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# APPROXIMATE DECONVOLUTION MODELS FOR MAGNETOHYDRODYNAMICS 

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#### Abstract

We consider the family of approximate deconvolution models (ADM) for the simulation of the large eddies in turbulent viscous, incompressible, electrically conducting flows. We prove existence and uniqueness of solutions, we prove that the solutions to the ADM-MHD equations converge to the solution of the MHD equations in a weak sense as the averaging radii converge to zero, and we derive a bound on the modeling error. We prove that the energy and helicity of the models are conserved, and the models preserve the Alfvén waves. We provide the results of the computational tests, that verify the accuracy and physical fidelity of the models.


Key words. Large eddy simulation, magnetohydrodynamics, approximate deconvolution
AMS subject classifications. 65M12, 76F65, 76W05

## 1. Introduction.

Magnetically conducting fluids arise in important applications including climate change forecasting, plasma confinement, controlled thermonuclear fusion, liquid-metal cooling of nuclear reactors, electromagnetic casting of metals, MHD sea water propulsion. In many of these, turbulent MHD (magnetohydrodynamics [2]) flows are typical. The difficulties of accurately modeling and simulating turbulent flows are magnified many times over in the MHD case. They are evinced by the more complex dynamics of the flow due to the coupling of Navier-Stokes and Maxwell equations via the Lorentz force and Ohm's law.

The flow of an electrically conducting fluid is affected by Lorentz forces, induced by the interaction of electric currents and magnetic fields in the fluid. The Lorentz forces can be used to control the flow and to attain specific engineering design goals such as flow stabilization, suppression or delay of flow separation, reduction of near-wall turbulence and skin friction, drag reduction and thrust generation. There is a large body of literature dedicated to both experimental and theoretical investigations on the influence of electromagnetic force on flows (see e.g., $[16,24,25,14,15,36,10,37,17,33,4,11])$. The MHD effects arising from the macroscopic interaction of liquid metals with applied currents and magnetic fields are exploited in metallurgical processes to control the flow of metallic melts: the electromagnetic stirring of molten metals [26], electromagnetic turbulence control in induction furnaces [38], electromagnetic damping of buoyancy-driven flow during solidification [28], and the electromagnetic shaping of ingots in continuous casting [30].

Direct numerical simulation of a $3 d$ turbulent flow is often not computationally economical or even feasible. On the other hand, the largest structures in the flow (containing most of the flow's energy) are responsible for much of the mixing and most of the flow's momentum transport. This led to various numerical regularizations; one of these is Large Eddy Simulation (LES) [29], [18], [5]. It is based on the idea that the flow can be represented by a collection of scales with different sizes, and instead of trying to approximate all of them down to the smallest one, one defines a filter width $\delta>0$ and computes only the scales of size bigger than $\delta$ (large scales), while the effect of the small scales on the large scales is modeled. This reduces the number

[^0]of degrees of freedom in a simulation and represents accurately the large structures in the flow.

In [20] we considered the problem of modeling the motion of large structures in a viscous, incompressible, electrically conducting, turbulent fluid. We introduced a simple closed LES model, and performed full numerical analysis. This model can be also addressed as zeroth order Approximate Deconvolution Model - referring to the family of models in [1]. In this report we consider the family of the Approximate Deconvolution Models for MagnetoHydroDynamics (ADM for MHD); we perform the numerical analysis of the models and also verify their physical fidelity.

The mathematical description of the problem proceeds as follows. Assuming the fluid to be viscous and incompressible, the governing equations are the Navier- Stokes and pre-Maxwell equations, coupled via the Lorentz force and Ohm's law (see e.g. [32]). Let $\Omega=(0, L)^{3}$ be the flow domain, and $u(t, x), p(t, x), B(t, x)$ be the velocity, pressure, and the magnetic field of the flow, driven by the velocity body force $f$ and magnetic field force curl $g$. Then $u, p, B$ satisfy the MHD equations:

$$
\begin{align*}
& u_{t}+\nabla \cdot(u u)-\frac{1}{\mathrm{Re}} \Delta u+\frac{S}{2} \nabla\left(B^{2}\right)-S \nabla \cdot(B B)+\nabla p=f \\
& B_{t}+\frac{1}{\mathrm{Re}_{\mathrm{m}}} \operatorname{curl}(\operatorname{curl} B)+\operatorname{curl}(B \times u)=\operatorname{curl} g  \tag{1.1}\\
& \nabla \cdot u=0, \nabla \cdot B=0
\end{align*}
$$

in $Q=(0, T) \times \Omega$, with the initial data:

$$
\begin{equation*}
u(0, x)=u_{0}(x), \quad B(0, x)=B_{0}(x) \quad \text { in } \Omega \tag{1.2}
\end{equation*}
$$

and with periodic boundary conditions (with zero mean):

$$
\begin{equation*}
\Phi\left(t, x+L \mathrm{e}_{i}\right)=\Phi(t, x), i=1,2,3, \quad \int_{\Omega} \Phi(t, x) d x=0 \tag{1.3}
\end{equation*}
$$

for $\Phi=u, u_{0}, p, B, B_{0}, f, g$.
Here $\operatorname{Re}, \operatorname{Re}_{\mathrm{m}}$, and $S$ are nondimensional constants that characterize the flow: the Reynolds number, the magnetic Reynolds number and the coupling number, respectively. For derivation of (1.1), physical interpretation and mathematical analysis, see $[6,21,31,13]$ and the references therein.

The necessity of choosing different filtering widths has been justified computationally for coupled Navier-Stokes problems in [27]. If ${ }^{-\delta_{1}}$, ${ }^{-\delta_{2}}$ denote two local, spacing averaging operators that commute with the differentiation, then averaging (1.1) gives the following non-closed equations for $\bar{u}^{\delta_{1}}, \bar{B}^{\delta_{2}}, \bar{p}^{\delta_{1}}$ in $(0, T) \times \Omega$ :

$$
\begin{align*}
& \bar{u}_{t}^{\delta_{1}}+\nabla \cdot\left(\overline{u u}^{\delta_{1}}\right)-\frac{1}{\operatorname{Re}} \Delta \bar{u}^{\delta_{1}}-S \nabla \cdot\left(\overline{B B^{\delta_{1}}}\right)+\nabla\left(\frac{S}{2}{\overline{B^{2}}}^{\delta_{1}}+\bar{p}^{\delta_{1}}\right)=\bar{f}^{\delta_{1}} \\
& \bar{B}_{t}^{\delta_{2}}+\frac{1}{\operatorname{Re}_{\mathrm{m}}} \operatorname{curl}\left(\operatorname{curl} \bar{B}^{\delta_{2}}\right)+\nabla \cdot\left(\overline{B u}^{\delta_{2}}\right)-\nabla \cdot\left(\overline{u B}{ }^{\delta_{2}}\right)=\operatorname{curl} \bar{g}^{\delta_{2}}  \tag{1.4}\\
& \nabla \cdot \bar{u}^{\delta_{1}}=0, \quad \nabla \cdot \bar{B}^{\delta_{2}}=0
\end{align*}
$$

The usual closure problem which we study here arises because $\overline{u u}^{\delta_{1}} \neq \bar{u}^{\delta_{1}} \bar{u}^{\delta_{1}}, \overline{B B}^{\delta_{1}} \neq$ $\bar{B}^{\delta_{1}} \bar{B}^{\delta_{1}}, \overline{u B}^{\delta_{2}} \neq \bar{u}^{\delta_{1}} \bar{B}^{\delta_{2}}$. To isolate the turbulence closure problem from the difficult problem of wall laws for near wall turbulence, we study (1.1) hence (1.4) subject to
(1.3). The closure problem is to replace the tensors $\overline{u u}^{\delta_{1}}, \overline{B B}^{\delta_{1}}, \overline{u B}^{\delta_{2}}$ with tensors $\mathscr{T}\left(\bar{u}^{\delta_{1}}, \bar{u}^{\delta_{1}}\right), \mathscr{T}\left(\bar{B}^{\delta_{2}}, \bar{B}^{\delta_{2}}\right), \mathscr{T}\left(\bar{u}^{\delta_{1}}, \bar{B}^{\delta_{2}}\right)$, respectively, depending only on $\bar{u}^{\delta_{1}}, \bar{B}^{\delta_{2}}$ and not $u, B$. There are many closure models proposed in large eddy simulation reflecting the centrality of closure in turbulence simulation. Calling $w, q, W$ the resulting approximations to $\bar{u}^{\delta_{1}}, \bar{p}^{\delta_{1}}, \bar{B}^{\delta_{2}}$, we are led to considering the following model

$$
\begin{aligned}
& w_{t}+\nabla \cdot \mathscr{T}(w, w)-\frac{1}{\operatorname{Re}} \Delta w-S \mathscr{T}(W, W)+\nabla q=\bar{f}^{\delta_{1}} \\
& W_{t}+\frac{1}{\operatorname{Re}_{\mathrm{m}}} \operatorname{curl}(\operatorname{curl} W)+\nabla \cdot \mathscr{T}(w, W)-\nabla \cdot \mathscr{T}(W, w)=\operatorname{curl} \bar{g}^{\delta_{2}} \\
& \nabla \cdot w=0, \quad \nabla \cdot W=0
\end{aligned}
$$

