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CRAB-CAVITY BEAM DYNAMICS ISSUES FOR AN LHC UPGRADE

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Abstract

Modern colliders bring into collision a large number of bunches per pulse or per turn to achieve a high luminosity. The long-range beam-beam effects arising from parasitic encounters at such colliders are mitigated by introducing a crossing angle. Under these conditions, crab cavities can be used to restore effective head-on collisions and thereby to increase the geometric luminosity. Crab cavities have been proposed for both linear and circular colliders. The crab cavities are rf cavities operated in a transverse dipole mode, which imparts on the beam particles a transverse kick that varies with the longitudinal position along the bunch. The use of crab cavities in the LHC may not only raise the luminosity, but it could also complicate the beam dynamics, e.g. crab cavities might not only cancel synchro-betatron resonances excited by the crossing angle but they could also excite new ones, they could reduce the dynamic aperture for off-momentum particles, they could degrade the collimation cleaning efficiency, and so on. In this note, we summarize the results of several preliminary studies of beam-dynamics challenges associated with crab cavities in the LHC.

1 Introduction

The Large Hadron Collider (LHC), now being constructed at CERN (European Organization for Nuclear Research), will start beam commissioning in 2008. The LHC has four Interaction Points (IPs). Its design luminosity is $10^{34} \text{ cm}^{-2}\text{s}^{-1}$ at the two high-luminosity proton-proton experiments ATLAS (located at IP1) and CMS (located at IP5), with a centre-of-mass energy of 14 TeV [1]. Studies aimed to further raising the LHC luminosity are being carried out since 2001, from 2004 onwards jointly by the European CARE-HHH network and by US-LARP. These studies developed a roadmap for increasing the luminosity of LHC by a factor of 10 above its design value, to $10^{35} \text{ cm}^{-2}\text{s}^{-1}$ in the second half of the next decade. Recently, three upgrade scenarios have crystallized from various considerations. These three scenarios are called the early-separation (ES) scheme, the full crab crossing (FCC) scheme, and the Large-Piwinski angle (LPA) scheme [2]. For the ES and FCC approaches, crab cavities are an essential ingredient of the upgrade. As usual, the crab cavities will render the effective Piwinski angle equal to zero, and minimize the geometric luminosity loss which otherwise arises from the crossing angle. By contrast, the LPA scheme relies on a non-zero Piwinski angle. It could only make use of crab cavities if the latter do not restore the beam-beam tune shift to the value obtained for a head-on collision.

The crab cavity gives rise to a z-dependent horizontal or vertical kick on the beam particles (depending in the crossing plane), as well as to a change in the longitudinal momentum (or energy). In MAD coordinates, the Hamiltonian to describe the “thin” crab cavity is [3]

$$H_{crab} = \frac{q\hat{V}}{p_s} \sin\left(\phi_s + \frac{\omega x_3}{c}\right) \cdot x_1 \quad (1)$$

where, H_{crab} denotes the Hamiltonian, V the crab-cavity voltage, q the particle charge, p_s the momentum of particle, ϕ_s the “synchronous” phase of the crab cavity rf wave, ω the angular frequency of the crab cavity, c the velocity of light, x_1 the horizontal coordinate (x) and x_3 the longitudinal coordinate (z), respectively.

The horizontal and longitudinal kicks from the crab cavity can be written as:

$$\Delta p_1 = -\frac{\partial H_{crab}}{\partial x_1} = -\frac{q\hat{V}}{p_s} \sin\left(\phi_s + \frac{\omega x_3}{c}\right) \quad (2)$$

$$\Delta p_3 = -\frac{\partial H_{crab}}{\partial x_3} = -\frac{q\hat{V}}{p_s} \cos\left(\phi_s + \frac{\omega x_3}{c}\right) \cdot \left(\frac{\omega}{c}\right) \cdot x_1 \quad (3)$$

In circular colliders, crab cavities may be configured according to either one of two schemes, namely as “local” or “global” crab cavities. In the local scheme, which corresponds to the original classical proposals of Palmer, Oide and Yokoya, a pair of crab cavities is placed at both sides of one IP, with the phase advance between crab cavity and IP optimized to be $\pi/2$ in the crossing plane; in that case, for each value of z the crab cavities act like a local bump and the z -dependent closed orbit in the other parts of the ring is not affected. The voltage needed for the first crab cavity (to “rotate” the bunch so as to be head-on at the interaction point) and for the second crab cavity (to rotate the bunch back) can be calculated using the formulae (4) and (5) below [4]. For the global scheme, the crab cavity is located somewhere around the ring, at a place which satisfies certain phase advance requirements which involve the total betatron tune in addition to the phase advance from crab cavity to the IP; in the global case, the z -dependent closed orbit will be changed all around the ring with crab cavity active, and the crab cavity voltage for the global scheme can be calculated from formula (6) below [4]. The bunch will be “tilted” and the tilt will exhibit some kind of oscillation all around the ring such that at the interaction point it will be head-on. The global scheme is the one which has actually been implemented at KEKB [5].

The rf voltage required for the two local crab cavities installed around one IP are determined as

$$V_1 = \frac{c \cdot E \cdot \tan(\theta/2)}{\omega \cdot \sqrt{\beta^* \beta_{crab}} \cdot \sin(\Delta\varphi)} \quad (4)$$

$$V_2 = -R_{22} \cdot V_1 \quad (5)$$

where V_1 denotes the voltage of the first crab cavity, V_2 the voltage of the second crab cavity, E the particle energy, θ the full crossing angle at the IP, β^* the beta function at the IP, β_{crab} the beta function at the crab cavity location, and $\Delta\varphi$ the phase advance between the IP and the first crab cavity. Lastly, R_{22} signifies the (2,2) element of the optical transport matrix from the first crab cavity to the second crab cavity.

The rf voltage for the global crab cavity is given by

$$V = \frac{c \cdot E \cdot \tan(\theta/2)}{\omega \cdot \sqrt{\beta^* \beta_{crab}}} \cdot \left| \frac{2 \sin(\pi Q)}{\cos(\Delta\varphi - \pi Q)} \right| \quad (6)$$

where V designates the voltage of the global crab cavity, and Q the betatron tune.

In the following, we will study the beam dynamics associated with the presence of a crab cavity for two kinds of LHC optics: the “nominal” LHC optics and the so-called “low-beta-max” optics [6]. Relevant parameters of these two optics are listed in Table 1. Both optics are studied for a beam at top energy. If not specifically mentioned, the nominal LHC optics and beam 1 are considered.

