

Temperature dependence of η/s of strongly interacting matter: Effects of the equation of state and the parametric form of $(\eta/s)(T)$

Jussi Auvinen,^{1,*} Kari J. Eskola^{2,3}, Pasi Huovinen⁴, Harri Niemi,^{2,3} Risto Paatelainen⁵, and Péter Petreczky⁶

¹*Institute of Physics Belgrade, 11080 Belgrade, Serbia*

²*Department of Physics, University of Jyväskylä, P.O. Box 35, FI-40014 University of Jyväskylä, Finland*

³*Helsinki Institute of Physics, P.O. Box 64, FI-00014 University of Helsinki, Finland*

⁴*Institute of Theoretical Physics, University of Wrocław, 50-204 Wrocław, Poland*

⁵*Theoretical Physics Department, CERN, CH-1211 Geneva 23, Switzerland*

⁶*Physics Department, Brookhaven National Laboratory, Upton, New York 11973, USA*



(Received 24 June 2020; accepted 22 September 2020; published 21 October 2020)

We investigate the temperature dependence of the shear viscosity to entropy density ratio η/s using a piecewise linear parametrization. To determine the optimal values of the parameters and the associated uncertainties, we perform a global Bayesian model-to-data comparison on Au + Au collisions at $\sqrt{s_{\text{NN}}} = 200$ GeV and Pb + Pb collisions at 2.76 TeV and 5.02 TeV, using a 2 + 1D hydrodynamical model with the Eskola-Kajantie-Ruuskanen-Tuominen (EKRT) initial state. We provide three new parametrizations of the equation of state (EoS) based on contemporary lattice results and hadron resonance gas, and use them and the widely used $s95p$ parametrization to explore the uncertainty in the analysis due to the choice of the equation of state. We find that η/s is most constrained in the temperature range $T \approx 150\text{--}220$ MeV, where, for all EoSs, $0.08 < \eta/s < 0.23$ when taking into account the 90% credible intervals. In this temperature range the EoS parametrization has only a small $\approx 10\%$ effect on the favored η/s value, which is less than the $\approx 30\%$ uncertainty of the analysis using a single EoS parametrization. Our parametrization of $(\eta/s)(T)$ leads to a slightly larger minimum value of η/s than the previously used parametrizations. When we constrain our parametrization to mimic the previously used parametrizations, our favored value is reduced, and the difference becomes statistically insignificant.

DOI: [10.1103/PhysRevC.102.044911](https://doi.org/10.1103/PhysRevC.102.044911)

I. INTRODUCTION

The main goal of the ultrarelativistic heavy-ion collisions at the Large Hadron Collider (LHC) and the Relativistic Heavy-Ion Collider (RHIC) is to understand the properties of the strongly interacting matter produced in these collisions. In recent years the main interest has been in extracting the dissipative properties of this QCD matter from the experimental data (e.g., [1–5]), in particular its specific shear viscosity: the ratio of shear viscosity to entropy density η/s (for a review, see Refs. [6–9]). The field has matured to a level where a global Bayesian analysis of the parameters can provide statistically meaningful credibility ranges to the temperature dependence of η/s [10–12]. These credibility ranges agree with earlier results like those obtained using the Eskola-Kajantie-Ruuskanen-Tuominen (EKRT) model [13].

However, with the exception of papers like Refs. [14–16], the equation of state (EoS) is taken as given in the models used

to extract the η/s ratio from the data. Recent fluid-dynamical studies generally use an EoS based on contemporary lattice QCD results, but during the last decade many studies in the literature used the EoS parametrization $s95p$ [17]. This parametrization is based on by now outdated lattice data [18], and recent studies have reported an approximate 60% [16] or 30% increase [19] in the extracted value of η/s when switching from $s95p$ to a contemporary lattice-based EoS. Furthermore, even if the errors of the contemporary lattice QCD calculations overlap, there is still a small tension between the trace anomalies obtained using the highly improved staggered quark (HISQ) [20,21] and stout [22,23] discretization schemes. Consequently the EoSs differ, and if the procedure to extract η/s from the data is as sensitive to the details of the EoS as Refs. [16,19] claim, this tension may lead to additional uncertainties in the η/s values extracted from the heavy-ion collision data.

In the previously mentioned Bayesian analysis [10,11], where the EoS is based on contemporary lattice data [20], the temperature dependence of η/s was assumed to be monotonically increasing above the QCD transition temperature T_c . In a Bayesian analysis the slope parameter of such parametrization is always constrained to be non-negative, and limiting the final slope parameter to zero would require extremely strong constraints from the experimental data. Therefore, by construction, the analysis leads to an η/s increasing with

* auvinen@ipb.ac.rs

Published by the American Physical Society under the terms of the [Creative Commons Attribution 4.0 International](https://creativecommons.org/licenses/by/4.0/) license. Further distribution of this work must maintain attribution to the author(s) and the published article's title, journal citation, and DOI.

temperature above T_c , even if there is no physical reason to exclude a scenario where η/s is constant in a broad temperature range above T_c . A more flexible parametrization, which does not impose such constraints, is thus needed to determine the temperature dependence of η/s .

In this work we address both the sensitivity of the extracted η/s to the EoS used in the model calculation and the temperature dependence of η/s in the vicinity of the QCD transition temperature. We perform a Bayesian analysis of the results of EKRT + hydrodynamics calculations [13,24], and the data obtained in $\sqrt{s_{\text{NN}}} = 200$ GeV Au + Au collisions [25–27] and Pb + Pb collisions at 2.76 TeV [28–30] and 5.02 TeV [30,31]. To study the temperature dependence of η/s we use a piecewise linear parametrization in three parts: linearly decreasing and increasing regions at low and high temperatures are connected by a constant-value plateau of variable range. With this parametrization, data favoring a strong temperature dependence will lead to large slopes and a narrow plateau; conversely, an approximately constant η/s can be obtained with small slope parameter values and a wide plateau. To explore the sensitivity to the EoS, we use four different parametrizations: the well-known $s95p$ parametrization, and three new parametrizations based on contemporary lattice QCD results. A comparison of the final probability distributions of the parameters will tell whether the most probable parameter values depend on the EoS used, and whether that difference is significant when the overall uncertainty in the fitting procedure is taken into account.

II. EQUATION OF STATE

In lattice QCD the calculation of the equation of state (EoS) usually proceeds through the calculation of the trace anomaly, $\Theta(T) = \epsilon(T) - 3p(T)$, where ϵ and p are energy density and pressure, respectively. Thermodynamical variables are subsequently derived from it using the so-called integral method [32]. Therefore we base our EoS parametrizations on the trace anomaly and obtain pressure from the integral

$$\frac{p(T)}{T^4} - \frac{p(T_{\text{low}})}{T_{\text{low}}^4} = \int_{T_{\text{low}}}^T \frac{dT'}{T'^5} \Theta(T'). \quad (1)$$

Once the pressure is known, the energy and entropy densities can be calculated, $\epsilon(T) = \Theta(T) + 3p(T)$ and $s(T) = [\epsilon(T) + p(T)]/T$, respectively. To make a construction of chemical freeze-out at $T \approx 150$ MeV temperature possible, we use the hadron resonance gas (HRG) trace anomaly at low temperatures instead of the lattice QCD result. Equally important is that this choice allows for energy and momentum conserving switch from fluid degrees of freedom to particle degrees of freedom without any nonphysical discontinuities in temperature and/or flow velocity.¹ Furthermore, it gives

¹Energy and momentum conservation require that the fluid EoS is that of free particles, and that the degrees of freedom are the same in the fluid and particles [33]. If the dissipative corrections are small, switch from fluid consistent with the contemporary lattice QCD results [34] to particles in the UrQMD [35] or SMASH [36] hadron

use a consistent value for the pressure at T_{low} required for the evaluation of pressure [see Eq. (1)].

As a baseline, we use the $s95p$ parametrization [17], where HRG containing hadrons and resonances below $M < 2$ GeV mass from the 2004 Particle Data Group (PDG) summary tables [37] is connected to the parametrized hotQCD data from Ref. [18]. To explore the effects of various developments during the last decade, we first connect the HRG based on the PDG 2004 particle list [37] to parametrized contemporary lattice data obtained using the HISQ discretization scheme [20,21]. The lattice spacing, a , is related to the temperature and temporal lattice extent, N_t , as $a = 1/(N_t T)$. Since the lattice spacing (N_t) dependence is small for this action, we use these results at fixed lattice spacing $N_t = 8, 10$, and 12. We name our parametrizations according to the convention used to name $s95p$, and label this parametrization $s87h_{04}$. Here “ $s87$ ” signifies entropy density reaching 87% of its ideal gas value at $T = 800$ MeV, the letter “ h ” refers to the HISQ action, and the subscript “ 04 ” to the vintage of the PDG particle list (2004). Note that even if our parametrization differs from the lattice trace anomaly in the hadronic phase, it agrees with the contemporary lattice calculations which show that at $T = 800$ MeV the entropy density reaches 87%–88% of the ideal gas value (cf. Fig. 8 of Ref. [21]).

The number of well-established resonances has increased since 2004, so we base our parametrization $s88h_{18}$ on HRG containing all² strange and nonstrange hadrons and resonances in the PDG 2018 summary tables³ [40], and on the same HISQ lattice data [20,21] we used for $s87h_{04}$. Furthermore, there is a slight tension in the trace anomaly between the HISQ and stout discretization schemes. To explore whether this difference has any effect on hydrodynamical modeling, we construct the parametrization $s83s_{18}$ using PDG 2018 resonances, and the continuum extrapolated lattice data obtained using the stout discretization [22,23]. The second letter “ s ” in the label refers now to the stout action, and the subscript “ 18 ” to the vintage of the particle list. The details of these parametrizations are shown in Appendix A.

In the top and middle panels of Fig. 1, we show the parametrized trace anomalies, and the lattice data as used to make them: continuum extrapolated for the stout action, and at fixed lattice spacing for the HISQ action, since its lattice spacing (N_t) dependence is small. As seen in the topmost panel, the most noticeable change in the lattice results during the last decade is the reduction of the peak of the trace anomaly (cf. $s95p$ to others). Also, as mentioned, the lattice results obtained using the HISQ and stout actions slightly

cascades at $T = 150$ MeV temperature leads to roughly 9%–10% or 6%–7% loss in both total energy and entropy, respectively.

²With the exception of $f_0(500)$. See Refs. [38,39].

³Note that the PDG Meson Summary Table and Baryon Summary Table contain (almost) all states listed by the PDG, and are different from the PDG Meson Summary Tables and Baryon Summary Tables we use [41]. The PDG Baryon Summary Tables contain the three and four star resonance states. The PDG does not assign stars to meson states, but the Meson Summary Tables contain the states not labeled “Omitted from summary table” in the individual listings.

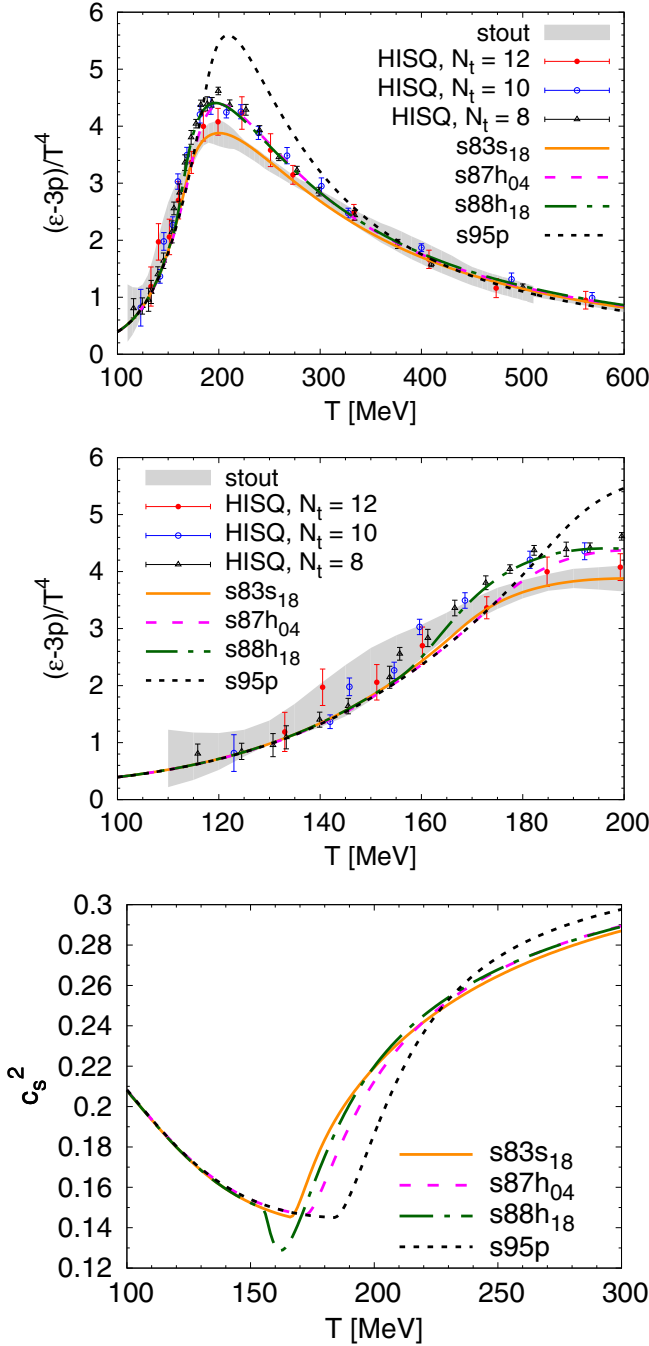


FIG. 1. The trace anomaly (top and middle) and the speed of sound squared (bottom) as functions of temperature in the four parametrizations of the EoS compared to the lattice data obtained using the HISQ [20,21] and stout [22,23] discretization schemes.

differ around the peak, and consequently $s83s_{18}$ differs from $s87h_{04}$ and $s88h_{18}$. The higher peak does not, however, mean a lower speed of sound. As shown in the lowest panel of Fig. 1, the speed of sound in the $s95p$ parametrization is not significantly lower than in the other parametrizations, but the temperature region where it is low is broader than in the other parametrizations. Thus we expect $s95p$ to be effectively softer than the other EoSs. On the other hand, the speed of

sound in $s88h_{18}$ depicts a characteristic dip below the speed of sound in the other parametrizations. This is a consequence of the parametrization of the trace anomaly in that temperature region.

