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Electromagnetism as an emergent phenomenon: a step-by-step guide

Carlos Barceló¹, Raúl Carballo-Rubio¹, Luis J Garay^{2,3} and Gil Jannes⁴

¹ Instituto de Astrofísica de Andalucía (IAA-CSIC), Glorieta de la Astronomía, 18008 Granada, Spain

² Departamento de Física Teórica II, Universidad Complutense de Madrid, 28040 Madrid, Spain

³ Instituto de Estructura de la Materia (IEM-CSIC), Serrano 121, 28006 Madrid, Spain

⁴ Modelling & Numerical Simulation Group, Universidad Carlos III de Madrid, Avda. de la Universidad 30, 28911 Leganés, Spain

E-mail: carlos@iaa.es, raulc@iaa.es, luisj.garay@ucm.es and gil.jannes@uc3m.es

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Abstract

We give a detailed description of electrodynamics as an emergent theory from condensed-matter-like structures, not only *per se* but also as a warm-up for the study of the much more complex case of gravity. We concentrate on two scenarios that, although qualitatively different, share some important features with the idea of extracting the basic generic ingredients that give rise to emergent electrodynamics and, more generally, to gauge theories. We start with Maxwell's mechanical model for electrodynamics, where Maxwell's equations appear as dynamical consistency conditions. We next take a superfluid ³He-like system as representative of a broad class of fermionic quantum systems whose low-energy physics reproduces classical electrodynamics (Dirac and Maxwell equations as dynamical low-energy laws). An important lesson that can be derived from both analyses is that the vector potential has a microscopic physical reality and only in the low-energy regime is this physical reality blurred in favor of gauge invariance, which in addition turns out to be secondary to effective Lorentz invariance.

Keywords: emergent gravity, electrodynamics, quantum gravity, gauge invariance, gravity



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1. Introduction

This work concerns emergent electromagnetism, but its motivation comes from trying to construct an emergent theory of gravity. Let us start at the beginning.

Given the classical behavior of general relativity, it is unavoidable to consider that a deeper-layer theory regularizing its singularities must exist. The search for such a theory, generically denoted as quantum gravity given the two main ingredients it should incorporate, is one of the cornerstones of modern theoretical physics. Merging quantum mechanics and general relativity encounters a number of difficulties; most arguably emanating from their different empathies toward the presence of background structures. Although quantum mechanics is easily implementable when a background structure exists, general relativity demands the absence of an *a priori* fixed background structure. Given this situation, one can try to adapt quantum mechanics so as to elevate background independence, i.e., a geometrical viewpoint, to a fundamental principle (e.g., consider the loop quantum gravity approach [1–3]).

However, the situation has also led some researchers to consider whether Einstein's theory could be just an emergent classical theory [4–6]. From this perspective, one does not have to strictly quantize general relativity, but to search for an underlying structure, containing in principle no geometric notions whatsoever, such that classical general relativity can emerge at a coarse-grained level. In this work, we use the word emergent in this sense. For example, we would consider the string theory approach to quantum gravity as emergent, but approaches such as causal dynamical triangulations, causal sets, or loop quantum gravity as non-emergent. Taking aside the much-developed string theory approach, there exist some much less explored emergent-gravity approaches based on condensed-matter-like systems [7, 8]. Contrarily to string theory, these latter approaches keep no relativistic trace at the fundamental level: even special relativity is emergent.

In the last 10 years, it has become clear that the appearance of metric structures controlling the propagation of effective fields within condensed matter systems, mostly in low-energy regimes, is quite simple and ubiquitous [9]. However, in general, these metric structures do not follow Einstein's equations, and it is difficult to force them to, even at a theoretical level [9, 10]. It is not clear what the fundamental origin of this difficulty is. In the context of an emergent dynamics in the style of Sakharov, this difficulty has been traced back to the ubiquitous non-relativistic behavior of the effective fields at the scale playing the role of Planck scale in these systems [7, 11].

Given the difficulties in constructing a complete emergent theory of gravity within this setting, we decided to explore, in detail, all steps involved in the construction of the much simpler case of emergent electromagnetism, to refresh our knowledge before returning to the gravitational problem. Moreover, to our knowledge, there does not exist a work of reference in which this construction in condensed matter systems (in particular, in Helium-3) is performed in a step-by-step fashion, making transparent all hypotheses and approximations involved. It is our intention that this work may also serve as a study guide for specialists in other approaches to quantum gravity.

Although our current experimental knowledge of quantum electrodynamics has not required a revision, it is still interesting to analyze the structure of a possible deeper layer underneath electrodynamics. From an exercising perspective, as we have said, it is always helpful to understand simpler systems before embarking on more complicated endeavors. From a more physical perspective, there are partial emergent models that suggest that gravity and

electromagnetism might emerge in a unified manner from a single underlying system [7] and, indeed, this is also the situation in string theory [12, 13]. If the very arena in which physics occurs—spacetime—has a discrete underlying structure, it is sensible to think that electromagnetism would also have such structure.

In this work, we present two models of emergent electromagnetism. One is originally due to Maxwell himself [14, 15]. We revise and slightly update Maxwell's hydrodynamical model in light of the physics we know today. The other model, which constitutes the bulk of the paper, is more sophisticated and is based on ideas coming from what we know about the superfluid phases of Helium-3. This construction follows the lead of the works of Volovik (see [10] and references therein), and, among other things, intends to make his ideas more accessible to non-specialists in condensed matter. Many steps in the construction have our own perspective, therefore any misjudgement or error can only be blamed on us.

Remarkably, Maxwell arrived to his unification of light and electromagnetism through the development of a mechanical model that could underlie all electromagnetic phenomena [14, 15]. He imagined the electromagnetic aether as consisting of an anisotropic and compressible fluid made of cells, capable of acquiring rotation, separated by a layer of small idle wheels or ball bearings capable of rotating and moving between the cells. The bodies would be immersed in this fluid as an iron ball is immersed in water; they would distort the fluid around them. He did not commit to this specific model as truly representing physical reality, but defended it on the grounds of a 'proof of principle' of the possibility of formulating electromagnetism as a mechanical model.

At present, we know that the physics of the microscopic world is controlled by quantum mechanical notions. The second model presented here assumes a quantum-mechanical substratum as fundamental (therefore we do not discuss the possibility that quantum mechanics itself could also be emergent). Then, it is developed to show how, at least classical, electrodynamics can emerge. Again, we consider this a proof of principle and not a commitment with the very form of the underlying physics. The analysis proceeds in a step-by-step basis. We try to discern which are the basic ingredients common to any emergent theory of electromagnetism and which appear as particular of this specific construction.

It has been argued that the most crucial step made by Maxwell was abandoning his mechanical model and focusing on the properties of the resulting coarse-grained effective field theory [16–18]. The field-theoretical point of view has since been a central theme in most developments in fundamental physics. Whereas nobody can deny the tremendous power and successes of this approach, it assumes many ingredients as a matter of principle, without a deeper explanation: e.g., why is there a maximum velocity for the propagation of signals? Why is there gauge invariance? Why are elementary particles within a class indistinguishable? Why are there no magnetic monopoles? As we will see, an emergent approach is capable of explaining many of these questions. On the other hand, an emergent perspective puts a stronger accent on the universal characteristics of possible microscopic theories than on the specifics of a particular implementation. We think the emergent approach complements the field-theoretical approach, together providing a much richer source of understanding.

2. An updated Maxwell fluid model

In Maxwell's time, people did not have a clear idea of what electric currents really were, not to mention the unknown atomic structure of matter. Given the present knowledge, we can propose

an updated fluid model for electromagnetism following closely Maxwell's proposal [14, 15]. For other modern viewpoints on Maxwell's hydrodynamical model, see e.g., [19–24].

Imagine a fluid made of two different elementary constituents: vortical cells and small ball bearings. A vortical cell is made of a topologically spherical and deformable membrane filled with a fluid. The details of this fluid are not important in what follows, so, to simplify matters, let us take it to be incompressible and highly viscous. The membrane provides a fixed constant tension in all its points. It supports tangential as well as normal tensions. In the case in which the membrane were put to rotate around an axis, the filling fluid would rapidly end up rotating with a uniform angular velocity around that axis. The total angular momentum of the vortical cell will be $I\Omega$, with I denoting its moment of inertia and Ω denoting its angular velocity, or Iv_e/r_e , with r_e , v_e denoting its equatorial radius and velocity.

The fluid inside the cell has an isotropic (hydrostatic) pressure. When it is non-rotating, this pressure is a constant p_0 throughout the cell. However, rotation provides centrifugal forces that change the pressure pattern. In the equator, the pressure will have an excess $p_0 + \frac{1}{2}\rho v^2$ with respect to the poles. Independently of the precise form of the cell, the cell as a point will exert a pressure excess in the directions orthogonal to the axis. This pressure excess will be proportional to the rotation velocity squared:

$$p_{\parallel} = p_0 + C_{\parallel}\Omega^2, \quad p_{\perp} = p_0 + C_{\perp}\Omega^2, \quad C_{\parallel} < C_{\perp}. \quad (1)$$

We can also write this excess as

$$\Delta p = p_{\perp} - p_{\parallel} = \mu_{\text{micro}}^{-1} B_{\text{micro}}^2. \quad (2)$$

At this stage, the dimensions of B_{micro} and μ_{micro} are not fixed; only the dimensions of the preceding product. For later convenience, let us choose B_{micro} to denote minus the average density of angular momentum in the vortical cell, multiplied by a typical length scale in the system, $B_{\text{micro}} = -(I\Omega/V)R$. Then, the quantity μ_{micro} is a constant with units $\{ML\}$ (mass times length). Although the dimensions have been fixed, one can still multiply μ_{micro} by a dimensionless number N and B_{micro} by \sqrt{N} with no effect, or in other words, one can change the length scale R that defines B_{micro} if one redefines μ_{micro} accordingly. One could also define B_{micro} with a reversed sign with no effect (in fact, we have chosen the negative sign for later convenience). A specific definition of B_{micro} will appear only when fixing an operational meaning for it. Let us advance here that, when later introducing the unit of charge, it will be natural to define B_{micro} with units $\{(J/L^3)(L/Q)\}$ and μ_{micro} with units $\{ML/Q^2\}$ (J is the angular momentum and Q the charge).

On the other hand, a ball bearing is a small spherical ball (much smaller than a vortical cell) that sticks to any membrane in such a way that it can move over it, but it cannot slide; that is, any movement must be accompanied either by rotation or by a tangential stretching of the membrane itself. This fluid of small balls is also endowed with a hydrostatic (isotropic) pressure. This pressure produces a displacement of the microscopic distribution of ball bearings with respect to the vortical cells that in turn produces microscopic restoration forces. This is because most ball bearings will be attached to at least two vortical cells so only way to move them is by creating a tangential distortion (and a subsequent tension) on the membranes (see figure 1). Thus, the hydrostatic pressure of the ball bearings combined with their stickiness results in an equilibrium state endowed with tensions, which we call a vacuum state.

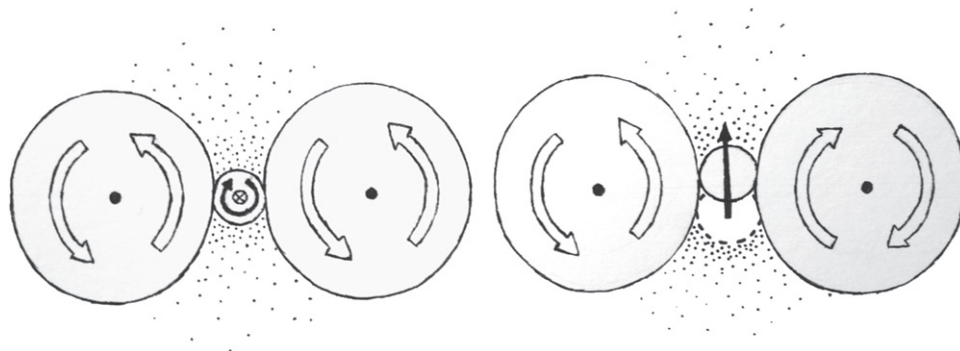


Figure 1. Diagram showing the transfer of rotation between the different elements in the fluid.

The complete description of a fluid made of a large number of vortical cells with an even larger number of ball bearings stuck to their surfaces, all combined in a box, will be tremendously complicated and uncontrollable in practice. However, from a coarse-grained perspective, we can use just a few macroscopic variables to characterize the state of the fluid, as in standard fluid mechanics. Consider one small part of the fluid that still contains a large number of constituents. At any coarse-grained point, the vortical cells will contribute with an overall hydrostatic pressure p_H plus some tension acting in a specific direction, the overall rotation axis. This will lead to an anisotropic pressure that can be written as

$$p_{ij} = \delta_{ij}p_H - \mu_0^{-1}B_iB_j. \quad (3)$$

Here, the vector \mathbf{B} is the macroscopic version of $\mathbf{B}_{\text{micro}}$ and therefore is proportional to the angular momentum density (total angular momentum in the coarse-grained point divided by its volume). The quantity μ_0 is a constant with units $\{ML\}$. The same redefinition ambiguities associated with the microscopic quantities apply to their macroscopic versions. These (p_H and \mathbf{B}) are our first macroscopic variables.

Now, non-vacuum states can have tangential displacements of the ball bearings (with their associated restoring tensions) beyond their equilibrium positions. We can characterize these tensions using a microscopic displacement vector field $\mathbf{D}_{\text{micro}}$ (displacement of each ball bearing with respect to its vacuum position). At the coarse-grained level, we can construct a displacement-density vector field \mathbf{D} and associate with it a force field $\mathbf{E} = \epsilon_0^{-1}\mathbf{D}$, with ϵ_0 representing a free constant with the appropriate dimensions. The real restoration force field will be proportional to the displacement and hence to this force field \mathbf{E} .

To recover (classical) electrodynamics from a fluid system like the one described, we still need one more ingredient: something must play the role of charge. In the vacuum state, ball bearings are all strongly stuck to vortical cells so that they cannot move from one vortical cell to another. However, out of this vacuum state, there can be movable ball bearings able to performing macroscopic displacements by jumping from cell to cell. To introduce movable ball bearings in the system, one could even break some of the strong links of the ball bearings characterizing the vacuum. In this way, ball bearings can be relocated in space. As we will see, the presence of regions with an overdensity or an underdensity of movable ball bearings, with respect to the vacuum state, can be associated with a positive and a negative charge density,

respectively. These overdensities and underdensities will in turn be responsible for the redistribution of tensions in non-vacuum states described previously.

