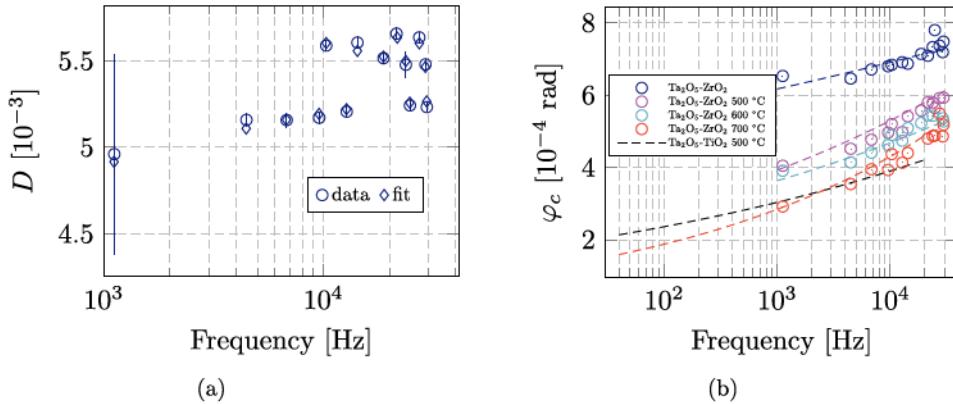


Figure 15. Measured mechanical properties of IBS $Ta_2O_5-ZrO_2$ thin films (LMA) as a function of frequency, as measured on a disk suspended with a GeNS system: (a) measured and best-fit simulated dilution factor D , for the first three families of disk mode shapes [32]; (b) coating loss angle φ_c , as measured after deposition and after annealing treatments at 500, 600 and 700 °C for 10 h, in air. Data of the annealed $Ta_2O_5-TiO_2$ layers of Advanced LIGO and Advanced Virgo is also shown, for comparison [32]. Dashed lines are least-squares fits using a power-law model $\varphi_c(f) = af^b$, best-fit parameters a and b are presented in table 7. Values at $f < 1$ kHz are model extrapolations.

can estimate coating Young modulus Y_c and Poisson ratio ν_c by fitting the results of finite-element simulations to measured values of D estimated via equation (6). Compared to nano-indentation, such method does not rely on the prior knowledge of ν_c , and does not depend on the nature of the substrate used nor on the model chosen for data analysis [32]. Figure 15(a) shows best-fit results for three mode families of different shapes and frequency [32], giving $Y_c = 120 \pm 3$ GPa and $\nu_c = 0.32 \pm 0.01$. Details about our fitting method and finite-element simulations can be found elsewhere [30, 32].

Figure 15(b) shows the coating loss angle before and after consecutive annealing treatments at 500, 600 and 700 °C for 10 h, in air. As expected, the annealing reduced the coating loss. However, the coating showed cracks after the 700 °C annealing. This issue is further discussed in section 5.

We used a power-law model $\varphi_c(f) = af^b$ to describe the observed frequency-dependent behavior of the coating internal friction [24, 53, 81], and applied least-squares linear regression in order to fit the data. Best-fit parameters are presented in table 7. It is worth noting that the b parameter increased after each annealing step, implying lower loss values in the most sensitive region (50–300 Hz) of GW detectors. Thus, by extrapolating our results to lower frequency, we obtain a loss angle of 1.8×10^{-4} at 100 Hz for the $Ta_2O_5-ZrO_2$ annealed at 700 °C. This value has to be compared to 2.4×10^{-4} for the $Ta_2O_5-TiO_2$ layers of GW detectors [32].

4.2. ECR-IBD

The coating mechanical loss at ambient temperature was characterized using fused-silica cantilevers. They were mounted in a stainless steel clamp within a vacuum chamber with residual pressure below 5×10^{-6} mbar, and excited using an electrostatic actuator [49]. The freely-decaying amplitude of the resonant motion was monitored using an optical shadow sensor. We then used equations (3) and (5) to calculate the coating loss for the cantilever bending modes [79], and the results are presented in section 4.5.

Table 7. Best-fit parameters of a power-law internal friction model $\varphi_c(f) = af^b$ for data sets of IBS $Ta_2O_5-ZrO_2$ thin films (LMA) measured on a disk suspended with a GeNS system, as function of annealing peak temperature T_a (as-deposited coating samples are denoted by $T_a = 0^\circ\text{C}$), shown in figure 15(b). Each annealing treatment was performed in air and lasted 10 h.

T_a ($^\circ\text{C}$)	a (10^{-4} rad Hz $^{-b}$)	b
0	4.4 ± 0.4	0.05 ± 0.01
500 $^\circ\text{C}$	1.6 ± 0.2	0.13 ± 0.01
600 $^\circ\text{C}$	1.7 ± 0.2	0.12 ± 0.01
700 $^\circ\text{C}$	0.8 ± 0.1	0.18 ± 0.02

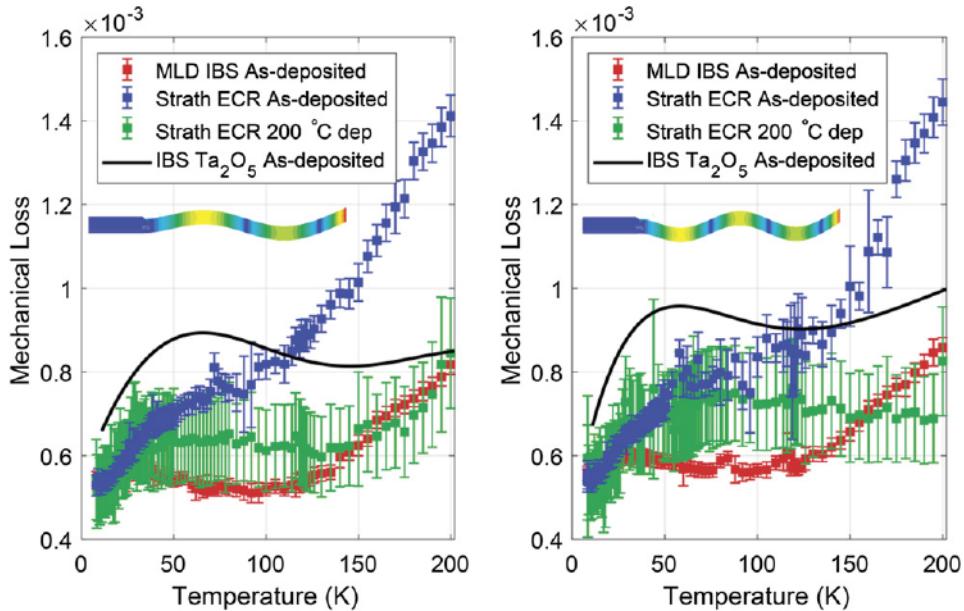


Figure 16. Cryogenic mechanical loss of as-deposited ECR-IBD $Ta_2O_5-ZrO_2$ coatings as a function of sample temperature, as measured on clamped $\sim 65\ \mu\text{m}$ thick silicon cantilevers. Left: (a) results for the third-order bending mode at $\sim 1.4\ \text{kHz}$; right: (b) results for the fourth-order bending mode at $\sim 2.7\ \text{kHz}$. Data of an ECR-IBD sample grown through standard deposition (blue markers) and that of an ECR-IBD grown on a substrate held at $200\ ^\circ\text{C}$ (green markers) are shown, and compared to the data of as-deposited IBS $Ta_2O_5-ZrO_2$ MLD2018 (see figure 14) and Ta_2O_5 [39] coating samples. Finite-element representations of the measured resonant mode are visible in each plot.

