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SPECTROSCOPIC ANALYSIS OF POTENTIAL ASTROMOLECULES  
VIA QUANTUM CHEMICAL QUARTIC FORCE FIELDS

By  
Mason Bradford Gardner

A thesis submitted to the faculty of The University of Mississippi in partial fulfillment of the requirements of the Sally McDonnell Barksdale Honors College.

Oxford, MS  
May 2021

Approved By

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## ABSTRACT

### MASON BRADFORD GARDNER: Spectroscopic Analysis of Potential Astromolecules Via Quantum Chemical Quartic Force Fields

Astrochemistry has been substantially aided by computational techniques, particularly through the use of Quartic Force Field (QFF) analysis. Several methods have proven useful at correlating computed spectroscopic data with experimental observations. The F12-TZ QFF correlated well with experimental data for silicon oxide compounds, particularly those potentially involved in development from rocky bodies to planetary masses <sup>[27]</sup>. Compared to argon matrix experimental data, the vibrational frequencies for the molecules SiO<sub>2</sub>, SiO<sub>3</sub>, Si<sub>2</sub>O<sub>3</sub>, and Si<sub>2</sub>O<sub>4</sub> become less accurate as the complexity of the molecules increases but should still be predictive of infrared characteristics of silicon oxides as they form clusters in space <sup>[27]</sup>. The CcCR QFF was found to be accurate for predicting B<sub>0</sub> and C<sub>0</sub> rotational constants within 35 MHz and vibrational frequencies within 5.7 cm<sup>-1</sup> for many molecules, including those with heavy atoms <sup>[21]</sup>. When used in conjunction, the F12-TZ and CcCR QFFs produced parallel data for predicting the brightest vibrational frequencies in relatively complex molecules containing noble gases; rotational constants produced by the CcCR QFFs also present evidence for future identification of such molecules.

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## LIST OF ABBREVIATIONS

SOFIA	Stratospheric Observatory for Infrared Astronomy
QFF	quartic force field
VPT2	second-order vibrational perturbation theory
CCSD(T)	coupled cluster theory at the singles, doubles, and perturbative triples level
F12-TZ	CCSD(T)-F12b/cc-pVTZ-F12
CcCR	CCSD(T)/CBS with core correlation and scalar relativity corrections
CBS	complete basis set extrapolation
MT	Martin-Taylor core correlating basis set
DK	Douglas-Kroll Hamiltonian
DFT	density functional theory analysis
JWST	James Webb Space Telescope
ZPVEs	zero-point vibrational energies
MAE	mean absolute error
MA%E	mean absolute percent errors



## CHAPTER I: QUANTUM CHEMICAL ASTROCHEMISTRY

### A. Introduction

Astrochemistry represents one of the widest fields of science encompassing a myriad of interrelated fields, including planetary geology, interstellar biology, and astronomy, not to mention the various subdisciplines of chemistry. Astrochemistry is directly concerned with the origin, abundance, and identity of molecules interacting in interstellar bodies, such as planets, stars, nebulae, etc. Originally, gas-phase experiments were conducted to compare observations between these “ground-based” experiments and spectroscopic data to determine the identity of astronomical molecules <sup>[1]</sup>; these determinations allow for conclusions and further research into chemical interactions in space and provide clues to astronomers as to the behavior of matter in these objects. Gas-phase experiments are a vital part of providing reliable and conclusive evidence about astrochemical data, but computational techniques offer solutions that expand the reach of astrochemistry <sup>[1]</sup>.

Since the introduction of computational methods in chemistry, the abilities of astrochemists have grown substantially. By using quantum chemical theory to computationally design molecules and calculate their observables, a parallel option to gas-phase experiments is readily available and effective based on accuracy between theoretical calculations and physical experiment <sup>[1]</sup> with examples in Chapters II and III. As spectroscopy is the most dependable form of identifying molecules in space and the only available for remote sensing, the computational methods in this thesis are used to determine the spectroscopic data that would be observed under laboratory observations, specifically vibrational frequencies and rotational constants.

## B. Spectroscopic Data

Vibrational spectroscopy measures the energy difference between quanta of molecular vibrational modes and are typically measured using infrared spectroscopy. Harmonic vibrational frequencies represent the energy differences that would be observed from molecules acting as perfect springs. However, molecules are anharmonic, and their energy levels vary significantly from harmonic modes. Specifically, anharmonic frequencies are usually lower than harmonic frequencies because the anharmonic, or Morse, behavior becomes asymptotic as bond length increases. Anharmonic vibrational frequencies are observed in spectroscopy, unlike harmonic vibrational frequencies. Hence, computational astrochemists must produce anharmonic vibrational frequencies for direct comparison to astronomical data [2].

Vibrational frequencies also have intensities, or brightness, if they have a change in dipole moment due to the vibration; these are infrared active frequencies. The intensity is also an observable characteristic of detectable astromolecules because it shows the activity levels of anharmonic vibrational modes, not harmonic modes [1] [3]. This is particularly noteworthy for the work reported in Chapter IV because of the high intensities exhibited by the noble gas structures.

Similarly, rotational constants affect the difference in energy between rotational states and are most often measured using microwave spectroscopy. These modes are just as unique as vibrational frequencies, and analyzing both of these observables makes astrochemical identification and analysis much more accurate than using either alone. Molecules have a rotational constant for each of the three principle axes—A, B, C. While linear molecules only have one rotational constant, noted as B, due to two degenerate and one infinite coordinate along the bonding plane. Rotational frequencies also follow a constant pattern based on the angular momentum while the step size between vibrational modes decreases with higher modes, an effect

of the anharmonic potential. In this way, rotational modes are easier to predict than vibrational modes [2].

Rotational and vibrational data can be used in conjunction as well. For example, vibrational frequencies are observed using infrared radiation, which is hindered by our planet's atmosphere; however, the microwave radiation used to observe rotational spectra are not hindered by the atmosphere. Therefore, rotational spectroscopy can be used first to determine where to direct the vibrational spectroscopes, improving the accuracy and decreasing the cost and errors occurring during vibrational studies. NASA uses this in its Stratospheric Observatory for Infrared Astronomy (SOFIA); SOFIA is an infrared spectroscope that is flown high into the sky in order to reduce the atmosphere's inhibition and to point the spectroscope in the direction previously determined by rotational spectroscopy [4].

### C. Quartic Force Fields: F12-TZ and CcCR

Quartic force field (QFF) analysis involves a fourth-order Taylor series expansion of a molecule's equilibrium geometry to produce a potential energy surface, represented by the equation below. The “ $F$ ” terms are force constants and “ $\Delta$ ” terms are displacements of the bond lengths by 0.005 Å and of the angles by 0.005 radians, respectively, within the molecule's optimized geometry:

$$V = \frac{1}{2} \left( \sum_{ij} F_{ij} \Delta_i \Delta_j \right) + \frac{1}{6} \left( \sum_{ijk} F_{ijk} \Delta_i \Delta_j \Delta_k \right) + \frac{1}{24} \left( \sum_{ijkl} F_{ijkl} \Delta_i \Delta_j \Delta_k \Delta_l \right). \quad (1)$$

This equation generates a potential energy surface of molecular displacements to find the minimum displaced geometry [5] [6]. Second-order vibrational perturbation theory (VPT2) uses the potential surface and force constants to calculate observable data about specific astromolecules, including the vibrational frequencies and rotational constants. QFFs have been

established as a reasonable method of determining these observables of molecules while maintaining a relatively low computational cost and high level of accuracy <sup>[7]</sup>. There are several available levels of theory to compute the desired spectroscopic data, and each has its own tradeoffs between cost and accuracy. The research in this thesis conducts QFFs based on coupled cluster theory at the singles, doubles, and perturbative triples (CCSD(T)) level <sup>[8][9]</sup>. CCSD(T) QFFs follow a common procedure:

- Step 1. Optimize molecular geometry
- Step 2. Generate the displacements that define the QFF
- Step 3. Calculate energies from displacements
- Step 4. Fit potential energy surface
- Step 5. Calculate anharmonic spectroscopic data
- Step 6. Correct for resonances

Two kinds of QFFs are used in these projects: F12-TZ and CcCR, which are defined by their corresponding levels of theory discussed below. Chapter II discusses the use of F12-TZ QFFs in the study of silicon oxide compounds while Chapters III and IV discuss the use of CcCR QFFs in rotational constants and noble gas molecules, respectively.

Once a geometry for any designated molecule has been constructed and optimized, the computations can proceed. This is the same starting point for any QFF, and the level of theory is where the F12-TZ and CcCR QFFs begin diverging in technique <sup>[5]</sup>.

- Step 1. Optimize molecular geometry

The F12-TZ QFF relies on the “CCSD(T)-F12b/cc-pVTZ-F12” basis set, which is an F12 explicit electron correlation function with a triple zeta level basis set <sup>[10][11]</sup>. Specifically, the geometry optimization uses the CCSD(T)-F12b/cc-pVTZ-F12 basis set <sup>[10][11]</sup> in the quantum chemistry software package MOLPRO <sup>[12][13]</sup>, which is how this method gets its abbreviated name of “F12-TZ”.

In contrast, CcCR QFFs optimize the geometry using canonical CCSD(T) with the aug-cc-pV5Z and Martin-Taylor (MT) basis sets in MOLPRO, whose complete descriptions are available in Table 1 below:

Table I.1: CcCR Geometry Optimization Basis Sets

<b>Basis Set</b>	<b>Abbreviation</b>
CCSD(T)/aug-cc-pV5Z	5Z
CCSD(T)/cc-pVTZ-MT with core electrons	MT <sub>c</sub>
CCSD(T)/cc-pVTZ-MT without core electrons	MT

The Martin-Taylor basis sets are used to correct the impact that core electrons have on the geometry optimization <sup>[14]</sup>. The composite geometry is constructed by combining the 5Z geometry with the effect of the core electrons, found by the difference between MT<sub>c</sub> and MT.

$$Composite = 5Z + MT_c - MT. \quad (2)$$

Step 2. Generate the displacements that define the QFF

Once the bond angles and bond lengths of any designated molecule are optimized, every bond length and bond angle are subjected to individual displacements in every directional place, by 0.005 Å for the bond lengths and by 0.005 radians for the bond angles. Each molecular structure will have a different total number of displacements reflecting the number of atoms and bonds. These displacements are collected and converted into Cartesian coordinates for use in later programs.

Step 3. Calculate energies from displacements

Each displacement increases the molecule's potential energy, which can be calculated using various basis sets. The energy displacements in F12-TZ also use the "CCSD(T)-F12b/cc-pVTZ-F12" basis set <sup>[10][11]</sup>. These energy displacements are calculated in MOLPRO <sup>[12][13]</sup>.

In the CcCR method, the composite geometry from above is used to generate seven sets of energy displacements, including a complete basis set extrapolation (CBS), Martin-Taylor basis set (MT) to correct for core electrons, and the Douglas-Kroll Hamiltonian (DK) to correct for scalar relativity<sup>[14][15][16]</sup>. The full basis sets are shown in Table 2 below. The CcCR name is representative of these energy displacements. The CBS energy is composed of 5Z, QZ, and TZ basis sets; the MT basis sets correct for core electrons; the DK basis sets correct for scalar relativity. This CCSD(T) computation can be summarized as CCSD(T)/CBS (“C”) with core correlation corrections (“cC”) and scalar relativity corrections (“R”), or “CcCR” when abbreviated all together<sup>[17]</sup>.

Table I.2: CcCR Energy Displacements Basis Sets

<b>Basis Set</b>	<b>Abbreviation</b>
CCSD(T)/aug-cc-pV5Z	5Z
CCSD(T)/aug-cc-pVQZ	QZ
CCSD(T)/aug-cc-pVTZ	TZ
CCSD(T)/cc-pVTZ-MT with core electrons	MT <sub>c</sub>
CCSD(T)/cc-pVTZ-MT without core electrons	MT
CCSD(T)/cc-pVTZ-DK with scalar relativity	DK <sub>r</sub>
CCSD(T)/cc-pVTZ-DK without scalar relativity	DK

The energies from these calculations are recombined using Equation 3 below, where " $E$ " represents the energy, " $l$ " represents the angular momentum, and the exponential terms reflect the Taylor series expansion. This equation completes the CBS extrapolation, and is the best extrapolation scheme currently available<sup>[17]</sup>.

$$E(l) = \left( 5Z + QZ \left( l + \frac{1}{2} \right)^{-4} + TZ \left( l + \frac{1}{2} \right)^{-6} \right) + (MT_c - MT) + (DK_r - DK). \quad (3)$$

While these two methods follow a similar QFF methodology to determine spectroscopic data for these molecules, F12-TZ QFFs have a smaller computational cost based on the use of

explicit correlation and are quicker, while CcCR QFFs have a higher computational cost because they account for other molecular properties and include more levels of theory <sup>[18] [19] [20] [21]</sup>.

#### Step 4. Fit potential energy surface

While the geometry optimization and energy displacements require different basis sets for F12-TZ and CcCR, the remaining steps are performed the same way. The single set of F12-TZ displacement energies are turned into relative energies based on the equilibrium energy, and these relative energies are then used to generate a potential energy surface in ANPASS using a least-squares procedure. The potential energy surface yields a new equilibrium geometry, and the force constants are found by refitting the potential to this geometry in ANPASS. Similarly, the components of the CcCR displacement energies are combined using Equation 3, made relative to the lowest combined energy, and fed into ANPASS in the same way described above <sup>[22]</sup>.

#### Step 5. Calculate anharmonic spectroscopic data

The force constants and new equilibrium geometry are then collected for use by INTDER <sup>[23]</sup> and SPECTRO <sup>[24]</sup>. INTDER converts the internal coordinate force constants to Cartesian coordinates for use in SPECTRO <sup>[23]</sup>. SPECTRO completes the VPT2 and generates the observable harmonic and corrected anharmonic spectroscopic data <sup>[24] [25]</sup>.

#### Step 6. Correct for resonances

This “raw” computed data is then used by SPECTRO again to correct for degenerative modes, Coriolis resonances, Fermi resonances, Darling-Dennison resonances, and potential Fermi-resonance polyads <sup>[26]</sup>. The results from this step produce the harmonic and corrected anharmonic spectroscopic data desired.

## D. Applications

Both F12-TZ and CcCR QFFs can be applied to various astromolecules. As discussed in Chapter II, the F12-TZ QFF produces vibrational frequencies of silicon oxide compounds that were consistently within  $7\text{ cm}^{-1}$  of argon-matrix experimentally observed frequencies <sup>[27]</sup>. Chapter III discusses the CcCR method, showing that this method accurately produced vibrational frequencies to within an average  $5.8\text{ cm}^{-1}$  of experimental values for several small molecules known to be present in interstellar media <sup>[21]</sup>. In addition, CcCR rotational constants for the least anharmonic molecules are found to be accurate to within 34 MHz of experiment, and this accuracy is within the limit for directly observing these molecules in space <sup>[21]</sup>. Given the successful implementation of these QFFs in Chapters II and III, they are both utilized in Chapter IV. Molecules in Chapter IV are currently only theorized to exist as spectroscopic data has not yet been obtained for these molecules, but CcCR and F12-TZ produced similar data for these specific noble gas molecules, as discussed in Chapter IV.

Accuracies within  $7\text{ cm}^{-1}$  of vibrational frequencies and 34 MHz of rotational constants support the idea that both the F12-TZ QFF and CcCR QFF blend speed and accuracy well for effectively predicting and comparing spectroscopic data to observations, while F12-TZ is slightly less accurate regarding rotational data <sup>[21][27]</sup>. As interest in exotic and yet-to-be-observed interstellar molecules increases, many of these molecules will likely be difficult to synthesize for gas-phase experimental observation, making computational methods the best means to continue to generate highly accurate data for comparison to astronomical observation and exploration. By using computational methods in conjunction with physical experiments, the time and cost of arduous laboratory methods can be saved <sup>[1]</sup>.



This research also has direct implications on knowledge of the universe. By comparing computational data to spectroscopic observations, astrochemists can make determinations about many interstellar bodies <sup>[1]</sup>. In Chapter II, silicon oxide molecules are the focal point because they are found in rocky bodies in space and form distinct clusters on the way to becoming larger bodies, such as comets and planets <sup>[27]</sup>. In Chapter III, the accuracy of rotational constants in CcCR QFFs for several small molecules are highlighted due to the inherent importance of and predictability of rotational constants in spectroscopy, showing that CcCR is effective and produces results similar to those of gas-phase experiments <sup>[21]</sup>. Chapter IV concerns the application of F12-TZ and CcCR QFFs to theoretical molecules with noble gas bonding, which could have implications on further research into noble gas bonding in new molecules. These methods can be used in many other studies, such as determining the age of stars, development of planets, and molecular composition of nebulae. Unifying computational astrochemistry with experimental spectroscopy can offer deeper insights in the chemistry of interstellar space and increase the imaginative potential of new molecules yet to be discovered <sup>[1]</sup>.

## CHAPTER II: SILICON OXIDES

Reprinted from <sup>[27]</sup>

Gardner, M. B.; Westbrook, B. R.; Fortenberry, R. C. Spectral Characterization for Small Clusters of Silicon and Oxygen: SiO<sub>2</sub>, SiO<sub>3</sub>, Si<sub>2</sub>O<sub>3</sub>, & Si<sub>2</sub>O<sub>4</sub>. *Planetary and Space Science* 2020, 193, 105076.

1 Spectral Characterization for Small Clusters of Silicon and  
2 Oxygen: SiO<sub>2</sub>, SiO<sub>3</sub>, Si<sub>2</sub>O<sub>3</sub>, & Si<sub>2</sub>O<sub>4</sub>

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9 **Abstract**

10 Rocky bodies are made of minerals comprised largely of silicon and oxygen. How these  
11 minerals are formed from their constituent atoms is not fully known. The rovibrational  
12 IR spectral data produced in this work may help to observe small molecules containing  
13 silicon and oxygen so that these potential molecular intermediates can be observed. These  
14 molecules have strong absorption features between 7.0  $\mu\text{m}$  and 8.0  $\mu\text{m}$ , most notably,  
15 and are fully characterized in the IR from the present quantum chemical data. The  
16 antisymmetric Si–O stretches of small silicon oxide clusters also fall in this range and  
17 have large intensities. Hence, this quantum chemical analysis provides spectral data for  
18 such molecules that may be of significance for astrochemical classification and could play a  
19 role in the formation or degradation of mineral nanocrystals from or into their constituent  
20 atoms. Both explicitly computed anharmonic fundamental vibrational frequencies and  
21 those determined from scaled harmonic frequencies agree well with known experimental  
22 data, and spectroscopic constants are provided herein such that astronomical rotational  
23 spectral characterization may also be possible for the  $C_{2v}$  SiO<sub>3</sub> and Si<sub>2</sub>O<sub>3</sub> molecules.

24 *Keywords:* Rovibrational spectroscopy; quantum chemistry; astrochemistry; planet  
25 formation

26 **1. Introduction**

27 Quartz, sand, and silica are all forms for the most abundant material in the Earth's  
28 crust and mantle. Crystalline SiO<sub>2</sub> gives way to the molten form deeper into the mantle,  
29 and it mixes with magnesium, iron, and aluminum to form most of the material present  
30 between the crust and core [1]. Similar processes likely played out all over the early Solar  
31 System and even beyond since both oxygen and silicon are two of the most abundant  
32 atoms in the Universe [2, 3, 4]. Consequently, there are few inorganic materials as  
33 common as silicon dioxide.

