Dismantling the Deep Earth: Geochemical Constraints from Hotspot Lavas for the Origin and Lengthscales of Mantle Heterogeneity

by

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Tileinkað Önnu Maríu

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Submitted to the Department of Marine Geology and Geophysics, Massachusetts Institute of Technology-Woods Hole Oceanographic Institution in partial fulfillment of the requirements for the degree of Doctor of Philosophy

Abstract

Chapter 1 presents the first published measurements of Sr-isotope variability in olivinehosted melt inclusions. Melt inclusions in just two Samoan basalt hand samples exhibit most of the total Sr-isotope variability observed in Samoan lavas. Chapter 3 deals with the largest possible scales of mantle heterogeneity, and presents the highest magmatic ³He/⁴He (33.8 times atmospheric) discovered in Samoa and the southern hemisphere. Along with Samoa, the highest ³He/⁴He sample from each southern hemisphere high ³He/⁴He hotspot exhibits lower ¹⁴³Nd/¹⁴⁴Nd ratios than their counterparts in the northern hemisphere. Chapter 2 presents geochemical data for a suite of unusually enriched Samoan lavas. These highly enriched Samoan lavas have the highest ⁸⁷Sr/⁸⁶Sr values (0.72163) measured in oceanic hotspot lavas to date, and along with trace element ratios (low Ce/Pb and Nb/U ratios), provide a strong case for ancient recycled sediment in the Samoan mantle. Chapter 4 explores whether the eclogitic and peridotitic portions of ancient subducted oceanic plates can explain the anomalous titanium, tantalum and niobium (TITAN) enrichment in high ³He/⁴He ocean island basalts (OIBs). The peridotitic portion of ancient subducted plates can contribute high ³He/⁴He and, after processing in subduction zones, a refractory, rutile-bearing eclogite may contribute the positive TITAN anomalies.

Introduction

Unlike the earth's crust, it is difficult to sample the mantle directly. However, there are a few isolated (and geographically limited) instances where we can observe the uppermost regions of the mantle. Ophiolites—portions of oceanic plates that have been uplifted and emplaced (obducted) onto the edge of continents—expose the uppermost regions of the mantle on the surface (Dick, 1976). Similarly, at certain, ultra-slow spreading mid-ocean ridges, mantle peridotites are emplaced on the surface (Dick et al., 2003), where they can be recovered with deep submersibles or dredging operations. Additionally, ultramafic xenoliths, which are pieces of the upper mantle entrained in upwelling magmas, allow direct inspection of the composition of the upper mantle (but unlike ophiolites and abyssal peridotites recovered by submersibles, ultramafic xenoliths provide little spatial context for study of the mantle). Thus, except for these few rare instances, we cannot examine the composition of the mantle directly: The deepest hole ever drilled, only ~12.3 km deep, took 22 years and untold Soviet resources to complete-and got no where near mantle depths (unfortunately, the drill site was atop thick continental crust). While robots rove the surface of a planet, Mars, that is > 56,000,000km distant, we have not directly observed our own planet at depths >12 km!

Nonetheless, there are indirect methods for evaluating the composition of the deeper earth. Mantle geochemists often use lavas erupted on the surface as "windows" to the composition of the mantle below. Hotspot lavas erupted in oceanic settings, or oceanic island basalts (OIBs), are formed by partially melting and upwelling, solid mantle. During melting, several radiogenic isotope systems and a number of trace element ratios remain unfractionated (or little fractionated, in the case of trace elements with similar compatibilities, at least when the partition coefficients are much less than the degree of melting) from the original, unmelted mantle. OIBs exhibit a great deal of isotopic and trace element heterogeneity, indicating that the mantle from which they were derived is also quite heterogeneous (Zindler and Hart, 1986). This fundamental observation leads to some of the most important questions in the field of mantle

geochemistry: How did the mantle become heterogeneous? At what scale lengths do the heterogeneities exist? How long do the heterogeneities survive?

One common paradigm in mantle geochemistry assumes that oceanic plates, which form by melting and depleting the upper mantle at mid-ocean ridges, are subducted back into the mantle from which they formed. Covered with a veneer of oceanic/continental sediment, subducted oceanic plates inject sediment, oceanic crust and depleted peridotitic mantle lithosphere into the mantle (Hofmann and White, 1982; White and Hofmann, 1982). In this way, mantle heterogeneities are born. Following storage in the mantle, subducted plates and sediment are thought to be sampled by upwelling plumes that melt and erupt lavas on the surface. However, a counteracting mechanism chaotic mantle convection—stretches, thins, mixes and stirs, and thus homogenizes (or at least greatly attenuates) mantle heterogeneities on long timescales.

In Chapter 1, isotopic heterogeneities are explored at very short lengthscales in olivine-hosted melt inclusions in oceanic OIBs. Using a laser ablation system coupled to a MC-ICPMS (multi-collector inductively coupled plasma mass spectrometer), ⁸⁷Sr/⁸⁶Sr was measured in olivine-hosted melt inclusions recovered from Samoan basalts. Complementing the pioneering work on Pb-isotopes in olivine-hosted melt inclusions from two Polynesian hotspots (Saal et al., 1998), significant Sr-isotope heterogeneity was also observed in the melt inclusions from individual Samoan basalt hand samples. Melt inclusions in one Samoan lava exhibit a range of ⁸⁷Sr/⁸⁶Sr from 0.70686 to 0.70926. The isotopic diversity hosted in the melt inclusions from a single lava indicate that the size of the melting zone beneath a Samoan volcano can be larger than the lengthscales of mantle heterogeneities in the mantle upwelling beneath the hotspot. Furthermore, none of the 41 melt inclusions analyzed exhibit ⁸⁷Sr/⁸⁶Sr ratios lower than the least radiogenic whole-rock basalts in Samoa (⁸⁷Sr/⁸⁶Sr =0.7044). This ⁸⁷Sr/⁸⁶Sr data, combined with trace element data on the same melt inclusions, provide strong evidence against assimilation of oceanic crust as the source of the isotopic diversity in the melt inclusions.

Chapter 3 also deals with lengthscales of mantle heterogeneity, but by comparison to Chapter 1, Chapter 3 considers the largest possible lengthscales of heterogeneity in the Earth's mantle. The high ³He/⁴He (or FOZO, Focus Zone; Hart et al., 1992) mantle

reservoir, a domain that is considered to be one of the oldest (and deepest?) reservoirs in the mantle, is the focus of chapter 3. In this chapter, the highest magmatic ${}^{3}\text{He}/{}^{4}\text{He}$ ratios ever recorded in a southern hemisphere lava (33.8 times atmospheric) are reported in samples from the Samoan island of Ofu. These new measurements from Ofu Island place Samoa in the same category of high ³He/⁴He hotspots as Hawaii, Iceland and the Galapagos. Along with Samoa, the highest ³He/⁴He sample from each southern hemisphere high ³He/⁴He hotspot exhibits lower ¹⁴³Nd/¹⁴⁴Nd ratios than their counterparts in the northern hemisphere (excluding lavas erupted in continental, back-arc, and submarine ridge environments). The observation of a large-scale isotopic enrichment in the FOZO-A (austral) high ³He/⁴He mantle compared to the FOZO-B (boreal) high ³He/⁴He mantle is similar to the DUPAL anomaly, a globe-encircling feature of isotopic enrichment observed primarily in southern hemisphere ocean island basalts. The possible existence of hemispheric-scale heterogeneity in one of the oldest reservoirs in the mantle has important implications for mantle dynamics. It suggests that regions of the (lower?) mantle have escaped the rapid convection motions that dominate the upper mantle. However, the origin of the hemispheric-scale heterogeneity in the FOZO (and DUPAL) reservoir is unknown.

Nonetheless, having defined the variability that exists in the high ³He/⁴He mantle, Chapter 3 also explores whether or not the FOZO reservoirs are truly depleted, as is commonly suggested (Hart et al., 1992), or whether they have been re-enriched. The recent discovery of superchondritic ¹⁴²Nd/¹⁴⁴Nd ratios in terrestrial (Boyet and Carlson, 2005), martian and lunar (Caro et al., 2007) suggests that bulk silicate earth (BSE) may have superchondritic Sm/Nd ratios (¹⁴⁷Sm/¹⁴⁴Nd>0.209, the minimum ratio necessary to generate the terrestrial mantle ¹⁴²Nd/¹⁴⁴Nd anomaly relative to chondrites), and that the earth has a minimum ¹⁴³Nd/¹⁴⁴Nd of 0.51304. If this is true, then the FOZO reservoirs are actually enriched relative to BSE.

Chapter 2 explores a common paradigm in mantle geochemistry, that subduction of marine/continental sediments can generate geochemically enriched mantle domains that can be sampled by mantle upwellings. In this chapter, remarkably high ⁸⁷Sr/⁸⁶Sr ratios are reported in submarine lavas recovered from the flanks of the Samoan island of

Savai'i. These Savai'i lavas exhibit the highest ⁸⁷Sr/⁸⁶Sr ratios reported for ocean island basalts to date. The isotope and trace element data are consistent with the presence of a recycled sediment component (with a composition similar to the upper continental crust) in the Samoan mantle. Importantly, Pb-isotopes in the most enriched Samoan lavas preclude contamination by modern-marine sediment. The ultra-enriched Samoan lavas have most certainly been "contaminated" by sediment, but the sediment is of an ancient (>> 200 Ma) origin and has been recycled into the Samoan mantle source. In summary, Chapter 3 provides the strongest evidence yet that the sediment that goes down in subduction zones does come back up in OIBs. However, given the large mass of sediment that has been subducted into the mantle over geologic history, it is still a mystery why clear signatures of sediment recycling are so rare in OIBs (Hofmann, 1997).

Chapter 4 reports evidence for radiogenic ¹⁸⁷Os/¹⁸⁸Os and enrichment in Titanium, Tantalum and Niobium (TITAN) in high ³He/⁴He lavas globally. To explain these observations, the dominant paradigm for the formation of mantle heterogeneity is applied to the high ³He/⁴He reservoir: Can the subduction of oceanic plates (crust and peridotite) generate the geochemical signatures associated with the high ³He/⁴He reservoir? Radiogenic ¹⁸⁷Os/¹⁸⁸Os and TITAN enrichment are both geochemical signatures that are associated with recycled eclogite, suggesting that the high ³He/⁴He lavas were derived from a mantle source hosting a recycled slab component However, eclogites are quantitatively degassed in subduction zones and do not have intrinsically high ³He/⁴He. None-the-less, the peridotitic portion of recycled slabs has been suggested to preserve high ³He/⁴He over time (e.g., Parman et al., 2005). The eclogitic and peridotitic portions of subducted plates are intimately associated in space and time, a geometry that is conducive to later mixing in the mantle. Thus, together, the two lithologies can provide the "raw materials" for the formation of the high ³He/⁴He mantle.

Importantly, the TITAN enrichment in high ³He/⁴He mantle sampled by oceanic hotspot lavas may provide a clue about the location of the "missing" TITAN in the earth. Shallow geochemical reservoirs in the earth—continental crust and the depleted mid-ocean ridge basalt mantle (DMM)—have a shortage of the element Ti, Ta and Nb (TITAN) (McDonough, 1991; Rudnick et al., 2000). The observation of TITAN

enrichment in high ³He/⁴He OIB lavas suggests that the mantle domain hosting the Earth's "missing" TITAN is sampled by deep, high ³He/⁴He mantle plumes.

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Chapter 1

Strontium isotopes in melt inclusions from Samoan basalts: Implications for heterogeneity in the Samoan plume*

Abstract

We measured ⁸⁷Sr/⁸⁶Sr ratios on 41 olivine-hosted melt inclusions from nine Samoan basalts using laser ablation multi-collector (LA-MC) ICPMS. ⁸⁷Sr/⁸⁶Sr ratios are corrected for mass bias after eliminating major isobaric interferences from Rb and Kr. The external precision averages ± 320 ppm (2 σ) for the ⁸⁷Sr/⁸⁶Sr ratios on natural Samoan basalt glass standards of a similar composition to the melt inclusions.

All of the Sr-isotope ratios measured by LA-MC-ICPMS on Samoan melt inclusions fall within the range measured on whole rocks using conventional methods. However, melt inclusions from two Samoan basalt bulk rock samples are extremely heterogeneous in ⁸⁷Sr/⁸⁶Sr (0.70459–0.70926), covering 70% of the variability observed in ocean island basalts worldwide and nearly all of the variability observed in the Samoan island chain (0.7044–0.7089). Seven melt inclusions from a third high ³He/⁴He Samoan basalt are isotopically homogeneous and exhibit ⁸⁷Sr/⁸⁶Sr values from 0.70434 to 0.70469.

Several melt inclusions yield ⁸⁷Sr/⁸⁶Sr ratios higher than their host rock, indicating that assimilation of oceanic crust and lithosphere is not the likely mechanism contributing to the isotopic variability in these melt inclusions. Additionally, none of the 41 melt inclusions analyzed exhibit ⁸⁷Sr/⁸⁶Sr ratios lower than the least radiogenic basalts in Samoa (⁸⁷Sr/⁸⁶Sr=0.7044), within the quoted external precision. This provides an additional argument against assimilation of oceanic crust and lithosphere as the source of the isotopic diversity in the melt inclusions.

The trace element and isotopic diversity in Samoan melt inclusions can be modeled by aggregated fractional melting of two sources: A high ${}^{3}\text{He}/{}^{4}\text{He}$ source and an EM2 (enriched mantle 2) source. Melts of these two sources mix to generate the isotopic diversity in the Samoan melt inclusions. However, the melt inclusions from a basalt with the highest ${}^{3}\text{He}/{}^{4}\text{He}$ ratios in Samoa exhibit no evidence of an enriched component, but can be modeled as melts of a pure high ${}^{3}\text{He}/{}^{4}\text{He}$ mantle source.

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Strontium isotopes in melt inclusions from Samoan basalts: Implications for heterogeneity in the Samoan plume

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Abstract

We measured ⁸⁷St/⁸⁶St ratios on 41 olivine-hosted melt inclusions from nine Samoan basalts using laser ablation multi-collector (LA-MC) ICPMS. ⁸⁷St/⁸⁶St ratios are corrected for mass bias after eliminating major isobaric interferences from Rb and Kr. The external precision averages ± 320 ppm (2 σ) for the ⁸⁷St/⁸⁶St ratios on natural Samoan basalt glass standards of a similar composition to the melt inclusions.

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Keywords: 87Sr/86Sr; laser ablation; MC-ICPMS; melt inclusion; Samoa; EM2; PHEM; FOZO

1. Introduction

Ocean island basalts (OIBs) erupted at hotspots are thought to be the surface expression of buoyantly upwelling mantle plumes that sample the mantle's compositional heterogeneities at various depths and times

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[1,2]. The Samoan islands and seamounts, formed by a mantle plume impinging on the Pacific plate just north of the Tonga Trench, form a time-progressive hotspot track [3,4] which conforms reasonably well to Morgan's hotspot model [5]. Samoan lavas exhibit the highest 87 Sr/ 86 Sr ratios and the largest 87 Sr/ 86 Sr variation (0.7044–0.7089) measured in fresh OIBs [4,6], making them ideal for prospecting for diverse Sr-isotope compositions in melt inclusions.

Olivine-hosted melt inclusions in Samoan lavas provide snapshots of diverse magma chemistry before complete melt aggregation, providing an opportunity to see more of the isotopic heterogeneity which exists in the melt source but that is not detectable in whole rocks. However, the chemical variability in melt inclusions may be generated by a number of processes that obscure source variation, including pre-entrapment fractional crystallization, post-entrapment diffusive reequilibration, crustal assimilation, and degree, type and depth of melting [7-16].

Studies delineating Pb-isotope diversity in melt inclusions have demonstrated that heterogeneous melt source compositions are an important factor in generating compositional variability [17-21]. A landmark Pbisotope study of melt inclusions hosted in basalts from Mangaia Island in the Cook Islands revealed significantly more isotopic heterogeneity than is found in whole rocks from the island [17]. The results indicate the presence of an unradiogenic Pb-isotope endmember in the melt inclusions not discernable in whole-rock basalts. Problematically, this unradiogenic Pb endmember has been poorly characterized, owing partly to the large uncertainties associated with in situ Pb-isotope measurements: Pb-isotope data from melt inclusions generally are limited to ²⁰⁸Pb/²⁰⁶Pb versus ²⁰⁷Pb/²⁰⁶Pb isotope projections (due to the inability to collect precise ²⁰⁴Pb data on silicate melt inclusions), which place DMM (depleted MORB mantle, low ³He/⁴He, low ⁸⁷Sr/⁸⁶Sr), FOZO (Focus Zone, high ³He/⁴He, low ⁸⁷Sr/⁸⁶Sr), PHEM (Primitive Helium Mantle, high 3He/4He, middle-range 87 Sr/86 Sr) and EM2 (enriched mantle 2, low ³He/⁴He, high ⁸⁷Sr/⁸⁶Sr) in such close graphical proximity that they cannot be unequivocally resolved. The true pedigree of the unradiogenic Pb endmember in Mangaia is still unknown, and could be similar to any of these four endmembers.

An advantage to the Sr-isotope system is that the EM2 endmember has dramatically higher 87 Sr/ 86 Sr ratios (~0.7089) [4] than the DMM (0.7026) [22], PHEM (0.7045) [23] and FOZO (0.7030) [24] mantle endmembers, and can be readily differentiated from the three less radiogenic endmembers. PHEM hosts signif-

icantly more radiogenic Sr than DMM and FOZO and is easily resolved from these two components. Unfortunately, DMM and FOZO exhibit similar ⁸⁷Sr/⁸⁶Sr ratios and it will be difficult to differentiate between these two components as potential sources of the isotopic diversity in melt inclusions.

We present Sr-isotope data from olivine-hosted melt inclusions recovered from Samoan basalts, some of which lie near the EM2 mantle endmember, with the goal of better understanding the puzzling unradiogenic component sampled by melt inclusions. To this end, we also contribute Sr-isotope data measured on melt inclusions from a recently discovered high ³He/⁴He basalt from Samoa [25]. Our strategy is to analyze Sr isotopes in melt inclusions from EM2 and high ³He/⁴He role of the various components–EM2, PHEM, DMM and FOZO–that may be contributing to the Sr-isotope diversity in the Samoan plume.

2. Methods

A detailed description of the protocol used for in situ measurement of Sr isotopes in basaltic glasses (and melt inclusions) by LA-MC-ICPMS is provided in the Supplementary data. In order to measure Sr-isotope ratios in situ, we use a 213 nm NewWave laser ablation system coupled to a Themo-Finnigan Neptune MC-ICPMS, located in the Plasma Facility at Woods Hole Oceano-graphic Institution (WHOI). During analytical runs, the laser is run in aperture mode with 100% power, a pulse rate of 20 Hz and a spot size of 120 μ m. The raster pattern varies depending on the size and shape of the melt inclusion, and the line speed is 4 μ m/s. Surface contamination is removed by pre-ablation using the same raster and spot size, but with a pulse rate of 5 Hz, 45% power and a raster speed of 30 μ m/s.

During each analytical session, we measure intensities on masses 82 through 88. Raw data are exported to an offline data correction program (TweaKr) for correcting the Rb and Kr isobaric interferences. Runs with low intensities (i.e., <1 V on mass 88, due to small size or low Sr content) were discarded as they are prone to large systematic errors [26]. Masses 85 and 88 are pure Rb and Sr, respectively, with no significant known interferences, and require correction only for mass fractionation. We correct for Kr interferences on masses 84 and 86 so that the mass fractionation-corrected ⁸⁴Sr/⁸⁶Sr value is canonical (0.0565725). The protocol for correcting mass 87 for the Rb interference is the following: A Samoan basalt glass with known ⁸⁷Sr/⁸⁶Sr, from analysis by conventional Thermal Ionization Mass Spectrometry

															2	3	
Sample name and grain #	Disc #	Mass 82 meas'd-V	Mass 83 meas'd-V	Mass 84 meas'd-V	Mass 85 meas [°] d-V	Mass 86 meas'd-V	Mass 87 meas ² d-V	Mass 88 meas'd-V	⁸⁵ Rb/ ⁸⁷ Rb required	⁸⁷ Sr/ ⁸⁶ Sr corr'd ^a	⁸⁷ Sr/ ⁸⁶ Sr corr'd ^b	Int. prec.	Rb/Sr I	nt.	# Cycles	Type	Oliv Fo
AVON3-78-1											0.70890		0.1255				
(whole rock):																	
78-1Z	Yearl	0.0013	0.0011	0.0193693	0.158083	0.263023	0.250903	2.25759	2.58966	0.70791	0.70784	87	0.0746 1	.82	14°	0	84.6
78-1#7 ^d	Disc1	0.0020	0.0022	0.0193323	0.164332	0.188415	0.199670	1.60867	2.58811	0.70927	0.70926	123	0.1081 2	47	6	Η	84.6
78-1#9 ^d	Disc1	0.0019	0.0021	0.0178511	0.111482	0.163039	0.159876	1.38845	2.58811	0.70733	0.70732	103	0.0850 1	.90	9	Η	84.4
78-1#6 ^d	Disc2	0.0031	0.0034	0.0297219	0.341900	0.308400	0.354647	2.62064	2.58593	0.70846	0.70843	66	0.1397 2	.63	32	Η	83.6
78-1#9 ^d	Disc2	0.0031	0.0034	0.0316443	0.384820	0.340724	0.395010	2.89882	2.58593	0.70873	0.70870	94	0.1421 3	.08	30	Н	82.8
78-1#10 ^d	Disc2	0.0030	0.0031	0.0241696	0.094343	0.202069	0.179689	1.70410	2.58663	0.70845	0.70842	104	0.0593 3	21	16	Н	83.6
78-1#4 ^d	Disc3	0.0049	0.0057	0.0638080	0.704962	0.756968	0.817352	6.43584	2.58694	0.70869	0.70866	44	0.1179 0	09.0	40	Ċ	85.2
78-linside	Disc3	0.0048	0.0049	0.0311385	0.155536	0.152840	0.165959	1.25071	2.58564	0.70689	0.70686	139	0.1337 2	.20	16	0	84.3
78-1B	Disc3	0.0044	0.0043	0.0327454	0.189114	0.222802	0.230356	1.86119	2.58675	0.70791	0.70788	127	0.1089 5	.76	10	0	84.2
78-1C	Disc3	0.0041	0.0043	0.0408646	0.414636	0.401444	0.448986	3.40453	2.58675	0.70866	0.70863	38	0.1304 1	96.	24	0	83.9
78-1D	Disc3	0.0048	0.0048	0.0313504	0.130366	0.156852	0.158811	1.28485	2.58675	0.70699	0.70696	119	0.1091 3	.79	14	0	85.3
AVON3-71-2											0.70594		0.0754				
(whole rock):																	
71-2C ^d	Year1	0.0015	0.0012	0.0241555	0.131881	0.348997	0.301171	2.99950	2.58977	0.70488	0.70481	50	0.0469 4	66.	10°	5	88.8
71-2C°	Disc3	0.0049	0.0051	0.0423813	0.151006	0.357484	0.309961	3.00928	2.58772	0.70487	0.70484	09	0.0539 2	60"	19	9	88.8
71-2Da	Yearl	0.0014	0.0012	0.0190549	0.143364	0.258958	0.241512	2.22256	2.58977	0.70555	0.70549	53	0.0687 3	.92	12 ^c	c	90.1
71-2i_mi1 ^d	Disc1	0.0041	0.0042	0.0293387	0.151118	0.188888	0.190950	1.57393	2.58696	0.70605	0.70602	42	0.1028 0	1.89	19	9	90.7
71-2#6_saol1	Disc1	0.0040	0.0044	0.0392836	0.237282	0.399731	0.376067	3.39132	2.58675	0.70532	0.70529	63	0.0750 2	.45	32	c	90.3
71-2#6_saol2	Disc1	0.0038	0.0039	0.0311684	0.151330	0.259691	0.241924	2.18859	2.58675	0.70511	0.70508	06	0.0741 3	.15	15	0	90.5
71-2pentagon	Disc1	0.0035	0.0036	0.0274768	0.164455	0.221630	0.220515	1.86556	2.58653	0.70603	0.70600	69	0.0944 1	.40	35	c	91.1
71-2#S ^d	Disc3	0.0049	0.0055	0.0608667	0.440578	0.704362	0.672560	5.98503	2.58694	0.70537	0.70534	34	0.0792 1	-23	40	9	91.0
71-2#5°	Disc3	0.0049	0.0057	0.0600272	0.456791	0.688053	0.667289	5.84446	2.58694	0.70537	0.70534	34	0.0841 1	.21	40	9	91.0
71-2#6 ^d	Disc3	0.0050	0.0053	0.0461581	0.261970	0.420138	0.399032	3.54513	2.58674	0.70602	0.70600	53	0.0794 1	.35	40	9	85.7
71-2#6°	Disc3	0.0050	0.0054	0.0511361	0.320998	0.514222	0.489758	4.35389	2.58674	0.70600	0.70598	32	0.0793 1	.19	40	Ċ	85.7
71-2Z	Disc3	0.0041	0.0042	0.0286629	0.120538	0.176062	0.169532	1.46336	2.58809	0.70576	0.70573	16	0.0882 2	.60	21	0	86.1
71-2N	Disc3	0.0041	0.0043	0.0326106	0.171904	0.249420	0.242412	2.09517	2.58778	0.70554	0.70551	48	0.0879 2	14	26	0	505
71-2imi_run2	Disc3	0.0046	0.0048	0.0392641	0.162916	0.326152	0.292704	2.74335	2.58778	0.70520	0.70518	112	0.0638 2	.80	12	0	91.0
71-2V	Disc3	0.0034	0.0037	0.0358138	0.196791	0.370738	0.339729	3.15166	2.58754	0.70461	0.70459	122	0.0668 4	.32	13	c	87.4
OFU-04-06											0.70458		0.0458				
(whole rock): OFLL04_06_13 ^d	Ofi	0.0012	0.0004	0.0123598	0.087599	0.202878	0 179279	1 74009	2 58774	0.70458	0.70457	900	0.0541.6	104	9	Ħ	20.8
OFU-04-06-15 ^d	Ofi	0 0008	0.0006	0.0166800	0.116272	0.260130	0.231639	2.23575	2 58791	0.70443	0.70443	128	0.0556 1	3	2 12	: =	77.3

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H-										0.70550		0.0700				
ick):																
inter i an			0.000000			0000000									(
Disc1 0.0	0035	0.0035	0.0250240	0.083017	0.152405	0.138888	1.27387	2.28654	0.70653	0.70653	83	0.0691	1.22	16	5	\$3.0
Disc1 0.0	6100	0.0021	0.0204900	0.125897	0.216562	0.204083	1.85257	2.58857	0.70692	0.70692	109	0.0720	0.92	15	U	83.0
Disc3 0.0	0048	0.0048	0.0316250	0.103649	0.160366	0.150752	1.31592	2.58772	0.70683	0.70680	144	0.0846	5.70	8	U	81.4
										0.70539		0.0688				
ick):																
Disc1 0.0	0020	0.0021	0.0186133	0.117638	0.165627	0.163878	1.40909	2.58798	0.70625	0.70624	75	0.0884	3.01	20	G	87.9
Disc1 0.0	0038	0.0040	0.0305189	0.209211	0.258789	0.264567	2.18215	2.58653	0.70577	0.70575	87	0.1027	1.92	12	U	87.9
Disc1 0.0	0037	0.0041	0.0371899	0.294848	0.389761	0.392493	3.31142	2.58653	0.70577	0.70574	93	0.0954	1.73	6	C	87.9
Disc3 0.0	0041	0.0042	0.0302365	0.092078	0.197293	0.173102	1.64414	2.58754	0.70482	0.70480	144	0.0600	6.45	10	c	90.1
-2										0.70540		0.0679				
ick):																
Disc3 0.0	0049	0.0051	0.0641890	0.514757	0.759082	0.741869	6.46111	2.58815	0.70599	0.70596	58	0.0856	2.68	31	υ	1
Disc3 0.0	0047	0.0056	0.0601701	0.544474	0.708353	0.718034	6.03171	2.58815	0.70616	0.70613	69	0.0969	1.22	36	U	I
Disc3 0.0	0044	0.0045	0.0283814	0.086324	0.138106	0.128251	1.13104	2.58815	0.70522	0.70520	181	0.0818	4.84	Π	C	87.1
Disc3 0.0	0044	0.0045	0.0334163	0.133112	0.234231	0.215648	1.95929	2.58815	0.70528	0.70525	55	0.0728	3.15	23	c	86.6
ī.										0.70469		0.0484				
ick):																
Yearl 0.0	0015	0.0012	0.0174643	0.076749	0.224099	0.189719	1.91969	2.58895	0.70483	0.70477	157	0.0426	1.38	90	c	85.1
										0.70620		0.0597				
Disc3 01	1048	0.0048	0.0330020	0.078200	0 180374	0.161268	1 56507	2 58564	0.20500	0.70518	361	0.0530	613	11	C	76.5
		0.000		1000			in and in	-		0.70889	1	0.1108)	20
ick):																
Yearl 0.0	0014	0.0012	0.0184433	0.207247	0.241204	0.254752	2.06790	2.58966	0.70814	0.70807	61	0.1068	3.51	18 ^c	U	84.1
10 01	101.01	1 611	-	0 0 0	1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1		11.1.1		1 0 0	-		- 4	100			
, except for OFL	J-04-06 at	nd S11, are c	iredge sample	s from the A	VON3 dredg	ng cruise. M	elt inclusion	type refers to	o whether it	was homoge.	nized ()	I), glassy	(C) o	r crysta	lline (C). Host
								0 - 10								
-3 ock): Yearl 0.(, except for OFU	0014 1-04-06 ar	0.0012 ad S11, are d	0.0184433 iredge samples	0.207247 6 from the A	0.241204 VON3 dredgi	0.254752 ng cruise. M	2.06790 elt inclusion	2.58966 type refers to	0.70814 Whether it	0.7088 0.7080 vas hom	9 0ge	9 17 61 ogenized (F	9 0.1108 7 61 0.1068 ogenized (H), glassy	9 0.1108 7 61 0.1068 3.51 ogenized (H), glassy (G) α	 0.1108 61 0.1068 3.51 18^e ogenized (H), glassy (G) or crysta 	9 0.1108 7 61 0.1068 3.51 18 ^e C ogenized (H), glassy (G) or crystalline (C
, except for OFL	J-04-06 at	nd S11, are o	iredge sample	s from the A	VON3 dredgi	ng cruise. M	elt inclusic	10	on type refers to	on type refers to whether it v	on type refers to whether it was homoge	in type refers to whether it was homogenized (1	on type refers to whether it was homogenized (H), glassy	on type refers to whether it was homogenized (H), glassy (G) o	in type refers to whether it was homogenized (H), glassy (G) or crysta	in type refers to whether it was homogenized (H), glassy (G) or crystalline (C

for the isobaric Rb interference using the Samoan glass standard analyses from those 5 days (see Supplementary data). Melt inclusion Sr-isotope and Rb/Sr data from a particular sample are listed immediately below the equivalent whole-rock data.

^a Sr-isotopes are corrected for Kr and Rb interferences and fractionation-corrected.

^b After correction for Rb and Sr interferences and fractionation, melt inclusions are corrected for 987 standard solution runs for each analytical session. ^c Cycle length for melt inclusions from disc "Year1" are 16 s integrations during analysis. Cycles for all other measurements are 8 s integration.

^d Has major and/or trace element data in Table 2. ^e Replicate analysis of a melt inclusion is listed immediately below first analysis. Six melt inclusions were large enough to permit replicate analyses. Melt inclusion sample Yearl_71-2 was replicated over a 1-year period, and was renamed Disc3_71-2C for the replicate analysis. All other replicate analyses were made during the same analytical session.

(TIMS), is measured by laser ablation, Kr-corrected, and the ⁸⁵Rb/⁸⁷Rb ratio is adjusted until the known TIMS 87 Sr/86 Sr value of the glass is achieved. We adopted the strategy of using the required 85Rb/87Rb of the Samoan basalt glass standards to bracket the 85Rb/ ⁸⁷Rb of the basalt glass unknowns. In order to estimate the overall accuracy associated with this technique, we apply contiguous bracketing of the glass standard runs, and we are able to reproduce the known TIMS values to within an average of ± 320 ppm (2 σ standard deviation). However, due to uncertainty associated with the 85 Rb/ ⁸⁷Rb ratio (2.5875 \pm 0.00275, 2 σ), the final, corrected 87Sr/86Sr ratio exhibits an error magnification that is directly proportional to the Rb/Sr ratio of the sample. Samples with low Rb/Sr will exhibit less error from the Rb correction (145 ppm, Rb/Sr=0.04) than samples with high Rb/Sr (505 ppm, Rb/Sr=0.14) (see Supplementary data). However, over the range of Rb/Sr in the Samoan basalt glass standards (0.045-0.126), we find no relationship between the internal (in-run) precision of 87Sr/86Sr ratios (which average 45 ppm, 26 standard deviation) and Rb/Sr during analyses of basalt glasses by laser ablation. Similarly, there does not appear to be a relationship between Kr/Sr and the internal precision or reproducibility (external precision) of the ⁸⁷Sr/⁸⁶Sr ratios in Samoan glass standards over the range of ratios that we have observed during melt inclusion analysis (⁸²Kr/⁸⁸Sr from 0.00013 to 0.004). Finally, the reproducibility of 87Sr/86Sr measurement does not appear to be related to Sr intensity over the range of Sr intensities observed in lasering Samoan glasses, a range that encompasses the melt inclusion analyses.

Six melt inclusions were large enough for replicate analysis (one melt inclusion, 71-2C, was replicated over a one-year period), and five of their ⁸⁷Sr/⁸⁶Sr ratios were reproducible within the quoted precision. However, the replicate analysis of melt inclusion 71-11a was different by 550 ppm (see Table 1), while error resulting from the Rb correction is only 260 ppm (2σ) on Samoan glass standards with similar Rb/Sr ratios. The internal precision of the replicate analysis of this melt inclusion was ~100 ppm (2σ). Data from this melt inclusion indicates that larger-than-usual ⁸⁵Rb/⁸⁷Rb variations over time can occasionally generate uncertainties (above the 2σ level) in ⁸⁷Sr/⁸⁶Sr that are somewhat larger than error predicted by the data from Samoan glass standards.

An upper limit for the ⁸⁷Sr/⁸⁶Sr measurement precision on Samoan melt inclusions with low Rb/Sr can be inferred from the near-uniform ratios obtained on melt inclusions from the high ³He/⁴He Ofu basalt. The Rb/Sr values were among the lowest during analysis of Ofu melt inclusions, and the tight clustering of the Ofu melt inclusions may be partially explained by decreased error of ${}^{87}\text{Sr}{}^{86}\text{Sr}$ measurement for these samples compared to other, higher Rb/Sr Samoan glasses and melt inclusions from Vailulu'u and Malumalu. If we assume that the Ofu melt inclusions are isotopically homogeneous, then the external precision on these 7 melt inclusions is ± 335 ppm (2σ). Some of the apparent variability may be a result of error from the Rb correction, which is ± 190 ppm (2σ) at Rb/Sr ratios of 0.053, and may not reflect true variability. Additionally, internal precision varied from 60 to 226 ppm (2σ standard error) on the seven Ofu melt inclusions.

Masses 85 and 88 represent pure Rb and Sr, respectively, so that fairly precise measurement of Rb/Sr ratios can be generated. After correcting for mass fractionation ($\approx 1.5\%$ /amu), Rb/Sr ratios on Samoan basalt glasses measured by laser ablation are reproduceable to 17% (2σ , compared to ratios obtained by XRF/ICP techniques on the same samples), and precise (1.7%, 2σ) during multiple runs on the same glass (see Supplementary data).

Major element compositions of glassy and homogenized melt inclusions were obtained with a JEOL-733 automated electron microprobe at the Massachusetts Institute of Technology using an electron beam with current of 10 nA and accelerating potential of 15 kV focused to a spot of 1-2 µm in diameter for olivine analyses, and defocused to 10 µm for glass analyses. Trace element contents were determined with a Cameca IMS 3f ion microprobe following the techniques described in [27,28]. A small beam (5 µm diameter spot), combined with a high-energy filtering technique (80-100 eV window), was used to determine trace element concentrations. Precision for Sr, La, Zr, Y is estimated to be $\pm 15\%,$ and $\pm 20{-}30\%$ for Ba, Nb and Rb. Homogenization of olivine-hosted melt inclusions was performed in a furnace at 1187-1220 C (depending on olivine composition) at 1 atm pressure for 5 min in a graphite capsule.

To correct for the effects of crystallization of olivine in the glassy and homogenized melt inclusions, we add equilibrium olivine to the melt inclusions in 0.1% increments until equilibrium with mantle olivine (Fo₉₀) is achieved, assuming olivine-melt partitioning of Fe and Mg from [29]. Instead of correcting the melt inclusions to be in equilibrium with the host olivine, this correction scheme is chosen so that we can compare them to similarly corrected Samoan whole-rock lavas (after discarding data from the most evolved– MgO<6.5 wt.%–whole-rock samples).

3. Results

3.1. Sr-isotope variability in melt inclusions

Sr-isotopes were measured in melt inclusions from nine geochemically well-characterized basalt samples [4,25] from five islands and seamounts located along the Samoan hotspot track. Olivines (Fo_{76-91}) with large ellipsoidal melt inclusions ($50-250 \mu$ m diameter) were separated from the basalt samples for melt inclusion exposure and isotopic analysis. Most of the melt inclusions were crystalline and usually contained dendritic clinopyroxene in a glassy matrix, with spinel and rare sulfide globules; some ($\approx 5\%$) of the melt inclusions were glassy. The high ³He/⁴He basalt sample OFU-04-06 was unique in that amphibole and apatite were common melt inclusion phases, and carbonate was also observed.

Olivines (Fo₈₂₋₈₅) hosting melt inclusions were separated from Malumalu seamount dredge sample 78-1, a picrite with the highest 87 Sr/86 Sr ratio (0.7089) of any OIB [4]. Melt inclusion-rich olivines were also recovered from dredge sample 71-2 (Fo₈₄₋₉₁) from Vailulu'u, a seamount that displays intermediate enrichment relative to the other Samoan islands and seamounts (bulk rock 87Sr/86Sr isotope ratios from 0.7052 to 0.7067, n=20). From Ofu Island, olivines (Fo76-84) were recovered from ankaramite dike sample OFU-04-06, which exhibits the highest ³He/⁴He measured in a Samoan basalt [25]. Lavas from Ofu are the most isotopically homogeneous of the volcanoes in the Samoan chain (87Sr/86Sr from 0.70444 to 0.70480, n=12). Finally, a smaller number of olivines were separated for melt inclusion analysis from Vailulu'u dredge samples 71-11 (Fo₈₁₋₈₃), 63-2 (Fo₈₆₋₈₇) and 68-3 (Fo₈₇₋₉₀), Malumalu dredge sample 78-3 (Fo₈₄), Ta'u Island dredge sample 74-1 (Fo₈₅), and Savai'i subaerial post-crosional sample S11 (Fo76).

The Sr-isotope data from melt inclusions in just three Samoan whole-rock samples (OFU-04-06, 78-1 and 71-2) define a broad array that encompasses the entire spectrum of Sr-isotope variability (0.70434–0.70926) recorded in Samoan basalts (Fig. 1 and Table 1). Eleven melt inclusions from 78-1 display ⁸⁷Sr/⁸⁶Sr values of 0.70686–0.70926, and encompass over 30% of the isotope variability observed in the OIB mantle. Vailulu'u melt inclusions from dredge samples 71-2 (⁸⁷Sr/⁸⁶Sr=0.70459–0.70602, n=12 melt inclusions) and 68-3 (⁸⁷Sr/⁸⁶Sr=0.70480–0.70624, n=3) exhibit a smaller range of Sr-isotope values than 78-1. However, the magnitude of the Sr-isotope heterogeneity in the melt inclusions from these two samples is approxi-



Fig. 1. 87Sr/86Sr (upper panel) and Rb/Sr (lower panel) in melt inclusions from nine basalt samples compared to the total variation in each island/seamount (shown by boxes). References for whole-rock data can be found in [4]. The nine whole rocks from which the melt inclusions were separated are represented by larger symbols. No melt inclusions are more isotopically depleted than the most depleted lava measured in the island chain (87Sr/86Sr>0.7044). The range for EPR N-MORB [22] (averaged by segment) plots significantly below the Samoan whole rocks and melt inclusions. Replicate analyses on the same melt inclusion are averaged. The Rb/Sr for Savai'i whole-rock sample SAV1-25 is not included because of alteration. Internal precision (2σ, standard error) for 87Sr/86Sr is approximately the size of the data symbols. 87 Sr/86 Sr error bars represent error propagated from the Rb correction, as determined on Samoan glass standards, and dominates the error associated with 87Sr/86Sr measurement by LA-ICP-MS; maximum and minimum (500 and 150 ppm) errors are shown for reference. Error bars for Rb/Sr are $\pm 17\%$ (2 σ , standard deviation), and are based on the reproduceability of Rb/Sr measurement on Samoan glass standards; internal precision for Rb/Sr averages an order of magnitude better. Samples are listed in order of increasing whole-rock 87 Sr/86 Sr, from left to right.

mately equal to the variability observed in Vailulu'u seamount lavas, although the melt inclusions sample a more depleted component than observed in whole rocks from this seamount. Among the Vailulu'u samples, only sample 71-11 (87 Sr/ 86 Sr=0.70653-0.70692, *n*=2)

hosts inclusions that sample compositions similar to the most enriched isotopic compositions observed in Vailulu'u volcano. Like sample 71-11, isotopic analyses of melt inclusions from sample 63-2 (87Sr/ ⁸⁶Sr=0.70520-0.70613, n=3) lie completely within the isotopic range found in the whole rocks from Vailulu'u seamount. OFU-04-06 melt inclusions exhibit the least isotopic variability (87Sr/86Sr=0.70434-0.70469, n=7) among the samples with isotopic analyses on more than two different inclusions, and this variability is similar to the variability sampled by whole-rock lavas from the island. There is only a single melt inclusion analysis from each of basalt samples 78-3 (0.70807), 74-1 (0.70477) and S11 (0.70518); the Srisotope ratios lie within the range observed in the respective host volcanoes.

Although melt inclusions show more depleted isotopic compositions than the island or seamount from which they were recovered, they are not observed to sample compositions more depleted than whole rocks from the Samoan hotspot (87 Sr/ 86 Sr ≥ 0.7044). Therefore, the least radiogenic Sr-isotope ratios in Samoan basalts and melt inclusions are significantly more enriched than the depleted upper mantle sampled by MORB along the EPR (East Pacific Rise; 0.70228– 0.70287, N-MORB segment averages from [22]); the old oceanic crust upon which the Samoan island chain is constructed is probably isotopically similar to these modern EPR basalts.

Rb/Sr ratios measured by LA-ICP MS on the same set of melt inclusions tell a story similar to that of the isotopes. Melt inclusions in sample 78-1 exhibit the largest variation in Rb/Sr ratios (0.0593-0.1421), and the variability is similar to that observed in the whole rocks measured from Malumalu seamount. The Rb/Sr ratios in OFU-04-06 melt inclusions show some heterogeneity (Rb/Sr from 0.0476 to 0.0595, n=7), but this variability is smaller than the variability sampled by whole-rock lavas from the island. Vailulu'u melt inclusions from dredge samples 71-2 (Rb/Sr from 0.0469 to 0.1028, n=12 melt inclusions) and 68-3 (Rb/Sr from 0.060 to 0.1027, n=3) exhibit a range of Rb/Sr values that falls between 78-1 and OFU-04-06. Unlike the isotopes, however, the magnitude of the Rb/Sr heterogeneity in the melt inclusions from samples 71-2 and 68-3 is greater than the variability observed in Vailulu'u seamount lavas, and, within error of measurement, do not sample a component with lower Rb/Sr than observed in whole rocks from this seamount.

The present dataset suggests that the isotopic variability exhibited by the melt inclusions in a basalt sample may be a function of the whole-rock isotopic composition. In Fig. 2, the Sr-isotope variability of melt inclusions in three basalt samples-those with the largest number of melt inclusion analyses-is plotted against the bulk 87 Sr/ 86 Sr composition of the respective whole rocks. The 87 Sr/ 86 Sr variability, determined by the difference between the highest and the lowest 87 Sr/ 86 Sr



Fig. 2. Upper panel: Melt inclusion isotopic variability (in ppm) as a function of whole-rock 87Sr/86Sr ratio for the three Samoan lavas with the greatest number of melt inclusion analyses. Melt inclusion isotopic variability increases with increasing 87Sr/86Sr (and decreasing 3He/ ⁴He, not shown) in the three Samoan basalts. Variability is determined by the difference between the most isotopically extreme melt inclusions in a basalt sample. The maximum and minimum values for this variability are a result of the uncertainty introduced by the Rb correction, which is directly related to the Rb/Sr of the melt inclusion. For example, the maximum variability in basalt sample 78-1 is determined by the difference between the highest (0.7092 plus 391 ppm uncertainty) and lowest plausible 87Sr/86Sr (0.7072 minus 393 ppm uncertainty). The minimum variability is determined by the difference between the lowest plausible 87Sr/86Sr in the most enriched melt inclusion (i.e., 0.7089, or 0.7092 minus 389 ppm uncertainty) and the highest plausible 87Sr/86Sr in the most depleted melt inclusion (0.7074, or 0.7072 plus 393 ppm uncertainty). Lower panel: Melt inclusion Rb/Sr variability (absolute) as a function of whole-rock 87Sr/ ⁸⁶Sr ratio. Maximum and minimum variability is determined the same way, but assumes an uncertainty for Rb/Sr of 17%. The magnitude of variability due to internal precision is approximated by the size of the data symbols in both panels.

ratios from melt inclusions in a given basalt sample, exhibits a maximum and a minimum due to the uncertainty introduced by the Rb correction. Although the number of data points is limited, the data are consistent with melt inclusion isotopic diversity increasing with increasing 87 Sr/86 Sr (increasing EM2 component) and decreasing 3He/4He (not shown). The range of variability for Malumalu sample 78-1 is larger than, but overlaps with, the range of values from Vailulu'u sample 71-2. The melt inclusions from the high ³He/⁴He Ofu basalt exhibit the smallest range of probable 87Sr/ ⁸⁶Sr ratios, and they do not overlap with the range from samples 71-2 and 78-1. A similar observation can be made for the variability of Rb/Sr ratios in the melt inclusions, where Rb/Sr tends toward greater melt inclusion variability in samples 78-1 and 71-2; the Ofu sample has the smallest range of variability, and overlaps slightly with the lowest probable variability in sample 71-2. It is notable that OFU-04-06 exhibits the highest ³He/⁴He ratio (and low ⁸⁷Sr/⁸⁶Sr) found in Samoa, an observation that may be linked to the small

degree of isotopic and trace element variability in its melt inclusions.

3.2. Major and trace element characteristics of melt inclusions

Rb/Sr and ⁸⁷Sr/⁸⁶Sr ratios from the melt inclusions define a broad array that encompasses the entire spectrum of Sr-isotope and Rb/Sr variability recorded in Samoan basalts (Fig. 3). Curiously, the melt inclusions from Malumalu do not form an array by themselves, but plot over a broad region. The Rb/Sr and Sr-isotope data array form a crude mantle isochron of 1.1 Ga.

Major and trace elements were measured on melt inclusions from Vailulu'u, Malumalu and Ofu basalts, and they reveal a large range of compositions (Table 2). Although the trace element compositions of Vailulu'u melt inclusions are similar to whole-rock analyses from this seamount, melt inclusions from EM2 endmember basalt 78-1 record a greater degree of trace element variability than all of the whole rocks measured from Malumalu. One



Fig. 3. 8^{75}Cr^{86} Sr versus Rb/Sr ratios for Samoan whole rocks and melt inclusions, determined by laser ablation. The shaded regions represent the extent of melt inclusion variability for the volcances (Vailulu'u, Malumalu and Ofu). The open region encompasses the least evolved (MgO ≥ 6.5 wt.%) whole-rock measurements from these three volcances and Ta'u, thus encompassing islands and seamounts only on the eastern half of the Samoan hotspot track (where ~98% of the analyzed melt inclusions were recovered). Whole-rock data from Ofu is unpublished, and for the other volcances is from [4]. MORB range is limited to EPR N-MORB (M) segment averages [22]. High ³He/⁴He basalts from Hawaii (H) [37], Iceland (I) [38] and Baffin Island (B) [39] provide an approximation for FOZO (higher ³He/⁴He ratios exist for Hawaiian basalts [40], but ⁸⁷Sr/⁸⁵Sr data are not available for these samples; Rb/Sr data for high ³He/⁴He samples from Baffin Island are found in [41]). Error bars on symbols are internal precision of measurement (2σ , mean deviation). Error bars on periphery of figure denote external precision of measurements, as determined by Samoan glass standards with compositions similar to the melt inclusion. Replicate analyses of the same melt inclusion are shown. The melt inclusion data form a crude mantle isochron age of ~1.1 Ga.