The same method was applied to perform cryogenic loss measurements of coated silicon cantilevers, whose results are shown in figure 16 for the third- and fourth-order bending modes at $\sim 1.4\ \text{kHz}$ and $\sim 2.7\ \text{kHz}$, respectively, and compared to the results obtained for the IBS sample MLD2018 (see section 4.1.1 and figure 14). Below $\sim 100\ \text{K}$, both the ECR-IBD and IBS as-deposited $Ta_2O_5-ZrO_2$ coatings have a lower coating loss than that of as-deposited IBS Ta_2O_5 [39]. This shows that, prior to annealing, the addition of ZrO_2 already reduced the coating loss at low temperature. Below $30\ \text{K}$, the loss of the ECR-IBD coatings is comparable to that of IBS coatings.

Figure 16 also shows the data of coating loss for an ECR-IBD $\text{Ta}_2\text{O}_5-\text{ZrO}_2$ coating sample grown on a substrate held at 200 °C. By comparing the ECR-IBD samples, it can be seen that heating the substrate to a higher temperature had a considerable effect on the low-temperature coating loss, between 100 and 200 K: for instance, the coating loss of the sample grown at 200 °C is 1.8 times smaller at 200 K. Finally, we found that the coating loss of the ECR-IBD sample grown at 200 °C is similar to that of the IBS sample MLD2018 in its as-deposited state.

Coating Young modulus Y_c was measured using a Bruker Hysitron TriboScope nano-indenter system, with a three-sided Berkovich pyramid diamond tip ($Y_{\text{indenter}} = 1140 \text{ GPa}$, and $\nu_{\text{indenter}} = 0.007$) and found to be $Y_c = 176 \pm 8 \text{ GPa}$, assuming a coating Poisson ratio of $\nu_c = 0.297$ determined by molecular dynamics models [66]. Measurements were taken on as-deposited thin films grown on 1 mm thick JGS-1 fused-silica witness samples, following the Oliver–Pharr method [2].

4.3. RBTD

Coating mechanical loss, Young modulus and Poisson ratio were estimated by measuring suspended coated fused-silica disks. The disks were measured by the LIGO Laboratory group at Caltech, in a setup with four independent GeNS systems [25] installed in the same vacuum chamber to simultaneously measure all the accessible resonant modes of four samples [54].

Prior to coating deposition, the resonant frequencies and mechanical loss of the disks were characterized through specific measurements, which served as a reference and for background noise estimation for the subsequent series of measurements performed after deposition. All surfaces of the bare disks, including the edge, were polished to a standard optical grade to reduce the contribution of substrate surface losses [5, 24].

After coating deposition, the disks were measured again, and then repeatedly annealed at increasing peak temperature T_a , from 300 to 800 °C. After each annealing treatment, the resonant frequencies and mechanical loss of the disks were measured. Knowing the coating thickness and density from the previous characterizations (see sections 2 and 3), it was possible to compute the coating elastic properties, i.e. Young modulus Y_c and Poisson ratio ν_c , with a procedure similar to what explained in [32, 52] and in [57] and to what was done in analyzing the samples grown by LMA. The measured elastic properties, with their estimation uncertainties, were then used to compute the dilution factors D and extract the coating loss angle according to equation (3). The results are presented in table 6 and figures 18 and 19, and discussed in section 4.5.

4.4. HiPIMS/RF-MS

Coating mechanical loss, Young modulus Y_c and Poisson ratio ν_c were measured on suspended coated fused-silica disks by the LIGO Laboratory group at Caltech, using the same setup and analysis methods described in section 4.3.

HiPIMS coating samples with different cation ratios $\eta = \text{Zr}/(\text{Zr} + \text{Ta})$ were produced, in order to determine how ZrO_2 content affects the loss angle. Also, as mentioned in section 3.5, RF-MS coating samples were grown with the same Ar: O_2 gas ratio, which had higher oxygen content and hence lower densities and refractive index. Samples were then annealed at increasing peak temperature T_a by steps of 50 and 100 °C, starting from 400 °C, until they showed signs of crystallization. The annealing time for each soaking temperature T_a was 10 h. Depending on the cation ratio, the onset of crystallization was observed at temperatures which ranged from $T_a = 700^\circ\text{C}$ to $T_a = 850^\circ\text{C}$.

Figure 17 presents the coating loss angle as a function of the cation ratio η at 650°C, where it can be seen that most values are overlapping. This means that, within the uncertainties of

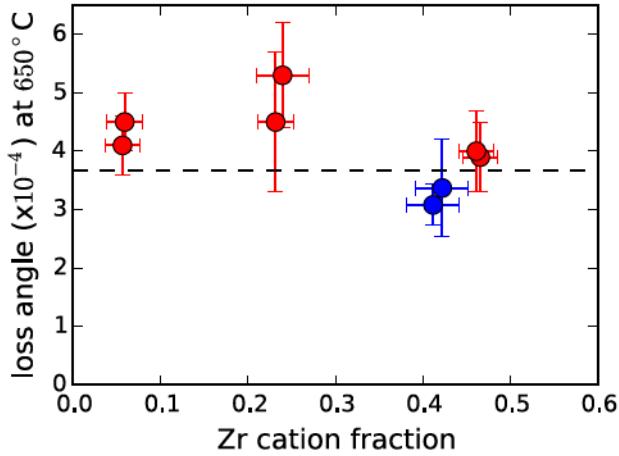


Figure 17. Coating loss of HiPIMS (red markers) and RF-MS (blue markers) $\text{Ta}_2\text{O}_5\text{-ZrO}_2$ thin films as a function of cation ratio $\eta = \text{Zr}/(\text{Zr} + \text{Ta})$, as measured on suspended fused-silica disks after annealing at 650°C . The dashed black line represents the loss angle value of the annealed $\text{Ta}_2\text{O}_5\text{-TiO}_2$ layers of Advanced LIGO and Advanced Virgo [32].

the measurements, η does not considerably affect the coating loss angle. Although still within measurement uncertainty, however, RF-MS samples feature a slightly lower coating loss angle. Therefore, even if HiPIMS samples were shown to be denser and closer to stoichiometry than RF-MS samples, this did not translate into a decrease of the coating loss angle.

