Geochemical Perspectives Letters

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■ Extreme ¹⁸O-enrichment in majorite constrains a crustal origin of transition zone diamonds

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Abstract

doi: 10.7185/geochemlet.1507

The fate of subducted oceanic lithosphere and its role in the planet-scale geochemical cycle is a key problem in solid Earth studies. Asthenospheric and transition zone minerals included in diamond have been interpreted as representing subducted oceanic crust based on inclusion REE patterns and strong ¹³C depletion of their host diamond (δ^{13} C as low as -23 ‰). This view/explanation, however, has been challenged by alternative interpretations that variable carbon isotopic compositions either result from high temperature fractionation involving carbides, or reflect primordial, unhomogenised mantle reservoirs. Here, we present the first oxygen isotope analyses of inclusions in such ultradeep diamonds - majoritic garnets in diamond from Jagersfontein (South Africa). The oxygen isotope compositions provide unambiguous evidence for derivation of the inclusions from subducted crustal materials. The $\delta^{18}O_{VSMOW}$ values of the majorites range from +8.6 % to +10.0 %, well outside that of ambient mantle ($\pm 5.5 \pm 0.4$ %) and indicate that the protoliths were very heavily weathered at relatively low temperatures. When this information is combined with the broadly eclogitic composition of the majoritic garnets, a derivation from subducted sea-floor basalts is implied. Based on the association between the heavy oxygen and light carbon, the light carbon isotope composition cannot relate to deep mantle processes and is also ultimately derived from the crust.

Received 13 May 2015 | Accepted 18 May 2015 | Published 9 June 2015

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Introduction

Letter

The volatile abundance, bulk composition, and nature of deep Earth are unresolved geochemical questions (e.g., Hirschmann and Dasgupta, 2009). Diamonds and their mineral inclusions are the only materials available from sublithospheric depths and provide the best geochemical information from these otherwise inaccessible regions (Stachel et al., 2005). Such ultra-deep diamonds have been interpreted as having formed in subducted oceanic crust based on trace element patterns of their inclusions showing negative Eu anomalies and ¹³C depletion of their diamond hosts (Tappert et al., 2005b). However, this interpretation is non-unique and controversial (Cartigny, 2005; Corgne et al., 2012). To understand Earth's deep carbon cycle it is, however, of key importance to constrain definitively or refute whether diamond acts as a carrier of crustal derived carbon in deeply subducting slabs. To resolve this question, we present the first analysis of oxygen isotope compositions of ultradeep inclusions – majoritic garnets in diamond from Jagersfontein (South Africa).

The carbon isotopic systematics of the mantle – which may contain more carbon than the crust-biosphere system (Hirschmann and Dasgupta, 2009) are not well understood. Much of the chemical information that we have on deep carbon comes from the study of diamonds and their inclusions (Deines, 1980). From studies of the minerals included in diamonds, it is clear that the vast majority are sourced from the lithospheric mantle. However, there is a small fraction of inclusion assemblages that imply trapping in the asthenosphere, transition zone, or even the lower mantle (Stachel et al., 2005). Much of the difficulty in understanding the nature of the deep carbon cycle revolves around this diamond material. The carbon isotope ratios (δ^{13} C values) in diamond have a wide range - from -41 ‰ to +3 ‰ (Deines, 1980; Cartigny, 2005; De Stefano et al., 2009) however, there is controversy over whether the observed variability represents sampling of distinct, unhomogenised, primitive carbon domains, fractionation at mantle conditions, or subduction of carbon that was fractionated by nearsurface processes (Cartigny, 2005). Though mainly formulated in the context of cratonic diamonds, these considerations apply equally to the origin of ultradeep diamonds. δ^{13} C values for sub-lithospheric diamond have a wide range (from 0 ‰ to -25 ‰; Stachel et al., 2002; Tappert et al., 2005a; Walter et al., 2011) and have often been interpreted as reflecting incorporation of subducted carbon, thereby documenting deep subduction and the subsequent return of originally crustal carbon. Nevertheless, this interpretation is not unique. For example, Mikhail et al. (2014) suggested that iron carbide inclusions in diamonds from Jagersfontein document isotopic fractionation ($\Delta^{13}C_{carbide-diamond} > 7\%$) sufficiently strong to cause the observed ¹³C depleted signature. Similarly, evidence based on the rare earth element signatures (observation of negative Eu anomalies) of majoritic garnet inclusions from Jagersfontein, interpreted to reflect feldspar fractionation in crustal protoliths (Tappert et al., 2005b), was disputed by Griffin and O'Reilly (2007) and Corgne et al. (2012) who considered Eu anomalies to reflect redoxrelated metasomatic signatures, or crystallisation from a melt at high pressures.