34 Silicon dioxide boils at nearly 3300 K and becomes isovalent with carbon dioxide in  
35 its triatomic molecular form. Such would likely be a necessary signature of silicon-based  
36 metabolism in exobiology. With regards to less exotic applications, silicates are known  
37 to condense under cold conditions creating rocky material in the first place [5], and these  
38 small molecules likely aggregate from silicon and oxygen atoms or the SiO monomer  
39 which has been observed in astrophysical environments since 1971 [6]. At the other end  
40 of the star's lifetime, the ablation of rocky materials as a star dies [7] or the stellar  
41 infall of most solid materials at any stage in a star's evolution will likely vaporize the  
42 crystalline or even molten silicates and quartzes creating small silicon oxide molecules  
43 for which little spectral data are well-known. A similar vaporization process would also  
44 take place terrestrially in foundries when high temperatures are utilized in large-scale  
45 industrial processes. Regardless of the circumstance, further spectral analysis of small  
46 silicon oxide clusters in the gas phase is necessary to produce data for the classification  
47 of silicon dioxide and its derivatives [8] such that the evolution of silicate materials from  
48 atoms to solids can be observed geophysically or industrially.

49 Argon matrix experiments have provided vibrational frequencies for the antisym-  
50 metric stretch and bending fundamentals of isolated SiO<sub>2</sub> at 1416 cm<sup>-1</sup> and 273 cm<sup>-1</sup>,  
51 respectively [9, 10, 11, 12]. Five of the six fundamental frequencies of SiO<sub>3</sub> have been ten-

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52 tatively assigned from similar experiments in conjunction with density functional theory  
53 analysis (DFT) [13], but follow-up work has not conclusively confirmed such attributions.  
54 Most notably, the  $a_1$  O–O stretch could not be attributed experimentally from the DFT  
55 computations. The larger structures have been observed through photoelectron spec-  
56 troscopy [14, 8], but the fundamental vibrational frequencies have not been conclusively  
57 determined in the laboratory. Three modes of  $\text{Si}_2\text{O}_4$  have been reported in the literature  
58 again from argon matrix experiments [15], but there has been no experimental or high-  
59 level theoretical follow-up in the intervening 30 years. Granted, the gas phase data are  
60 likely similar to the argon matrix results in each of these studies since the argon should  
61 interact far less with the silicon and oxygen atoms than they would with hydrogens, for  
62 instance, but corroboration of some variety for these modes is still lacking. While some  
63 are better understood than others, the full vibrational spectra and especially gas-phase  
64 rotational constants for each these molecules are not fully classified.

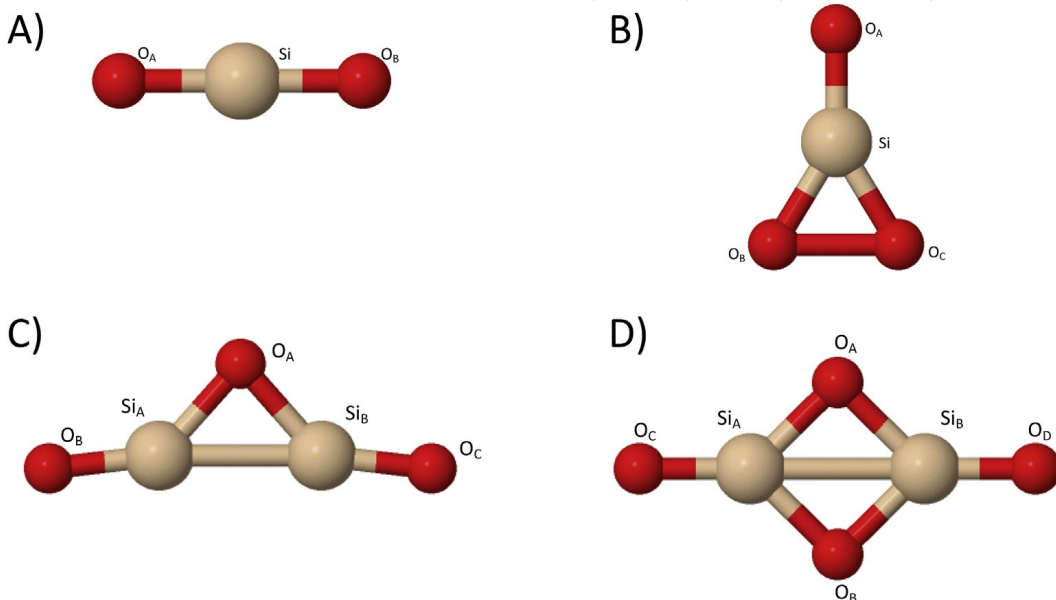
65     Consequently, high-accuracy spectroscopic predictions for the vibrational frequencies  
66 of these molecules will enhance spectral models of environments where they may be  
67 found both in the laboratory and in nature. Quartic force fields (QFFs) are fourth-order  
68 Taylor series approximations to the potential portion of the internuclear Hamiltonian [16].  
69 These have been shown to produce exceptional accuracy for determining anharmonic  
70 vibrational frequencies for relatively low computational cost, provided that a sufficient  
71 electronic structure method can be used to compute this specialized potential energy  
72 surface [16], most often coupled cluster theory at the singles, doubles, and perturbative  
73 triples [CCSD(T)][17, 18] level. The most accurate QFFs have produced vibrational  
74 frequencies to within  $1\text{ cm}^{-1}$  on occasion and most often within  $5\text{-}10\text{ cm}^{-1}$ . The accuracy  
75 of the rotational constants vary, but most are within 30 MHz of gas phase experiment  
76 [19, 20, 21, 22, 23, 24, 25, 26, 27, 28, 29, 30, 31].

77     However, the computational cost of this involved CCSD(T) approach including cor-  
78 rections for core electron correlation and complete basis set extrapolations is prohibitive  
79 for molecules containing more than five atoms. Recent work has shown that vibrational

80 frequencies of closed-shell molecules utilizing explicitly correlated CCSD(T)-F12b[32, 33]  
81 for the QFF energy points are within  $7 \text{ cm}^{-1}$  of the more expensive computations and  
82 cost orders of magnitude less time [34, 35, 36]. Some CCSD(T)-F12b anharmonic vibra-  
83 tional frequencies are actually closer to experiment than their more costly counterparts  
84 [35]. This approach has been utilized to predict anharmonic spectral data for  $\text{MgSiO}_3$   
85 and two isomers of  $\text{Mg}_2\text{SiO}_4$  (enstatite and forsterite monomers, respectively) for which  
86 no previous vibrational or rotational spectroscopic data, theoretical or otherwise, exist  
87 [37]. Most notably these inorganic oxides produce very large infrared intensities implying  
88 that relatively small amounts of material could still be observed [38, 39, 37, 40]. Fur-  
89 thermore, other studies have produced scaling factors for harmonic frequencies of small,  
90 inorganic oxides that promise to reduce the computational time further for similar species  
91 in the production of accurate, anharmonic vibrational frequencies [40].

92     Regardless, the present work will extend the data for these geophysically- and astrochemically-  
93 relevant silicon oxides by computing the anharmonic vibrational frequencies and spectro-  
94 scopic constants for four closed-shell singlet silicon oxide structures:  $\text{SiO}_2$ ,  $\text{SiO}_3$ ,  $\text{Si}_2\text{O}_3$ ,  
95 and  $\text{Si}_2\text{O}_4$ . The  $D_{3h}$  structures of both  $\text{CO}_3$  and  $\text{SiO}_3$  have proven problematic in the  
96 past [41]. Here, only the  $C_{2v}$   $\text{SiO}_3$  structure, akin to its carbon analogue which performed  
97 well in recent QFF examination [42], is considered. While a four-membered ring isomer  
98 of  $\text{Si}_2\text{O}_3$  is also known [43], we are leaving this structure for future study. The  $D_{2h}$  struc-  
99 ture of  $\text{Si}_2\text{O}_4$  has also been previously imposed to explain the results of photoelectron  
100 experiments [14] and the known vibrational frequencies thereof [15]. The rovibrational  
101 data provided herein will help to classify the building blocks or end products of silicates  
102 in regions where the solid form is known to exist [44, 7] especially with the growth in  
103 telescopic power most notably signified in the upcoming launch of the *James Webb Space*  
104 *Telescope (JWST)*.

Figure 1: The optimized geometries and atom labels of A)  $\text{SiO}_2$ , B)  $\text{SiO}_3$ , C)  $\text{Si}_2\text{O}_3$ , and D)  $\text{Si}_2\text{O}_4$ .



## 105 2. Computational Details

106 Precise optimization of the reference geometry is the first step to computing QFF-  
107 based anharmonic rovibrational data. The geometry optimization and all subsequent  
108 energy computations use CCSD(T)-F12b with the cc-pVTZ-F12 [45, 46] basis set (ab-  
109 breviated as F12-TZ from here on) in MOLPRO2015.1 [47, 48]. The optimized geometry  
110 is then used to compute the harmonic vibrational frequencies within MOLPRO for com-  
111 parison to those that result from the QFF. From the reference geometry of each silicon  
112 oxide molecule, coordinates are constructed to define the QFF with the bond lengths dis-  
113 placed by  $0.005 \text{ \AA}$  and the bond angles and dihedrals displaced by  $0.005$  radians. Each  
114 of the four molecules has its own unique coordinate system since each has a different  
115 number of bond lengths, angles, and dihedrals giving different numbers of total points  
116 necessary for the QFF.  $\text{SiO}_2$  has 57;  $\text{SiO}_3$  has 413;  $\text{Si}_2\text{O}_3$  has 1585; and  $\text{Si}_2\text{O}_4$  has 1973  
117 points. The coordinates for each system are defined below with atom labels in Figure 1.

118 The out-of-plane torsional modes are labeled as OPB. The coordinates for SiO<sub>2</sub>:

$$S_1(\sigma_g) = \frac{1}{\sqrt{2}}[(Si - O_A) + (Si - O_B)] \quad (1)$$

$$S_2(\sigma_u) = \frac{1}{\sqrt{2}}[(Si - O_A) - (Si - O_B)] \quad (2)$$

$$S_3(\pi_u) = \angle(O_A - Si - O_B) \quad (3)$$

$$S_4(\pi_u) = \angle(O_A - Si - O_B); \quad (4)$$

119 SiO<sub>3</sub>

$$S_1(a_1) = Si - O_A \quad (5)$$

$$S_2(a_1) = \frac{1}{\sqrt{2}}[(O_B - Si) + (O_C - Si)] \quad (6)$$

$$S_3(a_1) = \frac{1}{\sqrt{2}}[\angle(O_B - Si - O_A) + \angle(O_C - Si - O_A)] \quad (7)$$

$$S_4(b_2) = \frac{1}{\sqrt{2}}[(O_B - Si) - (O_C - Si)] \quad (8)$$

$$S_5(b_2) = \frac{1}{\sqrt{2}}[\angle(O_B - Si - O_A) - \angle(O_C - Si - O_A)] \quad (9)$$

$$S_6(b_1) = OPB(O_A - Si - O_B - O_C); \quad (10)$$

120 Si<sub>2</sub>O<sub>3</sub>

$$S_1(a_1) = Si_1 - Si_2 \quad (11)$$

$$S_2(a_1) = \frac{1}{\sqrt{2}}[(O_A - Si_A) + (O_A - Si_B)] \quad (12)$$

$$S_3(a_1) = \frac{1}{\sqrt{2}}[(Si_A - O_B) + (Si_B - O_C)] \quad (13)$$

$$S_4(a_1) = \frac{1}{\sqrt{2}}[\angle(O_B - Si_A - O_A) + \angle(O_C - Si_B - O_A)] \quad (14)$$

$$S_5(b_2) = \frac{1}{\sqrt{2}}[(O_A - Si_A) - (O_A - Si_B)] \quad (15)$$

$$S_6(b_2) = \frac{1}{\sqrt{2}}[(Si_A - O_B) - (Si_B - O_C)] \quad (16)$$

$$S_7(b_2) = \frac{1}{\sqrt{2}}[\angle(O_B - Si_A - O_A) - \angle(O_C - Si_B - O_A)] \quad (17)$$

$$S_8(b_1) = \frac{1}{\sqrt{2}}[\tau(O_B - Si_A - O_A - Si_B) - \tau(O_C - Si_B - O_A - Si_A)] \quad (18)$$

$$S_9(a_2) = \frac{1}{\sqrt{2}}[\tau(O_B - Si_A - O_A - Si_B) + \tau(O_C - Si_B - O_A - Si_A)]; \quad (19)$$



121 and  $\text{Si}_2\text{O}_4$

$$S_1(a_g) = \frac{1}{\sqrt{2}}[(O_A - O_B) + (Si_A - Si_B)] \quad (20)$$

$$S_2(a_g) = \frac{1}{\sqrt{2}}[(O_A - O_B) - (Si_A - Si_B)] \quad (21)$$

$$S_3(a_g) = \frac{1}{\sqrt{2}}[(Si_A - O_C) + (Si_B - O_D)] \quad (22)$$

$$S_4(b_{1u}) = \frac{1}{2}[(O_A - Si_A) - (O_A - Si_B) - (O_B - Si_A) + (O_B - Si_B)] \quad (23)$$

$$S_5(b_{1u}) = \frac{1}{2}[\angle(O_C - Si_A - O_A) - \angle(O_C - Si_A - O_B) - \angle(O_D - Si_B - O_A) + \angle(O_D - Si_B - O_B)] \quad (24)$$

$$S_6(b_{2u}) = \frac{1}{2}[(O_A - Si_A) + (O_A - Si_B) - (O_B - Si_A) - (O_B - Si_B)] \quad (25)$$

$$S_7(b_{2u}) = \frac{1}{2}[\angle(O_C - Si_A - O_A) + \angle(O_C - Si_A - O_B) - \angle(O_D - Si_B - O_A) - \angle(O_D - Si_B - O_B)] \quad (26)$$

$$S_8(b_{3g}) = \frac{1}{\sqrt{2}}[(Si_A - O_C) - (Si_B - O_D)] \quad (27)$$

$$S_9(b_{3g}) = \frac{1}{2}[(O_A - Si_A) - (O_A - Si_B) + (O_B - Si_A) - (O_B - Si_B)] \quad (28)$$

$$S_{10}(b_{3u}) = \frac{1}{\sqrt{2}}[(O_C)_z + (O_D)_z] \quad (29)$$

$$S_{11}(b_{3u}) = \frac{1}{2}[(O_A)_z + (O_B)_z - (Si_A)_z - (Si_B)_z] \quad (30)$$

$$S_{12}(b_{2u}) = \frac{1}{\sqrt{2}}[(O_C)_z - (O_D)_z], \quad (31)$$

122 where the last of these three has exhibited some questionable results in the low-frequency  
 123 range for the magnesium fluoride dimer [39]. The other coordinate systems have been  
 124 utilized successfully in previous studies on the magnesium hydride monomer,  $\text{HeHHe}^+$ ,  
 125 carbon dioxide,  $\text{CO}_3$ ,  $\text{NCNCN}^-$ , and  $\text{C}_2\text{O}_3$  [39, 49, 42, 50].

126 The resulting energies are fit to a least-squares polynomial giving a sum of squared  
 127 residuals on the order of  $10^{-17}$  a.u.<sup>2</sup> for all molecules but  $\text{Si}_2\text{O}_4$  which is  $10^{-13}$  a.u.<sup>2</sup>  
 128 This fit determines the actual minimum, or equilibrium, geometry. The final set of  
 129 force constants are generated by refitting the points to this new minimum; these are  
 130 given in the supplemental information (SI). The INTDER program [51] transforms these  
 131 force constants into Cartesian coordinates for more general implementation. Then, the  
 132 SPECTRO program [52] computes the harmonic and anharmonic frequencies including  
 133 anharmonic zero-point vibrational energies (ZPVEs) using vibrational perturbation the-  
 134 ory at second-order (VPT2) as well as spectroscopic constants making use of rotational

Table 1: The F12-TZ Vibrationally Averaged ( $_0$ ) and Equilibrium ( $_e$ ) Geometrical Parameters (in Å or Degrees) as Defined from Figure 1.

	SiO <sub>2</sub>	SiO <sub>3</sub>	Si <sub>2</sub> O <sub>3</sub>	Si <sub>2</sub> O <sub>4</sub>
$r_0$ (Si <sub>A</sub> -O <sub>A</sub> )	1.51249	1.50567	1.50675	1.50045
$r_0$ (Si <sub>A</sub> -O <sub>B</sub> )		1.62024	1.67351	1.66525 <sup>a</sup>
$\angle_0$ (O <sub>A</sub> -Si <sub>A</sub> -O <sub>B</sub> )		148.675	137.445	90.544
$r_0$ (Si <sub>A</sub> -Si <sub>B</sub> )			2.21681	2.36644
$\angle_0$ (O <sub>A</sub> -Si <sub>A</sub> -O <sub>C</sub> )				135.326
$r_e$ (Si <sub>A</sub> -O <sub>A</sub> )	1.51066	1.50395	1.50726	1.50337
$r_e$ (Si <sub>A</sub> -O <sub>B</sub> )		1.61659	1.66903	1.66135 <sup>a</sup>
$\angle_e$ (O <sub>A</sub> -Si <sub>A</sub> -O <sub>B</sub> )		148.814	137.313	90.673
$r_e$ (Si <sub>A</sub> -Si <sub>B</sub> )			2.21069	2.36327
$\angle_e$ (O <sub>A</sub> -Si <sub>A</sub> -O <sub>C</sub> )				135.337

<sup>a</sup>This is actually  $r_{e/0}$  (Si<sub>A</sub>-O<sub>C</sub>) because of the different atom labeling in Si<sub>2</sub>O<sub>4</sub> but is placed here for consistency of bond type.

135 perturbation theory at second-order [53, 54, 55] and Fermi resonance polyads [56].

136 The B3LYP/aug-cc-pVDZ double-harmonic intensities and dipole moments are com-  
 137 puted with Gaussian09 [57, 58, 59, 60, 61]. These have been shown to be in good agree-  
 138 ment with higher-level computations previously [62, 63]. Additionally, the scaling factors  
 139 for the Si-O stretching (0.98242) and bending (0.99261) frequencies determined previ-  
 140 ously [40] are applied to the harmonics computed directly in MOLPRO for comparison  
 141 of their performance in these related but not identical systems.

### 142 3. Results & Discussion

143 The geometries for each of the four molecules examined here are given in Table 1  
 144 with labels from Figure 1. Most notably, the Si<sub>A</sub>-O<sub>A</sub> bond lengths, those that have the  
 145 oxygen atom in a silaketone moiety (external Si=O group) are largely consistent with a  
 146 bond length on the order of 1.5 Å. SiO<sub>2</sub> is the exception to this where the longer bond  
 147 length likely arises from the weakening of the silicon atom’s electron donation due to  
 148 having two Si=O bonds instead of just one, and its magnitude here is in line with that  
 149 computed previously [13]. The Si<sub>A</sub>-O<sub>B/C</sub> bonds can be thought of as single bonds from  
 150 a carbon analogue [42] and are notably and consistently longer. The Si-Si bonds are  
 151 longer in Si<sub>2</sub>O<sub>4</sub> than in Si<sub>2</sub>O<sub>3</sub>, but the additional oxygen atom in the former naturally  
 152 increases the size of the ring thus making for a longer diagonal distance between silicon  
 153 atoms.

Table 2: The F12-TZ Vibrational Frequencies (in  $\text{cm}^{-1}$ ) and Intensities (in  $\text{km/mol}$  in parentheses) for  $\text{SiO}_2$ .

