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2 **Boron isotopic composition of olivine-hosted melt inclusions from Gorgona komatiites,**  
3 **Colombia: new evidence supporting wet komatiite origin**

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29

30 **Abstract** – A fundamental question in the genesis of komatiites is whether these rocks originate  
31 from partial melting of dry and hot mantle, 400–500°C hotter than typical sources of MORB and  
32 OIB magmas, or if they were produced by hydrous melting of the source at much lower  
33 temperatures, similar or only moderately higher than those known today. Gorgona Island,  
34 Colombia, is a unique place where Phanerozoic komatiites occur and whose origin is directly  
35 connected to the formation of the Caribbean Large Igneous Province. The genesis of Gorgona  
36 komatiites remains controversial, mostly because of the uncertain origin of volatile components  
37 which they appear to contain. These volatiles could equally result from shallow level magma  
38 contamination, melting of a “damp” mantle or fluid-induced partial melting of the source due to  
39 devolatilization of the ancient subducting plate. We have analyzed boron isotopes of olivine-  
40 hosted melt inclusions from the Gorgona komatiites. These inclusions are characterized by  
41 relatively high contents of volatile components and boron (0.2–1.0 wt.% H<sub>2</sub>O, 0.05–0.08 wt.%  
42 S, 0.02–0.03 wt.% Cl, 0.6–2.0 µg/g B), displaying positive anomalies in the overall depleted,  
43 primitive mantle (PM) normalized trace element and REE spectra ([La/Sm]<sub>n</sub> = 0.16–0.35;  
44 [H<sub>2</sub>O/Nb]<sub>n</sub> = 8–44; [Cl/Nb]<sub>n</sub> = 27–68; [B/Nb]<sub>n</sub> = 9–30, assuming 300 µg/g H<sub>2</sub>O, 8 µg/g Cl and  
45 0.1 µg/g B in PM; Kamenetsky et al., 2010. Composition and temperature of komatiite melts  
46 from Gorgona Island constrained from olivine-hosted melt inclusions. *Geology* 38, 1003–1006).  
47 The inclusions range in δ<sup>11</sup>B values from –11.5 to +15.6 ± 2.2‰ (1 SE), forming two distinct  
48 trends in a δ<sup>11</sup>B vs. B-concentration diagram. Direct assimilation of seawater, seawater-derived  
49 components, altered oceanic crust or marine sediments by ascending komatiite magma cannot  
50 readily account for the volatile contents and B isotope variations. Alternatively, injection of <3%  
51 of a <sup>11</sup>B enriched fluid to the mantle source could be a plausible explanation for the δ<sup>11</sup>B range  
52 that also may explain the H<sub>2</sub>O, Cl and B excess.

53  
54 **Keywords** Komatiites, Gorgona Island, Melt inclusions, Volatile components, Boron isotopes,  
55 Ion microprobe

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58 **1. Introduction**

59           The Island of Gorgona, Colombia, is a unique place where Phanerozoic komatiites have  
60 been first described and then studied in detail (Gansser, 1950; Echeverría, 1980; Aitken and  
61 Echeverría, 1984; Echeverría and Aitken, 1986; Kerr et al., 1996; Kerr, 2005; Sinton et al., 1998;  
62 Walker et al., 1991, 1999; Révillon et al., 2000, 2002; Serrano et al., 2011). The problem of the  
63 origin of Gorgona komatiites is directly connected to the origin of the Caribbean oceanic plateau  
64 (also called the Caribbean Large Igneous Province or CLIP; see review by Hastie and Kerr,  
65 2010). The plateau is the remnant of one or more oceanic melting anomalies active during the  
66 late Cretaceous in the Pacific realm and represents a large outpouring of mafic volcanism. The  
67 main question is whether the mafic volcanic rocks of the Caribbean plateau were erupted onto  
68 the Farallon plate and then subsequently transported during the late Cretaceous to the northeast,  
69 having collided with a large intra-oceanic arc (i.e., the Great Arc of the Caribbean) or whether  
70 formation of the CLIP was related to the opening a large slab window produced by the  
71 intersection of the proto-Caribbean spreading ridge with the Great Caribbean Arc (Pindell et al.,  
72 2006; Pindell and Kennan, 2009; Hastie and Kerr, 2010 and references therein). Previous age  
73 determinations of the Gorgona mafic rocks yielded ages ranging from ~86 to ~92 Ma (Aitken  
74 and Echeverría, 1984; Sinton et al., 1998; Walker et al., 1999), more consistent with the first  
75 hypothesis where the Galapagos plume head impacted the Farallon plate. However, a recent set  
76 of robust  $^{40}\text{Ar}$ - $^{39}\text{Ar}$  ages (Serrano et al., 2011) obtained for the main types of Gorgona mafic  
77 rocks (komatiites, peridotites, gabbros, picritic basalts) documents much longer magmatic  
78 activity on the island, spanning from ~99 to ~61 Ma. This data questions, in fact, the broadly  
79 accepted model suggesting that the Caribbean Large Igneous Province formed ultimately due to  
80 melting of a plume head at ~90 Ma. Furthermore, Serrano et al. (2011) have demonstrated that  
81 multiple magmatic pulses over several tens of Ma, which resulted in a long period of diffuse  
82 magmatism without a clear pattern of migration recognized as within a small area like Gorgona  
83 Island as in other Caribbean LIP areas, can be broadly coeval with the opening of the Caribbean  
84 slab window between ~100 and ~66 Ma.

85           The abundances of volatile components in komatiitic magmas represent a central issue in  
86 their origin because the presence of volatiles in the mantle source considerably affect peridotite

87 solidus, shifting it to lower temperature and pressure and thus this information is essential to  
88 constrain the thermal and chemical evolution of the deep Earth. The controversy remains as to  
89 whether komatiites represent the products of essentially dry and hot melting of the mantle or  
90 whether they result from melting of rising wet plumes, which gained their volatiles by slab  
91 dehydration during subduction (e.g., Allègre, 1982; Arndt et al., 1997; 1998; Parman et al., 1997;  
92 Litasov and Ohtani, 2002; Wilson et al., 2003; Grove and Parman, 2004; Herzberg, 1995;  
93 Herzberg et al., 2007; Berry et al., 2008 among others). Although many existing studies are  
94 traditionally skeptical about a wet origin of the Archaean komatiites (e.g., Arndt et al., 1998 and  
95 references therein), numerous petrological, geochemical and experimental data indicate, for  
96 instance, that komatiites from the 3.5 Ga Barberton Greenstone Belt and from 3.3 Ga  
97 Comondale, South Africa are hydrous, characterized by somewhat elevated SiO<sub>2</sub> contents,  
98 exhibit boninite-like trace element spectra and are believed to have formed due to subduction  
99 zone impact (e.g., Wilson et al., 2003; Grove and Parman, 2004). In addition, several recent  
100 studies utilizing the melt inclusion approach and focusing on the 2.7-Ga-old Belingwe  
101 komatiites, Zimbabwe, have argued that primary komatiitic magmas may contain up to 1 wt.%  
102 H<sub>2</sub>O, but the source of water appears to be controversial (shallow depth magma contamination  
103 versus melting of a hydrous mantle plume source; Shimizu et al., 2001; Berry et al., 2008; Kent  
104 et al., 2009).

105         Although melt inclusions in minerals are routinely used to assess temperature, pressure  
106 and redox conditions of magma crystallization, chemical composition and the volatile budget of  
107 many rock types (e.g., Sobolev and Danyshevsky, 1994; Sobolev and Chaussidon, 1996;  
108 Gurenko and Chaussidon, 1995, 1997; Kent et al., 1999; Danyushevsky et al., 2002a; Rose et al.,  
109 2001; Gurenko et al., 2005; Le Voyer et al., 2008), the melt inclusion approach has not been  
110 widely used to investigate komatiites, apart from a few studies (McDonough and Ireland, 1993;  
111 Shimizu et al., 2001; 2009; Danyshevsky et al., 2002b; Berry et al., 2008; Kent et al., 2009;  
112 Kamenetsky et al., 2010). In part, this is because (a) komatiites are always intrinsically altered  
113 with a significant proportion of their phenocrystic olivine having been fractured and replaced by  
114 serpentine, and (b) melt inclusions entrapped during rapid growth of the spinifex olivine often  
115 remain connected to groundmass so that their volatile concentrations can be biased by shallow

116 depth magma degassing. Recently, Shimizu et al. (2009) and Kamenetsky et al. (2010) described  
117 volatile-enriched melt inclusions in spinel and olivine from Gorgona komatiitic lavas, with  
118 contrasting interpretations. In particular, strong enrichment in volatile elements (CO<sub>2</sub>, F, H<sub>2</sub>O  
119 and Cl) accompanied by high and variable H<sub>2</sub>O/Ce = 307 ± 198, Cl/K<sub>2</sub>O = 1.07 ± 0.52 ratios (in  
120 spinel-hosted melt inclusions) and by exceptionally high H<sub>2</sub>O/Ce ratios of 2000–5000 (in  
121 olivine-hosted melt inclusions) were assigned to “seawater/brine assimilation” by the magma at  
122 upper mantle depth (Shimizu et al., 2009). In contrast, Kamenetsky et al. (2010) suggested the  
123 origin of similar volatile-rich olivine-hosted melt inclusions due to partial melting of a wet  
124 mantle plume, resulting in a parental komatiitic magma with ~17 wt.% MgO and up to 1.0 wt.%  
125 H<sub>2</sub>O, whose crystallization occurred at temperatures as low as 1330–1340°C.

126 Here, we extend our study and report new boron isotope data coupled with volatile and  
127 trace element concentrations obtained for a subset of homogenized melt inclusions reported by  
128 Kamenetsky et al. (2010). The central purpose of the present work is to shed more light on the  
129 origin of the Gorgona komatiitic magmas. These results can be used to better understand the  
130 origin of the Caribbean Large Igneous Province as well as that of komatiitic volcanism  
131 worldwide.

132

### 133 *1.1. Why boron isotopes?*

134 The light lithophile element boron is a very powerful tracer for crust recycling,  
135 slab–mantle wedge interaction or shallow depth contamination because boron has no natural  
136 redox chemistry and fractionation between the two isotopes of boron is almost entirely  
137 controlled by their partitioning between trigonal (near-neutral aqueous fluids preferentially  
138 containing <sup>11</sup>B) and tetrahedral species (coordination of B in many silicates, which preferentially  
139 incorporate <sup>10</sup>B) (Palmer et al., 1987, 1992; Palmer and Swihart, 1996; Peacock and Hervig,  
140 1999; Williams et al., 2001; Hervig et al., 2002; Tonarini et al., 2003; Schmidt et al., 2005;  
141 Wunder et al., 2005). The surface reservoirs are enriched in boron, from 2 ± 1 µg/g B in the  
142 lower continental crust through 10–15 µg/g B in the upper crust to ~100 µg/g B in the altered  
143 oceanic crust and marine sediments (Harder, 1978; Truscott et al., 1986; Spivack and Edmond,  
144 1986; Leeman et al., 1992; Ishikawa and Nakamura, 1992, 1993; Smith et al., 1995, 1997;

145 Kasemann et al., 2001; Savov et al., 2009), as compared to the Earth mantle having 0.05–0.25  
146  $\mu\text{g/g}$  B (Ryan and Langmuir, 1993; Chaussidon and Jambon, 1994; Chaussidon and Marty,  
147 1995).

148 Boron isotopes significantly fractionate during surface geochemical processes ( $\delta^{11}\text{B}$   
149 expressed relatively to NBS 951 standard range from  $-30$  to  $+60\%$  in natural samples from  
150 different geological environments; Palmer and Swihart, 1996). Their fractionation is negligible  
151 during high-temperature igneous processes, such as partial melting and fractional crystallization,  
152 because (a) boron is an incompatible element and remains mostly in the melt (Chaussidon and  
153 Libourel, 1993; Chaussidon and Jambon, 1994) and (b) the isotopic fractionation factor ( $\alpha_{\text{trigonal-}}$   
154  $\text{tetrahedral} = [^{11}\text{B}/^{10}\text{B}]_{\text{trigonal}}/[^{11}\text{B}/^{10}\text{B}]_{\text{tetrahedral}}$ ), as first demonstrated by Kotaka (1973) and Kakihana  
155 et al. (1977) using the example of  $\text{B}(\text{OH})_3$  (boric acid) and  $[\text{B}(\text{OH})_4]^-$  species in the aqueous  
156 solutions, decreases strongly from 1.0193 at 300 K to 1.00704 at 700 K but stays nearly constant  
157 at higher, magmatic temperatures (1.00393 at 1000 K and 1.00188 at 1500 K). Several “second-  
158 order” effects (e.g., differences in geometries and bonding energies between  $[\text{BO}_4]$ -tetrahedra in  
159 micas and  $[\text{B}(\text{OH})_4]^-$  species in fluids, presence of various, chemically different species of boron  
160 associated with K, Al, and Si complexes in melts and fluids at high pressures and temperatures)  
161 can also play an important role in fractionation of B isotopes (Tonarini et al., 2003; Schmidt et  
162 al., 2005; Wunder et al., 2005).

