

Commissioning of the ATLAS Pixel Detector with Cosmic Ray Data

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on behalf of the ATLAS Collaboration



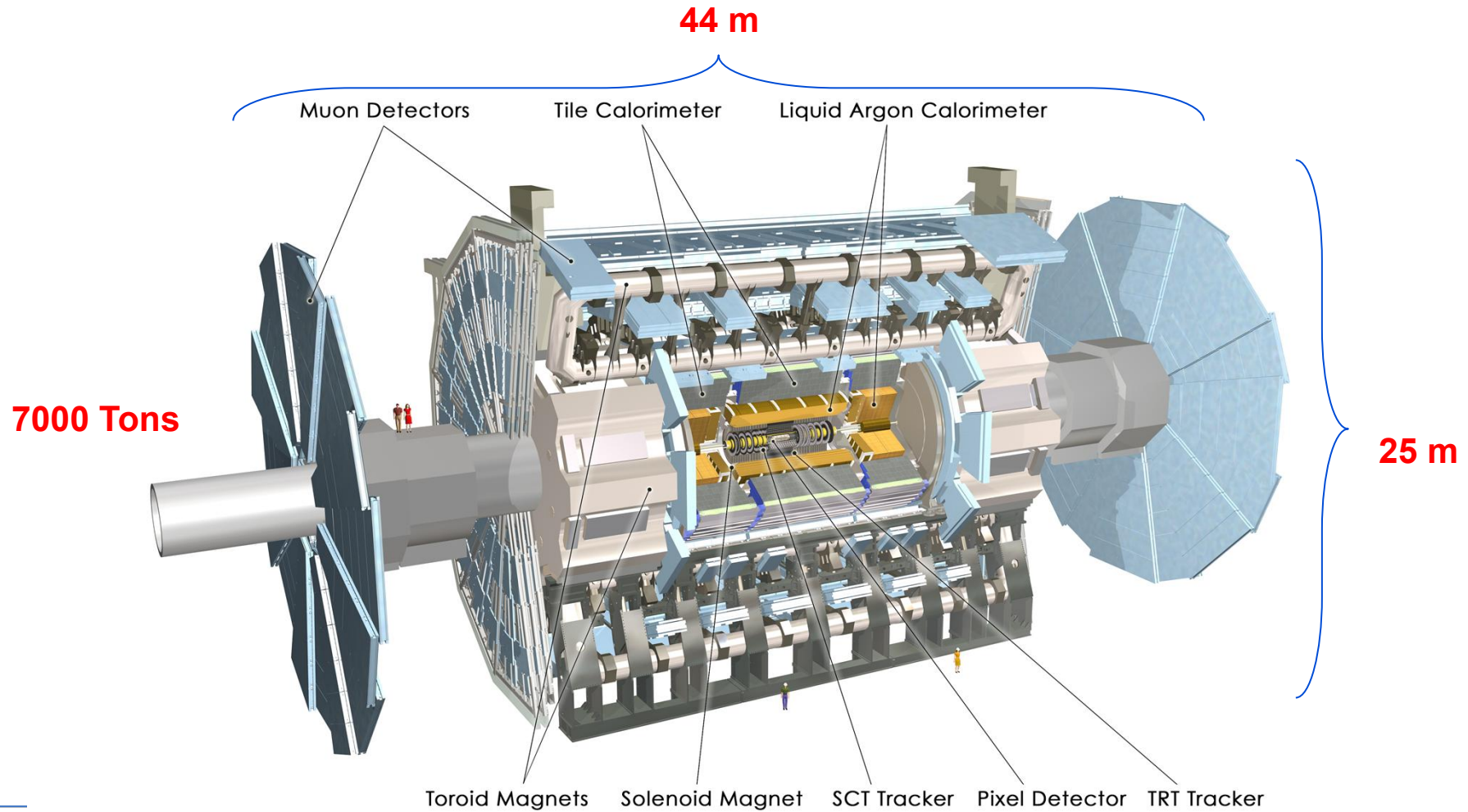
Outline

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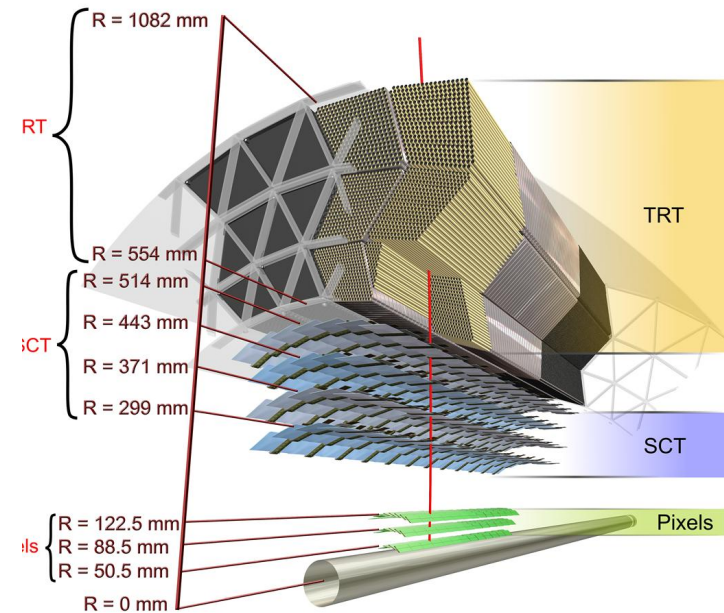
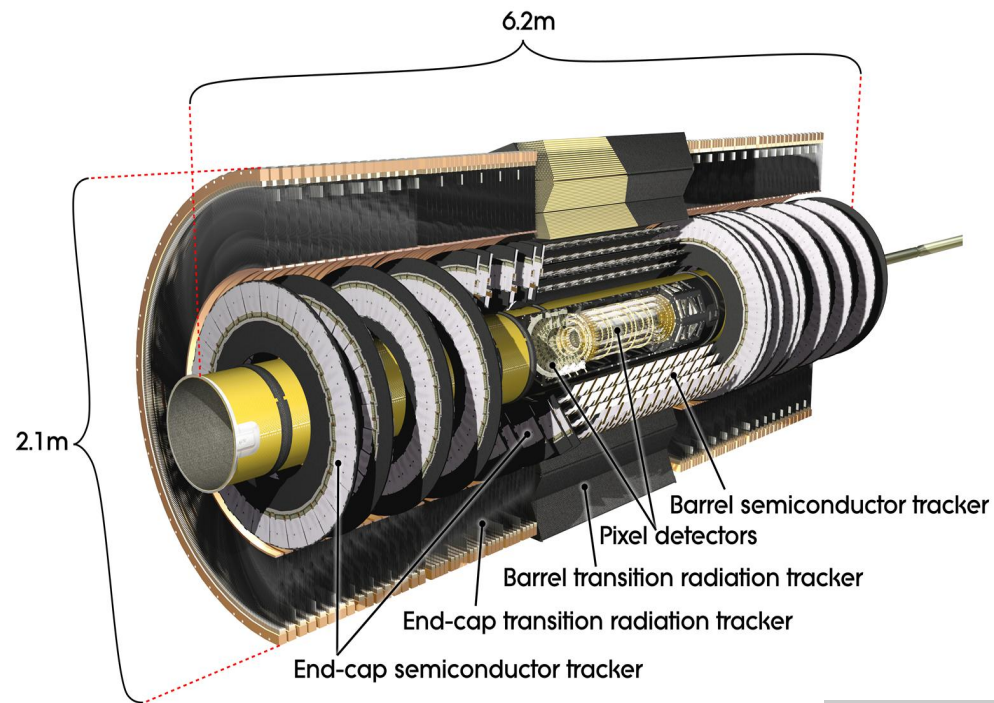
- ATLAS and its Inner Detector
- The ATLAS Pixel Detector and its commissioning
 - ▣ Calibration and Cosmic ray data taking
 - Threshold, noise and masked pixels
 - Time-over-Threshold and timing
 - Resolution, efficiency and noise occupancy
 - Lorentz angle
- Status and expectations
 - ▣ Readiness for collisions and long term operations



The ATLAS Detector



The ATLAS Inner Detector



- 2 T solenoidal magnetic field
- Acceptance $|\eta| < 2.5$
(transition radiation tracker $|\eta| < 2$)

- Momentum resolution
 $\sigma(p_T)/p_T = 0.05\% p_T [\text{GeV}/c] \oplus 1\%$
- Impact parameter resolution ($0.25 < |\eta| < 0.5$)
 $\sigma(d_0) = 10 \mu\text{m} \oplus 140 \mu\text{m} / p_T [\text{GeV}/c]$