As known, the HRG trace anomaly is below the lattice results [20,22,23] at low temperatures. This difference has been interpreted to indicate the existence of yet unobserved resonance states [34,42]. The need for further states has also been seen in the study of the strangeness baryon correlations on the lattice [43], and confirmed by the S -matrix-based virial expansion [44]. However, we do not include predicted states from any model⁴ in this work, since we do not know how they would decay, but use the states from the PDG summary tables only. Consequently the parametrized trace anomaly is slightly below even the most generous error bars of the lattice results around $T \approx 150$ – 160 MeV temperature, as shown in the middle panel of Fig. 1.

On the other hand, whether we use the PDG 2004 or 2018 particle list causes only a tiny difference in the trace anomaly. The main difference between the $s87h_{04}$ and $s88h_{18}$ parametrizations arises from the connection of the HRG to the lattice parametrization. When parametrizing $s88h_{18}$ we wanted the trace anomaly to reach its lattice values soon above $T_c = 155$ MeV, whereas we allowed $s87h_{04}$ to agree with lattice at larger temperature where the lattice trace anomaly drops below the HRG trace anomaly; for details, see Appendix A. Consequently the $s88h_{18}$ parametrization rises above the HRG values leading to the characteristic dip in the speed of sound (lowest panel in Fig. 1). Note that the $s83s_{18}$ parametrization does not depict a similar dip in the speed of sound, since the lower peak and larger errors of the continuum extrapolated stout action result allow the parametrized trace anomaly to drop below the HRG values immediately.

III. HYDRODYNAMICAL MODEL

We employ a fluid-dynamical model used previously in Refs. [13,24,46–48]. The spacetime evolution is computed numerically in $(2+1)$ dimensions [49], and the longitudinal expansion is accounted for by assuming longitudinal boost invariance. We also neglect here the bulk viscosity and the small net-baryon number. The evolution of the shear-stress tensor $\pi^{\mu\nu}$ is described by the second-order Israel-Stewart formalism [50], with the coefficients of the nonlinear second-order terms obtained by using the 14-moment approximation in the ultrarelativistic limit [51,52]. The shear relaxation time is related to the shear viscosity by $\tau_\pi = 5\eta/(\epsilon + p)$, where ϵ is energy density in the local rest frame, and p is thermodynamic pressure.

Transverse momentum spectra of hadrons are computed by using the Cooper-Frye freeze-out formalism at a constant-temperature surface, followed by all 2- and 3-body decays of unstable hadrons. The chemical freeze-out is encoded into the EoS as described in Ref. [53], and the fluid evolves from chemical to kinetic freeze-out in partial chemical equilibrium (PCE) [54]. The kinetic and chemical freeze-out temperatures

⁴As done in, e.g., Refs. [34,45].

T_{dec} and T_{chem} are left as free parameters to be determined from the experimental data through the Bayesian analysis. The dissipative corrections δf to the momentum distribution at the freeze-out are computed according to the usual 14-moment approximation $\delta f_{\mathbf{k}} \propto f_{0\mathbf{k}} k^\mu k^\nu \pi_{\mu\nu}$, where $f_{0\mathbf{k}}$ is the equilibrium distribution function, and k^μ is the four-momentum of the hadron.

The remaining input to fluid dynamics are the EoS, initial conditions, and shear viscosity. The different options for EoS were discussed in the previous section, and the initial conditions will be detailed in the next section. The temperature dependence of the shear viscosity η/s is parametrized in three parts, controlled by T_H , the lower bound of the temperature range where η/s has its minimum value, $(\eta/s)_{\text{min}}$, and the width of this temperature range, W_{min} :

$$(\eta/s)(T) = \begin{cases} S_{\text{HG}}(T_H - T) + (\eta/s)_{\text{min}}, & T < T_H, \\ (\eta/s)_{\text{min}}, & T_H \leq T \leq T_Q, \\ S_{\text{QGP}}(T - T_Q) + (\eta/s)_{\text{min}}, & T > T_Q, \end{cases} \quad (2)$$

where the additional parameters are the linear slopes below T_H and above $T_Q = T_H + W_{\text{min}}$, denoted by S_{HG} and S_{QGP} , respectively.

We note that bulk viscosity and chemical nonequilibrium are related [55,56]. Even if we ignore the bulk viscosity, some of its effects are accounted for by the fugacities in a chemically frozen fluid: At temperatures below T_{chem} the isotropic pressure is reduced compared to the equilibrium pressure due to the different chemical composition. Thus, introducing the chemical freeze-out changes not only the particle yields with respect to evolution in equilibrium, but similarly to the bulk viscosity, reduces the average transverse momentum of hadrons too. However, this affects the evolution only when temperature is below T_{chem} , and in contrast to the bulk viscosity, there is, e.g., no entropy production associated with the chemical freeze-out and subsequent chemical nonequilibrium [58].

Finally, we emphasize that we solve the spacetime evolution from the hot quark-gluon plasma (QGP) all the way to the kinetic freeze-out as a single continuous fluid-dynamical evolution. This is different from the hybrid models used, e.g., in Refs. [10–12], where the evolution below some switching temperature is solved with a microscopic hadron cascade. The advantage of the fluid-dynamical evolution without a cascade stage is that the transport properties are continuous in the whole temperature range. Note that in the hybrid models the switching from fluid dynamics to hadron cascade introduces an unphysical discontinuity in, e.g., η/s that is $O(1)$ in the cascade [57], but $O(0.1)$ in fluid-dynamical simulations at switching. Another advantage of our approach is that we can freely parametrize the viscosity in the hadronic matter too, and constrain it using the experimental data.

IV. INITIAL CONDITIONS

The initial energy density profiles are determined using the EKRT model [59–61] based on the next-to-leading-order (NLO) perturbative QCD computation of the transverse en-

ergy, and a gluon saturation conjecture. The latter controls the transverse energy production through a local semihard scale $p_{\text{sat}}(T_A T_A, \sqrt{s_{\text{NN}}}, A, K_{\text{sat}})$, where $T_A(x, y)$ is a nuclear thickness function at transverse location (x, y) . The essential free parameters in the EKRT model are the proportionality constant K_{sat} in the saturation condition, and the constant β controlling the exact definition of the minijet transverse energy at NLO [60]. The setup used here is identical to the one used in Refs. [13,24,48], where $\beta = 0.8$, and K_{sat} is left as a free parameter to be determined from the data. We note that K_{sat} is independent of the collision energy $\sqrt{s_{\text{NN}}}$ and nuclear mass number A , so that once K_{sat} is fixed the $\sqrt{s_{\text{NN}}}$ and A dependence of the initial conditions is entirely determined from the QCD dynamics of the EKRT model. With a given p_{sat} the local energy density at the formation time $\tau_p = 1/p_{\text{sat}}$ can be written as

$$\epsilon(x, y, \tau_p) = \frac{K_{\text{sat}}}{\pi} [p_{\text{sat}}(x, y)]^4. \quad (3)$$

This we further evolve to the same proper time $\tau_0 = 1/p_{\text{min}}$, where $p_{\text{min}} = 1$ GeV, at every point in the transverse plane where $p_{\text{sat}} > p_{\text{min}}$ by using $0 + 1$ dimensional Bjorken hydrodynamics with the assumption $\epsilon = 3p$.

In the EKRT model, fluctuations in the product of the nuclear thickness functions, $T_A T_A$, give rise to the event-by-event fluctuations in the energy density through p_{sat} in Eq. (3). Moreover, the centrality dependence of the initial conditions arises from the centrality dependence of $T_A T_A$. A full treatment of the dynamics in heavy-ion collisions would take the event-by-event fluctuations into account by evolving each event separately. However, to make the present study computationally feasible, we omit the evolution of such fluctuations here; instead, for each centrality class, we average a large number of these fluctuating initial states, and compute the fluid-dynamical evolution only for the averaged initial distributions.

The computed energy densities are not linear in K_{sat} or in $T_A T_A$, and different averaging procedures can lead to significantly different event-averaged initial conditions. In the previous event-by-event EKRT studies [13,24,48] a fair agreement was obtained between the data and the computed $\sqrt{s_{\text{NN}}}$, A , and centrality dependencies of the charged hadron multiplicity. To preserve as much as possible of this agreement, we average the initial conditions by averaging the initial entropy distributions: We compute first a large set of initial energy density profiles using the procedure detailed in Ref. [13]. Each of the generated energy density profiles is converted to an entropy density profile by using the EoS which will be used later during the evolution. The entropy density profiles are then averaged, and the average entropy density profile is converted to an average energy density profile using the same EoS.

In the event-by-event framework the centrality classes were determined from the final multiplicity distribution. However, this way of classifying events is not available here, as it would require fluid-dynamical evolution of each of the fluctuating initial conditions. Instead, we predetermine the centrality classes according to the number of wounded nucleons in the sampled Monte Carlo nuclear configurations, which were

used to construct the event-by-event initial conditions. The number of wounded nucleons is computed using the nucleon-nucleon cross section $\sigma_{\text{NN}} = 42, 64, \text{ and } 70 \text{ mb}$ for 200 GeV, 2.76 TeV, and 5.02 TeV collisions, respectively. We note that the nucleon-nucleon cross section does not enter in the computation of the initial conditions, but they are used here only in the centrality classification. In the context of the full event-by-event modeling we have tested that the final results are only weakly sensitive to the precise way of the centrality classification.

V. STATISTICAL ANALYSIS

The eight free parameters of our model, $\{K_{\text{sat}}, (\eta/s)_{\text{min}}, T_{\text{H}}, W_{\text{min}}, S_{\text{HG}}, S_{\text{QGP}}, T_{\text{dec}}, T_{\text{chem}}\}$, were introduced in Secs. III and IV. We want to tune them to achieve the best possible fit to an experimental data set of 90 data points. This set consists of the following observables at (10–20)%, (20–30)%, (30–40)%, (40–50)%, and (50–60)% centrality classes:⁵

- (i) The charged particle multiplicity at midrapidity, $dN_{\text{ch}}/d\eta$, and 4-particle cumulant p_T -averaged elliptic flow, $v_2\{4\}$, in Au + Au collisions at $\sqrt{s_{\text{NN}}} = 200 \text{ GeV}$ (RHIC) [25,27] and Pb + Pb collisions at $\sqrt{s_{\text{NN}}} = 2.76$ [28,30] and $\sqrt{s_{\text{NN}}} = 5.02 \text{ TeV}$ [30,31] (LHC).
- (ii) The multiplicities at midrapidity, dN_i/dy , and average transverse momenta $\langle p_T \rangle_i$, of pions (π^+), kaons (K^+), and protons (p) in Au + Au collisions at RHIC [26] and in Pb + Pb collisions at the lower LHC energy [29].

Let us consider each combination of the free parameters as a point \vec{x} in the 8-dimensional input (parameter) space, the model output $\vec{y}(\vec{x})$ as a corresponding point in the 90-dimensional output space (space of observables), and the experimental data \vec{y}^{exp} as the target point in the space of observables. With these definitions we can formulate the posterior probability distribution $P(\vec{x}|\vec{y}^{\text{exp}})$ of the best-fit parameter values by utilizing Bayes' theorem:

$$P(\vec{x}|\vec{y}^{\text{exp}}) \propto P(\vec{y}^{\text{exp}}|\vec{x})P(\vec{x}), \quad (4)$$

where $P(\vec{x})$ is the prior probability distribution of input parameters and $P(\vec{y}^{\text{exp}}|\vec{x})$ is the likelihood function

$$\begin{aligned} P(\vec{y}^{\text{exp}}|\vec{x}) &= \frac{1}{\sqrt{|2\pi\Sigma|}} \exp \left\{ -\frac{1}{2} [\vec{y}(\vec{x}) - \vec{y}^{\text{exp}}]^T \Sigma^{-1} [\vec{y}(\vec{x}) - \vec{y}^{\text{exp}}] \right\}. \end{aligned} \quad (5)$$

⁵Charged particle multiplicities at RHIC are averages over two adjacent PHENIX centrality classes; for example, at (10–20)% centrality N_{ch} is an average over (10–15)% and (15–20)% classes, (20–30)% is an average over (20–25)% and (25–30)% classes, and so on. This applies also for RHIC identified particle data at (10–20)% centrality.

⁶We consider an average of measured protons and antiprotons as the target value for the proton multiplicity at RHIC.

Here Σ is the covariance matrix representing the uncertainties related to the model-to-data comparison.

As a function with an eight-dimensional domain, the posterior probability distribution $P(\vec{x}|\vec{y}^{\text{exp}})$ is too complicated to evaluate and analyze fully. Instead, we produce samples of it with a parallel tempered Markov chain Monte Carlo [62] based on the EMCEE sampler [63]. An ensemble of random walkers is initialized in the input parameter space based on the prior probability⁷ and each proposed step in parameter space is accepted or rejected based on the change in the value of the likelihood function. At a large number of steps, the distribution of the taken steps is expected to converge to the posterior distribution.

