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1 **Paleomagnetism, magnetic anisotropy and U-Pb baddeleyite geochronology of the early**
2 **Neoproterozoic Blekinge-Dalarna dolerite dykes, Sweden**

3

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18

19 **Highlights**

- 20 • Paleomagnetism of Blekinge-Dalarna dolerite dykes demonstrates the reliability of a
21 951-935 Ma key pole for Baltica.
- 22 • The anomalous direction from 947 Ma Nornäs dyke is attributed to a partial
23 remagnetization.
- 24 • Baltica and Laurentia drifted from high to low latitude between 970-960 Ma and 950-
25 935 Ma, and returned back to high latitude by 920-870 Ma.
- 26 • The Blekinge-Dalarna dolerite dykes are unlikely a giant circumferential swarm
27 generated by a mantle plume.

28 **Abstract**

29 Paleogeographic proximity of Baltica and Laurentia in the supercontinent Rodinia has been
30 widely accepted. However, robust paleomagnetic poles are still scarce, hampering quantitative
31 tests of proposed relative positions of the two cratons. A recent paleomagnetic study of the
32 early Neoproterozoic Blekinge-Dalarna dolerite (BDD) dykes in Sweden provided a 946-935 Ma
33 key pole for Baltica, but earlier studies on other BDD dykes discerned large variances in
34 paleomagnetic directions that appeared to indicate more complicated motion of Baltica, or
35 alternatively, unusual geodynamo behavior in early Neoproterozoic time. We present combined
36 paleomagnetic, rock magnetic, magnetic fabric and geochronological studies on BDD dykes in
37 the Dalarna region, southern Sweden. Positive baked-contact and paleosecular variation tests
38 support the reliability of the 951-935 Ma key pole (Paleolatitude = -2.6°N , Paleolongitude =
39 239.6°E , $A_{95} = 5.8^{\circ}$, $N = 12$ dykes); and the ancient magnetic field was likely a stable geocentric
40 axial dipole at that time, based on a positive reversal test. Detailed analysis of the 947 Ma
41 Nornäs dyke, one of the dykes previously showing anomalous directions, suggests a partial
42 viscous remagnetization. Therefore, the observed large variances in nearly coeval BDD dykes
43 are suspected to result from present-day overprints that were not adequately removed in
44 earlier studies. In addition, we obtained a 971 Ma virtual geomagnetic pole (Paleolatitude = -
45 27.0°N , Paleolongitude = 230.4°E , $A_{95} = 14.9^{\circ}$, $N = 4$ dykes) for Baltica. Comparing similar-aged
46 poles from Laurentia, we suggest that Baltica and Laurentia drifted together from high to low
47 latitude between 970-960 Ma and 950-935 Ma, and returned back to high latitude by 920-870
48 Ma. In this scenario, the apparent polar wander paths of Baltica and Laurentia may be more
49 complicated than the previously proposed, solitary Sveconorwegian and Grenville loops. The
50 new U-Pb baddeleyite ages do not support BDD dykes as a giant circumferential swarm
51 generated by a mantle plume, and the prolonged timespan of dyke intrusion is likely associated
52 with the plate boundary forces as causing gravitational extension at the waning stage of the
53 Sveconorwegian orogeny.

54

55 **Keywords**

56 Blekinge-Dalarna dolerite (BDD) dykes; Sveconorwegian loop; Baltica; paleomagnetism;
57 magnetic anisotropy; U-Pb baddeleyite geochronology

58

59 **1. Introduction**

60 Although the existence of the Proterozoic supercontinent Rodinia has long been suggested,
61 its configuration is still highly debated, and new studies continue to paint different pictures
62 regarding the shape of Rodinia (e.g., [Slagstad et al., 2013](#); [Wen et al., 2017](#); [Wen et al., 2018](#)).
63 Nonetheless, the juxtaposition of Laurentia and Baltica in Rodinia is widely adopted in various
64 reconstruction models (e.g., [Dalziel, 1997](#); [Pisarevsky et al., 2003](#); [Cawood and Pisarevsky,](#)
65 [2006](#); [Li et al., 2008](#); [Evans, 2009](#)), in order to account for geological similarities between the
66 two cratons. As the only quantitative method to constrain the paleolatitude and the orientation
67 of pre-Pangea continents, paleomagnetism plays an important role in testing current
68 reconstruction models of Laurentia and Baltica. The apparent polar wander (APW) paths of
69 Laurentia and Baltica, including the Grenville and Sveconorwegian loops, respectively, show a
70 broad agreement during post-1.0-Ga Rodinia assembly ([Li et al., 2008](#)), supporting the
71 paleogeographic proximity of the two cratons. However, most paleomagnetic poles on the
72 Grenville and Sveconorwegian loops derive from high-grade metamorphic rocks, making the
73 determination of the age of remanence acquisition difficult ([Brown and McEnroe, 2012](#); [Brown](#)
74 [and McEnroe, 2015](#)). Compilation of paleomagnetic poles with ages between 1.3 Ga and 0.9 Ga
75 ([Veikkolainen et al., 2017](#)) shows that even though numerous poles have been generated for
76 Baltica and Laurentia, high-quality poles with quality criteria $Q_{V_{90}}$ value larger than 4 are scarce
77 ([Fig. 1a](#)), and most poles have no field tests to constrain the age of remanence ([Fig. 1b](#)). As a
78 result, large uncertainties remain regarding the shapes and younging directions of the Grenville
79 and Sveconorwegian loops ([Hyodo and Dunlop, 1993](#); [Elming et al., 1993](#); [Weil et al., 2006](#);
80 [Elming et al., 2014](#)).

81 Mafic intrusions are promising targets for yielding high-quality poles since they are usually
82 enriched in magnetite grains. Also, baddeleyite grains in mafic intrusions can be directly dated
83 by U-Pb method with high precision ([Söderlund et al., 2005](#)). A recent study of the Blekinge-
84 Dalarna dolerite (BDD) dykes in southern Sweden generated a 946-935 Ma low-latitude pole

85 that has been proposed as a key pole for Baltica (Elming et al., 2014). However, earlier studies
86 of BDD dykes show large variances in paleomagnetic directions (Patchett and Bylund, 1977;
87 Bylund, 1985; Bylund and Elming, 1992) that might indicate unusual cratonic motions or
88 complicated geomagnetic behavior. Here we present detailed paleomagnetic, rock magnetic
89 and magnetic fabric studies on a number of BDD dykes in the Dalarna region, southern Sweden,
90 in order to better understand the paleogeography of Baltica in early Neoproterozoic time. We
91 also present new U-Pb baddeleyite ages that shed light on the tectonic origin of BDD dykes.

92

93 **2. Geologic background, previous work and sampling**

94 Following the Svecofennian orogeny (2.0-1.75 Ga), Baltica grew outward as a result of
95 accretionary tectonics manifested by the 1.81-1.76 Ga Transscandinavian Igneous Belt (TIB;
96 Bogdanova et al., 2015). Afterwards, there was a protracted interval of mafic magmatism
97 peaked at 1.6 Ga, 1.57 Ga, 1.47-1.44 Ga, 1.27-1.26 Ga, 1.22 Ga, and 0.98-0.95 Ga, respectively
98 (Söderlund et al., 2005; Brander and Söderlund, 2007). The 1.47-1.44 Ga magmatism, referred
99 to as the Danopolonian event (Bogdanova et al., 2001), is largely coeval with dynamic high-
100 grade metamorphism in southwestern Sweden, and is suggested to be related to convergent
101 active margin processes called the Hallandian event (Christoffel et al., 1994; Söderlund et al.,
102 2002; Möller et al., 2007; Brander and Söderlund, 2007). After the Hallandian event, the 1.1-0.9
103 Ga Sveconorwegian orogeny (e.g., Bingen et al., 2008) extensively reworked the basement
104 rocks west of the Protogine Zone (PZ) and the Sveconorwegian Frontal Deformation Zone
105 (SFDZ) in southwest Scandinavia (Fig. 2; Wahlgren et al., 1994). Later, Caledonian allochthons
106 were thrust onto the northwest margin of Baltica at 0.6-0.4 Ga (Fig. 2; Gaál and Gorbatshev,
107 1987; Bingen and Solli, 2009).

108 Partly coincident with the Sveconorwegian orogeny, the early Neoproterozoic BDD dykes
109 intruded the TIB and Svecofennian rocks east of the PZ and SFDZ, over an extent of 750 km (Fig.
110 2). One prominent feature of BDD dykes is their arcuate shape, trending NE-SW in the Blekinge
111 region, and deflected ~60° to NW-SE in the Dalarna region (Fig. 2). The ages of BDD dykes are
112 well established by U-Pb baddeleyite geochronology and ⁴⁰Ar/³⁹Ar whole-rock dating, ranging
113 from 978 Ma to 939 Ma (Söderlund et al., 2005; Elming et al., 2014). The origin of BDD dykes is

114 debated. Different models have been proposed, including fracturing due to late
115 Sveconorwegian uplift (Patchett and Bylund, 1977), gravitational collapse at the final stage of
116 the Sveconorwegian orogeny (Pisarevsky and Bylund, 2006), and the giant circumferential
117 system of a mantle plume (Buchan and Ernst, 2016). Petrological studies show that BDD dykes
118 are fine- to medium-grained with slight alteration, and the major minerals consist of
119 plagioclase, olivine, clinopyroxene, orthopyroxene, biotite, ilmenite, and titanomagnetite
120 (Johansson and Johansson, 1990; Solyom et al., 1992).

121 Previous paleomagnetic work focused on the southern and central part of BDD dykes
122 (e.g., Patchett and Bylund, 1977; Bylund, 1985; Bylund and Elming, 1992), while the northern
123 part is relatively less studied. Recently, Elming et al. (2014) proposed a low-latitude pole for
124 Baltica at 946-935 Ma, combining the results of several BDD dykes in the Norrköping and Falun
125 areas with the 935 Ma Göteborg-Slussen dykes in southwestern Sweden (Fig. 2; Pisarevsky and
126 Bylund, 2006). The reliability of this low-latitude pole is supported by a positive baked-contact
127 test and the appearance of antipodal directions. However, large variances still remain in the
128 paleomagnetic results of BDD dykes, complicating the application of the low-latitude pole to
129 paleogeographic reconstruction of Baltica. For example, the 947 Ma Nornäs dyke yields a high-
130 latitude pole (Piper and Smith, 1980; Bylund, 1985), which apparently contradicts with the
131 result of Elming et al. (2014). It is also difficult to explain their difference by plate tectonics
132 because they are very similar in age and would imply extremely rapid continental motions.
133 Alternative interpretations have been suggested, such as late-stage selective remagnetization,
134 true polar wander, a non-dipole geomagnetic field or the non-averaged paleosecular variation
135 (Pesonen and Klein, 2013). However, none of these has been fully examined. Notably, the result
136 of the Nornäs dyke was obtained more than three decades ago when modern laboratory
137 treatment and data analysis of paleomagnetism had not been fully developed. Hence, it is
138 necessary to restudy the Nornäs dyke with more detailed and sophisticated techniques.

139 Since more than 90% of the bedrock in the Dalarna region is covered by glacial deposits,
140 aeromagnetic data (Ripa et al., 2012) and data from geological mapping of the Swedish
141 Geological Survey (Ripa, personal communication) were used to help delineate dykes in the
142 field. For detailed mapping of the outcrops, a portable magnetic susceptibility meter was used

143 to determine the extension of the dykes and contacts to host rocks. A number of NW-SE
144 trending dykes in the Dalarna region were sampled using a portable gasoline-powered rock drill
145 (Fig. 3). Host rocks (ca. 1.46 Ga Öje basalt and ca. 1.46 Ga Dala sandstone) were collected in two
146 sites for baked-contact tests, the former where a clear intrusive contact was observed, and the
147 latter where the concealed contact could be triangulated to within about a meter of the baked
148 host-rock samples. Core samples were oriented with a Brunton compass, and sun-compass
149 readings were also taken in order to correct any small-scale magnetic anomaly in the outcrop.
150 Block samples were collected from the central parts of dykes for U-Pb baddeleyite
151 geochronology.

152

153 3. Methods

154 3.1 U-Pb baddeleyite geochronology

155 All samples were crushed, and baddeleyite grains were separated using the Wilfley table
156 technique following Söderlund and Johansson (2002) at Lund University in Sweden. The
157 extracted baddeleyite grains are dark to moderately brown. Grains from all samples are fresh
158 without any trace of alteration. About 1-5 grains per fraction were combined and transferred to
159 clean Teflon capsules. Grains were washed in numerous steps using 3 M HNO₃, including one
160 step on a hotplate (~30 minutes). One drop of the ²⁰⁵Pb-²³³⁻²³⁶U tracer solution and 10 drops of
161 HF-HNO₃ (10:1) were added to each capsule. Baddeleyite grains were completely dissolved
162 after 3 days in an oven under high pressure at a temperature of ~190°C. The samples were
163 evaporated on a hotplate and then dissolved in 10 drops of 6.2 M HCl. One drop of 0.25 M
164 H₃PO₄ was added to each capsule before it dried down. U and Pb were loaded on an outgassed
165 Re filament together with a small amount of silica gel.

166 Thermal Ionization Mass Spectrometry (TIMS) was performed at the Laboratory of Isotope
167 Geology at the Swedish Museum of Natural History in Stockholm, Sweden, using a Thermo
168 Finnigan Triton TIMS system. An ETP SEM detector equipped with a RPQ filter was used to
169 measure the Pb and U isotope intensities in dynamic (peak-switching) mode. Pb-isotopes were
170 measured at a filament temperature in the 1200-1230°C range, while U-isotopes were
171 measured in dynamic mode on the SEM with filament temperatures exceeding 1300°C. Data

172 reduction was performed using the Excel add-in “Isoplot” of Ludwig (2003); decay constants for
173 ^{238}U and ^{235}U follow those of Jaffey et al. (1971). All errors in age and isotopic ratios are quoted
174 at 2σ . Initial Pb isotope compositions were corrected according to the global common Pb
175 evolution model of Stacey and Kramers (1975). U-Pb data are presented in Table 1 and the
176 fractions are plotted in the concordia diagrams in Fig. 4.

177

178 3.2 Magnetic measurements

179 Magnetic measurements were conducted in the Paleomagnetic Laboratory and
180 Archaeomagnetism Laboratory at Yale University, USA. In order to understand the pattern of
181 dyke propagation and to discern the possible alteration or deformation of dykes (Hrouda, 1982;
182 Rochette et al., 1992), anisotropy of magnetic susceptibility (AMS) was measured for each dyke
183 using an AGICO Kappabridge KLY-4S susceptibility meter. AMS data were analyzed in Anisoft42
184 software. To characterize magnetic mineralogy, representative samples were subjected to
185 thermomagnetic susceptibility analysis. Temperature ranges from 25°C to 700°C and is
186 controlled by a CS3 high temperature furnace apparatus. Bulk magnetic susceptibility was
187 measured during heating and cooling in an argon gas environment in order to subdue magnetic
188 phase transition. Magnetic grain size was inferred by the Day plot (Day et al., 1977) constructed
189 by magnetic parameters, which are determined by hysteresis loop measurement using a
190 MicroMag 2900 alternating gradient magnetometer (AGM). After the measurement of natural
191 remanent magnetization (NRM), samples were cooled to liquid nitrogen temperature (~77 K) in
192 a magnetic shielded container to demagnetize the remanence carried by multidomain grains
193 (Muxworthy and McClelland, 2000). Thermal demagnetization was conducted in an ASC
194 Scientific TD-48 thermal demagnetizer with stepwise heating up to 580°C in 15-20 steps in a
195 nitrogen gas environment. Sister samples from each dyke were demagnetized using alternating
196 field (AF) in a Molspin tumbler demagnetizer. After each thermal or AF demagnetization step,
197 remanence was measured by a 2G Enterprises cryogenic DC-SQUID magnetometer with
198 automatic sample-changing device (Kirschvink et al., 2008). Paleomagnetic vectors were
199 calculated using principal component analysis (Kirschvink, 1980) and the great circle method
200 (McFadden and McElhinny, 1988), and were plotted using vector-endpoint diagrams

201 (Zijderveld, 1967) in PaleoMag X software (Jones, 2002). Paleogeographic reconstruction was
202 carried out using GPlates software (Boyden et al., 2011).

