1	A Lagrangian study of interfaces at the edges of cumulus clouds
2	Vishnu Nair*
3	Department of Civil and Environmental Engineering, Imperial College London, London, UK.
4	Thijs Heus
5	Department of Physics, Cleveland State University, USA
6	Maarten van Reeuwijk
7	Department of Civil and Environmental Engineering, Imperial College London, UK

<sup>8</sup> \**Corresponding author*: Vishnu Nair, vs2016@imperial.ac.uk

#### ABSTRACT

Interfaces at the edge of an idealised, non-precipitating, warm cloud are studied using Direct 9 Numerical Simulation (DNS) complemented with a Lagrangian particle tracking routine. Once a 10 shell has formed, four zones can be distinguished: the cloud core, visible shell, invisible shell and 11 the environment. The boundary between the invisible shell and the environment is the Turbulent-12 NonTurbulent Interface (TNTI) which is typically not considered in cloud studies. Three million 13 particles were seeded homogeneously across the domain and properties were recorded along 14 individual trajectories. The results demonstrate that the traditional cloud boundary (separating 15 cloudy and non-cloudy regions using thresholds applied on liquid condensate or updraft velocity) 16 are some distance away from the TNTI. Furthermore, there is no dynamic difference between the 17 traditional liquid-condensate boundary and the region extending to the TNTI. However, particles 18 crossing the TNTI exhibit a sharp jump in enstrophy and a smooth increase in buoyancy. The 19 traditional cloud boundary coincides with the location of minimum buoyancy in the shell. The 20 shell pre-mixes the entraining and detraining air and analysis reveals a highly skewed picture of 21 entrainment and detrainment at the traditional cloud boundary. A preferential entrainment of 22 particles with velocity and specific humidity higher than the mean values in the shell is observed. 23 Large-eddy simulation of a more realistic setup detects an interface with similar properties using the 24 same thresholds as in the DNS, indicating that the DNS results extrapolate beyond their idealised 25 conditions. 26

### **1. Introduction**

Cumulus convection is one of the most important unresolved processes in weather and climate 28 forecasting. The parameterisation of different processes in convective schemes such as the rep-29 resentation of turbulent fluxes is a major source of uncertainty in Numerical Weather Prediction 30 (NWP) models and Global Circulation Models (Bony et al. 2015). Some of the most successful 31 convective parameterisations employ mass-flux schemes (bulk and spectral). In such schemes a 32 cloud model is employed to specify the upward mass flux through the cloud base. These mass 33 flux schemes are mostly of two types: bulk schemes where a single entraining-detraining plume 34 represents the combined effect of an ensemble of active cumuli, or spectral schemes where multiple 35 plumes are considered. Some examples (not an exhaustive list by any means) of the bulk mass flux 36 models are the schemes by Tiedtke (1989), Kain and Fritsch (1990), Bechtold et al. (2001), Gregory 37 and Rowntree (1990), Gerard (2015) and some popular spectral methods are the works of Arakawa 38 and Schubert (1974), Johnson (1976) and Moorthi and Suarez (1992). Two important parameters 39 which modify the mass flux and the vertical transport of temperature, moisture and momentum in 40 these models are the entrainment and detrainment rates. A multitude of parameterisations exist to 41 calculate these rates as summarized in de Rooy et al. (2013). Most of the schemes mentioned so 42 far calculate the change of mass flux with height using entrainment and detrainment rates which 43 have been divided into large scale dynamical rates, and smaller scale turbulent rates. 44

LES has become the primary tool to diagnose entrainment and detrainment rates and it has served the community well over the past years with the majority of the parameterisations arising from LES studies (de Rooy et al. (2013) and references therein). A traditional definition to diagnose

3

<sup>48</sup> bulk entrainment and detrainment rates through LES is (Siebesma 1998),

$$E_{\phi} = -\frac{1}{A\phi_{e}} \oint_{\partial\Omega_{E}} \rho \hat{\mathbf{n}} \cdot (\mathbf{u} - \mathbf{u}_{i})\phi dl, \qquad (1a)$$

$$D_{\phi} = \frac{1}{A\phi_c} \oint_{\partial \Omega_D} \rho \hat{\mathbf{n}} \cdot (\mathbf{u} - \mathbf{u}_i) \phi dl.$$
(1b)

where  $E_{\phi}$  and  $D_{\phi}$  are the entrainment and detrainment rates,  $\phi$  is a conserved scalar and the subscripts *c* and *e* denote averages over the cloud and environment respectively. The path integral is taken horizontally over a cloud area *A* at a constant vertical height. Here the cloud boundary  $\partial \Omega$ has been decomposed into an entraining and detraining region, denoted  $\partial \Omega_E$  and  $\partial \Omega_D$  respectively. The regions are defined by the sign of the net entrainment velocity across the interface  $\hat{\mathbf{n}} \cdot (\mathbf{u} - \mathbf{u}_i)$ , and clearly  $\partial \Omega = \partial \Omega_E \cup \partial \Omega_D$ .

Although the definitions above are rigorous, the entrainment/detrainment rates will depend on 55 the choice of the interface  $\partial \Omega$ . There are multiple interfaces that can be chosen. Most operational 56 parameterisations use the boundary of the cloud core as the interface over which entrainment is 57 calculated. Different definitions of the cloud core interface exist: often these are based on non-58 zero threshold values on liquid water content and vertical velocity (Romps 2010) or thresholds on 59 buoyancy and/or vertical velocity (de Roode and Bretherton 2003). Convective structures can also 60 be identified using decaying passive tracers that are constantly emitted from the surface (Couvreux 61 et al. 2010; Park et al. 2016, 2017). Siebesma and Cuijpers (1995) approximated fractional 62 entrainment and detrainment rates using LES and evaluated the effect of different definitions for 63 the cloudy area. They employ a single bulk active cloudy part and a passive environmental part and 64 define the active cloud domain in different ways: cloud (all grid cells with non-zero liquid water), 65 updraft (non-zero liquid water with positive vertical velocity) and core (non-zero liquid water, 66 positive vertical velocity and positive buoyancy). Their study reveals that the updraft sampling 67 gives the best approximations for the turbulent fluxes of the conserved variables  $\theta_l$  and  $q_t$  (for 68

the temperature and humidity respectively). This is because the updraft sampling includes the 69 negatively buoyant but decelerating part of the updraft rise as well. A different decomposition (as 70 originally proposed by Tiedtke (1989)) into an updraft, a downdraft and an environment was also 71 evaluated to check the role of downdrafts. This did not prove to be better than the updraft sampling 72 but it was stated that this would be relevant in the cases where the mass flux in the downdraft would 73 be comparable to that in the cloud updraft. A more recent study by Gu et al. (2020) has shown that a 74 bulk mass flux parameterization using a simple cloud-environment decomposition does not capture 75 the correct magnitude and vertical transport of turbulent fluxes of heat and humidity. Inclusion of 76 the downdraft significantly improved mass flux parameterizations (especially at the cloud top). 77

Romps (2010) recognizes the importance of defining the interface over which entrainment 78 takes place and defines two categories, active and inactive, whereby a Lagrangian parcel can be 79 considered to have entrained when it flips from inactive to active and vice-versa for detrainment. 80 The atmosphere is divided into these two categories using thresholds for a condensate mixing ratio 81 and the vertical velocity. The 'local' rates calculated in Romps (2010) revealed values twice as 82 high as that in the particle budget calculations in bulk plume schemes. Dawe and Austin (2011) 83 investigated the difference between the bulk (from bulk mass flux schemes) and local rates, and 84 attributed this to the presence of the moist and negatively buoyant subsiding shells around the cloud 85 cores and drier air at the edge of the cloud core. Bulk plume mass flux parameterisations define 86 the properties of entrained air as the horizontal slab average over the environment and that of the 87 detrained air as the slab average over the cloud core as shown in equations (1a) and (1b). However, 88 the presence of the shell means that the air entering and leaving the cloud is effectively pre-mixed. 89 As explained in de Rooy et al. (2013), it is important to note that the two approaches will result in 90 the correct turbulent transport as long as the mixing coefficients  $\epsilon$  and  $\delta$  are diagnosed and applied 91 in the relevant framework correctly. Smaller  $\epsilon$  ( $\delta$ ) values combined with the higher difference 92

<sup>93</sup> between the entraining (detraining) air and the cloud (environment) in the bulk approach should <sup>94</sup> result in the same lateral turbulent flux as in the direct approach. Dawe and Austin (2011) also <sup>95</sup> showed a preferential entrainment of parcels with average humidity and vertical velocity higher <sup>96</sup> than that of the mean shell properties (thus enhancing the flux transfer). This makes it highly <sup>97</sup> relevant to take a closer look at a cloud edge and focus on what effect the choice of interface has <sup>98</sup> on air parcels at the cloud edge. This is the primary motivation for the present work.

