

1 **Geochemistry of abyssal peridotites (Mid-Atlantic Ridge, 15°20'N,**
2 **ODP Leg 209): Implications for fluid/rock interaction in slow**
3 **spreading environments**

4
5 [Changes are marked in red]

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34 **Abstract**

35 Abyssal peridotite from the 15°20'N area of the Mid-Atlantic Ridge show
36 complex geochemical variations among the different sites drilled during ODP Leg 209.
37 Major element compositions indicate variable degrees of melt depletion and
38 refertilization as well as local hydrothermal metasomatism. Strongest evidence for
39 melt-rock interactions are correlated Light Rare Earth Element (LREE) and High Field
40 Strength Element (HFSE) additions at sites 1270 and 1271. In contrast, hydrothermal
41 alteration at Sites 1274, 1272, and 1268 causes LREE mobility associated with minor
42 HFSE variability, reflecting the low solubility of HFSE in aqueous solutions. Site 1274
43 contains the least-altered, highly refractory, peridotite with strong depletion in LREE
44 and shows a gradual increase in the intensity of isochemical serpentinization; except
45 for the addition of H₂O which causes a mass gain of up to 20 g/100 g. The formation
46 of magnetite is reflected in decreasing Fe²⁺/Fe³⁺ ratios. This style of alteration is
47 referred to as rock-dominated serpentinization. In contrast, fluid-dominated
48 serpentinization at Site 1268 is characterized by gains in sulfur and development of
49 U-shaped REE pattern with strong positive Eu anomalies which are also
50 characteristic for hot (350 to 400 °C) vent-type fluids discharging from black smoker
51 fields. Serpentinites at Site 1268 were overprinted by talc alteration under static
52 conditions due to interaction with high a_{SiO_2} fluids causing the development of smooth,
53 LREE-enriched patterns with pronounced negative Eu anomalies. These results
54 show that hydrothermal fluid-peridotite and fluid-serpentinite interaction processes
55 are an important factor regarding the budget of exchange processes between the
56 lithosphere and the hydrosphere in slow spreading environments.

57

58 Keywords: Serpentinization, slow spreading ridges, abyssal peridotite, hydrothermal
59 alteration, geochemistry, Ocean Drilling Program Leg 209

60

61 **Introduction**

62 Abyssal peridotites represent sections of the upper mantle which are exposed
63 on the seafloor in **and near** fracture zones and along slow and ultraslow spreading
64 ridges due to tectonic faulting associated with extension and crustal thinning. Current
65 estimates indicate that ridges with a spreading rate of less than 20 mm yr⁻¹ comprise
66 about one third of the 55,000 km global ridge system (Dick et al., 2003). Prominent
67 examples include the Southwest Indian Ridge (e.g., Sauter et al., 2004) the Gakkel
68 Ridge in the Arctic ocean (e.g., Cochran et al., 2002) and segments of the Mid-
69 Atlantic Ridge (MAR; Lagabrielle et al., 1998 and references therein). Hence, upper
70 mantle rocks are important components of the shallow oceanic lithosphere and an
71 integral part of hydrothermal interaction processes at divergent plate margins.

72 The significance of hydrothermal interaction processes involving ultramafic
73 rock has become increasingly evident. It was realized that fluid-peridotite interaction
74 has important consequences for the rheology of the oceanic lithosphere (Escartín et
75 al., 1997), geochemical budgets of the oceans (Snow and Dick, 1995; Thompson and
76 Melson, 1970) and microbial processes at and below the seafloor (Alt and Shanks,
77 1998; Kelley et al., 2005).

78 The most prominent manifestations of these processes are active, ultramafic-
79 hosted hydrothermal systems discharging hot (350 to 400 °C) metal-rich fluids and
80 generating black-smoker chimney fields on the seafloor (e.g., Logatchev and
81 Rainbow Sites, Douville et al., 2002; Mozgova et al., 1999). Lower temperature (40-
82 90 °C) hydrothermal venting has been recently discovered in the off axis environment

83 (Lost City Hydrothermal Field, Früh-Green et al., 2003; Kelley et al., 2001, 2005). On
84 a larger scale, chemical (CH₄ and Mn) and turbidity anomalies in the water column
85 along slow and ultraslow spreading ridges indicate that hydrothermal systems are
86 common in this environment (Charlou et al., 1993; Edmonds et al., 2003; German et
87 al., 1998). Some of these occurrences are clearly associated with avolcanic
88 spreading (Bach et al., 2002), while others have high ratios of CH₄ to Mn (Charlou et
89 al., 1993), indicating a significant contribution from peridotite-hosted hydrothermal
90 systems.

91 In general, hydrothermal alteration of peridotite is inferred to be dominated by
92 hydration reactions of olivine and pyroxene that lead to the formation of serpentine
93 minerals (“serpentinization *senso stricto*”; e.g., Janecky and Seyfried, 1986; Komor et
94 al., 1985; Miyashiro, et al., 1969; O’Hanley, D.S., 1996; Wicks and Whittaker, 1977).
95 However, Ocean Drilling Program (ODP) Leg 209 drill core of abyssal peridotite from
96 five sites along the MAR in the 15°20’N area show a remarkable diversity of
97 alteration intensities and mineral assemblages indicating that fluid-peridotite, and
98 fluid-serpentinite, interaction can take place at a wider range of temperature and
99 redox conditions than commonly appreciated (**Figs. 1 and 2, Table 1**; Bach et al.,
100 2004).

101 The purpose of this paper is to investigate the geochemical systematics of
102 these abyssal peridotite samples in order to determine the influence of hydrothermal
103 alteration on their composition. Hence, this study contributes to the characterization
104 of the geochemical budget of the lithosphere at slow and ultraslow spreading ridges.

105

106 **[INSERT Figures 1 and 2 and Table 1]**

107

108 ***Local geology and hydrothermal alteration of peridotite in the 15°20'N area***

109 Abundant peridotite and gabbroic rocks are exposed on both flanks of the
110 spreading axis of the slow spreading MAR to the north and south of the 15°20'N
111 Fracture Zone (full rate: 25 mm yr⁻¹; Escartín et al., 2003; Kelemen et al., 2004;
112 Lagabrielle et al., 1998). In this area, widespread fluid-peridotite interaction in the
113 sub-seafloor is evident from extensive CH₄ and H₂ anomalies in the water column
114 and high-T hydrothermal discharge at the active Logatchev black smoker field (**Fig.**
115 **1**; Batuev et al., 1994; Bogdanov et al., 1997; Charlou et al., 1998).

116 The ODP Leg 209 drill sites 1268, 1270, 1271, 1272, and 1274 are located
117 within the MAR axial valley and on the immediate valley walls between 14°40'N and
118 15°40'N and between 18 and 156 m of basement dominated by ultramafic rocks have
119 been penetrated (**Figs. 1 and 2 and Table 1**). Overall, mantle deformation fabrics
120 are weakly developed and deformation was localized along ductile shear zones
121 formed under granulite to greenschist facies conditions and late brittle faults. Intact
122 blocks of peridotite with protogranular fabrics were preserved between these zones
123 and underwent tectonic rotation (Kelemen et al., 2004; Kelemen et al., submitted;
124 Shipboard Scientific Party, 2004).

125 Detailed accounts of the lithologies drilled during ODP Leg 209 are presented
126 in the ODP Initial Results Volume (Shipboard Scientific Party, 2004). The style and
127 intensity of hydrothermal alteration varies between the sites and the fluid-rock
128 interaction processes are controlled by a variety of factors including reaction kinetics,
129 temperature, modal mineralogy of the protolith, fluid composition, redox variations,
130 and metasomatic processes (Bach et al., 2004). Here, we present a summary of the
131 lithological characteristics and the principal controls on hydrothermal alteration in
132 order to provide a framework for the interpretation of the geochemical variations.

133 Hole 1274A (**Fig. 2a**) contains the “least altered” harzburgite and dunite. The
134 intensity of serpentinization ranges from highly altered (~60 vol% secondary
135 minerals) in the upper part of the hole to completely altered (>95 vol% secondary
136 minerals) in the lower portion that is also characterized by abundant fault zones.
137 Brucite-serpentine-magnetite is the common alteration assemblage in dunite
138 whereas brucite is rare in harzburgite. This indicates that brucite is stabilized in
139 orthopyroxene-poor peridotite whereas formation of serpentine from orthopyroxene in
140 harzburgite releases SiO₂ to the fluid promoting the reaction of brucite to serpentine
141 (Bach et al., 2004). Late-stage interaction with ambient seawater is documented by
142 aragonite veinlets and associated red halos of oxidative alteration (Fe-oxyhydroxide-
143 clay alteration). Variations in the $\delta^{18}\text{O}_{\text{aragonite}}$ indicate increasing formation
144 temperatures from 2°C near the seafloor to 15 °C at 90 mbsf (meter below sea-floor;
145 Bach and Paulick, 2004).

146 The topmost ~50 m of Hole 1272A (**Fig. 2b**) consist of a diverse lithological
147 assemblage including diabase, vesicular basalt, gabbro, serpentinized peridotite and
148 carbonate-cemented breccia with serpentinite clasts which has been interpreted as a
149 mega-breccia (Shipboard Scientific Party, 2004). Below this unit, serpentinized
150 harzburgite with minor dunite consists of serpentine-magnetite±brucite±iowaite
151 assemblages. Iowaite ($\text{Mg}_4[\text{OH}]_8\text{Fe}^{3+}\text{OCl} \times 1-4 \text{H}_2\text{O}$) has been described previously
152 from submarine mud volcanoes (Heling and Schwarz, 1992) and serpentinites at the
153 Iberian margin (Gibson et al., 1996). Iowaite in Hole 1272A represents the first
154 documented occurrence of this mineral in a mid-ocean ridge setting (Shipboard
155 Scientific Party, 2004); however, the presence of iowaite veins in Hess Deep
156 serpentinites is reported in Früh-Green et al. (2004). In Hole 1272A, iowaite formed
157 under oxidizing conditions from Fe-bearing brucite during late-stage, low temperature
158 alteration (Bach et al., 2004).

159 Holes 1271A and B (**Fig. 2c**) recovered a complex mixture of ultramafic and
160 mafic lithologies including melt-impregnated (commonly amphibole-bearing) dunite,
161 amphibole gabbro, troctolite, and gabbroic intrusions. In addition, the dunite contains
162 abundant disseminated chromite as well as irregular veinlets and individual chromite
163 pods (Shipboard Scientific Party, 2004). Petrographic evidence for melt-rock
164 interaction includes mm- to sub mm-scale network-like veins, commonly forming rims
165 around spinel crystals, which are now composed of chlorite. The serpentinization of
166 dunite and minor harzburgite is dominated by serpentine-brucite-magnetite
167 assemblages; however, relict olivine is common in amphibole-bearing assemblages.

168 Four holes were drilled at Site 1270 on the eastern rift valley wall (**Fig. 1**).
169 Holes 1270C and 1270D are located immediately adjacent to each other whereas
170 Holes 1270B and 1270A are located at ~300 m and ~500 m down slope to the west
171 (Shipboard Scientific Party, 2004). Hole 1270A (**Fig. 2d**) consists of serpentinized
172 harzburgite with minor dunite and Hole 1270B (**Fig. 2e**) consists of gabbro with minor,
173 completely talc altered, harzburgite. Holes 1270C and D (**Fig. 2f**) consist of
174 serpentinized harzburgite and dunite, however, relict olivine and orthopyroxene are
175 locally preserved. In these less serpentinized areas, pyroxene is replaced by talc and
176 tremolite whereas olivine is weakly serpentinized documenting an initial high-
177 temperature alteration stage ($T > 350$ to 400°C) where replacement of pyroxene
178 proceeds at higher rates than alteration of olivine (Allen and Seyfried, 2003; Bach et
179 al., 2004). Furthermore, there are abundant gabbroic intrusions which are commonly
180 the focus of intense ductile deformation. These shear zones were locally invaded by
181 mafic and differentiated melts as well as fluids of hydrothermal and magmatic origin
182 causing the formation of amphiboles, local occurrences of zircon and apatite
183 (Shipboard Scientific Party, 2004).

184 Hole 1268A (**Fig. 2g**) contains serpentinized and talc-altered harzburgite and
185 dunite in the upper section whereas gabbro dominates below 105 mbsf. Here,
186 serpentinites contain up to 3 vol% pyrite, which is otherwise rare in the altered
187 peridotite drilled during ODP Leg 209. Talc alteration of serpentinites is interpreted as
188 the result of Si-metasomatism and fluids with high a_{SiO_2} could have been derived from
189 alteration of pyroxene in peridotite or gabbro at depth (Bach et al., 2004). Talc
190 alteration occurred under static conditions as suggested by serpentinite
191 microtextures that are perfectly pseudomorphed and preserved (Bach et al, 2004;
192 Shipboard Scientific Party, 2004).

193

194 **Methods**

195 The geochemical data base consists of 168 analyses of peridotite drill core
196 samples derived from ODP Sites 1268, 1270, 1271, 1272, and 1274. Examples of
197 the typical rock types from these Sites are presented in **Table 2**. One group of 85
198 samples have been investigated on-board the JOIDES-Resolution during Leg 209
199 and have been analyzed for major and some trace elements by ICP-AES. Volatiles
200 (H_2O , CO_2 , and S) were determined by combustion and element analyses. These
201 data and details of the analytical procedures are documented in Shipboard Scientific
202 Party (2004).

203 A second set of 28 samples have been analyzed for major elements by XRF at
204 the Open University (ARL 8420+) and volatile elements (CO_2 , S, N) have been
205 determined by elemental analyzers at GFZ-Potsdam. Whole rock trace element
206 concentrations have been analyzed at the ISTEEM of Montpellier University (France)
207 on a quadrupole VG-PQ2 Inductively Coupled Plasma-Mass Spectrometry (ICP-MS)
208 following the procedure described in Ionov et al. (1992) and in Godard et al. (2000).

209 The method involves dissolution of 100 mg aliquots in a HF-HClO₄ mixture and
210 dilution by a factor of 1000 for the analysis of trace elements, except for more
211 concentrated elements such as Li, Cu, Ni, Co and Sc that were analyzed as a
212 different batch with a dilution factor of 4000. In and Bi were used as internal
213 standards during ICP-MS measurements. The REE, U, Th, Sr, Zr, Hf, Rb, Ba, Li, Cu,
214 Ni, Co and Sc concentrations were determined by external calibration using multi-
215 element standard solutions (Merck). To avoid memory effects due to the introduction
216 of concentrated Nb-Ta solutions in the instrument, Nb and Ta concentrations were
217 determined by using, respectively, Zr and Hf as internal standards. This technique is
218 an adaptation to ICP-MS analysis of the method described by Jochum et al. (1990)
219 for the determination of Nb by spark-source mass spectrometry. The limits of
220 detection, the procedural blank contributions and the values obtained for the
221 international standards JP-1 and PCC-1 are reported in **Table 3 (Appendix)**.

222 A third set of 55 samples have been analyzed for major element
223 concentrations by XRF at Bonn University (Philips PW1480) and volatiles (CO₂, S, N)
224 have been determined at Freiberg University (VARIO EL gas analyzer). Trace and
225 REE contents of these samples were determined by ICP-MS (HP4500 [Agilent] with
226 Cetac ASX-500 autosample) at Göteborg University. About 100 mg of rock powder
227 was digested using a HNO₃-HF mixture and diluted to a factor of 3800. Re and In
228 were used as internal standards. Calibration was performed using four different multi-
229 element standards (Merck and Agilent). Accuracy was monitored using rock
230 standards (JP-1, JB-1 and two in-house standards) treated as unknowns, and was
231 within 10% for all analyzed except Hf and Ta (25%). Rocks that gave below-detection
232 limit values for REE (28 samples) were re-analyzed for their REE content using an
233 alternative procedure. Sample digestion was similar to the first method, instead that
234 the final dilution factor was 500, and special care was taken to obtain low digestion

235 blanks, including the use of ultra pure acids (Romil®) and pre-leached sample vials
236 and bottles. To minimize matrix effects, a calibration standard was prepared by
237 spiking an aliquot of one of the samples with a 100 ppb REE solution. Drift was
238 monitored by analyzing JB-2 every four samples. This procedure allowed measuring
239 REE concentrations of down to 0.001 chondrite. Reproducibility was measured on
240 five duplicates and is on average 5%, but increases to 15% at element of <2.5 ppb in
241 the sample. Rock standards JP-1, UB-N, NIM-D, NIM-P, NIM-N, and SARM47 were
242 used to monitor accuracy and values are on average between 5 and 10% of literature
243 values (Korotev, 1996; Makishima and Nakamura, 1997; Pin and Joannon 1997).

244 In addition, the Fe²⁺ concentrations of 102 samples (second and third set, and
245 shipboard samples of Hole 1268A) were determined by standard titration methods at
246 the GeoForschungsZentrum Potsdam and Bonn University (Germany). The XRF,
247 ICP-MS, volatile, and Fe²⁺ data of the second and third set (83 samples) analyzed at
248 the Universities of Bonn, Montpellier, and Göteborg and the Open University are
249 documented in an electronic supplement to this paper (**Table 4**).

250 The composition of the primary phases has been determined by electron
251 microprobe analysis at the University of Köln (JEOL JXA-8900RL Superprobe;
252 acceleration voltage: 20 kV for olivine and 15 kV for other phases; beam current: 50
253 nA for olivine and 20 nA for other phases; focused beam). In total, about 440
254 individual analyses have been obtained from 13 samples of Hole 1274A. These data
255 are documented in a report to the ODP Scientific Results Volume for Leg 209 (Moll et
256 al., submitted). Furthermore, point counting data were obtained from the Hole 1274A
257 samples, in order to constrain the modal proportions of mineral phases (1000 points
258 on a 0.2 x 0.2 mm grid). These results were combined with the microprobe data in
259 order to determine the protolith composition of the samples prior to serpentinization
260 (**Table 5, Appendix**).

261 **[INSERT Table 2]**

262

263 **Results**

264 In MgO/SiO₂ vs. Al₂O₃/SiO₂ space (**Fig. 3**) a 'terrestrial array' represents the
265 successive magmatic depletion of a primitive mantle and highly depleted
266 compositions are characterized by low Al₂O₃/SiO₂ values (<0.01; Hart and Zindler,
267 1986; Jagoutz et al., 1979). A global data set of abyssal peridotite presented by Niu
268 (2004) shows that these samples follow a similar trend that is systematically off set to
269 lower MgO/SiO₂ values due to MgO loss during seafloor weathering (Niu, 2004;
270 Snow and Dick, 1995). The increasing Al₂O₃/SiO₂ with decreasing MgO/SiO₂ in this
271 data set are at least partially attributable to melt impregnation processes within the
272 thermal boundary layer (Niu, 2004).

273 The drilled peridotites from the 15°20'N area show a considerable
274 compositional diversity in MgO/SiO₂ vs. Al₂O₃/SiO₂ space that is related to a number
275 of different processes. Similar to the most refractory peridotites from oceanic
276 environments (Bodinier and Godard, 2003), most peridotites are strongly melt-
277 depleted with low Al₂O₃/SiO₂ values (<0.02). Harzburgites from Holes 1272A and
278 1274A cluster tightly around the melting trend and preserve the high MgO/SiO₂ ratios
279 typical of refractory mantle rocks (>1). However, most dunites from Hole 1274A and
280 from Site 1271 plot systematically above the mantle array since their ratio of
281 olivine_{modal}/(olivine_{modal} + pyroxene_{modal}) is close to 1. This compositional trend is well-
282 known from ophiolitic and orogenic massifs (e.g., Godard et al., 2000). Al₂O₃/SiO₂
283 values > 0.02 at MgO/SiO₂ values around 1.2 (corresponding to Fo₉₀ olivine) in
284 dunites from Site 1271 are likely due to addition of chromite. In contrast to these
285 melt-rock interaction trends, serpentinized dunites and harzburgites of Sites 1268

286 and 1270 have MgO/SiO₂ values below the terrestrial array (0.8 to 1). These cannot
287 be the result of melt–rock interaction but are either due to MgO loss during
288 interaction with seawater or addition of silica. Talc alteration of serpentinites at Hole
289 1268A causes a further decrease in MgO/SiO₂ ratios that is likely due to Si-
290 metasomatism, as will be discussed later.

