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# **MODE** Tides<sup>1</sup>

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#### ABSTRACT

IGPP and AOML bottom pressure measurements at four MODE stations constitute a unique set of deepsea tidal measurements (although deployed for other purposes). A response analysis relative to a Bermuda reference has been optimized with regard to the number of complex weights and the makeup of gravitational and radiational inputs. Duplicate instrumentation on EDIE capsule gave 32.067, 2.5°; 32.074, 2.6° for M<sub>2</sub> amplitude (cm) and Greenwich epoch, thus attesting the reality of measured small station differences (order 1 cm, 1°). M<sub>2</sub> tidal currents (calculated from the M<sub>2</sub> surface and bottom slopes) have u and v speeds of 0.5 and 0.8 cm s<sup>-1</sup>, respectively, in rough agreement (both amplitude and phase) with preliminary estimates from current measurements. M<sub>2</sub> and K<sub>1</sub> tides are in accord with some existing cotidal and co-range charts. M<sub>3</sub> tides are a fraction of equilibrium magnitude, whereas M<sub>4</sub>, M<sub>5</sub> and M<sub>8</sub> (typically 0.07, 0.05, 0.03 cm) vastly exceed equilibrium values. Presumably these overtides are generated by nonlinear coupling in the world's shallow basins, from where they radiate into the global oceans to attain a level where radiative and dissipative processes are somehow balanced.

## 1. Introduction

Measurements of bottom pressure during the MODE (Mid-Ocean Dynamics Experiment) experiment, March to July 1973, were for the purpose of studying mesoscale eddies, and accordingly the emphasis was on periods longer and length scales much shorter than those typical of tides. In fact, one premise underlying the experiment was that the "tidal noise" could be effectively eliminated from the records (Brown *et al.*, 1975). In accomplishing this mission, we have performed a unique experiment of simultaneous deep-sea tide measurements. Information concerning instrumentation and experimental procedures has been described separately (Snodgrass *et al.*, 1975).

#### 2. Bermuda reference

The response method (Munk and Cartwright, 1966) was used, following the procedure by Cartwright *et al.* (1969) for the analysis of relatively short deep-sea records. This involves a two-step analysis: (i) the transfer functions of a reference station relative to the tidal potential, and (ii) the transfer functions of the deep-sea records relative to the reference station.

The 1950-60 tide record at Bermuda was used to derive the transfer functions relative to the input potentials; these transfer functions subsequently served as a basis of a Bermuda tide prediction for the MODE period. Table A (Appendix) shows the results<sup>4</sup> of a combined analysis of three 355-day series equally spaced in a period of lunar perigee (8.85 Julian years).

#### 3. MODE transfer functions

Transfer functions and harmonic constants for the seven deep-sea data series<sup>5</sup> are given in Tables B to H. Table 1 is a summary of principal constituents for the five series exceeding one month (thus excluding MERT and EDIE-MARCH). The transfer functions are based on two complex weights (0, -2 days) for diurnal constituents, and three complex weights  $(0, \pm 2 \text{ days})$  for semidiurnal constituents (see below). "Grav+rad" and "grav only" refer to separate analyses with regard to the nongravitational (radiational) solar effects; we

<sup>&</sup>lt;sup>4</sup> After all of the MODE tide analyses were completed, a small error was found in the analysis of the diurnal tides at Bermuda. Inasmuch as the variation was used consistently in arriving at both the Bermuda harmonic constants and the Bermuda prediction for the MODE period, the accuracy of the MODE tide analyses is not affected. For the record, Table A values and corrected values (in parentheses) are as follows:

	$H \cdot (cm)$	G (deg)
$Q_1$ :	1.13 (1.13)	188.2 (186.6)
O1:	5.30 (5.30)	192.0 (192.1)
<b>P</b> <sub>1</sub> :	2.02 (2.01)	187.7 (187.8)
<b>K</b> 1:	6.56 (6.55)	187.1 (187.0)

<sup>&</sup>lt;sup>6</sup> IGPP (Institute of Geophysics and Planetary Physics, University of California, San Diego) sea-floor pressure series are named REIKO-MAY, MERT, EDIE-MARCH, EDIE-MAY P1, and EDIE-MAY P2. AOML (Atlantic Oceanographic and Meteorological Laboratories, Miami, Fla.) series are named AOML1 and AOML3.

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TABLE 1. Admittances and tidal constants: R is amplitude ratio,  $\phi$  station lead (deg) relative to Bermuda, H amplitude (cm), and G Greenwich epoch (deg). The frequency (cycles per day) is given at the head of each column.

				Grav+rad			Grav	Grav only	
		O <sub>1</sub> 0.930	K <sub>1</sub> 1.003	(N <sub>2</sub> ) 1.895	M <sub>2</sub> 1.932	(S <sub>2</sub> ) 2.000	M <sub>2</sub> 1.932	(S <sub>2</sub> ) 2.000	
R	AOML3	1.142	1.138	0.956	0.943	0.877	0.943	0.752	
	AOML1	1.156	1.180	0.983	0.967	0.877	0.967	0.764	
	REIKO	1.174	1.187	1.003	0.970	0.838	0.969	0.706	
	EDIE-MAY P1	1.172	1.196	0.921	0.9000	0.787	0.900	0.666	
	EDIE-MAY P2	1.180	1.184	0.920	0.9002	0.782	0.900	0.663	
φ	AOML3 AOML1 REIKO EDIE-MAY P1 EDIE-MAY P2	4.9 5.6 4.0 7.9 8.0	-6.8 -7.6 -7.4 -10.3 -10.2	$0.2 \\ -2.1 \\ -1.4 \\ -3.8 \\ -4.0$	-0.6 -2.3 -2.4 -4.2 -4.3	$\begin{array}{r} - & 7.3 \\ - & 6.6 \\ - & 8.5 \\ -10.3 \\ -10.1 \end{array}$	-0.6 -2.3 -2.4 -4.1 -4.3	-9.1 -8.2 -12.6 -13.5 -13.3	
H	AOML3	6.05	7.47	7.82	33.60	7.09	33.60	6.08	
	AOML1	6.13	7.74	8.04	34.45	7.08	34.45	6.17	
	REIKO	6.22	7.79	8.20	34.57	6.77	34.53	5.70	
	EDIE-MAY P1	6.21	7.84	7.54	32.067	6.35	32.07	5.38	
	EDIE-MAY P2	6.25	7.77	7.53	32.074	6.32	32.07	5.36	
G	AOML3	196.9	193.9	337.5	358.9	31.5	358.9	33.3	
	AOML1	197.6	194.7	339.8	0.6	30.8	0.6	32.4	
	REIKO	196.0	194.5	339.1	0.7	32.7	0.7	36.8	
	EDIE-MAY P1	199.9	197.4	341.5	2.5	34.5	2.4	37.7	
	EDIE-MAY P2	200.0	197.3	341.7	2.6	34.3	2.6	37.5	

prefer the grav+rad values. In the past, the choice of the number of weights and the treatment of radiational tides have been subjective. On the basis of MODE and other recent measurements, Zetler and Munk (1975) have established some criteria. Here we shall give a brief review of how the selection was made.

