Neutrinos from spin dynam ics

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A bstract

We conjecture that neutrino physics might correspond to the spontaneous m agnetisation phase of an Ising-like spin m odel interaction coupled to neutrino chirality which operates at scales close to the Planck mass. We argue that this scenario leads to a simple extension of the Standard M odel with no additional param eters that dynam ically generates parity violation and spontaneous symmetry breaking for the gauge bosons which couple to the neutrino. $_{ew}^2 = 2M$ where $_{ew}$ is the electrow eak The neutrino m ass in them odel is m scale and M is the scale of the \spin-spin" interaction. For the ground state of the model the free energy density corresponding to the cosm ological constant 4 is

, consistent with observation. m

1 Introduction

The neutrino sector is one of the most puzzling in particle physics. Only left-handed neutrinos participate in weak interactions. Recent experim ents have revealed oscillations between the three fam ilies of neutrino plus sm all neutrino m asses, with the heaviest neutrino m ass 0:05 eV, much less than the masses of the charged leptons and quarks. The mass of the lightest neutrino is presently not well constrained although for a norm al hierarchy of neutrino masses with m₁ m₂ m_3 one nds 0:008 eV { for recent reviews see Refs.[1, 2, 3, 4, 5]. Possible exm 2 m₁ planations for the sm all neutrino m ass involve either right-handed sterile neutrinos (together with som e additional \new physics" to suppress the mass relative to the charged leptons) or Majorana mass terms with local coupling to scalar Higgs elds using the see saw mechanism [6] to push the mass of the right-handed neutrinos to a very high scale, thus connecting the neutrino sector with new physics at much higher mass scales.

It is interesting to ask whether there m ight be an alternative explanation: C an we construct a dynam ical m echanism which yields observed neutrino physics? Is there an analogue situation in other branches of physics that one m ight hope to learn from ?

Here we investigate a possible analogy with the Ising model of statistical mechanics and, more generally, spin glass. Suppose that we associate the \spins" in the Ising model with neutrino chirality and the \internal energy per spin" with the neutrino mass. The ground state of the Ising model exhibits spontaneous magnetisation where all the \spins" line up; the \internal energy per spin" and the free energy density of the spin system go to zero.

This observation suggests the possibility that, perhaps, neutrino physics m ight arise from collective spin m odel phenom ena at a large scale, logarithm ically close to the P lanck m ass. E lectrow eak interactions m ight be included as an \in purity" in the m odel: the ! eW process corresponds to a small but nite probability for one of the \spins" to turn o if the Ising interaction couples just to neutrinos. The spontaneous magnetisation phase then exhibits parity violation as the right-handed neutrino decouples from the physics. We argue that it also exhibits spontaneous breaking of the SU (2) gauge symmetry coupled to the neutrino. The neutrino m ass m in this picture is expected to be about $\frac{2}{eW} = 2M$ where eW is the electrow eak scale and M is the scale of the spin m odel interaction. The free energy density of the spin system behaves like a cosm obgical constant and for the low energy spontaneous m agnetisation phase is $\frac{1}{4}$ m + O (kT), where kT 0:0002 eV for the CM B tem perature of free space.

2 Ising model dynamics

We rst brie y outline the basics of the Ising model (for reviews see [7,8]) and then discuss similarities and possible application to neutrino physics. The Ising model uses a spin lattice to study ferrom agnetism for a spin system in therm al equilibrium. A pplications include crystals, lattice gases and spin glass. One assigns a \spin" (= 1) to each site and introduces a nearest neighbour \spin-spin" interaction. In two dimensions the Ham iltonian reads

$$H = J \begin{pmatrix} X \\ i; j & i+1; j + i; j+1 & i; j \end{pmatrix} \qquad \begin{pmatrix} X \\ h & i; j \end{pmatrix}$$
(1)

Here J is the bond energy and h = B where B denotes any external magnetic eld and is the magnetic moment. (In this paper we take h = B = 0.) One sum s over the possible spins _{ij}. Physical observables are calculated through the partition function

$$Z = \sum_{i=1}^{X} \exp(H_i):$$
 (2)

Here = 1=kT where k is Boltzm ann's constant and T is the tem perature; k = $1.38 \quad 10^{23}$ JK ¹ = $8.617 \quad 10^{5}$ eV K ¹ and kT j_{B00K} = $[38.68]^{1}$ eV. We can norm alise the energy by adding a constant so that neighbouring parallel spins give zero contribution. Then, the only positive contribution to the energy will be from neighbouring disjoint spins of 2J and the probability for that will be exp(2 J). Once a magnetisation direction is selected, it remains stable because of the in nite num ber of degrees of freedom in the therm odynam ic lim it.

