Calculation of the even-odd energy difference in superfluid Fermi systems using the pseudopotential theory

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PACS 74.20.Fg - BCS theory and its development

PACS 67.85.Lm – Degenerate Fermi gases

PACS 71.15.Dx – Computational methodology (Brillouin zone sampling, iterative diagonalization, pseudopotential construction)

Abstract – The pseudopotential theory is extended to the Bogoliubov-de Gennes equations to determine the excess energy when one atom is added to the trapped superfluid Fermi system with even number of atoms. Particular attention is paid to systems being at the Feshbach resonance point. The results for relatively small particle numbers are in harmony with the Monte Carlo calculations, but are also relevant for systems with larger particle numbers. Concerning the additional one quasiparticle state we define and determine two new universal numbers to characterize its widths.

The pseudopotential theory has proved to be an important tool in different areas, as for instance solid state physics (see, e.g., [1]), quantum chemistry (see, e.g., [2]), metal clusters (see, e.g., [3]). A general survey of the method regarding the interest of researchers in a wide range of applications is given in [4]. For a recent review about the approximations in electronic structure energy see ref. [5]. The idea was first formulated by Hellmann [6] and appeared also in Gombás' work [7]. The quantum mechanical foundation was started by Fényes [8] (for the early history of the pseudopotential theory see [9, 10]).

In the quantum mechanical treatment of many particle systems it is crucial to reduce the task from an N_{total} -particle problem to an N_{eff} particle one, where $N_{eff} \ll N_{total}$, when $N_{total} \gg 1$. In atoms, molecules and metals, e.g., N_{eff} is the number of valence electrons. The pseudopotential method works by taking the advantage of the fact that the large negative potential energy felt by a valence electron is almost completely cancelled by its large positive kinetic energy, which is due to the oscillations of its wave function inside the core. The method deals with pseudo wave functions which are smooth and easier to handle. One expects that similar situation arises in any fermion system put in an external potential when the analog of the valence electron problem in a broad sense is investigated. Characteristic differences also show up in case of trapped Fermi gases as a consequence of the shortrange nature of the interaction between the atoms and in particular when the superfluid state is realized, especially at the Feshbach resonance.

The main goal of the present paper is to formulate and use the pseudopotential method for trapped Fermi gases at zero temperature. For the sake of concreteness the internal degree of freedom of the atoms will be assumed to be twofold, one can have in mind for instance the ⁶Li isotopes as elements of a two component Fermi gas. The BCS side of the BCS-BEC crossover will be considered, where the interparticle interaction can be characterized by a negative *s*-wave scattering length *a*. We will focus in particular on the Feshbach resonance point, where universal properties set in, since $|a| \to \infty$ there, and the range of the interaction can be neglected (for a review see [11]).

We start by recalling the Bogoliubov-de Gennes (BdG) equations, which can be considered as the generalization of the Hartree-Fock equations to include pair correlations

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and conceived as equations of a mean-field theory [12, 13]. $-\epsilon$. The orthonormalization condition They are as follows:

$$[H + U_{int}] u_n + \Delta v_n = \epsilon_n u_n,$$

$$\Delta^* u_n - [H + U_{int}] v_n = \epsilon_n v_n, \qquad (1)$$

or in short

$$\hat{\Omega}\phi_n = \epsilon_n \phi_n \tag{2}$$

with $\phi_n = (u_n, v_n)$. The operator H can be written as

$$H = -\frac{\hbar^2}{2m}\nabla^2 + U_{ext}(\mathbf{r}) - \mu, \qquad (3)$$

where m stands for the mass of the atom, μ for the chemical potential and U_{ext} for the confining potential, which will be assumed to be harmonic in the applications. $U_{int}(\mathbf{r})$ denotes the diagonal part of the mean-field potential, while $\Delta(\mathbf{r})$ is the pair potential of the mean field theory. In a more general framework both $U_{int}(\mathbf{r})$ and $\Delta(\mathbf{r})$ contain contributions from correlation effects, but in the local density approximation (LDA), adopted also here, they are local quantities. We are interested in a system of particle number N + 1 with N being even. In particular we want to calculate the change in the ground state energy when adding one particle to the N-particle system. Since the theory adopted here is not working with fixed particle numbers it is more precise to speak about even- and odd-number parity states [12], but for sake of simplicity the therminology above is often used.

