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Abstract

For accurate line-by-line modeling of molecular cross sections several physical processes “beyond Voigt” have to be considered. For the speed-dependent Voigt and Rautian profiles (SDV, SDR) and the Hartmann-Tran profile the difference $w(iz_-) - w(iz_+)$ of two complex error functions (essentially Voigt functions) has to be evaluated where the function arguments z_{\pm} are given by the sum and difference of two square roots. These two terms describing z_{\pm} can be huge and the default implementation of the difference can lead to large cancellation errors. First we demonstrate that these problems can be avoided by a simple reformulation of z_- . Furthermore we show that a single rational approximation of the complex error function valid in the whole complex plane (e.g. by Humlíček, 1979 or Weideman, 1994) enables computation of the SDV and SDR with four significant digits or better. Our benchmarks indicate that the SDV and SDR functions are about a factor 2.2 slower compared to the Voigt function, but for evaluation of molecular cross sections this time lag does not significantly prolong the overall program execution because speed-dependent parameters are available only for a fraction of strong lines.

Keywords: Complex error function; Voigt profile; Hartmann-Tran profile; Rational approximations

1. Introduction

The Voigt profile [1] accounting for collision (pressure) and Doppler (thermal) broadening has been the standard for high-resolution line-by-line modeling of infrared (IR) and microwave molecular absorption. Inadequacies of this profile have been observed since decades in molecular laboratory spectroscopy [e.g. 2], and more recently discrepancies between theory and measurements have also become significant in atmospheric spectroscopy.

Collision induced changes of molecular velocity reduce the Doppler broadening, and several profiles have been suggested to describe this collisional (or Dicke) narrowing, e.g. the Rautian profile [3]. Moreover, the speed-dependence of the relaxation rates modifies the Lorentz line shape modeling the collision broadening [4, 5] and can be described by the Speed-Dependent Voigt (SDV) profile. Assuming that these processes are independent leads to the Speed-Dependent Rautian (SDR) profile. The partial correlation of these effects is considered by the partially Correlated quadratic-Speed-Dependent Hard-Collision profile (pCqSDHCP), originally developed by Tran, Ngo, and Hartmann [6] (therefore called Hartmann-Tran or HT profile for short) and recommended for high resolution spectroscopy in a recent IUPAC Technical Report [5].

Using these more sophisticated profiles has been shown to improve the analysis of ground-based Fourier transform IR spectra and space-borne limb occultation measurements of the Atmospheric Chemistry Experiment — Fourier Transform Spectrometer (ACE-FTS) [e.g. 7–9]. The stringent accuracy and precision requirements of current space-borne carbon dioxide and methane observations (e.g. the Greenhouse Gases Observing Satellite [GOSAT, 10], the Orbiting Carbon Observatory [OCO-2, 11] and the Sentinel-5p TROPOspheric Monitoring Instrument [TROPOMI, 12]) have also highlighted the need for improved molecular spectroscopy in the short-wave IR [e.g. 13–19]. Improved Atmospheric Spectroscopy databases (IAS) have been the objective of a study funded by ESA in the context of the Scientific Exploitation of Operational Missions (SEOM) initiative: SEOM-IAS provides data for CO, CH₄ and H₂O in the TROPOMI 2.3 μm region [20]. Moreover, the HITRAN database has started to include “beyond-Voigt” parameters for selected molecules [21].

Computational aspects of the SDV profile have been discussed in a previous paper [22] (henceforth “CASDV”). Here we extend this discussion to the SDR profile. After a brief survey of available line data we recall some definitions and relations in the following section. We continue with a discussion of some subtleties of the SDV function and then present comparisons of SDR evaluations against a reference code in section 3. Moreover, we provide an extensive assessment of the computational efficiency of the SDV and SDR (floating-point operation counts and time benchmarks). A summary is given in the final section 4. For simplicity, pressure-induced line shift, self-broadening,

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and line mixing will be ignored. Python modules of the SDV and SDR functions along with rational approximations for the complex error function are provided in supplemental files.

2. Line data and models

2.1. Line data

For an assessment of the performance of various line model algorithms and implementations it is useful to recap the range of the function arguments to be expected. In Schreier [23, 24] we concluded that the ratio of the Lorentz and Gauss width (the y parameter, see next subsection) can be as small as 10^{-8} for atmospheric IR spectroscopy. Even smaller values are possible for large wavenumbers (visible or UV), light molecules (e.g. H_2) and warm to hot atmospheres.

Considering the new parameters required by the speed-dependent profiles, a survey of the SEOM-IAS data indicates that the ratio of the collision broadening parameters has a mean value of $\gamma_2/\gamma_0 \approx 0.1$, see Fig. 1. Furthermore, the bottom-right plot of Fig. 1 shows that the frequency of velocity-changing collisions (Dicke narrowing) is about one tenth of the air-broadening parameter, i.e. $\nu_{\text{vc}} \approx 0.01 \text{ cm}^{-1} \approx 0.1\gamma_0$. Note that in SEOM-IAS the partial correlation between speed-dependence and velocity changes is not considered, i.e. the HT parameter η is assumed to be zero and the SDR is sufficient for this database. Furthermore, the line shift's speed dependence parameter δ_2 is zero for all carbon monoxide and water lines, whereas 12 of the 11905 methane lines have nonzero values with $-0.006545 \leq \delta_2 \leq 0.00316 \text{ cm}^{-1}$.

In the latest version of the HITRAN database, “beyond-Voigt” parameters are available for seven (of 49) molecules according to Table 3 of Gordon et al. [21]. However, these are not necessarily HT or SDV parameters and might describe, for example, line mixing. For H_2O there are 2016 (of about 52 000) lines with nonzero γ_2 in the spectral range $1852\text{--}3995 \text{ cm}^{-1}$ and a minimum line strength $S = 6.9 \cdot 10^{-25} \text{ cm}^{-1}/(\text{molec.cm}^{-2})$, see Fig. 2 (left) and Loos et al. [25, 26]; moreover the Dicke narrowing parameter is nonzero for 237 transitions with a mean of 0.0157 cm^{-1} . For N_2O 80 lines around 2200 cm^{-1} with $S \geq 10^{-19} \text{ cm}^{-1}/(\text{molec.cm}^{-2})$ are listed (Fig. 2). For both molecules the speed dependence of the line width is about one tenth of the speed-averaged width, and the correlation parameter η is zero for all transitions. (For H_2 see also [27].)

Similar to HITRAN and GEISA, the SEOM-IAS line parameters are given for a reference pressure $p_{\text{ref}} = 1 \text{ atm}$ and temperature $T_{\text{ref}} = 296 \text{ K}$. Both broadening parameters are assumed to have the same pressure and temperature dependence, i.e. $\gamma_{0,2} \sim pT^{-n}$ with the temperature exponent n (1/2 according to classical theory), whereas the narrowing parameter is inversely proportional to temperature, $\nu_{\text{vc}} \sim p/T$ [see also 28, 29]. In Earth's (or a

planetary) atmosphere these three parameters are varying over orders of magnitude, but the ratios γ_2/γ_0 and ν_{vc}/γ_0 are approximately constant over pressure (or altitude).

2.2. Line models

Following Varghese and Hanson [3] we distinguish between the Voigt profile g and its generalizations and the Voigt function K and its generalizations

$$g(\nu; \hat{\nu}, \gamma_L, \gamma_G, \dots) = \frac{\sqrt{\ln 2/\pi}}{\gamma_G} K(x, y, \dots). \quad (1)$$

The profiles g are defined in terms of the physical variables wavenumber ν , center position $\hat{\nu}$, and broadening parameters, all in units of reciprocal wavenumbers [cm^{-1}] (in the following all γ as well as ν_{vc} are considered for the actual pressure and temperature). Division by the Gaussian (Doppler) half width allows to introduce dimensionless variables that reduces the number of arguments by one and is more convenient from a mathematical point:

$$x \equiv \sqrt{\ln 2}(\nu - \hat{\nu})/\gamma_G \quad y \equiv \sqrt{\ln 2}\gamma_L/\gamma_G \quad (2a)$$

$$q \equiv \sqrt{\ln 2}\gamma_2/\gamma_G \quad \zeta \equiv \sqrt{\ln 2}\nu_{\text{vc}}/\gamma_G. \quad (2b)$$

From a computational point, these profiles and corresponding functions can be calculated readily from the complex error function [30–32]

$$w(z) = \frac{i}{\pi} \int_{-\infty}^{\infty} \frac{e^{-t^2}}{z - t} dt \quad (3)$$

$$= K(x, y) + iL(x, y) \quad \text{with } z = x + iy. \quad (4)$$

The real part is the “standard” Voigt function, essentially the convolution of a Lorentzian and Gaussian function,

$$K(x, y) = \frac{y}{\pi} \int_{-\infty}^{\infty} \frac{e^{-t^2}}{(x - t)^2 + y^2} dt, \quad (5)$$

where x and y are the standard Voigt function parameters (2a) characterizing the distance to the line center at wavenumber $\hat{\nu}$ and the ratio of the Lorentzian and Gauss widths.