Mode	Description	Symm.	Harm. Freq.	Anharm. (QFF)	Scale Factor	Scaled	Exp. <sup>a</sup>	$\Delta(\text{QFF-Scaled})$
1	1.00 $S_2$	$\sigma_u$	1439.7 (67)	1420.7	0.98242	1414.4	1416	6.3
2	1.00 $S_1$	$\sigma_g$	992.4 (0)	984.2	0.98242	975.0		9.2
3	1.00 $S_3$	$\pi_u$	290.6 (79)	289.0	0.99261	288.5	273	0.5
ZPVE			1361.4	1501.9				

<sup>a</sup>Argon matrix experimental data from [9, 10, 11, 12]

### 154 3.1. Anharmonic Frequencies

155 The F12-TZ vibrational frequencies for  $\text{SiO}_2$  are given in Table 2. While the gas phase  
156 values for this molecule are not known, the argon matrix results correlate exceptionally  
157 well with the explicit QFF values for the  $1420.7 \text{ cm}^{-1}$  and  $289.0 \text{ cm}^{-1}$  fundamentals  
158 compared to the  $1416 \text{ cm}^{-1}$  and  $273 \text{ cm}^{-1}$  experimental frequencies [12]. Interestingly,  
159 the scaled  $1414.4 \text{ cm}^{-1}$  and  $288.5 \text{ cm}^{-1}$  frequencies for the same modes are in slightly  
160 closer agreement with experiment. Even so, the difference in the explicit QFF anhar-  
161 monic frequencies and the scaled values are within the accuracy ( $7 \text{ cm}^{-1}$ ) of the F12-TZ  
162 approach implying that either is a valuable choice of method. The infrared inactive sym-  
163 metric stretch has a larger difference between the two anharmonic frequencies, but no  
164 experimental data can verify which is more accurate. The F12-TZ QFF produces nearly  
165 identical values as gas-phase experimental  $\text{CO}_2$  frequencies (differences of less than  $2$   
166  $\text{cm}^{-1}$  in every case) [42], and similar accuracies appear to be present in the silicon ana-  
167 logue, as well.

168 The strong correspondence between theory and argon matrix experiments continues  
169 with  $C_{2v}$   $\text{SiO}_3$ . As shown in Table 3, the difference between the F12-TZ  $\text{SiO}_3$  anharmonic  
170 frequencies and that for the argon matrix experiments[13] is never more than  $8 \text{ cm}^{-1}$ ,  
171 and  $\nu_3$  at  $857.1 \text{ cm}^{-1}$  is within  $2 \text{ cm}^{-1}$  of experiment. Granted, the matrix will shift  
172 these values relative to the gas phase values that would be observed in astrophysical  
173 contexts, but, again, these shifts should be small. Furthermore, the B3LYP/aug-cc-pVTZ  
174 double-harmonic intensities are in at least semi-quantitative agreement with experiment

Table 3: The F12-TZ QFF Vibrational Frequencies (in  $\text{cm}^{-1}$ ) and Intensities (in  $\text{km/mol}$  in parentheses) for  $C_{2v}$   $\text{SiO}_3$ .

Mode	Description	Symm.	Harm. Freq.	Anharm. (QFF)	Scale Factor	Scaled	Exp. <sup>a</sup>	$\Delta(\text{QFF-Scaled})$
1	$0.71S_1 + 0.29S_2$	$a_1$	1391.4 (114)	1371.0	0.98242	1366.9	1363.5 (100)	4.1
2	$0.72S_2 + 0.27S_1$	$a_1$	894.2 (5)	885.3	0.98242	878.5	877.1 (12)	6.8
3	$1.00S_4$	$b_2$	868.5 (90)	857.1	0.98242	853.2	855.3 (53)	3.9
4	$1.00S_3$	$a_1$	515.0 (25)	497.0	0.99261	511.2	–	-14.2
5	$1.00S_6$	$b_1$	309.7 (82)	305.3	0.99261	307.4	287.8 (96)	-2.1
6	$1.00S_5$	$b_2$	299.3 (60)	298.0	0.99261	297.1	292.0 (62)	0.9
ZPVE			2139.1	2129.5				

<sup>a</sup>Argon matrix data from [13].

175 showing which bands could be considered strong and which could be considered weak.  
 176 Unsurprisingly, the  $\nu_1$  terminal  $\text{Si}=\text{O}_A$  stretch at  $1371.0 \text{ cm}^{-1}$  is the brightest vibrational  
 177 mode, and the OPB bend for  $\text{O}_A$  is also bright lining up with experiment.

178 The previous  $\text{SiO}_3$  experiment could not identify the  $a_1 \nu_4$  symmetric bend (equiv-  
 179 alently described as the  $\text{O}-\text{O}$  stretch). This fundamental has an intensity greater than  
 180 the  $\nu_2$   $\text{Si}-\text{O}$  symmetric stretch previously characterized in the argon matrix experiments  
 181 [13]. The reason is likely that the previous DFT computations suggested that the fun-  
 182 damental should lie higher in frequency close to  $582 \text{ cm}^{-1}$ . However, the present work  
 183 strongly suggests that this fundamental is much lower in frequency at  $497.0 \text{ cm}^{-1}$ . Un-  
 184 fortunately, this region of the IR spectrum is not reported in this previous work negating  
 185 any *ex post facto* analysis. Regardless, the F12-TZ QFF values confirm the other argon  
 186 matrix fundamental vibrational frequency assignments for  $C_{2v}$   $\text{SiO}_3$ [13] and show that  
 187 the last remaining band is most likely lower in frequency than previously believed.

Table 4:  $\text{Si}_2\text{O}_3$  F12-TZ Harmonic and Fundamental Vibrational Frequencies (in  $\text{cm}^{-1}$ ) and Intensities (in  $\text{km/mol}$  in parentheses).

Mode	Description	Symm.	Harm. Freq.	Anharm. (QFF)	Scale Factor	Scaled	$\Delta(\text{QFF-Scaled})$
1	$0.86S_3+0.08S_2+0.07S_1$	$a_1$	1331.0 (17)	1303.3	0.98242	1307.6	-4.3
2	$0.96S_6+0.04S_5$	$b_2$	1274.0(152)	1245.6	0.98242	1251.6	-6.0
3	$0.85S_2+0.08S_4+0.07S_3$	$a_1$	859.1(122)	836.2	0.98242	844.0	-7.8
4	$0.88S_5+0.07S_7+0.05S_6$	$b_2$	663.2(52)	637.6	0.98242	651.6	14.0
5	$0.74S_1+0.18S_4+0.07S_3$	$a_1$	440.7(21)	432.5	0.99261	437.5	-5.0
6	$0.93S_7+0.08S_5$	$b_2$	248.7(16)	245.0	0.99261	246.9	-1.9
7	$1.00S_9$	$b_1$	245.7(62)	244.6	0.99261	243.9	0.7
8	$0.74S_4+0.19S_1+0.07S_2$	$a_1$	189.8(30)	188.2	0.99261	188.4	-0.2
9	$1.00S_8$	$a_2$	153.4(0)	153.7	0.99261	152.3	1.4
ZPVE			2643.3	2679.5			

188 The F12-TZ QFF fundamental vibrational frequencies for  $\text{Si}_2\text{O}_3$  have not been pre-  
189 viously explored experimentally or theoretically and are given here in Table 4. The  
190 silaketone stretches ( $\nu_1$  and  $\nu_2$ ) are in the same frequency range as that for  $\text{SiO}_3$ , and  
191 the antisymmetric  $\nu_2$  stretch at  $1245.6 \text{ cm}^{-1}$  has the greatest intensity of all the fun-  
192 damentals. This is also true for the related  $\text{C}_2\text{O}_3$  and  $\text{C}_2\text{N}_3^-$  molecules [42, 50]. The  
193  $a_1 \nu_3$  stretching of the central  $\text{O}_A$  atom produces the next-brightest fundamental at  
194  $836.2 \text{ cm}^{-1}$ . The permanent dipole moment is being extended in this case, giving a  
195 larger charge separation upon vibration. The remaining frequencies are dimmer but still  
196 non-negligible, and the five lowest frequency fundamentals are all below  $500 \text{ cm}^{-1}$ , as is  
197 common for such heavy atoms like silicon [38, 39, 37].

Table 5:  $\text{Si}_2\text{O}_4$  F12-TZ Harmonic and Fundamental Vibrational Frequencies (in  $\text{cm}^{-1}$ ) and Intensities (in  $\text{km/mol}$  in parentheses).

Mode	Description	Symm.	Harm. Freq.	Anharm. (QFF)	Scale Factor	Scaled	Exp. <sup>a</sup>	$\Delta(\text{QFF-Scaled})$
1	$0.84S_3+0.13S_1$	$a_g$	1354.4 (0)	1338.2	0.98242	1330.6		7.6
2	$0.86S_8+0.14S_9$	$b_{3g}$	1316.9 (382)	1303.6	0.98242	1293.7	1293.3	9.9
3	$0.89S_6+0.11S_7$	$b_{2u}$	904.8 (259)	911.7	0.98242	888.9	889.2	22.8
4	$0.70S_1+0.24S_2+0.06S_3$	$a_g$	871.6 (0)	861.3	0.98242	856.3		5.0
5	$0.86S_9+0.14S_8$	$b_{3g}$	804.0 (252)	856.8	0.98242	789.9	786.4	66.9
6	$1.00S_4$	$b_{1u}$	718.7 (0)	658.1	0.98242	706.1		-48.0
7	$0.73S_2+0.17S_1+0.10S_3$	$a_g$	488.0 (0)	485.9	0.98242	479.4		6.5
8	$0.94S_1+0.06S_10$	$b_{3u}$	465.4 (99)	456.1	0.99261	462.0		-5.9
9	$1.00S_5$	$b_{1u}$	313.3 (0)	281.7	0.99261	311.0		-29.3
10	$0.89S_7+0.11S_6$	$b_{2u}$	295.5 <sup>†</sup> (43)		0.99261	293.3		
11	$1.00S_12$	$b_{2u}$	236.6 (0)	230.5	0.99261	234.9		-4.4
12	$0.94S_10+0.06S_11$	$b_{3u}$	121.5 <sup>†</sup> (26)		0.99261	120.6		
ZPVE			3945.4	3796.3				

<sup>a</sup> Argon matrix experimental results from [15].

<sup>†</sup> Denotes a MOLPRO harmonic frequency

198 The 12 fundamental vibrational frequencies for  $D_{2h}$   $\text{Si}_2\text{O}_4$  are given in Table 5. The  
199 silaketone stretches are blue-shifted in this molecule compared to  $\text{Si}_2\text{O}_3$  and are present  
200 above  $1300 \text{ cm}^{-1}$ . The  $b_{3g} \nu_2$  antisymmetric stretch, again, has the largest intensity of  
201 all the fundamental frequencies for this molecule. Stretches within the ring ( $\nu_3$  and  $\nu_5$ )  
202 are the next-brightest with intensities above  $250 \text{ km/mol}$ , more than 3.5 times that of  
203 the antisymmetric stretch in water. Comparison to argon matrix data from [15] gives  
204 good agreement between the  $1293.3 \text{ cm}^{-1}$  value and the F12-TZ QFF  $\nu_2$  frequency at

205 1303.6  $\text{cm}^{-1}$ . This deteriorates slightly for  $\nu_3$  where experiment places this at 889.2  $\text{cm}^{-1}$   
206 and the QFF is 911.7  $\text{cm}^{-1}$ . The correlation is completely off for  $\nu_5$  with experiment  
207 attributing this ring deformation to a band at 786.4  $\text{cm}^{-1}$  and the QFF at 856.8  $\text{cm}^{-1}$ .  
208 Either the band has been misassigned in the experiment, or the computations are off.

209 In this case, the latter is most likely correct. The fitting of the points was the worst  
210 for  $\text{Si}_2\text{O}_4$  of the silicon oxides studied. Furthermore, two harmonic vibrational frequen-  
211 cies computed via the QFF do not align with those computed from within MOLPRO’s  
212 standard harmonic frequency computation,  $\omega_{10}$  and  $\omega_{12}$ . The OPB coordinate struggles  
213 to define the proper motion within the constraints of the QFF and VPT2, and this subse-  
214 quently affects the fitting of the force constants for the other coordinates. The potential  
215 for the OPB is likely flat reducing the capabilities of VPT2 as defined by the QFF. The  
216  $\nu_{10}$  and  $\nu_{12}$  fundamentals could not even be computed from the QFF data. Hence, the  
217 F12-TZ QFF VPT2 anharmonic vibrational frequencies for  $\text{Si}_2\text{O}_4$  below the silaketone  
218 stretches should be treated as suggestions.

219 However, all is not lost in the prediction of these anharmonic frequencies. The recent  
220 determination of scaling factors for M–O stretches and bends (where M is a second-row  
221 atom) can be applied to  $\text{Si}_2\text{O}_4$ . Doing so actually produces a fundamental frequency for  
222  $\nu_3$  at 888.9  $\text{cm}^{-1}$  and  $\nu_5$  at 789.9  $\text{cm}^{-1}$ . Both are within 3.5  $\text{cm}^{-1}$  of the argon matrix  
223 experiment. Furthermore, the  $\nu_2$  antisymmetric stretch is 1293.7  $\text{cm}^{-1}$  from the scaled  
224 values, 9.9  $\text{cm}^{-1}$  below the explicit QFF, and only 0.4  $\text{cm}^{-1}$  below the experimental value.  
225 Consequently, the scaled harmonics are likely producing more meaningful fundamental  
226 vibrational frequencies for this molecule than the QFF. The harmonic force field is much  
227 better behaved and less likely to suffer from noise contamination in these numerical  
228 derivatives [19, 20], and the amount of absolute anharmonicity is relatively small in the  
229 first place.

230 The scaled harmonic frequencies are also listed for the other three molecules giving  
231 slightly better correlation with experiment for  $\text{SiO}_2$  (Table 2) as discussed previously.  
232 Agreement between experiment and scaled harmonics of  $\text{SiO}_3$  (Table 3) is comparable

233 with the F12-TZ QFF VPT2 results. Some modes are better with the scaled values  
234 ( $\nu_1$ ) and some with the explicit anharmonicity computed ( $\nu_3$ ) implying that either is  
235 appropriate. Both also demonstrate that  $\nu_4$  is still lower in frequency than previous  
236 experiments explored. Since there are no experimental data for  $\text{Si}_2\text{O}_3$ , comparison be-  
237 tween the QFF VPT2 results and the scaled harmonics both with F12-TZ is necessary,  
238 but both are quite comparable with one another (Table 4). All modes agree to within 8  
239  $\text{cm}^{-1}$  save for the  $b_2 \nu_4$  antisymmetric stretch. The mean absolute error (MAE) between  
240 the QFF and scaled harmonics for the stretching frequencies is  $5.8 \text{ cm}^{-1}$  when removing  
241  $\nu_4$  and  $8.0 \text{ cm}^{-1}$  when including it. The bends and torsions are much closer with the  
242 MAE at  $1.8 \text{ cm}^{-1}$ , but, again, the magnitudes of the frequencies are smaller in the first  
243 place. Hence, these scaling factors are comparable to the QFF VPT2 fundamental vibra-  
244 tional frequencies implying that these heuristics could be used as a first-order guess to  
245 the fundamental vibrational frequencies of  $\text{Si}_2\text{O}_4$  and potentially even for larger silicon  
246 oxide clusters where QFFs or any anharmonic vibrational frequency computations are  
247 prohibitively large.

### 248 3.2. Rotational and Spectroscopic Constants

249 The spectroscopic constants for each of the molecules examined are given in Table 6.  
250 These include the pure rotational constants, the vibrationally-averaged rotational con-  
251 stants, the quartic and sextic (Watson S Hamiltonian) distortion constants, and even  
252 the dipole moments of the two  $C_{2v}$  molecules. While these may not be as accurate as  
253 methods including core electron correlation or other additive factors [35, 36], these rota-  
254 tional constants should serve as a good starting point for assessing the rotational spectra  
255 of these molecules. Strangely, a search of the literature did not yield any experimental  
256 rotational constants for  $\text{SiO}_2$  which are provided here for this nonpolar molecule. The  
257 other three molecules are all clearly near-prolate rotors especially for  $\text{Si}_2\text{O}_3$ .

258 The vibrationally-excited rotational constants (numbered in the same order as the  
259 fundamental vibrational frequencies) for  $\text{Si}_2\text{O}_4$  are given for the modes with the least  
260 questionable vibrational frequencies. While the pure rotational transitions of this molecule

Table 6: The F12-TZ QFF Spectroscopic Data for the Four Silicon Oxides.

	SiO <sub>2</sub>	SiO <sub>3</sub>	Si <sub>2</sub> O <sub>3</sub>	Si <sub>2</sub> O <sub>4</sub>
$A_0$ (MHz)		22360.3	21364.6	11474.0
$B_0$ (MHz)	6907.7	5489.9	1766.6	1636.1
$C_0$ (MHz)		4400.5	1630.8	1433.1
$A_1$ (MHz)		22333.1	21390.3	11469.1
$B_1$ (MHz)	6872.3	5466.1	1761.6	1632.6
$C_1$ (MHz)		4384.4	1626.6	1430.4
$A_2$ (MHz)		22292.4	21361.1	11465.8
$B_2$ (MHz)	6888.1	5480.1	1762.9	1633.0
$C_2$ (MHz)		4392.0	1627.6	1430.8
$A_3$ (MHz)		22446.5	21282.1	11497.3
$B_3$ (MHz)	6920.3	5469.6	1764.8	1633.3
$C_3$ (MHz)		4391.4	1628.1	1430.4
$A_4$ (MHz)		22001.5	20997.0	11458.5
$B_4$ (MHz)		5518.5	1768.1	1635.0
$C_4$ (MHz)		4392.3	1629.6	1431.8
$A_5$ (MHz)		22310.3	21576.9	11442.8
$B_5$ (MHz)		5493.9	1761.6	1635.2
$C_5$ (MHz)		4409.0	1627.2	1431.5
$A_6$ (MHz)		22408.0	21370.4	11446.5
$B_6$ (MHz)		5504.1	1768.2	1633.6
$C_6$ (MHz)		4400.9	1631.2	1430.7
$A_7$ (MHz)			21368.9	
$B_7$ (MHz)			1767.8	
$C_7$ (MHz)			1632.6	
$A_8$ (MHz)			21507.1	
$B_8$ (MHz)			1767.9	
$C_8$ (MHz)			1630.7	
$A_9$ (MHz)			21279.8	
$B_9$ (MHz)			1768.2	
$C_9$ (MHz)			1633.4	
$\Delta_J$ (Hz)	1.500	941.78	120.62	56.658
$\Delta_K$ (kHz)		133.62	363.47	9.985
$\Delta_{JK}$ (kHz)		15.616	-4.444	0.389
$\delta_J$ (Hz)		227.57	18.119	8.426
$\delta_K$ (kHz)		11.708	1.065	0.484
$\Phi_J$ ( $\mu$ Hz)	-8.885	518.99	2.540	2.770
$\Phi_K$ (Hz)		-0.240	9.387	0.022
$\Phi_{JK}$ (mHz)		80.382	0.920	0.242
$\Phi_{KJ}$ (mHz)		-233.60	-127.77	-4.098
$\phi_j$ ( $\mu$ Hz)		261.21	4.481	1.058
$\phi_{jk}$ (mHz)		45.749	0.152	0.152
$\phi_k$ (Hz)		1.292	0.502	0.009
$\mu$ (D)	-	0.87	0.66	-



261 will not be observed since it has no dipole moment, the  $A_2$ ,  $B_2$ , and  $C_2$  values, how-  
262 ever, will likely be important for rovibrational modeling since  $\nu_2$  of this molecule has the  
263 largest infrared intensity computed of the set. Finally, the dipole moments are reported  
264 at the bottom of Table 6.  $\text{SiO}_3$  is the most polar, but  $\text{Si}_2\text{O}_3$  has a smaller but similar  
265 magnitude dipole moment. This differs from the carbon analogues where  $\text{C}_2\text{O}_3$  is almost  
266 apolar [42] likely due to the larger electronegativity difference between oxygen and silicon  
267 as well as the longer Si–O bonds.