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Melt inclusion	n major	and trace	eleme	ıt data	(uncorre	cted for	r olivine	A DEVISION	(Property)	0.000	0.100000000000		TOTTON											1
Sample #	Disc #	Grain #	SiO2	TiO ₂	Al_2O_3	FeO	MnO	MgO	CaO	$\mathrm{Na_2O}$	$\mathbf{K}_2\mathbf{O}$	$\mathrm{P}_{2}\mathrm{O}_{5}$	Total	Mg #	OIFo	Rb	Sr .	X	Zr	Nb	Ba	La	$^{87}\mathrm{Sr}/^{86}\mathrm{Sr}^{a}$	Type
78-1	Disc1	上井	45.64	3.77	12.97	14.11	0.14	6.85	9.53	3.46	2.16	0.46	99.19	49.0	84.6	181	176	83.6	979	1	ī	ī	0.70926	Н
78-1	Disc1	6#	46.57	2.99	13.06	10.75	0.16	8.02	13.09	2.94	1.57	0.44	99.74	59.6	84.4	49	443	33.0	245	3	1	45.6	0.70732	Η
78-1	Disc1	#10-11	44.98	3.22	12.84	11.18	0.14	9.26	11.29	2.62	1.76	0.39	98.10	62.1	84.0	75	507	28.1	269	î.	į.	Ŀ	1	Η
78-1	Disc1	#12	46.37	3.76	13.09	11.51	0.17	8.03	11.38	2.93	1.91	0.40	39.65	58.0	84.4	59	545	33.2	339	73.1	390	Ţ	1	Η
78-1	Disc2	#5	45.16	3.42	12.62	11.09	0.17	9.15	12.42	2.46	1.42	0.42	98.68	62.0	843	130	127	75.5	568	132.9	796	96.1	Ĩ.	Η
78-1	Disc2	9#	47.22	3.45	13.60	11.80	0.23	7.10	10.17	3.13	2.10	0.42	99.38	54.4	83.6	115	926	54.1	543	122.4	999	93.8	0.70843	Н
78-1	Disc2	L#	47.20	2.92	13.65	9.94	0.14	7.83	11.36	2.73	2.00	0.40	98.49	60.9	83.9	1				ſ	1	E	ſ	Η
78-1	Disc2	#8	45.57	3.20	12.26	11.20	0.17	9.26	11.33	2.71	1.33	0.47	97.88	62.1	83.6	11	619	40.6	360	76.0	409	57.7	1	Η
78-1	Disc2	6#	46.81	3.46	14.03	11.81	0.16	7.34	10.54	2.93	1.91	0.43	99.59	55.2	82.8	76	655	38.7	400	84.8	482	72.3	0.70870	Η
78-1	Disc2	#10	46.32	3.46	13.16	11.01	0.15	7.65	11.60	4.80	1.27	0.46	100.01	57.9	83.6	33	631	37.3	346	73.4	450	65.1	0.70842	Η
78-1	Disc2	#11Le	46.32	3.24	12.74	11.04	0.14	8.18	12.72	3.18	1.33	0.39	79.62	59.5	83.6	55	598	36.7	281	56.1	401	45.5	1	Н
78-1	Disc2	#11RtS	I.	ī	t	ī	ł	î	į,	I	t	ī	I.	I	83.2	194	1694	100.5	871	193.5	1258	173.6	1	Η
78-1	Disc2	#11RtL	44.61	3.77	12.99	11.76	0.19	6.23	13.20	2.70	1.57	0.44	97.85	51.2	83.2	299	2480	158.7	1310	293.5	1887	246.6	1	Η
78-1	Disc2	#3	46.09	3.23	12.50	10.68	0.15	8.85	11.47	2.59	1.59	0.43	16.79	62.1	84.1	135	1038	62.5	579	126.8	719	100.6	1	Η
78-1	Disc2	14	46.45	2.99	12.17	10.32	0.15	9.22	12.92	2.37	1.37	0.40	98.80	63.9	83.9	104	106	55.1	465	110.9	686	78.9	1	Н
78-1	Disc3	#	46.07	3.81	13.22	10.47	0.12	6.00	12.52	2.93	1.80	0.42	97.81	53.1	852	69	555	26.4	297	64.8	370	55.4	0.70866	Ð
71-2	Disc1	8#3-4	48.50	2.66	12.55	9.41	6.1 4	7.38	14.63	2.33	0.84	0.35	99.13	60.8	87.4	26	384	23.0	185	30.4	124	22.4	1	9
71-2	Disc1		48.13	2.57	12.54	9.21	0.14	8.07	13.85	2.48	0.78	0.27	98.40	63.4	2.06	1	ĺ,	í,		ī	1	1	0.70602	G
71-2	Disc3	#3	46.12	3.48	13.06	10.94	0.15	5.81	14.09	2.55	1.29	0.32	98.23	513	84.3	50	517	29.5	237	52.7	347	40.9	1	5
71-2	Disc3	#2	46.71	3.10	12.72	9.66	0.15	7.90	13.10	2.45	1.24	0.32	97.72	61.8	91.0	45	504	27.1	243	58.8	314	35.3	0.70534	G
71-2	Disc3	9#	46.46	3.23	12.31	10.52	0.17	7.54	13.64	2.64	1.14	0.36	98.50	58.7	85.7	50	579	33.3	281	60.1	342	49.9	0.70599	Ð
71-2	Disc3	c	48.87	3.33	13.16	8.77	0.15	7.63	12.81	2.68	0.83	0.39	98.94	63.3	88.8	27	437	22.2	232	42.6	177	33.9	0.70482	5
11-17	Disc3	1#	49.01	3.13	12.53	11.05	0.18	6.15	12.01	2.85	0.94	0.40	98.51	52.4	81.4	34	419	27.1	232	41.4	205	35.6	0.70680	Ð
11-12	Disc1	R	49.28	2.82	12.70	10.94	0.14	6.89	11.63	2.60	1.01	0.37	98.73	55.5	\$3.0	36	409	30.3	223	40.9	214	33.0	0.70672	G
68-3	Disc1	#6mil	48.77	2.88	14.30	8.75	0.13	4.68	14.74	2.64	1.35	0.37	98.98	51.4	87.9	54	533	27.0	260	62.5	378	46.3	0.70624	Ċ
OFU-04-06	Ofu	13	44.25	3.31	12.60	10.10	0.11	7.99	13.38	3.40	0.94	0.50	96.75	61.4	3.67	25	856	41.6	464	76.8	343	106.2	0.70457	Н
OFU-04-06	Ofi	14	46.33	3.07	14.03	8.64	0.07	8.03	11.89	3.52	1.08	0.53	97.24	54.9	80.4	61	613	33.3	352	55.6	311	77.8	1	Н
OFU-04-06	Ofu	15	46.47	2.81	13.67	10.58	0.14	7.20	10.64	3.96	1.09	0.33	96.96	60.5	77.3	09	700	30.0	374	54.8	303	82.1	0.70443	Н
OFU-04-06	ofu	16	46.74	3.13	14.12	7.00	0.09	8.26	13.11	3.76	1.19	0.45	97.97	49.5	81.9	68	760	38.6	393	72.5	363	91.1	ĩ	Н
OFU-04-06	Oft	18	50.44	1.16	14.17	7.56	0.07	7.33	11.22	4.69	1.38	0.39	98.67	51.4	79.3	63	926	33.8	460	84.1	399	113.9	ĩ	Η
OFU-04-06	Ofu	19	47.15	2.14	11.15	17.06	0.21	13.83	2.72	1.56	1.7.1	0.46	98.00	75.2	76.6	1				1	1	1	1	Η
OFU-04-06	Ofit	20	47.36	2.90	14.50	9.00	0.13	7.59	10.84	4.02	1.06	0.39	97.84	55.1	78.6	81	606	39.0	438	70.3	402	32.5	I	Η
OFU-04-06	Ofu	24rep	46.35	3.58	13.47	8.58	60.0	8.11	13.70	3.23	0.72	0.29	98.13	55.8	\$3.0	1	ļ	Ù		1	1	T	1	Η
OFU-04-06	Ofu	25	45.74	3.16	13.55	11.98	0.12	17.71	11.15	3.54	0.89	0.42	98.33	63.6	772	59	629	38.1	383	48.7	243	ī	0.70454	Η
OFU-04-06	Ofu	26	46.14	2.51	12.70	10.60	01.0	8.04	13.61	3.01	0.74	0.34	97.93	62.3	76.5	45	573	37.7	338	42.3	191	a	0.70443	Η

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OFU-04-06	ofi	27	44.62	3.36	13.38	10.12	0.10	7.99	12.36	3.35	16.0	0.49	96.82 (0.03	9.5	5 8(17 40	2	1 15	5.8 3	21 11	1.4 -		Н
OFU-04-06	Ofu	28	47.60	2.70	13.87	9.87	0.08	7.49	10.92	4.18	0.94	0.33	98.13	58.5	8.4	12 8	4 34	.6 3.	82 7	1.1 3.	22 9	4.3 0.7	0434	Н
OFU-04-06	Ofu	30	47.23	3.03	13.82	7.23	0.09	7.78	12.39	3.74	1.15	0.24	96.79	\$ 6.03	2.3 6	3 6	32	8	24 5	4.0 2	83 7	- 9.9		Н
OFU-04-06	ofu	31	42.22	2.01	13.18	14.43	0.12	7.24	12.41	4.17	1.14	0.43	97.95	58.4 \$	80.3 8	1 100	16 46	.6 5	15 7	9.4 4	45 13	6.2 -		Η
OFU-04-06	ofa	33	41.65	2.83	13.21	16.38	0.17	6.96	10.68	3.77	1.06	0.34	97.17	1.17	7.7 6	96 90	11 39	4	79 8	1.9 3	91 10	5.9 -		Н
OFU-04-06	Ofu	0	46.18	3.00	13.70	9.93	0.13	7.53	11.26	3.70	66.0	0.32	96.80	6.85	6.5	0 8(10 35	5 4	99 8	1.5 3.	26 8	0.9 0.7	0469	Н
OFU-04-06	ofu	en	45.75	3.12	13.20	16.7	0.10	8.20	13.67	3.01	0.83	0.37	96.24	54.3 8	3.8 4	1 1	4 36	8	7 86	2.2 3	01 7	2.4 0.7	0443	н
OFU-04-06	ofu	4	45.71	3.05	13.40	10.95	0.06	7.62	10.88	3.93	0.93	0.33	96.89	51.8	7.2	1 9	2 33	.6 4	20 7	3.2 3.	26 7	0.2 -		н
OFU-04-06	Ofu	S	47.03	3.01	14.37	8.75	0.12	7.66	11.77	3.70	121	0.19	97.92	54.7 8	1.4	9 6	11 31	.6 3.	33 6	3.5 3	93 7	- 6.8		н
OFU-04-06	ofa	00	45.83	3.21	13.50	10.96	0.13	7.74	11.19	4.06	86.0	0.36	98.04 (51.6	7.7 6	3 8(61 40	4 0.	52 8	1.2 3	63 10	2.5 -		н
OFU-04-06	Ofu	6	44.84	3.37	13.52	12.21	0.15	6.89	10.62	4.25	1.04	0.32	97.31	1.12	7.4	7 100	16 42	7 5	28 9	8.4 4	44 10	4.7 -		H
OFU-04-06	ofu	22	38.04	4.65	10.75	12.83	0.18	9.84	17.18	1.15	0.12	0.59	95.55	70.3 8	6.08	8 116	11 72	4.	21 12	1.9 2	55 3	- 6.8		Η
OFU-04-06	Ofu	21	45.30	4.63	14.07	10.83	0.20	6.89	9.58	4.35	1.57	0.74	98.40	50.4	8.1 6	1	9 34	8.	17 5	1.7 2	71 5	1.4 -		Н
OFU-04-06	ofu	23	45.58	5.87	12.35	8.58	0.12	8.17	13.00	3.71	1.32	0.84	99.47	\$ 673	8 9.08	102	12 48	.6 5.	12 15	8.9 3	74 3	8.4 -		н
OFU-04-06	ofu	7(fir)	1	ł	1	į	1	ï	1	ī	ì	1	ĺ	4	=	0 8	14 21	5	39 IL	2.8 5	88 9	- 9.7		Η
OFU-04-06	ofu	7(sec)	Ĵ	ŧ	Ĩ	1	Ţ	ł	1	ä	1	12	3		ļ	5 80	9 37	5 4	28 7	4.1 3	65 10	4.7 -		н
Rb/Sr ratios	by LA-i	ICP-MS a	are cons	idered 1	to be mo	tre accu	rate than	those c	btained	by ion	probe.	Any dis	crepancy	/ betwe	en the l	Xb/Sr r	atios m	easured	by the	two m	ethods	is proba	bly due 1	o the
difficulty of 1	neasuri	ng Rb by	ion prot	be techi	niques. S	amples	78-1, 71	-2, 71-1	1 and 6	8-3 are	from the	e Samoa	IN AVON	3 dredg	ing cru	ise, and	I their s	ample 1	names a	are prec	eded by	/ the pre	fix "AVC	N3."
iviajor elemen	II Cala	are III WI.	No, and	l'ace el	cinents	ld un ane	DIII. FOIS	sterile co	ments 1	I ITE I	ATIO 1SO	mes (n	ro) are	also gu	15 12	ASSY (L) OF HIC	uagoun	II) Dazi	i.				
^a See data	in Tabl	a 1 for Si	r-isotope	data.																				

melt inclusion from sample 78-1 displays high Sr and Ba concentrations (2500 and 1900 ppm, respectively) that are 6 times more enriched than any whole-rock lavas examined from Malumalu. This ultra-enriched composition is significantly more enriched than observed in any Samoan basalts, and is derived from either an extremely enriched source or from very small degrees of melting.

Melt inclusions from OFU-04-06 also exhibit some unusual major and trace element compositions. One Ofu melt inclusion exhibits unusually low K_2O concentrations. Additionally, the OFU-04-06 melt inclusions exhibit a large range in SiO₂, including one sample with unusually low SiO₂ (38.0%) and high CaO (17.2%). These uncommon major element compositions may be due to the combination of unusual phases present in the OFU-04-06 melt inclusions prior to homogenization in the lab. Low glass totals in the OFU-04-06 inclusions are likely due to high volatile contents in the homogenized glasses, a hypothesis consistent with the volatile-rich phases in melt inclusions from this sample.

In addition to Rb/Sr ratios, several other major and trace element parameters correlate with 87 Sr/86 Sr in the melt inclusions. Melt inclusions from Ofu, Vailulu'u and Malumalu exhibit negative Ba anomalies that also are observed in Samoan shield-stage lavas [4] (Fig. 4). An approximation for this Ba anomaly is (Ba/Nb)_N (normalized to PUM, primitive upper mantle [30]), a ratio which correlates with 87 Sr/86 Sr in whole rocks and melt inclusions (Fig. 5). The (Ba/Nb)_N values are lowest (largest Ba-anomaly) in the basalts and melt inclusions with low 87 Sr/86 Sr, and highest in the basalts and melt inclusions with elevated ⁸⁷Sr/86Sr ratios. It is notable that basalts and melt inclusions associated with elevated ³He/⁴He ratios, and not the high ⁸⁷Sr/⁸⁶Sr EM2 endmember basalts, have the largest Ba anomalies. The mechanism that generates this anomaly is unknown [4], and the anomaly also exists in MORB and HIMU basalts, but not in basalts with EM1 characteristics.

The negative K_2O anomaly in the Samoan melt inclusions (Fig. 4) is commonly observed in OIBs. Curiously, however, PUM-normalized K exhibits little variability regardless of the degree of enrichment of the other trace elements (with the exception of a single Ofu melt inclusion). Despite the limited variability, olivine fractionation-corrected K_2O concentrations exhibit a relationship with ⁸⁷Sr/⁸⁶Sr in the melt inclusions (Fig. 5).

Nb/Zr also correlates with ⁸⁷Sr/⁸⁶Sr in Samoan basalts and melt inclusions (Fig. 5). The lowest Nb/Zr ratios are associated with unradiogenic ⁸⁷Sr/⁸⁶Sr values. Nb/Zr ratios correlate with Pb-isotopes [4] and inversely with ³He/



Fig. 4. PUM-normalized trace element patterns of Samoan basalts and melt inclusions; melt inclusion data is from ion probe analysis, except for K and Ti, which were measured by electron probe. Vaihulu'u melt inclusions are from three samples: 71-2 (grey lines), 68-3 (dashed line), 71-11 (solid black lines). Malumalu and Ofa melt inclusions are from samples 78-1 and OFU-04-06, respectively. All melt inclusions (except for one melt inclusion from 78-1 and two from OFU-04-06 that have no major element compositions, see Table 2) are corrected for olivine fractionation to be in equilibrium with mantle olivine of Fo_{90} (see text for correction scheme). The grey field encloses the range of whole-rock patterns from the least evolved basalts (>6.5 mt;% MgO) from Vaihulu'u, Ofu and Malumalu; whole-rock compositions have been corrected for olivine addition/fractionation. PUM values from [30].

⁴He (not shown) in Samoan basalts. Nb/Zr may serve as a proxy for ³He/⁴He isotopes in Samoan melt inclusions; therefore, it may be important that the Nb/Zr ratios in several of the Ofu-04-06 melt inclusions exhibit values lower than the ratios observed in whole rocks.

3.3. Melting models of the EM2 source

It is important to place constraints on the trace element variability introduced by melting processes. A plot of Sr versus Ti/Zr shows that the majority of the melt inclusions form an array that extends outside of the whole-rock field to enriched Sr and low Ti/Zr values (Fig. 6). The low Ti/Zr ratios (15-88) observed in the melt inclusions extend to lower values than observed in whole-rock basalts from the eastern Samoa islands and seamounts (63-130). Such low values cannot be produced by crystal fractionation of melts before olivine entrapment, and assimilation of MORB (Ti/Zr=88) also fails to produce low Ti/Zr values [31]. A model of the Samoan EM2 source composition [4] provides a robust estimate for the mantle source sampled by the extreme EM2 basalt sample 78-1, and variable degrees of aggregated modal fractional melts of this source (Sr=20.0 ppm, Ti/Zr=101.9) in the garnet and spinel stability fields can describe much of the melt inclusion array in Fig. 6. Consistent with the model of the EM2 source as a metasomatized harzburgite [4], we adopt a harzburgite source lithology (1% spinel, 3.6% clinopyroxene [cpx], 20.6% orthopyroxene [opx] and 74.8% olivine), with mineral modes from [32] and mineral/melt partition coefficients from [33]. We assume the mineral modes of a similar bulk composition for melting in the garnet stability field (3.8% garnet, 2.7% cpx, 17.7% opx and 75.8% olivine) using the spinel to garnet conversion from [34]. The two melting models follow similar trajectories, but melting in the spinel stability field is required to generate the exceptionally high Sr and low Ti/Zr observed in the ultra-enriched Malumalu melt inclusions. Interestingly, if the EM2 source [4] has a more cpx-rich lithology than the harzburgite in our melting model, it will not produce melts with the Sr concentrations observed in the most enriched melt inclusion at reasonable degrees (>1%) of melting (Fig. 7).

Other geochemical indicators, including Y/Zr ratios, more clearly resolve the relative roles of melting in the garnet and spinel stability fields. Due to the relative compatibility of Y in residual garnet, low Y/Zr ratios are consistent with melting in the presence of garnet. A role for melting in the garnet stability field is suggested in a plot of Y/Zr against Nb/Zr in Fig. 6. This is particularly true for the Ofu melt inclusions, which trend to the lowest Y/Zr ratios



Fig. 5. Nb/Zr, K₂O and Ba/Nb_N versus ⁸⁷Sr/⁸⁶Sr. K₂O measured by electron probe, and Nb, Zr and Ba by ion probe. (Ba/Nb)_N (normalized to PUM) is used as a proxy for the pervasive negative Ba anomaly in Samoan shield-stage basalts. The EM2 basalts and melt inclusions are associated with elevated Nb/Zr, K₂O and (Ba/Nb)_N (smaller negative Ba anomalies), and the low ⁸⁷Sr/⁸⁶Sr basalts exhibit lower Nb/Zr, K₂O and (Ba/Nb)_N (larger negative Ba anomalies). All melt inclusions are corrected for olivine fractionation. The grey fields enclose the range of whole-rock patterns from the least evolved basalts (>6.5 wt.% MgO) from Vailulu'u, Ofu and Malumalu (whole-rock compositions have been corrected for olivine addition/fractionation). Symbols are the same as Fig. 3. ⁸⁷Sr⁸⁶Sr internal precision (2 σ , standard error) is approximately the size of the symbol.

observed in the melt inclusions suite and straddle the gamet melting trend at 5% melt. This is similar to the Ti/Zr versus Sr plot, where the gamet melting curve trends through the Ofu melt inclusion field at $\sim 4\%$ melt. Malumalu and Vailulu'u melt inclusions (and Samoan whole rocks) are offset to higher Y/Zr ratios, perhaps suggesting a larger role for melting in the spinel stability field, an observation that is consistent with the same subset of melt inclusions in the Ti/ Zr versus Sr melt model.

No single melting model of the EM2 source perfectly describes the melt inclusion fields for all three volcanoes, but we find that a combination of melting and mixing satisfactorily reproduces the melt inclusion geochemical variability. The relative roles of melting and mixing of different components can be partially deconvolved in a plot of ⁸⁷Sr/⁸⁶Sr versus 1/Sr (Fig. 8), where two-component mixing trajectories are linear and variable degrees of melting result in horizontal trajectories. The Ofu melt inclusions lie on a horizontal trend, which can be described by various degrees of melting of a single source that exhibits a ⁸⁷Sr/⁸⁶Sr ratio of ~0.7045, and the Malumalu and Vailulu'u melt inclusions form a diagonal array that suggests a role for two-component mixing.

4. Discussion

4.1. A homogeneous source for PHEM basalts

Compelling evidence that the Samoan melt inclusions sample a heterogeneous source comes from Sr-isotope analysis of the melt inclusions from Vailulu'u and Malumalu basalts. However, the uniformly unradiogenic character of the Ofu melt inclusions precludes a significant contribution from an enriched, radiogenic (EM2) component. This suggests that Samoan melts with high 3He/4He sample a homogeneous source and do not mix with melts of an enriched component. By comparison, the melt inclusions from Malumalu and Vailulu'u span a large range of Sr concentrations and isotopic compositions, indicating that both variable degrees of melting and mixing have occurred. An aggregated fractional melt trajectory for the gamet stability field is plotted (Fig. 8) for the EM2 source [4], and the Malumalu and Vailulu'u melt inclusions form a broad array that trends diagonally away from the horizontal Ofu melting trajectory toward low degree ($\sim 1\%$) melts of the EM2 source; the Vailulu'u and Malumalu melt inclusions can be produced by aggregated melts of an Ofu source that then mix with aggregated fractional garnet melts of the EM2 source.

Unlike the EM2 source, the trace element source composition of the high ³He/⁴He, lower ⁸⁷Sr/⁸⁶Sr Ofu source component is less clear. Called PHEM [23], this



Fig. 6. Upper panel: Ti/Zr versus Sr concentration in melt inclusions and whole rocks with melting models. All melt inclusions and whole rocks are corrected for olivine fractionation to be in equilibrium with mantle olivine. Grey field is for whole-rock samples from Vailulu'u, Malumalu and Ofu. Lines with closed boxes represent aggregated fractional melting trends of the EM2 source [4] in the garnet and spinel stability fields, using partition coefficients from [33]. Harzburgite mineral modes are from [32]. All melting is non-modal. Tick marks are every 1% melting, beginning at 0% and increasing to the left. Lower panel: Y/Zr versus Nb/Zr in melt inclusions and whole rock, including melting models in the garnet and spinel stability fields. Tick marks are every 1% and the degree of melting increases to the left. Melting parameters and grey field are the same as upper panel.

component melts to form basalts and melt inclusions from Ofu. However, it is possible to bracket the source composition of this component and estimate the degree of melting captured in the Ofu melt inclusions. Although the Ofu basalts are more isotopically depleted than the EM2 basalts, an isotopically and trace element depleted DMM lherzolite source (7 ppm Sr) [35] fails to produce the high Sr concentrations observed in the Ofu melt inclusions (Fig. 7). This would suggest that the Ofu source is either more refractory or more trace element enriched, or both, than the lherzolitic [35] DMM source. The first option can be explored by invoking a more refractory, harzburgitic DMM source. However, only unreasonably low (F < 1%) degrees of melting can produce the most enriched Sr concentration observed an Of u melt inclusion. The second scenario can be tested by invoking the trace element enriched EM2 source, and assigning it a lherzolitic lithology that is similar to DMM. At reasonable degrees of melting (F=1.5%), such a source can generate melts with sufficiently high Sr contents to match the range observed in Ofu. Finally, a harzburgitic EM2 source, which is both more refractory and trace element enriched than DMM, can generate the most enriched melt inclusions from Malumalu and Ofu between 1% and 2% melting. The harzburgitic EM2 source can serve as a probable upper limit for the trace element enrichment of the PHEM source because we consider it unlikely for the less isotopically enriched PHEM component to exhibit greater trace element enrichment than EM2. However, M.G. Jackson, S.R. Hart / Earth and Planetary Science Letters 245 (2006) 260-277



Fig. 7. Sr in melt versus degree of melting (F) for various combinations of two lithologies and two trace element source abundances. Two dashed lines represent olivine fractionation-corrected Sr concentrations in the most enriched melt inclusions from Malumalu sample 78-1 (upper line) and Vailulu'u sample 71-2 (lower line). DMM trace element source and lherzolite lithology from [35], converted to equivalent garnet facies lithology using relationship from [34]. EM2 source from [4]. Harzburgite lithology from [32]. Note that only the EM2 source with a harzburgite lithology can produce the highest Sr concentration observed in the melt inclusion from sample 78-1 at $F \ge 1\%$. All melting is non-modal in the garnet stability field. Neither modal melting nor melting in the spinel stability field significantly changes our conclusions at $F \ge 1\%$.

a PHEM source with a lithology (and/or Sr concentrations) intermediate between DMM and EM2 cannot be ruled out. If PHEM has the same Sr content and lithology as the EM2 source, the Vailulu'u and Malumalu melt inclusion arrays can be explained as mixtures of $6.5\pm1.5\%$ melts of PHEM and ~1% melts of EM2. However, the degree of melting of the PHEM source that contributes to the Malumalu–Vailulu'u mixing array should be taken as a maximum.

Unlike the Ofu melt inclusions, which sample a pure PHEM source, no melt inclusions sample a pure EM2 melt (as calclulated by [4], using ultra-enriched Srisotope compositions from Samoan xenoliths [36]). Mixing lines in Fig. 8 between a 1% EM2 melt and 6.5±1.5% PHEM melts indicate that the subset of Vailulu'u and Malumalu melt inclusions with both 87 Sr/ 86Sr and Sr concentration data are dominated by a PHEM component, and exhibit less than a ~30% contribution from the EM2 component. However, one Malumalu melt inclusion exhibits a much larger contribution from an EM2 component, as indicated by its high (~1865 ppm) Sr content and low (~17) Ti/Zr ratio (Fig. 6). The PHEM-EM2 melt mixing lines in Fig. 8 suggest that this melt inclusion contains more than a 70% contribution from the EM2 melt component. The mixing model suggests that a 70% contribution from an EM2 melt would produce an extrapolated 87 Sr/86 Sr ratio of ~ 0.712 , which is a significantly higher ratio than has been observed in a Samoan basalt, but close to the 87 Sr/

⁸⁶Sr of cpx in metasomatized xenoliths from Savai'i in western Samoa. Unfortunately, this ultra-enriched melt inclusion was too small for isotopic analysis.

4.2. Isotopic variability in Samoan melt inclusions: MORB or FOZO?

Correlations between trace elements and Sr isotopes suggest that trace element variability in Samoan basalts and melt inclusions may reflect heterogeneity in the Samoan mantle. Rb/Sr, K2O, (Ba/Nb)N and Nb/Zr ratios correlate with Sr isotopes in Samoan lavas and melt inclusions, suggesting that these trace elements are heterogeneous in the Samoan mantle source. However, a combination of melting processes and variable source lithology may drive the observed correlations, limiting the role for source heterogeneity. For example, the relatively constant K₂O concentrations and negative (PUM-normalized) anomalies in Samoan lavas may be a result of residual phlogopite, which may cause K2O to behave more compatibly in the Samoan source during melting. However, K2O correlates with 87 Sr/86 Sr in the whole rocks and melt inclusions (Fig. 5), suggesting a role for K2O heterogeneity in the Samoan source, and that K₂O concentrations in Samoan basalts and melt inclusions may be controlled only partially by melting processes.

Assuming that Rb/Sr variability in Samoan melts reflects source variability, the array formed by the ⁸⁷Sr/



Fig. 8. ⁸⁷Sr/⁸⁶Sr versus 1/[Sr] for Samoan melt inclusions and basalts. Malumalu and Vailulu'u melt inclusions form a quasi-linear array extending away from the whole-rock basalts from Malumalu, Ofu and Vailulu'u (grey field, same as Fig. 6). Ofu melt inclusions fall off this array, plotting at lower ⁸⁷Sr/⁸⁶Sr at a given Sr concentration. Aggregated non-modal fractional melting trends of the EM2 and PHEM sources in the gamet stability field are shown as horizontal lines (tick marks are every 1% melting, beginning at 0% and increasing to the right). Two mixing lines, both extending diagonally from a 1% EM2 melt to 5% and 8% PHEM melts, are shown (each tick represents 10% mixture, starting at 0% EM2 and increasing to the upper left). Ofu melt inclusions lie on the melting trend for PHEM, and require no mixing with an EM2 component. Malumalu and Vailulu'u melt inclusions can be described by mixing of melts from the EM2 and PHEM sources. Sr concentrations in melt inclusions and whole rocks are corrected for olivine fractionation. PHEM and EM2 source compositions plot outside of the figure as indicated, at 1/Sr values of ~0.05 (Sr concentrations of 20.0 ppm). The published source for EM2 [4] is used in this figure. As a limit, PHEM is given the trace element source abundances and lithology of EM2 in the melt model. Symbols same as in Fig. 3.

⁸⁶Sr and Rb/Sr data can be modeled as binary mixing between a PHEM and an EM2 component, a model originally proposed by Farley et al. [23]. Such a model is consistent with the trace element melting/mixing model proposed above, which suggests that the Ofu melt inclusions sample only a high ³He/⁴He PHEM mantle reservoir and the Vailulu'u and Malumalu melt inclusions result from mixing melts from both the PHEM and EM2 mantle reservoirs.

Previous work on the Pb-isotopic variability in melt inclusions [17] suggested that the unradiogenic endmember in EM2 basalts from Tahaa may be depleted (MORB or FOZO?), and thus lie at even lower ⁸⁷Sr/⁸⁶Sr values than observed in the high ³He/⁴He PHEM lavas from Ofu (Fig. 3). In fact, an extrapolation of the Samoa melt inclusion and whole-rock basalt array in Fig. 3 does indeed trend toward one of two depleted components that are significantly less radiogenic than PHEM: MORB [22] or a common high ³He/⁴He mantle component, called FOZO [24] (Fig. 3). The FOZO component is represented by basalts with the highest ³He/⁴He from Hawaii [37], Iceland [38] and Baffin Island [39]. Both MORB and FOZO lie on a similar extension of the Samoa melt inclusion 87 Sr 86 Sr-Rb/Sr array, so it is difficult to distinguish which, if either, of these two components is sampled by the Samoan melt inclusions. If the depleted component is MORB, it may be entrained in melt inclusions by shallow anatexis due to preferential cooling and olivine crystallization near magma chamber and conduit walls [7,8,21]. However, the presence of FOZO (or any other high ³He/⁴He component) in the Samoan melt inclusions would require that the isotopic variability in melt inclusions reflect true source heterogeneity, assuming that a high ³He/⁴He component does not exist in the oceanic crust or lithosphere.

4.3. The case against MORB

It may be possible to look at other lines of geochemical evidence to discern whether MORB or FOZO play a role in augmenting the isotopic diversity in Samoan melt inclusions. Models suggesting that melt inclusion isotopic variability is caused only by contamination from unradiogenic oceanic crust and lithosphere at shallow levels do not explain how several Samoan melt inclusions have higher ⁸⁷Sr/⁸⁶Sr ratios than their host

bulk rock compositions (see Fig. 1). For example, Samoan whole-rock basalt sample 71-11 hosts several melt inclusions that have 87Sr/86Sr ratios (up to 0.70692) that are significantly more enriched than its host rock (0.70550). The presence of ⁸⁷Sr/⁸⁶Sr ratios in melt inclusions that are higher than the bulk rock require that at least some of the isotopic variability present in melt inclusions is derived from the mantle source, because the enriched component in melt inclusion 71-11a is too enriched to be found in the oceanic crust and lithosphere. Therefore, if assimilation of oceanic crust and lithosphere contributes heterogeneity to the Samoan melt inclusions, it cannot be the only means by which isotopic heterogeneity is produced in Samoan melt inclusions, and some contribution from the melt source must be involved as well.

On a different tack, the case for the less radiogenic melt inclusions sampling the depleted oceanic crust and lithosphere by assimilation is limited severely by the observation that, within analytical uncertainty, not a single melt inclusion has an 87 Sr/86 Sr ratio that is lower than the least radiogenic (87 Sr/ 86 Sr=0.7044) wholerock basalt measured in the Samoan islands (Fig. 1). The Samoan melt inclusions trace out a range of Sr-isotope variability that is confined to the region of Sr-isotope space defined by the Samoan whole-rock data (Fig. 3). On an island-by-island basis, the interpretation is more complicated, as melt inclusions from two Vailulu'u whole-rocks sample a component more depleted than found in whole rocks measured from the seamount. However, the least radiogenic component found in Vailulu'u melt inclusions is also found in lavas from nearby Samoan islands (e.g., Ta'u and Ofu), exhibits elevated ³He/⁴He ratios, and is thus known to exist in the Samoan plume. Many of the downstream Samoan seamounts also are dominated by 87 Sr/86 Sr between 0.7044 and 0.7049 [3]. It seems unnecessary, therefore, to invoke contamination from the oceanic crust and lithosphere to explain the presence of the less radiogenic component when it already exists inside the plume! Although the argument can be made that an insufficient number of melt inclusions have been analyzed to detect a component more depleted than what is found in whole rocks, the number of melt inclusions analyzed for Sr isotopes (n=41) is already significant, and is equal to ~30% of the number of published 87 Sr/86 Sr whole-rock analyses from the Samoan hotspot.

The high ³He/⁴He, unradiogenic Sr component in Samoan basalts (PHEM) is unique in that it exhibits ⁸⁷Sr/⁸⁶Sr ratios more enriched than in the high ³He/⁴He Hawaii, Iceland or Baffin Island basalts,

suggesting that the high ³He/⁴He reservoir in the mantle is at least mildly heterogeneous in ⁸⁷Sr/⁸⁶Sr ratios. The same line of reasoning that precludes the presence of entrained melts from oceanic lithosphere in the Samoan melt inclusions also minimizes the possibility that a traditional, depleted FOZO-like component (87Sr/86Sr=0.7030 [24]) serves as the unradiogenic Sr component: If melt inclusion diversity were a result of entrainment of a component (MORB or FOZO) more depleted than found in Samoan basalts, then the melt inclusions would extend to 87 Sr/ ⁸⁶Sr ratios lower than found in whole rocks (0.7044). However, the high ³He/⁴He Samoan sample OFU-04-06 defines the lowest 87 Sr/86 Sr portion of the Samoan whole-rock mixing array (see Fig. 3) and the melt inclusions are identical to the whole rock, suggesting that the least radiogenic Sr composition sampled by the Ofu basalts is the same component found in the melt inclusions. Therefore, we maintain that the unradiogenic Sr component in Samoan melt inclusions is more enriched than MORB or FOZO, and is likely the same PHEM component sampled by the high 3He/ ⁴He Samoan basalts, suggesting that a two-component EM2-PHEM mixing model may be the most appropriate for melt inclusions originating in the enriched Samoan mantle. Scatter around such a mixing model (Fig. 8) may be due to minor contributions from other components [4] that may exist in the Samoan mantle.

4.4. Implications for source heterogeneity (or lack thereof)

The results for Sr-isotope measurements in Samoan melt inclusions support an argument for an origin of the isotopic variability in the melt source, not contamination by oceanic crust and lithosphere. We assume that the isotopic variability (or lack thereof) in Samoan melt inclusions is not a product of variable degrees of homogenization in magma conduits and chambers before olivine entrapment, but rather that the isotopic variability in melt inclusions reflects the heterogeneity of the melt source: When the melt source is heterogeneous, melt inclusions capture the range of heterogeneity while the isotopic composition of the bulk rock lava represents an average of the heterogeneity sampled in the melt. By extension, we infer that the high ³He/⁴He whole-rock sample OFU-04-06 tends to sample a more homogeneous source, as the melt inclusions are nearly isotopically homogeneous and identical to the bulk rock. Perhaps, then, only melting of a pure PHEM source allows the high ³He/⁴He composition to persist in the

Ofu whole-rock lavas. Vailulu'u and Malumalu melt inclusions are more isotopically heterogeneous, and thus are inferred to sample a heterogeneous source that captures much of the mixing spectrum between the EM2 and PHEM components. The contribution of an EM2 component may explain the diminished ³He/⁴He composition in the lavas from these two volcanoes. Although 87 Sr/86 Sr analyses of the melt inclusions from high ³He/⁴He basalts from other localities are not yet available, perhaps the high 3He/4He mantle that these basalts sample is homogeneous and devoid of enriched domains. This hypothesis is consistent with the melting-mixing model above (see Fig. 8), which suggests that the isotopically homogeneous melt inclusions from the high 3He/4He basalt from Ofu exhibit no evidence of mixing with an EM2 component.

It is notable that while near-pure PHEM melts are observed in Samoan melt inclusions, pure EM2 melts (87Sr/86Sr=0.7128 [4,36]) were not unequivocally detected in this study (i.e., by measurement of Sr isotopes). Several melt inclusions with ultra-enriched trace element patterns were observed, but most of the melt inclusions are composed of <30% EM2 component. Perhaps this indicates that, compared to the PHEM component, the EM2 component in the Samoan plume is rare. Alternatively, the EM2 component may not be rare in the Samoan plume, but is more refractory and produces less melt than the PHEM component. In this way, perhaps, EM2 melts are less frequently sampled by melt inclusions. Future work on melt inclusions will help resolve the relative contributions of the enriched and high 3He/4He sources to OIB lavas.

5. Summary

The following conclusions can be drawn from this study:

- The Sr-isotopic diversity in melt inclusions from Samoan basalts does not extend significantly above or below the range defined by whole rocks from the Samoan hotspot.
- 2.) A few melt inclusions exhibit ⁸⁷Sr/⁸⁶Sr ratios significantly higher than their host whole rock. This is taken as evidence that assimilation of MORB lithosphere cannot be the only mechanism that contributes isotopic diversity to the melt inclusions.
- 3.) The ⁸⁷Sr/⁸⁶Sr ratio of the high ³He/⁴He basalt is essentially indistinguishable from the Sr-isotope ratios (0.7044) measured in its melt inclusions, and the ⁸⁷Sr/⁸⁶Sr ratios in melt inclusions from

other Samoan basalts do not exhibit ratios lower than 0.7044. This observation is consistent with the hypothesis that a high ³He/⁴He component, not MORB, is the unradiogenic Sr endmember in Samoan melt inclusions.

4.) Melt inclusions from a high ³He/⁴He Samoan basalt (with less radiogenic ⁸⁷Sr/⁸⁶Sr) are isotopically more homogeneous than the melt inclusions from basalts with higher ⁸⁷Sr/⁸⁶Sr (more contribution from an EM2 component). This may indicate that the Samoan high ³He/⁴He basalts sample a source that is more isotopically homogeneous than the source that produces basalts with an EM2 component.

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Appendix A. Supplementary data

Supplementary data associated with this article can be found, in the online version, at doi:10.1016/j.epsl. 2006.02.040.

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Appendix A. Supplementary data (as it appears, published, on the Elsevier website) Background Datasets Text,Figs&Tables Click here to download Background Datasets: SuppDataFigsTables.pdf
Supplementary Data (Jackson and Hart, 2006)

Sr isotopes by laser ablation PIMMS: Application to Samoan basaltic melt inclusions

1. Introduction

We have thoroughly, but not exhaustively, investigated the parameters necessary for precise and accurate Sr isotope analysis of basalt glasses by *in situ* laser ablation PIMMS. The lower limit of external precision attained under optimal conditions for the SRM987 Sr solution standard is \pm 5 ppm (>25 volt data, 1 σ), and is fairly comparable to good TIMS data. For lower intensity data (~ 1-5 volts), more comparable to our typical laser ablation work, the external solution precision is in the 20-80 ppm range. We have instituted a new method for Kr correction that is based on ⁸⁴Kr, and a method for Rb correction that utilizes basalt glass standards with significant Rb/Sr ratios and known ⁸⁷Sr/⁸⁶Sr ratios. Even in basalts requiring very large Rb corrections, we are able to obtain ⁸⁷Sr/⁸⁶Sr data with ~320 ppm external precision and 45 ppm internal (in-run) precision (2 σ standard deviation) on a suite of Samoan glasses with known ⁸⁷Sr/⁸⁶Sr ratios.

We developed an analytical protocol for the measurement of Sr isotope ratios by Laser Ablation PIMMS for analysis of melt inclusions in olivine phenocrysts in OIBs (ocean island basalts) from the EM2 (Samoa), EM1 (Pitcairn, Christmas) and HIMU (Mangaia) mantle end-members. We acknowledge a significant existing body of work on Sr isotope measurement by laser ablation PIMMS, but will not attempt to review this literature or compare it with the protocols we have installed on the NEPTUNE at the Woods Hole Oceanographic Institution [1-6].

2. Techniques

2.1. Instrumental Description

The data discussed here were obtained with a ThermoFinnigan NEPTUNE multicollector ICP-MS, coupled to a NEW Wave UP213 laser, housed in the Plasma Facility at the Woods Hole Oceanographic Institution. This NEPTUNE was installed in June 2003, replacing an earlier one that had been installed in February 2002, but damaged by the October 2002 fire in the Ion Probe Facility.

The typical operating parameters of the NEPTUNE and laser are given in Table 1 (note that our techniques have evolved over time, and some earlier data may have utilized somewhat different protocols). Currently, we aspirate clean dilute (5%) nitric acid during the lasering, to allow solution standards to be interspersed with laser runs when needed. The arrangement of Faraday cups, and the relevant isotopic masses that are collected, is given in Table 2. Amplifiers are not rotated, as this rotates the Faraday (off-mass) baselines as well. The mean raw Faraday intensities are transferred after all analyses to an offline data reduction program (TweaKr), for the various corrections.

We have opted to run the laser at 100% power and in apertured mode. While we could likely emplace higher energy density in focused mode, this tends to throw off larger chips and particles, particularly from basalt glass samples. We have not made any

investigations of the effects of particle size on mass bias, interferences, and the like. Each analysis takes about 6.5 minutes of lasering, including offpeak baseline integration for 64 seconds and 20 cycles of 16 second integrations (and, more recently, 40 cycles of 8 second integrations). With a 120 μ m spot and 200 μ m long raster lines (see Table 1), spaced at only 20 μ m, the final ablation pit is pyramidal, approximately 300 x 300 μ m, and 250 μ m deep. Typically, there will be little or no drift of ⁸⁸Sr intensity with time during ablation of homogeneous basalt glass standards, and the 1 σ standard deviation of intensity will fall in the range 5-15%.

2.2. Standards and Canonical Isotope Abundances

For Rb and Sr isotope standards, we have used NIST standards SRM984 and SRM987. The certified values for these standards are listed in Table 3, along with the quoted uncertainties. Note that ${}^{85}\text{Rb}{}^{87}\text{Rb}$ in SRM 984 is only certified to \pm 423 ppm, ${}^{87}\text{Sr}{}^{86}\text{Sr}$ and ${}^{85}\text{Sr}{}^{86}\text{Sr}$ to only 380 ppm, and ${}^{84}\text{Sr}{}^{86}\text{Sr}$ to 0.25%. It is likely that the uncertainties in ${}^{87}\text{Sr}{}^{86}\text{Sr}$ and ${}^{88}\text{Sr}{}^{86}\text{Sr}$ are not independent, but possibly related by some fractionation-dependence. Note also that the certified ${}^{86}\text{Sr}{}^{88}\text{Sr}$ value is 0.1193515, not the consent value of 0.119400 adopted by the community. We have corrected all of the SRM 987 certificate ratios, by exponential law, to be constent with the consent value of 0.1194. Ironically, this results in a "corrected" ${}^{87}\text{Sr}{}^{86}\text{Sr}$ ratio of 0.7101938, which is lower than (but marginally within errors of) the value commonly adopted by the community of 0.710240 (which is itself different from the "uncorrected" certificate value of 0.710339). It will obviously be important for published papers to be very clear as to their usage of SRM 987 standard values. For Kr, we adopt the isotope abundances compiled by Ozima and Podosek (2001).

2.3. Backgrounds and Baselines

The NEPTUNE software allows a choice of baseline protocols. PIMMS users frequently use "on-peak" baselines, after sample wash-out [1, 3-6]; the NEPTUNE also allows a "defocused" beam baseline. Our experience is that baselines can be measurably (> 50 uV) elevated across a broad mass region during actual sample analysis and thus different from those measured in a defocused "beam-off" state, or while running gas or acid blanks. We have chosen therefore to adopt the common TIMS procedure of running "off-peak" baselines while sample analysis is underway; while this consumes part of the sample, we see no other way to ensure reliability of baselines. (However, baselines on Samoan basalt standards and basaltic melt inclusions do not change significantly over the course of an analytical session, and we have adopted the practice of applying baselines from larger melt inclusions-run during the same analytical session-to exceptionally small melt inclusions to save precious material). We have chosen to measure baselines at a position 0.70 amu below each mass; this avoids potential baseline interferences from doubly-charged half-mass REE peaks, allows the Faradays that are intentionally set at 83.5 and 85.5 in order to monitor these REE⁺² to offset 0.2 amu below mass 83 and 85 during baseline measurement, and places the ⁸⁸Sr cup near the minimum in the valley between ⁸⁷Sr and ⁸⁸Sr during baseline measurement. With this protocol, the only baseline that will have a significant tail contribution is ⁸⁷Sr. With the measured abundance sensitivity at Sr mass of ~ 1ppm at 1 amu, the 87 Sr/ 86 Sr ratio will be elevated by < 12

ppm; this error is small compared to our external precision goal, and will also be normalized, to first order, through the use of the SRM987 standard.

2.4. Kr Correction

As is well known, there are interferences at ⁸⁴Sr and ⁸⁶Sr masses from Kr. Conventionally, the Kr is considered to be a contaminant in the Ar plasma gas, and is corrected by measurement of ⁸²Kr or ⁸³Kr. Over the past two years of measurement here, our ⁸³Kr intensity has varied from ~ 0.05 mV to ~ 20 mV, and we believe that not all of it comes in with the argon, nor is all of mass 83 actually Kr. Typically, after changing samples in the laser cell, there is a component of air Kr which slowly decreases as the chamber is purged. Ultimately, the Kr intensity with He flowing from the cell into the machine will decrease substantially below that in aspirated solutions (because the He has less Kr than the argon). It appears to us that there is a component of atmospheric Kr dissolved in aspirated solutions and this contributes to the Kr background. This could be possibly moderated by hermetically isolating the solutions from the atmosphere. Air entrainment by the plasma at the torch is also probable, but has proven difficult to constrain.

We have typically measured both 82 Kr and 83 Kr during all Sr analyses. Deviations of up to a factor of two from the canonical 83 Kr/ 82 Kr ratio of ~ 1 are very common in laser runs, with values most often above one, but also frequently below one (Fig. 1). Even solution runs on the 987 standard will frequently show 10-20% deviations from a ratio of 1, typically with the ratios being too low (interferences on mass 82). Clearly there exist isobaric interferences that seriously hamper efforts to use either of these masses to correct for Kr. We have adopted an alternative scheme which basically uses the most abundant Kr isotope at mass 84 (~57%) to make the Kr correction on mass 86. This mass has a Sr "interference" on it, but for ⁸⁸Sr intensities in the 1-2 volt range, the Kr makes up 35-75% of the mass 84 peak (i.e. the Kr and Sr are approximately equal in intensity). By "subtracting" Kr until the ⁸⁴Sr/⁸⁸Sr ratio equals the canonical value of 0.00675476 (while iterating the mass-bias correction), this allows a robust correction to be made on mass 86 (there is a large error demagnification in this process, due to the fact that ⁸⁶Kr/⁸⁴Kr is ~ 0.30, while the 86 Sr/ 84 Sr ratio is ~ 17.7). Obviously, this technique relies on the absence of any other significant isobaric interferences at mass 84; in any event, these are likely to be less fatal than those at mass 83, because of the 84 Kr / 83 Kr ratio of \sim 5. One indication that this calculation method is helpful is the observation that the external precision of samples and standards run at low-intensity (1-2 volts on mass 88) is typically improved by 30-50%, compared to the same data corrected with ⁸³Kr. Further discussion of isobaric interference issues may be found in section IIF, below.

We are able to correct for Kr interferences so that elevated Kr/Sr ratios do not noticeably diminish the precision or accuracy of glass analysis while using our protocol, provided that the ⁸⁸Sr intensity is > 1 V and ⁸²Kr/⁸⁸Sr is < 0.004 (Table 4). At exceptionally high Kr/Sr ratios, there may be a tendency for measurement precision to degrade. Therefore, analyses of unknowns that exhibit high Kr/Sr ratios (82 Kr/⁸⁸Sr > 0.004) are discarded.

Uncertainty in the isotopic ratios of Kr used in our correction scheme does not significantly contribute to diminishing the accuracy of 87 Sr/ 86 Sr analyses by laser ablation. At elevated Kr/Sr (82 Kr/ 88 Sr = 0.004) —where such uncertainties will play the

largest role in affecting the final Kr-corrected ${}^{87}\text{Sr}/{}^{86}\text{Sr}$ ratio—a 1% uncertainty in any or all of the Kr-isotope ratios will change the final corrected ${}^{87}\text{Sr}/{}^{86}\text{Sr}$ ratio by <15 ppm. Additionally, allowing the Kr-isotope ratios to fractionate by up to 2% (before or during entry into the plasma) changes the ${}^{87}\text{Sr}/{}^{86}\text{Sr}$ by <15 ppm.