291

292 **[INSERT Figure 3]**

293 **Sites 1274 and 1272**

294 Peridotites from Sites 1272 and 1274 have low concentrations in elements
295 such as Al (Al₂O₃: 0.1 to 1 wt%), Sc (2 to 9 ppm), and V (5 to 40 ppm; **Fig. 4 a and b**)
296 and are highly depleted in LREE with concentrations below 1% of the chondritic
297 values whereas HREE concentrations are gradually increasing from Gd to Lu
298 reaching Lu_N values of 0.3 (**Fig. 5a and b**). These compositions are similar to those
299 observed in the most depleted peridotites sampled in oceans and ophiolitic massifs
300 (Bodinier and Godard, 2003; Godard et al., 2000; Niu, 2004; Niu and Hekinian, 1997;
301 Parkinson et al., 1998; Pearce et al., 2002).

302 In Hole 1274A, where the least-altered peridotites were sampled, the positive
303 correlations of Al₂O₃ versus V, Sc, and Cr can be linked to the variations in the modal
304 proportions of pyroxene and olivine in the protolith since the dunites are more
305 refractory in character than the associated harzburgite (**Fig. 4 a, b, c**). Point counting
306 results confirm that the modal proportion of pyroxene (pyroxene_{modal}/[pyroxene_{modal} +
307 olivine_{modal}]) are correlated with the Al₂O₃ concentrations at this site (**Fig. 4e**). Since
308 the data from the more intensely altered sites follow the same geochemical trends it
309 can be inferred that Al, V, Sc, and Cr were immobile during hydrothermal processes.
310 The peridotites from Hole 1274A also define linear arrays showing a small decrease

311 in MgO and FeO_{total} and increase in SiO₂ with increasing Al₂O₃ (**Fig. 4 d, f, g**). This is
312 also consistent with the inferred relationship between modal pyroxene abundance
313 and bulk rock geochemistry.

314 The most prominent effect of serpentinization is the addition of water to the
315 rock and even the least-altered peridotites from Hole 1274A have H₂O contents of
316 ~10 wt% which increases to ~15 wt% for complete serpentinization (**Fig. 6a**).
317 Serpentinized dunites have even higher water contents (up to 17.5 wt%) due to the
318 formation of brucite. In terms of SiO₂ and MgO contents, most of the serpentinized
319 peridotites cluster around the composition of serpentine, however, some brucite-
320 bearing dunites have MgO concentrations up to 50 wt% (**Fig. 6b**). Another important
321 effect of serpentinization is the change in Fe-oxidation state (**Fig. 6c**). The samples
322 from Holes 1272A and 1274A cluster along a line defined by a constant Fe-budget
323 where changes in FeO and Fe₂O₃ concentrations are attributable to the oxidation of
324 ferrous iron in primary silicates to ferric iron in magnetite and/or serpentine. Data
325 from Hole 1274A show that with progressive alteration the FeO concentrations
326 decrease from 4 to 2 wt% and the most Fe₂O₃-rich samples have the Fe₂O₃/FeO-
327 ratio of magnetite. This indicates that for completely serpentinized samples the Fe-
328 budget is mainly controlled by the formation of magnetite.

329 The REE pattern of serpentinites from Holes 1272A and 1274A are largely
330 similar to the least-altered samples except from some variability in LREE
331 concentrations and the local development of a positive Eu-anomaly (**Figs. 5a, b**).
332 Also, the concentrations of HREE at Site 1272 are consistently below the
333 concentrations in least-altered samples from Site 1274 suggesting an even more
334 depleted precursor composition or dilution of REE concentrations due to mass
335 addition during serpentinization.

336

337 **[INSERT Figures 4, 5 and 6]**

338 **Sites 1270 and 1271**

339 The composition of serpentinites from Sites 1270 and 1271 is generally similar
340 to the serpentinites from Site 1274 and 1272 in terms of major element and most
341 trace elements (**Figs. 4 and 6**). However, chromite is abundant at Site 1271 which is
342 reflected in elevated Cr and FeO_{total} concentrations and lower SiO₂ contents (**Fig. 4c,**
343 **d, g; 6d**; Shipboard Scientific Party, 2004).

344 The abundance of mafic dikelets and thin section-scale melt impregnation
345 textures in some drill holes of Sites 1270 and 1271 is reflected in particular REE
346 systematics (**Fig. 5c, d**) and elevated concentrations of some incompatible elements
347 (Zr, Th, U; **Fig. 7**). In contrast to the LREE depleted pattern of the peridotites from
348 Holes 1274A and 1272A, samples from Holes 1270 B, C, and D (**Fig. 5c**) and most
349 samples from Hole 1271B (**Fig. 5d**) show flat to slightly LREE enriched pattern with
350 chondrite-normalized concentrations ranging from 0.2 to 10. These are almost
351 identical to analyses of mafic rocks sampled during ODP Leg 209 showing smooth,
352 flat to LREE enriched pattern (**Fig. 5c, d, e**). Furthermore, peridotites from Holes
353 1270C and D and 4 samples from Hole 1271B have elevated U contents ranging
354 from 0.4 to 1.4 ppm (**Fig. 6a**) and comparatively high concentrations of Th (0.02 to
355 0.46 ppm) and Zr (1 to 5 ppm, **Fig. 6b**). These characteristics suggest that the
356 primary REE signatures and Th, U, and Zr concentrations of these peridotite samples,
357 which probably had initial compositions similar to the refractory peridotite of Site 1274,
358 were modified due to interaction with mafic to differentiated melts.

359 Serpentinites from Hole 1270A, 1271A and two samples from Hole 1271B
360 deviate from these characteristics introduced by melt-rock interaction processes.
361 These samples show U-shaped REE patterns with positive Eu anomalies (**Figs 5, c,**

362 **d, e**) that are either due to the addition of plagioclase during melt-rock interaction
363 (e.g., Niu et al., 1997) or result from interactions with hydrothermal fluids similar to
364 the serpentinites at Hole 1268A.

365

366 **[INSERT Figure 7]**

367 **Site 1268**

368 The major and trace element concentrations of serpentinization at Hole 1268A
369 are similar to the other serpentinites sampled during ODP Leg 209. However, the
370 abundance of disseminated pyrite and pyrite veins is reflected in elevated S
371 concentrations (range: 0.1 to 2.1 wt%; average: 0.6 wt%, **Fig. 6d**) and low
372 Fe₂O₃/FeO ratios compared to serpentinites with magnetite as the principal Fe-
373 hosting mineral phase. Remarkably, serpentinization at Hole 1268A is characterized
374 by the development of strongly U-shaped REE patterns with pronounced positive Eu-
375 anomalies (**Fig. 5f**). The MREE and LREE show wide ranges of concentrations
376 (variations of up to two orders of magnitude) whereas the HREE are comparatively
377 constant (range of Lu_N: 0.11 to 0.27). The most extreme examples have positive
378 LREE/HREE ratios and show distinct positive Eu anomalies.

379 Talc alteration at Hole 1268A supersedes serpentinization and is
380 characterized by a decrease in H₂O (ranging mainly between 5 to 6 wt.% which
381 corresponds to the water content of talc, **Fig. 6a**). Furthermore, talc-altered rocks are
382 SiO₂-rich (60 to 65 wt%) with comparatively low MgO values (**Fig. 6b**) and
383 characterized by higher FeO concentrations (4 to 5 wt%) but considerably lower
384 Fe₂O₃ concentrations (1 to 3 wt%) than serpentinization. Also, the data plot below the
385 line defining a constant Fe-budget (**Fig. 7d**) which may indicate that some Fe was
386 lost or that the total Fe concentration was diluted due to a gain in SiO₂. This issue is

387 further examined below (Discussion) using a mass balance approach. In contrast to
388 serpentinization in Hole 1268A, the talc alteration is characterized by comparatively
389 smooth, flat to LREE-enriched patterns, and most samples show a well-developed
390 negative Eu-anomaly (**Fig. 5g**).

391

392 **[INSERT Figure 8]**

393 **Discussion**

394 Recent studies have suggested that asthenospheric melting and refertilization
395 by melt-rock interaction are the dominant processes controlling the major and trace
396 element characteristics of abyssal peridotites (Niu et al., 1997; Niu 2004). Our data
397 show that this may only be true for fluid-peridotite interaction under certain conditions
398 here referred to as rock-dominated serpentinization.

399 **Figure 8** summarizes the various processes affecting ascending refractory
400 peridotite as evidenced in different locales sampled during ODP Leg 209. Melts rising
401 through the lithosphere may interact with the peridotite and cause particular
402 geochemical modifications (e.g., addition of SiO₂, REE and HFSE). In some
403 instances, these melts may be channeled along ductile deformation zones that
404 develop in response to tectonic strain. Evidence for such processes is preserved at
405 Sites 1270 and 1271 where addition of mafic and differentiated melts imparted local
406 enrichments in some trace elements (e.g., U, Th, Zr) and REE.

407 An initial stage of high-temperature (>375-400°C) hydrothermal interaction of
408 peridotite with hydrothermal solutions is documented by relict textures of
409 orthopyroxene replacement by talc and tremolite at Sites 1270 and 1271 (Allen and
410 Seyfried, 2003; Bach et al., 2004). This stage is overprinted by widespread
411 serpentinization that may be largely isochemical (i.e., restricted to the addition of H₂O,

412 rock-dominated serpentinization) or associated with particular geochemical
413 modifications due to more intense interaction with hydrothermal fluids (fluid-
414 dominated serpentinization). Continued fluid-serpentinite interaction caused iowaite
415 formation at Site 1272 (Bach et al., 2004) and serpentine replacement by talc at Site
416 1268. The last stage of alteration is low-T interaction of the serpentinites with
417 ambient seawater that circulates in the near seafloor environment, generating
418 aragonite veinlets, Fe-oxyhydroxides, and clays.

419 The following discussion focuses on the geochemical processes in the
420 “hydrothermal regime”, excluding the high-T (replacement of orthopyroxene) and low-
421 T (seafloor alteration) end-member conditions (**Fig. 8**). However, in order to assess
422 element mobility under these hydrothermal conditions the influence of melt-rock
423 interaction processes on the composition of Leg 209 peridotites needs to be
424 considered first.

425

426 ***[INSERT Figure 9]***

427 ***REE and HFSE systematics related to melt-rock interaction (Sites 1270 and***
428 ***1271) and hydrothermal alteration (Sites 1274, 1272, and 1268)***

429 The peridotites drilled during ODP Leg 209 show remarkable compositional
430 variations and REE have been added during melt-rock interaction (**Fig. 5 c, d**) and
431 hydrothermal alteration (**Fig. 5 f, g**). In order to investigate the relative importance of
432 these processes at the different sites the behavior of HFSE and REE may be
433 considered. In aqueous solutions, the LREE are more readily transported than HREE
434 and HFSE whereas melt-rock interaction causes addition of LREE and HFSE to the
435 rock in about equal proportions (Niu, 2004). In **Figure 9** the global abyssal peridotite
436 data presented by Niu (2004) define positive trends in Nb vs. La and Th vs. Ce space

437 which indicates that melt-rock interaction is the dominant factor controlling the
438 compositional variation. However, the situation is considerably more complex in the
439 15°20'N area where two deviating trends can be observed. One trend is defined by
440 the HFSE enriched samples that are mainly derived from Holes 1270B, C, and D and
441 Hole 1271B. In line with the petrographic evidence for melt impregnation, these
442 samples follow the trend of the global data set characterizing melt-rock interaction as
443 the dominant process. A different trend is defined by data from Holes 1274A, 1272A
444 and 1268A which show LREE enrichment correlated with only a minor increase in
445 HFSE concentrations (Fig. 9a, b). This relationship indicates that LREE variability is
446 largely due to hydrothermal alteration, which affects HFSE, like Th, to lesser extents.
447 The most prominent example of hydrothermal alteration affecting REE contents (i.e.,
448 fluid-dominated alteration) are serpentinites and talc-altered rocks from Hole 1268A.

449 In addition to the increase in LREE concentrations during magmatic and
450 hydrothermal processes, increasing ratios of MREE/HREE (Gd_N/Lu_N) are apparent
451 when samples from Sites 1268, 1270, and 1271 are compared to samples from Sites
452 1272 and 1274 (Fig. 5). These changes are correlated with high HFSE
453 concentrations in many samples at Sites 1270 and 1271, whereas an array of
454 increasing Gd_N/Lu_N at low HFSE concentrations is defined by data from Sites 1268,
455 1272, and 1274 (Fig. 9c, d). This is consistent with the interpretation that MREE
456 have also been added locally by hydrothermal alteration. However, the relatively wide
457 range of Gd_N/Lu_N ratios at Hole 1274A (rock-dominated serpentinization) suggest
458 that primary variability in the peridotite protolith may also have played a role.

459 It is important to note that some samples from Sites 1270 and 1271 do not
460 follow the “melt-rock interaction trend” but show relatively low HFSE concentrations
461 similar to the samples defining the “hydrothermal alteration trend”. This may indicate

462 that melt-rock interaction, and hence melt transport, was heterogeneous at the scale
463 sampled by these drill holes.

464

465 ***[INSERT Figure 10]***

466 ***Rock-dominated serpentinization: Mass changes at Site 1274***

467 The peridotite at Site 1274 shows successive replacement of olivine and
468 pyroxene by serpentine ± brucite ± magnetite assemblages and serpentinization is
469 complete below 60 mbsf (**Fig. 2a**). Nevertheless, it can be demonstrated that
470 variations in the proportions of olivine and pyroxene in the protolith are responsible
471 for the variations in Al₂O₃, SiO₂, MgO and FeO_{total} (**Fig. 5**). Hence, the relative
472 proportions of these elements were preserved despite serpentinization indicating that
473 they were immobile during fluid/rock interaction (cf. Niu, 2004). This interpretation is
474 supported by microprobe data revealing distinctive compositions for serpentine
475 formed after orthopyroxene and serpentine formed after olivine (Moll et al.,
476 submitted).

477 If serpentinization was isochemical and simply related to the addition of H₂O
478 then an overall mass gain should cause a decrease in the measured concentrations
479 of immobile elements in the serpentinized rocks compared to their unaltered
480 precursor (Barrett and MacLean, 1994; Grant, 1986; MacLean, 1990; MacLean and
481 Barrett, 1993). However, the samples from Hole 1274A are heterogeneous in their
482 modal orthopyroxene to olivine ratios so that the precursor composition is distinctive
483 for each sample. Hence, a precursor composition must be determined for each
484 sample and compared to the measured composition of the altered rock in order to
485 determine the elemental fluxes during alteration. It is possible to calculate these
486 precursor compositions by combining the modal proportions of primary phases

487 (determined from point counting) with the compositional data of these phases
488 (microprobe analyses). In the case of Hole 1274A, trace amounts of fresh primary
489 phases are preserved even in the most altered samples. The results of this
490 procedure are presented in **Table 5 (Appendix)**.

491 As should be expected, the calculated composition of the precursor peridotites
492 show well-constrained correlations for MgO, SiO₂, Al₂O₃, and FeO_{tot} concentrations
493 with the modal proportions of pyroxene and olivine (**Fig. 10 a - d**). The dunite and
494 pyroxene-poor harzburgite (pyroxene_{modal}/[pyroxene_{modal} + olivine_{modal}] < 0.2) have
495 relatively low Al₂O₃ and SiO₂ concentrations but high MgO and FeO_{tot} values
496 compared to the harzburgite with high modal pyroxene contents
497 (pyroxene_{modal}/[pyroxene_{modal} + olivine_{modal}] = 0.3 to 0.4). These correlations indicate
498 that the data may be used to establish the mass changes during serpentinization at
499 Hole 1274A.

500 In general, the budget of mass addition or loss during alteration can be
501 quantified based on a comparison of the concentration of immobile elements in the
502 precursor with that in the altered equivalent, since any differences must be caused by
503 overall mass change. This relationship can be expressed as

$$504 \quad c_{\text{immobile_element}}^{\text{precursor_rock}} * EF = c_{\text{immobile_element}}^{\text{altered_rock}}$$

505 where c is the concentration and the enrichment factor (EF) is >1 if the rock lost
506 mass during alteration and <1 if mass has been added. Assuming that only H₂O was
507 added during serpentinization, SiO₂, MgO, FeO_{tot} and Al₂O₃ may be used to calculate
508 EF values for the variably altered peridotite (**Table 5**). The results show that EF_{SiO₂},
509 EF_{MgO} and EF_{FeO_{tot}} are within the range of 0.95 to 0.80 and values for each sample
510 are narrowly constrained (**Fig. 10 e**). Overall, there is a decrease in EF values with
511 increasing H₂O content which supports the interpretation that H₂O addition to the
512 rock during alteration was associated with successive mass gains. In contrast, the EF

513 values calculated on the basis of Al₂O₃ concentrations show a large scatter and are
514 generally lower than expected (**Fig. 10e**). This would suggest that Al₂O₃ was lost
515 from the peridotite during serpentinization, which is unlikely given the low levels of Al
516 in hydrothermal fluids venting from peridotite massifs (<3 micromolar; Douville et al.,
517 2002). Alternatively, it may be inferred that the Al₂O₃ concentrations in the precursor
518 have been overestimated. Since Al₂O₃ is concentrated in spinel, which is only a
519 minor phase in the samples, an overestimation of its modal abundance during the
520 point counting procedure has severe effects on the calculated concentrations in the
521 precursor. Also, Al₂O₃ concentrations are mainly below 1 wt% in the altered rocks,
522 hence, there are profound effects on the EF value if the calculated concentration of
523 Al₂O₃ precursor is overestimated by just 0.1 to 0.5 wt%. Consequently, we consider the
524 EF_{Al₂O₃} values less reliable than the EF_{SiO₂}, EF_{MgO} and EF_{FeO_{tot}} values.

525 In order to determine the total mass change during alteration, the EF values
526 calculated for each sample based on the SiO₂, MgO, and FeO_{total} concentrations
527 have been averaged (**Table 5; Fig. 10e**). The relationship between mass change,
528 measured and reconstructed concentrations and the calculated enrichment factor
529 can be expressed as

$$530 \quad \Delta X_{total} = (C_{total}^{altered_rock} / EF_{average}) - C_{total}^{precursor_rock}$$

531 where ΔX represent a mass change in [g/100g]. The results show that mass addition
532 was in the range of 10 to 20 g/100g (**Table 3, Fig. 10 f**) and a positive correlation of
533 these mass changes with the volume of hydrothermal veins determined by point
534 counting suggests that the additional mass was accommodated by volume expansion
535 and hydrothermal veining.

536 These mass balance calculations support the interpretation that geochemical
537 modifications of the peridotite in Hole 1274A were minimal and largely restricted to
538 the addition of H₂O. However, whereas the content of total iron has not been affected

539 by alteration, there is a considerable shift in oxidation state. The concentration of
540 Fe_2O_3 is increasing from 4 to 6.5 wt% with increasing H_2O values (**Fig. 11a**) which is
541 consistent with the presence of magnetite in the more strongly altered samples.

542

543 **[INSERT Figure 11]**

544 ***Rock-dominated alteration: REE systematics at Sites 1274 and 1272***

545 The least-altered peridotite and the completely serpentinized peridotite at Hole
546 1274A show generally similar LREE-depleted patterns which is consistent with the
547 interpretation of quasi-isochemical serpentinization. However, there are variations in
548 the HREE contents and the LREE/HREE ratios, and some samples show positive Eu
549 anomalies. At Hole 1272A, the serpentinites show narrowly constrained Lu_N values
550 below the concentrations in the least-altered peridotite. The LREE and MREE are
551 somewhat variable and positive Eu-anomalies are common.

552 Potentially, all these variations may be attributable to slight differences in the
553 primary mineral assemblages. The principal carrier of the REE signature in these
554 peridotites is clinopyroxene, which may vary in abundance and composition, and the
555 development of a positive Eu anomaly could be due to the presence of plagioclase,
556 potentially introduced by melt-rock interaction. However, there is evidence indicating
557 that hydrothermal processes are also important, in particular, for the development of
558 Eu anomalies.

559 With regard to the HREE concentrations it may be inferred that primary
560 composition and hydrothermal alteration both contribute to the well-constrained trend
561 of decreasing Lu_N concentrations with increasing degree of serpentinization reflected
562 in H_2O concentrations (**Fig. 11 b**). In dunite the high modal proportion of olivine
563 favors the formation of serpentine-brucite assemblages causing high H_2O

564 concentrations during alteration and the scarcity of ortho- and clinopyroxene in the
565 protolith could be responsible for low Lu concentrations in the precursor. Hence, the
566 low Lu_N concentrations (**Fig. 11 b**) are probably due to the primary modal
567 characteristics and dilution related to mass gain during serpentinization. However,
568 there is no correlation between La_N and H_2O (**Fig. 11c**), which could be explained by
569 invoking La mobility during serpentinization if a constant La_N/Lu_N ratio in the
570 unaltered precursors of the Hole 1274A samples is assumed. Elevated La_N/Lu_N
571 values are particularly common in the lower portion of the drill hole where intense
572 faulting has been observed (**Fig. 11d**) which may have facilitated serpentinization in
573 this area by providing fluid pathways. Nevertheless, similar relative enrichments
574 observed in highly refractory peridotites have been previously attributed to late
575 melt/rock interaction associated with melt extraction and transport (e.g., Godard et al.,
576 1995; Navon and Stolper, 1987). In fact, minor differences in the abundance and
577 composition of clinopyroxene may readily explain the variations of Lu_N/La_N ratios at
578 Hole 1274A.

