

This is the accepted manuscript version of the paper “Klébesz, R., Grácz, Z., Szanyi, G., Liptai, N., Kovács, I., Patkó, L., Pintér, Z., Falus, G., Wertzergom, V., and Szabó, C. (2015): Constraints on the thickness and seismic properties of the lithosphere in an extensional setting (Nógrád-Gömör Volcanic Field, Northern Pannonian Basin). *Acta Geodaetica et Geophysica*, 50/2, 133-149”

The final publication is available at [link.springer.com](http://link.springer.com): <http://link.springer.com/article/10.1007/s40328-014-0094-0>

1 Constraints on the thickness and seismic properties of the lithosphere in an extensional setting (Nógrád-  
2 Gömör Volcanic Field, Northern Pannonian Basin)

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11

## 12 Abstract

13 The Nógrád-Gömör Volcanic Field (NGVF) is one of the five mantle xenolith bearing alkaline basalt  
14 locations in the Carpathian Pannonian Region. This allows us to constrain the structure and properties  
15 (e.g. composition, current deformation state, seismic anisotropy, electrical conductivity) of the upper  
16 mantle, including the lithosphere-asthenosphere boundary (LAB) using not only geophysical, but also  
17 petrologic and geochemical methods.

18 For this pilot study, eight upper mantle xenoliths have been chosen from Bárna-Nagykö, the  
19 southernmost location of the NGVF. The aim of this study is estimating the average seismic properties of  
20 the underlying mantle. Based on these estimations, the thickness of the anisotropic layer causing the  
21 observed average SKS delay time in the area was modelled considering five lineation and foliation end-  
22 member orientations. We conclude that a 142-333 km thick layer is required to explain the observed SKS

23 anisotropy, assuming seismic properties calculated by averaging the properties of the eight xenoliths. It  
24 is larger than the thickness of the lithospheric mantle. Therefore, the majority of the delay time  
25 accumulates in the sublithospheric mantle. However, it is still in question whether a single anisotropic  
26 layer, represented by the studied xenoliths, is responsible for the observed SKS anisotropy, as it is  
27 assumed beneath the Bakony-Balaton Highland Volcanic Field (Kovács et al. 2012), or the sublithospheric  
28 mantle has different layers.

29 In addition, the depths of the Moho and the lithosphere-asthenosphere boundary ( $25\pm 5$ ,  $65\pm 10$  km,  
30 respectively) were estimated based on S receiver function analyses of data from three nearby  
31 permanent seismological stations.

32 **Keywords:** seismic anisotropy, mantle xenolith, S receiver functions, lithospheric mantle, LAB, Moho

### 33 **1 Introduction**

34 Study of the subcontinental lithospheric mantle, in many cases, is only possible through different  
35 geophysical methods (e.g. seismology, magnetotellurics). However, upper mantle rocks occurring at  
36 various geodynamical settings can provide direct information about the geochemical composition and  
37 deformation history of the lithospheric mantle. Peridotite massifs can be structurally and  
38 compositionally altered, due to their obduction and subsequent long exposure time on the surface,  
39 whereas mantle xenoliths could be more representative of the upper mantle. Xenoliths can record  
40 geochemical and physical events in the lithospheric mantle, such as melting, enrichment and  
41 deformation events. In addition, different physical properties (e.g. seismic properties, electrical  
42 conductivity) of the mantle beneath a given area can be estimated based on xenoliths, and subsequently,  
43 these can be compared to geophysical data. Several examples from the literature (e.g. Baptiste and  
44 Tommasi 2014; Bascou et al. 2011; Fullea et al. 2011, 2012; Jones et al. 2013; Kovács et al. 2012) show

45 that this integrated petrologic, geochemical and geophysical approach yields to a better understanding  
46 of the structure and composition of the lithospheric mantle.

47 One way to constrain the structure of the mantle is through seismic anisotropy studies. The significance of  
48 seismic anisotropy studies lies in the relationship between deformation processes and anisotropic  
49 structures. Deformation under ductile conditions leads to the development of crystal preferred  
50 orientation (CPO) in anisotropic mantle silicates (e.g. olivine, pyroxenes). Due to the strong anisotropy of  
51 olivine and its large proportion in the mantle, development of CPO at a large scale may be responsible  
52 for seismic anisotropy in the upper mantle. Therefore, characterization of the anisotropic structure, by  
53 geophysical methods and/or xenolith studies, can provide direct information on the geodynamic  
54 processes (e.g. Long and Becker 2010).

55 In the Carpathian Pannonian region (CPR) Plio-Pleistocene alkali basalts have sampled the upper mantle  
56 at five known volcanic fields, bringing xenoliths to the surface (Fig. 1). In the past few decades, these  
57 mantle xenoliths have been extensively studied (for reviews see Dobosi et al. 2010; Szabó et al. 2004).  
58 Therefore, there is a vast body of knowledge on the composition and geochemical evolution of the  
59 lithospheric mantle beneath the CPR. Thus, recent studies have already focused on the deformation  
60 state, in addition to the geochemistry, of the lithosphere (Falus et al. 2007, 2008; Hidas et al. 2007; Liptai  
61 et al. 2013). Kovács et al. (2012) recognized that syntheses of petrologic, geochemical, geophysical and  
62 structural geological data is essential to constrain the geodynamical history of the CPR.