Now we can explain how Maxwell's equations emerge from this fluid system.

- (i) Any rotational of the force field \mathbf{E} will exert a torque that will increase the angular momentum of the vortical cells (and thus decrease \mathbf{B} because we have defined it to be minus an angular momentum density). Once a specific meaning for \mathbf{B} is given (recall that one can redistribute a constant dimensionless factor N between \mathbf{B} and μ or, in other words, one has an initial flexibility in defining the length scale R), one can always find a specific ϵ_0 , so as to write

$$\frac{\partial \mathbf{B}}{\partial t} = -\nabla \times \mathbf{E}. \quad (4)$$

In other words, this equation can be interpreted as fixing the value of ϵ_0 with respect to μ_0 , and so fixing the relation between the force field \mathbf{E} and the displacement field \mathbf{D} .

- (ii) The presence of a ball bearing overdensity or underdensity produces a change in the displacement field. Assuming the displacements to be sufficiently small, unless the system is unnaturally fine tuned, we will always be able to write:

$$\nabla \cdot \mathbf{D} \approx \rho_Q \rightarrow \epsilon_0 \nabla \cdot \mathbf{E} \approx \rho_Q. \quad (5)$$

This charge density ρ_Q will be tightly related to the density of ball bearings with respect to the vacuum state. However, they do not need to perfectly coincide. The only quantity with a macroscopic operational meaning (at least, at this linear level) will be the charge density. At this stage, one could introduce some reference unit of charge, and accordingly change the units of all quantities by referring them to the effect of this reference charge.

- (iii) When the ball bearings move, they exert torques on the cells. This applies to both ball bearings strongly stuck to the cells (not movable to other cells), which produce a change in the displacement field, and to ball bearings movable between cells (associated with charge currents). Reciprocally, when the rotation field of the vortical cells \mathbf{B} acquires some rotational, it causes the ball bearings in the region to move (within their respective possibilities). We can encode this behavior in an equation of the form

$$\nabla \times (\mu_0^{-1} \mathbf{B}) = \mathbf{J}_T; \quad \mathbf{J}_T = \mathbf{J}_Q + \frac{\partial \mathbf{D}}{\partial t}. \quad (6)$$

The first term of the current \mathbf{J}_T is due to the movable ball bearings (a proper current of charge), whereas the second term is due to the displacement of the non-movable cells (hence displacement current). This equation fixes the value of μ_0 or, equivalently, the precise definition of \mathbf{B} (the equation determines the value of the length-scale constant R). It is interesting to note that there is a curious interplay between the displacement current term and the occurrence of relativistic dynamics, as this is the term that is absent in the magnetic limit of Galilean electrodynamics [25].

- (iv) Let us assume that the rotation field is divergenceless, although *a priori* there is no reason why this should be the case (more on this later). Then, we will have

$$\nabla \cdot \mathbf{B} = 0. \quad (7)$$

That it is, we have recovered all of Maxwell's equations from a mechanical fluid system. Let us make some observations here:

- (i) As already remarked by Maxwell, the important point here is not the specific details of this specific fluid model but its existence. The equations for the macroscopic fields will not depend strongly on these details. In the derivation, it has been necessary to make the assumption of smallness of all the perturbations with respect to the vacuum state. Beyond this regime, one would expect to observe non-linear effects. For example, one would expect non-linear pressures of the form

$$p_{ij} = \delta_{ij}p_0 - \mu^{-1}(B, E)B_iB_j. \quad (8)$$

In the linear limit, one can approximate $\mu(B, E)$ by a constant μ_0 .

- (ii) From Maxwell's equations, one immediately deduces that this system admits light-like perturbations. These perturbations will move with speed $c = \sqrt{\epsilon_0\mu_0}$. The fluid system can be perfectly described using Newtonian physics, in which there is no limitation to the velocity of the bodies. Nonetheless, light speed shows up directly from the elastic properties of the body. The crucial ingredient for generalized sound velocities to emerge is that variations in time of local properties depend on local gradients of these same properties.
- (iii) Given that $\nabla \cdot \mathbf{B} = 0$, one can always write $\mathbf{B} = \nabla \times \mathbf{A}$ locally. For instance, we could associate \mathbf{A} with the macroscopic version of the flow lines of the fluid within the vortical cells. On the other hand, in places in which magnetic fields are stationary, $\nabla \times \mathbf{E} = 0$; therefore, one can write $\mathbf{E} = -\nabla\phi$ locally. Then, the field ϕ represents a hydrostatic electric tension. A positively charged body will tend to move to places with smaller electric tension. In more general situations, due to the structure of Maxwell's equations, we can always write $\mathbf{E} = -\nabla\phi + \partial_t\mathbf{A}$. Knowing the coarse-grained structure of the fluid flow lines and hydrostatic electric tension, one knows ϕ and \mathbf{A} .

Now, one can realize that, regarding the values of \mathbf{E} , \mathbf{B} , the combination $\{\phi, \mathbf{A}\}$ and $\{\phi + \partial_t\chi, \mathbf{A} + \nabla\chi\}$ is equivalent. Within an emergence framework, an appropriate interpretation of the previous condition, the gauge invariance condition, is that although the flow structure of the vortical cells and the electric tension both have a specific reality, different macroscopic configurations related through a gauge transformation are operationally indistinguishable from the effective dynamical theory (see also [23, 24]). Gauge invariance appears because aspects of the system are 'invisible' to observers restricted to experience only the effective fields.

- (iv) In Maxwell's version of the fluid model, charges were associated directly with individual ball bearings, and currents with the movement of ball bearings from cell to cell. It seems completely unrealistic to have an electric current of this sort without some resistance or friction. But this resistance was perfectly accommodated in Maxwell's model by considering the currents as existing only within materials (conductors). Maxwell ascribed the ubiquitous resistance in a conducting wire to the collisions of the movable ball bearings when jumping between cells. However, this model will have difficulty dealing with a charged elementary particle in an otherwise empty space. There is no experimental evidence that the vacuum causes friction on a charge moving with uniform velocity. In

fact, the presence of an effect of this sort would immediately uncover the existence of a privileged reference frame, against all we know about the relativity principle.

However, we wanted our updated fluid framework to encompass also the movement of a free electron in an otherwise empty space. For that we proposed to associate an electron to an overdense region of ball bearings (this identification is of course not complete, as we have not attempted to specify the internal forces responsible for its structural stability). When considering an electron as a very localized overdense region, it appears difficult to find it in a uniform-velocity trajectory without dissipation of some sort. The movement of the ball bearings would be very noisy, with multiple collisions involved. However, if one imagines a pure plane-wave distribution of the overdense regions, it appears perfectly plausible for the propagation of the wave to occur without any appreciable friction: the propagation of the wave will not involve the presence of a macroscopic current of ball bearings. It is interesting to point out an analogy between this behaviour and that of free quantum particles. A position eigenstate of a wave function can propagate with a certain velocity, but as it travels, it diffuses in space. This diffusion might be seen as an analog in the quantum formulation of the noisy propagation expected in our fluid model. On the other hand, a momentum eigenstate is also an eigenstate of the Hamiltonian and it is not distorted by the propagation.

It is also interesting to realize that the system allows, in principle, pair-creation of particles from the vacuum. If one pulls out ball bearings in one place and moves them to another region, one would have created equivalent overdense and underdense regions. The appearance of these ‘quantum-like’ behaviors in our model might be interpreted as suggesting that the quantum mechanism itself could be an emergent phenomenon. Here, we just mention this possibility without pursuing it any further.

- (v) The absence, as far as we know, of magnetic monopoles in nature has always being striking. As mentioned previously, the fluid model *a priori* allows magnetic monopole configurations. However, when looking at the model carefully, one realizes that this kind of configurations does not seem to be favored by the system. Microscopically speaking, a magnetic monopole involves rotating cells with their angular momenta distributed radially. Any ball bearing located at the confluence of these cells will produce friction because the cells cause dragging forces incompatible with the no-sliding condition (see diagram in figure 2). It seems reasonable that the system would tend to avoid these configurations (the same argument applies at the macroscopic level). The divergence-free condition $\nabla \cdot \mathbf{B} = 0$ simply encodes the absence of these configurations.
- (vi) It is interesting to estimate how small the constituents of this fluid must be to pass unnoticed to current experiments. The smallest length scale ever tested is of the order of 10^{-19} m. So, in principle, a fluid structure several orders of magnitude beyond 10^{-19} m would remain undetected. Notice that the Planck length is 10^{-35} m, still 15 orders of magnitude ahead (equivalent to comparing a human with interstellar distances).
- (vii) It is also interesting to point out the different nature of light excitation and charged matter, even when considering it as wave-like. One can perfectly imagine a charge density with no overall velocity representing a charged particle or distribution of particles at rest. Light, however, is similar to a phonon excitation of a lattice; it always travels with its fixed velocity (of course, at very high energies, one would expect some dispersive effects to develop).

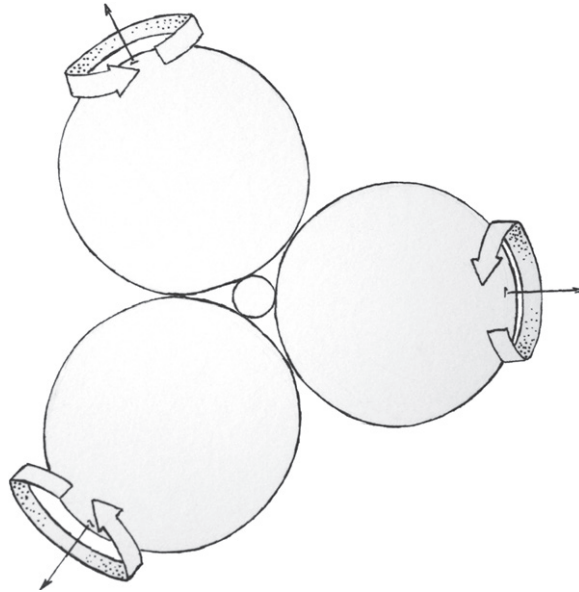


Figure 2. Diagram explaining the absence of magnetic monopoles. The system avoids these configurations because they would create friction for the central ball bearing at confluence of the rotating cells.

3. Model based on Helium-3

In this section, we present a model of emergent electrodynamics based on the well-established theoretical understanding of the physics of Helium-3. Our presentation of the model will follow a top-down scheme. Nonetheless, we would like to stress that these theoretical ideas were developed in close feedback with experiments and are proved to a great extent by them.

Most of the introductory material covered here is well understood nowadays but, as far as we know, has not been presented in a logical step-by-step order so as to lead to a final emergent electrodynamics. In the following, only specific references are quoted.

For a general discussion on superfluid Helium-3, one can draw on the review [26], or the books [27, 28] and references therein. Concerning the low-energy properties of this system and analogies with other branches of physics, including relativistic field theories, the seminal reference is [10].

3.1. Microscopic ^3He -like systems

Let us consider a quantum liquid composed of a large collectivity of spin-1/2 atoms. Here, we use the word ‘atom’ to mean that these spins need not be elementary objects (they need not be precisely ^3He atoms either). We require the interactions between these atoms to be short-range but otherwise they can be very complicated, including higher-than-two-body effects. We also require that the two-body interactions be characterized by a potential of Lennard–Jones type (which is rotationally invariant) plus possibly some interaction term involving the spins. Interactions in ^3He indeed possess these characteristics.

To solve a system of this sort in full detail is beyond human capacities. We need simpler theories that serve as approximate models of the exact microscopic theory. A first step in this direction is provided by Landau’s Fermi-liquid theory. This theory starts from the exact

description of a free Fermi gas. In the free theory, there appears the notion of Fermi surface and a notion of quasiatom and quasihole excitations. Landau's hypothesis is that, generically (at least under certain conditions of temperature and pressure), the N -particle ground state and the spectra of quasiatom and quasihole excitations (i.e., the spectra in the surroundings of the Fermi surface) of the preceding strongly interacting theories are in adiabatic one-to-one correspondence with that of the free theory [28]. Under this hypothesis, we can use the same labels for these states. There exists some microscopic justification of Landau's hypothesis [26].

Now, regarding all physics associated with low energies (vacuum state and excitations close to it), one can substitute the precise strongly interacting theory with an equivalent weakly interacting theory of quasiatoms. For instance, in second quantization language and in a momentum representation, one can write Landau's grand canonical Hamiltonian as

$$\begin{aligned} \hat{H}_L - \mu\hat{N} = & \sum_{p\alpha} \left(\frac{p^2}{2m^*} - \mu \right) \hat{a}_{p\alpha}^\dagger \hat{a}_{p\alpha} \\ & + \frac{1}{2} \sum_{pp'\alpha\beta} f(\mathbf{p}, \mathbf{p}', \alpha, \beta) \hat{a}_{p\alpha}^\dagger \hat{a}_{p\alpha} \hat{a}_{p'\beta}^\dagger \hat{a}_{p'\beta}. \end{aligned} \quad (9)$$

Here, $\hat{a}_{p'\beta}^\dagger$, $\hat{a}_{p'\beta}$ are, respectively, creation and annihilation operators of quasiatoms with α, β representing the spin degree of freedom. Following standard conventions, in the rest of the text we will write O instead of \hat{O} for any operator, leaving this notation for unit vectors. The expression (9) can be used in situations in which the number of quasiatoms (equal to that of atoms) is kept fixed. The chemical potential is $\mu = p_F^2 / (2m^*)$ where p_F denotes the Fermi momentum and m^* denotes the effective mass of the quasiatoms (this mass does not need to coincide with the mass of the initial atoms; in ^3He , it is a few times smaller). The function $f(\mathbf{p}, \mathbf{p}', \alpha, \beta)$ must be symmetric under the exchange $\mathbf{p}, \alpha \leftrightarrow \mathbf{p}', \beta$ and can be used to fit the specific interaction. Both m^* and f are in principle phenomenological quantities that depend on details of the microscopic interaction. This Hamiltonian model has proven to be very successful, for example, for the description of the normal phase of ^3He , in the temperature range between 1 and 003 K.