A detailed study of lateral mixing at the cloud edge and the dynamics of the shell was performed 99 by Nair et al. (2020) (hereafter NHvR20) using Direct Numerical Simulation (DNS). Cloud edge 100 studies using DNS are performed at moderate Reynolds numbers and consider idealized setups, 101 often using mixing layers as models for the cloud edge. One of the main drawbacks of such 102 mixing layer models is that the shell layer exhausts the cloud layer very rapidly and a very transient 103 negatively buoyant layer is formed which grows at the expense of the cloud layer. In NHvR20, 104 a canonical setup was developed which solves this problem by applying a forcing term over a 105 positively buoyant cloud layer which nudges cloud properties to predefined values. This ensures 106 the presence of an actively growing cloud throughout the simulation and such a numerical setup 107 is well suited to look at cloud edge interfaces. Furthermore the entrainment coefficient calculated 108 in NHvR20 was of the same order of magnitude as those calculated in LES studies (Yeo and 109 Romps 2012). This suggests that DNS can play an important role as a tool to study flow at cloud 110 edges. The DNS results for the buoyancy distribution in NHvR20 revealed the different possible 111 interfaces over which cloud-environment mixing can occur. The formation of the shell gives rise 112 to two interfaces which could possibly be of interest. The interface separating regions with zero 113 and non-zero liquid water lies within the shell and this is the first interface of interest. NHvR20 114 showed that this interface is at the point where the buoyancy in the shell is a minimum. We shall 115 refer to this interface as the Visible Shell Boundary (VSB) since it separates the visible part of the 116

shell (which has non-zero liquid water) from the invisible region of the shell (no liquid water). The 117 second interface of interest is the outer boundary of the shell which separates the turbulent region 118 at the cloud edge from the non-turbulent environment, i.e. the Turbulent-Non Turbulent Interface 119 (TNTI). The TNTI has been extensively studied at the edges of jets, plumes, wakes, mixing layers 120 and boundary layers as reviewed in da Silva et al. (2014). The interface layer is known to include 121 two adjacent layers, the Viscous Super Layer (VSL) where vorticity is introduced through diffusion, 122 and the Turbulent SubLayer (TSL) which matches the vorticity from the turbulent region to the 123 VSL. The TNTI is considered to be a surface with zero thickness between or within these two 124 layers. However, this interface has been rarely considered in most cloud-edge studies. LES studies 125 using a decaying scalar emitted from the surface to identify coherent structures in the convective 126 boundary layer (Couvreux et al. 2010; Park et al. 2017) report areas around the cloud with elevated 127 levels of those scalars, quickly dropping off when moving further away into the environment. It is 128 possible that theses areas are a proxy for the TNTI. In this paper we will show that, as far as this 129 idealised flow is concerned, the TNTI represents the outer edge of the cloud system. 130

The effectiveness of Lagrangian particle tracking in clouds was demonstrated by Heus et al. 131 (2008) who used LES and tracked massless Lagrangian particles that followed the flow to study 132 mixing between clouds and the environment. This study settled the debate surrounding the origin 133 of in-cloud air by tracing back cloud-air parcels, clearly revealing the absence of significant cloud 134 top mixing and that practically all mixing occurs laterally. It is also worth mentioning that in the 135 atmosphere, detrainment layers are often associated with locally increased stratification (de Rooy 136 et al. 2013). Yeo and Romps (2012) used Lagrangian particle tracking on an individual cumulus 137 cloud and calculated a higher rate of entrainment compared to the Eulerian direct measurements. 138 This was attributed to the fast re-circulation of air in and out of the core. It was also shown that 139

almost half of the air entrained by a cloud during its lifetime had been previously detrained. These
 were effects which could not be captured or resolved by Eulerian measurements.

The combination of DNS and a Lagrangian particle tracking routine facilitates a study of mixing 142 and entrainment over the different interfaces at cloud edges in detail. We will highlight the 143 importance of the TNTI and hence the importance of including the shell in cloud edge mixing 144 studies. By following individual particles as they entrain and detrain across the different interfaces, 145 we show that a Lagrangian parcel can be considered to be 'entrained' when it crosses the TNTI 146 and that the shell extends the cloud edge to this interface. We also aim to highlight the degree of 147 pre-mixing done by the shell to the entraining and detraining air parcels. The case setup of the 148 DNS is intentionally designed to avoid the complexity of a real cloud and only focus on the process 149 of lateral mixing and entrainment at the cloud edge. In order to investigate the extent to which the 150 findings are transferable to realistic clouds, we perform Large-Eddy Simulation of a BOMEX case 151 with no mean wind, and explore whether we can detect the TNTI there. Using conditional averages 152 of the enstrophy and mass flux we show that the different interfaces can be distinctly identified and 153 present results highlighting the net mass exchange between the different zones. 154

#### **2.** Case setup and simulation details

#### <sup>156</sup> a. Direct Numerical Simulations

The case is identical to that used in NHvR20 and is shown in figure 1. The domain is divided into a cloud and environmental layer. The cloud layer is moist and positively buoyant and the environment is dry and its buoyancy is defined to be zero. As in Abma et al. (2013), the dominant mixing is assumed to occur locally and hence the influence of the cloud top and base can be neglected which makes the system statistically homogeneous in the vertical direction  $\hat{z}$ . This

allows us to impose periodic boundary conditions on the top and bottom boundaries ( $\hat{z}$  direction) 162 and in the span-wise direction ( $\hat{y}$  direction) if the domain is large enough. Free-slip boundary 163 conditions are imposed along the  $\hat{x}$  direction. In order to prevent the evaporation of the cloud 164 due to turbulent mixing, a forcing is applied over the initial cloud layer from x = 0 to x = 1 m 165 to nudge the values of the vertical velocity w, the temperature  $\theta_l$  and humidity  $q_t$  to pre-defined 166 values  $w_c$ ,  $\theta_{l,c}$  and  $q_{t,c}$ , respectively. Details of the forcing scheme can be found in NHvR20. 167 A negatively buoyant shell forms due to the nudging in the cloud layer, which then develops in 168 a self-similar fashion. Two distinct flow phases were observed in NHvR20 within a negatively 169 buoyant turbulent cloud-environment mixture. The first is a 'drag' phase where the momentum 170 flux transfer (between the cloud core and the shell) dominates and the negatively buoyant shell is 171 dragged vertically upwards by the active cloud layer. The onset of the second 'buoyancy' phase 172 occurs when the negative buoyancy within the shell overcomes the drag and consequently the shell 173 starts descending. Another pertinent finding in NHvR20 is that the shell falls ballistically and the 174 mean velocity inside the shell is dynamically unimportant. 175

The code for direct numerical simulation, SPARKLE, solves the incompressible Navier-Stokes equations under the Boussinesq approximation, and transport equations for scalars to fourth order accuracy. The buoyancy b is given by

$$b = g \left( \frac{\theta - \theta_0}{\theta_0} - \left( 1 - \frac{R_v}{R_d} \right) (q_t - q_{t,0}) - \frac{R_v}{R_d} (q_l - q_{l,0}) \right),$$
(2)

<sup>179</sup> where  $\theta$  is the potential temperature,  $q_l$  is the liquid water specific humidity,  $\theta_0$ ,  $q_{l,0}$  and  $q_{t,0}$  are the <sup>180</sup> environmental values of the potential temperature and total water specific humidity respectively, <sup>181</sup>  $R_d = 287.0 \text{ J kg}^{-1} \text{ K}^{-1}$  and  $R_v = 461.5 \text{ J kg}^{-1} \text{ K}^{-1}$  are the gas constants for dry air and water vapour <sup>182</sup> respectively. The value of the liquid water specific humidity in the environmental layer,  $q_{l,0}$ , is <sup>183</sup> set to zero. A bulk condensation scheme developed by Sommeria and Deardorff (1976) is used to diagnostically calculate  $q_l$  in the cloud layer. Further details of the numerical method used in SPARKLE can be found in Craske and van Reeuwijk (2015).