63 The aim of this study is to contribute to our existing knowledge of the upper mantle beneath the CPR by  
64 using both xenoliths and seismological data. The Nógrád-Gömör Volcanic Field (NGVF), has been chosen  
65 for integrated petrologic, geochemical and geophysical studies. The recent systematical sampling and  
66 several ongoing studies of xenoliths in the area, in addition to the three nearby permanent seismological  
67 stations, makes the NGVF an excellent target area.

68

## 69 **2 Geological settings**

70 The central part of the CPR is the Pannonian Basin, which is characterized by anomalously thin  
71 lithosphere. The average thickness of the crust is ~30 km, and the lithosphere-asthenosphere boundary  
72 (LAB) is located at 60-100 km in depth (Horváth et al. 2006). The NGVF is located at the northern edge of  
73 the Pannonian Basin (Fig. 1). The basement of the NGVF consists of the Gemeric and Veporic units,  
74 which consist mostly of Paleozoic and Mesozoic sequences, and sheared and tectonized crystalline  
75 nappes (Tomek 1993). In the cover sequence Tertiary sediments and volcanic rocks such as Plio-  
76 Pleistocene alkali basalts and their pyroclasts and minor Miocene andesites occur. The alkaline basaltic  
77 volcanic centers are dispersed in a NNW-SSE orientation on an approx. 1000 km<sup>2</sup> area from Podrečany  
78 (Slovakia) to Bárna (Hungary; Fig. 1).

79 The formation of the effusive rocks and pyroclasts with basanitic composition is related to post  
80 extensional thermal relaxation of the asthenosphere (Embey-Isztin et al. 1993; Harangi 2001; Szabo et al.  
81 1992). The age of the volcanism at this area is 7.2-0.4 Ma, based on K/Ar ages (Balogh et al. 1986; Hurai  
82 et al. 2013; Pécskay et al. 2006). Xenoliths are abundant in the volcanics at NGVF along the Podrečany-  
83 Bárna zone, however they are absent in the volcanics east of Fil'akovo. Xenoliths reported from the area  
84 show large compositional range, such as Cr-diopside suite of ultramafic xenoliths (Konečný et al. 1995;  
85 Szabó and Taylor 1994), clinopyroxene-rich xenoliths that are interpreted as cumulates trapped at the  
86 mantle-crust boundary (Kovács et al. 2004; Zajacz et al. 2007), and few lower crustal granulite xenoliths  
87 (Kovács and Szabó 2005). This paper focuses on peridotite xenoliths from Bárna-Nagykő location, which  
88 is in the southern part of the NGVF (Fig1).

89

## 90 **3 Sample descriptions**

91 For this study, 8 lherzolite xenoliths have been used. These xenoliths are part of a larger collection and  
92 they were chosen as representative samples for electron back scattered diffraction (EBSD) studies during  
93 previous studies (Liptai 2013; Liptai et al. 2013; and unpublished data). The main lithological and  
94 deformation characteristics of the eight xenoliths relevant for this study, including CPO patterns of  
95 olivine and pyroxenes, are summarized here. However, the interpretation of these data is beyond the  
96 scope of this paper, it will be presented elsewhere (Liptai et al., in progress).

97 The peridotites are lherzolites, composed of olivine (70-88 vol.%), orthopyroxene (5-16 vol.%),  
98 clinopyroxene (5-13 vol.%)  $\pm$  spinel ( $\leq 1$  vol.%, Table 1). Three of the xenoliths of this study (NBN032A,  
99 NBN0311, NBN0321) have porphyroclastic texture (Table 1). The porphyroclasts are dominantly  
100 orthopyroxenes with curvilinear grain boundaries and sizes between 1.0-5.3 mm, whereas the neoblasts  
101 (i.e., olivine, orthopyroxene and clinopyroxene) are usually  $\leq 0.6$  mm (Liptai et al. 2013) and have  
102 polygonal shapes with straight boundaries. The other five xenoliths (NBN035, NBN0316, NBN0319,  
103 NBN9, NBN27) have equigranular texture (Table 1), with the typical grain size  $\leq 1.2$  mm (Liptai et al. 2013;  
104 Szabó and Taylor 1994). Olivine grains often show undulose extinction and low-angle subgrain walls.  
105 Melt pockets, consisting of glass and secondary minerals (spinel and clinopyroxene), are also present;  
106 Liptai et al. (2013) interpret these features as a result of increasing temperature in the deep prior to  
107 sampling.