Evaluating the output $\vec{y}(\vec{x})$ of the fluid-dynamical model at every point \vec{x} where the random walker might enter is a computationally impossible task. Therefore we approximate the output using Gaussian process (GP) emulators [64] (see Appendix B). Each GP is able to provide estimates for only one observable, so to keep the number of required emulators manageable, we perform a principal component analysis (PCA) to reduce the dimension of the output space from 90 observables into $k = 6$ most important principal components. Further details about the PCA are described in Appendix C. We utilize the “scikit-learn” Python module [65] and in particular the submodules “sklearn.gaussian_process” and “sklearn.decomposition.PCA” in the model emulation.

Thus, in our likelihood function (5), we replace $\vec{y}(\vec{x})$ with the GP estimate in the principal component space $\vec{z}^{\text{GP}}(\vec{x})$ (likewise \vec{y}^{exp} is transformed to \vec{z}^{exp}), and include the emulator estimation error into the covariance matrix:

$$\Sigma_z = \Sigma_z^{\text{exp}} + \Sigma_z^{\text{GP}}, \quad (6)$$

where Σ_z^{exp} is the (originally diagonal) experimental error matrix transformed to principal component space, and

$$\Sigma_z^{\text{GP}} = \text{diag}[\sigma_{z,1}^{\text{GP}}(\vec{x})^2, \sigma_{z,2}^{\text{GP}}(\vec{x})^2, \dots, \sigma_{z,k}^{\text{GP}}(\vec{x})^2] \quad (7)$$

is the GP emulator covariance matrix obtained from the emulator (see Appendix B).

To work, the GP emulators must be conditioned with a set of training points, $\{\vec{z}(\vec{x}_i)\}$, created by running the fluid-dynamical model with several different parameter combinations $\{\vec{x}_i\}$. For the present investigation, we have produced 170 training points for each EoS, distributed evenly in the input parameter space⁸ using minmax Latin hypercube sampling [66]. The emulation quality was then checked by using the trained emulator to predict the results at 30 additional test points, which were not part of the training data. An example of the results of this confirmation process is shown in Fig. 2 for 2.76 TeV Pb + Pb collisions using the $s95p$ parametrization.

⁷In the present case, the shape of the prior is a uniform hypercube with an additional restriction $T_{\text{dec}} < T_{\text{chem}}$. The prior ranges are shown in Figs. 3 and 4.

⁸The restriction $T_{\text{dec}} < T_{\text{chem}}$ does not apply to the training points.

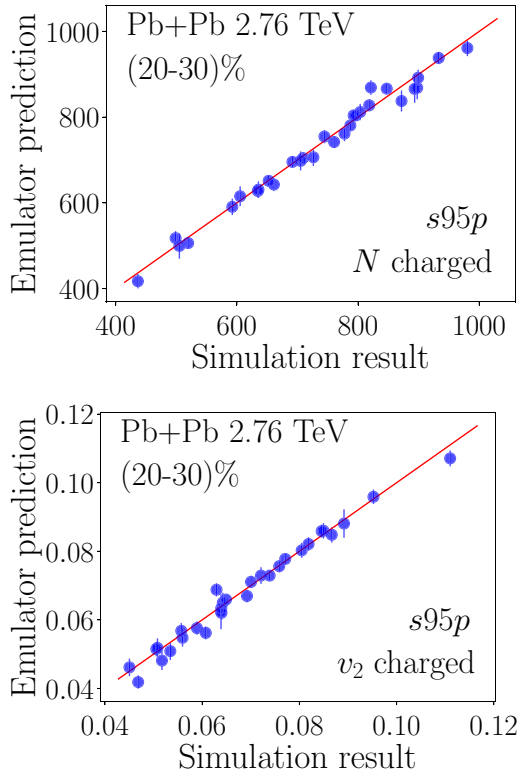


FIG. 2. Illustration of the quality of the Gaussian process emulation for 30 test points for simulations with the $s95p$ EoS. Upper panel: Charged particle multiplicity in (20–30)% most central Pb + Pb collisions at $\sqrt{s_{\text{NN}}} = 2.76$ TeV. Lower panel: Elliptic flow $v_2\{\text{RP}\}$ in (20–30)% most central Pb + Pb collisions at $\sqrt{s_{\text{NN}}} = 2.76$ TeV.

VI. RESULTS

The marginal posterior probability distributions for each parameter are obtained from the full 8-dimensional probability distribution (see Sec. V) by integrating over the other seven parameters. The resulting distributions when using the four investigated EoSs are shown in Figs. 3 and 4. In these figures the range of the x axis illustrates the prior range of values, with the exception of T_{chem} , the range of which depends on the EoS.⁹ The median values of these distributions provide a good approximation for the most probable values, and are listed both in the legends of the figures and in Table I. The 90% credible intervals—i.e., the range which covers 90% of the distribution around the median—are shown as errors in Table I. Two-dimensional projections of the probability distribution depicting correlations between parameter pairs are shown in Appendix D.

⁹For $s83s_{18}$, $s87h_{04}$, and $s88h_{18}$, the prior range is $120 \text{ MeV} < T_{\text{chem}} < T_0$, where T_0 is the temperature where the parametrization deviates from the HRG (see Appendix A). For $s95p$ the range is $120 < T_{\text{chem}}/\text{MeV} < 180$.

A. Nuisance parameters

The analysis involves three parameters which are not directly related to the transport properties of produced matter: K_{sat} , T_{dec} , and T_{chem} . The probability distributions for these three “nuisance” parameters, shown in Fig. 3, are nicely peaked, and the parameters have well defined constraints. For the chemical freeze-out temperature, the median is $T_{\text{chem}} = 153\text{--}155$ MeV, which is compatible with the values obtained using the statistical hadronization model [67]. Note that the difference in the median is not due to the resonance content of the EoS, but due to a complicated interplay of the softness of the EoS, shear, and build-up of the flow. Nevertheless, the particle ratios are the dominant factor in constraining T_{chem} .

For K_{sat} and T_{dec} , we see a common trend where $s95p$ gives a distribution which peaks at the lowest value of the four EoSs, followed by $s87h_{04}$, and the highest peak values are shared by $s83s_{18}$ and $s88h_{18}$ with almost identical distributions. The obtained values for the EKRT normalization parameter, $K_{\text{sat}} \approx 0.5$, are compatible with the values found previously [13], and the small differences between different EoS parametrizations result from slightly different entropy production during the evolution. Differences seen in the kinetic freeze-out temperature $T_{\text{dec}} = 126\text{--}132$ MeV are also small, and seem to follow the conventional rule of thumb: a softer EoS requires a lower freeze-out temperature to create hard enough proton p_T distributions. On the other hand, differences in the median values of all these three parameters are smaller than the credibility intervals, and thus not statistically meaningful.

B. $(\eta/s)(T)$

At first sight $(\eta/s)_{\text{min}}$ depicts the behavior described in Refs. [16,19]: the favored value is lower for $s95p$ than for the newer parametrizations (see Fig. 3 and Table I). However, the effect is noticeably smaller than seen in those studies—only $\approx 10\%\text{--}20\%$ —and well within the 90% credible intervals ($\pm \approx 30\%$) of the analysis. The comparison of η/s for different EoSs is further complicated by the large number of parameters controlling the temperature dependence of η/s . The probability distributions of parameters T_{H} , W_{min} , S_{HG} , and S_{QGP} , shown in Fig. 4, are very broad extending to the whole prior range in most cases, and thus do not possess any clearly favored values. However, the wide posterior distributions of the $(\eta/s)(T)$ parameters are partly caused by the inherent ambiguity in the chosen parametrization: for a given temperature T , multiple parameter combinations can generate similar values of η/s . For example, at low temperatures $(\eta/s)(T)$ is mostly determined by S_{HG} and T_{H} , but it is better constrained than either of these parameters. The reason is that S_{HG} and T_{H} are not independent, but slightly anticorrelated: the correlations between the pairs of parameters are shown in Appendix D. Thus it is more illustrative to construct the probability distribution for η/s values with respect to temperature, and plot the median and credibility intervals of this distribution as shown in Figs. 5 and 6.

In the upper panel of Fig. 5 we show the median of $(\eta/s)(T)$ for each EoS parametrization, and the union and intersection of the 90% credible intervals of all four distributions. The union of the credibility intervals provides insight

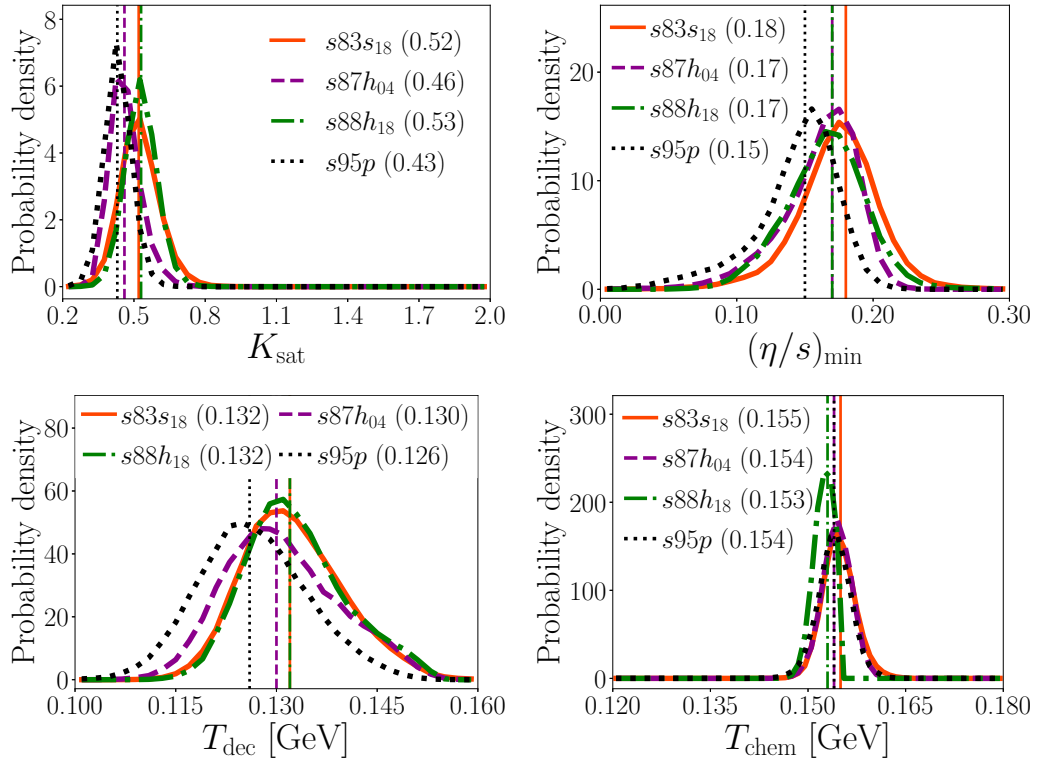


FIG. 3. Comparison of K_{sat} , $(\eta/s)_{\text{min}}$, T_{dec} , and T_{chem} marginal posterior probability distributions for the four investigated EoSs. Vertical lines (bracketed numbers in legend) indicate median values for the distributions. With the exception of T_{chem} , the range of the x axis in the plots is the original prior range.

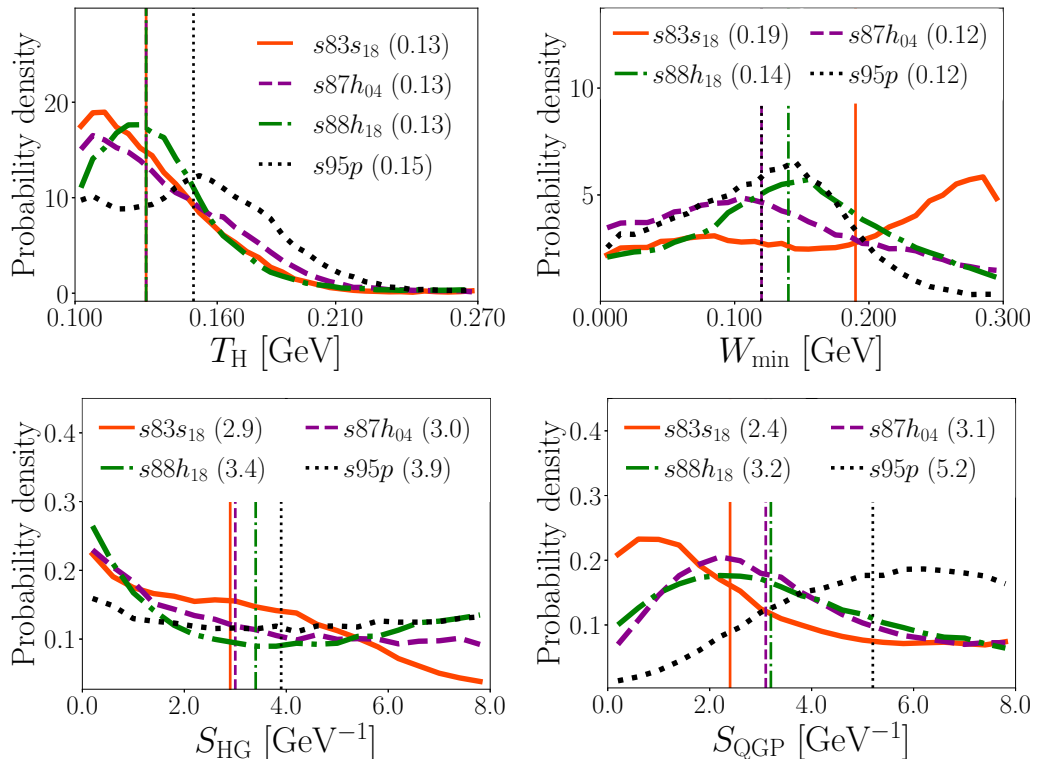


FIG. 4. As Fig. 3, but showing the marginal posterior probability distributions of T_{H} , W_{min} , S_{HG} , and S_{QGP} .

