

hep-ph/0101088
 January 2001
 DO-TH 2000/19
 MSUHEP-01205
 Cavendish-HEP-2001/01

Predictions for Neutrino Structure Functions

S. Kretzer,^{a,b} F.I. Olness,^c R.J. Scalise,^c R.S. Thorne,^{d,*} U.K. Yang^{e,f}

^aUniv. Dortmund, Dept. of Physics, D-44227 Dortmund, Germany

^bMichigan State University, Dept. of Physics and Astronomy, East Lansing, MI 48824-1116

^cSouthern Methodist University, Department of Physics, Dallas, TX 75275-0175

^dCavendish Laboratory, University of Cambridge, Madingley Road, Cambridge, CB3 0HE

^eUniversity of Chicago, Enrico Fermi Institute, Chicago, IL 60637-1434

^fUniv. of Rochester, Rochester, NY 14627

The first measurements of ΔxF_3 are higher than current theoretical predictions. We investigate the sensitivity of these theoretical predictions upon a variety of factors including: renormalization scheme and scale, quark mass effects, higher twist, isospin violation, and PDF uncertainties.

1. Introduction

Deep inelastic lepton-nucleon scattering experiments have provided precision information about the quark distributions in the nucleon. However, there has been a long-standing discrepancy between the F_2 structure functions extracted from neutrino and muon experiments in the small x range. Recently, a new analysis of differential cross sections and structure functions from CCFR ν_μ -Fe and $\bar{\nu}_\mu$ -Fe data was presented; in this study, the neutrino-muon difference is resolved by extracting the ν_μ structure functions in a physics model independent way.[1]

In previous analyses of ν_μ data,[2] structure functions were extracted by applying a slow rescaling correction to correct for the charm mass suppression in the final state. In addition, the ΔxF_3 term (used as input in the extraction) was calculated from a leading order charm production model. These resulted in physics model dependent (PMD) structure functions. In the new analysis,[1] slow rescaling corrections are not applied, and ΔxF_3 and F_2 were extracted from two parameter fits to the data.

The extracted physics model independent (PMI) values for F_2^ν are then compared with F_2^μ within the framework of NLO models for massive charm production; these are found to be in agreement, thus resolving the long-standing discrepancy between the two sets of data. However, the first measurements of ΔxF_3 are higher than current theoretical predictions. The objective of this paper is to investigate the sensitivity of ΔxF_3 upon a variety of factors including renormalization scheme and scale, quark mass effects, higher twist,

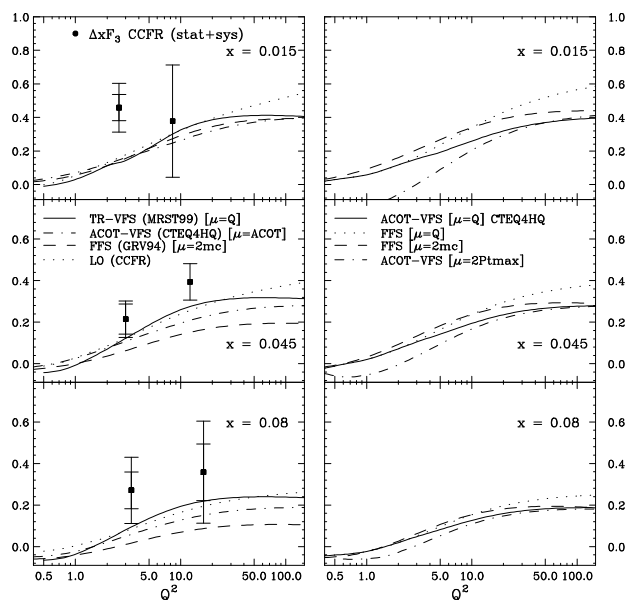


Figure 1. Comparison of ΔxF_3 data from CCFR with various theoretical predictions.

isospin violation, and PDF uncertainties.

In Fig. 1 we have plotted the quantity ΔxF_3 for an isoscalar target computed to order α_s^1 . For comparison, we also display data from the CCFR analysis.[1] While there is much freedom in the theoretical calculation, the difference between these calculations and the data at low Q values warrants further investigation. We will discuss and compare the different theoretical calculations, and examine the inherent uncertainty in each with respect to different input parameters. We

*Royal Society University Research Fellow

will also examine the experimental input, and assess uncertainties in this data.

1.1. Measurement of ΔxF_3

The structure functions are defined in terms of the neutrino-nucleon cross section via:

$$\frac{d^2\sigma^{\nu,\bar{\nu}}}{dx dy} = \frac{G_F^2 M E_\nu}{\pi} \left[F_2(1-y) + xF_1y^2 \pm xF_3y(1-\frac{y}{2}) \right] \quad (1)$$

where G_F is the Fermi weak coupling constant, M is the nucleon mass, E_ν is the incident energy, $y = E_h/E_\nu$ is the fractional energy transfer, and E_h is the final state hadronic energy.

The sum of ν_μ and $\bar{\nu}_\mu$ differential cross sections for charged current interactions on an isoscalar target is then:

$$\begin{aligned} F(\epsilon) &\equiv \left[\frac{d^2\sigma^\nu}{dx dy} + \frac{d^2\sigma^{\bar{\nu}}}{dx dy} \right] \frac{(1-\epsilon)\pi}{y^2 G_F^2 M E_\nu} \\ &= 2xF_1[1+\epsilon R] + \frac{y(1-y/2)}{1+(1-y)^2} \Delta xF_3 \end{aligned} \quad (2)$$

where $\epsilon \simeq 2(1-y)/[1+(1-y)^2]$ is the polarization of the virtual W boson. In the above equation, we have used the relation:

$$2xF_1(x, Q^2) = F_2(x, Q^2) \frac{1+4M^2x^2/Q^2}{1+R(x, Q^2)} \quad (3)$$

where $R = \sigma_L/\sigma_T$ is the ratio of the cross-sections of longitudinally- to transversely-polarized W -bosons, Q^2 is the square of the four-momentum transfer to the nucleon, and $x = Q^2/2ME_h$ is the Bjorken scaling variable. For $x < 0.1$, R in neutrino scattering is expected to be somewhat larger than R for muon scattering because of the production of massive charm quarks in the final state for the charged current neutrino production.

Using Eq. (2) and Eq. (3), F_2 and ΔxF_3 can be extracted separately. Because of the positive correlation between R and ΔxF_3 , the extracted values of F_2 are rather insensitive to the input R . If a large input R is used, a larger value of xF_3 is extracted from the y distribution, thus yielding the same value of F_2 . In contrast, the extracted values of ΔxF_3 are sensitive to the assumed value of R , which yields a larger systematic error, shown on the data. For R the QCD-inspired fit in [3] is used, but corrected for charged current neutrino scattering using a leading order slow rescaling model. This gives precisely the same type of correction as a full NLO calculation including the massive charm quark, as shown in Ref. [4], but leads to a somewhat higher normalization than the perturbative correction. It is arguable which prescription for R leads to the better fit to existing data, but the difference between the two leads to an uncertainty in ΔxF_3 of the order of the systematic error shown. Clearly a further reduction in the assumed value of R (even down to zero), as suggested by models including $\ln(1/x)$ resummations,

would still leave a discrepancy between the lowest Q^2 data points and the theoretical predictions.

