

The effects of Prandtl number on the nonlinear dynamics of Kelvin-Helmholtz instability in two dimensions

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(Received xx; revised xx; accepted xx)

It is known that the pitchfork bifurcation of Kelvin-Helmholtz instability occurring at minimum gradient Richardson number $Ri_m \simeq 1/4$ in viscous stratified shear flows can be subcritical or supercritical depending on the value of the Prandtl number, Pr . Here we study stratified shear flow restricted to two dimensions at finite Reynolds number, continuously forced to have a constant background density gradient and a hyperbolic tangent shear profile, corresponding to the ‘Drazin model’ base flow. Bifurcation diagrams are produced for fluids with $Pr = 0.7$ (typical for air), 3 and 7 (typical for water). For $Pr = 3$ and 7, steady billow-like solutions are found to exist for strongly stable stratification of Ri_m beyond $1/2$. Interestingly, these solutions are not a direct product of a Kelvin-Helmholtz instability, having half the wavelength of the linear instability, and arising through a superharmonic bifurcation. These short-wavelength states can be tracked down to at least $Pr \approx 2.3$ and act as instigators of complex dynamics, even in strongly stratified flows. Direct numerical simulations of forced and unforced two-dimensional flows are performed, which support the results of the bifurcation analyses. Perturbations are observed to grow approximately exponentially from random initial conditions where no modal instability is predicted by a linear stability analysis.

1. Introduction

Kelvin-Helmholtz instability (KHI) is believed to be important in geophysical flows found in both the oceans (Smyth & Moum 2012) and atmosphere (Fukao *et al.* 2011; Sun *et al.* 2015). Of particular importance is the generation of abyssal oceanic turbulence by the break down of shear instabilities, which is an area of significant uncertainty in climate modelling (Gregg *et al.* 2018). Direct observations in the atmosphere, such as of sheared clouds, are relatively straightforward to perform, whereas only a few studies have observed Kelvin-Helmholtz billows in the abyssal ocean (Van Haren & Gostiaux 2010). Amongst other parameters, the Prandtl number $Pr := \nu/\kappa$ (the ratio of kinematic viscosity ν to thermal diffusivity κ) involved in these two settings is different, making it important to understand any resulting differences in the dynamics. In the atmosphere, $Pr \simeq 0.7$ whereas in the ocean $Pr \simeq 7$ and when the diffusion of salt is important (described by a diffusivity κ_s), the equivalent Schmidt number $Sc := \nu/\kappa_s \simeq 700$ (Thorpe 2005).

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Several simple models of stratified mixing layers have been proposed which exhibit KHI. The two most commonly used, the Drazin (1958) and Holmboe (1960) models, are both found to be linearly stable in the inviscid case when the minimum gradient Richardson number Ri_m (as defined below) is greater than one quarter. This observation led to the celebrated Miles-Howard theorem (Miles 1961; Howard 1961), which shows that inviscid flows are always linearly stable when the gradient Richardson number is everywhere greater than one quarter. A longstanding challenge has been to determine whether significant *nonlinear* dynamics are also precluded for $Ri_m > 1/4$.

With viscosity, the Prandtl number enters the problem and there is a body of evidence suggesting this parameter has a significant impact on the behaviour of KHI (Klaassen & Peltier 1985*a*; Salehipour *et al.* 2015; Rahmani *et al.* 2016) and stratified turbulence generally (Brucker & Sarkar 2007). In particular, it has been shown that the bifurcation of KHI near (minimum gradient) Richardson number $1/4$ is subcritical when $Pr > 1$ and supercritical when $Pr < 1$ (Brown *et al.* 1981; Churilov & Shukhman 1987; Lott & Teitelbaum 1992; Mkhinini *et al.* 2013). (In this paper we use the dynamical systems convention that ‘subcritical’ refers to the stable region $Ri_m > Ri_c$ and ‘supercritical’ to the unstable region $Ri_m < Ri_c$.) Despite this, most simulations studying the nonlinear behaviour of KHI have concentrated on the degenerate value $Pr = 1$ (Klaassen & Peltier 1985*b*; Caulfield & Peltier 2000; Mashayek & Peltier 2011; Kaminski *et al.* 2017), which allows a coarser computational grid to be used compared with higher Pr .

Although the effect of Pr on the sub/supercriticality of the bifurcation is well documented, this gives only a weakly nonlinear understanding beyond classical linear stability analyses, and cannot predict the full nonlinear effects. It could be the case that full turbulence is possible through subcritical transition for flows with high minimum Richardson numbers, substantially above $1/4$, where turbulence is usually assumed to be suppressed (Thorpe 2005), or it could be that nontrivial, nonlinear states do not exist in flows with Ri_m significantly larger than $1/4$, and that the behaviour is simple and transient, as was found for $Pr = 1$ (Parker *et al.* 2019). Below, we argue for the former scenario by presenting direct evidence that 2-dimensional finite-amplitude billow-like states exist for $Ri_m \gtrsim 0.4$ - i.e. substantially above $1/4$ - for $Pr \gtrsim 2.3$ and indirect evidence that this situation continues below this threshold. Importantly, this implies that complicated temporal dynamics are possible for flows generally considered inert due to a lack of a Kelvin-Helmholtz linear instability.

To establish this key result, the paper proceeds as follows. In §2, the equations of our forced model and numerical methods are briefly presented while in §3, bifurcation diagrams of the forced two-dimensional flow are given for $Pr \in \{0.7, 3, 7\}$, and the differences and continuous change between these two values is discussed. Finally, §4 compares the time evolution of the forced and the equivalent unforced systems by performing a 2D direct numerical simulation (DNS) of the flow at various Richardson numbers, before concluding remarks are made in §5.

2. Methods

We study the Boussinesq equations for velocity \mathbf{u} and buoyancy b :

$$\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} = -\nabla p + Ri_b b \mathbf{e}_z + \frac{1}{Re} \nabla^2 \mathbf{u}, \quad (2.1a)$$

$$\frac{\partial b}{\partial t} + \mathbf{u} \cdot \nabla b = \frac{1}{RePr} \nabla^2 b, \quad (2.1b)$$

$$\nabla \cdot \mathbf{u} = 0. \quad (2.1c)$$

81 The non-dimensional parameters are the Reynolds number Re , quantifying the relative
 82 importance of inertia to viscosity, the Prandtl number Pr , quantifying the relative
 83 importance of diffusion of buoyancy to viscosity, and the bulk Richardson number
 84 Ri_b , quantifying the relative importance of buoyancy to shear. With a gravitational
 85 acceleration g , shear layer depth $2d^*$, velocity difference $2U^*$, reference density ρ^* ,
 86 reference density gradient $\Delta\rho^*/d^*$, and diffusivities ν and κ for momentum and density
 87 respectively, these are calculated as

$$Re := U^*d^*/\nu, \quad Pr := \nu/\kappa \quad \text{and} \quad Ri_b := \frac{g\Delta\rho^*d^*}{\rho^*U^{*2}}. \quad (2.2)$$

88 In this paper we consider the evolution of perturbations away from the flow $\mathbf{u} =$
 89 $\tanh z \mathbf{e}_x$, $b = z$. This is the so-called ‘Drazin model’ of a mixing layer, for which weakly-
 90 nonlinear analyses have been performed (Churilov & Shukhman 1987). Unlike the perhaps
 91 more commonly considered ‘Holmboe model’ with $b = \tanh z$, the Drazin model does not
 92 exhibit the viscous Holmboe instability discussed in Parker *et al.* (2020), which would
 93 complicate our picture. Using the Drazin model, the gradient Richardson number of the
 94 flow Ri_g is bounded below by Ri_b , since

$$Ri_g(z) := Ri_b \frac{db/dz}{(du/dz)^2} \geq Ri_m = Ri_b = Ri_g(0). \quad (2.3)$$

95 Therefore, for this flow, the dynamically significant *minimum* gradient Richardson number
 96 Ri_m corresponds to the *bulk* Richardson number Ri_b which appears as a coupling
 97 parameter in the governing equations. Furthermore, the Miles-Howard theorem thus
 98 implies linear stability for $Ri_b > 1/4$ at infinite Re .