The SDR function is essentially the quotient of the difference of two complex error functions

$$K_{\text{sdr}}(x, y, q, \zeta) = \text{Re} \left(\frac{w(iz_-) - w(iz_+)}{1 - \sqrt{\pi}\zeta(w(iz_-) - w(iz_+))} \right) \quad (6)$$

where the extra parameters q and ζ are the broadening's quadratic speed dependence γ_2 and the frequency of velocity-changing collisions ν_{vc} normalized by the Gaussian widths, respectively, Eq. (2b). The arguments of the complex error function are defined by ¹

$$z_{\pm} = \sqrt{\frac{y + \zeta - ix}{q} - \frac{3}{2} + \frac{1}{4q^2}} \pm \frac{1}{2q} \quad (7)$$

$$= \sqrt{X + Y} \pm \sqrt{Y}. \quad (8)$$

¹The upper case X and Y correspond to the HT notation [5, 6]

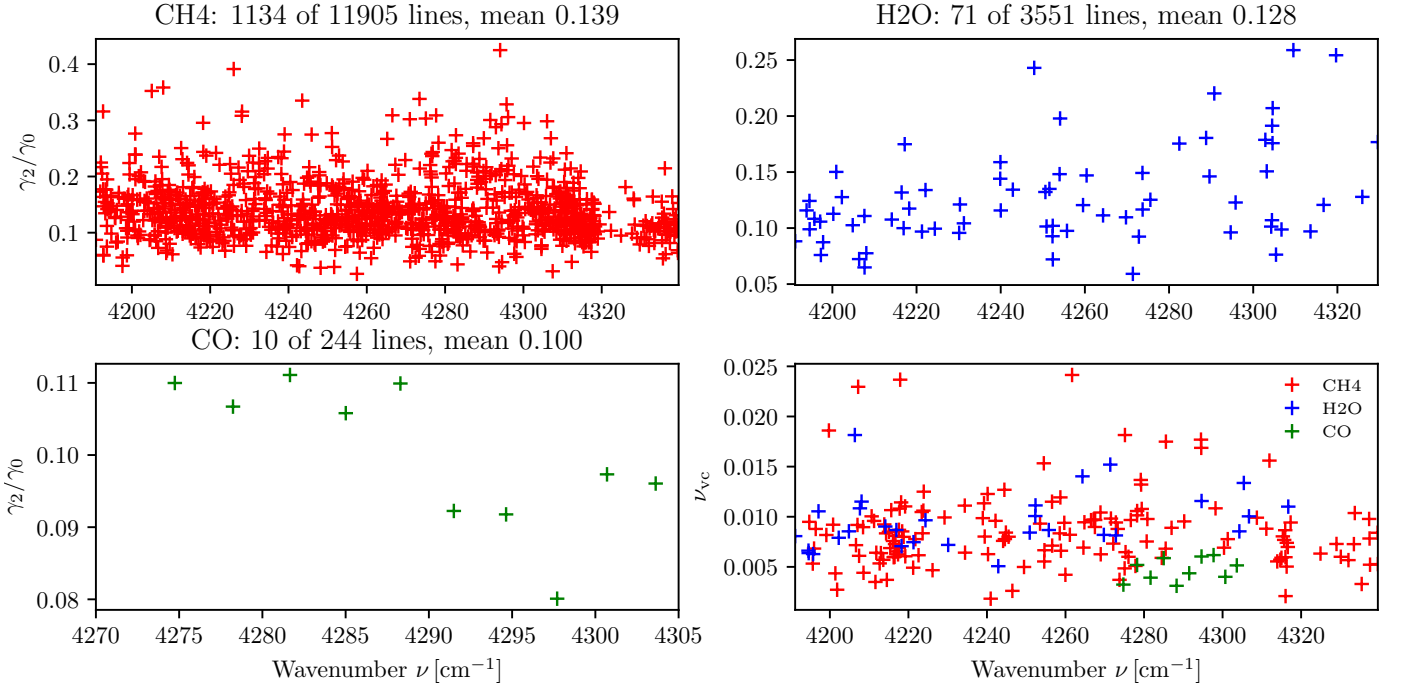


Figure 1: Survey of SEOM-IAS line data. Top and bottom left: ratio of the collision broadening parameters, γ_2/γ_0 (the title indicates the number of lines with nonzero γ_2 and the mean value of the ratio). Bottom right: Dicke narrowing parameter ν_{vc} .

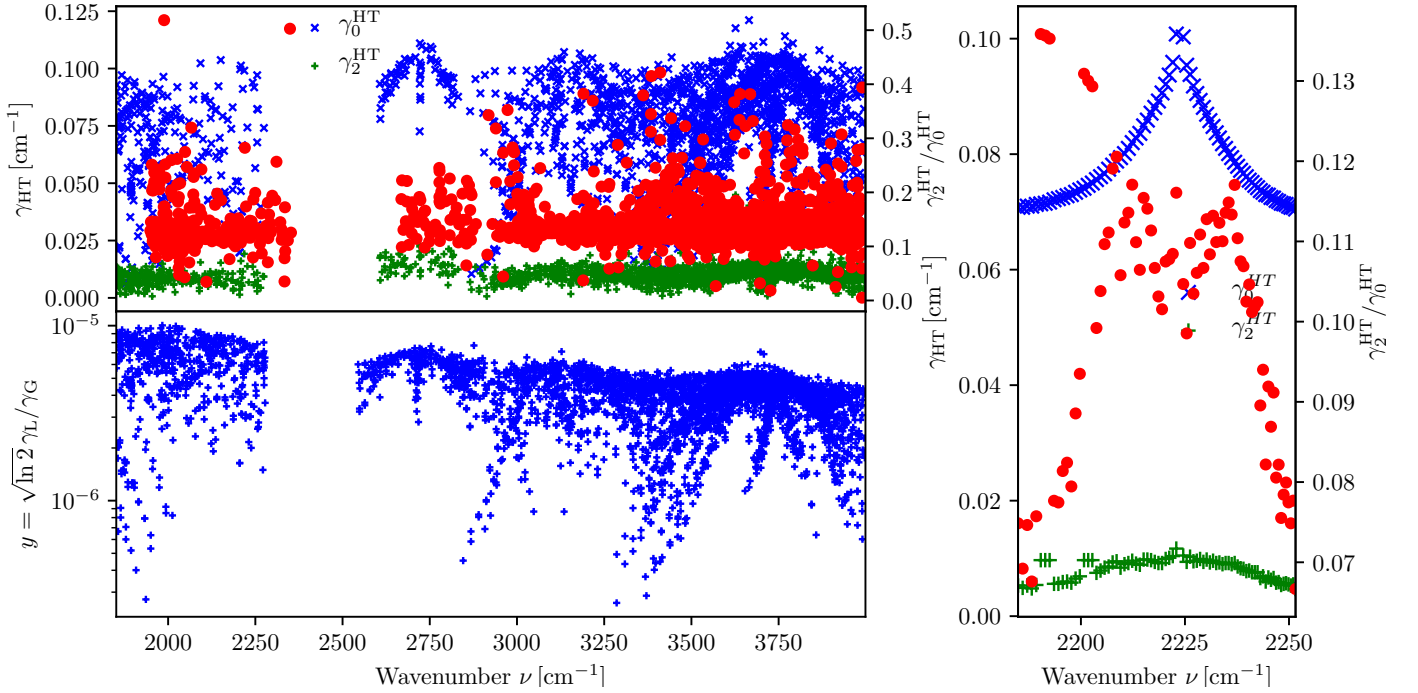


Figure 2: HITRAN H₂O (left) and N₂O (right) lines with HT widths defined: Top left: collision broadening parameters (x and + markers, left axis) and their ratio γ_2/γ_0 (red filled circles, right axis). Bottom left: the Voigt parameter y for ToA @ 120 km.

For vanishing $\nu_{vc} \propto \zeta = 0$ the SDR function reduces to the SDV function

$$K_{\text{sdv}}(x, y, q) = \text{Re}(w(iz_-) - w(iz_+)), \quad (9)$$

whereas for zero speed dependence $\gamma_2 \propto q = 0$ the Rautian function is obtained,

$$K_r(x, y, \zeta) = \text{Re} \left(\frac{w(z)}{1 - \sqrt{\pi} \zeta w(z)} \right). \quad (10)$$

A vast number of algorithms have been developed for the complex error function exploiting a wide variety of numerical techniques. In recent papers we have discussed rational approximations by Humlíček [34, 35] and Weideman [36] and recommended combinations of the asymptotic approximation given in the Humlíček [35] `w4` code (the quotient of a first and second degree polynomial) with the more accurate approximations of [34] or [36] for small arguments [23, 24].

3. Results

3.1. Computation of the complex square roots

The arguments iz_{\pm} , Eq. (7), of the complex error functions for (6) and (9) require the computation of a complex square root. In essence²,

$$r = \sqrt{\alpha + i\beta} \quad (11a)$$

$$= \frac{1}{\sqrt{2}} \sqrt{\sqrt{\alpha^2 + \beta^2} + \alpha} + \text{sign}(\beta) \frac{i}{\sqrt{2}} \sqrt{\sqrt{\alpha^2 + \beta^2} - \alpha} \quad (11b)$$

$$= \frac{1}{\sqrt{2}} \sqrt{\sqrt{\alpha^2 + \beta^2} + \alpha} + \frac{i\beta/\sqrt{2}}{\sqrt{\sqrt{\alpha^2 + \beta^2} + \alpha}} \quad (11c)$$

The three variants of (11) are all equivalent mathematically, but might be different numerically.

