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1 Shock effects in feldspars: an overview

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18 **ABSTRACT**

19 **Feldspars are the dominant mineral in the crust of most terrestrial planetary bodies,**
20 **including Earth, Earth's moon, and Mars, as well as in asteroids, and thus in meteorites.**

21 **These bodies have experienced large numbers of hypervelocity impact events and so it is**
22 **important to have a robust understanding of the effect that shock waves exert on feldspars.**
23 **However, due to their optical complexity and susceptibility to weathering, feldspars are**
24 **under-utilized as shock barometers and indicators of hypervelocity impact. Here, we**
25 **provide an overview of the work done on shocked feldspars so far, in an effort to better**
26 **frame the current strengths and weaknesses of different techniques, and to highlight some**
27 **gaps in the literature.**

28 **1. INTRODUCTION**

29 Hypervelocity impacts result in distinctive microstructural deformation in rocks and
30 minerals during passage of a shock wave (e.g., French and Short, 1968; Roddy et al., 1978;
31 Stöffler and Grieve, 2007; French and Koeberl, 2010). This shock metamorphism records the
32 pressure, temperature, and strain rate conditions experienced by geo-materials during crater
33 formation. Understanding the shock metamorphic overprint of planetary materials is crucial to
34 the study of meteorites, terrestrial impact structures, and samples returned to Earth by space
35 missions. Clear criteria for characterizing and quantifying the degree of shock metamorphism are
36 required together with an understanding of how shock metamorphism affects different
37 lithologies.

38 Shock metamorphic effects vary between rock types due differences in the mineralogy
39 and microstructural characteristics of the target (e.g., French and Koeberl, 2010; Ferrière and
40 Osinski, 2013 and references therein). This observation has been used to propose multiple shock
41 classification schemes for different rock types including meteorites (e.g., Fritz et al., 2017),
42 quartzofeldspathic rocks, sandstones, basaltic-gabbroic rocks, particulate rock material,
43 chondritic and dunitic rocks (Stöffler et al., 2018). Shock effects are mineralogically selective,

44 and heterogeneous in spatial distribution, such that neighbouring minerals, even if the same
45 composition, may record different effects. This is due to differences in the shock impedance, a
46 physical property of matter that describes how efficiently a shock wave propagates through a
47 material. The mineral grains and pore spaces composing a polymineralic rock have different
48 shock impedances, hence the shock wave propagates with variable efficiency, and this leads to a
49 heterogeneous distribution of the maximum shock pressure and resultant shock metamorphic
50 effects within a sample (e.g., Ogilvie et al., 2011). In addition, the degree and type of
51 deformation is affected by the mineral's crystal structure and composition (Stöffler, 1972), pre-
52 impact temperature (e.g., Langenhorst et al., 1992; Huffman et al., 1993; Langenhorst and
53 Deutsch, 1994; Fritz et al., 2011) and pre-impact stress conditions (Daniel et al., 1997; Sims et
54 al., 2019). Shock deformation of most minerals is evidenced by a progressive degradation of the
55 crystal structure until melting occurs.

56 Feldspar group minerals (alkali feldspar and plagioclase) are both chemically and
57 optically complex. Various publications address shock metamorphism for plagioclase, a major
58 component in planetary materials, including asteroids, Mars, Earth, and Earth's moon. However,
59 the K-rich alkali feldspars that are rare in meteorites but abundant in terrestrial rocks, have not
60 been studied in as much detail. Feldspar group minerals suffer from alteration in a hydrous
61 environment rendering them challenging to study using solely optical microscopy. Additionally,
62 pre-existing microstructures, such as twin planes, cleavage planes, and exsolution lamellae,
63 might be masking or possibly even precluding the formation of shock-generated planar features.
64 The aforementioned challenges to studying feldspar microstructures coupled with inconsistencies
65 in the features associated with different shock levels in different rock types (e.g., Stöffler et al.,
66 2018; Fritz et al., 2017) make the utility of feldspars as a shock thermo-barometer variable.

67 Despite their abundance in rocky planetary materials, feldspars have been under-utilized
68 for shock barometry compared to quartz and olivine. However, feldspar is becoming increasingly
69 popular for shock barometry (e.g., Jaret et al., 2014; Kayama et al., 2018; Fritz et al., 2019a). In
70 this paper, we discuss all impact cratering related effects on the feldspar group, but the reader
71 should note that only definitive planar deformation features (PDFs), and diaplectic glass are
72 considered *diagnostic shock effects* on their own (French and Koeberl, 2010). Certain high-
73 pressure polymorphs can also be diagnostic, but only given a proper geologic context. Diagnostic
74 shock effects are those that only form via passage of a shock wave with a strain rate and a
75 pressure-temperature-time (P-T-t) path that cannot be produced by endogenic processes.
76 Additional microstructures can form during impact cratering events but can also form by other
77 geological processes. For example, fractures, kink bands, mosaicism, and darkening can inform
78 the degree to which feldspars have been shocked, but these features can also be produced
79 endogenically. Even tectonically-induced amorphization is possible, although only under
80 extreme and rare conditions such as those reported for silica by Janssen et al. (2010).

81 **2. Feldspar Group Minerals**

82 Feldspars are the most abundant mineral group in crustal rocks on Earth (Deer et al.,
83 2001), as well as in the crustal rocks of terrestrial planetary bodies (Papike et al., 1991). The
84 feldspar group is composed of two solid solutions: alkali feldspar ($\text{NaAlSi}_3\text{O}_8$ - KAlSi_3O_8) with a
85 density of 2.62 - 2.55 g/cm^3 and plagioclase feldspar ($\text{NaAlSi}_3\text{O}_8$ - $\text{CaAl}_2\text{Si}_2\text{O}_8$) with a density of
86 2.62 - 2.76 g/cm^3 . Intergrowths of alkali and plagioclase feldspars are common, such as perthite
87 (albite lamellae in K-feldspar host) and antiperthite (K-feldspar lamellae in albite). The crystal
88 system is triclinic or monoclinic, depending on composition, thermal history, and crystallization
89 temperature. Planar microstructures that are common in feldspar are good cleavage ($\{001\}$ and

90 {010}) and the formation of twins (Seifert, 1964; Zoltai and Stout, 1984; Xu et al., 2016 and
91 references therein). Common twins in alkali-feldspar are Carlsbad, with the z-axis as twin axis,
92 and Baveno, with (021) as twin axis, and a combination of albite and pericline twinning,
93 resembling a "tartan" pattern (Deer et al., 2001).

94 On planetary bodies with active hydrospheres, plagioclase routinely alters to sericite,
95 zeolites, or clay group minerals. These products can locally be replaced by secondary calcite, due
96 to hydrothermal alteration (e.g. Leichmann et al., 2003; Pittarello et al., 2013). A typical product
97 of hydrothermal alteration of plagioclase is "saussurite", an old term used to indicate a mixture
98 of zoisite, epidote, sericite, and other components.

99 **3. Shock effects in Feldspars**

100 Shocked feldspar from impact structures on Earth, meteorites, and shock recovery
101 experiments displays a range of shock-induced microstructures including undulatory extinction,
102 mosaicism, planar microstructures, diaplectic glass, and melted (i.e., flow-textured and/or
103 vesiculated) feldspar glass (Table 1, e.g., Stöffler, 1967; Stöffler et al., 2018). These effects
104 result from progressive breakdown of the crystal lattice, and are linked with a specific range of
105 shock pressures, temperatures, and anorthite content, as determined by high-pressure
106 experiments (Fig. 1).

107 The state of matter under shock compression can be described with Hugoniot curves
108 (e.g., Melosh, 1989; Langenhorst, 2002; Langenhorst and Deutsch, 2012; Fritz et al., 2017),
109 which are depicted, for a specific mineral, in a density (or specific volume) vs. pressure plot.
110 Figure 2 shows experimental data on anorthosite (rock composed mainly of Ca-rich plagioclase)
111 derived from particle and shock wave velocity measurements at given shock pressures, with the

112 specific volume calculated via the Rankine-Hugoniot relations. The Hugoniot is not a
113 representation of a gradual increase in pressure and density, but describes the possible states
114 achieved by a material that discontinuously jumps across the shock front from the uncompressed
115 to the compressed state. The Hugoniot Elastic Limit (HEL) of a material indicates the transition
116 from a purely elastic state to an elastic-plastic state. For feldspars, the HEL has been measured at
117 3.5-4.5 GPa. Pressures in excess of this range cause permanent changes in feldspar crystal
118 structures (Ahrens et al., 1973; Grady and Murri, 1976; Huffman and Reimold, 1996). In
119 addition, the Hugoniot of feldspar is subdivided into Regions I to III, which mark pressures
120 above which the compressibility of the material during shock changes substantially. These
121 regions correspond to specific shock metamorphic effects – the physical manifestation of the
122 pressure-volume work during shock. Hence, shock effects from fracturing, kinking, and
123 mosaicism, to formation of planar microstructures, diaplectic glass, and melting can be linked to
124 a range of shock pressures. In order to appropriately extrapolate a shock pressure estimate to the
125 whole rock, one must consider numerous grains throughout the sample, ideally in a variety of
126 minerals. In the following sections, we provide an overview of the effects produced in feldspars
127 during hypervelocity impact, both diagnostic (unique to impact) and non-diagnostic (formed both
128 during impact and endogenically).

129 **3.1. Fracturing**

130 Fracturing is pervasive in rocks affected by impact (Fig. 3). The density of impact-
131 induced fractures in quartz and feldspars has been found to increase with increasing pressure up
132 to ~20 GPa (Lambert 1979). However fracture density decreases as pressure increases above
133 ~20 GPa. Lambert (1979) concluded that the correlation between fracture density and pressure

134 was too weak to be used for quantitative pressure calibration. Fractures are rare or absent in
135 partly or fully isotropic feldspars (see Section 3.7).

136 **3.2. Undulatory extinction**

137 Undulatory (or undulose) extinction is the optical effect of a wave of extinction sweeping
138 through a single crystal as the microscope stage is rotated with the sample between crossed
139 polarizers; i.e., the entire crystal does not go in and out of extinction simultaneously (Fig. 4).
140 This optical effect is the result of bending of the crystal structure without fracturing. Undulatory
141 extinction is a common effect in minerals that have experienced non-uniform strain, and is not
142 diagnostic of shock metamorphism, however it can be informative. For example, the amount of
143 stress that a crystal has experienced can be qualitatively estimated based on how undulatory the
144 extinction is; i.e., approximately how much variation in extinction angle is displayed within a
145 single crystal? A higher amount of variation records higher strain. There is potential to quantify
146 this effect using a Universal stage (U-stage) or electron backscatter diffraction (EBSD)
147 measurements of the misorientation of the optic axes, however, we are not aware of any
148 systematic studies attempting to quantify undulatory extinction in feldspars. Such efforts would
149 be unlikely to develop a robust correlation given the ubiquitous nature of undulatory extinction
150 as a result of tectonic deformation ($P \ll 1$ GPa).

151 **3.3. Optical Mosaicism**

152 Optical mosaicism is an extinction pattern observed in crystals composed of numerous
153 subdomains, with differently oriented optic axes (e.g., Dacheille et al., 1968; Stöffler, 1972;
154 Stöffler and Langenhorst, 1994; French and Koeberl, 2010). This extinction pattern is different
155 from undulatory extinction in which extinction passes smoothly through the crystal during

156 rotation of the microscope stage. In mosaicism, subdomains are only a few micrometers in size,
157 hence several rotated subdomains overlap in a typical $\sim 30 \mu\text{m}$ thick thin section. Mosaicism is
158 best observed in transmitted light between crossed polarizers, where it appears as a patchwork of
159 extinction domains; with EBSD, as a patchwork of crystallographic orientations; or with single-
160 crystal or in situ X-ray diffraction (XRD), producing arcs of diffraction spots instead of single
161 spots. A similar pattern of extinction is also caused by endogenic processes (e.g., Spry, 1969;
162 Vernon, 1975), so optical mosaicism cannot be used as a unique diagnostic indicator of shock
163 metamorphism.

164 **3.4. Planar microstructures**

165 Shock related planar microstructures in feldspars include planar fractures (PFs), planar
166 deformation features (PDFs), deformation bands, and kink bands (e.g., Stöffler, 1967, 1972;
167 French and Short, 1968; French, 1998). PFs and PDFs in feldspar are not as well characterized as
168 those in quartz, so we start with a brief discussion of these features in quartz to frame the
169 discussion. In quartz, PFs are open cracks, typically $>3 \mu\text{m}$ wide, and spaced $\sim 15\text{-}20 \mu\text{m}$ apart;
170 PDFs are lamellae composed of amorphous (or recrystallized) SiO_2 , $< 2 \mu\text{m}$ wide, and spaced
171 $\sim 2\text{-}10 \mu\text{m}$ apart (e.g., Engelhardt and Bertsch, 1969; Stöffler and Langenhorst, 1994;
172 Langenhorst, 2002). Both PFs and PDFs are planar and form parallel to rational crystallographic
173 orientations (French and Short, 1968; French et al., 2004).

174 Early reports of what are now referred to as PDFs in feldspar were described as various
175 types of lamellae in plagioclase by both Stöffler (1967) and Dworak (1969). Stöffler (1967)
176 described $1\text{-}8 \mu\text{m}$ thick isotropic (or low birefringence) lamellae spaced a few micrometers apart,
177 similar to planar features described in quartz. Dworak (1969) described several types of lamellae,
178 equivalent to deformation bands, PDFs, and PFs. The lamellae described as being most similar to

179 quartz PDFs are 2-5 μm thick isotropic (or low birefringence) lamellae spaced about 10 μm
180 apart, sometimes decorated with small inclusions. Dworak (1969) described feldspar PFs as up to
181 10 μm wide open fractures spaced a few micrometers apart and forming preferentially at grain
182 margins.

183 In feldspar, PDFs occurring along with more widely spaced features (i.e., deformation
184 bands or twins) form a ladder texture (Fig. 5) (e.g., Stöffler, 1967, 1972; Engelhardt and Stöffler,
185 1968; French and Short, 1968). Pittarello et al. (2013) reported microstructures in plagioclase
186 that resemble feather features (FFs) previously described for quartz (Poelchau and Kenkmann,
187 2001), however no systematic study has yet been conducted on apparent FFs in feldspars. Sets of
188 thin alternating isotropic lamellae ($<10 \mu\text{m}$ wide) oriented parallel to the most common
189 plagioclase twins have been reported as shock-induced microtwins (e.g., Stöffler, 1966; Dworak,
190 1969; Stöffler, 1972), though it has yet to be confirmed whether the formation of the microtwins
191 is due to shock metamorphism, nor whether the currently isotropic lamellae originated as twin
192 lamellae.

193 A great many planar microstructures exist in feldspars to begin with, so using an optical
194 microscope alone to fully characterize putative PDFs is challenging, and it is better combined
195 with high-resolution techniques such as electron microscopy. Pickersgill et al. (2017)
196 investigated planar microstructures in Or-rich alkali feldspar from granitoid rocks from the
197 Chicxulub impact structure using EBSD combined with transmission electron microscopy
198 (TEM). EBSD highlighted semi-planar microstructures visible in transmitted light parallel to the
199 trace of $\{001\}$, $\{010\}$, and $\{100\}$. Follow up TEM work revealed that the “planes” are in fact
200 semi-coherent subgrain boundaries that are neither straight nor parallel (spacing varies from 0.2-
201 0.6 μm between adjacent boundaries), and have no indication of amorphous material, making

202 these subplanar microstructures distinct from the strict definition of PDFs as defined in quartz.
203 The microstructures parallel to $\{110\}$ form lamellar subgrains that, despite resembling strain-
204 induced twins, are too closely spaced and narrow to be twins. Likewise, they are not exsolution
205 lamellae as they show no chemical variation. Pittarello et al. (2020a) conducted EBSD on
206 shocked plagioclase grains in a meta-granite from the Manicouagan impact structure, Canada.
207 Results suggest that amorphization in plagioclase begins either within pre-shock twins or along
208 shock-induced lamellae that are similar to both PDFs and micro-twins, likely representing
209 structural failure of specific crystallographic planes during shock. Attempted indexing of these
210 features using EBSD and U-stage resulted in orientations of $\{001\}$, $\{010\}$, and $\{1\bar{2}0\}$ or $\{1\bar{3}0\}$
211 (ambiguous).

212 PDFs in feldspars and quartz are reported to start to form at about the same pressure (>8
213 GPa, Stöffler, 1971; Ostertag, 1983), and are no longer present in samples subjected to pressures
214 high enough to cause amorphization of the mineral (24-35 GPa). PDFs in feldspars have been
215 reported for various impact structures including Ries (Stöffler, 1966; Engelhardt and Stöffler,
216 1968); Sudbury (French, 1968); Manicouagan (Dworak, 1969; Dressler, 1990; White, 1993);
217 Gardnos (French et al., 1997); Tenoumer (French et al., 1970; Jaret et al., 2014); Dhala (Pati et
218 al., 2010); and El'gygytgyn (Pittarello et al., 2013). Overwhelmingly, these are noted in passing
219 as an accessory to PDFs in quartz, with rare attention or detail given to the PDFs in feldspar. As
220 a result, the compositional range of feldspars bearing PDFs are rarely noted, but from the limited
221 examples, PDFs appear to be confined to alkali feldspar or low-Ca ($<An_{50}$) plagioclase
222 (Langenhorst et al., 1995; Trepmann et al., 2003; Gibson and Reimold, 2005; Nagy et al., 2008;
223 Pittarello et al., 2013; Jaret et al., 2018), and are generally absent in higher Ca-plagioclase even
224 for samples in shock ranges expected to form PDFs. For example, at the Mistastin Lake impact

225 structure, PDFs are abundant in quartz, and completely absent in plagioclase (An_{31-55})
226 (Pickersgill et al., 2015b). Similarly, at the Lonar Crater, India, which is a basaltic target with
227 highly calcic plagioclase, no PDFs have been identified despite abundant examples of diaplectic
228 glass (Kieffer et al., 1976; Jaret et al., 2015). However, we are aware of two exceptions to the
229 high-Ca plagioclase resistance to PDFs, both in lunar anorthite, from the meteorite Oued Awlitis
230 001 (Wittmann et al., 2019), and Apollo sample 10046 (Dence et al., 1970). This indicates a
231 strong compositional (and therefore crystallographic) effect on whether or not PDFs form in
232 feldspars. The lack of observed planar features in experimentally shocked plagioclase (An_{63}) led
233 Gibbons and Ahrens (1977) to suggest that the presence of PDFs in plagioclase is not as
234 diagnostic of shock as maskelynite, and that PDFs would represent a purely local effect.

235 **3.5. Shock darkening**

236 Shock darkening is a reduction in the transparency of a mineral in transmitted light (e.g.,
237 Reimold and Miller, 1989; French et al., 1997; Jaret et al., 2018). This optical phenomenon is a
238 common shock feature in mafic silicates, but has also been observed in feldspars (Fig. 6). Shock
239 darkening of silicates is attributed to submicroscopic Fe-bearing droplets, that are mobilized and
240 concentrated at shock conditions prior to melting (e.g., Rubin, 1992; Moreau et al., 2017). Shock
241 darkening has been described in terrestrial shocked feldspars at the Gardnos (French et al.,
242 1997), Manson (Koeberl et al., 1996), Roter Kamm (Reimold and Miller, 1989), and Sudbury
243 (French, 1968) impact structures. Shock darkening has also been observed in experimentally
244 shocked andesine, albite, and bytownite (Jaret et al., 2018). In plane-polarized light samples
245 shocked between 17 and 50 GPa all show a pronounced darkening that is inherent to the samples
246 and not a surface coating, nor does it correspond to elemental composition at the spatial
247 resolution of the electron microprobe, which supports the attribution of darkening to sub-

248 micrometer Fe-droplets. To our knowledge no atomic level study of shock darkened feldspars
249 has been conducted. This effect is superficially similar to clouded feldspar, but can be
250 differentiated in that in clouded feldspar the features are usually resolvable with an optical
251 microscope (e.g., Poldervaart and Gilkey, 1954; Whitney, 1972; Smith and Brown, 1988; Putnis
252 et al., 2007).

253 **3.6. Alternate twin deformation**

254 An effect observed in some feldspars is the deformation of only alternate twins in, for
255 example, a polysynthetically twinned plagioclase. This alternate twin deformation includes the
256 formation of possible PDFs in every other twin (forming a ladder texture or alternate twin
257 lamellae, Fig. 5a,b), and the formation of diaplectic glass (alternate twin isotropisation).
258 Alternate twin effects are interpreted to occur in response to "favorable" crystallographic
259 orientation of one twin set with respect to the local shock wave (Stöffler, 1966, 1967; Taylor and
260 Dence, 1969; Robertson, 1973; Dressler, 1990; Jaret et al., 2014; Pickersgill et al., 2015b;
261 Pittarello et al., 2020a). Jaret et al. (2014) interpreted alternate twin deformation to indicate
262 relatively low pressure conditions where subtle orientation differences are enough to impede
263 shock deformation, whereas at higher shock conditions there is so much energy that slight
264 misorientations do not affect the deformation process. Likewise, Pickersgill et al., (2015a) used
265 in situ micro X-ray diffraction (μ XRD) to find that adjacent twin sets generally deform to the
266 same degree, which they interpreted to mean that the difference in lattice orientation between
267 twins relative to the shock wave that causes alternate twin deformation occurs over a very
268 narrow range of orientations.

269 **3.7. Diaplectic glass**

270 Diaplectic glasses, sometimes (in older papers) referred to as *thetomorphic glasses*, are
271 phases which are optically isotropic (i.e., extinct on rotation of the stage between crossed
272 polarisers); amorphous to Raman, X-ray diffraction, or other techniques that measure
273 crystallinity; but still retain the exact chemical composition, and morphology of the precursor
274 mineral thereby preserving the original rock fabric (Fig. 7) (e.g., Engelhardt et al., 1967; Binns,
275 1967; Engelhardt and Stöffler, 1968; Hörz and Quaide, 1973). Diaplectic glasses do not exhibit
276 flow textures or vesiculation, even when observed using high resolution imaging techniques such
277 as scanning electron microscopy (SEM). Preservation of the internal fabric is indicated by
278 preservation of, for example, magmatic pyroxene lamellae in maskelynite from the Martian
279 meteorite Shergotty (Fig. 7) and Fe-oxide “clouding” in diaplectic plagioclase glass from the
280 Mistastin Lake impact structure (Pickersgill et al., 2015b). Even chemical zonation is preserved
281 in diaplectic plagioclase glass from both terrestrial impact structures (Jaret et al., 2015) and
282 meteorites (Treiman and Treado, 1998). The transition from crystalline (i.e., birefringent)
283 feldspar to diaplectic feldspar glass is documented by grains in which some areas are amorphous,
284 and some areas remain birefringent. This transition can manifest as alternate twin isotropisation
285 (Stöffler, 1966, 1971; French, 1998), but is equally found unrelated to twins or other planar
286 features (Fig. 7C,D, e.g., Pickersgill et al., 2015b; Pittarello et al., 2020a, 2020b). Both partially
287 and fully isotropic diaplectic glasses are characterized by a lack of fracturing, even when
288 surrounded by highly fractured minerals (Fig. 7 a,b).

289 **3.7.1. Identification**

290 In transmitted light, initial suspicion of diaplectic glass comes from noting a grain that is
291 constantly extinct between crossed polarizers. In reflected light or during SEM of a polished
292 surface (thin or thick section), initial inference of diaplectic glass comes from observing a

293 notable pattern of fractures: fractures are rare or absent in partly and fully isotropic feldspar and
294 fractures in adjacent minerals such as olivine and pyroxene terminate at the interface with the
295 diaplectic glass (see Fig. 7 a,b).

296 In order to definitively identify diaplectic glasses, it is crucial to ensure that the phase
297 being examined is 1) amorphous, and 2) the same composition as feldspar. Permanent extinction
298 under an optical microscope can indicate a) an amorphous phase; b) orientation of an optic axis
299 of a crystal parallel to the line of sight of the microscope (perpendicular to the thin section
300 surface); and c) naturally isotropic minerals (e.g., garnet). It is therefore necessary to confirm an
301 amorphous state using additional techniques such as obtaining an interference figure using the
302 Bertrand lens of the microscope (no interference figure should be present with an amorphous
303 phase), or using a U-stage to rotate the sample out of extinction. If the phase is indeed
304 amorphous, in situ μ XRD will reveal a diffuse band across the detector rather than distinct
305 diffraction spots. Likewise, Raman spectra of amorphous feldspar have two broad peaks of
306 relatively low intensities at $\sim 400\text{-}600 \text{ cm}^{-1}$ and $\sim 1000\text{-}1100 \text{ cm}^{-1}$ (Fig 14). Composition
307 should be confirmed using an analytical SEM or electron probe microanalysis (EPMA).

308 **3.7.2. Structure**

309 The amorphous nature of diaplectic glasses suggests that they have lost the ordered
310 internal atomic arrangement of the precursor crystal. However, diaplectic feldspar glasses
311 display several properties that indicate a transitional state between a crystal and a glass produced
312 by quenching of a molten mineral (i.e., a monomineralic melt glass):

313 1) Diaplectic plagioclase glass has a density and refractive index (RI) intermediate between
314 a plagioclase crystal and a melt glass of identical chemical composition (Engelhardt et al.

315 1967; Ostertag 1983; Fritz et al. 2019a). Both density and RI decrease with increasing
316 shock level (Arndt et al., 1982, see Fig. 8);

317 2) Diaplectic glasses maintain substantial intermediate-range structural order, unlike melt
318 glasses (e.g., Milton and de Carli, 1963; Bunch et al., 1967; Arndt et al., 1982; Ostertag,
319 1983; Jaret et al. 2015).

320 3) Diaplectic plagioclase glass retains a structural memory of its former crystalline state
321 (Milton and de Carli, 1963; Bunch et al., 1967; Hörz and Quaide, 1973; Arndt et al.,
322 1982; Diemann and Arndt, 1984; Jaret et al., 2015).

323 When heated at ambient pressure, diaplectic glasses can recrystallize back into a single
324 crystal, which is distinctly different from melt glass, which always reverts to a polycrystalline
325 aggregate (Bunch et al., 1967; Arndt et al., 1982). Sometimes these recrystallized feldspars even
326 regain undulatory extinction. Further investigations into the memory effects of diaplectic glass
327 have been investigated using high-pressure experiments and are discussed in Section 4.

328 Structural differences between melt glass and diaplectic glass have been investigated by a
329 number of techniques. Synchrotron X-ray total scattering experiments of crystalline plagioclase,
330 natural diaplectic plagioclase glass, and fused plagioclase glass (all of labradorite composition)
331 show that both glasses are characterized by lack of periodicity beyond distances of 10 Å, unlike
332 crystalline plagioclase (Jaret et al., 2015). However, in the intermediate range (4–10 Å) the
333 diaplectic glasses show a higher degree of atomic order than melt glasses (Jaret et al., 2015).
334 Orientation sensitive measurements on diaplectic glass have suggested retention of orientation
335 and position of some atoms despite loss of long-range order. Micro-FTIR spectroscopy of
336 individual naturally and experimentally shocked plagioclase grains is also sensitive to

337 orientation, where peak positions and intensities of vibrational modes in plagioclase can shift by
338 ~40 wavenumbers, making it difficult to correlate changes in peak intensities or positions with
339 shock level (Fig. 9). However, the degree to which orientation affects the spectra decreases with
340 increasing shock level (Jaret et al., 2015). Orientation effects have been found to be a useful tool
341 for distinguishing between diaplectic and melt-produced glasses. For example, in infrared
342 spectroscopy diaplectic labradorite glass exhibits one Si-O stretching band, but this peak position
343 changes by 40 wavenumbers upon rotation of the grain (Fig. 9). Melt glass on the other hand
344 does not show effects of orientation on spectral position (Jaret et al., 2015). Nuclear magnetic
345 resonance (NMR) spectroscopy shows the overall spectral shape and pattern of diaplectic
346 labradorite glass is similar to that of crystalline plagioclase, suggesting a lack of large-scale
347 differences in structure or silica polymerization. Fused labradorite glass on the other hand has a
348 larger difference in chemical shift in NMR spectra compared to crystalline labradorite, further
349 supporting that diaplectic plagioclase glass and fused plagioclase glass are structurally different
350 (Jaret et al., 2015).

351 **3.7.3. Alteration**

352 In a hydrous environment, feldspars and diaplectic glass are susceptible to alteration. As
353 a result, recrystallized feldspars are typical in terrestrial impactites. The resultant textures are
354 described as spherulitic or plumose microcrystals, interpreted as evidence of altered diaplectic
355 glasses during post-shock heating (French, 1998). This effect occurs in the Manson impact
356 structure, where shocked plagioclase appears to have highly altered alternating twins (Short and
357 Gold, 1996). Likewise, at the Mistastin Lake impact structure, alternating twins in some samples
358 have altered to zeolites (Pickersgill et al., 2015b). In both cases, these textures are interpreted as
359 preferential alteration of diaplectic glass in grains that underwent alternate twin isotropization.

360 **3.7.4. Formation pressure**

361 Fully isotropic diaplectic feldspar glass forms by the passage of a shock wave with
362 pressures of more than 24-30 GPa and less than ~47 GPa (Table 1, Ostertag, 1983; Jaret et al.,
363 2018; Fritz et al., 2019a). Both meteorites and terrestrial impactites have documented a
364 relationship between the formation of diaplectic plagioclase glass from plagioclase in contact
365 with shock melt veins (Walton et al., 2016; Sharp et al., 2019). In these cases, plagioclase further
366 from melt veins, or in contact with thin (1-4 μm) veins remains crystalline, but plagioclase in
367 contact with melt veins has become isotropic. The formation of diaplectic glass is therefore
368 attributed to the elevated temperature gradient caused by contact with shock melt, combined with
369 the large deviatoric stresses experienced by minerals along vein margins. These natural
370 observations agree with experimental results that demonstrate that both elevated temperature and
371 shear stress reduce the pressure at which plagioclase transforms to diaplectic glass (e.g., Kubo et
372 al., 2010; Daniel et al., 1997). There is a strong compositional effect on the pressures associated
373 with amorphization, where more calcic plagioclase becomes amorphous at lower pressure
374 conditions (Ostertag, 1983; Angel, 1994; Fritz et al., 2019a; Jaret et al., 2018). Most likely this
375 trend reflects the Si-Al polymerization rather than Ca content directly (Angel, 1994; Jaret et al.,
376 2018). There has been significantly less work on quantifying the differences in shock response
377 between the alkali-series feldspars. Ostertag (1983) reported amorphization of experimentally
378 shocked sanidine and orthoclase at pressures of ~30 and ~32 GPa, whereas microcline retained
379 some birefringence up to pressures of 45 GPa.

