# THE STATISTICS OF LARGE EARTHQUAKE MAGNITUDE AND AN EVALUATION OF GREEK SEISMICITY 

by

## K.C. Makropoulos

(B.Sc., Univ. of Athens)

Thesis presented for the degree of Doctor of Philosophy of the University of Edinburgh in the Faculty of Science

1978

MARIA AND CHRISTOS

ETH
MAPIA KAI XPIETO

## ABSTRACT

The problem of the upper bound to earthquake magnitude occurrence is examined. It is demonstrated using simple frequency-magnitude and energy-magnitude laws, that it is possible to include an upper bound as an unknown parameter, and to calculate its value both analytically and graphically. It is shown that a finite upper bound to earthquake magnitude is necessary to preserve a finite rate of energy release. This upper bound is expressed in terms of the mean annual release of energy and the parameter $b$ of the frequency-magnitude law.

The third type asymptotic distribution of extreme values of Gumbel, which includes the upper bound as a parameter, is then determined using Marquardt's algorithm. The methods of Extreme-Values and strain energy release are then analytically related and both methods tested on the high seismicity of the circum-Pacific belt. Uncertainties on the extreme value parameters and related predictions are obtained using an error matrix, and uncertainties on the parameters from strain energy release are also determined. Both methods give similar results and are applicable to the estimation of seismic risk.

These methods are then applied to evaluate Greek seismicity and seismic risk. Greek earthquakes are relocated and magnitudes determined to produce a homogeneous catalogue using mainly instrumental data since 1901. Tectonic models of Greece are examined using the new hypocentres. Seismic risk maps are presented and these contour maps show the maximum expected earthquake magnitudes and accelerations in the next $T$ years at stated probability levels.

## ACKNOWLEDGEMENTS

I would like to express my sincere thanks to Professor A G Galanopoulos and Professor J Drakopoulos for their encouragement and advice to start this work, and for allowing me leave of absence from Athens University to study in Edinburgh.

I am grateful to Professor K M Creer, head of the Geophysics Department (University of Edinburgh) and Dr P L Willmore, head of the Global Seismology Unit (Institute of Geological Sciences), for accepting me to Edinburgh and allowing me full use of their facilities for my project. I am specially indebted to the Global Seismology Unit for having provided me with all possible assistance and support throughout my work.

I am very grateful to Dr P W Burton who supervised me during the course of this work. Without his continual guidance, criticism, encouragement and patience, especially during those dark days when nothing would work, this thesis would never have been completed. I was very lucky to have him as supervisor. He was for me a good friend and I am thankful for his invaluable help throughout this work.

I would like to express my gratitude to Dr P L Willmore and Dr S Crampin for their advice and constant assistance.

Many thanks are also due to all other members of the Global Seismology Unit who helped me in one way or another, in particular I thank R W McGonigle, R B Jones, G Neilson and D Booth who helped me with various discussions and advice on computing.

My student colleague M Assumpcao deserves special thanks for the useful discussions and the enjoyable time I have had in the past three years.

My special thanks are also due to Mrs $G$ Hall for typing the thesis.
This work was partly financed by the Greek Ministry of Coordination, under grant no. 627/32213. The University of Athens also financially supported this work.
Pages
ABSTRACT ..... iii
ACKNOWLEDGEMENTS ..... iv
CHAPTER I Introduction ..... 1
1.1 General statement of problem ..... 1
1.2 Specific problems and research goals ..... 4
CHAPTER II Review of statistical models ..... 6
2.1 Introduction ..... 6
2.2 Statistical models using the whole process ..... 6
2.2.1 Occurrence models ..... 6
2.2.2 Magnitude models ..... 10
2.2.3 Attenuation model ..... 11
2.3 Extreme-Value theory fitting, using the part process ..... 12
2.3.1 The three asymptotic distributions ..... 15
2.3.2 Mathematical meaning of the parameters ..... 16
2.3.3 Physical meaning of the parameters ..... 17
2.3.4 Probability papers. Plotting positions ..... 17
2.3.5 Estimation of the parameters ..... 20
2.3.6 Useful relations for forecasting procedure ..... 23
2.3.7 Applications of the Extreme-Value theory to seismic risk problems ..... 26
2.4 Summary ..... 28
CHAPTER III The upper bound for earthquake magnitude ..... 29
3.1 Statement of the problem ..... 29
3.2 Energy release and maximum magnitude earthquake ..... 30
3.2.1 Mathematical consideration ..... 30
3.2.2 Graphical methods of estimating $M_{2}$ and $M_{3}$ ..... 34
3.2.3 Testing the two methods ..... 35
3.2.4 Results and discussion ..... 36
3.3 Summary ..... 40
CHAPTER IV Third type asymptotic distribution of Gumbel and
strain energy release relations ..... 42
4.1 Introduction ..... 42
4.1.1 Uncertainties - weights ..... 42
4.1.2 Relations between parameters of strain energy release and the third type asymptote ..... 42
4.1.3 Testing region ..... 43
4.2 Estimation of the parameters ..... 43
4.2.1 Non-linear least-squares methods ..... 43
4.2.2 The Marquardt (1963) algorithm ..... 45
4.2.3 Computations ..... 48
4.3 Prediction uncertainties ..... 48
4.4 Energy release and the third type asymptotic distribution ..... 50
4.4.1 The mode ..... 50
4.4.2 The mean annual energy release ..... 50
4.4.3 The upper limit ..... 54
4.5 Testing the third type asymptotic distribution method ..... 54
4.5.1 Data ..... 54
4.5.2 Data treatment ..... 54
4.5.3 Results and discussion ..... 57
4.6 Comparing the results from energy release and third type asymptotic distribution methods ..... 62
4.7 Conclusions
An earthquake catalogue for the area of Greece$\xrightarrow{\mathrm{N}_{33}^{42.5}, \mathrm{E}_{19}^{29} \text { since } 1901}$67
5.1 Introduction ..... 67
5.2 Previous work in Greek earthquake cataloguing ..... 67
5.3 The ISS epicentres: reasons for inaccuracies ..... 71
5.4 Data sources ..... 72
5.4.1 Relocation ..... 72
5.4.2 Magnitude ..... 73
5.5 Earthquake relocation procedure ..... 74
5.5.1 Methods and computer programs chosen for relocation ..... 74
5.5.2 Procedure used ..... 74
5.6 Magnitude determination ..... 78
5.6.1 Magnitude scale chosen ..... 78
5.6.2 Magnitude determination procedure ..... 79
5.7 Completeness of the catalogue ..... 82
5.7.1 Temporal plot of grouped events ..... 82
5.7.2 Analysis of sample completeness ..... 84
5.8 Comparing the results ..... 88
5.9 Summary ..... 90
CHAPTER VI Greek tectonics and seismicity ..... 93
6.1 Introduction ..... 93
6.2 Morphologic, geologic and geophysic features of the area ..... 93
6.3 Principle tectonic models for Greece and the adjacent areas ..... 95
6.4 Spatial distribution of the earthquakes ..... 99
6.4.1 Shallow earthquakes ( $\mathrm{h}<60 \mathrm{~km}$ ) ..... 99
6.4.2 Intermediate earthquakes $(h \geqslant 60 \mathrm{~km})$ ..... 100
6.5 Isodepth maps from radial vertical cross-sections ..... 102
6.5.1 Procedure ..... 102
6.5.2 Results and discussion ..... 103
6.6 Summary ..... 106
CHAPTER VII Greek seismic risk evaluation ..... 108
7.1 Introduction ..... 108
7.2 Data and cities for seismic risk estimation ..... 109
7.3 Seismic risk based on the magnitude distribution ..... 110
7.3.1 Comparison of the methods applied ..... 110
7.3.2 Seismic risk evaluation - magnitude ..... 111
7.4 Seismic risk based on peak acceleration of ground motion ..... 114
7.4.1 Choosing the acceleration-distance formula ..... 114
7.4.2 Method used to fit the maximum acceleration data ..... 115
7.4.3 Procedure used to evaluate the seismic risk in T years at a given probability level ..... 115
7.4.4 Seismic risk evaluation - acceleration ..... 117
7.5 Spatial distribution of seismic risk in Greece ..... 117
7.5.1 Contour maps of seismic risk - Procedure used ..... 117
7.5.2 Results and discussion ..... 118
7.6 Summary ..... 122
CHAPTER VIII Summary and final conclusions ..... 125
APPENDIX A Computer program listing ..... 131
Al Risk Analysis Program (RAP) ..... 132
APPENDIX B Earthquake catalogue for Greece since 1901 ..... 147
REFERENCES ..... 182

## CHAPTER I

INTRODUCTION

As a result of increasingly complex industrial, commercial and residential developments, which cause centres of population to spread in ever widening circles, reliable estimations of seismic risk and seismic hazard, and developments of means of mapping them, are among the research problems in seismology which most urgently require answers.

Definition of seismic risk as "the probability of occurrence of an earthquake in the future" (Lomnitz, 1974), implies a degree of future uncertainty. Hence principles of probabilistic forecasting and decision making are essential in any seismic risk analysis. Models of seismic risk usually consist of:
i) empirical formulae based on available macroseismic data,
ii) statistical distribution laws for earthquake occurrence in time and magnitude, and
iii) attenuation laws describing the decay of seismic ground motion with focal distance.
1.1 General statement of problem

The distribution of earthquake magnitudes in time and in size is generally investigated by:
i) Using the whole available data - whole process.
ii) Using only the extreme value magnitudes - part process.

When models of the first category are applied to the experimental data, like the linear frequency-magnitude model of Gutenberg and Richter (1944), it becomes clear that they do not represent the real process for the large earthquakes. Most of the proposed alternative expressions such
as the quadratic or truncated frequency-magnitude formulae (Merz and Cornell, 1973, Cornell and Vanmarcke, 1969) do not recognize the inescapable existence of an upper bound to the magnitude that can be generated in a region (Esteva, 1976).

To implement earthquake-resistant design codes, it is usually necessary to know the maximum dynamic load to which a structure might be subject during its design life, or alternatively, the most probable return period of a specified design load. The inclusion of an upper bound to earthquake magnitude as an unknown parameter in a statistical model leads to more reliable estimates, especially for large earthquakes, because it is closer to the real process than that represented by the unlimited or truncated models.

One of the objectives of this study is to investigate the regional upper bound for earthquake magnitude. This is first attempted by using the strain energy release in the region, calculated from the linear energymagnitude law, combined with analytic expressions for the upper bound to magnitude.

In all the statistical models which use the whole process, inclusion of low magnitudes, which usually are incomplete and inhomogeneous, can bias the estimation of the prediction parameters. On the other hand, in earthquake engineering applications, the need to consider extreme value distributions separately from the statistics of the whole process is of primary importance. Thus, another prime objective of this study is to investigate the usefulness of the distribution of Extreme-Values, which Gumbel (1966) has called the third type asymptotic distribution of extremes, for estimating the seismic risk and return periods of largest earthquakes. This type of distribution is chosen because it holds for initial distributions which are limited towards the largest values, and it contains the upper bound of the distribution as an unknown parameter.

The theory of Extreme-Value statistics is formulated under the assumptions:
i) the prevailing conditions are valid in the future and, ii) the observed extreme values are independent of each other.

For the case of earthquake occurrence, our experience shows that earthquakes do not occur at the same level of magnitude and frequency all over the world. Practically $99 \%$ of all earthquakes occur along plate boundaries. Less than $3 \%$ of the earth's seismic energy release occurs on the midoceanic rises or in the interior of plates (Lomnitz, 1974).

Although aftershocks following large earthquakes are the most outstanding example of dependent events, Gumbel (1966) suggests that the influence of interdependence may vanish for largest values of a variate. Lomnitz (1966) points out that large earthquakes are indeed characterized by a high degree of randomness and independence in time.

It is then reasonable to assume for a specific region and sampling period that the behaviour of the largest earthquakes will usually be similar to that of the near past; although the distribution will vary over geological epochs. Consequently the theory of Extreme-Value statistics can be applied to establish a prediction procedure for the largest earthquake magnitude of the next $n$ years, by using the past $N$ years' earthquake data in a given region.

These two objectives form the first part of this thesis.
In the second part, the objective is to evaluate the seismicity and seismic risk of Greece, using the methods already developed, which by then will be seen to give reliable estimates of the future seismic activity of a region. The evaluation is obtained by estimating both the return periods for specified magnitudes and also the expected magnitudes and accelerations within a period of time, all at a given probability level. The results
are finally presented by mapping the geographic variation of earthquake risk in terms of maximum magnitude earthquakes and maximum ground motion accelerations expected to occur in the next $T$ years.

### 1.2 Specific problems and research goals

In any seismic risk analysis estimation of the uncertainties in the predictions is a vital factor for the final judgement of the results. A specific goal is then to develop a technique to compute the errors on the parameters of the third type asymptotic distribution, and on all related predictions, using an error matrix. A second goal is to explore any physical meaning of these parameters by linking them with the physical release of strain energy. This is attempted by relating physical quantities such as mean annual energy release, derived from the linear frequency-magnitude and energy-magnitude laws, with the same quantities obtained using the parameters of the third type asymptote. The strain energy release and the third type asymptotic distribution methods are then tested on the seismicity of the circum-Pacific belt, and then applied to the seismicity of Greece.

Seismically, Greece is one of the most active countries in the world and the most active country in Europe. About 3 to $4 \%$ of the seismic energy release in the world is contributed by Europe, and half of this by Greece (Galanopoulos, 1971a). The long documented seismic history of Greece reports many catastrophes due to earthquakes (Galanopoulos 1961, Lomnitz 1974). However, demands of statistical seismology for data which is as accurate and homogeneous as possible, implies that instrumentally recorded events are preferable. The present century is the only available period with data of sufficient reliability for our purposes. Even so, because the worldwide density of stations has increased markedly from decade to decade, the completeness of a sample of earthquake data is strongly dependent not only on the geographic area but also on the particular time interval covered. In
order to achieve a more accurate picture of seismic risk for Greece, an important aspect of this study must be the preparation of an homogeneous earthquake catalogue by relocating all the events for which there is sufficient data. For this purpose all earthquakes for the period 1917-1963 are selected from International and Greek sources and the hypocentral parameters will be recalculated, and the completeness of the data tested. The calculation of seismic risk, either in terms of strain energy release or acceleration, velocity etc. depends critically on the magnitude of the earthquakes considered and so homogeneous magnitudes are necessary. This requirement can be best fulfilled if the magnitudes are determined from the same instruments which ideally should have been operating all the time. Thus a complementary aspect of relocation is the calculation of magnitudes for every single earthquake which appears in the new catalogue.

Uppsala Wiechert amplitudes from Uppsala seismological bulletins will be used to determine surface-wave magnitude up to 1954, whereas from 1955 onwards, the Uppsala and Kiruna amplitudes from modern instruments will be used.

Thus, the sample of extremes that eventually forms the final basis for statistical analysis of the earthquake risk in Greece consists of the set of annual maxima of magnitudes and accelerations which are drawn as required from the new catalogue. The final product will be detailed maps and evaluations of seismic risk in Greece.

## REVIEW OF STATISTICAL MODELS

### 2.1 Introduction

The earthquake phenomenon has been analysed for many years in terms of specific characteristics such as its location, magnitude and focal depth. The set of these characteristics of historical earthquakes is called the "seismicity" of the region. With the development of the hypothesis of sea-floor spreading, however, the earthquake phenomenon and seismic activity in general began tc be regarded as a global process. Since then, the generation and propagation of seismic energy from source to the site are among the earthquake's diverse seismic properties which have been studied with rising interest.

The occurrence of earthquakes in space and time falls under the general category of stochastic processes, that is, mathematical models of a given physical system that changes in accordance with the laws of probability (Lomnitz, 1974). Hence, statistical models have to be used, and the validity of the model checked by its concordance with past observations.

The aim of this chapter is to review some of the existing statistical models of the occurrence of earthquakes in time, magnitude and attenuation which includes acceleration, velocity and displacement. Emphasis is given to the Extreme-Value models, of which the third-type asymptotic distribution will be one of the main subjects of this study.

### 2.2 Statistical models using the whole process

2.2.1 Occurrence models
a) Simple Poisson model

Let us consider the earthquake as an event which occurs along
a time axis. This model assumes that one event in a given magnitude
range and in any given volume of the earth crust is equally likely to be found in any unit time interval along the time axis, and it is independent of any other event. The probability of finding $n$ events in time $t$, if the mean rate of occurrence $k$ is known, follows the Poisson probability law:

$$
\begin{equation*}
P(n, k t)=\frac{e^{-k t}(k t)^{n}}{n!} \tag{2-1}
\end{equation*}
$$

The mean and the variance of the Poisson distribution are both equal to the mean rate $k$.

According to the Poisson model, the probability of observing no events within a time $t$, that is of finding a time larger than $t$ without events, is:

$$
\begin{equation*}
P(0, t)=e^{-k t} \tag{2-2}
\end{equation*}
$$

Then the probability of finding a time equal or less than $t$ with no events is:

$$
\begin{equation*}
P(t)=1-e^{-k t}, 0<t<\infty \tag{2-3}
\end{equation*}
$$

This time interval is the waiting time or return period between .two consecutive events. The corresponding probability density function is:

$$
\begin{equation*}
P^{\prime}(t)=k e^{-k t}, t>0 \tag{2-4}
\end{equation*}
$$

with mean $1 / k$ and variance $1 / k^{2}$, where $P^{\prime}(t)$ denotes the derivative of $P(t)$ with respect to time.

There are cases where the Poisson model seems to give satisfactory results, at least for main events or when dealing with large shocks throughout the world (Cornell and Kallberg, 1969, Sacuiu and Zorilescu, 1970). But many authors, for example, Knopoff (1964) and Vere-Jones (1970) find, after testing different regions in the world, that the Poisson model is
inadequate and does not explain the time distribution of low magnitude events, because it ignores the tendency for these events to cluster in space and time. In addition to this, the Poisson mode1 assumes that the distribution of the waiting time between events does not depend on knowledge of the time elapsed since the last one, that is the occurrence of a certain magnitude is completely independent of the occurrence of any other.
b) Clustering models

In order to improve the simple Poisson model the basic assumption of independence of events, especially at low magnitudes, has to be checked. Thus, in the case of simple dependence, the Markov model may be used, whereas the non-Markovian process may be the alternative approach for complex dependence.
(1) Markov model

The basic idea of using the Markov model for earthquake occurrence comes from the fact that according to the Elastic Rebound Theory (Reid, 1911), there is a storage of strain energy that has to build up before a new event takes place.

According to the Markov model, the probability of a future event depends on the past history of events so that (Oliveira, 1974):

$$
\begin{equation*}
P\left(n_{k}, t_{k} \mid n_{0}, t_{o} ; n_{1} t_{1} ; \ldots n_{k-1}, t_{k-1}\right)=P\left(n_{k}, t_{k} \mid n_{k-1}, t_{k-1}\right) \tag{2-5}
\end{equation*}
$$

$n_{i}, t_{i}$ being the number of events $n_{i}$ to occur in the time interval $t_{i}$; or the probability of being in a state*k, after considering all the
*States of the process are the regions in which a multidimensional continuum is divided. For example, the variables: energy, number of earthquakes, latitude, longitude, depth and time, define a 6-dimensional space: the earthquake state space (Lomnitz, 1974).
states from zero up to $k$, depends only on the probability of being in the state $k-1$. This is a first order Markov chain characterized by the transition probability:

$$
P\left(n_{k}, t_{k} \mid n_{k-1}, t_{k-1}\right)
$$

Authors like Shah and Vagliente (1972), Vagliente (1973), Benjamin and Cornell (1970) have emphasised the usefulness of Markov simulation but also some weaknesses of this model due to the difficulty in setting the initial conditions and because it requires complicated numerical treatment.
(2) Non-Markovian models

Applicability of some of these general models is discussed by Vere-Jones (1970). These models assume that earthquakes occur in clusters, and that the number of events in each cluster is stochastically independent of its origin time.

In these models the conditional probability of an event taking place during the interval $(t, t+d t)$, given that the cluster consists of $N$ shocks is equal (Esteva, 1976) to:

$$
\begin{equation*}
P(N, t)=N \ell(t) d t \tag{2-6}
\end{equation*}
$$

where

$$
\begin{equation*}
\ell(t)=\partial L(t) / \partial t \tag{2-7}
\end{equation*}
$$

L(t) being the cumulative distribution function of the time of an event corresponding to a given cluster, measured from the cluster origin time.

Most of these models are based on information about earthquakes with magnitudes above relatively low thresholds, recorded during time intervals of at most ten years, and as Esteva (1976) points out, the degree of clustering observed, and the distribution of times between clusters, cannot be extrapolated to higher magnitude thresholds and longer time interval without further study.

The usual expression relating earthquake magnitudes with their rates of occurrence is due to Gutenberg and Richter (1944). This relationship is called "magnitude-frequency formula" and states that in a certain region and in a given period of time:

$$
\begin{equation*}
\log N(m)=a-b m \tag{2-8}
\end{equation*}
$$

where $N(m)$ is the mean number of earthquakes per unit volume and per unit time having magnitude greater than $m$, and $a$ and $b$ are zone-dependent constants. a depends on the period of observation and on the level of seismicity of the region, and consequently, varies widely from region to region, while $b$ remains within a relatively narrow range.

Equation (2-8) can be normalized to yield the frequency distribution of magnitudes in a region. Thus, for $m=0$ equation (2-8) becomes:

$$
\begin{equation*}
\log N(0)=a \tag{2-9}
\end{equation*}
$$

and normalization is achieved by dividing through by $N(0)$ :

$$
\begin{equation*}
\log [1-F(m)]=\log \frac{N(m)}{N(0)}=b m \tag{2-10}
\end{equation*}
$$

which yields:
or

$$
\begin{align*}
& 1-F(m)=e^{-\beta m}(m \geqslant 0)  \tag{2-11}\\
& F(m)=1-e^{-\beta m}(m \geqslant 0) \tag{2-12}
\end{align*}
$$

where $\beta=b /$ loge $=1 / \bar{m}$, and $F(m)$ is the cumulative probability distribution of earthquake magnitudes. The frequency distribution $f(m)$ is the first derivative of $F(m)$ :

$$
\begin{equation*}
f(m)=\beta e^{-\beta m}(m \geqslant 0) \tag{2-13}
\end{equation*}
$$

[^0]Experimental data show that the linear form $(2-8)$ does not hold for very large earthquakes. Thus, alternative expressions to (2-8) have been proposed, attempting to represent more adequately the observed magnitude-recurrence data. Cornell and Vanmarcke (1969) impose a limitation or upper bound to the exponential distribution of magnitudes (2-13), by introducing a truncating factor $k\left(m_{1}\right)$ where $m_{r}$ is the largest possible Richter's magnitude, the value of which for the world has been suggested as 8.5 to 9.0 (Richter, 1958). Thus, equation (2-13) becomes:

$$
\begin{equation*}
f(m)=k\left(m_{1}\right) \beta e^{-\beta m}, 0 \leqslant m \leqslant m_{1} \tag{2-14}
\end{equation*}
$$

Sacuiu and Zorilescu (1970) and Merz and Cornell (1973) introduce a quadratic term for the magnitude into equation (2-8).

Most of the expressions, however, fail to recognize the existence of an upper bound to the magnitude that can be generated in a given region, the existence of which is inescapable (Esteva, 1976).

### 2.2.3 Attenuation model

The attenuation of seismic energy as it propagates from source to the site is a very important consideration for any seismic risk analysis. Attenuation models relate the changes of the characteristics of earthquake ground motion with distance. Seismic attenuation laws have been determined using the maximum acceleration, velocity and displacement. Esteva and Rosenblueth (1964) after considering data from earthquakes, the propagation of elastic waves and the definition of magnitude, proposed the general formula:

$$
\begin{equation*}
\mathrm{Y}=\mathrm{b}_{1} e^{\mathrm{b}_{2} \mathrm{~m}-\mathrm{b}_{3}} \tag{2-15}
\end{equation*}
$$

where $Y$, the peak ground motion, is connected with the magnitude $m$ and the focal distance $R ; b_{1}, b_{2}$, and $b_{3}$ are constants and their values
depend on the type of $Y$ considered. In Chapter VII, Table 7-7 summarizes some of the proposed formulae. Because of the limited number of strongmotion records and especially because most of these records come from the same area, these expressions often amount to modifications of the general formula (2-15) and allow for local conditions rather than large regional differences.

### 2.3 Extreme-Value theory fitting, using the part process

The previous models which use the whole process have a common "weak" point; there is a lack of accuracy, homogeneity and completeness of data sets used, at low magnitudes. These properties are very important for any statistical treatment and mainly depend on the sophistication of the instruments and networks used to record the earthquakes. Even now, after seventy years of continuous development, the problem still remains for smaller earthquakes. Another difficulty of applying the previous models is that in most cases the initial distribution function is not known. From the point of view of earthquake risk, however, the quantity usually needed is the maximum dynamic load to which a structure will be subject during its design life or, alternatively, the most probable return-period of a specified design load. Thus, what is of primary importance in earthquake engineering implies a need to consider the extreme value distribution separately from the statistics of the whole process.

The statistical theory which seems to overcome most of the abovementioned problems is the Extreme-Value Theory. This theory following Gumbel's developments and applications to flood analysis (Gumbel, 1935, 1966) has attracted widespread interest and has been adapted by geophysicists for hydrological computations, climatic evaluations (Jenkinson 1955, Gringorten 1963b, Krumbein and Lieblein, 1956), as well as for
analysis of earthquake events and seismic risk determination (Nordquist, 1945, Dick 1965, Epstein and Lomnitz 1966, Kárnik and Hubnerova 1968, Milne and Davenport 1969, Schenkova and Kárnik 1970, Stepp 1971, Shakal and Willis 1972, Curtis 1973, Yegulalp and Kuo 1974, Yegulalp 1974, Schenkova and Schenk 1975, Lilwall 1976, Willmore and Burton 1976, Schenkova and Kárnik 1976, 1978, Kárnik and Schenkova 1977, Burton 1978a, 1978b).

Some of the more important practical advantages of the ExtremeValue method are (Lomnitz 1974):
i) The extreme values of a geophysical variate are better known, more homogeneous in time, and more accurately determined than the average events in a time series of data.
ii) The method does not require a detailed knowledge of the parent distribution. The distribution of extremes depends only on the tail of the distribution of the variate, and certain important distributions behave in a similar way at large values; for example, the normal, exponential and log normal distributions.

While details of the theory and many references to basic original contributions will be found in the previous papers, it will be useful to summarize the main results. With regard to the largest value (a similar formulation applies to the smallest value) the problem is formulated as follows:

Suppose we have a sample consisting of N independent extremes, each being the largest of sample of size $n$, where $n$ is large, drawn from the same parent population. The series of extremes then forms a distribution called the Extreme-Value distribution. For the case of annual extremes, n is equal to 365 days, and N is equal to period in years considered.

Suppose $Y$ to be the largest earthquake occurring in a year; and define

$$
\begin{equation*}
y=\max \left(m_{1}, m_{2}, \ldots m_{N}\right) \tag{2-16}
\end{equation*}
$$

where the $m_{i}$ form a sequence of annual maximum earthquakes drawn at random from the cumulative distribution $F(m)$. Then the probability that $Y$ will be the largest among N independent samples is:

$$
\begin{align*}
\phi_{N}(y) & =P(Y \leqslant y)=P\left(a 11 m_{i} \leqslant y\right) \\
& =P\left(m_{1} \leqslant y, m_{2} \leqslant y, \ldots, m_{N} \leqslant y\right) \tag{2-17}
\end{align*}
$$

Since the $m_{i}$ are independent events, the probability of the largest event can be written using the multiplication rule as:

$$
\begin{align*}
\phi_{N}(y) & =P(Y \leqslant y)=P\left(m_{1} \leqslant y\right) P\left(m_{2} \leqslant y\right) \ldots P\left(m_{N} \leqslant y\right) \\
& =F_{m_{1}}(y) F_{m_{2}}(y) \ldots F_{m_{N}}(y) \text { and } \\
\phi_{N}(y) & =F_{m}^{N}(y) \tag{2-18}
\end{align*}
$$

The probability of a value to be equal or larger than $y$ is

$$
1-\phi_{N}(y)
$$

and its reciprocal

$$
\begin{equation*}
T_{N}(y)=\frac{1}{1-\phi_{N}(y)} \tag{2-19}
\end{equation*}
$$

is the return period of $y$, which is the mean number of intervals required for a largest value greater than or equal to $y$ to be observed.

If $K$ year extreme intervals have to be used, the distribution $\phi_{\mathrm{K}}(\mathrm{y})$ is related to $\phi_{1}(\mathrm{y})$ for the one year intervals by the formula:

$$
\begin{equation*}
\phi_{K}(y)=\phi_{1}^{K}(y) \tag{2-20}
\end{equation*}
$$

### 2.3.1 The three asymptotic distributions

Irrespective of the parent distribution, the distribution from which the extremes are sampled, the limiting extreme-value distribution (Davis 1970, p. 205) must take one of three forms. These we will call first, second and third type asymptotic distributions of Gumbel (1966). Each of these assumes a specific behaviour for absolute large values of the variable (Gumbel, 1963). The three asymptotic distributions are:

$$
\begin{array}{ll}
\text { I } & \phi^{(1)}(x)=\exp [-\exp (-a(x-u))], a>0 \\
\text { II } & \phi^{(2)}(x)=\exp \left[-\left(\frac{u-\varepsilon}{x-\varepsilon}\right)^{k}\right], k>0, x \geqslant \varepsilon, u>\varepsilon \geqslant 0 \\
\text { III } & \phi^{(3)}(x)=\exp \left[-\left(\frac{\omega-x}{\omega-u}\right)^{k}\right], k>0, \quad x \leqslant \omega, u<\omega
\end{array}
$$

and if we introduce the reduced variable $y$ as:

$$
\begin{equation*}
y=-\ln [-\ln \phi(x)] \tag{2-23}
\end{equation*}
$$

these become:

$$
\begin{align*}
& y_{I}=a(x-u) \text { with sign of } \frac{d^{2} x=0}{d y^{2}}  \tag{2-24}\\
& y_{I I}=k \ell n(x-\varepsilon)+\text { const with sign of } \frac{d^{2} x}{d y^{2}}>0  \tag{2-25}\\
& y_{\text {III }}=\text { const-kln(w-x) with sign of } \frac{d^{2} x}{d y^{2}}<0 \tag{2-26}
\end{align*}
$$

The first type asymptotic distribution (2-20) holds for initial distributions unlimited in both directions (exponential or normal), but tends to zero exponentially in the tails. The second type arises when the initial distribution is bounded below and exponentially approaches zero in the upper tail. The third type results when the initial distribution is bounded towards the right. The second type asymptotic distribution is therefore ruled out, and only the third type asymptotic
distribution, and for comparative reasons the first type, will be employed in the analysis of extreme magnitude earthquakes in this thesis.
2.3.2 Mathematical meaning of the parameters
a) For the first type asymptotic distribution:
a is the extremal intensity function
$\underline{u}$ is the characteristic largest value with the property

$$
\phi^{(1)}(u)=\frac{1}{e}
$$

and at the same time $u$ is the mode of the largest values.
b) For the third type asymptotic distribution
$\underline{\omega}$ is the upper limit with the property

$$
\phi^{(3)}(\omega)=1
$$

$\underline{k}$ is the shape parameter and
$\underline{u}$ is again the characteristic largest value with

$$
\phi^{(3)}(\mathrm{u})=\frac{1}{\mathrm{e}}
$$

but not the mode as in the first type asymptote.
The two asymptotic distributions have the common property

$$
\begin{equation*}
\phi^{(1)}(u)=\phi^{(3)}(u)=\frac{1}{e} \tag{2-27}
\end{equation*}
$$

Therefore, approximately $36 \%$ of the observations in all cases should be situated before the value $\mathrm{x}=\mathrm{u}$.
a) For the first type asymptotic distribution
$\underline{1 / a}$ is a measure of dispersion of the extreme values of magnitude from the mode $u$ where
$\underline{u}$ is the most probable observed annual (if the time interval is one year) maximum magnitude. For $u$ we have (see 2.3.6):

$$
u=\frac{a}{b}
$$

where $a$ and $b$ are the constants of the frequency-magnitude or GutenbergRichter's law.
b) For the third asymptotic distribution

W is the upper limit of the distribution (ie the largest earthquake magnitude that may occur in the area under consideration). Such an upper limit is likely to be a function of maximum source size in the Earth's crust and upper mantle (Yegulalp and Kuo, 1974).
$\underline{u}$ is the maximum magnitude which because of $(2-27)$ is exceeded in the long run about $63 \%$ of the time.
$\underline{k}$ is the shape parameter because its reciprocal $1 / k$ is a measure of the curvature of the asymptotic distribution curve. The value of $k$ shows how quickly the curve approaches the upper limit.

### 2.3.4 Probability papers - plotting positions

a) Probability papers

Probability papers are constructed to obtain approximate straight lines for the observed (cumulative) frequencies for an assumed distribution. These papers were first suggested by Powell (1943) and prepared by Gumbel (1945), and give a simple graphical method of testing the fit between theory and observations. The choice of a probability paper is identical with the choice of a distribution. Generally, the probability
paper is a rectangular grid where the observed variate x is plotted as ordinate and the reduced variate $y$ as abscissa, both in linear scales. In addition, the probability $\phi(x)$ is plotted on a scale parallel to the scale of $y$, and the return period $T(x)$ on an upper line parallel to the abscissa, going to the right. The scales $\phi(x)$ and $T(x)$ are not linear. Consider, for example, the first type asymptotic distribution of extremes:

$$
\begin{equation*}
\phi(\mathrm{x})=\exp [-\exp (-\mathrm{a}(\mathrm{x}-\mathrm{u}))] \tag{2-28}
\end{equation*}
$$

or taking the double logarithm

$$
\begin{equation*}
F(x)=-\ln [-\ln \phi(x)]=a(x-u)=y \tag{2-29}
\end{equation*}
$$

then $F(x)$ is a linear function of $x$ and plots as a straight line on extreme probability paper.
b) Plotting position

In order to use the probability papers, the $N$ annual extremes $x_{i}$ are arranged in ascending size such that $x_{1} \leqslant x_{2} \leqslant \cdots \leqslant x_{N}$ and a probability value $\phi\left(x_{i}\right)$ has to be assigned to each extreme $x_{i}$. The value of $\phi\left(x_{i}\right)$ is determined from the N observed extremes in such a way that

$$
\begin{equation*}
0<\phi\left(x_{1}\right)<\phi\left(x_{N}\right)<1 \tag{2-30}
\end{equation*}
$$

Gumbel (1954) has proposed the formula

$$
\begin{equation*}
\phi_{N}\left(x_{i}\right)=\frac{i}{N+1} \tag{2-31}
\end{equation*}
$$

where $i$ is the ordered value from a sample of size $N$. This formula gives an approximation to the expected probability for the ordered observations regardless of the initial distribution.

Kimball (1960) and Gringorten (1963a) have discussed alternative plotting positions and inherent bias in the various options. These
plotting positions are:

$$
\begin{gather*}
\phi\left(x_{i}\right)=\left(i-\frac{1}{2}\right) / N,  \tag{2-32}\\
\phi\left(x_{i}\right)=\left(i-\frac{3}{8}\right) /\left(N+\frac{1}{4}\right), \tag{2-33}
\end{gather*}
$$

or

$$
\begin{equation*}
\phi\left(x_{i}\right)=(i-0.44) /(N+0.12) \tag{2-34}
\end{equation*}
$$

All these plotting formulae can be expressed by the general formula

$$
\begin{equation*}
\phi\left(\mathrm{x}_{\mathrm{i}}\right)=(\mathrm{i}-\mathrm{a}) /(\mathrm{N}+1-2 \mathrm{a}) \tag{2-35}
\end{equation*}
$$

and the problem of selecting a suitable formula reduces to the problem of finding a suitable value of $a$. The equations (2-31) to (2-34) are derived from $(2-35)$ for $a=0, a=\frac{1}{2}, a=\frac{3}{8}$, and $a=0.44$.

The plotting position (2-32) is used by Jenkinson (1955) for the analysis of meteorological data. Kimball (1960) shows that this position produces a bias for the largest of the annual extremes which is opposite in direction to that of Gumbel's (2-31), which as Knopoff and Kagan (1978) show also produces a significant bias. Gringorten (1963a) shows that as $N$ becomes large, a approaches the value of

$$
\begin{equation*}
a=1-e^{-\gamma}=0.439 \tag{2-36}
\end{equation*}
$$

where $\gamma$, Euler's constant, is equal to $0.577 \ldots$, and he suggests that (2-34) is most suitable for the double exponential Gumbel's first type asymptotic distribution. This function minimizes the bias in the longreturn period end of the distribution. Hence, Gringorten formula is chosen in this study because we are more interested in obtaining best fit for the high magnitudes and long return periods at the right end of the distribution.

### 2.3.5 Estimation of the parameters

a) The first type asymptotic distribution
(1) Maximum likelihood method

Kimball (1946) estimates the parameters a and $u$ of the first type asymptotic distribution ( $2-20$ ), by the maximum likelihood method using complicated successive approximations. Stepp (1971) also uses this method for the earthquake risk analysis in the Puget Sound area (USA). The basic principle of the method is to maximize the product:

$$
\underset{i=1}{N=\prod_{i=1}^{N} f\left(x_{i} ; u, a\right)}
$$

With respect to $u$ and $a$, where $f(x)=\phi^{\prime}(x)$ is the probability function and $N$ the sample size. If we use the logarithm, then (2-37) becomes:

$$
L=\sum_{i=1}^{N} \log f\left(x_{i} ; u, a\right)
$$

The conditions of maxima $L$ with respect to $u$ and a are:

$$
\begin{equation*}
\frac{\partial L}{\partial u}=0 \text { and } \frac{\partial L}{\partial a}=0 \tag{2-39}
\end{equation*}
$$

By simultaneously solving these equations using successive approximations one can get the maximum likelihood estimates of $u$ and $a$.

The maximum likelihood method applied to ordered extremes, weights the lower extremes of the series (Kimball 1946) relatively much more than the upper extremes. However, when treating annual earthquake extremes, known to be incomplete especially at the low magnitudes, this property of the maximum likelihood method may influence the final estimates.
(2) Least-squares method

The advantage of probability paper is the transformation of the
theoretical curve $\phi(x)$, $x$ into a straight line. The linear reduction

$$
\begin{equation*}
y=a(x-u) \text { or } x=u+\frac{1}{a} y \tag{2-40}
\end{equation*}
$$

for which any probability paper is designed also allows the classical method of least-squares to be used. This method is used for the estimation of $u$ and a throughout this study.
b) The third type asymptotic distribution
(1) Method of moments

The parameters $\omega, u$, and $k$ of the third type asymptotic distribution (eq 2-22), may be estimated with the help of the first three sample moments (Gumbel 1966, Yegulalp and Kuo 1974)..

Equation (2-22), with the transformation

$$
\begin{equation*}
z=\frac{\omega-x}{\omega-u}, \quad x \leqslant \omega \tag{2-41}
\end{equation*}
$$

becomes

$$
\begin{equation*}
\phi(x)=e^{-z^{k}} \tag{2-42}
\end{equation*}
$$

Then, the reduced moment of order $z^{\ell}$ is:

$$
\begin{equation*}
z^{-\ell}=-\int_{0}^{\omega} z^{\ell / k} d e^{-z}=\frac{\ell \Gamma}{k}\left(\frac{\ell}{k}\right) \tag{2-43}
\end{equation*}
$$

Hence

$$
\begin{equation*}
(\omega-\mathrm{x})^{l}=(\omega-\mathrm{u})^{l} \Gamma\left(1+\frac{\ell}{\mathrm{k}}\right) \tag{2-44}
\end{equation*}
$$

$* \Gamma(x)$ is called Gamma Function and defined as:

$$
\Gamma(x)=\left\{\begin{array}{l}
\int_{0^{\infty} e^{-t} t^{x-1} d t ;} \text { for } x>0 \\
\lim _{n \rightarrow \infty} \frac{n!n-1}{x(x+1)(x+2) \ldots(x+n-1)}, \text { for } x \neq 0,-1,-2, \ldots
\end{array}\right.
$$

The three moments can be obtained from equation (2-44) by substituting 1, 2, 3 for \& respectively.

Then, the estimators of the parameters can be found from tables (Gumbe1, 1966 p. 282).

The estimation of the parameters can also be obtained with the help of the first two moments and the largest observed earthquake magnitude in a period of $N$ years. Thus, it is possible to avoid the third moment which has a large variance and the disadvantage of the previous consideration which does not assure that the upper limit $\omega$ is larger than the largest earthquake magnitude observed in $N$ years. An estimated upper limit for the largest magnitudes, which is smaller than the largest observed value, does not make sense.

The estimators of $\omega, u$, and $k$ can also be found from the tables of $k$ and $N$ (Gumbel, 1963).
(2) Method of least-squares

Because the previous method needs all the values of $x_{i}$ to be available and because earthquake catalogues are, in practice, incomplete, Yegulalp and Kuo (1974) decided to approach the problem using the least-squares method. The main points of this method are:

If $x_{1}, x_{2}, \ldots x_{N}$ are the observed maximum magnitudes in a given region, and $P_{1}, P_{2}, \ldots P_{N}$ are the corresponding plotting positions, then the problem consists of minimizing the sum of squares of the differences between the theoretical and observed maximum magnitudes:

$$
\begin{equation*}
n=\sum_{i=1}^{r}\left(x_{i}-x_{i}^{\prime}\right)^{2} \text {, for } \omega>0,0<u<\omega, k>0 \tag{2-45}
\end{equation*}
$$

where

$$
\begin{equation*}
x_{i}^{\prime}=\omega-(\omega-u)\left[-\ell n \phi\left(x_{i}\right)\right]^{1 / k} \tag{2-46}
\end{equation*}
$$

The necessary and sufficient conditions to minimize $n$ are:

$$
\begin{equation*}
\frac{\partial n}{\partial \omega}=\frac{\partial n}{\partial u}=\frac{\partial n}{\partial(1 / k)}=0 \tag{2-47}
\end{equation*}
$$

and
for

$$
\begin{gather*}
\frac{\partial^{2} n}{\partial \omega^{2}}, \frac{\partial^{2} n}{\partial \omega \partial u}, \frac{\partial^{2} n}{\partial u^{2}}, \frac{\partial^{2} n}{\partial(1 / k)^{2}}, \frac{\partial^{2} n}{\partial \omega \partial(1 / k)}, \frac{\partial^{2} n}{\partial u \partial(1 / k)}>0  \tag{2-48}\\
\omega>0, \quad 0<u<\omega, \quad k>0
\end{gather*}
$$

Thus, estimators of $\omega, u$, and $k$ can be obtained by satisfying these conditions.

While this method avoids rather than takes into account the incompleteness of earthquake catalogues, it does not consider the other important factor of accuracy of earthquake magnitudes. Thus we need more powerful estimation methods which allow weights to be assigned to each individual extreme. The variance-covariance, or error matrix, for the parameters should also be examined. Given the uncertainties of the parameters, we can then examine the stability of the system and accordingly the usefulness of this procedure when applied to forecasting. This is attempted in the present study (Chapter IV).

### 2.3.6 Useful relations for forecasting procedure

a) For the first type asymptotic distribution

Epstein and Lomnitz (1966) show that assuming a Poisson distribution for the number of earthquakes with magnitude exceeding zero in a year, and if $m$, the magnitude, is a random variable distributed with cumulative distribution function:

$$
\begin{equation*}
F(m)=1-e^{-\beta m}, m \geqslant 0 \tag{2-49}
\end{equation*}
$$

then, the largest annual earthquake magnitude is distributed with the cumulative distribution function:

$$
\begin{equation*}
G(m)=\exp [-\operatorname{aexp}(-\beta m)], \quad m \geqslant 0 \tag{2-50}
\end{equation*}
$$

But equation $(2-50)$ corresponds to equation $(2-20)$ which is the first asymptotic distribution of largest values. From (2-20) taking the logarithm twice, we have:

$$
\begin{equation*}
\ln [-\ln \phi(x)]=\text { ua-ax } \tag{2-51}
\end{equation*}
$$

and if we note that:

$$
\exp [-\mathrm{a}(\mathrm{x}-\mathrm{u})]
$$

is the expected number of earthquakes, $N_{x}$ in a given year which have magnitude exceeding $x$, it follows that:

$$
\begin{equation*}
\ell \operatorname{nN}_{x}=\mathrm{ua}-\mathrm{ax} \tag{2-52}
\end{equation*}
$$

Comparison between $(2-52)$ and the widely used Gutenberg-Richter's empirical formula

$$
\begin{equation*}
\log N_{x}=a^{\prime}-b x \tag{2-53}
\end{equation*}
$$

gives

$$
\begin{equation*}
a^{\prime}=u a / \ln 10 \tag{2-54}
\end{equation*}
$$

and

$$
\begin{equation*}
\mathrm{b}=\mathrm{a} / \ln 10 \tag{2-55}
\end{equation*}
$$

from which

$$
\begin{equation*}
\frac{a^{\prime}}{b}=u \tag{2-56}
\end{equation*}
$$

which is the mode or most probable annual maximum magnitude.
The same authors have derived the following relations:
The modal earthquake magnitude in a T years period is

$$
\begin{equation*}
\tilde{m}_{T}=u+\frac{\ell n T}{a} \text { or (Curtis, 1973) } \tilde{m}_{T}=\frac{a^{\prime}}{b}+\frac{\log T}{b} \tag{2-57}
\end{equation*}
$$

The value $m_{p}$ of the annual maximum magnitude which is exceeded with probability p is

$$
\begin{equation*}
m_{p}=u-\frac{\ell n[-\ell n(1-p)]}{a} \text { or } m_{p}=\frac{a^{\prime}}{b}-\frac{\log [-\ell n(1-p)]}{b} \tag{2-58}
\end{equation*}
$$

The value $m_{p r}$ of the maximum magnitude which is exceeded with probability $p$ in an $r$ year period is

$$
\begin{equation*}
m_{p r}=m_{p}+\frac{\ell n r}{a} \text { or } m_{p r}=m_{p}+\frac{\log r}{b} \tag{2-59}
\end{equation*}
$$

Finally, the probability $P_{m r}$ of occurrence of an earthquake of magnitude greater than $m$ in an $r$ year period is:

$$
\mathrm{P}_{\mathrm{mr}}=1-\exp [-\mathrm{rexp}(-\mathrm{a}(\mathrm{~m}-\mathrm{u}))]
$$

or

$$
\begin{equation*}
\mathrm{P}_{\mathrm{mr}}=1-\exp \left[-10^{\mathrm{a}^{\prime}} \operatorname{rexp}(-\mathrm{bm} \mathrm{\ell n} 10)\right] \tag{2-60}
\end{equation*}
$$

b) For the third type asymptotic distribution

While for the first type asymptote the characteristic largest value, $u$, coincides with the mode of the distribution, in the third type asymptote this property does not exist because of the asymmetry of the distribution.

Thus, the mode of the third type asymptote is (Gumbel, 1966, p 286)

$$
\begin{equation*}
\tilde{m}=\omega-(\omega-u)\left(1-\frac{1}{k}\right)^{1 / k} \tag{2-61}
\end{equation*}
$$

which for the next $T$ years will be (Yegulalp, 1974):

$$
\begin{equation*}
\tilde{\mathrm{m}}_{\mathrm{T}}=\omega-(\omega-\mathrm{u})\left[\left(1-\frac{1}{\mathrm{k}}\right) / \mathrm{T}\right]^{1 / \mathrm{k}} \tag{2-62}
\end{equation*}
$$

From the method of moments, (eq $2-44$ for $\ell=1$ ), the expected annual largest magnitude $\overline{\mathrm{m}}$ is:

$$
\begin{equation*}
\overline{\mathrm{m}}=E(\mathrm{~m})=\omega-(\omega-u) \Gamma\left(1+\frac{1}{\mathrm{k}}\right) \tag{2-63}
\end{equation*}
$$

and the expected largest magnitude for the next $T$ years, $E\left(m_{T}\right)$, will be:

$$
\begin{equation*}
\bar{m}_{T}=E\left(m_{T}\right)=\ddot{\omega}-(\omega-u) \Gamma\left(1+\frac{1}{k}\right) T^{1 / k} \tag{2-64}
\end{equation*}
$$

In addition to these point estimates, Yegulalp (1974) considers the interval in which the largest earthquakes of the next $T$ years will lie with a given probability $a$, by solving the equations:

$$
\begin{equation*}
1-\frac{\mathrm{a}}{2}=\exp \left[-\mathrm{T}\left(\left(\omega-\mathrm{m}_{\mathrm{up}}\right) /(\omega-\mathrm{u})\right)^{\mathrm{k}}\right] \tag{2-65}
\end{equation*}
$$

and

$$
\begin{equation*}
\frac{\mathrm{a}}{2}=\exp \left[-\mathrm{T}\left(\left(\omega-\mathrm{m}_{\ell}\right) /(\omega-\mathrm{u})\right)^{\mathrm{k}}\right] \tag{2-66}
\end{equation*}
$$

where $m_{u p}$ and $m_{\ell}$ are the upper and lower bounds respectively. Thus we have:

$$
\begin{equation*}
\mathrm{m}_{u p}=\omega-(\omega-u)\left[-\frac{1}{\mathrm{~T}} \ln \left(1-\frac{\mathrm{a}}{2}\right)\right]^{1 / k} \tag{2-67}
\end{equation*}
$$

and

$$
\begin{equation*}
\mathrm{m}_{\ell}=\omega-(\omega-\mathrm{u})\left[-\frac{1}{\mathrm{~T}} \ell \mathrm{n}\left(\frac{\mathrm{a}}{2}\right)\right]^{1 / \mathrm{k}} \tag{2-68}
\end{equation*}
$$

2.3.7 Applications of the Extreme-Value theory to Seismicrisk problems

Nordquist (1945) was the first to apply the theory of extreme values for the estimation of the probability of occurrence of maximum magnitude earthquakes. He demonstrates that the theory is applicable to the seismic hazard problem. Using world data for 1904 to 1939 from Gutenberg and Richter (1941), and data from Southern California, he shows
that the probability of occurrence of maximum magnitude earthquakes obeys fairly well the first type asymptotic distribution of the largest values. Dick (1965) applied the theory to New Zealand earthquake data for 1942-1961. Since 1966, when Epstein and Lomnitz pointed out the relation among the first type asymptote parameters and the GutenbergRichter's frequency-magnitude constants (see 2.3.6), many authors have published papers using the first type asymptote Lomnitz (1966, 1974), Kárnik and Hubnerova (1968), Milne and Davenport (1969), Schenkova and Kárnik (1970), Kárnik (1971), Shakal and Willis (1972), Curtis (1973), Schenkova and Schenk (1975), Rikitake (1976), Scenkova and Kárnik (1976)].

Yegulalp and Kuo (1966) approach the occurrence of maximum magnitude earthquakes by considering the third type asymptotic distribution. They use World Seismicity data for 1904-1952 from Gutenberg-Richter (1954), and find that the regional occurrence of maximum magnitude earthquakes favours the third rather than the first type asymptotic distribution. Makjanic (1972) proposes that the distribution of the intensities of earthquakes felt at Zagreb also follows the third type asymptotic distribution. He found an intensity of 8.83 on the Mercalli, Cancani, Sieberg (MCS) scale, to be the upper limit for intensities at that city. In a second paper by Yegulalp and Kuo (1974) the early results are improved by applying a test of predictability. They predict the magnitude corresponding to the return periods $2,3,5$ and 13 years for data in the interval 1953-1965 by using previously estimated parameters. Comparison between the number of exceedances with these observed for the same magnitude show remarkably good agreement. The establishment of a forecasting procedure for largest earthquake magnitudes is also attempted by Yegulalp (1974). Using Yegulalp and Kuo's method for estimation of the parameters $\omega$, $u$ and $k$, Lilwall (1976) and Willmore and Burton (1976) analyse the seismicity and seismic hazard in the UK, and Kárnik and Schenkova (1977) in the Balkan
earthquake provinces.
Burton (1978a and b) analysed the seismicity of UK and the European area using Marquardt's technique of estimating the parameters and their uncertainties (Marquardt, 1963). This technique is further developed in the present work (Chapter IV).
2.4 Summary

The review of the statistical models which use the whole process reveals several weak points caused by the inclusion of low magnitude data, such as lack of completeness, accuracy and homogeneity. The ExtremeValue theory using only extreme magnitudes, which are more complete and accurate, seems to overcome most of these problems. The first type asymptotic distribution assumes that the distribution of extremes is unlimited in both directions, whereas, the third type asymptotic distribution is limited towards the right. Since, for the earthquake magnitude, such an upper bound must be set by the strength of the crustal rocks (Richter, 1958), the third type asymptotic distribution will be used for the seismic risk analysis in this study. It will be compared with the first type asymptotic distribution and the energy release methods (Chapter IV). First, in the next chapter, the upper bound for the earthquake magnitude will be analysed in terms of energy release.

### 3.1 Statement of the problem

Since 1944 when Gutenberg and Richter obtained their well-known empirical frequency-magnitude relation (see equation $2-8$ ), the magnitude, $m$, of an earthquake has been recognised to be as important a parameter as the location and depth. Furthermore the relation

$$
\begin{equation*}
\log E=A+B m \tag{3-1}
\end{equation*}
$$

with A and B constants, has shown the physical meaning of magnitude to be a measure of the energy $E$, released during an earthquake of magnitude $m$. However, as in paragraph 2.2.2 is pointed out, cases where discrepancies between data and the simple frequency-magnitude law (2-8) become apparent especially for low or very high magnitude ranges, and expressions alternative to (2-8) have been proposed.

Richter (1958) considering the behaviour of the frequency-magnitude experimental data, affirmed that "a physical upper bound to the largest possible magnitude must be set by the strength of the crustal rocks, in terms of the maximum strain which they are competent to support without yielding". Yegulalp and Kuo (1974) have asserted that "it is apparent on physical grounds that there must exist an upper limit to the occurrence of a maximum magnitude earthquake in each region".

But while the alternative expressions are in general agreement with the observations, they do not identify what this upper bound is, nor do they include this upper bound as an unknown parameter in the resulting distribution function. However, the more or less arbitrary introduction of an upper bound, similar to those above (eq 2-14), without considering the dependency of such an upper bound on some physical
parameters characterising the region contributes to errors in the evaluation of the other parameters of the distribution.

In the next paragraph will be demonstrated that using the simple Gutenberg-Richter relation (2-8) it is possible to include an upper bound for magnitude of earthquakes, that this is necessary, and furthermore the value of such an upper bound for a given region will be found.

### 3.2 Energy release and the upper bound for earthquake magnitude

### 3.2.1 Mathematical consideration

Let: $M_{1}$ be the most probable annual maximum magnitude (mode)
$M_{2}$ be the magnitude which corresponds to the annual rate of energy release, and
$M_{3}$ the upper bound for the earthquake magnitude in the same region. From the relations $(2-8)$ and (3-1) we have:

$$
\begin{equation*}
\ell n N=a^{\prime}-b^{\prime} m \rightarrow N=e^{a^{\prime}-b^{\prime} m} \tag{3-2}
\end{equation*}
$$

and

$$
\begin{equation*}
\ell n E=A^{\prime}+B^{\prime} m \rightarrow E=e^{A^{\prime}+B^{\prime} m} \tag{3-3}
\end{equation*}
$$

where $a^{\prime}=a \ell n 10, b^{\prime}=b \ell n 10, A^{\prime}=A \ell n 10$ and $B^{\prime}=B \ell n 10$

The number of earthquakes per year with magnitude range dm is:

$$
N(m)-N(m+d m)=e^{a^{\prime}-b^{\prime} m}-e^{a^{\prime}-b^{\prime}(m+d m)}
$$

or

$$
\begin{equation*}
d N(m)=b^{\prime} e^{a^{\prime}-b^{\prime} m} d m \tag{3-5}
\end{equation*}
$$

The annual energy release for all earthquakes with magnitude in the range $\mathrm{dm}, \overline{\mathrm{E}}$, is then:

$$
\begin{equation*}
d \bar{E}=e^{A^{\prime}+B^{\prime} m} \cdot b^{\prime} e^{a^{\prime}-b^{\prime} m} d m \tag{3-6}
\end{equation*}
$$

and the total energy release, TE, is

$$
\begin{align*}
T E & =\int_{M_{0}}^{M_{E}} \bar{E}(m) d m=b^{\prime} \int_{M_{0}}^{M_{3}} e^{A^{\prime}+B^{\prime} m} \cdot e^{a^{\prime}-b^{\prime} m_{d m}} \\
& =\left.b^{\prime} e^{a^{\prime}+A^{\prime}} \cdot \frac{1}{B^{\prime}-b^{\prime}} \cdot e^{\left(B^{\prime}-b^{\prime}\right) m}\right|_{M_{0}} ^{M_{3}} \tag{3-7}
\end{align*}
$$

where $M_{0}$ is the earthquake magnitude threshold.
Equation (3-7) is equivalent to

$$
\begin{equation*}
T E=\frac{b^{\prime}}{B^{\prime}-b^{\prime}} \cdot e^{a^{\prime}+A^{\prime}}\left[e^{\left(B^{\prime}-b^{\prime}\right) M_{3}}-e^{\left(B^{\prime}-b^{\prime}\right) M_{O}}\right] \tag{3-8}
\end{equation*}
$$

Usually b $\sim 1$ and $B=1.44$ for surface wave magnitude $M_{s}$ used in this study (Bath, 1958), so $B-b \sim 0.5$. Hence, only case when $B>b$, which is almost always observed, will be considered hereafter.

Using the definition of $M_{2}$ equation (3-8) becomes:

$$
\begin{equation*}
T E \equiv e^{A^{\prime}+B^{\prime} M_{2}}=\frac{b^{\prime}}{B^{\prime}-b^{\prime}}\left[e^{a^{\prime}+A^{\prime}+\left(B^{\prime}-b^{\prime}\right) M_{3}}-e^{a^{\prime}+A^{\prime}\left(B^{\prime}-b^{\prime}\right) M_{0}}\right] \tag{3-9}
\end{equation*}
$$

and if we notice (eq 2-56) that:

$$
\begin{equation*}
{\frac{a^{\prime}}{b^{\prime}}}^{\prime} \frac{a}{b}=M_{1} \tag{3-10}
\end{equation*}
$$

from equation (3-9) we have:

$$
\begin{equation*}
e^{B^{\prime} M_{2}}=\frac{b^{\prime}}{B^{\prime}-b^{\prime}} \cdot e^{M_{1} b^{\prime}}\left[e^{\left(B^{\prime}-b^{\prime}\right) M_{3}}-e^{\left(B^{\prime}-b^{\prime}\right) M_{o}}\right] \tag{3-11}
\end{equation*}
$$

As $M_{o} \rightarrow-\infty$ (unlimited to the left) equation (3-11) becomes:

$$
\begin{equation*}
e^{B^{\prime} M_{2}}=\frac{b^{\prime}}{B^{\top}-b^{\top}} \cdot e^{M_{1} b^{\prime}} \cdot e^{\left(B^{\prime}-b^{\prime}\right) M_{3}} \tag{3-12}
\end{equation*}
$$

or equivalently:

$$
\begin{equation*}
B^{\prime} M_{2}=\ln \left(\frac{b^{\prime}}{B^{\prime}-b^{\prime}}\right)+M_{1} b^{\prime}+\left(B^{\prime}-b^{\prime}\right) M_{3} \tag{3-13}
\end{equation*}
$$

From (3-13) by solving for $M_{2}$ and $M_{3}$ we get:

$$
\begin{equation*}
M_{2}=\frac{b^{\prime}}{B^{\prime}} M_{1}+\frac{B^{\prime}-b^{\prime}}{B^{\prime}} M_{3}+\frac{1}{B^{\top}} \ln \left(\frac{b^{\prime}}{B^{\prime}-b^{\prime}}\right) \tag{3-14}
\end{equation*}
$$

and

$$
\begin{equation*}
M_{3}=\frac{1}{B^{\prime}-b^{\prime}}\left[B^{\prime} M_{2}-b^{\prime} M_{1}-\ln \left(\frac{b^{\prime}}{B^{\prime}-b^{\top}}\right)\right] \tag{3-15}
\end{equation*}
$$

Because the mean rate of energy release $\left(M_{2}\right)$ is finite (Knopoff and Kagan, 1978 ) the $M_{3}$ from the equation (3-15), must be finite. The relations (3-14) and (3-15), because of (3-4) become:

$$
\begin{align*}
& M_{2}=\frac{1}{B}\left[b M_{1}+(B-b) M_{3}+\log \left(\frac{b}{B-b}\right)\right]  \tag{3-16}\\
& M_{3}=\frac{1}{B-b}\left[B M_{2}-b M_{1}-\log \left(\frac{b}{B-b}\right)\right] \tag{3-17}
\end{align*}
$$

and

Thus:

$$
\begin{equation*}
M_{3}=C_{1}(b) M_{2}+C_{2}(b) M_{1}+c_{3}(b) \tag{3-18}
\end{equation*}
$$

where

$$
\begin{equation*}
C_{1}(b)=\frac{B}{B-b}, C_{2}(b)=-\frac{b}{B-b} \text {, and } C_{3}(b)=-\frac{1}{B-b} \log \left(\frac{b}{B-b}\right) \tag{3-19}
\end{equation*}
$$

In addition to the previous relations it can be shown that:

$$
\begin{align*}
M_{3}-M_{2} & =C_{1}(b) M_{2}-M_{2}+C_{2}(b) M_{1}+c_{3}(b)  \tag{3-20}\\
& =\left[C_{1}(b)-\square M_{2}+c_{2}(b) M_{1}+c_{3}(b)\right.
\end{align*}
$$

and because

$$
\begin{equation*}
c_{1}(b)-1=-C_{2}(b) \tag{3-21}
\end{equation*}
$$

it follows that:

$$
\begin{equation*}
M_{3}-M_{2}=-C_{2}(b)\left[M_{2}-M_{1}\right]+C_{3}(b) \tag{3-22}
\end{equation*}
$$

similarly,

$$
\begin{equation*}
M_{3}-M_{1}=c_{1}(b)\left[M_{2}-M_{1}\right]+c_{3}(b) \tag{3-23}
\end{equation*}
$$

From the previous consideration, several features are significant, The first feature to note is that a finite upper bound to the largest earthquake magnitude is necessary to preserve a finite rate of energy release. This is inescapable even if the usual linear frequency-magnitude law applies for a given region. A second significant feature comes from equation (3-18).
probable
The upper bound $M_{3}$ is a function of the most/ annual maximum magnitude $M_{1}$, and the/ annual rate of energy release $M_{2}$. Furthermore, $M_{3}$ depends on the b value which characterizes the region from seismotectonic point of view. The relation between $M_{3}$ and $b$ coupled with the property of $b$ being different from region to region (Duda, 1965), leads to the conclusion that each region must also have its own upper bound for earthquake magnitudes expected to occur within that region.

Thus, for $a$ given region having the $a$ and $b$ of equation (2-8) and calculating the mean annual rate of energy released simply by dividing the total energy released $\Sigma E$, for a given period, $T$, with that period (ie $\Sigma E / T$ ) we can calculate $M_{2}$ and $M_{1}$ from equations (3-1) and (3-10) respectively. Equation $(3-18)$ can then be used to estimate the maximum magnitude earthquate, $M_{3}$, which may occur within that region. From the above equations the uncertainties of $M_{1}, M_{2}$, and $M_{3}$ can also be estimated.

The stationarity of $M_{3}$ is, of course, related to the stationarity of the parameters $a$ and $b$. Generally, the larger the number of earthquakes available for analysis, the more reliable are the estimates of a and $b$ (Duda, 1965) and consequently of $M_{3}$.
3.2.2 Graphical method of estimating $M_{2}$ and $M_{3}$

The graphical method of estimating $M_{2}$ and $M_{3}$ for a given region consists of plotting the cumulative energy released as a function of time, It is based on the assumption that the rate of total energy accumulation and release in a given region with similar geological structure, remains fairly constant, provided that the period of observation is long enough to average out existing periodicities. Because it is intuitively evident that the strain producing forces do not change within a time span of a few hundred years, it is reasonable to assume that in a given region, the rate of strain or energy accumulation and the possibilities of energy storage per unit volume are everywhere similar within the region, Hence the total amount of energy that may be accumulated and released remains fairly constant.

Therefore, from the graph of cumulative energy released as a function of time (see Figure 3-1), we can derive:
i) the rate of energy released as the slope of the line, $\mathrm{SS}^{\prime}$, which connects the starting point $S$ ( 0.0 energy), with the final $S^{\prime},(E$ total). If the time interval is taken to be one year $(\Delta T=1$ year $)$, then the slope $(\Delta E / \Delta T)$ represents the annual rate of energy released, and the corresponding magnitude will be $\mathrm{M}_{2}$.
ii) Since the total energy that may be accumulated and released in a given region is constant, the two lines, $B B^{\prime}$ and $C C^{\prime}$ of maximum and minimum energy released, that is the lines which pass through the end, $B B^{\prime}$, and beginning, $C C^{\prime}$, points of the active periods, $\mathrm{PP}^{\prime}$, should run parallel to each other and to the SS'. Thus the vertical


Fig 3-1 Explanatory diagram for graphical representation of the energy release method of calculating upper bound magnitude and mean annual energy release.

SOUTH AMERICA


Fig 3-2 Cumulative energy release as a function of time for Region 1, and period 1897-1964. The insert shows the values of maximum possible energy release Emax, mean annual energy releasa $\bar{E} /$ year, and the values of $a$ and $b$ of the frequency-magnitude formula. The magnitudes which correspond to Emax, $\bar{E} /$ year, and annual mode (a/b) are also listed.
distance, $E_{1} E_{2}$, of these two enveloping parallel lines, indicates the total amount of energy that may be released in the region. Hence, the vertical distance (energy) $\mathrm{E}_{1} \mathrm{E}_{2}$ is the upper limit for the energy that can be observed in the region $E_{\max }$, if the accumulated energy can be released by a single earthquake. The corresponding magnitude of such an earthquake must correspond to $M_{3}$.
iii) The horizontal distance $\mathrm{T}_{1} \mathrm{~T}_{2}$, between the two parallel lines $B B^{\prime}$ and $C C^{\prime}$ indicates the minimum time, $\operatorname{Tr}$, required for the accumulation of the maximum energy if there were no earthquakes in the meantime. This time interval will be called the "waiting time".
3.2.3 Testing the two methods

In order to test both analytical and graphical methods, the circumPacific belt is chosen. This region is seismically the most active in the world. Only $24 \%$ of the earth's seismic energy release occurs outside of it (Duda, 1965).

Data used are those compiled by Duda (1965) and his catalogue supplies the most reliable data now available for large earthquakes ( $M \geqslant 7.0$ ) and seems homogeneous over the whole world for the magnitudes and period considered.

Following Duda's subdivision, the circum-Pacific belt is divided into eight subregions as follows:
(1) Region 1 (South America)
(2) Region 2 (North America)
(3) Region 3 (Aleutians, Alaska)
(4) Region 4 (Japan, Kurile, Kamchatka)
(5) Region 5 (New Guinea, Banda Sea, Celebes, Moluccas, Philippines)
(6) Region 6 (New Hebrides, Solomon, New Guinea)
(7) Region 7 (New Zealand, Tonga, Kermadec) and
(8) Region 8 (Caroline, Marianas)

The last subregion (Region 8) is rejected because of insufficient data.
For each of the seven remaining subregions, the constants $a$ and $b$ of the frequency-magnitude relation $(2-8)$ and the annual energy release using equation (3-1) with constants $A=12,24$ and $B=1,44$ (Båth, 1958) are calculated for shallow and shallow plus intermediate earthquakes. The period considered is from 1897 to 1964. The values of $a$ and $b$ and their standard deviations $\sigma_{a}$ and $\sigma_{b}$ are calculated by the least squares method,

Tables $3-1$ and $3-2$ tabulate the values of $a, \sigma_{a}, b, \sigma_{b}$, total energy released within the whole period TE , mean annual energy released $\mathrm{TE} / \mathrm{year}$, $M_{1}, \Delta M_{1}, M_{2}, \Delta M_{2}$ as well as the values of $M_{3}$ for both analytical $M_{3 A}, \Delta M_{3 A}$, and graphical $M_{3 G}$ methods for each subregion for shallow and shallow plus intermediate earthquakes respectively, Both tables also contain the ratios $M_{2} / M_{1}, M_{3} / M_{1}$, the waiting time $T R$, and the difference betwen $M_{3}$ and $M_{2}$ for each region. Table $3-2$ also includes the same parameters for Greece for the period 1952-1972 (Galanopoulos, 1972a) Turkey for the period 1918-1973 (Alsan et al., 1975), and for the world as a whole for the period 1897-1970 (Băth, 1973). For Turkey and the world, the graphs of energy released as a function of time published by these authors are used to obtain $M_{3}$, whereas for Greece only the analytical method is applied. The results of the graphical method used for Greece from 1950-1972 are from Galanopoulos (1972a) who used a graph of strain released as a function of time, to calculate $\mathrm{M}_{3}$. The graphical method for regions and periods noted by the captions is illustrated in Figures 3-2 to 3-11.

