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25 East Asian summer monsoon (EASM) precipitation affects the lives of billions of
26 people and impacts the stability of fragile desert ecosystems in central Asia [1].
27 Therefore, many studies have focused on understanding the variability of the EASM
28 and its relationship with insolation, ice sheet, and CO₂ forcings [1-3]. Evidence for
29 EASM variability is preserved in the eolian dust sequences on the Chinese Loess
30 Plateau (CLP), which highlight the importance of Northern Hemisphere (NH) ice sheets,
31 CO₂ levels, and insolation in controlling the strength of the EASM during the
32 Quaternary [2]. In contrast, Pliocene (5.3-2.6 Ma) proxy records of the EASM from the
33 CLP reveal weak orbital signals, suggesting weak sensitivity to ice sheets, CO₂ levels,
34 and insolation forcing during this sustained warm period [4]. These Pliocene proxy
35 records also contrast model simulations which suggest high sensitivity of the EASM to
36 orbital forcing (particularly precession) during this time [5].

37 In order to investigate the apparent lack of strong orbital cycles in the Pliocene
38 records from the CLP characterized by warm northern high latitudes and minor NH ice
39 sheets [6], we generate a high-resolution (3 ka) monsoon precipitation record from the
40 Loess Plateau (Fig. 1) using a recently proposed, promising magnetic parameter-based
41 precipitation proxy (χ_{fd}/HIRM), with larger values corresponding to higher precipitation
42 (see supplementary methods).

43 Using either established paleomagnetic age model of the Loess Plateau Chaona site
44 (supplementary methods and Fig. S1) or the orbitally-tuned age model (supplementary
45 methods and Fig. S2), the χ_{fd}/HIRM record consistently shows dominant 20-ka cycles,
46 in sharp contrast with the loess magnetic susceptibility record from the same section

47 (Fig. 2 and Fig. S3), which fails to resolve any orbital cycles during the 3.25-2.95 Ma.
48 The comparison of χ_{fd} /HIRM with June insolation gradient between 30° N and 30° S [7]
49 suggests that high χ_{fd} /HIRM values align well with the high June insolation gradient
50 (Fig. 2 and Fig. S3). We note that boreal summer insolation shows similar trends (Fig.
51 S4) to the June insolation gradient (30° N – 30° S), but modern climate data suggest that
52 the interhemispheric summer insolation gradient is a more likely forcing mechanism
53 for monsoon moisture than boreal summer insolation [8].

54 Our results suggest that precipitation decreases at four low insolation intervals
55 between 3.15 and 2.95 Ma (highlighted by pink bands, Fig. 2 and Fig. S3), although
56 these decreases are not as pronounced as the other low insolation intervals, suggesting
57 a non-linear response to insolation forcing. Interestingly, these four intervals correspond
58 to larger ice volume (Fig. 2 and Fig. S3), indicating a likely role that ice sheets played
59 in controlling monsoon precipitation. This inference is supported by comparing the 40-
60 ka band variations, where precipitation is out-of-phase with the June insolation gradient
61 (Fig. S5) but in-phase with the benthic oxygen isotope stack [9] (Fig. S2). The χ_{fd} /HIRM
62 record also shows non-orbital periodicities, such as the 30-ka and semi-precessional
63 signals associated with beats or harmonics of orbital cycles [10] (Fig. 2), confirming
64 that precipitation on the Loess Plateau had a non-linear response to insolation forcing.

65 In order to test whether both insolation and ice sheets are joint forcing for northern
66 China precipitation, we stacked the June insolation gradient with the benthic oxygen
67 isotope record (Fig. S6). The stacked records show similar variations and cyclicities to
68 the CLP precipitation proxy records, including main orbital cycles as well as semi-

69 precessional cycles (Fig. S6), providing further support for the inference that both
70 insolation and ice sheets are joint forcing for northern China precipitation variations
71 during the middle Piacenzian.

72 Our results provide an opportunity to robustly understand precipitation variations
73 in northern China and potential forcing during the Pliocene warmth at orbital bands. We
74 discussed the implications of our findings as follows.

75 First, although the middle Piacenzian period is characterized by a stable warm
76 climate, our record reveals that precipitation was highly sensitive to insolation forcing.
77 This finding is consistent with the South China Sea K/Si-based monsoon record [11],
78 and together, the marine and terrestrial records consistently reveal high sensitivity of
79 precipitation to insolation forcing in sustained warm periods. Second, our record
80 suggests that Antarctic ice sheets likely played an important role in affecting northern
81 China precipitation. The benthic oxygen isotope stack should indicate global ice volume
82 variations during the Quaternary after onset of intensive NH glaciations, with NH ice
83 sheet size variations playing a key role [2]. During the warm Pliocene, however, before
84 the intensive onset of NH glaciation, variations in the benthic oxygen isotope stack
85 should mainly reflect Antarctica ice sheet variations. During this time, pollen evidence
86 from a Russian Arctic lakes indicates late Pliocene summer temperatures up to 8°C
87 warmer than today, with mean temperature in the warmest month consistently above
88 10°C, leaving little room for permanent ice sheets to exist [6]. Furthermore, drilling
89 results from the Ross Sea reveal clear evidence for rapid cooling and Antarctica ice
90 sheet size increase during the middle Piacenzian [12], supporting that middle

