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Cephid Velocity Curves

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ABSTRACTS

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30.03 <u>Cepheid Velocity Curves</u>, N. R. SIMON, A. S. LEE and T. J. TEAYS, <u>Univ. of Neb.-Lincoln</u>. Fourier decompositions are made for a number of observed velocity curves of classical Cepheids. The observations are fit with Fourier series of the form

$v_{obs} = A_0 - A_i \sin(i\omega t + \phi_i)$,

where the index i runs from 1 to 4 or 1 to 8 depending upon the requirements of the data. Although the sample of stars is small, we show that the Hertzsprung progression expresses itself quantitatively in terms of the low-order Fourier coefficients, particularly the quantity $\phi_{21} \equiv \phi_2 - 2\phi_1$. This result complements a similar finding for the light curves. When the Fourier decompositions of the velocity curves are compared with those of some theoretical models (Vermury and Stothers 1978, <u>Ap. J. 225</u>, 939), new evidence is uncovered favoring a resonance explanation for the "bump" sequence. This research was partially supported by the National Science Foundation under Grant No. AST79-05290.

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