Now, there exist many systems that are different in the details of their interactions but are undistinguishable from a low-energy point of view. They form part of the same low-energy universality class. These universality classes are characterized by topological properties of the vacuum state. Therefore, when working out a theory of emergent electromagnetism, one is really obtaining a family of theories with the same low-energy behavior. The same operators $a_{p\alpha}$ can represent different physical quasiatoms in different strongly interacting spin-fluid systems. These operators can also represent the proper atoms of a weakly interacting spin-gas system. In the following, we will analyze the properties of a specific weakly interacting theory, independently of any specific physical realization one could have in mind. Thus, we will address only atoms, always remembering they could be equivalently quasiatoms.

3.2. A weakly interacting gas

Let us focus on a weakly interacting theory of spin-1/2 atoms. One can go one step further than Landau's Fermi-liquid theory and analyze a more general interaction.

One can introduce the atom field ψ . Then, in second quantization, the grand canonical Hamiltonian for the system of spin-1/2 atoms with two-body interactions reads

$$\begin{aligned}
H - \mu N &:= \int d^3x \psi^\dagger(\mathbf{x}) \left(-\frac{\hbar^2}{2m^*} \nabla^2 - \mu \right) \psi(\mathbf{x}) \\
&\quad + \frac{1}{2} \int d^3x d^3x' V(\mathbf{x} - \mathbf{x}') \psi^\dagger(\mathbf{x}) \psi^\dagger(\mathbf{x}') \psi(\mathbf{x}') \psi(\mathbf{x}).
\end{aligned} \tag{10}$$

We have assumed for the time being that the interaction potential does not depend on the spin. In the momentum representation, $\psi_\alpha = \sum_{\mathbf{p}} a_{\mathbf{p}\alpha} e^{i\mathbf{p}\cdot\mathbf{x}}$, we have

$$\begin{aligned}
H - \mu N &:= \sum_{\mathbf{p}\alpha} \left(\frac{p^2}{2m^*} - \mu \right) a_{\mathbf{p}\alpha}^\dagger a_{\mathbf{p}\alpha} \\
&\quad + \frac{1}{2} \sum_{\mathbf{p}_1 + \mathbf{p}_2 = \mathbf{p}_3 + \mathbf{p}_4, \alpha\beta} \tilde{V} \left(\frac{\mathbf{p}_1 - \mathbf{p}_2 + \mathbf{p}_3 - \mathbf{p}_4}{2} \right) a_{\mathbf{p}_4\beta}^\dagger a_{\mathbf{p}_3\alpha}^\dagger a_{\mathbf{p}_2\alpha} a_{\mathbf{p}_1\beta},
\end{aligned} \tag{11}$$

with

$$\tilde{V}(\mathbf{p}) := \frac{p_F^3}{\hbar^3} \int d^3r e^{i\mathbf{p}\cdot\mathbf{r}/\hbar} V(\mathbf{r}) \quad \text{and} \quad V(-\mathbf{r}) = V(\mathbf{r}). \tag{12}$$

We have taken this definition so that \tilde{V} has dimensions of energy. Our notation in what follows assumes a finite volume and thus a discrete sum in momentum space; the infinite volume limit can be obtained by adding the appropriate dimensionful constants. Notice that the potential term in equation (11) is invariant under a Galilean boost transformation of the reference frame. We should remember this property, which can apparently be lost under certain approximations, which will be made in the following. This sum contains different interaction channels: the Hartree channel [which contains the previous Landau terms, equation (9)], the Fock channel, and the pairing channel [28]. Of special relevance in what follows is the pairing channel that appears for interactions satisfying $\mathbf{p}_1 = -\mathbf{p}_2 =: \mathbf{p}$ and $\mathbf{p}_3 = -\mathbf{p}_4 =: \mathbf{p}'$. The pairing terms control the form of the vacuum state of the theory (see Leggett's discussion in [28]). The pairing Hamiltonian reads

$$\begin{aligned}
H_p - \mu N &:= \sum_{\mathbf{p}\alpha} \left(\frac{p^2}{2m^*} - \mu \right) a_{\mathbf{p}\alpha}^\dagger a_{\mathbf{p}\alpha} \\
&\quad + \frac{1}{2} \sum_{\mathbf{p}\mathbf{p}'\alpha\beta} \tilde{V}[(\mathbf{p}' + \mathbf{p})] a_{-\mathbf{p}'\beta}^\dagger a_{\mathbf{p}'\alpha}^\dagger a_{-\mathbf{p}\alpha} a_{\mathbf{p}\beta}.
\end{aligned} \tag{13}$$

If the potential does not depend on the orientation, it can only depend on $|\mathbf{p}' + \mathbf{p}| = p^2 + p'^2 + 2pp'\hat{\mathbf{p}} \cdot \hat{\mathbf{p}}'$. Then, we can always write it as an expansion of the form [26]

$$\tilde{V}(|\mathbf{p} + \mathbf{p}'|) = \sum_l \tilde{V}_l(p, p') P_l(\hat{\mathbf{p}} \cdot \hat{\mathbf{p}}'), \tag{14}$$

where P_l represents Legendre polynomials (the converse is not true: not all expansions can be put in exact correspondence with $V(r)$ potentials). As we are always interested in the surroundings of the Fermi surface (where the low-energy excitations reside), we can take the potential to depend mainly on the angle $\hat{\mathbf{p}} \cdot \hat{\mathbf{p}}'$ and not in the norms, which will be $p, p' \simeq p_F$.

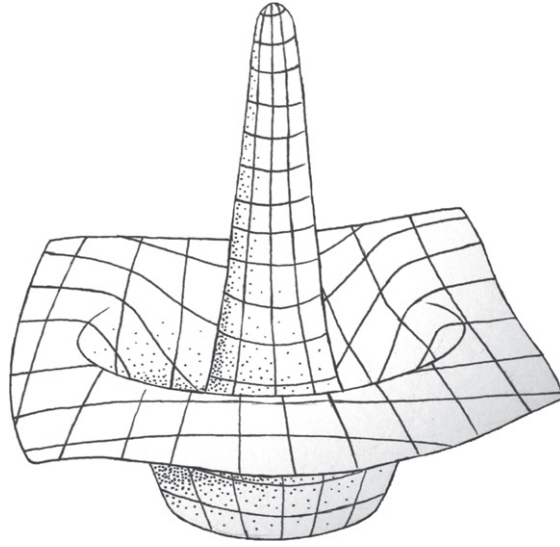


Figure 3. Diagram showing the qualitative form of the interaction potential.

Take now a microscopic interaction such that $|\tilde{V}_1| \gg |\tilde{V}_{l \neq 1}|$. Then $g := -\tilde{V}_1(p_F, p'_F)$ will be a positive constant because of the binding character of the potential. The potential will be written as

$$\tilde{V} \simeq -g \hat{\mathbf{p}} \cdot \hat{\mathbf{p}}' \simeq -\frac{g}{p_F^2} \mathbf{p} \cdot \mathbf{p}'. \quad (15)$$

For instance, the simplest interaction of this kind is the one provided by

$$V(\mathbf{x} - \mathbf{x}') = -\frac{g}{8p_F^2} (\nabla - \nabla')^2 \delta(\mathbf{x} - \mathbf{x}') + g \delta(\mathbf{x} - \mathbf{x}'). \quad (16)$$

This interaction has $\tilde{V}_{l \geq 2} = 0$. Near the Fermi surface, the remaining components verify $|\tilde{V}_1| \gg |\tilde{V}_0|$ so that the potential approximately behaves as equation (15). This interaction is the distributional limit of potentials of the form shown in figure 3. These potentials exhibit a repulsive hard core and an attractive tail (precisely the type of interaction between ^3He atoms). As it is not possible to construct a translation-invariant interaction potential with only $\tilde{V}_1 \neq 0$ (it would fail to be invariant under constant shifts in momentum space), equation (16) is the best approximation to an interaction of the form (15) one can find.

For properties involving long wavelengths compared with the interparticle distance, the model potential (16) will be perfectly appropriate as representative of an entire microscopic class. Taking this potential, the grand canonical pairing Hamiltonian finally reads

$$H_p - \mu N := \sum_{p\alpha} \left(\frac{p^2}{2m^*} - \mu \right) a_{p\alpha}^\dagger a_{p\alpha} - \frac{g}{2p_F^2} \sum_{pp'\alpha\beta} (\mathbf{p}' \cdot \mathbf{p}) a_{-p'\beta}^\dagger a_{p'\alpha}^\dagger a_{p\alpha} a_{-p\beta}. \quad (17)$$

This is the system we will work with in the next subsections.

3.3. Condensation and order parameters

The model interaction described in the previous subsection is called a p -wave spin-triplet pairing interaction. Below a critical temperature, it enforces the formation of anisotropic Cooper pairs (as opposed to the isotropy of the Cooper pairs in classical superconductivity). The spatial anisotropy of these pairs is associated with the fact that they possess angular momentum. Given the antisymmetric structure of the orbital part of the wave function, its spin structure must be symmetric and thus belongs to the triplet space of the spin product. These pairs condense, acquiring a macroscopic occupation. The macroscopic wave function or order parameter associated with the condensed pairs will be

$$\Psi_{\alpha\beta} := \frac{g}{p_F} \left\langle \sum_{\mathbf{p}} \mathbf{p} \cdot a_{\mathbf{p}\alpha} a_{-\mathbf{p}\beta} \right\rangle. \quad (18)$$

As a consequence of the spin-dependence and of the dominance of anisotropic p -wave interaction, this order parameter is not a scalar (as in the case of classical superconductivity or Bose-Einstein condensation) but a matrix, with spin indices α, β . There is also an implicit orbital index i because of the \mathbf{p} -dependence of $\Psi_{\alpha\beta}$. The normal-liquid phase has as symmetry group $\text{SO}(3)_L \times \text{SO}(3)_S \times \text{U}(1)$, i.e., independent rotations of the coordinate and spin spaces plus a phase-invariance symmetry associated with the conservation of the number of atoms. Pair condensation amounts to the spontaneous (partial) breaking of this symmetry. The order parameters appearing in this p -wave spin-triplet condensation are symmetric in the spin indices and therefore can always be written as:

$$\Psi_{\alpha\beta}^i = i(\sigma_a \sigma_2)_{\alpha\beta} d^{ai}, \quad (19)$$

where σ_a are Pauli matrices and d^{ai} is in general a complex vector in both spin and position space. The set $\{\sigma_a \sigma_2\}_{a=1,2,3}$ forms a basis for all 2×2 symmetric matrices. The imaginary prefactor i is introduced by convention to make this a real matrix basis and put all the complexity into the vector d^{ai} .

Depending on the details of the interaction, the order parameter can acquire different structures. The precise form of the order parameter is obtained by a minimization principle. In the microscopic theory the quantity to be minimised is the expectation value of the Hamiltonian (17) in the Fock vacuum state. One can alternatively use a minimization within Ginzburg–Landau theory, which is a special case of the phenomenological Landau-Lifshitz theory of second-order phase transitions. In this approach, one must minimize the free-energy functional of the order parameter. The order parameter (19) is zero above a certain critical temperature T_c but takes a finite value for $T < T_c$. The thermodynamic potential of interest in the experimental situation (constant temperature T and volume V) is the Helmholtz free energy. If we suppose that near T_c the free energy is analytic in the order parameter and obeys the symmetries of the microscopic Hamiltonian, then one can write a Taylor expansion near the critical temperature. The symmetries of the interaction dictate the type of terms that can appear in this expansion. The precise values of the coefficients in Ginzburg–Landau theory depend on the microscopic theory and can be derived from it, for example in the Bardeen–Cooper–Schrieffer theory of superconductivity [29].

Appendix appendix A provides some details regarding the minimization procedure in the case in which one neglects the spin-spin interactions. Under certain conditions, four solutions

are found to the minimization problem. The (Balian–Werthamer) BW and Anderson–Brinkman–Morel (ABM) states are associated by confrontation with experiments with the superfluid phases B and A, respectively. The other two states are the so-called planar and polar states. The planar state and the ABM state are topologically characterised by the presence of Fermi points. The BW state is fully gapped, while the polar state has a Fermi manifold of dimension 1. As we will see, Fermi points give rise to relativistic low-energy excitations. It is easy to understand why this is the case: near these points, the dispersion relation of quasiparticles is linear to leading order, and is three-dimensional, unlike in the case of Fermi manifolds of higher dimension [30].

For our purposes, we are especially interested in the planar state; we will show why in the next subsections. Its order parameter is

$$d_{\text{planar}}^{ai}(T) := \Delta(T) (\hat{s}^a \hat{m}^i + \hat{s}'^a \hat{n}^i), \quad (20)$$

where \hat{m} , \hat{n} are unit vectors in position space and \hat{s} and \hat{s}' are unit vectors in spin space subject to the orthogonality conditions $\hat{m} \cdot \hat{n} = 0$, $\hat{s} \cdot \hat{s}' = 0$. In this expression, the scalar function $\Delta(T)$ is the gap parameter that contains the dependence of the order parameters on the temperature T and the interaction constant g . At zero temperature, its value is approximately $\Delta_0 := \Delta(0) \simeq k_B T_c$, where T_c is the critical temperature.

The planar state has not yet been observed in nature among the superfluid phases of ${}^3\text{He}$. If one neglects dipole-dipole interactions, then this state is never the lowest energy state of the system. However, when considering these interactions, which in ${}^3\text{He}$ are rather feeble, this state should be the global minimum in a narrow temperature band in phase space (see [31]). Here, we are not considering real ${}^3\text{He}$ but a system constructed with atoms adapted to our needs. Thus, we assume that there exist some additional atom-atom interactions beyond the Lennard–Jones potentials such that they select the planar state as the natural vacuum.

It is sometimes instructive to remember the other well-known states of this system: the ABM and BW states. Their order parameters are respectively

$$d_{\text{ABM}}^{ai}(T) := \Delta(T) \hat{s}^a (\hat{m}^i + i \hat{n}^i), \quad (21)$$

$$d_{\text{BW}}^{ai}(T) := \Delta(T) \delta^{ai}. \quad (22)$$

Here, there is also an orthogonality condition $\hat{m} \cdot \hat{n} = 0$.

Before closing this subsection, let us comment that, within the interpretation of a strongly interacting system of atoms, the realization of any of these condensed phases takes us beyond the strict limits of applicability of Landau’s Fermi-liquid hypothesis. The Fermi surface of the free system has been deformed so strongly that it no longer survives. It has been either completely eliminated (BW state) or reduced to just some points (planar and ABM states). However, it is remarkable that a weak-interaction model of quasiatoms is able to describe correctly the condensation and low-energy excitation of these systems. For the interpretation in which one directly starts from a weakly interacting system of atoms, the previous comment is irrelevant: in this case, the weakly interacting theory is already the very microscopic theory.