The ATLAS Pixel Detector

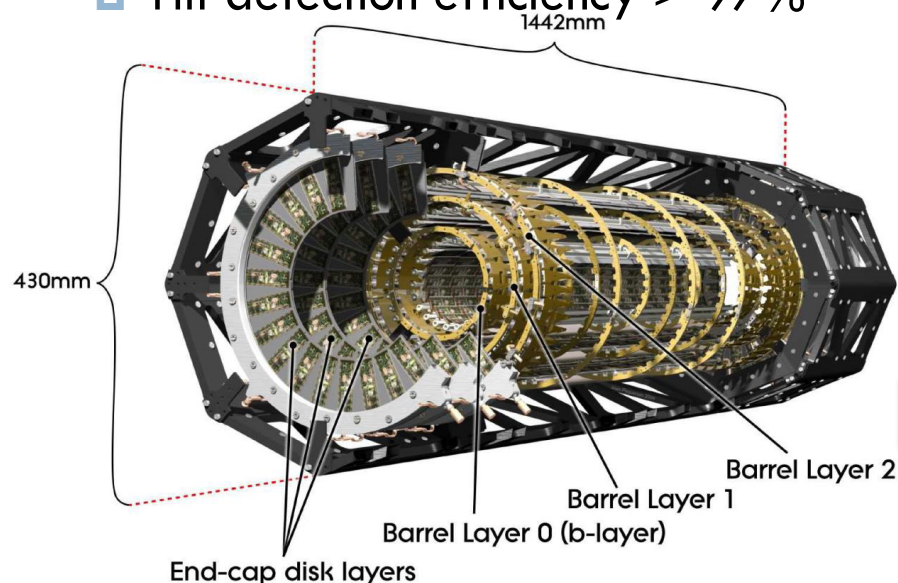
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□ Requirements:

- Position resolution in $r\phi$ -direction $< 15 \mu\text{m}$
- 3 track points for $|\eta| < 2.5$
- Time resolution $< 25 \text{ ns}$
- Hit detection efficiency $> 97\%$

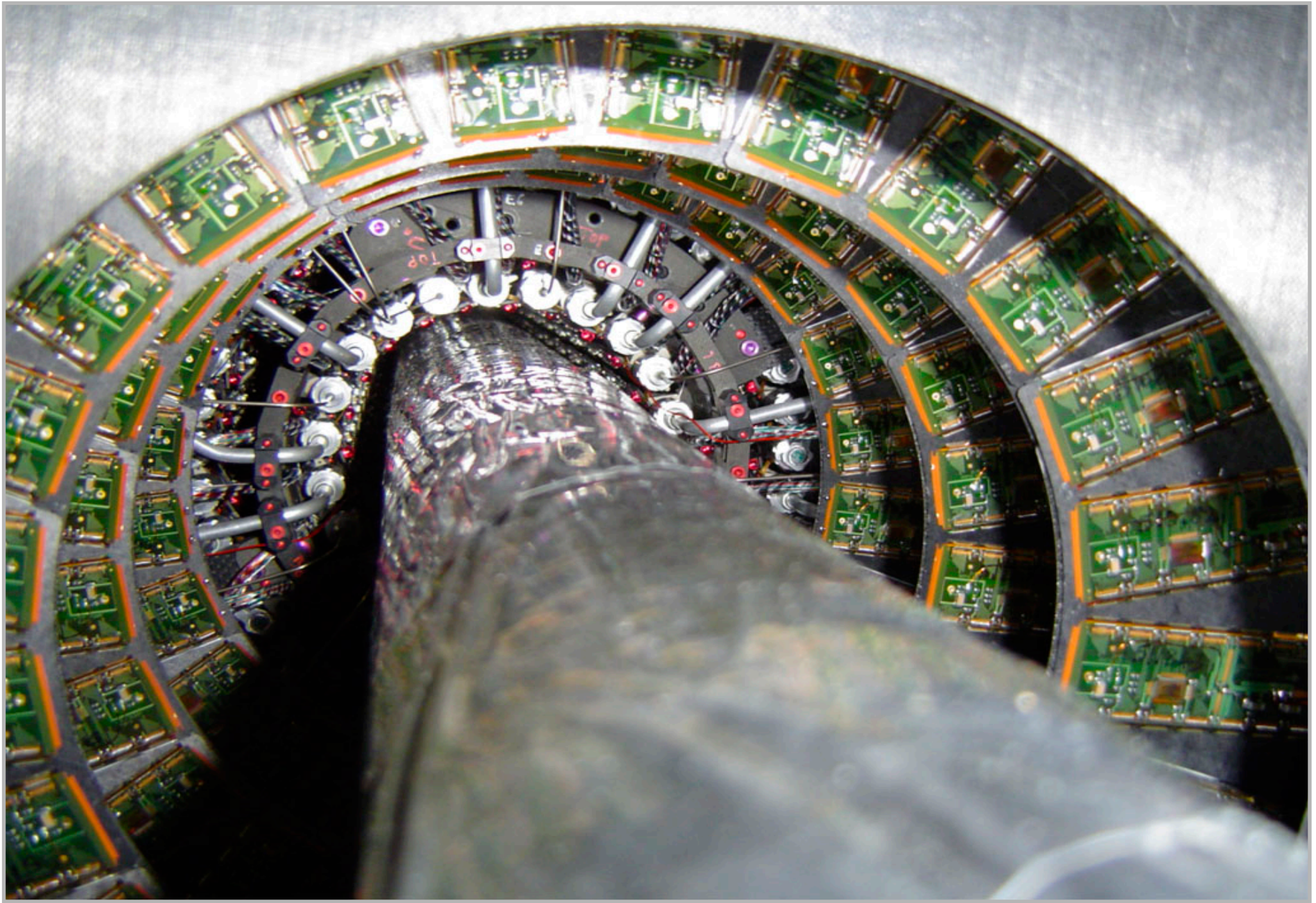
□ Basic Properties:

- 1744 Pixel Modules on three barrel layers and 2 x 3 disks
- 80M readout channels
- Innermost layer at 5 cm
 - Radiation tolerance
 $500 \text{ kGy} / 10^{15} \text{ 1 MeV n}_{\text{eq}}\text{cm}^{-2}$
- Evaporative C_3F_8 cooling integrated in local support structure \rightarrow Module temperature below $0 \text{ }^\circ\text{C}$