91 Piacenzian ice sheet variations mainly occurred in Antarctica. The link between
92 Antarctic ice sheets and CLP precipitation, based on the in-phase relationship between
93 precipitation and the benthic oxygen isotope stack at the 40-ka band, can be achieved
94 by at least two processes. I) Model simulations suggest that larger Antarctic ice sheets
95 can intensify upwelling of circum-Antarctica deep water, which can reduce Southern
96 Hemisphere (SH) sea ice and warm surface seawater [13]. Subsequently, northward
97 propagation of SH warm seawater can result in a warmer Eurasia, more water vapor
98 transport from surrounding oceans to continental China, and a larger sea-land pressure
99 gradient associated with an amplified sea-land thermal contrast [13]. All of these factors
100 would be able to promote EASM precipitation increase. II) It has been proposed that
101 larger Antarctic ice sheets tend to push the Mascarene high and Australia high
102 northward to intensify cross-equatorial moisture transport and Asian monsoon
103 precipitation [14]. We note, however, that Antarctic ice volume fluctuations in the
104 Pliocene could be a symptom of an underlying global forcing from CO₂ amplification
105 well known for the late Pleistocene [15]. Therefore, we consider both ice-volume and
106 CO₂ to be important forcing mechanisms for the observed precipitation variations at the
107 40-ka band.

108 In summary, the χ_{fd} /HIRM data presented here provides a high-resolution proxy
109 record of precipitation from northern China between 3.25-2.95 Ma. 20-ka (precessional)
110 cycles dominate the record, challenging past research suggesting weak sensitivity of
111 northern China precipitation to insolation forcing during this sustained warm period in
112 the late Pliocene. In addition, this record suggests Antarctic ice sheet growth and/or

113 global atmospheric CO₂ may influence precipitation at 40-ka timescales during this time.
114 These findings highlight the importance of orbital forcing on precipitation even during
115 times of relatively warm climate, such as the middle Piacenzian or during future climate
116 scenarios.

117 **Conflict of interest**

118 The authors declare that they have no conflict of interest.

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127 **Author contributions**

128 Junsheng Nie designed the experiments. Zeng Luo, Rui Zhang and Zhao Wang
129 performed experiments. All authors analyzed data. Junsheng Nie, Zeng Luo and
130 Richard V. Heermance wrote the paper with the help of other authors.