1.2. Quark Parton Model Relations

Now that we have outlined the experimental method used in the extraction of the F structure functions, it is instructive to recall the simple leading-order correspondence between the F 's and the PDF's:¹

$$\begin{aligned} F_2^{(\nu,\bar{\nu})N} &\simeq x \{u + \bar{u} + d + \bar{d} + 2s + 2c\} \\ xF_3^{(\nu,\bar{\nu})N} &\simeq x \{u - \bar{u} + d - \bar{d} \pm 2s \mp 2c\} \end{aligned} \quad (4)$$

Therefore, the combination ΔxF_3 yields:

$$\Delta xF_3 = xF_3^{\nu N} - xF_3^{\bar{\nu} N} \simeq 4x\{s - c\} \quad (5)$$

Note that since this quantity involves the parity violating structure function F_3 , this measurement has no analogue in the neutral current photon-exchange process. Also note that since, at leading-order, ΔxF_3 is directly sensitive to the strange and charm distributions, this observable can be used to probe the heavy quark PDF's, and to understand heavy quark (charm) production. We discuss these possibilities further in the following subsection.

1.3. Implications for PDF's

We have illustrated in Eq. (5) how ΔxF_3 is closely tied to the heavy quark PDF's. The question is: given the present knowledge base, should we use ΔxF_3 to determine the heavy quark PDF's, or vice versa. To answer this question, we briefly review present measurements of heavy quark PDF's, and assess their uncertainty.

1.3.1. Tevatron $W + Q$ Production

The precise measurement of W plus heavy quark ($W + Q$) events provides important information on heavy quark PDF's; additionally, such signals are a background for Higgs and squark searches.[5, 6]

Unfortunately, a primary uncertainty for $W + Q$ production comes from the heavy quark PDF's. Given that ΔxF_3 is sensitive to these heavy quark PDF's, we see at least two scenarios. One possibility is that new analysis of present data will resolve this situation prior to Run II, and provide precise distributions as an input to the Tevatron data analysis. If the situation remains unresolved, then new data from Run II may help to finally solve this puzzle. In the far future, a neutrino experiment from the neutrino factory at a linear collider would be an ideal tool to measure any neutrino structure function.[7]

¹To exhibit the basic structure, the above is taken in the limit of 4 quarks, a symmetric sea, and a vanishing Cabibbo angle. Of course, the actual analysis takes into account the full structure.[1]

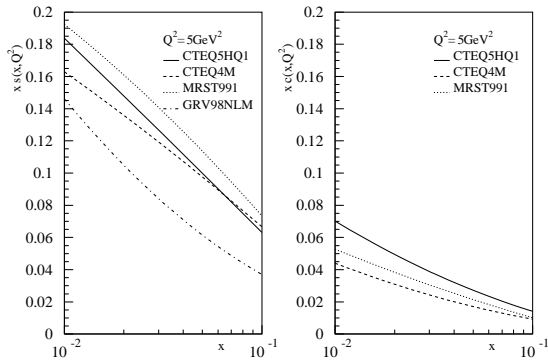


Figure 2. a) We display several strange sea PDFs in the relevant x range at $Q^2 = 5 \text{ GeV}^2$. b) Same as above for the charm sea.

1.3.2. DIS Di-muon production

The strange distribution is directly measured by dimuon production in neutrino-nucleon scattering.² The basic sub-process is $\nu s \rightarrow \mu^- c X$ with a subsequent charm decay $c \rightarrow \mu^+ X'$.

DIS dimuon measurements have safely established the breaking of the $SU(3)$ flavor symmetry,

$$\kappa \equiv \frac{\int dx x s(x)}{\int dx x [\bar{u}(x) + \bar{d}(x)]/2} \Big|_{Q^2=\mathcal{O}(10\text{GeV}^2)} \lesssim \frac{1}{2} \quad (6)$$

in the nucleon sea. Still, there remain large uncertainties for the s -quark distribution in the kinematic regions relevant for $\Delta x F_3$, even though the (x, Q) -range of the CCFR dimuon measurements [9, 13] is comparable. CCFR recorded 5044 ν_μ and 1062 $\bar{\nu}_\mu$ $\mu^\mp \mu^\pm$ events with $P_{\mu_1} \geq 9 \text{ GeV}$, $P_{\mu_2} \geq 5 \text{ GeV}$, $30 \geq E_\nu \geq 600 \text{ GeV}$, and $\langle Q^2 \rangle = 22.2 \text{ GeV}^2$. The more recent NuTeV experiment recorded a similar sample of events, and these are presently being analysed through a MC simulation based on NLO quark- and gluon-initiated corrections at differential level [10, 14, 15]. A complete NLO analysis of this data together with a global analysis may help to further constrain the strange quark distribution.

At present, PDF sets take strangeness suppression into account by imposing the constraint in Eq. (6) as $s(x) = \kappa[\bar{u}(x) + \bar{d}(x)]/2$ at the PDF input scale $Q_0 \simeq 1 \text{ GeV}$, [16, 17] or by evolving $s(x)$ from a vanishing input at a lower scale.[18] The residual uncertainty can be large, as can be seen from the collection of strange seas in Fig. 2.

In the following, we will take the uncertainty in the strange PDF into account by relaxing the experimental constraint in Eq. (6) as implemented in the MRST partons.[16] This will be discussed in Sec. [3.2].

²Presently, there are a number of LO analyses, and one NLO ACOT analysis.[8–11]. Results of a recent LO analysis by NO-MAD [12] are in line with these experiments.

1.3.3. Charged and Neutral Current DIS

The strange distribution can also be extracted indirectly using a combination of charged (F_2^ν) and neutral (F_2^μ) current structure functions; however, the systematic uncertainties involved in this procedure make an accurate determination difficult.[1] The basic idea is to use the (leading-order) relation

$$\frac{F_2^\mu}{F_2^\nu} \simeq \frac{5}{18} \left\{ 1 - \frac{3}{5} \frac{(s + \bar{s}) - (c + \bar{c}) + \dots}{q + \bar{q}} \right\} \quad (7)$$

to extract the strange distribution. Here, $q + \bar{q}$ represents a sum over all quark flavors. This method is complicated by a number of issues including the $x F_3$ component which can play a crucial role in the small- x region—precisely the region where we observe the discrepancy. From the corresponding relation:

$$\frac{5}{18} F_2^\nu - F_2^\mu \simeq \frac{1}{12} \Delta x F_3 \quad (8)$$

we see that these problems are not independent; however, this information, together with the exclusive dimuon events, may provide a more precise determination of the strange quark sea, and help to resolve our puzzle.

Prior to the DIS dimuon data, the 1992 CTEQ1 analysis [19] found that a combination of NC structure functions from NMC[20] and the physics-model-dependent charged current structure functions from CCFR [9] seem compatible with approximate $SU(3)$ -symmetry, *i.e.*, $\kappa \sim 1$ in Eq. (6). Recent dimuon measurements now exclude an $SU(3)$ -symmetry $\kappa = 1$.

We shall explore the effect of κ on $\Delta x F_3$ in Sec. [3.2].

2. Dependence of $\Delta x F_3$ on input parameters

We now systematically investigate the sensitivity of the theoretical predictions of $\Delta x F_3$ upon a variety of factors including: renormalization scheme and scale, quark mass effects, higher twist, isospin violation, and PDF uncertainties.

