Body rocking or lift off in flow



Te Whare Wānanga o Waitaha CHRISTCHURCH NEW ZEALAND

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30 January 2014

Configuration

Small time properties

O(1) time, no flow

O(1) time, flow

Lift off

Discussion

Rock 'n' roll on Mars





HiRISE/MRO/LPL/NASA (Barchans $\sim 200m$)

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Carneiro *et al.*, Phys. Rev. Lett. (2013)

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HiRISE/MRO/LPL/NASA (Barchans $\sim 200m$)

Small time properties

O(1) time, no

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Other motivations





Cox et al., J. Geol. (2012)

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Governing equations

$$H(x, t) = h(t) + (x - x_c)\theta(t) - F(x);$$

$$H = H_x = 0 \text{ at } x = x_0(t).$$



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$$H = H_x = 0 \text{ at } x = x_0(t).$$

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In both gaps:

$$\begin{aligned} H_t + (uH)_x &= 0, \\ u_t + uu_x &= -p_x; \\ u &= \dot{x}_0(t) \quad \text{at } x = x_0(t), \end{aligned}$$



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$$H(x, t) = h(t) + (x - x_c)\theta(t) - F(x);$$

$$H = H_x = 0 \text{ at } x = x_0(t).$$

In both gaps:

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$$\begin{array}{ll} H_t + (uH)_x = 0, & p + \frac{1}{2}u^2 = \frac{1}{2} & \text{at } x = 0, \\ u_t + uu_x = -p_x; & \text{in Gap I, but in Gap II} \\ = \dot{x}_0(t) & \text{at } x = x_0(t), & p = 0 & \text{at } x = 1. \end{array}$$

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$$H(x, t) = h(t) + (x - x_c)\theta(t) - F(x);$$

$$H = H_x = 0 \text{ at } x = x_0(t).$$

In both gaps:

$$H_t + (uH)_x = 0,$$
 $p + \frac{1}{2}u^2 = \frac{1}{2}$ at $x = 0,$
 $u_t + uu_x = -p_x;$ in Gap I, but in Gap II
 $u = \dot{x}_0(t)$ at $x = x_0(t),$ $p = 0$ at $x = 1.$

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$$egin{aligned} & M\ddot{h}(t) = \int_{0}^{1} p(x,t) \, \mathrm{d}x + N(t) - Mg^{+}, \ & I\ddot{ heta}(t) = \int_{0}^{1} (x-x_{c})p(x,t) \, \mathrm{d}x + (x_{0}-x_{c})N(t). \end{aligned}$$



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Numerical solutions of small-t equations



Figure: Sinusoidal body $F(x) = \sin(\pi x)$.



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Figure: Elliptical body $F(x) = \sqrt{x - x^2}$.



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Figure: Smooth-cornered body.

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Figure: Constant-curvature body F(x) = x(1 - x).



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Figure: Constant-curvature body with $x_0(0) = 0.7$.



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Figure: Constant-curvature body with $x_0(0) = 0.7$.



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O(1) time, negligible fluid effects

Mass- and moment of inertia-dominated.

$$\begin{array}{rcl} M\ddot{h}(t) &=& N(t) - Mg^+ \ , \\ I\ddot{\theta}(t) &=& (x_0 - x_c)N(t) \ , \\ H(x,t) &=& h(t) + (x - x_c)\theta(t) - F(x) \ , \\ H &=& H_x = 0 \ \text{at} \ x = x_0(t) \ . \end{array}$$

In subsequent analysis, a key equation:

$$\alpha \ddot{x}_0 + \beta \dot{x}_0^2 = (x_0 - x_c)g^+.$$

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Here, α , β depend on I, M, and body shape.



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Evolution of contact position



Figure: Sinusoidal body, with $g^+ = 10$, $x_c = 0.5$, varying initial contact point location, and zero initial contact point velocity.



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Figure: Elliptical body, with $g^+ = 10$, $x_c = 0.5$, varying initial contact point location, and zero initial contact point velocity.





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Figure: Analytical prediction of rocking behaviour obtained by asymptotic analysis. Sinusoidal body, conditions as before. Unscaled on the right. Dots are numerical solutions.



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Figure: Analytical prediction of rocking behaviour obtained by asymptotic analysis. Elliptical body, conditions as before. Unscaled on the right. Dots are numerical solutions.



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O(1) time, fluid effects



Figure: Sinusoidal body, M = I = 0.125, $x_0(0) = 0.33$, $\dot{x}_0(0) = 0$.



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Figure: Sinusoidal body, M = I = 0.125, $x_0(0) = 0.33$, $\dot{x}_0(0) = 0$.



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Figure: Sinusoidal body, M = I = 0.125, $\dot{x}_0(0) = 0$, and $x_0(0) = 0.33$ (solid lines) or $x_0(0) = 0.28$ (dashed lines).

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Figure: Elliptical body, M = I = 0.08, $x_0(0) = 0.25$, $\dot{x}_0(0) = 0$.



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Figure: Elliptical body, M = I = 0.05, $x_0(0) = 0.25$, $\dot{x}_0(0) = 0$. Lift off occurs at $t \approx 2$ (subsequent results not physically meaningful).



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Figure: Smooth body, M = I = 0.1, $x_0(0) = 0.25$, $\dot{x}_0(0) = 0$. Lift off occurs at $t \approx 1.6$ (subsequent results not physically meaningful).



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Transition from rocking to lift off is smooth

With negligible flow, a key equation was:

$$\alpha \ddot{x}_0 + \beta \dot{x}_0^2 = (x_0 - x_c)g^+,$$

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where α, β depended on M, I, and body shape.