108 Silicate minerals in the studied samples do not show visible elongation, thus lineation and foliation is  
109 defined by shape and position of spinel grains and melt pockets. For samples where lineation and  
110 foliation is observable (NBN032A, NBN0311, NBN0321, NBN035, NBN0316, NBN0319) thin sections were  
111 prepared parallel to the xz-plane (x is parallel to the lineation and z is normal to the foliation). Since  
112 these xenoliths have olivine [100] and [010] maxima parallel to x and z, respectively, samples in which  
113 lineation and/or foliation was not observable (NBN9, NBN27) were rotated so that olivine axes match the  
114 directions mentioned above.

115 Olivine [100] and [010] axes in the three porphyroclastic and two equigranular (NBN9, NBN27) textured  
116 xenoliths have clear maxima parallel to the lineation and normal to the foliation, respectively; with [010]  
117 usually showing higher maximum densities and [100] depicting girdle-like distribution in the plane of  
118 foliation. [001] axes are generally more scattered (Fig. 2). This CPO type is recognized as A-type  
119 orientation (e.g., Jung et al. 2006; Kovács et al. 2012). In case of the other three equigranular xenoliths  
120 (NBN035, NBN0316, NBN0319) olivine [100] axes display clear maxima parallel to the lineation, whereas  
121 [010] and [001] axes distribute in a girdle in a plane normal to the lineation (Fig. 2), which is referred to  
122 as D-type CPO (Jung et al. 2006). Orthopyroxene and clinopyroxene crystal axes show scattered  
123 distribution with occasional maxima at random angles, which is attributed to overrepresentation of  
124 certain grains due to incomplete indexation during EBSD analyses resulting in recording as multiple  
125 grains with similar orientation.

126 Strength of fabric is quantified by the J-index (Bunge 1982), which is a dimensionless index ranging  
127 between 1 (for random orientation) and infinity (for single crystal). J-indices of the studied xenoliths  
128 range from 2.44 to 3.56.

129

130

#### 131 **4 Previous studies of seismic anisotropy in the CPR**

132 Seismic anisotropy is commonly studied by the measurements of shear wave splitting using SKS phases.  
133 It can constrain the orientation of the fast polarization direction, which is usually believed to be parallel  
134 to the mantle flow direction, and the strength and geometry of the anisotropic structure. However, it  
135 does not tell much about the depth distribution of anisotropy (for review, see Long and Becker 2010 and  
136 references therein).

137 Shear wave splitting studies in the CPR concentrated mostly on the western part of the Pannonian Basin  
138 and along the Carpathians. Ivan et al. (2008) estimated  $\sim 1.2$  s mean delay time of the slow wave and fast

139 directions around 135° for stations in the proximity of the Carpathians. Ivan et al. (2002) calculated 141°  
140 and 133° fastest split wave direction and 0.62 and 0.73 s delay times using two different codes at the  
141 station PSZ (GE) (Fig. 2). Dricker et al. (1999) estimated fast directions ~130° and 1.5 s in and near the  
142 Carpathians. Stuart et al. (2007) and Kovács et al. (2012) estimated delay times between 0.5 and 1.5 s  
143 and observed NW-SE and E-W anisotropy orientations in the CPR and its surrounding area (Fig. 3). In  
144 some cases (e.g. Ivan et al. 2002; Ivan et al. 2008) the source of the SKS splitting cannot be clearly  
145 identified, however due to the similarities in the anisotropy orientations observed within the CPR, and  
146 the lack of correlation between the delay times and crustal/lithospheric thickness (Dricker et al. 1999)  
147 and shear wave anisotropies calculated based on xenoliths (Kovács et al. 2012) indicate that the  
148 anisotropy is not limited only to the lithosphere, but the asthenosphere also should have a major  
149 contribution to it.

150

## 151 **5 Methods**

### 152 *5.1 S receiver function analysis*

153 S receiver function (SRF) method utilizes steeply incident teleseismic S to P converted waves to constrain  
154 the velocity structure beneath the recording site (e.g., Farra and Vinnik 2000; Yuan et al. 2006). Wave  
155 conversion occurs at velocity discontinuities such as the Moho and the LAB. The components are rotated  
156 to theoretical radial and tangential directions (LQT local ray coordinate system). As a result P waves  
157 dominate the L component, vertically polarized S waves (SV) can be found on the Q component and  
158 horizontally polarized waves (SH) on the T component. SRFs are computed by deconvolving the S  
159 waveform on the Q component from the corresponding L and T components. Individual SRFs with  
160 common piercing points are summed to improve signal-to-noise ratio. Interfaces appear as peaks on the  
161 summed SRFs, at times related to their depth.

162 Using the data of three seismological stations (Fig.3) the Moho and the LAB depths beneath the study  
163 area were investigated with SRFs. Seismograms of earthquakes with magnitude (mb) larger than 5.7  
164 occurred between 2006 and 2013 have been collected for the stations BUD, PSZ and VYHS based on the  
165 focal parameters provided by the ANSS Comprehensive Catalog  
166 (<http://earthquake.usgs.gov/earthquakes/search/>). We calculated the piercing points for each station for  
167 a reference depth of 70 km. The receiver functions belonging to the piercing points which lie near to the  
168 study area were selected and summed (Fig. 3).