#### 268 4. Conclusions

269 The small silicon oxide clusters  $\text{SiO}_2$ ,  $\text{SiO}_3$ ,  $\text{Si}_2\text{O}_3$ , and  $\text{Si}_2\text{O}_4$  are shown here to be  
270 stable species with notably bright mid- to far-IR active fundamental vibrational frequen-  
271 cies. The antisymmetric silaketone stretch in the silicon dioxide dimer has the largest  
272 intensity of the set. The range between  $1420\text{ cm}^{-1}$  and  $1250\text{ cm}^{-1}$  ( $7.0\text{ }\mu\text{m}$  and  $8.0$   
273  $\mu\text{m}$ ) contains this most intense band and its counterparts from the other three oxides.  
274 The other infrared bands typically fall below  $700\text{ cm}^{-1}$  ( $>14.3\text{ }\mu\text{m}$ ) with many of the  
275 silicon oxides analyzed here having one or two bands around  $850\text{ cm}^{-1}$  ( $\sim 11.8\text{ }\mu\text{m}$ ).  
276 Each of these regions have notable bumps from astronomical spectra [64], implying that  
277 small, geochemically-relevant silicon oxides may be present in circumstellar media and  
278 protoplanetary disks. Upcoming *JWST* spectra could potentially resolve such peaks.  
279 Furthermore, the larger molecules have infrared features approaching that of known sil-  
280 icate dusts. The polar  $\text{SiO}_3$  and  $\text{Si}_2\text{O}_3$  clusters could be observed from the ground  
281 with radiotelescopes, and the present data will aid in the experimental characterization  
282 necessary to provide reference data for such observations.

283 Additionally, the previously derived scaling factors [40] show promise in treating sim-  
284 ilar inorganic oxides that are intractable for QFF computations. Comparison to experi-  
285 mental spectra for  $\text{SiO}_2$  and  $\text{SiO}_3$  shows that both the explicit anharmonic computations  
286 and the scaled harmonics are similarly accurate partly due to the small magnitudes of  
287 the anharmonicities of the molecules examined. However, these scaled harmonics also do

288 not suffer from coordinate issues observed in Si<sub>2</sub>O<sub>4</sub> making them also potentially useful  
289 for future exploration of larger mineralogically-relevant inorganic oxides.

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## 293 References

- 294 [1] White, W. M. *Geochemistry*, 1st ed.; Wiley: Hoboken, NJ, 2013; Chapter 10: The Big Picture:  
295 Cosmochemistry.
- 296 [2] Savage, B. D.; Sembach, K. R. Interstellar Abundances from Absorption-Line Observations with  
297 the *Hubble Space Telescope*. *Annu. Rev. Astron. Astrophys.* **1996**, *34*, 279–329.
- 298 [3] McCall, B. J. Dissociative Recombination of Cold H<sub>3</sub><sup>+</sup> and Its Interstellar Implications. *Phil. Trans.*  
299 *Royal Soc. A* **2006**, *364*, 2953–2963.
- 300 [4] Fortenberry, R. C. The Case for Gas-Phase Astrochemistry without Carbon. *Mol. Astrophys.* **2020**,  
301 *18*, 100062.
- 302 [5] Rouillé, G.; Jäger, C.; Krasnokutski, S. A.; Krebszb, M.; Henning, T. Cold Condensation of Dust  
303 in the ISM. *Farad. Discuss.* **2014**, *168*, 449–460.
- 304 [6] Wilson, R. W.; Penzias, A. A.; Jefferts, K. B.; Kutner, M.; Thaddeus, P. Discovery of Interstellar  
305 Silicon Monoxide. *Astrophys. J.* **1971**, *167*, L97–L100.
- 306 [7] McCarthy, M. C.; Gottlieb, C. A.; Cernicharo, J. Building Blocks of Dust: A Coordinated Labora-  
307 tory and Astronomical Study of the Archtype AGB Carbon Star IRC+10216. *J. Molec. Spectrosc.*  
308 **2019**, *356*, 7–20.
- 309 [8] Wang, L.-S.; Desai, S. R.; Wu, H.; Nicholas, J. B. Small Silicon Oxide Clusters: Chains and Rings.  
310 *Z. Phys. D* **1997**, *40*, 36–39.
- 311 [9] Schnöckel, H. IR Spectroscopic Detection of Molecular SiO<sub>2</sub>. *Angew. Chem. Int. Ed. Engl.* **1978**,  
312 *17*, 616–617.
- 313 [10] Schnöckel, V. H. Matrixreaktionen von SiO. IR-spektroskopischer Nachweis der Molekeln SiO<sub>2</sub> und  
314 OSiCl<sub>2</sub>. *Z. Anorg. Allg. Chem.* **1980**, *460*, 37–50.
- 315 [11] Andrews, L.; McCluskey, M. Bending Modes of SiO<sub>2</sub> and GeO<sub>2</sub> in Solid Argon. *J. Mol. Spectrosc.*  
316 **1992**, *154*, 223–225.
- 317 [12] Jacox, M. E. Vibrational and Electronic Energy Levels of Polyatomic Transient Molecules. *J. Phys.*  
318 *Chem. Ref. Data* **1994**, *23*, 1–461.
- 319 [13] Tremblay, B.; Roy, P.; Manceron, L.; Alikhani, M. E.; Roy, D. Vibrational Spectrum and Structure  
320 of Silicon Trioxide SiO<sub>3</sub>: A Matrix Isolation Infrared and Density Functional Theory Study. *J.*  
321 *Chem. Phys.* **1996**, *104*, 2773–2781.
- 322 [14] Wang, L.-S.; Wu, H.; Desai, S. R.; Fan, J.; Colson, S. D. A Photoelectron Spectroscopic Study of  
323 Small Silicon Oxide Clusters: SiO<sub>2</sub>, Si<sub>2</sub>O<sub>3</sub>, and Si<sub>2</sub>O<sub>4</sub>. *J. Phys. Chem.* **1996**, *100*, 8697–8700.
- 324 [15] Mehner, T.; Göckes, H. J.; Schunck, S.; Schnöckel, H. Dimeres SiO<sub>2</sub> Matrix-IR-Untersuchungen  
325 und ab Initio SCF-Rechnungen. *Z. Anorg. Allg. Chem.* **1980**, *580*, 121–130.
- 326 [16] Fortenberry, R. C.; Lee, T. J. Computational Vibrational Spectroscopy for the Detection of  
327 Molecules in Space. *Ann. Rep. Comput. Chem.* **2019**, *15*, 173–202.
- 328 [17] Raghavachari, K.; Trucks, G. W.; Pople, J. A.; Head-Gordon, M. A Fifth-Order Perturbation  
329 Comparison of Electron Correlation Theories. *Chem. Phys. Lett.* **1989**, *157*, 479–483.
- 330 [18] Shavitt, I.; Bartlett, R. J. *Many-Body Methods in Chemistry and Physics: MBPT and Coupled-*  
331 *Cluster Theory*; Cambridge University Press: Cambridge, 2009.
- 332 [19] Huang, X.; Lee, T. J. A Procedure for Computing Accurate *Ab Initio* Quartic Force Fields: Appli-  
333 cation to HO<sub>2</sub><sup>+</sup> and H<sub>2</sub>O. *J. Chem. Phys.* **2008**, *129*, 044312.
- 334 [20] Huang, X.; Lee, T. J. Accurate *Ab Initio* Quartic Force Fields for NH<sub>2</sub><sup>-</sup> and CCH<sup>-</sup> and Rovibra-  
335 tional Spectroscopic Constants for Their Isotopologs. *J. Chem. Phys.* **2009**, *131*, 104301.

- 336 [21] Huang, X.; Taylor, P. R.; Lee, T. J. Highly Accurate Quartic Force Field, Vibrational Frequencies,  
337 and Spectroscopic Constants for Cyclic and Linear  $C_3H_3^+$ . *J. Phys. Chem. A* **2011**, *115*, 5005–  
338 5016.
- 339 [22] Fortenberry, R. C.; Huang, X.; Francisco, J. S.; Crawford, T. D.; Lee, T. J. The *trans*-HOCO  
340 Radical: Fundamental Vibrational Frequencies, Quartic Force Fields, and Spectroscopic constants.  
341 *J. Chem. Phys.* **2011**, *135*, 134301.
- 342 [23] Fortenberry, R. C.; Huang, X.; Francisco, J. S.; Crawford, T. D.; Lee, T. J. Vibrational Frequencies  
343 and Spectroscopic Constants from Quartic Force Fields for *cis*-HOCO: The Radical and the Anion.  
344 *J. Chem. Phys.* **2011**, *135*, 214303.
- 345 [24] Fortenberry, R. C.; Huang, X.; Francisco, J. S.; Crawford, T. D.; Lee, T. J. Fundamental Vibrational  
346 Frequencies and Spectroscopic Constants of HOCS<sup>+</sup>, HSCO<sup>+</sup>, and Isotopologues via Quartic Force  
347 Fields. *J. Phys. Chem. A* **2012**, *116*, 9582–9590.
- 348 [25] Huang, X.; Fortenberry, R. C.; Lee, T. J. Protonated Nitrous Oxide, NNOH<sup>+</sup>: Fundamental Vibra-  
349 tional Frequencies and Spectroscopic Constants from Quartic Force Fields. *J. Chem. Phys.* **2013**,  
350 *139*, 084313.
- 351 [26] Fortenberry, R. C.; Huang, X.; Crawford, T. D.; Lee, T. J. Quartic Force Field Rovibrational  
352 Analysis of Protonated Acetylene,  $C_2H_3^+$ , and Its Isotopologues. *J. Phys. Chem. A* **2014**, *118*,  
353 7034–7043.
- 354 [27] Zhao, D.; Doney, K. D.; Linnartz, H. Laboratory Gas-Phase Detection of the Cyclopropenyl Cation  
355 (*c*- $C_3H_3^+$ ). *Astrophys. J. Lett.* **2014**, *791*, L28.
- 356 [28] Fortenberry, R. C.; Lee, T. J.; Müller, H. S. P. Excited Vibrational Level Rotational Constants for  
357 SiC<sub>2</sub>: A Sensitive Molecular Diagnostic for Astrophysical Conditions. *Molec. Astrophys.* **2015**, *1*,  
358 13–19.
- 359 [29] Kitchens, M. J. R.; Fortenberry, R. C. The Rovibrational Nature of Closed-Shell Third-Row Tri-  
360 atomics: HOX and HXO, X = Si<sup>+</sup>, P, S<sup>+</sup>, and Cl. *Chem. Phys.* **2016**, *472*, 119–127.
- 361 [30] Fortenberry, R. C. Quantum Astrochemical Spectroscopy. *Int. J. Quant. Chem.* **2017**, *117*, 81–91.
- 362 [31] Bizzocchi, L.; Lattanzi, V.; Laas, J.; Spezzano, S.; Giuliano, B. M.; Prudeniano, D.; Endres, C.;  
363 Sipilä, O.; Caselli, P. Accurate Sub-millimetre Rest Frequencies for HOCO<sup>+</sup> and DOCO<sup>+</sup> Ions.  
364 *Astron. Astrophys.* **2017**, *602*, A34.
- 365 [32] Adler, T. B.; Knizia, G.; Werner, H.-J. A Simple and Efficient CCSD(T)-F12 Approximation. *J.*  
366 *Chem. Phys.* **2007**, *127*, 221106.
- 367 [33] Knizia, G.; Adler, T. B.; Werner, H.-J. Simplified CCSD(T)-F12 Methods: Theory and Benchmarks.  
368 *J. Chem. Phys.* **2009**, *130*, 054104.
- 369 [34] Agbaglo, D.; Lee, T. J.; Thackston, R.; Fortenberry, R. C. A Small Molecule with PAH Vibrational  
370 Properties and a Detectable Rotational Spectrum: *c*-(C) $C_3H_2$ , Cyclopropenylidene Carbene. *As-*  
371 *trophys. J.* **2019**, *871*, 236.
- 372 [35] Agbaglo, D.; Fortenberry, R. C. The Performance of CCSD(T)-F12/aug-cc-pVTZ for the Compu-  
373 tation of Anharmonic Fundamental Vibrational Frequencies. *Int. J. Quantum Chem.* **2019**, *119*,  
374 e25899.
- 375 [36] Agbaglo, D.; Fortenberry, R. C. The Performance of Explicitly Correlated Wavefunctions  
376 [CCSD(T)-F12b] in the Computation of Anharmonic Vibrational Frequencies. *Chem. Phys. Lett.*  
377 **2019**, *734*, 136720.
- 378 [37] Valencia, E. M.; Worth, C. J.; Fortenberry, R. C. Enstatite (MgSiO<sub>3</sub>) and Forsterite (Mg<sub>2</sub>SiO<sub>4</sub>)  
379 Monomers and Dimers: Highly-Detectable Infrared and Radioastronomical Molecular Building  
380 Blocks. *Mon. Not. Royal Astron. Soc.* **2020**, *492*, 276–282.
- 381 [38] Kloska, K. A.; Fortenberry, R. C. Gas-Phase Spectra of MgO Molecules: A Possible Connection from  
382 Gas-Phase Molecules to Planet Formation. *Mon. Not. Royal Astron. Soc.* **2018**, *474*, 2055–2063.
- 383 [39] Palmer, C. Z.; Fortenberry, R. C. Rovibrational Considerations for the Monomers and Dimers  
384 of Magnesium Hydride (MgH<sub>2</sub>) and Magnesium Fluoride (MgF<sub>2</sub>). *J. Phys. Chem. A* **2018**, *122*,  
385 7079–7088.
- 386 [40] Westbrook, B. R.; Fortenberry, R. C. Anharmonic Frequencies of (MO)<sub>2</sub> & Related Hydrides for  
387 M = Mg, Al, Si, P, S, Ca, & Ti and Heuristics for Predicting Anharmonic Corrections of Inorganic  
388 Oxides. *J. Phys. Chem. A* **2020**, *124*, 3191–3204.
- 389 [41] Beste, A.; Bartlett, R. J. The Electronic Structure of SiO<sub>3</sub>: A Problematic Example for Coupled  
390 Cluster Methods. *Chem. Phys. Lett.* **2002**, *366*, 100–108.
- 391 [42] Fortenberry, R. C.; Peters, D.; Ferari, B. C.; Bennett, C. J. Rovibrational Spectral Analysis of CO<sub>3</sub>  
392 and C<sub>2</sub>O<sub>3</sub>: Potential Sources for O<sub>2</sub> Observed in Comet 67P/Churyumov-Gerasimenko. *Astrophys.*  
393 *J. Lett.* **2019**, *886*, L10.
- 394 [43] Wang, Y.; Chen, M.; Xie, Y.; Wei, P.; Schaefer, III, H. F.; von R. Schleyer, P.; Robinson, G. H.

- 395 Stabilization of Elusive Silicon Oxides. *Nature Chem.* **2015**, *7*, 509–513.
- 396 [44] Rho, J.; Gomez, H. L.; Boogert, A.; Smith, M. W. L.; Lagage, P.-O.; Dowell, D.; Clark, C. J. R.;  
 397 Peeters, E.; Cami, J. A Dust Twin of Cas A: Cool Dust and 21  $\mu\text{m}$  Silicate Dust Feature in the  
 398 Supernova Remnant G54.1+0.3. *Mon. Not. Royal Astron. Soc.* **2018**, *479*, 5101–5123.
- 399 [45] Hill, J. G.; Peterson, K. A. Correlation Consistent Basis Sets for Explicitly Correlated Wavefunc-  
 400 tions: Valence and Core-Valence Basis Sets for Li, Be, Na, and Mg. *Phys. Chem. Chem. Phys.*  
 401 **2010**, *12*, 10460–10468.
- 402 [46] Prascher, B. P.; Woon, D. E.; Peterson, K. A.; Dunning, T. H.; Wilson, A. K. Gaussian Basis Sets  
 403 for Use in Correlated Molecular Calculations. VII. Valence, Core-Valence, and Scalar Relativistic  
 404 Basis Sets for Li, Be, Na, and Mg. *Theor. Chem. Acc.* **2011**, *128*, 69–82.
- 405 [47] Werner, H.-J. et al. MOLPRO, Version 2015.1, a Package of *ab Initio* Programs. 2015; see  
 406 <http://www.molpro.net>.
- 407 [48] Werner, H.-J.; Knowles, P. J.; Knizia, G.; Manby, F. R.; Schütz, M. Molpro: A General-Purpose  
 408 Quantum Chemistry Program Package. *WIREs Comput. Mol. Sci.* **2012**, *2*, 242–253.
- 409 [49] Stephan, C. J.; Fortenberry, R. C. The Interstellar Formation and Spectra of the Noble Gas, Proton-  
 410 Bound HeHHe<sup>+</sup>, HeHNe<sup>+</sup> and HeHAr<sup>+</sup> Complexes. *Mon. Not. Royal Astron. Soc.* **2017**, *469*,  
 411 339–346.
- 412 [50] Dubois, D.; Sciamma-O’Brien, E.; Fortenberry, R. C. The Fundamental Vibrational Frequencies  
 413 and Spectroscopic Constants of the Dicyanoamine Anion, NCNCN<sup>-</sup> (C<sub>2</sub>N<sub>3</sub><sup>-</sup>): Quantum Chemical  
 414 Analysis for Astrophysical and Planetary Environments. *Astrophys. J.* **2019**, *883*, 109.
- 415 [51] Allen, W. D.; coworkers, 2005; *INTDER* 2005 is a General Program Written by W. D. Allen and  
 416 Coworkers, which Performs Vibrational Analysis and Higher-Order Non-Linear Transformations.
- 417 [52] Gaw, J. F.; Willets, A.; Green, W. H.; Handy, N. C. In *Advances in Molecular Vibrations and*  
 418 *Collision Dynamics*; Bowman, J. M., Ratner, M. A., Eds.; JAI Press, Inc.: Greenwich, Connecticut,  
 419 1991; pp 170–185.
- 420 [53] Mills, I. M. In *Molecular Spectroscopy - Modern Research*; Rao, K. N., Mathews, C. W., Eds.;  
 421 Academic Press: New York, 1972; pp 115–140.
- 422 [54] Watson, J. K. G. In *Vibrational Spectra and Structure*; Doring, J. R., Ed.; Elsevier: Amsterdam,  
 423 1977; pp 1–89.
- 424 [55] Papousek, D.; Aliev, M. R. *Molecular Vibration-Rotation Spectra*; Elsevier: Amsterdam, 1982.
- 425 [56] Martin, J. M. L.; Taylor, P. R. Accurate *ab Initio* Quartic Force Field for *trans*-HNNH and Treat-  
 426 ment of Resonance Polyads. *Spectrochim. Acta A* **1997**, *53*, 1039–1050.
- 427 [57] Becke, A. D. Density-Functional Thermochemistry. III. The Role of Exact Exchange. *J. Chem.*  
 428 *Phys.* **1993**, *98*, 5648–5652.
- 429 [58] Yang, W. T.; Parr, R. G.; Lee, C. T. Various Functionals for the Kinetic Energy Density of an  
 430 Atom or Molecule. *Phys. Rev. A* **1986**, *34*, 4586–4590.
- 431 [59] Lee, C.; Yang, W. T.; Parr, R. G. Development of the Colle-Salvetti Correlation-Energy Formula  
 432 into a Functional of the Electron Density. *Phys. Rev. B.* **1988**, *37*, 785–789.
- 433 [60] Dunning, T. H. Gaussian Basis Sets for Use in Correlated Molecular Calculations. I. The Atoms  
 434 Boron through Neon and Hydrogen. *J. Chem. Phys.* **1989**, *90*, 1007–1023.
- 435 [61] Frisch, M. J. et al. Gaussian 09 Revision D.01. 2009; Gaussian Inc. Wallingford CT.
- 436 [62] Yu, Q.; Bowman, J. M.; Fortenberry, R. C.; Mancini, J. S.; Lee, T. J.; Crawford, T. D.; Klem-  
 437 perer, W.; Francisco, J. S. The Structure, Anharmonic Vibrational Frequencies, and Intensities of  
 438 NNHNN<sup>+</sup>. *J. Phys. Chem. A* **2015**, *119*, 11623–11631.
- 439 [63] Finney, B.; Fortenberry, R. C.; Francisco, J. S.; Peterson, K. A. A Spectroscopic Case for SPSi  
 440 Detection: The Third-Row in a Single Molecule. *J. Chem. Phys.* **2016**, *145*, 124311.
- 441 [64] Molster, F. J.; Lim, T. L.; Sylvester, R. J.; Waters, L. B. F. M.; Barlow, M. J.; Beintema, D. A.;  
 442 Cohen, M.; Cox, P.; Schmitt, B. The Complete ISO Spectrum of NGC 6302. *Astron. Astrophys.*  
 443 **2001**, *372*, 165–172.