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In contrast, oxygen isotope compositions ($\delta^{18}O_{VSMOW}$ values) in mantle peridotites, primitive basalts, and the moon have a narrow range, deviating only slightly from +5.5 ‰ (Eiler, 2001). This gross homogeneity is a consequence of an initially well-mixed Earth (Pahlevan and Stevenson, 2007), and small values of mass dependent stable isotope fractionation at high temperatures which prevent mantle silicates from substantially fractionating oxygen isotopes among each other (Urey, 1947; Criss, 1991). Crustal rocks typically have higher δ^{18} O values mainly within the range +6 to +12 % (Simon and Lécuyer, 2005) due to lowtemperature reactions involving water. In rare cases, crustal rocks can become relatively depleted in ¹⁸O, with values below +5.5 % (Taylor, 1968). Thus, variation makes oxygen isotope compositions a uniquely robust tracer of crustal material in the mantle (Eiler et al., 2000; Schulze et al., 2003). This unique feature has previously been leveraged to test the origin of carbon isotope anomalies in lithospheric diamonds. Measuring the oxygen isotope composition of coesite (Schulze et al., 2003) and eclogitic garnet inclusions (Schulze et al., 2013; Ickert et al., 2013) in diamond, a coupled relationship of high δ^{18} O values (inclusions) and low δ^{13} C values (host diamonds) emerged, constraining a crustal origin of both the diamond carbon and the diamond substrate. For sublithospheric diamonds, similar evidence involving oxygen isotopic analyses of inclusions has so far been lacking.

Rare diamonds from Jagersfontein (South Africa) contain inclusions of majoritic garnet (Deines *et al.*, 1991; Tappert *et al.*, 2005b), an ultra-high pressure mineral reflecting the increasing solubility of pyroxene in the garnet structure at sub-lithospheric depth (Ringwood and Major, 1971). Experimental work (*e.g.*, Akaogi and Akimoto, 1977; Irifune *et al.*, 1986) demonstrated a linear relationship between the solubility of pyroxene (which can be parameterised by Si-excess over the available tetrahedral sites) and pressure, thereby providing crude minimum depth estimates. The studied inclusions from Jagersfontein have Si contents of up to 3.5 atoms per formula unit, indicating that some of the inclusions came from depths of at least 500 km (Tappert *et al.*, 2005b), well within the transition zone (Fig. 1). Based on low Cr contents (<0.4 wt. %) the majoritic garnets are classified as "eclogitic", or based on their low Na concentrations, metapyroxenitic (Gurney, 1984; Schulze, 2003; Kiseeva *et al.*, 2012). The relatively high molar Mg/(Mg+Fe) (0.68-0.83) of these inclusions preclude derivation from typical pelitic sediments.

Methods

Mineral inclusions in diamonds have generally presented a unique analytical challenge because of their small size. Non-destructive or microbeam techniques are preferred because they consume little or no sample. Although Lowry *et al.* (1999) employed a novel, bulk laser fluorination technique to measure oxygen isotope compositions of mineral inclusions in diamond, most subsequent workers have employed secondary ion mass spectrometry (SIMS) for δ^{18} O measurements (Schulze *et al.*, 2003; Smart *et al.*, 2012; Ickert *et al.*, 2013; Schulze *et al.*, 2013).