To simplify this analysis, we first examine the influence of these factors on the LO expression: $\Delta x F_3 = 4x\{s(\xi, \mu) - c(x, \mu)\}$; after using this as a “toy model,” we will then return to the full NLO calculation in the next section.³ For most variables, the simplified LO is sufficient to display the general behavior of the full NLO result. There are two exceptions: 1) the scheme dependence, and 2) the PDF dependence. These factors depend on the interplay of both the quark-initiated LO contributions as well as the NLO gluon-initiated contributions. For this reason, we will postpone discussion of these effects until the following section.

2.1. Charm Mass

We start by examining the effect of the charm mass, m_c , on $\Delta x F_3$. In Fig. 3, we plot the leading-order ³Note that a possible $s \neq \bar{s}$ asymmetry [21] would average out in $\Delta x F_3$.

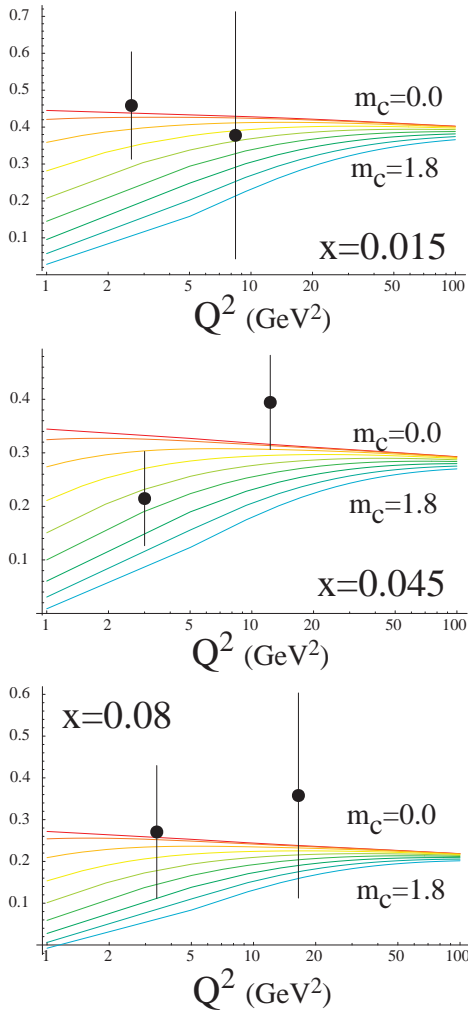


Figure 3. Variation of the leading-order expression $\Delta xF_3 = 4x\{s(\xi, \mu) - c(x, \mu)\}$ on the charm mass, m_c , plotted *vs.* μ . We take values of m_c in the range $m_c = [0, 1.8]$ GeV using steps of 0.2 GeV for three choices of x . We use $\xi = x(1 + m_c^2/Q^2)$, and $\mu^2 = (Q^2 + 1.6^2 \text{ GeV}^2)$.

expression $\Delta xF_3 = 4x\{s(\xi, \mu) - c(x, \mu)\}$ *vs.* Q^2 for values of m_c in the range $= [0, 1.8]$ GeV using steps of 0.2 GeV. Here, we define $\xi = x(1 + m_c^2/Q^2)$ which is a “slow-rescaling” type of correction[22,23] which (crudely) includes mass effects by shifting the x variable. “Slow rescaling” naturally arises at LO from the single-charm threshold condition ($W > m_c$) and is required at NLO for consistent mass factorization [23, 24].

Note, the result of this correction is most significant at low Q^2 . To isolate the kinematic m_c dependence, we have chosen the scale $\mu^2 = Q^2 + 1.6^2 \text{ GeV}^2$. (We will separately vary the scale in the following subsection.)

Note that this exercise is only altering the charm mass in one aspect of the calculation; to be entirely consistent it would be necessary to obtain parton dis-

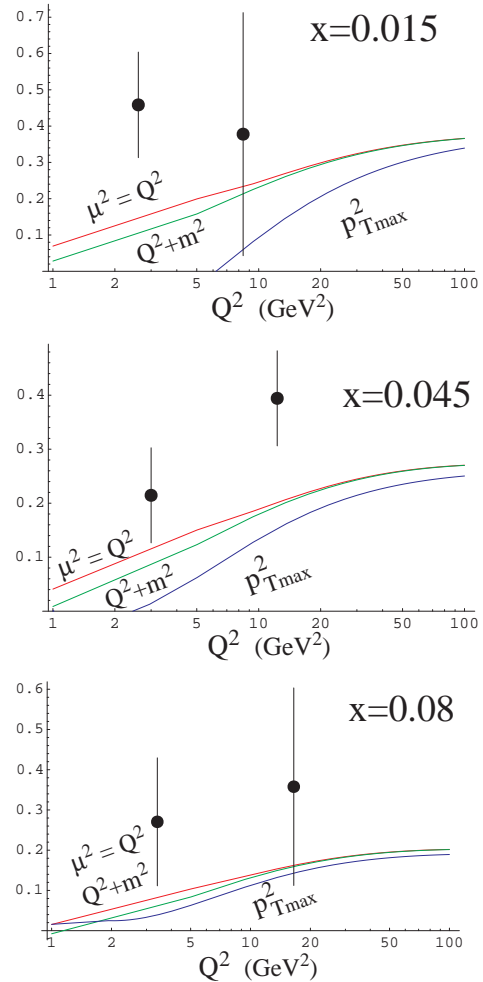


Figure 4. Variation of the leading-order ΔxF_3 on the renormalization scale, μ . We have chosen $\mu^2 = \{Q^2, Q^2 + m_c^2, P_{Tmax}^2\}$.

tributions (particularly the charm quark) using fits with different charm masses. For charm masses in the range 1.2 to 1.8 GeV, this will be a small effect; for charm masses below ~ 1.2 GeV, (which is below the experimentally allowed range[25,26]), such issues become important and the curves of Fig. 3 will be modified, i.e. lower m_c will lead to longer evolution for charm, a larger charm distribution, and a lowering of the curves in Fig. 3.

While taking $m_c \rightarrow 0$ does raise the theoretical curves in the regions where we observe the discrepancy (namely, the low Q region), varying m_c even within the wide range of $[1.2, 1.8]$ GeV (lower 4 curves) does not give us sufficient flexibility to match either the shape or normalization of the data.

2.2. μ Scale

Next, we investigate the variation of $\Delta xF_3 = 4x\{s(\xi, \mu) - c(x, \mu)\}$ with the renormalization/factorization scale μ . We use three choices of

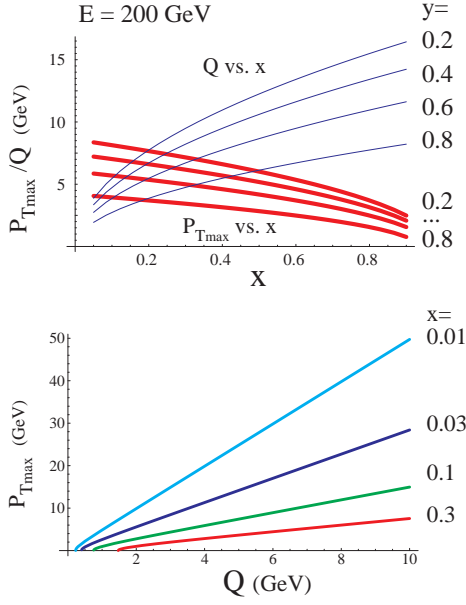


Figure 5. a) We display P_T^{max} and Q vs. x for $E=200$ GeV and $y = \{0.2, 0.4, 0.6, 0.8\}$. b) P_T^{max} vs. Q for $x = \{0.3, 0.1, 0.03, 0.01\}$.

the μ^2 scale:

- $\mu^2 = Q^2$,
- $\mu^2 = Q^2 + m_c^2$,
- $\mu^2 = P_{T_{max}}^2$.