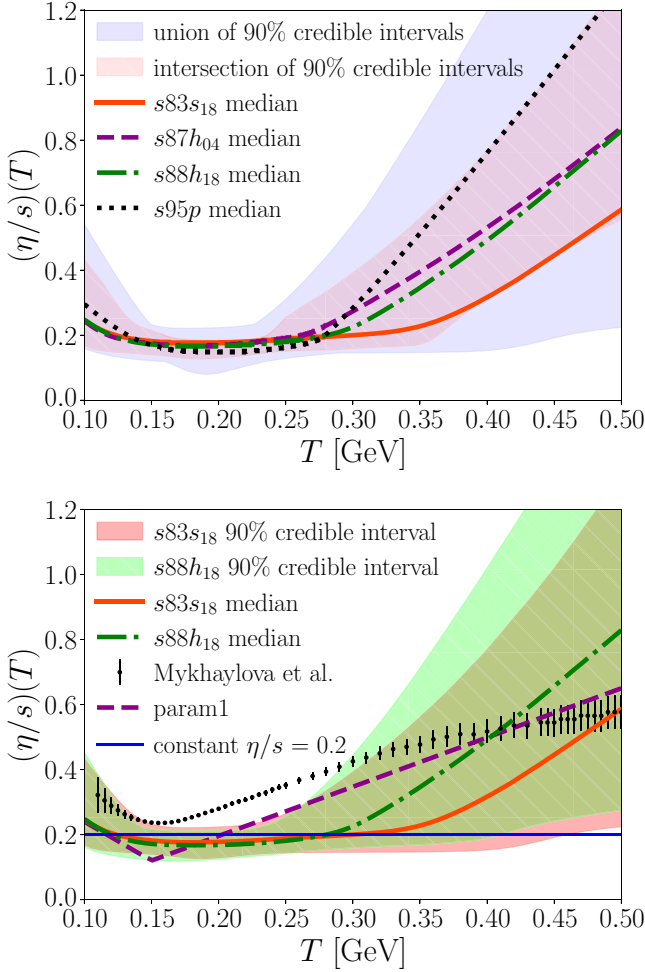


FIG. 5. Temperature dependence of η/s . Upper panel: Median of η/s with respect to T for each EoS with the union and intersection of the 90% credible intervals of the distributions. Lower panel: Median of η/s with respect to T for the $s83s_{18}$ and $s88h_{18}$ parametrizations with corresponding credibility intervals compared with two results from Ref. [13] ($\eta/s = 0.2$ and param1) and a recent quasiparticle model prediction by Mykhaylova *et al.* [68].

on the total uncertainty in the analysis including the uncertainty from the EoS parametrization, whereas the difference between the union and intersection illustrates how much of the uncertainty comes from the EoS parametrizations. To emphasize the result using state-of-the-art EoSs, the lower panel of Fig. 5 depicts the median and credibility intervals for the parametrizations $s83s_{18}$ and $s88h_{18}$ only. In the same panel two older results from Ref. [13], and a recent theoretical prediction from Ref. [68], are shown as well. To make it possible to distinguish the credibility intervals for each EoS separately, η/s for each EoS at various temperatures with associated uncertainties is shown in Fig. 6.

We obtain well constrained η/s in a temperature range $150 \lesssim T/\text{MeV} \lesssim 220$, where the median values of η/s are practically constant for all contemporary EoSs, and $s95p$ leads to modest temperature dependence well within the credibility intervals. Within this temperature range η/s is constrained between 0.08 and 0.23 by the 90% credible intervals. In

TABLE I. Estimated parameter values (medians) and uncertainties (90% credible intervals) from the posterior distributions.

Parameter	$s83s_{18}$	$s87h_{04}$	$s88h_{18}$	$s95p$
K_{sat}	$0.52^{+0.15}_{-0.12}$	$0.46^{+0.12}_{-0.09}$	$0.53^{+0.11}_{-0.10}$	$0.43^{+0.10}_{-0.09}$
$(\eta/s)_{\text{min}}$	$0.18^{+0.04}_{-0.06}$	$0.17^{+0.03}_{-0.07}$	$0.17^{+0.04}_{-0.06}$	$0.15^{+0.03}_{-0.07}$
T_H (GeV)	$0.13^{+0.05}_{-0.03}$	$0.13^{+0.06}_{-0.03}$	$0.13^{+0.06}_{-0.03}$	$0.15^{+0.06}_{-0.04}$
W_{min} (GeV)	$0.19^{+0.10}_{-0.17}$	$0.12^{+0.15}_{-0.11}$	$0.14^{+0.13}_{-0.12}$	$0.12^{+0.10}_{-0.10}$
S_{HG} (GeV^{-1})	$2.9^{+4.0}_{-2.7}$	$3.0^{+4.5}_{-2.8}$	$3.4^{+4.2}_{-3.2}$	$3.9^{+3.7}_{-3.6}$
S_{QGP} (GeV^{-1})	$2.4^{+4.9}_{-2.1}$	$3.1^{+4.2}_{-2.5}$	$3.2^{+4.1}_{-2.7}$	$5.2^{+2.5}_{-3.5}$
T_{dec} (MeV)	132^{+14}_{-11}	130^{+16}_{-12}	132^{+15}_{-10}	126^{+15}_{-12}
T_{chem} (MeV)	155^{+4}_{-4}	154^{+4}_{-3}	153^{+2}_{-3}	154^{+4}_{-4}

particular, for the state-of-the-art EoSs ($s83s_{18}$ and $s88h_{18}$), we obtain even tighter limits $0.12 < \eta/s < 0.23$ within this range, and the well constrained region extends to slightly higher temperature. For further details see Fig. 5 and Table II. Interestingly η/s at 130 MeV (or at 150 MeV in the case of $s95p$) temperature differs from the favored value (median) of the $(\eta/s)_{\text{min}}$ parameter (compare Tables I and II), even if the

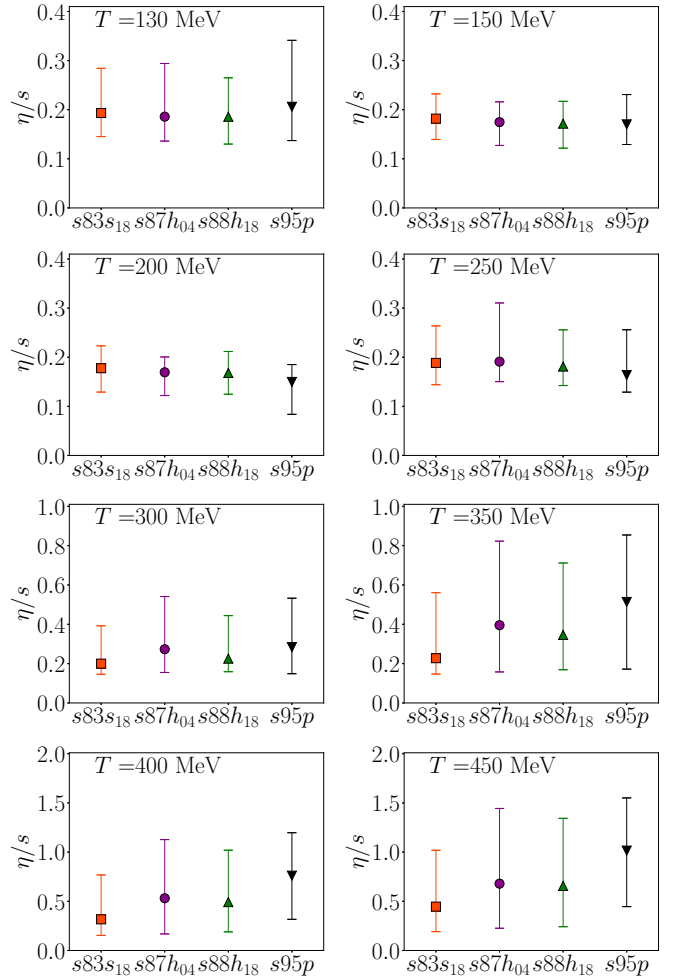


FIG. 6. Median values (filled markers) and 90% credible intervals (error bars) for η/s at temperatures $T = 130, 150, 200, 250, 300, 350, 400$, and 450 MeV.

TABLE II. Median values of η/s at various temperatures with associated uncertainties (90% credible intervals) from the posterior distributions. Values rounded to two significant figures.

T (MeV)	$s83s_{18}$	$s87h_{04}$	$s88h_{18}$	$s95p$
130	$0.19^{+0.09}_{-0.04}$	$0.19^{+0.10}_{-0.05}$	$0.19^{+0.07}_{-0.06}$	$0.21^{+0.13}_{-0.07}$
150	$0.18^{+0.05}_{-0.04}$	$0.17^{+0.05}_{-0.04}$	$0.17^{+0.05}_{-0.05}$	$0.17^{+0.06}_{-0.04}$
200	$0.18^{+0.04}_{-0.05}$	$0.17^{+0.03}_{-0.05}$	$0.17^{+0.04}_{-0.05}$	$0.15^{+0.04}_{-0.07}$
250	$0.19^{+0.07}_{-0.05}$	$0.19^{+0.12}_{-0.04}$	$0.18^{+0.08}_{-0.04}$	$0.16^{+0.10}_{-0.03}$
300	$0.20^{+0.19}_{-0.05}$	$0.27^{+0.27}_{-0.11}$	$0.23^{+0.21}_{-0.07}$	$0.28^{+0.25}_{-0.13}$
350	$0.23^{+0.33}_{-0.08}$	$0.40^{+0.42}_{-0.24}$	$0.35^{+0.36}_{-0.18}$	$0.51^{+0.34}_{-0.34}$

avored value of the T_H parameter is 130 MeV (or 150 MeV) (see Fig. 4 and Table I). This seemingly counterintuitive behavior is due to the fat tails of T_H distributions extending to larger temperatures, and thus broadening the region where S_{HG} affects the η/s values. Consequently we see the lowest η/s values at $T \approx 200$ MeV temperature (Fig. 5 and Table II), where the effect of the lower $(\eta/s)_{\min}$ value of the $s95p$ parametrization is also visible.

It is not surprising that we get the best constraints on η/s in the temperature range $150 \lesssim T/\text{MeV} \lesssim 220$. As was shown in Ref. [46], the temperature range where v_2 is most sensitive to the shear viscosity is only slightly broader than this, and higher order anisotropies are sensitive to shear at even narrower temperature ranges.¹⁰

Even if the uncertainties remain large, we can see qualitative differences in the high-temperature behavior of η/s , where $s95p$ seems to favor earlier and more rapid rise of η/s with increasing temperature (Figs. 5 and 6), a difference which is visible in the S_{QGP} parameter as well (Fig. 4).

Considering earlier results in the literature, this is intriguing. Alba *et al.* [16] used an EoS based on contemporary stout action data called PDG16+/WB2 + 1, and observed that the reproduction of the LHC data ($\sqrt{s_{NN}} = 5.02$ TeV) required larger constant η/s for this EoS than for $s95p$. On the other hand, they were able to use the same value of constant η/s for both EoSs to reproduce the RHIC data. They interpreted this to mean that at large temperatures $s95p$ would necessitate lower values of η/s , but we see an opposite behavior. In a similar fashion we see a difference between the high-temperature behavior obtained using the HISQ ($s88h_{18}$ and $s87h_{04}$) and stout-action-based EoSs ($s83s_{18}$), but the differences are much smaller than the credibility intervals, and thus cannot be considered meaningful.

At temperatures below 150 MeV we again see expanding credibility intervals, and a tendency of η/s to increase with decreasing temperature, but hardly any sensitivity to the EoS. Anisotropies measured at RHIC energy are sensitive to the shear viscosity in the hadronic phase [46,47], and since Schenke *et al.* in Ref. [19] saw sensitivity to the EoS using RHIC data only, we would have expected some sensitivity to

the EoS at low temperatures. The difference may arise from the bulk viscosity which depended on the EoS as well in Ref. [19], or from a different EoS in the hadronic phase. As mentioned, our EoSs are based on known resonance states, whereas the EoSs in Refs. [16,19] follow the lattice results closely. A better fit to lattice results can be obtained by including predicted but unobserved resonance states in the HRG. We plan to study how the inclusion of these states might affect the results, once we have concocted a plausible scheme for their decays, so that we can evaluate their contribution to the EoS after chemical freeze-out in a consistent manner.

Furthermore, unlike in Ref. [19] where a hadron cascade was used to describe the evolution in the hadronic phase, in our approach the change in the EoS can also be partly compensated by a change in the freeze-out temperature instead of shear viscosity. As shown in Appendix D, there is indeed an anticorrelation between T_{dec} and $(\eta/s)_{\min}$. Therefore forcing the system to freeze out at the same temperature, independently of the EoS, would increase the difference in $(\eta/s)_{\min}$. However, the anticorrelation is rather weak ≈ -0.4 (-0.2) for $s88h_{18}$ ($s95p$), and thus requiring EoS independent T_{dec} would not change $(\eta/s)_{\min}$ very much.