Tran, Ngo, and Hartmann [6] (henceforth “TNH”) evaluate the square root of the complex argument using Fortran’s intrinsic function `CDSQRT`, essentially the right hand side of Eq. (11a), whereas the Boone, Walker, and Bernath [33] (henceforth “BWB”) Fortran implementation is based on Eq. (11b). In `CASDV` we have indicated that the alternative form (11c) might be advantageous computationally because possible inaccuracies of the difference evaluation can be avoided.

with $X \equiv [(i(\dot{\nu} - \nu) + \gamma_L + \nu_{vc})/\gamma_2 - \frac{3}{2}]^2$ and $Y \equiv (\gamma_G/2\sqrt{\ln 2}\gamma_2)^2$. Note that Y is real, a non-vanishing line shift speed-dependence parameter would lead to a complex Y . Boone et al. [33] define two variables α and β corresponding to the real and imaginary part of X (with $\nu_{vc} = 0$), a third variable δ is identical to Y . The function $\text{Re}(A)$ defined in Eq. (7) of Tennyson et al. [5] is identical to the “SDV function” K_{sdv} except for a factor $\sqrt{\pi \ln 2}/\gamma_G$. Note that in footnote 1 of [22] we have incorrectly stated that an Im operator is missing in Eq. (11) of Tran et al. [6].

²This is essentially the square root used by Boone et al. [33] without the δ term.

With respect to efficiency, it requires only two instead of three square root evaluations (but an extra division) and there is no need to evaluate the sign function. Note that the Voigt and generalized Voigt functions are symmetric w.r.t. $x \propto \beta \equiv \text{Im}(X)$, i.e. they depend on $|\beta|$ only and the `sign` function can be avoided.

The contour plots of Fig. 3 demonstrate that the three forms are equivalent except for very small y , i.e. the relative differences $|\Delta K|/K$ are small. However, discrepancies can be observed for $y < 10^{-6}$. Inspection of the SDV for $y = 5 \cdot 10^{-8}$ (Fig. 3 bottom right) indicates that (11b) produces discrete steps near the line center (small x), where the other two forms are indistinguishable. In the wings (larger x) the three forms show a zigzag behaviour; (11b) and (11c) are quite similar and somewhat smaller than (11a) that oscillates around the Voigt function. Fortunately the function values here are many orders of magnitude smaller compared to the center value and the apparent errors are likely to have no impact on molecular cross sections resulting from the superposition of numerous lines. Note that TNH carefully discuss limiting asymptotic cases and have implemented dedicated algorithms for small and large $|X|/Y$ in the supplemental Fortran source code.

3.2. Problems with tiny values of y and q .

Although the direct evaluation of the complex square (11a) appears to be better, the zigzag indicates further problems. A closer analysis indicates that these problems are related to the evaluation of z_- , i.e. the difference of the complex square root $\sqrt{X+Y}$ and \sqrt{Y} . For tiny y (and $q < y$) the real valued $Y = 1/4q^2$ becomes extremely big, much bigger than the complex valued X , and the standard floating point precision with about 16 significant digits fails to evaluate this difference. To test this hypothesis, the SDV has been implemented using the Python library `mpmath` [37] allowing real and complex floating-point arithmetic with arbitrary precision. With 32 significant figures (roughly quadruple precision) the square roots, their difference and the complex error function (computed by $w(z) = \exp(-z^2) \text{erfc}(-iz)$) can be evaluated reliably even for these extreme cases.

Fig. 4 (left) shows results for $y = 10^{-8}$ and $q = 0.1y$ (some specific numbers for $x = 10$ and $x = 12$ are listed in the appendix): Whereas $\alpha = \frac{y}{q} - \frac{3}{2}$ is a small number, $Y \equiv \delta = 1/4q^2$ is huge, and with the default double precision accuracy of Python (or C or Fortran) the real part of the first square root in (11) is essentially independent of α . For the x values considered here ($0 \leq x \leq 25$) the imaginary part of $X+Y$, i.e. x/q is a moderately large number. However, the real part of the complex square root $\sqrt{X+Y}$ is identical in double precision for $x = 10$ and $x = 12$ (but clearly distinct with enhanced precision). Moreover, because $|X| \ll Y$, the difference of two huge, approximately similar numbers cannot be evaluated correctly, resulting in a stepwise behaviour of $\text{Re}(\sqrt{X+Y} - \sqrt{Y})$ (Fig. 4

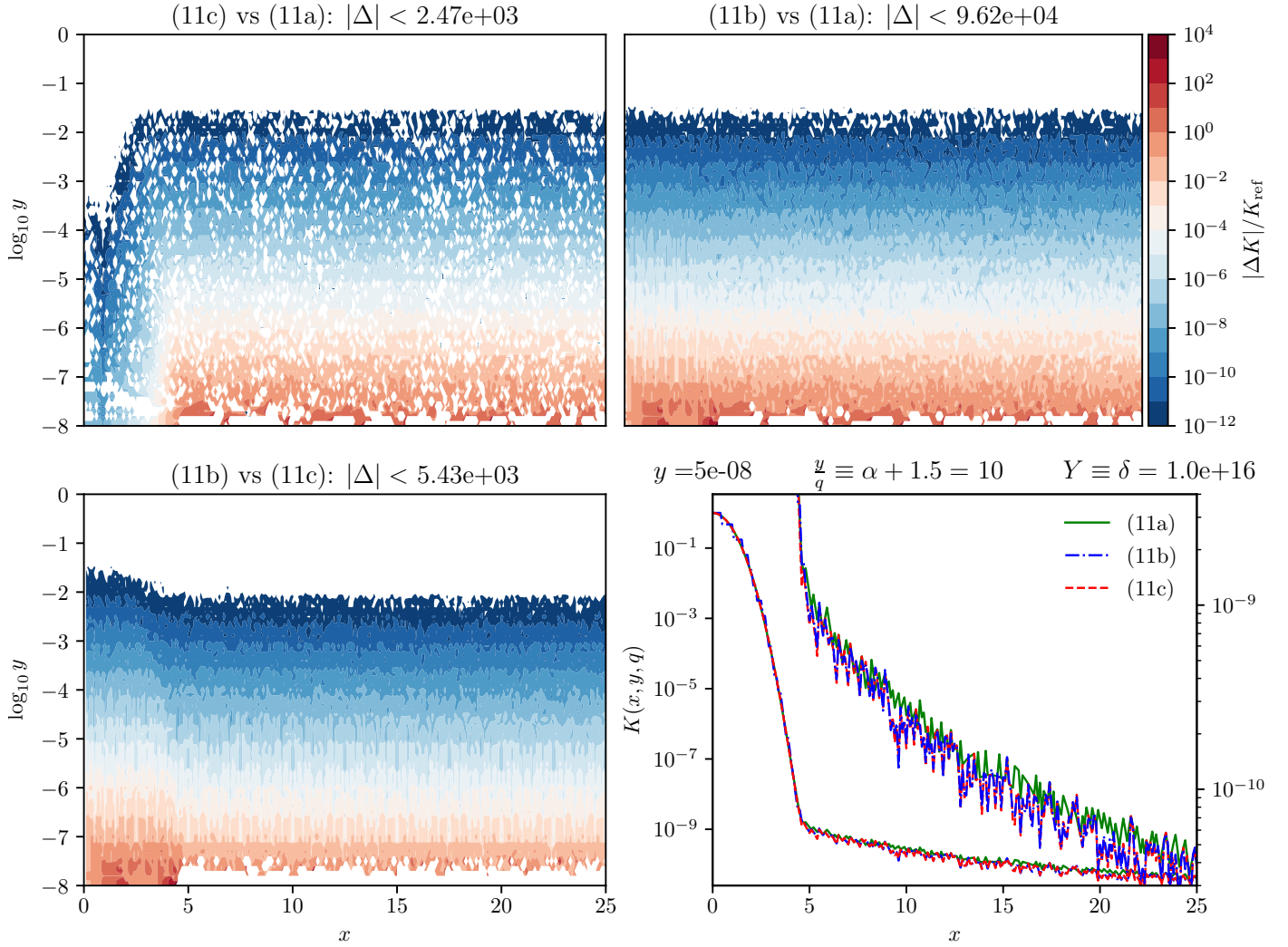


Figure 3: Contour plot of relative differences of the SDV function for $\gamma_L/\gamma_2 = 10$: mutual differences of the three versions shown in Eq. (11). Bottom right: the SDV functions for $y = 5 \cdot 10^{-8}$; the curves are plotted twice with the entire K range (left axis) and with the K range limited to small values (right axis).

top-left). As a consequence, the evaluation of $w(iz_-)$ is unreliable, too; in particular, the difference of the complex error functions $w(iz_-) - w(iz_+)$ can become negative, indicated by the missing values in Fig. 4 (center-left).