380 The transitional regime between diaplectic and melt glasses results in shock-temperatures
381 in plagioclase that are not high enough to cause melting of the whole grain, but are just high

382 enough to cause eutectic melting along plagioclase/pyroxene interfaces (Fig. 7) and assimilation
383 of minerals with a lower melting point.

384 **3.8. Melting**

385 At higher shock pressure monomineralic melt glass forms with the same composition as
386 the original feldspar (e.g., Stöffler, 1972, 1967). An example of fully melted plagioclase glass is
387 shown in Fig. 10, in the strongly shocked Martian meteorite ALH 77005. Petrographically
388 melted feldspar glass is significantly different from diaplectic feldspar glass in that melt glass
389 displays flow textures and/or vesicles.

390 Upon cooling, if the pressure pulse is longer than the undercooling, high pressure phases
391 can crystallize from the melt (e.g., Sharp and DeCarli, 2006; Walton et al., 2014). Small
392 occurrences of plagioclase-composition melt glass can be identified by careful high-resolution
393 backscatter electron imaging (El Goresy et al., 2013). Care must be taken to ensure that apparent
394 feldspar-composition glasses are not whole-rock melts using EPMA or analytical SEM. Feldspar
395 melting is associated with local peak shock pressure of ~ 45-50 GPa, however melting can occur
396 in porous targets shocked to lower pressures, where the shock kinetic energy is mostly converted
397 into heat (e.g., Kieffer, 1971; Davison et al., 2010).

398 **3.9. High pressure polymorphs**

399 The temperatures during shock are such that weakly to strongly shocked rocks can serve
400 as a heat sink allowing localized zones of shock melt to quickly cool by assimilation and thermal
401 conduction during elevated shock pressure conditions. Therefore, localized zones of shock melt
402 (10s to 100s of micrometers in size) provide the pressure-temperature-time path (P-T-t) allowing
403 for the formation and preservation of a variety of high-pressure phases (Tomioka and Miyahara,

404 2017; Fritz et al., 2017). These local zones of shock melt serve as natural crucibles mimicking
405 the conditions in the Earth's lower crust or mantle. High-pressure minerals are not common in
406 Earth's upper crust, but they are produced endogenically and therefore can only be related to
407 shock when observations are combined with relevant geologic context (French and Koeberl,
408 2010). A list of high-pressure minerals related to shocked feldspar is given in Table 2, together
409 with their chemical formulae, experimentally derived pressure and temperature (P-T) stability
410 fields and references. High-pressure phases have been identified in most types of meteorites, in
411 various rocks from terrestrial impact structures, and synthesized in high-pressure laboratory
412 experiments. This indicates that formation and preservation of high-pressure phases is a typical
413 result of impact cratering events.

414 **3.10. Sieve-texture and checkerboard feldspar**

415 "Checkerboard" feldspar is a checkered-looking pattern of subgrain domains found
416 within feldspar clasts in impact melt rocks (Fig. 11). Checkerboard texture has the appearance of
417 crystallographically controlled individual rectangular domains separated by melt products (e.g.,
418 Grieve, 1975). Internal features such as twinning and extinction angle are preserved across the
419 grains demonstrating that they were once a single crystal (Fig. 11). This texture was originally
420 termed "checkerboard" by Grieve (1975), who described unshocked partly digested plagioclase
421 inclusions in a melt rock with a sieved, "checker-board" appearance due to thermal effects of the
422 surrounding melt. Bischoff and Stöffler (1984) first proposed a detailed scenario for
423 development of sieve texture in clasts in impact melt rock from the Lappajärvi impact structure
424 (Finland). Their scenario essentially regarded checkerboard texture as the result of rapid,
425 crystallographically-controlled recrystallization from diaplectic feldspar glass, comparable to the

426 suggested formation of ballen quartz (e.g., Ferrière et al., 2010). As such, the texture would
427 represent positive evidence of transient shock pressures exceeding 25-35 GPa.

428 Whitehead et al., (2002) revised the formation of checkerboard texture based on feldspar
429 clasts in melt rocks from the Popigai impact structure. Their scenario more closely resembles the
430 disequilibrium melting of plagioclase observed during the rapid decompression of intermediate
431 composition magmatic systems, resulting in typical (i.e., igneous) “sieve-textured” feldspar
432 (Nelson and Montana, 1992). Whitehead et al. (2002) noted similar textures in clinopyroxenes in
433 the Popigai impact melt rocks, which are also commonly associated with so-called “sieve-
434 textured” feldspars in volcanic rocks.

435 A texture similar to checkerboard feldspar occurs in volcanic rocks, and is also referred
436 to as “sieve texture”. The images that accompany the igneous definition of “sieve texture” (Klein
437 and Philpotts, 2016) imply crystallographic control of the melt channels. This can lead to
438 confusion when the terms “sieve-texture” and “checkerboard texture” are used interchangeably,
439 because there are two genetic interpretations for a similar appearance: i) volcanic chemically
440 driven disequilibrium melting (Nelson and Montana, 1992) and ii) impact-induced
441 crystallographically controlled disequilibrium melting (Whitehead et al., 2002).

442 Whitehead et al., (2002) attempted to distinguish impact melting from igneous melting
443 based on the optical crystallography and X-ray element mapping of subgrains within the feldspar
444 clasts. However, it is not clear from their observations that distinguishing these two processes is
445 possible (Harris et al., 2019). Therefore, it has not been convincingly shown that the sieve
446 texture of igneous petrologists and the sieve or checkerboard texture of impact petrologists are
447 the result of different processes. Consequently checkerboard or sieve-texture feldspar, while
448 common in both volcanic and impact melt rocks, should not be used as a unique diagnostic

449 criterion of either process unless significant progress is made in differentiating igneous sieve
450 texture from impact-induced sieve or checkerboard texture.

451 **4. Experimental Shock Calibrations**

452 High-pressure experiments enable formation conditions (i.e., pressure, temperature,
453 timing) to be assigned to rock and mineral samples based on the microscopic textures produced.
454 Such experiments are conducted using four main techniques: 1) shock experiments; 2) shock
455 recovery experiments; 3) rapid compression experiments; 4) static compression experiments.
456 These different techniques (1-4 in previous sentence, summarized in Table 3) create different
457 physical conditions. As a result, the same high-pressure effects (e.g., amorphization) may
458 develop at different pressure and temperature conditions depending on the experimental process
459 and the characteristics of the test material (i.e., initial grain size, composition, porosity, presence
460 of structural defects, etc.). Therefore, the reported experimental data need careful evaluation
461 before comparing results between different experiments and their application to natural shock
462 metamorphism. The experimental setups mentioned previously could influence the pressures at
463 which specific microstructural deformation is observed in several ways:

464 1) By studying the samples in a compressed or decompressed state, i.e., in situ during static
465 pressure experiments or after decompression to ambient pressure from static pressure
466 experiments (Daniel et al., 1997; Johnson et al., 2002a, 2003; Kubo et al., 2010; Jaret et
467 al., 2018).

468 2) By applying hydrostatic or non-hydrostatic conditions to the sample (Daniel et al., 1997)

469 3) By changing the rate at which the sample is compressed (e.g., Sims et al., 2020)

- 470 4) By changing the duration of the high-pressure conditions (Huffman et al., 1993; Kubo et
471 al., 2010; Ogilvie et al., 2011; Fritz et al., 2019a); and,
- 472 5) By selecting the pre-experiment or the syn-experiment temperature of the experimental
473 setup (Gratz et al., 1992; Huffman and Reimold, 1996; Kubo et al., 2010; Tomioka et al.,
474 2010; Fritz et al., 2011, 2019a).

475 In addition, experimental measurements of the shock wave and particle velocity
476 parameters allow for calculation of the pressure, density, and energy of the compressed material
477 using the Rankine–Hugoniot relations (Fig. 2). These different experimental approaches have
478 provided a variety of complementary information on material properties during shock
479 compression and after decompression to ambient pressures (see Section 3).

480 The wealth of experimental data provides space for a variety of opinions among shock
481 researchers. For simplicity, the discussion can be separated into two end-member interpretations:

- 482 1) *The shock metamorphic record in naturally shocked rocks is in conflict with experimental*
483 *results obtained by the different types of experiments* (Stöffler et al., 1991, 2018; Chen et
484 al., 1996; Sharp and DeCarli, 2006; Gillet et al., 2007; El Goresy et al., 2013). Various
485 researchers agree that there is a substantial difference between results from static pressure
486 and dynamic (shock) pressure experiments, with different authors favoring results from
487 either one or the other (Stöffler and Langenhorst, 1994; Sharp and DeCarli, 2006; Gillet
488 et al., 2007; Kubo et al., 2010; Jaret et al., 2018; Sims et al., 2020). The argument is that
489 static experiments are too slow and do not generate a shock wave, because the pressure
490 increases stepwise and is sustained over relatively long time scales of seconds to minutes.
491 The dynamic experiments occur over extremely short timescales (~ 1 ns – 1 μ s) and
492 cannot sustain the shock conditions for the same duration as in a natural impact event. In
493 the case of reverberation shock experiments, the maximum pressure is not achieved in a
494 single shot as is the case in natural shock waves (Langenhorst, 2002). Thus, neither static

495 nor dynamic experiments reproduce the duration and loading path as inferred for the
496 natural case (Stöffler and Langenhorst, 1994; Sharp and DeCarli, 2006; Gillet et al.,
497 2007), which is typically in the temporal range of milliseconds to rarely a few seconds
498 (Langenhorst, 2002; Sims et al., 2019; Bowling et al., 2020). These interpretations are
499 due to a well-discussed discrepancy between the pressures associated with
500 transformations observed in static experiments compared to those in dynamic
501 experiments. For example, diamond anvil cell (DAC) experiments and shock experiments
502 from the same starting material show a ~7-9 GPa difference in amorphization pressures
503 (Jaret et al., 2018; Sims et al., 2020), consistent with interpretations compiled from
504 literature of either isolated shock or static experiments (Daniel et al., 1997; Johnson et al.,
505 2002a, 2003). Sims et al. (2020) recognized memory effects in the DAC experiments
506 suggesting two amorphization points – one where the sample appears amorphous while
507 still under compression, and a higher point where the sample remains amorphous after
508 decompression. The higher pressure is closer to those of the shock experiments, which
509 may in part help reconcile the discrepancies. Recent “rapid compression” experiments
510 using membrane DAC setups that allow for control on strain-rate have shown that strain
511 rate has a significant effect on the pressures at which plagioclase deforms (Sims et al.,
512 2019).

513 2) *The shock metamorphic record in naturally shocked rocks agrees with the results*
514 *obtained by both shock and static pressure experiments* (Fritz et al., 2017). Shock
515 recovery experiments show that with increasing shock pressure different types of shock
516 deformation effects develop in olivine and plagioclase, respectively. These two different
517 thermo-barometers provide consistent results if studied in naturally shocked rocks such as
518 shocked igneous Martian meteorites (Fritz et al., 2005a). These consistent results mean,
519 that either both olivine and plagioclase are affected by potential differences between
520 shock recovery experiments and the natural case to the same degree, or that shock
521 recovery experiments accurately reproduce the conditions in a natural impact event.
522 Shock reverberation experiments and hydrostatic pressure experiments produce an
523 amorphous material from plagioclase and quartz as long as both types of experiments
524 achieved the same pressure and temperature conditions and the samples are studied in a
525 decompressed state (Fritz et al., 2019a). The shock-loading path achieved in shock

526 reverberation experiments is a reasonable analog for the diffuse shock front traveling
527 through lithological units composed of different types of minerals, porosity and fractures
528 during a natural impact event (Fritz et al., 2011). Shock reverberation experiments allow
529 well-defined shock pressures to be applied to a relatively large sample volume (Müller
530 and Hornemann, 1969; Fritz et al., 2011). Most problems regarding the shock thermo-
531 barometry of meteorites are not related to experimental limitations but mainly result from
532 inconsistent interpretation of the observational evidence (Fritz et al., 2017, 2019a). This
533 includes: a) Misunderstandings regarding the duration of isobaric shock pressure
534 conditions in naturally shocked rocks, which mostly implies that high-pressure phases
535 form in (and near) local shock melt zones during declining shock pressures, b) A
536 mineralogically inconsistent definition of the S6 shock level leading to an incorrect shock
537 pressure assignment of all chondrites containing high-pressure phases (Fritz et al., 2017).

538 Reconciling the different interpretations of the shock thermo-barometry of rocks remains an
539 ongoing area of active research and discussion. For example, a recent study of the shock stage
540 distribution of 2280 ordinary chondrites confirmed that all S6 chondrites need reclassification
541 (Bischoff et al., 2019). Despite these discussions, there is agreement on the types of shock
542 deformation effects in feldspar and the order in which these shock effects occur with increasing
543 pressure (see Section 3).

544 **5. Quantifying shock features in feldspars**

545 Multiple techniques have been applied to begin to quantify shock effects in feldspars.
546 These techniques are summarized in Table 4, and discussed in the following sections.

547 **5.1. Measuring planar microstructures in feldspars**

548 In quartz, the number and orientation of PDF sets have been correlated with specific
549 pressures through shock loading experiments allowing for refined pressure estimates from
550 measurement of PDF orientations (e.g., Hörz, 1968; Huffman and Reimold, 1996; Ferrière et al.,

551 2009). Such a correlation has not been developed for feldspars due to the complexities involved
552 in PDF formation, identification, and orientation measurement in this mineral group.

553 The universal stage (U-stage) is the most common method of determining PDF
554 orientations in quartz, and this method has been used with feldspars, though it is significantly
555 more challenging (Stöffler, 1967; Dworak, 1969; Pittarello et al., 2020a). Measuring the
556 crystallographic orientation of any planar element in feldspars is complicated compared to
557 measuring the same features in quartz because feldspar is optically biaxial, and the crystal
558 structure (and therefore optical properties) of feldspars varies with composition through the solid
559 solution from Ca-rich to Na-rich to K-rich end members (Doeglas, 1940; Haff, 1940; Fairbairn
560 and Podolsky, 1951). Therefore in addition to the measurements themselves being complex,
561 accurately indexing the measured planes requires use of the specific indexing template for the
562 composition of feldspar under investigation; examples of such indexing templates can be found
563 in Stöffler (1967) and Dworak (1969). These factors result in orientation measurements of planar
564 features in feldspars using solely a U-stage to be both time-consuming and error-prone. For
565 example, Dworak (1969) measured 113 lamellae in 51 crystals of labradorite using the U-stage,
566 but due to the inherent difficulty in using a U-stage with feldspars, the lowered birefringence of
567 some host crystals, and the changes in An content (even within a single zoned crystal) the
568 assignment of the orientation of these lamellae had a large error. The most common PDF
569 orientations measured in feldspars are summarized in Table 5. Both Stöffler (1967) and
570 Robertson (1966) came to the conclusion that planar deformation features in feldspar form over
571 such a narrow range of pressures that it would not make sense to correlate orientations with peak
572 pressures. By implication, the mere presence of planar deformation features in feldspar,
573 regardless of their orientation, would be a sensitive shock barometer.

574 An additional issue in investigating PDFs in feldspar with the U-stage is optimizing the
575 choice of glass hemispheres. The glass hemispheres each have a fixed refractive index, and for
576 accurate measurements the refractive index must be the same as the host crystal. However, the
577 refractive index of plagioclase ranges from 1.53 (albite) to 1.59 (anorthite) and for K-rich
578 feldspar, it ranges from 1.518 (orthoclase) and 1.539 (microcline). The hemispheres used for
579 measurements in quartz have refractive index 1.554, which is between the end member values
580 for plagioclase, but is too high for K-feldspar. Hemispheres with closer refractive indices exist,
581 but are not readily available to order and are difficult to match to the grain of interest. However,
582 the error induced by the difference in refractive index between the investigated grain and that of
583 the glass hemisphere pair can be calculated, and used to correct the data (e.g., Reinhard, 1931).

584 To circumvent some of the difficulties encountered by using the U-stage alone, EBSD
585 has recently been combined with U-stage or focused ion beam (FIB) milling and BSE imaging to
586 measure orientation of PDFs. EBSD provides crystal orientation of the host mineral
587 circumventing some of the challenges of defining a unique orientation to a biaxial mineral using
588 the U-stage alone. EBSD can also be used to determine a family of planes to which the trace of a
589 plane most likely belongs. However, to complete the orientation measurement in 3D, a “dip”
590 measurement is acquired using the U-stage (Pittarello and Koeberl, 2017; Pittarello et al., 2020a,
591 2020c) or by making a FIB trench perpendicular to the trace of the plane and measuring the dip
592 off the resultant image (Pickersgill and Lee, 2015). Thus far, using the FIB technique for this
593 purpose has been demonstrated only for quartz, however the same principle should work for
594 feldspars. Currently, transmission electron microscopy (TEM) is necessary for a definitive
595 identification of the crystallographic orientation along which PDFs develop in feldspar.

596 **5.2. Density, refractive index, and birefringence**

597 Density, refractive index (RI), and birefringence decrease as pressure increases until the
598 mineral becomes amorphous (Stöffler, 1974; Stöffler and Langenhorst, 1994). At shock
599 pressures <20 GPa the RI of shocked crystalline (birefringent) feldspar is only slightly lower
600 than the RI of unshocked feldspar (Kleeman, 1971; Stöffler, 1974; Robertson, 1975). At
601 pressures greater than ~24 to 35 GPa (depending on chemical composition), feldspars are
602 completely converted to an amorphous state (diaplectic glass). Once this has occurred, and for
603 each type of feldspar over a very narrow pressure range, the RI of diaplectic feldspar glass drops
604 sharply (Ostertag, 1983). Shock pressures of ~45 GPa produce an amorphous material with a RI
605 matching that of a melt glass of identical chemical composition to feldspar. As an exception,
606 diaplectic anorthite glass produced in 41.5 GPa shock experiments displayed a RI that was
607 higher than that of unshocked crystalline anorthite (Fritz et al., 2019a). The well-tested
608 relationship between the refractive index of diaplectic feldspar glass, the An content, and the
609 shock pressure is illustrated in Figure 9 (e.g., Binns, 1967; Engelhardt et al., 1970; Stöffler et al.,
610 1975; Robertson, 1975; Ostertag, 1982; Fritz et al., 2019a).

611 **5.3. X-Ray Diffraction (XRD)**

612 Two types of X-ray scattering experiments have been applied to shocked feldspars: X-ray
613 diffraction (XRD) experiments by powder or single-crystal XRD, in situ μ XRD (in thin section
614 and hand sample), and high-energy total X-ray scattering experiments using a synchrotron.
615 Contributions by total-scattering experiments are discussed in section 3 as they relate to
616 amorphous phase studies. At pressures too low to induce amorphization or phase-change, XRD
617 patterns show peak broadening. There are two types of peak that can broaden in this fashion,
618 peaks in the chi direction (visible with single-crystal or μ XRD) and peaks in the 2θ direction
619 (visible with single-crystal, μ XRD, and powder diffraction).

620 Shock experiments on andesine (4-10 GPa) and oligoclase (3-34 GPa) by Hörz and
621 Quaide (1973), showed that the degree of crystal lattice damage is correlated with shock
622 pressure. Furthermore, by using single crystal XRD to measure the length of streaks in the chi
623 direction (i.e., strain-related mosaicity) they could start to quantify the amount of strain recorded
624 by the crystals and use that as a proxy to quantify shock level. A correlation between streak
625 length and shock level was also documented in a single-crystal XRD study of plagioclase from
626 the Charlevoix impact structure (Walawender, 1977). Pickersgill et al. (2015a) built upon this
627 work by using in situ μ XRD to quantify strain-related mosaicity by measuring the full width at
628 half maximum (FWHM) of the intensity vs. chi angle pattern, and compared those measurements
629 to optical signs of shock in samples of labradorite and andesine from the Mistastin Lake impact
630 structure, and lunar anorthite collected during the Apollo program. As with the single-crystal
631 studies discussed previously, peak broadening in the chi dimension increased with increased
632 optical signs of shock, up to amorphization of the sample (i.e., formation of diaplectic glass) – at
633 which point the X-rays produced a diffuse band on the detector. Powder XRD analyses of
634 shocked and unshocked sanidine by Kayama et al. (2012) reported almost no change in
635 diffraction peak intensity between samples. However, with increased shock pressure they did
636 find slight peak broadening in 2θ , followed by a sudden change to no detectable peaks (i.e., an
637 amorphous pattern) for amorphous samples.

638 Strain-related mosaicity, as measured using in situ μ XRD or single-crystal XRD, shows
639 more variation in shock level of feldspars prior to amorphization than optical determinations, and
640 likely could be useful in subdividing the lower end of the shock scale (Pickersgill et al., 2015a).
641 However, in order to properly quantify shock via strain-related mosaicity, more work needs to be
642 done such as correlating results with those from other techniques (Raman for example) and via

643 application to a wider variety of feldspar compositions and, ideally, also to experimentally
644 shocked samples. Caution must be applied when using XRD to measure shock, as the features
645 described in the previous paragraphs indicate strain, and therefore, without contextual
646 information, it is impossible to differentiate the cause of the strain (i.e., shock or tectonism).

647 **5.4. Thermal Infrared Absorption Spectroscopy**

648 Thermal infrared absorption spectra of naturally shocked feldspars show lower intensity
649 and less spectral detail than unshocked feldspar. Both detail and intensity continue to decrease as
650 pressure increases as a result of increased glass content and progressive lattice disordering
651 (Lyon, 1963; Bunch et al., 1967; Stöffler and Hornemann, 1972; Stöffler, 1974; Arndt et al.,
652 1982; Ostertag, 1983; Johnson et al., 2002b, 2003; Jaret et al., 2015, 2018). Crystalline feldspars
653 exhibit infrared reflectance peaks between 950 and 1150 cm^{-1} (reststrahlen bands) that
654 correspond to Si-O stretching modes of the SiO_4 tetrahedra. Additional peaks occur between 700
655 and 850 cm^{-1} corresponding to Si-bridging oxygen modes, and between 400 and 600 cm^{-1}
656 corresponding to O-Si-O bending modes (Iiishi et al., 1971; Okuno, 2003). Bulk IR emissivity of
657 experimentally shocked albite shows a progression of change to the IR spectra in the form of
658 decrease in overall band intensity relative to background, as well as loss of the low wavenumber
659 peaks. Spectroscopically, plagioclase that has been transformed to amorphous material (either
660 by melting or as diaplectic glass) exhibits one broad peak reflecting the Si-O stretching vibration.

661 Micro-FTIR spectroscopy of individual plagioclase grains, both naturally and
662 experimentally shocked, has an added complexity in that infrared spectra are orientation-
663 sensitive, where peak positions and intensities of vibrational modes in plagioclase can shift by
664 ~40 wavenumbers, depending on orientation, making it difficult to correlate changes in peak
665 intensities or positions with shock level (Fig. 9). However, the degree to which orientation

666 affects the spectra decreases with increasing shock level (Jaret et al., 2018). Furthermore,
667 orientation effects have been found to be a useful tool for distinguishing between diaplectic and
668 melt-produced glasses. Diaplectic glass retains orientation effects in IR, despite its amorphous
669 nature. Melt glass on the other hand does not show effects of orientation on spectral position.

670 **5.5. Cathodoluminescence (CL)**

671 The main effects of shock metamorphism on cathodoluminescence (CL) spectra of
672 plagioclase are 1) decreasing luminescence with increasing amorphization (e.g., Fig. 12, Kaus
673 and Bischoff, 2000; Götte, 2009; Götze, 2009; Pittarello et al., 2015; Kayama et al., 2018;
674 Pittarello et al., 2020b); 2) the normal band emission at ~ 550-570 nm is often shifted to ~ 630
675 nm (Kayama et al., 2009a); 3) diaplectic plagioclase glass produces a broad band around 350 nm
676 (Götte, 2009; Pittarello et al., 2015; Gucsik et al., 2004). Intriguingly, and in contrast to
677 plagioclase, alkali feldspar shows increased CL intensity with increasing shock pressure
678 (Kayama et al., 2012).

679 CL spectral analysis has been proposed to quantify shock effects and correlate with
680 pressure (e.g., Kayama et al., 2012, 2018). However, the correlation between shock and signal
681 intensity is complex because the intensity of the luminescent centers is regulated by numerous
682 variables, including the abundance of activators, which can be trace elements, point defects, or
683 both, and the abundance of quenchers, which would reduce the effect of the activators. Even in
684 unshocked feldspars, the nature of activators is not always clear, so imaging studies must be
685 accompanied by detailed back-scattered electron imaging in order to properly interpret the
686 microstructures (e.g., Finch and Klein, 1999; Götze et al., 2000; Parsons et al., 2015). As a result
687 of the complex mechanisms involved in luminescence, and therefore the relationship between
688 luminescence and shock intensity, CL is generally viewed as less reliable than other techniques

689 (e.g., XRD or Raman) for quantifying shock intensity, though it remains one of the most useful
690 tools for investigating microstructures.

691 **5.6. Raman spectroscopy**

692 Raman spectroscopy allows identification of high-pressure phases (Section 3.9) and
693 spectra show that increased shock corresponds to a progressive decrease in peak intensity and
694 increase in peak width compared to unshocked feldspar (Fig. 13, e.g., Fritz et al., 2005b; Jaret et
695 al., 2014). Peak positions and width are somewhat dependent on composition, so it is challenging
696 to identify quantitative metrics in Raman spectra that uniquely identify shock.

697 At pressures < 29 GPa, shocked plagioclase has lower intensity and broader FWHM of
698 characteristic Raman bands compared to unshocked plagioclase. At higher pressure (>~40 GPa)
699 Raman spectra exhibit a broad plateau as a result of amorphization (Fig. 13, Jaret et al., 2018). At
700 much higher pressure (>45 GPa) high post-shock temperatures lead to recrystallization, and
701 reappearance of the characteristic Raman spectra (Fritz et al., 2005a,b).

702 **5.7. Photoluminescence spectroscopy**

703 Photoluminescence spectroscopy is a technique, complementary to Raman spectroscopy,
704 which has only been applied to shocked plagioclase by Pittarello et al. (2020b).

705 Photoluminescence allows quantification of the “damage” in the lattice (i.e., amorphization) as
706 percent of disordered structure within the investigated volume (a few μm^3). Similar to peak
707 broadening in Raman spectra, Nd^{3+} luminescence bands get wider and are less defined with
708 increased levels of amorphization.

709 **5.8. Nuclear Magnetic Resonance (NMR) Spectroscopy**

710 NMR spectroscopy uses nuclear spin transitions as a measure of the local chemical
711 bonding and coordination environment (Stebbins and Xue, 2014). Because NMR is sensitive to
712 the entire local chemical environment and not just long-range ordering, NMR is particularly
713 useful for measuring the structure of amorphous materials. In amorphous silicates, NMR spectra
714 of ^{29}Si (“ ^{29}Si NMR”) are sensitive to the number and length of Si-O bonds, types of nearest
715 neighbor atoms, and coordination of Si – all of which can change during shock deformation or if
716 high-pressure phases are produced (Jaret et al., 2015).

717 **6. Ongoing debates**

718 **6.1. Formation mechanism of diaplectic glass**

719 The details of diaplectic glass formation are the subject of much debate. There are, in
720 general, two main hypotheses: 1) rapid quenching of monomineralic melt at high pressure, or 2)
721 solid-state structural collapse/destruction of atomic ordering of the mineral. Some details of these
722 two formation hypotheses are explained here.

723 **6.1.1. Melt formation**

724 Grady (1977) interpreted diaplectic glass to form by quenching of a high-density melt
725 phase upon pressure release. Such a phase would have solidified so rapidly that flow, which
726 would normally erase the morphology and texture of the precursor minerals, did not have time to
727 occur. In this scenario, local temperature spikes due to heterogeneous shock compression, cause
728 local melting despite the temperature in the majority of the rock remaining below the liquidus
729 (Grady et al., 1975; Grady, 1980). A key aspect of Grady’s scenario for formation of diaplectic
730 glass is the dissipation of strain energy by the development of shear bands. Microstructures
731 similar to shear bands have been observed in recovered samples from rapid compression

732 experiments (Sims et al., 2019). Similarly, Arndt et al. (1982) suggested that diaplectic glass
733 forms from shock-induced melting, but that the duration of high temperature is short enough that
734 the liquid transition is incomplete and the disordered transitional state is ‘locked in’. Stöffler and
735 Langenhorst (1994) explained the formation of diaplectic quartz glass by shock melting and
736 quenching during decompression. The concept of diaplectic glass being a high-pressure melt was
737 later adopted to explain the formation of maskelynite by Chen and El Goresy (2000) and El
738 Goresy et al. (2013).

739 **6.1.2. Solid state formation**

740 The amorphous structure, and absence of flow textures, suggests a near instantaneous
741 change from a crystalline structure to glass without melting (e.g., De Carli and Jamieson, 1959;
742 Engelhardt and Stöffler, 1968). Ahrens et al. (1969) used shock experiments on plagioclase to
743 suggest that diaplectic glass forms by solid-state release from a high-pressure phase during
744 decompression. Williams and Jeanloz (1988) also favored a scenario where diaplectic glass is
745 produced by reversion from a higher-pressure phase, which, similar to the scenario of Hemley et
746 al. (1988), calls for high pressure-induced high-coordinated glasses. Hemley et al.’s assumption
747 of a high-coordinated phase during shock compression rests on the experimentally determined
748 Hugoniot curve where the steeper high-pressure branch (region III in Fig. 2) indicates a density
749 during shock compression similar to the density expected of a compressed hollandite-like
750 structured phase (Ahrens et al., 1969). A hollandite-like structure scenario would be supported
751 by the existence of high-pressure phases and/or high-coordinated glasses. However, evidence of
752 high-pressure phases was not observed in static experimentally produced diaplectic glass (Daniel
753 et al., 1997; Sims et al., 2020). Likewise, high-coordinated glasses were not observed in in situ
754 compression experiments or naturally shocked samples (Kubo et al., 2010; Jaret et al., 2015;

755 Sims et al., 2019). Furthermore, recent ^{29}Si NMR studies of natural diaplectic plagioclase, did
756 not detect structural remnants of high-pressure high-coordination glass, even though such
757 remnants would be resolvable with NMR analyses (Jaret et al., 2015). Ashworth and Schneider
758 (1985) proposed a mechanism for diaplectic SiO_2 glass that is similar to metamictization – a
759 process in which alpha particles physically displace atoms within the unit cell. In a diaplectic
760 glass, physical displacement of atoms would be caused by the shock wave, rather than alpha
761 particles.