Our result of a very slowly rising η/s with decreasing temperature in the hadronic phase (i.e., below $T \approx 150$ MeV) may look inconsistent with microscopic calculations predicting relatively large $\eta/s \sim 1$ in the hadronic phase [57,69,70]. However, our result is for a chemically frozen HRG, while the microscopic calculations usually give η/s in chemical equilibrium. In the transport model study of Ref. [71], it was shown that nonunit pion and kaon fugacities, i.e., chemical nonequilibrium, can significantly reduce η/s in hadron gas. Since, as a first-order approximation, η depends only weakly on the chemical nonequilibrium [72], the main effect is due to s : At a given temperature the entropy density s_{PCE} in a chemically frozen HRG can be significantly larger than the entropy density in chemical equilibrium s_{CE} , and as a consequence $(\eta/s)_{\text{PCE}}$ can be much smaller than $(\eta/s)_{\text{CE}}$. We may thus obtain an approximation for the η/s in a chemically equilibrated system as $(\eta/s)_{\text{CE}} = (\eta/s)_{\text{PCE}}(s_{\text{PCE}}/s_{\text{CE}})$ [13]. In our case, where $T_{\text{chem}} = 154$ MeV, the ratio of entropies in a chemically frozen to a chemically equilibrated system is ≈ 3.5 at $T = 100$ MeV (≈ 1.8 at $T = 130$ MeV), which is sufficient to bring our results to the level described in Ref. [70].

In Fig. 5 we also made a comparison to the earlier results of Ref. [13] and the recent quasiparticle model prediction from Ref. [68]. As expected, the earlier results from Ref. [13,24] are not far from the present analysis, and param1 is practically within the 90% credible interval in the whole temperature range. On the other hand, constant $\eta/s = 0.2$ is below the $s95p$ limits at high temperatures, but as discussed, the overall sensitivity to η/s at high temperatures is low. Interestingly the prediction of the quasiparticle model of Ref. [68] comes very close to our values for η/s around T_c , although the region where η/s is low is narrower than what we found here. This is intriguing, since the quasiparticle model was tuned to reproduce the stout action EoS, i.e., our EoS $s83s_{18}$, which in our analysis leads to the broadest region where η/s is almost constant.

¹⁰Note that the studies in Ref. [46] were carried out using the $s95p$ EoS. We have not checked how sensitive those results are to the EoS parametrization.

The small value of η/s and its weak temperature dependence in the temperature range $150 \lesssim T/\text{MeV} \lesssim 220$ may indicate that the QGP is strongly coupled not only in the immediate vicinity of T_c but in a broader temperature region. This was first proposed in Ref. [73], and agrees with the lattice QCD calculations that indicate the presence of hadron-like resonances in the QGP in a similar or slightly broader temperature interval [74–77]. The strongly coupled nature of the QGP can also be seen in the large value of the coupling constant defined in terms of the free energy of static quark-antiquark pairs [78]. In any case, our result for $(\eta/s)(T)$ is compatible with the lattice QCD calculations, which indicate that the weakly coupled QGP picture may be applicable only for $T > 350$ MeV [78–81].

C. The effect of the parametric form

When we use the state-of-the-art EoSs ($s88h_{18}$ and $s83s_{18}$), our result for the minimum value of η/s is higher than the result obtained in an earlier Bayesian analysis of Ref. [10]: $0.12 < \eta/s < 0.23$ vs $\eta/s = 0.07^{+0.05}_{-0.04}$. While the equations of state in both analyses are based on the latest lattice results, an important difference is that Ref. [10] assumed the minimum of η/s to occur at fixed $T = 154$ MeV temperature, and η/s to rise linearly above that temperature. Moreover, below $T = 154$ MeV they used a hadron cascade to model the evolution, and the transport properties of the hadronic phase were thus fixed.

To explore how much the results depend on the form of the $(\eta/s)(T)$ parametrization, we mimic the parametrization used in Ref. [10] by constraining the plateau in our parametrization to be very small ($0 < W_{\min}/\text{MeV} < 2$), and the minimum to appear close to T_c ($150 < T_H/\text{MeV} < 160$). The resulting temperature dependence of η/s for the $s88h_{18}$ and $s95p$ parametrizations is shown in Fig. 7, and compared to our full result (the behavior of the $s87h_{04}$ and $s83s_{18}$ parametrizations is similar to $s88h_{18}$).

The change in parametrization reduces the minimum value of η/s to $0.12^{+0.03}_{-0.03}$ for $s88h_{18}$, which is closest to the EoS used in Ref. [10]. The credibility interval now overlaps with the result from the earlier analysis [10], and the results are thus consistent. The remaining difference may result from the bulk viscosity, event-by-event-fluctuations, differences in the EoS parametrization scheme, or the transport description of the hadronic phase. As mentioned earlier, switching to the hadron cascade creates a discontinuity in $(\eta/s)(T)$. Enforcing a similar discontinuity in the $(\eta/s)(T)$ parametrization might bring closer the minimum values of η/s obtained using hybrid models and continuous fluid dynamics. For $s95p$ the minimum value drops to $0.06^{+0.04}_{-0.04}$, but since the $s95p$ parametrization is based on the older lattice data, comparing this value against Ref. [10] is not straightforward. However, due to the significant overlap of the credibility intervals, we consider both results consistent with Ref. [10], demonstrating the weak sensitivity of the extracted η/s to the EoS used in the calculations.

Another interesting change is seen in the high-temperature behavior. In the full analysis the $s95p$ parametrization leads to the largest η/s at large temperatures, but the restricted parametrization causes $s95p$ to favor the lowest η/s at large

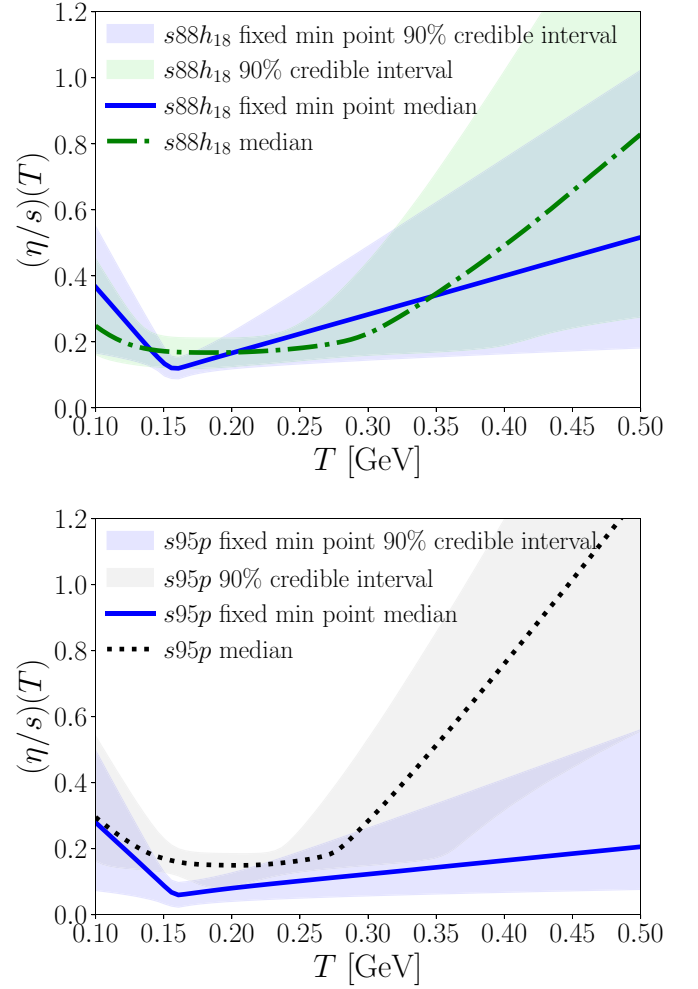


FIG. 7. Temperature dependence of η/s for the $s88h_{18}$ (top) and $s95p$ (bottom) EoSs using the full parametrization and a parametrization constrained to have a minimum at a fixed point in temperature (“fixed min point”).

temperatures. As seen previously, $s95p$ favors the lowest η/s at $200 < T/\text{MeV} < 250$ temperatures (see Fig. 6 and Table II), which in the restricted parametrization dictates the behavior at much higher temperatures as well.

Nevertheless, even if the results depend on the form of the parametrization, the credibility intervals overlap and the results are consistent. The only deviation from this rule is for the $s95p$ parametrization around $T = 160$ MeV temperature where the difference is statistically significant (see Fig. 7). We have also checked that when we use the favored parameter values, the typical differences in the fit to the data due to different parametric forms are only $\approx(1-3)\%$.

Similarly, we can mimic temperature-independent η/s by constraining the priors of the S_{HG} and S_{QGP} parameters close to zero. We have checked that such a choice does not increase the sensitivity of η/s to the EoS parametrization, and that the median values of the constant $\eta/s = (\eta/s)_{\min}$ were only $\approx 10\%$ larger than the median values for η/s at $T = 200$ MeV for the full parametrization—again, a sign of v_2 being

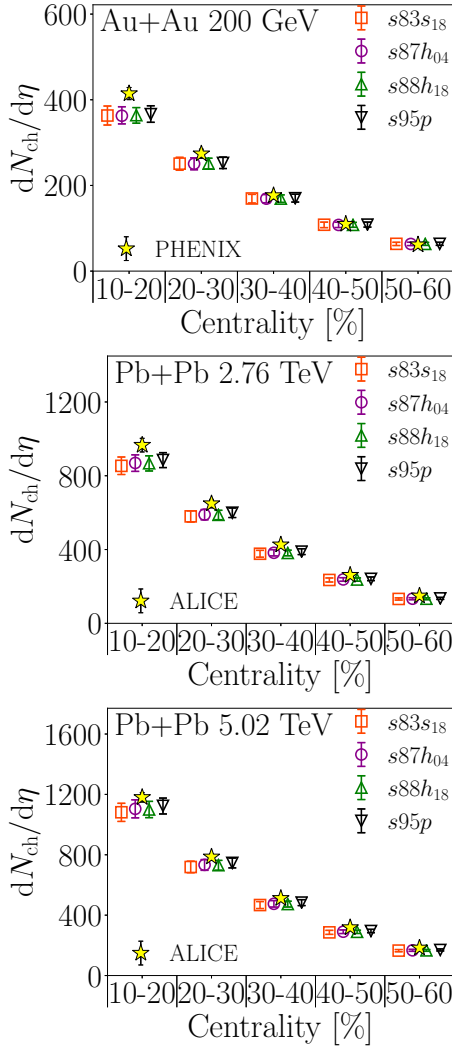


FIG. 8. Charged particle multiplicity at various centralities using 1000 samples from the posterior distribution of each EoS. Marker centers indicate median values, and error bars 90% credible intervals. Top panel: Au + Au at $\sqrt{s_{NN}} = 200$ GeV compared to PHENIX data [25]. Middle panel: Pb + Pb at $\sqrt{s_{NN}} = 2.76$ TeV compared to ALICE data [28]. Bottom panel: Pb + Pb at $\sqrt{s_{NN}} = 5.02$ TeV compared to ALICE data [31].

most sensitive to shear viscosity in the $150 \lesssim T/\text{MeV} \lesssim 220$ temperature range [46].

Thus, in the Bayesian analysis the parametric form of η/s does affect the results, and is therefore a kind of prior whose effects are difficult to quantify. On the other hand, the credibility intervals overlap in all the cases, which emphasizes their importance: The “true” value could be anywhere within the credibility interval, and there is still a 10% chance that it is outside of it.

D. Comparison with the data

Finally, as an overall quality check, we show how well the favored parameter combinations reproduce the experimental data. This is done by drawing 1000 samples from each pos-

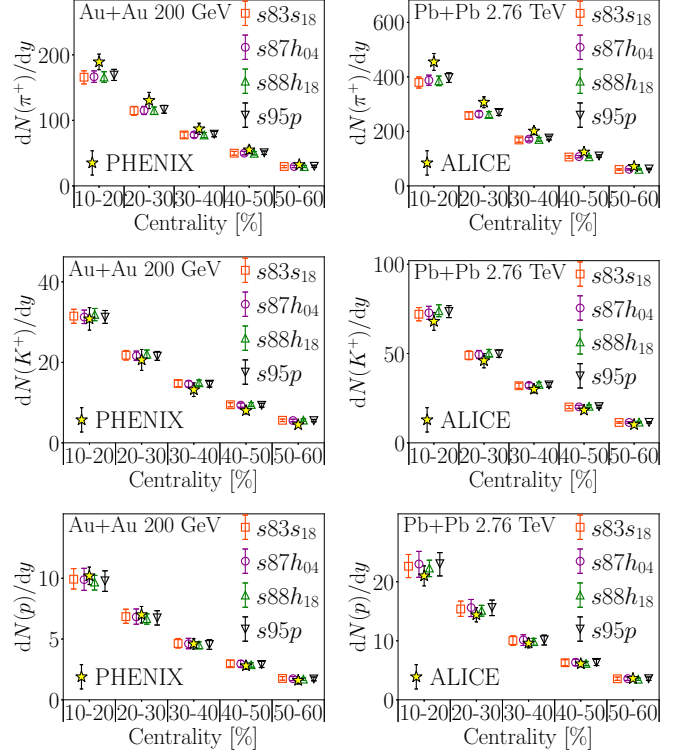


FIG. 9. Pion (upper panels), kaon (middle panels), and proton (lower panels) multiplicities at various centralities using 1000 samples from the posterior distribution of each EoS. Marker centers indicate median values, and error bars 90% credible intervals. Left panels: Au + Au at $\sqrt{s_{NN}} = 200$ GeV compared to PHENIX data [26]. Right panels: Pb + Pb at $\sqrt{s_{NN}} = 2.76$ TeV compared to ALICE data [29].

terior distribution and using the Gaussian process emulator to predict the simulation output for these values. The results for charged and identified particle multiplicities, identified particle $\langle p_T \rangle$, and the elliptic flow $v_2\{4\}$ are shown in Figs. 8–11, respectively.