203

204 4. Results

205 4.1 U-Pb baddeleyite geochronology

206 We report U-Pb baddeleyite TIMS ages of three NW-SE trending BDD dykes from the
207 Dalarna region. Data are summarized in Table 1, and concordia diagrams are shown in Fig. 4.
208 Three baddeleyite fractions of dyke G16S12, which is intrusive into the Öje basalt (G16S11), are
209 concordant at 951 ± 5 Ma (2σ , mean square weighted deviates [MSWD] = 1.4). This age is
210 calculated as the weighted mean of $^{207}\text{Pb}/^{206}\text{Pb}$ dates for these fractions. Three fractions of
211 dyke G16S22 were analyzed, of which two are concordant within uncertainty whereas one
212 analysis is slightly discordant. The weighted mean of $^{207}\text{Pb}/^{206}\text{Pb}$ dates is 971 ± 7 Ma (2σ , MSWD
213 = 0.61). Three fractions of dyke G16S37, which is probably the southern extension of the
214 Nornäs dyke, cluster at and just below the concordia curve. The weighted mean is 947 ± 4 Ma
215 (2σ , MSWD = 2.8). These age estimates are interpreted as dating the crystallization of the rocks.

216 Before this study, a total of 11 ages of BDD dykes/sills were published, with nine U-Pb
217 baddeleyite ages (Söderlund et al., 2004a; Söderlund et al., 2005) and two $^{40}\text{Ar}/^{39}\text{Ar}$ whole-rock
218 ages (Elming et al., 2014). We provide an additional three U-Pb baddeleyite ages. Collectively,
219 they demonstrate a ~40 million-year range of BDD intrusions with probably three magmatic
220 pulses, extending from 978 Ma to 939 Ma (Fig. 5).

221

222 4.2 Rock magnetism

223 AMS data show that the degree of anisotropy (P_j) of BDD dykes is typically below 6%, which
224 is a common value for primary fabric of igneous rocks (Hrouda, 1982). Low P_j values indicate
225 that these rocks have not experienced significant deformation, hence, have a reasonable
226 chance of retaining primary magnetization. The majority of dykes exhibit oblate fabrics with
227 K_{max} and K_{int} axes dispersed in the NW-SE oriented, vertical or sub-vertical plane (Fig. 6e). Some
228 dykes show prolate fabrics with K_{max} axis pointing to the NW-SE direction (Fig. 6a). The
229 orientation of magnetic anisotropy ellipsoid genuinely reflects the trends (NW-SE) of these

230 dykes (Knight and Walker, 1988), which is also supported by field observations and the
231 aeromagnetic anomaly map (Fig. 3; Ripa et al., 2012). AMS of Öje basalt is expected to have a
232 horizontal oblate fabric; the K_{\min} axis shows a transition from steep to shallow directions (Fig.
233 6c). Samples with shallow K_{\min} axes are close to the contact and the chilled margin of dyke
234 G16S12. It is suspected that some secondary magnetic minerals might grow along the basalt
235 contact zone due to the migration of reducing fluids, and cause anomalously shallow K_{\min} axes.
236 But the remanences of the secondary magnetic minerals are adequately removed by thermal
237 demagnetization and have no effect on primary paleomagnetic signals (see discussion below).
238 AMS of the Dala sandstone (G16S24) shows a typical depositional fabric (oblate and horizontal),
239 with K_{\min} axis perpendicular to the bedding plane and K_{\max} and K_{int} axes distributed parallel to
240 the bedding plane (Fig. 6h).

241 Thermomagnetic susceptibility analysis of BDD dykes shows that heating and cooling
242 curves are generally reversible (Fig. 7). During heating, the magnetic susceptibility decreases
243 substantially between 580°C and 600°C, which provides clear evidence for the presence of
244 magnetite. Minor drops in magnetic susceptibility are also noticed between 600°C and 700°C,
245 indicating small amounts of hematite or maghemite. Exceptions are Öje basalt (G16S11) and
246 dyke G16S12, which yields a distinct susceptibility hump and a large decline between 300°C and
247 400°C on the heating curve (Fig. 7c-d). This temperature range coincides with the Curie
248 temperature of magnetic sulfides (pyrrhotite or greigite). During heating, the magnetic sulfides
249 were broken down to form new magnetite, as suggested by the sharply increased susceptibility
250 at 580°C on the cooling curve. The magnetic sulfides are likely the reason for shallow K_{\min} axis
251 observed in the AMS data of Öje basalt (G16S11). Dala sandstone (G16S24) has a low magnetic
252 susceptibility and shows a gradual decrease in susceptibility between 600°C and 700°C (Fig. 7g),
253 suggesting the major magnetic phase is hematite or titanohematite.

254 The coercivity of remanence of BDD dykes determined by hysteresis loop measurement is
255 typically less than 30 mT, which is a normal value for magnetite. Only dyke G16S12 gives slightly
256 higher coercivity (~80 mT), showing the contribution from magnetic sulfides (Fig. 8b). The
257 typical grain size of magnetic phases in BDD dykes falls in pseudo-single domain (PSD) region on
258 the Day plot. According to Dunlop (2002)'s theoretical estimates, the magnetic grains are

259 mixtures of single domain (SD) and multidomain (MD) minerals with varying SD content ranging
260 from 20% to 60% (Fig. 8c).

261

262 4.3 Paleomagnetism

263 Paleomagnetic results show that during heating, most samples exhibit a significant decline
264 of remanence at $\sim 580^{\circ}\text{C}$, close to the unblocking temperature for pure magnetite (Fig. 9).
265 Samples that were subjected to liquid nitrogen bath show a large decrease of remanence,
266 indicating that the viscous remanence carried by MD grains has been effectively removed (Fig.
267 9). Some samples show a gradual loss of remanence at lower temperatures, which could be
268 either due to the demagnetization of larger size grains or magnetic sulfides. Most samples yield
269 a clear decay-to-origin component between 500°C and 580°C , which is defined as the
270 characteristic remanent magnetization (ChRM). Only samples from sites G16S17 and G16S18
271 (two sites were collected from the same dyke) were analyzed by the combination of principal
272 component analysis and great circle method, owing to the overlapping unblocking
273 temperatures of different-sized grains.

274 Usable paleomagnetic directions of BDD dykes fall into two statistically different groups
275 (Fig. 10a), one with shallower inclinations (Group A) and another with steeper inclinations
276 (Group B). Importantly, the dating results also show that the two groups are different in age.
277 Group A is about 951 ± 5 Ma, ~ 20 million years younger than Group B (971 ± 7 Ma). Both groups
278 show two polarities. Following Precambrian paleomagnetism database PALEOMAGIA's
279 nomenclature (Veikkolainen et al., 2017), we assign reverse polarity to sites with southeasterly
280 declinations (Groups Ar, Br) and normal polarity to sites with northwesterly declinations
281 (Groups An, Bn; Table 2). Paleomagnetic directions of Group A are resemble those of the Y1, Y2
282 and Falun dykes obtained in Elming et al. (2014). A positive baked-contact test supports the
283 primary origin of the NW-up direction (Group An; Elming et al., 2014), In addition, we have two
284 positive baked-contact tests for the SE-down direction (Group Ar). The baked areas of the Öje
285 basalt (G16S11) and the Dala sandstone (G16S24), which were intruded by dykes G16S12 and
286 G16S23 respectively, show similar remanence to BDD dykes, but are very different from their
287 primary (unbaked) directions (Piper and Smith, 1980; Fig. 10a). We performed a reversal test on

288 Group A, together with Y1, Y2 and Falun dykes from [Elming et al. \(2014\)](#) and the Tuve, Small,
289 Hjuvik and Slussen dykes from [Pisarevsky and Bylund \(2006\)](#); the test is demonstrated to be
290 positive with classification C ($\gamma/\gamma_c = 8.7^\circ/13.9^\circ$; [McFadden and McElhinny, 1990](#)). On the basis of
291 the similar ages and paleomagnetic results of [Pisarevsky and Bylund \(2006\)](#), [Elming et al.](#)
292 [\(2014\)](#), and our data, we obtained a new mean 951-935 Ma paleomagnetic pole: Paleolatitude
293 = -2.6°N , Paleolongitude = 239.6°E , $A_{95} = 5.8^\circ$ ($N = 12$ dykes; [Table 2](#)). There is no baked-contact
294 test for the four dykes in Group B, but their demagnetization patterns, such as the square-
295 shouldered thermal decay curve with the unblocking temperature closed to 580°C , suggest that
296 their ChRMs are probably primary ([Fig. 9](#)). The number of dykes in Group B is insufficient for the
297 reversal test, but quasi-antipodal directions have been observed, also supporting the notion
298 that dykes in Group B carry primary remanence. We calculated a 971 Ma virtual geomagnetic
299 pole (VGP) from Group B, which is Paleolatitude = -27.0°N , Paleolongitude = 230.4°E , $A_{95} = 14.9^\circ$
300 ($N = 4$ dykes; [Table 2](#)).

301 Paleomagnetic results of the Nornäs dyke (G16S05 and G16S06; 946.8 ± 1.2 Ma) show that
302 the natural remanent magnetization (NRM) direction is very close to the steep ChRM yielded
303 from previous work ([Piper and Smith, 1980](#); [Bylund, 1985](#)). The steep ChRM is also very close to
304 present-day field (PDF) direction ([Fig. 10b](#)). However, as temperature or AF intensity increases,
305 the inclination gradually decreases ([Fig. 11](#)). It is noteworthy that the direction of remanences
306 of some samples are able to migrate towards the upper hemisphere at $\sim 570^\circ\text{C}$ or ~ 30 mT. The
307 remanence becomes unstable when approaching the unblocking temperature or coercivity of
308 magnetite, so it is hard to isolate a decay-to-origin component. However, the directions move
309 towards the NW-up BDD-reference direction (Group An; [Fig. 10](#)). The pattern of vector-
310 endpoint diagrams clearly shows a partial remagnetization. Based on the aeromagnetic
311 anomaly map and geochronology, it is likely that sites G16S16 and G16S37 (947 ± 4 Ma) are
312 from a southeastward extension of the Nornäs dyke, and they also give identical paleomagnetic
313 results and similar ages. Therefore, we interpret that the Nornäs dyke consists of some PSD
314 grains that carry a PDF overprint, difficult to be adequately removed due to the strong overlap
315 of unblocking temperature or coercivity with SD magnetite grains. Since a clear decay-to-origin
316 component cannot be isolated from sites G16S05, G16S06, G16S16 and G16S37, we tried to use

317 great circle method for paleomagnetic analyses. Combining all great circles from these four
318 sites of the Nornäs dyke, we obtained a mean direction of Declination = 307.9°, Inclination = -
319 35.3°, $\alpha_{95} = 2.9^\circ$, which we named it as “Nornäs new”, the direction of which is very close to the
320 NW-up BDD-reference direction (Group An; Fig. 10; Table 2). The Nornäs new direction is likely
321 the primary remanence of the Nornäs dyke. However, without any sample bearing a clear
322 decay-to-origin component, we prefer to exclude this Nornäs new mean direction from the
323 paleomagnetic statistics of this study. The steep ChRM direction of previous work was
324 calculated from the minimum scatter in magnetic directions after demagnetization in AF
325 intensities of 10-20 mT (Piper and Smith, 1980; Bylund, 1985; Table 2), which is too low to
326 remove the PDF overprint. Another dyke, G16S14, yields similar demagnetization pattern as the
327 Nornäs dyke and is also likely affected by partial remagnetization in the PDF direction.

328

329 **5. Discussion**

330 5.1 Origin of BDD dykes – A giant circumferential swarm?

331 Arcuate-shaped swarms have been increasingly reported in different geological settings
332 with various intrusion ages (Denyszyn et al., 2009; Mäkitie et al., 2014; Buchan and Ernst,
333 2018). However, the physical mechanism explaining the unusual geometry is still under debate.
334 Interestingly, the coronae on Venus are also characterized by similar ring-shaped surface
335 expressions and are thought to be associated with tectono-magmatic processes (Squyres et al.,
336 1992). Are the arcuate-shaped swarms on Earth and Venusian coronae intrinsically related in
337 terms of their origin?

338 Venusian coronae have two components: a radiating system and a circumferential system,
339 both of which are presumably underlain by dykes (Ernst et al., 2003). In order to explain their
340 distinctive tectonic and topographic features, Stofan and Head (1990) suggested a mantle
341 plume origin for coronae. The radiating system is argued to be related to the upwelling of hot
342 magma, causing surface uplift and dyke intrusion, while the circumferential system is due to
343 gravitational collapse as the mantle upwelling ceases. If this mechanism is true, it is expected
344 that the radiating system would predate the circumferential system.

345 [Ernst and Buchan \(1998\)](#) first proposed that the arcuate-shaped swarms on the Earth could
346 be analogous to Venusian coronae. They defined any arcuate-shaped swarm as a giant
347 circumferential swarm if it has a primary circular or elliptical geometry with an arc of $> 45^\circ$ and
348 a diameter > 60 km ([Buchan and Ernst, 2016](#)). Based on these criteria, BDD dykes would be
349 classified as a giant circumferential swarm, accounting for their 750-km long, $\sim 60^\circ$ arcuate
350 geometry. The primary curved geometry is supported by tectonic coherence of southern
351 Sweden since Neoproterozoic time, and also the demonstrably high-quality BDD paleomagnetic
352 data, which discern no structural rotation following emplacement. Therefore, the critical test of
353 this hypothesis hinges on another two aspects. Is there a corresponding radiating system? And
354 if so, is the radiating system older than BDD dykes, as required by [Stofan and Head \(1990\)](#)'s
355 model?

356 The available ages of BDD dykes/sills suggest prolonged intrusions ([Fig. 5](#)). Given the
357 distribution of these ages, there seems to be 3 possible pulses of BDD intrusions, first pulse
358 from 980 Ma to 965 Ma; second from 955 Ma to 945 Ma; third from 942 Ma to 935 Ma ([Fig. 5](#)),
359 although definitive conclusion still needs more geochronological studies. Any radiating-system
360 candidate should be older or at least very close to the first pulse of BDD intrusions. The
361 Göteborg-Slussen dykes in southwestern Sweden were proposed as the radiating system linked
362 to the purported mantle plume, because their trends are sub-orthogonal to those of the BDD
363 dykes ([Fig. 2; Buchan and Ernst, 2016](#)). However, the Tuve dyke, which belongs to the
364 Göteborg-Slussen dyke suite, is dated to be 935 ± 3 Ma by the U-Pb baddeleyite method
365 ([Hellström et al., 2004](#)), which approaches equivalency to the youngest members of BDD dykes
366 but is tens of millions of years younger than the majority of BDD dykes. Another candidate for
367 the radiating system might be the Hunnedalen dykes in western Norway ([Fig. 2](#)), trending NE-
368 SW, sub-orthogonal to BDD dykes. But the Hunnedalen dykes are ~ 100 million years younger
369 than BDD dykes ([Walderhaug et al., 1999](#)), excluding their possibility of being the radiating
370 component. In general, geochronological data do not support [Buchan and Ernst \(2016\)](#)'s mantle
371 plume model. The prolonged intrusion interval of BDD dykes is more likely connected with the
372 plate boundary forces causing gravitational extension in the Baltic foreland during protracted

373 waning stages of the Sveconorwegian orogeny. Variable orientation of the regional stress field
374 might be the cause for the primary arcuate geometry (Wahlgren et al., 1994).