Table 1: Relevant parameters of the two LHC optics under study

parameter	symbol	nominal LHC optics	low-beta-max optics
protons per bunch	N_b [10^{11}]	1.15	1.15
rms bunch length	σ_z [cm]	7.55	7.55
rms energy spread	σ_e [10^{-4}]	1.1	1.1
beta function at IP1&5	β^* [m]	0.55	0.25
emittance	E [10^{-6} m]	3.75	3.75
full crossing angle	θ_c [μ rad]	285	381
horizontal beta function at s=0	β_x [m]	41	40
vertical beta function at s=0	β_y [m]	356	408

2 Horizontal crossing at IP5

In the LHC, the beams cross horizontally at IP5 and vertically at IP1. In this section we will consider the horizontal crossing case at IP5 only. Both the global crab cavity scheme and the local crab cavity scheme will be studied. This followed by some preliminary results of luminosity simulations with the code GUINEA-PIG and by a first assessment of the impact of crab cavities on the performance of the collimation system. In Section 3 the case of a vertical crossing at IP5 (for the real LHC the vertical crossing will be at IP1) will be examined. In all studies of this note, the starting point of the LHC sequence will be IP3, and tracking is performed with the well known-code SixTrack, which is the work horse of dynamic aperture studies at the LHC.

2.1 Global crab cavity for the nominal LHC optics

The location of the global crab cavity is selected to be upstream of IP5, outside D2, between Q4 and Q5. We selected the location there as beams 1 and 2 have reached their full arc separation and for local crab cavity case, to be studied afterwards, the phase advance to IP5 is near $\pi/2$ (the location of the global crab cavity is here taken to be the same as the first cavity location of the local crab cavity case; in reality the global cavity would more likely be installed in IR4 where the separation between the two beams is largest and cryogenics plus other needed infrastructure might be more easily available). The longitudinal distance between the selected position and IP5 is around 170~171 m, and the horizontal phase advance between IP5 and the global crab cavity is around 0.23~0.25. The detailed parameters of IP5 and of the crab cavity location are listed in Table 2. Using formulae (6), the voltage for an 800-MHz crab cavity is calculated to be 11.9087 MV. For 400-MHz cavity, the voltage value needs to be doubled.

Table 2: Parameters at IP5 and at the location of the global crab cavity

	Longitudinal S [m]	β_x [m]	β_y [m]	Phase advance x	Phase advance y
IP5	6664.6	0.55	0.55	15.4716	14.5941
crab cavity	6835.2	258	1547	15.7496	14.8553

Figure 1 shows the simulated longitudinal-horizontal bunch distribution at IP5 for three cases: with an 800-MHz crab cavity (CC), for a 400-MHz crab cavity, and assuming a totally linear kick crab cavity RF slope. From this figure it can be inferred that the linearity of the 400-MHz crab cavity is much better than that of the 800-MHz cavity. However,

the cavity size will be much larger for the 400-MHz cavity, which may make it difficult install such cavity on the beam line in the LHC tunnel.

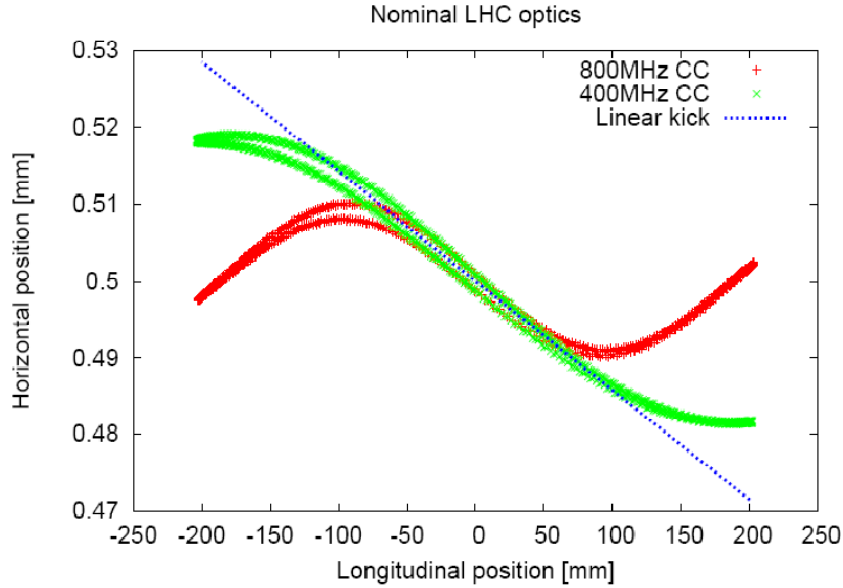


Figure 1: Bunch distribution at IP5 with global crab cavity

A sufficiently large dynamic aperture is important for achieving a good beam life time and a high integrated luminosity. Here we track two twin particles over 100,000 turns in SixTrack, to determine the minimum dynamic aperture. The beam-beam effect has not been taken into account. The initial relative momentum offset is set to 0.00027 (3/4 of the rf bucket; standard convention for all LHC dynamic-aperture studies). The results are shown in Fig. 2, which demonstrates that the 800-MHz crab cavity provides for a better dynamic aperture than the 400-MHz cavity. However, this need not always be true. A comparison of different cases will be shown in the following sections.

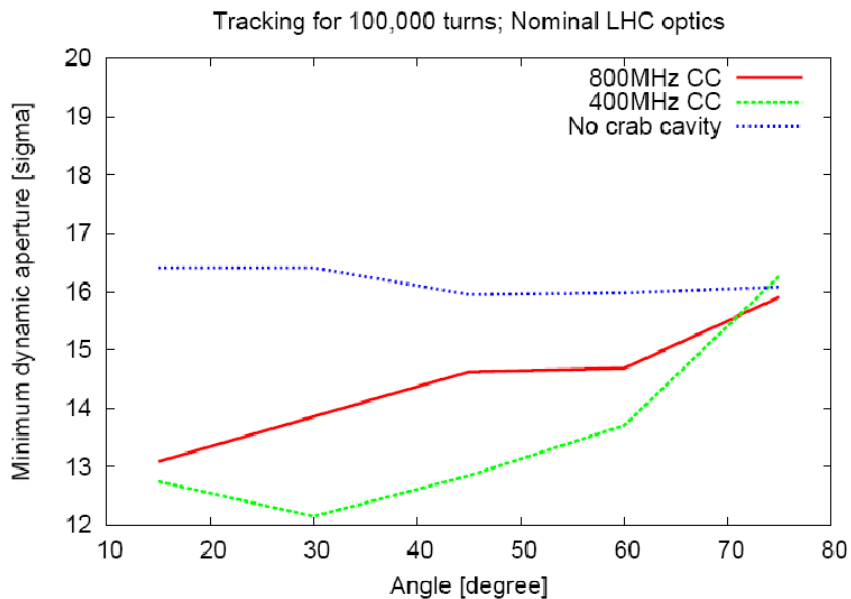


Figure 2: Dynamic aperture for the nominal LHC optics

For the purpose of our study, 168 BPMs are placed in the nominal LHC optics sequence file with equal physical separation between successive BPM elements. Then one particle with zero initial horizontal and vertical offset, and with energy offsets equal to 1, 2, 3 sigma are tracked for 1000 turns respectively. The horizontal peak to peak orbit is extracted from each set of BPM data and plotted in Figs. 3, 4 and 5. Comparing these three cases it shows that the 400-MHz global crab cavity introduces the largest disturbance of the horizontal orbit, while the 800-MHz case is much improved. The impact on the collimation system will be preliminarily analyzed in the following sections. The performances of the global crab cavity and the local crab cavity scheme will also be compared later.

