ARCSTONE: Calibration of Lunar Spectral Reflectance from Space

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SBIR programs: Phase-I & Phase-II

ARCSTONE: Team and Contributions

NASA LaRC

Mission concept & science
Project management *
Engineering coordination
Instrument electronics
Flight and ground software
Mechanical, Thermal & Structural
Environmental testing
* SSAI: sub-contract management

LASP

Laboratory for Atmospheric and Space Physics

Instrument concept
Component characterization
Radiometric calibration
Error budget

NASA GSFC

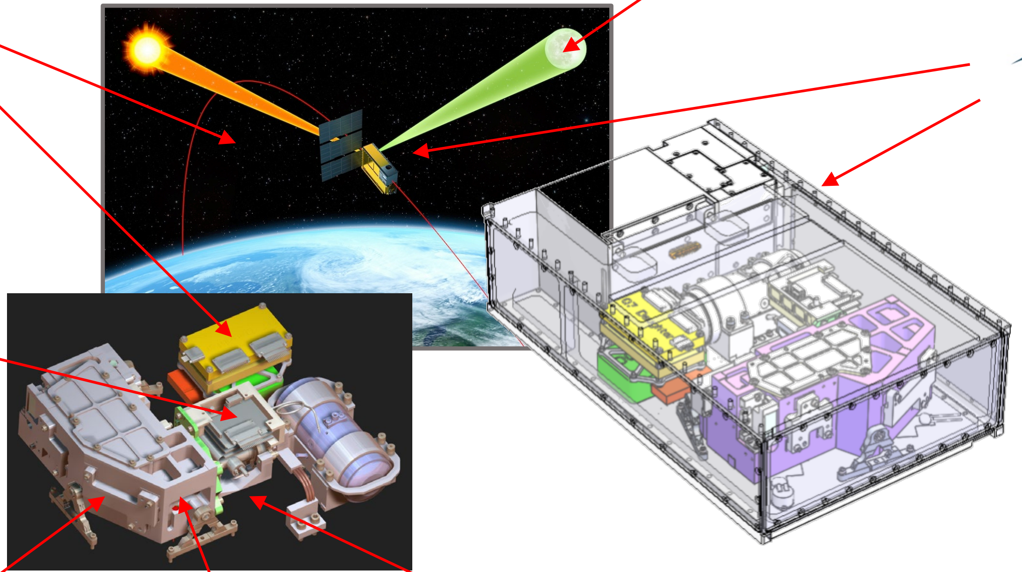
Optical black coating

USGS

Lunar calibration approach (ROLO)

BLUE CANYON TECHNOLOGIES

6U CubeSat Bus



RESONON

Instrument concept
Instrument design
Radiometric modeling
Fabrication
Assembly & alignment
Functional testing

QUARTUS ENGINEERING

Instrument Analysis (STOP, RV, TE)
Input to instrument design
Flexures design

ARCSTONE TEAM:

- NATIONWIDE COLLABORATION of EXPERTS !
- Collaboration with NIST & UMBC:
  Ground and Airborne lunar measurements

Moon: Potentially Accurate Source for Calibration On-orbit

- Measurement accuracy is directly related to the information content of the dataset. Measurement accuracy is critical to EOS! Current EOS cannot handle data gaps. Need overlapping observations: CERES, MODIS/VIIRS, Landsats, PACE/SeaWIFS, etc.

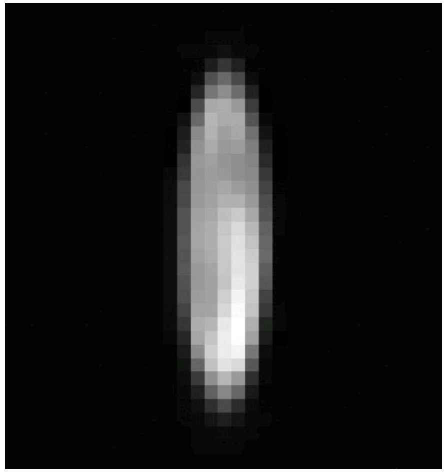
Calibration reference: Lunar Spectral Irradiance (entire disk)

Reflectance of Lunar surface stable to < 10^{-8} / year

- SeaWiFS gain stability: 0.13% (k=1) over 12 years
- Accuracy of current Lunar Model (ROLO): 5 – 10%

On-Orbit Calibration Need:

Absolute accurate spectral irradiance for all lunar phase angles and libration states.