163 Boron isotopic composition of fresh basalts is variable, ranging from  $-9.4 \pm 4.0\%$  in the  
164 mafic continental rift lavas through  $-7.3 \pm 4.0\%$  in oceanic island basalts (OIB) and  $-3.3 \pm$   
165  $2.2\%$  in mid-ocean ridge basalts (MORB) to  $-0.8 \pm 5.0\%$  in back arc basin basalts (BABB)  
166 (Chaussidon and Jambon, 1994; Chaussidon and Marty, 1995). Later, strongly negative  $\delta^{11}\text{B}$   
167 values of  $-7.2\%$  in the Central Andes and  $-17.9\%$  in Ecuadorian arc lavas were described by  
168 Rosner et al. (2003) and Le Voyer et al. (2008). The  $\delta^{11}\text{B}$  value of  $-10 \pm 2\%$  has been estimated  
169 for the OIB mantle source, and  $-4.0 \pm 1.6\%$  and  $-3.6 \pm 1.9\%$  for the N-MORB and E-MORB  
170 mantle sources, respectively (Chaussidon and Marty, 1995). However, somewhat higher values  
171 inferred for the N- and E-MORB sources obtained on the basis of pillow rim glass compositions  
172 were explained by probable interaction of the magmas with oceanic crust. Later, Gurenko and

173 Chaussidon (1997) demonstrated the overall  $^{11}\text{B}$ -depleted signature of the Icelandic mantle ( $\delta^{11}\text{B}$   
174  $= -11.3 \pm 1.9\%$ ), based on a study of very primitive olivine-hosted melt inclusions.

175 Systematically higher  $\delta^{11}\text{B}$  values reported for arc lavas compared to exhumed  
176 subduction related metamorphic rocks suggest that dehydration reactions significantly decrease  
177 the  $\delta^{11}\text{B}$  values of the subducting oceanic crust and sediments, implying strong fractionation of  
178  $^{11}\text{B}$  and  $^{10}\text{B}$  isotopes as boron partitions in the fluid (Peacock and Hervig, 1999). In particular,  
179 this is supported by experimental study of Wunder et al. (2005), which has shown that a wide  
180 range of  $\delta^{11}\text{B}$  values in volcanic arcs may result from continuous dehydration of micas. Another  
181 major carrier of boron, also a source of variable but generally positive  $\delta^{11}\text{B}$  values, is the  
182 serpentinized mantle wedge at the slab-mantle interface (e.g., Benton et al., 2001; Savov et al.,  
183 2005, 2007). In relation to these sources of B, the formation of a hydrated wedge due to  
184 infiltration of B-rich fluids progressively released from the down dragged slab at arc front depths  
185 may account for strongly varying crossarc boron isotopic signatures (Straub and Layne, 2002).  
186 On the other hand, Chaussidon and Jambon (1994), Chaussidon and Marty (1995) and Smith et  
187 al. (1995) have demonstrated that positive  $\delta^{11}\text{B}$  values accompanied by elevated  $\delta^{18}\text{O}$  of  $>5.5\%$   
188 and increasing radiogenic Sr isotope ratios in primitive MORB and OIB magmas can result from  
189 shallow-level assimilation of the oceanic crust, whereas the negative  $\delta^{11}\text{B}$  related with relatively  
190 low  $\delta^{18}\text{O}$  values of  $<5.5\%$  can be due to interaction of these magmas with the Layer 3 gabbroic  
191 crust (for reference, see oxygen isotope profile through the altered, sediment-covered oceanic  
192 crust of the Oman ophiolite; Gregory and Taylor, 1981). Finally, B isotope variations from  $-8.9$   
193 to  $-0.8\%$  in the Snake River Plain–Yellowstone rhyolites can be ascribed to the existence of  
194 distinctive sources of boron coming from either juvenile basalt-derived protoliths sitting in the  
195 crust or from strongly fluid-depleted, metamorphosed continental crust (Savov et al., 2009). All  
196 of the above information suggests that boron isotopes represent a powerful tool which can be  
197 effectively used to understand the origin of the Gorgona komatiitic magmas.

198

## 199 **2. Geological background and studied samples**

200 Gorgona is a small island (8.3 km long and 2.5 km wide) located approximately 50 km to  
201 the west off the Pacific coast of Colombia (**Fig. 1**). The geology of Gorgona was first described

202 in Gansser (1950) and Echeverría (1980) and interpreted as a sequence of tilted and faulted  
203 blocks of mafic and ultramafic rocks surrounding an axial ridge of cumulate peridotites. The  
204 central part of the island is an undeformed peridotite–gabbro complex. Basalts are by far the  
205 most abundant eruptive rock type. The komatiite lava flows (1.5 to 6 m thick), often having  
206 chilled zones ("polyhedrally jointed zones") at their upper margins, are interlayered with basaltic  
207 flows and sills (Kerr, 2005; **Fig. S1a, b**; *Supplementary online material*).

208 The upper polyhedrally jointed layers are usually 20 to 100 cm thick. They contain  
209 skeletal olivine microphenocrysts (0.2–0.5 mm) in a very fine grained matrix, grading into a 10  
210 to 40 cm thick layer composed of fine random spinifex-textured skeletal olivine plates, set in a  
211 groundmass of "feathery" pyroxenes. The spinifex textured layer consist of hollow elongated  
212 olivine crystals forming chains, blebby chains and plates up to 50 cm in length, which may be  
213 underlain by a thin layer (up to 30 cm thick) of horizontally-oriented, skeletal olivine plates up to  
214 10 cm long. Some, not all, flows possess a lower olivine cumulate layer. Where cumulate layers  
215 do exist, they are generally thin (20–30 cm) and discontinuous, representing no more than 30%  
216 of the flow. The cumulate olivine crystals are up to 1 mm in size, have a predominantly skeletal  
217 habit, though polyhedral crystals are also present, and are set in the groundmass composed of  
218 radiating needles of clinopyroxene, granular olivine, plagioclase laths and Cr-spinel octrahedra  
219 (**Fig. S1c**; *Supplementary online material*).

220 The studied samples were collected from fresh coastal exposures of komatiites on the NE  
221 side of the island (**Fig. 1**). Samples GOR94-3/4/17/44 are from the cumulate zone of komatiitic  
222 flows, GOR94-28 represents the jointed top of a flow, and GOR94-17 (Punta Trinidad) is from  
223 the cumulate zone of a flow having a vesicular zone. The bulk-rock chemical compositions of  
224 the studied samples are given in Kamenetsky et al. (2010).

225

### 226 **3. Results**

227 Olivine hosted melt inclusions (up to 60  $\mu\text{m}$ ) have rounded, elliptic to tubular  
228 morphologies and appear as either clear glass with a shrinkage bubble or microcrystalline  
229 aggregates with several bubbles (**Fig. S1d, e**; *Supplementary online material*). Some inclusions  
230 exhibit a visible rim of olivine crystallized on the walls, suggesting extensive post-entrapment



231 crystallization. Olivine phenocrysts containing melt inclusions were re-heated and homogenized  
232 prior to analysis with electron and ion microprobes (for more detail see *Supplementary online*  
233 *material*). The concentrations of major and trace elements and the volatile contents of heated,  
234 quenched to glass and corrected for Fe loss melt inclusions were reported by Kamenetsky et al.  
235 (2010). Here we present only the composition of melt inclusions which were analyzed for boron  
236 isotopes, including several additional inclusions analyzed for volatile and trace elements (**Table**  
237 **1**).

238 The inclusions are characterized by relatively high contents of volatile components  
239 (0.18–1.03 wt.% H<sub>2</sub>O, 0.052–0.077 wt.% S, 0.022–0.031 wt.% Cl) and boron (0.61–2.02 µg/g  
240 B), displaying positive anomalies in the overall depleted, primitive mantle (PM) normalized  
241 trace element and REE spectra ( $[La/Sm]_n = 0.16–0.35$ ;  $[H_2O/Nb]_n = 8–44$ ;  $[Cl/Nb]_n = 27–68$ ;  
242  $[B/Nb]_n = 9–30$ , assuming 300 µg/g H<sub>2</sub>O, 8 µg/g Cl and 0.1 µg/g B in PM as normalizing  
243 values). The melt inclusions form a compact field of basaltic composition, range from 13 to 18  
244 wt.% MgO, following the common trend of olivine fractionation and accumulation (Fig 2 in  
245 Kamenetsky et al., 2010). All inclusions exhibit a very strong enrichment (ca. 10 to 50 times) in  
246 H<sub>2</sub>O, Cl and B relatively to other trace elements in the multi-element diagram (**Fig. 2**). Their  
247 trace element spectra reflect an overall depleted signature ( $[La/Sm]_n = 0.16–0.38$ ), corresponding  
248 to the field of bulk-rock lava compositions.

249 The B/K ratios ranging from 0.0019 to 0.0063 are systematically higher than those  
250 known for primitive N-MORB (B/K of ~0.001) and OIB (B/K of ~0.0003) melt inclusions,  
251 pillow rim glasses and lavas (Ryan and Langmuir, 1993; Chaussidon and Jambon, 1994;  
252 Gurenko and Chaussidon, 1997). No clear correlations of H<sub>2</sub>O, Cl and B contents with other  
253 major, trace and especially fluid-mobile incompatible (such as Ba, K, La, Sr) element  
254 concentrations are observed. Similarly, the magnitude of B enrichment does not correlate with  
255 H<sub>2</sub>O or Cl enrichment. We explain this by the very narrow concentration ranges of trace  
256 elements in the studied melt inclusions (in particular, 3.1–4.7 µg/g Ba i.e. 12%, 228–343 µg/g K  
257 i.e. 11%, 0.5–0.9 µg/g La i.e.13%, and 73–87 µg/g Sr i.e. 5.5% of the average element  
258 concentration), which in fact do not exceed the common analytical errors of SIMS measurements  
259 for trace elements (the uncertainty ranges between 10 and 30% of 1RSD = 1 relative standard

260 deviation; Kamenetsky et al, 2010). In addition, the absence of correlations may be consistent  
261 with a serpentinized arc source of B (and H<sub>2</sub>O and Cl) in the included melts which represents, as  
262 demonstrated by Savov et al. (2007), “large slab inventory depletions of B (~75%), Cs (~25%),  
263 As (~15%), Li (~15%), and Sb (~8%); surprisingly low (generally less than 2%) depletions of  
264 Rb, Ba, Pb, U, Sr; and no depletions in REE and the high field strength elements (HFSE)” (see  
265 discussion below).

266 The  $\delta^{11}\text{B}$  values of the studied melt inclusions ranging from  $-11.5$  to  $+15.6$  ‰ at B-  
267 concentration of  $0.6$ – $2$   $\mu\text{g/g}$  also show only subtle correlations with some incompatible elements  
268 such as K, Sr, Hf and some REEs, if at all. On the other hand, melt inclusions form two distinct  
269 trends based on  $\delta^{11}\text{B}$  vs. B relationships (**Fig. 3**). The first trend includes seven of 15 inclusions  
270 with relatively low  $\delta^{11}\text{B}$  ranging from  $-11.5$  to  $-7.3$ ‰ (average  $-9.0 \pm 1.5$ ‰), spanning over the  
271 entire range of B concentrations and approaching the “primary” value of the Earth’s mantle ( $-10$   
272  $\pm 2$ ‰; Chaussidon and Marty, 1995). Given that igneous processes such as magma fractionation  
273 or partial melting do not cause significant fractionation of boron isotopes, we ascribe this trend  
274 to olivine crystallization that may occur either in the magma plumbing system or represent  
275 incomplete melting of olivine from the inclusion walls during laboratory experiments. The  
276 second trend is characterized by extension of  $\delta^{11}\text{B}$  to highly positive values (up to  $+15.6$ ‰) with  
277 increasing boron concentrations up to  $1.6$   $\mu\text{g/g}$  B (**Fig. 3**). Possible reasons of the second trend  
278 and its’ implications for the origin of Gorgona komatiites are discussed below.