Of course, ΔxF_3 with μ^2 scales of Q^2 and $Q^2 + m_c^2$ differ only at lower values of Q^2 ; ΔxF_3 with μ^2 scale of $P_{T_{max}}^2$ is comparable to Q^2 and $Q^2 + m_c^2$ at larger x , but lies below for smaller x . The scale choice $Q^2 + m_c^2$ leads to an improvement over Q^2 by providing a lower bound on μ to keep the scale in the perturbative region.

The choice of $P_{T_{max}}$ is motivated, in part, by some observations by Collins.[27] To display the relationship between $P_{T_{max}}$ and Q , in Fig. 5 we plot both $P_{T_{max}}$ and Q vs. x for 4 choices of y ; note[13] that the x -dependence of $P_{T_{max}}$ is opposite that of Q . While tying the scale choice of μ to $P_{T_{max}}$ has some interesting intuitive interpretations,⁴ for small x this scale clearly becomes too large for the relevant physics. In any case, it cannot help us with our ΔxF_3 problem as the scale choice of $P_{T_{max}}$ moves the theory curves *away* from the data.

2.3. Higher Twist

We now illustrate the potential effects due to higher twist contributions. We parameterize such contributions by multiplying the leading-twist terms by a cor-

⁴Technically, the interpretation is in terms of the characteristic P_T of the partonic subprocess, but this is unobservable; therefore $P_{T_{max}}$ is used instead.[27]

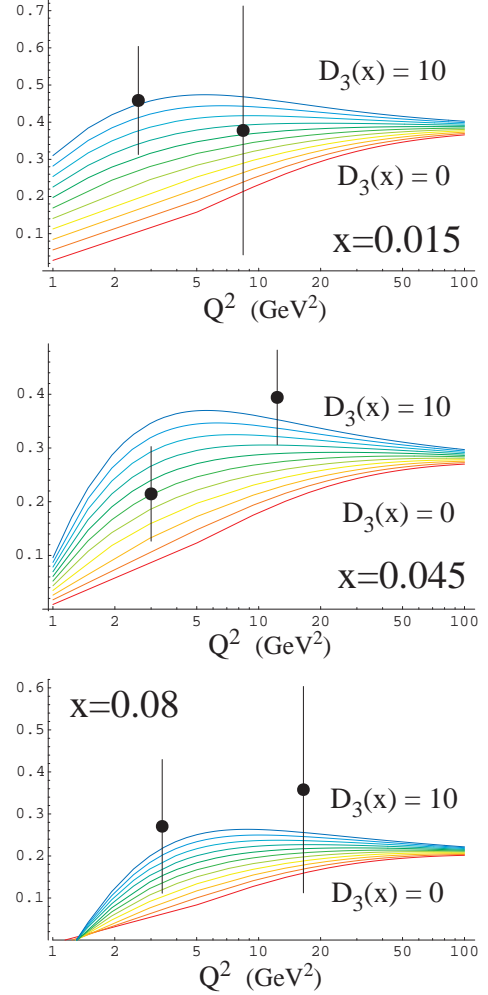


Figure 6. Variation of the LO ΔxF_3 on higher twist. The functional form is $(1 + D_3(x)/Q^2)$ where we take $D_3(x)$ in the range $D_3(x) = [0, 10]$ GeV^2 in steps of 1 GeV^2 .

rection factor that will increase at low Q^2 . More specifically we use:[28]

$$F_i^{HT}(x, Q^2) = F_i^{LT}(x, Q^2) \left(1 + \frac{D_i(x)}{Q^2} \right) \quad (9)$$

where $i = \{1, 2, 3\}$. We vary $D_3(x)$ over the range $[0, 10]$ GeV^2 in steps of 1 GeV^2 . (For the purposes of our simple illustration, it is sufficient to take $D_3(x)$ to be independent of x .) We will find that this range is well beyond what is allowed by experiment; however, we display this exaggerated range to make the effect of the higher twist contributions evident.

To normalize this choice with the allowable range consistent with data, we compare with the MRST higher twist fit which extracted a limit on the function $D_2(x)$. We see from the table of $D_2(x)$ that the allowed contribution from the higher twist terms is quite small. In addition, we note that the sign for $D_2(x)$ obtained

for the relevant small- x region tends to be negative—exactly the opposite sign that is needed to move the theory toward the data.

As we have shown constraints on $D_2(x)$, the obvious question is should we expect $D_3(x)$ to be substantially different? A calculation of the power corrections using renormalons[29] suggests that while both $D_1(x)$ and $D_3(x)$ are of the same order of magnitude as $D_2(x)$ and have similar x -dependence, that $D_1(x)$ and $D_3(x)$ are even more negative than $D_2(x)$; again, this trend would move the theory farther from the data. Therefore, we conclude that that $D_2(x)$ represents a conservative limit for the $D_3(x)$ power corrections.

Examining Fig. 6, we note that it would take an enormous higher twist contribution to bring the normalization of the curves in the range of the data points, and still the shape of the Q^2 -dependence is not well matched; hence, we conclude that this is not a compelling solution.

x	LO	NLO	NNLO
0 – 0.0005	-0.4754	0.0116	-0.0061
0.0005 – 0.005	-0.2512	-0.0475	0.0437
0.005 – 0.01	-0.2481	-0.1376	-0.0048
0.01 – 0.06	-0.2306	-0.1271	-0.0359
0.06 – 0.1	-0.1373	-0.0321	0.0167
0.1 – 0.2	-0.1263	-0.0361	0.0075
0.2 – 0.3	-0.1210	-0.0893	-0.0201
0.3 – 0.4	-0.0909	-0.1710	-0.1170
0.4 – 0.5	0.1788	-0.0804	-0.0782
0.5 – 0.6	0.8329	0.3056	0.1936
0.6 – 0.7	2.544	1.621	1.263
0.7 – 0.8	6.914	5.468	4.557
0.8 – 0.9	19.92	18.03	15.38

Table 1

Values of the higher-twist coefficient $D_2(x)$ extracted from the LO, NLO and NNLO fits. Units are GeV^2 . Table taken from Ref. [30].

2.4. Isospin Violations

The naive parton model identity in Eq. (5) is modified if the full (non-diagonal) CKM structure, NLO QCD radiative corrections, and QCD based charm production are taken into account. This expression is also modified *even in leading-order* if we have a violation of exact $p \leftrightarrow n$ isospin-symmetry, (or charge symmetry); *e.g.*, $u_n(x) \not\equiv d_p(x)$. In deriving Eq. (5), isospin-symmetry was necessary to guarantee that the u, d -contributions cancel out in the difference, thereby leaving only the $s(x)$ and $c(x)$ contributions.⁵

⁵Note we have not investigated shadowing corrections. The CCFR data on Fe is converted with an isoscalar correction, and

