

2D Derivative Phase Induced Growth of 3D All Inorganic Perovskite Micro Nanowire Array **Based Photodetectors**

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13

Abstract: A large number of derivative phases in inorganic perovskites have been 14 reported with special structures and shown extraordinary performances in 15 photoelectronic device applications. The reverse phase transition between derivative 16 phases and perovskites usually induces recrystallization or forms mixed components. 17 In this work, we report derivative phase induced growth of the CsPbBr₃ micro-nanowire 18 (MW) array by utilizing phase transition of the two-dimensional CsPb₂Br₅ phase. 19 Owning to its layered structure and phase transition, annealing of CsPb₂Br₅ at a 20 temperature of 550 °C combined with solvent quenching led to a templating effect to 21 form a high-quality CsBr MW array. Subsequent PbBr₂ deposition and the second 22 annealing were employed to form the aligned CsPbBr₃ MW arrays. Based on this 23 method, a CsPbBr₃ MW array based photodetector was fabricated. The large grain size, 24 less grain boundaries (GBs) and lower surface potential of the CsPbBr₃ MW array led 25 to high device performance with a responsivity of 7.66 A W⁻¹, a detectivity of $\sim 10^{12}$ 26 Jones and long-term operational stability over 1900 min. 27

28

29 Introduction

Inorganic halide perovskite materials CsPbX₃ (X=Cl, Br and I) have attracted great attention because of their applications in solar cells, light emitting diodes (LEDs), lasers and photodetectors (PDs).^[1-4] Their unique properties such as tunable bandgap, high carrier mobility, long carrier diffusion length and high absorption coefficient make them strong candidates for optoelectrical device applications.^[5-8] To date, the development of perovskites has evolved from initially three/two-dimensional (3D/2D) bulk films^{[9,}

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^{10]} to low dimensional nano-structures, for example, one-dimensional (1D)
 nanowires/microwires (NWs/MWs)^[11, 12] and zero-dimensional (0D) quantum dots^[13, 3]
 ^{14]}.

In comparison with bulk films and quantum dots with many small grains and long-4 chain ligands, the 1D NWs/MWs exhibit well-aligned structures, lower lateral 5 dimension, less GBs and confined carrier transport in 1D channels/structures, which 6 leads to high crystallinity, less recombination and superior electronic transport 7 behaviors.^[15] Additionally, multiple synthesis methods making use of forward growth 8 have been developed such as the solution method^[11, 16] (hot-injection,^[17] anion 9 exchange^[18] etc.) and chemical vapor deposition (CVD).^[19, 20] The solution process is 10 a facile, low-cost technique, and has been employed in fabrication of MAPbI₃ 11 perovskite nanowires, because of the low synthesis temperature and a precise ratio of 12 precursor materials.^[2, 21] However, the existence of residue organic solvents such as 13 DMF, DMSO will affect the quality of perovskites and accelerate the 14 decomposition/degradation of perovskite materials.^[22-24] In contrast, the vapor 15 deposition enables high purity of the materials and easy controllability without the 16 adverse effects of residue solvents to ensures long-term stability.^[2, 25, 26] In addition, 17 low solubility of all-inorganic perovskites is difficult due to the lower solubility of 18 inorganic perovskites and more complex Cs-Pb-X phase diagram.^[21, 23] In parallel, the 19 vapor deposition technique (e.g., all vacuum deposition, CVD) provides a non-solvent 20 environment and high temperature/vacuum system, which allows preparation of highly 21 crystalline NWs without the impurity phase.^[21, 27] These strategies have shown showing 22 great advantages in light sources and photodetectors. For example, Wang, Pan and 23 coworkers developed wavelength-tunable in-plane aligned CsPbX₃ perovskite NWs 24 lasers with a low pumping threshold of 4 µJ cm⁻².^[28] Moreover, Meng and coworkers 25 reported a vapor-liquid-solid growth technique to grow high quality CsPbX₃ NWs and 26 realized PDs in the visible light regime with an impressive responsivity of 4489 A W⁻¹ 27 and detectivity over 7.9×10^{12} Jones.^[29] Although perovskite NWs have shown 28 promising performance in optoelectronic devices, random dispersion and short length 29 make it difficult to integrate them into optoelectronic devices. Recently, NW arrays 30