4.5. Comparative summary

We discuss here the effects of different cation ratios $\eta = \text{Zr}/(\text{Zr} + \text{Ta})$ and annealing peak temperatures T_a on the mechanical loss of our coating samples, limiting our analysis to the results obtained at ambient temperature so that we can compare samples from all the growth techniques presented in this work. To ease comparison, an average coating loss $\bar{\varphi}$ can be taken from every set of measured resonant modes of each sample.

Figure 18 shows the resulting $\bar{\varphi}$ as a function of cation ratio η , where it can be seen that η does not affect significantly the coating mechanical loss. Furthermore, coating loss seems to be fairly independent of the technique used to grow the samples.

The samples were progressively annealed through steps of increasingly high peak temperature T_a to test the effect of annealing on coating loss. For these experiments, the fused silica substrate is typically annealed to temperatures higher than 900°C before being coated. The high temperature annealing of the bare substrate lowers its mechanical loss ($\varphi_s < 10^{-7}$) such that it is negligible compared to the coating loss. In addition, the substrate loss will not change with lower temperature annealing cycles, so that any change in the coated sample loss may be attributed solely to the coating.

Figure 19 shows that annealing significantly reduced the coating mechanical loss, from about 10^{-3} for as-deposited samples down to $(2\text{--}4) \times 10^{-4}$ for samples treated at 800°C .

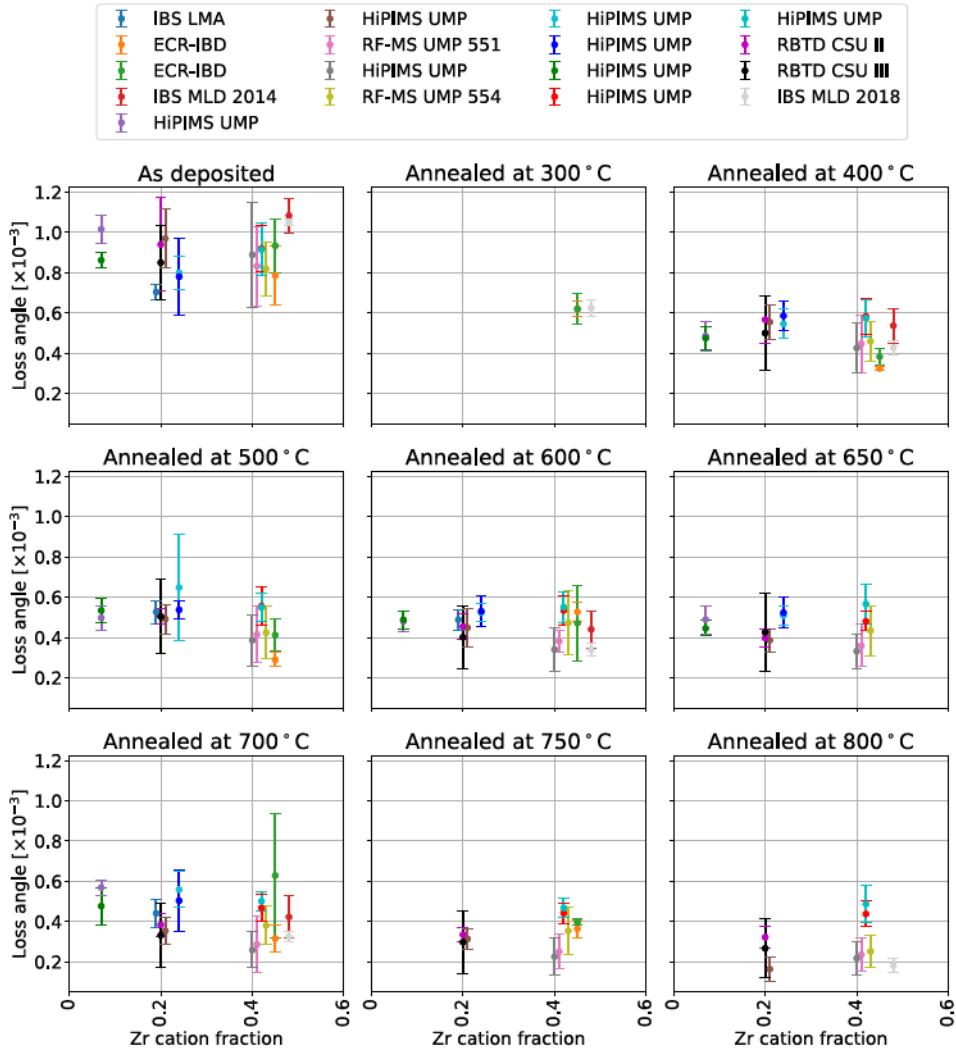


Figure 18. Frequency-averaged coating loss angle $\bar{\varphi}$ of all the $Ta_2O_5-ZrO_2$ coating samples considered in this work, as a function of cation ratio $\eta = Zr/(Ta + Zr)$. Each panel corresponds to a different annealing peak temperature T_a .

5. Effects of annealing

As expected, co-sputtering zirconia proved to be an efficient way to frustrate crystallization in tantalum thin films, allowing for a substantial increase of the maximum annealing temperature T_a . Figure 20 shows that higher cation ratios $\eta = Zr/(Zr + Ta)$ allowed to increase T_a up to 800 °C, without crystallization occurring. More specifically, IBS samples have been tested up to 800 °C and no sign of crystallization was observed. RBTD sample CSU II ($\eta = 0.23$) was amorphous after treatment at 800 °C. HiPIMS/RF-MS samples showed signs of crystallization for temperatures ranging from 700 to 850 °C, depending on the cation ratio η .

Figures 18 and 19 show that annealing significantly reduced the coating mechanical loss, from about $\bar{\varphi} \sim 10^{-3}$ for the as-deposited samples down to $\bar{\varphi} \sim (2 - 4) \times 10^{-4}$ for the samples