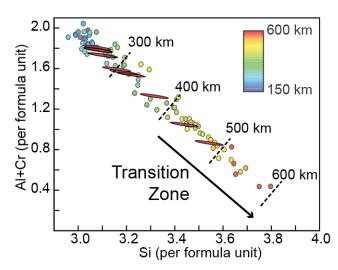


Figure 1 The chemical composition of the Jagersfontein majoritic garnet inclusions based on a formula unit containing twelve oxygens, and their relationship to minimum entrapment depth. The elongate ellipses are the chemical compositions of the garnets and the size and shape of the ellipse represents the 95 % confidence limit based on electron microprobe uncertainties. The small circles are experimental results (from the compilation of Collerson et al., 2010, with the addition of data from Irifune et al., 1986), with colour shading corresponding to the apparent depth that their pressure of equilibration represents. For clarity, uncertainties are not shown on the experimental results; however they are approximately the same size as the uncertainties for the natural garnets. Depth contours are derived from a linear regression of the experimental data relating Si to equilibration pressure.

We made high-precision SIMS δ^{18} O measurements on the majorite inclusions in diamond originally studied by Tappert et al. (2005b), who published trace element measurements of the majorities and carbon isotope compositions of their host diamonds. Analytical procedures follow Ickert and Stern (2013), with modifications described in the Supplementary Information. As in previous studies (e.g., Schulze et al., 2013), the inclusions, which are <200 µm in their largest dimension, were released by crushing the host diamond, and mounted in epoxy grain mounts along with reference materials. The data are summarised in Table 1, and the full dataset is available in the Supplementary Information Table S-1. Each majorite sample was analysed twice, and in each case the two spots were identical within uncertainty. A major obstacle to accurate SIMS analysis is the dependence of the instrumental mass fractionation on the chemical composition and structure of the target material, the so called "matrix effect" (Russell et al., 1980; Eiler et al., 1997; Vielzeuf et al., 2005; Page et al., 2010). We employ the technique described by Page et al. (2010) to correct for the bias between the primary reference material and the majorite samples based on Ca-content. However, we have no reference material for which we can correct for the majorite component. Since we observe



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a uniformly heavy oxygen isotope composition at a wide range of abundances of majorite component (including nearly no majorite), we infer that the matrix effect associated with the majorite substitution is small and is not responsible for the heavy oxygen compositions. The range in majorite compositions mainly reflects the variable trapping depth, and therefore the solubility of pyroxene in the garnet. The total uncertainty for each spot analysis is approximately ± 0.4 % (2σ), which is mainly due to overdispersion of the primary reference material (± 0.3 %) and the matrix correction (± 0.2 %), but includes the uncertainty in the primary reference material and the uncertainty in the session-level mass fractionation correction (see Ickert and Stern, 2013 for details).

Table 1 Oxygen and carbon isotope compositions of majoritic inclusions and their diamond hosts, respectively. The carbon isotope compositions are from Tappert *et al.* (2005b). Uncertainties on δ^{18} O values include repeatability of the reference material (RM), the calibration uncertainty, and the uncertainty in the bulk values of the RM. Note that diamonds JF-01 and JF-37 have two inclusions each.

Sample	Spot ID	δ ¹⁸ O _{VSMOW}	Mean	±95 % c.l.	$\delta^{13} C_{VPDB}$	
JF-01A	S1318_1	8.78	8.88	0.37	-21.2	
	S1318_2	8.97	0.00	0.37	-21.2	
JF-01B	S1319_1	9.07	9.20	0.37	-21.2	
	S1319_2	9.32	9.20	0.57		
JF-09A	S1320_1	9.94	10.00	0.37	-21.8	
Jr-09A	S1320_2	10.06	10.00	0.57	-21.8	
IF-22A	S1321_1	9.02	9.06	0.27	17.0	
JF-22A	S1321_2	9.09	9.06	0.37	-17.3	
JF-37A	S1322_1	9.28	0.22	0.37	-21.7	
	S1322_2	9.36	9.32	0.37		
JF-37B	S1323_1	9.71	9.72	0.37	-21.7	
	S1323_2	9.73	9.72	0.57		
IE 20A	S1324_1	9.19	9.09	0.37	-17.9	
JF-39A	S1324_2	8.98	9.09	0.37		
JF-42A	S1325_1	9.26	0.21	0.37	-17.4	
	S1325_2	9.15	9.21	0.37	-1/.4	
JF-44B	S1326_1	8.62	8.74	0.38	-17.7	
	S1326_2	8.85	8.74	0.38		
JF-55A	S1328_1	8.62	0.57	0.20	17.6	
	S1328_2	8.52	8.57	0.38	-17.6	
JF-58B	S1331_1	9.14	9.21	0.38	22.0	
	S1331_2	9.28	9.21	0.38	-23.0	