How close to the beam pipe? See next slide





B. Di Girolamo - Commissioning of the ATLAS Pixel Detector - RD09 30 September 2009

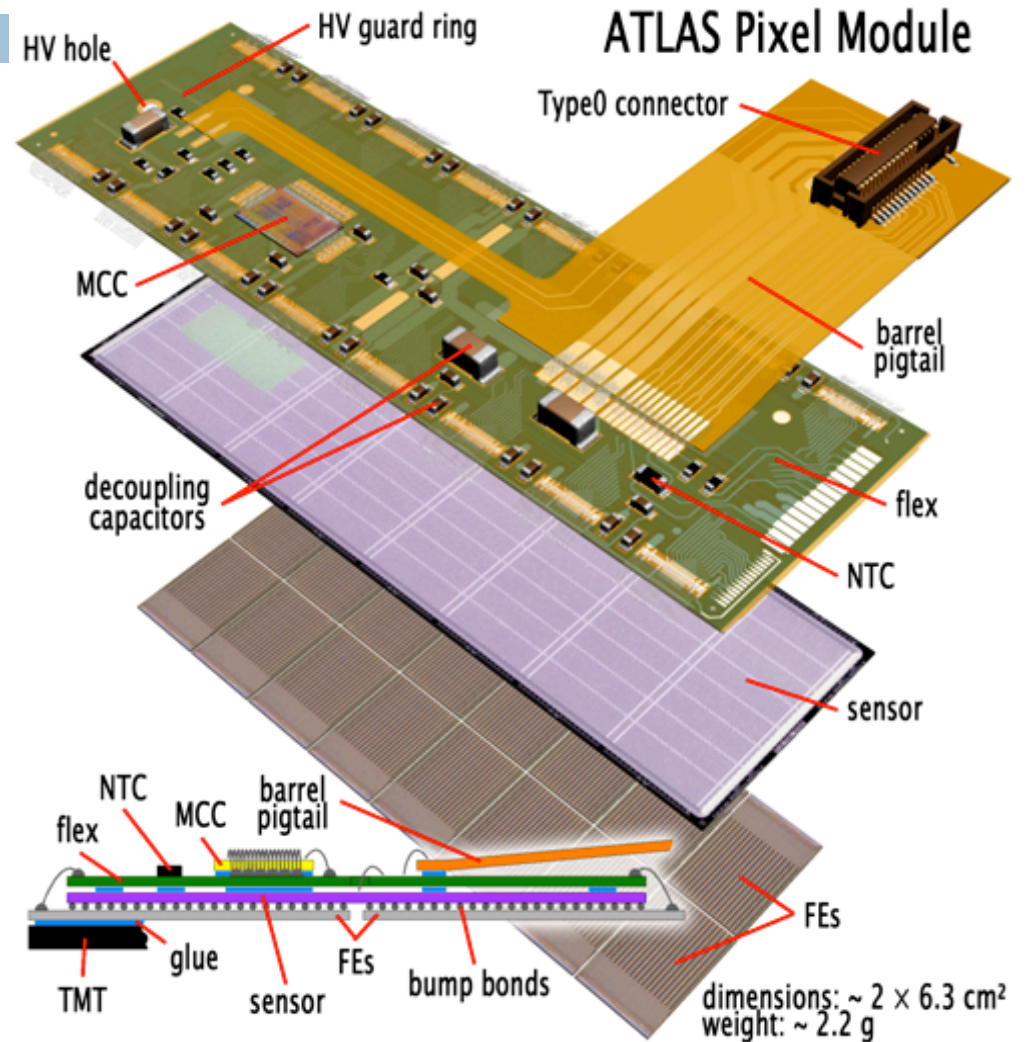


The ATLAS Pixel Module

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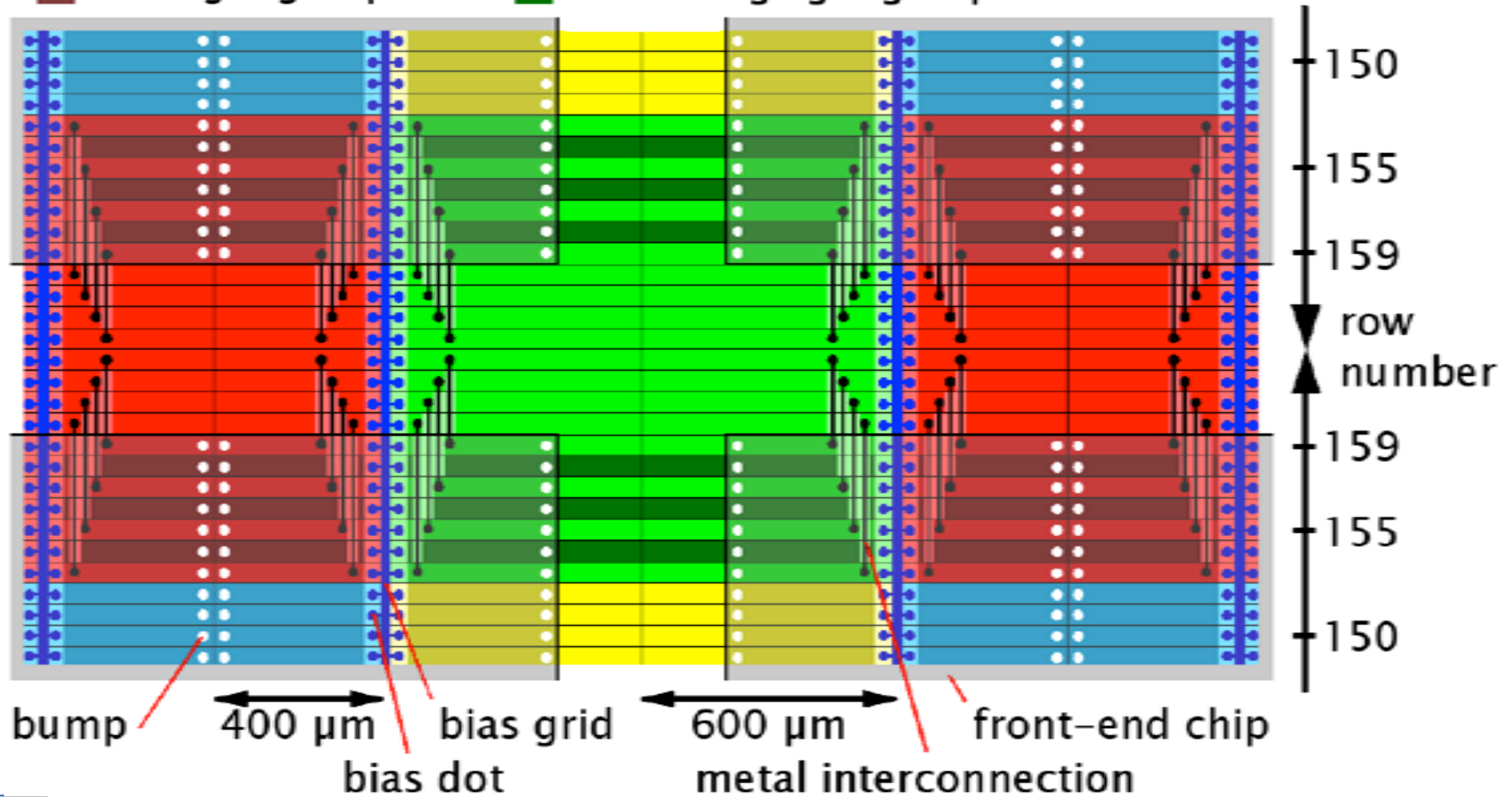
- Sensor
 - ▣ 250 μm thick n-on-n Si sensor
 - ▣ 47232 (328 x 144) Pixels (46080 read channels)
 - ▣ Typical pixel size 50 x 400 μm (50 x 600 μm pixels in gaps between FE chips)
 - ▣ Bias voltage 150 – 600 V

- Readout
 - ▣ 16 FE Chips with 2880 pixels each
 - ▣ Pulse height measured by means of Time over Threshold
 - ▣ Zero suppression in the FE chips, MCC chip builds module event
 - ▣ Data transfer 40 to 160 MHz depending on layer (occupancy)



Inter-chip Region

- pixel
- ganged pixel
- inter-ganged pixel
- long pixel
- long+ganged pixel
- inter-long+ganged pixel



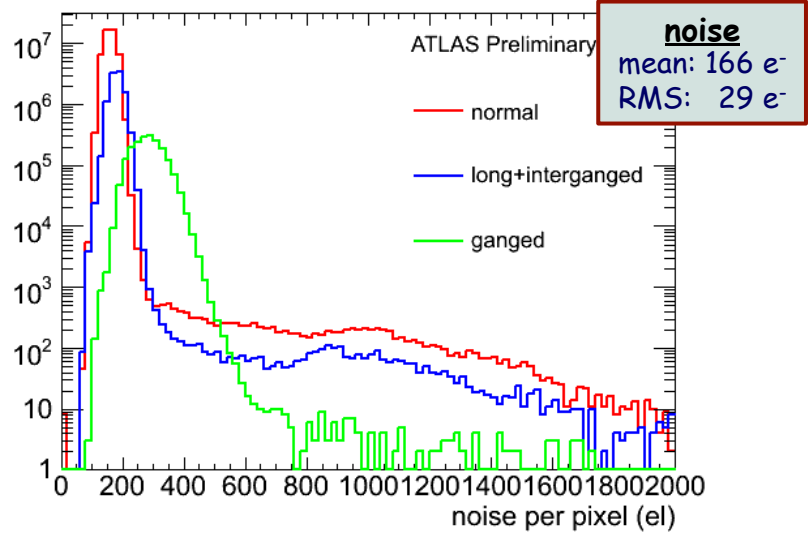
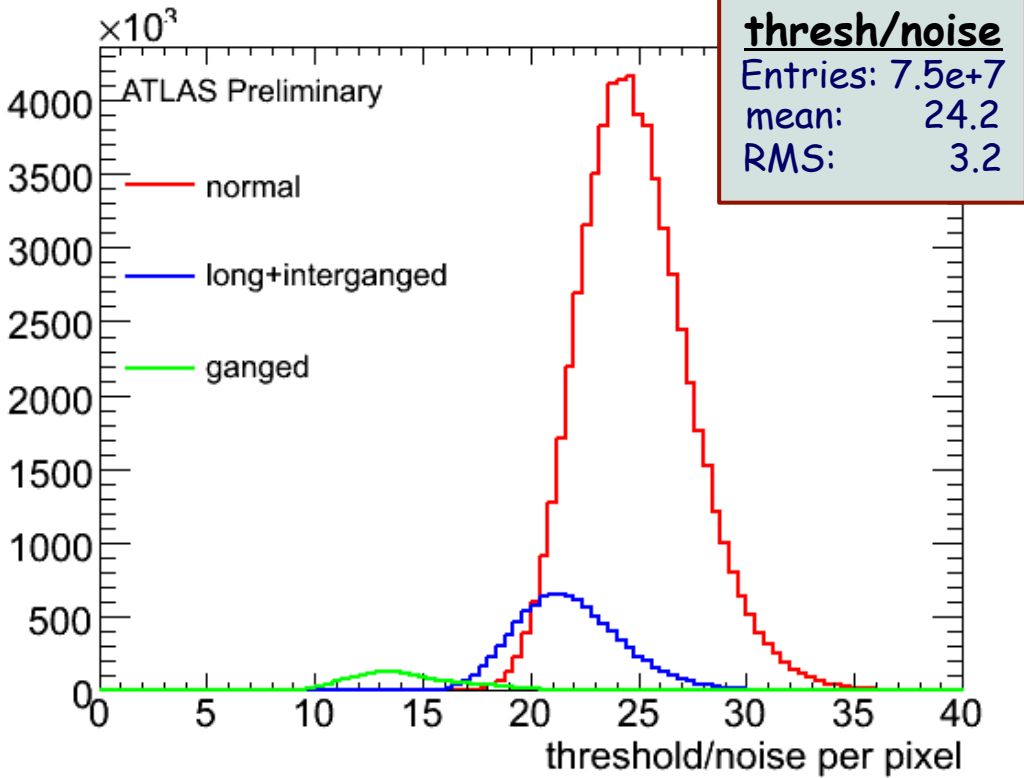
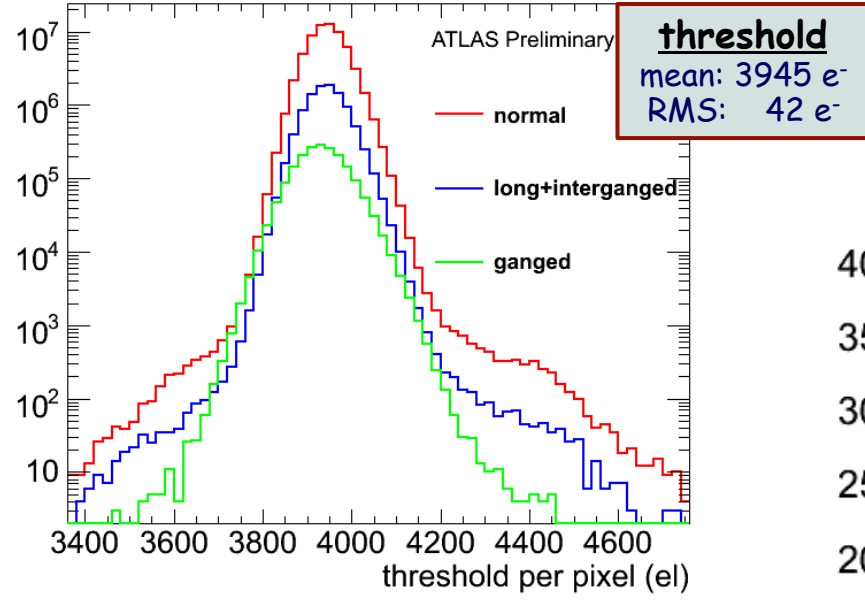
The Pixel Detector commissioning