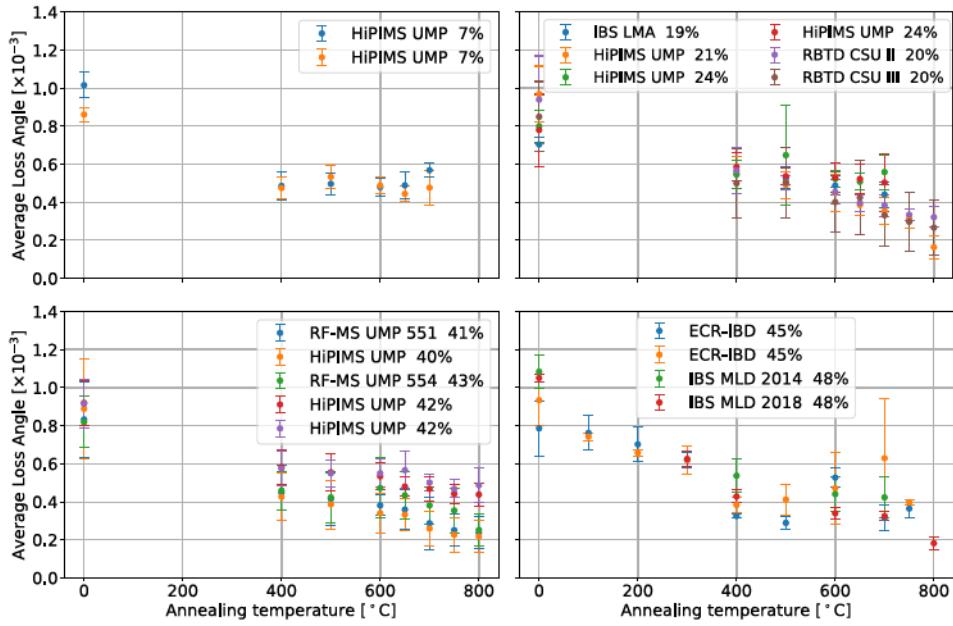


Figure 19. Frequency-averaged coating loss angle $\bar{\varphi}$ of all the $Ta_2O_5-ZrO_2$ coating samples considered in this work, as a function of the annealing peak temperature T_a (as-deposited samples are denoted by $T_a = 0^\circ\text{C}$). The four panels group together different samples with similar cation ratio $\eta = \text{Zr}/(\text{Ta} + \text{Zr})$.

treated at 800°C , whereas, within the measurement uncertainties, the doping does not affect the loss angle for a given T_a .

However, cracks appeared on IBS (MLD), RBTD and HiPIMS/RF-MS samples after annealing at 500°C , and on single-layer IBS samples (LMA) after annealing at 700°C . By way of example, figure 21 shows the cracks observed on an IBS sample (LMA) with an optical microscope; such cracks covered the whole surface of the sample, and their width was of the order of 10^{-5} m . Likely, this happened because of a mismatch between the thermal expansion coefficients of the thin films and substrates, causing high thermal stress. Although in principle our mechanical loss measurements are sensitive to such defects, we found no evidence that our results were affected. At least for some IBS samples (LMA), we avoided the formation of cracks by decreasing the annealing heating and cooling rates down to 7°C h^{-1} . Such samples could be annealed up to 800°C without any cracks forming at their surface.

Nevertheless, starting from $T_a = 500^\circ\text{C}$, bubble-like defects of different number and size, detected with an optical microscope, appeared at variable depth in the IBS HR coating samples, likely due to the presence of incorporated argon within the layers [29, 43, 70]. We observed no trace of such defects in the annealed IBS single-layer samples deposited under identical conditions, indicating that this phenomenon only occurs when layers are stacked. It is not clear how to avoid this issue yet, and we are currently working to solve this problem.

At the same time, annealing at $T_a = 500^\circ\text{C}$ decreased the optical absorption of the IBS HR samples down to 0.5 ppm.

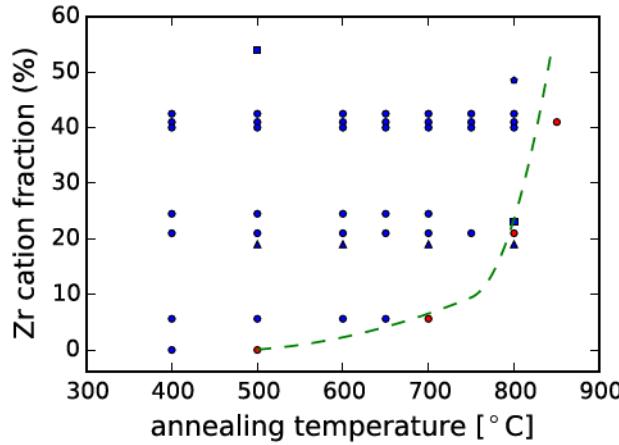


Figure 20. State of IBS, RBTD and HiPIMS/RF-MS Ta_2O_5 – ZrO_2 thin films after annealing at temperature T_a , as a function of cation ratio $\eta = \text{Zr}/(\text{Ta} + \text{Zr})$. Blue markers denote amorphous samples, while red ones indicate that the transition to a polycrystalline phase was observed. RBTD samples are denoted by squares, IBS samples by triangles (LMA) and pentagons (MLD), HiPIMS/RF-MS samples by circles. The dashed curve shows the approximate limit of the phase transition as a function of η .



Figure 21. Cracks observed on an IBS Ta_2O_5 – ZrO_2 thin film (LMA) after annealing at $T_a = 700^\circ\text{C}$.

6. Raman spectroscopy

Raman spectroscopy probes the vibrational modes of a material constituents. In previous studies on amorphous silicon (a-Si), Roorda *et al* found a correlation between structural relaxation and the width of a peak in its Raman spectrum [51]. A sharpening of the peak was interpreted as the results of a decrease in the bond angle distribution upon annealing. Granata *et al* also

found a correlation between the mechanical loss and the normalized area of a Raman line in IBS and bulk silica [31]. Thus, it is of interest to see if there is a correlation between the Raman spectra of $Ta_2O_5-ZrO_2$ thin films and their loss angle, via structural relaxation.

Raman spectroscopy was performed on the coated silica disks with a Renishaw Invia Reflex microscope, with an Ar laser at a wavelength of 514 nm. The measurement was performed three times at different locations on each sample, then the spectra of each sample were averaged. It was carried out on all the MS as-deposited samples and on seven MS samples after annealing at 800 °C. Note that two of these annealed samples showed signs of crystallization, with significant change in the coating Young modulus and loss angle measurements (see table 6 and sections 4.4 and 4.5). Figure 22(a) presents an example of the Raman spectrum measured on a crystallized sample, where the arrows identify peaks associated with the crystal phase [45].

Tantala–zirconia has a complex structure, which translates into several peaks in the Raman spectrum, some of them overlapping, and makes the data analysis non-trivial. In addition, $Ta_2O_5-ZrO_2$ thin films are transparent at 514 nm, so that peaks from the underlying silica substrate also contribute to the measured spectra. Therefore, the initial step of the analysis was to decorrelate the contribution of the thin film from that of the substrate. Figure 22(a) shows a Raman spectrum of an amorphous sample (red curve), with, superimposed, a model (black) resulting from different contributions summed together: (i) the spectrum of a bare silica substrate (blue), (ii) the spectrum of a pure tantalum thin film (green) obtained from a previous study [55], (iii) a background contribution represented by an exponential (dashed gray), (iv) additional contributions required so that the model correctly fit the data (purple), which are then assumed to be due to the addition of zirconia and which were modeled using four Lorentzian functions. Thanks to this model, we can identify where the thin film is the main contributor to the measured spectrum: referring to the purple curve of figure 22(a), this can be seen to happen in the interval from 600 to 750 cm^{-1} . An additional contribution from the thin film is observed near 900 cm^{-1} , but is less clearly separated from the substrate signal and was disregarded in the following analysis.