With any reasonable averaging operator, the true averages $\bar{u}^{\delta_{1}}, \bar{B}^{\delta_{2}}, \bar{p}^{\delta_{1}}$ are smoother than $u, B, p$. We consider the family of closure models, pioneered by Stolz and Adams [1]. These Approximate Deconvolution Models (ADM) use the deconvolution operators $G_{N}^{1}$ and $G_{N}^{2}$, that will be defined in Section 2. The ADM for the MHD reads

$$
\begin{align*}
& w_{t}+\nabla \cdot{\left.\overline{\left(G_{N}\right.}{ }_{N} w\right)\left(G_{N}^{1} w\right)}^{\delta_{1}}-\frac{1}{\operatorname{Re}} \Delta w-S \nabla \cdot{\left.\overline{\left(G_{N}\right.}{ }^{2} W\right)\left(G_{N}^{2} W\right)}^{\delta_{1}}+\nabla q=\bar{f}^{\delta_{1}}  \tag{1.5a}\\
& W_{t}+  \tag{1.5b}\\
& \quad \frac{1}{\operatorname{Re}_{\mathrm{m}}} \operatorname{curl}(\operatorname{curl} W)+\nabla \cdot\left(\overline{\left(G_{N}^{2} W\right)\left(G_{N}^{1} w\right)}{ }^{\delta_{2}}\right)-\nabla \cdot\left(\overline{\left(G_{N}^{1} w\right)\left(G_{N}^{2} W\right)}{ }^{\delta_{2}}\right) \\
& \quad=\operatorname{curl} \bar{g}^{\delta_{2}}  \tag{1.5c}\\
& \nabla \cdot w=0, \quad \nabla \cdot W=0
\end{align*}
$$

subject to $w(0, x)=\bar{u}_{0}^{\delta_{1}}(x), W(0, x)=\bar{B}_{0}^{\delta_{2}}(x)$ and periodic boundary conditions (with zero means).

We shall show that the ADM MHD model (1.5) has the mathematical properties expected of a model derived from the MHD equations by an averaging operation and which are important for practical computations. Note that $N=0$ in (1.5) leads to the model discussed in [20, 19].

The model considered can be developed for quite general averaging operators, see e.g. [1]. The choice of averaging operator in (1.5) is a differential filter, defined as follows. Let the $\delta>0$ denote the averaging radius, related to the finest computationally feasible mesh. (In this report we use different lengthscales for the Navier-Stokes and Maxwell equations). Given $\phi \in L_{0}^{2}(\Omega), \bar{\phi}^{\delta} \in H^{2}(\Omega) \cap L_{0}^{2}(\Omega)$ is the unique solution of

$$
\begin{equation*}
A_{\delta} \bar{\phi}^{\delta}:=-\delta^{2} \Delta \bar{\phi}^{\delta}+\bar{\phi}^{\delta}=\phi \quad \text { in } \Omega \tag{1.6}
\end{equation*}
$$

subject to periodic boundary conditions. Under periodic boundary conditions, this averaging operator commutes with differentiation, and with this averaging operator, the model (1.5) has consistency $O\left(\delta^{2 N+2}\right)$, i.e.,

$$
\begin{aligned}
& \overline{u u}^{\delta_{1}}={\overline{G_{N}^{1}} \bar{u}^{\delta_{1}} G_{N}^{1} \bar{u}^{\delta_{1}}}^{\delta_{1}}+O\left(\delta_{1}{ }^{2 N+2}\right), \\
& \overline{B B}^{\delta_{1}}={\overline{G_{N}} \bar{B}^{\delta_{2}} G_{N}^{2} \bar{B}^{\delta_{2}}}^{\delta_{1}}+O\left(\delta_{2}^{2 N+2}\right), \\
& \overline{u B}^{\delta_{2}}={\overline{G_{N}} \bar{u}^{\delta_{1}} G_{N}^{2} \bar{B}^{\delta_{2}}}^{\delta_{2}}+O\left(\delta_{1}^{2 N+2}+\delta_{2}{ }^{2 N+2}\right),
\end{aligned}
$$

for smooth $u, B$. We prove that the model (1.5) has a unique, strong solution $w, W$ that converges in the appropriate sense $w \rightarrow u, W \rightarrow B$, as $\delta_{1}, \delta_{2} \rightarrow 0$.

In Section 2 we address the global existence and uniqueness of the solution for the closed MHD model. Section 3 treats the questions of limit consistency of the model and verifiability. The conservation of the kinetic energy and helicity for the approximate deconvolution model is presented in Section 4. Section 5 shows that the model preserves the Alfén waves, with the velocity tending to the velocity of Alfvén waves in the MHD, as the radii $\delta_{1}, \delta_{2}$ tend to zero. The computational results in Section 6 confirm the accuracy and the physical fidelity of the models.
2. Existence and uniqueness for the ADM MHD equations. Introduce the family of the approximate deconvolution operators $G_{N}^{1}, G_{N}^{2}$, that are used in the ADM models (1.5).

Definition 2.1 (Approximate Deconvolution Operator). For a fixed finite $N$, define the $N$ th approximate deconvolution operators $G_{N}^{1}$ and $G_{N}^{2}$ by

$$
G_{N}^{i} \phi=\sum_{n=0}^{N}\left(I-A_{\delta_{i}}^{-1}\right)^{n} \phi, \text { for } i=1,2 .
$$

Note that since the differential filter $A_{\delta_{i}}$ is self adjoint, $G_{N}^{i}$ is also. $G_{N}^{i}$ was shown to be an $O\left(\delta_{i}^{2 N+2}\right)$ approximate inverse to the filter operator $A_{\delta_{i}}^{-1}$ (see [8]). Finally, it is easy to show that since $A_{\delta_{i}}$ commutes with differentiation, so does $G_{N}^{i}$.

LEMMA 2.2. The operator $G_{N}^{i}$ is compact, positive, and is an asymptotic inverse to the filter $A_{\delta_{i}}^{-1}$, i.e., for very smooth $\phi$ and as $\delta_{i} \rightarrow 0$ satisfies

$$
\begin{align*}
& \phi=G_{N}^{1} \bar{\phi}^{\delta_{1}}+(-1)^{N+1} \delta_{1}^{2 N+2} \Delta^{N+1} A_{\delta_{1}}^{-(N+1)} \phi,  \tag{2.1}\\
& \phi=G_{N}^{2} \bar{\phi}^{\delta_{2}}+(-1)^{N+1} \delta_{2}^{2 N+2} \Delta^{N+1} A_{\delta_{2}}^{-(N+1)} \phi .
\end{align*}
$$

The proof of Lemma 2.2 can be found in [8].
Lemma 2.3. $\|\cdot\|_{G_{N}^{i}}$ defined by $\|v\|_{G_{N}^{i}}=\left(v, G_{N}^{i} v\right)$ is a norm on $\Omega$, equivalent to the $L^{2}(\Omega)$ norm, and $(\cdot, \cdot)_{G_{N}^{i}}$ defined by $(v, w)_{G_{N}^{i}}=\left(v, G_{N}^{i} w\right)$ is an inner product on $\Omega$.

For the proof see [5].
We shall use the standard notations for function spaces in the space periodic case (see [35]). Let $H_{p}^{m}(\Omega)$ denote the space of functions (and their vector valued counterparts also) that are locally in $H^{m}\left(\mathbb{R}^{3}\right)$, are periodic of period $L$ and have zero mean, i.e. satisfy (1.3). We recall the solenoidal spaces

$$
\begin{aligned}
& H=\left\{\phi \in H_{2}^{0}(\Omega), \nabla \cdot \phi=0 \text { in } \mathscr{D}(\Omega)^{\prime}\right\}^{2} \\
& V=\left\{\phi \in H_{2}^{1}(\Omega), \nabla \cdot \phi=0 \text { in } \mathscr{D}(\Omega)^{\prime}\right\}^{2}
\end{aligned}
$$

We define the operator $\mathscr{A} \in \mathscr{L}\left(V, V^{\prime}\right)$ by setting

$$
\begin{equation*}
\left\langle\mathscr{A}\left(w_{1}, W_{1}\right),\left(w_{2}, W_{2}\right)\right\rangle=\int_{\Omega}\left(\frac{1}{\operatorname{Re}} \nabla w_{1} \cdot \nabla w_{2}+\frac{1}{\operatorname{Re}_{\mathrm{m}}} \operatorname{curl} W_{1} \operatorname{curl} W_{2}\right) d x \tag{2.2}
\end{equation*}
$$

for all $\left(w_{i}, W_{i}\right) \in V$. The operator $\mathscr{A}$ is an unbounded operator on $H$, with the domain $D(\mathscr{A})=\{(w, W) \in V ;(\Delta w, \Delta W) \in H\}$ and we denote again by $\mathscr{A}$ its restriction to $H$.