The \internal energy per spin" corresponding to the Ham iltonian in Eq.(1) is

$$(J) = 2J \tanh(2 J) + \frac{K}{d} \frac{dK}{d} \int_{0}^{2} d \frac{\sin^{2}}{(1+)}$$
(3)

where

$$K = \frac{2}{\cosh(2 J) \coth(2 J)}$$
(4)

and

$$= \frac{q}{1 \quad K^2 \sin^2} :$$
 (5)

The \free energy per spin" or free energy density for the system is

F () =
$$\frac{1}{2} \ln (2 \cosh 2 J) + \frac{1}{2} \int_{0}^{2} d \ln \frac{1}{2} (1 + \frac{q}{1 - K^{2} \sin^{2}}) :$$
 (6)

The internal energy and the free energy are related through

$$= \frac{2}{2} (F)$$
$$= F + TS$$
(7)

where S is the entropy. The pressure for the spin system is p = F. The Ising model has a second order phase transition and exhibits spontaneous magnetisation. In two dimensions the critical coupling $(J)_c$ is determined through the equation fsinh 2($J)_c$ = 1g. For values of (J) (J) the magnetisation per spin is M where

$$M = 1 [\sinh(2 J)]^4 = (8)$$

The magnetisation vanishes for (J) < (J)_c. In the limit J ! 1 , one nds

$$(J) = 2J \quad 24J \exp(8 J) + :::$$
 (9)

$$F = 2J + 3kT \exp(8 J) + ...$$
 (10)

and

$$M = 1 \quad 2\exp((8 J) + :::$$
 (11)

As advertised above, the zero-point energy is then renorm alised by adding + 2J to

(J) and F (or 2JN to the Ham iltonian where N is the number of spin sites) so that the \internal energy per spin" and free energy density vanish in the ground state with spontaneous magnetisation: (1) = 0. In general, the critical coupling depends on the number of dimensions. For 1, 2, 3 and 4 Euclidean dimensions the critical coupling (J)_c is 1, 0.441, 0.167 and 0.150 respectively.

O ne can extend the model to spin glasses [9] by introducing a probability distribution over the parameters of the model, e.g. the bond energies J and the magnetic moments.

3 Neutrinos

How can we construct an analogy with neutrinos and particle physics?

First, the Ising-like interaction itselfm ust be non-gauged otherw ise it will average to zero and there will be no spontaneous symmetry breaking and no spontaneous magnetisation (see e.g. page 51 of [7]).

Second, it is necessary to set a mass scale for J. If the Ising model analogy is to have connection with particle physics it is in portant to note that the coupling constant for the \spin-spin" interaction is proportional to the mass scale J. It therefore cannot correspond to a renormalisable interaction suggesting that uctuations around the scale J occur only near the extrem e high-energy or high-tem perature lim it of particle physics near the Planck mass M. We consider the elect of taking J = +M. The combination J is then very large making it almost certain that, if the analogy is applicable, the spontaneous magnetisation phase is the one relevant to particle physics phenomiena. (Fermion generations in particle physics might be a further hint at some kind of spin related dynamics at a very high mass scale.) The

exponential suppression factor e 2 ^J ensures that uctuations associated with the Ising-like interaction are negligible, thus preserving renorm alisability for all practical purposes.

