1	Origin and geodynamic relationships of the Late Miocene to Quaternary alkaline basal	
2	volcanism in the Pannonian Basin, eastern-central Europe	
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14	Abstract	
15	Alkaline basaltic volcanism has been taking place in the Carpathian-Pannonian Region	
16	since 11 Ma and the last eruptions occurred only at 100-500 ka. It resulted in scattered low-	
17	magma volume volcanic fields located mostly at the margins of the Pannonian basin. Many of	
18	the basalts have compositions close to those of the primitve magmas and therefore can be	
19	used to constrain the conditions of the magma generation. Low degree (2-3%) melting could	
20	occur in the convective asthenosphere within the garnet-spinel transition zone. Melting started	
21	at about 100 km depth and continued usually up to the base of the lithosphere. Thus, the final	
22	melting pressure could indicate the ambient lithosphere-asthenosphere boundary. The	
23	asthenospheric mantle source regions of the basalts were heterogeneous, presumably in small	
24	scale, and included either some water or pyroxenite/eclogite lithology in addition to the fertile	
25	to slightly depleted peridotite. Based on the prevailing estimated mantle potential temperature	

(1300-1400°C) along with number of further observations we exclude the existence of mantle 26 27 plume or plume fingers beneath this region. Instead, we propose that plate tectonic processes controlled the magma generation. The Pannonian basin acted as a thin-spot after the 20-12 Ma 28 29 syn-rift phase and provided suction in the sublithospheric mantle, generating asthenospheric flow from below the adjoining thick lithospheric domains. A near vertical upwelling along the 30 31 steep lithosphere-asthenosphere boundary beneath the western and northern margin of the 32 Pannonian basin could result in decompressional melting producing low-volume melts. The 33 youngest basalt volcanic field (Perşani) in the region is inferred to have been formed due to the dragging effect of the descending lithospheric slab beneath the Vrancea zone that could 34 35 result in narrow rupture at the base of the lithosphere. Continuation of the basaltic volcanism cannot be excluded as inferred from the still fusible condition of the asthenospheric mantle. 36 37 This is reinforced by the detected low-velocity seismic anomalies in the upper mantle beneath 38 the volcanic fields.

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40 Keywords: basalt, spinel, Carpathian-Pannonian region, magma genesis, monogenetic
41 volcanic field, asthenosphere flow

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43 Introduction

Origin of the basaltic magmas in monogenetic volcanic fields is a key issue to understand the nature of low-magma-flux, episodic, but long lasting volcanic activity (Connor and Conway, 2000; Wang et al., 2002; Aldanmaz et al., 2006; Jung et al., 2006; 2012; Valentine and Perry, 2007; Blondes et al., 2008; Haase and Renno, 2008; Brenna et al., 2012; McGee et al., 2013). Each volcano in such volcanic system is considered to represent eruption of a discrete magma batch ascending rapidly from the mantle source, although recent studies revealed that composition of erupted magmas could change even within the eruptive sequence

of a monogenetic volcano (Strong and Wolff, 2003; Brenna et al., 2010; Erlund et al., 2010; 51 McGee et al., 2012). Apart of these exceptions, basalts of monogenetic volcanoes provide a 52 53 unique snapshot of the nature of the mantle source regions and could reveal their subtle 54 compositional differences. Furthermore, basaltic products having composition close to the primitive magmas can be used to infer the condition of the mantle generation (pressure and 55 temperature, mineral and chemical composition of the melted mantle rock, degree of melting, 56 mixing of melts etc.). Using this knowledge, a better understanding can be achieved about the 57 58 triggering mechanism of the melt generation and the geodynamic relationships of the volcanism. Nevertheless, recent studies have revealed that evolution of basaltic magmas could 59 60 be far more complex than previously thought (Woodland and Jugo, 2007; Smith et al., 2008; 61 Jankovics et al., 2012; 2013) and careful investigation is necessary to select the most appropriate samples, which can be used to gain insight into the mantle and the melting 62 63 process.

The Carpathian-Pannonian region (CPR; Fig. 1.) is a key natural laboratory for 64 65 studying the interaction of deep mantle processes, igneous activity and tectonic events (Szabó 66 et al., 1992; Seghedi et al., 1998; 2004a; 2004b; 2005; Harangi, 2001a; Konečný et al., 2002; Harangi and Lenkey, 2007; Seghedi and Downes, 2011). Alkaline basalt monogenetic 67 68 volcanic fields have developed since 11 Ma, partly overlapping and partly postdating the 69 extensive calc-alkaline andesitic-dacitic volcanism (Pécskay et al., 2006; Lexa et al., 2010). 70 The alkaline basaltic volcanism is considered as part of the widespread Neogene to 71 Quaternary volcanism in western and central Europe (Wilson and Downes, 1991; 2006; 72 Embey-Isztin and Dobosi, 1995; Lustrino and Wilson, 2007), but it shows many similarities also with the alkaline basaltic volcanism occurred with close spatial and temporal association 73 74 of calc-alkaline and potassic-ultrapotassic volcanism in the Mediterranean (Wilson and Bianchini, 1999; Harangi et al., 2006; Beccaluva et al., 2011). The origin of the basaltic 75

volcanism in the CPR is still controversial (Embey-Isztin and Dobosi, 1995; Embey-Isztin et
al., 2001; Harangi, 2001a; 2001b; Seghedi et al., 2004a; Harangi and Lenkey, 2007; Ali and
Ntaflos, 2011; Ali et al., 2013; Harangi et al., 2013), largely due to its occurrence after the
main extensional phase of the Pannonian basin. Eruption of basaltic magmas took place even
at a couple of 100's ka, thus, understanding the mechanism of melt generation is crucial also
to evaluate rigorously whether this kind of volcanic activity could continue in the future.

In this paper, we discuss the conditions of the basaltic magma generation in the CPR using primarily those samples, which have chemical composition close to primitive magmas. We combine major and trace element data of the host basalts as well as compositional features of spinels, one of the liquidus minerals in these rocks, to constrain the origin of the magmas and the nature of the source regions. Then, we propose a possible mechanism that could be responsible for the formations of the basaltic volcanoes in this region.

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89 Geological setting

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91 The Carpathian-Pannonian region lies in the northeastern part of the Alpine-Mediterranean region and is characterized by an arcuate orogenic belt (Carpathians) with a 92 basin area behind it (the Pannonian basin; Fig. 1A). The Pannonian basin is surrounded by the 93 94 Alpine, Carpathian and Dinarides mountain belts. It is underlain by thin lithosphere (50-80 km) and crust (22-30 km) coupled with high heat flow (>80 mW/m²; Csontos et al., 1992; 95 96 Horváth, 1993; Tari et al., 1999; Lenkey et al., 2002; Horváth et al., 2006). It is regarded as a special type of back-arc basin formed by heterogeneous stretching of the lithosphere. The 97 thinning of the lithosphere started in the Early Miocene (around 20 Ma) and the syn-rift phase 98 99 was finished by the end of Miocene (around 12 Ma; Horváth et al., 2006). Tari et al. (1999) 100 defined an initial core-complex style extension during the Early Miocene followed by a wide-

101 rift type extension. Huismans et al. (2001) distinguished two main rifting events, involving a 102 Middle Miocene passive rifting followed by a Late Miocene active rifting episode. As a 103 consequence, the lithospheric mantle underwent a dramatic attenuation (stretching factor of 104 5-50, whereas the crust was only moderately thinned (stretching factor of 1.2-2.5) leading to 105 the suggestion of an inhomogeneous stretching model (Sclater et al., 1980; Horváth et al., 106 1988; 2006). As an extreme case, Horváth et al. (2006) assumed the disappearance of the 107 whole mantle lithosphere beneath the Pannonian Basin at the beginning of the synrift phase. 108 The thinning of the lithosphere was accompanied by asthenospheric mantle updoming as 109 shallow as 50-60 km (Stegena et al., 1975; Huismans et al., 2001; Konečný et al., 2002) and 110 therefore, the Pannonian basin can be considered as one of the hottest basin areas of Europe 111 (Cloetingh et al., 2007). The syn-rift phase was followed by post-rift thermal subsidence, 112 when several thousands meters of Late Miocene to Quaternary sediments filled parts of the 113 basin areas (Great Hungarian Plain, Little Hungarian Plain).

114 There are diverse views on the driving mechanism of the lithospheric extension such 115 as (a) formation of an overthickened hot lithosphere, which underwent gravitational collapse 116 during the Early Miocene (Tari et al., 1999; Horváth et al., 2006); (b) suction of the slab roll-117 back beneath the retreating subduction along the Eastern Carpathians (Csontos et al., 1992); 118 (c) dripping the lower, gravitationally instable lithospheric material into the deep mantle 119 (Lorinczi and Houseman, 2000) and (d) asthenospheric flow either via the "Istria window" 120 beneath the Dinarides (Horváth and Faccenna, 2011) or from the Adria-European collision 121 zone and ensuing Alpine orogeny (Kovács et al., 2012). While the models (a) and (b) are 122 supported by geophysical and geological observations and are widely accepted by the scientific community, the novel models (c) and (d) have still a number of inconsistencies and 123 124 therefore need more supporting evidences.

125 Since the Pliocene, the CPR has been in a tectonic inversion stage (Horváth and 126 Cloetingh, 1996; Bada et al., 1999) due to the continuous push from west (Adria-push from 127 the Alps–Dinarides belt) and blocking the area at east by the thick Ukrainian lithosphere. The 128 entire basin system has been landlocked and this led to a gradual structural inversion in the 129 form of multi-scale folding and fault reactivation. Some parts of the CPR (Great Hungarian 130 Plain, Little Hungarian Plain and Danube basin, Vienna basin, Styrian basin and Brasov 131 basin) are still subsiding, while other parts show uplift (Dombrádi et al., 2010). This 132 differential vertical movement is accommodated by reactivation of older faults. Major strike-133 slip faults have a SW-NE orientation at the northern and central part of the CPR, whereas to 134 the south a characteristic E-W trending strike-slip fault system was identified (Horváth et al., 135 2006; Fig. 1B).

The nature of the upper mantle beneath the CPR is inferred from seismic tomographic models. These show a high-velocity body in the transition zone (400-670 km) interpreted as accumulated subducted slab material (Wortel and Spakman, 2000; Piromallo and Morelli, 2003; Hetényi et al., 2009), a near-vertical high-velocity slab beneath the southeastern margin of the CPR (Vrancea zone), regarded as the trace of final-stage subduction (Sperner et al., 2001; 2004) and relatively low velocity material between the transition zone and the base of lithosphere.

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144 Areal distribution of the basalts

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In the CPR, four main monogenetic volcanic fields were developed (Styrian, BakonyBalaton Upland, Nógrád-Gemer, Perşani; Fig. 1B). In addition, localized basaltic centres were
developed in a few areas (Pauliberg-Oberpullendorf in Burgenland, Lucaret-Sanovita in
Banat), where no precursor and subsequent volcanism occurred. Basalt volcanic centres of the

eastern part of the Little Hungarian Plain (called here Kemenes Volcanic Field, Fig. 1B) and
around Štiavnica (ST in Fig. 1B) can be considered to form ultra-low-magma-flux volcanic
fields, where volcanic activity was sporadic (<10 volcanic centres) during a considerable (2-6
Myr) timescale and characterized by several 100's ka (several Myr in ST) long quiescence
time between the active events.

155 The basaltic volcanism was controlled by external conditions, such as tectonics, 156 topography, presence and style of aquifer in the basement (Martin and Németh, 2004; 157 Kereszturi et al., 2011). Tuff rings and maar structures were developed as a result of 158 phreatomagmatic explosive eruptions followed occasionally by pure magmatic explosive 159 volcanism and in a few cases lava lakes were formed in the late-stage effusive phase. Large 160 volcanic edifices, such as shield volcanoes were built up in each volcanic field (Stradner 161 Kogel in Styria, Kab, Agár and Fekete hills in the Bakony-Balaton Upland, Medvedia in 162 Nógrád-Gemer) except for the Perşani. The volcanism was dominantly basaltic, the only 163 exception is the Kemenes Volcanic Field, where a large trachyandesite-trachyte volcano 164 developed around 10-11 Ma (Pásztori volcano; Harangi et al., 1995; Harangi, 2001b). This 165 volcano subsided subsequently and was covered by thick (over 2000 m) Late Miocene to 166 Quaternary sediments.

167 The geographic location of the volcanic fields in the CPR (Fig. 1B) is remarkable, 168 since they are mostly at the western and northern margins of the Pannonian basin and not in 169 the central area, where there is the thinnest lithosphere, whereas the youngest one is found at 170 the southeastern part of the region, close to the seismically active Vrancea zone. The two 171 largest volcanic fields in the CPR are the Bakony-Balaton Upland Volcanic Field (Martin and 172 Németh, 2004) and the Nógrád-Gemer Volcanic Field (Konečný et al., 1995; 1999), each of 173 them involves about 50 eruptive centres, have relatively long timescale of the volcanism (6-8 174 Ma) and includes intermittent active phases with relatively long (several 100's ka) repose

time. Kereszturi et al. (2011; 2013) calculated the recurrence rate for both volcanic fields and 175 got about 0.11 and 0.08 Ma/event, respectively. Although the volcanic edifices are deeply 176 eroded in both areas, they estimated also the magma-flux that is 0.9 km³/Ma (calculated only 177 for the last 5 active phases, excluding the initial phase) and 0.4 km³/Ma, respectively (the total 178 erupted magma volume is estimated to be about 4 and 3.5 km³, respectively). These numbers 179 180 are comparable with the low magma-flux monogenetic volcanic fields, such as the Southwestern Nevada Volcanic Field (Valentine and Perry, 2007) and classify them as typical 181 182 tectonically-controlled, time predictive volcanic fields. The Styrian and Persani Volcanic 183 Fields (Seghedi and Szakács, 1994; Downes et al., 1995; Ali and Ntaflos, 2013; Harangi et al., 184 2013) contain less eruptive centres, although in Styria, the eruptive volume could be relatively large (>1 km³ total volume) due to the significant amount of effusive volcanic products 185 186 (Stradner Kogel, Klöch, Steinberg). The Persani Volcanic Field is the youngest one in the 187 CPR and is located in a 22 km long and 8 km wide area (Seghedi and Szakács, 1994) at the 188 northwestern periphery of the intramontane Braşov basin (Fig. 1B.). This area is characterized 189 by NE-SW trending normal faults, resulting from a NW-SE extensional regime (Gîrbacea et 190 al., 1998). The eruptive centres appear to be structurally controlled and show a rough NE-SW 191 trending alignment. Popa et al. (2012) recorded subcrustal seismicity beneath the Persani area 192 and, using seismic tomography modelling, suggested that a low-velocity anomaly could be 193 inferred at the crust-mantle boundary. The youngest basalt volcano (Putikov vŕšok; Šimon and Halouzka, 1996) of the CPR is located near Banska Štiavnica upon the Late Pleistocene 194 195 terrace of the river Hron. It does not rigorously belong to the Nógrád-Gemer Volcanic Field 196 and was formed after more than 6 Ma quiescence in this area. There are two 'satellite 197 volcanoes" in the CPR: The alkaline basalts of Pauliberg and Oberpullendorf are found at the 198 Alps-Pannonian Basin transition zone in Burgenland (Harangi, 2001b; Ali and Ntaflos, 2011) 199 and represent a single volcanic episode at about 11 Ma. Harangi et al. (1995) suggested that

they could have some compositional similarities with the basaltic dyke rocks in the Pásztori
volcano. In the southern part of the CPR, the alkaline basalt of the Lucaret-Sanovita belongs
also to a single eruptive event without any precursor magmatism. It is located close to a major
west-east trending strike-slip zone (Fig. 1B) and at the boundary of subsiding and uplifting
areas (Dombrádi et al., 2010). Remarkably, further single – potassic-ultrapotassic - volcanoes
(Bár, Uroi) were formed also along this tectonic line during the last 2.5 Ma.