were proposed to solve this issue because of their controlled alignment and high 1 uniformity in large area. Feng and coworkers reported a 1D CsPbBr₃ perovskite NW 2 array based photodetector with a high responsivity exceeding 1000 A W⁻¹ by using a 3 micropillar-structured template with asymmetric wettability via the solution method.^[11] 4 In addition, Waleed and coworkers reported a vertical CsPbI₃ NW array based 5 photodetector with a responsivity of 6.7 mA W⁻¹ by combining chemical vapor 6 deposition (CVD) with anodic alumina membrane (AAM) templates.^[30] It is worth 7 noting that the formation of NW arrays usually depends on the templates such as 8 AAO/AAM templates^[30, 31], CD/DVD disks^[32, 33] and photoetching^[34] or special 9 substrates like M-template substrates^[21, 28] and asymmetric wettability substrates^[11, 35]. 10 Use of these templates not only leads to increased cost but also makes it difficult to 11 upscale. A facile and template-free method is desirable for upscaling of NW arrays. 12

Herein, we propose a two-step method of facile phase induced growth of CsPbBr₃ 13 MW arrays by using CsBr as the intermediate product and controlling the 14 decomposition of the CsPb₂Br₅ perovskite derivative phase. In the first step, a height of 15 16 1.1 µm CsBr MW arrays is obtained by controlling the decomposition from CsPb₂Br₅ film (CsPb₂Br₅ \rightarrow CsBr+2PbBr₂ \uparrow). Then, the CsBr MW arrays are post-treated with 17 extra PbBr₂, leading to CsPbBr₃ MW arrays with large grain sizes (over 5 µm) and a 18 larger height of 1.72 μ m (CsBr+PbBr₂ \rightarrow CsPbBr₃). By investigating the growth 19 mechanism of CsBr MW and CsPbBr3 MW arrays, we find that the layered structure of 20 21 CsPb₂Br₅^[36] serves as a nominal template, which makes the intermediate product of CsBr MWs possible to grow along the CsPb₂Br₅ template to form arrays. Finally, a 22 high-performance photodetector is achieved with an impressive responsivity of 7.66 A 23 W^{-1} , high detectivity of ~10¹² Jones and long-term operational stability over 1900 min 24

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Figure 1. a) Schematic illustration of formation of CsBr MW arrays and CsPbBr₃ MW arrays.
b) Scanning electron microscope (SEM) image, c) Atomic force microscopy (AFM)
morphology image, d) X-ray diffraction pattern (XRD) and e) energy-dispersive X-ray
spectroscopy (EDS) mapping of CsBr MW arrays.

6

7 Result and discussion

Figure 1a illustrates the growth process of CsBr MW arrays and CsPbBr₃ MW arrays. 8 The CsPb₂Br₅ precursor layer with a thickness of 1 µm is first deposited on the Si/SiO₂ 9 substrate by a sequential vapor deposition method (more details can be found in 10 Experiment Section). Then the CsPb₂Br₅ films are moved to a hotplate (400 °C) rapidly 11 and annealed at a high temperature of 550 °C for 20 min. Solvent quenching is carried 12 out by immersing the samples into an isopropanol (IPA) solution to cool down the 13 14 samples to room temperature quickly in order to maintain the structure of long MWs. The well-aligned CsBr MW arrays with an average length of $\sim 175 \,\mu m$ can be obtained 15 (Figure 1b, S1) as follows: $CsPb_2Br_5 \rightarrow CsBr+2PbBr_2\uparrow$. Notably, energy-dispersive X-16 ray spectroscopy (EDS) mapping in Figure 1e, S2 shows that Cs and Br are uniformly 17 distributed in the MW arrays. The absence of Pb indicates that PbBr₂ has evaporated 18 completely. The Cs:Br composition ratio of 1:1 as seen in the corresponding EDS line 19 scan in Figure S3, S4 agree well with the stoichiometric ratio of CsBr. The high 20