Lunar image by SeaWIFS

Expected Impacts:
- Quality of data products
- Long-term consistency
- Handling data gaps
- Reduces instrument size, mass, power
- Reduce complexity
- Accurate CubeSat sensors

Applications of the Lunar Calibration Approach
(satellite operators worldwide !)

Team	*Satellite*	*Sensor*	*G/L*	*Dates*	*Number of obs*	*Phase angle range (°)*
CMA	FY-3C	MERSI	LEO	2013-2014	9	[43 57]
CMA	FY-2D	VISSR	GEO	2007-2014		
CMA	FY-2E	VISSR	GEO	2010-2014		
CMA	FY-2F	VISSR	GEO	2012-2014		
JMA	MTSAT-2	IMAGER	GEO	2010-2013	62	[-138,147]
JMA	GMS5	VISSR	GEO	1995-2003	50	[-94,96]
JMA	Himawari-8	AHI	GEO	2014-	-	
EUMETSAT	MSG1	SEVIRI	GEO	2003-2014	380/43	[-150,152]
EUMETSAT	MSG2	SEVIRI	GEO	2006-2014	312/54	[-147,150]
EUMETSAT	MSG3	SEVIRI	GEO	2013-2014	45/7	[-144,143]
EUMETSAT	MET7	MVIRI	GEO	1998-2014	128	[-147,144]
CNES	Pleiades-1A	PHR	LEO	2012	10	[+/-40]
CNES	Pleiades-1B	PHR	LEO	2013-2014	10	[+/-40]
NASA-MODIS	Terra	MODIS	LEO	2000-2014	136	[54,56]
NASA-MODIS	Aqua	MODIS	LEO	2002-2014	117	[-54,-56]
NASA-VIIRS	NPP	VIIRS	LEO	2012-2014	20	[50,52]
NASA-OBPG	SeaStar	SeaWiFS	LEO	1997-2010	204	(<10, [27-66])
NASA/USGS	Landsat-8	OLI	LEO	2013-2014	3	[-7]
NASA	OCO-2	OCO	LEO	2014		
NOAA-STAR	NPP	VIIRS	LEO	2011-2014	19	[-52,-50]
NOAA	GOES-10	IMAGER	GEO	1998-2006	33	[-66, 81]
NOAA	GOES-11	IMAGER	GEO	2006-2007	10	[-62, 57]
NOAA	GOES-12	IMAGER	GEO	2003-2010	49	[-83, 66]
NOAA	GOES-13	IMAGER	GEO	2006	11	
NOAA	GOES-15	IMAGER	GEO	2012-2013	28	[-52, 69]
VITO	Proba-V	VGT-P	LEO	2013-2014	25	[-7]
KMA	COMS	MI	GEO	2010-2014	60	
AIST	Terra	ASTER	LEO	1999-2014	1	-27.7
ISRO	OceanSat2	OCM-2	LEO	2009-2014	2	
ISRO	INSAT-3D	IMAGER	GEO	2013-2014	2	

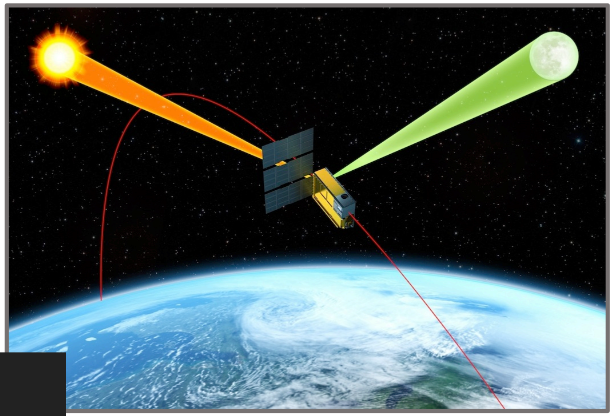
From GSICS (Global Space-based Inter-Calibration System) Lunar Calibration Workshop, December 2014, EUMETSAT.

