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1 **Deep carbon through time:**
 2 **Earth's diamond record and its implications for**
 3 **carbon cycling and fluid speciation in the mantle**

4
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25
 26
 27 **Abstract**

28
 29 Diamonds are unrivalled in their ability to record the mantle carbon cycle and mantle
 30 fO_2 over a vast portion of Earth's history. Diamonds' inertness and antiquity means
 31 their carbon isotopic characteristics directly reflect their growth environment within
 32 the mantle as far back as ~3.5 Ga. This paper reports the results of a thorough
 33 secondary ion mass spectrometry (SIMS) carbon isotope and nitrogen concentration
 34 study, carried out on fragments of 144 diamond samples from various locations, from
 35 ~3.5 to 1.4 Ga for P [peridotitic]-type diamonds and 3.0 to 1.0 Ga for E [eclogitic]-
 36 type diamonds. The majority of the studied samples were from diamonds used to
 37 establish formation ages and thus provide a direct connection between the carbon
 38 isotope values, nitrogen contents and the formation ages. In total, 908 carbon isotope
 39 and nitrogen concentration measurements were obtained. The total $\delta^{13}C$ data range
 40 from -17.1 to -1.9 ‰ (P = -8.4 to -1.9 ‰; E = -17.1 to -2.1 ‰) and N contents range
 41 from 0 to 3073 at. ppm (P = 0 to 3073 at. ppm; E = 1 to 2661 at. ppm). In general,
 42 there is no systematic variation with time in the mantle carbon isotope record since >3
 43 Ga. The mode in $\delta^{13}C$ of peridotitic diamonds has been at -5 (± 2) ‰ since the earliest
 44 diamond growth ~3.5 Ga, and this mode is also observed in the eclogitic diamond
 45 record since ~3 Ga. The skewness of eclogitic diamonds' $\delta^{13}C$ distributions to more

46 negative values, which the data establishes began around 3 Ga, is also consistent
47 through time, with no global trends apparent.

48 No isotopic and concentration trends were recorded within individual samples,
49 indicating that, firstly, closed system fractionation trends are rare. This implies that
50 diamonds typically grow in systems with high excess of carbon in the fluid (i.e.
51 relative to the mass of the growing diamond). Any minerals included into diamond
52 during the growth process are more likely to be isotopically reset at the time of
53 diamond formation, meaning inclusion ages would be representative of the diamond
54 growth event irrespective of whether they are syngenetic or protogenetic. Secondly,
55 the lack of significant variation seen in the peridotitic diamonds studied is in keeping
56 with modeling of Rayleigh isotopic fractionation in multicomponent systems
57 (RIFMS) during isochemical diamond precipitation in harzburgitic mantle. The
58 RIFMS model not only showed that in water-maximum fluids at constant depths
59 along a geotherm, fractionation can only account for variations of <1 ‰, but also that
60 the principal $\delta^{13}\text{C}$ mode of -5 ± 1 ‰ in the global harzburgitic diamond record occurs
61 if the variation in $f\text{O}_2$ is only 0.4 log units. Due to the wide age distribution of P-type
62 diamonds, this leads to the conclusion that the speciation and oxygen fugacity of
63 diamond forming fluids has been relatively consistent. The deep mantle has therefore
64 generated fluids with near constant carbon speciation for 3.5 Ga.

65

66

67 **Keywords**

68

69 Mantle carbon; isotope fractionation; mantle oxygen fugacity; diamond growth; deep
70 carbon cycle; subduction.

71

72

73 **Introduction**

74

75 Carbon and other elements that form volatile species in Earth's mantle, such as
76 nitrogen and hydrogen, are major constituents of our planet and, along with oxygen,
77 dominate the composition of the atmosphere and hydrosphere. The cycling of these
78 elements from the surface into the deep Earth via subduction, and then back to the
79 surface via volcanism has been integral to regulating climate, ocean volume and the
80 habitability of our planet. While many of these volatile cycles in the atmosphere,
81 biosphere and crust are understood, the deeper aspects of these cycles in the mantle

82 and core are rather poorly constrained (Hazen & Schiffrics, 2013). Of interest is the
83 comparison between constancy of the surficial carbon isotope record (e.g. Des Marais,
84 2001) and that of the mantle through time because the mantle is - by more than an
85 order of magnitude - a larger reservoir of C than the crust (e.g. Kelemen & Manning,
86 2015). By studying the fluxes between Earth's reservoirs we gain valuable insights
87 into the geological processes of subduction, partial melting, degassing and
88 metasomatism, and by tracing these processes back through time we gain further
89 insight into the evolution of our planet.

90

91 The upper mantle's redox state is intimately linked to the evolution of Earth's
92 atmosphere through volcanic degassing (Kasting et al., 1993). Potential variations in
93 the mantle's oxidation state through time therefore have significant consequences, yet
94 this has been a controversial topic for many decades (McCammon, 2005; Frost &
95 McCammon, 2008; Foley, 2011). Most of the controversy lies in the fact that oxygen
96 fugacity information obtained from the study of volcanic rocks does not agree with
97 that obtained from the study of mantle peridotites (see Figure 1 of Foley (2011) and
98 references therein). Determining mantle fO_2 from $Fe^{3+}/\Sigma Fe$ analyses on ancient
99 peridotites and mid-ocean ridge basalts (MORB) is problematic due to the
100 independent or combined effects of alteration and metasomatism. As a result, trace
101 element systematics, such as V, Cr, Sc and Ce, have been employed to estimate
102 oxygen fugacity (fO_2 ; Canil, 1997; Li and Lee, 2004; Trail et al., 2011). Many studies
103 of upper mantle-derived rocks have concluded that mantle fO_2 has remained constant,
104 within ~1 log unit of the fayalite-magnetite-quartz (FMQ) oxygen buffer at 1 GPa,
105 since 3.5 Ga (Canil, 1997; Canil and Fedortchouk, 2001; Delano, 2001; Li and Lee,
106 2004; Berry et al., 2008; Trail et al., 2011; Hibbert et al., 2012; Rollinson et al., 2017;
107 Nicklas et al., 2018). However, the comprehensive review by Foley (2011) along with
108 some recent studies (Aulbach & Viljoen, 2015; Aulbach & Stagno, 2016; Shu et al.,
109 2016; Aulbach et al., 2017; Nicklas et al. 2019) have suggested that the Archean
110 mantle was more reducing than that of the present day. These contrasting results
111 highlight the uncertainty in our knowledge of fO_2 through time.

112

113 Along with potential temporal variation in the mantle's oxygen fugacity, there are
114 also vertical variations to consider. The effects of pressure cause the mantle to
115 become more reducing with depth (Ballhaus & Frost, 1994). The implications of this

116 variation for carbon are that at more oxidised (higher fO_2) conditions in the shallow
117 mantle, carbonate and CO_2 would be the stable / dominant species, changing to CH_4
118 and metallic carbide with decreasing fO_2 (and depth). Therefore variation of mantle
119 fO_2 through time has implications for the speciation of carbon-related volatiles in the
120 upper mantle.

121

122 Diamonds are unrivalled in their ability to record the mantle carbon cycle over a vast
123 portion of Earth's history. They are the product of ascending, cooling, carbon-
124 saturated metasomatic fluids-melts and/or redox reactions, predominantly within
125 peridotitic and eclogitic domains in the mantle lithosphere (see Stachel & Luth (2015)
126 and references therein). Their occurrence with inclusions of CO_2 (Schrauder &
127 Navon, 1993), carbonate (e.g. Wang et al., 1996), methane (Smit et al., 2016) and
128 carbide (e.g. Mikhail et al., 2014; Smith et al., 2016), show their formation over a
129 significant range of oxygen fugacities. However, the lack of buffering capacity of
130 peridotitic subcratonic mantle (Luth & Stachel, 2014) suggests that redox reactions
131 may not be responsible for diamond growth in this part of the mantle. Luth & Stachel
132 (2014) showed that the speciation of carbon in mantle fluids is actually the controlling
133 variable when it comes to the oxygen fugacity of the peridotitic lithospheric mantle,
134 which is counter to the traditional view of the fluids adjusting to the prevalent fO_2
135 rather than buffering them (e.g. Woodland et al. 2006). Understanding the role of
136 carbon-bearing melts in eclogitic domains and their relationship to fO_2 is inherently
137 more complex and the subject of ongoing study.

138

139 A key characteristic that makes diamond such a powerful tool for studying the mantle
140 is that once grown, they do not chemically re-equilibrate with their surroundings. This
141 means their carbon isotopic composition, as well as their potential load of nitrogen,
142 hydrogen, boron and nickel impurities, directly reflect their growth environment.
143 Given that diamond growth events have been dated as far back as ~ 3.5 Ga (Figure 1:
144 see Supplementary Material Table S1 and references therein), no other mineral can
145 provide such detailed insight into mantle fluid processes and the deep carbon cycle
146 throughout Earth's evolution.

147

148 The historical database of diamond carbon isotope compositions (e.g., Cartigny et al.,
149 2014) is based on the bulk analysis of diamond fragments. Bulk C-isotopic analyses,

150 combined with FTIR nitrogen analyses, have been successfully employed in some
151 instances to track Rayleigh fractionation processes during diamond formation
152 (Cartigny et al., 2001; Thomassot et al., 2007; Stachel et al., 2009). The homogeneity
153 of carbon isotope values ($\pm 2\%$) in the majority of diamonds studied has been shown
154 by comparing data from multiple fragments of the same samples (see figure 5 of
155 Cartigny et al., 2004). Stepped combustion analysis of individual samples also
156 revealed limited heterogeneity in carbon isotopes values (Mikhail et al., 2014) but
157 lacked spatial resolution. The development of secondary ion mass spectrometry
158 (SIMS) methods applicable to measuring carbon isotopes of diamonds *in situ* from
159 very small analytical volumes (Harte & Otter, 1992; Fitzsimmons et al., 1999; Hauri
160 et al., 2002), showed clearly that a few gem diamonds had very heterogeneous carbon
161 isotopic values depending on their growth history (Harte et al., 1999; Bulanova et al.,
162 2002; Zedgenizov & Harte, 2004). Considering the observed diamond internal
163 heterogeneity on the micron scale (e.g. Fitzsimmons et al., 1999), the most accurate
164 approach to determine the possible fluid speciation in diamond-forming fluids is to
165 measure spatially resolved growth-directional trends (i.e. core to rim) in the carbon
166 isotopic and nitrogen concentration systematics in individual diamonds (e.g. Smart et
167 al., 2011; Wiggers de Vries et al., 2013; Smit et al., 2016; 2019a).

168

169 This paper reports the results of a major SIMS carbon isotope study, carried out on
170 144 diamond samples, spanning a large range of geological time (~3.5 to 1.4 Ga for
171 P-type diamonds, 3.0 to 1.0 Ga for E-type diamonds). The sample collection studied
172 provides a direct connection between the ages of diamond formation, obtained from
173 inclusion studies, and the carbon isotope values of the host diamonds. The spatial
174 resolution of the analyses, together with the readily identifiable core to rim growth
175 directions evident in most samples, allows trends in the variability of C isotopic
176 values and nitrogen concentrations to be examined, with the goal to search for
177 possible systematic shifts in the speciation of carbon-bearing mantle fluids through
178 time.

179

180

181 **Samples**

182

183 Fragments of 144 diamond crystals, from nine localities in Canada, Australia,
184 Southern Africa and Russia, have been analysed for their carbon isotopic values and
185 nitrogen contents. Of those, 88 are defined as having a peridotitic association (P-type)
186 based upon their inclusion paragenesis, and 56 formed in eclogitic substrates (E-type).
187 The details of the sample source locations are provided in **Table 1**, along with
188 references to the original studies that dated the inclusions. With the exception of 18
189 samples from Diavik (Slave Craton, Canada; Donnelly et al., 2007), 11 from Wawa
190 (Superior Craton, Canada; Stachel et al., 2006), 13 from Ellendale (Kimberley Craton,
191 Australia; Jaques et al., 1989; 1994), and 10 from Argyle (Kimberley Craton,
192 Australia; Stachel et al., 2018), the remaining majority of the diamonds studied here
193 are fragments, recovered after breakage to obtain inclusions for dating studies.
194 Therefore, they can be directly linked to the age of the diamond growth event
195 documented in the original dating study.

196

197 For peridotitic diamonds from De Beers Pool, Udachnaya and Venetia, diamond
198 formation ages come from radiogenic isotope analysis of pooled garnet inclusions
199 (Richardson et al., 1984; Richardson & Harris, 1997; Richardson et al., 2009). For
200 eclogitic diamonds from De Beers Pool (Bultfontein, De Beers, Dutoitspan, and
201 Wesselton kimberlites), Jwaneng and Orapa, ages were obtained by Re-Os analysis of
202 individual sulphide inclusions (Richardson et al., 2001; 2004; Shirey et al., 2009). In
203 the case of the Diavik, Argyle and Ellendale diamonds, which are not directly derived
204 from age dating studies on these diamond populations (Aulbach et al., 2009;
205 Richardson, 1986; Smit et al., 2010), the inclusion compositions and N aggregation
206 characteristics are similar to the original suites used for geochronology and thus the
207 published ages are applied. The diamonds from Wawa have not been dated directly
208 but are constrained by the age of the rocks they are found in (~2.7 Ga; Stachel et al.,
209 2006) as being Neoproterozoic, which obviously represents a minimum age. All of the
210 diamonds (with the possible exception of two Wawa diamonds with majoritic garnets)
211 are lithospheric (140 – 250 km) in origin.