The simulation is performed on a domain of 30m x 15m x 15m using a grid of 1536 x 768 x 768 186 points. The initial profiles are the same as simulation A03 in NHvR20 and the simulation is run 187 for 80 s. The liquid water specific humidity and vertical velocity in the cloud is,  $q_{l,c} = 3 \text{ g kg}^{-1}$  and 188  $w_c = 0.81 \text{ m s}^{-1}$ , respectively. The resulting cloud buoyancy is  $b_c = 0.046 \text{ m s}^{-2}$ . A top-hat profile 189 is implemented for  $\theta_l$  and  $q_t$  with  $\Delta \theta_l = -5.9$  K and  $\Delta q_t = 5.5$  g kg<sup>-1</sup>. The kinematic viscosity v 190 is  $4 \times 10^{-4}$  m<sup>2</sup> s<sup>-1</sup>. The Taylor Reynolds number ( $Re_{\lambda}$ ) at the end of the simulation is 91 and the 191 grid resolution  $r = \frac{\Delta x}{\eta} = 1.3$  where  $\eta$  is the Kolmogorov length scale. The integral shell time scale 192 for the simulation  $\tau = l_e/u'$  is 14.4 s and the Kolmogorov time scale  $t_\eta = (v/\varepsilon)^{0.5}$  is 0.45 s (where 193 the integral length scale  $l_e = u'^3/\varepsilon$ ). Here u' is the rms velocity and  $\varepsilon$  is the dissipation rate. It is 194 important to mention here that  $\varepsilon$  is calculated as an average over the width of the shell and not the 195 entire cloud domain. Hence the time scales  $\tau$  and  $t_{\eta}$  are representative of the characteristic time 196 scales of the largest and smallest eddies inside the shell. 197

A Lagrangian particle tracking routine has been implemented to solve the equation of motion for massless particles

$$\frac{\mathrm{d}\mathbf{x}}{\mathrm{d}t} = \mathbf{u}(\mathbf{x}, t). \tag{3}$$

Here,  $\mathbf{u} = (u, v, w)$  is the velocity vector with u, v and w as the horizontal, transverse and vertical components respectively, and  $\mathbf{x} = (x, y, z)$  is the particle position vector. A tricubic Hermite interpolation scheme is used to calculate the particle velocities from the Eulerian flow velocity fields. A third order Adams-Bashforth time-stepping scheme is used to integrate the particle locations at each time step. SPARKLE uses a two-dimensional domain decomposition <sup>205</sup> and is parallelized using MPI. The particles on processor boundaries are communicated across <sup>206</sup> processors after each time integration step following the flow field (Perrin and Jonker 2015).

A total of 3 million particles are seeded into the numerical domain after the flow has reached 207 a self-similar state at  $t_0 = 68$  s. Importantly, by this stage, the flow is in the buoyancy regime 208 and there is a descending shell present in the domain (at earlier times, the shell is only visible 209 in the turbulence but the cloud boundary is still dragged up due to the turbulence; see NHvR20). 210 Figure 2 shows the mean buoyancy and vertical velocity profiles at the time of seeding. The 211 profiles are normalized by their respective minimum values  $b^*$  and  $w^*$ . The particles are seeded 212 homogeneously across the x, y and z axes within a volume bounded by [0, 9 m], [0, 15 m], and [0, 10 m], [213 30 m] respectively. Since the outer boundary of the shell is at approximately 7.5 m at the end of 214 the simulation, the particles are seeded only up to a distance of 9 m across the x axis. The region 215 from 9 m to the boundary at 15 m is the quiescent environment and the results are not affected by 216 not seeding within this region. Data for the particles is written every 0.2 s. 217

Since the flow in the shell has been shown to be self-similar in NHvR20, we expect the results to hold for different initial values/parameters. The sensitivity of the results is analysed in appendix B by performing a simulation with different initial parameters (simulation A06 in NHvR20), and the plots reveal that the particle behaviour are independent of the initial flow parameters.

222

#### 223 b. Large Eddy Simulations

To explore the feasibility of an interface in more realistic set ups, we use the MicroHH LES (van Heerwaarden et al. 2017) to run an LES of shallow cumulus convection based on the BOMEX case (Siebesma et al. 2003) at an isotropic 4.1 m grid spacing, using  $800 \times 800 \times 800$  grid points and a 3.2 km horizontal domain size. Note that this grid spacing is much finer than the 25 m that Heus and Jonker (2008) found was sufficient for a converging mass flux in the subsiding shell. With this
resolution, the internal dynamics of the shell are better resolved and the enstrophy approaches a
meaningful value. We use a simulation time of 10 h, of which the first 3 h are discarded as spin-up.
In order to separate shear driven turbulence from convective turbulence, we set the geostrophic
wind to zero. As it turns out, shear does not seem to make a major difference to our results, other
than enlarging the shell size.

#### **3. Interfaces at the cloud edge**

As mentioned in the introduction, multiple interfaces exist at the cloud edge over which we can 235 consider entrainment and detrainment. In this section we clearly define and visualize the two 236 interfaces that exist due to the presence of the shell, i.e. the Visible Shell Boundary (VSB) and 237 the TNTI. We also include the Cloud Core Boundary (CCB) which is commonly used in most 238 parameterisation studies. The CCB and the VSB are defined by applying thresholds on the buoyancy 239 b and the liquid water specific humidity  $q_l$ . The TNTI is defined by applying a threshold on the 240 enstrophy  $\omega^2$ . The enstrophy is a scalar quantity which is defined as  $\omega^2 = \boldsymbol{\omega} \cdot \boldsymbol{\omega}$ , where  $\boldsymbol{\omega} = \nabla \times \mathbf{u}$ 241 is the vorticity. We apply a threshold value of  $q_{l,th} = 10^{-5} \text{ kg kg}^{-1}$  and  $\omega_{th}^2 = 10^{-6} \text{ s}^{-2}$ . An analysis 242 on the sensitivity of the TNTI location to the choice of enstrophy threshold value is performed in 243 appendix A. The three interfaces divide the domain into four different zones which can be defined 244 as: 245

- 1. Cloud core (CC), where  $b > 0, q_l \ge q_{l,th}$ ;
- 247 2. Visible Shell (VS), where  $b < 0, q_l \ge q_{l,th}$ ;
- <sup>248</sup> 3. Invisible Shell (IS), where b < 0,  $q_l < q_{l,th}$ ,  $\omega^2 \ge \omega_{th}^2$ ;
- 4. Environment (E), where b = 0 and  $\omega^2 < \omega_{th}^2$ .

These zones are shown in figure 3, depicting the CC (red), VS (blue) IS (yellow) and the E (white). The CCB is defined on the interface between the CC and the VS, the VSB on the interface between the VS and the IS and the TNTI on the interface between the IS and E. The snapshot represents the flow at  $t_0$ . For reference, the boundary considered by Romps (2010) would be closer to the CCB than the VSB due to the high threshold applied to the vertical velocity in that study.

Interestingly, the VSB coincides with the location where the buoyancy is minimum. Evaporation of cloud liquid water due to mixing and entrainment results in latent heat absorption. This evaporative cooling is maximum at the point where all the liquid water has evaporated, i.e. at the VSB. This can be explained using equation (2). When  $q_l$  drops to zero due to complete evaporation, the consequent temperature drop results in a negative value for  $(\theta - \theta_0)/\theta_0$  with the buoyancy falling to a minimum (NHvR20).

In the rest of the manuscript, all references to the 'shell' indicate a union/combination of both the IS and the VS.

#### **4. Lagrangian particles**

Lagrangian particles are seeded uniformly across the four zones at  $t = t_0$  over the interval 0 < x < 9264 m, 0 < y < 15 m, 0 < z < 30 m. A total of 3 million particles were divided among the different 265 zones with 578,964 in the CC, 502,801 in the VS, 347,348 in the IS and 1,572,052 in the E. Since 266 the particles are passive, they follow the flow. A snapshot of the particles at a normalized time 267  $t^* = (t - t_0)/\tau = 0.67$  is shown in figure 4. The particles are colored according to the zones in 268 which they are initially seeded at  $t_0$ : CC (red), VS (blue), IS (yellow) and E (pink). This figure 269 provides a qualitative idea of the source and destination of entraining and detraining particles. The 270 CC contains particles from the VS and IS, there are indications of CC particles in the IS and even 271 indications of E particles in the VS. 272

Figure 5 quantifies the movement of the particles between the zones by showing an origin-273 destination matrix for the initial and final zones of all particles that leave their zone of origin. Each 274 square in the matrix represents a particular zone. The rows represent the zone of origin and the 275 columns represent the destination zone at  $t^* = 0.67$ . The colors (and the numbers) indicate the 276 percentage of particles that originated in the zone represented by the row and has a final destination 277 in the zone represented by the column. The percentages are calculated over the total number of 278 particles (in all four zones) at  $t^*$ . Hence the sum of values in all the cells in the matrix is equal 279 to 100. It is necessary to exercise caution when generating such a matrix since the percentages 280 will be dependent on the initial seeding density and the domain size. For instance, an analysis 281 of the particles that started and finished in their zones of origin revealed that a high percentage 282 simply do not make any crossing at all. In the time interval from  $t^* = 0$  to 0.67, the following 283 percentage (number) of particles remain in their zones of origin without making any crossing: CC 284 -89% (385,146), VS -73% (221,989), IS -90% (163,162) and E -99.9% (1,380,563). Hence all 285 particles that remain in their zones of origin are excluded to give a better understanding of mixing 286 and crossing between different zones. We still have a significant number of particles to obtain 287 reliable statistics for figure 5 : CC - 193,818, VS - 280,812, IS - 184,186 and E - 191,489. 288 Particles starting in the CC detrain into the VS (16.4%) and very rarely make it to the IS in the 289 time interval considered (0.4%). Particles which entrain back to the CC (5.9%) were originally 290 detrained. Particles originating in the VS cross over to the both the CC and IS, of which 9.7% 291 entrain back. Entrainment into the CC (17.7%) is much more dominant than detrainment into the 292 IS (5.8%). The CC-VS mixing is nearly symmetrical, showing nearly equal transport of particles 293 between the two layers. Conversely, VS - IS mixing is highly skewed with particles originating in 294 the IS showing a very high preference to entrain into the VS (16.3%) compared to particles starting 295 from the VS detraining into the IS. 296