375

376 5.2 Implications for paleogeography of Baltica

377 Large variances have been observed in the paleomagnetic results of BDD dykes (Fig. 10b;
378 Table 2), which have been interpreted differently in previous studies (e.g., Pesonen and Klein,
379 2013). Here, we summarize available paleomagnetic and geochronological data of BDD dykes in
380 the literature (Fig. 5; Table 2), in an attempt to examine each proposed model, and then discuss
381 their implications. First, if we assume that all BDD results are reliable without any
382 remagnetization/contamination, the variance in BDD dyke remanences would be substantial
383 (Table 2), for instance, the very steep inclinations of 947 Ma Nornäs dyke and 954 Ma
384 Karlshamn dyke, and the very shallow inclinations of 946 Ma Fäjö dyke and 948 Ma Bräkne-
385 Hoby dyke (Fig. 10b). This variance, occurring within a fairly short time interval (< 10 Ma), if
386 interpreted as plate motion, would require unrealistically fast drift rates of Baltica,
387 contradicting geologically younger plate tectonic speeds. Even true polar wander, which might
388 occur as fast as $\sim 6^\circ$ per Myr (Rose and Buffett, 2017), is insufficient to reconcile the variances in
389 BDD remanences. Another possibility might be a non-GAD field, but even then, the entire BDD
390 dataset would be difficult to explain unless departures from GAD were extreme. For example,
391 Pesonen et al. (2012) show that with 11% octupole field (relative to a dominant GAD field), the
392 inclination shallowing effect will approach a maximum $\sim 10^\circ$ at mid-paleolatitudes. However,
393 the inclination differences among BDD remanences are mostly larger than 10° , and some even
394 reach $70\text{-}80^\circ$ (Table 2). Only if the transient ancient magnetic field was totally dominated by
395 octupole component (Tauxe, 2005) or by an ephemeral equatorial dipole field (Abrajevitch and
396 Van der Voo, 2010) can the large inclination differences be explained.

397 After the experience gained from careful demagnetization procedures in our study, we are
398 suspicious that the large variances of remanence in the entire BDD dataset likely result from
399 selective remagnetization. The 947 Ma Nornäs dyke, after being subjected to detailed
400 demagnetization, shows that the remanence gradually moves away from the previously
401 determined ChRM and towards the NW-up BDD-reference direction (Figs. 10 and 11). In fact,

402 based on the great circle analyses, the primary remanence of the Nornäs dyke should be the
403 same as that of Group An, and the low-temperature or low-AF remanence component of
404 Nornäs dyke is NNW and steep down, very close to the PDF direction. Hence, it is very likely
405 that the Nornäs dyke was affected by partial remagnetization and component mixing due to
406 overlapping coercivity and/or unblocking temperature spectra. That concept impels us to doubt
407 the robustness of other anomalous directions, especially the 954 Ma Karlshamn dyke, for which
408 the published ChRM direction is very close to the PDF direction (Fig. 10b). Also, as the
409 remanences of 946 Ma Lösen-Fäjö and 948 Ma Bräkne-Hoby dykes are half-way between the
410 PDF direction and the NW-up BDD-reference direction (Group An), they could possibly carry
411 substantial PDF overprints as well. Besides, the previously published directions of the Nornäs,
412 Lösen-Fäjö, Karlshamn and Bräkne-Hoby dykes were calculated using the minimum scatter of
413 data at 10 mT, 20 mT, 30 mT or 40 mT (Patchett and Bylund, 1977; Piper and Smith, 1980;
414 Bylund, 1985; Table 2). AF demagnetization below 30 mT is generally too low to be pertinent to
415 Precambrian paleomagnetic remanence preservation. If two components have strongly
416 overlapping demagnetization spectra, higher AF intensities do not necessarily guarantee a
417 successful removal of a partial overprint. With more sophisticated laboratory equipment and
418 more accurate analytical methods (principal component analysis, great circle method etc.) now
419 available, we suggest a re-study of these dykes before using their directions for geophysical
420 interpretations.

421 Because we suspect the reliability of some results of BDD dykes in previous studies, either
422 due to the inadequate demagnetization (e.g., low AF field) or outdated analytical methods (e.g.,
423 least scatter method for averaging paleomagnetic directions), for subsequent discussion we
424 only focus on the results that we can be assured to lack an overprint/contamination. This would
425 yield 12 dykes in Group A (6 dykes in Group An and 6 dykes in Group Ar; Table 2) and 4 dykes in
426 Group B (1 dyke in Group Bn and 3 dykes in Group Br; Table 2). We admit that by this treatment
427 the paleomagnetic dataset of BDD dykes is reduced, but we can test for averaging out the
428 paleosecular variation (PSV) by calculating the angular dispersion (S) for each of the two poles
429 we obtained from Groups A and B, respectively, following the approximation equation:

430
$$S = \frac{81^\circ}{\sqrt{k}}$$

431 where k is the best estimate of precision parameter (Butler, 1992).

432 The calculated S values are plotted against the 1.5-0.5 Ga model G curve fitted by
433 Veikkolainen and Pesonen (2014). It is noted that the 951-935 Ma pole obtained from Group A
434 matches the model G curve very well (Fig. 12; Table 3), which means that even though we used
435 a smaller dataset, the number of dykes seems sufficient to average out the PSV. In contrast, the
436 971 Ma pole obtained from Group B falls below the model G curve (Fig. 12; Table 3). Since there
437 are only 4 dykes in Group B, this pole is considered merely as a VGP, and more dykes in this
438 group are needed to provide a paleomagnetic pole in future studies. Further geochronology on
439 Group B dykes would also be useful to assess whether they are restricted to a narrow age
440 range.

441 The early Neoproterozoic apparent polar wander (APW) path of Baltica, known as the
442 Sveconorwegian loop, is under debate in terms of its shape and form. Previously, the
443 Sveconorwegian loop (as shown in the South Pacific) was generally assumed to exhibit
444 counterclockwise motion (Fig. 13a; Elming et al., 1993). However, some have suggested that
445 the southernmost part of the Sveconorwegian loop represents a post-900 Ma delayed
446 remanence acquisition during slow exhumation of the deep-seated igneous rocks (Walderhaug
447 et al., 1999; Brown and McEnroe, 2004). Pisarevsky and Bylund (2006) speculated that the
448 delayed acquisition is probably caused by low-temperature chemical alteration. Elming et al.
449 (2014) proposed a clockwise motion of the south-Pacific polarity representation of the
450 Sveconorwegian loop, incorporating the 945 Ma equatorial pole and assuming that the
451 southernmost part of the loop is post-900 Ma (Fig. 13a). Similarly, the early Neoproterozoic
452 APW path of Laurentia (also viewed in the south-Pacific polarity representation) is also
453 interpreted as either clockwise (Hyodo and Dunlop, 1993) or counterclockwise (Weil et al.,
454 2006). Regardless of uncertainties in the APW paths of two cratons, geological evidence
455 supports a proximity of Baltica and Laurentia during the Neoproterozoic time (Cawood and
456 Pisarevsky, 2006).

457 Our results from the BDD dykes support the equatorial pole for Baltica at 951-935 Ma,
458 similar to the result of Elming et al. (2014). In addition, we obtained a 971 Ma VGP for Baltica,
459 suggesting a high-paleolatitude position. Brown and McEnroe (2012) studied paleomagnetism

460 of the igneous and metamorphic rocks in the Adirondack Highlands of Laurentia. Performing
461 careful rock-magnetic and petrological studies, they generated several poles with modeled
462 cooling ages of 990-960 Ma. We also performed the PSV test of the 970 Ma and 960 Ma poles
463 from [Brown and McEnroe \(2012\)](#), which suggests that the PSV has been averaged out from
464 these two poles ([Fig. 12](#); [Table 3](#)). Cratonic reconstruction using these poles and our new 971
465 Ma VGP permits a close position between Baltica and Laurentia at 970-960 Ma. The
466 reconstructed positions of two cratons allow the Sveconorwegian and Grenville orogenies be a
467 continuous belt, which has been suggested in many other studies (e.g., [Li et al., 2008](#); although
468 see [Gower et al., 2008](#), for cautionary details). The Sveconorwegian and Grenville loops seem to
469 have some oscillatory components ([Fig. 13](#)), which could be attributed to plate motions or true
470 polar wander ([Evans, 2009](#)). In this scenario, the paleogeographic evolution of the two cratons
471 is characterized by high- to low-latitude drift between 970-960 Ma and 950-935 Ma, and a
472 return from low to high latitude by 920-870 Ma ([Fig. 14](#)). Notably, all three pole groups include
473 at least one result that appears to average paleosecular variation adequately ([Fig. 12](#)). The
474 implied drifting speeds of ~100-150 km/Ma, which is fast but is a reasonable rate for either
475 plate tectonics or true polar wander. Deconvolving those two processes will require more
476 detailed paleomagnetic work from Baltica, Laurentia, and other cratons with suitably complete
477 early Neoproterozoic geological records.

478

479 **6. Conclusions**

480 We present a detailed paleomagnetic, rock magnetic and anisotropy of magnetic
481 susceptibility, and geochronological study of the Blekinge-Dalarna dolerite dykes, which leads
482 to following conclusions:

483 (1) Positive baked-contact, reversal and PSV tests support the reliability of the equatorial
484 paleomagnetic pole for Baltica at 951-935 Ma (Paleolatitude = -2.6°N , Paleolongitude = 239.6°E ,
485 $A_{95} = 5.8^{\circ}$, $N = 12$ dykes), which can be used as a key pole to constrain the paleogeography of
486 Baltica.

487 (2) The anomalous paleomagnetic direction obtained from the 947 Ma Nornäs dyke is
488 probably due to a PDF overprint that was not adequately removed by low alternating-field

489 demagnetization levels in previous studies, instead of originating from true polar wander, or
490 abnormal geomagnetic field behaviors. PDF component contamination is suspected in other
491 anomalously-directed BDD dykes.

492 (3) A well-dated 971 Ma VGP (Paleolatitude = -27.0°N , Paleolongitude = 230.4°E , $A_{95} =$
493 14.9°) from four BDD dykes, in concert with same-age poles from Laurentia, suggests a high-
494 latitude position for Baltica in proto-Rodinia. Paleogeographic reconstruction demonstrates
495 that Baltica and Laurentia drifted together towards low latitude between 970-960 Ma and 950-
496 935 Ma, and moved back to high latitude by 920-870 Ma. In this scenario, the apparent polar
497 wander path would be more complicated than either the Sveconorwegian and Grenville loops
498 considered in isolation.

499 (4) Based on published ages of BDD dykes and adjacent dykes sub-orthogonal to it, it seems
500 that the requirement of a single mantle plume model is not satisfied. The BDD dykes more likely
501 result from plate boundary forces associated with the Sveconorwegian orogeny. The arcuate
502 geometry could be associated with a spatially varying regional stress distribution.

503

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512

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731 **Figure captions**

732 Fig.1 Compilation of paleomagnetic poles from Baltica and Laurentia within the 1.3-0.9 Ga
733 interval (Veikkolainen et al., 2017). (a) Sum of quality criteria $Q_{(6)}$ (Van der Voo, 1990), excluding
734 the seventh criterion as in the Precambrian paleomagnetism database PALEOMAGIA. (b)
735 Individual quality criteria (Van der Voo, 1990). N = number of paleomagnetic poles.

736

737 Fig. 2 Geologic map of southern Sweden. GÄZ = Göta Älv Zone, MZ = Mylonite Zone, PZ =
738 Protogine Zone, SFDZ = Sveconorwegian Frontal Deformation Zone. Inset map shows the major
739 tectonic divisions of southern Sweden (modified from Söderlund et al., 2004b). GS = Göteborg-
740 Slussen dykes, Hu = Hunnedalen dykes, RIC = Rogaland Igneous Complex. Small boxes show the
741 locations of Y1, Y2 (Y), and Falun (F) dykes in Elming et al. (2014); and Lösen-Fäjö (L), Bräkne-
742 Hoby (B), Karlshamn (K), and Härsjön (H) dykes in Patchett and Bylund (1977); and Nornäs dyke
743 (N) in this study.

744

745 Fig. 3 Geologic map of the western Dalarna region (modified from Lundmark and Lamminen,
746 2016). Proterozoic dyke locations are based on Ripa et al. (2012). Sites yield interpretable
747 paleomagnetic results (white) and sites give paleomagnetically unstable directions/partial
748 remagnetization (black) are differentiated by colors.

749

750 Fig. 4 U-Pb concordia diagrams of dated BDD dykes.

751

752 Fig. 5 Summary of ages from 980-930 Ma BDD-related intrusions in southern Sweden. Age
753 references are listed in Table 2. Shaded areas show age ranges of possible pulses of BDD
754 intrusions.

755

756 Fig. 6 Representative stereonet projections of the anisotropy of magnetic susceptibility (AMS)
757 data. Squares, triangles and circles show the principal axes of AMS ellipsoids. Grey arrows
758 indicate the trends of dykes inferred from aeromagnetic anomalies (Ripa et al., 2012). Red
759 arrow represents the evolution of AMS ellipsoid of Öje basalt.

760

761 Fig. 7 Results of thermomagnetic susceptibility analysis. Heating and cooling curves are
762 represented by red and blue colors, respectively.

763

764 Fig. 8 Hysteresis ratios of BDD dykes displayed in the Day plot (Day et al., 1977). The dashed line
765 shows the SD/MD theoretical mixing curve (Dunlop, 2002). Inset plots are representative
766 hysteresis loops after the correction of paramagnetic slope.

767

768 Fig. 9 Representative thermal demagnetization results of BDD dykes. Vector-endpoint diagrams
769 are shown (Zijderveld, 1967), with equal-area stereonet plots and remanence intensity (J/J_0)
770 plots. ChRMs are plotted with blue and red arrows representing declinations and inclinations,
771 respectively.

772

773 Fig. 10 (a) Equal-area stereonet projection summarizing the ChRMs of BDD dykes with
774 corresponding 95% confidence cones. Unbaked paleomagnetic direction of Öje basalt and Dala
775 sandstone is suggested by the purple diamond. Present-day field direction of sampling area is
776 indicated by the black square. Closed and open circles show the downwards and upwards
777 paleomagnetic inclinations. PDF = present-day field. (b) Equal-area stereonet projection of
778 anomalous BDD dyke directions. Red stars indicate the BDD reference directions.

779

780 Fig. 11 Typical thermal and alternating-field demagnetization behaviors of the Nornäs dyke.
781 Vector-endpoint diagrams (Zijderveld, 1967), equal-area stereonet plots and remanence
782 intensity (J/J_0) plots are shown. Red solid/dashed lines indicate the trending of remanent
783 directions in lower/upper hemisphere.

784

785 Fig. 12 Paleosecular variation (PSV) test. The thick black line is the 1.5-0.5 Ga model G curve
786 from Veikkolainen and Pesonen (2014). The gray area is the corresponding error limits (2σ).
787 Numbers are paleomagnetic poles used in the paleogeographic reconstruction, which are listed
788 in Table 3.

789

790 Fig. 13 (a) Sveconorwegian loop for Baltica. Counterclockwise motion is from [Elming et al.](#)
791 [\(1993\)](#) and clockwise motion is from [Elming et al. \(2014\)](#). (b) Grenville loop for Laurentia.
792 Counterclockwise motion is from [Weil et al. \(1998\)](#) and clockwise motion is from [Hyodo and](#)
793 [Dunlop \(1993\)](#). (c) Coeval 1000-850 Ma paleomagnetic poles from Baltica and Laurentia in
794 present North America reference frame. White arrows indicate the younging directions of
795 apparent polar wander paths. Blue dash lines indicate new loop proposed by this study. Red
796 poles are from Baltica and green poles are from Laurentia. Poles' numbers are listed in [Table 3](#).

797

798 Fig. 14 Paleogeographic reconstructions of Baltica and Laurentia. (a) 970-960 Ma (Euler pole of
799 Laurentia to absolute reference: 22.4°N, 100.6°E, 127.6°; Euler pole of Baltica to Laurentia:
800 75.8°N, 95.7°E, -59.2°); (b) 950-935 Ma (Euler pole of Laurentia to absolute reference: -37.9°N, -
801 65.5°E, -99.8°; Euler pole of Baltica to Laurentia: 75.8°N, 95.7°E, -59.2°); (c) 920-870 Ma (Euler
802 pole of Laurentia to absolute reference: -15.6°N, -93.5°E, -125.9°; Euler pole of Baltica to
803 Laurentia: 75.8°N, 95.7°E, -59.2°). Paleomagnetic poles used for reconstruction are listed and
804 numbered in [Table 3](#). Dark and light gray poles are from Baltica and Laurentia, respectively.

805

806 **Table captions**

807 Table 1 Results of U-Pb baddeleyite geochronology.

808 Table 2 Summary of paleomagnetic results of BDD dykes.

809 Table 3 Paleomagnetic poles constrain the early Neoproterozoic paleogeographic

810 reconstruction of Baltica and Laurentia.