169 SRF calculations (e.g. Farra and Vinnik 2000) were carried out for events with epicentral distances  
170 between 60° and 85°. A time window of 150 s in length was selected (100 s before the S wave arrival  
171 time and 50 s after it) and the data were bandpass filtered between 4 and 20 s. The ZNE components  
172 were rotated to local LQT ray coordinate system, where the rotation was performed using the  
173 theoretical backazimuth value and the incidence angle was determined by the method of Kumar et al.  
174 (2006). The individual SRFs were moveout corrected for a reference slowness of 6.4 s/° based on the  
175 IASP91 velocity model (Kennett and Engdahl 1991). In order to make the SRF comparable to the P wave  
176 receiver functions, the polarity and time axis of the SRFs were reversed, thus the positive values of the  
177 receiver functions indicate interfaces of increasing velocity with depth and vice versa.

178

## 179 *5.2 Calculation of seismic properties based on CPO measurements*

180 Seismic properties and their 3D distribution of the Bárna-Nagykő lherzolite xenoliths were calculated  
181 based on the olivine, orthopyroxene, clinopyroxene CPO, and on the modal composition (Mainprice  
182 1990). The CPO were measured by EBSD at University of Montpellier II (Montpellier, France), using a  
183 JEOL JSM-5600 scanning electron microscope equipped with an EBSD system, producing crystal  
184 orientation maps, which covered the entire thin section. Modal composition of the xenoliths was  
185 determined based on the phase maps of the thin sections obtained during the EBSD analyses. For details



186 of the EBSD data acquisition and treatment, see e.g. Falus et al. (2008). For olivine, orthopyroxene and  
187 clinopyroxene, the single crystal elastic tensors of Abramson et al. (1997), Jackson et al. (2007), and Isaak  
188 et al. (2006) at ambient conditions were used. A Voigt-Reuss-Hill averaging was applied in all  
189 calculations. The calculated seismic properties of all eight xenoliths are summarized in Table 1.

190 The average seismic properties of the mantle beneath the southern part of the NGVF were estimated by  
191 averaging the calculated elastic tensors of each xenolith. By using this approach, we assume that the  
192 orientation of the foliation and lineation is the same for all samples, therefore, it will result in an upper  
193 bound for the estimated anisotropy. By using all xenoliths with equal weight in averaging, we also  
194 assume that these lherzolite samples accurately represent the lithospheric mantle. However, considering  
195 published results (Liptai et al. 2013; Szabó and Taylor 1994), and unpublished xenoliths, we conclude  
196 that about 20% of all samples have secondary recrystallized texture that are not considered in this study  
197 due to the lack of EBSD data. The amount of uncertainty this assumption introduces, though, cannot be  
198 determined and is neglected in this study.

199

200

## 201 **6 Results**

### 202 *6.1 SRF*

203 In the resulting SRF two significant, large amplitude phases can be seen at 3.2 s and at 7.6 s (Fig. 4). The  
204 first, positive peak corresponds to the Moho, whereas the second, negative peak is related to the LAB.  
205 Based on the IASP91 model (Kennett and Engdahl 1991), the depth of the Moho and the LAB can be  
206 estimated as 25 ( $\pm$  5) km and 65 ( $\pm$  10) km, respectively. The estimated errors take into account the  
207 effects of 5% variation of seismic velocity for the IASP91 model in the crust and some additional errors  
208 due to lateral heterogeneities and noise (Mohammadi et al. 2013).

209

## 210 6.2 Seismic anisotropy

211 Olivine CPO symmetry patterns and 3D distribution of the calculated seismic properties are reported in  
212 Table 1 and in the online resource. The 3D distribution of the average seismic properties calculated  
213 based on the eight xenoliths of this study is also reported in Table 1, and shown in Fig. 5. All xenoliths,  
214 including the average, have seismic anisotropy patterns with some similar characteristics. The fastest P  
215 wave direction is always aligned with the olivine [100] axis maxima, which corresponds to the lineation.  
216 The fast S (S1) wave polarization planes all include the lineation and therefore, the projection of the S1  
217 polarisation direction on the surface will be parallel to the lineation, which marks the fossil flow  
218 direction. S1 wave velocity is minimum for waves propagating at high angles ( $\geq 60^\circ$ ) to the lineation and  
219 the foliation, and the highest at  $\sim 45^\circ$  to the lineation in the foliation plane, except in xenolith NBN0321,  
220 where it is highest normal to the lineation in the foliation plane. The  $V_p/V_{s1}$  ratios are the highest for  
221 waves propagating at low angle ( $\leq 30^\circ$ ) to the lineation and low for waves propagating at high angles  
222 ( $\geq 45^\circ$ ) to the lineation. In some cases the minimum can be identified for propagation directions normal  
223 or close to normal to both the lineation and the foliation. The  $V_p/V_{s2}$  ratios are the highest for waves  
224 propagating in the foliation plane, in most cases a clear maximum is observed parallel to the lineation  
225 and low velocities at  $\geq 60^\circ$  to the foliation.