Discussion and Conclusions

The measured δ^{18} O values for the inclusions fall into a small range from +8.6 % to +10.0 % (Fig. 2). These values far exceed the range of upper mantle peridotite and derivative melts ($\pm 5.5 \pm 0.4$ %; Mattey et al., 1994); in the crust, values this high (in silicate rocks) are only associated with strong interactions between rocks and water at low temperature (Taylor and Sheppard, 1986). High temperature processes, such as CO₂ escape from carbon-rich melts, can fractionate both carbon and oxygen isotope ratios (Deines, 1980; Ickert et al., 2013), however this will only occur under unusual circumstances, and the magnitudes of the fractionations are relatively small and cannot account for >4 % shifts. Moreover, if the carbon isotope data are also considered, there are no processes that can simultaneously create depletion in ¹³C and enrichment in ¹⁸O (Ickert et al., 2013). Cartigny (2010) suggested that the depleted carbon isotope signatures of the Jagersfontein transition zone-suite of diamonds were derived from a primordial mantle reservoir and not from crustal materials. The presence of a previously undetected high- δ^{18} O, primordial reservoir in the deep Earth is unlikely, given the homogeneity of the mantle and the near identical oxygen isotope systematics between the Earth and Moon (Wiechert et al., 2001), so we suggest that this hypothesis is now untenable.

These new data establish a link between the high $\delta^{18}O$ values of inclusions and the low ^{13}C values of the diamond hosts. This relationship is firm evidence for a subducted protolith and makes it extremely unlikely that the associated negative $\delta^{13}C$ values are related to deep mantle isotopic fractionation (Mikhail et al., 2014). In a subduction scenario, the low $\delta^{13}C$ values (Fig. 2) imply that the carbon originated not as carbonate but as organic carbon – either residual from biological processes or produced abiotically by Fischer-Tropsch-type reactions. The $\delta^{18}O$ values require that the host rocks were strongly weathered, and probably underwent substantial secondary hydration on the sea floor.

This $\delta^{18}\text{O}-\delta^{13}\text{C}$ relationship implies that the recovered ultradeep inclusions were derived from intensely hydrated (e.g., $\delta^{18}\text{O} >> +7$ %) and typically oxidised basaltic rock close to the seawater interface, and that the diamond carbon was closely associated with this protolith. Such rocks have a low relative abundance in the oceanic crust, and the coincidence of their rarity and their apparent association with diamonds most likely attests to a causal relationship. That reduced, rather than oxidised, carbon is involved suggests that after thermal cracking, the carbon is subducted as metastable graphite. Since conversion of metastable graphite to diamond is not expected to produce macrocrystals (Sung, 2000), diamond formation during a fluid/melt aided dissolution – precipitation process is likely. The volatile-rich nature of the protolith is consistent with the recent suggestion by Harte (2010) that ultradeep diamonds are associated with devolatilisation and possibly melting reactions.



Letter

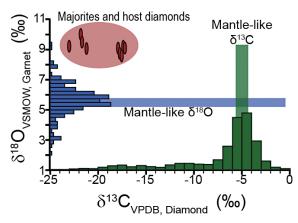


Figure 2 The δ^{18} O compositions of the majoritic garnets and the δ^{13} C values of their host diamonds. The majoritic garnets and their diamond hosts plot well away from the mantle field. The histogram on the Y-axis is of garnets from eclogite xenoliths (compilation of Ickert *et al.*, 2013), the histogram on the X-axis is of cratonic diamonds (Stachel *et al.*, 2009).