#### a. Weights

In the response analysis, the measured series is approximated (in the least-square sense) as a weighted sum of the reference series for various leads or lags. For example, a single complex weight (1+0i) for zero lag corresponds to identical series; 0+2i corresponds to a measured series in quadrature with the reference series and of twice its amplitude. The vertical lines (solid and dashed) in the bottom left corner of Fig. 1 show a diurnal amplitude ratio of 1.18 and phase lag of 8° of EDIE-MAY P1 relative to Bermuda, obtained for a single (complex) weight.

Munk and Cartwright (1966) recommend lag intervals of 2 days. For two complex weights (lags 0, 2 days) the admittance is now a smoothly varying function of frequency: e.g., the amplitude ratio and phase lag are somewhat larger at higher frequency for the diurnals (see also  $O_1$  and  $K_1$  in Table 1). For additional weights, the admittances become increasingly more wiggly, in part (one surmises) as a result of the noise content. The trick is to terminate when one's credo on the smoothness of oceanic admittances is violated.

To obtain objective criteria, the series was divided into sections A and B. The top left panel gives the variance in the residual of section A predicted from an analysis of A and B, respectively. For self-prediction (A residuals from A weights) one expects, and finds, that the residuals diminish with increasing number of weights. However, after three weights, the improvement is slow and one suspects that the analysis is responding more to noise than to signal. At approximately that point one expects, and finds, that A residuals from B weights should deteriorate. The results are similar for B residuals, except that the overall residual is lower. On this basis the decision was made to stop at two weights. For the semidiurnal tides the "turning point" is a bit later, hence the decision to stop at three weights. The decisions are in general accord with the wiggliness of the admittances, taking into account also the resemblance between admittances from A and B residuals.

## b. Radiational tides

Radiational tides are periodic variations in sea level primarily related to meteorological changes such as the semi-daily cycle in barometric pressure and daily land and sea breezes. These cyclical variations match the frequencies of solar (not lunar) gravitational tidal constituents. Cartwright (1966) and Zetler (1971) found the average ratio of radiational amplitude to gravitational amplitude at the  $S_2$  frequency to be about 0.17.

The bandwidths of the radiational tides are much narrower than those of the gravitational tides, essentially covering tidal lines separated by cycles per year (cpy) rather than cycles per month (cpm). Furthermore, the relative amplitudes within the 1 cpm range are quite different for gravitational and radiational tides. In particular, equilibrium  $K_2/S_2$  (grav) is 0.27, whereas the rad ratio is 0.09. A response analysis for a sufficiently long record (as the 10-year Bermuda record) satisfactorily resolves these different bandwidths and amplitude relationships, and the gravitational admittances across tidal bands are expected to be smooth. Inasmuch as traditional tidal analysis for any length of series and response analysis for short series cannot distinguish between gravitational and radiational contributions to a constituent, but instead solve for their vectorial sum, discontinuities in admittances at the  $S_2$  frequency are inevitable. In fact, Zetler (1971) used these discontinuities as a simple means of calculating the radiational  $S_2$  from harmonic constants computed by traditional analysis.

Gravitational and radiational weights were resolved<sup>6</sup> for the Bermuda record and separate gravitational and radiational predictions were prepared for the MODE period. Given these reference predictions for Bermuda, there are three options for analyzing the MODE data: (i) compute gravitational and radiational admittances separately in a combined analysis; (ii) use the gravitational predictions only; and (iii) sum the gravitational and radiational predictions for each species and use the summed complex predicted series as reference.

All three methods were tried. For the first case, it quickly became evident that a much longer deep-sea series was required (probably 10 years) to separate gravitational and radiational contributions. An unstable matrix in solving for the weights resulted in some absurd results (such as a predominance of radiational tides). An extrapolation was made from the (lunar) N<sub>2</sub> and M<sub>2</sub> frequencies to the (solar) S<sub>2</sub> frequency to determine the extent of the discontinuity at the latter frequency, using both the second and third options with EDIE-MAY P1 (Table 1):

	N <sub>2</sub>	M <sub>2</sub>		$S_2$	
Fre- quency (cpd)	1.895	1.932	Extra- polated	2.000 Grav +rad	Grav only
<i>R</i> ф	0.92 -3.8°	0.90 -4.2°	0.86 ⊷4.9°	0.79 10.3°	0.67 

<sup>6</sup> Only the gravitational admittances are listed in Table A. The gravitational and radiational constants and their sums are as follows:

	Gravit	ational	Radia	ational	Sum		
	H	G	H	G	H	G	
	(cm)	(deg)	(cm)	(deg)	(cm)	(deg)	
$P_1$	2.16	187.7	0.15	5.9	2.01	187.8	
$K_1$	6.70	187.0	0.15	5.9	6.55	187.0	
$S_2$	9.26	21.7	1.24	185.3	8.08	24.2	
$K_2$	2.31	21.9	0.11	185.3	2.21	22.7	

In a comparison of the Bermuda harmonic constants with those obtained for a different period by traditional harmonic analysis (IHB Spec. Publ. No. 26, Sheet 600), the discrepancies in amplitude are less than 4%, in epoch less than  $2^{\circ}$ .

TABLE 2. Ratio of residual to recorded variance.

		Station									
	AOML3	AOML1	REIKO	EDIE- MAY P1	EDIE- MAY P2						
Total (0 to 1	2 cpd)										
grav+rad	0.00931	0.01338	0.00483	0.00680	0.00948						
grav only	0.00937	0.01340	0.00485	0.00680	0.00949						
Diurnal (1 cp	d±4½ cpi	m)									
grav+rad	0.00207	0.00212	0.00051	0.00107	0.00135						
grav only	0.00206	0.00192	0.00054	0.00105	0.00136						
Semi diurnal	$(2 \text{ cpd} \pm 4)$	1/2 cpm)									
grav+rad	0.00066	0.00051	0.00014	0,00020	0.00021						
grav only	0.00068	0.00056	0.00019	0.00022	0.00024						

The discontinuity is smaller for the grav+rad option. Furthermore, residual variances for grav+rad are generally somewhat smaller than for grav only (Table 2). The better results with the grav+rad option indicate a common origin for Bermuda and MODE radiational tides, with a similar relation to the gravitational tides. We would expect the grav only procedure to be superior if Bermuda radiational tides were the result of (i) local effects such as diurnal winds from island heating, or (ii) inverted barometer response to atmospheric pressure (absent from bottom pressure readings). Evidently this is not the case.

## c. Admittance

Inasmuch as the transducers for EDIE P1 and P2 were only 10 cm apart, a comparison of the two records is a measure of reproducibility of results (Snodgrass *et al.*, 1975). For  $M_2$  the admittances differ by only  $2 \times 10^{-4}$  in amplitude and 0.1° in phase. For the four diurnal tidal constituents listed, the amplitude discrepancy is always less than 1% (the largest being 0.07 cm for  $K_1$ ) and the phases all agree within 0.1°.

Changes between constituents within each species (relative to Bermuda) are smooth. We find that for  $O_1$  and  $K_1$  relative amplitudes at any one station are within 2%, relative phases within 3°; for  $N_2$  and  $M_2$  the values are 3% and 1°. Yet, the Bermuda reference (Table A) shows marked variation across the semidiurnal band, possibly the result of a free mode. The evidence is for a resonance of large areal extent, rather than highly localized.