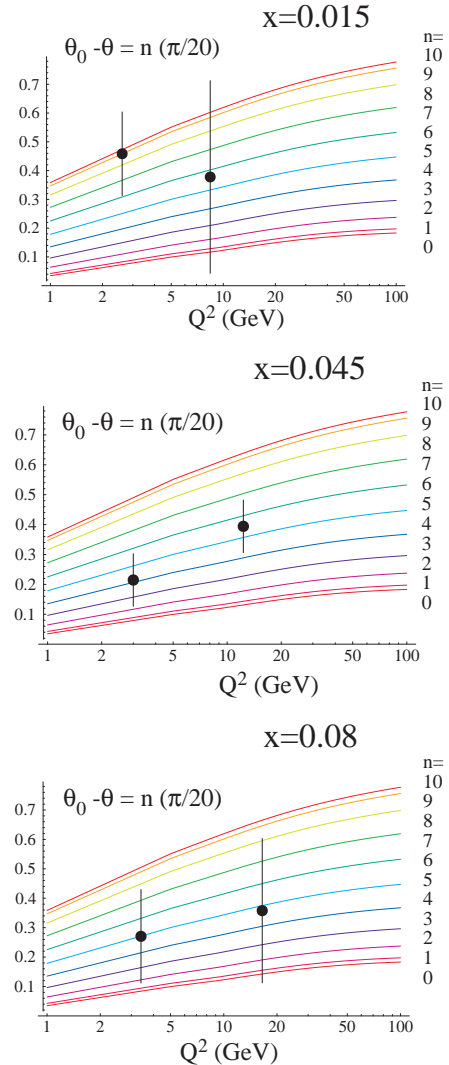


Figure 7. Impact of a one-parametric charge symmetry violation (CSV) toy model of Eq. (10) and Eq. (11) on the LO ΔxF_3 of Eq. (5). θ is varied over the range $\theta_0 - \theta = [0, \pi/2]$ in steps of $(\pi/20)$.

The validity of exact *charge symmetry* (CS) has recently been reexamined.[31] Residual u, d -contributions to ΔxF_3 from *charge symmetry violation* (CSV) would be amplified due to enhanced valence components $\{u_v(x), d_v(x)\}$, and because the $d \rightarrow u$ transitions are not subject to slow-rescaling corrections which strongly suppress the $s \rightarrow c$ contribution to ΔxF_3 , as illustrated in Fig. 3.

We will examine possible contributions to ΔxF_3 by considering a “toy” model to parameterize the CSVs. In isospin space, we can parameterize a general transformation as a rotation:

$$|q\rangle_n^{\text{CSV}} = N_q \sum_{q'} R_{qq'}(\theta) |q'\rangle_p \quad (10)$$

the corresponding uncertainties are included in the data.[2]

where R is a rotation matrix, and N_q is the normalization factor. For example, in this model the u -distribution in the neutron would be related to the proton distributions via the relation:

$$u_n^{\text{CSV}}(x, Q^2) = N_u^2 [\cos^2(\theta)u_p(x, Q^2) + \sin^2(\theta)d_p(x, Q^2)] \quad (11)$$

For $\theta = \pi/2$, we recover the symmetric limit $u_p(x, Q^2) = d_n(x, Q^2)$. Note that we define the normalization N_u such that this model preserves the sum rule:

$$\int dx u_n^{\text{CSV}}(x, Q^2) - \bar{u}_n^{\text{CSV}}(x, Q^2) = 1 \quad (12)$$

Note that Eq. (10) should not be considered a serious theory,⁶ but rather a simple one-parameter (θ) model which is flexible enough to illustrate a range of CSV effects for ΔxF_3 .

In Fig. 7 we vary θ over the its maximum range $[0, \pi/2]$ in steps of $\pi/20$. The exact charge symmetry limit ($\theta_0 = \pi/2$) of Eq. (11) corresponds to the lowest curve in Fig. 7. Note the effect of the CSV contribution monotonically increases as θ deviates from the charge symmetry (CS) limit $\theta_0 = \pi/2$. From the plot we observe that a violation of $\theta_0 - \theta \gtrsim (3/10)(\pi/2)$ is required to bring the theory in the neighborhood of the data.

At the relevant x values of interest for ΔxF_3 , this translates (via Eq. (11) and its analogue for d_n) into a $\sim 20\%$ symmetry violation for u_n , and a $\sim 10\%$ symmetry violation for d_n . Specifically,

$$\frac{d_p(x)}{u_n^{\text{CSV}}(x)} \simeq 1.2 \quad (13)$$

$$\frac{u_p(x)}{d_n^{\text{CSV}}(x)} \simeq 0.9 \quad (14)$$

Since the electroweak couplings are flavor-independent, ($V_q = A_q = 1 \forall q$), ΔxF_3 is in principle insensitive to a re-shuffling of the CSV contributions between u_n and d_n . In particular, if we define the shift due to CSV in ΔxF_3 as $\delta\Delta xF_3$, we find:

$$\delta\Delta xF_3 \equiv \frac{x}{4}(\delta d + \delta\bar{d} - \delta u - \delta\bar{u}) \quad (15)$$

where

$$\delta u(x, Q^2) \equiv u_n^{\text{CSV}}(x, Q^2) - d_p(x, Q^2) \quad (16)$$

The expression for $\delta\Delta xF_3(x, Q^2)$ of Eq. (15) is evaluated at $Q^2 = 4 \text{ GeV}^2$ from Eq. (11) with $\theta_0 - \theta = (3/10)(\pi/2)$ and is plotted in Fig. 8.⁷

Although this level of isospin violation certainly improves the description of ΔxF_3 , it is necessary to consider precisely what level of violation is actually allowed by other experimental data. For instance, it

⁶ *E.g.*, it does strictly speaking not commute with evolution.

⁷ Again, the detailed x -shape from Eq. (11) should *not* be taken as a serious model prediction.

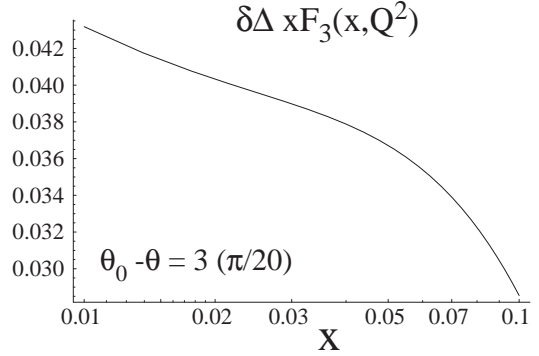


Figure 8. $\delta\Delta xF_3(x, Q^2)$ with $Q^2 = 4 \text{ GeV}^2$ as defined in Eq. (15) and Eq. (11) for $\theta_0 - \theta = 3(\pi/20)$.

has previously been suggested that the discrepancy between F_2 from neutrino and muon data itself may be due to isospin violation.[31] This type of violation required that \bar{u} be very different to \bar{d} in the region of interest.

The measurement of the lepton charge asymmetry in W decays from the Tevatron [32, 33] places tight constraints on the up and down quark distributions in the range $0.007 < x < 0.24$, constraining them to be approximately as specified in the parton sets obtained by the global analyses. While only strictly telling us about parton distributions in the proton, this data rules out an isospin violation of this type to about 5% as demonstrated in [33].