212

213 Whilst Thomassot et al (2009) argue that at Jwaneng silicate- and sulphide-bearing
214 diamonds form in separate events, the presence of silicate and sulphide inclusions (in
215 part even touching; Richardson et al., 2004) in the same P and E-type diamonds at
216 Jwaneng strongly suggests coeval formation. Even if silicate- and sulphide-bearing

217 diamonds formed during different growth events, this would have very limited
218 implications for this study because all of the peridotitic diamond samples that are
219 directly tied to the dating studies contained only silicate inclusions, while all directly
220 dated eclogitic samples are sulphide-bearing. For the three P-type sample collections
221 with indirect ages (Diavik, Ellendale and Wawa), only 4 of the diamonds contained
222 sulphides (1 from Diavik, 3 from Ellendale). For the E-type sample collections with
223 indirect ages (Diavik and Argyle), all 10 Argyle samples are silicate-bearing (in
224 keeping with the dating study which used silicates) and 4 of the 9 Diavik samples are
225 silicate-bearing (not in keeping with the sulphide derived age). So with the exception
226 of 8 samples, the remainder are all in keeping with the inclusion type used to obtain
227 the age. The peridotitic samples represent silicate-bearing diamond formation events,
228 while excluding Argyle, the remaining eclogitic samples represent sulphide-bearing
229 diamond formation events.

230

231 Given that all of the diamonds studied here are fragments, they do not record the
232 crystals' entire growth history. During selection and mounting, every effort was made
233 to discern an outer crystal surface, so that the subsequent analysis would have a
234 context of growth direction and data transects would represent the evolution of the
235 crystal's growth. Data from 96 samples have confirmed growth direction context,
236 supported by cathodoluminescence (CL) imaging of their growth structure.

237

238 Fibrous and polycrystalline diamonds are assumed to form during unusually rapid
239 crystal growth driven by strong supersaturation in carbon. Fibrous diamonds are most
240 likely of young (Phanerozoic) geological age (Sunagawa, 1984; Gurney et al., 2010),
241 while no polycrystalline diamonds have ever been dated. Such young and undated
242 diamond types cannot be used to constrain variations in the carbon isotopic
243 composition of diamonds through time, and they are therefore not included in the
244 present study, nor are the literature data derived from those types of samples.
245 Similarly, microdiamonds (generally defined as <0.5 mm) are also not included as age
246 data are not available.

247

248 **Methods**

249

250 Mount preparation, CL imaging and secondary ion mass spectrometry (SIMS) were
251 carried out at the Canadian Centre for Isotopic Microanalysis (CCIM), University of
252 Alberta. Four SIMS mounts were created and analyzed in this study (M1369, M1406,
253 M1410, M1520). The rough diamond fragments were cast together in groups of 20 to
254 50 in 25 mm epoxy discs, and ground and polished using electroplated diamond pads
255 on rotary equipment to expose flat sections through grain interiors. These epoxy discs
256 containing polished diamonds were trimmed to mm-scale blocks and pressed into
257 SIMS indium mounts (M1369, M1410, M1520) along with a piece of CCIM diamond
258 reference material (RM) S0270 and vitreous carbon S0233A. For epoxy mount
259 M1406, reference materials were pressed into an integrated indium slot within an
260 epoxy mount. In all cases, the RMs and unknowns were spaced closely together on
261 each mount. The SIMS mounts were coated with ~20 nm of Au prior to scanning
262 electron microscopy (SEM) using a Zeiss EVO MA15 instrument, operating at 15 kV
263 and ~3 nA beam current. CL SEM images were obtained using a parabolic mirror
264 coupled to a high-sensitivity, broadband photomultiplier detector. The mounts were
265 subsequently coated with additional Au prior to SIMS analysis.

266

267 Isotopes of carbon ($^{13}\text{C}/^{12}\text{C}$) were determined using the IMS-1280 multi-collector ion
268 microprobe at CCIM. Analytical methods and reference materials are detailed in Stern
269 et al. (2014). Primary beam conditions included the use of 20 keV $^{133}\text{Cs}^+$ ions focused
270 to a diameter ~ 15 μm and beam currents of ~ 2.0 nA. The primary beam was rastered
271 across a 20 x 20 μm area for 30 s prior to analysis, to clean the surface of Au and
272 contaminants, and implant Cs. The normal incidence electron gun was not utilized, as
273 most diamonds are conducting under ion bombardment. Negative secondary ions
274 were extracted through 10 kV to the grounded secondary column (Transfer section).
275 Automated tuning of the secondary ions in the Transfer section preceded each
276 analysis. Secondary ion collection conditions for C-isotopes included an entrance slit
277 width of 100 μm , field aperture of 5 x 5 mm, a field aperture-to-sample magnification
278 of 100 x, and a fully-open energy slit. Both $^{12}\text{C}^-$ and $^{13}\text{C}^-$ were analyzed
279 simultaneously in Faraday cups (L'2 and FCs using $10^{10} \Omega$ and $10^{11} \Omega$ amplifier
280 circuits) at mass resolutions of 2000 and 2900, respectively. Faraday cup baselines
281 were determined once at the start of the session. Mean count rates for $^{12}\text{C}^-$ and $^{13}\text{C}^-$
282 were typically $1.1\text{--}1.6 \times 10^9$ and $1.2\text{--}1.8 \times 10^7$ counts/s, respectively, determined over
283 a 75 s counting interval, with total analysis time of 210 s. The analytical sequence

284 interspersed measurements of unknowns with the natural diamond S0270, having
285 $\delta^{13}\text{C}_{\text{VPDB}} = -8.88 \pm 0.10 \text{ ‰}$ in a 4:1 ratio. Instrumental mass fractionation (IMF) for
286 $^{13}\text{C}/^{12}\text{C}$ is typically about -24 ‰, determined precisely for each session from
287 utilizing all the replicate analyses of S0270 diamond time-corrected back to the start
288 of the sub-session. The standard deviation of the $^{13}\text{C}/^{12}\text{C}$ values per session was
289 $\pm 0.05 - 0.07 \text{ ‰}$ after correction for systematic within-session IMF drift of $\leq 0.5 \text{ ‰}$
290 over several hours. Uncertainties of individual $\delta^{13}\text{C}_{\text{VPDB}}$ analyses propagate within-
291 spot ($\sim \pm 0.04 \text{ ‰}$, 1σ), between-spot ($\pm 0.05 \text{ ‰}$, 1σ), and between-session errors
292 ($\pm 0.01 \text{ ‰}$, 1σ), and average $\pm 0.14 \text{ ‰}$ (2σ).

293

294 Nitrogen abundances were determined for some diamonds immediately following C-
295 isotope analysis from the identical spot locations. The primary beam was rastered for
296 30 –60 s prior to analysis to clean the adjacent area. Secondary ion collection
297 conditions included an entrance slit width of 45 μm , field aperture of 3 x 3 mm, and
298 energy slit width of 40 eV transmitting low-energy ions. The molecular ion ratios of
299 $[^{12}\text{C}^{14}\text{N}^-]/[^{12}\text{C}^{12}\text{C}^-]$ or $[^{13}\text{C}^{14}\text{N}^-]/[^{12}\text{C}^{13}\text{C}^-]$ were analyzed simultaneously using either a
300 Faraday cup–Faraday cup (L'2–FC2) or Faraday cup–electron multiplier combination
301 (L'2–EM). Mass resolution was $\sim 6300 - 7000$ for CN^- molecular ions, 2100 for
302 $^{12}\text{C}^{12}\text{C}^-$, and 6000 for $^{12}\text{C}^{13}\text{C}^-$. Electron multiplier counts were corrected for
303 background and deadtime (40 ns). Total analysis time was 200 – 240 s, including pre-
304 analysis raster, secondary ion centering, and peak counting time of 25 – 50 s. The
305 sensitivity factor for N in diamond was determined by analysis of S0280 diamond
306 having nitrogen concentration ($[\text{N}]$) = 1670 at. ppm located on a separate mount
307 (Stern et al., 2014), or analysis of S0270 diamond with ~ 2150 at. ppm located on the
308 same mount. Uncertainties in concentration are $\pm 5 \text{ ‰}$ absolute (95% conf.)

309

310

311 **Results**

312

313 In total, 908 combined carbon isotope and nitrogen concentration measurements have
314 been obtained for peridotitic ($n = 504$) and eclogitic ($n = 404$) inclusion-bearing
315 diamonds. The entire dataset is provided in the **Supplementary Material**, grouped
316 according to inclusion paragenesis (**Table S2**). Histograms of the $\delta^{13}\text{C}$ data (**Figure 2**)
317 are given per location and age, while the statistical characteristics are provided in

318 **Table 2.** The $\delta^{13}\text{C}$ data range from -17.1 to -1.9 ‰ (P -8.4 to -1.9 ‰; E -17.1 to -2.1
319 ‰) and N contents range from 0.4 to 3073 at. ppm (P 0.4 to 3073 at. ppm; E 1 to 2661
320 at. ppm).

321

322 As a different number of analyses were recorded per sample, there is potential for a
323 sample bias to become incorporated in the population data through overrepresentation
324 of one sample's characteristics relative to another sample. For example, in the
325 literature the presence of highly negative (<-15 ‰) portions within a diamond are
326 commonly studied by SIMS in much greater detail than the often more dominant
327 mantle like (~-5 ‰) volumes as they are of greater potential interest in defining the
328 source of the carbon, resulting in a strongly bimodal distribution in some datasets,
329 which is proportionally misleading. To check for this within this study, the average of
330 each sample's C and N characteristics were calculated as well as their median values,
331 and then averaged to compare against the whole population per location. It is clear
332 from the values shown in **Table 2** that overrepresentation is not a significant issue for
333 the $\delta^{13}\text{C}$ and N concentration data reported here.

334

335 From the CL images it is possible to tell that the vast majority of diamonds with
336 sufficient nitrogen content to luminesce are dominated by octahedral growth. Some
337 samples from Diavik ($n = 3$), Venetia ($n = 1$), De Beers Pool ($n = 1$), Jwaneng ($n = 4$)
338 and Orapa ($n = 1$) also reveal some portions of cuboid growth, identified by having a
339 CL response consistent with previous studies of cuboid growth (e.g. Welbourn et al.,
340 1989; Howell et al., 2012). However, as there is no evidence of carbon isotopic
341 fractionation between cuboid and octahedral sectors in mixed-habit diamonds (Howell
342 et al., 2013) this observation does not affect the present data. [N.B. No fibrous
343 diamond material was analysed within this study.] The CL images also reveal that the
344 sample collection is quite varied in terms of the growth history of individual
345 diamonds (**Figure 3**). Nearly half of the diamonds studied here (71 of 144) clearly
346 exhibit either pulsed growth (repetitive subtle variations in CL response; not clear if
347 single or multiple growth events) or clear discrete multiple growth events (noted by a
348 marked difference in CL response as well as resorption features). Complex histories
349 involving multiple growth stages are not necessarily accounted for in the defined ages
350 of the samples, given the rarity of multiple inclusion-bearing diamonds with
351 inclusions in more than one growth zone and that many of the P-type diamond

352 formation ages were determined from pooled silicate inclusions from many diamonds,
353 as opposed to techniques of dating individual sulphide or silicate inclusions (e.g.
354 Pearson et al., 1998; Timmerman et al., 2017).

355

356 Four of the samples exhibiting multiple growth events (all E-type; ddmi-37, JWR7,
357 JWR13, JWR22) show a jump in the $\delta^{13}\text{C}$ values between two layers of >3 ‰; the
358 lowest isotopic values are found in the core of the stone, with subsequent overgrowth
359 having less negative $\delta^{13}\text{C}$ values. If the $\delta^{13}\text{C}$ data for these four samples are removed
360 from the dataset presented in [Table 2](#), then the average Δ values (where Δ represents
361 the variation within individual samples) for E-type diamonds from Diavik (1.9 Ga)
362 and Jwaneng (3 Ga and 2.1 Ga) change to 0.4 ‰ (from 1.0 ‰), 1.5 ‰ (from 3.7 ‰)
363 and 1.4 ‰ (from 2.5 ‰), respectively. This reduces the average Δ value of the E-type
364 population to 1.2 ‰ from 1.6 ‰ (calculated from the average Δ values of the nine
365 diamond growth events from the five localities, as opposed to the average Δ values of
366 all 52 individual samples). Given that the average Δ value of the P-type population is
367 0.8 ‰, this implies that intra-sample variation is typically slightly higher in E-type
368 diamonds than in the P-type diamonds studied here.

369

370 To investigate whether the presence of multiple growth events in half of the diamonds
371 affects the dataset as a whole, the data from samples revealing only a single growth
372 event ([Table 3](#)) were plotted as histograms (for the P-type and E-type populations)
373 and compared against the total dataset ([Figure 4](#)). The single-phase growth peridotitic
374 data (52 samples out of the 88) show a very similar distribution to the total P-type
375 data set. The mean (-4.7 ‰) and median (-4.8 ‰) values of the single growth data are
376 only slightly higher for than the total dataset (-4.8 ‰ and -4.9 ‰, respectively), with
377 their standard deviations being very similar; 1.1 (total dataset) to 1.2 (single growth).
378 For the E-type single-phase growth data (21 samples out of 56), the second mode in
379 the $\delta^{13}\text{C}$ distribution at -12 to -9 ‰ is much more prominent than in the total dataset.
380 The main reason for this is that the second mode around -10 ‰ is predominantly
381 governed by the Argyle samples, which show less evidence for multiple growth
382 events (4 out of 10 samples), while the Orapa (11 out of 13 samples) and Jwaneng (11
383 out of 16 samples) data, which are centred around -5 ‰, are greatly reduced (only 51
384 out of 233 analyses remain). Despite this more prominent bimodality in the single

385 growth data set, the mean of the single growth data is slightly higher (-6.6 ‰) than for
386 the total dataset (-6.7 ‰) whilst the median is slightly lower (-5.9 ‰ versus -5.7 ‰).
387 From this comparison, we conclude that the P-type dataset may be interpreted within
388 a temporal context without significant concern that the carbon signal could be biased
389 by multiple growth events being mis-classified into single time intervals. For the E-
390 type samples, the coupling of the published age data to the new carbon data is clearly
391 less robust due to a prevalence of multi-stage growth; this will need to be kept in mind
392 when interpreting the dataset as a whole, although the differences between the mean
393 and median of the single growth subset compared with the total dataset are small.