M otivated by these observations, suppose we start with a gauge theory based on

coupled to quarks and leptons with no chiral dependent couplings and unbroken local gauge invariance. We then turn on the non-gauged Ising interaction coupled just to the neutrino in the upper component of the SU (2) isodoublet with the coupling J M _s; _{ew}; (the QCD, SU (2) weak and QED couplings). It seems reasonable that the Ising interaction here exhibits the same two-phase picture with spontaneous magnetisation. Then, in the symmetric phase where $J < (J)_c$ the theory is symmetric under exchange of left and right handed neutrino chiralities and we have unbroken local gauge invariance. In the spontaneous magnetisation phase the neutrino vacuum is \spin"-polarised, a choice of chirality is made and the right-handed neutrino decouples from the physics. Parity is spontaneously broken and the gauge theory coupled to the leptons becomes

$$SU(2)_{L} U(1)$$
 (13)

It is reasonable to believe that the SU (2) gauge symmetry coupled to the neutrino is now spontaneously broken. To understand how the W and Z⁰ gauge bosons might acquire mass, rst consider the issue of con nement. Before we turn on the Ising interaction we have QCD quark and gluon con nement plus a vector SU (2) gauge theory with electron and neutrino con nem ent. C on nem ent is intim ately connected with dynam icalchiral symmetry breaking and the generation of Dirac mass terms for constituent quarks and (at this stage) analogue constituent leptons. (See R ef. [10] for a recent discussion in QCD¹ and Ref. [11] for an overview how this connection is in plemented in the phenomenological Bag model of con nement.) Next turn on the Ising interaction and go to the spontaneous magnetisation phase. The righthanded neutrino decouples from the physics: the scalar chiral condensate for the neutrino \m elts" and the con ning solution for the neutrino should disappear. The left-handed neutrino will want to escape the con nem ent radius whereas the SU (2) gauge bosons will want to remain con ned { in contradiction to the fundam ental SU (2) symmetry. Some modication of the propagators must happen, viz. mass generation for the gauge bosons and the transition from the con nem ent to Higgs phases of the model so that the Coulom b force is replaced by a force of nite range

 $^{^1{\}rm P}\,{\rm ure}\,{\rm Y}\,{\rm ang}$ -M ills theory and Yang-M ills theory coupled to ferm ions are both con ning theories but the mechanism is di erent for each. Recent calculations [10] of quenched QCD in the covariant Landau gauge suggest that if chiral sym metry is restored, then the quark con nem ent solution disappears.

with nite mass scale and the issues associated with infrared slavery are avoided 2 . In this scenario experimental investigations of electroweak symmetry breaking will be probing fundamental properties of strong coupling dynamics. The W and Z⁰ gauge bosons which couple to the neutrino are massive and the QED photon and QCD gluons are massless.

W hat about the electroweak scale? Is there anything which stabilises it? First, in this scenario the SU (2) gauge symmetry coupled to the neutrino is dynamically broken. Second, dynamics associated with the Ising \spin" interaction are suppressed in the spontaneous magnetisation phase by the exponential factor e ^M where is a nite inverse mass-scale like an inverse small-tem perature factor or a mass scale typical of laboratory experiments, suggesting a natural hierarchy of scales. Further, the enorm ous excitation energy for right-handed neutrinos M rem inds us of the in nite energy required to excite a free-quark in QCD because of con nement. The in nite free-quark excitation energy compares with the nite chiral symmetry breaking scale F 100 M eV. Likewise, the electroweak mass-scale

250 G eV is much less than the right-handed neutrino excitation energy.

W hat about nite neutrino m asses? W eak interactions m ean that we have two basic scales in the problem : J M and the electrow eak scale $_{ew}$ induced by spontaneous symmetry breaking. (R ealistic accessible neutrino kinetic energies will be m uch less than J M and hence a m inor correction to the total energy and H am iltonian.) First, m ake the usual assumption that electrow eak symmetry breaking generates D irac m ass terms for fermions which participate in electrow eak interactions, e.g. as \in purities" in the \spin" system . Next, suppose that we approximate the two-phase spin-m agnetisation system by left-handed and right-handed neutrinos with M a jorana m ass terms so that the right-handed neutrino appears only at scales

 $O\ (M\)$ and electroweak processes contribute a regular $D\ irac\,m\,ass\,term$. Then the <code>\internalenergies</code> per spin" read in <code>m</code> atrix form as

where the rst row and rst column refer to the left-handed states of the neutrino and the second row and second column refer to the right-handed states. D is gonalising this matrix for M_{ew} gives the light mass eigenvalue

m
$$\frac{2}{ew} = 2M$$
 (15)

after the usual chiral rotation. Substituting the values of the P lanck m ass M $_{P1}$ hc=G $_{N}$ 1:2 10° G eV and the electroweak scale $_{ew}$ 250 G eV into Eq.(15)