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- August-December 2008
 - ▣ Functionality checks, calibrations and cosmic ray data
 - ▣ 240 k tracks with field off, 190 k tracks with field on
- May-July 2009
 - ▣ Short calibration period and cosmic ray data
 - ▣ 90 k tracks with field off, 180 k tracks with field on
- Restarted mid-August 2009
 - ▣ 5 weeks for calibration and very soon in continuous cosmic ray data taking until beam comes



Threshold and Noise

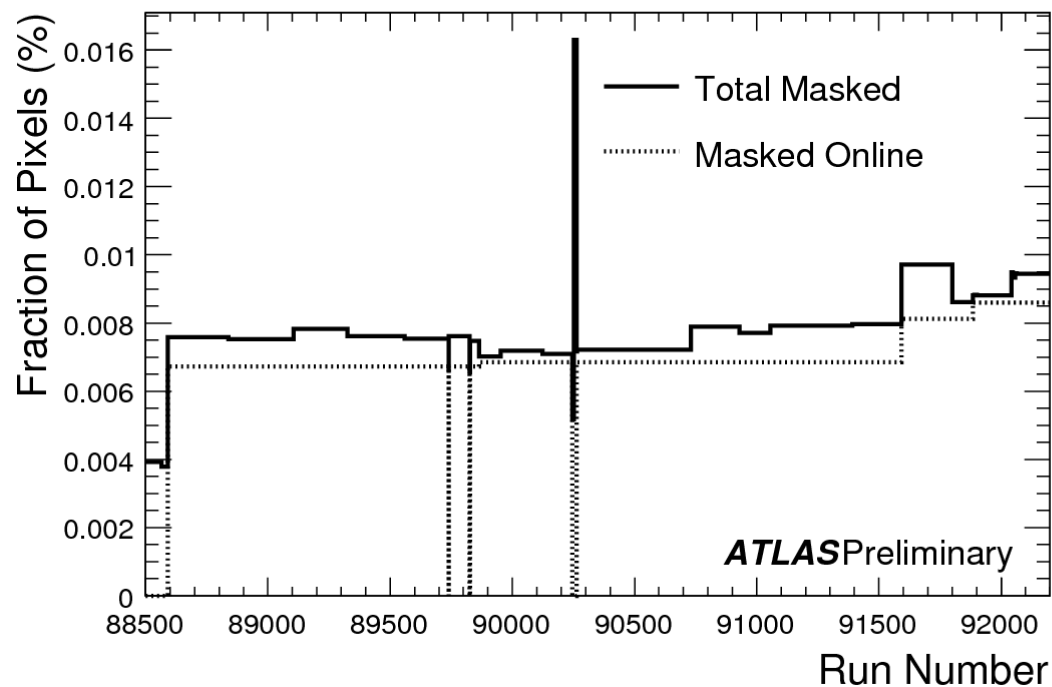
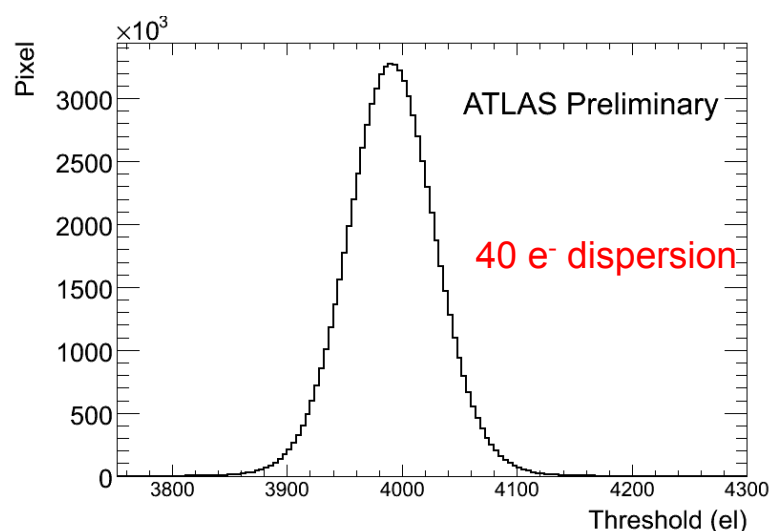


75M pixels in the plots!
 Noise approx. 200 e⁻ for most pixels,
 slightly higher for special pixels
Threshold/Noise approx. 25 for normal pixels