It is assumed in the following that the N-body problem is solved. It might mean the solution of the BdG equations in the BCS mean-field approach, in which case $U_{int}(\mathbf{r})$ is generally put zero [11, 14] and only $\Delta(\mathbf{r})$ should be solved self-consistently. It is argued that in the meanfield description this is a consistent choice [15]. At unitarity, however, there exists an extension of the density functional theory, which determines also $U_{int}(\mathbf{r})$ selfconsistently [16–18]. In our general considerations there is no need to specify which case is adopted. We do not use the results explicitly, suppose only their existence in principle and do not restrict ourselves to such particle numbers for which calculations have been realized. We avoid the details of the direct interaction U_{int} by fixing U_{int} and Δ using the solution of the N-body problem. In a more precise treatment one ought to solve the (N + 1)particle problem self-consistently. (In a normal system, as the electrongas in an atom for instance, this amounts to neglecting the core polarization by the valence electron.) In our model pseudopotential at unitarity universal constants will carry all the informations concerning the Natom problem. Therefore, we do not need to specify a regularisation procedure, as in refs. [14, 18].

It will be important in the following that eqs. (1) provide positive and negative eigenvalues as well representing the charge conjugation symmetry [17,19]. More concretely if $\phi^{(+)} = (u, v)$ is an eigenfunction with eigenvalue $\epsilon > 0$ then $\phi^{(-)} = (-v^*, u^*)$ is an eigenfunction with eigenvalue

$$\int d^3r \left[u_n^* u_m + v_n^* v_m\right] = \delta_{n,m} \tag{4}$$

leads to the conclusion that eigenfunctions belonging to eigenvalues of different sign are orthogonal.

Let us denote by $E^{(N)}$ the ground state energy of the N particle system, then, see, e.g., ref. [12]

$$E^{(N+1)} - E^{(N)} = \epsilon_{min} + \mu, \quad N \text{ is even.}$$
 (5)

For μ the chemical potential of the N-body system should be taken. The calculation of ϵ_{min} as the smallest positive eigenvalue of eq. (2) raises severe technical difficulties for large N, since the corresponding eigenfunction becomes highly oscillating. To avoid it and to make the determination of ϵ_{min} manageable even in such a case we generalize the pseudopotential method to the BdG equations by introducing the modified equation

$$\left(\hat{\Omega} + \hat{U}_{ps}\right)\tilde{\phi}_n = \epsilon_n\tilde{\phi}_n,\tag{6}$$

where ϕ_n stands for the pseudo wave function $\phi_n(\mathbf{r}) =$ $(\tilde{u}_n(\mathbf{r}), \tilde{v}_n(\mathbf{r}))$, and \tilde{U}_{ps} is the nonlocal pseudopotential

$$\hat{U}_{ps} = \sum_{t} \left\{ (\epsilon_{n} - \epsilon_{t}) \begin{pmatrix} u_{t}(\mathbf{r}) \\ v_{t}(\mathbf{r}) \end{pmatrix} (u_{t}^{*}(\mathbf{r}'), v_{t}^{*}(\mathbf{r}')) + (\epsilon_{n} + \epsilon_{t}) \begin{pmatrix} -v_{t}^{*}(\mathbf{r}) \\ u_{t}^{*}(\mathbf{r}) \end{pmatrix} (-v_{t}(\mathbf{r}'), u_{t}(\mathbf{r}')) \right\}.$$
(7)

Here ϵ_t is positive by definition. $\phi_t(\mathbf{r}) = (u_t(\mathbf{r}), v_t(\mathbf{r}))$'s are selected of those solutions of eq. (1) that are involved in building the functions $U_{int}(\mathbf{r})$ and $\Delta(\mathbf{r})$.