The satisfactory agreement at one site assures us of the reality of the small difference between REIKO near the central mooring (28°00'N, 69°40'W) and EDIE-MAY about 150 km to the south-southeast, a small separation when reckoned in tidal dimensions. (The locations of all bottom pressure sensors are shown in Fig. 2). The reasonable fit between the harmonic constants for REIKO and AOML1, roughly 20 km apart near the central mooring, encourages us to discuss jointly results from the IGPP and AOML data. Nevertheless, even though there is no obvious calibration

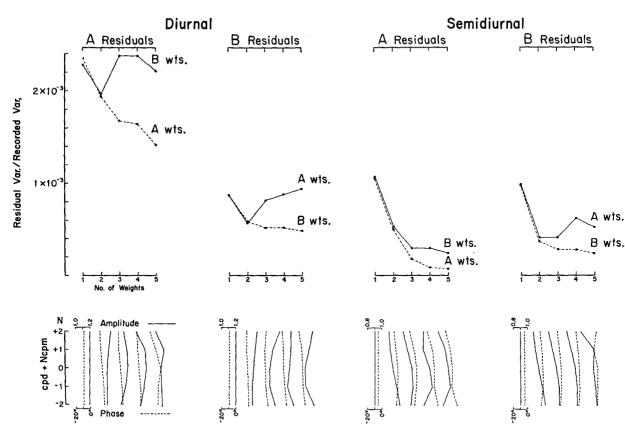


FIG. 1. Analysis of variance for EDIE-MAY P1. The left upper panel gives the ratio of residual to recorded diurnal variance for the first 29 days (record A), with prediction weights based on the A record (self-prediction) and B record (last 29 days), respectively. The remaining upper panel gives the B residuals and corresponding semidiurnal ratios. In each case the dashed lines refer to "self-predictions," the solid lines to future (B residuals from A weights) or past (A residuals from B weights) predictions, respectively. The lower panels give the corresponding amplitude ratios (solid) and phase lags (dashed) relative to Bermuda reference, as a function of frequency (1  $cpd\pm0$ , 1, 2 cpm for diurnals, 2  $cpd\pm0$ , 1, 2 cpm for semidiurnals), for 1, 2, ..., 5 complex weights. With increasing number of weights the self-prediction residuals (but not necessarily those for the future predictions) diminish, and the admittances become increasingly wiggly.

difference between the IGPP and AOML gauges, it seems prudent to use the REIKO, EDIE capsules for comparison between central mooring with the area to the south and the AOML1, AOML3 capsules for comparison between central mooring with the area to the east.

Station differences are consistent for different constituents within each species. At the central mooring the semidiurnal tides are about 8% larger and 2° earlier than at EDIE; the diurnal tides are about equal in amplitude and 3°-4° earlier than at EDIE. Comparing the AOML data, at the central mooring the semidiurnal tides are about 2% larger and 2° later than at AOML3; the diurnal tides slightly smaller and earlier at AOML3. These comparisons imply that both the 1 and 2 cpd tides progress from the northeast, arriving first at AOML3, then the central mooring, and finally at EDIE.

## 4. Comparison with cotidal charts

There is a general impression that the state-of-the-art in preparing cotidal and co-range charts leaves much to be desired; certainly there are considerable variations between published charts. Most charts deal with only the largest tidal constituent,  $M_2$ , sometimes labeling the chart as applying to semidiurnal tides. Similarly, if a chart is designated "diurnal tides," ordinarily it refers to the  $K_1$  constituent. Fig. 2 is a comparison of Dietrich's (1944) charts for  $M_2$  and  $K_1$  with our long series, AOML1 and 3, REIKO, and EDIE-MAY P1 and P2.

# a. $M_2$ tides

Dotted lines show our inference of Dietrich's  $0^{h}15^{m}$ and  $0^{h}30^{m}$  cotidal lines. All four MODE stations and Bermuda lie between  $0^{h}$  and  $0^{h}15^{m}$  lines (solar hours, not component hours), so their Greenwich epochs should be between  $0^{\circ}$  and 7.2°. The indicated Bermuda epoch (from our analysis) is 358.3° (IHB sheet 600 gives  $0^{\circ}$ ). Dietrich's  $0^{h}$  line would need only small displacements to fit the Bermuda values. AOML1, EDIE and REIKO fit Dietrich's chart and are relatively consistent; AOML3 is roughly  $3^{\circ}$  (6 min in time) early

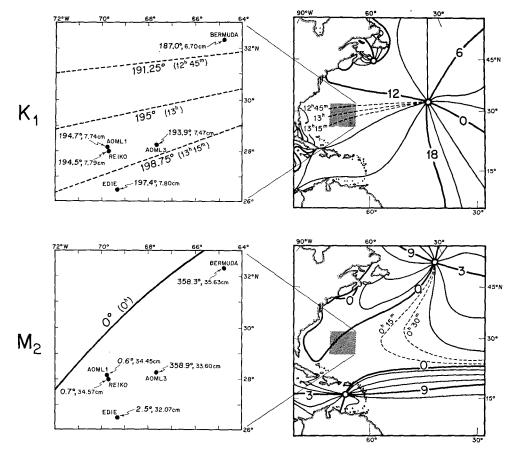


FIG. 2. Right panels show Dietrich's (1944) cotidal lines in the North Atlantic for  $K_1$  and  $M_2$  tides, respectively. Values are in solar hours, with dashed curves designating our interpolation. The MODE area falls within the shaded square which is shown on an enlarged scale in the left panels, for comparison with the results at MODE stations ( $\bullet$ ).

relative to the other MODE epochs and to the Dietrich configuration. Dietrich does not show co-range lines and the  $M_2$  cotidal lines are too complex in this region to infer readily whether the  $M_2$  amplitudes are consistent with distance from an amphidrome.

Although the Tiron *et al.* (1967)  $M_2$  cotidal chart has the XII<sup>h</sup> (same as 0<sup>h</sup>) in a similar position and orientation as Dietrich's, one would interpret his progression in time in the MODE area to be to the northwest, contrary to Dietrich's chart and to the values for three MODE stations. The orientation of the Tiron *et al.* XII<sup>h</sup> line would call for the AOML3 epoch to be between the epochs of EDIE and the central mooring stations (REIKO and AOML1); thus, on this chart, too, the AOML3 epoch seems to be slightly low (early). The orientation and progression of the Tiron *et al.* co-range lines call for the largest MODE amplitude at the central mooring, as is found to be the case. AOML3 amplitude conforms in that its value falls between those for EDIE and the central mooring.

# b. K1 tides

All stations are roughly  $2^{\circ}-3^{\circ}$  early relative to the cotidal lines, yet they are consistent in that they pro-

gress toward the south-southeast with spacing matching that of the inferred cotidal lines. As expected, amplitudes increase with distance from the amphidrome.

## c. Baroclinic noise

In making detailed comparisons with cotidal charts it should be recognized that the measurements contain a baroclinic contribution of roughly 1% amplitude,  $\frac{1}{4}^{\circ}$ in phase. The characteristic wavelength of the baroclinic tides is 100 km, and at separations of several hundred kilometers the baroclinic contribution to the measured vector sum will differ almost randomly from station to station. This effect exceeds the instrumental noise. If we may assume that barotropic and baroclinic tidal *currents* are of the same order (Section 6), then the pressure gradients and associated surface slopes are of the same order, 10<sup>-2</sup> cm km<sup>-1</sup> for M<sub>2</sub>. To obtain baroclinic surface amplitudes we multiply by  $\lambda/2\pi$ , with  $\lambda = 160$  km for the gravest baroclinic mode. This gives a near-bottom baroclinic amplitude of 0.3 cm, and for a random orientation the phase shift is by roughly  $\frac{1}{4}^{\circ}$ . If there had been surface measurements, the baroclinic contribution would have been several times larger.