### 3.2.4 Results and discussion

a) General features

Figures $3-2$ to $3-11$ and table $3-1$ and $3-2$ reveal several general features. The first is that as the energy release decreases and gets closer to the lower parallel bound, the possibility of having a large earthquake increases and vice versa. Thus, the lower (upper) parallel bound is

Parameters computed from shallow earthquakes

| Region | a | b | $M_{1}$ | TE*/year | $M_{2}$ | $\mathrm{M}_{3 \mathrm{~A}}$ | $M_{3 G}$ | $M_{2} / M_{1}$ | $M_{3 A} / M_{1}$ | $M_{3 A} / M_{2}$ | $\mathrm{M}_{3 A^{-M}}{ }^{\text {a }}$ | Tr(years) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| (1) South America | 5.18 $\pm .58$ | 0.74 $\pm .08$ | $\begin{aligned} & 6.96 \\ & \pm .05 \end{aligned}$ | 5.72 | $\begin{aligned} & 7.99 \\ & \pm .13 \end{aligned}$ | $\begin{aligned} & 9.05 \\ & \pm .32 \end{aligned}$ | 9.01 | 1.15 | 1.30 | 1.13 | 1.06 | 27 |
| (2) North America | $\begin{aligned} & 8.40 \\ & \pm .62 \end{aligned}$ | $\begin{aligned} & 1.15 \\ & \pm .07 \end{aligned}$ | $\begin{aligned} & 7.29 \\ & \pm .04 \end{aligned}$ | 4.50 | $\begin{aligned} & 7.93 \\ & \pm .09 \end{aligned}$ | $\begin{aligned} & 8.87 \\ & \pm .56 \end{aligned}$ | 9.00 | 1.09 | 1.22 | 1.12 | 0.94 | 31 |
| (3) Aleutians-Alaska | $\begin{aligned} & 5.86 \\ & \pm .68 \end{aligned}$ | $\begin{aligned} & 0.85 \\ & \pm .08 \end{aligned}$ | $\begin{aligned} & 6.89 \\ & \pm .05 \end{aligned}$ | 3.70 | $\begin{aligned} & 7.86 \\ & \pm .12 \end{aligned}$ | $\begin{aligned} & 9.03 \\ & \pm .37 \end{aligned}$ | 8.97 | 1.14 | 1.31 | 1.15 | 1.17 | 21 |
| (4) Japan, Kuril, Kamchatka | $\begin{aligned} & 8.14 \\ & \pm .51 \end{aligned}$ | $\begin{aligned} & 1.10 \\ & \pm .06 \end{aligned}$ | $\begin{array}{r} 7.40 \\ \pm .03 \end{array}$ | 8.50 | $\begin{aligned} & 8.11 \\ & \pm .09 \end{aligned}$ | $\begin{aligned} & 9.19 \\ & \pm .48 \end{aligned}$ | 8.91 | 1.10 | 1.24 | 1.13 | 1.08 | 25 |
| (5) N. Gun. Bunda Sea, Celebes | $\begin{aligned} & 9.17 \\ & \pm .62 \end{aligned}$ | $\begin{aligned} & 1.24 \\ & \pm .07 \end{aligned}$ | $\begin{aligned} & 7.39 \\ & \pm .03 \end{aligned}$ | 6.10 | $\begin{aligned} & 8.02 \\ & \pm .09 \end{aligned}$ | $\begin{aligned} & 9.04 \\ & \pm .83 \end{aligned}$ | 8.92 | 1.09 | 1.22 | 1.13 | 1.02 | 19 |
| (6) N. Hebr. Solom. N. Guinea | $\begin{aligned} & 9.27 \\ & \pm .81 \end{aligned}$ | $\begin{aligned} & 1.27 \\ & \pm .10 \end{aligned}$ | $\begin{aligned} & 7.27 \\ & \pm .04 \end{aligned}$ | 2.99 | $\begin{aligned} & 7.80 \\ & \pm .06 \end{aligned}$ | $\begin{aligned} & 8.61 \\ & \pm .84 \end{aligned}$ | 8.70 | 1.08 | 1.18 | 1.10 | 0.81 | 16 |
| (7) N. Zeal., Tonga, Kermadec | 6.52 $\pm .54$ | $\begin{aligned} & 0.94 \\ & \pm .07 \end{aligned}$ | $\begin{array}{r} 6.93 \\ \pm .04 \end{array}$ | 2.99 | $\begin{aligned} & 7.80 \\ & \pm .16 \end{aligned}$ | $\begin{aligned} & 8.96 \\ & \pm .54 \end{aligned}$ | 8.93 | 1.12 | 1.29 | 1.15 | 1.16 | 41 |

* Units are $10^{23} \mathrm{erg}$

Parameters computed from shallow plus intermediate earthquakes

| Region | a | b | $M_{1}$ | TE*/year | $M_{2}$ | $\mathrm{M}_{3 \mathrm{~A}}$ | $M_{3 G}$ | $\mathrm{M}_{2} / \mathrm{M}_{1}$ | $M_{3 A} / M_{1}$ | $\mathrm{M}_{3 \mathrm{~A}} / \mathrm{M}_{2}$ | $M_{3 A}-M_{2}$ | Tr (years) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| (1) South America | 7.54 $\pm .49$ | 1.04 $\pm .06$ | 7.24 $\pm .04$ | 6.44 | $\begin{aligned} & 8.03 \\ & \pm .12 \end{aligned}$ | $\begin{aligned} & 9.18 \\ & \pm .54 \end{aligned}$ | 9.10 | 1.11 | 1.27 | 1.14 | 1.15 | 25 |
| (2) North America | 8.74 $\pm .81$ | 1.19 $\pm .10$ | 7.34 $\pm .03$ | 4.74 | $\begin{aligned} & 7.94 \\ & \pm .10 \end{aligned}$ | $\begin{aligned} & 8.86 \\ & \pm .55 \end{aligned}$ | 9.00 | 1.08 | 1.21 | 1.12 | 0.92 | 33 |
| (3) Aleutians-Alaska | 6.73 $\pm .53$ | 0.95 $\pm .07$ | 7.07 $\pm .04$ | 4.06 | $\begin{aligned} & 7.89 \\ & \pm .11 \end{aligned}$ | $\begin{aligned} & 9.01 \\ & \pm .41 \end{aligned}$ | 8.78 | 1.12 | 1.27 | 1.14 | 1.12 | 20 |
| (4) Japan, Kuril, Kamchatka | 8.42 $\pm .55$ | 1.12 $\pm .07$ | 7.52 $\pm .03$ | 10.43 | $\begin{aligned} & 8.18 \\ & \pm .07 \end{aligned}$ | $\begin{aligned} & 9.11 \\ & \pm .42 \end{aligned}$ | 8.96 | 1.09 | 1.21 | 1.11 | 0.93 | 14 |
| (5) N. Gun. Bunda Sea, Celebes | 8.23 $\pm .37$ | 1.10 $\pm .05$ | $\begin{array}{r} 7.48 \\ \pm .03 \end{array}$ | 8.62 | $\begin{aligned} & 8.12 \\ & \pm .06 \end{aligned}$ | $\begin{aligned} & 9.00 \\ & \pm .36 \end{aligned}$ | 9.04 | 1.09 | 1.20 | 1.11 | 0.88 | 22 |
| (6) N. Hebr. Solom. N. Guinea | 10.29 $\pm .44$ | $\begin{array}{r} 1.40 \\ \pm .06 \end{array}$ | 7.35 $\pm .04$ | 4.19 | $\begin{aligned} & 7.90 \\ & \pm .05 \end{aligned}$ | $\begin{array}{r} 8.83 \\ \pm 2.80 \end{array}$ | 8.83 | 1.07 | 1.20 | 1.12 | 0.93 | 21 |
| (7) N. Zeal., Tonga, Kermadec | 7.36 $\pm .31$ | $\begin{aligned} & 1.04 \\ & \pm .04 \end{aligned}$ | $\begin{aligned} & 7.08 \\ & \pm .03 \end{aligned}$ | 3.63 | $\begin{aligned} & 7.86 \\ & \pm .13 \end{aligned}$ | $\begin{aligned} & 9.03 \\ & \pm .55 \end{aligned}$ | 8.97 | 1.11 | 1.27 | 1.15 | 1.17 | 30 |
| Greece (1950-1972) | $\begin{aligned} & 6.89 \\ & \pm .14 \end{aligned}$ | $\begin{aligned} & 1.05 \\ & \pm .02 \end{aligned}$ | $\begin{aligned} & 6.58 \\ & \pm .05 \end{aligned}$ | 1.37 | $\begin{aligned} & 7.45 \\ & \pm .12 \end{aligned}$ | $\begin{aligned} & 8.60 \\ & \pm .57 \end{aligned}$ | 8.52 | 1.13 | 1.31 | 1.15 | 1.15 | 30 |
| Turkey (1913-1973) | 4.11 $\pm .07$ | $\begin{aligned} & 0.68 \\ & \pm .01 \end{aligned}$ | 6.04 $\pm .05$ | 0.33 | $\begin{aligned} & 7.19 \\ & \pm .09 \end{aligned}$ | $\begin{aligned} & 8.11 \\ & \pm .23 \end{aligned}$ | 7.95 | 1.18 | 1.34 | 1.14 | 0.92 | 14 |
| World (1897-1970) | 10.44 $\pm .80$ | $\begin{aligned} & 1.29 \\ & \pm .10 \end{aligned}$ | 8.09 $\pm .03$ | 44.00 | $\begin{aligned} & 8.68 \\ & \pm .04 \end{aligned}$ | $\begin{aligned} & 9.24 \\ & \pm .53 \end{aligned}$ | 9.52 |  |  |  |  | 18 |

* Units are $10^{23}$ erg

NORTH AMERICA


Fig 3-3 Cumulative energy release as a function of time for Region 2, and period 1897-1964. For explanation of the insert see Fig 3-2.

ALEUT I ANS, ALASKA


Fig 3-4 Cumulative energy release as a function of time for Region 3, and period 1897-1964. For explanation of the insert see Fig 3-2.

JAPAN, KURILE, KAMCHATKA


Fig 3-5 Cumulative energy release as a function of time for Region 4, and period 1897-1964. For explanation of the insert see Fig 3-2.
N.GUIN., BANDA, CELEB., MOLUC., PHILL.


Fig 3-6 Cumulative energy release as a function of time for Region 5, and period 1897-1964. For explanation of the insert see Fig 3-2.


Fig 3-7 Cumulative energy release as a function of time for Region 6, and period 1897-1964.
For explanation of the insert see Fig 3-2.


Fig 3-8 Cumulative energy release as a function of time for Region 7, and period 1897-1964. For explanation of the insert see Fig 3-2.


[^1]

Fig 3-10 Cumulative energy release as a function of time for Turkey for
the period $1913-1973$ (after Alsan et al, 1975). the period 1913-1973 (after Alsan et al, 1975).


Fig 3-11 Cumulative energy release as a function of time for the world, for the period 1897-1970 (after B9th, 1973).
the bound of higher (lower) seismic risk for the region, because it is the line of maximum (minimum) storage of energy which may be released. Then, the maximum possible energy to be released in a year is the difference between $E_{\max }$ and the level of energy which already has been released during the recent past. The second feature to note is that on graphs like these the energy of large events dominates because of the logarithmic nature of equation (3-1). For the circum-Pacific belt, for example, it is clear that the very active period of the first decades of the present century dominates the analysis of the seismic behaviour of this part of the world,

A third significant feature is the very good agreement between $M_{3}$ obtained from analytical and graphical methods. This agreement shows that the assumption that the vertical distance between the two parallel lines is equivalent to the maximum possible energy which may be released, is a realistic one.

Probably the most significant feature is the close relation between $M_{1}, M_{2}$, and $M_{3}$. From table $3-1$ for shallow earthquakes we can derive the following relations:

$$
\begin{align*}
& M_{2}=(1.11 \pm 0.04) M_{1}, \\
& M_{3}=(1.25 \pm 0.05) M_{1},  \tag{3,24}\\
& M_{3}=(1.13 \pm 0.02) M_{2}, \text { and } \\
& M_{3}-M_{2}=1.03 \pm 0.13,
\end{align*}
$$

When the same procedure is applied to the parameters of Table 3-.2 for shallow plus intermediate earthquakes, we find exactly the same proportionality

$$
\begin{align*}
& M_{2}=(1,11 \pm 0.06) M_{1}, \\
& M_{3}=(1.25 \pm 0.07) M_{1},  \tag{3-25}\\
& M_{3}=(1.13 \pm 0.03) M_{2} \text { and } \\
& M_{3}-M_{2}=1.04 \pm 0.13,
\end{align*}
$$

The fact that the relations $(3-24)$ and $(3-25)$ are almost identical may be due to the same mechanism, characterizing both shallow and intermediate earthquakes in the depth range $0-400 \mathrm{~km}$ (Båth and Duda, 1963).

Because Duda's subdivision is based on the distribution and number of shallow earthquakes rather than on tectonic evidence, the relations (3-24) and (3-25) seem to be valid for tectonically very different regions and for a wide range of magnitudes. Considering the ease with which $M_{1}$ and $M_{3}$ may be derived for a region, equations $(3-24)$ and $(3-25)$ will be of great assistance for regional seismic risk considerations, particularly if, as seems likely, they have a universal character.
b) Regional features
(1) Circum-Pacific belt

From figures $3-2$ to $3-8$ it is clear that for all the circum-Pacific belt subregions after the high seismic activity during the first two decades of this century, a general pattern of decreasing activity is apparent. The fact that in 1965 in almost all cases the level of energy released is close to the lower bound means that a period of increasing seismic activity may be about to start. This, in fact, seems to be the case. Looking back to the international data file, for the period 19651974, we can see that:

Region 1 has experienced 27 quakes with magnitude range 7.0 to 8.1

| Region 2 | $"$ | $"$ | 13 | $"$ | $"$ | $"$ |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- |
| Region $3 "$ | $"$ | 13 | $"$ | $"$ | $"$ | $"$ |
| Region 4 | $"$ | $"$ | 27 | $"$ | $"$ | $"$ |
| to 8.5 |  |  |  |  |  |  |
| Region 5 | $"$ | $"$ | 35 | $"$ | $"$ | $"$ |
| Region 6 | $"$ | $"$ | 49 | $"$ | $"$ | $"$ |
| to 7.9 |  |  |  |  |  |  |


| Region 1 | $:$ | 27 years |
| :--- | :--- | :--- |
| Region 2 | $:$ | 31 years |
| Region 3 | $:$ | 21 years |
| Region 4 | $:$ | 15 years |
| Region 5 | $:$ | 19 years |
| Region 6 | $:$ | 16 years |
| Region 7 | $:$ | 41 years |

The two most active regions are situated diagonally opposite each other. These are the north-western part (Region 4) and south-eastern part (Region 1) of the circum-Pacific belt, with $M_{3}=9.2$ and $M_{3}=9.1$ respectively. Region 4 also has the shorter waiting time which means that a period of 15 years without any large earthquake is enough to accumulate energy for an earthquake with magnitude as big as 9.2 . (2) Greece and Turkey

From Figures 3-9 (Greece) and Figure 3-10 (Turkey) and Table 3-2: we can conclude the following.

The difference in $b$ values for these two seismically very active neighbouring countries shows that we expect to have a larger proportion of strong earthquakes in Turkey (smaller $b$ value) than in Greece, but the maximum expected earthquake is more likely to occur in Greece $\left(M_{3}=8.6\right)$ than in Turkey $\left(M_{3}=8.1\right)$. For both regions a new active period may be about to start. Values of the expected annual maximum and for mean rate of annual energy release, for both Greece and Turkey, are also tabulated in Table 3-.2.
(3) World

Table 3-2 also contains the parameters for shallow and intermediate earthquakes for the world as a whole. Since these earthquakes occurred in a broad variety of tectonic plates, the parameters $b$ and $M_{1}$ are only
of theoretical interest. $M_{3 G}$, however, which does not depend on the value of $b$, has a very significant meaning. It is an indication of the global upper bound for earthquake magnitudes, and gives a value of 9.5.

### 3.3 Summary

The analytical method, which uses the simple frequency-magnitude law, demonstrates the necessary existence of an upper limit to maximum magnitude earthquakes, and from both analytical and graphical methods it is possible to estimate the size of such an upper limit. As this size depends on the values of $b$, which are different from region to region, each region must also have its own upper limit. Thus, earthquake magnitude statistical models must include this upper limit as an unknown parameter, in order to be as close to reality as possible. The advantage of the methods described is the easy way of getting the size of the upper limit $M_{3}$, simply by having the $M_{1}$ and $M_{2}$ through the relations (3-15) or (3-17).

The two methods are tested in the circum-Pacific belt using Duda's catalogue and subdivision. The results show a very good agreement between $M_{3}$ obtained from both analytical and graphical methods.

The empirically obtained relations (3-24) and (3-25) are almost identical which may be due to the fact that shallow and intermediate earthquakes have the same mechanism in the depth range $0-400 \mathrm{~km}$ (Bath and Duda, 1963). If the close relation between $M_{1}, M_{2}$, and $M_{3}$ derived from tectonically very different regions has a universal character, equations (3-24) and (3-25) will be of great assistance for regional seismic risk considerations. The upper bound to magnitude and the annual rate of energy release differ by one magnitude unit.

From the regional features it is concluded that the two most active regions in the circum-Pacific belt are Region 4 (Japan, Kurile, Kamchatka)
and Region 1 (South America). These regions are situated diagonally opposite to each other and they have an upper bound of $M_{3}=9.2$ and $M_{3}=9.1$ respectively. Region 4 also has the shorter waiting time ( 15 years). In almost all the circum- Pacific belt a general pattern of decreasing activity is observed. However, since the cumulative energy release is in all cases close to the lower bound (maximum energy storage) a period of increasing seismic activity started after 1965.

For Greece, Turkey, using published graphs, it is found that the upper bound to earthquake magnitude is $M_{3}=8.6$ and $M_{3}=8.1$ respectively.

For the world as a whole, an upper bound with the value of $M_{3}=9.5$ is found to indicate the global upper bound for earthquake magnitudes.

In the next chapter, the problem of the upper bound for earthquake magnitude is analysed with the Extreme-Value statistical models, the third type asypmtotic distribution of which seems to fulfil almost all the conditions necessary to be a useful tool in obtaining estimates of the frequencies and recurrence times for large magnitude earthquakes, A method of estimating the parameters, which allows variance-covariance matrix for the parameters to be calculated is described. Thus, comparisons between parameters obtained by the analytical and graphical methods of this chapter, and the same parameters from the Extreme-Value statistics method are made, and relations between the simple frequency-magnitude law and third type asymptote results are established.

THIRD TYPE ASYMPTOTIC DISTRIBUTION OF GUMBEL AND STRAIN ENERGY RELEASE RELATIONS

### 4.1 Introduction

A method of determining the parameters of Gumbel's third type asymptotic distribution is described in this chapter, and these parameters are then related to physical strain energy release.

### 4.1.1 Uncertainties - weights

The need for a statistical model which includes the largest expected earthquake magnitude in a given region as an unknown parameter arises from the arguments in Chapter III. Hence the third type asymptotic distreibution of Gumbel, being such a statistical model, is used for seismic risk evaluation in many regional seismic studies. But when reviewing these studies, it soon becomes apparent that none of them considers the uncertainties in the parameters involved, nor are the uncertainties of the data used taken into account. However, these factors are very important for any statistical treatment.

In this chapter it is demonstrated that using the non-1inear leastsquares fitting method, it is possible to calculate the uncertainties of the estimated parameters by extending it to obtain variance-covariance or error, matrices. The method described in this chapter also allows weight for each individual extreme magnitude to be assigned. Thus, the calculated parameters with their uncertainties may constitute an improved basis for seismic risk estimations.

### 4.1.2 Relations between parameters of strain energy release and the <br> third type asymptote

The physical meaning of the parameters of the first type asymptote is related to that of the parameters of the frequency-magnitude formula as shown by Epstein and Lomnitz (1966) (see paragraph 2.3.6). The lack
of any relation between the parameters of the third asymptote, and the parameters of other formulae with clear physical meaning (eg frequencymagnitude), has previously meant that no physical interpretation of these parameters has been attempted.

Although a direct comparison between the parameters of the non-linear third type asymptote and those of these linear formulae cannot be made, it is still possible to use the procedure of Chapter III to relate these parameters with the parameters of strain energy release such as $M_{1}, M_{2}$, and $M_{3}$. So, the parameters of the third type asymptote can be expressed in terms of physical quantities like $M_{1}, M_{2}$, and $M_{3}$.

### 4.1.3 Testing region

The circum-Pacific belt is again chosen as a region for testing both the method and the relations between the expressions which correspond to those of $M_{1}, M_{2}$, and $M_{3}$ derived in Chapter III. It is then demonstrated that using the third type asymptotic distribution of Gumbel, with the uncertainties of the parameters computed, it is possible to establish a forecasting procedure for the maximum magnitude earthquakes likely to occur in the most seismically active region in the world.
4.2 Estimation of the parameters

### 4.2.1 Non-1inear least-squares methods

Because of non-1inearity in the parameters $\omega, u$ and $k$ of the third type asymptotic distribution of Gumbel (see paragraph 2.3)

$$
\begin{equation*}
\phi(x)=\exp ^{0}\left[-\left(\frac{\omega-x}{\omega-u}\right)^{k}\right] \tag{4-1}
\end{equation*}
$$

the conventional linear least-squares method cannot be directly applied to estimate them. The problem is approached by using the non-1inear leastsquares method. The data input consists of earthquake maximum magnitudes and some additional requirements are met by this method. These requirements
are:
i) to allow weight for each individual earthquake magnitude
ii) To take into account the years for which earthquake magnitudes are absent or not available, and
iii) to compute the variance-covariance or error matrices from which the uncertainties of the parameters can be calculated.

The non-linear least-squares methods such as grid search, gradient search, linearization of fitting function etc, are like the least-squares methods based on the following two principles:
i) A measure of goodness of fit between the data and postulated curve, 2 $x$, can be defined as:

$$
\begin{equation*}
x^{2}=\sum_{i=1}^{N}\left[\frac{1}{\sigma_{i}^{2}}\left[y_{i}-y_{i}\left(x_{i}\right)\right]^{2}\right] \tag{4-2}
\end{equation*}
$$

where $\sigma_{i}$ are the uncertainties in the $N$ data points $y_{i}$
ii) According to the method of least-squares, the optimum values of the parameters $a_{j}$ are obtained by minimizing $x^{2}$ with respect to each of the parameters simultaneously. This gives
$\frac{\partial}{\partial a_{j}} x^{2}=\frac{\partial}{\partial a_{j}} \sum_{i=1}^{N}\left[\frac{1}{\sigma_{i}^{2}}\left[y_{i}-y_{\phi}\left(x_{i}\right)\right]^{2}\right]=0, j=1, n$
where $n$ is the number of parameters.

It is generally not convenient to derive an analytical expression for calculating the parameters of a non-linear function $y(x)$. Instead, $x^{2}$ must be considered a continuous function of the $n$ parameters $a_{j}$, describing a hypersurface in n-dimensional space, and the space must be searched for the minimum of $x^{2}$.

There are a number of ways of finding this minimum value: by searching parameter space using a grid or gradient search, approximate analytical methods such as parabolic extrapolation of $x^{2}$, or linearization of the fitting function using a Taylor expansion. The gradient search method and the method of linearizing the function are combined in the algorithm chosen in this study.

One disadvantage inherent in the analytical methods is that while they converge quite rapidly to the point of minimum $\chi^{2}$ from points nearby, they cannot be relied on to approach the minimum with any accuracy from a point outside the region where the $\chi^{2}$ hypersurface is approximately parabolic. In contrast, the gradient search is ideally suited for approaching the minimum from far away, but it does not converge rapidly when in the immediate vicinity of the minimum. Therefore we need an algorithm which behaves like a gradient search for the first portion of a search and behaves like an analytical solution as the search converges. Such an algorithm eventually chosen is the Marquardt algorithm.

### 4.2.2 The Marquardt (1963) algorithm

The main features of this algorithm (Bevington 1969) are obtained as follows.

Expand the fitting function $y(x)$ to first order in a Taylor expansion as a function of the parameters $a_{j}$

$$
\begin{equation*}
y(x)=y_{0}(x)+\sum_{j=1}^{n}\left[\frac{\partial y_{0}(x)}{\partial a_{j}} \delta a_{j}\right] \tag{4-4}
\end{equation*}
$$

where $y_{0}(x)$ is the value of the fitting function at the starting point $x$. This gives a function which is linear in the parameter increments $\delta a_{j}$. To this approximation, $X^{2}$ can be expressed explicitly as a function of the parameter increments $\delta a_{j}$ :

$$
x^{2}=\sum_{i=1}^{n}\left[\frac{1}{\sigma_{i}^{2}}\left[y_{i}-y_{0}\left(x_{i}\right)-\sum_{j=1}^{n}\left[\frac{\partial y_{0}\left(x_{i}\right)}{\partial a_{j}} \delta a_{j}\right]\right]^{2}\right] \quad \text { (4-5) }
$$

Following the method of linear least-squares, $\chi^{2}$ is minimised with respect to each of the parameter increments $\delta a_{j}$ by setting the derivatives equal to 0 ,

$$
\begin{equation*}
\frac{\partial x^{2}}{\partial \delta a_{k}}=-2 \sum_{i=1}^{N}\left[\frac { 1 } { \sigma _ { i } ^ { 2 } } \left[y_{i}-y_{0}\left(x_{i}\right)-\sum_{j=1}^{n}\left[\frac{\partial y_{0}\left(x_{i}\right)}{\partial a_{j}} \delta a_{j}\right]\left[\frac{\partial y_{0}\left(x_{i}\right)}{\partial a_{k}}\right]=0\right.\right. \tag{4-6}
\end{equation*}
$$

This yields a set of $n$ simultaneous equations, which we can treat as a matrix equation
or

$$
\begin{align*}
b_{k} & =\sum_{j=1}^{n}\left(\delta a_{j} A_{j k}\right), k=1, n \\
b & =\delta a A \tag{4-7}
\end{align*}
$$

where b is a row matrix whose elements are

$$
\begin{equation*}
b_{k}=\sum_{i=1}^{N}\left[\frac{1}{\sigma_{i}^{2}}\left[y_{i}-y_{0}\left(x_{i}\right)\right] \frac{\partial y_{0}\left(x_{i}\right)}{\partial a_{k}}\right] \tag{4-8}
\end{equation*}
$$

and $A$ is a symmetric matrix of order $n$ whose elements are

$$
\begin{equation*}
A_{j k}=\sum_{i=1}^{N}\left[\frac{1}{\sigma_{i}^{2}} \frac{\partial y_{0}\left(x_{i}\right)}{\partial a_{j}} \frac{\partial y_{0}\left(x_{i}\right)}{\partial a_{k}}\right] \tag{4-9}
\end{equation*}
$$

and this is called the curvature matrix because of its relationship to the curvature of $\chi^{2}$ in parameter space.

The Marquardt algorithm combines the gradient search with the method of linearizing the fitting function by increasing the diagonal terms of the curvature matrix $A$ by a factor which controls the interpolation of the algorithm between the two extremes. Thus equation (4-7) becomes

$$
\begin{align*}
b & =\delta a A^{\prime} \\
A_{j k}^{\prime} & = \begin{cases}A_{j k}^{(\mu+1)} & \text { for } j=k \\
A_{j k} & \text { for } j \neq k\end{cases} \tag{4-10}
\end{align*}
$$

If $\mu$ is very small (near to the minimum), equations (4-10) are similar to the solution of equations (4-7) developed from a Taylor's expansion. If $\mu$ is very large (far from the minimum), the diagonal terms of the curvature matrix dominate, and the matrix equation degenerates in $n$ separate equations

$$
\begin{equation*}
b_{j}=\mu \delta a_{j} A_{j} \tag{4-11}
\end{equation*}
$$

which yields increments $\delta a_{j}$ in the same direction as the gradients $b_{j}$ of equation (4-8) but with lengths scaled by $A_{j j}$ and reduced by a factor of $\mu$. The solution for the parameter increments $\delta a_{j}$ follows from equation (4-10)

$$
\begin{equation*}
\delta a_{j}=\sum_{k=1}^{n}\left(b_{k} E^{\prime} j k\right) \tag{4-12}
\end{equation*}
$$

where $b_{k}$ are given in equation (4-8) and the matrix $E^{\prime}$ is the inverse of the matrix $A$ whose elements are given in equations (4-9).

Thus when the starting points are far away from the point of minimum $\chi^{2}$ the gradient search method brings them quite rapidly near to it, and when this method starts to suffer markedly as the search approaches the minimum, the linearization of fitting function method, more suitable for points nearby, continues the trial until the minimum $\chi^{2}$ has been reached. Finally, the uncertainties of the parameters can be obtained from
the inverse of the curvature matrix which, in the limits of the approximation we have made, is

$$
\begin{equation*}
\sigma^{2} a_{j}=E^{\prime}{ }_{j j} \tag{4-13}
\end{equation*}
$$

$E^{\prime}$ is called the error matrix because it contains most of the information needed to estimate the errors. For $\mathrm{n}=3$ it becomes:

$$
\left[\begin{array}{ccc}
\operatorname{Var} a_{1} & \operatorname{Cov} a_{1} a_{2} & \operatorname{Cov} a_{1} a_{3}  \tag{4-14}\\
\operatorname{Cov} a_{2} a_{1} & \operatorname{Var} a_{2} & \operatorname{Cov} a_{2} a_{3} \\
\operatorname{Cov} a_{3} a_{1} & \operatorname{Cov} a_{3} a_{2} & \operatorname{Var} a_{3}
\end{array}\right]
$$

### 4.2.3 Computations

For the purpose of this study a computer program has been written using the previous algorithm. A full description of the computation procedure and a complete annotated listing of this program is given in Appendix A.

### 4.3 Prediction Uncertainties

In section 2.3.6 the most probable maximum modal magnitude, $\tilde{m}(\mathrm{n})$, for the next $n$ years, and the interval in which the maximum magnitude will lie with a given probability leve1 ' $\alpha$ ', were found to be:

$$
\begin{align*}
\tilde{m}(n) & \simeq \omega-(\omega-u)\left[\left(1-\frac{1}{k}\right) / n\right]^{1 / k}  \tag{4-15}\\
m_{u p}(n) & \simeq \omega-(\omega-u)\left[-\frac{1}{n} \ln \left(1-\frac{\alpha}{2}\right)\right]^{1 / k} \tag{4-16}
\end{align*}
$$

and

$$
\begin{equation*}
\mathrm{m}_{\ell}(\mathrm{n}) \simeq \omega-(\omega-u)\left[-\frac{1}{\mathrm{n}} \ln \left(\frac{\alpha}{2}\right)\right]^{1 / k} \tag{4-17}
\end{equation*}
$$

where $\tilde{m}(n)$ is the mode of the next $n$ years and $m_{u p}(n)$ and $m_{l}(n)$ the upper and lower bounds, or the interval, at probability level $\alpha$.

It is obvious that, using equations $(4-16)$ and $(4-17)$, the $m_{u p}(n)$ and $m_{l}(n)$ tend towards $w$ as $n \rightarrow \infty$. Thus the upper bound, $m_{u p}(n)$ cannot exceed the parameter $\omega$, despite the lack of precision in estimates of $\omega$ (Burton, 1978b). However, we can overcome this disadvantage by assigning uncertainties to the upper and lower bounds. It is possible to use the approximation formula:
$\sigma^{2}(m) \simeq \sigma^{2} \omega\left(\frac{\partial m}{\partial \omega}\right)^{2}+\sigma^{2} u\left(\frac{\partial m}{\partial u}\right)^{2}+\sigma^{2}(1 / k)\left(\frac{\partial m}{\partial(1 / k)}\right)^{2}+2 \sigma^{2} \omega u\left(\frac{\partial m}{\partial u}\right)\left(\frac{\partial m}{\partial u}\right)+\ldots$
where $\sigma^{2}$ is the variance or covariance of the parameter involved, and the partial derivatives of equations $(4-18)$ are calculated according to the particular form of prediction formula for $m(n)$. Using equations (4-18) for the uncertainties of the bounds, it can be shown that

$$
\begin{gathered}
\sigma^{2}\left[m_{u p}(n)\right], \sigma^{2}\left[m_{\ell}(n)\right] \rightarrow \sigma^{2}(\omega) \\
n \rightarrow \infty
\end{gathered}
$$

A11 the partial derivatives of the equations $(4-16)$ and (4-17) with respect to $u$ and $1 / k$ tend to zero, because of a factor $\frac{1}{n}$, whereas the partial derivatives with respect to $\omega$ tend to unity. Making use of the variance covariance matrix of $(4-14)$, it is possible to use equation ( $4-18$ ) and to assign uncertainties on $(4-16)$ and $(4-17)$ which have the advantage that they include the probability level as a parameter.

A full description of the above calculation, as well as a computer program which has been written for this purpose, are also included in the main program (Appendix A).

Because the third type asymptotic distribution is a three-parameter curve, its parameters cannot be directly related to those of linear frequencymagnitude or energy-magnitude formulae. However, it is possible to compare. results derived from the expressions used to describe the same physical quantities, such as most probable expected maximum magnitude, or annual energy release etc. So relations among the parameters of the different models can be established.

### 4.4.1 The mode

For the third type asymptotic distribution, the most probable annual maximum or annual mode, $\tilde{\mathrm{x}}$, is given (see eq $2-61$ ) as:

$$
\begin{equation*}
\tilde{x}=\omega-(\omega-u)(1-1 / k)^{1 / k} \tag{4-20}
\end{equation*}
$$

The same quantity, using the whole available data set is $M_{1}=a / b$. Then $\tilde{x}$ may be compared with $M_{1}$, that is

$$
\begin{equation*}
M_{1}=\frac{a}{b} \simeq \tilde{x}=\omega-(\omega-u)(1-1 / k)^{1 / k} \tag{4-21}
\end{equation*}
$$

which for the $T$ year mode (see eq $2-42$ and $2-47$ ) becomes

$$
\begin{equation*}
\mathrm{M}_{\mathrm{T}}=\frac{\mathrm{a}}{\mathrm{~b}}+\frac{\log \mathrm{T}}{\mathrm{~b}} \simeq \tilde{\mathrm{x}}_{\mathrm{T}}=\omega-(\omega-\mathrm{u})[(1-1 / \mathrm{k}) / \mathrm{T}]^{1 / \mathrm{k}} \tag{4-22}
\end{equation*}
$$

### 4.4.2 The mean annual energy release

The expected yearly number of events, $N_{x}$, over some magnitude $x$, is connected with its return period $T_{x}$, (Epstein and Lomnitz, 1966) by the relation

$$
\begin{equation*}
N_{x}=\frac{1}{T_{x}} \tag{4-23}
\end{equation*}
$$

For the third type asymptotic distribution we have from equations (4-22) and (4-23):

$$
\begin{equation*}
\left(\frac{\omega-x}{\omega-u}\right)=\left(\frac{1-1 / k}{T_{x}}\right)^{1 / k}=\left[(1-1 / k) N_{x}\right]^{1 / k} \tag{4-24}
\end{equation*}
$$

and

$$
\begin{equation*}
N_{x}=\left(\frac{\omega-x}{\omega-u}\right)^{k} /\left(1-\frac{1}{k}\right) \tag{4.25}
\end{equation*}
$$

Equation (4-25) for $x=\tilde{x}$ becomes: $N_{x}=1$, because the annual mode is the earthquake magnitude which is expected to be exceeded once in a given year. The number of earthquakes with magnitude in the range $d x$ in a given year is:

$$
\begin{equation*}
d N_{x}=\frac{k}{(1-1 / k)} \cdot \frac{1}{(\omega-u)^{k}}(\omega-x)^{k-1} d x \tag{4-26}
\end{equation*}
$$

and the annual energy release, $E_{d x}$, for all earthquakes with magnitude in the range $d x$ is

$$
\begin{aligned}
E_{d x} & =e^{A+B x_{d N}} \\
& =e^{A+B x} \cdot \frac{k}{(1-1 / k)} \cdot \frac{1}{(\omega-u)^{k}} \cdot(\omega-x)^{k-1} d x
\end{aligned}
$$

where $A$ and $B$ are the constants of the energy-magnitude equation

$$
\begin{equation*}
\ell n E=A+B x \tag{4-28}
\end{equation*}
$$

Then the total annual energy release, TE, from (4-27) becomes

$$
\begin{equation*}
T E=\int_{-\infty}^{\omega} e^{A+B x^{\prime}} d N=\int_{-\infty}^{\omega} e^{A+B x} \cdot C \cdot(\omega-x)^{k-1} d x \tag{4-29}
\end{equation*}
$$

where

$$
\begin{equation*}
\epsilon=\frac{k}{(1-1 / k)} \cdot \frac{1}{(\omega-u)^{k}} \tag{4-30}
\end{equation*}
$$

or

$$
\begin{equation*}
T E=C \cdot e^{A} \int_{-\infty}^{\omega} e^{B x}(\omega-x)^{k-1} d x \tag{4-31}
\end{equation*}
$$

If we put

$$
\begin{equation*}
\omega-x=y \rightarrow x=\omega-y \tag{4-32}
\end{equation*}
$$

then, by noting that

$$
\begin{align*}
& x \rightarrow-\infty, y \rightarrow \infty \\
& x \rightarrow \infty, y \rightarrow 0  \tag{4-33}\\
& d x=-d y
\end{align*}
$$

and

$$
e^{B x}=e^{B \omega} \cdot e^{-B y}
$$

the equation ( $4-31$ ) becomes

$$
\begin{equation*}
T E=C \cdot e^{A+B \omega} \int_{0}^{\infty} y^{k-1} e^{-B y} d y \tag{4-34}
\end{equation*}
$$

But

$$
\begin{equation*}
\int_{0}^{\infty} y^{k-1} e^{-B y} d y=\frac{\Gamma(k)}{B^{k}}(B>0, k-1>-1) \tag{4-35}
\end{equation*}
$$

where $\Gamma(k)$ is the Gamma-function of $k$. Equation ( $4-34$ ) becomes:

$$
\begin{equation*}
T E=C \cdot e^{A+B \omega} \cdot \frac{\Gamma(k)}{B^{k}} \equiv e^{A+B X_{2}} \tag{4-36}
\end{equation*}
$$

where $X_{2}$ is the magnitude which corresponds to this TE. By re-substituting $C$ from equation $(4-30)$ we have:

$$
\begin{equation*}
\frac{k^{2}}{k-1} \cdot \frac{1}{(\omega-u)^{k}} \cdot e^{A} \cdot e^{B \omega} \cdot \frac{\Gamma(k)}{B^{k}} \equiv e^{A} \cdot e^{B X_{2}} \tag{4-37}
\end{equation*}
$$

or

$$
\begin{equation*}
e^{B X_{2}}=\frac{k^{2}}{(k-1)} \cdot \frac{\Gamma(k)}{(\omega-u)^{k} \cdot B^{k}} \cdot e^{B \omega} \tag{4-38}
\end{equation*}
$$

Taking logarithms on both sides of $(4-38)$ and solving for $X_{2}$ we have:

$$
\begin{equation*}
X_{2}=\omega-\frac{k \ln B}{B}+\frac{1}{B} \ln \left[\frac{k^{2}}{k-1} \cdot \frac{\Gamma(k)}{(\omega-u)^{k}}\right] \tag{4-39}
\end{equation*}
$$

If we use Bath's (1958) energy-magnitude constants in equation (4-28) this gives $B=1.44 \ln 10$, and equation (4-39) finally becomes

$$
\begin{equation*}
X_{2}=\omega-0.3615 k+\frac{1}{1.44} \log \left[\frac{k^{2}}{k-1} \cdot \frac{\Gamma(k)}{(\omega-u)^{k}}\right] \tag{4-40}
\end{equation*}
$$

Then, the $M_{2}$ which corresponds to the mean annual energy release by using the whole data set (see paragraphs 3.2 .1 and 3.2 .2 ), should be comparable with $X_{2}$, which is the same magnitude derived from the third type asymptotic distribution parameters. That is:
$M_{2}=\frac{1}{B}\left[b M_{1}+(B-b) M_{3}+\log \left(\frac{b}{B-b}\right)\right] \simeq X_{2}=\omega-0.3615 k+\frac{1}{1.44} \log \left[\frac{k^{2}}{k-1} \cdot \frac{\Gamma(k)}{(\omega-u)^{k}}\right]$

The $M_{3}$ upper limit for earthquake magnitudes, which is derived from the whole process, should be comparable with $\omega$ of the third type asymptotic distribution.
4.5 Testing the Third Type Asymptotic Distribution Method
4.5.1 Data

Verification of the relations ( $4-21$ ) and (4-41) depends on the agreement of the results when these two models are applied to the same area. The circum-Pacific belt is again chosen as the testing area. Periods of investigation are from 1897 to 1964 , as in Chapter III, and from 1897 to 1975 inclusive. This is the longest time span available.

For the above periods of investigation the data sets used are:
i) those compiled by Duda (1965)
ii) Gutenberg and Richter's (1954) catalogue, for years in which no earthquakes have been reported in Duda's catalogue, and
iii) since 1956 from the Institute of Geological Sciences seismicity file (Burton 1978c).
4.5.2 Data treatment
a) Missing years

For each of the seven subregions of the circum-Pacific belt, the largest yearly observed earthquake magnitude is taken and these are ranked in increasing size. Then the plotting position of the ith observation, $x_{i}$, is defined as:

$$
\begin{equation*}
p_{i}=\frac{i-0.44}{N+0.12} \tag{4-42}
\end{equation*}
$$

where i is the rank, counted from below, and N is the total number of observations.

The couples ( $p_{i}, x_{i}$ ) can be plotted on Gumbel's probability paper.

Here the parameters are computed with the help of the least-squares methods described in paragraphs 2.3.5 (first type) and 4.2.1 (third type). However, even for these very active regions, there are "missing years" without recorded earthquakes. This is due to the threshold of the magnitudes recorded in the catalogues, or to the instruments' detectability during operation. So the problem of filling the missing years arises, and this becomes even greater as the method is applied to smaller or less active regions.

It is possible to reduce the number of missing years by taking instead of one year extremes, two or more year extremes. Equation (2-20) converts back to one year maxima. But although this method may reduce the number of missing years, it still does not ensure that there will be no such intervals. An alternative solution may be the filling of all these empty intervals with a specific magnitude, for example with the magnitude which corresponds to intensity $I=V$ (Schenkova and Kárnik, 1976, 1978). However, as the number of these artificially created maximum values becomes large the curve is forced to pass through them because of their cumulative weight. Furthermore, as these points are clustered at one part of the distribution, they may influence the slope of the curve and, consequently, the estimation of the parameters.

In this study Yegulalp and Kuo's (1974) consideration is adopted. If during $N$ years there are j missing extremes, the first actual observed extreme is ranked as $j+1$, assuming the first $j$ of the $N$ observations are not available. The advantages of this approach are:
i) The rank of the remaining magnitudes is the same as it would be if there were no missing years, provided that during these missing years, no earthquake with magnitude greater than the first actual observed magnitude occurred.

It allows the completeness of the data set to be taken into account since this method of treatment does not disturb the distribution of the actual observed magnitudes. Thus, magnitudes which prove to be incompletely reported, see Chapter V, can be omitted from the calculations without affecting the distribution of the remaining extremes.
b) Weighting the data

One of the refinements which the method of estimation of the parameters allows, is the possibility of assigning weight, or uncertainty to each individual extreme magnitude. Such an uncertainty may simply reflect the date of the observations, the size of the earthquakes and the sensitivity of instruments used, or general improvement in magnitude determiation during this century.

The main sources for Duda's catalogue are Gutenberg's (1956) work for the period 1897-1903, and Gutenberg and Richter's (1954) catalogue for the period 1904-1952. For both periods the magnitudes are those revised by Richter (1958).

Considering the comments (Gutenberg, 1954, page 609, and Gutenberg and Richter, 1954, page 10) about the accuracy of the magnitudes listed, in these two works, and also because Duda's magnitudes are those of Richter (1958) converted from unified magnitude $m$, to the surface wave magnitude Ms, the system used for weighting the annual extremes chosen in this study is:

Period (Duda's original source)

1. 1897-1903 (Gutenberg, 1956) and

1904-1917 (Gutenberg and Richter, 1954) : $\pm 0.6$
2. 1904-1917 (Duda's addition of 146 events): $\pm 0.4$
3. 1918-1953 (Gutenberg and Richter, 1954)

3a. when a magnitude is assigned to the tenth of a unit$\pm 0.3$

Standard deviation assigned

```
    3b}\mathrm{ . when it is assigned to the nearest
    quarter : }\pm0.
    3c. when as in 3a with the addition of }\pm : \pm0.
    3d. when as in 3b with the addition of }\pm: : \0.
    4. 1954-1975 (Duda and IGS file) : 
```


### 4.5.3 Results and discussion.

a) The parameters

The estimated parameters with their uncertainties for both first and third type asymptotic distributions are tabulated in Tables 4-1 and 4-2 respectively. These include the seven subregions of the circum-Pacific belt for sampling periods 1897 to 1964 and 1897 to 1975. Figures 4-1a to $4-7 \mathrm{~b}$, inclusive, show the two distribution curves and the observed annual maximum magnitudes.

In both tables there are two additional columns. One contains the number of missing annual extremes, labelled "missing years", and the other, "chi-square", contains the difference between the reduced chi-square for the first, $\rho_{1}$, and the same quantity for the third type asymptotic distribution, $\rho_{3}$. The reduced chi-square $\rho$, for $F$ degrees of freedom, given by $x^{2} / F$, is taken as a measure of goodness of fit.

Tables $4-1$ and $4-2$ and figures $4-1$ a to $4-7 \mathrm{~b}$ inclusive, show several significant features. First feature to note is that for each region the characteristic value $u$, with the probability $\phi(u)=1 / e$ of not being exceeded during a year, is well determined.

Secondly, when the data in a region shows little curvature, and $\lambda=1 / k$ is small, and therefore $\omega$ is high, as in regions 1,5 , and 6 , then these parameters are accompanied by large uncertainties. This may indicate that the time span proves insufficient to establish curvature. In fact, when comparing the two tables, it is apparent that as the sampling duration lengthens, and the observational data improves, the values usually tend to stabilize.

Estimated Parameters of Asymptotic distributions (1897-1964)

| Region | Third type |  |  |  |  |  | First type |  |  |  | Missing years | chi square$\rho_{1}^{-\rho} 3$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | $\omega$ | $\sigma_{\omega}$ | u | $\sigma_{i}$ | $\lambda=\frac{1}{\mathrm{k}}$ | ${ }^{\sigma} \lambda$ | $u$ | $\sigma_{v}$ | $\frac{1}{\mathrm{a}}$ | ${ }^{\circ} \frac{1}{\mathrm{a}}$ |  |  |
| (1) South America | 10.16 | 1.13 | 7.08 | 0.04 | 0.197 | 0.091 | 7.08 | 0.04 | 0.477 | 0.028 | 8 | +0.094 |
| (2) North America | 9.14 | 0.51 | 7.14 | 0.04 | 0.320 | 0.110 | 7.11 | 0.03 | 0.501 | 0.031 | 3 | +0.186 |
| (3) Aleutians, Alaska | 9.66 | 0.63 | 6.78 | 0.04 | 0.260 | 0.083 | 6.84 | 0.04 | 0.486 | 0.029 | 13 | +0.216 |
| Japan <br> (4) Kurile <br> Kamchatka | 9.30 | 0.34 | 7.38 | 0.03 | 0.327 | 0.076 | 7.34 | 0.03 | 0.437 | 0.024 | 2 | +0.348 |
| New Guinea, Banda Sea <br> (5) Celebes, Moluccas Philippines | 10.00 | 1.11 | 7.42 | 0.03 | 0.194 | 0.098 | 7.40 | 0.03 | 0.425 | 0.027 | 2 | +0.069 |
| New Hebrides <br> (6) Solomon New Guinea | 9.44 | 1.11 | 7.23 | 0.03 | 0.220 | 0.125 | 7.22 | 0.04 | 0.397 | 0.031 | 6 | +0.060 |
| New Zealand <br> (7) Tonga Kermadec | 8.95 | 0.35 | 6.89 | 0.04 | 0.357 | 0.091 | 6.95 | 0.04 | 0.441 | 0.027 | 10 | +0.316 |
| (8) World | 9.23 | 0.25 | 8.17 | 0.03 | 0.358 | 0.056 | 8.12 | 0.03 | 0.285 | 0.025 | 0 | +0.299 |

Estimated Parameters of Asymptotic distributions ( $1897-1975_{\text {sep }}$ )

| Region | Third type |  |  |  |  |  | First type |  |  |  | chi square$\rho_{1}^{-\rho} 3$ | Missing years | Observed Maximum |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | $\omega$ | $\sigma_{\omega}$ | u | $\sigma_{u}$ | $\lambda=\frac{1}{\mathrm{k}}$ | $\sigma_{\lambda}$ | u | ${ }_{v}$ | $\frac{1}{\mathrm{a}}$ | $\sigma \frac{1}{\mathrm{a}}$ |  |  |  |
| (1) South America | 9.92 | 0.94 | 7.09 | 0.03 | 0.208 | 0.086 | 7.10 | 0.03 | 0.457 | 0.025 | +0.101 | 10 | 8.9 |
| (2) North America | 9.01 | 0.39 | 7.11 | 0.03 | 0.352 | 0.094 | 7.10 | 0.03 | 0.471 | 0.027 | +0.257 | 6 | 8.6 |
| (3) Aleutians, Alaska | 9.77 | 0.69 | 6.79 | 0.04 | 0.235 | 0.077 | 6.85 | 0.04 | 0.469 | 0.027 | +0.178 | 16 | 8.7 |
| (4) Japan <br>  Kamchatka | 9.33 | 0.34 | 7.38 | 0.03 | 0.307 | 0.069 | 7.34 | 0.03 | 0.429 | 0.022 | +0.329 | 2 | 8.9 |
| New Guinea  <br> (5) Banda Sea, Celebes <br>  Maluccas, Philippines | 9.58 | 0.72 | 7.41 | 0.03 | 0.235 | 0.097 | 7.38 | 0.03 | 0.414 | 0.025 | +0.091 | 4 | 8.7 |
| New Hebrides <br> (6) Solomon <br> New Guinea | 9.56 | 1.14 | 7.25 | 0.03 | 0.198 | 0.112 | 7.24 | 0.03 | 0.381 | 0.027 | +0.050 | 6 | 8.6 |
| New Zealand <br> (7) Tonga <br> Kermadec | 8.91 | 0.33 | 6.87 | 0.04 | 0.359 | 0.089 | 6.95 | 0.04 | 0.429 | 0.025 | +0.294 | 14 | 8.7 |
| World | 9.13 | 0.15 | 8.12 | 0.03 | 0.395 | 0.054 | 8.07 | 0.03 | 0.299 | 0.022 | +0.326 | 0 | 8.9 |

SOUTH AMERICA 1


Fig 4-la Asymptotic distribution curves of extreme values of magnitude for Region 1, for the period 1897-1964. Straight line indicates the first type of extreme value distribution, curved line indicates the third type, + indicates observed annual maximum magnitude. Subsidiary $x$ axis represents the probability of a magnitude being an annual extreme, and its return neriod in yoars

SOUTH AMERICA 2


Fig 4-1b Asymptotic distribution curves of extreme values of magnitude for Region 1, for the period 1897-1975. (Explanation of symbols as in Fig 4-1a).

## NORTH AMERICA 1



Fig 4-2a Asymptotic distribution curves of extreme values of magnitude for Region 2, for the period 1897-1964. (Explanation of symbols as in Fig 4-1a).


Fig 4-2b Asymptotic distribution curves of extreme values of magnitude for Region 2, for the period 1897-1975. (Explanation of symbols as in Fig 4-1a).

## ALEUTIANS, ALASKA 1



Fig 4-3a Asymptotic distribution curves of extreme values of magnitude for Region 3, for the period 1897-1964. (Explanation of symbols as in Fig 4-1a).


Fig 4-3b Asymptotic distribution curves of extreme values of magnitude for Region 3, for the period 1897-1975. (Explanation of symbols as in Fig 4-1a).


Fig 4-4a Asymptotic distribution curves of extreme values of magnitude for Region 4, for the period 1897-1964. (Explanation of symbols as in Fig 4-1a).

JAPAN, KURILE, KAMCHATKA
2


Fig 4-4b Asymptotic distribution curves of extreme values of magnitude for Region 4, for the period 1897-1975. (Explanation of symbols as in Fig 4-1a).


Fig 4-5a Asymptotic distribution curves of extreme values of magnitude for Region 5, for the period 1897-1964. (Explanation of symbols as in Fig 4-1a).


Fig 4-5b Asymptotic distribution curves of extreme values of magnitude for Region 5, for the period 1897-1975. (Explanation of symbols as in Fig 4-1a).


Fig 4-6a Asymptotic distribution curves of extreme values of magnitude for Region 6, for the period 1897-1964. (Explanation of symbols as in Fig 4-la).


Fig 4-6b Asymptotic distribution curves of extreme values of magnitude for Region 6, for the period 1897-1975. (Explanation of symbols as in Fig 4-1a).


Fig 4-7a Asymptotic distribution curves of extreme values of magnitude for Region 7, for the period 1897-1964. (Explanation of symbols as in Fig 4-1a).


Fig 4-7b Asymptotic distribution curves of extreme values of magnitude for Region 7, for the period 1897-1975. (Explanation of symbols as in Fig 4-1a).

On the other hand, although the absolute values of $\omega$ for almost all the regions are higher in both tables than that for the world as a whole, when their uncertainties are considered, all regions have a common range for $\omega$ corresponding to the one for the world (9.0-9.5).

These regions are the most seismically active in the world, with observed earthquakes with magnitude as high as 8.9 (Regions 1 and 4), and largest earthquakes with magnitude at least 8.6 (Regions 2 and 6). This range of $9-9.5$ as an upper bound to future events seems to be realistic.

Furthermore, the regional asymptotic distribution curves are clearly upper bounded by the asymptotic curve for the world. This is well shown in figures $4-8 \mathrm{a}$ and $4-8 \mathrm{~b}$ where all the regional curves are plotted along with the curve for the world.

They are lower bounded by the curve of Region 7 (New Zealand, Tonga, Kermadec), which is the least active region of the seven subregions, having the smallest number of shallow earthquakes.

The most significant feature which tables $4-1$ and $4-2$ reveal may be that in all regions the reduced chi-square of the first type asymptote fit is greater than that of the third asymptotic distribution. The minimum difference of $\rho_{1}-\rho_{2}$, is 0.05 for Region 6 which has also the minimum value for $\lambda$ (minimum curvature). The significance of such a small difference in the two distributions fitted can be seen in figure $4-6$ b for Region 6 , and in figure $4-5 \mathrm{a}$ for Region $5\left(\rho_{1} \rho_{2}=0.07\right)$. The maximum difference between $\rho_{1}$ and $\rho_{3}$ is 0.348 for Region 4 which shows a well-formed third asymptotic distribution curve.

From the above features it can be concluded that the third type asymptotic distribution is preferable as a general model for the statistical behaviour of the occurrence of maximum magnitude earthquakes. With the help of the parameter uncertainties now computed, it is possible to establish a more realistic forecasting procedure based on more information about the


Fig 4-8a Asymptotic distribution curves of extreme values of magnitude for the circum-Pacific belt and the world as a whole, for the period 1897-1964. (Explanation of symbols as in Fig 4-1a).


Fig 4-8b Asymptotic distribution curves of extreme values of magnitude for the circum-Pacific belt and the world as a whole, for the period 1897-1975. (Explanation of symbols as in Fig 4-1a).
data sample.
b) Statistical stability of the system

Because one of the basic assumptions for this model is that the future seismic behaviour of a region will be similar to that in the past, the statistical stability of the system must be examined. Only if stability exists, is it possible to make valuable applications of the procedures derived using the calculated parameters.

As a test parameter the most expected maximum magnitude, mode, with return period 75 years, $M_{75}$ is chosen. This approximates to the time span of the largest sample period. The results for sample periods of 35,45 , 55 and 75 years along with the range in which the $M_{75}$ will lie with $95 \%$ probability level are tabulated in Table 4-3. These are calculated using equations (4-16) and (4-17).

As expected for these regions characterized by continuous high seismic activity, it is clear from Table 4-3 that statistical stability is effectively achieved over the whole range of the intervals chosen. This stability tends to increase as the sample period lengthens and the mode, $M_{75}$, becomes stable. c) Regional features

In view of the statistical stability of the data sample, the parameters derived from the longest available period (ie 1897-1975) may now be used to predict future occurrence of maximum magnitude earthquakes. Because the prediction should be made for a period comparable with the longest available period, Table $4-4$ contains the mode which is expected to be exceeded at least once during the next $1,10,20,50$ and 100 years for each of the seven subregions and for the world as a whole. The range of this magnitude at the 95\% probability level is included.

In addition to these quantities, it is possible to determine uncertainties in them using the variance-covariance matrices and equation (4-16).

In fact, negative covariance of $\omega$ and $\lambda$ exists and so predicted

Table 4-3

Test of Statistical Stability $\tilde{M}(75)$

| Sample <br> Reriod <br> Region (years) | 35 | 45 | 55 | 65 | 75 |
| :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | $9.2 \pm 1.3^{*}$ | $9.1 \pm 1.4$ | $9.0 \pm 1.1$ | $8.9 \pm 1.0$ | $8.9 \pm 1.0$ |
| 2 | $8.9 \pm 1.0$ | $8.8 \pm 1.1$ | $8.7 \pm 0.8$ | $8.7 \pm 0.7$ | $8.7 \pm 0.8$ |
| 3 | $8.7 \pm 0.9$ | $8.9 \pm 1.2$ | $8.8 \pm 0.9$ | $8.8 \pm 0.9$ | $8.8 \pm 1.0$ |
| 4 | $9.0 \pm 0.8$ | $8.9 \pm 0.8$ | $8.9 \pm 0.9$ | $8.9 \pm 0.9$ | $8.9 \pm 0.7$ |
| 5 | $9.0 \pm 0.9$ | $9.0 \pm 0.8$ | $8.9 \pm 1.0$ | $8.9 \pm 0.9$ | $8.9 \pm 1.0$ |
| 7 | $8.8 \pm 0.9$ | $8.9 \pm 1.0$ | $8.7 \pm 1.0$ | $8.7 \pm 1.1$ | $8.7 \pm 1.2$ |
| 7 | $8.8 \pm 1.0$ | $8.8 \pm 0.9$ | $8.7 \pm 0.9$ | $8.6 \pm 0.9$ | $8.6 \pm 0.6$ |

* The uncertainties are the ranges in which the mode with return period $T=75$ years (M75) will lie with probability $95 \%$. (See equation 4-16).

Predicted most probable largest earthquake magnitude (mode)
and upper and lower bounds of the interval in which the maximum magnitude will lie
with probability $95 \%$ for return periods $1,10,20,50$ and 100 years

|  | 1 |  |  | 10 |  |  | 20 |  |  | 50 |  |  | 100 |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | L. Bound | Mode | U. Bound | L. Bound | Mode | U. Bound | L. Bound | Mode | U. Bound | L. Bound | Mode | U. Bound | L. Bound | Mode | U. Bound |
| (1) South America | $6.2 \pm .1$ | $7.2 \pm .1$ | $8.7 \pm .1$ | $7.7 \pm .1$ | $8.3 \pm .1$ | $9.2 \pm .3$ | $8.0 \pm .1$ | $8.5 \pm .1$ | $9.3 \pm .3$ | $8.3 \pm .1$ | $8.8 \pm .1$ | $9.5 \pm .4$ | $8.6 \pm .1$ | $9.0 \pm .2$ | $9.5 \pm .5$ |
|  | $6.2 \pm .1$ | $7.2 \pm .1$ | $8.6 \pm .1$ | $7.7 \pm .1$ | $8.3 \pm .1$ | $9.2 \pm .2$ | $8.0 \pm .1$ | $8.5 \pm .1$ | $9.2 \pm .3$ | $8.3 \pm .1$ | $8.7 \pm .1$ | $9.4 \pm .4$ | $8.5 \pm .1$ | $8.9 \pm .1$ | $9.4 \pm .4$ |
| (2) North America | $6.1 \pm .2$ | $7.4 \pm .1$ | $8.5 \pm .1$ | $7.7 \pm .1$ | $8.3 \pm .1$ | $8.9 \pm .3$ | $8.0 \pm .1$ | $8.5 \pm .1$ | $8.9 \pm .3$ | $8.3 \pm .1$ | $8.6 \pm .2$ | $9.0 \pm .3$ | $8.4 \pm .1$ | $8.7 \pm .2$ | $9.0 \pm .4$ |
|  | $6.0 \pm .2$ | $7.4 \pm .1$ | $8.5 \pm .1$ | $7.7 \pm .1$ | $8.3 \pm .1$ | $8.8 \pm .2$ | $8.0 \pm .1$ | $8.4 \pm .1$ | $8.8 \pm .2$ | $8.3 \pm .1$ | $8.6 \pm .1$ | $8.9 \pm .3$ | $8.4 \pm .1$ | $8.7 \pm .2$ | $8.9 \pm .3$ |
| (3) Aleutians, Alaska ${ }^{\text {a }}$ | $5.6 \pm .2$ | $7.0 \pm .1$ | $8.5 \pm .1$ | $7.5 \pm .1$ | $8.2 \pm .1$ | $9.1 \pm .2$ | $7.8 \pm .1$ | $8.4 \pm .1$ | $9.2 \pm .3$ | $8.2 \pm .1$ | $8.7 \pm .1$ | $9.3 \pm .3$ | $8.4 \pm .1$ | $8.9 \pm .2$ | $9.3 \pm .3$ |
|  | $5.7 \pm .2$ | $7.0 \pm .1$ | $8.5 \pm .1$ | $7.5 \pm .1$ | $8.2 \pm .1$ | $9.0 \pm .2$ | $7.8 \pm .1$ | $8.4 \pm .1$ | $9.1 \pm .3$ | $8.2 \pm .1$ | $8.7 \pm .1$ | $9.3 \pm .3$ | $8.4 \pm .1$ | $8.8 \pm .1$ | $9.3 \pm .3$ |
| Japan a | $6.4 \pm .1$ | $7.6 \pm .1$ | $8.7 \pm .1$ | $8.0 \pm .1$ | $8.5 \pm .1$ | $9.0 \pm .2$ | $8.2 \pm .1$ | $8.7 \pm .1$ | $9.1 \pm .2$ | $8.5 \pm .2$ | $8.8 \pm .1$ | $9.1 \pm .2$ | $8.7 \pm .1$ | $8.9 \pm .1$ | $9.2 \pm .2$ |
| (4) Kurile $\quad$ Kamehatka | $6.4 \pm .1$ | $7.6 \pm .1$ | $8.7 \pm .1$ | $7.9 \pm .1$ | $8.5 \pm .1$ | $9.0 \pm .2$ | $8.2 \pm .1$ | $8.6 \pm .1$ | $9.1 \pm .2$ | $8.5 \pm .1$ | $8.8 \pm .1$ | $9.1 \pm .2$ | $8.6 \pm .1$ | $8.9 \pm .1$ | $9.2 \pm .2$ |
| N Guínea, Banda Sea a | $6.7 \pm .1$ | $7.5 \pm .1$ | $8.7 \pm .2$ | $7.9 \pm .1$ | $8.5 \pm .1$ | $9.2 \pm .3$ | $8.2 \pm .1$ | $8.6 \pm .1$ | $9.3 \pm .4$ | $8.5 \pm .1$ | $8.8 \pm .2$ | $9.4 \pm .4$ | $8.6 \pm .1$ | $9.0 \pm .2$ | $9.5 \pm .5$ |
| (5) Celebes, Moluccas Philippines | $6.6 \pm .1$ | $7.5 \pm .1$ | $8.7 \pm .1$ | $7.9 \pm .1$ | $8.4 \pm .1$ | $9.1 \pm .2$ | $8.1 \pm .1$ | $8.6 \pm .1$ | $9.1 \pm .3$ | $8.4 \pm .1$ | $8.8 \pm .1$ | $9.2 \pm .3$ | $8.6 \pm .1$ | $8.9 \pm .1$ | $9.3 \pm .4$ |
| N. Hebrides a | $6.5 \pm .1$ | $7.4 \pm .1$ | $8.5 \pm .2$ | $7.7 \pm .1$ | $8.2 \pm .1$ | $8.9 \pm .4$ | $7.9 \pm .1$ | $8.4 \pm .1$ | $8.9 \pm .4$ | $8.2 \pm .1$ | $8.6 \pm .2$ | $9.0 \pm .5$ | $8.4 \pm .1$ | $8.7 \pm .3$ | $9.1 \pm .6$ |
| (6) Solomon N. Guinea | $6.6 \pm .1$ | $7.4 \pm .1$ | $8.4 \pm .2$ | $7.7 \pm .1$ | $8.2 \pm .1$ | $8.9 \pm .3$ | $7.9 \pm .1$ | $8.4 \pm .1$ | $8.9 \pm .4$ | $8.2 \pm .1$ | $8.5 \pm .2$ | $9.0 \pm .5$ | $8.4 \pm .1$ | $8.7 \pm .2$ | $9.1 \pm .5$ |
| N. Zealand a | $5.7 \pm .2$ | $7.2 \pm .1$ | $8.4 \pm .1$ | $7.6 \pm .1$ | $8.2 \pm .1$ | $8.7 \pm .2$ | $7.8 \pm .1$ | $8.4 \pm .1$ | $8.8 \pm .2$ | $8.2 \pm .1$ | $8.5 \pm .1$ | $8.8 \pm .3$ | $8.3 \pm .1$ | $8.6 \pm .2$ | $8.8 \pm .3$ |
| (7) Tonga <br> Kermadec | $5.7 \pm .2$ | $7.2 \pm .1$ | $8.4 \pm .1$ | $7.5 \pm .1$ | $8.2 \pm .1$ | $8.7 \pm .2$ | $7.8 \pm .1$ | $8.3 \pm .1$ | $8.7 \pm .2$ | $8.1 \pm .1$ | $8.5 \pm .1$ | $8.8 \pm .2$ | $8.3 \pm .1$ | $8.6 \pm .1$ | $8.8 \pm .2$ |
| World | $7.5 \pm .1$ | $8.3 \pm .1$ | $8.9 \pm .1$ | $8.5 \pm .1$ | $8.8 \pm .1$ | $9.1 \pm .1$ | $8.6 \pm .1$ | $8.9 \pm .1$ | $9.1 \pm .2$ | $8.8 \pm .1$ | $9.0 \pm .1$ | $9.1 \pm .2$ | $8.9 \pm .1$ | $9.0 \pm .1$ | $9.2 \pm .2$ |
|  | $7.4 \pm .1$ | $8.3 \pm .1$ | $8.9 \pm .1$ | $8.5 \pm .1$ | $8.8 \pm .1$ | $9.0 \pm .1$ | $8.6 \pm .1$ | $8.9 \pm .1$ | $9.2 \pm .1$ | $8.7 \pm .1$ | $9.0 \pm .1$ | $9.1 \pm .1$ | $8.8 \pm .1$ | $9.0 \pm .1$ | $9.1 \pm .1$ |

a: Using parameters estimated from sample period: 1897-1964
b: Using parameters estimated from sample period: $1897-1975$ sep.
magnitudes are generally well determined with small uncertainty even when there are relatively large uncertainties in the parameters $\omega$ and $\lambda$. The observed maximum magnitude earthquake within each region is shown in the last column of Table $4-2$ to illustrate this point.

From Table 4-4 several features concerning the future seismicity of the circum-Pacific belt are apparent.
i) The annual mode for all the regions is greater than $m=7.0$ with the maximum for Region 4 (Japan, Kurile, Kamchatka) as high as $m=7.6$.
ii) During the next 10 years a maximum magnitude earthquake which will exceed $m=8.2$ is expected in almost every region in the circumPacific belt. This may be as high as $m=8.5$ for Region 4 . Likewise, for the next 20 years a maximum magnitude earthquake is expected which may exceed 8.3 (Regions 6 and 7), 8.4 (Regions 2 and 3), 8.5 (Regions 1 and 5) and 8.6 (Region 4).
iii) The regions in which events with predicted maximum magnitude expected to exceed $8.8-8.9$ during the next 100 vears are: Region 1 , Region 4 and Region 5. These regions are situated in the north-western (Region 4 and Region 5) and south-eastern (Region 1) part of the circum-Pacific be1t, diagonally opposite each other. This is comparable with the results of paragraph 3.2.4 using the strain energy release method.
iv) From figures $4-1$ a to $4-7 \mathrm{~b}$ it can be seen that the return period, and the number of exceedences over the next 100 years, for an earthquake with magnitude 8.0 or greater in each of the seven subregions is as follows:

| Region 1: 7 years, | $14-15$ occurrences |
| :--- | ---: |
| Region 2: 6.5 years, | $15-16$ occurrences |
| Region 3: 10 years, | 10 occurrences |


| Region 4: 4 years, | 25 occurrences |
| :--- | :--- | ---: |
| Region 5: 4.5 years, | $22-23$ occurrences |
| Region 6: 8 years, | $12-13$ occurrences |
| Region 7: 10 years, | 10 occurrences |

v) Kanamori (1978) points out that the surface-wave magnitude scale saturates for great earthquakes with faults of length greater than 60 km . In the circum-Pacific belt he found four "giant" earthquakes which, in his newly introduced magnitude scale $M_{W}$ have magnitudes exceeding $M_{W}=9.0$. These are:
the 1960 Chilean earthquake. (22.5.1960) with magnitude $M_{W}=9.6$ ( $\mathrm{Ms}=8.3$ )- Region 1
the 1964 Alaskan " $\quad$ " 28.3 .1064 ) $\quad M_{W}=9.2$ (Ms=8.4)-Region 3
the 1957 Aleutian " 09.3 .1957 ) " $\quad M_{W}=9.1$ (Ms=8.25)-Region 3
the 1952 Kamchatka $" \quad(04.11 .1952) \quad M_{W}=9.0$ (Ms=8.4) - Region 4
Comparison of $M_{w}$ with surface-wave magnitude Ms, for earthauakes with
a smaller fault dimension shows that $M_{W}$ agrees reasonably well with Ms.
To examine the effect which the exclusion of such a high magnitude may have on the results of the statistical analysis, the annual extremes, including the new magnitude 9.6 for the Chile earthquake of 1960, are taken and re-analysed for Region 1. It is found that:

$$
\omega=11.05 \pm 0.34, u=7.07 \pm 0.03 \text { and } \lambda=0.158 \pm 0.018
$$

and the values of the mode in $1,20,50$, and 100 years using these new parameters are:

$$
\tilde{\mathrm{m}}_{1}=7.2 \pm 0.1, \tilde{\mathrm{~m}}_{10}=8.4 \pm 0.1, \tilde{\mathrm{~m}}_{50}=8.9 \pm 0.2, \text { and } \tilde{\mathrm{m}}_{100}=9.2 \pm 0.2
$$

Comparing the above results with those of Table $4-4$ it is
apparent that the new predicted magnitudes do not differ significantly from the magnitudes which were derived without using $M_{w}$ for the 1960
annual maximum. Considering that Region 1 includes the most dramatic single adjustment to magnitude, that is, from $M_{S}=8.3$ to $M_{w}=9.6$, it seems that these few cases of saturation do not produce any significant bias in the prediction procedure using the third type asymptotic distribution.
4.6 Comparing the results from energy release and third type asymptotic distribution methods

The validity of the established relations, equations ( $4-21$ ) and (4-41), among the parameters of the frequency-magnitude and energy-magnitude laws, and those of the third type asymptote can be verified by comparing the results for the same quantity derived using the two different procedures.

For each of the seven subregions both the mode and the magnitude which corresponds to the mean annual energy release, are calculated using equations 4-20 and 4-40 respectively. These are tabulated in Table 4-5 along with the upper limit $\omega$. In the same table the values of $M_{1}, M_{2}$ and $M_{3}$, calculated in Chapter III (Table 3-2), are listed for each subregion. Then a comparison can be made between $M_{1}$ and $\tilde{x}, M_{2}$ and $X_{2}$, whereas $M_{3}$ must be within the range of $\omega$.

Table 4-5 reveals some of the most significant features of this study. From the remarkably similar results for $M_{1}$ and $\tilde{x}$, as well as for $M_{2}$ and $X_{2}$ we can draw several conclusions.
i) The relations obtained here among the parameters of the two different procedures used to describe the same phenomenon, are valid over a wide range of seismo-tectonic environments.
ii) These relations give a physical link to all the parameters of the third type asymptotic distribution. Equation (4-41) links the parameters with the mean annual energy release in the region under investigation, and equation (4-21) with the $a$ and $b$ of the magnitidefrequency formula.

Table 4-5

Comparison between the parameters derived from the strain energy release and third type asymptote methods

| Region <br> code | $\mathrm{M}_{1}$ | $\tilde{x}$ | $\mathrm{M}_{2}$ | $\mathrm{X}_{2}$ | $\mathrm{M}_{3}$ | $\omega \pm \delta \omega$ | $\rho_{3}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 7.25 | 7.21 | 8.03 | 8.13 | 9.10 | $10.16 \pm 1.2$ | 0.07 |
| 2 | 7.34 | 7.37 | 7.94 | 8.07 | 9.00 | $9.14 \pm 0.5$ | 0.03 |
| 3 | 7.07 | 7.00 | 7.89 | 8.02 | 8.78 | $9.66 \pm 0.6$ | 0.09 |
| 4 | 7.52 | 7.61 | 8.20 | 8.28 | 8.96 | $9.30 \pm 0.4$ | 0.07 |
| 5 | 7.48 | 7.53 | 8.12 | 8.23 | 9.04 | $10.00 \pm 1.1$ | 0.02 |
| 6 | 7.35 | 7.35 | 7.90 | 7.98 | 8.83 | $9.44 \pm 1.1$ | 0.08 |
| 7 | 7.08 | 7.19 | 7.86 | 7.91 | 8.97 | $8.95 \pm 0.4$ | 0.05 |
| World | 8.09 | 8.30 | 8.68 | 8.62 | 9.52 | $9.23 \pm 0.25$ | 0.03 |

iii) In all regions except the third, the values of $M_{3}$ are within the range of $\omega \pm \delta \omega$. However, the case of Region 3 has its own significance. From Table $4-5$ it can be seen that Region 3 has the maximum value for the chi-square fit of all seven regions, that is, relatively the worst fit. But this region has also the worst fit when Duda (1965) plots the regional recurrence curves using the frequency-magnitude formula. He points out that "this may be caused by the superposition of two natural populations of earthquakes". The third type asymptotic distribution for this region might be similarly influenced by such a seismic feature producing this relatively poor fit.

Comparing the regional seismicity resulting from both the methods of paragraphs 3.2 .4 and $4.5 \cdot 3$ shows that the corresponding seismic pictures are almost identical.

Finally, the remarkably similar results which are obtained from these two methods in a region with the most complete data set, like the circum-Pacific belt, suggests that the third type asymptotic distribution method will be of great assistance for forecasting procedures in less active regions where only the extreme earthquake magnitudes may be known.

### 4.7 Conclusions

The objectives of this chapter were:
i) to investigate the usefulness of the third type asymptotic distribution of extremes for predicting earthquake risk, and
ii) to relate the parameters of that asymptote with those of the frequency-magnitude and energy-magnitude laws, in order to explore their physical meaning and obtain a link with the physical release of strain energy.

Towards the first objective, a specific goal was to develop a computing procedure which would determine uncertainties in the parameters computed. It is assumed that an error analysis is vital to any seismic risk analysis.

Using Marquardt's algorithm which is based upon the non-linear leastsquares fitting, the covariance matrix among the three parameters has been obtained.

This computing procedure allows several important factors to be taken into account, such as an assigned weight for each individual extreme magnitude, the number of missing years may be considered and incompletely reported data may be omitted without disturbing the distribution of the observed maximum magnitudes.

The variance-covariance matrix among the parameters shows that the characteristic largest value $u$ is the most precisely known parameter. The variance for $\omega$ and $\lambda$ are usually large. This indicates that the upper limit to earthquake magnitude and the curvature of the distribution are often difficult to resolve with high precision.

The negative covariance between the upper limit and curvature implies that they are not independent parameters (Burton, 1978b). Yegulalp and Kuo (1974) also note that there is a correlation between these two parameters. Thus, it can be concluded that among the parameters of the third type asymptote only the characteristic value $u$ is independent.

It is observed that when the data shows little curvature $\lambda$, and, therefore, a high value of $\omega$, these parameters are usually accompanied by larger uncertainties than usual, which may indicate that the time span is insufficient to establish curvature. However, the existence of negative covariance between them leads to seismic risk calculations with reasonable uncertainties. Uncertainties have been computed for the upper and lower
bounds within which the predicted quantities will lie at a given probability leve1.

Comparison between the first and third type asymptotic distribution shows that in all cases the third has a better fit to the data than the first. This is reasonable because the third type asymptote is closer to the real process: it takes into account the existence of an upper magnitude threshold. It can be concluded that the third type asymptote is usually preferable as a general model for the statistical behaviour of the occurrence of maximum magnitude earthquakes.

The second objective is achieved by relating the third asymptote's parameters to the parameters of the frequency-magnitude and energy-magnitude laws. In particular, the $X_{2}$ (eq 4-41) gives the mean annual rate of energy release in the region in terms of the Gumbel parameters.

A second feature is the remarkably similar results which are obtained when the two different procedures, that is whole process and part process, are used to describe the same phenomenon in the same area.

Finally, the centres of highest seismic activity in the circum-Pacific belt are diagonally opposite each other, and this presumably relates to the tectonic movement of the Pacific plate (Duda, 1965). The seismicity of this region is expected to generate a maximum magnitude which may exceed $m=8.2$ in almost every region in the circum-Pacific belt during the next 10 years. Regions 1,4 , and 5 are the regions in which an earthquake with magnitude $m=8.8$ to 8.9 is expected to be exceeded at least once during the next 100 years.

