#### Technical University of Denmark



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# Spectroscopy for Industrial Applications: High-Temperature Processes

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## DTU Chemical Engineering

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From Lab to Field

## Outline

- Hot flow gas cell and FTIR/UV optical set up
- A road to In Situ measurements:
  - NH3 spectroscopy at high-temperatures: band assignment and spectra modelling
  - o NH3/H2O field measurements at a pilot scale 6MW gasifier
  - Phenol major trace gas from PAH's in low temperature gasification
  - Temperature-dependent UV absorption cross-sections
  - $\circ~$  Why In Situ measurements are important: comparison with "standard" tools
- How planets meet the Earth
- Conclusions

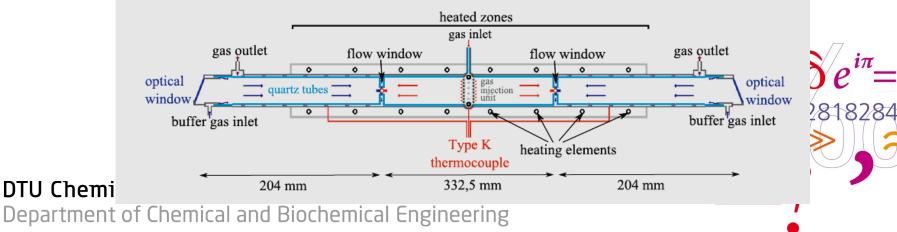
## DTU Chemical Engineering



Lab (Home work) NH3/Phenol: experimental set up



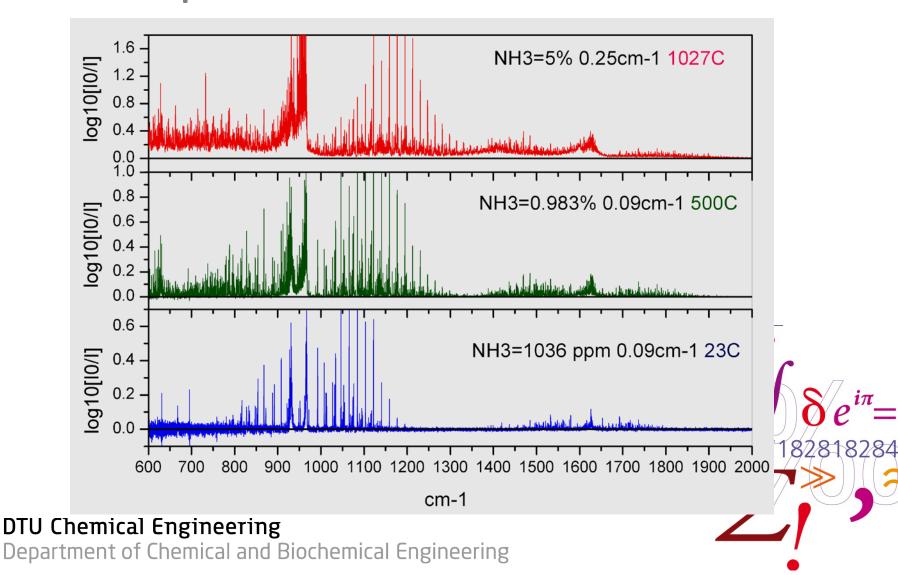
- 3-zones flow gas cell for corrosive gases;
- No internal windows;
- Stable uniform T-profile (±1.8C);
- Tmax= 525C
- L=33.25 cm
- P=1 bar
- suitable for UV-FIR optical measurements
- more details: H. Grosch et al. JQSRT 130 (2013) 392-399
- FTIR Spectrometer (Agilent 660), 0.09 cm-1
- an IR light source (up to 1500C)
- UV spectrometer (Acton 250i/CCD), 0.019 nm
- a highly stable D2-lamp



Lab (Home work)

# NH3 FTIR absorption spectra: changes with T

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# Lab (Home work)NH3 spectroscopy: line assignments, new results

#### Table 1: Lines assigned to previously observed bands

	Lines Assigned This Work				Previously measured				
Band	J <sub>max</sub>	K <sub>max</sub>	Frequency Range cm <sup>-1</sup>	Number of Lines	J <sub>max</sub>	K <sub>max</sub>	Number of Lines	Reference	
V <sub>4</sub>	17	17	1290 - 1868	277	15	15	1663	Cottaz 2000	500C
V <sub>2</sub>	20	20	634 - 1333	385	23	20	177	Yu 2010	1 bar
v <sub>2</sub> +v <sub>4</sub> -v <sub>2</sub>	12	12	1412 - 1818	83	10		384	Cottaz 2001	0.09cm-1
2v <sub>2</sub>	16	15	1407 - 1870	43	15	15	403	Cottaz 2000	
2v <sub>2</sub> -v <sub>2</sub>	18	18	607 - 1236	180	10	10	32	Singh 1988	
3v <sub>2</sub> -v <sub>2</sub>	12	12	1104 - 1652	18	10		132	Cottaz 2001	