394

395

396 *Literature Data*

397

398 In the discussion section below, the P and E-type data from this study are compared
399 with that from the literature. Literature data predominantly represent “bulk” (typically
400 fragments of ≤ 1 mg) combustion analyses. We compiled a database containing ~2600
401 monocrystalline samples (i.e. excluding fibrous, polycrystalline and micro-diamonds),
402 which subsequently is referred to as the “*total literature data*”.

403

404 Figure 1 shows that in addition to the present work there are more than 20 diamond
405 growth events (excluding fibrous growth) for which there is a geological age. To
406 expand both the number of growth events discussed and to compare the data
407 presented here with published data on the same deposits, we utilized data from the
408 literature. The key criterion for using literature data was that the mineral inclusion
409 paragenesis could be defined and clearly tied to a specific dated growth event. For
410 example, Finsch has an age for P-type diamond growth at 3.3 - 3.5 Ga (Richardson et
411 al., 1984) and for E-type growth at 1.5 Ga (Richardson et al., 1990), so carbon isotope
412 data for diamonds from this mine of known paragenesis (e.g. Appleyard et al., 2004;
413 Palot et al, 2013) can be attributed to either age. Conversely, Premier / Cullinan has
414 multiple growth ages for peridotitic paragenesis diamonds (see **Figure 1**) and since the
415 carbon isotope data cannot be tied to specific growth events, data from Cullinan
416 diamonds are not considered here.

417

418 A few studies with age constraints have utilized SIMS analysis, with multiple points
419 per sample, namely the Zimmi (Sierra Leone) data from Smit et al. (2019a), the
420 Koffiefontein data (Timmerman et al., 2019a) and some of the Finsch data (Palot et
421 al., 2013), the latter two both from South Africa. In these cases, the number of
422 samples is only 5, 3 and 10, but the number of individual analyses equate to 179, 12
423 and 113, respectively. The numbers of analyses and data statistics for the literature
424 data discussed below are provided in Table 4, while all of the original data are
425 compiled in the Supplementary Material (Table S3). This subset of literature data
426 (combining ‘bulk’ and SIMS data from age constrained deposits) will be referred to as
427 the “dated literature data”.

428

429

430 Discussion

431

432 *Carbon isotope systematics*

433

434 The range of carbon isotope values recorded in the global bulk diamond database (i.e.
435 total literature data) extends from -41 ‰ to +5 ‰ (see Cartigny et al., 2014 and
436 references therein). The tight clustering of $\delta^{13}\text{C}$ data for peridotitic diamonds (95 %
437 are within -5 ± 2.2 ‰) coincides with $\delta^{13}\text{C}$ data from other mantle sources (e.g.
438 carbonatite and kimberlite carbonates, mid-oceanic ridge basalts, volcanic CO_2
439 emissions, mantle xenoliths; see Deines (2002) and references therein), leading to a
440 general agreement that a $\delta^{13}\text{C}$ value of around -5 ‰ reflects the carbon isotopic
441 signature of the upper mantle. The negative skewness of $\delta^{13}\text{C}$ data for eclogitic
442 diamonds (~30 % are < -10 ‰) has been the source of debate for decades, and three
443 different explanations have been proposed:

444

445 (1) Primordial heterogeneity within the mantle (Deines et al., 1993).

446 (2) The fractionation of carbon isotopes in an open system. This can occur either
447 prior to diamond precipitation, where decarbonation may cause CO_2 enriched
448 in ^{13}C to escape, leaving behind a ^{13}C -depleted fluid (Cartigny et al., 1998;
449 2001), or it may be a consequence of carbon isotope fractionation during
450 diamond precipitation in a fluid-limited system (Deines, 1980).

451 (3) The recycling of crustal carbon via subduction, which has undergone
452 fractionation by near-surface process to produce a very large range of carbon
453 isotope values (Sobolev and Sobolev, 1980; Milledge et al., 1983; Kirkley et
454 al., 1991).

455

456 The concept of primordial heterogeneity within the mantle was based on the
457 observation that a single kimberlite could sample many sub-populations of diamonds,
458 with no apparent relationship between their physical and chemical characteristics
459 (Deines et al., 1993). It was therefore suggested that the heterogeneity observed in the
460 carbon isotope values was primordial, a signature acquired during Earth's accretion
461 and not subsequently homogenized by mantle convection. While support from this
462 model comes from similarity in the range of $\delta^{13}\text{C}$ data obtained from carbon in
463 meteorites and eclogitic diamonds (Deines 1980), it is implausible that ^{13}C depleted
464 primordial heterogeneities should be restricted almost exclusively to metabasaltic
465 diamond substrates. As such, this hypothesis will be excluded from further discussion
466 here, and the focus will be on understanding the arguments for and against the second
467 and third hypotheses.

468

469 Fractionation of carbon isotopes, both prior to and during diamond formation, have
470 been a significant part of the discussion surrounding the negative skewness of the
471 eclogitic diamond $\delta^{13}\text{C}$ record. Cartigny et al. (1998; 2001) proposed that carbon
472 isotopic fractionation would be driven by the separation of CO_2 from the diamond-
473 forming fluid/melt prior to diamond crystallisation, depleting the fluid/melt in ^{13}C .
474 While this process has not been ruled out, it has been calculated that only very small
475 amounts (<1 %) of the residual fluid/melt reaches a $\delta^{13}\text{C}$ value of <-14 ‰, and almost
476 none <-20 ‰ (Smart et al., 2011). So, it is thought unlikely that this process alone can
477 account for the eclogitic diamond $\delta^{13}\text{C}$ record.

478

479 The case for fractionation processes occurring during diamond growth are based upon
480 equilibrium fractionation factors for carbon isotopes at mantle temperatures, which
481 have been derived from theoretical calculations, experimental determination and
482 empirical measurements (Chacko et al., 2001; Satish-Kumar et al., 2011; Reutsky et
483 al., 2012; 2015; Horita & Polyakov, 2015). Diamond growth through redox reactions
484 (i.e. either the reduction of oxidized carbon species, or the oxidation of reduced

485 carbon species) produces characteristic $\delta^{13}\text{C}$ trends depending on the redox speciation
486 in the fluid/melt (see Cartigny et al., 2014). Diamonds growing from an oxidized
487 fluid/melt, containing CO_2 or CO_3^{2-} , will be offset to lower $\delta^{13}\text{C}$ values relative to the
488 source (the exact amount depends on the isotope fractionation factor between the fluid
489 carbon species and diamond), leading to successive enrichment in ^{13}C (Figure 5A, &
490 B). In a fluid-limited system, the resultant diamond population would exhibit a
491 positive skewness in the $\delta^{13}\text{C}$ distribution. Conversely, diamonds growing from a
492 reduced fluid/melt containing CH_4 , will be initially offset to heavier $\delta^{13}\text{C}$ values
493 relative to the source, but in a fluid-limited system will be successively enriched in
494 ^{12}C (Figure 5C). Empirical support for the fractionation hypothesis has come from
495 carbon isotope data obtained from both individual diamonds by SIMS (oxidized
496 fluid/melt – Zedgenizov et al., 2006; Smart et al., 2011; Palot et al., 2013; Mikhail et
497 al., 2014) and diamond populations via bulk analyses (reduced fluid/melt –
498 Thomassot et al., 2007), while the skewness of population distributions have also
499 been discussed (e.g. Deines, 1980; Stachel & Harris, 2009).

500

501 Recently, the effectiveness of redox reactions to account for diamond precipitation in
502 peridotitic substrates has been questioned. Luth & Stachel (2014) showed that the
503 redox buffering capacity of peridotitic mantle is insufficient to account for economic
504 quantities of diamonds to form. CHO fluids within depleted peridotitic mantle with
505 compositions near the water-maximum would not contain just one carbon species but
506 both oxidized and reduced (see Figure 1 of Stachel et al., 2017a; Stachel & Luth,
507 2015). Rayleigh isotope fractionation in multi-component systems is more complex
508 than in systems with a single carbon species; their modeling showed that regardless of
509 the dominant carbon species, diamond always progresses along ^{13}C enrichment paths
510 (Figure 5D). The predicted absence of progression towards more negative $\delta^{13}\text{C}$ values
511 even for fluid mixtures dominated by reduced carbon species ($\text{CH}_4/\text{CO}_2 > 1$), was
512 supported by SIMS data for Marange (Zimbabwe) diamonds that contained CH_4
513 micro-inclusions yet revealed core-to-rim trends towards heavier $\delta^{13}\text{C}$ values (Smit et
514 al., 2016; Stachel et al., 2017a). It is important, however, to highlight the limitations
515 and likely oversimplification of this model. Firstly, it is only truly applicable to
516 harzburgites (in lherzolites and eclogites conditions would be above their solidi,
517 therefore generating more complex melts), and secondly, pure CHO fluids are likely a
518 simplified starting point for the model - they have only been observed as inclusions in

519 diamonds very rarely. Analysis of more typical micro-inclusions containing diamond-
520 forming fluids/melts shows they contain significant proportions of dissolved material
521 (Weiss et al, 2009), which presently are not considered in this model. However,
522 thermodynamical modelling of hydrous solutions in equilibrium with peridotite or
523 eclogite at high pressures and temperatures up to 1000 °C (Sverjensky and Huang,
524 2015; Huang and Sverjensky, 2019) does predict the coexistence of a variety of
525 oxidized and reduced species together with high contents of major oxides.

526

527 The third hypothesis, which invokes recycled carbon as the source of the varied $\delta^{13}\text{C}$
528 values, is based around a subducted nature of eclogitic diamond substrates (e.g.
529 Helmstaedt et al., 1972). Altered oceanic crust (AOC; the igneous portion of oceanic
530 crust) shows considerable C isotopic variability. A recent study by Li et al. (2019)
531 suggests the AOC contains two carbon isotopic end-members; (1) a high $\delta^{13}\text{C}$ (~0 ‰)
532 carbonate component that forms during low-temperature alteration of the oceanic
533 crust, and (2) isotopically light biogenic carbonate. However, Valley & O'Neill
534 (1981) have shown that very negative $\delta^{13}\text{C}$ values can be formed through
535 fractionation (~600 – 800 °C) and may not necessarily require biogenic origins (see
536 Etiope & Sherwood Lollar, 2013). The occurrence of low $\delta^{13}\text{C}$ values in
537 sublithospheric diamonds, whose inclusions have been shown to represent a snapshot
538 of deep subduction (e.g. Deines et al., 1991; Tappert et al., 2005; Brenker et al., 2007;
539 Walter et al., 2011) provides empirical support for this subduction model, along with
540 oxygen isotope analyses of eclogitic garnet inclusions that clearly fall outside the
541 mantle range (Ickert et al., 2013; 2015; Schulze et al., 2013; Burnham et al., 2015).
542 Ickert et al. (2013) interpreted their $\delta^{13}\text{C}$ - $\delta^{18}\text{O}$ data to represent interaction (not
543 strictly two-component mixing) between fluids carrying 'mantle-like' carbon (i.e.
544 $\delta^{13}\text{C}$ ~-5 ‰) and subducted substrates with non-mantle-like $\delta^{18}\text{O}$; the only portion of
545 former AOC that carries enough carbon to not become overwhelmed in its carbon
546 isotope signature by the mantle-derived fluids are highly altered, and hence carbon-
547 enriched, former pillow basalts, where strongly negative $\delta^{13}\text{C}$ is invariably associated
548 with high $\delta^{18}\text{O}$ ($\delta^{18}\text{O} \geq 6$ ‰; Ickert et al. 2013). Li et al. (2019) propose that this link
549 between altered oceanic crust, mantle fluids, diamonds and their inclusions is also
550 capable of explaining the apparent decoupling of C and N isotopes in diamonds
551 (Mikhail et al., 2014; Hogberg et al., 2016).

552

553 In the present study, we will provide an impartial interrogation of the data reported,
554 along with a comparison with literature data, to evaluate how the above factors fit
555 within the realms of mantle fractionation and subduction models.

556

557 *Paragenetic Variations in the Diamond Record*

558

559 The $\delta^{13}\text{C}$ mode of both the P- and E-type paragenetic datasets is -5.0 ± 0.5 ‰; 39% of
560 this study's data falls within this tight range (37% for P-type, 41% for E-type) along
561 with 33% of the *total literature data* (obtained through bulk combustion analysis).

562 The proportions of data included increase significantly when only increasing the
563 modal range slightly; 61 % of this study's data fall within -5.0 ± 1.0 ‰ (68% for P-
564 type, 53% for E-type) along with 52% of total literature data; 81 % of this study's
565 data fall within -5.0 ± 2.0 ‰ (90% for P-type, 70% for E-type) along with 73% of
566 literature data. The isotopically light tail observed in the data is clearly confined in
567 this study to the E-type diamonds, with the lowest P-type $\delta^{13}\text{C}$ value being only -8.4
568 ‰.

569

570 Our extensive SIMS study (with little or no bias towards individual diamonds that are
571 isotopically heterogeneous and/or depleted in ^{13}C) has produced a dataset that is
572 largely in agreement with the *total literature database* obtained through bulk
573 combustion analyses. This provides significant support to the utility of bulk $\delta^{13}\text{C}$ data.
574 Bulk combustion averages out internal variations from multiple growth layers,
575 potentially missing out on reporting the full range of carbon isotope values present.
576 However, given that 66 % of our samples showed <1 ‰ internal $\delta^{13}\text{C}$ variation, with
577 86 % <2 ‰ and 96 % <3 ‰, it is clear that the bulk analysis of diamond fragments
578 will generally yield a good approximation of their carbon isotopic composition.