²A sim ilar e ect occurs in 1+1 dim ensional gauge theories such as the Schwingerm odel coupled to dynam ical ferm ions: con nem ent gives way to Higgs phenom ena if the ferm ion mass is set exactly to zero [12].

gives m $3 10^6$ eV, which is plausible for the mass of the lightest neutrino and respectable given the simple approximations used above.

The matrix in Eq.(14) looks like the see saw mechanism result [6] although the fundam ental physics is quite di erent. In the see-saw picture the left-handed and right-handed com ponents of a four-state D irac neutrino are split by M a prana m ass term s involving coupling to scalar Higgs elds into a pair of two-state M a prana neutrinos with di erent m asses. The m asses of these left-handed light and right- $^{2}_{\text{ev}} = M_{\text{D}}$ where m handed heavy Majorana neutrinos are related through m is the mass of the light neutrino and M $_{\rm D}$ is both the \new physics" scale and the mass of the heavy neutrino. For a light neutrino mass 0:05 eV this relation gives 10^{15} GeV, which is not so far (on a logarithm ic scale) from the Planck mass Мъ $M_{Pl} = 12$ 10^{9} G eV. In the two-phase picture suggested here the tiny mass for the neutrino originates from collective \spin dynam ics" near the P lanck scale instead of through local Yukawa couplings to elementary scalar Higgs elds. The connection to the see-saw matrix (14) suggests that, perhaps, the neutrino in this picture should be M a jorana and therefore have no vector current: its electric charge should vanish.

It is interesting that the vacuum energy density corresponding to the cosm o- $\log cal constant$ 0:002 eV [13] has a similar num erical value to the range of possible light neutrino m asses [2], prom pting the question whether related underlying dynam ics m ight be at work? In the Ising m odel the free energy density for the spin system (or energy available for work) is related to the internal energy per spin through Eq.(7): = Q(F)=Q. It follows that F $^{2}_{\text{\tiny CMT}} = 2M + O (kT)$ where T is the tem perature of the system, which for the present Universe is the CMB tem perature 2.73K or kT 0:0002 eV. The free energy density is suppressed by the \spin dynam ics" which generate the spontaneous magnetisation and which \spin"-polarise the neutrino vacuum . The small nite value re ects the relatively ew = 2M jscalariew god, associated with the electroweak sm all scalar com ponent, and QCD scales which is induced in an otherwise \spin"-polarised total vacuum, viz.

jvacuum i_{total} ($_{ew}$ = 2M) jscalar $i_{ew, gcd}$ + jpolarised i (16)

E lectroweak and QCD interactions couple to the scalar condensates associated with the scales $_{ew}$ and $_{qcd}$ whereas the total vacuum in this picture is dominated by the \spin"-polarised neutrino component with zero free-energy density. Without the \spin-polarised" component the free-energy density would be $^{\frac{1}{4}}$ $_{ew}$. The positive sign for the free-energy density corresponds to negative pressure.

In conclusion, there are clear sim ilarities between neutrino phenom enology and spontaneous magnetization in the Ising model. This suggests the conjecture that, perhaps, neutrinos are associated with the spontaneous magnetisation phase of a spin model interaction which operates at scales close to the Planck mass. The

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phase transition associated with the spontaneous magnetisation of neutrino chirality in the spin system would, in turn, lead to parity violation and, we have argued, also spontaneous breaking of the SU (2) gauge symmetry coupled to the neutrino. Further, there would be a rapid drop in the potential governing the electrice neutrino mass or \energy per spin" from a value close to M at temperatures kT M to a value $\frac{2}{e_W} = 2M$ in the ground state which might be connected to the potential needed for in ation. There is no elementary scalar eld in them odel. It is interesting that the observed cosm ological constant corresponds to a vacuum energy density $\frac{1}{4}$ com parable with the expected value of the light neutrino mass. For the spin system the corresponding free energy density goes as F m + O (kT).

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