#### Table 2: Lines assigned to previously unobserved bands\* with 10 or more lines assigned in this work.

		7				
	Band	J <sub>max</sub>	K <sub>max</sub>	Frequency Range cm <sup>-1</sup>	Number of Lines	$\Omega \int S e^{i\pi} =$
	v <sub>4</sub> -v <sub>2</sub>	11	11	622 - 1013	20	
	2v <sub>4</sub> -v <sub>4</sub>	9	5	1430 - 1792	10	{2.71828-18284
	$2v_4^2 - v_4$	8	5	1420 - 1805	10	
<b>DTU Che</b> i Departme	v <sub>2</sub> +v <sub>4</sub> -v <sub>4</sub>	13	13	680 - 1270	77	
	30-20-	14	12	628 - 1455	31	
	$3v_2 - 3v_2$	12	9	628 - 743	12	•

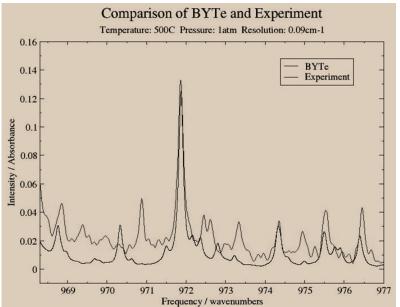
\*Have not found measurements in published works.

# NH3 spectroscopy: line assignments, an example

0.16

0.14

0.12



#### List of A

BYT

9

Lab (Home work)

59 970 971 97 Freque Assigned Lines	973 974 975 976 ncy / wavenumbers	0.1 2v <sub>2</sub> -v <sub>2</sub> Q(16) K = 16 0.04 0.04 0.02 977 0 969	$\begin{array}{c} \mathbf{K} = 2 \\ \mathbf{K} = 17 \\ \mathbf{K} = 1 \\ K$				
Te Frequency	Experimental Frequency	Obs - Calc	Upper Quantum	Lower Quantum			
			Numbers*	Numbers*	π		
968.761998	968.825639	0.063641	- 0 2 0 0 0 0 16 16	+010001616	=		
970.332898	970.874628	0.54173	-02000143	+010000133	_		
971.871137	971.868991	-0.002146	-0100002 1	+00000011	284		
972.159794	972.456569	0.296775	+02000142	-01000132	40		
972.363167	972.60723	0.244063	+020001717	-0100001717			
972.801729	973.330403	0.528674	-02000141	+010000131			
974.317864	974.354898	0.037034	-0101013 1	+0001012 1			
975.511534	975.530054	0.01852	+01010132	-00010122			
976.392929	976.449086	0.056157	- 0 2 0 0 0 0 18 18	+010001818			
				—			

Line Assignments Based on Comparison

9 Lines Assigned  $\Delta K = 0$ 

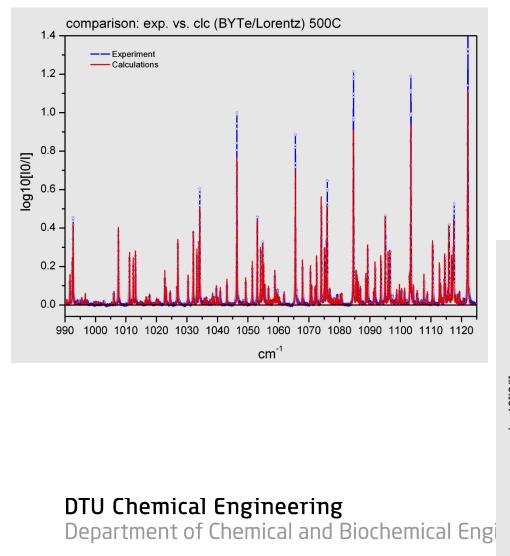
 $V_2$ R(1)

K = 1

\*Parity  $v_1 v_2 v_3 v_4 l_3 l_4 J K$ 

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# NH3: experiment (0.09cm-1) vs calculations (BYTe)