99 For finite Re , these choices of velocity and buoyancy profiles are not steady solutions of
 100 (2.1), but will diffuse away on an $O(Re)$ timescale. Nevertheless, the background profiles
 101 can be considered steady for perturbation dynamics over a shorter timescale. Therefore,
 102 when finding bifurcation diagrams (which require a non-decaying base state from which
 103 finite amplitude states can bifurcate), we study solutions of the related forced equations

$$\frac{\partial \mathbf{u}}{\partial t} + \mathbf{u} \cdot \nabla \mathbf{u} + \tanh z \frac{\partial \mathbf{u}}{\partial x} + w \operatorname{sech}^2 z = -\nabla p + Ri_b b \mathbf{e}_z + \frac{1}{Re} \nabla^2 \mathbf{u}, \quad (2.4a)$$

$$\frac{\partial b}{\partial t} + \mathbf{u} \cdot \nabla b + \tanh z \frac{\partial b}{\partial x} + w = \frac{1}{RePr} \nabla^2 b, \quad (2.4b)$$

$$\nabla \cdot \mathbf{u} = 0, \quad (2.4c)$$

104 where now \mathbf{u} , b and p represent the (possibly large) disturbances away from the
 105 background flow. Throughout, we take $Re = 1000$ which is relatively low compared with
 106 most modern direct numerical simulations, (see for example Salehipour *et al.* (2015)) but
 107 the high Pr combined with the computational intensity of the state tracking means that
 108 higher Re are not at present feasible. This limitation is discussed in §5.

109 The equations are solved on a two-dimensional domain periodic in the x direction with
 110 length L_x . Stress-free boundary conditions are imposed at $z = \pm L_z$. Both the solution of
 111 these equations and the finding and tracking of states and bifurcations largely uses the
 112 procedures presented in Parker *et al.* (2019). The key difference is that the non-uniform
 113 vertical grid has been modified to give a broader region of high resolution in the centre
 114 of the domain, in that we now use grid points located at

$$z_n = \frac{L_z}{3} \left[2 \left(\frac{2n - N_z - 1}{N_z - 1} \right)^7 + \left(\frac{2n - N_z - 1}{N_z - 1} \right) \right].$$

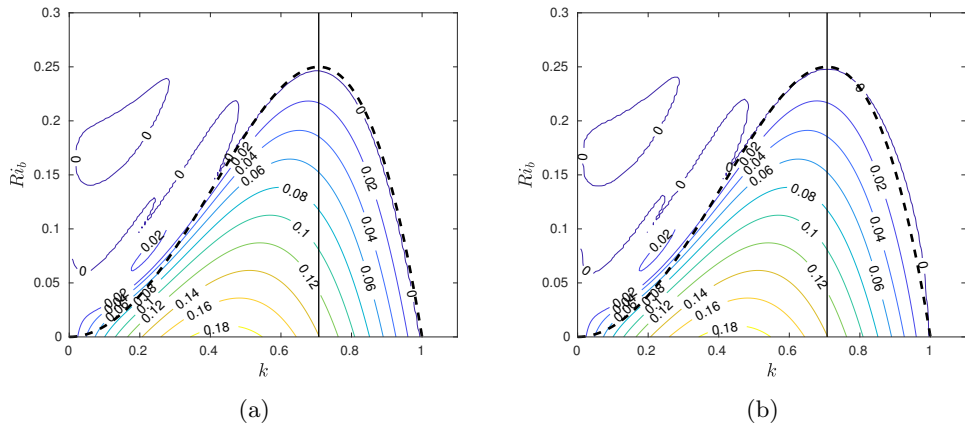


Figure 1: Linear stability diagrams of the flow at $Re = 1000$ for (a) $Pr = 0.7$ (b) $Pr = 7$, given as contours of the growth rate σ plotted against the wavenumber k and Ri_b , where the fastest growing mode of the form $e^{ik(x-ct)+\sigma t}\hat{\mathbf{u}}(z)$ has been found. The vertical line marks the wavenumber corresponding to a mode-1 disturbance in our domain of length $2\sqrt{2}\pi$. Note that mode- n , $n \geq 2$, are all stable for all Ri_b . The dashed line shows the stability boundary calculated by Drazin (1958) for $Re \rightarrow \infty$. Here, as with all the nonlinear calculations, the domain half-height is $L_z = 10$.

115 States are converged using Newton-GMRES, then followed as parameters vary using
 116 pseudo-arclength continuation. The bifurcation analysis of §3 uses a grid with $N_x =$
 117 64 horizontal grid points and $N_z = 512$ vertical grid points, which was shown to be
 118 sufficiently accurate by reconverging some of the points at $N_x = 256$, $N_z = 768$. For the
 119 direct numerical simulations of §4, for which much more complex spatial structures are
 120 possible, $N_x = 256$ and $N_z = 768$ is used.

121 For a state $X = (\mathbf{u}, b)$, we define the (energy-like) norm

$$\|X\| := \sqrt{\frac{1}{L_x} \int_{-L_z}^{L_z} dz \int_0^{L_x} dx (|\mathbf{u}|^2 + Ri_b b^2)}. \quad (2.5)$$

122 We also define a second function $m(X)$ of a given state, a measure of the component of
 123 the vertical velocity in the first Fourier mode

$$m(X) := \frac{1}{L_x} \int_{-L_z}^{L_z} dz \int_0^{L_x} dx u_z \sin \frac{2\pi x}{L_x}. \quad (2.6)$$

124 3. Bifurcation diagrams

125 Figure 1 shows the linear stability, calculated using a code from Smyth & Carpenter
 126 (2019), of the flows considered. The shape of the stability boundary is very close to the
 127 inviscid analytical result $Ri_b = k^2(1 - k^2)$ (Drazin 1958), which is overlaid. One curious
 128 difference is the presence of bands of instability at low wavenumbers. These have nonzero
 129 phase speed, and are similar to the ‘Holmboe instability’ mentioned in passing by Smyth
 130 & Peltier (1989) for a linear stratification and piecewise linear shear. The exact structure
 131 of these unstable bands is highly sensitive to the domain height, and they are believed to
 132 be caused by a resonance between discretised internal waves and the shear. This diagram

varies little as Pr is changed. However, as we demonstrate below, the nonlinear behaviour is strongly affected by Pr .

Henceforth we concentrate on the case of a domain of fixed streamwise length $L_x = 2\sqrt{2}\pi$. This is the wavelength of the marginally unstable mode at $Ri_b = 1/4$ in the inviscid, unbounded case, which is little modified in our viscous domain of finite height. The associated wavenumber $k_1 := 1/\sqrt{2}$ is marked on figure 1 as a vertical line. For $0.7 \leq Pr \leq 7$ the critical Richardson number Ri_c is close to, but slightly less than $1/4$ due to viscous effects: $Ri_c \approx 0.246$ for $Pr = 0.7$ and $Ri_c \approx 0.248$ for $Pr = 7$. Note that for this choice of domain size, only mode-1 disturbances (i.e. those which have one wavelength in the domain) are linearly unstable, as any mode with $k \geq 2k_1$ (and therefore any mode with two or more wavelengths in the domain) is linearly stable. A domain height of $L_z = 10$ was chosen, as this was assumed to be sufficiently large compared with L_x that the behaviour at large Ri_b is not significantly altered, while still being computationally efficient. At low Ri_b , this choice of L_z becomes significant, as discussed a little later.

Figure 2 shows the primary branch of steady KH states at $Pr = 0.7$ which bifurcates from the background flow at $Ri_b \approx 0.246$, in agreement with the linear stability analysis of figure 1a. The branch was found to be stable at $Ri_b = 0.24$, and a state was converged here using a simple timestepper. The rest of the branch was traced out using pseudo-arclength continuation. The pitchfork bifurcation is clearly supercritical, in agreement with weakly-nonlinear theory. Figure 2 also shows the bifurcation curve at $Pr = 1$ described in Parker *et al.* (2019). This is close to the degenerate case between super- and sub-criticality; it can just be made out that this case is very slightly subcritical.

Figure 3 shows the much more complicated situation at $Pr = 7$. The pitchfork bifurcation P_0 at $Ri_b \approx 0.247$ of the background flow is subcritical, in agreement with weakly nonlinear theory. The state which arises is therefore unstable, and was converged by a conventional edge-tracking procedure (e.g. Schneider *et al.* 2007). Edge-tracking was performed at $Ri_b = 0.26$, applying interval bisection with initial conditions of the upper branch state with wavenumber $k = k_1$ (see below), scaled to have lower amplitudes. At P_1 , two symmetric branches of wavenumber k_1 , which differ in phase by $\pi/2$, collide to give a state with wavenumber $k_2 := 2k_1$. The saddle-node bifurcations S_1 , S_2 and S_3 indicate the location of this mode-2 branch.