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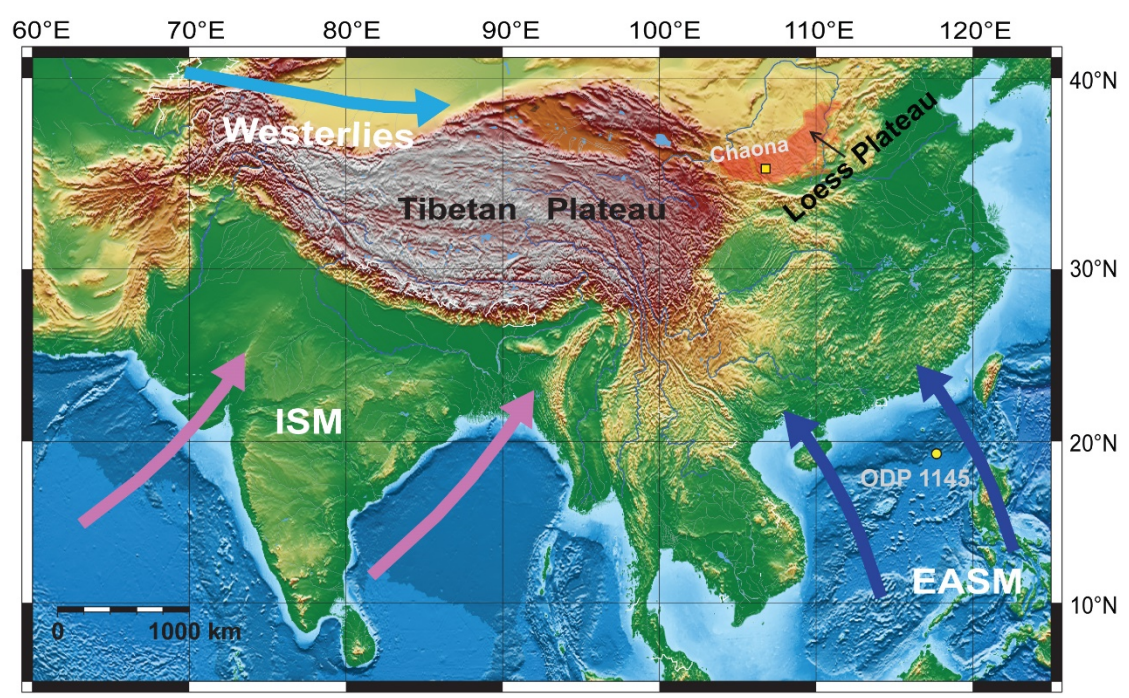
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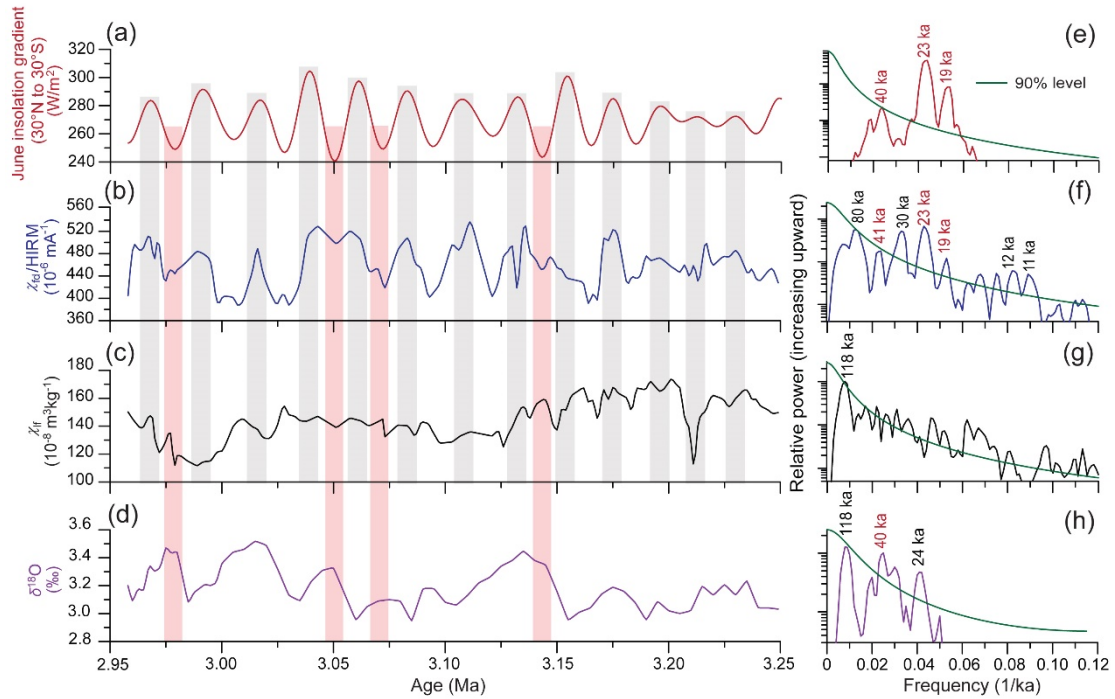
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174 **Fig. 1.** Map of study sites and modern Asian summer atmospheric circulation pattern.

175 EASM: East Asian summer monsoon. ISM: Indian summer monsoon.



176

177 **Fig. 2.** Chaona paleoclimate records and comparison with June insolation gradient and

178 benthic oxygen isotope stack. (a) June insolation gradient [7]. (b) Chaona $\chi_{fd}/HIRM$. (c)

179 Chaona χ_{lf} . (d) Benthic $\delta^{18}O$ stack [9]. (e)-(h) Power spectra of (a)-(d), respectively. The

180 grey bands highlight the intervals where $\chi_{fd}/HIRM$ peaks correspond to June insolation

181 gradient maxima. The pink bands highlight the complex relationship between $\chi_{fd}/HIRM$,

182 benthic $\delta^{18}O$, and insolation gradient. The y axes of the spectral plots are in log 10 scale.

183 The age model was based on orbital tuning.

184

185

24 **1. Abstract**

25 The processes controlling East Asian summer monsoon (EASM) variations during past
26 warm periods are poorly known but important for understanding its response to
27 prolonged warming. The eolian dust sequences on the Chinese Loess Plateau (CLP) play
28 a pivotal role in understanding EASM variations, revealing that Quaternary EASM was
29 forced by joint ice sheets, CO₂ levels, and insolation forcing. However, the paleoclimate
30 records from the Pliocene dust sequences on the CLP reveal weak sensitivity to any
31 recognized forcing, which is inconsistent with model simulations. To resolve this puzzle,
32 we present a 3-ka (thousand years) resolution precipitation record from the CLP,
33 focusing on the middle Piacenzian (3.264–3.025 Ma), the closest persistently warm
34 period in geological history, using a recently proposed precipitation proxy based on
35 magnetic parameters. The record reveals dominant 20-ka precessional cycles,
36 invalidating the prior hypothesis of a weak EASM sensitivity to insolation forcing
37 during the warm Pliocene. At the 40-ka band, the precipitation record shows an inverse
38 phase relationship with middle latitude summer insolation but an in-phase relationship
39 with Southern Hemisphere ice volume, which we attribute to CO₂ and/or ice sheet
40 control on northern China precipitation via atmosphere or marine processes. These
41 findings demonstrate a strong and highly variable EASM during the Pliocene, reshaping
42 our understanding of EASM variations in past warm periods.