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crystallinity of the CsBr MW arrays is further confirmed by surface investigation. 1 Atomic force microscopy (AFM) topography images of the CsBr MW arrays and 2 individual MW are shown in Figure 1c, S5, S6. The smooth surface with an average 3 height of 1.1 µm and uniform width of 1~2 µm suggest high crystallinity of the CsBr 4 MWs and MW growth of the CsBr MW array according to Ostwald ripening 5 mechanism during the formation, which will be discussed later. The X-ray diffraction 6 7 (XRD) pattern in Figure 1d, S7, S8 indicates that fabricated CsBr MW arrays exhibit 8 preferred orientation along (211).

9





Figure 2. a) SEM image, b) EDS mapping, c) XRD pattern of the CsPbBr₃ MW arrays. d)
Scanning transmission electron microscopy (STEM) image and e) simulated lattice pattern of
the single CsPbBr₃ MW. The inset shows the corresponding Fourier transform pattern. f) AFM
topography image and g) surface potential distribution of the single CsPbBr₃ MW. h) Time
resolved photoluminescence (TRPL) decay curve of the CsPbBr₃ MW arrays.

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17 Based on the as-synthesized CsBr MW arrays, the CsPbBr₃ MW arrays were

obtained (Figure 2a) after the post-treatment with an extra PbBr₂ film (Figure S9) 1 followed by annealing (CsBr+PbBr₂ \rightarrow CsPbBr₃), which show uniform distribution and 2 silkworm-like grains. The average grain size is over 5 µm with the maximal size up to 3 11 μ m (Figure S10). The EDS element map (Figure 2b, S11) shows that the MW arrays 4 are composed mainly of the Cs, Pb and Br elements. To extract the precise element ratio, 5 the EDS line scan was performed (Figure S12, S13) and the corresponding element 6 ratio of Cs, Pb and Br is close to 1:1:3, which agrees well with the stoichiometric ratio 7 of the CsPbBr₃ phase. In addition, the absence of Pb and Br in the gap between the 8 MWs (Figure S14) indicates that the complete evaporation of extra PbBr₂. Furthermore, 9 crystal structure characterization of MW arrays was conducted by XRD. The peaks at 10 15.19°, 21.55°, 26.48°, 30.64°, 34.37°, 37.77°, 43.89°, 46.69°, 49.38° and 54.51° in 11 Figure 2c can be assigned as (100), (110), (111), (200), (210), (211), (220), (300), (310) 12 and (222) diffraction peaks, which indicate that the cubic CsPbBr₃ perovskite (Figure 13 S8) is formed after the post-treatment by extra PbBr₂. Note that CsPbBr₃ MW arrays 14 exhibit a polycrystalline structure without oriented growth. To gain more insight into 15 16 the component and structure of the CsPbBr₃ MWs, scanning transmission electron microscopy (STEM) measurements were conducted. The STEM characterization and 17 corresponding simulation images in Figure 2d, 2e further demonstrate the crystalline 18 character of CsPbBr₃, which is indexed to be the [111] crystal zone of the cubic CsPbBr₃ 19 phase. AFM topography images in Figure 2f, S15 illustrate that the CsPbBr₃ MW has 20 a height of 1.72 µm and width of MWs is close to 6 µm. Large grain size and high 21 crystallinity of CsPbBr₃ grains can effectively reduce recombination at the grain 22 boundaries (GBs). To get a clear comparison with our CsPbBr₃ MW arrays, a CsPbBr₃ 23 film with a thickness of 1.79 µm was fabricated on the Si/SiO₂ substrates (Figure S16). 24 Kelvin probe force microscopy (KPFM) measurements were carried out (Figure 2g, 25 S17). In comparison with a conventional CsPbBr₃ film that shows an uneven surface 26 potential distribution in the film (Figure S18) under light illumination, the CsPbBr₃ 27 MW shows a uniform surface potential distribution. This difference can be ascribed to 28 the large and uniform perovskite grain in the MWs, which shows a reduced potential 29 barrier at the GBs and thus is expected to promote carrier transport at GBs as seen in 30