762 Shock recovery experiments at starting temperatures of 77-293 K showed that in this
763 temperature range the formation of diaplectic plagioclase glass is controlled by shock pressure
764 only, and not temperature (Fritz et al., 2019a). This observation, together with the structural
765 properties of diaplectic plagioclase glass (Jaret et al., 2015) and the recognition that diaplectic
766 glass forms via a gradual collapse of the lattice (Hörz and Quaide, 1973, Fritz et al., 2019a), and
767 the observation of isotropic twins alternating with crystalline twins in plagioclase (e.g., Stöffler,
768 1971; Pickersgill et al., 2015a,b; Pittarello et al., 2020a) further advocates for a solid-state
769 amorphization mechanism.

770 **6.2. Maskelynite and diaplectic glass**

771 The first description of maskelynite was by Tschermak (1872) in the Shergotty martian
772 meteorite, described as a previously unrecognised isotropic phase of near labradorite
773 composition. This new phase was then named after M. H. N. Story-Maskelyne, a famous English
774 mineralogist of the time. Tschermak (1883) then found the same phase in chondrites and, upon
775 realizing that it was pseudomorphous with plagioclase, changed his interpretation to a melted or
776 otherwise vitrified glass of plagioclase composition. A shock origin for maskelynite was
777 suggested by Binns (1967) based on differences in the RI of maskelynite and melt-glass. At the

778 same time, the term *diaplectic glass* was proposed to refer to “amorphous phases produced by
779 shock waves without melting, and [which] are distinguishable from ordinary molten glasses”
780 (Engelhardt et al., 1967; and Engelhardt and Stöffler, 1968). Since that time, *maskelynite* and
781 *diaplectic plagioclase glass* have often been used interchangeably, which attached a genetic
782 connotation to maskelynite. Historically, most isotropic plagioclase-composition material was
783 called maskelynite, but modern observations have enabled us to distinguish between phases that
784 have flowed (e.g., Chen and El Goresy, 2000) and “true” diaplectic glasses that show no
785 evidence of melting (e.g., Jaret et al., 2015, Diemann and Arndt, 1984), even within a single thin
786 section. Because there is confusion over how the term maskelynite is interpreted, it is imperative
787 that clear observations regarding the nature of the isotropic phase be included in any description
788 so that it is clear whether the authors are using “maskelynite” to refer to an apparently solid-state
789 glass or a melt product.

790 **7. Concluding remarks & remaining questions**

791 Feldspars are typically difficult to study due to the optical complexity of their crystal
792 structure and, when in the presence of water, their relatively rapid weathering rate.
793 Consequently, the feldspar group is often neglected in favor of quartz or olivine for use in shock
794 barometry. As a result, the shock scale for feldspar is limited and, essentially, qualitative.
795 However, feldspars will have particular utility when studying quartz-poor rocks such as basalt, a
796 dominant rock type on Mars and Earth’s moon, and an important surficial lithology on most
797 terrestrial bodies. There have been many advances on the effects of shock metamorphism on
798 feldspars over the past five decades but there are several major questions remain open, in
799 particular:

- 800 1. Determining the exact nature and formation mechanism of diaplectic glass. Recent work
801 has made significant strides, but confirmation and further studies would help to bolster
802 these results (Section 6.1);
- 803 2. Understanding the formation of PDFs in feldspars, including exactly why the An content
804 appears to be a controlling factor; the relationship between thin closely spaced shock-
805 induced microtwins and PDFs; whether FFs form in feldspars; and the development of
806 apparently shock-induced planar features that do not appear to be strict PDFs when
807 investigated using high resolution TEM and EBSD (Section 3.4);
- 808 3. Whether or not measuring the crystallographic orientations of shock-induced planar
809 features in feldspars can be improved upon and developed into a useful tool for shock
810 barometry (Section 5.1);
- 811 4. Understanding the exact mechanism by which only alternate twins deform in
812 polysynthetically twinned plagioclase, leaving the other set apparently undamaged, and
813 whether this could be used to determine the local orientation of shock wave propagation
814 (Section 3.6);
- 815 5. How, or if, the intensity of shock effects is affected by the presence of pre-existing
816 microstructures (e.g., cleavage, twinning, exsolution), the crystallographic orientation
817 with respect to the shock wave, and the presence of other mineral phases (Sections 3.4,
818 3.6);
- 819 6. How best to utilize and integrate multiple analytical techniques, which probe slightly
820 different aspects of feldspar composition/structure and deformation (Section 5);

821 7. How to reconcile the pressure conditions associated with crystallographic transformations
822 between various experimental techniques and naturally shocked materials, which is an
823 ongoing challenge in many fields of geoscience: do all types of minerals behave in the
824 same way during shock recovery experiments and natural impacts or do some mineral
825 develop the same shock effects at lower pressures? (Section 4)

826 8. How does shock influence radioisotopic age determination? A topic not touched upon in
827 this paper, because it is worth an entire volume on its own, but the interested reader can
828 examine, for example, Jessberger and Ostertag (1982) and Fernandes et al. (2009).

829 Because shocked feldspars are a significant constituent of most planetary materials, the
830 importance of understanding their shock behavior and ability to inform shock barometry will
831 only become more relevant as we gain increasing numbers of samples from other planetary
832 bodies.

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839 **9. REFERENCES CITED**

840 Agarwal, A., Reznik, B., Kontny, A., Heissler, S., and Schilling, F., 2016, Lingunite-a high-
841 pressure plagioclase polymorph at mineral interfaces in doleritic rock of the Lockne impact
842 structure (Sweden): Scientific Reports, v. 6, p. 25991, doi:10.1038/srep25991.

- 843 Ahrens, T.J., Petersen, C.F., and Rosenberg, J.T., 1969, Shock compression of feldspars: Journal
844 of Geophysical Research, v. 74, p. 2727–2746.
- 845 Ahrens, T.J., O’Keefe, J.D., and Gibbons, R. V, 1973, Shock compression of a recrystallized
846 anorthositic rock from Apollo 15*: Proceedings of the 4th Lunar Science Conference, v. 3,
847 p. 2575–2590.
- 848 Angel, R., 1988, High-pressure structure of anorthite: The American mineralogist, v. 73, p.
849 1114–1119.
- 850 Angel, R.J., 1994, Feldspars at High Pressure BT - Feldspars and their Reactions, in Parsons, I.
851 ed., Dordrecht, Springer Netherlands, p. 271–312, doi:10.1007/978-94-011-1106-5_7.
- 852 Arndt, J., Hummel, W., and Gonzalez-Cabeza, I., 1982, Diaplectic labradorite glass from the
853 Manicouagan impact crater; I, Physical properties, crystallization, structural and genetic
854 implications: Physics and Chemistry of Minerals, v. 8, p. 230–239.
- 855 Ashworth, J.R., and Schneider, H., 1985, Deformation and transformation in experimentally
856 shock-loaded quartz: Physics and Chemistry of Minerals, v. 11, p. 241–249,
857 doi:10.1007/BF00307401.
- 858 Beck, P., Gillet, P., Gautron, L., Daniel, I., and El Goresy, A., 2004, A new natural high-pressure
859 (Na,Ca)-hexaluminosilicate $[(\text{Ca}_x\text{Na}_{1-x})\text{Al}_3+x\text{Si}_{3-x}\text{O}_{11}]$ in shocked Martian meteorites:
860 Earth and Planetary Science Letters, v. 219, p. 1–12, doi.org/10.1016/S0012-
861 821X(03)00695-2.
- 862 Binns, R.A., 1967, Stony meteorites bearing maskelynite: Nature, v. 213, p. 1111–1112.
- 863 Bischoff, A., and Stöffler, D., 1984, Chemical and structural changes induced by thermal
864 annealing of shocked feldspar inclusions in impact melt rocks from Lappajaervi Crater,
865 Finland (W. V Boynton & G. Schubert, Eds.): Journal of Geophysical Research, v. 89,
866 Suppl., p. 645–656, doi:http://dx.doi.org/10.1029/JB089iS02p0B645.
- 867 Bischoff, A., Schleiting, M., and Patzek, M., 2019, Shock stage distribution of 2280 ordinary
868 chondrites—Can bulk chondrites with a shock stage of S6 exist as individual rocks?
869 Meteoritics and Planetary Science, v. 54, p. 2189–2202, doi:10.1111/maps.13208.
- 870 Bowling, T.J., Johnson, B.C., Wiggins, S.E., Walton, E.L. , Melosh, H.J. Sharp, T.G., 2020,
871 Dwell time at high pressure of meteorites during impact ejection from Mars: Icarus, 343,
872 113689, doi.org/10.1016/j.icarus.2020.113689.
- 873 Bunch, T.E., Cohen, A.J., and Dence, M.R., 1967, Natural terrestrial maskelynite: American
874 Mineralogist, v. 52, p. 244–253.
- 875 Chen, M., and El Goresy, A., 2000, The nature of maskelynite in shocked meteorites; not
876 diaplectic glass but a glass quenched from shock-induced dense melt at high pressures:
877 Earth and Planetary Science Letters, v. 179, p. 489–502.

- 878 Chen, M., Sharp, T.G., El Goresy, A., Wopenka, B., and Xie, X., 1996, The majorite-pyrope +
879 magnesiowüstite assemblage: Constraints on the history of shock veins in chondrites:
880 Science, v. 271, p. 1570–1573, doi:10.1126/science.271.5255.1570.
- 881 Dacheil, F., Gigl, P., and Simons, P.Y., 1968, Experimental and analytical studies of crystalline
882 damage useful for the recognition of impact structures, in French, B.M. and Short, N.M.
883 eds., Shock metamorphism of natural materials, Baltimore, MD, Mono Book Corp., p.
884 555–570.
- 885 Daniel, I., Gillet, P., McMillan, P.F., Wolf, G., and Verhelst, M.A., 1997, High-pressure
886 behavior of anorthite: Compression and amorphization: Journal of Geophysical Research:
887 Solid Earth, v. 102, p. 10313–10325, doi:10.1029/97jb00398.
- 888 Davison, T. M., Collins, G. S., and Ciesla, F. J., 2010, Numerical modelling of heating in porous
889 planetesimal collisions: Icarus, v. 208, p. 468-481.
- 890 De Carli, P.S., and Jamieson, J.C., 1959, Formation of an amorphous form of quartz under shock
891 conditions: The Journal of Chemical Physics, v. 31, p. 1675–1676, doi:10.1063/1.1730673.
- 892 Deer, W.A., Howie, R.A., and Zussman, J., 2001, Rock-forming minerals, Volume 4a:
893 Framework silicates - Feldspars: London, The Geological Society, 972 p.
- 894 Dence, M. R., Douglas, J. A. V., Plant, A. G., and Traill, R. J., 1970, Petrology, Mineralogy, and
895 deformation of Apollo 11 samples: Proceedings of the Apollo 11 Lunar Science
896 Conference, v. 1, p 315-340.
- 897 Diemann, E., and Arndt, J., 1984, Diaplectic Labradorite Glass from the Manicouagan Impact
898 Crater: II. X-Ray Diffraction Studies and Structural Model: Physics and Chemistry of
899 Minerals, v. 11, p. 178-181.
- 900 Doeglas, D.J., 1940, Reliable and rapid method for distinguishing quartz and untwinned feldspar
901 with the universal stage: American Mineralogist, v. 25, p. 286.
- 902 Dressler, B., 1990, Shock metamorphic features and their zoning and orientation in the
903 Precambrian rocks of the Manicouagan Structure, Quebec, Canada: Tectonophysics, v.
904 171, p. 229–245, doi:10.1016/0040-1951(90)90101-D.
- 905 Dworak, U., 1969, Stoßwellenmetamorphose des Anorthosits vom Manicouagan Krater, Quebec,
906 Canada: Contributions to Mineralogy and Petrology, v. 24, p. 306–347.
- 907 El Goresy, A., Chen, M., Gillet, P., and Dubrovinsky, L.S., 2000, Shock-induced high-pressure
908 phase-transition of labradorite to hollandite “(Na₄₇-Ca₅₁-K₂)” in Zagami and the
909 assemblage hollandite “(Na₈₀-Ca₁₂-K₈)” + jadeite in L chondrites: constraints to peak-
910 shock pressures (D. W. G. Sears et al., Eds.): Meteoritics & Planetary Science, v. 35, p.
911 A51–A51.
- 912 El Goresy, A., Gillet, P., Miyahara, M., Ohtani, E., Ozawa, S., Beck, P., and Montagnac, G.,
913 2013, Shock-induced deformation of Shergottites: Shock-pressures and perturbations of

- 914 magmatic ages on Mars: *Geochimica et Cosmochimica Acta*, v. 101, p. 233–262,
915 doi:10.1016/j.gca.2012.10.002.
- 916 Engelhardt, W. von, and Bertsch, W., 1969, Shock induced planar deformation structures in
917 quartz from the ries crater, Germany: *Contributions to Mineralogy and Petrology*, v. 21, p.
918 378, doi:10.1007/BF02672809.
- 919 Engelhardt, W. von, and Stöffler, D., 1968, Stages of shock metamorphism in crystalline rocks
920 of the Ries Basin, Germany., in French, B.M. and Short, N.M. eds., *Shock metamorphism
921 of natural materials*, Baltimore, MD, Mono Book Corp., p. 159–168.
- 922 Engelhardt, W. von, Arndt, J., Stöffler, D., Muller, W.F., Jeziorkowski, H., and Gubser, R.,
923 1967, Diaplektische Gläser in den Breccien des Ries von Nördlingen als Anzeichen für
924 Stosswellenmetamorphose: *Contributions to Mineralogy and Petrology*, v. 15, p. 93–102.
- 925 Engelhardt, W. von, Arndt, J., Mueller, W.F., and Stöffler, D., 1970, Shock metamorphism of
926 lunar rocks and origin of the regolith at the Apollo 11 landing site: *Proceedings of the
927 Apollo 11 Lunar Science Conference*, v. 1, p. 363–384.
- 928 Evans, W.J., Yoo, C.S., Lee, G.W., Cynn, H., Lipp, M.J., and Visbeck, K., 2007, Dynamic
929 diamond anvil cell (dDAC): A novel device for studying the dynamic-pressure properties
930 of materials: *Review of Scientific Instruments*, v. 78, doi:10.1063/1.2751409.
- 931 Fairbairn, H.W., and Podolsky, T., 1951, Notes on precision and accuracy of optic angle
932 determination with the universal stage: *American Mineralogist*, v. 36, p. 823.
- 933 Fernandes, V.A., Burgess, R., and Morris, A., 2009, ^{40}Ar - ^{39}Ar age determinations of lunar basalt
934 meteorites Asuka 881757, Yamato 793169, Miller Range 05035, LaPaz Icefield 02205,
935 Northwest Africa 479, and basaltic breccia Elephant Moraine 96008: *Meteoritics &
936 Planetary Science*, v. 44, p. 805–821. doi.org/10.1111/j.1945-5100.2009.tb00770.x.
- 937 Ferrière, L., and Osinski, G.R., 2013, Shock Metamorphism, in Osinski, G.R. and Pierazzo, E.
938 eds., *Impact Cratering: Processes and Products*, Wiley-Blackwell, p. 106–124.
- 939 Ferrière, L., Morrow, J.R., Amgaa, T., and Koeberl, C., 2009, Systematic study of universal-
940 stage measurements of planar deformation features in shocked quartz: Implications for
941 statistical significance and representation of results: *Meteoritics & Planetary Science*, v.
942 44, p. 925–940.
- 943 Ferrière, L., Koeberl, C., Libowitzky, E., Reimold, W.U., Greshake, A., and Brandstätter, F.,
944 2010, Ballen quartz and cristobalite in impactites: New investigations, in Gibson, R.L. and
945 Reimold, W.U. eds., *Large Meteorite Impacts and Planetary Evolution IV*, Geological
946 Society of America, v. 465, p. 0, doi:10.1130/2010.2465(29).
- 947 Finch, A.A., and Klein, J., 1999, The causes and petrological significance of
948 cathodoluminescence emissions from alkali feldspars: *Contributions to Mineralogy and
949 Petrology*, v. 135, p. 234–243, doi:10.1007/s004100050509.

- 950 Flemming, R.L., 2007, Micro X-ray diffraction (μ XRD); a versatile technique for characterization
951 of Earth and planetary materials: *Canadian Journal of Earth Sciences*, v. 44, p. 1333–1346.
- 952 French, B.M., 1968, Sudbury structure, Ontario: Some petrographic evidence for an origin by
953 meteorite impact, in French, B.M. and Short, N.M. eds., *Shock metamorphism of natural*
954 *materials*, Baltimore, Mono Book Corp., p. 383–411.
- 955 French, B.M., 1998, *Traces of catastrophe, a handbook of shock-metamorphic effects in*
956 *terrestrial meteorite impact structures*: Houston, TX, USA, Lunar and Planetary Institute,
957 120 p.
- 958 French, B.M., and Koeberl, C., 2010, The convincing identification of terrestrial meteorite
959 impact structures: What works, what doesn't, and why: *Earth-Science Reviews*, v. 98, p.
960 123–170, doi:10.1016/j.earscirev.2009.10.009.
- 961 French, B.M., and Short, N.M. (ed.), 1968, *Shock metamorphism of natural materials* (B. M.
962 French & N. M. Short, Eds.): Baltimore, Mono Book Corp., 644 p.
- 963 French, B., Hartung, J., Short, N., and Dietz, R., 1970, Tenoumer Crater, Mauritania. Age and
964 petrologic evidence for origin by meteorite impact: *Journal of Geophysical Research*, v. 75,
965 p. 4396–4406, doi:10.1029/jb075i023p04396.
- 966 French, B.M., Koeberl, C., Gilmour, I., Shirey, S.B., Dons, J.A., and Naterstad, J., 1997, The
967 Gardnos impact structure, Norway: Petrology and geochemistry of target rocks and
968 impactites: *Geochimica et Cosmochimica Acta*, v. 61, p. 873–904, doi:10.1016/S0016-
969 7037(96)00382-1.
- 970 French, B. M., Cordua, W. S., and Plescia, J. B., 2004, The Rock Elm meteorite impact structure,
971 Wisconsin: Geology and shock-metamorphic effects in quartz: *GSA Bulletin*, v. 116, p.
972 200–218.
- 973 Fritz, J., Greshake, A., and Stöffler, D., 2005a, Micro-Raman spectroscopy of plagioclase and
974 maskelynite in Martian meteorites: Evidence of progressive shock metamorphism:
975 *Antarctic Meteorite Research*, v. 18, p. 96–116.
- 976 Fritz J., Artemieva N. A., and Greshake A., 2005b, Ejection of Martian meteorites: *Meteoritics &*
977 *Planetary Science*, v. 40, p. 1393–1411.
- 978 Fritz, J., Wünnemann, K., Reimold, W.U., Meyer, C., and Hornemann, U., 2011, Shock
979 experiments on quartz targets pre-cooled to 77 K: *International Journal of Impact*
980 *Engineering*, v. 38, p. 440–445, doi:10.1016/j.ijimpeng.2010.10.014.
- 981 Fritz, J., Greshake, A., and Fernandes, V.A., 2017, Revising the shock classification of
982 meteorites: *Meteoritics and Planetary Science*, v. 52, p. 1216–1232,
983 doi:10.1111/maps.12845.
- 984 Fritz, J., Assis Fernandes, V., Greshake, A., Holzwarth, A., and Böttger, U., 2019a, On the
985 formation of diaplectic glass: Shock and thermal experiments with plagioclase of different

- 986 chemical compositions: *Meteoritics and Planetary Science*, v. 54, p. 1533–1547,
987 doi:10.1111/maps.13289.
- 988 Fritz, J., Greshake, A., Klementova, M., Wirth, R., Palatinus, L., Assis Fernandes, V., Bottger,
989 U., and Ferriere, L., 2019b, Donwilhelmsite, IMA 2018-113. . CNMNC Newsletter No. 47,
990 February 2019, page 199: *European Journal of Mineralogy*, v. 31, p. 197–202.
- 991 Gibbons, R., and Ahrens, T., 1977, Effects of shock pressures on calcic plagioclase: *Physics and*
992 *Chemistry of Minerals*, v. 1, p. 95–107.
- 993 Gibson, R.L., and Reimold, W.U., 2005, Shock pressure distribution in the Vredefort impact
994 structure, South Africa, in Kenkmann, T., Hörz, F.P., and Deutsch, A. eds., *Special Paper -*
995 *Geological Society of America*, Geological Society of America (GSA), Boulder, CO, v.
996 384, p. 329–349.
- 997 Gillet, P., Chen, M., Dubrovinsky, L., and El Goresy, A., 2000, Natural NaAlSi₃O₈-hollandite in
998 the shocked sixiangkou meteorite: *Science (New York, N.Y.)*, v. 287, p. 1633–6.
- 999 Gillet, P., El Goresy, A., Beck, P., and Chen, M., 2007, High-pressure mineral assemblages in
1000 shocked meteorites and shocked terrestrial rocks: Mechanisms of phase transformations
1001 and constraints to pressure and temperature histories, in Ohtani, E. ed., *Advances in High-*
1002 *Pressure Mineralogy*, Geological Society of America, v. 421, p. 0,
1003 doi:10.1130/2007.2421(05).
- 1004 Götte, T., 2009, Petrological modifications in continental target rocks from terrestrial impact
1005 structures: evidence from cathodoluminescence, in Gucsik, A. ed., *Cathodoluminescence*
1006 *and its application in the planetary sciences*, Heidelberg, Springer-Verlag, p. 45–60.
- 1007 Götze, J., 2009, Cathodoluminescence microscopy and spectroscopy of lunar rocks and minerals
1008 BT - Cathodoluminescence and its application in the planetary sciences, *in* Gucsik, A. ed.,
1009 Berlin, Heidelberg, Springer Berlin Heidelberg, p. 87–110, doi:10.1007/978-3-540-87529-
1010 1_5.
- 1011 Götze, J., and Kempe, U., 2009, Physical principles of cathodoluminescence (CL) and its
1012 applications in geosciences, *in* Gucsik, A. ed., *Physical principles of Cathodoluminescence*
1013 *(CL) and its applications in geosciences*, Berlin, Springer-Verlag, p. 1–22.
- 1014 Götze, J., Krbetschek, M.R., Habermann, D., and Wolf, D., 2000, High-resolution
1015 cathodoluminescence studies of feldspar minerals BT - Cathodoluminescence *in*
1016 *geosciences*, in Pagel, M., Barbin, V., Blanc, P., and Ohnenstetter, D. eds., Berlin,
1017 Heidelberg, Springer Berlin Heidelberg, p. 245–270, doi:10.1007/978-3-662-04086-7_10.
- 1018 Grady, D.E., 1977, Processes occurring in shock wave compression of rocks and minerals, *in*
1019 Manghnani, M.H. and Akimoto, S.I. eds., *High-pressure research: Applications in*
1020 *Geophysics*, New York, Academic Press, p. 389–438.
- 1021 Grady, D.E., 1980, Shock deformation of brittle solids: *Journal of Geophysical Research*, v. 85,
1022 p. 913–924, doi:http://dx.doi.org/10.1029/JB085iB02p00913.

- 1023 Grady, D.E., and Murri, W.J., 1976, Dynamic unloading in shock compressed feldspar:
1024 Geophysical Research Letters, v. 3, p. 472–474.
- 1025 Grady, D.E., Murri, W.J., and De Carli, P.S., 1975, Hugoniot sound velocities and phase
1026 transformations in two silicates: Journal of Geophysical Research, v. 80, p. 4857–4861,
1027 doi:<http://dx.doi.org/10.1029/JB080i035p04857>.
- 1028 Gratz, A.J., Nellis, W.J., Christie, J.M., Brocious, W., Swegle, J., and Cordier, P., 1992, Shock
1029 metamorphism of quartz with initial temperatures -170 to + 1000°C: Physics and
1030 Chemistry of Minerals, v. 19, p. 267–288, doi:10.1007/BF00204005.
- 1031 Grieve, R.A.F., 1975, Petrology and chemistry of the impact melt at Mistastin Lake crater,
1032 Labrador: Geological Society of America Bulletin, v. 86, p. 1617–1629.
- 1033 Gucsik, A., Ninagawa, K., Nishido, H., Toyoda, S., Tsuchiyama, A., Bidló, A., and
1034 Brezsnýánszky, K., 2004, Cathodoluminescence (CL) spectral study of experimentally
1035 shock-deformed plagioclase, *in* Workshop on Chondrites and Protoplanetary Disk,
1036 Houston, TX, United States (USA), Lunar and Planetary Institute, Houston, TX, p.
1037 Abstract #9012.
- 1038 Haff, J.C., 1940, Use of the Wulff net in mineral determination with the Universal stage:
1039 American Mineralogist, v. 24, p. 689–707.
- 1040 Harris, R., Schultz, P., and Jaret, S., 2019, Critical comparison of textures observed in impact
1041 melt and volcanic rocks: Geological Society of America Abstracts with Programs,
1042 doi:10.1130/abs/2019AM-338073.
- 1043 Hemley, R.J., Jephcoat, A.P., Mao, H.K., Ming, L.C., and Manghnani, M.H., 1988, Pressure-
1044 induced amorphization of crystalline silica: Nature, v. 334, p. 52–54,
1045 doi:10.1038/334052a0.
- 1046 Hörz, F., 1968, Statistical measurements of deformation structures and refractive indices in
1047 experimentally shock loaded quartz, *in* French, B.M. and Short, N.M. eds., Shock
1048 metamorphism of natural materials, Baltimore, MD, Mono Book Corp., p. 243–253.
- 1049 Hörz, F., and Quaide, W.L., 1973, Debye-Scherrer investigations of experimentally shocked
1050 silicates: The Moon, v. 6, p. 45–82.
- 1051 Huffman, A.R., and Reimold, W.U., 1996, Experimental constraints on shock-induced
1052 microstructures in naturally deformed silicates: Tectonophysics, v. 256, p. 165–217,
1053 doi:10.1016/0040-1951(95)00162-x.
- 1054 Huffman, A.R., Brown, J.M., Carter, N.L., and Reimold, W.U., 1993, The microstructural
1055 response of quartz and feldspar under shock loading at variable temperatures: Journal of
1056 Geophysical Research, v. 98, p. 22171–22197, doi:10.1029/93JB01425.
- 1057 Iiishi, K.T., Tomisaka, T., Kato, T., and Umegaki, Y., 1971, Isomorphous substitution and
1058 infrared and far infrared spectra of the feldspar group: Neues Jahrbuch fuer Mineralogie.
1059 Abhandlungen, v. 115, p. 98–119.

- 1060 Irifune, T., Ringwood, A. E., and Hibberson, W. O., 1994, Subduction of continental crust and
1061 terrigenous and pelagic sediments: an experimental study: *Earth and Planetary Science*
1062 *Letters*, v. 126, p. 351-368.
- 1063 James, O.B., 1969, Jadeite: Shock-induced formation from oligoclase, Ries Crater, Germany:
1064 *Science*, v. 165, p. 1005 LP – 1008, doi:10.1126/science.165.3897.1005.
- 1065 Janssen, C., Wirth, R., Rybacki, E., Naumann, R., Kemnitz, H., Wenk, H.-R., and Dresen, G.,
1066 2010, Amorphous material in SAFOD core samples (San Andreas Fault): Evidence for
1067 crush-origin pseudotachylytes? *Geophysical Research Letters*, v. 37,
1068 doi:10.1029/2009GL040993.
- 1069 Jaret, S.J., Kah, L.C., and Harris, R.S., 2014, Progressive deformation of feldspar recording low-
1070 barometry impact processes, Tenoumer impact structure, Mauritania: *Meteoritics &*
1071 *Planetary Science*, v. 49, p. 1007–1022, doi:10.1111/maps.12310.
- 1072 Jaret, S.J., Woerner, W.R., Phillips, B.L., Ehm, L., Nekvasil, H., Wright, S.P., and Glotch, T.D.,
1073 2015, Maskelynite formation via solid-state transformation: Evidence of infrared and X-ray
1074 anisotropy: *Journal of Geophysical Research : Planets*, v. 120, p. 570-587.
- 1075 Jaret, S.J., Johnson, J.R., Sims, M., DiFrancesco, N., and Glotch, T.D., 2018, Microspectroscopic
1076 and petrographic comparison of experimentally shocked albite, andesine, and bytownite:
1077 *Journal of Geophysical Research: Planets*, v. 123, p. 1701–1722,
1078 doi:10.1029/2018JE005523.
- 1079 Jessberger, E., and Ostertag, R., 1982, Shock-effects on the K-Ar system of plagioclase feldspar
1080 and the age of anorthosite inclusions from North-Eastern Minnesota: *Geochimica et*
1081 *Cosmochimica Acta*, v. 46, p. 1465 – 1471.
- 1082 Johnson, J.R., 2003, Visible/near-infrared spectra of experimentally shocked plagioclase
1083 feldspars: *Journal of Geophysical Research*, v. 108, p. 5120, doi:10.1029/2003JE002127.
- 1084 Johnson, J.R., Hörz, F., Lucey, P.G., and Christensen, P.R., 2002a, Thermal infrared
1085 spectroscopy of experimentally shocked anorthosite and pyroxenite; implications for
1086 remote sensing of Mars: *Journal of Geophysical Research*, v. 107, p. 14,
1087 doi:http://dx.doi.org/10.1029/2001JE001517.
- 1088 Johnson, J.R., Staid, M.I., and Titus, T.N., 2002b, Shocked plagioclase signatures in thermal
1089 emission spectrometer data of Mars: *Abstracts of Papers Submitted to the Lunar and*
1090 *Planetary Science Conference*, v. 33, p. 0-Abstract 1345.
- 1091 Johnson, J.R., Hörz, F., and Staid, M.I., 2003, Thermal infrared spectroscopy and modeling of
1092 experimentally shocked plagioclase feldspars: *American Mineralogist*, v. 88, p. 1575–
1093 1582.
- 1094 Kaus, A., and Bischoff, A., 2000, Cathodoluminescence (CL) properties of shocked plagioclase:
1095 *Meteoritics & Planetary Science*, v. 35, p. A86.