Analysis of the particles that originate and finish in the IS (2.3%) reveals that these are made 297 up almost entirely (99%) of particles that entrain and detrain back from the VS. A negligible 298 percentage of particles detrain into the environment and entrain back. The almost one-sided 299 crossing of particles at the TNTI is also revealed by the high percentage of particles (21.4%)300 that originate from the environment into the IS and the negligible percentage moving in the 301 opposite direction. Since the shell thickness is increasing linearly (NHvR20), we can consider the 302 TNTI as a moving interface that would uniformly entrain particles resident in the environment, 303 i.e.  $\Delta N_e = \rho_p A \Delta h$ . Here  $\Delta N_e$  is the number of particles entrained when the interface moves a 304 horizontal distance  $\Delta h$  covering an area A, and  $\rho_p$  is the particle density. Therefore the number of 305 particles entrained over a time step dt is given by  $dN_e/dt = \rho_p A dh/dt$ . A comparison between the 306 number of particles entraining across the TNTI calculated directly from the DNS data, and from 307 the model show a very good agreement with less than 3% error between (not shown). A moving 308 interface results in a constant number of particles entrained from the environment in our setup with 309 negligible detrainment across the TNTI. This explains the skewness in crossings at the TNTI. Yeo 310 and Romps (2012) have shown that a high percentage of air entrained at the cloud boundary had 311 previously been detrained from the cloud. However, entrainment at the TNTI in our setup is almost 312 entirely from the environment. A similar argument can be used to explain the skewness at the VSB. 313 Since the VSB is moving as a function of time as the visible shell thickens, the entrainment from 314 the IS to the VS dominates the detrainment from VS to IS. 315

Another striking result from figure 5 is that a negligible percentage of particles travel all the way from the CC to the environment and vice-versa. This will be addressed further in section 6 where we investigate the degree of premixing done by the shell and how it acts as a buffer layer between the CC and the environment.

A better perspective of the path followed by the particles can be obtained by looking at the 320 average times taken by particles to travel across different zones. This quantity is determined by 321 calculating how long a particle resides in a zone after entering it via an interface. The residence 322 times normalized by the shell time scale  $\tau$  are shown in figure 6. Calculating the average time 323 after seeding to cross the nearest interface from each zone is quite deceptive since this is highly 324 dependant on the seeding density. We will hence avoid this calculation and focus on residence 325 times after entrainment or detrainment across an interface. For particles originating in the CC, the 326 nearest interface will be the VSB, whilst for those originating in the VS, the CCB and the VSB are 327 the nearest interfaces. For particles from the IS, the VSB and the TNTI are the nearest interfaces 328 and for those from the environment, the TNTI is the sole interface of interest. 329

<sup>330</sup> Particles originating in the CC have a mean residence time of  $0.44\tau$  in the VS after crossing <sup>331</sup> the CCB and before they detrain into the IS. However it should be noted that the mean times are <sup>332</sup> dependent on the time duration over which the particle data is collected (12 s in this simulation). <sup>333</sup> Particles residing for a longer period (greater than 12 s) are not included, which could result in a <sup>334</sup> higher magnitude of the mean residence time. Since very few particles in the IS cross the TNTI, <sup>335</sup> the residence time in the IS is not calculated.

<sup>336</sup> A majority of the particles originating from the IS cross the VSB into the VS, where they reside <sup>337</sup> for  $0.29\tau$  before crossing into the CC. Particles originating in the environment cross the TNTI and <sup>338</sup> remain in the IS for  $0.46\tau$ . A small percentage (5%) cross over to the VS and reside there for 0.55 <sup>339</sup>  $\tau$ .

#### **5. Entrainment at the cloud edge**

In this section we analyse the time histories of particles that cross the three interfaces. This will reveal particle behaviour before, at and after crossing each interface, and is information that

is impossible to obtain with Eulerian statistics. For this analysis, we consider each interface 343 individually and look at all particle crossings across it. A particle is considered to have entrained 344 when it moves to a different zone from right to left (towards the CC) and to have detrained when 345 it moves to a different zone from left to right (towards the environment). The time variable  $t_e$ 346 and  $t_d$  represent the time at which a particle entrains or detrains across an interface respectively. 347 In case of multiple entrainment events by the same particle across a particular interface, the time 348 of occurrence of the first entrainment event is taken as the entrainment time  $t_e$ . For detrainment 349 we follow a slightly different procedure. A detrainment event is considered to take place only if 350 a particle moves to a different zone and resides there, i.e. detraining particles that are entrained 351 back are not considered in the analysis while particles that detrain-entrain-detrain are counted. In 352 short, for a particle to be considered in the detrainment analysis, the last crossing between zones 353 should have been from right to left. The time of the last crossing from right to left is taken as the 354 detrainment time  $t_d$ . We adopted such an approach since entraining particles show a tendency to 355 detrain after one or two time steps immediately after entrainment, before entraining again at the 356 next time step. Counting such short term detrainment events contaminates the plot and are hence 357 these are not considered in the detrainment plot. 358

Figure 7 shows box plots of the enstrophy  $\omega^2$ , and buoyancy b of all entraining particles. In 359 box plots, groups of particle data are expressed using their quartiles. The tops and bottoms of 360 the box are the first and third quartiles (or the 25th and 75th percentiles), and the red line is the 361 median of the particle data group. The x-axis represents the time before and after a crossing and 362 is normalized with the shell time scale  $\tau$ . The whiskers (showing variability from the first and 363 third quartiles) and outlier points for the box plot have been removed from the figure for the sake 364 of clarity. A sharp jump in the enstrophy values is observed as particles cross the TNTI as shown 365 in figure 7(a). This is consistent with vorticity jumps that are observed across the TNTI in several 366

flows (da Silva et al. 2014). Once entrained into the IS, the particle acquires enstrophy rapidly. The buoyancy shows a very smooth increase for particles entraining across the TNTI as shown in figure 7(b). As shown in section 2, the outer boundary of the (invisible) shell where the buoyancy drops to zero coincides with the TNTI. The value of buoyancy is zero at the environment and since almost all the particles crossing the TNTI originate from the environment (as was also shown in section 4) there is a smooth decrease in the buoyancy as particles cross into the IS.

No significant change in the particle enstrophy is observed as they entrain into the VS through 373 the VSB (figure 7(c)). The VSB resides deep in the shell where turbulence intensities are high, 374 and the particles are not subjected to sharp gradients across this interface. The enstrophy values 375 before entrainment predominantly indicate particles originating from the IS and almost none from 376 the environment. In sharp contrast to the enstrophy, there is a kink in the buoyancy as the particles 377 entrain into the VS. This is due to the VSB coinciding with the buoyancy minimum. The majority 378 of the particles crossing into the VS are from the IS as seen by the negative values of buoyancy 379 before entrainment. This is also consistent with the origin-destination matrix in figure 5. Positive 380 values for the buoyancy can be observed for  $(t - t_e)/\tau > 0.6$  which indicates that particles cross 381 over into the CC. 382

Particles crossing the CCB show a similar behaviour as those crossing the VSB with the enstrophy remaining almost the same across the interface. The buoyancy shows a kink at the point where the particles cross over from the negatively buoyant VS into the positively buoyant CC. The positive and negative values for the buoyancy before entrainment indicate that entraining particles include those which originate from the VS as well as those which have been recently detrained from the core. At about  $0.3\tau$  before entrainment, the buoyancy is almost entirely negative indicating only particles from the VS. The finite jump in enstrophy values at the TNTI combined with the relatively flat behaviour across the VSB and the CCB highlights the relevance of the TNTI as the entraining cloud interface. The VSB and the CCB can be considered to be a cross-over point where particle buoyancy decreases significantly (to a minimum at the VSB and zero at the CCB) as a result of evaporative cooling.