579

580 While both the P and E-type datasets have a mode around -5 ‰, it is clear from the
581 population statistics, and from previous studies, that eclogitic diamonds show more
582 variation and a greater prevalence of ^{13}C depleted isotopic values. The standard
583 deviation about the mean for the E-type data set (± 2.5 ‰) is more than double that for
584 the P-type data (± 1.1 ‰) and the E-type distribution is skewed to lower $\delta^{13}\text{C}$ values (-
585 1.3; Pearson's Coefficient of Skewness, see [Table 3](#)) compared to a negligible

586 positive skewness of the P-type data (0.2). This is broadly in keeping with the
587 literature data. The P-type literature data are, however, far more tightly distributed
588 around the -5 ‰ mode and nearly 34 % of published E-type $\delta^{13}\text{C}$ data fall below -10
589 ‰ (only ~16 % in this study), compared to only around 1 % of P-type diamonds (0 %
590 in this study). The average variation observed in individual P-type diamonds in this
591 study is 0.85 ‰, compared to a value of 1.53 ‰ in the E-type diamonds (which
592 includes the 4 samples with highly negative cores). Therefore, E-type diamonds
593 typically do exhibit more internal variation in $\delta^{13}\text{C}$ values compared with P-type
594 diamonds.

595 The E-type dataset's distribution has a more bimodal appearance than the literature
596 data (see Figure 4 of Cartigny et al., 2014). The main reason for this is the over-
597 representation of data from Argyle. With its prominently ^{13}C -depleted distribution (68
598 % of the Argyle data in this study is lower than -9 ‰, for the literature it is ~86 %) it
599 is over-represented in this study (18 %) compared to the literature (~9 %). However,
600 Argyle is not the only deposit in the world to have a signature of predominantly ^{13}C
601 depleted diamond carbon. Guaniamo (Venezuela), Jagersfontein (South Africa),
602 Dachine (French Guiana) and Jericho (Canada), all have population modes that are
603 even lower than seen in Argyle diamonds (see Figure 5 of Cartigny et al. (2014) and
604 references therein), so the bimodality may persist even if the dataset is expanded.

605

606 *Isotopic Trends within Individual Diamonds*

607

608 The advantage of *in situ* SIMS analysis of diamond is the power to reveal isotopic
609 core-to-rim trends within individual samples, which can provide evidence of Rayleigh
610 fractionation processes and potentially reveal the speciation of the diamond-forming
611 fluids. Diamond core-to-rim fractionation trends have been observed in several
612 diamond suites (Zedgenizov et al., 2006; Thomassot et al., 2007; Smart et al., 2011;
613 Palot et al., 2013; Wiggers de Vries et al., 2013; Petts et al., 2015; Smit et al., 2016;
614 2019a), however these trends are typically associated with small overall $\delta^{13}\text{C}$
615 variations <3 ‰. When we consider the data from the present study, there are no
616 convincing trends observed in any of the 144 samples examined in detail. **Figure 6**

617 shows typical traverses observed in the diamonds studied here, with non-systematic
618 variations in $\delta^{13}\text{C}$ and/or N contents. Two thirds of the samples analysed showed <1
619 ‰ variation in $\delta^{13}\text{C}$ (one third <0.5 ‰). Overall, our data set shows that diamonds
620 with recognizable fractionation trends are generally very rare. Despite a number of
621 publications reporting clear fractionation trends (see above), our globally
622 representative sample set documents that diamonds with systematic and significant
623 co-variations in $\delta^{13}\text{C}$ and [N] are the exception rather than the rule. This suggests that
624 diamond is a more passive recorder of changing $\delta^{13}\text{C}$ rather than being the cause of
625 $\delta^{13}\text{C}$ variability. The agreement in the modes between the bulk analytical data and our
626 *in situ* data strongly supports this view and we conclude that significant closed system
627 isotopic fractionation did not operate for the majority of individual growth events.
628 This implies that diamond growth typically occurs in temporally and spatially
629 localized systems with high excess of carbon in the fluid, i.e. relative to the mass of
630 growing diamonds.

631 This conclusion has important implications for the discussion surrounding the
632 syngeneity or protogeneity of inclusions in diamond (see Nestola et al. (2017) and
633 references therein), and the ramifications on diamond growth ages obtained from
634 these inclusions. If diamond inclusions are protogenetic, then in a process with high
635 fluid:rock ratios, the likelihood of the inclusions being isotopically reset at the time of
636 diamond formation is high. This would lead to inclusion ages being representative of
637 the diamond growth event (i.e. being synchronous), irrespective of whether they are
638 syngenetic. Although detailed study has shown that mineral inclusion size and
639 residence temperature play an important role (Nestola et al., 2019). Isotopic resetting
640 has been categorically shown for sulphide inclusions from Zimmi (Smit et al., 2016;
641 2019b) where inclusions in three different diamonds provide the same isochron ages,
642 yet their Re-Os and sulphur isotopes both show that the sulphides were existing in the
643 lithosphere prior to the diamond formation event.

644 The work of Stachel et al. (2017a) on Rayleigh isotopic fractionation in
645 multicomponent systems (RIFMS) during isochemical precipitation showed that in
646 water-maximum fluids within harzburgitic domains, fractionation could only account
647 for variations of <1 ‰. This is in keeping with the lack of significant variation seen in
648 the peridotitic diamonds of this and former studies, but is not the only potential

649 explanation (see following paragraph). In the simplest water-maximum RIFMS
650 model, with only CO₂ and CH₄ as carbon-bearing species, the fluid speciation ($X_{[CO_2]}$
651 = CO₂/[CO₂+CH₄]) and the dependent variable oxygen fugacity exert a strong control
652 on the carbon isotope composition of the precipitated diamond, irrespective of growth
653 mode (Figure 5D). For both isochemical and wall-rock buffered redox scenarios, the
654 carbon isotope composition of the first precipitated diamond only depends on the
655 $X_{[CO_2]}$ of the fluid and its assumed starting composition. Carbon isotopic evolution of
656 the fluid would differ depending on whether diamond formation occurs isochemically
657 or through wall rock-buffered redox reactions and whether the system is fluid-limited
658 or not. However, the variation is limited as the effective fractionation factor is the
659 weighted average of the factors of CH₄ and CO₂, which oppose each other. As both
660 species are consumed in equal amount the variation in fO_2 is also limited. The model
661 shows that the first precipitated diamond varies by 3.7 ‰ in $\delta^{13}C$ over a range of
662 $X_{[CO_2]}$ values from 0.9 to 0.1, which corresponds to a total increase in $\Delta \log fO_2$
663 (FMQ) of only about 0.8 log units. Assuming diamond-forming fluids with a constant
664 mantle-like $\delta^{13}C$ value of -5 ‰, it is, therefore, possible to account for the principal
665 mode of -5 ± 1 ‰ in the global harzburgitic diamond record with a variation in fO_2 of
666 just 0.4 log units, based on modelling of diamond formation at constant depth along a
667 geotherm.

668

669 The water-maximum fluids model is somewhat simplified (see earlier discussion in
670 *Carbon Isotope Systematics* section). However, as more realistic thermodynamic
671 models of subsolidus fluids also predict the presence of CO₂, carbonates and more
672 reduced organic species (e.g. Sverjensky and Huang, 2015; Tiraboschi et al., 2018)
673 the conclusions of the above modelling should still be applicable. If dissolved
674 carbonate joins the assemblage, it should act to further limit the fractionation. While
675 subsolidus fluids are expected in harzburgitic rocks along the 40 mW/m² model
676 geotherm, lherzolitic and eclogitic mantle would be above the hydrous solidus. Saline
677 and carbonatitic melts similar to the high-density fluids trapped in fibrous diamonds
678 may also infiltrate all these rocks. Still, as long as the fluids and melts carry both
679 oxidized carbonate and/or CO₂ together with reduced species, fractionation is
680 expected to be rather limited.

681

682 If diamond precipitation has only a limited ability to fractionate the carbon isotopic
683 composition of mantle fluids, and if their isotopic composition reflects that of the
684 fluid to within ± 1 ‰, then, the pronounced mode of diamond $\delta^{13}\text{C}$ values through
685 time at about -5 ‰ leads to the conclusion that the speciation and consequently
686 oxygen fugacity of diamond forming mantle fluids has been relatively consistent since
687 ~ 3.5 Ga. This also means that the deep (sublithospheric?) mantle has provided fluids
688 with near constant carbon speciation for the last 3.5 billion years.

689

690 At present there is no single model that is widely accepted, to account for the
691 increased variability of $\delta^{13}\text{C}$ values in eclogitic diamonds. The cause of this greater
692 variability in the $\delta^{13}\text{C}$ record of eclogitic diamonds may reflect either (1) diamond-
693 forming fluids with significantly more varied carbon characteristics - stemming from
694 subducted material containing highly variable biogenic and/or abiogenic carbon
695 sources, or (2) a growth environment more prone to significant isotopic fractionation
696 between the diamond and fluid/melt, or (3) a combination of both of these scenarios.
697 We note that an isotopically light carbon surface reservoir was available on Earth for
698 recycling to the mantle depths of diamond growth since 3.5 Ga (e.g. Des Marais,
699 2001) removing the need to produce this signal solely in the mantle by fractionation.

700

701

702 *The Diamond Record Through Time*

703

704 **Figure 7** places the diamond carbon isotope record into a chronological context using
705 data from this study and the *dated literature data*. The restricted nature of peridotitic
706 $\delta^{13}\text{C}$ values (as discussed previously) appears consistent through time, as does the
707 occurrence of isotopically light carbon (< -10 ‰) in eclogitic diamonds. However, the
708 mode of eclogitic $\delta^{13}\text{C}$ values is still at -5 ‰, consistent with that of the peridotitic
709 diamonds. Isotopically light carbon only represents a very small proportion of the
710 data; 6 % of the data from this study and 10 % of the *dated literature data* (excluding
711 the SIMS data) shown in **Figure 7**. The tail towards ^{13}C -depleted carbon in the
712 eclogitic diamonds also shows no signs of systematic variation through time (**Figure**
713 **7**), with its first occurrence in the record from 3 Ga onwards. With the oldest eclogitic
714 diamonds being tied to the onset of plate tectonics (Helmstaedt et al., 2010; Shirey &
715 Richardson, 2011), if the low $\delta^{13}\text{C}$ carbon signature is related to a biogenic carbon

716 contribution in altered oceanic crust (Li et al., 2019), this would imply that this
717 contribution, with respect to both amount and $\delta^{13}\text{C}$ values, has been fairly consistent
718 through 3 Ga. This is supported by the Precambrian marine carbonate isotope
719 database (Shields and Veizer, 2002). But the data in Figure 7 also shows that
720 isotopically light carbon, the potential key biomarker in altered oceanic crust, was not
721 recycled to mantle keel depths as extensively and consistently before 3 Ga as it was
722 after 3 Ga. Whether this was due to a geodynamic cause such as the absence of
723 eclogite recycling or a biologic cause such as the absence of biogenic components
724 deep in the oceanic crust or both is unknown. The important point is that the onset of
725 the temporal appearance of deeply-sourced, isotopically light carbon had only been
726 previously inferred from mineral inclusion paragenesis and general E-Type diamond
727 carbon isotope histograms. For the first time, this study has established the onset and
728 persistence of deep recycling of potentially biogenic carbon in the rock record by
729 direct analysis of the geochronologically dated carbon itself.

730

731 Looking at the $\delta^{13}\text{C}$ and N data together provides additional insight into potential
732 variations through time. Here we used multidimensional scaling (MDS) and principal
733 component analysis (PCA) to better visualise similarities and differences in the data.
734 MDS was only carried out on the $\delta^{13}\text{C}$ data as a function of their locations, while PCA
735 was carried out on the combined $\delta^{13}\text{C}$ and [N] datasets (see [Supplementary Materials](#)
736 for detailed information about this data processing methodology).

737

738 For the P-type data, MDS of the $\delta^{13}\text{C}$ values shows strong overlap between the
739 various populations with one exception ([Figure 8a](#)). There is clear discrimination
740 between Wawa and the two older diamond growth events at Diavik and De Beers
741 Pool ([Figure 8a](#)), while the three younger growth events (Venetia, Udachnaya and
742 Ellendale) exhibit more spread on the x -axis and overlap with all three of the older
743 populations. This observation is repeated in the PCA plot ([Figure 8b](#)). Despite N
744 concentration being the most significant variable in creating the spread in the data (i.e.
745 it is weighted most heavily in principal component 1), the discrimination caused by
746 just the $\delta^{13}\text{C}$ values is still retained in the PCA. This discrimination is the result of
747 $\delta^{13}\text{C}$ data for Wawa diamonds having a mode at -4‰ and its distribution having the
748 most positive skewness of all the P-type populations, while the older diamonds from
749 Diavik and De Beers Pool have strong modes at -5‰ and distributions with more

750 negative skewness. Despite this one minor difference in the data from 6 populations,
751 the PCA and MDS strongly reinforce the findings from the earlier interrogation of the
752 data (Figures 2 and 7), that there is no significant variation in $\delta^{13}\text{C}$ with geological
753 time for peridotitic diamonds.

754

755 Despite diamond growth at Wawa only being constrained by a minimum age, it is
756 interesting to consider the data in the broader context of the Superior Craton on which
757 the diamonds occur. Diamonds from the Renard (Quebec), Kyle Lake and
758 Attawapiskat (Ontario) kimberlites of the eastern and central Superior Craton,
759 respectively, show a temporal variation in $\delta^{13}\text{C}$. The Neoproterozoic Renard and T1
760 (Kyle Lake) kimberlites, along with Wawa (Neoproterozoic), carry diamond populations
761 with a mode in $\delta^{13}\text{C}$ of -4 ‰ (Hunt et al., 2012; Smit et al., 2014; Stachel et al., 2006)
762 while the Jurassic U2 (Attawapiskat) and Victor pipes exhibit a mode of -5 ‰ (Smit
763 et al., 2014; Stachel et al., 2017b). Between the emplacement of the Kyle Lake and
764 Attawapiskat kimberlites, the southern Superior craton experienced the major
765 Midcontinent Rift event at 1.1 Ga, which is considered to have been diamond-
766 destructive beneath the Attawapiskat area (Smit et al., 2014; Aulbach et al., 2018;
767 Stachel et al. 2018). Globally, our data show no discernible variation of mantle carbon
768 through time; yet, it is clear that detailed studies on the regional / craton scale have
769 the potential to reveal subtle temporal variations. However, it is not possible to
770 discern whether this change beneath the Superior Craton is caused by a small change
771 in the starting carbon isotopic composition, or the carbon speciation (and the
772 dependent variable oxygen fugacity) of the diamond-forming fluid.

