1	Estimation of sinking velocities using free-falling dynamically scaled
2	models: foraminifera as a test case
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12 Summary Statement

- 13 We developed a novel method to determine the sinking velocity of biologically important micro-
- 14 scale particles using 3D printed scale models.

15 Abstract

16 The velocity of settling particles is an important determinant of distribution in extinct and extant 17 species with passive dispersal mechanisms, such as plants, corals, and phytoplankton. Here we adapt dynamic scaling, borrowed from engineering, to determine settling velocities. Dynamic scaling 18 19 leverages physical models with relevant dimensionless numbers matched to achieve similar dynamics to the original object. Previous studies have used flumes, wind tunnels, or towed models to examine 20 21 fluid flows around objects with known velocities. Our novel application uses free-falling models to 22 determine the unknown sinking velocities of planktonic foraminifera – organisms important to our 23 understanding of the Earth's current and historic climate. Using enlarged 3D printed models of 24 microscopic foraminifera tests, sunk in viscous mineral oil to match their Reynolds numbers and drag 25 coefficients, we predict sinking velocities of real tests in seawater. This method can be applied to study 26 other settling particles such as plankton, spores, or seeds.

27 Introduction

28 The transport of organisms and biologically derived particles through fluid environments strongly 29 influences their spatiotemporal distributions and ecology. In up to a third of terrestrial plants (Willson et al., 1990), reproduction is achieved through passive movement of propagules (e.g., seeds) on the 30 31 wind. In aquatic environments, propagules of many sessile groups from corals (Jones et al., 2015) to 32 bivalves (Booth, 1983) are dispersed by ambient currents, eventually settling out of the water column 33 to their final locations. Furthermore, most dead aquatic organisms (from diatoms to whales) sink, 34 transporting nutrients to deeper water and contributing to long term storage of carbon (De La Rocha 35 & Passow, 2007). In the case of microfossils, sinking dynamics of the original organisms even influences our reconstructions of the Earth's paleoclimate (Van Sebille et al., 2015). Crucially, the 36 37 horizontal distances over which all these biological entities are transported, and therefore their 38 distributions, are affected by their settling velocities (Ali et al., 2011).

39 Measuring the individual settling velocities of small particles directly is challenging, especially when 40 they are too small to be imaged easily without magnification (e.g. Walsby and Holland, 2006). Here we apply dynamic scaling, an approach commonly used in engineering, to circumvent this difficulty 41 42 and accurately quantify the kinematics of sub-millimeter scale free-falling particles using enlarged 43 physical models. We use scaled-up physical models in a high-viscosity fluid, enabling easy 44 measurements of settling speed, orientation, and other parameters using inexpensive standard high-45 definition web cameras. While dynamically scaled models have previously been employed to study a 46 number of problems in biological fluid mechanics (e.g. Vogel, Ellington and Kilgore, 1973; Vogel, 1987; 47 Vogel, 1994; Koehl, 2003), the study of freely-falling particles of complex shape – for which settling speed is the key unknown parameter - presents a unique challenge to experimental design that we 48 49 overcome in this work.

50 Engineering problems such as aircraft and submarine design often are approached using scaled-down 51 models in wind tunnels or flumes to examine fluid flows around the model and the resulting fluid 52 dynamic forces it is subjected to. To ensure that the behaviour of the model system is an accurate 53 representation of real life, similarity of relevant physical phenomena must be maintained between 54 the two. If certain dimensionless numbers (i.e., ratios of physical quantities such that all dimensional 55 units cancel) that describe the system are equal between the life-size original and the scaled-down model, "similitude" is achieved and all parameters of interest (e.g., velocities and forces) will be 56 57 proportional between prototype and model (Zohuri, 2015). Intuitively, the model and real object must 58 be geometrically similar (i.e., have the same shape), so that the dimensionless ratio of any length between model and original, $Length_{model}/Length_{real}$, is constant – this is the scale factor (S) of the 59 model. Less obvious is the additional requirement of dynamic similarity, signifying that the ratios of 60 61 all relevant forces are constant. For completely immersed objects sinking steadily at terminal velocity 62 (achieved quickly for most small particles, see Time to Terminal Velocity), dynamic similarity is achieved by matching the Reynolds number (Re). 63

Re is a measure of the ratio of inertial to viscous forces in the flow (Batchelor, 2000; within a biological
context Vogel, 1994), and is typically defined as:

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$$Re = \frac{L U \rho_{fluid}}{\mu}$$

where ρ_{fluid} is density of the fluid (kg m⁻³); L is a characteristic length (m) of the object; U is the 68 object's velocity (m s⁻¹); and μ is the dynamic viscosity (N s m⁻², or Pa s) of the fluid. In cases where 69 $L U \rho_{fluid}$ is large compared to μ , e.g. whales, birds, and fish ($Re \approx 3 \times 10^9$ - 3×10^6 , Vogel, 1994), 70 71 inertial forces dominate. In cases where $L U \rho_{fluid}$ is relatively small compared to μ , e.g. sperm, bacteria (Re \approx 3×10⁻² - 1×10⁻⁵), viscous forces dominate. Finally, when L U ρ_{fluid} is of comparable 72 magnitude to μ , Re is intermediate and one cannot discount either inertial or viscous forces. If the 73 74 scaled model and original system exhibit identical Re, the relative importance of inertial versus viscous 75 forces is matched between the two and any qualitative features of the flows (e.g. streamlines) will 76 also be identical.

77 Dynamically scaled physical models exhibiting the same Re as the original systems have been used in 78 a number of biological studies. Vogel and La Barbera (1978) outline the principles of dynamic scaling: 79 to obtain the same Re when enlarging small organisms, the fluid flow must be slower and/or the fluid 80 more viscous, and when making smaller models of large organisms, the fluid flow must be faster 81 and/or the fluid less viscous. For instance, Vogel (1987) used air in place of water flowing at lower 82 speeds when investigating the refilling of the squids mantle during swimming by scaling a model up 83 1.5 times relative to the animal's actual size. More recently, Stadler et al (2016) investigated sand 84 inhalation in skinks with 3D-printed enlarged models, using helium instead of air (thereby increasing 85 viscosity) as the experimental fluid. Koehl and colleagues have studied crustacean antennule flicking 86 (lobsters (Reidenbach et al., 2008), mantis shrimp (Stacey et al., 2002) and crabs (Waldrop et al., 87 2015)) as well as the movements of copepod appendages (Koehl, 1995) with enlarged models, using mineral oil in place of water. Finally, perhaps the largest change in scale was employed by Kim et al 88 89 (2003), who modelled the bundling of E. coli flagella at a scale factor of ~61,000, submerged in silicone

90 oil (10⁵ times more viscous than water), and rotated at 0.002 rpm compared to the 600 rpm observed

91 in real bacteria (Sowa & Berry, 2008).