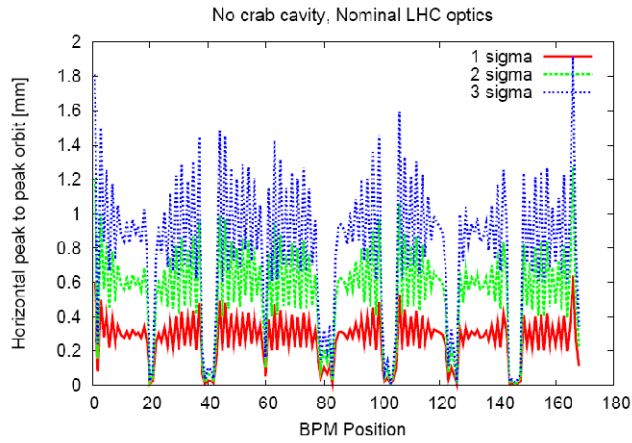


Figure 3: Horizontal peak-to-peak orbit for the nominal LHC optics without crab cavity for various momentum offsets

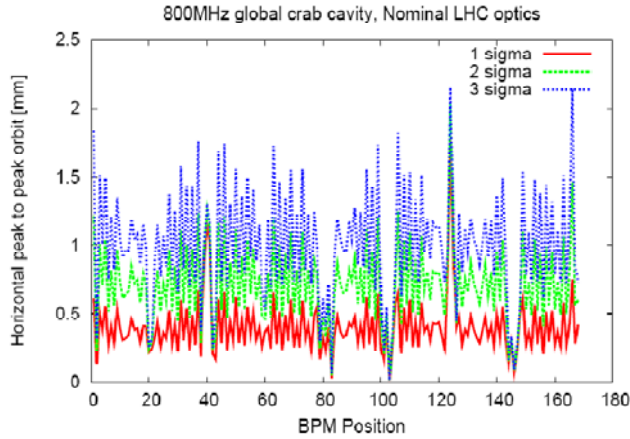


Figure 4: Horizontal peak-to-peak orbit for the nominal LHC optics with an 800-MHz global crab cavity for various initial momentum offsets

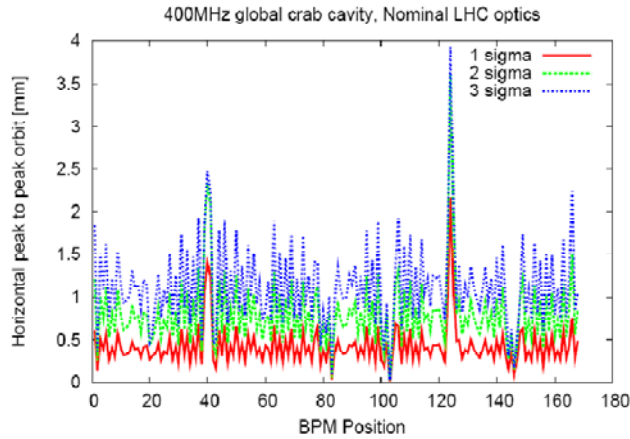


Figure 5: Horizontal peak-to-peak orbit for the nominal LHC optics with a 400-MHz global crab cavity for various initial momentum offsets

In Figure 5, there are two peaks, in IR5 (the first peak) and in IR1 (the second peak). They are due to the closed orbit disturbance induced by the crab cavities and the physics behind remains to be studied in detail.

Synchrotron resonances are caused by different transverse beam-beam forces at different longitudinal positions (synchrotron oscillation modulation) in the colliding bunches, an effect which is introduced by the transverse crossing angle at IP [7]. It was suggested that crab cavities could eliminate this kind of resonance for circular colliders. (It was also predicted – long ago - that crab-cavities could restore an effective head-on collision at the IP for both linear colliders [8] and circular colliders [9].) Here, without beam-beam effects, we simply compute the FFT of simulated BPM data by using the SUSSIX code, to obtain the spectrum shown in Fig. 6. A particle is tracked for 10,000 turns to minimize the influence of noise on the FFT results. Both the horizontal initial offset and the longitudinal initial offset are set to be one sigma. A detailed complementary study including beam-beam force is still ongoing. Three cases will be compared, namely only the “head-on” beam-beam effects with a crossing angle, the beam beam effects plus a crab cavity, and only crab cavity without beam-beam. We expect to see that for the case with beam-beam effects plus crab cavity the synchrotron resonance strength is the lowest.

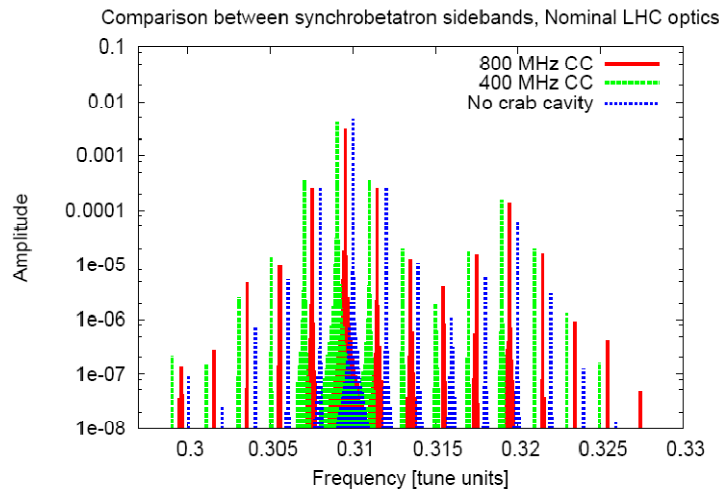


Figure 6: Comparison of synchrotron sidebands with two different global crab cavities and without any crab cavity for the nominal LHC optics; no beam-beam effects are included

To render the comparison without beam-beam effect clearer, in Fig. 6 the frequency of the 800-MHz global crab cavity case was shifted downward by -0.0005 tune units, and that for the 400-MHz case by -0.001 tune units. In Fig. 6 it can be seen that the fractional part of the horizontal tune is equal to 0.31 from the FFT, and the fractional part of the vertical tune equals 0.32, which are just the official tunes for the nominal LHC optics at top energy. Also, the amplitudes of the synchrotron resonances are larger with crab cavity than for the case with no crab cavity, especially considering higher order sidebands, which represent synchrotron resonances excited by the crab cavity. These resonances are generated because the crab cavity exerts a z -dependent horizontal kick on the particles, which is similar to the beam-beam crossing effect at the IP (and which should partially cancel the latter).