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# 5. Tides of higher order

The high reproducibility of the principal tidal constituents encouraged us to analyze MODE pressures for the relatively small components of higher order. These are derived by the response method, using the equilibrium potential  $G_3^3$  as reference function for  $M_3$ , and the nonlinear input functions  $G_2^2G_2^2$ ,  $G_2^1G_2^2G_2^2$ ,  $G_2^2G_2^2G_2^2$ (rather than  $G_4^4$ ,  $G_5^5$ ,  $G_6^6$ ) for  $M_4$ ,  $M_5$ ,  $M_6$ , for the reasons stated below. For comparison, Table 3 includes the principal diurnal and semidiurnal tides, equilibrium amplitudes, and some Bermuda estimates. The amplitudes for the higher-order tides at MODE stations correspond to total species energy and therefore are somewhat higher than comparable values for the specific constituents.

 $M_3$  amplitudes at MODE stations and Bermuda are within a factor of 2, and phases lie within one octant. There is nothing extraordinary about this result other than the satisfaction of detecting a 1 mm constituent in deep-sea records. Amplitudes are somewhat less than the equilibrium value, about the same ratio as for the diurnals.

 $M_4$ ,  $M_5$ ,  $M_6$  amplitudes are tiny (typically 0.7, 0.5, 0.3 mm) yet very much *larger* than the equilibrium values  $(3 \times 10^{-2}, 4 \times 10^{-4}, 3 \times 10^{-6} \text{ mm})$ . One's first inclination is to attribute these results to experimental noise. We note that IHB values are generally high, perhaps because they include the noise continuum, whereas other estimates reflect only the coherent signal. AOML3 amplitudes are generally higher. Yet in spite

of these inconsistencies we believe the estimates for the higher-order harmonics to be significant, for the reason that phases are not randomly distributed, especially if one allows for the fact that distances between stations are no longer negligible as compared to wavelengths.

There is then the question as to why the  $M_4$ ,  $M_5$ ,  $M_6$ amplitudes are so much larger than equilibrium amplitudes. Garrett and Munk (1971) (see also Gallagher and Munk, 1971) have suggested that quadratic and higher-order interactions in the world's shallow basins produce multiple frequencies which leak into the global oceans and attain an equilibrium level for which radiative and dissipative processes are somehow balanced. For these reasons the response method was used with double and triple products of the appropriate semidiurnal and diurnal tide potentials as input functions.

## 6. Inferred tidal currents

Tidal measurements at two points determine the tidal variations in the average pressure gradient between these points, and impose some conditions on the tidal currents. The tidal displacement of the seafloor (and bottom pressure gauge) must here be taken into account (Munk *et al.*, 1970). We use  $\hat{\varsigma}^{\text{s}}$  and  $\hat{\varsigma}^{\text{B}}$  for the radial (upward) displacement, relative to the center of the Earth, of the sea surface and bottom, respectively;  $\zeta = \hat{\varsigma}^{\text{s}} - \hat{\varsigma}^{\text{B}}$  is the surface displacement relative to the sea bottom, as observed.

The gravitational forces of the Moon and Sun are conveniently expressed in terms of the *equilibrium* 

	K	K <sub>1</sub>		M <sub>2</sub>		3	$M_4$		$M_5$		$M_6$	·
	<i>Н</i> (ст)	G (deg)	<i>Н</i> (ст)	G (deg)	<i>H</i> (cm)	G (deg)	H (cm)	G (deg)	<i>H</i> (cm)	G (deg)	H (cm)	G (deg)
Equilibrium												
Normalized* At MODE latitude 28°N	36.9 11.8		63.2 19.0		0.76 0.22		$1.1 \times 10^{-2}$ $3.0 \times 10^{-3}$		$1.5 \times 10^{-4}$ $3.7 \times 10^{-5}$		$1.5 \times 10^{-6}$ $3.4 \times 10^{-7}$	_
Bermuda												
St. George (IHB p. 600) Wunsch (1972)** Garrett and Munk (1971)	<u>6.4</u> —	189 	35.5  36.2	000	0.15 0.13	40 74	0.40 0.16 0.18	232 260			0.18	83 
MODE***												
AOML3 AOML1 REIKO EDIE-MAY P1 EDIE-MAY P2	7.5 7.7 7.8 7.8 7.8	194 195 195 197 197	33.6 34.5 34.6 32.1 32.1	359 001 001 003 003	$\begin{array}{c} 0.078 \\ 0.155 \\ 0.155 \\ 0.136 \\ 0.104 \end{array}$	69 36 38 40 55	$\begin{array}{c} 0.134 \\ 0.065 \\ 0.074 \\ 0.077 \\ 0.074 \end{array}$	354 330 278 271 253	$\begin{array}{c} 0.58 \\ 0.021 \\ 0.033 \\ 0.050 \\ 0.045 \end{array}$	321 277 262 302 295	$\begin{array}{c} 0.031 \\ 0.003 \\ 0.021 \\ 0.032 \\ 0.019 \end{array}$	255 102 125 19 59

TABLE 3. Amplitudes and Greenwich epochs of higher-order tides.

\* Using the normalization by Munk and Cartwright (1966).  $K_1$ ,  $M_2$  and  $M_3$  are from Cartwright and Taylor (1971);  $M_4$ ,  $M_5$  and  $M_6$  are from a computer-generated potential. Latitude factors are

$$(-1)^m \left[\frac{2n+1}{4\pi}\right]^{\frac{1}{2}} \left[\frac{(n-m)!}{(n+m)!}\right]^{\frac{1}{2}} p_n^m (\cos\theta)$$

where  $\theta = 90^{\circ} - \text{latitude}$ .

\*\* Some values have been modified in accordance with a recent personal communication.

\*\*\*  $M_3$ ,  $M_4$ ,  $M_5$   $M_6$  are derived from total species energy. Epochs for  $M_4$ ,  $M_5$ ,  $M_6$  are for the summed tides within the appropriate cpd bands.

TABLE 4. Computed and measured tidal currents.\*

	$u \operatorname{comp}$	ponent	v compone		
	Speed (cm s <sup>-1</sup> )	G (deg)	Speed (cm s <sup>-1</sup> )	G (deg)	
From pressure gradients [Eq. (4)]	0.49	110	0.80	258	
Hendry central mooring					
Measured	0.30	72	0.31	233	
Barotropic (0)	0.60	110	0.55	328	
Baroclinic (1)	0.37	306	0.64	184	
(2)	0.37	218	0.36	108	
Hendry average, all depths and moorings	0.49	106	0.38	299	

\* Values in parentheses are modes obtained by a mode decomposition.

response

$$\begin{cases} \hat{\varsigma}_{e}^{\mathrm{S}} = (1+k)U/g, \quad \hat{\varsigma}_{e}^{\mathrm{B}} = hU/g \\ \varsigma_{e} = \hat{\varsigma}_{e}^{\mathrm{S}} - \hat{\varsigma}_{e}^{\mathrm{B}} = (1+k-h)U/g \end{cases}$$
(1)

of surface, bottom and surface relative to bottom, respectively, and this essentially defines Love numbers k, h (k allows for the self-attraction of the tidal bulge). The procedure does not allow for non-equilibrium selfattraction and loading (Hendershott, 1973).