Similar to the reformulation discussed above (Eqs. (11b) and (11c)) the difference of the two roots (8) can also be computed in a reliable way to avoid the subtraction of two similar numbers [38]³,

$$z_- \equiv \sqrt{X+Y} - \sqrt{Y} = \frac{X}{\sqrt{X+Y} + \sqrt{Y}} = \frac{X}{z_+}. \quad (12)$$

Fig. 4 shows that this safe variant avoids the steps in $\text{Re}(z_-)$ and gives SDV values close to the quadruple precision values; whereas the relative error of the original version is in the percentage range, the cancellation-safe implementation has a maximum relative error of about 10^{-5}

(bottom-left). The superiority of this approach is also confirmed by the two examples shown on the right of Fig. 4, where the original “naive” computation delivers oscillations with frequent negative SDV values.

3.3. The SDV function again

Both TNH and BWB have emphasized that the computation of the SDV and HT profiles needs to be done carefully when different complex error function approximations are used for the two arguments iz_{\pm} . In particular, the TNH code implements the Humlíček [34] `cpf12` subroutine (with about 10^{-6} maximum relative error) that utilizes two different approximations for large and small arguments (in fact augmented with a 15-term asymptotic expansion of $w(z)$ for $|z| > 8$). When the two arguments are close to each other, TNH ensure that they are evaluated with the same method. BWB exploit the Humlíček [35] `w4` code (10^{-4} relative accuracy) with four different approximations: their code uses the higher degree rational

³See also <https://herbie.uwplse.org/>

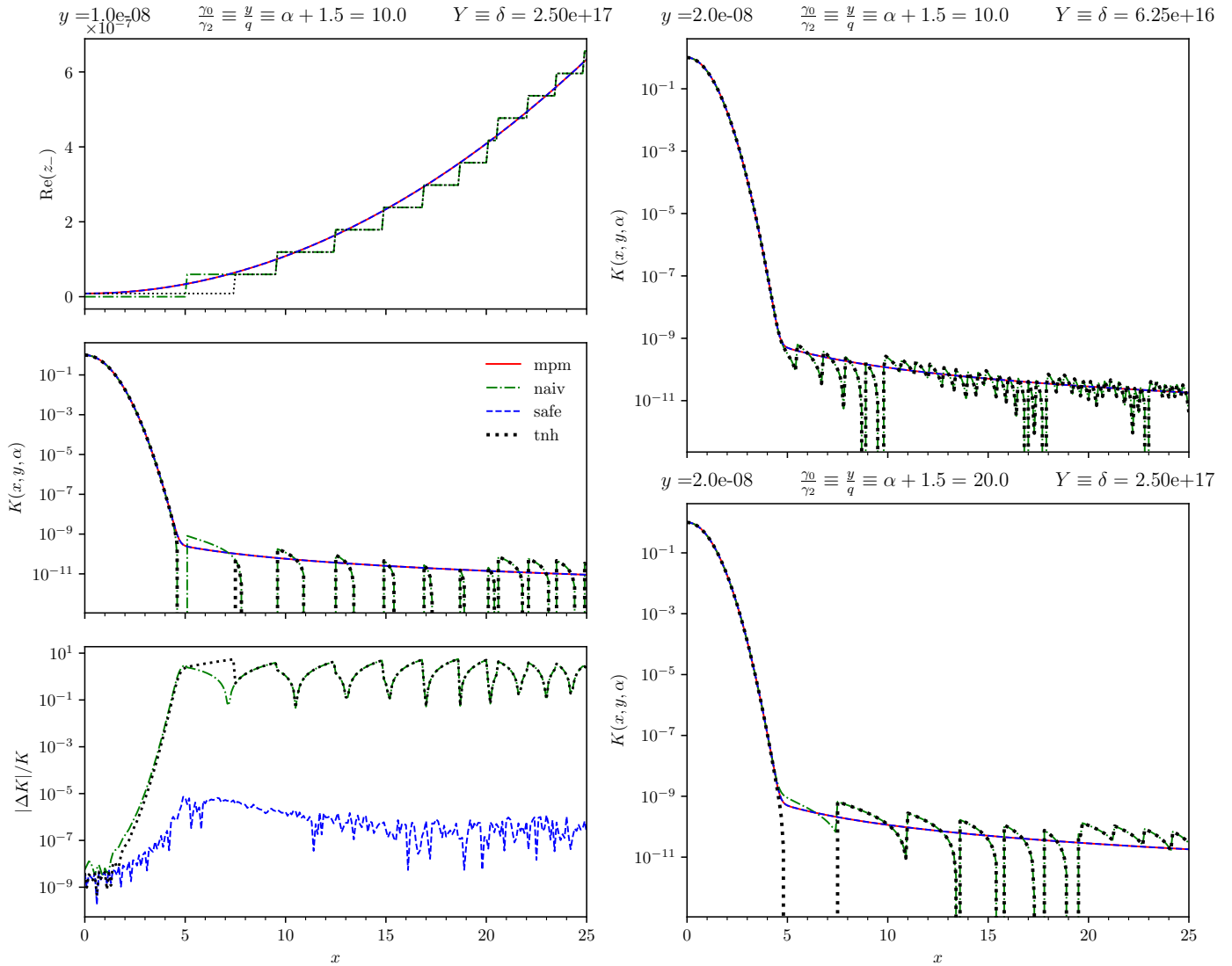


Figure 4: The SDV function evaluated for tiny values of y . Top left: the real part of the complex error function argument z_- for $y = 1 \cdot 10^{-8}$ and $q = 1 \cdot 10^{-9}$. Center left: the corresponding SDV function values. Bottom left: the relative error w.r.t. the “mpm” quadruple precision evaluation. Right: the SDV functions for $y = 2 \cdot 10^{-8}$ and two values of q .

approximation for both arguments. However, as discussed in CASDV, this approach is dangerous because the higher degree approximation is not necessarily valid in the region of large function arguments.

Obviously, these problems could be avoided if a complex error function algorithm is used that is valid for the entire x, y range. Weideman [36] has developed rational approximations that can be used in the entire complex plane. As demonstrated in [23], the $n = 24$ approximation has a maximum relative error less than 10^{-4} for $y > 10^{-4}$ only; with $n = 32$ terms the code is better than $9 \cdot 10^{-5}$ for all $0 \leq x \leq 20$ and $10^{-6} \leq y \leq 10^2$ for both the real and imaginary part of w . The region I approximation of the Humlíček [34] `cpf12` code cannot be used for $y < 10^{-2}$, but a generalization of this 12-term rational approximation applicable to all (x, y) considerably improves the performance [24]: the 16-term approximation (`cpf16`) achieves

an accuracy of about 10^{-4} for $y > 10^{-8}$, and with 20 terms (`cpf20`) the error can be further reduced by more than a factor ten.

Fig. 5 depicts the relative errors of the SDV computed with these “global” complex error function algorithms (the 16- and 20-term Humlíček [34] and the 24- and 32-term Weideman [36] approximations) for $\gamma_L/\gamma_2 = y/q = 10$. As in our previous works we use the `wofz` complex error function provided by SciPy (module `scipy.special`) with a stated accuracy of 13 significant digits as a reference code [see also 39, 40]. All implementations work reasonably good for $y > 10^{-4}$, but for smaller y problems become apparent especially for the Weideman $n = 24$ approximation. The Weideman $n = 32$ and the Humlíček 16-term approximation can be used safely for $y > 10^{-6}$, but only `cpf20` is able to evaluate the SDV with errors less than $5 \cdot 10^{-5}$ for $0 \leq x \leq 25$ and $10^{-8} \leq y \leq 10^2$. Note that the error