- 1096 Kayama, M., Gucsik, A., Nishido, H., Ninagawa, K., Tsuchiyama, A., and Goetze, J., 2009a,
1097 Cathodoluminescence and Raman spectroscopic characterization of experimentally
1098 shocked plagioclase: LPI Contribution, p. 49.
- 1099 Kayama, M., Nakazato, T., Nishido, H., Ninagawa, K., Gucsik, A., and Goetze, J., 2009b,
1100 Cathodoluminescence characterization of maskelynite and alkali feldspar in shergottite:
1101 LPI Contribution, p. 50.
- 1102 Kayama, M., Nishido, H., Ninagawa, K., Gucsik, A., and Goetze, J., 2009c,
1103 Cathodoluminescence of shocked plagioclase and alkali feldspar from Ries Crater,
1104 Germany: LPI Contribution, p. 51.
- 1105 Kayama, M., Nishido, H., Sekine, T., Nakazato, T., Gucsik, A., and Ninagawa, K., 2012, Shock
1106 barometer using cathodoluminescence of alkali feldspar: Journal of Geophysical Research,
1107 v. 117, p. E09004.
- 1108 Kayama, M., Sekine, T., Tomioka, N., Nishido, H., Kato, Y., Ninagawa, K., Kobayashi, T., and
1109 Yamaguchi, A., 2018, Cathodoluminescence of high-pressure feldspar minerals as a shock
1110 barometer: Meteoritics and Planetary Science, v. 53, p. 1476–1488,
1111 doi:10.1111/maps.13092.
- 1112 Kieffer, S.W., 1971, Shock metamorphism of the Coconino sandstone at Meteor Crater, Arizona:
1113 Journal of Geophysical Research, v. 76, p. 5449–5473.
- 1114 Kieffer, S.W., Schaal, R.B., Gibbons, R., Horz, R., Milton, D.J., and Dube, A., 1976, Shocked
1115 basalt from Lonar impact Crater, India, and experimental analogues: Lunar and Planetary
1116 Science Conference Proceedings, v. 1, p. 1391–1412.
- 1117 Kleeman, J.D., 1971, Formation of diaplectic glass by experimental shock loading of orthoclase:
1118 Journal of Geophysical Research, v. 76, p. 5499–5503, doi:10.1029/JB076i023p05499.
- 1119 Klein, C., and Philpotts, A., 2016, Introduction to mineralogy and petrology: Cambridge,
1120 Cambridge University Press.
- 1121 Knight, K.S., Price G.D., 2008, Powder Neutron-diffraction studies of clinopyroxenes. I. The
1122 crystal structure and thermoelastic properties of jadeite between 1.5 and 270 K: The
1123 Canadian Mineralogist, v. 46, p. 1593–1622.
- 1124 Koeberl, C., Reimold, W.U., Kracher, A., Träxler, B., Vormaiier, A., and Körner, W., 1996,
1125 Mineralogical, petrological, and geochemical studies of drill core samples from the
1126 Manson impact structure, Iowa, *in* Koeberl, C. and Anderson, R.R. eds., The Manson
1127 impact structure, Iowa; anatomy of an impact crater, Geological Society of America, v.
1128 302, p. 0, doi:10.1130/0-8137-2302-7.145.
- 1129 Kubo, T., Kimura, M., Kato, T., Nishi, M., Tominaga, A., Kikegawa, T., and Funakoshi, K.,
1130 2010, Plagioclase breakdown as an indicator for shock conditions of meteorites: Nature
1131 Geoscience, v. 3, p. 41–45, doi:10.1038/NNGEO704.

- 1132 Kubo, T., Kono, M., Imamura, M., Kato, T., Uehara, S., Kondo, T., Higo, Y., Tange, Y., and
1133 Kikegawa, T., 2017, Formation of a metastable hollandite phase from amorphous
1134 plagioclase: A possible origin of lingunite in shocked chondritic meteorites: *Physics of the*
1135 *Earth and Planetary Interiors*, v. 272, p. 50–57, doi:10.1016/j.pepi.2017.09.006.
- 1136 Lambert, P., 1979, Fractures induced by shock in quartz and feldspar: *Mineralogical Magazine*,
1137 v. 43, p. 527–533, doi:10.1180/minmag.1979.043.328.13.
- 1138 Langenhorst, F., 2002, Shock metamorphism of some minerals: Basic introduction and
1139 microstructural observations: *Bulletin of the Czech Geological Survey*, v. 77, p. 265–282.
- 1140 Langenhorst, F., and Deutsch, A., 1994, Shock experiments on pre-heated α - and β -quartz: I.
1141 Optical and density data: *Earth and Planetary Science Letters*, v. 125, p. 407–420,
1142 doi:10.1016/0012-821X(94)90229-1.
- 1143 Langenhorst, F., and Deutsch, A., 2012, Shock metamorphism of minerals: *Elements*, v. 8, p.
1144 31–36, doi:10.2113/gselements.8.1.31.
- 1145 Langenhorst, F., and Dressler, B., 2003, First observation of silicate hollandite in a terrestrial
1146 rock: *LPI Contribution*, v. 1167, p. 0-Abstract 4046.
- 1147 Langenhorst, F., and Poirier, J.-P., 2000, Anatomy of black veins in Zagami; clues to the
1148 formation of high-pressure phases: *Earth and Planetary Science Letters*, v. 184, p. 37–55,
1149 doi:http://dx.doi.org/10.1016/S0012-821X(00)00317-4.
- 1150 Langenhorst, F., Deutsch, A., Stöffler, D., and Hornemann, U., 1992, Effect of temperature on
1151 shock metamorphism of single-crystal quartz: *Nature*, v. 356, p. 507–509,
1152 doi:10.1038/356507a0.
- 1153 Langenhorst, F., Joreau, P., and Doukhan, J.C., 1995, Thermal and shock metamorphism of the
1154 Tenham chondrite: A TEM examination: *Geochimica et Cosmochimica Acta*, v. 59, p.
1155 1835–1845, doi:10.1016/0016-7037(95)00086-F.
- 1156 Lee, M.R., Parsons, I., Edwards, P.R., and Martin, R.W., 2007, Identification of
1157 cathodoluminescence activators in zoned alkali feldspars by hyperspectral imaging and
1158 electron-probe microanalysis: *American Mineralogist*, v. 92, p. 243–253,
1159 doi:10.2138/am.2007.2160.
- 1160 Leichmann, J., Broska, I., and Zachovalová, K., 2003, Low-grade metamorphic alteration of
1161 feldspar minerals: A CL study: *Terra Nova*, v. 15, p. 104–108, doi:10.1046/j.1365-
1162 3121.2003.00467.x.
- 1163 Letoulllec, R., Pinceaux, J.P., and Loubeyre, P., 1988, The membrane diamond anvil cell: A new
1164 device for generating continuous pressure and temperature variations: *high pressure*
1165 *research*, v. 1, p. 77–90, doi:10.1080/08957958808202482.
- 1166 Liu, L.-G., 1978, High-pressure phase transformations of albite, jadeite and nepheline: *Earth and*
1167 *Planetary Science Letters*, v. 37, p. 438–444, doi:10.1016/0012-821X(78)90059-6.

- 1168 Liu, X., 2006, Phase relations in the system $\text{KAlSi}_3\text{O}_8\text{--NaAlSi}_3\text{O}_8$ at high pressure–high
1169 temperature conditions and their implication for the petrogenesis of lingunite: *Earth and*
1170 *Planetary Science Letters*, v. 246, p. 317-325.
- 1171 Liu, L.-G., and El Gorse, A., 2007, High-pressure phase transitions of the feldspars, and further
1172 characterization of lingunite: *International Geology Review*, v. 49, p. 854–860,
1173 doi:10.2747/0020-6814.49.9.854.
- 1174 Lyon, R.J.P., 1963, Evaluation of infrared spectrophotometry for compositional analysis of lunar
1175 and planetary soils: NASA Technical Note D-1871.
- 1176 Ma, C., Tschauner, O., Beckett, J.R., Liu, Y., Rossman, G.R., Zhuravlev, K., Prakapenka, V.,
1177 Dera, P., and Taylor, L.A., 2015, Tissintite, $(\text{Ca},\text{Na}, \oplus)\text{AlSi}_2\text{O}_6$, a highly-defective, shock-
1178 induced, high-pressure clinopyroxene in the Tissint martian meteorite: *Earth and Planetary*
1179 *Science Letters*, v. 422, p. 194–205, doi:10.1016/j.epsl.2015.03.057.
- 1180 Ma, C., Tschauner, O. and Beckett, J.R., 2017, A new high-pressure calcium aluminosilicate
1181 $(\text{CaAl}_2\text{Si}_{3.5}\text{O}_{11})$ in martian meteorites: another after-life for plagioclase and connections to
1182 the CAS phase: *Lunar and Planetary Science XLVIII*, 1128.
- 1183 Ma C., Tschauner, O., Beckett, J.R., Rossman G. R., Prescher C., Prakapenka, V.B. Bechtel, H.
1184 A., and MacDowell, A., 2018, Liebermannite, KAlSi_3O_8 , a new shock-metamorphic, high-
1185 pressure mineral from the Zagami Martian meteorite: *Meteoritics and Planetary Science*, v.
1186 53, p. 50-61.
- 1187 Marsh, S. P. (Ed.), 1980, LASL shock Hugoniot data: University of California Press, Berkeley.
- 1188 Melosh, H.J., 1989, Impact cratering: A geologic process: New York, Oxford University Press,
1189 245 p.
- 1190 Milton, D.J., and de Carli, P.S., 1963, Maskelynite: formation by explosive shock: *Science*, v.
1191 140, p. 670 LP – 671, doi:10.1126/science.140.3567.670.
- 1192 Moreau, J., Kohout, T., and Wünnemann, K., 2017, Shock-darkening in ordinary chondrites:
1193 Determination of the pressure-temperature conditions by shock physics mesoscale
1194 modeling: *Meteoritics and Planetary Science*, v. 52, p. 2375–2390,
1195 doi:10.1111/maps.12935.
- 1196 Müller, W.F., and Hornemann, U., 1969, Shock-induced planar deformation structures in
1197 experimentally shock-loaded olivines and in olivines from chondritic meteorites: *Earth and*
1198 *Planetary Science Letters*, v. 7, p. 251–264, doi:10.1016/0012-821X(69)90062-4.
- 1199 Myers, S.A., Cygan, R.T., Assink, R.A., and Boslough, M.B., 1998, ^{29}Si MAS NMR relaxation
1200 study of shocked Coconino Sandstone from Meteor Crater, Arizona: *Physics and*
1201 *Chemistry of Minerals*, v. 25, p. 313–317, doi:10.1007/s002690050120.
- 1202 Nagy, S., Gucsik, A., Bérczi, S., Ninagawa, K., Nishido, H., Kereszturi, A., Hargitai, H., and
1203 Okumura, T., 2008, K-feldspar and biotite as shock indicator minerals from Bosumtwi

- 1204 impact crater: Abstracts of Papers Submitted to the Lunar and Planetary Science
1205 Conference, v. 39, Abstract 1144.
- 1206 Nelson, S.T., and Montana, A., 1992, Sieve-textured plagioclase in volcanic rocks produced by
1207 rapid decompression: *American Mineralogist*, v. 77, p. 1242–1249.
- 1208 Ogilvie, P., Gibson, R.L., Reimold, W.U., Deutsch, A., and Hornemann, U., 2011, Experimental
1209 investigation of shock metamorphic effects in a metapelitic granulite: The importance of
1210 shock impedance contrast between components: *Meteoritics and Planetary Science*, v. 46,
1211 p. 1565–1586, doi:10.1111/j.1945-5100.2011.01250.x.
- 1212 Ohtani, E., Kimura, Y., Kimura, M., Takata, T., Kondo, T., and Kubo, T., 2004, Formation of
1213 high-pressure minerals in shocked L6 chondrite Yamato 791384: constraints on shock
1214 conditions and parent body size: *Earth and Planetary Science Letters*, v. 227, p. 505–515,
1215 doi:10.1016/j.epsl.2004.08.018.
- 1216 Okuno, M., 2003, Structural evolution of quartz and feldspar crystals and their glasses by shock
1217 compression, *in* Davison, L., Horie, Y., and Sekine, T. eds., *High-pressure shock
1218 compression of solids V. Shock wave and high pressure phenomena*, New York.
- 1219 Ostertag, R., 1982, Annealing behaviour of diaplectic plagioclase glass, *in* *Lunar and Planetary
1220 Science Conference*, p. 607–608.
- 1221 Ostertag, R., 1983, Shock experiments on feldspar crystals: *Journal of Geophysical Research*, v.
1222 88, Suppl., p. B364–B376, doi:10.1029/JB088iS01p0B364.
- 1223 Papike, J.J., Taylor, L., and S., S., 1991, Lunar minerals, *in* Heiken, G.H., Vaniman, D., and
1224 French, B.M. eds., *Lunar sourcebook: A user's guide to the Moon*, Cambridge, Cambridge
1225 University Press, p. 736.
- 1226 Parsons, I., and Lee, M.R., 2009, Mutual replacement reactions in alkali feldspars I:
1227 Microtextures and mechanisms: *Contributions to Mineralogy and Petrology*, v. 157, p.
1228 641–661, doi:10.1007/s00410-008-0355-4.
- 1229 Parsons, I., Steele, D.A., Lee, M.R., and Magee, C.W., 2008, Titanium as a cathodoluminescence
1230 activator in alkali feldspars: *American Mineralogist*, v. 93, p. 875–879,
1231 doi:10.2138/am.2008.2711.
- 1232 Parsons, I., Gerald, J.D.F., and Lee, M.R., 2015, Routine characterization and interpretation of
1233 complex alkali feldspar intergrowths: *American Mineralogist*, v. 100, p. 1277–1303,
1234 doi:10.2138/am-2015-5094.
- 1235 Pati, J.K., Jourdan, F., Armstrong, R.A., Reimold, W.U., and Prakash, K., 2010, First SHRIMP
1236 U-Pb and $^{40}\text{Ar}/^{39}\text{Ar}$ chronological results from impact melt breccia from the
1237 Paleoproterozoic Dhala impact structure, India, *in* Gibson, R.L. and Reimold, W.U. eds.,
1238 *Large Meteorite Impacts and Planetary Evolution IV*, Geological Society of America, v.
1239 465, doi:10.1130/2010.2465(27).

- 1240 Pickersgill, A.E., and Lee, M.R., 2015, Identifying planar deformation features using EBSD and
1241 FIB, in *Bridging the Gap III: Impact Cratering In Nature, Experiments, and Modeling*, p.
1242 1056.
- 1243 Pickersgill, A.E., Flemming, R.L., and Osinski, G.R., 2015a, Toward quantification of strain-
1244 related mosaicity in shocked lunar and terrestrial plagioclase by in situ micro-X-ray
1245 diffraction: *Meteoritics and Planetary Science*, v. 50, p. 1851–1862,
1246 doi:10.1111/maps.12514.
- 1247 Pickersgill, A.E., Osinski, G.R., and Flemming, R.L., 2015b, Shock effects in plagioclase
1248 feldspar from the Mistastin Lake impact structure, Canada: *Meteoritics and Planetary*
1249 *Science*, v. 50, p. 1546–1561, doi:10.1111/maps.12495.
- 1250 Pickersgill, A.E., Lee, M.R., Daly, L., Mark, D.F., and IODP-ICDP Expedition 364 Science
1251 Party, 2017, Planar microstructures (lamellar subgrains) in feldspar from the Chicxulub
1252 impact structure, in *80th Annual Meeting of the Meteoritical Society*, p. 6182.
- 1253 Pittarello, L., and Koeberl, C., 2017, Shock-induced planar features in plagioclase: a project on
1254 measurements and investigations on their occurrence in relation with the An content, in
1255 *80th Annual Meeting of the Meteoritical Society*.
- 1256 Pittarello, L., Schulz, T., Koeberl, C., Hoffmann, J.E., and Münker, C., 2013, Petrography,
1257 geochemistry, and Hf-Nd isotope evolution of drill core samples and target rocks from the
1258 El'gygytyn impact crater, NE Chukotka, Arctic Russia: *Meteoritics and Planetary*
1259 *Science*, v. 48, p. 1160–1198, doi:10.1111/maps.12088.
- 1260 Pittarello, L., Roszjar, J., Mader, D., Debaille, V., Claeys, P., and Koeberl, C., 2015,
1261 Cathodoluminescence as a tool to discriminate impact melt, shocked and unshocked
1262 volcanics: A case study of samples from the El'gygytyn impact structure: *Meteoritics and*
1263 *Planetary Science*, v. 50, p. 1954–1969, doi:10.1111/maps.12559.
- 1264 Pittarello, L., Daly, L., Pickersgill, A.E., Ferrière, L., and Lee, M.R., 2020a, Shock
1265 metamorphism in plagioclase and selective amorphization: *Meteoritics & Planetary*
1266 *Science*, v. 55, p. 1103–1115, doi:10.1111/maps.13494
- 1267 Pittarello, L., Fritz, J., Roszjar, J., Lenz, C., Chanmuang, C.N., and Koeberl, C., 2020b, Partial
1268 amorphization of experimentally shocked plagioclase: A spectroscopic study: *Meteoritics*
1269 *and Planetary Science*, v. 55, p. 669–678, doi:10.1111/maps.13445.
- 1270 Pittarello, L., Ferrière, L., Feignon, J.-G., Osinski, G. R., and Koeberl, C., 2020c, Preferred
1271 orientation distribution of shock-induced planar microstructures in quartz and feldspar:
1272 *Meteoritics & Planetary Science*, v. 55, p. 1082–1092, doi: 10.1111/MAPS.13490-3304
- 1273 Poelchau, M.H., and Kenkmann, T., 2011, Feather features; a low-shock-pressure indicator in
1274 quartz: *Journal of Geophysical Research*, v. 116, p. B02201,
1275 doi:http://dx.doi.org/10.1029/2010JB007803.

- 1276 Poldervaart, A., and Gilkey, A.K., 1954, On clouded plagioclase: *American Mineralogist*, v. 35,
1277 p. 75–91.
- 1278 Prior, D.J. et al., 1999, The application of electron backscatter diffraction and orientation
1279 contrast imaging in the SEM to textural problems in rocks: *American Mineralogist*, v. 84,
1280 p. 1741–1759, doi:10.2138/am-1999-11-1204.
- 1281 Putnis, A., Hinrichs, R., Putnis, C. V., Golla-Schindler, U., and Collins, L.G., 2007, Hematite in
1282 porous red-clouded feldspars: Evidence of large-scale crustal fluid-rock interaction: *Lithos*,
1283 v. 95, p. 10–18, doi:10.1016/j.lithos.2006.07.004.
- 1284 Reeder, R.J., and Michel, F.M., 2013, Application of total X-ray scattering methods and pair
1285 distribution function analysis for study of structure of biominerals, in *Methods in*
1286 *Enzymology*, doi:10.1016/B978-0-12-416617-2.00020-5.
- 1287 Reimold, W.U., and Miller, R.M., 1989, The Roter Kamm impact crater, SWA/Namibia: *Lunar*
1288 *and Planetary Science Conference Proceedings*, v. 19, p. 711–732.
- 1289 Reinhard, M., 1931, *Universal Drehtischmethoden : Einführung in die kristallographischen*
1290 *Grundbegriffe und die Plagioklasbestimmung*: Basel, Verlag, von B. Wepf & Cie., 119 p.
- 1291 Reynard, B., Okuno, M., Shimada, Y., Syono, Y., and Willaime, C., 1999, A Raman
1292 spectroscopic study of shock-wave densification of anorthite (CaAl₂Si₂O₈) glass: *Physics*
1293 *and Chemistry of Minerals*, v. 26, p. 432–436.
- 1294 Ringwood, A.E., Reid, A.F., and Wadsley, A.D., 1967, High pressure transformation of alkali
1295 aluminosilicates and aluminogermanates: *Earth and Planetary Science Letters*, v. 3, p. 38–
1296 40, doi:http://dx.doi.org/10.1016/0012-821X(67)90008-8.
- 1297 Robertson, P.B., 1973, *Shock metamorphism of potassic feldspars*: Durham University, 326 p.
- 1298 Robertson, P.B., 1975, Experimental shock metamorphism of maximum microcline: *Journal of*
1299 *Geophysical Research (1896-1977)*, v. 80, p. 1903–1910, doi:10.1029/JB080i014p01903.
- 1300 Robertson, P.B., Dence, M.R., and Vos, M.A., 1968, Deformation in rock-forming minerals from
1301 Canadian craters, *in* French, B.M. and Short, N.M. eds., *Shock metamorphism of natural*
1302 *materials*, Baltimore, MD, Mono Book Corp., p. 433–452.
- 1303 Roddy, D.J., Pepin, R.O., and Merrill, R.B. (Eds.), 1978, *Impact and explosion cratering*:
1304 Toronto, Pergamon Press, 1301 p.
- 1305 Rubin, A.E., 1992, A shock-metamorphic model for silicate darkening and compositionally
1306 variable plagioclase in CK and ordinary chondrites: *Geochimica et Cosmochimica Acta*, v.
1307 56, p. 1705–1714, doi:10.1016/0016-7037(92)90236-C.
- 1308 Rucks, M.J., Whitaker, M.L., Glotch, T.D., Parise, J.B., Jaret, S.J., Catalano, T., and Dyar, M.D.,
1309 2018, Making tissintite: Mimicking meteorites in the multi-anvil: *American Mineralogist*,
1310 v. 103, p. 1516–1519, doi:10.2138/am-2018-6539.

- 1311 Seifert, K.E., 1964, The genesis of plagioclase twinning in the Nonnewaug Granite: *The American*
1312 *Mineralogist*, v. 49, p. 297–320.
- 1313 Sharp, T.G., and DeCarli, P., 2006, Shock effects in meteorites, in *Meteorites and the early solar*
1314 *system II*, p. 653–677.
- 1315 Sharp, T.G., Walton, E.L., Hu, J., Agee, C., 2019, Shock conditions recorded in NWA 8159
1316 martian augite basalt with implications for the impact cratering history on Mars:
1317 *Geochimica et Cosmochimica Acta*, v. 246, p. 197-212.
- 1318 Short, N.M., and Gold, D.P., 1996, Petrography of shocked rocks from the central peak at the
1319 Manson impact structure, *in* Koeberl, C. and Anderson, R.R. eds., *Special Paper -*
1320 *Geological Society of America*, Geological Society of America (GSA), Boulder, CO, v.
1321 302, p. 245–265.
- 1322 Sims, M. et al., 2019, Pressure-induced amorphization in plagioclase feldspars: A time-resolved
1323 powder diffraction study during rapid compression: *Earth and Planetary Science Letters*, v.
1324 507, p. 166–174, doi:10.1016/j.epsl.2018.11.038.
- 1325 Sims, M., Jaret, S.J., Johnson, J.R., Whitaker, M.L., and Glotch, T.D., 2020, Unconventional
1326 high-pressure Raman spectroscopy study of kinetic and peak pressure effects in plagioclase
1327 feldspars: *Physics and Chemistry of Minerals*, v. 47, p. 1–10, doi:10.1007/s00269-020-
1328 01080-z.
- 1329 Smith, J. V, and Brown, W.L., 1988, *Feldspar minerals*: Berlin, Springer-Verlag, 828 p.
- 1330 Spry, A.H., 1969, *Metamorphic textures*: Oxford, New York, Pergamon Press, 350
- 1331 Stähle, V., Altherr, R., Nasdala, L., and Ludwig, T., 2011, Ca-rich majorite derived from high-
1332 temperature melt and thermally stressed hornblende in shock veins of crustal rocks from
1333 the Ries impact crater (Germany): *Contributions to Mineralogy and Petrology*, v. 161, p.
1334 275–291, doi:10.1007/s00410-010-0531-1.
- 1335 Stebbins, J.F., and Xue, X., 2014, NMR spectroscopy of inorganic Earth materials: *Reviews in*
1336 *mineralogy and geochemistry*, v. 78, p. 605–653, doi:10.2138/rmg.2014.78.15.
- 1337 Stöffler D., 1966, Zones of Impact Metamorphism in the Crystalline Rocks of the Nördlinger
1338 Ries Crater: *Contributions to Mineralogy and Petrology*, v. 12, p. 15-24.
- 1339 Stöffler, D., 1967, Deformation und Umwandlung von Plagioklas durch Stoßwellen in den
1340 Gesteinen des Nördlinger Ries: *Contributions to Mineralogy and Petrology*, v. 16, p. 51–
1341 83, doi:10.1007/BF00371608.
- 1342 Stöffler, D., 1971, Progressive metamorphism and classification of shocked and brecciated
1343 crystalline rocks at impact craters: *Journal of Geophysical Research*, v. 76, p. 5541–5551.
- 1344 Stöffler, D., 1972, Deformation and transformation of rock-forming minerals by natural and
1345 experimental shock processes: I. Behavior of minerals under shock compression:
1346 *Fortschritte der Mineralogie*, v. 49, p. 50–113.

- 1347 Stöffler, D., 1974, Deformation and transformation of rock-forming minerals by natural and
1348 experimental shock processes: II. Physical properties of shocked minerals: *Fortschritte der*
1349 *Mineralogie*, v. 51, p. 256–289.
- 1350 Stöffler, D., and Grieve, R.A.F., 2007, Impactites, *in* Fettes, D. and Desmons, J. eds.,
1351 *Metamorphic Rocks: A classification and glossary of terms*, Recommendations of the
1352 International Union of Geological Sciences, Cambridge, UK, Cambridge University Press,
1353 p. 82–92, and 111–242.
- 1354 Stöffler, D., and Hornemann, U., 1972, Quartz and feldspar glasses produced by natural and
1355 experimental shock: *Meteoritics*, v. 7, p. 371–394.
- 1356 Stöffler, D., and Langenhorst, F., 1994, Shock metamorphism of quartz in nature and
1357 experiment: I. Basic observation and theory: *Meteoritics*, v. 29, p. 155–181.
- 1358 Stöffler, D., Schulien, S., and Ostertag, R., 1975, Rock 61016: multiphase shock and
1359 crystallization history of a polymict troctolitic-anorthositic breccia: *Lunar and Planetary*
1360 *Science Conference Proceedings*, v. 1, p. 673–692.
- 1361 Stöffler, D., Ostertag, R., Jammes, C., Pfannschmidt, G., Sen Gupta, P. R., Simon, S. B., Papike,
1362 J. J., and Beauchamp, R. H., 1986, Shock metamorphism and petrography of the Shergotty
1363 achondrite: *Geochimica et Cosmochimica Acta* v. 50, p. 889–903.
- 1364 Stöffler, D., Keil, K., and Scott, E.R.D., 1991, Shock metamorphism of ordinary chondrites:
1365 *Geochimica et Cosmochimica Acta*, v. 55, p. 3845–3867,
1366 doi:[http://dx.doi.org/10.1016/0016-7037\(91\)90078-J](http://dx.doi.org/10.1016/0016-7037(91)90078-J).
- 1367 Stöffler, D., Hamann, C., and Metzler, K., 2018, Shock metamorphism of planetary silicate rocks
1368 and sediments: Proposal for an updated classification system: *Meteoritics and Planetary*
1369 *Science*, v. 53, p. 5–49, doi:10.1111/maps.12912.
- 1370 Stöffler, D., Hamann, C., and Metzler, K., 2019, Addendum to “Stöffler, D., Hamann, C., and
1371 Metzler, K., Shock metamorphism of planetary silicate rocks and sediments: Proposal for
1372 an updated classification system. *Meteoritics & Planetary Science* 53, 5–49. *Meteoritics &*
1373 *Planetary Science* 54, 946–949. Stöffler, D., and Hornemann, U., 1972, Quartz and feldspar
1374 glasses produced by natural and experimental shock: *Meteoritics*, v. 7, p. 371–394.
- 1375 Taylor, F.C., and Dence, M.R., 1969, A probable meteorite origin for Mistastin Lake, Labrador:
1376 *Canadian Journal of Earth Sciences*, v. 6, p. 39–45.
- 1377 Thiel, M. van (Ed.), 1977, Compendium of shock wave data, (Livermore: Lawrence Livermore
1378 Laboratory Report UCRL-50108, 1977), p. 710–712.
- 1379 Tomioka, N., and Miyahara, M., 2017, High-pressure minerals in shocked meteorites:
1380 *Meteoritics and Planetary Science*, v. 52, p. 2017–2039.
- 1381 Tomioka, N., Mori, H., and Fujino, K., 2000, Shock-induced transition of NaAlSi₃O₈ feldspar
1382 into a hollandite structure in a L6 chondrite: *Geophysical Research Letters*, v. 27, p. 3997–
1383 4000.

- 1384 Tomioka et al., 2007, Static amorphization of plagioclase: Comparison to the formation pressure
1385 of diaplectic glass in shock experiments: *Meteoritics and Planetary Science*, v. 42, p.
1386 A148
- 1387 Tomioka, N., Kondo, H., Kunikata, A., and Nagai, T., 2010, Pressure-induced amorphization of
1388 albitic plagioclase in an externally heated diamond anvil cell: *Geophysical Research*
1389 *Letters*, v. 37, p. 1–5, doi:10.1029/2010GL044221.
- 1390 Treiman, A., and Treado, P., 1998, Martian maskelynite? Raman spectra of plagioclase-
1391 composition glasses from ALH84001, EETA 79001, and ALHA 77005, in 29th Lunar and
1392 Planetary Science Conference, p. Abstract 1196.
- 1393 Trepmann, C., Whitehead, J., and Spray, J., 2003, Shock effects in target rocks from the
1394 Charlevoix impact structure, Quebec, Canada: Abstracts of Geological Society of America
1395 Northeastern Section – 38th Annual Meeting, p. 7– 17.
- 1396 Tschermak, G., 1883, Beitrag zur Klassifikation der Meteoriten: Sitzungsberichte der
1397 Mathematisch-Naturwissenschaftlichen Klasse der Kaiserlichen Akademie der
1398 Wissenschaften, Wien, v. 88, p. 347–371.
- 1399 Tschermak, G., 1872, Die Meteoriten von Shergotty und Gopalpur: Sitzungsberichte der
1400 Mathematisch-Naturwissenschaftlichen Klasse der Kaiserlichen Akademie der
1401 Wissenschaften, Wien, v. 65, p. 122–145.
- 1402 Turner, F.J., 1947, Determination of plagioclase with the four-axis universal stage: *American*
1403 *Mineralogist*, v. 32, p. 389–410.
- 1404 Vernon, R.H., 1975, Deformation and recrystallization of a plagioclase grain: *American*
1405 *Mineralogist*, v. 60, p. 884–888.
- 1406 Vinet, N., Flemming, R.L., and Higgins, M.D., 2011, Crystal structure, mosaicity, and strain
1407 analysis of Hawaiian olivines using in situ X-ray diffraction: *American Mineralogist*, v. 96,
1408 p. 486–497.
- 1409 Wackerle, J., 1962, Shock-wave compression of quartz: *Journal of Applied Physics*, v. 33, p.
1410 922, doi:10.1063/1.1777192.
- 1411 Walawender, M.J., 1977, Shock-produced mosaicism in plagioclase, Charlevoix structure,
1412 Quebec: *Canadian Journal of Earth Sciences = Revue Canadienne des Sciences de la Terre*,
1413 v. 14, p. 74–81.
- 1414 Walton, E.L., Sharp, T.G., Hu, J., and Filiberto, J., 2014 Heterogeneous mineral assemblages in
1415 martian meteorite Tissint as a result of a recent small impact event on Mars. *Geochimica et*
1416 *Cosmochimica Acta*. 140: 334-348
- 1417 Walton, E.L., Sharp, T.G., Hu, J., 2016, Frictional melting processes and the generation of shock
1418 veins in terrestrial impact structures: Evidence from the Steen River impact structure,
1419 Alberta, Canada: *Geochimica et Cosmochimica Acta*, v. 180, p. 256-270.