We define also a continuous tri-linear form $\mathscr{B}_{0}$ on $V \times V \times V$ by setting

$$
\begin{align*}
& \mathscr{B}_{0}\left(\left(w_{1}, W_{1}\right),\left(w_{2}, W_{2}\right),\left(w_{3}, W_{3}\right)\right)=\int_{\Omega}\left(\nabla \cdot{\left.\overline{\left(G_{N}\right.} w_{2}\right)\left(G_{N}^{1} w_{1}\right)}^{\delta_{1}} w_{3}\right.  \tag{2.3}\\
& -S \nabla \cdot{\overline{\left(G_{N}^{2} W_{2}\right)\left(G_{N}^{2} W_{1}\right)}}^{\delta_{1}} w_{3}+\nabla \cdot{\left.\overline{\left(G_{N}^{2} W_{2}\right)\left(G_{N}^{1} w_{1}\right)}{ }^{\delta_{2}} W_{3}-\nabla \cdot{\overline{\left(G_{N}\right.} 1}_{1} w_{2}\right)\left(G_{N}^{2} W_{1}\right)}^{\delta_{2}} W_{3} d x
\end{align*}
$$

and a continuous bilinear operator $\mathscr{B}(\cdot): V \rightarrow V$ with

$$
\left\langle\mathscr{B}\left(w_{1}, W_{1}\right),\left(w_{2}, W_{2}\right)\right\rangle=\mathscr{B}_{0}\left(\left(w_{1}, W_{1}\right),\left(w_{1}, W_{1}\right),\left(w_{2}, W_{2}\right)\right)
$$

for all $\left(w_{i}, W_{i}\right) \in V$.
The following properties of the trilinear form $\mathscr{B}_{0}$ hold (see [23, 31, 12, 9])

$$
\begin{align*}
& \mathscr{B}_{0}\left(\left(w_{1}, W_{1}\right),\left(w_{2}, W_{2}\right),\left(A_{\delta_{1}} G_{N}^{1} w_{2}, S A_{\delta_{2}} G_{N}^{2} W_{2}\right)\right)=0 \\
& \mathscr{B}_{0}\left(\left(w_{1}, W_{1}\right),\left(w_{2}, W_{2}\right),\left(A_{\delta_{1}} G_{N}^{1} w_{3}, S A_{\delta_{2}} G_{N}^{2} W_{3}\right)\right)  \tag{2.4}\\
& \quad=-\mathscr{B}_{0}\left(\left(w_{1}, W_{1}\right),\left(w_{3}, W_{3}\right),\left(A_{\delta_{1}} G_{N}^{1} w_{2}, S A_{\delta_{2}} G_{N}^{2} W_{2}\right)\right),
\end{align*}
$$

for all $\left(w_{i}, W_{i}\right) \in V$. Also

$$
\begin{align*}
& \left|\mathscr{B}_{0}\left(\left(w_{1}, W_{1}\right),\left(w_{2}, W_{2}\right),\left(w_{3}, W_{3}\right)\right)\right|  \tag{2.5}\\
& \quad \leq C\left\|\left(G_{N}^{1} w_{1}, G_{N}^{2} W_{1}\right)\right\|_{m_{1}}\left\|\left(G_{N}^{1} w_{2}, G_{N}^{2} W_{2}\right)\right\|_{m_{2}+1}\left\|\left({\overline{w_{3}}}^{\delta_{1}},{\overline{W_{3}}}^{\delta_{2}}\right)\right\|_{m_{3}}
\end{align*}
$$

for all $\left(w_{1}, W_{1}\right) \in H^{m_{1}}(\Omega),\left(w_{2}, W_{2}\right) \in H^{m_{2}+1}(\Omega),\left(w_{3}, W_{3}\right) \in H^{m_{3}}(\Omega)$ and

$$
\begin{array}{ll}
m_{1}+m_{2}+m_{3} \geq \frac{d}{2}, \quad \text { if } m_{i} \neq \frac{d}{2} \text { for all } i=1, \ldots, d, \\
m_{1}+m_{2}+m_{3}>\frac{d}{2}, \quad \text { if } m_{i}=\frac{d}{2} \text { for any of } i=1, \ldots, d
\end{array}
$$

In terms of $V, H, \mathscr{A}, \mathscr{B}(\cdot)$ we can rewrite (1.5) as

$$
\begin{align*}
& \frac{d}{d t}(w, W)+\mathscr{A}(w, W)(t)+\mathscr{B}((w, W)(t))=\left(\overline{\boldsymbol{f}}^{\delta_{1}}, \operatorname{curl} \overline{\boldsymbol{g}}^{\delta_{2}}\right), t \in(0, T)  \tag{2.6}\\
& (w, W)(0)=\left(\bar{u}_{0}^{\delta_{1}}, \bar{B}_{0}^{\delta_{2}}\right)
\end{align*}
$$

where $(\boldsymbol{f}, \operatorname{curl} \boldsymbol{g})=P(f, \operatorname{curl} g)$, and $P: L^{2}(\Omega) \rightarrow H$ is the Hodge projection.
ThEOREM 2.4. For any $\left(\bar{u}_{0}^{\delta_{1}}, \bar{B}_{0}{ }^{\delta_{2}}\right) \in V$ and $\left(\bar{f}^{\delta_{1}}, \operatorname{curl} \bar{g}^{\delta_{2}}\right) \in L^{2}(0, T ; H)$ there exists a unique strong solution to (1.5) $(w, W) \in L^{\infty}\left(0, T ; H^{1}(\Omega)\right) \cap L^{2}\left(0, T ; H^{2}(\Omega)\right)$ and $w_{t}, W_{t} \in L^{2}((0, T) \times \Omega)$. Moreover, the following energy equality holds:

$$
\begin{equation*}
\mathscr{E}(t)+\int_{0}^{t} \varepsilon(\tau) d \tau=\mathscr{E}(0)+\int_{0}^{t} \mathscr{P}(\tau) d \tau, \quad t \in[0, T] \tag{2.7}
\end{equation*}
$$

where
$\mathscr{E}(t)=\frac{\delta_{1}^{2}}{2}\|\nabla w(t, \cdot)\|_{G_{N}^{1}}^{2}+\frac{1}{2}\|w(t, \cdot)\|_{G_{N}^{1}}^{2}+\frac{\delta_{2}^{2} S}{2}\|\nabla W(t, \cdot)\|_{G_{N}^{2}}^{2}+\frac{S}{2}\|W(t, \cdot)\|_{G_{N}^{2}}^{2}$,
$\varepsilon(t)=\frac{\delta_{1}^{2}}{\operatorname{Re}}\|\Delta w(t, \cdot)\|_{G_{N}^{1}}^{2}+\frac{1}{\operatorname{Re}}\|\nabla w(t, \cdot)\|_{G_{N}^{1}}^{2}+\frac{\delta_{2}{ }^{2} S}{\operatorname{Re}_{\mathrm{m}}}\|\Delta W(t, \cdot)\|_{G_{N}^{2}}^{2}+\frac{S}{\operatorname{Re}_{\mathrm{m}}}\|\nabla W(t, \cdot)\|_{G_{N}^{2}}^{2}$,
$\mathscr{P}(t)=\left(f(t), G_{N}^{1} w(t)\right)+S\left(\operatorname{curlg}(t), G_{N}^{2} W(t)\right)$.

Proof. (Sketch) The proof follows from [20], using a semigroup approach and the machinery of nonlinear differential equations of accretive type in Banach spaces. The key to the model, as in MHD, is to make the nonlinear terms to vanish by an appropriate choice of test function. We observe that by (2.4)

$$
\mathscr{B}_{0}\left((w, W),(w, W),\left(A_{\delta_{1}} G_{N}^{1} w, S A_{\delta_{2}} G_{N}^{2} W\right)\right)=0
$$

thus taking the inner product of (2.6) with $\left(A_{\delta_{1}} G_{N}^{1} w, S A_{\delta_{2}} G_{N}^{2} W\right)$ and integrating by parts we get

$$
\begin{aligned}
& \frac{1}{2} \frac{d}{d t}\left(\|w\|_{G_{N}^{1}}^{2}+\delta_{1}^{2}\|\nabla w\|_{G_{N}^{1}}^{2}+S\|W\|_{G_{N}^{2}}^{2}+\delta_{2}^{2} S\|\nabla W\|_{G_{N}^{2}}^{2}\right) \\
& \quad+\frac{1}{\operatorname{Re}}\left(\|\nabla w\|_{G_{N}^{1}}^{2}+\delta_{1}^{2}\|\Delta w\|_{G_{N}^{1}}^{2}\right)+\frac{S}{\operatorname{Re}_{\mathrm{m}}}\left(\|\nabla W\|_{G_{N}^{2}}^{2}+\delta_{2}^{2} S\|\Delta W\|_{G_{N}^{2}}^{2}\right) \\
& =\left(f, G_{N}^{1} w\right)+S\left(\operatorname{curl} g, G_{N}^{2} W\right)
\end{aligned}
$$

The pressure is recovered from the weak solution via the classical DeRham theorem (see [22]).