Next, we wanted to estimate the width of the Raman peak attributed to the thin film. However, trying to subtract substrate contributions or using a fit model would likely influence the shape of the peak, hence its width. To avoid this, we rather relied directly on the experimental data: we calculated the maximum slope of the raw signal on one side of the peak, at a location where the coating contribution to the spectrum is dominant. Indeed, the slope of a peak edge is a proxy for the peak width: for both Lorentzian and Gaussian functions, the maximum slope is proportional to the inverse squared width. Thus narrower peaks, which might be attributed to a more relaxed material, should feature a steeper slope on each side. To this end, the raw spectra were smoothed with a bi-cubic spline and normalized so that their maximum, in the 550 cm^{-1} to 750 cm^{-1} spectral region, where the contribution of the thin film is at the highest, is at unity. We then computed the derivative, an example of the result is found in figure 22(b).

In all panels of figure 22, the position of interest, i.e. the point where the slope of the signal related to the Zr contribution is the steepest, is represented by a vertical red line. In figure 22(c), we can see two raw spectra of the same sample, acquired before and after annealing. The slope computed using the spline fit at the Raman shift of interest is also plotted as a straight line, for each spectrum. We can observe from the raw data that the slope around 705 cm^{-1} changes upon annealing and is well represented by the first derivative deduced from the method illustrated in figure 22(b). If Raman peaks sharpen with relaxation, and relaxation decreases the loss angle, we expect a higher value of the derivative with decreasing loss angle.

Figure 23 shows the absolute value of the first derivative around 705 cm^{-1} as a function of the loss angle, where the color scale represent the cation ratio $\eta = \text{Zr}/(\text{Zr} + \text{Ta})$. The red markers correspond to the two crystallized samples which show characteristic peaks in their

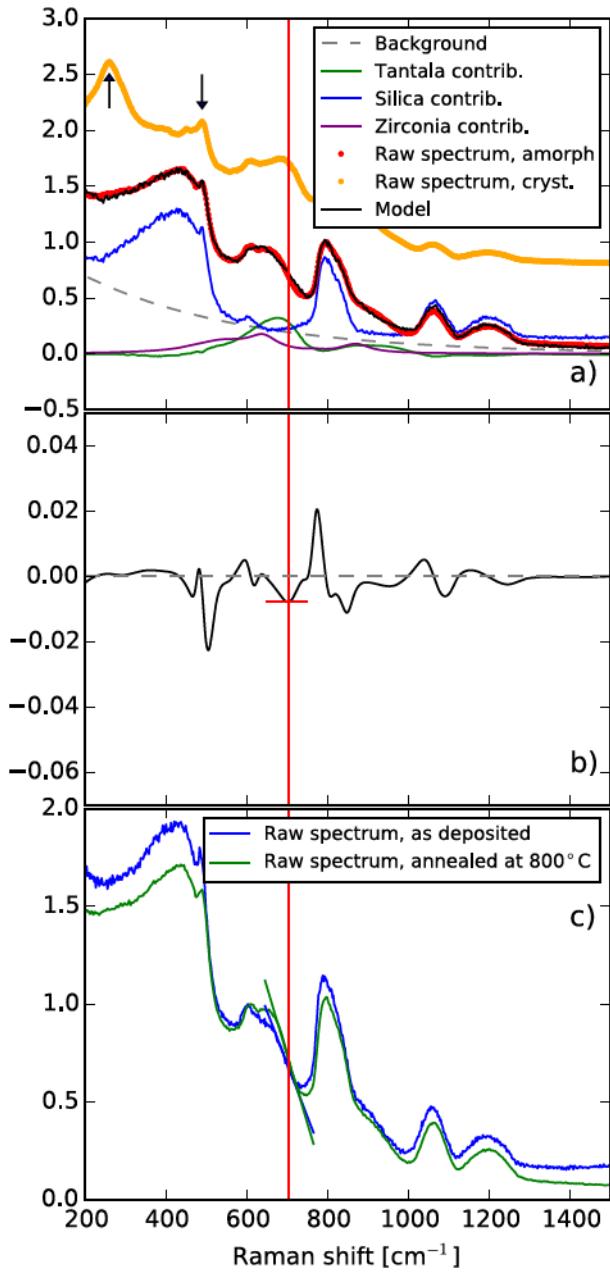


Figure 22. Measured Raman spectra of HiPIMS/RF-MS Ta₂O₅–ZrO₂ thin films and their analysis: (a) example of a spectrum obtained for an amorphous sample with, superimposed, a model resulting from different contributions summed together (see text for details); (b) derivative of the raw Raman spectrum, smoothed using bi-cubic splines; (c) raw, normalized spectrum of the same sample, as-deposited and annealed. The slopes extracted using the derivative method shown in (b) are represented by straight lines. The vertical red line across all panels is the position of the minimum of the first derivative of the Zr contribution shown in (a).

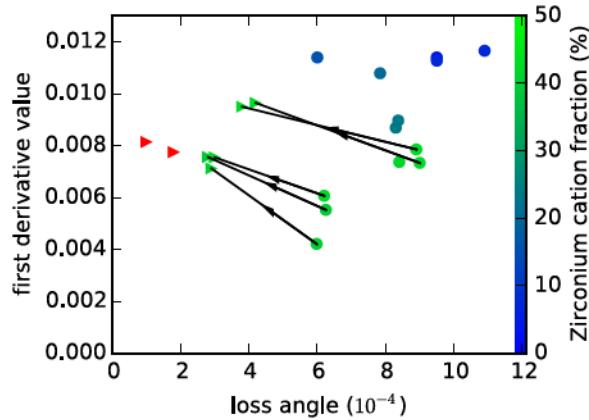


Figure 23. Absolute value of the minimum of the derivative of the Raman spectrum near 705 cm^{-1} as a function of the loss angle, for HiPIMS/RF-MS $\text{Ta}_2\text{O}_5\text{-ZrO}_2$ thin films. The color scale corresponds to the cation ratio $\eta = \text{Zr}/(\text{Zr} + \text{Ta})$, circles denote spectra acquired on the as-deposited samples and triangles denote those acquired on the annealed samples, the same sample being linked by a black arrow. Red markers represent samples showing signs of crystallization.

Raman spectrum, and most probably some delamination, making their loss angle measurements unreliable. The circles are for measurements carried out on as-deposited samples while the triangles are for the annealed ones, and the same samples are linked by an arrow pointing toward the annealed one.