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Transition from rocking to lift off is smooth

With negligible flow, a key equation was:

$$\alpha \ddot{x}_0 + \beta \dot{x}_0^2 = (x_0 - x_c)g^+,$$

where α, β depended on *M*, *I*, and body shape. With fluid effects, the corresponding equation becomes:

$$\alpha \ddot{x}_0 + \beta \dot{x}_0^2 = (x_0 - x_c)g^+ + \frac{i_2 - (x_0 - x_c)i_1}{M},$$

where

$$i_1 = \int_0^1 p(x,t) \, dx$$
, $i_2 = \int_0^1 (x-x_c)p(x,t) \, dx$.

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Transition from rocking to lift off is smooth

The integrated flow-pressure contributions i₁, i₂ move us into a part of solution space unobtainable in no-flow case.

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Transition from rocking to lift off is smooth

- The integrated flow-pressure contributions i₁, i₂ move us into a part of solution space unobtainable in no-flow case.
- For small times $t \to t_{LO}-$, consider the elliptical body with x_0 close to the leading edge, so that $x = \epsilon X$, $\epsilon \ll 1$ (similar arguments hold at trailing edge).

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Transition from rocking to lift off is smooth

- The integrated flow-pressure contributions i₁, i₂ move us into a part of solution space unobtainable in no-flow case.
- For small times $t \to t_{LO}-$, consider the elliptical body with x_0 close to the leading edge, so that $x = \epsilon X$, $\epsilon \ll 1$ (similar arguments hold at trailing edge).
- Analysis shows that u ~ e^{5/4} leading to a pressure response being an O(e^{5/2}) perturbation from the value 1/2.

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Transition from rocking to lift off is smooth

- The integrated flow-pressure contributions i₁, i₂ move us into a part of solution space unobtainable in no-flow case.
- For small times $t \to t_{LO}-$, consider the elliptical body with x_0 close to the leading edge, so that $x = \epsilon X$, $\epsilon \ll 1$ (similar arguments hold at trailing edge).
- Analysis shows that u ~ e^{5/4} leading to a pressure response being an O(e^{5/2}) perturbation from the value 1/2.
- Thus the lift-off generating mechanism of i₁, i₂ an effect of "added mass" and evolution in the fluid-body interaction is still dominated by the O(1) global contributions.

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Discussion Smith & Wilson, J. Fluid Mech. (2013)

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• Flow-dominated bodies tend to lift off immediately.



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Discussion Smith & Wilson, J. Fluid Mech. (2013)

- Flow-dominated bodies tend to lift off immediately.
- Without flow, bodies tend to rock rather than lift off.

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Rocking is well-understood analytically.



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Discussion Smith & Wilson, J. Fluid Mech. (2013)

- Flow-dominated bodies tend to lift off immediately.
- Without flow, bodies tend to rock rather than lift off.
 Rocking is well-understood analytically.
- For full fluid-structure interaction problem, added-mass effect and flow evolution lead to either rocking or lift off.

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□ The parameter space is subtle.



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- □ The parameter space is subtle.
- Rocking transitions smoothly to lift off.



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 Rocking is well-understood analytically.
- For full fluid-structure interaction problem, added-mass effect and flow evolution lead to either rocking or lift off.
 - □ The parameter space is subtle.
 - Rocking transitions smoothly to lift off.
- A body is "light" and lifts off immediately, or is "heavy" and needs a push from pressures in the narrowing gaps.

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Future work.



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- Future work.
 - □ Generalise upstream flow shear, boundary layers, etc.

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□ Post-lift off: flow through gap equilibrates pressure.



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 - Irregular body shape, groups of bodies; body flex; surface shape, curvature, roughness; surface fluid; 3D effects.



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- A body is "light" and lifts off immediately, or is "heavy" and needs a push from pressures in the narrowing gaps.
- Future work.
 - □ Generalise upstream flow shear, boundary layers, etc.
 - □ Post-lift off: flow through gap equilibrates pressure.
 - Irregular body shape, groups of bodies; body flex; surface shape, curvature, roughness; surface fluid; 3D effects.
 - Reptation and clashes, multiple bodies
 - see also Smith & Wilson, Proc. Roy. Soc. A (2011).



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Lift off on Mars

• Martian gravitational acceleration \sim 0.38 that of Earth.



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• Martian gravitational acceleration \sim 0.38 that of Earth.

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• Martian atmospheric density \sim 0.0167 that of Earth.



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Lift off on Mars

- Martian gravitational acceleration \sim 0.38 that of Earth.
- Martian atmospheric density \sim 0.0167 that of Earth.
 - Dimensional threshold wind speed for lift-off on Mars is 2–3 times that on Earth.

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• Threshold speed $\propto \sqrt{\text{particle size}}$.



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- Martian gravitational acceleration \sim 0.38 that of Earth.
- Martian atmospheric density \sim 0.0167 that of Earth.
 - Dimensional threshold wind speed for lift-off on Mars is 2–3 times that on Earth.
 - Threshold speed $\propto \sqrt{\text{particle size}}$.
 - Both results consistent with Martian observations in Wang, Int. J. Land Processes Arid Enviro. (2012).

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Figure: Elliptical body, M = I = 0.05, initially $x_0(0) = 0.25$ and $\dot{x}_0(0)$ is either -0.2, 0, or 0.2. Also shown: grid-independence.



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Figure: Elliptical body, M = I = 0.08, $x_0(0) = 0.25$, $\dot{x}_0(0) = 0$.



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Figure: The influence of the gravity parameter g^+ (value in parentheses in legend) on the behaviour of the constant-curvature body. Here, M = I = 0.1, $x_0(0) = 0.7$, $\dot{x}_0(0) = 0$. Results corresponding to values of x_0 greater than unity are not physical, but are still shown.

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