Separately to this, a stable upper branch state from $Pr = 3$ (where the system gives a simpler subcritical bifurcation, see below) was continued up in Pr to give rise to the mode-1 states of wavenumber k_1 which join at the pitchfork P_2 . At this value of Pr , none of this branch is stable. In fact, numerous other pitchfork and Hopf bifurcations, the precise locations of which were not determined, were found to exist on all branches, so that only a small section of the k_2 branch is stable. These secondary bifurcations give rise to the complex and apparently chaotic behaviour of the system discussed in §4. A systematic stability analysis of all the states in the figures was not performed, but none of a sample of states at $Pr = 7$ was found to be stable using a simple Arnoldi algorithm (see Parker *et al.* 2019).

As the states in figures 2 and 3 are traced to lower Ri_b and their amplitude and therefore physical extent becomes sufficiently large, the states begin to ‘feel’ the effects of the boundaries at $z = \pm L_z = \pm 10$. At this point, the structure changes dramatically, with the branches folding back to higher Ri_b , and the results are no longer physically relevant to unbounded flows. We have therefore chosen to exclude these sections from the diagrams. In an unbounded or sufficiently tall domain, the unstable states presumably continue past $Ri_b = 0$, as the unstratified Kelvin-Helmholtz instability saturates as a finite amplitude ‘billow’, although whether this also occurs for the k_2 branch is unclear.

Figure 4 depicts three low amplitude states on the branch between the pitchfork

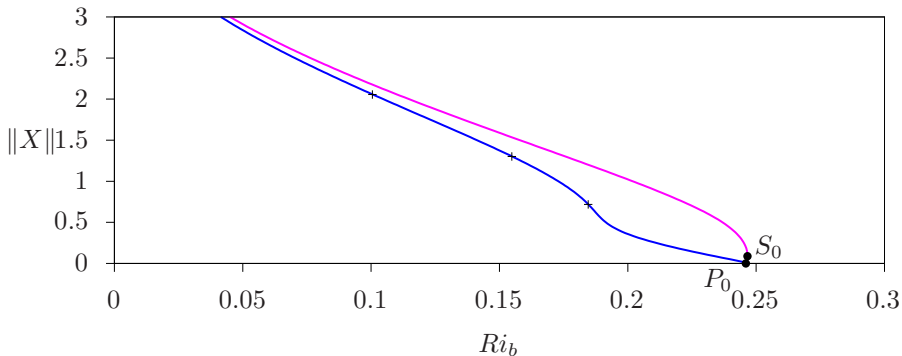


Figure 2: Bifurcation diagram for the Drazin model with a domain width of $2\sqrt{2}\pi$, with $Re = 1000$ and $Pr = 0.7$ (blue) and $Pr = 1$ (pink). The line represents a steady state solution with magnitude shown on the vertical axis. The crosses mark points reconverged at a higher resolution.

183 bifurcations P_0 and P_1 . Figure 4a is relatively close to the primary pitchfork P_0 , and shows
 184 a clear mode-1 structure of wavenumber k_1 , in agreement with the unstable eigenmode of
 185 the background flow, which the structure closely resembles. Figure 4b is further along the
 186 branch and there is now a noticeable mode-2 signal, manifesting as a structure emerging
 187 between the two ‘billows’. The amplitude has also increased. There is a natural transition
 188 therefore between the eigenmode and the pure mode-2 structure at P_1 , as shown in figure
 189 4c. A similar transition, at significantly higher amplitude, with structures much more
 190 closely resembling classic KH billows, is observed on the upper branch, as Ri_b increases
 191 towards P_2 (figure 5).

192 Figure 6 shows the mode-2 structures, i.e. those with wavenumber k_2 , at the three
 193 saddle-node bifurcation points. They are all qualitatively different. S_1 and S_3 , in figures
 194 6a and 6c respectively, are both highly reminiscent of classical KH billows, with a clear
 195 elliptical vortex. At S_1 the billows are significantly separated spatially, but at S_3 they
 196 are much more closely backed, but still with a distinctive ‘braid’ region connecting them.
 197 At S_2 , a low amplitude state intermediate between S_1 and S_3 , the structure is different
 198 again, and much less familiar.

199 The bifurcation points labelled in figure 3 can themselves be converged using a Newton-
 200 GMRES method, and tracked as Pr is varied, in a way identical to the tracking of
 201 bifurcation points as Re varies in Parker *et al.* (2019). The basic (mode-1) saddle-node
 202 bifurcation found in that paper, which we call S_0 , was continued to larger values of Pr
 203 just as those of figure 3 were continued to smaller values of Pr . The primary pitchfork
 204 P_0 , which exists for $Pr < 1$ too, can be found using this method or from linear stability
 205 analysis of the background flow. The results are shown in figure 7. S_1 and S_3 were found
 206 to be difficult to converge and continue, due to the presence of several marginally stable
 207 eigenvalues nearby, but were located directly at $Pr = 7$ and $Pr = 3$. S_0 could not be
 208 continued beyond $Pr = 3.8$, and there is no obvious bifurcation point which corresponds
 209 to S_0 in figure 3. P_1 , P_2 and S_2 all stopped converging below $Pr = 2.3$ and they appear
 210 to collide and disappear.

211 To clarify the situation, the intermediate value $Pr = 3$ was studied in detail (figure 8).
 212 The main (mode-1) branch, with $k = k_1$ and which connects to the fundamental pitchfork
 213 P_0 , is a simple subcritical curve, extending up to $Ri_b \approx 0.3$. Completely disconnected
 214 from this, extending to higher Ri_b , is a mode-2 loop (with $k = k_2$), which is a continuation

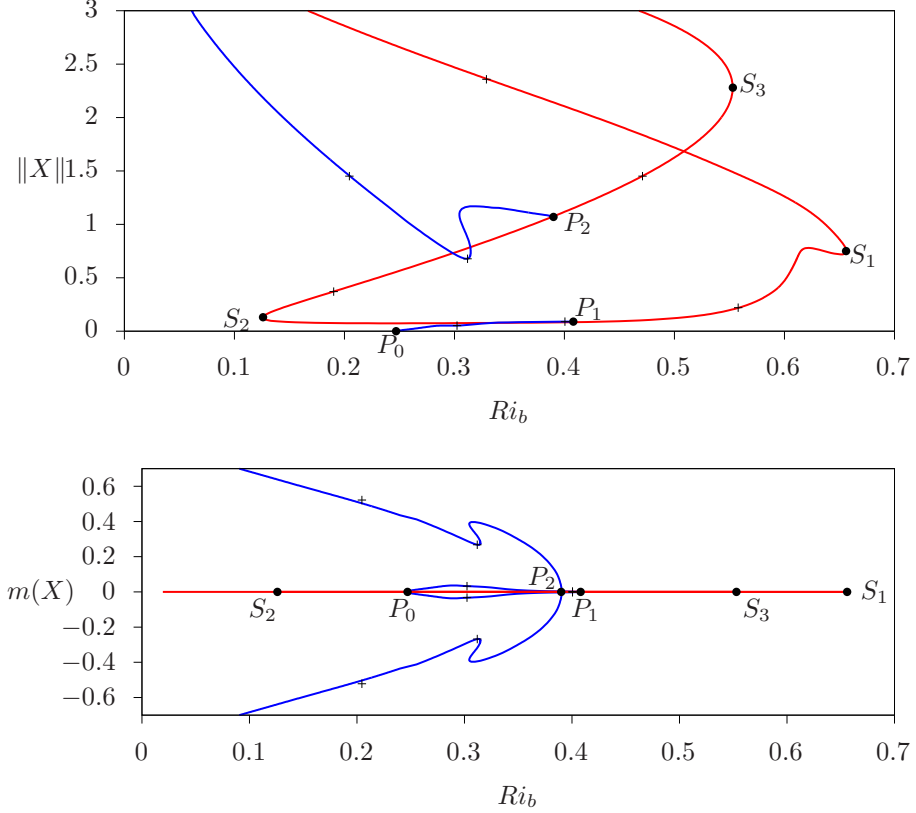


Figure 3: Top: As for figure 2, but with $Pr = 7$. Bottom: the same data, showing the contribution of the first Fourier mode in the streamwise direction to the states. The blue lines shows states with wavenumber $k_1 := 1/\sqrt{2}$, in agreement with the linear instability of the background flow. The red lines shows states with wavenumber $k_2 := 2k_1$, which arise at the pitchfork bifurcation P_1 . The crosses mark points reconverged at a higher resolution.