Theorem 2.5. Let $m \in \mathbb{N},\left(u_{0}, B_{0}\right) \in V \cap H^{m-1}(\Omega)$ and $(f$, curlg $) \in L^{2}(0, T$; $\left.H^{m-1}(\Omega)\right)$. Then there exists a unique solution $w, W, q$ to the equation (1.5) such that

$$
\begin{aligned}
& (w, W) \in L^{\infty}\left(0, T ; H^{m+1}(\Omega)\right) \cap L^{2}\left(0, T ; H^{m+2}(\Omega)\right) \\
& q \in L^{2}\left(0, T ; H^{m}(\Omega)\right)
\end{aligned}
$$

Proof. The result is already proved when $m=0$ in Theorem 2.4. For any $m \in \mathbb{N}^{*}$, we assume that

$$
\begin{equation*}
(w, W) \in L^{\infty}\left(0, T ; H^{m}(\Omega)\right) \cap L^{2}\left(0, T ; H^{m+1}(\Omega)\right) \tag{2.9}
\end{equation*}
$$

so it remains to prove

$$
\left(D^{m} w, D^{m} W\right) \in L^{\infty}\left(0, T ; H^{1}(\Omega)\right) \cap L^{2}\left(0, T ; H^{2}(\Omega)\right)
$$

where $D^{m}$ denotes any partial derivative of total order $m$. We take the $m^{\text {th }}$ derivative of (1.5) and have

$$
\begin{aligned}
& D^{m} w_{t}-\frac{1}{\operatorname{Re}} \Delta D^{m} w+D^{m} \overline{\left(G_{N}^{1} w \cdot \nabla G_{N}^{1} w\right)}{ }^{\delta_{1}}-S D^{m} \overline{\left(G_{N}^{2} W \cdot \nabla G_{N}^{2} W\right)}{ }^{\delta_{1}}+\nabla D^{m} q=D^{m} \bar{f}^{\delta_{1}}, \\
& D^{m} W_{t}+\frac{1}{\operatorname{Re}_{\mathrm{m}}} \nabla \times \nabla \times D^{m} W+D^{m}{\overline{\left(G_{N}^{1} w \cdot \nabla G_{N}^{2} W\right)}}^{\delta_{2}}-D^{m}{\overline{\left(G_{N}^{2} W \cdot \nabla G_{N}^{1} w\right)}}^{\delta_{2}} \\
& \quad=\nabla \times D^{m} \bar{g}^{\delta_{2}}, \\
& \nabla \cdot D^{m} w=0, \nabla \cdot D^{m} W=0, \\
& D^{m} w(0, \cdot)=D^{m}{\overline{u_{0}}}^{\delta_{1}}, D^{m} W(0, \cdot)=D^{m}{\overline{B_{0}}}^{\delta_{2}},
\end{aligned}
$$

with periodic boundary conditions and zero mean, and the initial conditions with zero
divergence and mean. Taking $A_{\delta_{1}} D^{m} w, A_{\delta_{2}} D^{m} W$ as test functions we obtain

$$
\begin{align*}
& \frac{1}{2} \frac{d}{d t}\left(\left\|D^{m} w\right\|_{0}^{2}+\delta_{1}^{2}\left\|\nabla D^{m} w\right\|_{0}^{2}+S\left\|D^{m} W\right\|_{0}^{2}+S \delta_{2}^{2}\left\|\nabla D^{m} W\right\|_{0}^{2}\right)  \tag{2.10}\\
& \quad+\frac{1}{\operatorname{Re}}\left(\left\|\nabla D^{m} w\right\|_{0}^{2}+\delta_{1}^{2}\left\|\Delta D^{m} w\right\|_{0}^{2}\right)+\frac{1}{\operatorname{Re}_{\mathrm{m}}}\left(\left\|\nabla D^{m} W\right\|_{0}^{2}+\delta_{2}^{2}\left\|\Delta D^{m} W\right\|_{0}^{2}\right) \\
& =\int_{\Omega}\left(D^{m} f D^{m} w+\nabla \times D^{m} g D^{m} W\right) d x-\mathscr{X}
\end{align*}
$$

where

$$
\begin{aligned}
& \mathscr{X}=\int_{\Omega}\left(D^{m}\left(G_{N}^{1} w \cdot \nabla G_{N}^{1} w\right)-S D^{m}\left(G_{N}^{2} W \cdot \nabla G_{N}^{2} W\right)\right) D^{m} w \\
&+\left(D^{m}\left(G_{N}^{1} w \cdot \nabla G_{N}^{2} W\right)-D^{m}\left(G_{N}^{2} W \cdot \nabla G_{N}^{1} w\right)\right) D^{m} W d x
\end{aligned}
$$

Now we apply (2.5) and use the induction assumption (2.9)

$$
\begin{aligned}
\mathscr{X}= & \sum_{|\alpha| \leq m}\binom{m}{\alpha} \sum_{i, j=1}^{3} \int_{\Omega}\left(D^{\alpha} G_{N}^{1} w_{i} D^{m-\alpha} D_{i} G_{N}^{1} w_{j}-S D^{\alpha} G_{N}^{2} W_{i} D^{m-\alpha} D_{i} G_{N}^{2} W_{j}\right) D^{m} w_{j} \\
& \quad+\left(D^{\alpha} G_{N}^{1} w_{i} D^{m-\alpha} D_{i} G_{N}^{2} W_{j}-D^{\alpha} G_{N}^{2} W_{i} D^{m-\alpha} D_{i} G_{N}^{1} w_{j}\right) D^{m} W_{j} d x \\
\leq & C(m)\left(\left\|G_{N}^{1} w\right\|_{m}^{3 / 2}\left\|G_{N}^{1} w\right\|_{m+1}^{1 / 2}+\left\|G_{N}^{2} W\right\|_{m}^{3 / 2}\left\|G_{N}^{2} W\right\|_{m+1}^{1 / 2}\right)\|w\|_{m} \\
& +\left(\left\|G_{N}^{1} w\right\|_{m}\left\|G_{N}^{2} W\right\|_{m}^{1 / 2}\left\|G_{N}^{2} W\right\|_{m+1}^{1 / 2}+\left\|G_{N}^{2} W\right\|_{m}\left\|G_{N}^{1} w\right\|_{m}^{1 / 2}\left\|G_{N}^{1} w\right\|_{m+1}^{1 / 2}\right)\|W\|_{m}
\end{aligned}
$$

Integrating (2.10) on $(0, T)$, using the Cauchy-Schwarz and Hölder inequalities, Lemma 2.2, 2.3 and the assumption (2.9) we obtain the desired result for $w, W$. We conclude the proof mentioning that the regularity of the pressure term $q$ is obtained via classical methods, see e.g. [34, 3].

## 3. Accuracy of the model.

We address first the question of consistency, i.e., we show that the solution of the closed model (1.5) converges to a solution of the MHD equations (1.1) when $\delta_{1}, \delta_{2}$ tend zero.
Let $\tau_{u}, \tau_{B}, \tau_{B u}$ denote

$$
\begin{equation*}
\tau_{u}=G_{N}^{1} \bar{u}^{\delta_{1}} G_{N}^{1} \bar{u}^{\delta_{1}}-u u, \tau_{B}=G_{N}^{2} \bar{B}^{\delta_{2}} G_{N}^{2} \bar{B}^{\delta_{2}}-B B, \tau_{B u}=G_{N}^{2} \bar{B}^{\delta_{2}} G_{N}^{1} \bar{u}^{\delta_{1}}-B u \tag{3.1}
\end{equation*}
$$

where $u, B$ is a solution of the MHD equations obtained as a limit of a subsequence of the sequence $w_{\delta_{1}}, W_{\delta_{2}}$.

We prove in Theorem 3.2 that the model's consistency errors $\left\|\bar{u}^{\delta_{1}}-w\right\|_{L^{\infty}\left(0, T ; L^{2}(Q)\right)}$, $\left\|\bar{B}^{\delta_{2}}-W\right\|_{L^{\infty}\left(0, T ; L^{2}(Q)\right)}$ are bounded by $\left\|\tau_{u}\right\|_{L^{2}\left(Q_{T}\right)},\left\|\tau_{B}\right\|_{L^{2}\left(Q_{T}\right)},\left\|\tau_{B u}\right\|_{L^{2}\left(Q_{T}\right)}$.

### 3.1. Limit consistency of the model.

Theorem 3.1. There exist two sequences $\delta_{1}^{n}, \delta_{2}^{n} \rightarrow 0$ as $n \rightarrow 0$ such that

$$
\left(w_{\delta_{1}^{n}}, W_{\delta_{2}^{n}}, q_{\delta_{1}^{n}}\right) \rightarrow(u, B, p) \quad \text { as } \delta_{1}^{n}, \delta_{2}^{n} \rightarrow 0
$$

where $(u, B, p) \in L^{\infty}(0, T ; H) \cap L^{2}(0, T ; V) \times L^{\frac{4}{3}}\left(0, T ; L^{2}(\Omega)\right)$ is a solution of the MHD equations (1.1). The sequences $\left\{w_{\delta_{1}^{n}}\right\}_{n \in \mathbb{N}},\left\{W_{\delta_{2}^{n}}\right\}_{n \in \mathbb{N}}$ converge strongly to $u, B$ in $L^{\frac{4}{3}}\left(0, T ; L^{2}(\Omega)\right)$ and weakly in $L^{2}\left(0, T ; H^{1}(\Omega)\right)$, respectively, while $\left\{q_{\delta_{1}^{n}}\right\}_{n \in \mathbb{N}}$ converges weakly to $p$ in $L^{\frac{4}{3}}\left(0, T ; L^{2}(\Omega)\right)$.

Proof. The proof follows that of Theorem 3.1 in [20], and is an easy consequence of Theorem 3.2 and Proposition 3.3.