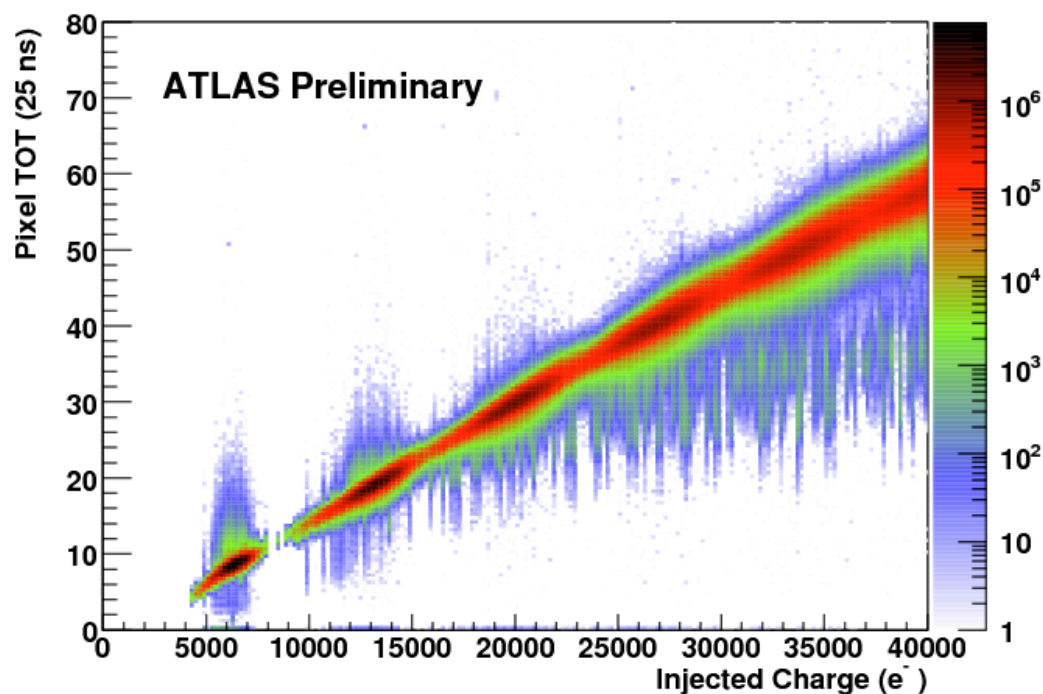
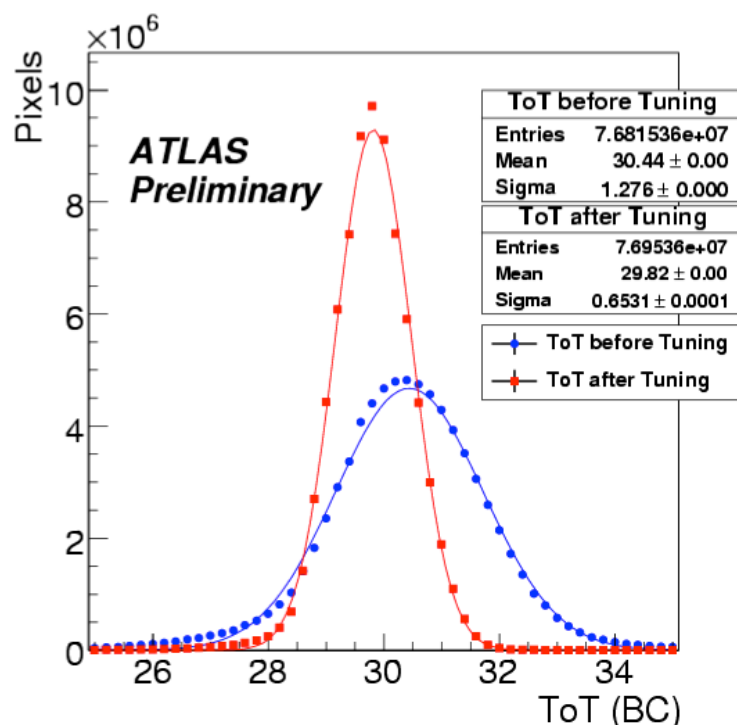
Threshold and Noise: Masked Pixels

- Threshold setting: 4000 e^-
- Threshold tuned pixel by pixel, threshold dispersion $\sim 40 e^-$
- Fraction of masked pixel $\sim 0.02\%$



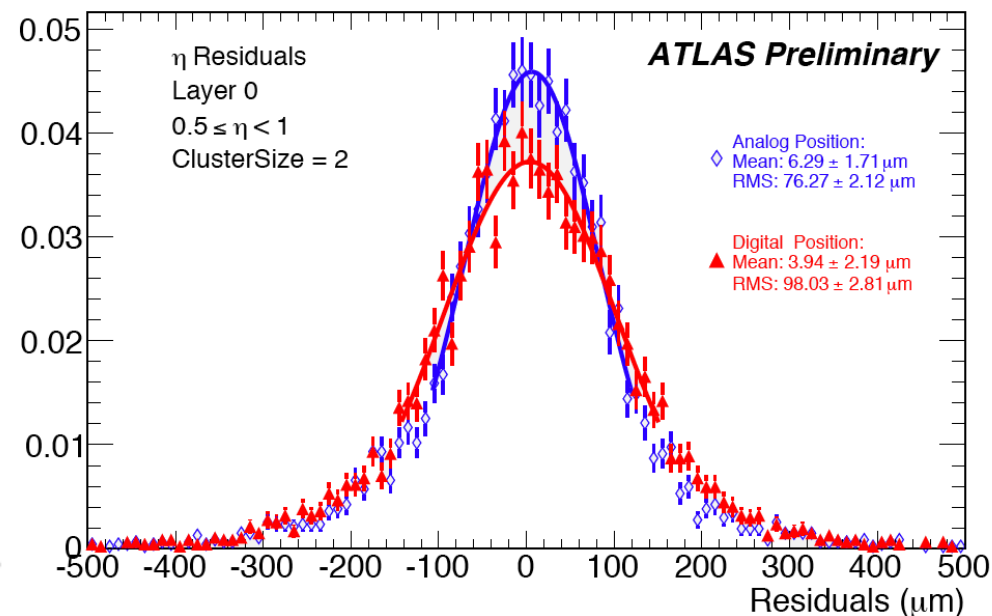
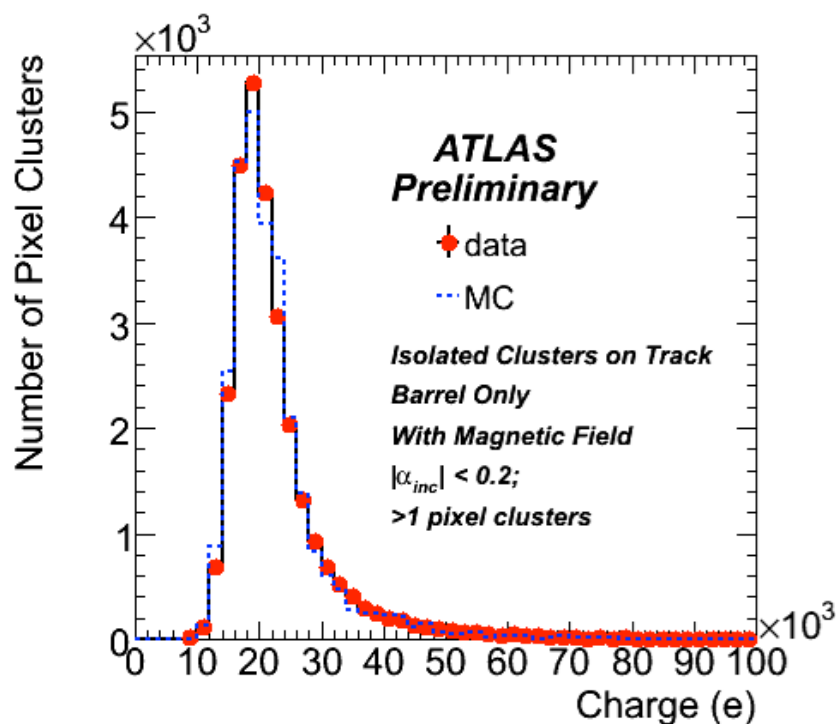
Time over Threshold

- FE Chips provide Time over Threshold information for each hit
 - ▣ Nearly linear dependence on deposited charge
- Pixel-by-pixel tuning; chosen tuning: 30 BC for 20 ke⁻
- Calibration by means of test charge injection