92 In all the above studies, basic kinematics such as speeds in the original system were relatively easy to 93 measure, and the experiments aimed to reveal the forces involved (e.g. hydrodynamic drag) or details 94 of the fluid flow such as the pattern of streamlines. Since the representative speed U of the original 95 system was known, designing experiments to achieve similitude was relatively straightforward 96 because the Re was also known a priori – in these cases, the model size, speed, and working fluid 97 properties were simply interrelated through Re (Eqn 1). For instance, once a working fluid and the 98 model size were chosen, the required towing speed was obvious. However, in the case of 99 sedimentation of small particles (e.g. spores, seeds, plankton), the sinking speed (U) is the key 100 unknown. With an unknown sinking speed, the operating Re is also unknown, so it is not 101 straightforward to design experiments that achieve similitude with the original system. Here we 102 present an iterative methodology leveraging 3D printed dynamically scaled models that allows 103 determination of the sinking speeds of small objects of arbitrarily complex shape.

104 We use planktonic foraminifera as an example of a small ($200 - 1500 \,\mu$ m) biological particle for which 105 the settling velocity is important and typically unknown. Foraminifera are a phylum of marine ameboid 106 protists (B. K. Sen Gupta (ed.), 2002; Schiebel & Hemleben, 2005). By secreting calcium carbonate, 107 foraminifera produce a multi-chambered shell (test) which, in planktonic foraminifera, can grow up to 108 1500 µm in diameter, and which frequently exhibits complex shape (Table 1). Once the organism dies 109 or undergoes reproduction, the empty test sinks to the ocean floor, and so oceanic sediments contain 110 substantial numbers of foraminifera tests. Foraminifera account for 23-56% of the oceans production 111 of carbonate (CO₃) (Schiebel, 2002), an important factor in climate change models (Passow & Carlson, 2012). Of particular interest for climate predictions is calculating the flux of tests reaching the ocean 112 floor (Schiebel, 2002; Jonkers & Kučera, 2015). While there are more than 30 extant species and over 113 114 600 species in the fossil record, settling velocities are known for only 14 species of foraminifera (Fok-115 Pun & Komar, 1983; Takahashi & Be, 1984; Caromel et al., 2014, $3.41 \times 10^{-4} - 6.8 \times 10^{-2}$ m s⁻¹, Re $\approx 18 - 10^{-2}$ m s⁻¹ 55). 116

117 Materials and Methods

118 Similitude and Settling Theory

119 We assume that the size (i.e., L – defined as the maximum length parallel to the settling direction, A120 - defined as the projected frontal area, and V - the particle volume not including any fluid-filled 121 cavities), 3D shape (Ψ , here treated as a categorical variable due to our consideration of arbitrarily 122 complex morphologies, see Table 1), and density ($\rho_{particle}$) of the original sinking particle are known, 123 while the sinking speed (U) is unknown. The properties of the fluid surrounding the original particle 124 (i.e. ρ_{fluid} , μ) are also known, and our goal is to design experiments in which we sink a scaled-up 125 model particle in a working fluid of known ρ_{fluid} and μ in order to determine the model particle's 126 sedimentation speed and, via similitude, U of the original particle.

127 While previous work (Berger et al., 1972; Fok-Pun & Komar, 1983; Takahashi & Be, 1984; Caromel et 128 al., 2014) suggests that the *Re* of sinking forams should be $10^{0} - 10^{2}$, the exact value of *Re* for 129 morphology Ψ is assumed to be unknown. Hence, it is not immediately clear what size the model 130 should be (i.e., the scale factor $S = L_{model}/L_{real}$) in order to match this *Re* in the experiments and 131 ensure similitude. Solving for both *Re* and *S* simultaneously requires additional mathematical 132 relationships beyond Eqn 1.

133 Throughout, we use a superscript O to refer to the *Original* values of dimensioned variables at life size 134 (e.g., L^{O} , V^{O} , A^{O} , $\rho_{particle}^{O}$, U^{O}) and Re^{O} , C_{D}^{O} for the values of the dimensionless Reynolds number 135 and drag coefficient (defined below) corresponding to real particles sinking in the original fluid (e.g., 136 seawater of ρ_{fluid}^{O} , μ^{O}). While the fluid dynamics of flow around a particle of particular shape Ψ can 137 be considered theoretically over a range of Re, only the dynamics at Re^{O} and C_{D}^{O} will represent the 138 "operating point" corresponding to the life size particle settling speed U^{O} .

When a particle is sinking steadily at its terminal velocity, the sum of the external forces acting on the particle is zero (Eqn 2); that is, the upward drag force (F_{drag} , Eqn 3) and buoyant force ($F_{buoyancy}$, Eqn 4) must balance the weight of the particle (F_{weiaht} , Eqn 5):

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$$\Sigma F = F_{drag} + F_{buoyancy} - F_{weight} = 0$$

144 where

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 $F_{drag} = \frac{1}{2} C_D(\Psi, Re) \rho_{fluid} U^2 A$

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Eqn 3 introduces the drag coefficient $C_D(\Psi, Re)$, a dimensionless descriptor of how streamlined an 147 object is. Both C_D and Re must be matched to achieve similitude. C_D depends on the shape of the 148 object Ψ , including its orientation relative to the freestream flow – for instance, C_D of a flat plate 149 150 oriented parallel to laminar flow is as low as 0.003 while C_D of a flat plate oriented perpendicular to the flow is ~ 2.0 (Munson et al., 1994). However, in addition to object geometry, C_D also depends on 151 qualitative characteristics of the flow, such as whether it is laminar or turbulent – that is, C_D also 152 153 depends on Re. C_D of a sphere decreases from about 200 at Re = 0.1 to about 0.5 at Re = 1000; C_D 154 generally decreases with Re for most shapes (Munson et al., 1994; Morrison, 2013). While C_D does 155 not depend on object size directly, larger objects generally experience higher drag forces and this is 156 captured by the inclusion of particle area (A) in the expression for F_{drag} (Eqn 3). For brevity, we will 157 omit Ψ hereafter and write the drag coefficient as $C_D(Re)$.