Detrainment plots show similar behaviours. We only show plots for the VSB and the CCB. This 394 is because we observed very few qualifying detrainment events at the TNTI compared to the VSB 395 and the CCB as shown in section 4. This is an indication of the pre-mixing done by the shell. 396 Particles originating from the CC and detraining all the way to the environment is an extremely 397 rare event within the time scale studied  $(0.67\tau)$ . Similar to the entrainment plots, there are no finite 398 jumps in enstrophy during detrainment across the VSB and the CCB. Across the VSB, the majority 399 of detraining particles are from the VS (entirely negative buoyancy at start of the plot). But across 400 the CCB, there are particles that originate from both the CC and the VS. The detraining particles 401 that originate from the CC are probably those at the edge of the CC which mix with the negatively 402 buoyant VS as a result of which they lose their positive buoyancy and is detrained from the CC. A 403 closer look at the properties of entraining and detraining particles is taken in the next section. 404

#### **6. Shell pre-mixing and preferential entrainment at interfaces**

In this section we look at how effective the shell is in premixing the entraining and detraining parcels. All particles crossing the CCB, VSB and the TNTI at normalized time  $t^* = 0.67$  are considered irrespective of their origin. As mentioned in the introduction, Dawe and Austin (2011) attributed the difference between the bulk and local entrainment rates to be essentially due to the presence of the subsiding shell around the CC. This means that the properties of entraining and detraining air can no longer be the same as the mean values in the environment and the CC respectively. This is explored in figure 9 which shows histograms of the total water specific <sup>413</sup> humidity  $q_t$  (a, c, e) and vertical velocity w (b, d, f) for all particles that have entrained and <sup>414</sup> detrained across the interfaces. Figure 9 shows entrainment and detrainment across the CCB (a,b), <sup>415</sup> the VSB (c,d), and the TNTI (e,f).

The histograms in figure 9 (a) and (b) are clearly skewed and suggest that the properties (vertical 416 velocity w and total water specific humidity  $q_t$ ) of the entraining and detraining particles do 417 not have a mean value equal to the horizontal slab average values of the environment and CC 418 respectively. The particles entraining (cyan) across the CCB have a mean value (cyan dashed line) 419 of w and  $q_t$  closer to those at the inner edge of the VS. The mean of w and  $q_t$  is also higher than 420 the mean for all particles in the VS (cyan dotted line), i.e. there is clear evidence for preferential 421 entrainment of particles which have a  $q_t$  and w higher than the mean values in the VS. Dawe and 422 Austin (2011) explained this preferential entrainment by the presence of negatively buoyant regions 423 which still had positive vertical velocity and condensed liquid water. As these parcels rise, there is 424 latent heat release due to further condensation, thus making the parcels positively buoyant which 425 leads to them being entrained into the CC. This is true in the current simulation as well. In the 426 buoyancy phase, even though the mean value of vertical velocity in the shell is negative, there still 427 are upward moving negatively buoyant parcels with liquid water (especially in the VS). We would 428 not expect to see this effect to happen at the VSB and this is indeed true as seen in figure 9 (c) and 429 (d). However, the mean values of  $q_t$  (cyan dashed line) and  $\theta_l$  (not shown) of particles crossing 430 into the VS coincides with the mean saturation value  $\overline{q_s}$  and  $\overline{\theta_s}$ .<sup>1</sup> This again highlights the degree 431 to which the shell premixes the entraining air. Entrainment across the TNTI is shown in figure 432 9 (e) and (f) and involves particles only from the environment as was shown in section 4. The 433

<sup>&</sup>lt;sup>1</sup>Considering a classical  $q_t - \theta_l$  mixing diagram and assuming linear mixing between a saturated cloud and unsaturated environment parcel, the intermediate thermodynamic states of the cloud environment mixture can be assumed to lie on a straight line connecting the two initial states. This means that the mean properties of the saturation mixture,  $\overline{q_s}$  and  $\overline{\theta}_s$ , are the coordinates of the point where the mixing line crosses the saturation curve (fig 2(a) in NHvR20)

mean value of  $q_t$  of particles entraining into the IS coincides with the mean of the Environment 434 as expected. Entraining particles also have a mean vertical velocity that is slightly negative. A 435 possible reason for this could be the rapid entrainment of particles that were detrained from the IS 436 and have negative velocities. Figure 9 also shows the properties of particles detraining across the 437 CCB and the VSB (red). Figure 9(a) clearly shows that it is the drier air at the edges of the CC 438 that is detraining across the CCB. For the CC,  $\overline{q_t} = 12.5 \text{ g kg}^{-1}$  (red dotted line) while the mean 439 of the detraining particles is significantly lower at around 11.1 g kg<sup>-1</sup> (red dashed line). Another 440 interesting observation is that the mean vertical velocity of the detraining parcels is positive as 441 shown in figure 9(b). These could be the parcels that are negatively buoyant at the edges of the 442 CC but still moving up due to the positive vertical momentum of the fluid in the CC. The vertical 443 velocity histogram also shows a very similar behaviour to the entrainment. Detrainment across the 444 VSB (figure 9 c and d) shows a very small sample range for  $q_t$ . Very few particles detrain into the 445 IS which is consistent with the percentages seen in the matrix in figure 5. The mean of the vertical 446 velocity of the detraining particles is negative and is very close to that of the IS (blue dotted line). 447 Detrainment across the TNTI is almost entirely negligible with very few particles crossing from 448 the IS. 449

#### **7. Presence of the interface in LES**

The DNS setup is highly idealized. In order to verify whether the results for the DNS extrapolate to situations in which the full cloud life cycle is represented, we ran a high resolution LES of shear-free BOMEX case (Holland and Rasmusson 1973) at a resolution of 4.1 m. Figure 10 shows a contour plot of the horizontal cross-section of enstrophy at a mid-cloud layer (z = 1000 m). A visual inspection suggests that most of the IS is indeed localized, and usually centered around the clouds (denoted in black contours). Typical values of enstrophy drop from  $10^{-1}$  in the clouds to <sup>457</sup> below 10<sup>-3</sup> in the environment. Figure 11 shows the mass flux density (a) and enstrophy (b) plotted <sup>458</sup> as a function of the distance to the nearest cloud edge. The plots clearly show that the shell itself <sup>459</sup> is confined to within less than 200m from the nearest cloud edge. The additional empty IS areas at <sup>460</sup> plot locations (600m, 2200m) and (1300m,1500m), are likely related to a recently-dissipated cloud <sup>461</sup> and do not necessarily contribute to the negative mass flux. Clouds will also generate internal <sup>462</sup> waves which in turn produce vorticity in the environment (Fodor and Mellado 2020). However, <sup>463</sup> detangling both is subtle and is beyond the scope of this work.

Figure 12 shows the conditionally sampled enstrophy over the cloud layer, with each layer defined 464 similar to the definitions in section 3, but with the adjusted thresholds  $q_{l,th} = 0$  and  $\omega_{th}^2 = 10^{-3} \text{ s}^{-2}$ . 465 For these and the consecutive graphs, we have used the LES output between the 3rd and 10th hour, 466 with a sample time of 0.5hr. The IS is defined with  $\omega_{th}^2 = 10^{-3} \text{ s}^{-2}$  and with no threshold on the 467 buoyancy. Throughout the cloud layer, the enstrophy in the cloud and its immediate surroundings 468 is an order of magnitude higher than the enstrophy in the environment. This enstrophy jump across 469 the two zones shows that similar to the TNTI observed in the DNS, we can observe a clear interface 470 between the IS and the environment as well. Because the ambient fluid is still turbulent, this is 471 now actually a Turbulent-Turbulent Interface (TTI). According to Kankanwadi and Buxton (2020), 472 the adjustment in enstrophy across the TTI is analogous to the TNTI. 473

The choice of the enstrophy threshold  $\omega_{th}^2 = 10^{-3} \text{ s}^{-2}$  can be justified by looking at the cumulative mass flux for all non-cloudy grid cells plotted against the enstrophy as shown in figure 13. Starting from zero at high enstrophy (clouds), there is a small bump (perhaps evaporating cloud tops), after which a steep decrease in mass flux and an inflection point at  $\omega^2 = 10^{-3} \text{ s}^{-2}$  is seen. In other words, the mass flux per unit enstrophy is the highest in the IS which is centered around the clouds (see figure 12). Finally, figure 14 shows the conditionally sampled mass flux profile. In the middle of the cloud layer, far away from the lifting condensation level and the level of neutral buoyancy, the

IS is responsible for the majority of the negative mass flux which remains almost constant from 481 the cloud base at 0.5 km up to 1.1 km. Highly negative values of mass flux in the environment at 482 around 1.2 km to 1.5 km are observed. One possible reason could be that the cloud top generates 483 internal waves with a net negative mass flux because the initial upward motion is still part of the 484 high enstrophy region. However, between the height range of 0.75 km to 1.1 km, the mass flux in 485 the environment is close to zero. Overall, the results suggest that, in the mid layer of the cloud (for 486 the conditions considered in this simulation), the upwards mass flux in the cloud core is balanced 487 by the negative mass flux in the IS with no lateral mass exchange at the TTI, or, the net exchange 488 rate across the TTI is zero. This is consistent with the results in Jonker et al. (2008) where the 489 mass flux in the shell (within 200 m of the cloud edge) was shown to compensate for about 80% of 490 the in-cloud mass flux. 491

<sup>492</sup> Our DNS study can be considered to correspond to the LES results between 0.5 km and 0.8 km. In <sup>493</sup> this region, the gradient of the mass flux in the environment shows a negative slope which represents <sup>494</sup> air being entrained from the environment. The DNS study reveals a similar net entrainment at the <sup>495</sup> TNTI as mentioned in section 4.