2.5. Rb Correction

The analysis of basaltic melt inclusions typically involves samples in which 20-50% of intensity on mass 87 is due to Rb, and large Rb corrections are required. Our goal was to devise a protocol where the Rb-corrected ⁸⁷Sr/⁸⁶Sr ratios are accurate to within a few hundred ppm. We adopted the strategy of using natural basalt glasses (Table 5) with precisely known ⁸⁷Sr/⁸⁶Sr ratios (obtained by TIMS) to bracket the basalt glass unknowns; for each "standard" glass, the ⁸⁵Rb/⁸⁷Rb ratio required to give the TIMS ⁸⁷Sr/⁸⁶Sr is calculated, and these bracketing ⁸⁵Rb/⁸⁷Rb ratios are then used for the Rb corrections in the unknown glasses. This method depends only on the Rb mass bias being relatively invariant on short time scales (20 minutes), and not being a function of the particular major element composition of the various glasses analyzed.

To assess these issues, we performed laser ablation analyses of a suite of 14 Samoan basalt glasses of known ⁸⁷Sr/⁸⁶Sr, with significant variability in major and trace element composition. The results of nine one-day analysis campaigns are shown in Fig. 2; interspersed with these laser ablation analyses during the first two days were a series of runs on mixed 984-987 Rb-Sr solutions. While the total variability of the "Required ⁸⁵Rb/⁸⁷Rb" (basically the variability of the Rb mass bias factor) is substantial (2317 ppm, excluding the 987-984 Rb-Sr solution runs), the variation with time is relatively smooth, such that the bracketing technique will be fairly effective (note that the error on ⁸⁷Sr due to the Rb effect is de-magnified by a factor of ~2.59 due to the ⁸⁵Rb/⁸⁷Rb ratio).

Several things may be noted. First, the "empirically" determined Rb ratio in lasered basalts has an average value of 2.58745, which is some 2100 ppm lower than the canonical value certified for the SRM984 Rb standard (the stated uncertainty in SRM984 is \pm 770 ppm). Whether this reflects the existence of natural isotope variations in Rb, or simply an underestimate of the SRM certified value is unclear. If the former, then it will perhaps be important to use basalt glass standards that are petrogenetically related the unknowns. Secondly, there is a clear tendency for the Rb-Sr solution runs to lie at the high side of the data series, averaging 2.5900 (1200 ppm lower than the canonical value for the SRM984 Rb standard). This would suggest a slight difference in Rb mass bias for solution runs versus laser ablation runs (and this would not be surprising, given the much larger plasma loading from the laser runs).

To assess the overall accuracy of this correction scheme, we have "corrected" each run in two ways: first, using the contiguous bracketing runs, and secondly, correcting each run with the global average "required" ⁸⁵Rb/⁸⁷Rb ratio of 2.58745. The resulting "corrected" ⁸⁷Sr/⁸⁶Sr ratios are then compared with the known values (note that the internal precision of the solution runs averaged about \pm 10 ppm (1 σ , standard error); that of the laser runs averaged \pm 17 ppm, with 88 intensities ranging from 1-12 volts). For the "bracketing technique", the mean deviation (measured against TIMS) for the nine analytical sessions is 127 ppm (320 ppm, 2 σ standard deviation, assuming a Gaussian distribution); for the "global average" technique, the mean deviation is 155 ppm (Fig. 3). We adopt the first scheme to correct unknowns and when discussing the accuracy of

standard glass (and melt inclusion) runs; this technique has the advantage of capturing the downward drift of the ⁸⁵Rb/⁸⁷Rb required values over time (see Fig. 2). While obviously not as precise as TIMS analyses, this level of reproducibility is excellent for an *in situ* technique, and will allow us to embark on a realistic program of melt-inclusion analysis.

Due to the uncertainty of the ⁸⁵Rb/⁸⁷Rb required value, there will be an error magnification on the corrected ⁸⁷Sr/⁸⁶Sr ratios of basaltic unknowns as Rb/Sr increases. In order to model this effect, we selected several glass standard runs with different Rb/Sr ratios and similar, low Kr/Sr ratios, and varied the 85Rb/87Rb ratio by 2 standard deviations about the mean global value of 2.58745 (±0.00278, 2σ standard deviation). The propagated error increases linearly to 630 ppm when Rb/Sr is 0.14, and the accuracy appears to scale with the Rb/Sr such that the uncertainty in ⁸⁵Rb/⁸⁷Rb magnifies the error on the final ⁸⁷Sr/⁸⁶Sr (Fig. 4). However, the "bracketing method" for the Rb correction reduces the uncertainty in the 85 Rb/87 Rb required value by ~20%, so that the propagated error from the reduced uncertainty translates to an improvement in the accuracy of the ⁸⁷Sr/⁸⁶Sr r (i.e., to 505 ppm when Rb/Sr is 0.14). The internal precision for ⁸⁷Sr/⁸⁶Sr measurements on lasered glasses is generally an order of magnitude better than the accuracy (this is also true for the Rb/Sr measurements), and does not appear to worsen with increasing Rb/Sr ratios. However, the internal precision of the 87Sr/86Sr, but not the accuracy, does vary with Sr intensities (Fig. 5), and, to a lesser extent, the number and length (8 or 16 second integrations) of cycles of analysis. The precision on the melt inclusion analyses also varies as a function of Sr intensity and the number and length of cycles. However, because the accuracy of the ⁸⁷Sr/⁸⁶Sr of Samoan glasses is not related to Sr intensity over the range of 1-12 Volts on mass 88 (Table 4), somewhat lower Sr intensities during the melt inclusion runs (1-6.5 Volts) should not affect the overall accuracy of 87Sr/86Sr.

2.6. Other Interferences

We have directly measured possible isobaric mass interferences resulting from Ca dimers and argides, FeO2, doubly-charged Er and Yb and KrH. The Ca dimers and argides invoke errors of less than <10 ppm, for Ca/Sr ratios typical of alkali basalts; there is no straightforward way to monitor or correct for these. Similarly, in a typical basalt laser run (with 500 ppm Sr), the total propagated error in ⁸⁷Sr/⁸⁶Sr from FeO₂ will then be less than 40 ppm (and could be nil). Additionally, the REE^{**}invoke errors of <30 ppm for typical Sr/REE ratios in alkali basalts; these can in principal be corrected for by monitoring the 83.5 and 85.5 half-mass peaks due to ¹⁶⁷Er⁺⁺ and ¹⁷¹Yb⁺⁺. Finally, there seems to be some evidence for the formation of KrH in the plasma, with Kr/KrH exhibiting ratios from 35-70. Kr hydrides create isobaric interferences on masses 83, 84, 85 and 87, and they may play an important role when Kr/Sr ratios are elevated. In practice, however, it appears that there are frequently other unexplained interferences at these masses, as well as at the Kr and Rb masses, so that correction for these isobaric interferences is not always successful. These problems not only limit the attainable precision of laser ablation analyses, but can limit the precision of straightforward Sr solution analyses as well, even of the SRM987 standard. A continuing investigation of these issues (and a fuller discussion of interferences from Ca dimers, argides, FeO₂, KrH and doubly charged Er and Yb) is underway (Hart et al., in prep).

2.7. Measurement of Rb/Sr by LA-ICP-MS

Masses 85 and 88 represent pure Rb and Sr, respectively, so that fairly precise measurement of Rb/Sr ratios can be generated. After correcting for mass fractionation during each run, Rb/Sr ratios on Samoan basalt glasses measured by laser ablation are accurate to 7% (1 mean deviation, compared to ratios obtained by XRF/ICP techniques on the same glasses), and precise (0.67%, 1 mean deviation) during multiple runs on a suite of Samoan glasses (Fig. 6). We note that this technique does not appear to work when running a "dry" plasma; Rb and Sr are strongly fractionated from each other, so that the measured Rb/Sr ratios are up to 80% higher than in the standard glass (see last 4 analyses in Table 4).

2.8. Replicability

We have not yet made a comprehensive study of reproducibility of sample analyses. We have done an analysis of a group of 50 runs on the SRM987 standard, spread over a 17 month time period, using solutions varying in concentration from 10 ppb to 600 ppb. All errors discussed here will be given at the 10 level. Considering first the 200 and 600 ppb solutions, the average internal precision for these was 7 ppm (for ⁸⁸Sr intensities varying from 8-28 volts). The average external precision, calculated from the variations within a single day's analysis session, was 12 ppm (and the number of standards run during each of these sessions varied from 2-6). While the overall ratio of external to internal precision was ~1.8, there was not a significant correlation between external and internal precision on a session-to-session basis (i.e. the internal precision on individual runs is not a good guide to the expected external precision). For the daily means of 11 sessions over the 17 month time period, the average ⁸⁷Sr/⁸⁶Sr was 0.710255, with a 15 standard deviation of a single analysis of 15 ppm. We should note that this time period involved a number of different baseline protocols and cup configurations, so the statistics may not be representative of our current procedures. There is one obvious conclusion, however, and that is that the precision on a daily basis is similar to the longterm precision; in other words, the variability in standard runs has almost as much "daily" scatter as it does "yearly" scatter.

Within this data set, there is a fair correlation between Sr intensity and the precision of the data; external precision for a given session is ± 5 ppm for >25 volt data, \pm 7-20 ppm for 9-15 volt data and \pm 50-80 ppm for 0.3-1 volt data. Overall, the 1 σ external precision in ppm, as a function of ⁸⁸Sr intensity in volts, may be empirically expressed as: ~ [80/(volts)^{0.8}].

3. Summary

The Finnigan NEPTUNE multi-collector ICP-MS has proven to be an excellent instrument for developing robust Sr isotope analysis protocols. Among its advantages are the stability of the Faraday-amplifier system, the ability to run with intensities up to 50 volts, and the stability of the mass bias for Sr and Rb. While the external precision of solution analyses are not yet comparable to the best TIMS techniques, the *in situ* laser ablation technique appears to be a reliable and very promising tool for the study of small scale-length isotopic heterogeneities, even in samples with a significant Rb component.

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Figure Captions

Fig. 1. Variation of intensity ratio of mass 82 to mass 83 over time. Series of Samoan basalt glasses analyzed by laser ablation (colored symbols) and mixed Sr-Rb standard solutions (SRM987-SRM984, black diamonds). The certificate value for ⁸²Kr/⁸³Kr is 1.004. We observe no relationship between mass 82/mass 83 and the precision or accuracy of ⁸⁷Sr/⁸⁶Sr measurement by LA-MC-ICPMS.

Fig. 2. ⁸⁵Rb/⁸⁷Rb ratio required to give correct ⁸⁷Sr/⁸⁶Sr of a series of Samoan basalt glasses (analyzed by TIMS, open diamonds) and mixed Sr-Rb standard solutions (SRM987-SRM984, filled diamonds). The certificate value for ⁸⁵Rb/⁸⁷Rb in SRM984 is 2.593 ± 0.002 ; all of the "calculated" Rb ratios are lower than the certificate value, and outside quoted error limits. Runs using a dry plasma not shown.

Fig. 3. Reproducibility of 87 Sr/ 86 Sr for 12 Samoan basalt glass standards (with known 87Sr/86Sr by TIMS) by LA-MC-ICPMS. Two data trends represent reproducibility (external precision) using the "global average" and the "bracketing" correction schemes for the isobaric interference of Rb on mass 87. The bracketing method gives better overall external precision, and is adopted as the correction scheme in the manuscript. Mixed NBS987 and NBS984 solution runs not included in the figure. The downward drift of 85 Rb/ 87 Rb (see Fig. 2) is noted in the downward drift of error using the "global average" Rb correction scheme. Error bars on symbols are internal (in-run) precision (2 mean deviations). Dashed lines mark external precision of ± 320 ppm (2 σ , standard deviation). Larger scatter in later runs is due to higher Rb/Sr ratios of the glass standards analyzed. Better external precision in early runs is due to smaller variation in 85 Rb/ 87 Rb over time. The glass standard runs at the beginning and the end of each analytical session are not correctible using the bracketing technique and are not plotted. However, all glass standards are correctible using the "global average" Rb correction scheme, so more data points are plotted for this latter correction scheme.

Fig. 4. Rb/Sr vs. propagated error on 87 Sr/ 86 Sr measurement due to uncertainty of the 85 Rb/ 87 Rb required value. The average 85 Rb/ 87 Rb required values (used for correcting the 87 Rb isobaric interference on 87 Sr) during all laser runs on Samoan glass standards is 2.58745 (±0.00278, 2 σ), and this uncertainty generates larger errors on the final 87 Sr/ 86 Sr at higher Rb/Sr ratios. The upper model line shows the error (2 σ , standard deviation) on final 87 Sr/ 86 Sr ratio when using the "global average" technique for the Rb correction (see text for description). Symbols on line represent error propagation on (from the uncertainty on the 85 Rb/ 87 Rb) using actual basalt analyses by laser ablation. Alternatively, error using the "bracketing" method for the Rb correction is 20% lower, and is described by the lower line. The Rb correction contributes the vast majority of the error on the final 87 Sr/ 86 Sr in our protocol for measuring basal glasses by LA-MC-ICPMS. Therefore, we consider the lower model curve to be a good approximation of the reproducibility of our method.

Fig. 5. Relationship between intensity on mass 88 and internal precision during laser ablation analysis of Samoan glass standards. Various Neptune and laser operating conditions and variations in basalt and plasma chemistry, inclusing Rb/Sr and Kr/Sr, show little or no relationship with measurement precision. The curve models the relationship between intensity and internal precision (1 mean deviation), and is described by the following relationship: Precision = $45*(\sqrt{Intensity})/Intensity$.

Fig. 6. Samoan glass Rb/Sr measured by LA-MC-ICPMS Neptune plotted against Rb/Sr measured by standard techniques (XRF or ICP) on powders of the same glasses. Error bars (2σ standard deviation from the mean of measured values) are shown on samples with at least 9 laser analyses. All other glasses had one or two analyses. One-to-one line is solid, and weighted least squares regression (including equation) is dashed.

Table 1. Typical NEPTUNE and Laser Operating Parameters

RF Power	1200 W
Argon cooling gas, flow rate	15L/min
Ar auxiliary gas, flow rate	0.8 L/min
He sample gas, flow rate	0.5 L/min
Interface cones	X-cones
Mass analyzer pressure	5x10 ⁻⁹ mbar
Resolution mode	Low (400)
Abundance sensitivity	~ 1 ppm at -1 amu, mass 88 (RPQ off)
Detection system	Faraday cups (9)
Signal analysis set-up	8 sec or 16 sec integration/cycle, 40 or 20 cycles
Background-baseline protocol	1 min integration pre-analysis, -0.7 amu all Faradays
Nebuliser and uptake rate	Elemental Scientific, Inc., PFA 20-50 microliters per minute
Spray chamber	Elemental Scientific, Inc., Stable Introduction System (SIS)
Sensitivity, solution Sr	75 V ⁸⁸ Sr per ppm Sr in 5% HNO ₃
Sensitivity, laser Sr	$\sim 1 \ \rm volt$ $^{\rm 88} \rm Sr \ per \ 100 \ ppm \ Sr \ in \ basalt \ at \ 1 \ mj$
Laser type	New Wave UP213, quad Nd YAG 213 nm
Carrier gas	Helium
Beam optics	Apertured mode
Spot size	120 um (60 um or less in high-Sr carbonates)
Raster pattern	200 x 200 um, 20 um line spacing, 4 um/sec
Pulse rate	20 Hz
Power setting	100% (lower for high-Sr carbonates)
Power delivered	~ 1.0-1.3 mj
Pre-ablation	same raster and spot size, 5 Hz, 45% power, 30 um/sec

Cup	Mass
Cup	IVIASS
H4	88
H3	87
H2	86
H1	85
С	84
L1	83
L2	82.5
L3	82
L4	81.5

Table 3. Adopted Isotopic Abundances.													
<u>Krypton</u>	Isotopic Abuna	nce											
78	0.003469												
80	0.02257												
82	0.11523												
83	0.11477												
84	0.56998												
86	0.17398												
From Ozima and Podo	sek, 2001												
Strontium SRM987	Isotopic Abunda	nce**	Isotopic Abundance*										
84	0.005578		0.005574										
86	0.098600		0.098566										
87	0.070029		0.070015										
88	0.825793		0.825845										
	Isotopic Ratio	S**	Isotopic Ratios*										
84/86	0.056573		0.056549	±2529	ppm								
87/86	0.710194		0.710339	±367	ppm								
88/86	8.375209		8.37861	±388	ppm								
84/88	0.006755		0.006749	±2529	ppm								
86/88	0.119400		0.119352	±388	ppm								
*NIST Certificate Value	s for SRM987 with	quote	d uncertainties.										
**Renormalized to 0.11	94 using exponent	ial law	; this value has been u	sed by									
community consensus	for decades. The	corres	sponding ⁸⁷ Sr/ ⁸⁶ Sr value	adopted by									
most of the communit	v is 0.71024.												
Rubidium SRM984													
85	0.72168												
87	0.27832												
85/87*	2.593	±771	ppm										
*NIST Certificate Value	s for SRM984, with	quote	ed uncertainties.										

100100 1001					10 00 00 M	0	8	1.000-00101	1000		10.000
Date Glass	Mass 82	Mass 83 Mass 84 Mass 85 Mass 86 Mass 87 Mass 88 ^{co} Rb/ ^{or} Rb Mass ^{or} Sr/ ^{or} Sr ^{or} S	Srl ^m Sr ^w Srl ^m Sr ^w	Srl [®] Sr	St/"St "St	Irl"Sr	Sr/"Sr Rb/Sr	Rb/Sr	Rb/Sr	Rb/Sr mas	jo# s
Name ^a	Meas'd	Meas'd Meas'd Meas'd Meas'd Meas'd Meas'd Required [®] Bias [®] Laser [®] La	Laser ^d Laser ^m I.	. Prec. [®]	E. Prec. ¹ E.	Prec.	TIMS [§] Laser ⁴	I. Prec. ^e E	. Prec. ^f	XRF ⁹ 82/8	8" Cycles'
01/26/04 75-10	0.0019	0.0015 0.070160 0.383076 1.161550 0.981164 9.99332 2.58799 1.45 0.70451 0.7	.70450	₽	41	1	.704533 0.041	1.18	8.4	0.045 0.000	119 20
01/26/04 75-10	0.0019	0.0015 0.068130 0.410905 1.130253 0.970173 9.72817 2.58762 1.47 0.70452 0.7	.70452 0.70454	Ŧ	19	12 0	.704533 0.045	0.42	6.0	0.045 0.000	120 20
01/26/04 75-10	0.0019	0.0015 0.056972 0.316560 0.928578 0.788522 7.99055 2.58761 1.48 0.70452 0.7	70452 0.70456	10	16	39 0	.704533 0.042	0.97	5.4	0.045 0.000	123 20
01/26/04 75-10	0.0018	0.0014 0.038589 0.210680 0.591571 0.505435 5.08531 2.58867 1.49 0.70447 0.7	.70446 0.70450	Ŧ	97	41 0	.704533 0.044	1.00	Ę	0.045 0.000	35 20
01/26/04 75-10	0.0019	0.0015 0.059566 0.310559 0.978671 0.821938 8.42300 2.58865 1.48 0.70448 0.7	.70447 0.70456	6	85	34	.704533 0.039	3.32	12.0	0.045 0.000	122 20
01/26/04 76-13	0.0017	0.0014 0.055907 0.544459 0.915110 0.871425 7.87805 2.58963 1.50 0.70620 0.7	.70619 0.70633	13	285	89 0	.706395 0.074	0.30	4.5	0.077 0.000	122 20
01/26/04 76-8	0.0017	0.0014 0.054455 0.526363 0.889211 0.845639 7.65417 2.58922 1.49 0.70621 0.7	70621 0.70639	Ŧ	231	22 0	.706374 0.073	0.16	8.6	0.081 0.000	123 20
01/26/04 63-13	0.0017	0.0013 0.045648 0.441395 0.731734 0.698360 6.29660 2.58914 1.50 0.70536 0.7	.70536 0.70547	÷	226	71 0	.705520 0.075	0.90	4.7	0.071 0.000	028 20
01/26/04 68-3	0.0018	0.0013 0.045375 0.373755 0.727534 0.668448 6.26096 2.58796 1.50 0.70535 0.7	.70534 0.70542	15	62	51 0	705388 0.064	0.60	7.3	0.069 0.000	20 20
01/26/04 72-2	0.0019	0.0013 0.048745 0.436044 0.789158 0.737371 6.79224 2.58771 1.50 0.70537 0.7	70537 0.70538	19	37	16 0	.705395 0.069	0.37	15.1	0.060 0.000	128 20
01/26/04 73-12	0.0019	0.0013 0.045637 0.450865 0.733196 0.704204 6.30990 2.58719 1.50 0.70668 0.7	70667 0.70670	14	28	63 0	.706653 0.076	0.58	0.1	0.076 0.000	30 18
01/26/04 71-2	0.0019	0.0013 0.045247 0.412129 0.726657 0.683560 6.25381 2.58762 1.50 0.70593 0.7	.70592 0.70597	44	27	37 0	.705943 0.070	0.22	6.8	0.075 0.000	330 20
01/26/04 73-1	0.0018	0.0013 0.050270 0.533292 0.818134 0.797957 7.04356 2.58866 1.51 0.70660 0.7	.70660 0.70666	10	175	86 0	.706720 0.081	0.49	0.3	0.081 0.000	126 20
01/26/04 70-1	0.0016	0.0010 0.035357 0.270199 0.577103 0.519443 4.96111 2.58846 1.43 0.70530 0.7	.70530 0.70540	14	107	41 0	.705371 0.058	0.70	2.1	0.060 0.000	333 20
01/26/04 71-11	0.0016	0.0010 0.036493 0.316405 0.590123 0.547012 5.07213 2.58908 1.43 0.70526 0.7	.70526 0.70537	12	194	27 0	.705394 0.067	0.27	4.7	0.070 0.000	331 20
01/26/04 71-22	0.0017	0.0013 0.041431 0.362145 0.673708 0.625087 5.79313 2.58923 1.44 0.70533 0.7	.70532	14	211	1	.705473 0.067	0.55	6.1	0.071 0.000	129 20
01/27/04 73-1	0.0015	0.0010 0.042357 0.501491 0.715500 0.713457 6.18177 2.58963 1.65 0.70655 0.7	.70648	14	336	1	.706720 0.086	0.54	7.0	0.081 0.000	24 20
01/27/04 76-13	0.0012	0.0009 0.041502 0.427055 0.703095 0.674792 6.07676 2.58938 1.67 0.70628 0.7	70621 0.70641	14	259	25 0	.706395 0.075	0.45	3.4	0.077 0.000	17 120
01/27/04 63-13	0.0011	0.0009 0.035692 0.355703 0.599686 0.571649 5.18396 2.58951 1.69 0.70539 0.7	.70533 0.70553	17	270	8	.705520 0.073	1.43	1.8	0.071 0.000	121 20
01/27/04 78-1	0.0011	0.0008 0.041195 0.639397 0.702572 0.761301 6.07662 2.58977 1.70 0.70862 0.7	.70856 0.70875	16	462 1	185 0	.708886 0.112	1.26	11.1	0.126 0.000	118 20
01/27/04 75-10	0.0011	0.0008 0.042356 0.280778 0.725189 0.631261 6.27295 2.58813 1.70 0.70455 0.7	.70449 0.70463	12	61	137 0	.704533 0.047	0.41	6.1	0.045 0.000	117 20
01/27/04 76-13	0.0016	0.0013 0.055752 0.552520 0.923794 0.881225 7.95726 2.58977 1.51 0.70624 0.7	.70618 0.70632	12	304	111 0	.706395 0.074	0.27	4.1	0.077 0.000	10 10
01/27/04 76-13	0.0016	0.0013 0.055497 0.554683 0.920418 0.879848 7.93062 2.58968 1.53 0.70625 0.7	.70619 0.70641	10	295	17 0	.706395 0.075	0.26	3.5	0.077 0.000	120 20
01/27/04 76-13	0.0016	0.0013 0.058108 0.578687 0.970494 0.925562 8.36646 2.58986 1.55 0.70624 0.7	.70617 0.70638	ŧ	315	24 0	.706395 0.074	0.41	0.8	0.077 0.000	119 20
01/27/04 76-13	0.0015	0.0013 0.049800 0.506317 0.817493 0.786891 7.04391 2.58966 1.54 0.70625 0.7	.70618	12	301	1	706395 0.077	0.42	4.6	0.077 0.000	122 20
02/13/04 71-11	0.0020	0.0020 0.052127 0.412318 0.785318 0.723659 6.73813 2.58794 1.41 0.70542 0.7	.70535	13	62	1	705394 0.065	0.21	6.5	0.070 0.000	30 20
02/13/04 78-8	0.0020	0.0019 0.049567 0.579772 0.740226 0.759702 6.35140 2.58835 1.42 0.70757 0.7	.70750	13	159	1	.707614 0.098	0.40	9.0	0.097 0.000	32 20
03/26/04 70-1	0.0019	0.0016 0.046911 0.327674 0.704302 0.633370 6.06027 2.59680 1.56 0.70545 0.7	.70541	17	60	1	705371 0.058	0.27	3.2	0.060 0.000	332 20
12/09/04 75-10	0.0027	0.0026 0.062513 0.321199 0.902189 0.770682 7.75525 2.58656 1.56 0.70465 0.7	.70458	6	64	1	.704533 0.044	0.34	1.4	0.045 0.000	35 49
12/09/04 75-10	0.0025	0.0025 0.062778 0.327432 0.917974 0.785039 7.90096 2.58672 1.61 0.70464 0.7	.70457 0.70459	9	52	82 0	.704533 0.044	0.17	1.5	0.045 0.000	332 48
12/09/04 76-13	0.0025	0.0024 0.063843 0.557662 0.950839 0.902924 8.19690 2.58901 1.67 0.70632 0.7	70625 0.70629	10	203	144 0	.706395 0.072	0.13	6.6	0.077 0.000	30 49

ies 82 to 88 5 Value ŝ --Incel data 1 blized⁸⁷Sr/⁸⁶Sr Table 4. Corrected and nor

00030 48	00043 50	00035 50	00037 37	00036 38	00069 38	00059 40	00046 40	00133 38	00106 40	00129 27	00043 40	00073 30	00046 40	00133 31	00061 40	00073 40	00056 40	00066 40	00052 40	00055 40	00070 40	00074 40	00075 40	00087 40	00043 40	00043 40	00101 40	00106 26	00067 40	00067 40	CONC. I STORE SHOWS	00053 40
9 0.077 0.	0 0.045 0.	2 0.045 0.	2 0.077 0.	3 0.077 0.	7 0.075 0.	5 0.075 0.	9 0.077 0.	2 0.075 0.	1 0.075 0.	2 0.075 0.	2 0.077 0.	5 0.077 0.	3 0.077 0.	3 0.075 0.	1 0.075 0.	\$ 0.075 0.	3 0.075 0.	6 0.077 0.	8 0.077 0.	2 0.077 0.	3 0.075 0.	3 0.077 0.	1 0.075 0.	7 0.075 0.	2 0.077 0.	6 0.077 0.	2 0.075 0.	5 0.075 0.	0 0.075 0.	2 0.075 0.		0.0/0 0.
3 6.	4 7.4	9 11.	7 2.	9 1.1	4	0 4.5	6 12	1	4 3.	8	8 12	8 1.	0 13.	4 3.1	8	8	3 16.	5 19.	6 25.	7 28.	4 1.	0 11.	9 1.	7 4.	3 20.	6 26.	9	1 5.4	8 7.	4 10.		0
72 0.1	42 1.6	40 2.0	76 0.1	76 0.1	72 0.2	72 0.3	87 0.3	76 0.5	78 0.3	79 0.6	37 0.3	78 0.9	38 0.3	78 0.8	82 0.6	33 0.5	38 0.6	32 0.6	97 0.6	99 0.4	77 0.5	36 0.6	76 0.4	79 0.4	93 0.2	38 0.4	79 0.3	30 0.7	81 0.5	83 0.2	00 01	200
0.706395 0.0	0.704533 0.0	0.704533 0.0	0.706395 0.0	0.706395 0.0	0.705943 0.0	0.705943 0.0	0.706395 0.00	0.705943 0.0	0.705943 0.0	0.705943 0.0	0.706395 0.01	0.706395 0.0	0.706395 0.01	0.705943 0.0	0.705943 0.01	0.705943 0.01	0.705943 0.01	0.706395 0.0	0.706395 0.0	0.706395 0.0	0.705943 0.0	0.706395 0.01	0.705943 0.0	0.705943 0.0	0.706395 0.0	0.706395 0.0	0.705943 0.0	0.705943 0.01	0.705943 0.01	0.705943 0.01	0 705043 0.0	0.0 0+000 /.0
215	175	I	1	76	95	132	223	81	75	278	360	290	338	1	1	36	39	53	48	126	149	148	79	131	135	227	142	55	22	÷	10	\$
203	133	32	136	76	158	54	258	5	56	268	83	233	44	291	26	60	135	135	245	255	1	128	43	42	260	315	78	126	63	43		-
6	16	12	12	Ħ	15	12	13	19	26	43	19	47	19	43	18	17	16	19	16	16	17	16	20	17	10	ŧ	12	26	15	14		2
.69 0.70632 0.70625 0.70624	.77 0.70469 0.70463 0.70466	.82 0.70462 0.70456	.73 0.70650 0.70649	74 0.70645 0.70645 0.70634	70 0.70606 0.70605 0.70601	69 0.70599 0.70598 0.70585	66 0.70658 0.70658 0.70655	.73 0.70595 0.70595 0.70589	78 0.70591 0.70590 0.70600	.77 0.70576 0.70575 0.70575	76 0.70646 0.70645 0.70665	71 0.70624 0.70623 0.70619	78 0.70643 0.70643 0.70663	.82 0.70574 0.70574	.29 0.70599 0.70596	.33 0.70601 0.70599 0.70592	38 0.70607 0.70604 0.70597	.44 0.70652 0.70649 0.70636	.43 0.70659 0.70657 0.70643	42 0.70660 0.70658 0.70648	29 0.70597 0.70595 0.70584	.27 0.70651 0.70649 0.70650	31 0.70594 0.70591 0.70589	29 0.70594 0.70591 0.70585	27 0.70661 0.70658 0.70649	31 0.70664 0.70662 0.70656	29 0.70592 0.70589 0.70584	40 0.70588 0.70585 0.70590	42 0.70593 0.70590 0.70596	42 0.70594 0.70591 0.70594	35 0 70597 0 70504 0 70801	
2.58901 1	2.58560 1	2.58695 1	2.58638 1	2.58683 1	2.58615 1	2.58697 1	2.58571 1	2.58737 1	2.58781 1	2.58933 1	2.58686 1	2.58910 1	2.58712 1	2.58951 1	2.58694 1	2.58698 1	2.58651 1	2.58655 1	2.58594 1	2.58591 1	2.58735 1	2.58653 1	2.58772 1	2.58771 1	2.58577 1	2.58552 1	2.58798 1	2.58832 1	2.58785 1	2.58770 1	1 00200	C.000.7
4 0.0024 0.062189 0.541662 0.926303 0.879131 7.98773	1 0.0022 0.041088 0.190853 0.563339 0.478042 4.85665	0 0.0019 0.044719 0.210583 0.648179 0.547229 5.59980	9 0.0032 0.062592 0.558316 0.908212 0.873053 7.83316	8 0.0031 0.062148 0.561306 0.908499 0.874702 7.83946	3 0.0035 0.045848 0.326355 0.560609 0.529545 4.81303	5 0.0038 0.053988 0.403808 0.691899 0.654374 5.94536	6 0.0044 0.065799 0.643348 0.910299 0.907215 7.83305	3 0.0034 0.031707 0.179591 0.292804 0.278502 2.49353	9 0.0019 0.019831 0.129680 0.205503 0.197576 1.75832	9 0.0020 0.018789 0.111868 0.176592 0.169506 1.50584	1 0.0027 0.040442 0.401984 0.569223 0.567925 4.90895	1 0.0026 0.027671 0.212827 0.335122 0.323659 2.87663	9 0.0025 0.034968 0.346399 0.485332 0.485705 4.18601	8 0.0018 0.016824 0.097919 0.154975 0.148658 1.32174	6 0.0051 0.069247 0.570788 0.881891 0.852357 7.52080	5 0.0051 0.058418 0.470999 0.717218 0.695719 6.11468	0 0.0048 0.061762 0.582067 0.830511 0.822118 7.10423	1 0.0035 0.042900 0.404831 0.547640 0.550584 4.68347	4 0.0041 0.054000 0.586498 0.753766 0.770557 6.45857	5 0.0044 0.053429 0.584683 0.737877 0.758293 6.31971	6 0.0049 0.064039 0.471377 0.775869 0.736914 6.60756	9 0.0053 0.064899 0.528307 0.775994 0.759615 6.60372	8 0.0050 0.063217 0.452777 0.746571 0.708570 6.35639	0 0.0052 0.059835 0.419486 0.671236 0.641146 5.70452	2 0.0066 0.099696 1.041701 1.410914 1.417626 12.0573	0 0.0065 0.096919 1.069271 1.373301 1.402463 11.7439	8 0.0050 0.053745 0.351423 0.566159 0.539079 4.80269	2 0.0044 0.045297 0.297450 0.470912 0.450860 4.00124	2 0.0045 0.058865 0.471070 0.730507 0.706057 6.24085	9 0.0044 0.054667 0.450610 0.678735 0.661178 5.79870	0 0 0055 0 070054 0 674603 4 063063 4 023263 0 06473	CITED'S COLATA 1 2000001 2001 1000 4000 100 000000 0
0.0024	0.0021	0.0020	0.0029	0.0028	0.0033	0.0035	0.0036	0.0033	0.0019	0.0019	0.0021	0.0021	0.0019	0.0018	0.0046	0.0045	0.0040	0.0031	0.0034	0.0035	0.0046	0.0049	0.0048	0.0050	0.0052	0:0050	0.0048	0.0042	0.0042	0.0039	01000	0.0040
12/09/04 76-13	12/09/04 75-10	12/09/04 75-10	01/10/05 76-13	01/10/05 76-13	01/10/05 71-2	01/10/05 71-2	01/10/05 76-13	01/10/05 71-2	01/10/05 71-2	01/10/05 71-2	01/10/05 76-13	01/10/05 76-13	01/10/05 76-13	01/10/05 71-2	01/11/05 71-2	01/11/05 71-2	01/11/05 71-2	01/11/05 76-13	01/11/05 76-13	01/11/05 76-13	01/11/05 71-2	01/11/05 76-13	01/11/05 71-2	01/11/05 71-2	01/11/05 76-13	01/11/05 76-13	01/11/05 71-2	01/11/05 71-2	01/11/05 71-2	01/11/05 71-2	04/11/DE 71-2	

08/09/05 71-2	0.0010	0.0012 0.039667 0.410998 0.676025 0.646988 5.81840 2.58823 1.44 0.70587 0.70587 0.705	0593 21	108	20	0.705943 0.076	0.77	0.1	0.075 0.00017	40
08/09/05 75-10	0.0010	0.0009 0.041922 0.267759 0.717313 0.618485 6.17226 2.58758 1.43 0.70453 0.70452 0.704	0460 13	15	96	0.704533 0.046	1.00	3.8	0.045 0.00016	4
08/09/05 71-2	0.0009	0.0009 0.034364 0.350906 0.581416 0.555389 5.00466 2.58932 1.46 0.70577 0.70577 0.705	0574 16	249	288	0.705943 0.075	0.42	0.7	0.075 0.00019	40
08/09/05 75-10	0.0010	0.0012 0.044398 0.305593 0.761972 0.665350 6.55471 2.58663 1.41 0.70458 0.70458 0.704	0468 14	67	209	0.704533 0.050	0.81	11.6	0.045 0.00015	4
08/09/05 76-13	0.0009	0.0011 0.040285 0.445631 0.686601 0.668344 5.90585 2.58878 1.41 0.70626 0.70626 0.706	0619 18	192	291	0.706395 0.081	0.51	4.5	0.077 0.00016	4
08/09/05 76-13	0.0011	0.0021 0.058298 0.720332 1.014084 1.011925 8.72086 2.58676 1.38 0.70647 0.70646 0.706	9663 16	97	331	0.706395 0.088	0.79	14.5	0.077 0.00013	4
08/09/05 71-2	0.0005	0.0005 0.017692 0.167519 0.288194 0.271203 2.45862 2.58909 1.03 0.70580 0.70580 0.705	0570 22	210	347	0.705943 0.074	0.79	2.2	0.075 0.00020	4
08/09/05 75-10	0.0006	0.0008 0.029739 0.191258 0.502920 0.433086 4.29411 2.58584 1.03 0.70463 0.70462	- 18	127	ł	0.704533 0.048	1.71	7.8	0.045 0.00015	4
08/10/05 71-2	0.0008	0.0008 0.030262 0.298953 0.507454 0.483010 4.38372 2.58702 1.65 0.70598 0.70598	- 16	48	1	0.705943 0.072	0.63	4.0	0.075 0.00018	4
08/10/05 75-10	0.0008	0.0010 0.038555 0.249734 0.660744 0.572473 5.70962 2.58474 1.65 0.70469 0.70469 0.704	0462 17	217	123	0.704533 0.046	0.83	3.9	0.045 0.00015	6
08/10/05 76-13	0.0010	0.0017 0.046609 0.547945 0.808282 0.798610 6.98314 2.58548 1.63 0.70659 0.70659 0.706	17 12	280	92	0.706395 0.083	0.77	8.0	0.077 0.00014	4
08/10/05 71-2	0.0008	0.0009 0.031222 0.316861 0.530073 0.506826 4.58580 2.58748 1.72 0.70594 0.70594 0.705	0574 22	10	290	0.705943 0.073	0.28	2.9	0.075 0.00018	4
08/10/05 75-10	0.0008	0.0009 0.038968 0.201179 0.669617 0.559369 5.78558 2.58498 1.64 0.70464 0.70464 0.704	0466 12	157	177	0.704533 0.037	4.72	17.4	0.045 0.00013	4
08/10/05 71-2	0.0007	0.0006 0.018731 0.177420 0.298953 0.284978 2.58205 2.58792 1.70 0.70590 0.70590 0.705	0570 26	67	346	0.705943 0.073	0.65	3.4	0.075 0.00029	38
08/10/05 76-13	0.0010	0.0021 0.053202 0.659724 0.930221 0.931071 8.04115 2.58547 1.65 0.70660 0.70660	- 20	295	1	0.706395 0.087	0.60	12.8	0.077 0.00013	4
08/11/05 71-2	0.0008	0.0008 0.028474 0.280326 0.481040 0.457445 4.16744 2.58755 1.80 0.70593 0.70593	- 19	19	1	0.705943 0.071	0.54	5.7	0.075 0.00019	4
08/11/05 71-2	0.0008	0.0009 0.025628 0.249742 0.428075 0.407073 3.70673 2.58701 1.78 0.70598 0.70598 0.705	0597 21	49	43	0.705943 0.071	0.58	5.5	0.075 0.00021	4
08/11/05 71-2	0.0008	0.0009 0.027774 0.267394 0.463111 0.439103 4.00773 2.58715 1.75 0.70596 0.70596 0.705	0592 20	31	33	0.705943 0.071	1.25	6.4	0.075 0.00021	33
08/11/05 71-2	0.0006	0.0008 0.025284 0.255547 0.433576 0.413754 3.75980 2.58677 1.83 0.70600 0.70600 0.705	1597 23	80	36	0.705943 0.072	0.48	4.8	0.075 0.00016	4
08/11/05 71-2	0.0007	0.0008 0.026528 0.274040 0.455528 0.436904 3.94988 2.58697 1.83 0.70598 0.70598 0.705	1597 26	56	42	0.705943 0.073	1.15	2.9	0.075 0.00017	38
08/11/05 71-2	0.0007	0.0006 0.022004 0.216339 0.367855 0.350677 3.18822 2.58782 1.83 0.70590 0.70591 0.705	19 19	53	93	0.705943 0.072	0.40	2.0	0.075 0.00021	4
08/11/05 71-2	0.0007	0.0008 0.024642 0.251123 0.417276 0.400127 3.61669 2.58720 1.82 0.70596 0.70596 0.705	19 19	25	40	0.705943 0.073	0.51	2.8	0.075 0.00019	4
08/11/05 71-2	0.0007	0.0008 0.025226 0.258615 0.427220 0.410275 3.70300 2.58721 1.82 0.70596 0.70596 0.705	0598 21	25	25	0.705943 0.074	0.34	2.2	0.075 0.00019	6
08/11/05 71-2	0.0007	0.0007 0.023358 0.232289 0.395225 0.376807 3.42680 2.58804 1.84 0.70589 0.70589	- 17	81	1	0.705943 0.072	0.31	5.1	0.075 0.00019	40
01/16/06 78-1	0.0062	0.0061 0.070911 0.689532 0.762508 0.818940 6.52057 2.58430 1.69 0.70947 0.70932	- 13	610	1	0.708886 0.112	0.65	10.6	0.126 0.00096	45
01/16/06 78-1	0.0059	0.0058 0.050415 0.365045 0.405885 0.432392 3.43846 2.58490 1.72 0.70939 0.70924 0.708	3891 22	494	36	0.708886 0.113	1.03	10.3	0.126 0.00172	45
01/16/06 78-1	0.0058	0.0057 0.089039 1.0088890 1.140008 1.219198 9.79527 2.58586 1.68 0.70925 0.70909 0.708	3881 16	294	106	0.708886 0.109	0.72	12.9	0.126 0.00060	45
01/16/06 78-1	0.0056	0.0055 0.062132 0.584565 0.661023 0.704447 5.65166 2.58572 1.70 0.70927 0.70911 0.708	0878 26	323	146	0.708886 0.110	06.0	12.6	0.126 0.00098	39
01/16/06 78-1	0.0053	0.0053 0.042605 0.276774 0.311453 0.329490 2.62646 2.58406 1.71 0.70951 0.70935 0.709	3913 21	654	340	0.708886 0.112	0.32	11.0	0.126 0.00201	45
01/16/06 78-1	0.0050	0.0050 0.033419 0.140970 0.164176 0.169158 1.35428 2.58587 1.72 0.70925 0.70910 0.708	33 33	296	214	0.708886 0.110	0.58	12.1	0.126 0.00371	45
01/16/06 78-1	0.0050	0.0050 0.036845 0.206490 0.230955 0.243559 1.93289 2.58548 1.70 0.70931 0.70916 0.709	0910 35	382	306	0.708886 0.113	0.85	9.7	0.126 0.00259	27
01/16/06 78-1	0.0085	0.0085 0.049816 0.121730 0.146879 0.145002 1.15711 2.58816 1.77 0.70894 0.70878 0.708	3854 52	148	489	0.708886 0.111	0.56	11.3	0.126 0.00732	45
01/16/06 71-2	0.0155	0.0157 0.091273 0.139273 0.263105 0.227057 2.06673 2.58584 1.76 0.70624 0.70608 0.706	005 28	197	147	0.705943 0.071	0.37	5.4	0.075 0.00752	45
01/16/06 71-2	0.0183	0.0185 0.123513 0.333251 0.611060 0.551464 5.04318 2.58586 1.72 0.70623 0.70608	- 13	190	1	0.705943 0.070	0.37	7.2	0.075 0.00364	45

4	45	45	45
0.00273	0.00169	0.00225	0.00217
0.126	0.126	0.075	0.075
51.3	71.1	74.8	79.4
0.19	0.82	0.55	0.38
0.190	0.215	0.132	0.135
0.708886	0.708886	0.705943	0.705943
I	1	I	4
602	80	43	14
37	24	27	20
ı	;	ł	1
0.70846	0.70888	0.70591	0.70595
.70862	.70904	70607.	.70611
1.48 0	1.44 0	1.45 0	1.45 0
2.58924	2.58742	2.58759	2.58734
0.0105 0.0098 0.071513 0.683739 0.460786 0.592918 3.84615	0.0111 0.0104 0.090872 1.314651 0.775036 1.069266 6.54519	0.0119 0.0112 0.086823 0.653507 0.631526 0.700407 5.29883	0.0126 0.0119 0.093511 0.738091 0.694135 0.778190 5.82947
01/16/06 78-1	01/16/06 78-1	01/16/06 71-2	01/16/06 71-2

Analyses are listed in chronological order. Cycles are 16 second integrations of measurement during analysis until 3/26/04. Cycles are 8 second integrations thereafter.

^a All samples are dredge samples from the AVON3 dredging cruise (Workman et e/., 2004) [7]

bis Rbi²⁷Rb required values are obtained after correcting for Kr and Rb interferences, fractionation correction, and normalization to NBS 987 solution run:

during each analytical session. The average $^{\infty}$ Rb $^{
m K}$ Rb required value is 2.58745, which is used to obtain the Rb correction using the "global average" technique

^c Sr-isotopes are corrected for Kr and Rb interferences and fractionation corrected

⁴After correction for Rb and Sr interferences and fractionation, gass standard runs are corrected for 987 standard solution runs for each analytical session

Internal (in-run) precision (I. Prec.) is quoted as 10 standard error, and is calculated offline using the TweeKr program. In the above dataset, the average internal precision is 17 ppm for ⁸⁷5/⁸⁶Sr, and 0.67% for Rb/Sr. External precision (E. Prec.) is calculated assuming that the TIMS²⁰ St/⁴⁶Sr (ppm) and XRF Rb/Sr (%) values obtained on bulk powders are correct. The average external precision of ²⁰St/⁴⁶Sr measurements on the above glass standards is 155 ppm mean deviation (1o), which translates to 388 ppm standard deviation (2o) if a Gaussian distribution is assumed. The average external precision on Rb/Sr

measurements on glass standards is 6.7% (mean deviation, 1 d), which translates to a standard deviation of 16.7% (20). Uses global average 16 Rb 16 Rb value (2.68745) for isobaric Rb interference.

^{3 17}Sh⁷⁶Sr and Rb/Sr values obtained by conventional methods (XRF or ICP) are reported in Workman*et al.*, 2004 [7]. [™] Raw voltages on mass 82 divided by mass 88. This is a proxy for Kr/Sr during analysis

¹ Samples are run at less than 100% laser power (between 45-90% power)

These dry runs are not used in calculations or in the figures, and were performed for comparative purposes only. We find that the Rb is strongly fractionated from Sr, and that the Rb/R Rb required Unlike all other Samoan glass standard and melt inclusion analyses, these four analyses were performed with a "dry" plasma (i.e., no aspiration of the 5% nitric solution during laser analysis). and mass bias tend to me lower than similar runs made with a wet plasma during the same analytical session.

The amount of mass bias correction required to obtain the canonical value for⁸⁶ Sr^{Fig}Sr. Units are %/amu.

External precision is calculated assuming that the TIINS²⁷SA²⁶S r is correct. External precision is calculated using the "tracketing" technique for the correction of the isobaric interference from Rb on mass 87. The

external precision is somewhat improved (127 ppm, 1 mean deviation) compared to the global average technique for the Rb correction (155 ppm) when comparing the same set of glass analyses.

Bracketing is a useful technique for individual. 1-day analytical sessions only, as 📽 Rb^{/27}R b values can change significantly between analytical sessions. Therefore,

³ Same as previous column, but ⁶² S/⁶⁵S values calculated by correcting for the Rb interferece using the "bracketing method", our preferred correction scheme and the one adopted in the manuscript the first and last runs of the Samoan glass standards of each analytical session cannot be corrected with the bracketing technique. Runs with a dry plasma are not examined.

for correcting unknowns.