773

774 While the MDS and PCA plots of the E-type data show more spread (Figure 9a & b)
775 than their corresponding P-type plots, the data from Argyle appear to be a slight
776 outlier. Even though the Argyle data are by no means fully discrete from any other
777 data set, these are the only diamonds with data entirely outside the main cluster. Since
778 the earliest proposals of subducted organic matter playing a role in diamond formation
779 (Sobolev & Sobolev, 1980; Milledge et al., 1983), Argyle has been the primary case
780 study to support this model (e.g., Jaques et al., 1989; Stachel et al., 2018) with its
781 dominance of eclogitic inclusions (~90%) and the normal distribution of $\delta^{13}\text{C}$ data
782 around a mode of -11 ‰. Detailed study of the diamonds and their inclusions have led
783 to the conclusion that eclogitic diamond growth predominantly occurred at unusually

784 high temperatures (~1250 – 1400 °C), indicating a depth at the very base of the
785 lithospheric mantle (Bulanova et al., 2018; Stachel et al., 2018). Recent noble gas
786 analyses also support a subducted origin for the Argyle eclogitic diamonds which
787 typically have low R/Ra values (<0.5; where $R/Ra = {}^3\text{He}/{}^4\text{He}_{\text{sample}} / {}^3\text{He}/{}^4\text{He}_{\text{air}}$) with the
788 noble gas subduction signatures focused at the base of the lithosphere, suggesting less
789 extensive fluid migration and interaction at higher levels with the sub-continental
790 lithospheric mantle (Timmerman et al., 2019b). Argyle reflects a strong case for
791 subduction-related diamond growth, but what potentially makes Argyle unique is the
792 amount of subducted carbon involved. Whether there was extensive mixing at greater
793 than average depths, of a large amount of isotopically light carbon with either mantle
794 carbon or isotopically heavier subducted carbonates, it created a fairly homogenized
795 diamond-forming fluid with an intermediate $\delta^{13}\text{C}$ signature. This contrasts with the
796 more common observation of a minor isotopically light (~-25 ‰) $\delta^{13}\text{C}$ signature
797 coupled with a dominant *mantle-like* (~-5 ‰) $\delta^{13}\text{C}$ signature.

798

799 There are two deposits, Orapa and Jwaneng, in our E-type dataset that have multiple
800 dated diamond growth events. This provides some insights into potential variations of
801 carbon through time within small sections of the mantle. The PCA of Orapa data
802 subdivided by age (Figure 9c) reveals some discrimination between the
803 subpopulations, with potential mixing between two end-members. The 2.1 Ga
804 samples have some overlap with the 3 Ga samples, but largely fall in a distinct group,
805 while the 1 Ga samples seemingly spread between the two. Potential endmember (1)
806 has more negative $\delta^{13}\text{C}$ values (-8 to -12 ‰) coupled with lower nitrogen
807 concentrations (<400 ppm), while endmember (2) has more *mantle-like* $\delta^{13}\text{C}$ values (-
808 4 to -8 ‰) and higher nitrogen concentrations (>400 ppm). This indicates potential
809 variation in the source of diamond forming fluids beneath Orapa through time, in
810 agreement with a similar conclusion by Timmerman et al. (2017) based upon carbon
811 and strontium isotope analyses of Orapa and Letlhakane diamonds. Although
812 Timmerman et al. (2017) discerned three potential endmembers, it is important to bear
813 in mind that only 2 of the 13 Orapa samples in this present study did not show
814 evidence of containing multiple growth events. Therefore, it is possible that the dating
815 of individual growth zones combined with detailed SIMS analyses may reveal a
816 clearer distinction as to identifying the potential sources of carbon through time at
817 Orapa. Interestingly, the PCA and MDS processing of the Jwaneng data reveal no

818 such subdivision with age. In fact, Orapa, and nearby Letlhakane, (Timmerman et al.,
819 2017) are the only deposits where a variation in the diamonds' carbon isotope record
820 with time has been documented (N.B. this excludes the consideration of fibrous and
821 polycrystalline diamond growth events, as well as those without multiple defined ages
822 at a single location). Despite the subtle variations seen at Orapa and the offset to
823 lower $\delta^{13}\text{C}$ seen at Argyle, the remaining E-type data reveal no systematic variations
824 in $\delta^{13}\text{C}$ through time on a global scale. This applies to the occurrence of both the $\delta^{13}\text{C}$
825 mode of -5 ‰ as well as the skewness of distributions to the more negative $\delta^{13}\text{C}$
826 values.

827

828 *Summary*

829

830 This comprehensive statistical interrogation of the diamond record for carbon isotope
831 values and nitrogen content through time revealed that, despite potential local
832 fluctuations, in general there is no systematic variation with time in the mantle carbon
833 isotope record over 3 billion years. The mode $\delta^{13}\text{C}$ of peridotitic diamonds has been at
834 -5 (± 2) ‰ since the earliest diamond growth ~ 3.5 Ga, and it is also observed in the
835 eclogitic diamond record since ~ 3 Ga. The skewed distribution of the $\delta^{13}\text{C}$ in eclogitic
836 diamonds to more negative values is also consistent through time, with no discernible
837 global trends apparent. However the onset of recycled, isotopically light C that was
838 suggested from Sm-Nd and Re-Os data on dated inclusions (Shirey & Richardson,
839 2011) is corroborated by a complete C isotope dataset.

840

841 On a local level, it is clear that individual deposits can have carbon isotope
842 distributions that are subtly (Wawa) or more obviously (Argyle) different to the global
843 norm. There is also scope to reveal variations over time in the diamond populations at
844 a single locality (Orapa). What is surprising, however, is that these variations are not
845 more common. The $\delta^{13}\text{C}$ range observed in eclogitic diamonds is a reflection of
846 complex and varied growth environments, potentially involving more significant
847 amounts of fractionation, and/or multiple, isotopically distinct carbon sources. A
848 marked divergence from the global norm is seemingly very limited, with Argyle being
849 the most notable (and most studied) exception. The more limited $\delta^{13}\text{C}$ range observed
850 in peridotitic diamond populations likely reflects a simpler growth environment,
851 associated with the influx of mantle derived CHO-fluids, with less inherent variability

852 compared to subduction-related scenarios. This more stable and consistent
853 environment potentially allows for more subtle variations in the $\delta^{13}\text{C}$ data to be
854 attributed to minor shifts in the chemistry of the diamond-forming fluids (e.g. Wawa,
855 Kyle Lake and the Attawapiskat diamonds on the Superior Craton).

856

857 Despite this potential for variation and heterogeneity, mantle carbon as recorded by
858 diamond has remained remarkably consistent over the course of Earth's history.

859

860 **Conclusions**

861

862 This study has utilized detailed *in situ* SIMS analyses of a large collection of
863 diamonds with established formation ages, providing insight into the deep carbon
864 record over a 3.5 Ga time period, from various cratons worldwide. This work reveals
865 a lack of discernible systematic variations in the global carbon isotope composition of
866 Earth's mantle through time as recorded by peridotitic diamonds. Despite eclogitic
867 diamonds having more varied C isotopic values on an individual crystal basis, on a
868 deposit-scale basis and globally, their C isotopic variability is generally consistent
869 overall and also lacks any discernible systematic temporal variation. While there are
870 clear examples in both the P-type (e.g. Wawa) and E-type (e.g. Argyle) populations of
871 individual deposits whose carbon records sit outside the global norm, the overall
872 consistency of the diamond record is remarkable. Examples of temporal variation
873 within individual diamond deposits (E-type diamond record at Orapa; Timmerman et
874 al., 2017) and on the craton-scale (Archean versus Neoproterozoic P-type diamonds
875 from the Superior Craton) do exist but are the exception. Despite these anomalies, the
876 global $\delta^{13}\text{C}$ diamond record for diamond growth in both peridotitic and eclogitic
877 substrates is remarkably consistent through time.

878

879 The fact that the *in situ* dataset reported here provides very good agreement with the
880 global "bulk" diamond record is a very important finding, which confirms earlier
881 combustion studies of internal heterogeneity (Galimov, 1984; Cartigny et al., 2004;
882 Mikhail et al., 2014). Despite single combustion analyses not recording internal
883 heterogeneity within individual diamonds, the global bulk dataset is representative of
884 the diamond record and interpretations based upon it have a solid foundation. A key
885 reason for this agreement is that the typical variation of $\delta^{13}\text{C}$ values in individual

886 diamond fragments sampled in this study is small, with two thirds of the samples
887 varying by <1 ‰. Additionally, no clear fractionation trends were observed in any of
888 the diamonds studied here, indicating that fluid-limited conditions during diamond
889 growth are very rare and proportionally over-represented in the literature. Instead,
890 diamond growth appears to typically occur in systems with excess of carbon in the
891 fluid, precluding significant Rayleigh fractionation during continued diamond
892 crystallization. A significant implication of this is that inclusions are very likely to be
893 isotopically reset at the time of diamond formation in a fluid-rich environment, either
894 through fluid-assisted equilibration/diffusion or more extensive partial melting,
895 meaning that the radiogenic isotope systematics of diamond inclusions reflect
896 formation ages irrespective of the inclusions being syngenetic or protogenetic with
897 respect to diamond growth (Smit et al., 2016; 2019b; Nestola et al., 2019).

898

899 The global literature diamond database has a remarkably consistent mode in $\delta^{13}\text{C}$ at -5
900 ‰ over 3.5 Ga. While the majority of samples in the database were not directly dated,
901 this present study has provided the temporal context to show that this mode in the
902 $\delta^{13}\text{C}$ is also consistent through time. Focusing on peridotitic diamonds that mostly
903 would have grown under subsolidus conditions (Stachel & Luth, 2015), the data from
904 this study show a tight grouping around this mode (90% fall within -5 ± 2 ‰), which
905 is also consistent with literature data. Modelling of this dataset through growth from
906 water-maximum CHO fluids reveals that this range can be accounted for by only a
907 small variation in the carbon speciation of the fluid (0.4 log units), which in turn
908 reflects a remarkably consistent redox state of these diamond-forming fluids for 3.5
909 Ga.

910

911 Despite eclogitic diamond growth not necessarily occurring in a subsolidus
912 environment (Stachel & Luth, 2015), the E-type diamond record also has a consistent
913 $\delta^{13}\text{C}$ mode of -5 ‰ (70% of data fall within -5 ± 2 ‰). Total $\delta^{13}\text{C}$ variation in
914 individual diamonds is generally minor (typically <2 ‰); when the internal variation
915 is larger, it is often occurring due to multiple discrete growth events. Even when
916 carbon isotope fractionation trends have been recorded in some rare diamond
917 examples, the total variation has always been <3 ‰. This suggests that the
918 environment for diamond growth is not conducive to generating significant variations
919 through fractionation processes alone. In view of the strong negative skewness of the

920 E-type diamond record, down to values of ~ -40 ‰ (De Stefano et al., 2009), this
921 range of $\delta^{13}\text{C}$ values is considered to be predominantly the result of variations in the
922 carbon signal of the primary fluids/melt. The subducted nature of eclogites,
923 containing some isotopically light abiogenic and/or biogenic carbon, is the prevalent
924 model to explain the cause of the negative skew in eclogitic diamond $\delta^{13}\text{C}$ values
925 (Sobolev & Sobolev, 1980; Milledge et al., 1983; Kirkley et al., 1991).

926

927 The earliest eclogitic diamonds have been tied to the onset of subduction and plate
928 tectonics (Helmstaedt et al., 2010; Shirey & Richardson, 2011). The appearance of
929 isotopically light carbon at 3 Ga and its persistence in the eclogitic paragenesis after
930 that time affirms a secular change in recycled carbon at this time. If this is indeed the
931 case, then ultimately the net effect of subduction has not altered the ambient signal of
932 the mantle's carbon record. Whether this is due to the net carbon isotopic composition
933 of a subducted plate being about -5 ‰ (Shilobreeva et al., 2011), or the fact that
934 recycled carbon is just a very minor contribution in the overall mantle carbon budget,
935 remains to be proven.

936

937

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1224

1225

1226 **Figure Captions**

1227

1228 **Figure 1:** Illustration showing the occurrence of diamond growth events throughout
1229 Earth's history. The deposits are grouped based upon their location; Australia (left),
1230 Canada, Russia, and Africa (right). White symbols represent deposits studied in this
1231 paper. Square symbols represent peridotitic diamond growth events (Hz and Lh
1232 denote harzburgitic and lherzolitic classifications where reported), while diamond
1233 symbols represent eclogitic diamond growth events. Fibrous diamond growth events
1234 (circle symbols) noted here reflect the kimberlite eruption age. A table containing all
1235 of the dates and references are provided in the Supplementary Material (**Table S1**).

1236

1237 **Figure 2:** Histograms for all of the analyses of peridotitic (right-hand side) and
1238 eclogitic (left-hand side) diamonds studied here. Where available, published literature
1239 data (combustion data only, provided in **Table S3**) are shown in grey. The
1240 representation of the two datasets is not additive (one on top of the other) but one is
1241 set in front of the other. Note that the scales on the y-axis vary between plots. The
1242 histograms use 0.5 ‰ binning.

1243

1244 **Figure 3:** Cathodoluminescence images of a selection of samples showing varied
1245 growth histories. (A) JWR 8; Jwaneng, (B) ORS 2; Orapa, (C) DP-4; De Beers Pool,
1246 (D) DDMI 199; Diavik, (E) JWR 18; Jwaneng. All scale bars are 200 μm long.