The overall agreement with the data is quite good for all observables, and the analysis is able to find equally good data fits for all four EoS. As normal for thermal models, the charged particle multiplicities tend to be underestimated due to the tension between pion multiplicity on one hand, and kaon and proton multiplicities on the other hand. As the analysis makes a compromise between too few pions and too many kaons and protons, the overall charged particle multiplicity (which is dominated by pions) will remain below the data. Also the mean transverse momentum of pions is slightly too large, which may prove difficult to alleviate without the introduction of bulk viscosity [4] and/or improved treatment of resonances during the hadronic phase [82].

VII. SUMMARY

In this work, we have introduced three new parametrizations of the equation of state based on the contemporary lattice data:

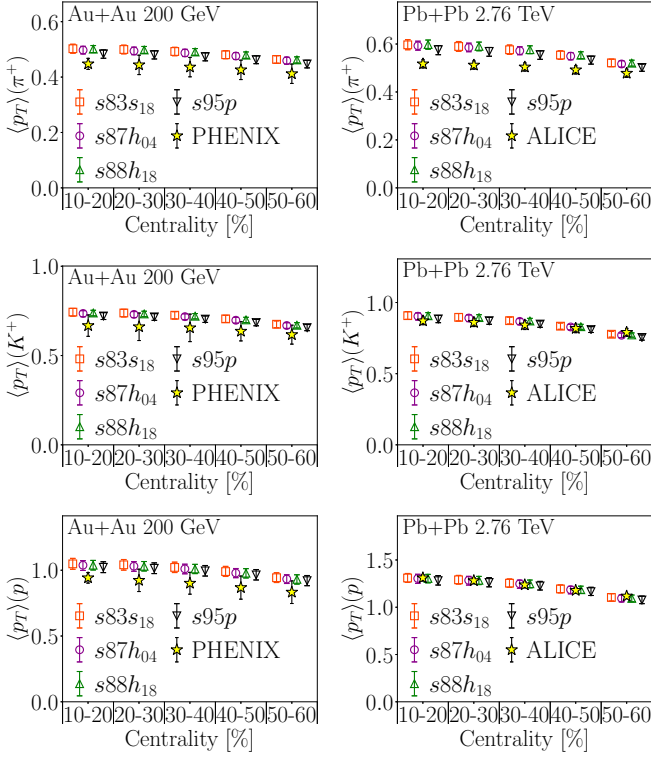


FIG. 10. As Fig. 9 but for the mean transverse momentum.

- (i) $s87h_{04}$ connects the HRG based on the PDG 2004 particle list to parametrized lattice data obtained using the HISQ discretization scheme.
- (ii) $s88h_{18}$ is based on the HRG containing all strange and nonstrange hadrons and resonances in the PDG 2018 summary tables, and the same HISQ lattice data as $s87h_{04}$.
- (iii) $s83s_{18}$ is constructed using the PDG 2018 resonances, and the continuum extrapolated lattice data obtained using the stout discretization.

We used these new parametrizations and the older $s95p$ parametrization to examine how sensitive the shear viscosity over entropy density ratio η/s is to the equation of state. We

assumed a piecewise linear parametrization for $(\eta/s)(T)$, and determined the probability distributions of the best-fit parameter values within the EKRT framework using a Bayesian statistics approach.

Using charged and identified particle multiplicities, identified particle mean transverse momenta, and elliptic flow at three different collision energies as calibration data, we were able to constrain the value of η/s to be between 0.08 and 0.23 with 90% credibility in the temperature range $150 \lesssim T/\text{MeV} \lesssim 220$ when all EoS parametrizations are taken into account. When we constrain the EoSs to the most contemporary parametrizations $s83s_{18}$ and $s88h_{18}$, we obtain $0.12 < \eta/s < 0.23$ in the above mentioned temperature range. As the differences between the EoSs are well covered by the 90% credible intervals, the earlier results obtained using the $s95p$ parametrization remain valid. The weak sensitivity to the EoS is consistent with the old ideal fluid results for flow and EoS: Based on flow alone, it is difficult to distinguish an EoS with a smooth crossover from an EoS without phase transition [83]. Thus when the differences between EoSs are just details in the crossover, the differences in flow, which should be compensated by different shear viscosity, are small, and consequently differences in the extracted η/s are small.

The overall agreement with the data is quite good, and similar to Refs. [13,24], where event-by-event fluctuations were included to the framework of EKRT initial conditions and fluid dynamics, albeit without the Bayesian analysis. The good agreement achieved here is partly due to the EKRT initial conditions. In particular the centrality and $\sqrt{s_{\text{NN}}}$ dependence of hadron multiplicities follow mainly from the QCD dynamics of the EKRT model. A noticeable difference to the earlier event-by-event analysis is that here we used identified hadron multiplicities as constraint, which led to the chemical freeze-out temperature $T_{\text{chem}} \approx 154$ MeV, and a slight overshoot of the pion average p_T compared to the data. In the earlier analysis $T_{\text{chem}} \approx 175$ MeV was used to reproduce the average p_T data, which in turn led to too large proton multiplicity. It is possible to solve this tension by introducing bulk viscosity [4], but that is left for a future work. We emphasize that compared to the (in principle) more detailed hydro + cascade models our hydro + partial chemical equilibrium ap-

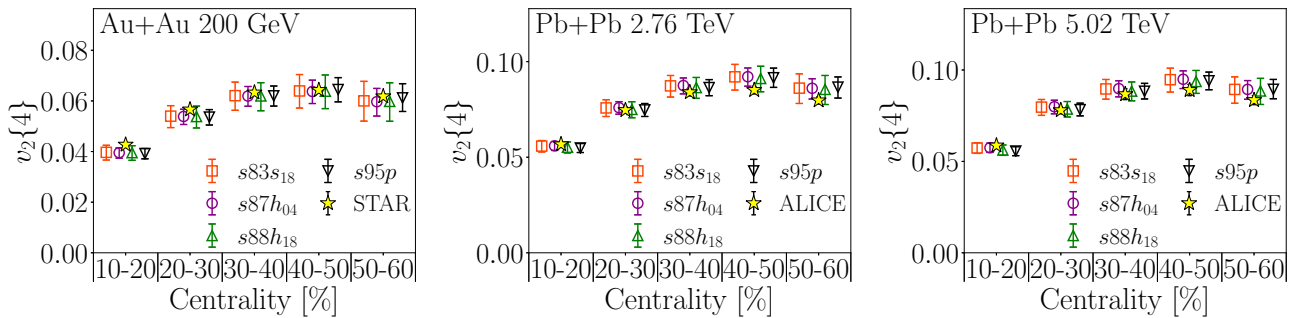

 FIG. 11. Charged particle elliptic flow $v_2\{4\}$ at various centralities using 1000 samples from the posterior distribution of each EoS. Marker centers indicate median values, and error bars 90% credible intervals. Left panel: Au + Au at $\sqrt{s_{\text{NN}}} = 200$ GeV compared to STAR data [27]. Middle panel: Pb + Pb at $\sqrt{s_{\text{NN}}} = 2.76$ TeV compared to ALICE data [30]. Right panel: Pb + Pb at $\sqrt{s_{\text{NN}}} = 5.02$ TeV compared to ALICE data [30].

TABLE III. The values of parameters for different fits of the trace anomaly.

	d_0	d_1 (GeV ²)	d_2 (GeV ⁴)	d_3 (GeV ^{n₃})	d_4 (GeV ^{n₄})	d_5 (GeV ^{n₅})	n_3	n_4	n_5	T_0 (MeV)
$s83s_{18}$	5.688×10^{-3}	0.3104	-6.217×10^{-3}	-6.680×10^{-32}	1.071×10^{-32}		41	42		166
$s87h_{04}$	5.669×10^{-2}	0.2974	-4.184×10^{-3}	-5.146×10^{-8}	1.420×10^{-33}		10	42		172
$s88h_{18}$	4.509×10^{-2}	0.3082	-5.136×10^{-3}	-1.150×10^{-10}	2.076×10^{-32}	-3.021×10^{-33}	13	41	42	155
$s95p$		0.2660	2.403×10^{-3}	-2.809×10^{-7}	6.073×10^{-23}		10	30		183.8

proach has two major advantages: It allows us to parametrize $(\eta/s)(T)$ so that it is continuous in the whole temperature range, and at the same time it gives us a possibility to constrain the viscosity also in the hadronic phase.

Inclusion of event-by-event fluctuations to the analysis would provide access to several new flow observables such as higher flow harmonics v_n , and flow correlations, which may give tighter constraints in broader temperature interval on $(\eta/s)(T)$. However, within the current uncertainties of the fitting procedure, we cannot exclude the possibility that the effect of the EoS remains negligible even when η/s at $T > 220$ MeV becomes better under control.

Since the sensitivity of flow to shear viscosity at high temperatures is low, observables based on high- p_T particles may be useful to constrain not only the pre-equilibrium dynamics [84–86] but also the properties of the fluid when it is hottest.

ACKNOWLEDGMENTS

We thank V. Mykhaylova for sharing her quasiparticle results with us. J.A. and P.H. were supported by the European Research Council, Grant No. ERC-2016-COG:725741; P.H. was also supported by National Science Center, Poland, under grant Polonez DEC-2015/19/P/ST2/03333 receiving funding from the European Union’s Horizon 2020 research and innovation program under Marie Skłodowska-Curie Grant Agreement No. 665778; K.J.E. and H.N. were supported by the Academy of Finland, Project No. 297058; and P.P. was supported by the US Department of Energy under Contract No. DE-SC0012704. We acknowledge the CSC–IT Center for Science in Espoo, Finland, for the allocation of computational resources.

APPENDIX A: EOS PARAMETRIZATION

At high temperature the trace anomaly can be well parametrized by the inverse polynomial form. Therefore we will use the following ansatz for the high-temperature region:

$$\frac{\epsilon - 3p}{T^4} = d_0 + \frac{d_1}{T^2} + \frac{d_2}{T^4} + \frac{d_3}{T^{n_3}} + \frac{d_4}{T^{n_4}} + \frac{d_5}{T^{n_5}}. \quad (\text{A1})$$

This form does not have the right asymptotic behavior in the high-temperature region, where we expect $(\epsilon - 3p)/T^4 \sim g^4(T) \sim 1/\ln^2(T/\Lambda_{\text{QCD}})$, but it works well in the temperature range of interest. Furthermore, it is flexible enough to match the HRG result in the low-temperature region. We match this ansatz to the HRG model at temperature T_0 by requiring that the trace anomaly and its first and second derivatives are continuous. This requirement provides constraints for three

parameters, d_0, d_1 , and d_2 , and leaves the remaining seven, $d_3, d_4, d_5, n_3, n_4, n_5$, and T_0 , to be fixed by minimizing a χ^2 fit to the data. Fitting the powers n_3 – n_5 would be a highly nonlinear problem, but we simplify the problem by requiring that the powers be integers, and using brute force: We make a fit with all the integer values $5 \leq n_3 \leq 40$, $n_3 < n_4 \leq 41$, and $n_4 < n_5 \leq 42$, and choose the values n_3, n_4 , and n_5 which lead to the smallest χ^2 . When the powers and T_0 are kept fixed, minimizing χ^2 requires only a simple matrix inversion. Thus to fix T_0 we are able to cast χ^2 as a function of only a single parameter, T_0 . We require that $155 \leq T_0/\text{MeV} \leq 190$, and search for the value of T_0 which minimizes χ^2 .

To obtain the continuum limit in the lattice calculations of the trace anomaly, one has to perform interpolation in the temperature, and then perform continuum extrapolations (see, e.g., [23]). This procedure can introduce additional uncertainties when providing parametrization of the lattice results. As mentioned in the main text, the lattice spacing (N_t) dependence of the lattice results is small in the case of the HISQ discretization scheme for $N_t \geq 8$. In fact, for $T > 230$ MeV and $T < 170$ MeV there is no statistically significant N_t dependence, so in these temperature ranges we can use the HISQ lattice results with $N_t = 8, 10$, and 12 . In the peak region, $170 < T/\text{MeV} < 230$, the $N_t = 8$ HISQ results are slightly higher than the $N_t = 10$ and $N_t = 12$ results, and therefore have been omitted from the fits. At temperatures above 800 MeV only lattice results with $N_t = 6$ and 4 are available [20,21]. To take the larger discretization errors of the $N_t = 6$ and 4 results into account, we follow Ref. [21], scale them by factors 1.4 and 1.2, and include systematic errors of 40% and 20%, respectively. Contrary to the HISQ action results, we employ the continuum extrapolated stout action results [22,23] for simplicity. The resulting parameters are shown in Table III. We find that only the parametrization $s88h_{18}$ requires the use of all six terms in Eq. (A1). In the cases of $s83s_{18}$ and $s87h_{04}$ we are able to obtain equally good fits with only five terms, and thus set d_5 to zero by hand.