Acknowledgements

Full datasets and methods can be found in the Supplementary Information. The diamonds for this project were gifted to JWH by the Diamond Trading Company, a member of the DeBeers Group of Companies. DeBeers are sincerely thanked for this material. RBI was supported by a Killam Memorial Postdoctoral Fellowship. The Canadian Centre for Isotopic Microanalysis (CCIM) was created through Canadian Foundation for Innovation and Alberta Science and Research Investment Program grants to TS. TS acknowledges support through a Natural Sciences and Engineering Research Council of Canada Discovery grant and the Canada Research Chairs program. This is a contribution of CCIM project P1014.

Editor: Bruce Watson

Additional Information

Supplementary Information accompanies this letter at www.geochemicalperspectivesletters.org/article1507

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Cite this letter as: Ickert, R.B., Stachel, T., Stern, R.A., Harris, J.W. (2015) Extreme ¹⁸O-enrichment in majorite constrains a crustal origin of transition zone diamonds. *Geochem. Persp. Let.* 1, 65-74.

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Supplementary Information

The Supplementary Information includes:

- ➤ Supplementary Methods
- ➤ Figures S-1 to S-3
- ➤ Table S-1

SI-I

> Supplementary Information References

Supplementary Methods

Majoritic garnets were embedded in epoxy resin along with reference garnets S0068 and S0088B (Ickert and Stern, 2013). Using the University of Alberta ims1280 at the Canadian Centre for Isotopic Microanalysis, a ~15 mm diameter $^{133}\text{Cs}^+$ (20 keV impact energy) beam was used to sputter 1-2 mm deep pits in inclusions that were mounted in epoxy with garnet reference material. Garnet S0068 was used as the primary reference material for the $^{18}\text{O}/^{16}\text{O}$ measurements on the day of the analyses (Fig. S-1). However, the measured $\delta^{18}\text{O}$ value of the majoritic garnets will also depend on the difference in chemical composition between S0068 and the majoritic garnets. This "matrix effect" is corrected by referencing the garnet measurements to a calibration defined by garnets of known isotopic composition and varying Ca-Mg-Fe abundances, in a manner similar to Page *et al.* (2010) and described by Ickert and Stern (2013). The magnitude of this secondary effect is small and ranges from +0.1 ‰ to +1.1 ‰ and in all cases produces a lower $\delta^{18}\text{O}$ value than an "uncorrected" value (Fig. S-2).

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The accuracy of the secondary matrix correction is assessed by the measurement of secondary reference material S0088B, which requires a substantial (>3 ‰) correction (Fig. S-2). The mean δ^{18} O value measured for S0088B was +4.06 ‰, in excellent agreement with the conventionally determined δ^{18} O value of +4.13 ± 0.09 ‰ (Fig. S-2; Ickert and Stern, 2013). The repeatability of S0068 provided a measure of single spot precision. On the two mounts on which the majoritic garnets were embedded, S0068 yielded repeatabilities of ±0.16 ‰ and ±0.17 ‰ (1SD, n_1 = 16, n_2 = 14, no rejections). Each individual grain was measured twice, and the difference between the two spots varied from 0.02 ‰ to 0.26 ‰, well within the 2 σ uncertainty for each spot (Fig. S-3). Final uncertainties are a combination of the repeatability of S0068 (which is the dominant source of uncertainty), the correction for session-related mass bias, the uncertainty in the matrix correction, and the uncertainty in the bulk δ^{18} O value of S0068, and are better than ±0.38 ‰ (2 σ).

We have made no correction for any potential matrix effects associated with the majorite component. However, we note that despite the wide range of majorite component in these crystals (including grains that include only very small amounts of majorite), there is no correlation between the measured $\delta^{18}\text{O}$ values and the amount majorite in the garnet. This strongly suggests that any matrix effect related to majorite is negligible. This is expected, given that the compositional and crystal-chemical differences between pyrope-rich Ca-Mg-Fe²+ garnets and these majoritic garnets are small.