Table 5.	Composit	ions of Sa	moan gia	isses ("sta	ndards")	used in de	etermining	RD/ R	required	values in	this study	/.		
lass Nam	76-8	78-1	78-8	76-13	75-10	63-13	70-1	71-11	71-22	72-2	73-1	73-12	71-2	68-3
volcano	Malumalu	Malumalu	Malumalu	Malumalu	Tau	Vailulu'u	Vailulu'u	Vailulu'u	Vailulu'u	Vailulu'u	Vailulu'u	Vailulu'u	Vailulu'u	Vailulu'u
Sr/ Sr	0.706374	0.708886	0.707614	0.706395	0.704533	0.705520	0.705371	0.705394	0.705473	0.705395	0.706720	0.706653	0.705943	0.705388
SIO2	47.04	45.54	47.33	46.35	47.10	47.16	47.57	47.57	45.21	45.53	47.14	47.10	43.85	47.84
Al ₂ 03	13.69	8.94	14.19	13.55	16.29	14.76	13.63	13.71	8.13	14.10	13.97	12.07	6.37	11.86
TiO2	3.88	1.96	3.29	3.86	4.02	3.01	2.91	2.75	1.68	3.31	2.80	2.46	1.44	2.55
FeO*	11.93	11.89	11.45	12.78	12.56	11.59	11.44	10.64	10.77	12.13	11.26	10.25	10.80	10.51
MnO	0.18	0.17	0.17	0.20	0.20	0.18	0.18	0.18	0.17	0.19	0.18	0.18	0.17	0.17
CaO	11.86	8.23	11.47	11.80	9.11	12.37	13.05	12.72	9.99	11.88	12.93	12.80	8.07	12.75
MgO	6.27	19.68	6.93	6.42	5.16	6.57	7.56	8.60	21.79	8.56	7.52	11.70	27.44	10.68
K ₂ O	1.92	1.41	2.05	1.94	1.46	1.28	0.98	1.15	0.67	1.40	1.37	1.15	0.59	0.99
Na ₂ O	2.75	1.92	2.74	2.61	3.47	2.72	2.35	2.35	1.39	2.54	2.50	2.02	1.11	2.36
P205	0.49	0.27	0.40	0.49	0.63	0.37	0.34	0.33	0.20	0.37	0.34	0.28	0.17	0.29
Pre-total	99.27	100.27	99.74	99.12	99.58	99.01	98.81	98.74	98.71	98.39	99.10	99.13	98.50	99.49
Ni	65	616	97	71	5	59	57	135	778	110	79	220	1080	285
Cr	137	1289	106	130	15	110	98	316	1687	227	202	768	2141	822
V	356	194	342	384	282	339	338	324	212	367	325	295	177	300
Ga	22	14	21	26	26	20	21	19	14	23	14	17	9	20
Cu	72	49	79	95	21	58	82	91	156	79	70	59	138	82
Zn	121	102	104	125	140	95	92	92	78	102	86	77	77	89
Cs	0.51	0.39	0.60	0.5	0.39	0.28	0.31	0.34	0.21	0.39	0.38	0.31	0.19	0.30
Rb	44 3	418	51.4	42.8	27.6	33.6	25.0	30.4	19.0	29.3	38.1	30.4	16.9	26.6
Ba	355	255	372	341	264	279	203	247	149	287	347	282	128	218
Th	64	60	63	53	49	52	42	44	2.6	5.0	61	49	23	4.0
	1 30	1.05	1 36	1 29	1 10	1.07	0.93	0.95	0.56	1 11	1 15	0.87	0.49	0.87
Nb	56.7	39.0	59.0	57.5	51.3	45.1	37.1	40.5	24.7	47.5	46.7	38.2	20.7	36.4
Ta	3.9	2.6	3.8	3.8	3.6	31	25	28	16	3.2	31	25	1.4	2.5
1.0	40.4	20.4	52.0	40.6	43.4	37.4	31.0	32.0	20.2	20.2	44.2	26.1	17.4	20.2
Co	40.4	79.0	00.0	40.0	43.1	74.4	S1.0	92.9 0E 0	20.2	70.0	44.2	20.1	24.4	29.3
Db	4.2	13.0	30.0	50.3	30.5	74.4	00.0	00.0	40.4	70.3	03.0	00.0	1.9	00.7
Pu	4.2	4.2	4.0	0.2	3.3	2.0	2.2	2.5	1.0	2.0	4.0	0.2	1.0	2.0
PI	11.2	7.9	11.5	11.8	11.1	8./	7.0	1.8	4./	9.1	9.4	7.8	4.0	0.8
Na	40.5	31.0	44.2	41.2	48.0	35.4	31.5	31.0	19.0	31.2	30.0	30.2	10.5	20.3
51	070	333	529	554	017	4/0	420	434	207	492	4/3	398	224	387
21	2/0	183	200	281	327	218	198	180	117	221	212	1//	98	108
nı	1.2	4./	0.4	/.1	8.3	5.8	5.4	5.2	3.1	5.8	5.8	4.9	2.0	4.4
Sm	10.5	6.5	9.1	10.2	11.6	7.8	7.5	1.3	4.4	8.4	7.8	0.7	3.7	0.5
Eu	3.2	1.9	2.5	2.9	3.8	2.5	2.3	2.3	1.4	2.6	2.4	2.1	1.1	2.0
Gđ	9.5	5.8	8.0	9.1	10.7	7.1	6.8	6.4	3.9	7.3	6.8	6.0	3.2	5.8
TB	1.4	0.8	1.1	1.3	1.6	1.1	1.0	1.0	0.6	1.1	1.1	0.9	0.5	0.9
Dy	7.4	4.4	5.8	6.6	8.8	5.9	5.7	5.4	3.3	6.0	5.6	5.0	2.7	4.8
Ho	1.3	0.8	1.0	1.2	1.6	1.1	1.0	1.0	0.6	1.1	1.0	0.9	0.5	0.9
Y	32.2	19.6	30.5	33.1	40.2	27.5	26.4	25.9	15.6	29.3	27.3	23.5	12.6	22.5
Er	3.0	1.8	2.5	2.8	3.8	2.7	2.5	2.4	1.4	2.7	2.5	2.2	1.2	2.1
Tm	0.38	0.23	0.34	0.37	0.50	0.35	0.33	0.32	0.19	0.35	0.33	0.28	0.16	0.28
Yb	2.05	1.24	1.97	2.13	2.80	2.02	1.87	1.83	1.11	2.02	1.90	1.64	0.87	1.63
Lu	0.29	0.19	0.29	0.31	0.40	0.30	0.27	0.27	0.16	0.29	0.28	0.23	0.13	0.24
Sc	31.0	25.2	28.5	25.0	20.5	33.4	41.9	39.2	36.0	38.0	39.1	42.7	29.8	40.9



Fig. 1



Fig. 2



Fig.3



Supp Data Fig. 4



Fig. 5



Fig. 6

Chapter 2

The return of subducted continental crust in Samoan lavas*

Abstract

Substantial quantities of terrigenous sediments are known to enter the mantle at subduction zones, but little is known about their fate in the mantle¹. Subducted sediment may be entrained in buoyantly upwelling plumes and returned to the Earth's surface at hotspots²⁻⁵, but the proportion of recycled sediment in the mantle is small, and clear examples of recycled sediment in hotspot lavas are rare^{6,7}. Here we report remarkably enriched ⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd isotope signatures in Samoan lavas from three dredge locations on the underwater flanks of Savai'i island, Western Samoa. The submarine Savai'i lavas represent the most extreme ⁸⁷Sr/⁸⁶Sr isotope compositions reported for ocean island basalts to date. The data are consistent with the presence of a recycled sediment component (with a composition similar to the upper continental crust) in the Samoan mantle. Trace-element data show affinities similar to those of the upper continental crust—including exceptionally low Ce/Pb and Nb/U ratios⁸—that complement the enriched ⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd isotope signatures. The geochemical evidence from these Samoan lavas significantly redefines the composition of the EM2 (enriched mantle 2; ref. 9) mantle endmember, and points to the presence of an ancient recycled upper continental crust component in the Samoan mantle plume.

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The return of subducted continental crust in Samoan lavas

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Substantial quantities of terrigenous sediments are known to enter the mantle at subduction zones, but little is known about their fate in the mantle¹. Subducted sediment may be entrained in buoyantly upwelling plumes and returned to the Earth's surface at hotspots²⁻⁵, but the proportion of recycled sediment in the mantle hotspots2 is small, and clear examples of recycled sediment in hotspot lavas are rare^{6,7}. Here we report remarkably enriched ⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd isotope signatures in Samoan lavas from three dredge locations on the underwater flanks of Savai'i island, Western Samoa. The submarine Savai'i lavas represent the most extreme ⁸⁷Sr/⁸⁶Sr isotope compositions reported for ocean island basalts to date. The data are consistent with the presence of a recycled sediment component (with a composition similar to the upper continental crust) in the Samoan mantle. Trace-element data show affinities similar to those of the upper continental crustincluding exceptionally low Ce/Pb and Nb/U ratios⁸—that complement the enriched ⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd isotope signatures. The geochemical evidence from these Samoan lavas significantly redefines the composition of the EM2 (enriched mantle 2; ref. 9) mantle endmember, and points to the presence of an ancient recycled upper continental crust component in the Samoan mantle plume.

The Earth's mantle, as sampled by ocean island basalts erupted at hotspots, is chemically and isotopically heterogeneous. However, the origin of the geochemical heterogeneity of the mantle is not well understood. One model for the geochemical evolution of the mantle assumes that much of the chemical diversity is a result of subduction, a tectonic process that introduces enriched oceanic crust and compositionally heterogeneous sediment into a largely primitive (or slightly depleted) mantle^{5,10,11}. Following subduction, these surface materials mix with a peridotitic mantle, thus imprinting their enriched chemical and isotopic signatures on its various domains. A number of isotopically distinct geochemical reservoirs, as sampled by ocean island basalts, have resulted from this process. The isotopic endmembers are often referred to as HIMU (high $\mu = {}^{238}\text{U}/{}^{204}\text{Pb}$), EM1 (enriched mantle 1) and EM2 (enriched mantle 2) and DMM (depleted mid-ocean-ridge basalt mantle)9. Although the most radiogenic Pb isotope ratios observed in the HIMU component have been proposed to result from a contribution of recycled oceanic crust^{9,12}, most models for the creation of the EM1 and EM2 mantle reservoirs invoke a small portion of lithologically distinct sediments that have been recycled into the mantle9,13.

The volcanically active Samoan islands and seamounts define a hotspot track with a classical EM2 pedigree^{7,14,15}. The first high-precision ⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd measurements from Samoan lavas were interpreted as evidence of sediment recycling⁵. Recently,

however, the proposed recycled sediment origin of the enriched Samoan basalts has been questioned (see Supplementary Discussion), and an alternative model favouring source enrichment by metasomatic processes was proposed⁷. The extreme isotopic and chemical enrichment in the new Samoan EM2 lavas exhibit distinctly continental fingerprints, and argue for a role for a component similar to ancient recycled upper continental crust (UCC) in the Samoan plume (see Supplementary Discussion for the ALIA 2005 cruise dredge locations and geochemical data).

The most isotopically enriched Samoan whole-rock ⁸⁷Sr/⁸⁶Sr signature (0.720469, Mg# = 57.2) is recorded in a trachyandesite, dredge sample D115-21, which was taken from the southwestern flank of Savai'i. Clinopyroxene mineral separates from the same sample yielded an even higher ⁸⁷Sr/⁸⁶Sr ratio (0.721630). A trachybasalt (D115-18) hosts the second-most-enriched ⁸⁷Sr/⁸⁶Sr (0.718592, Mg# = 58.7), and clinopyroxene mineral separates from the sample also gave more enriched ratios (0.702032–0.720830). Six other lavas recovered in the same dredge also exhibit enriched ⁸⁷Sr/⁸⁶Sr ratios (0.708175–0.716394, Mg# = 52.0–65.1). Dredge D118, located on the far western end of the Savai'i lineament, contained an alkali basalt with enriched ⁸⁷Sr(⁸⁶Sr (0.710337, measured on fresh clinopyroxene). Dredge D128, taken on the northeastern flanks of Savai'i, yielded a transitional basalt with a high ⁸⁷Sr fratio (0.712500, Mg# = 70.5) and several other basalts with less enriched ⁸⁷Sr (0.706397–0.708170, Mg# = 61.2–63.9). Dredge D114, taken on the southwestern flanks of Savai'i, provided younger shield basalts of transitional chemistry and normal ⁸⁷Sr/⁸⁶Sr (0.705422–0.705435, Mg# = 67.2 and 76.3).

The 87 Sr/ 86 Sr isotopes in the basalts from all three ultra-enriched sampling localities are complemented by enriched (low) 143 Nd/ 144 Nd and the lowest 3 He/ 4 He ratios (4.31–4.93 Ra, or ratio to atmosphere) observed in Samoan basalts. Together, the new data extend the Samoan isotope array to a region outside the global ocean island basalt field (Fig. 1). Highly enriched EM2 signatures have previously been observed only in metasomatized xenoliths from Savai'i (67 Sr/ 86 Sr up to 0.712838; ref. 16), and the Samoan EM2 basalts provide the first evidence that the enriched component hosted in these xenoliths also occurs as erupted basalts. The enriched 87 Sr/ 86 Sr and 143 Nd/ 144 Nd isotope ratios, coupled with the low 3 He/ 4 He, are consistent with a recycled UCC component in the mantle source of the Samoan EM2 basalts.

The UCC reservoir exhibits several diagnostic trace-element characteristics that can be useful for detecting its presence in Samoan EM2 lavas. Compared to ocean island basalt and mid-ocean-ridge basalt lavas, UCC displays exceptional depletion in Nb (and Ta), Ti and Eu, and enrichment in Pb (Fig. 2). Samoan basalts have

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Figure 1 | ⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd isotope ratios of new enriched Samoan lavas. The values are compared with other Samoan shield basalts⁷, global ocean island basalt compositions and GLOSS (global subducting sediment). Analyses in which whole-rock (w) powders and dimopyroxene (c) analyses are performed on the same sample are connected by a tie-line. A model mixing line between depleted Ta'u peridotite and UCC is marked at 1% intervals, with increasing contribution from the latter component. The hypothetical UCC mixing endmember lies outside the figure. Approximately 5% UCC is required to produce the spidergram of sample D115-18 (see Supplementary Discussion), and ~6% is required to generate the ⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd in D115-21.

trace-element characteristics that are increasingly similar to UCC with more enriched ⁸⁷Sr¹⁸⁶Sr and ¹⁴⁵Nd/¹⁴⁴Nd values (Fig. 3). Although the most isotopically depleted basalts from Samoa show slight positive anomalies in Nb and Ti, the magnitude of these anomalies decreases monotonically towards the most enriched Samoan EM2 basalts. Similarly, a correlation exists between greater



Figure 2 | Primitive-mantle-normalized"' trace-element patterns for the Samoan EM2 endmember. The EM2 spidergram is the most isotopically enriched Samoan lava, sample D115-21 ($^{87}Sr)^{86}Sr = 0.720469$, Mg = 57). The other mantle endmembers (corrected to Mg # numbers of 60–62) and UCC⁷ are plotted for comparison. Similar to UCC, the Samoan EM2 lava exhibits large negative Ti and Nb (and Ta) anomalies and an excess of Pb (and K).

Pb enrichment and increasing isotopic enrichment in Samoan basalts. Importantly, the Eu anomaly is increasingly negative in the most isotopically enriched Samoan EM2 lavas (excluding basalts with MgO < 6.5 wt%), and the Rb/Sr and U/Pb are too low in the lower (or middle) continental crust¹⁷ to be consistent with the new Samoan Sr and Pb isotope data; these observations rule out the involvement of lower (or middle) continental crust. Furthermore, rare xenoliths with enriched ⁸⁷Sr/⁶⁶Sr and ¹⁴³Nd/¹⁴⁴Nd from the subcontinental lithospheric mantle⁸ suggest that this mantle domain can be isotopically enriched. However, the subcontinental lithospheric mantle does not appear to exhibit the trace-element anomalies observed in the most isotopically enriched Samoan lavas¹⁸. Instead, isotope ratios



Figure 3 87Sr/86Sr and 143Nd/144Nd ratios suggest the presence of a UCC component in Samoan EM2 lavas. The ⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd ratios are plotted against diagnostic UCC trace-element indicators in Samoan basalts: the more isotopically enriched Samoan basalts exhibit trace-element characteristics that are increasingly similar to UCC. UCC17 plots outside the panels, and its trace-element and isotopic composition is indicated by the level (and direction) of the arrows. a-c, All submarine Savai'i samples are plotted, as are other Samoan shield lavas with MgO > 6.5 wt%. d, Submarine Savai'i samples with MgO < 6.5% are excluded (avoiding possible effects of plagioclase fractionation). All trace-element data shown are by ICP-MS (inductively coupled plasma mass spectrometry). Element anomalies are calculated as follows (where subscript N means normalized as indexs (metric subscript is metric) in the matrix of the primitive mantle^2); Ti/Ti = Ti_N/ (Nd_N^{0.0555} \times Sm_N^{0.333} \times Gd_N^{0.722}); Nb/ Nb^{+} = Nb_N/ \sqrt{(Th_N \times La_N); Eu/Eu^{*} = Eu_N/} $(\mathrm{Sm}_{\mathrm{N}} \times \mathrm{Gd}_{\mathrm{N}}).$

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and trace-element anomalies (Nb, Ti, Eu and Pb) in Samoan basalts generate arrays that trend towards a composition similar to UCC.

We can exclude a shallow origin for the anomalous enrichment observed in the Samoan EM2 lavas. Owing to the close proximity of the Tonga trench, located only 120 km south of Savai'i, rapid cycling of sediment from the subduction zone into the Samoan plume was proposed as a mechanism for generating the extreme isotopic enrichment in Samoan lavas¹⁴. However, at the time that the submarine Savai'i lavas were erupted 5 Myr ago³¹, plate reconstructions indicate that the northern terminus of the Tonga trench was located 1,300 km to the west of Savai'i (ref. 19), and sediment input from the Tonga trench can be ruled out as a source of enrichment in these lavas.

Evidence from Pb isotopes suggests that it is unlikely that shallowlevel contamination by modern marine sediments is responsible for the isotopic enrichment in the Samoan EM2 basalts. In 7 Pb/ 204 Pb – Δ^{208} Pb/ 204 Pb isotope space, Samoan basalts and global marine sediments1 exhibit non-overlapping fields with diverging trends (Fig. 4). Moreover, three composite cores taken from the Samoan region, and a single ferromanganese crust from the flanks of Savai'i, plot in the global marine sediment field and exhibit no geochemical relationship with the extremely enriched Samoan lavas. It is also unlikely that the Samoan plume has been contaminated by stranded continental crust, such as was found beneath the Kerguelen plateau20 and the southern Mid-Atlantic Ridge21, or by ancient limestone blocks like those discovered in the Romanche fracture zone2 The tectonic history of the Samoan region places it neither at the locus of continental rifting, which was responsible for the marooned Kerguelen and southern Atlantic continental blocks, nor in proximity to any Pacific fracture zones23

Large quantities of sediment derived from UCC have entered the mantle at subduction zones over geologic time¹, and such a reservoir is ideally suited as an enriched source for the Samoan plume. The array formed by the Samoan EM2 basalts in ¹⁴⁵Nd/¹⁴¹Nd⁻⁸⁷Sr/⁸⁶Sr isotope space is anchored on the depleted end by basalts from



Figure 4 | ΔPb isotope compositions of Samoan lavas and marine sediment samples indicate that the Samoan EM2 lavas are not contaminated with modern marine sediment. Samoan basalts show no overlap with oceanic sediments contributing to GLOSS¹ (purple squares, where the GLOSS average composition is represented by a green crossed 'average' square), composite sections from three sediment cores taken in the Samoan region (blue circles), and a Samoan ferromanganese rind (blue diamond). Pbisotope data for mantle endmembers DMM (using data from mid-ocean ridge basalt normal segments, as defined and catalogued by ref. 28), EM1 (Pitcairn), and HIMU (Mangaia and Tubuai) are from the literature (see Supplementary Information for reference citations). The use of ΔPb isotope notation²⁹ to identify sediment components in ocean island basalts is discussed elsewhere¹⁰. All Samoan data shown are determined using a highprecision Tl-spike protocol⁷.

Ta'u, one of the youngest, easternmost Samoan islands. The $^{143}\rm Nd/^{144}\rm Nd-^{87}Sr/^{86}Sr$ array suggests mixing between this dominant, slightly depleted Ta'u component and a rare, enriched component that exhibits isotope and trace-element characteristics similar to UCC. The proportion of the enriched component in the Samoan EM2 lavas can be estimated by calculating trace-element concentrations in the depleted Ta'u mantle and mixing this composition with UCC (see Supplementary Discussion). A contribution of 5% UCC to the depleted Ta'u mantle generates a composition that, after mixing and melting, produces a trace-element pattern similar to that observed in Samoan EM2 sample D115-18 (with 87Sr/86Sr of 0.718592). Fixing the proportions of the depleted and UCC components in the Samoan EM2 source in this way then defines the $^{87}{\rm Sr}/^{46}{\rm Sr}$ and $^{143}{\rm Nd}/^{144}{\rm Nd}$ isotopic composition of this material as 0.7421 and 0.5117, respectively. The most isotopically enriched Samoan lavas have higher 87 Sr/ 86 Sr than the average continental crust inferred from suspended river sediments (87 Sr/ 86 Sr = 0.716; ref. 24) and global marine sediments (~0.717; ref. 1), values that are biased towards younger continental crust. However, composites of directly sampled ancient continental shield rocks show isotopically enriched compositions25 that bracket the calculated composition of the recycled UCC sediment in the Samoan mantle. The new ultra-enriched EM2 lavas suggest an unusually enriched recycled protolith in the Samoan mantle.

Despite the large volumes of sediment entering the mantle at subduction zones (estimated at 0.5-0.7 km³ yr⁻¹; ref. 1), isotopic signatures associated with recycled UCC are rare in ocean island basalts26. This enriched component is also uncommon in the Samoan plume, where the highly enriched Samoan EM2 lava D115-18 is calculated to have only 5% recycled UCC (and 95% depleted Ta'u source), and 90% of the remaining Samoan basalts exhibit depleted $^{143}\rm Nd/^{144}\rm Nd$ ratios (>0.512638). In addition to being rare in other ocean island basalts, recycled UCC may exist in low abundance in the Samoan plume. The reason for this may be that most subducted sediment melts and is rapidly returned to the surface in subduction zone volcanoes, or is simply scraped off onto the forearc and is never subducted. Alternatively, if a significant portion of UCC has been subducted over the past 4 Gyr ($0.5-0.7 \text{ km}^3 \text{ yr}^{-1}$) and has survived subduction zone melting, the resulting accumulated reservoir in the mantle will constitute only ~0.15% of its mass. Such a small reservoir may be diluted by the ambient mantle after convective stirring, a mechanism that efficiently attenuates mantle heterogeneities. Therefore, recycled crustal signatures can be greatly diluted and difficult to detect. By contrast, the recycled UCC component in the Samoan plume is an anomalous survivor in a chaotic mantle.

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Supplementary Information is linked to the online version of the paper at .nature.com/nature

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Author Contributions M.G.J. performed most of the experimental work, developed the model and wrote the paper. S.R.H. and H.S. conceived the project, and were co-chiels of the ALIA expedition. A.A.P.K. and J.K. were responsible for the cruise bathymetry, and A.A.P.K. greatly improved the figures. J.B. and M.K. provided analytical assistance and access to facilities. A.A.P.K., J.B., J.K. and J.A.R. helped with sample preparation. All authors participated in the discussion and interpretation of results, and commented on the manuscript.

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SUPPLEMENTARY INFORMATION

Contents of Supplementary Information (in a merged PDF titled jacksonsuppinfo.pdf; Size: 0.7 MB total):

- 1. Supplementary Discussion
- 2. Supplementary Figure 1 and caption
- 3. Supplementary Figure 2 and caption
- 4. Supplementary Figure 3 and caption
- 5. Supplementary Figure 4 and caption
- 6. Supplementary Table 1
- 7. Supplementary Table 2
- 8. Supplementary Table 3

The file contains a Supplementary Discussion that gives a model of the Samoan EM2 source, including calculations and assumptions. The file also contains a Supplementary Table 1 and a Supplementary Table 2 that together provide the model parameters used to calculate the depleted Ta'u and EM2 Samoan sources, respectively. The file also contains a Supplementary Table 3 that provides the new Samoan geochemical data. Four supplementary figures also are included in the file: Supplementary Figure 1 shows new helium isotope data from the remarkably enriched Samoan lavas; Supplementary Figure 2 is an expanded view of main text Figure 1, and shows the composition of the Samoan lavas in the context of individual samples from the upper continental crust; Supplementary Figure 3 compares the spidergram of a model melt of the EM2 source with the spidergram of an isotopically-enriched Samoan EM2 lava. Supplementary Figure 4 provides a model that explains the correlation between SiO₂ and ⁸⁷Sr/⁸⁶Sr in lavas from dredge ALIA D115.

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1. Supplementary Discussion

Recycled UCC in the Samoan EM2 mantle. The standard model for the genesis of the EM2 reservoir suggests a role for recycled sediment, probably terrigenous in origin¹⁻⁷. However, a number of problems with the standard model were outlined in ref. 8, and an alternative origin of the Samoan EM2 source (recycled, metasomatized lithosphere) was proposed. In light of the remarkably enriched Samoan lavas presented in this paper, we reconsider the arguments against a recycled terrigenous component in the EM2 plume. We demonstrate that recycled upper continental crust (UCC) can have an important role in the origin of the EM2 source. However, we maintain that it is (most likely) not "upper continental crust" (UCC) that is being recycled, but sediment with a composition resembling that of UCC. While the sediment recycling model presented here does not invalidate the metasomatic model for the lavas presented earlier⁸, the new lavas presented in this study are consistent with a recycled sediment component.

The problems with recycled marine sediment in the Samoan EM2 mantle source (as outlined in ref. 8) are as follows. The variation in ¹⁸⁷Os/¹⁸⁸Os in Samoan lavas is not consistent with mixing between recycled marine sediment and depleted MORB mantle (DMM). Such a scenario would require 35% marine sediment in the EM2 source, a quantity not observed in the trace element patterns of Samoan basalts. Additionally, the smooth trace element spidergrams observed in Samoan EM2 lavas were suggested to be inconsistent with a sediment component in the Samoan EM2 source. This is because, unlike Samoan EM2 lavas presented earlier⁸, marine sediments exhibit "jagged" spidergrams, marked by large negative anomalies for Nb (and Ta), Ti and Eu and large positive anomalies for Pb (and K). It was also noted that EM2 lavas exhibit negative Ba-

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anomalies, a feature not characteristic of marine sediments⁸. Furthermore, because EM2 lavas exhibit high ³He/⁴He ratios (~8 Ra, ratio to atmosphere, or ~1.4×10⁻⁶), a trait not shared with the low ³He/⁴He ratios in sediments (0.05-1 Ra)⁹ and continental crust (0.007 Ra)¹⁰, it was suggested that the Samoan EM2 source does not host a sediment component⁸. Finally, Pb-isotope composition of modern marine sediment was observed to be unsuitable as an endmember for the enriched Samoan basalts⁸: Modern marine sediments exhibit ²⁰⁸Pb/²⁰⁴Pb ratios that are too low at a given ²⁰⁶Pb/²⁰⁴Pb to serve as mixing endmembers for the Samoan EM2 lavas.

Below, we suggest that, instead of a marine sediment composition such as GLOSS (Global Subducting Sediment¹¹), sediment with a composition like UCC (a composition here approximated by UCC from ref. 12) is more suitable for generating a source sampled by the most isotopically-enriched Samoan lavas. A small portion (~5%) of (sediment with the composition of) UCC mixed with a depleted Samoan plume component generates a peridotite that, when melted, produces a spidergram similar to that observed in the most isotopically-enriched Samoan lavas. Like Samoan EM2 lavas, UCC exhibits a negative Ba anomaly, a feature not shared with GLOSS¹¹. Furthermore, published Os-isotopes⁸ in enriched Samoan EM2 lavas are not inconsistent with a UCC component in the plume. The Os-isotope signature in Samoan basalts is not likely a result of mixing between DMM (¹⁸⁷Os/¹⁸⁸Os = 0.125) and marine sediment; however, a mixture of UCC (¹⁸⁷Os/¹⁸⁸Os = 1.05, [Os] = 30 ppt¹³) and a depleted Samoan plume component generates a mixing trend that describes the Os-isotope data for Samoan shield basalts (excluding samples with <100 ppt Os, and assuming measurement precision of ±1.5% in data from ref. 8). This mixing scenario uses measured Ta'u ¹⁸⁷Os/¹⁸⁸Os ratios

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of 0.129⁸ and assumes that the Ta'u peridotite source also has Os concentrations that are depleted relative to primitive mantle (>3,900 ppt¹⁴). Moreover, new helium isotope data for Samoan EM2 lavas exhibit low ratios (<5 Ra) that plot on a trajectory that trends to UCC, and are consistent with the existence of a recycled UCC component in the Samoan plume (Supplementary Figure 1). Finally, our model for the formation of the EM2 mantle agrees with Pb-isotope constraints. UCC exhibits a large range of Pb-isotope values, and the Samoan EM2 lavas exhibit Pb-isotope values that plot in the field previously defined for UCC¹⁵.

Model for the EM2 source. In order to model the generation of the mantle source sampled by the new Samoan EM2 lavas, we take advantage of the array formed by Samoan basalts in ⁸⁷Sr/⁸⁶Sr-¹⁴³Nd/¹⁴⁴Nd isotope space. The array suggests that the most enriched Samoan basalts were formed as products of mantle-mixing between a depleted component (here represented by the mantle source sampled by the isotopically-depleted and remarkably homogeneous lavas from Ta'u island) and a component similar to UCC (Supplementary Figure 2). In order to determine the proportion of UCC in the Samoan plume, we first calculate a trace element peridotite source for the depleted Ta'u endmember that is consistent with the radiogenic isotopes of Nd, Hf, Sr and Pb. We then determine the amount of UCC that must be added to the depleted Ta'u peridotite source so that the final mixture, an enriched peridotite, can be melted to generate a spidergram similar to the new Samoan EM2 lavas.

In the following modeling exercise, we generate a model for the EM2 source that describes the array formed by the most isotopically-enriched submarine Savai'i lavas from dredge D115, and we make no attempt to model the other components previously

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identified⁸ in lavas from the Samoan hotspot. We emphasize that the model is just one possible model that is consistent with the isotopes and trace elements in Samoan EM2 lavas. The model is presented only to demonstrate that recycling sediment into Samoan EM2 lavas is possible.

Composition of the depleted Ta'u peridotite source. Following ref. 8, we generate an average olivine fractionation corrected trace element budget for Ta'u lavas (Supplementary Figure 3 and Supplementary Table 1). In order to determine a trace element mantle source sampled by the average Ta'u lava composition, we first assume a peridotite source lithology, an aggregated fractional melting model, and we adopt mineral-melt partition coefficients from ref. 16 (Supplementary Table 1). We also assume a two-stage isotope model for the evolution of the depleted Ta'u source, and that this differentiation event of a primitive mantle composition occurred at 1.8 Ga. This age is commonly quoted as the average mantle differentiation age¹⁷, and is an age that is consistent with the array formed by Ta'u lavas in ²⁰⁷Pb/²⁰⁴Pb vs. ²⁰⁶Pb/²⁰⁴Pb isotope space (see Fig. 7 in ref. 8). Given these assumptions, a Ta'u source is calculated so that parentdaughter ratios-Sm/Nd and Lu/Hf-will generate the present-day depleted ¹⁴³Nd/¹⁴⁴Nd (Ta'u average is 0.512789⁸) and ¹⁷⁶Hf/¹⁷⁷Hf (Ta'u average is 0.282987¹⁸) isotopes measured in Ta'u lavas given the two-stage isotope model. The Ta'u source is not very sensitive to its age of formation: ages of 1.0-2.5 Ga require only small variations in melting-4.5 to 5.5%-to generate the average Ta'u lava. However, assuming a formation age of 1.8 Ga, a 5.1% melt (with ~50.6% garnet melting) of the hypothetical Ta'u source will generate a model melt with a spidergram that is both identical to the average measured Ta'u lava and consistent with isotopic constraints. These highly

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specific melting parameters are presented only to generate an average Ta'u source that is consistent with isotopic constraints, and by presenting them we are not suggesting such precise knowledge of the actual mantle "plumbing" beneath Ta'u island.

The ⁸⁷Sr/⁸⁶Sr of primitive mantle is unconstrained, so it is not possible to evaluate whether the ⁸⁷Sr/⁸⁶Sr measured in Ta'u lavas is enriched or depleted relative to primitive mantle. However, the average 87 Sr/86 Sr in Ta'u lavas (0.704650) and the Rb/Sr of the Ta'u source (0.0268) can be modeled as having evolved from primitive mantle at 1.8 Ga if the present-day primitive mantle ⁸⁷Sr/⁸⁶Sr is 0.70508, a value that is in the range typically assigned to primitive mantle. In order for the two-stage isotope model to produce the observed average Ta'u 206Pb/204Pb (19.271) and 206Pb/204Pb (15.597) (excluding T14 and considering only TI-spiked data⁸), the proportion of sulfide¹⁹ in the Ta'u source mineralogy is adjusted to obtain an appropriate parent-daughter U/Pb source ratio. The resulting calculated Th/Pb source ratio is within error of the Th/Pb ratio required to produce the average Ta'u 208 Pb/204 Pb composition (39.424) in 1.8 Ga. While non-unique (a different melt model could be chosen or the Ta'u source formation age may be different, etc.), the trace element source calculated from Ta'u lavas is consistent (within the uncertainties of the data) with constraints from radiogenic isotopes. Generation of a source for the Samoan EM2 lavas. Determining the precise nature of the recycled component contributing to the enrichment in the Samoan plume is not straightforward. UCC rocks and individual marine sediment cores show a large degree of trace element heterogeneity that varies considerably with geography and provenance^{11,12}, and the composition of a recycled sediment may depend on the geography of the subduction zone. More problematic is the issue of temporal variability^{20,21}: the trace

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SUPPLEMENTARY INFORMATION

element budgets of sediments encountered in modern oceans may be poor analogues for the sediments subducted in the past. Finally, the poorly constrained processes operating in subduction zones, including fluid loss and/or melting of sediments, may modify the composition of the subducted sediment component²²⁻²⁵.

With these caveats aside, we take a simple approach. We assume that the compositions of subducted material are conserved in the subduction zone. Perhaps fluids and melts are removed from the slab and inoculated into the mantle wedge; but the resulting slab residue and the fertilized mantle wedge may in some cases stay together as a package. This package is what enters the general circulation, to eventually be remixed to end up looking like "closed system" slab recycling. While the composition of the original material sent into the subduction zone is underconstrained, we explore whether or not a modern UCC composition, when mixed with the depleted Ta'u source, can generate a suitable source for the Samoan EM2 lava D115-18. We assume that after the depleted Ta'u source mixed completely with the enriched UCC component, the resulting peridotite (treated here as a single lithology) was melted in a modal, aggregated fractional melting system. The modal abundances of the mantle phases, the contribution of UCC to the depleted Ta'u source, and the degree of melting (and the proportion of garnet and spinel melting) of the resulting mixture (the EM2 source) are all adjusted to generate a trace element spidergram that is similar to the enriched Samoan EM2 basalts. The combination of these parameters that generates a "best-fit" spidergram to the Samoan EM2 lava composition is given in Supplementary Table 2. The agreement between the model spidergram and the spidergram for Samoan EM2 lava D115-18 is optimized (the

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fit is within 14% for all the trace elements considered, or \sim 11% if U is excluded) when the contribution of the UCC component is \sim 5%.

Having fixed the proportions of UCC (5%) and the depleted Ta'u source (95%) in the EM2 source sampled by lava D115-18 (with ⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd of 0.718592 and 0.512314, respectively), we use trace element budgets and isotopic constraints to calculate the ⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd of the UCC component. Assuming a measured isotopic (⁸⁷Sr/⁸⁶Sr = 0.704650 and ¹⁴³Nd/¹⁴⁴Nd = 0.512789) and calculated trace element composition (Supplementary Table 1) of the Ta'u source and a trace element composition of UCC¹², we calculate the Sr and Nd isotopic composition of the UCC component in the Samoan plume to be 0.7421 and 0.5117, respectively. The ⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd values calculated for the UCC endmember in the Samoan plume are within the range of values measured in ancient UCC shield rocks²⁶ (Supplementary Figure 2).

While the addition of UCC to the depleted Ta'u source generates a spidergram that is a close match to Samoan EM2 lavas, the fit is not perfect. In particular, the element that exhibits the least perfect fit to the data is U. Samoan EM2 lavas have lower U concentrations than the model result, and the disagreement may be a result of U-loss during weathering (note the high Th/U in sample D115-18; Supplementary Table 3). Addressing contamination by marine sediment. Implicit in the model of the EM2 source is the assumption that the sediment signature in the Samoan EM2 lavas is a primary mantle signal and not a result of shallow-level sediment contamination. A plot of Δ^{207} Pb/²⁰⁴Pb- Δ^{208} Pb/²⁰⁴Pb (see Main Text Figure 4) indicates that sediment contamination is not an issue. Further evidence comes from ⁸⁷Sr/⁸⁶Sr measurements in sediments from the Samoan region, which exhibit ⁸⁷Sr/⁸⁶Sr ratios (0.70614 to 0.70824)

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that are much lower than the most enriched Samoan EM2 lavas (0.72047 in whole-rock sample D115-21). It is very difficult to generate the high ⁸⁷Sr/⁸⁶Sr observed in the new Samoan EM2 lavas by sediment contamination. For example, the GLOSS average ⁸⁷Sr/⁸⁶Sr composition is only 0.7173 (327 ppm Sr)¹¹, a value that is much lower than the most isotopically-enriched lavas from Samoa. Even if the most isotopically-enriched marine sediment (⁸⁷Sr/⁸⁶Sr = 0.73493, 251 ppm Sr) in the compilation from ref. 11 were added to the least isotopically-enriched submarine Savai'i lava (⁸⁷Sr/⁸⁶Sr = 0.705435, 374 ppm Sr), over 60% sediment assimilation would be required to generate the most radiogenic ⁸⁷Sr/⁸⁶Sr observed in the Samoan lavas. Such large quantities of sediment are not visible in the trace element spidergrams of the most enriched Samoan lavas.

In general, Samoan lavas do not show a correlation between ⁸⁷Sr/⁸⁶Sr and SiO₂ (Supplementary Figure 4). However, ⁸⁷Sr/⁸⁶Sr ratios in whole-rock lavas from dredge D115 do correlate with SiO₂. Having ruled out sediment assimilation using Sr and Pb isotope data, the ⁸⁷Sr/⁸⁶Sr – SiO₂ array formed by ALIA D115 samples is interpreted to be a result of magma mixing between an evolved (high SiO₂), isotopically-enriched magma and a less evolved, less isotopically-enriched magma. Other major and trace element data are consistent with this scenario (see Supplementary Figure 4). Perhaps the isotopically-enriched magmatic endmember evolved by crystal fractionation in a magma chamber and, just before eruption, mixed with a later pulse of a less evolved, less isotopically-enriched magma. Magma mixing is not an uncommon phenomenon; heterogeneous ⁸⁷Sr/⁸⁶Sr ratios recorded in olivine-hosted melt inclusions from individual Samoan basalt samples suggest that mixing of magmas from isotopically-distinct sources is not uncommon in Samoa²⁷.

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Data sources for HIMU and EM1 lavas. In Fig. 2 and Fig. 4 of the main text, trace element and Pb-isotope data for the HIMU endmember are from samples collected at Mangaia and Tubuai islands and are reported in refs. 41 and 42. Trace element and Pb-isotope data from the EM1 endmember are from samples collected at Pitcairn and include data from ref. 43 and unpublished data (S.R. Hart and E.H. Hauri). Only data from the freshest, most isotopically-extreme samples were used.

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2. Supplementary Figures



Supplementary Figure 1 87 Sr/ 86 Sr vs. 3 He/ 4 He isotopes in Samoan basalts. New 3 He/ 4 He ratios in the enriched Samoan EM2 basalts are low (<5 Ra), and plot on a trajectory that extrapolates to a UCC component with low 3 He/ 4 He (0.007 Ra) 10 and high 87 Sr/ 86 Sr (0.7421). Data for the other Samoan islands and seamounts are from refs 8 and 40.

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Supplementary Figure 2⁸⁷Sr/⁸⁶Sr vs. ¹⁴³Nd/¹⁴⁴Nd of Samoan shield basalts compared with samples from upper continental crust (UCC) from ref. 26. The grey field (with solid outline) encompasses Sr and Nd isotopic analyses on UCC samples²⁶, and extends outside of the figure to values of 1.1862 and 0.51023, respectively. The grey field (with dashed outline) encompasses the non-Samoan OIB field. The orange field describes Samoan shield basalts observed in previous studies; the new Samoan EM2 basalts (red) extend well into the field for UCC. Tic marks represent addition of the UCC component to the depleted peridotite Ta'u source. The isotopic composition of the UCC endmember in the Samoan plume is calculated to have ⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd of 0.7421 and 0.5117, respectively (light blue box, see Supplementary Discussion), and lies in the range of values previously measured on UCC rocks²⁶. 5% of this hypothetical UCC composition is required to generate the spidergram of Samoan sample D115-18 (and an estimated 6% UCC is required to generate the ⁸⁷Sr/⁸⁶Sr ratios in the most enriched cpx from sample D115-21).

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Supplementary Figure 3 Primitive mantle²⁸ normalized trace element patterns for Samoan basalts and peridotite sources. These compositions are used in a sediment recycling model for generating a Samoan EM2 mantle source. Ta'u lavas are assumed to be melts of a depleted plume component, and the average, olivine-fractionation corrected Ta'u lava composition is plotted. A Ta'u source for the average Ta'u lava is calculated to satisfy isotopic constraints. 5% of UCC¹² is added to 95% of the Ta'u source to make the EM2 source sampled by the Samoan lava D115-18, and the model melt of the EM2 peridotite source is plotted. The spidergram of the model melt is similar to sample D115-18 (the second most isotopically-enriched Samoan sample). The most isotopicallyenriched lava, D115-21, is too evolved to reliably reconstruct its trace element composition. Plotted compositions can be found in Supplementary Tables 1 & 2.

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Supplementary Figure 4 A model for the co-variation of SiO₂ and ⁸⁷Sr/⁸⁶Sr in ALIA D115 samples. Samoan lavas do not exhibit a one-to-one relationship between SiO₂ (or Mg#) and ⁸⁷Sr/⁸⁶Sr; evolved Samoan lavas exhibit both high (up to 0.7205) and low (0.7044) 87 Sr/86 Sr ratios. However, lavas from dredge ALIA D115 exhibit a correlation between ⁸⁷Sr/⁸⁶Sr and SiO₂. We explain this correlation in ALIA D115 lavas as a result of mixing between a highly evolved, isotopically-enriched magma and a less evolved, less isotopically-enriched magma. In this scenario, the mixing endmembers are a Savai'i submarine lava sample (ALIA D114-01; Supplementary Table 3) and a highly evolved lava from Tutuila (91TP-128; SiO₂ = 66.30 wt.%, MgO = 0.41 wt.%, FeO = 4.06 wt.%, CaO = 1.62 wt.%, Al₂O₃ = 17.03 wt.%, Sr = 395 ppm; Natland unpubl. data). The measured ⁸⁷Sr/86Sr of the Tutuila lava is 0.7055358. However, for the sake of argument, this sample is given an ⁸⁷Sr/⁸⁶Sr ratio of 0.730 in the mixing model in order to determine whether mixing between an evolved magma and a less evolved magma can generate a mixing array consistent with the data from dredge ALIA D115. Sc was not measured on the Tutuila lava, but its concentration is assumed to be 4 ppm. The reduction of Sc and CaO/Al₂O₃ (including reduced Mg#'s and increased SiO₂) in the most isotopically-enriched lavas from ALIA D115 is consistent with cpx fractionation. Crystal fractionation of an isotopically-enriched magma followed by mixing with a less evolved, less isotopically-enriched magma describes the geochemical data from dredge ALIA D115.

3. Supplementary Tables

Supplementary Table 1. Calculation of the depleted component in the Samoan plume, using

	Primitive	Avg. Ta'u lava"	Variability in 18 Ta'u lavas	Bulk partition coefs	Bulk partition coefs	Depleted Ta'u source"
	mantle	(olivine corr'd to Fogo)	(1o, std. dev., %)	(garnet field)	(spinel field)	(calculated from avg. Ta'u lava)
Rb	0.6	14.9	16.5	0.000144	0.000124	0.759
Ba	6.6	140	11.9	0.000144	0.000124	7.14
Th	0.0795	2.71	13.4	0.00219	0.00178	0.138 (0.117)5
U	0.0203	0.61	12.7	0.00074	0.00018	0.0309
Nb	0.658	28.53	11.5	0.00393	0.00302	1.45
Та	0.037	1.97	11.7	0.00393	0.00302	0.10
к	240	5860	9.8	0.00150	0.00128	299
La	0.648	23.27	11.0	0.0114	0.00983	1.20
Ce	1.675	48.93	10.3	0.0191	0.0161	2.64
Pb	0.15	1.49	18.4	0.0834	0.0816	0.161
Pr	0.254	6.09	10.0	0.0314	0.0247	0.368
Nd	1.25	26.58	9.1	0.0463	0.0357	1.88
Sr	19.9	354.1	9.9	0.0665	0.0364	28.3
Zr	10.5	184	9.9	0.0867	0.0371	16.5
Hf	0.283	4.74	9.0	0.114	0.0607	0.538
Sm	0.406	6.64	7.8	0.0868	0.0574	0.657
Eu	0.154	2.10	7.9	0.121	0.0671	0.253
Ti	1205	17190	7.4	0.176	0.122	2960
Gd	0.544	6.32	7.8	0.179	0.0769	0.965
Tb	0.099	0.94	8.1	0.254	0.0876	0.18
Dy	0.674	5.17	8.2	0.316	0.0964	1.18
Ho	0.149	0.93	8.1	0.402	0.0972	0.25
Y	4.30	23.86	7.3	0.475	0.102	7.40
Er	0.438	2.24	7.3	0.481	0.105	0.70
Tm	0.068	0.29	7.2	0.673	0.110	0.12
Yb	0.441	1.65	7.1	0.865	0.115	0.84
Lu	0.0675	0.24	7.5	1.09	0.124	0.148

¹Primitive mantle is from ref 28.

²Ta'u basalts (with MgO > 6.5 wt.%) used in the average lava composition include: T10, T16, T19, T22, T23, T25, T30, T32, T33, T44, T45, T46, T47, T48, T51, T54, T55, and 74-1. T14 is suspected of Pb contamination and is not included. The 18 Ta'u basalts were individually corrected for olivine fractionation to be in equilibrium with a mantle olivine composition of Fo₈₀ (assuming Fe²/Fe_{tural} is 0.90), trace element corrected, and then averaged.

mantle olivine composition of Pos (assuming Fe⁻/Pe_{strat} is 0.90), trace element corrected, and then averaged. ³Bulk partition coefficients are calculated by assuming the following modal abundances: spinel stability field: 3.72% spinel, 17.8% dinopyroxene, 26% orthopyroxene, 52.39% olivine, 0.162% sulfde; gamet stability field: 10.7% gamet, 20.8% dinopyroxene, 13.4% orthopyroxene, 55.1% olivine, 0.161% sulfde. Mineral-melt partition coefficients are from ref. 16, with the following additions and modifications: Ba and Ta are assumed to have the same bulk partition coefficients as Rb and Nb, respectively; Tm and Sr are assumed to have a bulk partition coefficient that is the average of the elements that bracket them on the spidergram; the sulfde-melt partition coefficient for Pb is assumed to be 42. The mineral modes in the gamet and spinel stability fields are calculated to agree with a primitive mantle modes in the gamet and spinel stability failed as re calculated to agree with a primitive mantle modes in the gamet and spinel stability failed are calculated to agree with a primitive mantle major element composition. Although the Ta'u source is slightly depleted, and therefore has suffered a small amount of melt extraction during its history, the low degree of melt extraction will not significantly change the calculated mineral modes. The Ta'u cause assumed a part and spine factional melt with a 50.6% melt contribution.

The Ta's uscreasing and the the average, olivine-fractionation corrected Ta's lava resulted from a 5.1% modal aggregated fractional melt with a 50.6% melt contribution from the garnet stability field (and the remainder from the spinel stability field). This precise set of melting parameters allows the parent-daughter ratios (LU/Hf and Sm/Nd) in the Ta's source to be consistent with the ¹²⁶Hf¹¹⁷Hf and ¹⁴³Nd/¹⁴⁴Nd ratios measured in the Ta's lavas (given a two-stage evolution model, with differentiation from primitive mantle at 1.8 Ga).

²⁰⁷ he Th concentration in parenthesis, which yields a Th/Pb ratio of 0.727, provides a value required for the Ta'u source to be consistent with constraints from ²⁰⁷ bb²⁰⁴ Pb isotopes measured in Ta'u lavas. The other Th value for the Ta'u source is calculated by adjusting the modal sulfide abundance to give U/Pb ratios that are consistent with ²⁰⁶ Pb/²⁰⁴ Pb and ²⁰⁷ Pb/²⁰⁴ Pb. This Th svalue (0.138) generates a Th/Pb ratio (0.855) that is within error (18%) of the measured Th/Pb variability (2323) in Ta'u lavas, and is therefore (within uncertainty of the data available on Ta'u lavas) consistent with measured ²⁰⁶ Pb/²⁰⁴ Pb ratios.

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3. Supplementary Tables

Supplementary Table 1. Calculation of the depleted component in the Sam	noan plume, using lavas from Ta'u island
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	Primitive	Avg. Ta'u lava*	Variability in 18 Ta'u lavas	Bulk partition coefs"	Bulk partition coefs"	Depleted Ta'u source"
	mantle	(olivine corr'd to Fogo)	(1o, std. dev., %)	(garnet field)	(spinel field)	(calculated from avg. Ta'u lava)
Rb	0.6	14.9	16.5	0.000144	0.000124	0.759
Ba	6.6	140	11.9	0.000144	0.000124	7.14
Th	0.0795	2.71	13.4	0.00219	0.00178	0.138 (0.117)5
U	0.0203	0.61	12.7	0.00074	0.00018	0.0309
Nb	0.658	28.53	11.5	0.00393	0.00302	1.45
Та	0.037	1.97	11.7	0.00393	0.00302	0.10
к	240	5860	9.8	0.00150	0.00128	299
La	0.648	23.27	11.0	0.0114	0.00983	1.20
Ce	1.675	48.93	10.3	0.0191	0.0161	2.64
Pb	0.15	1.49	18.4	0.0834	0.0816	0.161
Pr	0.254	6.09	10.0	0.0314	0.0247	0.368
Nd	1.25	26.58	9.1	0.0463	0.0357	1.88
Sr	19.9	354.1	9.9	0.0665	0.0364	28.3
Zr	10.5	184	9.9	0.0867	0.0371	16.5
Hf	0.283	4.74	9.0	0.114	0.0607	0.538
Sm	0.406	6.64	7.8	0.0868	0.0574	0.657
Eu	0.154	2.10	7.9	0.121	0.0671	0.253
Ti	1205	17190	7.4	0.176	0.122	2960
Gd	0.544	6.32	7.8	0.179	0.0769	0.965
Tb	0.099	0.94	8.1	0.254	0.0876	0.18
Dy	0.674	5.17	8.2	0.316	0.0964	1.18
Ho	0.149	0.93	8.1	0.402	0.0972	0.25
Y	4.30	23.86	7.3	0.475	0.102	7.40
Er	0.438	2.24	7.3	0.481	0.105	0.70
Tm	0.068	0.29	7.2	0.673	0.110	0.12
Yb	0.441	1.65	7.1	0.865	0.115	0.84
Lu	0.0675	0.24	7.5	1.09	0.124	0.148

¹Primitive mantle is from ref 28.