### 3.2. Verifiability of the model.

Theorem 3.2. Suppose that the true solution of (1.1) satisfies the regularity condition $(u, B) \in L^{4}(0, T ; V)$. Then the consistency errors $e=\bar{u}^{\delta_{1}}-w, E=\bar{B}^{\delta_{2}}-W$ satisfy

$$
\begin{align*}
& \|e(t)\|_{0}^{2}+S\|E(t)\|_{0}^{2}+\int_{0}^{t}\left(\frac{1}{\operatorname{Re}}\|\nabla e(s)\|_{0}^{2}+\frac{S}{\operatorname{Re}_{\mathrm{m}}} \| \operatorname{curl}\left(E(s) \|_{0}^{2}\right) d s\right. \\
& \quad \leq C \Phi(t) \int_{0}^{t}\left(\operatorname{Re}\left\|\tau_{u}(s)+S \tau_{B}(s)\right\|_{0}^{2}+\operatorname{Re}_{\mathrm{m}}\left\|\tau_{B u}(s)-\tau_{B u}(s)\right\|_{0}^{2}\right) d s, \tag{3.2}
\end{align*}
$$

where $\Phi(t)=\exp \left\{\operatorname{Re}^{3} \int_{0}^{t}\|\nabla u\|_{0}^{4} d s, \operatorname{Re}_{\mathrm{m}}^{3} \int_{0}^{t}\|\nabla u\|_{0}^{4} d s+\operatorname{Re}_{\mathrm{m}} \operatorname{Re}^{2} \int_{0}^{t}\|\nabla B\|_{0}^{4}\right\}$.
Proof. The errors $e=\bar{u}^{\delta_{1}}-w, E=\bar{B}^{\delta_{2}}-W$ satisfy in variational sense

$$
\begin{aligned}
& e_{t}+ \nabla \cdot\left(\overline{G_{N}^{1} \bar{u}^{\delta_{1}} G_{N}^{1} \bar{u}^{\delta_{1}}-G_{N}^{1} w G_{N}^{1} w}{ }^{\delta_{1}}\right)-\frac{1}{\operatorname{Re}} \Delta e+S \nabla \cdot\left(\overline{G_{N}^{2} \bar{B}^{\delta_{2}} G_{N}^{2} \bar{B}^{\delta_{2}}-G_{N}^{2} W G_{N}^{2} W}{ }^{\delta_{1}}\right) \\
&+\nabla\left(\bar{p}^{\delta_{1}}-q\right)=\nabla \cdot\left(\bar{\tau}_{u}^{\delta_{1}}+S \bar{\tau}_{B}^{\delta_{1}}\right), \\
& E_{t}+ \frac{1}{\operatorname{Re}_{\mathrm{m}}} \nabla \times \nabla \times E+\nabla \cdot{\overline{G_{N}^{2}} \bar{B}^{\delta_{2}} G_{N}^{1} \bar{u}^{\delta_{1}}-G_{N}^{2} W G_{N}^{1} w}^{\delta_{2}}-\nabla \cdot \overline{G_{N}^{1} \bar{u}^{\delta_{1}} G_{N}^{2} \bar{B}^{\delta_{2}}-G_{N}^{1} w G_{N}^{2} W} \\
& \quad \delta_{2} \\
& \quad=\nabla \cdot\left(\bar{\tau}_{B u}^{\delta_{2}}-\bar{\tau}_{u B}^{\delta_{2}}\right),
\end{aligned}
$$

and $\nabla \cdot e=\nabla \cdot E=0, e(0)=E(0)=0$. Taking the inner product with $\left(A_{\delta_{1}} G_{N}^{1} e\right.$, $\left.S A_{\delta_{2}} G_{N}^{2} E\right)$ we get as for (2.7) the following

$$
\begin{aligned}
& \frac{1}{2} \frac{d}{d t}\left(\|e\|_{G_{N}^{1}}^{2}+S\|E\|_{G_{N}^{2}}^{2}+\delta_{1}^{2}\|\nabla e\|_{G_{N}^{1}}^{2}+\delta_{2}^{2} S\|\operatorname{curl} E\|_{G_{N}^{2}}^{2}\right) \\
& +\frac{1}{\operatorname{Re}}\|\nabla e\|_{G_{N}^{1}}^{2}+\frac{S}{\operatorname{Re}_{\mathrm{m}}}\|\operatorname{curl} \mathrm{E}\|_{G_{N}^{2}}^{2}+\frac{\delta_{1}^{2}}{\operatorname{Re}}\|\Delta e\|_{G_{N}^{1}}^{2}+\frac{\delta_{2}^{2} S}{\operatorname{Re}_{\mathrm{m}}}\|\operatorname{curl} \operatorname{curl} E\|_{G_{N}^{2}}^{2} \\
& +\int_{\Omega}\left(\nabla \cdot\left(G_{N}^{1} \bar{u}^{\delta_{1}} G_{N}^{1} \bar{u}^{\delta_{1}}-G_{N}^{1} w G_{N}^{1} w\right) G_{N}^{1} e+S \nabla \cdot\left(G_{N}^{2} \bar{B}^{\delta_{2}} G_{N}^{2} \bar{B}^{\delta_{2}}-G_{N}^{2} W G_{N}^{2} W\right) G_{N}^{1} e\right. \\
& \left.+S \nabla \cdot\left(G_{N}^{2} \bar{B}^{\delta_{2}} G_{N}^{1} \bar{u}^{\delta_{1}}-G_{N}^{2} W G_{N}^{1} w\right) G_{N}^{2} E-S \nabla \cdot\left(G_{N}^{1} \bar{u}^{\delta_{1}} G_{N}^{2} \bar{B}^{\delta_{2}}-G_{N}^{1} w G_{N}^{2} W\right) G_{N}^{2} E\right) d x \\
& =-\int_{\Omega}\left(\left(\tau_{u}+S \tau_{B}\right) \cdot \nabla G_{N}^{1} e+S\left(\tau_{B u}-\tau_{u B}\right) \cdot \nabla G_{N}^{2} E\right) d x \\
& \leq \frac{1}{2 \operatorname{Re}}\|\nabla e\|_{0}^{2}+\frac{S}{2 \operatorname{Re}_{\mathrm{m}}}\|\operatorname{curl} E\|_{0}^{2}+\frac{\operatorname{Re}}{2}\left\|\tau_{u}+S \tau_{B}\right\|_{0}^{2}+\frac{\operatorname{Re}_{\mathrm{m}}}{2 S}\left\|\tau_{B u}-\tau_{u B}\right\|_{0}^{2} .
\end{aligned}
$$

Using the identity $G_{N}^{1} \bar{u}^{\delta_{1}} G_{N}^{1} \bar{u}^{\delta_{1}}-G_{N}^{1} w G_{N}^{1} w=G_{N}^{1} e G_{N}^{1} \bar{u}^{\delta_{1}}+G_{N}^{1} w G_{N}^{1} e$, Lemmas 2.2, 2.3, the divergence free condition and (2.5) we have

$$
\begin{aligned}
& \frac{d}{d t}\left(\|e\|_{0}^{2}+S\|E\|_{0}^{2}+\delta_{1}^{2}\|\nabla e\|_{0}^{2}+S \delta_{2}^{2}\|\operatorname{curl} E\|_{0}^{2}\right) \\
& \quad+\frac{1}{\operatorname{Re}}\|\nabla e\|_{0}^{2}+\frac{S}{\operatorname{Re}_{\mathrm{m}}}\|\operatorname{curl} \mathrm{E}\|_{0}^{2}+\frac{\delta_{1}^{2}}{\operatorname{Re}}\|\Delta e\|_{0}^{2}+\frac{\delta_{2}^{2} S}{\operatorname{Re}_{\mathrm{m}}}\|\operatorname{curl} \operatorname{curl} E\|_{0}^{2} \\
& \leq \int_{\Omega}\left(-G_{N}^{1} e \cdot \nabla G_{N}^{1} \bar{u}^{\delta_{1}} G_{N}^{1} e-S \nabla \cdot\left(G_{N}^{2} E G_{N}^{2} \bar{B}^{\delta_{2}}\right) G_{N}^{1} e-S \nabla \cdot\left(G_{N}^{2} E G_{N}^{1} \bar{u}^{\delta_{1}}\right) G_{N}^{2} E\right. \\
& \left.\quad+S G_{N}^{1} e \cdot \nabla G_{N}^{2} \bar{B}^{\delta_{2}} G_{N}^{2} E\right) d x+\operatorname{Re}\left\|\tau_{u}+S \tau_{B}\right\|_{0}^{2}+\operatorname{Re}_{\mathrm{m}}\left\|\tau_{B u}-\tau_{u B}\right\|_{0}^{2} \\
& \leq C\left(\|\nabla e\|_{0}^{3 / 2}\|e\|_{0}^{1 / 2}\left\|\nabla \bar{u}^{\delta_{1}}\right\|_{0}+2 S\|E\|_{0}^{1 / 2}\|\nabla E\|_{0}^{1 / 2}\left\|\nabla \bar{B}^{\delta_{2}}\right\|_{0}\|\nabla e\|_{0}\right. \\
& \left.\quad+S\|E\|_{0}^{1 / 2}\|\nabla E\|_{0}^{3 / 2}\left\|\nabla \bar{u}^{\delta_{1}}\right\|_{0}\right)+\operatorname{Re}\left\|\tau_{u}+S \tau_{B}\right\|_{0}^{2}+\operatorname{Re}_{\mathrm{m}}\left\|\tau_{B u}-\tau_{u B}\right\|_{0}^{2} .
\end{aligned}
$$