Figure S17b. In contrast, the uneven grain size distribution indicates that overgrown 1 grains exist in the perovskite film (Figure S18c), which can result in carrier 2 accumulation at the GBs because of existence of the potential barrier. Carrier 3 accumulation at GBs would significantly lower the valence and conduction band of the 4 perovskite films as illustrated in Figure S19, which induces carrier recombination at 5 the GBs.^[37, 38] Furthermore, the steady-state photoluminescence (PL) measurements 6 were carried out for CsPbBr₃ MW arrays and films. As seen in Figure S20, the 7 8 enhanced PL intensity of CsPbBr₃ MW arrays can be attributed to high quality, less defects and reduced non-radiative recombination in the perovskite grains.^[23] The 9 lifetime of charge transport was investigated by time resolved photoluminescence 10 (TRPL). By fitting with the following equation:^[39, 40] 11

12
$$\tau_{ave} = \frac{A_1 \tau_1^2 + A_2 \tau_2^2}{A_1 \tau_1 + A_2 \tau_2}$$

where A_1 and A_2 are the amplitudes, τ_1 and τ_2 represent non-radiative recombination and radiative recombination, respectively. The average lifetime is up to 84.14 ns (**Figure 2h**), which is much longer than the CsPbBr₃ film (36.38 ns, **Figure S18d**), indicating that the traps in MWs is effectively suppressed by reducing the GBs and obtaining a large grain size.

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Figure 3. a) XRD patterns and b) corresponding SEM images showing the conversion of the
CsPb₂Br₅ precursor layer to the CsBr MW array at different temperatures. c) Schematic
illustration of formation of the CsBr MW arrays and CsPbBr₃ MW arrays.

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1

How to understand the formation of CsBr and CsPbBr₃ MW arrays? We first 6 investigated the inverse and forward reactions in CsPb₂Br₅. As mentioned above, CsBr 7 reacted with PbBr₂ in a molar ratio of 1:2 to form CsPb₂Br₅. In contrast, CsPb₂Br₅ 8 would easily decompose to CsPbBr3 or CsBr because of the thermal instability at the 9 temperature above 300 °C.^[41] As seen in Figure 3a, 3b, the pristine CsPb₂Br₅ film 10 shows high crystallinity and full coverage (Figure 3b, Pristine). Perovskite grains would 11 quickly melt on the hotplate at 400 °C and grow into large ones in order to reduce the 12 surface energy (Figure 3b). Meanwhile, part of the CsPb₂Br₅ phase decomposed to 13 CsPbBr₃ with evaporation of PbBr₂ because of the low melting point of PbBr₂ (357 °C) 14 (Figure 3a, 400 °C).^[37] This procedure can be explained by the following equation: 15 $CsPb_2Br_5 \rightarrow CsPbBr_3 + PbBr_2\uparrow$. As the annealing temperature was increased to 500 °C, 16 the grains would grow larger according to the Ostwald ripening mechanism,^[42] and 17