- 1420 White, J.C., 1993, Shock-induced amorphous textures in plagioclase, Manicouagan, Quebec,
1421 Canada: *Contributions to Mineralogy and Petrology*, v. 113, p. 524–532.
- 1422 Whitehead, J., Grieve, R.A.F., and Spray, J.G., 2002, Mineralogy and petrology of melt rocks
1423 from the Popigai impact structure, Siberia: *Meteoritics and Planetary Science*, v. 37, p.
1424 623–647, doi:10.1111/j.1945-5100.2002.tb00844.x.
- 1425 Whitney, P.R., 1972, Spinel inclusions in plagioclase of metagabbros from the Adirondack
1426 Highlands: *American Mineralogist*, v. 57, p. 1429–1436.
- 1427 Williams, Q., and Jeanloz, R., 1988, Spectroscopic evidence for pressure-induced coordination
1428 changes in silicate glasses and melts: *Science*, v. 239, p. 902 LP – 905,
1429 doi:10.1126/science.239.4842.902.
- 1430 Wittmann, A., Korotev, R.L., Jolliff, B.L., Nishizumi, K., Jull, A.J.T., Caffee, M.W., Zanetti, M.,
1431 Irving, A.J., 2019, Petrogenesis of lunar impact melt rock meteorite Oued Awlitis 001:
1432 *Meteoritics and Planetary Science*, v. 54, p. 2167-2188.
- 1433 Xu, C., Zhao, S.R., Li, C., and He, X., 2016, Plagioclase twins in a basalt: An electron
1434 backscatter diffraction study: *Journal of Applied Crystallography*, v. 49, p. 2145–2154,
1435 doi:10.1107/S1600576716015739.
- 1436 Yagi, A., Suzuki, T., and Akaogi, M., 1994, High pressure transitions in the system KAlSi_3O_8 -
1437 $\text{NaAlSi}_3\text{O}_8$: *Physics and Chemistry of Minerals*, v. 21, p. 12–17.
- 1438 Zoltai, T., and Stout, J.H., 1984, *Mineralogy: concepts and principles*: Minneapolis, Burgess
1439 Publishing Company, 505 p.

1440

1441 **10. FIGURE CAPTIONS**

1442 Figure 1: Plot showing the relationship between anorthite content and pressure needed to
1443 form partly to fully isotropic plagioclase. The overall trend demonstrates that plagioclase of
1444 higher mole % An content results in complete amorphization at lower pressures than plagioclase
1445 of lower mole % An content. The graph is based on shock recovery experimental data
1446 summarized by Fritz et al. (2019a).

1447

1448 Figure 2: Shock pressure versus specific volume Hugoniot plot of a representative
1449 plagioclase (An_{-50}) illustrating the three distinct phase regimes (I, low pressure; II, mixed; and

1450 III, high pressure) and the range of the Hugoniot elastic limit (HEL). Data from Marsh (1980)
1451 and van Thiel (1977).

1452

1453 Figure 3: A) Transmitted light photomicrograph of heavily fractured anorthite in Apollo
1454 sample 15415,90; fractures are highlighted by offset twins when viewed between crossed
1455 polarizers. B) Backscattered electron image of fractured plagioclase (albite, dark grey) from the
1456 El'gygytgyn impact structure, with planar features (black arrows) in the lower right corner.

1457

1458 Figure 4: Undulatory extinction in Apollo samples 76335,55 (A) and 60215,13 (B)
1459 between cross-polarized light, as the stage is rotated a wave of extinction passes through the
1460 crystal.

1461

1462 Figure 5: Planar microstructures in oligoclase from the Tenoumer impact structure (A,
1463 B); note the planar features in (B) only exist in alternating twins; after Jaret et al. (2014). C)
1464 Andesine (An_{23}) from the Manicouagan impact structure showing multiple sets of planar
1465 features. D) K-rich feldspar with albite exsolution (beige, labeled Ab) and tartan twins spread
1466 over the whole grain from the Gardnos impact structure; planar features on the right side of the
1467 image (parallel to white lines) with the same orientation as polysynthetic twins.

1468

1469 Figure 6: Plane-polarised photomicrographs of andesine experimentally shocked to 25
1470 GPa (A) and 56 GPa (B) showing progressive degrees of shock-darkening. Electron probe
1471 microanalysis, Raman, and FTIR analyses suggest this darkening is purely optical, and does not
1472 correspond with observable chemical or structural changes. After Jaret et al. (2018).

1473

1474 Figure 7: Diaplectic plagioclase glass. A) Illustration of maskelynite in the Shergotty
1475 martian meteorite, highlighting preserved external crystal morphology and irregular fractures
1476 extending from maskelynite grain margins into surrounding igneous pyroxene (after Tschermak
1477 1872, drawn at 75x magnification, estimated grain size is $\sim 400 \mu\text{m}$). B) Composite reflected light
1478 (left) and transmitted light between crossed polarizers (right) image showing a partly isotropic
1479 plagioclase crystal enclosed by olivine in martian meteorite Chassigny, though the plagioclase is
1480 only partly isotropic, the same fracture pattern is observed as in the illustration by Tschermak
1481 (A). C,D) Partially isotropic plagioclase in PPL (C) and XPL (D) from Apollo 15 sample
1482 15684,4. Black areas in XPL remain extinct on rotation of the stage. E,F) Completely isotropic
1483 plagioclase in basalt from Lonar crater, India (after Jaret et al. 2015). G) Plane-polarised
1484 transmitted light image of maskelynite from the Shergotty meteorite ($\sim 30 \text{ GPa}$) which includes a
1485 perfectly preserved magmatic pyroxene. H) Plane-polarised transmitted light image of
1486 maskelynite from the Los Angeles Martian meteorite ($\sim 45 \text{ GPa}$) with a pyroxene fragment
1487 partially affected by eutectic melting.

1488

1489 Figure 8: Plots demonstrating the relationship between refractive index, pressure, and
1490 composition of diaplectic feldspar glasses. A) Refractive index vs. An content. Displayed are
1491 isobars indicating the shock pressures that result in a plagioclase normative phase with a given
1492 index of refraction. Fields indicating the optical properties, birefringent or isotropic, are
1493 indicated by grey lines; after Fritz et al. (2005b) using data from Stöffler et al. (1986). B)
1494 Refractive indices of diaplectic feldspar glasses from experiments in the pressure range from 28-
1495 45 GPa. After Ostertag (1983).

1496

1497 Figure 9: IR spectra of 2 orientations from (A) crystalline, (B) diaplectic glass, and (C)
1498 melt glass of labradorite composition. Crystalline labradorite exhibits multiple peaks, all
1499 showing shifts with orientation indicating the anisotropic nature of the material. Both melt glass
1500 and diaplectic glass are amorphous (one broad peak) but diaplectic glass retains orientation
1501 effects not observed in the melt glass sample. After Jaret et al. (2015).

1502

1503 Figure 10: Plane-polarised light photomicrograph of plagioclase glass displaying bubbles
1504 and flow textures surrounded by brown stained olivine, in some areas the plagioclase glass has
1505 developed a rim of recrystallized plagioclase. From the strongly shocked ALH (Allan Hills)
1506 77005 martian meteorite.

1507

1508 Figure 11: Photomicrographs of sieve or checkerboard feldspar. A, B) Plagioclase from
1509 Gow Lake impact structure, showing preserved polysynthetic twins between crossed polarizers
1510 (B). C) Plane-polarized light photomicrographs showing magnified view of checkerboard
1511 feldspar in a clast from Gow Lake impact melt rock; D) sieve-textured feldspar from a volcanic
1512 andesite (image credit: Dr Alessandro Mommio).

1513

1514 Figure 12: Cathodoluminescence images and spectra of plagioclase (An_{55}) experimentally
1515 shocked to 28 GPa by Fritz et al., (2019a). A) Transmitted light photomicrograph between
1516 crossed polarizers (XPL) showing heterogeneous distribution of isotropic domains (black). B)
1517 Optical microscope cathodoluminescence (OM-CL) image of the same grain as in (A), showing
1518 three spectra locations (1-3). Areas of higher crystallinity (point 1) have higher luminescence

1519 than isotropic domains (points 2 and 3). C) CL spectra corresponding to points 1-3 in (B), further
1520 showing that the domains with remaining crystallinity have higher luminescence than those that
1521 are isotropic, after Pittarello et al. (2020b).

1522

1523 Figure 13: Raman spectra of shocked feldspar. A) Raman spectra of andesine (An_{43}):
1524 unshocked (0 GPa), shocked (up to 29 GPa), and diaplectic glass (>29 GPa). For clarity, three
1525 spectra of 0 GPa (top line), 26 GPa (middle line), and 56 GPa (bottom line), representing the
1526 aforementioned categories (unshocked, shocked, diaplectic glass) are shown here. Up to
1527 pressures of 29 GPa (middle line), increasing shock corresponds to an overall decrease in peak
1528 intensity and causes the $480/510 \Delta \text{cm}^{-1}$ peak ratio to approach 1. Above 29 GPa (bottom line),
1529 the sample appears amorphous, and exhibits a different Raman pattern, with two broad peaks at
1530 $\sim 500 \Delta \text{cm}^{-1}$ and $580 \Delta \text{cm}^{-1}$. Data from Jaret et al. (2018). B) Raman spectra of diaplectic
1531 feldspar glass with three different An contents (top: An_{15} – naturally shocked L6 chondrite;
1532 middle: An_{50} – experimentally shocked troctolite; bottom: An_{94} – experimentally shocked
1533 gabbro). The spectral position of the hump in the region of 950 to 1150 cm^{-1} varies with An
1534 content. Data from Fritz et al. (2019a).

1535

Table 1: Summary table of impact-related metamorphism in feldspar group minerals, with key to the section of the paper that discusses them in more detail (Column 1: “Sctn.”) and relevant figures (Fig.). Reported pressures are in GPa and are derived from ^[1]Ostertag (1983) and ^[2]Jaret et al. (2018), both of which conducted shock recovery (gas gun) experiments. These pressures should act as guides and are indicative of trends in the order of formation of shock features, however the exact values may not be directly applicable to natural samples for reasons discussed in Section 4.

Sctn.	Feature	Fig.	Orthoclase	Sanidine	Microcline	Albite	Oligoclase	Andesine	Labradorite	Bytownite			
			Single crystal ^[1]	Single crystal ^[1]	Single crystal ^[1]	An ₀₂ , rock ^[2]	Single crystal ^[1]	An ₄₃ , rock ^[2]	Single crystal ^[1]	An ₇₇ ^b , rock ^[1] , [2]			
3.1	Irregular fracturing ^a	3	<10.5 to 32	<10.5 to 30	<10.5 to 34	<i>N.R.</i>	<10.5 to 29	<i>N.R.</i>	10.5 to 27	10.5 to 26 ^a			
3.2	Undulatory extinction Wave of extinction passes through the grain	4	<i>No limits explicitly stated</i>										
3.3	Mosaicism Patchwork of slightly different extinction angles within a grain	<i>b</i>	10.5 to 28	14 to 30	14 to 45	<i>Ambiguous</i>	~18 to 31	<i>Ambiguous</i>	18 to 28	22 to 27 ^[1]			
3.4	Planar features/microstructures Open parallel planar fractures (PFs) and/or Closely spaced parallel planar lamellae (PDFs)	5	~10.5 to 36	10.5 to 31	~10.5 to 42	<i>N.O.</i>	~14 to 33	<i>N.O.</i>	14 to 23	14 to 27 ^[1] <i>N.O.</i> ^[2]			
3.5	Shock darkening Darkening of mineral in plane-polarised transmitted light	6	<i>N.R.</i>	<i>N.R.</i>	<i>N.R.</i>	>~24	<i>N.R.</i>	>~24	<i>N.R.</i>	>~22 ^[1]			
3.7	Diaplectic glass Amorphous phase that maintains external crystal form and internal features such as chemical zoning	7		Partial		>26-28	>26-28	>26-28	~50-56	~26-28	~28-30	>26-28	~25-27 ^[2]
				Complete		>32-34	>32-34	>45 ^c	>~55 ^d	>32-34	>~47 ^d	>28-30	>28-30 ^[1] ; >~38 ^[2] _d
3.8	Melt glass Amorphous phase that shows flow textures and/or vesiculation	10	45-50 ^e	45-50 ^e	45-50 ^e	<i>N.O.</i>	>42 ^e	<i>N.O.</i>	>42 ^e	>45-50 ^[1] _e ; <i>N.O.</i> ^[2]			
3.9	High pressure phases Phases related to feldspar composition that are stable at higher pressure conditions than found on Earth’s surface		<i>See Table 2</i>										
3.10	Sieve and checkerboard feldspar Feldspar clasts in melt rocks made up of individual rectangular domains separated by melt products	11	<i>Temperature rather than pressure is the controlling factor</i>										
3.6	Alternate twin deformation Special cases of planar features or diaplectic glass confined to only alternate twins	5a,b	<i>This is a broad category of effects that each seem to occur over a narrow range of pressure conditions, somewhat lower than the equivalent full grain effect. e.g., alternate diaplectic glass twins fall into the category above of “partial isotropization”.</i>										

N. R. = Not reported – features were not explicitly discussed in the manuscript so presence in those samples is ambiguous. **N.O.** = None observed - explicitly stated that the author looked but did not find that feature in the sample. **Ambiguous** = Jaret et al. (2018) found a texture reminiscent of mosaicism, but were ambiguous in their interpretation, so did not explicitly report pressure conditions for mosaicism in those samples. ^aOstertag (1983) found that fracturing decreased over 22 GPa. ^bNote: no images of mosaicism are included because it is challenging to accurately capture that feature in a still image, and no high quality images were found. ^cOstertag (1983) found that microcline maintained reduced birefringence up to 45 GPa, where their experiments stopped. ^dNote: Different pressures for the amorphisation of plagioclase are reported in the literature, the table shows data from Ostertag (1983) and Jaret et al. (2018) and Figure 1 uses the data compiled by Fritz et al. (2019a), as a result the values reported in this table for complete amorphisation from Jaret et al. (2018) are higher (~55 GPa, ~47 GPa, ~38 GPa) than those shown in

Figure 1 (<35 GPa), likely due to differences in experimental setup and/or plagioclase composition. °Ostertag (1983) did not detect melt glass unambiguously, so assumed values to be in accordance with Stöffler and Hornemann (1972).

Table 2: High-pressure phases commonly associated with feldspathic compositions in highly shocked lithologies, including name, chemical composition, type locality, and the experimentally derived stability fields. The pressure [P] stability field of high-pressure phases strongly depends on chemical composition of the environment, and the prevailing temperatures [T], so the presented P-T conditions serve only as a guide and further details can be found in the cited references.

Name	Formula	Type Locality	Synthesized	Meteorites	Terrestrial impact structures	P-T	Refs.
Donwilhelmsite	CaAl ₄ Si ₂ O ₁₁	Lunar meteorite Oued Awlitis 001	Yes	Yes	No	13 – 36 GPa, >1200 °C	1-3
Jadeite	NaAlSi ₂ O ₆	Myanmar		Yes	Yes	2.5 GPa; 1000 °C	4-9
Liebermannite (K-Hollandite)	KAlSi ₃ O ₈	Martian meteorite Zagami	Yes			12 GPa, 1026 °C	10-16
Lingunite (Na-Hollandite)	NaAlSi ₃ O ₈	Sixiangkou L6 chondrite	Yes	Yes	Yes	Stable at 22-23 GPa; >1200 °C Metastable at ~20-24 GPa; ~1100-1300 °C	11-15, 17-22
Tissintite	(Ca,Na,□)AlSi ₂ O ₆	Martian meteorite Tissint	Yes			6 - 8.5 GPa; 1000-1350 °C	23, 24
Zagamiite	CaAl ₂ Si _{3.5} O ₁₁	Martian meteorite Zagami		Yes			23-26

[1] Fritz et al. (2019b), [2] Irifune et al. (1994), [3] Beck et al. (2004), [4] James (1969), [5] Ohtani et al. (2004), [6] Tomioka et al. (2007), [7] Knight and Price (2008), [8] Kubo et al. (2010), [9] Stähle et al. (2011), [10] Ringwood et al. (1967), [11] El Goresy et al. (2000), [12] Gillet et al. (2000), [13] Langenhorst and Poirier (2000), [14] Tomioka et al. (2000), [15] Langenhorst and Dressler (2003), [16] Ma et al. (2018), [17] Liu (1978), [18] Yagi et al. (1994), [19] Liu (2006), [20] Liu and El Goresy (2007), [21] Agarwal et al. (2016), [22] Kubo et al. (2017), [23] Ma et al. (2015), [24] Rucks et al. (2018), [25] Walton et al. (2014), [26] Ma et al. (2017).

Table 3: Summary of different experimental techniques for calibrating shock features and parameters.

Technique	Timescale	Method	Purpose	References
Shock	Nanoseconds	In situ measurements of shockwaves	Determination of the equation of state (EoS) using Rankine-Hugoniot relations (Fig. 2)	Wackerle (1962)
Shock recovery	Shock pulses of microsecond duration	Shock impedance or reverberation technique	Investigations of shock effects in materials	Müller and Hornemann (1969); Hörz and Quaide (1973); Fritz et al. (2011)
Rapid compression	Seconds – 10s mins	Membrane and dynamic diamond anvil cell	Determination of phase stability Typically collect simultaneous XRD during compression and decompression	Letoullec et al. (1988); Evans et al. (2007); Sims et al. (2019)
Static compression	Minutes to hours	Diamond anvil cell	Determination of phase stability Typically coupled with either XRD or Raman spectroscopy after each compression step	Angel (1988); Williams and Jeanloz (1988); Daniel et al. (1997); Sims et al. (2020)

Table 4: Summary of analytical techniques used for assessing shock in feldspars. Optical microscopy is not included as it is assumed that primary characterization has already been conducted.

Technique	Sample requirements	General technique use	Limitations	Application to studies of shock metamorphism	Refs.
Universal-stage (U-stage)	Thin section	Measuring 3D orientation of planar and linear elements (i.e., twins, PDFs, optic axes)	Complex operation for biaxial solid solution minerals like feldspars (see Section 5.1) Only possible to view about 2/3 of a rectangular thin section U-stage does not fit on most modern microscopes	Indexing planar microstructures	(1-4)
Electron Backscattered Diffraction (EBSD)	Highly polished surface (i.e., thin or thick section, polished mount) Crystalline material Sample must fit in SEM	Analysing the distribution of orientations within or between minerals Measuring deformation within a crystal Identifying phases	Cannot uniquely index planar features without the addition of U-stage measurements or FIB cross-sectioning (see Section 5.1) Only applicable to crystalline material (no patterns provided by amorphous material)	Indexing planar microstructures (Section 5.1, complementary to other techniques) Increasing internal misorientation with increasing shock pressure No diffraction from diaplectic and melt glasses	(5-8)
X-ray diffraction (XRD)	Powder Single-crystal In situ μ xRD	Powder (whole rock or mineral separates) Individual mineral separate Thin sections Polished mount Hand sample	Identifying phases Quantifying strain Refining crystal structure Identifying lattice breakdown (e.g., evaluating degree of shock)	Cannot distinguish amorphous phases from each other Peaks broaden in chi direction (single-crystal or in situ μ XRD only) at relatively low pressure Peaks broaden in the 2 θ direction at higher pressure No diffraction from diaplectic and melt glasses (they produce a broad diffuse band)	(9-12)
High-energy total X-ray scattering	Powdered mineral (~mg) Single crystal	Quantifying atomic distances and positions (e.g., evaluating degree of ordering)	Requires synchrotron beamtime	Diaplectic glasses show intermediate order but no long-range order and orientation effects Melt glasses show no ordering >10 Å and no orientation effects	(13-14)
Cathodoluminescence (CL)	Thin section Thick section Polished mount	Imaging microstructures and defects	Luminescence is controlled by multiple variables, such as the abundance of activators (trace elements, point defects, both), and the abundance of quenchers	Plagioclase: Normal band emission at ~550-570 nm commonly shifted to ~630 nm with increasing pressure Diaplectic glass results in a broad band ~ 350 nm Increasing amorphisation causes decreasing luminescence Alkali-feldspar: Increasing pressure results in increasing luminescence	(15-31)
Raman spectroscopy	Thin section Thick section Polished mount	Identifying phases (i.e., high-pressure polymorphs; see Section 3.9) Evaluating crystallinity	Libraries and databases for mineral identification are limited Difficult to assign individual peaks to specific modes	Bands progressively broaden and decrease in intensity resulting in increased FWHM with increasing pressure Diaplectic and melt glasses result in broad peaks about 500 and 1000 cm^{-1} (dependent on composition, see Fig. 14)	(31-36)
Photoluminescence spectroscopy	Thin section Thick section Polished mount	Revealing lattice defects	Need to have endmembers for quantitative damage calibration (unshocked to completely amorphous)	Nd^{3+} luminescence bands broaden with increasing amorphization	(27)

Thermal infrared absorption spectroscopy	Powder (whole rock or mineral separates) (mgs – 10 mgs)	Identifying phases Measuring crystal lattice structure	Averages grain orientations, so no orientation effects are detectable Limited sensitivity to minor components of the mixture	Decreased spectral detail and intensity with increasing pressure Loss of low wavenumber peaks (as a result of lattice disordering and increased glass content) with increasing pressure	(14, 35, 37–44)
Micro FTIR spectroscopy	Thin section Thick section Polished mount		Orientation sensitive measurements (e.g., peak height) can be hard to match to databases of non-oriented measurements	Diaplectic and melt glass result in one broad peak Decreased intensity of peaks due to loss of crystallinity with increasing pressure Diaplectic glass preserves orientation effects Melt glass does not preserve orientation effects	(14, 35)
Nuclear magnetic resonance (NMR) spectroscopy	Powder (whole rock or mineral separates) (mgs – 10 mgs)	Identifying phases Determining local chemical environment	Sample must include an NMR active isotope (e.g., ²⁹ Si in feldspars)	Diaplectic glass lacks long range order, results in a spectral shape similar to melt glass, but peak centre and FWHM are more similar to crystalline plagioclase Melt glass lacks long range order, and results in broader spectral peak than crystalline plagioclase or diaplectic plagioclase glass	(14, 45)

[1] Reinhard (1931), [2] Turner (1947), [3] Stöffler (1967), [4] Dworak (1969), [5] Prior et al. (1999), [6] Pickersgill et al. (2017), [7] Pickersgill and Lee (2015), [8] Pittarello et al. (2020a), [9] Hörz and Quaide (1973), [10] Flemming (2007), [11] Vinet et al. (2011), [12] Pickersgill et al. (2015a), [13] Reeder and Michel (2013), [14] Jaret et al. (2015), [15] Lee et al. (2007), [16] Parsons et al. (2008), [17] Parsons and Lee (2009), [18] Finch and Klein (1999), [19] Götze et al. (2000), [20] Parsons et al. (2015), [21] Kaus and Bischoff (2000), [22] Götte (2009), [23] Gucsik et al. (2004), [24] Götze (2009), [25] Pittarello et al. (2015), [26] Kayama et al. (2018), [27] Pittarello et al. (2020b), [28] Kayama et al. (2012), [29] Kayama et al. (2009a), [30] Kayama et al. (2009b), [31] Kayama et al. (2009c), [32] Fritz et al. (2005a), [33] Fritz et al. (2019a), [34] Jaret et al. (2014), [35] Jaret et al. (2018), [36] Reynard et al. (1999), [37] Lyon (1963), [38] Bunch et al. (1967), [39] Stöffler and Hornemann (1972), [40] Stöffler (1974), [41] Arndt et al. (1982), [42] Ostertag (1983), [43] Johnson et al. (2002a), [44] Johnson (2003), [45] Myers et al. (1998).

Table 5: Summary table of orientations of shock-induced planar microstructures in feldspars measured by U-stage.

Feldspar composition	Reported orientation(s)	Crater	References
Labradorite	(001), (010), (111), (203), (101)	Manicouagan	Dworak (1969)
Andesine (\sim An ₃₀)	(001) 25%, (010) 11%, (100) 10%, ($\bar{1}\bar{2}0$) 10%, (012) 7%, (130) 6%, others occurring at less than 2% frequency	Ries	Stöffler (1967)
Plagioclase	(001), (010), ($\bar{1}02$), and ($\bar{1}\bar{2}1$)	Several Canadian craters	Robertson et al. (1968)
An ₉₈	(101), (010), (021)	Apollo breccia 10046	Dence et al. (1970)