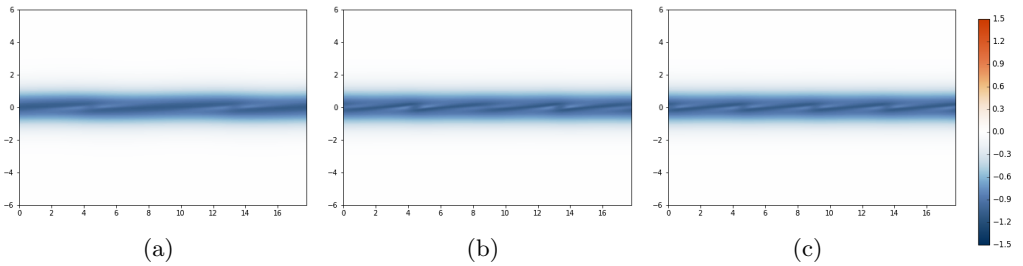


Figure 4: Vorticity fields of the steady perturbation states at $Pr = 7$ on the mode-1 branch connecting P_0 and P_1 . (a) $Ri_b \approx 0.3$, (b) $Ri_b \approx 0.4$, (c) at P_1 , $Ri_b \approx 0.41$. Here, and in all other such figures, two domain widths have been plotted to show the periodic structure.

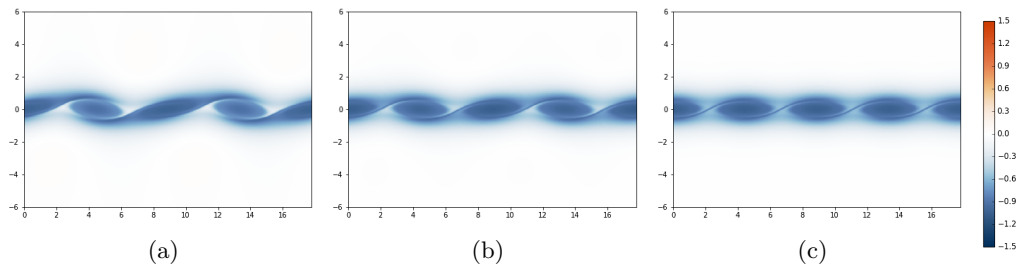


Figure 5: Vorticity fields on the upper mode-1 ($k = k_1$) branch at $Pr = 7$. (a) $Ri_b \approx 0.34$, (b) $Ri_b \approx 0.38$, (c) at P_2 , $Ri_b \approx 0.39$.

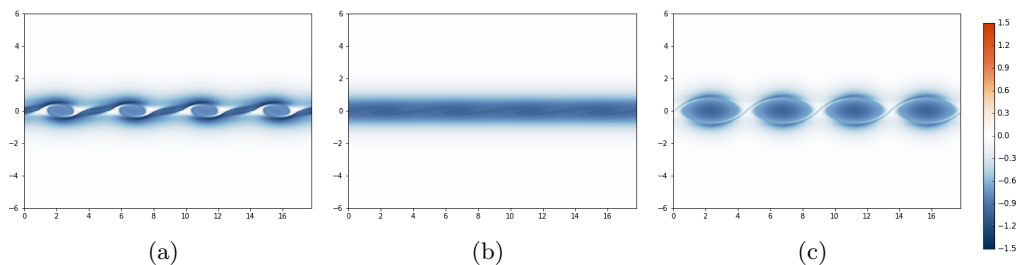


Figure 6: Vorticity fields of the mode-2 ($k = k_2$) steady states at $Pr = 7$ at the saddle-node bifurcations (a) S_1 , (b) S_2 , (c) S_3 .

215 of the similar curve shown in figure 3. There is also a mode-1 branch ($k = k_1$) connected
 216 to this, which links P_1 and P_2 . Between $Pr = 3$ and $Pr = 7$, this mode-1 branch
 217 collides with the fundamental mode-1 branch to give the situation in figure 3. Below
 218 $Pr = 3$, it appears that this disconnected curve closes at $Pr \approx 2.3$, though the picture
 219 is incomplete, since the behaviour of the states at high amplitude is unknown. The most
 220 natural explanation would be that the k_2 branch is a closed loop, but no evidence of
 221 this has been found up to amplitudes for which the finite vertical domain size becomes
 222 important and obscures the results.

223 4. Direct numerical simulations

224 As mentioned in §2, the equations (2.4) are an approximation for large but finite Re ,
 225 which ignores the fact that the background profiles diffuse. This is not a problem for
 226 rapidly changing perturbations to the background flow, but many of the connections
 227 between the steady states found in §3 appear to be very slow dynamically. In particular,
 228 although the KHI grows rapidly from small disturbances to the background, it took
 229 exceptionally long time integrations, of non-dimensional times an order of magnitude
 230 larger than Re , before the billow states were steady enough for the Newton iteration to
 231 converge on the stable states. For this reason, it is unwise to draw conclusions about the
 232 unforced system directly from the results of §3. The steady states of the forced system
 233 do not correspond to steady states in the unforced system, and a bifurcation analysis in
 234 the same way is not possible. Therefore, we explore the behaviour of the unforced system
 235 (2.1) using (two-dimensional) direct numerical simulation.

236 Direct numerical simulations started from randomly perturbed states may follow
 237 chaotic trajectories and visit states much more spatially complex than the simple steady
 238 states discussed in §3. Therefore, a much higher resolution is required to avoid ‘ringing’

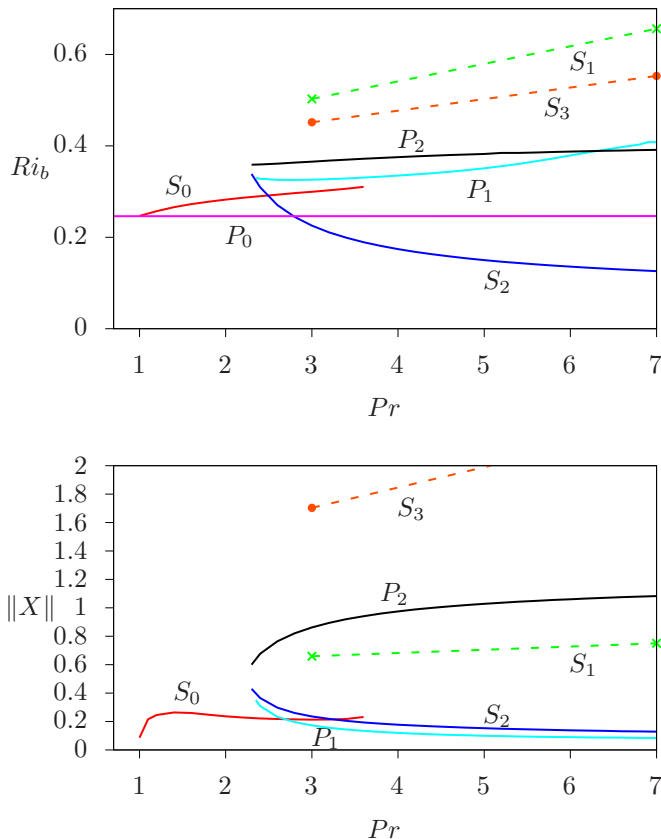


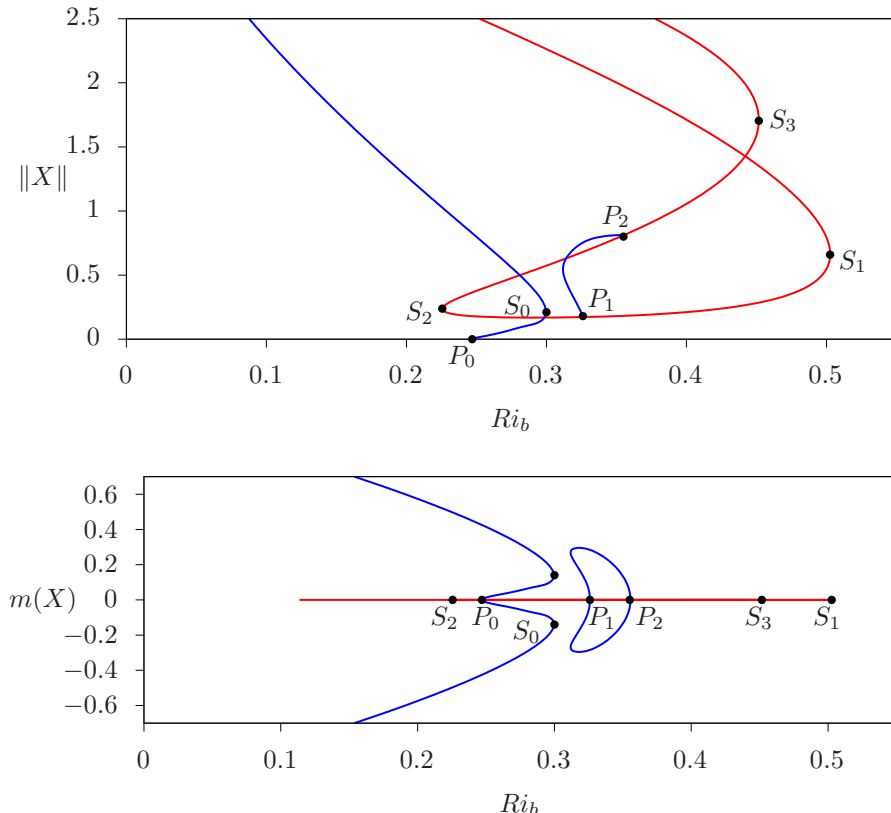
Figure 7: Tracking of the various bifurcation points shown in figures 2 and 3 as Pr varies. S_1 and S_3 were not tracked, but their locations at $Pr = 3$ and $Pr = 7$ have been marked and interpolated with dashed lines.