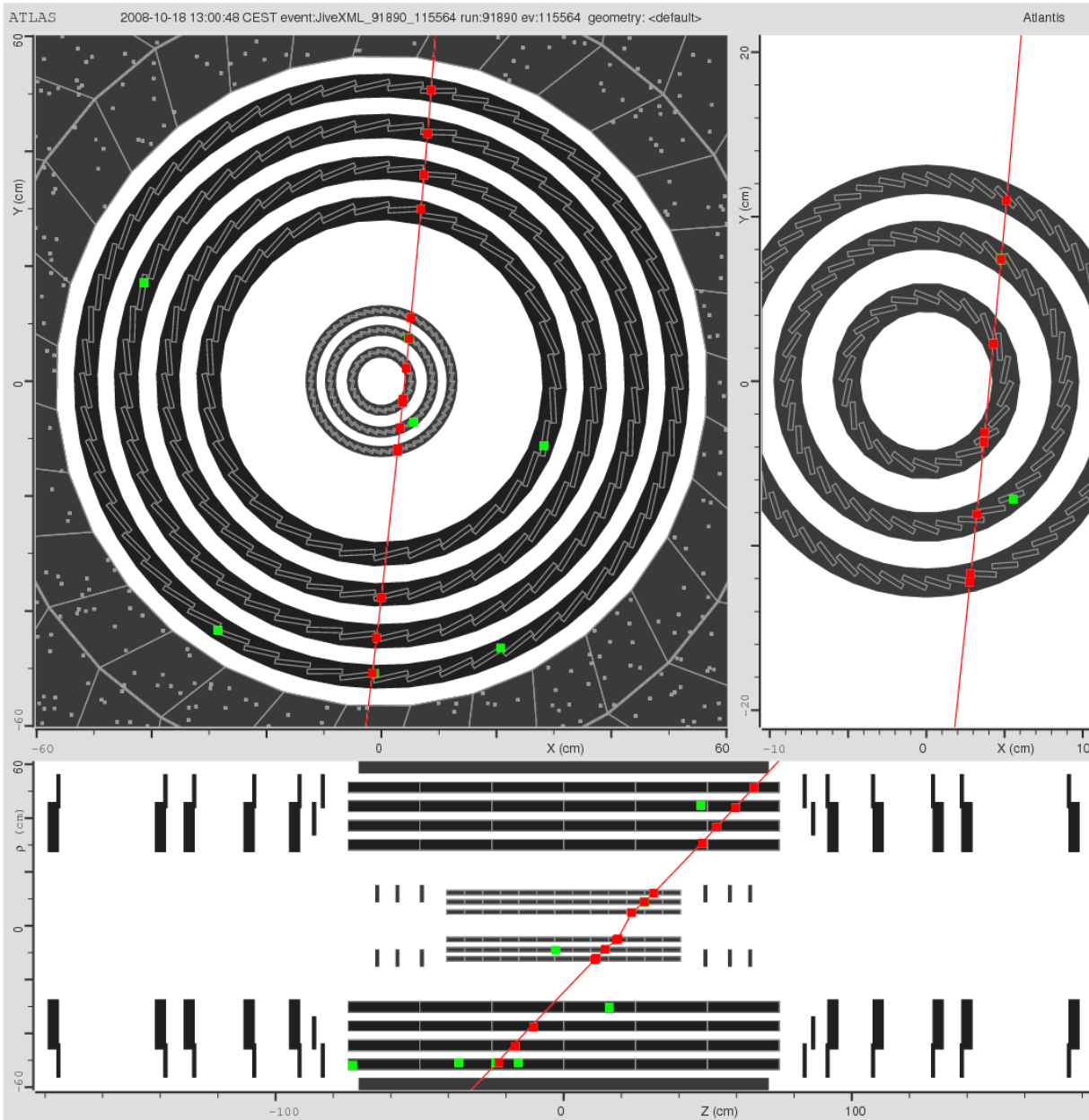
Time over Threshold and Resolution

- Charge measurement with ToT in cosmic ray data taking
 - “Landau” peak at 18300 e⁻ (Simulation 19000 e⁻): Confirms ToT Calibration
- Impact on resolution: still limited by statistics, but noticeable



$\sigma : 98 \mu\text{m} \rightarrow 76 \mu\text{m}$





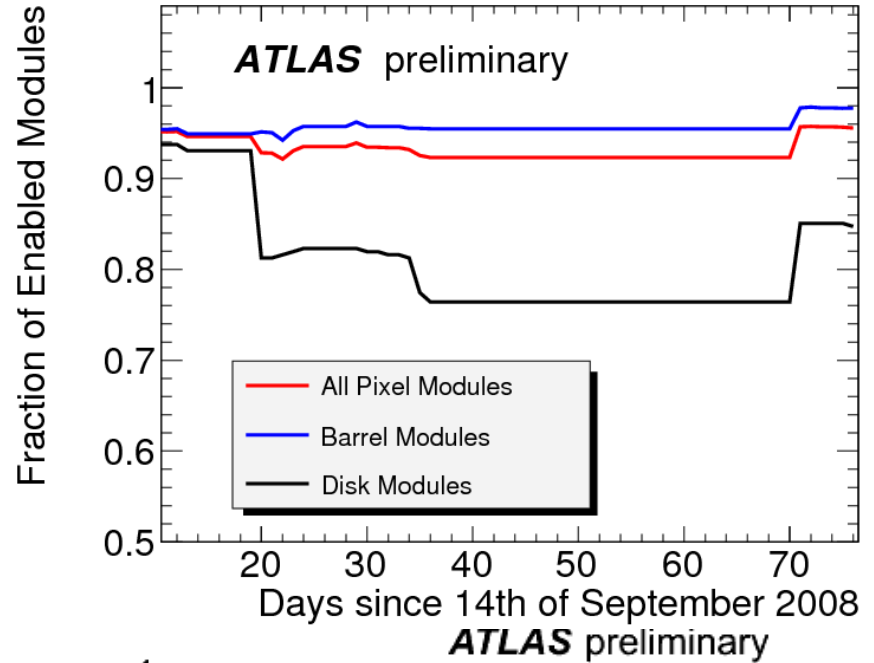
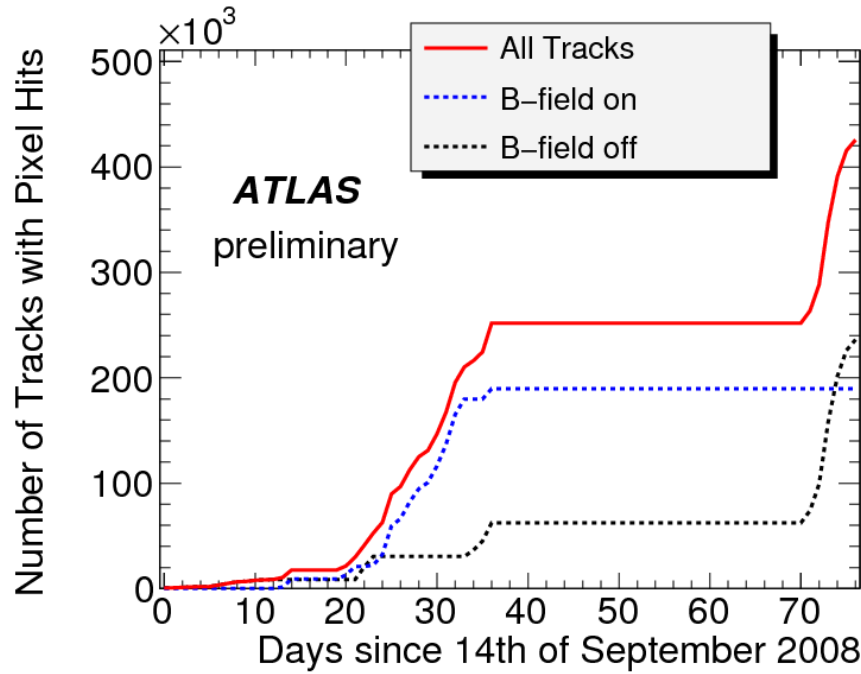
Tracks

- Track with 8 pixel hits on track (2 x 2hits in module overlap regions)
- **Red**: hits on track
- **Green**: isolated hits (noise)
- Noise occupancy:

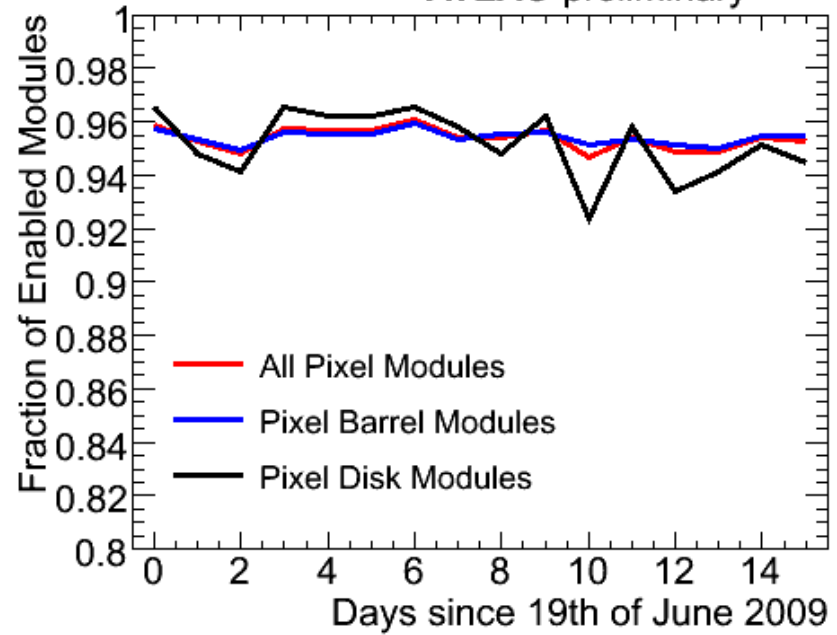
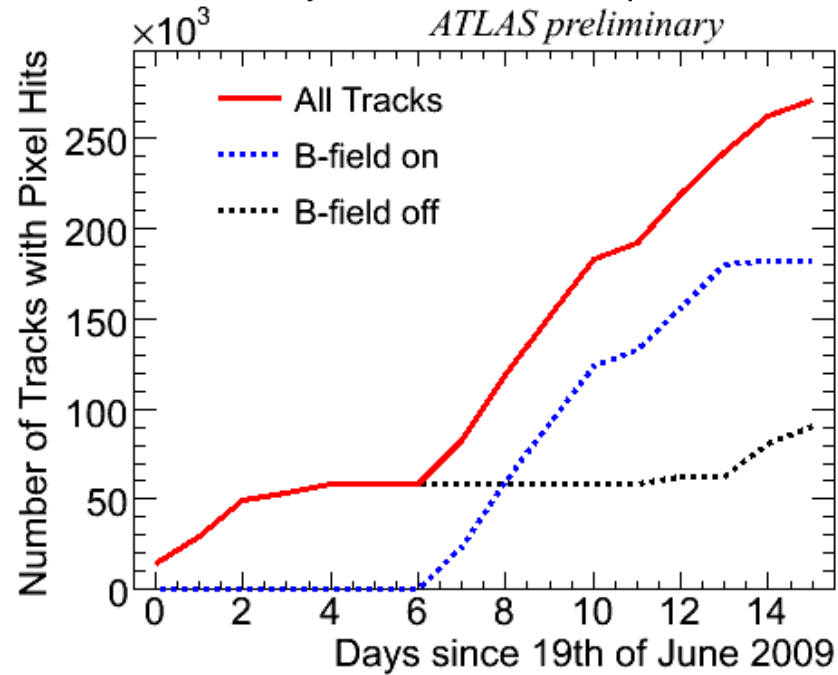
~ 10^{-10} hits/pixel/BC