579 Several samples from Holes 1274A and 1272A show a positive Eu-anomaly
580 (**Fig. 5a and b**) which could be an indication for the former presence of plagioclase.
581 There are three samples from Hole 1274A showing elevated Sr concentrations (4 to
582 8 ppm, **Fig. 12a**) which could be supporting evidence for such a conclusion. However,
583 all the samples from Hole 1272A have Sr concentrations <1.5 ppm **and serpentinites**
584 **from Hole 1268A, where serpentinization is characterized by strongly positive Eu-**
585 **anomalies, are also Sr poor.** Also, no relicts of plagioclase were observed in thin
586 sections. In addition, there is no correlation between Eu/Eu^* and Sr and therefore
587 relict plagioclase is probably not important in controlling the Eu-systematics for most
588 samples. Consequently, it seems more likely that interaction with hydrothermal fluids
589 during serpentinization resulted in local addition of Eu. This suggests that the

590 serpentinizing fluid leached Eu during fluid-rock interaction prior to reacting with the
591 sampled peridotites. Eu mobility is much increased compared to that of the trivalent
592 REE under highly reducing conditions and high Cl-contents (Allen and Seyfried,
593 2005). The positive Eu-anomaly may therefore have been imposed during
594 serpentinization in Holes 1274A and 1272A, even though alteration was largely rock-
595 dominated.

596

597 **[INSERT Figure 12 and 13]**

598 ***Fluid-dominated serpentinization: Site 1268***

599 The serpentinites showing S addition (Hole 1268A; **Fig. 7c**) and U-shaped
600 REE patterns with strong positive Eu anomalies (Hole 1268A and some serpentinites
601 of Sites 1270 and 1271; **Fig. 5**) are interpreted to have experienced fluid-dominated
602 serpentinization. It is unlikely that the positive Eu-anomaly can be attributed to relict
603 plagioclase since Sr concentrations are below 2 ppm for most samples (**Fig. 12a**).

604 The REE patterns of the Hole 1268A serpentinites deviate strongly from the
605 least-altered peridotites but show similarities to the REE characteristics of
606 hydrothermal fluids discharging at black smokers of ultramafic-hosted systems at the
607 seafloor, such as the Rainbow and Logatchev hydrothermal sites (**Fig. 13**). End-
608 member compositions of these hot (350 to 400°C) metal and S-rich fluids are
609 characterized by positive La_N/Lu_N ratios and strong positive Eu anomalies (Douville et
610 al., 2002; 1997). It has been suggested that these REE characteristics are due to
611 equilibration with plagioclase-bearing lithologies during circulation within the oceanic
612 crust (Douville et al., 2002; Klinkhammer et al., 1994). However, recent experimental
613 studies demonstrate that the chlorinity and redox potential of the fluid is a major
614 controlling factor on LREE and Eu complexation and transportation (Allen and

615 Seyfried, 2005) and that the presence of plagioclase is not required for the
616 generation of LREE-enriched fluid compositions with positive Eu anomalies.

617 Apparently, interaction of refractory peridotite with a hot, black-smoker type
618 hydrothermal fluid will strongly influence the REE systematics of the resulting
619 serpentinite, given sufficient fluid/rock ratios. Clearly, the LREE and Eu are likely to
620 be affected most strongly since HREE concentrations in the fluids are 3 to 4 orders of
621 magnitude lower than in the peridotite (**Fig. 13**). This may explain why the HREE of
622 serpentinites at Hole 1268A are fairly constant whereas LREE and MREE are highly
623 variable (**Fig. 5f**). Interaction with black-smoker type hydrothermal fluids may
624 generate successively more modified REE pattern in the serpentinites and the
625 samples with highest La_N/Lu_N and Gd_N/Lu_N ratios probably represent the most mature
626 stage of this process.

627 These considerations imply that LREE and MREE were incorporated in the
628 rock during serpentinization. However, it is difficult envisage that the LREE and
629 MREE were preferentially incorporated in serpentine minerals since the effective
630 atomic radius of HREE are comparable to Mg (0.89) whereas LREE are similar to Ca
631 (1.12; octahedral coordination; Shannon, 1976). Either the fluid/rock ratio was so
632 high that the REE content of serpentine was entirely controlled by the fluid or there
633 may be trace amounts of hydrothermal REE-bearing phases imparting the particular
634 characteristics onto the rocks. This issue could be addressed by in-situ analytical
635 techniques (LA-ICP-MS) investigations regarding the microscale variations in the
636 REE systematics of the serpentinites. In any case, high fluid fluxes are required to
637 add sufficient amounts of REEs to the rock in order to erase the LREE depleted
638 character of the protolith. **Hence, based on the available data, it can be concluded**
639 **that the time-averaged signal of serpentinization involving vent-type hydrothermal**

640 fluids imparts a particular; LREE and Eu enriched REE pattern on the rock and this
641 style of alteration is referred to as fluid-dominated serpentinization.

642

643 **[INSERT Figure 14]**

644 **Talc alteration: Site 1268**

645 Talc alteration at Site1268 overprinted serpentinization and clearly modified
646 the bulk rock compositions in terms of the Mg/Si ratios, H₂O contents, and REE
647 systematics. Quantification of the geochemical budget of this process is difficult since
648 both the composition of the serpentinites and the talc alteration in Hole 1268A show
649 significant variability that can be attributed to differences in the modal composition of
650 the protoliths. This is particularly evident for Al₂O₃ concentrations which range from
651 0.1 to 1.2 wt% for both serpentinization and talc alteration suggesting substantial
652 primary variability of the precursor peridotite in terms of the concentration of spinel
653 and the modal proportions of pyroxene and olivine (**Fig. 4**). However, in order to gain
654 a first order approximation of the general trends in the mass budget of the major
655 components (SiO₂, Al₂O₃, FeO, Fe₂O₃, FeS₂, MgO, H₂O) the average composition of
656 the serpentinites may be compared to the average composition of the talc alteration.
657 For this exercise three different scenarios have been considered: 1) talc alteration of
658 serpentinization took place without mass change, 2) MgO was immobile during
659 alteration, and 3) SiO₂ was immobile (**Table 6 [Appendix]; Fig. 14**).

660 Assuming that no mass change took place during talc alteration (EF = 1) the
661 ΔX can be calculated by simple subtraction of the concentration of the component in
662 the altered rock (talc alteration) from the concentration in the precursor
663 (serpentinization). In this model, the rock would have gained SiO₂ (18 g/100g) and
664 lost H₂O, MgO and total iron (mainly in the form of Fe₂O₃ and FeS₂). An alternative

665 scenario is provided by the assumption that MgO remained immobile during talc
666 alteration (EF = 1.315). Due to the higher Si/Mg ratio of talc compared to serpentine
667 this calculation results in a large mass gain of SiO₂ (37 g/100g). This is partially offset
668 by loss of H₂O (-5 g/100g), however, the calculated total mass gain is substantial (31
669 g/100g). Interestingly, the iron budget is almost neutral and gains in FeO are offset
670 by the loss in Fe₂O₃. Hence, iron may have remained largely immobile during talc
671 alteration and the calculated mass changes for the different iron species reflect
672 adjustments to the new physico-chemical conditions. This model is supported by
673 considerations regarding the fluid chemistry of talc alteration which indicate that Si-
674 rich and Mg saturated fluids are responsible for the replacement of the serpentine
675 (Bach et al., 2004). However, the high total mass gain calculated for this process
676 invokes substantial volume expansion. Assuming a simple reaction of
677 $1 \text{ Mg}_3\text{Si}_2\text{O}_5(\text{OH})_4 \text{ (serpentine)} + 2 \text{ SiO}_{2,\text{aq}} = 1 \text{ Mg}_3\text{Si}_4\text{O}_{10}(\text{OH})_2 \text{ (talc)} + \text{H}_2\text{O}$
678 a volume increase of 27% can be calculated based on the molar volumes of talc (140
679 cm³/mol) and chrysotile (110 cm³/mol; Robie et al., 1979). In this regard, it is
680 important to consider the relatively high abundance of hydrothermal veins estimated
681 from core logging data at Hole 1268A (**Table 1**, Shipboard Scientific Party, 2004;
682 Bach et al., 2004). Here, 8.9 % of the drill core consist of macroscopic hydrothermal
683 veins and about half of these are filled by talc. Potentially, the mass addition
684 associated with Si metasomatism during talc alteration of serpentinite is
685 accommodated by volume expansion and associated veining.

686 An additional calculation is based on the assumption that SiO₂ was immobile
687 during the replacement of serpentinite by talc (EF = 0.698; **Fig 14**). In this case, high
688 losses in all other major components are required in order to modify the elemental
689 proportions appropriately resulting in a total mass loss of -30g/100g. In particular,
690 substantial MgO would need to be removed from the rock (-18g/100g) which appears

691 to be an unlikely scenario since the solubility of hydrous Mg-silicates in hydrothermal
692 fluids is extremely low (e.g., Saccocia et al., 1994).

693 Overall, it may be concluded that talc alteration of serpentinites at Hole 1268A
694 was associated with large total mass gains due to SiO_2 addition from a high a_{SiO_2} fluid
695 which may have been accommodated by volume expansion associated with intense
696 hydrothermal veining. The high a_{SiO_2} fluids may have been derived from pyroxene
697 destructive alteration of peridotite or gabbro at depth (Bach et al., 2004). **Similarly,**
698 **interactions with gabbro may well be the source of sulfides in rocks from Hole 1268A,**
699 **as suggested by Alt and Shanks (2003) for sulfides in serpentinites from the 23°N**
700 **area on the Mid-Atlantic Ridge. Alternatively, fluids with high sulfur fugacities could**
701 **be produced during desulfurization of primary sulfides at a serpentinization front deep**
702 **in the system.**

703 The REE signature of talc-alteration deviates substantially from the
704 characteristics of serpentinization at Hole 1268A (**Fig. 4**). Serpentinization is
705 characterized by U-shaped patterns with strong positive Eu anomalies, whereas the
706 talc alteration results in negative Eu anomalies and gently sloping or flat patterns.
707 Also, the overall concentrations of REE increased despite mass addition to the
708 serpentinites, which would have the tendency to dilute the concentration of any
709 element considered to be immobile, such as the HREE. Apparently, LREE, MREE,
710 and HREE have been added during talc alteration in particular proportions generating
711 comparatively smooth patterns. However, Eu does not adhere to this trend and
712 Eu/Eu^* values are generally <1 .

713 It is difficult to conclude how the particular REE signature of talc alteration was
714 generated. Maybe Eu resided in a different phase than the other REE in the
715 serpentinite protolith. Upon talc-alteration, the Eu-bearing phase may have been
716 destroyed while the other REE-bearing phase is stable. Alternatively, it may be

717 possible that REE, including Eu, are all substituted into the serpentine structure and
718 that recrystallization released Eu because it does not fit into the crystal lattice of talc.
719 However, talc rocks with prominent positive Eu anomalies are known from submarine
720 hydrothermal breccias and hydrothermally altered gabbroic rocks (D’Orazio et al.,
721 2004).

722 Boschi et al. (2006) observed considerable REE and HFSE variations and
723 variable Eu anomalies in schistous talc- and amphibole-rich metasomatized
724 ultramafic rocks from Atlantis Massif and concluded that these might be due to fluid-
725 rock interactions, although melt impregnation processes cannot be ruled out. While
726 hydrothermal mobilization of REE and HFSE seems plausible in schistous rocks from
727 detachment faults that focused fluid flow, such mass transfers are more difficult to
728 envision in situations of static alteration such as recorded in Hole 1268A. It is
729 possible, however, that the talc-alteration in Hole 1268A took place immediately
730 adjacent to shear zones as suggested by Boschi et al. (2006) for static
731 metasomatisms at Atlantis Massif. In that instance, the variability in REE contents in
732 rocks from Hole 1268A may reflect different degrees of infiltration by synkinematic
733 metasomatic fluids that migrated along a detachment fault.

734

735 **Conclusions**

736 The mantle section drilled in the 15°20’ N area shows that there are
737 considerable compositional heterogeneities within a MAR segment of less than 100
738 km in length due to magmatic and hydrothermal processes. Overall, the peridotites
739 are depleted in several incompatible elements (e.g., Al, Sc, V) indicating a refractory
740 starting composition.

741 At Sites 1270 and 1271, melt-rock interaction processes are prevalent
742 generating smooth, LREE enriched pattern similar to the pattern of gabbroic rocks.
743 Here, the elevated LREE concentrations and LREE/HREE ratios are correlated with
744 increasing HFSE contents indicating that both were introduced by a melt phase.

745 In contrast, samples of altered peridotite from Sites 1274, 1272, and 1268
746 define different trends where increasing LREE concentrations correlate with only
747 slightly increasing HFSE contents. Since the solubility of LREE in aqueous solutions
748 is higher than the solubility of HFSE it is inferred that this trend is due to
749 hydrothermal alteration processes.

750 Serpentinization took place under variable conditions which can be described
751 as “rock-dominated” and “fluid-dominated”. Fluid-dominated serpentinization at Hole
752 1268A is characterized by addition of S, dominantly in the form of pyrite, and the
753 formation of U-shaped REE pattern with strong positive Eu-anomalies. These REE
754 pattern are similar to the REE characteristics of hot (350 to 400°C) hydrothermal
755 fluids discharging at black smoker sites of ultramafic hosted hydrothermal systems
756 such as Rainbow and Logatchev. Hence, it is inferred that serpentinization involving
757 vent-type fluids and high fluid/rock ratios imposed the REE signature of the fluid onto
758 the serpentinites.

759 Rock-dominated serpentinization at Sites 1274 and 1272 is characterized by
760 little geochemical deviation from the precursor compositions except for the addition of
761 H₂O causing an overall mass gain of up to 20g/100g. The REE patterns are
762 essentially similar to the least-altered peridotite sampled at the upper portion of Hole
763 1274A. Some of the variability in LREE concentrations could be related to
764 hydrothermal alteration and/or localized melt/rock interaction. The presence of some
765 samples with positive Eu-anomalies is most likely caused by addition of Eu during
766 fluid/rock interaction.

767 Talc alteration under static conditions of serpentinites at Hole 1268A was due
768 to interaction with high a_{SiO_2} fluids and associated with substantial mass addition (in
769 the order of 30 g/100 g). Furthermore, REE are apparently added to the rock
770 generating smooth REE patterns, however, strong negative Eu-anomalies are
771 prevalent.

772 Overall, it can be concluded that hydrothermal processes are capable of
773 locally generating a wide spectrum of REE patterns controlled by variations in
774 alteration conditions. The evolution of the REE characteristics can be envisioned as a
775 continuum from LREE depleted (Holes 1274A and 1272A) to hydrothermal vent fluid
776 type patterns (serpentinites at Hole 1268A) on to LREE and MREE enriched patterns
777 (talc alteration at Hole 1268A). These findings have important bearings on the
778 interpretation of REE data from completely serpentinitized abyssal peridotite. Clearly,
779 variations in LREE alone are unreliable as indicators for the nature of the mantle
780 protolith or the melt-rock interaction processes. Also, the budget of geochemical
781 exchange processes between the hydrosphere and lithosphere at ultraslow
782 spreading ocean ridges need to take into account that serpentinites may represent
783 an important sink for Eu and LREE under fluid-dominated alteration conditions.
784 However, it should be emphasized that the precise mechanism of the incorporation of
785 REE in serpentinites is poorly understood. It needs to be established whether the
786 REE are hosted within the serpentine minerals or whether there are accessory
787 phases present that control the REE systematics. Such research would require the
788 application of in-situ analyses such as LA-ICP-MS and may also elucidate the
789 controls on the development of the particular REE pattern associated with talc
790 alteration.

791

792

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802 **Appendix**

803 ***[Insert Tables 3, 4, and 5]***

804 **References**

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997

998 **Figure captions:**

999

1000 Fig. 1: ODP Leg 209 recovered variably altered peridotite at five sites in the vicinity of
1001 the 15°20'N Fracture Zone. A: Location of the 15°20'N Fracture Zone in the
1002 Atlantic Ocean. B: Location of the southern and northern working area of ODP
1003 Leg 209. C and D: Location of ODP sites 1268, 1270, 1271, 1272, and 1274
1004 (stars). The location of the Logatchev active hydrothermal site (triangle) is also
1005 shown. Bathymetry from Lagabrielle et al. (1998).

1006

1007 Fig. 2: Graphic logs for ODP drill holes from Sites 1274, 1272, 1271, 1270 and 1268.
1008 The columns show the recovery rate, lithological characteristics, and the
1009 intensity of alteration and veining. The distribution of primary and alteration
1010 minerals is highly variable.

1011 (a) Hole 1274A contains the least-altered peridotite with up to 50 vol% olivine and
1012 pyroxene in the upper section, however, serpentinization is complete below 100
1013 mbsf.

1014 (b) The topmost ~50 m of Hole 1272A are interpreted as a tectonic mega-breccia
1015 dominated by basalt and diabase (Shipboard Scientific Party, 2004) whereas
1016 the lower section consists of completely serpentinized harzburgite with
1017 abundant iowaite ($\text{Mg}_4[\text{OH}]_8\text{Fe}^{3+}\text{OCl} \times 1-4 \text{H}_2\text{O}$).

1018 (c) Holes 1271A and 1271B are 75 m apart and recovered a complex mixture of
1019 serpentine-brucite-magnetite altered, chromite-bearing dunite and minor
1020 harzburgite. Furthermore, there are amphibole gabbro, troctolite, and gabbroic
1021 intrusions. Ultramafic samples from these holes show textural evidence for melt-
1022 rock interaction on hand specimen and thin section scale.

1023 (d, e, and f) Four holes were drilled at Site 1270: Holes 1270C and 1270D are
1024 located immediately adjacent to each other whereas Holes 1270B and 1270A
1025 are located at ~300 m and ~500 m down slope to the west (Shipboard Scientific
1026 Party, 2004). Hole 1270A consists mainly of serpentized harzburgite with
1027 occasional gabbroic intrusion whereas Hole 1270B consists of gabbro and
1028 gabbronorite with localized occurrences of completely talc altered harzburgite.
1029 Serpentized peridotite in Holes 1270C and D are heavily intruded by gabbroic
1030 dikelets.

1031 (g) Hole 1268A recovered completely serpentized peridotite associated with
1032 pyrite-bearing veinlets (up to 3 vol%) that have been overprinted by pervasive
1033 talc alteration. Magnetite is generally absent in talc altered serpentinite.
1034 Abbreviations: Altn: Alteration; Amp: Amphibole; Bru: Brucite; Carb: Carbonate;
1035 Chl: Chlorite; Cr-sp: Cr-spinel; Fe-oxi: Fe-oxyhydroxide; Mt: Magnetite; Opx:
1036 Orthopyroxene; Py: Pyrite; Serp: Serpentine (lizardite and/or chrysotile). Depth
1037 in meter below seafloor (mbsf); recovery in % per core barrel.

1038

1039 Fig. 3: In MgO/SiO_2 vs. $\text{Al}_2\text{O}_3/\text{SiO}_2$ space the data of the abyssal peridotite from the
1040 $15^\circ 20' \text{N}$ area show a considerable variability whereas analyses of a global set
1041 of abyssal peridotite presented by Niu (2004) define a trend parallel to the
1042 terrestrial array (Jagoutz et al., 1979). The geochemistry of the $15^\circ 20' \text{N}$
1043 samples is controlled by a variety of processes including modal mineralogical
1044 composition and hydrothermal alteration.

1045

1046 Fig. 4: Major and trace element concentration in variably altered peridotite recovered
1047 from ODP Sites 1268, 1270, 1271, 1272, and 1274. Major element
1048 concentrations are recalculated to 100% anhydrous. **Symbols as in Fig. 3.**

- 1049 (a) Al_2O_3 vs. V
 1050 (b) Al_2O_3 vs. Sc
 1051 (c) Al_2O_3 vs. Cr
 1052 (d) Al_2O_3 vs. $\text{FeO}_{\text{total}}$
 1053 (e) Al_2O_3 vs. modal proportion of pyroxene at Site 1274A (point counting data:
 1054 modal pyroxene is preserved pyroxene + pyroxene pseudomorphed by
 1055 serpentine; modal olivine is preserved olivine + olivine replaced by serpentine)
 1056 (f) Al_2O_3 vs. MgO
 1057 (g) Al_2O_3 vs. SiO_2

1058

1059 Fig. 5: The REE-Y patterns of variably altered peridotite recovered from ODP sites
 1060 1268, 1270, 1271, 1272, and 1274 are highly diverse reflecting primary
 1061 characteristics, melt-rock interaction and effects of hydrothermal alteration
 1062 (serpentinization and talc alteration).