43 **2. Materials and methods**

44 *2.1. Magnetic proxy*

45 Magnetic minerals produced during weathering are sensitive to climate in arid-semi
46 arid region [16-19]. Hematite and nanometer-scale ferrimagnetic mineral (magnetite,

47 which is quickly oxidized to maghemite due to high surface to volume ratio) are two
48 commonly produced minerals during weathering [16, 19, 20]. It is generally accepted
49 that hematite forms in dry environment [20]. In comparison with hematite, weathering-
50 produced nanometer-scale magnetite requires periods of reducing conditions in
51 sediment so that Fe^{3+} can be reduced to Fe^{2+} , which requires a wetter climate and higher
52 precipitation [16, 21]. Therefore, the ratio of nanometer-scale magnetite-maghemite
53 over hematite (i.e., the content of magnetic minerals generated during wetter climate
54 divided by those generated during drier climate) is more sensitive to precipitation
55 variations in the arid-semi arid regions than any single magnetic mineral alone. The ratio
56 is represented by frequency-dependent magnetic susceptibility (χ_{fd})/hard isothermal
57 remanent magnetization (HIRM), where χ_{fd} indicates content of pedogenic nanometer-
58 scale ferrimagnetic grains and HIRM reflects hematite content.

59 In order to measure the χ_{fd} /HIRM values, 108 bulk samples were collected at 5 cm
60 intervals in the Chaona section. We measured the magnetic susceptibility (χ) of these
61 samples at a lower frequency of 976 Hz and a higher frequency of 15616 Hz, using the
62 AGICO multi-function spinner Kappabridge (MFK1-FA). χ_{fd} was then calculated based
63 on the equation of $\chi_{\text{fd}} = \chi_{976\text{Hz}} - \chi_{15616\text{Hz}}$. Isothermal remanent magnetization (IRM)
64 intensity was measured at a magnetic field of 1.2 T and -0.3 T, respectively. HIRM (hard
65 IRM) was calculated as $(\text{IRM}_{1.2\text{T}} + \text{IRM}_{-0.3\text{T}})/2$. Finally, we calculated χ_{fd} /HIRM values
66 of 108 samples. These experiments were carried out at the Paleomagnetism Laboratory
67 in the China University of Geosciences, Beijing, China.