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207 Temporal evolution

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Monogenetic basalt volcanic fields are typically active over millions of years. The 209 210 volcanism is characterized by intermittent eruptions gathering often in active phases, which 211 are separated by long repose periods. Recurrence rate of volcanism is particularly important in 212 evaluating the volcanic hazard (Condit and Connor, 1996; Valentine and Perry, 2007), 213 although precise dating of eruptions is difficult for basaltic rocks. In the Pannonian Basin, 214 eruption ages were determined mostly by K/Ar radiometric technique (Balogh et al., 1982; 215 1986; 1990; 1994; Borsy et al., 1986). Ar/Ar dating was performed for the basaltic rocks of 216 the Bakony-Balaton Upland volcanic fields that corroborated the reliability of the former 217 K/Ar data (Wijbrans et al., 2007). In contrast, the new Ar/Ar age data (Panaiotu et al., 2013) 218 refined significantly the former knowledge (Panaiotu et al., 2004) on the temporal evolution 219 of the Persani volcanic field. So far, the available K/Ar data base provides a solid framework 220 for the temporal evolution of the alkaline basaltic volcanism in the Pannonian basin, although 221 more careful studies are necessary to define the recurrence rates in individual volcanic fields 222 more precisely and constrain the eruption age of the youngest volcanic episodes.

In the Pannonian Basin, eruption of alkali basaltic magmas started in the westernmostmargin, where two outcrops are known, at Pauliberg and at Oberpullendorf in Burgenland

(Ali and Ntaflos 2011) and their ages were determined as 11.7 to 10.5 Ma (Balogh et al., 225 226 1994). Similar ages were published for the basalts found in drilling cores around Kecel 227 (central Pannonian basin; 8.1-10.4 Ma; Balázs and Nusszer 1987). Although these basalts are 228 strongly altered, their Late Miocene ages were corroborated by stratigraphic data. 229 The basaltic volcanism in the two largest volcanic fields (Bakony-Balaton Upland and 230 Nógrád-Gemer) lasted for more than 6 Myr. Noteworthy, eruption of basaltic magmas in both 231 areas commenced at 7.8-8.0 Ma (Balogh and Németh, 2005; Balogh et al., 2005; Chernyshev 232 et al., 2013) and continued after 2 Ma quiescence in the Bakony-Balaton Upland and after 233 about 1 Ma repose time in the Nógrád-Gemer area. The volcanism in both areas was 234 characterized by closely packed active phases separated by long repose periods (Konečný et 235 al., 1995; 1999; Wijbrans et al., 2007). Kereszturi et al. (2013) estimated the peak magma 236 output times being at 3-4 Ma in the Bakony-Balaton Upland volcanic field and 1.5-3.0 Ma in 237 the Nógrád-Gemer area. The alkaline basalt occurrences cropping out within the area of the Middle Miocene eroded andesitic Štiavnica composite volcanic complex are spatially 238 239 separated from the Nógrád-Gemer volcanic field. They show highly sporadic eruption times, 240 such as 6.6-7.8 Ma in Banska Štiavnica and Dobrá Niva and 0.1-0.5 Ma for the Putikov 241 volcano at the Hron river (Balogh et al., 1981; Konečný et al., 1999; Šimon and Halouzka, 1996; Šimon and Maglay, 2005; Chernyshev et al., 2013). The age of the Putikov vŕšok 242 243 volcano has a particular importance, since it indicates the latest alkaline basalt eruption in the 244 CPR. Balogh et al. (1981) and later, Chernyshev et al. (2013) provided 430-530 ka eruption 245 ages using K/Ar technique, whereas Šimon and Halouzka (1996) considered the stratigraphic 246 position of the lava flow and interpreted it as sitting on the Late Pleistocene (Riss) terrace of 247 the Hron river and this points to an age of 120-150 ka. This younger age was supported by 248 optically stimulated luminescence (OSL) dating of the underlying sediments that yielded an 249 age of 102 ± 11 ka. Further investigation is necessary and highly important to constrain the

exact eruption age of this basalt. The youngest eruption ages in the Nógrád-Gemer volcanic
field require also detailed investigation. In the last intense eruptive phase between 1.1-1.5 Ma,
mostly effusive activity was recognized (Konečný et al. 1995; 1999). The latest eruptions in
this area was previously thought to have occurred around 400 ka (Hodejov and Fil'akovo
maars; Konečný et al. 1995), however, new (U-Th)/He zircon dating (Hurai et al., 2013) did
not confirm this young age.

In the western margin of the CPR (Styrian basin; Ali et al., 2013), alkaline basaltic
volcanism occurred between 3.9 and 1.9 Ma (Balogh et al., 1994) after a long quiescence
period since the 14.0-17.5 Ma high-K trachyandesitic volcanic activity (Harangi et al., 1995).
The latest eruption resulted in a relatively large volume shield volcano composed of
nephelinites at Stradner Kogel.

The youngest monogenetic volcanic field of the CPR is located at the western margin of the Perşani Mts., between the Transylvanian basin and the Brasov basin (Seghedi and Szakács, 1994; Downes et al., 1995; Harangi et al., 2013). Although former K/Ar and palaeomagnetic dating (Panaiotu et al., 2004) implied two, well separated eruption phases, the new Ar/Ar dating (Panaiotu et al., 2013) revised this and indicated a fairly continuous, but very scattered volcanism from 1.22 Ma to 0.52 Ma.

Finally, a unique basalt occurrence is found at Lucaret (southern Pannonian basin), where no precursor volcanic activity took place and the age of this basalt was determined as 2.4 Ma (Downes et al., 1995).

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271 Petrologic and geochemical characteristics

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Petrologic and geochemical features of the alkaline basaltic rocks of the CPR have
been published in several papers (Embey-Isztin et al., 1993; 2001; Seghedi and Szakács,

1994; Dobosi et al., 1995; Downes et al., 1995; Harangi et al., 1995; 2013; Harangi, 2001b;
Harangi and Lenkey, 2007; Ali and Ntaflos, 2011; Tschegg et al., 2010; Jankovics et al.,
2012; 2013; Ali et al., 2013) and we completed them with new analysis, mostly from the
Nógrád-Gemer Volcanic Field (Sági et al., 2013). Representative composition of the basalts
from different locations is given in Table 1. In this section, we provide a brief summary on the
main features of the basalts.

281 The basaltic rocks are classified as alkali basalts, trachybasalts, basanite, 282 phonotephrite and nephelinite (Fig. 2). Evolved rocks types occur in the buried Pásztori 283 volcano, they are trachyandesite and trachyte (Harangi et al., 1995; Harangi, 2001b). Most of 284 the basalts are olivine-phyric, while clinopyroxene phenocryst is found in a few localities of 285 the Bakony-Balaton Upland and Kemenes Volcanic Field, but is more common in the basalts of the Styrian Volcanic Field and in the Nógrád-Gemer Volcanic Field. The olivine 286 287 phenocrysts with Fo>84 mol% often contain spinel inclusions and these two minerals can be 288 regarded as liquidus phases. Both minerals have a wide compositional range (Embey-Isztin et 289 al., 2007; Ali and Ntaflos, 2011; Jankovics et al., 2012; 2013; Ali et al., 2013; Harangi et al., 290 2013). The most magnesian olivines have typically high Mn content compared with the trend 291 defined by Herzberg (2011) for olivines crystallized from peridotite-derived magmas. They are characterized by low Fe/Mn ratio (Fe/Mn=40-75); this feature along with the relatively 292 293 low Ni content (<2500 ppm) is not consistent with derivation of the primary magmas from 294 pyroxenite source as defined by Sobolev et al. (2007) and Herzberg (2011). Clinopyroxene 295 phenocrysts could have crystallized mostly at the crust mantle boundary zone as shown by the 296 geobarometric calculations. Their complex zoning often found in basalts from the Nógrád-297 Gemer and the Bakony-Balaton Upland indicates mixing of magma batches at depth (Dobosi, 298 1989; Dobosi and Fodor, 1992). Mantle-derived ultramafic and lower crustal granulite

xenoliths are common in basalts of each volcanic field (Embey-Isztin et al., 2001; Szabó et
al., 2004).

301 The CPR basalts are almost exclusively silica-undersaturated and have mg-numbers 302 ranging from 0.5 to 0.8 (MgO=5-15 wt%). The most magnesian basalts contain, however, 303 significant amount of xenocrysts (e.g., Bondoró and Füzes at Bakony-Balaton Upland), which 304 strongly modify the bulk rock composition (Jankovics et al., 2012; 2013). Majority of the 305 basalts have an mg-number >0.62 suggesting only a minor to moderate effect of fractional 306 crystallization. Moderately fractionated basalts are found mostly in Styria and in the Nógrád-307 Gemer areas. For this study, we selected those samples that have mg-number between 0.64 308 and 0.72 (MgO=8-13 wt%) and therefore their compositions are close to that of the primitive 309 magmas. In these samples, variation of major elements indicates the controlling factor of 310 various amounts of olivine fractionation. However, characteristic differences can be 311 recognized between the most primitive basalts. The Pauliberg basanite has a distinct 312 composition with low Al₂O₃ and high TiO₂ and FeO contents (Fig. 3). Experimental studies 313 suggest that such composition could be obtained by low-degree melting of carbonated, silica-314 deficient eclogite (Dasgupta et al., 2006; Kogiso and Hirschmann, 2006). The basalts from Oberpullendorf (Burgenland) and Putikov vŕšok (Štiavnica) have a transitional major element 315 316 composition between the Pauliberg basanite and the rest of the CPR basalts.

The trace element composition of basalts resembles those commonly found in intraplate magmas and is similar to the Neogene to Quaternary alkaline mafic rocks in Western and Central Europe (Wilson and Downes, 1991; Lustrino and Wilson, 2007). The FTS elements, such as Ni, Cr and Sc have relatively high abundances (Ni=100-300 ppm; Cr=100-500 ppm; Sc=10-30 ppm; Fig. 3) consistent with the slight olivine±spinel-dominated fractional crystallization. Based on the primitive mantle (McDonough and Sun, 1995) normalized trace element patterns, Harangi and Lenkey (2007) defined two main

compositional groups: the first one is characterized by stronger enrichment in incompatible
trace elements and a typical negative potassium anomaly, whereas the second one has
smoother trace element patterns and no negative potassium anomaly (Fig. 4). The strongly
incompatible trace elements have a wide compositional range (La=25-90 ppm, Nb=30-130
ppm, Th=3-13 ppm; Fig. 3) and show a positive correlation with an intercept of zero. Same
covariance with Ba can be also observed although the basalts from Pauliberg and Styria show
a distinct trend.

331 Sr-Nd-Pb isotope ratios of the basalts were published by Embey-Isztin et al. (1993), 332 Dobosi et al. (1995), Downes et al. (1995) and Harangi et al. (1995). All of the basalts fall into the depleted field in Sr-Nd isotope space, similarly to majority of the Neogene to 333 334 Quaternary alkaline mafic rocks in Western and Central Europe (Wilson and Downes, 1991; Wilson and Patterson, 2001). The most depleted end-member of this variation trend is a 335 336 mixture of HIMU and DMM. In the Pb-Pb isotopic diagrams, they show a relatively large variation, many of the samples have relatively high ²⁰⁷Pb/²⁰⁴Pb ratio (>15.6) causing a vertical 337 shift from the NHRL. The highest 206 Pb/ 204 Pb ratio (19.6-19.7) is shown by the Pauliberg 338 339 basanite.

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Most of the alkaline basalts of the CPR show chemical composition (MgO>8 wt%; Ni=130-350 ppm; Fig. 3) that could be close to that of the primitive magmas. Using only these samples, the condition of the melting process as well as the nature of the source region can be constrained that could contribute to our understanding on the reason of mantle melting beneath the area. Trace element ratios such as Nb/Th (5–15) and Nb/La (1–2) of the basalts show values that are higher than those observed in continental crustal rocks (Rudnick and

349 Fountain, 1995) and are typical of mantle-derived melts, mostly for those derived from the 350 asthenosphere. Therefore, we can conclude that crustal contamination did not significantly 351 modify the trace element composition of the CPR mafic magmas. Most of the selected basalts 352 are olivine-phyric and the olivine phenocrysts often contain spinel inclusions indicating that 353 these mineral phases could have been liquidus minerals as have been suggested for the usual 354 evolution of alkaline silica-undersaturated mafic magmas (Roeder et al., 2006). High pressure 355 clinopyroxene crystallization could modify the magma composition as well as the 356 composition of coexisting evolving phases and this should be considered even in the absence 357 of clinopyroxene phenocrysts in the rocks (Smith et al., 2008). This would decrease the CaO 358 content of the magma and possibly also leads to lowering the CaO content of the olivine phenocrysts. Nevertheless, our geothermobarometric calculations on selected clinopyroxenes 359 in some basalts (Harangi et al., 2013) suggest that they crystallized later then the olivines. The 360 361 relatively high Sc content of the basalts (Sc>15 ppm except for the samples from Lucaret) and 362 the uniform Cr-number of the spinels along with various Fo content of coexisting olivines 363 also imply that no significant early-stage clinopyroxene fractionation occurred that would 364 significantly modify the magma composition. Taking all of this into account, the composition of the alkaline basalts with MgO>8 wt% can be used to infer the melting condition and the 365 366 nature of the source rock.