1063 (a) The least-altered peridotite (heavy gray lines) from the upper 50 mbsf at Hole
 1064 1274A are strongly depleted in LREE, in keeping with the refractory character of
 1065 common mid-ocean ridge mantle. Serpentinized harzburgite (white squares)
 1066 and dunite (black squares) have similar LREE depleted pattern than least-
 1067 altered peridotite. However, there is some variability in HREE and some
 1068 samples appear to have gained LREE and Eu. A zone consisting of soft
 1069 serpentine mud at 132 mbsf is interpreted as a fault zone (crosses) and shows
 1070 a flat REE pattern at chondritic concentrations possibly due to localized melt-
 1071 rock interaction processes.

1072 (b) The iowaite-bearing serpentinites from Hole 1272A are characterized by
 1073 variable MREE and LREE concentrations and the development of positive Eu-

1074 anomalies. The HREE contents are constant and lower compared to the least-
1075 altered peridotite from Hole 1274A (gray lines).

1076 (c) The smooth REE pattern of serpentinites from Holes 1270C and D (crosses)
1077 are flat to LREE enriched and show a remarkable similarity to mafic rocks
1078 sampled during ODP Leg 209. One sample of talc altered harzburgite from Hole
1079 1270B shows the same characteristics (black cross on grey background).
1080 These features are consistent with petrographic characteristics indicating
1081 significant melt-rock interaction. In contrast, two samples from Hole 1270A
1082 (white cross on black background) show pronounced U-shaped pattern with
1083 strong positive Eu-anomalies. These features are similar to the “fluid-dominated
1084 serpentinization” at Hole 1268A (**Fig. 5f**).

1085 (d) Similar to the samples from Holes 1270B, C, and D the REE data for most of
1086 the samples from Hole 1271B indicate interaction with a melt phase causing
1087 REE addition and flat to LREE enriched REE shapes. However, two samples
1088 from the upper 40 mbsf show U-shaped pattern with a strong positive Eu-
1089 anomaly similar to the “fluid-dominated serpentinization” at Hole 1268A (**Fig. 5f**).

1090 (e) The Hole 1271A was drilled immediately adjacent to Hole 1271B but reached
1091 only 40 mbsf. In this domain, the effects of melt-rock interaction processes
1092 appear to be less pronounced than further below in Hole 1271B. The REE
1093 pattern of Hole 1271A are strongly enriched in LREE compared to the least
1094 altered peridotite sampled in Hole 1274A (**Fig. 5a**), however, most samples
1095 show U-shaped pattern and positive Eu anomalies. It is possible that melt-rock
1096 interaction and subsequent hydrothermal processes are both responsible for
1097 these characteristics.

1098 (f) Serpentinization at Hole 1268A generated peculiar U-shaped REE pattern with
1099 strong LREE enrichments and pronounced positive Eu-anomalies. Whereas

1100 LREE and MREE show variations of about two orders of magnitude the HREE
1101 concentrations are remarkably constant. Heavy grey lines are least-altered
1102 peridotite from Hole 1274A.

1103 (g) Talc alteration at Hole 1268A that overprinted serpentinization, shows smooth,
1104 flat to LREE enriched pattern and most samples have a negative Eu-anomaly.
1105 Overall REE concentrations are elevated compared to serpentinization. Heavy
1106 grey lines are least-altered peridotite from Hole 1274A.

1107

1108 Fig. 6: In terms of major elements the effects of alteration are particularly apparent in
1109 variations of MgO, FeO_{total}, SiO₂, H₂O and S concentrations and the oxidation
1110 state of iron. Symbols as in Fig. 3.

1111 (a) MgO vs. H₂O

1112 (b) MgO vs. SiO₂

1113 (c) FeO_{total} vs. S

1114 (d) Fe₂O₃ calculated vs. FeO_{measured}

1115 Major element concentrations are recalculated to 100% anhydrous. The H₂O
1116 concentrations were measured directly on board the RV Joides-Resolution
1117 during ODP Leg 209 for 85 samples. For the remaining samples the H₂O
1118 contents have been calculated using the loss on ignition and the data for CO₂, S,
1119 and NO₃ (Table 5, electronic supplement).

1120

1121 Fig. 7: Some serpentinites from Sites 1270 and 1271 show relatively high immobile
1122 trace element concentration indicative for melt-rock interaction processes.

1123 Symbols as in Fig. 3.

1124 (a) U vs. Th

1125 (b) Zr vs. Th

1126

1127 Fig. 8: Flow chart illustrating the processes affecting the composition of refractory
1128 peridotite within the thermal boundary layer and in the near-seafloor
1129 environment. The various pathways that are documented in the samples
1130 derived from ODP Leg 209 are indicated and the sites showing the relevant
1131 evidence are marked. A magmatic regime where ductile deformation and melt-
1132 rock interaction processes are dominant is contrasted to a hydrothermal regime
1133 where temperatures are $<450^{\circ}\text{C}$ and fluid-peridotite (and/or serpentinite)
1134 interaction causes the formation of various alteration assemblages.

1135

1136 Fig. 9: Deviating trends for melt-rock interaction and hydrothermal alteration are
1137 defined in LREE vs. HFSE space. This is due to the higher solubility of LREE in
1138 aqueous solutions compared to HFSE. In contrast, LREE and HFSE are equally
1139 soluble in silicic melts.

1140 (a) Nb vs. La

1141 (b) Th vs. Ce

1142 (c) Nd vs. Gd_N/Lu_N

1143 (d) Th vs. Gd_N/Lu_N

1144

1145 Fig. 10: Results of mass balance calculations for Hole 1274A. **White squares:**
1146 **serpentinized harzburgite; black squares: serpentinized dunite.**

1147 (a, b, c, and d) The reconstructed bulk rock geochemical composition of the
1148 precursor peridotites at Hole 1274A shows well constrained correlations with
1149 the modal proportions of pyroxene and olivine. This indicates that the
1150 recalculation process generated realistic results.

- 1151 (e) Enrichment factors have been calculated for each sample using SiO_2 , Al_2O_3 ,
1152 MgO , and $\text{FeO}_{\text{total}}$ as immobile monitor elements. The calculations based on
1153 SiO_2 , MgO , and $\text{FeO}_{\text{total}}$ yield consistent results and increasing H_2O contents of
1154 the samples are correlated with decreasing EF values. The EF values
1155 calculated based on the Al_2O_3 data show wide scatter.
- 1156 (f) There is a positive correlation among the calculated mass gains for the
1157 individual samples and the observed volume proportions of micro veins in thin
1158 sections. This indicates that the mass gain was accommodated by a volume
1159 increase during serpentinization.

1160

1161 Fig. 11: Effects of hydrothermal alteration on Fe oxidation state and REE systematics
1162 at Hole 1274A. **White squares: serpentinized harzburgite; black squares:**
1163 **serpentinized dunite.**

1164 (a) A positive correlation of $\text{Fe}_2\text{O}_3_{\text{calculated}}$ (determined from XRF analyses and FeO
1165 measurements using titration methods) and H_2O content demonstrates that
1166 increasing intensity of serpentinization is coupled with oxidation of iron. This is
1167 also reflected in the formation of magnetite at an advanced stage of
1168 serpentinization.

1169 (b) A negative trend of Lu_N and H_2O suggests that mass addition during
1170 serpentinization may have diluted the concentration of HREE which were
1171 immobile during hydrothermal alteration. However, it is difficult to ascertain that
1172 HREE concentrations were identical for all the samples prior to alteration and,
1173 hence, primary and secondary factors may both be important in generating this
1174 relationship.

1175 (c) Whereas a well-constrained trend exists for Lu_N and H_2O (**Fig. 11b**) the
1176 variations in La_N are not correlated with the degree of serpentinization. This

1177 suggests that La may have been mobile during alteration. Alternatively, it could
1178 be argued that La_N/Lu_N ratios were variable in the unaltered protoliths of Hole
1179 1274A.

1180 (d) Elevated La_N/Lu_N ratios are common in the deeper portions of Hole 1274A
1181 where serpentinization is complete.

1182

1183 Fig. 12: Variations in the Eu/Eu^* values show no correlation with Sr concentrations,
1184 hence, excluding the possibility that trace amounts of plagioclase could be
1185 responsible for the high Eu concentrations observed in samples representing
1186 fluid dominated serpentinization. Symbols as in **Fig. 3**.

1187

1188 Fig. 13: Comparison of the REE pattern of serpentinites from Hole 1268A (fluid
1189 dominated serpentinization) with the composition of vent type hydrothermal
1190 fluids sampled at the Rainbow and Logatchev ultramafic hosted hydrothermal
1191 systems. Fluid data from Douville et al., 1997 and 2002.

1192

1193 Fig. 14: Results of mass change calculations for talc alteration of serpentinite at Hole
1194 1268A. Three different scenarios have been considered: no mass change,
1195 immobility of SiO_2 and immobility of MgO. Based on textural observations and
1196 considerations regarding reaction processes (Bach et al., 2004) it is inferred
1197 that immobility of MgO is the most likely scenario. Mass addition of SiO_2 is at
1198 least partially accommodated by hydrothermal veining and volume expansion.
1199 Data are presented in Tables 5 and 6.

1200

1201

1202

1203 Tables:

1204

1205 **Table 1: Summary of drill holes, ODP Sites 1268, 1270, 1271, 1272, and 1274.**

1206 Table 2: Geochemical composition of variably altered peridotite from ODP Sites 1268,
1207 1270, 1271, 1272, and 1274.

1208 Table 3: Analytical limits and analyses for the international standards JP-1, PCC-1,
1209 and NIM-N by the geochemical laboratories in Montpellier (France) and
1210 Göteborg (Sweden).

1211 Table 4: Bulk-rock geochemical data of 83 samples from ODP Leg 209 (electronic
1212 supplement).

1213 Table 5: Hole 1274A – reconstructed primary mineralogy, composition, and mass
1214 change during serpentinization.

1215 Table 6: Calculation of mass changes due to talc alteration of serpentinite at Hole
1216 1268A.

1217

Figure01

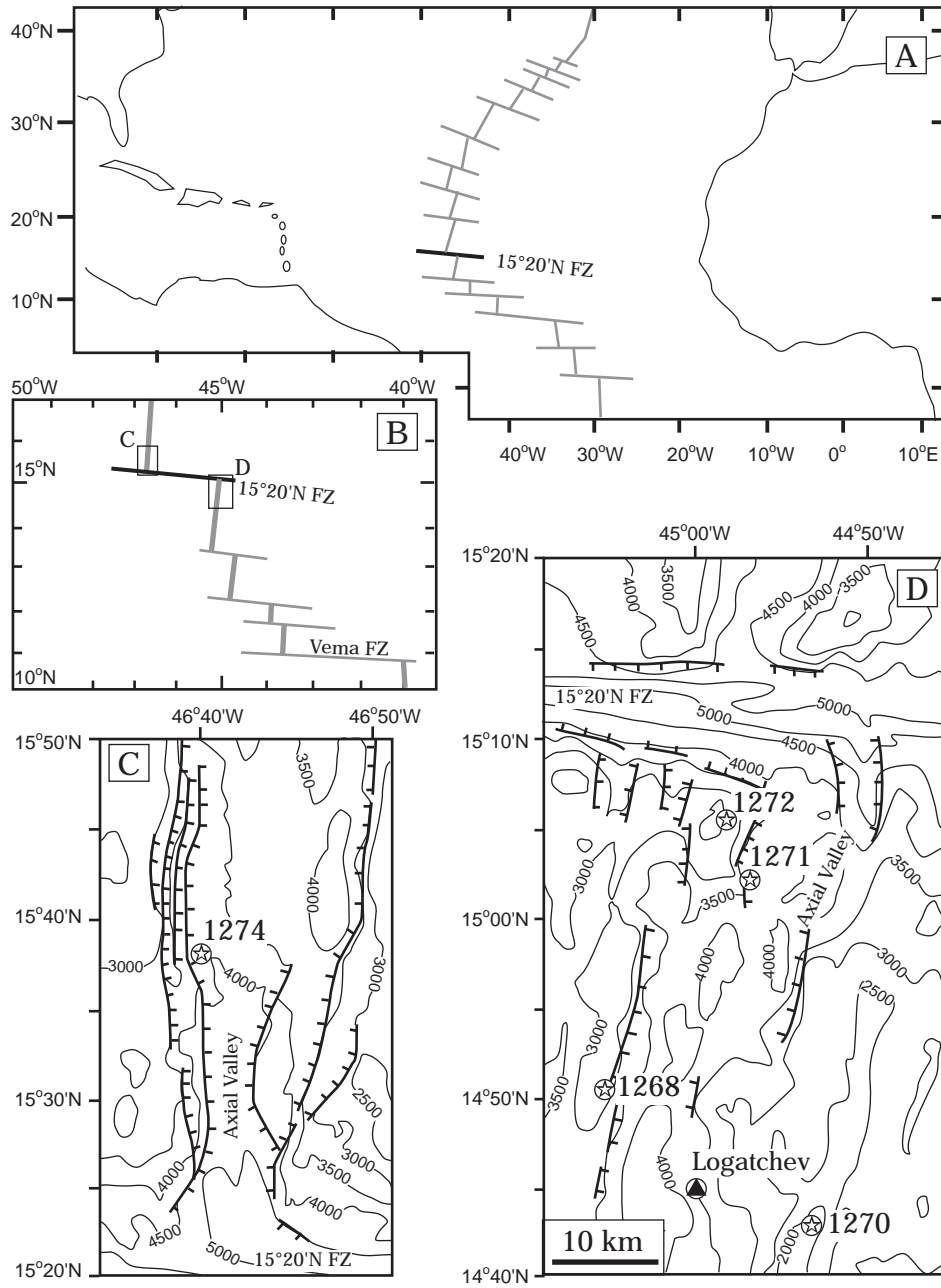
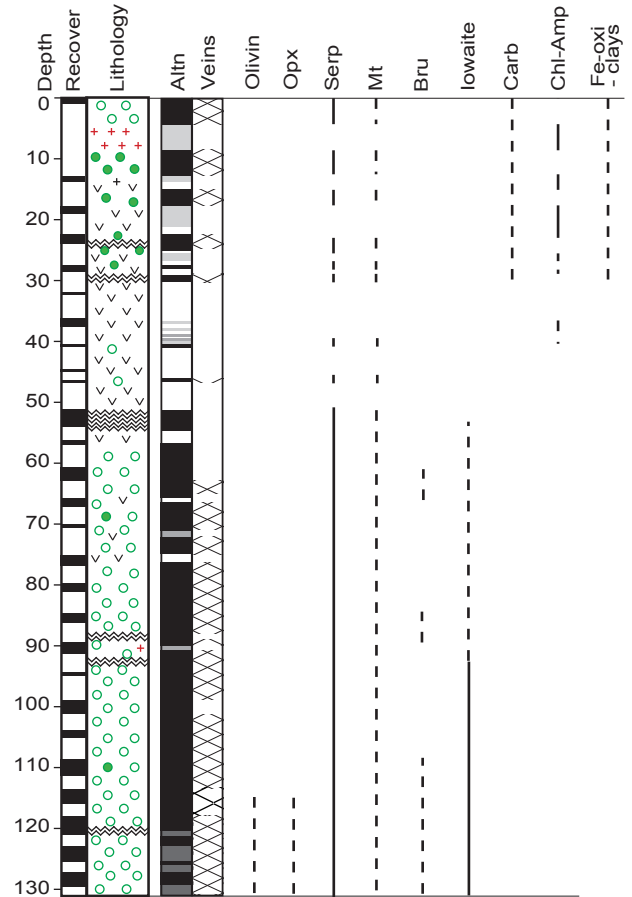
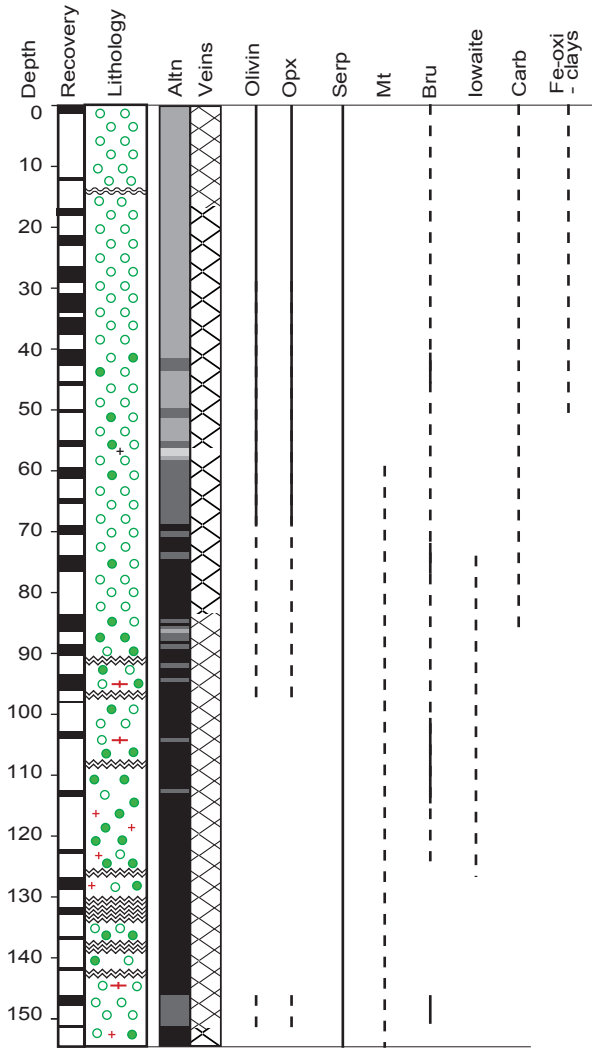