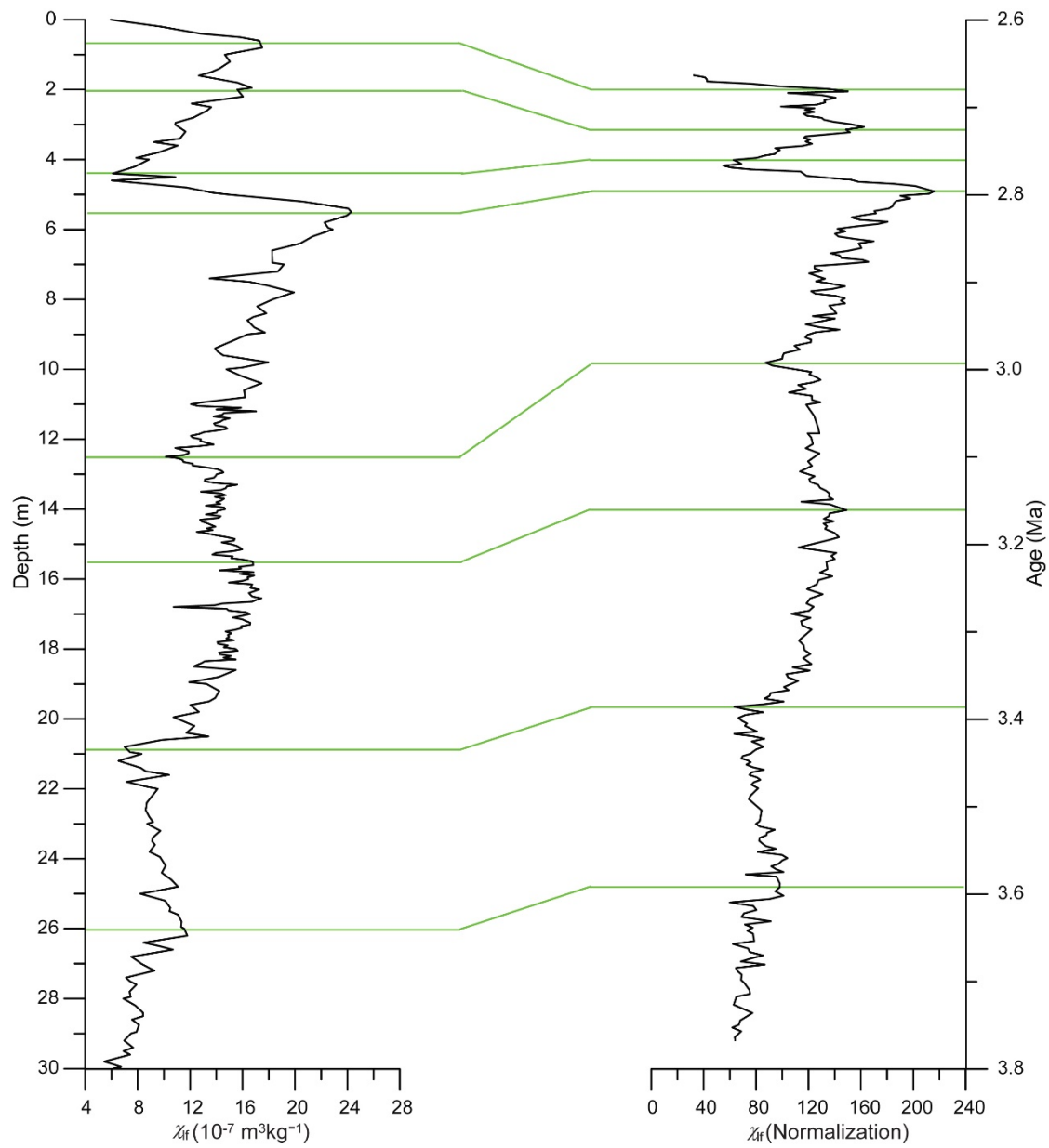
68 2.2. Establishing the initial age model of the Chaona section on the Loess Plateau

69 The age model of the Chaona section (35.1°N, 107.2°E) on the CLP was established
70 using paleomagnetism dating [22] in 2001. To reconstruct the orbital timescale
71 precipitation variations of the middle Piacenzian, we collected bulk samples at 5 cm
72 intervals for magnetic susceptibility analysis in September 2014. The good match
73 between new magnetic susceptibility data and previous magnetic susceptibility data [22]
74 in the Chaona section provides the age control points for establishing the age model of
75 resampling Chaona section (Fig. S1). Based on the 8 age control points, we obtain the
76 initial age model of new magnetic susceptibility data through piecewise linear
77 interpolation.

78 2.3. Establishing the orbitally-tuned age model for the Chaona sections

79 The strong 20-ka cycles in the χ_{fd} /HIRM record and the similar pattern with
80 insolation provide an opportunity to improve the precision of the age model from each
81 section (Fig. S3f). Only minimal tuning was needed for each section (Fig. S2b). We note
82 that the observed precipitation cyclicities and their comparison with insolation and ice
83 sheet records are consistent between the paleomagnetic age model and the tuned age
84 model (Fig. S2), indicating that tuning does not influence the results.