However, there are other strong constraints on isospin violation. For example, we note that while the toy model above leaves the neutron singlet combination $q + \bar{q}$ invariant at the $\lesssim 2\%$ level in the region $x \in [0.01; 0.1]$, it would lower the NC observable:

$$\left[\frac{4}{9}(u + \bar{u}) + \frac{1}{9}(d + \bar{d}) \right]_n \Big|_{x \in [0.01; 0.1]} \quad (17)$$

by about 10%. An effect of this size would definitely be visible in NMC F_2^n/F_2^p data which has an uncertainty of order a few percent in this kinematic region, and acts as a major constraint.[20]

At this point, one could play clever games to evade the constraints of specific experiments. For example, a re-shuffling of CSV contributions between the individual $\delta q, \delta\bar{q}$ in Eq. (15) according to:

$$\begin{aligned} x u_n &= x d_p - \frac{2}{5} \delta\Delta xF_3 \\ x d_n &= x u_p + \frac{8}{5} \delta\Delta xF_3 \\ x \bar{q}_n &: \text{analogous} \end{aligned} \quad (18)$$

would keep Eq. (17) invariant. However, this would in turn raise the CC observable Eq. (15) by $\gtrsim 5\%$; though it would help to explain the excess in δxF_3 , it would spoil the new-found compatibility between neutral current and charged current data.

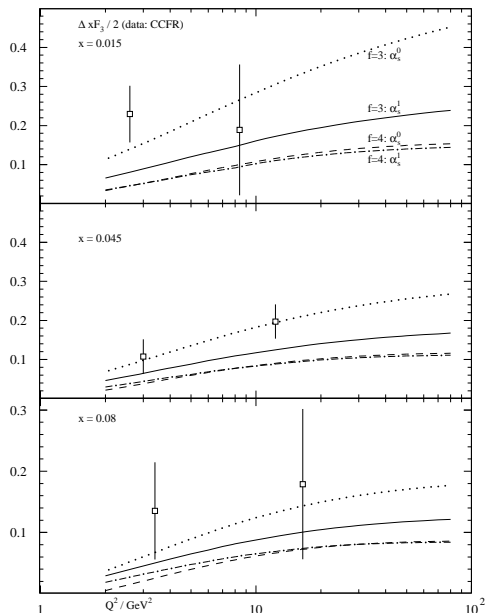


Figure 9. LO and NLO calculations for ΔxF_3 vs. Q^2 in 3 and 4 flavor schemes for three values of x . The dotted curve ($f=3, \alpha_s^0$) corresponds to the 3-flavor LO result. The dashed curve ($f=4, \alpha_s^0$) corresponds to the 4-flavor LO result. The solid and dot-dash curves correspond to the 3 and 4 flavor NLO QCD calculations, respectively.

In addition, there are also fixed-target Drell-Yan experiments[34, 35] such as NA51 and E866 which precisely measure d/\bar{u} in the range[35] $0.04 < x < 0.27$, and are also sensitive to isospin violating effects.

We therefore conclude that the many precise data sets which constrain different combinations of the PDF's probably leave no room for CSV contributions of the magnitude necessary to fully align the theory curves with the ΔxF_3 data. However, we should add the caveat that an exhaustive investigation of the interplay of these different data sets and their influence on ΔxF_3 will only be possible within a global PDF analysis

3. NLO calculation of $\Delta xF_3/2$

Now that we have used the leading-order expression for ΔxF_3 to systematically investigate dependence of this observable on various parameters, in this section we now turn to the full NLO calculation.

3.1. Contributions to the NLO Calculation

In Fig. 9, we have plotted the LO and NLO calculations for ΔxF_3 vs. Q^2 on an isoscalar target in 3 and 4 flavor schemes. The 3-flavor LO calculation ($f=3$,

α_s^0) involves primarily the strange quark contribution, $s(x)$, as the charm distribution is excluded in this case. When the higher order terms are included ($f=3, \alpha_s^1$), this result moves (substantially) toward the predictions of the 4-flavor scheme. We note that while the 3-flavor LO calculation ($f=3, \alpha_s^0$) appears consistent with the data, we cannot take this result as a precise theoretical prediction as this simplistic result is highly dependent on scheme and scale choices; a result that is verified by the large shift in going from LO to NLO.

The pair of curves in the 4-flavor scheme (using the CTEQ4HQ distributions) nicely illustrates how the charm distribution $c(x, \mu^2)$ evolves as $\ln(Q^2/m_c^2)$ for increasing Q^2 ; note, $c(x, \mu^2)$ enters with a negative sign so that the 4-flavor result is below the 3-flavor curve. For the scale choice, we take $\mu = \sqrt{Q^2 + m_c^2}$. While the scale choice $\mu = Q$ is useful for instructive purposes such as demonstrating the matching of the 3- and 4-flavor calculations at $\mu = Q = m_c$, the choice $\mu = \sqrt{Q^2 + m_c^2}$ is more practical as it provides a lower bound on μ which is important for the PDF's and $\alpha_s(\mu)$. (Cf., Sec. [3.3], and Ref. [36].)

Additionally we note the stability of the 4-flavor scheme in contrast to the 3-flavor scheme. The shift of the curves when including the NLO contributions is quite minimal, particularly when compared with the 3-flavor result.[37, 38] This suggests that organizing the calculation to include the charm quark as a proton constituent can be advantageous even at relatively low values of the energy scale.

3.2. PDF Uncertainties: $s(x)$, ...

In Fig. 10, we show the variation of NLO calculation of ΔxF_3 on the strange-quark PDF. To obtain a realistic assessment of the $s(x)$ dependence, we have use the NLO calculation with PDF's based on the MRST set which are re-fit with the value of $\kappa = 2s/(\bar{u} + \bar{d})$ constrained to be $\kappa = \{0.50, 0.78, 1.00\}$. Note, by re-fitting the PDF's with the chosen value of κ we are assured to have an internally consistent set of PDF's with appropriate matching between the quarks and gluon, and with the sum-rules satisfied.⁸

The choice $\kappa = 0.50$ is in line with the many experimental determinations of κ , cf., Table. 8; as expected, this prediction lies farthest from the data points.

The choice $\kappa = 0.78$ is taken as an extreme upper limit given the experimental constraints; actually, in light of the results of Table. 8, this is arguably *beyond* present experimental bounds. This prediction is marginally consistent at the outer reach of the systematic + statistical error bars.

Finally, we take an $SU(3)$ symmetric set ($\kappa = 1$)

⁸Note, κ is certainly Q -dependent, and the values for κ quoted above correspond to the Q_0 of the evolution. While κ compares the integral of $s(x)$ to the sea-quarks, there is also the possibility of an x -dependent variation. This has been studied in the fits of the strange-sea[9, 10, 13, 39]; we shall find that such subtle effects can play no role in resolving the ΔxF_3 issue.

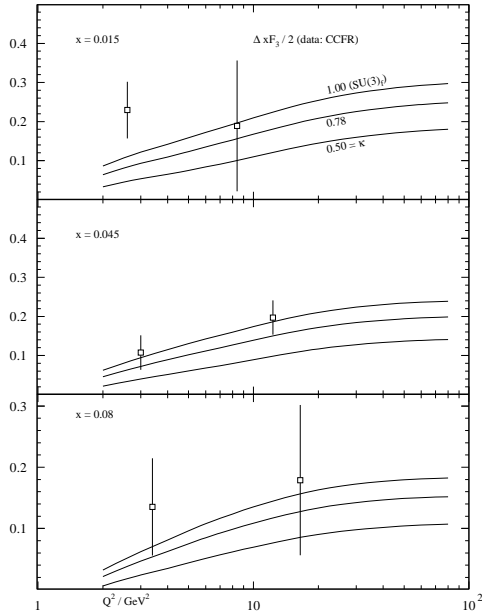


Figure 10. Variation of ΔxF_3 on the strange-quark PDF. The NLO calculation is used with three sets of PDF's. These PDF sets are re-fit based on the MRST set, and the value of κ is constrained to be $\kappa = \{0.50, 0.78, 1.00\}$.

purely for illustrative purposes. It is interesting to note that even this extreme value is still below the central value of the data points at the higher x values.