The buoyant force ($F_{buoyancy}$, Eqn 4) and weight (F_{weight} , Eqn 5) are both expressed using particle volume (V), gravitational acceleration (g), and density of the fluid (ρ_{fluid}) or particle ($\rho_{particle}$), respectively: 161

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$$F_{weight} = V \rho_{particle}$$

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165 expression for the drag coefficient obtained through a force balance (indicated by a superscript \mathcal{F}): 166

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$$C_D^{\mathcal{F}}(Re) = \frac{2\left(\rho_{particle} - \rho_{fluid}\right) V g}{\rho_{fluid} U^2 A} = \left(\frac{2\left(\rho_{particle} - \rho_{fluid}\right) V g}{\rho_{fluid} A}\right) \left(\frac{\rho_{fluid} L}{Re \mu}\right)^2 = \frac{2\left(\rho_{particle} - \rho_{fluid}\right) \rho_{fluid} V g L^2}{A Re^2 \mu^2}$$

169 Note that this expression can be simplified further upon identification of the dimensionless Archimedes number $Ar = g L_{Ar}^3 \rho_{fluid} (\rho_{particle} - \rho_{fluid})/\mu^2$ if the cubed length scale $L_{Ar}^3 =$ 170 VL^2/A , yielding $C_D^{\mathcal{F}}(Re) = Ar/Re^2$, as previously highlighted by others (e.g. Karamanev, 1996). 171 172 However, we will proceed with the original form of Eqn 6 to keep key variables such as L explicit.

If C_D were known for a particular morphology, we could simply substitute values corresponding to the 173 original test in seawater into Eqn 6 and solve for $Re = Re^{O}$ and thus U^{O} via Eqn 1, immediately solving 174 the problem of unknown settling speed. Unfortunately, the complex shapes of foraminifera (Table 1) 175 coupled with the implicit dependence of C_D on Re means that both variables are generally unknown, 176 and thus far we have only one constraining relationship between C_D and Re. More information is 177 required to determine where along this constraint curve the operating C_D^0 and Re^0 are located. This 178 information can come from experiments in which the sinking speeds of scaled-up model particles of 179 180 various sizes (i.e., scale factors S) in a viscous fluid are measured directly, allowing us to calculate Revia Eqn 1 and then $C_D^{\mathcal{F}}(Re)$ via Eqn 6 for the models, with appropriate values substituted for each 181 experiment. For clarity, we can rewrite Eqn 6 for a model in terms of S and the original test parameters 182 (L^{O}, A^{O}, V^{O}) : 183

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$$C_{D}^{\mathcal{F}}(Re) = \frac{2\left(\rho_{particle} - \rho_{fluid}\right)\rho_{fluid}(V) g (S L^{0})^{2}}{(S^{2} A^{0}) Re^{2} \mu^{2}}$$

where we use the fact that for a model, $L = S L^0$ and $A = S^2 A^0$. While one would also expect V =186 $S^3 V^0$ for 3D printed models, limitations of our 3D printer led to variation in V that we overcame using 187 a more general empirical relationship between S and V based on mass measurements – see 3D Printer 188 Limitations. Eqn 7 represents a constraining relationship between C_D and Re for the sinking particle, 189 which we use to collect (Re, C_D) experimental data points at several S. Once sufficient data are 190 191 collected, we can construct a new, empirical relationship (e.g., a cubic spline fit) between C_D and Refor a particular particle shape, which we term $C_D^E(Re)$. Finally, we can solve for the operating Re^0 , 192 C_D^O , and U^O by finding the intersection point between the $C_D^{\mathcal{F}}(Re)$ constraint curve specific to life-size 193 particles sinking in seawater (i.e., Eqn 7 with S = 1 and $\rho_{particle}^{0}$, ρ_{fluid}^{0} , μ^{0}) and our empirical $C_{D}^{E}(Re)$ 194

- 195 spline curve valid for a particular particle shape moving steadily through any fluid. MATLAB code can
- 196 be downloaded from https://github.com/matthewwalkerbio/Dynamic-scaling.

197 Study Species

To construct an empirical $C_D^E(Re)$ curve for a particular test morphology, we started with 3D scans of 198 199 individual specimens from 30 different species (Table 1). The majority of the species were selected 200 from the University of Tohoku museum's database, eforam Stock 201 (http://webdb2.museum.tohoku.ac.jp/e-foram/), with a micro computed tomography (µCT) scan 202 resolution between 2.5 and 3.6 pixels per µm, and were exported as 3D triangular mesh (STL format) 203 files. Specimens of an additional three species were scanned using synchrotron radiation based 204 micro-computed tomography (SRµCT). Imaging was performed at the Imaging Beamline P05 (IBL) (Greving et al., 2014; Haibel et al., 2010; Wilde et al., 2016) operated by the Helmholtz-Zentrum-205 206 Geesthacht at the storage ring PETRA III (Deutsches Elektronen Synchrotron – DESY, Hamburg, 207 Germany). Specimens were imaged at an photon energy of 14 keV and with a sample to detector 208 distance of 17 mm. For each tomographic scan, 900 projections at equal intervals between 0 and 209 π were recorded. Tomographic reconstruction was done via a classical filtered back projection using 210 the RECLBL library (Huesman et al., 1977). For processing, raw projections were binned two times 211 resulting in an effective pixel size of the reconstructed volume of 1.44 μ m. These scans were segmented and rendered using SPIERS (Sutton et al., 2012), and again exported in STL format and are 212 213 avaliable from Morpho Source. Meshes of all foraminifera were manually checked in Meshlab (Callieri 214 et al., 2012) for integrity.

For species where more than one scan was available, the scan that contained the best-preserved specimen was chosen. By only including one specimen per species, this approach neglects phenotypic plasticity which is demonstrated in planktonic foraminifera (e.g. Lohmann, 1983; Morard *et al.*, 2013), but was chosen due to limitations of µCT scan availability and time constraints on the project.

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220 3D printing and model preparation

221 The 3D scans allowed us to easily fabricate scaled-up (scale factor S) physical models of each specimen 222 using a FormLabs Form1+ (Formlabs, Somerville, Massachusetts, USA) 3D printer, using FormLabs 223 Clear Resin Version 2 (Formlabs, Somerville, Massachusetts, USA) with a layer thickness of 50 µm (see 224 Error! Reference source not found.D for examples) and x-y resolution of 200 µm. Models were 225 washed and flushed with isopropanol to remove excess resin following Formlabs' guide and allowed 226 to air dry. Support material was removed (Error! Reference source not found.D.vi), and the models 227 lightly sanded with 400 grit Wet 'n' Dry paper, followed by a final isopropanol wash to remove any 228 remaining residue. Once dry, models were filled with mineral oil in preparation for sinking. Clear resin 229 was chosen to allow each model to be checked for bubbles (which would increase the buoyancy of 230 the model). Any bubbles were removed using a 30-gauge needle and syringe.