In this discussion, we attempt to constrain the origin of the magmas using the following principle: we consider the simplest model over the complex models and take a complex model only when no simple model works (Niu, 2005). Thus, we simplified the mantle source material considering to be dominantly peridotitic in lithology, knowing, however, that involvement of more fusible material, such as pyroxenite or eclogite could be possible. There are some inferences for the existence of such material in the mantle source

373 region, but we think that further, careful investigation is necessary to test their influence on374 the melting models and the composition of the primary magmas.

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376 Asthenospheric vs. lithospheric mantle source region

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378 Origin of the continental intraplate basaltic rocks is often debated. Alkaline silica-379 undersaturated mafic magmas are considered being generated by low-degree of melting either 380 in the upwelling asthenosphere (Bradshaw et al., 1993; Wang et al., 2002; Niu et al., 2011; 381 Gazel et al., 2012) or by melting of metasomatic, usually amphibole-rich veins in the 382 lithosphere (Fitton et al., 1991; Ormerod et al., 1991; DePaolo and Daley, 2000; Beccaluva et 383 al., 2007; Valentine and Perry, 2007; Bianchini et al., 2008; Pilet et al., 2008; Mayer et al., 384 2013; 2014). Variation of trace element and isotopic compositions of the Neogene to 385 Quaternary basaltic rocks in Europe led Wilson and Downes (1991; 2006) to conclude that 386 both mantle regions could have contributed to the magma generation. This was reinforced by 387 the more specific studies in various parts of volcanic areas in Europe (Haase et al., 2004; 388 Haase and Renno, 2008; Jung et al., 2012), including the Pannonian basin (Embey-Isztin and 389 Dobosi, 1995; Harangi, 2001a; Seghedi et al., 2004) and was suggested also for the origin of 390 the basaltic magmas in the western USA (Fitton et al., 1988; 1991), in the Auckland volcanic 391 field, New Zealand (McGee et al., 2013) and western Syria (Ma et al., 2011). However, an 392 important question is why the colder lithospheric mantle undergoes melting. For this, the 393 presence of fusible material (metasomatic veins) and either significant thinning of the 394 lithosphere or a thermal perturbation (heating) are necessary. The fusible material is usually 395 considered to be amphibole-bearing veins, which originated from low-volume asthenospheric 396 melts and are frozen in the lower lithosphere (McKenzie, 1989; Niu, 2008; Pilet et al. 2008; 397 Humphreys and Niu, 2009). The presence of amphibole in the source region is usually

inferred based on the negative K-anomaly in the primitive mantle normalized trace element
patterns (Class and Goldstein, 1997; Späth et al., 2001; Jung et al., 2005; Haase and Renno,
2008). Amphibole is stable up to about 1100°C temperature and up to about 2.5-3.0 GPa (ca.
100 km depth) and is regarded to be confined to the lithospheric mantle (Wallace and Green,
1991).

403 The negative K-anomaly as expressed by the K/K* ratio (calculated as $K_N/\sqrt{(Nb_N*La_N)}$, where K_N , Nb_N and La_N are primitive mantle normalized values) is a 404 characteristic feature of many of the CPR basalts and shows a wide continuous range 405 406 (K/K*=0.3-1.2; Fig. 5). It has a negative correlation with the La/Y and (Sm/Yb)_N ratios, 407 which are the indicators of the degree and the relative depth of melting. This could suggest 408 that a K-bearing phase, presumably amphibole was present in the source of all of the CPR 409 magmas and remained in the residual mantle during low degree melting and entered into the 410 melt in the case of larger melting. In that situation, the source of the magmas should be 411 entirely the lithospheric mantle. Melting of the lower lithosphere in the post-rift phase of the 412 Pannonian basin cannot be explained by decompression-driven process, but only by thermal 413 perturbation, i.e. via heat flux from the underlying asthenospheric mantle. Taking this view, 414 Seghedi et al. (2004b) postulated localized mantle plumes (plume fingers) beneath the basalt 415 volcanic fields of the CPR. However, there is lack of physical evidences for the generation of 416 localized hot mantle upwelling scattered beneath the Carpathian-Pannonian region and there 417 is a lack of explanation why this does not occur beneath the central Pannonian basin, where 418 there is the thinnest lithosphere. Furthermore, the lowest K/K* value is shown by the basalts 419 having the highest $(Sm/Yb)_N$ ratios and therefore reflect garnet in the source region. This 420 would mean melting at depth greater than 80 km, which is already the asthenospheric mantle 421 beneath most part of the region. Nevertheless, theoretically, it is possible that the high 422 (Sm/Yb)_N ratio is inherited in the metasomatic vein (Pilet et al., 2008). Although Pilet et al.

(2008) claimed that preferential melting of this material would yield magma composition akin 423 424 to alkaline mafic magmas, others suggest that in this case potassic magmas are generated 425 (McKenzie, 1989; Wilson and Downes, 1991; Cebriá and López-Ruiz, 1995). Furthermore, 426 the K/K* ratio in the CPR basalts has no correlation with the La/Nb ratio that is used as an 427 indicator of the lithospheric mantle (Fitton et al., 1991). In addition, the La/Nb ratio of the 428 CPR basalts is less than 1 which is characteristic of asthenosphere-derived magmas. Finally, 429 many alkaline OIB, such as St. Helena-type basalts (Chaffey et al., 1989; Kawabata et al., 430 2011) show negative K-anomalies, typically those formed by low-degree melting. Although, 431 melting of metasomatized lithosphere cannot be excluded as a potential origin of such basalts 432 (Pilet et al., 2005; 2008), most authors consider them as derived from upwelling hot mantle 433 beneath the oceanic lithosphere. Thus, this compositional feature could be rather general both 434 in oceanic and continental realms and could be an inheritance in the mantle source regions, 435 possibly representing the most fusible part (eclogite or pyroxenite with negative K-anomaly 436 and HIMU/FOZO characteristics; Sobolev et al., 2005; 2007; Fekiacova et al., 2007; 437 Dasgupta et al., 2010; Timm et al., 2009) of the melted mantle material. In this context, it is 438 noteworthy, that a correlation between the K/K* value and the radiogenic isotope ratios (⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd; Fig. 5) is observed within the volcanic fields and that can be more 439 440 readily explained by derivation of the magmas with compositionally distinct mantle sources. 441 Lower degrees of melting would produce magmas from the more fusible parts of the source region that are characterized by low ⁸⁷Sr/⁸⁶Sr and high ¹⁴³Nd/¹⁴⁴Nd isotope ratios similar to 442 443 HIMU/FOZO (Stracke et al., 2005) and negative K-anomaly (low K/K*), whereas higher 444 degree of melting would lead to mixing with melts coming from compositionally distinct 445 mantle domains (Harangi and Lenkey, 2007).

In summary, derivation of the basaltic magmas solely from the lithospheric mantledoes not seem to be a plausible model in the CPR. In turn, we suggest that composition of the

448 CPR basalts is consistent with melting predominantly or even entirely of the asthenospheric 449 mantle. Since there are no convincing evidences for the existence of mantle plume or plume 450 fingers beneath this region, nor progressive lithospheric extension, we think that it is 451 physically difficult to generate significant amount of alkaline sodic magmas in the lithosphere 452 during the post-rift and tectonic inversion stages unless considering local detachment and 453 sinking of lithospheric fragments into the asthenosphere (Timm et al., 2010). At this stage, we 454 think that composition of the CPR basaltic rocks reflects mostly the nature of the source 455 region of the asthenospheric mantle and the heterogeneity within it.

456

457 Constrains on melting columns

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459 Considering that the basaltic magmas in the CPR could derive predominantly from the 460 asthenospheric mantle, an upward mantle flow is required to initiate melting. As the solid 461 mantle material moves upwards adiabatically, melting starts when the ambient temperature 462 intersects the solidus (initial melting pressure; Langmuir et al., 1992). This depth depends on 463 the mantle potential temperature as well as the nature of the mantle material (Langmuir and 464 Forsyth, 2007), but for alkaline mafic magmas it is usually within the garnet stability field, i.e. 465 deeper than 80 km (Niu et al., 2011). Melt production continues during further upwelling and 466 as a result, the initial magma composition will change due to the increasing extent of melting 467 and according to the changing mineral assemblage of the melted mantle rocks (from garnet to 468 spinel). Thus, the early-stage melt shows strong garnet-signature (high Ce/Y, Nb/Y and 469 [Sm/Yb]_N ratios; Ellam, 1992; Salters and Hart; 1989; Putirka, 1999; Niu et al., 2011), but it 470 becomes progressively diluted in the strongly incompatible elements and looses their garnet 471 fingerprint. The extent of the melting column (defined by the initial and final melting 472 pressure, denoted as po and pf, respectively) depends primarily on the thickness of the

lithosphere (Ellam, 1992; Langmuir et al., 1992; Haase, 1996; Putirka, 1999; Niu et al., 2011).
Melting beneath thick lithosphere stops at greater depth and therefore produces smaller
amount of magma, which are enriched in strongly incompatible trace elements and have a
pronounced garnet-signature, whereas in areas where the lithosphere is thinner, melting could
continue even in the spinel-stability field and produces more melt with different incompatible
element abundances. Thus, the p_f can be used as a proxy for the thickness of the lithosphere,
i.e. defining the depth of the lithosphere-asthenosphere boundary (LAB).

480 The pressure of the melting can be constrained by various methods. Major element 481 composition of the primitive magmas is influenced by the pressure of melting and the extent 482 of melting (Takahashi and Kushiro, 1983; Langmuir et al., 1992; Hirose and Kushiro, 1993; Baker and Stolper, 1994; Walter, 1998; Niu et al., 2011). MgO and FeO^t increase, SiO₂ 483 484 decreases with increasing pressure of melting, whereas the Na₂O content of the primary 485 magma depends mostly on the degree of melting. The CPR basalts have a relatively wide 486 range of FeO content (FeO^t=8-12 wt%; Fig. 3) showing an inverse correlation with SiO₂ 487 indicating that the primary magmas could have formed at various pressures. In order to 488 quantify the pressure of melting, i.e. the depth of the magma generation first, correction of 489 fractional crystallization has to be performed. This can be achieved by adding as much olivine 490 to the bulk rock composition until the mg-number is around 0.72, consistent with a primitive 491 melt composition that is in equilibrium with a melted peridotite with olivine having Fo-492 content around 0.9 mol%. For the selected mafic rocks of the CPR, correction of 5-15% 493 olivine fractionation was necessary.

Langmuir et al. (1992) developed a model where the FeO^t and Na₂O contents of the primary magmas are used to constrain the intitial and final pressure of melting. Although this was constructed for mid-ocean ridge basalts, Wang et al. (2002) demonstrated that it provides reliable results for most melting within the depth interval of 50 and 110 km. The melting

498 model lines were compared with experimentally determined peridotite melts and a good 499 reproduction was obtained between 1.0 and 4.0 GPa. Furthermore, Wang et al. (2002) 500 emphasized that for magmas having Na₂O>3 wt% that is relevant for the CPR basalts, the 501 model is not dependant upon the melt pooling mechanism. We used this model calculation, 502 considering accumulated fractional melting process to determine the melting columns under 503 the alkaline basalt volcanic field of the CPR. In the Fig. 6A, each curve represents a different 504 initial melting pressure, i.e. where the ambient mantle temperature intersects the solidus. The 505 total iron content of the magma reflects the condition of the onset of melting, i.e. the initial 506 pressure that has a close relationship also with the mantle temperature. As melting proceeds, 507 the volume of magma increases and as a consequence the concentration of incompatible 508 elements as well as Na₂O decreases, while the total iron concentration does not change 509 significantly. Therefore, the final values of the FeO^t and the Na₂O of the accumulated melt vield information on the initial melting pressure and the cessation of the melting, i.e. the final 510 511 melting pressure. These pressure values are then converted to depth, assuming average densities of crust and mantle, 2.85 g/cm³ and 3.25 g/cm³, respectively as suggested by Wang 512 513 et al. (2002). The crustal thickness was taken as 32 km beneath Burgenland (western CPR) 514 and 35 km beneath Perşani (eastern CPR), whereas 28 km was used for the rest of the areas. 515 Note that the pressure/depth estimate has an about 10-20 km uncertainty due to various effects such as the estimate of the primary composition, mantle fertility and the neglecting Fe^{3+} in the 516 517 iron content of the primitive melts (Wang et al., 2002).

We obtained p_0 and p_f values for 13 basalt localities that cover all the monogenetic volcanic fields of the CPR (Fig. 6A). Within them, only one basalt group from the Nógrád-Gemer Volcanic Field shows a melting column with a depth interval that seems to be unusually shallow (2.0 to 1.55 GPa, i.e. melting between 66 and 52 km). In all other cases the initial and final melting pressure values are greater then 2 GPa; thus, considering the

523 lithospheric thickness beneath the Pannonoan basin, they indicate magma generation in the 524 sublithospheric mantle. Generation of the primary magmas started predominantly at around 3 525 GPa, i.e. at 100 km depth. This corresponds with a depth of the spinel-garnet transitional field 526 (Robinson and Wood, 1998; Klemme and O'Neill, 2000; Klemme, 2004). At such depth, dry 527 peridotite could not melt up to a mantle potential temperature of 1450°C, therefore presence 528 of water or pyroxenite/eclogite with lower solidus is expected in the source region that may 529 lead to lower the calculated pressure (Wang et al., 2002). Nevertheless, the final depth of 530 melting columns of the CPR basalts is between 2.1 and 2.5 GPa (70-83 km depth), what is in 531 good agreement with the geophysically determined thickness of the lithosphere (Horváth et 532 al., 2006) beneath the volcanic fields (Fig. 7). In addition to the usual melting depth of the 533 CPR basalts, the alkaline basalts from Burgenland (Pauliberg and Oberpullendorf) fall into the 534 3.4 and 3.8 GPa melting lines (Fig. 6A) that corresponds with an onset of melting at greater 535 depth (110-125 km). This fits, however, well with the deep structure beneath these basaltic 536 volcanoes, i.e. they are located in an area characterized by thick lithosphere (ca. 110 km 537 present day thickness; Fig. 7). The relatively shallow melting column beneath the Persani 538 volcanic field (86-70 km depth interval) seems to be remarkable, since formerly Dérerova et 539 al. (2006) and Horváth et al. (2006) suggested a thick lithosphere here (about 140 km). 540 However, Harangi et al. (2013) pointed out that the Persani basaltic magmas were originated 541 by melting of asthenospheric mantle and therefore the LAB should be much shallower, i.e. not 542 deeper than 60 km. This is consistent with the geophysical result of Martin et al. (2006) and 543 implies that a narrow rupture could have developed in the lower lithosphere beneath the 544 Perşani area, probably as a result of the dragging effect of the descending dense lithospheric 545 slab beneath the Vrancea zone located about 100 km from the volcanic field (Harangi et al., 546 2013). In two cases (Putikov vŕšok and Lucaret), we got a melting column fairly deep in 547 comparison with the lithospheric thickness. Partial melting could start here deeper than 100

km and terminated shortly after 10-15 km mantle upwelling. A possible explanation for this
would be the exhaustion of the fusible material such as hydrous peridotite or garnetpyroxenite/eclogite.