From the results obtained in this chapter it is apparent that the energy release method developed in Chapter III and the third type asymptotic distribution described here, are capable of describing the seismic feature of a region. The method of calculating the uncertainties in the parameters and in the prediction quantities shows that the third type asymptotic
distribution method can be a useful model for prediction procedure. Thus, in the next part of this study these methods will be applied to Greek seismicity in an attempt to evaluate the seismic risk of the area. This area, although it has a seismicity lower than the circum-Pacific belt, is seismically the most active area in Europe. Before this can be done, an earthquake catalogue of the area is presented which will be the main source for the ensueing risk evaluation.

CHAPTER V
AN EARTHQUAKE CATALOGUE FOR THE AREA OF GREECE $\left(N_{33}^{42.5}, \mathrm{E}_{19}^{29}\right)$ SINCE 1901
5.1 Introduction

Among the demands of modern seismology for seismic hazard or zoning considerations, is the existence of an earthquake data set as accurate, homogeneous and complete as possible. Thus, although Greece has the seismic privilege to be the most active area of Europe, and its long history can report catastrophes due to earthquakes as far back as 2100 B.C. (Sieberg, 1932; Galanopoulos, 1961), the requirements for accuracy and completeness restrict usage to mainly instrumentally recorded events. Therefore, the earliest starting point for a statistical treatment founded on instrumental measurements is the beginning of the present century, whereas the whole seismic history remains a vital background factor for the seismic behaviour of the region.

The steady improvement of number and quality of seismological stations from decade to decade, however, makes it almost impossible to prepare a catalogue which would be equally complete and accurate over the whole period of instrumental observations. With the data available, homogeneity can only be achieved by consistent treatment of all source parameters.

For Greece, such an effort is attempted in the present study by recalculating all source parameters, when there are enough data to justify this treatment, by using single and joint epicentre determination methods of calculating source parameters. Magnitudes are recalculated according to a consistent scheme for the longest possible part of the period under consideration. The results, then, are tested for completeness and compared with the ISS original locations and with macro-seismic epicentres.

### 5.2 Previous work in Greek earthquake cataloguing

A1though the need for earthquake cataloguing was recognized in Greece
as early as the middle of the nineteenth century, and a number of Greek earthquake catalogues have existed since 1879 (Schmidt, 1879, for 1840-1878; Galanopoulos, 1953, for 1879-1892; Annales de 1'Observatoire National d'Athènes from 1893 up to 1936; and monthly and annual bulletins since 1949), the first systematic attempt to accumulate and classify all earthquakes over a certain magnitude $(M \geqslant 5)$ was made by Galanopoulos in 1960. By careful examination of all available sources, he published details of a large number of earthquakes with macroseismic information which occurred in the period 1801-1958. This catalogue and a more recent version for the period 1700-1960 (Galanopoulos, 1963) which includes a long list of further references, are the basis for most of the later publications concerning Greek seismicity.

In 1969, Kárnik, in his publication for Europe (1901-1955), made the first attempt to present an earthquake catalogue with magnitude determinations. The basic sources for Greek epicentres came from the previous Galanopoulos publications as well as from Gutenberg and Richter (1954). In this catalogue there are four types of magnitudes: (a) Mor $M_{L H}$ forthquakes with depths $1-5 \mathrm{~km}$ or $5-60 \mathrm{~km}$ (sup. and n respective 1 y ); (b) $\mathrm{M}_{B}$ for depths $60-300 \mathrm{~km}$; (c) $M$ in brackets for magnitudes converted from macroseismic parameters $I_{o}$ and $r$; and (d) $M$ when this magnitude is taken from national catalogues and corrected for differences between it and his $M_{\text {LH }}$ standards.

In 1970 the UNDP/UNESCO Survey of the Seismicity of the Balkan Region began, and it gave first priority to the compilation of an earthquake catalogue. This catalogue (Shebalin, Karnik and Hadžievski, 1974) represents a "collection of main earthquake parameters" as pointed out in its preface. It covers the period 1901-1970 and contains earthquakes with $M \geqslant 4$ or $I_{o} \geqslant V I$. Basic sources for Greece were lists of earthquake parameters from Galanopoulos (1960, 1963), Kárnik (1969, 1971), and Papazachos and Comninakis (1971),
coupled with International Summaries and Catalogues.
The UNESCO's catalogue has been widely used for many investigations for the Balkan region because it is the most complete in terms of the number of earthquake entries. It is useful to briefly describe the procedure used for earthquake parameter determinations by quoting from its introduction and explanatory text.
i) Origin time "The estimation of the origin time was based mainly on results reported by ISS, BCIS, ISC, etc., and for the beginning of the century from reported P -arrivals and macroseismic observations. The final estimation was, usually, taken as average for all the accessible determinations" (page 51).
ii) Epicentre coordinates
"As a base for determinations of normal epicentres ( $\mathrm{h}<60 \mathrm{~km}$ ) in the catalogue the macroseismic data were chosen. For the intermediate shocks $(60 \leqslant h \leqslant 300)$ the position of epicentres were taken from instrumental determinations. For the aftershocks, the coordinates of the main shock were repeated" (pages 52, 53).
iii) Focal depth
"For the whole investigated period (1901-1970), the only possibility for the determination of focal depth of 'normal' shocks was the use of macroseismic data. All instrumental determinations of focal depth of 'normal' shocks published by ISS, BCIS, etc., for the period before 1965 were neglected. Since 1965 the instrumental determinations of $h$ made by ISC and CGS (NEIS), became the source of information. The determination of focal depth for intermediate sources was based mainly on the instrumental data" (pages 56, 57).

```
"The \(M_{\text {LH }}\) (Karnik et al., 1962) was taken as the basic one for
'normal' shocks. Generally the instrumental determinations
of \(M_{\text {LH }}\) made by Karnik (1969) were used for the period 1901-1963.
For \(M_{B}=4-5\) the published \(M_{B}\) from ISC and US were taken,
because for this interval \(M_{B} \approx M_{L H}\) can be assumed. In the
total absence of instrumental data, the magnitude corresponding
to \(M_{L H}\) was determined from macroseismic data. For intermediate
shocks the magnitudes \(M_{B}\) determined by Karnik (1964) were used
or, exceptionally, were taken from international summaries"
(pages 58, 59).
```

This catalogue will be referred to as UNS from now on.
The first attempt to relocate epicentres reported by the International Seismological Summary (ISS), using the single-event method and the computer program SPEEDY of Douglas, Young and Lilwall (1974) for the Marmara region in north-west Turkey, which is included in our area of investigation, was made by Crampin and Uçer (1975). They relocated all sixty earthquakes which took place within this area from 1913-1963. Without applying any station adjustinent the relocated epicentres were significantly different from the ISS locations, with an average shift of 40 km for events before 1957, and 5 km for events from 1956-1963 when the ISS epicentres were calculated by computer.

Alsan, Tezuçan and Båth (1975) published the first computerized earthquake catalogue for the whole of Turkey. This catalogue, in order to provide better continuity with corresponding catalogues, includes events from beyond the Turkish borders. Because the same criterion is also applied in the present study, there is an overlapping zone of $33.5^{\circ} \mathrm{N}-$ $42.0^{\circ} \mathrm{N}, 25.5^{\circ} \mathrm{E}-29.0^{\circ} \mathrm{E}$. For the relocation of epicentre parameters, a
computer program based on the single-event method was used. This program was developed in 1972 at the BCIS (Rothé et al., 1972).

Magnitudes were redetermined using Swedish station amplitudes. Üçer et al. (1975) point out that the shift from ISS locations for the first decade after 1918 gives an average change in position of 137 km . This average shift is lessened for the later events being 42 km for the decade 1948-1958, whereas for the period $1959-1963$ it is less than 10 km . This catalogue will be referred to as ATB (ie the initials of the authors).

Finally, Galanopoulos (1977) includes the most recent catalogue of Greek earthquakes. This catalogue contains a list of earthquakes since 1902 with $\mathrm{M} \geqslant 5 \frac{1}{2}$. It is a compilation from his previous work ( 1960,1963 ), Kárnik (1969), Gutenberg and Richter (1954), UNS and ATB, and since 1961 all earthquake parameters, except magnitude, are those calculated by ISS and the International Seismological Centre (ISC). All magnitudes are averaged surface wave magnitudes, determined using macro and microseismic information.

It is apparent that none of the existing catalogues for Greece as a whole fulfil the important objective of homogeneity either for locations or for magnitudes. All of them have inconsistencies because they are compiled from many different sources. Furthermore, the ISS locations are influenced by the period in which the events were recorded and by changes in travel-time models used (see below), as is apparent from the work of Crampin and Üçer (1975) for the Marmara area and Ucer et al. (1975) for the results of ATB.
5.3 The ISS epicentres: reasons for inaccuracies

The only Bulletins which tabulate phase arrival and epicentres from 1917 until 1963 are those of the International Seismological Summary. ISS used different travel-time tables during the whole period (TurnerZoppritz from 1917 until 1929, Jeffreys-Bullen revised tables from

1930-1936, and since then Jeffreys-Bullen with ellipticity corrections).
Epicentres were determined by hand-operated mechanical calculators, except for the last seven years when an electronic computer was used, and although the full scale least-squares procedure used was similar to the computer procedure used today, it frequently adopted old locations to fresh sets of arrivals to reduce the prodigious amount of work (Crampin and Uçer, 1975).

The lack of adequate travel-time tables for the whole period and the technique of adopting old locations, leads to inaccuracies in ISS epicentre determinations. Even for recent events (ie since 1954) the use of the world-wide travel-time tables, without station adjustments for the particular region, (acceptable when making international routine calculations), makes the ISS locations less precise than is possible. These inaccuracies can be allowed for by relocating using consistent travel times, coupled with station adjustments to account for station, travel-time, network and source effects. An attempt to eliminate these inaccuracies is made in this study.

### 5.4 Data sources

5.4.1 Relocation

Because both SPEEDY and JED programs (Douglas, Young and Lilwall, 1974) which are used for the relocation procedure, are designed for first arrival readings (ie P-waves), the following data sources are used:
i) For the period 1913-1917, the monthly bulletins of the British Association for the Advancement of Science.
ii) For 1918-1963, the bulletins of the ISS.
iii) For 1964-1976, the bulletins of the ISC.

Arrival data were collected for those earthquakes with epicentres not only within Greek territory but slightly beyond.

The area of investigation is limited to latitudes $33^{\circ} \mathrm{N}$ to $42.5^{\circ} \mathrm{N}$ and longitudes $19^{\circ} \mathrm{E}$ to $29^{\circ} \mathrm{E}$ north of the $38^{\circ} \mathrm{N}$ parallel and $30^{\circ} \mathrm{E}$ south of it, in order to cover the Dodecanese Islands.

An attempt to recalculate events prior to 1917 using source (i) was not successful because of the poor quality and quantity of the readings reported as first arrivals. The stations were too widely spread and too limited in number. For the period 1964-1975, where data from source (iii) is available, several test recomputations showed that the shift between the new and old location was on an average less than 10 km to within the 95\% confidence limits. Furthermore ISC gives standard deviations in origin time, coordinates and focal depth determinations, which indicate the quality of the solution given. Therefore no recomputations for this period are made.

Earthquakes for the period 1917-1963 inclusive are relocated here, using first arrival data from source (ii) exclusively. For the 605 earthquakes which are relocated, 45,000 first arrivals (cards) were punched and used as input to both programs.
5.4.2 Magnitude

In order to determine the surface-wave magnitude, $M$ (see 5.7) the following data sources are used:
i) For the period 1908-1959, the annual bulletins of the Seismological Institute at Uppsala (SIU) for readings of Uppsala station (UPP).
ii) For 1951-1955, the annual bulletins of SIU for readings of Kiruna station (KIR).
iii) For 1956-1963, the monthly bulletins of SIU for all the Swedish network.
iv) For 1901-1970, the UNS catalogue.
v) For 1964-1976, the ISC magnitude determinations.
vi) For 1976-1978, the United States National Earthquake Information Centre (NEIC) magnitude determinations.

Earthquakes are today generally located by one of two methods. The first method is called the "single-event" or Geiger's method (Geiger, 1910); the second is the Joint Epicentre Determination (JED) method (Douglas, 1967).
5.5.1 Methods and computer programs chosen for relocation

The computer programs SPEEDY and JED written by Douglas, Young and Lilwall (1974) are used to relocate the source parameters of Greek earthquakes. Both programs are designed to accept first-arrival readings only, and a set of travel-time tables. In this study the ISS first arrivals and Herrin's "68" travel-time tables are the main input data to both programs. SPEEDY is a "single-event" based program, whereas JED is a "group-event" based program. A detailed description of the two methods can be found in Lilwall (1969).

Although JED is a more accurate method of epicentre determination than "single-event" methods, it is a very costly process in terms of computing time. So JED is used here as the first step of the relocation procedure to determine a set of station adjustments which are then retained. Thus, having a set of station adjustments derived from a group of major and well-recorded events using JED, the "single-event" based program SPEEDY can then be used, and these adjustments applied to the travel-time of other events.
5.5.2 Procedure used

Applying station adjustments to travel-times in a "single-event" location, assumes that the corrections for a particular station are constant for all epicentres. This is not necessarily the case for the large area of Greece, which has a very complicated tectonic structure. The whole area was initially divided into three regions, but not divided in time:

Region A: Western Greece (west of $22^{\circ} \mathrm{E}$ ), with mainly shallow earthquakes.

Region B: Southern Greece (south of $38^{\circ} \mathrm{N}$ ), with the majority of intermediate earthquakes.

Region C: Central and Northern Greece (the remaining part of Greece).

For Region A, 19 well-recorded events (minimum number of stations 65, average number 121) were relocated using SPEEDY. This facilitated the selection of.a master event, which was restrained in a subsequent JED relocation of the same 19 events. A set of station adjustments, corresponding to the earthquakes of region A were obtained for 260 stations which had recorded more than 5 events. These adjustments were subsequently used in individual SPEEDY relocations of the 19 events, including the previous master event. This resulted in a new slightly changed position for the master event (in all cases the shift was less than 5 km ). The master event was then restrained at its new position and a final set of adjustments for the region determined by JED. This procedure was repeated for the other two regions.

Before applying station adjustments in all subsequent relocations, the effectiveness of these "terms" in the precision of the final solution was checked. This check was made using the latest and presumably more accurate set of data from 1957 until 1963.

As a measure of calculated precision the $95 \%$ confidence area around the epicentre has been chosen (Flinn, 1965). The 202 events of this period were relocated using SPEEDY both with and without station adjustments. Table 5-1 tabulates the results and shows that 174 of the total 202 events $(86.1 \%)$, relocated using station adjustments, have smaller $95 \%$ confidence areas than those to which adjustments were not applied. The $17.8 \%$ overall average improvement is surprisingly high considering the quality
of the data set used. It is expected that as the quality of first arrivals increases, and the number of stations omitted because of bad readings decreases, the application of station adjustments should improve the precision and accuracy of epicentre parameters, although the difference in confidence areas should decrease. In fact, the 28 cases of larger confidence areas are almost all associated with poor quality first arrivals. The application of station adjustments results in some stations which have been truncated in the first solution being included in later solutions, if these stations correspond to poor quality first arrivals the confidence area may increase; but because of the larger number of stations, the solution will probably be less subject to station network bias.

Table 5-1
Contribution of station adjustments into final earthquake location

| Region | No of events | No of <br> improved <br> cases <br> $\left(A_{W}<A_{0}\right) *$ | Percentage <br> of improved <br> $\%$ | No of <br> worst <br> cases <br> $\left(A_{w} \geqslant A_{o}\right)$ | Percentage <br> of worst <br> $\%$ | Average <br> improved <br> $\left(1-\frac{A_{w}}{A_{0}}\right)$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| A | 71 | 62 | 87.3 | 9 | 18.3 | 22.1 |
| B | 48 | 37 | 77.1 | 11 | 22.9 | 11.2 |
| C | 83 | 75 | 90.4 | 8 | 9.6 | 20.1 |
| TOTAL | 202 | 174 | 86.1 | 28 | 13.9 | 17.8 |

* $A_{W}: 95 \%$ confidence area with station adjustments
$A_{0}: 95 \%$ confidence area without station adjustments

It was also found that the differences in type, sensitivity and accuracy of the instrument used during the whole period combined with the uneven distribution of the stations, to minimize the effect of the station
adjustment to a more accurate location. So, the final procedure chosen divides the data both geographically and in time as follows. First, the whole region was divided into two subregions: $A+C=R_{1}$ and $B=R_{2}$. This is done not only in order to save computing time, but also because $A$ and $\mathbb{C}$ subregions show similarities in earthquake depths. The vast majority of earthquakes in subregions $A$ and $C$ have their origin in the upper crust, whereas region $B$ includes the vast majority of intermediate earthquakes (ie $\mathrm{h}>60 \mathrm{~km}$ ). Secondly, the time-span of each group is chosen short enough to avoid large differences in station distribution and instrumental characteristics, but long enough to include at least one or two large and well-recorded events with as much macroseismic information as possible, the latter being the main criterion because of the need for a restrained master event. After 1953, when the quality of stations and first arrivals remarkably improved and macroseismic and microseismic locations, in most cases, became close to each other, the limit in number of events within a group and first arrivals, put by JED and computer storage capability, governs the choice.

The periods chosen are: $1917-1925,1926-1930,1931-1935,1936-1940$, $1941-1945,1946-1949,1950-1952,1953,1954-1956,1957-1958,1959,1960-1969$, and 1952-1963. Consequently, 26 ( $2 \times 13$ ) runs of JED were necessary, and 605 earthquakes have been relocated for the period 1917-1963.

For each of the 26 groups a master event has to be chosen and restrained. The criterion for choosing a master event for a group is that this event must have the best solution, that is the smallest $95 \%$ confidence area derived from SPEEDY with station adjustments applied, coupled with supporting macroseismic information. However, especially for shallow earthquakes when discrepancies exist, more weight is given to the macroseismic hypocentre than the SPEEDY epicentre. Table 5-2 lists the epicentres of the 26 master events chosen, along with their uncertainties; Intensities and macroseismic epicentres taken from UNS are also included.

The 26 Master events used for the relocation procedure

| Date | $\begin{array}{\|c} \mathrm{La}_{\mathrm{o}}^{\mathrm{o}} \\ \hline \end{array}$ | $\stackrel{\mathrm{Lon}}{\mathrm{O}}^{\mathrm{E}}$ | $\begin{gathered} \text { Depth } \\ \mathrm{Km} \\ \hline \end{gathered}$ | No Obs. | Period | Region | Intensity | Macroseismic Epicentre |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 13.11.24 | $39.2 \pm 0.2$ | $20.9 \pm 0.1$ | $85 \pm 20$ | 29 | 1917-1925 | R1 | 7 | $\left(\begin{array}{ll}39.3 & 20.7\end{array}\right)$ |
| 06.07.25 | $37.8 \pm 0.1$ | $21.9 \pm 0.1$ | $70 \pm 20$ | 53 | 1917-1925 | R2 | 7 | $\left(\begin{array}{ll}37.8 & 22.1\end{array}\right)$ |
| 30.08 .26 | $36.8 \pm 0.1$ | $23.2 \pm 0.1$ | $26 \pm 12$ | 75 | 1926-1929 | R2 | 8 | $\left(\begin{array}{ll}36 \frac{3}{4} & 23 \frac{1}{4}\end{array}\right)$ |
| 18.04.28 | $42.3 \pm 0.1$ | $25.3 \pm 0.1$ | $7 \pm 4$ | 95 | 1926-1929 | R1 | 10 | $\left(\begin{array}{ll}42.2 & 25.1\end{array}\right)$ |
| 26.09.32 | $40.4 \pm 0.1$ | $23.8 \pm 0.1$ | $5 \pm 4$ | 134 | 1930-1934 | R1 | 10 | $\left(\begin{array}{ll}40.5 & 23 \frac{3}{4}\end{array}\right)$ |
| 09.11 .34 | $36.5 \pm 0.1$ | $25.4 \pm 0.1$ | $132 \pm 14$ | 64 | 1930-1934 | R2 | - | - |
| 04.01.35 | $40.8 \pm 0.1$ | $27.5 \pm 0.1$ | $13 \pm 10$ | 84 | 1935-1940 | R2 | 9 | $\left(\begin{array}{ll}40.5 & 27.5\end{array}\right)$ |
| 20.07 .38 | $38.3 \pm 0.1$ | $23.7 \pm 0.1$ | $42 \pm 15$ | 81 | 1935-1940 | R1 | 8 | $\left(\begin{array}{ll}38.5 & 23.8\end{array}\right)$ |
| 27.08 .42 | $41.6 \pm 0.1$ | $20.5 \pm 0.1$ | $12 \pm 7$ | 50 | 1941-1946 | R1 | 8 | $\left(\begin{array}{ll}41.7 & 20.5\end{array}\right)$ |
| 02.09 .45 | $34.4 \pm 0.1$ | $28.6 \pm 0.1$ | $62 \pm 9$ | 68 | 1941-1946 | R2 | - | - |
| 30.06 .48 | $39.0 \pm 0.1$ | $20.5 \pm 0.1$ | $36 \pm 11$ | 96 | 1947-1949 | R1 | 9 | $\left(\begin{array}{ll}38.8 & 20.5\end{array}\right)$ |
| 23.07 .49 | $38.7 \pm 0.1$ | $26.3 \pm 0.1$ | $17 \pm 7$ | 124 | 1947-1949 | R2 | 9 | $\left(\begin{array}{ll}38.5 & 26.3\end{array}\right)$ |
| 05.04.51 | $37.5 \pm 0.1$ | $20.3 \pm 0.1$ | $41 \pm 14$ | 78 | 1950-1952 | R1 | - | - |
| 17.12 .52 | $34.5 \pm 0.1$ | $24.2 \pm 0.1$ | $17 \pm 9$ | 232 | 1950-1952 | R2 | - | - |
| 07.02 .53 | $34.8 \pm 0.1$ | $24.1 \pm 0.1$ | $33 \pm 9$ | 138 | 1953 | R2 | - | - |
| 11.08 .53 | $38.4 \pm 0.1$ | $20.7 \pm 0.1$ | $11 \pm 6$ | 244 | 1953 | R1 | 10 | $\left(\begin{array}{ll}38.2 & 20.7\end{array}\right)$ |
| 03.01 .55 | $39.2 \pm 0.1$ | $22.3 \pm 0.1$ | $41 \pm 9$ | 100 | 1954-1956 | R1 | 7 | $\left(\begin{array}{lll}39.2 & 22.1\end{array}\right)$ |
| 16.07 .55 | $37.7 \pm 0.1$ | $27.2 \pm 0.1$ | $31 \pm 9$ | 232 | 1954-1956 | R2 | 8 | $\left(\begin{array}{ll}37.5 & 27.1\end{array}\right)$ |
| 24.04 .57 | $36.4 \pm 0.1$ | $28.6 \pm 0.1$ | $69 \pm 5$ | 255 | 1957-1958 | R2 | - | - |
| 27.11 .57 | $39.4 \pm 0.1$ | $22.7 \pm 0.1$ | $42 \pm 5$ | 111 | 1957-1958 | R1 | 6 | $\left(\begin{array}{ll}39.2 & 22.6\end{array}\right)$ |
| 25.04 .59 | $37.0 \pm 0.1$ | $28.6 \pm 0.1$ | $35 \pm 7$ | 149 | 1959 | R2 | 8 | $\left(\begin{array}{lll}37.0 & 28.7\end{array}\right)$ |
| 07.10 .59 | $41.0 \pm 0.1$ | $19.8 \pm 0.1$ | $28 \pm 4$ | 161 | 1959 | R1 | 8 | $\left(\begin{array}{ll}41.0 & 19.8\end{array}\right)$ |
| 26.05 .60 | $40.6 \pm 0.1$ | $20.6 \pm 0.1$ | $20 \pm 4$ | 194 | 1960-1961 | R1 | 9 | $\left(\begin{array}{ll}40.6 & 20.7\end{array}\right)$ |
| 23.05 .61 | $36.8 \pm 0.1$ | $28.4 \pm 0.1$ | $74 \pm 5$ | 212 | 1960-1961 | R2 | 8 | $\left(\begin{array}{ll}36.5 & 28.6\end{array}\right)$ |
| 28.04 .62 | $36.2 \pm 0.1$ | $26.8 \pm 0.1$ | $56 \pm 5$ | 173 | 1962-1963 | R2 | - | - |
| 28.08 .62 | $37.8 \pm 0.1$ | $22.9 \pm 0.1$ | $95 \pm 3$ | 226 | 1962-1963 | R1 | 7 | $\left(\begin{array}{ll}37.7 & 22.6\end{array}\right)$ |

For the remaining events of the group, and in those cases where SPEEDY failed to give reliable depth determinations, depths were restrained to values derived from (M, $\left.I_{0}, h\right)$-relations of Shebalin et al (1974), using self-determined magnitudes described below.

### 5.6 Magnitude determination

For magnitude determination of Greek earthquakes, the method used by Alsan, Tezucan and Båth (ATB) for Turkish earthquakes (ATB - 1975) is adopted for the following reasons:
i) Swedish annual and monthly bulletins with ground amplitudes are available from 1908.
ii) The majority of ATB catalogue earthquakes are within the common Aegean area (ie $35.5^{\circ} \mathrm{N}-42.5^{\circ} \mathrm{N}, 25.5^{\circ} \mathrm{E}-30.0^{\circ} \mathrm{E}$ ). Thus the converting equations derived for that region are also valid for Greece (see below). This common area is tectonically the most complicated, the remaining parts being characterised by containing the vast majority of shallow earthquakes.
iii) For both Turkish and Greek earthquakes, the dominating surfacewave periods of Swedish records are around $10-15 \mathrm{sec}$. , and hence the validity of converting equations becomes even stronger.
iv) The uniform way of magnitude determinations for these seismically very active neighbouring countries in which there is no seismicity boundaries, enables further and larger scale investigations to be made.

### 5.6.1 Magnitude scale chosen

The only instrument which has been in operation during the whole period of investigation, with almost unchanged characteristics, is the Uppsala's Wiechert seismograph. There are far too few cases where the Wiechert has recorded body phases to justify calculations of the body wave
magnitude $M_{b}$ alone, so the surface-wave magnitude $M$ is chosen as the standard magnitude required for the whole period. The value of $M$ is calculated from

$$
\begin{equation*}
M=\log \frac{A}{T}+1.66 \log \Delta^{0}+3.3 \tag{5-1}
\end{equation*}
$$

where $T$ is the period in the range $10-30$ seconds, A the ground amplitude in microns, and $\Delta$ the epicentral distance in degrees (ATB, 1975). Values of $M$ generally exhibit greater stability than $M_{b}$ values when only one or few stations are available (Båth, 1969).
5.6.2 Magnitude determination procedure

A detailed description of the procedure used can be found in the ATB catalogue. The basic steps and conversion equations used for magnitude determinations of the Greek earthquakes are as follows.

The ATB procedure is applied and the magnitude is taken as the average of $M$ from Uppsala (UPP) derived from long-period Benioff instruments, and M from Kiruna (KIR), derived from Galitzin instruments. For a consistent calculation of $M$ over the whole period of investigation, the following regression equations, derived from parallel recordings on the instruments since 1955 are used:
and

$$
\begin{array}{ll}
M(U P P)=1.01 M(K I R)-0.17 & \text { for } N=221  \tag{5-2}\\
M(K I R)=0.91 M(U P P)+0.58 & \text { for } N=221
\end{array}
$$

where $N$ is, the number of pairs of observations (ie number of earthquakes). Then, the average $\bar{M}$, in terms of $M(U P P)$ or in $M(K I R)$ alone, is:
$\bar{M}=\frac{1}{2}[M(U P P)+M(K I R)]=0.95 M(U P P)+0.29=1.01 M(K I R)-0.08$

Likewise:

$$
\begin{equation*}
\bar{M}=0.85 M(W)+1.04 \quad \text { for } N=51 \tag{5-4}
\end{equation*}
$$

where $M(W)$ is the Wiechert magnitude.
When surface-wave records are unavailable, then:

$$
\begin{equation*}
\bar{M}_{b}=\frac{1}{2}\left[M_{b}(U P P)+M_{b}(K I R)\right] \tag{5-5}
\end{equation*}
$$

where $M_{b}$ is determined from short-period vertical-component $P$-wave records, using the formula

$$
\begin{equation*}
M_{b}=\log \frac{A}{T}+q(\Delta, h) \tag{5-6}
\end{equation*}
$$

where the calibrating term $\mathrm{q}(\Delta, \mathrm{h})$ is taken from Gutenberg and Richter (1956).

The regression equation of $\bar{M}$ on $\bar{M}_{b}$ is:

$$
\begin{equation*}
\bar{M}=1.46 \bar{M}_{b}-2.91 \text { for } N=63 \text { and } h \leqslant 45 \mathrm{~km} \tag{5-7}
\end{equation*}
$$

When only $M_{b}$ (UPP) or $M_{b}$ (KIR) is available, then:

$$
\left.\begin{array}{ll}
\bar{M}=1.30 M_{b}(U P P)-1.91 & \text { for } N=90  \tag{5-8}\\
\bar{M}=1.45 M_{b}(K I R)-3.04 & \text { for } N=66
\end{array}\right\}
$$

Equations (5-2) to (5-8) are taken from the ATB catalogue and applied to earthquakes from 1908 until 1968.

The ISC took over the service from ISS in 1964 and started to determine body-wave magnitudes $M_{b}$, with continuously decreasing magnitude threshold. A regression equation $\bar{M}$ and $M_{b}$ (ISC) is derived for the five overlapping years 1964-1968 so that the smaller Greek earthquakes, which are too distant from the Swedish network to ensure detection, are included without sacrificing the achieved homogeneity. This is:

$$
\begin{equation*}
\bar{M}=1.37 M_{b}(I S C)-1.74 \text { for } N=187 \tag{5-9}
\end{equation*}
$$

with a standard deviation on $\bar{M}$ of $\pm 0.27$.

It is necessary to check if any systematic bias exists in magnitude determinations from the Swedish network. A regression between $M_{b}$ (ISC) values and $M$ for the same events reported from all available agencies is derived for the period 1964 to 1975. This is:

$$
\begin{equation*}
M=1.31 M_{b}(I S C)-1.41 \text { for } N=126 \tag{5-10}
\end{equation*}
$$

with a standard deviation on M of $\pm 0.41$.
These two lines plotted in Figure $5-1$ are almost identical, so any disadvantage caused by using only a local network is diminished.

Equations (5-9) and (5-10) have their own limitations for extrapolation because the values of $M_{b}$ used are not greater than 6.1 to 6.3 , the vast majority being between $M_{b}=4.4-5.5$.

In the ATB catalogue the corresponding equation to (5-9) is:

$$
\begin{equation*}
\bar{M}=1.55(\text { ISC })-2.49 \text { for } N=110 \tag{5-11}
\end{equation*}
$$

This equation is very similar to equation (5-9), especially in the range $M_{b}=4.3-5.5$ for which the conversion equations (5-2) to (5-8) are needed. This is probably because the overlapping area of the two investigations includes the Aegean Sea, which is seismically very active, resulting in similar conversion equations for both regions.

The number of observations may be used as an additional check on calculated magnitude, especially for events with $M \leqslant 5$. The number of observations, $N$, serves as a measure of recording distance and so $M$ depends on $\log \mathrm{N}$. For the period 1964-1968, the following equation is derived:

$$
\begin{equation*}
\bar{M}=1.51 \log N^{\prime}+2.04 \text { for } N=187 \tag{5-12}
\end{equation*}
$$

with a standard deviation on the calculated $M$ of $\pm 0.34$, whereas for $M_{b}$ (ISC) the corresponding equation is:


Fig 5-1 Surface-wave magnitudes Ms of Greek earthquakes (derived from amplitudes of the Swedish network and reported from other agencies), versus body-wave magnitude Mb (reported by ISC), for the same events (eq 5-9 and 5-10 in the text respectively).

$$
\begin{equation*}
M_{b}(I S C)=0.94 \log N^{\prime}+3.06 \text { for } N=187 \tag{5-13}
\end{equation*}
$$

with a standard deviation on the $M$ of $\pm 0.24$.
Although the values of the coefficients of these equations certainly depend on the years used, equation (5-12) is used as an additional check for small earthquakes (ie $M \leqslant 5$ ) since 1964.

### 5.7 Completeness of the catalogue

The consistent method of calculation of all source parameters, including magnitude, is designed to give a high degree of homogeneity of data included in the present catalogue. But like any other catalogue, its completeness depends upon data availability, and this is far from homogeneous, the usual bias being against small shocks, particularly in the earliest years. This inhomogeneity can be assessed by finding the magnitude above which the catalogue can be considered as homogeneous and reasonably complete, or alternatively by assigning time intervals in which a certain magnitude range is likely to be completely reported.
5.7.1 Temporal plot of grouped events

Table 5-3 and Figure 5-2 illustrate incompleteness in the data available for Greece. Figure 5-2 shows the number of earthquakes per decade. (The last group covers the period 1971-1977). Events are grouped in six magnitude ranges: $M \leqslant 4.7,4.8<M \leqslant 5.2,5.3 \leqslant M \leqslant 5.7,5.8<M \leqslant 6.2,6.3 \leqslant M \leqslant 6.7$, and $\mathrm{M} \geqslant 6.8$; and they are plotted along with the total number of events per decade. The numerical data corresponding to Figure 5-2 are listed in Table 5-3, for the complete period from 1901-1977 (earthquakes for 19011907 are taken from the UNS catalogue).

The first feature to note from Figure $5-2$ is that there is no reason to question the completeness of the large earthquakes with magnitude greater than 5.8. The fluctuation in the number of magnitude 5.8 and larger earthquakes recorded per decade, shows no trend in the 77 -year sample period from


Fig 5-2 Reported Greek earthquakes grouped in six magnitude ranges for each decade since 1901.

1901 to 1977. Events of magnitude $M=5.8$ are well recorded because they have an average $20 \mu$ ground amplitudes in Uppsala's Wiechert seismograph.

## Table 5-3

Number of earthquakes in the area of Greece reported in each decade since 1901

| Period | $M \leqslant 4.7$ | $4.8 \leqslant M \leqslant 5.2$ | $5.3 \leqslant M<5.7$ | $5.8 \leqslant M \leqslant 6.2$ | $6.3 \leqslant M \leqslant 6.7$ | $M \geqslant 6.8$ | Total |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $1901-1910$ | 0 | 4 | 17 | 13 | 5 | 4 | 43 |
| $1911-1920$ | 0 | 19 | 27 | 16 | 6 | 3 | 71 |
| $1921-1930$ | 0 | 40 | 37 | 17 | 4 | 6 | 104 |
| $1931-1940$ | 0 | 48 | 29 | 9 | 6 | 4 | 96 |
| $1941-1950$ | 0 | 55 | 46 | 16 | 6 | 3 | 126 |
| $1951-1960$ | 48 | 149 | 74 | 18 | 7 | 10 | 306 |
| $1961-1970$ | 360 | 158 | 56 | 28 | 9 | 3 | 614 |
| $1971-1977$ | 282 | 79 | 28 | 7 | 3 | 0 | 399 |

It is likely that these earthquakes have been completely recorded by the Swedish network for at least the past 70 years.

Secondly, for earthquakes with $5.3 \leqslant M \leqslant 5.7$, there is a gradual increase in the number of shocks throughout the whole period. Two interpretations can be given for this behaviour. First is that completeness of data increases with an increasing sample period, and second, that the observed behaviour is due to the statistical fluctuation of activity. Although the first interpretation cannot be rejected on the basis of figure 5-2 alone, the fact that in our magnitude determinations the Uppsala's Wiechert ground amplitude reports are used from 1908 to 1951 , coupled with its unchanged characteristics, supports the possibility of a temporal trend in activity for this range of magnitudes. For these moderate earthquakes Drakopoulos (1976a)also postulates a trend in activity, deduced from observations of
intensities reported in the UNS catalogue for Greece. However the lower part of the range $(5.3-5.4)$ is apparently influenced by the time in which it was recorded. It is likely that earthquakes with magnitude $\mathrm{M} \geqslant 5.5$ are completely reported during the past 60 to 70 years.

A third important feature of Figure $5-2$ is that the most significant jump in the total number of reported events occurs for the last three periods (1951-1977). Although the first decade (1951-1960) was very active (17 earthquakes with $M \geqslant 6.3$, whereas the average number for the remaining period was 9 earthquakes), the huge increment in the total number is due to the contribution of smaller events with $M \leqslant 5.2$. Since 1951 the use of Uppsala and Kiruna's more sensitive instruments affects the threshold of magnitude detectability which, with the use of the ISC determinations and conversion formula (5-9) is, finally, down to $M=4.0$.

### 5.7.2 Analysis of sample completeness

The previous analysis suggests that the present catalogue is severely incomplete below magnitude $M=5.5$ before 1950. But earthquakes with magnitudes less than 5.5 may represent maximum annual magnitude earthquakes. The inclusion of such earthquakes for years in which the reporting is incomplete may affect the results of an extreme value statistical treatment, or any other statistical analysis of the data. An analytical method of assigning time intervals in which a certain magnitude class is likely to be completely reported is needed, and Stepp's (1971) method is chosen here. This method is based on the behaviour of the mean rates of occurrence $\lambda .(M)=N(M) /$ year, where N is the number of events with magnitudes within the class of magnitude M.

Each magnitude class is modelled as a point process in time. The variance of the estimate of a sample mean is inversely proportional to the number of observations in the sample.

The variance of the sample mean, is obtained by assuming that the
earthquake sequence can be modelled by the Poisson distribution. If $\mathrm{k}_{1}$, $k_{2}, \ldots k_{n}$ are the number of quakes per unit time interval, then an unbiased estimate of the mean rate per unit time interval of this sample is (Hamilton, 1964, p 90):

$$
\begin{equation*}
\lambda=\frac{1}{n} \sum_{i=1}^{n} k_{i} \tag{5-14}
\end{equation*}
$$

and its variance is

$$
\begin{equation*}
\sigma_{\lambda}^{2}=\frac{\lambda}{n} \tag{5-15}
\end{equation*}
$$

That is, the variance is equal to the mean where $n$ is the number of unit time intervals. Taking the unit time to be one year gives:

$$
\begin{equation*}
\sigma_{\lambda}=\sqrt{\lambda / T} \tag{5-16}
\end{equation*}
$$

as the standard deviation of the estimate of the mean, where $T$ is the sample length.

Assuming stationarity, the mean, variance and other moments of each observation stay the same, and we expect that $\sigma_{\lambda}$ behaves as $1 / \sqrt{T}$ in the sub-interval in which the mean rate of occurrence in a magnitude class is constant. Departure from $1 / \sqrt{T}$ behaviour, that is from the line parallel to $\sigma_{\lambda}=1 / \sqrt{T}$ through the points ( $\sigma_{\lambda}, \mathrm{T}$ ) of a magnitude class, for a stable mean rate of occurrence, means that either the sub-interval is not long enough to give a good estimate of the mean or that these sub-intervals include periods in which reports are incomplete.

Table 5-4 tabulates the rates of earthquake occurrence as a function of time interval ( $\lambda=N / T$ ) for each of the five chosen magnitude classes, $4.2 \leqslant M \leqslant 4.7,4.8 \leqslant M \leqslant 5.2,5.3 \leqslant M \leqslant 5.7,5.8<M \leqslant 6.2$ and $M \geqslant 6.3$. The difference from the previous magnitude division (see 5.8 .1 ) being the reduction of the two higher classes to one, because there is no question about the

Values of $\lambda$ and $\sigma_{\lambda}$ for five classes of magnitude and time interval $T$

| Period | $\begin{gathered} \mathrm{T} \\ \text { Years } \end{gathered}$ | $1 / \sqrt{T}$ | $4.2 \leqslant M \leqslant 4.7$ |  |  | $4.8 \leqslant M \leqslant 5.2$ |  |  | $5.3 \leqslant M \leqslant 5.7$ |  |  | $5.8 \leqslant \mathrm{M} \leqslant 6.2$ |  |  | $\mathrm{M} \geqslant 6.3$ |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  | $\lambda=\frac{N}{T}$ |  |  | $\lambda=\frac{\mathrm{N}}{\mathrm{T}}$ |  |  | $\lambda=\frac{N}{T}$ | ${ }_{\lambda}$ |  | $\lambda=\frac{\mathrm{N}}{\mathrm{~T}}$ | ${ }^{\circ} \lambda$ |  | $\lambda=\frac{N}{T}$ | ${ }_{\lambda}{ }_{\lambda}$ |
| 1977-1973 | 5 | 0.45 | 210 | 42.0 | 2.90 | 60 | 12.0 | 1.55 | 20 | 4.0 | 0.89 | 5 | 1.0 | 0.45 | 2 | 0.40 | 0.28 |
| 1977-1968 | 10 | 0.32 | 422 | 42.2 | 2.05 | 134 | 13.4 | 1.66 | 51 | 5.1 | 0.71 | 19 | 1.9 | 0.44 | 5 | 0.50 | 0.22 |
| 1977-1963 | 15 | 0.26 | 630 | 42.0 | 1.67 | 212 | 14.1 | 0.97 | 76 | 5.1 | 0.58 | 31 | 2.1 | 0.37 | 12 | 0.80 | 0.23 |
| 1977-1958 | 20 | 0.22 | 669 | 33.5 | 1.29 | 275 | 13.7 | 0.83 | 103 | 5.2 | 0.51 | 41 | 2.1 | 0.32 | 18 | 0.90 | 0.21 |
| 1977-1953 | 25 | 0.20 | 687 | 27.5 | 1.05 | 367 | 14.7 | 0.77 | 147 | 5.9 | 0.48 | 52 | 2.1 | 0.29 | 31 | 1.24 | 0.22 |
| 1977-1948 | 30 | 0.18 | 690 | 23.0 | 0.88 | 413 | 13.8 | 0.68 | 170 | 5.7 | 0.44 | 56 | 1.9 | 0.25 | 37 | 1.23 | 0.20 |
| 1977-1943 | 35 | 0.17 | 690 | 19.7 | 0.75 | 413 | 12.3 | 0.59 | 194 | 5.6 | 0.40 | 61 | 1.7 | 0.22 | 40 | 1.14 | 0.18 |
| 1977-1938 | 40 | 0.16 | 690 | 17.3 | 0.66 | 454 | 11.4 | 0.53 | 211 | 5.3 | 0.36 | 74 | 1.9 | 0.21 | 42 | 1.05 | 0.16 |
| 1977-1933 | 45 | 0.15 | 690 | 15.3 | 0.58 | 476 | 10.6 | 0.48 | 223 | 5.0 | 0.33 | 75 | 1.7 | 0.19 | 48 | 1.07 | 0.15 |
| 1977-1928 | 50 | 0.14 | 690 | 13.8 | 0.53 | 503 | 10.1 | 0.45 | 245 | 4.9 | 0.31 | 84 | 1.7 | 0.18 | 55 | 1.10 | 0.15 |
| 1977-1923 | 55 | 0.13 | 690 | 12.5 | 0.48 | 521 | 9.5 | 0.42 | 261 | 4.8 | 0.29 | 90 | 1.7 | 0.17 | 59 | 1.07 | 0.14 |
| 1977-1918 | 60 | 0.13 | 690 | 11.5 | 0.44 | 540 | 9.0 | 0.39 | 282 | 4.7 | 0.28 | 100 | 1.7 | 0.17 | 62 | 1.03 | 0.13 |
| 1977-1913 | 65 | 0.12 | 690 | 10.6 | 0.40 | 546 | 8.4 | 0.36 | 292 | 4.5 | 0.26 | 109 | 1.7 | 0.16 | 64 | 0.99 | 0.12 |
| 1977-1908 | 70 | 0.12 | 690 | 9.8 | 0.37 | 551 | $7 . .9$ | 0.33 | 300 | 4.3 | 0.25 | 115 | 1.6 | 0.15 | 71 | 1.01 | 0.12 |
| 1977-1903 | 75 | 0.11 | 690 | 9.2 | 0.35 | 552 | 7.4 | 0.31 | 313 | 4.2 | 0.24 | 122 | 1.6 | 0.15 | 78 | 1.04 | 0.12 |
| 1977-1901 | 77 | 0.11 | 690 | 8.9 | 0.34 | 552 | 7.2 | 0.30 | 314 | 4.1 | 0.23 | 124 | 1.6 | 0.14 | 79 | 1.03 | 0.11 |

completeness for events of $M \geqslant 6.3$ for the whole period. $N$ is the cumulative number of earthquakes in the time interval $T$. For each rate $\lambda$, the standard deviation $\sigma_{\lambda}$, is computed using equation (5-16). Values of $\sigma_{\lambda}$ for the five magnitude classes as a function of sample length are plotted on Figure 5-3. The "reference-line" $\sigma_{\lambda}=1 / \sqrt{T}$ is plotted in the top right-hand corner of this figure.

Figures 5-3 and Table 5-4 reveal several features relevant to the statistical treatment of our earthquake catalogue. First, the postulated behaviour of $\sigma_{\lambda}$ is observed, at least over a sub-interval of the total 77year period interval, for all magnitude classes. Secondly, a minimum time interval is required to reach a stable estimate of the mean recurrence rate. This interval is a function of magnitude class, being successively longer with each higher maximum magnitude class. For earthquakes with magnitude $4.2 \leqslant M \leqslant 4.7,5$ to 10 years of homogeneous observations are sufficient to establish a stable mean rate; for a maximum magnitude of 5.2 the minimum observation period is between 10 to 15 years. A stable estimate of the mean recurrence rate of maximum magnitude 5.7 is obtained in about 15 to 20 years, and for maximum magnitude 6.2 in about 30 to 40 years, while for magnitude greater than 6.3 an interval of about $40-50$ years of homogeneous observations is required for a stable estinate of the mean recurrence rate. Thirdly, departure of observed values of $\sigma_{\lambda}$ from $1 / \sqrt{T}$ behaviour with increasing sample length occurs for all magnitudes except those greater than 6.3. The latter departure, can again be explained by incomplete reporting of earthquakes as early data are incorporated into the sample. An alternative explanation is that this is caused by a trend towards increasing frequency in the data. However if the latter is the case, departure from $1 / \sqrt{T}$ behaviour would be expected to occur at the same time in all magnitude classes. The fact that Figure 5-3 does not show a certain departure at the same time for all classes, gives rise to the explanation that there is not an overall


Fig 5-3 Completeness: standard deviation of the estimate of the mean of the annual number of events as a function of sample length. The insert shows the "reference line" $\sigma_{\lambda}=1 \sqrt{T}$ to which the lines must be parallel (see paragraph 5.7.2).
trend of activity, and confirms that departure is due to incomplete reporting.

It is concluded that maximum magnitude 4.7 events are completely reported only during the most recent 15 year interval, events with $4.8 \leqslant M \leqslant 5.2$ during the most recent 30 year interval, events with $5.3 \leqslant M \leqslant 5.7$ during at least the past 60 years. The period in which events with magnitude $5.7 \leqslant M \leqslant 6.2$ are completely reported is during the past 67 years, whereas events with $M \geqslant 6.3$ are completely reported over the whole 77 year sample interval. Table 5-5 summarises the results after applying Stepp's test for completeness.

Table 5-5

Results from test of complet异筑ess for Greece (1901-1977)

| Magnitude class | Time required for stable <br> mean recurrence rate (years) | Period of completely <br> reported events |  |
| :---: | :---: | :---: | :---: |
| $4.2 \leqslant M \leqslant 4.7$ | $5-10$ | 15 | $(1977-1963)$ |
| $4.8 \leqslant M \leqslant 5.2$ | $10-15$ | 30 | $(1977-1948)$ |
| $5.3 \leqslant M \leqslant 5.7$ | $15-20$ | 60 | $(1977-1918)$ |
| $5.8 \leqslant M \leqslant 6.2$ | $30-40$ | 67 | $(1977-1911)$ |
| $M \geqslant 6.3$ | $40-50$ | 77 | $(1977-1901)$ |

The results of Table $5-5$ show that it is possible to create an artificially homogeneous data sample by determining intervals over which earthquakes in different magnitude classes are completely reported. So we can avoid the problem of estimating the recurrence rates from the whole available data sample which for uncorrected data results in the estimated recurrence rates of large earthquakes being overestimated, the recurrence rates for small earthquakes being underestimated.

Appendix B contains the complete list of Greek earthquakes since 1901 along with the explanatory text. There are 1806 earthquakes for which magnitudes are assigned using the procedure described in paragraph 5.6

For the earliest period 1901 until 1907 no ground amplitudes are available, and so only the 32 largest earthquakes with all parameters adopted from the UNS-catalogue are included. UNS's determinations are adopted because its magnitudes for large earthquakes usually closely resemble our determinations. Since 1917 the recalculated locations are detailed along with the total shift in distance (km), and azimuth (degr.), from the ISS locations. As a comparison between ISS and recalculated locations, the average total shift as a function of the period recorded and the percentage of earthquakes which shifted, are calculated and plotted on Figures 5-4 and 5-5 respectively.

The first feature to note is that the largest annual average shifts are obtained for the first decade after 1917, and the average change in position for 78 earthquakes is 165 km . In the second decade the average shift for 95 shocks is 72 km . In the following third and fourth decades and the remaining seven years (1957-1963) with 84,182 and 166 shocks, the average shifts decrease rapidly with values of $59 \mathrm{~km}, 35 \mathrm{~km}$ and 17 km respectively. The second significant feature is that changes in epicentres from the ISS locations are more than 40 km for $40 \%$ of the total number of recalculated quakes, and more than 50 km for $34 \%$.

The method chosen to relocate the events, by using as master events earthquakes with as much macroseismic information as possible, coupled with station adjustments derived from each group of events, strengthens the possibility that these large changes reflect the degree of inaccuracies in ISS locations, especially for the early events. The fact that these changes rapidly decrease from decade to decade, as more stations report arrivals,


Fig 5-4 Mislocation errors: average total epicentre shift in distance (km) from ISS location to the recalculated location, as a function of the period recorded.

## SHIFT



Fig 5-5 Mislocation errors: percentage of total epicentre shift in distance (km) from ISS location to the recalculated location exceeding a certain shift.
supports this explanation. As a further test, Table 5-6 tabulates the 39 worst cases (largest shifts) in which the UNS catalogue gives different locations, usually based on macroseismic information, along with the total shifts from the ISS and UNS locations. The average shift from UNS locations, for the same events is only 39.9 km . Again this supports the explanation that these large shifts are due to ISS mislocations for reasons discussed in paragraph 5.3.

Figure $5-6$ plots the shift in distance versus change in azimuth for each relocated earthquake showing that except for five of the earliest events $(24.12 .1917,27.12 .1917,09.02 .1918,22.03 .1919$, and 10.05 .1921$)$ which have an azimuthal change around $220^{\circ}-240^{\circ}$, there is no correlation between the shift in direction and the shift in distance. These five earthquakes were previously located in the north-eastern part of Greeece, but macroseismic information (Galanopoulos 1961, UNS 1974) shows that the new locations are very close to macroseismic epicentres (see also Table 5-6).

Differences in locations between ISS and the new epicentres certainly reflect differences in the methods used. But while for the earlier events it is obvious that ISS mislocations are due to different travel-time models used, or because of adoption of old positions for later events, the 17 km average total shift found for the latest period (1956-1963), using many more stations all around the world with very sensitive instruments, is a significant figure. During this period, the major differences between JED and ISS methods of event location are the different travel-time tables used, and the station adjustments which JED applies to the first arrivals.

However, Crampin and Üçer (1975) using Jeffreys-Bullen travel times and the SPEEDY program without station adjustments point out that for the Marmara Sea area the average shift in position for the period 1956 to 1963 is about 5 km from the ISS locations. On the other hand, in the ATB catalogue for Turkey (Alsan et al, 1975) when "Herrin" travel-time tables are

Cases of large shifts from ISS and UNS locations

| DATE | Shift from in distance (km) | SS locations in azimuth(deg) | Shift from in distance (km) | NS locations in azimuth(deg) |
| :---: | :---: | :---: | :---: | :---: |
| 14.03.1917 | 232.4 | 305.7 | 42.4 | 142.3 |
| 24.12.1917 | 410.8 | 241.3 | 18.0 | 17.4 |
| 09.02.1918 | 445.4 | 237.5 | 63.4 | 62.6 |
| 22.08.1919 | 580.6 | 233.3 | 203.2 | 197.6 |
| 29.11.1920 | 126.2 | 60.0 | 67.5 | 84.0 |
| 30.03.1921 | 254.7 | 283.0 | 47.2 | 221.0 |
| 10.05.1921 | 415.5 | 237.7 | 26.6 | 303.7 |
| 10.08.1921 | 157.4 | 10.8 | 58.6 | 68.6 |
| 13.09.1921 | 98.6 | 22.3 | 24.1 | 275.5 |
| 14.09.1921 | 157.6 | 204.9 | 54.8 | 184.6 |
| 09.02 .1922 | 144.7 | 187.6 | 24.9 | 113.9 |
| 20.05.1923 | 135.8 | 213.7 | 40.1 | 232.3 |
| 01.08 .1923 | 129.8 | 106.0 | 49.4 | 137.9 |
| 13.11.1924 | 117.9 | 138.7 | 21.1 | 122.7 |
| 12.04.1925 | 598.9 | 307.2 | 66.9 | 201.4 |
| 01.09.1925 | 226.3 | 225.6 | 51.3 | 162.9 |
| 26.02.1926 | 140.5 | 286.5 | 33.4 | 80.2 |
| 26.06.1926 | 123.8 | 312.2 | 54.4 | 300.9 |
| 05.07.1926 | 134.7 | 218.7 | 40.0 | 278.1 |
| 30.08 .1926 | 83.5 | 170.1 | 7.5 | 331.9 |
| 01.07.1927 | 87.7 | 188.8 | 14.5 | 80.5 |
| 18.04.1928 | 90.6 | 314.7 | 35.8 | 77.2 |
| 28.04.1928 | 91.9 | 291.7 | 25.4 | 116.3 |
| 11.09 .1931 | 154.3 | 9.4 | 20.5 | 293.1 |
| 23.09.1931 | 226.3 | 347.8 | 37.5 | 128.7 |
| 26.09.1932 | 105.6 | 0.0 | 28.8 | 0.0 |
| 01.06.1933 | 221.8 | 7.23 | 20.8 | 7.23 |
| 02.07.1933 | 325.7 | 358.0 | 79.4 | 264.7 |
| 03.09.1935 | $-\quad 161.7$ | 5.2 | 41.1 | 97.6 |
| 08.08.1936 | 130.3 | 15.6 | 54.5 | 28.3 |
| 03.06 .1938 | 66.1 | 226.5 | 2.1 | 238.7 |
| 18.09.1938 | 116.5 | 268.8 | 30.4 | 354.9 |
| 23.02 .1940 | 81.6 | 323.7 | 23.4 | 267.3 |
| 14.02.1943 | 114.7 | 333.2 | 24.9 | 2.1 |
| 27.05.1944 | 147.3 | 222.7 | 37.3 | 311.3 |
| 04.06 .1947 | 58.9 | 22.2 | 11.6 | 341.1 |
| 29.11 .1947 | 104.0 | 353.8 | 12.8 | 286.5 |
| 10.08.1948 | 174.5 | 2.3 | 46.3 | 279.9 |
| 17.09.1949 | 62.8 | 72.3 | 14.7 | 303.9 |

Average total shift ISS: 185.6 km
Average total shift UNS: $\quad 39.9 \mathrm{~km}$


Fig 5-6 Mislocation errors: shift in distance (km) versus change in azimuth (degr.) for the relocated events compared to the ISS locations.
used, it is concluded that there are no significant differences between ISC locations, calculated with Jeffreys-Bullen travel-time tables, and their locations.

It seems reasonable to conclude that a significant part of the 17 km shift in position found for the period since 1956 is due to the station adjustments applied to the first arrival data by the JED method. In fact, for the 202 earthquakes of this period which are located using SPEEDY, with and without station adjustments, the changes in relative position (ie, distance between the two positions of the same event with and without station adjustments) have an average shift of 8.6 km .

The results of these tests show that the relocated /centres are significantly different from those of ISS with the average shift decreasing from decade to decade since 1917. Even for the most recent period of investigation since 1956 there is still a significant change in position, with an average shift of 17 km .
5.9 Summary

Because none of the existing earthquake catalogues for Greece $\left(N_{33}^{42.5}\right.$, $\mathrm{E}_{19}^{29}$ ) fulfil the demands of modern seismology for accuracy, homogeneity and completeness, this chapter has attempted to reduce these inaccuracies as far as available data permits. Considering the data available for recalculation of source parameters, the catalogue presented here contains:
i) Earthquakes for which all parameters are calculated using the JED method (Period: 1917-1963).
ii) Earthquakes for which magnitudes are determined using the Swedish network ground amplitude records (Period: 1908-1977).
iii) Earthquakes adopted from the UNS catalogue for all parameters, because neither first arrival nor ground amplitude data is available (Period: 1901-1907). not recalculations of the other parameters because of limited number of arrivals. For these earthquakes, parameters other than magnitude are adopted from other sources after special investigation for as much macroseismic information as possible (Period: 1908-1977).

The last two cases are included in the catalogue for completeness. Although the accuracy of the adopted earthquakes is not known, the criterion for adoption using macroseismic information eliminates large errors in location.

For all these earthquakes magnitudes are determined using the same procedure as for the rest of the data sample, and so magnitude homogeneity is retained.

When Stepp's test of completeness is applied, the results show that only earthquakes with magnitude greater than 6.3 are completely reported during the whole period of investigation (1901-1977), whereas earthquakes with a maximum magnitude $M \leqslant 4.7$ are completely reported only during the most recent 15 years (1963-1977). The time required for stable mean recurrence rate is found to be $40-50$ years of homogeneous observations for magnitude greater than 6.3 , whereas for earthquakes with magnitude between 4.2 and 4.7, only 5 to 10 years of homogeneous observations are sufficient to establish a stable mean recurrence rate.

The JED method chosen to relocate the events has the advantage of using master events, which coupled with its capability of detecting source, travel-times, network and station bias and the facility to combine these into a single "term" (ie station adjustment), guarantees the highest possible accuracy for the relocated earthquake epicentres. The consistent treatment of all available data, particularly magnitude determinations, ensures a high degree of homogeneity for the whole period of investigation.

Completeness of the catalogue is limited by data availability, but
using the results of Table $5-4$, it is possible to determine intervals over which earthquakes in different magnitude classes are completely reported.

Finally, because the accuracy of the following seismic risk investigation is related to the earthquake catalogue used, comparisons between ISS original locations and those of the present catalogue are worthwhile.

The comparison tests show that ISS locations are severely biased. These large average total shifts imply that it is not possible to describe the detailed seismicity of Greece by just using the ISS data sample for the whole period. This earthquake catalogue for Greece has a high degree of homogeneity, accuracy and completeness, and permits more detailed seismotectonic studies to be made on the basis of a long, instrumentally recorded data sample.

In the next chapter existing tectonic models will be tested to see if they are in accord with the recalculated parameters of the earthquakes. This will then be followed by the application of the statistical techniques of the first part of this thesis to the new earthquake catalogue to estimate seismic risk for Greece.

## GREEK TECTONICS AND SEISMICITY

6.1 Introduction

The spatial distribution of earthquakes in a region shows its present active tectonics, and the size of earthquake magnitudes is a measure of the degree of the activity. Hence, maps with the spatial distribution of the epicentres can reveal the tectonic features of the region with as much precision as the accuracy of the earthquake parameters used.

Greece and the adjacent areas (ie the Greek mainland, the Aegean Sea and western Turkey) have the highest seismic activity in the whole Mediterranean and European area (Kárnik, 1969; Galanopoulos, 1971a). The high seismic activity shows that this area is tectonically very active. This, coupled with the fact that it is a part of the Alpine-Himalayan zone, which is the only continental region where large scale shortening is now taking place (McKenzie, 1978), makes it a region of great interest for geologists and geophysicists.

In this chapter, an attempt is made to examine the validity of existing tectonic models using the recalculated parameters of the earthquakes in the area (Appendix B). Furthermore, using the recalculated depths and radial vertical distance-depth cross-sections, three dimensional isodepth contouring maps are produced. These maps reveal several significant features of the tectonic process in the region.
6.2 Morphologic, geologic and geophysic feature of the area

The main morphologic and geologic features of the area of Greece and the adjacent areas are, from south to north (see Fig. 6-1):

1. the Mediterranean ridge (or chain)
2. the Hellenic trench (or trough)


MERCATOR

Fig 6-1 Summary map of the Aegean region, showing morphologic and geologic trends in a schematic way.
3. the Hellenic arc, and
4. the northern Aegean Sea.

The Mediterranean ridge has irregular topography and extends from the Ionian Sea to Cyprus. It is not a mid-ocean ridge, and Finetti (1976) investigating its tectonic features in detail suggests the name "east Mediterranean chain".

The Hellenic trench consists of a series of depressions to a depth of 5100 m which parallels the Hellenic arc.

The Hellenic arc is formed by the outer sedimentary arc, a link between the southern Dinarides and the Turkish Taurides, and the inner volcanic arc which parallels the sedimentary arc. Between these two arcs is the Cretan trough with water depth to about 2000 m . The outer sedimentary arc consists of Paleozoic to Tertiary rocks folded and faulted in several phases of the Alpine orogeny, while the inner volcanic arc consists of recent andencitic volcanism at Santorini, Nisyros, Milos and Kos.

The Aegean Sea immediately north of the volcanic arc is a rather stable block of folded Paleozoic and granitoid masses. The extreme north includes the northern Aegean trough with water depth to about 1500 m , the northeast extension of which is probably the small depression of the Marmara Sea (Papazachos and Comninakis, 1976).

Greece was surveyed gravimetrically and magnetically in the years 1971-1973 (Makris, 1975). Along the Greek mainland the Bouguer anomalies have negative values with a gravity minimum of -140 mGal situated at the Pindos Mountains. The Aegean Sea is characterized by positive Bouguer anomalies with a maximum of +175 mGal at the central trough of the Cretan Sea (Makris, 1975), while in the central and northern Aegean it is about +50 mGal . A belt of negative free-air anomalies down to -200 mGal follows the Hellenic trench, while the Bouguer anomalies are positive up to +180 mGal (More11i et al, 1975).

Positive magnetic anomalies have been determined in several parts of the Aegean Sea. The strongest of these anomalies have been observed along the volcanic arc, in the northern Aegean trough and in the Cretan trough (Vogt and Higgs, 1969; Makris, 1973). The magnetic field is undisturbed in the Mediterranean Sea south of Crete (Vogt and Higgs, 1969).

Heat flow is relatively high in the Aegean Sea floor ( 2.1 HFU ) in the volcanic arc of the southern Aegean and Jongsma (1974), has interpreted it as due to underthrusting of oceanic crust.

Seismic refraction studies and experiments (Papazachos, 1969 ; Makris, 1973 , 1976a) have indicated that the crust thins from about 50 km below the Peloponnesus and the Pindus Mountains toward the Aegean ( $25-30 \mathrm{~km}$ ) and that the central part of the Cretan Sea crust is only 20 km thick (Makris, 1976b; see Fig. 6-2).
6.3 Principle tectonic models for Greece and the adjacent areas By definition (McKenzie and Parker, 1967), seismic belts mark the boundaries of stable plates, and focal mechanisms indicate the relative motions of adjacent plates. Focal-mechanism studies (Constantinescu et al, 1966; Papazachos and Delibasis, 1969; McKenzie, 1970, 1972; Ritsema, 1974) suggest southerly to westerly thrusting of the arc over the Mediterranean.

McKenzie $(1970,1972)$ was the first to delineate a small, rapidly moving plate, which contains the Aegean, part of Greece, Crete and part of western Turkey (see Fig 6-3). He called it the "Aegean plate". The south-western boundaries were well defined, and earthquake fault plane solutions show that the motion between the Aegean and African plates is in a north-south direction. The northern boundary was defined by extensional and transform faults, and he concluded that it was a continuation of the North Anatolia fault (but see below). The boundary with the other


Fig 6-2 Contour map of the Moho discontinuity from gravity and seismic data (after Makris, 1973).


Fig 6-3 Sketch of plate boundaries and motions in the Aegean area obtained from seismicity and fault plane solutions (after McKenzie, 1972).
plate, the "Turkish plate", was poorly defined.
After McKenzie's work, contemporary plate tectonics in the area, and its problems, were discussed by Lort (1971), Papazachos and Comninakis (1971, 1976, 1978), Galanopoulos (1972b, 1973, 1974, 1975), Comninakis and Papazachos(1972, 1976), Alvarez (1973), Dewew et al (1973), Papazachos (1973, 1974, 1976a, 1976b, 1977), Makris (1973, 1975, 1976a, 1976b, 1978), Gregersen (1977) and others. The common point of almost all these studies is that the African plate underthrusts Greece and the adjacent areas along the Hellenic arc. The mean dipping angle is about $35^{\circ}$ (Papazachos and Comninakis, 1971; Galanopoulos, 1973; Agarwal et a1, 1976; Gregersen, 1977). However, McKenzie's model has been critized by a number of authors (Papazachos, 1973, 1976a, 1976b, 1977; Crampin and Üçer, 1975; Mercier et al, 1976) for its simpicity and its definition of its northern and western boundaries.

From the definition of the boundaries of a plate (see above), it is difficult to talk about truly stable aseismic microplates in this region. All the maps of spatial distribution of epicentres show that several small aseismic blocks exist. Hence, most geologists and geophysicists now prefer the name "Aegean area" rather than "plate" because of its real complexity.

As Figure $6-3$ shows, the northern boundary of McKenzie's plate consists of transform faults, but Mercier et al (1976), after extensive investigations in central Greece, found no evidence of a transform fault. The continuation of that northern boundary towards the North Anatolia fault has also been debated (Papazachos, 1976a, 1977; Crampin and Üçer, 1975).

Papazachos (1976a) using focal mechanisms and the spatial distribution of earthquakes in the northern Aegean, has concluded that there is an amphitheatrical Benioff-zone which, although less well defined compared with the similar one in the south Aegean, is dipping towards a
thrust region which includes the northernmost part of the Aegean and part of the Marmara area.

A different model for the Aegean area has been suggested by Makris (1976b, 1978). According to his model, the deformation of the region is the surface expression of a hot mantle plume which extends to the base of the lithosphere and has been mobilized through compressional processes that forced the lithosphere to sink into the asthenosphere. The model is based on refractional-seismic data from which a low velocity of the compressional waves of $7.7 \mathrm{~km} / \mathrm{sec}$ for the upper mantle has been determined, on gravity measurements, which show that density lower than normal is extending to the base of the lithosphere, and on the high values of heat flow in the Aegean area.

This model explains that the crustal thickening along the Hellenic arc is due to the crustal down-buckling which is thickening at the compressional front. This collision is responsible for the high seismicity along the arc. The Hellenic trench is the result of the upwards movement of the Aegean crust which is forced to override part of the IonianEast Mediterranean crust and lithosphere towards Africa. This movement causes a subduction zone to develop at the collision front. According to this model, the deep seismicity is caused by crust and upper mantle fragments dislocated from their original positions and subducted into the soft, low Q asthenosphere. Thus more complicated Benioff-zones develop, which differ from the Benioff-zones of the Pacific, because the interaction is between continental-continental or continental-subcontinental blocks.

Recently, McKenzie (1978) published another model for the Aegean Sea and surrounding regions. This model is a modification of the previous one of Fig 6-3. It is based on new fault plane solutions of earthquakes taken from USGS and NOAA, Landsat photographs, and seismic refraction records. The main points of this model are the following:
i) Rapid extension is taking place in the northern and eastern parts of the Aegean Sea region, whereas the southern part is relatively inactive.
ii) From the fact that the pre-Miocene geology of the islands of the Aegean closely resembles that of Greece and Turkey, and from the thin ( $\sim 30 \mathrm{~km}$ ) crust of the southern part of the Aegean compared with the thick ( $\sim 50 \mathrm{~km}$ ) crust beneath Greece and Turkey (see also Fig. 6-2), it is concluded that the thin crust of the Aegean has been produced by stretching the orogeny belt by a factor of two since the Miocene. This stretching can account for the high heat flow, while the sinking slab produced by subduction along the Hellenic arc maintains the motion.
iii) In north-western Greece and Albania, where he found both thrust and normal faults, while his data did not show any sinking slab, the term "blob" of cold mantle detaching from the lower half of the lithosphere is introduced. These blobs are produced by thermal instability when the lithosphere is thickened by thrusting.
iv) The direction of relative motion between the southern Aegean region and Africa determined from fault plane solutions is $211^{\circ}$ E.
v) The new information suggests that the North Anatolia Fault does not extend beneath the Aegean, and that the motions are taken up on several structures with components of normal faulting.
vi) There is no evidence to support the suggestion made in his previous model (1972) that the Anatolian Trough is connected to the Gulf of Corinth.

The main differences between his previous model (Fig 6-3) and the new one are summarized in Fig 6-4, taken from his recent paper (Mackenzie, 1978, Fig 18).


Fig 6-4 Summary of the present deformation of theAegean area after McKenzie (1978). LLong curved lines show normal faults. Lines with open semicircles show thrust faults. Solid dots mark epicentres of shocks for which mechanisms are used. Arrows show the direction of motion obtained from fault plane solutions. The long heavy arrow shows the direction of relative motion between the Aegean and Africa. Heavy Vs mark sites of recent volcanism.)

McKenzie (1978) also comments that "the theory of plate tectonics is of little value in regions such as northern Greece and Turkey where the deformation is spread over a zone". Dewey and Sengor (1978) also point out that plate tectonics is not useful in the Aegean area where normal faulting is not confined to a narrow zone. These two comments and the wide criticism which McKenzie directed against almost all the proposed models for the region may reflect the real complexity of the Aegean area.

### 6.4 Spatial distribution of the epicentres

6.4.1 Shallow earthquakes ( $h<60 \mathrm{~km}$ )

Figure 6-5 is a map of the epicentres of shallow earthquakes for Greece and the adjacent areas ( $33.5^{\circ} \mathrm{N}$ to $42.5^{\circ} \mathrm{N}, 18.5^{\circ} \mathrm{E}$ to $29.5^{\circ} \mathrm{E}$ ) for the period 1901-1978, based on the recalculated parameters (1917-1963) and those of UNS and ISC. A total of 1492 events is mapped which, because of the consistent treatment (see Chapter $V$ ), is believed to represent the most complete, accurate and homogeneous data sample available for this purpose.

Comparing this map with similar ones previously published, it is apparent that, although the general pattern of all maps is the same because of the high activity of the region, a better delineation of the seismic zones is achieved. A well defined seismic belt which runs parallel to the Hellenic Arc, and contains the majority of shallow earthquakes, can be seen in Figure 6-5. To the north, this zone runs well within Albania in an almost north-south direction, whereas its south-east part joins the diffused zone of Asia Minor (the west coast of Turkey).

A second well defined zone with a ENE-WSW trend runs from Leukas Island, where it meets the first zone, and through central Greece reaching the east coast of Greece near Volos. After that, it is divided into two less well defined branches. One turns to the north and covers the



Fig 6-5 Spatial distribution of the epicentres of shallow depth earthquakes for Greece since 1901.

Chalkidiki peninsula, and the other continues through the Sporades Islands and, in an east-west direction, joins Asia Minor after reaching the north coast of Lesvos Island. However, this branch seems to be divided near the eastern Sporades, and a new narrow zone is developed with an almost north-south direction, which passes through the west coast of Limnos Island and ends, quite sharply, near the south coast of Thassos Island. From the distribution of shallow earthquakes in the northeast part of the region (Marmara area), it is not clear whether the North Anatolian Fault extends towards the Aegean Sea, or diverts into western Turkey, or both.

A third well defined zone is that which follows the Saronikos and Corinth gulfs. This zone at the west end and, in the middle of the gulf of Patras, is curved, and joins the previous zone (central Greece) in a north-south direction, rather than continuing and meeting the arcuated first zone (along the Hellenic arc).

From Figure 6-5, it is clear that at least three well defined aseismic blocks exist: the attikocycladic block, which is part of what McKenzie (1972) calls the "Aegean plate", the block formed by north-eastern Greece and the aseismic block in central Greece around the Ptolemais basin. The existence of these aseismic blocks indicates that it is difficult to talk about a simple plate model for the region. It means that the lithosphere is very fragmented, which is illustrated by the existence of these several small aseismic blocks.
6.4.2 Intermediate earthquakes ( $\mathrm{h}>60 \mathrm{~km}$ )

Figure $6-6$ is a map of the spatial distribution of the intermediate depth earthquakes. Different symbols are used for the different depth ranges, and the size of the symbols is proportional to magnitude ranges, as the caption of the figure describes.

From Figure 6-6 it can be seen that most of the earthquakes with



Fig 6-6 Spatial distribution of the epicentres of intermediate depth earthquakes for Greece since 1901.
depth $\mathrm{h}<100 \mathrm{~km}$ are associated with the Hellenic arc.
In previously published maps, nearly all large intermediate earthquakes in the area of Greece were limited to the region south of $38.5^{\circ} \mathrm{N}$ (Drakopoulos, 1976b). However, from Figure 6-6 it is clear that there is a significant number of intermediate earthquakes in the north part of the region which had been reported as shallow. From both Figures 6-5 and 6-6 it can be seen that the distribution of the hypocentres is not well represented by a simple dipping Benioff zone, either to the south or to the north. Most parts have both shallow and intermediate earthquakes, without any clear or systematic change in depth as the distance from the Hellenic arc lengthens towards the volcanic arc.