In conclusion we note that increasing the strange quark distribution does succeed in moving the theory toward the data; however, our consistent NLO analysis presented here suggests that we have only limited freedom to increase $s(x)$, and that this alone is not sufficient to obtain good agreement between theory and data.

3.3. Scheme Choice

In our final section, we present the best theoretical predictions presently available to demonstrate the scheme dependence of $\Delta xF_3(x, Q)$. Specifically, in Fig. 11 we show predictions for:

- NLO FFS GRV.[18]
- NLO VFS TR calculation.[41]
- NLO VFS ACOT calculation with CTEQ4 PDF's.[23, 42, 43]
- NLO VFS ACOT calculation with CTEQ5 PDF's.[17, 23, 42]

All these calculations use NLO matrix elements, and are matched with appropriate global PDF's which are

Experiment	Order	κ	Ref.
CDHS	LO	$0.47 \pm 0.08 \pm 0.05$	[8]
FMMF	LO	$0.41^{+0.075}_{-0.075} {}^{+0.103}_{-0.069}$	[40]
CHARM II	LO	$0.39^{+0.07}_{-0.06} {}^{+0.07}_{-0.07}$	[11]
CCFR*	NLO	$0.477^{+0.046}_{-0.044} {}^{+0.023}_{-0.024}$	[13]
CCFR†	NLO	$0.468^{+0.061}_{-0.046} {}^{+0.024}_{-0.025}$	[13]
CCFR†	LO	$0.373^{+0.048}_{-0.041} \pm 0.018$	[9]
NOMAD	LO	$0.48^{+0.09}_{-0.07} {}^{+0.17}_{-0.12}$	[12]
NuTeV	LO	$0.42 \pm 0.07 \pm 0.06$	[10]

Table 2

Next-to-leading-order and leading-order fit results. Errors are statistical and systematic. This table is displayed to *estimate* the upper limits allowed by experiment; a full comparison must take into account scheme and scale choices, and the shape parameters.

*Collins-Spiller Fragmentation.

† Peterson Fragmentation.

fitted in the proper scheme.

The first observation we make is how closely these four predictions match, especially given the wide variation displayed in previous plots such as Fig. 9. In hindsight, this result is simply a consequence of the fact that while different renormalization schemes can produce different results, this difference can only be higher order.⁹ Thus, the difference between these curves is indicative of terms of order α_s^2 which have yet to be calculated.¹⁰ When terms of order α_s^N are included, the span of these predictions will be systematically reduced to order α_s^{N+1} .

In Fig. 11, we note the very close agreement among the VFS calculations, particularly the TR calculation and the ACOT calculation with CTEQ4 PDF's. The ACOT calculation with the two CTEQ curves show primarily the effect of the charm distribution, as CTEQ4 uses $m_c = 1.6$ and CTEQ5 uses $m_c = 1.3$. The GRV calculation shows the effect of using yet a different scheme, in this case a FFN scheme, with its appropriately matched PDF. Were we to use MRST or CTEQ PDF's, the spread of these theory curves would decrease; however, this would most likely represent an underestimate of the true theoretical uncertainties arising from both the hard cross section and PDF's.¹¹

While we consider it a triumph of QCD that different schemes truly yield comparable results (higher order terms aside), we should be cautious and note that the spread of these curves can only underestimate the true theoretical uncertainty. Note that GRV has a

⁹To be precise, different renormalization schemes can differ by i) terms of higher order in the perturbation series, and ii) terms of higher twist which do not factorize.[44, 45]

¹⁰For asymptotic results at order α_s^2 , see Ref. [37].

¹¹The computation of PDF errors is a complex subject. For some recent approaches to this topic see: Refs.[16, 17, 39, 46].

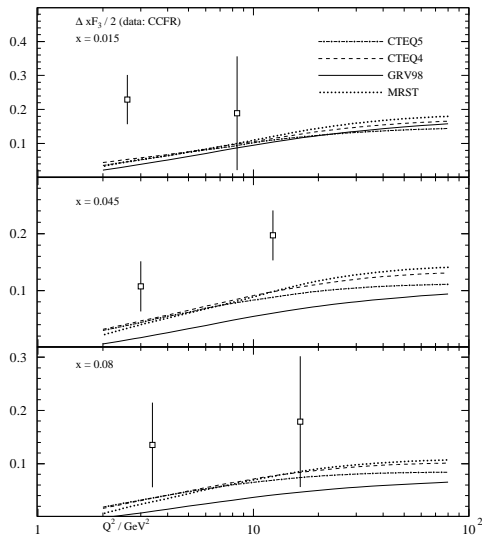


Figure 11. Variation of ΔxF_3 on the renormalization scheme. All curves use NLO calculations, and appropriately matched PDF's. We note that the theoretical predictions are consistent within the theoretical uncertainty—as they should be.

rather different strange distribution due to a different philosophy of obtaining this distribution rather than due to a different scheme.

4. Conclusions and Outlook

Comprehensive analysis of the neutrino data sets can provide incisive tests of the theoretical methods, particularly in the low Q^2 regime, and enable precise predictions that will facilitate new particle searches by constraining the PDF's. This document serves as a progress report, and work on these topics will continue in the future.

Theoretical predictions for ΔxF_3 undershoot preliminary fixed target data at the $\sim 1\sigma$ -level at low x and Q . The neutrino structure function ΔxF_3 is obviously sensitive[47] to the strange sea of the nucleon and the details of deep inelastic charm production. A closer inspection reveals, however, considerable dependence upon factors such as the charm mass, factorization scale, higher twists, contributions from longitudinal W^\pm polarization states, nuclear shadowing, charge symmetry violation, and the PDF's. This makes $\Delta xF_3(x, Q^2)$ an excellent tool to probe both perturbative and non-perturbative QCD.

We have explored the variation of $\Delta xF_3(x, Q^2)$ on the above factors and found none of these to be capa-

ble of resolving the discrepancy between the data and theory.

Although we have not eliminated the possibility that the entire set of parameters conspires to align the theory with the data, we have demonstrated this to be extremely unlikely. Of course, a definitive answer can only be obtained by a global analysis which combines the neutrino data for dimuons, ΔxF_3 , $F_2^\nu(PMI)$, and $F_2^{\mu,e}$.

As the situation stands now, this $\Delta xF_3(x, Q^2)$ puzzle poses an important challenge to our understanding of QCD and the related nuclear processes in an important kinematic region. The resolution of this puzzle is important for future data analysis, and the solution is sure to be enlightening, and allow us to expand the applicable regime of the QCD theory.

Acknowledgments

This work is supported by the Royal Society, the U.S. Department of Energy, the National Science Foundation, the Lightner-Sams Foundation, and the 'Bundesministerium für Bildung, Wissenschaft, Forschung und Technologie', Bonn.

We thank J. Bluemlein, A. Bodek, J. Conrad, J. Morfin, S. Kuhlmann, R.G. Roberts, H. Schellman, M. Shaevitz, J. Smith, and W.-K. Tung, for valuable discussions.