226 The changes in olivine CPO symmetry cause small variations in the seismic anisotropy patterns. The  
227 lherzolite xenoliths displaying A-type olivine CPO pattern and xenolith NBN035 show a distinct  $V_p$   
228 minimum at normal or close to normal to the foliation. S wave anisotropy is minimum at  $45^\circ$  to both the  
229 lineation and the foliation and it is the highest at  $\sim 45^\circ$  to the lineation in the foliation plane. S2 wave  
230 velocity in case of xenoliths displaying A-type fabric is minimum for waves that are propagating normal  
231 to the lineation in and perpendicular to the foliation, and the highest at  $\sim 45^\circ$  to the lineation and the  
232 foliation plane, except in xenolith NBN032A.

233 Xenoliths NBN 0316 and NBN0319 (D-type) do not show a clear  $V_p$  minimum, P wave velocity is low at  
234 every direction normal to the lineation. In case of all D-type xenoliths and NBN032A, S wave anisotropy is  
235 minimum for waves propagating at low angle to the lineation and the highest normal to the lineation in  
236 the foliation plane. S2 wave velocity is low for waves propagating normal or close to normal to the  
237 lineation and the highest at low angles ( $\leq 30^\circ$ ) to the lineation.

238 The average sample show patterns intermediate between the two types described above. The P wave  
239 velocity pattern is similar to those with A-type olivine CPO symmetry, in contrast S2 wave velocity  
240 pattern resembles to those with D-type olivine CPO symmetry. S wave anisotropy is minimum at  $45^\circ$  to  
241 both the lineation and the foliation, like for xenoliths with A-type symmetry, but it is highest normal to  
242 the lineation in the foliation plane as it is for xenoliths with D-type symmetry.

243 No wide variations in seismic properties of the individual samples were observed, the main variation is in  
244 the intensity of the anisotropy (Table 1). A clear linear correlation between the J-index and the maximum  
245 anisotropies (P wave and S1 wave anisotropy and S wave splitting) was not recognized, xenoliths with  
246 higher J-index, though, tend to have stronger anisotropy (Table 1).

247

248

## 249 **7 Discussion**

### 250 *7.1 Estimated Moho and LAB depths*

251 The crustal and lithospheric thickness maps of the CPR constructed by Horváth et al. (2006) is generally  
252 accepted, therefore we compared our results to these maps in order to evaluate them. These maps are  
253 the improved versions of the maps of Horváth (1993). The crustal thickness map is based on data  
254 obtained by traditional refraction methods, reflection seismic profiling and gravity modelling studies  
255 (Horváth 1993, 2006). Based on these maps the estimated depth of the Moho beneath the NGVF is  $\sim 27.5$

256 km, which is in agreement with the results of recent deep crustal seismic profiling (Grad et al. 2006;  
257 Hrubcová et al. 2010; Tomek 1993) and the result of this study within the estimated error.

258 The crustal thickness is well constrained in the CPR, however, significantly less data is available on the  
259 lithospheric thickness. The lithospheric thickness map of Horváth (1993) was constructed based on P  
260 wave travel time residuals and magnetotelluric soundings, and it estimates the LAB at 60-80 km in depth  
261 beneath the NGVF. However, the lithospheric thickness map of Horváth et al. (2006), which was  
262 improved by new magnetotelluric results, estimates the LAB at greater depth (80-100 km) beneath the  
263 study area. Our result ( $65 \pm 10$  km) is in good agreement with the result of Horváth (1993), however it  
264 indicates a shallower LAB than the currently generally accepted values of Horváth et al. (2006).

265 Plomerová and Babuška (2010) presented a uniform updated model of the European LAB based on P  
266 wave residuals. They predict an even deeper LAB beneath the NGVF, at least  $100 (\pm 10)$  km. Geissler et  
267 al. (2010) used SRF obtained at 78 European permanent broad-band stations to estimate the thickness of  
268 the European lithosphere. Most of the Sp piercing points for an 80 km deep LAB were located  $\sim 80$  km N-  
269 E from the stations in the study of Geissler et al. (2010). The NGVF is in the proximity of the Sp piercing  
270 points of the BUD station, therefore we assume that the BUD station can be used for comparison.

271 Geissler et al. (2010) estimated 74 km for LAB depth and 28 km for Moho depths based on the data  
272 obtained at the BUD station, which is in good agreement with our result within the estimated errors.