Figure 6-7 shows the shift in position and depth for the earthquakes which occurred during the last ten years of the recalculated period (19531963). Two sizes of symbols are used. The smaller size symbols show the locations taken from the ISS data file, while the larger ones show the recalculated locations. The different symbols for different depth ranges are the same as in Figures $6-5$ and $6-6$, where depths for the ISS locations are those of UNS's catalogue because ISS usually did not calculate depth values.

From Figure 6-7 it can be seen that the majority of the recalculated depths for intermediate earthquakes along the Hellenic arc have shallower depths than those of UNS, especially in the southwest part of Crete. In that part of the Hellenic arc, no earthquake with a depth greater than 100 km is found. The opposite is the case for the southeast part of the arc (eastern Crete, Karpathos, Rodos) and the volcanic arc... Generally a tendency of increasing depths is observed. This may be due to the fact that UNS's depth determinations were based upon macroseismic information using relations among $I, M$ and $h$ (UNS, 1974). As is pointed out in the UNS catalogue the intermediate depth earthquakes seem to have a very big


Fig 6-7 Shift in position and changes in depth between the recalculated and ISS locations for earthquakes of the period 1953-1963. Large depth
symbols show the recalculated positions, small depth symbols show the positions reported by ISS.



Fig 6-8 Spatial distribution of all earthquakes for Greece since 1901.
vertical extension combined with the inclination of the main rupture, especially for the Aegean Sea. In such cases all the instrumental data forces us to treat the earthquakes as of intermediate depth, but all the macroseismic data presents them as shallow. The spatial distribution of all earthquakes for the period under investigation is mapped in Figure 6-8. 6.5 Isodepth maps from radial vertical cross-sections 6.5.1 Procedure

Using the parameters of the new earthquake catalogue, two and threedimensional isodepth maps are produced in an attempt to examine the validity of the existing models, and to distinguish places where further work must be done, rather than to present another model. These maps result from continuous radial vertical cross-sections using the following procedure:
i) The approximate centre of the volcanic arc is chosen as a reference point. Its coordinates are $38.9^{\circ} \mathrm{N}, 26.0^{\circ} \mathrm{E}$.
ii) Radial vertical planes are drawn with a common axis to the one passing through the reference point, and differing by an azimuth of $10^{\circ}$ (36 vertical planes).
iii) In each plane the epicentral distance from the centre versus focal depth is plotted for all earthquakes within $\pm 10^{\circ}$ azimuth from that plane. Thus, 36 radial vertical cross-sections are produced, each of which overlaps the one adjacent by an azimuth of $10^{\circ}$.
iv) For each of these 36 cross-sections, a smooth curve following the lower part (deeper earthquakes) of the projected events is drawn. Earthquakes which have a depth less than 50 km are excluded. Thus, the coordinates of the points taken from these curves, that is distance from the reference point and depth, constitute the data to be contoured. Figures 6-9a to 6-9f illustrate this procedure.

Cross-section owro ver. Mave seoce rav.
 Fig. 6-9a Four vertical radial planes passing through the reference point ( $38.9{ }^{\circ} \mathrm{N}, 26.0^{\circ}$ E). The hypocentres projected
onto each plane have azimuths between $\pm 10^{\circ}$ from the plane. Heavy lines are those from which the points for onto each plane have azimuths between $\pm 10^{\circ}$ from the plane. Heavy lines are those from which the points for the isodepth maps are taken.

enoss-section owra vent.teme lzooke rne



Cnoss-section owro rent.mane iseote in
Fig 6-9b For explanation see caption of Fig 6-9a.





Fig 6-9c For explanation see caption of Fig 6-9a.


Cnoss-section owro rent.mane zeoote inf





Fig 6-9d For explanation see caption of Fig 6-9a.



Cnoss-section onta vent.plane zseoce rar


## CNoss-section anto vent.mane zeooce FNr



## CNoss-section outa vent. PLawe 27e0es me

Fig $6-9 \mathrm{e}$ For explanation see caption of Fig 6-9a.



cnoss-sectrow onto vent.rinuc zacoce tse


## chass-scetiom anta rent.rlant seoses tra


anoss-rection owto verr.prane 31eoce mp
Fig 6-9f For explanation see caption of Fig 6-9a.

Figures $6-10$ and $6-11$ a to $6-11 \mathrm{~h}$, resulting from the previously described procedure, show the following features:
a) The Hellenic trench

All the tectonic models of the region agree that the Hellenic arc is a collision front between Africa and the Aegean area and that Africa underthrusts the Aegean area causing a subduction zone to develop at the collision front. This subduction zone is easily seen in all the figures.

The isodepth of $85-115 \mathrm{~km}$ runs almost parallel to the Hellenic trench, with increasing depth towards the two ends. In the north-western end of the trench this isodepth continues to the north, and runs well into Albania, where it has not previously been mapped.

An isodepth of $115-145 \mathrm{~km}$ is developed in the convex north-western end of the Hellenic trench which includes the north-western part of Greece, Corfu Island and the southwest coast of Albania. From the isodepths alone it is difficult to conclude if there is a "blob" (McKenzie, 1978) or a poorly defined Benioff zone (Papazachos and Comninakis, 1976), or if this thickening is caused by the existence of the rigid Apullian block beneath the sea to the west, which is stronger than the heavily deformed belts of Albania and Greece (McKenzie, 1978). However, the clear continuation of this arcuated zone north of Leukas Island and the existence of thrust faults (see Fig 6-4), may suggest that subduction is taking place. This area is certainly among those where much more work is needed in order to understand the present tectonic process.

At the south-eastern end of the Hellenic trench (east of Crete) the depth distribution shows that the subduction zone continues and meets western Turkey and Cyprus, through the Pliny and Strabo trenches, rather than the coast of western Turkey near Rodos Island. From the isodepth


Fig 6-10 Isodepth map for Greek earthquakes obtained using radial vertical cross-sections shown in Figs 6-9a - 6-9f. Symbols for focal depth ranges are indicated in the insert.


Fig 6-11a Three dimensional isodepth maps for Greek earthquakes obtained using vertical cross-sections shown in Figs 6-9a - 6-9f. The depth scale in km is indicated in the insert. R.P: reference point for all cross-iectional diagrams.


$$
\begin{array}{ll}
\text { azimuth }=225 & \text { altitude }=45 \\
\text { *width }=10.00 & \text { rheight }=2.00 \\
\text { smoothigas }=-1.00 &
\end{array}
$$

* before foreshortening 30/08/78

Fig 6-11b For explanation see caption of Fig 6-11a.
$\begin{array}{ll}\text { azimuth }=45 & \text { attirude }=60 \\ \text { nwidt } h=10.00 & \text { wheight }=2.00\end{array}$

* before foreshortening 19/10/78
\(\left.\begin{array}{l}1.00 <br>
0.50 <br>

0.00\end{array}\right]\)| -60.00 |
| ---: |
| -130.00 |
| -200.00 |

Fig 6-11c For explanation see caption of Fig 6-11a.

$\begin{array}{ll}\text { A 3D ISODEPTH MAP FOR GREECE VIEWED } \\ \text { azimuth }=45 & \text { alt itude }=45 \\ * \text { widt }=10.00 & * \text { ight }=2.00 \\ \text { smoothings }=-1.00 & \\ * \text { before foreshort ening } & 30 / 08 / 78\end{array}$
Fig 6-11d For explanation see caption of Fig 6-11a


Fig 6-11e For explanation see caption of Fig 6-11a.


A 30 I SODEPTH MAP FOR GREECE VIEWED FROM THE NORTH-WEST
$\begin{array}{ll}\text { azimuth }=135 & \text { altitude }=45 \\ \text { *width }=10.00 & \text { nheight }=2.00 \\ \text { smoothings }=-1.00 & \end{array}$

* before fureshortening 30/08/78

Fig 6-11f For explanation see caption of Fig 6-11a.


Fig 6-11g For explanation see caption of Fig 6-11a.


A 30 I SODEPTH MAP FOR GREECE VIEWED FROM THE SOUTH-EAST $\begin{array}{ll}\text { azimuth }=315 & \text { altitude }=45 \\ \text { *widrh }=1 \mathrm{C} .00 & \text { *height }=2.00 \\ \text { smath:ngs }=-1.00 & \end{array}$
smorth:ngs $=-1.00$

* befure fureshortening 30/0я/78

Fig 6-11h For explanation see caption of Fig 6-11a.
maps it is apparent that south of Rodos the most extensive and deepest seismicity of the whole area is found, and earthquakes with depths within $180-210 \mathrm{~km}$ dominate the area.

Although the majority of the intermediate shocks are related to the subduction zone now occurring along the Hellenic arc, from the isodepth maps it can be seen that there is no clear increase in depth with distance from the thrust zone, and the distribution of deeper earthquakes does not follow the volcanic arc. Thus, it is difficult to define a simple Benioff zone in the southern part of the Aegean region.

The less active part of the arc is between the Peloponnesus and western Crete, but if the relative motion between the Aegean area and Africa is in a sousthwest direction, as the results of McKenzie's fault plane solution shows (big arrow in Fig 6-4), this part of the arc should be the most active. Furthermore, the dip of the slab of subduction in that direction should be greater beneath southern Greece than beneath the south-eastern part of the Aegean. McKenzie (1978) points out that this may indicate a seismic gap or changes in direction during the period of subduction. Only further work on determining the history of deformation behind the Hellenic arc may explain the relative movements along this arc.
b) The Aegean area

In the southern part of the Aegean area there are four places with maximum isodepths:
i) The north-western part of the Peloponnesus, the Saronikos gulf and the eastern part of the Corinth gulf. Papazachos (1977) interprets the existence of thrust faulting for three intermediate earthquakes, and their difference from the surrounding shallow earthquakes in this region, as due to a sinking slab from the Ionian Sea to the Aegean. However, McKenzie's fault plane
solutions (1978), and field mapping of Mercier et al (1976), show that most of the deformation in this area is produced by normal faulting (see also Fig 6-4).
ii) The south-eastern part of the Peloponnesus.
iii) South of Rodos Island, and
iv) West of Kos Island.

The general picture of the south part of the region is an arcuated distribution but with no clear increasing depths with distance from the trench, the deepest part of the region being the part south of Rodos which probably continues towards western Turkey.

The main feature in the central part of the Aegean area, is that central Greece and the Aegean Sea are characterised by shallow seismic activity. This, coupled with the existence of two aseismic blocks (see 6.4.1), shows that there is no subcrustal evidence that the Northern Anatolian fault is connected with the seismic zone of central Greece, or the gulf of Corinth.

In the northern part of the Aegean the features are more complicated than elsewhere. The intermediate shocks are fewer than in the southern part. The fact that thrust and normal faulting exist (Ritsema, 1974; Papazachos, 1976a), and the low depth seismicity of the region, lead Papazachos (1976a) to suggest a northerly sinking slab produced by a subduction zone in the northern Aegean. McKenzie (1978) points out that there is no evidence of thrusting on the scale required for such a suggestion, and that these shocks may lie within material subducted at a trench which is no longer active. From the depth distribution alone none of these hypotheses can be rejected. This is another region for which more geophysical data is necessary for a better understanding of its deep tectonic process.

The new earthquake catalogue and the spatial and depth distribution maps based on it reveal that Greece, and the adjacent areas of the Aegean are tectonically more complicated than had previously been recognised. The seismicity of the region is mainly due to the collision between the underthrusting African plate and the Aegean area along the Hellenic arc. Using the parameters of the new catalogue, a better delineation of the seismic activity of the region is achieved, and the spatial and depth distribution of earthquakes show that none of the proposed tectonic models completely explain the observed activity over the whole area.

The existence of small aseismic blocks of which three are well defined shows that the lithosphere is very fragmented, and the region can not be modelled by a simple plate. The earthquake depth distribution and the clear continuation of the seismic activity within Albania along the Hellenic arc may suggest that in the north-western end of the Hellenic trench subduction is taking place. At the south-eastern end of this trench the depth distribution shows that the subduction zone continues and meets Turkey much further south than the existing models suggest. It is difficult to define a simple Benioff zone either in the south part or in the north part of the Aegean Sea, because there is no clear increase in depth with distance from the thrust zone, and the distribution of deeper earthquakes does not follow the volcanic arc. To fully understand the tectonic process further work is still necessary, especially in the following places:
i) The north-western part of the Hellenic arc (north-western Greece, south-western coast of Albania).
ii) The south-eastern end of the arc (Eastern Crete, Karpathos and Rodos Islands).
iii) The northern Aegean Sea and north-western part of Turkey (Marmara Sea area).

In the next Chapter, the seismic risk of Greece will be examined using the earthquake parameters of the new catalogue, the Extreme-Value statistical method of the third type asymptotic distribution combined with strain energy release method already described. Special emphasis will be given to six of the largest cities of Greece.

## GREEK SEISMIC RISK EVALUATION

Several attempts have been made to map seismicity and seismic risk of Greece. Galanopoulos (1968) calculated the seismic risk expression in recurrence rates of shallow earthquakes with $m>5 \frac{1}{2}, 6,6 \frac{1}{2}$, and 7 in each square degree of Greece. Comninakis (1975) defined the seismic risk in terms of the most probable annual maximum magnitude from the $a$ and $b$ values per $\frac{1}{4}$ square degree and for the sample period 1911 to 1970. Algermissen et al (1976) used the UNS catalogue to compile seismic risk maps of the Balkan region, depicting acceleration and velocity with $70 \%$ probability of not being exceeded in 25 and 200 year periods.

However, the calculation of seismic risk in terms of expected magnitude, acceleration, velocity or displacement, depends critically on the time span considered, and also on the earthquake parameters used. If the time span is not sufficient to establish stable estimates of risk, then ensuejng risk maps will differ for different periods of observation. Thus, comparison between the two seismic maps of Galanopoulos (1968) and Comninakis (1975) shows contradictory results for the area of the north Aegean, because the first map was compiled before the strong activity in the north Aegean during 1965-1967. On the other hand, any uncertainty or significant revision in magnitude produces a related uncertainty in, or requires a recalculation of, the seismic risk parameters.

The new earthquake catalogue for Greece is used in this chapter to evaluate Greek seismic risk in terms of magnitude, by using both the energy release and third type asymptotic distribution methods. The results are presented in the form of contour maps of annual and 80 -year most probable
maximum magnitudes, and magnitudes with $70 \%$ probability of not being exceeded in the next 50 and 100 years. Risk calculations are then made in terms of acceleration using an "average formula" for acceleration attenuation which is derived from most of the commonly used formulae. An attempt to apply the third asymptotic distribution method to the annual maximum observed accelerations, results in poor convergence with values of $\lambda \rightarrow 0.0$, which shows that this asymptotic curve tends to the first type asymptotic distribution. Thus the risk in terms of acceleration is computed using the first type asymptotic distribution only, and contour maps of acceleration with $70 \%$ probability of not being exceeded in the next 50,100 , and 200 years, illustrate the results.

Seismic risk for six of the heavily industrial and highly populated centres of Greece is examined in more detail.
7.2 Data, and cities for seismic risk estimation

The new earthquake catalogue of Appendix B is used for the seismic risk evaluation of Greece as a whole and for the six following cities of Greece:

| Athens $:$ | $37.97^{\circ} \mathrm{N}, 23.72{ }^{\circ} \mathrm{E}$ |
| :--- | :--- | :--- |
| Thessaloniki: | $40.64^{\circ} \mathrm{N}, 22.93^{\circ} \mathrm{E}$ |
| Patra $:$ | $38.23^{\circ} \mathrm{N}, 21.75^{\circ} \mathrm{E}$ |
| Corinth : | $37.92^{\circ} \mathrm{N}, 22.93^{\circ} \mathrm{E}$ |
| Heraklion : | $35.35^{\circ} \mathrm{N}, 25.18^{\circ} \mathrm{E}$ |
| Rodos $:$ | $36.43^{\circ} \mathrm{N}, 28.27^{\circ} \mathrm{E}$ |

For each of these cities the distribution of shallow plus intermediate earthquakes for a radius of 100 km and 150 km from their centres, is taken and analysed, with both energy release and Extreme-Value methods. These radii are chosen in order to obtain estimates of the seismic risk more applicable to normal and tall structures. Normal size buildings are mainly
subject to the seismic risk associated with strong local earthquakes with high frequency seismic waves ( $N \geqslant 3 \mathrm{c} / \mathrm{sec}$ ), whereas tall buildings with longer resonant periods can be seriously affected by more distant earthquakes.

### 7.3 Seismic risk based on the magnitude distribution

Table 7-1 tabulates the three parameters of the third type asymptotic distribution, along with their uncertainties and the reduced chi-square. These parameters are calculated using annual extreme magnitudes and the method described in Chapter IV. The column "missing years" contains the number of years without reported earthquakes.

The results for $M_{1}, M_{2}$ and $M_{3}$ using the method of energy release, which is described in Chapter III, and the comparable quantities derived from the parameters of Table $7-1$ using the equations (4-21) and (4-41) are tabulated in Table 7-2. Figures 7-1a to 7-7d illustrate both methods.

### 7.3.1 Comparison of the methods applied

The first feature to note from Tables $7-1$ and $7-2$ is the close relation between the results of the two different methods. In places in which the cumulative energy release graphs include at least one well defined cycle of periodicity, as in Figures $7-2 \mathrm{a}$ and $7-2 \mathrm{c}$ (Athens for 100 km and 150 km radii), Figures 7-5a and 7-5c (Corinth), Figure 7-4a (Patra for 100 km radius), the parameters of the third type asymptotic distribution are well defined, and they have small uncertainties. In places where it is not clear if the periodic cycle is completed, as in Figures $7-3 a$ and $7-3 c$ (Thessaloniki) and Figure $7-7 a$ (Rodos for 100 km radius), the parameters are accompanied by larger uncertainties.

A second feature is that in almost all cases the magnitude distribution has a remarkably good third type asymptotic behaviour. This is apparent from the figures of the third type asymptotic curves and the observed data

Estimated parameters of the third type asymptote

| Place | $\omega$ | $\sigma_{\omega}$ | $u^{\prime}$ | $\sigma_{u}$ | $\lambda$ | $\sigma_{\lambda}$ | Reduced <br> chi-square | Missing <br> years |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :---: | :---: |
| Athens (100) | 6.80 | $\pm 0.39$ | 2.98 | $\pm 0.52$ | 0.595 | $\pm 0.193$ | 0.027 | 47 |
| Athens (150) | 7.35 | $\pm 0.58$ | 4.19 | $\pm 0.21$ | 0.402 | $\pm 0.151$ | 0.042 | 38 |
| Thessaloniki (100) | 8.19 | $\pm 1.17$ | 2.39 | $\pm 0.77$ | 0.363 | $\pm 0.169$ | 0.232 | 58 |
| Thessaloniki (150) | 8.57 | $\pm 0.83$ | 3.58 | $\pm 0.30$ | 0.346 | $\pm 0.119$ | 0.159 | 45 |
| Patra (100) | 6.69 | $\pm 0.46$ | 4.12 | $\pm 0.26$ | 0.504 | $\pm 0.206$ | 0.035 | 40 |
| Patra (150) | 8.23 | $\pm 0.99$ | 5.12 | $\pm 0.07$ | 0.238 | $\pm 0.111$ | 0.039 | 21 |
| Corinth (100) | 6.75 | $\pm 0.32$ | 3.66 | $\pm 0.41$ | 0.671 | $\pm 0.220$ | 0.046 | 43 |
| Corinth (150) | 7.15 | $\pm 0.44$ | 4.68 | $\pm 0.11$ | 0.432 | $\pm 0.147$ | 0.074 | 29 |
| Heraklion (100) | 7.86 | $\pm 1.17$ | 3.73 | $\pm 0.32$ | 0.291 | $\pm 0.158$ | 0.128 | 46 |
| Heraklion (150) | 8.93 | $\pm 1.97$ | 4.87 | $\pm 0.09$ | 0.176 | $\pm 0.117$ | 0.046 | 27 |
| Rodos (100) | 9.23 | $\pm 3.29$ | 3.88 | $\pm 0.31$ | 0.170 | $\pm 0.158$ | 0.146 | 48 |
| Rodos (150) | 9.39 | $\pm 2.02$ | 4.53 | $\pm 0.12$ | 0.176 | $\pm 0.062$ | 0.046 | 38 |
| Greece | 8.73 | $\pm 0.65$ | 6.21 | $\pm 0.04$ | 0.236 | $\pm 0.073$ | 0.023 | 0 |

Estimated parameters of the energy release method. $X_{1}, X_{2}$ and $\omega$ are from equations $(4-21),(4-41)$ and Table 7-1.

| Place | $\mathrm{M}_{1}$ | $\mathrm{X}_{1}$ | $\mathrm{M}_{2}$ | $\mathrm{X}_{2}$ | $\mathrm{M}_{3}$ | $\omega$ | Waiting time (years) |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Athens (100) | 4.2 $\pm .1$ | $\begin{aligned} & 4.5 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 5.7 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 5.9 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 6.7 \\ & \pm .3 \end{aligned}$ | $\begin{aligned} & 6.8 \\ & \pm .4 \end{aligned}$ | 33 |
| Athens (150) | $\begin{aligned} & 4.6 \\ & \pm .1 \end{aligned}$ | $\begin{gathered} 4.8 \\ \pm .1 \end{gathered}$ | $\begin{aligned} & 6.0 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 6.1 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 7.1 \\ & \pm .3 \end{aligned}$ | $\begin{aligned} & 7.4 \\ & \pm .6 \end{aligned}$ | 34 |
| Thessaloniki (100) | $\begin{aligned} & 3.7 \\ & \pm .2 \end{aligned}$ | $\begin{gathered} 3.4 \\ \pm .4 \end{gathered}$ | $\begin{aligned} & 6.0 \\ & \pm .2 \end{aligned}$ | $\begin{aligned} & 6.3 \\ & \pm .2 \end{aligned}$ | $\begin{aligned} & 7.3 \\ & \pm .4 \end{aligned}$ | $\begin{gathered} 8.2 \\ \pm 1.17 \end{gathered}$ | 58 |
| Thessaloniki (150) | $\begin{aligned} & 4.3 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 4.3 \\ & \pm .2 \end{aligned}$ | $\begin{aligned} & 6.6 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 6.7 \\ & \pm .2 \end{aligned}$ | $\begin{aligned} & 7.9 \\ & \pm .4 \end{aligned}$ | $\begin{aligned} & 8.6 \\ & \pm .8 \end{aligned}$ | 61 |
| Patra (100) | $\begin{aligned} & 4.7 \\ & \pm .1 \end{aligned}$ | $\begin{array}{r} 4.9 \\ \pm .1 \end{array}$ | $\begin{aligned} & 5.7 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 5.8 \\ & \pm .2 \end{aligned}$ | $\begin{aligned} & 6.9 \\ & \pm .3 \end{aligned}$ | $\begin{aligned} & 6.7 \\ & \pm .5 \end{aligned}$ | 45 |
| Patra (150) | $\begin{aligned} & 5.2 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 5.3 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 6.3 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 6.4 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 7.5 \\ & \pm .4 \end{aligned}$ | $\begin{array}{r} 8.2 \\ \pm 1.0 \end{array}$ | 49 |
| Corinth (100) | $\begin{aligned} & 4.7 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 5.2 \\ & \pm .2 \end{aligned}$ | $\begin{aligned} & 5.8 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 6.1 \\ & \pm .3 \end{aligned}$ | $\begin{aligned} & 6.9 \\ & \pm .3 \end{aligned}$ | $\begin{aligned} & 6.8 \\ & \pm .3 \end{aligned}$ | 36 |
| Corinth (150) | $\begin{aligned} & 5.0 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 5.3 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 6.1 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 6.2 \\ & \pm .2 \end{aligned}$ | $\begin{aligned} & 7.1 \\ & \pm .3 \end{aligned}$ | $\begin{aligned} & 7.2 \\ & \pm .4 \end{aligned}$ | 29 |
| Heraklion (100) | $\begin{aligned} & 4.3 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 4.2 \\ & \pm .2 \end{aligned}$ | $\begin{aligned} & 5.9 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 5.9 \\ & \pm .2 \end{aligned}$ | $\begin{aligned} & 7.1 \\ & \pm .4 \end{aligned}$ | $\begin{array}{r} 7.9 \\ \pm 1.9 \end{array}$ | 54 |
| Heraklion (150) | $\begin{aligned} & 5.0 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 5.0 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 6.3 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 6.3 \\ & \pm .2 \end{aligned}$ | $\begin{aligned} & 7.5 \\ & \pm .4 \end{aligned}$ | $\begin{array}{r} 8.9 \\ \pm 1.9 \end{array}$ | 52 |
| Rodos (100) | $\begin{aligned} & 4.4 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 4.1 \\ & \pm .3 \end{aligned}$ | $\begin{aligned} & 6.0 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 6.1 \\ & \pm .3 \end{aligned}$ | $\begin{aligned} & 7.2 \\ & \pm .4 \end{aligned}$ | $\begin{array}{r} 9.2 \\ \pm 3.3 \end{array}$ | 64 |
| Rodos (150) | $\begin{aligned} & 4.8 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 4.7 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 6.4 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 6.5 \\ & \pm .2 \end{aligned}$ | $\begin{aligned} & 7.5 \\ & \pm .4 \end{aligned}$ | $\begin{array}{r} 9.4 \\ \pm 2.0 \end{array}$ | 39 |
| Greece | 6.2 $\pm .1$ | $\begin{aligned} & 6.4 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 7.1 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 7.2 \\ & \pm .1 \end{aligned}$ | $\begin{aligned} & 8.2 \\ & \pm .4 \end{aligned}$ | $\begin{aligned} & 8.7 \\ & \pm .6 \end{aligned}$ | 33 |

GREECE 1901-1978


Fig 7-1a Cumulative energy release as a function of time for Greece (1901-1978).


Fig 7-1b Third type asymptotic distribution curve for Greece (1901-1978), + indicates observed annual maximuin magnitude.

## ATHENS 100KM



Fig 7-2a Cumulative energy release as a function of time for an area of 100 km radius from the city of Athens (1901-1978).


Fig 7-2b Third type asymptotic distribution curve for an area of 100 km radius from the city of Athens (1901-1978).

ATHENS 150kM


Fig 7-2c Cumulative energy release as a function of time for an area of 150 km radius from the city of Athens (1901-1978).

ATHENS $R=150 \mathrm{KM}$


Fig 7-2d Third type asymptotic distribution curve for an area of 150 km radius from the city of Athens (1901-1978).

THESSALONIKI 100KM


Fig 7-3a Cumulative energy release as a function of time for an area of 100 km radius from the city of Thessaloniki (1901-1978).

THESSALONIKI $R=100 \mathrm{KM}$


Fig 7-3b Third type asymptotic distribution curve for an area of 100 km radius from the city of Thessaloniki (1901-1978).
Fig

$$
\begin{aligned}
& \text { ! } 7 \\
& \text { nj }
\end{aligned}
$$

PATRA 100 KM


Fig 7-4a Cumulative energy release as a function of time for an area of 100 km radius from the city of Patra (1901-1978)


Fig 7-4b Third type asymptotic distribution curve for an area of 100 km radius from the city of Patra (1901-1978).

## PATRA 150KM



Fig 7-4c Cumulative energy release as a function of time for an area of 150 km radius from the city of Patra (1901-1978).

## PATRA $R=150 \mathrm{KM}$



Fig 7-4d Third type asymptotic distribution curve for an area of 150 km radius from the city of Patra (1901-1978).

## CORINTH 100km



Fig 7-5a Cumulative energy release as a function of time for an area of 100 km radius from the city of Corinth (1901-1978).

CORINTH $\quad R=100 K M$


Fig 7-5b Third type asymptotic distribution curve for an area of 100 km radius from the city of Corinth (1901-1978) .


Fig 7-5c Cumulative energy release as a function of time for an area of 150 km radius from the city of Corinth (1901-1978).


Fig 7-5d Third type asymptotic distribution curve for an area of 150 km radius from the city of Corinth (1901-1978).


Fig 7-6a Cumulative energy release as a function of time for an area of 100 km radius from the city of Heraklion (1901-1978).


Fig 7-6b Third type asymptotic distribution curve for an area of 100 km radius from the city of Heraklion (1901-1978).

## HERAKLION 150KM



Fig 7-6c Cumulative energy release as a function of time for an area of 150 km radius from the city of Herak1ion (1901-1978).

HERAKLION R=150KM


Fig 7-6d Third type asymptotic distribution curve for an area of 150 km radius from the city of Heraklion (1901-1978).

## RODOS 100KM



Fig 7-7a Cumulative energy release as a function of time for an area of 100 km radius from the city of Rodos (1901-1978).


Fig 7-7b Third type asymptotic distribution curve for an area of 100 km radius from the city of Rodos (1901-1978).

## RODOS 150KM



Fig 7-7c Cumulative energy release as a function of time for an area of 150 km radius from the city of Rodos (1901-1978).


Fig 7-7d Third type asymptotic distribution curve for an area of 150 km radius from the city of Rodos (1901-1978).
points in Figures $7-2$ to $7-7$, as well as from the small values of chisquare (Table 7-1).

From Table 7-1 and Figures $7-3 b$ and $7-3 d$, it can be seen that Thessaloniki has the largest chi-square values and poorest fit for both 100 km and 150 km radii. The distribution of the observed data points, however, may suggest that this is caused by the superposition of two natural populations of earthquakes, as in Figure $4-3 a$ and $4-3 b$ for the Aleutians and Alaska region. In fact, the area around Thessaloniki, in the north Aegean area, is tectonically much more debated than any other part of Greece (see paragraph 6.3).

### 7.3.2 Seismic risk evaluation - magnitude

From the parameters of the third type asymptotic distribution, the return period in years for earthquake magnitudes $m=5.0$ to $m=8.0$, with a step of $\Delta \mathrm{m}=0.5$, are calculated and listed in Table 7-3. The uncertainties can be computed from equation (4-1) using the formula
$\sigma^{2} \phi=\sigma^{2} \omega\left(\frac{\partial \phi}{\partial \omega}\right)^{2}+\sigma^{2} u\left(\frac{\partial \phi}{\partial u}\right)+\sigma^{2} \lambda\left(\frac{\partial \phi}{\partial \lambda}\right)^{2}+2 \sigma^{2} \omega u\left(\frac{\partial \phi}{\partial \omega}\right)\left(\frac{\partial \phi}{\partial u}\right)+\ldots$
and the complete variance-covariance matrix (4-14) for the parameters at each place. Then, the uncertainties for the different return periods for the given place can be derived from equation (2-19).

Tables 7-4 and 7-5 tabulate the number of exceedances during the next 50 and 100 years for each of the magnitudes previously listed in Table 7-3. The results for Greece and the city of Athens are discussed here in detail, whereas, the values of all parameters for the remaining five places are also tabulated in the same tables and plotted as for the special cases of Greece and the city of Athens.

## a) Greece

For the area of Greece the upper bound for earthquake magnitudes

Return period (years) for given magnitudes ( $M_{s}$ )

| Place | Ms | M.0 | 5.5 | 6.0 | 6.5 | 7.0 | 7.5 |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- |
| Athens (100) | 4.1 | 6.6 | 14.3 | 71.1 | - | - | - |
| Athens (150) | 2.6 | 4.3 | 8.7 | 26.0 | 222.9 | - | - |
| Thessaloniki (100) | 5.7 | 8.7 | 15.0 | 30.0 | 77.8 | 343.6 | - |
| Thessaloniki (150) | 3.2 | 4.6 | 7.3 | 13.2 | 28.8 | 86.2 | 530.6 |
| Patra (100) | 2.8 | 5.2 | 14.3 | 184.1 | - | - | - |
| Patra (150) | 1.4 | 2.3 | 4.6 | 12.3 | 49.8 | 442.2 | - |
| Corinth (100) | 2.9 | 4.4 | 8.7 | 42.9 | - | - | - |
| Corinth (150) | 1.9 | 3.1 | 6.4 | 22.5 | 654.0 | - | - |
| Heraklion (100) | 4.1 | 7.4 | 16.0 | 46.1 | 222.2 | - | - |
| Heraklion (150) | 1.8 | 3.2 | 6.9 | 19.1 | 69.3 | 378.8 | - |
| Rodos (100) | 4.4 | 8.6 | 18.9 | 49.3 | 156.4 | 664.3 | - |
| Rodos (150) | 2.6 | 4.1 | 8.3 | 19.8 | 57.3 | 216.1 | - |
| Greece | 1.0 | 1.1 | 1.3 | 2.4 | 5.5 | 21.9 | 198.4 |

Table 7-4

Number of exceedances during the next 50 years v . Magnitude $\left(M_{\mathrm{s}}\right)$

| City | 5.0 | 5.5 | 6.0 | 6.5 | 7.0 | 7.5 | 8.0 |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Athens (100) | $12-13$ | $7-8$ | $3-4$ | $0-1$ |  |  |  |
| Athens (150) | $19-20$ | $11-12$ | $5-6$ | $1-2$ | - | - | - |
| Thessaloniki (100) | $8-9$ | $5-6$ | $3-4$ | $1-2$ | 0.1 | - | - |
| Thessaloniki (150) | $15-16$ | $10-11$ | $6-7$ | $3-4$ | $1-2$ | $0-1$ | - |
| Patra (100) | $17-18$ | $9-10$ | $3-4$ | 0.1 | - | - | - |
| Patra (150) | $34-35$ | $22-22$ | $10-11$ | $4-5$ | $1-2$ | - | - |
| Corinth (100) | $17-18$ | $11-12$ | $5-6$ | $1-2$ | - | - | - |
| Corinth (150) | $25-26$ | $16-17$ | $7-8$ | $2-3$ | - | - | - |
| Heraklion (100) | $12-13$ | $6-7$ | $3-4$ | $1-2$ | $0-1$ | - | - |
| Herak1ion (150) | $28-29$ | $15-16$ | $7-8$ | $2-3$ | $0-1$ | - | - |
| Rodos (100) | $11-12$ | $5-6$ | $2-3$ | $1-2$ | - | - | - |
| Rodos (150) | $19-20$ | $12-13$ | $6-7$ | $2-3$ | $0-1$ | - | - |
| Greece | $49-50$ | $47-48$ | $37-38$ | $20-21$ | $9-10$ | $2-3$ | $0-1$ |

Table 7-5

Number of exceedances during the next 100 years v. Magnitude $\left(M_{s}\right)$

|  | 5.0 | 5.5 | 6.0 | 6.5 | 7.0 | 7.5 | 8.0 | Maximum <br> Observed <br> (78 years |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Athens (100) | $24-25$ | $15-16$ | $7-8$ | $1-2$ | - | - | - | 6.6 |
| Athens (150) | $38-39$ | $23-24$ | $11-12$ | $3-4$ | 0.1 | - | - | 7.0 |
| Thessaloniki (100) | $17-18$ | $11-12$ | $6-7$ | $3-4$ | $1-2$ | - | - | 7.1 |
| Thessaloniki (150) | $31-32$ | $21-22$ | $13-14$ | $7-8$ | $3-4$ | $0-1$ | -1 | 7.8 |
| Petra (100) | $35-36$ | $19-20$ | $7-8$ | $0-1$ | - | - | - | 6.6 |
| Petra (150) | $68-69$ | $43-44$ | $21-22$ | $8-9$ | $2-3$ | $0-1$ | - | 7.3 |
| Corinth (100) | $34-35$ | $22-23$ | $11-12$ | $2-3$ | - | - | - | 6.6 |
| Corinth (150) | $51-52$ | $32-33$ | $15-16$ | $4-5$ | 0.1 | - | - | 7.0 |
| Heraklion (100) | $24-25$ | $13-14$ | $6-7$ | $2-3$ | 0.1 | - | - | 7.1 |
| Heraklion (150) | $56-57$ | $31-32$ | $14-15$ | $5-6$ | $1-2$ | 0.1 | - | 7.2 |
| Rodos (100) | $22-23$ | $11-12$ | $5-6$ | $2-3$ | $0-1$ | - | - | 7.1 |
| Rodos (150) | $38-39$ | $24-25$ | $12-13$ | $5-6$ | $1-2$ | $0-1$ | - | 7.3 |
| Greece | $99-100$ | $94-95$ | $75-76$ | $40-41$ | $18-19$ | $-4-5$ | $0-1$ | 8.0 |

is found to be

$$
8.16 \pm 0.43 \text { or } 8.73 \pm 0.65
$$

for the energy release and third type asymptotic distribution methods respectively. During the 78 years of instrumental data the maximum earthquake magnitude had a value of 8.0 (11.8.1903). Galanopoulos (1972a) using the graphical method for strain energy release (see Fig 3-10) estimated the upper bound and found a value of $m=8 \frac{1}{2}$. Considering the seismic history of the area (Galanopoulos, 1960, 1961; Richter, 1958; Lomnitz, 1974) the value of $m=8.73 \pm 0.65$ as an upper bound for surface-wave magnitude is acceptable. The "waiting time" (see paragraph 3.2.2) for an earthquake equal to the upper bound is 33 years (Fig 7-1a).

The most probable annual maximum (mode) is $m=6.4 \pm 0.1$. Galanopoulos (1972a) and Comninakis (1975), using the frequency-magnitude law, estimated the mode as 6.5 and 6.4 respectively.

The annual rate of energy release corresponds to an earthquake with magnitude $m=7.2 \pm 0.1$ (Table 7-2 and Figure 7-1a).

Table 7-3 shows that in Greece an earthquake with magnitude greater than or equal to 5.5 is expected to be the annual maximum magnitude almost every year (return period $=1.06$ years). The period required for an earthquake with magnitude greater than or equal to 7.0 to be observed as an annual maximum is 5.5 years, whereas an interval of about 200 years is needed for an earthquake with magnitude greater than or equal to the largest observed in the region (11.8.1903, $m=8.0$ ) to be an annual extreme magnitude. However, the return period for an earthquake with magnitude 7.5 is only 22 years.

Table 7-6 contains the number of exceedances of given maximum magnitude earthquakes which is predicted for Greece, along with the number observed, during the 78 year sample period. This shows that for both low

Predicted and Observed number of exceedances

| Magn m $\geqslant$ | Greece |  |  | Athens ( $\mathrm{R}=100 \mathrm{~km}$ ) |  |  | Athens ( $\mathrm{R}=150 \mathrm{~km}$ ) |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | No of exceedances |  |  | No of exceedances |  |  | No of exceedances |  |  |
|  | Predicted |  | Observed | Predicted |  | Observed | Predicted |  | Observed |
|  | 50 years | 100 years | 78 years | 50 years | 100 years | 78 years | 50 years | 100 years | 78 years |
| 5.0 | 49-50 | 99-100 | 78 | 11-14 | 24-25 | 22 | 19-20 | 38-39 | 36 |
| 5.5 | 47-48 | 94-95 | 76 | 7-8 | 15-16 | 12 | 11-12 | 23-24 | 24 |
| 6.0 | 37-38 | 75-76 | 63 | 3-4 | 7-8 | 6 | 5-6 | 11-12 | 14 |
| 6.5 | 20-21 | 40-41 | 36 | 0-1 | 1-2 | 2 | 1-2 | 3-4 | 4 |
| 7.0 | 9-10 | 18-19 | 19 | - | - | - | - | 0-1 | 1 |
| 7.5 | 2-3 | 4-5 | 5 | - | - | - | - | - | - |
| 8.0 | 0.1 | 0-1 | 1 | - | - | - | - | - | - |

and high magnitudes the predicted numbers agree with the maximum magnitudes observed. The close agreement between observed and predicted extreme magnitudes not only illustrates the plausibility of the method, but also shows that the sample period considered is long enough for stable estimates.
b) Athens $\left(37.97^{\circ} \mathrm{N}, 23.72^{\circ} \mathrm{E}\right)$

For the city of Athens the upper bound for earthquake magnitude has
the value of

$$
m=6.80 \pm 0.39(100 \mathrm{~km}) \text { or } m=7.35 \pm 0.58(150 \mathrm{~km})
$$

The "waiting times" for these magnitudes are 33 and 34 years respectively (Fig 7-2a and 7-2c). During the 78 year sample period the maximum observed earthquake magnitudes were 6.6 (28.08.1962) and $7.0(30.08 .1926)$ within a distance of 100 km and 150 km from Athens. The most probable annual maximum magnitude within the same distance is: $4.5 \pm 0.1$ and $4.8 \pm 0.1$ respectively. The mean annual rates of energy release correspond to earthquakes with magnitudes $5.7 \pm 0.1$ and $6.0 \pm 0.1$. The same quantities from the third type asymptotic distribution analysis have the values of $5.9 \pm 0.1$ and $6.0 \pm 0.1$. From Table $7-6$ it can be seen that for both radii there is also a remarkably good agreement between predicted and observed number of exceedances over the whole range of magnitude.

However, the example of Athens has its own unusual significance. It shows that although the number of missing years is greater than the number of observed extremes, the assumption that the first observed extreme is ranked as $j+1$, assuming the first $j$ of the $N$ observations are not available (see paragraph 4.5 .2 (a)), leads to realistic results from the third type asymptotic distribution method.

### 7.4.1 Choosing the acceleration-distance formula

Since 1964 when Esteva and Rosenblueth proposed a general formula
(2-15) for ground motion, most of the existing formulae are modifications of (2-15) rather than different models of the changes of the characteristics of earthquake ground motion with distance. The main reason for this is a lack of large numbers of strong motion records necessary for regional studies.

Table 7-7 tabulates some of the best known attenuation formulae in terms of maximum acceleration with remarks on their validity. From Figure $7-8$, where these formulae are plotted for $m=7.5$ and $h=10 \mathrm{~km}$, it is apparent that they present a large dispersion and can not be universally accepted. Thus, for the estimation of seismic risk in terms of maximum acceleration for Greece, it is thought that an average formula derived from the relations of Table 7-7 may be a more reliable way to approach the problem. The validity of such a formula can then be evaluated by the degree of its concordance with the existing local strong motion records. From the formulae of Table 7-7 the "average formula" for maximum acceleration is found, by trial and error, to be:

$$
\begin{equation*}
A=2164 e^{0.70 m}(R+20)^{-1.80} \mathrm{~cm} / \mathrm{sec}^{2} \tag{7-2}
\end{equation*}
$$

with uncertainties $\Delta \mathrm{b}_{2}= \pm 0.03$ and $\Delta \mathrm{b}_{3}= \pm 0.02$. This formula is also plotted in Figure 7-8. To demonstrate the validity of this formula the upper, lower and average values for the whole range of the eight curves, and for epicentral distances from 10 km to 120 km for $\mathrm{m}=7.5$ and $\mathrm{h}=10 \mathrm{~km}$, are tabulated in Table 7-8 along with the values from equation (7-2). The next step is to check the proposed formula with observations from the existing Greek strong motion records. Since 1972, when the first accelerograph was installed in Greece, there are eight accelerograms

Maximum acceleration formulae from which the "average formula" (7-2)
is derived

| 1. $\mathrm{A}=1080 \cdot \mathrm{e}^{0.5 \mathrm{M}} \cdot(\mathrm{R}+25)^{-1.32}$ | Donovan, 1973 | in $\mathrm{cm} / \mathrm{sec}^{2}$, more than 20 f soil overlying the rock |
| :---: | :---: | :---: |
| 2. $A=6.6 \cdot 10^{-2} \cdot 10^{0.4 M_{L}} \cdot R^{-1.39}$ | Orphal and Lahoud, 1974 | in $g$, hard rock. $\Delta \mathrm{b}_{2}= \pm 0.076, \Delta \mathrm{~b}_{3}= \pm 0.063$ |
| 3. $A=5600 \cdot e^{0.8 M} \cdot(R+40)^{-2}$ | Esteva, 1974 | in $\mathrm{cm} / \mathrm{sec}^{2}$, hard rock |
| 4. $A=5000 \cdot e^{0.8 M} \cdot(R+40)^{-2}$ | Shah et a1, 1975 | in $\mathrm{cm} / \mathrm{sec}^{2}$, hard rock |
| 5. $A=1230 e^{0.8 M}(R+13)^{-2}$ | Ahorner and Rosenhaur, 1975 | in $\mathrm{cm} / \mathrm{sec}^{2}$, hard rock |
| 6. $A=1.03 \mathrm{~h}^{0.6} \cdot 10^{0.54 \mathrm{M}} \cdot \mathrm{R}^{-1.5}$ | $\text { Bath, } 1975$ | in $\mathrm{cm} / \mathrm{sec}^{2}$ |
| 7. $\log A=2.308-1637 \log (\mathrm{R}+30)+0.411 \mathrm{M}$ | Katayama, 1974 | in $\mathrm{cm} / \mathrm{sec}^{2}$ |
| 8. $\log A_{p}=M+\log A(R)-\log A_{o}(M, s, p, v)$ | Trifunac, 1976 | ```in cm/sec}\mp@subsup{}{}{2 p: conf. leve1, s: type of soil, v: component``` |



Fig 7-8 Peak acceleration as a function of epicentral distance using $m=7.5$ and $h=10 \mathrm{~km}$, for the eight acceleration attenuation formulae of Table 7-7 and the derived average formula (7-2).

Values of maximum acceleration ( $\mathrm{cm} / \mathrm{sec}^{2}$ ) for epicentral distance $R(\mathrm{~km})$ and derived from the formulae of Table $7-7$, using $M=7.5$ and $h=10 \mathrm{~km}$

| Formula | 10 km | 20km | 30 km | 40 km | 50km | 60km | 70 km | 80km | 90 km | 100 km | 110 km | 120 km |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Esteva 1975 | 771 | 581 | 440 | 342 | 273 | 222 | 184 | 155 | 132 | 114 | 100 | 88 |
| Bath 1975 | 865 | 435 | 259 | 174 | 126 | 97 | 77 | 64 | 53 | 45 | 40 | 35 |
| Donavan 1973 | - | - | 253 | 181 | 151 | 128 | 111 | 98 | 87 | 78 | 70 | 64 |
| Orphal et al 1974 | 1661 | 878 | 542 | 375 | 280 | 218 | 177 | 148 | 125 | 108 | 95 | 84 |
| Shah et al 1975 | 688 | 519 | 393 | 305 | 244 | 198 | 164 | 138 | 118 | 102 | 89 | 78 |
| Ahorner et al 1975 | 673 | 397 | 249 | 168 | 121 | 91 | 71 | 57 | 46 | 39 | 33 | 28 |
| $\begin{array}{lr} \text { Trifunac } & 1976 \\ (\mathrm{p}=0.5, \quad \mathrm{~s}=2, & \mathrm{h} \leqslant 15) \\ \hline \end{array}$ | 922 | 546 | 310 | 180 | 112 | 78 | 58 | 45 | 38 | 34 | 30 | 27 |
| Katayama 1974 | 501 | 380 | 288 | 229 | 186 | 151 | 128 | 109 | 95 | 85 | 74 | 68 |

Table 7-8b

Range and averages of acceleration values ( $\mathrm{cm} / \mathrm{sec}^{2}$ ) $v$ distance ( km ) bracketed by the eight formulae of Table 7-8a, and comparable values derived from the proposed formula of equation 7-2.

| Upper - Lower | 922-673 | 581-397 | 440-249 | 342-168 | 280-112 | 222-78 | 184-58 | 155-45 | 132-38 | 115-34 | 98-30 | 88-27 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Average | $798 \pm 125$ | $490 \pm 92$ | $345 \pm 100$ | $255 \pm 87$ | $196 \pm 84$ | $150 \pm 72$ | $121 \pm 63$ | $100 \pm 55$ | $85 \pm 47$ | $75 \pm 40$ | $64 \pm 34$ | $58 \pm 30$ |
| Proposed formula | $716 \pm 215$ | $486 \pm 148$ | $340 \pm 105$ | $250 \pm 79$ | $191 \pm 61$ | $151 \pm 48$ | $123 \pm 40$ | $102 \pm 33$ | $86 \pm 28$ | $74 \pm 24$ | $64 \pm 21$ | $56 \pm 19$ |

available. Table 7-9 contains the maximum recorded accelerations for these shocks taken from Drakopoulos (1976b), whereas the other parameters are from the catalogue presented in this study. The last column gives the values of the maximum acceleration derived from equation(7-2). From this table it can be seen that values from the "average formula" (7-2) agree with mo $t$ of the observed values of maximum acceleration. 7.4.2 Method used to fit the maximum acceleration data

Equation (7-2) is applied to each earthquake, and the ranked annual maximum accelerations form the data sample for the Extreme-Value method, similarly to that used for maximum magnitudes. An attempt to apply the third type asymptotic distribution to the annual maximum accelerations results in poor convergence with values of $\omega$ as high as 10 g and values of $\lambda$ close to 0.0 . However, when the curvature parameter $\lambda$ tends to zero, then the third type asymptotic curve becomes the first type asymptotic distribution.

An explanation for this may be that the value of peak acceleration from equation (7-2) depends not only on the magnitude, but also on the focal distance from the point of interest. Because of the nature of attenuation (Fig 7-8) which is rapid for focal distances less than 40 km , and slow towards the longest distances for which the strong motion is highly attenuated, the data points are concentrated at low values of acceleration with occasional high values for near earthquakes. In this situation the straight line (first asymptote) seems to fit the data better than the three parameter curve of the third asymptote, and so the first type asymptote is applied to the extreme accelerations for Greece.
7.4.3 Procedure used to evaluate the seismic risk in $T$ years at a given probability level

The seismic risk of the six selected places is calculated using

Table 7-9

Observed peak accelerations from strong motion records of Greek earthquakes compared with the values predicted by equation $7-2$

| Date | Origin time <br> h:m:s | Station | $M_{s}$ | $R_{k m}$ | A observed <br> $\mathrm{cm} / \mathrm{sec}^{2}$ | A from <br> $\mathrm{cm} / \mathrm{sec}^{2}$ |
| :---: | :---: | :--- | :--- | :--- | :--- | :--- | :---: |
| 1972 Sep. 17 | $14: 07: 15.3$ | Argostolion | 5.9 | 29 | 170 | 122 |
| 1972 Oct. 30 | $14: 32: 10.7$ | Argostolion | 5.4 | 28 | 110 | 90 |
| 1973 Nov. 4 | $15: 52: 12.6$ | Leukas | 5.9 | 20 | 180 | 175 |
| 1973 Nov. 4 | $16: 11: 38.7$ | Leukas | 4.9 | 20 | 80 | 87 |
| 1974 Jan. 29 | $15: 12: 44.8$ | Patra | 4.3 | 30 | 40 | 38 |
| 1975 Apr. 4 | $05: 16: 16.5$ | Patra | 5.7 | 56 | 58 | 48 |
| 1975 May 13 | $00: 22: 53.0$ | Xylokastron | 4.6 | 46 | 74 | 30 |
| 1975 Oct. 12 | $08: 23: 12.6$ | Corinth | 5.0 | 35 | 33 | 47 |

equation (2-59) in terms of maximum acceleration $A_{P T}$, with $70 \%$ probability of not being exceeded in a T year period. That is:

$$
\begin{equation*}
A_{P T}=A_{P 1}+\frac{\ell n T}{a}=u-\frac{\ell n[-\ell n(1-P)]}{a}+\frac{\ell n T}{a} \tag{7-3}
\end{equation*}
$$

where $A_{P 1}$ is the annual maximum acceleration which is exceeded with probability $P, u$ and a are the parameters of the first type asymptotic distribution using $P=0.30$ ( $1-P=0.70$ probability to be the largest acceleration value in a T year period). Table 7-10 tabulates $A_{P T}$ for $25,50,100$ and 200 year periods.

Using the same procedure, the maximum expected velocity and displacement for the same probability and periods is also tabulated in Table $7-10$. These values are derived using the equation of Orphal and Lahoud (1974):

$$
\begin{align*}
& \mathrm{V}=7.26 \cdot 10^{-1} \cdot 10^{0.52 \mathrm{~m}} \cdot \mathrm{R}^{-1.39} \mathrm{~cm} / \mathrm{sec} \\
& \mathrm{D}=4.71 \cdot 10^{-2} \cdot 10^{0.57 \mathrm{~m}} \cdot \mathrm{R}^{-1.18} \mathrm{~cm} \tag{7-5}
\end{align*}
$$

For these places, the magnitude $\mathrm{m}_{\mathrm{TP}}$ with the probability $P$ of being the maximum during a $T$ year period can also be derived from equation (2-20). That is:

$$
\begin{equation*}
P_{T}(m)=P_{1}^{T}(m)=\exp \left[-T\left(\frac{\omega-m}{\omega-u}\right)^{1} / \lambda\right] \tag{7-6}
\end{equation*}
$$

and

$$
\begin{equation*}
\mathrm{m}_{\mathrm{TP}}=\omega-(\omega-\mathrm{u})\left(-\ell \mathrm{nP} \mathrm{~T}_{\mathrm{T}}\right)^{\lambda} / \mathrm{T}^{\lambda} \tag{7-7}
\end{equation*}
$$

Then the return period $R P$ in years from equation (7-6) is

$$
\begin{equation*}
R P=\frac{1}{1-P^{1 / T}} \tag{7-8}
\end{equation*}
$$

which for $P=0.70$ and $T=25,50,100$, and 200 years corresponds to: $R P=70,140,280$, and 560 years. This means that a magnitude with $70 \%$ probability of not being exceeded in 25 years has a return period equal to 70 years. These values are also tabulated in Table 7-10.
7.4.4 Seismic risk evaluation - acceleration

From Table 7-10 it can be seen that the maximum acceleration for the short-term (25 years) risk and for the long-term (200 years) risk expected in the cities of Thessaloniki and Corinth is $A \simeq 120 \mathrm{~cm} / \mathrm{sec}^{2}$ and $180 \mathrm{~cm} / \mathrm{sec}^{2}$ respectively. The same cities also have the highest values for the expected velocities and displacements, although these values are not necessarily associated with the same earthquake.

The difference between the risk determined from the extreme magnitudes and accelerations at a particular place reflects the fact that in the attenuation models, the focal distance of each earthquake is taken into account. Thus, Athens and Corinth have almost the same seismic risk in terms of expected magnitude, but they differ significantly in terms of expected acceleration, velocity, and displacement. Because these two places are characterized by similar earthquake depth distribution and are near each other ( $\sim 50 \mathrm{~km}$ ), the difference shows that the seismic risk in the city of Athens is due to relatively more distant earthquakes than in the city of Corinth.

On the other hand, the cities of Heraklion and Rodos are characterized by intermediate depth earthquakes and have the lowest seismic risk in terms of expected maximum accelerations although the expected earthquakes may have large magnitudes.
7.5 Spatial distribution of seismic risk in Greece $\left[\mathrm{N}_{33}^{42.5}, \mathrm{E}_{19}^{29}\right]$
7.5.1 Contour maps of seismic risk - Procedure used

Close agreement is obtained between the observed and the predicted extreme magnitude occurrences for Greece as a whole, and for the six

Amplitudes which have $70 \%$ probability of not being exceeded in $T$ years

| Amplitude of: Magnitude* |  |  |  |  | Acceleration ( $\mathrm{cm} / \mathrm{sec}^{2}$ ) |  |  |  | Velocity ( $\mathrm{cm} / \mathrm{sec}$ ) |  |  |  | Displacement (cm) |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Period (years) | : 25 | 50 | 100 | 200 | 25 | 50 | 100 | 200 | 25 | 50 | 100 | 200 | 25 | 50 | 100 | 200 |
| Athens | $\begin{aligned} & 6.50 \\ & 0.16 \end{aligned}$ | $\begin{aligned} & 6.60 \\ & 0.21 \end{aligned}$ | $\begin{aligned} & 6.67 \\ & 0.25 \end{aligned}$ | $\begin{aligned} & 6.71 \\ & 0.30 \end{aligned}$ | 79.93 | 92.39 | 104.85 | 117.32 | 6.89 | 8.01 | 9.12 | 10.24 | 1.98 | 2.31 | 2.63 | 2.95 |
| Thessaloniki | $\begin{aligned} & 6.95 \\ & 0.25 \end{aligned}$ | $\begin{aligned} & 7.22 \\ & 0.32 \end{aligned}$ | $\begin{aligned} & 7.44 \\ & 0.35 \end{aligned}$ | $\begin{aligned} & 7.61 \\ & 0.43 \end{aligned}$ | 122.47 | 143.16 | 163.85 | 184.54 | 11.95 | 14.05 | 16.15 | 18.25 | 3.25 | 3.82 | 4.39 | 4.97 |
| Patra | $\begin{aligned} & 6.39 \\ & 0.18 \end{aligned}$ | $\begin{aligned} & 6.48 \\ & 0.25 \end{aligned}$ | $\begin{aligned} & 6.54 \\ & 0.28 \end{aligned}$ | $\begin{aligned} & 6.58 \\ & 0.31 \end{aligned}$ | 102.40 | 117.16 | 131.92 | 146.68 | 8.10 | 9.30 | 10.51 | 11.71 | 2.11 | 2.42 | 2.74 | 3.05 |
| Corinth | $\begin{aligned} & 6.57 \\ & 0.16 \end{aligned}$ | $\begin{aligned} & 6.64 \\ & 0.20 \end{aligned}$ | $\begin{aligned} & 6.68 \\ & 0.24 \end{aligned}$ | $\begin{aligned} & 6.70 \\ & 0.29 \end{aligned}$ | 117.87 | 136.27 | 154.67 | 173.07 | 10.21 | 11.88 | 13.54 | 15.20 | 2.62 | 3.04 | 3.46 | 3.89 |
| Heraklion | $\begin{aligned} & 6.66 \\ & 0.19 \end{aligned}$ | $\begin{array}{r} 6.88 \\ 0.23 \end{array}$ | $\begin{aligned} & 7.06 \\ & 0.27 \end{aligned}$ | $\begin{aligned} & 7.21 \\ & 0.35 \end{aligned}$ | 55.93 | 63.73 | 71.52 | 79.32 | 4.94 | 5.69 | 6.44 | 7.19 | 1.46 | 1.69 | 1.92 | 2.15 |
| Rodos | $\begin{aligned} & 6.65 \\ & 0.25 \end{aligned}$ | $\begin{aligned} & 6.94 \\ & 0.31 \end{aligned}$ | $\begin{aligned} & 7.19 \\ & 0.40 \end{aligned}$ | $\begin{aligned} & 7.42 \\ & 0.45 \end{aligned}$ | 63.88 | 73.15 | 82.41 | 91.68 | 6.38 | 7.42 | 8.46 | 9.50 | 1.90 | 2.22 | 2.54 | 2.86 |

* Magnitude values are those for distance within 100 km from the cities, derived from the third type asymptote.
selected places, using the Extreme-Value method. Thus, a more detailed evaluation of seismic risk in terms of maximum magnitudes and acceleration for every part of Greece may be attempted, and contour maps of the distribution of the seismic risk can be prepared.

Greece is divided into cells of $0.5^{\circ}$ Latitude $\times 0.5^{\circ}$ Longitude, and a mesh of grid points with spacing of $0.5^{\circ}$ Lat, $0.5^{\circ}$ Lon, is created for the whole area. All earthquakes occurring within a circle of $1^{\circ}$ radius, with its centre at a particular grid point, are then collected from the new earthquake catalogue, and their annual maximum observed magnitudes are analysed with the third type asymptotic distribution. With a similar computing procedure, but for radius of $2^{\circ}$, the first type asymptotic distribution is applied to the maximum accelerations. For every grid point the parameters and their uncertainties are computed. Each set of parameters corresponds to an area which overlaps the adjacent one by about $3 / 4$ or $7 / 8$ of the area, respectively. Using equations (2-61), (2-62), and (7-7), the annual and 80 year mode, and the maximum magnitude which is expected with $70 \%$ probability of not being exceeded in 50 and 100 years, are then estimated and contoured. Similarly, the parameters of the first type asymptote are used in equation (7-3) producing contour maps of seismic risk in terms of acceleration with $70 \%$ probability of not being exceeded in 50,100 , and 200 years.

For these calculations the computer program used in the calculations of Chapter IV is extended and listed in Appendix A. For the contours, the General Purpose Contouring Programme (GPCP) of Edinburgh Regional Computing Centre (ERCC) is used.

### 7.5.2 Results and discussion

a) Magnitude distribution

Figures $7-9$ to $7-12$ show the distribution of seismic risk in terms
of maximum expected magnitudes. The values in Figures $7-11$ and $7-12$

Fig 7-9 annual mode cont. for Greece


Fig 7-9 Most probable annual maximum earthquake magnitude (mode) for Greece.

Fig 7-10 80 year mode cont.

## for Greece



Fig 7-10 80 year most probable maximum earthquake magnitude for Greece.

Fig 7-11 magn. for 70\% prob. of not being exceeded in 50 years


Fig 7-11 Maximum expected earthquake magnitude with $70 \%$ probability of not being exceeded in 50 years for Greece.

Fig 7-12 magn. with $70 \%$ prob. of not being exceeded in 100 years


Fig 7-12 Maximum expected earthquake magnitude with $70 \%$ probability of not being exceeded in 100 years for Greece.
correspond to return periods of 140 and 280 years (eq 7-8), and these combined with Figures $7-9$ and $7-10$ allow the short-term (annual mode) and long-term risk in every part of the area to be evaluated.

Comparing these figures shows that the values of maximum annual magnitude are significantly different from the values for the 80 year mode, and the two overall pictures are also different in regional detail. The reason for this is that every place has its own distribution curvature for magnitude occurrence, which shows how quickly the distribution approaches the upper bound. Places in which the distribution of maximum magnitudes has greater curvature (larger value of $\lambda$ ), than other places, show greater difference between the regional contour maps of the annuai and 80 -year modes. As longer return periods are considered these differences become small, because the expected magnitudes are close to the upper bound and so the overall pictures are then similar. Figures 7-9 to 7-12 illustrate this point.

Figure 7-10 describes the most expected maximum magnitude during a period of 80 years. Because this period is almost equal to the time span of the data used (1901-1978), Figure 7-10 is the figure most easily compared with the observed distribution in Figure 6-8. The two figures are very similar.

The three main aseismic blocks discussed tectonically in paragraph 6.4.1 are clearly defined in Figure $7-10$ as the areas of the lowest seismic risk in Greece. Areas with high values of maximum expected magnitudes are also areas where the observed magnitudes have large values. It is expected that during the next 80 years earthquakes with magnitude $\geqslant 7.5$ will occur at least once in the area of the Greek-Yugoslavia borders, the Chalkidiki peninsula, and the Marmara Sea in north-western Turkey. These areas have experienced earthquakes with magnitudes $m=7.8(4.4 .1904), m=7.4(8.11 .1905)$, and $m=7.4$ (18.3.1953) respectively.

Figures $7-11$ and 7-12 which describe the risk in terms of maximum magnitude with $70 \%$ probability of not being exceeded in 50 and 100 years show a similar picture of seismicity. This is because these values correspond to return periods of about 140 and 280 years, and so they become close to their own upper bound with only small changes with increasing return periods. In both figures, an area with high risk, which in Figure $7-10$ was poorly defined, becomes clear. It is the area of the south-eastern part of the Peloponnesus and Kithera Island. This area has experienced the largest intermediate depth earthquake with magnitude $m=8.0$ (11.8.1903, $h=120 \mathrm{~km})$.

In the next 100 years there is a probability of $70 \%$ that the maximum expected earthquake magnitude will not be greater than $m=8.0$ in any of these high risk areas and consequently in the whole area of Greece.
b) Acceleration distribution

The seismic risk in terms of values of maximum acceleration with $70 \%$ probability to be the largest during the next 50,100 , and 200 years is described in Figures 7-13 to 7-15.

The overall picture of these three figures show an almost identical pattern. This is expected because it is apparent from equation (7-3) that the values of maximum acceleration, with a given probability of not being exceeded up to a certain time, increase as a linear function of the logarithm of time. So as the time increases, the acceleration value at a particular point increases proportionally to its previous value, and therefore the shape of the isovalue lines does not change.

Comparing the figures which show the risk in terms of maximum magnitude with those of maximum acceleration, the pictures differ significantly mainly in the places where intermediate earthquakes dominate. This is expected because in the strong motion attenuation formulae, the focal distance from the point of interest is an important factor in calculations of the

Fig 7-13 acceleration with $70 \%$ prob of not being exceeded in 50 years


Fig 7-13 Maximum expected acceleration with $70 \%$ probability of not being exceeded in 50 years for Greece.

Fig 7-14 acceleration with $70 \%$ prob of not being exceeded in 100 years


Fig 7-14 Maximum expected acceleration with $70 \%$ probability of not being exceeded in 100 years for Greece.

Fig 7-15 acceleration with $70 \%$ prob of not being exceeded in 200 years


Fig 7-15 Maximum expected acceleration with $70 \%$ probability of not being exceeded in 200 years for Greece.
final maximum acceleration observed at that point. Thus, from Figures 7-13 to 7-15 two new high risk areas are well defined. These are the areas around Cephalonia and Leukas Islands, and around Lesvos and the eastern Sporades Islands, with values of maximum accelerations at the $70 \%$ probability level. for the next 50 years approximately equal to $200 \mathrm{~cm} / \mathrm{sec}^{2}(0.2 \mathrm{~g})$. The areas where an acceleration of 0.3 g is expected to be the maximum acceleration at the $70 \%$ probability level in the next 200 years are the Marmara Sea area, and the area around Cephalonia and Leukas Islands.

As a part of the UNESCO Survey of the Seismicity of the Balkan region, maximum acceleration risk maps were prepared by a group of seismologists at the U.S. Geological Survey, Denver, Colorado Centre (Algermissen et al, 1976). These maps depict acceleration and velocity with $70 \%$ probability of not being exceeded in 25 and 200 year periods using data from the UNS's catalogue, and attenuation formulae derived from those of Schnabel and Seed (1973). Comparing the 200 year map for accelerations with Figure 7-15, the values of Algermissen et al's map are significantly larger than those of Figure 7-15. The main reason for the high values found in the previous work, for exampe $A=0.6 \mathrm{~g}$ around Cephalonia Island, seems to be the way the attenuation formulae were applied. They modified Schnabel and Seed's formula in such a way that it can be applied for two depth intervals, 15 and 110 km . Then all earthquakes with a depth less than 50 km were considered to have occurred at a depth of 15 km , and earthquakes with a depth more than 50 km were considered to have occurred at 110 km . However, the vast majority of earthquakes in Greece, as the recalculated depth parameter shows, have their origin at a shallow depth. 1492 out of a total of 1805 earthquakes have a depth less than 50 km , with an average depth of $28-35 \mathrm{~km}$. Thus, the above considerations may lead to serious overestimation of the values computed. Drakopoulos (1976b)also points out that Algermissen et al's values for the maximum acceleration appear to be relatively high for

Greece.

Considering the computing procedure followed here, that is every individual parameter is taken into account, and the agreement between the observed and calculated values of maximum accelerations, using equation (7-2), Figures $7-13$ to $7-15$ present an improved evaluation of the seismic risk in terms of maximum acceleration in Greece.
7.6 Summary

The objective of this Chapter was to evaluate seismic risk in Greece, and at six selected cities, in terms of both maximum expected magnitude and accelerations.

Assessing seismic risk in terms of maximum magnitudes is achieved by applying the method of strain energy release and the method of the third type asymptotic distribution. The results obtained show that where the energy release graphs include at least one well defined cycle of periodicity then the parameters of the third type asymptotic distribution have small uncertainties. From the figures presented, and the values of chi-square, it is shown that in all places analysed the magnitude distribution is well fitted by third type asymptotic behaviour.

From Table 7-6, where the predicted number of exceedances for different magnitudes is compared with those observed during the sample period (19011978), the close agreement also shows that the third type asymptotic distribution method, with the assumption made for missing years, reveals the seismic picture of a region.

The seismic risk for the selected cities is obtained in terms of both the return period for earthquakes with magnitude from $m=5.0$ to $m=8.0$, and also the number of expected exceedances of these magnitudes during the next 50 and 100 years.

The value of $m=8.73 \pm 0.65$ is found as an upper bound for earthquake magnitudes of Greece, and this is in agreement with the seismic history
of the area. An earthquake with magnitude equal to or greater than $m=5.5$ is expected as an annual extreme almost every year. An earthquake with magnitude equal to or greater than $m=7.5$ is expected to occur in Greece every 22 years, whereas the return period for an earthquake with magnitude greater than $m=8.0$, observed once in Greece (11.9.1903), is about 200 years.

For Athens the upper bound magnitude is $m=6.8 \pm 0.4$ within an area of 100 km radius from the city, whereas the most probable annual maximum magnitude is $m=4.51 \pm 0.08$.

Seismic risk is then obtained in terms of maximum accelerations. There is no universally accepted formula of strong motion attenuation, and the existing formulae show large dispersion in terms of maximum acceleration, and so an average formula is derived from the eight published formulae of Table 7-7. This formula agrees fairly well with the observed values of maximum acceleration recorded in Greece. An attempt to use the third type asymptotic method for maximum accelerations, however, results in poor convergence with values of $\lambda \approx 0.0$ and an unacceptable upper limit $\omega \approx 10 \mathrm{~g}$. A possible explanation for this may be the tendency of the observed maximum accelerations to cluster towards the two ends of the distribution as a result of the nature of the attenuation of the motion with focal distance. The first type asymptotic distribution appears to be a better representation of the distribution of the maximum accelerations.

Seismic risk in terms of maximum expected acceleration at a $70 \%$ probability level for the six places are tabulated in Table 7-8. The same table also contains the results for maximum velocities and displacement derived using the same procedure with equations (7-4) and (7-5) respectively. To complete the seismic picture, values of the maximum expected magnitude, at the same probability level, are also included in the table.

The same detailed evaluation is applied to the whole area of Greece by dividing it into cells of $0.5^{\circ}$ Lat $\times 0.5^{\circ}$ Lon. These methods are then
applied at each grid point for an area extending to $1^{\circ}$ radius from it, and the following values are calculated: the annual and 80 year mode, maximum magnitude at the $70 \%$ probability level of not being exceeded in 50 and 100 years, and maximum acceleration for the $70 \%$ probability level of not being exceeded in 50,100 and 200 years. These values are then contoured, and the maps derived (Fig 7-9 to 7-15) illustrate the spatial distribution of seismic risk within Greece.

A comnon feature in all these maps is the existence of three well defined aseismic blocks. These are: the atticocycladic block, the ptolemais basin block and the block formed by the north-eastern part of Greece. These aseismic blocks of low seismic risk correlate with the tectonics described in the previous chapter and indicate that the region can not be modelled by a simple tectonic plate. Also well defined are areas of high seismic risk which correlate with the tectonically most active areas, and these are:
i) Along the Hellenic arc: Greek-Albania-Yugoslavia borders, Leukas and Chephalonia Islands in the Ionian Sea, and the south-eastern end of the arc around Rodos Island.
ii) The western end of the Northern Anatolian fault in the Marmara Sea area and
iii) The northern Aegean Sea: Chalkidiki peninsula and Sporades Islands.

The objectives of this study were:
i) to examine the upper bound for earthquake magnitude occurrence,
ii) to investigate the usefulness of the third type asymptotic distribution for predicting earthquake risk, and
iii) to evaluate the seismicity and seismic risk of Greece.

The first two objectives are studied in the first part of this thesis and the third is achieved in the second part.

Reviewing the existing statistical models for earthquake occurrence and magnitude distribution shows that Extreme-Value methods have some important practical advantages compared to the methods which use the whole process. Because the sample consists of extreme values only, they are more complete, accurate and homogeneous than the entire range of events, and so Extreme-Value methods avoid some weak points of the othermethods, particularly those caused by the inclusion of low magnitudes.

Experimental data shows that the linear frequency-magnitude formula does not hold for very large earthquakes. However, most of the proposed alternative expressions fail to recognize the existence of an upper bound to the magnitude that can be generated in a given region, or do not include this upper bound as an unknown parameter in the resulting distribution function. When the problem of the upper bound for earthquake magnitude was examined, it was demonstrated using simple frequency-magnitude and energy-magnitude laws that it is possible to include such an upper bound as an unknown parameter, and to calculate its value both analytically and graphically. It was also shown that a finite upper bound to the earthquake magnitude is necessary to preserve a finite rate of energy release. From the analytical
method this upper bound is found to be a function of the mode $M_{1}$, the mean rate of annual energy release $M_{2}$, and the parameter $b$ of the frequencymagnitude law. Because each region has its own value of $b$ (Duda, 1965), each region must also have its own upper bound to the magnitude of earthquakes which can occur within the region.

The graphical and analytical methods of strain energy release are tested in the most seismically active region in the world, that is, the circum-Pacific belt. The results show remarkably good agreement for the values found for the upper bound from both methods. The empirically obtained relations between $M_{1}$ and $M_{2}$ show that the upper bound $M_{3}$, and the annual mean rate of energy release $M_{2}$, differ by one magnitude unit.

The advantage of the Extreme-Value methods, combined with the necessity for the existence of an upper bound to the earthquake magnitude occurrence, shows that the third type asymptotic distribution of extreme values, which includes the upper bound as a parameter, is a statistical model with which realistic predictions for large events can be obtained.

The second objective was then to investigate the applicability of the third type asymptotic distribution method. Because it is assumed that an error analysis is vital to any seismic risk analysis, a specific goal towards this objective was to develop a technique capable of computing the errors of the distribution parameters calculated and of all prediction quantities. A computer program using Marquardt's algorithm, written to do this, and the parameters with their error matrix obtained, reveal that:
i) The characteristic largest value $u$ is the most precisely known parameter.
ii) The variances for $\omega$ and $\lambda$ are usually large and the existence of negative covariance between them shows that they are not independent.
iii) When the data shows little curvature $\lambda$, and therefore a high value
for $\omega$, these parameters are usually accompanied by large uncertainties. However, this negative covariance leads to seismic risk calculations with acceptable uncertainties.