239 artifacts and be confident that the equations are being solved accurately. It was found to
 240 be sufficient to use 256 horizontal modes and 768 grid points vertically. All the simulations
 241 are performed at $Re = 1000$, with a domain half-height $L_z = 10$, in agreement with the
 242 calculations of the previous section.

243 4.1. DNS of exact states

244 We directly compare DNS of states found in §3, with and without the background
 245 forcing and an additional perturbation. Our aim is to determine how much the forcing
 246 affects the dynamics, rather than a complete characterisation of the dynamics without
 247 forcing. Therefore, we concentrate on one choice of parameters, for which we have a
 248 number of interesting exact states, $Pr = 7$ and $Ri_b = 0.3$. We initialise the flows with
 249 the $k = k_1$ and $k = k_2$ states at $Ri_b = 0.3$ which both have $\|X\| \approx 0.75$. To these we add
 250 a random perturbation of energy $\frac{1}{2}\|X\|^2 = 0.001$.

251 The results are shown in figures 9-12, as well as the supplemental movies. As expected,
 252 the forced, unperturbed simulations (figures 9c-12c) show perfectly steady states. With-
 253 out the artificial forcing (figures 9a-12a), the states gradually decay, with only slow
 254 changes in form. This suggests that the dynamics of the forced system are, in some

Figure 8: As for figure 3, but with $Pr = 3$.

255 sense, orthogonal to the diffusion of the background flow. When a perturbation is added
 256 to the k_2 state, chaotic behaviour develops in both the unforced (figures 11b and 12b)
 257 and forced (figures 11d and 12d) cases. This takes the form of a k_1 billow, though of a
 258 significantly higher amplitude than the k_1 steady state. This is an example of a ‘billow
 259 pairing’ subharmonic instability (Winant & Browand 1974; Klaassen & Peltier 1989).
 260 In the perturbed simulations of the k_1 steady state, there is no such energetic activity,
 261 suggesting that the state is fairly stable. A linear stability analysis shows that it is in fact
 262 weakly unstable, perhaps explaining why, in the unforced case, a k_2 billow is beginning
 263 to develop at the end of the $t = 100$ time window. Overall, good agreement between
 264 the forced and unforced cases is observed, and the differences can be attributed to the
 265 obvious decay of energy, as well as the random nature of the perturbations.

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4.2. DNS of random initial conditions

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In the previous subsection, the initial conditions in the unforced simulations were billow
 structures, so it is no surprise that billows are observed later in the simulations. However,
 from those results, it is not clear that KH billows can develop ‘naturally’ (i.e. from random
 perturbations of sufficient amplitude) in the subcritical regions of parameter space, in
 what might be called a nonlinear KH instability. Therefore, here we additionally perform
 DNS using completely random, large-amplitude perturbations to the one-dimensional
 background flow.

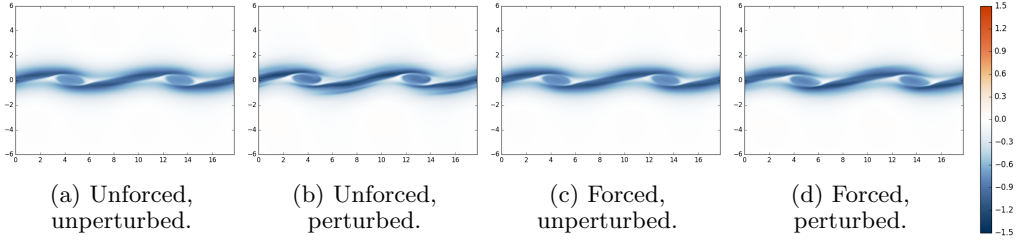


Figure 9: Total vorticity field of simulations at time $t = 20$ for the k_1 exact state at $Pr = 7$, $Ri = 0.3$.

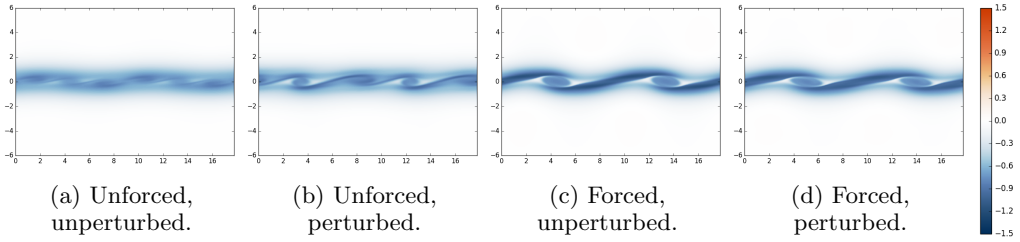


Figure 10: Vorticity at $t = 100$ for the k_1 exact state at $Pr = 7$, $Ri = 0.3$.

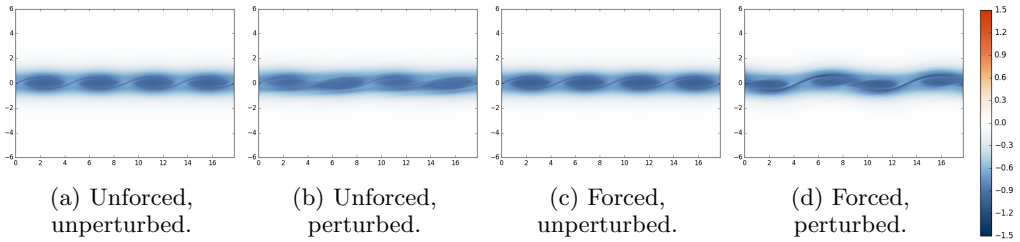


Figure 11: Vorticity at $t = 20$ for the k_2 exact state at $Pr = 7$, $Ri = 0.3$.

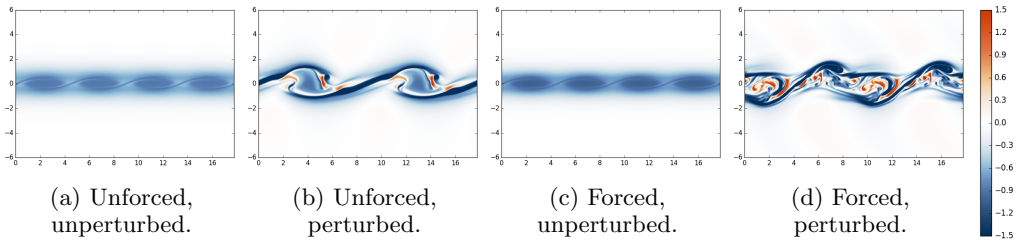


Figure 12: Vorticity at $t = 100$ for the k_2 exact state at $Pr = 7$, $Ri = 0.3$.