## CHAPTER III: ROTATIONAL CONSTANTS

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1    Highly-Accurate Quartic Force Fields for the Prediction of  
2            Anharmonic Rotational Constants and Fundamental  
3            Vibrational Frequencies

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---

12    **Abstract**

13            The CcCR quartic force field (QFF) methodology is capable of computing  $B_0$  and  $C_0$   
14    rotational constants to within 35 MHz (0.14%) of experiment for triatomic and larger  
15    molecules with at least two heavy atoms. Additionally, the same constants for molecules  
16    with four or more atoms agree to within 20 MHz (0.12%) of experiment for the current  
17    test set. This work also supports previous claims that the same QFF methodology can  
18    produce fundamental vibrational frequencies with a deviation less than  $5.7 \text{ cm}^{-1}$  from  
19    experiment. Consequently, this approach of augmenting complete basis set extrapolated  
20    energies with treatments of core electron correlation and scalar relativity produces some  
21    of the most accurate rovibrational spectroscopic data available.

22            **Keywords:** Quantum Chemistry; Quartic Force Fields; Rotational Spectroscopy;  
23    Vibrational Spectroscopy; Coupled Cluster Theory

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24    **1. Introduction**

25            Quartic force fields (QFFs) making use of coupled cluster theory, specifically at the  
26    singles, doubles, and perturbative triples [CCSD(T)] level of theory [1], and other post-  
27    CCSD(T) effects have been shown to produce exceptionally accurate vibrational frequen-  
28    cies and rotational constants [2, 3]. These accuracies are reported to be within  $1.0 \text{ cm}^{-1}$

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29 on occasion and often within  $5.0 \text{ cm}^{-1}$  of gas-phase experiment for vibrational frequen-  
30 cies, and within 30 or so MHz for the  $B$  and  $C$  rotational constants [4–15]. While these  
31 anecdotal values are promising for the composite method employed, a more systematic  
32 analysis of how the QFF performs for the prediction of these observable spectroscopic  
33 values is necessary.

34 The above-mentioned accurate QFF employs CCSD(T) in a complete basis set (CBS)  
35 extrapolation scheme [16] utilizing the aug-cc-pVTZ, aug-cc-pVQZ, and aug-cc-pV5Z  
36 basis sets [17–19, 16, 20]. The difference in the CCSD(T) energy with and without  
37 the inclusion of core electrons is used to correct the CBS energy for the effects of core  
38 correlation. These computations most often utilize the Martin-Taylor (MT) core elec-  
39 tron basis set [21], but the more standard aug-cc-pCVTZ basis set has been shown to  
40 be just as viable [4, 22, 23]. Additional corrections for scalar relativity are incorpo-  
41 rated using the Douglas-Kroll (DK) Hamiltonian [24, 25] by taking the difference in the  
42 CCSD(T)/cc-pVTZ-DK energies including and excluding the relativistic terms. This  
43 use of the CCSD(T)/CBS energy (“C”) including corrections for core correlation (“cC”)  
44 and scalar relativity (“R”) is often called the CcCR QFF [26]. Additional corrections  
45 for higher-order electron correlation (“E”) and even quantum electrodynamics (“Q”) can  
46 be added for presumed higher accuracy. For most molecules, save for those with higher  
47 bond orders, such terms do not improve the accuracy of the computations for molecules  
48 containing third-row or smaller atoms by more than  $1.0 \text{ cm}^{-1}$  in many cases but increase  
49 the computational cost of the QFF by more than a factor of 2, especially if higher-order  
50 coupled cluster truncations beyond full triples are included [4, 6, 26, 23].

51 Consequently, the CcCR QFF has been touted as a sweet spot for accuracy and com-  
52 putational cost in the determination of anharmonic vibrational frequencies and rotational  
53 constants. While errors in vibrational fundamentals have rarely reached above  $10 \text{ cm}^{-1}$   
54 for the CcCR QFF, some cases have arisen where the rotational constants are in error by  
55 more than 100 MHz [4, 5]. As such, the present work will analyze a set of small molecules  
56 for which there are known, highly-accurate experimental data available for comparison  
57 to computed CcCR QFF rotational constants. Some vibrational frequencies will also  
58 be characterized as a part of this study, but rotational constants remain the principle  
59 focus. Other groups have established various means of computing quantum chemical  
60 rotational constants with much success even aiding in the detection of new molecules  
61 in space via radiotelescopes [27–34]. However, the present work will focus solely on the  
62 CcCR approach.

63 In particular, less computationally intensive perturbation theory (PT) approaches  
64 exist for obtaining highly-accurate rotational constants within 0.05% of gas-phase exper-  
65 imental values [35, 36]. The level of theory required for accurate vibrational corrections  
66 to equilibrium rotational constants is also quite low, with even B3LYP [37–39] offering  
67 reasonable accuracy for well-behaved systems [40]. However, these previous results rely  
68 on fortuitous characteristics of the systems investigated, primarily molecules composed  
69 of hydrogen, carbon, nitrogen, and oxygen. CcCR, on the other hand, is a more rigorous  
70 theoretical technique that is applicable to broader classes of molecules. Additionally,  
71 the CcCR methodology uses composite energies to make up the force field rather than  
72 correcting the observables afterwards. Previous work has clearly demonstrated the reli-  
73 ability of CcCR fundamental frequencies [3], but the goal of the present study is to  
74 analyze its efficacy in the computation of rotational constants. Such constants are of-  
75 ten produced as a byproduct of a given vibrational study, so a better understanding of

76 their typical accuracy is needed to accompany future CcCR investigations and to further  
77 contextualize previous ones.

78 Variational approaches to the determination of rotational data also exist [41, 42], but  
79 these are typically much more costly, raising issues as to whether the additional cost  
80 is worth the accuracy gains [3]. Again, the focus here is on assessing the established  
81 CcCR/PT approach, not reviewing theoretical rovibrational spectroscopic techniques.  
82 With this in mind, the selection of molecules has been drawn from the CcCR/PT liter-  
83 ature rather than from a broader pool of computed spectra. The literature selection is  
84 further augmented by new work on HCN, HCO<sup>+</sup>, HNC, HMgNC, HNO, NH<sub>3</sub>, H<sub>2</sub>CO,  
85 and HOCN, which helps to tease out trends in the existing data. To further limit the  
86 scope of the present work, all of the molecules studied herein are covalently bound. More  
87 weakly-bound structures lead to very flat potential energy surfaces, which are poorly  
88 treated by QFFs in general [43, 44], especially those relying upon numerical differentia-  
89 tion. Thus, the present work will serve to benchmark the performance of the CcCR/PT  
90 QFF for various classes of small (3-6 atoms), covalent molecules, for which accurate  
91 gas-phase experimental rotational data are available.

## 92 2. Computational Details

93 The computation of CcCR QFFs begins with geometry optimizations of the desired  
94 molecule via CCSD(T)/aug-cc-pV5Z. The geometry is then corrected for changes re-  
95 sulting from differences in CCSD(T)/MT optimizations with and without core electrons  
96 included. These reference geometries are then displaced by 0.005 Å and 0.005 radians  
97 per each step per each symmetry-internal coordinate. The QFF is of the form:

$$V = \frac{1}{2} \sum_{ij} F_{ij} \Delta_i \Delta_j + \frac{1}{6} \sum_{ijk} F_{ijk} \Delta_i \Delta_j \Delta_k + \frac{1}{24} \sum_{ijkl} F_{ijkl} \Delta_i \Delta_j \Delta_k \Delta_l, \quad (1)$$

98 where the  $F_{ij\dots}$  terms are the force constants and the  $\Delta_i \Delta_j \dots$  terms are the displaced  
99 distances for coordinates  $i$ ,  $j$ , and so forth [45, 46]. Then, at each point ( $\Delta_i \Delta_j \dots$ ), the  
100 seven energies necessary for the CcCR QFF are computed. These are then fit via a least-  
101 squares procedure and refit in order to produce zero gradients, the resulting equilibrium  
102 geometry, and the other force constants ( $F_{ij\dots}$ ) for the potential function in Equation 1  
103 [4, 6].

104 The force constants are then transformed into generic Cartesian coordinates for gen-  
105 eral usage through the INTDER2005 program [47]. The Cartesian force constants are  
106 utilized in the SPECTRO program [48] to produce the rotational constants and anhar-  
107 monic vibrational frequencies. Coriolis and Fermi resonances, as well as Fermi resonance  
108 polyads [49], are then included in the second-order perturbation theory [50–52] treatment  
109 of the vibrational and rotational Hamiltonians within SPECTRO in order to produce the  
110 most accurate spectroscopic data possible within a perturbation theory approach.

111 Once the CcCR QFF data are determined for the present work, these are compared  
112 with experimental data taken from the chemical literature. Other, theoretical CcCR ro-  
113 tational constants and fundamental vibrational frequencies are taken from the literature,  
114 as well, in order to provide a larger data set. For the present work, most of the molecules  
115 explored are closed-shell, but some are radicals providing a breadth of molecules. The  
116 full list is given in Tables 1 and 2. In all cases, the experimental rotational constants



117 and fundamental vibrational frequencies are subtracted from the computed CcCR QFF  
118 second-order perturbation theory values to yield the reported differences. The absolute  
119 values of these differences are averaged herein in order to produce the mean absolute  
120 error or MAE. The percent error is calculated, using

$$\frac{|\text{Diff.}|}{\text{Experiment}} \times 100\%, \quad (2)$$

121 where  $|\text{Diff.}|$  is the absolute value of the aforementioned difference. The rotational con-  
122 stants are also divided into subgroups in order to tease out the behavior of certain types  
123 of molecules with certain properties or characteristics. In addition to MAEs for these  
124 groups, the mean absolute percent errors (MA%Es) are reported in Table 3.

### 125 3. Results and Discussion

126 The CcCR quartic force field approach provides exceptional accuracy in the prediction  
127 of vibrationally-averaged rotational constants in most cases. The full set of principle  
128 rotational constants considered in this work is given in Table 1. The sources of the  
129 data are listed in Table 2 with eight CcCR molecular datasets original to this work:  
130 HCN, HCO<sup>+</sup>, HNC, HMgNC, HNO, NH<sub>3</sub>, H<sub>2</sub>CO, and HOCN. Clearly from Table 1,  
131 many systems have excellent correspondence between CcCR QFF theory and gas-phase  
132 experiment while some are less accurate. The MAE for the entire set of rotational  
133 constants is 537.4 MHz (0.20%) as listed in Table 3. Most evidently, the  $A_0$  rotational  
134 constants vary by significantly greater amounts than  $B_0$  and  $C_0$ , both in terms of absolute  
135 errors and even for the percent errors. Many of the molecules in this set are near-prolate,  
136 and the larger magnitude of the  $A$  constant contributes to this difference. Besides the  
137 oblate  $c$ -C<sub>3</sub>H<sub>3</sub><sup>+</sup>, the least prolate molecule,  $c$ -C<sub>3</sub>H<sub>2</sub>, has the smallest error for the CcCR  
138 QFF  $A_0$  compared to experiment at 64.0 MHz. Such behavior in near-prolate molecules  
139 is known within the quantum chemistry community [27, 28], and this is borne out in the  
140 MAE for the  $A$  constants alone at 1577.5 MHz (0.31%).

141 However, such a large error incorrectly implies that CcCR QFF computations are  
142 insufficient for predicting spectroscopic constants. Removing the  $A$  constants from con-  
143 sideration certainly improves the average rotational constant and also even lowers the  
144 percent error. Linear molecules have degenerate  $B$  and  $C$  constants, and averaging over  
145  $B_0$  and  $C_0$  requires counting the  $B$  constant in the linear molecules twice. Such an aver-  
146 age of  $B_0$  and  $C_0$  with  $A_0$  removed reduces the MAE by an order of magnitude to 92.7  
147 MHz (0.15%) in line two of Table 3.

148 Closer inspection of Table 1 shows that several molecules appear to be outliers with  
149 absolute errors for  $B_0$  and  $C_0$  in the range of 100+ MHz. Most of these are hydrides with  
150 a lone heavy atom: NH<sub>2</sub><sup>-</sup>, H<sub>2</sub>O, and NH<sub>3</sub>. Removing the rotational constant errors for  
151 these three molecules lowers the MAE by roughly one-third to 34.6 MHz (0.14%).

152 The error in the computed versus experimental rotational constants of water has  
153 been known since the formulation of the CcCR family of composite QFF schema [4], and  
154 this work shows that such simple systems are often the most difficult to model. The  
155 reason for this poor behavior in light molecules is multifaceted. First, these molecules  
156 have larger rotational constants implying that regular percent errors manifest themselves  
157 with higher magnitude differences from experiment. Second, these molecules have large

Table 1: The CcCR Equilibrium & Vibrationally-Averaged and Gas Phase Experimental Rotational Constants in MHz.

Molecule	CcCR			Experiment	Difference	Diff.	% Error
	Equil.	Vib. Avg.					
HCN	<i>B</i>	44580.6	44386.4	44316	70.4	70.4	0.16
HCO <sup>+</sup>	<i>B</i>	44851.4	44611.1	44594.4	16.7	16.7	0.04
HNC	<i>B</i>	45571.5	45405.3	45332	73.3	73.3	0.16
C <sub>2</sub> H	<i>B</i>	44187.6	43702.2	43674.5	27.7	27.7	0.06
C <sub>2</sub> H <sup>-</sup>	<i>B</i>	41735.7	41568.7	41639.2	-70.5	70.5	0.17
HMgNC	<i>B</i>	5436.4	5471.0	5481.4	-10.4	10.4	0.19
NH <sub>2</sub> <sup>-</sup>	<i>A</i>	688371.4	689654.6	691045.6	-1391.0	1391.0	0.20
	<i>B</i>	392881.0	391702.8	391780.8	-78.0	78.0	0.02
	<i>C</i>	250125.8	243476.4	243269.6	206.8	206.8	0.09
H <sub>2</sub> O	<i>A</i>	820382.1	831542.7	835840.3	-4297.6	4297.6	0.51
	<i>B</i>	438356.5	436042.4	435351.7	690.7	690.7	0.16
	<i>C</i>	285702.2	278664.6	278138.7	525.9	525.9	0.19
HNO	<i>A</i>	566875.5	560818.2	553898.6	6919.6	6919.6	1.25
	<i>B</i>	42534.6	42430.6	42312.8	117.8	117.8	0.28
	<i>C</i>	39566.0	39281.6	39165.1	116.5	116.5	0.30
HSO	<i>A</i>	302261.2	300113.4	299483.9	629.5	629.5	0.21
	<i>B</i>	20715.7	20597.8	20502.8	95.0	95.0	0.46
	<i>C</i>	19387.0	19219.9	19135.7	84.2	84.2	0.44
HSS	<i>A</i>	299200.8	297770.1	296974.4	795.7	795.7	0.27
	<i>B</i>	8030.6	8009.9	7996.4	13.5	13.5	0.17
	<i>C</i>	7820.7	7790.4	7776.7	13.7	13.7	0.18
HPSi	<i>A</i>	306054.4	296974.9	297187	-212.1	212.1	0.07
	<i>B</i>	8211.2	8169.4	8169	0.4	0.4	0.00
	<i>C</i>	7996.6	7939.1	7936.7	2.4	2.4	0.03
SiC <sub>2</sub>	<i>A</i>	52343.2	52602.5	52473.7	128.8	128.8	0.25
	<i>B</i>	13264.2	13145.8	13158.7	-12.8	12.8	0.10
	<i>C</i>	10582.5	10444.8	10442.6	2.1	2.1	0.02
NH <sub>3</sub>	<i>A</i>	299988.5	297539.8	298107	-567.2	567.2	0.19
	<i>C</i>	190683.3	186445.2	185751.4	693.8	693.8	0.37
H <sub>2</sub> CO	<i>A</i>	285857.2	282684.0	281963.8	720.2	720.2	0.26
	<i>B</i>	39055.5	38898.6	38832.1	66.5	66.5	0.17
	<i>C</i>	34361.0	34058.7	34002.5	56.2	56.2	0.17
HOCN	<i>A</i>	673758.4	677180.7	674308	2872.7	2872.7	0.43
	<i>B</i>	10622.2	10592.4	10577	15.4	15.4	0.15
	<i>C</i>	10457.4	10414.3	10398	16.3	16.3	0.16
<i>cis</i> -HOCO	<i>A</i>	141896.9	143151.8	142944.9	206.9	206.9	0.14
	<i>B</i>	11856.5	11757.3	11739.6	17.7	17.7	0.15
	<i>C</i>	10942.1	10846.5	10830	16.5	16.5	0.15
<i>trans</i> -HOCO	<i>A</i>	167470.4	168266.3	167768.1	498.2	498.2	0.30
	<i>B</i>	11535.7	11448.5	11433.2	15.3	15.3	0.13
	<i>C</i>	10792.5	10702.0	10686.7	15.3	15.3	0.14
HOCO <sup>+</sup>	<i>A</i>	765443.3	784759.5	789951	-5191.5	5191.5	0.66
	<i>B</i>	10818.6	10787.1	10773.6	13.5	13.5	0.13
	<i>C</i>	10667.2	10623.7	10609.4	14.3	14.3	0.13
NNOH <sup>+</sup>	<i>A</i>	624453.7	625221.3	625957.716	-736.4	736.4	0.12
	<i>B</i>	11357.9	11306.9	11301.5628	5.3	5.3	0.05
	<i>C</i>	11155.0	11090.2	11084.28	6.0	6.0	0.05
<i>c</i> -C <sub>3</sub> H <sub>2</sub>	<i>A</i>	35377.2	35156.6	35092.6	64.0	64.0	0.18
	<i>B</i>	32420.0	32243.1	32212.9	30.2	30.2	0.09
	<i>C</i>	16917.1	16767.7	16749.3	18.4	18.4	0.11
<i>c</i> -C <sub>3</sub> H <sub>3</sub> <sup>+</sup>	<i>A</i>	30956.6	30761.7	30753.9	7.8	7.8	0.03
	<i>B</i>	30956.6	30761.7	30753.9	7.8	7.8	0.03
	<i>C</i>	15478.3	15342.8	15338.9	3.9	3.9	0.03
Average					63.1	537.4	0.20

Table 2: The CcCR and Gas Phase Experimental References.

Molecule	CcCR	Experiment
HCN	This Work	53
HCO <sup>+</sup>	This Work	54-59
HNC	This Work	60, 61
C <sub>2</sub> H	62	63-66
C <sub>2</sub> H <sup>-</sup>	5	67, 68
HMgNC	This Work	69
NH <sub>2</sub> <sup>-</sup>	5	70
H <sub>2</sub> O	4	71
HNO	This Work	72
HSO	73	74, 31
HSS	73	75
HPSi	76	27
SiC <sub>2</sub>	11	77, 78
NH <sub>3</sub>	This Work	79, 80
H <sub>2</sub> CO	This Work	80
HOCN	This Work	81, 82
<i>cis</i> -HOCO	83	84
<i>trans</i> -HOCO	26	84
HOCO <sup>+</sup>	8	85
NNOH <sup>+</sup>	9	86, 87
<i>c</i> -C <sub>3</sub> H <sub>2</sub>	15	88
<i>c</i> -C <sub>3</sub> H <sub>3</sub> <sup>+</sup>	6	7

Table 3: Average Errors for CcCR vs. Experiment for Various Cases.