continuously decompose at such a high temperature. Due to the layered structure of 1 CsPb₂Br₅, the outmost layers would decompose faster than the inner layers (Figure 3c, 2 step 3), which results in stepped structured grains (Figure 3b, S21, 500 °C). When the 3 treated temperature was further increased to 550 °C, the Cs₄PbBr₆ peaks could be 4 detected in the XRD patterns, which suggests the continuous decomposition of CsPbBr₃ 5 follows: $4CsPbBr_3 \rightarrow Cs_4PbBr_6+3PbBr_2 \uparrow$. In addition, the Cs_4PbBr_6 MW arrays 6 (Figure S22) could be formed based on the CsPb₂Br₅ layer template. Prolonged 7 8 annealing time (10 min and 20 min) induced the complete disappearance of Cs₄PbBr₆ phase by further extraction of PbBr₂ (Cs₄PbBr₆ \rightarrow 4CsBr + PbBr₂ \uparrow) and the 9 intermediate products of CsBr MW arrays could be obtained because of the high 10 melting point of CsBr over 660 °C and evaporation of PbBr₂ (Figure 3b and 3c, step 11 4). This decomposition procedure makes the CsBr MW possible to keep the layered 12 structure to construct arrays after fast evaporation of the low melting point component 13 (PbBr₂). Although the MWs appeared to grow along a certain direction within a local 14 region of the substrate, on the scale of the whole substrate the MWs grew along different 15 16 directions. As seen in Figure S1, in this specific sample location, some MWs showed the same orientation along the x direction. However, one of the MWs (Figure S1) was 17 grown along the y direction. To realize the CsPbBr₃ MW arrays, extra PbBr₂ was 18 deposited on the surface of the CsBr arrays (Figure S9). Then the as-prepared samples 19 were annealed at a temperature of 350 °C for 30 min to ensure the further reaction 20 between CsBr and PbBr₂ to form highly crystalline CsPbBr₃. Finally, silkworm-like 21 CsPbBr₃ MWs were obtained with an average grain size about 5 µm. 22



Figure 4. a) Schematic illustration and b) responsivity spectrum of the CsPbBr₃ MW arrays
photodetector. c) I-V and d) I-t curves of the CsPbBr₃ MW array based photodetector under
dark and light illumination with different wavelengths. The inset is a photograph of the device.
e) I-V and f) I-t curves of the CsPbBr₃ MW array based photodetector under light illumination
with a wavelength of 400 nm with different power intensities. g) Photocurrent and h)
responsivity/detectivity of the device as a function of incident power intensities. i) Response
time of the photodetector during one cycle.

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The CsPbBr₃ MW array based PDs (Figure 4a) were fabricated by coating 80 nm 10 gold as the electrode using an aperture mask of 60 µm (Figure 4c, inset). The 11 responsivity spectrum of the device (Figure 4b) shows a high photoresponsivity in the 12 visible region, which is consistent with the absorption spectrum of CsPbBr3 in our 13 previous works.^[7, 37] The current-voltage (I-V) curves in Figure 4c show a low dark 14 current of -1.59×10⁻¹⁰ A at a bias voltage of -5 V and photocurrent at a fixed power 15 intensity of 1.2 mW cm⁻² under light illumination with different wavelengths. The high 16 photocurrent of -113.20 nA was achieved at a bias voltage of -5 V under light 17 illumination with a wavelength of 400 nm, which was much higher than the same 18 thickness of CsPbBr₃ film photodetector (-50.76 nA, Figure S23). In parallel, the 19 photocurrent under illumination light of 365 nm and 450 nm are -90.14 nA and -76.06 20 nA, respectively. The responsivity (R) and detectivity (D^{*}) of the PDs were extracted 21

1 from the following equations: [43-45]

$$2 \qquad R = \frac{\Delta I}{P_{\lambda} S} \tag{1}$$

3
$$D^* = R \sqrt{\frac{S}{2qI_{dark}}} = \frac{R\sqrt{BS}}{i_n}$$
 (2)