Two remarks are in order here. One has to include states of positive and negative eigenvalues to maintain the "charge conjugation" symmetry. The summation may extend only to a restricted set to avoid divergences, for instance to states

$$\left[\langle u_t | H + U_{int} | u_t \rangle + \langle v_t | H + U_{int} | v_t \rangle\right] < 0.$$
(8)

This choice will be assumed in the following. One can easily see that the solution of eq. (6) is not unique, any linear combination of $\phi_t^{(+)}$ and $\phi_t^{(-)}$ (satisfying the condition (8)) can be added to $\tilde{\phi}_n$ without changing the eigenvalue ϵ_n . As a consequence the most general form of U_{ps} can be written as

$$U_{ps} = \sum_{t} \left\{ \left| \phi_t^{(+)} \right\rangle \left\langle F_t^{(+)} \right| + \left| \phi_t^{(-)} \right\rangle \left\langle F_t^{(-)} \right| \right\}.$$
(9)

One can easily convince oneself that

$$\left\langle F_t^{(\pm)} \middle| \tilde{\phi}_n \right\rangle = (\epsilon_n \pm \epsilon_t) \left\langle \phi_t^{(\pm)} \middle| \tilde{\phi}_n \right\rangle;$$
 (10)

 $F^{(\pm)}$ are quite general functions. This expression of the pseudopotential is the generalization of the Austin, Heine Sham proposal [20]. A natural requirement might be that the pseudo wave function be as smooth as possible (formally, it has been proposed to minimize the kinetic energy [21,22], see also [2]). An obvious condition is that the pseudopotential should cancel a large part of the external and the local mean-field potentials (see e.g., [1,2,23]).

In this spirit we choose the generalized Austin type [1, 20] form

$$\hat{U}_{ps} = \begin{pmatrix} U_{ps}^{H} & U_{ps}^{\Delta} \\ U_{ps}^{\Delta *} & -U_{ps}^{H*} \end{pmatrix}$$
(11)

with

$$U_{ps}^{H} = -\sum_{t}' \left(u_{t}(\mathbf{r}) \left[U(\mathbf{r}') - \mu \right] u_{t}^{*}(\mathbf{r}') + v_{t}^{*}(\mathbf{r}) \left[U(\mathbf{r}') - \mu \right] v_{t}(\mathbf{r}') \right), \quad (12)$$

$$U_{ps}^{\Delta} = -\sum_{t}' \Big(v_t(\mathbf{r}) \left[U(\mathbf{r}') - \mu \right] u_t^*(\mathbf{r}') \\ -u_t^*(\mathbf{r}) \left[U(\mathbf{r}') - \mu \right] v_t(\mathbf{r}') \Big).$$
(13)

The prime on \sum means that the summation extends to states specified above and

$$U = U_{ext} + U_{int}.$$
 (14)

This expression of the pseudopotential corresponds to the choice

$$F_t^{(+)}(\mathbf{r}) = -(U(\mathbf{r}) - \mu) \left(u_t(\mathbf{r}), v_t(\mathbf{r}) \right), \qquad (15)$$

$$F_t^{(-)}(\mathbf{r}) = -(U(\mathbf{r}) - \mu) \left(-v_t^*(\mathbf{r}), u_t^*(\mathbf{r}) \right).$$
(16)

We consider in the following a trapped gas at the Feshbach resonance, where universal properties characterize the system. Let's define the Thomas-Fermi region (TFR), where the conditions for the Thomas-Fermi approach are fulfilled (here and in the following, of course, the generalized Thomas-Fermi theory including paircorrelations is meant supposing that the particle number is large enough [24, 25]). An essential contribution to $\sum_{t}' (u_t(\mathbf{r})u_t^*(\mathbf{r}') + v_t(\mathbf{r})v_t^*(\mathbf{r}'))$ arises only in the TFR then and it can be taken zero outside this region. Furthermore, it can be replaced to a good approximation, by a Diracdelta function $\delta(\mathbf{r} - \mathbf{r}')$ within the TFR if the number of particles is large enough, which in other words means that the nonlocality of the pseudopotential is disregarded. Concerning (13) it is decisive that in LDA u_t and v_t are plane waves and the condition (8) selects wave numbers in pairs $\mathbf{k}, -\mathbf{k}$, that altogether leads to $U_{ps}^{\Delta} = 0$. The pseudopotential becomes diagonal in this approximation

$$\hat{U}_{ps} = \begin{cases} \begin{pmatrix} \mu - U(\mathbf{r}) & 0\\ 0 & U(\mathbf{r}) - \mu \end{pmatrix} \delta(\mathbf{r} - \mathbf{r}'), & \mathbf{r} \in \text{TFR} \\ \mathbf{0}, & \text{otherwise.} \end{cases}$$
(17)