For the sake of completeness, we also parametrize the HRG part of the trace anomaly as

$$\frac{\epsilon - 3p}{T^4} = a_1 T^{m_1} + a_2 T^{m_2} + a_3 T^{m_3} + a_4 T^{m_4}. \quad (\text{A2})$$

To carry out the fit we evaluate the HRG trace anomaly in the temperature interval $70 < T/\text{MeV} < T_{\text{high}}$, where T_{high} depends on the parametrization, with 1 MeV steps assuming that each point has equal “error.” The limits have entirely utilitarian origin: in hydrodynamical applications the system decouples well above 70 MeV temperature and only a rough approximation of the EoS, $p = p(\epsilon)$, is needed at lower tem-

TABLE IV. The values of parameters for different fits of the HRG trace anomaly.

	a_1 (GeV $^{-m_1}$)	a_2 (GeV $^{-m_2}$)	a_3 (GeV $^{-m_3}$)	a_4 (GeV $^{-m_4}$)	m_1	m_2	m_3	m_4	T_{high} (MeV)
$s83s_{18}$	0.1850	1.985×10^4	1.278×10^5	-1.669×10^7	0	5	7	10	170
$s87h_{04}$	4.654	-879	8081	-7.039×10^6	1	3	4	10	190
$s88h_{18}$	0.1844	2.043×10^4	8.550×10^5	-2.434×10^7	0	5	8	10	169
$s95p$	4.654	-879	8081	-7.039×10^6	1	3	4	10	190

peratures. On the other hand we expect to switch to the lattice parametrization below T_{high} , and the HRG EoS above that temperature is not needed either. We fix the powers in Eq. (A2) again using brute force. We require them to be integers, go through all the combinations $0 \leq l_1 < l_2 < l_3 < l_4 \leq 10$, fit the parameters a_1, a_2, a_3, a_4 to the HRG trace anomaly evaluated with 1 MeV intervals, and choose the values l_1, l_2, l_3 , and l_4 which minimize the χ^2 . We end up with parameters shown in Table IV. To obtain the EoS, one also needs the pressure at the lower limit of the integration [see Eq. (1)] $T_{\text{low}} = 0.07$ GeV: $p(T_{\text{low}})/T_{\text{low}}^4 = 0.1661$. Our EoSs are available in a tabulated form at arXiv as ancillary files for this paper, and at Ref. [87]. These tables also include the option of a chemically frozen hadronic stage, and a list of resonances included in the hadronic stage with their properties and decay channels.

APPENDIX B: PREDICTING MODEL OUTPUT WITH GAUSSIAN PROCESSES

Let us assume that we do not know exactly what the model's output y for a particular input parameter \vec{x} is, but we know its most probable value $\mu(\vec{x})$. We postulate that the probability distribution for the output value $P(y)$ is a normal distribution with mean $\mu(\vec{x})$ and so far unknown width σ . Thus the probability distribution for a set Y_a of N model outputs for observable a , corresponding to a set X of N points in the parameter space, is a multivariate normal distribution:

$$\mathcal{G} : X \rightarrow Y_a \sim \mathcal{N}(\boldsymbol{\mu}, \mathbf{C}), \quad (\text{B1})$$

where $\boldsymbol{\mu} = \mu(X) = \{\mu(\vec{x}_1), \dots, \mu(\vec{x}_N)\}$ is the mean of the distribution, and \mathbf{C} is the covariance matrix defined by the covariance function $c(\vec{x}, \vec{x}')$:

$$\mathbf{C} = \mathcal{C}_{X,X} = \begin{pmatrix} c(\vec{x}_1, \vec{x}_1) & \dots & c(\vec{x}_1, \vec{x}_N) \\ \vdots & \ddots & \vdots \\ c(\vec{x}_N, \vec{x}_1) & \dots & c(\vec{x}_N, \vec{x}_N) \end{pmatrix}. \quad (\text{B2})$$

As we are only interested in interpolating within the training data, we may set $\mu(X) \equiv 0$, and construct the covariance function $c(\vec{x}, \vec{x}')$ in such a way that the probability distribution is narrow at the training points nevertheless. This way we minimize our *a priori* assumptions about the model behavior in regions of parameter space not covered by the training data.¹¹ Our chosen covariance function is a radial-basis func-

tion (RBF) with a noise term

$$c(\vec{x}, \vec{x}') = \theta_0 \exp\left(-\sum_{i=1}^n \frac{(x_i - x'_i)^2}{2\theta_i^2}\right) + \theta_{\text{noise}} \delta_{\vec{x}\vec{x}'}. \quad (\text{B3})$$

The hyperparameters $\vec{\theta} = (\theta_0, \theta_1, \dots, \theta_n, \theta_{\text{noise}})$, where n is the dimension of the input parameter space, are not known *a priori* and must be estimated from training data, consisting of simulation output U computed at training points T , by maximizing the log-likelihood (see Chapter 5 of [64])

$$\log P(U|T, \vec{\theta}) = -\frac{1}{2} U^T \mathbf{C}^{-1}(T, \vec{\theta}) U - \frac{1}{2} \log |\mathbf{C}(T, \vec{\theta})| - \frac{N}{2} \log(2\pi). \quad (\text{B4})$$

Emulator prediction for the model output y_0 at a point \vec{x}_0 can then be determined by writing a joint probability distribution for the output at various points in parameter space:

$$\begin{pmatrix} y_0 \\ U \end{pmatrix} \sim \mathcal{N}\left(\begin{pmatrix} 0 \\ \vec{0} \end{pmatrix}, \begin{pmatrix} \mathcal{C}_{0,0} & \mathcal{C}_{0,T} \\ \mathcal{C}_{T,0} & \mathcal{C}_{T,T} \end{pmatrix}\right), \quad (\text{B5})$$

from which we can derive the conditional predictive mean $y^{\text{GP}}(\vec{x}_0)$ and associated variance $\sigma^{\text{GP}}(\vec{x}_0)^2$ as (see, e.g., Appendix A.2 of [64])

$$y^{\text{GP}}(\vec{x}_0) = \mathcal{C}_{0,T} \mathbf{C}_{T,T}^{-1} U, \quad \sigma^{\text{GP}}(\vec{x}_0)^2 = \mathcal{C}_{0,0} - \mathcal{C}_{0,T} \mathbf{C}_{T,T}^{-1} \mathcal{C}_{T,0}. \quad (\text{B6})$$

Note that we use the training data U twice: first in Eq. (B4) to determine the hyperparameters $\vec{\theta}$ of the covariance function $c(\vec{x}, \vec{x}')$ and then in Eq. (B6) as a condition for the GP prediction.

APPENDIX C: PRINCIPAL COMPONENT ANALYSIS

We reduce the number of Gaussian processes needed for model emulation with principal component analysis (PCA), which transforms the data in the directions of maximal variance.

We represent the model output with an $N \times m$ matrix Y , where N is the number of simulation points and m the number of observables. In preparation for the PCA, the data columns are normalized with the corresponding experimental values to obtain dimensionless quantities, and centered by subtracting the mean of each observable from the elements of each column; we denote this scaled and shifted data matrix by \hat{Y} .

We then want to find an eigenvalue decomposition of the covariance matrix $\hat{Y}^T \hat{Y}$:

$$\hat{Y}^T \hat{Y} = V \Lambda V^T, \quad (\text{C1})$$

¹¹Note that we use Gaussian process to estimate the model output of the principal components, not the actual observables; see Appendix C.

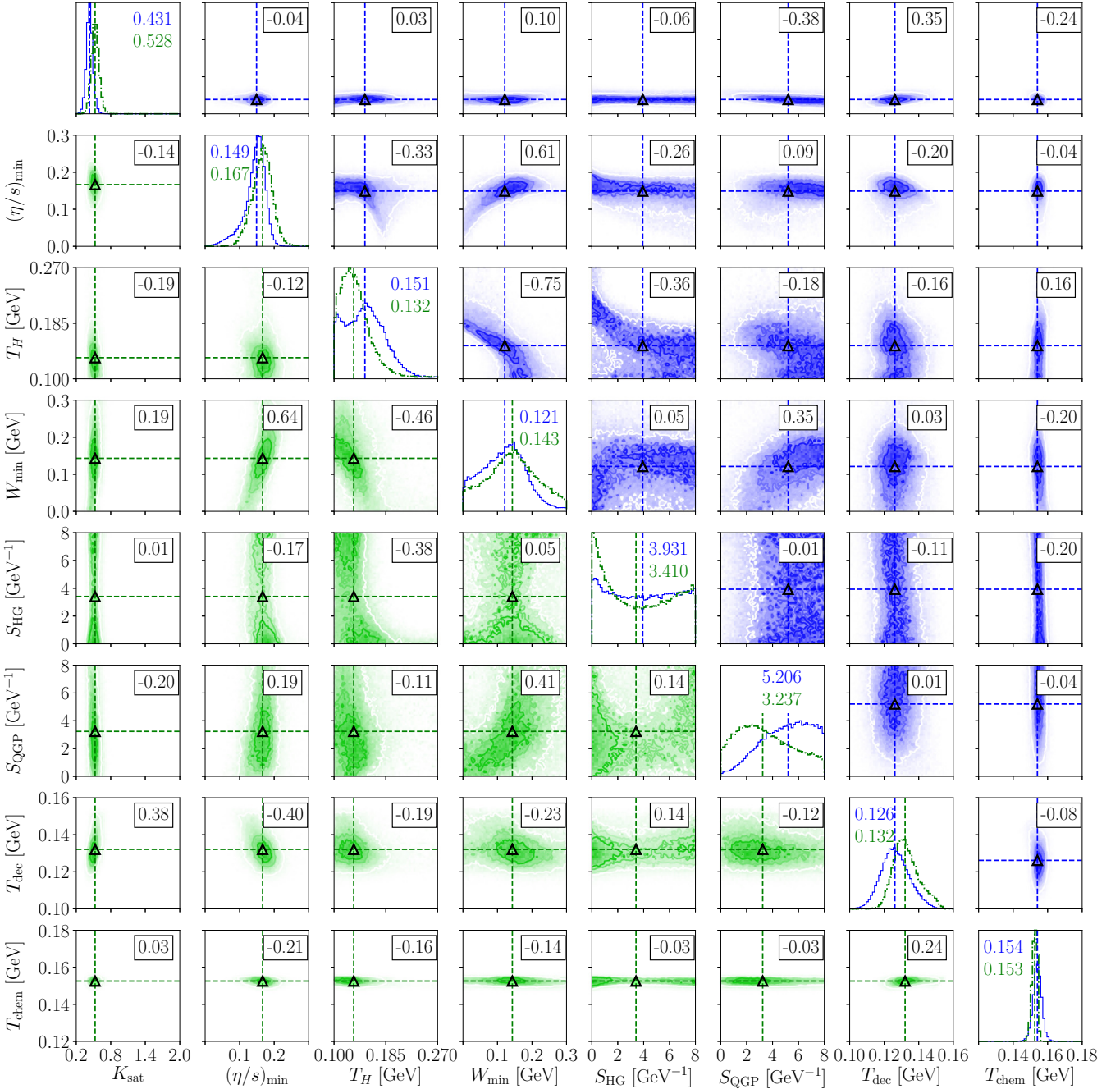


FIG. 12. Posterior probability distribution for the $s88h_{18}$ (lower triangle, green color) and $s95p$ (upper triangle, blue color) EoSs. Diagonal panels: Marginalized 1-D distributions for each parameter. Solid blue line: $s95p$. Dash-dotted green line: $s88h_{18}$. Dashed lines and numbers indicate median value, with upper number corresponding to $s95p$ and lower number to $s88h_{18}$. Off-diagonal panels: 2-D projections of the posterior distributions. Dashed lines indicate median values for each parameter, while the framed numbers refer to Spearman rank correlation coefficients for each parameter pair.

where Λ is the diagonal matrix containing the eigenvalues $\lambda_1, \dots, \lambda_m$ and V is an orthogonal matrix containing the eigenvectors of the covariance matrix.

The eigenvalue decomposition is found by factorizing \hat{Y} via the singular value decomposition:

$$\hat{Y} = USV^T, \quad (\text{C2})$$

where S is a diagonal matrix containing the singular values (square roots of the eigenvalues of $\hat{Y}^T \hat{Y}$) and V contains the

right-singular vectors of \hat{Y} (eigenvectors of $\hat{Y}^T \hat{Y}$); these are the principal components (PCs). Matrix U contains the left-singular vectors of \hat{Y} , which are eigenvectors of $\hat{Y} \hat{Y}^T$.

The eigenvalues are proportional to the total variance of the data. Since $\lambda_1 \geq \lambda_2 \geq \dots \geq \lambda_m$, the fraction of the total variance explained by the k th principal component, $\lambda_k / (\sum_{j=1}^m \lambda_j)$, becomes negligible starting from some index $k < m$. This allows us to define a lower-rank approximation of the original transformed data matrix $Z =$

$\hat{Y}V$ as $Z_k = \hat{Y}V_k$, where V_k contains the first k columns of V .

The transformation of a vector \vec{y} from the space of observables to a vector \vec{z} in the reduced-dimension principal component space is thus defined as

$$\vec{z} = \vec{y}V_k, \quad (\text{C3})$$

while for matrices [such as the covariance matrix in the likelihood function (5)] the transformation is

$$\Sigma_z = V_k^T \Sigma_y V_k. \quad (\text{C4})$$

To compare an emulator prediction \vec{z}^{GP} against physical observables, we use the inverse transformation

$$\vec{y}^{\text{GP}} = \vec{z}^{\text{GP}} V_k^T. \quad (\text{C5})$$

APPENDIX D: CORRELATIONS BETWEEN THE MODEL PARAMETERS

Figure 12 provides a more detailed view of the 8-dimensional posterior probability distribution, using the

analysis results for the $s88h_{18}$ and $s95p$ EoSs as an example. The diagonal panels show the marginalized one-dimensional distributions for each parameter, which were summarized in Figs. 3 and 4 in Sec. VI. The off-diagonal panels illustrate the correlations between each parameter pair (X, Y) . The correlation strength is quantified with the Spearman rank correlation coefficient [88], which is the Pearson correlation coefficient between the rank values r_X and r_Y :

$$\rho = \frac{C(r_X, r_Y)}{\sigma(r_X)\sigma(r_Y)}, \quad (\text{D1})$$

where C refers to covariance and σ to standard deviation. This relaxes the assumption of a linear relationship, present in the Pearson correlation coefficient, and is instead a measure of the monotonic relationship between the two parameters.