2.2 Local crab cavity for nominal LHC optics

For the local crab cavities, the location of the first crab cavity is chosen at exactly the same place as for the global crab cavity. We recall that this location is found upstream of IP5, outside D2, and between Q4 and Q5. The second crab cavity is placed symmetrically at the other side of the IP5. Relevant parameters are listed in Table 3.

Table 3 Parameters at IP5 and at local crab cavities

	Longitudinal S [m]	β_x [m]	β_y [m]	Phase advance x	Phase advance y
Crab cavity 1	6493.9	1523	255	15.2225	14.3283
IP5	6664.6	0.55	0.55	15.4716	14.5941
Crab cavity 2	6835.2	258	1547	15.7496	14.8553

Applying formulae (4) and (5), and also using the transfer matrix from crab cavity 1 to crab cavity 2, the 800-MHz cavity voltages are calculated to be 5.1051 MV and 2.6751 MV for the two cavities, respectively. Similar to Fig. 1, the simulated longitudinal-horizontal bunch distribution for the local crab cavity case is benchmarked with the linear kick model, in Fig. 7. Again the 400-MHz case clearly reveals a better linearity. Following the same analysis procedures as in Section 2.1, here we simply present the results for the dynamic aperture, for the horizontal peak-to-peak orbit, and for synchrotron sidebands, in Figs. 8, 9 and 10, respectively.

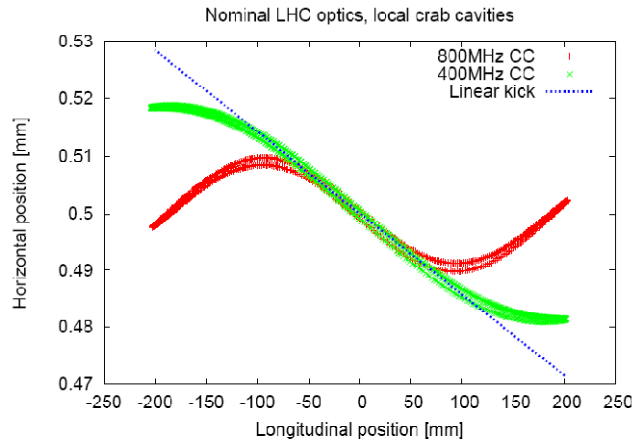


Figure 7: Bunch distribution at IP5 with local crab cavities

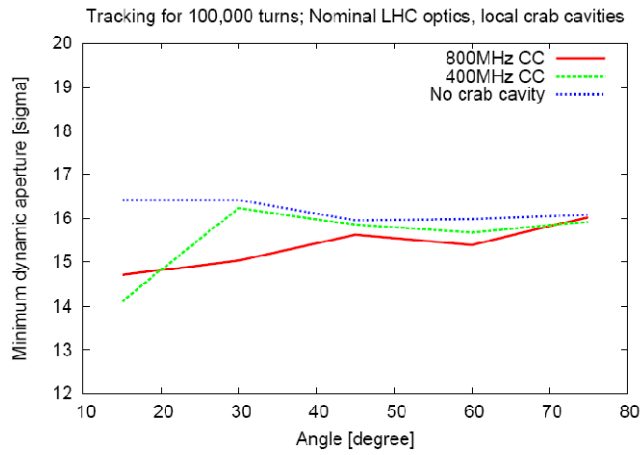


Figure 8: Dynamic aperture for the nominal LHC optics with local crab cavities

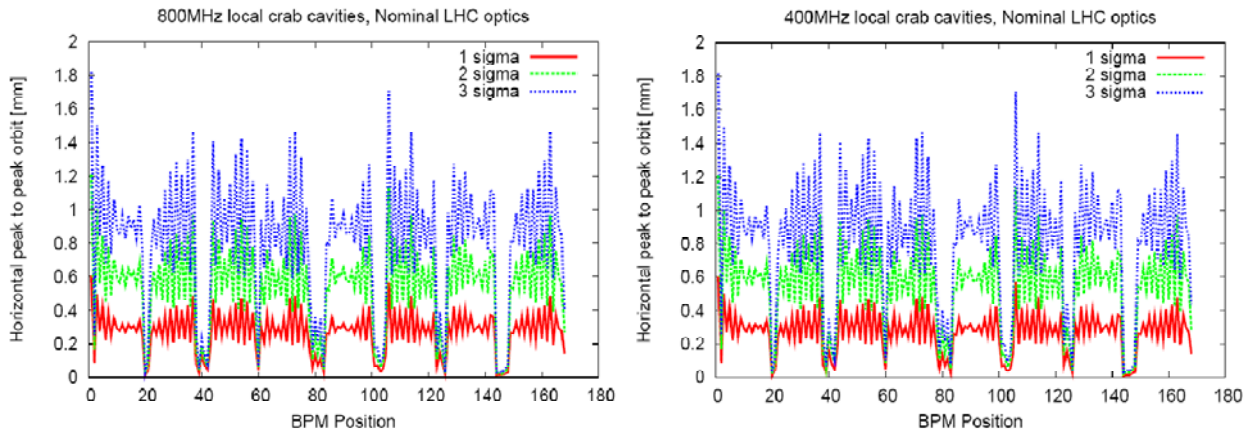


Figure 9: Horizontal peak-to-peak orbit for the nominal LHC optics with local crab cavities operating at 800 MHz (left) or 400 MHz (right).

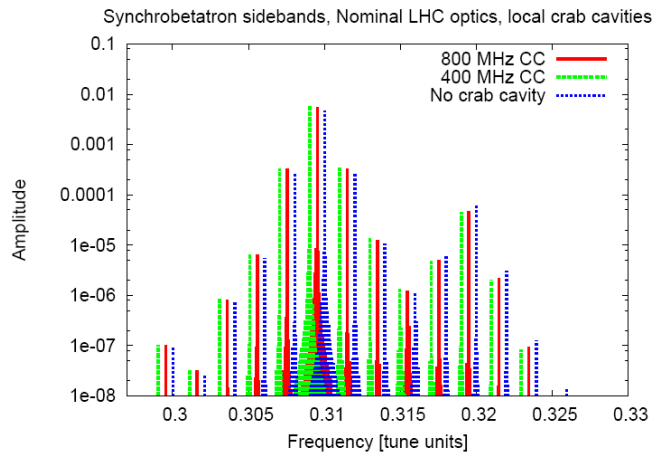


Figure 10: Comparison of synchrotron sidebands for the nominal LHC optics with local crab cavities

For the local crab cavity case, the dynamic aperture is much better than for the global crab cavities, at maximum 1~2 sigma degradation compared to the case without crab cavity. Figure 9 confirms that local crab cavities do not affect the closed orbit around the ring. Figure 10 shows more clearly that the amplitude of each sideband for these three cases is almost the same, which is different from the situation in Fig. 6 (the frequency shift added for plotting purposes is the same as in Figure 6).