Figure 1

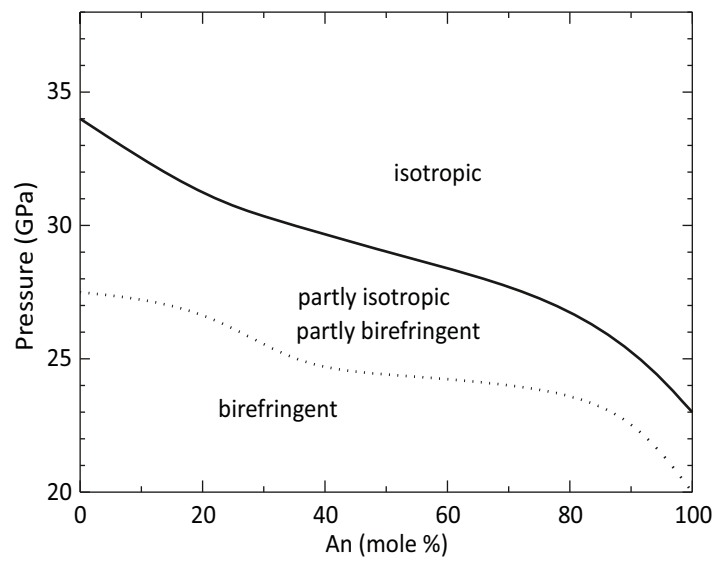


Figure 2

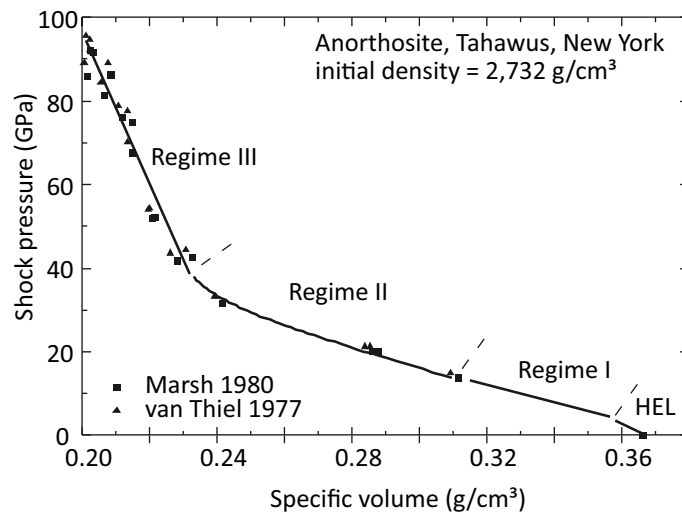
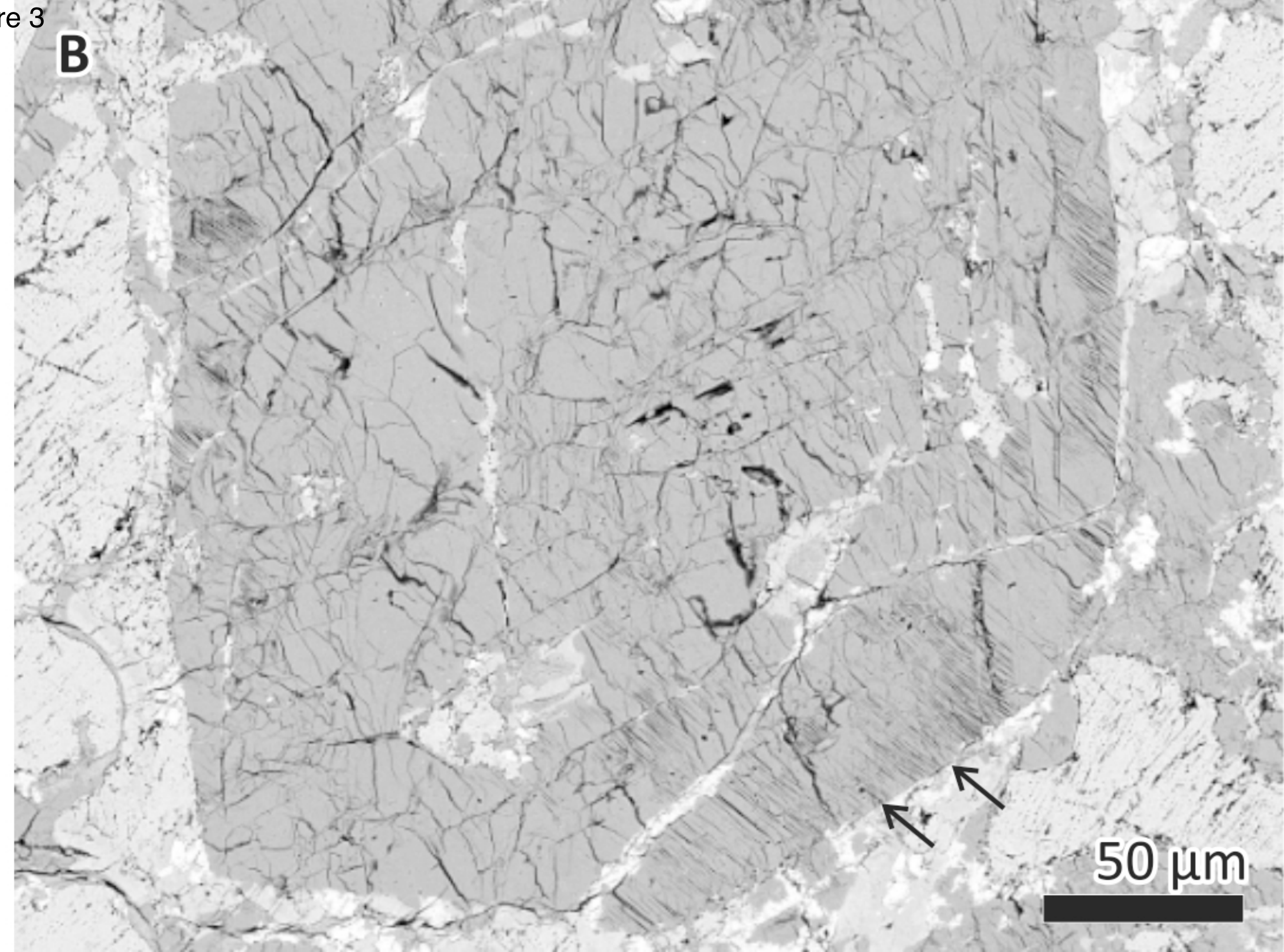
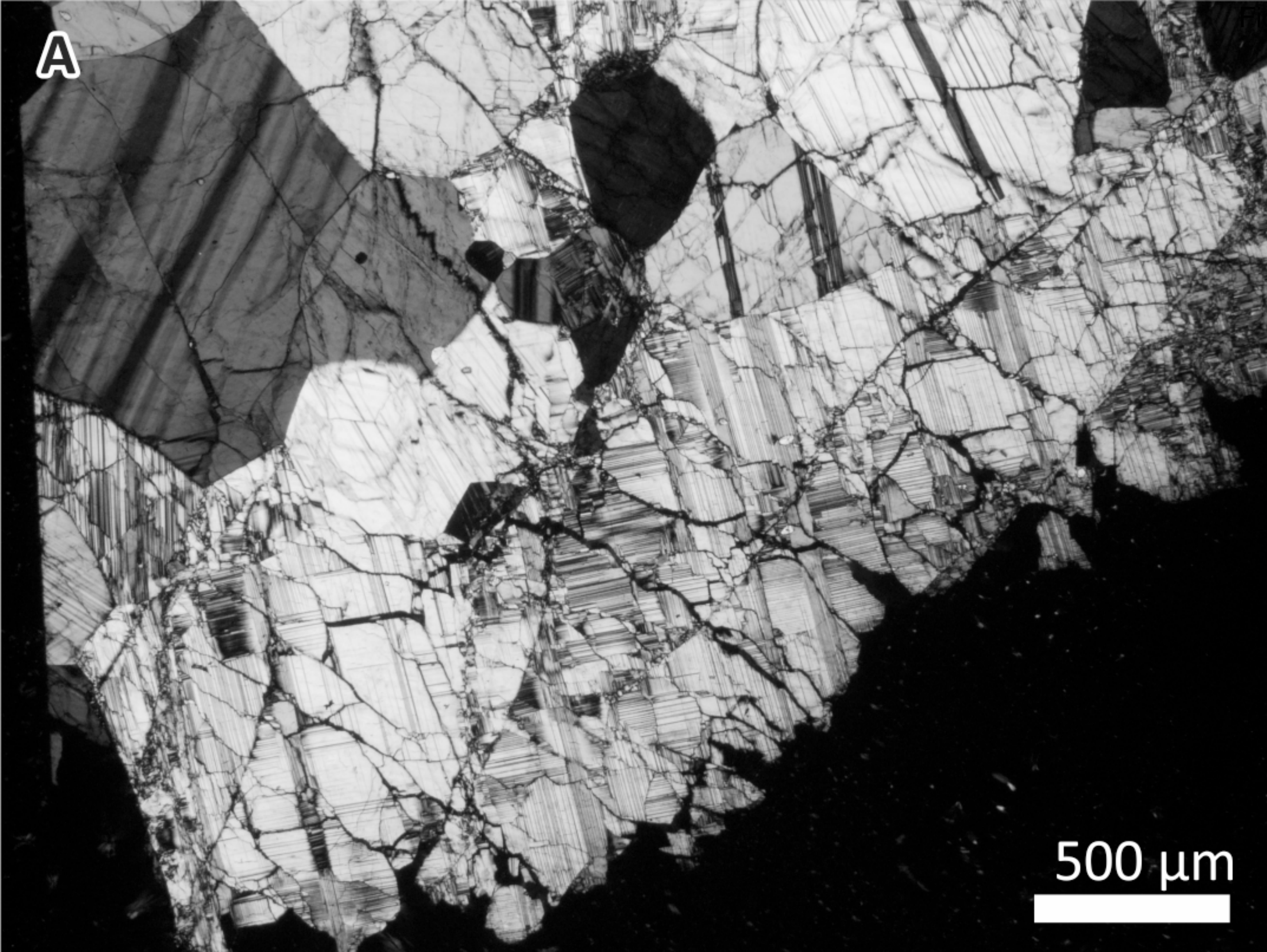


Figure 3



A

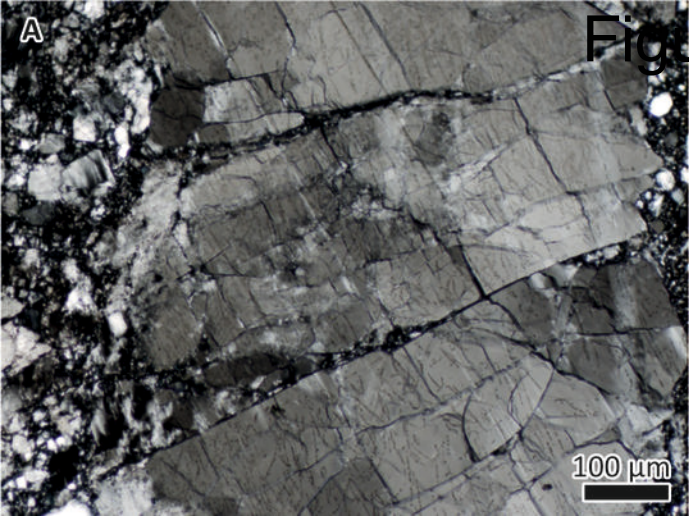


Figure 4

B

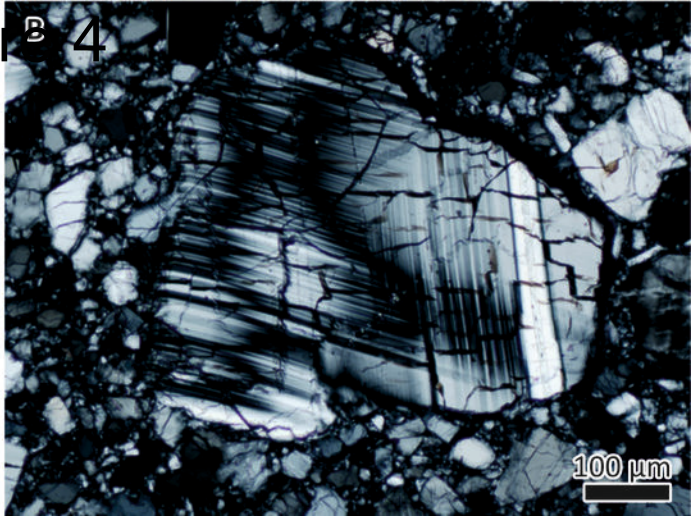


Figure 5

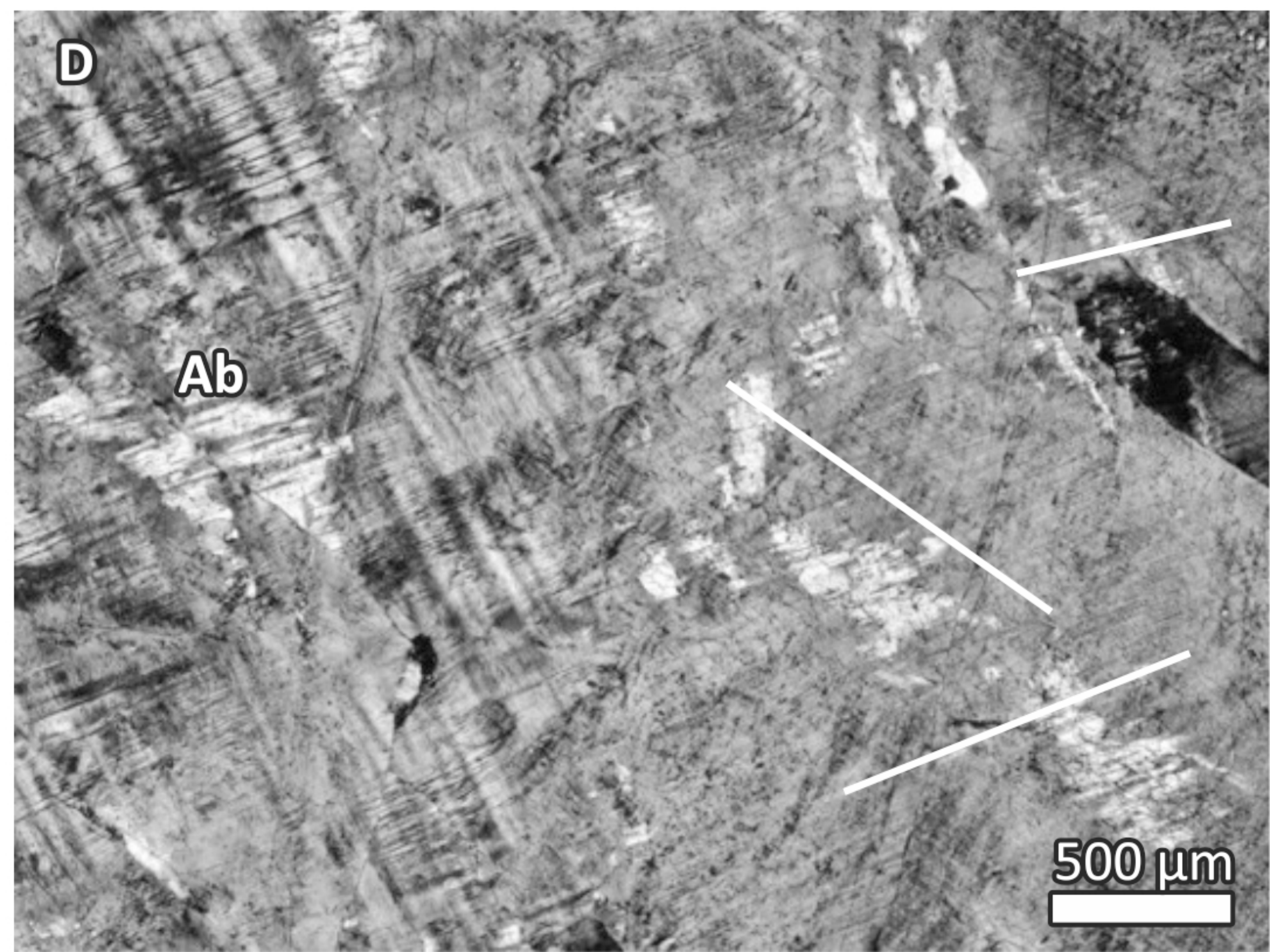
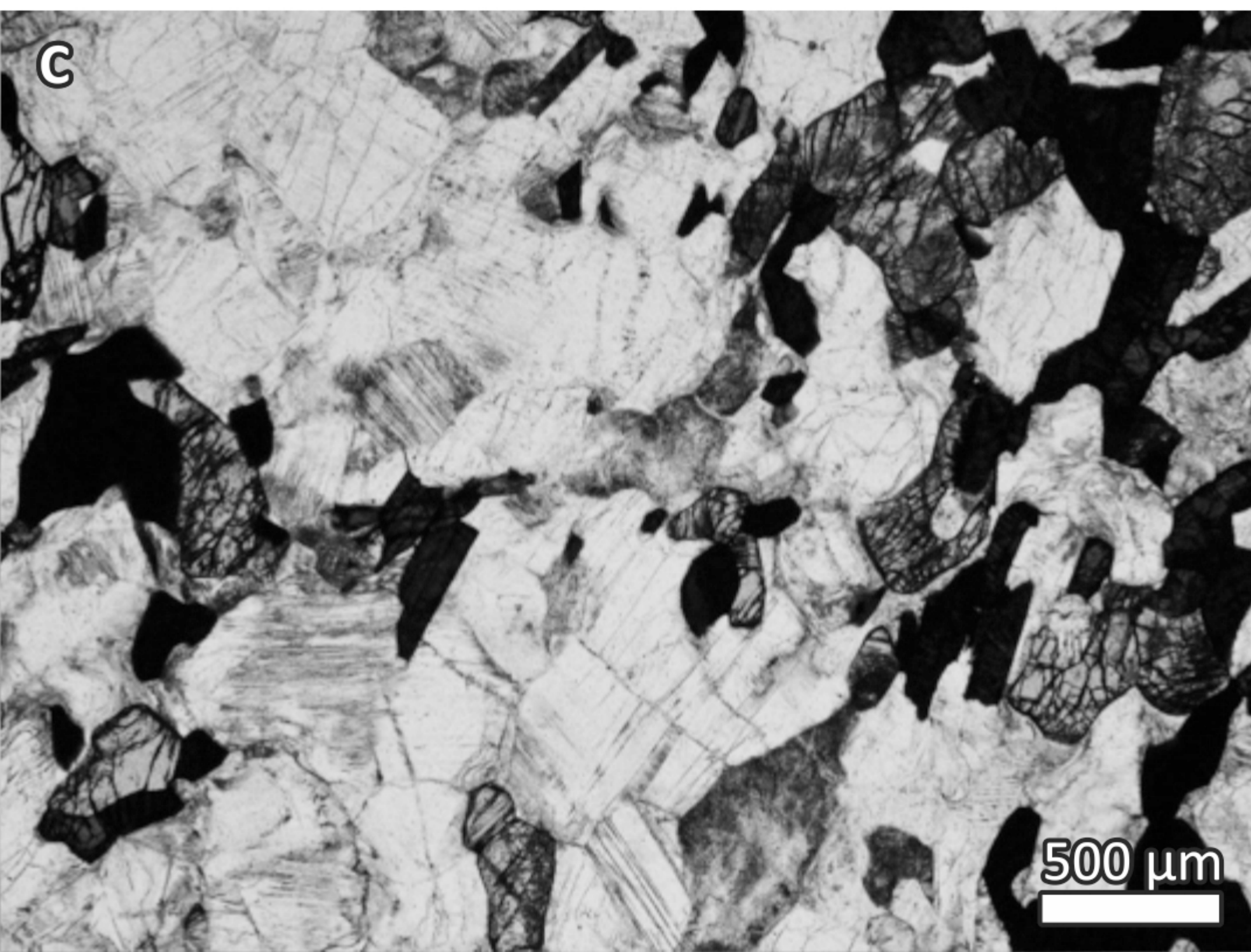
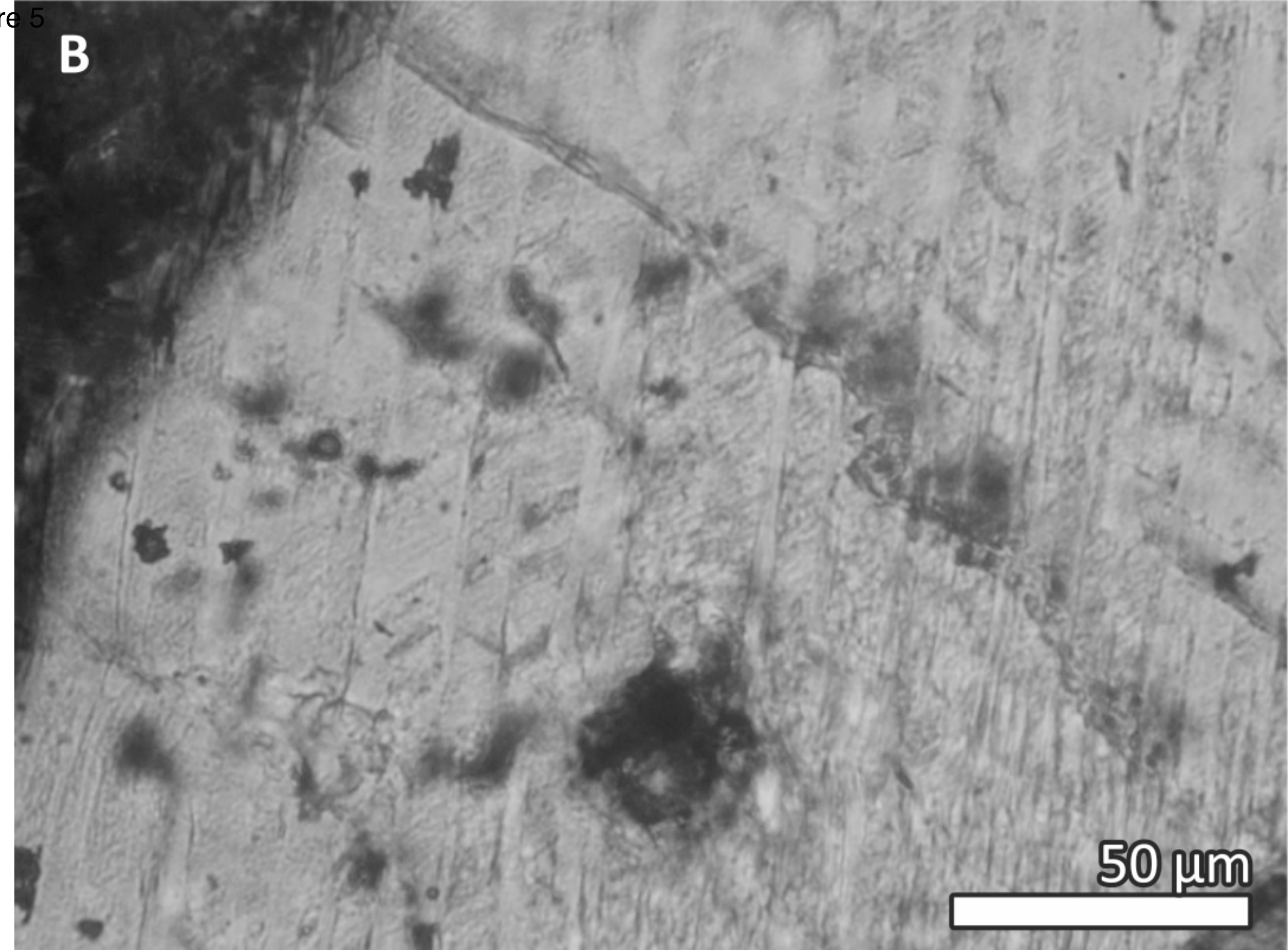
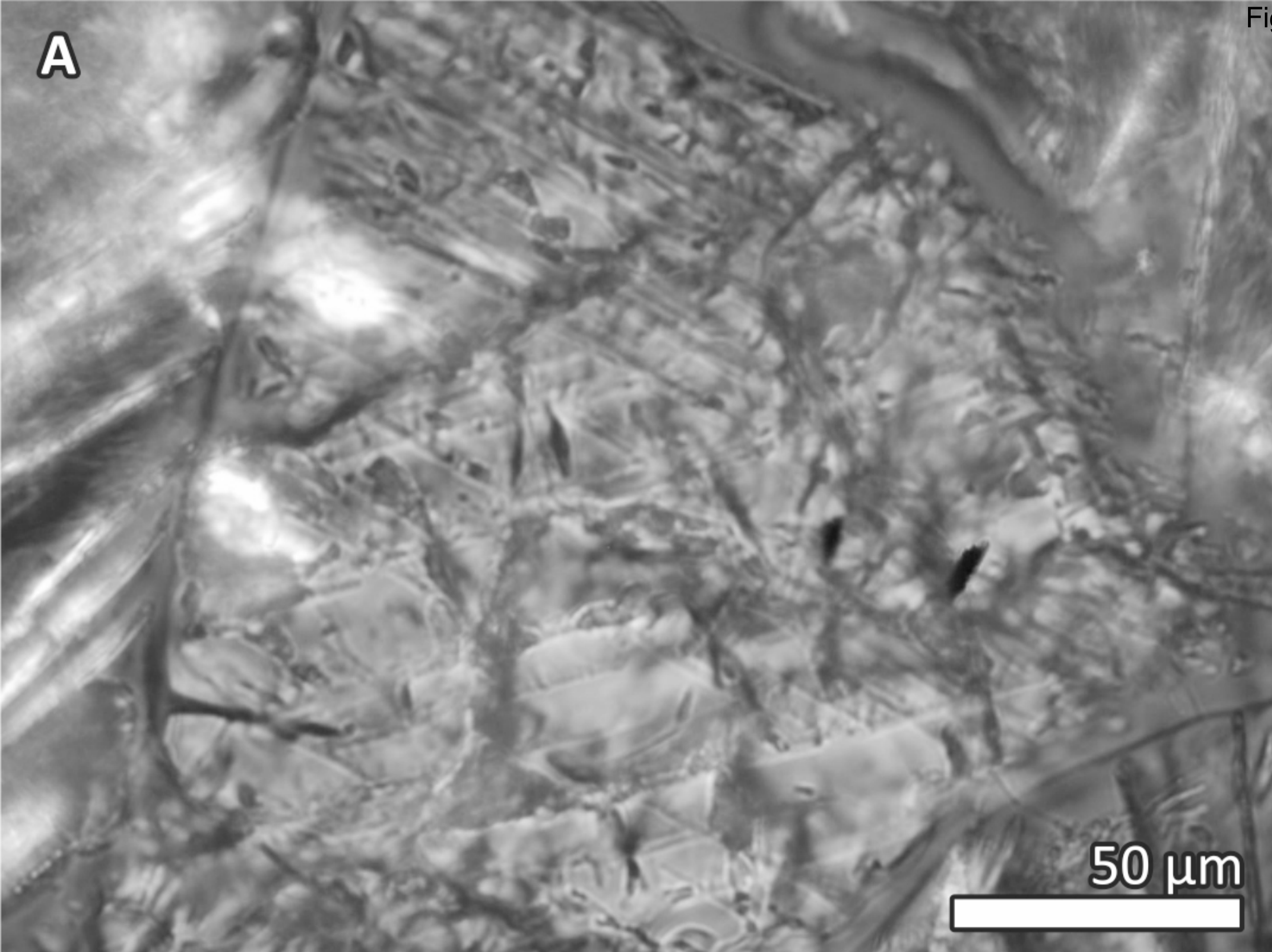


Figure 6

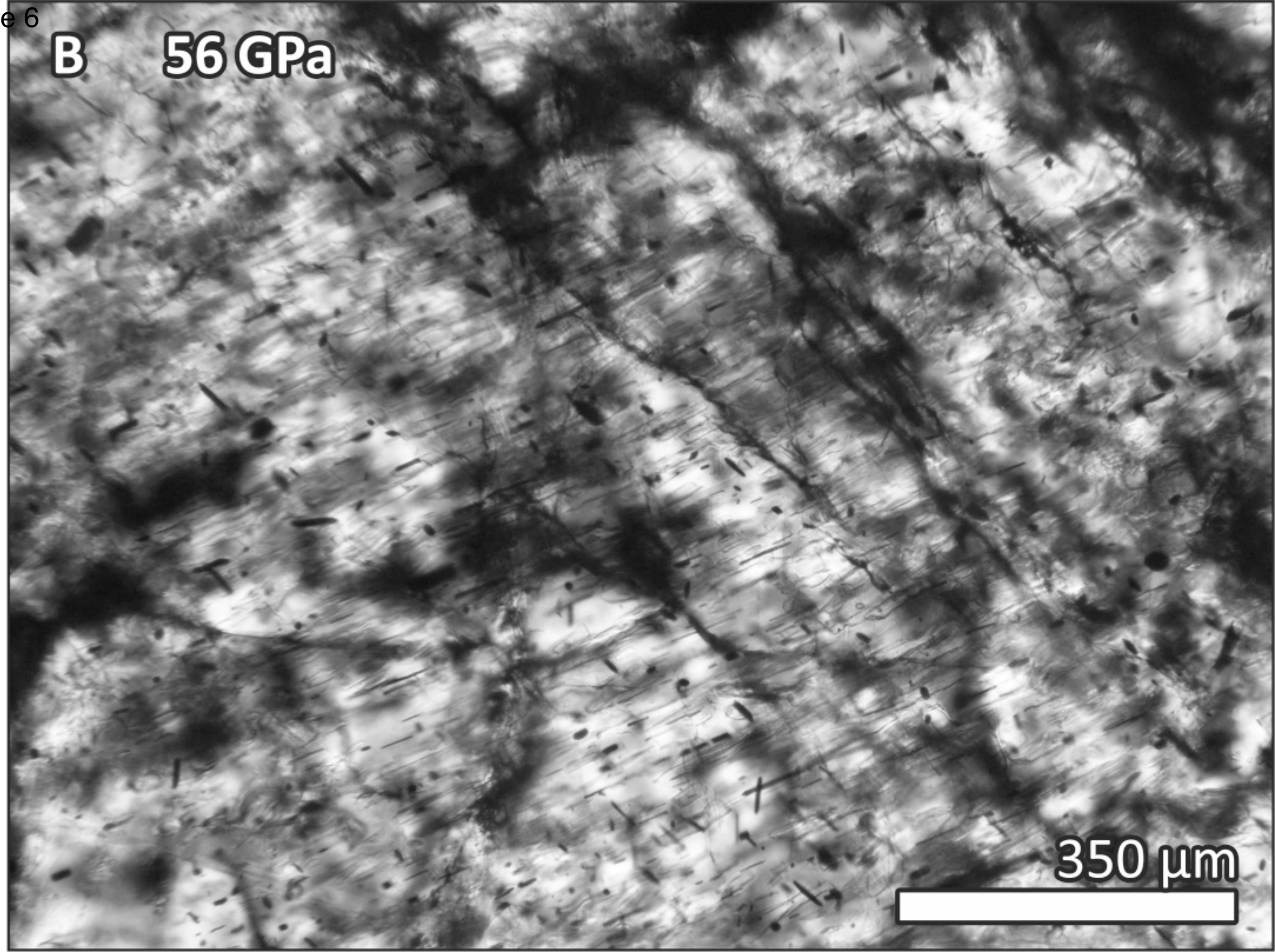
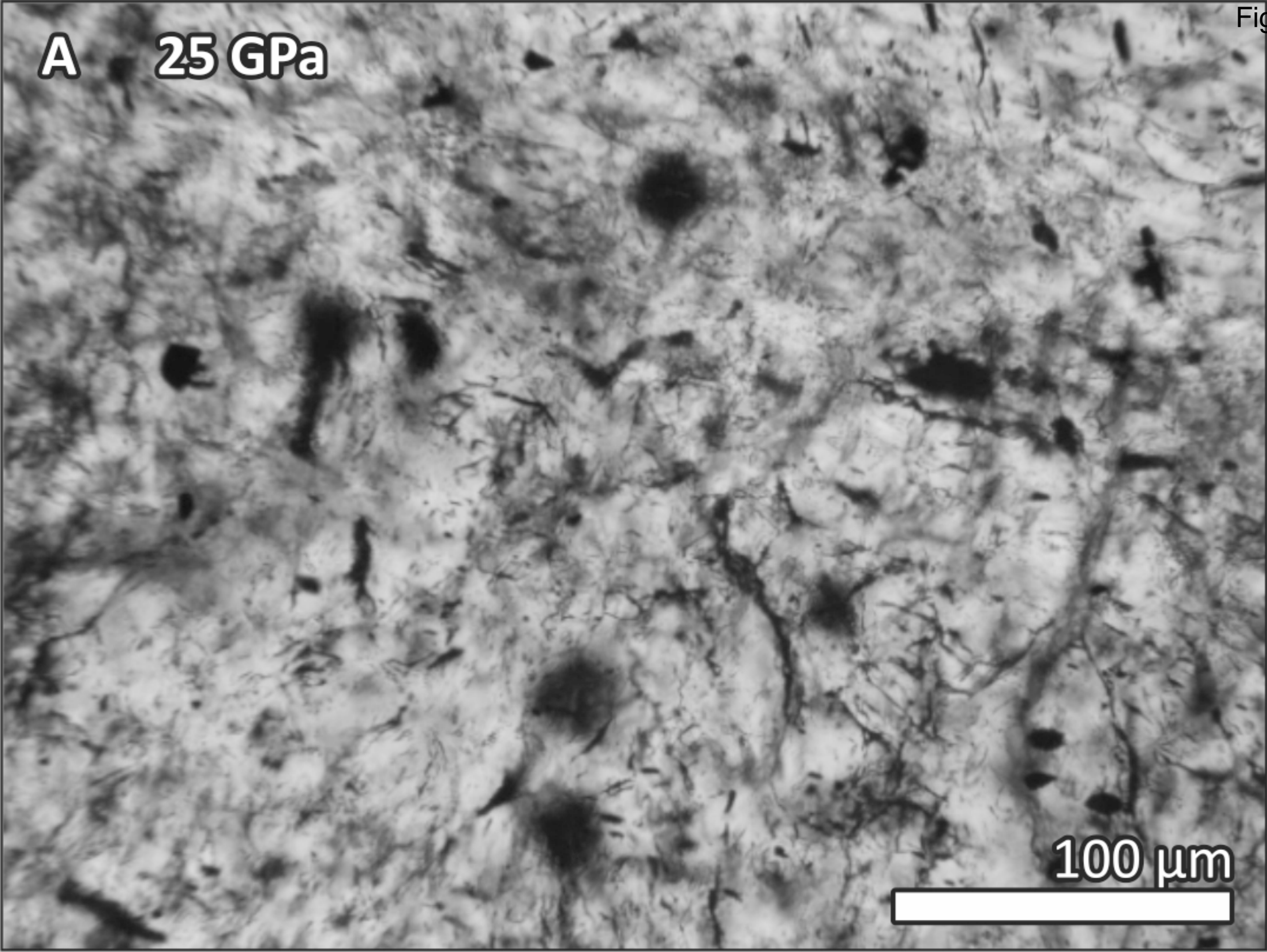