551 An independent test of the result provided by the Langmuir et al. (1992) model can be 552 achieved with comparison of the primary magma compositions of the CPR basalts with 553 experimental data sets (Sakuyama et al., 2009) and using other geobarometric calculations (Scarrow and Cox, 1995; Lee et al., 2009). Sakuyama et al. (2009) compiled results of various 554 555 melting experiments and suggested that SiO₂ content of the primary melt is the most sensitive 556 parameter on melting pressure in the range of 1.0-3.0 GPa and it does not depend significantly 557 on the degree of melting, the initial bulk composition and modal mineralogy. Water content of 558 the source peridotite could slightly increase the SiO₂ content of the melt at constant pressure 559 (Hirose and Kawanoto, 1995) that may lead to slight underestimation of the pressure if dry 560 peridotite melting is considered. Plotting the primary magma compositions of the CPR basalts 561 on the ol'-ne'-qtz' triangular diagram along with the isobaric melting lines as determined by 562 Sakuyama et al. (2009) based on experimental data sets, 2.5-3.5 GPa melting pressures are 563 obtained (Fig. 6B) that are in good agreement with the constrains on the melting columns 564 described above, particularly with the initial melting pressure. The only exception is the 565 basalts from Lucaret for which we got melting pressure of 2.5 GPa, significantly lower than 566 from the Langmuir et al. (1992) model (3.20-2.75 GPa). Scarrow and Cox (1995) developed a 567 simple equation for the melting pressure estimate based on the SiO₂ content of primary rocks. 568 It gives slightly lower pressure values than the ol'-ne'-qtz' diagram, whereas the corrected 569 SiO₂-geobaromether by Wang et al. (2002) appears to strongly overestimate the pressure 570 particularly at higher pressure (>2.5 GPa). Finally, we applied also the peridotite-melt 571 barometer of Lee et al. (2009), which gives consistent result with the SiO₂-barometer of 572 Scarrow and Cox (1995), but also with the final pressure estimate of the Langmuir et al.

(1992) model (Fig. 8). There is also a good correlation between the pressure obtained by the
Lee et al. (2009) barometry and the qtz'/(ol'-ne'-qtz') ratio, which was suggested as a
sensitive parameter of depth of melting (Sakuyama et al., 2009). The only exception is the
basalts from Pauliberg that form a parallel trend at slightly higher pressure range, probably
due to the difference of the source lithology.

578 In summary, we can conclude that melting of the upwelling asthenospheric mantle 579 material could start at about 100 km depth beneath most of the volcanic fields of the CPR and 580 the final melting pressure could have been controlled by the thickness of the lithosphere, i.e. 581 the lithosphere-asthenosphere boundary. Partial melting occurred at greater depth beneath 582 Pauliberg, which is underlain by thicker lithosphere, whereas beneath the Persani area melting 583 could have taken place at shallower depth due to a local rupture of the lower lithosphere. The 584 youngest basalt beneath Putikov vŕšok generated at greater depth (100-120 km), whereas the 585 2.4 Ma Lucaret basalt could have been derived also from a relatively shallow depth in spite of 586 the higher FeO content of the basalts.

587

588 Conditions of melt generation

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We proposed that partial melting of the upwelling asthenospheric mantle material 590 591 could lead to the generation of basaltic magmas, which fed the volcanic eruptions in the 592 monogenetic volcanic fields of the CPR. A further constrain on the melt generation process 593 can be obtained from trace element petrogenetic modelling (Fig. 9). We used different initial 594 compositions (depleted MORB mantle and slightly enriched primitive mantle; Salters and 595 Stracke, 2004; McDonough and Sun, 1995) and different mineralogical assemblage (garnet-596 lherzolite, spinel lherzolite) of the source material and different melting models (batch and 597 fractional melting; Shaw, 1970; critical or dynamic melting; Sobolev and Shimizu, 1992; Zou,

598 1998; Shaw, 2000) with the mineral-melt distribution coefficients proposed by Kostopoulos 599 and James (1992). Trace element ratios of Zr/Nb and La/Y were chosen, because their values 600 are not dependant on early-stage crystal fractionation and are sensitive (La/Y) on the presence 601 of spinel and garnet in the mantle source. As we discussed earlier, it is unlikely that 602 amphibole was also present in the source region of the CPR basalts and therefore, it was not 603 considered in the modelling calculations. The result of the petrogenetic modelling is shown in 604 Figure 9.

605 The primary magmas of the CPR basalts were formed from trace element enriched 606 mantle source (1.5-4.0-times of the primitive mantle composition; McDonough and Sun, 607 1995) within the spinel-garnet transitional zone (Fig. 9.). The degree of melting is somewhat 608 model dependant, but it is between 1.5% and 5%. The best fit of the CPR basalts with the 609 melting model lines is given assuming continuous dynamic melting or decompressional 610 critical melting (Sobolev and Shimizu, 1992; Zou, 1998; Shaw, 2000) process. This model is 611 consistent with the melting process described in the previous section, i.e. commencing of 612 melting in a mantle source, where garnet prevails over spinel and then continues in the spinel 613 stability field. In this case, 2-3% melting is required to obtain the composition of the primary 614 basaltic magmas of the CPR. Noteworthy, the Lucaret basalts fall on the spinel-peridotite 615 melting line implying shallow melting and suggest that the FeO content of the bulk rock 616 overestimates the melting pressure and the SiO₂-based geobarometries provide more reliable 617 result. On the other hand, the basalts from the western Pannonian Basin (Pauliberg, 618 Oberpullendorf and Klöch from the Styrian basin) fall on the garnet-peridotite melting line 619 consistent with their derivation from a deeper source.

620 The depth of melting for the CPR basalts as shown by various models is between 60 621 and 120 km. The associating melting temperature can be determined using the peridotite-melt 622 thermometer by Lee et al. (2009; using molar $Fe^{3+}/Fe^{tot}=0.13$) and the PRIMELT2 software of

Herzberg and Asimow (2008; using Fe₂O₃=TiO₂). They resulted in fairly similar temperature 623 values up to 1470°C (Fig. 7), where the Lee et al. (2009) calculation tends to give much 624 625 higher (>100°C) temperature than the PRIMELT2 (mostly for the Pauliberg basalts). The 626 lowest mantle potential temperature is shown by the Persani basalts (1310-1380°C) that 627 records indeed shallow melting as indicated by the geobarometric calculations, whereas in the 628 other volcanic fields, we got melting temperature in a range between 1350 and 1450°C. The 629 highest potential temperature values were obtained for the basalts at Pauliberg (>1440°C) and 630 at Putikov vŕšok (1420-1450°C).

631 Taken together the obtained pressure and temperature values, the Figure 10 shows that 632 the melting columns of the basaltic magmas of the CPR are just below the dry peridotite 633 solidus. This requires either some water or a more fusible material, such as pyroxenite or 634 eclogite in the mantle source. The relatively low CaO content (<10 wt%) of the calculated 635 primary basaltic magmas as well as the low Ca also in the olivine phenocrysts (Harangi et al., 636 2013) are features that can be explained by pyroxenite melting (Herzberg, 2011). However, 637 the olivine in the CPR basalts (Embey-Isztin and Dobosi, 2007; Jankovics et al., 2012; 2013; 638 Harangi et al., 2013) has relatively low Ni and high Mn concentrations that is just the opposite 639 of what required in the case of pyroxenite melting (Sobolev et al., 2007). These features could 640 be explained, however, also by the lid-effect, i.e. by variations in lithospheric thickness (Niu 641 et al., 2011). Magmas generated under a thin lithosphere could crystallize olivines with 642 relatively low Ni and high Mn and corresponding low Ni/Mg, low Ni/(Mg/Fe) and high 643 Mn/Fe and high Ca/Fe ratios. Although the Ca deficiency of the olivine and the host rocks of 644 the CPR is still unresolved, it appears that the basaltic magmas beneath the CPR could have 645 originated dominantly from a peridotitic mantle source. In this case, however, some water 646 (50-200 ppm; Hirschmann et al., 2009) should be present in the source region. The origin of 647 this volatile content can be explained either by the supposed intense Palaeogene-Miocene

subduction around the Carpathians (Csontos et al., 1992) or dewatering of the accumulated 648 649 subducted material in the transitional zone (Hetényi et al., 2009). An important point coming 650 from the Figure 10. is that the upper mantle beneath the thin continental lithosphere of the 651 Pannonian basin has a temperature close to the peridotite solidus and the mantle could be 652 partially hydrous that makes it potentially still fusible. Eruptions of basaltic magmas in the 653 last 2 Ma could confirm the potentially fusible state of the mantle. The low seismic velocity 654 anomaly down to the transitional zone beneath the northern part of the Pannonian basin 655 (Dando et al., 2011) could provide a further supporting evidence of this.

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657 Mantle source characteristics

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659 Chromian spinels enclosed by olivine phenocrysts could provide important 660 information about the mantle sources of the CPR basalts. They occur in each volcanic field 661 and therefore it can be used to gain insights into the nature of the sublithospheric mantle on a 662 regional scale. Chromian spinel is often the liquidus phase along with olivine in basaltic melts 663 (Fisk and Bence, 1980; Luhr and Carmichael, 1985), therefore it reflects the primary 664 characteristics of the basaltic melt. It can be used for deciphering the original magmatic 665 characteristics even of heavily altered basalts (Allan, 1992). Composition of the liquidus 666 spinels is considered to be close to that of the spinels in the residual mantle source region, 667 although other controlling factors such as magma differentiation, crystallization pressure, 668 redox condition and the degree of partial melting also influence the composition of spinels 669 (Hill and Roeder, 1974; Sigurdsson and Schilling, 1976; Dick and Bullen, 1984; Allan et al., 670 1988; Roeder and Reynolds, 1991; Arai, 1994; Clynne and Borg, 1997; Kamenetsky et al., 671 2001; Roeder et al., 2003).

The spinel inclusions found in the CPR basalts show a wide compositional range (cr-672 673 number=0.15-0.68, mg-number=0.2-0.8). They are mostly chromian spinels thought to be 674 formed as early-stage mineral phases, however, as soon as they became out of the narrow 675 temperature interval where they could appear as liquidus crystals (Sigurdsson and Schilling, 676 1976; Fisk and Bence, 1980; Luhr and Carmichael, 1985), they react with the differentiated 677 melt and the spinel component is rapidly replaced by magnetite and ulvöspinel components increasing their FeO and Fe₂O₃, and TiO₂ contents (Luhr and Carmichael, 1985; Clynne and 678 679 Borg, 1997; Arai, 1992; Roeder et al., 2003). Therefore, we selected spinels enclosed by the 680 most magnesian (Fo>84 mol%) olivines and then, we filtered out the analyzed spinel compositions using their TiO₂ and Fe³⁺-number in order to focus on the most primitive spinel 681 682 population. Instead of using a definite limit of these parameters (e.g., Clynne and Borg, 1987), we plot all the data on the $Cr-Fe^{3+}$ -Al triangular diagram and compared them with the natural 683 684 basalt-hosted spinel compositional range (Barnes and Roeder, 2001). The majority of the 685 MORB spinels show a wide variation in the Cr-Al trend, whereas many spinels from OIB 686 have a slight enrichment in Ti and Fe. We selected the spinel population thought to represent liquidus mineral phases primarily based on the overlap with the 50th percentiles of the OIB 687 688 and MORB spinels and then we considered the possible differentiation trends. The whole data set and the selected "liquidus-spinel" compositions are shown in Fig. 11. The primary spinels 689 form four coherent groups in the Cr-Fe³⁺-Al diagram that could represent different mantle 690 691 domains beneath the CPR. Three of them appear in all major volcanic fields of the CPR, 692 whereas the forth one belongs solely to the Pauliberg location. In general, spinel composition 693 closely mimics the host rock composition (as shown by the correlation between the cr-number of spinels and the $Fe_2O_3^{t}$, Al_2O_3 and TiO_2 contents of the host rock). The basanites of 694 695 Pauliberg have a specific composition within the Miocene-Quaternary alkaline basalts of the 696 CPR (Embey-Isztin and Dobosi, 1995; Ali and Ntaflos, 2011) showing depletion in Al and

enrichment in Fe and Ti compared with the others. Spinels in these rocks are also depleted in
Al (thus, they have relatively high cr-number) and slightly enriched in Fe and Ti, plotting
outside from the MORB and OIB spinel fields. Nevertheless, considering the specific
composition of their host rock, they could also belong to the "liquidus-spinel" population and
reflect a particular mantle source region of the Pauliberg magmas. Tentatively, this might be a
sign of involvement of eclogite during melting, a hypothesis that requires further studies.

703 Compositional characters of the spinels are controlled by many factors as listed before. 704 Influence of the magma differentiation could be filtered out by limiting the data set to the 705 most primitive spinels. The pressure of crystallization affects predominantly the Al content of 706 the chromian spinels (Dick and Bullen, 1984) showing positive correlation between them. The 707 large compositional variation of the spinel compositions in the CPR basalts, however, could 708 be unlikely related to the variable crystallization pressures, since the lowest Al content in 709 spinels was recorded in those basalts which were interpreted as derived from the deepest 710 source (Pauliberg, Putikov, Lucaret). Spinel composition is dependant also on the degree of 711 partial melting as demonstrated by Arai (1994) and Baker and Stolper (1994). The cr-number 712 of spinel in the mantle restite rock increases with increasing degree of melting. Spinels with 713 0.10-0.15 cr-number are considered to represent fertile mantle regions, whereas spinels with 714 cr-number of about 0.6 could indicate more than 25% degree of melting (Baker and Stolper 715 1994). However, no correlation can be recognized between the variation of spinel cr-number 716 and the incompatible trace element ratios of the CPR basalts that are sensitive on the 717 progressive melting process (e.g., La/Y, [Sm/Yb]_N). Furthermore, our petrogenetic model 718 calculation suggests that the mafic magmas of CPR could have originated by 2-3% of melting 719 and therefore, the large compositional variation of spinels cannot be attributed to various 720 degrees os partial melting. Taking all of these into account, we can conclude that the spinels 721 in the CPR basalts reflect the nature of the mantle source region, more precisely the character

of the residual mantle rocks and indicates heterogeneous, variably depleted sublithosphericmantle beneath the Pannonian basin.