279

## 280 **4. Discussion**

### 281 *4.1. Shallow depth magma contamination*

282 A main argument against a hydrous origin of the Gorgona komatiites is their extreme  
283 depletion in incompatible trace elements, which progressively increases from basalt through  
284 komatiite to picrite, all being related to the same type of the depleted mantle source (e.g., Arndt  
285 et al., 1997; Revillon et al., 2000, 2002). In particular, Arndt et al. (1997, p. 297) have noted:  
286 “Because the extreme depletion of incompatible elements makes it very unlikely that volatiles  
287 were present in the mantle source (they would have been removed along with other incompatible  
288 elements during the melt extraction), the parental magma of the tuffs and picrites must have

289 gained water during passage towards the surface, perhaps through assimilation of hydrated crust  
290 in a shallow-level magma chamber”. Later, Shimizu et al. (2009) described CO<sub>2</sub>-rich (40–4000  
291 µg/g) komatiitic melt inclusions with highly variable H<sub>2</sub>O (0.03–0.9 wt.%) and Cl (6–1056 µg/g)  
292 contents in Cr-spinels collected in beach sands from Gorgona. They argued that the volatile  
293 contents in melt inclusions could be biased prior to their entrapment by magma degassing or  
294 interaction with seawater or brine or, alternatively, by possible post-entrapment modification in  
295 the inclusions itself. Our first task is therefore to assess whether the compositional signature of  
296 the olivine hosted melt inclusions studied during the present work (i.e., varying B isotope  
297 composition, positive H<sub>2</sub>O, Cl and B anomalies in the incompatible trace element spectra) could  
298 result from shallow depth magma contamination.

299 To answer this question, we examined a two-component mixing scenario, calculating the  
300 effects of possible interaction of presumably “uncontaminated” komatiitic magma with seawater,  
301 15%- and 50%-NaCl brine and altered oceanic crust in a similar way as it has been done by Kent  
302 et al. (1999) and Shimizu et al. (2009). The isotopic composition of boron together with volatile-  
303 to-trace or trace-to-trace element ratios (Cl/K<sub>2</sub>O, B/K<sub>2</sub>O, H<sub>2</sub>O/K<sub>2</sub>O), as used by Kent et al.  
304 (1999) and Shimizu et al. (2009), were selected for model calculations. Due to the lack of data  
305 on the composition of altered oceanic crust, marine sediments and other possible contaminating  
306 agents specific to the Gorgona region, most of the end-member compositions represent broad  
307 estimates available in the submarine environment. We note that the assigned elemental and  
308 isotopic compositions of the selected end-members are fixed values taken from literature, though  
309 the number of possible contaminants and their compositional range could be more variable. The  
310 employment of the advanced assimilation-fractional crystallization model (e.g., DePaolo, 1981)  
311 is not applicable to our particular case because all studied melt inclusions were trapped in Mg-  
312 rich olivines of Fo<sub>89–91</sub>, representing very primitive komatiitic magma, which experienced  
313 negligible olivine ± Cr-spinel fractionation. Early fractionation of olivine ± spinel cannot  
314 significantly affect the volatile contents in the magma since H<sub>2</sub>O, Cl, B and K are strongly  
315 incompatible in these minerals. Thus, fractionation will cause equivalent enrichment of magma  
316 by these elements without affecting their ratios. The following mixing end-members were  
317 selected (**Table 2**):

318

319 1. “Uncontaminated” komatiitic magma. The concentrations of K (309  $\mu\text{g/g}$ ), Ba (4.1  $\mu\text{g/g}$ )  
320 and Nb (0.44  $\mu\text{g/g}$ ) were defined as average composition of seven  $^{11}\text{B}$ -depleted melt  
321 inclusions giving average  $\delta^{11}\text{B}$  value of  $-9.0 \pm 1.5\%$ . The concentrations of  $\text{H}_2\text{O}$ , Cl and  
322 B were calculated from the concentrations of neighboring elements of similar  
323 incompatibility:  $([\text{H}_2\text{O}]_n = ([\text{La}]_n \times [\text{Ce}]_n)^{0.5}$ , being equal to 0.042 wt.%  $\text{H}_2\text{O}$ ,  $[\text{Cl}]_n =$   
324  $([\text{Nb}]_n \times [\text{K}]_n)^{0.5}$ , equal to 7.38  $\mu\text{g/g}$  Cl and  $[\text{B}]_n = ([\text{K}]_n \times [\text{La}]_n)^{0.5}$ , equal to 0.118  $\mu\text{g/g}$  B,  
325 so that no enrichment would be visible in the trace element spectra.

326 2. Seawater. The elemental composition of seawater at 3.5% salinity having 4.5  $\mu\text{g/g}$  B,  
327 19350  $\mu\text{g/g}$  Cl, 392  $\mu\text{g/g}$  K, 0.021  $\mu\text{g/g}$  Ba and  $1.5\text{E}-05$   $\mu\text{g/g}$  Nb was taken from  
328 Turekian (1968). Boron isotopic composition ( $\delta^{11}\text{B} = +39.5\%$ ) was taken from Leeman  
329 and Sisson (1996).

330 3. 15%- and 50%-NaCl brines. The composition of the brine (15%-NaCl: 85 wt.%  $\text{H}_2\text{O}$ ,  
331 99000  $\mu\text{g/g}$  Cl, 2076  $\mu\text{g/g}$  K, 100  $\mu\text{g/g}$  B; 50%-NaCl: 50 wt.%  $\text{H}_2\text{O}$ , 303000  $\mu\text{g/g}$  Cl,  
332 6476  $\mu\text{g/g}$  K) was taken from Kent et al. (1999), who utilized the data of Berndt and  
333 Seyfried (1990) obtained for experimental phase-separated brines. The concentrations of  
334 Ba = 0.111  $\mu\text{g/g}$  and Nb =  $7.94\text{E}-05$   $\mu\text{g/g}$  in the 15%-NaCl brine and the concentrations  
335 of B = 312  $\mu\text{g/g}$ , Ba = 0.347  $\mu\text{g/g}$  and Nb =  $2.48\text{E}-04$   $\mu\text{g/g}$  in the 50%-NaCl brine were  
336 calculated by proportional extrapolation of the known concentrations of B in the 15%-  
337 NaCl brine given by Kent et al. (1999) and Ba and Nb in seawater to the higher K  
338 concentrations by keeping constant their B/K, Ba/K and Nb/K ratios. Boron isotopic  
339 composition of natural NaCl brines varies strongly, ranging from +25.5 to +48.7‰ in the  
340 brines from the Australian salt lakes and from +55.7 to +57.4‰ in the surface brines of  
341 Red Sea, Israel (e.g., Vengosh et al., 1991a, b). We arbitrary assigned the extreme  $\delta^{11}\text{B}$   
342 values of +25.5‰ and +57.4‰ to the 15%- and 50%-NaCl brines, respectively, because,  
343 as shown below, such difference in isotopic and trace element composition between  
344 15%- and 50%-NaCl brines only marginally affects the calculation results.

345 4. Altered oceanic crust (AOC). The chemical composition of AOC, especially that of  
346 volatile components, varies strongly between different segments of the ocean floor (e.g.,

347 Michael and Cornell, 1998). These authors demonstrated that MORBs from the  
348 Galapagos spreading center have strongly varying Cl and K concentrations (25–4050  
349  $\mu\text{g/g}$  Cl, 249–4151  $\mu\text{g/g}$  K; Cl/K = 0.06–1.09). For our modeling, we have chosen the  
350 concentrations of Cl and K representing mean values of the above ranges (2040  $\mu\text{g/g}$  Cl  
351 and 2200  $\mu\text{g/g}$  K, giving Cl/K = 0.93). The concentrations of Ba = 22.6  $\mu\text{g/g}$ , Nb = 1.22  
352  $\mu\text{g/g}$ , H<sub>2</sub>O = 5 wt.%, B = 5.2  $\mu\text{g/g}$  and  $\delta^{11}\text{B} = +3.7\text{‰}$  were taken from Smith et al.  
353 (1997), Kent et al. (1999) and Hochstaedter et al. (2001).

354 5. Siliceous marine sediment (SED). The concentrations of K = 1245  $\mu\text{g/g}$ , H<sub>2</sub>O = 8.2 wt.%,  
355 Ba = 1950  $\mu\text{g/g}$ , and Nb = 2.44  $\mu\text{g/g}$  correspond to the Colombia sediment from  
356 DSDP/ODP drill Hole 504 located 200 km south of the Costa Rica Rift in the eastern  
357 equatorial Pacific (Table 1 in Plank and Langmuir, 1998). We note that SED end-  
358 member was selected exclusively as one possible, if not a single, end-member to account  
359 for the negative  $\delta^{11}\text{B}$  ratios in <sup>11</sup>B-depleted melt inclusions. In contrast to marine  
360 carbonates, which are usually characterized by positive and wide range of  $\delta^{11}\text{B}$  from ~0  
361 to +30‰, the terrigenous, siliceous sediments may have  $\delta^{11}\text{B}$  as low as –15‰ (Leeman  
362 and Sisson, 1996 and references therein). In particular, the concentration of B = 120  $\mu\text{g/g}$   
363 and  $\delta^{11}\text{B} = -10\text{‰}$  (Smith et al., 1997) were assigned to SED. Although the  
364 concentrations of Cl in marine sediments are expected to be strongly varying, we also  
365 arbitrary assigned 1000  $\mu\text{g/g}$  Cl to SED because, as stated above, Cl does not play an  
366 essential role in our calculations, and the SED mixing lines on the diagrams, except to  
367 one panel outlining relationships between  $\delta^{11}\text{B}$  and B/K<sub>2</sub>O, are given only for  
368 comparison with other mixing end-members.

369  
370 As follows from the modeling (**Fig. 4**), direct contamination of the ascending komatiitic  
371 magma at shallow depth by crustal rocks, as suggested by Arndt et al. (1997), or by seawater or  
372 seawater-derived components (saline brines; Shimizu et al., 2009), is unable to account for the  
373 H<sub>2</sub>O, Cl and B enrichment observed in the studied melt inclusions. We emphasize that if one  
374 would consider mixing relations in coordinates of Cl/K<sub>2</sub>O vs. H<sub>2</sub>O/K<sub>2</sub>O ratios only, following  
375 Shimizu et al. (2009), such an approach may lead to an erroneous conclusion. Indeed, to account

376 for the compositions of melt inclusions with  $H_2O/K_2O = 4.5\text{--}29.6$ ,  $Cl/K_2O = 0.46\text{--}1.04$ , would  
377 require  $<1.5$  wt.% seawater or even  $<0.7$  wt.% of 15%-NaCl brine or  $<0.2$  wt.% of 50%-NaCl  
378 brine, respectively (**Fig. 4a**). No effect of AOC or oceanic sediments can be recognized in this  
379 diagram. Similarly, mixing the komatiite end-member with 0.5–1.8 wt.% of 15%-NaCl brine or  
380 0.2–0.5 wt.% of 50%-NaCl brine may effectively explain the entire range of B/K<sub>2</sub>O ratios found  
381 in the studied melt inclusions. In contrast, if we plot the results of this mixing in coordinates of  
382 B/K<sub>2</sub>O vs. Cl/K<sub>2</sub>O (**Fig. 4c**), the essential role of seawater, 15%- and 50%-NaCl brines (the most  
383 effective contamination agents in **Fig. 4a, b**), as well as the role of altered oceanic crust and  
384 marine sediments are no longer justified. Mixing relations in coordinates  $\delta^{11}B$  vs. B/K<sub>2</sub>O (**Fig.**  
385 **4d**) also provide no evidence for magma contamination. Although admixture of marine  
386 sediments (less than 2 wt.% of SED) to the komatiitic magma may potentially account for the  
387  $\delta^{11}B$  range in the <sup>11</sup>B-depleted melt inclusions, we rule out this possibility because there is no  
388 evidence for a significant role of marine sediments from the volatile-trace element relations (**Fig.**  
389 **4a-c**). In conclusion, we note that our data are also in a good agreement with Walker et al. (1999,  
390 p. 722) who have concluded: “Although, the most radiogenic Gorgona komatiites have the  
391 lowest Os concentrations in the suite, mass balance arguments weigh heavily against lithospheric  
392 contamination of their parental magmas accounting for the enrichments in <sup>187</sup>Os”.