2008



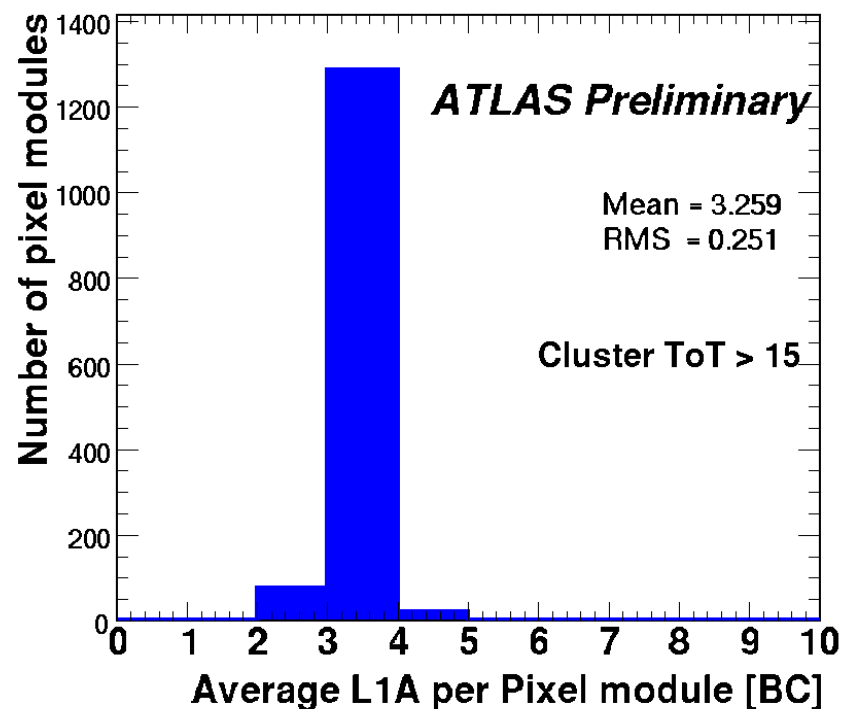
2009



Timing

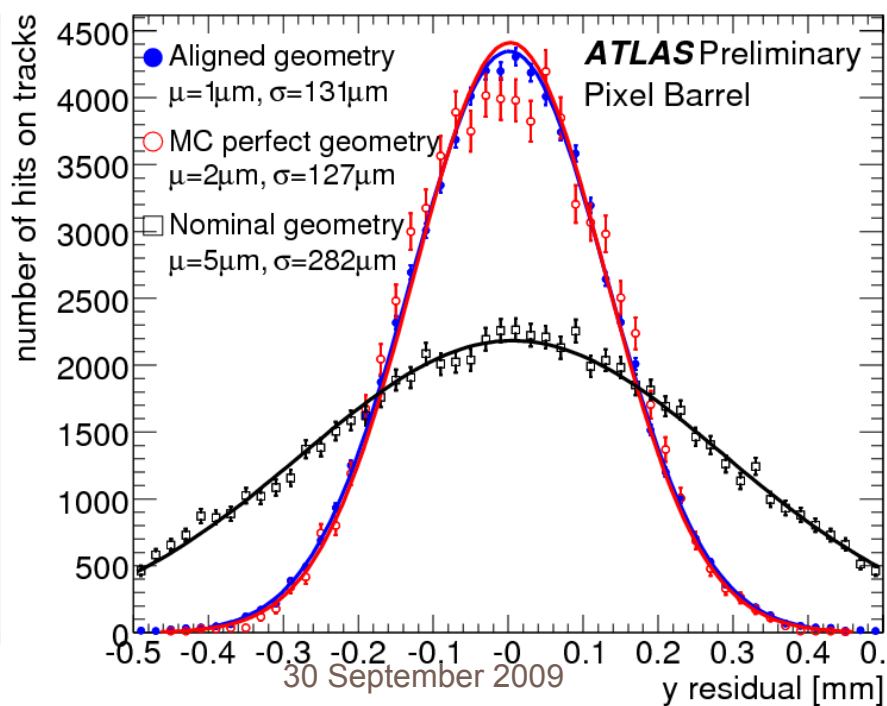
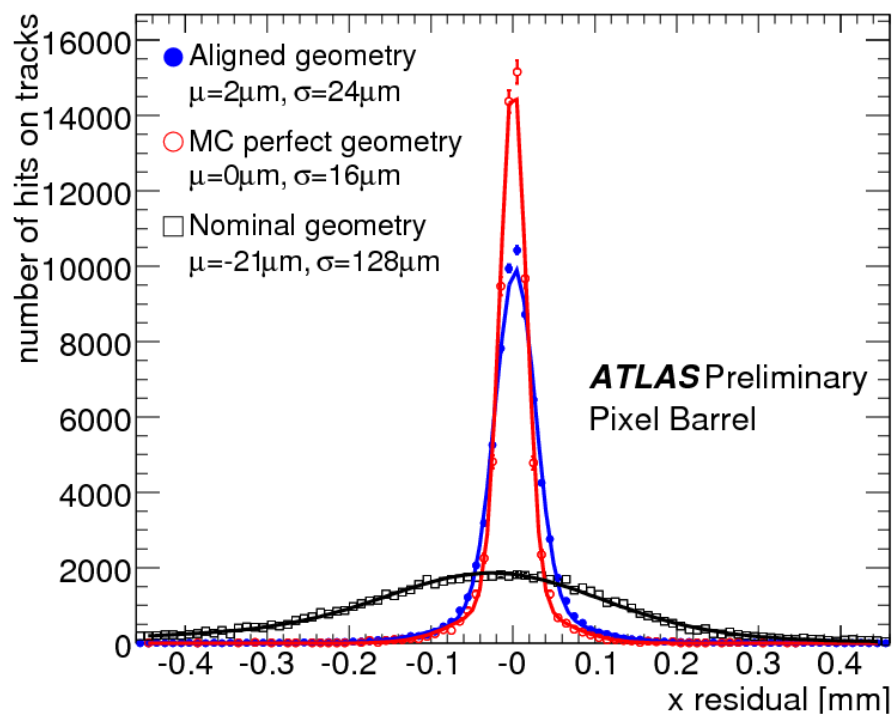
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- Each hit has to be assigned to the correct bunch crossing (25 ns)
- Module clocks have to be precisely aligned with the bunch crossing clock
- In cosmic ray data taking: readout of 8 consecutive BCs (plot shows hit time w.r.t. beginning of readout window)
- Correction of propagation delays:
 - ▣ First step: time alignment of readout crates (oscilloscope measurements)
 - ▣ Second step: time alignment of modules (using cable length data)
- Remaining effects:
 - ▣ Trigger jitter
 - ▣ Random phase of cosmics
 - ▣ Timewalk; “in-time” for less than $\sim 5000 e^-$
- Plan to start data taking with 5 BC, later reduce readout window to 3 BC and 1 BC