724 Spinel compositions suggest that at least three main compositional domains (and a 725 specific one beneath Pauliberg) could exist in the mantle source region and they are well 726 distributed beneath each volcanic field. Although the composition of the liquidus spinels is 727 close to the spinels in the restitic mantle rocks rather than in the intitial source material, the 728 small degree of melting characterizing the CPR basalts could not significantly modify the 729 compositional feature of the initial spinels as also emphasized by Clynne and Borg (1997) for 730 the Lassen basalts. Therefore, the composition of the restite spinels could be close to that of 731 the initial ones. Noteworthy, the youngest (<600 ka) basalts could have derived either from 732 the most fertile mantle source domain (Perşani basalts; spinel cr-number=0.23-0.32; Harangi et al., 2013) or from a mantle domain akin to enriched OIB (Putikov vŕšok; spinel cr-733 734 number=0.55-0.61; TiO₂=2.0-3.5 wt%; Sági et al., in prep). This might mean that the 735 sublithospheric mantle beneath CPR could be potentially still capable to produce basaltic 736 magmas, i.e. contains fusible parts of material.

737

738 Geodynamic relationships

739

The origin of the alkaline basaltic volcanism in the CPR seems to be enigmatic since it occurred after the main phase of the lithospheric extension of the Pannonian basin (20-15 Ma; Horváth et al., 2006), during the stage of post-rift thermal subsidence and tectonic inversion. Melt generation has taken place intermittently since 11 Ma with a peak period between 5 and 1 Ma. In such situations, hot mantle upwelling beneath the thinned lithospheric realm seems to be a viable mechanism to produce low-volume basaltic magma (Embey-Isztin et al., 2001; Harangi, 2001b). Composition of the Neogene alkaline mafic rocks in Europe shares many

747 similarities, particularly in the radiogenic isotope ratios. Taking this into account and 748 correlating it with seismic anomalies in the sublithospheric mantle beneath Europe, Hoernle et 749 al. (1995) and Granet et al. (1995) proposed a common mantle reservoir of the mafic magmas 750 (called European Asthenospheric Reservoir; EAR by Cebriá and Wilson, 1995) that fed the 751 basaltic volcanism via localized active mantle plumes ("plume finger") in different parts of 752 Europe, such as in the Massif Central, the Rhenish area, the Eger graben, Bohemia and also in 753 the Pannonian basin (Granet et al., 1995; Wilson and Patterson, 2001). The plume finger 754 model was put forward for the Pannonian basin by Embey-Isztin et al. (2001), Seghedi et al. 755 (2004b) and Seghedi and Downes (2011). Seghedi et al. (2004b) distinshuished at least three 756 plume fingers beneath the CPR, primarily based on the HIMU-like isotopic character of the 757 basalts. However, we have to emphasize that the HIMU-like isotopic ratios have solely a 758 descriptive message and this does not necessarily mean plume-origin. Nevertheless, the 759 plume-related basaltic magmatism in the CPR was proposed also by other authors based on 760 geological, geochemical and geophysical considerations (Goes et al., 1999; Buikin et al., 761 2005; Burov and Cloetingh, 2009). In contrast, Harangi and Lenkey (2007) argued against a 762 mantle plume beneath the Pannonian basin.

763 There are a couple of observations that are not consistent with the plume model. 764 Beneath the Pannonian basin, in the depth range of 400-650 km, a positive seismic velocity 765 anomaly is recognized (Wortel and Spakman, 2000; Piromallo et al., 2001; Piromallo and 766 Morelli, 2003; Hetényi et al. 2009). The origin of this anomaly was postulated as 767 accumulation of thick, cold material, possibly subducted residual material supplied from 768 several subduction episodes (Penninic in the Cretaceous, Vardar in the Palaeogene, and the 769 East Carpathian subduction in the Miocene; Hetényi et al. 2009). This vast amount of cold 770 material can form a barrier against the rise of any lower mantle-derived plume and prevents 771 the initiation of upwelling of hot material from the Transition Zone. However, it has been

772 shown that the Transitional Zone could be one of the wettest parts of the upper mantle, 773 particularly when subducted material is accumulated there for a long time (Hirschmann, 2006; 774 Karato, 2011). Thus, it could serve as the source for wet plumes instead of hot plumes. In this 775 scenario, wet melting in the upper part of the Transition Zone (Hirschmann, 2006) could 776 provide small volume hydrous magmas that initiate further melting beneath the continental 777 lithosphere. Although this is a reasonable scenario it does not explain the observation why the 778 basaltic volcanic fields are at the marginal parts of the Pannonian basin and not in the central 779 areas. The characteristic features of hot mantle plumes are also missing in the CPR. The broad 780 topographic updoming such as detected in the Massif Central (Wilson and Patterson, 2001), or 781 in classic hot spot areas (Courtillot et al., 2003), is absent; on the contrary, some parts of the 782 Pannonian Basin are still subsiding. The high heat flow in the CPR (Lenkey et al., 2002) can 783 be readily explained by the thinning of the lithosphere during basin formation and by the still 784 shallow asthenosphere (Royden et al., 1982). As far as the HIMU-like isotopic ratios of some 785 CPR basalts are concerned, they could have also a metasomatic origin (e.g., Hart, 1988; Sun 786 and McDonough, 1989; Halliday et al., 1995; Pilet et al., 2005). The sporadic distribution and 787 the low-magma output type of the volcanic fields in the CPR also do not support any hot 788 mantle plume origin. Therefore, there are no supporting evidences that favour hot mantle 789 plume or plume finger (Seghedi et al., 2004) origin of the basaltic volcanism in the CPR. 790 Generation of basaltic magmas in the asthenosphere requires upward movement of 791 solid mantle material, i.e. a mantle flow. In this context, the spatial distribution of the basalt 792 volcanic fields at the peripheral areas of the CPR could have relevance. They occur above a 793 lithosphere-asthenosphere boundary with sharp depth variation between the orogenic Alps 794 and cratonic North European Platform with thick (>160 km) lithospheric roots and the 795 stretched Pannonian basin with thin (< 70 km) lithosphere within relatively short distances 796 (Fig. 12.). The Pannonian basin could therefore act as a thin-spot providing suction in the

797 sublithospheric mantle and generating mantle flow from below the thick lithospheric roots 798 (Harangi and Lenkey, 2007; Harangi, 2009). The SKS splitting measurements show a 799 dominant west-east direction beneath the western part of the Pannonian basin with a clear 800 continuation from the Alps and this direction changes to a northwest-southeast trend toward 801 the central and northern areas (Kovács et al. 2012). On the western side of the Alps, similar, 802 east-west directed SKS splitting directions were recognized that turn gradually in north-south 803 direction northward and this was considered as relatively recent asthenospheric flow (Barroul 804 et al. 2011). Remarkably, a similar seismic pattern was observed also in the Lake Baikal area, 805 where it was interpreted as sublithospheric mantle flow from beneath the thick Siberian craton 806 towards the thin Baikal–Mongolian area (Lebedev et al., 2006). The low-volume, scattered, 807 and sporadic basalt volcanism in the Baikal rift could have been connected with this mantle 808 flow by decompression melting when the asthenospheric material ascended from the 200-km 809 depth beneath the craton to shallower depths beneath the rift (Lebedev et al., 2006). A similar 810 triggering mechanism was put forward by Niu (2005) explaining the basaltic volcanism in the East China lowland. Thus, it appears that a large contrast in the LAB architecture could 811 812 enhance mantle flow that results in intraplate basaltic volcanism with many similarities to OIB-type compositions. This appears to be the case also in the Alps-Pannonian basin 813 814 transition zone, where relatively hot asthenospheric material could ascend from beneath the 815 Alps along the steep LAB (Harangi, 2009; Fig. 13). The important role of asthenospheric 816 mantle flow as a driving force for the Neogene extension and extrusion events is getting to 817 have an increasing attention in the geodynamic model of the Pannonian basin. Konečný et al. 818 (2002) invoked an asthenospheric flow due to subduction roll-back mechanism and explained 819 the magma generation by this process. Kovács et al. (2012) proposed that this mantle flow is 820 the result of the convergence between the Eurasian and African plates along the Alpine belt, 821 whereas Horváth and Faccenna (2011) suggested that mantle material was pumped through

the Istria slab window beneath the CPR mantle as a result of slab rollback along the eastern 822 823 margin of the region. An important difference between these models and the explanation put 824 forward in this paper is the time of the mantle flow. According to Konečný et al. (2002), 825 Kovács et al. (2012) and Horváth and Faccenna (2011) this process could have been taking 826 place since the Oligocene/Miocene and act as a major role in the geodynamic evolution of the 827 CPR, whereas here, we propose that it is a passive process commenced just after the Mid-828 Miocene formation of the Pannonian basin with strongly thinned lithosphere and in this case 829 this thin spot condition acted as a driving force of the mantle flow.

830 In the northern part of the Pannonian basin, there has been a long lasting volcanic 831 activity since about 16 Ma (Harangi et al., 2007). The alkaline basalt volcanism followed an 832 intense calc-alkaline volcanic period (Konečný et al., 1995; 2002). Nevertheless, no traces of 833 subduction-related mantle source are observed in the composition of the basaltic products. It 834 can be explained by effective change in the sublithospheric mantle beneath the Stiavnica and 835 Nógrád-Gemer area that could require again mantle flow mechanism. In the seismic 836 tomography model sections published by Dando et al. (2011) a low-velocity anomaly can be 837 observed beneath this region that has a continuation into the deeper mantle towards north. Tentatively, this can be interpreted as a sign of an upward flow beneath the steep LAB. The 838 839 low magma output and the sporadic eruption events over several million years long period (8-840 0.1 Ma) could suggest, however, a tectonic control of the melt generation-magma ascent 841 processes. Strain localization in the upper mantle beneath major faults that penetrate the upper 842 mantle (Leloup et al. 1995; Vauchez et al., 2012) could be a possible mechanism of this. 843 Reactivation of transtensional tectonic lines could have a close link with the tectonic inversion of the Pannonian basin (Horváth and Cloetingh, 1996) that started in the Late Miocene when 844 845 the peak magma-output phase occurred in the Bakony-Balaton Upland and the Štiavnica and

846 Nógrád-Gemer volcanic fields. Further research could focus on the interrelationship of847 tectonic processes along pre-existing lines and the volcanic events.

848 The youngest basalt volcanic field was developed at the eastern part of the CPR 849 (Perşani area; Seghedi and Szakács, 1994; Downes et al., 1995) from 1.2 to 0.6 Ma (Panaiotu 850 et al., 2013). Seghedi et al. (2011) invoked toroidal asthenospheric flow around the 851 downgoing slab beneath Vrancea to explain the localized alkali basalt volcanism, whereas 852 Harangi et al. (2013) proposed that this could have occurred as a response to the formation of 853 a narrow rupture in the lower lithosphere, possibly as a far-field effect of the dipping of dense 854 continental lithospheric material beneath the Vrancea zone. Upper crustal extensional stress 855 field with reactivation of normal faults at the eastern margin of the Transylvanian basin could 856 enhance the rapid ascent of the mafic magmas.

Finally, it is remarkable that local low-velocity mantle anomaly has been detected beneath both the Štiavnica and the Perşani areas (Dando et al., 2011; Popa et al., 2012) that along with our interpretation about the presence of mantle region containing still fusible material suggest that further volcanic eruptions cannot be unambiguously excluded in these areas. The rapid ascent of the magmas (Harangi et al., 2013; Jankovics et al., 2013) implies that the quiet situation could change rapidly.

863

864 Summary

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Evaluation of the major and trace element content of the most primitive basalts in the CPR along with the interpretation of the most primitive spinel compositions indicates that the primary magmas could have generated in the convective asthenosphere. Trace element modelling indicates incompatible trace element-enriched source composition and 2-3% melting within the garnet-spinel transition zone. Melting started at about 100 km depth where

garnet prevails over spinel and continued usually up to the base of the lithosphere. Thus, the 871 872 final melting pressure could indicate the ambient lithosphere-asthenosphere boundary (LAB). 873 The asthenospheric mantle source regions of the basalts were heterogeneous, presumably at 874 small scale. Spinel compositions show a wide range, although four coherent groups can be 875 distinguished that are interpreted as reflect the main mantle domains beneath this region. They 876 can be recognized at almost each volcanic field. Two of these have fertile and slightly 877 depleted character, respectively, whereas the two others are enriched in Fe and Ti and overlap 878 the spinels found in the basalts of Hawaii. They might indicate involvement of 879 pyroxenitic/eclogitic lithology in the melting process that requires further investigation. Major 880 element composition of some of the basalts (low Al and high Ti and Fe; e.g., Pauliberg, 881 Putikov) could be also consistent with melting of carbonated, silica-deficient eclogite. Indeed, 882 considering the depths of the estimated melting columns of the CPR basaltic magmas, a more 883 fusible material than the dry peridotite is required in the source regions of the magmas. 884 Presence of water or pyroxenite/eclogite could explain melting at such depth. 885 The prevailing estimated mantle potential temperature (1300-1400°C) along with 886 number of further observations do not require anomalously hot asthenospheric mantle. Thus, 887 we can exclude the presence of any mantle plume or mantle plume fingers beneath the 888 Pannonian basin. The locations of the monogenetic basalt volcanic fields at the western and 889 northern margin of the Pannonian basin could indicate a close relationship with the steep LAB 890 beneath these areas. We propose that asthenospheric flow from beneath the adjoining thick 891 lithospheric domains (Alps, cratonic North European Platform) to the thinned Pannonian 892 basin could play a triggering role in the melt generation. A near vertical upwelling along the 893 steep LAB could result in decompressional melting producing low-volume melts. The 894 Pannonian basin could therefore act as a thin-spot after the Mid-Miocene syn-rift phase (20-895 12 Ma) and provided suction in the sublithospheric mantle generating asthenospheric flow
896 from below the thick lithospheric roots. In addition to this dominant process, single basaltic 897 volcanic events in the inner part of the Pannonian basin could have been related to strain-898 localization in the upper mantle beneath major strike-slip faults that penetrate the upper 899 mantle. Further southeast, close to the Vrancea zone, the youngest basalt volcanic field in the 900 CPR could have been formed as a plate tectonic response of the dipping dense continental 901 lithospheric material. The suction of the descending lithospheric slab could result in narrow 902 rupture at the base of the lithosphere. Decompression melting of the upwelling asthenospheric 903 mantle could result in magma generation, while reactivation of normal faults enhanced the 904 rapid ascent of the magmas.