3.4. Low-energy quasiparticle excitations

In this subsection, we will analyze the form of the quasiparticle excitations living right above the vacuum of the planar state. These are new types of quasiparticles, specific combinations of the atoms and holes of Landau's theory. We call them Bogoliubov quasiparticles.

Once the system has settled to a condensed state, the pairing interaction can be expanded up to quadratic order in perturbations around the condensed state (the so-called Gor'kov factorization [32]). It is easy to see that the resulting quadratic Hamiltonian reads

$$H_p - \mu N := \sum_{p\alpha} M(\mathbf{p}) a_{p\alpha}^\dagger a_{p\alpha} + \frac{1}{2p_F} \sum_{p\alpha\beta} \mathbf{p} \cdot \boldsymbol{\Psi}_{\alpha\beta} a_{-p\beta}^\dagger a_{p\alpha}^\dagger + \frac{1}{2p_F} \sum_{p\alpha\beta} \mathbf{p} \cdot \boldsymbol{\Psi}_{\beta\alpha}^* a_{p\alpha} a_{-p\beta}, \quad (23)$$

where we have defined $M(\mathbf{p}) = p^2/(2m^*) - \mu$. Consider now the order parameter to be a homogeneous planar state characterized by the vectors $\hat{\mathbf{s}}$, $\hat{\mathbf{s}}'$, $\hat{\mathbf{m}}$, and $\hat{\mathbf{n}}$. Let us choose a system of coordinates adapted to the pairs-spin-space Cartesian trihedral

$$\hat{\mathbf{x}} = \hat{\mathbf{s}}, \quad \hat{\mathbf{y}} = \hat{\mathbf{s}}', \quad \hat{\mathbf{z}} = \hat{\mathbf{s}} \times \hat{\mathbf{s}}'. \quad (24)$$

Calculating the commutator between quasiparticle operators and $H_p - \mu N$ shows that the evolution equations of quasiparticle operators particularized to the planar order parameter (20) are

$$i\dot{a}_{p\uparrow} = M(\mathbf{p})a_{p\uparrow} - c_\perp \mathbf{p} \cdot (\hat{\mathbf{m}} - i\hat{\mathbf{n}})a_{-p\uparrow}^\dagger, \quad (25)$$

$$i\dot{a}_{p\downarrow} = M(\mathbf{p})a_{p\downarrow} + c_\perp \mathbf{p} \cdot (\hat{\mathbf{m}} + i\hat{\mathbf{n}})a_{-p\downarrow}^\dagger. \quad (26)$$

Here, we have introduced the parameter $c_\perp := \Delta_0/p_F$ with dimensions of velocity. The evolution of the two spin populations is decoupled. This property permits us to consider the spin populations separately, simplifying the treatment.

Let us first anticipate the appearance of Fermi points in this vacuum state. Acting with the operator $i\partial_t$ on equation (25), one finds that the dependence on $a_{-p\uparrow}^\dagger$ vanishes, and one can write

$$\begin{aligned} (i\partial_t)^2 a_{p\uparrow} &= \left\{ M^2(\mathbf{p}) + c_\perp^2 \left[(\mathbf{p} \cdot (\hat{\mathbf{m}} - i\hat{\mathbf{n}}))(\mathbf{p} \cdot (\hat{\mathbf{m}} + i\hat{\mathbf{n}})) \right] \right\} a_{p\uparrow} \\ &= \left\{ M^2(\mathbf{p}) + c_\perp^2 (\mathbf{p} \times \hat{\mathbf{l}})^2 \right\} a_{p\uparrow}, \end{aligned} \quad (27)$$

where $\hat{\mathbf{l}} := \hat{\mathbf{m}} \times \hat{\mathbf{n}}$. The eigenvalues of the evolution operator vanish only in the so-called Fermi points in momentum space⁵,

$$\mathbf{p}_{F,\pm} = \pm p_F \hat{\mathbf{l}}, \quad (28)$$

as these eigenvalues are given by the square root of

$$M^2(\mathbf{p}) + c_\perp^2 (\mathbf{p} \times \hat{\mathbf{l}})^2. \quad (29)$$

⁵ Also often called Weyl points. We use the term 'Fermi point' in accordance with [10]. 'Fermi point' can be understood as the generic term for topological point nodes, which includes Weyl points when the underlying manifold is 3+1 dimensional, and Dirac points for 2+1 dimensions. ³He-A is then an example of the Weyl category of Fermi points.

We can now see that the dispersion relation is relativistic near these points in momentum space. Considering a plane wave with momentum $\mathbf{p} = qp_F \hat{\mathbf{l}} + \mathbf{p}$ with $q = \pm 1$ and \mathbf{p} as a small deviation with respect to the corresponding Fermi point, we obtain the frequency

$$\omega^2 = c_{\parallel}^2 \mathbf{p}_l^2 + c_{\perp}^2 (\mathbf{p}_m^2 + \mathbf{p}_n^2), \quad (30)$$

where $c_{\parallel} = p_F / m^*$ (recall that $p_F = \sqrt{2 m^* \mu}$). We use the subindices m, n, l to denote projections on the pairs Cartesian trihedral $\hat{\mathbf{m}}, \hat{\mathbf{n}}, \hat{\mathbf{l}}$, i.e., $\mathbf{p}_m = \mathbf{p} \cdot \hat{\mathbf{m}}$, etc. This linear dispersion relation is valid below the energy scale

$$E_{\perp} := m^* c_{\perp}^2. \quad (31)$$

A look at equation (30) reveals that the parameters c_{\parallel} and c_{\perp} correspond to the propagation velocity of low-energy quasiparticles in the directions parallel and perpendicular to the anisotropy axis, respectively.

Now we diagonalize the Hamiltonian concentrating on two regions in momentum space surrounding the two Fermi points (these regions contain the real low-energy excitations of the system). In other words, we shall find new annihilation operators $\alpha_{p\alpha q}$, over which the action of the Hamiltonian is diagonal. As labels for these operators, we use the deviation \mathbf{p} with respect to any of the Fermi points, $\mathbf{p} = \pm p_F \hat{\mathbf{l}} + \mathbf{p}$, the spin index α , and a subscript $q = u$, with d indicating the Fermi point near which it is localized (in momentum space): the u Fermi point ($+p_F \hat{\mathbf{l}}$) or the d Fermi point ($-p_F \hat{\mathbf{l}}$). In terms of the label \mathbf{p} , this leads to an apparent doubling of the degrees of freedom. Alternatively, one can work directly with operators $a_{p\alpha q}$ (the only difference between these two sets of operators is a linear change of basis) defined as

$$a_{\mathbf{p}\uparrow u} := a_{p_F \hat{\mathbf{l}} + \mathbf{p}, \uparrow}, \quad a_{\mathbf{p}\uparrow d} := a_{-p_F \hat{\mathbf{l}} + \mathbf{p}, \uparrow}. \quad (32)$$

Focusing first on the \uparrow spin projection, we can write the corresponding equations of motion as

$$\begin{aligned} i\dot{a}_{\mathbf{p}\uparrow u} &= c_{\parallel} \mathbf{p}_l a_{\mathbf{p}\uparrow u} - c_{\perp} (\mathbf{p}_m - i\mathbf{p}_n) a_{-\mathbf{p}\uparrow d}^{\dagger}, \\ i\dot{a}_{-\mathbf{p}\uparrow d}^{\dagger} &= -c_{\parallel} \mathbf{p}_l a_{-\mathbf{p}\uparrow d}^{\dagger} - c_{\perp} (\mathbf{p}_m + i\mathbf{p}_n) a_{\mathbf{p}\uparrow u}, \end{aligned} \quad (33)$$

remembering that the equality sign is strictly speaking an approximately equal sign, and that it is not valid for momenta too far from the Fermi point. The two previous equations can be written in a compact manner as

$$i\partial_t \chi_{\mathbf{p}\uparrow} = \mathcal{H}_{\mathbf{p}\uparrow} \chi_{\mathbf{p}\uparrow}, \quad \chi_{\mathbf{p}\uparrow} = \begin{pmatrix} a_{\mathbf{p}\uparrow u} \\ a_{-\mathbf{p}\uparrow d}^{\dagger} \end{pmatrix}, \quad (34)$$

with

$$\mathcal{H}_{\mathbf{p}\uparrow} := c_{\parallel} \mathbf{p}_l \sigma_3 - c_{\perp} \mathbf{p}_m \sigma_1 - c_{\perp} \mathbf{p}_n \sigma_2. \quad (35)$$

This is a linear spinor equation for a Weyl spinor with helicity $+$ (calculated as the product of the three factors ± 1 that appear in front of the Pauli matrices). Before continuing, let us notice that the evolution equations for all the $a_{p\uparrow}$ in equation (25) are not linear in the complex plane due to the presence of complex conjugate terms. However, they have a different quasilinear symmetry. The system is invariant if one multiplies the $a_{p\uparrow}$ with \mathbf{p} in the u hemisphere by a

complex constant c and those in the d hemisphere by its complex conjugate c^* . This symmetry has allowed us to write a linear equation for the previous spinor $\chi_{p\uparrow}$. This spinor contains information about both Fermi points.

The same arguments can be applied to the \downarrow projection of the real (atomic) spin of Landau's quasiparticles to obtain

$$\chi_{p\downarrow} := \begin{pmatrix} a_{p\downarrow u} \\ a_{-p\downarrow d}^\dagger \end{pmatrix}, \quad (36)$$

and Hamiltonian operator

$$H_{p\downarrow} := c_{\parallel} p_l \sigma_3 + c_{\perp} p_m \sigma_1 - c_{\perp} p_n \sigma_2, \quad (37)$$

with helicity -1 in this case. Notice that the only difference between the two spin projections is a sign accompanying \hat{m} in the order parameter [see also equations (25) and (26)], which translates into a different helicity in the low-energy corner. For this reason, the atomic spin projection index can be thought of as a helicity index for the low-energy Bogoliubov excitations. We will explicitly check this later.

As a final step, one can arrange the two spinors to form a bispinor that obeys the following evolution equation:

$$i\partial_t \begin{pmatrix} \chi_{p\uparrow} \\ \chi_{p\downarrow} \end{pmatrix} = e_b^a Y^b p_a \begin{pmatrix} \chi_{p\uparrow} \\ \chi_{p\downarrow} \end{pmatrix}, \quad (38)$$

with $a, b = 1, 2, 3$ and

$$Y^1 = \begin{pmatrix} -\sigma_1 & 0 \\ 0 & \sigma_1 \end{pmatrix}, \quad Y^2 = \begin{pmatrix} -\sigma_2 & 0 \\ 0 & -\sigma_2 \end{pmatrix}, \quad Y^3 = \begin{pmatrix} \sigma_3 & 0 \\ 0 & \sigma_3 \end{pmatrix}. \quad (39)$$

The only non-zero components of e^b_a are

$$e^1_{\perp 1} := c_{\perp}, \quad e^2_{\perp 2} := c_{\perp}, \quad e^3_{\parallel 3} := c_{\parallel}. \quad (40)$$

Now we can find a matrix X such that the set $\{X, XY^1, XY^2, XY^3\}$ is a representation of the Dirac matrices. Considering that the matrices (39) verify the properties

$$(Y^a)^2 = I_4, \quad \{Y^a, Y^b\} = 2\delta^{ab}I_4, \quad (41)$$

(with I_4 the 4×4 identity), which follow directly from the properties of the Pauli matrices, such a matrix X must verify

$$X^2 = I_4, \quad \{X, Y^a\} = 0. \quad (42)$$

One can verify a solution to these equations is given by

$$X = \begin{pmatrix} 0 & \sigma_1 \\ \sigma_1 & 0 \end{pmatrix}. \quad (43)$$

The corresponding representation of the Dirac matrices is

$$\begin{aligned}\gamma^0 &= \begin{pmatrix} 0 & \sigma_1 \\ \sigma_1 & 0 \end{pmatrix}, & \gamma^1 &= \begin{pmatrix} 0 & I_2 \\ -I_2 & 0 \end{pmatrix}, \\ \gamma^2 &= \begin{pmatrix} 0 & -i\sigma_3 \\ -i\sigma_3 & 0 \end{pmatrix}, & \gamma^3 &= \begin{pmatrix} 0 & -i\sigma_2 \\ -i\sigma_2 & 0 \end{pmatrix}.\end{aligned}\quad (44)$$

Therefore, we can conclude that the low-energy excitations of the planar phase are massless Dirac spinors in Minkowski spacetime, thus satisfying the evolution equation

$$e_f^\mu \gamma^I \bar{\mathbf{p}}_\mu \psi_p = 0, \quad \psi_p := \begin{pmatrix} \chi_{p\uparrow} \\ \chi_{p\downarrow} \end{pmatrix}, \quad (45)$$

where we have taken the Fourier time transform and defined $\bar{\mathbf{p}}^\mu := (\omega, \mathbf{p})$. The components of the tetrad e_f^μ , $\mu = 1, 2, 3, 4$, $I = 1, 2, 3, 4$ are given by equation (40), complemented by

$$e^0_0 := 1. \quad (46)$$

Spacetime is therefore Minkowskian because the tetrad field has been taken to be constant. The constant velocities c_{\parallel} and c_{\perp} can be absorbed into a rescaling of the coordinates. This laboratory anisotropy would in any case be unobservable for low-energy ‘internal’ observers (see the next section and [33]).

The occurrence of four components in the low-energy fermionic object ψ (whose Fourier components are defined in equation (45)) is tied up to the existence of two degrees of freedom, one for each Fermi point, for each projection of the spin. The spin projection must be considered as the helicity eigenvalue in the low-energy description: let us evaluate the chirality operator in this representation

$$\gamma^5 := i\gamma^0\gamma^1\gamma^2\gamma^3 = \begin{pmatrix} I_2 & 0 \\ 0 & -I_2 \end{pmatrix}. \quad (47)$$

That is, χ_{\uparrow} and χ_{\downarrow} have a well-defined chirality, as

$$\frac{1 + \gamma^5}{2} \psi = \psi_{\uparrow} = \begin{pmatrix} \chi_{\uparrow} \\ 0 \end{pmatrix}, \quad \frac{1 - \gamma^5}{2} \psi = \psi_{\downarrow} = \begin{pmatrix} 0 \\ \chi_{\downarrow} \end{pmatrix}. \quad (48)$$

Summarizing, the natural low-energy excitations that appear in this model are massless Dirac fermions. To reproduce electrodynamics of electrons and positrons, one would need to generate some small mass gap for the excitations. This might require a more complicated system, with, for example, some Yukawa couplings, and is beyond the analysis carried out here.