where ΔI is photocurrent ($\Delta I = I_{light} - I_{dark}$, I_{light} and I_{dark} are the current with light 4 illumination and in dark, respectively). P_{λ} is the light illumination power intensity on 5 the device and S is the active area of the device; q, B and in are the elementary charge, 6 electrical bandwidth of noise measurement and noise current, respectively. The 7 CsPbBr₃ MW arrays can be regarded as a rectangular-shaped structure and the 8 corresponding active area is estimated as follows:^[29] $S = n \times a \times b$, where n is the number 9 of MWs in the device (n=4), and a and b are the width of the MWs (6 µm) and length 10 of the channel (60 µm), respectively. The corresponding responsivity/detectivity of the 11 devices under light illumination of 365 nm, 400 nm and 450 nm were 6.13 A W-12 ¹/3.25×10¹² Jones, 7.66 A W⁻¹/4.05×10¹² Jones and 5.18 A W⁻¹/2.74×10¹² Jones, 13 respectively. Time-dependent curves (I-t) in Figure 4d suggest that high repeatability 14 of the devices under light illumination with different wavelengths. Moreover, the 15 intensity dependent properties of the devices were also investigated by changing the 16 irradiation power intensity from 1.02 to 17.46 mW cm⁻². As the incident power intensity 17 increased, the photocurrent of the device was greatly increased (Figure 4d) because 18 more carriers were excited and separated. The corresponding I-t curves in Figure 4e 19 also exhibited excellent reproducibility even at high power intensity. By fitting the 20 curve in Figure 4f with the formula of $I=P^{\theta}$, where I, P and θ are the photocurrent, 21 power intensity and proportionality constant, respectively, we found that θ was 0.48 for 22 400 nm incident light, which was not close to the ideal value of 1 because θ was related 23 to recombination processes of photo-generated carriers.^[46, 47] This suggests that some 24 traps and recombination centers exist in the device when the devices are operated under 25 a high incident light intensity, but the devices still exhibit a good linearity for photo-26 detection.^[46, 48] In addition, we also conducted the light power intensity investigation at 27 365 nm and 450 nm in Figure S24, S25, which also exhibits an excellent dependence 28 on the power intensity with a good fitting showing a linearity relationship for light 29

detection. Figure S26 depicts the response characteristic under different operational 1 biases from -1 V to -5 V, as the corresponding photocurrent increased from 14.4 nA to 2 367.1 nA, which indicates that the devices could work well under even low bias 3 voltages. Beside responsivity and detectivity, the response time is also an important 4 parameter for photodetectors, representing the photo-sensitivity of the device and can 5 be defined as the photocurrent increase from 10% to 90% (trise) and decrease from 90% 6 to 10% (t_{fall}) of the maximum value.^[49] We chose one cycle from Figure 4f. It is clear 7 8 that the CsPbBr₃ MW array based PD shows a fast response time of 275/550 ms during ON/OFF operation under 400 nm (Figure 4i). 9

10



11

Figure 5. a) Time-dependent response of the CsPbBr₃ MW array based PD after 100 cycles of
operation. b) Operational stability of CsPbBr₃ MW array based PDs under continuous
illumination (@400 nm, -5 V, 17.46 mW cm⁻²).

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Furthermore, we also investigated the operational stability of our PD devices. Different from conventional storage stability under dark condition without bias, the operational stability was conducted under continuous light illumination and constant voltage, which reflected the more realistic performance of the PD devices during

operation. First, we switched ON/OFF by repeatedly turning light (400 nm, 17.46 mW 1 cm^{-2}) on and off under a bias voltage of -5 V (Figure 5a). The device exhibited excellent 2 reproducibility and stability. After 100 cycles, the photocurrent of the device 3 maintained the same value, which yielded a high ON/OFF ratio of $\sim 10^3$. In parallel, 4 continuously light illumination was also performed under 400 nm (~17.46 mW cm⁻²) 5 at a fixed bias voltage of -5 V. The photocurrent of the device dropped rapidly at the 6 7 first 600 min and then maintained at a stable value (Figure 5b). The devices remained 8 60% of the initial photocurrent even after 1900 min, which represents better operational stability than previous reports as seen in Table 1. The high stability performance can 9 be explained by the improved grain crystallinity, larger grain size (about 5 µm), thicker 10 layer (about 1.72 µm) and fewer grain boundaries. Decomposition/degradation of 11 perovskite devices often starts from the grain boundaries, which serve as a pathway for 12 moisture/oxygen penetration and gradually induce the phase transition/decomposition 13 process toward the grain interiors because chemical binding at grain boundaries with 14 low crystallinity is much weaker than that within the crystal domain interior.^[50, 51] On 15 16 the other hand, charge trapping at the GBs induced by light soaking in the moisture condition can trigger the irreversible degradation of perovskite.^[52] In comparison with 17 many GBs in the thin film, high crystallinity and fewer GBs in the CsPbBr₃ MW arrays 18 alleviate moisture attacks and charge accumulation at the GBs under long-term 19 illumination under the moisture condition. In parallel, the outmost layer of the thicker 20 CsPbBr₃ crystallites in the MW arrays naturally protect the inner crystallites from 21 decomposition/degradation.^[50] 22

23 24

 Table 1 Comparison of inorganic perovskite nanowires from previous works and this work.