REFERENCES

1. U. K. Yang *et al.* [CCFR-NuTeV Collaboration], hep-ex/0009041.
2. W. G. Seligman *et al.*, CCFR Collab., Phys. Rev. Lett. **79**, 1213 (1997); W. G. Seligman, Ph.D. Thesis, Columbia University, Nevis-292 (1997).
3. L. W. Whitlow, E. M. Riordan, S. Dasu, S. Rock and A. Bodek, Phys. Lett. **B282**, 475 (1992).
4. C. Boros, F.M. Steffens, J.T. Londergan and A.W. Thomas, Phys. Lett. **B468**, 161 (1999).
5. W. T. Giele, S. Keller and E. Laenen, Nucl. Phys. Proc. Suppl. **51C**, 255 (1996); Phys. Lett. **B372**, 141 (1996); hep-ph/9408325.
6. R. Demina *et al.*, hep-ph/0005112. L. de Barbaro *et al.*, hep-ph/0006300.
7. R.D. Ball, D.A. Harris, and K.S. McFarland, hep-ph/0009223.
8. H. Abramowicz *et al.*, CDHSW Collab., Z. Phys. **C15**, 19 (1982).
9. S.A. Rabinowitz *et al.*, CCFR Collab., Phys. Rev. Lett. **70**, 134 (1993).
10. T. Adams *et al.* [NuTeV Collaboration], hep-ex/9906037.
11. P. Vilain *et al.* [CHARM II Collaboration], Eur. Phys. J. **C11**, 19 (1999).
12. P. Astier *et al.* [NOMAD Collaboration], Phys. Lett. **B486** (1990) 35.

13. A. O. Bazarko *et al.* [CCFR Collaboration], *Z. Phys.* **C65**, 189 (1995); A. O. Bazarko, Ph.D. Thesis. NEVIS-1504
14. M. Glück, S. Kretzer and E. Reya, *Phys. Lett.* **B398**, 381 (1997); **B405**, 392 (1997) (E).
15. S. Kretzer and F. Olness, in preparation.
16. A.D. Martin, R.G. Roberts, W.J. Stirling, R.S. Thorne, *Eur. Phys. J.* **C14**, 133 (2000).
17. H. L. Lai *et al.*, *Eur. Phys. J.* **C12**, 375 (2000).
18. M. Gluck, E. Reya and A. Vogt, *Eur. Phys. J.* **C5**, 461 (1998).
19. J. Botts *et al.*, CTEQ1, *Phys. Lett.* **B304**, 159 (1993).
20. NMCollaboration: M. Arneodo *et al.*, *Nucl. Phys.* **B483**, 3 (1997).
21. S.J. Brodsky and B.-Q. Ma, *Phys. Lett.* **B381**, 317 (1996).
22. R.M. Barnett, *Phys. Rev. Lett.* **36**, 1163 (1976); H. Georgi and H.D. Politzer, *Phys. Rev.* **D14**, 1829 (1976).
23. F. Olness, W.K. Tung, *Nucl. Phys.* **B308** (1988) 813; M. Aivazis, F. Olness, W.K. Tung, *Phys. Rev.* **D50** (1994) 3085; M. Aivazis, J.C. Collins, F. Olness, W.K. Tung, *Phys. Rev.* **D50** (1994) 3102.
24. T. Gottschalk, *Phys. Rev.* **D23**, 56 (1981); M. Glück, S. Kretzer and E. Reya, *Phys. Lett.* **B380**, 171 (1996); **B405**, 391 (1996) (E).
25. J. Breitweg *et al.* [ZEUS Collaboration], *Eur. Phys. J.* **C12**, 35 (2000); C. Coldewey [H1 and ZEUS Collaborations], *Nucl. Phys. Proc. Suppl.* **74**, 209 (1999).
26. D.E. Groom *et al.*, *The European Physical Journal* **15** (2000) 1, available on the PDG WWW pages (URL: <http://pdg.lbl.gov/>).
27. J. C. Collins, ANL-HEP-CP-90-58 *Proc. of 25th Rencontre de Moriond: High Energy Hadronic Interactions, Les Arcs, France, Mar 11-17, 1990*.
28. A. D. Martin, R. G. Roberts, W. J. Stirling and R. S. Thorne, *Phys. Lett.* **B443**, 301 (1998).
29. M. Dasgupta and B. R. Webber, *Phys. Lett.* **B382**, 273 (1996).
30. A. D. Martin, R. G. Roberts, W. J. Stirling and R. S. Thorne, *Eur. Phys. J.* **C18**, 117 (2000).
31. C. Boros, F.M. Steffens, J.T. Londergan and A.W. Thomas, *Phys. Lett.* **B468**, 161 (1999); C. Boros, J.T. Londergan, A.W. Thomas, *Phys. Rev.* **D59**, 074021 (1999); C. Boros, J.T. Londergan, A.W. Thomas, *Phys. Rev. Lett.* **81**, 4075 (1998) and related references therein.
32. F. Abe *et al.* [CDF Collaboration], *Phys. Rev. Lett.* **81**, 5754 (1998).
33. A. Bodek, Q. Fan, M. Lancaster, K.S. McFarland and U.K. Yang, *Phys. Rev. Lett.* **83**, 2892 (1999).
34. A. Baldit *et al.* [NA51 Collaboration], *Phys. Lett.* **B332**, 244 (1994).
35. E. A. Hawker *et al.* [FNAL E866/NuSea Collaboration], *Phys. Rev. Lett.* **80** (1998) 3715.
36. C. Schmidt, hep-ph/9706496; J. Amundson, C. Schmidt, W. K. Tung, X. Wang, *JHEP***0010**, 031 (2000); J. Amundson, F. Olness, C. Schmidt, W. K. Tung, X. Wang, FERMILAB-CONF-98-153-T, Jul 1998.
37. M. Buza and W.L. van Neerven, NPB 500, 301 (1997). M. Buza, Y. Matiounine, J. Smith, R. Migneron and W. L. van Neerven, *Nucl. Phys.* **B472**, 611 (1996).
38. A. Chuvakin, J. Smith, W.L. van Neerven, *Phys. Rev. D* **61**, 096004 (2000); *Phys. Rev. D* **62**, 036004 (2000); A. Chuvakin, J. Smith, hep-ph/9911504.
39. V. Barone, C. Pascaud and F. Zomer, *Eur. Phys. J.* **C12**, 243 (2000).
40. B. Strongin *et al.*, *Phys. Rev.* **D43**, 2778 (1991).
41. R.S. Thorne, R.G. Roberts, hep-ph/0010344; *Phys. Rev.* **D57** (1998) 6871.
42. S. Kretzer, I. Schienbein *Phys. Rev.* **D56** (1997) 1804; *Phys. Rev.* **D58** (1998) 094035; *Phys. Rev.* **D59** (1999) 054004.
43. H. L. Lai *et al.*, *Phys. Rev.* **D55** (1997) 1280.
44. J.C. Collins, W.-K. Tung *Nucl. Phys.* **B278**, 934 (1986); J.C. Collins, *Phys. Rev.* **D58** (1998) 094002.
45. M. Krämer, F. Olness, D. Soper, *Phys. Rev. D* **62**, 096007 (2000);
46. S. Alekhin, *Eur. Phys. J.* **C10**, 395 (1999); W. T. Giele and S. Keller, *Phys. Rev. D* **58**, 094023 (1998); W. T. Giele, S. Keller and D. A. Kosower, "Parton distributions with errors," *In *La Thuile 1999, Results and perspectives in particle physics* 255-261*; J. Pumplin, D. R. Stump and W. K. Tung, hep-ph/0008191.
47. V. Barone, U. D'Alesio and M. Genovese, *Phys. Lett.* **B357**, 435 (1995).