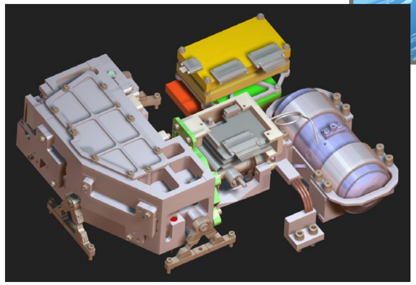
- **Instruments with lunar calibration capabilities participating in the GSICS GIRO (*GSICS implementation of the ROLO model*) program**
- ***List includes sensors with lunar observations submitted to the database at EUMATSAT as of December 2014.***
- ***Next GSICS Lunar Calibration Workshop: November 2020, virtual (?)***

ARCSTONE Objectives:

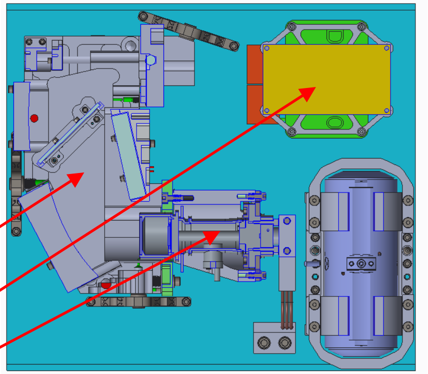
- *To enable on-orbit high-accuracy absolute calibration for the past, current, and future reflected solar sensors* in LEO and GEO* by providing lunar spectral irradiance as function of satellite viewing geometry and specified wavelength.

- To design, build, calibrate and validate a prototype instrument, demonstrate *form-fit-function for a 6U observatory with compliance in size, mass, power, and thermal performance*.





ARCSTONE Concept:
Accurate measurements of Lunar Irradiance from Space with an Instrument flying on 6U CubeSat (courtesy BCT) in LEO.



Progress of ARCSTONE instrument Design

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* Planetary instruments: OSIRIS Rex Camera suite [Golish et al., 2020]

$TRL_{current} = 4$ $TRL_{out} = 5$

ARCSTONE Mission Concept

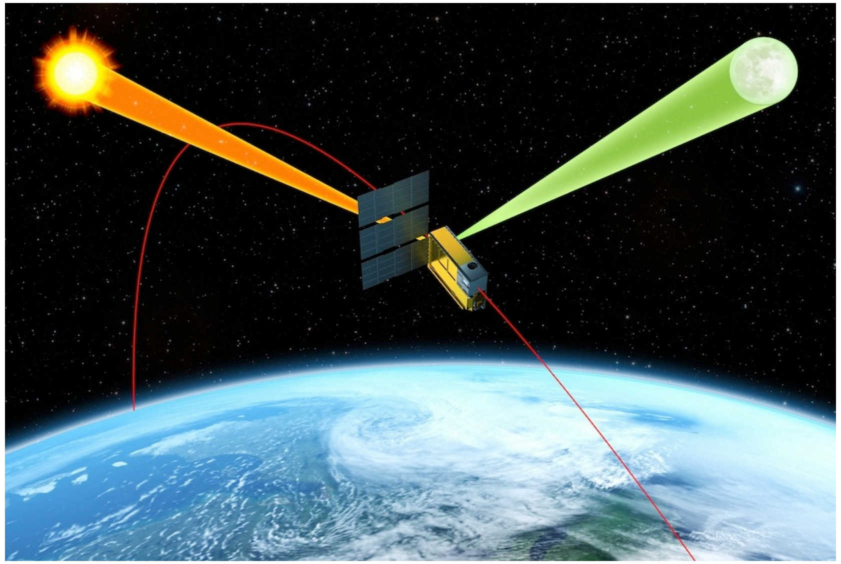
Concept of Operations and Data Products:

- Data to collect: Lunar spectral irradiance every 12 hours, 10 minutes
- Data to collect: Solar spectral irradiance for calibration (daily)
- Combined uncertainty < 0.5% (k=1)
- Spectrometer with single-pixel field-of-view about 0.7° (no scanning !)
- Sun synchronous orbit at 500 – 600 km altitude
- Spectral range from 350 nm to 2300 nm, spectral sampling at 4 nm

1 year: Improvement of current Lunar Calibration Model (factor of 2 – 4);
3+ years: New Lunar Irradiance Model, improved accuracy level (factor of 10).