Figure 23 clearly shows that, upon annealing, each individual sample sees its peak slope increase and its loss angle decrease at the same time. Hence, the Raman peak sharpens upon relaxation, similar to what is observed in a-Si [51]. However, if one considers only the subset of the as-deposited samples (circles), it appears that the lower the loss angle, the lower the slope, which means the Raman peaks are wider. Also, samples that have the lowest loss angle are the ones with the highest cation ratio η (greener data points). Hence, the shape of the Raman peak near 705 cm^{-1} cannot be attributed solely to relaxation, as it is also clearly influenced by the micro-structural modifications that large amounts of Zr induce into the thin films, which in turn have an impact on the vibrational states distribution and hence on the shape of the Raman spectrum.

Therefore, these results show that Raman spectroscopy could be used as a probe to monitor the relaxation state of tantalum-based alloys, and that relaxation is linked to the loss angle. Similar conclusions were reached using x-ray diffraction on tantalum [55], where the x-ray peak of tantalum sharpened with decreasing loss angle. However, the Raman peak shape is also influenced by the amount of doping, so while it can be used to monitor the evolution of a specific sample, it would be difficult to use it more generally as a probe of the state of relaxation in a material or a predictor of the loss angle.

7. Summary and conclusions

In order to decrease coating Brownian noise in current GW detectors, we developed a set of co-sputtered tantalum-zirconia thin films. We extensively characterized these coatings for optical

and mechanical properties, to test their performances against the very stringent requirements of GW detectors.

As expected, co-sputtering zirconia proved to be an efficient way to frustrate crystallization in tantalum thin films, allowing for a substantial increase of the maximum annealing temperature T_a , hence for a decrease of coating mechanical loss. The lowest average loss, $\bar{\varphi}_c = 1.8 \times 10^{-4}$ rad, was measured on IBS sample MLD2018 with cation ratio $Zr/(Ta + Zr) = 48.5 \pm 0.4\%$ annealed at 800 °C. Higher annealing temperatures were prevented by the onset of crystallization.

The lowest loss value measured on co-sputtered tantalum–zirconia thin films has to be compared to $\varphi_c = (2.3–3.4) \times 10^{-4}$ rad of tantalum–titania layers in current GW detectors [32, 82], and to other high-index oxide coatings, either simple or co-sputtered [29, 33, 59–62]. Further coating loss reduction, down to 1×10^{-4} rad or lower, might be obtained with non-oxide coatings [33].

As measured on IBS HR samples with the same design of those of Advanced LIGO and Advanced Virgo, we observed as little optical absorption (0.5 ppm) as in current GW detectors [27, 46]. The same samples, however, featured an unusually high value of scattering, 45 ppm, compared to that of present detectors (5 ppm) [27, 46]. Indeed, part of the observed excess scatter might be ascribed to the presence of Ar bubbles formed upon annealing [29, 43, 70].

For some samples, we measured a refractive index at 1064 and 1550 nm higher than that of pure tantalum thin films. This implies that HR stacks containing zirconia-doped tantalum layers could be even thinner than HR coatings of present detectors, further decreasing coating Brownian noise [33].

For use in future GW detectors, further development will be needed to decrease scattering and to avoid the formation of bubble-like defects upon annealing. We are currently working to find appropriate solutions to these issues.

Acknowledgments

The research performed at the Laboratoire des Matériaux Avancés was partially supported by the Virgo Coating Research and Development (VCR & D) Collaboration. The work performed at Université de Montréal and École Polytechnique de Montréal was supported by the Natural Sciences and Engineering Research Council of Canada (NSERC), the Canadian foundation for innovation (CFI) and the Fonds de recherche Québec, Nature et technologies (FQRNT) through the Regroupement Québécois sur les matériaux de pointe (RQMP). These authors thank S Roorda, M Chicoine, L Godbout and F Turcot for fruitful discussions and technical support. The work performed at the University of Glasgow was supported by the Science and Technology Facilities Council (STFC, ST/N005422/1 and ST/I001085/1) and the Royal Society (UF100602 and UF150694). The research performed at Hobart and William Smith Colleges was supported by National Science Foundation Grant awards PHY-1307423, PHY-1611821, and PHY-1707863.

Data availability statement

The data that support the findings of this study are available upon reasonable request from the authors.

ORCID iDs

- M Abernathy  <https://orcid.org/0000-0001-9805-4311>
 A Amato  <https://orcid.org/0000-0001-9557-651X>
 R Bassiri  <https://orcid.org/0000-0001-8171-6833>
 M Canepa  <https://orcid.org/0000-0002-5148-1233>
 A Di Michele  <https://orcid.org/0000-0002-0357-2608>
 M A Fazio  <https://orcid.org/0000-0002-9057-9663>
 M Granata  <https://orcid.org/0000-0003-3275-1186>
 A M Gretarsson  <https://orcid.org/0000-0003-3438-9926>
 E J Hough  <https://orcid.org/0000-0003-3242-3123>
 I W Martin  <https://orcid.org/0000-0001-7300-9151>
 B Maynard  <https://orcid.org/0000-0003-4799-1545>
 P G Murray  <https://orcid.org/0000-0002-8218-2404>
 S Penn  <https://orcid.org/0000-0003-4956-0853>
 S Rowan  <https://orcid.org/0000-0002-0666-9907>
 F Schiettekatte  <https://orcid.org/0000-0002-2112-9378>
 G Vajente  <https://orcid.org/0000-0002-7656-6882>

References

- [1] Callen H B and Welton T A 1951 Irreversibility and generalized noise *Phys. Rev.* **83** 34–40
- [2] Oliver W C and Pharr G M 1992 An improved technique for determining hardness and elastic modulus using load and displacement sensing indentation experiments *J. Mater. Res.* **7** 1564
- [3] Çetinörgü E, Baloukas B, Zabeida O, Klemburg-Sapieha J E and Martinu L 2009 Mechanical and thermoelastic characteristics of optical thin films deposited by dual ion beam sputtering *Appl. Opt.* **48** 4536
- [4] Ceresoli D and Vanderbilt D 2006 Structural and dielectric properties of amorphous ZrO₂ and HfO₂ *Phys. Rev. B* **74** 125108
- [5] Penn S D 2006 Frequency and surface dependence of the mechanical loss in fused silica *Phys. Lett. A* **352** 3–6
- [6] Jellison G E and Modine F A 1996 Parameterization of the optical functions of amorphous materials in the interband region *Appl. Phys. Lett.* **69** 371
- [7] Aasi J et al 2015 Advanced LIGO *Class. Quantum Grav.* **32** 074001
- [8] Abbott B P et al 2016 Binary black hole mergers in the first advanced LIGO observing run *Phys. Rev. X* **6** 041015
- [9] Abbott B P et al 2016 Observation of gravitational waves from a binary black hole merger *Phys. Rev. Lett.* **116** 061102
- [10] Abbott B P et al 2017 GW170817: observation of gravitational waves from a binary neutron star inspiral *Phys. Rev. Lett.* **119** 161101
- [11] Abbott B P et al 2017 Multi-messenger observations of a binary neutron star merger *Astrophys. J.* **848** L12
- [12] Abbott B P et al 2018 GW170817: measurements of neutron star radii and equation of state *Phys. Rev. Lett.* **121** 161101
- [13] Abbott B P et al 2019 GWTC-1: a gravitational-wave transient catalog of compact binary mergers observed by LIGO and Virgo during the first and second observing runs *Phys. Rev. X* **9** 031040
- [14] Abbott B P et al 2021 GWTC-2: compact binary coalescences observed by LIGO and Virgo during the first half of the third observing run *Phys. Rev. X* **11** 021053
- [15] Acernese F et al 2014 Advanced Virgo: a second-generation interferometric gravitational wave detector *Class. Quantum Grav.* **32** 024001
- [16] Alexandrovski A et al 2009 Photothermal common-path interferometry (pci): new developments *Solid State Lasers XVIII: Technology and Devices* vol 7193 (Bellingham, WA: SPIE Optical Engineering Press) p 71930D