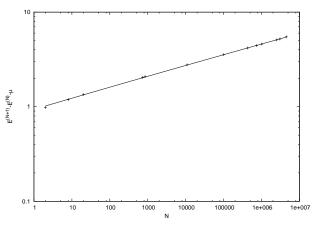


Fig. 1: Energy difference $E^{(N+1)} - E^{(N)} - \mu$ in units of $\hbar\omega$ as a function of N. Both axes are logarithmic, $\delta = 1.16$. Straight line shows an $N^{1/9}$ behavior.

This pseudopotential can be regarded as a model one, since the approximation breaks down near the border of the TFR. The total hamiltonian of the model reads as

$$\hat{\Omega} + \hat{U}_{ps} = \begin{pmatrix} -\frac{\hbar^2 \nabla^2}{2m} & \Delta(\mathbf{r}) \\ \Delta(\mathbf{r}) & \frac{\hbar^2 \nabla^2}{2m} \end{pmatrix}, \quad \mathbf{r} \in \text{TFR}, \quad (18)$$

and in the region $\mathbf{r} \notin \text{TFR}$:

$$\hat{\Omega} + \hat{U}_{ps} = \begin{pmatrix} -\frac{\hbar^2 \nabla^2}{2m} + U_{ext} - \mu & 0\\ 0 & \frac{\hbar^2 \nabla^2}{2m} - U_{ext} + \mu \end{pmatrix}.$$
(19)

Here we used the fact that $U_{int} = 0$ and $\Delta(\mathbf{r}) = 0$ in LDA outside the TFR.

At unitarity

$$\Delta(\mathbf{r}) = \delta\left(\mu - U_{ext}(\mathbf{r})\right),\tag{20}$$

where δ is a universal constant (see, e.g., [11]). In our calculations δ is an input. In contrast, δ is obtained as an output from the methods used in refs. [14, 18].

We have carried out the calculation for a spherical symmetric harmonic oscillator trap potential

$$U_{ext}(\mathbf{r}) = \frac{1}{2}m\omega^2 r^2 \equiv \mu x^2, \qquad (21)$$

where x = r/R with R being the Thomas-Fermi radius. First we have looked for a spherically symmetric solution (i.e., l = 0) to get the smallest eigenvalue of eqs. (6),(18),(19). In principle one can solve the BdG equations for $\mathbf{r} \in \text{TFR}$ and $\mathbf{r} \notin \text{TFR}$ by supposing that $\tilde{\phi}_n$ is finite and vanishes as $\mathbf{r} \to \infty$, and by carefully matching the wave functions at the border of TFR. But from the numerical point of view it is more controllable the method of expansion in a given basis set. In the numerics we expand both components of $\tilde{\phi}_n$ in the 3D harmonic oscillator basis $\phi_{nlm}(\mathbf{r})$. In the isotropic case l and m are good quantum numbers. Thus, the expansion is over the different n(radial quantum number) values. Matrix elements of the kinetic energy and U_{ext} in this basis are known exacly. Numeric integration is applied for the matrix elements of $\Delta(\mathbf{r})$ and \hat{U}_{ps} (which are nonzero for $\mathbf{r} \in TFR$). After calculating the matrix elements of $\hat{\Omega} + \hat{U}_{ps}$ a simple diagonalization yields the eigenvalue ϵ_n and the expansion coefficients of u_n and v_n . Special attention is paid to the size of the truncated basis $\{\phi_{nlm}(\mathbf{r})|0 \leq n \leq n_{max}\}$. (In the calculation we choose for n_{max} such a value for which the classical turning point of $\phi_{n_{max}lm}(\mathbf{r})$ is at least 1.5 times bigger than the Thomas-Fermi radius. This ensures that we have enough basis functions which "feel" both regions $\mathbf{r} \in \text{TFR}$ and $\mathbf{r} \notin \text{TFR}$.) In figs. 1–3 we had $n_{max} = 100$ and tested that choosing n_{max} for a slightly bigger value (cf. $n_{max}=150$) the calculated ϵ_n , u_n and v_n are practically the same as for $n_{max} = 100$.