²Ta'u basalts (with MgO > 6.5 wt.%) used in the average lava composition include: T10, T16, T19, T22, T23, T25, T30, T32, T33, T44, T45, T46, T47, T48, T51, T54, T55, and 74-1. T14 is suspected of Pb contamination and is not included. The 18 Ta'u basalts were individually corrected for olivine fractionation to be in equilibrium with a mantle olivine composition of Fo₈₀ (assuming Fe²/Fe_{tural} is 0.90), trace element corrected, and then averaged.

mantle olivine composition of Pos (assuming Fe⁻/Pe_{strat} is 0.90), trace element corrected, and then averaged. ³Bulk partition coefficients are calculated by assuming the following modal abundances: spinel stability field: 3.72% spinel, 17.8% dinopyroxene, 26% orthopyroxene, 52.39% olivine, 0.162% sulfde; gamet stability field: 10.7% gamet, 20.8% dinopyroxene, 13.4% orthopyroxene, 55.1% olivine, 0.161% sulfde. Mineral-melt partition coefficients are from ref. 16, with the following additions and modifications: Ba and Ta are assumed to have the same bulk partition coefficients as Rb and Nb, respectively; Tm and Sr are assumed to have a bulk partition coefficient that is the average of the elements that bracket them on the spidergram; the sulfde-melt partition coefficient for Pb is assumed to be 42. The mineral modes in the gamet and spinel stability fields are calculated to agree with a primitive mantle modes in the gamet and spinel stability failed as re calculated to agree with a primitive mantle modes in the gamet and spinel stability failed are calculated to agree with a primitive mantle major element composition. Although the Ta'u source is slightly depleted, and therefore has suffered a small amount of melt extraction during its history, the low degree of melt extraction will not significantly change the calculated mineral modes. The Ta'u cause assumed a part and spine factional melt with a 50.6% melt contribution.

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²⁰⁷ he Th concentration in parenthesis, which yields a Th/Pb ratio of 0.727, provides a value required for the Ta'u source to be consistent with constraints from ²⁰⁷ bb²⁰⁴ Pb isotopes measured in Ta'u lavas. The other Th value for the Ta'u source is calculated by adjusting the modal sulfide abundance to give U/Pb ratios that are consistent with ²⁰⁶ Pb/²⁰⁴ Pb and ²⁰⁷ Pb/²⁰⁴ Pb. This Th svalue (0.138) generates a Th/Pb ratio (0.855) that is within error (18%) of the measured Th/Pb variability (2323) in Ta'u lavas, and is therefore (within uncertainty of the data available on Ta'u lavas) consistent with measured ²⁰⁶ Pb/²⁰⁴ Pb ratios.

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BHVO-1																	1	0.10	1.6	130	127	0.43	17.2	1.17	0.01	2.03	5.44	24.6	385	167	444	2.08	619	0.93	533	1.00	20.4	033	1.90	0.27	31.9							T	
WF2																	3	0.89	20.4	171	2.15	0.51	6.84	0.46	R D C C C	171	3.06	13.0	192	83	2.56	108	3.58	0.61	3.86	0.82		032	2.01	0.31	35.6								
1118-22C/URT 13.325 173.430 2660 Fe-Mn crust																		0.15	8.1	1761	9.95	8.08	157.4	939	1.011	0.00.2	16.24	69.48	1062	842	16.41	4.10	17.03	2.89	20.42	4.74	113.60	2.08	13.18	2.07	2.11.7				18.764	15.629	38.757	ude ralues	102) (2)
HC-12-207 14.617 168.5 5037 5037 sediment																		0.92	27.1	276	3.61	0.74	24.84	1 89	24.95	7.87	6.25	26.72	382	167	4 20	21.0	6 33	0.98	5.58	1.03	78.07	1034	1.98	0.30	26.3	0.706140			18 932	15.628	38.888	ole. exceptions incl entrations. Dri 18, Ba/Rb	enfored samp site (D115-2 b and Cs
VM-18-253 12.233 169 5286 sediment																	2	1.05	32.6	360	4.64	1.07	40.83	2.89	25 38	21.00	8.26	34.92	690	208	11.6	979	7.92	1.16	6.35	1.16	25.12	2.0.2	2.10	0.32	24.0	206202.0			18.849	15.624	38.849	ind in the tal ref. 8. The I U and U con from dredge	in the most ighest Rb/C le than U, R
RC 13-43* 12:25 171:683 4872 sediment	26.67	3.85	2.23	8.28	0.23	12.81	6.86	1.26	3.78	0.31	77 18	18 48	166	203	15	98	108	0.79	26.1	369	4.00	0.99	34.75	2.56	01.67	8 29	6.87	29.08	209	168	4 22	R9.0	6.55	0.98	5.43	0.99	20.07	0.32	1.87	0.28	21.7	0.708247			18.844	15.624	38.850	nanganese r reported in between Th	salmo rativo le with the h or less mobi
D128-21 13.213 172.093 2560 Basaft	40.05	1211	2.28	9.67	21'0	3.81	11.65	52	21.2	0.75	57.8R	1.43	286	190	10	4	115	0.41	28.4	240	4.29	0.72	31.26	2.17	07.75	4 77	8.55	34.67	363	179	5.02	11 C	6.86	1.02	5.49	1.03	81.67	1.41	1.98	0.28	28.9	0.712600	0 610460	004710-0	18.963	15.628	38.273	I the ferron nal basalts bon exists 15-07 and	a the samp the samp ts equally
D128-19 13.213 172.083 2560 Basatt	A0.22	12.61	2.85	9.91	0.16	10.54	7.88	1,56	2.40	98.0	10.05	3,05	106	180	19	82	114	0.56	42.3	328	6.28	1.31	48.45	3.33	40.44	5 13	11 32	44.23	481	242	6.36	0.80	7.76	1.14	6.26	1.13	27.00	0.36	2.11	0.32	31.0	0.708170			19.151	15.640	39.465	e lavas and and subaer gh correlat	ere consure Excluding
D128-17 13.213 172.093 2560 Basalt	6778	12.07	3.09	11.31	71.0	10.60	10.10	1.33	2.15	0.35	78.85	0.75	202	319	21	59	116	0.4	32.4	257	521	1.10	50.79	3.47	41.65	403	9.32	37.85	466	227	5.88	17.8	7.35	1.10	6.17	111	70.07	0.37	2.09	0.31	31.7	0.706397	0 510650	200710-0	19,150	15.630	39.418	ale ID of the submarine sples, a rou Excluding :	om ref. 29 suggest t
D128-01 13.213 172.093 2560 Basalt	A0.0A	12.68	2.90	9.73	0.17	10.68	822	1.58	241	15.0	80.78	2.00	142	294	20	56	108	0.59	44.2	345	6.53	1 38	48.97	3.43	26.14 00.00	537	11.60	44.89	496	248	6.38	P T C	781	41.1	6.27	1.14	58.17	037	2.14	0.32	31.2	0.708117			19.150	15.639	39.465	n the samp d in other s rese 4 sam	of B5-95 fe and RbVCs
D118-23 13.325 173.430 2660 Alkali basalt	CC CV	7.81	2.21	10.08	0.11	7.26	16.74	0.99	197	0.48	10.01	8.97	528	169	4	45	66	4.19	28.5	701	12.83	1.56	74.63	4.55	09.0R	0.771	18.42	87.46	317	243	6.37	312	894	1.16	5.83	0.98	DR.CT	0.28	1.47	0.22	21.9	0.711808	0.710337	0.512453	18.967	15.643	39.738	4.71 lefeted from 5) observes colluding th tos greate	ical range U, Ba/Rb
D118-22 13.325 173.430 2660 Alkai basalt	20.71	742	3.20	10.82	0.15	13.68	17.13	132	2/1	190	18.48	4.72	472	283	4	47	96	0.61	44.5	477	14.04	1.97	98.75	6.64	16.85	242	22.12	82.00	627	260	6.94	18.61	10.15	1.27	6.27	1.05	76.07	0.28	1.55	0.22	28.5	0.707902	1120120	1107100	19.557	15.657	40.285	has been o ge (4.5±1. U loss). E e Ba/Rb ra	In high source of the canon (The endine)
D118-13 13.325 173.430 2660 Alicali basalt	20.02	7.66	3.46	12.06	0.19	12.70	16.63	101	RG1	0.81	18.58	3.63	478	280	12	02	08	0.33	33.2	216	15.45	1.48	107.52	7.18	10.611	2.802	24.97	92.59	515	284	1.34	4 16	11.40	1.45	7.28	1.24	01.05	0.36	1.95	0.30	31.0		0120120	7107100	19.557	15.658	40.290	efix "ALLA" e same rar suggesting 18-23) hav	t is dose to boies of all
D115-28 14.090 172.898 32.20 Hawaite	60.00	13.44	3.30	7,86	0.12	9.62	6.37	2.73	0.10	0.52	01.78	2.51	180	040	24	25	103	0.66	81.4	510	8.83	1.77	16.43	3.61	10.10	11 81	13.76	52.43	645	BIB	88.7	3 14	9.01	1.24	6.65	1.15	00 0	80.7	1.89	0.27	24.4	0.703985	012020	0/07100	19.034	15.634	39.419	na. The pro ally fall in the 1.1 to 10.4.	positive con ue of 96 that ise three pr
D115-26 14.090 172.898 3220 Trachy- basak	61 00	14.26	2.71	7.46	0.09	1.27	8	352	3.50	142	07 08	4.42	108	100	27	23	150	0.58	83.8	989	10.31	2.10	47.73	3.23	PC'RC	8 84	14.39	55.43	528	320	8.26	67 LL	9.20	1,28	99.9	1.15	B/ Br	0.33	1.87	0.27	19.8	0.716394	0000120	2007100	19.009	15.644	39.452	/ Kilo Moa lies genera nge from 7 ues (D115	ons ano a verage val opes. The
D115-22 14.090 172.898 3220 Basalt	47.20	12.84	3.43	8.65	0.12	14.1	7.47	2.31	11.5	0.52	11 12	5.56	145	258	22	52	142	0.28	57.4	466	8.26	1.67	52.26	3.50	78.60	14.2	13.43	51,35	583	302	1.60	24.01	8.89	1.24	6.56	1.12	010	0.33	1.82	0.26	19.6	0.709450	1000120	100710-0	19.070	15.637	39.419	e of the R/ noan samy ere Th/U ra	concerte au , with an a and Sr-isol
D115-21 14.090 172.898 3220 Trachy- andeste	67.00	14.63	1.72	5.35	0.05	3.21	3.61	6.05	3.68	131	78.55	4.88	0 0	88	27	17	28	1.07	155.6	683	17.46	2.89	48.14	3.08	61.33 • FAE	13.58	16.97	58.94	478	380	18.8	2 82	7.19	0.97	4.86	0.82	20.12	0.24	1.38	0.20	8.3	0.720469	0.721630	1077 10 0	18.954	15.646	39.424	ALIA cruis he new Sar D118 (wh th the high	e of 66-200 entrations
D115-20 14.090 172.898 3220 Hawaite	40.06	13.16	3.55	9.13	0.13	7.89	3 I 9	2.57	3.32	0.61	80.33	4 27	181	274	25	56	122	0.81	74.0	488	10,13	R N	58.73	3.95	10.40	6.47 6.47	15.17	58.29	662	358	18.8	3.50	8.82	1.41	7.38	1.26	20.03	787	1.99	0.27	18.8	0.708175	100010	100710-0	19.063	15.629	39.396	g the 2005 e Th/U of t om dredge samples w	e correran In the rang th Rb cont
D115-18 rep ¹ 14.090 17.2.898 3220																																										0.718583	0.212000	0707100				ari island durin quite good. Th rree samples fr loss. The two	oceitent negam b/Cs ratios fall correlations w
Di 15-18 14.090 172.898 3220 Trachybasait	60.40	14.20	2.45	2.09	20.0	6.08	5.09	4,01	50.00	0.36	87 GR	4,40	116	154	27	23	M	0.59	5 16	646	10.74	8	47.48	211E	78'00	1.021	15.02	55.89	511	341	88.8	3 00	8 34	1.13	5.87	78.0	CU. 82	0.28	1.47	0.21	15.2	0.718592	0.720830 (gr); 0.720232 (bl) 0.647244	1,512285 (ar): 0.512300 (bl)	18.957	15.643	39,399	4.31 4.31 evaluan for most samples is throm dredge D115) and all th id lavas may be a result of U.	C from ret. Jay and show at a amples from dredge D118, R w Samoan lavas exhibit good
D115-07 14.080 172.898 3220 Basalt	AB 20	11.53	3.13	9.19	0.13	8.75	863	2.68	2.41	0.45	17.58	6.25	365	926	21	52	119	89/0	80.0	1736	7,94	3.89	49.64	939	00.00	7.95	12.99	49.91	714	289	1.35	19.0	8.16	1.16	5.93	101	10.07	0.28	1.54	0.23	19.3	0.711481	0 617544	0	19.042	15,638	39,418	dredged from state of presi is the highest ically-enriche	Excluding si tios in the new
D115-03 14.090 172.898 3220 Hawaite	AD 10	12.59	3.50	9.74	0.13	8.17	7.97	3.08	587	19.0	EL IR	1.90	222	137	24	4	120	0.70	83.1	668	7.13	1.63	54.12	3.68	10101	7.77	11.97	48.31	617	305	8.02	314	9.02	1.25	6.44	1.06	07.17	030	1.59	0.22	19.8	0.711408	171807.0	0.512588	19.011	15.633	39.347	4.93 crust were setting, the its sample nost isotop	atio in neo M2 source. 9 Rb/Cs rat
D1 14-03 13 975 172 724 2490 Picrite	AA 02	7.42	3 00	12.33	0.18	8.89	20.07	0.83	1.00	0.36	20.88	ulu 1	820	258	13	92	127	0.22	14.2	177	3.28	0.63	33.64	231	cn/97	80.00 90.0	7.75	32.01	374	186	4.80	001	6 20	0.90	4.89	0.86	90'07	0.25	141	0.20	29.9	0.706435	0.610700	0017100	19 322	15619	39.475	anganese abmarine s he LOI in th	e D118, th
D114-01 13.975 172.724 2480 Basalt	AB 20	10.00	4.00	11.78	0.18	11.08	12.16	1.13	-R0.2	850	87.68	0.29	116	328	20	a	125	030	22.5	244	4.44	96.0	47.81	3.34	38.13	10.97	9.59	40.69	498	253	09.9	79.5	8.61	1.23	6.52	1.14	9 20 6	0.33	1.86	0.27	34.6	0.706422	0.610706	00/7100	19.317	15.618	39.473	the ferrom pled in a s AU (~5.8) h	the highly of tram
Sample ID Latitude (*5) Longitude (*W) Water Depth (m) Lithology	08	AlsO ₃	TIO2	FeOr*	MnO	CaO	MgO	KeO	Na ₂ O	F20%	Pre-total	-Be	Z (5 >	Ga	Ou	Zn	ő	82	Ва	£	Þ	gz	e .	5 2	5 6	Pr.	PN	ŝ	12	Ŧ,	E =	Po	٩	ò	P :	⊢ ů	2, ت	ЧY	Lu	Sc	"Str"Sr (WR.)	"SrimSr (cpc)	(xor) by Wey (cox)	(NR) dq by (NR.)	207 Pb/204 Pb (W.R.)	(.N.N.) department	² He//He (Ra) Basalt samples and In spite of being sam sample D115-07 (Th However, the high Th	to be a signature of and all the samples i

In house sandards (BCRP, BM-Q1 and MCH.01) that are desolved and processed with each batch of uninnows. The calchadon-values for these three standards were determined in 2001 using 17 international reference materials, including BCR-1 (Basab), VA2 (Dhatese), C-2 (Gannito), DC1 (Diatabate), DC2 (C) (Diatabate), DC2 (Diatabate), DC2 (Diatabate), DC3 (Diatabate), www.nature.com/nature

ratio to a minosphere, is 1384x10²). All helium data were actained by orushing of epx and stivner in vacuum. It is assumed here that radogenic contributions are negligable. However, without total host and taba, we cannot exclude the possibility of informaticas and approximate the tradogenic contributions are negligable. However, without total host and taba, we cannot exclude the possibility of informaticas

ement was made on a leached powder prepared from a different batch of rock chips from the same rock sample. the basalt matrix. measu eplicate

Cores viables at Innor-Dolery Earth Observatory core repository, and all cores were taken in the region of the Sancoanistands and seamcunds. Care was taken to sample all throtoges and, when making a composite of each core, the throtoges samped were combined in sproximately the same proportions as found in the respective cores. Composites were not subject to leading prior to isoarona and samonas to according to the ferromanguates find was taken from the rim of basist sample D116.22. The Min-Ind was powdered before analysis and not subjected to leaching. "All Fe reported as FeO." Mage = molar ratio of MgCI(MgO + 0.8/FeO).

Chapter 3

New Samoan lavas from Ofu Island reveal a hemispherically heterogeneous high ³He/⁴He mantle *

Abstract

New measurements of high ³He/⁴He ratios in Samoan lavas from Ofu Island (19.5–33.8 times atmospheric) extend the known range for ³He/⁴He in the southern hemisphere mantle. The Ofu data suggest that the high ³He/⁴He mantle component thought to be common to all oceanic hotspots, called FOZO (Focus Zone), is not homogeneous. Sr, Nd and Pb isotopes in Ofu lavas indicate that the Samoan high 3He/4He component is isotopically distinct from the high ³He/⁴He lavas from Hawaii, Iceland and Galapagos. Along with Samoa, the highest ³He/⁴He sample from each southern hemisphere high ³He/⁴He hotspot exhibits lower ¹⁴³Nd/¹⁴⁴Nd ratios than their counterparts in the northern hemisphere (excluding lavas erupted in continental, back-arc, and submarine ridge environments). The observation of a large-scale isotopic enrichment (generally higher ⁸⁷Sr/⁸⁶Sr and lower ¹⁴³Nd/¹⁴⁴Nd) in the FOZO-A (austral) high 3He/4He mantle compared to the FOZO-B (boreal) high ³He/⁴He mantle is similar to the DUPAL anomaly, a globe-encircling feature of isotopic enrichment observed primarily in southern hemisphere ocean island basalts. The recent discovery that terrestrial samples have ¹⁴²Nd/¹⁴⁴Nd ratios higher than chrondrites has potentially important implications for the origin of the FOZO reservoirs, and suggest that the high ³He/⁴He mantle has been re-enriched.

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New Samoan lavas from Ofu Island reveal a hemispherically heterogeneous high ³He/⁴He mantle

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11 Abstract

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New measurements of high ³He/⁴He ratios in Samoan lavas from Ofu Island (19.5 33.8 times atmospheric) extend the known 12 range for ³He/⁴He in the southern hemisphere mantle. The Ofu data suggest that the high ³He/⁴He mantle component thought to be 13 common to all oceanic hotspots, called FOZO (Focus Zone), is not homogeneous. Sr, Nd and Pb isotopes in Ofu lavas indicate that 14 the Samoan high ³He⁴He component is isotopically distinct from the high ³He⁴He lavas from Hawaii, Iceland and Galapagos. 15 Along with Samoa, the highest ³He/⁴He sample from each southern hemisphere high ³He/⁴He hotspot #exhibits lower ¹⁴³Nd/¹⁴⁴Nd 16 ratios than their counterparts in the northern hemisphere (excluding lavas erupted in continental, back-arc, and submarine ridge 17 environments). The observation of a large-scale isotopic enrichment (generally higher ⁸⁷Sr/86Sr and lower ¹⁴³Nd/¹⁴⁴Nd) in the 18 FOZO-A (austral) high ³He/⁴He mantle compared to the FOZO-B (boreal) high ³He/⁴He mantle is similar to the DUPAL anomaly, 19 a globe-encircling feature of isotopic enrichment observed primarily in southern hemisphere ocean island basalts. The recent 20 discovery that terrestrial samples have 142Nd/144Nd ratios higher than chrondrites has potentially important implications for the 21origin of the FOZO reservoirs, and suggest that the high ³He/⁴He mantle has been re-enriched. 22 © 2007 Published by Elsevier B.V. 23

23 24

25 Keywords: Samoa; FOZO; PHEM; C; 3He/4He; mantle; basalt; geochemistry; DUPAL; hotspot

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27 1. Introduction

Oceanic lavas with high ³He/⁴He signatures are rare, and derive from ancient reservoirs in the earth's mantle. Volcanically active hotspots with high ³He/⁴He lavas, such as Samoa and Hawaii, sample melts of buoyantly upwelling regions of the deep mantle where the high ³He/⁴He reservoir is thought to reside (Kurz et al., 1982; 33 Hart et al., 1992; Class and Goldstein, 2005). Conse- 34 quently, ocean island basalts (OIBs) erupted at hotspots 35 provide a unique tool for probing the composition and 36 history of the deep mantle. Radiogenic isotopes in OIBs 37 are commonly used as tracers for the various mantle 38 components revealed at hotspots, and show a diverse 39 range of compositions, or endmembers, including DMM 40 (depleted mid-ocean ridge basalt [MORB] mantle), 41 HIMU (high 'µ', or ²³⁸U/²⁰⁴Pb mantle), EM1 and EM2 42 (enriched mantle 1 and 2) (Zindler and Hart, 1986). 43

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² upweining regions of the teep manife where the high

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Emerging from this taxonomic diversity, a unifying 44 theory in mantle geochemistry maintains that a single 45 mantle component exists that is common to all hotspots 46 (Hart et al., 1992): Mixing arrays from individual ocean 47 48 islands originate near the mantle endmembers in various radiogenic isotope spaces and converge on a region 49 50 characterized by depleted isotope ratios that is distinct from normal MORB (Hart et al., 1992). Lavas plotting 51 in this region of convergence often exhibit elevated 52 ³He/⁴He ratios, and are suggested to sample a com-53 ponent in the mantle common to all hotspots. Variably 54 called FOZO (Hart et al., 1992), PHEM (Primitive 55 Helium Mantle; Farley et al., 1992), or C (Common; 56 Hanan and Graham, 1996), the high ³He/⁴He common 57 58 component is thought to be a relatively less degassed region of the (lower?) mantle (Kurz et al., 1982; Class 59and Goldstein, 2005). 60

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New geochemical data from the Samoan hotspot are 61 not entirely consistent with this view of the mantle. 62 The new ${}^{3}\text{He}/{}^{4}\text{He}$ ratios (up to 33.8±0.2 Ra, ratio to 63 atmosphere) from the Samoan Island of Ofu are the 64 highest yet recorded in the southern hemisphere, and are 65 significantly higher than ³He/⁴He ratios previously 66 measured (25.8 Ra) in Samoan basalts and xenoliths 67 (Farley et al., 1992; Poreda and Farley, 1992; Workman 68 et al., 2004). The new helium isotope data from Samoa 69 extend the range of observed ³He/⁴He up to values 70 comparable to those found in Hawaii (32.3 Ra; Kurz 71 et al., 1982), Iceland (37.7 Ra; Hilton et al., 1999) and the 72Galapagos (30.3 Ra; Kurz and Geist, 1999; Saal et al., 73 2007), referred to here as HIG. The Ofu lavas are 74 isotopically more enriched (higher 87 Sr/86 Sr and lower 75 ¹⁴³Nd/¹⁴⁴Nd) than the high ³He/⁴He samples from HIG 76 and exhibit elevated incompatible trace element con-77 centrations. Due to this isotopic and trace element 78enrichment relative to HIG lavas, the new data from 79 Samoan high ³He/⁴He lavas are inconsistent with recent 80 models that describe the evolution of the high ³He/⁴He 81 mantle, and the Ofu data suggest that the high ³He/⁴He 82 mantle domain is isotopically heterogeneous. 83

84 2. Methods and results

85 2.1. Sample location and state of preservation

Ofu Island is located in the eastern province of the Samoan archipelago, an age-progressive hotspot track (Hart et al., 2004; Koppers et al., submitted for publication) located just north of the northern terminus of the Tonga subduction zone. The samples were collected at various locations on the perimeter of Ofu and Olosega islands (Ofu hereafter; sample location map is available in the supplementary data in the Appendix). 93 Tholeiitic lavas in Samoa are rare (Natland, 1980; 94 Workman et al., 2004), and with the exception of a 95 cumulate (OFU-04-14) and a gabbro (OFU-04-17), 96 the Ofu lavas presented in this study are alkali basalts 97 (Table 1). The Ofu samples are generally quite fresh. 98 With the exception of sample OFU-04-12, which has a 99 Th/U ratio of 4.8 (and may indicate U-loss during sub- 100 aerial weathering), the range of Th/U ratios in the Ofu 101 sample suite is 4.0-4.4. The Ba/Rb ratios for Ofu sam- 102 ples (9.1±1.1 at 1σ) are similar to the values for young 103 Samoan basalts reported previously (Workman et al., 104 2004), and somewhat lower than the canonical value of 105 ~12 for fresh OIB lavas (Hofmann and White, 1983). 106 Excluding the cumulate sample OFU-04-14, which has a 107 Rb/Cs ratio of 280, the range of Rb/Cs values from the 108 Ofu samples (from 73 to 137) is close to the canonical 109 range of 85-95 (Hofmann and White, 1983). These 110 weathering proxies indicate that elements equally or less 111 mobile than U, Rb and Cs yield useful petrogenetic 112 information. 113

2.2. He, Sr, Nd and Pb isotopes in Ofu lavas 114

New ³He/⁴He values (19.5 to 33.8 Ra) were measured 115 at Woods Hole Oceanographic Institution on olivine 116 and clinopyroxene (cpx) phenocrysts separated from 12 117 hand samples (Table 1). Measurements were made by 118 crushing and fusion in vacuo, following the protocol 119 reported in (Kurz et al., 2004). The sample with the 120 highest ³He/⁴He value, OFU-04-06, was taken from an 121 ankaramite dike exposed at 2 m depth in a recent road 122 cut. Olivines from this sample are relatively gas rich 123 $(67.5 \times 10^{-9} \text{ cm}^3 \text{ STP g}^{-1}$, the sum of crushing and 124 fusion) and yielded similar ³He/⁴He ratios on two sepa- 125 rate crushes of the same olivine separate (OFU-04-06cr1 126 followed by OFU-04-06cr2). Following the crushing 127 experiments, a fusion extraction of the resulting olivine 128 powder (OFU-04-06fus) yielded lower ³He/⁴He, indi- 129 cating the presence of ingrown radiogenic helium, prob- 130 ably implanted from the Th and U-rich matrix. Two 131 different olivine populations (lighter and darker olivines) 132 separated from sample OFU-04-06 yielded similar 133 ³He/⁴He ratios (33.4 and 33.6 Ra). These fusion and 134 crushing experiments, coupled with sampling depths, 135 preclude the influence of cosmogenic helium for 136 this important sample. The lava with the second highest 137 ³He/⁴He value measured by crushing-sample OFU-04- 138 15-also yielded lower ³He/⁴He ratios upon fusion 139 (OFU-04-15fus) of the crushed olivine powder. Addi- 140 tionally, crushing experiments of sample OFU-04-03 141 (OFU-04-03cr1 followed by OFU-04-03cr2) yielded 142

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ORU-04-10 Olivine Aniastratite Flow 3.2 25.6 0.5 0.70-643 0.51233 19.16 15.577 39.206 3.25 503 0.20 131.6 5.4 0.20 ORU-04-10 Olivine Alkail Basalt Flow 8.0 27.0 0.2 9.135 19.157 39.206 3.55 39.09 107.6 9.09 107.6 9.09 107.6 9.09 108.0 5.3 0.20 131.6 5.4 0.20 ORU-04-11 Orivine Alkail Basalt Flow NA NA NA 0.70-6453 0.512304 19.184 15.554 39.206 139.0 131.6 5.4 0.20 131.6 5.4 0.20 131.6 5.4 0.20 131.6 5.4 0.20 131.6 5.4 0.20 131.6 5.4 0.20 0.31.6 5.4 0.20 0.31.6 5.4 0.20 0.20 131.6 5.559 3.25 5.0 0.20 0.70-513.6 5.1158 </td <td>$\begin{array}{c c c c c c c c c c c c c c c c c c c$</td> <td>$\begin{array}{c ccccccccccccccccccccccccccccccccccc$</td> <td>$\begin{array}{c c c c c c c c c c c c c c c c c c c$</td> <td>$\begin{array}{c cccc} 0.01-0.410 & 01vine & Antarmire Flow & 3.2 & 2.56 & 0.5 & 0.70-533 & 0.51335 & 19.16 & 15.577 & 9.206 & 3.25 & 503 & 4.03 & 10.20 & 1316 & 5.4 & 0.46 \\ 0.01-0.410 & 0rx & Alkali Basali Flow NA & NA & NA & 0.70-653 & 0.512815 & 19.187 & 15.583 & 9.206 & 5.6 & 9.29 & 4.30 & 10.79 & 730 & 5.8 & 0.43 \\ 0.01-0.4+11 & None & Takyh Basali Flow NA & NA & NA & 0.70-653 & 0.512810 & 19.187 & 15.583 & 9.206 & 5.0 & 29 & 4.30 & 0.70 & 5.3 & 0.42 \\ 0.01-0.4+12 & 01vine & Alkali Basali Flow NA & NA & NA & N7-655 & 0.512810 & 19.187 & 15.584 & 9.405 & 3.29 & 311 & 4.78 & 10.36 & 1183 & 6.2 & 0.42 \\ 0.01-0.4+12 & 01vine & Alkali Basali Flow NA & NA & NA & N7-655 & 0.512810 & 19.126 & 15.584 & 9.405 & 3.29 & 311 & 4.78 & 10.36 & 1183 & 6.2 & 0.42 \\ 0.01-0.4+13 & None & Hawaire Boulder & 4.3 & 2.50 & 0.2 & 0.70-551 & 0.512810 & 19.126 & 15.584 & 9.168 & 2.39 & 2.49 & 0.37 \\ 0.01-0.4+13 & 01vine & Antaramire Boulder & 4.3 & 2.96 & 0.2 & 0.70-551 & 0.51284 & 9.116 & 15.583 & 9.169 & 2.38 & 2.04 & 0.40 \\ 0.01-0.4+16 & 01vine & Antaramire Boulder & 1.3 & 2.46 & 9.30 & 8.01 & 2.93 & 8.0 & 4.20 & 0.37 \\ 0.01-0.4+16 & 01vine & Autaramire Boulder & 1.2 & 0.2 & 0.70-551 & 0.51284 & 9.116 & 15.583 & 9.169 & 2.93 & 4.39 & 0.51 & 2.42 & 0.42 \\ 0.01-0.4+16 & 01vine & Autaramire Boulder & 1.2 & 0.70-458 & 0.512844 & 9.216 & 15.583 & 9.169 & 2.38 & 0.42 & 0.90 & 0.37 & 0.42 & 0.44 & 0.44 & 0.44 & 13.4 & 14.4 & 14$</td> <td>$\begin{array}{cccccccccccccccccccccccccccccccccccc$</td> <td>$\begin{array}{c ccccccccccccccccccccccccccccccccccc$</td> <td> ORU-04-10 Olivine Antananie Flow 3.2 ORU-04-10 Olivine Antananie Flow 8.0 ORU-04-10 Olivine Antani Basalt Flow 8.0 ORU-04-10 Olivine Antani Basalt Flow 8.0 ORU-04-10 Olivine Antani Basalt Flow 8.0 ORU-04-11 None Atkali Basalt Flow 8.0 ORU-04-12 Olivine Antani Basalt Flow 8.0 ORU-04-13 Olivine Antani Basalt Flow 8.0 ORU-04-14 Olivine Antani Basalt Flow 8.0 ORU-04-15 Olivine Antani Basalt Flow 8.0 ORU-04-14 Olivine Antani Basalt Flow 8.0 ORU-04-15 Olivine Antanani Boulder 4.5 Olivine Antanani Boulder 4.5 Olivine Antanani Boulder 4.5 Olivine Antanani Boulder 1.4 Olivine Antanani Boulder 1.1 Olivine Antana Antana Antana Antana Antana Antana</td> <td>10.01 000</td> <td>15 586</td> <td>162.65</td> <td>4 06</td> <td>513</td> <td>411</td> <td>8 78</td> <td>27.5</td> <td>57</td> <td>0.40</td>	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c cccc} 0.01-0.410 & 01vine & Antarmire Flow & 3.2 & 2.56 & 0.5 & 0.70-533 & 0.51335 & 19.16 & 15.577 & 9.206 & 3.25 & 503 & 4.03 & 10.20 & 1316 & 5.4 & 0.46 \\ 0.01-0.410 & 0rx & Alkali Basali Flow NA & NA & NA & 0.70-653 & 0.512815 & 19.187 & 15.583 & 9.206 & 5.6 & 9.29 & 4.30 & 10.79 & 730 & 5.8 & 0.43 \\ 0.01-0.4+11 & None & Takyh Basali Flow NA & NA & NA & 0.70-653 & 0.512810 & 19.187 & 15.583 & 9.206 & 5.0 & 29 & 4.30 & 0.70 & 5.3 & 0.42 \\ 0.01-0.4+12 & 01vine & Alkali Basali Flow NA & NA & NA & N7-655 & 0.512810 & 19.187 & 15.584 & 9.405 & 3.29 & 311 & 4.78 & 10.36 & 1183 & 6.2 & 0.42 \\ 0.01-0.4+12 & 01vine & Alkali Basali Flow NA & NA & NA & N7-655 & 0.512810 & 19.126 & 15.584 & 9.405 & 3.29 & 311 & 4.78 & 10.36 & 1183 & 6.2 & 0.42 \\ 0.01-0.4+13 & None & Hawaire Boulder & 4.3 & 2.50 & 0.2 & 0.70-551 & 0.512810 & 19.126 & 15.584 & 9.168 & 2.39 & 2.49 & 0.37 \\ 0.01-0.4+13 & 01vine & Antaramire Boulder & 4.3 & 2.96 & 0.2 & 0.70-551 & 0.51284 & 9.116 & 15.583 & 9.169 & 2.38 & 2.04 & 0.40 \\ 0.01-0.4+16 & 01vine & Antaramire Boulder & 1.3 & 2.46 & 9.30 & 8.01 & 2.93 & 8.0 & 4.20 & 0.37 \\ 0.01-0.4+16 & 01vine & Autaramire Boulder & 1.2 & 0.2 & 0.70-551 & 0.51284 & 9.116 & 15.583 & 9.169 & 2.93 & 4.39 & 0.51 & 2.42 & 0.42 \\ 0.01-0.4+16 & 01vine & Autaramire Boulder & 1.2 & 0.70-458 & 0.512844 & 9.216 & 15.583 & 9.169 & 2.38 & 0.42 & 0.90 & 0.37 & 0.42 & 0.44 & 0.44 & 0.44 & 13.4 & 14.4 & 14$	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	 ORU-04-10 Olivine Antananie Flow 3.2 ORU-04-10 Olivine Antananie Flow 8.0 ORU-04-10 Olivine Antani Basalt Flow 8.0 ORU-04-10 Olivine Antani Basalt Flow 8.0 ORU-04-10 Olivine Antani Basalt Flow 8.0 ORU-04-11 None Atkali Basalt Flow 8.0 ORU-04-12 Olivine Antani Basalt Flow 8.0 ORU-04-13 Olivine Antani Basalt Flow 8.0 ORU-04-14 Olivine Antani Basalt Flow 8.0 ORU-04-15 Olivine Antani Basalt Flow 8.0 ORU-04-14 Olivine Antani Basalt Flow 8.0 ORU-04-15 Olivine Antanani Boulder 4.5 Olivine Antanani Boulder 4.5 Olivine Antanani Boulder 4.5 Olivine Antanani Boulder 1.4 Olivine Antanani Boulder 1.1 Olivine Antana Antana Antana Antana Antana Antana	10.01 000	15 586	162.65	4 06	513	411	8 78	27.5	57	0.40
ORU-04-10 Offware Alkali Basalt Flow 8.0 2.20 0.2 *	$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	834 10165	15 577	30.08	30.5	203	4.03	10.20	131.6	44	0.46
OFU-04-10 opx Alkali Basalt Flow 1.0 19.7 0.6 0.704638 0.512815 19.187 15.585 39.261 3.65 4.39 10.79 73.0 5.8 0.2 OFU-04-11 None Trachybrasht Flow NA NA NA 0.704635 0.512804 19.184 15.584 39.261 3.65 4.39 109.79 73.0 5.8 0.2 OFU-04-11 None Trachybrasht Flow NA NA 0.704635 0.512800 19.201 15.594 39.261 3.93 4.78 0.99 183 6.2 0.2 0.704635 0.512800 19.202 15.594 39.164 2.93 3.13 4.17 30.09 9.28 0.1 9.26 0.2 0.704635 0.512810 19.126 15.584 39.164 2.93 3.9164 2.93 2.913 4.28 0.1 OFU-04-15 Oftwine Ankaramite Boulder 1.43 2.29 0.704535 0.512810 9.126 1.925	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	$\begin{array}{c c c c c c c c c c c c c c c c c c c $									ir	
DTU-04-11 None Trachybracht How NA NA NA 0.704635 0.512804 19184 15.594 39.299 6.60 929 4.30 909 108.0 5.3 0.1 DTU-04-12 Olivine Alkali Basalt Flow 3.5 2112 0.3 0.704595 0.512810 19.202 15.594 3.9405 3.29 311 4.78 10.36 118.3 6.2 0.2 0.704595 0.512810 19.202 15.581 3.02 4.07 9.09 9.83 5.0 0.2 0.704517 0.512810 19.105 15.581 3.026 5.02 13.83 6.2 0.2 0.704517 0.704517 0.512810 19.105 15.581 3.026 5.02 13.83 6.2 0.2 DTU-04-15 None Hawaire Boulder 14.3 22.0 0.704539 0.512821 19.126 15.581 39.164 2.32 38.30 4.8 0.2 0.2 0.704535 5.12834 19.126 3.29 18.2 18.2 0.2 DTU-04-15 Ohvine Alkani Basalt <th< td=""><td>$\begin{array}{c c c c c c c c c c c c c c c c c c c$</td><td>$\begin{array}{c c c c c c c c c c c c c c c c c c c$</td><td>$\begin{array}{c c c c c c c c c c c c c c c c c c c$</td><td>$\begin{array}{c c c c c c c c c c c c c c c c c c c$</td><td>$\begin{array}{c c c c c c c c c c c c c c c c c c c$</td><td>$\begin{array}{c c c c c c c c c c c c c c c c c c c$</td><td> DTU-04-11 None Tradiybasalt Flow NA NA NA NA 0.704655 0.5123 DTU-04-12 Olivine Alkali Basalt Flow 3.5 21.2 0.3 0.704755 0.5123 DTU-04-13 None Hawaire Flow NA NA 0.704655 0.5123 DTU-04-14 Olivine Aukanite Boulder 4.5 2.5.0 0.2 0.704517 0.5123 DTU-04-15 Olivine Ankaranic Boulder 1.4.5 29.6 0.2 0.704517 0.5123 DTU-04-16 cpx Ankaranic Boulder 1.4.5 29.6 0.2 0.704517 0.5123 DTU-04-15 cpx Ankaranic Boulder 1.4.5 20.2 0.704517 0.5123 DTU-04-16 cpx Ankaranic Boulder 1.4.5 20.2 0.704519 0.5123 DTU-04-16 cpx Ankaranic Boulder 1.0 2 0.704519 0.5123 DTU-04-17 Olivine Gabbio Boulder 1.0 2 0.70441 0.704488 0.5123 Motora)^b Ankoranic Ankaranic NA 1.3 1.70 1.6 0.704681 0.5123 Motora)^b Samples were collected at Oft Island, American Samoa in August of 2004, All "He"He ratios are reported reletive to a poetcometry and associated measurement precision are reported in Hart and Bhuschin (2006). Si, Nd and Po-isotope an indextender dower were by CPC 4.04 0.704488 0.5123 Motora)^b Samples were collected at Oft Island, American Samoa in August of 2004, All "He"He ratios are reported reletive talges at the was interpreted for the more the cross the cross by CPC 4.04 0.704484 0.404 core site target target to a state that the one target target target target target target target to the one target target</td><td>815 19.187</td><td>15.585</td><td>39.261</td><td>3.65</td><td>495</td><td>4.39</td><td>10.79</td><td>73.0</td><td>5.8</td><td>0.43</td></th<>	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	 DTU-04-11 None Tradiybasalt Flow NA NA NA NA 0.704655 0.5123 DTU-04-12 Olivine Alkali Basalt Flow 3.5 21.2 0.3 0.704755 0.5123 DTU-04-13 None Hawaire Flow NA NA 0.704655 0.5123 DTU-04-14 Olivine Aukanite Boulder 4.5 2.5.0 0.2 0.704517 0.5123 DTU-04-15 Olivine Ankaranic Boulder 1.4.5 29.6 0.2 0.704517 0.5123 DTU-04-16 cpx Ankaranic Boulder 1.4.5 29.6 0.2 0.704517 0.5123 DTU-04-15 cpx Ankaranic Boulder 1.4.5 20.2 0.704517 0.5123 DTU-04-16 cpx Ankaranic Boulder 1.4.5 20.2 0.704519 0.5123 DTU-04-16 cpx Ankaranic Boulder 1.0 2 0.704519 0.5123 DTU-04-17 Olivine Gabbio Boulder 1.0 2 0.70441 0.704488 0.5123 Motora)^b Ankoranic Ankaranic NA 1.3 1.70 1.6 0.704681 0.5123 Motora)^b Samples were collected at Oft Island, American Samoa in August of 2004, All "He"He ratios are reported reletive to a poetcometry and associated measurement precision are reported in Hart and Bhuschin (2006). Si, Nd and Po-isotope an indextender dower were by CPC 4.04 0.704488 0.5123 Motora)^b Samples were collected at Oft Island, American Samoa in August of 2004, All "He"He ratios are reported reletive talges at the was interpreted for the more the cross the cross by CPC 4.04 0.704484 0.404 core site target target to a state that the one target target target target target target target to the one target target	815 19.187	15.585	39.261	3.65	495	4.39	10.79	73.0	5.8	0.43
DFU-04-12 Ofivine Alkali Basalt Flow 3.5 21.2 0.3 0.704795 0.512800 19.202 15.594 39.405 3.29 31 4.78 10.36 118.3 6.2 0. DFU-04-13 None Hawaire Flow NA NA NA NA 0.704795 0.512810 19.176 15.581 3.29 31 4.78 10.36 118.3 6.2 0. DFU-04-13 Nome Hawaire Boulder 4.5 2.50 0.2 0.704517 0.512819 19.156 15.581 39.164 2.39 288 28.0 4.8 0. DFU-04-15 Olivine Ankarmire Boulder 14.3 2.9 0.704519 0.512821 19.141 15.589 39.164 2.92 2.99 9.23 2.9 0. DFU-04-15 Olivine Ankarmire Boulder 14.3 2.9 0.7 15.581 19.156 15.581 39.169 3.29 4.9 0.9 9.2 4.8 0. 0. DU-04.15 Olivine <td< td=""><td>$\begin{array}{c c c c c c c c c c c c c c c c c c c$</td><td>$\begin{array}{c c c c c c c c c c c c c c c c c c c$</td><td>$\begin{array}{c c c c c c c c c c c c c c c c c c c$</td><td>$\begin{array}{c c c c c c c c c c c c c c c c c c c$</td><td>$\begin{array}{cccccccccccccccccccccccccccccccccccc$</td><td>$\begin{array}{c c c c c c c c c c c c c c c c c c c$</td><td> DFU-04-12 Olivine Alkali Basalt Flow 3.5 21.2 0.3 0.704795 0.5123 DFU-04-13 None Hawaire Flow NA NA NA NA 0.704625 0.5123 DFU-04-14 Olivine Cumulate Boulder 4.5 2.5.0 0.2 0.704517 0.5123 DFU-04-15 Givine Ankaramic Boulder 1.4.3 2.9.6 0.2 0.704519 0.5123 DFU-04-16 Gray Ankaramic Boulder 1.4.3 2.9.6 0.2 0.704519 0.5123 DFU-04-16 Gray Ankaramic Boulder 1.4.3 2.9.6 0.2 0.704519 0.5123 DFU-04-16 Gray Ankaramic Boulder 1.2 2.2 2 7 0.704519 0.5123 DFU-04-16 Gray Ankaramic Boulder 1.0 2 1.9 1 0.704438 0.5123 DFU-04-17 Olivine Gabboo Boulder 1.0 2.0.13 1.7.0 1.6 0.704638 0.5123 Motoral^b Ankaramic NA 1.3 1.7.0 1.6 0.704638 0.5123 Aloreal^b Ankaramic NA 1.3 1.7.0 1.6 0.704631 0.5123 Aloreal^b Samples were collected at Ofti Island, American Samoa in August of 2004, All "He/He mitos are reported relative to a pettometry and associated measurement precision are reported in Hart and Bluszhijn (2006). 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The "He/He ratio</td><td>804 19.184</td><td>15.594</td><td>39.299</td><td>6.60</td><td>929</td><td>4.30</td><td>9.09</td><td>108.0</td><td>5.3</td><td>0.45</td></td<>	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	 DFU-04-12 Olivine Alkali Basalt Flow 3.5 21.2 0.3 0.704795 0.5123 DFU-04-13 None Hawaire Flow NA NA NA NA 0.704625 0.5123 DFU-04-14 Olivine Cumulate Boulder 4.5 2.5.0 0.2 0.704517 0.5123 DFU-04-15 Givine Ankaramic Boulder 1.4.3 2.9.6 0.2 0.704519 0.5123 DFU-04-16 Gray Ankaramic Boulder 1.4.3 2.9.6 0.2 0.704519 0.5123 DFU-04-16 Gray Ankaramic Boulder 1.4.3 2.9.6 0.2 0.704519 0.5123 DFU-04-16 Gray Ankaramic Boulder 1.2 2.2 2 7 0.704519 0.5123 DFU-04-16 Gray Ankaramic Boulder 1.0 2 1.9 1 0.704438 0.5123 DFU-04-17 Olivine Gabboo Boulder 1.0 2.0.13 1.7.0 1.6 0.704638 0.5123 Motoral^b Ankaramic NA 1.3 1.7.0 1.6 0.704638 0.5123 Aloreal^b Ankaramic NA 1.3 1.7.0 1.6 0.704631 0.5123 Aloreal^b Samples were collected at Ofti Island, American Samoa in August of 2004, All "He/He mitos are reported relative to a pettometry and associated measurement precision are reported in Hart and Bluszhijn (2006). Sr, Nd and Pb-isotope an indextangle When an olivine samo in August of 2004, All "He/He mitos are reported relative to a spectometry and associated measurement precision are reported for them case, the crusis experiment is labeled as "ci-a" a Required a black correction of 46% due to low gas consentitations. The "He/He ratio on ci-dis ethnopyroxer measi and the consection of 46% due to low gas consentitations. The "He/He ratio on ci-dis ethnopyroxer measi and the sectometry and sectometre measurement precision are reported more than on ci-dis ethnopyroxer measi and the consection of 46% due to low gas consentitations. The "He/He ratio on ci-dis ethnopyroxer measi and the sectometre more measine measurement on ci-dis ethnopyroxer measine and the sectometrations. The "He/He ratio on ci-dis ethnopyroxer measine and the sectometrations. The "He/He ratio on ci-dis ethnopyroxer measine and the sectometrations. The "He/He ratio on ci-dis ethnopyroxer measine and the sectometrations. The "He/He ratio	804 19.184	15.594	39.299	6.60	929	4.30	9.09	108.0	5.3	0.45
DFU-04-13 None Hawnite Flow NA NA 0.704625 0.512816 19.176 15.581 39.264 5.02 7.03 4.07 9.09 9.22 4.8 0.0 DFU-04-14 Olivine Cumulate Boulder 4.5 250 0.2 0.704517 0.512819 19.126 15.584 39.164 2.95 9.88 280 280 280 280 280 280 280 280 280 280 280 280 29.166 2.95 9.92 4.8 0.0 0.704517 0.512821 19.141 15.580 39.169 3.83 280 4.8 0.2 0.704518 0.512824 19.141 15.580 39.169 3.8 0.9 87.8 0.2 0.704518 0.512834 19.216 15.587 39.276 3.9 601 4.96 5.3 0.2 0.704438 0.512834 19.160 15.587 39.276 3.9 60 129.65 5.0 0.704438 0.512844 <td>$\begin{array}{c c c c c c c c c c c c c c c c c c c$</td> <td>$\begin{array}{c c c c c c c c c c c c c c c c c c c$</td> <td>$\begin{array}{c c c c c c c c c c c c c c c c c c c$</td> <td>$\begin{array}{c c c c c c c c c c c c c c c c c c c$</td> <td>$\begin{array}{c c c c c c c c c c c c c c c c c c c$</td> <td>$\begin{array}{c c c c c c c c c c c c c c c c c c c$</td> <td> DFU-04-13 None Havaire Flow NA NA NA 0.704625 0.5123 DFU-04-14 Olivine Cuandate Boulder 4.5 25.0 0.2 0.704517 0.5132 DFU-04-15 Olivine Cuandate Boulder 4.5 25.0 0.2 0.704519 0.5123 DFU-04-15 Olivine Ankaramic Boulder 1.4.3 29.6 0.2 0.704519 0.5123 DFU-04-15 Olivine Ankaramic Boulder 1.4.3 29.6 0.2 0.704438 0.5123 DFU-04-17 Olivine Ankaramic Boulder 1.0 2.6.4 0.4 0.704438 0.5123 DFU-04-17 Olivine Gabbio Boulder 1.0 2.6.4 0.4 0.704438 0.5123 MOJ-01 Olivine Ankaramic NA 1.3 17.0 1.6 0.704438 0.5123 MOJ-01 Olivine Ankaramic NA 1.3 17.0 1.6 0.70461 0.512 MOD-01 Olivine Ankaramic NA 1.3 17.0 1.6 0.70461 0.512 Mooteab⁵ were collected at Oft Island, American Samoa in August of 2004 All *He/He ratios are reported relative to a performent precision are reported in Hart and Bhazhin (2006) Si, Nia and Pb-isolope an indexched whole rock powders by ICP at the Washington State University CocAmbylical 1.4.0. Aradi associated measurement precision are reported for Hart and Bhazhin (2006) Si, Nia and Pb-isolope an indexched whole rock powders by ICP at the Washington State University CocAmbylical 1.4.0. Atomote State University are reported for the mote the crust septentent is labeled as "critoxe are argoned were studied more than one choic setting the last as "a strategian" are reported for the mote than one crito and physical physicas are reported as the one of 4.6% is the low physicas than one choice than are strated as the one state choree than the area one one than one choice there are area one for the mote state correction on the stander as "a required a for "critome stander as the one state choree mote area area to a for the mote state correction are state as "a". </td> <td>800 19.202</td> <td>15.594</td> <td>39,403</td> <td>3.29</td> <td>331</td> <td>4.78</td> <td>10.36</td> <td>118.3</td> <td>6.2</td> <td>0.42</td>	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	 DFU-04-13 None Havaire Flow NA NA NA 0.704625 0.5123 DFU-04-14 Olivine Cuandate Boulder 4.5 25.0 0.2 0.704517 0.5132 DFU-04-15 Olivine Cuandate Boulder 4.5 25.0 0.2 0.704519 0.5123 DFU-04-15 Olivine Ankaramic Boulder 1.4.3 29.6 0.2 0.704519 0.5123 DFU-04-15 Olivine Ankaramic Boulder 1.4.3 29.6 0.2 0.704438 0.5123 DFU-04-17 Olivine Ankaramic Boulder 1.0 2.6.4 0.4 0.704438 0.5123 DFU-04-17 Olivine Gabbio Boulder 1.0 2.6.4 0.4 0.704438 0.5123 MOJ-01 Olivine Ankaramic NA 1.3 17.0 1.6 0.704438 0.5123 MOJ-01 Olivine Ankaramic NA 1.3 17.0 1.6 0.70461 0.512 MOD-01 Olivine Ankaramic NA 1.3 17.0 1.6 0.70461 0.512 Mooteab⁵ were collected at Oft Island, American Samoa in August of 2004 All *He/He ratios are reported relative to a performent precision are reported in Hart and Bhazhin (2006) Si, Nia and Pb-isolope an indexched whole rock powders by ICP at the Washington State University CocAmbylical 1.4.0. Aradi associated measurement precision are reported for Hart and Bhazhin (2006) Si, Nia and Pb-isolope an indexched whole rock powders by ICP at the Washington State University CocAmbylical 1.4.0. Atomote State University are reported for the mote the crust septentent is labeled as "critoxe are argoned were studied more than one choic setting the last as "a strategian" are reported for the mote than one crito and physical physicas are reported as the one of 4.6% is the low physicas than one choice than are strated as the one state choree than the area one one than one choice there are area one for the mote state correction on the stander as "a required a for "critome stander as the one state choree mote area area to a for the mote state correction are state as "a". 	800 19.202	15.594	39,403	3.29	331	4.78	10.36	118.3	6.2	0.42
DFU-04-14 Olivine Cumulate Boulder 4.5 25.0 0.2 0.704517 0.512819 19.126 15.584 39.164 2.95 292 4.39 9.88 280.2 4.9 0. DFU-04-15 fts Olivine Ankaramite Boulder 14.3 2.96 0.2 0.70459 0.512822 19.141 15.580 39.169 3.89 601 4.02 9.28 8.80 4.8 0. DFU-04-15 fts Olivine Ankaramite Boulder 12 0.2 19 0.2 0.70459 0.512824 19.216 15.587 39.216 3.54 639 3.96 860 12.96 5.3 0. DFU-04-16 epi to the Ankaramite NA 1.3 17.0 1.6 0.704431 0.51274 19.216 15.503 38.931 2.11 381 3.75 11.14 2704 8.5 0. Offore Olivine Ankaramite NA 1.3 17.0 1.6 0.704621 0.512774 19.216 15.603 38.933 2.11 381 3.75 11.14 2704 8.5 0. Adoreal ^b (Adorea) ^b	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c ccccc} \text{JFU-04-14} & \text{Olivine} & \text{Cumulate} & \text{Boulder} & 45 & 2.0 & 0.2 & 0.704517 & 0.512819 & 19.126 & 15.584 & 39.164 & 2.95 & 292 & 439 & 9.88 & 2.802 & 4.9 & 0.37 \\ \text{JFU-04-15} & \text{Olivine} & \text{Ankaramic} & \text{Boulder} & 1.4 & 3 & 2.96 & 0.2 & 0.704539 & 0.512822 & 19.141 & 15.580 & 39.169 & 3.89 & 601 & 4.02 & 9.28 & 830 & 4.8 & 0.43 \\ \text{JFU-04-15} & \text{Olivine} & \text{Ankaramic} & \text{Boulder} & 1.2 & 2.2 & 1.2 & 1.2 & 2.2 & 19.141 & 15.580 & 39.169 & 3.89 & 601 & 4.02 & 9.28 & 830 & 4.8 & 0.43 \\ \text{JFU-04-16} & \text{Olivine} & \text{Ankaramic} & \text{Boulder} & 1.2 & 2.2 & 1.2 &$	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	DFU-04-14 Olivine Cumulate Boulder 4.5 2.5.0 0.2 0.704517 0.5128 DFU-04-15 Olivine Ankaramic Boulder 14.3 2.9.6 0.2 0.704519 0.5128 DFU-04-16 cpx Ankaramic Boulder 14.3 2.9.6 0.2 0.704539 0.5128 DFU-04-16 cpx Ankaramic Boulder 1.2 12 0.2 " 3.001438 DFU-04-16 cpx Ankaramic Boulder 1.0 2.6.4 0.4 0.704438 0.5128 DFU-04-10 Olivine Gabba Boulder 1.0 2.6.4 0.4 0.704463 0.5128 (G001-01 Olivine Gabbar 1.3 1.7.0 1.6 0.704621 0.5127 (Moocea) ^b Ankaramic NA 1.3 1.7.0 1.6 0.704621 0.5127 (Moocea) ^b Kamples were collected at Oft. Island, American Senoa in August of 2004, All ³ He/ ⁴ He ratios are reported relative to a pectrometry and associated measurement precision are reported in Hart and Bluschin (2006), Sr, Nd and Pb-isotope an underbeic whole rock powders by JCP at the Washingon Sine University GeoAmbAhal Lab.	816 19.176	15.581	39.264	5.02	723	4.07	9.09	99.2	4.8	0.40
DFU-04-15 Olivine Ankaramire Boulder 14.3 29.6 0.2 0.704559 0.512822 19.141 15.580 39.169 3.89 601 4.02 9.28 88.0 4.8 0. DFU-04-15 fits Olivine Ankaramire Boulder 3.2 21.2 0.2 2 21.2 0.2 2 21.2 0.2 2 2 21.2 0.2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2 2	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	$ \begin{array}{c ccccc} \text{JU-04-15} & \text{Olivnia} & \text{Ankaramire} & \text{Boulder} & 14.3 & 29.6 & 0.2 & 0.704559 & 0.512822 & 19.141 & 15.580 & 39.169 & 3.89 & 601 & 4.02 & 9.28 & 88.0 & 4.8 & 0.43 \\ \text{JU-04-15} & \text{Olivnia} & \text{Ankaramire} & \text{Boulder} & 3.2 & 112 & 0.2 & 3.19 & 101 & 15.587 & 39.276 & 3.54 & 6.9 & 3.96 & 8.6 & 3.8 & 0.48 \\ \text{JU-04-15} & \text{Orivnia} & \text{Ankaramire} & \text{Boulder} & 3.2 & 102 & 19 & 1 & 70.4438 & 0.512844 & 19.16 & 15.567 & 39.276 & 3.54 & 6.9 & 3.96 & 8.6 & 3.8 & 0.48 \\ \text{JU-04-17} & \text{Orivnia} & \text{Ankaramire} & \text{Boulder} & 10 & 2.64 & 0.4 & 0.704488 & 0.512844 & 19.16 & 15.563 & 39.276 & 3.54 & 6.9 & 3.96 & 8.6 & 1.36 & 3.0 & 4.8 & 5.5 & 0.50 \\ \text{JU-04-17} & \text{Orivnia} & \text{Gabbon} & \text{Boulder} & 10 & 2.64 & 0.4 & 0.704488 & 0.512844 & 19.16 & 15.563 & 3.93 & 3.11 & 3.75 & 11.14 & 2.704 & 8.5 & 0.50 \\ \text{JOOOU-10} & \text{Olivnia} & \text{Ankaramire} & \text{NA} & 1.3 & 1.70 & 1.6 & 0.704621 & 0.512744 & 19.216 & 15.563 & 3.89 & 330 & 3.11 & 2.704 & 8.5 & 0.50 \\ \text{JOOOCORD}^{\text{ID}} & \text{Ankaramire} & \text{NA} & 1.3 & 1.70 & 1.6 & 0.704621 & 0.512744 & 19.216 & 15.563 & 3.89 & 331 & 3.15 & 3.11.4 & 2.704 & 8.5 & 0.50 \\ \text{JOOOCORD}^{\text{ID}} & \text{Ankaramire} & \text{NA} & 1.3 & 1.70 & 1.6 & 0.704621 & 0.512774 & 19.216 & 15.563 & 3.89 & 331 & 3.15 & 3.11.4 & 2.704 & 8.5 & 0.50 \\ \text{JOOOCORD}^{\text{ID}} & \text{Ankaramire} & \text{NA} & 1.3 & 1.70 & 1.6 & 0.704621 & 0.512774 & 19.216 & 15.563 & 3.89 & 331 & 3.11.4 & 2.704 & 8.5 & 0.50 \\ \text{JOOOCORD}^{\text{ID}} & \text{Ankaramire} & \text{NA} & 1.3 & 1.70 & 1.6 & 0.704621 & 0.512774 & 19.216 & 15.563 & 3.80 & 3.60 & 4.6 & 5.5 & 0.50 \\ \text{JOOOCORD}^{\text{ID}} & \text{Ankaramire} & \text{NA} & 1.3 & 1.70 & 1.6 & 0.704621 & 0.512774 & 19.216 & 15.563 & 3.80 & 3.80 & 3.60 & 4.6 & 5.5 & 0.50 \\ \text{Aloxeral}^{\text{ID}} & \text{Ankaramire} & \text{Ankaramire} & \text{NA} & 1.8 & 1.007 & 0.019 & \text{NA} & 1.6 & 0.704621 & 0.512774 & 19.216 & 15.563 & 0.660 & 0.70666 & 0.706666 & 0.706666 & 0.706666 & 0.706666 & 0.706666 & 0.706666 & 0.706666 & 0.706666 & 0.706666 & 0.706666 & 0.706666 & 0.706666 & 0.706666 & 0.70666 & 0.706666 & 0.$	$ \begin{array}{c ccccc} \operatorname{DFU}(-4+15 & \operatorname{O}(\operatorname{krine} \ \operatorname{Ankaramic} \ \operatorname{Boulder} \ 14.3 & 29.6 & 0.2 & 0.704559 & 0.512822 & 19.141 & 15.580 & 39.169 & 3.89 & 601 & 4.02 & 9.28 & 88.0 & 4.8 & 0.43 \\ \operatorname{DFU}(-4+15 & \operatorname{CW} & \operatorname{Alkainamic} \ \operatorname{Boulder} \ 3.2 & 212 & 0.2 & \\ \operatorname{DFU}(-4+17 & \operatorname{CW} & \operatorname{Alkainamic} \ \operatorname{Boulder} \ 3.2 & 212 & 0.2 & \\ \operatorname{DFU}(-4+17 & \operatorname{O}(\operatorname{krine} \ \operatorname{Ankaramic} \ \operatorname{Boulder} \ 10 & 2.2 & 212 & 0.4 \\ \operatorname{DFU}(-4+17 & \operatorname{O}(\operatorname{krine} \ \operatorname{Alkainamic} \ \operatorname{Boulder} \ 10 & 2.6 & 0.4 & 0.704438 & 0.512834 & 19.216 & 15.587 & 39.176 & 3.34 & 6.93 & 3.96 & 8.6 & 12.96 & 5.3 & 0.46 \\ \operatorname{DFU}(-4+17 & \operatorname{O}(\operatorname{krine} \ \operatorname{Ankaramic} \ \operatorname{Boulder} \ 10 & 2.6 & 0.4 & 0.704438 & 0.512344 & 19.16 & 15.567 & 39.175 & 3.14 & 3.91 & 31.4 & 2.704 & 8.5 & 0.50 \\ \operatorname{Aloxea})^{\mathrm{b}} & \operatorname{Ankaramic} \operatorname{Ankaramic} \operatorname{NA} & 1.3 & 1.70 & 1.6 & 0.704638 & 0.512744 & 19.216 & 15.603 & 38.933 & 2.11 & 381 & 3.75 & 11.14 & 2704 & 8.5 & 0.50 \\ \operatorname{Aloxea})^{\mathrm{b}} & \operatorname{Ankaramic} \operatorname{Ankaramic} \operatorname{NA} & \operatorname{Ankaramic} \operatorname{NA} & \operatorname{Ankaramic} \operatorname{NA} & 1.3 & 1.70 & 1.6 & 0.704638 & 0.512744 & 19.216 & 15.603 & 38.933 & 2.11 & 381 & 3.75 & 11.14 & 2704 & 8.5 & 0.50 \\ \operatorname{Aloxea})^{\mathrm{b}} & \operatorname{Ankaramic} \operatorname{Ametican} \operatorname{Samoa} \operatorname{in} \operatorname{August} \operatorname{of} 2.004, \operatorname{All}^{\mathrm{He}}/\mathrm{He} \operatorname{ratios} \operatorname{are} \operatorname{or} \operatorname{Porker} + 19.216 & 15.603 & 38.933 & 2.11 & 381 & 3.75 & 11.14 & 270.4 & 8.5 & 0.50 \\ \operatorname{Aloxea} \operatorname{Ankaramic} \operatorname{Ametican} \operatorname{Samoa} \operatorname{in} \operatorname{Ametican} \operatorname{Ametican} \operatorname{Ametican} \operatorname{Samoa} \operatorname{in} \operatorname{Ametican} \operatorname{Samoa} \operatorname{Ametican} \operatorname{Ametican} \operatorname{Ametican} \operatorname{Ametican} \operatorname{Ametican} \operatorname{Ametican} \operatorname{Ametican} \operatorname{Ametican} \operatorname{Ametican} Ame$	$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	JFU-04-15 Olivine Ankaramire Boulder 14.3 29.6 0.2 0.704559 0.5123 JFU-04-15 fits Olivine Ankaramire Boulder 14.3 29.6 0.2 0.704559 0.5123 JFU-04-16 cpx Ankaramire Boulder 0.2 19 1 0.704438 0.5123 JFU-04-17 Olivine Gabbon Boulder 1.0 26.4 0.4 0.704438 0.5123 (4001-01 Olivine Gabbon Boulder 1.0 26.4 0.4 0.704438 0.5123 (4000-01 Olivine Ankaramire NA 1.3 1.7,0 1.6 0.704621 0.5127 (Adoorea) ^b (Adoorea) ^b (Adoorea) ^b (Adoorea) ^c (Adoorea) ^c	819 19.126	15.584	39.164	2.95	292	4.39	9.88	280.2	4.9	0.37
DFU-04-15 fits Olivine Ankaranite Boulder 3.2 21.2 0.2 " DFU-04-16 cpx Akail Basalt Boulder 0.2 19 1 0.704438 0.512834 19.216 15.587 39.276 3.54 639 3.96 8.60 1296 5.3 0. DFU-04-17 Olivine Gabbox Boulder 1.0 2.64 0.4 0.704498 0.512844 19.160 15.576 39.173 2.18 310 4.14 7.30 118.2 51 0. MO01-01 Olivine Ankaranite NA 1.3 17.0 1.6 0.704621 0.512774 19.216 15.603 38.933 2.11 381 3.75 11.14 2704 8.5 0. (Moorea) ^b	DEU-04-15 fits Olivine Ankaramie Boulder 3.2 21.2 0.2 " " " " " " " " " " " " " " " " " " "	DFU-04-15 fits Olivine Ankaramic Boulder 3.2 21.2 0.2 "	DFU-04-15 fits Olivine Ankaramite Boulder 3.2 21.2 0.2 " " " " " " " " " " " " " " " " " " "	DFU-04-15 fits Olivine Ankaramite Boulder 3.2 21.2 0.2 " " " " " " " " " " " " " " " " " " "	DFU-04-15 fits Olivine Ankaramire Boulder 3.2 21.2 0.2 " " " " " " " " " " " " " " " " " " "	DFU-04-15 fits Olivine Ankaramic Boulder 3.2 21.2 0.2 " </td <td>DFU-04-15 fits Olivine Ankaramic Boulder 3.2 21.2 0.2 " " DFU-04-16 opx Alkal Basalt Boulder 0.2 119 1 0.70-438 0.5123 DFU-04-17 Olivine Gabbro Boulder 1.0 26.4 0.4 0.70-438 0.5123 DFU-04-17 Olivine Gabbro Boulder 1.0 26.4 0.4 0.70-438 0.5123 (Moorea)^b Ankaramic NA 1.3 17.0 1.6 0.70-621 0.5123 (Moorea)^b Ankaramic NA 1.3 17.0 1.6 0.70-621 0.5123 (Moorea)^b Samples were collected at Oft Island, American Samoa in August of 2004, All ³He³/He ratios are reported reletive to a pectrometry and associated measurement precision are reported in Hart and Bhaschip (2006), Si, Nd and Pb-isotope an indexbed whole rock owders by ICP at the Wasinggor Sine University GeoAmbytral Lab. The ratio are apprecisioned and measured nore than once, the crush experiment is labeled as "at" ^a Required a blank correction of 46% due to low gas concentrations. The ³He³/He³ He ratio on chis clinopyroxere measu ^a</td> <td>822 19.141</td> <td>15.580</td> <td>39,169</td> <td>3.89</td> <td>601</td> <td>4.02</td> <td>9.28</td> <td>88.0</td> <td>4.8</td> <td>0.43</td>	DFU-04-15 fits Olivine Ankaramic Boulder 3.2 21.2 0.2 " " DFU-04-16 opx Alkal Basalt Boulder 0.2 119 1 0.70-438 0.5123 DFU-04-17 Olivine Gabbro Boulder 1.0 26.4 0.4 0.70-438 0.5123 DFU-04-17 Olivine Gabbro Boulder 1.0 26.4 0.4 0.70-438 0.5123 (Moorea) ^b Ankaramic NA 1.3 17.0 1.6 0.70-621 0.5123 (Moorea) ^b Ankaramic NA 1.3 17.0 1.6 0.70-621 0.5123 (Moorea) ^b Samples were collected at Oft Island, American Samoa in August of 2004, All ³ He ³ /He ratios are reported reletive to a pectrometry and associated measurement precision are reported in Hart and Bhaschip (2006), Si, Nd and Pb-isotope an indexbed whole rock owders by ICP at the Wasinggor Sine University GeoAmbytral Lab. The ratio are apprecisioned and measured nore than once, the crush experiment is labeled as "at" ^a Required a blank correction of 46% due to low gas concentrations. The ³ He ³ /He ³ He ratio on chis clinopyroxere measu ^a	822 19.141	15.580	39,169	3.89	601	4.02	9.28	88.0	4.8	0.43
DFU-04-16 cpx Alkali Basalt Boulder 0.2 19 1 0.704438 0.512834 19.216 15.587 39.276 3.54 639 396 8.60 1296 5.3 0. DFU-04-17 Olivine Gabboo Boulder 1.0 2.6.4 0.4 0.704498 0.512344 19.160 15.576 39.173 2.18 310 4.14 7.30 118.2 51 0. MO01-01 Olivine Ankaranite NA 1.3 1.7.0 1.6 0.704621 0.512774 19.216 15.603 38.933 2.11 381 3.75 11.14 2704 8.5 0. 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All the far cauos are reported relative to a spectrometry and associated measurement prevision are reported in Hart and Bluszhin (2006). Si; Nd and Po-isotope an inteached whole nork powders by ICP at the Washington State University GeoAnalytical Lab. (Errors associated with trace effects and be found in Ackson et al (2010). Only two samples (OFU-04-06 fits and y crushing. When an olivine sample was curshed and measured more than once, the crush experiment is labeled as "cri ³ Required a black correction of 46% due to low gas concentrations. The 'He/He ratio on this clinopyroxere measu	A CHI	and a state	when the state	1.5	1000	0-0-0	N 10 11	IC STORE	the second second	100