Set	Units	MAE	MA%E
A <sub>0</sub> , B <sub>0</sub> , & C <sub>0</sub>	MHz	537.4	0.20
B <sub>0</sub> & C <sub>0</sub>	MHz	92.7	0.15
B <sub>0</sub> & C <sub>0</sub> w/o H <sub>2</sub> O, NH <sub>2</sub> <sup>-</sup> , & NH <sub>3</sub>	MHz	34.6	0.14
B <sub>0</sub> & C <sub>0</sub> w/o Triatomics & NH <sub>3</sub>	MHz	18.9	0.12
A <sub>0</sub>	MHz	1577.5	0.32
$\nu_x$	cm <sup>-1</sup>	5.7	0.70

158 anharmonicities, as well. This shifts the zero-point structure and, subsequently, the  
159 rotational constants by more than is typical for treatment from second-order rotational  
160 perturbation theory. A large anharmonicity is also present in HNO which is known to  
161 have an extreme anharmonicity for the N–H stretch [89]. Here the  $B_0$  and  $C_0$  are in  
162 error from experiment by 117.8 MHz and 116.5 MHz, respectively, marking the largest  
163 deviations for molecules in this set with at least two heavy atoms. This is also borne out  
164 in the percent errors for this molecule which are the largest for both sets inclusive and  
165 exclusive of  $A$ . Formaldehyde is roughly the same mass as HNO and also has relatively  
166 large errors for  $B_0$  and  $C_0$ , but these are roughly half the magnitude and percent error  
167 of that in HNO. Hence, the anharmonicity and large rotational constants present in  
168 the three single-heavy-atom hydrides are driving this error beyond that of molecules  
169 containing multiple heavy atoms.

170 Further constraining the data set by removing all of the triatomic species and am-  
171 monia produces an MAE of only 18.9 MHz (0.12%). The increase in mass and decrease  
172 in the rotational constant values continue to drive the error down for the comparison  
173 between theory and experiment. The nine molecules remaining in the present set range  
174 in shape from linear to oblate. They contain standard  $p$ -block elements, in addition  
175 to Mg, and consist of both open- and closed-shell molecules. Formaldehyde has the  
176 largest errors of 66.5 MHz and 56.2 MHz while the next largest come from  $c$ -C<sub>3</sub>H<sub>2</sub> at  
177 30.2 MHz and 18.4 MHz. Save for HMgNC, all of the computed rotational constants are  
178 larger than their experimental counterparts. Consequently, the CcCR VPT2 QFF  $B_0$   
179 and  $C_0$  rotational constants can be said to be within 20 MHz or 0.12% of experiment for  
180 four-atom and larger molecular systems. Such accuracy should be sufficient for compar-  
181 ison with high-resolution gas-phase experiment and potentially even radioastronomical  
182 observation.

183 Since QFFs have been utilized to compute the rotational constants from CcCR, fun-  
184 damental vibrational frequencies are also produced in this methodology. Consequently,  
185 the present work provides a novel means for analyzing the performance of CcCR QFF  
186 second-order vibrational perturbation theory (VPT2) in predicting fundamental vibra-  
187 tional frequencies. Fewer molecules of our sample set have gas-phase, experimental vi-  
188 brational spectral data available, but 12 molecules have at least one high-resolution  
189 experimentally known, gas-phase fundamental vibrational frequency giving 36 frequen-  
190 cies for comparison as given in Table 4. For this set, the MAE per fundamental is 5.7  
191  $\text{cm}^{-1}$ . Consequently, the typical error reported previously for the CcCR QFF VPT2  
192 fundamental vibrational frequencies as mentioned in the introduction is verified for a  
193 larger set of molecules. Furthermore, of the 36 frequencies contained in this set, only one  
194 has an error of greater than 15  $\text{cm}^{-1}$  (SiC<sub>2</sub> at 21.0  $\text{cm}^{-1}$ ), and 23 are less than 5  $\text{cm}^{-1}$   
195 with errors for seven fundamentals of less than 2  $\text{cm}^{-1}$ .

#### 196 4. Conclusions

197 Quantum chemically-computed rotational constants are accurately represented from  
198 CcCR QFFs provided that the molecules of interest contain more than two heavy atoms.  
199 The MAE for the  $B_0$  and  $C_0$  constants for our set of explored molecules is 34.6 MHz  
200 (0.14%) when the smallest hydrides are excluded. Larger ( $\geq 4$  atoms) and less anhar-  
201 monic molecules have an even better MAE at 18.9 MHz (0.12%). The smallest  $B_0$  and  
202  $C_0$  errors from our set are from the relatively massive HPSi molecule (0.4 MHz and 2.4

Table 4: The CcCR and Gas Phase Experimental Fundamental Vibrational Frequencies in  $\text{cm}^{-1}$ .

Molecule		CcCR	Experiment	Difference	Diff.]	% Error
HCN	$\nu_1$	3313.1	3311.47	1.6	1.6	0.05
	$\nu_2$	719.7	711.97	7.7	7.7	0.45
	$\nu_3$	2106.2	2096.84	9.3	9.3	1.09
HCO <sup>+</sup>	$\nu_1$	3089.8	3088.73951	1.0	1.0	0.03
	$\nu_2$	2189.9	2183.94961	6.0	6.0	0.27
	$\nu_3$	833.4	829.72	3.7	3.7	0.44
HNC	$\nu_1$	3655.0	3652.65	2.3	2.3	0.06
	$\nu_2$	2027.7	2023.86	3.8	3.8	0.19
	$\nu_3$	474.8	462.72	12.1	12.1	2.61
H <sub>2</sub> O	$\nu_1$	3659.7	3657.05	2.7	2.7	0.07
	$\nu_2$	1595.8	1597.75	-1.9	1.9	0.12
	$\nu_3$	3758.0	3755.93	2.1	2.1	0.06
HNO	$\nu_1$	2688.8	2684	4.8	4.8	0.18
	$\nu_2$	1577.1	1565	12.1	12.1	0.77
	$\nu_3$	1510.8	1501	9.8	9.8	0.65
SiC <sub>2</sub>	$\nu_1$	1750.5	1746.0	4.5	4.5	0.26
	$\nu_2$	844.7	840.6	4.1	4.1	0.49
	$\nu_3$	175.4	196.37	-21.0	21.0	0.68
NH <sub>3</sub>	$\nu_1$	3445.5	3444	1.5	1.5	0.04
	$\nu_2$	3346.0	3337	9.0	9.0	0.27
	$\nu_3$	1628.7	1627	1.7	1.7	0.10
	$\nu_4$	974.2	950	24.2	24.2	2.55
H <sub>2</sub> CO	$\nu_1$	2832.6	2843	-10.4	10.4	0.37
	$\nu_2$	2782.5	2782	0.5	0.5	0.02
	$\nu_3$	1751.1	1746	5.1	5.1	0.29
	$\nu_4$	1501.8	1500	1.8	1.8	0.12
	$\nu_5$	1251.0	1249	2.0	2.0	0.16
	$\nu_6$	1171.5	1167	4.5	4.5	0.39
HOCN	$\nu_1$	3623.0	3610	13.0	13.0	0.36
	$\nu_2$	2298.5	2302	-3.5	3.5	0.15
	$\nu_3$	1231.3	1227	4.3	4.3	0.35
	$\nu_4$	1087.4	1082	5.4	5.4	0.50
	$\nu_6$	454.6	460	-5.4	5.4	1.17
	$\nu_7$	3371.2	3375.37413	-4.2	4.2	0.12
HOCO <sup>+</sup>	$\nu_1$	3332.0	3330.91	1.1	1.1	0.03
<i>c</i> -C <sub>3</sub> H <sub>2</sub>	$\nu_4$	1278.8	1278.8	0.0	0.0	0.00
	$\nu_5$	1065.1	1061.5	3.6	3.6	0.34
	$\nu_7$	888.6	886.4	2.2	2.2	0.25
	$\nu_9$	772.8	787.4	-14.6	14.6	1.85
<i>c</i> -C <sub>3</sub> H <sub>3</sub> <sup>+</sup>	$\nu_4$	3131.7	3131.1447	0.6	0.6	0.02
Average				2.7	5.7	0.70

203 MHz); the next smallest are for  $c\text{-C}_3\text{H}_3^+$  at 7.8 MHz and 3.9 MHz implying that the  
204 CcCR approach is largely agnostic to atom type at this point. The largest of the well-  
205 behaved set are for formaldehyde at 66.5 MHz and 56.2 MHz, but most are less than 20  
206 MHz, roughly the MAE. Such accuracies are close to the limits that can be utilized for di-  
207 rect astronomical observation. While a forest of lines are present in any spectral window  
208 observed towards interesting astronomical objects, the CcCR QFF theoretical rotational  
209 spectral progressions will be shifted only slightly when compared to such observations.

210 Additionally, the CcCR QFF is benchmarked to predict fundamental vibrational fre-  
211 quencies to within  $5.7\text{ cm}^{-1}$ , further showcasing the accuracy of this method in use for  
212 the past decade or so. While CCSD(T)-F12/cc-pVTZ-F12 has been shown to mirror the  
213 accuracies of the CcCR composite QFF in computing vibrational frequencies [90, 91],  
214 explicitly correlated theory cannot provide the same relative errors for the rotational  
215 constants. While composite schema for explicitly correlated methods may yet produce  
216 more accurate rotational constants [92], the canonical-CCSD(T)-based CcCR QFF re-  
217 mains one of the most accurate approaches for computing rotational constants developed  
218 thus far.

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## 224 References

- 225 [1] Raghavachari, K.; Trucks, G. W.; Pople, J. A.; Replogle, E. Highly Correlated Systems: Structure,  
226 Binding Energy and Harmonic Vibrational Frequencies of Ozone. *Chem. Phys. Lett.* **1989**, *158*,  
227 207–212.
- 228 [2] Császár, A. G. Anharmonic Molecular Force Fields. *WIREs Comput. Mol. Sci.* **2012**, *2*, 273–289.
- 229 [3] Fortenberry, R. C.; Lee, T. J. Computational Vibrational Spectroscopy for the Detection of  
230 Molecules in Space. *Ann. Rep. Comput. Chem.* **2019**, *15*, 173–202.
- 231 [4] Huang, X.; Lee, T. J. A Procedure for Computing Accurate *Ab Initio* Quartic Force Fields: Appli-  
232 cation to  $\text{HO}_2^+$  and  $\text{H}_2\text{O}$ . *J. Chem. Phys.* **2008**, *129*, 044312.
- 233 [5] Huang, X.; Lee, T. J. Accurate *Ab Initio* Quartic Force Fields for  $\text{NH}_2^-$  and  $\text{CCH}^-$  and Rovibra-  
234 tional Spectroscopic Constants for Their Isotopologs. *J. Chem. Phys.* **2009**, *131*, 104301.
- 235 [6] Huang, X.; Taylor, P. R.; Lee, T. J. Highly Accurate Quartic Force Field, Vibrational Frequencies,  
236 and Spectroscopic Constants for Cyclic and Linear  $\text{C}_3\text{H}_3^+$ . *J. Phys. Chem. A* **2011**, *115*, 5005–  
237 5016.
- 238 [7] Zhao, D.; Doney, K. D.; Linnartz, H. Laboratory Gas-Phase Detection of the Cyclopropenyl Cation  
239 ( $c\text{-C}_3\text{H}_3^+$ ). *Astrophys. J. Lett.* **2014**, *791*, L28.
- 240 [8] Fortenberry, R. C.; Huang, X.; Francisco, J. S.; Crawford, T. D.; Lee, T. J. Quartic Force Field  
241 Predictions of the Fundamental Vibrational Frequencies and Spectroscopic Constants of the Cations  
242  $\text{HOCO}^+$  and  $\text{DOCO}^+$ . *J. Chem. Phys.* **2012**, *136*, 234309.
- 243 [9] Huang, X.; Fortenberry, R. C.; Lee, T. J. Protonated Nitrous Oxide,  $\text{NNOH}^+$ : Fundamental Vibra-  
244 tional Frequencies and Spectroscopic Constants from Quartic Force Fields. *J. Chem. Phys.* **2013**,  
245 *139*, 084313.
- 246 [10] Fortenberry, R. C.; Huang, X.; Crawford, T. D.; Lee, T. J. Quartic Force Field Rovibrational  
247 Analysis of Protonated Acetylene,  $\text{C}_2\text{H}_3^+$ , and Its Isotopologues. *J. Phys. Chem. A* **2014**, *118*,  
248 7034–7043.

- 249 [11] Fortenberry, R. C.; Lee, T. J.; Müller, H. S. P. Excited Vibrational Level Rotational Constants for  
250 SiC<sub>2</sub>: A Sensitive Molecular Diagnostic for Astrophysical Conditions. *Molec. Astrophys.* **2015**, *1*,  
251 13–19.
- 252 [12] Kitchens, M. J. R.; Fortenberry, R. C. The Rovibrational Nature of Closed-Shell Third-Row Tri-  
253 atomics: HOX and HXO, X = Si<sup>+</sup>, P, S<sup>+</sup>, and Cl. *Chem. Phys.* **2016**, *472*, 119–127.
- 254 [13] Bizzocchi, L.; Lattanzi, V.; Laas, J.; Spezzano, S.; Giuliano, B. M.; Prudenzeno, D.; Endres, C.;  
255 Sipilä, O.; Caselli, P. Accurate Sub-millimetre Rest Frequencies for HOCO<sup>+</sup> and DOCO<sup>+</sup> Ions.  
256 *Astron. Astrophys.* **2017**, *602*, A34.
- 257 [14] Fortenberry, R. C. Quantum Astrochemical Spectroscopy. *Int. J. Quant. Chem.* **2017**, *117*, 81–91.
- 258 [15] Fortenberry, R. C.; Novak, C. M.; Layfield, J. P.; Matito, E.; Lee, T. J. Overcoming the Failure  
259 of Correlation for Out-of-Plane Motions in a Simple Aromatic: Rovibrational Quantum Chemical  
260 Analysis of *c*-C<sub>3</sub>H<sub>2</sub>. *J. Chem. Theor. Comput.* **2018**, *14*, 2155–2164.
- 261 [16] Martin, J. M. L.; Lee, T. J. The Atomization Energy and Proton Affinity of NH<sub>3</sub>. An *Ab Initio*  
262 Calibration Study. *Chem. Phys. Lett.* **1996**, *258*, 136–143.
- 263 [17] Dunning, T. H. Gaussian Basis Sets for Use in Correlated Molecular Calculations. I. The Atoms  
264 Boron through Neon and Hydrogen. *J. Chem. Phys.* **1989**, *90*, 1007–1023.
- 265 [18] Kendall, R. A.; Dunning, T. H.; Harrison, R. J. Electron Affinities of the First-Row Atoms Revisited.  
266 Systematic Basis Sets and Wave Functions. *J. Chem. Phys.* **1992**, *96*, 6796–6806.
- 267 [19] Peterson, K. A.; Dunning, T. H. Benchmark Calculations with Correlated Molecular Wave Func-  
268 tions. VII. Binding Energy and Structure of the HF Dimer. *J. Chem. Phys.* **1995**, *102*, 2032–2041.
- 269 [20] Fortenberry, R. C.; Huang, X.; Francisco, J. S.; Crawford, T. D.; Lee, T. J. Fundamental Vibrational  
270 Frequencies and Spectroscopic Constants of HOCS<sup>+</sup>, HSCO<sup>+</sup>, and Isotopologues via Quartic Force  
271 Fields. *J. Phys. Chem. A* **2012**, *116*, 9582–9590.
- 272 [21] Martin, J. M. L.; Taylor, P. R. Basis Set Convergence for Geometry and Harmonic Frequencies.  
273 Are *h* Functions Enough? *Chem. Phys. Lett.* **1994**, *225*, 473–479.
- 274 [22] Fortenberry, R. C.; Huang, X.; Crawford, T. D.; Lee, T. J. The 1 <sup>3</sup>A' HCN and 1 <sup>3</sup>A' HCO<sup>+</sup>  
275 Vibrational Frequencies and Spectroscopic Constants from Quartic Force Fields. *J. Phys. Chem.*  
276 *A* **2013**, *117*, 9324–9330.
- 277 [23] Cheng, Q.; Fortenberry, R. C.; DeYonker, N. J. Towards a Quantum Chemical Protocol for the  
278 Prediction of Rovibrational Spectroscopic Data for Transition Metal Molecules: Exploration of  
279 CuCN, CuOH, and CuCCH. *J. Chem. Phys.* **2017**, *147*, 234303.
- 280 [24] Douglas, M.; Kroll, N. Quantum Electrodynamical Corrections to the Fine Structure of Helium.  
281 *Ann. Phys.* **1974**, *82*, 89–155.
- 282 [25] Jansen, G.; Hess, B. A. Revision of the Douglas-Kroll Hamiltonian. *Phys. Rev. A* **1989**, *39*, 6016–  
283 6017.
- 284 [26] Fortenberry, R. C.; Huang, X.; Francisco, J. S.; Crawford, T. D.; Lee, T. J. The *trans*-HOCO  
285 Radical: Fundamental Vibrational Frequencies, Quartic Force Fields, and Spectroscopic constants.  
286 *J. Chem. Phys.* **2011**, *135*, 134301.
- 287 [27] Lattanzi, V.; Thorwirth, S.; Halfen, D. T.; Mück, L. A.; Ziurys, L. M.; Thaddeus, P.; Gauss, J.;  
288 McCarthy, M. C. Bonding in the Heavy Analogue of Hydrogen Cyanide: The Curious Case of  
289 Bridged HPSi. *Angew. Chem. Int. Ed. Engl.* **2010**, *49*, 5661–5664.
- 290 [28] Barone, V.; Biczysko, M.; Puzzarini, C. Quantum Chemistry Meets Spectroscopy for Astrochem-  
291 istry: Increasing Complexity toward Prebiotic Molecules. *Acc. Chem. Res.* **2015**, *48*, 1413–1422.
- 292 [29] Cernicharo, J.; McCarthy, M. C.; Gottlieb, C. A.; Agúndez, M.; Prieto, L. V.; Baraban, J. H.;  
293 Changala, P. B.; Guééin, M.; Kahane, C.; Martin-Drumel, M. A.; Patel, N. A.; Reilly, N. J.;  
294 Stanton, J. F.; Quintana-Lacaci, G.; Thorwirth, S. Discovery of SiCSi in IRC+10216: A Missing  
295 Link between Gas and Dust Carriers of Si-C Bonds. *Astrophys. J. Lett.* **2015**, *806*, L3.
- 296 [30] McCarthy, M. C.; Baraban, J. H.; Changala, P. B.; Stanton, J. F.; Martin-Drumel, M.-A.; Thor-  
297 wirth, S.; Gottlieb, C. A.; Reilly, N. J. Discovery of a Missing Link: Detection and Structure of the  
298 Elusive Disilicon Carbide Cluster. *J. Phys. Chem. Lett.* **2015**, *6*, 2107–2111.
- 299 [31] Cazzoli, G.; Lattanzi, V.; Kirsch, T.; Gauss, J.; Belén Tercero and, J. C.; Puzzarini, C. Laboratory  
300 Measurements and Astronomical Search for the HSO Radical. *Astron. Astrophys.* **2016**, *591*, A126.
- 301 [32] Finney, B.; Fortenberry, R. C.; Francisco, J. S.; Peterson, K. A. A Spectroscopic Case for SPSI  
302 Detection: The Third-Row in a Single Molecule. *J. Chem. Phys.* **2016**, *145*, 124311.
- 303 [33] Demaison, J.; Craig, N. C.; Gurusinghe, R.; Tubergen, M. J.; Rudolph, H. D.; Coudert,  
304 L. H.; Szalay, P. G.; Császár, A. G. Fourier Transform Microwave Spectrum of Propene-3-*d*<sub>1</sub>  
305 (CH<sub>2</sub>=CHCH<sub>2</sub>D), Quadrupole Coupling Constants of Deuterium, and a Semiexperimental Equi-  
306 librium Structure of Propene. *J. Phys. Chem. A* **2017**, *121*, 3155–3166.
- 307 [34] Puzzarini, C.; Barone, V. Diving for Accurate Structures in the Ocean of Molecular Systems with