The equations of motion in the center-of-earth system are

$$\partial_{\iota} u - fv = g \partial_{x} (\hat{\zeta}_{e}^{\mathrm{S}} - \hat{\zeta}^{\mathrm{S}}), \quad \partial_{\iota} v + fu = g \partial_{y} (\hat{\zeta}_{e}^{\mathrm{S}} - \hat{\zeta}^{\mathrm{S}}), \quad (2)$$

where  $f = 2\Omega \cos\theta$  is the Coriolis parameter. We now set

$$\hat{\zeta}^{\mathrm{B}} = \hat{\zeta}^{\mathrm{B}}_{e},\tag{3}$$

thus assuming that the bottom distortion (unlike the surface) can be adequately represented by the equilibrium configuration. It follows that  $\zeta_e^{\rm S} - \hat{\zeta}^{\rm S} = \zeta_e - \zeta$ . For harmonic oscillations  $\exp(-i\omega t)$ , Eqs. (2) can be written in the bottom-of-the-sea system (observation coordinates) as

$$\begin{bmatrix} u \\ v \end{bmatrix} = \frac{g}{f^2 - \omega^2} \begin{bmatrix} -i\omega\partial_x + f\partial_y \\ -f\partial_x - i\omega\partial_y \end{bmatrix} (\zeta_e - \zeta).$$
(4)

Let the real part of  $\zeta_m = A_m \exp -i(\omega t - G_m)$  designate the tidal elevation at station m, and

$$\zeta_{mn} = (\mathbf{A}_n e^{iG_n} - \mathbf{A}_m e^{iG_m}) e^{-i\omega t} = \mathbf{A}_{mn} e^{iG_{mn}} e^{-i\omega t}$$

the station difference between m and n. For  $M_2$ , we have the following:

	AOML1	AOML3	513	EDIE	REIKO	ζer
$\begin{array}{c} A & ({ m cm}) \\ G & ({ m deg}) \end{array}$	34.45 0.6	33.60 358.9				2.72 338.3

The average tidal gradient over the MODE area has

been obtained from the station differences according to

$$\zeta_{13} = L_{13} (\partial_x \zeta \cos \phi_{13} + \partial_y \zeta \sin \phi_{13}) \zeta_{\text{ER}} = L_{\text{ER}} (\partial_x \zeta \cos \phi_{\text{ER}} + \partial_y \zeta \sin \phi_{\text{ER}})$$
(5)

with  $L_{13}=218$  km,  $L_{ER}=171$  km,  $\phi_{13}=3^{\circ}$ ,  $\phi_{ER}=102^{\circ}$  designating the appropriate station separations and bearings for AOML1 and 3, REIKO and EDIE. The corresponding equilibrium gradients derived from

$$U/g = (24.39 \text{ cm}) \sin^2\theta \exp[-i(\omega t + 2\lambda)]$$

are as follows:

$$\frac{\partial_{x\zeta_{e}} = R^{-1}(1+k-h)(2\times 24.39)}{\times \sin\theta \exp\left[-i(\omega t+2\lambda+\pi/2)\right]} \\ \frac{\partial_{y\zeta_{e}} = R^{-1}(1+k-h)(24.39)}{\times \sin 2\theta \exp\left[-i(\omega t+2\lambda+\pi)\right]}$$
(6)

where k=0.29, h=0.59 are the Love numbers,  $\theta=62^{\circ}$  is colatitude,  $\lambda=-70^{\circ}$  is east longitude, and R the equatorial radius of the earth. The numerical values from (5) and (6) are

 $M_2$  tidal currents were then computed from (4), using  $f=6.8\times10^{-5}$  s<sup>-1</sup>. Listed with them in Table 4 are some analyses of current records by Ross Hendry, MIT: (i) the measured near-bottom currents; (ii) the contributions to near-bottom currents from various modes, using a mode decomposition of current profiles at the central mooring; (iii) an average over all depths and for many moorings.

In principle, the currents computed from (4) are to be interpreted as the *total* tidal current (barotropic plus baroclinic) at the depth of the array. The difficulty is that the spacing between stations ( $\sim 200$  km), though small compared to the barotropic wavelength, is comparable to the baroclinic wavelength ( $\lambda = 160$  km for mode 1), so that the baroclinic contributions are substantially reduced in an average taken between stations. We may then look for a first-order agreement between the computed values and the barotropic component (whether derived from mode separation or by averaging), and this is roughly what is found. But the baroclinic contributions at a given point are comparable to the barotropic contribution, as has long been known, and for a detailed comparison it would be necessary to resolve the baroclinic cotidal field by closely spaced stations.

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.

## **Results of MODE Tide Measurements**

#### TABLE A.

Station: Bermuda standard tide gauge

32°24'N, 64°42'W First series starts 1950 May 21 (441652–450171)† Second series starts 1953 May 2 (467500–476019) Third series starts 1956 April 14 (493372–501891)

Combined solution using meaned matrices

Reference: Gravitational and radiational potentials:  $G_2^1(0, \pm 2, \pm 4)$ ,  $G_2^2(0, \pm 2, \pm 4)$ ,  $G_3^1(0)$ ,  $G_3^2(0)$ ,  $R_1^1(0)$ ,  $R_2^1(0)$ ,  $R_2^2(0)$ 

	.1	+ (Station /D)		Principal harmonic constituents							
		s‡ (Station/Ro cpm (0.03660			-		Stati	on		ce (gravi- ul only)	
Frequency (cpd)	Real	Imaginary	R	φ* (deg)		Frequency (cpd)	II (cm)	G* (deg)	H (cm)	G* (deg)	
0,8929346	0,2221	-0.0316	0.2244	-8.1	(Q1)	0.8932441	1.13	188.2**	5.02	0	
0.9295357	0,1976	-0.0421	0.2020	-12.0	O <sub>1</sub>	0.9295357	5.30	192.0**	26.22	0	
0.9661368	0.1820	-0.0250	0.1837	-7.8	-						
					$(\mathbf{P}_1)$	0.9972621	2,02**	187.7	12.21	0	
1.0027379	0.1823	-0.0214	0.1835	-6.7	K <sub>1</sub>	1.0027379	6.56**	187.1**	36.88	0	
1,0393390	0,1956	-0.0433	0.2003	-12.5							
1.8590714	0.6266	0.3772	0.7314	31.0							
1.8956725	0.6259	0.2589	0.6773	22.5	$(N_2)$	1.8959820	8.18	337.7	12,10	0	
1.9322736	0.5636	0.0169	0.5369	1.7	$M_2$	1.9322736	35.63	358.3	63,19	0	
1.9688747	0.4340	-0.1248	0.4516	-16.0	-						
					$(S_2)$	2.0000000	8.08	24.2	29.40	0	
2.0054758	0.2677	-0.1078	0.2885	-21.9	$\mathbf{K}_{2}$	2.0054758	2.21	22.7	8.00	0	

† Greenwich hours since 1900 January 1, 0h.

‡ Dimensionless.

Station: REIKO-MAY Pressure

\* G is Greenwich epoch,  $\phi$  is station lead.

\*\* Small corrections for those values listed in the text.