SDV with $y/q = 10$

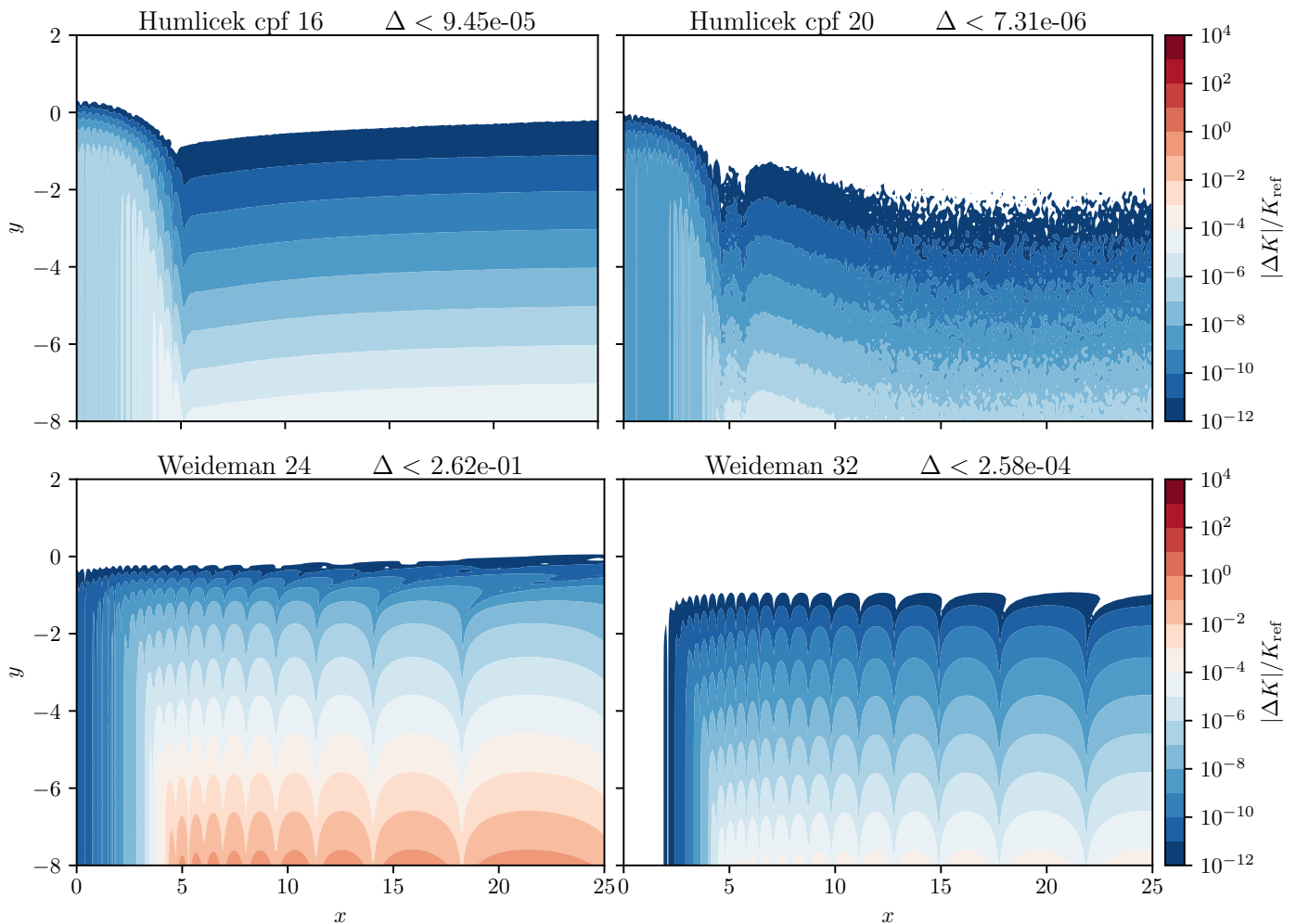


Figure 5: Contour plot of relative differences of the SDV function computed with the Humlíček (top) and Weideman complex error function codes for $\gamma_L/\gamma_2 = 10$. The number in the title indicates the maximum relative error.

patterns for the SDV deviations are similar to the pattern of the complex error function deviations (cf. Fig. 7 in [23] and Fig. 2 and 4 in [24]).

3.4. Beyond-Voigt vs. Voigt

Before analyzing the performance of various complex error function algorithms for the evaluation of the SDR it is instructive to see the relative change of the function due to the inclusion of velocity-changing collisions. Figure 6 shows that the relative change $|K_{\text{sdr}} - K_{\text{vgt}}|/K_{\text{vgt}}$ can be as large as some percent. Differences between SDR and Voigt up to about 4% can be observed in the line center region for $y \approx 1$, whereas for large $x \gg x_{1/2}$ the relative difference is less than 10^{-4} , i.e. the advanced profiles essentially resemble the Voigt profile asymptotically. Here the Voigt function half width is given by the Whiting [41] approximation $x_{1/2} = \frac{1}{2} \left(y + \sqrt{y^2 + 4 \ln 2} \right)$.

3.5. The SDR function

The relative error of the SDR function computed with the Humlíček [34] or Weideman [36] code is shown in Fig. 7. Interestingly the patterns are quite similar to those of Fig. 5. In particular, for the Weideman $n = 24$ rational approximation the errors can become significant already for moderately small y (low pressure), whereas for 32 terms the approximation shows problems only for tiny y . Both versions of the Humlíček complex error function code give reliable results with four and five significant digits for the 16 and 20 term approximation, respectively.

3.6. Performance

For an assessment of the computational efficiency of various implementations we have used IPython’s builtin “magic function” `%timeit` similar to the tests described in Schreier [24]. First we evaluate the SDV for the “matrix” of 251 x grid points and 101 y grid points used for all contour plots shown so far (see Appendix B). Tab. 1 shows

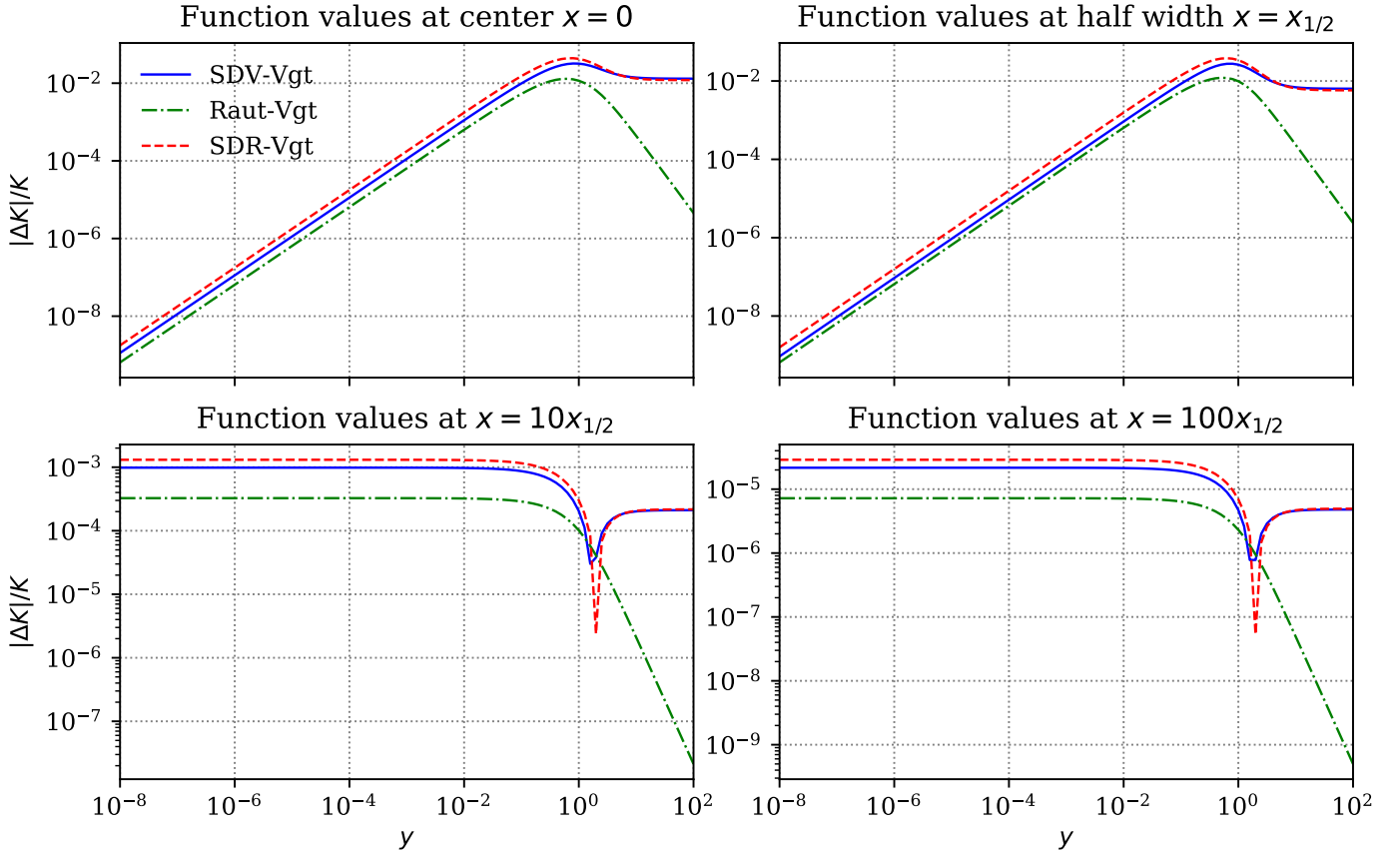


Figure 6: Relative differences of the SDR, SDV, and Rautian relative to the Voigt function as a function of y with $y/q = y/r = 0.1$.

Table 1: Execution time for various SDV implementations. (All tests have been performed on a single Intel Core i5-9600 with 3.1 GHz.)

function	time
cpf16	1.46 ms
sdv_real_bwb (11b)	4.37 ms
sdv_real_casdv (11c)	4.47 ms
sdv_naive (11a)	4.91 ms
sdv_safe (12)	4.94 ms

that the times required for the three different versions of the complex square root (11) and the “cancellation-safe” alternative (12) are approximately identical. Both division and square root are classified as “special function” by Goedecker and Hoisie [42] from a computational point of view, hence the costs of the (11b) and (11c) versions should be roughly identical.

Obviously these advanced line shapes are computationally more expensive. For the following counts of floating-point operations only those required for every x value (corresponding to a wavenumber grid point ν) are considered, i.e. operations executed only once per line (e.g. division by γ_G) are ignored. For both SDV and SDR two complex error functions have to be evaluated and in addition a complex square root (z_+ according to (7)) and a division (z_- according to (12)) is required, a second division for the

SDR (6), whereas for the Rautian function (10) there is a single extra division. Tran et al. [6] have discussed the speed of the HT profile and concluded that “the maximum ratio of the computer time needed for this calculation to that for the corresponding Voigt profile can thus reach $2 \times 2.5 = 5$.” Note that for the HT function no further “special function” evaluations are required compared to the SDR.