905 Origin of basaltic magmas with fairly primitive composition for the last 1 Ma implies 906 that fusible material could still exists in the mantle, which have a potential temperature close 907 to the solidus. This inferred condition and the low-velocity anomalies detected in the mantle 908 depths as well as the still high localized heat flows in these areas all suggest that continuation 909 of the basaltic volcanism in the CPR cannot be unambiguously excluded.

910

911 Acknowledgement

912

913 Our ideas about the basaltic volcanism in the Carpathian-Pannonian region have 914 refined during number of instructive and sometimes heavy discussions with several colleagues 915 such as Hilary Downes, Theodoros Ntaflos, Ioan Seghedi, Frank Horváth, Károly Németh, 916 Csaba Szabó, István Kovács, Orlando Vaselli, László Lenkey, Marge Wilson, Michele 917 Lustrino, Jaroslav Lexa, Kadosa Balogh, Zoltán Pécskay and László Fodor. These were 918 promoted by great workshops started with the PANCARDI meetings in the 1990's and early 919 2000's, followed by the ILP PLUME workshop at Mt. Sainte-Odile in 2006, the EMAW 920 conference in Ferrara, 2007 and the Basalt2013 in Görlitz in 2013. Discussions about the

921	mantle potential temperature calculations with Claude Herzberg and Keith Putirka were
922	particularly stimulated. Terry Plank kindly provided the calculation scheme of the Langmuir
923	et al. (1992) melting model. Participation of the young scientists, M. Éva Jankovics, Balázs
924	Kiss and Réka Lukács was supported in this research by the European Union and the State of
925	Hungary, co-financed by the European Social Fund in the framework of TÁMOP-4.2.4.A/2-
926	11/1-2012-0001 `National Excellence Program'. Constructive comments provided by Stefan
927	Jung and an anonymous reviewer helped us to clarify our views.
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1496 Figure captions

1498	Figure 1. $-$ (A) Geological setting and (B) location of the Miocene to Quaternary basalts in
1499	the Carpathian-Pannonian Region with the main tectonic lines (after Horváth et al.,
1500	2006). SVF=Styrian Volcanic Field; BU=Burgenland basalts; KVF=Kemenes
1501	Volcanic Field; BBVF=Bakony-Balaton Upland Volcanic Field; ST= Štiavnica
1502	basalts; KE=Kecel basalts (buried); LU=Lucaret basalt; PVF=Perşani Volcanic Field.
1503	Figure 2. – Classification of the basalts based on the TAS diagram (Le Bas et al. 1986). The
1504	used data set is found in the supplementary table along with the references to the
1505	sources of the data. Data of the Pásztori trachyte is from Harangi et al. (1995) and
1506	Harangi (2001b).
1507	Figure 3 Selected major elements (in wt%) and trace elements (in ppm) vs. MgO for the
1508	CPR basalts. Symbols as in Fig. 2.
1509	Figure 4. – Primitive mantle (McDonough and Sun, 1995) normalized trace element patterns
1510	for representative CPR basalts. Note that two groups can be distinguished based on the
1511	incompatible trace element abundances and the existence of negative K-anomaly.
1512	Figure 5. – $(La/Yb)_N$ vs. K/K* and ¹⁴³ Nd/ ¹⁴⁴ Nd vs. K/K* diagrams for the CPR basalts. Note
1513	the negative correlation between these variables suggesting formation of the strongest
1514	negative K-anomaly in the melts generated at small degree of melting where garnet
1515	remained in the residue. The covariance between 143 Nd/ 144 Nd and K/K* in the
1516	different basalt volcanic fields of the CPR implies that the negative K-anomaly could
1517	be inherited in the mantle source regions. The K/K* is calculated as $K_N/\sqrt{(Nb_N*La_N)}$,
1518	where N denotes normalization to the primitive mantle (McDonough and Sun, 1995)
1519	values. Symbols as in Fig. 2.

Figure 6. – (A) Estimation of the initial and final pressures (p_0 and p_f , respectively) of the 1520 1521 melting column for selected CPR basalts using the adiabatic decompression model of Langmuir et al. (1992). The FeO^t and the Na₂O of the calculated primary magmas of 1522 1523 the CPR basalts yield information on the initial melting pressure and the cessation of 1524 the melting, i.e. the final melting pressure. Each curve represents polybaric mantle 1525 melting lines with different initial melting pressure, i.e. where the ambient mantle 1526 temperature intersects the solidus. Each tick mark represents 0.1 GPa of 1527 decompression. The compositions of the primary magmas at the melting curves provide the values of the final pressure. 1528

(B) Determination of melting pressure based on the calculated primary magma

1530 compositions of the CPR basalts using the ol'-ne'-qtz' plot of Hirose and Kushiro

1531 (1993). ol'=ol+0.75opx; ne'=ne+0.6ab; qtz'=qtz+0.4ab+0.25opx, where ol, opx, ne,

ab, qtz are CIPW normative mineral components. Melting pressure lines are afterSakuyama et al. (2009).

Figure 7. – Calculated melting columns for different basalts (based on Langmuir et al., 1992 model) compared with the present lithosphere-asthenpshere boundary (LAB) beneath the volcanic fields that is inferred from interpretation of seismic data (Horváth et al., 2006). Mantle potential temperatures are given using the calculation scheme of Lee et al. (2009) using molar $Fe^{3+}/Fe^{tot}=0.13$. Beneath the Perşani volcanic field Dérerova et al. (2006) and Horváth et al. (2006) suggested a thick lithosphere (about 140 km),

1540 whereas Martin et al. (2006) indicated much thinner lithosphere. Our melting column

1541 calculation for the Perşani basalts is consistent with the proposed lithosphere thickness1542 of Martin et al. (2006).

Figure 8. – Comparison of the melting pressure obtained by the calculation procedure of Lee
et al. (2009) with the final melting pressure (p_f) provided by the Langmuir et al. (1992)

model. There is a reasonably good correlation between these pressure values. Symbolsas in Fig. 2.

1547	Figure 9 Trace element modelling for the melt generation of the CPR basaltic magmas.
1548	Model parameters: critical melting model (critical melt fraction is 0.01) proposed by
1549	Sobolev and Shimizu (1992) with the following source rocks: spinel-lherzolite –
1550	olivine (57%), orthopyroxene (25.5%), clinopyroxene (15%), spinel (2.5%); garnet-
1551	lherzolite - olivine (60.1%), orthopyroxene (18.9%), clinopyroxene (13.7%), garnet
1552	(7.3%). Melting modes are ol1.21opx8.06cpx76.37sp14.36 for sp-lherzolite and
1553	ol1.3opx8.7cpx36gt54 for garnet-lherzolite, respectively. Source rock composition: La
1554	and Nb - 4 x primitive mantle values (2.59 and 2.63 ppm, respectively), Zr - 2 x
1555	primitive mantle values (21 ppm) and $Y - 1.5 x$ primitive mantle values (6.45 ppm,
1556	where primitive mantle values are from McDonough and Sun, 1995). Distribution
1557	coefficients are from Kostopoulos and James (1992). Tick marks indicate the degree
1558	of melting at each melting model. Symbols are explained in Fig. 2.
1559	Figure 10. – Melting conditions beneath the Carpathian-Pannonian region in a pressure-
1560	temperature diagram. The position of the melting columns is based on the result of the
1561	Langmuir et al. (1992) model combined with the mantle potential temperature and
1562	melting pressure values calculated by the thermbarometer of Lee et al. (2009). The
1563	melting columns of the basaltic magmas are just below the dry peridotite solidus (as
1564	defined by Hirschmann, 2000). This requires either some water or a more fusible
1565	material, such as pyroxenite or eclogite in the mantle source. Hydrous peridotite solidi:
1566	Hirschmann (2006); pyroxenite (MIX1G) pyroxenite solidus: Kogiso et al. (2003);
1567	MORB-like pyroxenite solidus: Pertermann and Hirschmann (2003). UMX: position
1568	of the ultramafic peridotite xenoliths as given by Embey-Isztin et al. (2001) and
1569	Kovács et al. (2012).

1570	Figure 11 Compositional characteristics of spinel inclusions in olivine phenocrysts of the
1571	CPR basalts based on the Cr-Fe ³⁺ -Al diagram. The most primitive spinels form four
1572	compositionally coherent groups that can be related to the existence of different
1573	mantle domains in the source regions of the basaltic magmas. Note that three of these
1574	groups are present beneath each volcanic fields of the CPR. MORB and OIB spinel
1575	fields are as defined by Barnes and Roeder (2001). Compositional data of spinels are
1576	from Jankovics et al. (2012), Harangi et al. (2013) and Sági et al. (in prep).
1577	Figure 12. – Position of the basalt volcanic fields in the CPR compared with the seismically
1578	defined lithosphere-asthenosphere boundary lines (Horváth et al., 2006). Note that the
1579	volcanic fields are located at the margin of the Pannonian basin and not in the central
1580	parts where the lithosphere is the thinnest.
1581	Figure 13. – Conceptual model for the triggering of melt generation beneath the Pannonian
1582	basin. See text for explanation.
1583	
1584	Table 1 – Representative major and trace element composition of the basalts from different
1585	volcanic fields of the CPR. Reference: 1: Ali et al. (2013), 2: Ali and Ntaflos (2011),
1586	3: this study, 4: Harangi et al. (2013). p_f and p_o are the calculated final and initial
1587	melting pressures using the Langmuir et al. (1992) model. T and p are the mantle
1588	potential temperature and the melting pressure using the thermobarometer of Lee et al.
1589	(2009) with molar $Fe^{3+}/Fe^{tot}=0.13$.

Figure 1A Click here to download high resolution image



Figure 1B Click here to download high resolution image





Figure 3 Click here to download high resolution image


















Temperature (°C)