1247

1248 **Figure 4:** Histograms showing the data obtained from all of the samples, compared to
1249 those where only diamonds whose CL indicates single growth event are counted,
1250 separated based upon inclusion type.

1251

1252 **Figure 5:** Theoretical fractionation plots for diamond growth from a fluid with a
1253 starting $\delta^{13}\text{C}$ value of -5 ‰ (yellow star; f_c represents the remaining fraction of
1254 carbon-bearing species in the diamond-forming fluid). The offset between the initial
1255 diamond growth and the fluid is governed by the carbon species in the fluid. The plots
1256 show diamond growth from (a) a CO_2 (oxidized) fluid, (b) a CO_3 (oxidized) fluid, and
1257 (c) a CH_4 (reduced) fluid. Plot (d) shows a more complex multicomponent system,
1258 with the fluid containing both oxidized and reduced species (namely CO_2 and CH_4) in
1259 varying proportions (10 mol% CO_2 - blue, 50 % - black, 89 % - red lines). Note that
1260 in these multicomponent calculations, the diamond's $\delta^{13}\text{C}$ values always trend to less
1261 negative values irrespective of the dominant carbon species. Plot (d) after Stachel et
1262 al. (2017a).

1263

1264 **Figure 6:** CL images of (A) DBP-363_5, (B) V-336, (C) DBP-9, (D) DBP-460_9,
1265 with their SIMS $\delta^{13}\text{C}$ analysis locations marked and red arrows showing growth
1266 direction. The graphs show the $\delta^{13}\text{C}$ values plotted against N concentration. The lines
1267 connect the analyses in order, arrows showing the growth direction.

1268

1269 **Figure 7:** Box and Whisker plot showing the $\delta^{13}\text{C}$ data from this study (bold colours),
1270 as well as data from the literature (combustion plus SIMS data; faint colours), in a

1271 chronological order. Where single outliers are more than 3 ‰ from the remaining
1272 data, they have been marked as single data points (in Panda, Diavik and Venetia). The
1273 literature data plotted here is provided in **Table 3** (with the total dataset, including
1274 references, provided in **Table S3**).

1275

1276 **Figure 8:** MDS and PCA plots for peridotitic diamonds. (a) MDS plot showing only
1277 $\delta^{13}\text{C}$ data. Note how the Wawa data (red diamonds) do not overlap with the older
1278 Diavik (black circles) and De Beers Pool (black squares) data. (b) PCA plot of $\delta^{13}\text{C}$ -
1279 N data. Given that MDS is a transformation that assigns input information a unitless
1280 (arbitrary unit) value (as it preserves only the relationships between data rather than
1281 specific information), we have not placed a scale on either axis of the MDS plot.

1282

1283 **Figure 9:** MDS and PCA plots for eclogitic diamonds. (a) MDS plot showing only
1284 $\delta^{13}\text{C}$ data. (b) PCA plot of $\delta^{13}\text{C}$ -N data. The dashed line represents a contour which
1285 contains 71% of all the data. Note how the Argyle data falls almost exclusively
1286 outside of this cluster. (c) PCA plot showing $\delta^{13}\text{C}$ - N data from only the Orapa
1287 samples.

1288

1289 **Table 1:** Details of the diamond samples

1290 Information regarding the source locations and quantities of the diamonds studied in
1291 this paper. References provided are the original works that dated the inclusions. In
1292 cases where two references are provided, the second notes the study where the
1293 diamonds have been sourced from which is separate to the original dating study.

1294

1295 **Table 2:** Summary of the $\delta^{13}\text{C}$ and N data

1296 The statistics generated from the $\delta^{13}\text{C}$ and N datasets for both peridotitic and eclogitic
1297 diamonds in this study. The Δ symbol represents the ‘variation per sample’. The ‘Pear
1298 Skew’ value is the Pearson Coefficient of Skewness which is calculated as $[3 \times$
1299 $(\text{average} - \text{median})] / \text{standard deviation}$. The “average of samples’ averages /
1300 medians” rows of data have been obtained by taking either the average or median
1301 values from each sample individually, and then taking an average of them per deposit.
1302 This was done to see if over analysis of an individual diamond was affecting the
1303 distribution of the data as a whole. The final rows show the differences between the

1304 data as a whole and of the individual stones. The average of these differences are ~0.1
1305 ‰ for the P-type data, and ~0.2 ‰ for the E-type data. Only the Orapa 3 Ga and 2.1
1306 Ga samples exceed differences of 0.25 ‰, showing that overall individual samples
1307 have not been over-analysed in this study.

1308

1309 **Table 3:** Summary of single growth vs multiple growth data

1310 The statistics of the total P-type and E-type $\delta^{13}\text{C}$ and N datasets compared against the
1311 reduced datasets obtained from samples that exhibit only a single growth event (as
1312 interpreted from their CL imagery).

1313

1314 **Table 4:** Literature $\delta^{13}\text{C}$ data

1315 Table showing the statistics of the literature data presented in **Figure 7**. *n* represents
1316 the number of diamonds, with each diamond providing a single data point. * Three
1317 locations utilize SIMS data where the first number is the number of diamonds, while
1318 the subsequent number in brackets represents the total number of analyses. The total
1319 dataset summarized here is provided in **Table S3**, including all relevant references.

1320

1321

1322 SUPPLEMENTARY MATERIAL:

1323

1324 Table S1) References for Figure 1.

1325 This table contains all of the references used to make Figure 1, showing all of the
1326 documented diamond growth events in Africa, Russia, Canada and Australia.

1327

1328 Table S2) Total $\delta^{13}\text{C}$ and N data from this study

1329 This table contains all of the $\delta^{13}\text{C}$ and N data obtained in this study, with the diamond
1330 samples separated based upon their inclusion paragenesis.

1331

1332 Table S3) Literature $\delta^{13}\text{C}$ data

1333 This table contains all of the literature data, referenced in the paper and its figures,
1334 along with the references from which the data were obtained.

1335

1336 Supplementary Material 1: Detailed methodology used for the multi-dimensional
1337 scaling (MDS) and principal component analysis (PCA) presented in this study.

1338

1339 Supplementary Material 2: Reference list for Table S1 and S3.

1340

1341

Table 1

Location	Craton	Paragenesis	Number of samples	Age
Argyle	Kimberley	Eclogitic	10	1.58 Ga
De Beers Pool	Kaapvaal	Eclogitic	8	2.9 Ga
De Beers Pool	Kaapvaal	Peridotitic	21	3.2 Ga
Diavik	Slave	Eclogitic	9	1.86 Ga
Diavik	Slave	Peridotitic	9	~ 3.3 Ga
Ellendale	Kimberley	Peridotitic	13	1.4 Ga
Jwaneng	Kaapvaal	Eclogitic	16	1.0 / 2.1 / 3.0 Ga
Orapa	Kaapvaal	Eclogitic	13	1.0 / 2.1 / 3.0 Ga
Udachnaya	Siberian	Peridotitic	15	2.0 Ga
Venetia	Limpopo Belt	Peridotitic	19	< 2.3 Ga
Wawa	Superior	Peridotitic	11	> 2.7 Ga

Table 2

	Orapa 3 Ga		Jwaneng 3 Ga		De Beers P
	$\delta^{13}\text{C}$ (VPDB)	[N] at. ppm	$\delta^{13}\text{C}$ (VPDB)	[N] at. ppm	$\delta^{13}\text{C}$ (VPDB)
No. of samples	6	6	6	6	8
No. of analyses	51	51	52	52	51
average	-6.10	596	-5.96	747	-5.61
min	-9.24	8	-17.09	1	-6.84
max	-4.38	1249	-2.98	2661	-3.30
average Δ	0.86	701	3.71	937	1.42
min Δ	0.28	318	0.43	231	0.53
max Δ	1.34	1166	12.13	2516	2.55
Q1	-7.70	171	-5.37	397	-6.11
Q3	-4.92	920	-4.83	1060	-5.27
median	-5.35	748	-5.13	891	-5.64
st dev	1.55	397	3.10	519	0.71
Pear Skew	-1.45	-1.15	-0.80	-5.15	0.15
Average of samples'					
averages	-6.83	406	-6.18	714	-5.52
medians	-6.84	408	-6.52	606	-5.44
Difference between					
averages	0.73	190	0.23	33	-0.09
medians	1.49	340	1.39	285	-0.20
	Diavik ~3.3 Ga		De Beers Pool 3.2 Ga		Wawa :
	$\delta^{13}\text{C}$ (VPDB)	[N] at. ppm	$\delta^{13}\text{C}$ (VPDB)	[N] at. ppm	$\delta^{13}\text{C}$ (VPDB)
No. of samples	9	9	21	21	11
No. of analyses	56	57	135	135	65
average	-5.12	132	-5.45	76	-3.58
min	-6.29	0	-8.39	2	-4.48
max	-4.19	1131	-2.87	1477	-1.87
average Δ	0.55	334	1.20	230	0.35
min Δ	0.18	46	0.19	4	0.08
max Δ	1.47	1124	2.51	1463	0.95
Q1	-5.39	11	-5.72	10	-4.16
Q3	-4.52	120	-4.98	45	-2.96
median	-5.27	44	-5.26	20	-3.76
st dev	0.55	236	0.84	195	0.72
Pear Skew	0.81	-0.56	-0.66	0.86	0.75
Average of samples'					
averages	-5.04	131	-5.47	5	-3.35
medians	-5.05	80	-5.45	28	-3.36
Difference between					
averages	-0.08	0	0.03	71	-0.23

Table 3

	TOTAL P-TYPE DATA		SINGLE GROWTH P-TYPE		TOTAL E-T
	$\delta^{13}\text{C}$ (VPDB)	[N] at. ppm	$\delta^{13}\text{C}$ (VPDB)	[N] at. ppm	$\delta^{13}\text{C}$ (VPDB)
No. of samples	88	88	45	45	56
No. of analyses	504	506	244	245	404
average	-4.79	221	-4.73	171	-6.74
min	-8.39	0	-8.39	2	-17.09
max	-1.87	3073	-1.87	1275	-2.07
average Δ	0.85	369	0.49	147	1.53
min Δ	0.07	4	0.07	4	0.10
max Δ	2.80	3066	2.56	1013	12.13
Q1	-5.42	19	-5.38	21	-7.90
Q3	-4.07	222	-4.02	204	-5.09
median	-4.87	52	-4.76	45	-5.65
st dev	1.12	369	1.16	263	2.48
Pear Skew	0.23	1.37	0.08	1.44	-1.32

Table 4

Location	PANDA	DIAVIK	FINSCH	DE BEERS POOL	DE BEERS POOL
Age (Ga)	3.52	3.5-3.3	3.3-3.3	3.3-3.2	2.9
Paragenesis	P-TYPE	P-TYPE	P-TYPE	P-TYPE	E-TYPE
n=	129	110	70	213	53
min	-14.05	-11.50	-8.57	-6.98	-15.97
Q1	-5.45	-5.44	-6.59	-5.42	-5.62
median	-5.12	-5.13	-6.12	-5.18	-5.13
Q3	-4.86	-4.62	-5.64	-4.84	-4.59
max	-2.98	-2.10	-2.62	-0.93	-3.06