274 Eight different simulations were performed. We study the cases of $Pr = 0.7$ and
 275 $Pr = 7$, modelling air and water; $Ri_b = 0.1$ and $Ri_b = 0.3$ for the supercritical
 276 and subcritical regions; and initial disturbance wavenumbers k_1 , for which the linear
 277 instability is approximately maximised, and k_2 , for which no linear instability is predicted
 278 but for which we found nonlinear steady states. The simulations of equations (2.1) are
 279 started from the Drazin model plus a random perturbation,

$$\mathbf{u} = \tanh z \mathbf{e}_x + \mathbf{u}', \quad b = z + b', \quad (4.1)$$

280 where the perturbation $X = (\mathbf{u}', b')$ has components only in the first 42 Fourier modes

281 horizontally (with even-numbered modes only for k_2) and first four Hermite polynomials
 282 vertically, as in Parker *et al.* (2020). This perturbation is entirely random, and does not
 283 correspond to the modes found by bifurcation analysis, except insofar as the streamwise
 284 wavelength of the disturbances are the same, as they are required to be by the periodic
 285 boundary conditions imposed on all domains considered here. The initial perturbations
 286 are scaled to have amplitude $\|X\| = 0.3$, a relatively large disturbance, which is signifi-
 287 cantly greater than that of the lowest branch of states in figure 3, and therefore should
 288 be sufficient to push the dynamical system out of the basin of attraction of the laminar
 289 background flow. Due to the random nature of the initial conditions, it is possible that
 290 no instability is detected even when the parameters are favourable. The results presented
 291 here represent a single realisation of the random initial conditions, and since reasonable
 292 agreement was found with our bifurcation results, no attempt has been made to more
 293 systematically sample the possible results.

294 The relative phases and amplitudes of the individual Fourier modes within the initial
 295 conditions are likely to have a significant impact on which structures ultimately develop,
 296 in a situation such as that at $Pr = 7$, where several different steady states are known to
 297 exist in the forced model. One particular consequence of choosing the initial conditions in
 298 this way is that the random perturbation in general adds a mean streamwise velocity to
 299 the flow, so that billows appear to propagate through the domain. These do not represent
 300 intrinsically moving structures, but are merely a symmetry of the system which was
 301 suppressed in the previous section.

302 For perturbations with $k = k_2$ at $Pr = 0.7$, no significant nonlinear behaviour was
 303 observed at either value of Ri_b . Figures 13a and 13c both show S-shaped vorticity streaks
 304 characteristic of the transient, linear Orr mechanism at $t = 20$. By $t = 100$, as shown
 305 in figures 14a and 14c, these have diffused away to give simple shear layers, which are
 306 slightly asymmetric due to the random nature of the initial perturbations. These results
 307 are unsurprising, since no linear instability exists at this wavelength and we did not
 308 detect any nonlinear modes at this Pr either.

309 For perturbations with $k = k_1$ at $Pr = 0.7$, long-lived, nonlinear billow structures
 310 are observed at both $Ri_b = 0.1$ (figures 13b and 14b) and $Ri_b = 0.3$ (figures 13d and
 311 14d). The former is to be expected since a linear instability exists, but the latter is more
 312 surprising, as the base flow is linearly stable and the results of §3 show the bifurcation to
 313 be a simple supercritical one. The existence of a finite amplitude steady state in the forced
 314 model should be expected to imply nontrivial dynamics in the unforced simulations, but
 315 the converse is not necessarily true. We speculate further on this case in §5.

316 The $k = k_2$ simulation at $Pr = 7$ and $Ri_b = 0.3$ shows what we believe to be the
 317 most novel result reported here, namely that Kelvin-Helmholtz-like billows can exist in
 318 domains too narrow to support a linear instability. Figures 15c and 16c show the slow
 319 development of a higher amplitude state, which is very similar to the exact solution
 320 shown in figure 6a. Figure 15a with $Ri_b = 0.1$ appears to show only the results of the Orr
 321 mechanism on the initial perturbation, but by $t = 100$ shown in figure 16a one can just
 322 discern a long-lived, low-amplitude structure which is reminiscent of the lower branch of
 323 solutions found in §3, as shown in figure 6b.

324 Figures 15b and 16b show the large billow which develops at $Pr = 7$ and $Ri_b = 0.1$.
 325 This is despite the fact that we also found steady states with double this wavenumber in
 326 the forced model, but since all the states we found at these parameters were unstable,
 327 it is difficult to draw conclusions. Similarly at $Ri_b = 0.3$ in figures 15d and 16d, a small
 328 billow of wavenumber k_1 is observed. It could be the case that the initial perturbation
 329 determines whether a mode-1 or mode-2 structure develops in the wider domain, since

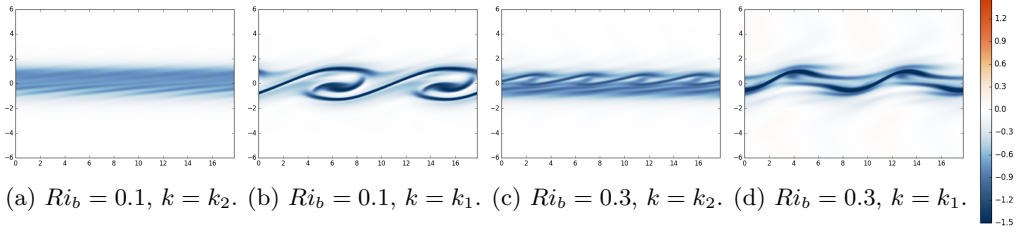


Figure 13: Total vorticity field of the unforced flow at time $t = 20$ for the Drazin model plus a random perturbation. Parameter values: $Re = 1000, Pr = 0.7$.

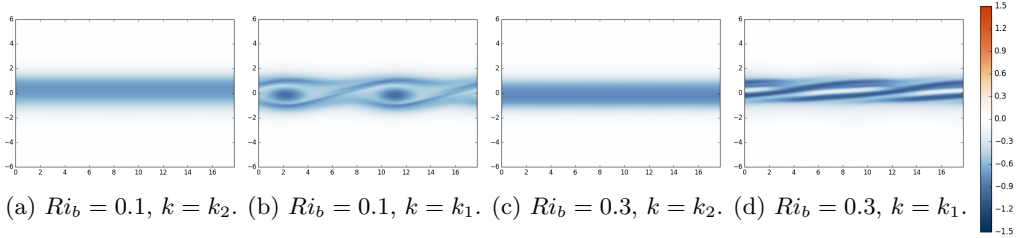


Figure 14: Vorticity at $Re = 1000$ and $Pr = 0.7$ at $t = 100$.

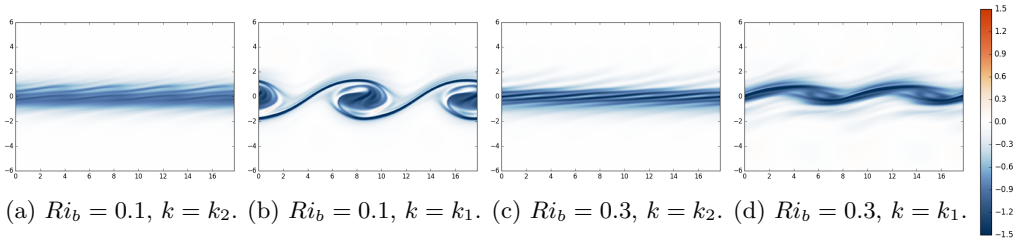


Figure 15: Vorticity at $Re = 1000$ and $Pr = 7$ at $t = 20$.

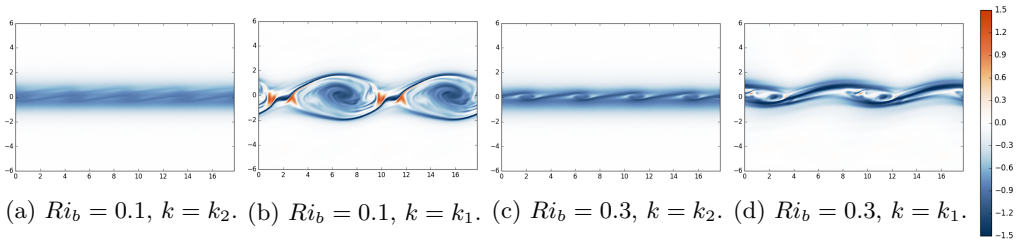


Figure 16: Vorticity at $Re = 1000$ and $Pr = 7$ at $t = 100$.

330 the initial amplitude is rather large and the results are noisy, or this could be evidence
 331 that the mode-1 structure is, in some sense, more stable.