#### TABLE B.

#### 1 June-5 July 1973 27°58.2'N, 69°40.4'W (643496 to 644368 Greenwich hours since 1900 January 1 0<sup>h</sup>) (ex prediction: $G_2^1+G_2^1+R_1^1+R_2^1$ (0, -2), $G_2^2+G_2^2+R_3^2$ (0, +2)

Ref	erence : B	ermuda	a summed	complex	prediction:	$G_2^i + G_3^i +$	$R_1^{i} + R_2^{i}$ (	0, -2),	$G_2^2 + G_3^2 +$	$-R_2^{*}(0,\pm 2)$
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А	dmittance	s† (Station/R	eference)			Princi	pal harmoni	c constituer	nts	
	Intervals 1	cpm (0.03660	)11 cpd)				Sta	tion	Reference	
Frequency (cpd)	Real	Imaginary	R	φ* (deg)		Frequency (cpd)	H (cm)	G* (deg)	<i>II</i> (cm)	G* (deg)
0.8929346	1,1459	-0.0560	1,1473	-2.8	(Q1)	0.8932441	1.30	191.0	1.13	188.2
0.9295357	1.1713	-0.0823	1.1742	-4.0	O1	0.9295357	6.22	196.0	5.30	192.0
0.9661368	1.1824	-0.1172	1,1882	- 5.7						
					(P <sub>1</sub> )	0.9972621	2.40	194.9	2.02	187.7
1.0027379	1.1768	-0.1533	1.1868	-7.4	$K_1$	1.0027379	7.79	194,5	6.56	187.1
1.0393390	1.1558	-0.1832	1.1702	-9.0						
Recorded v	ariance : 85	5.3 cm <sup>2</sup>			4					
Residual va Ratio: 0.00		43 cm²			-4					
1.8590714	0,9946	0.0372	0,9953	-2.1						
1.8956725	1.0025	-0.0252	1.0028	-1.4	$(N_2)$	1.8959820	8.20	339,1	8.18	337.7
1.9322736	0.9693	-0.0400	0.9702	-2.4	$M_2$	1.9322736	34.57	0.7	35,63	358.3
1.9688747	0,9020	-0.0787	0,9055	-5.0						
					$(S_2)$	2.0000000	6.77	32.7	8.08	24.2
2.0054758	0.8145	-0.1331	0,8253	-9,3	$K_2$	2.0054758	1.82	32.0	2.21	22.7
Recorded v Residual va Ratio: 0.00	riance: 0.0									

† Dimensionless.

\* G is Greenwich epoch,  $\phi$  is station lead.

# Station: MERT-Pressure

17 March-1 April 1973 27°59.3'N, 69°40.3'W

(641741 to 642089 Greenwich hours since 1900 January 1 0<sup>h</sup>) Reference: Bermuda summed complex prediction:  $G_2^1+G_3^1+R_1^1+R_2^1$  (0, -2),  $G_2^2+G_3^2+R_2^2$  (0, -2)

rvals 1	cpm (0.03660	)11 cpd)							
		ni cpay				Sta	tion	Refe	erence
Real	Imaginary	R	φ* (deg)		Frequency (cpd)	<i>H</i> (cm)	G* (deg)	<i>Н</i> (ст)	G* (deg)
.3560	-0.1951	1.3699	-8.2	(Q <sub>1</sub> )	0.8932441	1.55	196.4	1.13	188.2
.2718	-0.2041	1.2881	-9.1	01	0.9295357	6.83	201.1	5.30	192.0
.1924	0.1747	1.2051	-8.3						
				$(P_1)$	0.9972621	2.32	193.9	2.02	187.7
.1342	-0.1132	1.1399	-5.7	$K_1$	1.0027379	7.48	192.8	6.56	187.1
.1094	-0.0323	1.1099	-1.7						
.9496	0.0889	0.9537	5.3						
.9647	0.0361	0.9654	2.1	$(N_2)$	1.8959820	7.90	335.6	8.18	337.7
.9549	-0.0178	0.9550	-1.1	$M_2$	1.9322736	34.03	359,4	35.63	358.3
.9221	-0.0618	0.9241	-3.8						
				$(S_2)$	2,0000000	7.15	29.7	8.08	24.2
.8732	-0.0867	0.8775	5.7	$\mathbf{K}_2$	2.0054758	1.94	28.4	2.21	22.
	.3560 .2718 .1924 1342 .1094 nce : 34 ce : 0,30 9496 9647 9549 9221	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	Real         Imaginary $R$ (deg)           .3560 $-0.1951$ $1.3699$ $-8.2$ .2718 $-0.2041$ $1.2881$ $-9.1$ .1924 $-0.1747$ $1.2051$ $-8.3$ .1342 $-0.1132$ $1.1399$ $-5.7$ .1094 $-0.0323$ $1.1099$ $-1.7$ nce: $34.0$ cm <sup>2</sup> .9496 $0.0889$ $0.9537$ $5.3$ .9647 $0.0361$ $0.9654$ 2.1           .9549 $-0.0178$ $0.9550$ $-1.1$ .9221 $-0.0618$ $0.9241$ $-3.8$	Real         Imaginary $R$ (deg)           .3560 $-0.1951$ $1.3699$ $-8.2$ (Q <sub>1</sub> )           .2718 $-0.2041$ $1.2881$ $-9.1$ O <sub>1</sub> .1924 $-0.1747$ $1.2051$ $-8.3$ (P <sub>1</sub> )           .1342 $-0.1132$ $1.1399$ $-5.7$ K <sub>1</sub> .1094 $-0.0323$ $1.1099$ $-1.7$ nce: $34.0 \text{ cm}^2$ ce: $0.302 \text{ cm}^2$ 9496 $0.0889$ $0.9537$ $5.3$ .9647 $0.0361$ $0.9654$ 2.1         (N <sub>2</sub> )           .9549 $-0.0178$ $0.9550$ $-1.1$ M <sub>2</sub> .9221 $-0.0618$ $0.9241$ $-3.8$ (S <sub>2</sub> )	Real         Imaginary $R$ (deg)         (cpd)           .3560 $-0.1951$ $1.3699$ $-8.2$ (Q <sub>1</sub> ) $0.8932441$ .2718 $-0.2041$ $1.2881$ $-9.1$ $O_1$ $0.9295357$ .1924 $-0.1747$ $1.2051$ $-8.3$ (P <sub>1</sub> ) $0.9972621$ .1342 $-0.1132$ $1.1399$ $-5.7$ $K_1$ $1.0027379$ $1094$ $-0.0323$ $1.1099$ $-1.7$ $nce: 34.0 \text{ cm}^2$ $e: 0.302 \text{ cm}^2$ 9496 $0.0889$ $0.9537$ $5.3$ $99647$ $0.0361$ $0.9654$ $2.1$ $(N_2)$ $1.8959820$ .9549 $-0.0178$ $0.9550$ $-1.1$ $M_2$ $1.9322736$ .9221 $-0.0618$ $0.9241$ $-3.8$ $(S_2)$ $2.0000000$	Real         Imaginary $R$ (deg)         (cpd)         (cm)           3560 $-0.1951$ $1.3699$ $-8.2$ (Q <sub>1</sub> ) $0.8932441$ $1.55$ 2718 $-0.2041$ $1.2881$ $-9.1$ $O_1$ $0.9295357$ $6.83$ $.1924$ $-0.1747$ $1.2051$ $-8.3$ (P <sub>1</sub> ) $0.9972621$ $2.32$ $.1342$ $-0.1132$ $1.1399$ $-5.7$ $K_1$ $1.0027379$ $7.48$ $.1094$ $-0.0323$ $1.1099$ $-1.7$ $-1.7$ $-1.7$ nce: $34.0 \text{ cm}^2$ $ee: 0.302 \text{ cm}^2$ $-9537$ $5.3$ $99647$ $0.0361$ $0.9654$ $2.1$ $(N_2)$ $1.8959820$ $7.90$ $9549$ $-0.0618$ $0.9241$ $-3.8$ $(S_2)$ $2.0000000$ $7.15$	Real         Imaginary $R$ (deg)         (cpd)         (cm)         (deg)           3560 $-0.1951$ $1.3699$ $-8.2$ (Q <sub>1</sub> ) $0.8932441$ $1.55$ $196.4$ 2718 $-0.2041$ $1.2881$ $-9.1$ $O_1$ $0.9295357$ $6.83$ $201.1$ $.1924$ $-0.1747$ $1.2051$ $-8.3$ (P <sub>1</sub> ) $0.9972621$ $2.32$ $193.9$ $.1342$ $-0.1132$ $1.1399$ $-5.7$ $K_1$ $1.0027379$ $7.48$ $192.8$ $.1094$ $-0.0323$ $1.1099$ $-1.7$ $nce: 34.0 \text{ cm}^2$ $ee: 0.302 \text{ cm}^2$ 9496 $0.0889$ $0.9537$ $5.3$ $9647$ $0.0361$ $0.9654$ $2.1$ $(N_2)$ $1.8959820$ $7.90$ $335.6$ $9549$ $-0.0178$ $0.9550$ $-1.1$ $M_2$ $1.9322736$ $34.03$ $359.4$ $9221$ $-0.0618$ $0.9241$ $-3.8$ $(S_2)$ $2.0000000$ $7.15$ $29.7$	Real         Imaginary $R$ (deg)         (cpd)         (cm)         (deg)         (cm)           3560 $-0.1951$ $1.3699$ $-8.2$ (Q <sub>1</sub> ) $0.8932441$ $1.55$ $196.4$ $1.13$ 2718 $-0.2041$ $1.2881$ $-9.1$ $O_1$ $0.9295357$ $6.83$ $201.1$ $5.30$ $1924$ $-0.1747$ $1.2051$ $-8.3$ $P_1$ $0.9295357$ $6.83$ $201.1$ $5.30$ $11342$ $-0.1132$ $1.1399$ $-5.7$ $K_1$ $1.0027379$ $7.48$ $192.8$ $6.56$ $1094$ $-0.0323$ $1.1099$ $-1.7$ $-1.7$ $-0.0323$ $1.1099$ $-1.7$ nce: $34.0 \text{ cm}^2$ $ce: 0.302 \text{ cm}^2$ $-9.03257$ $5.3$ $9.947$ $0.0361$ $0.9654$ $2.1$ $(N_2)$ $1.8959820$ $7.90$ $335.6$ $8.18$ $9549$ $-0.0178$ $0.9550$ $-1.1$ $M_2$ $1.9322736$ $34.03$ $359.4$ $35.63$