2.3 Global crab cavity for the low-beta-max optics

For the low-beta-max optics, the crab cavities are placed at similar locations as for the nominal LHC optics, and the results we get are indeed also quite similar to the corresponding results for the nominal LHC optics. So we only briefly introduce the exact location of crab cavity and the other associated parameters. The location of the global crab cavity is selected to be upstream of IP5, outside D2, and between Q4 and Q5. The detailed parameters of IP5 and of the crab cavity location are listed in Table 4. Using formulae (6), the voltage for 800-MHz crab cavity is calculated to be 9.468 MV.

Table 4: Parameters at IP5 and at the location of the global crab cavity for the low-beta-max optics

	longitudinal S [m]	β_x [m]	β_y [m]	phase advance x	phase advance y
IP5	6664.6	0.25	0.25	15.6213	14.367
crab cavity	6835.7	1023	2685	15.86	14.5988

It is interesting to mention that for the low-beta-max optics, the horizontal peak-to-peak orbit exhibits four maxima around the ring, namely in the two triplets at both sides of IP1 and IP5.

2.4 Local crab cavities for the low-beta-max optics

The local crab cavity locations are also similar to those for the nominal LHC optics case. Some details are listed in Table 5, and the 800-MHz cavity voltages for crab cavity 1 and crab cavity 2 are calculated to be 4.9568 MV and 3.1006 MV, respectively.

Table 5: Parameters at IP5 and at the locations of the local crab cavities for the low-beta-max optics

	Longitudinal S [m]	β_x [m]	β_y [m]	Phase advance x	Phase advance y
Crab cavity 1	6493.4	2616.8	1079.6	15.3511	14.0901
IP5	6664.6	0.25	0.25	15.6213	14.367
Crab cavity 2	6835.7	1023	2685	15.86	14.5988

2.5 Comparison of dynamic aperture and beta-beating

The tracking conditions are the same as those in Section 2.1. Beam beam elements are added in the sequence file of Beam 2 with several longitudinal slices. At IP1 and IP5, both head-on and parasitic collisions are included. At IP2 and IP8, only head-on collisions are taken into account. The particles are tracked for 100,000 turns using SixTrack, with two error seeds. The minimum dynamic aperture for different conditions is listed in Table 6.

Table 6: Minimum dynamic aperture for different cases

Case	Nominal LHC optics [sigma]	Lowbetamax [sigma]
No beam-beam, No CC	16	15.9
Beam-beam (BB)	8.2	13
Global CC 800 (400) MHz	13 (12.1)	12.4 (14.3)
Global CC 800 (400) MHz + BB		
Local CC 800 (400) MHz	14.7 (14.1)	16 (15.5)
Local CC 800 (400) MHz + BB		

The off-momentum beta beating is widely studied in LHC, as it is very important for the collimation system, and the safe operation of the superconducting ring. Figure 11 compares the off-momentum beta beating for the nominal LHC optics and for the low-beta-max optics used in our study. It can be seen that for the nominal LHC optics, the tune is optimized for minimum off-momentum beta beating in the region between IP1 to IP5, while for the low-beta-max optics, which has a larger beating, it is the contrary.

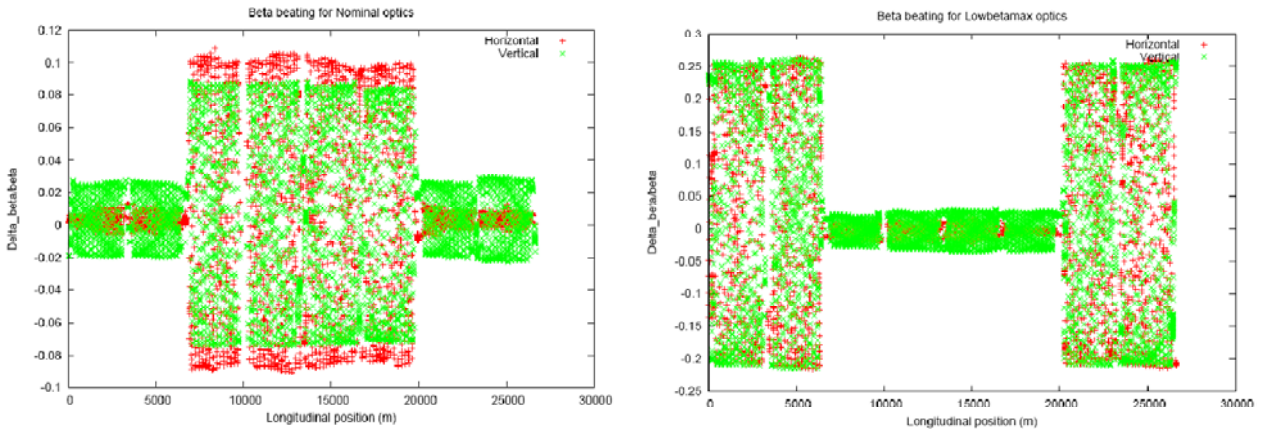


Figure 11: Off-momentum beta beating for the nominal LHC optics (left) and for the low-beta-max optics (right) with a relative momentum offset of 0.00027

Crab cavities introduce another kind of beta beating, which depends on the longitudinal position inside the bunch. Here we consider the global 800-MHz crab cavity cases discussed earlier in Sections 2.1 and 2.3. The effect of the crab cavity is modelled by adding a horizontal corrector (with z-dependent strength) at the location of the crab cavity in SIXTRACK. With this trick we can study this new kind of beta beating for a particle with one sigma longitudinal offset. According to the results shown in Fig. 12, the additional beta beating caused by the crab cavity is comparatively small with respect to the existing off-momentum beta beating. For the nominal LHC optics, it is less than ± 0.004 .