Following convention when defining the area $A_{particle}$ used in the definition of C_D (Eqn 3), we 231 232 measured projected area of the sinking foraminifera. Referring to high resolution images of the sinking model (Error! Reference source not found.D), a digital model of the foraminifera was manually 233 234 aligned to measure the projected area in a plane perpendicular to the sinking direction (Error! 235 Reference source not found.D). We used the same procedure to measure the maximum length 236 parallel to the flow (L) for the calculation of Re (Error! Reference source not found.D). These choices facilitated objective comparisons of C_D across morphologically diverse species, to be detailed in a 237 future study. 238

239 3D Printer Limitations

Whilst in principle, the volume of a printed model should simply scale according to $V = S^3 V^0$, due to 240 241 inherent limitations of the 3D printer as well as difficulty in removing excess resin from small models, we found that this expectation was usually not satisfied, and weighing the models showed that 242 $M/\rho_{particle} > S^3 V^0$ where M is particle mass (Error! Reference source not found.C). Therefore, we 243 estimated V of each model by weighing on an Entris 224-1S mass balance (±0.001 g) and assuming 244 245 $\rho_{particle}$ was 1121.43 ± 13.73 kg m⁻³, based on the average mass of five 1 cm³ cubes of printed resin. Furthermore, whenever a predicted value for V at a given scale factor S was needed, i.e. in Eqn 7 (see 246 247 Remaining iterations), we based this on cubic spline interpolation of our V(S) data for existing models when sufficient data were available, with extrapolation based on cubic scaling of V(S) if required (see 248 249 Error! Reference source not found.C):

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$$V^{predict}(S) = \begin{cases} H(S), N \ge 3\\ S^3 V^0, N < 3 \text{ or } S = 1 \end{cases}$$

where *N* is the number of existing volume measurements and *H* represents the cubic spline fit of *V* vs *S*. Note that because we always directly measured *V* by weighing after printing each model, and it is not necessary to achieve the exact *Re* and C_D of the operating point (*Re*⁰ and C_D^0) in the experiments (see Remaining iterations), the empirical spline-based volume prediction was not strictly required for our method to succeed. It merely aids in improving the rate of convergence of our iterative approach by reducing the difference between our anticipated and actual *Re*, *C*_D for each experiment.

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258 Settling tank

259 The models were released in a cylindrical acrylic tank (0.9 m in diameter and 1.2 m in height) of mineral 260 oil ("Carnation" white mineral oil, Tennants Distribution Limited, Cheetham, Manchester, UK; ρ = 830 kg m⁻³, μ = 0.022 Pa S) filled to a depth of 1.18 m (approximately 750 L). The tank was fitted with a 261 262 custom net and net retrieval system (Error! Reference source not found.A) to allow easy retrieval of the models after their descent, allowing each model to be sunk 5 times. Integrated to the net retrieval 263 264 system was the release mechanism, which was held centrally over the tank, with the grasping parts 265 submerged below the oil level. This ensured that each model was released in a controlled and 266 repeatable fashion.

267 Particle imaging

To minimise reflections, the tank was surrounded by a black fabric tent-like structure. This also served as a dark background to facilitate visualisation of the model during descent. The tank was illuminated with a single 800 lumen LED spotlight placed underneath the tank and, as the Formlabs' Clear Resin is UV-fluorescent, two 20W "Blacklight" UV fluorescent tubes were placed above the tank.

272 The sinking models were recorded using two Logitech C920 HD webcams (Logitech, Lausanne, 273 Switzerland), placed at 90° to each other (Error! Reference source not found.A) and recording at 960 274 pixels x 720 pixels and ~30 frames per second, allowing monitoring of the position and orientation of 275 the particle in 3D as it fell. As these consumer-grade webcams use a variable frame-rate system, a 276 custom MATLAB script was used to initiate camera recording, recording both frames and frame 277 timestamps. Videos were recorded for 500 frames (~17 s). Sinking velocity was calculated over a 278 central 0.8 m depth range, ensuring the model was at terminal velocity (see Time to Terminal Velocity) 279 whilst also avoiding end effects which could slow the model as it reached the bottom of the tank. 280 Based on observations of suspended dust, there was no discernable convection in the tank during any 281 trials that might potentially affect sinking velocities. The curved walls of the tank introduced distortion, 282 which was removed using the MATLAB toolbox "Camera Calibrator" (Mcandrew, 2004). Pixel size was 283 1.06 pixels per millimetre with a mean reprojection error of 0.5 pixels, therefore distance 284 measurements (for calculating sinking velocities) were accurate to within 0.5 mm (0.06% of the 285 traversed depth).

286 Velocity calculation

287 Models were tracked in distortion-corrected frames using a modified version of Trackbac (Guadayol 288 et al., 2017; Guadayol, 2016). The per-frame centroid coordinates obtained were then paired with the 289 timestamp values recorded to calculate average settling velocity components in 2D for each camera 290 (below, U_x is horizontal speed from camera one, U_y is horizontal speed from camera two, and $U_{z,1}$ 291 and $U_{z,2}$ are the vertical speeds corresponding to the two cameras). A resultant velocity magnitude 292 was then calculated for each camera, and these two values averaged to yield a single estimate for U293 per experiment (Eqn 9).

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$$U = \frac{1}{2} \left(\sqrt{U_x^2 + U_y^2 + U_{z,1}^2} + \sqrt{U_x^2 + U_y^2 + U_{z,2}^2} \right)$$

Each model was sunk five times and a mean U was calculated from these replicates. Replicates beyond a threshold of ±5% of the median sinking velocity were discarded from this average. Each model was dropped one additional time and photographed using a Canon 1200D DSLR camera (Tokyo, Japan) mounted on a tripod close to the tank, to obtain high resolution (18 megapixels) images which were used to determine model orientation (and thus L and A) during settling (**Error! Reference source not found.**D).