† Dimensionless. \* G is Greenwich epoch,  $\phi$  is station lead.

#### TABLE D.

# Station: EDIE-March Pressure

20 March-9 April 1973 26°26.9'N, 69°19.0'W (641810 to 642282 Greenwich hours since 1900 January 1 0h) Reference: Bermuda summed complex prediction:  $G_2^1+G_3^1+R_1^1+R_2^1$  (0, -2),  $G_2^2+G_3^2+R_2^2$  (0, -2)

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А	dmittance	s† (Station/R	eference)			nts					
]	Intervals 1	cpm (0.03660	011 cpd)			Station				Reference	
Frequency (cpd)	Real	Imaginary	R	φ* (deg)		Frequency (cpd)	H (cm)	G* (deg)	H (cm)	G*. (deg)	
0.8929346	1.0705	-0.1095	10.761	-5.8	(Q1)	0.8932441	1,22	194.0	1.13	188.2	
0.9295357	1.1343	-0.1430	1.1433	-7.2	O1	0.9295357	6.06	199.2	5.30	192.0	
0.9661368	1.1766	-0.2014	1.1938	-9.7							
					(P <sub>1</sub> )	0.9972621	2.46	200.1	2.02	187.7	
1.0027379	1.1886	-0.2724	1.2195	-12.9	$\mathbf{K}_{1}$	1.0027379	8.00	200.0	6.56	187.1	
1.0393390	1.1679	-0.3415	1.2167	-16.3							
Recorded v	ariance : 54	.2 cm <sup>2</sup>									
Residual va	riance: 0.0	98 cm <sup>2</sup>									
Ratio: 0.00	18										
1.8590714	0.8566	0.0725	0.8597	4.8							
1.8956725	0.8797	0.0254	0.8800	1.7	$(N_2)$	1.8959820	7.20	336.1	8.18	337.7	
1.9322736	0.8794	-0.0270	0.8798	-1.8	$M_2$	1,9322736	31.35	0.1	35.63	358.3	
1,9688747	0.8559	-0.0738	0.8591	-4.9							
					$(S_2)$	2.0000000	6.69	31.3	8.08	24.2	
2.0054758	0.8140	-0.1054	0.8208	-7.4	$\mathbf{K}_{z}$	2.0054758	1.81	30.1	2.21	22.7	
Recorded v Residual va Ratio: 0.000	ariance: 51 riance: 0.1	.8 cm <sup>2</sup>	0.0200	-7.7	Δ2	2,0037750	1.01	50.1	2.21	4	

† Dimensionless.

\* G is Greenwich epoch,  $\phi$  is station lead.

#### TABLE E.

# Station: EDIE-May P1

:---

12 May-9 July 1973 26°27.8'N, 69°19.6'W

(643072 to 644468 Greenwich hours since 1900 January 1 0<sup>h</sup> Reference: Bermuda summed complex prediction:  $G_2^1+G_3^1+R_1^1+R_2^1$  (0, -2,)  $G_2^2+G_3^2+R_2^2$  (0, ±2)

А	dmittance	s† (Station/R	eference)		Principal harmonic constituents						
Intervals 1 cpm (0.0366011 cpd)							Sta	tion	Refe	erence	
Frequency (cpd)	Real	Imaginary	R	$\phi^*$ (deg)		Frequency (cpd)	H (cm)	G* (deg)	U (cm)	G* (deg)	
0.8929346	1.1378	-0.1465	1,1472	-7.3	(Q1)	0.8932441	1,30	195.5	1,13	188.2	
0.9295357	1.1609	-0.1621	1.1722	-7.9	$\hat{O}_{I}$	0,9295357	6.21	199.9	5,30	192.0	
0.9661368	1.1748	-0.1864	1.1894	-9.0							
					$(\mathbf{P}_1)$	0.9972621	2,41	197,8	2.02	187.7	
1.0027379	1.1764	0.2142	1.1957	- 10.3	K,	1.0027379	7.84	197.4	6,56	187.1	
1.0393390	1.1654	-0.2399	1.1899	-11.6							
Recorded v Residual va Ratio: 0,00	riance : 0,0										
1,8590714	0.9013	-0.0845	0,9053	-5.4							
1,8956725	0.9192	-0.0608	0.9213	-3.8	$(N_2)$	1,8959820	7.54	341.5	8.18	337.7	
1.9322736	0.8976	-0.0652	0.9000	-4.2	M 2	1,9322736	32.07	2.5	35.63	358.3	
1,9688747	0.8410	-0.0968	0.8465	6.6							
					$(S_2)$	2,0000000	6.35	34.5	8.08	24.	
2.0054758	0.7611	-0.1490	0.7755	-11.1	$K_2$	2,0054758	1.71	33.8	2.21	22.3	

† Dimensionless.