-
- [1] M. Luzum and P. Romatschke, *Phys. Rev. C* **78**, 034915 (2008); **79**, 039903(E) (2009).
 - [2] P. Bozek, *Phys. Rev. C* **81**, 034909 (2010).
 - [3] H. Song, S. A. Bass, and U. Heinz, *Phys. Rev. C* **83**, 054912 (2011); **87**, 019902(E) (2013).
 - [4] S. Ryu, J.-F. Paquet, C. Shen, G. S. Denicol, B. Schenke, S. Jeon, and C. Gale, *Phys. Rev. Lett.* **115**, 132301 (2015).
 - [5] I. A. Karpenko, P. Huovinen, H. Petersen, and M. Bleicher, *Phys. Rev. C* **91**, 064901 (2015).
 - [6] U. Heinz and R. Snellings, *Annu. Rev. Nucl. Part. Sci.* **63**, 123 (2013).
 - [7] C. Gale, S. Jeon, and B. Schenke, *Int. J. Mod. Phys. A* **28**, 1340011 (2013).
 - [8] P. Huovinen, *Int. J. Mod. Phys. E* **22**, 1330029 (2013).
 - [9] C. Shen, [arXiv:2001.11858](https://arxiv.org/abs/2001.11858).
 - [10] J. E. Bernhard, J. S. Moreland, S. A. Bass, J. Liu, and U. Heinz, *Phys. Rev. C* **94**, 024907 (2016).
 - [11] S. A. Bass, J. E. Bernhard, and J. S. Moreland, *Nucl. Phys. A* **967**, 67 (2017).
 - [12] J. E. Bernhard, J. S. Moreland, and S. A. Bass, *Nat. Phys.* **15**, 1113 (2019).
 - [13] H. Niemi, K. J. Eskola, and R. Paatelainen, *Phys. Rev. C* **93**, 024907 (2016).
 - [14] S. Pratt, E. Sangaline, P. Sorensen, and H. Wang, *Phys. Rev. Lett.* **114**, 202301 (2015).
 - [15] J. S. Moreland and R. A. Soltz, *Phys. Rev. C* **93**, 044913 (2016).
 - [16] P. Alba, V. Mantovani Sarti, J. Noronha, J. Noronha-Hostler, P. Parotto, I. Portillo Vazquez, and C. Ratti, *Phys. Rev. C* **98**, 034909 (2018).
 - [17] P. Huovinen and P. Petreczky, *Nucl. Phys. A* **837**, 26 (2010).
 - [18] A. Bazavov *et al.*, *Phys. Rev. D* **80**, 014504 (2009).
 - [19] B. Schenke, C. Shen, and P. Tribedy, *Phys. Rev. C* **99**, 044908 (2019).
 - [20] A. Bazavov *et al.* (HotQCD Collaboration), *Phys. Rev. D* **90**, 094503 (2014).
 - [21] A. Bazavov, P. Petreczky, and J. H. Weber, *Phys. Rev. D* **97**, 014510 (2018).
 - [22] S. Borsanyi, G. Endrodi, Z. Fodor, A. Jakovac, S. D. Katz, S. Krieg, C. Ratti, and K. K. Szabo, *J. High Energy Phys.* **11** (2010) 077.
 - [23] S. Borsanyi, Z. Fodor, C. Hoelbling, S. D. Katz, S. Krieg, and K. K. Szabo, *Phys. Lett. B* **730**, 99 (2014).
 - [24] H. Niemi, K. J. Eskola, R. Paatelainen, and K. Tuominen, *Phys. Rev. C* **93**, 014912 (2016).
 - [25] S. Adler *et al.* (PHENIX Collaboration), *Phys. Rev. C* **71**, 034908 (2005).
 - [26] S. S. Adler *et al.* (PHENIX Collaboration), *Phys. Rev. C* **69**, 034909 (2004).
 - [27] J. Adams *et al.* (STAR Collaboration), *Phys. Rev. C* **72**, 014904 (2005).
 - [28] K. Aamodt *et al.* (ALICE Collaboration), *Phys. Rev. Lett.* **106**, 032301 (2011).
 - [29] B. Abelev *et al.* (ALICE Collaboration), *Phys. Rev. C* **88**, 044910 (2013).
 - [30] J. Adam *et al.* (ALICE Collaboration), *Phys. Rev. Lett.* **116**, 132302 (2016).
 - [31] J. Adam *et al.* (ALICE Collaboration), *Phys. Rev. Lett.* **116**, 222302 (2016).
 - [32] G. Boyd, J. Engels, F. Karsch, E. Laermann, C. Legeland, M. Lutgemeier, and B. Petersson, *Nucl. Phys. B* **469**, 419 (1996).
 - [33] C. Anderlik, L. P. Csernai, F. Grassi, W. Greiner, Y. Hama, T. Kodama, Z. I. Lázár, V. K. Magas, and H. Stöcker, *Phys. Rev. C* **59**, 3309 (1999).
 - [34] P. Huovinen and P. Petreczky, *PoS (Confinement2018)*, 145 (2019).
 - [35] S. A. Bass *et al.*, *Prog. Part. Nucl. Phys.* **41**, 255 (1998); M. Bleicher *et al.*, *J. Phys. G* **25**, 1859 (1999).
 - [36] J. Weil, V. Steinberg, J. Staudenmaier, L. G. Pang, D. Oliinychenko, J. Mohs, M. Kretz, T. Kehrenberg, A. Goldschmidt, B. Bäuchle *et al.*, *Phys. Rev. C* **94**, 054905 (2016).
 - [37] S. Eidelman *et al.* (Particle Data Group), *Phys. Lett. B* **592**, 1 (2004).
 - [38] R. Venugopalan and M. Prakash, *Nucl. Phys. A* **546**, 718 (1992).

- [39] W. Broniowski, F. Giacosa, and V. Begun, *Phys. Rev. C* **92**, 034905 (2015).
- [40] M. Tanabashi *et al.* (Particle Data Group), *Phys. Rev. D* **98**, 030001 (2018).
- [41] Non-strange-meson summary tables: <http://pdg.lbl.gov/2018/tables/rpp2018-tab-mesons-light.pdf>; strange mesons: <http://pdg.lbl.gov/2018/tables/rpp2018-tab-mesons-strange.pdf>; p , n , and N resonances: <http://pdg.lbl.gov/2018/tables/rpp2018-tab-baryons-N.pdf>; Lambda, Lambda resonances: <http://pdg.lbl.gov/2018/tables/rpp2018-tab-baryons-Lambda.pdf>; Sigma, Sigma resonances: <http://pdg.lbl.gov/2018/tables/rpp2018-tab-baryons-Sigma.pdf>; Xi, Xi resonances: <http://pdg.lbl.gov/2018/tables/rpp2018-tab-baryons-Xi.pdf>; Omega, Omega resonances: <http://pdg.lbl.gov/2018/tables/rpp2018-tab-baryons-Omega.pdf>.
- [42] A. Majumder and B. Muller, *Phys. Rev. Lett.* **105**, 252002 (2010).
- [43] A. Bazavov *et al.*, *Phys. Rev. Lett.* **113**, 072001 (2014).
- [44] C. Fernández-Ramírez, P. M. Lo, and P. Petreczky, *Phys. Rev. C* **98**, 044910 (2018).
- [45] P. Alba, V. M. Sarti, J. Noronha-Hostler, P. Parotto, I. Portillo-Vazquez, C. Ratti, and J. M. Stafford, *Phys. Rev. C* **101**, 054905 (2020).
- [46] H. Niemi, G. S. Denicol, P. Huovinen, E. Molnar, and D. H. Rischke, *Phys. Rev. C* **86**, 014909 (2012).
- [47] H. Niemi, G. S. Denicol, P. Huovinen, E. Molnar, and D. H. Rischke, *Phys. Rev. Lett.* **106**, 212302 (2011).
- [48] K. J. Eskola, H. Niemi, R. Paatelainen, and K. Tuominen, *Phys. Rev. C* **97**, 034911 (2018).
- [49] E. Molnar, H. Niemi, and D. H. Rischke, *Eur. Phys. J. C* **65**, 615 (2010).
- [50] W. Israel and J. M. Stewart, *Ann. Phys.* **118**, 341 (1979).
- [51] G. S. Denicol, H. Niemi, E. Molnar, and D. H. Rischke, *Phys. Rev. D* **85**, 114047 (2012); G. S. Denicol, *ibid.* **91**, 039902 (2015).
- [52] E. Molnar, H. Niemi, G. S. Denicol, and D. H. Rischke, *Phys. Rev. D* **89**, 074010 (2014).
- [53] P. Huovinen, *Eur. Phys. J. A* **37**, 121 (2008).
- [54] T. Hirano and K. Tsuda, *Phys. Rev. C* **66**, 054905 (2002).
- [55] K. Paech and S. Pratt, *Phys. Rev. C* **74**, 014901 (2006).
- [56] K. Dusling and T. Schäfer, *Phys. Rev. C* **85**, 044909 (2012).
- [57] J. B. Rose, J. M. Torres-Rincon, A. Schäfer, D. Oliinychenko, and H. Petersen, *Phys. Rev. C* **97**, 055204 (2018).
- [58] H. Bebie, P. Gerber, J. Goity, and H. Leutwyler, *Nucl. Phys. B* **378**, 95 (1992).
- [59] K. J. Eskola, K. Kajantie, P. V. Ruuskanen, and K. Tuominen, *Nucl. Phys. B* **570**, 379 (2000).
- [60] R. Paatelainen, K. J. Eskola, H. Holopainen, and K. Tuominen, *Phys. Rev. C* **87**, 044904 (2013).
- [61] R. Paatelainen, K. J. Eskola, H. Niemi, and K. Tuominen, *Phys. Lett. B* **731**, 126 (2014).
- [62] W. D. Vousden, W. M. Farr, and I. Mandel, *Mon. Not. R. Astron. Soc.* **455**, 1919 (2016).
- [63] D. Foreman-Mackey, D. W. Hogg, D. Lang, and J. Goodman, *Publ. Astron. Soc. Pac.* **125**, 306 (2013).
- [64] C. E. Rasmussen and C. K. I. Williams, *Gaussian Processes for Machine Learning* (MIT Press, Cambridge, MA, 2006).
- [65] F. Pedregosa *et al.*, *J. Mach. Learn. Res.* **12**, 2825 (2011).
- [66] pyDOE: Design of Experiments for Python, <https://pythonhosted.org/pyDOE/randomized.html>.
- [67] A. Andronic, P. Braun-Munzinger, K. Redlich, and J. Stachel, *Nature (London)* **561**, 321 (2018).
- [68] V. Mykhaylova, M. Bluhm, K. Redlich, and C. Sasaki, *Phys. Rev. D* **100**, 034002 (2019).
- [69] M. Prakash, M. Prakash, R. Venugopalan, and G. Welke, *Phys. Rep.* **227**, 321 (1993).
- [70] L. P. Csernai, J. I. Kapusta, and L. D. McLerran, *Phys. Rev. Lett.* **97**, 152303 (2006).
- [71] N. Demir and S. A. Bass, *Phys. Rev. Lett.* **102**, 172302 (2009).
- [72] A. Wiranata, M. Prakash, P. Huovinen, V. Koch, and X. Wang, *J. Phys.: Conf. Ser.* **535**, 012017 (2014).
- [73] E. V. Shuryak and I. Zahed, *Phys. Rev. C* **70**, 021901(R) (2004).
- [74] I. Wetzorke, F. Karsch, E. Laermann, P. Petreczky, and S. Stickan, *Nucl. Phys. B, Proc. Suppl.* **106-107**, 510 (2002).
- [75] F. Karsch, S. Datta, E. Laermann, P. Petreczky, S. Stickan, and I. Wetzorke, *Nucl. Phys. A* **715**, 701 (2003).
- [76] M. Asakawa, T. Hatsuda, and Y. Nakahara, *Nucl. Phys. B, Proc. Suppl.* **119**, 481 (2003).
- [77] S. Mukherjee, P. Petreczky, and S. Sharma, *Phys. Rev. D* **93**, 014502 (2016).
- [78] A. Bazavov, N. Brambilla, P. Petreczky, A. Vairo, and J. H. Weber (TUMQCD Collaboration), *Phys. Rev. D* **98**, 054511 (2018).
- [79] A. Bazavov, H. T. Ding, P. Hegde, F. Karsch, C. Miao, S. Mukherjee, P. Petreczky, C. Schmidt, and A. Velytsky, *Phys. Rev. D* **88**, 094021 (2013).
- [80] R. Bellwied, S. Borsányi, Z. Fodor, S. D. Katz, A. Pásztor, C. Ratti, and K. K. Szabó, *Phys. Rev. D* **92**, 114505 (2015).
- [81] H. T. Ding, S. Mukherjee, H. Ohno, P. Petreczky, and H. P. Schadler, *Phys. Rev. D* **92**, 074043 (2015).
- [82] P. Huovinen, P. M. Lo, M. Marczenko, K. Morita, K. Redlich, and C. Sasaki, *Phys. Lett. B* **769**, 509 (2017).
- [83] P. Huovinen, *Nucl. Phys. A* **761**, 296 (2005).
- [84] J. Noronha-Hostler, B. Betz, J. Noronha, and M. Gyulassy, *Phys. Rev. Lett.* **116**, 252301 (2016).
- [85] C. Andres, N. Armesto, H. Niemi, R. Paatelainen, and C. A. Salgado, *Phys. Lett. B* **803**, 135318 (2020).
- [86] D. Zigic, B. Ilic, M. Djordjevic, and M. Djordjevic, *Phys. Rev. C* **101**, 064909 (2020).
- [87] See <https://osf.io/thazn/wiki/home>.
- [88] C. Spearman, *Am. J. Psychol.* **15**, 72 (1904).