85

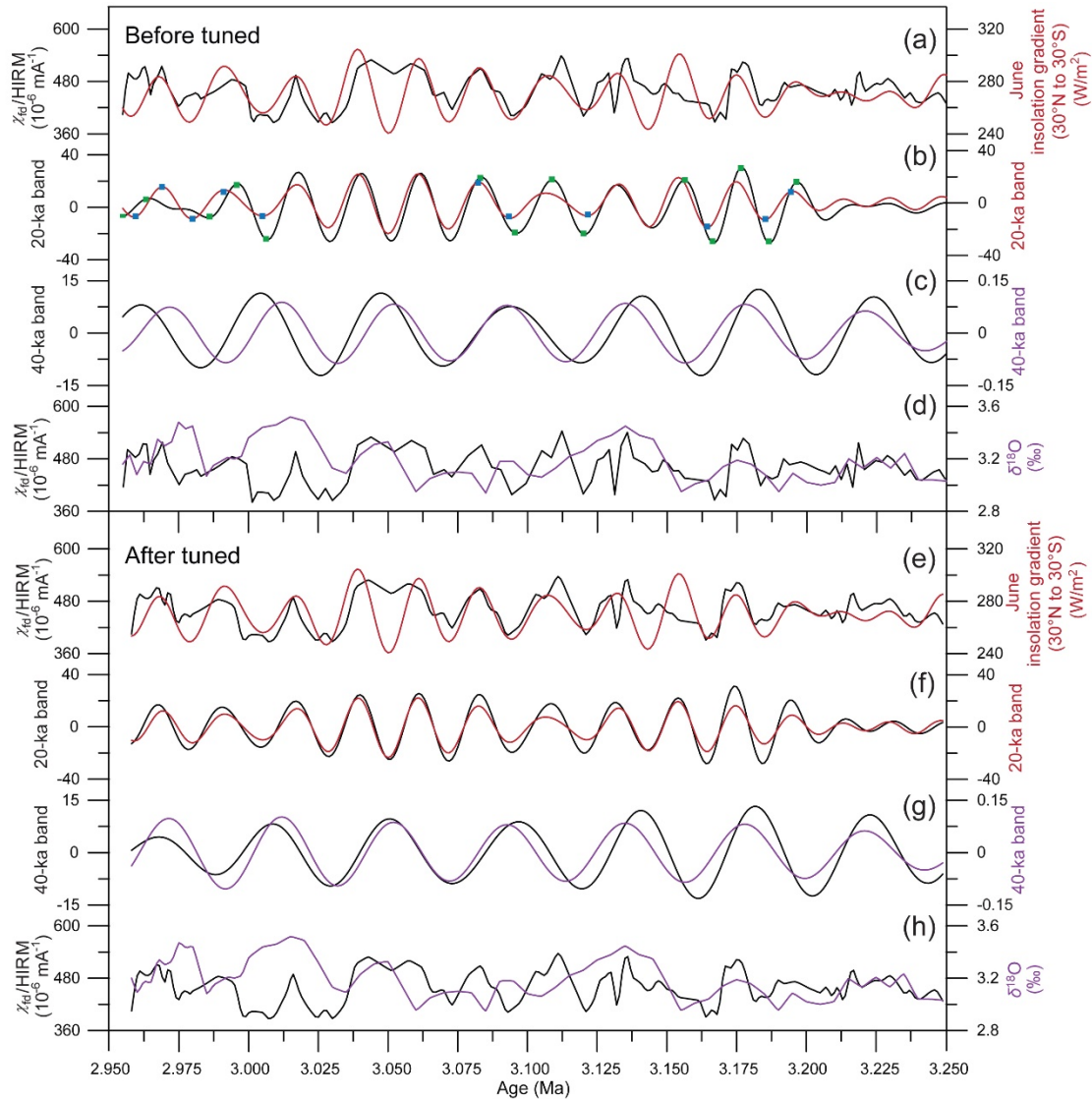


86

87 **Fig. S1.** The age control points for tuning the Chaona magnetic susceptibility. The right
 88 side is the new Chaona magnetic susceptibility data with depth and the left side is the
 89 previous Chaona magnetic susceptibility with age model [22]. χ_{lf} : Magnetic
 90 susceptibility at low frequency.

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95 **Fig. S2.** A comparison of Chaona precipitation proxy record with June insolation96 gradient and benthic oxygen isotope stack. (a) and (e) $\chi_{fd}/HIRM$ and June insolation97 gradient [7]. (b) and (f) $\chi_{fd}/HIRM$ and June insolation gradient at filtered 20-ka band. (c)98 and (g) $\chi_{fd}/HIRM$ and benthic $\delta^{18}\text{O}$ stack [9] at filtered 40-ka band. (d) and (h) $\chi_{fd}/HIRM$ 99 and benthic $\delta^{18}\text{O}$ stack. (b) show how the orbital tuning was done for Chaona section:

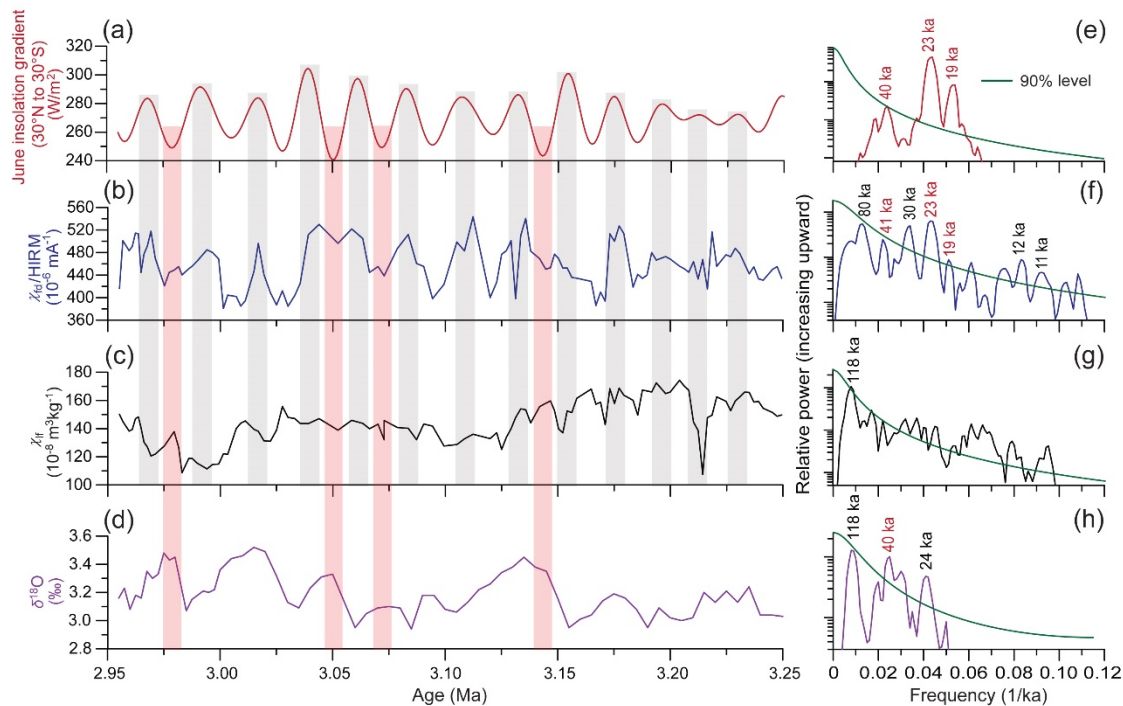
100 green points in (b) were tuned to the closest insolation reference points (blue points).