Key Technologies to Enable the Concept:

- Approach to orbital calibration via referencing Sun (TSIS measurements): Demonstration of lunar and solar measurements with *the same optical path using integration time to reduce solar signal -- Major Innovation !*
- Pointing ability of spacecraft now permits obtaining required measurements *with instrument integrated into spacecraft.*



6U CubeSat Spacecraft Bus:
courtesy of Blue Canyon Technologies (BCT)

BCT 6U XB6 Spacecraft pointing:
Accuracy 0.002° (1-sigma) in 3 axis
Stability 1 arc-sec over 1 sec

ARCSTONE Mission: Key Performance Parameters

Key Parameters	Threshold Value	Goal Value
Accuracy (reflectance)	1.0% (k=1)	0.5% (k=1)
Stability	< 0.15% (k=1) per decade	< 0.1% (k=1) per decade
Orbit	Sun-synch orbit	Sun-synch orbit
Time on-Orbit	1 year	3 years
Frequency of sampling	24 hours	12 hours
Instrument pointing	< 0.2° combined	< 0.1° combined
Spectral Range	380 nm – 900 nm	350 nm – 2300 nm
Spectral Sampling	8 nm	4 nm

* * Threshold Values considered as success criteria

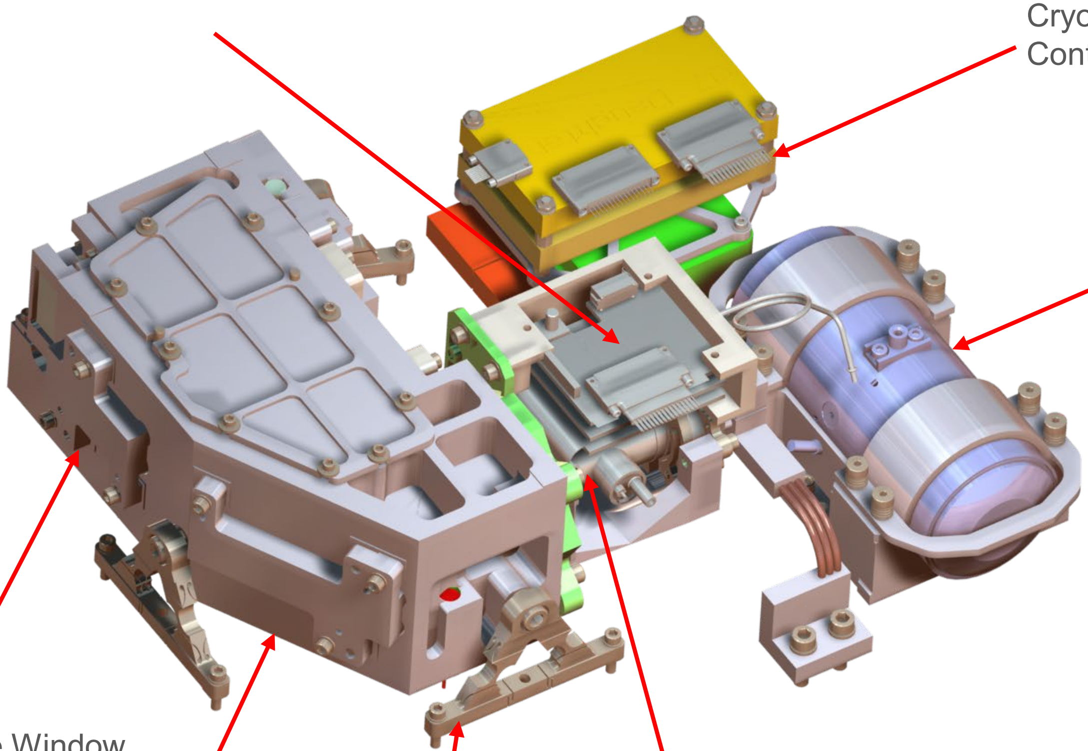
Reference for radiometric requirements (ROLO, T. Stone):
Lunar Phase Angle = 75°;
Irradiance = 0.6 (micro W / m^2 nm)
Wavelength = 500 nm