3.5. Internal low-energy observers and the emergence of charge

We have seen that the concept of atomic spin acquires a different meaning in the low-energy corner in which the fermionic quasiparticles are described by Dirac’s theory: they play the role of a charge label. The Dirac equation in equation (45) is invariant under a U(1) transformation of ψ (in fact, it is invariant under transformations of ψ_{\uparrow} and ψ_{\downarrow} separately). The corresponding conserved charge is

$$Q := Q_{\uparrow} + Q_{\downarrow} = N_u - N_d, \quad (49)$$

where the operators N_u and N_d represent the number of excitations associated with the positive Fermi point (with positive $q = 1$ charge) and with the negative Fermi point (with negative $q = -1$ charge), respectively. As we will explain, a notion of charge has emerged in the low-energy theory, owing to the duplicity of Fermi points. In fact, we will see later that when coupling these Dirac quasiparticles to an effective electromagnetic field, this charge plays the role of an electric charge.

For an external observer (or laboratory observer), charge conservation is nothing more than momentum conservation. Imagine a scattering process involving two quasiparticles from *the same Fermi point*. Momentum conservation only tells us that

$$\mathbf{p}_1 + \mathbf{p}_2 = 2qp_F \hat{\mathbf{l}} + \mathfrak{p}_1 + \mathfrak{p}_2 = \mathbf{p}_3 + \mathbf{p}_4. \quad (50)$$

As the products of the scattering event must be quasiparticles, the solutions of this equation must be of the form

$$\mathbf{p}_3 = qp_F \hat{\mathbf{l}} + \mathfrak{p}_3, \quad \mathbf{p}_4 = qp_F \hat{\mathbf{l}} + \mathfrak{p}_4. \quad (51)$$

The momentum conservation condition is thus equivalent to the conservation of the deviations \mathfrak{p} :

$$\mathfrak{p}_1 + \mathfrak{p}_2 = \mathfrak{p}_3 + \mathfrak{p}_4. \quad (52)$$

If the scattering is instead between quasiparticles from *different Fermi points*, the conservation of momentum reads

$$\mathbf{p}_1 + \mathbf{p}_2 = \mathfrak{p}_1 + \mathfrak{p}_2 = \mathbf{p}_3 + \mathbf{p}_4, \quad (53)$$

which implies that

$$\mathbf{p}_3 = qp_F \hat{\mathbf{l}} + \mathfrak{p}_3, \quad \mathbf{p}_4 = -qp_F \hat{\mathbf{l}} + \mathfrak{p}_4. \quad (54)$$

The resulting quasiparticles must live in different Fermi points. Again, the momentum conservation condition is then equivalent to the conservation of the deviations \mathfrak{p} ,

$$\mathfrak{p}_1 + \mathfrak{p}_2 = \mathfrak{p}_3 + \mathfrak{p}_4. \quad (55)$$

In both cases, the energy-conservation condition is equivalent at low energies to the usual massless relativistic condition in the deviations \mathfrak{p} , so it does not distinguish between the two types of processes.

However, in the low-energy description, the conserved charge is associated with an intrinsic property of the Dirac field and its conservation has nothing to do, from this point of view, with momentum conservation. The effective duplication of degrees of freedom in the low-energy theory with respect to the degrees of freedom in the initial quasiparticle field is just apparent. This is apparent in figure 4.

In this paper, we construct an effective low-energy world that cannot be distinguished operationally from the world of electrodynamics. We therefore must discuss the important meaning of internal observer. In relation with the emergence of charge, here we can already discuss the difference between two types of potential internal observers. One can be called an internal Fermi-point observer and the other an internal low-energy observer.

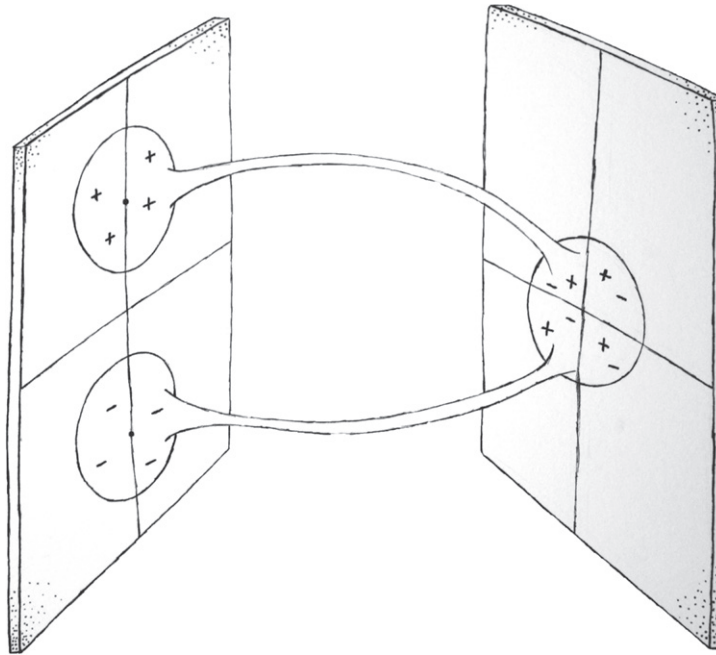


Figure 4. Diagram explaining the effective duplication of the degrees of freedom. The left-hand side shows the quasiparticles living near the two Fermi points. The right-hand side shows the effective low-energy description for an internal observer. An internal observer is insensitive to the origin (upper or lower Fermi point) of the quasiparticles in terms of momentum, but sees them as having opposite charges.

An internal Fermi-point observer is an observer living in one specific Fermi point q . We can associate a momentum $qp_F \hat{\mathbf{I}}$ to that observer. This observer sees the momentum region as a low-energy world full of spinor waves (these will not be Weyl spinors but specific superpositions of them). This observer's world would have half the degrees of freedom compared to the Dirac bispinors. In addition, this observer will see quasiparticles coming from the other Fermi point, which will have a tremendous relative momentum $2qp_F$, although with low energies. To obtain a standard electrodynamic world for these kind of observers, our model lacks two ingredients: (i) the quasiparticles from different Fermi points should not interact with each other (in the model, we are discussing they do) and (ii) one should duplicate in some way the number of degrees of freedom associated with that Fermi point (such as producing some fragmented Fock state condensation).

An internal low-energy observer on the other hand is an observer who sees *all* the low-energy excitations. It is reasonable that they will use as a natural momentum label the deviations \mathbf{p} . These momenta can properly describe the scattering events between all the quasiparticles as long as these observers confer an additional property to these quasiparticles, which is conserved in the interaction process. This property is charge, even though for the external observers this is nothing but the difference between the quasiparticle number around both Fermi points, as we have seen.

Let us discuss one final point regarding the nature of the low-energy excitations of the system (the previously described spin waves or Dirac quasiparticles). Consider a spin wave with exactly the Fermi point momentum

$$e^{ip_F \hat{l} \cdot \mathbf{x} / \hbar}. \quad (56)$$

We have seen that this oscillation pattern carries no energy. This oscillation is stationary so it cannot carry momentum either. This oscillation pattern is, in reality, part of the vacuum state. If the momentum of these spin waves has a small departure \mathbf{p} from the Fermi point momentum, then we have seen that they do carry an energy $E \simeq c|\mathbf{p}|$. As we have explained, the effective spacetime is Minkowskian, so the anisotropic velocities in the laboratory will not have any operational meaning for low-energy internal observers. We will just define a single constant c relating their space and time dimensions. The momentum carried by one such wave is precisely \mathbf{p} . Its direction marks the direction of the propagation of the spin wave. Its modulus can be seen as derived from E/c . Therefore, the momentum $p_F \hat{l} + \mathbf{p}$ is not the real momentum carried by the spin wave, relevant to experiments measuring impulse transfers within the liquid. The real momentum of the spin wave is \mathbf{p} .

A momentum $|\mathbf{p}|$ has an associated wavelength $\lambda = 2\pi\hbar/|\mathbf{p}|$. These wavelengths, which are much larger than the mean interparticle distance $\lambda_1 = 2\pi\hbar/p_F$, are the actual ‘observable’ wavelengths of the spin waves in the liquid. In the case of a classical (large amplitude) spin wave, this ‘observability’ will match our intuitive sense of observation of a wave.

3.6. Textures and electromagnetic fields

Let us analyze now what happens in the case in which the order parameter, instead of being completely homogeneous, contains a perturbation in position space with respect to the situations of the previous subsection. Such inhomogeneities in this kind of order parameters are typically called textures. These variations develop over a scale that is large compared to the effective size of a Cooper pair, or healing length, whose zero-temperature limit is

$$\xi_0 := \frac{\hbar c_{\parallel}}{p_F c_{\perp}} \simeq \frac{\hbar p_F}{m^* k_B T_C}. \quad (57)$$

In other words, the Fourier modes of the variations of the order parameter have wave numbers given by

$$k \ll k_{\max} := \frac{2\pi}{\xi_0}. \quad (58)$$

This means that all wavelength variations must be much larger than the healing length (57), because shorter wavelength variations are not consistent with the very existence of a local order parameter. An equivalent restriction applies to the rapidity of temporal variations of the order parameter. If we define a natural time scale as $t_0 := \xi_0/c_{\perp}$, then for consistency, we must be sure the temporal variations of the order parameter are slower than this time scale.

The textures we will consider are of two kinds. The first one is usually called orbital texture [28] and is given by the bending of the direction \hat{l} of the angular momentum of the pairs. They amount to two degrees of freedom. In addition to this, the planar-phase order parameter has, in general, the possibility of rotating around the angular momentum axis. This leads to one additional degree of freedom. In simple situations, when only this rotation is present, the additional degree of freedom is just a phase from which one can define a superfluid velocity and momentum as

$$\mathbf{v}_s(\mathbf{x}) := \frac{1}{2m^*} \nabla \phi, \quad \mathbf{p}_s(\mathbf{x}) := m^* \mathbf{v}_s(\mathbf{x}). \quad (59)$$

In more complicated situations, the superfluid velocity need not be irrotational (see [34]), but the important thing is that, in any case, the superfluid velocity contributes, with one single additional degree of freedom, to the physics of the system, on top of the two degrees of freedom of $\delta\hat{l}$. In the simplest case in which \mathbf{p}_s is a constant vector, we can again work in momentum space to analyze the form of the low-energy excitations.

The selection of a specific inertial frame in which the Fermi fluid is at rest can be seen as a peculiar example of spontaneously broken symmetry. Two relatively moving states have different energies with respect to a third inertial observer (e.g., the laboratory observer). However, if there was no interaction at all between the fluid and external objects in a particular frame, there would be no physical reason to select one specific uniform fluid velocity rather than another. In practice, the tiny interactions between the fluid and some specific inertial environment (typically the laboratory environment/frame) induces this very frame as the rest frame of the fluid. Then, the condensed vacuum state incorporates this same frame selection: the pairs are at rest with respect to this specific frame.

In what follows we take the operational view that a specific frame with a constant velocity \mathbf{v}_s with respect to the laboratory has been selected, regardless of the origin of this selection. This means that the pairing has occurred between $\mathbf{p} + \mathbf{p}_s$ and $-\mathbf{p} + \mathbf{p}_s$ atoms. This implies that equations (25) and (26) should now be written as

$$i\dot{a}_{\mathbf{p}+\mathbf{p}_s,\uparrow} = \left[M(\mathbf{p}) + \mathbf{p} \cdot \mathbf{v}_s \right] a_{\mathbf{p}+\mathbf{p}_s,\uparrow} - c_{\perp} \mathbf{p} \cdot (\hat{\mathbf{m}} - i\hat{\mathbf{n}}) a_{-\mathbf{p}+\mathbf{p}_s,\uparrow}^{\dagger}, \quad (60)$$

$$i\dot{a}_{\mathbf{p}+\mathbf{p}_s,\downarrow} = \left[M(\mathbf{p}) + \mathbf{p} \cdot \mathbf{v}_s \right] a_{\mathbf{p}+\mathbf{p}_s,\downarrow} + c_{\perp} \mathbf{p} \cdot (\hat{\mathbf{m}} + i\hat{\mathbf{n}}) a_{-\mathbf{p}+\mathbf{p}_s,\downarrow}^{\dagger}. \quad (61)$$

To reach these equations, one needs to perform an active Galilean transformation under which any momentum label is shifted by $+\mathbf{p}_s$, and take into account the transformation laws for the different objects appearing in the evolution operator. This is best understood from the Galilean transformation of the grand canonical Hamiltonian (11). Recall that the potential term is invariant under such transformation and that the kinetic term acquires two extra terms: a Doppler contribution $\mathbf{p} \cdot \mathbf{v}_s$ and a global shift $\mathbf{p}_s^2 / (2m^*)$, which can be absorbed in the chemical potential for the moving system [35],

$$\bar{\mu} = \mu + \frac{\mathbf{p}_s^2}{2m^*}; \quad (62)$$

$$M(\mathbf{p}) + \mathbf{p} \cdot \mathbf{v}_s = \frac{(\mathbf{p} + \mathbf{p}_s)^2}{2m^*} - \bar{\mu}. \quad (63)$$

If we consider that the pairing channel is given by $\mathbf{p}_1 + \mathbf{p}_2 = \mathbf{p}_3 + \mathbf{p}_4 = 0$, we immediately reach the conclusion that the transformed pairing Hamiltonian is precisely equation (17) with a Doppler shift and with all the labels shifted by \mathbf{p}_s . Finally, the order parameter is just bookkeeping the statistics of the pairs and hence depends only on the relative momenta of the members of the pairs and not on their global motion, which means that the order parameter is unchanged by the Galilean transformation.