332 Since in this unforced version the background flow diffuses away, the energy in the
 333 perturbation to this background, i.e. the energy in the billow states, is also expected to
 334 diffuse away. Figures 17 and 18 show the evolution of the total energy of the perturbation
 335 $\frac{1}{2}\|X\|^2$ for these simulations. Though in several cases there is an initial growth of energy
 336 before it decreases, there is no one clear energy level or steady state to which the state is
 337 attracted, and so direct comparison with the amplitudes on the bifurcation diagrams in
 338 section 3 is not fruitful. The k_1 simulations show wavy lines at large energy, in agreement

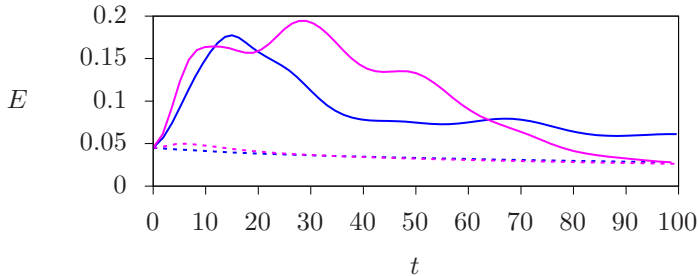


Figure 17: Perturbation energy from unforced DNS at $Pr = 0.7$, as depicted in figures 13 and 14. Blue: $Ri_b = 0.1$, pink: $Ri_b = 0.3$. Solid: $k = k_1$, dashed: $k = k_2$.

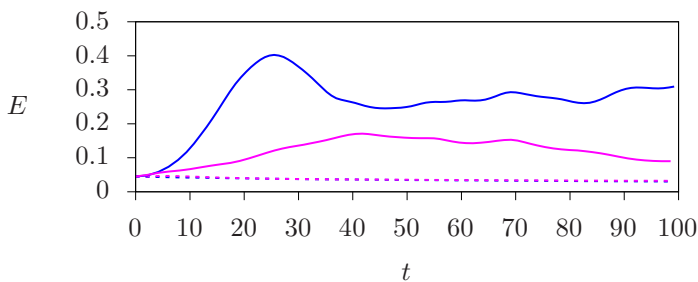


Figure 18: Perturbation energy from unforced DNS at $Pr = 7$, as depicted in figures 15 and 16. Blue: $Ri_b = 0.1$, pink: $Ri_b = 0.3$. Solid: $k = k_1$, dashed: $k = k_2$.

339 with the simulations in §4.1, for which chaotic k_1 billows were found – in that case
 340 triggered by perturbing k_2 exact states. The simulations restricted to k_2 instead show
 341 slow decay, regardless of whether long-lived billow states develop or not, indicating that
 342 the clear k_2 structures visible in the simulations are potentially less physically relevant
 343 than the k_1 structures.

344 Movies of all eight of these simulations are available in the supplementary material.
 345 In the movies, a clear distinction is visible between the strongly unstable cases, with
 346 $k = k_1$, for which the initial billow growth leads to energetic and chaotic behaviour, and
 347 the remaining cases, for which the initial structures, if they develop at all, merely diffuse
 348 away without any strong overturning. We note again that in some situations the billows
 349 are observed to propagate through the domain; this is not evidence of a Holmboe wave
 350 type instability with phase speed significantly different from the mean flow speed, but
 351 rather a consequence of the large amplitude initial perturbation having a net effect on
 352 the mean flow.

353 5. Conclusion

354 This paper presents a systematic study of the nonlinear behaviour of the Drazin
 355 model of a two-dimensional finite Reynolds-number stratified shear layer - a hyperbolic
 356 tangent shear and constant density gradient - at three different values of Pr , using
 357 both the tracking of exact coherent structures in the forced system and direct numerical
 358 simulations of the forced and unforced systems.

359 In the $Pr = 0.7$ case, we found a simple, supercritical pitchfork bifurcation, with the

360 resulting steady-state Kelvin-Helmholtz billows increasing in amplitude as (minimum)
 361 Richardson number is decreased, so far as we could track them. This agrees with weakly-
 362 nonlinear results which predict a supercritical bifurcation for $Pr < 1$. Despite the fact
 363 that we have found no finite amplitude steady states at $Ri_b > 1/4$ when $Pr = 0.7$, the
 364 unforced simulations of §4 showed clear nonlinear billow-like structures at $Ri_b = 0.3$.
 365 This could mean that there are additional steady states which are either connected to
 366 the primary instability by a bifurcation of the upper branch, or disconnected, perhaps
 367 through a homotopic continuation of the disconnected states found at $Pr = 3$ (see figure
 368 8). It could also be the case that these structures appear on trajectories which do not
 369 have an associated steady state, but rather represent an excitable system, for which the
 370 base state is stable but fast/slow dynamics nevertheless allow rapid transient growth.
 371 The observation of this structure means we are unable to state categorically whether
 372 significant nonlinear behaviour – which could lead to turbulence and mixing in the three-
 373 dimensional case – is likely to occur for $Ri_b > 1/4$ in gases, although these results and
 374 the work of Kaminski *et al.* (2017) are highly suggestive that there is more to discover
 375 at $Pr \lesssim 1$.

376 We observed a strongly subcritical pitchfork bifurcation in the flow modelling water
 377 with $Pr = 7$, as expected from the weakly-nonlinear predictions. Significantly, states
 378 were found to exist well above $Ri_b = 0.5$. Moreover, the fact that the mode-1 structure
 379 bifurcates in a superharmonic instability into a hitherto-unknown mode-2 structure
 380 implies that billow structures exist at wavelengths which are linearly stable. In section
 381 4, we demonstrated good agreement between the forced model used for the bifurcation
 382 diagrams, and an unforced model, which may be seen as more realistic for geophysical
 383 flows (the other approximations notwithstanding). In particular, we observed that random
 384 initial conditions can trigger both k_1 and k_2 billows at both $Ri_b = 0.1$ and $Ri_b = 0.3$.
 385 These results clearly indicate that in oceanic flows, the Miles-Howard criterion for linear
 386 stability does not preclude complicated mixing dynamics on times short compared to
 387 viscous diffusion.

388 The transition between $Pr = 0.7$ and $Pr = 7$ was studied in the forced model, to
 389 understand how the structures relate to one another. $Pr = 1$ and $Pr = 3$ both show
 390 the primary branch of billow states to be a simple subcritical one, but at $Pr = 3$,
 391 disconnected mode-1 states were also found, connecting to the mode-2 states at $Pr = 7$,
 392 and apparently disappearing below $Pr = 2.3$. Increasing the Prandtl number above 3,
 393 the disconnected mode-1 branch collides at some point (< 7) with the primary mode-1
 394 branch to fundamentally change the mode-1 solution topology. Given this microcosm of
 395 behaviour, it is entirely plausible that (a) further loops of mode-1 solutions exist off the
 396 mode-2 branch and survive down below $Pr \approx 2.3$ as well as (b) the mode-2 branch itself
 397 reaches to much lower Pr . In fact, it is not inconceivable that the mode-2 branch exists
 398 at $Pr = 1$ but is not at all connected to the primary mode-1 branch of Kelvin-Helmholtz
 399 instability tracked in Parker *et al.* (2019).

400 The results presented here add to a body of literature considering the dependence on
 401 Pr of the behaviour of KHI, with possible consequences in oceanographic applications.
 402 Previous authors have found that mixing efficiency decreases with Pr when Re and Ri_b
 403 are kept fixed; Brucker & Sarkar (2007) showed this for a DNS initialised with turbulence
 404 and Salehipour *et al.* (2015) for an idealised KH billow. No clear reason for this is known,
 405 though it has been suggested it could be attributed to higher stratification near the
 406 centreline, reduced lengthscales, or higher isotropy, as Pr is increased. The existence of
 407 the $k = k_2$ structures we have found at higher Pr is further evidence of these reduced
 408 length scales, in addition to shorter wavelength secondary instabilities documented by
 409 Salehipour *et al.* (2015).

410 It should be clear that there are numerous natural extensions to the present study. It
411 would be of interest to see how the results vary with Re , as $Re = 1000$ is much lower
412 than in geophysically relevant flows. It is assumed that if complex behaviour exists at
413 $Re = 1000$ for given Pr and Ri_b , it will also do so for higher Re - in Parker *et al.*
414 (2019) it was shown that increasing Re corresponds to an increase in the maximum Ri_b
415 of steady states, at least for $Pr = 1$. Much higher values of Pr , as would be relevant
416 to salt-stratified water, could also be an area for future study. Our results suggest that
417 the dynamics only get more complex with increasing Pr , and higher Ri_b can give rise to
418 steady states. Increasing either Re or Pr significantly would require a finer discretisation
419 of the domain, necessitating either much more computational resources or a different
420 strategy from that pursued here.