#### **8.** Discussion and Concluding remarks

<sup>497</sup> A numerical study of the different interfaces at the edge of a cumulus cloud was performed using <sup>498</sup> DNS and LES. The DNS study reveals the presence of four distinct zones which can be detected <sup>499</sup> by applying thresholds on the enstrophy  $\omega^2$  and the specific humidity of liquid water  $q_l$ . The <sup>500</sup> four zones have been defined as the cloud core, a visible and invisible shell and the environment <sup>501</sup> layer. The different zones give rise to three distinct interfaces: the cloud core boundary, visible <sup>502</sup> shell boundary (traditionally considered the cloud edge in parameterization studies) and a turbulent non-turbulent interface. Massless Lagrangian particles were introduced at the cloud edge and their
 trajectories and properties were tracked.

One of the main findings of this work is the detection of an interface (in both the DNS and LES 505 studies) between the invisible shell and the environment, separating regions of different turbulence 506 intensities and across which a finite jump in enstrophy is observed. This layer extends beyond 507 the traditional cloud boundary defined by the liquid water specific humidity. For the DNS, the 508 environment is quiescent and the interface is representative of the classical Turbulent-Non Turbulent 509 Interface (da Silva et al. 2014). The LES studies, which are substantially more complex and do not 510 have a quiescent environment, reveal an interface separating zones with distinctly different levels of 511 enstrophy, and are hence more appropriately described by a turbulent-turbulent interface, with the 512 majority of the downward mass flux contained within this interface. Pertinently, the DNS results 513 indicate there is no dynamic distinction between the visible shell and the invisible shell, indicating 514 the two are part of the same system. It remains to be seen whether the IS plays a dynamical role 515 in the evolution of cloud boundaries, but the current study clearly indicates that the cloud edge 516 extends beyond the visible shell. 517

The DNS results reveal that Lagrangian particles experiences finite jumps in its enstrophy and a smooth increase in buoyancy when it crosses the TNTI. The traditional cloud boundary considered in LES entrainment studies, the visible shell boundary, is essentially a cross-over station which does not significantly affect entraining or detraining parcels. It coincides with the location of minimum buoyancy within the shell and parcels entraining across this boundary have saturation values that can be predicted from the mixing diagram.

<sup>524</sup> We also observe preferential entrainment across the CCB as shown by Dawe and Austin (2011), <sup>525</sup> where particles with total humidity  $q_t$  and vertical velocity w higher than the mean of all parcels <sup>526</sup> in the shell more likely to entrain. Also, drier air that is close to the CCB is more likely to detrain.

24

This drier air is also likely to be negatively buoyant but moving with a positive vertical velocity in the cloud core.

The DNS study is performed using a highly idealized setup. The simplifications include an 529 infinitely long cloud interface without stratification of the environment and no other sources of 530 turbulence such as wind shear. It is also important to highlight the fact that by imposing a periodic 531 boundary condition on a domain size of 30 m, larger eddies in the cloud. However, the LES 532 simulations are performed using a more realistic setup and was able to detect a TTI using the same 533 thresholds as in the DNS which shows that it is possible to extend the notion of a TNTI to turbulent 534 environments. Indeed, da Silva et al. (2014) describes the interface layer as a thin region with a 535 finite thickness that separates either (a) regions with different turbulent intensity or (b) turbulent 536 and (external) irrotational flow regions. Conditionally averaged mass flux in the different zones in 537 the LES results reveal that within the mid cloud layer, there is no net exchange across the TTI. The 538 upward mass flux in the cloud core is compensated by the mass flux in the IS with negligible mass 539 flux in the environment. 540

<sup>541</sup> While we do not observe any detrainment across the TNTI, this does not mean that no air can be <sup>542</sup> transported from the cloud into the environment. The cloud will eventually dissipate, and so will <sup>543</sup> the IS. During this process, the humid air of the cloud will moisten the environment. However, this <sup>544</sup> is an entirely different mechanism than any direct mixing between the cloud and the environment, <sup>545</sup> with different results. For instance, if detrained air remains close to the cloud boundary, subsequent <sup>546</sup> entrainment events will do less to dilute the cloud, resulting in stronger updrafts. We also speculate <sup>547</sup> that a slowly dissipating IS may allow for preconditioning certain regions for subsequent convection; <sup>548</sup> further research is needed to confirm or falsify such speculation.

25

Acknowledgments. Vishnu Nair and Maarten van Reeuwijk acknowledge funding from the Marie Sklodowska Curie Actions under the European Union's Horizon 2020 research and innovation
 programme (Grant no 675675). Computational resources on the UK super-computing facility
 ARCHER via the UK Turbulence Consortium (EP/R029326/1) and the Imperial College HPC
 services are gratefully acknowledged. Thijs Heus was supported by the U.S. Department of Energy's Atmospheric System Research, an Office of Science, Office of Biological and Environmental
 Research program, under Grant DE-SC0017999.

Data availability statement. All data generated in this study will be uploaded to the OpenAc cess database of the Marie-Sklodowska Curie Action COMPLETE with unrestricted access from
 September 2020.

#### APPENDIX A

559

#### 560

#### Sensitivity to threshold values

<sup>561</sup> A very popular method used to detect a TNTI is by applying a threshold on enstrophy which <sup>562</sup> allows one to separate the turbulent region from the approximately irrotational, non-turbulent <sup>563</sup> region (da Silva et al. 2014). This value can be selected from a range since statistics have been <sup>564</sup> shown to be insensitive to the exact threshold value (Bisset et al. 2002). We perform an analysis <sup>565</sup> to verify if the location of the TNTI is indeed insensitive to the choice of threshold values in our <sup>566</sup> case setup.

<sup>567</sup> We check a range of magnitudes for  $\omega_{th}^2$  from  $10^{-3}$  to  $10^{-6}$  s<sup>-2</sup>. In figure A1, the interfaces <sup>568</sup> obtained by applying the different thresholds are plotted. The overlapping interfaces obtained <sup>569</sup> from this range of threshold values point to the fact that these ranges lie in the viscous superlayer observed in interfacial layers and any threshold magnitude within this range can be used as a robust technique to detect the TNTI. Hence, we settle on a value of  $\omega_{th}^2 = 10^{-6} s^{-2}$  for this study.

Furthermore, the TNTI can be seen to coincide exactly with the outer boundary of the shell as determined from the isoline of the buoyancy field b = 0. The inner interface corresponding to the isoline b = 0 coincides with the CCB.

# APPENDIX B

576

575

### Sensitivity to initial conditions

The sensitivity to the DNS initial conditions for the different results presented in the paper are analysed by running a second simulation with different initial parameters. The initial profiles are similar to simulation A06 in NHvR20 with  $\Delta \theta_l = 1.8K$ , and  $\Delta q_t = 1.9gkg^{-1}$ . A similar domain size and grid resolution is used. The simulation is run for 148s and 3 million particles are seeded after 120s. The Taylor Reynold's number ( $Re_\lambda$ ) at the end of the simulation is 57. The integral time scale  $\tau$  is 24.3s and  $t_n$  is 1.04s.

<sup>583</sup> The origin-destination matrix is plotted in figure B1. All the results shown for the main simulation <sup>584</sup> hold for this case as well: symmetric mixing between CC and VS, highly skewed mixing between <sup>585</sup> VS and IS and almost negligible number of particles originating from the environment and from <sup>586</sup> the CC making it all the way to the CC and environment respectively in the time interval of approx <sup>587</sup>  $1\tau$ .

Figure B2 shows the behaviour of entraining and detraining particles at the different interfaces. Entraining particles exhibit a finite jump in the enstrophy and a smooth increase in buoyancy after crossing the TNTI (figure B2(a,b)). Both entraining and detraining particles do not show a noticeable increase in enstrophy while crossing the VSB (figures B2(c) and (e) respectively), with a kink observed in the buoyancy values (figures B2(d,f)). <sup>593</sup> Figures B1, B2 and B3 hence confirm that all the results proposed in the manuscript hold for the <sup>594</sup> new simulation as well and the results are insensitive to the DNS initial conditions.