As an alternative to the treatment in section 3.4, we will first combine the excitations in a bispinor and then concentrate on the excitations close to the Fermi points. The results are independent of the order of operations, but in this case it is simpler to proceed this way. Starting with the \uparrow spin projection, the equations of motion are given by:

$$i\partial_t \begin{pmatrix} a_{\mathbf{p}+\mathbf{p}_s, \uparrow} \\ a_{-\mathbf{p}+\mathbf{p}_s, \uparrow}^\dagger \end{pmatrix} = H_{\mathbf{p}, \mathbf{p}_s, \uparrow} \begin{pmatrix} a_{\mathbf{p}+\mathbf{p}_s, \uparrow} \\ a_{-\mathbf{p}+\mathbf{p}_s, \uparrow}^\dagger \end{pmatrix}, \quad (64)$$

where

$$H_{\mathbf{p}, \mathbf{p}_s, \uparrow} := M(\mathbf{p})\sigma_3 + \mathbf{p} \cdot \mathbf{v}_s \sigma_0 - c_\perp \mathbf{p}_m \sigma_1 - c_\perp \mathbf{p}_n \sigma_2. \quad (65)$$

Similar manipulations with the \downarrow spin projection permit us to write the following evolution equation for bispinors:

$$i\partial_t \begin{pmatrix} a_{\mathbf{p}+\mathbf{p}_s, \uparrow} \\ a_{-\mathbf{p}+\mathbf{p}_s, \uparrow}^\dagger \\ a_{\mathbf{p}+\mathbf{p}_s, \downarrow} \\ a_{-\mathbf{p}+\mathbf{p}_s, \downarrow}^\dagger \end{pmatrix} = H_{\mathbf{p}, \mathbf{p}_s} \begin{pmatrix} a_{\mathbf{p}+\mathbf{p}_s, \uparrow} \\ a_{-\mathbf{p}+\mathbf{p}_s, \uparrow}^\dagger \\ a_{\mathbf{p}+\mathbf{p}_s, \downarrow} \\ a_{-\mathbf{p}+\mathbf{p}_s, \downarrow}^\dagger \end{pmatrix}, \quad (66)$$

where the 4×4 evolution operator is

$$H_{\mathbf{p}, \mathbf{p}_s} := M(\mathbf{p})Y^3 + \mathbf{p} \cdot \mathbf{v}_s Y^0 + c_\perp \mathbf{p}_m Y^1 + c_\perp \mathbf{p}_n Y^2. \quad (67)$$

The set of matrices $\{Y^1, Y^2, Y^3\}$ were defined in equation (39), and $Y^0 := I_4$. Now if we concentrate on excitations near the Fermi points (i.e., linearize around $\mathbf{p} = +p_F \hat{\mathbf{l}}$; this linearization is sufficient as the equation is already encompassing all the degrees of freedom), one obtains a Dirac equation in momentum space:

$$\left(e_f^\mu \gamma^I \bar{\mathbf{p}}_\mu + \gamma^0 p_F \hat{\mathbf{l}} \cdot \mathbf{v}_s \right) \psi_p = 0, \quad (68)$$

where ψ_p is the bispinor in (66), $\mathbf{p} = \mathbf{p} - p_F \hat{\mathbf{l}}$, and $\bar{\mathbf{p}}^\mu = (\omega, \mathbf{p})$. The non-zero components of the tetrad e_f^μ are given by

$$e^0_0 := 1, \quad e^1_1 := c_\perp, \quad e^2_2 := c_\perp, \quad e^3_3 := c_\parallel, \quad e^i_0 := v_s^i. \quad (69)$$

The corresponding metric components are

$$g^{\mu\nu} = \eta^{IJ} e_f^\mu e_f^\nu, \quad \rightarrow \quad g^{\mu\nu} = \begin{pmatrix} -1 & -v_s^i \\ -v_s^i & D^{ij} - v_s^i v_s^j \end{pmatrix}, \quad (70)$$

with

$$D^{ij} = \begin{pmatrix} c_{\perp}^2 & 0 & 0 \\ 0 & c_{\perp}^2 & 0 \\ 0 & 0 & c_{\parallel}^2 \end{pmatrix}. \quad (71)$$

This is an acoustic metric [36] which, given that we are assuming a uniform background velocity \mathbf{v}_s , corresponds to a flat Minkowski spacetime. Equation (68) is completely equivalent to equation (45) in the homogeneous case. The only difference is a constant shift in the energy of quasiparticles.

To discuss inhomogeneities, it is better to work in position space. Equation (68) would then be written as

$$e^{\mu} \gamma^I (i\partial_{\mu} - B_{\mu})\psi = 0, \quad (72)$$

where $\partial_{\mu} := (\partial_t, \nabla)$ is the derivative operator including time, and $B_{\mu} := (p_{\text{F}} \hat{\mathbf{l}} \cdot \mathbf{v}_s, p_{\text{F}} \hat{\mathbf{l}})$ is a constant background. The content of this equation is then completely equivalent to one with $B_{\mu} = 0$, because a constant background value can be absorbed into unobservable offsets of energy and momentum. However, here we consider fluctuations of the background $\delta \hat{\mathbf{l}}$ and $\delta \mathbf{v}_s$. These fluctuations act as an effective vector field, which affects the evolution of quasiparticles:

$$e^{\mu} \gamma^I (i\partial_{\mu} - \nu \bar{A}_{\mu})\psi = 0. \quad (73)$$

Here, ν is a constant that controls the dimensions of the field \bar{A}_{μ} (recall that in standard electrodynamics, the vector potential has dimensions of momentum per unit charge). The kind of coupling of the fermionic quasiparticles to the vector field \bar{A}^{μ} suggests that we identify it as an effective electromagnetic gauge field, as in other inhomogeneous situations in condensed matter physics (see [37]). However, we should put the metric in its standard Minkowskian form before carrying out this identification. To do so, we transform to comoving coordinates so that the $v_s^i \partial_i$ term in Dirac's equation (73) vanishes. Then the vector field is identified as

$$\bar{A} := \frac{1}{\nu} \delta (p_{\text{F}} \hat{\mathbf{l}}), \quad \bar{A}_0 = \frac{1}{\nu} p_{\text{F}} \hat{\mathbf{l}} \cdot \delta \mathbf{v}_s. \quad (74)$$

The object \bar{A} is a genuine vector, with three degrees of freedom: two originate from the variations $\delta \hat{\mathbf{l}}$ of the order parameter, and the other one from density fluctuations δp_{F} . On the other hand, \bar{A}_0 contains just one degree of freedom independent of these.

To end this subsection, let us also note the inhomogeneities in the order parameter make the acoustic metric (70) non-flat. Thus, when considering higher-than-first-order effects, the same degrees of freedom making up this effective electromagnetic potential will be responsible for some partial curved-spacetime effects.

3.7. Gauge symmetry and dynamics

Our discussion so far has shown that the low-energy description of the system contains features that are not included in the original theory, such as the notion of electric charge and chirality. In this section, we address another emergent property: gauge symmetry. When gauge fields are emergent entities, the discussion naturally splits in two aspects: on the one hand, the

kinematical invariance of the theory under gauge transformations and, on the other hand, the *dynamical* preservation of this symmetry [38, 39]. The study of analog gravity setups, where the relevant gauge group is composed by diffeomorphisms, has shown that condensed-matter analogies usually fail to achieve the second point [36, 40]. This section is devoted to an analysis of these issues in the context of the model developed in this article, where the gauge group is simpler.

By kinematical gauge invariance, we refer to a property of the way in which the low-energy quasiparticle excitations, the Nambu–Gor’kov spinors in equations (34) and (36), react under the presence of different given fields \bar{A}_μ , independent of their origin. As we have seen, the fields \bar{A}_μ are associated with spatial and temporal variations of the orbital part of the order parameter, which is represented by a trihedral $\{\hat{m}, \hat{n}, \hat{l}\}$. Kinematic gauge invariance occurs when there are equivalent classes of \bar{A}_μ leading to essentially the same effect upon the quasiparticles.

Consider, as an example, the following static texture:

$$\delta\hat{m} = \delta\hat{m}(\mathbf{x} \cdot \hat{m}_0), \quad \delta\hat{n} = \delta\hat{n}(\mathbf{x} \cdot \hat{n}_0), \quad \delta\mathbf{p}_s = 0. \quad (75)$$

In this case, one can find a local phase transformation of the ermionic fields that transforms the evolution equation (73) into a free Dirac equation for the new spinor field. That is, for internal observers, the configuration (75) would be equivalent to the absence of textures if they identify the physical objects with equivalence classes defined by these gauge transformations. A spinor field wave packet is not deflected in any way by the previous texture and one could consider that as a defining feature of the equivalent class of configurations. As we could have anticipated, two textures differing in the gradient of a scalar, $\bar{A}'_\mu - \bar{A}_\mu = \partial_\mu\varphi$, lead to the same type of effects in the spinor field; the function φ can always be absorbed locally into the spinor’s local phase:

$$\psi \longrightarrow \exp[i\varphi(t, \mathbf{x})]\psi. \quad (76)$$

Recall that, in the same way, in Maxwell’s model (section 2 and [14, 15]), electromagnetic potentials have also a reality but some of their properties are not relevant at low energies. At this point, it is important to remark again that this picture is only partial: the description of the system is simple because we are looking only at low-energy phenomena. In particular, this gauge invariance will be violated at some point when the low-energy description breaks down, for instance, when the effective Lorentz invariance disappears. At some point, even the condensation and thus the very existence of the field \bar{A}_μ would disappear. Moreover, we are not considering the excitation of other collective modes, assuming that they are frozen. To derive the low-energy Dirac equation (73) with vector potential (74), we assumed that these other collective modes are not excited (for example, the clapping modes that can be associated with gravitons; see [41] for a general discussion of the different collective modes and their significance, and [42] for the surprising relation between these clapping modes and the effective cosmological constant in $^3\text{He-A}$).

Let us now discuss the issue of dynamical gauge invariance. In principle, it could be the case that the kinematical gauge invariance was not preserved by the dynamics. By looking at the interaction of two spinor wave packets, through a mediator field \bar{A}_μ , one could detect differences beyond the introduction of a local phase. This amounts to the possibility of distinguishing between different members of the kinematical equivalence class. The emergence

of a dynamical gauge invariance will definitely signal the irrelevance of certain degrees of freedom of \bar{A}_μ in the self-consistent low-energy physics of the system.

The obtention of a dynamical gauge invariance turned out to be an issue much more subtle than expected. Following the literature, one is first naturally led to believe that the key of the question resides in an induction mechanism in the manner of Sakharov [43] (adapted to electromagnetism by Zel'dovich [44]). However, we found this path to be paved with problems (for a summary of those problems, see appendix C). Rethinking the problem, we realized that the key may well reside in the very emergence of Lorentz invariance. Here, we describe the logic of the emergence of dynamical gauge invariance along this line of thought.

Thus far, we have not inquired about the origin of the inhomogeneities in the order parameter, or in other words, of the \bar{A}_μ fields. For instance, a specific texture could be forced upon the system by using external forces. However, here we want a closed self-consistent system. Then, at low energy, the inhomogeneities of the order parameter can only be induced by the presence of fermionic quasiparticles. For that, it is useful to think of coherent states of fermionic quasiparticles, i.e., macroscopic spin waves. The source of \bar{A}_μ exhibits Lorentz invariance below the energy scale (31). We used the notation \bar{A}_μ to denote the objects that appear directly in Dirac's equation when written in comoving Lab coordinates. From the Lab perspective, we know that \bar{A}_i is a vector and \bar{A}_0 a scalar under spatial changes of coordinates. However, the solutions of this Dirac equation under the presence of \bar{A}_μ will be connected with those worked out in a Lorentz transformed coordinate frame but now under the presence of a Lorentz transformed object $A_\nu = \Lambda_\nu^\mu \bar{A}_\mu$. From the kinematic perspective, internal low-energy observers of quasiparticles are not able to pick out any special or privileged Minkowskian inertial frame. Thus, they will construct their world view using an, in principle, generic object A_μ with the transformation properties of a Lorentz covariant four-vector.

However, one question is how the object A_μ transforms under a Lorentz change of coordinates. Another question is whether the entire system is invariant under active Lorentz transformations or, in other words, whether the dynamics of A_μ is Lorentz invariant. As we said previously, we are allowed to consider only those A_μ produced by the presence of quasiparticles. It is then a reasonable hypothesis that two Lorentz-related sources lead to two Lorentz-related A_μ s, i.e., that the Lorentz covariant structure of the spinor waves is passed over to the texture field. We do not have a proof of this reasonable hypothesis but, for the moment, let us assume it is indeed valid in our system. Under this hypothesis, internal observers will write down a generic Lagrangian for the system of the form

$$\mathcal{L}(\psi, A_\mu) := -\frac{1}{4\mu_0} F^{\mu\nu} F_{\mu\nu} + \frac{\xi}{2\mu_0} (\partial_\mu A^\mu)^2 + \frac{m^2}{2\mu_0} A_\mu A^\mu + \mathcal{L}_D(\psi, A_\mu). \quad (77)$$

Let us first examine the possible value of the mass constant in the effective Lagrangian density (77). A simple argument shows that m cannot be different from zero. A non-zero value for the mass parameter would mean that to create a texture, there would always have to be an energetic gap, no matter how smooth the texture may be (i.e., no matter how large the associated wavelengths). These variations of the order parameter explore the degeneracy manifold of the planar order parameter (20). Thus, in the limit of very long wavelengths, the construction of a texture should cost no energy. This is clear when we consider orbital textures, but it might appear less clear for perturbations of the superfluid velocity field. The problem is that, from the point of view of the order parameter, a constant velocity already imposes a

specific length scale for the variations of the phase. The perturbations of the velocity are encoded in second derivatives of the phase. If perturbations of the velocity field did have a gap, then, among other things, our Lorentz-invariant hypothesis would be broken, as the A_μ would have an anisotropic mass (indeed, this is what occurs in the A-phase of ^3He when spin-orbit interactions are considered [10]). However, it is well known that assuming a fixed constant velocity background, the extra energy associated with the introduction of acoustic waves is such that its dispersion relation is gapless. Therefore, there is no mass term for A_μ in this theory.