Using $a b \leq \varepsilon a^{4 / 3}+C \varepsilon^{-3} b^{4}$ we obtain

$$
\begin{aligned}
& \frac{d}{d t}\left(\|e\|_{0}^{2}+S\|E\|_{0}^{2}+\delta_{1}^{2}\|\nabla e\|_{0}^{2}+S \delta_{2}^{2}\|\operatorname{curl} E\|_{0}^{2}\right) \\
& \quad+\frac{1}{\operatorname{Re}}\|\nabla e\|_{0}^{2}+\frac{S}{\operatorname{Re}_{\mathrm{m}}}\|\operatorname{curl} \mathrm{E}\|_{0}^{2}+\frac{\delta_{1}^{2}}{\operatorname{Re}}\|\Delta e\|_{0}^{2}+\frac{\delta_{2}^{2} S}{\operatorname{Re}_{\mathrm{m}}}\|\operatorname{curl} \operatorname{curl} E\|_{0}^{2} \\
& \leq C\left(\operatorname{Re}^{3}\|e\|_{0}^{2}\left\|\nabla \bar{u}^{\delta_{1}}\right\|_{0}^{4}+\operatorname{Re}_{\mathrm{m}} \operatorname{Re}^{2}\|E\|_{0}^{2}\left\|\nabla \bar{B}^{\delta_{2}}\right\|_{0}^{4}+\operatorname{Re}_{\mathrm{m}}^{3}\|E\|_{0}^{2}\left\|\nabla \bar{u}^{\delta_{1}}\right\|_{0}^{4}\right) \\
& \quad+\operatorname{Re}\left\|\tau_{u}+S \tau_{B}\right\|_{0}^{2}+\operatorname{Re}_{\mathrm{m}}\left\|\tau_{B u}-\tau_{u B}\right\|_{0}^{2}
\end{aligned}
$$

and by the Gronwall inequality we deduce

$$
\begin{aligned}
& \|e(t)\|_{0}^{2}+S\|E(t)\|_{0}^{2}+\int_{0}^{t}\left(\frac{1}{\operatorname{Re}}\|\nabla e(s)\|_{0}^{2}+\frac{S}{\operatorname{Re}_{\mathrm{m}}}\|\operatorname{curl} E(s)\|_{0}^{2}\right) d s \\
& \leq C \Psi(t) \int_{0}^{t}\left(\operatorname{Re}\left\|\tau_{u}(s)+S \tau_{B}(s)\right\|_{0}^{2}+\operatorname{Re}_{\mathrm{m}}\left\|\tau_{B u}(s)-\tau_{u B}(s)\right\|_{0}^{2}\right) d s
\end{aligned}
$$

where

$$
\Psi(t)=\exp \left\{\operatorname{Re}^{3} \int_{0}^{t}\left\|\nabla \bar{u}^{\delta_{1}}\right\|_{0}^{4} d s, \operatorname{Re}_{\mathrm{m}}^{3} \int_{0}^{t}\left\|\nabla \bar{u}^{\delta_{1}}\right\|_{0}^{4} d s+\operatorname{Re}_{\mathrm{m}} \operatorname{Re}^{2} \int_{0}^{t}\left\|\nabla \bar{B}^{\delta_{2}}\right\|_{0}^{4} d s\right\}
$$

Using the stability bounds $\left\|\nabla \bar{u}^{\delta_{1}}\right\|_{0} \leq\|\nabla u\|_{0},\left\|\nabla \bar{B}^{\delta_{2}}\right\|_{0} \leq\|\nabla B\|_{0}$ we conclude the proof.
3.3. Consistency error estimate. The bounds on the errors (3.1) are given in the following proposition.

Proposition 3.3. Let

$$
(u, B) \in L^{4}((0, T) \times \Omega) \cap L^{4}\left(0, T ; H^{2 N+2}(\Omega)\right), N \geq 0
$$

Then

$$
\begin{aligned}
& \left\|\tau_{u}\right\|_{L^{2}(Q)} \leq C \delta_{1}^{2 N+2} \\
& \left\|\tau_{B}\right\|_{L^{2}(Q)} \leq C \delta_{2}^{2 N+2} \\
& \left\|\tau_{B u}\right\|_{L^{2}(Q)} \leq C\left(\delta_{1}^{2 N+2}+\delta_{2}^{2 N+2}\right)
\end{aligned}
$$

where $C=C\left(\|(u, B)\|_{L^{4}((0, T) \times \Omega)},\|(u, B)\|_{L^{4}\left(0, T ; H^{2 N+2}(\Omega)\right)}\right)$. The proof uses Lemma 2.2 and follows the outline of the proofs in Section 3.3 of [20].
4. Conservation laws. As our model is some sort of a regularizing numerical scheme, we would like to make sure that the model inherits some of the original properties of the 3D MHD equations.

It is well known that kinetic energy and helicity are critical in the organization of the flow.

The energy $E=\frac{1}{2} \int_{\Omega}(v(x) \cdot v(x)+B(x) \cdot B(x)) d x$, the cross helicity $H_{C}=$ $\frac{1}{2} \int_{\Omega}(v(x) \cdot B(x)) d x$ and the magnetic helicity $H_{M}=\frac{1}{2} \int_{\Omega}(\mathbb{A}(x) \cdot B(x)) d x$ (where $\mathbb{A}$ is the vector potential, $B=\nabla \times \mathbb{A}$ ) are the three invariants of the MHD equations (1.1) in the absence of kinematic viscosity and magnetic diffusivity ( $\frac{1}{\operatorname{Re}}=\frac{1}{\operatorname{Re}_{\mathrm{m}}}=0$ ).

Introduce the characteristic quantities of the model (1.5)

$$
\begin{aligned}
& E_{A D M}=\frac{1}{2}\left[\left(A_{\delta_{1}} w, w\right)_{G_{N}^{1}}+\left(A_{\delta_{2}} W, W\right)_{G_{N}^{2}}\right] \\
& H_{C, A D M}=\frac{1}{2}\left(A_{\delta_{1}} w, A_{\delta_{2}} W\right), \text { and } \\
& H_{M, A D M}=\frac{1}{2}\left(A_{\delta_{2}} W, \overline{\mathbb{A}}^{\delta_{2}}\right)_{G_{N}^{2}}, \text { where } \overline{\mathbb{A}}^{\delta_{2}}=A_{\delta_{2}}^{-1} \mathbb{A} .
\end{aligned}
$$

This section is devoted to proving that these quantities are conserved by (1.5) with the periodic boundary conditions and $\frac{1}{\mathrm{Re}}=\frac{1}{\mathrm{Re}_{\mathrm{m}}}=0$. Also, note that

$$
E_{A D M} \rightarrow E, H_{C, A D M} \rightarrow H_{C}, H_{M, A D M} \rightarrow H_{M}, \quad \text { as } \delta_{1,2} \rightarrow 0
$$

ThEOREM 4.1. The following conservation laws hold, $\forall T>0$

$$
\begin{align*}
& E_{A D M}(T)=E_{A D M}(0)  \tag{4.1}\\
& H_{C, A D M}(T)=H_{C, A D M}(0)+C(T) \max _{i=1,2} \delta_{i}^{2 N+2} \tag{4.2}
\end{align*}
$$

and

$$
\begin{equation*}
H_{M, A D M}(T)=H_{M, A D M}(0) \tag{4.3}
\end{equation*}
$$

REmARK 4.1. Note that the cross helicity $H_{C, A D M}$ of the model is not conserved exactly, but it possesses two important properties:

$$
H_{C, A D M} \rightarrow H_{C} \text { as } \delta_{1,2} \rightarrow 0
$$

and

$$
H_{C, A D M}(T) \rightarrow H_{C, A D M}(0) \text { as } N \text { increases. }
$$

In the case of equal radii, $\delta_{1}=\delta_{2}$, the following cross helicity is exactly conserved:

$$
H_{\times, A D M}(w, W)(t)=\frac{1}{2}\left((w, W)_{N}+\delta^{2}(\nabla w, \nabla W)_{N}\right)
$$

Proof. The proof follows the outline of the corresponding proof in [20]. Consider (1.5) with $\frac{1}{\operatorname{Re}}=\frac{1}{\operatorname{Re}_{\mathrm{m}}}=0$.

Start by proving (4.1). Multiply (1.5a) by $A_{\delta_{1}} G_{N}^{1} w$, and multiply (1.5b) by $A_{\delta_{2}} G_{N}^{2} W$. Integrating both equations over $\Omega$ gives

$$
\begin{align*}
& \frac{1}{2} \frac{d}{d t}\left(A_{\delta_{1}} w, w\right)_{G_{N}^{1}}=\left(\left(\nabla \times G_{N}^{2} W\right) \times G_{N}^{2} W, w\right)_{G_{N}^{1}}  \tag{4.4}\\
& \frac{1}{2} \frac{d}{d t}\left(A_{\delta_{2}} W, W\right)_{G_{N}^{2}}-\left(G_{N}^{2} W \cdot \nabla G_{N}^{1} w, W\right)_{G_{N}^{2}}=0 \tag{4.5}
\end{align*}
$$

Adding (4.4)-(4.5) and using the identity

$$
\begin{equation*}
((\nabla \times v) \times u, w)=(u \cdot \nabla v, w)-(w \cdot \nabla v, u) \tag{4.6}
\end{equation*}
$$

we obtain

$$
\begin{aligned}
& \frac{1}{2} \frac{d}{d t}\left[\left(A_{\delta_{1}} w, w\right)_{G_{N}^{1}}+\left(A_{\delta_{2}} W, W\right)_{G_{N}^{2}}\right] \\
& =\left(G_{N}^{2} W \cdot \nabla G_{N}^{2} W, G_{N}^{1} w\right)-\left(G_{N}^{1} w \cdot \nabla G_{N}^{2} W, G_{N}^{2} W\right)+\left(G_{N}^{2} W \cdot \nabla G_{N}^{1} w, G_{N}^{2} W\right)=0
\end{aligned}
$$

which yields (4.1).
To prove (4.2), multiply (1.5a)-(1.5b) by $A_{\delta_{1}} G_{N}^{2} W$ and $A_{\delta_{2}} G_{N}^{1} w$, respectively, and integrate over $\Omega$ to get

$$
\begin{align*}
& \left(\frac{\partial A_{\delta_{1}} w}{\partial t}, W\right)_{G_{N}^{2}}+\left(G_{N}^{1} w \cdot \nabla G_{N}^{1} w, W\right)_{G_{N}^{2}}=0  \tag{4.7}\\
& \left(\frac{\partial A_{\delta_{2}} W}{\partial t}, w\right)_{G_{N}^{1}}+\left(G_{N}^{1} w \cdot \nabla G_{N}^{2} W, w\right)_{G_{N}^{1}}=0 \tag{4.8}
\end{align*}
$$