302 Wall effects

303 At low Re, the effects of artificial walls in an experimental (or computational) system can be 304 nonintuitively large and lead to substantial errors if not accounted for (Vogel, 1994). Acting as an 305 additional source of drag, the walls several tens of particle diameters away can slow down a sinking 306 particle and increase its apparent drag coefficient. We designed our experiments to minimise wall 307 effects by using an 0.8 m diameter tank (Error! Reference source not found.A) and model diameters 308 on the order of 1 cm. To reduce potential errors further, we applied the method of Fayon & Happel 309 (1960; summarised in Clift et al., 1978) to convert between the apparent drag coefficient when walls are present (C_D^{wall}) and the desired drag coefficient in an unbounded domain (C_D^{∞}): 310

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$$C_D^{\infty} \approx C_D^{walls} - \frac{24}{Re} \left(K(\lambda) - 1 \right)$$

313 where

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$$K(\lambda) = \frac{1 - 0.75857 \cdot \lambda^5}{1 - 2.1050 \,\lambda + 2.0865 \,\lambda^3 - 1.7068 \,\lambda^5 + 0.72603 \,\lambda^6}$$

Here, $\lambda = d/D$ where d is the diameter of the sinking particle and D is the tank diameter; we take d = L. While Eqn 10 is not exact, it substantially reduces the error otherwise incurred if one were to neglect wall effects entirely. Note that Eqn 10 is only valid up to about Re = 50, beyond which different corrections can be used (Clift et al., 1978).

We applied this correction by taking any experimentally determined C_D to equal C_D^{walls} , and using C_D^{∞} estimated according to Eqn 10 for subsequent calculations as detailed below. In our experiments, λ ranged from 0.0027 – 0.0173, yielding *K* between 1.0057 – 1.0377. Wall effects were therefore quite small, with C_D^{∞}/C_D^{walls} ranging from 0.993 – 0.994.

323 Iterative approach

324 First iteration

To construct an empirical cubic spline $C_D^E(Re)$ needed to solve for U^O , at least three experimental data points (corresponding to three scale factors) are needed. These first three *S* were chosen by using an existing empirical $C_D(Re)$ relationship for a sphere, valid for $0 < Re < 10^6$ (Morrison, 2013, Fig. 8.13, page 625):

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$$C_D^M(Re) = \frac{24}{Re} + \frac{2.6 \left(\frac{Re}{5.0}\right)}{1 + \left(\frac{Re}{5.0}\right)^{1.52}} + \frac{0.411 \left(\frac{Re}{263,000}\right)^{-7.94}}{1 + \left(\frac{Re}{263,000}\right)^{-8.00}} + \left(\frac{Re^{0.80}}{461,000}\right)^{-8.00}$$

While morphologically complex particles such as foraminifera tests (Table 1) are not expected to behave like ideal spheres, Eqn 11 should be sufficient to provide initial guesses, after which we iterate to find the solution. We note that if the particle shapes of interest were all most similar to some other

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- well-studied geometry (e.g., cylinders, discs, etc), using a known $C_D(Re)$ relationship for that shape
- could provide better initial guesses and faster convergence.

Substituting Eqn 11 into Eqn 7 (with S = 1, $V = V^{O}$, and $\rho_{particle}^{O}$, μ^{O} substituted) and moving 336 all terms to one side, we can numerically solve (MATLAB: fzero function) for our first estimate of the 337 338 operating Re^{O} . Substituting this Re back into Eqn 7 or Eqn 11 yields an estimate of the operating C_{D}^{O} . 339 We aimed to reproduce this Re and C_D in the first experiment, excepting that we accounted for wall effects by distinguishing between C_D^{∞} and C_D^{walls} expected to occur in the tank. Hence, we could again 340 substitute this Re into Eqn 7 but now with $ho_{particle}$ corresponding to the resin model and ho_{fluid} and 341 μ corresponding to mineral oil, and combine this expression with Eqn 10, assuming our estimated 342 $C_D = C_D^{\infty}$, $C_D^{\mathcal{F}} = C_D^{walls}$, and $\lambda = S L^O / D$. The resulting expression can be solved numerically for the 343 first scale factor, termed S_1 . Two more scale factors (S_2 and S_3), one smaller and one larger than S_1 , 344 were chosen to span expected Re values for forams from published literature (e.g. Fok-Pun & Komar, 345 1983; Takahashi & Be, 1984; Caromel et al., 2014) as well as Re⁰ for other species which had reached 346 347 convergence. This procedure was intended to bound the correct S value that reproduces the operating Re^{O} and C_{D}^{O} of the settling particle. The three models were printed, their actual volumes V measured 348 via weighing, and their settling velocities U experimentally measured as detailed in Appendix 3. 349

- An empirical cubic spline curve $C_D^E(Re)$ can now be fitted (D'Errico, 2009) to these three initial (*Re*, C_D) data points, constrained to be monotonically decreasing and concave up within the limits of the data to match expectations for drag on objects at low to moderate *Re*. Three optimally spaced spline knots were used since this yielded excellent fits to the data as the number of data points increased. These details of the spline as well as its order (i.e., cubic vs linear) are somewhat arbitrary but we ensured that our results were sufficiently converged as to be insensitive to them (see Remaining iterations).
- The operating point (Re^{O}, C_{D}^{O}) corresponding to the particle settling in the natural environment can be visually represented as the intersection point of the $C_{D}^{\mathcal{F}}(Re)$ curve defined by Eqn 7 (with S = 1and $\rho_{particle}^{O}, \rho_{fluid}^{O}, \mu^{O}$) and the empirical $C_{D}^{E}(Re)$ relationship based on our experimental data. Algebraically, the operating point is the solution to $C_{D}^{\mathcal{F}}(Re) = C_{D}^{E}(Re)$. We solved for Re^{O} numerically using a root finding algorithm (MATLAB's *fzero*) on the objective function $C_{D}^{\mathcal{F}}(Re) - C_{D}^{E}(Re) = 0$ and then obtained C_{D}^{O} by substituting Re^{O} into Eqn 7. Finally, U^{O} was easily determined from the definition of Re^{O} (Eqn 1 with U^{O} , L^{O} , and ρ_{fluid}^{O} substituted).
- Since our first three empirical data points and fitted spline C_D^E corresponded to guessed model scale factors *S*, our initial operating point prediction (Re^O , C_D^O) often was not located near any of these initial points or sometimes even within the bounds of these data (in which case linear extrapolation of C_D^E was used to estimate the operating point). Therefore, to ensure the accuracy of our predicted U^O , we continued iterating with additional experiments.