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reproducible results. The reproducibility of the measurements for Ofu samples, the relatively high helium
concentrations, the absence of higher ³He/⁴He ratios
on melting, and shielding of many of the samples all
indicate that cosmogenic ³He is not a factor in generating
the remarkably high ³He/⁴He ratios.

The Sr, Nd, and Pb chemistry, mass spectrometry, associated measurement precision and standard normalizations are reported in Hart and Blusztajn (2006) and references therein. Basalt Sr, Nd and Pb-isotope analyses were performed on the same powder following 1 h of leaching in 6.2 N HCl at 100 °C. Sr and Nd chemistry



was done with conventional ion chromatography using 184 DOWEX 50 cation resin and HDEHP-treated teflon for 185 Nd separation (Taras and Hart, 1987). An HBr-HNO3 186 procedure (Galer, in press; Abouchami et al., 1999) was 187 used for Pb chemistry, with a single column pass. Sr, Nd, 188 and Pb-isotope analyses were done on the NEPTUNE 189 multi-collector ICP-MS at WHOI. Internal precision is 190 5–10 ppm (2 σ) for Sr and Nd-isotope measurements. 191 Adjusting to 0.710240 (SRM987 Sr standard) and 192 0.511847 (La Jolla Nd standard) gives an estimated 193 external precision for Sr and Nd of 15–25 ppm (2σ). 194 The internal precision for Pb-isotope ratios is better 195 than 15-30 ppm, and using SRM997 Tl as an internal 196 standard, the external reproducibility following full 197 chemistry ranges from ~ 20 ppm (2 σ) for 207 Pb/ 206 Pb 198 to ~120 ppm (2 σ) for ²⁰⁸Pb/²⁰⁴Pb (Hart and Blusztajn, 199 2006). Pb-isotope ratios are adjusted to the SRM 981 200 values of Todt et al. (1996).

Ofu lavas exhibit slight variations in the radiogenic 202 isotope ratios of 87 Sr/86 Sr (0.704438-0.704795), 203 143Nd/144Nd (0.512800 to 0.512844) and 206Pb/204Pb 204 (19.126 to 19.257) (Fig. 1 and Table 1). Together with 205 Ofu, data from Samoan lavas in general define a wedge- 206 like shape in 3He/4He-87Sr/86Sr isotope space: lavas with 207 the highest (more enriched) 87 Sr/86 Sr ratios exhibit low 208 ³He/⁴He ratios and samples with the least radiogenic 209 ⁸⁷Sr/⁸⁶Sr ratios, found primarily at Ofu, are associated 210 with the highest ³He/⁴He ratios (Fig. 1). Similarly, 211 Samoan lavas with the most enriched (lowest) 143 Nd/ 212 144Nd have low 3He/4He ratios. However, Nd and Pb 213 isotopes measured in Ofu lavas do not define end-member 214 values for Samoa, but instead fall toward the upper end 215 found in the hotspot. 216

In order to further constrain the nature of elevated 217 ³He/⁴He lavas associated with EM2 hotspots, we present 218 Sr, Nd and Pb isotopes for a high ³He/⁴He sample from 219

Fig. 1. The new Ofu data indicate that the Samoan high ³He/⁴He lavas have more enriched 87 Sr/86 Sr and 143 Nd/144 Nd than the highest 3He/4He lavas from Hawaii, Iceland and Galapagos (abbreviated HIG). The field of Samoan shield basalts is shaded grey, and the field of Samoan lavas from Ofu Island is shaded green (see Table 1). HIG lavas trend to a more isotopically depleted high 3He/4He component than lavas from Ofu. A single sample from Baffin Island is an exception to the Sr and Ndisotope separation between the Ofu and HIG lavas, but was indicated by Stuart et al. (2003) to be crustally contaminated. Data sources are summarized in Graham (2002) or noted in Table 2. In the isotope projections shown in this figure, the separation between the highest ³He/⁴He samples from the northern and southern hemisphere high ³He/⁴He hotspots is not perfect. For example, the Pb-isotopes of the highest ³He/⁴He lavas from Samoa and the Galapagos overlap. However, the isotopic and hemispheric separation of FOZO-A and FOZO-B samples (Table 2) is observed in the isotope projections shown in Fig. 2. (For interpretation of the references to color in this figure legend, the reader is referred to the web version of this article.)



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Fig. 2. The highest ${}^{3}\text{He}^{4}\text{He}$ lavas from northern hemisphere (FOZO-B) hotspots separate from the highest ${}^{3}\text{He}^{4}\text{He}$ samples from southern hemisphere (FOZO-A) hotspots in a.) ${}^{87}\text{Sr}{}^{86}\text{Sr}{}^{-143}\text{Nd}{}^{144}\text{Nd}$ space, b.) the mantle tetrahedron, c.) a "helium balloon" pinplot showing ${}^{87}\text{Sr}{}^{86}\text{Sr}{}^{-143}\text{Nd}{}^{144}\text{Nd}{}^{-3}\text{He}^{4}\text{He}$ isotope spaces. The criteria for selecting high ${}^{3}\text{He}{}^{4}\text{He}{}$ lavas with Sr-Nd-Pb isotopes most representative of the OIB mantle are described in Table 2 and Sections 3 and 4.1 of the text. The EDR and the non-chondritic BSE are described in Fig. 6. Although the Baffin Island lava (square marked with a "B") is crupted in a continental setting, and is thus excluded from the compilation of high ${}^{3}\text{He}{}^{4}\text{He}{}$ lavas (see Table 2), it hosts the highest magmatic ${}^{3}\text{He}{}^{4}\text{He}{}$ on record and is included in the figure for reference. The highest ${}^{3}\text{He}{}^{4}\text{He}{}$ lavas (onsistent with earlier observations (Hart et al., 1992). Compared to other southern hemisphere high ${}^{3}\text{He}{}^{4}\text{He}{}$ lavas, Samoan high ${}^{3}\text{He}{}^{4}\text{He}{}$ lavas Sort exhibit anomalous Sr, Nd and Pb isotopes.

220 Moorea island in the Societies hotspot. The high ³He/⁴He

sample from Moorea, MO01-01, has a ³He/⁴He ratio of

- 222 17.0 Ra (Hanyu and Kancoka, 1997), and has Sr, Nd and
- 223 Pb-isotope ratios similar to Ofu basalts (Table 1).

224 3. Ofu in a global context

The ⁸⁷Sr/⁸⁶Sr and ¹⁴³Nd/¹⁴⁴Nd compositions of Ofu lavas show that the high ³He/⁴He reservoir sampled by the Samoan hotspot is more enriched than the high 3 He/ 4 He 227 reservoir sampled by HIG lavas. Data in Fig. 1 show a 228 clear "valley" between the highest 3 He/ 4 He Samoan and 229 HIG lavas; samples with 3 He/ 4 He greater than ~ 22 Ra 230 have not been found in this gap. The Ofu samples have 231 87 Sr/ 86 Sr ratios greater than 0.7044, while the high 232 3 He/ 4 He HIG lavas have 87 Sr/ 86 Sr less than 0.7038; 233 a single Baffin Island sample, which is considered by 234 Stuart et al. (2003) to be affected by contamination from 235

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9.1	Table 7
164 - L	Table 2

t2.2

6

Summary of the radiogenic isotope data for the highest ³He/⁴He samples used to define FOZO-A and FOZO-B

t2.3	Hotspot	Island/ seamount/ region	ID of sample with highest ³ He/ ⁴ He used	3 He/ 4 He ($\pm 1\sigma$)	⁸⁷ Sr/ ⁸⁶ Sr	¹⁴³ Nd/ ¹⁴⁴ Nd	²⁰⁶ Pb/ ²⁰⁴ Pb	²⁰⁷ Pb/ ²⁰⁴ Pb	²⁰⁸ Pb/ ²⁰⁴ Pb	Epsilon ¹⁴³ Nd
t2.4	FOZO-A									
t2.5	Samoa ^a	Ofu	Ofu-04-06	33.8 ± 0.2	0.704584	0.512827	19.189	15.571	39.202	4.2
t2.6	Societies ^a	Moorea	MO-01-01	17.0 ± 1.6	0.704621	0.512774	19.216	15.603	38.933	3.2
t2.7	Macdonald ^a	Macdonald	SO-47 64DS2	15.8 ± 0.3	0.703755	0.512777	19.46	15.61	39.27	3.2
t2.8	Juan Fernandez ^{b, c}	Mas a Tierra	PIN-8	17.2 ± 0.3	0.70363	0.512840	19.13	15.61	38.98	4.5
t2.9	Reunion ^b	Reunion	VP1931	13.6 ± 0.3	0.704130	0.51284	18.88	15.59	38.96	4.5
t2.10	Kerguelen ^d	Heard	HB24A	18.3 ± 0.5	0.704862	0.512705	18.776	15.588	39.170	1.8
t2.11	Bouvet ^b	Bouvet	WJ8B	12.4 ± 1	0.70371	0.51284	19.588	15.653	39.243	4.5
t2.12	Marquesas ^a	Hiva Oa	HO-AT-1	14.4 ± 1.4	0.70471	0.512834	19.017	15.570	38.823	4.4
t2.13										
t2.14	FOZO-B									
t2.15	Hawaii ^{a, c}	Loihi	KK18-8	32.3 ± 0.6	0.703680	0.512945	18.448	15.477	38.189	6.5
t2.16	Iceland ^a	Westfjords	SEL-97	37.7±2.0	0.703465	0.512969	18.653	15.473	38.453	7.0
t2.17	Galapagoes ^a	Fernandina	NSK 97-214	$30.3{\pm}0.2$	0.703290	0.512937	19.080	15.537	38.711	6.4
t2.18	Cape Verde ^b	San Nicolau	SN-10	15.7 ± 0.1	0.703050	0.51296	19.538	15.586	38.964	6.7
t2.19	Azores ^{a, c}	Terciera	T2	11.3 ± 0.8	0.703520	0.512960	19.883	15.630	39.310	6.8

If Sr, Nd and Pb isotopes are unavailable for a sample with high 3 He/ 4 He, they are estimated based on the criteria below and the estimated values are t2.20 italicized in the table.

High ³He⁴He lavas erupted at deep submarine ridge environments are excluded, and include the following locations as summarized in Graham (2002): Southeast Indian ridge near Amsterdam and St. Paul (14.1 Ra), east and west rifts of the Easter microplate (11.7 Ra), Gulf of Tadjoura near Afar (14.7 Ra), southern Mid-Atlantic ridge near Shona and Discovery (12.3 and 15.2 Ra, respectively), Southwest Indian Ridge near Bouvet Island (14.9 Ra) and the Manus Basin back-are spreading center (15.1 Ra). However, sub-aerially erupted high ³He⁴He lavas sampled at ridge-centered hotspots, including Bouvet Island (Kurz et al., 1998) and Iceland (Hilton et al., 1999), are included in the dataset. Additionally, high ³He⁴He lavas erupted in continental settings are not considered, and include: Baffin Island (Stuart et al., 2003) and West Greenland picrites (Graham et al., 1998), Yellowstone (Graham et al., 2006) and Afar (Scarsi and Craig, 1996). Finally, high ³He⁴He lavas erupted in back-are environments, such as

t2.21 Rochambeau Bank (Poreda and Craig, 1992) and Manus Basin (Macpherson et al., 1998), are excluded from our treatment of the high ³He⁴He mantle. ^a Many of the high ³He⁴He lavas in this table have a full complement of lithophile radiogenic isotope analyses measured on the same samples, and are from the following locations: The Galapagos (Kurz and Geist, 1999; Saal et al., 2007), Hawaii (Kurz et al., 1983; Staudigel et al., 1984), Samoa (Table 1), Iceland (Hilton et al., 1999), Societies (Hanyu and Kaneoka, 1997; Table 1), Marquesas (Castillo et al., 2007), Macdonald

t2.22 (Hémond et al., 1994; Moreira and Allègre, 2004) and Azores (Turner et al., 1997; Moreira et al., 1999).
^b A full complement of Sr, Nd and Pb-isotope data do not exist for all the high ³He/⁴He samples listed in the table above. Values for the "missing" lithophile radiogenic isotopes (data in italics) are generated based on the following criteria. The high ³He/⁴He Cape Verde lava has Sr and Pb isotopes but lacks Nd isotopes (Doucelance et al., 2003). Sr and Nd-isotope data correlate well in the Cape Verde islands, and the Nd-isotope ratio for the Cape Verde sample is made by regression through existing data. The high ³He/⁴He Bouvet Island sample (Kurz et al., 1998) also has Sr and Pb-isotope data, but lacks a Nd-isotope measurement. The missing Nd isotopic value is estimated by averaging existing Bouvet island data from O'Nions et al. (1977). A high ³He/⁴He sample from Reunion with Sr isotope data (Graham et al., 1990) lacks Nd and Pb-isotope data; Reunion lavas are isotopically uniform, and the GEOROC database (http://georoc.mpch-mainz.gwdg.de/georoc/) was used to estimate the Nd and Pb isotope ratios (0.51284) in the Juan Fernandez sample exhibits one of the least radiogenic Nd-isotope ratios (0.51284) in the Juan Fernandez suite, so Pb-isotope compositions (²⁰⁶Pb/²⁰⁴Pb from 19.045 to 19.214) from three other Juan Fernandez samples (Gerlach et al., 1986) with

t2.23 the least radiogenic Nd isotopes (0.51282-0.51284) were averaged to generate a Pb-isotope composition for the high ³He⁴He lava. ⁶ When the ³He⁴He ratios quoted above are lower than the maximum from a hotspot, it is because Sr, Nd and/or Pb-isotope ratios could not be estimated reliably for the sample, or because higher reported ³He⁴He ratios were not measured as precisely (or were suspected of having a cosmogenic ³He influence). For example, a ³He⁴He value of 14.8 Ra was reported for an Azores lava (Madureira et al., 2005), but lithophile radiogenic isotopes were not reported and a cosmogenic ³He influence cannot be ruled out for this sample. Similarly, a value of up to 35.3 Ra has been reported from Loihi t2.24. (Valbracht et al., 1997), but analytical uncertainties on this measurement were large and lithophile isotopes were not reported.

t2.24 (Valbracht et al., 1997), but analytical uncertainties on this measurement were large and lithophile isotopes were not reported. ^d The Sr, Nd and Pb isotopes of the Heard sample (Barling and Goldstein, 1990; Hilton et al., 1995) were not measured on the same sample as the

t2.25 ³He^AHe, but they were measured on a sample (69244) from the same flow.

continental crust, is an exception to the separation of Ofu
 and HIG lavas. Ofu's trend to a more enriched region
 of Nd-isotope space than that sampled by the HIG lavas
 is also clear (¹⁴³Nd/¹⁴⁴Nd<0.51285 in Ofu versus ¹⁴³Nd/
 ¹⁴⁴Nd>0.51290 in HIG lavas). Previous work suggested

that the Samoan data trends toward a similar high 3 He/ 4 He 241 component as the HIG lavas (Farley et al., 1992; Hart 242 et al., 1992). The new isotope data argue against this 243 hypothesis because Ofu has 3 He/ 4 He as high as HIG but 244 with more enriched 87 Sr 86 Sr and 143 Nd/ 144 Nd. 245

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4. Discussion

4.1. FOZO-A and FOZO-B: some important caveats

isotope heterogeneity of the high 3He/4He mantle reser-247 voir (Fig. 2), we have compiled a dataset of Sr, Nd and Pb 248 isotopes for lavas that are representative of the high 249 ³He/⁴He OIB mantle. We include only the hotspots that 250 have ³He/⁴He>11 Ra, thereby excluding hotspots such as 251 Tristan, Gough, and St. Helena, etc., that do not sample a 252 component from the high ³He/⁴He mantle. We also limit 253 our discussion to just the single highest ³He/⁴He sample 254 from each high ³He^AHe hotspot, in an attempt to define 255the Sr-Nd-Pb-isotope composition of the highest ³He/ 256 ⁴He reservoir sampled by a hotspot. Finally, high ³He/⁴He 257 lavas erupted at continental, back-arc and deep submarine 258 ridge environments are excluded (see Table 2 for a list of 259 specific samples from these three environments). 260

In order to place constraints on the Sr, Nd and Pb-

246

261 Using the above criteria, Table 2 lists the high ³He/⁴He samples that best represent the Sr, Nd and Pb 262 compositions of the high ³He/⁴He OIB mantle reservoir. 263 Combined with the new Ofu data, the high ³He/⁴He OIB 264 dataset in Table 2 suggests the existence of two iso-265 topically distinct high ³He/⁴He reservoirs in the mantle. 266 These two high ³He/⁴He reservoirs separate in the 267 northern (boreal) and southern (austral) hemispheres 268 (Fig. 2). The highest ³He/⁴He samples from each of the 269 southern hemisphere hotspots, Macdonald seamount, 270 Bouvet, Kerguelen, Juan Fernandez, Societies, Reunion 271 and Marquesas appear to sample the more isotopically 272 enriched (or less isotopically depleted, see Section 4.5), 273 high 3He/4He mantle component found in Ofu lavas. By 274 contrast, HIG lavas and two other northern hemisphere 275 hotspots-Cape Verde and Azores-sample a more iso-276 topically depleted, high ³He/⁴He component than 277 observed in southern hemisphere high ³He/⁴He lavas. 278 The highest ³He/⁴He samples from the boreal and austral 279 high ³He/⁴He hotspots separate toward the depleted 280 and enriched ends, respectively, of the global 87 Sr/86 Sr 281

¹⁴³Nd/¹⁴⁴Nd OIB array (Fig. 2). Fig. 2 shows that there 282 are two separate 3He/4He peaks in Sr-Nd-He isotope 283 space: One peak is formed by the highest ³He/⁴He 284 northern hemisphere samples and another peak, an-285 chored by the high ³He/⁴He Ofu lava, is formed by the 286 highest ³He/⁴He southern hemisphere samples. Samples 287with the highest ³He/⁴He from each of the boreal and 288 austral FOZO hotspots also are distinguished isotopi-289 cally inside the mantle tetrahedron formed by the 290 isotopes of Sr, Nd and Pb. The boreal (FOZO-B) domain 291 plots closer to the DMM-HIMU join while the austral 292 (FOZO-A) domain plots closer to the EM1-EM2 join 293 (Fig. 2). FOZO-A and FOZO-B also clearly separate in 904 a mantle tetrahedron constructed from the isotopes of Pb 205 (a figure of the Pb-isotope tetrahedron is available as 296 supplementary data in the Appendix). 297

Unfortunately, there are limited numbers of hotspots 300 with moderately high ³He/⁴He (from 11.3-18.3 Ra, 9 301 hotspots) and very high 3He/4He (>30 Ra, 4 hotspots); 302 lithophile isotope data on lavas from many of these high 303 ³He/⁴He hotspots are also limited (see Table 2). There- 304 fore, we cannot exclude the possibility that the apparent 305 hemispheric separation is related to sparse data, and that 306 existing He, Sr, Nd and Pb-isotope data do not yet 307 adequately characterize the terrestrial mantle. Further- 308 more, the isotopic and hemispheric separation of the two 309 FOZOs is not observed in all isotope projections: In 310 ³He/⁴He vs. ²⁰⁶Pb/²⁰⁴Pb and ³He/⁴He vs. ⁸⁷Sr/⁸⁶Sr iso- 311 tope spaces (not shown), the FOZO-A and FOZO-B 312 fields overlap. However, in ³He/⁴He vs. ¹⁴³Nd/¹⁴⁴Nd- 313 isotope space the two FOZOs are completely resolved 314 (Fig. 2). 315

An additional characteristic of the "two FOZOs" 316 model is that the isotopic and hemispheric distinction is 317 not always observed in the highest ³He/⁴He lavas erupted 318 in continental, back-arc and deep submarine mid-ocean 319 ridge settings. One possible explanation is that mantle 320 plume interaction with the shallow geochemical reser- 321 voirs in these three environments can decouple the deep 322 mantle 3He/4He signatures from the associated Sr, Nd and 323 Pb isotopes. If so, then the high 3He/4He ratios measured 324 in lavas from these three settings cannot be traced to the 325 same source components as the Sr, Nd and Pb isotopes. 326 High ³He/⁴He lavas erupted in continental settings may 327 have suffered crustal assimilation such that the Sr, Nd and 328 Pb isotopes are not representative of the high ³He/⁴He 329 mantle. For example, a high ³He/⁴He Baffin Island 330 sample (43.9 Ra; Stuart et al., 2003) with the least 331 depleted 143 Nd/144 Nd in the suite-0.512730-is inferred 332 to be crustally contaminated, and is not interpreted to be a 333 northern hemisphere expression of FOZO-A. High 334 ³He/⁴He lavas erupted in back-arc environments may 335 not have lithophile isotope ratios that reflect the high 336 ³He/⁴He mantle source, owing to possible involvement of 337 slab-derived fluids that may affect the Sr, Nd and Pb 338 isotopes (even though such fluids may have a negligible 339 effect on the ³He/⁴He ratios; Macpherson et al., 1998). 340 The high ³He/⁴He signatures observed in deeply erupted 341 submarine ridge lavas may host Sr, Nd and Pb isotopes 342 that reflect entrainment of the shallow depleted mantle. 343 However, we presume that the high plume flux neces- 344 sary for the generation of a sub-aerially-exposed, ridge- 345 centered hotspot volcano enhances the plume contribution 346 in the erupted lavas, thus overwhelming the depleted 347

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³⁴⁸ upper mantle component. Therefore, sub-aerially erupted ³⁴⁹ Bouvet Island and Iceland high ³He/⁴He lavas are ³⁵⁰ included in the database (both lavas host Sr, Nd and Pb ³⁵¹ isotopes that more enriched than nearby ridges, and are ³⁵² taken to reflect a plume component).

8

How can high ³He/⁴He, deep mantle signatures be 353 preserved and become decoupled from the deep mantle 354 Sr, Nd and Pb signatures in continental, back-arc and 355 deep submarine ridge environments? As a stimulus for 356 further investigation, we suggest that the He/(Sr,Nd,Pb) 357 ratios may be higher in the high ³He/⁴He mantle 358 (or mantle melt) than in the shallow contaminating 359 reservoirs found in these three geological settings. In this 360 way, high ³He/⁴He signatures often may be little affected 361 by contamination from shallow reservoirs while Sr, Nd 362 363 and Pb isotopes in such lavas can be strongly decoupled from the original high ³He/⁴He mantle source. 364

365 4.2. FOZO-A vs. FOZO-B: is the difference sediment?

The isotopic enrichment observed in FOZO-A lavas, 366 particularly from Ofu, relative to the FOZO-B reservoir 367 is not easily reconciled with existing models for the 368 evolution of the high 3He/4He mantle. One hypothesis for 260 the generation of the relative lithophile isotope and trace 370 element enrichment in high 3He/4He lavas from Samoa 371 maintains that the high 3He/4He mantle beneath this 372 hotspot was recently contaminated by rapidly cycled sedi-373 ment from the nearby Tonga trench (Class and Goldstein, 374 375 2005). However, evidence from Pb-isotopes rules out modern marine sediment contamination of the Samoan 376 high ³He/⁴He mantle. The sediment-OIB discriminating 377 properties (Hart, 1988) of Δ^{207} Pb/²⁰⁴Pb- Δ^{208} Pb/²⁰⁴Pb-378 isotope space (Hart, 1984) show that Samoan basalts and 379 modern global marine sediments (Plank and Langmuir, 380 1998) exhibit non-overlapping fields with diverging 381 trends (Fig. 3) (Jackson et al., 2007). Pb-isotope data 382 thus preclude the presence of modern marine sediment 282 (including sediments outboard of the Tonga trench) in the 384 FOZO-A mantle sampled by Ofu lavas. 385

386 Sr, Nd and Pb-isotope compositions for the high ³He/⁴He Moorea lava (Hanyu and Kaneoka, 1997) 387 388 provide further evidence that the more enriched radiogenic isotope compositions of Samoan high 3He/4He 389 lavas (compared to HIG lavas) are not necessarily a result 390 of rapidly cycled, subducted sediment (Table 1). Moorea 391 island is not located near a subduction zone, yet the high 392 ³He/⁴He Moorea sample has ⁸⁷Sr/⁸⁶Sr, ¹⁴³Nd/¹⁴⁴Nd and 393 ²⁰⁶Pb/²⁰⁴Pb-isotope ratios similar to Ofu basalts. 394

The observation of elevated trace element concentrations in Samoan high ³He/⁴He lavas (relative to HIG high ³He/⁴He lavas) has been used as evidence for recent



Fig. 3. In Δ^{207} Pb/²⁰⁴Pb- Δ^{208} Pb/²⁰⁴Pb-isotope space (Hart 1984, 1988), Samoan basalts (including Ofu basalts) and global marine sediments exhibit non-overlapping fields with diverging trends. The Samoan data are plotted with oceanic sediments (the Tonga sediment is circled) contributing to global subducting sediment (GLOSS) from Plank and Langmuir (1998). Samoan lavas from Ofu Island plot close to the highest ³He/⁴He lavas from Hawaii, Iceland and Galapagos. Samoan xenoliths from Savai'i island are plotted as stars (Hauri et al., 1993). Other plotted Pb-isotope data include endmember MORB (average of normal ridge segments; Su, in press), EM1 (Pitcairn; Eisele et al., 2002; Hart and Hauri, unpubl. data) and HIMU (Mangaia and Tubuai; Hauri and Hart, 1993; Woodhead, 1996). The figure is adapted from Jackson et al. (2007), and data for sediments from the Samoan region (piston cores) and the ferromanganese rind are from the same source. Pb-isotope data preclude the presence of modern marine sediment (including sediment recently subducted into the Tonga trench) in the FOZO-A mantle sampled by Ofu lavas.

sediment contamination of the Samoan high ³He/⁴He 398 mantle (Class and Goldstein, 2005). However, our 399 favored explanation for the elevated trace element 400 concentrations in the Ofu (and Moorea) high ³He/⁴He 401 lavas is that they are products of low degrees of melting, 402 a mechanism that can greatly increase the trace element 408 concentrations in lavas relative to their mantle source. 404 Ofu and Moorea high ³He/⁴He lavas are alkaline, 405 a petrologic feature that is uncommon among high 406 ³He/⁴He HIG lavas and one that is likely to be a result of 407 low degrees of mantle melting. Therefore, the alkaline 408 nature of Samoan high ³He/⁴He lavas may help resolve 409 the apparent paradox of high U and Th concentrations in 410 Samoan lavas that also host high ³He/⁴He ratios (Fig. 4). 411

Although the Samoan plume exhibits a (low ³He/⁴He) 412 component with recycled sediment, the high ³He/⁴He 413 Samoan lavas from Ofu exhibit no evidence for an ancient 414 recycled sediment component. The most enriched lavas 415 from Samoa display evidence for an ancient recycled 416 sediment component (White and Hofmann, 1982; Wright 417

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Fig. 4. ³He⁴He–Th for Ofu lavas compared to other OIBs and MORBs. Ofu basalts and the high ³He⁴He basalt from Moorea (MO-01-01) exhibit higher Th concentrations than high ³He⁴He basalts from other hotspots (see Table 1). The field for global OIBs (excluding Samoa) and MORBs is from Class and Goldstein (2005). All Ofu and Moorea Th data shown are by ICP, and the Ofu Th data are not corrected for olivine or cpx fractionation. The high Th concentrations in the Ofu and Moorea high ³He⁴He basalts are interpreted as resulting from lower degrees of melting, a hypothesis that is consistent with the alkaline nature of the lavas from these two islands.

and White, 1987; Farley et al., 1992), including high 418 ⁸⁷Sr/⁸⁶Sr ratios (up to 0.7216) and low ³He/⁴He ratios 419 (<5 Ra; Jackson et al., 2007). By contrast, 87Sr/86Sr and 420 143Nd/144Nd isotopes measured in Ofu lavas are among 421 the most depleted in Samoa (Table 1). Furthermore, the 422 positive Pb anomalies and high Ba/Nb ratios that are 423 diagnostic of sediment are observed in the most iso-424 topically enriched Samoan lavas, but are absent in the 425 isotopically depleted Ofu lavas (Fig. 5). The Ba/Nb and 426 Pb/Pb* measured in the HIG (FOZO-B) high ³He/⁴He 427 lavas are indistinguishable from the Ofu lavas, providing 428 further evidence that the Ofu lavas do not host an ancient 429 recycled sediment component (Fig. 5). 430

431 4.3. Of a lavas from recycled harzburgite?