To explore the physical meaning of the parameters of the third type asymptote they are linked with the physical release of strain energy through $M_{1}, M_{2}$, and $M_{3}$. The circum-Pacific belt was again chosen as the testing area. Remarkably similar results were then obtained from these two different procedures, strain energy release and extreme values, showing that the third type asymptotic distribution is practical for prediction purposes. The first type asymptotic distribution was also applied to the same area and the values of chi-square show that in all cases the third type asymptote gives a better fit to the data, and so it is preferable as a general model for the statistical behaviour of the occurrence of maximum magnitude earthquakes.

Strain energy release and third type asymptotic distribution methods when applied to the circum-Pacific belt show that the centres of highest seismic activity are diagonally opposite each other. These are: South America and Japan, Kurile and Kamchatka. This presumably relates to the tectonic movement of the Pacific plate (Duda, 1965).

In the second part of this study the final objective was achieved by applying these methods to the relatively lower seismicity of Greece. Because the results of any statistical method depend on the quality of data used, a specific goal towards the evaluation of Greek seismic risk was the preparation of an earthquake catalogue as accurate and homogeneous as possible.

The Joint Epicentre Determination (JED) technique of calculating earthquake hypocentres was used to recalculate all the source parameters for the period 1917-1963. The recalculated positions are significantly different from those reported by the ISS with the average shift decreasing from decade to decade since 1917. The large average total shifts imply
that it is not possible to describe the detailed seismicity of Greece by simply using the ISS data for the whole period.

Homogeneity of the earthquake magnitudes was achieved by recomputing them according to a consistent scheme for the longest possible part of the period under consideration (1907-1978). The recomputation for the period 1907-1968 was made by using ground amplitudes for Greek earthquakes reported by the Swedish network and regression equations derived from parallel recordings of Swedish instruments. For 1968 onwards, regression equations were used to convert the body-wave magnitudes $M_{b}$ reported by ISC into surface wave magnitudes $M_{s}$ of the Swedish network.

When Stepp's test of completeness was applied, the results show that only earthquakes with magnitude $m \geqslant 6.3$ are completely reported during the whole period of investigation, whereas earthquakes with a maximum magnitude $4.2 \leqslant m \leqslant 4.7$ are completely reported only during the most recent 15 years. The time required for estimates of the mean recurrence rate to become stable was found to be 40-50 years of homogeneous observations for magnitudes $\geqslant$ 6.3 , whereas for earthquakes with magnitudes in the range $4.2 \leqslant m \leqslant 4.7$ only 5 to 10 are required.

The catalogue presented in this study has a high degree of homogeneity and accuracy permitting more detailed seismotectonic studies to be made on the basis of a long instrumentally recorded data sample. The existing tectonic models of Greece were examined and their accordance with the recalculated parameters of the earthquakes was tested. From the maps describing the spatial and depth distribution of the recalculated earthquakes it is clear that the area is tectonically more complicated than had previously been recognised. None of the proposed tectonic models explain the observed seismic activity sufficiently over the whole area. Places where further work is necessary to understand fully the present tectonic process are:
i) the north-western part of the Hellenic arc (north-western Greece and Albania),
ii) the south-eastern part of the Hellenic arc (eastern Crete, Karpathos and Rodos Islands), and
iii) the northern Aegean Sea and north-western part of Turkey.

Finally, within the framework of Greek local tectonics, the evaluation of seismic risk was achieved by applying the statistical techniques of the first part of this study to the new earthquake catalogue. The evaluation was made in terms of maximum expected magnitude and acceleration. To obtain the maximum magnitude distribution both the strain energy release and the third type asymptotic distribution methods were applied. However, the first type asymptotic distribution is better than the third for representing the observed distribution of maximum accelerations. A possible explanation for this may be found in the tendency for most maximum accelerations to lie within the linear part of the acceleration attenuation curve, because the attenuation of ground motion only varies non-1inearly for unusually short focal distances. For the seismic risk analysis derived in terms of acceleration an average formula for the attenuation of maximum acceleration was derived from most of the commonly used formulae. This formula gives values which agree with the observed maximum accelerations recorded from eight Greek earthquakes.

Greece as a whole and six heavily industrial and highly populated centres were first selected for a detailed evaluation of seismic risk. The risk values obtained are consistent with the observed values during the period of investigation. For Greece the upper bound for earthquake magnitude is $8.73 \pm 0.65$. The return period for an earthquake of magnitude equal or greater than the maximum observed $m=8.0$ is about 200 years. The most probable annual maximum magnitude is $m=6.41 \pm 0.04$. In an area with

100 km radius from the city of Athens, the upper bound is equal to $\mathrm{m}=6.8$ $\pm 0.4$, the most probable annual maximum $\tilde{m}=4.51 \pm 0.08$. All regions are well fitted by the third type asymptotic behaviour, and there is close agreement in detail between the predicted and observed number of exceedances of maximum magnitudes for Greece and Athens.

A detailed spatial seismic risk evaluation was obtained for the whole area of Greece by dividing it into cells of $0.5^{\circ}$ Lat. and $0.5^{\circ}$ Lon. The maximum expected magnitude and acceleration was evaluated at each grid point by applying the third and first type asymptotic distribution methods respectively. These results are presented as contour maps and using these the seismic risk at any locality can be evaluated. The common feature in all these maps is the existence of three well-defined aseismic blocks; these are:
i) the attikocycladic block,
ii) the block around the Ptolemais basin and,
iii) the block formed by the north-eastern part of Greece.

Also well defined are areas of high seismic risk; these are:
i) the Greek-Yugoslavia borders,
ii) the Chalkidiki peninsula and part of the north Aegean Sea,
iii) the north-western part of Turkey,
iv) the Cephalonia and Leukas Islands and,
v) the Eastern Sporades and Lesvos Islands area.

## APPENDIX A

COMPUTER PROGRAM LISTING

This program is written in Fortran IV. Its purpose is to compute the parameters, and their uncertainties, of the first and third type asymptotic distribution of extreme values of the following variables: magnitude, acceleration, velocity or displacement. Using these parameters it then estimates prediction parameters and calculates probability levels.

The computer procedure is
i) It creates a mesh of equally spaced grid points in the area of interest.
ii) It selects all the earthquakes within an area with specified radius from each grid point, finds the annual maximum magnitude, acceleration, velocity or displacement, and ranks them considering the first observed maximum as being the $j+1$ th where $j$ is the number of missing years.
iii) For each data set derived from the previous step, it applies the least-squares method and calculates the parameters and their uncertainties. For the first type asymptotic distribution (eq 2-20) the regression equation is:

$$
\begin{equation*}
m=u+\frac{1}{a} y \tag{A-1}
\end{equation*}
$$

and the parameters are computed using the linear least squares method (subroutine LINFIT). For the third type asymptotic distribution (eq 2-22) the regression equation is:

$$
\begin{equation*}
m=\omega-(\omega-u)[(-\ln (\phi(\mathrm{m})))]^{\lambda} \tag{A-2}
\end{equation*}
$$

where $\phi(m)$ is the plotting position of $m$ given by the equation (2-35). The parameters are then computed using Marquardt's (1963) algorithm
as described in paragraph 4.2.2 (subroutine CURFIT).
iv) With the parameters and uncertainties computed it calculates prediction parameters such as: annual mode (eq 2-56 or $2-61$ ), T year mode (eq 2-57 or 2-62), upper and lower bounds (eq 2-66 and $2-68$ ), extreme values with given probability $P$ of not being exceeded in T years (eq 2-59 or 7-7). The listing of the program starts with a comprehensive block of comment cards.

```
THTS PROGRAM COMPUTES TIGE PARAMETERS AND THEIR UNCERTAINTIES
OF THE FIRST AND THIRD TYPE ASYMPTOTIC DISTRIBUTION OF EXTREMES.
USTNG THESE PARANETERS AND THE ERROR MATRIX IT COMPUTES PREDICTION
PARAMETERS SUCH AS : ANNUAL MODE,T-YEARS MODE,NAXIMUM EXPECTED
HAGNITUDE OR ACCELERATION OF NOT BEING EXCEEDED IN T YEARS,AT A
STATED PROBABILITY LEVEL.IT ALSO COMPUTES UPPER AND LOWER BOUNDS
FOR MAXIMUM EXFECTED MAGNITUDES,ACCELERATIONS,VELOCITIES,OR
DISPLACENENTS FOR GIVEN PRUBABILITY LEVELS.
THE PARAMETERS OF THE THIRD TYPE ASYMPTOTE ARE COMPUTED USING
MARQURDT'S(1963) ALGORITHM,BY LINEARISING THE FITTING FUNCTION:
    M=W=(W-U)*((~ALOG(P(M))))** LAMDA SUBROUTINE CURVFIT.
STARTING WITH INITIAL TRIAL VALUES OF W,U,AND LAMDA (A(1),A(2),A(3))
THE GOOONESS OF FIT TO THE N OBSERVALS IS MESURED BY THE REDUCED
CHI-SQUARE WHICH IS MININISED WITH RESPECT TO EACH PARAMETER,
LEADING TO THE LINEAR MATRIX EQUATION:
    BETA=DELTAA(I)*ALFA
THE UNCERTAINTIES ON A(1),A(2),AND A(3) ARE THEN CALCULATED FRCM
DELTAA(I)=BETA*E
WHFRE E IS THE INVERSE MARTIX OF ALFA. E IS THE ERROR MATRIX AND
ITS ELEIIENTS ARE THE VARIANCE AND COVARIANCE OF THE PARAMETERS W,
U, AND LANDA.
THF PARAMETERS OF THE FIRST TYPE ASYMPTOTE ARE COMPUTED USING
LINEAR LEAST-SQUARES METHOD WITH THE FITTING FUNCTION:
M=U+1/A*Y
    NHERE
        Y=-ALOG(~ALOG(P(M))) : THE REDUCED VARIABLE.
DESCRIPTION OF THE INPUT DATA
DATA STREAM 5 (CARDS 1-3 AND EARTHQUAKES DATA FILE )
FIRST CARD: FORNAT(6F6.2)
EMIN : BOTTOM LATITUDE OF THE AREA OF INVESTIGATION
EHAX : TOP LATITUDE
OMIN: LEFT LONGITUDE
BMAX : RIGHT LONGITUDE
DER : STEP CF SHIFT IN LAT AND LON
SIZE : RADIUS IN DEGREES OF THE AREA FROM EACH GRID POINT
SECOND CARD: FCRNAT(TI5)
1) IDENT :
```

```
                =1 FCR YAXIMUM ACCELERATION DISTRIBUTION
                    =2 FOR MAXIMUM VELOCITY
                    =3 FOR NAXIMUM OISPLACEMENT
                    =4 FOR NAXIMUM MAGNITUDE
2) MCDE ;
            =-1 WEIGHT=1/X(I)
            =0 EQUAL WEIUHTS IN ALL THE INPUT DATA
            =1 WEIGHT=1/SIGMAX(I)
    3-4) MAXT-HINT :PERIOD OF INVESTIGATICN
    5-6) LIST1,LIST2 OPTIONS FOR PRINTOUT (O OR 1 FOR FULL PRINTOUT )
    7)
            INT:
                =1 FOR THE FIRST TYPE ASYMPTOTE
                    =2 FOR THE THIRD TYPE ASYMPTOTE
THIRD CARD FORMAT(2FG.2)
    EMMIN : FINIMUM MAGNITUDE TO BE CONCIDERED
    PROB : PROBABILITY LEVEL AT WHICH THE PREDICTION PARAMETER
                        IS EXPECTED TO BE EXCEEDED IN T YEARS
                            AFTER THESE CARDS THE EARTHQUAKE DATA FILE FOLLOWS.IT CONTAINS YEAR,
CRIGIN TIME,LAT,LON,DEPTH,AND MAGNITUDE.
DATA STREAM 3 (CARDS 1-3)
EACH OF THESE CARDS CONTAINS THE INITIAL VALUE AND THE ST.
DEVIATION OF THE PARAHETER A(I).
FIRST CARD:W : OR-DELTAW
UECOND CARD:U+OR=DELTAU
THIRD CARD:LAHDA+OR=DELTALAMDA
FORMAT(2F7.3)
    RFAD 101. EMIN,EMAX,BMIN,BMAX,DEE,SIZE
    READ 102, IDENT,MCDE,NINT,MAXT,LIST2,LIST1,INT
    RFAD 101, EMKMIN,PROB
        DIMENSION A (11,37,85),GY(85),YM(85),Y(85),RY(4),Q(6),TITLE(4)
    1,PRC(85),YML(85),A1(5),DELTAA1(5),SIGMAA1(5),YFIT(85),
    2 SIGMAB(10), DERIV (10), ARRAY(10,10),B(10),SIGMAG(10),WEITH(85),
    3 ALPHA(10,10),BETA(10),SIGYML(85),YT(10),YPD(10)
    PI=3.141592
    RAD=180./PI
    DFBY=1. /DES
    FIAMDA=0.C01
    i)TERMS =3
    i) }11\textrm{J}=1.NTERM
    READ(3,103) A^(J),DELTAA1(J)
103 FORNAT(2F7.3)
```

```
    11 CONTINHE
    AA1=A1(1)
    AA2=A1(2)
    AAB=A1(3)
    IA=(EMAX-EMIN)*DEBY+1.
    JA=(BMAX-B!IN)*DEBY+1.
    KA=MAXT-MINT+1
    PRINT ?
    PRINT 5000
    DO 10 I=1,IA
    DO 10 J=1,JA
    DO 10 K=1,KA
    10 A(1,J,K)=0.
    NO=0
    20 READ 200.IYEAR,TITLE,OZ,EN,BO,IDEPTH,EM
    IF(EM.EQ.O.) GO TO 9O
    IF(IYEAR.LT.MINT.OR.IYEAR.GT.MAXT) GO TO 20
    IF(EM.LT.EHNIN) GO TO 2O
    NO=NO+1
    DFPTH=IDEPTH
    ENL=EHAX
    DO 15 I=1, 1A
    E1=ENL-SIZE
    E2=ENL+SIZE
    IF(EN.GE.EZ) GO TO 15
    IF(EN.LT.E1) GO TO 15
    BOY=BMIN
    DO 25 J=1,JA
    81=8OY-SIZE
    BZ=BOY+SIZE
    IF(BO.LT.B1) GO TO 25
    IF(BO.GE.BZ) GO TO 25
            CALL OIRCOS(EN,BO,AE,BE,CE)
            CALL DIRCOS(ENL,BCY,AP,BP,CP)
    S=(AE-AP)**2+(BE-BP)**2+(CE-CP)** 2
    IF(S.GT.O.) GO TO 27
    DTST=1.0
    GO TO 28
    27S=1.-S*0.5
        S=SQRT(1.-S*S)/S
        DTST=ATAN(S)*RAD*111.11
        RS=SIZE*111.11
        IF(DIST.GT.RS) GC TO 25
    23R=SQRT(DIST*DIST+DEPTH*DEPTH)
        GO TO (21,22,23,24), IDENT
        THF FOLLOWING FORNULA USED FOR ACCELERATION,
        A=2164*EXP(EN*0.7)*(R+20)**-1.80 SEE TEXT EQUATION 7-2
    21 US=-1.80
    AMP=2164.*EXP(EM*0.7)*(R+20.)**US
    GO TO 26
C
```

```
C THE FOLLOWIAG FCRHULA USED FOR VELCCITY,
C
    22 US=-1.34
        AMP=0.726*10.**(0.52*EM)*(R**US)
        GO TO 26
c
C
    ... AND FCR OISPLACEMENT CALCULATION
    23Us=-1.18
        AMP}=0.0471*10.**(EMM* 0.57)*(R**US
        GO TO 2t
    24 AMP=EM
    26 K=IYEAR-MINT+1
        IF(AMP.GT,A(I,J,K)) A(I,J,K)=AMP
        IF(LIST1.EG.0) GC TO 25
C PRINT 2100,K,IYEAR,TITLE,O3,EN,BO,IDEPTH,EM,
    1DIST,R,ANP,A(I,J,K)
    25 BOY=BOY + DEB
    15 ENL=ENL-DEB
        GO TO 20
    9O PRINT 900, NO
        GO TO (110,120,130,140), IDENT
    110 PRINT 1500
        GO TO 150
    120 PRINT }160
        GO TO 150
    130 PRINT 1700
        GO TO 150
    140 PRINT 1800
    150 CONTINUE
        PRINT 250, RS
        PRINT }
        RYEAR=25.
        PROB=ALOG(-ALOG(1.-PROB))
        OO 55 I=1.4
        RY(I)=ALOG(RYEAR)
    55 RYEAR=RYEAR*2.
        ENL=EMAX
        DO 50 I=1,IA
        B\capY=BMIN
        DO 60 J=1.JA
        PRINT 400, ENL,BOY
        L=0
        DO }70\textrm{K}=1,\textrm{KA
        IF(A(I,J,K).EQ.O.) GOTO 70
        L=L+1
    7\cup }YM(K)=A(I,J,K
        IF(L.LT.17) GO TO 60
        SK=L+1
        LL=SK-1
        IF(LISTZ.GT.O) PRINT 800
            CALL RANK (K,YM,Y)
        H=KA-L
```

```
        L=0
        Dत 95 K=1,KA
        LYEAR=MINT+K-1
        IF(A(I,J,K).EQ.O.) GC TO 95
        L=L+1
        SJ=(MAXT-MINT+1)-SK+1+L-0.44
        PRO1=SJ/(NAXT-HINT+1.12)
        GY(L)=-ALCG(-ALCG(PRO1))
        M=M+1
        PRO(L)=PRO1
        YML1=Y(M)
        YM(L)=YML1
        YML(L)=YML1
        GYK=GY(L)
        IF(LIST2.EQ.O) GO TO 95
        PRINT 950, K,LYEAR,A(I,J,K),YML1,GYK,PRO1,L
        G5 CONTINUE
        PRINT 350, L,MINT,LYEAR
        DO 56 II =1,LL
        SIGYML(II) =0.3
    56 CONTINUE
        GO TO (12,13).INT
C
    12 CALL LINFIT(GY,YM,SIGYML,L,MODE,ALFA,SIGMAA,BHTA,SIGMAB,R)
        YMOD=ALFA
        YP=ALFA=(PROB*BHTA)
        PRINT 2
        KYEAR=1
        PRINT 1300, YMOD,YP,KYEAR
        KYEAR=25
        DO }75\textrm{N}=1,
        RB=RY(iN)*BHTA
        YT(N)=YMOD+RB
        YPO(N)=YP+RB
            PRINT 1400, YT,YPD,KYEAR
        75 KYEAR=KYEAR*2
        WRITE(8,131) ENL,BOY,YNOD,(YT(N),N=1,4)
    131 FORMAT(7(1X,F7.2))
        WRITE(9.131)ENL,BOY,YP,(YPD(N),N=1,4)
        PRINT 2
        PRINT 1
        GO TO 60
    13 CALL CUPFIT(PRO,YML,SIGYML,LL,NTERMS,MODE,A1,DELTAA1,SIGMAA1,
        1FLAMDA,YFIT,CHISQR,SIGMAB,SIGMAG)
    14 Z=CHISQR
        CALL CURFIT(PRO,YML,SIGYML,LL,NTERMS,MODE,AY,DELTAA1,SIGMAA1,
        1FIAMDA,YFIT,CHISQR,SIGHAB,SIGMAG)
        U=Z-CHISQR
        IF(U.GT.0.01) GO TO 14
        WRITE(6.31)CHISQR
    31 FORMAT(1H,'FINAL CHISQR=',F10.5)
C
    PRINT PARALSETERS AND ST.DEVIATIONS
```

```
                WRITE(6,16)A1(1),SIGMAA1(1). OF W=1,F7.4)
            WRITE(0,17)A1(2),SIGMAA1(2)
    17 FORMAT(1H,'U=',F7.4,3X,'SD.OF U=',F7.4)
C
c
C
C
c
C
C
C
C
    WRITE(6,18)A1(3),SIGMAA1(3)
    18 FORMAT(1H,'L=',F7.4,3X,'SD.OF L=',F7.4)
        DO 19 JJ=1.3
        WRITE(6,32)JJ,SIGMAB(JJ),JJ,SIGMAG(JJ)
    32 FORMAT(1H,'CCV2',I1,'=',F7.4,2x,'COV1',I1,'=',F7.4)
    19 CONTINUE
        C=0.05
    ANNUAL MCDE
        ZM1=A1(1)-(A1 (1) -A1(2))*(1.-A1(3))**A1(3)
    N-YEAR MODE
        ZMN=A1(1)-(A1(1)-A1(2))*((1.-A1(3))/100.)**A1(3)
        UPPER BOUND FGR MAXIMUM EXPECTED MAGNITUDE IN N YEARS
        ZUN=A1(1)-(A1(1)-A1(2))*((-1./100.)*ALOG(1.- C/2.))***A1(3)
        Magnitude WITH 70% probability to be the maximum anNual magnitude
        ZM170=A1(1)-(A1(1)-A1(2))*((-ALOG(.70))**A1(3))
    MAGNITUDE WITH 7O%PROB.TO BE THE MAXIMUM IN THE NEXT N YEARS
        ZMN70=A1(1)-(A1(1)-Z11170)/(100.**A1(3))
        WRITE(6,527)ZM1,ZMN,ZUN
    527 FORMAT(1H,'ANNUAL MODE=',F6. 2. 2X,'100 YEAR MODE=',
    1F6.2,2X,'UP.BOUND OF 1OO YEAR MODE WITH 95 CON.LEV.=',FG.2)
        WRITE (7,523)ENL,BOY,ZN1,ZM17OO,ZMN,ZUN,ZMN7O,YNL1,L
    528 FORMAT (8F8.2.16)
        A1(1)=AA1
        A1 (2) = AAL2
        A1(3)=AA3
    to BOY=BOY+DEB
    50 ENL=ENL=DEG
        1 FORMAT(1H1)
    2 FORNAT(1H)
    101 FORMAT(6F6.2)
    102 FORMAT(715)
200 FORMAT(1X,I4,4A4,F4.1,2F8.2,I5,9X,F3.1)
    250 FORMAT(5X,'SIZE OF EARTHQUAKE SOURCE REGION'.
        1F8.2,' KHS.',1)
    300 FORNAT(5X,2I3,15,3X,213,F6.1,F6.2,2(3X,F5.2,F6.2),
        1I5,F6.1)
```

```
350 FORMAT(/19X,'NUMBER, OF CBSERVED,SHOCKS',15,1.19x,
    1'RETWEEN',IT,' - ',I5,' YEARS')
    19x,'X INTENSITY')
55u FORMAT(6(2x,F8.1),' RETURN P.')
600 FORMAT(GF10.2)
650 FORMAT(6F10.7,' SHOCKS/Y.')
800 FORMAT(13X,'K YEAR',5X,'AMP.',5X,'RANKED'.6X,'G(Y)',
    15X,'PROB.',3X,'L')
900 FORMAT(5X,'NUMBER OF PRCCESSED EARTHQUAKE DATA',I5)
950 FORMAT(10X,I4,I5,4F10.3,14)
1300 FORMAT(/16X,2(2X,F9.2),15,' YEAR')
1400 FORMAT( 16X,2(2X,F9.2),I5,' YEARS')
1500 FORMAT(5X,'RISK ANALYSIS BASED ON THE ACCELERATION VALUES
    1 IN CH./SEC.**2')
1600 FORMAT'(5X,'RISK ANALYSIS BASED ON THE VELOCITY VALUES IN'
    1' CM./SEC.')
1700 FORMAT(5X,'RISK ANALYSIS BASED ON THE DISPLACEMENT VALUES',
    1' IN CM.')
1800 FORMAT(5X,'RISK ANALYSIS BASED ON THE MAGNITUDE VALUES')
2100 FORMAT(1X,I2,I4,4A4,F4.1,2F8.2,15,F5.1,2F7.1,2F7.2)
5000 FORMAT'/gX,'RISK ANALYSIS IN A GIVEN REGION BASED ON THE'/,
    15X.'GUMBEL"S STATISTICAL THEORY OF EXTREME VALUES, '/)
        STOP
        END
```

SUBROUTINE DIRCOS (RLA,RLO,A,B,C)
$\mathrm{P}_{\mathrm{T}}=3.141592$
RAD $=\mathrm{PI} / 180$.
$E N=R L A * R A D$
$B \cap=R L O * R A D$
$E N=\operatorname{ATAN}(0.99238 * \operatorname{SIN}(E N) / C O S(E N))$
$C=\operatorname{SIN}(E N)$
$X=-\cos \left(E_{N}\right)$
$D=\operatorname{SIN}(B C)$
$E=-\operatorname{Cos}(B O)$
$A=X * E$
$B=-D * X$
RETURIN
END
SUBROUTINE RANK (N,Y,X)
DIMENSION X $(85), Y(85)$
YMAX $=1 . E 38$
$X 1=Y M A X$
DO $10 \quad J=1, N$
YMIN $=1 . E 37$
UO $20 \quad \mathrm{I}=1$, iv
IF(Y(I).GE.YMIN) GO TC 20
YHIN=Y(I)
$K=1$
$\operatorname{IF}(Y(I) . G T . X 1)$ GOTC 20

```
    YMIN=Y(I)
    20 CONTINUE
    X(J)=YMIN
    X1=YMIN
    10 Y(K)=YMAX
    RETURN
    END
c
            SUBROUTINE CURVFIT (X,Y,SIGMAY,NPTS, NTERNS,MODE,A,DELTAA,
            SIGMAA,FLANOA,YFIT, CHISQR)
    PURPOSE
        MAKE A LEAST-SQUARES FIT TO A NON LINEAR FUNCTION
            WITH A LINEARISATION OF THE FITTING FUNCTION
        SUBROUTINE CURFIT (X,Y,SIGMAY,NPTS,NTERMS,MODE,A,DELTAA,
    1 SIGMAA,FLAMDA,YFIT,CHISQR,SIGMAB,SIGMAG)
    DOUBLE PRECISION ARRAY
    DTMENSION X (100),Y(100),SIGMAY(100),A(10),DELTAA(10),SIGMAA(10),
    1 YFIT(100),SIGMAB(10),SIGMAG(10)
    DIMENSION WEIGHT(100),ALPHA(10,10),BETA(10), DERIV(10),
    1 ARRAY(10,10),B(10)
    11 NFREE=NPTS-NTERMS
    IF(NFREE) 13,13,20
    13 CHISQR=0.
    GO TO 110
            eVAlUATE WEIGHTS
    20 DO 30 I=1,NPTS
    21 IF (MODE) 22.27.29
    22 IF (Y(I)) 25.27.23
    23 WEIGHT(I)=1./Y(I)
    GO TO 30
    25 WFIGHT(I)=1./(-Y(I))
    GO TO 30
    27 WEIGHT(I)=1.
    GO TO 30
    29 WEIGHT(I)=1./SIGMAY(I)**2
    3u CONTINUE
C
c
C
    evaluate alpha and Beta matrices
    31 DO 34 J=1.NTERMS
    BETA(J)=0.
    DO 34 K=1,J
34 ALPHA}(J,K)=C
41 DO 50 I=1,NPTS
    CALL FDERIV (X,I,A,DELTAA,NTERMS,DERIV)
    DO 46 J=1, iNTERNS
    BFTA(J)=BETA(J) +WEIGHT(I) *(Y(I)-FUNCTN(X,I,A)) \starDERIV(J)
    DO 46 K=1.J
40́ ALPHA}(J,K)=ALPHA(J,K) +WEIGHT(I)*DERIV(J)*DERIV(K
50 CONTINUE
51 ON 53 J=1, NTERMS
    OO 53 k=1,J
```

```
    53 ALPHA(K,J)=ALPHA(J,K)
C
    EVALUATE CHI SQUARE AT STARTING POINT
C
        61 DO 62 I=1,NPTS
        6 2 ~ Y F I T ( I ) = F U N C T N ( X , I , A )
        63 CHISQ1=FCHISQ(Y,SIGMAY,NPTS,NFREE,MODE,YFIT)
C WRITE(6,66) CHISQ1
    66 FORMAT(1H,'CHISQ1=',F10.5)
C
    INVERT MODIFIED CURVATURE MATRIX TO FIND NEW PARAMETERS
C
        71 00 74 J=1.NTERMS
            DO }73\textrm{K}=1\mathrm{ , NTERHS
        73 ARRAY(J,K)=ALPHA(J,K)/SQRT(ALPHA(J,J)*ALPHA(K,K))
        74 ARRAY (J,J)=1.+FLAMDA
        BU CALL MATINV (ARRAY,NTERNS,DET)
        81 DO 84 J=1,NTERMS
            B(J)=A(J)
            ON 84 K=1.NTERMS
        84 B(J)=8(J) +BETA(K) *ARRAY(J,K)/SQRT(ALPHA(J,J)*ALPHA(K,K);
C
C
        91 DO G2 I=1.NPTS
        92 YFIT(I) =FUNCTN(X,I,B)
        93 CHISQR=FCHISQ(Y,SIGMAY,NPTS,NFREE,MODE,YFIT)
C WRITE(6.999) CHISQR
    999 FORMAT(1H ,'CHISQR=',F10.5)
C WRITE(6,998) A(1),A(2),A(3)
    998 FORMAT(1H,'W=',F7.4,3X,'U=',F7.4,3X,'L=',F7.4)
        IF (CHISQ1-CHISQR) 95.101.101
        95 FLAMDA=10.*FLAMDA
        GO TO 71
c
C EVALUATE PARAMETERS AND UNCERTAINTIES
C
    101 DO 103 J=1,NTERHS
        A(J)=B(J)
        STGMAB}(J)=ARRAY(J,2)/SQRT(ALPHA(J,J)*ALPHA(2,2))
        SIGMAG(J)=ARRAY(J,1)/SQRT(ALPHA}(J,J)*ALPHA(1,1)
    103 SIGMAA(J)= OSQRT(ARRAY(J,J)/ALPHA(J,J))
        FLAMDA=FLAMDA/10.
    110 RFTURN
        END
c
    FUNCTION FUNCTN (FOR THE THIRD ASYMPTOTIC DISTRIBUTION)
c FUNCTN(X,I,A)=W-(W-U)*Z(I)**L
C
        FUNCTION FUNCTN(X,I,A)
        DIMENSION X(100),A(10)
        XT = X(I)
```

```
        Z1=-ALOG(XI)
        Z2=Z1**A(3)
        Z.3=(A(1)-A(2))*Z2
        FUNCTN=A(1)=Z3
    20 RFTURN
        END
c
c
c PURPOSE
c
    EVALUATE REDUCED CHI SQUARE FOR FIT TO DATA
        FCHISQ=SUM( (Y-YFIT)**2/SIGMA**2)/NFREE
        FUNCTION FCHISQ(Y,SIGMAY,NPTS,NFREE,MODE,YFIT)
        DOUBLE PRECISION CHISQ,WEIGHT
        DIMENSION Y(100),SIGMAY(100), YFIT(100)
    11 CHISQ=O.
    1 2 ~ I F ~ ( N F R E E ) ~ 1 3 , 1 3 , 2 0
    13 FCHISQ=0.
        GO.TO40
C
C ACCUMULATE CHI SQUARE
C
    20 00 30 I=1,NPTS
    21 IF (MODE) 22,27,29
    22 IF(Y(I)) 25.27.23
    23 WEIGHT=1./Y(I)
        GO TO 30
    25 WEIGHT=1./(-Y(I))
        GO TO 3O
    27 WEIGHT=1.
        GO TO 30
    29 WEIGHT=1./SIGMAY(I) ** 2
    30 CHISQ=CHISQ+WEIGHT*(Y(I)-YFIT(I))**2
    31 FREE=NFREE
    32 FCHISQ=CHISQ/FREE
    4 0 ~ R E T U R I N ~
        END
        SUBRQUTINE FDERIV: ANALYTICAL:
    PURPOSE
        evaluate derivatives of FUNCTION FOR LEAST - SQUARES SEARCH
        FOR ARBITRARY FUNCTION GIVEN BY FUNCTN
        SUBROUTINE FDERIV (X,I,A,DELTAA,NTERMS,DERIV)
        DIMENSION X (100),A(10),DELTAA(10),DERIV(10)
        XI=x(I)
        Z1=-ALOG(XI)
        z2=21**A(3)
        DERIV(1)=1.-z2
        DFRIV(2)=22
```

```
    DERIV(3)=-(A(1)-A(2))*Z2*ALOG(Z1)
    RETURN
    END
C
C
C
C
    INVERT A SYMMETRIC MATRIX AND CALCULATE ITS DETERMINANT
        SUBROUTINE MATINV (ARRAY, NORDER,DET)
    DOUBLE PRECISION ARRAY,AMAX,SAVE
    DIMENSION ARRAY(10,10),IK(10),JK(10)
    DET=1.
    DO 100 K=1, NORDER
C
C
    AMAX=0.
    21 DO 30 I=K,NCRDER
        DO 3O J=K, NORDER
    23 IF (DABS (AMAX)=DABS(ARRAY(I,J))) 24.24.30
    24 AMAX=ARRAY(I,J)
        IK(K)=I
        Jk(k)=J
    3O CONTINUE
C
    INTERCHANGE ROWS AND COLUMNS TO PUT AMAX IN ARRAY(K,K)
    31 IF (AMAX) 41.32.41
    32 DFT=0.
        GO TO 140
    41I=IK(K)
        IF (I-K) 21.51.43
    43 DO 5O J=1,NORDER
        SAVE=ARRAY(K,J)
        ARRAY(K,J)=ARRAY(I,J)
    SO ARRAY(I,J)=-SAVE
    51 J=JK(K)
        IF (J-K) 21.61,53
    53 DO 60 I=1,NORDER
        SAVE=ARRAY(I,K)
        ARRAY(I,K)=ARRAY(II,J)
    GO ARRAY(I,J)=-SAVE
C
C
    61 DO 70 I=1,NORDER
        IF (I=K) 63.70.63
    63 ARRAY(I,K)=пARRAY(I,K)/AHAX
    7O CONTINUE
    71 DO 80 I=1,NORDER
        DO 80 J=1,NCRDER
        IF (I-K) 74,80,74
    74 1F (J=K) 75,80.75
```

```
    75 ARRAY(I,J) #ARRAY(I,J) +ARRAY(I,K)*ARRAY(K,J)
    80 CONTINUE
    81 DO 90 J=1,NORDER
    IF (J-K) 83.90,83
    83 ARRAY(K,J) =ARRAY (K,J)/AMAX
    90 CONTINUE
    ARRAY (K,K)=1./AMAX
    100 DET=DET*AMAX
c
C RESTORE ORDERING OF MATRIX
    101 DO 130 L=1,NORDER
        K=NORDER=L+1
        J=IK(K)
        IF (J-K) 111,111,105
    105 DO 110 I=1,NORDER
        SAVE=ARRAY(I,K)
        ARRAY(I,K)=-ARRAY(I ,J)
    110 ARRAY(I,J)=SAVE
    111 I =JK(K)
    IF (I-K) 130.130.113
    113 DO 120 J=1,NORDER
    SAVE=ARRAY(K ,J)
    ARRAY(K,J) =-ARRAY(I,J)
    120 ARRAY(I,J)=SAVE
    130 CONTINUE
    140 RETURN
    END
C
    SUBROUTINE LINFIT(X,Y,SIGMAY,NPTS,MODE,A,SIGMAA,B,SIGMAB,R)
    2OUBLE PRECISION SUM,SUMX,SUMY,SUMXZ,SUMXY,SUNYZ
    DOUBLE PRECISION XI,YI,WEIGHT,DELTA,VARNCE
    DIMENSION X(100),Y(100),SIGMAY(100)
C
    11 SUM=0.
    SUMX=0.
    SUMY=0.
    SUMX2=0.
    SUMXY=0.
    SUMY2=0.
    21 00 50 I=1.NPTS
    XI=X(I)
    YT=Y(I)
    IF(MODE) 31.36.38
    31 IF(YI) 34.36.32
    32 WFIGHT=1./YI
    GO TO 41
    34 WFIGHT=1./(-YI)
    GO TO 41
    30 WEIGHT=1.
    GO TO 41
    38 WEIGHT=1./SIGMAY(I) **2
```

```
    41 SUM=SUHP WEIGHT
    SUMX=SUMX WWEIGHT*XI
    SUMY=SUMY +WEIGHT*YI
    SUMX2=SUMXZ +WEIGHT*XI*XI
    SUMXY=SUMXY +NEIGHT*XI*YI
    SUMYZ=SUMYZ +WEIGHT*YI*YI
SO CONTINUE
C
C CALCULATE COEFFICIENTS AND ST. DEVIATIONS
C
    5 1 \text { DELTA=SUM*SUIFX 2*SUMX*SUMX}
    A=(SUMX2*SUMY-SUMX*SUMXY)/DELTA
    53 B=(SUMXY*SUM-SUMX*SUMY)/DELTA
    61 IF(MODE) 62,64,62
    6 2 ~ V A R N C E = 1 . ~
        GO TO 67
    6 4 C = N P T S - 2
        VARNCE=(SUMYZ + A*A*SUM+B*B*SUMX2-2.*(A*SUMY*B*SUMXY=A*B*SUMX))/C
    67 SIGMAA=DSQRT(VARNCE*SUMX2/DELTA)
    6 3 \text { SIGMAB=OSQRT (VARNCE*SUN/DELTA)}
    71R=(SUM*SUMXY=SUMX*SUMY)/DSQRT(DELTA*(SUM*SUMYZ=SUMY*SUMY))
        WRITE(6,100)A,SIGMAA,B,SIGMAB,R
        FORMAT(1H,'U='F7.4,2X,'S.D.OF U=',F6.4,'1/A=',F7.4,
        12x,'S.D.OF 1/A=',F6.4,30X,'R=',F10.5)
        RETURN
        END
C
C
C EXAMPLE OF INPUT DATA
C
    -39.95-45.28-19.00-28.00-00.50--1.00
    ----1----1-1900-1970-----1----1----1
    --4.20--0.30
    -1920-APR 15---09-20-37.3---39.99---20.25--123----------5.7
    STREAM 3
    ---8.55---0.80
    -=-7.30---0.08
    -=-0.20=--0.04
```

$C$
$C$

APPENDIX B

Earthquake catalogue for Greece since 1901

The geographic region studied is $33.0^{\circ} \mathrm{N}$ to $42.5^{\circ} \mathrm{N}$ and $19.0^{\circ} \mathrm{E}$ to 29. $0^{\circ} \mathrm{E}$. The parameters listed for each earthquake are: i) Date, ii) Origin time, iii) Latitude, iv) Longitude, v) focal depth, vi) Number of reported stations, vii) Surface-wave magnitude, and viii) Shift of recalculated epicentres from the ISS locations in distance (km) and in azimuth (degrees). The catalogue presented here contains:
i) Earthquakes for which all parameters are recalculated (1917-1963). For these earthquakes, the focal depth error is estimated to be less than $\pm 10 \mathrm{~km}$ for $h<50 \mathrm{~km}$ and $\pm 15 \mathrm{~km}$ for $h \geqslant 50 \mathrm{~km}$.
ii) Earthquakes which do not permit relocations for the period 19011963. For these earthquakes, instead of the number of reported stations, the appropriate reference is given.
iii) Earthquakes for which relocations are not attempted (1964-1968). For these earthquakes all parameters, except magnitude, are those given by ISC.

Since 1907 the surface-wave magnitude for all earthquakes in the catalogue are determined using the Swedish network ground amplitude records and the conversion formulae $(5-4),(5-7)$, and (5-9). The standard deviation of these magnitudes may, in general, be estimated as around $\pm 0.3$ units.

The abbreviations used for references are:

UNS: Earthquake catalogue of Shebalin et al, 1974

ATB: Earthquake catalogue of Alsan et al, 1975, and
ROT: Earthquake catalogue of Rothé, 1969.

Complete details of the method by which this catalogue was produced are given in Chapter $V$.

|  | DATE |  | ORIG.TIME |  | LAT | LON | DEPTH | OBS |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | ---: | MAG


| DATE |  | ORIG.TIME | LAT | LON | DEPTH | OBS | MAG | SHIFT |  |  |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- |
|  |  | GMT | N | E | KM |  | MS | DIST KM AZIM DEG |  |  |
| 19 | AUG | 10 | O9 | 23 | 00 | 40.6 | 27.1 | 15 | UNS | 6.5 |



| DATE |  |  | $\begin{gathered} \text { ORIG.TIME } \\ \text { GMT } \end{gathered}$ |  |  | $\begin{aligned} & \text { LAT } \\ & \mathrm{N} \end{aligned}$ | $\begin{gathered} \text { LON } \\ \text { E } \end{gathered}$ | $\begin{aligned} & \text { DEPTH } \\ & \text { KM } \end{aligned}$ | OBS | $\begin{gathered} \text { MAG } \\ \text { MS } \end{gathered}$ | DIST ${ }_{\text {K }}$ | HIFT |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1926 | MAR | 18 |  | 52 | 52.6 | 35.86 | 30.05 | 61 | 36 | 5.2 | 103.25 | 66.90 |
| 1926 |  | 19 | 00 | 28 | 30.0 | 36.23 | 29.93 | 25 | 30 | 5.0 | 116.78 | 45.77 |
| 1926 |  | 24 |  | 04 | 42.9 | 35.90 | 28.97 | 90 | 24 | 5.5 | 44.75 | 356.51 |
| 1926 |  | 31 |  | 06 | 45 | 36.0 | 29.0 | 15 | UNS | 5.0 |  |  |
| 1926 | APR | 22 |  | 11 | 54.1 | 35.99 | 29.23 | 140 | ATB | 5.2 |  |  |
| 1926 | JUN | 10 | 19 | 16 | 24.4 | 38.84 | 21.20 | 76 | 25 | 5.1 | 119.38 | 288.85 |
| 1926 |  | 26 | 19 | 46 | 42.1 | 36.75 | 26.98 | 109 | 96 | 7.3 | 123.79 | 312.58 |
| 1926 | JUL | 05 | 09 | 21 | 57.3 | 36.55 | 26.56 | 175 | 19 | 5.1 | 134.66 | 218.71 |
| 1926 | AUG | 18 |  | 05 | 02.3 | 38.08 | 20.93 | 56 | 42 | 5.5 | 39.10 | 76.63 |
| 1926 |  | 30 |  | 38 | 04.5 | 36.76 | 23.16 | 26 | 75 | 7.0 | 83.50 | 170.13 |
| 1926 | SEP | 19 | 01 | 04 | 01.9 | 36.09 | 22.08 | 71 | 49 | 5.9 | 66.07 | 6.28 |
| 1926 | OCT | 23 |  | 58 | 55.1 | 41.00 | 20.00 | 15 | 30 | 5.2 | 111.19 | 0.00 |
| 1926 | DEC | 17 |  | 31 | 11.1 | 41.26 | 20.01 | 20 | 39 | 5.7 | 51.87 | 55.84 |
| 1926 |  | 17 |  | 39 | 58.2 | 41.11 | 19.79 | 15 | 44 | 5.8 | 27.65 | 63.29 |
| 1927 | MAR | 24 | 14 | 46 | 47.5 | 35.45 | 26.39 |  | 35 | 5.7 | 61.47 | 35.31 |
| 1927 | JUN | 30 | 22 | 59 | 49.6 | 39.33 | 20.81 | 2 | 46 | 5.7 | 109.28 | 289.98 |
| 1927 | JUL | 01 |  | 19 | 01.0 | 36.72 | 22.85 | 45 | 68 | 6.5 | 87.71 | 188.80 |
| 1927 |  | 28 |  | 49 | 57.2 | 40.33 | 20.12 | 70 | 12 | 5.0 | 137.06 | 326.90 |
| 1927 | AUG | 07 |  | 33 | 50 | 42.4 | 19.5 | 6 | UNS | 5.0 |  |  |
| 1928 | Jan | 22 | 00 | 18 | 26.0 | 38.83 | 22.60 | 12 | 13 | 5.1 | 37.97 | 13.33 |
| 1928 | MAR | 31 |  | 29 | 47.7 | 38.01 | 27.92 | 12 | 68 | 7.0 | 55.04 | 187.36 |
| 1928 |  | 31 | 05 | 12 | 37.7 | 39.49 | 27.74 | 10 | ATB | 5.2 |  |  |
| 1928 | APR | 10 | 01 | 03 | 18 | 37.4 | 26.1 | 15 | UNS | 5.1 |  |  |
| 1928 |  | 14 | 08 | 59 | 58.0 | 42.34 | 26.02 | 19 | 88 | 7.0 | 74.95 | 342.03 |
| 1928 |  | 14 |  | 23 | 47.1 | 42.29 | 26.05 | 21 | 18 | 5.5 | 68.89 | 342.54 |
| 1928 |  | 18 |  | 22 | 51.2 | 42.27 | 25.35 | 7 | 95 | 7.1 | 101.19 | 309.12 |
| 1928 |  | 18 |  | 40 | 56 | 42.2 | 25.1 | 45 | UNS | 5.6 |  |  |
| 1928 |  | 18 |  | 05 | 45 | 42.0 | 26.0 | 36 | UNS | 5.5 |  |  |
| 1928 |  | 18 |  | 14 | 53.1 | 42.27 | 25.52 | 12 | 39 | 5.7 | 90.65 | 314.66 |
| 1928 |  | 22 | 19 | 59 | 29.4 | 38.40 | 23.34 | 28 | 22 | 5.4 | 46.81 | 342.53 |
| 1928 |  | 22 | 20 | 13 | 55.9 | 38.08 | 23.12 | 8 | 56 | 6.5 | 34.85 | 285.02 |
| 1928 |  | 25 | 09 | 25 | 54.4 | 42.54 | 26.23 | 12 | 41 | 5.9 | 93.64 | 356.47 |
| 1928 |  | 28 |  | 59 | 05.5 | 42.00 | 25.27 | 10 | 27 | 5.6 | 91.94 | 291.63 |
| 1928 |  | 29 |  | 49 | 20.7 | 37.71 | 23.08 | 84 | 28 | 5.5 | 34.22 | 198.20 |
| 1928 | MAY | 26 |  | 55 | 30.2 | 39.84 | 19.80 | 42 | 18 | 5.0 | 25.08 | 223.97 |
| 1928 | JUL | 15 |  | 33 | 60.7 | 37.91 | 27.57 | 13 | 49 | 5.8 | 26.21 | 112.75 |
| 1928 | DEC | 10 |  | 03 | 07.5 | 36.32 | 24.59 | 110 | 39 | 5.5 | 180.33 | 228.78 |
| 1929 | JAN | 17 | 00 | 06 | 40 | 40.6 | 19.6 | 15 | UNS | 5.1 |  |  |
| 1929 |  | 23 | 11 | 14 | 28.5 | 35.20 | 24.52 | 28 | 42 | 5.2 | 56.47 | 322.09 |
| 1929 | MAR | 27 | 07 | 41 | 46.5 | 36.63 | 26.68 | 106 | 25 | 5.1 | 25.37 | 139.50 |
| 1929 |  | 27 |  | 061 | 10 | 35.0 | 20.0 | 15 | UNS | 5.0 |  |  |
| 1929 | APR | 17 |  | 48 | 18.5 | 36.55 | 24.43 | 11 | 14 | 5.1 | 147.76 | 200.22 |
| 1929 | NOV | 11 |  | 35 | 02.7 | 36.68 | 26.21 | 15 | 15 | 5.0 | 29.51 | 242.88 |
| 1929 | DEC | 20 |  | 19 | 34 | 40.2 | 23.8 |  | UNS | 5.1 |  |  |
| 1930 | JAN | 15 | 23 | 58 | 22.9 | 36.93 | 28.25 | 45 | 11 | 5.0 | 224.64 | 17.31 |
| 1930 |  | 23 | 10 | 53 | 58.6 | 35.44 | 27.31 | 52 | 16 | 5.0 | 52.01 | 340.54 |
| 1930 | FEB | 14 |  | 38 | 18.6 | 35.96 | 24.71 | 91 | 81 | 6.2 | 26.94 | 260.45 |
| 1930 |  | 23 | 18 | 19 | 20.7 | 39.86 | 22.75 | 70 | 65 | 6.1 | 98.01 | 347.37 |
| 1930 | MAR | 06 | 08 | 21 | 47.0 | 34.78 | 26.31 | 101 | 29 | 5.5 | 32.55 | 345.15 |
| 1930 |  | 06 |  | 18 | 34.2 | 35.03 | 24.73 | 87 | 50 | 5.7 | 57.82 | 205.31 |
| 1930 |  | 31 | 12 | 33 | 51.4 | 39.70 | 23.34 | 10 | 66 | 6.1 | 65.84 | 239.68 |
| 1930 | APR | 17 | 20 | 06 | 49.2 | 37.80 | 23.17 | 66 | 70 | 6.1 | 77.89 | 358.05 |
| 1930 | AUG | 05 | 23 | 23 | 08.6 | 34.79 | 26.71 | 38 | 23 | 5.0 | 91.74 | 343.16 |
| 1930 | SEP | 13 | 20 | 06 | 03.2 | 37.85 | 22.77 | 183 | 14 | 5.3 | 91.66 | 335.57 |
| 1930 | NOV | 21 | 02 | 00 | 29.5 | 40.28 | 19.64 | 42 | 71 | 6.1 | 33.65 | 20.95 |


| date |  |  | $\underset{\text { GMT }}{\substack{\text { ORIG.TIME } \\ \text { ( }}}$ |  | $\begin{aligned} & \text { LAT } \\ & \mathrm{N} \end{aligned}$ | $\begin{gathered} \text { LON } \\ \text { E } \end{gathered}$ | $\begin{aligned} & \text { DEPTH } \\ & \text { KM } \end{aligned}$ | OBS | $\begin{aligned} & \text { MAG } \\ & \text { MS } \end{aligned}$ | SHIFT |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1930 | Nov | 21 | 192 | 2600 | 40.2 | 19.6 | 16 | UNS | 5.4 |  |  |
| 1931 | Jan | 04 | 00 | 0052.5 | 38.22 | 23.27 | 8 | 52 | 5.7 | 25.55 | 14.09 |
| 1931 |  | 11 | 191 | 1943 | 40.2 | 19.9 | 8 | UNS | 5.0 |  |  |
| 1931 |  | 28 | 05 | 5513.6 | 40.89 | 20.60 | 6 | 61 | 5.6 | 88.29 | 5.49 |
| 1931 | MAR | 07 | 001 | 1650.3 | 41.50 | 22.48 | 38 | 69 | 6.2 | 55.82 | 358.28 |
| 1931 |  | 08 | 015 | 5020.3 | 41.44 | 22.61 | 6 | 118 | 6.8 | 50.02 | 10.65 |
| 1931 | APR | 20 | 20 | 3340 | 35.0 | 27.0 | 15 | UNS | 5.1 |  |  |
| 1931 |  | 26 | 06 | 2455 | 38.5 | 26.2 | 10 | UNS | 5.1 |  |  |
| 1931 | JUN | 30 | 102 | 2401.4 | 36.29 | 22.87 | 103 | 34 | 5.4 | 26.49 | 206.63 |
| 1931 | JUL | 12 | 22 | 2438.9 | 39.72 | 24.83 | 38 | 36 | 5.2 | 103.53 | 284.03 |
| 1931 | AUG | 18 | 094 | 4710 | 40.8 | 23.5 | 15 | UNS | 5.1 |  |  |
| 1931 | SEP | 11 | 16 | 2322.7 | 38.87 | 23.29 | 77 | 31 | 5.0 | 154.28 | 9.40 |
| 1931 |  | 23 | 13 | 2816.2 | 39.99 | 19.94 | 6 | 29 | 5.0 | 226.28 | 347.79 |
| 1931 | NOV | 23 | 23 | 3213.0 | 36.99 | 21.28 | 64 | 21 | 5.1 | 58.03 | 340.20 |
| 1932 | MAR | 09 | 10 | 1652.3 | 38.23 | 20.62 | 11 | 53 | 5.4 | 28.08 | 22.36 |
| 1932 | MAY | 14 | 03 | 4506.8 | 35.88 | 28.65 | 78 | 35 | 5.2 | 14.59 | 99.26 |
| 1932 | JUN | 12 | 23 | 2424.3 | 36.43 | 25.19 | 122 | 14 | 5.0 | 240.71 | 333.19 |
| 1932 |  | 29 | 02 | 3022.3 | 36.35 | 26.72 | 85 | 33 | 5.3 | 123.43 | 320.16 |
| 1932 |  | 29 | 18 | 3345.2 | 35.53 | 26.70 | 155 | 24 | 5.6 | 81.86 | 272.60 |
| 1932 | AUG | 09 |  | 4448.2 | 36.71 | 27.73 | 110 | 18 | 5.1 | 246.15 | 4.79 |
| 1932 |  | 15 | 04 | 3440.1 | 39.10 | 22.17 | 51 | 41 | 5.7 | 19.05 | 52.91 |
| 1932 | SEP | 26 | 19 | 2043.0 | 40.39 | 23.81 | 05 | 134 | 7.1 | 65.69 | 0.74 |
| 1932 |  | 26 | 21 | 2702.1 | 40.75 | 23.80 | 35 | 58 | 5.8 | 105.63 | 0.00 |
| 1932 |  | 28 | 16 | 5212.6 | 40.64 | 23.31 | 16 | 57 | 5.8 | 27.03 | 1.82 |
| 1932 |  | 29 | 03 | 5724.4 | 40.83 | 23.46 | 25 | 87 | 6.4 | 49.87 | 15.78 |
| 1932 |  | 30 | 06 | 1219.3 | 35.94 | 22.60 | 43 | 50 | 5.5 | 12.05 | 233.60 |
| 1932 | ОСт | 09 | 06 | 2456.8 | 40.00 | 23.45 | 81 | 18 | 5.2 | 37.52 | 306.69 |
| 1932 |  | 23 | 13 | 3644.7 | 35.51 | 27.24 | 21 | 44 | 5.5 | 33.14 | 272.05 |
| 1932 | NOV | 01 | 16 | 1933.5 | 40.55 | 23.37 | 57 | 53 | 5.5 | 18.14 | 19.59 |
| 1932 | DEC | 07 | 07 | 5552.9 | 37.37 | 27.60 | 83 | 12 | 5.0 | 64.09 | 7.97 |
| 1933 | FEB | 25 | 23 | 1953.8 | 34.04 | 22.15 | 68 | 22 | 5.2 | 29.51 | 232.49 |
| 1933 | MAR | 14 | 01 | 1955.1 | 38.84 | 25.18 | 54 | 53 | 5.4 | 50.11 | 347.96 |
| 1933 |  | 22 | 18 | 1441.0 | 38.06 | 20.45 | 59 | 25 | 5.1 | 7.98 | 326.73 |
| 1933 | APR | 23 | 05 | 5741.8 | 36.76 | 27.17 | 44 | 117 | 6.7 | 30.24 | 261.53 |
| 1933 |  | 28 | 22 | 2852.4 | 35.09 | 27.10 | 64 | 37 | 5.4 | 14.42 | 317.60 |
| 1933 | MAY | 08 | 01 | 1350.7 | 40.65 | 23.17 | 74 | 13 | 5.0 | 53.02 | 301.89 |
| 1933 |  | 11 |  | 0948.5 | 40.76 | 23.67 | 16 | 100 | 6.5 | 40.38 | 356.37 |
| 1933 |  | 15 | 20 | 0137.9 | 36.35 | 26.80 | 10 | ATB | 5.2 |  |  |
| 1933 |  | 31 | 19 | 5550.1 | 40.48 | 23.47 | 98 | 18 | 5.0 | 17.60 | 58.31 |
| 1933 | JUN | 01 | 02 | 4041.5 | 40.68 | 23.83 | 10 | 25 | 5.1 | 221.76 | 7.23 |
| 1933 | JUL | 02 | 12 | 1949.8 | 40.43 | 22.87 | 156 | 16 | 5.2 | 325.56 | 358.06 |
| 1933 |  | 09 | 21 | 4247.1 | 36.95 | 20.43 | 83 | 27 | 5.1 | 8.34 | 228.21 |
| 1933 | AUG | 17 | 06 | 2442.9 | 37.32 | 28.90 | 60 | 16 | 5.0 | 39.96 | 26.51 |
| 1933 | SEP | 24 | 13 | 2125.7 | 35.65 | 28.62 | 145 | 15 | 5.2 | 94.08 | 79.50 |
| 1934 | FEB | 04 | 09 | 3525.6 | 41.54 | 19.42 | 10 | 54 | 5.6 | 19.17 | 32.77 |
| 1934 |  | 21 | 00 | 4026.9 | 34.77 | 22.73 | 153 | 24 | 5.1 | 70.32 | 25.52 |
| 1934 |  | 21 | 11 | 3728.3 | 34.60 | 22.29 | 74 | 55 | 5.7 | 45.83 | 347.19 |
| 1934 | NOV | 09 | 13 | 4103.3 | 36.47 | 25.41 | 132 | 64 | 6.3 | 79.10 | 222.04 |
| 1934 |  | 21 | 22 | 2630.0 | 33.67 | 25.89 | 120 | 25 | 5.3 | 41.33 | 358.71 |
| 1935 | JAN | 04 | 14 | 4131.3 | 40.76 | 27.53 | 13 | 84 | 6.6 | 84.60 | 1.72 |
| 1935 |  | 04 | 15 | 1854.6 | 40.12 | 27.65 | 12 | 12 | 5.5 | 57.57 | 158.19 |
| 1935 |  | 04 | 16 | 2008.3 | 40.69 | 27.54 | 18 | 81 | 6.6 | 76.83 | 2.53 |
| 1935 | FEB | 18 |  | 4009.8 | 40.33 | 23.64 | 8 | 28 | 5.2 | 30.24 | 104.95 |
| 1935 |  | 25 | 02 | 5130.5 | 36.07 | 24.83 | 67 | 108 | 7.1 | 17.87 | 296.93 |
| 1935 | MAR | 18 |  | 4047.2 | 36.0 | 27.30 | 83 | 60 | 6 | 70.01 | 22.76 |


| DATE |  |  | ORIG.TIME GMT |  |  | $\begin{gathered} \text { LAT } \\ \mathrm{N} \end{gathered}$ | $\begin{gathered} \text { LON } \\ \text { E } \end{gathered}$ | $\begin{gathered} \text { DEPTH } \\ \text { KM } \end{gathered}$ | OBS | $\begin{array}{r} \text { MAG } \\ \text { MS } \end{array}$ | SHIFT |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1935 | MAR | 31 | 03 | 21 | 35.4 | 41.18 | 19.88 | 40 | 55 | 5.7 | 44.87 | 281.67 |
| 1935 | AUG | 20 | 08 | 53 | 49.0 | 34.81 | 26.92 | 47 | 22 | 5.0 | 23.89 | 4.49 |
| 1935 | SEP | 03 | 17 | 35 | 39.3 | 39.45 | 20.67 | 106 | 47 | 5.4 | 161.74 | 5.19 |
| 1935 | OCT | 22 | 07 | 29 | 43.3 | 40.13 | 27.18 | 34 | 26 | 5.0 | 15.55 | 353.26 |
| 1935 | NOV | 07 | 04 | 37 | 34.9 | 40.82 | 20.00 | 24 | 43 | 5.4 | 90.64 | 332.22 |
| 1936 | JAN | 14 | 15 | 11 | 25.8 | 36.10 | 22.30 | 74 | 27 | 5.2 | 90.86 | 232.55 |
| 1936 |  | 29 | 15 | 55 | 41.0 | 42.02 | 20.10 | 47 | 24 | 5.1 | 54.86 | 310.73 |
| 1936 | FEB | 12 | 10 | 57 | 20.8 | 33.94 | 23.53 | 14 | 29 | 5.2 | 79.43 | 157.36 |
| 1936 | APR | 08 | 04 | 17 | 09.0 | 40.66 | 23.09 | 15 | 30 | 5.2 | 93.74 | 175.34 |
| 1936 |  | 15 | 16 | 03 | 04.0 | 37.63 | 20.38 | 25 | 18 | 5.0 | 42.71 | 194.47 |
| 1936 |  | 28 | 23 | 15 | 29.3 | 36.15 | 26.35 | 115 | 25 | 5.1 | 23.58 | 256.22 |
| 1936 | AUG | 08 | 04 | 12 | 57.0 | 34.43 | 26.28 | 103 | 50 | 5.5 | 130.26 | 15.57 |
| 1936 | OCT | 24 | 14 | 06 | 15.0 | 36.04 | 22.54 | 61 | 37 | 5.2 | 15.71 | 287.16 |
| 1937 | JAN | 02 | 14 | 04 | 02.1 | 34.04 | 24.95 | 39 | 43 | 5.2 | 8.11 | 145.37 |
| 1937 | MAY | 23 | 10 | 57 | 28.8 | 38.74 | 27.54 | 8 | 40 | 5.6 | 82.40 | 2.42 |
| 1937 | DEC | 16 | 17 | 35 | 36.9 | 36.09 | 23.75 | 51 | 61 | 5.5 | 66.07 | 48.77 |
| 1938 | JAN | 02 | 10 | 54 | 44 | 35.4 | 26.0 | 14 | UNS | 5.2 |  |  |
| 1938 |  | 16 | 13 | 36 | 35.0 | 35.34 | 27.92 | 131 | 21 | 5.2 | 31.65 | 328.44 |
| 1938 | FEB | 10 | 20 | 37 | 57.4 | 34.59 | 26.15 | 25 | 52 | 5.5 | 24.19 | 191.14 |
| 1938 | MAR | 11 | 14 | 51 | 06.1 | 39.23 | 20.52 | 55 | 54 | 5.3 | 48.49 | 351.77 |
| 1938 |  | 13 | 17 | 45 | 24.2 | 38.62 | 20.64 | 14 | 59 | 5.8 | 20.87 | 170.11 |
| 1938 | MAY | 12 | 22 | 09 | 43.9 | 35.15 | 26.24 | 23 | 58 | 5.8 | 39.29 | 5.36 |
| 1938 | JUN | 03 | 16 | 37 | 56.7 | 34.49 | 26.48 | 52 | 24 | 5.0 | 66.07 | 226.48 |
| 1938 | JUL | 02 | 12 | 26 | 45.3 | 40.48 | 27.79 | 6 | 12 | 5.0 | 165.43 | 353.81 |
| 1938 |  | 20 | 00 | 23 | 42.5 | 38.30 | 23.66 | 42 | 81 | 6.1 | 13.20 | 270.05 |
| 1938 |  | 27 | 01 | 29 | 18.7 | 38.32 | 23.79 | 44 | 19 | 5.0 | 2.39 | 338.60 |
| 1938 | AUG | 15 | 11 | 02 | 12.3 | 40.27 | 20.34 | 52 | 32 | 5.0 | 41.97 | 43.92 |
| 1938 | SEP | 18 | 03 | 50 | 40.9 | 38.27 | 22.47 | 53 | 72 | 5.9 | 116.52 | 268.77 |
| 1938 | DEC | 26 | 22 | 02 | 30.3 | 37.53 | 20.92 | 105 | 40 | 5.3 | 38.42 | 47.88 |
| 1939 | JAN | 02 | 04 | 36 | 17.9 | 39.74 | 27.86 | 100 | UNS | 5.7 |  |  |
| 1939 | MAY | 20 | 09 | 35 | 30.4 | 40.95 | 19.69 | 60 | 38 | 5.1 | 37.13 | 116.79 |
| 1939 |  | 31 | 00 | 24 | 05 | 38.0 | 22.0 | 14 | UNS | 5.0 |  |  |
| 1939 | JUN | 02 | 14 | 11 | 43.0 | 38.65 | 22.09 | 148 | 18 | 5.2 | 94.51 | 359.47 |
| 1939 | JUL | 28 | 10 | 12 | 53.0 | 35.47 | 25.21 | 54 | 22 | 5.0 | 131.29 | 8.35 |
| 1939 |  | 28 | 16 | 06 | 10.6 | 35.06 | 25.23 | 10 | 14 | 5.0 | 87.02 | 13.97 |
| 1939 | AUG | 09 | 03 | 30 | 34.1 | 40.52 | 19.38 | 8 | 20 | 5.1 | 66.39 | 29.12 |
| 1939 | SEP | 20 | 00 | 19 | 33.6 | 38.02 | 20.86 | 50 | 61 | 5.6 | 13.38 | 280.28 |
| 1939 |  | 22 | 00 | 36 | 34.2 | 38.78 | 26.73 | 05 | 81 | 7.0 | 28.93 | 211.18 |
| 1940 | JAN | 06 | 19 | 04 | 39.7 | 35.34 | 25.53 | 55 | 43 | 5.5 | 52.38 | 220.15 |
| 1940 | FEB | 01 | 06 | 20 | 06.0 | 41.31 | 24.43 | 15 | 14 | 5.2 | 164.30 | 56.25 |
| 1940 |  | 23 | 00 | 40 | 04.3 | 40.59 | 19.43 | 52 | 33 | 5.4 | 81.63 | 323.69 |
| 1940 |  | 29 | 16 | 07 | 47.5 | 34.84 | 25.48 | 43 | 77 | 6.1 | 102.89 | 201.95 |
| 1941 | JAN | 09 | 18 | 13 | 35.4 | 38.15 | 27.29 | 54 | 26 | 5.5 | 25.27 | 312.16 |
| 1941 | MAR | 01 | 03 | 52 | 55.2 | 39.73 | 22.46 | 25 | 61 | 6.3 | 26.12 | 352.35 |
| 1941 | MAY | 14 | 08 | 36 | 29.4 | 39.74 | 22.64 | 35 | 44 | 5.3 | 53.29 | 22.83 |
| 1941 |  | 16 | 01 | 27 | 58.2 | 39.62 | 22.41 | 34 | 35 | 5.1 | 35.87 | 1.38 |
| 1941 |  | 23 | 19 | 51 | 59.5 | 37.15 | 28.14 | 37 | 58 | 6.2 | 15.86 | 248.72 |
| 1941 |  | 23 | 22 | 34 | 17.9 | 37.18 | 28.29 | 47 | 40 | 5.7 | 2.39 | 201.71 |
| 1941 |  | 23 | 23 | 00 | 47.8 | 37.22 | 28.35 | 48 | 12 | 5.0 | 4.95 | 63.34 |
| 1941 | JUN | 23 | 08 | 00 | 38.5 | 37.95 | 27.81 | 10 | ATB | 5.1 |  |  |
| 1941 |  | 24 | 15 | 15 | 58 | 40.5 | 21.0 | 15 | UNS | 5.0 |  |  |
| 1941 | JUL | 13 | 15 | 39 | 42.7 | 38.25 | 26.46 | 69 | 49 | 6.2 | 92.71 | 38.61 |
| 1941 | SEP | 01 | 14 | 18 | 50.5 | 42.17 | 24.72 | 25 | 14 | 5.1 | 66.07 | 16.01 |
| 1941 |  | 21 | 22 | 40 | 31.1 | 37.50 | 28.29 | 70 | ATB | 5.3 |  |  |
| 1941 | NOV | 21 | 12 | 12 | 24 | 39.7 | 23.8 | 15 | UNS | 5.0 |  |  |


| DATE |  |  | ORIG.TIME GMT |  | $\begin{gathered} \text { LAT } \\ \mathrm{N} \end{gathered}$ | $\begin{gathered} \text { LON } \\ \mathrm{E} \end{gathered}$ | $\begin{aligned} & \text { DEPTH } \\ & \text { KM } \end{aligned}$ | OBS | $\begin{array}{r} \text { MAG } \\ \text { MS } \end{array}$ | DIST S | HIFT |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 194 | DEC | 13 | 061 | 1607.0 | 37.23 | 27.99 | 28 | 34 | 6.1 | 28.08 | 276.99 |
| 1942 | FEB | 05 | 011 | 1601.1 | 36.87 | 27.94 | 36 | 23 | 5.1 | 236.57 | 181.30 |
| 1942 | MAY | 09 | 043 | 3717.3 | 36.11 | 26.30 | 137 | 20 | 5.0 | 58.28 | 38.32 |
| 1942 |  | 21 | 034 | 4233.8 | 36.98 | 20.18 | 21 | 27 | 5.0 | 51.73 | 226.55 |
| 1942 | JUN | 01 | 090 | 0118.0 | 38.99 | 22.11 | 65 | 24 | 5.3 | 42.82 | 216.16 |
| 1942 |  | 01 | 091 | 1745.0 | 38.98 | 22.56 | 68 | 30 | 5.6 | 38.35 | 158.67 |
| 1942 |  | 16 | 044 | 4741.8 | 34.40 | 26.29 | 41 | 37 | 5.5 | 69.46 | 343.82 |
| 1942 |  | 16 | 054 | 4235.3 | 40.82 | 27.70 | 18 | 41 | 6.0 | 53.34 | 331.53 |
| 1942 |  | 21 | 043 | 3844.2 | 36.05 | 26.96 | 88 | 41 | 5.3 | 55.69 | 225.66 |
| 1942 | AUG | 12 | 203 | 3847.4 | 39.22 | 27.79 | 55 | 23 | 5.1 | 30.72 | 323.43 |
| 1942 |  | 27 | 061 | 1416.7 | 41.59 | 20.45 | 12 | 45 | 6.0 | 4.30 | 255.04 |
| 1942 | SEP | 01 | 094 | 4216.1 | 35.19 | 26.73 | 22 | 45 | 6.0 | 147.38 | 204.48 |
| 1942 | OCT | 28 |  | 3153.6 | 39.48 | 28.11 | 12 | 13 | 5.4 | 75.08 | 44.47 |
| 1942 |  | 28 |  | 2252.7 | 39.27 | 27.87 | 37 | 26 | 6.0 | 44.16 | 46.75 |
| 1942 |  | 28 | 024 | 4158.0 | 39.48 | 27.75 | 43 | 15 | 5.5 | 57.74 | 21.97 |
| 1942 | NOV | 15 | 170 | 0122.9 | 39.55 | 28.58 | 10 | ATB | 6.1 |  |  |
| 1943 | JAN | 07 | 111 | 1446.9 | 37.92 | 20.55 | 90 | 29 | 5.3 | 69.11 | 356.34 |
| 1943 |  | 07 | 22 | 3607.6 | 37.52 | 21.32 | 81 | 26 | 5.0 | 68.44 | 68.82 |
| 1943 |  | 08 | 235 | 5643.7 | 40.92 | 28.10 | 20 | ATB | 5.0 |  |  |
| 1943 |  | 11 |  | 5620.4 | 36.55 | 27.26 | 26 | 16 | 5.3 | 117.58 | 232.43 |
| 1943 | FEB | 14 |  | 2829.3 | 38.22 | 20.01 | 16 | 35 | 5.7 | 114.68 | 333.20 |
| 1943 | MAR | 25 | 025 | 5106.2 | 40.41 | 21.89 | 259 | 23 | 5.5 | 278.47 | 15.75 |
| 1943 | APR | 09 | 194 | 4649.6 | 34.55 | 28.01 | 8 | 23 | 5.0 | 61.20 | 0.86 |
| 1943 | MAY | 22 | 22 | 0552.7 | 38.36 | 20.43 | 76 | 27 | 5.2 | 64.09 | 308.85 |
| 1943 | JUN | 14 |  | 4703.2 | 38.76 | 20.42 | 78 | 21 | 5.2 | 98.52 | 329.18 |
| 1943 |  | 27 | 10 | 0542.4 | 35.14 | 24.26 | 32 | 31 | 5.1 | 85.11 | 345.09 |
| 1943 | JUL | 22 |  | 0930.2 | 38.84 | 20.39 | 15 | 35 | 5.3 | 19.30 | 283.76 |
| 1943 | OCT | 16 | 130 | 0857.5 | 36.31 | 27.89 | 95 | 63 | 5.8 | 10.99 | 185.14 |
| 1943 | NOV | 15 |  | 4308.9 | 36.81 | 28.84 | 83 | 14 | 5.5 | 11.43 | 160.33 |
| 1943 |  | 20 |  | 0159.4 | 36.55 | 28.36 | 35 | 33 | 5.5 | 55.47 | 225.47 |
| 1944 | JAN | 05 | 050 | 0503 | 36.4 | 27.4 | 150 | UNS | 5.1 |  |  |
| 1944 |  | 05 | 074 | 4414.1 | 36.61 | 27.61 | 69 | 32 | 5.6 | 30.32 | 38.85 |
| 1944 | MAY | 27 |  | 5235.7 | 36.22 | 27.19 | 91 | 38 | 5.7 | 147.27 | 222.70 |
| 1944 | JUN | 25 | 041 | 1629.3 | 38.74 | 29.00 | 69 | 51 | 6.2 | 31.87 | 235.81 |
| 1944 |  | 25 | 065 | 5753.2 | 38.97 | 29.55 | 57 | 26 | 5.5 | 22.54 | 261.34 |
| 1944 | JUL | 20 | 103 | 3730.9 | 36.06 | 27.02 | 53 | 23 | 5.4 | 78.11 | 36.96 |
| 1944 |  | 30 |  | 0045.6 | 37.14 | 22.27 | 85 | 35 | 5.6 | 53.20 | 337.30 |
| 1944 | AUG | 09 |  | 3638.2 | 35.84 | 27.07 | 137 | 19 | 5.6 | 64.13 | 53.67 |
| 1944 |  | 17 | 132 | 2815.1 | 35.67 | 26.80 | 98 | 17 | 5.4 | 33.43 | 55.16 |
| 1944 | OCT | 06 | 023 | 3448.5 | 39.46 | 26.43 | 26 | 76 | 7.0 | 24.59 | 286.08 |
| 1944 |  | 06 | 072 | 2826.2 | 39.37 | 26.06 | 40 | ATB | 5.1 |  |  |
| 1944 |  | 07 | 21 | 3428.6 | 39.40 | 26.49 | 16 | 27 | 5.5 | 18.53 | 270.07 |
| 1945 | JAN | 08 | 22 | 4223.3 | 39.17 | 20.47 | 53 | 25 | 5.4 | 42.88 | 344.71 |
| 1945 | AUG | 27 | 162 | 2656.7 | 36.13 | 26.61 | 162 | 21 | 5.0 | 60.88 | 64.93 |
| 1945 | SEP | 02 | 115 | 5404.6 | 34.43 | 28.61 | 62 | 68 | 6.4 | 27.30 | 277.20 |
| 1945 |  | 12 | 16 | 2934.4 | 40.10 | 19.81 | 133 | 20 | 5.1 | 20.28 | 304.46 |
| 1946 | APR | 05 |  | 5407.0 | 35.29 | 23.65 | 40 | 52 | 5.6 | 31.42 | 47.13 |
| 1946 |  | 12 | 07 | 3702.7 | 36.72 | 26.97 | 78 | 34 | 5.5 | 141.93 | 17.22 |
| 1946 |  | 16 | 11 | 4356.5 | 41.22 | 19.86 | 39 | 48 | 5.4 | 4.12 | 303.59 |
| 1946 | JUL | 16 | 05 | 2634.5 | 34.20 | 25.65 | 17 | 61 | 6.0 | 55.12 | 35.99 |
| 1946 | AUG | 20 | 172 | 2642.1 | 40.77 | 19.82 | 48 | 32 | 5.1 | 48.44 | 188.05 |
| 1946 | OCT | 13 | 21 | 2442.6 | 34.28 | 25.81 | 15 | 42 | 5.2 | 83.18 | 310.08 |
| 1946 | NOV | 21 | 01 | 4338.9 | 38.96 | 20.37 | 51 | 40 | 5.5 | 27.21 | 311.74 |
| 1947 | MAR | 21 | 23 | 00 04.2 | 34.92 | 23.30 | 21 | 40 | 5.1 | 22.54 | 204.60 |
| 1947 | APR | 12 | 14 | 0513.4 | 39.86 | 25.01 | 8 | 40 | 5.3 | 63.18 | 233.33 |