ARCSTONE MISSION CONOPS:

1. Lunar spectral irradiance observations:
   - Every 12 hours
   - Close to polar locations
   - Multiple measurements within 5– 10 minutes to improve SNR
2. Solar Spectral Irradiance observations (solar calibration):
   - Multiple measurements to get required SNR
   - This is radiometric calibration to the TSIS reference
3. Dark images:
   - Multiple measurements with closed shutter
   - Before every lunar and solar observations
4. Dark field (to calibrate out shutter temp):
   - Multiple measurements of dark space
5. Field-of-view sensitivity characterization:
   - Calibration of instruments alignment
6. Spectral calibration:
   - On-board spectral calibration
7. Spacecraft pointing calibration and other checks:
   - Defined by the BCT for calibration of spacecraft functions
8. Stand by mode:
   - Mode between observations
9. Data Downlink Mode
10. Safe Mode (if required)

* 6U CubeSat accommodation Study is completed

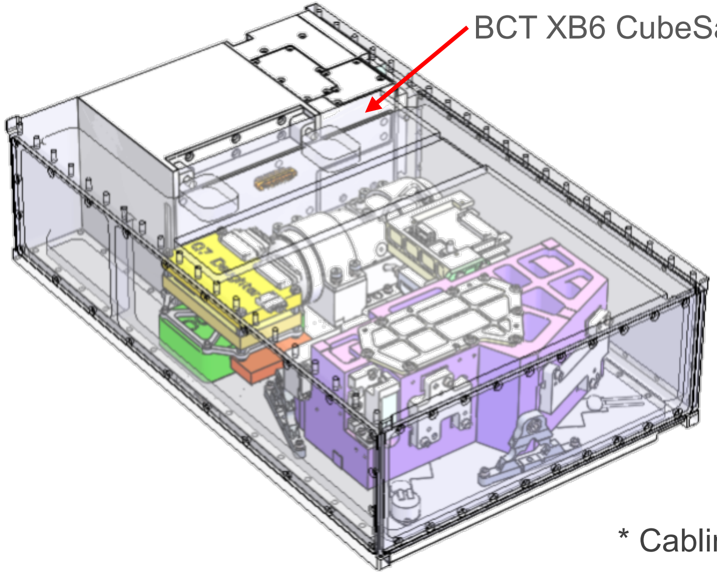
ARCSTONE Instrument in Fabrication



Volume:
Fits within up-to-date spacecraft bus CAD from BCT with at least 0.5mm clearance from all payload walls/features

Mass:
4.13 kg (*6 kg payload allowable*)