Station: EDIE-May P2

\* G is Greenwich epoch,  $\phi$  is station lead.

TABLE F.

#### 12 May-9 July 1973 26°27.8'N 69°19.6'W (643072 to 644468 Greenwich hours since 1900 January 1 0h

Reference: Bermuda summed complex prediction:  $G_2^1+G_3^1+R_1^1+R_2^1$  (0, -2),  $G_2^2+G_3^2+R_2^2$  (0, ±2)

А	dmittance:	s† (Station/Re	eference)		Principal harmonic constituents					
	Intervals 1 cpm (0.0366011 cpd)						Sta	tion	Refe	rence
Frequency (cpd)	Real	Imaginary	R	φ* (deg)		Frequency (cpd)	II (cm)	G* (deg)	II (cm)	G* (deg)
0.8929346	1.1543	-0.1456	1,1635	-7.2	(Q <sub>1</sub> )	0.8932441	1.31	195.4	1.13	188.2
0.9295357	1.1681	-0.1644	1.1796	-8.0	$O_1$	0,9295357	6.25	200.0	5.30	192.0
0.9661368	1.1721	-0.1872	1.1870	-9.1						
					$(\mathbf{P}_1)$	0.9972621	2.39	197.7	2.02	187.3
1.0027379	1.1656	-0.2095	1.1843	-10.2	$K_1$	1.0027379	7.77	197.3	6.56	187.1
1.0393390	1.1498	-0.2266	1.1720	11,1						
Recorded va Residual va Ratio: 0.001	riance: 0,1									
1.8590714	0.8971	-0.0861	0.9012	5.5						
1.8956725	0.9182	-0.0638	0.9204	-4.0	$(N_2)$	1.8959820	7.53	341.7	8.18	337.7
1.9322736	0.8977	-0.0677	0.9002	-4.3	$M_2$	1.9322736	32.07	2.6	35.63	358.3
1.9688747	0.8399	-0.0969	0.8454	-6.6						
					$(S_2)$	2.0000000	6.32	34.3	8.08	24.2
2.0054758	0.7568	-0.1454	0.7707	-10.9	$\mathbf{K}_2$	2.0054758	1.70	33.6	2.21	22.
Recorded v: Residual va Ratio : 0.000	riance : 0,1									

† Dimensionless.

\* G is Greenwich epoch,  $\phi$  is station lead.

#### TABLE G.

## Station: AOML1 Pressure

# 12 March-29 June 1973 28°08.4'N, 69°45.2'W

(641593 644225 Greenwich hours since 1900 January 1 0<sup>h</sup>) Reference: Bermuda summed complex prediction:  $G_2^1+G_3^1+R_1^1+R_2^1$  (0, -2),  $G_2^2+G_3^2+R_2^2$  (0, ±2)

А	dmittance	s† (Station/R	eference)		Principal harmonic constituents						
Intervals 1 cpm (0.0366011 cpd)							Sta	tion	Ref	erence	
Frequency (cpd)	Real	Imaginary	R	φ* (deg)		Frequency (cpd)	II (cm)	G* (deg)	II (cm)	G* (deg)	
0.8929346	1.1282	-0.1014	1.1327	-5.1	(Q1)	0.8932441	1.28	193.3	1.13	188.2	
0.9295357	1.1502	-0.1127	1.1557	-5.6	O <sub>1</sub>	0.9295357	6.13	197.6	5,30	192.0	
0.9661368	1.1650	-0.1326	1.1725	-6.5							
					$(P_1)$	0,9972621	2.38	195.2	2.02	187.1	
1.0027379	1.1693	-0.1569	1.1798	-7.6	$K_1$	1,0027379	7.74	194.7	6.56	187.	
1.0393390	1.1625	-0.1807	1.1764	- 8.8							
Recorded va Residual va Ratio : 0.002	riance: 0.1										
1.8590714	0.9678	-0.0589	0.9695	-3.5							
1.8956725	0.9824	-0.0366	0.9831	-2.1	$(N_2)$	1.8959820	8.04	339.8	8.18	337.2	
1,9322736	0,9663	-0.0386	0.9670	-2.3	M.	1,9322736	34.45	0,6	35,63	358,3	
1,9688747	0.9228	-0.0645	0.9251	- 4.0							
					$(S_2)$	2,0000000	7.08	30.8	8.08	24.	
2,0054758	0.8610	-0.1088	0.8679	-7.2	$K_2$	2.0054758	1.92	29.9	2.21	22.	
Recorded va Residual va Ratio: 0.000	riance: 0.3										

† Dimensionless. \* G is Greenwich epoch,  $\phi$  is station lead.

TABLE H.

Station: AOML3 pressure

#### 11 April-5 July 1973 28°14.2'N, 67°32.2'W (642325-644367 Greenwich hours since 1900 January 1 0h) Reference: Bermuda summed complex prediction: $G_2^1+G_3^1+R_1^1+R_2^1$ (0, -2), $G_2^2+G_3^2+R_2^2$ (0, ±2)

А	dmittance	s† (Station/R	eference)			Princi	c constituer	nts		
]	Intervals 1 cpm (0.0366011 cpd)						Sta	tion	Ref	erence
Frequency (cpd)	Real	Imaginary	R	$\phi^*$ (deg)		Frequency (cpd)	II (cm)	G* (deg)	<i>П</i> (ст)	G* (deg)
0.8929346	1.1281	-0,0812	1,1310	-4.1	(Q1)	0.8932441	1.28	192.3	1.13	188.2
0.9295357	1.1372	-0.0980	1.1414	-4.9	O1	0.9295357	6.05	196,9	5,30	192.0
0.9661368	1,1379	-0.1171	1.1440	-5.9						
					$(P_1)$	0,9972621	2.30	194.4	2.02	187.7
1.0027379	1.1301	-0.1346	1.1381	-6.8	$K_1$	1,0027379	7.47	193.9	6,56	187.1
1.0393390	1.1153	-0.1467	1.1250	-7.5						
Recorded va	ariance : 60	),2 cm <sup>2</sup>								
Residual va Ratio: 0.002		25 cm²					,			
1.8590714	0.9454	-0.0213	0.9457	-1.3						
1.8956725	0.9557	0.0024	0.9557	0.1	$(N_2)$	1.8959820	7.82	337.5	8.18	337.7
1.9322736	0.9428	-0.0092	0.9429	-0.6	$M_2$	1.9322736	33.60	358,9	35.63	358.3
1.9688747	0,9095	-0.0540	0,9111	-3.4						
		0.0010			$(S_2)$	2.0000000	7.09	31.5	8.08	24.2
2.0054758	0.8626	-0.1225	0.8712	-8.1	$\mathbf{K}_2$	2.0054758	1,93	30.8	2.21	22.7

† Dimensionless. \* G is Greenwich epoch,  $\phi$  is station lead.

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