Two pairs of inclusions (JF-01A, JF-01B and JF-37A, JF37B) are each from the same diamond. In both cases, the measured δ^{18} O values are identical within uncertainty.

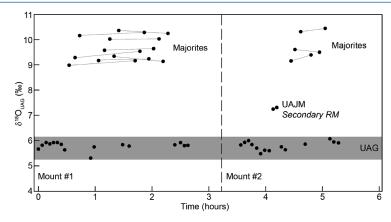


Figure S-1 Time series of the majorite analyses. All data are plotted as raw isotope ratios relative to 0.0020052, corrected only for amplifier gains and backgrounds. There are no corrections needed for drift or to eliminate outliers. Tie lines connect analyses of the same grain.

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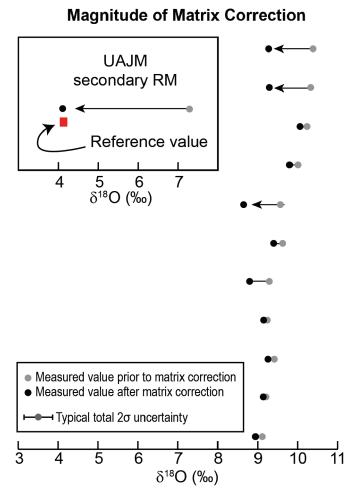


Figure S-2 Magnitude of the matrix correction. Each pair of points represents a single analysis, with the grey point representing the value prior to correction for the matrix effect, and the black point representing the value after the matrix correction. For reference, secondary reference material UAJM is plotted, which, after making a much larger correction for matrix effects than needed for the majorities, agrees within uncertainty with the nominal reference value.

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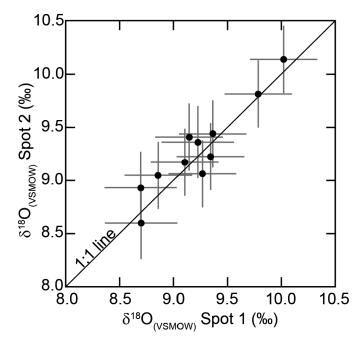


Figure S-3 Plot of each pair of analyses from the same grain. The analyses all agree within uncertainty, suggesting that the uncertainties have been estimated correctly.

Table S-1 Full analytical data for oxygen isotope measurements of majorites.

ID	Session	Time (hours)	$\delta^{18}O_{raw}$	1σ1	$\delta^{18}O_{UAG}$	X _{Ca*}	δ ¹⁸ O _{VSMOW}
S0068_1@28	IP10155	0.92	8.15	0.05			
S0068_1@29	IP10155	0.98	8.60	0.05			
S0068_2@07	IP10155	0.07	8.67	0.05			
S0068_2@08	IP10155	0.13	8.78	0.04			
S0068_2@09	IP10155	0.20	8.72	0.05			
S0068_2@10	IP10155	0.27	8.76	0.04			
S0068_2@11	IP10155	0.33	8.77	0.04			
S0068_2@12	IP10155	0.40	8.70	0.07			
S0068_2@13	IP10155	0.47	8.48	0.04			
S0068_2@14	IP10155	1.48	8.69	0.06			
S0068_2@15	IP10155	1.60	8.63	0.04			
S0068_2@16	IP10155	2.40	8.68	0.07			
S0068_2@17	IP10155	2.50	8.77	0.05			
S0068_2@18	IP10155	2.57	8.65	0.04			
S0068_2@19	IP10155	2.63	8.66	0.06			
S0068_2@6	IP10155	0.00	8.51	0.07			
S1318_1	IP10155	0.55	11.85	0.05	8.93	0.16	8.78
S1318_2	IP10155	1.70	12.04	0.05	9.11	0.16	8.97
S1319_1	IP10155	0.65	12.14	0.06	9.22	0.16	9.07
S1319_2	IP10155	1.78	12.40	0.06	9.47	0.16	9.32
S1320_1	IP10155	0.73	13.02	0.05	10.09	0.16	9.94
S1320_2	IP10155	1.87	13.14	0.05	10.21	0.16	10.06
S1321_1	IP10155	1.07	12.03	0.04	9.10	0.15	9.02
S1321_2	IP10155	1.95	12.10	0.05	9.17	0.15	9.09
S1322_1	IP10155	1.17	12.43	0.06	9.50	0.17	9.28
S1322_2	IP10155	2.03	12.50	0.06	9.58	0.17	9.36
S1323_1	IP10155	1.25	12.85	0.04	9.92	0.17	9.71
S1323_2	IP10155	2.12	12.88	0.05	9.94	0.17	9.73
S1324_1	IP10155	1.33	12.19	0.06	9.26	0.15	9.19
S1324_2	IP10155	2.20	11.99	0.07	9.06	0.15	8.98
S1325_1	IP10155	1.42	13.22	0.05	10.29	0.30	9.26
S1325_2	IP10155	2.28	13.11	0.06	10.18	0.30	9.15