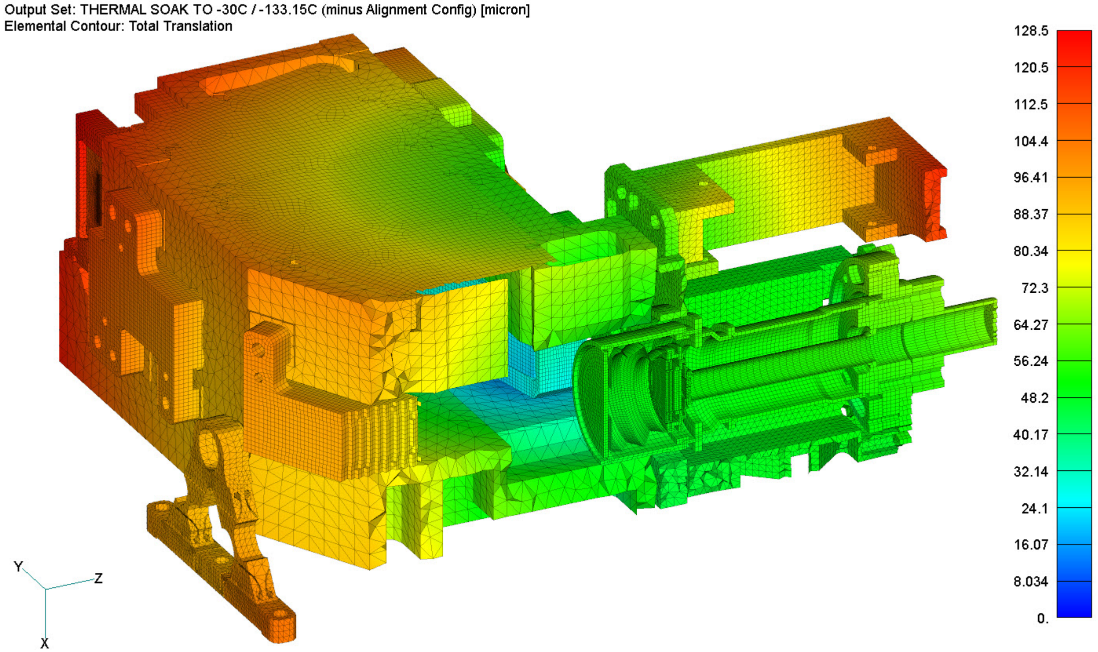
Power (all worst cases):
Science Mode: 23.83 W (*118 W peak allowable*)
Data Downlink Mode: 34.07 W
Stand By Mode: 15.5 W