- the Help of Spectroscopy and Quantum Chemistry. *Acc. Chem. Res.* **2018**, *51*, 548–556.
- [35] Puzzarini, C.; Heckert, M.; Gauss, J. The Accuracy of Rotational Constants Predicted by High-Level Quantum-Chemical Calculations. I. Molecules Containing First-Row Atoms. *J. Chem. Phys.* **2008**, *128*, 194108.
- [36] Alessandrini, S.; Gauss, J.; Puzzarini, C. Accuracy of Rotational Parameters Predicted by High-Level Quantum-Chemical Calculations: Case Study of Sulfur-Containing Molecules of Astrochemical Interest. *J. Chem. Theory Comput* **2018**, *14*, 5360–5371.
- [37] Yang, W. T.; Parr, R. G.; Lee, C. T. Various Functionals for the Kinetic Energy Density of an Atom or Molecule. *Phys. Rev. A* **1986**, *34*, 4586–4590.
- [38] Lee, C.; Yang, W. T.; Parr, R. G. Development of the Colle-Salvetti Correlation-Energy Formula into a Functional of the Electron Density. *Phys. Rev. B.* **1988**, *37*, 785–789.
- [39] Becke, A. D. Density-Functional Thermochemistry. III. The Role of Exact Exchange. *J. Chem. Phys.* **1993**, *98*, 5648–5652.
- [40] Piccardo, M.; Penocchio, E.; Puzzarini, C.; Biczysko, M.; Barone, V. Semi-Experimental Equilibrium Structure Determinations by Employing B3LYP/SNSD Anharmonic Force Fields: Validation and Application to Semirigid Organic Molecules. *J. Phys. Chem. A* **2015**, *119*, 2058–2082.
- [41] Carter, S.; Bowman, J. M.; Handy, N. C. Extensions and Tests of “Multimodes”: a Code to Obtain Accurate Vibration/Rotation Energies of Many-Mode Molecules. *Theor. Chem. Acc.* **1998**, *100*, 191–198.
- [42] Bowman, J. M.; Carter, S.; Huang, X. MULTIMODE: a Code to Calculate Rovibrational Energies of Polyatomic Molecules. *Int. Rev. Phys. Chem.* **2003**, *22*, 533–549.
- [43] Yu, Q.; Bowman, J. M.; Fortenberry, R. C.; Mancini, J. S.; Lee, T. J.; Crawford, T. D.; Klemperer, W.; Francisco, J. S. The Structure, Anharmonic Vibrational Frequencies, and Intensities of NNHNN<sup>+</sup>. *J. Phys. Chem. A* **2015**, *119*, 11623–11631.
- [44] Bassett, M. K.; Fortenberry, R. C. Symmetry Breaking and Spectral Considerations of the Surprisingly Floppy *c*-C<sub>3</sub>H Radical and the Related Dipole-Bound Excited State of *c*-C<sub>3</sub>H<sup>-</sup>. *J. Chem. Phys.* **2017**, *146*, 224303.
- [45] Cramer, C. J. *Essentials of Computational Chemistry: Theories and Models*, 2nd ed.; Wiley: West Sussex, England, 2004.
- [46] Lee, T. J.; Martin, J. M. L.; Taylor, P. R. An Accurate *ab Initio* Quartic Force Field and Vibrational Frequencies for CH<sub>4</sub> and Its Isotopomers. *J. Chem. Phys.* **1995**, *102*, 254–261.
- [47] Allen, W. D.; coworkers, 2005; *INTDER* 2005 is a General Program Written by W. D. Allen and Coworkers, which Performs Vibrational Analysis and Higher-Order Non-Linear Transformations.
- [48] Gaw, J. F.; Willets, A.; Green, W. H.; Handy, N. C. In *Advances in Molecular Vibrations and Collision Dynamics*; Bowman, J. M., Ratner, M. A., Eds.; JAI Press, Inc.: Greenwich, Connecticut, 1991; pp 170–185.
- [49] Martin, J. M. L.; Taylor, P. R. Accurate *ab Initio* Quartic Force Field for *trans*-HNNH and Treatment of Resonance Polyads. *Spectrochim. Acta A* **1997**, *53*, 1039–1050.
- [50] Mills, I. M. In *Molecular Spectroscopy - Modern Research*; Rao, K. N., Mathews, C. W., Eds.; Academic Press: New York, 1972; pp 115–140.
- [51] Watson, J. K. G. In *Vibrational Spectra and Structure*; Daring, J. R., Ed.; Elsevier: Amsterdam, 1977; pp 1–89.
- [52] Papousek, D.; Aliev, M. R. *Molecular Vibration-Rotation Spectra*; Elsevier: Amsterdam, 1982.
- [53] Mellau, G. C. Complete experimental rovibrational eigenenergies of HCN up to 6880 cm<sup>-1</sup> above the ground state. *J. Chem. Phys.* **2011**, *134*, 234303.
- [54] Amano, T. The  $\nu_1$  Fundamental Band of HCO<sup>+</sup> by Difference Frequency Laser Spectroscopy. *J. Chem. Phys.* **1983**, *79*, 3595.
- [55] Foster, S. C.; McKellar, A. R. W.; Sears, T. J. Observation of the  $\nu_3$  Fundamental Band of HCO<sup>+</sup>. *J. Chem. Phys.* **1984**, *81*, 578.
- [56] Davies, P. B.; Hamilton, P. A.; Rothwell, W. J. Infrared Laser Spectroscopy of the  $\nu_3$  Fundamental of HCO<sup>+</sup>. *J. Chem. Phys.* **1984**, *81*, 1598–1599.
- [57] Kawaguchi, K.; Yamada, C.; Saito, S.; Hirota, E. Magnetic Field Modulated Infrared Laser Spectroscopy of Molecular Ions: The  $\nu_2$  Band of HCO<sup>+</sup>. *J. Chem. Phys.* **1985**, *82*, 1750–1755.
- [58] Linnartz, H.; Speck, T.; Maier, J. P. High-Resolution Infrared Spectrum of the  $\nu_3$  Band in Ar–HCO<sup>+</sup>. *Chem. Phys. Lett.* **1998**, *288*, 504–508.
- [59] Cazzoli, G.; Cludi, L.; Buffa, G.; Puzzarini, C. Precise THz Measurements of HCO<sup>+</sup>, N<sub>2</sub>H<sup>+</sup>, and CF<sup>+</sup> for Astrophysical Observations. *Astrophys. J. Suppl. Ser.* **2012**, *203*, 1–9.
- [60] Mellau, G. C. The  $\nu_1$  band system of HNC. *J. Molec. Spectrosc.* **2010**, *264*, 2–9.
- [61] Mellau, G. C. Complete experimental rovibrational eigenenergies of HNC up to 3743 cm<sup>-1</sup> above



- the ground state. *J. Chem. Phys.* **2010**, *133*, 164303.
- [62] Morgan, W. J.; Fortenberry, R. C. Theoretical Rovibronic Treatment of the  $\tilde{X}^2\Sigma^+$  and  $\tilde{A}^2\Pi$  States of  $C_2H$  &  $\tilde{X}^1\Sigma^+$  State of  $C_2H^-$  from Quartic Force Fields. *J. Phys. Chem. A* **2015**, *119*, 7013–7025.
- [63] Stephens, J. W.; Yan, W.-B.; Richnow, M. L.; Solka, H.; Curl, R. F. Infrared Kinetic Spectroscopy of  $C_2H$  and  $C_2D$ . *J. Molec. Struct.* **1988**, *190*, 41–60.
- [64] Endo, Y.; Kanamori, H.; Hirota, E. Millimeter- and Submillimeter-Wave Spectra of the Vibrationally Excited CCD Radical. *Chem. Phys. Lett.* **1987**, *160*, 280–284.
- [65] Kanamori, H.; Seki, K.; Hirota, E. Infrared Diode Laser Kinetic Spectroscopy of the CCH Radical  $\nu_3$  Band. *J. Chem. Phys.* **1987**, *87*, 73–76.
- [66] Kanamori, H.; Hirota, E. Vibronic Bands of the CCH Radical Observed by Infrared Diode Laser Kinetic Spectroscopy. *J. Chem. Phys.* **1988**, *88*, 3962–3969.
- [67] Brünken, S.; Gupta, H.; Gottlieb, C. A.; McCarthy, M. C.; Thaddeus, P. Detection of the Carbon Chain Negative Ion  $C_8H^-$  in TMC-1. *Astrophys. J.* **2007**, *664*, L43–L46.
- [68] Ervin, K.; Lineberger, W. C. Photoelectron Spectra of  $C_2$  and  $C_2H^-$ . *J. Phys. Chem.* **1991**, *95*, 1167–1177.
- [69] Cabezas, C.; Cernicharo, J.; Alonso, J. L.; Agúndez, M.; Mata, S.; Guélin, M.; Peña, I. Laboratory and Astronomical Discovery of Hydromagnesium Isocyanide. *Astrophys. J.* **2013**, *775*, 133.
- [70] Tack, L. M.; Rosenbaum, N. H.; Owrutsky, J. C.; Saykally, R. J. Velocity Modulation Infrared Laser Spectroscopy and Structure of the Amide Anion ( $NH_2^-$ ). *J. Chem. Phys.* **1986**, *85*, 4222–4227.
- [71] Tennyson, J.; Zobov, N. F.; Williamson, R.; Polyansky, O. L. Experimental Energy Levels of the Water Molecule. *J. Phys. Chem. Ref. Data* **2001**, *30*, 735.
- [72] Johns, J. W. C.; McKellar, A. R. W.; Weinberger, E. The Infrared Spectrum of HNO. *Canadian. J. Phys.* **1983**, *61*, 1106–1119.
- [73] Fortenberry, R. C.; Francisco, J. S. On the Detectability of the  $\tilde{X}^2A'$  HSS, HSO, and HOS Radicals in the Interstellar Medium. *Astrophys. J.* **2017**, *835*, 243.
- [74] Yoshikawa, T.; Watanabe, A.; Sumiyoshi, Y.; Endo, Y. Laser Spectroscopy of the  $\tilde{A}^2A' - \tilde{X}^2A'$  System for the HSO Radical. *J. Molec. Spectrosc.* **2009**, *254*, 119–125.
- [75] Yamamoto, S.; Saito, S. Microwave Spectrum and Molecular Structure of the  $HS_2$  Radical. *Can. J. Phys.* **1994**, *72*, 954–962.
- [76] Fortenberry, R. C.; Francisco, J. S. Factors Affecting the Spectroscopic Observation of Bridged HPSi and Linear HSiP in Astrophysical/Interstellar Media. *Astrophys. J.* **2017**, *843*, 124.
- [77] Gottlieb, C. A.; Vrtilek, J. M.; Thaddeus, P. Laboratory Measurement of the Rotational Spectrum of SiCC. *Astrophys. J.* **1989**, *349*, L29–L32.
- [78] Butenhoff, T. J.; Rohlffing, E. A. Laser-Induced Fluorescence Spectroscopy of Jet-Cooled SiC<sub>2</sub>. *J. Chem. Phys.* **1991**, *95*, 1–8.
- [79] Dowling, J. M. The Rotation-Inversion Spectrum of Ammonia. *J. Molec. Spectrosc.* **1968**, *27*, 527–538.
- [80] Shimanouchi, T. *Tables of Molecular Vibrational Frequencies*, 39th ed.; National Standards Reference Data System: Washington, DC, 1972; Vol. 1.
- [81] Jacox, M. E. Vibrational and Electronic Energy Levels of Polyatomic Transient Molecules. *J. Phys. Chem. Ref. Data* **1994**, *23*, 1–461.
- [82] Brünken, S.; Gottlieb, C. A.; McCarthy, M. C.; Thaddeus, P. Laboratory Detection of HOCN and Tentative Identification in Sgr B2. *Astrophys. J.* **2009**, *697*, 880.
- [83] Fortenberry, R. C.; Huang, X.; Francisco, J. S.; Crawford, T. D.; Lee, T. J. Vibrational Frequencies and Spectroscopic Constants from Quartic Force Fields for *cis*-HOCO: The Radical and the Anion. *J. Chem. Phys.* **2011**, *135*, 214303.
- [84] Oyama, T.; Funato, W.; Sumiyoshi, Y.; Endo, Y. Observation of the pure rotational spectra of *trans*- and *cis*-HOCO. *J. Chem. Phys.* **2011**, *134*, 174303.
- [85] Bogey, M.; Demuynck, C.; Destombes, J. L.; Krupnov, A. Molecular Structure of HOCO<sup>+</sup>. *J. Molec. Spectrosc.* **1988**, *190*, 465–474.
- [86] Amano, T. Spectroscopic Detection of Protonated N<sub>2</sub>O. *Chem. Phys. Lett.* **1986**, *127*, 101–104.
- [87] Bogey, M.; Demuynck, C.; Destombes, J. L. Millimeter and Submillimeter Wave Spectroscopy of Protonated and Deuterated Nitrous Oxide. *J. Chem. Phys.* **1988**, *88*, 2108–2111.
- [88] Thaddeus, P.; Vrtilek, J. M.; Gottlieb, C. A. Laboratory and Astronomical Detection of Cyclopropanylidene, C<sub>3</sub>H<sub>2</sub>. *Astrophys. J.* **1985**, *296*, L63–L66.
- [89] Dateo, C. E.; Lee, T. J.; Schwenke, D. W. An Accurate Quartic Force-Field and Vibrational Frequencies for HNO and DNO. *J. Chem. Phys.* **1994**, *101*, 5853–5859.

- 425 [90] Agbaglo, D.; Fortenberry, R. C. The Performance of CCSD(T)-F12/aug-cc-pVTZ for the Compu-  
426 tation of Anharmonic Fundamental Vibrational Frequencies. *Int. J. Quantum Chem.* **2019**, *119*,  
427 e25899.
- 428 [91] Agbaglo, D.; Fortenberry, R. C. The Performance of Explicitly Correlated Wavefunctions  
429 [CCSD(T)-F12b] in the Computation of Anharmonic Vibrational Frequencies. *Chem. Phys. Lett.*  
430 **2019**, *734*, 136720.
- 431 [92] Huang, X.; Valeev, E. F.; Lee, T. J. Comparison of One-Particle Basis Set Extrapolation to Ex-  
432 plicitly Correlated Methods for the Calculation of Accurate Quartic Force Fields, Vibrational Fre-  
433 quencies, and Spectroscopic Constants: Application to H<sub>2</sub>O, N<sub>2</sub>H<sup>+</sup>, NO<sub>2</sub><sup>+</sup>, and C<sub>2</sub>H<sub>2</sub>. *J. Chem.*  
434 *Phys.* **2010**, *133*, 244108.

## CHAPTER IV: NOBLE GASES IN RELATIVELY COMPLEX MOLECULES

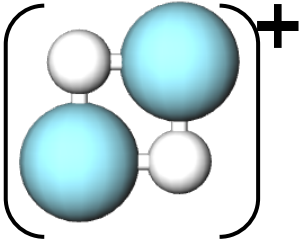
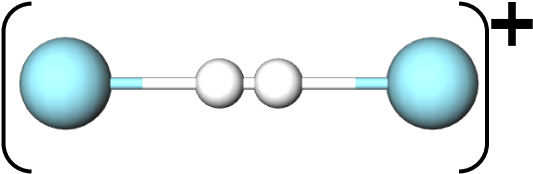
### A. Introduction

Noble gases have not been found to participate often in molecular bonding, but there are some significant bonds that include noble gas atoms. One particularly interesting molecule is  $\text{HeH}^+$  because it is composed of the most abundant elements in the universe even though Helium is not commonly involved in molecular interactions on Earth. Thorfin R. Hogness and E. G. Lunn first synthesized the helium hydride ( $\text{HeH}^+$ ) molecule in 1925 <sup>[28]</sup>, from which chemists in the 1970s speculated that it could be observed in interstellar bodies, and eventually deduced that it was the first molecule formed in the universe <sup>[29]</sup>. Recently, a spectral line matching that expected for  $\text{HeH}^+$  was observed in planetary nebula NGC 7027 using NASA's SOFIA telescope <sup>[29]</sup> <sup>[30]</sup>. This accomplishment shows many areas of astrochemistry being used in conjunction as discussed in Chapter I, combining laboratory experiments, computational predictions, and directing the vibrational spectroscopy SOFIA based on ground-based rotational spectroscopy <sup>[28]</sup> <sup>[29]</sup> <sup>[30]</sup>. The unique spectra observed in the  $\text{HeH}^+$  molecule and the proton-bound  $\text{HeHHe}^+$  <sup>[31]</sup> implies that larger combinations of these bonds could create additional distinctive spectra in interstellar bodies like NGC 7027 <sup>[30]</sup>. However, synthesizing noble gas molecules on Earth is incredibly difficult, evidenced by the lack of experimental data. Computational chemistry ideally can fill in some of the gaps on the path toward astronomically observing such rare molecules.

Helium, Neon, and Argon are chosen to study in this project because they are small enough and have few enough electrons to compute using the methods described previously. Additionally, they are also known to be among the ten most abundant atoms in the universe <sup>[32]</sup>.

Based on the bonding of  $\text{HeH}^+$  [30] and brightness of  $\text{HeHHe}^+$  [31], similar patterns of bonding are designed with common characteristics: two hydrogen atoms and two noble gas (Ng) atoms. For example,  $\text{HeHHHe}^+$  represents a dimer of the  $\text{HeH}^+$  molecule, which has already been observed and determined to be of importance to astrochemistry [29] [30]. Additionally, it contains the most common molecule in the universe in  $\text{H}_2$ , and the second-most abundant atoms in the universe in He [32], making any potential molecule based on these moieties highly likely to exist in the universe even if relatively rare compared to their constituents. Two main structural isomers are investigated for geometrical optimization: a linear and cyclic combination. These two combinations offer multiple molecules of study by combining the three noble gases included in this project: He, Ne, Ar. This yielded twelve possible molecules shown in Table 3 below:

Table IV.1: Noble Gas Molecules of Interest

Structure	Molecular Example	Possible Formulas
Cyclic $\text{Ng}_2\text{H}_2^+$		$\text{He}_2\text{H}_2^+$
		$\text{Ne}_2\text{H}_2^+$
		$\text{Ar}_2\text{H}_2^+$
		$\text{NeHeH}_2^+$
		$\text{ArHeH}_2^+$
		$\text{ArNeH}_2^+$
Linear $\text{NgHHNg}^+$		$\text{HeHHHe}^+$
		$\text{NeHHNe}^+$
		$\text{ArHHA}^+$
		$\text{NeHHHe}^+$
		$\text{ArHHHe}^+$
		$\text{ArHHNe}^+$

Ball-and-stick models are used to represent the large noble gas atoms and small hydrogen atoms. All six combinations were studied in both the cyclic and linear structures.

## B. Methods

After narrowing the list of twelve molecules to six linear molecules, the six linear structures were analyzed using the complete CcCR method, followed by the F12-TZ method.