Fig. 1

Figure02

A: Hole 1274A

B: Hole 1272A



Lithology



harzburgite



olivine gabbro, troctolite



melt-impregnated dunite



basalt, diabase



dunite



(micro)gabbro, gabbronorite



amphibole gabbro



magmatic dikelets



brittle and ductile deformation zone

Alteration



completely altered (>95 %) □



highly altered (>40 - 80 %) □



fresh to slightly altered (0 - 10 %) □



moderately veined (>1 to 5 vol%) □



very highly altered (>80 - 95 %) □



moderately altered (>10 - 40 %) □



poorly veined (>0.1 to 1 vol%) □



intensely veined (>5 vol%) □

Distribution of minerals

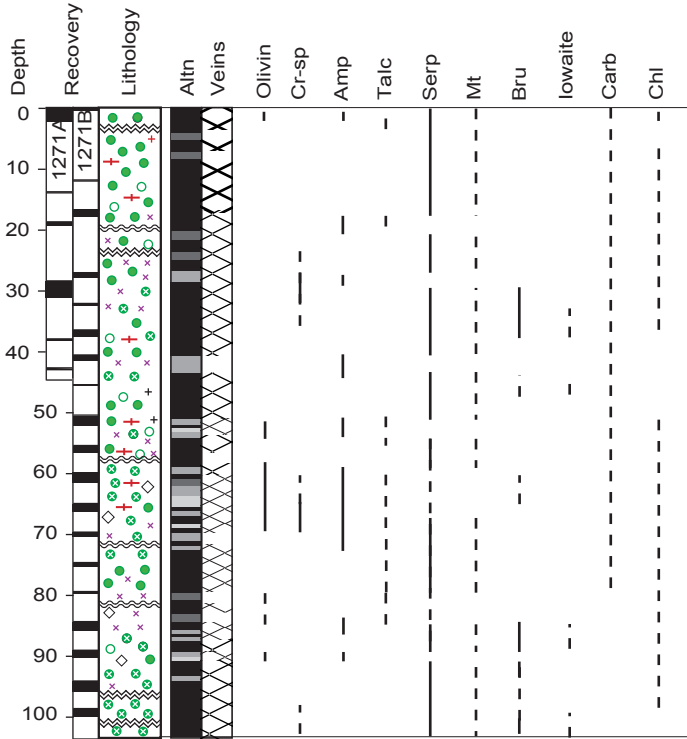


common to abundant

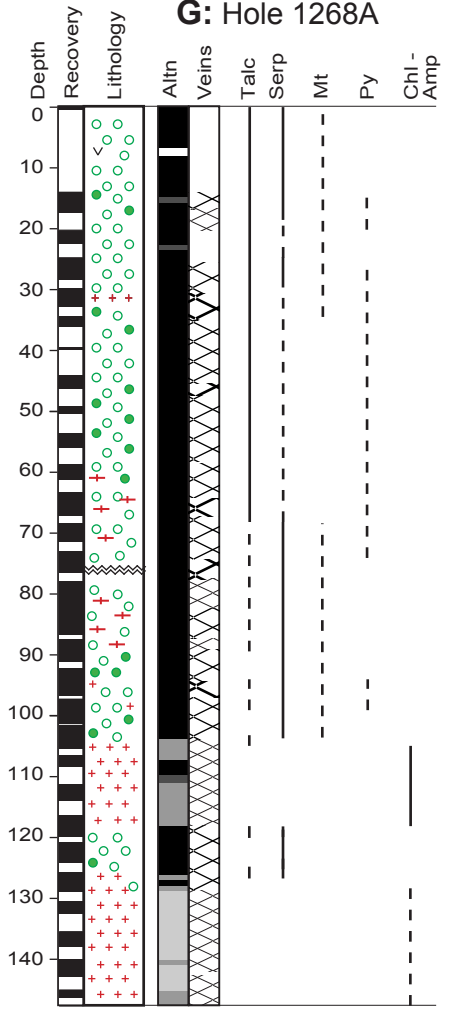


minor

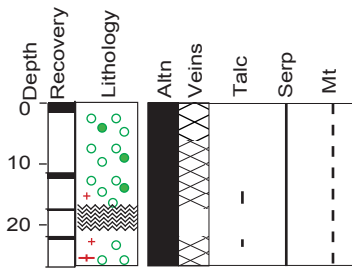
C: Holes 1271A and B



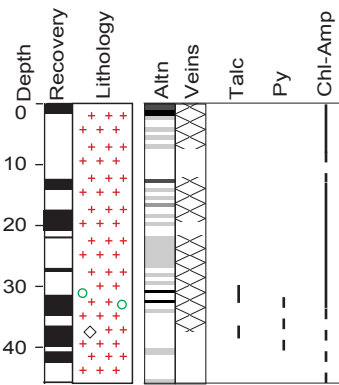
G: Hole 1268A



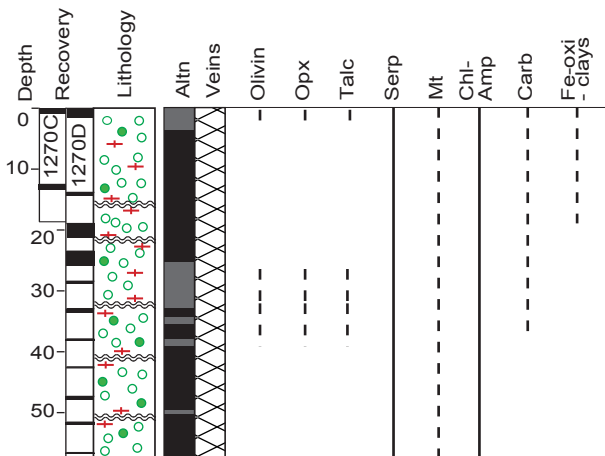
D: Hole 1270A

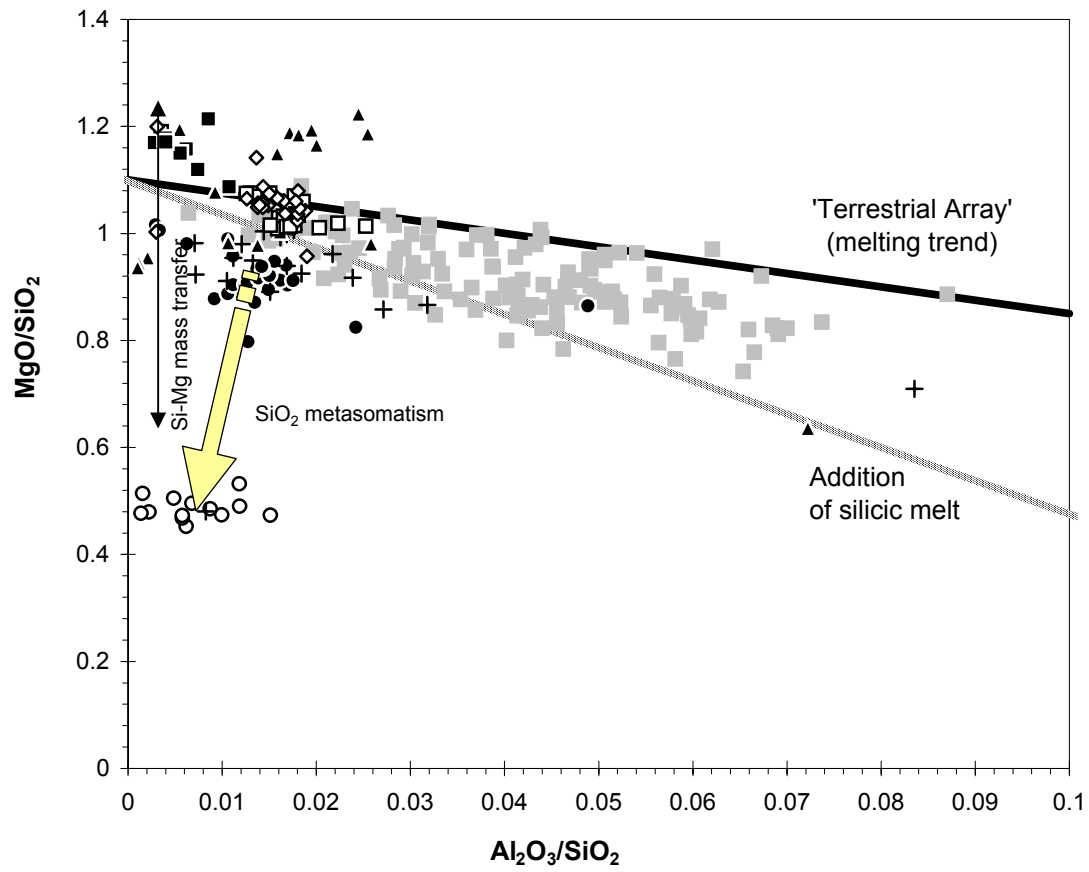


E: Hole 1270B



F: Holes 1270C and D





- Site 1268 - Serpentinite
- Site 1268 - Talc alteration
- + Site 1270 - Serpentinite (+/- melt-impregnation)
- ▲ Site 1271 - Serpentinite (chromite-bearing dunite +/- melt impregnation)
- ◇ Site 1272 - lowaite-bearing serpentinite
- Site 1274 - Variably serpentinized harzburgite
- Site 1274 - Variably serpentinized dunite
- Global abyssal peridotites (Niu, 2004)

Fig. 3

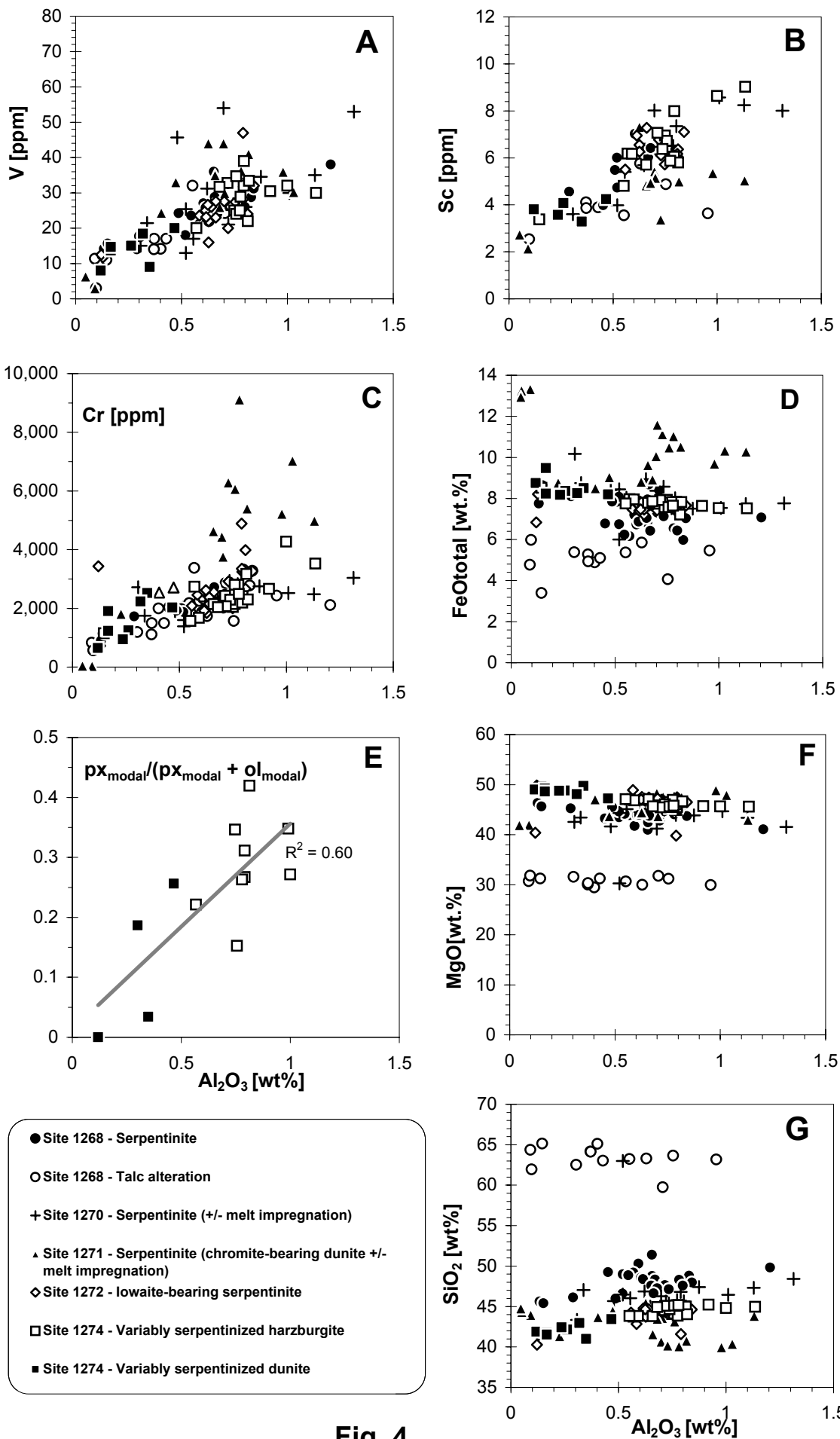


Fig. 4

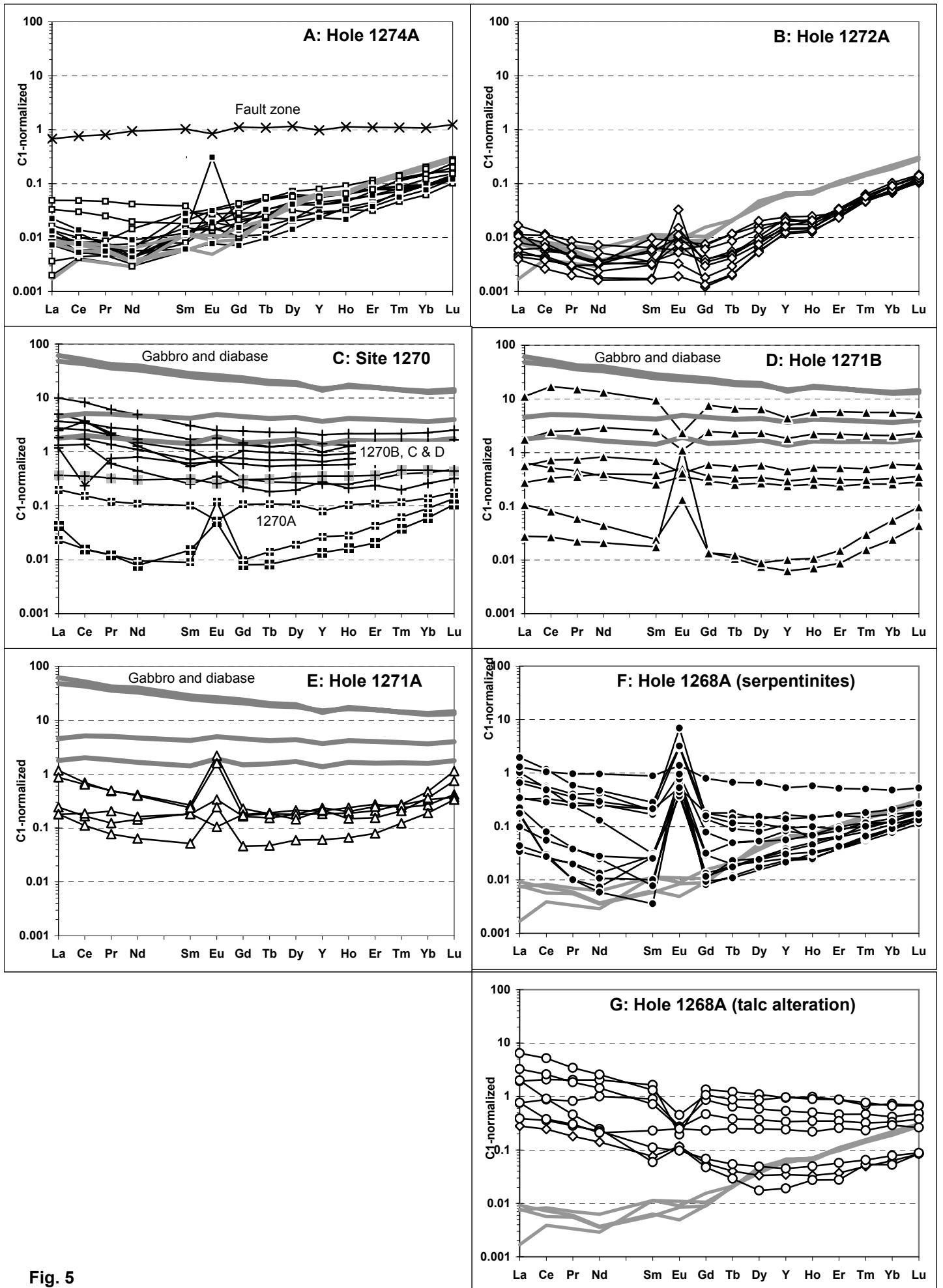


Fig. 5

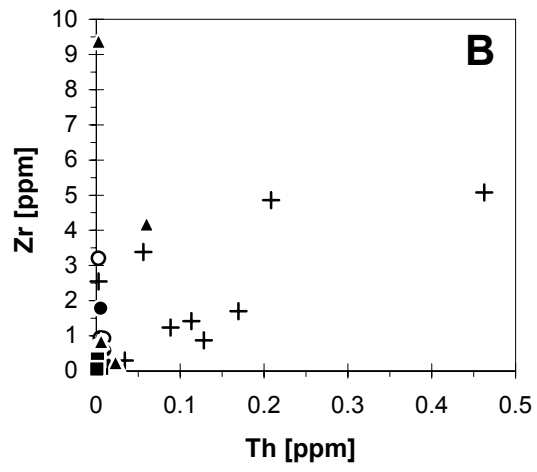
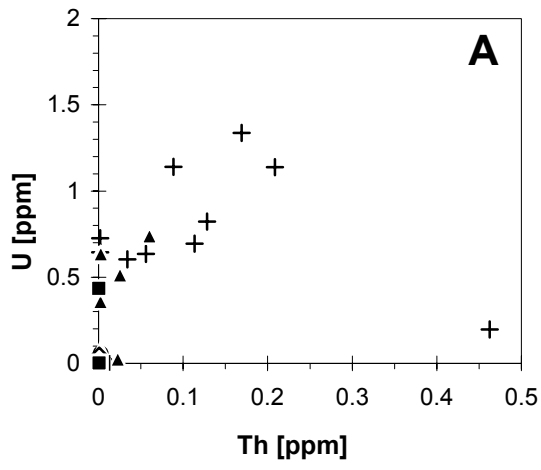