101 (a), (b), (c) and (d) show the results before tuned and (e), (f), (g) and (h) show the results

102 after tuned. The 20-ka central frequency = 0.05 ka^{-1} and bandwidth = 0.012 ka^{-1} . The
103 40-ka central frequency = 0.024 ka^{-1} and bandwidth = 0.004 ka^{-1} .

104

105



106

107 **Fig. S3.** Chaona paleoclimate records and comparison with June insolation gradient and

108 benthic oxygen isotope stack based on paleomagnetic dating. (a) June insolation

109 gradient [7]. (b) Chaona χ_{fd}/HIRM . (c) Chaona χ_{lf} . (d) Benthic $\delta^{18}\text{O}$ stack [9]. (e)-(h)

110 Power spectra of (a)-(d), respectively. The grey bands highlight the intervals where

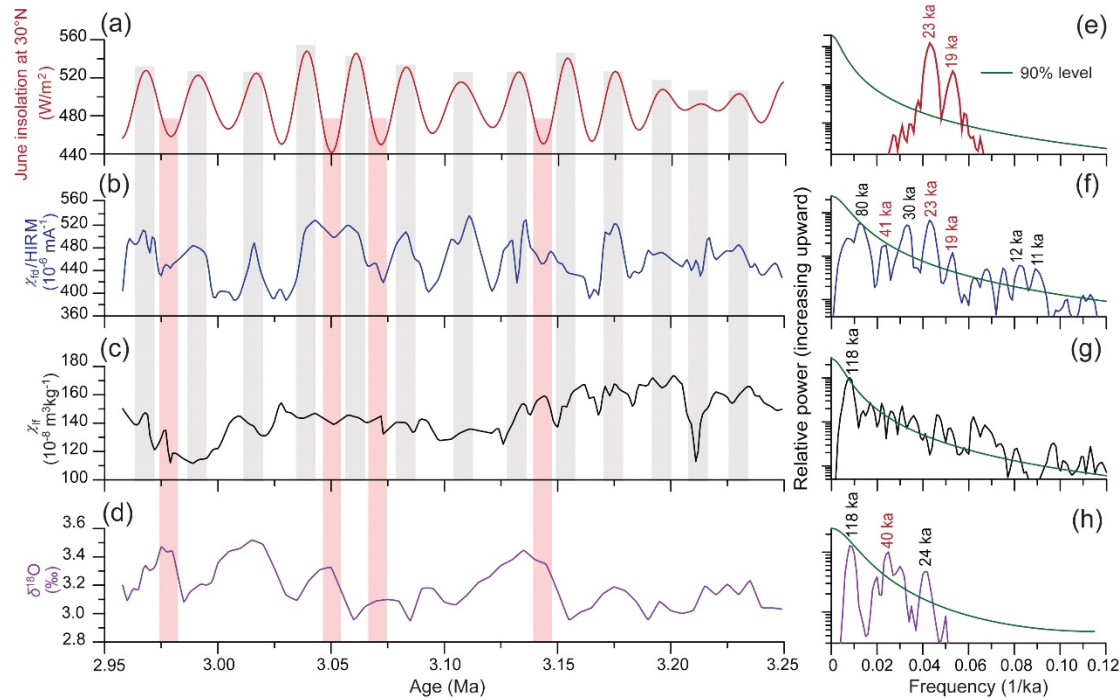
111 χ_{fd}/HIRM peaks correspond to June insolation gradient maxima. The pink bands

112 highlight the complex relationship between χ_{fd}/HIRM , benthic $\delta^{18}\text{O}$, and insolation

113 gradient. The y axes of the spectral plots are in log 10 scale.