Figure 1

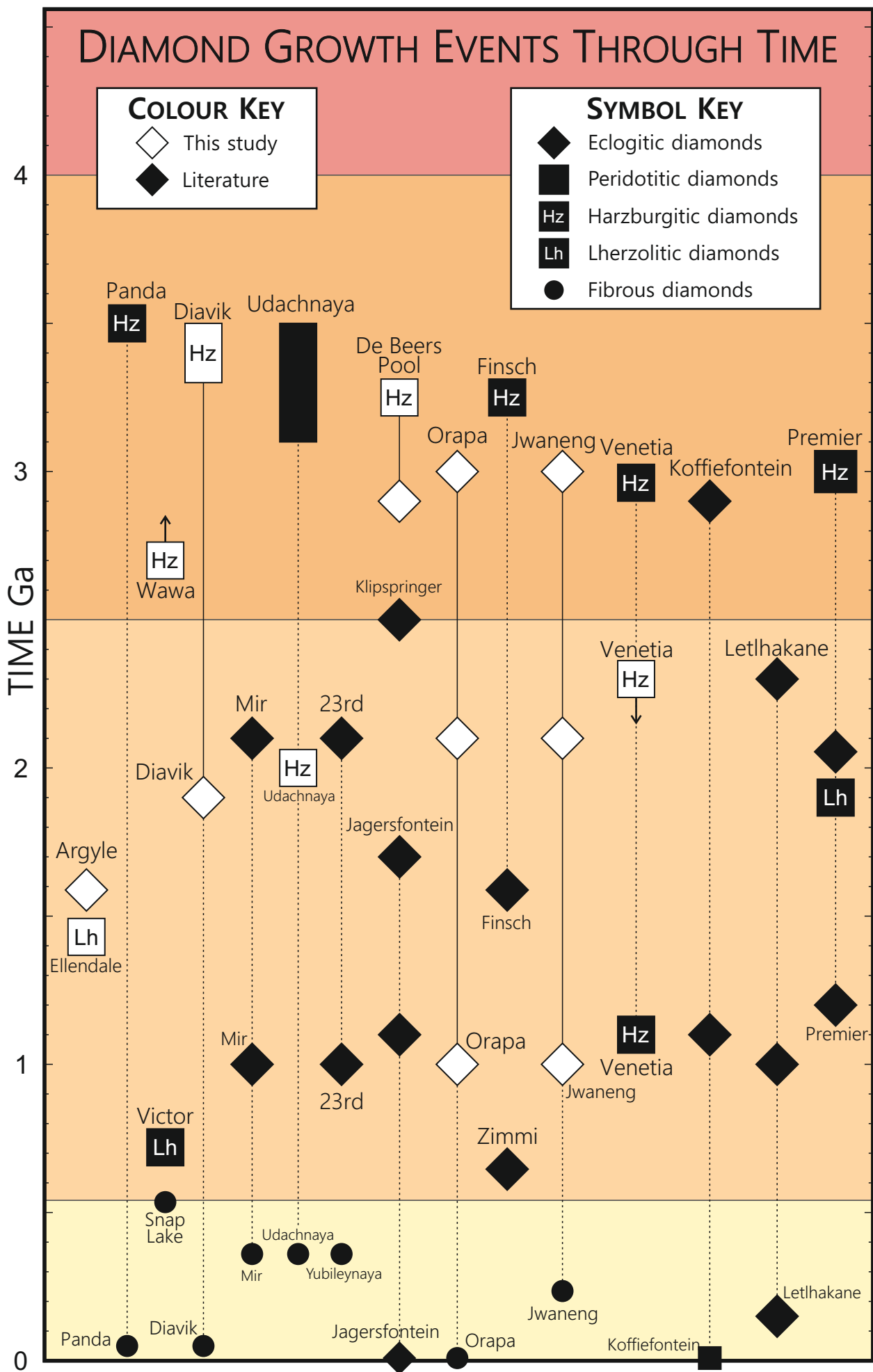


Figure 2

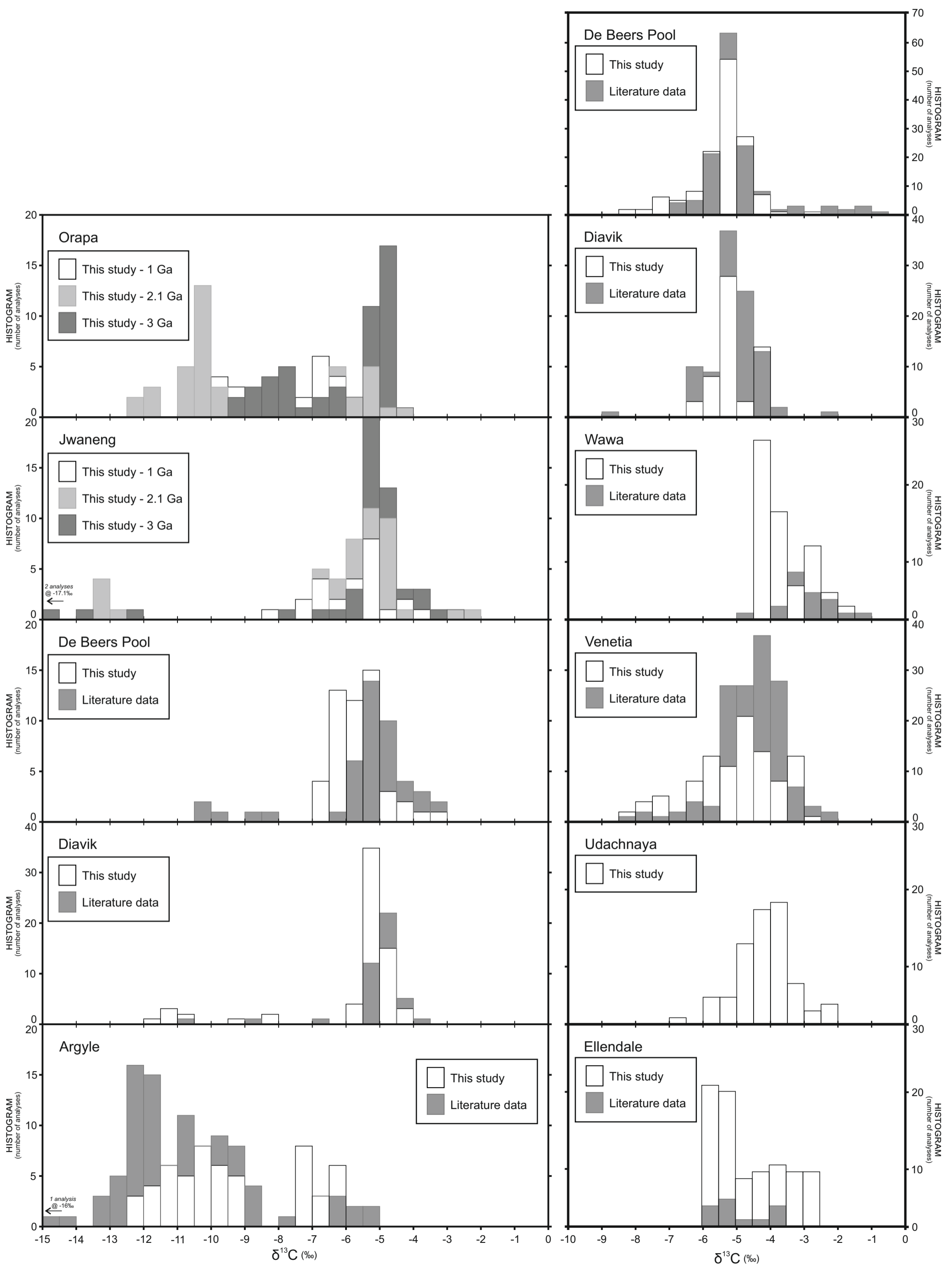


Figure 3

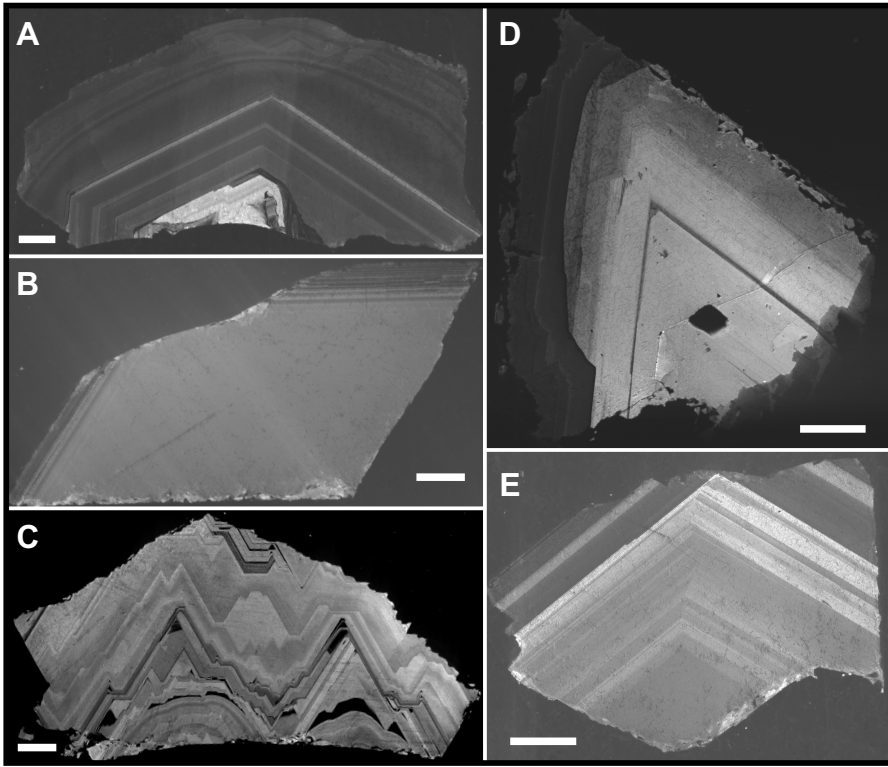


Figure 4

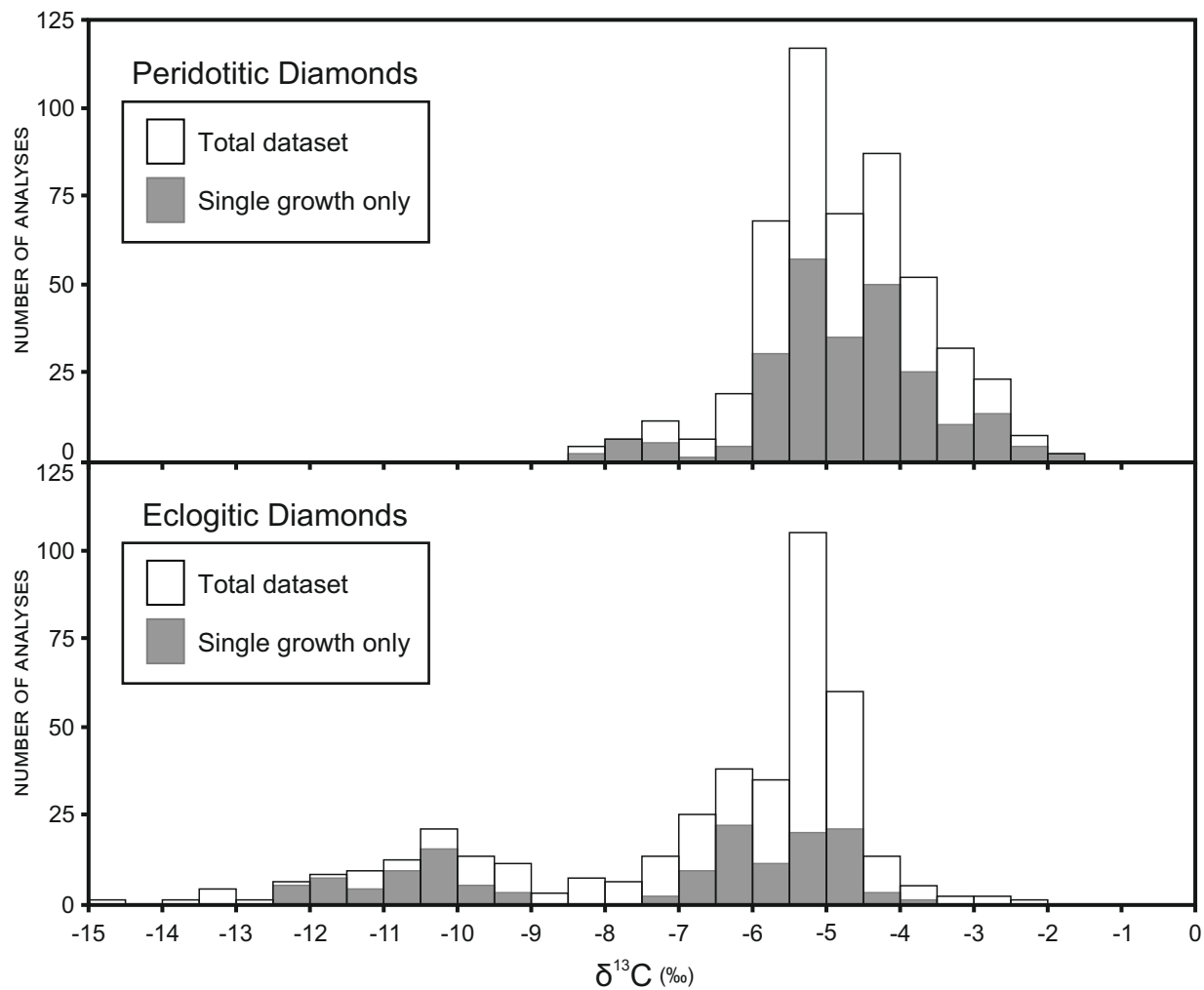


Figure 5

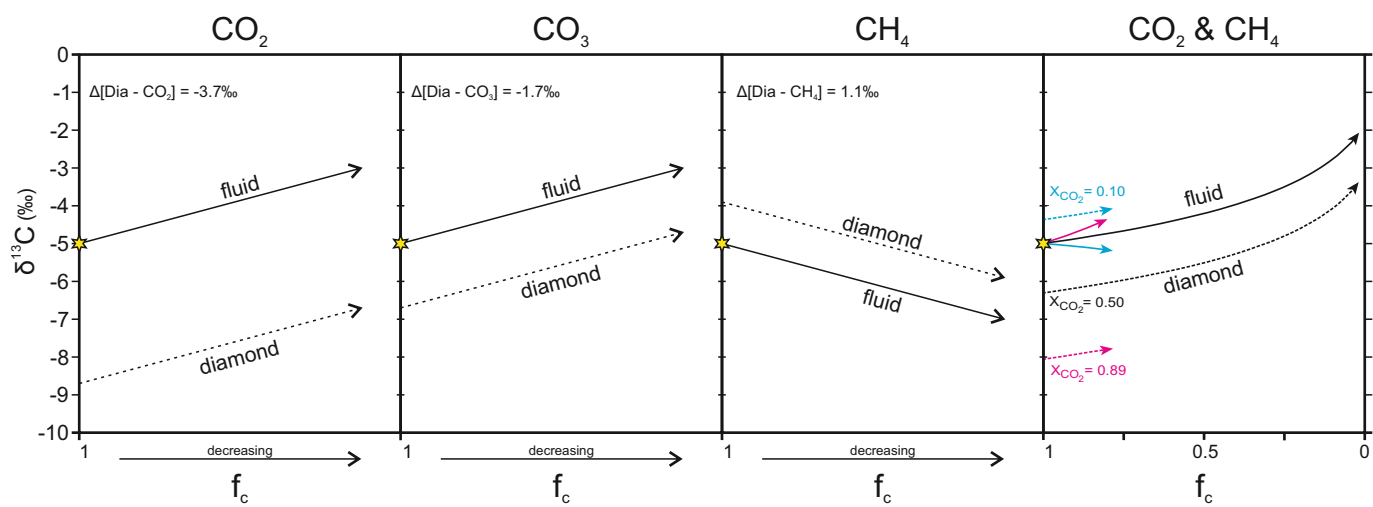


Figure 6

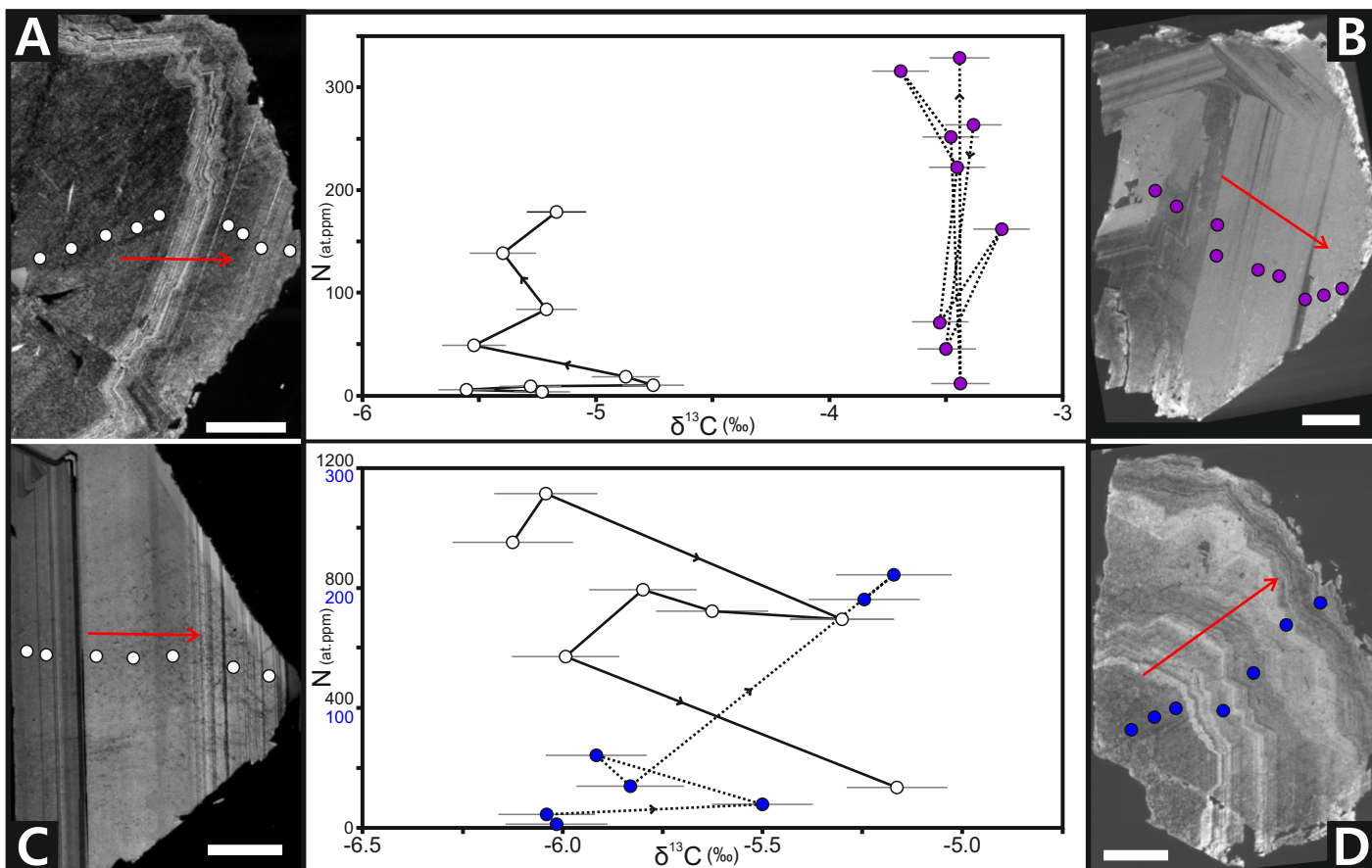
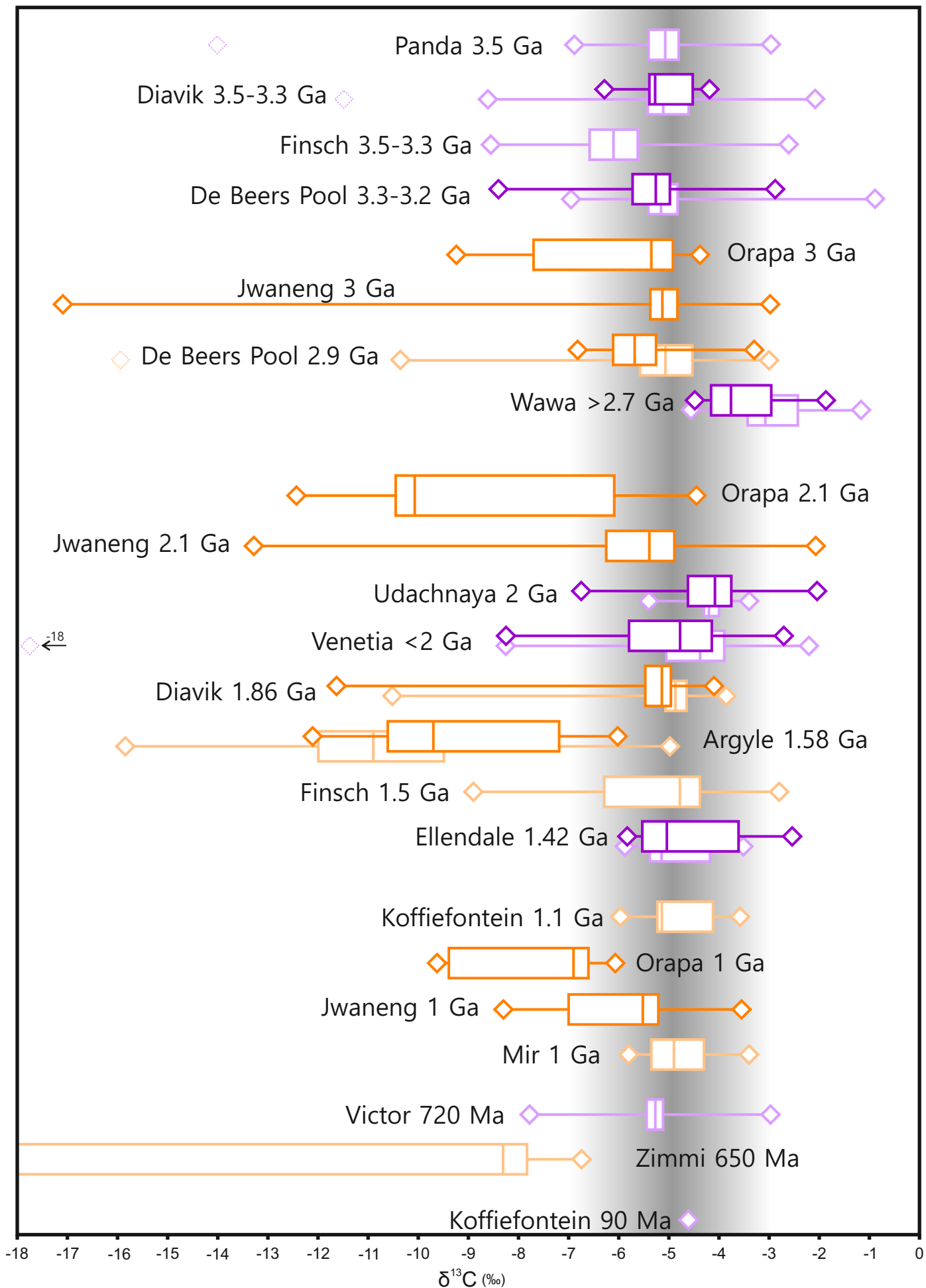


Figure 7



Eclogitic data - This Study

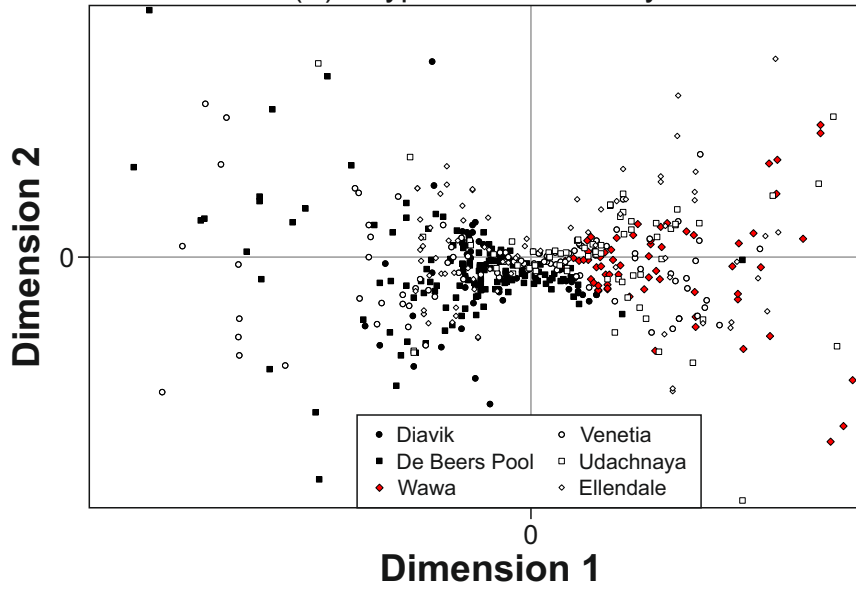