393

#### 394 *4.2. Fluid-induced mantle melting: was it due to subduction impact?*

395 Since shallow contamination of the Gorgona komatiitic magmas can not account for  
396 volatile enrichment in the melt inclusions, implying a primary juvenile origin of H<sub>2</sub>O, Cl and B  
397 (Kamenetsky et al., 2010), another process capable of changing B isotopic composition in the  
398 included melts, but not major and trace element concentrations, has to be investigated. Relatively  
399 high H<sub>2</sub>O and Cl abundances in komatiitic melt inclusions have previously been reported for the  
400 Belingwe komatiites, Zimbabwe (0.18–0.26 wt.% H<sub>2</sub>O, Danyushevsky et al., 2002b; 1.1–1.7  
401 wt.% H<sub>2</sub>O, Shimizu et al., 2001; 0.4–0.6 wt.% H<sub>2</sub>O, 500–700 μg/g Cl, Cl/K<sub>2</sub>O ~0.8–1.5, Kent et  
402 al., 2009). In particular, Kent et al. (2009) have suggested generation of the Belingwe komatiitic  
403 magmas “from melting of “damp” mantle, possibly with analogies to modern back arc  
404 environments”.

405           Although the model proposing that komatiites are generated by dry and hot mantle  
406 plumes rising from the thermal boundary layer within the Earth is widely recognized (Arndt et  
407 al., 1997; Herzberg, 1995; Kerr, 2005; Kerr et al., 1995, 1996; Walter, 1998; Révillon et al.,  
408 2000, 2002; Hastie and Kerr, 2010), an alternative scenario which favors a subduction related  
409 origin, recently reviewed in detail by Parman and Grove (2005), apparently could be viable to  
410 explain substantial H<sub>2</sub>O contents found in several types of komatiites, including those from the  
411 Island of Gorgona studied here. As stated above, a recent set of <sup>40</sup>Ar-<sup>39</sup>Ar ages obtained for the  
412 main types of Gorgona mafic rocks (Serrano et al., 2011) documents much longer magmatic  
413 activity on the island (from 98.7 ± 7.7 to 61.4 ± 4.8 Ma), as compared to the previous age  
414 determinations of about 86–92 Ma (Aitken and Echeverría, 1984; Sinton et al., 1998; Walker et  
415 al., 1999). Serrano et al. (2011) concluded that the diffuse magmatism occurred in Gorgona  
416 Island shows no clear pattern of migration, being broadly coeval with the opening of the  
417 Caribbean slab window and question, in fact, the broadly accepted origin of the Caribbean LIP  
418 ultimately due to melting of a plume head at ~90Ma. Then, if the model advocated by Serrano et  
419 al. (2011) is correct, the fluids released from the subducting slab or from metasomatically  
420 reworked mantle wedge might have triggered partial melting in the upwelling mixed mantle,  
421 resulting in a wide compositional range of magmas variously enriched or depleted in volatile  
422 components and having a range of  $\delta^{11}\text{B}$  values. This scenario is in agreement with the recent  
423 study of Kamenetsky et al. (2010), who also have questioned the ultimately "hot" signature of  
424 the Gorgona komatiites.

425           A significant fraction of initial boron and <sup>11</sup>B isotope inventory is contained in the  
426 uppermost few km of the subducting oceanic crust and sediments, being released progressively  
427 as subduction proceeds (e.g., Leeman and Sisson, 1996). Boron mobilized from the slab due to  
428 its dehydration will be systematically heavier than the parental AOC, resulting in +5 to +10‰  
429 enrichment of the fluid at 400–500°C (Ishikawa and Nakamura, 1992; Leeman and Sisson, 1996;  
430 Rose et al. 2001, Rosner et al. 2003; Leeman et al., 2004; Le Voyer et al., 2008). Since the  
431 composition of slab-derived fluid may vary strongly, depending closely on the composition of  
432 the subducting crust, on one hand, and the degree of its devolatilization, on another, a wide  $\delta^{11}\text{B}$   
433 range from –7‰ through +2‰ up to +15‰ (assuming +10‰ enrichment of the fluid phase

434 relative to the upper end of  $\delta^{11}\text{B}$  values given for AOC by Smith et al., 1997) is very probable  
435 (Rose et al. 2001; Leeman et al., 2004; Gurenko et al., 2005; Le Voyer et al., 2008; Tonarini et  
436 al., 2011 and references therein). The recent study by Tonarini et al. (2011) reports boron isotope  
437 data for the South Sandwich Island arc lavas, emphasizing their very positive  $\delta^{11}\text{B}$  values (+12 to  
438 +18‰, being the heaviest yet observed for subduction-related rocks), substantial enrichment in B  
439 concentrations (3 to 25  $\mu\text{g/g}$  B) and decoupling of B from other fluid-immobile incompatible  
440 elements. All these signatures are very similar to what we observe for the Gorgona melt  
441 inclusions, except for relatively high B concentrations. The authors propose multi-stage  
442 recycling of high- $\delta^{11}\text{B}$  and high-B serpentinite (mixed with arc crust and sediment material) as a  
443 possible source of extremely  $^{11}\text{B}$ -rich fluids. Presence of serpentinite in the slab is particularly  
444 important because (a) minimal B isotopic fractionation is expected during breakdown of  
445 serpentinite minerals, (b) high- $\delta^{11}\text{B}$  and high-B concentration fluids can be hosted in serpentine  
446 minerals up to 3–4 GPa and 650 °C and (c) when released, these fluids may account for  
447 decoupling of B and trace elements such as Rb, Ba, Pb, U, Sr (Ulmer and Trommsdorff, 1995;  
448 Benton et al., 2001; Savov et al., 2005, 2007; Tonarini et al., 2011).

449 Similar processes of fluid-enhanced partial melting of a hot mantle plume can be  
450 considered to account for the  $\delta^{11}\text{B}$  range in the Gorgona melt inclusions. The effects of possible  
451 interaction of slab-derived fluids with depleted MORB-like mantle source are illustrated on  
452 diagrams in coordinates of  $\delta^{11}\text{B}$  vs. B/Nb and Ba/Nb ratios (**Fig. 5**). Given that  $^{10}\text{B}/^{11}\text{B}$ , B/Nb  
453 and Ba/Nb ratios do not change significantly during partial melting or magma crystallization  
454 processes, the ratios of these elements calculated for the source can be directly applied to the  
455 resulting partial melts. The mixing end-members were defined as below (**Table 2**):

456

- 457 1. Mantle (MANT). A depleted trace element and radiogenic isotope signature of the  
458 Caribbean plume giving rise to the Gorgona komatiites has been proposed by Arndt et al.  
459 (1997). Taking into account the overall trace element depleted composition of the studied  
460 melt inclusions, a depleted MORB-like mantle component containing 1.2  $\mu\text{g/g}$  Ba, 0.21  
461  $\mu\text{g/g}$  Nb, 0.06  $\mu\text{g/g}$  B but having  $\delta^{11}\text{B} = -10\text{‰}$  (Chaussidon and Marty, 1995; Salters and  
462 Stracke, 2004) has been selected.



- 463 2. Recycled crust (REC). The concentrations of trace elements and B in the recycled crust,  
464 whose potential presence in the source of the Gorgona komatiites can be deduced from  
465 Re-Os isotope systematics (Walker et al., 1991, 1999), was taken to be the same as that  
466 of AOC previously used in modeling. The B isotopic composition of REC was fixed at  
467  $-12.8\text{‰}$ , corresponding to  $\delta^{11}\text{B}$  value of the residual slab (Tonarini et al., 2011).
- 468 3. Serpentinite (SERP). The forearc serpentinized peridotites are characterized by a wide  
469 range of boron contents (6.6–126  $\mu\text{g/g}$  B), boron isotopic composition (+10.4 to  
470 +25.3 $\text{‰}$ ) and trace element concentrations (Benton et al., 2001; Savov et al., 2005,  
471 2007). The presence of a serpentinized mantle wedge at the slab-mantle interface and its  
472 admixture into the ascending hot mantle plume may potentially modify the composition  
473 of the source rocks whereas the plume penetrates through a gap in the slab. We thus  
474 consider serpentinized peridotite as a possible mixing end-member, assigning the  
475 following composition: 3.2  $\mu\text{g/g}$  Ba, 0.1  $\mu\text{g/g}$  Nb, 20  $\mu\text{g/g}$  B and  $\delta^{11}\text{B} = 14.8\text{‰}$  based on  
476 the data by Benton et al. (2001) and Savov et al. (2005, 2007).
- 477 4. Subducting slab fluid (SSF). Three subducting slab fluid compositions (SSF1 through  
478 SSF3) with contrasting B concentrations and isotopic compositions (SSF1: 1000  $\mu\text{g/g}$  Ba,  
479 200  $\mu\text{g/g}$  B,  $\delta^{11}\text{B} = 15\text{‰}$ ; SSF2: 500  $\mu\text{g/g}$  Ba, 100  $\mu\text{g/g}$  B,  $\delta^{11}\text{B} = 2\text{‰}$ ; SSF3: 250  $\mu\text{g/g}$   
480 Ba, 50  $\mu\text{g/g}$  B,  $\delta^{11}\text{B} = -7\text{‰}$  at constant Nb = 0.36  $\mu\text{g/g}$ ) were compiled based on the  
481 Rosner et al. (2003), Leeman et al. (2004), Gurenko et al. (2005), Le Voyer et al. (2008)  
482 and Tonarini et al. (2011) data. The decrease of B and Ba contents and  $\delta^{11}\text{B}$  values in the  
483 fluid components qualitatively reflects the increase of slab temperature and degree of its  
484 devolatilization. We note that the fluids released from ultramafic slab  $\pm$  serpentinite and  
485 recycled forearc serpentinites (Tonarini et al., 2011) are within the range of SSF1 to  
486 SSF3 compositions currently used and thus not considered as individual end-members.

487  
488 As follows from the calculations, the major agent affecting the chemical composition of  
489 the hybrid mantle source and, consequently, the magmas formed by its partial melting is the  
490 composition of a fluid component released from the slab (in the case of subduction related  
491 scenario) or the composition of the crust in the magma source (recycling scenario). The presence

492 of ancient recycled crust in the source region (~10%, depending on the B/Nb or Ba/Nb variables  
493 used; **Fig. 5**), is possible but accounts only for a restricted part of the most  $\delta^{11}\text{B}$ -depleted melts.  
494 On the other hand, this amount is in a very good agreement with Re-Os isotopic data of Walker  
495 et al. (1999), who proposed that only a minor amount (<10%) of the recycled material might be  
496 present in the source of Gorgona komatiites. The role of serpentinized peridotites, which may  
497 occasionally be entrained by the hot rising plume, is also limited. These may account only for the  
498 most  $\delta^{11}\text{B}$ -enriched melts, requiring 1–10% of admixture.

499 The injection of the slab-derived fluids into melting region appears to be the most  
500 effective mechanism, as for producing B isotope variability because there is a wide range of  $\delta^{11}\text{B}$   
501 and B/Nb and Ba/Nb ratios in the fluid phase, as for decreasing the temperature of peridotite  
502 solidus. Less than 1 wt.% of SSF1 or SSF2 fluid has to be injected in the ascending hot mantle  
503 material in order to explain the observed  $\delta^{11}\text{B}$  enrichment up to ~16‰. The majority of low- $\delta^{11}\text{B}$   
504 melt inclusions would require admixture of less than 3 wt.% of the SSF3 fluid, representing the  
505 most high temperature product of slab devolatilization. We note that such a small amount of this  
506 fluid will not result in a significant enrichment in fluid mobile incompatible trace elements, such  
507 as Ba, B (both used in calculation), K, La, Ce and Sr, which are typical in subduction-related  
508 magmas, and thus may argue for the immediate role of subduction-derived aqueous fluids in the  
509 origin of H<sub>2</sub>O, Cl and B enrichment. Finally, our data support the conclusion by Kamenetsky et  
510 al. (2010) who demonstrated that in the presence of 0.2–1.0 wt.% H<sub>2</sub>O in the melt, the initial  
511 crystallization temperature of the Gorgona komatiitic magma could be as low as 1330–1340°C,  
512 though being still higher than the temperature of common mid-ocean ridge magmas.