Adding (4.7) and (4.8), we obtain

$$
\begin{equation*}
\left(\frac{\partial A_{\delta_{1}} w}{\partial t}, G_{N}^{2} W\right)+\left(\frac{\partial A_{\delta_{2}} W}{\partial t}, G_{N}^{1} w\right)=0 \tag{4.9}
\end{equation*}
$$

From Corollary 2.2 it follows that

$$
\begin{align*}
& G_{N}^{1} w=A_{\delta_{1}} w+(-1)^{N} \delta_{1}^{2 N+2} \Delta^{N+1} A_{\delta_{1}}^{-N} w  \tag{4.10}\\
& G_{N}^{2} W=A_{\delta_{2}} W+(-1)^{N} \delta_{2}^{2 N+2} \Delta^{N+1} A_{\delta_{2}}^{-N} W
\end{align*}
$$

Then (4.9) gives

$$
\begin{align*}
& \frac{d}{d t}\left(A_{\delta_{1}} w, A_{\delta_{2}} W\right)=\left(\frac{\partial A_{\delta_{1}} w}{\partial t}, A_{\delta_{2}} W\right)+\left(\frac{\partial A_{\delta_{2}} W}{\partial t}, A_{\delta_{1}} w\right)  \tag{4.11}\\
& =\left(\frac{\partial A_{\delta_{1}} w}{\partial t},(-1)^{N+1} \delta_{2}^{2 N+2} \Delta^{N+1} A_{\delta_{2}}^{-N} W\right)+\left(\frac{\partial A_{\delta_{2}} W}{\partial t},(-1)^{N+1} \delta_{1}^{2 N+2} \Delta^{N+1} A_{\delta_{1}}^{-N} w\right) \\
& =(-1)^{N+1} \delta_{2}^{2 N+2}\left(\frac{\partial A_{\delta_{1}} w}{\partial t}, \Delta^{N+1} A_{\delta_{2}}^{-N} W\right)+(-1)^{N+1} \delta_{1}^{2 N+2}\left(\frac{\partial A_{\delta_{2}} W}{\partial t}, \Delta^{N+1} A_{\delta_{1}}^{-N} w\right)
\end{align*}
$$

which proves (4.2).
Next, we prove (4.3). By multiplying (1.5b) by $A_{\delta_{2}} G_{N}^{2} \overline{\mathbb{A}}^{\delta_{2}}$, and integrating over $\Omega$ we get

$$
\begin{align*}
& \frac{1}{2} \frac{d}{d t}\left(\nabla \times A_{\delta_{2}} \overline{\mathbb{A}}^{\delta_{2}}, G_{N}^{2} \overline{\mathbb{A}}^{\delta_{2}}\right)  \tag{4.12}\\
& \quad+\left(G_{N}^{1} w \cdot \nabla G_{N}^{2} W, G_{N}^{2} \overline{\mathbb{A}}^{\delta_{2}}\right)-\left(G_{N}^{2} W \cdot \nabla G_{N}^{1} w, G_{N}^{2} \overline{\mathbb{A}}^{\delta_{2}}\right)=0
\end{align*}
$$

Since the cross-product of two vectors is orthogonal to each of them

$$
\begin{equation*}
\left(\left(\nabla \times G_{N}^{2} \overline{\mathbb{A}}^{\delta_{2}}\right) \times G_{N}^{1} w, \nabla \times G_{N}^{2} \overline{\mathbb{A}}^{\delta_{2}}\right)=0 \tag{4.13}
\end{equation*}
$$

it follows from (4.13) and (4.6) that

$$
\begin{equation*}
\left(G_{N}^{1} w \cdot \nabla G_{N}^{2} \overline{\mathbb{A}}^{\delta_{2}}, \nabla \times G_{N}^{2} \overline{\mathbb{A}}^{\delta_{2}}\right)=\left(\left(\nabla \times G_{N}^{2} \overline{\mathbb{A}}^{\delta_{2}}\right) \cdot \nabla G_{N}^{2} \overline{\mathbb{A}}^{\delta_{2}}, G_{N}^{1} w\right) \tag{4.14}
\end{equation*}
$$

Since $G_{N}^{2} W=\nabla \times G_{N}^{2} \overline{\mathbb{A}}^{\delta_{2}}$, we obtain from (4.12) and (4.14) that (4.3) holds.
5. Alfvén waves. In this section we prove that our model possesses a very important property of the MHD: the ability of the magnetic field to transmit transverse inertial waves - Alfvén waves. We follow the argument typically used to prove the existence of Alfvén waves in MHD, see, e.g., [7].

Using the density $\rho$ and permeability $\mu$, we write the equations of the model (1.5) in the form

$$
\begin{align*}
& w_{t}+\nabla \cdot\left({\overline{\left(G_{N}^{1} w\right)\left(G_{N}^{1} w\right)}}^{\delta_{1}}\right)+\nabla \bar{p}^{\delta_{1}}=\frac{1}{\rho \mu}{\overline{\left(\nabla \times G_{N}^{2} W\right) \times G_{N}^{2} W^{1}}-\nu \nabla \times(\nabla \times w)}_{\delta_{1}}^{\frac{\partial W}{\partial t}={\overline{\nabla \times\left(\left(G_{N}^{1} w\right) \times\left(G_{N}^{2} W\right)\right)}}^{\delta_{2}}-\eta \nabla \times(\nabla \times W)}  \tag{5.1a}\\
& \nabla \cdot w=0, \quad \nabla \cdot W=0 \tag{5.1b}
\end{align*}
$$

where $\nu=\frac{1}{\operatorname{Re}}, \eta=\frac{1}{\operatorname{Re}_{\mathrm{m}}}$.
Assume a uniform, steady magnetic field $W_{0}$, perturbed by a small velocity field $w$. We denote the perturbations in current density and magnetic field by $j_{\text {model }}$ and $W_{p}$, with

$$
\begin{equation*}
\nabla \times W_{p}=\mu j_{\text {model }} \tag{5.2}
\end{equation*}
$$

Also, the vorticity of the model is

$$
\begin{equation*}
\omega_{\text {model }}=\nabla \times w \tag{5.3}
\end{equation*}
$$

Since $G_{N}^{1} w \cdot \nabla G_{N}^{1} w$ is quadratic in the small quantity $w$, it can be neglected in the Navier-Stokes equation (5.1a), and therefore

$$
\begin{equation*}
\frac{\partial w}{\partial t}+\nabla \bar{p}^{\delta_{1}}=\frac{1}{\rho \mu} \overline{\left(\nabla \times G_{N}^{2} W_{p}\right) \times G_{N}^{2} W_{0}}{ }^{\delta_{1}}-\nu \nabla \times(\nabla \times w) . \tag{5.4}
\end{equation*}
$$

The leading order terms in the induction equation (5.1b) are

$$
\begin{equation*}
\frac{\partial W_{p}}{\partial t}={\overline{\nabla \times\left(G_{N}^{1} w \times G_{N}^{2} W_{0}\right)}}^{\delta_{2}}-\eta \nabla \times\left(\nabla \times W_{p}\right) \tag{5.5}
\end{equation*}
$$

Following the argument of [20] and using the approximating result of Corollary 2.2, we obtain that in the case of a perfect fluid $(\nu=\eta=0)$ and in the case $\nu=0$, $\eta \gg 1$ a transverse wave is recovered. The group velocity of the wave is equal to

$$
\tilde{v}_{a}=v_{a}+O\left(\delta_{1}^{2 N+2}+\delta_{2}^{2 N+2}\right)
$$

where $v_{a}$ is the Alfvén velocity $W_{0} / \sqrt{\rho \mu}$.
We conclude that our model (1.5) preserves the Alfvén waves and the group velocity of the waves $\tilde{v}_{a}$ tends to the true Alfvén velocity $v_{a}$ as the radii tend to zero.
6. Computational results. In this section we present computational results for the ADM models of zeroth, first and second order. The convergence rates are presented and the fidelity of the models is verified by comparing the quantities, which are conserved in the ideal inviscid case. The computations are made for the twodimensional problem, where the energy and enstrophy of the models are compared to those of the averaged MHD.

Consider the MHD flow in $\Omega=(0.5,1.5) \times(0.5,1.5)$. The Reynolds number and magnetic Reynolds number are $R e=10^{5}, R e_{m}=10^{5}$, the final time is $T=1 / 4$, and the averaging radii are $\delta_{1}=\delta_{2}=h$.