368 *Remaining iterations*

The model scale factor for the Nth experiment was chosen by combining Eqns 7, 8, and 10 with $Re = Re^{O}$ and $C_{D}^{walls} = C_{D}^{O}$, (from the previous iteration), $C_{D}^{\mathcal{F}} = C_{D}^{\infty}$, and $V = V^{predict}$, and

- numerically solving for *S*. A model close to this new scale was printed and sunk, its settling velocity *U* recorded and *Re* and C_D computed, and a more accurate spline C_D^E constructed by including this new data point. The calculation of (Re^0 , C_D^0) detailed in the previous section was then repeated, yielding a more accurate operating point. Overall, the aim was to tightly bound the predicted operating point with experimental data to maximize confidence in the fitted spline in this region.
- The iterative process (visualized as a flowchart, **Error! Reference source not found.**B, with a specific example of convergence given in Fig. 2 C & D) was repeated until:
- 378 1) the predicted operating point was not extrapolated beyond our existing data,
- 379 2) the variation in calculated U⁰ between the fitting of a linear spline and cubic spline was no
 380 greater than 5%, and
- 381 3) the variation between the predicted *Re⁰* and the closest experimentally measured *Re* was
 382 less than 15%.
- 383 In many cases, the difference between results based on four versus three data points was very small
- 384 (Fig. 2 C, D), indicating rapid convergence and the possibility of streamlining the method further in
- the future. Through this method we calculated the sinking velocities of 30 species of planktonic
- 386 foraminifera (Table 1).

387 Method Validation

Our basic methodology was first validated by 3D printing a series of spherical models (10 - 20 mm in)388 389 diameter) for which the theoretical $C_D(Re)$ relationship is already well-known. In order to achieve low 390 density (and thus low sinking velocity and low Re), these spheres were hollow and filled with oil via two small holes (of diameter 0.8% of the sphere diameter). Our empirically generated $C_D^E(Re)$ curve 391 compares favourably with the theoretical $C_D^M(Re)$ curve (Morrison, 2010) (R²=0.875, Error! Reference 392 393 **source not found.**A), with the distance between the curves approximately constant above $Re \approx 25$. 394 While the error grows larger at lower Re, we expected most foraminifera species to operate at $Re \approx$ 18 – 55 based on previous work (Berger et al., 1972; Fok-Pun & Komar, 1983; Takahashi & Be, 1984; 395 396 Caromel et al., 2014).

To quantify errors in our approach even more directly, we then considered hypothetical hollow 397 spherical particles with the same material density ($\rho_{particle}^{0}$) as foraminifera tests and a range of sizes 398 $(L^{0} = 750 - 1150 \ \mu m$, similar to the species we studied) settling in seawater. This size range 399 corresponds to Re = 12 - 27, the area where our $C_D^E(Re)$ curve is most divergent from $C_D^M(Re)$. We 400 compared predictions of the operating U^{O} based on our empirical $C_{D}^{E}(Re)$ curve versus the theoretical 401 $C_D^M(Re)$ curve for spheres as outlined above, substituting Eqn 11 for $C_D^E(Re)$ in the latter case. 402 Maximum relative error in predicted U^0 was 11.5% at Re = 16 (corresponding to a sphere 860 μ m in 403 diameter) while the minimum difference was 6.5% at Re = 27 (corresponding to a sphere of 1150 μ m 404 405 in diameter, Error! Reference source not found.A). This level of error is much smaller than the variation in U^0 we predicted across the 30 foraminifera species we investigated (Table 1). 406

407 Time to Terminal Velocity

408 This study was concerned with predicting steady sinking speeds, but in our experiments, each model

409 for a finite amount of time to accelerate from rest at the point of release to its

410 terminal sinking velocity. Since this transient portion of the sinking trajectory could introduce errors

411 into our analysis, it is important to determine whether it affected any of our recorded data.

412 During the transient acceleration phase, Eqn 2 does not hold. Instead, we can revert to the more

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413 general form of Newton's second law:

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$$\Sigma F = F_{drag} + F_{buoyancy} - F_{weight} = -Ma = -M\frac{dU}{dt}$$

416 where $M = V \rho_{particle}$ is particle mass, and the acceleration a can be equated to the time derivative 417 of instantaneous velocity $\frac{dU}{dt}$. A negative sign appears on the right-hand side so that we can define 418 the downward movement as positive for convenience. We can then substitute expressions for each 419 force as before:

420
421
$$\frac{1}{2}C_D(\Psi, Re) \rho_{fluid} U^2 A + V \rho_{fluid} g - V \rho_{particle} g = -M \frac{dU}{dt}$$

While thus far we have not assumed anything about the particle shape, to proceed further we require knowledge of $C_D(\Psi, Re)$ from vanishingly small Re (when the particle is at rest) up to the terminal velocity. Hence we will assume a spherical particle as an approximation to the model forams, so that Morrison's empirical equation (Eqn 11) can then be substituted for $C_D(\Psi, Re)$:

Here we have isolated $\frac{dU}{dt}$ on the right-hand side. If the definition of $Re = \frac{L U \rho_{fluid}}{\mu}$ is inserted into 14 (not shown for brevity), one obtains an ordinary differential equation (ODE) for the unsteady velocity U(t). The depth of the sphere Z(t) can then be obtained by solving a second much simpler ODE:

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432
$$U = \frac{dZ}{dt}$$

Both ODEs are easily solved numerically by e.g. MATLAB's *ode45* subject to the initial conditions U(t = 0) = 0 and Z(t = 0) = 0.

It is well known that as *Re* approaches zero in the limit of inertia-less Stokes flow, unsteadiness can only occur due to time-varying boundary conditions. Thus, a microorganism that stops actively swimming will almost instantly come to a stop, and a heavy micro-particle released from rest will

- almost instantly begin sinking at its terminal velocity (Purcell, 1977). As *Re* increases and inertia
 becomes increasingly important, the transient period of acceleration becomes longer. Therefore, a
 reasonable worst-case to examine here is the foraminifera model that sank at the highest *Re*.
- 441 We found *Globorotalia* (*Truncorotalia*) truncatulinoides to operate at Re = 42 (Table 1) but here we conservatively chose the largest scale model used to generate its $C_D^E(Re)$ spline for which S = 16 and 442 *Re* = 90. Inserting this model's length *L*, area *A*, and measured volume *V* into 14, we obtain solutions 443 444 for the time varying speed and depth of a sphere approximating this model's geometry (Error! 445 Reference source not found.B). The depth corresponding to where speed equals 99.9% of the terminal 446 velocity is approximately 4.6 cm, which is much smaller than the 19 cm between where the models 447 were released and the edge of the cameras' fields of view for data collection. Hence, the transient 448 acceleration of each model foraminifera should have had no effect on our data or results. Most of our 449 models should have reached terminal velocity even sooner since they sank at lower Re, e.g. within 2.2 450 cm for *C. dissimilis* operating at Re = 36 (Table 1).