273 Jones et al. (2010) pointed out that there can be significant differences in the estimated LAB depths  
274 based on the method used (i.e. magnetotellurics, SRF, analysis of P travel time residuals). Our results are  
275 similar to those obtained by the same method, i.e. SRF (Geissler et al. 2010). The differences between  
276 our estimated depths and the previously published lithospheric maps (Horváth 1993, 2006; Plomerová  
277 and Babuška 2010) might only be due to the used methods and data based on which the maps were  
278 constructed.

279 Another possible way of evaluating the estimated depths is to compare them to the estimated  
280 originating depths of the xenoliths from the NGVF. The originating depth of the xenoliths can be  
281 estimated based on the calculated equilibrium temperature and the appropriate heat flow values of the  
282 area, with an uncertainty of  $\pm 12$  km (Kovács et al. 2012). Liptai et al. (in progress) estimated that the  
283 xenoliths from Bárna-Nagykő are from 30-35 km, whereas xenoliths from the central part (Babi Hill and  
284 Medves Plateau) of the NGVF (Fig 1.) are from  $\sim 40$ -50 km. However, they argue, that incipient melting  
285 preserved in the xenoliths might have caused a compositional change in the pyroxene. Consequently, the  
286 estimated equilibrium temperatures, and hence the originating depth, can only be considered as a  
287 minimum estimate. The originating depth range of the xenoliths, considering also the uncertainty of the  
288 estimation and possible underestimation, is within our estimated range of the lithospheric mantle  
289 (between  $25\pm 5$  and  $65\pm 10$  km in depth).

290

## 291 *7.2 Seismic anisotropy*

292 The original, in-situ orientation of the xenoliths is unknown due to their transport to the surface,  
293 therefore we are unable to constrain the orientation of the foliation and lineation in the anisotropic  
294 layer. However, we are able to estimate the thickness of the anisotropic structure that could cause the  
295 observed delay times (e.g. Baptiste and Tommasi 2014; Ben-Ismaïl et al. 2001; Kovács et al. 2012;  
296 Michibayashi et al. 2006; Pera et al. 2003). The thickness (T) of an anisotropic layer is given by equation  
297 (1), where dt is the delay time of S waves,  $\langle V_s \rangle$  is the average velocity of the fast and slow velocities,  
298 and AVs is the anisotropy for a specific propagation direction expressed as a percentage (e.g. Pera et al.  
299 2003).

300

$$T = 100dt \cdot \frac{\langle V_s \rangle}{AV_s} \quad (1)$$

301 Ignoring the potential effect of crustal anisotropy (such as in e.g. Kovács et al. 2012), the calculated  
302 seismic properties of the average mantle beneath Bárna-Nagykő (Fig. 4 and Table 1) was used to  
303 estimate the thickness of the anisotropic layer. Five end-member orientations (Fig. 6) were considered  
304 for these estimations, following the example of Baptiste and Tommasi (2014), such as horizontal foliation  
305 and lineation (case 1), vertical foliation but horizontal lineation (case 2), vertical foliation and lineation  
306 (case 3), 45° dipping foliation and lineation (case 4), 45° dipping foliation and horizontal lineation (case  
307 5). Model calculations were carried out for two different scenarios with two different dt for all five end-  
308 member orientations. In the first scenario, we assumed  $dt \sim 1.1$  s by considering all observations of the  
309 CPR (Dricker et al. 1999; Ivan et al. 2002, 2008; Kovács et al. 2012; Stuart et al. 2007), in the second we  
310 assumed  $dt \sim 1.3$  s based on the observed delay times at the three seismological stations closest to the  
311 NGVF (data from: Pizskéstető - Ivan et al. 2002; and Pizskéstető, Bükk Mts., Central Slovakia in Fig. 2;  
312 data from Kovács et al. 2012).

313 For the five end-members the S wave polarization anisotropy is 1.88, 3.74, 2.0, 0.5 and 2.75 % for case 1  
314 to 5, respectively. The average S wave velocity is 4.82 km/s. Hence, the estimated thickness of the  
315 anisotropic structure is 282, 142, 265, 1061 and 193 km, and 333, 168, 313, 1254 and 228 km for case 1  
316 to 5, considering  $dt = 1.1$  s then 1.3 s, respectively (Fig. 6). Case 1 and 3 gives similar results, whereas case  
317 2 and 5 requires thinner, case 4 requires much thicker anisotropic layer to produce the same delay time.  
318 Based on global datasets, depth dependent of anisotropy in the upper mantle was recognized (Kustowski  
319 et al. 2008; Long and Becker 2010; Wenk 2004). The global average upper mantle anisotropy is  
320 significant in the upper  $\sim 200$ -250 km, and then gradually weakens between 250-400 km (Kustowski et al.  
321 2008; Long and Becker 2010). Therefore, assuming that a single anisotropic layer causes the observed  
322 delay time, we can conclude that only case 4, a layer with foliation and lineation close to 45°, is unlikely  
323 considering the global average upper mantle anisotropy. Without further seismic evidence, however, it is  
324 not possible to determine the orientation of the lineation and the foliation, but in the other four cases a

325 142-282 or 168-333 km thick layer is needed to produce the observed delay times. These thicknesses are  
326 significantly greater than the estimated lithospheric mantle thickness (~40 km), therefore at least ~100  
327 km thick sublithospheric mantle is required with the same structure to account for the seismic  
328 observations.