Figure 1

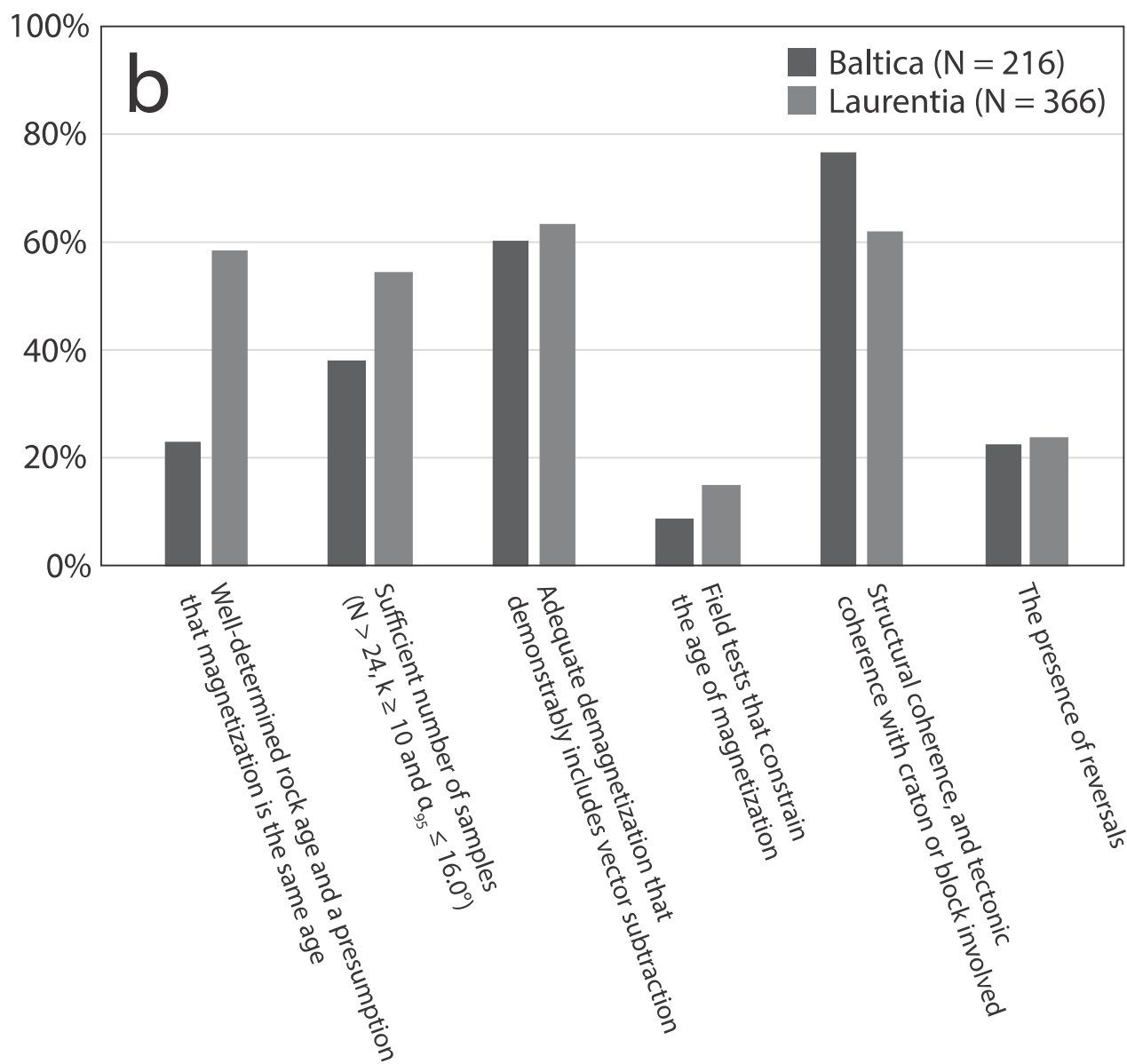
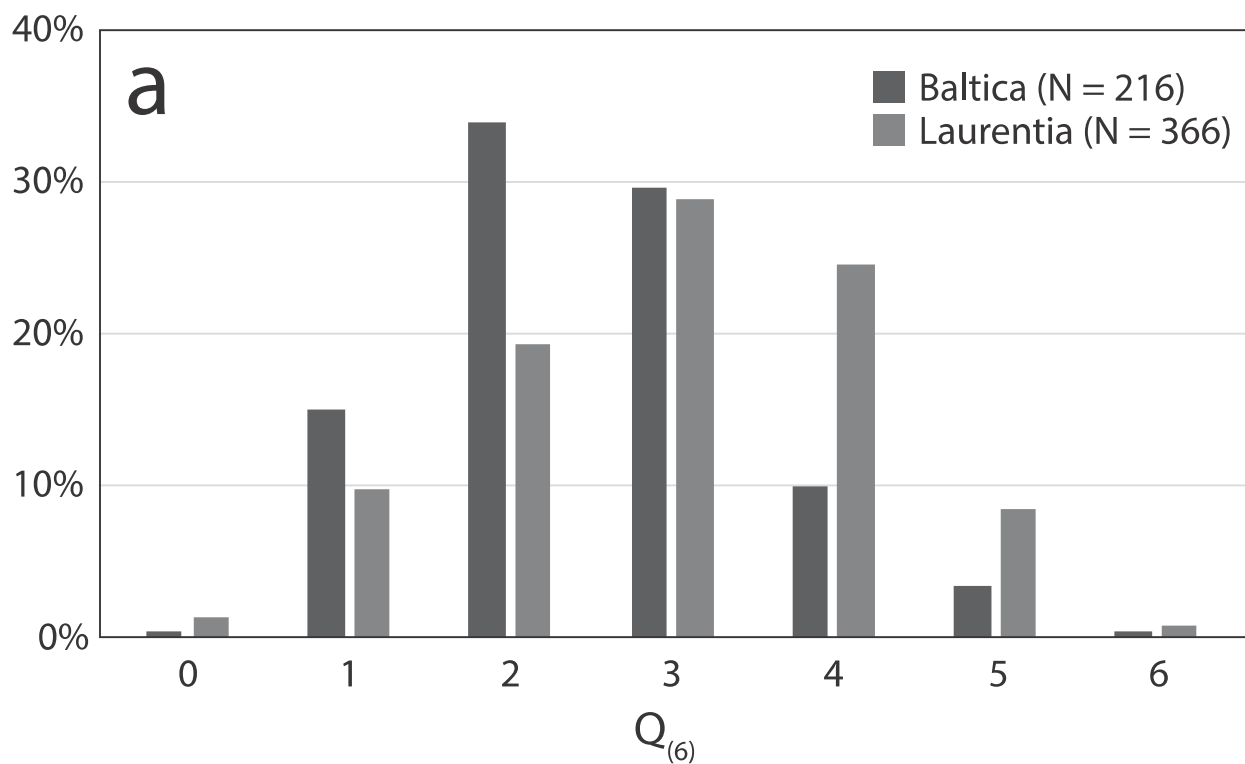