451 Results & Discussion

452 Here we present a novel method of determining settling speed by leveraging dynamically scaled 453 models falling under gravity rather than being towed at a controlled speed. Applying our method to 454 foraminifera-inspired spherical particles (Error! Reference source not found.A), we predict settling 455 speeds within 11.5% of theoretical expectations (Error! Reference source not found.E). In Error! Reference source not found.C & D we present an example of convergence of our method to the 456 operating Re^{O} , C_{D}^{O} , and U^{O} of a typical foraminifera species. There was little variation in the number 457 of iterations required to reach convergence (mean 4, range 3-6, see Table 1), despite the 458 459 morphological complexity of some species (e.g. Globigerinoidesella fistulosa). We suspect the higher 460 end of this range was due to these species having forms that were particularly challenging to clean 461 residual resin from, or the incomplete removal of air bubbles once submerged in oil.

462Table 1.Predicted sinking speeds U^0 for the 30 species of planktonic foraminifera included in this study, with each species463shown in both spiral view and 90° rotation (so that the spiral view is facing to the left). ¹ indicates species scanned high464resolution. Scans of the remaining 27 species were obtained from The University of Tohoku museum's database. Operating465 Re^0 and C_D^0 predicted for each species are also presented, and for comparison, the theoretical C_D^M of a sphere at the same466 Re^0 . The number of model iterations required to achieve convergence of U^0 is listed (iter). Species synonyms checked at

467 World Foramaninifera Database (Hayward et al., 2018).

Image	Species	<i>L⁰</i> (μm)	U ⁰ (cm s ⁻¹)	Re ⁰	C_D^0	$C_D^M(Re^0)$	lter	Image	Species	<i>L⁰</i> (μm)	U ⁰ (cm s ⁻¹)	Re ⁰	C_D^0	$C_D^M(Re^0)$	lter
	Neogloboquadrina acostaensis (Blow, 1959)	307	3.6	9.3	2.5	4.2	4		Sphaeroidinella dehiscens (Parker & Jones, 1865)	460	5.1	32.8	1.7	1.9	4
	Globigerinella adamsi (Banner & Blow, 1959)	596	3.0	9.6	2.6	4.1	4		Sphaeroidinella dehiscens ¹ (Parker & Jones, 1865)	269	2.5	8.3	5.0	4.5	4
	Dentoglobigerina altispira (Cushman & Jarvis, 1936)	282	6.1	37.3	1.6	1.8	5		<i>Catapsydrax dissimilis</i> (Cushman & Bermúdez, 1937)	380	6.2	36.4	1.7	1.8	4
	Globoturborotalita apertura (Cushman, 1918)	650	2.8	7.4	4.2	4.8	3		Neogloboquadrina dutertrei (d'Orbigny, 1839)	259	4.0	17.6	2.0	2.8	3
	Globigerina bulloides (d'Orbigny, 1826)	296	3.3	13.4	2.4	3.3	5		Globigerinoidesella fistulosa (Schubert, 1910)	392	2.6	11.2	3.0	3.7	4
	Globigerinoides conglobatus ¹ (Brady, 1879)	445	1.8	5.8	5.9	5.7	4		Globorotaloides hexagonus (Natland, 1938)	321	2.7	8.3	3.4	4.5	4
	Globorotalia (Truncorotalia) crassaformis (Galloway and Wissler, 1927)	250	3.3	9.5	2.3	4.1	4		Globigerinella obesa (Bolli, 1957)	539	2.8	9.4	3.2	4.1	4
A O	Neogloboquadrina humerosa (Takayanagi & Saito, 1962)	360	5.0	22.7	2.0	2.4	4		Globorotalia praemenardii (Cushman & Stainforth, 1945)	569	3.9	14.0	2.0	3.2	4
	Globoconella inflata (d'Orbigny, 1839)	422	3.9	21.4	2.5	2.5	4		Globoconella puncticulata (d'Orbigny in Deshayes, 1832)	269	4.0	14.5	2.8	3.1	3
	<i>Fohsella lobata</i> (Bermúdez, 1949)	312	4.2	13.1	1.5	3.3	3		Fohsella robusta (Bolli, 1950)	312	3.5	11.2	2.3	3.7	4
	<i>Globorotalia margaritae (</i> Bolli & Bermúdez, 1965)	258	2.9	8.1	2.8	4.6	3		Dentoglobigerina tripartita (Koch, 1926)	303	5.7	33.6	2.0	1.9	3
	Paragloborotalia mayeri (Cushman & Ellisor, 1939)	298	3.1	8.1	3.2	4.5	4		Paragloborotalia siakensis (LeRoy, 1939)	366	3.4	10.4	2.6	3.9	4
06	Globoturborotalita nepenthes (Todd, 1957)	353	4.2	16.4	2.4	2.9	3		Globoconella sphericomiozea (Walters, 1965)	627	4.3	18.8	2.3	2.7	4
	Pulleniatina obliquiloculata (Parker & Jones, 1865)	350	4.1	24.6	2.1	2.3	3		Globorotalia (Truncorotalia) truncatulinoides (d'Orbigny, 1839)	325	6.3	42.4	1.5	1.6	6
	Fohsella peripheroronda (Blow & Banner, 1966)	422	3.6	10.1	2.4	4.0	3		Orbulina universa¹ (d'Orbigny, 1839)	567	1.0	3.3	10.7	8.7	6
A68	Praeorbulina curva (Blow, 1956)	433	4.8	17.8	2.2	2.8	3								

469 Our predicted sinking speeds of foraminifera fall within aggregated existing data for 14 species (Error! 470 Reference source not found., Fok-Pun & Komar, 1983; Takahashi & Be, 1984; Caromel et al., 2014) 471 and compare well with known speeds for other particles of comparable size and density (e.g. faecal 472 pellets, (Table 3, Iversen & Ploug, 2010), phytoplankton (Fig. 1 in Smayda, 1971)). However, it should 473 be noted that our predicted speeds are higher than published values for five out of the seven 474 foraminifera species for which direct comparisons are possible (Error! Reference source not found.). This could be due to our ability to observe enlarged models of sinking foraminifera more accurately 475 476 compared to actual specimens, and the lack of control for wall effects in previous work, which would 477 tend to underestimate sinking speeds. There could also be considerable natural variation, which our 478 single specimen per species (excluding S. dehiscens) does not capture.