In fig. 1. the energy difference $E^{(N+1)} - E^{(N)} - \mu$ as given in eq. (5) is plotted. It follows the $N^{1/9}$ law predicted by Son [26]. Concerning the accuracy of the prefactor, it is remarkable that even for a relatively small number of particles as N = 20 our result lies between the Monte Carlo findings [27,28] and the result of the superfluid density functional calculation [16]. (Note that at this particle number one expects that the extra particle is in the *s*state). In fig. 2. the components of the eigenfunction are drawn. They have nodes, the appearence of which can be traced back to the fact that the operator (18), (19) couples the bare (normal) states of positive and negative eigenvalues.

The Thomas-Fermi approach breaks down in the vicinity of the TF-radius R, which circumstance may question the applicability of the pseudopotential (17) there. One can show, however, that the width of the solution of eqs. (6), (18), (19) is $\delta r \propto R(d/R)^{4/3}$, where d is the oscillator length $d = \sqrt{\hbar/(m\omega)}$, while the width of the surface region (where gradient corrections are important) scales as $\propto R(d/R)^4$ [25]. At large particle numbers $R \gg d$ is valid, which ensures that the error is small when extending the Thomas-Fermi density to the border. This makes possible to determine two new universal numbers. In fig. 3. the functions $w_u = A_u(\hbar\omega/\mu)^{+2/3}$ and $w_v = A_v(\hbar\omega/\mu)^{+2/3}$ are drawn together with the numerically determined widths of \tilde{u} and \tilde{v} , in units of R. The fitting provides $A_u = 1.2774$ and $A_v = 1.14616$.

For the solution in case of $l \neq 0$ one needs a refinement of the model (Then, of course, ϵ_{min} is to be understood for the given l). First of all we quote that at resonance

$$U(\mathbf{r}) - \mu = \frac{1}{\xi} \left(U_{ext}(\mathbf{r}) - \mu \right), \qquad (22)$$

where ξ is a universal constant, whose recent values are 0.372(0.005) [29] and 0.376(5) [30]. The complete cancellation of $U(\mathbf{r}) - \mu$ by the pseudopotential occurs now only in a more restricted region, namely in between the zeros $0 < r_1 < r_2 < R$ of $V(\mathbf{r}, \xi)$ defined as

$$V(\mathbf{r},\xi) = \frac{1}{\xi} \left(U_{ext}(\mathbf{r}) - \mu \right) + \frac{\hbar^2}{2m} \frac{l(l+1)}{r^2}.$$
 (23)

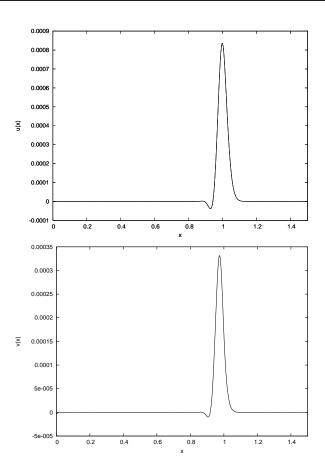


Fig. 2: The \tilde{u} (upper part) and the \tilde{v} (lower part) components of the pseudo wave function as a function of the dimensionless variable of x = r/R. Parameters: $\mu = 100 \hbar \omega$, $\epsilon_{min} = 4.437 \hbar \omega$, $\delta = 1.16$.

Outside this region, but within the Thomas-Fermi radius R the total potential is $V(\mathbf{r}, \xi)$ in the radial equation while outside R it is $V(\mathbf{r}, 1)$. Semiclassically in the region $r_1 < r < r_2$ the radial kinetic energy is positive. Note that in the semiclassical treatment the factor l(l + 1) in the centrifugal energy ought to be replaced by $(l + 1/2)^2$. We keep the quantum-mechanical expression, however, to be able to extrapolate the results to small *l*-values.