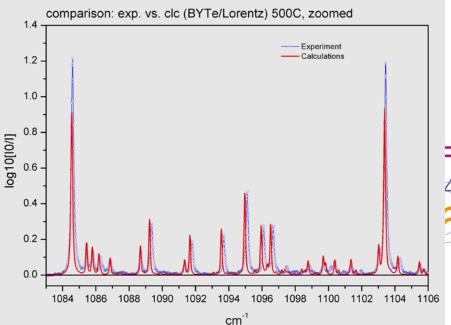


Lab (Home work)

#### Can we use BYTe at 500C for practical apps?

- in general a good agreement
- some difficulties with strong line intensities
- some frequency shifts in line positions

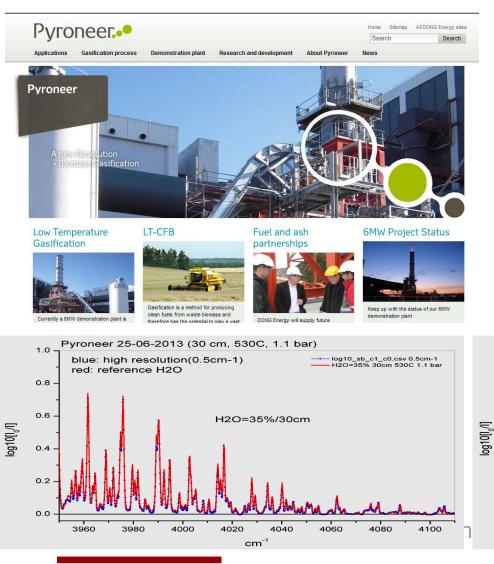
More work to do at even higher T (>500C)





From Lab to Field

# In Situ measurements on Pyroneer (6MW) gasifier



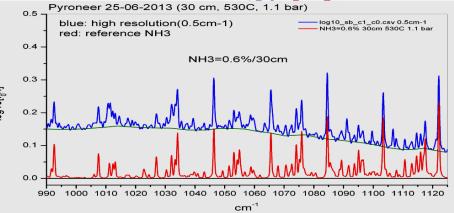
- Very complex producer gas composition (CO2, H2O, CO, H2, HC, PAH, tars) + particles
- Producer gas is feeded into an industrial burner of a power plant

#### Why to do it? (examples):

- H2O (related to mass balance)
- NH3 (related to NOx formation)

#### How?: In Situ IR abs measurements: no gas extraction

- Tough: out of the building on a platform (safety) with limited space (practical issues);
- Tgas about 530C;
- optical measurements over 30 cm;
- very strong any (UV-IR) light attenuation.

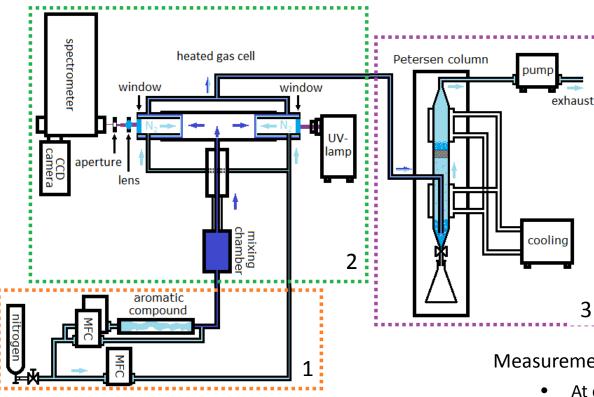




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Lab (Home work)

# Phenol UV absorption cross-sections: experimental set up



- 1. Gas mixing unit
  - N<sub>2</sub> (industrial standard)
  - molten aromatic crystals in tube
  - $\Rightarrow$  concentration unknown
  - admixture of N<sub>2</sub> for different concentration
- 2. Gas cell and optics
- 3. Petersen column
  - sampling in acetone
  - Sampling time 30 min
  - analysis with GC/MS