Figure 7

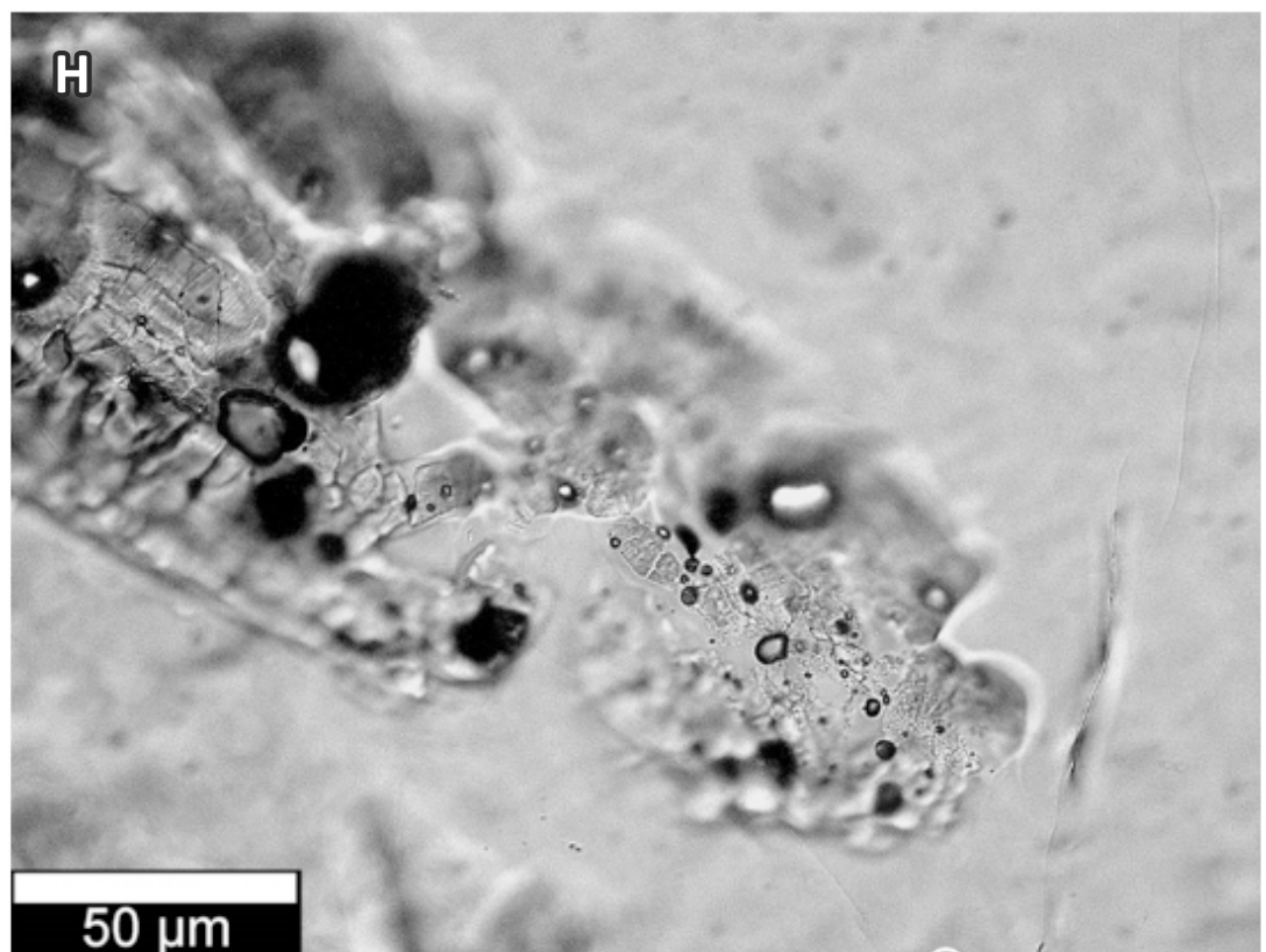
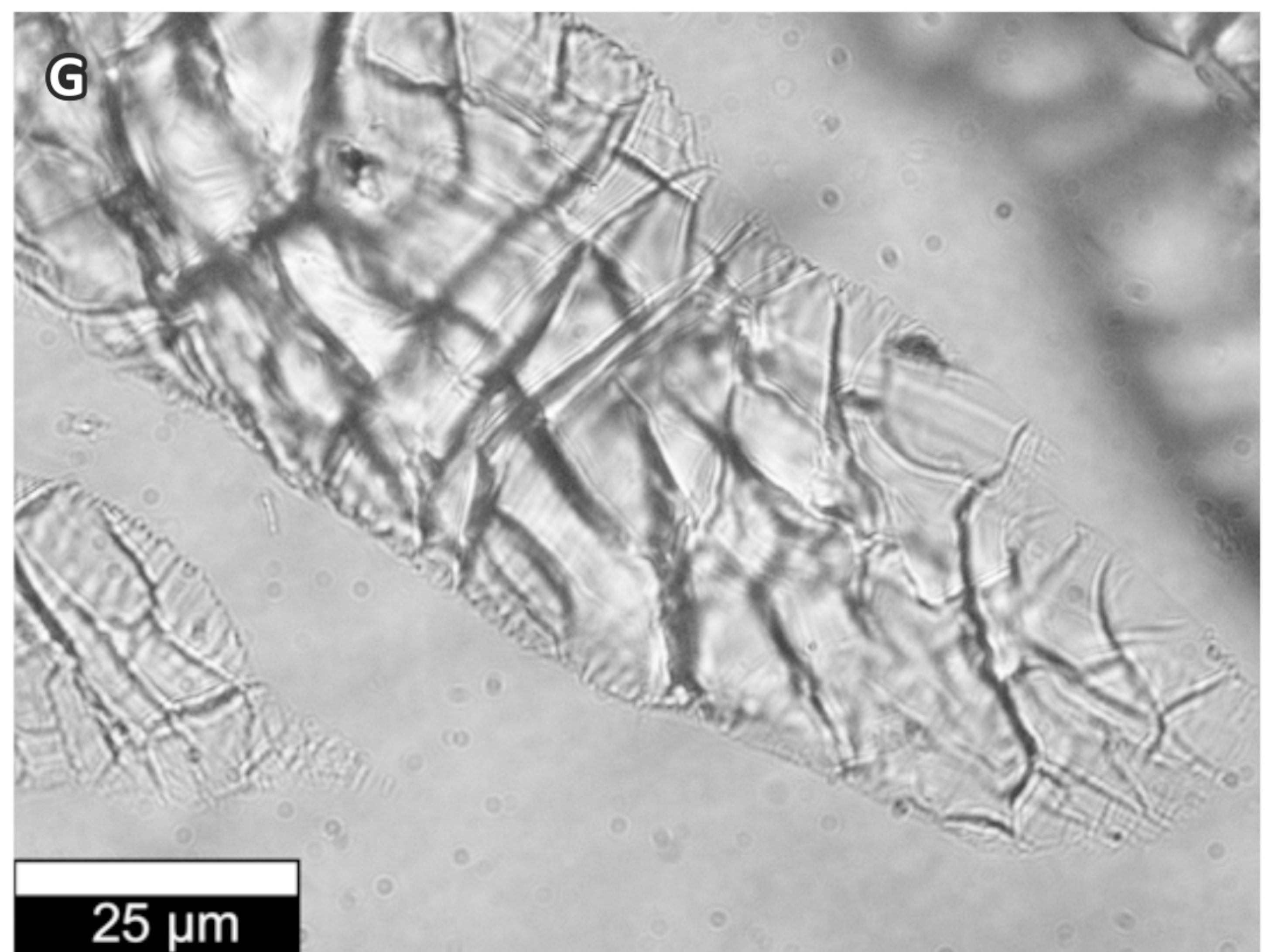
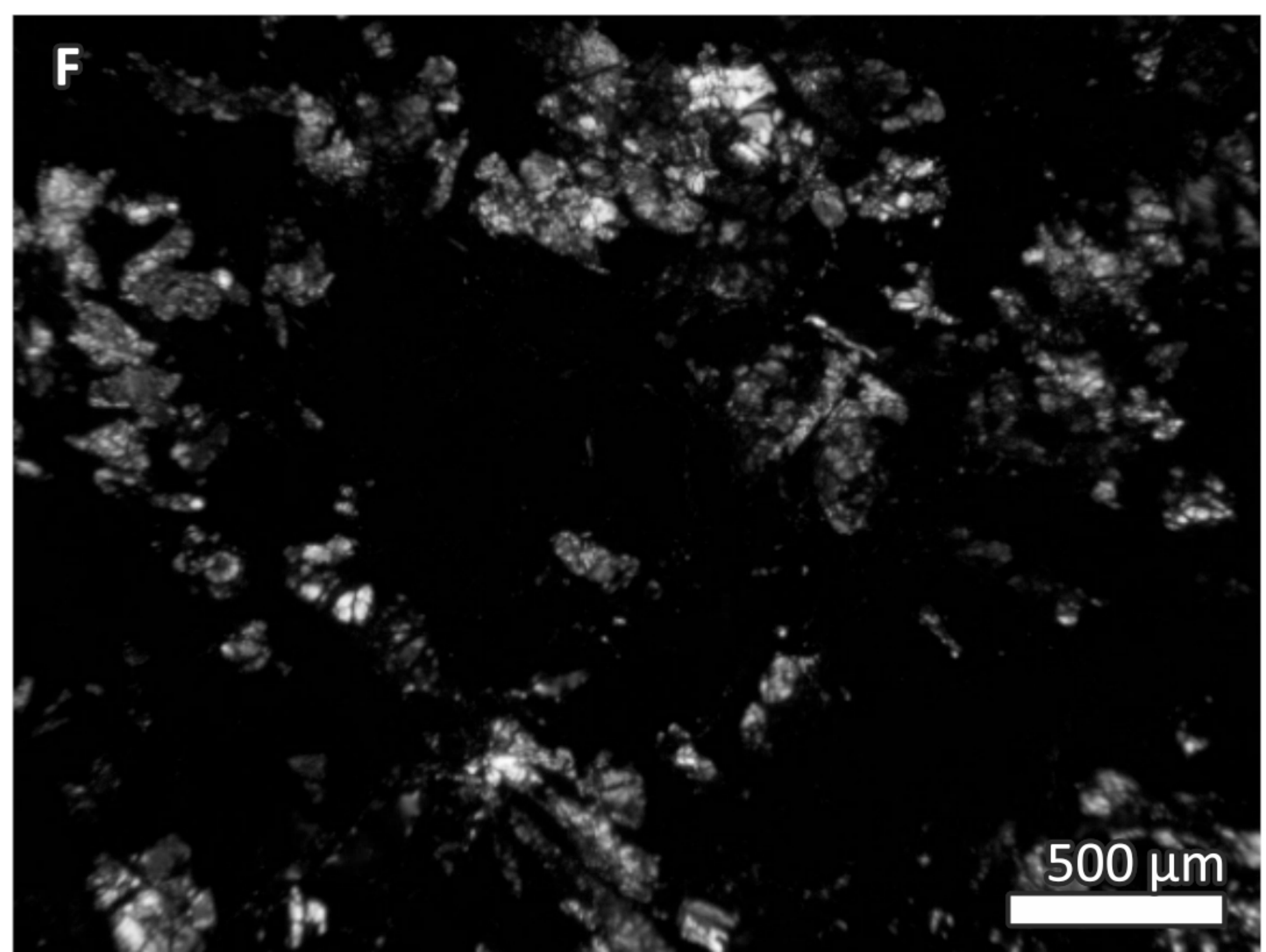
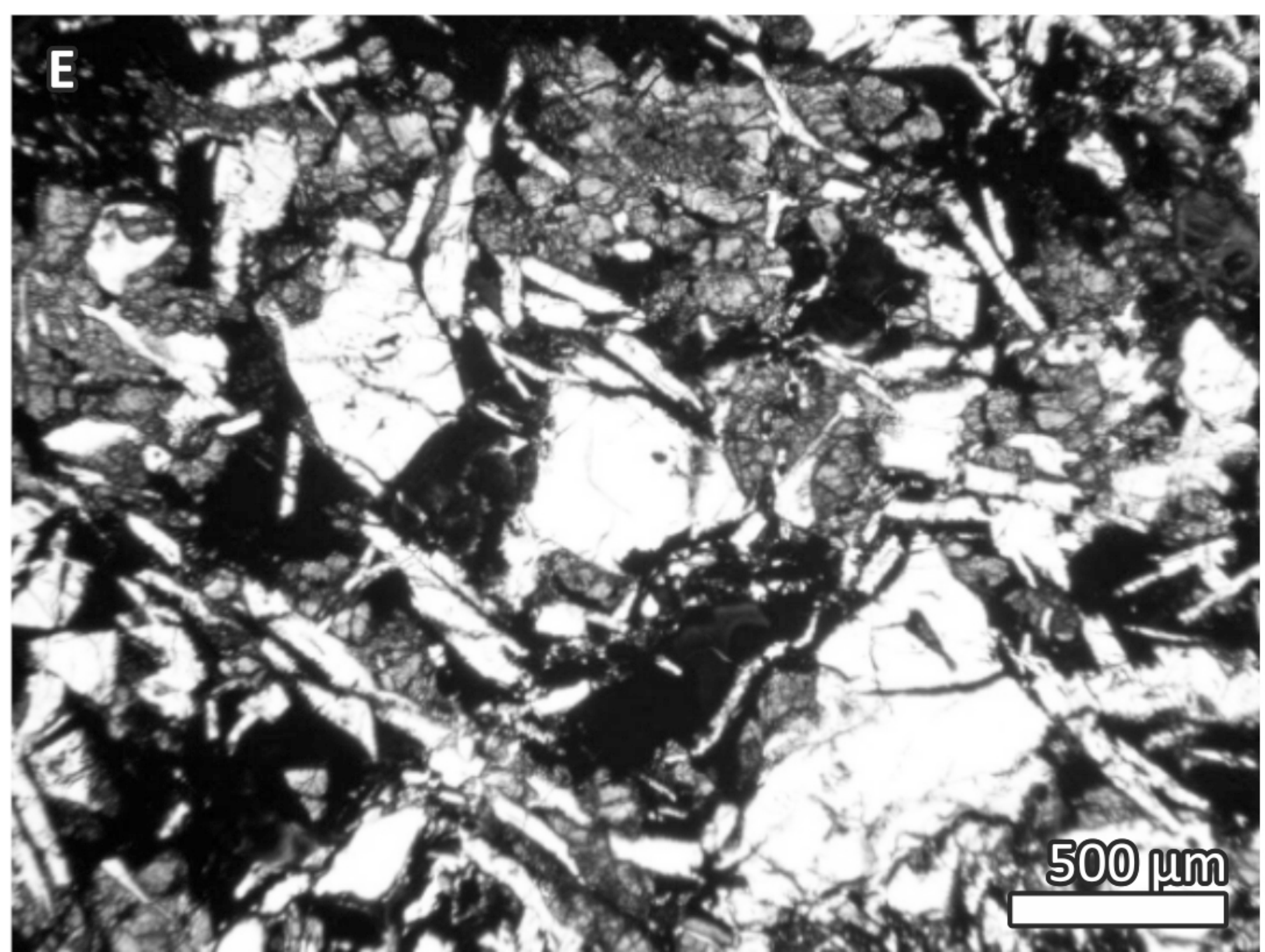
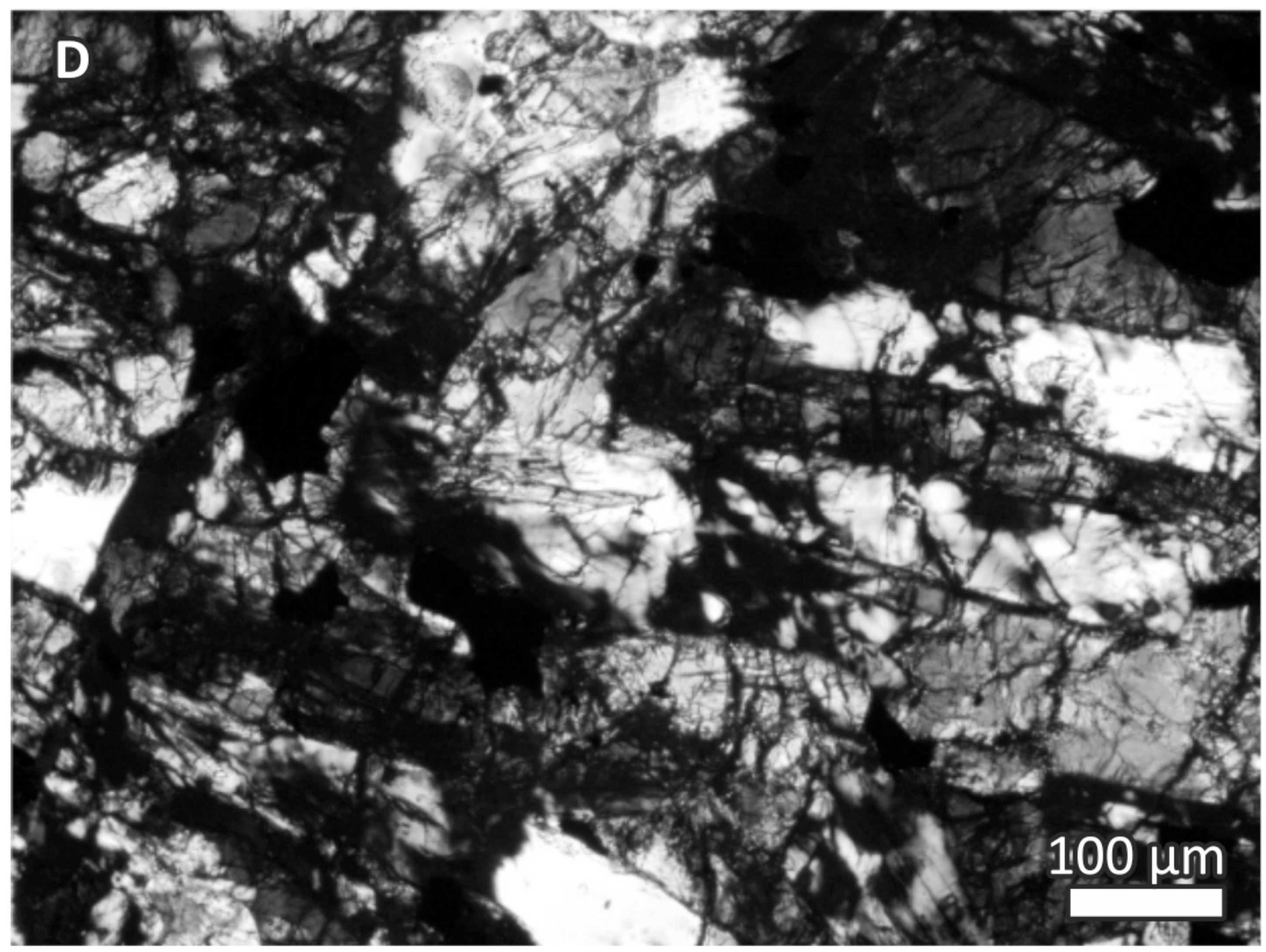
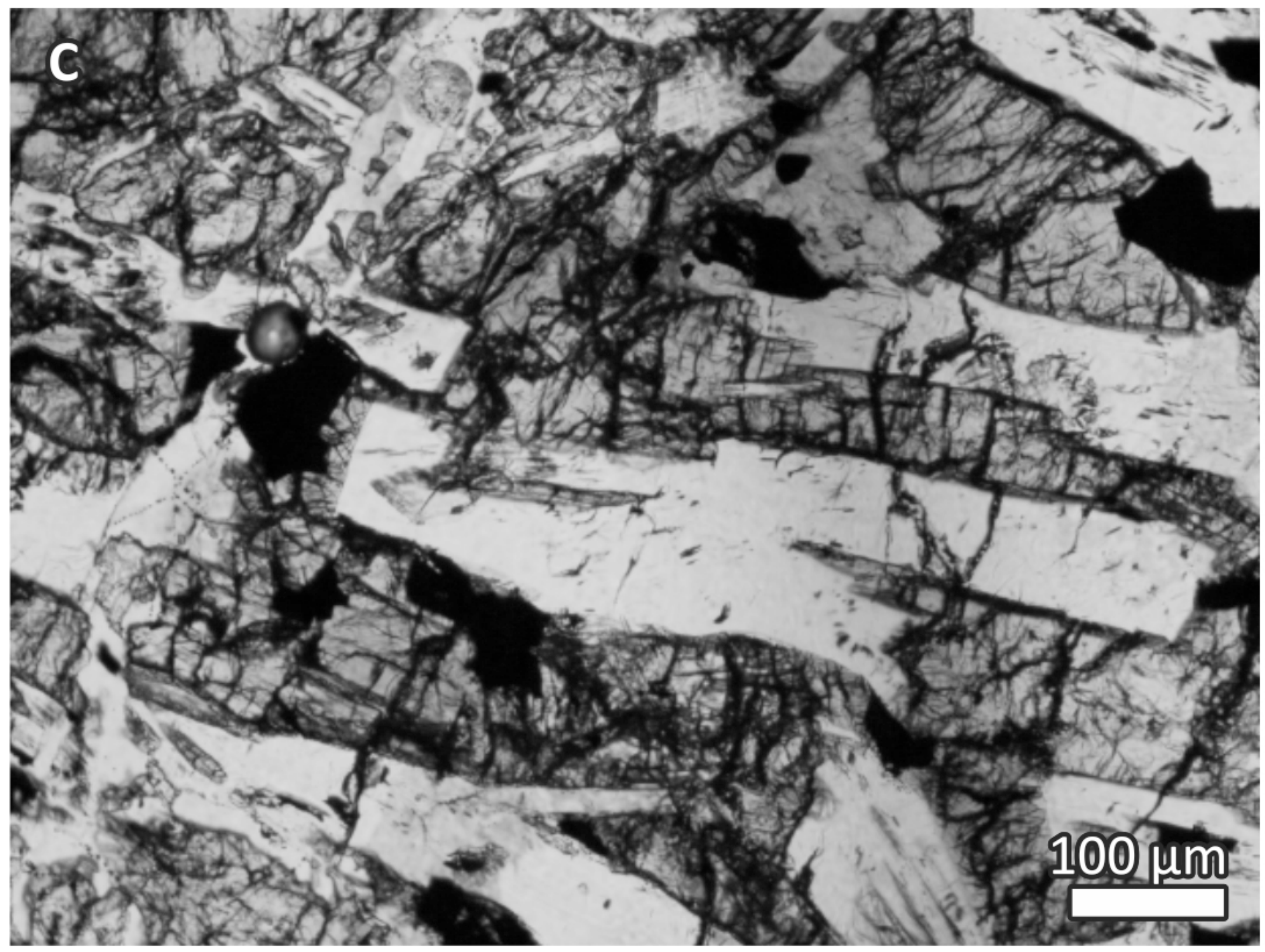
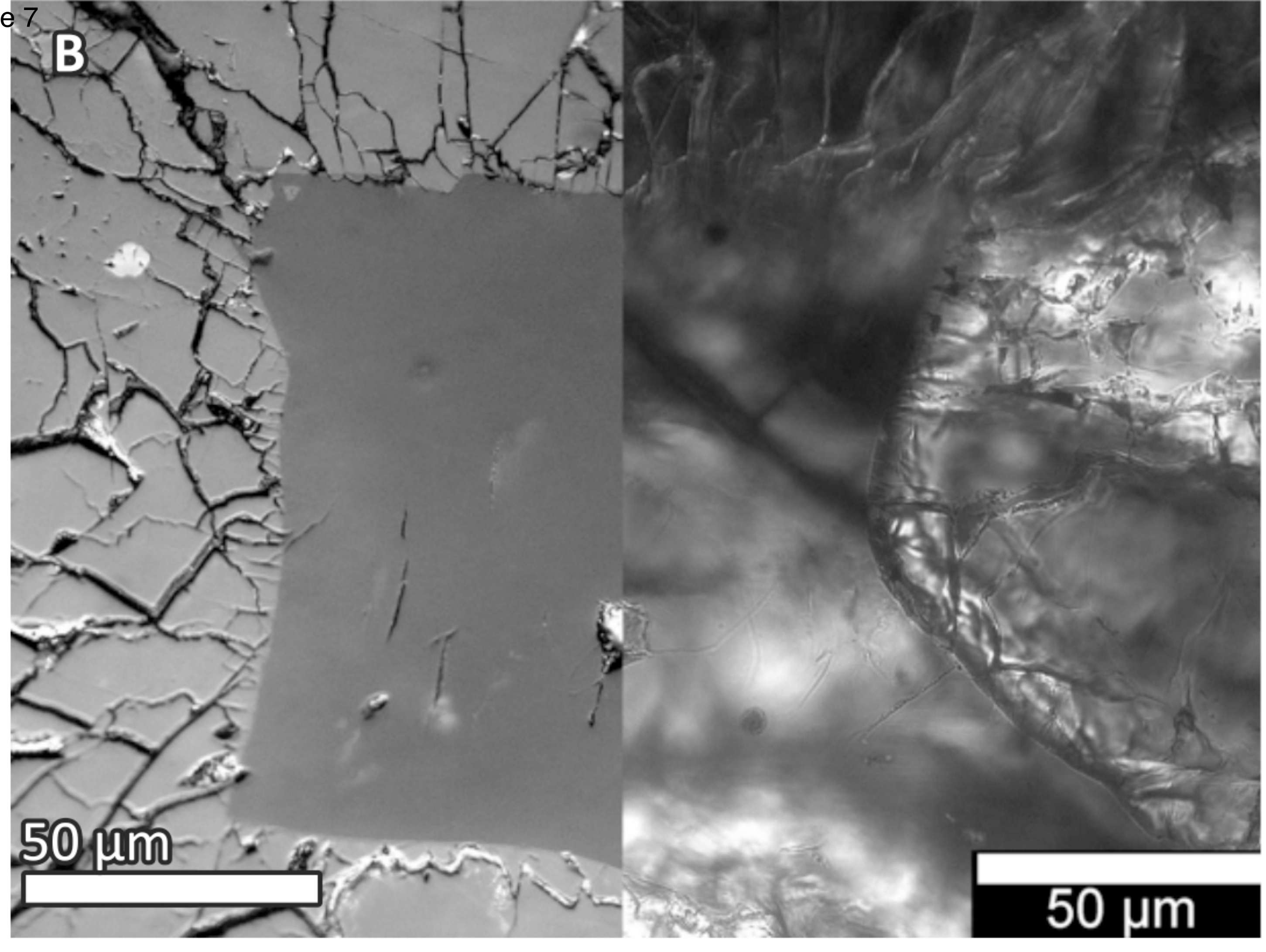
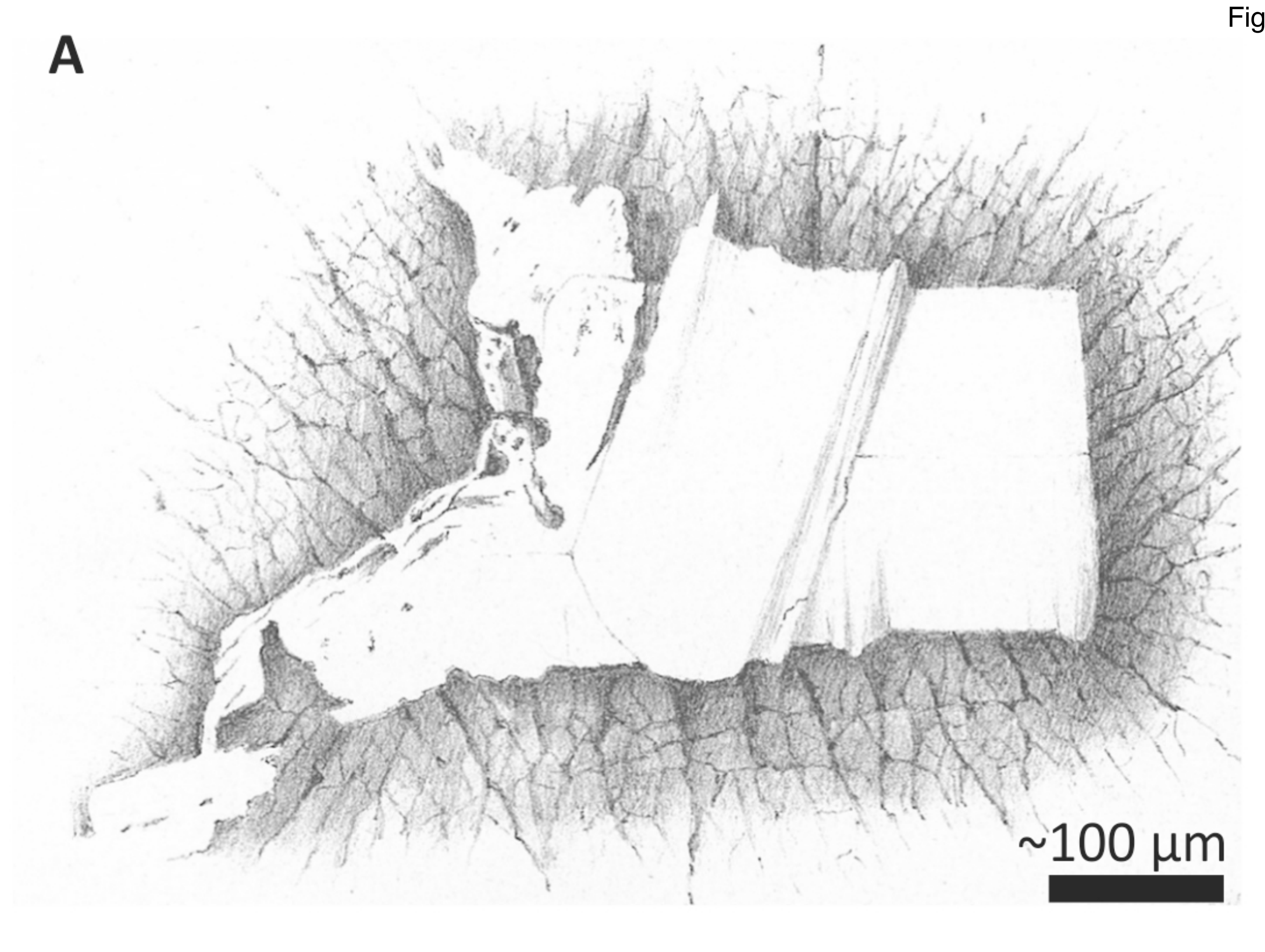