513

## 514 **5. Conclusions**

515 1. The olivine-hosted melt inclusions from Gorgona komatiites are characterized by a wide  
516 range of  $\delta^{11}\text{B}$  values from –11.5 to  $+15.6 \pm 2.2\text{‰}$  (1 SE), forming two distinct trends as a  
517 function of B concentration. The first major trend includes melts with relatively low and  
518 constant  $\delta^{11}\text{B}$  values ranging from –11.5 to –7.3‰ (giving an average of  $-9.0 \pm 1.5\text{‰}$ ),  
519 while the second trend shows a drastic increase of  $\delta^{11}\text{B}$  up to +15.6‰, at similar but  
520 varying (0.6–2  $\mu\text{g/g}$ ) concentrations of B.

- 521 2. Direct assimilation of seawater, seawater-derived components, altered oceanic crust or  
522 marine sediments by ascending komatiite magma cannot readily account for the volatile  
523 contents and B isotope variations.
- 524 3. Injection of the subduction slab fluid into the rising mantle plume in amount of <3 wt.%  
525 provides a plausible explanation for the  $\delta^{11}\text{B}$  range, and also may explain the origin of  
526  $\text{H}_2\text{O}$ , Cl and B enrichment of the Gorgona olivine hosted melt inclusions and slightly  
527 elevated, as compared to MORB, but still relatively low initial temperatures of the  
528 Gorgona komatiite magmas (Kamenetsky et al., 2010).

529

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545

546

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781

782 **Figure captions**

783 Fig. 1. Geological sketch map of Gorgona Island (modified after Echeverría, 1980; Kerr et al.,  
784 1996; Révillon et al., 2000). Sampling locations are marked by sample numbers referred to in the  
785 text.

786  
787 Fig. 2. Multi-element spectra of the olivine-hosted melt inclusions from Gorgona komatiites,  
788 demonstrating positive H<sub>2</sub>O, Cl and B enrichment relative to other trace and rare earth elements.  
789 Concentrations of trace elements in primitive mantle used for normalization are taken from  
790 Hofmann (1998); the normalization values for boron, chlorine and water are 0.3 µg/g B, 17 µg/g  
791 Cl and 300 µg/g H<sub>2</sub>O, as in Kamenetsky et al. (2010). Shaded field represents the composition of  
792 Gorgona picrites, komatiites and D-basalts from Kerr et al. (1996), Arndt et al. (1997), Révillon  
793 et al. (2000) and Serrano et al. (2011).

794  
795 Fig. 3. Boron contents and isotopic composition of the olivine-hosted melt inclusions from  
796 Gorgona komatiites. Bars represent ±1 SE analytical precision, shown individually for each B  
797 isotope measurement (Y axis) and corresponding to the average ±10% analytical precision in the  
798 case of B concentration. The box labeled *Mantle* represents the range of OIB mantle source  
799 ( $\delta^{11}\text{B} = -10 \pm 2\%$  at 0.05–0.25 µg/g B), as suggested by Ryan and Langmuir (1993),  
800 Chaussidon and Jambon (1994), Chaussidon and Marty (1995). Dashed arrows visually  
801 emphasize two possible trends of magma evolution: (1) the increase of B concentration at nearly  
802 constant boron isotopic composition that may be ascribed to high temperature olivine  
803 crystallization, and (2) concurrent increase of B concentration and  $\delta^{11}\text{B}$  values that may result  
804 either from magma contamination at shallow depth or from interaction of B-rich subduction-  
805 related fluid with mantle source rocks (see discussion).

806  
807 Fig. 4. Diagrams of (A) Cl/K<sub>2</sub>O vs. H<sub>2</sub>O/K<sub>2</sub>O, (B) B/K<sub>2</sub>O vs. H<sub>2</sub>O/K<sub>2</sub>O, (C) Cl/K<sub>2</sub>O vs. B/K<sub>2</sub>O  
808 and (D)  $\delta^{11}\text{B}$  vs. B/K<sub>2</sub>O (given as weight ratios, except for  $\delta^{11}\text{B}$  given in ‰) illustrate the effects  
809 of addition of different types of contamination products to presumably “uncontaminated”  
810 Gorgona komatiitic magma. Mixing lines between presumably “uncontaminated” komatiitic

811 magma (*KOM*) and seawater (*SW*), 15%- and 50%-NaCl saline brines (labeled *15%-NaCl* and  
812 *50%-NaCl*, respectively), altered oceanic crust (*AOC*) and siliceous marine sediment (*SED*)  
813 calculated from the compositions given in **Table 2** are shown (see text for explanation).  
814 Numbers on mixing lines refer to wt% of component added to parent komatiite. Note that mixing  
815 lines calculated for 15%- and 50% NaCl brine end-members almost completely overlap between  
816 each other in Panel C. The calculated mixing trends allow us to conclude that direct  
817 contamination of the Gorgona komatiitic magmas by altered basaltic rocks at shallow depth, as  
818 suggested by Arndt et al. (1997), or by seawater or seawater-derived components (saline brine),  
819 as discussed in Shimizu et al. (2009), are unable to account for the H<sub>2</sub>O, Cl and B enrichment  
820 observed in the melt inclusions studied.

821  
822 Fig. 5. Diagrams of (A)  $\delta^{11}\text{B}$  vs. B/Nb and (B)  $\delta^{11}\text{B}$  vs. Ba/Nb summarize the proposed  
823 interpretation of B-isotope variations in olivine-hosted melt inclusions from the Gorgona  
824 komatiites. Mixing lines between the depleted MORB-type mantle (MANT), recycled crust  
825 (REC), serpentized peridotite (SERP) and subduction slab fluids (SSF1, SSF2 and SSF3) are  
826 shown. Numbers on mixing lines refer to wt% of component added to MANT end-member.  
827