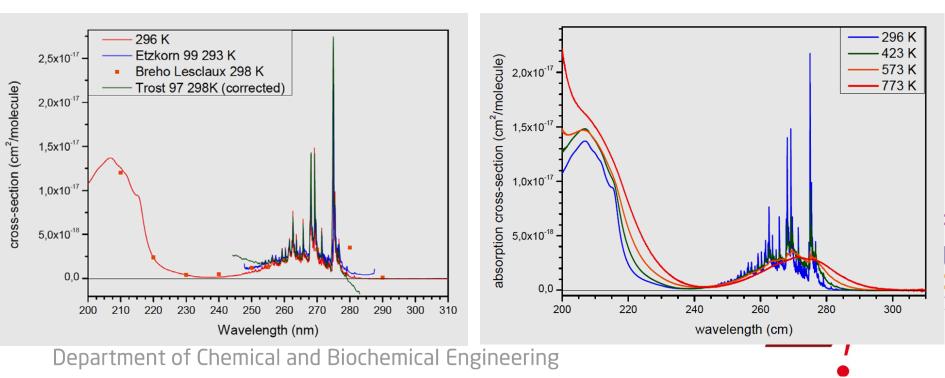
Measurements strategy:

- At each T two phenol concentrations
- At each concentration two sample
- During each sampling three UV spectra and three double concentration determination

# DTU Chemical Engineering

Lab (Home work)Phenol UV absorption cross-sections: temperature effects

- Not too many reference data available even at low T (about 23C)
- An excellent agreement with published data at low T
- Significant changes in the fine structure of the cross-section spectra with T



Low-temperature abs cross-sections: comparison

Abs cross-sections: from 23C to 500C



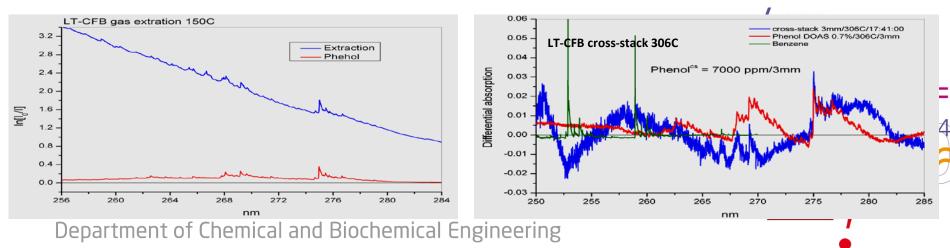


From Lab to Field

# In Situ measurements on LT-CFB (100kW) gasifier



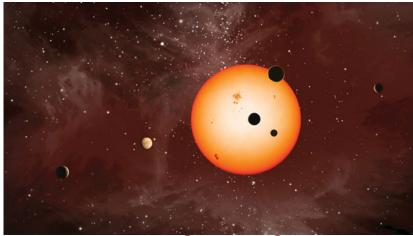
- Focus on trace gases in low- and high-temperature gasification processes;
- Producer gas issues:
  - corrosion (boilers)
  - reduced gas quality (fuel cells, gas grids)
- Phenol major trace gas from PAH's in the producer gas (LT-CFB process);
- Tgas = 300-500C; In Situ UV abs measurements over 3 mm;
- Phenol measurements by various techniques:
  - GC/MS (Petersen column (30 min) 215 ppm (±5%)
  - Gas extraction, 150 C: 407 ppm (±5%) (3 min)
  - In Situ, 306 C (DOAS approach): 7700 ppm (±10%) (3 min)





# Industry and Universities How other planets meets the Earth

- Far away planets on a global scale (e.g. exoplanets, stars) and current Earth's problems on a local scale (energy, emissions , taxes)
- Spectroscopy of hot planets and high-temperature processes: the same gases/temperatures of interest;
- DTU's projects about optical measurements in combustion (SO2, SO3, NH3, etc), gasification (trace gases, Cl- compounds) and waste utilization in collaboration with industry (DONG Energy, Vattenfall and Babcock &Wilcox Vølund)
- UCL's and DTU's common PhD/postdocs projects: SO3/SO2 and Cl-compounds (KCl, HCl, CH3Cl, CH4, H2CO)





From Lab to Field

## Conclusions

## In general

- You can find a lot inspirations for the work on the Earth
- Different research areas can have the same origin
- Scientists can make industry guys happy

## In particular:

- Excellent experimental tools are available for (VUV) UV-FIR optical measurements
- Temperature range can be also negative (e.g. gases at low T)
- New data/lines for NH3 BYTe extension and development
- o New data for phenol
- Try always In Situ and avoid any Ex Situ (extraction) measurements

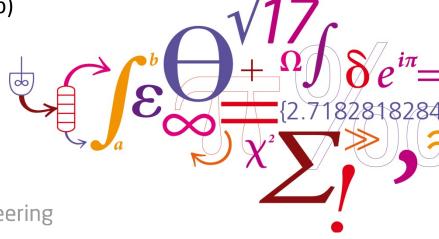
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# **DTU Chemical Engineering** Department of Chemical and Biochemical Engineering



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