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ID	Session	Time (hours)	$\delta^{18}O_{raw}$	1σ1	$\delta^{18}O_{UAG}$	X _{Ca*}	δ ¹⁸ O _{VSMOW}
S0068_1@1	IP10156	3.57	9.02	0.05			
S0068_1@2	IP10156	3.63	9.12	0.04			
S0068_1@3	IP10156	3.70	9.19	0.04			
S0068_1@4	IP10156	3.77	9.04	0.04			
S0068_2@1	IP10156	3.85	8.91	0.06			
S0068_2@10	IP10156	5.28	9.11	0.04			
S0068_2@2	IP10156	3.92	8.69	0.05			
S0068_2@3	IP10156	3.98	8.82	0.05			
S0068_2@4	IP10156	4.07	8.79	0.04			
S0068_2@5	IP10156	4.28	8.95	0.05			
S0068_2@6	IP10156	4.35	8.83	0.06			
S0068_2@7	IP10156	4.70	9.06	0.04			
S0068_2@8	IP10156	5.13	9.26	0.05			
S0068_2@9	IP10156	5.20	9.15	0.05			
S0088B_1@1	IP10156	4.13	10.44	0.08	7.16	0.97	4.02
S0088B_1@2	IP10156	4.20	10.51	0.07	7.23	0.97	4.09
S1326_1	IP10156	4.45	12.38	0.05	9.10	0.21	8.62
S1326_2	IP10156	4.80	12.61	0.05	9.33	0.21	8.85
S1328_1	IP10156	4.52	12.83	0.04	9.54	0.19	8.62
S1328_2	IP10156	4.95	12.72	0.05	9.43	0.19	8.52
S1331_1	IP10156	4.62	13.52	0.04	10.23	0.12	9.14
S1331_2	IP10156	5.05	13.66	0.06	10.37	0.12	9.28

 $\delta^{18}O_{raw}$ — The $^{18}O/^{16}O$ ratio measured on the instrument, corrected only for amplifier gains and baslines

 $\delta^{18}O_{UAG}$ — The $^{18}O/^{16}O$ ratio corrected for mass bias using the UAG garnet $\delta^{18}O_{VSMOW}$ — The $^{18}O/^{16}O$ ratio corrected for Ca-related (matrix effect) mass bias

		S0068				
Session	SD	n	SE	1σ (δ ¹⁸ Ο) ¹	$1\sigma_{matrix}^{2}$	2σ total
IP10155	0.16	16	0.04	0.065	0.09	0.37
IP10156	0.17	14	0.05	0.065	0.07	0.38

¹ The uncertainty on the determination of the bulk δ^{18} O value of the reference material by laser fluorination, from Ickert and Stern (2013).

² The uncertainty on the matrix correction (Ickert and Stern, 2013). The uncertainty is a function of the Ca-abundance, and for these garnets varies from 0.05-0.09 ‰, and for simplicity, the largest value is used for all garnets.



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Supplementary Information References

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