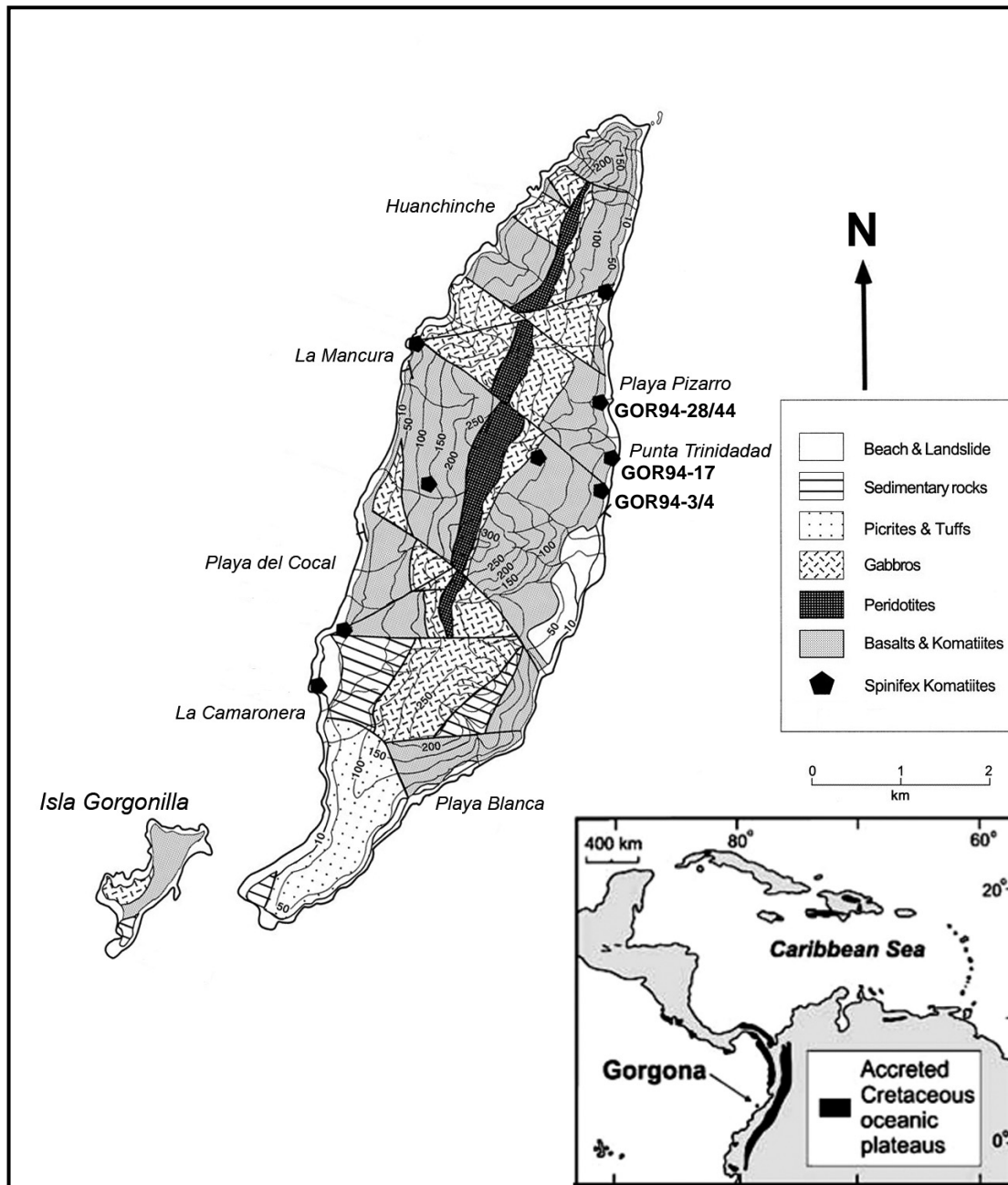


Fig. 1. Gurenko and Kamenetsky

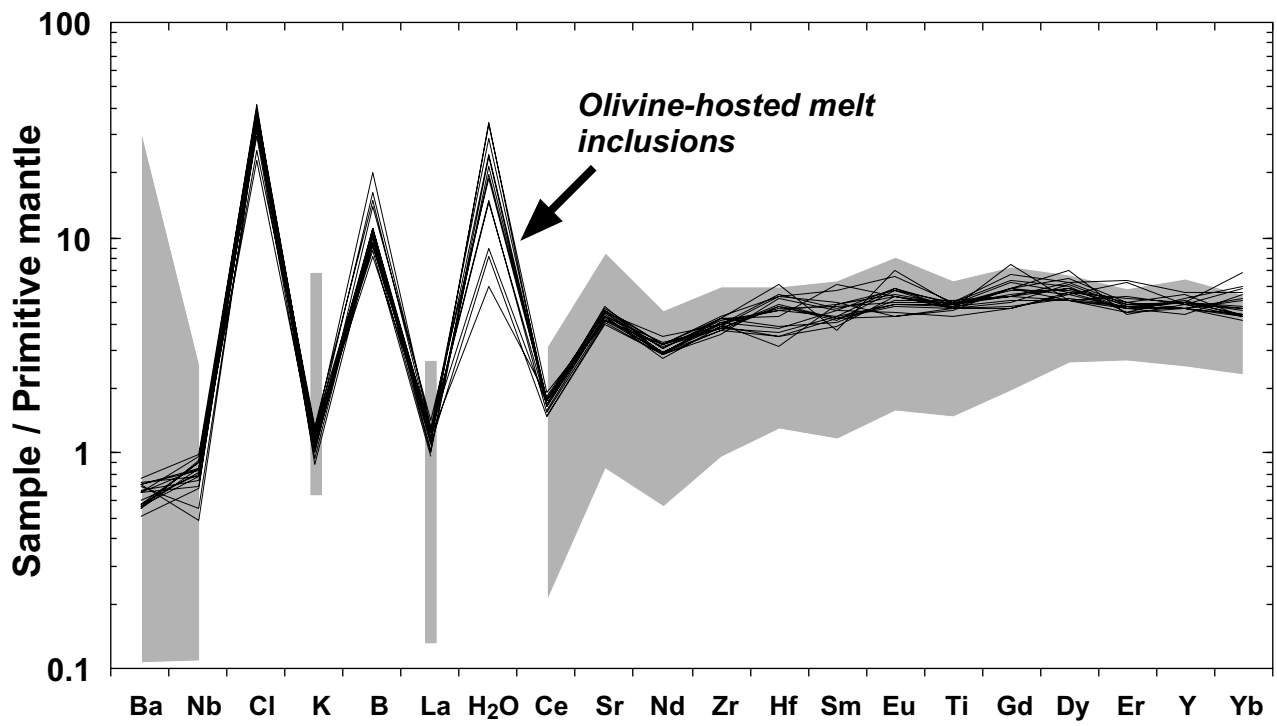


Fig. 2. Gurenko d Kamenetsky



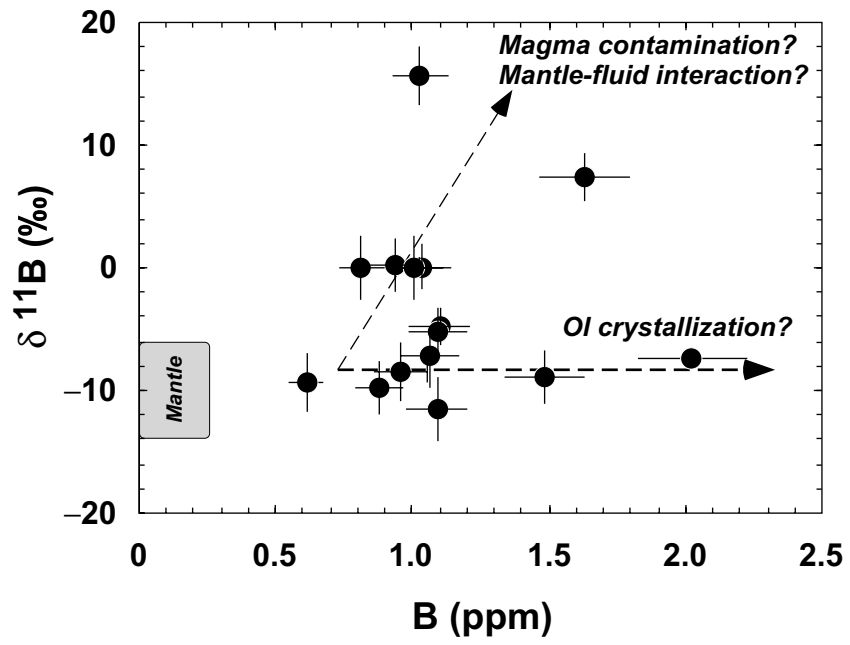


Fig. 3. Gurenko and Kamenetsky

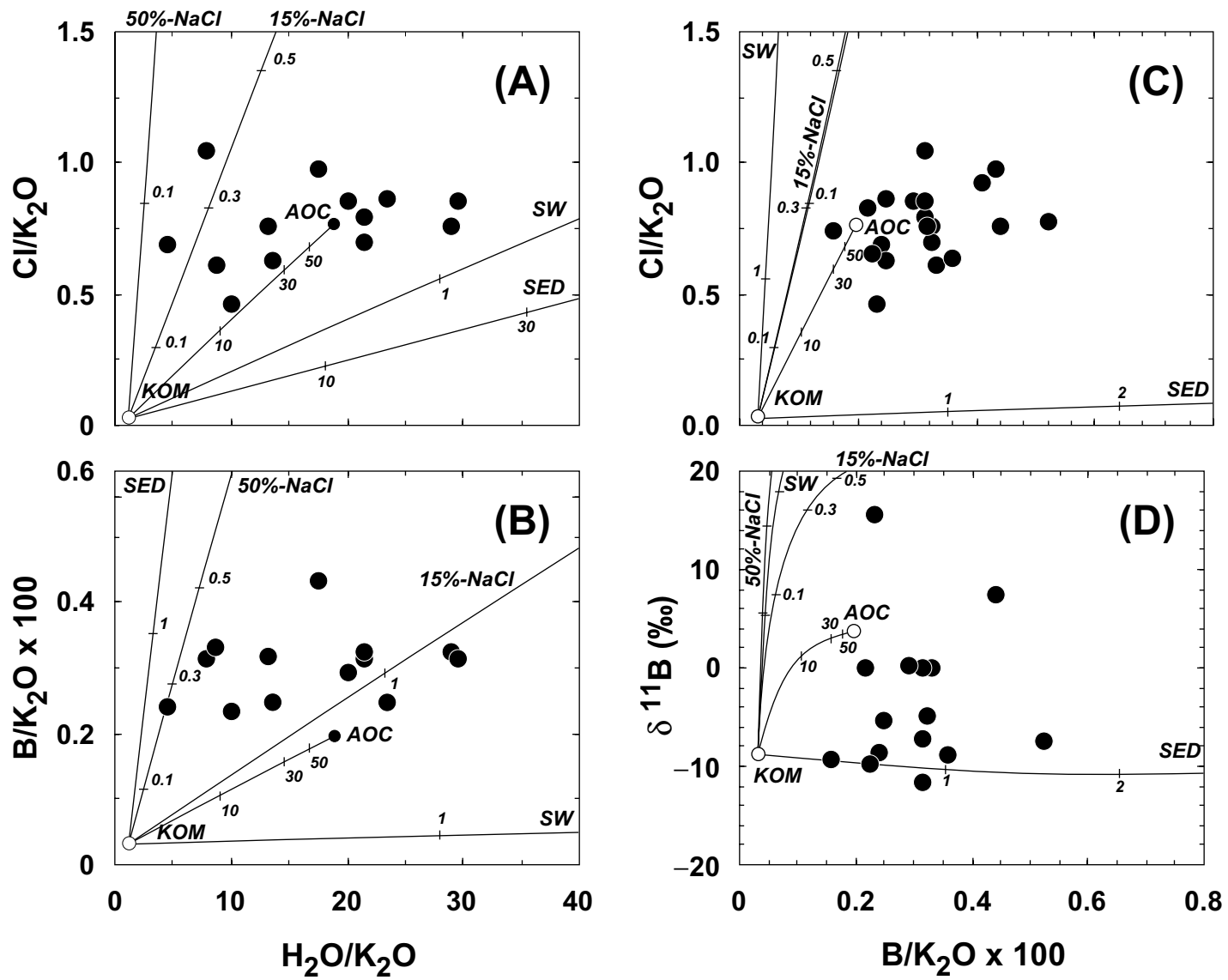


Fig. 4. Gurenko and Kamenetsky

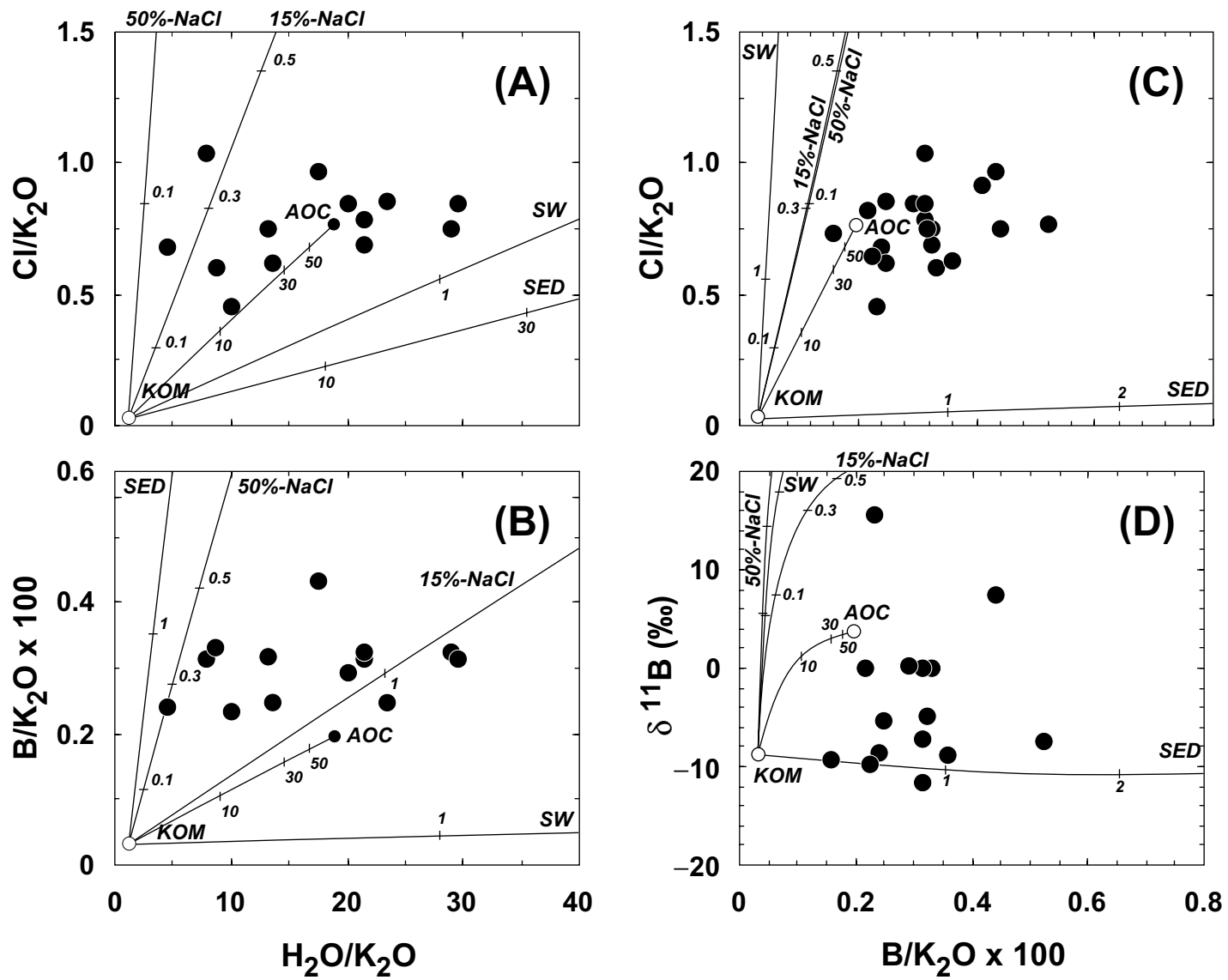


Fig. 4. Gurenko and Kamenetsky

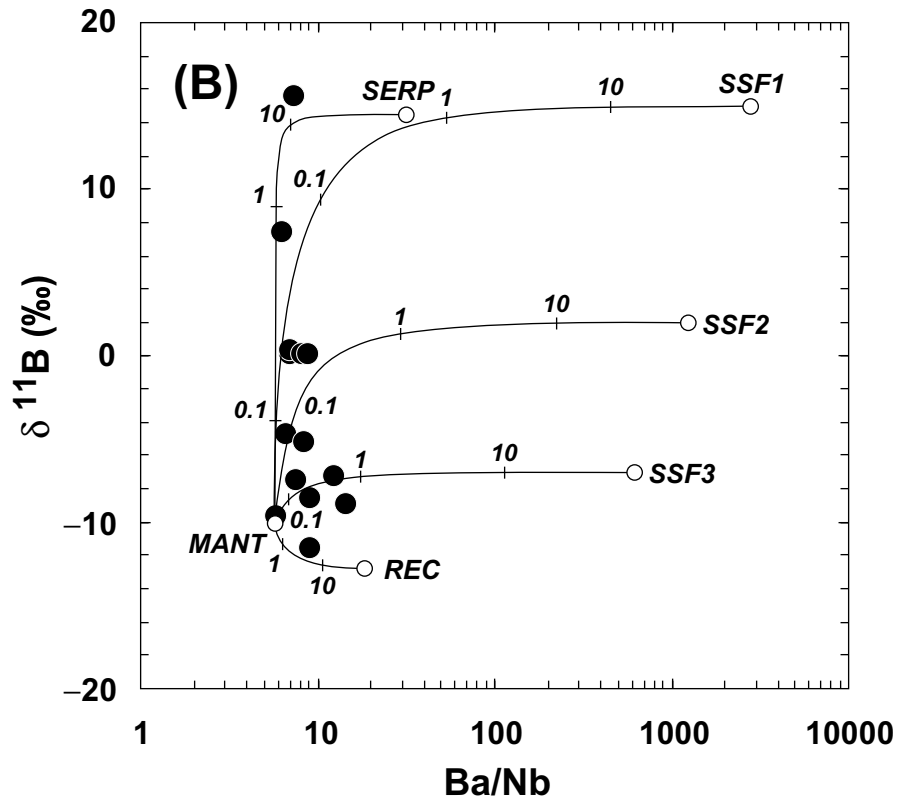
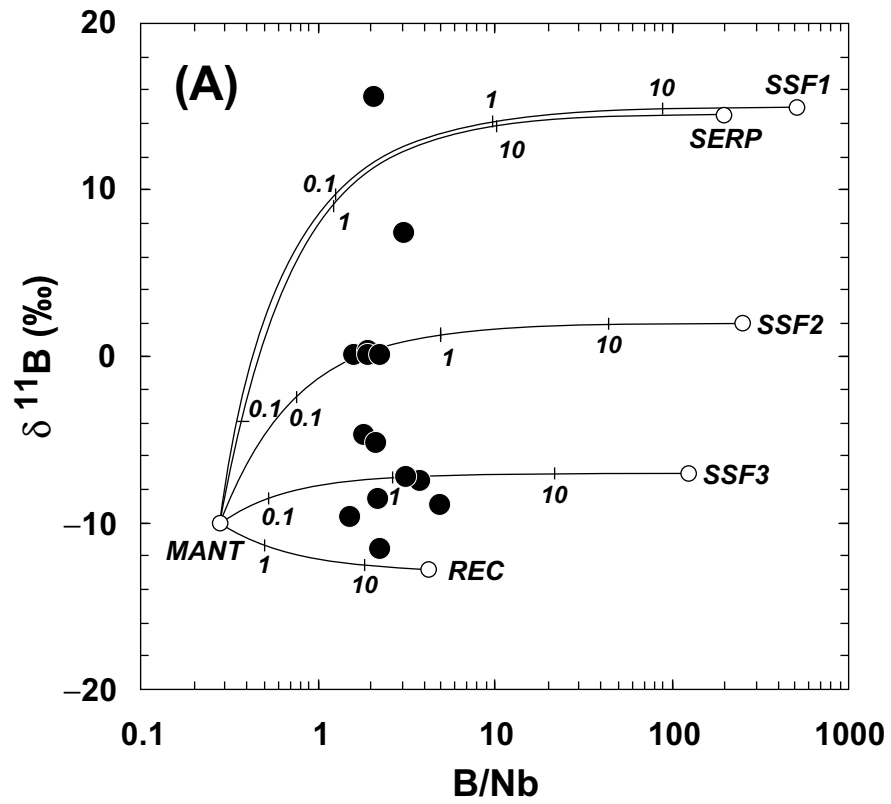


Fig. 5. Gurenke and Kamenetsky

**Table 1.**

Compositions of the laboratory heated melt inclusions in olivine from Gorgona komatiites, Colombia

Grain No	94-3-30a	94-3-32	94-3-33	94-4-20	94-4-21	94-17-10a	94-17-11	94-17-13	94-17-14	94-17-15	94-17-16	94-28-1	94-28-2	94-28-3a	94-28-7	94-28-8a	94-28-9	94-28-36	94-44-24	94-44-27
Sample	GOR94-3			GOR94-4			GOR94-17			GOR94-28			GOR94-44			GOR94-44				
<i>Heated melt inclusions<sup>a</sup></i>																				
SiO <sub>2</sub>	48.25	48.43	48.64	49.02	49.16	48.92	48.61	48.14	48.05	48.59	49.11	49.42	47.98	48.55	48.20	48.84	48.37	47.94	48.94	49.08
TiO <sub>2</sub>	0.85	0.85	0.83	0.85	0.87	0.85	0.87	0.86	0.88	0.84	0.84	0.95	0.88	0.86	0.88	0.88	0.87	0.86	0.87	0.85
Al <sub>2</sub> O <sub>3</sub>	15.90	15.85	15.54	15.58	15.91	15.75	15.68	15.60	15.94	15.50	15.46	17.17	16.11	16.12	16.58	16.01	15.79	15.85	15.27	15.50
FeO	6.14	6.91	6.38	6.48	6.23	6.71	6.62	7.11	6.79	7.34	7.57	7.15	7.94	8.37	7.54	8.40	8.90	8.24	6.57	6.56
MnO	0.13	0.13	0.10	0.09	0.09	0.12	0.12	0.15	0.13	0.14	0.15	0.11	0.16	0.18	0.14	0.14	0.13	0.14	0.08	0.11
MgO	11.70	11.45	11.91	11.03	10.76	10.90	10.39	10.76	9.84	10.81	10.95	8.39	9.73	9.92	9.56	9.66	10.06	9.94	10.31	10.43
CaO	13.64	13.75	13.40	13.89	14.17	14.22	14.36	13.88	14.54	13.84	13.79	15.00	14.62	14.07	14.46	14.37	14.13	14.12	14.07	13.95
Na <sub>2</sub> O	1.92	1.95	1.89	1.80	1.85	1.83	1.79	1.93	1.92	1.88	1.84	2.12	1.96	1.98	2.06	1.88	1.85	1.89	1.79	1.79
K <sub>2</sub> O	0.044	0.030	0.033	0.035	0.044	0.034	0.034	0.034	0.032	0.037	0.032	0.039	0.037	0.039	0.033	0.038	0.040	0.032	0.041	0.039
P <sub>2</sub> O <sub>5</sub>	0.046	0.054	0.055	0.058	0.061	0.062	0.044	0.052	0.047	0.052	0.054	0.070	0.061	0.062	0.067	0.057	0.056	0.065	0.051	0.046
S	0.052	0.068	0.066	0.056	0.056	0.058	0.052	0.056	0.071	0.063	0.064	0.071	0.077	0.073	0.071	0.072	0.076	0.076	0.052	0.059
Cl	0.023	0.025	0.023	0.029	0.025	0.025	0.026	0.025	0.025	0.022	0.022	0.029	0.028	0.030	0.031	0.031	0.029	0.027	0.026	0.026
Total	98.69	99.49	98.87	98.92	99.22	99.39	98.62	98.60	98.17	99.03	99.88	100.52	99.58	100.24	99.51	100.39	100.31	99.09	98.08	98.44
<i>Host olivine</i>																				