| DATE |  |  | ORIG.TIME GMT |  |  | $\underset{\mathrm{N}}{\mathrm{LAT}}$ | $\begin{gathered} \text { LON } \\ \mathrm{E} \end{gathered}$ | $\begin{gathered} \text { DEPTH } \\ \text { KM } \end{gathered}$ | OBS | MAG MS | DIST KM | FT <br> AZIM DEG |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1947 | APR | 19 |  | 29 | 44.4 | 39.33 | 23.55 | 30 | 35 | 5.3 | 66.43 | 27.13 |
| 1947 | JUN | 01 |  | 18 | 45.3 | 36.74 | 21.78 | 62 | 52 | 5.6 | 29.84 | 58.09 |
| 1947 |  | 04 |  | 29 | 57.6 | 40.09 | 23.96 | 45 | 65 | 6.0 | 58.94 | 22.16 |
| 1947 | JUL | 07 | 22 | 35 | 44.1 | 37.37 | 21.05 | 18 | 49 | 5.2 | 29.51 | 49.52 |
| 1947 |  | 21 | 09 | 36 | 36.3 | 37.55 | 22.99 | 60 | 29 | 5.0 | 103.99 | 24.63 |
| 1947 | AUG | 17 | 15 | 04 | 20.9 | 37.55 | 19.98 | 13 | 31 | 5.1 | 17.60 | 353.94 |
| 1947 |  | 30 | 22 | 21 | 41.9 | 35.50 | 23.37 | 34 | 88 | 6.2 | 44.75 | 356.49 |
| 1947 | SEP | 13 | 15 | 11 | 21.9 | 37.54 | 20.05 | 15 | 36 | 5.3 | 16.75 | 15.87 |
| 1947 | OCT | 06 | 19 | 55 | 36.3 | 36.71 | 21.79 | 2 | 118 | 6.7 | 28.60 | 221.70 |
| 1947 | NOV | 29 | 10 | 14 | 04.8 | 39.23 | 23.67 | 15 | 30 | 5.1 | 104.02 | 353.80 |
| 1947 | DEC | 09 | 23 | 19 | 04.6 | 42.20 | 19.75 | 50 | 22 | 5.0 | 127.93 | 16.91 |
| 1948 | JAN | 17 | 02 | 26 | 30.6 | 38.25 | 20.94 | 68 | 14 | 5.0 | 49.53 | 276.63 |
| 1948 | FEB | 09 | 12 | 58 | 17.9 | 35.32 | 27.15 | 25 | 124 | 7.2 | 21.21 | 192.83 |
| 1948 |  | 10 | 15 | 59 | 00.2 | 35.38 | 27.40 | 57 | 28 | 5.2 | 23.07 | 126.19 |
| 1948 |  | 11 | 22 | 31 | 24 | 35.5 | 27.1 | 15 | UNS | 5.0 |  |  |
| 1948 |  | 12 | 22 | 27 | 19.6 | 35.91 | 27.35 | 70 | 45 | 5.4 | 47.74 | 16.57 |
| 1948 |  | 15 | 17 | 55 | 02.0 | 35.37 | 27.34 | 58 | 46 | 5.4 | 19.92 | 138.57 |
| 1948 | MAR | 06 | 20 | 12 | 59.1 | 35.26 | 25.93 | 48 | 41 | 5.2 | 59.43 | 30.45 |
| 1948 |  | 26 | 03 | 02 | 09.5 | 40.60 | 21.47 | 39 | 34 | 5.0 | 87.38 | 152.87 |
| 1948 |  | 29 | 02 | 33 | 04.5 | 35.69 | 27.23 | 89 | 50 | 5.4 | 21.88 | 7.34 |
| 1948 |  | 29 | 10 | 22 | 48.3 | 35.28 | 23.32 | 47 | 63 | 5.3 | 21.88 | 339.98 |
| 1948 | APR | 22 | 10 | 42 | 49.7 | 38.73 | 20.38 | 12 | 100 | 6.7 | 21.21 | 247.96 |
| 1948 | MAY | 07 | 14 | 57 | 20.8 | 39.05 | 18.90 | 10 | 35 | 5.1 | 105.95 | 184.69 |
| 1948 |  | 22 | 05 | 07 | 56.8 | 34.65 | 24.31 | 38 | 31 | 5.5 | 167.46 | 26.33 |
| 1948 |  | 27 | 07 | 32 | 43.9 | 36.53 | 23.31 | 162 | 16 | 5.1 | 112.98 | 289.28 |
| 1948 | JUN | 17 | 06 | 52 | 35.9 | 37.66 | 21.81 | 53 | 34 | 5.1 | 30.24 | 238.80 |
| 1948 |  | 30 | 12 | 21 | 21.3 | 38.96 | 20.53 | 36 | 96 | 6.7 | 19.42 | 341.15 |
| 1948 | JUL | 24 | 06 | 03 | 10.9 | 34.49 | 24.49 | 20 | 135 | 6.4 | 10.99 | 354.74 |
| 1948 |  | 26 | 11 | 26 | 33.3 | 35.70 | 27.43 | 45 | 21 | 5.0 | 30.87 | 43.14 |
| 1948 | AUG | 10 | 13 | 27 | 10.2 | 38.47 | 28.88 | 79 | 33 | 5.1 | 174.48 | 2.29 |
| 1948 |  | 27 | 10 | 44 | 16.8 | 42.06 | 19.38 | 41 | 55 | 5.4 | 41.50 | 346.05 |
| 1948 |  | 27 | 11 | 24 | 27.0 | 41.89 | 19.54 | 20 | UNS | 5.0 |  |  |
| 1948 | SEP | 11 | 08 | 52 | 44.0 | 37.38 | 23.28 | 88 | 75 | 6.2 | 21.89 | 19.53 |
| 1948 |  | 20 | 18 | 00 | 00.5 | 34.60 | 26.58 | 74 | 42 | 5.2 | 51.07 | 229.25 |
| 1948 |  | 21 | 17 | 54 | 01.1 | 36.45 | 21.60 | 59 | 37 | 5.6 | 19.42 | 151.69 |
| 1948 | OCT | 10 | 17 | 43 | 10.5 | 35.43 | 23.54 | 43 | 62 | 5.6 | 39.04 | 19.14 |
| 1948 |  | 18 | 09 | 00 | 02.0 | 35.73 | 27.21 | 43 | 58 | 5.6 | 26.12 | 2.03 |
| 1948 |  | 19 | 03 | 04 | 39.1 | 35.62 | 27.88 | 86 | 13 | 5.0 | 63.29 | 77.62 |
| 1948 | NOV | 13 | 04 | 44 | 58.1 | 41.08 | 28.19 | 28 | 48 | 5.6 | 185.89 | 320.37 |
| 1949 | JAN | 14 | 15 | 53 | 58.7 | 38.76 | 25.23 | 12 | 40 | 5.7 | 7.52 | 233.77 |
| 1949 | FEB | 05 | 00 | 28 | 22.5 | 39.98 | 29.47 | 35 | 42 | 5.0 | 23.38 | 330.95 |
| 1949 |  | 05 | 15 | 24 | 22.5 | 38.03 | 21.69 | 57 | 27 | 5.0 | 60.84 | 86.64 |
| 1949 | JUN | 17 | 04 | 21 | 06.1 | 34.42 | 28.32 | 77 | 65 | 5.8 | 17.46 | 277.70 |
| 1949 |  | 26 | 05 | 42 | 34.4 | 39.80 | 20.47 | 52 | 47 | 5.1 | 111.71 | 354.27 |
| 1949 | JUL | 07 | 12 | 21 | 15.8 | 35.95 | 27.14 | 67 | 51 | 5.2 | 50.54 | 353.81 |
| 1949 |  | 23 | 15 | 03 | 35.2 | 38.71 | 26.27 | 17 | 124 | 7.0 | 13.56 | 347.94 |
| 1949 |  | 30 | 17 | 47 | 14.3 | 38.72 | 26.37 | 48 | 33 | 5.2 | 15.39 | 24.56 |
| 1949 | SEP | 17 | 11 | 30 | 15.8 | 37.07 | 22.67 | 42 | 43 | 5.0 | 62.76 | 72.25 |
| 1949 | OCT | 04 | 17 | 33 | 33.9 | 38.63 | 22.08 | 111 | 31 | 5.0 | 28.76 | 59.32 |
| 1949 | NOV | 23 | 16 | 51 | 02.6 | 38.58 | 26.22 | 25 | 52 | 5.6 | 7.30 | 252.26 |
| 1949 | DEC | 07 | 16 | 13 | 39.8 | 34.72 | 24.18 | 19 | 51 | 5.2 | 7.64 | 73.09 |
| 1950 | FEB | 12 | 09 | 43 | 52.0 | 34.40 | 24.26 | 29 | 17 | 5.0 | 36.67 | 156.14 |
| 1950 | MAY | 03 | 07 | 13 | 50.8 | 38.96 | 27.35 | 60 | 25 | 5.0 | 50.50 | 37.16 |
| 1950 |  | 30 | 09 | 52 | 42.6 | 35.69 | 27.50 | 50 | 21 | 5.0 | 34.85 | 52.13 |
| 1950 | SEP | 23 | 06 | 23 | 48.5 | 34.90 | 25.81 | 49 | 87 | 5.3 | 22.65 | 59.93 |


| date |  |  | $\underset{\text { GMT }}{\text { ORIG.TME }}$ |  | $\underset{\mathrm{N}}{\mathrm{LAT}}$ | $\begin{gathered} \text { LON } \\ \mathrm{E} \end{gathered}$ | $\begin{aligned} & \text { DEPTH } \\ & \text { KM } \end{aligned}$ | OBS | $\begin{gathered} \text { MAG } \\ \text { MS } \end{gathered}$ | DIST K | FT <br> AZIM DEG |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1950 | OCT | 22 | 05 | 5211.0 | 34.94 | 26.28 | 47 | 27 | 4.9 | 17.87 | 25.20 |
| 1950 | NOV | 28 | 17 | 5323.0 | 39.53 | 28.19 | 49 | 33 | 5.1 | 48.29 | 148.72 |
| 1950 | DEC | 28 | 22 | 3137.6 | 35.63 | 27.47 | 97 | 16 | 5.2 | 28.85 | 59.41 |
| 1951 | Jan | 09 | 00 | 2802.5 | 38.04 | 20.26 | 61 | 50 | 5.3 | 85.33 | 201.72 |
| 1951 | APR | 05 | 03 | 1529.4 | 37.46 | 20.30 | 41 | 78 | 5.3 | 34.15 | 78.69 |
| 1951 | AUG | 20 | 22 | 5154.7 | 35.06 | 24.07 | 62 | 74 | 5.1 | 40.32 | 356.08 |
| 1951 |  | 24 | 10 | 2734.1 | 37.22 | 21.43 | 51 | 67 | 5.1 | 15.24 | 127.55 |
| 1951 |  | 31 | 12 | 2945.4 | 35.89 | 22.45 | 30 | 100 | 5.6 | 53.84 | 323.90 |
| 1951 |  | 31 | 20 | 1838.9 | 35.51 | 22.73 | 23 | 68 | 5.1 | 6.43 | 279.96 |
| 1951 | SEP | 15 | 22 | 5212.3 | 40.23 | 27.69 | 7 | 40 | 5.6 | 32.77 | 234.48 |
| 1951 | OCT | 01 | 01 | 2641.4 | 34.60 | 26.62 | 49 | 70 | 5.1 | 16.61 | 44.76 |
| 1951 | NOV | 05 |  | 4355 | 36.0 | 29.0 |  | ATB | 5.2 |  |  |
| 1951 | DEC | 13 | 20 | 4620.5 | 40.38 | 26.82 | 215 | 24 | 5.1 | 105.79 | 78.70 |
| 1951 |  | 20 | 19 | 1205.6 | 38.07 | 20.05 | 27 | 39 | 5.3 | 70.66 | 249.00 |
| 1952 | FEB | 03 | 20 | 4501.1 | 40.47 | 25.85 | 77 | 25 | 5.0 | 37.13 | 35.24 |
| 1952 | MAR | 09 | 04 | 4529.9 | 37.48 | 20.49 | 47 | 26 | 5.2 | 73.32 | 218.07 |
| 1952 |  | 13 | 06 | 3001.8 | 41.02 | 28.14 | 11 | 27 | 5.4 | 4.02 | 56.49 |
| 1952 |  | 19 | 01 | 2727.7 | 39.61 | 28.60 | 33 | 116 | 5.8 | 23.38 | 202.16 |
| 1952 |  | 25 | 03 | 3521.4 | 34.87 | 23.32 | 38 | 37 | 4.9 | 26.94 | 196.00 |
| 1952 | APR | 03 | 03 | 2011.6 | 38.21 | 20.49 | 54 | 18 | 4.9 | 50.64 | 297.68 |
| 1952 | JUN | 09 | 14 | 4843.3 | 36.94 | 27.62 | 42 | 19 | 4.9 | 63.22 | 18.10 |
| 1952 |  | 12 | 11 | 0015.6 | 34.67 | 26.56 | 56 | 36 | 5.6 | 36.34 | 113.52 |
| 1952 |  | 13 | 01 | 0730.2 | 37.31 | 21.98 | 55 | 64 | 5.3 | 11.85 | 276.01 |
| 1952 |  | 27 | 13 | 0923.8 | 40.68 | 23.32 | 16 | 41 | 4.8 | 16.01 | 261.75 |
| 1952 | JUL | 08 | 20 | 5844.1 | 36.00 | 21.94 | 50 | 15 | 4.6 | 22.86 | 170.77 |
| 1952 | AUG | 21 | 04 | 1829.0 | 35.32 | 25.45 | 62 | 28 | 4.6 | 44.59 | 162.07 |
| 1952 |  | 24 | 20 | 4429.3 | 35.35 | 27.29 | 80 | 41 | 4.9 | 19.30 | 153.82 |
| 1952 | SEP | 23 | 20 | 3052.0 | 36.90 | 29.90 | 35 | 30 | 5.0 | 35.13 | 50.25 |
| 1952 | OCT | 05 | 10 | 2119.1 | 37.07 | 20.92 | 24 | 43 | 5.0 | 49.09 | 167.39 |
| 1952 |  | 05 | 10 | 5457.9 | 37.41 | 20.61 | 9 | 106 | 5.7 | 20.04 | 239.34 |
| 1952 |  | 07 | 16 | 0834.7 | 37.02 | 20.72 | 28 | 19 | 4.7 | 53.88 | 187.61 |
| 1952 |  | 10 | 11 | 5158.6 | 37.23 | 20.61 | 9 | 46 | 5.2 | 34.64 | 209.38 |
| 1952 |  | 13 | 16 | 4232.7 | 39.18 | 23.40 | 15 | 55 | 5.2 | 45.88 | 22.27 |
| 1952 |  | 22 | 04 | 1512.1 | 36.86 | 27.09 | 107 | 35 | 5.2 | 74.59 | 284.04 |
| 1952 | DEC | 17 | 23 | 0402.0 | 34.47 | 24.22 | 17 | 232 | 6.6 | 27.30 | 286.87 |
| 1952 |  | 22 | 23 | 5147.5 | 35.28 | 25.16 | 74 | 24 | 5.2 | 48.44 | 195.29 |
| 1952 |  | 31 | 14 | 4852.8 | 35.75 | 25.95 | 80 | 102 | 5.5 | 36.21 | 39.16 |
| 1952 |  | 31 | 17 | 1852.9 | 35.66 | 25.97 | 54 | 100 | 5.7 | 30.72 | 53.97 |
| 1953 | JAN | 07 |  | 0128 | 41.3 | 20.6 | 35 | UNS | 5.2 |  |  |
| 1953 |  | 07 |  | 1857 | 41.3 | 20.6 | 19 | UNS | 5.4 |  |  |
| 1953 |  | 10 | 23 | 2911.9 | 38.28 | 25.23 | 20 | ATB | 5.0 |  |  |
| 1953 | FEB | 05 | 22 | 4205 | 35.7 | 22.7 | 15 | UNS | 4.9 |  |  |
| 1953 |  | 07 | 22 | 3113.2 | 34.83 | 24.11 | 33 | 138 | 5.7 | 15.24 | 3.63 |
| 1953 |  | 14 |  | 4321.8 | 36.07 | 27.04 | 96 | 129 | 5.3 | 80.07 | 37.51 |
| 1953 |  | 22 |  | 2623 | 37.7 | 21.2 | 80 | UNS | 5.3 |  |  |
| 1953 | MAR | 18 | 19 | 0616.8 | 40.20 | 27.52 | 8 | 259 | 7.4 | 29.51 | 40.12 |
| 1953 |  | 18 | 21 | 1811.1 | 40.04 | 27.52 | 28 | 57 | 5.7 | 20.04 | 76.62 |
| 1953 |  | 19 |  | 5345.2 | 40.00 | 28.04 | 50 | 18 | 4.9 | 63.45 | 89.76 |
| 1953 |  | 19 | 21 | 1359.5 | 40.05 | 27.28 |  | 50 | 5.0 | 5.81 | 342.98 |
| 1953 |  | 26 | 15 | 1030.5 | 39.94 | 27.48 | 10 | ATB | 5.1 |  |  |
| 1953 |  | 31 | 00 | 5552.3 | 40.51 | 19.86 | 17 | 48 | 5.0 | 24.59 | 209.37 |
| 1953 | APR | 01 | 01 | 4740.3 | 40.13 | 27.48 | 23 | 42 | 5.3 | 21.77 | 46.72 |
| 1953 |  | 02 | 08 | 2146.1 | 38.88 | 25.32 | 10 | ATB | 5.0 |  |  |
| 1953 |  | 05 | 03 | 2156.1 | 40.61 | 19.96 | 8 | 25 | 4.5 | 11.43 | 198.72 |
| 1953 | MAY | 01 |  | 0647.0 | 38.74 | 26.66 | 48 | 33 | 4.9 | 35.40 | 63.50 |


| DATE |  |  | ORIG.TIME GMT |  | $\begin{gathered} \text { LAT } \\ \mathrm{N} \end{gathered}$ | $\underset{\mathrm{E}}{\mathrm{LON}}$ | $\begin{aligned} & \text { DEPTH } \\ & \text { KM } \end{aligned}$ | OBS | $\begin{gathered} \text { MAG } \\ \text { MS } \end{gathered}$ | $\text { DIST } \stackrel{\text { S }}{\mathrm{S}}$ | HIFT |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1953 | MAY | 02 | 0541 | 59.3 | 38.87 | 27.04 | 55 | 34 | 5.2 | 71.18 | 64.79 |
| 1953 |  | 02 | 1006 | 51.1 | 38.61 | 27.16 | 100 | ATB | 4.8 |  |  |
| 1953 |  | 02 | 1837 | 45.5 | 38.80 | 26.72 | 57 | 47 | 5.4 | 43.05 | 58.58 |
| 1953 |  | 14 | 1300 | 29.8 | 38.17 | 26.61 | 100 | ATB | 5.0 |  |  |
| 1953 | JUN | 03 | 1605 | 28.0 | 40.19 | 28.72 | 28 | 52 | 5.7 | 13.01 | 325.72 |
| 1953 |  | 07 | 1352 | 57.8 | 35.95 | 27.11 | 80 | ATB | 4.8 |  |  |
| 1953 |  | 09 | 1628 | 27.8 | 39.43 | 28.14 | 22 | 37 | 4.9 | 15.24 | 122.78 |
| 1953 |  | 13 | 1838 | 58 | 38.1 | 22.6 | 4 | UNS | 5.1 |  |  |
| 1953 |  | 18 | 0544 | 11.6 | 41.84 | 26.54 | 17 | 110 | 5.4 | 17.04 | 342.23 |
| 1953 |  | 21 | 0811 | 25 | 37.6 | 20.6 | 24 | 27 | 4.9 |  |  |
| 1953 |  | 23 | 0153 | 20.8 | 36.06 | 24.69 | 100 | 70 | 5.8 | 68.23 | 306.13 |
| 1953 | JUL | 22 | 1509 | 39.8 | 39.35 | 28.24 | 12 | 85 | 5.6 | 36.08 | 321.11 |
| 1953 | AUG | 09 | 0741 | 12.6 | 38.24 | 20.80 | 21 | 171 | 6.1 | 6.67 | 180.00 |
| 1953 |  | 11 | 0332 | 26.7 | 38.35 | 20.74 | 11 | 244 | 6.8 | 7.63 | 316.74 |
| 1953 |  | 11 | 0432 | 25 | 38.1 | 20.8 | 10 | UNS | 5.0 |  |  |
| 1953 |  | 11 | 1243 | 32.1 | 38.50 | 20.53 | 36 | 53 | 5.4 | 32.85 | 313.35 |
| 1953 |  | 11 | 1311 | 09.7 | 38.25 | 20.87 | 17 | 47 | 5.1 | 8.26 | 132.28 |
| 1953 |  | 12 | 0608 | 11.5 | 38.39 | 20.59 | 43 | 61 | 4.8 | 21.44 | 298.62 |
| 1953 |  | 12 | 0923 | 55.4 | 38.13 | 20.74 | 11 | 257 | 7.3 | 20.16 | 195.58 |
| 1953 |  | 12 | 1007 | 38 | 38.1 | 20.8 | 10 | UNS | 5.5 |  |  |
| 1953 |  | 12 | 1133 | 52.3 | 38.06 | 20.81 | 40 | 69 | 5.4 | 27.03 | 178.11 |
| 1953 |  | 12 | 1205 | 25.6 | 37.88 | 20.76 | 18 | 160 | 6.3 | 25.36 | 237.81 |
| 1953 |  | 12 | 1339 | 28.1 | 38.09 | 20.81 | 28 | 72 | 5.8 | 20.04 | 300.98 |
| 1953 |  | 12 | 1408 | 44.1 | 38.12 | 20.84 | 20 | 95 | 6.0 | 20.98 | 170.04 |
| 1953 |  | 12 | 1608 | 38.2 | 38.06 | 20.84 | 36 | 68 | 5.5 | 16.31 | 295.42 |
| 1953 |  | 13 | 0322 | 10.5 | 38.29 | 20.88 | 10 | 65 | 5.5 | 7.07 | 99.04 |
| 1953 |  | 13 | 1016 | 50 | 38.1 | 20.8 | 10 | UNS | 5.3 |  |  |
| 1953 |  | 17 | 0212 | 28.8 | 38.12 | 20.99 | 37 | 41 | 5.0 | 14.25 | 356.23 |
| 1953 | SEP | 05 | 0108 | 12.9 | 36.96 | 29.35 | 80 | ATB | 5.1 |  |  |
| 1953 |  | 05 | 1418 | 46.0 | 37.88 | 23.17 | 18 | 95 | 5.7 | 20.51 | 131.66 |
| 1953 |  | 14 | 1456 | 17.8 | 38.38 | 20.78 | 07 | 73 | 5.2 | 10.08 | 348.87 |
| 1953 | OCT | 10 | 2129 | 18.2 | 38.08 | 20.98 | 22 | 73 | 5.3 | 29.51 | 147.09 |
| 1953 |  | 16 | 2144 | 49.6 | 38.14 | 20.70 | 41 | 38 | 5.0 | 20.39 | 206.28 |
| 1953 |  | 21 | 1131 | 10.7 | 38.38 | 20.70 | 14 | 79 | 5.4 | 13.38 | 315.48 |
| 1953 |  | 21 | 1839 | 57.2 | 38.30 | 20.59 |  | 172 | 6.4 | 19.05 | 270.07 |
| 1953 |  | 21 | 2344 | 01 | 38.3 | 20.8 | 15 | UNS | 5.0 |  |  |
| 1953 | Nov | 03 | 2229 | 25 | 37.9 | 21.2 | 4 | UNS | 4.9 |  |  |
| 1953 |  | 08 | 1445 | 54.4 | 38.98 | 23.99 | 22 | 33 | 5.0 | 23.07 | 233.84 |
| 1953 |  | 20 | 1913 | 57 | 38.4 | 20.8 | 12 | UNS | 5.0 |  |  |
| 1953 |  | 28 | 2017 | 36.1 | 37.49 | 20.70 | 37 | 60 | 5.3 | 62.64 | 205.13 |
| 1953 |  | 30 | 1321 | 03.9 | 38.32 | 21.60 | 33 | 45 | 4.8 | 27.03 | 138.77 |
| 1953 | DEC | 20 | 1756 | 20.1 | 35.99 | 27.77 | 40 | ATB | 4.9 |  |  |
| 1953 |  | 28 | 0238 | 49.6 | 38.30 | 20.56 | 17 | 75 | 5.3 | 21.66 | 270.08 |
| 1954 | JAN | 02 | 0113 | 41.3 | 36.98 | 27.12 | 140 | ATB | 5.2 |  |  |
| 1954 |  | 18 | 1416 | 14.8 | 37.62 | 21.60 | 37 | 41 | 5.1 | 67.76 | 128.42 |
| 1954 |  | 24 | 1332 | 54.2 | 37.38 | 20.46 | 13 | 25 | 4.6 | 33.29 | 246.23 |
| 1954 | MAR | 08 | 0817 | 21.9 | 38.06 | 20.61 |  | 64 | 5.2 | 6.73 | 7.47 |
| 1954 |  | 23 | 1258 | 53.3 | 40.58 | 27.12 | 10 | ATB | 4.9 |  |  |
| 1954 | APR | 17 | 2052 | 51.5 | 37.99 | 22.98 | 19 | 71 | 5.1 | 13.20 | 35.12 |
| 1954 |  | 30 | 1302 | 39.5 | 39.23 | 22.28 | 16 | 211 | 6.7 | 11.43 | 138.35 |
| 1954 |  | 30 | 1933 | 30 | 39.3 | 22.2 | 26 | UNS | 4.7 |  |  |
| 1954 | MAY | 01 | 1524 | 59.3 | 37.79 | 27.07 | 42 | 35 | 5.0 | 12.83 | 31.67 |
| 1954 |  | 01 | 2053 | 34.6 | 37.81 | 26.95 | 54 | 60 | 5.5 | 13.91 | 340.17 |
| 1954 |  | 03 | 0524 | 55 | 36.0 | 21.5 | 15 | UNS | 4.8 |  |  |
| 1954 |  | 03 | 0851 | 17 | 36.0 | 21.5 | 15 | UNS | 5.0 |  |  |


| DATE |  |  | $\begin{aligned} & \text { ORIG.TIME } \\ & \text { GMT } \end{aligned}$ |  | $\begin{aligned} & \text { LAT } \\ & \mathrm{N} \end{aligned}$ | $\begin{gathered} \text { LON } \\ \text { E } \end{gathered}$ | $\begin{aligned} & \text { DEPTH } \\ & \text { KM } \end{aligned}$ | OBS | $\begin{gathered} \text { MAG } \\ \text { MS } \end{gathered}$ | dIST | SHIFT |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1954 | MAY | 03 | 13 | 2945.0 | 35.28 | 27.23 | 12 | 52 | 4.9 | 25.08 | 264.92 |
| 1954 |  | 04 | 16 | 4326.0 | 39.26 | 22.14 | 27 | 80 | 5.5 | 6.81 | 229.27 |
| 1954 |  | 04 |  | 4532.7 | 39.24 | 22.35 | 28 | 71 | 5.4 | 15.24 | 117.18 |
| 1954 |  | 04 | 23 | 4454 | 39.3 | 22.2 | 20 | UNS | 4.8 |  |  |
| 1954 |  | 12 | 021 | 1633 | 37.7 | 21.8 | 5 | UNS | 4.9 |  |  |
| 1954 |  | 15 | 12 | 2434 | 36.2 | 21.7 | 15 | UNS | 5.0 |  |  |
| 1954 |  | 25 | 22 | 0337.1 | 39.26 | 22.30 | 22 | 91 | 5.3 | 10.54 | 117.20 |
| 1954 | JUL | 18 | 14 | 4237.2 | 37.68 | 21.18 | 30 | 54 | 5.1 | 10.55 | 348.76 |
| 1954 | AUG | 03 | 181 | 1811.8 | 40.28 | 24.28 | 35 | 118 | 5.8 | 38.56 | 53.61 |
| 1954 |  | 05 | 041 | 1251 | 40.2 | 25.0 | 26 | UNS | 5.0 |  |  |
| 1954 |  | 05 |  | 3917.2 | 35.89 | 27.42 | 42 | 50 | 4.7 | 13.38 | 324.14 |
| 1954 |  | 06 | 11 | 3351 | 36.8 | 23.2 | 20 | UNS | 5.0 |  |  |
| 1954 |  | 06 |  | 0100 | 39.8 | 25.0 | 28 | UNS | 4.8 |  |  |
| 1954 | SEP | 02 | 015 | 5438.7 | 41.96 | 19.68 | 15 | 48 | 4.9 | 19.05 | 212.62 |
| 1954 |  | 04 | 041 | 1923.4 | 36.63 | 27.10 | 160 | ATB | 4.7 |  |  |
| 1954 | OCT | 24 | 23 | 3719.1 | 40.46 | 27.53 | 10 | ATB | 5.0 |  |  |
| 1954 |  | 26 | 103 | 3428.6 | 40.56 | 27.52 | 10 | ATB | 4.6 |  |  |
| 1954 | NOV | 23 | 23 | 2254.3 | 35.89 | 27.60 | 40 | ATB | 5.0 |  |  |
| 1954 | DEC | 23 | 162 | 2725.1 | 37.87 | 21.19 | 38 | 69 | 5.4 | 8.57 | 112.89 |
| 1954 |  | 30 | 02 | 0726.5 | 40.59 | 22.84 | 14 | 39 | 4.9 | 17.60 | 306.45 |
| 1954 |  | 30 | 11 | 0559.8 | 36.15 | 21.79 | 9 | 63 | 5.2 | 10.99 | 55.57 |
| 1955 | JAN | 03 | 01 | 0710.9 | 39.19 | 22.27 | 41 | 100 | 5.6 | 26.12 | 34.83 |
| 1955 |  | 08 | 075 | 5309.0 | 39.27 | 22.17 | 52 | 56 | 5.0 | 17.32 | 62.05 |
| 1955 |  | 28 | 074 | 4206.2 | 33.91 | 23.54 | 17 | 40 | 4.7 | 5.64 | 281.35 |
| 1955 | FEB | 21 | 194 | 4644 | 39.4 | 23.1 | 4 | UNS | 4.7 |  |  |
| 1955 |  | 22 | 094 | 4300 | 39.4 | 23.1 | 7 | UNS | 4.8 |  |  |
| 1955 | MAR | 28 |  | 4552.5 | 37.60 | 21.24 | 9 | 78 | 5.1 | 12.44 | 162.35 |
| 1955 | APR | 13 | 20 | 4551.3 | 37.29 | 22.50 | 19 | 99 | 5.2 | 10.31 | 262.91 |
| 1955 |  | 19 | 16 | 4723.8 | 39.31 | 23.06 | 15 | 139 | 6.2 | 11.64 | 199.05 |
| 1955 |  | 21 | 07 | 1818 | 39.3 | 23.0 | 5 | UNS | 5.8 |  |  |
| 1955 |  | 22 |  | 0233.0 | 34.75 | 23.77 | 46 | 46 | 4.7 | 22.54 | 284.82 |
| 1955 | JUN | 02 | 23 | 3439.4 | 40.37 | 25.59 | 10 | 72 | 5.7 | 1879 | 259.49 |
| 1955 | JUL | 09 | 23 | 5348.5 | 40.82 | 22.42 | 32 | 72 | 5.0 | 23.38 | 54.27 |
| 1955 |  | 16 | 07 | 0717.2 | 37.68 | 27.20 | 31 | 232 | 7.0 | 10.31 | 360.00 |
| 1955 |  | 18 | 03 | 0611.1 | 37.75 | 27.72 | 40 | ATB | 4.9 |  |  |
| 1955 | AUG | 28 | 13 | 3928.8 | 37.36 | 27.02 | 48 | 41 | 5.3 | 29.59 | 238.00 |
| 1955 | NOV | 11 | 18 | 2740.5 | 37.54 | 26.97 | 10 | ATB | 4.7 |  |  |
| 1956 | JAN | 06 | 12 | 1546.1 | 40.51 | 26.33 | 10 | 134 | 5.7 | 13.56 | 11.76 |
| 1956 |  | 27 | 01 | 1332.1 | 36.40 | 23.75 | 56 | 40 | 4.6 | 23.07 | 270.08 |
| 1956 | MAY | 05 | 20 | 4200.3 | 36.99 | 28.63 | 40 | ATB | 4.7 |  |  |
| 1956 |  | 15 | 18 | 3416.6 | 37.25 | 20.89 | 15 | 61 | 5.0 | 5.62 | 189.04 |
| 1956 |  | 15 | 22 | 5657.3 | 37.28 | 20.95 | 20 | 63 | 5.2 | 10.25 | 185.41 |
| 1956 |  | 18 | 22 | 0836.7 | 39.03 | 22.63 | 52 | 69 | 5.1 | 15.71 | 282.80 |
| 1956 | JUN | 11 | 01 | 1131.5 | 34.26 | 26.02 | 36 | 49 | 4.8 | 18.40 | 291.92 |
| 1956 | JUL | 09 | 031 | 1143.7 | 36.64 | 25.91 | 15 | 270 | 7.4 | 14.91 | 135.41 |
| 1956 |  | 09 | 03 | 2416.5 | 36.45 | 25.51 | 95 | 57 | 7.2 | 26.76 | 220.48 |
| 1956 |  | 09 | 06 | 1916.9 | 36.66 | 25.70 | 70 | ATB | 5.0 |  |  |
| 1956 |  | 09 |  | 2259.0 | 36.71 | 25.60 | 78 | 50 | 5.3 | 27.12 | 255.67 |
| 1956 |  | 09 | 07 | 3630.0 | 36.27 | 25.89 | 30 | ATB | 4.8 |  |  |
| 1956 |  | 09 | 09 | 4510.0 | 36.60 | 25.93 | 10 | ATB | 4.8 |  |  |
| 1956 |  | 09 | 11 | 3055.7 | 36.54 | 26.32 | 40 | ATB | 4.9 |  |  |
| 1956 |  | 09 | 20 | 1358.9 | 36.62 | 25.84 | 14 | 61 | 5.3 | 16.75 | 231.45 |
| 1956 |  | 09 | 20 | 4808.3 | 36.45 | 26.09 | 60 | ATB | 4.8 |  |  |
| 1956 |  | 09 | 21 | 2851.6 | 36.52 | 25.81 | 61 | 46 | 4.7 | 16.61 | 236.58 |
| 1956 |  | 10 | 03 | 0135.2 | 36.82 | 26.15 | 55 | 78 | 5.5 | 16.01 | 323.85 |


| DATE |  | ORIG. TIME |  | LAT | LON | DEPTH | OBS | MAG | SHIFT |  |
| :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- | :--- |
|  |  | GMT | N | E | KM |  | MS | DIST KM | AZIM DEG |  |
| 1956 | JUL | 22 | O3 | 29 | 06.7 | 36.89 | 26.32 | 40 | 57 | 4.8 |


| DATE |  |  | $\begin{aligned} & \text { ORIG.TIME } \\ & \text { GMT } \end{aligned}$ |  | $\begin{gathered} \text { LAT } \\ \mathrm{N} \end{gathered}$ | $\begin{gathered} \text { LON } \\ \mathrm{E} \end{gathered}$ | $\begin{aligned} & \text { DEPTH } \\ & \text { KM } \end{aligned}$ | OBS | $\begin{gathered} \text { MAG } \\ \text { MS } \end{gathered}$ | DIST K | FT <br> AZIM DEG |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1958 | APR | 03 |  | 2348.4 | 41.19 | 19.76 | 29 | 116 | 5.4 | 10.54 | 306.34 |
| 1958 |  | 03 | 071 | 1841.8 | 34.90 | 27.29 | 40 | 91 | 5.2 | 15.86 | 282.86 |
| 1958 |  | 04 |  | 1856.5 | 41.24 | 19.74 | 37 | 36 | 4.4 | 27.65 | 80.49 |
| 1958 |  | 24 | 08 | 0039.5 | 36.76 | 26.55 | 10 | ATB | 4.8 |  |  |
| 1958 | MAY | 03 | 201 | 1820.8 | 36.19 | 21.73 | 16 | 107 | 5.0 | 4.53 | 348.60 |
| 1958 |  | 09 | 02 | 4056.8 | 36.61 | 27.60 | 67 | 133 | 5.5 | 21.09 | 336.89 |
| 1958 |  | 27 | 18 | 2748.0 | 36.86 | 26.67 | 163 | 124 | 5.1 | 7.45 | 287.34 |
| 1958 | Jun | 05 |  | 2948.6 | 37.20 | 20.73 | 29 | 72 | 4.8 | 24.59 | 14.93 |
| 1958 |  | 10 |  | 2857.5 | 41.15 | 19.81 | 41 | 45 | 4.6 | 37.07 | 62.93 |
| 1958 |  | 30 | 08 | 4247.1 | 36.44 | 27.28 | 112 | 148 | 5.3 | 10.77 | 338.01 |
| 1958 | JU | 15 | 075 | 5925.1 | 35.54 | 23.58 | 38 | 88 | 4.8 | 21.77 | 2.46 |
| 1958 |  | 17 |  | 3711.4 | 40.72 | 23.39 | 19 | 134 | 5.5 | 13.38 | 11.73 |
| 1958 | AUG | 27 | 151 | 1634.6 | 37.45 | 20.67 | 9 | 183 | 6.5 | 4.41 | 270.00 |
| 1958 |  | 30 | 07 | 3544.9 | 37.35 | 20.60 | 11 | 56 | 4.6 | 15.71 | 41.27 |
| 1958 | SEP | 02 | 011 | 1325.5 | 37.44 | 20.62 | 13 | 86 | 5.4 | 5.53 | 232.93 |
| 1958 |  | 04 | 00 | 0300.5 | 36.54 | 26.70 | 35 | 86 | 5.0 | 6.67 | 180.02 |
| 1958 |  | 04 | 02 | 5106.2 | 36.40 | 27.01 | 140 | ATB | 4.9 |  |  |
| 1958 | NOV | 15 | 05 | 4240.5 | 37.45 | 21.73 | 31 | 123 | 5.5 | 37.00 | 220.77 |
| 1959 | Jan | 03 | 075 | 5923.9 | 35.26 | 29.04 | 80 | ATB | 4.8 |  |  |
| 1959 |  | 06 | 14 | 2840.9 | 36.66 | 29.11 | 30 | ATB | 4.7 |  |  |
| 1959 |  | 07 | 22 | 2203.2 | 36.79 | 29.14 | 26 | 48 | 4.7 | 8.39 | 301.98 |
| 1959 |  | 09 |  | 5508.2 | 36.15 | 21.67 | 16 | 70 | 4.9 | 13.38 | 290.81 |
| 1959 |  | 11 |  | 2735.1 | 36.77 | 29.07 | 61 | 52 | 4.7 | 14.25 | 356.17 |
| 1959 |  | 26 | 11 | 3843.9 | 36.78 | 29.02 | 47 | 42 | 5.0 | 12.24 | 299.50 |
| 1959 | FEB | 07 | 20 | 0825.7 | 37.56 | 20.90 | 50 | 59 | 4.6 | 26.48 | 340.11 |
| 1959 | MAR | 08 | 111 | 1718.5 | 40.21 | 19.89 | 47 | 39 | 4.5 | 3.40 | 90.00 |
| 1959 |  | 13 | 19 | 0811.2 | 34.43 | 26.48 | 29 | 59 | 4.5 | 15.71 | 14.30 |
| 1959 |  | 29 | 23 | 0724.5 | 37.39 | 23.81 | 61 | 49 | 4.6 | 11.64 | 19.52 |
| 1959 | APR | 08 | 19 | 0237.3 | 36.57 | 26.80 | 160 | ATB | 4.7 |  |  |
| 1959 |  | 19 | 17 | 3904.4 | 37.37 | 20.94 | 87 | 66 | 5.0 | 2.39 | 21.66 |
| 1959 |  | 25 |  | 2646.5 | 37.03 | 28.57 | 35 | 149 | 5.9 | 10.31 | 43.06 |
| 1959 |  | 25 |  | 0547.7 | 37.00 | 28.59 | 35 | 94 | 5.3 | 5.63 | 9.08 |
| 1959 |  | 30 | 22 | 4439.4 | 36.22 | 26.68 | 100 | ATB | 4.8 |  |  |
| 1959 | MAY | 14 | 00 | 5558.0 | 39.95 | 22.89 | 9 | 40 | 4.6 | 35.06 | 290.72 |
| 1959 |  | 14 | 06 | 2711.6 | 35.28 | 24.54 | 73 | 50 | 4.6 | 14.25 | 348.42 |
| 1959 |  | 14 |  | 3659.3 | 35.11 | 24.65 | 23 | 200 | 6.1 | 7.19 | 117.66 |
| 1959 |  | 14 |  | 2232.6 | 40.17 | 23.44 | 97 | 48 | 4.9 | 36.41 | 349.14 |
| 1959 |  | 20 | 16 | 3701.6 | 36.81 | 26.53 | 62 | 42 | 4.8 | 12.05 | 292.55 |
| 1959 | JUN | 09 | 11 | 2119.6 | 36.81 | 29.08 | 20 | ATB | 4.7 |  |  |
| 1959 |  | 10 |  | 1609.0 | 35.67 | 23.57 | 37 | 130 | 5.1 | 13.01 | 0.00 |
| 1959 | JUL | 12 | 16 | 5231.5 | 36.03 | 26.28 | 80 | ATB | 5.1 |  |  |
| 1959 |  | 26 | 17 | 0706.5 | 40.94 | 27.60 | 9 | 77 | 5.4 | 10.08 | 50.49 |
| 1959 | AUG | 11 | 23 | 2811.6 | 41.31 | 23.02 | 43 | 35 | 4.2 | 11.64 | 121.05 |
| 1959 |  | 16 | 18 | 4209.5 | 37.23 | 22.38 | 63 | 72 | 5.1 | 18.14 | 17.73 |
| 1959 |  | 17 | 01 | 3318.2 | 40.97 | 19.73 | 15 | 147 | 5.8 | 13.56 | 313.83 |
| 1959 |  | 17 | 04 | 2908.4 | 40.92 | 19.69 | 46 | 74 | 4.8 | 38.66 | 10.12 |
| 1959 |  | 18 |  | 0405.6 | 41.00 | 19.73 | 7 | 63 | 4.6 | 7.36 | 232.87 |
| 1959 | SEP | 01 | 11 | 3746.3 | 40.95 | 19.63 | 26 | 175 | 6.1 | 12.44 | 286.97 |
| 1959 |  | 03 | 04 | 0207.4 | 40.78 | 19.67 | 22 | 58 | 4.5 | 4.74 | 314.73 |
| 1959 |  | 16 | 05 | 1358.0 | 35.03 | 25.76 | 55 | 85 | 5.5 | 29.51 | 237.83 |
| 1959 | OCT | 05 | 20 | 3410.3 | 40.90 | 19.67 | 27 | 101 | 5.2 | 16.01 | 274.24 |
| 1959 |  | 07 | 08 | 3046.7 | 40.97 | 19.77 | 28 | 161 | 5.7 | 10.77 | 297.84 |
| 1959 | Nov | 06 | 07 | 3723.4 | 41.94 | 20.21 | 17 | 48 | 4.6 | 59.67 | 312.39 |
| 1959 |  | 15 | 17 | 0847.6 | 37.78 | 20.46 | 20 | 269 | 6.9 | 5.62 | 188.98 |
| 1959 |  | 19 |  | 0032.0 | 38.98 | 26.55 |  | 93 | 5.3 | 8.51 | 336.05 |


| DATE |  |  | $\underset{\text { GMT }}{\substack{\text { ORIG. TIME } \\ \text { GMT }}}$ |  | $\stackrel{\text { LAT }}{\mathrm{N}}$ | $\begin{gathered} \text { LON } \\ \text { E } \end{gathered}$ | $\begin{aligned} & \text { DEPTH } \\ & \text { KM } \end{aligned}$ | OBS | $\begin{gathered} \text { MAG } \\ \text { MS } \end{gathered}$ | DIST KM | FT <br> AZIM DEG |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1959 | NOV | 27 | 00 | 2226.9 | 37.74 | 20.14 | 15 | 92 | 5.4 | 3.77 | 332.21 |
| 1959 |  | 27 | 00 | 2615.3 | 37.80 | 20.16 | 10 | 41 | 5.6 | 4.15 | 122.33 |
| 1959 |  | 29 | 23 | 4948.3 | 36.05 | 23.54 | 16 | 35 | 4.5 | 27.03 | 282.07 |
| 1959 | DEC | 01 | 12 | 3850.9 | 37.83 | 20.13 | 18 | 130 | 5.4 | 53.16 | 350.42 |
| 1959 |  | 08 | 09 | 3518.7 | 36.95 | 29.00 | 53 | 31 | 5.0 | 5.55 | 233.10 |
| 1959 |  | 27 | 05 | 2251.3 | 35.16 | 25.95 | 83 | 61 | 5.0 | 39.35 | 322.70 |
| 1960 | Jan | 09 | 03 | 5855.2 | 37.07 | 28.90 | 49 | 80 | 4.9 | 21.21 | 345.05 |
| 1960 |  | 26 | 13 | 0545.5 | 37.00 | 28.93 | 72 | 69 | 5.2 | 18.92 | 336.84 |
| 1960 | FE | 01 | 11 | 5947.2 | 35.27 | 22.90 | 35 | 97 | 5.4 | 8.24 | 262.24 |
| 1960 |  | 22 | 21 | 0425.7 | 39.13 | 20.62 | 57 | 41 | 4.6 | 23.48 | 15.25 |
| 1960 |  | 23 | 00 | 3108.3 | 39.13 | 20.57 | 51 | 47 | 4.7 | 11.84 | 355.55 |
| 1960 |  | 23 | 07 | 3438.1 | 39.09 | 20.66 | 38 | 100 | 5.4 | 14.42 | 35.04 |
| 1960 |  | 23 | 07 | 4758.1 | 38.94 | 20.86 | 42 | 68 | 5.0 | 23.98 | 67.65 |
| 1960 | MAR | 12 | 11 | 5405.9 | 41.91 | 21.00 | 24 | 146 | 5.8 | 11.43 | 29.26 |
| 1960 | APR | 10 | 22 | 0533.5 | 37.79 | 27.67 | 27 | 68 | 4.8 | 14.59 | 324.48 |
| 1960 |  | 12 | 04 | 2244.9 | 37.80 | 27.60 | 13 | 50 | 4.6 | 18.66 | 303.75 |
| 1960 |  | 25 | 16 | 2840.6 | 38.60 | 25.34 | 12 | 58 | 4.9 | 14.25 | 7.45 |
| 1960 |  | 28 | 16 | 3328.2 | 34.51 | 26.45 | 46 | 52 | 5.4 | 29.26 | 354.54 |
| 1960 |  | 30 | 10 | 1248.7 | 35.98 | 26.34 | 100 | ATB | 4.7 |  |  |
| 1960 | MAY | 21 | 06 | 4119.4 | 37.72 | 20.21 | 26 | 66 | 4.8 | 8.93 | 354.35 |
| 1960 |  | 26 | 05 | 1016.6 | 40.56 | 20.63 | 20 | 194 | 6.4 | 8.11 | 46.76 |
| 1960 | JuN | 09 | 09 | 2408.5 | 40.44 | 20.00 | 26 | 60 | 4.8 | 11.43 | 18.76 |
| 1960 |  | 18 | 02 | 0418.7 | 34.42 | 26.21 | 52 | 40 | 4.6 | 18.27 | 28.04 |
| 1960 | JUL | 09 | 22 | 4257.4 | 40.74 | 20.71 | 32 | 50 | 4.7 | 15.24 | 23.92 |
| 1960 |  | 13 | 13 | 0107.1 | 40.61 | 23.45 | 24 | 128 | 5.4 | 8.74 | 50.54 |
| 1960 | AUG | 08 | 20 | 3625.4 | 35.42 | 27.20 | 63 | 62 | 4.8 | 15.55 | 334.49 |
| 1960 |  | 27 | 10 | 1723.2 | 34.58 | 26.19 | 50 | 84 | 5.1 | 16.46 | 290.73 |
| 1960 |  | 29 | 18 | 0039.6 | 34.39 | 26.23 | 92 | 44 | 4.6 | 48.49 | 323.74 |
| 1960 | SEP | 10 | 00 | 1914.7 | 34.58 | 26.23 | 17 | 88 | 5.2 | 16.31 | 328.20 |
| 1960 | OCT | 01 | 05 | 3046.1 | 35.38 | 25.97 | 77 | 77 | 5.1 | 16.46 | 336.20 |
| 1960 | NOV | 05 | 20 | 2053.8 | 39.12 | 20.63 | 22 | 156 | 5.8 | 4.22 | 37.81 |
| 1960 |  | 11 | 05 | 3133.5 | 38.84 | 20.81 | 31 | 103 | 5.2 | 7.83 | 6.35 |
| 1960 |  | 18 | 06 | 0357.8 | 35.17 | 27.83 | 199 | 53 | 4.7 | 27.91 | 263.14 |
| 1960 | DEC | 29 | 18 | 1940.5 | 34.86 | 22.17 | 27 | 59 | 4.9 | 33.58 | 254.59 |
| 1961 | JAN | 07 | 10 | 3057.6 | 35.53 | 26.09 | 84 | 88 | 5.5 | 12.44 | 320.75 |
| 1961 |  | 07 | 15 | 5259.9 | 37.70 | 21.14 | 44 | 76 | 4.7 | 10.77 | 158.33 |
| 1961 |  | 28 | 07 | 1817 | 39.4 | 22.0 | 33 | UNS | 4.9 |  |  |
| 1961 | FEB | 16 | 03 | 4446.9 | 40.22 | 19.87 | 37 | 57 | 4.5 | 20.75 | 187.28 |
| 1961 |  | 21 | 03 | 0201.3 | 36.50 | 22.91 | 60 | 73 | 4.7 | 21.10 | 336.86 |
| 1961 |  | 23 | 21 | 4555.4 | 36.75 | 27.02 | 42 | 59 | 5.1 | 6.23 | 270.00 |
| 1961 |  | 23 | 21 | 5653.3 | 36.83 | 27.07 | 32 | 38 | 4.8 | 36.54 | 297.48 |
| 1961 |  | 27 | 21 | 4008.2 | 36.68 | 26.95 | 64 | 55 | 5.0 | 32.55 | 358.41 |
| 1961 |  | 27 | 21 | 5438.9 | 36.66 | 26.95 | 48 | 46 | 5.1 | 20.63 | 279.67 |
| 1961 | MAR | 13 | 15 | 3201.8 | 36.21 | 26.43 | 10 | ATB | 4.7 |  |  |
| 1961 |  | 13 | 19 | 1719.3 | 34.57 | 26.58 | 25 | 97 | 5.0 | 15.55 | 325.91 |
| 1961 | MAY | 23 | 02 | 4524.1 | 36.82 | 28.40 | 74 | 212 | 6.4 | 14.25 | 336.32 |
| 1961 |  | 25 | 13 | 1147.9 | 36.72 | 26.66 | 60 | ATB | 4.8 |  |  |
| 1961 | JUN | 21 | 16 | 0452.5 | 37.93 | 28.82 | 64 | 60 | 5.3 | 29.02 | 300.44 |
| 1961 | JUL | 12 | 02 | 4834.7 | 39.26 | 23.79 | 38 | 45 | 4.6 | 32.70 | 58.84 |
| 1961 |  | 19 | 23 | 0058.9 | 37.71 | 20.18 | 22 | 96 | 5.2 | 5.28 | 90.01 |
| 1961 |  | 27 | 18 | 3544.8 | 34.83 | 25.16 | 17 | 54 | 4.5 | 22.75 | 4.71 |
| 1961 | AUG | 27 | 22 | 0851.9 | 35.67 | 23.41 | 60 | 99 | 5.0 | 10.08 | 360.00 |
| 1961 | SEP | 18 | 05 | 0835.4 | 34.48 | 25.82 | 36 | 47 | 4.4 | 17.04 | 307.85 |
| 1961 |  | 02 | 07 | 2145.1 | 36.66 | 21.86 | 19 | 138 | 5.4 | 4.79 | 338.14 |
| 1961 | NOV | 28 | 08 | 5847.9 | 40.21 | 26.30 | 76 | 53 | 5.2 | 71.65 | 11.69 |


| DATE |  |  | ORIG.TIME GMT |  |  | $\begin{gathered} \text { LAT } \\ \mathrm{N} \end{gathered}$ | $\begin{gathered} \text { LON } \\ \text { E } \end{gathered}$ | DEPTH KM | OBS | MAGMS | SHIFT |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  |  |  |  | ST KM |  |  |  |  | AZIM DEG |
| 1961 | DEC | 10 | 08 | 39 | 11.6 |  | 34.56 | 25.60 | 51 | 57 | 4.7 | 40.32 | 1.32 |
| 1961 |  | 11 |  | 53 | 13.7 | 36.42 | 23.44 | 75 | 80 | 4.9 | 6.65 | 70.45 |
| 1962 | JAN | 10 | 12 | 36 | 35.7 | 35.82 | 22.48 | 77 | 48 | 5.0 | 5.61 | 233.50 |
| 1962 |  | 19 | 19 | 38 | 02.7 | 38.35 | 22.25 | 35 | 90 | 5.3 | 22.51 | 210.50 |
| 1962 |  | 19 | 22 | 18 | 28.0 | 38.38 | 22.15 | 40 | 63 | 4.9 | 34.21 | 10.41 |
| 1962 |  | 26 | 08 | 17 | 39.8 | 35.32 | 22.78 | 11 | 198 | 5.8 | 11.21 | 27.12 |
| 1962 | MAR | 18 | 15 | 30 | 33.0 | 40.69 | 19.59 | 15 | 187 | 6.0 | 5.18 | 257.59 |
| 1962 |  | 26 | 09 | 22 | 12.6 | 38.44 | 20.86 | 67 | 93 | 5.1 | 49.62 | 22.87 |
| 1962 | APR | 04 | 20 | 51 | 05 | 34.6 | 25.5 | 20 | UNS | 4.7 |  |  |
| 1962 |  | 04 | 20 | 59 | 41.5 | 34.41 | 25.16 | 47 | 87 | 4.9 | 17.73 | 15.44 |
| 1962 |  | 10 | 21 | 37 | 10.3 | 37.76 | 20.09 | 5 | 196 | 6.3 | 3.77 | 207.78 |
| 1962 |  | 10 |  | 10 | 50.0 | 37.79 | 20.13 | 11 | 60 | 5.1 | 12.24 | 218.45 |
| 1962 |  | 11 |  | 47 | 29.3 | 37.65 | 20.16 | 6 | 146 | 5.4 | 4.93 | 116.81 |
| 1962 |  | 16 | 00 | 15 | 16.9 | 37.70 | 20.12 | 46 | 81 | 4.8 | 46.56 | 256.29 |
| 1962 |  | 16 |  | 19 | 06.0 | 36.15 | 27.23 | 140 | ATB | 5.2 |  |  |
| 1962 |  | 17 |  | 15 | 30 | 37.60 | 20.10 | 25 | UNS | 4.8 |  |  |
| 1962 |  | 17 | 11 | 33 | 55.2 | 37.78 | 20.09 | 33 | 85 | 5.2 | 14.92 | 61.59 |
| 1962 |  | 19 | 02 | 05 | 57.2 | 37.74 | 20.26 | 42 | 65 | 4.5 | 16.16 | 30.77 |
| 1962 |  | 28 |  | 19 | 02.9 | 36.20 | 26.84 | 56 | 173 | 6.0 | 16.89 | 16.14 |
| 1962 |  | 28 | 12 | 43 | 52.0 | 36.13 | 26.87 | 51 | 151 | 5.5 | 2.11 | 58.26 |
| 1962 | MAY | 01 | 11 | 53 | 59 | 38.20 | 20.50 | 90 | UNS | 4.9 |  |  |
| 1962 |  | 08 | 23 | 53 | 59.0 | 35.23 | 24.12 | 79 | 79 | 4.8 | 6.35 | 314.41 |
| 1962 | JUN | 28 | 06 | 51 | 07.7 | 40.76 | 20.69 | 40 | 95 | 5.0 | 7.13 | 200.74 |
| 1962 | JUL | 06 | 09 | 16 | 17.1 | 37.79 | 20.12 | 37 | 168 | 5.8 | 3.68 | 287.55 |
| 1962 |  | 06 | 15 | 54 | 25.9 | 37.79 | 20.31 | 6 | 56 | 4.4 | 6.25 | 79.76 |
| 1962 |  | 10 | 10 | 06 | 02.4 | 38.42 | 25.92 | 10 | ATB | 4.8 |  |  |
| 1962 | AUG | 28 | 10 | 59 | 57.4 | 37.80 | 22.88 | 95 | 226 | 6.6 | 2.39 | 201.57 |
| 1962 | SEP | 10 | 09 | 36 | 28.7 | 34.74 | 26.62 | 45 | 156 | 5.4 | 13.01 | 337.57 |
| 1962 | OCT | 04 | 19 | 46 | 12.1 | 37.93 | 22.36 | 53 | 123 | 5.0 | 6.72 | 7.49 |
| 1962 |  | 26 | 11 | 26 | 11.8 | 33.52 | 27.55 | 11 | 102 | 5.5 | 7.29 | 242.80 |
| 1962 | DEC | 13 | 22 | 45 | 34.7 | 35.22 | 27.97 | 65 | 59 | 4.7 | 35.06 | 7.54 |
| 1963 | JAN | 31 | 15 | 07 | 06.3 | 35.94 | 22.02 | 52 | 64 | 4.8 | 13.01 | 79.25 |
| 1963 | FEB | 15 | 10 | 18 | 25.7 | 40.15 | 19.89 | 31 | 64 | 4.6 | 27.91 | 143.89 |
| 1963 |  | 22 | 14 | 12 | 54.8 | 40.34 | 20.11 | 30 | 104 | 5.1 | 12.24 | 125.46 |
| 1963 | MAR | 04 | 15 | 10 | 20.8 | 34.96 | 25.18 | 39 | 134 | 5.3 | 5.85 | 18.15 |
| 1963 |  | 11 |  | 27 | 24.3 | 37.96 | 29.14 | 40 | 74 | 5.5 | 3.76 | 152.27 |
| 1963 |  | 17 | 14 | 17 | 24.0 | 39.35 | 20.89 | 41 | 94 | 5.1 | 8.19 | 132.73 |
| 1963 |  | 29 | 03 | 09 | 14.6 | 40.54 | 26.46 | 2 | 69 | 5.1 | 31.41 | 41.05 |
| 1963 | APR | 28 | 00 | 41 | 52.2 | 39.32 | 27.82 | 30 | ATB | 4.6 |  |  |
| 1963 | MAY | 06 | 19 | 30 | 32.6 | 39.32 | 20.34 | 53 | 60 | 4.5 | 48.04 | 263.50 |
| 1963 |  | 15 | 11 | 15 | 42.3 | 41.56 | 20.13 | 49 | 73 | 4.5 | 19.30 | 161.81 |
| 1963 | JUN | 04 | 22 | 11 | 35.0 | 38.92 | 20.54 | 46 | 90 | 5.0 | 2.82 | 113.18 |
| 1963 | JUL | 08 | 16 | 02 | 35.4 | 36.63 | 27.82 | 78 | 80 | 4.7 | 27.03 | 315.16 |
| 1963 |  | 10 | 07 | 19 | 41.3 | 39.80 | 23.90 | 127 | ROT | 4.6 |  |  |
| 1963 |  | 26 | 04 | 17 | 16.1 | 42.04 | 21.43 | 14 | 192 | 6.1 | 4.12 | 90.00 |
| 1963 |  | 26 | 19 | 46 | 39.1 | 36.84 | 28.76 | 80 | ATB | 5.0 |  |  |
| 1963 | SEP | 18 | 16 | 58 | 13.5 | 40.71 | 29.02 | 48 | 184 | 6.3 | 14.42 | 222.94 |
| 1963 |  | 24 | 02 | 10 | 45.8 | 40.91 | 28.90 | 11 | 71 | 4.8 | 6.88 | 14.14 |
| 1963 |  | 29 | 13 | 35 | 49.2 | 36.44 | 29.00 | 60 | ATB | 4.8 |  |  |
| 1963 | OCT | 02 | 21 | 05 | 16.5 | 35.02 | 23.49 | 58 | 67 | 4.5 | 21.88 | 7.40 |
| 1963 | NOV | 12 | 07 | 06 | 35.8 | 35.48 | 29.61 | 83 | 97 | 5.1 | 12.44 | 346.21 |
| 1963 | DEC | 16 | 13 | 47 | 57.4 | 36.97 | 20.96 | 15 | 162 | 5.8 | 10.99 | 214.52 |
| 1964 | JAN | 10 | 03 | 34 | 24.0 | 38.70 | 21.00 | 40 | 19 | 4.6 |  |  |
| 1964 |  | 30 | 17 | 45 | 57.0 | 37.41 | 29.89 | 59 | 123 | 5.5 |  |  |
| 1964 |  | 31 | 09 | 23 | 15.5 | 37.68 | 22.51 | 13 | 17 | 4.0 |  |  |


| DATE |  |  | ORIG.TIME GMT |  |  | $\stackrel{\text { LAT }}{\mathrm{N}}$ | $\begin{gathered} \mathrm{LON} \\ \mathrm{E} \end{gathered}$ | $\begin{aligned} & \text { DEPTH } \\ & \text { KM } \end{aligned}$ | OBS | $\begin{array}{r} \text { MAG } \\ \text { MS } \end{array}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1964 | FEB | 18 | 08 | 35 | 15.0 | 37.80 | 23.70 | 52 | 11 | 4.4 |
| 1964 |  | 23 | 22 | 41 | 03.9 | 39.21 | 23.73 | 10 | 131 | 5.4 |
| 1964 |  | 24 | 23 | 21 | 13.0 | 38.10 | 24.00 | 16 | 14 | 4.6 |
| 1964 |  | 24 | 23 | 30 | 28.0 | 39.09 | 23.80 | 41 | 59 | 4.7 |
| 1964 | MAR | 30 | 03 | 27 | 13.3 | 34.94 | 24.02 | 65 | 54 | 4.6 |
| 1964 |  | 31 | 09 | 33 | 12.3 | 36.43 | 28.78 | 57 | 34 | 4.5 |
| 1964 | APR | 08 | 14 | 12 | 28.5 | 35.04 | 24.29 | 64 | 119 | 5.1 |
| 1964 |  | 11 | 16 | 00 | 43.0 | 40.30 | 24.83 | 33 | 136 | 5.6 |
| 1964 |  | 15 | 20 | 54 | 27.4 | 39.04 | 23.71 | 44 | 52 | 4.4 |
| 1964 |  | 17 | 18 | 11 | 39.6 | 38.20 | 20.30 | 17 | 25 | 4.4 |
| 1964 |  | 20 | 18 | 37 | 32.6 | 35.09 | 24.46 | 68 | 30 | 4.4 |
| 1964 |  | 25 | 12 | 44 | 15.0 | 35.46 | 27.70 | 61 | 67 | 4.7 |
| 1964 |  | 29 | 04 | 21 | 05.1 | 39.25 | 23.72 | 20 | 137 | 5.5 |
| 1964 |  | 29 | 17 | 00 | 01.3 | 39.14 | 23.55 | 15 | 99 | 5.0 |
| 1964 | MAY | 13 | 17 | 06 | 14.8 | 36.28 | 28.21 | 82 | 27 | 4.4 |
| 1964 |  | 14 | 17 | 00 | 27.8 | 38.52 | 20.44 | 36 | 17 | 5.0 |
| 1964 |  | 18 | 20 | 03 | 14.2 | 36.95 | 24.29 | 109 | 13 | 4.3 |
| 1964 | JUN | 08 | 16 | 49 | 03.5 | 36.26 | 28.26 | 62 | 28 | 4.4 |
| 1964 |  | 12 | 07 | 46 | 21.0 | 37.34 | 29.93 | 5 | 27 | 4.6 |
| 1964 | JUL | 04 | 11 | 11 | 17.9 | 41.96 | 23.43 | 2 | 66 | 4.8 |
| 1964 |  | 17 | 02 | 34 | 26.7 | 38.05 | 23.63 | 155 | 241 | 6.0 |
| 1964 |  | 18 | 03 | 40 | 19.4 | 36.13 | 26.01 | 99 | 139 | 5.2 |
| 1964 | AUG | 17 | 00 | 17 | 48.5 | 35.28 | 25.90 | 64 | 95 | 4.6 |
| 1964 |  | 24 | 21 | 42 | 46.2 | 40.51 | 19.20 | 41 | 23 | 5.0 |
| 1964 |  | 25 | 11 | 11 | 52.0 | 35.75 | 28.84 | 51 | 125 | 5.3 |
| 1964 |  | 25 | 07 | 08 | 14.9 | 35.50 | 28.76 | 28 | 31 | 4.5 |
| 1964 |  | 25 | 08 | 05 | 01.8 | 35.28 | 28.67 | 55 | 38 | 4.4 |
| 1964 |  | 25 | 11 | 42 | 55.8 | 35.35 | 28.58 | 12 | 47 | 4.6 |
| 1964 |  | 25 | 14 | 37 | 33.6 | 35.55 | 28.82 | 35 | 70 | 5.6 |
| 1964 |  | 27 | 19 | 31 | 59.7 | 35.56 | 28.84 | 38 | 113 | 5.4 |
| 1964 |  | 28 | 12 | 06 | 18.4 | 37.80 | 19.95 | 61 | 55 | 4.6 |
| 1964 |  | 29 | 19 | 37 | 55.2 | 35.29 | 28.72 | 35 | 42 | 4.7 |
| 1964 |  | 31 | 19 | 35 | 39.0 | 36.10 | 20.20 | 21 | 21 | 4.9 |
| 1964 | SEP | 13 | 22 | 53 | 22.9 | 41.71 | 20.60 | 00 | 21 | 4.9 |
| 1964 |  | 18 | 00 | 08 | 47.6 | 35.69 | 29.07 | 40 | 95 | 5.3 |
| 1964 |  | 30 | 04 | 39 | 45.4 | 34.51 | 23.26 | 49 | 92 | 4.9 |
| 1964 | OCT | 06 | 14 | 29 | 57.9 | 40.24 | 28.16 | 23 | 151 | 5.1 |
| 1964 |  | 06 | 14 | 31 | 23.0 | 40.30 | 28.23 | 34 | 210 | 7.0 |
| 1964 |  | 07 | 23 | 07 | 53.9 | 40.19 | 28.36 | 31 | 45 | 4.4 |
| 1964 |  | 13 | 10 | 30 | 09.2 | 36.94 | 28.29 | 76 | 10 | 4.3 |
| 1964 |  | 17 | 09 | 50 | 28.0 | 35.02 | 25.43 | 18 | 117 | 5.0 |
| 1964 |  | 20 | 08 | 47 | 56.0 | 40.00 | 28.60 | 00 | 10 | 4.8 |
| 1964 | NOV | 05 | 20 | 55 | 45.8 | 35.11 | 24.13 | 27 | 40 | 4.7 |
| 1964 | DEC | 01 | 10 | 21 | 03.3 | 38.53 | 22.45 | 48 | 15 | 4.7 |
| 1964 |  | 09 | 18 | 28 | 46.0 | 41.57 | 20.92 | 78 | 66 | 4.4 |
| 1964 |  | 09 | 19 | 06 | 21.4 | 41.20 | 20.92 | 55 | 41 | 4.7 |
| 1964 |  | 15 | 21 | 03 | 15.7 | 40.02 | 28.79 | 26 | 43 | 4.5 |
| 1964 |  | 31 | 16 | 18 | 02.2 | 35.76 | 25.51 | 89 | 110 | 5.2 |
| 1965 | JAN | 02 | 13 | 47 | 43.4 | 36.46 | 26.10 | 59 | 12 | 4.9 |
| 1965 |  | 07 | 10 | 22 | 17.2 | 36.50 | 26.85 | 35 | 79 | 4.7 |
| 1965 |  | 09 | 04 | 11 | 51.0 | 36.00 | 27.40 | 63 | 17 | 4.3 |
| 1965 |  | 10 | 08 | 02 | 51.7 | 38.70 | 22.00 | 46 | 18 | 4.3 |
| 1965 |  | 17 | 03 | 39 | 32.5 | 34.58 | 27.83 | 29 | 33 | 4.6 |
| 1965 |  | 29 | 23 | 39 | 02.5 | 34.91 | 27.60 | 25 | 48 | 4.7 |
| 1965 | FEB | 06 | 03 | 47 | 57.8 | 35.41 | 27.04 | 71 | 59 | 4.7 |