329 Our results are compared to the thickness of the anisotropic layer beneath the Bakony-Balaton Highland  
330 Volcanic Field (BBHVF), in order to see if there is any resemblance between the NGVF and the BBHVF,  
331 which is the closest area where similar studies has been carried out (Kovács et al. 2012), and the BBHVF  
332 is also the part of the same tectonic unit (ALCAPA). The thickness of the anisotropic layer beneath the  
333 BBHVF was recalculated by using equation (1) based on the A-type xenolith reported by Kovács et al.  
334 (2012), considering the five end-member orientations described above. In the calculations, ~1 s surface  
335 delay time was assumed, which was measured at the proximity of BBHVF sites (Kovács et al. 2012). The  
336 thickness of anisotropic layer is 123, 85, 196, 140, 109 km for cases 1 to 5, respectively. The minimum  
337 thickness (85 km) is observed in case of vertical foliation and horizontal lineation (case 2), similarly to the  
338 estimates for the NGVF. The thickness of the sublithospheric part of the anisotropic layer is ~60-160 km,  
339 assuming a ~35 km thick lithospheric mantle, which was estimated beneath the BBHVF (Kovács et al.  
340 2012 and references therein). This thickness (~60-160 km) is considerably smaller than the thickness of  
341 the sublithospheric part of the anisotropic layer beneath the NGVF (~100-240 or ~130-290 km), assuming  
342 a single anisotropic layer.

343 In the BBHVF there is compelling evidence that the A-type xenoliths derive from the upper part of a  
344 mantle domain which represents asthenospheric material, lithospherized after the Miocene extension  
345 (Kovács et al. 2012). Therefore it is reasonable to assume that a single anisotropic layer, sampled by the  
346 A-type xenoliths, is responsible for the observed SKS delay times. However, based on the data presented  
347 here, it is not yet possible to assess whether the same scenario is true for the NGVF. Consequently, as a  
348 second approach, we assumed that the calculated average mantle represents only the lithospheric

349 mantle. This allow us to calculate, by rearranging equation (1), the portion of the delay time that  
350 accumulates in the lithospheric mantle. Based on the results of this study and previous data from the  
351 literature, 25 km thick crust and 55 km thick lithospheric mantle was assumed. As a results, 0.2 s, 0.4 s,  
352 0.2 s, 0.1 s and 0.3 s was calculated for cases 1 to 5, respectively. Even considering the upper limit of the  
353 crustal contribution to the delay time, which is typically 0.1 s per 10 km (Barruol and Mainprice 1993), it  
354 has become evident that the majority ( $\geq 50\%$ ) of the delay time may accumulate in the sublithospheric  
355 mantle. Further geochemical and deformation studies will be carried out, which might help us constrain  
356 the origin of the lithospheric mantle and its connection with the sublithospheric mantle, and hence the  
357 origin of the observed anisotropy.

358

## 359 **8 Summary**

360 The NGVF proved to be an excellent area for integrated geophysical and petrologic and geochemical  
361 studies. Data from three nearby seismological stations could be used to estimate the depths of the Moho  
362 and the LAB ( $25\pm 5$ ,  $65\pm 10$  km, respectively). Mantle xenoliths can be a powerful tool for estimating the  
363 3D distribution of the seismic properties. Relying on these estimated properties and the published SKS  
364 delay times, the thickness of the anisotropic structure beneath the NGVF was constrained based on eight  
365 mantle xenoliths from the southernmost location, Bárna-Nagykő. The thickness of a single anisotropic  
366 structure was estimated at least  $\sim 140$  km and maximum  $\sim 330$  km. The thickness of the anisotropic  
367 structure in case of each foliation and lineation orientations is larger beneath the NGVF than under the  
368 BBHVF. This could indicate differences in the anisotropic structures beneath the two area. At this point  
369 there is not enough evidence to assume that the delay time accumulates in a single anisotropic layer  
370 beneath the NGVF. However, we can conclude that the majority of the delay time accumulates in the  
371 sublithospheric mantle. Geochemical studies in the future may give constraints on the link between the  
372 lithospheric and sublithospheric mantle, and hence the possible source of the SKS anisotropy.



373

374

### 375 **Acknowledgements**

376 The authors thank V. Baptiste for helpful discussion. We are grateful for the thorough review and  
377 constructive comments of K. Hidas and an anonymous reviewer. This research was carried out in the  
378 framework of the cooperation agreement (TTK/6109/1/2014 and Sz/156/2014) between the Lithosphere  
379 Fluid Research Lab at Department of Petrology and Geochemistry of Eötvös University and the Geodetic  
380 and Geophysical Institute of the MTA Research Centre for Astronomy and Earth Sciences. This study was  
381 partially supported by the TAMOP-4.2.2.C-11/1/KONV-2012-0015 (Earth-system) project sponsored by  
382 the EU and European Social Foundation. IK was supported by the Bolyai Postdoctoral Fellowship Program  
383 and a Marie Curie International Reintegration Grant (NAMS-230937).