Some recent models advocate a role for ancient, 432 433 depleted, high ³He/(U+Th) harzburgitic mantle lithosphere in the formation of the high ³He/⁴He mantle 434 435 domain (Anderson, 1998; Parman et al., 2005; Heber et al., 2007). This assumes that cpx-poor melt residues 436 may have the property of increased ³He/(U+Th) relative 437 to the initial, unmelted source (Parman et al., 2005), and 438 thus may preserve high ³He/⁴He ratios over time. Trace 439 element budgets in abyssal peridotites provide a critical 440 test of the hypothesis that recycled harzburgite is involved 441 in the source of high ³He/⁴He OIB lavas. While a harz-442 burgite lithology is consistent with the cpx-poor nature of 443 the upper oceanic lithosphere, the abyssal peridotites that 444

comprise the upper oceanic lithosphere are too depleted in 445 incompatible trace elements (Workman and Hart, 2005) to 446 act as a source for the high ${}^{3}\text{He}{}^{4}\text{He}$ Ofu basalts. 447 For example, the maximum Sr concentration that can be 448 generated by extremely low degrees of aggregated frac- 449 tional melting of modern harzburgitic abyssal peridotites 450 (samples with <5% cpx, with an average reconstructed 451 bulk rock Sr concentration of ~0.02 ppm; see figure 452 available as supplementary data in the Appendix) is 453 ~5 ppm, which is far below the Sr concentrations 454 (~600 ppm) of primary Ofu lavas. Trace element budgets 455 in modern abyssal peridotites suggest that Ofu basalts are 456

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Fig. 5. ³He⁴He variation with Ba/Nb and Pb/Pb* in Samoan lavas and high ³He⁴He samples from Hawaii (Kurz et al., 1982), Iceland (Hilton et al., 1999) and the Galapagos (Kurz and Geist, 1999; Saal et al., 2007). Ba/Nb ratios in high ³He⁴He lavas from all four hotspots are low, and do not indicate input from marine sediment (GLOSS; Plank and Langmuir, 1998) or upper continental crust (UCC; Rudnick and Gao, 2003). Pb/Pb* values in high ³He⁴He lavas are also unlike GLOSS and UCC. On the other hand, the enriched (EM2) Samoan lavas have low ³He⁶He, high Ba/Nb and high Pb/Pb*, values that suggest a recycled sediment signature (Jackson et al., 2007). All trace element data were measured by ICP, but trace element data from the Iceland and Hawaii samples are unpublished. Pb/Pb[#]= Pb_N/ $\sqrt{(Ce_N \times Nd_N)}$, where N signifies normalization to primitive mantle (McDonough and Sun, 1995).

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457 not the melts of ancient, recycled (depleted) oceanic upper458 mantle lithosphere.

The base of the oceanic mantle lithosphere, which is 459 thought to have a lherzolitic composition similar to DMM 460 461 (Workman and Hart, 2005), can produce the Sr concentrations observed in Ofu lavas if it is melted at low degrees. 462 Lherzolitic melt residues also may have the property of 463 increased ³He/(U+Th) relative to the initial, unmelted 464 465 source (Parman et al., 2005). However, results from recent helium partitioning experiments yield conflicting results 466 (Heber et al., 2007; see Section 4.5 below), and suggest 467 that lherzolitic melt residues will have greatly diminished 468 ³He/(U+Th). The observation of radiogenic ¹⁸⁷Os/¹⁸⁸Os 469 in high ³He/⁴He lavas from Iceland is inconsistent with 470 a purely peridotitic origin of the high ³He/⁴He mantle 471 472(Brandon et al., in press), and suggest that helium partitioning studies based on melting pure peridotite 473 474 compositions may not be applicable to melting of the high ³He/⁴He reservoirs in the earth's mantle. 475

476 4.4. Of a lavas from chondritic primitive (undegassed)477 mantle or regassed mantle?

An alternative hypothesis suggests that high ³He/⁴He 478 Samoan lavas are derived from chondritic primitive mantle 479 (referred to as PHEM; Farley et al., 1992). However, such a 480 model fails to explain the non-chondritic 143 Nd/144 Nd and 481 ²⁰⁶Pb/²⁰⁴Pb ratios obtained in Ofu lavas: the ¹⁴³Nd/¹⁴⁴Nd 482 and ²⁰⁶Pb/²⁰⁴Pb ratios of Ofu lavas, which are 0.51283 483 484 and 19.19, respectively, are much higher than the corresponding chondritic primitive mantle values, which 485 are 0.512611 (Boyet and Carlson, 2005) and 17.675, 486 respectively. 487

Another, more recent model attempts to reconcile 488 the non-chondritic lithophile isotopes in high ³He/⁴He 489 lavas assumes that a depleted upper mantle source was 490 "regassed" by mixing with a small proportion of helium-491 rich, high ³He/⁴He chondritic mantle (Stuart et al., 2003; 492 Ellam and Stuart, 2004). This model suggested that 493 regassed, depleted mantle is in the source of the high 494 495 ³He/⁴He ratios found in basalts associated with the proto-Iceland plume (PIP). However, Ofu lavas fall well outside 496 of the array formed by PIP lavas in ³He/⁴He-¹⁴³Nd/¹⁴⁴Nd-497 isotope space (Ellam and Stuart, 2004), highlighting the 498 need for a model that includes a heterogeneous high 499 ³He/⁴He mantle. 500

501 4.5. Implications of 142 Nd/ 144 Nd for "re-enrichment" of 502 the high 3 He/ 4 He mantle

The discovery of ¹⁴²Nd/¹⁴⁴Nd ratios in accessible terrestrial rocks that are higher than chondrite (Boyet and

Carlson, 2005) has important consequences for the 505 origin and evolution of the two FOZOs. The terrestrial 506 142Nd/144Nd anomaly indicates that all measured terres- 507 trial rocks were derived from a reservoir that had 508 superchronditic Sm/Nd during the lifetime of 146Sm, as 509 imperfect mixing of nucleosynthetic material in the solar 510 nebula (Ranen and Jacobsen, 2006) does not explain the 511 terrestrial excess in 142Nd (Hidaka et al., 2003; Andreasen 512 and Sharma, 2006; Carlson et al., 2007; Wombacher and 513 Becker, 2007). For example, if Sm/Nd ratios were 514 heterogeneously distributed in the solar nebula at the 515 time of accretion, bulk silicate earth (BSE) may have 516 acquired higher Sm/Nd and thus higher time-integrated 517 142Nd/144Nd ratios than chondrites (Boyet and Carlson, 518 2006). If the assumptions for a non-chondritic BSE are 519 valid (Boyet and Carlson, 2006), then both FOZO-A and 520 FOZO-B are isotopically enriched relative to BSE (Fig. 6, 521 top panel). Alternatively, if BSE has chondritic Sm/Nd, 522 the terrestrial 142Nd/144Nd anomaly in accessible terres- 523 trial mantle rocks could have been generated by an early, 524 global terrestrial differentiation event within 30 Myr of 525 accretion (Boyet and Carlson, 2005). In this case, the 526 resulting early depleted reservoir (EDR) has superchon- 527 dritic Sm/Nd and evolves superchondritic 142Nd/144Nd (a 528 complementary hidden early enriched reservoir, or EER, 529 evolved sub-chondritic 142Nd/144Nd; Boyet and Carlson, 530 2005). The FOZO reservoirs exhibit 143 Nd/144 Nd ratios 531 that are lower (more enriched) than the minimum 532 143Nd/144Nd of the EDR (Fig. 6, bottom panel). If the 533 FOZO reservoirs (like all accessible terrestrial mantle 534 rocks) were ultimately derived from the EDR at some 535 point in earth's history, then the high ³He/⁴He FOZO 536 mantle has been re-enriched since the early differentiation 537 event. In summary, if the terrestrial 142Nd/144Nd anomaly 538 relative to chondrites is due to the decay of ¹⁴⁶Sm, the 539 observed 143Nd/144Nd ratios in FOZO-A and FOZO-B 540 lavas require that they were re-enriched relative to either a 541 non-chondritic BSE or the EDR. 549

There are a number of mechanisms by which the high 543 ³He/⁴He mantle could have been re-enriched. However, 544 the enriched material added to the high ³He/⁴He res-545 ervoirs must also have the property of preserving a high 546 ³He/⁴He signature over time. Thus, re-enrichment by 547 addition of recycled sediments seems unlikely, as lavas 548 exhibiting clear evidence of sediment recycling, such as 549 the high ⁸⁷Sr/⁸⁶Sr Samoan lavas from Samoa, exhibit 550 low ³He/⁴He (Jackson et al., 2007). On the other hand, 551 re-enrichment by addition of recycled (oceanic crust) 552 eclogite plums to a high ³He/⁴He mantle source may not 553 necessarily diminish the high ³He/⁴He signature (Bran-554 don et al., in press). As an alternative to re-enrichment by 555 eclogite addition, depleted oceanic mantle lithosphere 556

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that has been re-enriched (metasomatized) with melt may 557 serve as a source for high ³He/⁴He basalts (White, 2005). 558 However, if the re-enriching melt is a result of imperfect 559 melt extraction near a mid-ocean ridge (Workman et al., 560 2004), the ${}^{3}\text{He}/(U+Th)$ of the melt will have to be 561 similar to or higher than DMM to preserve high ³He/⁴He 562 in the recycled upper oceanic mantle lithosphere over 563 time (i.e., following subduction and isolation in the 564 lower mantle). This condition requires that the compat-565 ibility of helium is similar to or lower than U and Th 566 during peridotite melting. 567

568 However, such partitioning behavior is inconsistent with the helium partitioning results of Parman et al. 569 (2005). Recent helium partitioning results of Heber et al. 570 (2007) do suggest that He is less compatible than U and 571 Th during peridotite melting (assuming a DMM 572 lherzolite lithology from Workman and Hart (2005) 573 and U and Th partition coefficients from Kelemen et al. 574 (2004)). If the partitioning results of Heber et al. (2007) 575 are correct, then oceanic mantle lithosphere hosting 576 trapped melt may preserve high ³He/⁴He over time. 577 Thus, long-term isolation of melt-impregnated oceanic 578



mantle lithosphere may provide a viable mechanism for 579 preserving a high ³He/⁴He signature, and may have 580 potential for describing the radiogenic isotope composi-581 tions and trace element budgets observed in global high 582 ³He/⁴He lavas. 583

11

4.6. Origin of the hemispheric separation of FOZO-A 584 and FOZO-B: hints from the DUPAL anomaly 585

The observation of a southern hemisphere high 586 ³He/⁴He domain that is isotopically more enriched 587 than its northern hemisphere counterpart is reminiscent 588 of the DUPAL anomaly (Hart, 1984)–a globe-encircling 589 feature of isotopic enrichment observed primarily in the 590 southern hemisphere mantle–and indicates a long-term 591 separation of the earth's northern and southern hemi-592 sphere high ³He/⁴He mantle. While there are a number of 593 processes by which the re-enriched high ³He/⁴He mantle 594 can be generated, the mechanisms responsible for the 595 hemispheric separation, and long-term preservation, of 596 two isotopically distinct FOZO reservoirs are elusive. 597

Like the high ³He/⁴He reservoir (Kurz et al., 1982; 598 Hart et al., 1992) the DUPAL anomaly was suggested to 599 be an ancient feature residing in the lower mantle (Hart, 600 1988; Castillo, 1988). The DUPAL anomaly is the only 601

Fig. 6. Implications of terrestrial ¹⁴²Nd/¹⁴⁴Nd anomalies for the 143 Nd/144 Nd evolution of BSE and the origin of the two FOZOs. If the observed superchondritic terrestrial mantle 142Nd/144Nd ratios (Boyet and Carlson, 2005) were generated by ¹⁴⁶Sm decay, either 1.) (top panel) the earth accreted from non-chondritic material and BSE has superchondritic Sm/Nd, or 2.) (bottom panel) BSE is chondritic but underwent an early depletion event and all available terrestrial mantle rocks derive from an early depleted reservoir (EDR) with superchondritic Sm/Nd. All calculations are after Boyet and Carlson (2005, 2006), and assume a chondrite average ¹⁴⁷Sm/¹⁴⁴Nd of 0.1948, 143 Nd/144 Nd of 0.512611, a solar system initial 146 Sm/144 Sm of 0.008, and an identical age of 4.567 Ga for the accretion time of the earth (top panel) and for the early differentiation event (bottom panel). Given these assumptions, the non-chondritic BSE and the EDR require 147 Sm/ 144 Nd ratios of ~0.209 to generate the 142 Nd/ 144 Nd anomalies that are 20 ppm higher than chondrite (later formation times for the accretion of the earth, or for the early differentiation event, require even higher Sm/Nd ratios). Thus, the lowest present-day 143 Nd/144 Nd value for a non-chondritic BSE and the EDR is 0.51304 (ε^{143} Nd=+8.4). Both FOZOs exhibit 143Nd/144Nd values lower than the non-chondritic BSE (top panel) and the EDR (bottom panel), and both FOZOs thus exhibit evidence for re-enrichment. Starting at 3 Ga, the DMM reservoir evolved by continuous extraction of continental crust (Workman and Hart, 2005) from the non-chondritic BSE (top panel) or from the EDR (bottom panel); DMM calculations are identical to Boyet and Carlson (2006). The trajectories of CC and EER are estimated. CC is not extracted as a single event, but is continuously extracted after 3 Ga. The MORB field is from Boyet and Carlson (2005) and FOZO-A and FOZO-B are from Table 2; the Baffin Island (Baff. Is.) high ³He/⁴He lava (Stuart et al., 2003) is plotted for reference. $\varepsilon_{Nd}(t) = [^{143}Nd/$ $^{144}Nd_{Sample}(t)/^{143}Nd/^{144}Nd_{Chondrite}(t) - 1] \times 10^4$.



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other mantle domain suggested to occupy similar, 602 hemispheric proportions. The DUPAL anomaly shows 603 that the southern hemisphere exhibits generally more 604 isotopically enriched mantle domains than the northern 605 606 hemisphere, and we find a similar southern hemisphere enrichment in the high 3He/4He FOZO-A reservoir; the 607 depletion in the northern hemisphere DUPAL reference 608 reservoir is mirrored in the FOZO-B reservoir. The 609 geographic and geochemical similarities between the 610 FOZO and DUPAL reservoirs suggest that their origin 611 may have been linked. 612

One hypothesis for the formation of the DUPAL 613 domain is that oceanic mantle and crustal lithosphere were 614 injected into the mantle, thereby enriching this region 615 616 of the mantle. This model has the advantage of being 617 compatible with two of the possible re-enrichment mechanisms suggested for the FOZO reservoirs in Section 618 619 4.5: 1.) oceanic crustal eclogite plums were added to the depleted high ³He/⁴He mantle via subduction, or 2.) re-620 enriched (metasomatized by melt near a mid-ocean ridge) 621 oceanic mantle lithosphere was subducted and isolated in 622 the (deep?) mantle and became the high ³He/⁴He mantle 623 domain. These injection models may produce a random 624 pattern of isotopic enrichment throughout the mantle, with 625 no coherent hemispheric pattern (Hart, 1984). However, 626 if injection were focused around the perimeter of a 627 supercontinent during a period of anomalous subduction, 628 a hemispheric pattern in the FOZO and DUPAL reservoirs 629 might emerge. This model would suggest that the 630 631 hemispheric and isotopic separation may be a surviving artifact of the paleo-arrangement of the subduction zones 632 and continents during the formation of the FOZO reser-633 voirs. While the timing of the formation of the FOZO 634 reservoirs is unknown, we note that the Northern 635 Hemisphere Reference Line (NHRL) separates the two 636 FOZOs in ²⁰⁷Pb/²⁰⁴Pb ²⁰⁶Pb/²⁰⁴Pb-isotope space (see 637 figure included in the supplementary data in the 638 639 Appendix), possibly suggesting a similar formation time (~1.8 Ga) for the DUPAL and FOZO reservoirs (Pb-640 isotope ratios in high 3He/4He lavas are displaced from 641 642 the Geochron and preclude coeval formation of the FOZO and the hypothetical EDR reservoirs). 643

644 4.7. FOZO-A and FOZO-B: implications for mantle 645 dynamics

A clear implication of the isotopic and geographic separation of the two high ³He/⁴He reservoirs is that they had to be isolated from each other for long timescales. The long-term separation of the two FOZO domains requires preservation despite convective stirring, a mechanism that efficiently attenuates mantle heterogeneities (van Keken et al., 2002). However, the hemispheric heterogeneity in 652 the high ³He/⁴He reservoir is apparently not a feature 653 preserved in the convecting upper mantle sampled by 654 mid-ocean ridges (see Sections 34.1). Thus, the shallow 655 mantle may not be an ideal location for the preservation of 656 the FOZO-A and FOZO-B mantle domains. However, the 657 less rapid convective motions of the lower mantle may 658 make it a more suitable home for the FOZOs (Hart et al., 659 1992; Macpherson et al., 1998; van Keken et al., 2002; 660 Class and Goldstein, 2005). 661

One mechanism for the preservation of hemispheric- 662 scale heterogeneity may be the isolation of the high 663 ³He/⁴He domains in a dense boundary layer at the core- 664 mantle boundary (CMB). Seismic tomography suggests 665 that some hotspots, like Samoa and Hawaii, may well 666 originate as upwellings from this region of the mantle 667 (Montelli et al., 2006), and many hotspots with high 668 ³He/⁴He have been associated with velocity anomalies in 669 the deep mantle (Courtillot et al., 2003). The post- 670 perovskite phase proposed to exist at this depth may 671 exhibit a sufficiently large density contrast to isolate it 672 from the overlying convecting mantle (Guignot et al., 673 2007), and a new seismic technique may allow detection 674 of this phase at the base of the mantle (van der Hilst et al., 675 2007). Exploiting this new seismic tool, it may be pos- 676 sible to better resolve the spatial relationships between 677 velocity anomalies at the CMB and high 3He/4He ocean 678 islands at the earth's surface. Thus, a confluence of 679 geochemical and geophysical observations may ulti- 680 mately reveal the mechanism responsible for the long- 681 term preservation of the hemispheric heterogeneity in the 682 deep mantle, a feature that hinges on new observations of 683 high ³He/⁴He in Samoan lavas from Ofu. 684

5. Conclusions

From this study we draw the following conclusions: 686

685

- (1) New high ³He/⁴He ratios (19.5–33.8 times 687 atmospheric) from Ofu island are the highest 688 from Samoa (and the southern hemisphere), and 689 place Samoa in the same category of high ³He/⁴He 690 hotspots as Hawaii, Iceland and the Galapagos. 691
- (2) The new Ofu data reveal that at least two distinct 692 high ³He/⁴He reservoirs—one more isotopical-693 ly enriched than the other—exist in the earth's 694 mantle. 695
- (3) The two high ³He/⁴He reservoirs separate in the 696 earth's northern (FOZO-B, boreal) and southern 697 (FOZO-A, austral) hemispheres. 698
- (4) The trace element budgets and isotopic composi- 699 tions of the new high ³He⁴He samples from Samoa 700

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- 701 are inconsistent with derivation from recycled harzburgite. Additionally, the Samoan high 702 ³He/⁴He mantle does not appear to be contaminated 703 with rapidly cycled sediment from the Tonga sub-704 705 duction zone. Furthermore, the Nd and Pb isotopes of the highest 3He/4He Samoan basalts demonstrate 706 that they were not derived from a chondritic prim-707 itive mantle. 708 (5) If the terrestrial ¹⁴²Nd/¹⁴⁴Nd anomaly relative to
- (5) If the terrestrial ¹⁴²Nd/¹⁴⁴Nd anomaly relative to chondrites is due to the decay of ¹⁴⁶Sm, the observed ¹⁴³Nd/¹⁴⁴Nd ratios in FOZO-A and FOZO-B lavas require that they were re-enriched relative to either a non-chondritic BSE or the EDR.
- (6) The high ${}^{3}\text{He}/{}^{4}\text{He}$ mantle is an ancient reservoir, 714 and the discovery of isotopically distinct northern 715 and southern hemisphere high ³He/⁴He mantle 716 domains suggests that these regions of the mantle 717 718 escaped the convective mixing and stirring that has efficiently attenuated heterogeneities in the upper 719 mantle. This observation provides an important 720 constraint for future dynamic and isotopic models 721 describing the evolution of the earth's mantle. 722

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738 Appendix A. Supplementary data

Supplementary data associated with this article can
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Appendix A. Supplementary data (as it appears, published, on the Elsevier website) Supplementary material for on-line publication only Click here to download Supplementary material for on-line publication only: Supplement.pdf



Supplementary Material Fig. 1. Upper panel: Location map for samples taken from Ofu and Olosega islands that are reported in this study. Samples were taken from the major geologic formations recorded by Stice and McCoy (1968). Lower Panel: Ofu Island relative to the subaerial islands (Savai'i, Upolu, Tutuila and Ta'u) and submarine volcanoes (Muli, Malumalu and Vailulu'u) of the Eastern volcanic province (map from Workman et al. (2004)). Papatua and Uo Mamae are isolated, "off-axis" seamounts. Note Samoa's proximity to the northern termination of the Tonga arc/trench system.



Supplementary Material Fig. 2. The separation of FOZO-A and FOZO-B in Pb-isotope space. Top panels: In two dimensional Pb-isotope space (²⁰⁸Pb/²⁰⁴Pb vs. ²⁰⁶Pb/²⁰⁴Pb and ²⁰⁷Pb/²⁰⁴Pb vs ²⁰⁶Pb/²⁰⁴Pb isotope spaces), the two FOZOs can be separated by a line. However, the Northern Hemisphere Reference Line (NHRL) from Hart (1984) does not appear to divide the two FOZO's in ²⁰⁸Pb/²⁰⁴Pb – ²⁰⁶Pb/²⁰⁴Pb space, but does divide them in ²⁰⁷Pb/²⁰⁴Pb – ²⁰⁶Pb/²⁰⁴Pb space. Open black diamonds are global OIB data. Lower panel: FOZO-A and FOZO-B separate in 3D Pb-isotope space, indicating that Pb-isotopes are consistent with the hemispheric separation of the two high ³He/⁴He reservoirs. Pb-isotope data used to define FOZO-A and FOZO-B are found in Table 2.



Supplementary Material Fig. 3. Modal abundance of cpx compared to reconstructed whole-rock Sr concentrations in abyssal peridoties, a proxy for recycled, depleted, oceanic upper mantle lithosphere. Sr concentrations in abyssal peridotites diminish rapidly with small reductions in cpx modal abundance, a result of melt extraction from DMM in a fractional melting regime. Harzburgites and other cpx-poor lithologies that are produced in the mantle lithosphere at mid-ocean ridges are extremely trace element depleted. These peridotites may not be good candidates for a source that generates the enriched trace-element budgets observed in Ofu lavas. The abyssal peridotite compilation accompanies Workman and Hart (2005), and the estimate for the Sr concentration in DMM concentrations is from the same source. Samples with modal plagioclase have been excluded.
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Chapter 4

High ³He/⁴He hotspot lavas expose the Earth's "missing" titanium, tantalum and niobium (TITAN): The missing link between continental crust and depleted mantle found?

Abstract

A shortage of the elements titanium, tantalum and niobium (TITAN) exists in the Earth's shallow geochemical reservoirs—the depleted MORB (mid-ocean ridge basalt) mantle and continental crust—and the location of these missing elements is unknown. Here we report evidence for a global, TITAN-enriched reservoir sampled by OIBs (ocean island basalts) with high ³He/⁴He ratios, an isotopic signature associated with the deep mantle. Excesses of Ti (and to a lesser degree Nb and Ta) correlate remarkably well with ³He/⁴He in a dataset of global OIBs. The observation of TITAN enrichment in high ³He/⁴He OIB lavas suggests that the mantle domain hosting the Earth's "missing" TITAN is sampled by deep, high ³He/⁴He mantle plumes. The TITAN enrichment in the high ³He/⁴He reservoir has profound implications for the origin of the high ³He/⁴He mantle component, and suggests that, far from being a primitive reservoir, or simply a depleted peridotite reservoir, the high ³He/⁴He mantle sampled by OIBs appears to host a component of recycled, refractory, rutile-bearing oceanic crust that was processed in subduction zones.

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1. Introduction

The standard model for the evolution of the silicate earth maintains that the depleted MORB (mid-ocean ridge basalt) mantle (DMM) is the residue of continental crust (CC) extraction from an early primitive mantle (Jacobsen and Wasserburg, 1979; O'Nions et al., 1979; Allegre et al., 1980; Hofmann, 1988, 1997). If the earth has chondritic abundances of the refractory elements, DMM and CC must be geochemically complementary reservoirs within the earth. However, the TITAN trio of elements are prominently depleted in the continents (Rudnick and Gao, 2003), and their absence is not balanced by a corresponding enrichment in DMM (Workman and Hart, 2005). Thus, another deeper reservoir hosting the missing TITAN elements has been proposed to exist in the earth (McDonough, 1991; Rudnick et al., 2000; Kamber and Collerson, 2000).

Oceanic plates are formed by melting and depletion of the upper mantle at midocean ridges. The resulting oceanic lithosphere, composed of mafic oceanic crust and the uppermost region of the depleted peridotite mantle, is subducted back into the mantle at trenches, thereby contributing to its compositional heterogeneity (Hofmann and White, 1980, 1982; Chase, 1981; Zindler and Hart, 1986; van Keken et al., 2002). During subduction, the mafic portion of the plate is dehydrated and may be partially melted, and the resulting lavas erupted at subduction zone volcanoes are depleted in the TITAN elements. Incompatible elements are largely lost to the overlying mantle during dehydration and melting of the eclogite portion of slabs. By contrast, titanium-rich phases, such as rutile, may preferentially sequester the TITAN elements in the mafic portion of the downgoing slab, balancing the depletion observed in subduction zone lavas (Green and Pearson, 1986; Ryerson and Watson, 1987; Brennan et al., 1994; Foley et al., 2000; Schmidt et al., 2004; Kessel et al., 2005). Refractory, rutile-bearing eclogites have been subducted in large quantities over geologic time, and may form a reservoir in the mantle that hosts the Earth's missing TITAN (McDounough, 1991; Rudnick et al., 2000).

TITAN enrichment in hotspot lavas also has enormous potential as a geochemical tracer for recycled oceanic plates (Rudnick et al., 2000). If TITAN-enriched refractory eclogites are returned to the surface in mantle plumes, their presence would be evident as TITAN enrichment in hotspot lavas. A common mantle dynamics paradigm maintains

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that mantle plumes complete the process of recycling by transporting plate remnants from the core-mantle boundary back to the surface, where they melt and erupt as ocean island basalts (OIBs) (Hofmann and White, 1980, 1982; Chase, 1981). However, the reservoir hosting the Earth's missing TITAN has been difficult to detect in hotspot lavas (McDonough, 1991; Rudnick et al., 2000).

Rare, high ³He/⁴He (>30 Ra, ratio to atmosphere) ratios in lavas erupted at some hotspots, including Hawaii, Iceland, Galapagos and Samoa, are thought to be tracers of buoyantly upwelling mantle plumes that sample an ancient reservoir residing in the mantle (e.g., Kurz et al., 1982; Hart et al., 1992). Variously called FOZO (Focus Zone; Hart et al., 1992; Jackson et al., 2007a), PHEM (Primitive Helium Mantle; Farley et al., 1992) or C (Common; Hanan and Graham, 1996), the precise location (shallow or deep mantle) of the high ³He/⁴He reservoir is a source of intense debate (e.g., Anderson, 1998). Nonetheless, there is a growing consensus that this reservoir hosts a significant component of depleted mantle peridotite (e.g., Hart et al., 1992; Anderson, 1998; Parman et al., 2005; Heber et al., 2007). Recent work suggests that the high ³He/⁴He mantle may also host a component of recycled eclogite (Dixon et al., 2001; Brandon et al., 2007). In this paper, we argue that the high ³He/⁴He mantle sampled by OIBs hosts a component of recycled, refractory eclogite, and that this component balances Earth's budget for the elements titanium, tantalum and niobium (TITAN).

2. New Data and Observations

We report new trace element data by ICP-MS (inductively coupled plasma mass spectrometer) on the highest ³He/⁴He lavas from Hawaii (32.3 Ra; Kurz et al., 1982), Iceland (37.7 Ra; Hilton et al., 1999) and Samoa (33.8 Ra; Jackson et al., 2007a) (see Table 1). In Fig. 1, these new data are presented together with previously published ICP-MS trace element data for the highest ³He/⁴He Galapagos lava (30.2 Ra; Kurz and Geist, 1999; Saal et al., 2007). These high ³He/⁴He lavas exhibit Ti, Ta, and Nb excesses, or positive anomalies, relative to elements of similar compatitibility in peridotite (on a primitive mantle normalized basis, see Fig. 1). While the association of positive TITAN anomalies and high ³He/⁴He (or plume) signatures were previously observed regionally in

Hawaii (Dixon et al., 2001), Iceland (Fitton, 1997) and the Galapagos (Kurz and Geist, 1999; Saal et al., 2007), we suggest that the large, positive TITAN anomalies are a global phenomenon in high ³He/⁴He OIBs. The primitive-mantle normalized trace element patterns (spidergrams) of the highest ³He/⁴He lavas from Hawaii, Iceland, Galapagos and Samoa all share prominent, anomalous enrichment in the TITAN elements compared to elements of similar compatibility in peridotite (Fig. 1). In fact, the Nb/U ratios in the high ³He/⁴He OIB lavas are all higher than the average Nb/U value of 47 previously proposed for OIBs and MORBs (Hofmann et al., 1986).

By contrast, the mantle endmembers (Zindler and Hart, 1986) with low 3 He/ 4 He, including HIMU (high ' μ ', or 238 U/ 204 Pb; Graham et al., 1992; Hanyu and Kaneoka, 1997), EM1 (enriched mantle 1; Honda and Woodhead, 2005), EM2 (enriched mantle 2; Workman et al., 2004; Jackson et al., 2007b) and DMM exhibit spidergrams (Hart and Gaetani, 2006) that lack such pronounced TITAN anomalies (Fig. 1). While HIMU basalts can have positive Nb and Ta anomalies (Weaver et al., 1987; Weaver et al., 1991; Chauvel et al., 1992), they generally have lower anomalies than high 3 He/ 4 He lavas, and can even have negative Nb anomalies (Sun and McDonough, 1989). Importantly, HIMU lavas exhibit flat or negative Ti-anomalies (McDonough, 1991), and thus lack the positive Ti-anomalies observed in high 3 He/ 4 He lavas.

Available data (see Appendix A) also indicate that the TITAN enrichment is enhanced with increasing ³He/⁴He in OIB lavas (Fig. 2): Large, positive Ti (high Ti/Ti*), Nb (elevated Nb/Nb*) and Ta (Ta/Ta*, not shown) anomalies are observed in the highest ³He/⁴He basalts. While all high ³He/⁴He lavas (>30 Ra) have large, positive TITAN anomalies, not all lavas with positive TITAN anomalies have high ³He/⁴He. In the plots of TITAN anomalies vs. ³He/⁴He, the OIB data outline a "wedge-shaped" pattern. For example, Cape Verde lavas have large, positive Nb/Nb* values, but have low ³He/⁴He (Fig. 2). However, high ³He/⁴He lavas with negative TITAN anomalies are absent in the available dataset.

High ³He/⁴He lavas exhibit moderately radiogenic Os-isotopes (¹⁸⁷Os/¹⁸⁸Os>0.135), an observation that is contrary to previous suggestions that this reservoir hosts unradiogenic ¹⁸⁷Os/¹⁸⁸Os ratios (Hauri et al., 1996) similar to DMM

(Standish et al., 2002) (Fig. 2). Radiogenic ¹⁸⁷Os/¹⁸⁸Os in the high ³He/⁴He reservoir is consistent with the positive correlation observed between ³He/⁴He and ¹⁸⁷Os/¹⁸⁸Os in Icelandic hotspot lavas (Brandon et al., 2007). However, radiogenic ¹⁸⁷Os/¹⁸⁸Os is not unique to the high ³He/⁴He mantle: the low ³He/⁴He mantle endmembers EM1 and HIMU also have radiogenic ¹⁸⁷Os/¹⁸⁸Os. Importantly, unradiogenic ¹⁸⁷Os/¹⁸⁸Os is not observed in high ³He/⁴He OIB lavas.

3. The case for a refractory, rutile-bearing eclogite component in the high ³He/⁴He mantle sampled by OIBs

3.1. TITAN enrichment and high ¹⁸⁷Os/¹⁸⁸Os: Evidence for refractory eclogite.

The lack of large TITAN anomalies in DMM (Fig. 1) demonstrates that phases contained in upper mantle peridotites do not preferentially sequester the TITAN elements relative to other incompatible lithophile elements. By contrast, subduction zone lavas are TITAN-depleted, indicating that processes operating in their mantle source can fractionate TITAN from the other lithophile trace elements. It was suggested that eclogite melting in subduction zones may generate rutile-bearing residues that are residually-enriched in TITAN elements (e.g., McDonough, 1991). Experimental studies indicate that the TITAN elements are strongly partitioned into rutile during eclogite melting (Green and Pearson, 1986; Ryerson and Watson, 1987; Ayers, 1998; Stalder et al., 1998; Foley et al., 2000; Schmidt et al., 2004; Kessel et al., 2005), thereby generating positive TITAN anomalies in refractory, rutile-bearing slab residues that balance the depletion observed in subduction zone lavas.

Like TITAN-enrichment, radiogenic ¹⁸⁷Os/¹⁸⁸Os is not a geochemical signature typically associated with a peridotite reservoir. While peridotites tend to have low Re/Os and ¹⁸⁷Os/¹⁸⁸Os, mafic igneous rocks generally exhibit elevated Re/Os and ¹⁸⁷Os/¹⁸⁸Os ratios, (Walker et al., 1989; Reisberg et al., 1991; Hauri and Hart, 1993; Reisberg et al., 1993; Snow and Reisberg, 1995; Becker, 2000). Oceanic crust enters subduction zones with initially high Re/Os, and owing to moderate compatibility of Re in garnet, oceanic crust may retain high Re/Os (and with time, radiogenic ¹⁸⁷Os/¹⁸⁸Os) ratios during

subduction zone processing in the garnet stability field (Righter and Hauri, 1998). While it has been suggested that Os and Re are extracted from the slab in the subduction zone (e.g., Brandon et al., 1996; McInnes et al., 1999; Becker et al., 2000), we note that the Re/Os ratios of altered oceanic crust (187 Re/ 188 Os averages of 349 and 353 in composites of two separate drill cores; Peuker-Ehrenbrink et al., 2003) and oceanic crust gabbros (187 Re/ 188 Os average of 472, Hart et al., [1999]) are similar to the ratios found in metabasalts metamorphosed in paleosubduction zones (median 187 Re/ 188 Os = 326, including eclogites, blueschists and mafic granulites; Becker, 2000). Thus, there is ample evidence supporting the contention that high Re/Os ratios can be preserved in the slab during subduction zone metamorphism. Like TITAN enrichment, the observation of moderately radiogenic 187 Os/ 188 Os in high 3 He/ 4 He lavas is consistent with a refractory eclogite component in their mantle sources, and not consistent with the high 3 He/ 4 He reservoir being primitive mantle.

3.2. Depletion in the ⁴He-producing elements (U and Th) and TITAN enrichment.

While recycled oceanic crust has been suggested to be ubiquitous in the mantle sources beneath hotspots (Sobolev et al., 2007), the close association of a refractory, TITAN-enriched mafic component with the high ³He/⁴He mantle may appear contradictory since eclogites are quantitatively degassed in subduction zones (Staudacher and Allegre, 1988; Moreira and Kurz, 2001; Moreira et al., 2003). Far from hosting the high ³He/⁴He signature in the mantle source of high ³He/⁴He OIB lavas, the eclogite will instead contribute ⁴He (via alpha decay of U and Th) and generate low time-integrated ³He/⁴He ratios. However, the long-term ⁴He production of an eclogite can be greatly reduced by melt (or fluid) extraction of highly incompatible elements like U and Th from the slab during subduction processing. By contrast, the TITAN elements will be conserved if melt (or fluid) extraction occurs in the presence of rutile (e.g., McDonough, 1991). In this scenario, a positive TITAN anomaly indicates U and Th (including other incompatible trace elements, like La, Tb and Sm) depletion in a slab. The positive TITAN anomaly is formed by relative enrichment of TITAN due to conservation of Ti,

Ta and Nb and concomitant loss of U and Th. Therefore, the refractory eclogite model of McDonough (1991) is potentially consistent with the preservation of a high 3 He/ 4 He in the mantle. A U- and Th-depleted, TITAN-enriched eclogite will not produce significant post-subduction radiogenic 4 He ingrowth (see below), and may explain why TITAN enrichment is associated with high 3 He/ 4 He signatures.

By contrast, subducted oceanic crust that conserves much of its original U and Th budget through the subduction zone will not possess positive TITAN anomalies and will produce more ⁴He over time than a U and Th-depleted slab. Consequently, the abundant ⁴He produced by an undepleted slab would infect the surrounding mantle with radiogenic ⁴He due to rapid diffusion of helium in the mantle (Hart, 1984; Trull and Kurz, 1993; Hart et al., 2007), and is not compatible with the preservation of a high ³He/⁴He signature in the mantle. This hypothesis is consistent with the observation that high ³He/⁴He lavas never have negative (or flat) TITAN anomalies.

3.3. Refractory eclogite and high ${}^{3}\text{He}/{}^{4}\text{He}$ peridotite: the raw materials for the high ${}^{3}\text{He}/{}^{4}\text{He}$ OIB mantle.

While refractory, rutile-bearing eclogite possesses positive TITAN anomalies, it does not have intrinsically high ³He/⁴He ratios: high ³He/⁴He signatures in the TITANenriched, high ³He/⁴He OIB source must be derived from another lithology. Ancient mantle peridotites can potentially preserve elevated ³He/⁴He ratios over time (Hart et al., 1992; Anderson, 1998; Parman et al., 2005; Heber et al., 2007; Jackson et al., 2007a) if they were isolated from the convecting mantle early in Earth's history. However, mantle peridotites do not generally exhibit positive TITAN anomalies (McDonough, 1991). Alone, neither eclogite nor peridotite can contribute both TITAN-enrichment and high ³He/⁴He to the mantle source sampled by high ³He/⁴He OIB lavas. Both refractory eclogite and the high ³He/⁴He peridotite are required to generate such a mantle source. Thus, an important question is how the TITAN-enriched eclogites came to be associated (i.e., mixed) with high ³He/⁴He peridotites in the high ³He/⁴He mantle sampled by OIBs. 3.4. Peridotite and eclogite portions of ancient subducted slabs: A high ³He/⁴He, TITAN-enriched "package".

One possible explanation for the generation of a hybrid lithology in the high ³He/⁴He OIB mantle is that the crust and mantle components of ancient, subducted oceanic lithosphere contain the raw materials required for the generation of a mantle source sampled by high ³He/⁴He basalts. TITAN-enriched refractory eclogite exists at the top of the slab, and ancient high ³He/⁴He asthenospheric peridotite (with the same composition as contemporary DMM) is coupled to the underside of the downgoing plate. After processing in subduction zones, the crustal portion of the subducted plate hosts TITAN enrichment, and the asthenospheric DMM coupled to the bottom of the downgoing plate remains unscathed by subduction zone processes. The two components, TITAN-enriched eclogite and high ³He/⁴He asthenospheric peridotite, should be intimately associated as a "package" in space and time within a subducted plate, a geometry that is conducive to later mixing in the mantle.

We present a simple model for the generation of the high ³He/⁴He mantle sampled by OIBs, whereby the eclogitic and peridotitic components of the oceanic plate are subducted and isolated from the convecting mantle at 3 Ga and mixed during storage in the lower mantle (Fig 3). We assume that the upper mantle began with a primitive composition (McDonough and Sun, 1995) at 4.4 Ga, and evolved by continuous depletion to the present-day DMM composition of Workman and Hart (2005). Oceanic plates were continuously injected into the mantle over this time period, and the eclogitic and peridotitic portions of these subducted plates were thoroughly mixed in the lower mantle and were later sampled by upwelling mantle plumes and erupted at hotspots.

In order to develop a quantitative geochemical model for the ³He/⁴He, ¹⁸⁷Os/¹⁸⁸Os, and TITAN of the two lithologies (peridotite and eclogite) in a subducted plate, we make the following assumptions about their compositions over time:

i. 3 He/ 4 He of ancient asthenospheric DMM peridotite and refractory eclogite: The 3 He/ 4 He of DMM is thought to have decreased significantly over time (Fig. 3). However, portions of asthenospheric DMM peridotite that are isolated with the downgoing plate

(i.e., removed from the upper mantle) are not subject to further depletion. In the model, the subducted asthenospheric DMM peridotite "locked in" the high ${}^{3}\text{He}/{}^{4}\text{He}$ and ${}^{3}\text{He}/{}^{238}\text{U}$ ratios of ancient, less-depleted DMM, and evolved by closed-system decay of U and Th. By contrast, if the subducted portion of ancient asthenospheric DMM had remained as part of the convecting asthenosphere, it would have evolved by continuous melt extraction to become modern DMM with low ${}^{3}\text{He}/{}^{4}\text{He}$ and ${}^{3}\text{He}/{}^{238}\text{U}$ (see Fig. 3 and Appendix B for details of ${}^{3}\text{He}/{}^{4}\text{He}$ of the isolated, ancient DMM rapidly diverged from (and preserved higher ${}^{3}\text{He}/{}^{4}\text{He}$ than) its upper mantle counterpart, which continued to be depleted (in He relative to U and Th, see Appendix B) by continental and oceanic crust extraction (Fig. 3).

By comparison, we model the ³He/⁴He evolution of the subducted eclogite using the trace element composition of a hypothetical refractory eclogite calculated by McDonough (1991) (see Fig. 4 for a spidergram of the refractory eclogite). The eclogite is assumed to start with a ³He/⁴He ratio equal to DMM at its time of isolation, and following over 99% degassing in the subduction zone, the U and Th of the eclogite generates ⁴He by decay, thus rapidly diminishing the ³He/⁴He of the eclogite over time (Fig. 3, Table 2). Degassing of the eclogite portion of the subducted slab is simulated by increasing the ²³⁸U/³He of the eclogite by a factor of 1,000 relative to contemporary DMM.

ii. ¹⁸⁷Os/¹⁸⁸Os of ancient asthenospheric DMM peridotite and refractory eclogite: The ¹⁸⁷Os/¹⁸⁸Os of DMM and Primitive Mantle are not very different (0.12-0.13; Standish et al., 2002; Meisel et al., 2001), and thus the isolated, ancient asthenospheric peridotite portions of the subducted plates are assumed to have an intermediate present-day composition (0.125). DMM is also assumed to have had an Os concentration of 3000 ppt over geologic time. In contrast to the asthenospheric DMM peridotite, the eclogitic portion of the subducted plate likely evolved extremely radiogenic ¹⁸⁷Os/¹⁸⁸Os over time (Fig. 3). The ¹⁸⁷Os/¹⁸⁸Os evolution of the refractory eclogite is modeled using the median ¹⁸⁷Re/¹⁸⁸Os (325) and Os (6 ppt) from eclogites metamorphosed in paleosubduction zones

reported by Becker (2000): earlier subduction injection of eclogites with this composition yielded higher ¹⁸⁷Os/¹⁸⁸Os in the present day. The high calculated present-day ¹⁸⁷Os/¹⁸⁸Os ratios in the recycled eclogites are similar to the most radiogenic eclogites presented by Becker (2000).

iii. TITAN anomalies of ancient asthenospheric DMM peridotite and refractory

eclogite: The trace element content of the present-day asthenospheric DMM peridotite is assumed to be the same as the DMM compositions calculated by Workman and Hart (2005). However, DMM has likely become increasingly depleted throughout geologic time. We assume that DMM began with a primitive mantle composition (McDonough and Sun, 1995) and evolved by continuous depletion until the present day. Using the continuous transport equations in Appendix B, the trace element budget of DMM is calculated at various times, and "snapshots" of DMM compositions through time are plotted as spidergrams in Fig. 4 (see Table 2 for compositions). By contrast, the subducted eclogites are assumed to have the same present-day trace element composition as the hypothetical refractory eclogite from McDonough (1991), regardless of the isolation time (see Table 2). This hypothetical eclogite has a trace element composition similar to the eclogites with the largest positive TITAN anomalies presented in Becker et al. (1999) (see Table 2).

3.5. Mixing asthenospheric DMM peridotite with refractory eclogite.

The high ³He/⁴He, moderately radiogenic ¹⁸⁷Os/¹⁸⁸Os, and positive TITAN anomalies observed in the highest ³He/⁴He OIB lavas can be generated by mixing refractory eclogite and the asthenospheric DMM peridotite that were subducted together (in the same plate) at 3 Ga. When the proportion of refractory eclogite in the mixture is between 20 and 25%, the model generates ³He/⁴He, ¹⁸⁷Os/¹⁸⁸Os and positive TITAN anomalies that are similar to the highest ³He/⁴He OIB lavas (Fig. 2). If the proportion of refractory eclogite is increased, the present-day ³He/⁴He in the resulting mixture is diminished and the positive TITAN anomalies and ¹⁸⁷Os/¹⁸⁸Os ratios are both increased to values above those observed in high ³He/⁴He OIB lavas.

Fig. 4 shows that the addition of 20% eclogite to a 3 Ga DMM composition generates a hybrid eclogite-peridotite spidergram that is similar in shape to the spidergram of the highest ³He/⁴He lava from Iceland. (Sr is a poor fit, however, due to the positive Sr anomalies in the Icelandic lava, a possible result of interaction with shallow lithospheric gabbros [Gurenko and Sobolev, 2006]). This hybrid spidergram may thus serve as a plausible melt source for the high ${}^{3}\text{He}/{}^{4}\text{He}$ lava. It is important to note that the TITAN elements are "bracketed" by elements of similar compatibility (as long as rutile is absent) in the spidergram, and these bracketing elements are used to calculate the TITAN anomalies. Thus, the magnitude of TITAN anomalies are little affected by partial melting beneath a hotspot, and the positive TITAN anomalies in the mantle source of high ³He/⁴He lavas are reflected in the erupted hotspot lavas. Clearly, this assumes that the rutile present in the downgoing slab is no longer present in the high ${}^{3}\text{He}/{}^{4}\text{He}$ mantle that melts beneath hotspots. One way to destabilize rutile is to completely mix the (smaller proportion of) eclogite and the (larger proportion of) peridotite in the ancient recycled slab. Alternatively, if eclogite is still present in the source of the high ${}^{3}\text{He}/{}^{4}\text{He}$ OIB mantle source, it must be melted to a sufficiently high degree to eliminate rutile as a phase in the residue of melting (Gaetani et al., 2007).

This model for the generation of high ³He/⁴He, TITAN-enriched mantle can also generate a low ³He/⁴He that has positive TITAN anomalies. Subduction is a continuous process that has operated for much of geologic time, and the ³He/⁴He ratio of the upper mantle has likely decreased significantly (Fig. 3). Thus, the peridotite portion of more recently subducted oceanic plates will trap and preserve a lower ³He/⁴He upper mantle signature than ancient subducted plates, and the refractory eclogite portion of recently subducted plates will also host positive TITAN anomalies. For example, the lower ³He/⁴He, TITAN-enriched mantle source sampled by Cape Verde lavas can be generated by mixing refractory eclogite (10-30% by mass) and asthenospheric DMM components of a plate subducted between 1-2 Ga (Fig. 2).

The "wedge-shaped" outline of the OIB data in Fig. 2 highlights a striking absence of high ³He/⁴He lavas with negative (or even small positive) TITAN anomalies. Why do all high ³He/⁴He lavas exhibit TITAN-enrichment, or rather, why have high

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³He/⁴He-TITAN depleted (or only mildly TITAN enriched) lavas not been found? Continental crust and arc lavas (and associated sediments) compose the only known reservoir to exhibit TITAN depletion. Due to its extremely high U and Th contents, admixture of CC with high ³He/⁴He peridotite may not be conducive to the preservation of high ³He/⁴He (Jackson et al., 2007b), and could explain why TITAN-depleted lavas always exhibit low ³He/⁴He (Fig. 2). OIB lavas lacking TITAN anomalies may host an eclogite component that was not U and Th-depleted (and thus did not acquire positive TITAN anomalies) in a subduction zone, and may also produce significant ⁴He.

On the other hand, the absence of high ³He/⁴He lavas with flat (or even slightly positive) TITAN anomalies may be explained by the intimate spatial and temporal association between the TITAN-enriched eclogites and high ³He/⁴He peridotites suggested by this model: The peridotite and refractory eclogite components—the raw materials for the formation of the high ³He/⁴He, TITAN-enriched mantle—are always together in subducting plates, and it may not be possible to melt pure high ³He/⁴He peridotite slabs are stretched, thinned and folded in the dynamic mantle (Allegre and Turcotte, 1986), such that the diminished thickness of the (eclogite and peridotite) slab is less than the width of melting zones beneath hotspots. Such a process might explain why an eclogite signature (positive TITAN anomalies and radiogenic ¹⁸⁷Os/¹⁸⁸Os) is invariably present in high ³He/⁴He lavas.