For a first idea of the computational speed of the various line shapes we count the number of divisions and square roots, i.e. “special function” evaluations in the sense of Goedecker and Hoisie [42]. Assuming a single rational approximation for the complex error function (Humlíček’s `cpf16` in the following) valid on the whole complex plane one division is required for the Voigt function. Table 2 (second column) indicates that for the SDV and SDR functions a factor of four to five penalty compared to the Voigt function can be expected.

For the timings reported in Tab. 2 (last column) a somewhat more realistic scenario is considered: For line-by-line (lbl) molecular cross sections the contribution of a single spectral transition is usually computed in a finite interval with a fixed wavenumber cutoff (for example up to distances of 25 cm^{-1} from the line center) or a cut at a multiple of the line half width. Here a uniform x -grid from 0 to $100x_{1/2}$ and a grid point spacing defined by a

SDR with $y/q = 10$ and $y/\zeta = 10$

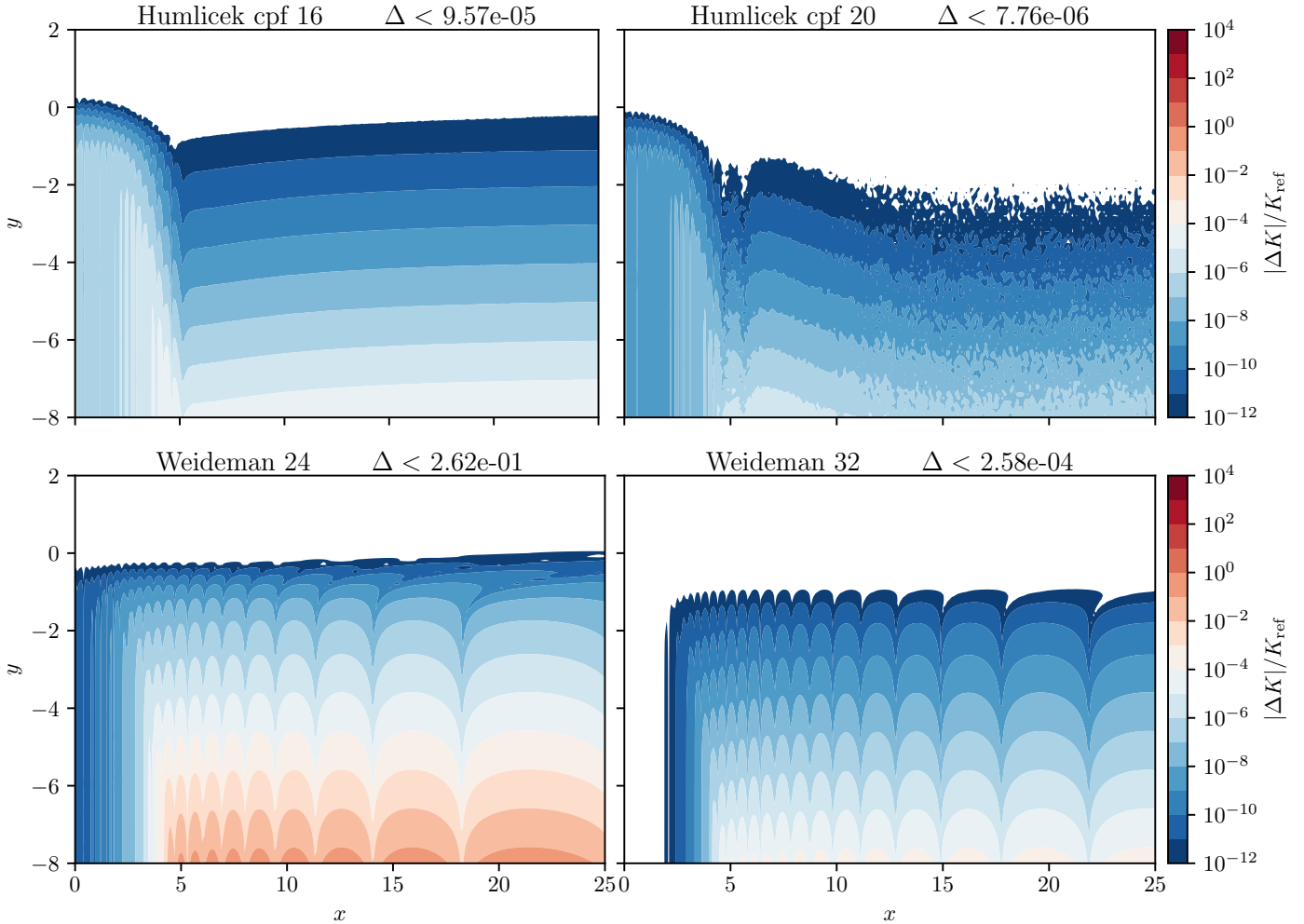


Figure 7: Contour plot of relative differences of the SDR function computed with the Humlíček (top) and Weideman complex error function codes compared to the `wofz` reference code.

fifth of the half width is used. The results demonstrate that the Rautian function is only slightly slower than the Voigt function, and the SDV and SDR are roughly a factor two slower, in contrast to the estimate based on “special function” counts.

These results suggest that the additions and multiplications should also be considered. In fact, with the Horner scheme 23 add-multiply operations are required to compute a Voigt function value with the Humlíček [34] rational approximation generalized to 16 terms [24]; for the SDV and SDR the number of extra multiplications (in addition to the 46 multiplications for the two `cpf16` calls) is small (Tab. 2, third column).

Finally we compute cross sections

$$k(\nu, p, T) = \sum_l S_l(T) g(\nu - \hat{\nu}_l, \gamma_l^{(L)}(p, T), \gamma_l^{(G)}(T), \dots) \quad (13)$$

in the SWIR with line data (position $\hat{\nu}_l$, strength S_l and broadening parameters) taken from the SEOM-IAS

Table 2: Execution time for Voigt, Rautian, SDV, and SDR functions. The second and third columns give the number of “special functions” and multiplications required. The `cpf16` rational approximation of the complex error function is used for all functions.

function	sp. fct.	mult.	time
Voigt	1	23	9.04 ms
Rautian	2	24	9.49 ms
SDV	4	49	19.4 ms
SDR	5	50	19.8 ms

database. The SWIR channels of TROPOMI [12] are exploited for remote sensing of atmospheric carbon monoxide (CO) and methane (CH₄). For the retrieval of CO vertical column densities with the Beer InfraRed Retrieval Algorithm (BIRRA) a spectral interval 4277–4303 cm⁻¹ is used [18, 43], and in the following we report execution times for cross sections evaluated for a 20 level mid latitude summer atmosphere (with top of atmosphere at 100 km, [44]). For consistency with the continuum cor-

Table 3: Execution time (seconds per line) for lbl molecular cross sections in the SWIR. In the first block the total number of lines in 4250–4330 cm⁻¹ and the number of lines with “beyond Voigt” parameters is given. For methane 62 lines have both non-zero speed-dependence and Dicke narrowing parameter. In the second block times for the `cpf16` complex error function code are listed, in the third block times for the `hum1cpf16` code.

function	CO	CH ₄	H ₂ O	total
lines	117	6516	1262	7795
$\gamma_2 > 0$	10	637	39	
$\nu_{vc} > 0$	10	66	113	
Voigt	0.0521	0.0411	0.0417	0.0413
Rautian	0.0545	0.0436	0.0442	0.0438
SDV	0.0580	0.0466	0.0437	0.0463
SDR	0.0599	0.0462	0.0457	0.0463
Voigt	0.0092	0.0068	0.0069	0.0069
Rautian	0.0121	0.0090	0.0090	0.0090
SDV	0.0126	0.0097	0.0076	0.0094
SDR	0.0157	0.0122	0.0100	0.0119

rections and to allow a convolution of the monochromatic spectrum with the instrumental line shape function an extended spectral interval is considered. The cross sections are computed using the `lb12xs` function of the Py4CATS package [PYthon for Computational ATMospheric Spectroscopy, 45].

The results listed in Tab. 3 confirm that the Rautian line shape is slightly more expensive w.r.t. computational speed. SDV and SDR times are approximately identical and larger than the Voigt times. The modest time lag incurred by the speed-dependent profiles appears to be reasonable because only a small fraction of the lines has γ_2 and/or ν_{vc} defined, i.e. the standard Voigt profile can be used for the majority of weak lines.

The high quality of the Humlíček [34] 16-term rational approximation is not required in the line wings, and further tests have therefore been conducted with a combination of this `cpf16` code and the Humlíček [35] asymptotic rational approximation for large complex error function arguments $|z| > 15$ (termed `hum1zpf16` in [24]). This combined code provides a significant speed-up for all line shapes, with the relative performance (Voigt vs. speed-dependent profiles) being similar to the previous tests (Tab. 3 third block).

4. Summary

Several line shapes have been recommended as a generalization of the standard Voigt function to model subtle broadening effects. Various computational aspects of the SDV considering the speed-dependence of the line broadening had been discussed in a recent paper [22] and here we continue this study including the more advanced SDR that also models collisional narrowing effects. First we present a brief survey of data available in the widely used HITRAN database [21] and in the new SEOM-IAS database [20] including a discussion of the parameter range to be expected.