421 We focussed on the case of a fixed domain size corresponding to one wavelength of the
422 most unstable mode at $Ri_b = 1/4$ (see figure 1). This leaves the possibility of different
423 behaviour at different wavelengths, but also more importantly ignores the interplay of
424 different wavelengths of instability with one another. The subharmonic ‘pairing’ insta-
425 bility of KH billows is widely documented in laboratory experiments and computational
426 simulations, and has not been studied here as the behaviour cannot be captured in our
427 short domain. Previous authors (Mashayek & Peltier 2011; Salehipour *et al.* 2015) have
428 demonstrated that such subharmonic merging instabilities are suppressed at sufficiently
429 high Re , which may explain why they are not observed in geophysical applications.
430 The short domain size also means we capture only one discretised unstable wavelength
431 rather than a range, and there could be significant interaction between these, leading
432 to important dynamics (see, for example, Scinocca & Ford (2000)). This gap between
433 idealised simulations of single KH billows and the messy turbulence seen in GFD settings
434 and larger DNS studies remains an important area for future research.

435 Even at the parameters we studied, much remains unclear. To what other states
436 do the secondary bifurcations give rise? Hopf bifurcations were detected, so periodic
437 orbits as well as steady states are expected. What new dynamics does a third, spanwise
438 dimension add to the flow? Certainly all two-dimensional states we have found will exist
439 in three dimensions, but many more secondary instabilities will exist and we expect
440 those states found to be stable in two dimensions to become unstable in three. From
441 direct numerical simulations, three-dimensional flows prone to primary Kelvin-Helmholtz
442 instability are known to behave very differently, quickly breaking down into turbulence,
443 without long-lived coherent billows; most of the mixing associated with KHI is due to
444 this billow breakdown in three dimensions. There is no guarantee that the states we have
445 found in two dimensions will be sufficiently stable to be realisable in three dimensions.
446 Nevertheless, the existence of the structures implies the possibility for complex behaviour
447 and mixing in geophysical flows at these parameters even if billows do not directly
448 develop.

449 6. Acknowledgements

450 J.P.P. was supported by an EPSRC DTA studentship. The data used to generate figures
451 2, 3, 7 and 8 is available at <https://doi.org/10.17863/CAM.64232>. The DNS and
452 continuation source code can be found at <https://github.com/Jezz0r/Stratiflow>.

453 7. Declaration of Interests

454 The authors report no conflict of interest.

REFERENCES

- 455 BROWN, S. N., ROSEN, A. S. & MASLOWE, S. A. 1981 The evolution of a quasi-steady critical
456 layer in a stratified viscous shear layer. *Proc. Roy. Soc. Lond. A* **375** (1761), 271–293.
- 457 BRUCKER, KYLE A & SARKAR, SUTANU 2007 Evolution of an initially turbulent stratified shear
458 layer. *Phys. Fluids* **19** (10), 105105.
- 459 CAULFIELD, C. P. & PELTIER, W. R. 2000 The anatomy of the mixing transition in
460 homogeneous and stratified free shear layers. *J. Fluid Mech.* **413**, 1–47.
- 461 CHURILOV, S. M. & SHUKHMAN, I. G. 1987 Nonlinear stability of a stratified shear flow: a
462 viscous critical layer. *J. Fluid Mech.* **180**, 1–20.
- 463 DRAZIN, P. G. 1958 The stability of a shear layer in an unbounded heterogeneous inviscid fluid.
464 *J. Fluid Mech.* **4** (2), 214–224.
- 465 FUKAO, S., LUCE, H., MEGA, T. & YAMAMOTO, M. K 2011 Extensive studies of large-
466 amplitude Kelvin–Helmholtz billows in the lower atmosphere with VHF middle and upper
467 atmosphere radar. *Q. J. Roy. Meteorol. Soc.* **137** (657), 1019–1041.
- 468 GREGG, M. C., D’ASARO, E. A., RILEY, J. J. & KUNZE, E. 2018 Mixing efficiency in the
469 ocean. *Annu. Rev. Mar. Sci.* **10**, 443–473.
- 470 HOLMBOE, J. 1960 Unpublished lecture notes.
- 471 HOWARD, L. N. 1961 Note on a paper of John W. Miles. *J. Fluid Mech.* **10** (4), 509–512.
- 472 KAMINSKI, A. K., CAULFIELD, C. P. & TAYLOR, J. R. 2017 Nonlinear evolution of linear
473 optimal perturbations of strongly stratified shear layers. *J. Fluid Mech.* **825**, 213–244.
- 474 KLAASSEN, GP & PELTIER, WR 1989 The role of transverse secondary instabilities in the
475 evolution of free shear layers. *J. Fluid Mech.* **202**, 367–402.
- 476 KLAASSEN, G. P & PELTIER, W. R 1985a The effect of Prandtl number on the evolution and
477 stability of Kelvin-Helmholtz billows. *Geophys. Astrophys. Fluid Dyn.* **32** (1), 23–60.
- 478 KLAASSEN, G. P. & PELTIER, W. R. 1985b Evolution of finite amplitude Kelvin-Helmholtz
479 billows in two spatial dimensions. *J. Atmos. Sci.* **42** (12), 1321–1339.
- 480 LOTT, F. & TEITELBAUM, H. 1992 Nonlinear dissipative critical level interaction in a stratified
481 shear flow: Instabilities and gravity waves. *Geophys. Astrophys. Fluid Dyn.* **66** (1-4), 133–
482 167.
- 483 MASHAYEK, A. & PELTIER, W. R. 2011 Three-dimensionalization of the stratified mixing layer
484 at high Reynolds number. *Phys. Fluids* **23** (11), 111701.
- 485 MILES, J. W. 1961 On the stability of heterogeneous shear flows. *J. Fluid Mech.* **10** (4), 496–508.
- 486 MKHININI, N., DUBOS, T. & DROBINSKI, P. 2013 On the nonlinear destabilization of stably
487 stratified shear flow. *J. Fluid Mech.* **731**, 443–460.
- 488 PARKER, J. P., CAULFIELD, C. P. & KERSWELL, R. R. 2019 Kelvin-Helmholtz billows above
489 Richardson number 1/4. *J. Fluid Mech.* **879**, R1.
- 490 PARKER, J. P., CAULFIELD, C. P. & KERSWELL, R. R. 2020 The viscous Holmboe instability
491 for smooth shear and density profiles. *J. Fluid Mech.* **896**, A14.
- 492 RAHMANI, M., SEYMOUR, B. R. & LAWRENCE, G. A. 2016 The effect of Prandtl number on
493 mixing in low Reynolds number Kelvin-Helmholtz billows. *Phys. Fluids* **28** (5), 054107.
- 494 SALEHIPOUR, H., PELTIER, W. R. & MASHAYEK, A. 2015 Turbulent diapycnal mixing in
495 stratified shear flows: the influence of Prandtl number on mixing efficiency and transition
496 at high Reynolds number. *J. Fluid Mech.* **773**, 178–223.
- 497 SCHNEIDER, T. M., ECKHARDT, B. & YORKE, J. A. 2007 Turbulence transition and the edge
498 of chaos in pipe flow. *Phys. Rev. Lett.* **99** (3), 034502.
- 499 SCINOCCA, JF & FORD, R 2000 The nonlinear forcing of large-scale internal gravity waves by
500 stratified shear instability. *J. Atmos. Sci.* **57** (5), 653–672.
- 501 SMYTH, W. D. & CARPENTER, J. R. 2019 *Instability in Geophysical Flows*. Cambridge
502 University Press.
- 503 SMYTH, W. D. & MOUM, J. N. 2012 Ocean mixing by Kelvin-Helmholtz instability.
504 *Oceanography* **25** (2), 140–149.
- 505 SMYTH, W. D. & PELTIER, W. R. 1989 The transition between Kelvin-Helmholtz and Holmboe
506 instability: an investigation of the overreflection hypothesis. *J. Atmospheric Sci.* **46** (24),
507 3698–3720.
- 508 SUN, J., NAPPO, C. J., MAHRT, L., BELUŠIĆ, D., GRISOGONO, B., STAUFFER, D. R., PULIDO,

- 509 M., STAQUET, C., JIANG, Q., POUQUET, A. & OTHERS 2015 Review of wave-turbulence
510 interactions in the stable atmospheric boundary layer. *Rev. Geophys* **53** (3), 956–993.
- 511 THORPE, S. A. 2005 *The Turbulent Ocean*. Cambridge University Press.
- 512 VAN HAREN, H. & GOSTIAUX, L. 2010 A deep-ocean Kelvin-Helmholtz billow train. *Geophys.*
513 *Res. Lett.* **37** (3).
- 514 WINANT, CLINTON D & BROWAND, FRED K 1974 Vortex pairing: the mechanism of turbulent
515 mixing-layer growth at moderate Reynolds number. *J. Fluid Mech.* **63** (2), 237–255.