Figure 2

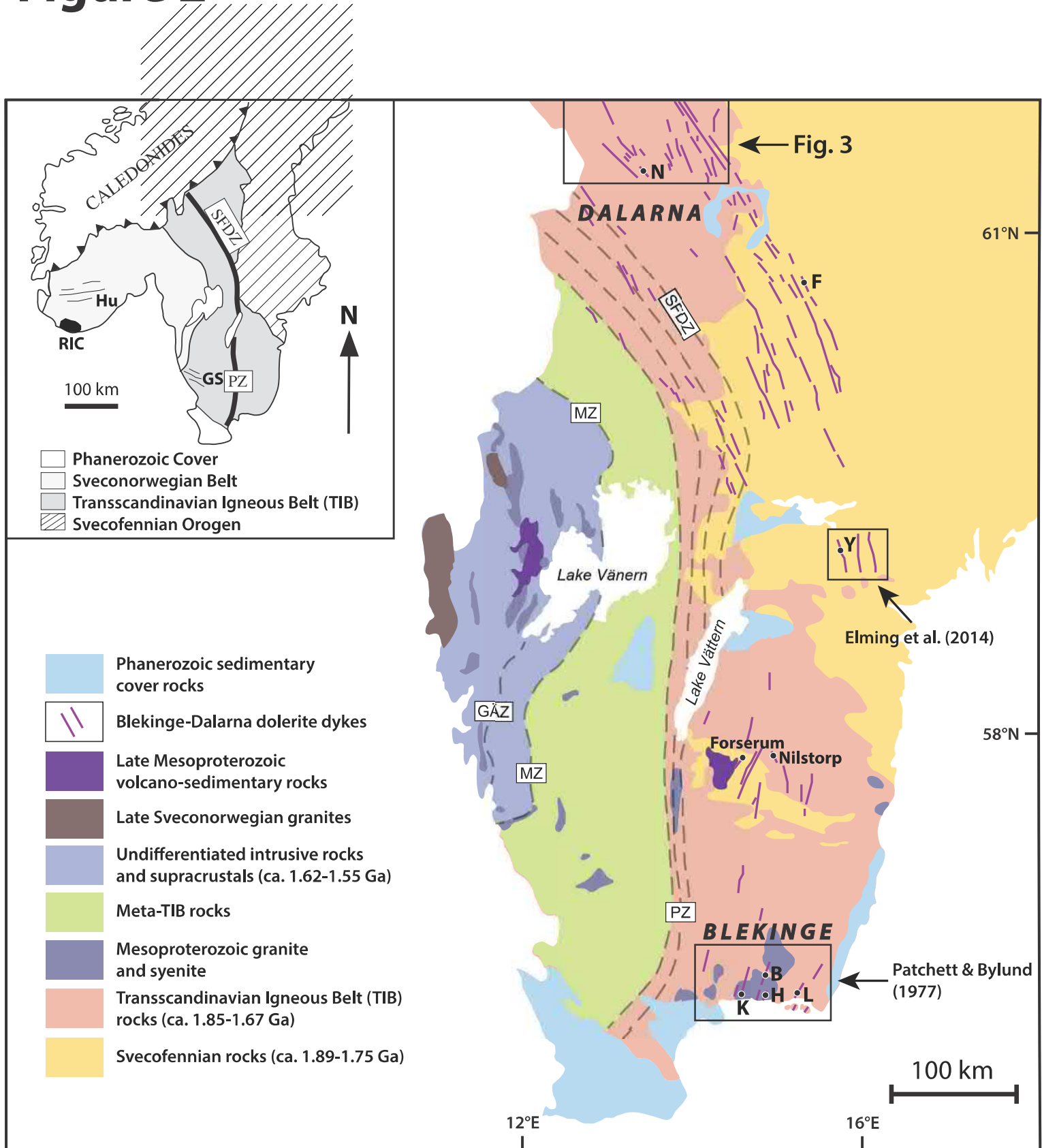


Figure 3

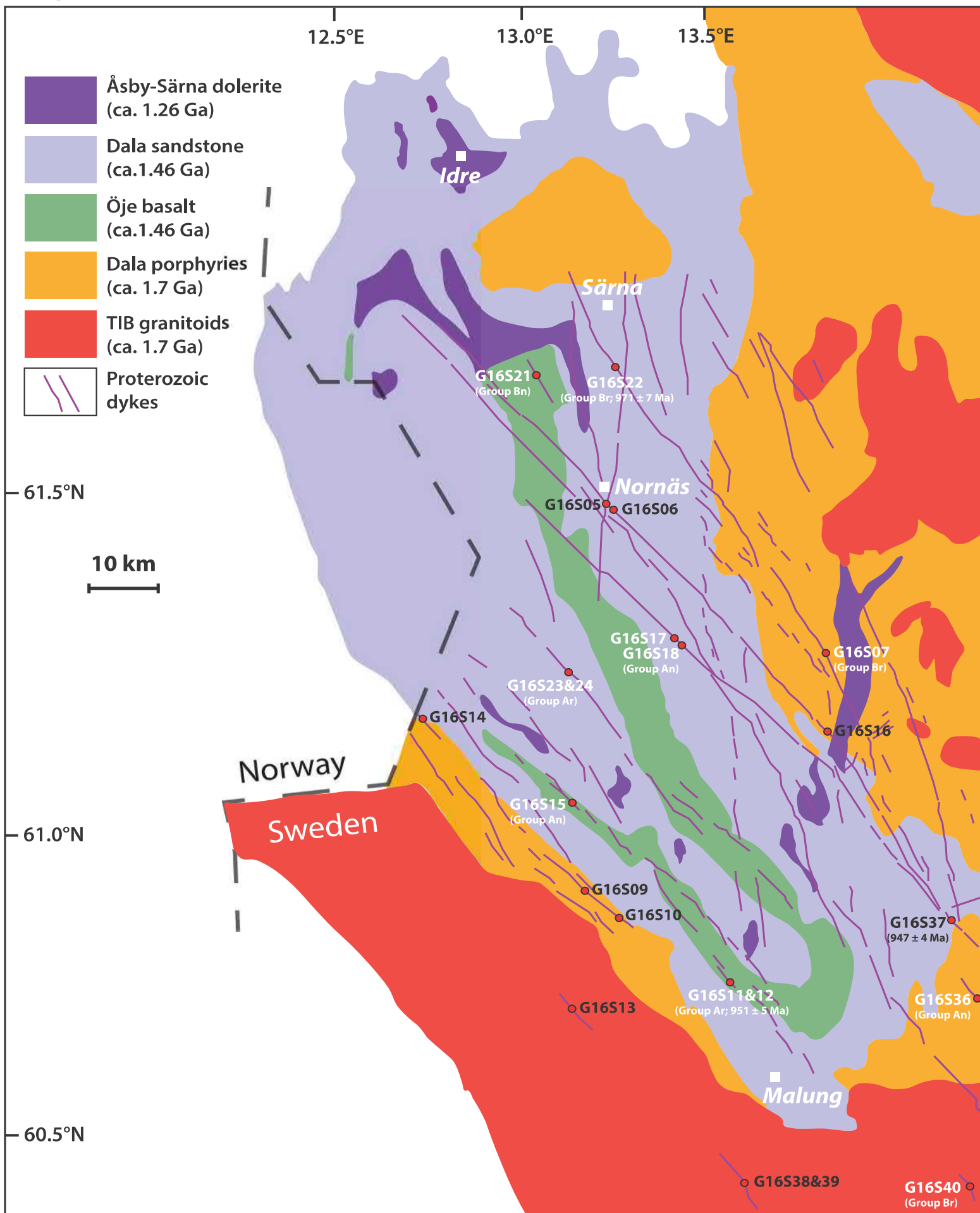


Figure 4

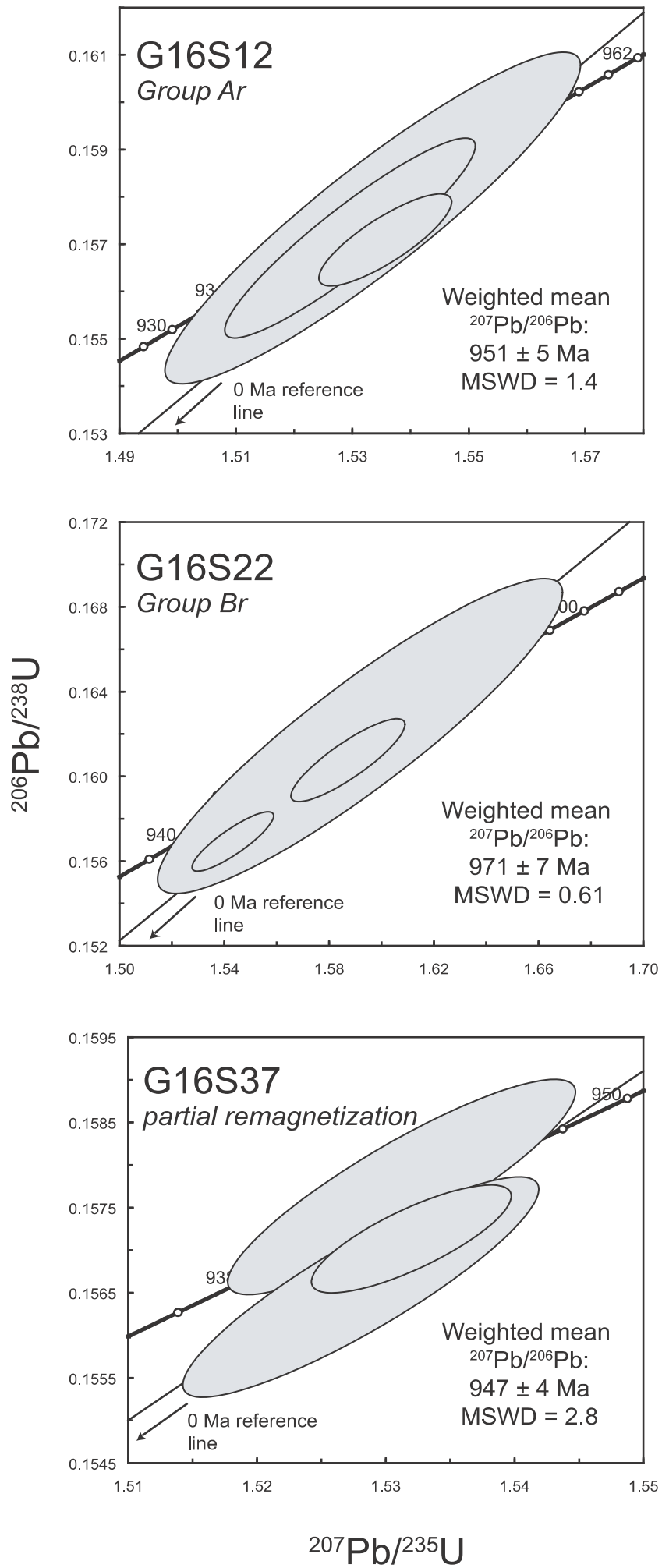


Figure 5

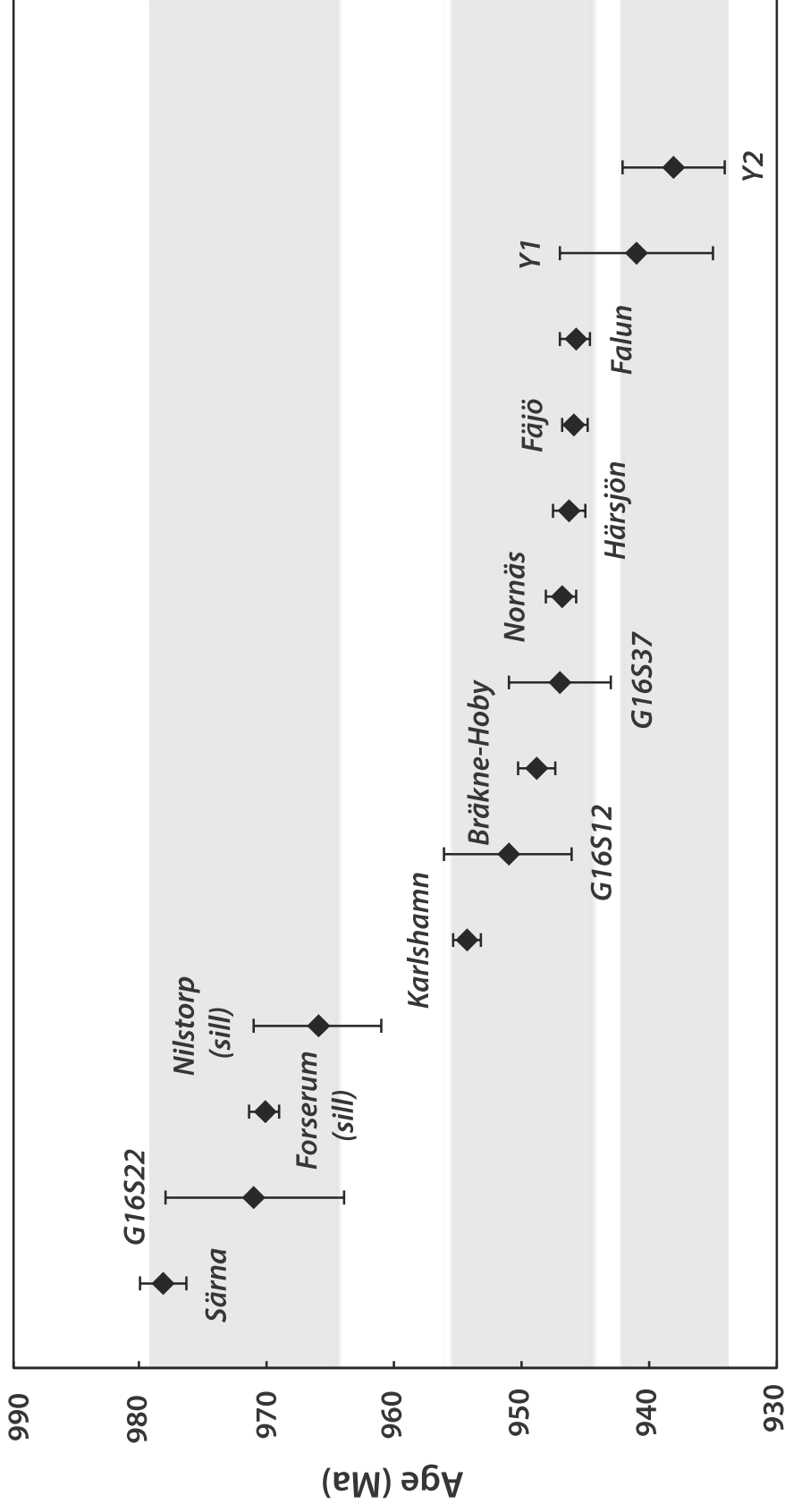


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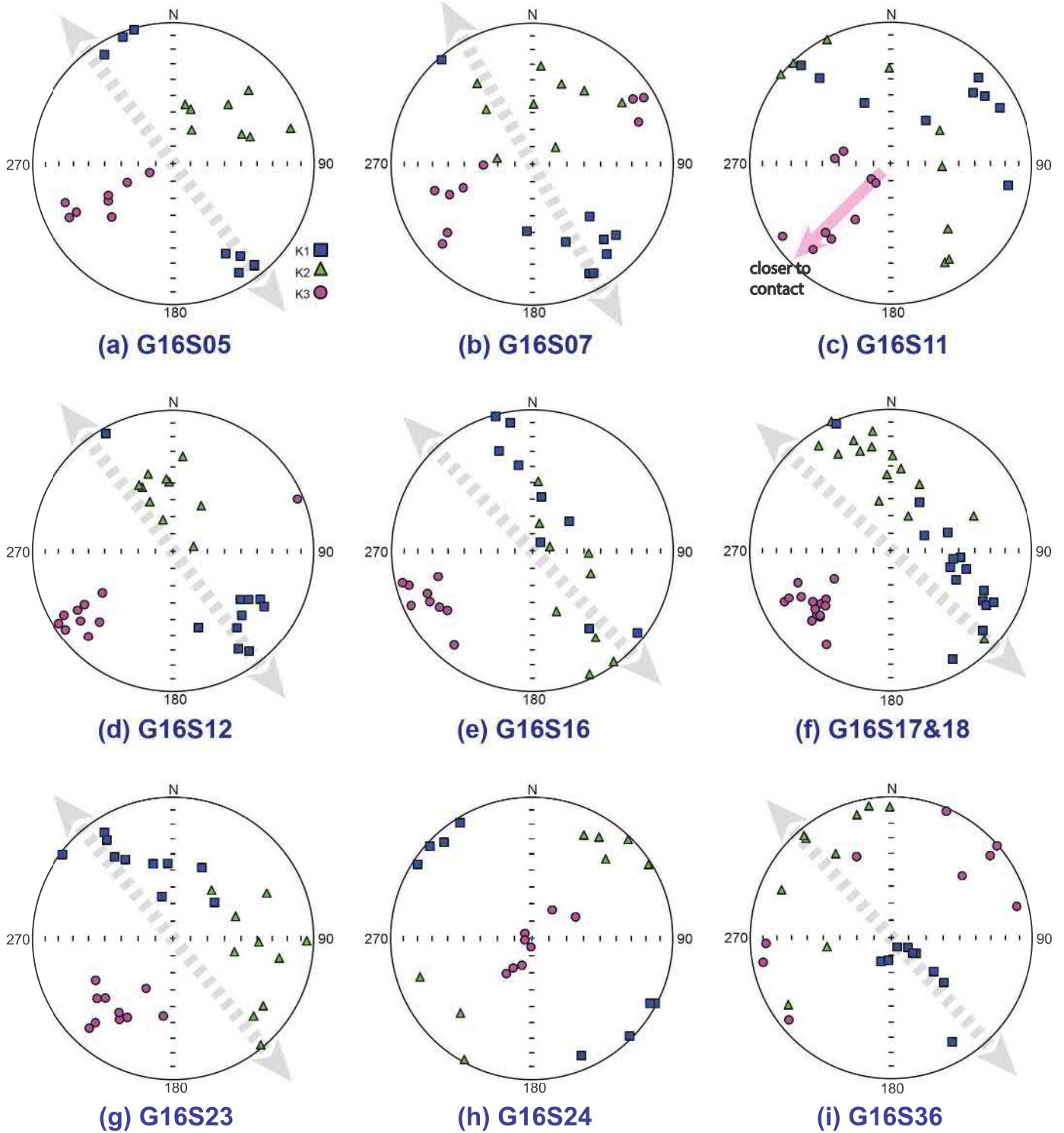


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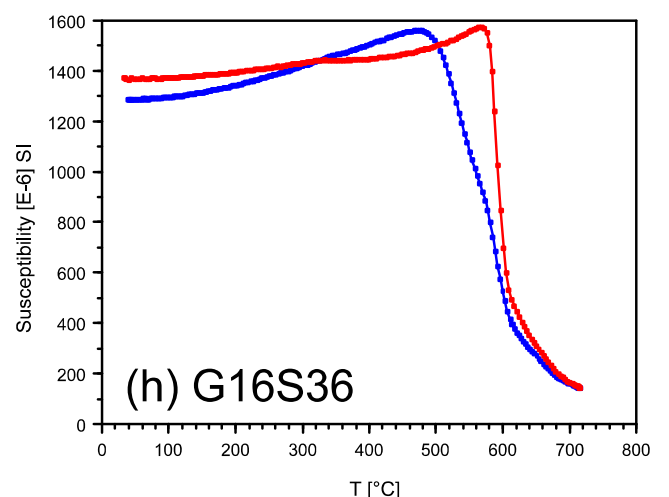
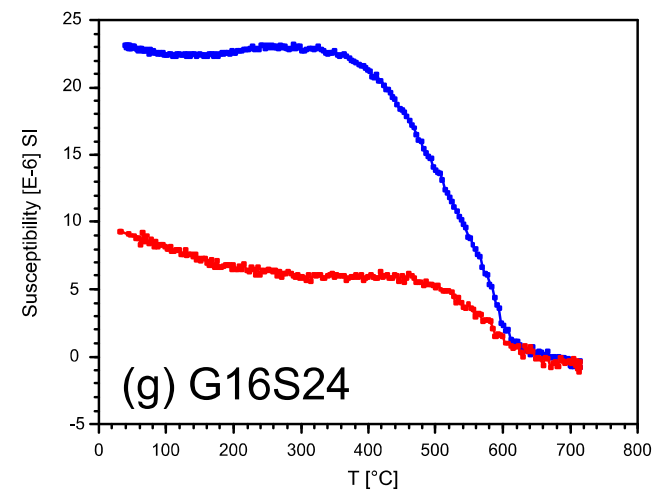
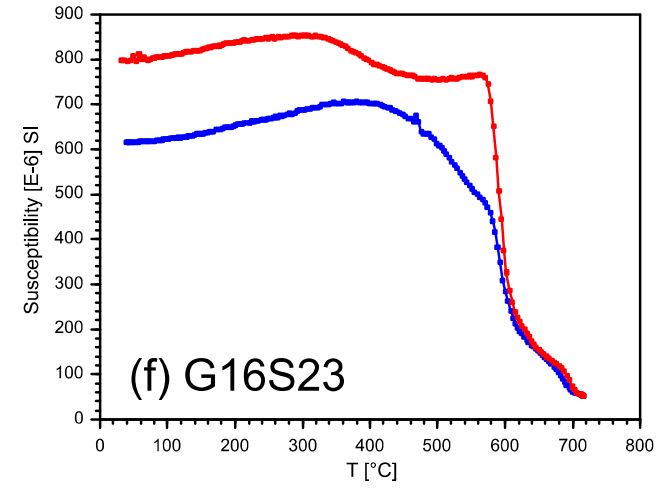
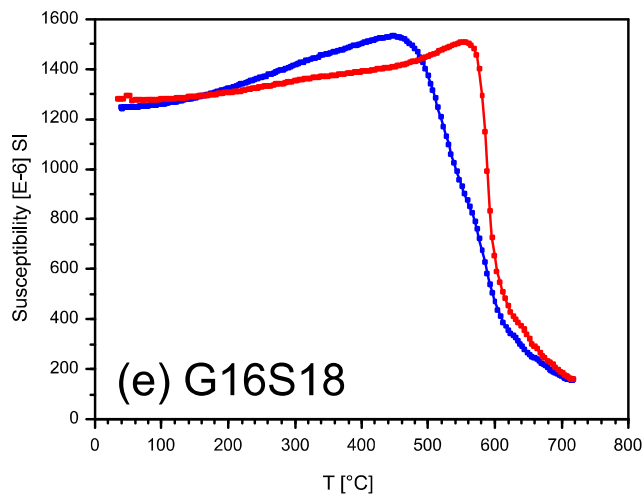
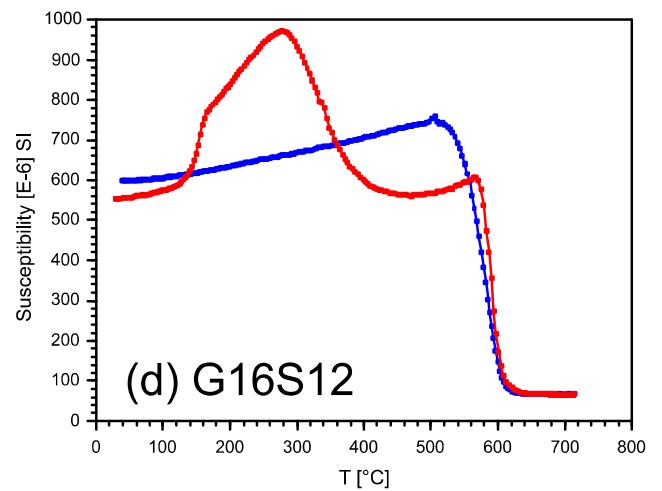
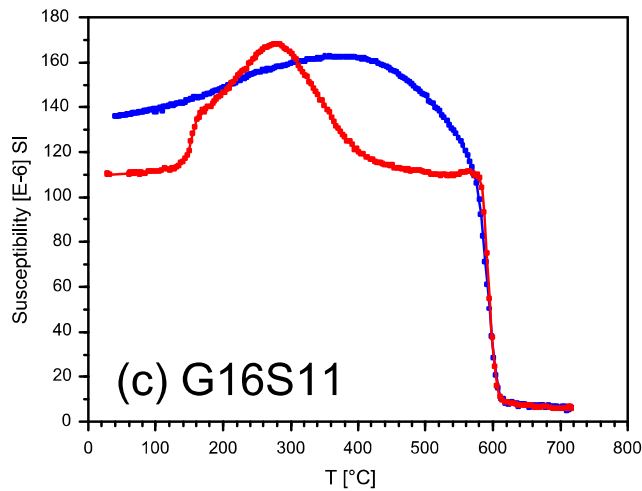
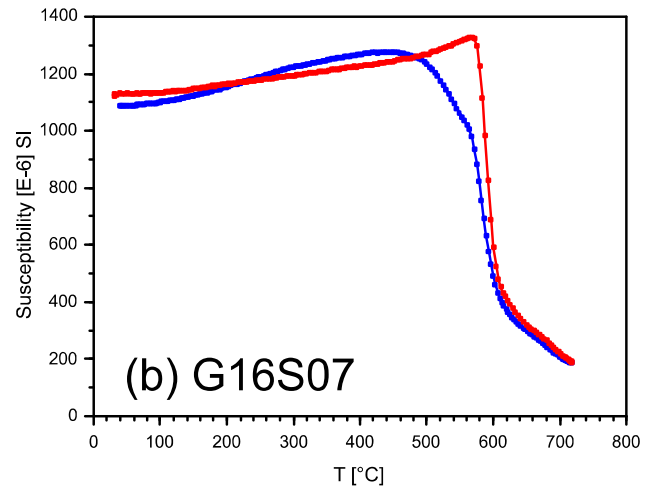
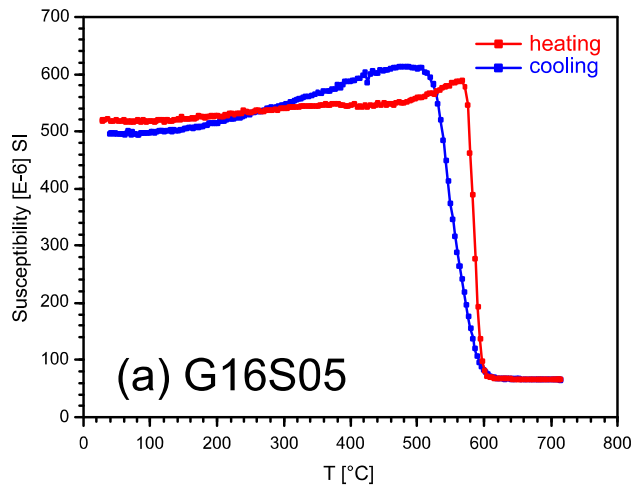
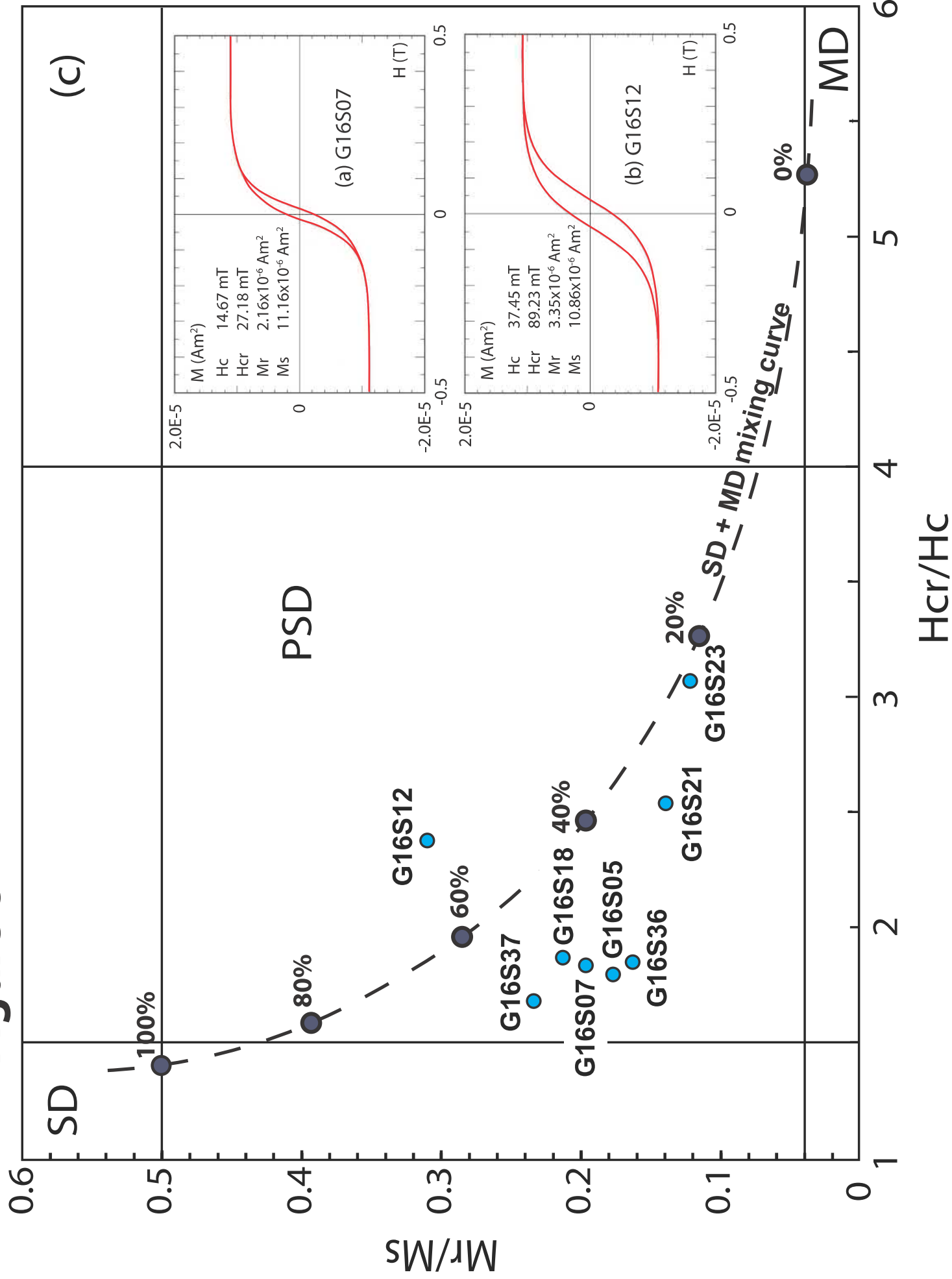