- [17] Amato A *et al* 2018 High-reflection coatings for gravitational-wave detectors: state of the art and future developments *J. Phys.: Conf. Ser.* **957** 012006
- [18] Amato A *et al* 2019 Optical properties of high-quality oxide coating materials used in gravitational-wave advanced detectors *J. Phys. Mater.* **2** 035004
- [19] Amato A *et al* 2020 Observation of a correlation between internal friction and urbach energy in amorphous oxides thin films *Sci. Rep.* **10** 1670
- [20] Birney R *et al* 2018 Amorphous silicon with extremely low absorption: beating thermal noise in gravitational astronomy *Phys. Rev. Lett.* **121** 191101
- [21] Boccara A C, Jackson W, Amer N M and Fournier D 1980 Sensitive photothermal deflection technique for measuring absorption in optically thin media *Opt. Lett.* **5** 377
- [22] Bräuer G, Szyszka B, Vergöhl M and Bandorf R 2010 Magnetron sputtering—milestones of 30 years *Vacuum* **84** 1354
- [23] Buikema A *et al* 2020 Sensitivity and performance of the Advanced LIGO detectors in the third observing run *Phys. Rev. D* **102** 062003
- [24] Cagnoli G *et al* 2018 Mode-dependent mechanical losses in disc resonators *Phys. Lett. A* **382** 2165
- [25] Cesarini E, Lorenzini M, Campagna E, Martelli F, Piergiovanni F, Vetrano F, Losurdo G and Cagnoli G 2009 A ‘gentle’ nodal suspension for measurements of the acoustic attenuation in materials *Rev. Sci. Instrum.* **80** 053904
- [26] Danilishin S L and Khalili F Y 2012 Quantum measurement theory in gravitational-wave detectors *Living Rev. Relativ.* **15** 5
- [27] Degallaix J *et al* 2019 Large and extremely low loss: the unique challenges of gravitational wave mirrors *J. Opt. Soc. Am. A* **36** C85
- [28] Evans M, Ballmer S, Fejer M, Fritschel P, Harry G and Ogin G 2008 Thermo-optic noise in coated mirrors for high-precision optical measurements *Phys. Rev. D* **78** 102003
- [29] Fazio M A, Vajente G, Ananyeva A, Markosyan A, Bassiri R, Fejer M M and Menoni C S 2020 Structure and morphology of low mechanical loss TiO₂-doped Ta₂O₅ *Opt. Mater. Express* **10** 1687
- [30] Granata M *et al* 2016 Mechanical loss in state-of-the-art amorphous optical coatings *Phys. Rev. D* **93** 012007
- [31] Granata M *et al* 2018 Correlated evolution of structure and mechanical loss of a sputtered silica film *Phys. Rev. Mater.* **2** 053607
- [32] Granata M *et al* 2020 Amorphous optical coatings of present gravitational-wave interferometers *Class. Quantum Grav.* **37** 095004
- [33] Granata M *et al* 2020 Progress in the measurement and reduction of thermal noise in optical coatings for gravitational-wave detectors *Appl. Opt.* **59** A229
- [34] Hálá M, Vernhes R, Zabeida O, Bousser E, Klembberg-Sapieha J E, Sargent R and Martinu L 2014 Growth and properties of high index Ta₂O₅ optical coatings prepared by HiPIMS and other methods *Surf. Coat. Technol.* **241** 33
- [35] Hálá M, Vernhes R, Zabeida O, Klembberg-Sapieha J-E and Martinu L 2014 Reactive HiPIMS deposition of SiO₂/Ta₂O₅ optical interference filters *J. Appl. Phys.* **116** 213302
- [36] Harry G M *et al* 2002 Thermal noise in interferometric gravitational wave detectors due to dielectric optical coatings *Class. Quantum Grav.* **19** 897
- [37] Harry G M *et al* 2006 Titania-doped tantalum/silica coatings for gravitational-wave detection *Class. Quantum Grav.* **24** 405
- [38] Klembberg-Sapieha J E *et al* 2004 Mechanical characteristics of optical coatings prepared by various techniques: a comparative study *Appl. Opt.* **43** 2670
- [39] Martin I W *et al* 2010 Effect of heat treatment on mechanical dissipation in Ta₂O₅ coatings *Class. Quantum Grav.* **27** 225020
- [40] Miller J *et al* 2015 Prospects for doubling the range of Advanced LIGO *Phys. Rev. D* **91** 062005
- [41] Murray P *et al* 2021 Mechanical loss reduction from crystallization in a Ta₂O₅–TiO₂ film (in preparation)
- [42] Nakagawa N, Gretarsson A M, Gustafson E K and Fejer M M 2002 Thermal noise in half-infinite mirrors with nonuniform loss: a slab of excess loss in a half-infinite mirror *Phys. Rev. D* **65** 102001
- [43] Paolone A *et al* 2021 Effects of the annealing of amorphous Ta₂O₅ coatings produced by ion beam sputtering concerning the effusion of argon and the chemical composition *J. Non-Cryst. Solids* **557** 120651