#### 595 **References**

- Abma, D., T. Heus, and J. Mellado, 2013: Direct Numerical Simulation of evaporative cooling at the lateral boundary of shallow cumulus clouds. *J. Atmos. Sci.*, **70**, doi:10.1175/JAS-D-12-0230.1.
- Arakawa, A., and W. H. Schubert, 1974: Interaction of a cumulus cloud ensemble with the large-scale environment, part i. *J. Atmos. Sci.*, **31** (**3**), 674–701.
- Bechtold, P., E. Bazile, F. Guichard, P. Mascart, and E. Richard, 2001: A mass-flux convection
   scheme for regional and global models. *Quart. J. Roy. Meteor. Soc.*, **127** (**573**), 869–886, doi:
   10.1002/qj.49712757309.
- <sup>603</sup> Bisset, D. K., J. C. R. Hunt, and M. M. Rogers, 2002: The turbulent/non-turbulent interface <sup>604</sup> bounding a far wake. *J. Fluid Mech.*, **451**, 383–410, doi:10.1017/S0022112001006759.
- Bony, S., and Coauthors, 2015: The role of convective-scale precipitation downdrafts in cumulus
   and synoptic-scale interactions. *Nature Geoscience*, 8, 261–268, URL https://doi.org/10.1038/
   ngeo2398.

```
    <sup>608</sup> Couvreux, F., F. Hourdin, and C. Rio, 2010: Resolved versus parametrized boundary-layer plumes.
    <sup>609</sup> part i: A parametrization-oriented conditional sampling in large-eddy simulations. Boundary
    <sup>610</sup> Layer Meteorology, 134, 441–458, doi:10.1007/s10546-009-9456-5.
```

<sup>611</sup> Craske, J., and M. van Reeuwijk, 2015: Energy dispersion in turbulent jets. part 1. Direct simulation <sup>612</sup> of steady and unsteady jets. *J. Fluid Mech.*, **763**, 500–537, doi:10.1017/jfm.2014.640.

28

613	da Silva, C., J.	Hunt, I.	Eames, and	1 J.	Westerweel,	2014:	Interfacial	layers	between	regions	of
614	different turb	ulence in	tensity. Anr	u. R	Rev. Fluid Me	ch., <b>46</b>	, 567–590.				

- <sup>615</sup> Dawe, J., and P. Austin, 2011: The influence of the cloud shell on tracer budget measurements of LES cloud entrainment. *J. Atmos. Sci.*, **68**, 2909–2920, doi:10.1175/2011JAS3658.1.
- de Roode, S. R., and C. S. Bretherton, 2003: Mass-Flux Budgets of Shallow Cumulus Clouds. J.

Atmos. Sci., **60** (1), 137–151, doi:10.1175/1520-0469(2003)060<0137:MFBOSC>2.0.CO;2.

- <sup>619</sup> de Rooy, W. C., and Coauthors, 2013: Entrainment and detrainment in cumulus convection: an <sup>620</sup> overview. *Quart. J. Roy. Meteor. Soc.*, **139** (**670**), 1–19, doi:10.1002/qj.1959.
- Fodor, K., and J. P. Mellado, 2020: New insights into wind shear effects on entrainment in convective boundary layers using conditional analysis. *Journal of the Atmospheric Sciences*,
  77 (9), 3227 3248, doi:10.1175/JAS-D-19-0345.1, URL https://journals.ametsoc.org/view/
  journals/atsc/77/9/jasD190345.xml.
- Gerard, L., 2015: Bulk Mass-Flux Perturbation Formulation for a Unified Approach of Deep Convection at High Resolution. *Mon. Wea. Rev.*, **143** (**10**), 4038–4063, doi:10.1175/ MWR-D-15-0030.1.
- Gregory, D., and P. R. Rowntree, 1990: A Mass Flux Convection Scheme with Representation
   of Cloud Ensemble Characteristics and Stability-Dependent Closure. *Mon. Wea. Rev.*, 118 (7),
   1483–1506, doi:10.1175/1520-0493(1990)118<1483:AMFCSW>2.0.CO;2.
- Gu, J.-F., R. S. Plant, C. E. Holloway, T. R. Jones, A. Stirling, P. A. Clark, S. J. Woolnough, and T. L.
- <sup>632</sup> Webb, 2020: Evaluation of the Bulk Mass Flux Formulation Using Large-Eddy Simulations. J.
- Atmos. Sci., 77 (6), 2115–2137, doi:10.1175/jas-d-19-0224.1.

- Heus, T., and H. Jonker, 2008: Subsiding shells around shallow cumulus clouds. *J. Atmos. Sci.*,
   65, 1003–1018, doi:10.1175/2007JAS2322.1.
- Heus, T., G. van Dijk, H. Jonker, and H. van den Akker, 2008: Mixing in shallow cumulus clouds
   studied by lagrangian particle tracking. *J. Atmos. Sci.*, 65, 2581–2597.
- Holland, J. Z., and E. M. Rasmusson, 1973: Measurements of the atmospheric mass, energy, and momentum budgets over a 500-kilometer square of tropical ocean. *Monthly Weather Review*, 101 (1), 44–55, doi:10.1175/1520-0493(1973)101<0044:MOTAME>2.3.CO;
- <sup>641</sup> 2, URL https://journals.ametsoc.org/view/journals/mwre/101/1/1520-0493\_1973\_101\_0044\_ <sup>642</sup> motame\_2\_3\_co\_2.xml.
- Johnson, R. H., 1976: The role of convective-scale precipitation downdrafts in cumulus and synoptic-scale interactions. *J. Atmos. Sci.*, **33** (10), 1890–1910.
- <sup>645</sup> Jonker, H. J., T. Heus, and P. P. Sullivan, 2008: A refined view of vertical mass transport by <sup>646</sup> cumulus convection. *Geophysical Research Letters*, **35** (**7**), 1–5, doi:10.1029/2007GL032606.
- Kain, J. S., and J. M. Fritsch, 1990: A one-dimensional entraining/detraining plume model and its
- application in convective parameterization. J. Atmos. Sci., 47 (23), 2784–2802.
- Kankanwadi, K. S., and O. R. H. Buxton, 2020: Turbulent entrainment into a cylinder wake from
   a turbulent background. *JFM*, **905**, A35, doi:10.1017/jfm.2020.755.
- Moorthi, S., and M. J. Suarez, 1992: Relaxed Arakawa-Schubert. A Parameterization of Moist
- 652 Convection for General Circulation Models. Mon. Wea. Rev., 120 (6), 978–1002, doi:10.1175/
- <sup>653</sup> 1520-0493(1992)120<0978:RASAPO>2.0.CO;2.
- <sup>654</sup> Nair, V., T. Heus, and M. van Reeuwijk, 2020: Dynamics of subsiding shells in actively growing
- clouds with vertical updrafts. J. Atmos. Sci., 77 (4), 1353–1369, doi:10.1175/JAS-D-19-0018.1.

- Park, S., P. Gentine, K. Schneider, and M. Farge, 2016: Coherent structures in the boundary and
   cloud layers: Role of updrafts, subsiding shells, and environmental subsidence. *J. Atmos. Sci.*,
   **73**, 1789–1814.
- Park, S.-B., T. Heus, and P. Gentine, 2017: Role of convective mixing and evaporative cooling in
   shallow convection. *J. Geophys. Res.*, **122**, 5351–5363, doi:10.1002/2017JD026466.
- Perrin, V., and H. Jonker, 2015: Lagrangian droplet dynamics in the subsiding shell of a cloud using
   Direct Numerical Simulations. *J. Atmos. Sci.*, **72**, 4015–4028, doi:10.1175/JAS-D-15-0045.1.
- <sup>663</sup> Romps, D., 2010: A direct measure of entrainment. J. Atmos. Sci., 67, 1908–1927, doi:10.1175/
   <sup>664</sup> 2010JAS3371.1.
- Siebesma, A., and J. Cuijpers, 1995: Evaluation of parametric assumptions for shallow cumulus
   convection. *J. Atmos. Sci.*, **52**, 650–666.
- <sup>667</sup> Siebesma, A., and Coauthors, 2003: A Large Eddy Simulation intercomparison study of shallow <sup>668</sup> cumulus convection. *J. Atmos. Sci.*, **60**, 1201–1219.
- Siebesma, A. P., 1998: *Shallow Cumulus Convection*, 441–486. Springer Netherlands, Dordrecht,
   doi:10.1007/978-94-011-5058-3\_19.
- <sup>671</sup> Sommeria, G., and J. Deardorff, 1976: Subgrid-scale condensation in models of nonprecipitating <sup>672</sup> clouds. *J. Atmos. Sci.*, **34**, 344–346.
- Tiedtke, M., 1989: A comprehensive mass flux scheme for cumulus parameterization in large-scale
  models. *Mon. Wea. Rev.*, **177**, 1779–1800, doi:10.1175/1520-0493(1989)117.
- van Heerwaarden, C. C., B. J. H. van Stratum, T. Heus, J. A. Gibbs, E. Fedorovich, and J. P. Mellado,
- <sup>676</sup> 2017: Microhh 1.0: a computational fluid dynamics code for direct numerical simulation and

- <sup>677</sup> large-eddy simulation of atmospheric boundary layer flows. *Geoscientific Model Development*,
- **10 (8)**, 3145–3165, doi:10.5194/gmd-10-3145-2017.
- <sup>679</sup> Yeo, K., and D. M. Romps, 2012: Measurement of Convective Entrainment Using Lagrangian
- Particles. J. Atmos. Sci., **70** (1), 266–277, doi:10.1175/jas-d-12-0144.1.