Fig. 6

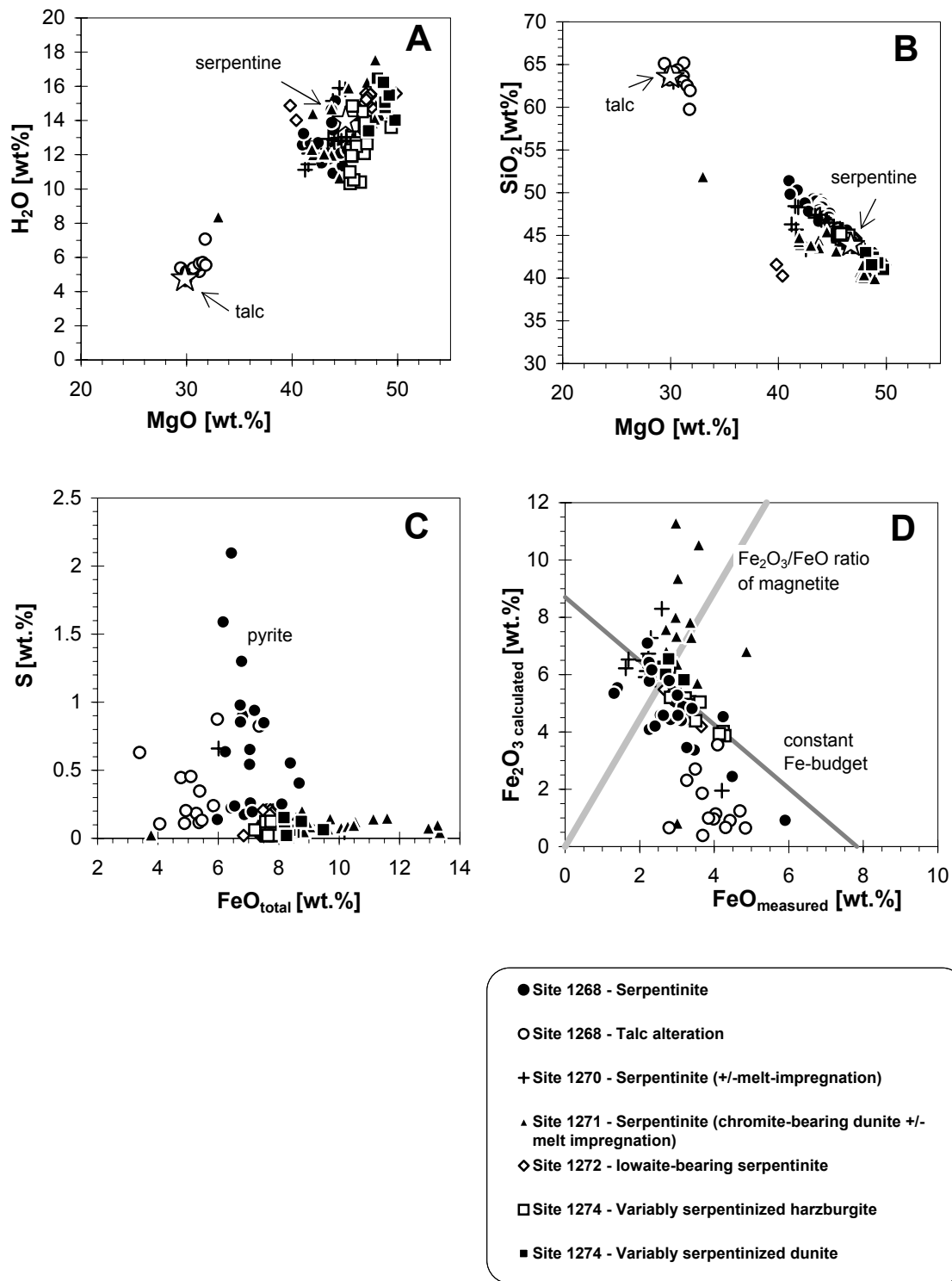


Fig. 7

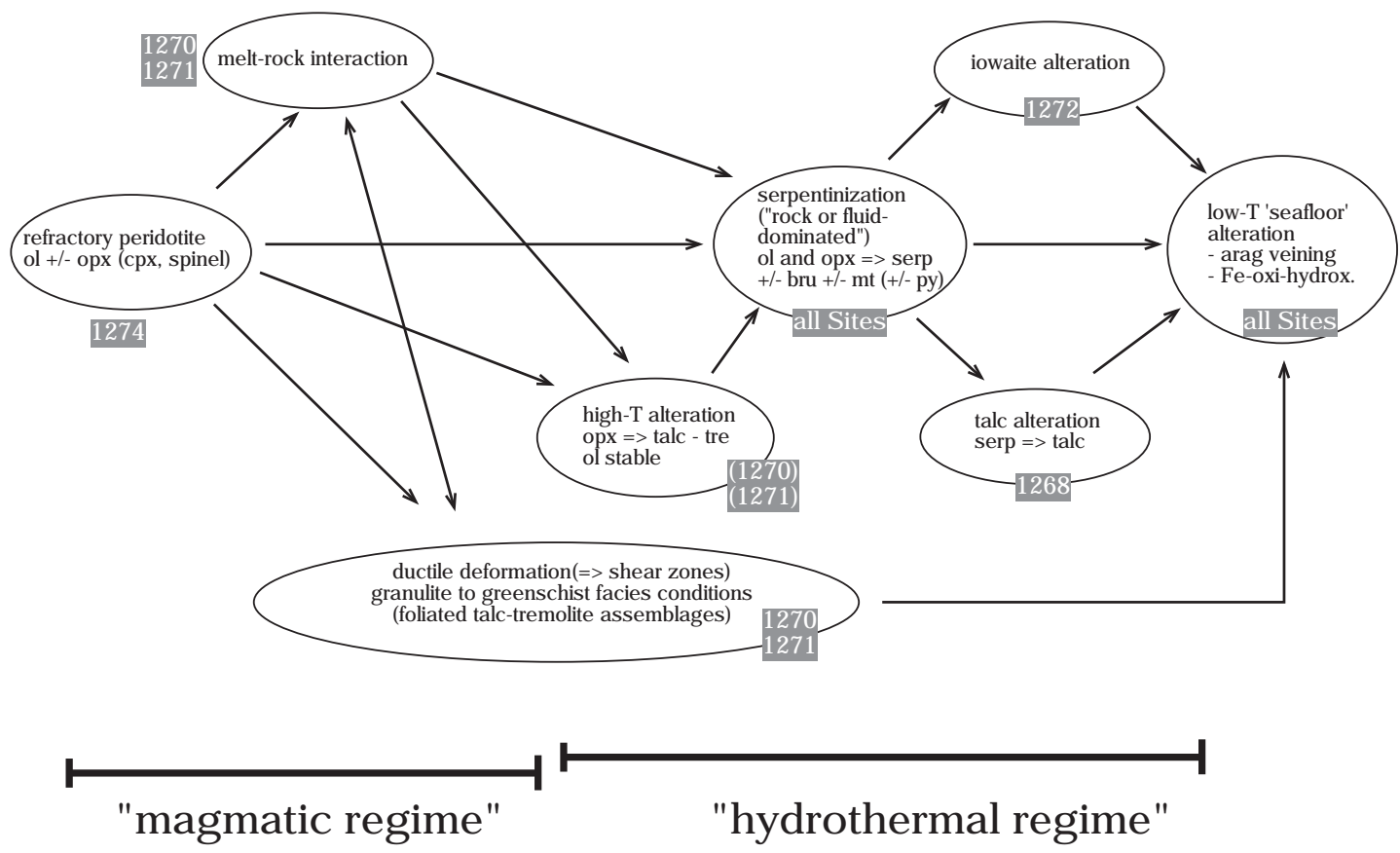
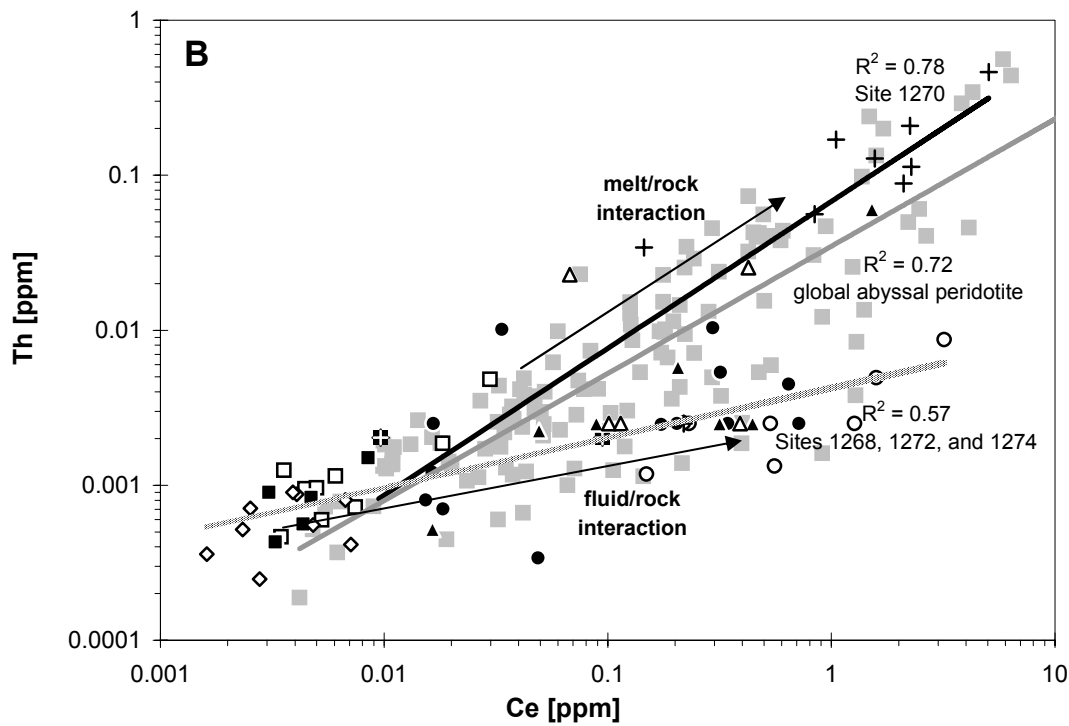
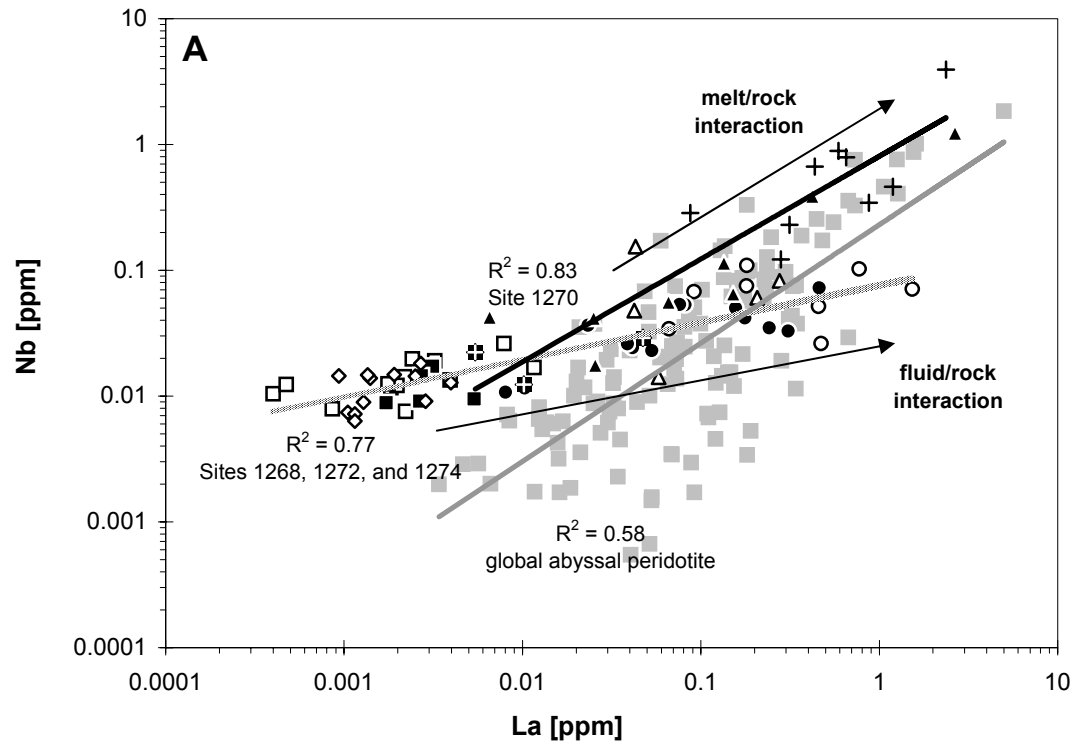


Fig. 8



- Site 1268 - Talc alteration
- Site 1268 - Serpentinite
- Hole 1270A - Serpentinite
- + Holes 1270B, C, & D - Serpentinite (+/- melt impregnation)
- △ Hole 1271A - Serpentinite (+/- melt impregnation)
- ▲ Hole 1271B - Serpentinite (+/- melt impregnation)
- Site 1274 - Variably serpentinized harzburgite
- Site 1274 - Variably serpentinized dunite
- ◇ Site 1272 - Serpentinite (+/- iowaite)
- global abyssal peridotite (Niu, 2004)

Fig. 9

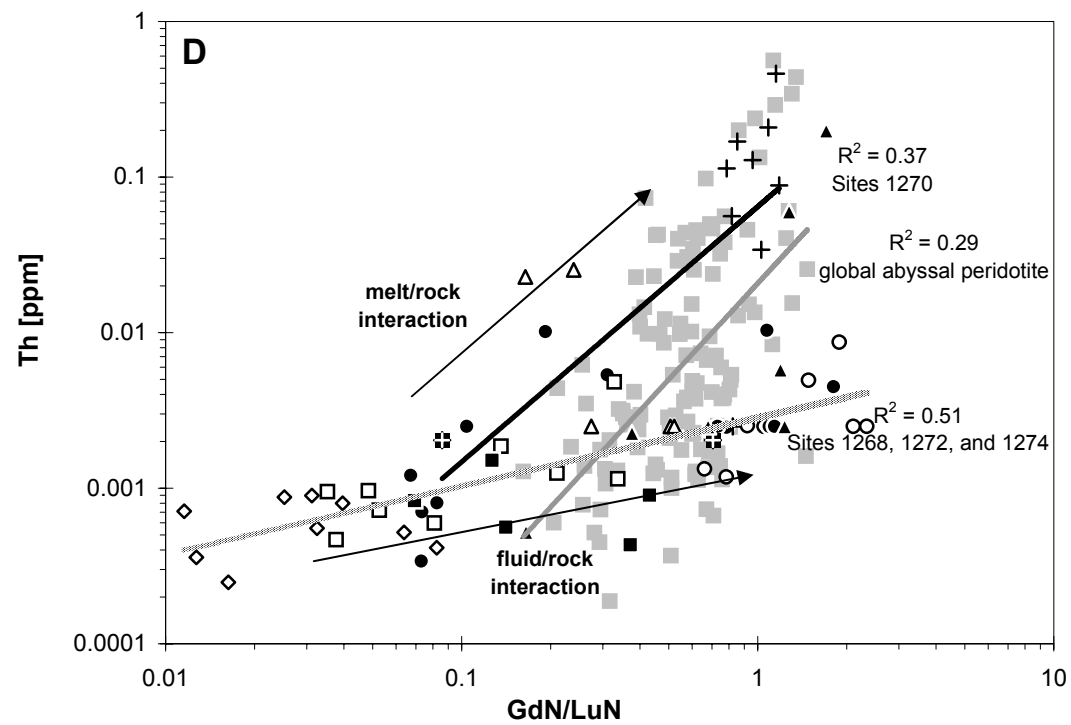
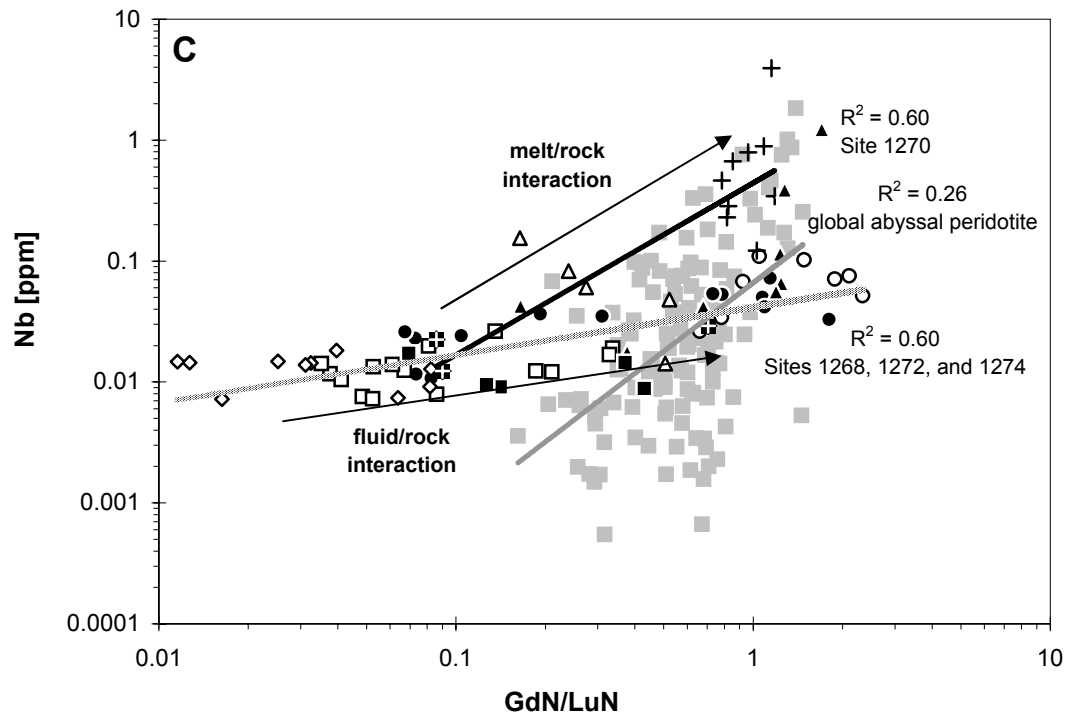