The solutions of the radial equations corresponding to the potential discussed above have interesting features. The minimal excitation energy scales as $\sim \mu^{-1} \sim N^{-1/3}$ as predicted by Son [26], but the prefactor depends on the value of ξ and is smaller than the result one gets from the estimated centrifugal energy at the Thomas-Fermi radius R. A detailed analysis shows that this effect is mainly due to the two component nature of the wave function. Furthermore, for larger *l*-values the excitation energy is no more proportional to l(l+1). At about $l \approx l_{max}$, defined by the condition $r_1 = r_2$, the energy curve merges into the excitation spectrum for l = 0, which might detect a breakdown of the mean-field-type theory at such excitation energies.

One can show that by replacing the wave function of

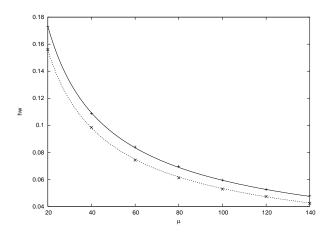


Fig. 3: The width of \tilde{u} (+) and \tilde{v} (×) belonging to the pseudo wave function for ϵ_{min} as a function of μ . The widths are measured in unit of R and μ in $\hbar\omega$. The solid and dashed lines show the w_u and w_v curves (see text).

the (N + 1)-th atom by the pseudo wave function the wave function of the total system does not alter. This requirement was the starting point in [31] working within the Hartree-Fock theory (see also ref. [2] for the generalization to the (N+2)-particle problem). The ground state of the N-particle system can be written as

$$|\Psi_0\rangle = \prod_t \alpha_t^{(+)}|0\rangle = \prod_t \alpha_t^{(-)^+}|0\rangle, \qquad (24)$$

where $\alpha^{(-)^+}$ and $\alpha^{(+)^+}$ are the quasiparticle creation operators in negative and positive energy states, respectively, with $\alpha^{(\pm)}$ the corresponding destruction operators. $|0\rangle$ stands for the vacuum of the atoms.

The state of the (N + 1)-body system can be given as a one-quasiparticle state

$$\alpha_n^+ |\Psi_0\rangle. \tag{25}$$

The corresponding state written in terms of the pseudo wave function reads as

$$\left(\alpha_n^+ + \sum_t c_t^{(+)} \alpha_t^{(+)} + \sum_t c_t^{(-)^*} \alpha_t^{(-)^+}\right) |\Psi_0\rangle, \qquad (26)$$

which is equivalent to (25) as can be seen from the expression of $|\Psi_0\rangle$. Note that normalization constants have not been included in the discussion above. The coefficients in (26) are

$$c_t^{(\pm)} = (\epsilon_n \pm \epsilon_t) \langle \phi_t^{(\pm)} | \tilde{\phi}_n \rangle.$$
(27)

In conclusion we have calculated the energy of an extra particle in trapped Fermi gases at unitarity by generalizing the pseudopotential theory to the BdG-type equation. The background N-particle problem has been treated within the Thomas-Fermi theory. This approximation could be improved by including the Weizsäcker-type correction as determined in case of the unitary system [25]. Along with such an extension it would be appropriate to take into account also the nonlocality of the pseudopotential, a feature lost when obtaining (11). All these are left for future work.

Some final remarks are in order. In our treatment all about the neighbouring even-number state are compressed into two universal numbers δ and ξ . The recent value of ξ is somewhat smaller than previously used in different matching processes, which might influences the other parameter values. This makes our choice to take the bare atomic mass in our calculation reasonable. Concerning δ we have taken its mean-field value by the same reason and by the expectation that it alters only slightly when using different models.

Applications of more elaborated forms of the pseudopotential to other systems along with details of the present work will be published elsewhere.

The present work has been partially supported by the Hungarian Scientific Research Fund under Grant Nos. OTKA 77534/77629 and OTKA 75529.

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