Take

$$
\begin{aligned}
& f=\binom{\frac{1}{2} \pi \sin (2 \pi x) e^{-4 \pi^{2} t / R e}-x e^{2 t}}{\frac{1}{2} \pi \sin (2 \pi y) e^{-4 \pi^{2} t / R e}-y e^{2 t}} \\
& \nabla \times g=\binom{e^{t}\left(x-(\cos \pi x \sin \pi y+\pi x \sin \pi x \sin \pi y+\pi y \cos \pi x \cos \pi y) e^{-2 \pi^{2} t / R e}\right)}{e^{t}\left(-y-(\sin \pi x \cos \pi y+\pi x \cos \pi x \cos \pi y+\pi y \sin \pi x \sin \pi y) e^{-2 \pi^{2} t / R e}\right)}
\end{aligned}
$$

The solution to this problem is

$$
\begin{aligned}
u & =\binom{-\cos (\pi x) \sin (\pi y) e^{-2 \pi^{2} t / R e}}{\sin (\pi x) \cos (\pi y) e^{-2 \pi^{2} t / R e}}, \\
p & =-\frac{1}{2}(\cos (2 \pi x)+\cos (2 \pi y)) e^{-4 \pi^{2} t / R e}, \\
B & =\binom{x e}{-y e} .
\end{aligned}
$$

Hence, although the theoretical results were obtained only for the periodic boundary conditions, we apply the family of ADMs to the problem with Dirichlet boundary conditions.

The results presented in the following tables are obtained by using the software FreeFEM + +. The velocity and magnetic field are sought in the finite element space of piecewise quadratic polynomials, and the pressure in the space of piecewise linears. In order to draw conclusions about the convergence rate, we take the time step $k=h^{2}$. We compare the solutions $(w, W)$, obtained by the ADM models, to the true solution $(u, B)$ and the average of the true solution $(\bar{u}, \bar{B})$. The second order accuracy in approximating the true solution $(u, B)$ is expected for ADM models of any order, whereas the accuracy in approximating the averaged solution $(\bar{u}, \bar{B})$ should increase as the order of the model increases.

The solution, computed by the zeroth order ADM, approximates both the true solution $(u, B)$ and the average of the true solution $\left(\bar{u}=\left(-\delta_{1}^{2} \Delta+I\right)^{-1} u, \bar{B}=\left(-\delta_{2}^{2} \Delta+\right.\right.$ $I)^{-1} B$ with the second order accuracy. The accuracy in approximating the averaged solution increases as the order of the model is increased.

Hence, the computational results verify the claimed accuracy of the model.
TABLE 6.1
Approximating the true solution, $R e=10^{5}, R e_{m}=10^{5}$, Zeroth Order ADM

| $h$ | $\\|w-u\\|_{L^{2}\left(0, T ; L^{2}(\Omega)\right)}$ | rate | $\\|W-B\\|_{L^{2}\left(0, T ; L^{2}(\Omega)\right)}$ | rate |
| :---: | :---: | :---: | :---: | :---: |
| $1 / 4$ | 0.0862904 |  | 0.0253257 |  |
| $1 / 8$ | 0.0515562 | 0.7431 | 0.0268628 | -0.085 |
| $1 / 16$ | 0.0204763 | 1.3322 | 0.0132399 | 1.0207 |
| $1 / 32$ | 0.00611337 | 1.7439 | 0.00412013 | 1.6841 |
| $1 / 64$ | 0.00163356 | 1.9039 | 0.001116 | 1.8844 |

TABLE 6.2
Approximating the true solution, $R e=10^{5}$, Re $e_{m}=10^{5}$, First Order ADM

| $h$ | $\\|w-u\\|_{L^{2}\left(0, T ; L^{2}(\Omega)\right)}$ | rate | $\\|W-B\\|_{L^{2}\left(0, T ; L^{2}(\Omega)\right)}$ | rate |
| :---: | :---: | :---: | :---: | :---: |
| $1 / 4$ | 0.086748 |  | 0.0219869 |  |
| $1 / 8$ | 0.0504853 | 0.781 | 0.0146218 | 0.5885 |
| $1 / 16$ | 0.0196045 | 1.3647 | 0.00401043 | 1.8663 |
| $1 / 32$ | 0.00589278 | 1.7342 | 0.00078723 | 2.3489 |
| $1 / 64$ | 0.00159084 | 1.8892 | 0.000170555 | 2.2065 |

Since the flow is not ideal (nonzero power input, nonzero viscosity/magnetic diffusivity, non-periodic boundary conditions), the energy and enstrophy are not conserved.

Table 6.3
Approximating the true solution, $R e=10^{5}, R e_{m}=10^{5}$, Second Order ADM

| $h$ | $\\|w-u\\|_{L^{2}\left(0, T ; L^{2}(\Omega)\right)}$ | rate | $\\|W-B\\|_{L^{2}\left(0, T ; L^{2}(\Omega)\right)}$ | rate |
| :---: | :---: | :---: | :---: | :---: |
| $1 / 4$ | 0.0854318 |  | 0.0229699 |  |
| $1 / 8$ | 0.0500093 | 0.7726 | 0.0170217 | 0.4324 |
| $1 / 16$ | 0.0194169 | 1.3649 | 0.00472331 | 1.8495 |
| $1 / 32$ | 0.00587995 | 1.7234 | 0.000856363 | 2.4635 |
| $1 / 64$ | 0.00159835 | 1.8792 | 0.000167472 | 2.3543 |

TABLE 6.4
Approximating the average solution, $R e=10^{5}, R e_{m}=10^{5}$, Zeroth Order ADM

| $h$ | $\\|w-\bar{u}\\|_{L^{2}\left(0, T ; L^{2}(\Omega)\right)}$ | rate | $\\|W-\bar{B}\\|_{L^{2}\left(0, T ; L^{2}(\Omega)\right)}$ | rate |
| :---: | :---: | :---: | :---: | :---: |
| $1 / 4$ | 0.0247837 |  | 0.0253257 |  |
| $1 / 8$ | 0.0245241 | 0.0152 | 0.0268628 | -0.085 |
| $1 / 16$ | 0.0131042 | 0.9042 | 0.0132399 | 1.0207 |
| $1 / 32$ | 0.00434599 | 1.5923 | 0.00412013 | 1.6841 |
| $1 / 64$ | 0.00120907 | 1.8458 | 0.001116 | 1.8844 |

Table 6.5
Approximating the average solution, Re $=10^{5}, R e_{m}=10^{5}$, First Order ADM

| $h$ | $\\|w-\bar{u}\\|_{L^{2}\left(0, T ; L^{2}(\Omega)\right)}$ | rate | $\\|W-B\\|_{L^{2}\left(0, T ; L^{2}(\Omega)\right)}$ | rate |
| :---: | :---: | :---: | :---: | :---: |
| $1 / 4$ | 0.0228254 |  | 0.0219869 |  |
| $1 / 8$ | 0.015202 | 0.5864 | 0.0146218 | 0.5885 |
| $1 / 16$ | 0.0043297 | 1.8119 | 0.00401043 | 1.8663 |
| $1 / 32$ | 0.000867986 | 2.3185 | 0.00078723 | 2.3489 |
| $1 / 64$ | 0.000192121 | 2.1757 | 0.000170555 | 2.2065 |

Table 6.6
Approximating the average solution, $R e=10^{5}, R e_{m}=10^{5}$, Second Order ADM

| $h$ | $\\|w-\bar{u}\\|_{L^{2}\left(0, T ; L^{2}(\Omega)\right)}$ | rate | $\\|W-B\\|_{L^{2}\left(0, T ; L^{2}(\Omega)\right)}$ | rate |
| :---: | :---: | :---: | :---: | :---: |
| $1 / 4$ | 0.0236209 |  | 0.0229699 |  |
| $1 / 8$ | 0.0172027 | 0.4574 | 0.0170217 | 0.4324 |
| $1 / 16$ | 0.00506669 | 1.7635 | 0.00472331 | 1.8495 |
| $1 / 32$ | 0.000956194 | 2.4057 | 0.000856363 | 2.4635 |
| $1 / 64$ | 0.000194768 | 2.2955 | 0.000167472 | 2.3543 |

But we expect the energy and enstrophy of the models to approximate the energy and enstrophy of the averaged MHD.

The enstrophy of the first and second order models approximates the enstrophy of the averaged MHD better than the zeroth order model's enstrophy, see Figure 6.1.

Figure 6.2 shows that the graph of the models energy is hardly distinguishable from that of the averaged MHD.

Zooming in at the final time $t=0.25$ we verify that the ADM energy approximates the averaged MHD energy better as the model's order increases, see Figure 6.3.


Fig. 6.1. ADM Enstrophy vs. averaged MHD


Fig. 6.2. ADM Energy vs. averaged $M H D$
7. Conclusions. A Large Eddy Simulation approach to the MagnetoHydroDynamic Turbulence was considered. The Approximate Deconvolution Models were introduced for the incompressible MHD equations, and this family of models was analyzed. We proved the existence and uniqueness of solutions, and their convergence in the weak sense to a solution of the MHD equations, as the filtering widths are


Fig. 6.3. ADM Energy vs. averaged MHD: zoom in
decreased to zero. We proved the accuracy of the model both theoretically (by establishing an á priori bound on the model's consistency error) and numerically (the results of the computational tests are listed in the previous section).

Also, all models in the family of the ADMs are proven to possess the physical properties of the MHD - the energy and helicity of the models are conserved, and the models are also proven to preserve the Alfvén waves, a unique feature of the MHD equations. The physical fidelity of the models was also verified computationally. The test results prove that both the solution and the energy of the averaged MHD equations are approximated better, as one increases the models' order N (from zeroth ADM to the first ADM, and from the first to the second ADM). This gives a freedom of choosing the model's order N , based on the desired accuracy of approximation and the available computational power. Finally, the tests demonstrate that in the situations when the direct numerical simulation is no longer available (flows with high Reynolds and magnetic Reynolds numbers), the solution can still be obtained by the ADM approach.

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