- [44] Penn S D *et al* 2003 Mechanical loss in tantalum/silica dielectric mirror coatings *Class. Quantum Grav.* **20** 2917
- [45] Perez I, Carrejo J L E, Sosa V, Perera F G, Mancillas J R F, Galindo J T E and Rodríguez C I R 2017 Evidence for structural transition in crystalline tantalum pentoxide films grown by RF magnetron sputtering *J. Alloys Compd.* **712** 303
- [46] Pinard L *et al* 2017 Mirrors used in the LIGO interferometers for first detection of gravitational waves *Appl. Opt.* **56** C11
- [47] Prato M, Chincarini A, Gemme G and Canepa M 2011 Gravitational waves detector mirrors: spectroscopic ellipsometry study of Ta₂O₅ films on SiO₂ substrates *Thin Solid Films* **519** 2877
- [48] Punturo M *et al* 2010 The Einstein telescope: a third-generation gravitational wave observatory *Class. Quantum Grav.* **27** 194002
- [49] Reid S *et al* 2006 Mechanical dissipation in silicon flexures *Phys. Lett. A* **351** 205–11
- [50] Reitze D *et al* 2019 Cosmic explorer: the U.S. contribution to gravitational-wave astronomy beyond LIGO (arXiv:1907.04833)
- [51] Roorda S, Sinke W C, Poate J M, Jacobson D C, Dierker S, Dennis B S, Eaglesham D J, Spaepen F and Fuoss P 1991 Structural relaxation and defect annihilation in pure amorphous silicon *Phys. Rev. B* **44** 3702
- [52] Tianjun L *et al* 2014 Measurements of mechanical thermal noise and energy dissipation in optical dielectric coatings *Phys. Rev. D* **89** 092004
- [53] Travasso F, Amico P, Bosi L, Cottone F, Dari A, Gammaiton L, Vocca H and Marchesoni F 2007 Low-frequency internal friction in silica glass *Europhys. Lett.* **80** 50008
- [54] Vajente G, Ananyeva A, Billingsley G, Gustafson E, Heptonstall A, Sanchez E and Torrie C 2017 A high throughput instrument to measure mechanical losses in thin film coatings *Rev. Sci. Instrum.* **88** 073901
- [55] Vajente G *et al* 2018 Effect of elevated substrate temperature deposition on the mechanical losses in tantalum thin film coatings *Class. Quantum Grav.* **35** 075001
- [56] Vajente G *et al* 2019 Chapter three—precision interferometry for gravitational wave detection: current status and future trends *Advances in Atomic, Molecular, and Optical Physics* vol 68 ed L F Dimauro, H Perrin and S F Yelin (New York: Academic) p 75
- [57] Vajente G *et al* 2020 Method for the experimental measurement of bulk and shear loss angles in amorphous thin films *Phys. Rev. D* **101** 042004
- [58] Villar A E *et al* 2010 Measurement of thermal noise in multilayer coatings with optimized layer thickness *Phys. Rev. D* **81** 122001
- [59] Yang L, Randel E, Vajente G, Ananyeva A, Gustafson E, Markosyan A, Bassiri R, Fejer M M and Menoni C S 2019 Investigation of effects of assisted ion bombardment on mechanical loss of sputtered tantalum thin films for gravitational wave interferometers *Phys. Rev. D* **100** 122004
- [60] Yang L, Randel E, Vajente G, Ananyeva A, Gustafson E, Markosyan A, Bassiri R, Fejer M and Menoni C 2020 Modifications of ion beam sputtered tantalum thin films by secondary argon and oxygen bombardment *Appl. Opt.* **59** A150–4
- [61] Yang L *et al* 2020 Structural evolution that affects the room-temperature internal friction of binary oxide nanolaminates: implications for ultrastable optical cavities *ACS Applied Nano Mater.* **3** 12308–13
- [62] Yang L *et al* 2021 Enhanced medium range order in vapor deposited germania glasses at elevated temperatures (arXiv:2102.08526)
- [63] Zhurin V V, Kaufman H R, Kahn J R and Hylton T L 2000 Biased target deposition *J. Vac. Sci. Technol. A* **18** 37
- [64] Fujiwara H 2007 *Spectroscopic Ellipsometry: Principles and Applications* (New York: Wiley)
- [65] Arago D F J and Biot J-B 1806 *Mémoire sur les affinités des corps pour la lumière: et particulièrement sur les forces refringentes des différens gaz* (Paris: Bachelier)
- [66] Trinastic J personal communication
- [67] Tewg J-Y, Kuo Y and Lu J 2005 Suppression of crystallization of tantalum oxide thin film by doping with zirconium *Electrochim. Solid State Lett.* **8** G27
- [68] Chu W K 1978 *Backscattering Spectrometry* (New York: Academic) (<https://doi.org/10.1016/B978-0-121-73850-1.X5001-7>)
- [69] Wood D L and Nassau K 1982 Refractive index of cubic zirconia stabilized with yttria *Appl. Opt.* **21** 2978
- [70] Cevro M and Carter G 1995 Ion-beam and dual-ion-beam sputter deposition of tantalum oxide films *Opt. Eng.* **34** 596

- [71] Fairley M 2005 <http://www.casaxps.com>
- [72] Fejer M M 2016 Thermal noise in mirror coatings for gravitational wave detection *Optical Interference Coatings 2016* (Optical Society of America) p MB.1
- [73] Gretarsson E M and Gretarsson A M 2018 Three methods for characterizing thermo-optic noise in optical cavities *Phys. Rev. D* **98** 122004
- [74] Mayer M 1999 SIMNRA, a simulation program for the analysis of NRA, RBS and ERDA *AIP Conf. Proc.* **475** 541
- [75] Kaiser N and Pulker H K 2003 *Optical Interference Coatings* (Berlin: Springer)
- [76] NIST Standard Reference Database Number 20 2000 <https://srdata.nist.gov/xps/>
- [77] Robie R 2018 Characterisation of the mechanical properties of thin-film mirror coating materials for use in future interferometric gravitational wave detectors *PhD Thesis* University of Glasgow
- [78] Saulson P R 1994 *Fundamentals of Interferometric Gravitational Wave Detectors* (Singapore: World Scientific)
- [79] Berry B S and Pritchett W C 1975 Vibrating reed internal friction apparatus for films and foils *IBM J. Res. Dev.* **19** 334
- [80] Ferlaudo A S, Ferreira G M, Pearce J M, Wronski C R, Collins R W, Deng X and Ganguly G 2002 Analytical model for the optical functions of amorphous semiconductors from the near-infrared to ultraviolet: applications in thin film photovoltaics *J. Appl. Phys.* **92** 2424
- [81] Gilroy K S and Phillips W A 1981 An asymmetric double-well potential model for structural relaxation processes in amorphous materials *Phil. Mag. B* **43** 735
- [82] Gras S and Evans M 2018 Direct measurement of coating thermal noise in optical resonators *Phys. Rev. D* **98** 122001
- [83] Nowick A S and Berry B S 1972 *Anelastic Relaxation in Crystalline Solids* (New York: Academic)
- [84] Reid S and Martin I 2016 Development of mirror coatings for gravitational wave detectors *Coatings* **6** 61
- [85] Bright T J, Watjen J I, Zhang Z M, Muratore C, Voevodin A A, Koukis D I, Tanner D B and Arenas D J 2013 Infrared optical properties of amorphous and nanocrystalline Ta₂O₅ thin films *J. Appl. Phys.* **114** 083515
- [86] Hong T 2013 Brownian thermal noise in multilayer coated mirrors *Phys. Rev. D* **87** 082001