Figure 8



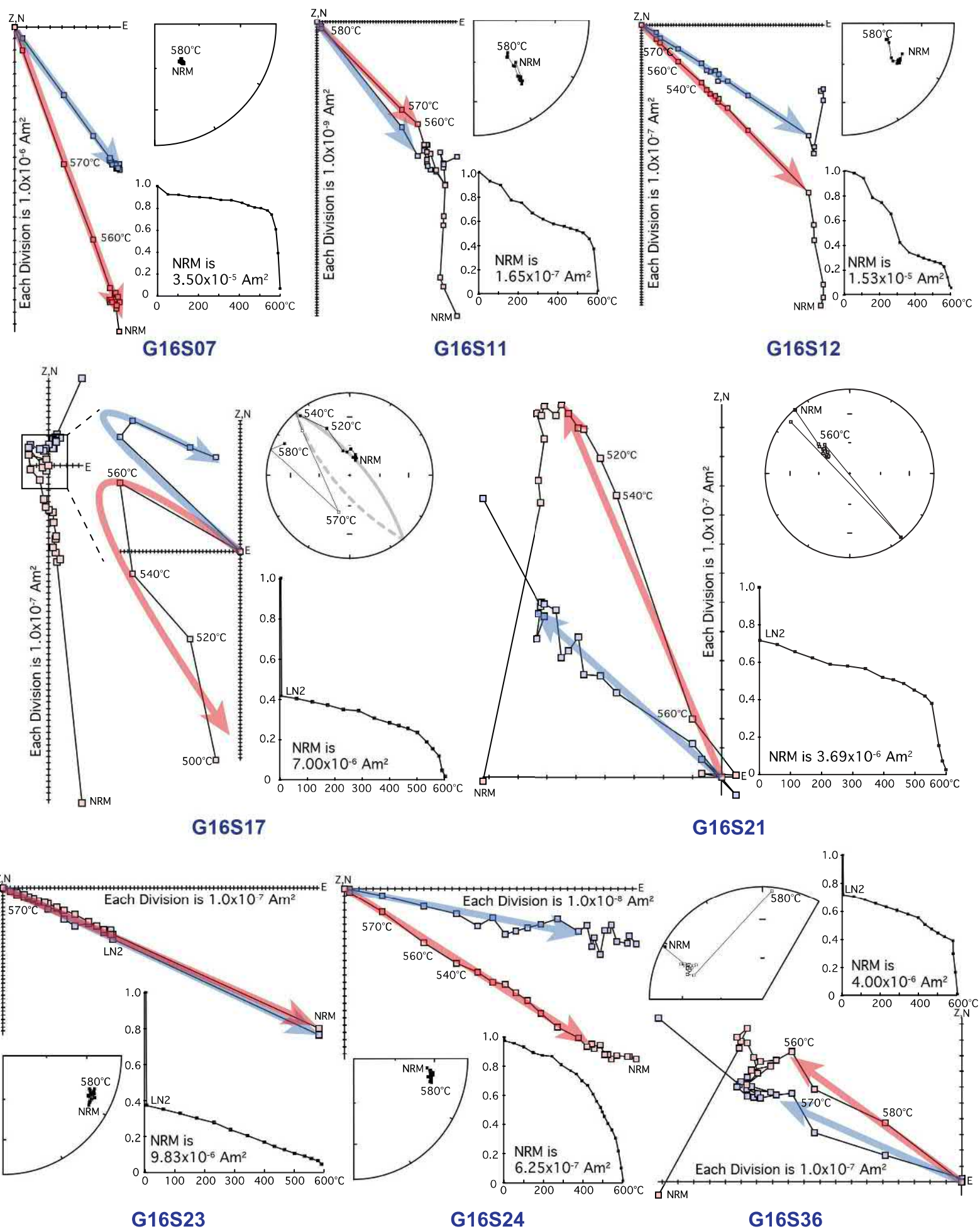


Figure 9

Figure 10

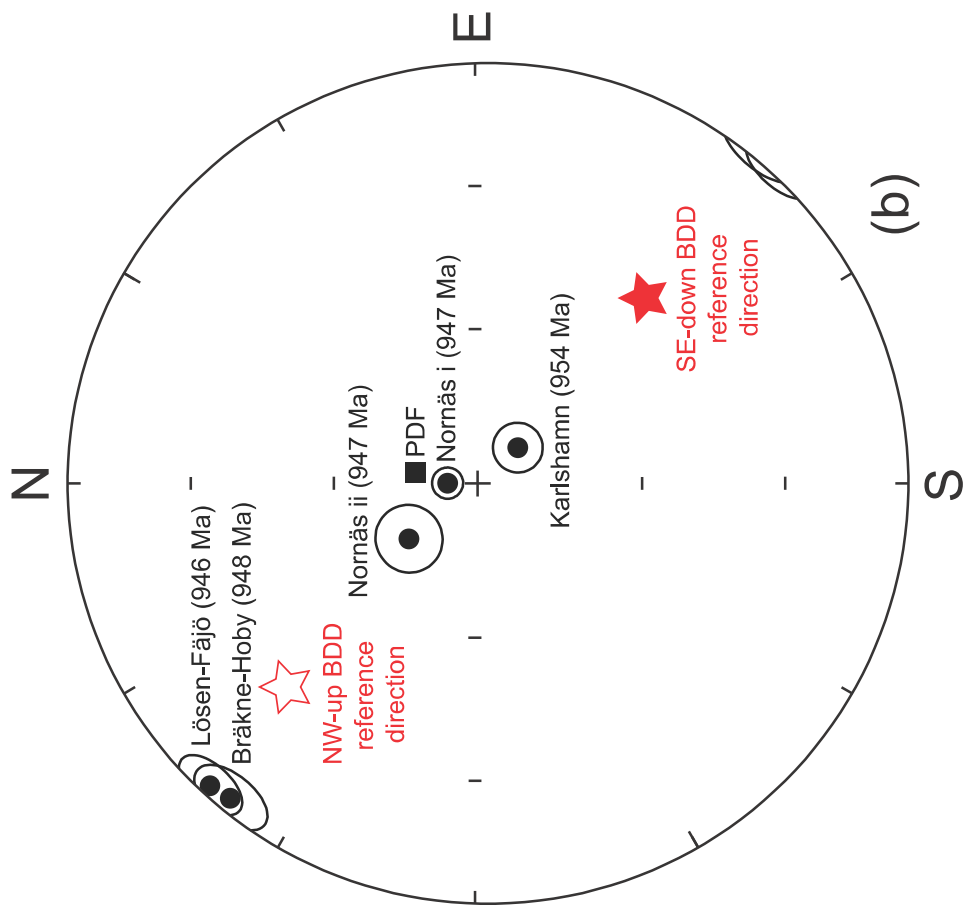
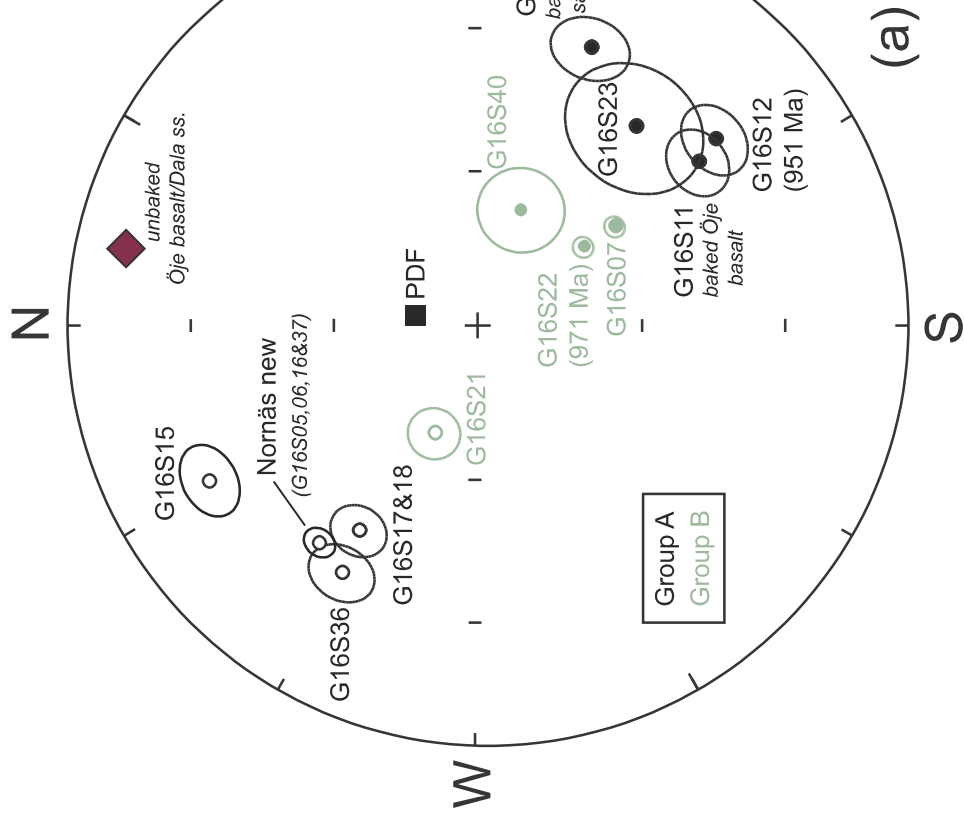
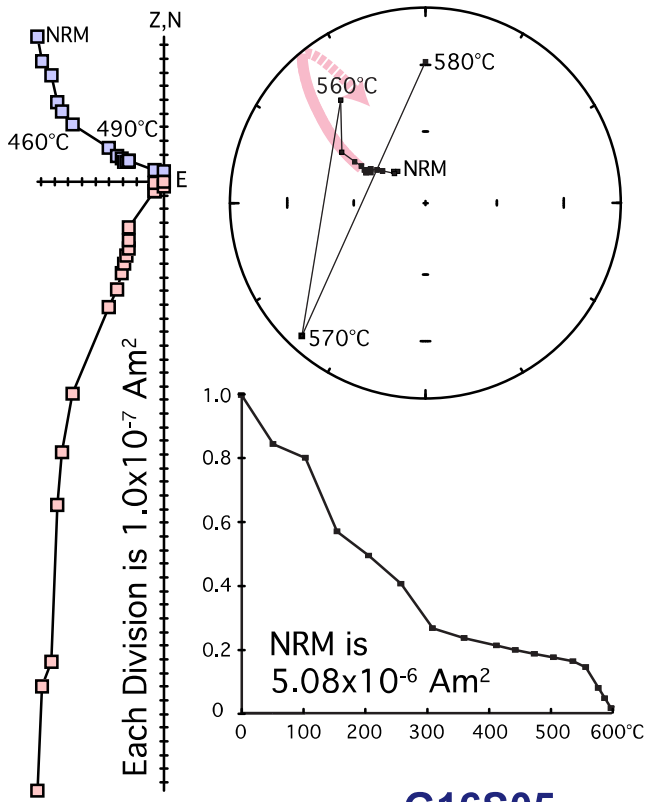
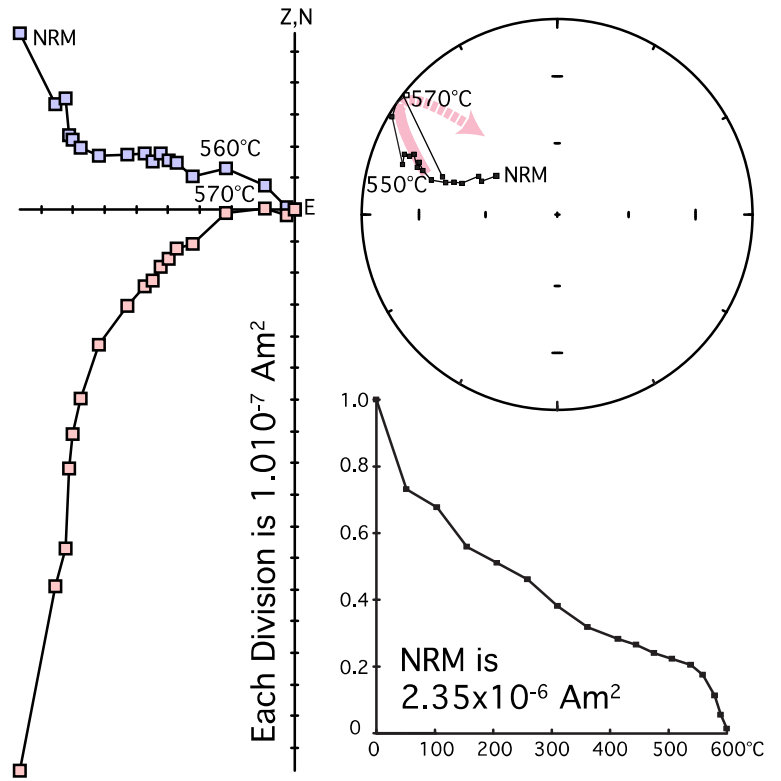