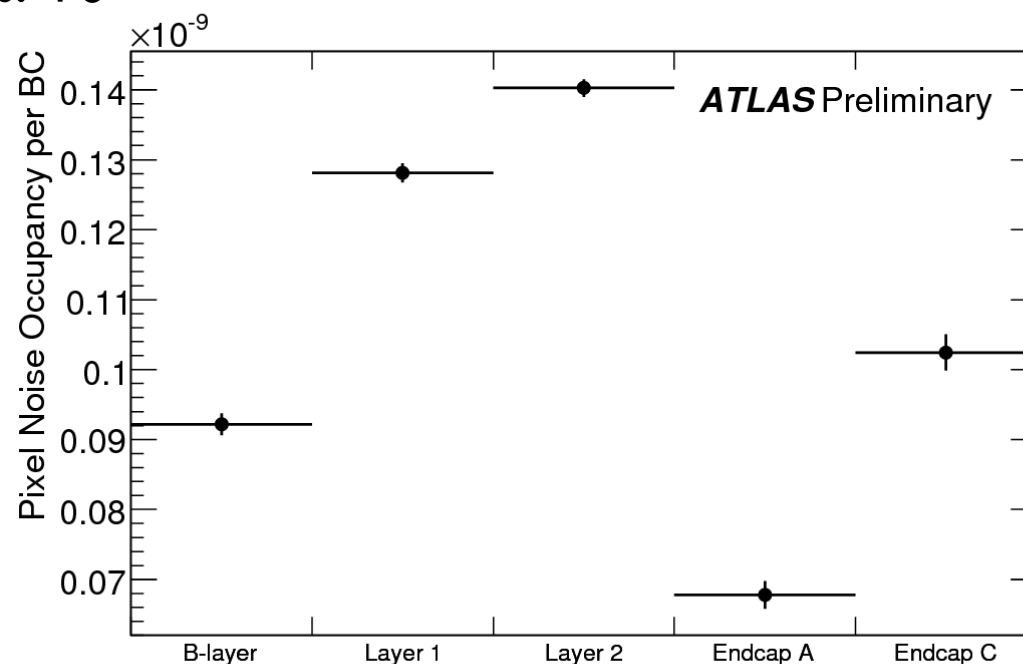
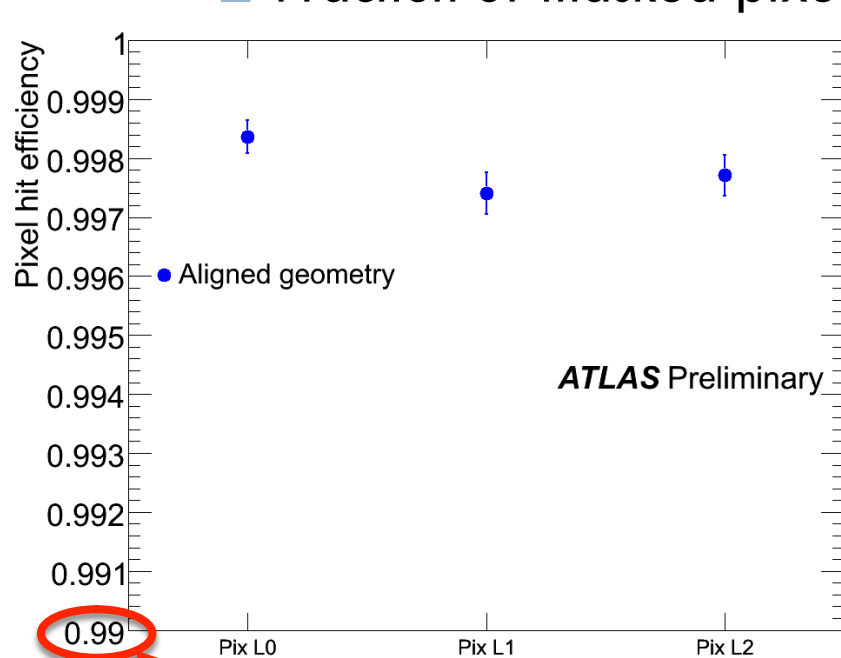
Alignment

- Alignment of pixel barrel modules from cosmic data
 - ▣ Beam data needed for end-cap alignment
- Alignment not yet perfect due to limited statistics, but large improvement w.r.t. nominal geometry and good starting point for alignment with beam:
 - ▣ Precision direction: $128\ \mu\text{m} \rightarrow 24\ \mu\text{m}$ beam direction: $282\ \mu\text{m} \rightarrow 131\ \mu\text{m}$



Efficiency and Noise Occupancy

- After alignment measured efficiency is $> 99.7\%$ for active modules
- Noise occupancy after masking of noisy pixels: $\sim 10^{-10}$
 - Fraction of masked pixels: 10^{-4}



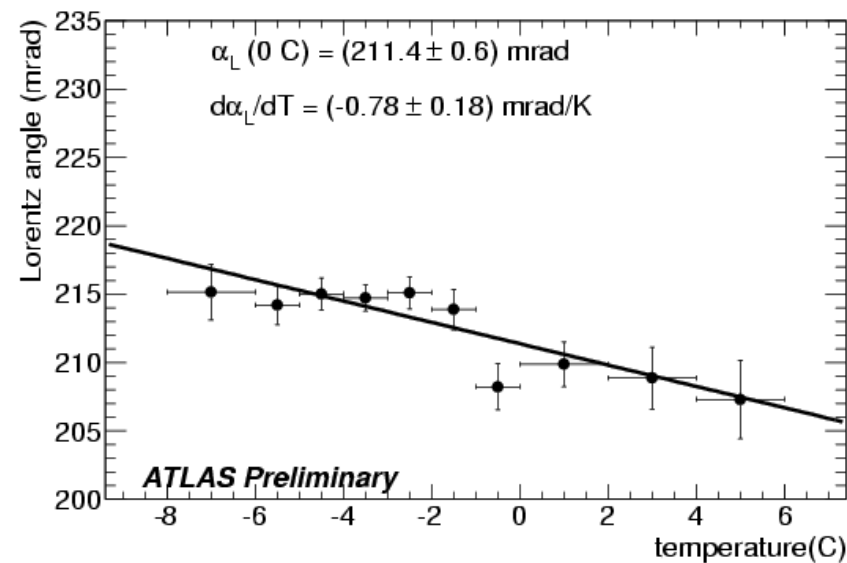
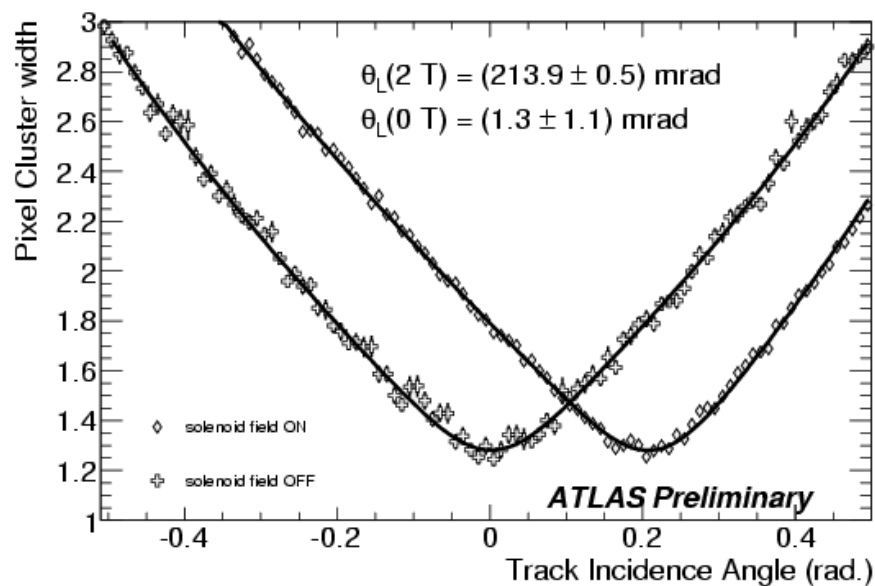
Zoom!



Lorentz Angle Measurement

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- Cluster size vs. track angle with and without magnetic field → Measurement of the Lorentz angle
- Measured value close to expected value (225 mrad)
- Theoretically expected dependence on mobility can be nicely seen when including modules of different temperature
 - ▣ Measured: (-0.78 ± 0.18) mrad/K, expected: -0.74 mrad/K



Status and expectations

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- The ATLAS Pixel Detector has been commissioned in a relatively short time
- Noise $\sim 200 e^-$, threshold tuned at $4000 e^-$ with a dispersion $\sim 40 e^-$, m.i.p. signal at $\sim 19000 e^-$
- The cosmic data taking has been extremely useful
 - ▣ Timing already in good shape: plan to reduce the readout window rapidly
 - ▣ Resolution $\sim 24 \mu\text{m}$ in the precision direction, efficiency $> 99.7 \%$
 - ▣ Noise occupancy $\sim 10^{-10}$, 10^{-4} fraction of masked pixels
- Starting from this year all cooling loops are operated
 - ▣ Modest amount of coolant leak: studying the effects under irradiation
- 1.6 % of the detector is not functional due to on-detector failures
- Tuning at lower thresholds for beam related studies

The Pixel Detector with $\geq 98\%$ working modules is ready for LHC