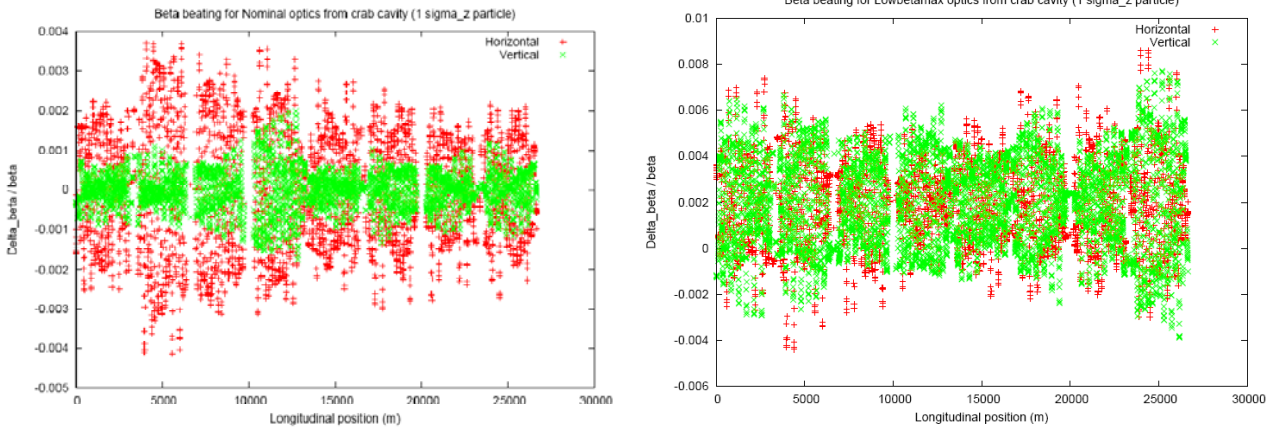


Figure 12: z-dependent “beta beating” due to the global crab cavity for the nominal LHC optics (left) and for the low-beta-max optics (right)

2.6 Luminosity simulation (IP1 or IP5) with GUINEA-PIG

GUINEA-PIG is a computer code for simulating the luminosity of electron-positron colliders [10]. Here we apply this code to a storage ring, in order to simulate the single-bunch geometric luminosity at the LHC. The total luminosity can then be calculated by multiplying with the bunch number 2808. For the LHC optics without crab cavity, two bunches at the IP1 or IP5, each consisting of 10,000 macro-particles, are generated with a Gaussian distribution in all six phase-space coordinates $(x, x', y, y', ct, \Delta p)$. The rms values for the Gaussian distributions are calculated from the design beam emittance, the design beam energy and the design beta functions at IP. It should be mentioned here that, as the proton bunches are much larger than the typical electron bunches in linear colliders, the “cut beam size” should be set to a large enough value in the code file “acc.dat”. To make the tracking results converge better, the longitudinal slice number of the bunch (“n_z” in the code) is chosen to be larger than for linear-collider simulations, say equal to 200. For the case with crab cavity, we assume that the phase advance between crab cavity and IP is exactly $\pi/2$, so that $\alpha=0$ at both the IP and at the crab-cavity location. Then, applying formula (2), a horizontal kick from the crab cavity is added for each particle and a new bunch distribution is obtained. Two example distributions in 3-D and 2-D at the IP are displayed in Figs. 13 and 14, respectively.

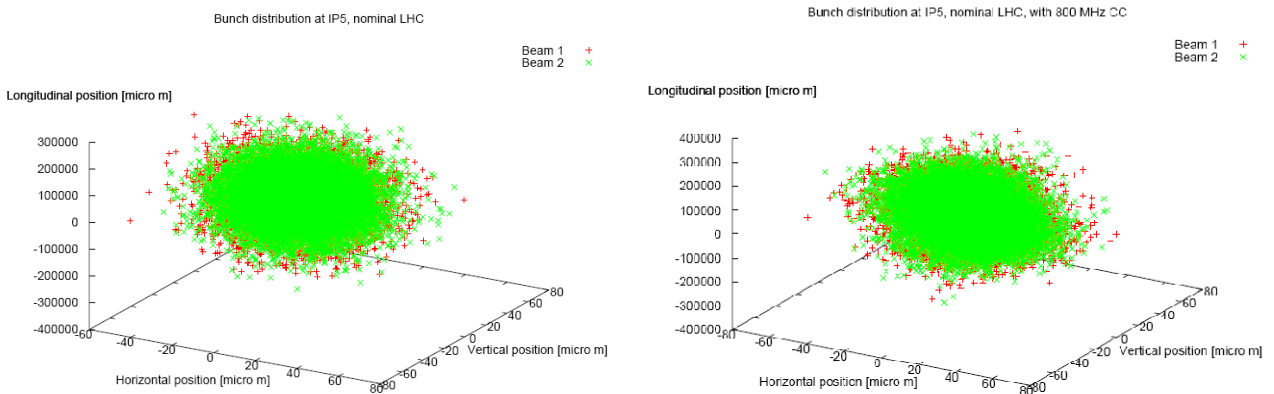


Figure 13: Distribution of the two colliding bunches for the nominal LHC without crab cavity (left) and with an additional 800-MHz crab cavity (right)

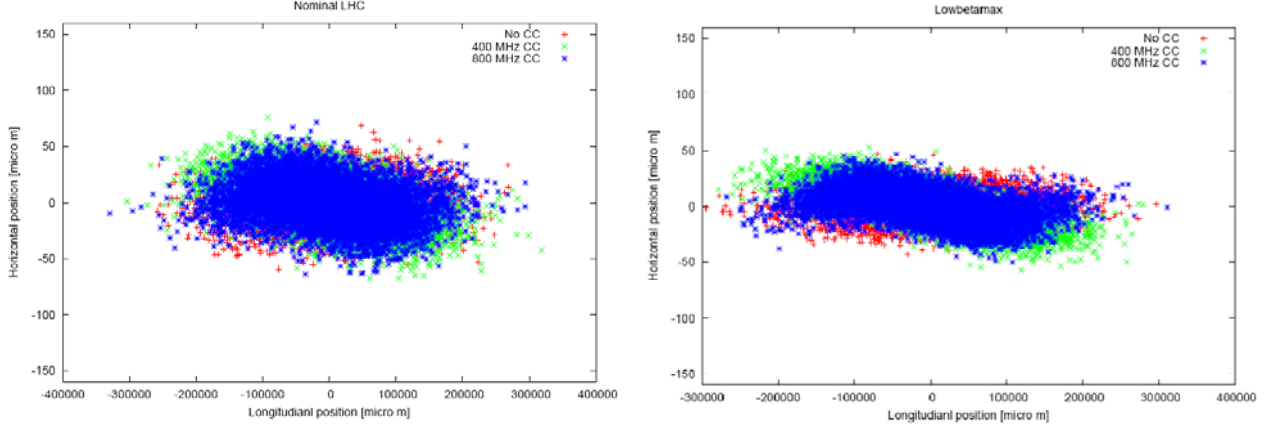


Figure 14: Bunch distribution for the nominal LHC optics (left) and for the low-beta-max optics (right) without crab cavity (red), with a 400-MHz crab cavity (green), and with an 800-MHz crab cavity (blue).

Simulation results for the LHC luminosity under different conditions are compiled in Table 7. For the nominal LHC optics without crab cavity, the simulated luminosity is in good agreement with the design value. A fictitious 20-MHz crab cavity is used to check the correctness of the transverse kick from the crab cavity. In the near future, more seeds will be simulated to determine the average luminosity with reduced error. A more precise tracking with the real optics parameters will be done later.