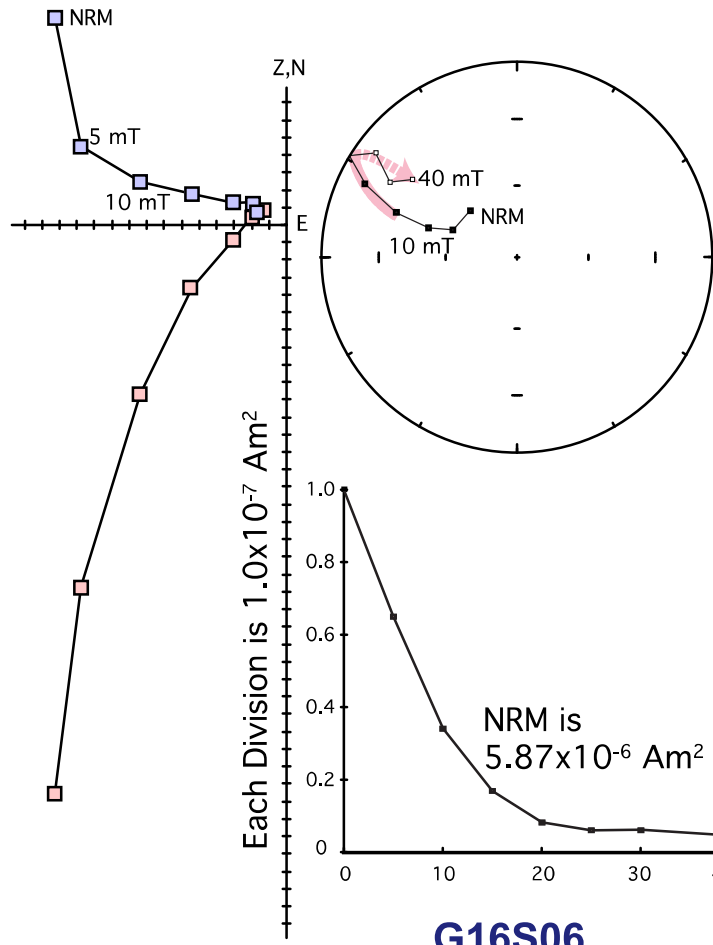
Figure 11



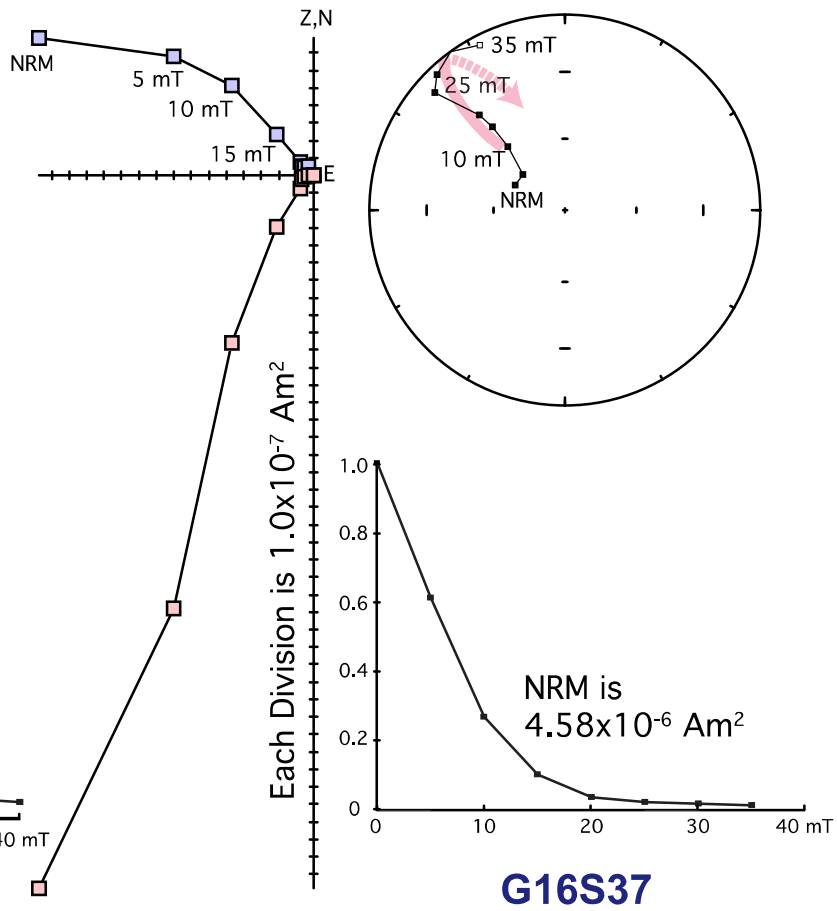
G16S05



G16S16



G16S06



G16S37

Figure 12

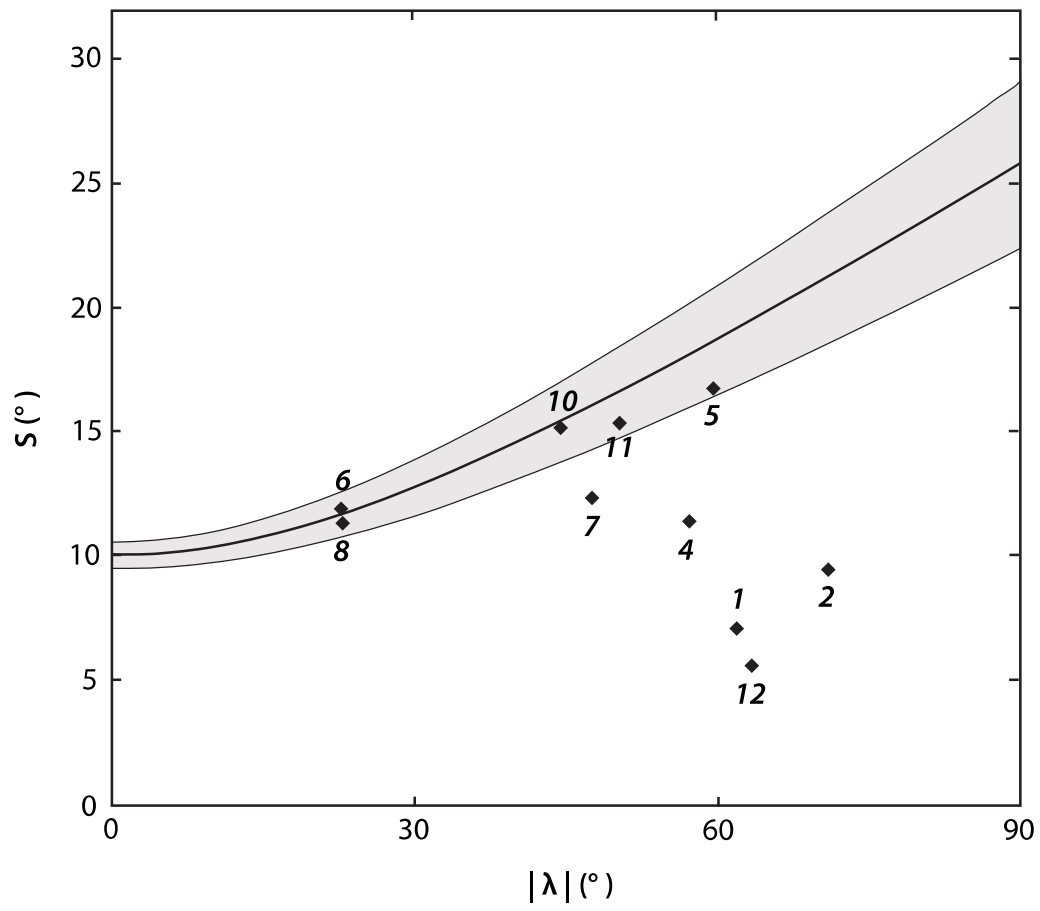


Figure 13

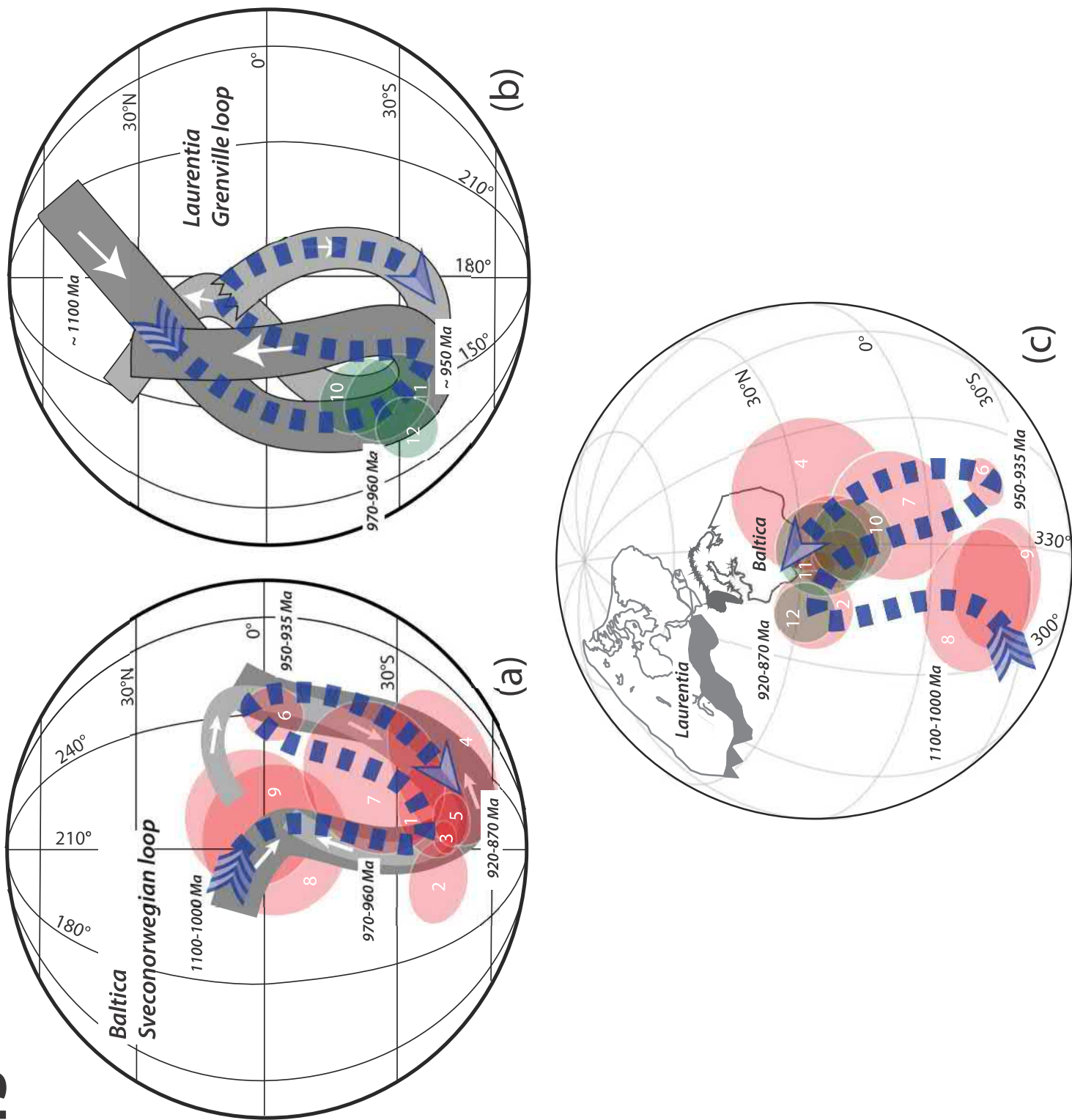


Figure 14

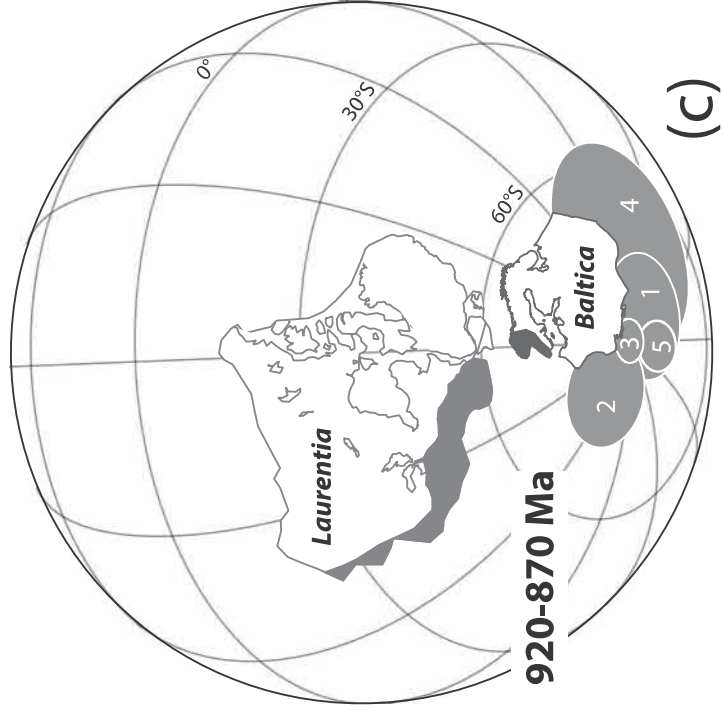
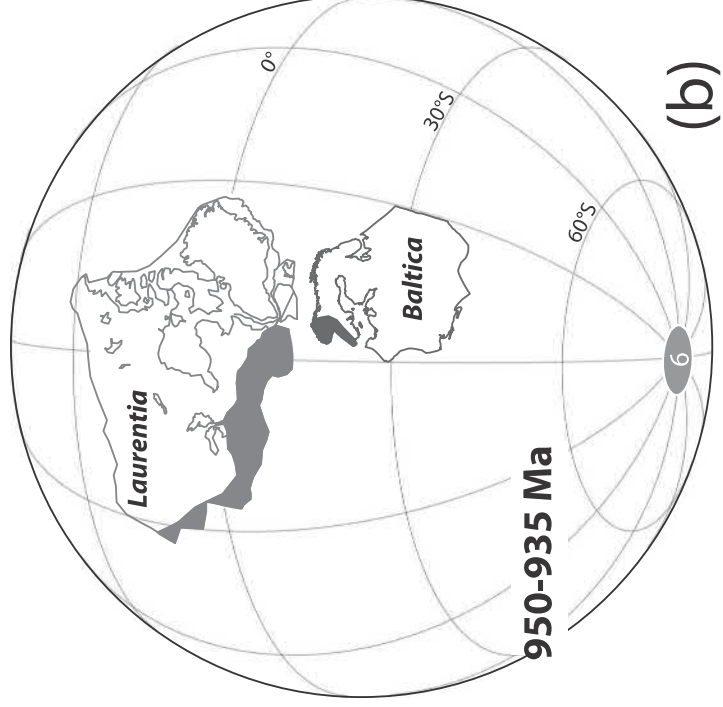
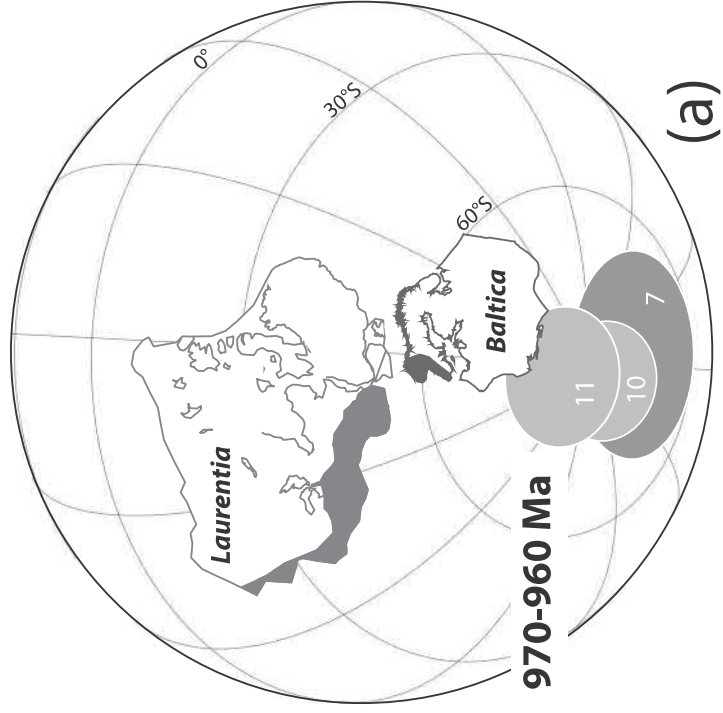