3.6. Two alternative models for a hybrid high ³He/⁴He mantle: "Eclogite Injection" and diffusion of "Ghost" primordial helium.

An alternative process for generating the hybrid eclogite-peridotite high ³He/⁴He mantle assumes that subducted slab peridotite contributes little to the helium budget of the high ³He/⁴He mantle sampled by OIBs. In this model, the refractory slab eclogite penetrates into the lower mantle and mixes with a hypothetical lower mantle high ³He/⁴He peridotite reservoir (i.e., not associated with the peridotite portion of slabs). The hybrid mixture then rises in a plume where it is melted beneath a hotspot. By comparison to the slab peridotite-eclogite "package" model describe in section 3.4 and 3.5 above, this

alternative "eclogite injection" model does not guarantee an intimate spatial and temporal association of the high ³He/⁴He peridotite and TITAN-enriched eclogite components. For example, deep mantle high ³He/⁴He peridotite could upwell in a plume without first being inoculated with refractory eclogite, in which case the erupted high ³He/⁴He lavas would lack positive TITAN anomalies. Such lavas have not been observed. While this model is not explored further here, one possible solution could be that, due to mixing by stretching and thinning of slabs in a convecting mantle over geologic time (Allegre and Turcotte, 1986), recycled eclogite has become pervasive in the mantle and is distributed at lengthscales smaller than the melting zones beneath hotspot volcanoes. In this way, volcanoes fed by upwelling high ³He/⁴He mantle plumes will inevitably sample subducted eclogite.

Instead of mechanically mixing the high ³He/⁴He peridotite and TITAN-enriched eclogite, as suggested in the "slab package" and "eclogite injection" models above, it may be possible to diffusively mix helium from a high ${}^{3}\text{He}/{}^{4}\text{He}$ peridotite into a degassed, U and Th-poor pyroxenite (i.e., refractory eclogite) (Albarede and Kaneoka, 2007). Due to the higher diffusivity of helium compared to non-volatile major and trace elements, helium isotopes may become decoupled from other lithophile isotope tracers (Hart et al., 2007), and primordial helium may become associated with recycled materials like refractory eclogites (Albarede and Kaneoka, 2007). Albarede and Kaneoka (2007) propose that helium from deep (high ${}^{3}\text{He}/{}^{4}\text{He}$) mantle peridotites can diffuse into embedded, tightly folded layers of stretched and thinned refractory eclogite. They suggest that changing the duration of the diffusion process, as well as the U and Th contents of the refractory eclogite layers, can generate mantle sources for both high and low ³He/⁴He hotspots. U and Th-poor refractory eclogite that was processed in subduction zones will have positive TITAN anomalies (McDonough, 1991), and because such eclogites will produce little ⁴He over time, they are perfect "containers" for preserving diffusively acquired high ${}^{3}\text{He}/{}^{4}\text{He}$ signatures. If these eclogites become sufficiently thinned (to <1-2 km thickness, by mantle mixing), they could acquire high ${}^{3}\text{He}/{}^{4}\text{He}$ signatures by diffusion from the ambient deep mantle peridotite (Hart et al., 2007). Thus, the "ghost" helium model of Albarede and Kaneoka (2007) may offer a

resolution to the paradoxical association of high ³He/⁴He signatures in lavas with strong eclogite signatures.

Nonetheless, the "ghost helium" model suffers from the same spatial and temporal issues as the "eclogite injection" model: There is no obvious mechanism preventing ambient lower mantle (eclogite-free) peridotite from upwelling and melting beneath a hotspot, thus generating a high ³He/⁴He lava that lacks positive TITAN anomalies. However, one possible resolution is that plume-entrained, high ${}^{3}\text{He}/{}^{4}\text{He}$ lower mantle peridotites are too refractory to contribute significantly to melting beneath OIBs (i.e., Albarede and Kaneoka [2007] invoke dunites and harzburgites), so that pure peridotites are never melted. Alternatively, if eclogite layers are pervasive in the mantle at lengthscales smaller than those sampled by melting zones (<100 km), it may be inescapable that eclogites always contribute to mantle melts. An additional problem with the "ghost helium" model is that rutile may still be stable in a pure eclogite mantle component that is upwelling beneath a hotspot; the residual rutile will hold back TITAN in the source, and as a result, positive TITAN anomalies will not be observed in the erupted lavas. However, if the eclogite is melted to sufficiently high degrees beneath a hotspot, the rutile will be completely consumed and not present in the residue (Gaetani et al., 2007).

3.7. High ³He/⁴He lavas without positive TITAN anomalies?

If a high ³He/⁴He peridotite could be melted in pure (no eclogite) form, there would be no positive TITAN-anomalies in the erupted lavas. There would also be no contribution of ⁴He ingrowth from the eclogite, and an even higher ³He/⁴He might be expected in the melts of such a mantle source. Such lavas have not been identified. However, with the highest magmatic ³He/⁴He values on record, Baffin Island lavas (Stuart et al., 2003) may provide an important test case for this hypothesis.

3.8. TITAN anomalies due to partitioning between lower mantle phases?

It is difficult to rule out the possibility of TITAN fractionation in lower mantle materials. Experimental studies of high pressure partitioning and mineralogy are in the early phases. However,Ca-perovskite in peridotitic and basaltic systems shows negative Ti and Nb partitioning patterns compared to Th, U and the rare earth elements (REEs) (Hirose et al., 2004). This means that a Ca-perovskite bearing solid assemblage would have negative anomalies, but melt equilibrated with Ca-perovskite could have positive anomalies. It if it possible to generate Ca-perovskite melts at the appropriate pressures and temperatures (for example, D''), and extract them from the lower mantle (the inferred home of the high ³He/⁴He domain), then Ca-perovskite melting in the lower mantle may offer a potential explanation for the origin of TITAN-enriched, high ³He/⁴He lavas.

4. The high ³He/⁴He, TITAN-enriched mantle: A reservoir for the "missing" TITAN elements in the earth?

In addition to offering insights into the composition and generation of the high ³He/⁴He mantle reservoir, TITAN-enrichment in high ³He/⁴He OIBs may provide information about the location and composition of the reservoir hosting the earth's missing TITAN elements. If the refractory lithophile trace elements have chondritic abundances in BSE (bulk silicate earth), mass balance constraints require that the TITAN elements missing in the shallow earth reservoirs—DMM and CC—must exist in the deeper earth (McDonough, 1991; Rudnick et al., 2000), perhaps in a deep reservoir composed of subducted oceanic plates (Christensen and Hofmann, 1994). In this section, we calculate the trace element budget of a hypothetical missing TITAN-rich reservoir that, when added together with DMM and CC, generates a BSE (McDonough and Sun, 1995) spidergram. We show that, like high ³He/⁴He OIB lavas, the spidergram of the TITAN-rich reservoir has positive TITAN anomalies.

In order to estimate the trace element composition of this deep, TITAN-enriched reservoir, we assume that the composition of BSE can be approximated with just three reservoirs: CC, DMM and subducted plates (here called PLATE, which is composed of oceanic crust, mantle lithosphere, and everything else--like sediment--subducted along with downgoing plates). The trace element budget of the PLATE reservoir is calculated using the following mass balance relationship:

$$[x]_{CC} \times M_{CC} + [x]_{DMM} \times M_{DMM} + [x]_{PLATE} \times M_{PLATE} = [x]_{BSE} \times M_{BSE} \quad (eq. 1)$$

where [x] represents the concentration of the element x in the four reservoirs, and M is the mass of the reservoirs. Assuming the mass of CC is fixed at 0.6% of BSE, and given that the mass proportions of the DMM and PLATE reservoirs are unknown, the following relationships are used to calculate possible trace element budgets for the PLATE reservoir:

$$M_{CC} = 0.006 \times M_{BSE}$$
 (eq. 2)
 $M_{PLATE} = \beta$ (eq. 3)
 $M_{DMM} = 0.994 - \beta$ (eq. 4)

where the mass of the three reservoirs, PLATE, CC, and DMM sum to the mass of the total silicate mantle. Employing the trace element budgets previously derived for CC (Rudnick and Gao, 2003), DMM (Workman and Hart, 2005) and BSE (McDonough and Sun, 1995), the trace element budget of the subducted PLATE reservoir is calculated for different values of M_{PLATE} (β). The results of the mass balance model are plotted in Fig. 5 (assuming that $\beta \ge 0.4$). The calculated PLATE spidergram changes as a function of its mass proportion (β) of BSE. For large values of β , the PLATE reservoir is more trace element depleted, and for smaller values of β , the PLATE reservoir is increasingly trace element enriched. The most important observation is that, regardless of the value of β , the PLATE reservoir exhibits positive TITAN anomalies.

It is possible to place constraints on the minimum size of the subducted PLATE reservoir and show that its mass proportion of BSE is unlikely to be small. Following Rudnick et al. (2000), we assume that the present mass of the oceanic crust $(5.3 \times 10^{21} \text{ kg})$ has been subducted every 100 Ma for the past 2.5 Ga, and the total mass of the reservoir comprised of subducted oceanic crust is 1.3×10^{23} kg. Assuming that the portion of the subducted oceanic plate that is oceanic mantle lithosphere is 10 times thicker and 10% denser than the oceanic crustal lithosphere, then the total mass of oceanic plates

subducted over the past 2.5 Ga is $\sim 1.4 \times 10^{24}$ kg ($\sim 8\%$ of which is oceanic crust), which is $\sim 40\%$ of the mass of BSE. We consider this a minimum estimate for the size of the subducted PLATE reservoir, as subduction probably operated before 2.5 Ga, and the Archaen mantle was hotter, thereby leading to higher rates of plate formation and subduction (Rudnick et al., 2000). Thus, the PLATE reservoir composes $\geq 40\%$ of BSE, and all possible spidergrams for the plate reservoir are more depleted than the PLATE spidergram shown in Fig. 5.

Therefore, mass balance considerations (equations 1-4) and estimates of plate recycling budgets suggest that a large mantle reservoir host for the missing TITAN elements. The spidergram of the calculated PLATE reservoir exhibits hints of the positive TITAN-anomalies observed in high ³He/⁴He OIBs. Thus, it is not implausible that high ³He/⁴He lavas are sampling the "missing" TITAN hosted in the PLATE reservoir.

While there are many similarities, there is some disagreement between the shape of the spidergrams of the calculated PLATE reservoir and the high ³He/⁴He OIB lavas. Most significantly, the positive TITAN anomalies in the PLATE reservoir are not as large as those observed in the high ³He/⁴He lavas from Hawaii, Iceland, Galapagos and Samoa. One solution to this discrepancy is the following: The PLATE reservoir is likely to be heterogeneous, as it includes everything in BSE that is not DMM or CC. These other components in the PLATE reservoir will dilute the large, positive TITAN-anomalies contributed from recycled, rutile-bearing eclogites. While the PLATE reservoir is not purely a high ³He/⁴He reservoir, its spidergram balances the Earth's budget of TITAN, and recycled, refractory, rutile-bearing eclogite likely contributes the positive TITAN anomalies to the PLATE reservoir (McDonough, 1991). It is this TITAN-enriched domain of the PLATE reservoir that may be associated with the source for high ³He/⁴He lavas.

In section 3.5 it was shown that, if mixed with high ${}^{3}\text{He}/{}^{4}\text{He}$ peridotite, approximately 20 to 25% refractory, rutile-bearing eclogite can generate the ${}^{3}\text{He}/{}^{4}\text{He}$, ${}^{187}\text{Os}/{}^{188}\text{Os}$ and TITAN anomalies in high ${}^{3}\text{He}/{}^{4}\text{He}$ OIBs. However, this percentage of eclogite greatly exceeds the percentage of eclogite in the PLATE reservoir (~8%).

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Perhaps this inconsistency can be explained by the fact that eclogites have lower solidus temperatures than pure, unadulterated peridotites. Thus, when melting the high ³He/⁴He portion of the PLATE reservoir beneath a hotspot, the eclogite's (be it pure eclogite or eclogite completely mixed into a peridotite) contribution to the melt will exceed its mass proportion in the mantle source.

5. Implications for ¹⁴²Nd/¹⁴⁴Nd measurements on terrestrial mantle rocks

A recent discovery demonstrated that the ¹⁴²Nd/¹⁴⁴Nd ratios in all measured terrestrial mantle rocks are 20 ppm higher than chondrite (Boyet and Carlson, 2005, 2006), indicating that these rocks were derived from a reservoir that had superchronditic Sm/Nd during the lifetime of ¹⁴⁶Sm (the first few hundred million years following accretion). If the BSE has chondritic abundances of the refractory elements, then the superchondritic ¹⁴²Nd/¹⁴⁴Nd ratios observed in the accessible terrestrial rocks suggest that they sample the depleted residue (early depleted reservoir, or EDR) of an early differentiation event. In this model, a complementary hidden "early enriched reservoir" (EER) with subchondritic ¹⁴²Nd/¹⁴⁴Nd must exist in the deep earth (Boyet and Carlson, 2005, 2006; Andreasen and Sharma, 2006; Carlson et al., 2007). Alternatively, if BSE accreted from non-chondritic materials, an early differentiation of the silicate earth is not required (Caro et al., 2007).

If the earth accreted from non-chondritic materials, the superchondritic $^{142}Nd/^{144}Nd$ ratios in the terrestrial mantle could be a result of superchondritic $^{147}Sm/^{144}Nd$ (≥ 0.209) ratios in BSE (Boyet and Carlson, 2006). However, a non-chondritic earth does not necessarily obviate the need for a deep, TITAN-enriched SLAB reservoir, because we see evidence for its presence in high $^{3}He/^{4}He$ OIBs.

Alternatively, if BSE does have chondritic abundances of the refractory elements, neither of the two early formed reservoirs—EDR and EER—have the appropriate isotopic and trace element characteristics to be the mantle source of high ³He/⁴He lavas. The EDR was suggested to be the high ³He/⁴He reservoir variously called FOZO, PHEM

or C (Boyet and Carlson, 2006). While the EDR does have higher ¹⁴³Nd/¹⁴⁴Nd than chondrite, consistent with globally high (superchondritic) ¹⁴³Nd/¹⁴⁴Nd ratios in high ³He/⁴He lavas (Jackson et al., 2007a), the EDR is calculated to have negative TITAN anomalies (Boyet and Carlson, 2005) and is not consistent with being the mantle source of TITAN-enriched, high ³He/⁴He OIB lavas. By contrast, the EER does have positive TITAN anomalies like those observed in high ³He/⁴He OIB lavas. An early crust isolated at the bottom of the mantle has been suggested to host high ³He/⁴He ratios (Tolstikhin and Hofmann, 2005). However, the EER cannot be the mantle source of the high ³He/⁴He as it has Sm/Nd ratios lower (more enriched) than chondrite, and will generate lower ¹⁴³Nd/¹⁴⁴Nd ratios than observed in high ³He/⁴He OIBs. Thus, neither of the two initial reservoirs suggested by Boyet and Carlson (2005) describe both the isotopic and trace element characteristic of the high ³He/⁴He OIB mantle. The depleted peridotite and refractory, rutile-bearing eclogite in ancient plates that were subducted and stored in the mantle are ideally suited to be a mantle source for high ³He/⁴He OIB lavas.

6. The fate of slabs and the high ³He/⁴He reservoir

The presence of recycled eclogites in the mantle source of high ³He/⁴He lavas has important implications for the helium isotope evolution of the mantle, since subduction zone processing likely plays an important role in determining the composition of recycled eclogites. Specifically, the uniqueness of the thermal regimes of different subduction zones may affect the composition of eclogites in dramatically different ways. Slab melting and dehydration in ancient, hot subduction zones may residually enrich the slab in TITAN elements while depleting it in U and Th, a process that is conducive to the formation and preservation of a high ³He/⁴He, TITAN-enriched mantle reservoir. Alternatively, cooler subduction zones may not generate TITAN anomalies (or deplete the ⁴He -producing elements, U and Th) in the slab, a scenario that may produce mantle reservoirs with small positive TITAN anomalies and low ³He/⁴He. Thus, the fate of the ³He/⁴He evolution of the various mantle reservoirs may hinge on the processes operating in subduction zones.

Appendix A: ³He/⁴He, ¹⁸⁷Os/¹⁸⁸Os and trace element data compilation.

³He/⁴He, ¹⁸⁷Os/¹⁸⁸Os and trace element data for representative OIB samples (plotted in fig. 2) are from the GEOROC database (http://georoc.mpch-mainz.gwdg.de/), from the helium database of Abedini et al. (2006), and from the literature. Some of the ³He/⁴He and ¹⁸⁷Os/¹⁸⁸Os data from Iceland are unpublished, as are ¹⁸⁷Os/¹⁸⁸Os data for Samoan samples with ${}^{3}\text{He}/{}^{4}\text{He} > 20$ Ra (the protocol's for measuring ${}^{187}\text{Os}/{}^{188}\text{Os}$ are the same as Skovgaard et al. [2001] for the unpublished Icelandic data and Workman et al. [2004] for the unpublished Samoan data). The ${}^{3}\text{He}/{}^{4}\text{He}$ data from OIBs (plotted in Fig. 2) were obtained by both crushing and fusion of olivine, clinopyroxene and glass, and are not filtered based on helium concentrations. However, samples suggested to have suffered shallow contamination by crust (e.g., several samples in Macpherson et al., 2005) were not included. Additionally, very evolved rocks (MgO<5.3 wt.%) were excluded, so as to preclude the effects of extensive fractional crystallization on the various trace element ratios (e.g., Ti/Ti* and Nb/Nb*). Using Ba/Rb as a filter for alteration, samples with high ratios (Ba/Rb > 25) were not considered. Only trace element (Th, La, Sm, Tb) data measured by ICP-MS and neutron activation are included, thereby eliminating samples with low-precision trace element measurements.

Appendix B: Helium isotope model.

The precise timing of the genesis of the high 3 He/ 4 He reservoirs cannot be calculated using helium isotopes because the degassing history of the DMM reservoir, the initial 4 He/ 3 He and 3 He abundance of a (hypothetical) undegassed mantle (and DMM), and the present-day 3 He abundance in DMM are not well known. Thus, the age of the reservoir (3 Ga), as sampled by the high 3 He/ 4 He lava from Iceland (sample SEL 97, see Table 1), is poorly constrained. For example, in the model presented in Section 3.5 above, a present-day undegassed mantle 238 U/ 3 He ratio of 70 was assumed (Table 2). However, if the initial 238 U/ 3 He ratio of the mantle is increased to a value of ~250, then our model can generate the TITAN anomalies, 187 Os/ 188 Os and 3 He/ 4 He of the mantle sampled by the Icelandic lava by subducting and isolating a plate at 4 Ga (in this case, the amount of eclogite required to generate the high 3 He/ 4 He, 187 Os/ 188 Os and positive

TITAN anomalies in this older, high ${}^{3}\text{He}/{}^{4}\text{He}$ mantle reservoir increases slightly). In order to generate a TITAN-enriched, high ${}^{3}\text{He}/{}^{4}\text{He}$ mantle at 2 Ga, unrealistically low ${}^{238}\text{U}/{}^{3}\text{He}$ ratios (${}^{238}\text{U}/{}^{3}\text{He} < 10$) are required.

Our preferred model for the time evolution of ³He/⁴He in DMM and primitive mantle are as follows. If an initial ²³⁸U/³He and ³He/⁴He of the undegassed mantle are assumed (see Table 2 for assumed values), and if DMM was formed by continuous depletion by extraction of oceanic and continental crust from a chondritic primitive mantle over Earth's history, the known ³He/⁴He of present day DMM (8 Ra) can be used to calculate the present-day ²³⁸U/³He of DMM. Thus, assuming that the continuous transport equations (Allegre, 1969; Hart and Brooks, 1970; Workman and Hart, 2005) accurately model the continuously depleting upper mantle (DMM), the ²³⁸U/³He and ³He/⁴He of DMM can then be calculated at any time from 4.4 Ga to the present day (see Table 2 for list of these values for DMM at different times in earth's history). The helium isotope evolution of the continuously depleting DMM reservoir (shown in Fig. 3) is modeled using the following continuous transport equations:

$${}^{4}\text{He}/{}^{3}\text{He}_{t} = {}^{4}\text{He}/{}^{3}\text{He}_{T} + \\ 8\lambda_{238}/(\lambda_{238}+k_{238})({}^{238}\text{U}/{}^{3}\text{He})_{T}(1-\exp(-1(T-t)(\lambda_{238}+k_{238}))) + \\ 7\lambda_{235}/(\lambda_{235}+k_{235})({}^{235}\text{U}/{}^{3}\text{He})_{T}(1-\exp(-1(T-t)(\lambda_{235}+k_{235}))) + \\ 6\lambda_{232}/(\lambda_{232}+k_{232})({}^{232}\text{Th}/{}^{3}\text{He})_{T}(1-\exp(-1(T-t)(\lambda_{232}+k_{232})))$$
(eq. 5)
where
$$k_{238} = -1\ln(({}^{238}\text{U}/{}^{3}\text{He})_{0}/({}^{238}\text{U}/{}^{3}\text{He})_{T})/(T-t) - \lambda_{238};$$
(eq. 6)
$$k_{235} = -1\ln(({}^{235}\text{U}/{}^{3}\text{He})_{0}/({}^{235}\text{U}/{}^{3}\text{He})_{T})/(T-t) - \lambda_{235};$$
(eq. 7)

$$k_{232} = -1\ln((^{232}\text{Th}/^{3}\text{He})_{0}/(^{232}\text{Th}/^{3}\text{He})_{T})/(T-t) - \lambda_{232}$$
(eq. 8)

where $\lambda_{238}(1.55 \times 10^{-10} \text{ y}^{-1})$, $\lambda_{235}(9.85 \times 10^{-10} \text{ y}^{-1})$ and $\lambda_{232}(4.95 \times 10^{-11} \text{ y}^{-1})$ are the decay constants for ²³⁸U, ²³⁵U, and ²³²Th, respectively, and k₂₃₈, k₂₃₅ and k₂₃₂ are the continuous transport coefficients for the U-Th-He system in DMM. The k-value is the difference in transport coefficients for U (or Th) and He, and is related to the difference in bulk partition coefficients between U (or Th) and He; negative k-values in the model indicate that He is transported from the mantle more efficiently than U and Th.

When $k_{238,235,232}$ equals zero, equations 5-8 describe closed-system ³He/⁴He evolution. It is assumed that the undegassed mantle has been closed to degassing since 4.4 Ga. Thus, the closed-system model starts at time T=4.4 Ga, where "t" is the time before present day (and T-t equals elapsed time). The ⁴He/³He_T (initial ratio) of primitive mantle is unconstrained, but is assumed to be 5,995 (or 120 Ra) and is assumed to increase to 7,224 (or ~100 Ra) today. The ²³²Th/²³⁸U₀ (present-day) ratio of primitive mantle is assumed to be 4.05 (McDonough and Sun, 1995). The primitive mantle ²³⁸U/³He₀ (present-day) ratio is then 70. From the ²³⁸U/³He₀ ratio, the ³He can be calculated by employing a present-day primitive mantle U concentration of 0.0203 ppm (McDonough and Sun, 1995), and is 7.3×10¹¹ atoms/g.

The model for the ${}^{3}\text{He}/{}^{4}\text{He}$ and ${}^{238}\text{U}/{}^{3}\text{He}$ of DMM starts at time T=4.4 Ga. It is assumed that DMM and primitive mantle had the same composition (${}^{4}\text{He}/{}^{3}\text{He}_{T}$, ${}^{238}\text{U}/{}^{3}\text{He}_{T}$ and ${}^{232}\text{Th}/{}^{238}\text{U}_{T}$) at 4.4 Ga, and that DMM began forming immediately by melt extraction from primitive mantle starting at 4.4 Ga. It is further assumed that DMM has evolved to exhibit present-day ${}^{232}\text{Th}/{}^{238}\text{U}_{0}$ and ${}^{4}\text{He}/{}^{3}\text{He}_{0}$ values of 2.55 (similar to the value for average DMM in Workman and Hart [2005]) and ~89,900 (8 Ra), respectively. Equations 5 through 8 are then solved for ${}^{238}\text{U}/{}^{3}\text{He}_{0}$, which is calculated to be ~54,000. For this solution, k₂₃₈ and k₂₃₅ are both -1.51×10⁻⁹ y⁻¹ and k₂₃₂ is -1.41×10⁻⁹ y⁻¹. If a DMM U concentration of 0.0032 ppm is assumed (Workman and Hart, 2005), the ${}^{3}\text{He}$ of present-day DMM is calculated to be 1.5×10⁸ atoms/g.

The ³He abundance in DMM calculated with the continuous depletion model is within a factor of 3 to 20 of the ³He abundances inferred from MORB samples and ³He flux from ridges (see Ballentine et al. [2002] for summary). Assuming 10% melting of the mantle source, ³He concentrations for DMM were derived from CO₂ concentrations and canonical mantle $CO_2/^3$ He ratios in MORB melt inclusions ($4.52 \times 10^8 \pm 1.93$ atoms ³He/g (Saal et al., 2002), "popping" rock (> 2.69×10^9 atoms ³He /g [Moreira et al., 1998]), and flux of ³He out of mid-ocean ridges (1.18×10^9 atoms ³He /g [Farley et al., 1995; Ballentine et al., 2002]).

The increase in 238 U/³He in DMM over the age of the earth (Table 2) requires that He is less compatible than U and Th during melting of this mantle reservoir. However, the assumption of increased 238 U/³He in DMM over time is realistic given the lherzolitic lithology of DMM, a mantle reservoir that hosts an estimated cpx modal abundance of ~13% (Workman et al., 2005). Results from a recent helium partitioning study (Heber et al., 2007) are consistent with helium being less compatible than U and Th (assuming U and Th partition coefficients from a recent compilation [Kelemen et al., 2003]) during mantle melting of a lherzolite lithology. However, helium partitioning during mantle melting is a controversial subject. Parman et al. (2005) reported olivine-melt partition coefficients for helium suggesting that helium may be more compatible than U and Th during melting of a cpx-poor lherzolite or harzburgite. However, Heber et al. (2007) report values that are over an order of magnitude smaller (less compatible), suggesting the helium may be more compatible than U and Th only when melting cpx-poor harzburgites or dunites. The discrepancy in olivine-melt helium partition coefficients between these two studies is not yet resolved.

Finally, the continuous transport equations can be written to calculate the concentrations of any element in DMM at any time in earth's history, assuming that DMM formed by continuous depletion of BSE starting at 4.4 Ga:

$$X_{DMM,t} = X_{BSE,0}(exp(-1(\alpha_X)[T-t]))$$
 (eq. 9)
where
 $\alpha_X = -1\ln(X_{DMM,0}/X_{BSE,0})/(T)$ (eq. 10)

where α_X is proportional to the transport of element X out of DMM over time; $X_{BSE,0}$ and $X_{DMM,0}$ are the present-day concentrations of element X in BSE and DMM, respectively; "t" is time before the present day and T = 4.4 Ga, and $X_{DMM,t}$ is the concentration of element X in DMM at any time "t" before the present day. Using equations 9 and 10, concentrations of trace elements that are known in BSE and present-day DMM can be used to calculate their time-dependent concentrations in DMM. For example, the present-day Ti concentrations in DMM (716 ppm) and BSE (1205 ppm) yield an α_{Ti} value of 1.18×10^{-10} y⁻¹ in equation 10. Thus, using this α_{Ti} value and solving for Ti_{DMM,t} in equation 10, the concentration of Ti in DMM can be calculated at any time in earth's history. More incompatible elements have larger values for α . For example, $\alpha_{Th}=5.25\times 10^{-10}$ y⁻¹ and $\alpha_{U}=4.20\times 10^{-10}$ y⁻¹, where Th is more incompatible than U. In Table 2, the abundances of several trace elements in DMM are provided at various times in earth's history.

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Figure 1. Titanium, tantalum and niobium (TITAN) are highly enriched in ocean island basalts (OIBs) with high ³He/⁴He. Top panel: Primitive mantle (PM) normalized trace element data (spidergrams) for basalts with the highest ${}^{3}\text{He}/{}^{4}\text{He}$ (>30 Ra) from 4 different hotspots. All high ³He/⁴He samples exhibit TITAN enrichment relative to neighboring elements on the spidergrams. Middle panel: Spidergrams for the low ${}^{3}\text{He}/{}^{4}\text{He}$ (<9 Ra) mantle endmembers. EM1 lavas are from Pitcairn, EM2 lavas are from Samoa, HIMU lavas are from Mangaia and Tubuaii. Spidergrams for the mantle endmembers and MORB are from Hart and Gaetani (2006) and (for EM2) Jackson et al. (2007b). TITAN enrichment is not as pronounced, or is absent, in the average spidergrams of the low ${}^{3}\text{He}/{}^{4}\text{He}$ mantle endmembers. Bottom panel: Spidergrams for DMM (Workman and Hart, 2005) and continental crust (Rudnick and Gao, 2003) indicate that these two shallow earth reservoirs are not entirely geochemically complementary: TITAN depletion in the continents is not balanced by corresponding enrichment in DMM (see Fig. 5 for mass balance details). However, the TITAN-enrichment in high ³He/⁴He lavas suggests that these lavas sample a (deeper) mantle reservoir that hosts the TITAN that is "missing" in the shallow geochemical reservoirs, DMM and continental crust. Sources of the ${}^{3}\text{He}/{}^{4}\text{He}$ data for the mantle endmembers and the high ³He/⁴He lavas can be found in the text and in Table 1.





Figure 2. Relationships between TITAN anomalies, ¹⁸⁷Os/¹⁸⁸Os and ³He/⁴He in representative hotspot lavas. High ³He/⁴He lavas all have elevated Ti/Ti* (top panel), Nb/Nb* (middle panel) and ¹⁸⁷Os/¹⁸⁸Os (bottom panel). With increasing positive TITAN anomalies and 187 Os/ 188 Os ratios in the OIBs, the maximum observed 3 He/ 4 He increases. All samples with high ${}^{3}\text{He}/{}^{4}\text{He}$ (>20 Ra) host radiogenic ${}^{187}\text{Os}/{}^{188}\text{Os}$ (> 0.130). The model curves describe mixing between the high ${}^{3}\text{He}/{}^{4}\text{He}$ (low TITAN anomaly) peridotitic and low ³He/⁴He (high TITAN anomaly) refractory eclogite portions of ancient oceanic plates. More recently isolated plates, and/or more contribution from eclogite, both tend to generate lower ${}^{3}\text{He}/{}^{4}\text{He}$ ratios in the peridoite-eclogite mixture over time. The model curves are not meant to describe the global OIB array. Instead, the model curves are only intended to constrain the mixing proportions of eclogite and peridotite in the high ${}^{3}\text{He}/{}^{4}\text{He}$ mantle sampled by the highest ${}^{3}\text{He}/{}^{4}\text{He}$ lavas (>30 Ra). In the model, more recently isolated plates (the 1 Ga plate example is shown) can generate TITAN anomalies, but cannot generate high ${}^{3}\text{He}/{}^{4}\text{He}$. However, the mixing model does match the ³He/⁴He ratios, moderately radiogenic ¹⁸⁷Os/¹⁸⁸Os and large positive TITAN anomalies observed in the highest ³He/⁴He lava (sample SEL 97 from Iceland, see Table 1) when the eclogitic (20-25%) and peridotitic (75-80%) portions of a 3 Ga plate are mixed. Mixing is marked at 10% intervals, with increasing contribution from eclogite. The most extreme lava in the figure (sample SEL 97) is the highest ${}^{3}\text{He}/{}^{4}\text{He}$ lava available that also has a complete suite of trace elements measured by ICP. Unfortunately, ¹⁸⁷Os/¹⁸⁸Os is not available for this sample, and the ¹⁸⁷Os/¹⁸⁸Os of the highest ${}^{3}\text{He}/{}^{4}\text{He}$ mantle is instead approximated using lavas (with ${}^{3}\text{He}/{}^{4}\text{He} > 30$ Ra) from Hawaii and Samoa. Samoan post-erosional (from Savai'i) data are excluded; all ¹⁸⁷Os/¹⁸⁸Os data in the figure are >50 ppt Os, except Cape Verde (>10 ppt) and Pitcairn (>20 ppt). Data sources for the highest ${}^{3}\text{He}/{}^{4}\text{He}$ mantle endmembers can be found in Table 1 and Appendix A. Nb/Nb*=Nb_N/(Th_N × La_N)^{0.5} and Ti/Ti* = Ti_N/(Sm × Tb)^{0.5}, where N means normalized to primitive mantle.


Figure 3. Time evolution of ³He/⁴He and ¹⁸⁷Os/¹⁸⁸Os of DMM, ancient subducted (asthenosperic) DMM and eclogite. Top panel: At 4.4 Ga, DMM starts forming by melt extraction from primitive mantle to form continental and oceanic crust, and the ${}^{3}\text{He}/{}^{4}\text{He}$ trajectories of DMM and the (hypothetical) primitive undegassed mantle separate immediately. Due to continuous depletion by melt extraction, DMM evolves low ${}^{3}\text{He}/{}^{4}\text{He}$. Portions of the ancient (asthenospheric) DMM, coupled to the bottomside of downgoing slabs, are sent into the lower mantle throughout geologic time (model curves are shown at 1 Ga intervals, and this is not intended to imply that the isolation of oceanic plates is episodic). These isolated peridotite portions of the downgoing slabs are modeled as having exactly the same U, Th/U, ²³⁸U/³He and ³He/⁴He as ambient DMM at the time of isolation, and they preserve higher ³He/⁴He than DMM due to their isolation from further melt depletion. Also shown is the concomitant subduction of the eclogitic portions of the same slabs, which begin with the same ³He/⁴He as DMM at the time of subduction; the U and Th of the subducted refractory eclogite are from McDonough (1991) and shown in Table 2. To simulate degassing, the 238 U/ 3 He of the eclogite is increased by a factor of 1000 (Parman et al., 2005) relative to the contemporary DMM composition, and as a result the ³He/⁴He ratios of the subducted eclogites rapidly decrease. Bottom Panel: The ¹⁸⁷Os/¹⁸⁸Os of Primitive Mantle (0.130; Meisel et al., 2001) and DMM (0.125; Standish et al., 2002) are very similar, thus ancient subducted asthenospheric DMM is well-approximated by the trajectory of DMM in the figure (a continuous depletion model is not used to describe the evolution of ¹⁸⁷Os/¹⁸⁸Os of DMM). The present-day ¹⁸⁷Re/¹⁸⁸Os of the refractory eclogite is from Becker (2000), and this composition generates increasingly radiogenic present-day ¹⁸⁷Os/¹⁸⁸Os for earlier isolation times. Refer to Appendix B and Table 2 for all parameters used in the model. When 20-25% of a refractory eclogite from an oceanic plate subducted at 3 Ga is mixed together with 75-80% of the asthenospheric portion of the same plate, the models generate the ${}^{3}\text{He}/{}^{4}\text{He}$ and ${}^{187}\text{Os}/{}^{188}\text{Os}$ of the mantle sampled by the highest ${}^{3}\text{He}/{}^{4}\text{He}$ (>30 Ra) OIBs (see Fig. 2 for mixing results). Note that the very radiogenic Os of the eclogite is vastly reduced in the mixture due to the very low Os contents of the eclogite (6 ppt) and the high Os contents of the isolated DMM peridotite (3,000 ppt).



Figure 4. Spidergram of McDonough's (1991) refractory eclogite used in the modeling, including spidergrams demonstraing the time-dependent trace element composition of DMM. In the mixing models shown in Fig. 2, the asthenospheric DMM and refractory eclogitic portions of a 3 Ga subucted plate are mixed together such that the final mixture has 20-25% eclogite. This mixing calculation is shown in the spidergram of this figure, where 20% refractory eclogite has been added to the composition of DMM at 3 Ga (See Table 2 for compositions of the refractory eclogite and the 3 Ga DMM peridotite). Excluding Sr (see text for discussion of shallow lithospheric contamination, and the resulting Sr-anomalies in some Icelandic lavas), the shape of the hybrid peridotite-eclogite spidergram, including the positive Nb and Ti anomalies, is very similar in shape to the spidergram of the high ³He/⁴He Icelandic lava. This similarity is consistent with the hybrid peridotite being the source of the high ³He/⁴He mantle sample by OIBs.



Figure 5. Spidergram of a reservoir composed of recycled oceanic plates (called PLATE) hosting the elements missing in the CC and DMM reservoirs. If BSE (McDonough and Sun, 1995) is chondritic, then CC (continental crust; Rudnick and Gao, 2003) and DMM (depleted MORB mantle; Workman and Hart, 2005) are not wholly complementary reservoirs in the earth. Regardless of the relative proportions of CC and DMM, there is a shortage of the elements Ti, Ta and Nb in the silicate earth. The trace element composition of a PLATE reservoir is calculated to balance the budget of elements missing in CC and DMM (the PLATE spidergram is calculated such that the three reservoirs, PLATE, CC and DMM, sum to the total silicate mantle). The PLATE reservoir is thus composed of everything in the silicate earth that is not in the DMM and CC reservoirs, and is likely composed of mostly subducted plates and (a small portion of) sediment. The plotted spidergram of the PLATE reservoir assumes that it constitutes 40% of the mass of the silicate earth, the minimum mass of oceanic crust and lithosphere subducted over the past 2.5 Ga. If the PLATE reservoir constitutes more than 40% of BSE, its spidergram becomes more depleted (note direction of arrows in figure), but it will still exhibit positive TITAN anomalies. The PLATE spidergram shares the positive TITAN anomalies with the high ³He/⁴He OIB lavas, consistent with the hypothesis that high ${}^{3}\text{He}/{}^{4}\text{He}$ OIBs sample the domain of the PLATE reservoir composed of subducted, refractory, rutile-bearing eclogites.

Hotspot Volcano Sample No. rock type	lceland Selardalur SEL 97 Tholeiite	Samoa Ofu OFU-04-06 Alkali basalt	Hawaii Loihi KK18-8 Tholeiite	Galapagos Fernandina NSK 97-214 Tholeiite
SiO ₂	45.94	44.84	48.43	49.44
Al ₂ O ₃	9.72	11.04	12.62	14.77
TiO	1.19	4.95	2.88	3.12
FeOT	10.37	12.78	11.90	11.3
MnO	0.15	0.18	0.18	0.19
CaO	10.15	12.42	12.82	11.6
MgO	18.74	9.81	8.06	6.41
K ₂ O	0.05	1.14	0.50	0.39
Na ₂ O	1.18	2.24	2.34	2.46
P_2O_5	0.06	0.59	0.50	0.31
⁸⁷ Sr/ ⁸⁶ Sr	0 703465	0 704584	0.703680	0 703290
¹⁴³ Nd/ ¹⁴⁴ Nd	0.512969	0.512827	0.512945	0.512937
²⁰⁶ Ph/ ²⁰⁴ Ph	18 653	19 189	18 448	19 080
²⁰⁷ Ph/ ²⁰⁴ Ph	15 473	15 571	15.447	15 537
²⁰⁸ Ph/ ²⁰⁴ Ph	38 453	39 202	18 189	38 710
³ He ^A He	37.7	33.8	32.3	30.3
110/110			02.0	
Ni	759	201	80	49
Cr	1728	533	384	155
V	203	364	350	366
Ga		21		26
Cu Za		80 130		83
Cs	0.02	0.33	0.09	0.07
Rb	0.915	25.64	9.12	7.60
Ba	19.03	248.9	123.1	89.0
Th	0.382	3.956	0.994	1.310
U	0.072	0.991	0.285	0.380
Nb	6.50	50.56	16.35	20.00
la	0.434	3.84	12.02	1.50
La	3.85	37.93	13.82	14.1
Ph	0.38	2.61	1 25	0.93
Pr	1.23	9.99	4.62	4.23
Nd	5.80	44.08	20.65	19.5
Sr	173	599	349	351
Zr	57	293	125	155
Hf	1.65	7.56	3.48	4.15
Sm	1.78	10.75	5.33	5.77
Eu	2.05	0.03	1.63	2.04
Th	0.35	1.40	0.87	1.03
Dv	2.16	7.34	4.89	6.08
Ho	0.42	1.23	0.92	1.21
Y	10.03	30.29	21.87	30.1
Er	1.06	2.70	2.24	3.04
Tm	0.15	0.32	0.29	0.41
YD	0.88	1.66	1.69	2.46
Lu	0.13	0.23	0.24	0.30
50	33.5	51.2	55.4	44.5

Supplementary Table 1. Major and trace element and isotopic composition of the highest ³He/⁴He lavas from 4 hotspots.

Trace element data (excluding Ni, Cr, V, Ga, Cu, Zn and Sc) are by ICP-MS and are here reported for the first time the on the high ³He/⁴He samples from Hawaii, Iceland and Samoa; the trace element data on the Galapagoes sample are reported elsewhere. (Saal et al., 2007). The isotope data for all four lavas are also reported elsewhere (Hilton et al., 1999; Jackson et al., 2007a; Saal et al., 2007; Kurz et al., 1983; Staudigel et al., 1984). Major element data for the Samoan lava are reported here for the first time; major element data for the Hawaii, Iceland and Galapagoes samples are available elsewhere (Hilton et al., 1999; Saal et al., 2007; Frey and Clague, 1983). Samoan major element (and Ni, Cr, V, Ga, Cu, Zn and Sc) data reported here were measured by XRF. Trace element (for Samoa, Hawaii and Iceland) data reported here were analyzed by ICP-MS. Major and trace element data reported here were analyzed at the Geoanalytical Lab at Washington State University. Internal and (estimated) external precision for major and trace element analyses at the Geoanalytical Lab are given in Jackson et al. (2007b). The Loihi sample was powdered in Tungsten-carbide, and Ta is therefore not reported for this sample.

Table 2. Components used to model I	He, Os and T	ITAN anon	alies in t	he high ³ h	He/ ⁴ He man	tle sample	d by OIB	ŝ												
	²³⁸ U/ ³ He ¹	зне	³ He/ ⁴ He	⁴ He	U ²³² Th/ ²³⁸ U	¹⁸⁷ Re/ ¹⁸⁸ Os	ő	0 ⁸¹ /SO ¹⁸⁸ 0	s Th	5	qN	La	s	PN	Zr	Sm	E	Ξ	Gd	q1
		atoms/g ¹	(Ra) ⁵	atoms/g			(ppt)		(mdd)	(mdd)	(mdd)	(mdd)	(mqq	(mdd)	(mdd)) (mqq) (mqc) (mqq) (mdc	(mqq
Undegassed Mantle ²	7.00E+01	7.3E+11	9.66	5.3E+15	4.05	0.396	3400	0.130	0.0795	0.0203	0.658	0.648	19.9	1.25	10.5	0.406 0	154	1205	0.54 0	660.0
Present-day DIMIN	5.40E+04	1.5E+08	8.0	1.3E+13	2.55	0.314	3000	0.120	0.0079	0.0032	0.149	0.192	7.664	0.581	5.08	0.239 0	0.096	716	0.36	0/0/0
Pure peridotite component ⁴ DMM peridotite (isolated at 1 Ga)	1.19E+04	1.0E+09	12.8	5.8E+13	2.81	0.314	3000	0.125	0.0134	0.0049	0.207	0.253	9.52	0.692	5.99	0.270 0	107	908	0.39 (0.076
DMM peridotite (isolated at 2 Ga)	2.63E+03	7.1E+09	27.7	1.8E+14	3.12	0.314	3000	0.125	0.0226	0.0074	0.291	0.334	11.83	0.823	7.07	0.304 0	0.119	202	0.43	0.082
DMM peridotite (isolated at 3 Ga)	5.80E+02	4.9E+10	55.6	6.3E+14	3.47	0.314	3000	0.125	0.0381	0.0113	0.409	0.440	14.69	0.980	8.34	0.343 0	132	1021	0.48 0	0.089
DMM peridotite (isolated at 4 Ga)	1.28E+02	3.4E+11	88.6	2.7E+15	3.85	0.314	3000	0.125	0.0644	0.0172	0.574	0.580	18.25	1.166	9.83	0.387 0	.148	1149	0.52 0	0.096
<u>Pure eclogite component^{6,7}</u> Refrac. eclogite (isolated at 1 Ga)	1.19E+07	9.4E+06	0.026	2.6E+14	2.81	325	9	5.5	0.11	0.044	4.5	1.7	83	5.0	50	1.9	0.7	0200	2.7	0.5
Refrac. eclogite (isolated at 2 Ga)	2.63E+06	4.2E+07	0.053	5.8E+14	3.12	325	9	11.0	0.11	0.044	4.5	1.7	83	5.0	50	1.9	0.7	0600	2.7	0.5
Refrac. eclogite (isolated at 3 Ga)	5.80E+05	1.9E+08	0.137	1.0E+15	3.47	325	9	16.6	0.11	0.044	4.5	1.7	83	5.0	50	1.9	0.7	0600	2.7	0.5
Refrac. eclogite (isolated at 4 Ga)	1.28E+05	8.7E+08	0.385	1.6E+15	3.85	325	9	22.3	0.11	0.044	4.5	1.7	63	5.0	50	1.9	0.7 1	0500	2.7	0.5
The understand of the processor of the processor that the undegassed mattle has been a dr (McDonough and Sun, 1995) and a ³ He (in Class and Goldstein (2005) and ~1x10)	s provided in osed system t concentration	this table. to all eleme of 7.3x10 ¹¹ the D" layer	nts, volati atoms/g, r in Tostik	le and non a ³ He valu	-volatile, sir that is w ofmann [200	nce 4.4 Ga. ell within the 06]). To sim	The ²³⁶ U	l ¹ ³ He of un uggested f 3% degass	degassed for the un sing of the	degasas slab in	is base ed mar the sub	ed on a the in the duction	primitiv brimitiv e litera zone, 1	e mantl ture (1. the ²³⁸ U	e U cor 1x10 ¹¹ /ÅHe va	ncentrat atoms/g	ion ion	j,		
(refractory) eclogite is taken to be a facto	or of 1,000 tim	ies greater t	than the 2	³⁸ U/ ³ He of	the contern	porary DMN	1 peridoti	te (Parmar	n et al., 2	005).										
² The trace element concentrations for u are from McDonough and Sun (1995).	ndegassed m	antle (exclu	Iding Os	and He) ar	e from McD	onough and	Sun (19	95). The	Osal /SO/a	s is fror	n Meise	et al.	(2001),	and the	e Os an	d ^{18/} Re/	SOs			
³ The trace element concentrations (excl ⁴ I lot the continuous transport equation	luding He and	d Os) for pre	esent-day	DMM are	from Workn	nan and Har	rt (2005).	The ¹⁸⁷ O	s/ ¹⁸⁸ Os o'	f DMM is	from t	standist	during	(2002).	due jo e	dottot	oi pue	- dotion		
all other trace elements concentrations (except Os) ar	e calculated	the sam	e way. He	fium concer	ntrations are	derived	from the c	alculated	238U/ ³ H	and U	concel	tration	o.						
⁵ At 4.4 Ga, the mantle is assumed to ha	ave had a ³ He	A ⁴ He ratio o	f 120 Ra,	a ratio tha	t is similar t	o the preser	nt-day atr	nosphere	of Jupiter	(Niema	nn et al	, 1996	; given	the						
assumed ""U/"He and ""Th/""U primitiv Appendix B. the "He/ ⁴ He ratios of isolater	/e mantle rati d DMM perido	os, closed s otite (asther	ystem ev losphere)	olution yiel throughou	ds a preser It time are ti	it-day 'He/'F he same as	He ratio o ambient	of 99.6 Ra upper mar	(see App otte DMM	endix B) at the ti	. Using	ubducti	ntinuou on and	s transp isolatio	oort equ	ations i	c			
⁶ The ¹⁸⁷ Re/ ¹⁸⁸ Os and Os concentrations recycled eclogites are then calculated as	for refractory suming differ	r eclogite ar ent isolation	e the med times of	lian values the same	reported for eclogite.	or metamorp	hosed m	etabasalts	s in Becke	er (2000)	. The	arious	187 Osl ¹⁸	⁸ Os val	ues shi	own for	the			
⁷ Four of the trace elements for the refract	ctory eclogite	(Nb, La, Zr	and Ti) w	ere provid	ed explicitly	in McDonot	ugh (199	1). Th, Sr,	, Nd, Sm,	Eu and	Gd wei	e estim	ated fro	om Fig.	5 in the	same				
source. U concentrations for the refracto concentration provided; Tb (used to calcu	ory eclogite ar ulate Ti/Ti*) is	e not provide	led by McE	Donough (1)	(1991). U c (991), and it	oncentration is concentral	is are est	timated as e refractor	suming a	Th/U ra	tio of 2 y a line	5 and u ar (on a	Primiti	e Th ve Man	tle nom	nalized	basis)			
extrapolation from Gd and Sm, which are (proxies for Nb and Ti anomalies) in the e elements of interest (Nb, Th, Ti, Sm, U, e Sr = 88.7 ppm, Zr = 152 ppm, Sm = 2.16	 provided. Tl eclogite datas etc.), the aver ppm., Ti = 10 	he trace ele set of Becke age trace el 0,080 ppm.	ment con r et al. (2 lement co	nposition o 000). Con mposition	f McDonou sidering onl of the TITA	gh's (1991) } y the eclogit N-enriched o	hypotheti es with T eclogites	cal eclogit 'l/Sm > 30' in Becker	e is quite 00 and NI et al. (20	similar b/Th>24 00) is: 7	to the e, and e h = 0.1	dogite: (cluding 59 ppn	with th all san ; U = 0	nples th 165 pp	est Nb/T nat lack m; Nb	h and 7 data fo = 10 pp	'i/Sm r all the m; Nd :	= 7.58 µ	;mqc	