| DATE |  |  | $\begin{aligned} & \text { ORIG.TIME } \\ & \text { GMT } \end{aligned}$ |  |  | $\begin{gathered} \text { LAT } \\ \mathrm{N} \end{gathered}$ | $\begin{gathered} \text { LON } \\ \text { E } \end{gathered}$ | $\begin{gathered} \text { DEPTH } \\ \text { KM } \end{gathered}$ | OBS | $\begin{array}{r} \text { MAG } \\ \text { MS } \end{array}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1965 | FEB | 09 | 20 | 38 | 41.7 | 37.92 | 20.25 | 08 | 54 | 4.8 |
| 1965 |  | 09 | 23 | 32 | 56.4 | 37.79 | 20.85 | 23 | 28 | 4.9 |
| 1965 |  | 20 | 22 | 47 | 08.6 | 38.40 | 22.10 | 06 | 21 | 4.3 |
| 1965 | MAR | 02 | 22 | 00 | 07.2 | 38.47 | 28.33 | 42 | 115 | 5.6 |
| 1965 |  | 03 | 01 | 37 | 18.3 | 38.27 | 28.47 | 42 | 18 | 4.5 |
| 1965 |  | 08 | 23 | 02 | 55.9 | 37.85 | 21.00 | 55 | 24 | 4.6 |
| 1965 |  | 09 | 17 | 57 | 54.5 | 39.34 | 23.82 | 18 | 215 | 6.3 |
| 1965 |  | 09 | 19 | 46 | 58.7 | 39.12 | 23.86 | 19 | 66 | 5.0 |
| 1965 |  | 09 | 21 | 20 | 04.5 | 39.19 | 23.87 | 07 | 50 | 4.8 |
| 1965 |  | 09 | 18 | 37 | 54.6 | 39.28 | 23.93 | 33 | 85 | 5.0 |
| 1965 |  | 09 | 22 | 19 | 06.4 | 39.17 | 23.96 | 13 | 30 | 4.6 |
| 1965 |  | 09 | 22 | 35 | 15.3 | 39.26 | 23.84 | 18 | 60 | 4.9 |
| 1965 |  | 10 | 01 | 36 | 05.8 | 39.08 | 23.77 | 18 | 101 | 5.1 |
| 1965 |  | 10 | 21 | 50 | 19.8 | 39.35 | 23.94 | 37 | 34 | 4.4 |
| 1965 |  | 13 | 04 | 08 | 40.6 | 39.11 | 23.97 | 11 | 62 | 4.9 |
| 1965 |  | 13 | 04 | 09 | 37.9 | 39.03 | 23.68 | 33 | 78 | 5.3 |
| 1965 |  | 13 | 15 | 42 | 16.5 | 39.14 | 23.90 | 18 | 23 | 4.6 |
| 1965 |  | 14 | 06 | 04 | 49.3 | 39.90 | 20.20 | 05 | 9 | 4.5 |
| 1965 |  | 15 | 23 | 08 | 30.9 | 39.16 | 24.00 | 33 | 24 | 4.7 |
| 1965 |  | 19 | 04 | 35 | 45.4 | 41.50 | 23.10 | 12 | 23 | 4.4 |
| 1965 |  | 19 | 23 | 37 | 31.9 | 41.39 | 22.88 | 33 | 17 | 4.4 |
| 1965 |  | 22 | 03 | 22 | 22.2 | 39.13 | 23.84 | 01 | 38 | 4.6 |
| 1965 |  | 31 | 09 | 47 | 26.3 | 38.38 | 22.26 | 45 | 279 | 6.6 |
| 1965 |  | 31 | 12 | 01 | 11.7 | 38.47 | 22.23 | 78 | 27 | 4.9 |
| 1965 |  | 31 | 20 | 08 | 25.5 | 39.20 | 24.10 | 33 | 32 | 4.7 |
| 1965 | APR | 03 | 05 | 19 | 18.0 | 37.70 | 23.80 | 15 | 9 | 4.3 |
| 1965 |  | 03 | 14 | 30 | 48.2 | 38.24 | 20.50 | 25 | 66 | 4.6 |
| 1965 |  | 05 | 03 | 12 | 54.6 | 37.75 | 22.00 | 34 | 218 | 6.0 |
| 1965 |  | 07 | 04 | 16 | 39.6 | 37.10 | 22.30 | 36 | 25 | 4.3 |
| 1965 |  | 09 | 23 | 57 | 02.0 | 35.06 | 24.31 | 39 | 210 | 6.1 |
| 1965 |  | 10 | 00 | 19 | 59.7 | 34.90 | 24.37 | 55 | 47 | 4.7 |
| 1965 |  | 19 | 06 | 46 | 33.7 | 34.56 | 28.36 | 33 | 24 | 4.4 |
| 1965 |  | 27 | 14 | 09 | 05.6 | 35.63 | 23.53 | 37 | 182 | 5.5 |
| 1965 |  | 29 | 09 | 46 | 56.8 | 37.14 | 26.89 | 08 | 56 | 4.8 |
| 1965 | MAY | 01 | 01 | 59 | 43.9 | 37.18 | 26.91 | 15 | 48 | 4.7 |
| 1965 |  | 02 | 22 | 33 | 25.4 | 35.61 | 23.52 | 56 | 41 | 4.6 |
| 1965 |  | 07 | 14 | 42 | 21.7 | 36.74 | 26.86 | 162 | 30 | 4.4 |
| 1965 |  | 13 | 21 | 09 | 16.7 | 39.22 | 20.73 | 58 | 20 | 4.4 |
| 1965 |  | 16 | 01 | 35 | 56.0 | 35.26 | 27.85 | 41 | 80 | 4.6 |
| 1965 |  | 29 | 01 | 47 | 48.0 | 35.13 | 22.64 | 56 | 33 | 4.6 |
| 1965 |  | 29 | 04 | 14 | 56.1 | 35.19 | 22.57 | 43 | 69 | 4.7 |
| 1965 | JUN | 03 | 18 | 31 | 51.0 | 39.72 | 23.21 | 33 | 107 | 4.8 |
| 1965 |  | 10 | 15 | 24 | 17.1 | 36.44 | 26.64 | 142 | 50 | 4.5 |
| 1965 |  | 29 | 15 | 40 | 31.5 | 34.20 | 26.23 | 33 | 67 | 4.6 |
| 1965 | JUL | 06 | 03 | 18 | 42.1 | 38.37 | 22.40 | 18 | 244 | 6.4 |
| 1965 |  | 06 | 13 | 34 | 14.8 | 34.73 | 25.64 | 61 | 37 | 4.7 |
| 1965 |  | 10 | 08 | 09 | 46.1 | 34.73 | 23.30 | 07 | 30 | 4.3 |
| 1965 |  | 13 | 14 | 19 | 01.0 | 37.50 | 27.80 | 35 | 22 | 4.4 |
| 1965 | AUG | 04 | 19 | 15 | 04.6 | 35.30 | 26.50 | 52 | 35 | 4.6 |
| 1965 |  | 14 | 04 | 47 | 51.7 | 38.45 | 21.60 | 30 | 24 | 4.3 |
| 1965 |  | 23 | 14 | 08 | 58.6 | 40.51 | 26.17 | 33 | 170 | 6.1 |
| 1965 |  | 24 | 01 | 11 | 07.2 | 35.67 | 23.50 | 54 | 56 | 4.7 |
| 1965 |  | 24 | 23 | 57 | 35.4 | 40.39 | 26.20 | 18 | 48 | 4.7 |
| 1965 |  | 25 | 04 | 57 | 45.7 | 34.72 | 25.08 | 10 | 95 | 4.8 |
| 1965 | SEP | 11 | 04 | 49 | 12.8 | 39.07 | 22.09 | 42 | 31 | 4.3 |


| DATE |  |  | ORIG.TIME GMT |  |  | $\begin{gathered} \text { LAT } \\ \mathrm{N} \end{gathered}$ | $\begin{gathered} \text { LON } \\ \text { E } \end{gathered}$ | $\begin{gathered} \text { DEPTH } \\ \text { KM } \end{gathered}$ | OBS | $\begin{array}{r} \text { MAG } \\ \text { MS } \end{array}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1965 | OCT | 11 | 06 | 37 | 03.0 | 42.00 | 21.50 | 05 |  | 4.6 |
| 1965 |  | 18 | 14 | 32 | 48.3 | 38.83 | 27.83 | 36 | 36 | 4.6 |
| 1965 |  | 28 | 04 | 27 | 12.9 | 38.41 | 22.37 | 29 | 31 | 4.3 |
| 1965 |  | 28 | 14 |  | 28.5 | 41.67 | 19.30 | 28 | 24 | 4.3 |
| 1965 | NOV | 02 | 03 | 27 | 07.4 | 39.48 | 25.32 | 05 | 94 | 5.0 |
| 1965 |  | 28 | 05 | 26 | 05.3 | 36.12 | 27.43 | 73 | 209 | 5.6 |
| 1965 | DEC | 04 | 16 | 39 | 57.5 | 34.26 | 26.25 | 12 | 68 | 4.5 |
| 1965 |  | 07 | 01 | 00 | 57.0 | 36.30 | 27.00 |  | 4 | 4.5 |
| 1965 |  | 07 | 08 | 25 | 03.2 | 35.60 | 27.50 |  | 4 | 4.8 |
| 1965 |  | 08 | 11 | 22 | 05.0 | 37.30 | 28.50 |  | 7 | 4.7 |
| 1965 |  | 13 | 17 | 44 | 08.4 | 40.25 | 19.82 | 07 | 45 | 4.6 |
| 1965 |  | 20 | 00 | 08 | 16.0 | 40.21 | 24.82 | 33 | 175 | 6.0 |
| 1965 |  | 20 | 00 | 30 | 57.6 | 40.01 | 24.80 | 42 | 31 | 4.7 |
| 1965 |  | 22 | 08 | 43 | 44.5 | 37.10 | 28.10 |  | 9 | 4.7 |
| 1965 |  | 25 | 12 | 15 | 33.1 | 39.84 | 25.00 | 41 | 31 | 4.7 |
| 1965 |  | 25 | 15 | 10 | 30.0 | 37.31 | 21.06 | 04 | 28 | 4.3 |
| 1966 | JAN | 02 | 23 | 12 | 18.0 | 37.67 | 23.18 | 12 | 92 | 4.7 |
| 1966 |  | 14 | 18 | 39 | 31.0 | 34.72 | 27.00 | 22 | 54 | 4.4 |
| 1966 |  | 15 | 18 | 07 | 46.5 | 36.72 | 23.09 | 37 | 60 | 4.7 |
| 1966 |  | 16 | 20 | 15 | 30.0 | 35.61 | 25.80 | 47 | 73 | 4.5 |
| 1966 |  | 17 | 08 | 39 | 42.6 | 40.09 | 20.57 | 46 | 29 | 4.7 |
| 1966 |  | 17 | 20 | 04 | 58.6 | 38.12 | 22.00 | 62 | 47 | 4.4 |
| 1966 |  | 18 | 21 | 20 | 02.6 | 35.12 | 23.49 | 60 | 47 | 4.4 |
| 1966 |  | 20 | 00 | 39 | 00.6 | 39.20 | 24.44 | 12 | 49 | 4.4 |
| 1966 |  | 26 | 13 | 30 | 28.0 | 38.94 | 21.47 | 46 | 36 | 4.6 |
| 1966 |  | 30 | 06 |  | 03.0 | 38.87 | 21.65 | 48 | 43 | 4.6 |
| 1966 |  | 31 | 04 | 30 | 57.0 | 39.05 | 21.90 | 51 | 24 | 4.4 |
| 1966 | FEB | 04 | 08 | 38 | 03.0 | 34.37 | 23.94 | 33 | 94 | 4.7 |
| 1966 |  | 05 | 02 | 01 | 45.3 | 39.10 | 21.74 | 16 | 261 | 6.2 |
| 1966 |  | 05 | 02 | 11 | 08.0 | 39.17 | 21.89 | 21 | 39 | 4.9 |
| 1966 |  | 05 | 02 | 58 | 01.2 | 39.11 | 21.91 | 50 | 130 | 5.0 |
| 1966 |  | 08 | 20 | 08 | 04.0 | 41.08 | 24.97 | 21 | 105 | 4.7 |
| 1966 |  | 08 | 13 | 16 | 22.2 | 36.23 | 28.11 | 79 | 62 | 4.4 |
| 1966 |  | 09 | 05 | 36 | 23.1 | 41.11 | 24.92 | 48 | 14 | 4.3 |
| 1966 |  | 10 | 13 | 21 | 45.9 | 38.95 | 21.70 | 39 | 31 | 4.3 |
| 1966 |  | 11 | 06 |  | 37.0 | 39.15 | 21.45 | 24 | 29 | 4.3 |
| 1966 |  | 12 | 13 | 36 | 22.2 | 38.83 | 21.43 | 46 | 55 | 4.9 |
| 1966 |  | 14 | 17 | 57 | 50.1 | 34.94 | 27.11 | 43 | 115 | 4.8 |
| 1966 |  | 14 | 20 | 16 | 58.0 | 38.82 | 21.42 | 39 | 52 | 4.3 |
| 1966 |  | 17 | 10 | 41 | 25.8 | 38.89 | 21.88 | 38 | 18 | 5.3 |
| 1966 |  | 19 | 10 | 22 | 27.0 | 39.04 | 21.65 | 08 | 16 | 4.4 |
| 1966 | MAR | 08 | 18 | 51 | 47.5 | 38.87 | 21.42 | 44 | 65 | 4.7 |
| 1966 |  | 11 | 20 | 01 | 45.0 | 34.40 | 24.23 | 30 | 126 | 5.0 |
| 1966 |  | 14 | 14 | 08 | 41.2 | 39.07 | 21.36 | 45 | 62 | 4.6 |
| 1966 |  | 29 | 00 | 08 | 42.8 | 37.00 | 26.80 | 33 | 15 | 4.5 |
| 1966 | APR | 01 | 13 | 15 | 05.2 | 38.72 | 21.49 | 45 | 91 | 4.7 |
| 1966 |  | 03 | 11 | 36 | 26.1 | 38.94 | 21.53 | 34 | 134 | 5.2 |
| 1966 |  | 07 | 03 | 25 | 45.0 | 37.83 | 21.14 | 25 | 106 | 4.8 |
| 1966 |  | 11 | 06 | 43 | 46.4 | 35.61 | 27.08 | 00 | 12 | 4.6 |
| 1966 |  | 13 | 20 | 44 | 08.7 | 36.80 | 28.50 | 00 | 3 | 4.8 |
| 1966 |  | 14 | 18 | 51 | 44.0 | 34.55 | 23.86 | 14 | 83 | 4.8 |
| 1966 |  | 21 | 06 | 45 | 26.9 | 34.49 | 25.69 | 51 | 99 | 5.1 |
| 1966 | MAY | 04 | 06 | 36 | 59.0 | 38.94 | 21.47 | 27 | 141 | 5.2 |
| 1966 |  | 04 | 21 | 49 | 01.8 | 37.74 | 27.71 | 37 | 126 | 5.2 |
| 1966 |  | 07 | 13 | 08 | 16.9 | 37.75 | 27.79 | 09 | 146 | 5.3 |


| DATE |  |  | $\begin{aligned} & \text { ORIG.TIME } \\ & \text { GMT } \end{aligned}$ |  |  | $\begin{gathered} \text { LAT } \\ \mathrm{N} \end{gathered}$ | $\begin{gathered} \text { LON } \\ \text { E } \end{gathered}$ | $\begin{gathered} \text { DEPTH } \\ \text { KM } \end{gathered}$ | OBS | $\begin{array}{r} \text { MAG } \\ \text { MS } \end{array}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1966 | MAY | 09 | 00 | 42 | 53.0 | 34.43 | 26.44 | 13 | 206 | 5.9 |
| 1966 |  | 09 | 06 | 08 | 29.6 | 34.31 | 26.44 | 43 | 88 | 4.7 |
| 1966 |  | 10 | 18 | 44 | 38.0 | 36.50 | 27.40 | 94 | 9 | 4.3 |
| 1966 |  | 11 | 10 | 21 | 41.8 | 34.30 | 26.40 | 06 | 22 | 4.7 |
| 1966 |  | 11 | 15 | 06 | 02.5 | 34.37 | 26.42 | 39 | 92 | 4.7 |
| 1966 |  | 12 | 20 | 31 | 02.5 | 38.56 | 25.82 | 33 | 54 | 4.5 |
| 1966 |  | 13 | 13 | 11 | 50.9 | 34.47 | 26.47 | 37 | 86 | 4.6 |
| 1966 |  | 14 | 23 | 00 | 44.7 | 37.00 | 22.02 | 40 | 71 | 4.5 |
| 1966 |  | 15 | 10 | 11 | 08.0 | 35.17 | 27.16 | 34 | 61 | 4.4 |
| 1966 |  | 16 | 17 | 30 | 56.1 | 34.48 | 26.46 | 41 | 91 | 4.6 |
| 1966 |  | 22 | 07 | 37 | 29.0 | 38.70 | 27.92 | 23 | 57 | 4.8 |
| 1966 |  | 24 | 09 | 39 | 26.5 | 37.33 | 21.89 | 34 | 128 | 4.9 |
| 1966 |  | 24 | 11 | 09 | 25.4 | 37.37 | 22.02 | 43 | 99 | 4.9 |
| 1966 |  | 24 | 17 | 43 | 32.3 | 34.87 | 24.62 | 43 | 63 | 4.7 |
| 1966 |  | 25 | 09 | 06 | 57.0 | 40.32 | 19.82 | 21 | 82 | 4.7 |
| 1966 | JUN | 02 | 22 | 51 | 28.0 | 38.50 | 27.23 | 30 | 43 | 4.5 |
| 1966 |  | 04 | 06 | 16 | 57.5 | 36.63 | 20.97 | 82 | 125 | 4.9 |
| 1966 |  | 05 | 20 | 52 | 02.5 | 37.24 | 21.94 | 35 | 47 | 4.3 |
| 1966 |  | 11 | 10 | 21 | 55.4 | 38.84 | 21.50 | 43 | 109 | 4.8 |
| 1966 |  | 11 | 12 | 05 | 02.7 | 37.37 | 21.08 | 47 | 91 | 4.7 |
| 1966 |  | 13 | 04 | 59 | 24.0 | 38.30 | 28.50 | 00 | 10 | 4.7 |
| 1966 |  | 19 | 17 | 55 | 30.0 | 38.55 | 27.35 | 09 | 82 | 4.8 |
| 1966 |  | 24 | 22 | 34 | 26.1 | 38.73 | 21.53 | 34 | 99 | 4.8 |
| 1966 |  | 25 | 06 | 20 | 46.9 | 38.54 | 26.90 | 00 | 8 | 4.4 |
| 1966 |  | 28 | 17 | 01 | 04.0 | 39.00 | 27.00 | 49 | 10 | 4.5 |
| 1966 |  | 29 | 00 | 49 | 35.0 | 41.29 | 20.47 | 16 | 43 | 4.2 |
| 1966 |  | 30 | 19 | 21 | 29.0 | 41.18 | 20.85 | 19 | 42 | 4.3 |
| 1966 | JUL | 12 | 02 | 56 | 22.0 | 35.50 | 22.49 | 07 | 144 | 5.4 |
| 1966 |  | 15 | 23 | 50 | 12.1 | 38.90 | 21.65 | 34 | 40 | 4.2 |
| 1966 |  | 19 | 02 | 52 | 33.0 | 38.30 | 27.10 | 00 | 9 | 4.5 |
| 1966 |  | 20 | 10 | 16 | 06.0 | 38.83 | 21.39 | 22 | 64 | 4.3 |
| 1966 |  | 24 | 01 | 27 | 39.0 | 38.98 | 21.94 | 15 | 17 | 4.3 |
| 1966 |  | 31 | 04 | 22 | 17.0 | 35.70 | 22.30 | 71 | 17 | 4.6 |
| 1966 |  | 31 | 11 | 03 | 21.0 | 41.20 | 21.20 | 31 | 10 | 4.6 |
| 1966 | AUG | 06 | 18 | 32 | 32.0 | 37.90 | 22.20 | 25 | 43 | 4.3 |
| 1966 |  | 07 | 14 | 30 | 46.0 | 36.34 | 22.31 | 49 | 50 | 4.6 |
| 1966 |  | 09 | 03 | 34 | 15.1 | 40.22 | 19.86 | 38 | 92 | 4.9 |
| 1966 |  | 10 | 15 | 22 | 40.2 | 36.40 | 22.22 | 39 | 80 | 4.6 |
| 1966 |  | 11 | 00 | 23 | 40.8 | 37.65 | 20.99 | 48 | 73 | 4.6 |
| 1966 |  | 11 | 04 | 34 | 13.0 | 38.74 | 21.76 | 06 | 84 | 4.6 |
| 1966 |  | 16 | 03 | 53 | 41.7 | 40.16 | 19.75 | 20 | 113 | 4.9 |
| 1966 |  | 21 | 01 | 30 | 43.5 | 40.33 | 27.40 | 12 | 154 | 5.5 |
| 1966 | SEP | 01 | 14 | 22 | 56.9 | 37.46 | 22.16 | 15 | 198 | 5.4 |
| 1966 |  | 01 | 12 | 35 | 34.0 | 38.03 | 22.81 | 39 | 47 | 4.6 |
| 1966 |  | 06 | 12 | 31 | 57.3 | 36.66 | 26.63 | 158 | 28 | 4.5 |
| 1966 |  | 10 | 10 | 55 | 16.7 | 36.53 | 26.90 | 146 | 33 | 4.5 |
| 1966 |  | 22 | 20 | 14 | 39.4 | 39.83 | 23.92 | 35 | 19 | 4.6 |
| 1966 |  | 23 | 23 | 47 | . 58.1 | 38.60 | 21.73 | 47 | 31 | 4.5 |
| 1966 |  | 27 | 10 | 54 | 53.0 | 36.98 | 24.16 | 42 | 26 | 4.9 |
| 1966 | OCT | 21 | 16 | 17 | 04.0 | 39.53 | 22.11 | 57 | 64 | 4.6 |
| 1966 |  | 22 | 05 | 38 | 24.0 | 41.96 | 23.09 | 13 | 28 | 4.7 |
| 1966 |  | 29 | 02 | 39 | 24.8 | 38.90 | 21.10 | 01 | 266 | 5.8 |
| 1966 |  | 29 | 12 | 13 | 06.8 | 34.74 | 27.54 | 64 | 42 | 4.6 |
| 1966 |  | 30 | 02 | 10 | 14.0 | 38.75 | 21.58 | 26 | 72 | 4.7 |
| 1966 | NOV | 09 | 15 | 12 | 28.0 | 39.18 | 20.54 | 35 | 57 | 4.5 |


| DATE |  |  | ORIG.TIME GMT |  |  | $\stackrel{\text { LAT }}{\mathrm{N}}$ | $\begin{gathered} \text { LON } \\ \text { E } \end{gathered}$ | DEPTH KM | OBS | $\begin{array}{r} \text { MAG } \\ \text { MS } \end{array}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1966 | NOV | 19 | 07 | 12 | 38.0 | 35.03 | 23.46 | 17 | 176 | 5.2 |
| 1966 | DEC | 18 | 07 | 42 | 20.0 | 35.10 | 26.92 | 33 | 88 | 4.7 |
| 1967 | JAN | 01 | 22 | 17 | 52.0 | 37.80 | 20.00 | 5 | 18 | 4.8 |
| 1967 |  | 04 | 05 | 58 | 52.5 | 38.37 | 22.04 | 1 | 172 | 5.5 |
| 1967 |  | 04 | 07 | 10 | 14.0 | 38.29 | 22.13 | 24 | 34 | 4.3 |
| 1967 | FEB | 09 | 14 | 08 | 18.2 | 39.92 | 20.26 | 1 | 207 | 5.6 |
| 1967 |  | 14 | 07 | 22 | 24.0 | 38.80 | 27.70 | 0 | 36 | 4.7 |
| 1967 |  | 20 | 09 | 11 | 38.0 | 34.68 | 24.74 | 48 | 24 | 4.4 |
| 1967 |  | 28 | 14 | 21 | 51.3 | 37.53 | 21.18 | 46 | 85 | 5.0 |
| 1967 | MAR | 04 | 17 | 58 | 09.0 | 39.25 | 24.60 | 60 | 329 | 6.8 |
| 1967 |  | 04 | 18 | 38 | 01.0 | 38.99 | 24.80 | 15 | 55 | 4.7 |
| 1967 |  | 28 | 00 | 04 | 28.0 | 38.44 | 25.42 | 29 | 76 | 4.6 |
| 1967 | APR | 04 | 16 | 59 | 06.2 | 35.59 | 23.56 | 73 | 93 | 4.7 |
| 1967 |  | 04 | 03 | 47 | 17.0 | 40.32 | 26.20 | 32 | 24 | 4.0 |
| 1967 | MAY | 01 | 07 | 09 | 03.0 | 39.60 | 21.29 | 34 | 301 | 6.2 |
| 1967 |  | 01 | 09 | 50 | 08.2 | 39.51 | 21.30 | 33 | 110 | 4.9 |
| 1967 |  | 01 | 08 | 15 | 46.9 | 39.75 | 21.42 | 38 | 46 | 4.7 |
| 1967 |  | 01 | 09 | 47 | 40.0 | 39.46 | 21.23 | 10 | 35 | 4.4 |
| 1967 |  | 01 | 14 | 38 | 02.0 | 39.36 | 21.31 | 21 | 49 | 4.4 |
| 1967 |  | 02 |  | 27 | 20.4 | 39.56 | 21.20 | 35 | 47 | 4.3 |
| 1967 |  | 02 | 08 | 11 | 55.9 | 39.45 | 21.29 | 39 | 40 | 4.4 |
| 1967 |  | 03 | 18 | 41 | 47.2 | 39.53 | 21.34 | 37 | 117 | 5.3 |
| 1967 |  | 04 | 13 | 31 | 07.8 | 39.63 | 21.26 | 39 | 72 | 4.7 |
| 1967 |  | 05 | 06 | 26 | 37.9 | 39.56 | 21.29 | 57 | 80 | 4.9 |
| 1967 |  | 09 | 08 | 00 | 47.3 | 39.72 | 21.39 | 53 | 55 | 4.7 |
| 1967 |  | 14 | 04 | 15 | 59.9 | 37.70 | 21.17 | 48 | 132 | 4.9 |
| 1967 |  | 15 | 08 | 12 | 57.9 | 34.53 | 26.64 | 35 | 178 | 5.2 |
| 1967 |  | 30 | 23 | 53 | 31.6 | 34.17 | 28.67 | 35 | 59 | 4.4 |
| 1967 | JUN | 01 | 10 | 39 | 23.5 | 36.81 | 29.26 | 43 | 148 | 5.0 |
| 1967 |  | 07 | 15 | 54 | 36.0 | 34.78 | 26.68 | 52 | 51 | 4.7 |
| 1967 |  | 11 | 05 | 35 | 05.0 | 38.14 | 22.91 | 40 | 71 | 4.4 |
| 1967 |  | 12 | 02 | 51 | 05.8 | 38.15 | 22.77 | 35 | 160 | 5.0 |
| 1967 |  | 12 | 01 | 29 | 09.5 | 38.08 | 22.90 | 47 | 78 | 4.4 |
| 1967 |  | 12 | 11 | 00 | 16.0 | 38.04 | 22.75 | 05 | 21 | 4.3 |
| 1967 |  | 12 | 18 | 12 | 46.6 | 39.06 | 21.27 | 46 | 72 | 4.6 |
| 1967 |  | 18 | 05 | 28 | 53.9 | 36.78 | 29.32 | 35 | 50 | 4.2 |
| 1967 |  | 20 | 16 | 37 | 23.4 | 38.23 | 20.77 | 39 | 32 | 4.4 |
| 1967 | JUL | 05 | 00 | 53 | 16.8 | 36.73 | 21.50 | 50 | 140 | 4.9 |
| 1967 |  | 06 | 08 | 21 | 51.3 | 36.67 | 21.43 | 43 | 64 | 4.6 |
| 1967 |  | 13 | 14 | 38 | 58.4 | 40.66 | 19.67 | 73 | 62 | 4.7 |
| 1967 |  | 19 | 09 | 06 | 22.2 | 38.10 | 28.87 | 41 | 105 | 4.9 |
| 1967 |  | 19 | 16 | 18 | 32.0 | 38.03 | 20.95 | 06 | 29 | 4.3 |
| 1967 |  | 20 | 19 | 03 | 30.4 | 40.72 | 19.88 | 58 | 58 | 4.4 |
| 1967 |  | 25 | 11 | 03 | 54.0 | 37.80 | 28.60 | 75 | 13 | 4.5 |
| 1967 |  | 25 | 12 | 39 | 28.0 | 37.90 | 28.70 | 101 | 12 | 4.5 |
| 1967 |  | 31 | 07 | 12 | 05.0 | 40.60 | 27.62 | 4 | 36 | 4.4 |
| 1967 | AUG | 09 | 00 | 33 | 15.0 | 36.98 | 28.40 | 64 | 19 | 4.2 |
| 1967 |  | 15 | 04 | 35 | 52.9 | 36.54 | 19.28 | 33 | 51 | 4.8 |
| 1967 |  | 28 | 17 | 36 | 41.0 | 36.73 | 26.74 | 169 | 67 | 4.7 |
| 1967 | SEP | 03 | 07 | 46 | 21.8 | 38.23 | 22.00 | 45 | 29 | 4.4 |
| 1967 |  | 05 | 08 | 31 | 02.2 | 36.72 | 29.33 | 24 | 22 | 4.6 |
| 1967 |  | 06 | 04 | 59 | 23.0 | 35.06 | 23.09 | 20 | 104 | 4.8 |
| 1967 |  | 07 | 00 | 32 | 22.0 | 40.75 | 19.58 | 13 | 35 | 4.3 |
| 1967 |  | 08 | 02 | 04 | 45.0 | 40.60 | 20.08 | 01 | 110 | 5.3 |
| 1967 |  | 08 | 09 | 51 | 42.8 | 39.08 | 21.40 | 40 | 65 | 4.4 |


| DATE |  |  | $\begin{aligned} & \text { ORIG.TIME } \\ & \text { GMT } \end{aligned}$ |  |  | $\underset{\mathrm{N}}{\mathrm{LAT}}$ | $\begin{gathered} \text { LON } \\ \text { E } \end{gathered}$ | $\begin{aligned} & \text { DEPTH } \\ & \text { KM } \end{aligned}$ | OBS | $\begin{array}{r} \text { MAG } \\ \text { MS } \end{array}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1967 | SEP | 12 | 14 | 46 | 42.0 | 39.23 | 21.46 | 25 | 56 | 4.7 |
| 1967 |  | 24 | 22 | 11 | 20.4 | 40.86 | 19.70 | 35 | 74 | 4.6 |
| 1967 |  | 26 | 05 | 05 | 37.4 | 41.53 | 20.94 | 39 | 49 | 4.3 |
| 1967 |  | 27 | 07 | 24 | 34.0 | 34.42 | 26.60 | 49 | 80 | 4.6 |
| 1967 | OCT | 05 | 12 | 00 | 53.7 | 37.74 | 20.80 | 37 | 145 | 5.1 |
| 1967 |  | 11 | 07 | 48 | 45.0 | 36.07 | 27.12 | 34 | 46 | 4.5 |
| 1967 |  | 22 | 05 | 38 | 04.7 | 36.90 | 21.10 | 05 | 24 | 4.9 |
| 1967 |  | 24 | 06 | 14 | 44.7 | 38.91 | 21.96 | 37 | 41 | 4.4 |
| 1967 |  | 26 | 04 | 55 | 39.3 | 37.22 | 29.05 | 46 | 150 | 5.1 |
| 1967 | NOV | 05 | 00 | 26 | 13.8 | 38.12 | 20.34 | 33 | 31 | 4.7 |
| 1967 |  | 06 | 10 | 32 | 58.0 | 39.05 | 20.61 | 01 | 44 | 4.6 |
| 1967 |  | 13 | 06 | 50 | 34.9 | 37.78 | 28.83 | 34 | 36 | 4.5 |
| 1967 |  | 18 | 02 | 31 | 36.0 | 35.25 | 23.05 | 34 | 97 | 4.7 |
| 1967 |  | 26 | 03 | 24 | 57.4 | 39.40 | 20.49 | 37 | 51 | 5.1 |
| 1967 |  | 30 | 07 | 23 | 50.4 | 41.41 | 20.44 | 21 | 312 | 6.5 |
| 1967 |  | 30 | 07 | 42 | 52.0 | 41.43 | 20.49 | 21 | 62 | 4.7 |
| 1967 |  | 30 | 08 | 13 | 17.5 | 41.40 | 20.50 | 30 | 18 | 4.4 |
| 1967 | DEC | 01 | 18 | 30 | 57.1 | 41.37 | 20.27 | 16 | 24 | 4.6 |
| 1967 |  | 01 | 20 | 07 | 51.0 | 41.28 | 20.28 | 28 | 39 | 4.7 |
| 1967 |  | 02 | 12 | 44 | 42.0 | 41.32 | 20.29 | 16 | 172 | 5.4 |
| 1967 |  | 02 | 00 | 24 | 13.0 | 41.31 | 20.34 | 08 | 54 | 5.5 |
| 1967 |  | 02 | 09 | 27 | 08.0 | 41.20 | 20.08 | 19 | 33 | 4.6 |
| 1967 |  | 02 | 14 | 18 | 04.0 | 41.29 | 20.29 | 42 | 28 | 4.4 |
| 1967 |  | 03 | 17 | 59 | 25.0 | 41.25 | 20.20 | 25 | 24 | 4.3 |
| 1967 |  | 04 | 00 | 48 | 51.0 | 41.17 | 20.66 | 10 | 19 | 4.4 |
| 1967 |  | 05 | 05 | 20 | 03.1 | 36.53 | 26.85 | 137 | 83 | 5.9 |
| 1967 |  | 06 | 00 | 01 | 56.0 | 41.30 | 20.40 | 42 | 30 | 4.4 |
| 1967 |  | 07 | 18 | 03 | 35.0 | 41.27 | 20.24 | 32 | 33 | 4.7 |
| 1967 |  | 14 | 02 | 54 | 54.0 | 34.49 | 26.27 | 57 | 85 | 4.5 |
| 1967 |  | 14 | 08 | 35 | 23.7 | 34.74 | 24.61 | 45 | 36 | 4.6 |
| 1967 |  | 19 | 08 | 32 | 32.3 | 41.49 | 20.43 | 29 | 110 | 4.9 |
| 1967 |  | 29 | 19 | 49 | 24.1 | 41.41 | 20.27 | 46 | 101 | 4.9 |
| 1967 |  | 30 | 21 | 27 | 20.3 | 40.66 | 21.47 | 34 | 32 | 4.6 |
| 1968 | JAN | 09 | 23 | 15 | 42.8 | 35.52 | 22.54 | 46 | 69 | 5.0 |
| 1968 | FEB | 07 | 22 | 22 | 19.0 | 36.65 | 26.74 | 153 | 146 | 5.8 |
| 1968 |  | 19 | 22 | 45 | 42.4 | 39.40 | 24.94 | 7 | 333 | 7.2 |
| 1968 |  | 19 | 23 | 09 | 46.4 | 39.36 | 24.70 | 15 | 25 | 4.6 |
| 1968 |  | 19 | 23 | 12 | 32.0 | 39.62 | 25.50 | 15 | 34 | 4.5 |
| 1968 |  | 19 | 23 | 53 | 51.0 | 39.55 | 25.30 | 33 | 36 | 4.6 |
| 1968 |  | 20 | 00 | 39 | 15.7 | 39.73 | 25.37 | 37 | 88 | 5.1 |
| 1968 |  | 20 | 01 | 28 | 29.0 | 39.40 | 25.60 | 46 | 18 | 4.2 |
| 1968 |  | 20 | 02 | 21 | 52.0 | 39.56 | 25.45 | 8 | 90 | 5.5 |
| 1968 |  | 20 | 02 | 29 | 28.0 | 39.30 | 24.90 | 33 | 8 | 4.6 |
| 1968 |  | 20 | 06 | 15 | 46.0 | 39.30 | 25.50 | 32 | 30 | 4.2 |
| 1968 |  | 20 | 09 | 35 | 51.6 | 39.41 | 24.88 | 33 | 98 | 4.8 |
| 1968 |  | 20 | 09 | 41 | 09.0 | 39.35 | 24.95 | 33 | 116 | 5.0 |
| 1968 |  | 20 | 16 | 50 | 44.8 | 36.15 | 27.39 | 64 | 121 | 5.0 |
| 1968 |  | 20 | 21 | 05 | 23.6 | 39.25 | 25.05 | 33 | 23 | 4.8 |
| 1968 |  | 21 | 00 | 17 | 28.0 | 39.56 | 24.97 | 2 | 22 | 4.6 |
| 1968 |  | 21 | 07 | 18 | 50.0 | 39.30 | 25.00 | 6 | 18 | 4.3 |
| 1968 |  | 21 | 12 | 35 | 55.3 | 39.61 | 25.30 | 5 | 27 | 4.6 |
| 1968 |  | 22 | 02 | 16 | 39.0 | 39.66 | 25.72 | 6 | 43 | 4.6 |
| 1968 |  | 22 | 04 | 57 | 47.0 | 39.39 | 25.02 | 19 | 78 | 4.5 |
| 1968 |  | 24 | 12 | 55 | 03.0 | 41.44 | 20.18 | 24 | 25 | 4.4 |
| 1968 |  | 26 | 05 | 43 | 30.4 | 39.39 | 24.79 | 5 | 12 | 4.3 |


| DATE |  |  | ORIG.TIME GMT |  |  | $\begin{gathered} \text { LAT } \\ \mathrm{N} \end{gathered}$ | $\begin{gathered} \text { LON } \\ \text { E } \end{gathered}$ | $\begin{gathered} \text { DEPTH } \\ \text { KM } \end{gathered}$ | OBS | $\begin{array}{r} \text { MAG } \\ \text { MS } \end{array}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1968 | FEB | 27 | 13 | 20 | 15.7 | 39.59 | 25.51 | 36 | 19 | 4.3 |
| 1968 |  | 27 | 13 | 37 | 45.4 | 39.61 | 25.51 | 35 | 67 | 4.7 |
| 1968 |  | 29 | 11 | 46 | 42.0 | 39.50 | 26.00 | 5 | 19 | 4.0 |
| 1968 |  | 29 | 12 | 47 | 33.5 | 39.12 | 24.32 | 18 | 34 | 4.3 |
| 1968 | MAR | 06 | 05 | 14 | 49.0 | 39.34 | 25.04 | 10 | 29 | 4.4 |
| 1968 |  | 10 | 06 | 48 | 17.1 | 39.10 | 24.36 | 33 | 72 | 4.4 |
| 1968 |  | 10 | 07 | 10 | 59.0 | 39.13 | 24.23 | 9 | 167 | 5.4 |
| 1968 |  | 11 | 17 | 32 | 46.9 | 39.50 | 25.56 | 5 | 36 | 4.3 |
| 1968 |  | 16 | 18 | 11 | 05.8 | 39.38 | 24.94 | 43 | 86 | 4.5 |
| 1968 |  | 21 | 09 | 42 | 51.0 | 38.80 | 27.60 | 52 | 13 | 4.4 |
| 1968 |  | 21 | 16 | 09 | 23.8 | 39.76 | 25.49 | 19 | 53 | 4.8 |
| 1968 |  | 23 | 17 | 16 | 35.8 | 39.78 | 25.64 | 5 | 31 | 4.3 |
| 1968 |  | 23 | 17 | 25 | 55.0 | 39.76 | 25.48 | 33 | 127 | 5.3 |
| 1968 |  | 28 | 07 | 39 | 59.5 | 37.84 | 20.89 | 23 | 255 | 5.9 |
| 1968 |  | 28 | 16 | 37 | 47.3 | 39.49 | 20.38 | 18 | 129 | 5.3 |
| 1968 | APR | 05 | 15 | 54 | 32.7 | 39.76 | 25.55 | 18 | 50 | 4.1 |
| 1968 |  | 08 | 08 | 59 | 09.0 | 39.68 | 25.50 | 9 | 15 | 4.2 |
| 1968 |  | 18 | 03 | 08 | 03.4 | 41.25 | 20.22 | 36 | 54 | 4.3 |
| 1968 |  | 24 | 08 | 18 | 03.3 | 39.33 | 24.88 | 20 | 227 | 5.5 |
| 1968 | MAY | 28 | 21 | 31 | 41.3 | 38.78 | 23.57 | 10 | 33 | 4.2 |
| 1968 |  | 30 | 17 | 40 | 26.0 | 35.45 | 27.88 | 27 | 222 | 5.9 |
| 1968 | JUN | 12 | 09 | 05 | 04.0 | 35.30 | 27.89 | 16 | 77 | 4.6 |
| 1968 |  | 24 | 10 | 17 | 31.0 | 38.00 | 20.80 | 42 | 31 | 4.3 |
| 1968 | JUL | 04 | 21 | 47 | 53.6 | 37.76 | 23.23 | 20 | 226 | 5.5 |
| 1968 |  | 08 | 17 | 41 | 06.4 | 34.47 | 25.08 | 38 | 194 | 5.5 |
| 1968 |  | 08 | 18 | 18 | 11.4 | 34.29 | 25.20 | 57 | 44 | 4.4 |
| 1968 |  | 09 | 15 | 00 | 47.4 | 34.39 | 25.10 | 49 | 80 | 4.6 |
| 1968 |  | 13 | 19 | 34 | 07.0 | 35.56 | 28.05 | 0 | 7 | 4.5 |
| 1968 |  | 25 | 22 | 05 | 29.0 | 40.95 | 20.09 | 23 | 79 | 4.5 |
| 1968 |  | 27 | 02 | 45 | 51.0 | 35.43 | 27.92 | 29 | 213 | 5.2 |
| 1968 |  | 31 | 09 | 21 | 56.0 | 37.84 | 21.14 | 34 | 58 | 4.3 |
| 1968 |  | 31 | 19 | 29 | 29.7 | 35.54 | 28.00 | 49 | 126 | 4.7 |
| 1968 | AUG | 04 | 23 | 24 | 22.2 | 37.81 | 21.02 | 62 | 53 | 4.6 |
| 1968 |  | 15 | 02 | 29 | 43.1 | 35.18 | 26.70 | 48 | 159 | 5.1 |
| 1968 | SEP | 05 | 18 | 42 | 32.0 | 36.30 | 26.70 | 1 | 24 | 4.5 |
| 1968 |  | 15 | 04 | 55 | 58.4 | 34.70 | 25.05 | 17 | 144 | 4.9 |
| 1968 |  | 16 | 02 | 55 | 52.0 | 38.05 | 20.66 | 11 | 23 | 4.4 |
| 1968 |  | 18 | 04 | 01 | 59.0 | 34.74 | 25.01 | 30 | 96 | 4.6 |
| 1968 |  | 28 | 00 | 53 | 28.0 | 40.49 | 26.38 | 28 | 65 | 4.7 |
| 1968 | OCT | 03 | 18 | 18 | 34.8 | 40.13 | 19.85 | 58 | 22 | 4.7 |
| 1968 |  | 06 | 15 | 06 | 43.0 | 36.96 | 26.38 | 17 | 99 | 4.7 |
| 1968 |  | 10 | 05 | 16 | 26.0 | 36.50 | 29.20 | 0 | 15 | 4.7 |
| 1968 |  | 11 | 03 | 02 | 36.0 | 36.54 | 25.87 | 33 | 26 | 4.4 |
| 1968 |  | 17 | 23 | 56 | 04.0 | 38.21 | 20.17 | 17 | 81 | 4.5 |
| 1968 |  | 19 | 15 | 34 | 54.0 | 35.24 | 23.40 | 6 | 133 | 4.9 |
| 1968 |  | 21 | 18 | 16 | 41.0 | 35.25 | 23.35 | 1 | 106 | 4.9 |
| 1968 |  | 28 | 12 | 54 | 30.0 | 38.89 | 25.82 | 4 | 67 | 4.7 |
| 1968 |  | 31 | 03 | 22 | 14.0 | 36.62 | 27.01 | 2 | 159 | 5.4 |
| 1968 | NOV | 03 | 04 | 49 | 33.7 | 42.10 | 19.35 | 28 | 214 | 5.4 |
| 1968 |  | 03 | 18 | 44 | 08.0 | 38.60 | 28.70 | 56 | 15 | 4.3 |
| 1968 |  | 04 | 20 | 05 | 59.0 | 36.44 | 26.98 | 35 | 55 | 4.5 |
| 1968 |  | 10 | 12 | 50 | 37.0 | 34.44 | 23.77 | 1 | 39 | 5.1 |
| 1968 |  | 10 | 14 | 29 | 33.0 | 34.55 | 23.86 | 5 | 30 | 4.3 |
| 1968 |  | 11 | 23 | 34 | 21.5 | 36.61 | 27.15 | 23 | 22 | 5.0 |
| 1968 |  | 11 | 23 | 53 | 07.0 | 36.61 | 27.10 | 33 | 50 | 4.1 |


| DATE |  | ORIG.TIME GMT |  |  | $\frac{\text { LAT }}{\mathrm{N}}$ | $\begin{gathered} \text { LON } \\ \text { E } \end{gathered}$ | DEPTH KM | OBS | $\begin{gathered} \text { MAG } \\ \text { MS } \end{gathered}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1968 NOV | 12 | 03 | 37 | 39.0 | 36.74 | 27.11 | 26 | 104 | 5.0 |
| 1968 | 12 | 06 | 08 | 55.6 | 36.64 | 27.16 | 24 | 92 | 5.0 |
| 1968 DEC | 04 | 18 | 43 | 28.0 | 36.34 | 26.98 | 43 | 52 | 4.4 |
| 1968 | 04 | 19 | 37 | 22.0 | 36.50 | 27.02 | 32 | 87 | 4.7 |
| 1968 | 05 | 07 | 52 | 11.1 | 36.60 | 26.92 | 31 | 206 | 5.6 |
| 1968 | 21 | 00 | 36 | 40.0 | 36.60 | 27.07 | 30 | 88 | 4.9 |
| 1968 | 25 | 12 | 17 | 19.1 | 34.99 | 24.31 | 58 | 156 | 5.1 |
| 1969 JAN | 10 | 04 | 32 | 03.4 | 39.23 | 19.97 | 37 | 44 | 4.4 |
| 1969 | 13 | 05 | 46 | 40.4 | 38.31 | 22.52 | 46 | 75 | 4.9 |
| 1969 | 13 | 07 | 57 | 07.6 | 34.57 | 24.93 | 60 | 56 | 4.4 |
| 1969 | 14 | 23 | 12 | 06.2 | 36.11 | 29.19 | 22 | 263 | 5.9 |
| 1969 | 31 | 14 | 40 | 04.0 | 34.29 | 26.14 | 31 | 78 | 4.9 |
| 1969 | 31 | 15 | 34 | 28.0 | 39.10 | 20.43 | 04 | 37 | 4.6 |
| 1969 FEB | 13 | 15 | 09 | 32.0 | 34.70 | 22.30 | 25 | 28 | 4.6 |
| 1969 | 17 | 09 | 11 | 46.0 | 34.11 | 25.31 | 26 | 33 | 4.7 |
| 1969 | 21 | 18 | 39 | 57.0 | 39.14 | 21.87 | 33 | 51 | 4.6 |
| 1969 | 26 | 12 | 35 | 49.1 | 36.66 | 27.18 | 33 | 38 | 4.3 |
| 1969 MAR | 03 | 00 | 59 | 10.5 | 40.09 | 27.50 | 6 | 225 | 5.9 |
| 1969 | 05 | 14 | 41 | 16.4 | 40.06 | 27.56 | 33 | 92 | 4.7 |
| 1969 | 22 | 18 | 00 | 55.0 | 39.10 | 28.67 | 28 | 37 | 4.7 |
| 1969 | 23 | 21 | 08 | 42.1 | 39.14 | 28.48 | 9 | 256 | 5.9 |
| 1969 | 24 | 01 | 59 | 34.0 | 39.11 | 28.51 | 30 | 156 | 5.1 |
| 1969 | 24 | 02 | 58 | 49.0 | 39.15 | 28.60 | 4 | 19 | 4.4 |
| 1969 | 24 | 08 | 13 | 05.4 | 39.02 | 28.41 | 43 | 55 | 4.7 |
| 1969 | 24 | 11 | 34 | 34.0 | 39.17 | 28.70 | 37 | 72 | 4.6 |
| 1969 | 24 | 12 | 13 | 17.0 | 39.08 | 28.65 | 20 | 34 | 4.7 |
| 1969 | 25 | 13 | 21 | 12.0 | 39.06 | 28.41 | 28 | 104 | 5.0 |
| 1969 | 25 | 13 | 21 | 34.2 | 39.25 | 28.44 | 37 | 197 | 5.8 |
| 1969 | 25 | 13 | 28 | 50.1 | 38.78 | 28.51 | 40 | 25 | 4.9 |
| 1969 | 25 | 14 | 18 | 52.1 | 39.17 | 28.49 | 34 | 120 | 4.9 |
| 1969 | 25 | 16 | 13 | 30.4 | 39.08 | 28.44 | 42 | 65 | 4.7 |
| 1969 | 26 | 03 | 31 | 26.5 | 39.03 | 28.27 | 37 | 56 | 4.6 |
| 1969 | 27 | 18 | 07 | 03.0 | 39.12 | 28.20 | 51 | 25 | 4.4 |
| 1969 | 28 | 01 | 48 | 29.5 | 38.55 | 28.46 | 4 | 296 | 6.4 |
| 1969 | 28 | 10 | 02 | 17.4 | 39.13 | 28.45 | 37 | 124 | 5.0 |
| 1969 APR | 02 | 04 | 57 | 30.0 | 38.13 | 20.12 | 20 | 38 | 4.4 |
| 1969 | 03 | 22 | 12 | 21.9 | 40.66 | 19.98 | 21 | 197 | 5.6 |
| 1969 | 06 | 03 | 49 | 33.9 | 38.47 | 26.41 | 16 | 229 | 5.9 |
| 1969 | 08 | 15 | 48 | 50.4 | 40.67 | 19.77 | 17 | 128 | 4.9 |
| 1969 | 09 | 16 | 27 | 49.0 | 38.16 | 19.99 | 11 | 28 | 4.4 |
| 1969 | 14 | 05 | 11 | 45.5 | 38.90 | 21.79 | 36 | 82 | 4.5 |
| 1969 | 16 | 04 | 54 | 12.8 | 35.30 | 27.90 | 55 | 143 | 4.9 |
| 1969 | 16 | 22 | 55 | 40.5 | 35.32 | 27.77 | 52 | 173 | 5.3 |
| 1969 | 16 | 23 | 21 | 06.2 | 35.23 | 27.72 | 58 | 181 | 5.3 |
| 1969 | 17 | 00 | 54 | 38.2 | 35.19 | 27.83 | 55 | 132 | 4.9 |
| 1969 | 21 | 20 | 36 | 40.0 | 39.42 | 25.09 | 1 | 83 | 4.7 |
| 1969 | 24 | 14 | 45 | 48.8 | 36.35 | 28.73 | 53 | 79 | 4.7 |
| 1969 | 27 | 10 | 58 | 26.0 | 36.54 | 28.21 | 33 | 72 | 4.7 |
| 1969 | 30 | 20 | 20 | 32.0 | 39.12 | 28.52 | 8 | 141 | 5.1 |
| 1969 MAY | 01 | 18 | 02 | 16.4 | 35.41 | 27.68 | 51 | 180 | 5.3 |
| 1969 | 01 | 20 | 06 | 45.4 | 35.39 | 27.73 | 67 | 145 | 4.7 |
| 1969 | 02 | 18 | 38 | 15.0 | 34.25 | 26.21 | 38 | 68 | 4.4 |
| 1969 | 03 | 03 | 25 | 36.3 | 35.21 | 28.03 | 81 | 48 | 4.6 |
| 1969 | 03 | 20 | 31 | 14.4 | 35.17 | 27.76 | 53 | 52 | 4.7 |
| 1969 | 13 | 17 | 48 | 02.1 | 39.03 | 28.57 | 35 | 42 | 4.6 |


| DATE |  |  | ORIG.TIME GMT |  |  | $\underset{\mathrm{N}}{\mathrm{LAT}}$ | $\begin{gathered} \text { LON } \\ \mathrm{E} \end{gathered}$ | DEPTH KM | OBS | MAG |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1969 | MAY | 14 | 10 | 05 | 17.1 | 35.33 | 27.72 | 43 | 181 | 5.3 |
| 1969 |  | 14 | 23 | 57 | 35.5 | 39.15 | 28.49 | 36 | 28 | 4.6 |
| 1969 |  | 15 | 12 | 05 | 56.8 | 35.28 | 27.73 | 46 | 143 | 4.9 |
| 1969 |  | 16 | 05 | 09 | 34.4 | 35.00 | 24.51 | 51 | 70 | 4.6 |
| 1969 |  | 16 | 07 | 27 | 01.1 | 39.13 | 21.82 | 39 | 153 | 5.1 |
| 1969 | JUN | 07 | 15 | 31 | 09.0 | 37.85 | 20.19 | 9 | 104 | 4.6 |
| 1969 |  | 12 | 15 | 13 | 30.9 | 34.43 | 25.04 | 22 | 288 | 6.2 |
| 1969 |  | 12 | 18 | 00 | 30.8 | 34.23 | 25.22 | 63 | 53 | 4.9 |
| 1969 |  | 13 | 01 | 23 | 14.6 | 34.35 | 25.14 | 41 | 89 | 4.7 |
| 1969 |  | 14 | 13 | 47 | 26.4 | 34.34 | 25.05 | 21 | 154 | 5.1 |
| 1969 |  | 14 | 14 | 32 | 57.4 | 34.34 | 25.26 | 50 | 43 | 4.4 |
| 1969 |  | 15 | 05 | 58 | 43.2 | 34.30 | 25.13 | 35 | 67 | 4.6 |
| 1969 |  | 16 | 16 | 06 | 25.6 | 38.11 | 20.58 | 40 | 82 | 4.6 |
| 1969 |  | 17 | 05 | 18 | 43.0 | 38.23 | 20.21 | 11 | 57 | 4.4 |
| 1969 |  | 19 | 06 | 52 | 36.7 | 34.29 | 25.23 | 49 | 45 | 4.6 |
| 1969 |  | 25 | 06 | 11 | 51.8 | 35.98 | 27.60 | 48 | 54 | 4.7 |
| 1969 | JUL | 03 | 09 | 42 | 02.0 | 38.41 | 22.05 | 28 | 71 | 4.5 |
| 1969 |  | 04 | 10 | 13 | 52.0 | 35.32 | 27.89 | 30 | 33 | 4.7 |
| 1969 |  | 08 | 08 | 09 | 13.0 | 37.50 | 20.31 | 30 | 212 | 5.8 |
| 1969 |  | 11 | 01 | 45 | 26.0 | 35.32 | 28.10 | 40 | 20 | 4.4 |
| 1969 |  | 20 | 15 | 51 | 56.5 | 37.94 | 20.41 | 38 | 64 | 4.5 |
| 1969 |  | 24 | 23 | 21 | 19.3 | 34.93 | 26.00 | 60 | 81 | 4.5 |
| 1969 | AUG | 13 | 04 | 06 | 03.0 | 38.37 | 21.75 | 24 | 66 | 4.5 |
| 1969 |  | 14 | 21 | 51 | 05.3 | 39.52 | 27.87 | 21 | 56 | 4.7 |
| 1969 |  | 26 | 02 | 15 | 37.1 | 41.73 | 20.03 | 28 | 118 | 5.0 |
| 1969 | SEP | 04 | 19 | 25 | 26.6 | 35.11 | 27.17 | 43 | 83 | 4.9 |
| 1969 |  | 06 | 20 | 30 | 40.3 | 36.73 | 28.35 | 72 | 137 | 5.1 |
| 1969 |  | 22 | 08 | 17 | 43.4 | 36.57 | 28.01 | 86 | 42 | 4.7 |
| 1969 |  | 28 | 22 | 54 | 08.0 | 34.30 | 25.15 | 29 | 203 | 5.5 |
| 1969 | OCT | 02 | 23 | 13 | 40.6 | 38.47 | 22.29 | 45 | 50 | 4.7 |
| 1969 |  | 07 | 05 | 09 | 12.0 | 39.20 | 28.40 | 13 | 148 | 5.0 |
| 1969 |  | 07 | 18 | 49 | 02.6 | 39.16 | 28.54 | 49 | 32 | 4.5 |
| 1969 |  | 12 | 13 | 34 | 19.9 | 39.76 | 20.55 | 46 | 163 | 5.1 |
| 1969 |  | 13 | 01 | 02 | 30.8 | 39.78 | 20.59 | 27 | 236 | 5.7 |
| 1969 | DEC | 01 | 20 | 18 | 03.8 | 34.85 | 24.22 | 35 | 136 | 5.3 |
| 1969 |  | 19 | 23 | 54 | 40.5 | 38.92 | 22.00 | 65 | 65 | 4.3 |
| 1969 |  | 20 | 17 | 40 | 36.3 | 36.59 | 23.46 | 90 | 65 | 5.1 |
| 1969 |  | 21 | 22 | 01 | 06.8 | 36.66 | 28.42 | 69 | 47 | 4.6 |
| 1969 |  | 23 | 02 | 13 | 49.0 | 39.37 | 23.80 | 6 | 19 | 4.9 |
| 1969 |  | 27 | 07 | 31 | 54.5 | 39.22 | 23.82 | 42 | 44 | 4.6 |
| 1969 |  | 28 | 22 | 02 | 35.6 | 40.67 | 19.62 | 51 | 44 | 4.6 |
| 1969 |  | 31 | 05 | 08 | 10.1 | 34.26 | 26.37 | 51 | 22 | 4.5 |
| 1969 |  | 31 | 05 | 37 | 05.6 | 34.44 | 26.11 | 54 | 108 | 5.1 |
| 1970 | JAN | 24 | 15 | 43 | 54.0 | 37.21 | 23.45 | 105 | 42 | 4.7 |
| 1970 | FEB | 11 | 19 | 01 | 18.9 | 37.59 | 22.67 | 79 | 125 | 5.0 |
| 1970 |  | 17 | 00 | 16 | 28.3 | 39.34 | 20.62 | 53 | 89 | 4.6 |
| 1970 |  | 17 | 04 | 51 | 10.3 | 38.83 | 21.68 | 22 | 68 | 4.6 |
| 1970 |  | 20 | 20 | 19 | 32.0 | 36.55 | 27.26 | 20 | 39 | 4.9 |
| 1970 |  | 22 | 15 | 48 | 31.0 | 35.21 | 25.24 | 43 | 58 | 5.3 |
| 1970 |  | 22 | 15 | 52 | 17.1 | 35.38 | 25.27 | 34 | 75 | 4.9 |
| 1970 | MAR | 03 | 16 | 35 | 46.2 | 38.28 | 20.74 | 49 | 44 | 4.3 |
| 1970 |  | 04 |  | 51 | 30.7 | 34.47 | 26.48 | 44 | 60 | 4.7 |
| 1970 |  | 17 |  | 00 | 56.8 | 41.40 | 21.07 | 43 | 43 | 4.6 |
| 1970 |  | 23 |  | 56 | 01.0 | 39.04 | 20.49 | 7 | 119 | 5.0 |
| 1970 |  | 29 |  | 37 | 19.6 | 38.74 | 27.83 | 56 | 94 | 4.4 |


| DATE |  |  | ORIG.TIME GMT |  |  | $\underset{\mathrm{N}}{\mathrm{LAT}}$ | $\begin{gathered} \text { LON } \\ \text { E } \end{gathered}$ | DEPTH KM | OBS | $\begin{array}{r} \text { MAG } \\ \text { MS } \end{array}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1970 | MAR | 29 | 14 | 40 | 26.6 | 38.73 | 28.00 | 47 | 46 | 4.4 |
| 1970 |  | 30 | 06 | 46 | 24.9 | 39.09 | 29.03 | 23 | 57 | 4.4 |
| 1970 | APR | 05 | 04 | 55 | 39.5 | 34.68 | 25.07 | 35 | 68 | 4.5 |
| 1970 |  | 07 | 09 | 18 | 44.0 | 34.57 | 26.14 | 20 | 111 | 5.0 |
| 1970 |  | 08 | 13 | 50 | 28.3 | 38.34 | 22.56 | 23 | 281 | 6.2 |
| 1970 |  | 11 | 01 | 03 | 11.3 | 38.17 | 22.77 | 55 | 82 | 4.4 |
| 1970 |  | 16 | 22 | 39 | 31.3 | 40.67 | 23.45 | 20 | 159 | 5.0 |
| 1970 |  | 20 | 15 | 39 | 31.6 | 38.27 | 22.66 | 38 | 193 | 5.3 |
| 1970 |  | 23 |  | 29 | 48.2 | 37.51 | 22.73 | 74 | 133 | 4.9 |
| 1970 |  | 23 | 09 | 01 | 26.6 | 39.13 | 28.65 | 28 | 208 | 5.4 |
| 1970 |  | 24 |  | 37 | 20.0 | 36.75 | 28.66 | 34 | 61 | 4.6 |
| 1970 | MAY | 08 | 18 | 30 | 42.7 | 38.67 | 22.30 | 58 | 70 | 4.4 |
| 1970 |  | 12 | 22 | 49 | 03.2 | 38.21 | 22.55 | 39 | 146 | 4.9 |
| 1970 |  | 24 | 11 | 03 | 01.0 | 36.05 | 25.49 | 37 | 26 | 5.0 |
| 1970 | JUN | 08 | 06 | 51 | 03.0 | 41.44 | 20.40 | 29 | 31 | 4.4 |
| 1970 |  | 09 | 20 | 43 | 30.2 | 36.16 | 25.56 | 63 | 44 | 4.4 |
| 1970 |  | 19 | 22 | 27 | 01.3 | 39.48 | 20.56 | 58 | 77 | 4.5 |
| 1970 |  | 27 | 18 | 57 | 15.0 | 41.49 | 19.39 | 48 | 79 | 4.5 |
| 1970 |  | 30 | 18 | 21 | 22.0 | 38.80 | 20.57 | 22 | 89 | 4.6 |
| 1970 | JUL | 02 | 07 | 50 | 14.0 | 38.72 | 20.59 | 27 | 109 | 5.9 |
| 1970 |  | 02 | 14 | 12 | 55.9 | 38.65 | 20.21 | 17 | 20 | 4.7 |
| 1970 |  | 03 | 00 | 41 | 00.0 | 38.77 | 20.42 | 24 | 82 | 4.9 |
| 1970 |  | 11 | 23 | 29 | 19.6 | 38.86 | 20.57 | 36 | 76 | 4.5 |
| 1970 |  | 13 | 00 | 46 | 45.0 | 38.73 | 20.56 | 19 | 84 | 4.5 |
| 1970 | AUG | 08 | 12 | 13 | 19.8 | 37.79 | 21.80 | 77. | 51 | 4.4 |
| 1970 |  | 18 | 17 | 40 | 17.9 | 39.16 | 21.78 | 38 | 60 | 4.5 |
| 1970 |  | 19 | 02 | 01 | 51.6 | 41.08 | 19.77 | 21 | 207 | 5.3 |
| 1970 |  | 29 | 10 | 42 | 17.2 | 41.49 | 19.45 | 33 | 45 | 4.4 |
| 1970 | SEP | 01 | 01 | 06 | 40.0 | 38.94 | 20.21 | 7 | 132 | 4.8 |
| 1970 |  | 03 | 05 | 32 | 10.2 | 39.60 | 28.78 | 22 | 167 | 5.1 |
| 1970 |  | 18 | 16 | 53 | 38.0 | 34.33 | 26.26 | 12 | 62 | 4.7 |
| 1970 |  | 24 | 21 | 25 | 15.0 | 34.17 | 26.18 | 42 | 21 | 4.3 |
| 1970 |  | 27 | 15 | 56 | 35.5 | 39.18 | 20.40 | 53 | 53 | 4.8 |
| 1970 |  | 28 | 19 | 54 | 09.0 | 37.09 | 28.59 | 24 | 36 | 4.4 |
| 1970 | OCT | 01 | 22 | 21 | 56.9 | 38.04 | 22.85 | 35 | 115 | 4.9 |
| 1970 |  | 01 | 22 | 38 | 37.2 | 38.02 | 22.77 | 43 | 131 | 5.3 |
| 1970 |  | 08 | 22 | 14 | 22.4 | 38.10 | 20.29 | 46 | 86 | 4.6 |
| 1970 |  | 10 | 13 | 48 | 26.0 | 38.07 | 20.19 | 35 | 73 | 4.4 |
| 1970 |  | 11 | 02 | 35 | 29.0 | 38.04 | 20.25 | 15 | 71 | 4.4 |
| 1970 |  | 31 | 16 | 07 | 39.4 | 42.10 | 19.35 | 39 | 75 | 4.6 |
| 1970 |  | 26 | 01 | 57 | 39.7 | 34.50 | 24.05 | 47 | 87 | 4.7 |
| 1970 | NOV | 30 | 09 | 49 | 02.3 | 39.06 | 21.94 | 38 | 46 | 4.4 |
| 1970 | DEC | 28 | 03 | 42 | 15.0 | 37.06 | 29.02 | 7 | 27 | 4.4 |
| 1970 |  | 28 | 17 | 00 | 46.0 | 35.92 | 28.21 | 28 | 65 | 4.6 |
| 1970 |  | 29 | 00 | 49 | 04.0 | 35.05 | 23.32 | 36 | 52 | 4.5 |
| 1970 |  | 29 | 12 | 47 | 10.0 | 35.05 | 23.36 | 49 | 37 | 4.7 |
| 1970 |  | 29 | 21 | 03 | 38.0 | 36.03 | 28.34 | 26 | 39 | 4.6 |
| 1970 |  | 30 | 18 | 54 | 44.0 | 36.97 | 28.94 | 23 | 51 | 4.7 |
| 1971 | JAN | 02 | 00 | 46 | 15.9 | 35.12 | 23.17 | 42 | 95 | 4.6 |
| 1971 |  | 02 | 03 | 25 | 36.0 | 37.07 | 29.04 | 07 | 52 | 4.3 |
| 1971 |  | 03 |  | 18 | 43.1 | 34.63 | 26.32 | 47 | 220 | 5.4 |
| 1971 |  | 17 |  | 18 | 47.0 | 38.08 | 20.51 | 29 | 19 | 4.4 |
| 1971 |  | 18 |  | 35 | 50.0 | 37.51 | 20.40 | 40 | 43 | 4.4 |
| 1971 |  | 19 |  | 33 | 56.7 | 34.30 | 24.06 | 34 | 143 | 5.0 |
| 1971 | FEB | 03 | 18 | 28 | 12.8 | 38.53 | 21.67 | 39 | 37 | 4.3 |


| DATE |  |  | ORIG.TIME GMT |  |  | $\begin{gathered} \text { LAT } \\ \mathrm{N} \end{gathered}$ | $\begin{gathered} \text { LON } \\ \text { E } \end{gathered}$ | $\begin{gathered} \text { DEPTH } \\ \text { KM } \end{gathered}$ | OBS | $\begin{gathered} \text { MAG } \\ \text { MS } \end{gathered}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1971 | FEB | 09 | 21 | 20 | 35.3 | 38.13 | 22.77 | 40 | 67 | 4.4 |
| 1971 |  | 11 | 16 | 57 | 09.0 | 39.82 | 20.92 | 32 | 29 | 4.4 |
| 1971 |  | 23 | 19 | 41 | 23.0 | 39.62 | 27.32 | 10 | 214 | 5.3 |
| 1971 | MAR | 03 | 19 | 01 | 01.3 | 36.45 | 22.29 | 67 | 52 | 4.3 |
| 1971 |  | 09 | 04 | 58 | 41.3 | 38.74 | 20.44 | 36 | 108 | 4.7 |
| 1971 |  | 15 | 15 | 23 | 19.8 | 37.29 | 24.14 | 41 | 78 | 4.7 |
| 1971 |  | 18 | 16 | 08 | 02.1 | 36.32 | 26.98 | 141 | 81 | 4.3 |
| 1971 |  | 24 | 05 | 11 | 10.0 | 37.89 | 20.37 | 6 | 112 | 4.6 |
| 1971 |  | 25 | 15 | 26 | 34.1 | 34.43 | 24.14 | 44 | 70 | 4.6 |
| 1971 |  | 30 | 00 | 30 | 13.6 | 38.73 | 20.50 | 38 | 32 | 4.4 |
| 1971 |  | 30 | 19 | 40 | 13.3 | 38.98 | 20.79 | 46 | 72 | 4.7 |
| 1971 | APR | 09 | 22 | 09 | 21.5 | 34.76 | 24.23 | 42 | 81 | 4.6 |
| 1971 |  | 19 | 02 | 43 | 50.5 | 38.81 | 20.54 | 08 | 156 | 5.2 |
| 1971 |  | 22 | 09 | 28 | 27.8 | 41.89 | 20.38 | 40 | 97 | 4.7 |
| 1971 | MAY | 01 | 13 | 45 | 27.4 | 40.95 | 27.99 | 13 | 133 | 4.6 |
| 1971 |  | 05 | 01 | 15 | 35.0 | 41.87 | 20.28 | 11 | 68 | 4.4 |
| 1971 |  | 26 | 07 | 09 | 26.0 | 37.10 | 21.70 | 33 | 20 | 4.9 |
| 1971 | JUN | 05 | 19 | 55 | 52.0 | 38.40 | 21.86 | 02 | 29 | 4.3 |
| 1971 |  | 07 | 13 | 34 | 19.1 | 34.32 | 22.80 | 42 | 67 | 4.7 |
| 1971 | JUL | 03 | 04 | 05 | 55.4 | 35.15 | 27.89 | 40 | 116 | 4.5 |
| 1971 |  | 16 | 05 | 50 | 23.9 | 35.11 | 23.07 | 39 | 83 | 4.6 |
| 1971 | AUG | 08 | 19 | 39 | 28.7 | 38.44 | 21.69 | 36 | 50 | 4.3 |
| 1971 |  | 11 | 05 | 37 | 27.3 | 36.81 | 23.96 | 109 | 131 | 5.1 |
| 1971 | SEP | 03 | 13 | 17 | 00.7 | 36.81 | 28.79 | 0 | 22 | 4.6 |
| 1971 |  | 09 | 06 | 51 | 09.1 | 38.21 | 20.16 | 03 | 28 | 4.3 |
| 1971 |  | 11 | 02 | 03 | 11.5 | 38.87 | 22.31 | 5 | 58 | 4.4 |
| 1971 |  | 26 | 05 | 44 | 31.7 | 37.83 | 21.99 | 48 | 45 | 4.3 |
| 1971 |  | 29 | 21 | 02 | 34.3 | 37.02 | 23.28 | 60 | 44 | 4.4 |
| 1971 | OCT | 03 | 23 | 19 | 41.2 | 34.10 | 26.08 | 35 | 111 | 4.7 |
| 1971 |  | 04 | 16 | 35 | 09.5 | 34.16 | 26.18 | 17 | 103 | 5.0 |
| 1971 |  | 13 | 03 | 26 | 26.1 | 34.24 | 26.06 | 17 | 183 | 5.1 |
| 1971 |  | 16 | 09 | 45 | 35.8 | 36.63 | 28.54 | 61 | 55 | 4.9 |
| 1971 | NOV | 04 | 14 | 25 | 49.2 | 35.26 | 22.88 | 41 | 66 | 4.5 |
| 1971 |  | 12 | 12 | 30 | 50.9 | 36.61 | 27.09 | 23 | 55 | 5.3 |
| 1971 |  | 22 | 19 | 26 | 45.7 | 35.34 | 27.81 | 34 | 100 | 4.9 |
| 1971 |  | 27 | 03 | 54 | 28.4 | 39.75 | 25.66 | 24 | 61 | 4.6 |
| 1971 | DEC | 02 | 09 | 40 | 58.4 | 39.23 | 26.45 | 35 | 32 | 4.4 |
| 1971 |  | 17 | 02 | 06 | 04.6 | 34.94 | 23.96 | 25 | 104 | 4.9 |
| 1971 |  | 18 | 02 | 33 | 29.7 | 36.76 | 23.02 | 41 | 76 | 4.6 |
| 1972 | JAN | 12 | 13 | 51 | 20.0 | 35.01 | 23.61 | 46 | 149 | 5.1 |
| 1972 |  | 20 | 00 | 52 | 19.4 | 36.64 | 27.15 | 16 | 79 | 4.6 |
| 1972 |  | 20 | 02 | 15 | 06.9 | 36.64 | 27.23 | 34 | 83 | 4.9 |
| 1972 | FEB | 02 | 21 | 19 | 51.7 | 38.78 | 21.32 | 62 | 72 | 4.5 |
| 1972 |  | 13 | 11 | 27 | 39.8 | 36.07 | 23.98 | 77 | 42 | 4.3 |
| 1972 |  | 13 | 13 | 07 | 12.0 | 36.97 | 24.08 | 27 | 51 | 4.4 |
| 1972 |  | 16 | 00 | 42 | 24.9 | 37.03 | 24.17 | 24 | 49 | 4.3 |
| 1972 |  | 20 | 21 | 38 | 38.8 | 36.84 | 21.75 | 63 | 25 | 4.3 |
| 1972 |  | 28 | 10 | 52 | 47.5 | 37.06 | 24.09 | 31 | 19 | 4.7 |
| 1972 | MAR | 04 | 01 | 09 | 37.3 | 36.78 | 23.11 | 53 | 23 | 4.3 |
| 1972 |  | 16 | 03 | 35 | 35.9 | 37.89 | 23.43 | 142 | 63 | 4.3 |
| 1972 |  | 25 | 06 | 16 | 08.8 | 36.67 | 27.51 | 55 | 32 | 4.3 |
| 1972 |  | 31 | 02 | 58 | 08.0 | 36.43 | 21.26 | 19 | 132 | 4.7 |
| 1972 |  | 31 | 20 | 32 | 01.2 | 36.62 | 27.09 | 18 | 86 | 4.3 |
| 1972 | APR | 23 | 05 | 13 | 33.2 | 38.20 | 20.78 | 46 | 82 | 4.6 |
| 1972 |  | 26 | 06 | 30 | 23.2 | 39.43 | 26.36 | 18 | 144 | 5.1 |