SiO <sub>2</sub>	40.84	41.02	40.38	40.34	40.28	40.68	40.80	40.74	40.69	40.32	40.42	40.35	41.34	40.66	41.24	40.22	40.40	40.64	40.68	40.69
FeO	8.48	9.18	8.61	9.37	9.03	9.61	9.72	9.54	9.81	10.39	10.08	9.21	9.02	8.90	8.85	8.97	8.95	9.25	8.65	8.88
MnO	0.12	0.15	0.14	0.14	0.16	0.15	0.17	0.13	0.15	0.16	0.17	0.15	0.15	0.15	0.15	0.14	0.14	0.15	0.12	0.14
MgO	49.99	49.71	49.26	48.81	48.88	48.86	49.06	49.19	48.86	48.05	48.36	49.01	50.02	49.48	50.17	48.56	48.91	49.32	49.45	49.44
CaO	0.31	0.34	0.32	0.32	0.32	0.32	0.33	0.33	0.33	0.34	0.34	0.32	0.33	0.32	0.32	0.32	0.32	0.31	0.32	0.31
NiO	0.44	0.45	0.43	0.42	0.44	0.42	0.40	0.41	0.42	0.39	0.40	0.43	0.44	0.44	0.45	0.44	0.43	0.42	0.44	0.44
Cr <sub>2</sub> O <sub>3</sub>	0.12	0.13	0.12	0.12	0.11	0.11	0.10	0.12	0.10	0.09	0.11	0.11	0.13	0.12	0.12	0.14	0.14	0.10	0.12	0.11
Total	100.31	100.98	99.25	99.52	99.21	100.14	100.57	100.46	100.36	99.74	99.87	99.56	101.43	100.07	101.30	98.79	99.29	100.19	99.78	100.00
Fo	91.3	90.6	91.1	90.3	90.6	90.1	90.0	90.2	89.9	89.2	89.5	90.5	90.8	90.8	91.0	90.6	90.7	90.5	91.1	90.9
<i>Compositions and temperatures of the trapped melts (normalized to 100%)<sup>b</sup></i>																				
SiO <sub>2</sub>	46.4	46.7	46.8	47.3	47.0	47.1	47.1	46.9	46.8	47.4	47.4	46.7	46.2	46.5	46.2	46.7	46.6	46.6	47.2	47.2
TiO <sub>2</sub>	0.7	0.7	0.7	0.7	0.7	0.7	0.7	0.7	0.7	0.7	0.7	0.7	0.7	0.7	0.7	0.7	0.7	0.7	0.7	0.7
Al <sub>2</sub> O <sub>3</sub>	12.7	13.2	12.6	12.9	12.8	13.1	13.1	13.2	13.4	13.6	13.3	13.2	13.0	13.0	13.1	12.9	13.0	13.2	12.1	12.4
Fe <sub>2</sub> O <sub>3</sub>	1.4	1.4	1.4	1.4	1.4	1.38	1.4	1.4	1.4	1.4	1.4	1.4	1.4	1.4	1.4	1.4	1.4	1.4	1.4	1.4
FeO	9.6	9.6	9.6	9.6	9.6	9.56	9.6	9.6	9.6	9.6	9.6	9.6	9.6	9.6	9.6	9.6	9.6	9.6	9.6	9.6
MnO	0.1	0.1	0.1	0.1	0.1	0.1	0.1	0.1	0.1	0.1	0.1	0.1	0.2	0.2	0.1	0.1	0.1	0.1	0.1	0.1
MgO	16.6	15.3	16.3	14.9	15.4	14.5	14.4	14.6	14.1	13.3	13.8	15.0	15.4	15.6	15.8	15.3	15.4	15.0	16.3	15.9
CaO	10.9	11.4	10.9	11.6	11.4	11.9	12.0	11.8	12.2	12.2	11.9	11.6	11.8	11.4	11.4	11.6	11.7	11.8	11.2	11.2
Na <sub>2</sub> O	1.5	1.6	1.5	1.5	1.5	1.5	1.5	1.6	1.6	1.7	1.6	1.6	1.6	1.6	1.6	1.5	1.5	1.6	1.4	1.4
K <sub>2</sub> O	0.03	0.03	0.02	0.03	0.03	0.03	0.03	0.03	0.03	0.04	0.03	0.03	0.03	0.03	0.02	0.03	0.03	0.03	0.03	0.03
P <sub>2</sub> O <sub>5</sub>	0.04	0.04	0.04	0.05	0.05	0.05	0.03	0.04	0.04	0.04	0.04	0.05	0.05	0.05	0.06	0.05	0.05	0.05	0.04	0.04
T <sub>calc</sub> °C	1383	1358	1376	1347	1358	1339	1337	1343	1332	1314	1325	1352	1362	1365	1370	1357	1359	1352	1375	1368
Ol-add	24.0	18.9	21.7	19.4	22.7	18.2	19.1	17.7	19.2	13.0	14.2	27.1	22.8	22.4	25.5	21.9	20.0	19.9	26.2	24.7