Fig. 9 continued

Figure10

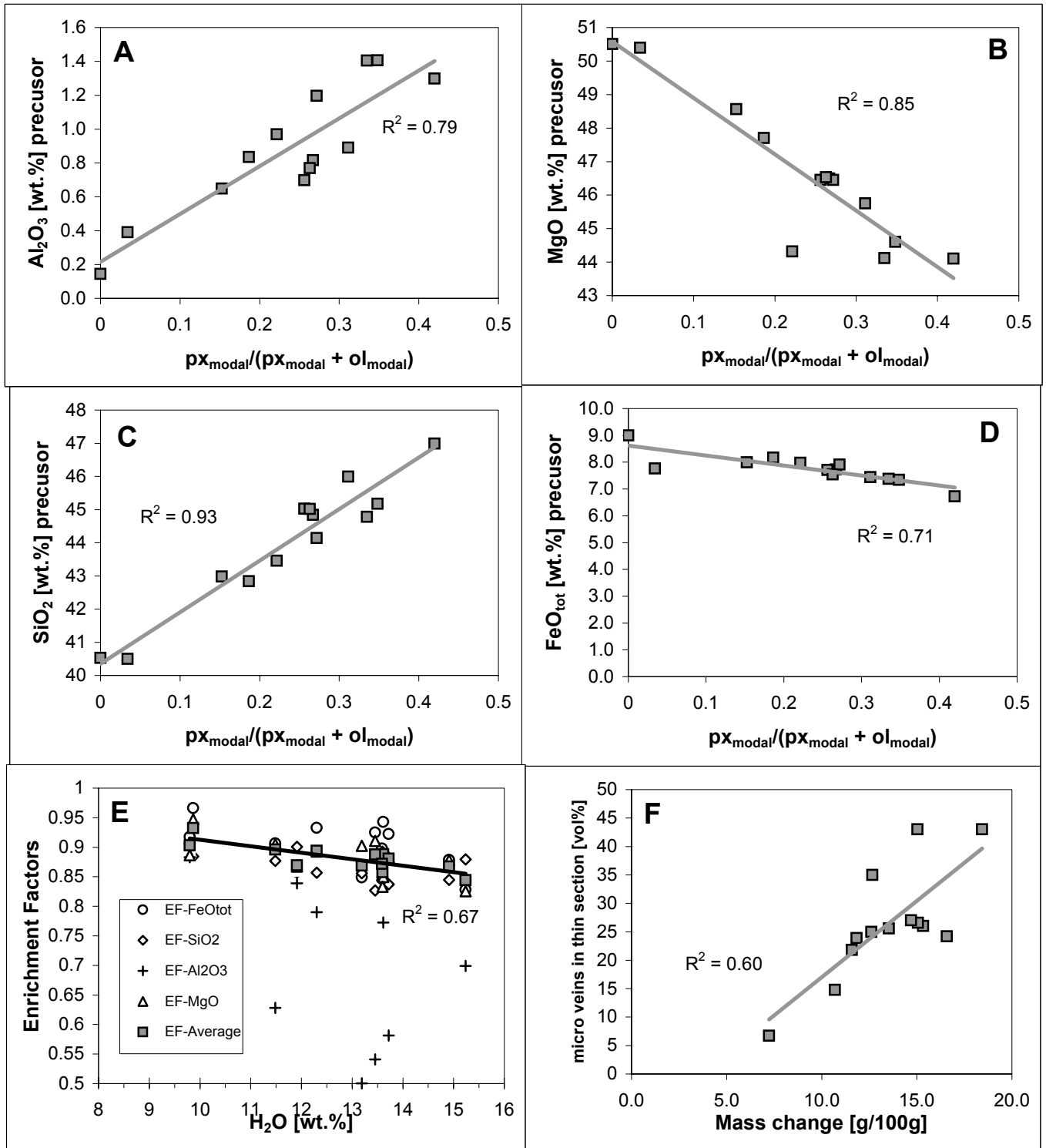


Fig. 10

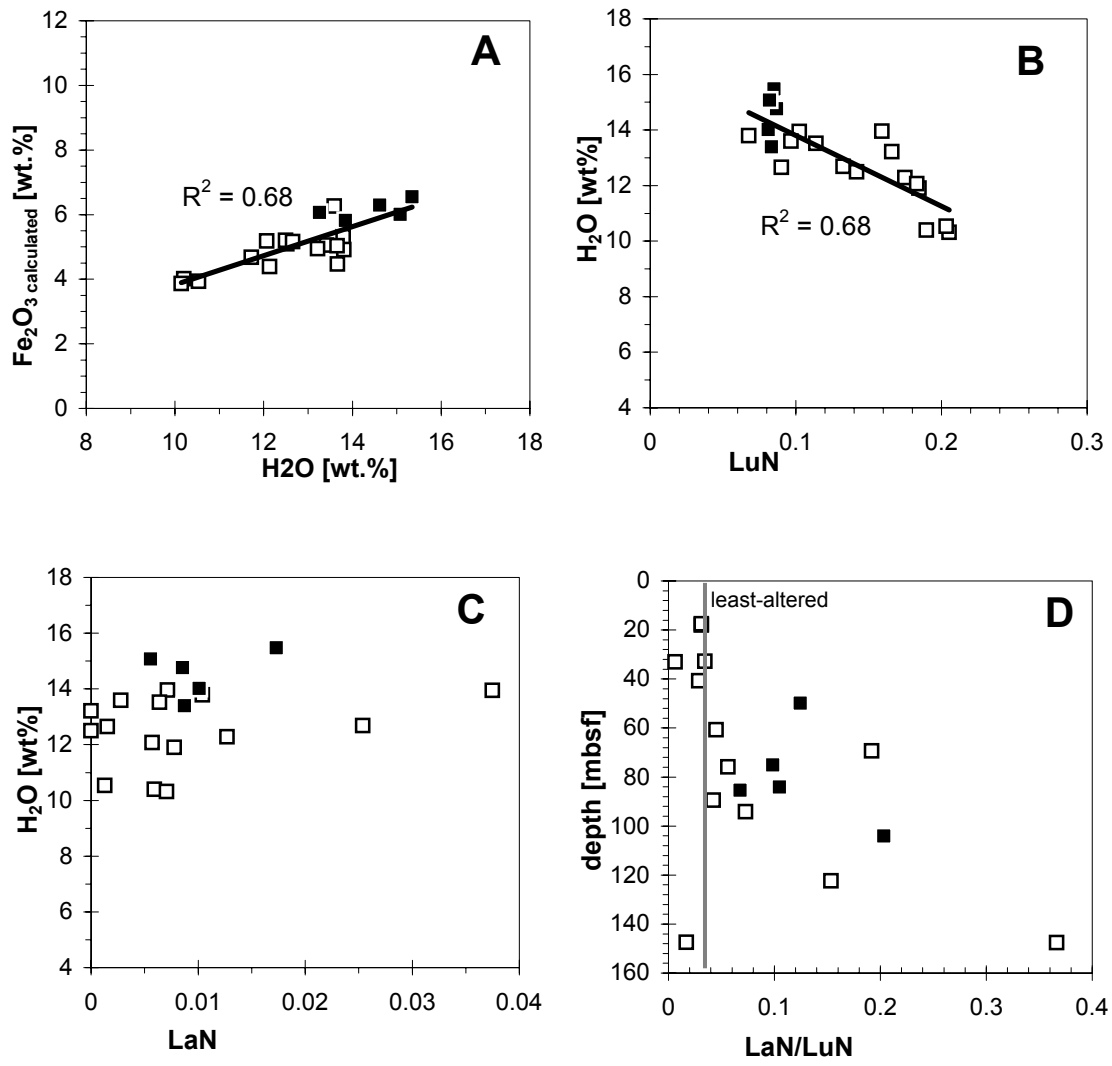


Fig. 11

Figure12

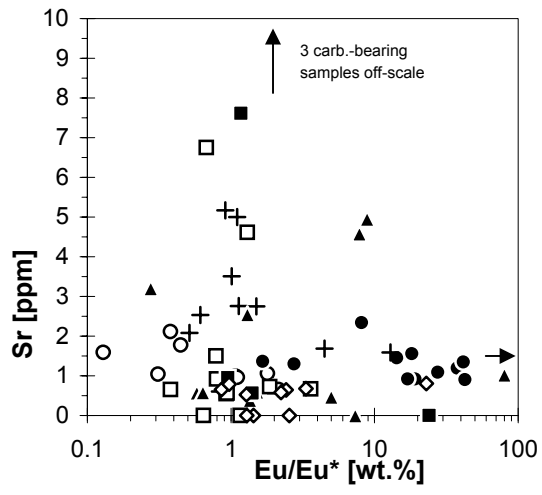


Fig. 12

Figure13

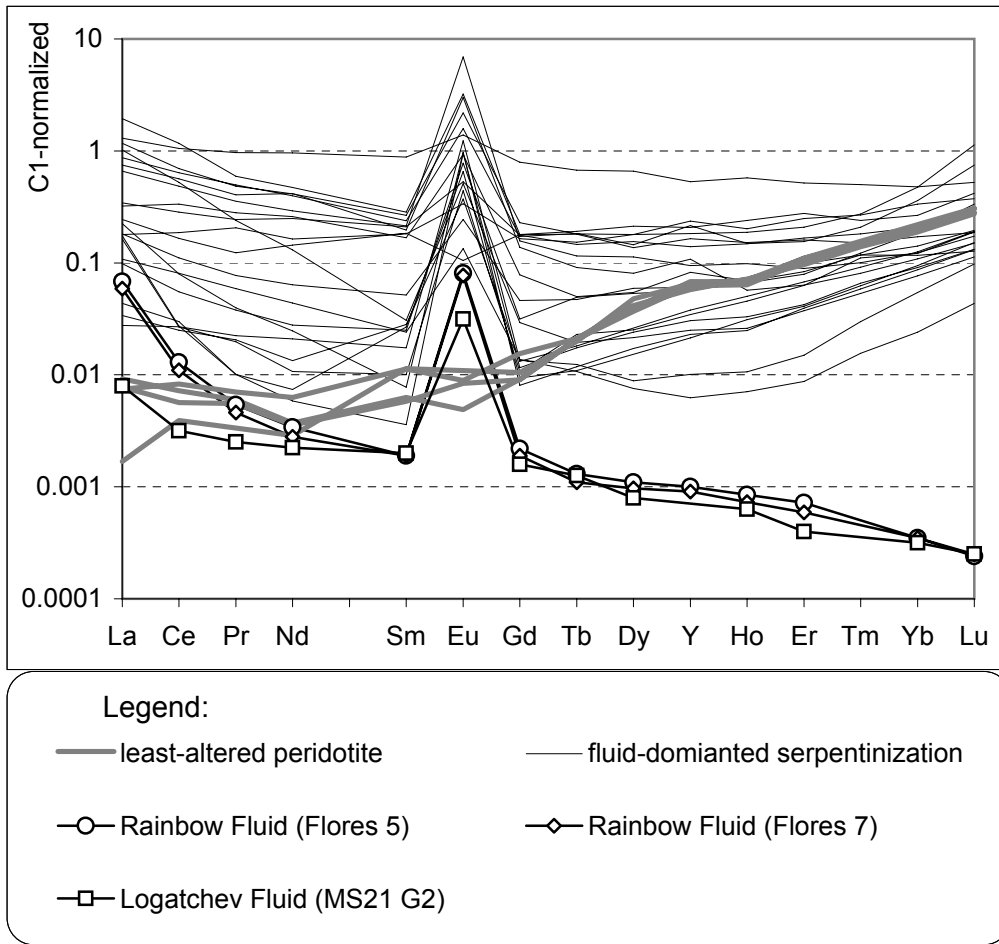


Fig. 13

Figure14

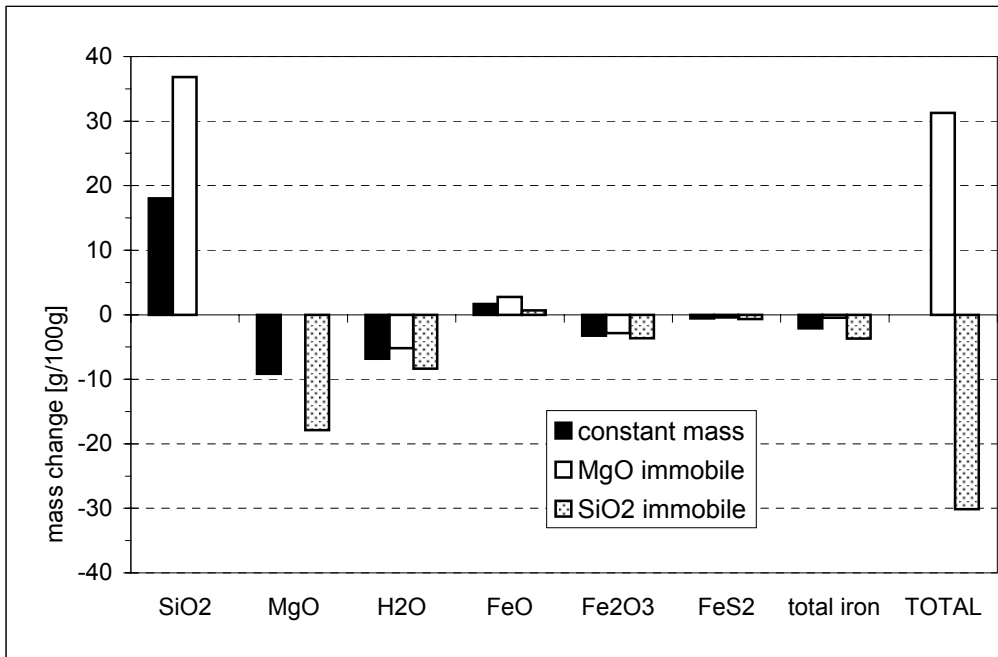


Fig. 14

Table01

[Click here to download Table: Table01_Drilling_summary.pdf](#)

Table 1: Summary of drill holes, ODP Sites 1268, 1270, 1271, 1272, and 1274

Holes	1268A	1270A	1270B	1270C	1270D	1271A	1271B	
Latitude	14°50.755'N	14°43.342'N	14°43.265'N	14°43.284'N	14°43.270'N	15°02.222'N	15°02.189'N	
Longitude	45°04.641'W	44°53.321'W	44°53.225'W	44°53.091'W	44°53.084'W	44°56.887'W	44°56.912'W	
Water depth (m)	3007	1951	1910	1822	1817	3612	3585	
Basement penetration (m)	147.6	26.9	45.9	18.6	57.3	44.8	103.6	
Core recovered (m)	78.7	3.3	17.2	2.0	7.7	5.8	15.9	
Lithology	Harzburgite	63%	89%	0.5%	81%	91%	1%	9%
	Dunite	11%	5%	-	17%	7%	98%	56%
	Gabbroic	26%	4%	99.5 %	1%	2%	1%	35%
	Fault gouge	-	2%	-	1%	-	-	-
Alteration of peridotite	Serpentine, talc, pyrite, hematite, magnetite	Serpentine, magnetite	talc	Serpentine, magnetite	Serpentine, magnetite	Serpentine, brucite, magnetite	Serpentine, brucite, magnetite	
Veining	Talc, serpentine, sulfide, oxide	Serpentine, talc, magnetite	talc, chlorite, serpentine, amphibole, sulfide	Serpentine, oxide, talc, carbonate	Serpentine, oxide, talc, carbonate, sulfide, magnetite	Serpentine, talc, amphibole, magnetite, carbonate	Serpentine, magnetite, talc, carbonate, amphibole	
Abundance of metamorphic veins	8.9%	1.2%	0.5%	2.0%	1.9%	6.8%	3.1%	
Alteration intensity (peridotite)	98-100%	98-100%	100%	98-100%	50-100%	30-100%	30-100%	

Modified from Bach et al., 2004

Table 2: Geochemical composition of variably altered peridotite from ODP Leg 209.

Hole	1268A	1268A	1268A	1268A	1268A	1268A	1270D	1270D
Core	19	2	19	8	4	3	3	3
Section	1	2	3	1	3	1	2	2
Depth ^a (cm)	34	108	6	28	26	29	39	85
Depth (mbsf)	97.3	16.48	99.88	44.3	28.04	20.49	20.6	21.01
Rock type	FD serp, Hz	FD serp, Hz	FD serp, Du	talc altn, Hz	talc altn, Hz	talc altn, Hz	serp Hz, MRI	serp Hz, MRI
Lab code	GR-06	AP-003	AP-023	GR-03	AP-008	AP-004	GR-09	AP-046
XRF Lab	OU	UB	UB	OU	UB	UB	OU	UB
[wt%]								
SiO ₂	40.42	40.47	40.83	59.27	60.31	60.6	42.14	41.05
TiO ₂	0.01	<0.01	<0.01	0.02	<0.01	<0.01	0.05	<0.01
Al ₂ O ₃	0.44	0.68	0.56	0.90	0.6	0.35	1.34	0.98
Fe ₂ O ₃ ^{tot}	7.55	6.08	6.72	5.69	6.19	5.54	7.11	7.45
MgO	38.83	38.07	38.12	28.06	28.57	28.35	36.51	37.65
MnO	0.07	0.07	0.09	0.08	0.04	0.03	0.11	0.09
CaO	0.03	0.18	0.05	0.10	0.03	<0.02	0.48	0.15
Na ₂ O	0.08	0.1	0.1	0.24	0.16	0.16	0.16	0.13
K ₂ O	0.01	<0.01	<0.01	0.02	<0.01	<0.01	0.05	0.02
P ₂ O ₅	0.01	0.01	<0.01	0.01	<0.01	<0.01	0.02	<0.01
LOI	12.61	13.25	12.42	4.36	4.58	4.68	12.31	12.06
Total	100.06	98.91	98.89	98.74	100.48	99.71	100.28	99.58
FeO ^a	2.64	1.20	2.56	4.57	4.50	4.21	2.66	2.11
Fe ₂ O ₃ ^b	4.62	4.74	3.87	0.61	1.19	0.86	4.15	5.10
Fe ²⁺ ratio	0.39	0.22	0.42	0.89	0.81	0.84	0.42	0.32
CO ₂	0.09	<0.45	<0.45	0.04	<0.45	<0.45	0.16	<0.45
S	0.21	2.09	0.65	0.13	0.24	0.18	0.05	0.04
N	bdl	<0.03	<0.03	bdl	<0.03	<0.03	0.009	<0.03
[ppm]								
Cr _{XRF}	2005	3319	2236	2431	1736	1109	2469	2481
Ni _{XRF}	2138	2087	2102	1715	1634	1599	1999	1999
ICP-MS Lab	UM	UG	UG	UM	UG	UG	UM	UG
Li	0.13	nd	nd	bdl	nd	nd	7.04	nd
Ba	0.04	3.44	0.58	0.10	bdl	2.11	6.22	3.48
Cs	0.002	<0.013	<0.013	0.002	<0.013	<0.013	0.011	<0.013
Co	96	92	106	91	79	75	93	98
Ni	2021	2029	2075	1589	1500	1534	1990	1994
As	nd	1.07	0.33	nd	0.45	0.40	nd	0.70
V	nd	29	22	nd	26	13	nd	35
Cr	nd	3006	1429	nd	1851	1126	nd	2420
Cu	13	15	35	11	546	82	4	3
Pb	0.24	7.79	0.59	0.11	0.32	0.37	0.03	0.13
Zn	nd	38	36	nd	100	66	nd	50
Sc	5.5	6.4	6.4	3.6	5.8	4.1	6.2	8.2
Rb	0.06	0.19	0.06	0.07	0.07	0.06	0.66	0.19
Sr	0.92	2.34	1.31	2.11	1.59	1.04	5.17	3.50
Y	0.06	0.13	0.04	1.52	1.48	0.84	3.28	2.23
Zr	bdl	0.58	0.09	0.93	0.66	3.20	5.07	0.87
Nb	0.01	0.05	0.02	0.07	0.05	0.08	3.94	0.79
Hf	0.00	<0.1	<0.1	0.02	<0.1	<0.1	0.18	<0.1
Ta	0.001	0.035	0.014	0.004	0.019	0.015	0.223	0.067
Th	0.001	0.003	0.000	0.009	0.003	0.003	0.462	0.128
U	0.0009	0.0025	0.0025	0.0075	0.0025	0.0025	0.1968	0.8221
La	0.0081	0.0760	0.0532	1.5309	0.4573	0.1805	2.3686	0.6516
Ce	0.0153	0.2044	0.0488	3.1947	1.2717	0.5349	5.0475	1.5654
Pr	0.0019	0.0267	0.0038	0.3276	0.1932	0.0780	0.5820	0.1952
Nd	0.0062	0.1210	0.0116	1.2065	0.9556	0.4706	2.2942	0.8047
Sm	0.0043	0.0256	0.0012	0.2015	0.2526	0.1379	0.4716	0.2044
Eu	0.0215	0.0540	0.0562	0.0261	0.0112	0.0158	0.1452	0.0789
Gd	0.0027	0.0161	0.0017	0.2190	0.2784	0.1766	0.5027	0.2771
Tb	0.0007	0.0019	0.0004	0.0329	0.0456	0.0242	0.0858	0.0465
Dy	bdl	0.0139	0.0044	0.2191	0.2796	0.1483	0.5834	0.3415
Ho	0.0028	0.0039	0.0014	0.0505	0.0560	0.0284	0.1229	0.0738
Er	0.0112	0.0163	0.0065	0.1457	0.1441	0.0767	0.3608	0.2369
Tm	0.0023	0.0030	0.0015	0.0196	0.0175	0.0119	0.0560	0.0377
Yb	0.0214	0.0199	0.0148	0.1141	0.1231	0.0707	0.3852	0.2525
Lu	0.0048	0.0032	0.0033	0.0171	0.0175	0.0123	0.0642	0.0424

Table 2: Geochemical composition of variably altered peridotite from ODP Leg 209.

Hole	1270D	1271A	1271B	1271B	1271B	1272A	1272A	1272A
Core	9	4	10	12	17	7	21	27
Section	1	2	1	1	1	1	1	2
Depth ^a (cm)	44	5	30	46	61	116	27	12
Depth (mbsf)	47.84	29.8	50.8	60.6	85.1	38.1	99.17	128.62
Rock type	serp Hz, MRI	serp Hz	serp Du, MRI	Gabbro	serp Du, MRI	Diabase	serp-iow Hz	serp-iow Hz
Lab code	AP-050	GR-11	AP-063	GR-13	GR-15	GR-17	AP-077	AP-083
XRF Lab	UB	OU	UB	OU	OU	OU	UB	UB
[wt%]								
SiO ₂	39.78	33.93	40.65	39.36	34.27	52.10	37.51	36.95
TiO ₂	<0.01	0.02	<0.01	0.07	0.01	1.48	<0.01	<0.01
Al ₂ O ₃	0.48	0.61	0.56	3.90	0.19	15.36	0.61	0.53
Fe ₂ O ₃ ^{tot}	7.72	10.45	8.77	9.03	8.08	9.31	7.27	7.34
MgO	38.99	40.20	39.77	34.79	40.95	7.20	39.81	40.16
MnO	0.07	0.13	0.1	0.12	0.10	0.15	0.09	0.1
CaO	<0.02	0.04	0.28	1.53	0.03	10.97	<0.02	0.05
Na ₂ O	0.14	0.06	0.17	0.46	0.06	2.62	0.17	0.11
K ₂ O	<0.01	0.03	<0.01	0.18	0.01	0.58	<0.01	<0.01
P ₂ O ₅	0.01	0.01	<0.01	0.01	0.01	0.23	<0.01	<0.01
LOI	12.77	13.79	10.28	10.27	14.55	0.88	14.60	14.57
Total	99.96	99.27	100.58	99.73	98.26	100.88	100.06	99.81
FeO ^a	2.05	4.16	2.71	3.81	2.97	6.75	2.49	2.48
Fe ₂ O ₃ ^b	5.44	5.83	5.76	4.80	4.78	1.81	4.50	4.58
Fe ²⁺ ratio	0.30	0.44	0.34	0.47	0.41	0.81	0.38	0.38
CO ₂	<0.45	0.18	<0.45	0.27	0.21	0.25	<0.45	<0.45
S	0.05	0.14	0.05	0.07	0.20	0.10	0.21	0.18
N	<0.03	0.002	<0.03	bdl	bdl	bdl	<0.03	<0.03
[ppm]								
Cr _{XRF}	1753	6298	2549	4567	1813	284	2399	2013
Ni _{XRF}	2350	2182	2249	1580	2549	97	2111	2163
ICP-MS Lab	UG	UM	UG	UM	UM	UM	UG	UG
Li	nd	bdl	nd	3.40	11.00	3.20	nd	nd
Ba	bdl	0.83	3.89	8.29	0.13	113.94	3.14	2.27
Cs	<0.013	0.002	0.016	0.021	bdl	0.099	<0.013	<0.013
Co	114	121	111	143	45	50	100	98
Ni	2309	2020	2217	1580	253	124	2115	2094
As	4.34	nd	1.26	nd	nd	nd	0.40	0.44
V	23	nd	40	nd	nd	nd	21	18
Cr	1670	nd	2450	nd	nd	nd	1443	1319
Cu	5	6	4	4	3	67	6	6
Pb	0.16	0.05	0.23	0.15	0.27	0.86	0.17	0.06
Zn	49	nd	61	nd	nd	nd	45	45
Sc	5.4	3.4	7.3	13.1	23.2	32.0	6.8	6.3
Rb	0.06	0.43	0.07	1.49	0.08	7.76	bdl	0.05
Sr	2.75	0.47	2.55	4.53	0.39	95.49	0.65	0.81
Y	0.44	0.10	0.38	2.15	0.47	23.37	0.03	0.02
Zr	3.38	0.25	0.29	1.52	0.84	125.96	0.10	0.09
Nb	0.23	0.15	0.07	0.10	0.06	19.68	0.01	0.01
Hf	<0.1	0.01	<0.1	0.06	0.02	3.18	<0.1	<0.1
Ta	0.021	0.007	0.019	0.002	0.002	1.115	<0.005	0.005
Th	0.056	0.023	0.003	0.041	0.006	1.216	0.001	0.001
U	0.6360	0.0262	0.6362	0.1460	0.0065	0.3867	0.0511	0.0025
La	0.3140	0.0430	0.1514	0.4250	0.0658	11.3345	0.0019	0.0014
Ce	0.8431	0.0677	0.3166	1.2277	0.2058	26.3415	0.0041	0.0025
Pr	0.0574	0.0073	0.0430	0.1737	0.0343	3.4116	0.0005	0.0003
Nd	0.2085	0.0299	0.1697	0.7720	0.1869	15.5050	0.0016	0.0008
Sm	0.0390	0.0079	0.0396	0.2175	0.0596	3.7672	0.0005	0.0003
Eu	0.0206	0.0141	0.0209	0.1116	0.0295	1.2988	0.0004	0.0019
Gd	0.0451	0.0095	0.0608	0.3044	0.0745	4.3021	0.0006	0.0003
Tb	0.0068	0.0018	0.0095	0.0578	0.0126	0.6893	0.0002	0.0001
Dy	0.0495	0.0151	0.0688	0.4339	0.0875	4.5085	0.0019	0.0014
Ho	0.0119	0.0037	0.0145	0.0934	0.0187	0.9227	0.0012	0.0007
Er	0.0394	0.0131	0.0395	0.2645	0.0522	2.5732	0.0056	0.0042
Tm	0.0050	0.0031	0.0066	0.0413	0.0080	0.3637	0.0015	0.0012
Yb	0.0443	0.0321	0.0443	0.2670	0.0554	2.2780	0.0153	0.0126
Lu	0.0081	0.0085	0.0072	0.0450	0.0092	0.3643	0.0036	0.0032

Table 2: Geochemical composition of variably altered peridotite from ODP Leg 209.