Table 1 Results of U-Pb baddeleyite geochronology

Sample (number of grains)	U/Th	Pb_b/Pb_{tot}^1	$\frac{^{206}Pb/^{204}Pb}{[Raw]^2}$	$\frac{^{207}Pb/^{235}U}{\pm 2s\% \text{ Error}}$	$\frac{^{206}Pb/^{238}U}{[Corr]^3}$	$\pm 2s\% \text{ Error}$	$\frac{^{207}Pb/^{235}U}{\pm 2s\% \text{ Error}}$	$\frac{^{206}Pb/^{238}U}{[Age, Ma]}$	$\frac{^{207}Pb/^{238}U}{\pm 2\sigma}$	$\frac{^{207}Pb/^{206}Pb}{\pm 2\sigma}$	Concordance
<i>G16S12</i>											
Bd-1 (2 grains)	27.4	0.085	818.2	1.5298	1.15	1.10	0.15711	940.7	946.5	9.2	0.994
Bd-2 (4 grains)	14.3	0.046	1449.3	1.5358	0.61	0.51	0.15707	940.4	955.1	7.0	0.985
Bd-3 (3 grains)	7.3	0.174	375.2	1.5337	1.90	1.82	0.15753	943.1	946.1	15.4	0.997
<i>G16S22</i>											
Bd-1 (2 grains)	5.9	0.311	214.3	1.5922	3.96	3.75	0.16185	967.1	967.4	33.6	1.000
Bd-2 (5 grains)	12.3	0.080	820.0	1.5437	0.82	0.73	0.15688	939.4	968.0	8.4	0.970
Bd-3 (5 grains)	9.0	0.168	365.5	1.5875	1.12	0.99	0.16072	960.8	975.8	11.6	0.985
<i>G16S37</i>											
Bd-1 (1 grains)	13.0	0.033	2026.6	1.5314	0.41	0.33	0.15714	940.8	948.2	5.2	0.992
Bd-2 (2 grains)	17.1	0.070	947.4	1.5281	0.74	0.68	0.15656	937.6	951.4	6.9	0.986
Bd-3 (2 grains)	7.6	0.119	520.8	1.5312	0.72	0.65	0.15773	944.1	940.3	7.0	1.004

¹ Pb_c = common Pb, Pb_{tot} = total Pb (radiogenic + blank + initial).

² Measured ratio, corrected for fractionation and spike.

³ Isotopic ratios corrected for fractionation (0.1% per amu for Pb), spike contribution, blank (0.4 pg Pb and 0.04 pg U), and initial common Pb. Initial common Pb corrected with isotopic compositions from the model of [Stacey and Kramers \(1975\)](#) at the age of the sample.

Table 2 Summary of paleomagnetic results of 800 dykes (from North to South)

Dyke	Stat (°N)	Slon (°E)	Age (Ma)	Dec (°)	Inc (°)	α_{95} (°)	N/n	Plat (°N)	Plon (°E)	A_{95} (°)	Group	Reference	Comment
G16522	61.6043	13.2260	971 ± 7	140.6	66.1	2.0	1/10	-34.5	220.8	3.0	Br	This study	NW-SE trending (324°)
G16521	61.5937	13.0047		296.5	-66.5	5.0	1/8	-32.1	236.3	7.5	Bn	This study	NW-SE trending
Normås ii	61.42	13.21	946.8 ± 1.2	325.0	71.3	6.5	1/4	77.0	282.0	10.5	-	Bylund (1985); Söderlund et al. (2005)	remnant at 10 mT
Normås i	61.42	13.21		0.4	82.2	3.0	1/6	77.0	14.0	6.0	-	Piper and Smith (1980); Söderlund et al. (2005)	-
G16506	61.4230	13.2150		Normås dyke, partial remagnetization and vector component contamination									
G16506	61.4207	13.2202		Normås dyke, partial remagnetization and vector component contamination									
G16507	61.2557	13.7790		142.1	58.7	2.0	1/9	-15.3	223.2	2.6	Br	This study	NW-SE trending (315°)
G16517	61.2504	13.4103		305.5	-45.5	4.2	1/4	-8.5	240.6	4.3	-	This study	NW-SE trending (315°), same dyke as G16518
G16518	61.2464	13.4186		300.5	-39.8	9.4	1/5	-6.4	246.6	8.7	-	This study	NW-SE trending (315°), same dyke as G16517
G16517&18 comb.				302.2	-41.7	5.2	1/9	-7.0	244.6	5.0	An	This study	-
G16524	61.1999	13.1505	~1460	110.3	30.4	6.8	1/8	-5.0	257.8	5.6	-	This study	baked Dala sandstone
G16523	61.1992	13.1515		126.6	40.9	12.9	1/7	-4.9	240.8	12.2	Ar	This study	NW-SE trending (320°)
G16514	61.1455	12.7273		demagnetization behavior like Normås dyke, partial remagnetization and vector component contamination									
G16516	61.1118	13.8530		southern extension of Normås dyke, partial remagnetization and vector component contamination									
G16515	61.0949	13.0151		330.7	-24.7	6.2	1/3	12.4	222.1	4.7	An	This study	NW-SE trending (325°)
G16500	60.9265	13.1784		inconsistent directions within dyke									
G16537	60.9181	14.1043	947 ± 4	southern extension of Normås dyke, partial remagnetization and vector component contamination									
Normås new (G16505,06,16&37 comb.)				307.9	-35.3	2.9	4/22	-0.8	241.7	2.6	-	This study	Mean direction calculated using great circle method
G16510	60.8878	13.3006		inconsistent directions within dyke									
G16536	60.8284	14.2725		300.6	-32.0	5.9	1/8	-1.4	249.5	5.0	An	This study	NW-SE trending (305°), same dyke as G16509
G16511	60.8077	13.5783	~1460	140.7	30.9	6.5	1/9	-14.9	223.0	2.6	-	This study	NW-SE trending
G16512	60.8077	13.5783	951 ± 5	142.2	36.9	6.2	1/9	3.1	228.7	5.6	Ar	This study	baked Öje basalt
Andersbo	60.76	15.41		143.0	42.0	9.0	1/7	1.0	229.0	7.3	-	Bylund and Elming (1992)	-
G16513	60.7147	13.3904		153.2	8.3	7.3	1/8	21.8	222.4	5.2	-	This study	NW-SE trending
Ejen	60.67	14.79		128.0	24.0	9.0	1/8	6.0	245.8	6.6	-	Bylund (1985)	remnant at 20 mT
Falun (new)	60.59	15.58	945.7 ± 1.2	126.7	37.9	3.0	1/9	-1.7	242.3	2.6	Ar	Elming et al. (2014); Söderlund et al. (2005)	NW-SE trending (330°)
Falun (old)	60.59	15.58	945.7 ± 1.2	131.4	45.8	5.6	1/10	-6.1	237.6	5.7	-	Patchett and Bylund (1977); Söderlund et al. (2005)	remnant at 20 mT
Gällsjön	60.51	14.53		136.0	48.0	12.0	1/5	18.0	241.0	8.5	-	Bylund and Elming (1992)	-
G16540	60.3185	14.2210		106.0	67.0	8.3	1/9	-35.0	243.7	12.5	Br	This study	NW-SE trending (330°)
G16539	60.3024	13.8662		inconsistent directions within dyke									
G16538	60.3016	13.8672		inconsistent directions within dyke									
Årby	59.27	16.46		142.5	52.4	7.0	1/6	-7.3	227.4	8.0	-	Patchett and Bylund (1977)	remnant at 20 mT
Marbystrand	58.61	16.48		127.5	31.1	3.6	1/9	3.3	246.0	3.0	-	Elming et al. (2014)	NW-SE trending (310°)
Y1	58.58	16.33	938 ± 4	311.2	-43.3	5.7	1/13	-3.4	239.0	4.0	An	Elming et al. (2014)	NE-SE trending (40°)
Y2	58.58	16.33	941 ± 6	313.3	-40.7	7.4	1/14	-0.5	238.3	7.0	An	Elming et al. (2014)	NE-SE trending (40°)
Slussen	58.2	12.2		300.9	-26.5	11.5	1/6	3.2	248.5	9.2	An	Pisarevsky and Bylund (2006)	-
Tove	57.8	11.8	935 ± 3	122.2	53.7	3.2	2/23	-14.0	237.9	3.7	Ar	Pisarevsky and Bylund (2006); Hellström et al. (2004)	-
Small	57.8	11.8		124.5	54.6	7.4	1/3	-13.9	235.8	8.8	Ar	Pisarevsky and Bylund (2006)	-
Hjuvik	57.7	11.7		118.8	37.1	11.5	1/12	-3.3	246.9	10.3	Ar	Pisarevsky and Bylund (2006)	-
Sjunnaryd (sill)	57.7	14.8		317.1	-28.9	15.1	1/4	8.8	246.4	12.4	-	Pesonen and Klein (2014)	-
Niörp (sill)	57.64	14.83	966 ± 5	315.2	-27.4	9.7	1/6	9.0	238.5	7.8	-	Patchett and Bylund (1977); Söderlund et al. (2005)	remnant at 40 mT
Tarnö	56.27	15.06		312.8	-41.9	4.8	1/5	0.3	237.0	4.6	-	Patchett and Bylund (1977)	remnant at 30 mT
Bräkne-Hoby	56.26	15.18	948.1 ± 1.4	309.3	3.6	6.1	1/7	22.2	251.9	4.3	-	Patchett and Bylund (1977); Söderlund et al. (2005)	remnant at 40 mT
Lösen-Fåjö	56.17	15.69	945.8 ± 1.0	312.6	2.6	5.4	1/6	23.3	248.9	3.8	-	Patchett and Bylund (1977); Söderlund et al. (2005)	remnant at 20 mT
Våby	56.17	15.22		137.0	33.0	10.0	1/3	8.0	236.0	8.4	-	Poorter (1975)	remnant at 10 mT
Karlshamn	56.15	14.86	954.2 ± 1.1	129.0	81.0	4.8	1/9	-43.2	213.5	9.1	-	Patchett and Bylund (1977); Söderlund et al. (2004a)	remnant at 30 mT
Group A Mean	-	-	-	127.6	65.4	9.7	12/116	-2.6	239.6	5.8	-	-	Averaging 12 dykes in Group An/Ar listed in this table
Group B Mean	-	-	-	128.8	39.6	6.5	4/36	-27.0	230.4	14.9	-	-	Averaging 4 dykes in Group Bn/Br listed in this table

Note: Stat/Slon = site geographic latitude/longitude, Dec = declination, Inc = inclination, Plat/Plon = paleomagnetic pole latitude/longitude, N/n = number of site/sample, α_{95} = radius of 95% confidence cone of the mean direction, A_{95} = radius of 95% confidence cone of the paleomagnetic pole

*least scatter of data at alternating field step used for mean vector calculation

Table 3 Paleomagnetic poles used to constrain the paleogeographic reconstruction of Baltica and Laurentia at ca. 1000-850 Ma

#	Paleomagnetic Pole	Dec (°)	Inc (°)	α_{95} (°)	k	N/n	Plat (°N)	Plon (°E)	A_{95} (°)	S (°)	$ \lambda $ (°)	Age (Ma)	1	2	3	4	5	6	7	Q	Reference
Baltica																					
1	Hunedalen dykes	294.0	-75.0	6.0	115	6/69	-41.0	222.0	10.5	7.2	61.8	848 ± 27 ^a ; 855 ± 59 ^b	1	1	1	0	1	0	1	5	Walderhaug et al. (1999)
2	Egersund-Ogna anorthosite	325.9	-80.1	4.9	73	13/69	-42.1	200.4	9.0	9.5	70.8	900 ^c	1	1	1	0	0	0	1	4	Brown and McEnroe (2004)
3	Egersund anorthosite	-	-	-	-	76/-	-43.5	213.7	3.6	-	-	900 ^c	0	1	1	0	0	0	1	3	Stearn and Piper (1984); Walderhaug et al. (1999)
4	Rogaland Igneous Complex	269.0	-72.0	11.0	49	5/24	-45.9	238.4	18.2	11.5	57.0	869 ± 14 ^d	1	0	1	1	0	0	1	4	Walderhaug et al. (2007)
5	Bjerkreim-Sokndal intrusion	303.4	-73.5	3.7	24	66/354	-35.9	217.9	6.0	16.8	59.4	916 ^c	1	1	1	0	1	0	1	5	Brown and McEnroe (2015)
6	Mean 951-935 Ma Baltica pole	308.8	-39.6	6.5	42	12/116	-2.6	239.6	5.8	12.0	22.5	951-935 ^d	1	1	1	1	1	1	1	7	This study
7	971 Ma BDD dykes VGP	307.6	-65.4	14.2	68	4/36	-27.0	230.4	14.9	12.4	47.5	971 ± 7 ^d	1	0	1	0	1	1	1	5	This study
8	Lamila-Ristijärvi dykes	355.5	-40.0	17.5	51	3/7	-2.1	212.2	16.4	11.4	22.8	1042 ± 50 ^d	0	0	1	1	1	0	0	3	Mertanen et al. (1996)
9	Bamble intrusion mean	-	-	-	-	-	3.0	217.0	15.0	-	-	1100-1040	0	1	1	0	0	0	0	2	Meert and Torsvik (2003)
Laurentia																					
10	Adirondack microcline gneisses	289.2	-62.8	7.6	29	14/80	-18.4	151.1	10.5	15.2	44.2	960 ^c	1	1	1	0	0	1	1	5	Brown and McEnroe (2012)
11	Adirondack metamorphic anorthosites and other rocks	283.9	-67.3	7.7	28	14/68	-25.1	149.0	11.6	15.4	50.1	970 ^c	1	1	1	0	0	1	1	5	Brown and McEnroe (2012)
12	Adirondack fayalite granites	297.0	-75.8	3.9	199	8/40	-28.4	132.7	6.9	5.7	63.2	990 ^c	1	1	1	0	0	1	1	5	Brown and McEnroe (2012)

Note: Dec = declination, Inc = inclination, α_{95} = radius of 95% confidence cone of the mean direction, k = precision parameter, N/n = number of site/sample, Plat/Plon = paleomagnetic pole latitude/longitude, A_{95} = radius of 95% confidence cone of the paleomagnetic pole, S = angular dispersion of VGPs, $|\lambda|$ = absolute value of paleolatitude, Q = sum of quality criteria (Van der Voo, 1990)

^a Ar-Ar biotite ages

^b Sm-Nd whole rock ages

^c Cooling ages

^d U-Pb baddeleyite ages