Figure 13 Click here to download high resolution image



locality Klöch Steinberg Pauliberg Oberpul- lendor Ság Sátorna Kapolcs volcanic field SVF SVF BU BU KVF BBVF BBVF Reference 1 1 2 2 3 3 3 SIO2 44.97 43.87 46.99 49.07 48.19 47.24 46.53 TIO2 2.25 2.12 3.76 2.73 2.31 2.18 2.24 Al2O3 15.58 14.92 11.13 12.75 14.10 15.51 15.37 Fe2O3t 10.03 11.02 12.59 11.88 1.79 10.35 9.03 9.61 CaO 10.19 10.80 9.95 10.13 9.04 8.80 9.26 Na2O 4.88 5.47 2.78 2.65 2.83 3.65 3.18 PO5 0.72 1.06 0.62 0.44 0.50 0.79 0.67 LO	sample	KL-4	ST-11	PLB-4	OB-1	SAG08-2	S-1	KK-1
volcanic field SVF SVF BU BU KVF BBVF BBVF Reference 1 1 2 2 3 3 3 SiO2 44.97 43.87 46.99 49.07 48.19 47.24 46.53 TiO2 2.25 2.12 3.76 2.73 2.31 2.18 2.24 Al2O3 15.58 14.92 11.13 12.75 14.10 16.51 15.51 F2O3t 10.03 11.02 12.59 11.58 11.79 10.35 0.66 MgO 9.40 8.19 10.54 9.14 9.55 9.03 9.61 CaO 10.19 10.80 9.95 10.13 9.04 8.80 9.26 NaZO 4.58 5.47 2.78 2.65 2.83 3.65 3.13 K2O 0.72 1.06 0.62 0.44 0.50 0.67 0.67 LOT 0.98 0.23 <t< td=""><td>locality</td><td>Klöch</td><td>Steinberg</td><td>Pauliberg</td><td>Oberpul-</td><td>Ság</td><td>Sátorma</td><td>Kapolcs</td></t<>	locality	Klöch	Steinberg	Pauliberg	Oberpul-	Ság	Sátorma	Kapolcs
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$\begin{array}{c c c c c c c c c c c c c c c c c c c $	SiO2	44.97	43.87	46.99	49.07	48.19	47.24	46.53
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MnO 0.17 0.20 0.15 0.20 0.14 0.16 0.15 MgO 9.40 8.19 10.54 9.14 9.55 9.03 9.61 CaO 10.19 10.80 9.95 10.13 9.04 8.80 9.26 NaZO 2.58 2.65 2.83 3.65 3.13 K2O 2.11 2.37 1.48 1.31 1.53 2.30 2.18 P2O5 0.72 1.06 0.62 0.44 0.50 0.79 0.67 LOI 0.98 0.23 0.11 2.12 2.1 1.6 1.4 mg-number 0.69 0.63 0.66 0.65 0.65 0.67 0.67 SL -27.25 -38.32 -3.19 2.40 -0.44 -12.19 -10.31 Ni 138 123 287 356 246 260 V 215 206 172 200 198 182 176	Fe2O3t	10.03	11.02	12.59	11.58	11.79	10.35	10.86
$\begin{array}{llllllllllllllllllllllllllllllllllll$	MnO	0.17	0.20	0.15	0.20	0.14	0.16	0.15
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	MgO	9.40	8.19	10.54	9.14	9.55	9.03	9.61
Na2O 4.58 5.47 2.78 2.65 2.83 3.65 3.13 K2O 2.11 2.37 1.48 1.31 1.53 2.30 2.18 P2O5 0.72 1.06 0.62 0.44 0.50 0.79 0.67 LOI 0.98 0.23 0.11 2.12 2.1 1.6 1.4 mg-number 0.69 0.63 0.66 0.65 0.65 0.67 0.67 SL -27.25 -38.32 -3.19 2.40 -0.44 -12.19 -10.31 Ni 136 121 332 287 356 246 260 V 215 206 172 200 198 182 176 Sc 19.4 13.6 16.4 20.9 21 19 22 Co 38.1 36 51.5 45.6 48.7 41.4 41.4 Rb 52.8 62.2 39.8 21.2 26.5 56.8 50.7 Sr 1008 1203 881	CaO	10.19	10.80	9.95	10.13	9.04	8.80	9.26
K2O2.112.371.481.311.532.302.18P2O50.721.060.620.440.500.790.67LOI0.980.230.112.122.11.61.4mg-number0.690.630.660.650.670.67S.I27.25-38.32-3.192.40-0.44-12.19-10.31Ni138123274324228200172Cr136121332287356246260V215206172200198182176Sc19.413.616.420.9211922Co38.13651.545.648.741.441.4Rb52.862.239.821.226.556.850.7Ba58171.3479305602877627Sr10081203881577808.11159.21385.8Zr368556498364187.8282.9194.7Nb94.211172.449.840.980.153.3Y28.534.331.427.321.127.521.5Th9.912.65.263.644.78.86.7La56.675.443.425.929.852.936.2Ce10413594.864.259.	Na2O	4.58	5.47	2.78	2.65	2.83	3.65	3.13
P205 0.72 1.06 0.62 0.44 0.50 0.79 0.67 LOI 0.98 0.23 0.11 2.12 2.1 1.6 1.4 mg-number 0.69 0.63 0.66 0.65 0.65 0.67 0.67 SL -27.25 -38.32 -3.19 2.40 -0.44 -12.19 -10.31 Ni 138 121 332 287 356 246 260 V 215 206 172 200 198 182 176 Sc 19.4 13.6 16.4 20.9 21 19 22 Co 38.1 36 51.5 45.6 48.7 41.4 41.4 Rb 52.8 62.2 39.8 21.2 26.5 56.8 50.7 Ba 581 713 479 305 602 877 627 Sr 1008 1203 881 577 808.1 1159.2 1385.8 Zr 368 564 498 364	K2O	2.11	2.37	1.48	1.31	1.53	2.30	2.18
$\begin{array}{c c c c c c c c c c c c c c c c c c c $	P2O5	0.72	1.06	0.62	0.44	0.50	0.79	0.67
mg-number 0.69 0.63 0.66 0.65 0.65 0.67 0.67 S.I. -27.25 -38.32 -3.19 2.40 -0.44 -12.19 -10.31 Ni 138 123 274 324 228 200 172 Cr 136 121 332 287 356 246 260 Sc 19.4 13.6 16.4 20.9 21 19 22 Co 38.1 36 51.5 45.6 48.7 41.4 41.4 Rb 52.8 62.2 39.8 21.2 26.5 56.8 50.7 Ba 511 713 479 305 602 877 627 Sr 1008 1203 881 577 808.1 1159.2 1385.8 Zr 368 556 498 364 187.8 282.9 194.7 Nb 94.2 111 72.4 49.8 40.9 80.1 53.3 Y 28.5 34.3 31.4 27.3 21.1 27.5 21.5 Th 9.9 12.6 5.26 3.64 4.7 8.8 6.7 La 56.6 75.4 43.4 25.9 92.8 52.9 36.2 Ce 104 135 94.8 64.2 59.8 101.3 72.1 Pr 11.4 14.5 12.1 6.9 7.81 12.2 9.11 Nd 462 59 49.7 <	LOI	0.98	0.23	0.11	2.12	2.1	1.6	1.4
S.I. -27.25 -38.32 -3.19 2.40 -0.44 -12.19 -10.31 Ni138123274324228200172Cr136121332287356246260V215206172200198182176Sc19.413.616.420.9211922Co38.13651.545.648.741.441.4Rb52.862.239.821.226.556.850.7Ba581713479305602877627Sr10081203881577808.11159.21385.8Zr368556498364187.8282.9194.7Sr10081203881577808.11159.21385.8Zr368556498364187.8282.9194.7Sr1008120384157.321.127.521.5Th9.912.65.263.644.78.86.7La56.675.443.425.929.852.936.2Ce10413594.864.259.8101.372.1Pr11.414.512.16.166.878.136.69Eu2.242.713.071.922.132.362.05Gd6.688.099.2 <td>mg-number</td> <td>0.69</td> <td>0.63</td> <td>0.66</td> <td>0.65</td> <td>0.65</td> <td>0.67</td> <td>0.67</td>	mg-number	0.69	0.63	0.66	0.65	0.65	0.67	0.67
Ni138123274324228200172Cr136121332287356246260V215206172200198182176Sc19413.616.420.9211922Co38.13651.545.648.741.441.4Rb52.862.239.821.226.556.850.7Ba581713479305602877627Sr10081203881577808.11159.21385.8Zr368556498364187.8282.9194.7Nb94.211172.449.840.980.153.3Y28.534.331.427.321.127.521.5Th9.912.65.263.644.78.86.7La56.675.443.425.929.852.936.2Ce10413594.864.259.8101.372.1Pr11.414.512.16.97.8112.29.11Nd46.25.755.894.114.374.864.73Sm7.268.9710.56.166.878.136.69Eu2.242.713.071.922.132.362.05Gd6.688.099.25.825.86	S.I.	-27.25	-38.32	-3.19	2.40	-0.44	-12.19	-10.31
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	Ni	138	123	274	324	228	200	172
V215206172200198182176Sc19.413.616.420.9211922Co38.13651.545.648.741.441.4Rb52.862.239.821.226.556.850.7Ba581713479305602877627Sr10081203881577808.11159.21385.8Zr368556498364187.8282.9194.7Nb94.211172.449.840.980.153.3Y28.534.331.427.321.127.521.5Th9.912.65.263.644.78.86.7La56.675.443.425.925.936.2Ce10413594.864.259.8101.372.1Pr11.414.512.16.97.8112.29.11Nd46.25949.730.731.648.737.8Sm7.268.9710.56.166.878.136.69Eu2.242.713.071.922.132.362.05Gd6.688.099.25.825.666.765.67Tb0.871.021.230.8060.850.80.73Dy4.455.155.894.114.37	Cr	136	121	332	287	356	246	260
Sc19.413.616.420.9211922Co 38.1 36 51.5 45.6 48.7 41.4 41.4 Rb 52.8 62.2 39.8 21.2 26.5 56.8 50.7 Ba 581 713 479 305 602 877 627 Sr 1008 1203 881 577 808.1 1159.2 1385.8 Zr 368 556 498 364 187.8 282.9 194.7 Nb 94.2 111 72.4 49.8 40.9 80.1 53.3 Y 28.5 34.3 31.4 27.3 21.1 27.5 21.5 Th 9.9 12.6 5.26 3.64 4.7 8.8 6.7 La 56.6 75.4 43.4 25.9 29.8 52.9 36.2 Ce 104 135 94.8 64.2 59.8 101.3 72.1 Pr 11.4 14.5 12.1 6.9 7.81 12.2 9.11 Nd 46.2 59 49.7 30.7 31.6 48.7 37.8 Sm 7.26 8.97 10.5 6.16 6.87 8.13 6.69 Eu 22.4 2.71 3.07 1.92 2.13 2.36 2.05 Gd 6.68 8.09 9.2 5.82 5.86 6.76 5.67 Tb 0.87 1.02 1.23 0.806 <	V	215	206	172	200	198	182	176
Co 38.1 36 51.5 45.6 48.7 41.4 41.4 Rb 52.8 62.2 39.8 21.2 26.5 56.8 50.7 Ba 581 713 479 305 602 877 627 Sr 1008 1203 881 577 808.1 1159.2 1385.8 Zr 368 556 498 364 187.8 282.9 194.7 Nb 94.2 111 72.4 49.8 40.9 80.1 53.3 Y 28.5 34.3 31.4 27.3 21.1 27.5 21.5 Th 9.9 12.6 5.26 3.64 4.7 8.8 6.7 La 56.6 75.4 43.4 25.9 29.8 52.9 36.2 Ce 104 135 94.8 64.2 59.8 101.3 72.1 Pr 11.4 14.5 12.1 6.9 7.81 12.2 9.11 Nd 46.2 59 49.7 30.7 31.6 48.7 37.8 Sm 7.26 8.97 10.5 6.16 6.87 8.13 6.69 Eu 2.24 2.71 3.07 1.92 2.13 2.36 2.05 Gd 6.68 8.09 9.2 5.82 5.86 6.76 5.67 Tb 0.87 1.02 1.23 0.806 0.85 0.8 0.73 Dy 4.45 5.15 5.89	Sc	19.4	13.6	16.4	20.9	21	19	22
Rb 52.8 62.2 39.8 21.2 26.5 56.8 50.7 Ba 581 713 479 305 602 877 627 Sr 1008 1203 881 577 808.1 1159.2 1385.8 Zr 368 556 498 364 187.8 282.9 194.7 Nb 94.2 111 72.4 49.8 40.9 80.1 53.3 Y 28.5 34.3 31.4 27.3 21.1 27.5 21.5 Th 9.9 12.6 5.26 3.64 4.7 8.8 6.7 La 56.6 75.4 43.4 25.9 29.8 52.9 36.2 Ce 104 135 94.8 64.2 59.8 101.3 72.1 Pr 11.4 14.5 12.1 6.9 7.81 12.2 9.11 Nd 46.2 59 49.7 30.7 31.6 48.7 37.8 Sm 7.26 8.97 10.5 6.16 6.87 8.13 6.69 Eu 2.24 2.71 3.07 1.92 2.13 2.36 2.05 Gd 6.68 8.09 9.2 5.82 5.86 6.76 5.67 Tb 0.87 1.02 1.23 0.806 0.85 0.8 0.73 Dy 4.45 5.15 5.89 4.11 4.37 4.86 4.22 Ho 0.82 0.94 1.0	Со	38.1	36	51.5	45.6	48.7	41.4	41.4
Ba581713479305602877627Sr10081203881577808.11159.21385.8Zr368556498364187.8282.9194.7Nb94.211172.449.840.980.153.3Y28.534.331.427.321.127.521.5Th9.912.65.263.644.78.86.7La56.675.443.425.929.852.936.2Ce10413594.864.259.8101.372.1Pr11.414.512.16.97.8112.29.11Nd46.25949.730.731.648.737.8Sm7.268.9710.56.166.878.136.69Eu2.242.713.071.922.132.362.05Gd6.688.099.25.825.866.765.67Tb0.871.021.230.8060.850.80.73Dy4.455.155.894.114.374.864.22Ho0.820.941.040.7550.740.840.78Er2.12.432.531.871.872.211.95Tm0.280.320.3120.2490.280.310.28Yb1.82.041.91	Rb	52.8	62.2	39.8	21.2	26.5	56.8	50.7
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	Ba	581	713	479	305	602	877	627
Zr368556498364187.8282.9194.7Nb94.211172.449.840.980.153.3Y28.534.331.427.321.127.521.5Th9.912.65.263.644.78.86.7La56.675.443.425.929.852.936.2Ce10413594.864.259.8101.372.1Pr11.414.512.16.97.8112.29.11Nd46.25949.730.731.648.737.8Sm7.268.9710.56.166.878.136.69Eu2.242.713.071.922.132.362.05Gd6.688.099.25.825.866.765.67Tb0.871.021.230.8060.850.80.73Dy4.455.155.894.114.374.864.22Ho0.820.941.040.7550.740.840.78Er2.12.432.531.871.872.211.95Tm0.280.320.3120.2490.280.310.28Vb1.82.041.911.531.491.881.52Lu0.250.280.2460.210.210.280.24Hf5.968.829.32	Sr	1008	1203	881	577	808.1	1159.2	1385.8
Nb94.211172.449.840.980.153.3Y28.534.331.427.321.127.521.5Th9.912.65.263.644.78.86.7La56.675.443.425.929.852.936.2Ce10413594.864.259.8101.372.1Pr11.414.512.16.97.8112.29.11Nd46.25949.730.731.648.737.8Sm7.268.9710.56.166.878.136.69Eu2.242.713.071.922.132.362.05Gd6.688.099.25.825.866.765.67Tb0.871.021.230.8060.850.80.73Dy4.455.155.894.114.374.864.22Ho0.820.941.040.7550.740.840.78Er2.12.432.531.871.872.211.95Tm0.280.320.3120.2490.280.310.28Yb1.82.041.911.531.491.881.52Lu0.250.280.2460.210.210.280.24Hf5.968.829.326.164.964.8Ta5.126.823.43 <td>Zr</td> <td>368</td> <td>556</td> <td>498</td> <td>364</td> <td>187.8</td> <td>282.9</td> <td>194.7</td>	Zr	368	556	498	364	187.8	282.9	194.7
Y28.534.331.427.321.127.521.5Th9.912.65.263.644.78.86.7La56.675.443.425.929.852.936.2Ce10413594.864.259.8101.372.1Pr11.414.512.16.97.8112.29.11Nd46.25949.730.731.648.737.8Sm7.268.9710.56.166.878.136.69Eu2.242.713.071.922.132.362.05Gd6.688.099.25.825.866.765.67Tb0.871.021.230.8060.850.80.73Dy4.455.155.894.114.374.864.22Ho0.820.941.040.7550.740.840.78Er2.12.432.531.871.872.211.95Tm0.280.320.3120.2490.280.310.28Vb1.82.041.911.531.491.881.52Lu0.250.280.2460.210.210.280.24Hf5.968.829.326.164.964.8Ta5.126.823.432.692.14.53.3U2.543.031.57 <t< td=""><td>Nb</td><td>94.2</td><td>111</td><td>72.4</td><td>49.8</td><td>40.9</td><td>80.1</td><td>53.3</td></t<>	Nb	94.2	111	72.4	49.8	40.9	80.1	53.3
$\begin{array}{c c c c c c c c c c c c c c c c c c c $	Y	28.5	34.3	31.4	27.3	21.1	27.5	21.5
La56.675.443.425.929.852.936.2Ce10413594.864.259.8101.372.1Pr11.414.512.16.97.8112.29.11Nd46.25949.730.731.648.737.8Sm7.268.9710.56.166.878.136.69Eu2.242.713.071.922.132.362.05Gd6.688.099.25.825.866.765.67Tb0.871.021.230.8060.850.80.73Dy4.455.155.894.114.374.864.22Ho0.820.941.040.7550.740.840.78Er2.12.432.531.871.872.211.95Tm0.280.320.3120.2490.280.310.28Yb1.82.041.911.531.491.881.52Lu0.250.280.2460.210.210.280.24Hf5.968.829.326.164.964.8Ta5.126.823.432.692.14.53.3U2.543.031.570.9170.92.31.8p_f2.83.12.52.352.22.2p_o2.93.83.832.8<	Th	9,9	12.6	5.26	3.64	4.7	8.8	6.7
La101101101101101101101Ce10413594.864.259.8101.372.1Pr11.414.512.16.97.8112.29.11Nd46.25949.730.731.648.737.8Sm7.268.9710.56.166.878.136.69Eu2.242.713.071.922.132.362.05Gd6.688.099.25.825.866.765.67Tb0.871.021.230.8060.850.80.73Dy4.455.155.894.114.374.864.22Ho0.820.941.040.7550.740.840.78Er2.12.432.531.871.872.211.95Tm0.280.320.3120.2490.280.310.28Yb1.82.041.911.531.491.881.52Lu0.250.280.2460.210.210.280.24Hf5.968.829.326.164.964.8Ta5.126.823.432.692.14.53.3U2.543.031.570.9170.92.31.8p_f2.83.12.52.352.22.2p_o2.93.83.832.8 <td< td=""><td>La</td><td>56.6</td><td>75.4</td><td>43.4</td><td>25.9</td><td>29.8</td><td>52.9</td><td>36.2</td></td<>	La	56.6	75.4	43.4	25.9	29.8	52.9	36.2
Pr11.414.512.16.97.8112.29.11Nd46.25949.730.731.648.737.8Sm7.268.9710.56.166.878.136.69Eu2.242.713.071.922.132.362.05Gd6.688.099.25.825.866.765.67Tb0.871.021.230.8060.850.80.73Dy4.455.155.894.114.374.864.22Ho0.820.941.040.7550.740.840.78Er2.12.432.531.871.872.211.95Tm0.280.320.3120.2490.280.310.28Yb1.82.041.911.531.491.881.52Lu0.250.280.2460.210.210.280.24Hf5.968.829.326.164.964.8Ta5.126.823.432.692.14.53.3U2.543.031.570.9170.92.31.8p_f2.83.12.52.352.22.2p_o2.93.83.832.82.9T13901472141613761343	Ce	104	135	94.8	64.2	59.8	101.3	72.1
Nd 46.2 59 49.7 30.7 31.6 48.7 37.8 Sm 7.26 8.97 10.5 6.16 6.87 8.13 6.69 Eu 2.24 2.71 3.07 1.92 2.13 2.36 2.05 Gd 6.68 8.09 9.2 5.82 5.86 6.76 5.67 Tb 0.87 1.02 1.23 0.806 0.85 0.8 0.73 Dy 4.45 5.15 5.89 4.11 4.37 4.86 4.22 Ho 0.82 0.94 1.04 0.755 0.74 0.84 0.78 Er 2.1 2.43 2.53 1.87 1.87 2.21 1.95 Tm 0.28 0.32 0.312 0.249 0.28 0.31 0.28 Yb 1.8 2.04 1.91 1.53 1.49 1.88 1.52 Lu 0.25 0.28 0.246 0.21 0.21 0.28 0.24 Hf 5.96 8.82 9.32 6.16 4.9 6 4.8 Ta 5.12 6.82 3.43 2.69 2.1 4.5 3.3 U 2.54 3.03 1.57 0.917 0.9 2.3 1.8 p_f 2.8 3.1 2.5 2.35 2.2 2.2 p_o 2.9 3.8 3.8 3 2.8 2.9	Pr	11.4	14.5	12.1	6.9	7.81	12.2	9.11
Sm 7.26 8.97 10.5 6.16 6.87 8.13 6.69 Eu 2.24 2.71 3.07 1.92 2.13 2.36 2.05 Gd 6.68 8.09 9.2 5.82 5.86 6.76 5.67 Tb 0.87 1.02 1.23 0.806 0.85 0.8 0.73 Dy 4.45 5.15 5.89 4.11 4.37 4.86 4.22 Ho 0.82 0.94 1.04 0.755 0.74 0.84 0.78 Er 2.1 2.43 2.53 1.87 1.87 2.21 1.95 Tm 0.28 0.32 0.312 0.249 0.28 0.31 0.28 Yb 1.8 2.04 1.91 1.53 1.49 1.88 1.52 Lu 0.25 0.28 0.246 0.21 0.21 0.28 0.24 Hf 5.96 8.82 9.32 6.16 4.9 6 4.8 Ta 5.12 6.82 3.43 2.69 2.1 4.5 3.3 U 2.54 3.03 1.57 0.917 0.9 2.3 1.8 p_f 2.8 3.1 2.5 2.35 2.2 2.2 p_o 2.9 3.8 3.8 3 2.8 2.9	Nd	46.2	59	49.7	30.7	31.6	48.7	37.8
Eu 2.24 2.71 3.07 1.92 2.13 2.36 2.05 Gd 6.68 8.09 9.2 5.82 5.86 6.76 5.67 Tb 0.87 1.02 1.23 0.806 0.85 0.8 0.73 Dy 4.45 5.15 5.89 4.11 4.37 4.86 4.22 Ho 0.82 0.94 1.04 0.755 0.74 0.84 0.78 Er 2.1 2.43 2.53 1.87 1.87 2.21 1.95 Tm 0.28 0.32 0.312 0.249 0.28 0.31 0.28 Yb 1.8 2.04 1.91 1.53 1.49 1.88 1.52 Lu 0.25 0.28 0.246 0.21 0.21 0.28 0.24 Hf 5.96 8.82 9.32 6.16 4.9 6 4.8 Ta 5.12 6.82 3.43 2.69 2.1 4.5 3.3 U 2.54 3.03 1.57 0.917 0.9 2.3 1.8 p_f 2.8 3.1 2.5 2.35 2.2 2.2 p_o 2.9 3.8 3.8 3 2.8 2.9	Sm	7.26	8.97	10.5	6.16	6.87	8.13	6.69
Gd 6.68 8.09 9.2 5.82 5.86 6.76 5.67 Tb 0.87 1.02 1.23 0.806 0.85 0.8 0.73 Dy 4.45 5.15 5.89 4.11 4.37 4.86 4.22 Ho 0.82 0.94 1.04 0.755 0.74 0.84 0.78 Er 2.1 2.43 2.53 1.87 1.87 2.21 1.95 Tm 0.28 0.32 0.312 0.249 0.28 0.31 0.28 Yb 1.8 2.04 1.91 1.53 1.49 1.88 1.52 Lu 0.25 0.28 0.246 0.21 0.21 0.28 0.24 Hf 5.96 8.82 9.32 6.16 4.9 6 4.8 Ta 5.12 6.82 3.43 2.69 2.1 4.5 3.3 U 2.54 3.03 1.57 0.917 0.9 2.3 1.8 p_f 2.8 3.1 2.5 2.35 2.2 2.2 p_o 2.9 3.8 3.8 3 2.8 2.9	Eu	2.24	2.71	3.07	1.92	2.13	2.36	2.05
Tb 0.87 1.02 1.23 0.806 0.85 0.8 0.73 Dy 4.45 5.15 5.89 4.11 4.37 4.86 4.22 Ho 0.82 0.94 1.04 0.755 0.74 0.84 0.78 Er 2.1 2.43 2.53 1.87 1.87 2.21 1.95 Tm 0.28 0.32 0.312 0.249 0.28 0.31 0.28 Yb 1.8 2.04 1.91 1.53 1.49 1.88 1.52 Lu 0.25 0.28 0.246 0.21 0.21 0.28 0.24 Hf 5.96 8.82 9.32 6.16 4.9 6 4.8 Ta 5.12 6.82 3.43 2.69 2.1 4.5 3.3 U 2.54 3.03 1.57 0.917 0.9 2.3 1.8 p_f 2.8 3.1 2.5 2.35 2.2 2.2 p_o 2.9 3.8 3.8 3 2.8 2.9 T 1390 1472 1416 1376 1343	Gd	6.68	8.09	9.2	5.82	5.86	£.36	5.67
10^{-1} 102^{-1} <	Th	0.87	1.02	1.23	0.806	0.85	0.8	0.73
D_{y} D_{110} D_{110} D_{100} D_{110} D_{100} D_{120} Ho 0.82 0.94 1.04 0.755 0.74 0.84 0.78 Er 2.1 2.43 2.53 1.87 1.87 2.21 1.95 Tm 0.28 0.32 0.312 0.249 0.28 0.31 0.28 Yb 1.8 2.04 1.91 1.53 1.49 1.88 1.52 Lu 0.25 0.28 0.246 0.21 0.21 0.28 0.24 Hf 5.96 8.82 9.32 6.16 4.9 6 4.8 Ta 5.12 6.82 3.43 2.69 2.1 4.5 3.3 U 2.54 3.03 1.57 0.917 0.9 2.3 1.8 p_f 2.8 3.1 2.5 2.35 2.2 2.2 p_o 2.9 3.8 3.8 3 2.8 2.9 T 1390 1472 1416 1376 1343	Dv	4 4 5	5.15	5.89	4 11	4 37	4 86	4 22
Index 0.02 0.071 1.07 0.012 0.171 0.017 0.017 0.017 Er 2.1 2.43 2.53 1.87 1.87 2.21 1.95 Tm 0.28 0.32 0.312 0.249 0.28 0.31 0.28 Yb 1.8 2.04 1.91 1.53 1.49 1.88 1.52 Lu 0.25 0.28 0.246 0.21 0.21 0.28 0.24 Hf 5.96 8.82 9.32 6.16 4.9 6 4.8 Ta 5.12 6.82 3.43 2.69 2.1 4.5 3.3 U 2.54 3.03 1.57 0.917 0.9 2.3 1.8 p_f 2.8 3.1 2.5 2.35 2.2 2.2 p_o 2.9 3.8 3.8 3 2.8 2.9 T 1390 1472 1416 1376 1343	Ho	0.82	0.94	1.04	0.755	0.74	0.84	0.78
InInInInInInTm 0.28 0.32 0.312 0.249 0.28 0.31 0.28 Yb 1.8 2.04 1.91 1.53 1.49 1.88 1.52 Lu 0.25 0.28 0.246 0.21 0.21 0.28 0.24 Hf 5.96 8.82 9.32 6.16 4.9 6 4.8 Ta 5.12 6.82 3.43 2.69 2.1 4.5 3.3 U 2.54 3.03 1.57 0.917 0.9 2.3 1.8 p_f 2.8 3.1 2.5 2.35 2.2 2.2 p_o 2.9 3.8 3.8 3 2.8 2.9 T 1390 1472 1416 1376 1343	Fr	2.1	2 43	2 53	1.87	1.87	2.21	1.95
Yh 0.20 0.32 0.312 0.213 0.20 0.31 0.20 Yb 1.8 2.04 1.91 1.53 1.49 1.88 1.52 Lu 0.25 0.28 0.246 0.21 0.21 0.28 0.24 Hf 5.96 8.82 9.32 6.16 4.9 6 4.8 Ta 5.12 6.82 3.43 2.69 2.1 4.5 3.3 U 2.54 3.03 1.57 0.917 0.9 2.3 1.8 p_f 2.8 3.1 2.5 2.35 2.2 2.2 p_o 2.9 3.8 3.8 3 2.8 2.9 T 1390 1472 1416 1376 1343	Tm	0.28	0.32	0.312	0.249	0.28	0.31	0.28
Lu 0.25 0.28 0.246 0.21 0.21 0.28 0.24 Hf 5.96 8.82 9.32 6.16 4.9 6 4.8 Ta 5.12 6.82 3.43 2.69 2.1 4.5 3.3 U 2.54 3.03 1.57 0.917 0.9 2.3 1.8 p_f 2.8 3.1 2.5 2.35 2.2 2.2 p_o 2.9 3.8 3.8 3 2.8 2.9 T 1390 1472 1416 1376 1343	Yh	1.8	2.04	1.91	1.53	1 49	1.88	1.52
Ha 0.25 0.26 0.21 0.21 0.21 0.26 0.21 Hf 5.96 8.82 9.32 6.16 4.9 6 4.8 Ta 5.12 6.82 3.43 2.69 2.1 4.5 3.3 U 2.54 3.03 1.57 0.917 0.9 2.3 1.8 p_f 2.8 3.1 2.5 2.35 2.2 2.2 p_o 2.9 3.8 3.8 3 2.8 2.9 T 1390 1472 1416 1376 1343	In	0.25	0.28	0.246	0.21	0.21	0.28	0.24
III 5.50 6.62 7.52 6.16 4.5 6 4.6 Ta 5.12 6.82 3.43 2.69 2.1 4.5 3.3 U 2.54 3.03 1.57 0.917 0.9 2.3 1.8 p_f 2.8 3.1 2.5 2.35 2.2 2.2 p_o 2.9 3.8 3.8 3 2.8 2.9 T 1390 1472 1416 1376 1343	Hf	5.96	8.82	9.32	6.16	19	6	1.8
$1a$ 5.12 0.02 5.43 2.07 2.11 4.5 5.5 U 2.54 3.03 1.57 0.917 0.9 2.3 1.8 p_{-f} 2.8 3.1 2.5 2.35 2.2 2.2 p_{-0} 2.9 3.8 3.8 3 2.8 2.9 T 1390 1472 1416 1376 1343	Та	5.12	6.82	3.43	2.69) 21	4.5	+.0 3 3
p_f 2.8 3.1 2.5 2.35 1.6 p_o 2.9 3.8 3.8 3 2.8 2.9 T 1390 1472 1416 1376 1343	II II	2.12 2.54	3.02	1 57	0.017	0.0	J 2 2	1.8
p_1 2.0 3.1 2.3 2.35 2.2 2.2 p_0 2.9 3.8 3.8 3 2.8 2.9 T 1390 1472 1416 1376 1343	n f	2.54	5.05	3.1	25	2 35	2.5	2.0
T 1390 1472 1416 1376 1343	r	2.0		3.1	2.5	2.55	2.2	2.2
	Р_У Т	1390		1472	1416	1376	1343	2.)