Next we demonstrate that for argument values possible in the upper atmosphere the SDV and SDR function values might suffer from cancellation errors; these numerical problems can be avoided by a simple reformulation, thus eliminating the need for case selections and Taylor expansions etc. Furthermore, the impact of various complex error function algorithms is studied: the Humlíček rational approximation generalized to 16 terms or the Weideman 32-term approximation deliver SDV and SDR values with four or more significant digits. (Note that in databases such as HITRAN, GEISA, or SEOM-IAS the line parameters are listed with a few digits accuracy only.) Finally an assessment of the execution speed is provided. First we discuss the number of floating point operations required for the various line shape functions: because of the extra divisions and square roots both SDV and SDR are computationally more expensive compared to the Voigt function. However, for “real world” applications (molecular cross sections) the impact on the total code execution time is expected to be small.

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References

- [1] B.H. Armstrong. Spectrum line profiles: The Voigt function. *J. Quant. Spectrosc. & Radiat. Transfer*, 7:61–88, 1967. doi: 10.1016/0022-4073(67)90057-X.
- [2] R. H. Dicke. The effect of collisions upon the Doppler width of spectral lines. *Phys. Rev.*, 89:472–473, 1953. doi: 10.1103/PhysRev.89.472.
- [3] P.L. Varghese and R.K. Hanson. Collisional narrowing effects on spectral line shapes measured at high resolution. *Appl. Opt.*, 23(14):2376–2385, 1984. doi: 10.1364/AO.23.002376.
- [4] J.M. Hartmann, C. Boulet, and D. Robert. *Collisional Effects on Molecular Spectra*. Elsevier, 2008.
- [5] J. Tennyson, P.F. Bernath, A. Campargue, A.G. Császár, L. Daumont, R.R. Gamache, J.T. Hodges, D. Lisak, O.V. Naumenko, L.S. Rothman, H. Tran, N.F. Zobov, J. Buldyreva, C.D. Boone, M.D. De Vizia, L. Gianfrani, J.-M. Hartmann, R. McPheat, D. Weidmann, J. Murray, N.H. Ngo, and O.L. Polyansky. Recommended isolated-line profile for representing high-resolution spectroscopic transitions (IUPAC technical report). *Pure Appl. Chem.*, 86(12):1931–1943, 2014. doi: 10.1515/pac-2014-0208.
- [6] H. Tran, N.H. Ngo, and J.-M. Hartmann. Efficient computation of some speed-dependent isolated line profiles. *J. Quant. Spectrosc. & Radiat. Transfer*, 129:199 – 203, 2013. doi: 10.1016/j.jqsrt.2013.06.015. Erratum: JQSRT 134, 104 (2014).

Table 4: Numerical values for $y = 10^{-8}$ and $q = 10^{-9}$, hence $\alpha = 8.5$ and $\delta \equiv Y = 2.5 \cdot 10^{17}$. The first and fourth row correspond to the “naive” version, the second and fifth row to the “safe” version” and the reference values in the third and sixth row are obtained by mpmath.

x	$\beta = \text{Im}(X)$	$\text{Re}(\sqrt{X+Y})$	$\text{Re}(iz_-)$	$\text{Re}(w(iz_-))$	K_{adv}
10	5000000000.0	500000000.0000001	1.1920928955078125e-07	6.829164078387363e-10	1.187268242909783e-10
	5000000000.0	500000000.0000001	1.0849999999999996e-07	6.215659511546557e-10	5.737636760689775e-11
	5000000000.0	500000000.000001085	0.0000001084999999999994482775	6.215658550296477893431e-10	5.7376254492165825521444e-11
12	6000000000.0	500000000.0000001	1.1920928955078125e-07	4.720118241361795e-10	-9.217775941157818e-11
	6000000000.0	500000000.0000001	1.5249999999999992e-07	6.038271231846435e-10	3.9637539636885786e-11
	6000000000.0	500000000.000001525	0.0000001524999999999988890375	6.03827085623295707495819e-10	3.96375257362927384521548e-11

- [7] M. Schneider and F. Hase. Improving spectroscopic line parameters by means of atmospheric spectra: Theory and example for water vapor and solar absorption spectra. *J. Quant. Spectrosc. & Radiat. Transfer*, 110(17):1825 – 1839, 2009. doi: 10.1016/j.jqsrt.2009.04.011.
- [8] M. Schneider, F. Hase, J.-F. Blavier, G.C. Toon, and T. Leblanc. An empirical study on the importance of a speed-dependent Voigt line shape model for tropospheric water vapor profile remote sensing. *J. Quant. Spectrosc. & Radiat. Transfer*, 112(3):465–474, 2011. doi: 10.1016/j.jqsrt.2010.09.008.
- [9] C.D. Boone, K.A. Walker, and P.F. Bernath. Speed-dependent Voigt profile for water vapor in infrared remote sensing applications. *J. Quant. Spectrosc. & Radiat. Transfer*, 105:525–532, 2007. doi: 10.1016/j.jqsrt.2006.11.015.
- [10] A. Kuze, H. Suto, M. Nakajima, and T. Hamazaki. Thermal and near infrared sensor for carbon observation Fourier-transform spectrometer on the Greenhouse Gases Observing Satellite for greenhouse gases monitoring. *Appl. Opt.*, 48(35):6716–6733, 2009. doi: 10.1364/AO.48.006716.
- [11] D. Crisp, R.M. Atlas, F.-M. Breon, L.R. Brown, J.P. Burrows, P. Ciaia, B.J. Connor, S.C. Doney, I.Y. Fung, D.J. Jacob, C.E. Miller, D. O’Brien, S. Pawson, J.T. Randerson, P. Rayner, R.J. Salawitch, S.P. Sander, B. Sen, G.L. Stephens, P.P. Tans, G.C. Toon, P.O. Wennberg, S.C. Wofsy, Y.L. Yung, Z. Kuang, B. Chudasama, G. Sprague, B. Weiss, R. Pollock, D. Kenyon, and S. Schroll. The orbiting carbon observatory (OCO) mission. *Adv. Space Res.*, 34(4):700 – 709, 2004. doi: 10.1016/j.asr.2003.08.062.
- [12] J.P. Veefkind, I. Aben, K. McMullan, H. Förster, J. de Vries, G. Otter, J. Claas, H.J. Eskes, J.F. de Haan, Q. Kleipool, M. van Weele, O. Hasekamp, R. Hoogeveen, J. Landgraf, R. Snel, P. Tol, P. Ingmann, R. Voors, B. Kruizinga, R. Vink, H. Visser, and P.F. Levelt. TROPOMI on the ESA Sentinel-5 Precursor: A GMES mission for global observations of the atmospheric composition for climate, air quality and ozone layer applications. *Remote Sensing of Environment*, 120:70 – 83, 2012. doi: 10.1016/j.rse.2011.09.027. The Sentinel Missions - New Opportunities for Science.
- [13] A.V. Nikitin, O.M. Lyulin, S.N. Mikhailenko, V.I. Perevalov, N.N. Filippov, I.M. Grigoriev, I. Morino, T. Yokota, R. Kumazawa, and T. Watanabe. GOSAT-2009 methane spectral line list in the 5550 – 6236 cm^{-1} range. *J. Quant. Spectrosc. & Radiat. Transfer*, 111(12-13):2211–2224, 2010. doi: 10.1016/j.jqsrt.2010.05.010.
- [14] A.V. Nikitin, O.M. Lyulin, S.N. Mikhailenko, V.I. Perevalov, N.N. Filippov, I.M. Grigoriev, I. Morino, Y. Yoshida, and T. Matsunaga. GOSAT-2014 methane spectral line list. *J. Quant. Spectrosc. & Radiat. Transfer*, 154:63 – 71, 2015. doi: 10.1016/j.jqsrt.2014.12.003.
- [15] F. Oyafuso, V.H. Payne, B.J. Drouin, V.M. Devi, D.C. Benner, K. Sung, S. Yu, I.E. Gordon, R. Kochanov, Y. Tan, D. Crisp, E.J. Mlawer, and A. Guillaume. High accuracy absorption coefficients for the Orbiting Carbon Observatory-2 (OCO-2) mission: Validation of updated carbon dioxide cross-sections using atmospheric spectra. *J. Quant. Spectrosc. & Radiat. Transfer*, 203:213 – 223, 2017. doi: 10.1016/j.jqsrt.2017.06.012. HITRAN2016 Special Issue.
- [16] A. Galli, S. Guerlet, A. Butz, I. Aben, H. Suto, A. Kuze, N. M. Deutscher, J. Notholt, D. Wunch, P. O. Wennberg, D. W. T. Griffith, O. Hasekamp, and J. Landgraf. The impact of spectral resolution on satellite retrieval accuracy of CO_2 and CH_4 . *Atmos. Meas. Tech.*, 7(4):1105–1119, 2014. doi: 10.5194/amt-7-1105-2014.
- [17] R. Checa-Garcia, J. Landgraf, A. Galli, F. Hase, V.A. Velasco, H. Tran, V. Boudon, F. Alkemade, and A. Butz. Mapping spectroscopic uncertainties into prospective methane retrieval errors from Sentinel-5 and its precursor. *Atmos. Meas. Tech.*, 8(9):3617–3629, 2015. doi: 10.5194/amt-8-3617-2015.
- [18] P. Hochstaffl and F. Schreier. Impact of molecular spectroscopy on carbon monoxide abundances from SCIAMACHY. *Remote Sensing*, 12(7):1084, 2020. doi: 10.3390/rs12071084.
- [19] P. Hochstaffl and F. Schreier. Impact of molecular spectroscopy on carbon monoxide abundances from TROPOMI. *Remote Sensing*, October 2020. Manuscript accepted subject to minor revisions.
- [20] M. Birk, G. Wagner, J. Loos, D. Mondelain, and A. Campargue. ESA SEOM-IAS - spectroscopic parameters database 2.3 μm region [data set]. Zenodo, 2017. URL <http://doi.org/10.5281/zenodo.1009126>.
- [21] I.E. Gordon, L.S. Rothman, et al. The HITRAN2016 molecular spectroscopic database. *J. Quant. Spectrosc. & Radiat. Transfer*, 203:3 – 69, 2017. doi: 10.1016/j.jqsrt.2017.06.038.
- [22] F. Schreier. Computational aspects of speed-dependent Voigt profiles. *J. Quant. Spectrosc. & Radiat. Transfer*, 187:44–53, 2017. doi: 10.1016/j.jqsrt.2016.08.009.
- [23] F. Schreier. Optimized implementations of rational approximations for the Voigt and complex error function. *J. Quant. Spectrosc. & Radiat. Transfer*, 112(6):1010–1025, 2011. doi: 10.1016/j.jqsrt.2010.12.010.
- [24] F. Schreier. The Voigt and complex error function: Humlíček’s rational approximation generalized. *Mon. Not. Roy. Astron. Soc.*, 479(3):3068–3075, 2018. doi: 10.1093/mnras/sty1680.
- [25] J. Loos, M. Birk, and G. Wagner. Measurement of air-broadening line shape parameters and temperature dependence parameters of H_2O lines in the spectral ranges 1850–2280 cm^{-1} and 2390–4000 cm^{-1} . *J. Quant. Spectrosc. & Radiat. Transfer*, 203:103 – 118, 2017. doi: 10.1016/j.jqsrt.2017.03.033.
- [26] J. Loos, M. Birk, and G. Wagner. Measurement of positions, intensities and self-broadening line shape parameters of H_2O lines in the spectral ranges 1850–2280 cm^{-1} and 2390–4000 cm^{-1} . *J. Quant. Spectrosc. & Radiat. Transfer*, 203:119 – 132, 2017. doi: 10.1016/j.jqsrt.2017.02.013.
- [27] P. Wcislo, I.E. Gordon, H. Tran, Y. Tan, S.-M. Hu, A. Campargue, S. Kassi, D. Romanini, C. Hill, R.V. Kochanov, and L.S. Rothman. The implementation of non-Voigt line profiles in the HITRAN database: H_2 case study. *J. Quant. Spectrosc. & Radiat. Transfer*, 177:175–191, 2016. doi: 10.1016/j.jqsrt.2016.01.024.
- [28] R.R. Gamache and B. Vispoel. On the temperature dependence of half-widths and line shifts for molecular transitions in the microwave and infrared regions. *J. Quant. Spectrosc. & Radiat. Transfer*, 217:440 – 452, 2018. doi: 10.1016/j.jqsrt.2018.05.019.
- [29] N. Stolarczyk, F. Thibault, H. Cybulski, H. Jóźwiak, G. Kowzan, B. Vispoel, I.E. Gordon, L.S. Rothman, R.R. Gamache, and P. Wcislo. Evaluation of different parameterizations of temperature dependences of the line-shape parameters based on ab initio calculations: Case study for the HITRAN