**Table 1 (Continued)**

Grain No	94-3-30a	94-3-32	94-3-33	94-4-20	94-4-21	94-17-10a	94-17-11	94-17-13	94-17-14	94-17-15	94-17-16	94-28-1	94-28-2	94-28-3a	94-28-7	94-28-8a	94-28-9	94-28-36	94-44-24	94-44-27	
Sample	GOR94-3			GOR94-4		GOR94-17						GOR94-28							GOR94-44		
<i>Boron isotopes, volatiles and trace elements<sup>c</sup></i>																					
$\delta^{11}\text{B}$ , ‰	15.6	0.1	0.1	-11.5	-5.3	ND	-4.8	-7.3	ND	ND	0.3	-9.3	7.4	-7.5	ND	0.0	-8.5	ND	-8.9	-9.7	
1 $\sigma$ SE	2.4	2.6	1.8	2.7	2.0	ND	1.5	2.6	ND	ND	2.2	2.5	2.0	1.9	ND	2.6	2.3	ND	2.2	2.2	
H <sub>2</sub> O	0.44	0.27	0.71	1.03	0.59	1.00	0.73	0.44	0.57	0.87	0.64	ND	ND	ND	ND	ND	0.18	0.25	ND	ND	
Cl	205	184	262	296	277	260	236	254	315	321	273	285	280	300	305	313	272	330	263	257	
B	1.03	1.01	1.04	1.09	1.10	1.1	1.10	1.07	1.4	0.9	0.93	0.61	1.63	2.02	1.4	0.81	0.95	1.0	1.48	0.88	
K	334	270	284	299	308	275	246	265	262	313	290	ND	307	321	274	313	288	228	343	327	
Ti	5081	5369	5129	5357	5266	5383	5120	5198	5215	5599	5097	ND	5413	5333	5333	5313	5395	4713	5022	5130	
V	405	406	406	453	412	484	369	388	378	462	397	ND	ND	ND	ND	ND	435	300	ND	ND	
Cr	686	706	825	736	579	ND	751	444	741	716	484	ND	ND	ND	ND	ND	401	778	ND	ND	
Sr	77.0	79.2	78.5	82.2	80.2	75.8	75.2	75.0	81.3	84.3	75.0	ND	86.4	83.3	84.9	86.8	82.3	72.7	78.1	73.7	
Y	19.8	19.6	18.7	18.6	19.6	20.7	18.4	18.7	20.4	19.5	19.4	ND	22.0	19.9	22.0	21.5	19.4	17.4	19.5	18.3	
Zr	37.9	39.1	38.0	38.2	39.2	38.5	36.5	37.0	38.8	41.0	35.7	ND	42.3	40.9	42.0	41.3	40.4	34.2	37.1	35.9	
Nb	0.49	0.52	0.46	0.48	0.51	0.56	0.60	0.34	0.42	0.61	0.48	ND	0.53	0.53	0.42	0.50	0.44	0.56	0.30	0.57	
Ba	3.67	4.31	4.06	4.46	4.29	3.52	3.98	4.27	3.07	4.65	3.42	ND	3.35	4.00	4.07	3.49	3.94	3.40	4.36	3.37	
La	0.77	0.77	0.63	0.69	0.71	0.71	0.77	0.77	0.87	0.80	0.61	ND	0.81	0.79	0.46	0.77	0.73	0.68	0.77	0.59	
Ce	2.71	2.70	2.63	2.65	2.87	3.09	2.81	2.47	2.83	2.83	2.80	ND	2.82	2.72	2.35	2.70	2.94	2.35	2.54	2.38	
Pr	0.59	0.55	0.51	0.64	0.56	0.69	0.57	0.60	0.64	0.53	0.55	ND	0.67	0.50	0.66	0.51	0.70	0.44	0.52	0.33	
Nd	3.52	3.40	3.26	3.46	3.46	3.81	3.84	3.72	4.17	3.80	3.50	ND	3.81	3.61	3.12	3.43	3.61	3.42	3.89	3.45	
Sm	1.92	1.68	1.51	1.89	1.80	1.84	1.68	1.62	1.83	1.57	1.64	ND	2.18	1.44	1.79	2.34	1.88	1.77	1.62	1.60	
Eu	0.63	0.71	0.77	0.80	0.85	0.73	0.84	0.72	0.77	0.75	0.69	ND	0.96	1.03	0.60	0.78	0.83	0.65	0.62	0.82	
Gd	2.46	3.15	2.96	2.74	2.77	3.23	2.55	2.61	2.97	2.74	2.41	ND	3.44	2.93	2.41	2.98	2.76	2.43	2.95	3.85	
Tb	0.51	0.59	0.58	0.51	0.56	0.55	0.52	0.51	0.51	0.57	0.56	ND	0.64	0.59	0.51	0.54	0.56	0.54	0.49	0.44	
Dy	3.33	3.52	3.29	3.67	3.71	3.81	3.89	3.28	3.46	3.95	3.47	ND	3.96	3.64	3.57	4.09	3.29	3.55	4.47	3.30	
Ho	0.67	0.66	0.71	0.74	0.74	0.84	0.62	0.63	0.59	0.71	0.75	ND	0.81	0.75	0.84	0.91	0.78	0.69	0.65	0.53	
Er	2.00	2.16	1.94	2.08	1.98	1.86	2.01	1.89	1.97	2.05	2.61	ND	2.64	2.05	2.00	2.09	2.20	1.94	1.85	2.00	
Tm	0.25	0.24	0.29	0.30	0.25	0.27	0.25	0.30	0.27	0.25	0.33	ND	0.34	0.29	0.25	0.32	0.29	0.24	0.38	0.28	
Yb	1.83	1.98	2.15	1.83	1.92	2.49	2.10	1.72	1.80	1.96	2.40	ND	2.29	2.88	2.76	1.80	1.94	2.24	1.79	1.79	
Lu	0.26	0.28	0.32	0.31	0.35	0.30	0.26	0.23	0.27	0.32	0.32	ND	0.31	0.28	0.24	0.30	0.30	0.30	0.22	0.28	
Hf	1.43	1.22	0.94	0.84	1.02	1.25	0.98	1.28	1.46	1.03	0.93	ND	1.44	1.63	1.53	1.16	1.24	1.40	1.31	1.27	
Pb	0.40	0.16	0.06	0.11	ND	0.30	ND	0.59	0.71	0.37	0.36	ND	0.10	0.20	0.30	0.09	0.12	0.38	ND	0.05	
Th	ND	0.02	0.02	0.02	0.01	0.06	0.02	0.04	0.03	0.03	ND	ND	0.03	0.03	ND	0.02	0.03	0.02	ND	0.02	
B/K	0.0031	0.0037	0.0037	0.0036	0.0036	0.0040	0.0045	0.0040	0.0053	0.0029	0.0032	0.0019	0.0053	0.0063	0.0049	0.0026	0.0033	0.0043	0.0043	0.0027	
[La/Sm] <sub>n</sub>	0.25	0.29	0.26	0.23	0.25	0.24	0.29	0.30	0.30	0.32	0.24	-	0.23	0.35	0.16	0.21	0.24	0.24	0.30	0.23	
[H <sub>2</sub> O/Nb] <sub>n</sub>	19	10	32	44	24	37	25	27	28	29	28	-	-	-	-	-	8	9	-	-	
[Cl/Nb] <sub>n</sub>	32	27	44	47	42	36	31	58	57	41	44	-	41	44	56	48	48	46	68	35	
[B/Nb] <sub>n</sub>	13	12	14	14	13	12	11	19	20	9	12	-	19	24	20	10	14	11	30	10	

**Notes for Table 1.**

<sup>a</sup> Major element concentrations of the laboratory heated melt inclusions and their host olivines (given in wt%) were analyzed using the JEOL JXA-8200 electron microprobe at the Max Planck Institute for Chemistry (Mainz, Germany) at standard conditions (Kamenetsky et al., 2010). A set of reference materials (i.e. natural and synthetic oxides, minerals and glasses; Micro-Analysis Consultants Ltd, Cambridgeshire, UK) and the Smithsonian Institution standard set for electron microprobe analysis (Jarosewich et al., 1980) were used for routine calibration and instrument stability monitoring. Typical analytical uncertainties (2RSD = 2 $\sigma$  relative standard deviation) are 1.5–3.0% for SiO<sub>2</sub>, Al<sub>2</sub>O<sub>3</sub>, FeO, MgO, CaO, TiO<sub>2</sub>; 4–6% for Na<sub>2</sub>O, 10% for K<sub>2</sub>O, 15% for P<sub>2</sub>O<sub>5</sub>, and 30% for MnO. As a monitor sample for S and Cl measurements, we also used the USNM 111240/52 VG-2 basaltic glass (recommended values of 0.134–0.137 wt.% S; 0.030 wt.% Cl; Dixon et al., 1991; Thordarson et al., 1996; N. Metrich, personal communication, 2003). The concentrations of 0.140  $\pm$  0.023 wt.% S and 0.029  $\pm$  0.007 wt.% Cl ( $\pm$ 2 $\sigma$  SD = 2-sigma standard deviation, N = 37) were obtained during this study (for more detail see Kamenetsky et al., 2010). Fo = mol% of forsterite component of the host olivine.

<sup>b</sup> Calculated compositions and temperatures of the trapped komatiitic melts are based on the compositions of the laboratory heated melt inclusions, corrected for “Fe-loss” and post-entrapment crystallization of host mineral (for detail see Kamenetsky et al., 2010); Ol-add = total amount of olivine (wt%) “added” to the included melts during correction.

<sup>c</sup> Boron isotope composition (in permil units) calculated relative to the NBS 951 standard (<sup>11</sup>B/<sup>10</sup>B = 4.04558  $\pm$  0.00033; Spivack and Edmond, 1986) given at  $\pm$ 1 $\sigma$  SE analytical precision (see text for detail). The contents of volatile components (H<sub>2</sub>O given in wt%, Cl in  $\mu$ g/g) and trace element concentrations (given in  $\mu$ g/g) were analyzed using the CAMECA IMS 3f ion microprobe and the single collector sector-field ICP-MS Element 2 instrument equipped with the New Wave Research UP213 Nd-YAG (213 nm) laser, both located in the Max Planck Institute for Chemistry, Mainz, Germany (raw values uncorrected for post-entrapment olivine crystallization in melt inclusions are given; for more detail see Kamenetsky et al., 2010). The [La/Sm]<sub>n</sub>, [H<sub>2</sub>O/Nb]<sub>n</sub>, [Cl/Nb]<sub>n</sub> and [B/Nb]<sub>n</sub> are element ratios normalized to primitive mantle (PM; assuming 300  $\mu$ g/g H<sub>2</sub>O, 8  $\mu$ g/g Cl and 0.1  $\mu$ g/g B, while Nb concentration of 0.6175  $\mu$ g/g is taken from Hofmann, 1988). ND = not determined, – = no value.

**Table 2.**

Compositions of the end-members used for calculation of mixing lines (see text for more details).

Component Unit	B μg/g	$\delta^{11}\text{B}$ ‰	H <sub>2</sub> O wt%	Cl μg/g	K μg/g	Ba μg/g	Nb μg/g
KOM <sup>(a)</sup>	0.118	-9.0	0.042	7.38	309	4.1	0.44
SW <sup>(b)</sup>	4.5	+39.5	97.5	19350	392	0.021	1.50E-05
15%-NaCl <sup>(c)</sup>	100	+25.5	85	99000	2076	0.111	7.94E-05
50%-NaCl <sup>(c)</sup>	312	+56.6	50	303000	6476	0.347	2.48E-04
AOC <sup>(d)</sup>	5.2	+3.7	5	2040	2200	22.6	1.22
REC <sup>(e)</sup>	5.2	-12.8	-	-	-	22.6	1.22
SED <sup>(f)</sup>	120	-10.0	8.2	1000	1245	1950	2.44
MANT <sup>(g)</sup>	0.06	-10.0	116	0.51	60	1.2	0.21
SERP <sup>(h)</sup>	20	+14.5	-	-	-	3.2	0.1
SSF1 <sup>(i)</sup>	200	+15.0	-	-	-	1000	0.36
SSF2 <sup>(i)</sup>	100	+2.0	-	-	-	500	0.36
SSF3 <sup>(i)</sup>	50	-7	-	-	-	250	0.36

<sup>a</sup> KOM = “uncontaminated” komatiitic magma representing average composition of seven, the most <sup>11</sup>B-depleted melt inclusions (i.e., 94-28-1, -28-3mi1, -28-9, -4-20, -44-27, -17-13, -44-24; Table 1). The concentrations of H<sub>2</sub>O, Cl and B were calculated from the concentrations of neighboring elements of similar incompatibility:  $[\text{H}_2\text{O}]_n = ([\text{La}]_n \times [\text{Ce}]_n)^{0.5}$ ,  $[\text{Cl}]_n = ([\text{Nb}]_n \times [\text{K}]_n)^{0.5}$  and  $[\text{B}]_n = ([\text{K}]_n \times [\text{La}]_n)^{0.5}$ .

<sup>b</sup> SW = seawater composition at 3.5% salinity.

<sup>c</sup> 15%- and 50%-NaCl brines.

<sup>d</sup> AOC = altered oceanic crust.

<sup>e</sup> REC = recycled crust.

<sup>f</sup> SED = siliceous marine sediment.

<sup>g</sup> MANT = mantle component of the Gorgona komatiite magma source.

<sup>h</sup> SERP = recycled forearc serpentinite.

<sup>i</sup> SSF1-SSF3 = subducting slab fluids.