## 681 LIST OF FIGURES

682 683	Fig. 1.	Numerical setup. A volumetric forcing is applied over the grey cloud region from $x = 0$ to $x = 1m$ .		35
684 685	Fig. 2.	Mean buoyancy $\overline{b}$ and mean vertical velocity $\overline{w}$ normalized by their respective minimum values $b^*$ and $w^*$ at $t_0$ when particles are seeded.		36
686 687 688 689	Fig. 3.	The different zones at the edge of the cloudy domain at the time of seeding $(t_0)$ . The color scheme is: red - Cloud Core (CC), blue - Visible Shell (VS), yellow - Invisible Shell (IS) and white - Environment (E). Also shown are the interfaces: Cloud Core Boundary (CCB), Visible Shell Boundary(VSB) and Turbulent Non-Turbulent Interface (TNTI).		37
690 691 692	Fig. 4.	Instantaneous particle positions on a two dimensional slice of the flow $t^* = 0.67$ . Also shown are the three interfaces. Particle colors represent the zone in which they were initially seeded : red (CC), blue (VS), yellow (IS) and pink (E).		38
693 694 695	Fig. 5.	An origin-destination matrix for the initial and final zones of all particles that leave their zone of origin. Figure shows the percentage wise distribution of particles based on the zone of origin at the time of seeding and the zone where they end up at time $t^* = 0.67$ .		39
696 697 698 699	Fig. 6.	Average time taken by all entraining/detraining particles to travel across a zone. The values shown are averages of the residence times in a zone normalized with $\tau$ . Asterisk indicates insufficient number of particles travelling from interface to interface to get a reliable mean value.		40
700 701 702	Fig. 7.	Box plot showing the time histories for $\omega^2$ and b for entraining particles crossing the TNTI (a,b), VSB(c,d) and CCB(e,f). Outlier points and whiskers for extreme values have been removed for the sake of clarity.		41
703 704 705	Fig. 8.	Box plot showing the time histories for $\omega^2$ and b for detraining particles crossing the VSB(a,b) and CCB(c,d). Outlier points and whiskers for extreme values have been removed for the sake of clarity.		42
706 707 708 709 710 711	Fig. 9.	Histograms showing the properties of particles that have E- entrained (cyan) and D- detrained (red) across the CCB (a, b), VSB (c,d) and the TNTI (e,f) at time $t^* = 0.67$ . The properties shown are (a,c,e) $q_t$ , and (b,d,f) $w$ . The dashed and dot-dashed cyan lines indicate mean magnitudes of the entrained particles and all particles in the entraining zone respectively and the red dashed and dotted line indicates the mean of detraining particles and all the particles in the detraining zone.		43
712 713	Fig. 10.	Contour plot of enstrophy at a mid-cloud layer ( $z = 1000m$ ). Black contours denote the cloud boundary.		44
714	Fig. 11.	Mass flux density and enstrophy as a function of the distance to the nearest cloud edge	•	45
715 716 717	Fig. 12.	Conditionally sampled enstrophy in each of the zones, averaged between hour 4-10 of the simulation, with a sample time of 0.5hr. Red: Cloud core, Blue: Visible shell, Yellow: Invisible shell, Black: Environment.		46
718 719	Fig. 13.	Cumulative mass flux vs enstrophy for non-cloudy regions, averaged between hour 4-10 of the simulation, with a sample time of 0.5hr.		47

720 721 722	Fig. 14.	Conditionally sampled mass flux in each zone, averaged between hour 4-10 of the simulation, with a sample time of 0.5hr. Red: Cloud core, Blue: Visible shell, Yellow: Invisible shell, Black: Environment.	48
723 724	Fig. A1.	Isolines of $b = 0$ (black) at the time of seeding ( $t_0$ ). Superimposed are the different interfaces obtained by applying different threshold values to the enstrophy field $\omega_{th}^2$ .	49
725 726 727	Fig. B1.	Origin-Destination matrix for particle final locations showing percentage-wise distribution of particles based on zone of origin at the time of seeding and the final destination zone at $t*=0.99$ .	50
728 729	Fig. B2.	Time histories for $\omega^2$ and b for entraining particles crossing the TNTI (a,b) and VSB (c,d), and for detraining particles crossing the CCB (e,f).	51
730 731 732 733	Fig. B3.	Average time taken by all entraining/detraining particles to travel across a zone. The values shown are averages of the residence times in a zone normalized with $\tau$ . Asterisk indicates insufficient number of particles travelling from interface to interface to get a reliable mean value.	52



FIG. 1. Numerical setup. A volumetric forcing is applied over the grey cloud region from x = 0 to x = 1m.



FIG. 2. Mean buoyancy  $\overline{b}$  and mean vertical velocity  $\overline{w}$  normalized by their respective minimum values  $b^*$  and  $w^*$  at  $t_0$  when particles are seeded.



FIG. 3. The different zones at the edge of the cloudy domain at the time of seeding ( $t_0$ ). The color scheme is: red - Cloud Core (CC), blue - Visible Shell (VS), yellow - Invisible Shell (IS) and white - Environment (E). Also shown are the interfaces: Cloud Core Boundary (CCB), Visible Shell Boundary(VSB) and Turbulent Non-Turbulent Interface (TNTI).



FIG. 4. Instantaneous particle positions on a two dimensional slice of the flow  $t^* = 0.67$ . Also shown are the three interfaces. Particle colors represent the zone in which they were initially seeded : red (CC), blue (VS), yellow (IS) and pink (E).



Fig. 5. An origin-destination matrix for the initial and final zones of all particles that leave their zone of origin. Figure shows the percentage wise distribution of particles based on the zone of origin at the time of seeding and the zone where they end up at time  $t^* = 0.67$ .



FIG. 6. Average time taken by all entraining/detraining particles to travel across a zone. The values shown are averages of the residence times in a zone normalized with  $\tau$ . Asterisk indicates insufficient number of particles travelling from interface to interface to get a reliable mean value.



FIG. 7. Box plot showing the time histories for  $\omega^2$  and *b* for entraining particles crossing the TNTI (a,b), VSB(c,d) and CCB(e,f). Outlier points and whiskers for extreme values have been removed for the sake of clarity.



FIG. 8. Box plot showing the time histories for  $\omega^2$  and *b* for detraining particles crossing the VSB(a,b) and CCB(c,d). Outlier points and whiskers for extreme values have been removed for the sake of clarity.



FIG. 9. Histograms showing the properties of particles that have E- entrained (cyan) and D- detrained (red) across the CCB (a, b), VSB (c,d) and the TNTI (e,f) at time  $t^* = 0.67$ . The properties shown are (a,c,e)  $q_t$ , and (b,d,f) *w*. The dashed and dot-dashed cyan lines indicate mean magnitudes of the entrained particles and all particles in the entraining zone respectively and the red dashed and dotted line indicates the mean of detraining particles and all the particles in the detraining zone.



Fig. 10. Contour plot of enstrophy at a mid-cloud layer (z = 1000m). Black contours denote the cloud boundary.



FIG. 11. Mass flux density and enstrophy as a function of the distance to the nearest cloud edge.



FIG. 12. Conditionally sampled enstrophy in each of the zones, averaged between hour 4-10 of the simulation, with a sample time of 0.5hr. Red: Cloud core, Blue: Visible shell, Yellow: Invisible shell, Black: Environment.



FIG. 13. Cumulative mass flux vs enstrophy for non-cloudy regions, averaged between hour 4-10 of the simulation, with a sample time of 0.5hr.



FIG. 14. Conditionally sampled mass flux in each zone, averaged between hour 4-10 of the simulation, with a sample time of 0.5hr. Red: Cloud core, Blue: Visible shell, Yellow: Invisible shell, Black: Environment.



Fig. A1. Isolines of b = 0 (black) at the time of seeding ( $t_0$ ). Superimposed are the different interfaces obtained by applying different threshold values to the enstrophy field  $\omega_{th}^2$ .



Fig. B1. Origin-Destination matrix for particle final locations showing percentage-wise distribution of particles based on zone of origin at the time of seeding and the final destination zone at t \* = 0.99.



Fig. B2. Time histories for  $\omega^2$  and *b* for entraining particles crossing the TNTI (a,b) and VSB (c,d), and for detraining particles crossing the CCB (e,f).



Fig. B3. Average time taken by all entraining/detraining particles to travel across a zone. The values shown are averages of the residence times in a zone normalized with  $\tau$ . Asterisk indicates insufficient number of particles travelling from interface to interface to get a reliable mean value.