- database. *J. Quant. Spectrosc. & Radiat. Transfer*, 240:106676, 2020. doi: 10.1016/j.jqsrt.2019.106676.
- [30] M. Abramowitz and I.A. Stegun. *Handbook of Mathematical Functions*. National Bureau of Standards, AMS55, New York, 1964.
- [31] F.W.J. Olver, D.W. Lozier, R.F. Boisvert, and C.W. Clark, editors. *NIST Handbook of Mathematical Functions*. Cambridge University Press, New York, NY, 2010. Print companion to [32].
- [32] DLMF. *NIST Digital Library of Mathematical Functions*. National Institute of Standards and Technology. URL <http://dlmf.nist.gov/>. Online companion to [31].
- [33] C.D. Boone, K.A. Walker, and P.F. Bernath. An efficient analytical approach for calculating line mixing in atmospheric remote sensing applications. *J. Quant. Spectrosc. & Radiat. Transfer*, 112(6):980 – 989, 2011. doi: 10.1016/j.jqsrt.2010.11.013.
- [34] J. Humlíček. An efficient method for evaluation of the complex probability function: the Voigt function and its derivatives. *J. Quant. Spectrosc. & Radiat. Transfer*, 21:309–313, 1979. doi: 10.1016/0022-4073(79)90062-1.
- [35] J. Humlíček. Optimized computation of the Voigt and complex probability function. *J. Quant. Spectrosc. & Radiat. Transfer*, 27:437–444, 1982. doi: 10.1016/0022-4073(82)90078-4.
- [36] J.A.C. Weideman. Computation of the complex error function. *SIAM J. Num. Anal.*, 31:1497–1518, 1994. doi: 10.1137/0731077.
- [37] Fredrik Johansson et al. *mpmath: a Python library for arbitrary-precision floating-point arithmetic (version 0.18)*, December 2013. <http://mpmath.org/>.
- [38] P. Panchekha, A. Sanchez-Stern, J.R. Wilcox, and Z. Tatlock. Automatically improving accuracy for floating point expressions. In *PLDI '15: Proceedings of the 36th ACM SIGPLAN Conference on Programming Language Design and Implementation*. Association for Computing Machinery, 2015. doi: 10.1145/10.1145/2737924.2737959. URL <https://herbie.uwplse.org/pldi15-paper.pdf>.
- [39] S.G. Johnson. Faddeeva package, Last access August 2020. URL <http://ab-initio.mit.edu/Faddeeva>.
- [40] S.G. Johnson and J. Wuttke. libcerf, numeric library for complex error functions, Last access August 2020. URL <https://jugit.fz-juelich.de/mlz/libcerf>.
- [41] E.E. Whiting. An empirical approximation to the Voigt profile. *J. Quant. Spectrosc. & Radiat. Transfer*, 8:1379–1384, 1968. doi: 10.1016/0022-4073(68)90081-2.
- [42] Stefan Goedecker and Adolfo Hoisie. *Performance Optimization of Numerically Intensive Codes*. SIAM, Philadelphia, PA, 2001.
- [43] S. Gimeno García, F. Schreier, G. Lichtenberg, and S. Slijkhuis. Near infrared nadir retrieval of vertical column densities: methodology and application to SCIAMACHY. *Atmos. Meas. Tech.*, 4(12):2633–2657, 2011. doi: 10.5194/amt-4-2633-2011.
- [44] G.P. Anderson, S.A. Clough, F.X. Kneizys, J.H. Chetwynd, and E.P. Shettle. AFGL atmospheric constituent profiles (0 – 120 km). Technical Report TR-86-0110, AFGL, 1986.
- [45] F. Schreier, S. Gimeno García, P. Hochstaffl, and S. Städt. Py4CAAtS — PYthon for Computational ATmospheric Spectroscopy. *Atmosphere*, 10(5):262, 2019. doi: 10.3390/atmos10050262.

Appendix A. Example of the SDV calculation for small values of y and q

To further illustrate the discussion of subsection 3.2 a comparison for two values of x and for $y = 10^{-8}$ and $q = 0.1y$ is given in Tab. 4. The numbers shown here are copied from the (I)Python shell output without any

formatting of the `print` statement. Note that for the `mpmath` calculations shown here a “triple precision” has been exploited (`mp.dps=24`), in contrast to the quadruple precision used for Fig. 4.

Appendix B. Code snippets used for time benchmarks

For the first test assessing the performance of the various ways to evaluate the complex square root and the difference thereof the SDV has been computed on the “matrix” of x and y values used for all contour plots (note that Python is case sensitive):

```
In [1]: from sdv import *
In [2]: # standard Voigt variables
...: xx=np.linspace(0.,25.,251) # linear grid
...: yy=np.logspace(-8,2,101) # logarithmic grid
...: # kind of 'matrices' required for contour plots
...: XX, YY = np.meshgrid(xx, yy)
...: ZZ = XX +1j*YY
In [3]: %timeit sdv_safe(XX, YY, 0.1*YY, cpf16p)
```

In the second test the SDR has been calculated for the 101 y values as defined above with an equidistant x grid in the interval defined by the Voigt function half width:

```
xx=np.linspace(0.,25.,251)
yy=np.logspace(-8,2,101)
%%timeit
for y in yy:
    xHalf = whiting(y); xx=arange(0.0,100*xHalf,xHalf/5)
    sdr(xx, y, 0.1*y, 0.1*y, cpf16p)
```