The equations of motion for $m = 0$ are

$$\square A_\mu - (1 + \xi)\partial_\mu\partial_\nu A^\nu = j_\mu. \quad (78)$$

The source of this equation of motion is the identically conserved fermionic current j_μ . If one takes the divergence in the last equation, one finds for $\xi \neq 0$, the following equation ($\xi = 0$ leads to a trivial identity):

$$\square(\partial^\mu A_\mu) = 0. \quad (79)$$

In this way, the divergence $\partial_\mu A^\mu$ effectively behaves as a free scalar field, not coupled to the rest of fields (note that when gravity is included, this is no longer true; in fact, the existence of such a scalar degree of freedom could have non-trivial cosmological implications [45]). The absence of sources for this construction makes it natural to impose a zero value of this field or, in other words, the Lorenz gauge condition

$$\partial^\mu A_\mu = 0. \quad (80)$$

Instead of working with this specific gauge fixing condition, one could choose to work from the start with a theory without the $(\partial_\mu A^\mu)^2$ term in the Lagrangian. Then, the theory will exhibit standard gauge invariance and one could proceed with any gauge fixing one likes. Both ways of proceeding will lead to the same physical results. Here, let us make an interesting observation. To construct a fundamental theory of massless relativistic spin-1 particles, it is compulsory to introduce gauge invariance (see [46]). The A_μ field cannot be observable in a theory with a fundamental Lorentz invariance (unbroken at all energies). However, when a Lorentz invariance is effective and appears only at low energies, the underlying theory can associate a physical reality to the A_μ , but the low-energy observers are oblivious to some of its properties⁶.

Regarding the value of the remaining constant, a dimensional analysis shows that the vacuum permeability would be given by the following expression:

$$\mu_0 = \frac{4\pi m^* \hbar}{\nu^2 p_F} \alpha, \quad (81)$$

where α is a dimensionless constant. The notation is not accidental: it corresponds to the effective fine-structure constant of the theory. Moreover, the only dimensionless quantity one can construct from the constants in the problem is the quotient c_\perp/c_\parallel . This means that the

⁶ It is interesting to notice that, in the case in which the gauge potential is identified with the four-vector describing the flow of matter, both the Lorenz and Coulomb gauge-fixing conditions can be physically interpreted in terms of properties of the underlying fluid [23, 24]. However, in this work, we propose a different identification of the relevant vector gauge field at low energies.

effective fine-structure constant must be calculable as a function of this quantity, i.e.,

$$\alpha = \alpha \left(\frac{c_{\perp}}{c_{\parallel}} \right). \quad (82)$$

This is all we can say with certainty about this quantity. Just as in effective field theories, the concrete form of the fine-structure constant must be determined by comparing a process (e.g., scattering of two quasiparticles) in both the low-energy theory and the condensed-matter theory with Hamiltonian (17), in which all the constants are explicit. However, the occurrence of the condensation could hinder this comparison, as it implies a non-trivial resummation of the perturbative contributions at different orders. This is beyond the scope of this paper, although interesting. Notice that it is natural to expect a behavior which guarantees that this scattering amplitude tends to zero when there is no interaction between fermions [$g \rightarrow 0$ limit in (17)].

What about the very presence of Lorentz invariance in the A_{μ} sector? We do not have a definitive argument that this should be the case. Imagine, for example, that the A_{μ} field could propagate faster than the effective speed of light as defined by the fermionic Lorentz symmetry. Then, the condensate would probably be unstable, as it would be energetically favorable for the particles \mathbf{p} , $-\mathbf{p}$ in the pair to become unpaired quasiparticles (producing some sort of Cherenkov radiation). Another argument is that the inhomogeneous perturbations of the condensed state might be seen as a coherent field of particle pairs moving on top of a homogeneous background condensate. Within that interpretation, it would be natural to expect that the velocity of these pairs would follow the same dispersion relation as their free particle cousins.

In summary, if a Lorentz invariance appears below the energy scale (31), as it is tied up to the existence of Fermi points, the resulting dynamical theory would be gauge invariant and thus indistinguishable from standard electromagnetism. From the perspective of electrodynamics as a fundamental theory, the imposition of Lorentz invariance and the fact that the interaction is mediated by a massless vector field that couples to the fermion current density are necessary and sufficient to obtain a gauge-invariant theory. This extends to the emergent scenario, thus fixing these as the relevant conditions one must set up to completely reproduce electrodynamics at low energies. One can wonder whether this result, i.e., the secondary character of the principle of gauge invariance, is particular to electrodynamics or not. In this respect, a detailed study of the non-abelian case and/or higher-spin fields (especially the spin-2 case) will be presented elsewhere.

4. Summary and conclusions

In this work, we recommend understanding electrodynamics from an emergent point of view. We do not commit with the specifics of the models presented. On the contrary, we want to focus on the generic characteristics of these models, those likely to be shared by any emergent model of electrodynamics. We also tried to convey the power contained in emergent constructions: very simple elementary components and interactions can lead to an enormously rich phenomenology, as described by the effective theory.

Although the constituents of both models are of different nature, it is not difficult to draw parallels. Maxwell's proposal contained two kinds of elements: vortical cells, whose most salient property is their ability to acquire rotation, and ball bearings, from which one constructs

the analog of charged matter. These two elements are also present in the model inspired by ${}^3\text{He}$. The role of movable ball bearings is now played by fermionic quasiparticles, low-energy excitations of a fundamental system of fermions subject to particular kinds of interactions. These low-energy excitations, or quasiparticles, evolve following Dirac's equation. On the other hand, when the fundamental fermions are paired up and condensed, they act as vortical cells, which possess intrinsic rotational properties due to the finite value of angular momentum characteristic of the p -wave condensation. The electromagnetic fields analyzed here arise as the coarse-grained view of these effective bosons, i.e., as perturbations of the condensed phase.

In both models, the velocity of light is emergent. Because both theories have been formulated as Galilean theories, there is, in principle, no obstruction for the elementary component to travel at arbitrarily large velocities. The speed of light appears as a 'sound' speed; the velocity of wave-like excitations in the system. In the case of the superfluid model, this velocity and its independence of the wavelength is strongly tied up to the occurrence of Fermi points where the dispersion relation becomes relativistic. The physics could all be described by a privileged external observer by using Newtonian notions. However, internal low-energy observers would tend to develop ways to understand their low-energy world that do not assume external structures. This epistemological choice is certainly valid, but would require the necessary assumption of some features as fundamental principles, and hence, a loss of explanatory power.

Another interesting parallelism is that, in both cases, the electromagnetic potential has a physical reality in terms of specific properties of the system under consideration. It is only at low energies that some of these degrees of freedom become effectively invisible and the internal gauge symmetry appears.

Beyond these parallelisms, the superfluid model goes further than Maxwell's model.

- (i) While in Maxwell's construction, the two substances making up the system, charged matter and electromagnetic fields, are independently postulated. In the superfluid framework, they arise from the same single set of underlying elements.
- (ii) The superfluid model can consider the spinorial and quantum-mechanical properties of matter. The notion of charge cannot emerge here from the quantum-mechanical quasiparticle density that is always non-negative ($\psi^\dagger\psi \geq 0$). However, two signs existing for the charge is a nice logical consequence of the appearance in pairs of the Fermi points.
- (iii) Moreover, it seems possible to include quantum features of the electromagnetic field in the ${}^3\text{He}$ -like model. Individual photons would correspond to tiny fluctuations of the condensed phase—so tiny that they involve only one of these effective bosons composed by a pair of fermions. This model suggests photons should not be viewed as fundamental particles, but as composite structures emerging from the same fundamental ingredients as the fermionic quasiparticles (there are other examples in the literature in which photons and electrons arise from the same underlying system, although in those constructions even Fermi–Dirac statistics are emergent [47]). At this point, this is only a (natural) conjecture. But in future work, we plan to analyze to what extent this can be rigorously formulated.

Emergent views of the kind analyzed in this paper always imply that the low-energy properties, for instance Lorentz invariance, will eventually break up at some high-energy scale. Thus, it is important to stress here that deviations from Lorentz invariance need not occur at the Planck scale (and indeed Lorentz violations at the Planck scale are almost excluded by

experimental observations; see [48, 49]). On the contrary, there are strong arguments suggesting that, if general relativity is an emergent theory, Lorentz symmetry must be very accurately respected at the Planck scale [11, 50], and the characteristic energy scale of Lorentz symmetry breaking should be several orders of magnitude higher than the Planck scale. We intend to approach the gravitational emergent problem in a future work.

Although the models presented here are Newtonian at high energies, we are far from suggesting that high-energy physics should be Newtonian. What these examples show is that high-energy physics will likely incorporate ingredients rather distinct from those of its low-energy incarnation. The emergent perspective is capable of providing tantalizing explanations of principles of physics without relying on the specifics of the high-energy theory. We thus think that, in our search for a deeper understanding of nature, an emergent point of view is a useful, and probably even necessary, complement to an analysis based on fundamental principles.

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Appendix A. Ginzburg–Landau minimization

The validity of the Ginzburg–Landau theory is restricted to temperatures near the critical transition temperature T_c . However, it is much easier to handle the calculations within this restricted setup, and then generalize them to the whole range of temperature by using the microscopic theory. In either case, the structure of the order parameter (18) is obtained using a minimization principle. In the microscopic theory, the quantity to be minimized is the expectation value of the Hamiltonian (17) in the corresponding Fock vacuum state. In the Ginzburg–Landau theory (for constant temperature T and volume V), the Helmholtz free energy functional of the order parameter, constructed as follows, is the quantity to be minimized. The order parameter (18) is zero above a certain critical temperature T_c but takes a finite value for $T < T_c$. If we suppose that, near T_c , the free energy is analytic in the order parameter and obeys the symmetries of the microscopic Hamiltonian, then one can write a Taylor expansion near the critical temperature. To have a non-trivial minimization problem of this free energy, one only needs to consider the first two non-zero orders. In our case, as the free energy must be invariant under rotations in both coordinate and spin spaces, these terms will be second-order and fourth-order. Given the order parameter (19), there is one possible second-order term and five fourth-order terms:

$$I_0 := \sum_{ai} d_{ai} d_{ai}^*, \quad (\text{A.1})$$

$$I_1 := \sum_{ai} \sum_{bj} d_{ai} d_{ai} d_{bj}^* d_{bj}^*, \quad (\text{A.2})$$

$$I_2 := \sum_{ai} \sum_{bj} d_{ai} d_{bj} d_{ai}^* d_{bj}^*, \quad (\text{A.3})$$

$$I_3 := \sum_{ai} \sum_{bj} d_{ai} d_{aj} d_{bi}^* d_{bj}^*, \quad (\text{A.4})$$

$$I_4 := \sum_{ai} \sum_{bj} d_{ai} d_{bj} d_{aj}^* d_{bi}^*, \quad (\text{A.5})$$

$$I_5 := \sum_{ai} \sum_{bj} d_{ai} d_{bi} d_{aj}^* d_{bj}^*. \quad (\text{A.6})$$

Then the general form of the free energy is

$$F = F_n + \alpha_0 (T - T_c) I_0 + \frac{1}{2} \beta (T_c) \sum_{s=1}^5 \beta_s I_s. \quad (\text{A.7})$$

Terms proportional to gradients of the order parameter are neglected because the variations of the order parameter are considered smooth enough. Here F_n is the free energy of the normal phase, which is independent of the order parameter so it is an irrelevant constant to our purposes. The form of the coefficients we are considering is enforced by the behavior of the order parameter near the critical temperature (see section 5.7 in [28] for a detailed discussion).

As it stands, this minimization problem is not analytically solvable. For this reason, the unitarity condition,

$$\sum_{b,c} \epsilon_{abc} d_{bi}^* d_{cj} = 0, \quad (\text{A.8})$$

is imposed. Although there is no theoretical argument to impose this condition (apart from simplicity of certain expressions), there is some experimental justification because it seems that the states of ^3He that are realized in nature are all unitary in this sense. Consideration of non-unitary states could of course be interesting for other purposes. In our case, although the state we are most interested in (the planar state) does not seem to be realized in nature, it is nevertheless also unitary.

Appendix B. Comments regarding the ABM order parameter

In this appendix we briefly analyze the quasiparticle evolution equations for the ABM state with order parameter (21) and

$$\hat{x} = \hat{m}, \quad \hat{y} = \hat{n}', \quad \hat{z} = \hat{s}, \quad (\text{B.1})$$

along the lines of the analysis performed for the planar state in the main text, to show the differences between both states. The reason for the choice of axes (B.1) is that the alignment of \hat{s} and $\hat{m} \times \hat{n}$ is favored by the action of nuclear dipole interactions [51, 52]. The arguments and conclusions in this section do not depend on this choice.

Because of this choice of axes, it is better to write the evolution equations of quasiparticle operators in the spin basis in \hat{x} direction, defined as

$$a_{p\rightarrow} := \frac{a_{p\uparrow} + a_{p\downarrow}}{\sqrt{2}}, \quad a_{p\leftarrow} := \frac{a_{p\uparrow} - a_{p\downarrow}}{\sqrt{2}}. \quad (\text{B.2})$$

When these equations are linearized around the Fermi point $+p_{\text{F}}\hat{\boldsymbol{l}}$ and represented in position space, one obtains the analog of equation (72) but split for both spin projections. One can verify that one obtains similar equations directly in the basis \uparrow, \downarrow when $\hat{\boldsymbol{s}} = \hat{\boldsymbol{x}}$. If, instead of that, one takes $\hat{\boldsymbol{s}} = \hat{\boldsymbol{y}}$, the equations are almost the same, with only a change of the sign of the two last terms in the last equation. The evolution operators are given by the following expressions:

$$\begin{aligned} \mathcal{H}_{\rightarrow} &:= c_{\parallel}\hat{\boldsymbol{l}} \cdot (-i\nabla - p_{\text{F}}\hat{\boldsymbol{l}}) + c_{\perp}(\sigma_1\hat{\boldsymbol{m}} - \sigma_2\hat{\boldsymbol{n}}) \cdot (-i\nabla), \\ \mathcal{H}_{\leftarrow} &:= c_{\parallel}\hat{\boldsymbol{l}} \cdot (-i\nabla - p_{\text{F}}\hat{\boldsymbol{l}}) - c_{\perp}(\sigma_1\hat{\boldsymbol{m}} + \sigma_2\hat{\boldsymbol{n}}) \cdot (-i\nabla). \end{aligned} \quad (\text{B.3})$$

One can realize that these equations both have the same chirality, by multiplying the ± 1 factors in front of the Pauli matrices. For this reason, it is better to apply a linear transformation to one of the equations, say the second, to change its chirality. Such a transformation is given by

$$\psi_{\leftarrow} \longrightarrow i\sigma_2\psi_{\leftarrow}^*. \quad (\text{B.4})$$