479 Sedimentation of microscale plankton has been measured both in situ (e.g. Waniek, Koeve and Prien, 480 2000) and in the laboratory (e.g. Smayda, 1971; Miklasz and Denny, 2010). By settling dense 481 suspensions of microorganisms, these studies provided a population sinking rate (Bienfang, 1981) 482 which could be two to three times lower than the settling velocity of an isolated particle in the typically 483 dilute ocean (Miklasz & Denny 2010). Other studies have, like us, used enlarged models of microscale 484 plankton to facilitate observations. Padisak et al. (2003) used handmade models of plankton to 485 examine drag, but there was no attempt at accurately matching Re. Holland (2010) used mechanical 486 pencil leads as models of sinking diatom chains, keeping Re < 1 in an improvement over Padisak et al. 487 (2003). However, neither authors calculated sinking velocities for real organisms. Our dynamic scaling 488 approach ensures that we accurately recreate the fluid flows around settling organisms - a 489 requirement for the correct prediction of sinking speeds. We also improve on previous methodologies 490 by effectively eliminating wall effects, basing our models on µCT scans, and using inexpensive cameras 491 to observe natural sinking orientation.

492 By design, our dynamic scaling approach yields an interpolated $C_D(Re)$ curve that describes the flow 493 dynamics (and thus sinking speeds) that would occur if various fluid and/or particle parameters were 494 varied, offering a degree of flexibility not seen in other studies. For example, phytoplankton blooms 495 can increase both the density and viscosity of water due to exudates (Jenkinson et al., 2015), while 496 increasing global temperatures have the opposite effect. The density and viscosity of seawater also 497 naturally vary with latitude. Understanding how these variations affect sinking rates can offer insights 498 into the evolutionary pressures on plankton. Our approach also allows us to isolate the effects of shape on sinking, even across species of widely varying size, density, etc, by comparing C_D of different 499 500 species all hypothetically sinking at the same *Re*; a manuscript focused on such biological questions 501 relating to foraminifera is currently in preparation. Differential settling speeds of foraminfera also

have implications for nutrient cycling, paleoclimate reconstruction (Kucera, 2007), and the marinecalcite budget (Schiebel, 2002).

504 Our method can easily be modified to study sedimenting particles operating at any Re, providing the 505 system's Re range can be experimentally replicated. Other sinking marine particles include diatoms $(Re \approx 10^{-2} - 1,Botte, Ribera D'Alcalà and Montresor, 2013)$ and radiolaria $(Re \approx 10 - 200,Takahashi and Montresor, 2013)$ 506 507 Honjo, 1983), for which one could use digital models as we have in conjunction with a suitably viscous 508 fluid (high viscosity silicone oil, see SI Further Applications) to enable sufficiently large models to be 509 produced (25 cm, see SI: Further Applications). The method can also be applied to terrestrial systems 510 such as settling spores ($Re \approx 50$ e.g. Noblin, Yang and Dumais, 2009) and dispersing seeds ($Re \approx 10^3$ 511 Azuma and Yasuda, 1989), again by using 3D printed models based on (often existing) μCT data.

512 Whilst our method pertains to settling in a quiescent fluid, one could conduct similar experiments using a flume to calculate threshold resuspension velocities (i.e. the horizontal flow speed required to 513 514 lift a particle off the substrate), important in the study of wind erosion and particle transport and 515 deposition (Bloesch, 1995; Bagnold, 1971). Similarly, studying particles suspended in shear flow could 516 be achieved using a treadmill-like device (e.g. Durham et al (2009)) or a Taylor-Couette apparatus (e.g. Karp-Boss & Jumars (1998)). While additional dimensionless groups beyond Re and C_D would need to 517 518 be matched to achieve similitude in these systems, we hope that our study provides a starting point 519 for the experimental study of these and other more complex problems.

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529 Competing interests

530 The authors have no competing interests to declare.

531 Author contributions

SH & MW were responsible for conceptualisation of the project, MW, SH and RS contributed to 532 developing the methodology, MW collected data and JUH & FW scanned the three foraminifera at 533 534 high resolution (See Table 1). RS & MW were responsible for MATLAB code, data curation and 535 visualisation, formal validation of the method and the initial draft of the manuscript. TH segmented the high resolution scans, generated models used for 3D printing, and assisted with 3D printing of all 536 models. MW, SH & RS contributed to editing and reviewing the manuscript, FW & JUH provided detials 537 538 on methodology of high resolution scanning. SH & RS supervised MW for the duration of the project, 539 and SH provided funding, resources and administration for the project.

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763 Figure & Table Legends

764 Figure 1

- 765 Fig. 1.A) Diagram of relevant forces and parameters between the model (left) and real life (right).
- 766 B) Summary of the full method; details are discussed in main text. Boxes with thicker lines represent a decision, square boxes
- 767 are data inputs, rounded square boxes are manual processes, and circles are computational steps.
- 768 Figure 2

769 Fig. 2 A) Comparison of our empirically generated $C_D^E(Re)$ curve for 3D printed spheres versus the theoretical $C_D^M(Re)$ curve

(Morrison, 2010). Goodness of fit of $C_D^E(Re)$ to $C_D^M(Re)$ is $R^2 = 0.857$. Sphere diameter is indicated for each model.

771 *C* & *D*) An example of our iterative solution process for *C*. dissimilis showing best estimates of operating values (including 772 the required model scale *S* to achieve similitude) based on experimental data from 3 (*C*), vs 4 (*D*) models. For reference, the 773 theoretical $C_D^M(Re)$ curve (Morrison, 2010) for a sphere is also shown. In *C*, *S* corresponding to the operating point is 774 estimated as 13.84. After an additional model was sunk at *S* = 13 (*D*), slightly more accurate estimates of the operating *S*, 775 Re^o , and C_D^o were obtained. When scaling the model for 3D printing only 1 decimal place was used rather than the 2 shown. 776 *D*-1) Models of foraminifera 3D printed in clear and black resin. Models were used for public engagement but demonstrate

the fidelity of the printer compared to the scan data shown Table 1.

779 Table 1

780Table 1.Predicted sinking speeds U^0 for the 30 species of planktonic foraminifera included in this study, with each species781shown in both spiral view and 90° rotation (so that the spiral view is facing to the left). ¹ indicates species scanned high782resolution. Scans of the remaining 27 species were obtained from The University of Tohoku museum's database. Operating783 Re^0 and C_D^0 predicted for each species are also presented, and for comparison, the theoretical C_D^M of a sphere at the same784 Re^0 . The number of model iterations required to achieve convergence of U^0 is listed (iter). Species synonyms checked at

785 World Foramaninifera Database (Hayward et al., 2018).