Hole	1274A	1274A	1274A	1274A
Core	6	6	18	16
Section	2	3	1	2
Depth ^a (cm)	128	24	83	26
Depth (mbsf)	32.73	33.1	94.13	85.4
Rock type	least-alt Hz	least-alt Hz	serp Hz	serp Du
Lab code	AP-086	GR-23	AP-096	GR-26
XRF Lab	UB	OU	UB	OU
[wt%]				
SiO ₂	40.05	39.62	39.21	35.72
TiO ₂	<0.01	0.01	<0.01	0.01
Al ₂ O ₃	0.71	0.64	0.99	0.20
Fe ₂ O ₃ ^{tot}	7.77	7.65	7.28	7.65
MgO	40.64	40.77	39.74	41.08
MnO	0.11	0.11	0.12	0.09
CaO	0.77	0.85	0.52	0.13
Na ₂ O	0.05	0.05	0.06	0.06
K ₂ O	<0.01	0.01	<0.01	0.01
P ₂ O ₅	<0.01	0.01	<0.01	0.01
LOI	9.80	10.48	11.91	15.36
Total	99.90	100.19	99.83	100.31
FeO ^a	3.86	3.71	3.08	2.29
Fe ₂ O ₃ ^b	3.48	3.53	3.86	5.10
Fe ²⁺ ratio	0.55	0.54	0.47	0.33
CO ₂	<0.45	0.25	<0.45	0.36
S	0.04	0.07	0.08	0.15
N	<0.03	0.002	<0.03	0.003
[ppm]				
Cr _{XRF}	2415	2427	3529	943
Ni _{XRF}	2146	2134	2056	2453
ICP-MS Lab	UG	UM	UG	UM
Li	nd	bdl	nd	0.04
Ba	1.24	0.14	1.12	0.001
Cs	0.016	0.001	<0.013	90
Co	97	97	100	108
Ni	2054	1930	2043	2273
As	0.43	nd	0.49	nd
V	31	nd	31	nd
Cr	1636	nd	2583	nd
Cu	6	8	6	13
Pb	bdl	bdl	bdl	0.06
Zn	45	nd	45	nd
Sc	8.0	7.1	9.0	3.6
Rb	bdl	0.04	bdl	0.01
Sr	0.10	0.26	0.67	0.95
Y	0.11	0.10	0.10	0.08
Zr	0.06	bdl	0.10	0.05
Nb	0.01	0.01	0.01	0.01
Hf	<0.1	0.00	<0.1	0.01
Ta	<0.005	0.001	<0.005	0.001
Th	0.001	bdl	0.001	0.001
U	0.0025	bdl	0.0025	0.0014
La	0.0022	0.0004	0.0039	0.0017
Ce	0.0044	0.0024	0.0074	0.0031
Pr	0.0006	bdl	0.0007	0.0005
Nd	0.0017	0.0013	0.0036	0.0030
Sm	0.0010	0.0017	0.0012	0.0043
Eu	0.0003	bdl	0.0020	0.0019
Gd	0.0019	0.0022	0.0024	0.0091
Tb	0.0008	0.0007	0.0009	0.0019
Dy	0.0100	0.0121	0.0092	0.0148
Ho	0.0038	0.0040	0.0036	0.0040
Er	0.0182	0.0182	0.0169	0.0129
Tm	0.0040	0.0038	0.0036	0.0020
Yb	0.0375	0.0346	0.0321	0.0149
Lu	0.0078	0.0077	0.0067	0.0031

Abbreviations: altn: alteration; bdl: below detection limit; FD serp: fluid-dominated serpentinization; Du: dunite; Hz: harzburgite; least-alt: least altered; LOI: loss on ignition; MRI: melt-rock interaction REE signature; nd: not determined; serp: serpentinized; serp-iow Hz: iowaite-bearing serpentinized harzburgite;

Laboratories:
UB: University of Bonn; UG: University of Göteborg; UM: University of Montpellier; OU: Open University.

^a: FeO determined by titration

^b: Fe₂O₃ recalculated from XRF and titration data

Table03

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Table 3: Analyses of international standards at the universities of Montpellier (France) and Göteborg (Sweden)

	JP-1					PCC-1			NIM-N			LOD Montpellier	LOD Göteborg	Batch blanks Montpellier	Batch blanks Göteborg
	Average Montpellier	Standard Deviation	Average Göteborg	Standard Deviation	Preferred Values ^a	Average Montpellier	Standard Deviation	Preferred Values ^b	Average Göteborg	Standard Deviation	Preferred Values ^c				
N analyses	2		2			3			3					9	4
ppm															
Li	1.46		1.71		1.8	1.12		1.2	6		6	0.003	0.025	b.d.l	
Sc	6.37	0.06	6.57	0.13	7.24	7.1	0.4	8.4	39		39.8	0.001	0.014	b.d.l	12.65
Ti	21.3	0.4	21.5			29	1	29	1152		1200	0.05	0.22	396	
Co	116	1	109	0	116	110	4	112	59		58	0.01	0.003	11	
Ni	2425	21	2302	15	2460	2325	80	2380	122		120	0.03	0.037	137	
V	na		20	0		na			224		220		0.013		
Cr	na		1576	20		na			29		30		0.37		
Cu	na		7	0		na			12		14		1.5		
Zn	na		45	1		na			61		68		0.31		
As	na		1	0		na			0.4				0.009		
Sn	na		0.026			na			0.7		1		0.05		
Sb	na		<dl			na			0.2		0.06		0.005		
ppb															
Rb	308	9	289	11	340	59	7	66	4191		4900	0.99	5.36	1.69	
Sr	662	75	<500		570	375	12	380	256191		260000	1.13	22.30	9.86	
Y	98	3	90	1	100	78	3	70	5150	216	6500	1.25	0.41	b.d.l	1.17
Zr	6039	442	6149	64	5340	128	10	134	11550		23000	1.51	6.89	6.51	
Nb	49	2	54	10	36	21.0	1.3	29	500		2000	0.39	1.08	0.39	
Cs	40	2	58	7	35	7.3	0.6	5.5	241		240	0.40	0.53	b.d.l	
Ba	11514	605	10345	54	9800	1468	89	900	82178		84000	1.22	194	177	
La	38	10	28		30	28.9	1.0	35	2878	150	2840	0.44	0.12	0.54	0.57
Ce	76	19	60		54	52.5	1.2	72	5912	261	5610	0.20	0.41	0.89	1.27
Pr	10	2	8		7.1	7.3	0.4	8.6	736	36	710	0.15	0.04	0.17	0.19
Nd	38	9	34		33	26.7	1.4	35	3301	139	3000	0.83	0.25	1.79	0.89
Sm	11	2	8		13	6.2	0.9	6.6	814	32	836	0.96	0.25	b.d.l	1.04
Eu	1.89	0.04	3.27		3.1	1.0	0.1	1.2	583	22	588	0.28	0.03	b.d.l	0.21
Gd	10.2	0.7	11.4		13	6.16	1.07	8.66	943	33	940	0.63	0.09	b.d.l	0.64
Tb	2.1	0.1	2.0		2.6	1.15	0.03	2	156	5	164	0.12	0.01	b.d.l	0.10
Dy	18.3	2.5	13.6		18	11.3	0.7	11	1027	35	1100	0.54	0.14	0.60	0.42
Ho	4.2	0.5	3.1		4.3	3.2	0.1	3	217	7	240	0.09	0.02	b.d.l	0.11
Er	13.9	0.9	12.0		14	11.95	0.97	12	632	21	660	0.72	0.24	b.d.l	0.49
Tm	2.5	0.1	2.1		2.7	2.54	0.17	2.87	94	2	99	0.21	0.04	b.d.l	0.10
Yb	21.7	1.5	20.1		21	22.8	1.3	24	641	16	656	0.48	0.18	b.d.l	0.56
Lu	4.6	0.3	4.2		4.7	5.2	0.3	5.7	96	2	101	0.21	0.02	b.d.l	0.11
Hf	143	6	317	155	120	4.2	0.3	6	495		360	0.47	0.98	b.d.l	
Ta	4.3	0.2	68.0	21.9	3.6	1.2	0.3	1.8	102		64	0.19	0.28	b.d.l	
Pb	78	7	na	80	90	8043	184	8500	na		7000	1.94		21.7	
Th	14	1	12	14	12	10.9	0.2	10	304	5	340	0.16	0.17	b.d.l	0.43
U	13.0	0.4		15.5	12	4.5	0.4	4.3	264		400	0.17	0.63	b.d.l	

Abbreviations: b.d.l.: below detection limit; LOD: Limit of detection; n.a.: not analyzed

Preferred Values ^a : after Govindaraju (1994) and Godard et al (2000); Preferred Values ^b : after Govindaraju (1994) and Takazawa et al (2003); Preferred Values ^c : after Govindaraju (1994) and Korotev (1996).

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Table05

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Table 5: Hole 1274A – reconstructed primary mineralogy, composition, and mass change during serpentinization.

Hole	1274A	1274A	1274A	1274A	1274A	1274A	1274A	1274A
Core	3	6	10	12	14	15	15	16
Section	1	2	1	1	1	1	2	1
Top (cm)	111	128	3	134	30	106	39	44
Bottom (cm)	120	135	10	142	36	114	46	52
Depth (mbsf)	18.01	32.73	49.93	60.74	69.3	75.06	75.86	84.14
Piece #	9	4	2	10	5	15A	5	7
Rock type	least-alt Hz	least-alt Hz	serp Du	serp Hz	serp Hz	serp Du	serp Hz	serp Du
modal data [vol%]								
olivine	20.6	20	7.6	2.2	5.2	0.4	3.1	2.1
orthopyroxene	13.4	8.3	0.3	1.5	3.9	0.1	1	0.6
clinopyroxene	0.3	0.6	0.2	1.2	2.1	0.7	0.8	1.3
spinel	1.6	0.3	1.2	0.5	0.1	1.3	1.5	0.2
magnetite	0.9	1.8	9.6	2.6	3.7	13.4	7	6.9
serp after ol	30	28.8	25	36.4	25.4	22.6	24.9	21.5
brucite after ol	9.8	9.7	29.5	22.5	17.2	24.1	23.7	24
serp after opx	16.7	12.1	0.7	7.8	15.5	8.7	11.4	14.3
serp, origin? ^a	0	3.6	1.1	1.1	3	2.7	0	0
serp in vein	6.3	11.9	17.1	18.9	19.6	17.1	20.9	14.3
mgt in vein	0.4	2.9	7.9	5.3	4.3	8.9	5.7	11.3
total:	100	100	100	100	100	100	100	100
prim mineral ^b								
total olivine	65.37	73.31	96.58	84.74	68.88	81.35	77.86	74.38
total opx	32.58	25.56	1.56	12.90	27.95	15.20	18.67	23.28
cpx	0.32	0.75	0.20	1.66	3.03	1.21	1.20	2.03
spinel	1.73	0.38	1.87	0.69	0.14	2.25	2.26	0.31
total	100.00	100.00	100.00	100.00	100.00	100.00	100.00	100.00
n EMP ^c								
olivine	18	17	20	13	15	23	1	16
orthopyroxene	9	8	4	7	4	2	7	4
clinopyroxene	9	13	8	16	6	9	14	12
spinel	7	5	8	3	8	6	4	7
reconst comp ^d								
[wt%]								
SiO ₂	44.58	44.27	40.29	42.39	45.15	42.24	43.93	44.45
TiO ₂	0.01	0.01	0.00	0.01	0.02	0.01	0.05	0.03
Al ₂ O ₃	1.40	0.81	0.39	0.64	0.87	0.82	0.98	0.69
Cr ₂ O ₃	0.97	0.38	1.09	0.42	0.31	1.15	1.30	0.36
FeO _{tot}	7.34	7.63	7.72	7.88	7.30	8.06	8.06	7.61
MgO	43.92	45.91	50.14	47.89	44.92	47.03	44.80	45.86
MnO	0.13	0.13	0.07	0.13	0.13	0.11	0.18	0.12
CaO	1.64	0.85	0.29	0.64	1.28	0.56	0.69	0.85
Na ₂ O	0.01	0.00	0.00	0.00	0.01	0.01	0.01	0.04
K ₂ O	0.00	0.00	0.00	0.00	0.00	0.00	0.00	0.00
total	100.00	100.00	100.00	100.00	100.00	100.00	100.00	100.00
altered rock ^e								
SiO ₂	39.43	40.12	35.11	37.82	38.69	35.67	37.60	37.22
TiO ₂	0.01	0.01	0.00	0.01	0.02	0.01	0.01	0.02
Al ₂ O ₃	0.67	0.71	0.30	0.65	0.69	0.22	0.49	0.40
Cr ₂ O ₃	0.60	0.71	0.74	0.63	0.64	0.37	0.80	0.59
FeO _{tot}	7.09	7.00	7.28	6.68	6.81	7.08	6.85	7.02
MgO	41.58	40.71	42.61	39.89	40.10	41.30	40.43	40.47
MnO	0.11	0.11	0.11	0.11	0.11	0.10	0.10	0.09
CaO	0.55	0.77	0.18	0.50	0.53	0.17	0.43	0.39
Na ₂ O	0.07	0.05	0.06	0.09	0.08	0.08	0.09	0.06
K ₂ O	0.01	0.01	0.00	0.01	0.01	0.01	0.01	0.01
LOI	9.88	9.81	13.60	13.62	12.32	15.00	13.21	13.73
total	100.00	100.00	100.00	100.00	100.00	100.00	100.00	100.00
EF values ^f								
EF _{SiO2}	0.88	0.91	0.87	0.89	0.86	0.84	0.86	0.84
EF _{MgO}	0.95	0.89	0.85	0.83	0.89	0.88	0.90	0.88
EF _{FeOtot}	0.97	0.92	0.94	0.85	0.93	0.88	0.85	0.92
EF _{Al2O3}	0.48	0.88	0.77	1.02	0.79	0.27	0.50	0.58
Average EF ^g	0.93	0.90	0.89	0.86	0.89	0.87	0.87	0.88
total mass change[g/100 g]	7.24	10.70	12.61	16.58	11.83	15.33	15.06	13.53

Table 5: Hole 1274A – reconstructed primary mineralogy, composition, and mass change during serpentinization.

Hole	1274A	1274A	1274A	1274A	1274A
Core	17	18	20	22	27
Section	1	1	1	1	2
Top (cm)	121	83	121	24	5
Bottom (cm)	129	93	126	32	11
Depth (mbsf)	89.51	94.13	104.11	122.34	147.65
Piece #	21	15	23	4	2
Rock type	serp Hz	serp Hz	serp Du	serp Hz	serp Hz
modal data [vol%]					
olivine	7.3	4.1	0	4.6	2.9
orthopyroxene	0.9	1.5	0	5.3	1.4
clinopyroxene	2.2	0.5	0	0.2	0.7
spinel	1.6	1.3	0	0.7	0.1
magnetite	10.4	0.2	1.5	0.5	1.5
serp after ol	27.7	23.5	0	23.8	10.8
brucite after ol	9.2	15.3	53.5	8.8	39
serp after opx	18.9	12.7	0	20.7	16.6
serp, origin? ^a	0	0	0	0	0
serp in vein	10.8	27.9	30	34.6	22.1
mgt in vein	11	15	13	0.8	4.9
total:	100	100	100	100	100
prim mineral ^b					
total olivine	65.19	72.84	100.00	58.03	73.71
total opx	29.20	24.11	0.00	40.56	25.17
cpx	3.24	0.85	0.00	0.31	0.98
spinel	2.36	2.21	0.00	1.09	0.14
total	100.00	100.00	100.00	100.00	100.00
n EMP ^c					
olivine	16	14	1	17	9
orthopyroxene	5	5	0	8	5
clinopyroxene	15	10	0	5	3
spinel	5	8	4	4	7
reconst comp ^d					
[wt%]					
SiO ₂	44.57	43.47	40.25	46.61	44.54
TiO ₂	0.01	0.01	0.00	0.01	0.01
Al ₂ O ₃	1.39	1.18	0.14	1.29	0.76
Cr ₂ O ₃	1.27	0.99	0.22	0.84	0.30
FeO _{tot}	7.24	7.79	8.93	6.67	7.46
MgO	44.01	45.75	50.16	43.75	46.04
MnO	0.13	0.13	0.13	0.13	0.13
CaO	1.38	0.68	0.17	0.67	0.74
Na ₂ O	0.01	0.01	0.00	0.02	0.01
K ₂ O	0.00	0.00	0.00	0.00	0.00
	100.00	100.00	100.00	100.00	100.00
altered rock ^e					
SiO ₂	39.08	39.15	35.40	38.54	37.59
TiO ₂	0.01	0.00	0.01	0.00	0.00
Al ₂ O ₃	0.87	0.99	0.10	0.70	0.66
Cr ₂ O ₃	1.25	1.03	0.19	0.92	0.81
FeO _{tot}	6.56	6.54	7.40	6.17	6.70
MgO	39.82	39.68	41.41	39.85	40.22
MnO	0.12	0.12	0.10	0.10	0.10
CaO	0.73	0.52	0.02	0.21	0.35
Na ₂ O	0.04	0.06	0.08	0.10	0.08
K ₂ O	0.01	0.00	0.01	0.00	0.00
LOI	11.51	11.90	15.29	13.40	13.48
total	100.00	100.00	100.00	100.00	100.00
EF values ^f					
EF _{SiO2}	0.88	0.90	0.88	0.83	0.84
EF _{MgO}	0.90	0.87	0.83	0.91	0.87
EF _{FeOtot}	0.91	0.84	0.83	0.93	0.90
EF _{Al2O3}	0.63	0.84	0.70	0.54	0.87
Average EF ^g	0.90	0.87	0.84	0.89	0.87
total mass change[g/100 g]	11.59	15.03	18.43	12.67	14.70

Abbreviations: cpx: clinopyroxene; Du: dunite; EF: enrichment factor; EMP: electron microprobe; Hz: harzburgite; least-alt: least altered; mgt: magnetite; ol: olivine; opx: orthopyroxene; prim mineral: primary mineralogy; reconst comp: reconstituted composition; serp: serpentinized; tot: total.

For the modal data, 1000 points have been counted in the thin sections on a 0.33 mm x 0.33 mm grid.

- ^a: Serpentinization of uncertain precursor
- ^b: Modal proportions of primary minerals (excluding vein material and serpentinization of unknown origin). Total olivine includes serpentinization after olivine. Total orthopyroxene includes serpentinization after orthopyroxene.
- ^c: Number of microprobe analyses of primary phases. These analytical data are documented in a data report to the Ocean Drilling Program Leg 209 Scientific Results volume (Moll et al., in prep.).
- ^d: The geochemistry of the precursors has been reconstructed by combining the modal proportions of the primary phases and their composition (averages of the microprobe analyses).
- ^e: Normalized composition of the altered rock (analytical data documented in Table 5).
- ^f: Enrichment factors calculated for SiO₂, MgO, FeO_{tot}, and Al₂O₃. Assumptions and procedures involved in this calculations are discussed in the text.
- ^g: Average enrichment factor based on EF_{SiO2}, EF_{MgO}, and EF_{FeOtot}.

Table 6: Calculation of mass changes due to talc alteration of serpentinite at Hole 1268A.

	Serpentinization (n=29)	Talc alteration (n=13)	constant mass	MgO immobile	SiO ₂ immobile
	average "precursor"	average "altered rock"	mass change [g/100g]	mass change [g/100g]	mass change [g/100g]
SiO ₂	41.74	59.77	18.03	36.84	0.00
MgO	38.14	29.01	-9.13	0.00	-17.88
H ₂ O	12.00	5.19	-6.81	-5.18	-8.38
FeO	1.75	3.43	1.68	2.76	0.65
Fe ₂ O ₃	4.58	1.31	-3.27	-2.86	-3.67
FeS ₂	0.97	0.54	-0.43	-0.26	-0.59
total iron ^a	7.30	5.39	-1.91	-0.21	-3.54
TOTAL	99.71	99.74	0.03	31.42	-30.06

The EF_{MgO} is 1.315 and the EF_{SiO_2} is 0.698. For the constant mass scenario the EF is 1.

^a: Sum of FeO + Fe₂O₃ + FeS₂. n = number of analyses. The geochemical data used for calculating the average compositions are documented in Table 5 and in Shipboard Scientific Party, 2004.