* Cabling is not shown

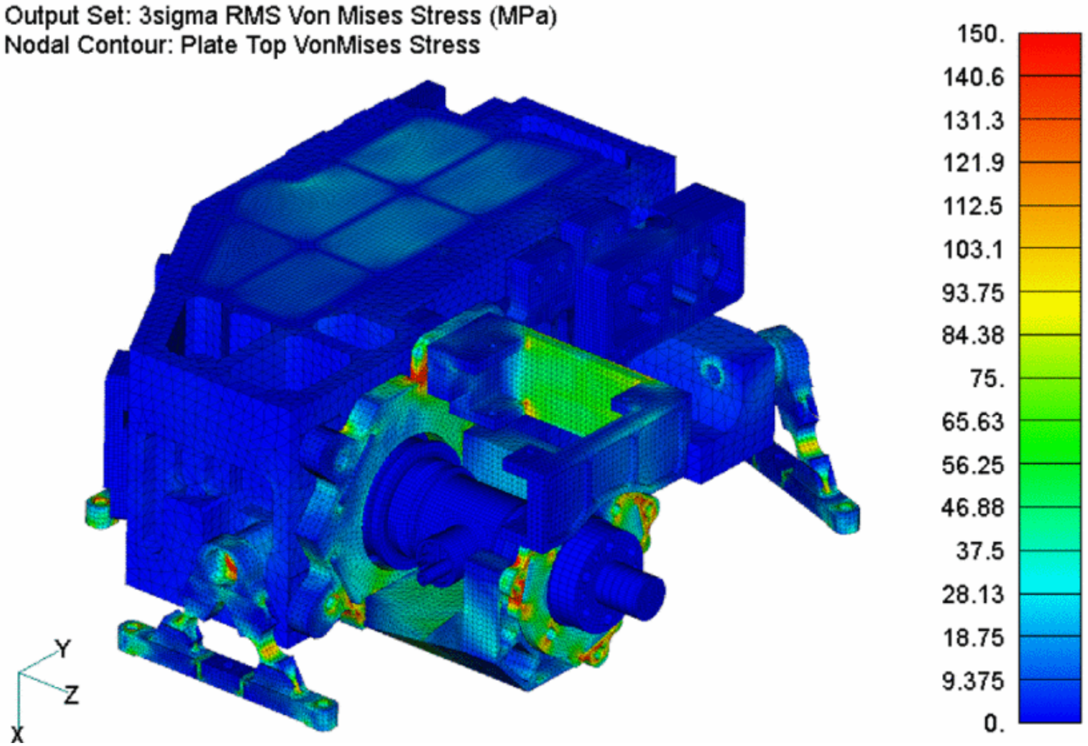
ARCSTONE Instrument Analysis

Performed Analysis:
STOP, Thermoelastic, Random Vibe



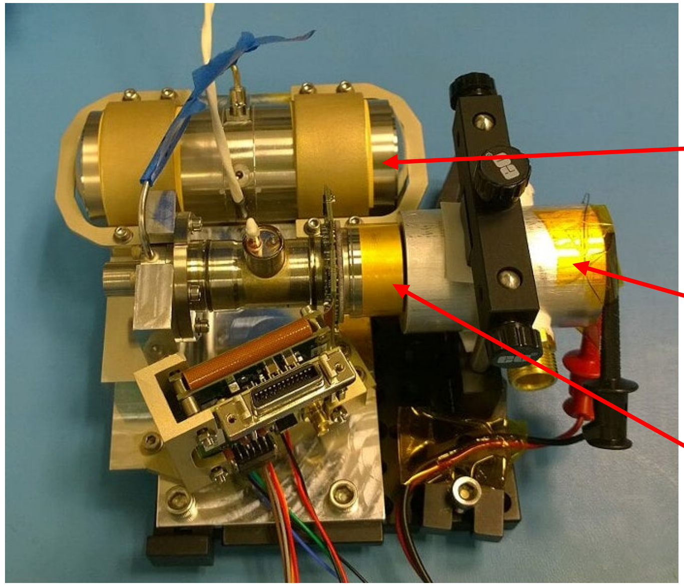
Optic bench displacements [microns] at −30°c.
Cutaway shows interior of camera dewar/cold finger.

Optic bench random vibration analysis.



ARCSTONE: SWIR IDCA (1 – 2.3 μm) Characterization

- **Sensor is uniform**
  - 745 hot/dead pixels
  - Only 2 pixels with no normal surrounding pixels
- **Vertical banding apparent in both dark and light images**
  - Eliminated through dark subtraction



Cryocooler

Cold Source

MCT Detector

Major Credits:
- IDCA selection/acceptance: Mike Cooney (NASA LaRC)
- Mechanical design: Trevor Jackson (NASA LaRC)
- IDCA characterization: Paul Smith (LASP, CU)

Integration time from 10^{-4} to 3.3 seconds !

SWIR IDCA Characterization Conclusions:

(1) SWIR IDCA *usable at 0.3% - 0.4% uncertainty level:*
- *Primary contributor to uncertainty is variation in the offset value between its measurements (repeatability over a few days).*
- *Offset value variation is a systematic uncertainty that cannot be mitigated through increased averaging, but may be lower during real data collecting operations, e.g. measuring offset before every lunar observation.*

(2) Camera linearity: *better than expected at 0.1% !*

(3) Initial Vacuum tests: *positive results !*

Full Spectral Range IDCA is essentially the same as SWIR IDCA (except for detector, OB filter, and integration time extended to 16 seconds)

ARCSTONE IIP: Status and Next Steps

Status:

- Design and STOP analysis completed for EDU instrument
- 6U CubeSat accommodation study completed
- Fabrication of instrument is in progress

Next Steps:

- Complete 6U CubeSat/Payload thermal study (September 2020)
- Complete fabrication of instrument (October 2020)
- Characterize Full Spectral Range IDCA (January 2021)
- Assemble instrument (February 2021)
- Calibrate instrument (May 2021)
- Field-test instrument with Sun and Moon measurements (TRL5, June 2021)

Testing ARCSTONE field equipment at NASA LaRC

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Recent Publications:

Swanson, R., C. Lukashin, M. Kehoe, M. Stebbins, H. Courrier, T. Jackson, M. Cooney, G. Kopp, P. Smith, C. Buleri, T. Stone, “The ARCSTONE Project to Calibrate Lunar Reflectance,” *IEEE Aerospace Proceedings,* 2020

Available online: https://ieeexplore.ieee.org/abstract/document/9172629

Stone, T.C., H. Kieffer, C. Lukashin, K. Turpie, “The Moon as a Climate-Quality Radiometric Calibration Reference,” *Remote Sens.,12*, 1837, 2020

Available online at https://www.mdpi.com/2072-4292/12/11/1837

ARCSTONE: Calibration of Lunar Spectral Reflectance from Space

http://arcstone.larc.nasa.gov

THANK YOU !