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PUT-2	CSAM1	LU52	BARC
Putikov	Belinsky	Lucaret	Barc
	vrch		
SE	NGVF	LU	PVF
3	3	3	4
44.32	48.43	48.90	46.82
2.53	1.94	2.23	1.78
13.64	15.07	14.45	16.01
12.11	10.91	10.82	10.16
0.19	0.16	0.15	0.17
11.32	10.03	8.54	9.26
10.62	7.82	8.58	9.48
2.91	3.86	4 00	3.93
1.32	1 45	1.00	1.92
1.02	0.32	0.57	0.48
2.4	0.52	-0.4	-0.1
0.69	0.2	-0.4	0.1
11.85	6.60	8.17	16 11
-11.85	-0.00	-0.17	170
230	234 101 5766	130	215
363 216	191.3700	1//	104
210	10	144	194
24 47.9	19 51 2	10.1	20
47.8	22	41.4	41.1
27.8	32	36	43.1
617	445	464	/58
1011	684.5	627	775
262.9	152	183	185.3
82.6	38.5	51.4	51.8
23.8	19.4	22.6	23.8
6.2	3.3	4.72	6
48.1	25.8	29.5	32.8
93.4	48.9	60	63.9
11.35	5.57	7.61	7.59
46.9	22.5	31.5	34
8.65	4.6	7.12	6.04
2.49	1.48	2.23	1.86
6.86	4.42	6.28	5.14
0.83	0.71	0.94	0.7
4.91	3.92	4.92	4.35
0.8	0.7	0.94	0.83
1.99	1.91	2.38	2.26
0.27	0.27	0.31	0.34
1.6	1.65	2	2.21
0.22	0.24	0.28	0.31
5.6	3.6		4.5
4.8	2.4	2.85	3.1
2	0.8	1.35	1.7
3.35	2.55	2.75	2.15
3.6	3	3.2	2.6
1418	1386	1395	1378

2.91 2.00 1.99 2.21

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