Table 7: LHC luminosity from GUINEA-PIG

	nominal LHC optics [$\text{cm}^{-2}\text{s}^{-1}$]	low-beta-max [$\text{cm}^{-2}\text{s}^{-1}$]
design (from LHC design report)	1.00E+34	
head-on collision	1.18E+34	2.5596E+34
normal crossing	1.00256E+34	1.50118E+34
crossing with local 20-MHz CC	1.1796E+34	2.4899E+34
crossing with local 400-MHz CC	1.1716E+34	2.42838E+34
crossing with local 800-MHz CC	1.1144E+34	2.146E+34

At the same time, we study the luminosity formulae with crab cavity analytically, to compare with the simulation results of GUINEA-PIG. We start from the original formulae shown in [11], where the bunch distribution is assumed to be Gaussian in the longitudinal direction and also in the plane of the crossing angle, namely

$$\rho_x(x) = \frac{1}{\sigma_x\sqrt{2\pi}} \exp\left(-\frac{x^2}{2\sigma_x^2}\right) \quad (7)$$

$$\rho_z(z) = \frac{1}{\sigma_z\sqrt{2\pi}} \exp\left(-\frac{z^2}{2\sigma_z^2}\right) \quad (8)$$

In comparison with the treatment in [11], here a more appropriate expression of the particles' coordinates is used, to correctly describe the kick from the crab cavity. Taking the rf sinusoidal wave shape of the crab cavity into account, and neglecting the x' dependent change of the particle energy (also assuming $R_{22}=0$), the coordinates of the two colliding beams can be written as

$$x_1 = x \cos \frac{\theta_c}{2} - s \sin \frac{\theta_c}{2} + \frac{1}{k_{cr}} \sin[k_{cr}(s - ct)] \sin \frac{\theta_c}{2} \quad (9)$$

$$x_2 = x \cos \frac{\theta_c}{2} + s \sin \frac{\theta_c}{2} - \frac{1}{k_{cr}} \sin[k_{cr}(s + ct)] \sin \frac{\theta_c}{2} \quad (10)$$

$$s_1 = s \cos \frac{\theta_c}{2} + x \sin \frac{\theta_c}{2} \quad (11)$$

$$s_2 = s \cos \frac{\theta_c}{2} - x \sin \frac{\theta_c}{2} \quad (12)$$

where θ_c denotes the crossing angle, c the velocity of light, and $k_{cr}=2\pi f_{crab}/c$ the rf wave number of crab cavity.

Writing the luminosity without crossing angle as

$$L_0 = \frac{N_b^2 f_{rev} n_b}{4\pi \sigma_x \sigma_y} \quad (13)$$

with N_b the number of particles inside a bunch, f_{rev} the revolution frequency, n_b the total number of bunches, σ_x the horizontal bunch size at IP, and σ_y the vertical bunch size at IP, the luminosity with finite crossing angle can be expressed (neglecting the hourglass effect) as

$$L = \frac{c N_b^2 f_{rev} n_b}{\sqrt{\pi} \sigma_y} \cos^2 \left(\frac{\theta_c}{2} \right) \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} \int_{-\infty}^{+\infty} \rho_x(x_1) \rho_z(s_1 - ct) \rho_x(x_2) \rho_z(s_2 + ct) dx ds dt \quad (14)$$

Using formulae (7)~(12) and formula (14), the luminosity reduction factor in comparison with the head-on collision case can be expressed as

$$R = \frac{\cos \frac{\theta_c}{2} c}{\pi \sigma_z^2} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \exp \left(-\frac{c^2 \cdot t^2}{\sigma_z^2} - \frac{s^2 \cdot \cos^2 \frac{\theta_c}{2}}{\sigma_z^2} - \frac{\sin^2 \frac{\theta_c}{2}}{4k_{cr}^2 \sigma_x^2} \cdot \left(2 + 4k_{cr}^2 \cdot s^2 - \cos(2k_{cr}(s - c \cdot t)) - \cos(2k_{cr}(s + c \cdot t)) - 8k_{cr} \cdot s \cdot \cos(k_c \cdot c \cdot t) \sin(k_c \cdot s) - 4\cos^2(k_c \cdot s) \cdot \sin^2(k_c \cdot c \cdot t) \right) \right) ds dt \quad (15)$$

Then we perform the numerical integration, using Mathematica, over s in the range between minus infinity and plus infinity, and over t in the range between minus $5\sigma_z$ and plus $5\sigma_z$, for two cases (400-MHz crab cavity and 800-MHz crab cavity) respectively. In Figure 15, the results from the analytical luminosity formula are presented by the four curves. Some simulation results from GUINEA-PIG are superimposed with dots. A good agreement between the analytical calculation and the Guinea-Pig simulation is found.

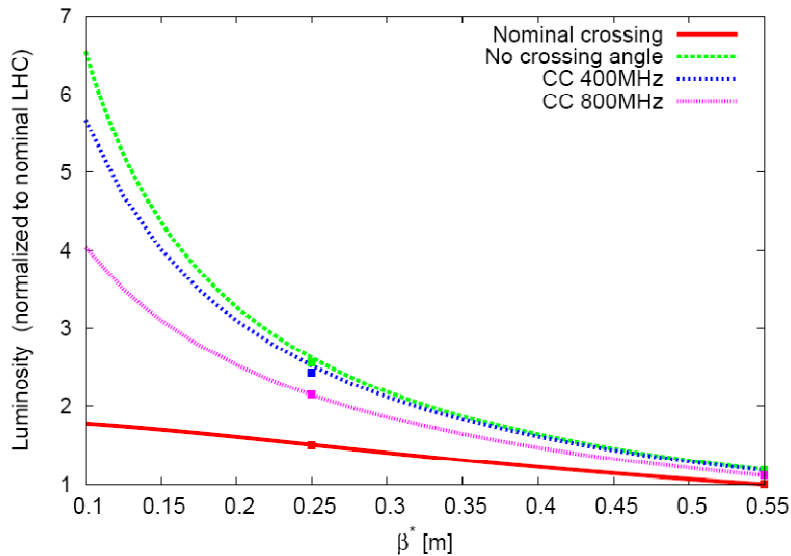


Figure 15: Normalized luminosity from analytical formulae (curves) and GUINEA-PIG simulations (dots)

2.7 Estimated impact on collimation system

LHC will run with super-conducting magnets and super-conducting RF cavities. Considering the huge beam power, a tiny loss of beam particles would result in a quench and may even damage the accelerator. The LHC collimation system is designed to protect the accelerator and to absorb the beam halo outside of a specified transverse beam size. There are two main regions where LHC collimators are concentrated: IR3 (momentum cleaning) and IR7 (betatron cleaning). Several kinds of collimators exist in the LHC. Among these the two main groups are the primary collimators and the secondary collimators. The difference between the jaw opening of the primary and secondary collimators is one beam sigma. Given the condition that the secondary collimators should never be hit by the primary beam halo, the closed orbit shall be strictly controlled in order to set the right jaw opening of collimators. As the global crab cavity could change the particle's z-dependent closed orbit, it will be interesting to see how the crab cavity will affect the collimation system, especially the difference of the closed orbit between the primary collimator and the next (or a few) secondary collimator(s). Here, some preliminary estimation results will be reported on the crab cavity impact on the collimation system. A more detailed study, addressing loss maps along the ring, heat deposition and collimation inefficiencies will be performed with the help of the collimation team.