Material	Formation	Method	Template	Length	R	D*	Operational	Ref
	direction			(µm)	(A/W)	(Jones)	Stability	
							[min]	
CsPbBr ₃ MW	Inverse	Vapor	No	>200	7.66	4.05×10 ¹²	1900	This
arrays								work
CsPbBr ₃ NW	Forward	Vapor	No	<30		_		[53]
arrays								
CsPbBr ₃ NW	Forward	Solution	Yes	>28	1377			[11]
arrays								

\ \ /	II	E	V	11	n	1
V V	11	$-\mathbf{E}$	Ι.	۰v	C	П

CsPbI ₃ NW	Forward	Vapor	Yes	~10	0.0067	1.5×10 ¹²	—	[30]
nanoarrays								
CsPbI ₃ /CsPbBr ₃	Forward	Solution	Yes	>750	0.125		_	[32]
NW arrays								
MAPbI ₃ MW	Forward	Solution	Yes	>400	13.57	5.25×10 ¹²	—	[15]
arrays								
MAPbI ₃ NW	Forward	Solution	Yes	>130	0.012	7.3×10^{12}	10	[33]
arrays								
MAPbI ₃ NW	Forward	Solution	No	>200	4.95	2×10 ¹³		[54]
arrays								
MAPbI ₃ MW	Forward	Solution	No	>650	0.16	1.3×10^{12}	—	[48]
arrays								
MAPbI ₃ NW	Forward	Solution	Yes	>180	6660	6.85×10 ¹²	_	[55]
arrays								
$MAPb(I_{1-x}Br_x)_3$	Forward	Solution	Yes	>230	12500	1.73×10 ¹¹	_	[34]
NW arrays								
FAPbI ₃ NW	Forward	Solution	Yes	>75	5282	1.45×10^{14}		[35]
arrays								
CsPbBr ₃ film	Forward	Vapor			0.375	2.96×10 ¹¹		[45]
CsPbBr ₃ film	Forward	Solution			0.010	4.56×10 ⁸		[56]
CsPbBr ₃ single	Forward	Solution			2		360	[57]
crystal								
CsPbBr ₃ single	Forward	Solution			0.028			[58]
omistal								

crystal

1 *Operational stability is defined as the stability when the device is under a bias voltage

2 and continuous light illumination.

3

4 Conclusion

In summary, we reported a 2D derivative phase CsPb₂Br₅ induced growth of 3D all 5 6 inorganic perovskite CsPbBr₃ MW arrays. Decomposition of layered structure of 2D 7 CsPb₂Br₅ induces the formation of single-crystalline CsBr MW arrays. The CsBr MW arrays are post-treated by extra PbBr₂ to convert to CsPbBr₃ phase. Impressively, the 8 high performance of a high responsivity of 7.66 A W⁻¹ and long-term operational 9 stability over 1900 min are simultaneously realized in the CsPbBr₃ MW array 10 photodetectors because of the large grain size, fewer grain boundaries and lower surface 11 potential barrier. 12

13

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6	The a	uthors declare no conflict of interest.
7		
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9	stabil	ity
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5	2D derivative phase induced growth of 3D all inorganic perovskite micro-
6	nanowire array based photodetectors
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	$\begin{array}{c} substrate \\ \hline \\ substrate \\ \hline \\ \\ \hline \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ \\ $



A well aligned CsPbBr₃ micro-nanowire (MW) array was synthesized by controlling
growth of the intermediate CsBr MW array originated from the deposition of the 2D
CsPb₂Br₅ phase. Furthermore, a high-performance photodetector was demonstrated
based on the CsPbBr₃ MW array.