This decision is made in the hopes of showing consistency between the CBS CcCR extrapolation and the explicit correlation of F12-TZ.

The intensities for each molecule’s infrared active modes are computed using the MP2/6-31+G level of theory, which optimizes the geometry and computes the intensities of harmonic vibrational frequencies [33]. This level of theory was chosen based on the high speed of computation because the intensity does not rely heavily on the accuracy of the geometry and harmonic frequencies.

### C. Results and Conclusions

Table IV.2: F12-TZ and CcCR Vibrational Frequencies ( $\text{cm}^{-1}$ ) and Intensities ( $\text{km mol}^{-1}$ )

Molecule	Mode	Symmetry	Description	F12-TZ Harmonic $\text{cm}^{-1}$	CcCR Harmonic $\text{cm}^{-1}$	CcCR Anharmonic $\text{cm}^{-1}$	IR Intensity $\text{Km/mol}$
HeHHHe <sup>+</sup>	$\omega_1$	$\sigma_g$	H-H stretch	2314	2314	2074	
	$\omega_2$	$\sigma_g$	symmetric He-H-H bend	397	399	269	303
	$\omega_3$	$\sigma_u$	symmetric He-H stretch	279	280		
	$\omega_4$	$\pi_u$	antisymmetric He-H stretch	706	706		918
	$\omega_5$	$\pi_g$	antisymmetric He-H-H bend	231	228		
NeHHNe <sup>+</sup>	$\omega_1$	$\sigma_g$	H-H stretch	2355	2356	2109	
	$\omega_2$	$\sigma_u$	symmetric Ne-H stretch	504	499	—	
	$\omega_3$	$\sigma_g$	symmetric Ne-H-H bend	220	219		235
	$\omega_4$	$\pi_u$	antisymmetric Ne-H stretch	709	679		2111
	$\omega_5$	$\pi_g$	antisymmetric Ne-H-H bend	193	186		
ArHHAr <sup>+</sup>	$\omega_1$	$\sigma_g$	H-H stretch	2280	2274	2429	
	$\omega_2$	$\sigma_u$	symmetric Ar-H stretch	924	925	807	
	$\omega_3$	$\sigma_g$	symmetric Ar-H-H bend	215	214		62
	$\omega_4$	$\pi_u$	antisymmetric Ar-H stretch	800	801		2983
	$\omega_5$	$\pi_g$	antisymmetric Ar-H-H bend	192	179		
NeHHHe <sup>+</sup>	$\omega_1$	$\sigma$	H-H stretch	2162	2170	1742	817
	$\omega_2$	$\sigma$	Ne-H stretch	760	743	606	1165
	$\omega_3$	$\sigma$	He-H stretch	148	149		16
	$\omega_4$	$\pi$	Ne-H-H bend	677	650		12
	$\omega_5$	$\pi$	He-H-H bend	145	144		229
ArHHHe <sup>+</sup>	$\omega_1$	$\sigma$	Ar-H stretch	1549	1550	1441	1964
	$\omega_2$	$\sigma$	H-H stretch	1137	1136	872	394
	$\omega_3$	$\sigma$	He-H stretch	79	78		8
	$\omega_4$	$\pi$	Ar-H-H bend	662	648		14
	$\omega_5$	$\pi$	He-H-H bend	34	26		89
ArHHNe <sup>+</sup>	$\omega_1$	$\sigma$	Ar-H stretch	1621	1613	1308	1805
	$\omega_2$	$\sigma$	H-H stretch	1157	1155	1054	379
	$\omega_3$	$\sigma$	Ne-H stretch	64	62		20
	$\omega_4$	$\pi$	Ar-H-H bend	673	650		10
	$\omega_5$	$\pi$	Ne-H-H bend	64	43		99

The results from these six linear molecules’ F12-TZ and CcCR QFFs are compiled in Table 4 above. Immediately of note is that only the first two anharmonic CcCR vibrational frequencies for each molecule were reliable enough to include. The other nonincluded frequencies were inconsistent and showed signs of being highly influenced by noise than by clear

computation, but the reported modes were consistent across both QFFs and justified in including in the table.

Given that F12-TZ and CcCR are known to produce vibrational frequencies within  $7\text{ cm}^{-1}$  of experimental data, while CcCR produces better rotational constants to within 34 MHz of experimental data <sup>[18] [19] [20] [21] [27]</sup>, Table 4 shows that the F12-TZ Harmonic and CcCR Harmonic frequencies are highly similar, often within a few wavenumbers of each other. This agreement suggests that the harmonic force constants in each method are similar, and the methods likely diverge in the anharmonic treatment in SPECTRO. However, only the first two vibrational modes from these molecules were reported. The other three seemed to reflect large amounts of noise rather than real data because of significant differences between the F12-TZ and CcCR results in addition to non-physical results and questionable behavior. This is likely due to the weak bonding present in many of these systems between the noble gas atom, especially Neon, and the small Hydrogen atom. This is also a consequence of the QFF only reaching the fourth-order expansion; with higher order, the accuracy should increase <sup>[5] [6]</sup>.

A positive anharmonicity is observed in  $\nu_1$  for  $\text{ArHHAr}^+$ . This is of concern given the expectation that anharmonic frequencies should be smaller than harmonic frequencies <sup>[2]</sup>, but this result could likely be another effect of the previously described weak bonding in the molecule. Further study should be used to study these lower frequency modes, especially within the vibrational bends of these molecules.

These results are likely useful because they follow the predicted effect of noble gases on known frequencies. Particularly, the anharmonic H-H stretch in the  $\text{H}_2^+$  molecule has been found at  $2191\text{ cm}^{-1}$  <sup>[34]</sup>, and the addition of noble gas ligands to this structure causes the frequency to drop substantially, as noted across all six molecules' CcCR anharmonic frequencies. This agrees

with quantum chemical theory according to the inverse relationship between mass and frequency. However, the magnitude of this change is harder to predict and did not show a correlation between ligand mass and change in frequency, likely due to other characteristics like bond strength and length. Therefore, these computed frequencies are likely reliable enough to begin comparing with astronomical observations.

Several unique intensities are noticed for these molecules as well. In the three molecules of  $\text{HeHHHe}^+$ ,  $\text{NeHHNe}^+$ , and  $\text{ArHHAr}^+$ , only two frequencies are infrared active due to the internal symmetry exhibited. The same vibrational frequency is the most intense in all six molecules, however, the stretch between a noble gas and hydrogen, and this intensity only increases as the mass of the noble gas increases. This suggests that spectroscopic observations of these molecules will show incredibly bright modes at these corresponding frequencies, which can be incredibly useful for molecular identification. The largest intensity across all molecules is found to be 2983 (km/mol) for the antisymmetric Ar-H stretch in  $\text{ArHHAr}^+$ , which is over one order of magnitude larger than the most intense 54-72 km/mol bending mode of water <sup>[35]</sup>. This  $\text{ArHHAr}^+$  brightness is uncommonly strong and implies that this stretch in particular will be evident if directly observed using infrared spectroscopy. Some molecules may produce vibrational frequencies in the similar range of the vibrations discussed above so their unique infrared brightness can directly narrow the possible molecules to the ones that occupy both that frequency and intensity.

Table IV.3: CcCR Rotational Constants

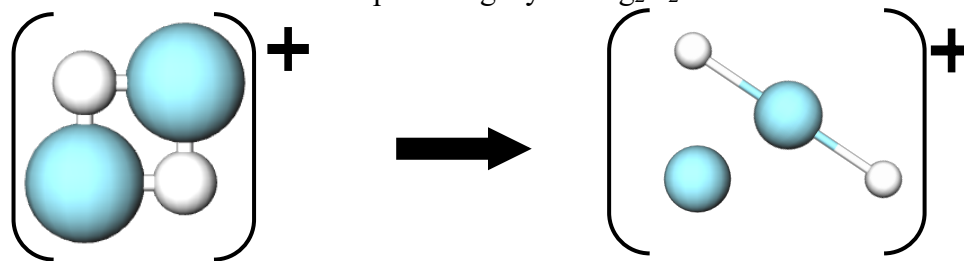
Molecule	$B_e$ MHz	$B_0$ MHz	$B_1$ MHz	$B_2$ MHz	$D_e$	$H_e$
$\text{HeHHHe}^+$	19749	18097	17277	17511	214 kHz	-2274 mHz
$\text{NeHHNe}^+$	3358	3282	3150	3057	3522 Hz	-2733 $\mu\text{Hz}$
$\text{ArHHAr}^+$	1525	1530	1479	1481	344 Hz	14 $\mu\text{Hz}$
$\text{NeHHHe}^+$	9451	8696	7272	9331	141 kHz	-12 Hz
$\text{ArHHHe}^+$	5870	5913	5437	5432	135 kHz	-16 Hz
$\text{ArHHNe}^+$	1840	1824	1571	1780	7058 Hz	-108 mHz

Given the fact that CcCR QFFs produce more accurate rotational constants than F12-TZ QFFs, only the CcCR rotational constants are included in the table above. The data in this table are likely reliable given that all six molecules share the basic linear geometry and differ in the ligand masses; therefore, theory suggests that the rotational constants should decrease with increased mass, which is consistent from the least massive HeHHHe<sup>+</sup> to the most massive ArHHAr<sup>+</sup> of the six molecules. Performing another type of QFF, possibly B3LYP, could help support this data being used for identification of these molecules in space.

By using further corrections for higher-order electron correlation and quantum electrostatics in CcCR QFFs, perhaps the results could become clearer, but such results are unlikely because these changes are most frequently noted in molecules containing atoms beyond the third row while increasing the computational cost and time immensely. Therefore, adding these corrections was deemed unnecessary and inefficient for the project at hand. Computing a global potential energy surface rather than a QFF in this project is one step that should hopefully improve these results.

Upon optimization of the CcCR geometry, the first step in effective QFFs, only the linear structures would actually engage in complete bonding. All six cyclic structures failed to optimize because one noble gas atom bonded to both hydrogen atoms while the other noble gas atom effectively trailed off without bonding. As shown in Figure 1 below, this effectively turns the cyclic structure into another linear structure, HNgH<sup>+</sup>.

Figure IV.1: Dissociation while Optimizing Cyclic Ng<sub>2</sub>H<sub>2</sub><sup>+</sup> Molecules





The inability of the proposed cyclic isomers to form an optimized geometry suggests that this cyclic structure is more of an intermediate or transition state of the collision between two  $\text{NgH}^+$  molecules than a bonded molecule. While this does not provide useful results for this project, further studies could investigate this collision and other dissociations of these molecules. For example, the differences between  $\text{HeHHHe}^+$  and  $\text{HeHH}^+$  may help determine which parts of these molecules are muddying the anharmonic frequencies. The dissociation energies of components like  $\text{HeHH}^+$ ,  $\text{HeH}^+$ , and  $\text{He}$  could also show the strength of the molecule's full bond. By determining how the frequencies change upon these dissociations, the interaction of these molecules and components may be better observed in highly energetic media, such as nebulae, which would lend insight into possible observation of the complete  $\text{NgHHNg}^+$  molecules.

## CHAPTER V: CONCLUSION

The F12-TZ and CcCR methods have been shown to have applicability in astrochemistry through accurate computation of observable characteristics of astromolecules. F12-TZ is reliable enough for computing vibrational frequencies while CcCR is even more reliable for computing rotational constants of various astromolecules <sup>[21]</sup> <sup>[27]</sup>. The accuracy between experimental and computational determination of these characteristics show that computational methods are reliable and effective for the time and cost associated with them. This fine balance has been successful for predicting vibrational and rotational spectra of silicon oxides and various interstellar molecules. This accuracy can also be extended to rare noble gas molecules discussed previously, which suggests that computational methods can be applied to an endless array of molecules that may not have even been imagined yet. Computational astrochemistry has implications not only on finding molecules in tremendous interstellar bodies but also on expanding the breadth of future molecules that can be imagined and eventually synthesized in laboratories.

## LIST OF REFERENCES:

1. Barone, V.; Biczysko, M.; Puzzarini, C. Quantum Chemistry Meets Spectroscopy for Astrochemistry: Increasing Complexity toward Prebiotic Molecules. *Accounts of Chemical Research* 2015, 48 (5), 1413–1422.
2. Sathyanarayana, D. N. *Vibrational spectroscopy: theory and applications*; New Age International: New Delhi, 2004.
3. Galabov, B.; Dudev, T.; Durig, J. R.; Orville-Thomas, W. J. Computations in Vibrational Intensity Spectroscopy. *Journal of Molecular Structure* 1988, 173, 111–128.
4. NASA Under the Spotlight: The Decade-Late, Over-Budget Arrival of SOFIA Shows That NASA's Practices Need to Change. *Nature* 2010, 466 (7305), 413–413.
5. C.J. Cramer, *Essentials of Computational Chemistry: Theories and Models*, second ed., Wiley, West Sussex, England, 2004.
6. T.J. Lee, J.M.L. Martin, P.R. Taylor, An accurate ab initio quartic force field and vibrational frequencies for CH<sub>4</sub> and its isotopomers, *J. Chem. Phys.* 102 (1995) 254-261.
7. Fortenberry, R.C., Lee, T.J., 2019. Computational vibrational spectroscopy for the detection of molecules in Space. *Annu. Rep. Comput. Chem.* 15, 173–202.
8. Raghavachari, K., Trucks, G.W., Pople, J.A., Head-Gordon, M., 1989. A fifth-order perturbation comparison of electron correlation theories. *Chem. Phys. Lett.* 157, 479–483.
9. Shavitt, I., Bartlett, R.J., 2009. *Many-body Methods in Chemistry and Physics: MBPT and Coupled-Cluster Theory*. Cambridge University Press, Cambridge.
10. Hill, J.G., Peterson, K.A., 2010. Correlation consistent basis sets for explicitly correlated wavefunctions: valence and core-valence basis sets for Li, Be, Na, and Mg. *Phys. Chem. Chem. Phys.* 12, 10460–10468.

11. Prascher, B.P., Woon, D.E., Peterson, K.A., Dunning, T.H., Wilson, A.K., 2011. Gaussian basis sets for use in correlated molecular calculations. VII. Valence, core-valence, and scalar relativistic basis sets for Li, Be, Na, and Mg. *Theor. Chem. Acc.* 128, 69–82.
12. Werner, H.-J., Knowles, P.J., Knizia, G., Manby, F.R., Schütz, M., 2012. Molpro: a general-purpose quantum Chemistry program package. *WIREs Comput. Mol. Sci.* 2, 242–253.
13. Werner, H.-J., et al., 2015. MOLPRO, version 2015.1, a package of ab Initio Programs. see. <http://www.molpro.net>.
14. J.M.L. Martin, P.R. Taylor, Basis set convergence for geometry and harmonic frequencies. Are h functions enough?, *Chem. Phys. Lett.* 225 (1994) 473–479.
15. M. Douglas, N. Kroll, Quantum electrodynamical corrections to the fine structure of helium, *Ann. Phys.* 82 (1974) 89–155.
16. G. Jansen, B.A. Hess, Revision of the Douglas-Kroll Hamiltonian, *Phys. Rev. A* 39 (1989) 6016–6017.
17. R.C. Fortenberry, X. Huang, J.S. Francisco, T.D. Crawford, T.J. Lee, The trans- HOCO radical: fundamental vibrational frequencies, quartic force fields, and spectroscopic constants, *J. Chem. Phys.* 135 (2011) 134301.
18. D. Agbaglo, R.C. Fortenberry, The performance of CCSD(T)-F12/aug-cc-pVTZ for the computation of anharmonic fundamental vibrational frequencies, *Int. J. Quant. Chem.* 119 (2019) e25899.
19. D. Agbaglo, R.C. Fortenberry, The performance of explicitly correlated wavefunctions [CCSD(T)-F12b] in the computation of anharmonic vibrational frequencies, *Chem. Phys. Lett.* 734 (2019) 136720.

20. X. Huang, E.F. Valeev, T.J. Lee, Comparison of one-particle basis set extrapolation to explicitly correlated methods for the calculation of accurate quartic force fields, vibrational frequencies, and spectroscopic constants: application to H<sub>2</sub>O, N<sub>2</sub>H<sup>+</sup>, NO<sup>+</sup>, and C<sub>2</sub>H<sub>2</sub>, *J. Chem. Phys.* 133 (2010) 244108.
21. Gardner, M. B.; Westbrook, B. R.; Fortenberry, R. C.; Lee, T. J. Highly-Accurate Quartic Force Fields for the Prediction of Anharmonic Rotational Constants and Fundamental Vibrational Frequencies. *Spectrochimica Acta Part A: Molecular and Biomolecular Spectroscopy* 2021, 248, 119184.
22. X. Huang, T.J. Lee, A procedure for computing accurate Ab initio quartic force fields: application to HO<sub>2</sub><sup>+</sup> and H<sub>2</sub>O, *J. Chem. Phys.* 129 (2008) 044312.
23. W.D. Allen, coworkers, INTDER 2005 is a General Program Written by W.D. Allen and Coworkers, which Performs Vibrational Analysis and Higher-Order Non-Linear Transformations, 2005..
24. J.F. Gaw, A. Willets, W.H. Green, N.C. Handy, in: J.M. Bowman, M.A. Ratner (Eds.), *Advances in Molecular Vibrations and Collision Dynamics*, JAI Press Inc, Greenwich, Connecticut, 1991, pp 170–185..
25. D. Papousek, M.R. Aliev, *Molecular Vibration-Rotation Spectra*, Elsevier, Amsterdam, 1982.
26. J.M.L. Martin, P.R. Taylor, Accurate ab initio quartic force field for trans-HNNH and treatment of resonance polyads, *Spectrochim. Acta A* 53 (1997) 1039– 1050.
27. Gardner, M. B.; Westbrook, B. R.; Fortenberry, R. C. Spectral Characterization for Small Clusters of Silicon and Oxygen: SiO<sub>2</sub>, SiO<sub>3</sub>, Si<sub>2</sub>O<sub>3</sub>, & Si<sub>2</sub>O<sub>4</sub>. *Planetary and Space Science* 2020, 193, 105076.

28. Helium hydride. <https://www.acs.org/content/acs/en/molecule-of-the-week/archive/h/helium-hydride.html> (accessed Mar 24, 2021).
29. Lemonick, S. Universe's First Molecules Finally Found in Space. *C&EN Global Enterprise* 2019, 97 (16), 5–5.
30. Güsten, R.; Wiesemeyer, H.; Neufeld, D.; Menten, K. M.; Graf, U. U.; Jacobs, K.; Klein, B.; Ricken, O.; Risacher, C.; Stutzki, J. Astrophysical Detection of the Helium Hydride Ion HeH<sup>+</sup>. *Nature* 2019, 568 (7752), 357–359.
31. Fortenberry, R. C.; Wiesenfeld, L. A Molecular Candle Where Few Molecules Shine: HeHHe<sup>+</sup>. *Molecules* 2020, 25 (9), 2183.
32. McCall, B.J., 2006. Dissociative recombination of cold H<sub>3</sub><sup>+</sup> and its interstellar implications. *Phil. Trans. Royal Soc. A* 364, 2953–2963.
33. Balci, K.; Akyuz, S. A Vibrational Spectroscopic Investigation on Benzocaine Molecule. *Vibrational Spectroscopy* 2008, 48 (2), 215–228.
34. CCCBDB Listing of experimental vibrational data for H<sub>2</sub><sup>+</sup> (Hydrogen cation). <https://cccbdb.nist.gov/expvibs2x.asp#webbook> (accessed Mar 24, 2021).
35. Bégué, D.; Baraille, I.; Garrain, P. A.; Dargelos, A.; Tassaing, T. Calculation of IR Frequencies and Intensities in Electrical and Mechanical Anharmonicity Approximations: Application to Small Water Clusters. *The Journal of Chemical Physics* 2010, 133 (3), 034102.