114

115



116

117 **Fig. S4.** Chaona paleoclimate records and comparison with June insolation at 30°N and

118 benthic oxygen isotope stack based on orbitally-tuned age model. (a) June insolation at

119 30°N [7]. (b) Chaona $\chi_{fd}/HIRM$. (c) Chaona χ_{lf} . (d) Benthic $\delta^{18}O$ stack [9]. (e)-(h) Power

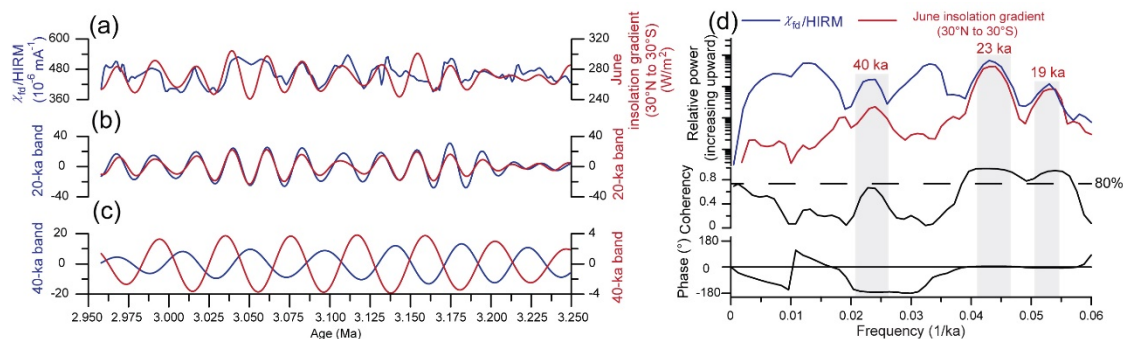
120 spectra of (a)-(d), respectively. The grey bands highlight the intervals where $\chi_{fd}/HIRM$

121 peaks correspond to June insolation maxima. The pink bands highlight the complex

122 relationship between $\chi_{fd}/HIRM$, benthic $\delta^{18}O$, and June insolation at 30°N. The y axes

123 of the spectral plots are in log 10 scale.

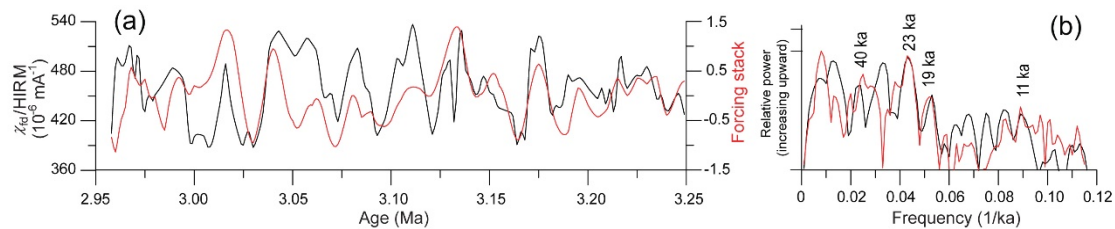
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126 **Fig. S5.** Comparison of χ_{fd} /HIRM from Chaona section with June insolation gradient
 127 and their cross-spectral results based on orbitally-tuned age model. (a)-(c) Comparison
 128 of χ_{fd} /HIRM from Chaona section with June insolation gradient [7] and their filtered
 129 results at 20-ka and 40-ka bands. (d) Cross-spectral comparison of χ_{fd} /HIRM and June
 130 insolation gradient. The power spectral is plotted on log scale. The coherence is plotted
 131 on an arctanh scale (80% = 0.749). The grey bands show the dominant orbital cycles.
 132 At the 40-ka band, χ_{fd} /HIRM and June insolation gradient are out of phase. This cross-
 133 spectral analysis was based on ARAND Time-Series Analysis Software [23]. The 20-
 134 ka central frequency = 0.05 ka^{-1} and bandwidth = 0.012 ka^{-1} . The 40-ka central
 135 frequency = 0.024 ka^{-1} and bandwidth = 0.004 ka^{-1} .

136



137

138 **Fig. S6.** Comparison of Loess Plateau χ_{fd} /HIRM with hypothesized insolation and ice
 139 sheet forcing stack based on orbitally-tuned age model. To produce the hypothesized
 140 forcing stack, we first normalized the June insolation gradient record and the benthic
 141 oxygen isotope record, then we multiplied the normalized benthic oxygen isotope record
 142 by 1.2 and added it with the normalized insolation record (hereafter forcing stack). (a)
 143 Comparison of Chaona χ_{fd} /HIRM with forcing stack. (c) The power spectral of (a),
 144 respectively. The y axes of the spectral plots are in log 10 scale.

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