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523

## 524 **Figure captions**

525 **Fig. 1 a)** Location of the mantle xenolith-bearing alkali basalt localities in the CPR (SBVF – Styrian Basin  
526 Volcanic Field; BBHVF – Bakony-Balaton Highland Volcanic Field; LHPVF – Little Hungarian Plain Volcanic  
527 Field; NGVF – Nógrád-Gömör Volcanic Field; PMVF – Persány Mountains Volcanic Field) **b)** Location of  
528 the upper mantle xenoliths, including the southernmost site, Bárna-Nagykő area within the NGVF

529 **Fig. 2** Pole figures of typical A- and D-type olivine CPO in the studied xenoliths, pictured in lower  
530 hemisphere equal area projections. Contours are 0.5 multiples of uniform distribution, lowest value  
531 contour is marked with a dashed line. Black square and white circle represent maximum and minimum  
532 axis densities, respectively, and n stands for the number of measured grains.

533 **Fig. 3** Location of the seismological stations (triangles) used in this study and the distribution of piercing  
534 points (crosses) of S receiver functions for 70 km depth. The color of the piercing points indicates the  
535 corresponding station. Operating organizations of the seismological stations: BUD - MTA CSFK Geodetic  
536 and Geophysical Institute, Hungary; PSZ - GEOFON Global Seismic Network, GFZ, Germany & MTA CSFK  
537 Geodetic and Geophysical Institute, Hungary; VYHS - Geophysical Institute, Slovak Academy of Sciences,  
538 Slovakia. Magnitude and direction of fast polarization direction of the near vertically propagating SKS  
539 phase are also shown. Data from i) red – Dando et al. 2011; Kovács et al. 2012; Stuart et al. 2007; ii)  
540 yellow - Kovács et al. 2012; Stuart et al. 2007; iii) blue – Ivan et al. 2002

541 **Fig. 4** Individual moveout corrected S receiver functions (lower panel) belonging to the piercing points  
542 displayed in Fig. 2 and their sum (upper panel). Two significant peaks (a positive and a negative) can be  
543 clearly observed. They correspond to the S-to-P conversions at the Moho and at the LAB, respectively

544 **Fig. 5** Seismic properties of the average sample obtained by averaging the elastic tensors of the 8 studied  
545 peridotite xenoliths. From left to right and top to bottom, schematic representation of the lineation and  
546 foliation reference frame used in this study, variation as a function of the propagation direction of the P  
547 wave velocities ( $V_p$  in km/s), of the shear wave polarization anisotropy ( $AVs$  in % =  $200 \times (V_{s1} - V_{s2}) /$   
548  $(V_{s1} + V_{s2})$ ), of the polarization of the fast shear wave  $S_1$  (coloring represent the intensity of  $AVs$ , as in  
549 the previous plot), of the two quasi-shear waves ( $V_{s1}$  and  $V_{s2}$ ) velocities, and of the  $V_p/V_{s1}$  and  $V_p/V_{s2}$   
550 ratios. Lower hemisphere stereographic projections

551 **Fig. 6** Calculated SKS anisotropy for the five different end-member orientations of the foliation and the  
552 lineation: (case 1) horizontal foliation and lineation, (case 2) vertical foliation with a horizontal lineation,  
553 (case 3) vertical foliation and lineation, (case 4) 45° dipping foliation and lineation, and (case 5) 45°  
554 dipping foliation with a horizontal lineation, after Baptiste and Tommasi 2014. Estimated thickness ( $T$ ) of



555 the anisotropic layer in case of  $dt=1.1$  s and  $dt=1.3$  s, and in case of Bakony-Balaton Highland Volcanic  
556 Field (BBHVF) (assuming  $dt=1$  s)

557 **Table**

558 **Table 1** Texture, modal composition (ol – olivine, opx – orthopyroxene, cpx- clinopyroxene, sp – spinel),  
559 rock type, CPO symmetry type, J index and seismic properties ( $V_p$  – P wave velocity,  $A_V$ s – shear waves  
560 polarization anisotropy,  $V_{s1}$  – velocity of the faster shear wave,  $V_{s2}$  – velocity of the slower shear wave,  
561  $dV_s$  – difference of the faster and slower shear wave,  $V_p/V_{s1}$  – ratio of the velocities of the P wave and  
562 the slower shear wave,  $V_p/V_{s2}$  – ratio of the velocities of the P wave and the faster shear wave) of the 8  
563 peridotite xenoliths from the study area (Bárna-Nagykő, Nógrád-Gömör Volcanic Field)

564

565