| DATE |  |  | ORIG.TIME GMT |  |  | $\begin{gathered} \text { LAT } \\ \mathrm{N} \end{gathered}$ | $\begin{gathered} \text { LON } \\ \text { E } \end{gathered}$ | $\begin{gathered} \text { DEPTH } \\ \text { KM } \end{gathered}$ | OBS | $\begin{array}{r} \text { MAG } \\ \text { MS } \end{array}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1972 | APR | 26 | 15 | 59 | 44.9 | 39.45 | 26.33 | 25 | 146 | 4.9 |
| 1972 |  | 29 | 18 | 29 | 38.3 | 34.80 | 24.66 | 48 | 178 | 5.3 |
| 1972 | MAY | 04 | 21 | 39 | 57.2 | 35.15 | 23.56 | 13 | 332 | 6.4 |
| 1972 |  | 08 | 09 | 20 | 55.5 | 41.69 | 23.64 | 12 | 168 | 5.1 |
| 1972 |  | 23 | 03 | 14 | 29.9 | 41.50 | 23.64 | 4 | 45 | 4.3 |
| 1972 | JUN | 05 | 10 | 44 | 59.6 | 37.83 | 21.38 | 57 | 2 | 4.3 |
| 1972 |  | 09 | 07 | 42 | 20.5 | 34.73 | 26.55 | 40 | 177 | 5.0 |
| 1972 |  | 15 | 00 | 33 | 24.9 | 38.34 | 22.20 | 33 | 5 | 5.1 |
| 1972 | JUL | 05 | 18 | 04 | 57.9 | 36.96 | 21.88 | 51 | 37 | 4.6 |
| 1972 |  | 08 | 05 | 46 | 15.3 | 41.56 | 23.68 | 38 | 39 | 4.7 |
| 1972 |  | 25 | 01 | 56 | 08.5 | 38.73 | 21.47 | 49 | 108 | 4.7 |
| 1972 | AUG | 12 | 23 | 47 | 57.9 | 41.09 | 22.69 | 12 | 108 | 4.6 |
| 1972 |  | 29 | 02 | 48 | 36.9 | 37.00 | 29.14 | 0 | 27 | 4.3 |
| 1972 | SEP | 03 | 08 | 38 | 46.3 | 39.16 | 27.98 | 30 | 118 | 4.6 |
| 1972 |  | 06 | 18 | 12 | 27.4 | 35.54 | 25.60 | 86 | 46 | 4.4 |
| 1972 |  | 13 | 04 | 13 | 19.7 | 37.96 | 22.38 | 75 | 344 | 6.2 |
| 1972 |  | 16 | 03 | 53 | 26.4 | 40.28 | 19.73 | 15 | 165 | 5.1 |
| 1972 |  | 16 | 14 | 06 | 26.7 | 41.35 | 20.68 | 6 | 86 | 4.4 |
| 1972 |  | 17 | 14 | 07 | 15.3 | 38.35 | 20.27 | 33 | 284 | 5.9 |
| 1972 |  | 17 | 14 | 44 | 10.4 | 38.39 | 20.36 | 39 | 62 | 4.6 |
| 1972 |  | 18 | 08 | 20 | 24.9 | 38.26 | 20.20 | 14 | 75 | 4.3 |
| 1972 |  | 23 | 01 | 53 | 16.5 | 42.25 | 25.31 | 24 | 89 | 4.6 |
| 1972 |  | 26 | 12 | 16 | 59.4 | 34.25 | 26.15 | 23 | 151 | 5.1 |
| 1972 | OCT | 10 | 04 | 31 | 40.3 | 35.24 | 25.42 | 33 | 92 | 4.6 |
| 1972 |  | 10 | 19 | 23 | 38.7 | 35.18 | 25.51 | 41 | 37 | 4.4 |
| 1972 |  | 15 | 22 | 02 | 54.3 | 37.99 | 21.03 | 66 | 103 | 4.4 |
| 1972 |  | 16 | 23 | 39 | 37.4 | 38.24 | 20.39 | 34 | 48 | 4.3 |
| 1972 |  | 23 | 09 | 56 | 27.5 | 37.78 | 26.32 | 28 | 35 | 4.3 |
| 1972 |  | 30 | 14 | 32 | 10.7 | 38.28 | 20.35 | 13 | 198 | 5.4 |
| 1972 | NOV | 05 | 19 | 25 | 42.6 | 35.03 | 24.77 | 31 | 198 | 5.3 |
| 1972 |  | 15 | 12 | 21 | 47.4 | 34.10 | 26.30 | 0 | 11 | 4.3 |
| 1972 |  | 17 | 02 | 42 | 35.8 | 37.39 | 20.33 | 41 | 90 | 4.3 |
| 1972 |  | 20 | 03 | 30 | 27.2 | 39.42 | 21.68 | 26 | 99 | 4.9 |
| 1972 |  | 24 | 03 | 48 | 34.2 | 39.39 | 20.43 | 09 | 209 | 5.4 |
| 1972 |  | 28 | 13 | 26 | 11.8 | 33.80 | 27.77 | 02 | 81 | 4.7 |
| 1972 | DEC | 02 | 13 | 28 | 22.8 | 35.28 | 27.06 | 36 | 168 | 5.3 |
| 1972 |  | 05 | 12 | 00 | 15.0 | 39.14 | 23.64 | 39 | 56 | 4.4 |
| 1972 |  | 14 | 17 | 50 | 21.7 | 37.99 | 20.39 | 10 | 66 | 4.3 |
| 1972 |  | 17 | 12 | 44 | 30.7 | 34.27 | 26.22 | 38 | 82 | 4.6 |
| 1972 |  | 19 | 19 | 34 | 30.0 | 35.29 | 27.74 | 40 | 120 | 4.7 |
| 1973 | JAN | 05 | 05 | 49 | 17.6 | 35.80 | 21.91 | 34 | 256 | 5.5 |
| 1973 |  | 10 | 03 | 24 | 12.0 | 37.69 | 21.42 | 45 | 148 | 4.9 |
| 1973 |  | 16 | 22 | 45 | 17.1 | 35.11 | 22.68 | 35 | 125 | 4.6 |
| 1973 |  | 23 | 11 | 46 | 43.0 | 34.28 | 24.98 | 37 | 101 | 4.7 |
| 1973 |  | 26 | 07 | 50 | 11.0 | 35.74 | 22.08 | 41 | 123 | 4.9 |
| 1973 | FEB | 20 | 05 | 55 | 15.0 | 34.38 | 23.88 | 19 | 101 | 4.5 |
| 1973 |  | 26 | 22 | 23 | 11.8 | 39.84 | 20.30 | 44 | 67 | 4.4 |
| 1973 | MAR | 12 | 20 | 30 | 43.9 | 35.86 | 21.81 | 45 | 114 | 4.7 |
| 1973 | APR | 06 | 14 | 13 | 57.3 | 34.41 | 25.18 | 37 | 199 | 5.3 |
| 1973 |  | 07 | 19 | 30 | 09.0 | 41.47 | 19.90 | 20 | 88 | 4.5 |
| 1973 |  | 16 | 00 | 05 | 42.2 | 34.64 | 25.01 | 44 | 131 | 4.5 |
| 1973 |  | 19 | 22 | 13 | 55.0 | 38.29 | 26.94 | 17 | 68 | 4.3 |
| 1973 |  | 22 | 13 | 39 | 44.4 | 35.07 | 23.45 | 46 | 118 | 4.5 |
| 1973 | MAY | 22 | 15 | 26 | 07.0 | 36.70 | 22.10 | 0 | 38 | 4.3 |
| 1973 | JUN | 12 | 11 | 01 | 52.3 | 34.18 | 26.15 | 47 | 63 | 4.3 |


| DATE |  |  | ORIG.TIME GMT |  |  | $\begin{aligned} & \text { LAT } \\ & \mathrm{N} \end{aligned}$ | $\begin{gathered} \text { LON } \\ \text { E } \end{gathered}$ | $\begin{gathered} \text { DEPTH } \\ \text { KM } \end{gathered}$ | OBS | $\begin{gathered} \text { MAG } \\ \text { MS } \end{gathered}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1973 | JUN | 26 | 19 | 05 | 23.4 | 34.36 | 26.13 | 50 | 158 | 4.9 |
| 1973 | JUL | 14 | 12 | 38 | 19.2 | 37.93 | 21.21 | 42 | 138 | 4.6 |
| 1973 |  | 21 | 12 | 51 | 55.0 | 34.94 | 24.72 | 33 | 73 | 4.4 |
| 1973 | AUG | 08 | 08 | 23 | 48.7 | 41.69 | 19.43 | 39 | 89 | 4.5 |
| 1973 | SEP | 12 | 09 | 36 | 50.5 | 40.72 | 21.01 | 91 | 75 | 4.4 |
| 1973 | OCT | 06 | 21 | 19 | 59.0 | 34.80 | 26.34 | 39 | 112 | 4.7 |
| 1973 |  | 10 | 11 | 05 | 34.8 | 34.34 | 28.45 | 63 | 78 | 4.6 |
| 1973 |  | 13 | 06 | 00 | 35.0 | 34.70 | 26.39 | 52 | 99 | 4.6 |
| 1973 |  | 14 | 18 | 07 | 06.4 | 34.68 | 26.31 | 51 | 152 | 4.8 |
| 1973 |  | 19 | 00 | 16 | 06.5 | 37.01 | 21.71 | 53 | 39 | 4.3 |
| 1973 | NOV | 04 | 15 | 52 | 12.6 | 38.87 | 20.54 | 13 | 277 | 5.9 |
| 1973 |  | 04 | 16 | 11 | 38.7 | 38.85 | 20.49 | 35 | 118 | 4.9 |
| 1973 |  | 12 | 00 | 07 | 11.3 | 35.35 | 27.74 | 47 | 160 | 4.7 |
| 1973 |  | 12 | 00 | 11 | 49.4 | 35.40 | 27.65 | 21 | 193 | 5.3 |
| 1973 |  | 14 | 09 | 33 | 57.4 | 35.29 | 27.74 | 42 | 102 | 4.5 |
| 1973 |  | 17 | 15 | 37 | 20.8 | 35.88 | 21.99 | 45 | 53 | 4.4 |
| 1973 |  | 19 | 07 | 28 | 59.9 | 35.34 | 27.71 | 60 | 118 | 4.7 |
| 1973 |  | 20 | 13 | 02 | 34.2 | 39.31 | 23.80 | 0 | 134 | 4.6 |
| 1973 |  | 29 | 10 | 57 | 44.3 | 35.18 | 23.81 | 37 | 309 | 5.9 |
| 1973 | DEC | 05 | 03 | 50 | 50.4 | 35.36 | 26.42 | 70 | 204 | 5.2 |
| 1973 |  | 24 | 13 | 53 | 54.8 | 35.08 | 27.69 | 53 | 68 | 4.3 |
| 1974 | JAN | 24 | 09 | 40 | 17.5 | 38.23 | 20.13 | 56 | 84 | 4.5 |
| 1974 |  | 27 | 21 | 06 | 17.4 | 35.04 | 25.38 | 35 | 119 | 4.7 |
| 1974 |  | 29 | 15 | 12 | 44.8 | 38.29 | 21.85 | 34 | 66 | 4.3 |
| 1974 | FEB | 01 | 00 | 01 | 02.2 | 38.55 | 27.22 | 24 | 176 | 5.4 |
| 1974 |  | 05 | 15 | 05 | 25.0 | 36.74 | 26.86 | 156 | 145 | 5.3 |
| 1974 |  | 23 | 01 | 28 | 47.0 | 38.02 | 22.00 | 40 | 82 | 4.4 |
| 1974 | MAR | 08 | 02 | 33 | 52.8 | 34.66 | 24.74 | 47 | 119 | 4.7 |
| 1974 |  | 09 | 04 | 12 | 07.5 | 34.59 | 24.80 | 42 | 96 | 4.4 |
| 1974 |  | 12 | 18 | 21 | 34.7 | 36.76 | 26.40 | 45 | 113 | 4.9 |
| 1974 |  | 13 | 17 | 20 | 45.2 | 34.60 | 24.75 | 46 | 144 | 4.6 |
| 1974 |  | 22 | 17 | 02 | 20.0 | 40.65 | 20.55 | 27 | 100 | 4.4 |
| 1974 | APR | 01 | 00 | 22 | 39.3 | 35.61 | 22.44 | 58 | 112 | 4.5 |
| 1974 |  | 07 | 14 | 22 | 48.7 | 34.75 | 24.70 | 38 | 204 | 4.7 |
| 1974 | MAY | 09 | 17 | 02 | 24.0 | 36.62 | 27.22 | 26 | 44 | 4.3 |
| 1974 |  | 19 | 22 | 01 | 09.7 | 35.47 | 26.31 | 84 | 238 | 4.9 |
| 1974 | JUN | 18 | 08 | 26 | 11.4 | 38.45 | 20.43 | 24 | 138 | 4.7 |
| 1974 |  | 22 | 23 | 30 | 12.1 | 41.25 | 23.05 | 08 | 159 | 5.1 |
| 1974 | JUL | 09 | 02 | 32 | 15.4 | 36.57 | 28.48 | 49 | 195 | 5.0 |
| 1974 | AUG | 14 | 16 | 05 | 20.1 | 35.44 | 23.04 | 64 | 53 | 4.3 |
| 1974 |  | 20 | 23 | 52 | 40.7 | 38.23 | 20.65 | 44 | 37 | 4.4 |
| 1974 | SEP | 05 | 11 | 34 | 37.4 | 35.71 | 24.75 | 53 | 118 | 4.4 |
| 1974 |  | 11 | 05 | 12 | 57.0 | 40.03 | 19.64 | 28 | 59 | 4.5 |
| 1974 |  | 13 | 18 | 24 | 57.4 | 40.48 | 23.39 | 8 | 91 | 4.4 |
| 1974 |  | 17 | 05 | 10 | 31.8 | 40.29 | 20.63 | 17 | 187 | 5.0 |
| 1974 |  | 18 | 09 | 07 | 02.0 | 40.21 | 20.78 | 3 | 56 | 4.3 |
| 1974 |  | 28 | 01 | 34 | 59.1 | 34.87 | 23.86 | 38 | 57 | 4.3 |
| 1974 | OCT | 20 | 11 | 25 | 50.3 | 39.57 | 18.83 | 33 | 163 | 5.0 |
| 1974 |  | 25 | 11 | 43 | 35.5 | 34.67 | 23.37 | 41 | 107 | 5.0 |
| 1974 | NOV | 09 | 06 | 00 | 45.0 | 38.92 | 20.32 | 24 | 83 | 4.9 |
| 1974 |  | 14 | 13 | 22 | 34.7 | 38.50 | 23.08 | 27 | 185 | 5.0 |
| 1974 |  | 14 | 14 | 26 | 46.6 | 38.48 | 23.01 | 6 | 182 | 5.1 |
| 1974 |  | 14 | 15 | 29 | 46.8 | 38.50 | 23.15 | 35 | 174 | 5.0 |
| 1974 |  | 23 | 07 | 52 | 28.2 | 39.77 | 18.81 | 38 | 36 | 4.3 |
| 1974 |  | 23 | 18 | 46 | 36.0 | 39.74 | 18.94 | 49 | 83 | 4.7 |


| DATE |  |  | $\begin{aligned} & \text { ORIG.TIME } \\ & \text { GMT } \end{aligned}$ |  |  | $\begin{gathered} \text { LAT } \\ \mathrm{N} \end{gathered}$ | $\begin{gathered} \text { LON } \\ \mathrm{E} \end{gathered}$ | $\begin{gathered} \text { DEPTH } \\ \text { KM } \end{gathered}$ | OBS | $\begin{array}{r} \text { MAG } \\ \text { MS } \end{array}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1974 | NOV | 29 | 06 | 35 | 33.4 | 35.40 | 27.89 | 49 | 107 | 4.6 |
| 1974 | DEC | 01 | 12 | 09 | 29.5 | 39.48 | 26.35 | 36 | 95 | 4.5 |
| 1974 |  | 02 | 23 | 02 | 19.0 | 38.42 | 22.31 | 34 | 128 | 4.9 |
| 1974 |  | 14 | 02 | 36 | 37.7 | 38.19 | 20.75 | 32 | 243 | 5.4 |
| 1974 |  | 14 |  | 29 | 22.0 | 38.40 | 20.40 | 37 | 122 | 4.5 |
| 1974 |  | 18 | 21 | 30 | 54.8 | 39.95 | 23.86 | 33 | 46 | 4.4 |
| 1974 |  | 20 | 15 | 09 | 32.6 | 39.64 | 20.53 | 47 | 159 | 4.8 |
| 1974 |  | 20 | 16 | 02 | 06.2 | 39.71 | 20.74 | 47 | 46 | 4.3 |
| 1974 |  | 26 | 14 | 01 | 51.1 | 35.10 | 22.75 | 40 | 57 | 4.3 |
| 1975 | JAN | 03 | 01 | 59 | 44.4 | 35.62 | 27.34 | 42 | 156 | 4.9 |
| 1975 |  | 08 | 19 | 32 | 34.1 | 38.24 | 22.65 | 26 | 246 | 5.7 |
| 1975 |  | 08 | 19 | 58 | 16.0 | 38.10 | 22.75 | 33 | 110 | 4.5 |
| 1975 |  | 09 | 18 | 53 | 44.3 | 34.78 | 24.03 | 41 | 140 | 4.6 |
| 1975 |  | 11 | 18 | 06 | 55.0 | 34.76 | 23.98 | 42 | 41 | 4.3 |
| 1975 |  | 24 | 16 | 33 | 04.4 | 41.14 | 19.77 | 46 | 111 | 4.5 |
| 1975 |  | 25 | 14 | 14 | 05.0 | 37.84 | 20.01 | 1 | 62 | 4.6 |
| 1975 |  | 26 | 05 | 30 | 52.0 | 36.72 | 24.44 | 32 | 152 | 4.9 |
| 1975 |  | 26 | 06 | 24 | 42.7 | 38.75 | 20.13 | 0 | 22 | 4.3 |
| 1975 | FEB | 02 | 21 | 12 | 20.2 | 40.48 | 21.39 | 40 | 69 | 4.4 |
| 1975 |  | 09 | 12 | 36 | 05.0 | 38.71 | 26.16 | 27 | 79 | 4.5 |
| 1975 |  | 15 | 10 | 23 | 21.0 | 35.77 | 26.95 | 46 | 132 | 4.7 |
| 1975 |  | 20 | 13 | 55 | 35.9 | 38.54 | 20.39 | 4 | 59 | 4.4 |
| 1975 |  | 28 | 19 | 51 | 09.0 | 40.66 | 22.52 | 29 | 87 | 4.5 |
| 1975 | MAR | 17 | 02 | 06 | 39.1 | 40.48 | 26.03 | 2 | 114 | 4.6 |
| 1975 |  | 17 | 05 | 11 | 16.5 | 40.48 | 25.95 | 22 | 219 | 5.0 |
| 1975 |  | 17 | 05 | 17 | 47.1 | 40.40 | 26.24 | 5 | 138 | 4.9 |
| 1975 |  | 17 | 05 | 35 | 17.6 | 40.48 | 26.08 | 18 | 252 | 5.8 |
| 1975 |  | 25 | 02 | 52 | 52.6 | 34.66 | 23.69 | 0 | 44 | 4.4 |
| 1975 |  | 27 | 05 | 15 | 07.9 | 40.45 | 26.12 | 15 | 337 | 6.7 |
| 1975 |  | 27 | 06 | 15 | 46.0 | 40.41 | 26.23 | 22 | 100 | 4.7 |
| 1975 |  | 27 | 19 | 42 | 42.5 | 40.48 | 26.08 | 5 | 80 | 4.6 |
| 1975 |  | 30 | 13 | 03 | 17.6 | 40.57 | 26.36 | 0 | 64 | 4.4 |
| 1975 | APR | 01 | 08 | 20 | 02.0 | 38.53 | 23.25 | 8 | 44 | 4.4 |
| 1975 |  | 04 | 05 | 16 | 16.5 | 38.11 | 21.98 | 56 | 290 | 5.7 |
| 1975 |  | 05 | 03 | 30 | 24.0 | 36.27 | 21.68 | 58 | 38 | 4.3 |
| 1975 |  | 23 | 01 | 08 | 08.0 | 40.40 | 26.04 | 20 | 69 | 4.4 |
| 1975 |  | 24 | 22 | 58 | 16.6 | 37.48 | 22.60 | 68 | 164 | 4.9 |
| 1975 | MAY | 03 | 03 | 16 | 02.1 | 37.70 | 21.18 | 0 | 48 | 4.3 |
| 1975 |  | 11 | 23 | 11 | 47.9 | 37.36 | 23.84 | 43 | 81 | 4.7 |
| 1975 |  | 13 | 00 | 22 | 53.0 | 38.19 | 22.72 | 45 | 105 | 4.6 |
| 1975 |  | 19 | 03 | 26 | 20.4 | 38.34 | 22.34 | 26 | 153 | 4.7 |
| 1975 |  | 19 | 23 | 25 | 40.9 | 39.61 | 19.74 | 47 | 33 | 4.5 |
| 1975 | JUN | 02 | 03 | 19 | 08.0 | 36.47 | 26.52 | 31 | 92 | 4.7 |
| 1975 |  | 07 | 17 | 36 | 36.9 | 34.32 | 26.22 | 51 | 61 | 4.5 |
| 1975 |  | 08 | 17 | 22 | 28.3 | 34.60 | 23.45 | 47 | 67 | 4.5 |
| 1975 |  | 30 | 13 | 26 | 54.6 | 38.49 | 21.62 | 3 | 230 | 5.1 |
| 1975 |  | 30 | 18 | 40 | 32.0 | 38.45 | 21.61 | 41 | 84 | 4.6 |
| 1975 | JUL | 25 | 19 | 17 | 11.9 | 38.41 | 21.87 | 38 | 176 | 4.9 |
| 1975 |  | 29 | 15 | 07 | 12.8 | 34.84 | 24.95 | 47 | 169 | 4.8 |
| 1975 | AUG | 21 | 15 | 29 | 18.5 | 40.14 | 19.80 | 46 | 43 | 4.4 |
| 1975 | SEP | 12 | 13 | 10 | 19.6 | 36.27 | 21.90 | 43 | 212 | 5.0 |
| 1975 |  | 13 | 14 | 30 | 37.6 | 38.47 | 22.02 | 40 | 150 | 4.9 |
| 1975 |  | 16 | 05 | 06 | 19.1 | 41.54 | 19.33 | 25 | 217 | 5.1 |
| 1975 |  | 16 | 18 | 45 | 48.2 | 41.52 | 19.28 | 46 | 79 | 4.6 |
| 1975 |  | 17 | 23 | 04 | 07.2 | 36.37 | 23.06 | 35 | 201 | 5.1 |


| DATE |  |  | $\begin{aligned} & \text { ORIG.TIME } \\ & \text { GMT } \end{aligned}$ |  |  | $\frac{\text { LAT }}{\mathrm{N}}$ | $\begin{gathered} \text { LON } \\ \text { E } \end{gathered}$ | $\begin{gathered} \text { DEPTH } \\ \text { KM } \end{gathered}$ | OBS | MAG MS |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1975 | SEP | 17 | 23 | 44 | 19.0 | 38.16 | 20.42 | 15 | 113 | 4.5 |
| 1975 |  | 20 | 05 | 22 | 18.2 | 34.60 | 26.41 | 60 | 89 | 4.9 |
| 1975 |  | 22 | 00 | 44 | 56.4 | 35.20 | 26.26 | 55 | 312 | 5.7 |
| 1975 |  | 23 | 21 | 34 | 14.1 | 36.60 | 26.76 | 158 | 86 | 4.6 |
| 1975 | OCT | 02 | 15 | 59 | 45.1 | 40.16 | 20.49 | 45 | 30 | 4.1 |
| 1975 |  | 06 | 21 | 27 | 54.0 | 34.13 | 25.22 | 24 | 80 | 4.3 |
| 1975 |  | 12 | 08 | 23 | 12.6 | 37.91 | 23.12 | 35 | 159 | 5.0 |
| 1975 |  | 16 | 17 |  | 49.3 | 38.28 | 21.95 | 40 | 40 | 4.0 |
| 1975 |  | 25 | 07 | 25 | 22.3 | 37.61 | 22.11 | 36 | 53 | 4.3 |
| 1975 |  | 28 | 02 | 43 | 23.8 | 35.34 | 23.19 | 67 | 31 | 4.0 |
| 1975 | NOV | 12 | 09 | 03 | 48.8 | 36.28 | 28.15 | 64 | 263 | 5.5 |
| 1975 |  | 13 | 03 | 07 | 20.5 | 33.42 | 22.84 | 0 | 203 | 5.3 |
| 1975 |  | 13 | 23 | 30 | 36.0 | 37.51 | 21.09 | 50 | 47 | 4.0 |
| 1975 |  | 17 | 14 | 36 | 41.0 | 34.29 | 23.34 | 2 | 64 | 4.1 |
| 1975 |  | 22 | 10 | 06 | 08.4 | 39.92 | 20.11 | 34 | 223 | 5.3 |
| 1975 |  | 30 | 04 | 20 | 46.8 | 37.02 | 21.25 | 52 | 47 | 4.3 |
| 1975 | DEC | 06 | 08 | 19 | 10.0 | 38.50 | 25.69 | 23 | 43 | 4.0 |
| 1975 |  | 08 | 23 | 03 | 38.0 | 36.43 | 27.90 | 5 | 18 | 4.1 |
| 1975 |  | 10 | 18 | 12 | 28.3 | 34.14 | 25.72 | 44 | 56 | 4.3 |
| 1975 |  | 16 | 08 | 08 | 29.4 | 39.44 | 20.45 | 50 | 26 | 4.4 |
| 1975 |  | 17 | 02 | 52 | 17.2 | 34.09 | 26.20 | 42 | 44 | 4.4 |
| 1975 |  | 21 | 15 | 37 | 16.6 | 35.62 | 26.78 | 98 | 62 | 4.4 |
| 1975 |  | 21 | 16 | 07 | 51.1 | 38.47 | 21.67 | 2 | 242 | 5.5 |
| 1975 |  | 21 | 21 | 05 | 15.0 | 38.38 | 21.85 | 0 | 41 | 4.2 |
| 1975 |  | 24 | 17 | 04 | 51.8 | 36.80 | 21.77 | 65 | 67 | 4.3 |
| 1975 |  | 31 | 06 | 34 | 04.0 | 38.42 | 21.70 | 28 | 53 | 4.1 |
| 1975 |  | 31 | 09 | 45 | 47.3 | 38.52 | 21.67 | 19 | 295 | 5.5 |
| 1975 |  | 31 | 14 | 53 | 42.0 | 38.45 | 21.70 | 26 | 80 | 4.5 |
| 1975 |  | 31 | 13 | 51 | 21.0 | 38.48 | 21.65 | 23 | 117 | 4.6 |
| 1975 |  | 31 | 15 | 14 | 36.2 | 38.38 | 21.76 | 0 | 54 | 4.2 |
| 1975 |  | 31 | 22 | 54 | 50.0 | 38.56 | 21.69 | 29 | 46 | 4.2 |
| 1976 | JAN | 01 | 00 | 04 | 06.0 | 38.42 | 21.72 | 18 | 122 | 4.7 |
| 1976 |  | 02 | 22 | 44 | 42.0 | 38.42 | 21.78 | 0 | 90 | 4.6 |
| 1976 |  | 03 | 05 | 54 | 35.1 | 38.37 | 21.54 | 27 | 33 | 4.2 |
| 1976 |  | 03 | 13 | 16 | 27.4 | 38.41 | 21.81 | 14 | 81 | 4.5 |
| 1976 |  | 03 | 13 | 50 | 08.0 | 38.36 | 21.77 | 2 | 35 | 4.0 |
| 1976 |  | 03 | 15 | 03 | 53.8 | 38.39 | 21.79 | 37 | 78 | 4.5 |
| 1976 |  | 10 | 07 | 11 | 20.0 | 36.80 | 27.92 | 31 | 70 | 4.5 |
| 1976 |  | 14 | 10 | 31 | 02.3 | 38.39 | 21.95 | 10 | 70 | 4.6 |
| 1976 |  | 18 | 15 | 10 | 28.4 | 38.81 | 20.51 | 5 | 254 | 5.7 |
| 1976 |  | 21 | 22 | 18 | 00.2 | 37.76 | 21.02 | 33 | 42 | 4.2 |
| 1976 | FEB | 01 | 14 | 46 | 07.0 | 36.12 | 22.28 | 60 | 50 | 4.3 |
| 1976 |  | 01 | 23 | 33 | 08.6 | 36.60 | 22.20 | 67 | 42 | 4.2 |
| 1976 |  | 02 | 12 | 13 | 01.0 | 39.78 | 20.60 | 36 | 43 | 4.4 |
| 1976 |  | 10 | 09 | 52 | 09.2 | 36.82 | 27.93 | 39 | 76 | 4.7 |
| 1976 |  | 18 | 11 | 41 | 37.9 | 38.70 | 20.44 | 48 | 87 | 4.5 |
| 1976 |  | 22 | 12 | 02 | 53.0 | 39.38 | 22.08 | 19 | 234 | 5.1 |
| 1976 |  | 22 | 22 | 01 | 48.8 | 39.39 | 22.13 | 34 | 143 | 4.8 |
| 1976 |  | 22 | 22 | 54 | 34.8 | 39.39 | 22.14 | 23 | 146 | 4.8 |
| 1976 |  | 22 | 22 | 56 | 34.2 | 39.33 | 21.91 | 69 | 31 | 4.3 |
| 1976 |  | 23 | 16 | 18 | 28.0 | 38.27 | 25.58 | 4 | 177 | 4.8 |
| 1976 |  | 23 | 17 | 12 | 18.8 | 34.90 | 26.90 | 70 | 33 | 4.1 |
| 1976 |  | 26 | 19 | 32 | 38.1 | 38.24 | 26.38 | 10 | 57 | 4.4 |
| 1976 | MAR | 02 | 19 | 41 | 34.1 | 40.66 | 19.59 | 11 | 191 | 4.7 |
| 1976 | APR | 04 | 22 | 26 | 27.4 | 34.83 | 26.42 | 36 | 44 | 4.2 |


| DATE |  |  | ORIG.TIME GMT |  |  | $\frac{\text { LAT }}{\mathrm{N}}$ | $\begin{gathered} \text { LON } \\ \text { E } \end{gathered}$ | $\begin{gathered} \text { DEPTH } \\ \text { KM } \end{gathered}$ | OBS | $\begin{array}{r} \text { MAG } \\ \text { MS } \end{array}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1976 | APR | 19 | 00 | 27 | 50.5 | 35.52 | 24.66 | 64 | 165 | 5.0 |
| 1976 |  | 30 | 16 | 09 | 30.2 | 35.97 | 24.66 | 98 | 98 | 4.8 |
| 1976 | MAY | 01 | 07 | 26 | 27.0 | 37.12 . | 27.72 | 33 | 24 | 4.2 |
| 1976 |  | 06 | 17 | 59 | 02.6 | 34.69 | 23.86 | 46 | 42 | 4.7 |
| 1976 |  | 11 | 16 | 59 | 48.2 | 37.56 | 20.35 | 33 | 202 | 6.3 |
| 1976 |  | 11 | 17 | 10 | 10.6 | 37.33 | 20.46 | 20 | 81 | 5.4 |
| 1976 |  | 13 | 00 | 44 | 15.0 | 39.72 | 20.32 | 59 | 32 | 4.6 |
| 1976 |  | 13 | 20 | 44 | 52.5 | 36.84 | 21.39 | 51 | 41 | 4.7 |
| 1976 |  | 13 | 22 | 18 | 05.3 | 37.37 | 20.54 | 52 | 24 | 4.1 |
| 1976 |  | 15 | 02 | 47 | 31.6 | 35.46 | 27.06 | 45 | 11 | 4.0 |
| 1976 |  | 15 | 03 | 03 | 08.3 | 36.33 | 23.30 | 52 | 33 | 4.4 |
| 1976 |  | 18 | 08 | 30 | 21.4 | 35.03 | 25.39 | 73 | 102 | 4.9 |
| 1976 |  | 30 | 16 | 26 | 42.3 | 37.44 | 20.63 | 33 | 14 | 4.0 |
| 1976 | JUN | 05 | 20 | 30 | 11.5 | 38.58 | 22.21 | 51 | 16 | 4.0 |
| 1976 |  | 05 | 22 | 21 | 55.1 | 37.82 | 21.93 | 79 | 17 | 4.3 |
| 1976 |  | 10 | 05 | 55 | 22.4 | 35.50 | 23.74 | 92 | 18 | 4.3 |
| 1976 |  | 11 | 18 | 26 | 14.7 | 39.40 | 20.39 | 53 | 26 | 4.6 |
| 1976 |  | 12 | 00 | 59 | 16.9 | 37.54 | 20.55 | 8 | 160 | 5.3 |
| 1976 |  | 12 | 02 | 41 | 43.5 | 37.39 | 20.56 | 38 | 19 | 4.2 |
| 1976 |  | 12 | 04 | 54 | 48.2 | 37.39 | 20.44 | 43 | 32 | 4.6 |
| 1976 |  | 13 | 00 | 20 | 00.5 | 37.48 | 20.61 | 48 | 21 | 4.3 |
| 1976 |  | 15 | 00 | 07 | 54.4 | 34.22 | 24.77 | 33 | 21 | 4.0 |
| 1976 |  | 15 | 12 | 46 | 52.5 | 37.42 | 20.59 | 41 | 19 | 4.4 |
| 1976 |  | 20 | 04 | 51 | 17.0 | 38.53 | 22.12 | 51 | 17 | 4.7 |
| 1976 |  | 21 | 10 | 59 | 14.0 | 34.67 | 24.12 | 22 | 19 | 4.2 |
| 1976 |  | 25 | 07 | 01 | 08.0 | 35.09 | 23.31 | 33 | 94 | 5.4 |
| 1976 | JUL | 02 | 05 | 16 | 42.4 | 39.23 | 21.72 | 36 | 31 | 4.9 |
| 1976 |  | 08 | 15 | 20 | 39.7 | 37.61 | 20.71 | 49 | 22 | 4.4 |
| 1976 |  | 13 | 20 | 37 | 25.8 | 37.39 | 20.49 | 49 | 19 | 4.3 |
| 1976 |  | 18 | 02 | 17 | 11.9 | 36.71 | 23.35 | 66 | 20 | 4.1 |
| 1976 |  | 18 | 13 | 30 | 47.4 | 38.64 | 20.42 | 31 | 21 | 4.4 |
| 1976 |  | 23 | 20 | 51 | 02.9 | 37.99 | 21.53 | 33 | 14 | 4.3 |
| 1976 |  | 02 | 05 | 15 | 35.0 | 35.55 | 25.98 | 125 | 49 | 4.7 |
| 1976 |  | 17 | 17 | 37 | 56.6 | 36.88 | 27.05 | 167 | 96 | 5.2 |
| 1976 |  | 18 | 00 | 58 | 06.2 | 36.59 | 26.94 | 154 | 26 | 4.2 |
| 1976 |  | 18 | 17 | 06 | 35.5 | 36.78 | 27.42 | 163 | 35 | 4.7 |
| 1976 |  | 19 | 01 | 12 | 36.7 | 37.70 | 28.89 | 3 | 94 | 5.1 |
| 1976 |  | 19 | 22 | 36 | 26.6 | 39.23 | 22.27 | 58 | 26 | 4.2 |
| 1976 |  | 22 | 13 | 28 | 49.0 | 39.33 | 29.08 | 14 | 111 | 4.9 |
| 1976 |  | 22 | 17 | 18 | 46.9 | 37.41 | 20.56 | 63 | 35 | 4.3 |
| 1976 |  | 23 | 03 | 15 | 15.5 | 38.34 | 20.67 | 52 | 52 | 4.4 |
| 1976 | SEP | 12 | 00 | 42 | 19.3 | 36.67 | 26.98 | 168 | 22 | 4.2 |
| 1976 |  | 15 | 03 | 36 | 40.3 | 40.49 | 27.48 | 10 | 9 | 4.4 |
| 1976 |  | 30 | 00 | 33 | 02.0 | 37.53 | 20.37 | 43 | 122 | 5.0 |
| 1976 |  | 30 | 17 | 09 | 58.6 | 38.48 | 21.62 | 16 | 18 | 4.3 |
| 1976 | OCT | 02 | 12 | 45 | 29.4 | 37.12 | 21.22 | 68 | 36 | 4.3 |
| 1976 |  | 17 | 21 | 26 | 36.9 | 35.11 | 25.51 | 33 | 10 | 4.4 |
| 1976 |  | 21 | 12 | 48 | 11.8 | 35.91 | 26.98 | 99 | 68 | 4.9 |
| 1976 |  | 23 | 12 | 08 | 26.9 | 34.25 | 25.44 | 81 | 14 | 4.2 |
| 1976 |  | 27 | 00 | 38 | 45.9 | 38.42 | 22.57 | 33 | 21 | 4.2 |
| 1976 |  | 27 | 09 | 28 | 48.0 | 34.25 | 25.93 | 58 | 22 | 4.2 |
| 1976 |  | 28 | 01 | 04 | 46.7 | 34.85 | 26.27 | 42 | 21 | 4.3 |
| 1976 |  | 31 | 08 | 59 | 35.6 | 38.11 | 22.48 | 29 | 14 | 4.2 |
| 1976 | NOV | 09 | 16 | 02 | 19.6 | 35.67 | 24.00 | 69 | 29 | 4.5 |
| 1976 |  | 11 | 17 | 08 | 00.8 | 35.05 | 22.99 | 33 | 10 | 4.2 |


| DATE |  |  | ORIG.TIME GMT |  |  | $\underset{\mathrm{N}}{\mathrm{LAT}}$ | $\begin{gathered} \text { LON } \\ \text { E } \end{gathered}$ | $\begin{gathered} \text { DEPTH } \\ \text { KM } \end{gathered}$ | OBS | $\begin{array}{r} \text { MAG } \\ \text { MS } \end{array}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1976 | NOV | 12 | 09 | 51 | 10.8 | 38.62 | 26.73 | 24 | 34 | 4.3 |
| 1976 |  | 12 | 09 | 55 | 33.4 | 38.57 | 26.71 | 6 | 54 | 4.7 |
| 1976 |  | 13 | 06 | 09 | 46.5 | 34.97 | 23.32 | 29 | 88 | 5.3 |
| 1976 |  | 13 | 11 | 54 | 48.1 | 38.61 | 26.72 | 33 | 15 | 4.4 |
| 1976 |  | 18 | 12 | 09 | 22.8 | 36.86 | 24.56 | 65 | 22 | 4.2 |
| 1976 |  | 21 | 23 | 10 | 33.3 | 38.38 | 26.86 | 33 | 13 | 4.1 |
| 1976 |  | 22 | 03 | 31 | 15.7 | 38.80 | 22.01 | 33 | 20 | 4.4 |
| 1976 |  | 22 | 11 | 53 | 06.1 | 37.23 | 20.14 | 33 | 10 | 4.1 |
| 1976 |  | 24 | 20 | 57 | 05.5 | 37.21 | 20.33 | 33 | 25 | 4.2 |
| 1976 |  | 26 | 21 | 30 | 43.4 | 36.28 | 27.27 | 42 | 10 | 4.2 |
| 1976 |  | 27 | 15 | 36 | 41.4 | 37.50 | 20.08 | 33 | 10 | 4.4 |
| 1976 |  | 27 | 18 | 44 | 45.4 | 37.21 | 20.10 | 33 | 12 | 4.2 |
| 1976 |  | 28 | 19 | 25 | 17.3 | 37.30 | 20.34 | 24 | 106 | 4.9 |
| 1976 |  | 29 | 17 | 16 | 07.3 | 34.86 | 25.69 | 46 | 17 | 4.8 |
| 1976 | DEC | 15 | 16 | 06 | 24.7 | 35.48 | 23.33 | 49 | 50 | 4.4 |
| 1976 |  | 24 | 21 | 48 | 39.7 | 36.24 | 26.76 | 160 | 14 | 4.3 |
| 1976 |  | 26 | 08 | 41 | 27.3 | 38.93 | 20.17 | 33 | 9 | 4.4 |
| 1976 |  | 27 | 07 | 54 | 13.3 | 39.13 | 20.56 | 32 | 78 | 4.9 |
| 1976 |  | 30 | 15 | 12 | 37.9 | 37.91 | 22.82 | 34 | 52 | 4.7 |
| 1976 |  | 31 | 00 | 39 | 57.0 | 37.91 | 22.91 | 54 | 9 | 4.4 |
| 1977 | JAN | 03 | 00 | 44 | 07.8 | 38.21 | 23.11 | 23 | 8 | 4.4 |
| 1977 |  | 16 | 09 | 16 | 49.1 | 37.89 | 22.93 | 52 | 49 | 4.6 |
| 1977 |  | 18 | 20 | 46 | 51.9 | 35.78 | 29.41 | 57 | 39 | 4.3 |
| 1977 |  | 24 | 06 | 38 | 04.3 | 34.85 | 25.85 | 106 | 13 | 4.0 |
| 1977 |  | 25 | 23 | 54 | 18.1 | 39.36 | 28.13 | 23 | 25 | 4.6 |
| 1977 | FEB | 17 | 03 | 17 | 09.6 | 35.50 | 22.20 | 33 | 8 | 4.0 |
| 1977 |  | 21 | 17 | 44 | 15.3 | 37.48 | 20.51 | 53 | 25 | 4.9 |
| 1977 |  | 23 | 20 | 21 | 18.0 | 36.99 | 21.96 | 80 | 46 | 4.2 |
| 1977 | MAR | 08 | 03 | 01 | 36.7 | 36.54 | 28.43 | 63 | 34 | 4.0 |
| 1977 |  | 22 | 20 | 02 | 13.0 | 38.41 | 20.80 | 84 | 14 | 4.2 |
| 1977 |  | 23 | 11 | 55 | 53.4 | 39.66 | 28.55 | 23 | 43 | 4.6 |
| 1977 |  | 26 | 05 | 48 | 44.0 | 37.81 | 23.24 | 37 | 32 | 4.7 |
| 1977 |  | 28 | 10 | 50 | 18.0 | 36.80 | 27.51 | 24 | 66 | 4.9 |
| 1977 | APR | 05 | 17 | 15 | 08.1 | 39.28 | 23.29 | 37 | 18 | 4.7 |
| 1977 |  | 05 | 19 | 50 | 48.2 | 35.03 | 26.32 | 54 | 37 | 4.1 |
| 1977 |  | 13 | 09 | 30 | 55.2 | 37.49 | 19.88 | 33 | 5 | 4.1 |
| 1977 |  | 22 | 23 | 57 | 07.2 | 38.93 | 21.16 | 70 | 43 | 4.2 |
| 1977 | MAY | 05 | 23 | 13 | 07.3 | 34.63 | 24.79 | 16 | 54 | 4.1 |
| 1977 |  | 13 | 16 | 14 | 34.7 | 39.14 | 23.65 | 33 | 22 | 4.7 |
| 1977 |  | 16 | 08 | 16 | 00.6 | 35.38 | 26.52 | 47 | 46 | 4.2 |
| 1977 |  | 21 | 23 | 22 | 49.0 | 36.49 | 27.09 | 104 | 6 | 4.0 |
| 1977 |  | 27 | 22 | 31 | 49.1 | 35.23 | 26.50 | 68 | 100 | 4.8 |
| 1977 | JUN | 02 | 17 | 20 | 19.5 | 35.20 | 27.68 | 33 | 39 | 4.2 |
| 1977 |  | 10 | 23 | 02 | 32.3 | 35.12 | 22.76 | 33 | 52 | 4.1 |
| 1977 |  | 17 | 15 | 41 | 45.4 | 38.44 | 20.35 | 59 | 17 | 4.2 |
| 1977 |  | 21 | 19 | 13 | 26.1 | 35.57 | 29.60 | 46 | 39 | 4.7 |
| 1977 |  | 27 | 22 | 53 | 44.4 | 35.72 | 27.30 | 33 | 34 | 4.1 |
| 1977 |  | 28 | 21 | 103 | 41.4 | 37.74 | 21.18 | 33 | 13 | 4.7 |
| 1977 | JUL | 01 | 12 | 40 | 38.5 | 40.69 | 20.79 | 33 | 17 | 4.2 |
| 1977 |  | 09 | 10 | ) 24 | 26.5 | 35.22 | 23.54 | 73 | 43 | 4.2 |
| 1977 |  | 12 | 13 | 32 | 56.6 | 36.63 | 26.96 | 157 | 37 | 4.2 |
| 1977 |  | 14 | 00 | 39 | 01.9 | 36.21 | 27.71 | 33 | 25 | 4.0 |
| 1977 |  | 18 | 10 | 09 | 15.6 | 41.56 | 20.07 | 42 | 21 | 4.9 |
| 1977 |  | 25 |  |  | 54.9 | 35.09 | 23.78 | 46 | 48 | 4.2 |
| 1977 |  | 27 | 23 | 49 | 31.7 | 34.16 | 26.08 | 33 | 18 | 4.7 |


| DATE |  |  | ORIG.TIME GMT |  |  | $\underset{\mathrm{N}}{\mathrm{LAT}}$ | $\begin{gathered} \text { LON } \\ \text { E } \end{gathered}$ | $\begin{gathered} \text { DEPTH } \\ \text { KM } \end{gathered}$ | OBS | $\begin{array}{r} \text { MAG } \\ \text { MS } \end{array}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1977 | JUL | 30 | 19 | 51 | 37.5 | 36.84 | 21.65 | 51 | 79 | 4.9 |
| 1977 | AUG | 05 | 13 | 19 | 54.8 | 34.27 | 25.80 | 24 | 56 | 4.2 |
| 1977 |  | 18 | 06 | 38 | 36.3 | 39.67 | 25.53 | 4 | 19 | 4.7 |
| 1977 |  | 18 | 09 | 27 | 40.0 | 35.22 | 23.39 | 42 | 125 | 5.3 |
| 1977 |  | 18 | 10 | 04 | 43.4 | 35.07 | 23.31 | 33 | 26 | 4.0 |
| 1977 |  | 25 | 03 | 03 | 09.3 | 35.00 | 28.28 | 10 | 14 | 4.1 |
| 1977 |  | 30 | 14 | 45 | 03.6 | 36.64 | 21.60 | 36 | 53 | 4.6 |
| 1977 |  | 30 | 20 | 51 | 50.2 | 36.38 | 21.56 | 33 | 13 | 4.6 |
| 1977 |  | 30 | 21 | 01 | 58.4 | 36.42 | 21.29 | 33 | 8 | 4.6 |
| 1977 |  | 31 | 08 | 22 | 15.3 | 37.74 | 21.24 | 73 | 43 | 4.7 |
| 1977 | SEP | 10 | 00 | 56 | 09.7 | 34.62 | 26.24 | 64 | 44 | 4.0 |
| 1977 |  | 10 | 06 | 31 | 41.8 | 34.93 | 23.01 | 33 | 74 | 5.1 |
| 1977 |  | 11 | 23 | 19 | 23.7 | 35.05 | 23.03 | 33 | 198 | 6.2 |
| 1977 |  | 12 | 02 | 57 | 55.0 | 34.99 | 23.17 | 36 | 106 | 4.5 |
| 1977 |  | 12 | 10 | 52 | 31.6 | 35.03 | 23.12 | 59 | 32 | 4.0 |
| 1977 |  | 12 | 23 | 10 | 32.3 | 35.65 | 24.17 | 165 | 14 | 4.1 |
| 1977 |  | 13 | 13 | 04 | 09.9 | 34.95 | 23.07 | 33 | 39 | 4.2 |
| 1977 |  | 14 | 18 | 49 | 07.6 | 34.99 | 23.06 | 33 | 80 | 4.5 |
| 1977 |  | 15 | 15 | 53 | 38.9 | 34.91 | 23.01 | 33 | 27 | 4.4 |
| 1977 |  | 18 | 05 | 57 | 19.3 | 34.92 | 23.31 | 33 | 17 | 4.2 |
| 1977 |  | 23 | 02 | 58 | 01.2 | 41.50 | 20.07 | 23 | 95 | 4.7 |
| 1977 |  | 24 | 20 | 43 | 08.9 | 35.06 | 23.21 | 64 | 54 | 4.2 |
| 1977 |  | 25 | 03 | 12 | 23.5 | 34.89 | 23.15 | 64 | 34 | 4.1 |
| 1977 | OCT | 07 | 12 | 42 | 51.3 | 38.76 | 20.63 | 33 | 17 | 4.3 |
| 1977 |  | 08 | 10 | 25 | 30.8 | 35.06 | 23.27 | 60 | 12 | 4.1 |
| 1977 |  | 10 | 08 | 49 | 42.5 | 35.40 | 23.38 | 81 | 17 | 4.3 |
| 1977 |  | 12 | 10 | 14 | 27.8 | 39.38 | 21.70 | 48 | 14 | 4.7 |
| 1977 |  | 12 | 20 | 37 | 34.3 | 34.98 | 23.99 | 33 | 8 | 4.1 |
| 1977 |  | 19 | 21 | 29 | 21.1 | 34.83 | 24.94 | 56 | 9 | 4.2 |
| 1977 |  | 22 | 10 | 02 | 09.1 | 34.95 | 23.16 | 33 | 107 | 5.1 |
| 1977 |  | 24 | 05 | 38 | 18.5 | 34.65 | 26.83 | 33 | 27 | 4.3 |
| 1977 |  | 27 | 22 | 43 | 32.5 | 37.96 | 27.88 | 24 | 90 | 4.7 |
| 1977 | NOV | 06 | 02 | 48 | 44.9 | 42.08 | 24.06 | 23 | 24 | 4.8 |
| 1977 |  | 17 | 06 | 28 | 09.1 | 42.02 | 24.08 | 12 | 22 | 4.7 |
| 1977 |  | 26 | 13 | 19 | 47.2 | 38.49 | 20.28 | 66 | 41 | 4.6 |
| 1977 |  | 28 | 02 | 59 | 10.8 | 36.05 | 27.76 | 85 | 229 | 5.9 |
| 1977 | DEC | 03 | 05 | 39 | 29.5 | 40.25 | 19.91 | 27 | 55 | 5.0 |
| 1977 |  | 09 | 15 | 53 | 36.7 | 38.35 | 27.19 | 19 | 94 | 4.6 |
| 1977 |  | 09 | 20 | 36 | 44.0 | 39.37 | 28.55 | 33 | 8 | 4.2 |
| 1977 |  | 08 | 00 | 40 | 43.6 | 35.21 | 23.38 | 62 | 42 | 4.4 |
| 1977 |  | 15 | 08 | 06 | 10.8 | 34.92 | 23.08 | 47 | 17 | 4.2 |
| 1977 |  | 16 | 07 | 37 | 30.1 | 38.43 | 27.22 | 34 | 72 | 4.8 |
| 1977 |  | 29 | 16 | 52 | 56.7 | 38.44 | 22.30 | 18 | 71 | 4.8 |
| 1978 | JAN | 29 | 10 | 23 | 44.3 | 35.18 | 25.94 | 33 |  | 5.7 |
| 1978 | MAR | 07 | 22 | 33 | 46.1 | 34.66 | 25.50 | 33 |  | 5.7 |
| 1978 | APR | 27 | 08 | 33 | 29.2 | 38.99 | 22.03 | 33 |  | 5.2 |
| 1978 | MAY | 23 | 23 | 34 | 15.3 | 40.78 | 23.41 | 33 |  | 5.3 |
| 1978 | JUN | 19 | 10 | 31 | 05.5 | 40.77 | 23.10 | 10 |  | 5.2 |
| 1978 |  | 19 | 10 | 48 | 10.7 | 40.68 | 23.06 | 10 |  | 4.8 |
| 1978 |  | 20 | 20 | 03 | 29.5 | 40.75 | 23.41 | 15 |  | 6.4 |
| 1978 |  | 20 | 21 | 51 | 03.4 | 40.64 | 23.02 | 10 |  | 4.1 |
| 1978 |  | 21 | 03 | 20 | 25.6 | 40.67 | 23.10 | 10 |  | 4.2 |
| 1978 |  | 21 | 06 | 00 | 05.7 | 40.67 | 23.19 | 10 |  | 4.2 |
| 1978 |  | 24 | 00 | 14 | 28.0 | 41.78 | 20.50 | 10 |  | 4.8 |
| 1978 |  | 26 | 00 | 03 | 48.8 | 42.37 | 20.40 | 10 |  | 4.6 |

## REFERENCES

Agarwal, N.K., Jacoby, W.R., and Berckhemer, H., 1976. Teleseismic P-wave traveltime residuals and deep structure of the Aegean region, Tectonophysics, 31, 33-57.

Ahorner, L., and Rosenhaur, W., 1975. Probability distribution of earthquake accelerations for the sites in Western Germany, Fifth European Conf. on Earth. Eng., Istanbul, Turkey.

Algermissen, S.T., Perkins, D.M., Isherwood, W., Gordon, D., Reagor, G., Howard, C., 1975. Seismic risk evaluation of the Balkan region, Proc. of the seminar on seismic zoning maps, UNESCO - Skopje, 2, 173-240.

Alsan, E., Tezucan, L., and Bath, M., 1975. An earthquake catalogue for Turkey for the interval 1913-1970. Common Report No. 75 of Kandilli Obs., Turkey and Seis. Inst. Uppsala, Sweden.
Alvarez, W., 1973. The application of plate tectonics to the Mediterranean region. In: Implications of Continental Drift to the Earth Sciences, 2. Academie Press, London, 893-908.

Bath, M., 1958. The energies of seismic body waves and surface waves, In: H. Benioff, M. Ewing, B.F. Howe11 Jr. and Press (Editors), Contributions in Geophysics, Pergamon, London, 1: 1-16.
Båth, M., 1969. Handbook on earthquake magnitude determinations. Seism. Inst., Uppsala, 158pp.
Båth, M, , 1973. Introduction to seismology, Birkhauser Verlag Basel, 395pp.
Bath, M., 1975. Seismicity of the Tanzania region, Tectonophysics, 27, 353-379.
Båth, M. and Duda, S.J., 1963. Strain release in relation to focal depth, Geofis. pura e app1. 56, 93-100.

Benjamin, J.R. and Corne11, C.A., 1970. Probability, statistics and decision for Civil Engineers, McGraw-Hill, New York, 684pp.
Bevington, P.R., 1969. Data reduction and error analysis for the physical sciences, McGraw-Hill, New York, 336pp.

Burton, P.W., 1978a. Perceptible earthquakes in the United Kingdom, Geophys. J.R.astr.Soc., 54: 475-479.

Burton, P.W., 1978b. The application of extreme value statistics to seismic hazard assessment in the European area, Proc. Symp, Anal. Seismicity and on Seismic Risk, Liblice, 17-22 October 1977. Academia, Prague 1978, 323-334.

Burton, P.W., 1978c. The IGS file of seismic activity and its use for hazard assessment, Inst. of Geol. Sciences, Seism. Bul. No. 6, 13 pp . Comninakis, P.E., 1975. A contribution to the investigation of the seismicity of the area of Greece. PhD. Thesis, Athens University, 110 pp . Comninakis, P.E., and Papazachos, B.C., 1972. Seismicity of the Eastern Mediterranean and some tectonic features of the Miditerranean Ridge, Geo1. Soc. Am. Bu1., 83, 1093-1102.

Comninakis, P.E., and Papazachos, B.C., 1976. Note on the crustal structure of the eastern Mediterranean, Ann. di Geofis., 24, 59-63.

Constantinescu, L., Ruprechtova, L., and Enescu, D., 1966. MediterraneanAlpine earthquake mechanisms and their seismotectonic implications, Geophys. J.R.astr.Soc., 10, 347-368.

Cornell, C.A., and Kallberg, T.K., 1969. Seismic risk in Southern California. M.I.T., Dept. of Civil Eng., Research Report, R69-31.

Corne11, C.A., and Vanmarcke, E.H., 1969. The major influences on seismic risk, 4 th World Conf. on Earth. Engineering, Chile.

Crampin, S., and Uçer, S.B., 1975. The seismicity of the Marmara Sea region of Turkey. Geophys. J.R.astr.Soc., 40, 269-288.
Curtis. J.W., 1973. A magnitude domain study of the seismicity of Papua, New Guinea, and the Solomon Islands, Bul. Seism. Soc. Am., 63, 787-806. Davis, H.A., 1970. Order Statistics, John Wiley and Sons, Inc., New York, 272pp.

Dewey, J.F., Pitman III, W.C., Ryan, W.B.F., and Bonnin, J., 1973. Plate tectonics and the evolution of the Alpine system. Geol. Soc. Am. Bull. 84, 3137-3180.

Dewey, J.F., and Sengor, A.M.C., 1978. Aegean and surrounding regions: complex multi-plate and continum tectonics in a convergent zone, Geol. Soc. Am. Bull. (in press).

Dick, I.D., 1965. Extreme Value theory and earthquakes, Proc. 3rd World Conf. on Earth. Eng., 1, 45-53.

Donovan, N.C., 1973. A statistical evaluation of strong motion data including the February 9, 1971 San Fernando earthquake, Proc. 5th World Conf. on Earth. Engineering, Rome.

Douglas, D., 1976. Joint epicentre determination, Nature, 215, 47-48. Douglas, A., Young, J.B., and Lilwal1, R.C., 1974. Computer programs for epicentre determination, AWRE Report No. 0 28/74.

Drakopoulos, J., 1976a. On the completeness of macroseismic data a) in the major area of Greece; b) in the Balkan area, Proc. of the seminar on seismic zoning maps, UNESCO - Skopje, 1, 132-155. Drakopoulos, J., 1976b. On the seismic zoning problems in Greece, Proc. of the seminar on seismic zoning maps, UNESCO - Skopje, 1, 300-335.

Duda, S.J., 1965. Secular seismic energy release in the Circum-Pacific Be1t, Tectonophysics, 2, 409-452.

Epstein, B., and Lomnitz, C., 1966. A model for the occurrence of large earthquakes, Nature, 211, 954-956.

Esteva, L., 1974. Geology and probability in the assessment of seismic risk, Proc. 2nd Int. Congr. Int-Assoc. Eng. Geol., Sau Paulo.

Esteva, L., 1976. Seismicity. In: Seismic Risk and Engineering Decisions, Lomnitz, C., and Rosenblueth, E., editors, E1sevier Scient. Publ. Comp., Amsterdam, 425pp.

Esteva, L., and Rosenblueth, E., 1964. Espectos de temblores a distancias moderadas y grandes, Bol. Soc. Mex. In. Sismica, 2, 1-18.

Finetti, I., 1976. Mediterranean Ridge: A young submerged chain associated with the Hellenic Arc, Bollet.Geof. Teor. ed Applic., 19, 31-65.

Flinn, E.A., 1965. Confidence region and error determinations for seismic event location, Revs. of Geophysics, 3, 157-185.

Galanopoulos, A.G., 1953. Katalog der Erdbeben in Griechenland fur die Zeit von 1879 bis 1892, Ann. Géol.d. Pays Hellén, 5, 114-229.

Galanopoulos, A.G., 1960. A catalogue of shocks with $I o \geqslant V I$ or $M \geqslant 5$ for the years 1801-1958, Athens, 119pp.

Galanopoulos, A.G., 1961. A catalogue of shocks with Io $\geqslant$ VII for the years prior to 1800 , Athens, 18 pp .

Galanopoulos, A.G., 1963. On mapping of seismic activity in Greece, Ann. di. Geof., 16, 37-100.

Galanopoulos, A.G., 1967. The seismotectonic regime in Greece, Ann. di. Geof., 20, 109-119.

Galanopoulos, A.G., 1968. On quantitative determination of earthquake risk, Ann. di. Geof., 21, 193-206.

Galanopoulos, A.G., 1971a. Introduction to seismology, Athens, 405pp. Galanopoulos, A.G., 1971b. Minimum and maximum magnitude threshold in the area of Attica; Greece, Ann. di. Geof., 24, 29-54.

Galanopoulos, A.G., 1972a. Annual and maximum possible strain accumulation in the major area of Greece, Ann. Géol.d. pays Hellén, 24, 467-480. Galanopoulos, A.G., 1972b. Plate tectonics in the area of Greece as reflected in the deep focus seismicity, Bull. Geol. Soc. Greece, 9, 266-285. Galanopoulos, A.G., 1973. On the difference of the stress field in the two centres of higher earthquake activity in the area of Greece, Ann. Géol. d. Pays Hellén, 25, 350-372.

Galanopoulos, A.G., 1974. On the tectonic processes along the Hellenic arc, Ann. di. Geof., 27, 429-442.

Galanopoulos, A.G., 1975. A new model accounting for the intermediate earthquakes at the convex side of the Hellenic arc, Ann. Géol. d. Pays He11én, 27, 355-370.

Galanopoulos, A.G., 1977. On the difference in the seismic risk for normal and tall structures at the same site, Nat. Obs. of Athens, 33pp.

Geiger, L., 1910. Herdbestimmung bei Erdbeben aus den Ankunflzeiten (The determination of earthquake centres from arrival times), K. Gese11. Wiss. Goett., 4, 331-349.

Gregeren, S., 1977. P-wave travel time residuals caused by a dipping plate in the Aegean arc in Greece, Tectonophysics, 37, 83-93.

Gringorten, I.I., 1963a. A plotting rule for extreme probability paper, J. Geophys. Research, 68, 813-814.

Gringorten, I.I., 1963b. Envelopes for ordered observations applied to meteorological extremes, J. Geophys. Research, 68, 815-826.

Gumbel, E.J., 1935. Les valeurs extrèmes des distribution statistiques, Ann. Inst. Henri Poincaré, 5, 115-158.

Gumbel, E.J., 1945. Simplified plotting of statistical observations, Trans. Geophys. Union, 26, 69.

Gumbel, E.J., 1954. Statistical theory of Extreme Value and some practical applications, U.S. Dept. Commerce, App. Math. Ser. 3, 51 pp .

Gumbe1, E.J., 1963. Statistical forecast of droughts, Bull. I.A.S.H. V111e Année 1: 5-23.

Gumbel, E.J., 1966. Statistics of extremes, Columbia Univ. Press, New York, 375pp.

Gutenberg, B., 1956. Great earthquakes 1896-1903. Trans. Am. Geophys. Union, 37, 608-614.

Gutenberg, B., and Richter, C.F., 1941. Seismicity of the earth, Geol. Soc. Am. Bull., 34, 1-131.

Gutenberg, B., and Richter, C.F., 1944. Frequency of earthquakes in California, Bull. Seism. Soc. Am., 34, 185-188.

Gutenberg, B., and Richter, C.F., 1954. Seismicity of the earth and associated phenomena, Princeton Univ. Press, Princeton, N.J. Gutenberg, B., and Richter, C.F., 1956. Magnitude and energy of earthquakes, Ann. di. Geof., 9, 1-15.

Hamilton, W.C., 1964. Statistics in physical science, The Ronald Press Co., New York, 230pp.

Herrin, E., Tucker, W., Taggart, D., Gordon, W., and Lobde11, J.L., 1968. Estimation of surface focus $P$ travel times, Bull. Seism. Soc. Am., 58, 1273-1291.

Jeffreys, H., and Bullen, E.K., 1940. Seismological tables, Brit. Assn. Gray-Milne Trust.

Jenkinson, A.F., 1955. The frequency distribution of the annual maximum (or minimum) values of meteorological elements, Q. Jour. Roy. Meteor. Soc., 87, 158-171.

Jongsma, D., 1974. Heat flow in the Aegean Sea, Geophys.J.R.astr.Soc., 37, 337-346.

Kanamori, H., 1978.Quantification of earthquakes, Nature, 271, 411-414. Kárnik, V., 1964. Magnitude-frequency relation and seismic activity in different regions of the European area, Bull. Intern. Inst. Seism. Earthq. Eng., 1, 9-32.

Kárnik, V., 1969. Seismicity of the European area, Part 1,"D. Reidel Publishing Company", Dordrecht-Holland, 364pp.

Kárnik, V., 1971. Seismicity of the European area, Part 2, "D Reidel Publishing Company", Dordrecht-Ho1land, 218pp.

Kárnik, V., Kondorskaya, N.V., Riznichenko, J.V., Savarensky, E.F., Shebalin, N.V., Soloviev, S.L., Vanek, J., and Zatopek, A., 1962, Standardization of the magnitude scale. Stud. Geod. Geoph., 1. Kárnik, V., and Hübnerova, Z., 1968. The probability of occurrence of largest earthquakes in the European area, Pure and App1. Geophys. 70, 61-73.

Karnik, V., and Schenkova, Z., 1977. The third asymptotic distribution in earthquake statistics, Proc. Symp. Anal. Seismicity and on Seismic Risk, Liblice, 17-22 October 1977, Academia, Prague 1978, 335-350.

Katayama, T., 1974. Statistical analysis of peak accelerations of recorded earthquake ground motions, SEISAN-KENKYU, Univ. of Tokyo, 26-1, 18-20.

Kimball, B.F., 1946. Assisgnment of frequencies to a completely ordered set of sample data, Trans. Am. Geophys. Union, 27, 843-846.

Kimball, B.F., 1960. On the choice of plotting positions on probability paper, Amer. Stat. Association Journal,55, 546-560.

Knopoff, i., 1964. The statistics of earthquakes in Southern California, Bull. Seism. Soc. Am., 54, 1871-1873.

Knopoff, L., and Kagan, Y., 1978. Analysis of the theory of extremes as applied to earthquake problems, J. Geophys. Res. 82, 5647-5675.

Krumbein, W.C., and Lieblein, J., 1956. Geological application of extreme value methods to interpretation of cobbles and boulders in gravel deposits, Trans. Amer. Geophys. Union, 37, 313-319.

Lilwall, R.C., 1969. The determination of epicentres and seismic travel times, PhD Thesis, Birmingham University, 110 pp .

Lilwall, R.C., 1976. Seismicity and seismic hazard in Britain, Inst. Geol. Seis. Bull. No. 4 (HMSO)

Lomnitz, C., 1966. Statistical prediction of earthquakes, Rev. Geophys., 4, 377-393.

Lomnitz, C., 1974. Global tectonics and earthquake risk, Elsev. Scient. Pub1. Comp., Amsterdam, 320pp.

Lort, J.M., 1971. The tectonics of the eastern Mediterranean: a geophysical review, Rev. Geophys. Space Phys., 9, 189-216.

McKenzie, D.P., 1970. Plate tectonics of the Mediterranean region, Nature, 226, 239-243.

McKenzie, D., 1972. Active tectonics of the Mediterranean region, Geophys. J.R.astr.Soc., 30, 109-185.

McKenzie, D., 1978. Active tectonics of the Alpine-Himalayan belt: the Aegean Sea and surrounding regions, Geophys. J.R.astr.Soc., 55, 217-254.

McKenzie, D., and Parker, R.L., 1967. The north Pacific: an example of tectonics on a sphere, Nature, $216,1276-1280$.

Makjanić, B., 1972. A contribution fo the statistical analysis of Zagreb earthquakes in the period 1869-1968, Pure and Applied Geophys., 95, 80-88.

Makris, J., 1973. Some geophysical aspects of the evolution of the Hellenides. Bull. Geol. Soc. Greece, 10, 206-213.

Makris, J., 1975. Crustal structure of the Aegean Sea and the Hellenides obtained from geophysical surveys, J. Geophys. Res., 41, 441-443.

Makris, J., 1976a. Geophysical investigations of the Hellenides. Hamb. Geophys. Einzelschr., Heft 27, 1-98.

Makris, J., 1976b. A dynamic model of the Hellenic arc deduced from Geophysical data, Tectonophysics, 36, 339-346.

Makris, J., 1978. The crust and upper mantle of the Aegean region from deep seismic so ndings, Tectonophysics, 46, 269-284.

Marquardt, D.W., 1963. An algorithm for least-squares estimation of nonlinear parameters, J.Soc. Indust. App. Maths., 11, 431-441.

Mercier, J.-L., Carey, E., Phillip, H. and Sorel, D., 1976. La néotectonique plio-quaternaire de 1 'arc égéen externe et de la mer Égée et ses relations avec seismicité, Bull. Soc. geol. Fr. 18, 159.

Merz, H., and Cornel1, C.A., 1973. Seismic risk analysis based on a quadratic magnitude-frequency law, Bull. Seism. Soc. Am., 63, 1999-2006.

Milne, A., and Davenport, A.G., 1969. Distribution of seismic risk in Canada, Bull. Seism. Soc. Am., 59, 729-754.

Morelli, C., Pisani, M., and Gantar, C., 1975. Geophysical studies in the Aegean Sea and in the Mediterranean, Boll. Geof. Teor. ed Applic., 18, 127-167.

Nordquist, J.M., 1945. Theory of large values applied to earthquake magnitudes, Trans. Amer. Geophys. Union, 26, 29-31.

Oliveira, C.S., 1974. Seismic risk analysis, Earth. Engin. Res. Centre, Report No. 74-1, 102pp.

Orpha1, D.L., and Lahoud, J.A., 1974. Prediction of peak ground motion from earthquakes, Bull. Seism. Soc. Am., 64, 1563-1574.

Papazachos, B.C., 1969. Phase velocities of Rayleigh waves in southeastern Europe and eastern Mediterranean Sea, Pure and Appl. Geophys., 75, 47-55.

Papazachos, B.C., 1973. Distribution of seismic foci in the Mediterranean and surrounding area and its tectonic implication, Geophys. JiR.astr. Soc., 33, 419-428.

Papazachos, B.C., 1974. Seismotectonics of the eastern Mediterranean area, Engineer, Seismol. and Earthquake Engineer, Noordhoff Leiden, Ed. J.S. Solnes, 1-31.

Papazachos, B.C., 1976a. Evidence of crustal shortening in the Northern Aegean region, Boll. Geof. Teor. ed Applic., 13, 66-71.

Papazachos, B.C., 1976b. Seismotectonics of the northern Aegean area, Tectonophysics, 33, 199-209.

Papazachos, B.C. 1977. A lithospheric model to interpret focal properties of intermediate and shallow shocks in Central Greece, Pure and App1. Geophys., 115, 655-666.

Papazachos, B.C., and Comninakis, P.E., 1971. Geophysical and tectonic features of the Aegean Arc, J. Geophys. Res., 76, 8517-8533.

Papazachos, B.C., and Comninakis, P.E., 1976. Modes of lithospheric interaction in the Aegean area, Symp. Internat. on the structural history of the Mediterranean basins, Split-Yugoslavia, Oct. 1976, 319-332.

Papazachos, B.C., and Comninakis, P.E., 1978. Deep structure and tectonics of the eastern Mediterranean, Tectonophysics, 46, 285-296.

Papazachos, B.C., and Delibasis, N.D., 1969. Tectonic stress field and seismic faulting in the area of Greece, Tectonophysics, 7, 231-255.

Powell, R.W., 1943. A simple method of estimating flood frequencies, Civ. Eng., 13, 105-107.

Reid, H.F., 1911. The elastic rebound theory of earthquakes, Univ. of California, Dept. of Geology, Bull. 6 (19).

Richter, C.F., 1958. Elementary Seismology, Freeman and Co., San Francisco, 766 pp .

Rikitake, T., 1976. Earthquake prediction, Elsevier Scient. Publ. Comp., Amsterdam, 357pp.

Ritsema, A.R., 1974. Earthquake mechanisms of the Balkan region. UNDP/UNESCO reports on Survey of the Seismicity of the Balkan Region, project REM/70/172.

Rothé, J.R., 1969. The seismicity of the earth 1953-1965, UNESCO, Earth Sciences 1, 336pp.
Rothé, J.P., 1972. Programme de détermination d'epicentres et method algorithmique de recherche d'un epicentre approche et de 1 'heure origine correspondante d'un seisme. Project PNUD-UNESCO, B.C.I.S. Strasbourg.

Sacuiu, I., and Zorilescu, D. 1970. Statistical analysis of seismic data on earthquakes in the area of the Vrancea focus, Bull. Seism. Soc. Am., 60, 1089-1099.

Schenkova, Z., and Kárnik, V., 1970. The probability of occurrence of largest earthquake in the European area, Part II, Pure and App1. Geophys., 33, 181-278.

Schenkova, Z., and Kárnik, V., 1976. Application of the Largest Values Theory to Balkan Earthquakes, Proc. of the Seminar on Seism. Zoning Maps, Skopje 27 Oct.- 4 Nov. 1975, 1, 193.

Schenkova, Z., and Kárnik, V., 1978. The third asymptotic distribution of largest magnitudes in the Balkam earthquake provinces, Pure and App1. Geophys., 116, 1314-1325.

Schenkova, Z., and Schenk, V., 1975. Return periods of earthquakes and trends of seismic activity, Pure and App1. Geophys., 43, 683-693.

Schmidt, J., 1879. Studien über Erdbeben, Leipzig.
Schnabel, P.B., and Seed, H.B., 1973. Accelerations in rock for earthquakes in the Western United States, Bull. Seism.Soc. Am., 63, 501-516.

Shah, H.C., and Movassate, M., i975. Seismic risk analysis - California state water project, Fifth European Conf. on Earth. Eng., Istanbu1, Turkey.

Shah, H.C., and Vag1iente, V.N., 1972. Forecasting the risk inherent in earthquake resistant design, Proc. Intern. Conf. on Microzonation, 2, Seattle.

Shakal, A.F. and Willis, D.E., 1972. Estimated earthquake probability in the North Circum-Pacific area, Bu11. Seism. Soc. Am., 1397-1410.

Shebalin, N.V., Kárnik, V., and Hadzievski, D., (Editors), 1974. Balkan region - Catalogue of earthquakes, UNESCO Project Office, Skopje.

Sieberg, A., 1932. Untersuchungen über Erdbeben und Bruchschollenbau in östlichen Mittelmeergebiet, Denkschr. Med. Naturw. Ges, 18, Jena.

Stepp, J.C., 1971. An investigation of Earthquake Risk in the Puget Sound Area by use of the type I Distribution of Largest Extremes, Ph.D Thesis, Pennsylvania State University, 131 pp .

Trifunac, M.D., 1976. Preliminary analysis of the peaks of strong earthquake ground motion - dependence of peaks on earthquake magnitude, epicentral distance, and recording site conditions, Bull. Seism. Soc. Am., 66, 189-219.
Ücer, B.S., Ayhan, E., and Alsan, E., 1975. Preliminary statistical results for the preparation of a seismic zoning map of Anatolia, Kandilli Obs. Istanbul, Turkey, 20pp.

Vag1iente, V.N., 1973. Forecasting the risk inherent in earthquake resistant design, Tech. Report No. 174, Dept. of Civil Engineering, Stanford University.

Vere-Jones, D., 1970. Stochastic models for earthquake occurrence, J.R. Stat. Soc., 32, 1-62.

Vogt, P.R., and Higgs, R.H., 1969. An aero-magnetic survey of the eastern Mediterranean Sea and its interpretation, Earth. Planet Sci. Letters, 5, 439-448.

Willmore, P.L., and Burton, P.W., The UK approach to hazard assessment, Conference proceedings (ed. Ritsema) "On nuclear reactors and seismic hazard", Luxemburg, Sept. 1975, 35-37.

Yegulalp, T.M., 1974. Forecasting for largest earthquakes, Management Science, 21, 418-421.

Yegulalp, T.M., and Kuo, J.T., 1966. Application of extremal statistics to the maximum magnitude earthquakes (Abstract), Trans. Am. Geophys. Union, 47, 163.

Yegulalp, T.M., and Kuo, 1974. Statistical prediction of the occurrence of maximum magnitude earthquakes, Bull. Seism. Soc. Am., 64, 393-414.


[^0]:    *Throughout this study the $\log$ is used for $\log _{10}$ whereas the expression ln is used for $\log _{e}$.

[^1]:    Fig 3-9 Cumulative strain energy release as a function of time for Greece for the period 1950-1972 (after Galanopoulos, 1972a).