Figure 16 shows the closed orbit at different z offsets for the nominal LHC optics with the 800-MHz global crab cavity introduced in Section 2.1. Single particles with zero transverse offset and only longitudinal z offset are tracked to get their closed orbit, again replacing the crab cavity by a horizontal corrector with the equivalent kick.

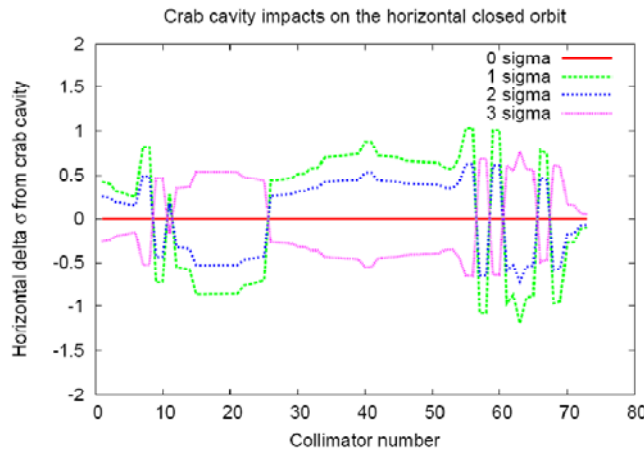


Figure 16: Closed orbit induced by the crab cavity

This closed orbit is checked against tracking results obtained with a real crab cavity in the sequence file. In the second case, a particle with the same initial condition as before is tracked for 1000 turns without the main RF so that the particle's longitudinal coordinate is approximately frozen. The horizontal position of this particle at each collimator for each turn is plotted together with the closed orbit (shown in Figure 16) at that collimator, in Fig. 17. The longitudinal position of this particle at the crab cavity location is checked to be nearly constant over this number of turns, and the energy change is seen to increase linearly with the number of turns. Figure 17 illustrates the case of one particle with zero transverse offset and 1 sigma longitudinal offset. Other cases are similar.

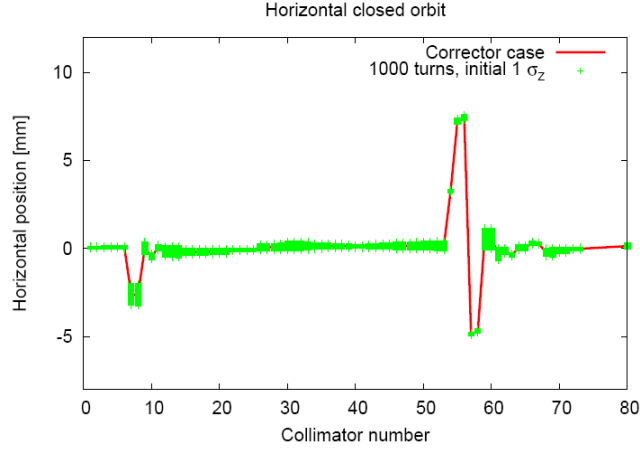


Figure 17: Comparison between closed orbit for a static corrector and oscillation average with real crab cavity

If this particle is launched also with a nonzero initial horizontal offset, this latter offset only adds to the amplitude of the betatron oscillations while the average value should still be the closed orbit shown in Fig. 16. Figure 18 confirms that this is the case.

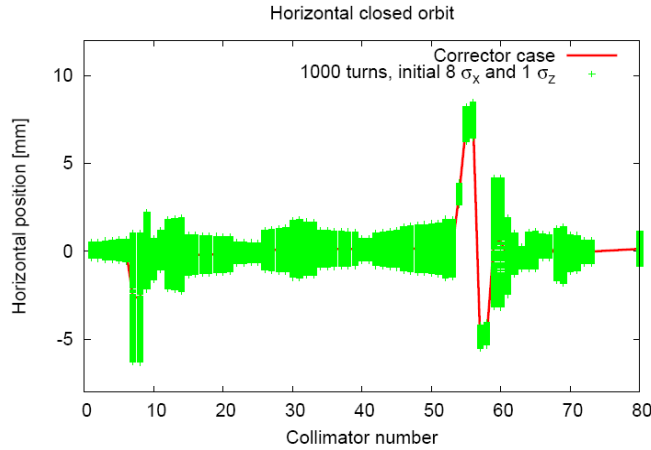


Figure 18: Comparison between closed orbit and oscillation (including betatron oscillation)

3 Vertical crossing at IP5

For the purpose of illustration, we assume to have a vertical beam crossing at IP5 (with the same crossing angle as for the horizontal case), and repeat the work done in Section 2. The locations of the global and local crab cavities are chosen exactly the same as for the horizontal crossing case. Here we just list the crab cavity voltage and the dynamic aperture tracking results. In Table 8 the required crab cavity voltage is listed, for the 285 micro-radian vertical crossing angle of the nominal LHC optics and a crossing angle of 381 micro-radians for the low-beta-max optics.

Table 8: Crab cavity voltage for different cases

	nominal LHC optics [MV]	low-beta-max optics [MV]
Global-800 MHz crab cavity	4.3283	5.8016
local 800-MHz, crab cavity one	2.0675	3.1122
local 800-MHz, crab cavity two	6.568	4.9087

Table 9 lists the results of dynamic-aperture studies for the cases with vertical crab cavities. The tracking conditions are the same as those described in Section 2.1.

Table 9: Minimum dynamic aperture for different cases

case	nominal LHC optics [σ]	low-beta-max [σ]
no beam-beam, no CC	16	15.9
Global CC 800 (400) MHz	14.2 (13.9)	12.7 (14)
local CC 800 (400) MHz	14.1 (14)	15.5 (16)

Comments

As a next step, some items discussed in this note should be explored in greater detail, such as the impact of crab cavities on the collimation system, and the interplay of a crab cavity with beam-beam effects (tune shift, synchrotron oscillations, luminosity). Also, a specific realistic location for a global crab cavity must be selected, probably in IR4, and the LHC optics may need to be adjusted for smaller crab cavity voltage (e.g. higher beta function at the crab cavity) and for a better performance with crab cavity.

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