Figure 8

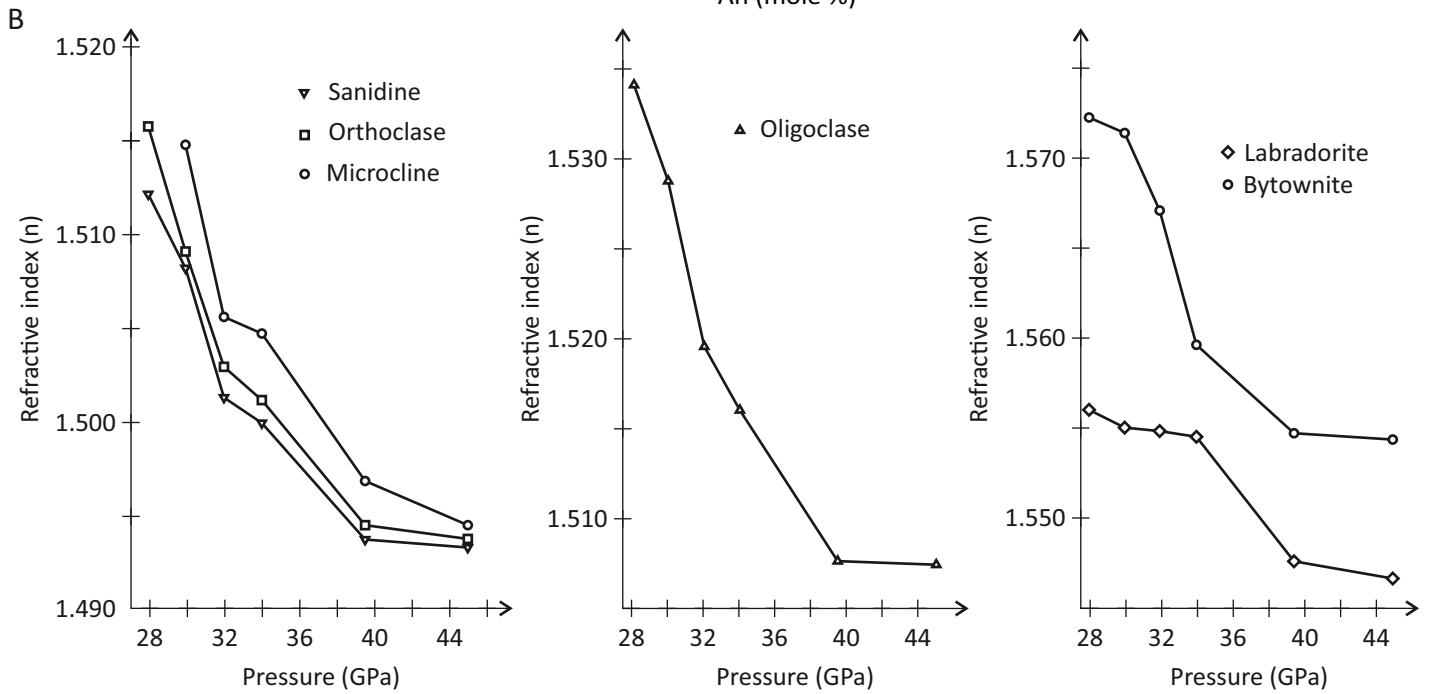
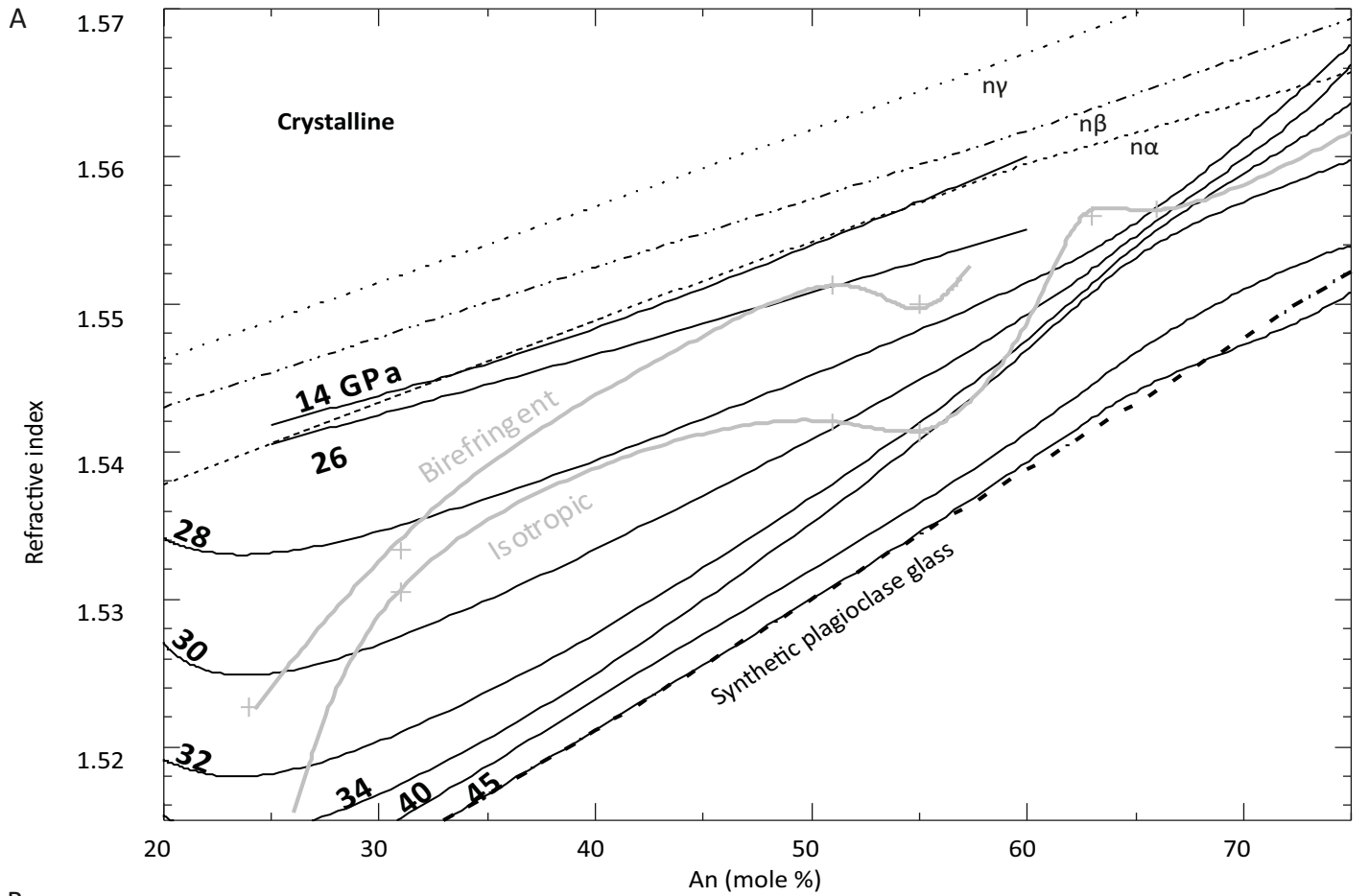


Figure 9

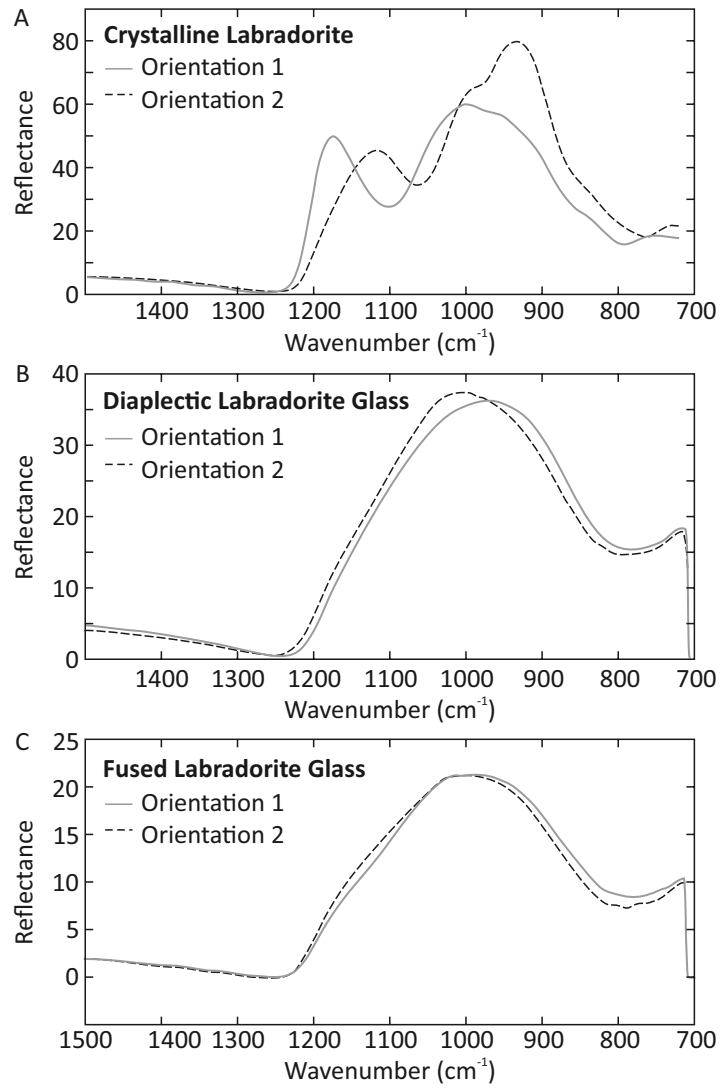
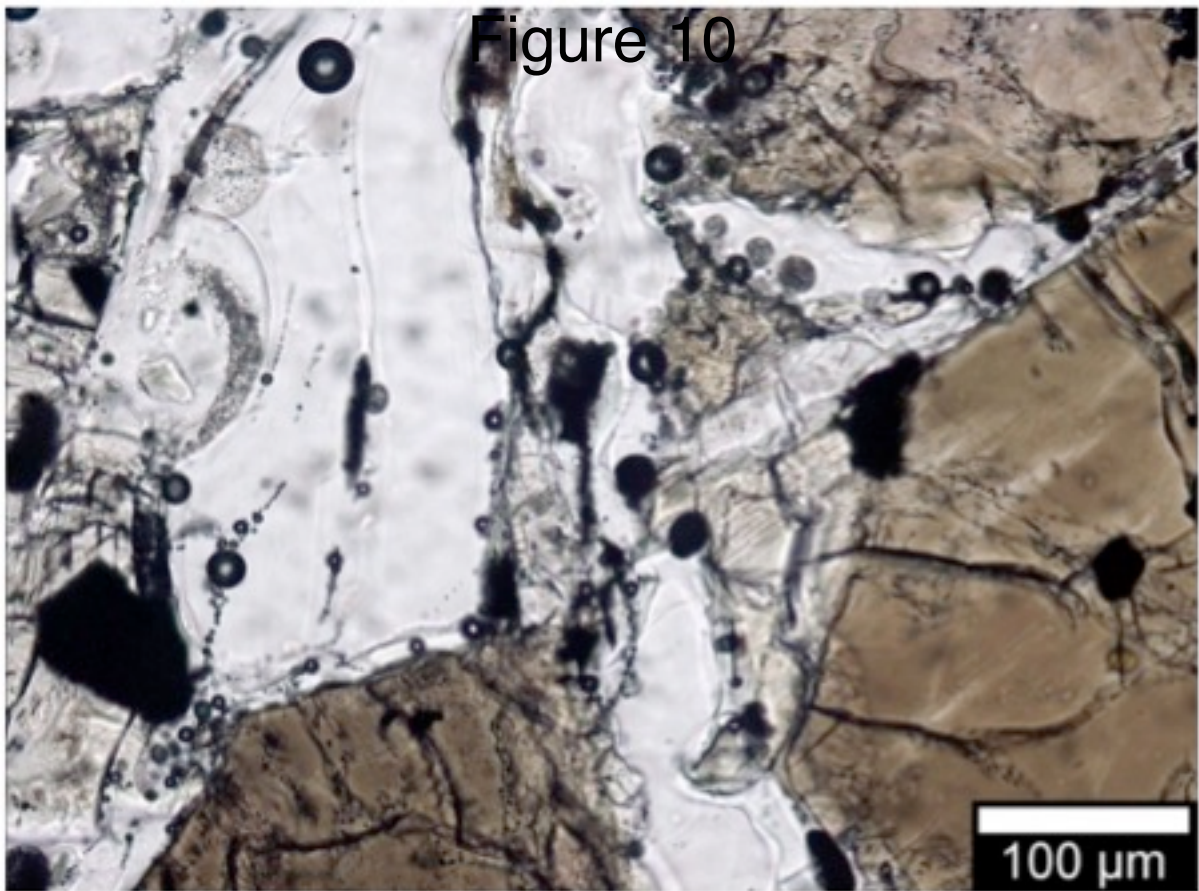


Figure 10



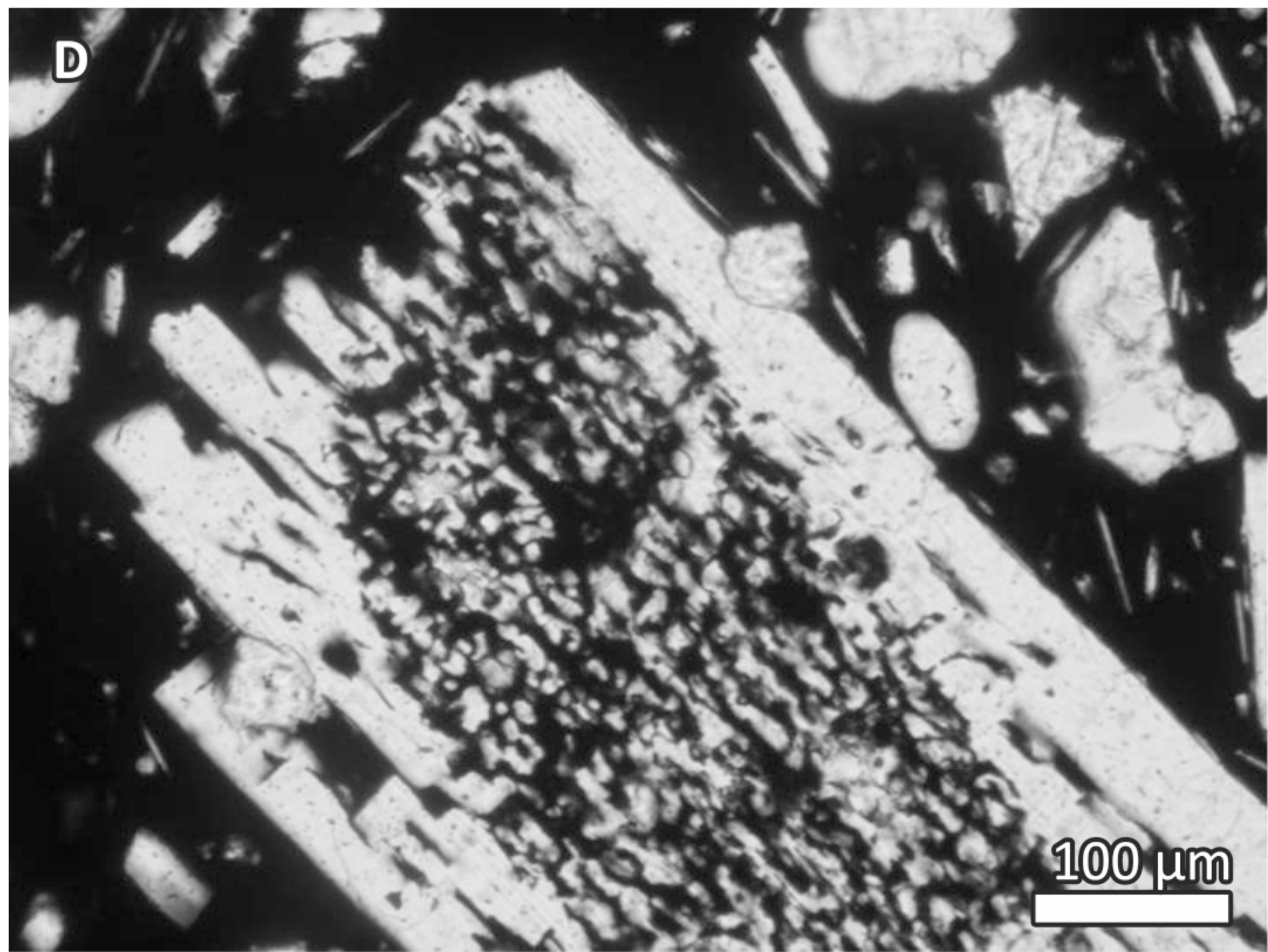
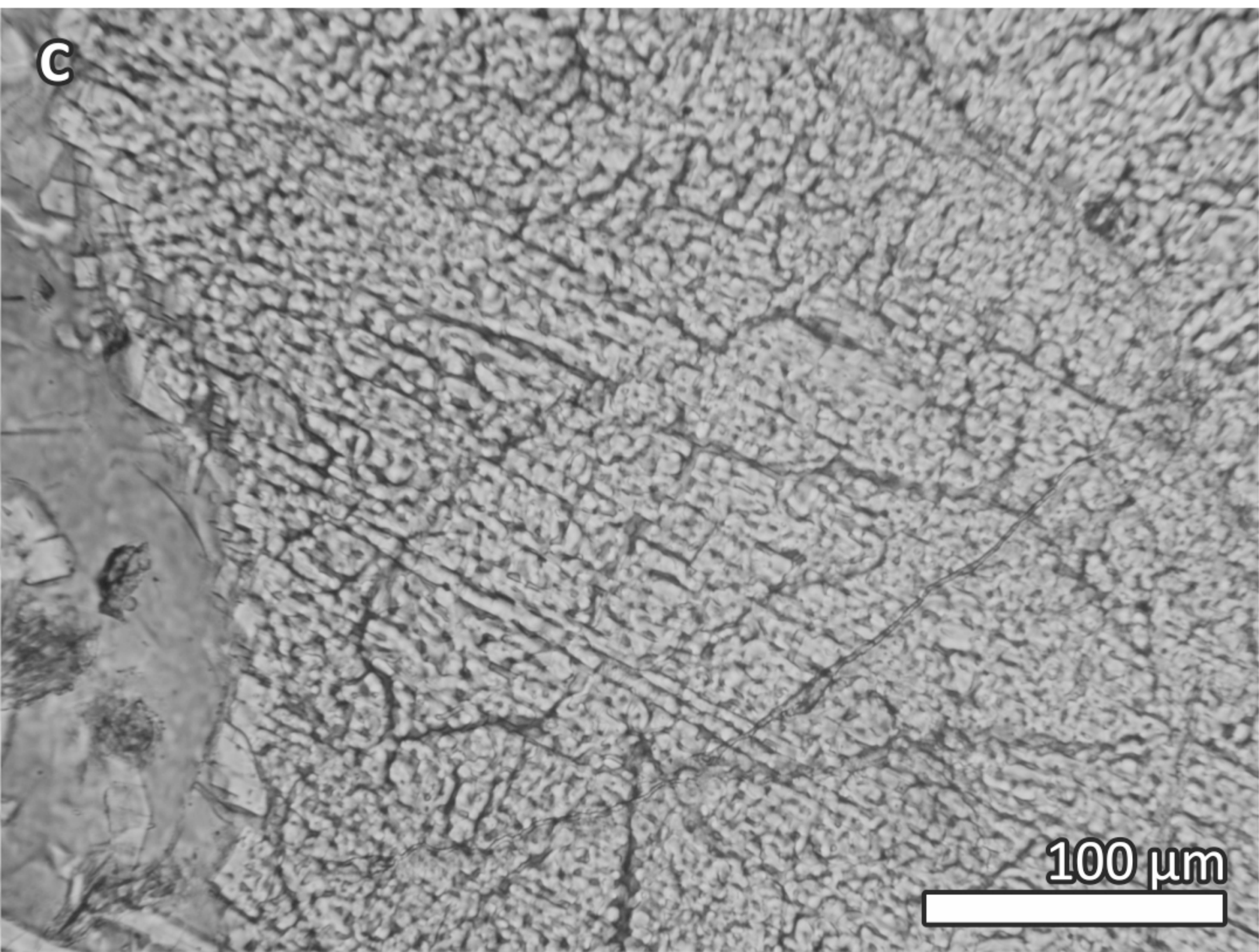
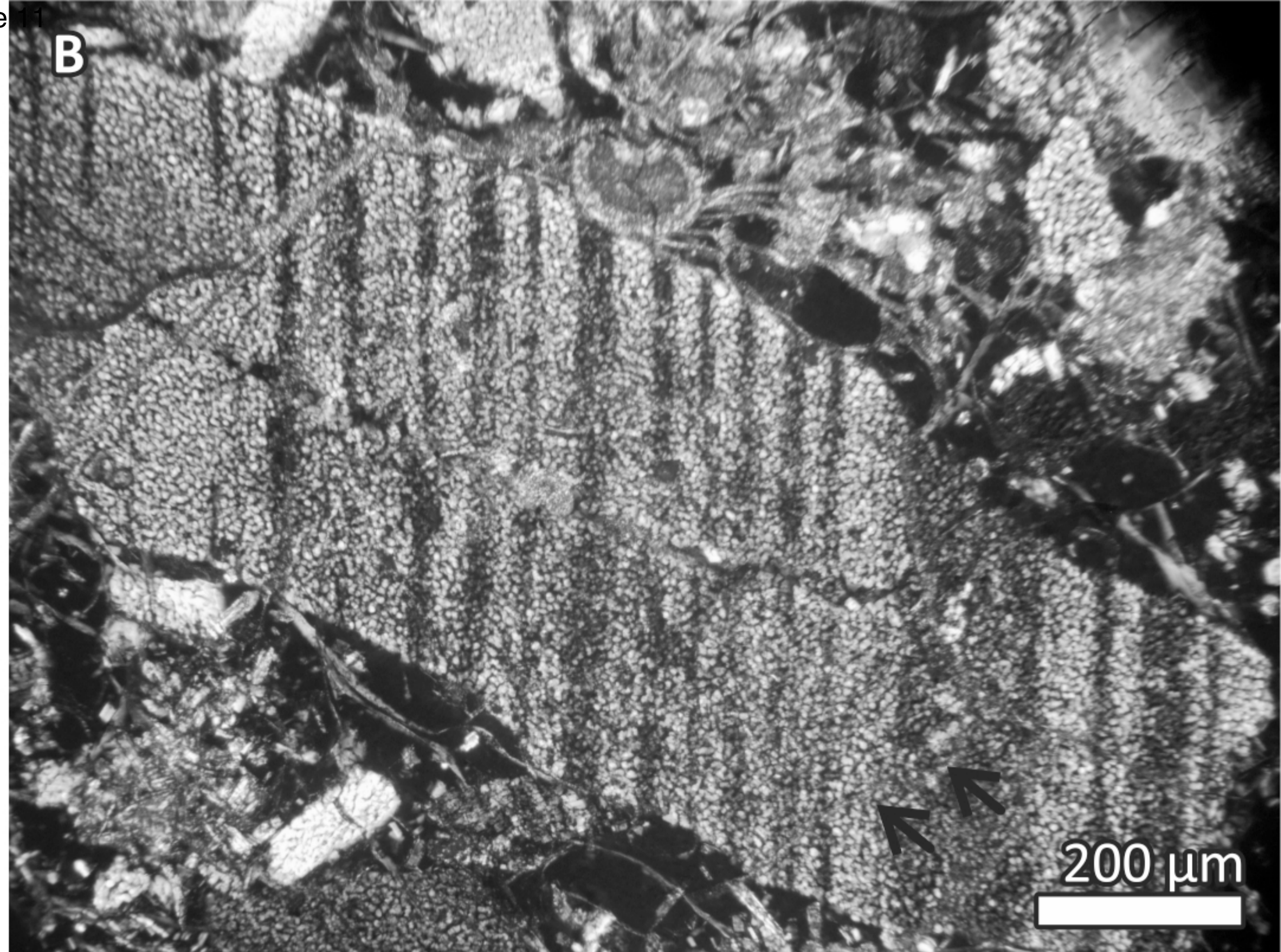
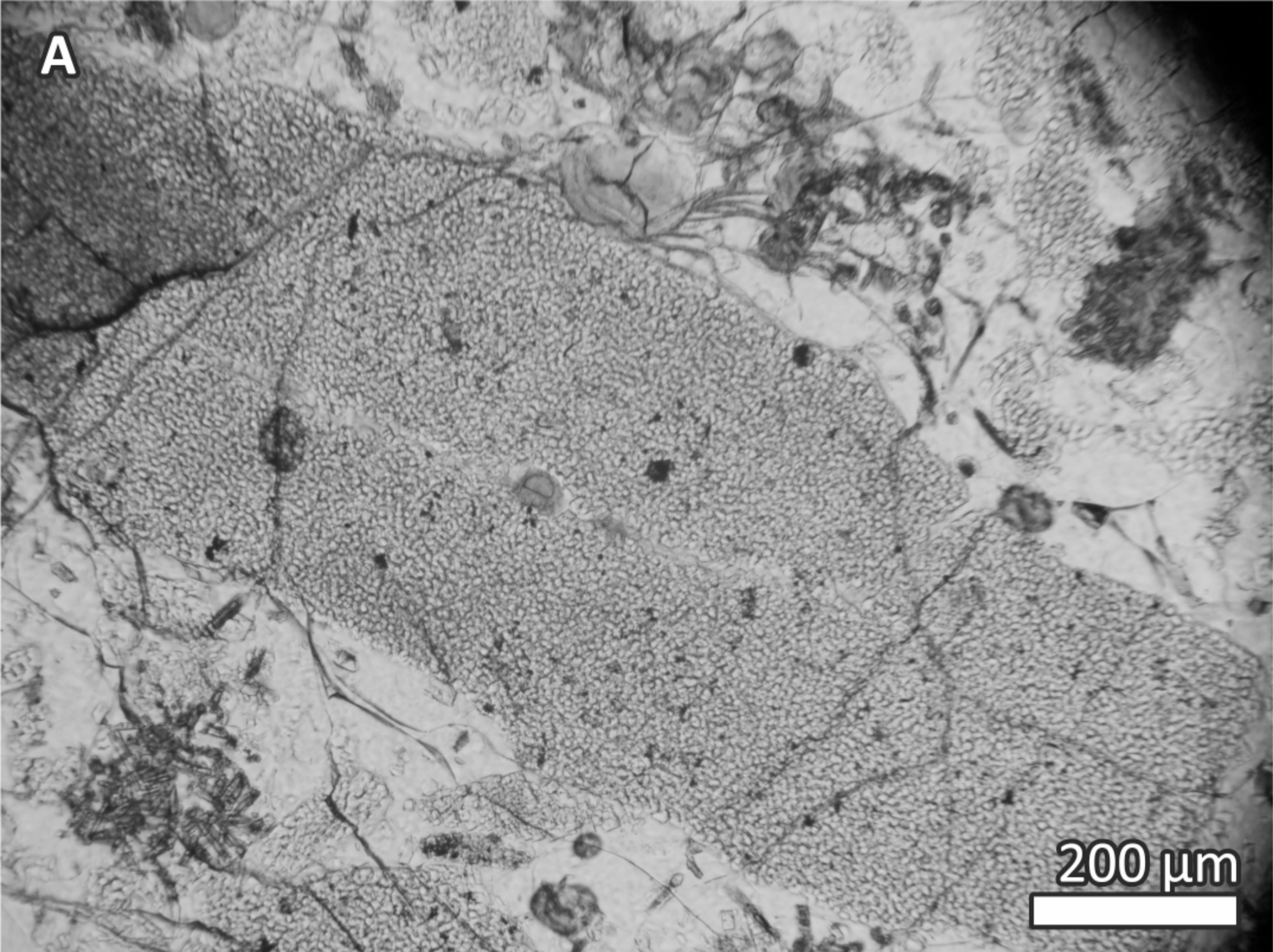


Figure 12

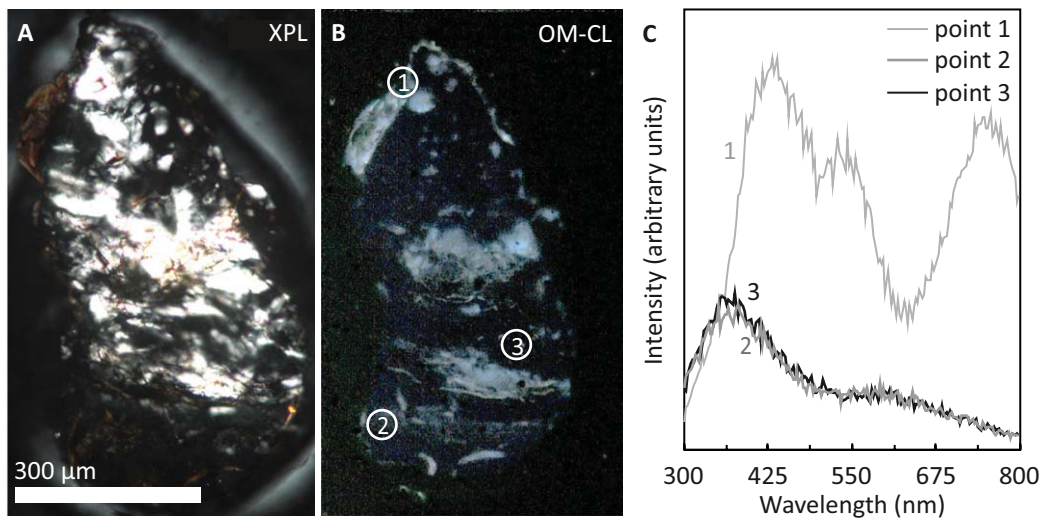


Figure 13