The transformed Hamiltonian is

$$\mathcal{H}_{\leftarrow'} := -\sigma_2\mathcal{H}_{\leftarrow}\sigma_2 = -c_{\parallel}\hat{\boldsymbol{l}} \cdot (-i\nabla + p_{\text{F}}\hat{\boldsymbol{l}}) + c_{\perp}(\sigma_1\hat{\boldsymbol{m}} - \sigma_2\hat{\boldsymbol{n}}) \cdot (-i\nabla). \quad (\text{B.5})$$

In the same way as with the planar phase, let us define the matrices

$$Z^1 = \begin{pmatrix} \sigma_1 & 0 \\ 0 & \sigma_1 \end{pmatrix}, \quad Z^2 = \begin{pmatrix} -\sigma_2 & 0 \\ 0 & -\sigma_2 \end{pmatrix}, \quad Z^3 = \begin{pmatrix} \sigma_3 & 0 \\ 0 & -\sigma_3 \end{pmatrix}. \quad (\text{B.6})$$

The problem of finding a representation of the gamma matrices $\{P, PZ^a\}_{a=1,2,3}$ is similar to the one studied for the case of the planar state. Here, a solution is given by

$$P := \begin{pmatrix} 0 & \sigma_3 \\ \sigma_3 & 0 \end{pmatrix}. \quad (\text{B.7})$$

However, now the two chiralities have a different coupling to the vector potential when the same perturbative analysis of section 3.6 is applied to the evolution equations of this section. Such a coupling implies that we cannot write the low-energy evolution equations as a Dirac field coupled to a vector potential. In fact, the coupling is now axial, with a term in the equation of motion proportional to

$$\gamma^5\gamma^{\mu}\bar{V}_{\mu}\psi. \quad (\text{B.8})$$

Here, we use the symbol \bar{V}_{μ} instead of \bar{A}_{μ} to denote the inhomogeneities of the orbital part of the order parameter. The reason is that, even if these objects are written in the same way in terms of the inhomogeneities, they should have different transformation properties under the usual symmetry transformations such as parity. This can be understood by looking at the structure of the Cooper pairs in both states, ABM and planar. In the first case, the vector $\hat{\boldsymbol{l}}$ shows the direction of the angular momentum of the Cooper pairs with positive as well as negative projection of spin, and thus, $\hat{\boldsymbol{l}}$ is an axial vector in this state. On the other hand, in the planar state, the two spin populations form Cooper pairs with opposite angular momentum,

implying that the planar state is not axial (for a more detailed discussion, see section 7.4 of [10]). This is consistent with the kind of coupling to a vector field that appears in each state.

This does not contradict the claim that the low-energy quasiparticle excitations of the Fermi liquid are determined by topology in momentum space. Both ABM and planar states are characterized by two Fermi points. In a homogeneous system, the low-energy fermionic excitations are the same in both states, and can be represented by a (free) Dirac field. However, the structure of the order parameter is different in both states, and so is the coupling of the fermionic excitations to the inhomogeneities of this order parameter. In other words, the only difference between these two states is their chirality. This is why the planar state serves better as an analog of the vacuum state of electrodynamics.

ABM and planar states are limiting cases of the family of axiplanar states [10]. In these states, the two spin populations are decoupled as far as the order parameter is concerned. For these two limiting states, one is considering perturbations with $\delta\hat{l}_\uparrow = \delta\hat{l}_\downarrow$ and $\delta\hat{l}_\uparrow = -\delta\hat{l}_\downarrow$, respectively. General axiplanar states can be analyzed with the same techniques to show that, in general, one has couplings to a polar as well as an axial vector, both constructed by different linear combinations of the independent variations $\delta\hat{l}_\uparrow$ and $\delta\hat{l}_\downarrow$, i.e., $(\delta\hat{l}_\uparrow \pm \delta\hat{l}_\downarrow)/\sqrt{2}$.

Appendix C. Inhomogeneous Ginzburg–Landau free energy and Zel’dovich picture of emergence

In our discussion, we have not followed the traditional approach in which the issue of the dynamics of the order parameter is approached (see [10]). Let us briefly describe it, as well as its shortcomings.

From a condensed-matter perspective, it is natural to expect that the dynamics of the order parameter can be determined by a generalized GinzburgLandau approach, analog to the one sketched in section 3.3. However, this time we retain terms containing derivatives of the order parameter. In a similar way, the temperature-dependent coefficients accompanying each term of the free action can be calculated from the microscopic theory. The result of this calculation would be an expression quadratic in the derivatives of the textures. If, for the moment, we restrict the discussion to orbital textures of the unit vector, $\delta\hat{l}$, the corresponding inhomogeneous part of the free action was worked out in [53]. At finite temperature T it is given by the following expression:

$$\frac{P_F c_{\parallel}}{12\pi^2 \hbar} \left[\log \left(\frac{\Delta_0}{k_B T} \right) \left[\hat{l} \times (\nabla \times \hat{l}) \right]^2 + \left[\hat{l} \cdot (\nabla \times \hat{l}) \right]^2 + (\nabla \cdot \hat{l})^2 \right]. \quad (\text{C.1})$$

In this expression, we keep the dominant terms in the zero-temperature limit $T \rightarrow 0$, as well as the first order in an expansion in the parameter c_{\perp}/c_{\parallel} . This expansion is usually carried out in the literature because of the smallness of this parameter in the experimental case of ^3He . The reason for the infrared divergence in the first term can be traced back to the existence of Fermi points in the fermionic spectrum. In laboratory realizations, the infrared divergence is always regulated by the temperature of the system. However, the other terms have coefficients that are constant in the limit $T \rightarrow 0$. Therefore one can always, in principle, lower the temperature sufficiently to make the first term in equation (C.1) dominant. For completeness, let us mention an assumption found in the literature concerning the existence of the following additional term in this expansion:

$$\frac{p_F c_{\parallel}}{12\pi^2 \hbar} \left(\frac{c_{\perp}}{c_{\parallel}} \right)^2 \log \left(\frac{\Delta_0}{k_B T} \right) \left[\hat{\mathbf{l}} \cdot (\nabla \times \hat{\mathbf{l}}) \right]^2. \quad (\text{C.2})$$

It has been claimed in [10] that this term is usually neglected because it is of quadratic order in the parameter c_{\perp}/c_{\parallel} , but that it appears in an explicit evaluation at this order [54]. From our point of view, however, there is no conclusive argument in this respect, as the definite relation between the evaluation of the Ginzburg–Landau energy and the technical procedure used in [54] (which indeed seems to be closer to Zel’dovich’s approach, as we discuss in this section) is not clear for us. Notice that the terms presented here, equation (C.1) plus equation (C.2), correspond to the potential energy of a theory with the usual kinetic term for the restricted kind of textures considered, $(\partial_t \delta \hat{\mathbf{l}})^2$. In the following, we will verify whether all these terms can be obtained in a simpler way from the perspective of the emergent relativistic theory we described in the main text.

Within the emergent relativistic field theory framework, a possible way an internal observer can determine the dynamics of the gauge fields is by integrating out fluctuations of the relativistic fermionic fields. This is nothing but the suggestion of Sakharov concerning gravity [43] (see [55] for a modern review), adapted to electrodynamics by Zel’dovich [44]. The integration over fermionic fluctuations, which technically amounts to an evaluation of a fermionic path integral in the presence of background fields, can be found in the literature carried out in different ways; see [54, 56]. A note of caution: these two approaches (Ginzburg–Landau and Sakharov–Zel’dovich) are, in principle, very distinct in nature. One is a finite-temperature analysis (implying we have a thermal distribution of fermionic quasiparticles), whereas the second is a zero-temperature calculation. We proceed with the comparison, but take this into account.

The only calculation we need is the evaluation of the one-loop polarization tensor, which characterizes the only divergent term in the fermionic path integral. This is a standard calculation that involves an ultraviolet regularization of the momentum integral with an upper cutoff Λ_+ as well as an infrared regularization by means of a similar quantity Λ_- . At this stage, the significance of these quantities is merely formal, although they gain a physical interpretation later. In the case in which Lorentz and gauge symmetries are preserved using the regularization method, the divergence is logarithmic in the limit $\Lambda_+/\Lambda_- \rightarrow \infty$ and corresponds to a term in the action

$$-\frac{\nu^2}{48\pi^2 \hbar} \log \left(\frac{\Lambda_+}{\Lambda_-} \right)^2 \int d^4x F^{\mu\nu} F_{\mu\nu}. \quad (\text{C.3})$$

In standard quantum field theory, this term would be absorbed by a suitable counterterm before taking the limit $\Lambda_+/\Lambda_- \rightarrow \infty$, leading to charge and photon field renormalization. However, in this construction, we have no definite tree-level kinetic term for the vector potential. Moreover, the ultraviolet divergence here is clearly an artefact of the extrapolation of the low-energy theory to higher energies. Following Zel’dovich, we can interpret this (finite) term as the leading kinetic term for the electromagnetic field induced at one-loop level:

$$-\frac{1}{4\mu_0} \int d^4x F^{\mu\nu} F_{\mu\nu}. \quad (\text{C.4})$$

In principle, the logarithmic behavior supports this interpretation because it permits this term to be dominant over any unknown, tree-level dynamical terms for the gauge field, at least for certain values of the quotient Λ_+/Λ_- . This would normally be understood as one-loop dominance [43, 55]. In this regime, it is reasonable to think that the dynamics of the order parameter should be well described by equation (C.4). In the resulting effective theory, as expected, the actual value of ν is not relevant because it can be changed by a redefinition of the vector potential; the only meaningful quantity is the combination

$$\alpha := \frac{\nu^2 \mu_0}{4\pi\hbar} = \frac{3\pi}{\log(\Lambda_+/\Lambda_-)}. \quad (\text{C.5})$$

Notice that the value of this dimensionless coupling constant, the effective fine-structure constant, is universal: it only depends on the value of the ultraviolet and infrared cutoffs.

The next step is the comparison of the term (C.3) with the corresponding limit of the Ginzburg–Landau approach, which contains in principle all the information about the evolution of textures. We complete it for restricted textures in which the superfluid velocity plays no role. At low temperatures, two terms in equations (C.1) and (C.2) are the dominant ones. At first order in the texture, for which

$$-\frac{\nu}{p_F} \mathbf{A} = \delta \hat{\mathbf{l}} = \delta \hat{\mathbf{m}} \times \hat{\mathbf{n}}_0 + \hat{\mathbf{m}}_0 \times \delta \hat{\mathbf{n}}, \quad (\text{C.6})$$

one can see that the sum of these terms is equivalent to the spatial part of the relativistic term, which can be written as

$$\begin{aligned} \frac{1}{2} F^{\mu\nu} F_{\mu\nu} &= c_{\perp}^2 (\partial_0 A_1)^2 + c_{\perp}^2 (\partial_0 A_2)^2 - c_{\perp}^4 (\partial_1 A_2 - \partial_2 A_1)^2 \\ &\quad - c_{\perp}^2 c_{\parallel}^2 (\partial_3 A_1)^2 - c_{\perp}^2 c_{\parallel}^2 (\partial_3 A_2)^2. \end{aligned} \quad (\text{C.7})$$

In this way, one can, in principle, accept that the picture of induction of dynamics captures the relevant dynamics of the system when the temperature is low enough. As long as Lorentz invariance is kept intact in the present scheme, one can argue in favor of the occurrence of the term (C.2) in the inhomogeneous Ginzburg–Landau free energy. Then, the matching of the low-energy relativistic theory with the Ginzburg–Landau approach tells us the value of the quotient of regulators:

$$\frac{\Lambda_+}{\Lambda_-} = \frac{\Delta_0}{k_B T}. \quad (\text{C.8})$$

This fixes the value of the fine-structure constant (C.5) in terms of parameters of the system and provides an interpretation of the ultraviolet and infrared regulators. They would be given by

$$\Lambda_+ \simeq \Delta_0, \quad \Lambda_- \simeq k_B T. \quad (\text{C.9})$$

The value of the infrared regulator simply implies that the energy scale associated with the temperature physically removes the infrared divergence. On the other hand, the value of the ultraviolet regulator is telling us we are performing the integration over fermions up to energies given by Δ_0 . The trouble with this observation is that this energy is much greater than the scale of violation of Lorentz symmetry $E_{\perp} = m_* c_{\perp}^2$ [recall the discussion around equation (31)]:

$$\frac{\Delta_0}{m^*c_{\perp}^2} = \frac{c_{\parallel}}{c_{\perp}} \gg 1. \quad (\text{C.10})$$

Thus, we have no definite argument that supports that the dynamics should be given by the standard relativistic term (C.4), and the whole argument falls apart. The only way to remedy this would be to evaluate the inhomogeneous Ginzburg–Landau free energy to verify the occurrence of equation (C.2). It is interesting to note there exists a different scale of violation of Lorentz symmetry for quasiparticles traveling along the anisotropy axis (only $p_{\parallel} \neq 0$ in equation 30) with higher characteristic energy, which could help to obtain this result. However, even if one succeeds, the picture would have additional negative features, as we argue next.

In the Ginzburg–Landau approach, there are additional terms that do not respect the low-energy emergent symmetries, relativistic, and gauge invariance. On the one hand, higher orders in the logarithmic term (109) are suppressed by inverse powers of the Fermi momentum, which is several orders of magnitude greater than the maximum momentum which can be carried by the electromagnetic field (recall equation (58)). Note that gauge invariance is a property only of the linearized version of (C.1), (C.2). On the other hand, terms that are quadratic in the texture but non-relativistic (the relevant symmetry group is the Galileo group) are suppressed in a weaker way by the logarithm in equations (C.1) and (C.2) or, equivalently, equation (C.3) at low temperatures. These terms come from the gradient expansion of the Goldstone variables so they are ultimately linked to the breaking of symmetries in the condensed phase. The first unsatisfactory feature of this argument is that, because this suppression is logarithmic, one must consider practically zero temperatures to ensure that the logarithm is large enough. But there is even a stronger argument against this picture: the value of the fine-structure constant (C.5) shows that the suppression of these terms is proportional to α . Thus in this effective theory the usual perturbative expansion in terms of the fine-structure constant makes no sense.

These arguments make it difficult to support a comparison between the Ginzburg–Landau free energy and the low-energy action for the emergent gauge fields. The mechanism which permits to obtain a dynamical implementation of gauge invariance cannot be a logarithmic suppression of the terms which violate this symmetry in the action (it is also difficult to reconcile this logarithmic suppression with the accuracy of known symmetries [57]). Our arguments are sufficiently general to apply to the original discussion of Zel'dovich [44], as it will be discussed elsewhere.

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