Peridotitic data - This Study

Eclogitic data - Literature

Peridotitic data - Literature

Figure 8

(A) P-type MDS: $\delta^{13}\text{C}$ only



(B) P-type PCA: $\delta^{13}\text{C}$ - N

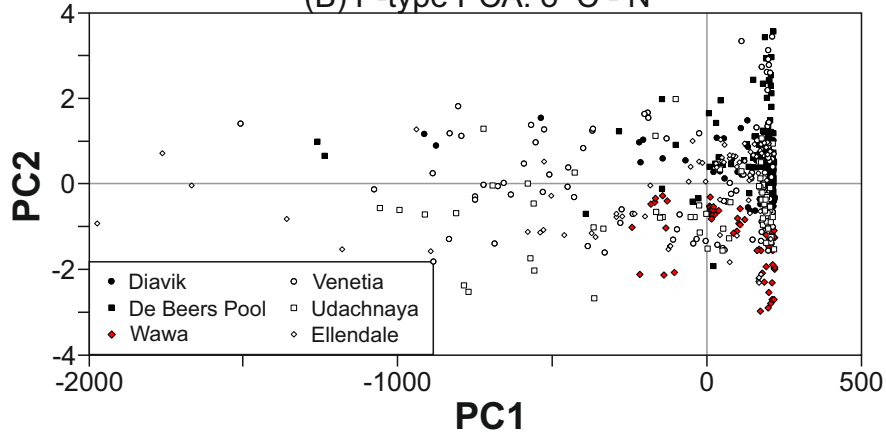
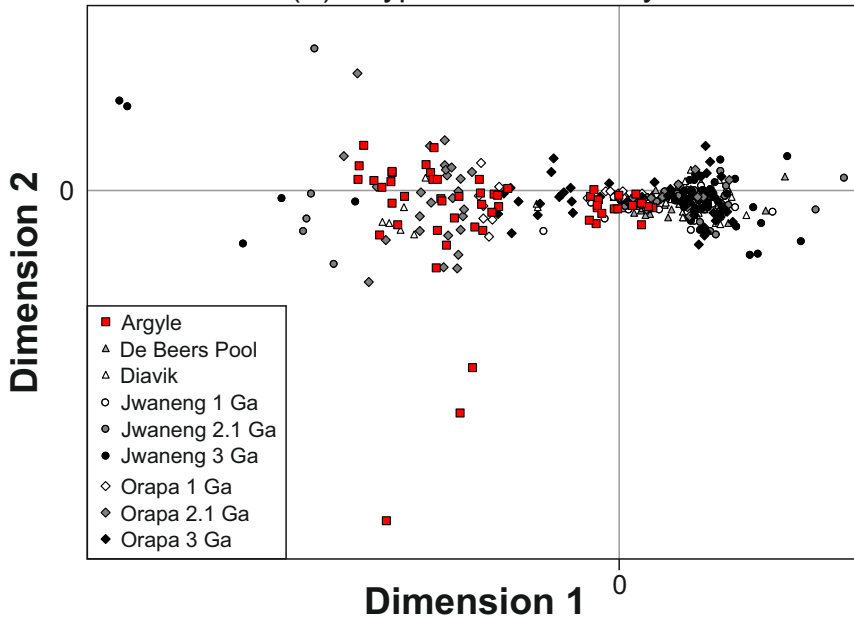
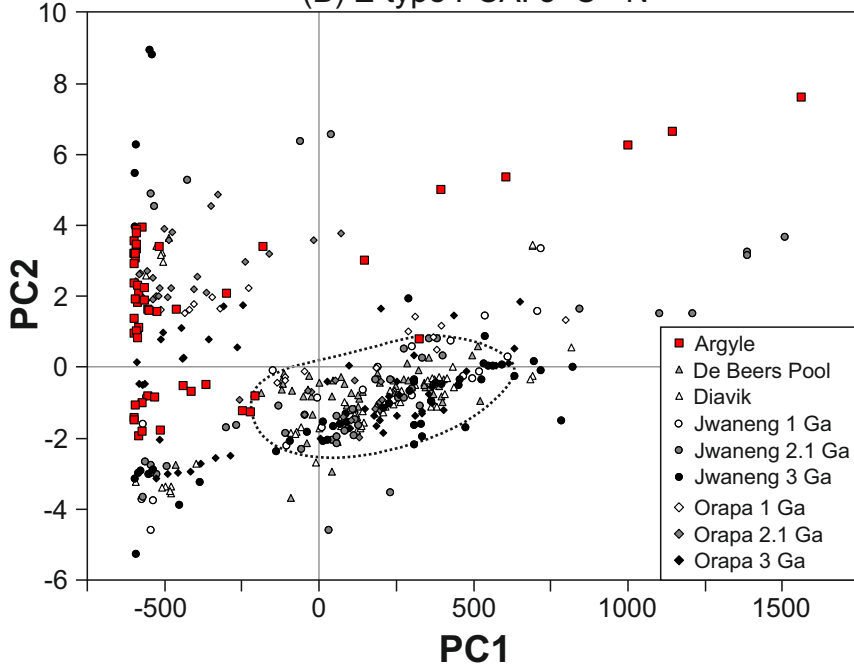


Figure 9

(A) E-type MDS: $\delta^{13}\text{C}$ only



(B) E-type PCA: $\delta^{13}\text{C}$ - N



(C) Orapa only PCA: $\delta^{13}\text{C}$ - N

