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ABSTRACT

6
7 The near-surface wind and temperature regime at three points in the Atacama Desert of
8 northern Chile is described using two-year multi-level measurements from 80-m towers lo-
9 cated in an altitude range between 2100 and 2700 m ASL. The data reveal the frequent
10 development of strong nocturnal drainage flows at all sites. Down-valley nose-shaped wind
11 speed profiles are observed with maximum values occurring at heights between 20 m and 60 m
12 AGL. The flow intensity shows considerable inter-daily variability and a seasonal modulation
13 of maximum speeds, which in the cold season can attain hourly average values larger than
14 20 m s^{-1} . Turbulent mixing appears significant over the full tower layer, affecting the cur-
15 vature of the nighttime temperature profile and possibly explaining the observed increase of
16 surface temperatures in the down-valley direction. Nocturnal valley winds and temperatures
17 are weakly controlled by upper-air conditions observed at the nearest aerological station.
18 Estimates of terms in the momentum budget for the development and the quasi-stationary
19 phases of the down-valley flows suggest that the pressure gradient force due to the near-
20 surface cooling along the sloping valley axes plays an important role in these drainage flows.
21 A scale for the jet nose height of equilibrium turbulent down-slope jets is proposed, based on
22 surface friction velocity and surface inversion intensity. At one of the sites this scale explains
23 about 70% of the case-to-case observed variance of jet nose heights. Further modeling and
24 observational work is needed, however, in order to better define the dynamics, extent and
25 turbulence structure of this flow system, which has significant wind-energy, climatic and
26 environmental implications.

1. Introduction

The Atacama Desert in northern Chile is renowned for its hyper-aridity caused by geographic and climatic factors (Garreaud and Aceituno 2007; Garreaud et al. 2010; Garreaud 2010). To its east, the Andes Cordillera rises to altitudes that exceed 6000 m precluding moisture transport from the Amazon basin. To its west, the subtropical South East Pacific (SEP) Anticyclone and the cold coastal waters of the Chile-Peru current provide for a very stable large-scale environment. The climate of the Atacama Desert is thus characterized by low synoptic variability, very infrequent precipitation, extremely clear skies (except near the coast where stratocumulus are common), intense solar radiation, large diurnal thermal amplitudes, strongly sloping topography, and a very arid surface with little vegetation and negligible moisture availability. In terms of surface winds, these characteristics constitute almost ideal conditions for the observation of terrain-induced thermal circulations. The existence in the Atacama Desert region of near-surface winds driven by the thermal contrasts of the cold ocean, the desert and the Andes has been described by Schmidt (1999) and documented through short-term field campaigns by Rutllant and Ulriksen (1979) and Rutllant et al. (2003). These winds have been hypothesized to affect the lower-tropospheric stability and regional hyper-aridity (Rutllant et al. 2003) and the transport of continental aerosols (natural and anthropogenic) towards the SEP stratocumulus layer over the ocean to the west (Chand et al. 2010). However, comprehensive and extended datasets describing surface winds in the Atacama Desert have only recently become available. Over the last few years Chilean energy authorities have conducted a series of projects aimed at evaluating the wind energy potential of the northern part of the country. As part of these efforts, three 80-m meteorological towers with multi-level wind and temperature measurements were installed in the Atacama Desert between 22 °S and 23 °S. The main purpose of the present communication is to document the strong nocturnal down-valley flows unveiled by these observations. In particular, our objectives are 1) to describe the mean characteristics and variability of the down-valley flow, with special emphasis on the local co-variability of wind and temperature

54 profiles at each site, and 2) to provide an observational evaluation of the regional and local
55 controls affecting these flows. The wind energy implications of these measurements, as well
56 as their numerical modeling and predictability assessment are being addressed in parallel
57 studies and will be reported elsewhere.

58 The variability of near-surface winds is affected by factors spanning a large range of
59 scales. At the synoptic-scale, pressure gradients provide for the day-to-day variability of
60 surface winds, while local obstacles and surface properties determine their finer-grade struc-
61 ture. In between, regional topography can be a powerful forcing of surface winds, playing a
62 prominent role in shaping their diurnal cycles in connection to mesoscale pressure gradients
63 ensuing from the daytime/nighttime heating/cooling of complex terrain. Whiteman (2000)
64 reviews the basic forms in which this thermal forcing manifests, distinguishing between slope
65 winds, along-valley winds, cross-valley winds, and mountain-plain winds, depending on the
66 temperature contrasts driving the pressure gradients (along a slope, within a valley or be-
67 tween a valley and a plain, across the valley sides, or between a mountain and a plain, for
68 each of the aforementioned cases, respectively). Here we describe the slope-wind and along-
69 valley wind mechanisms, which appear to be most relevant to the flows presented in this
70 work.

71 Nocturnal slope winds (also referred to as katabatic winds) are driven by an along-the-
72 slope pressure gradient arising as the air close to an inclined surface cools relative to the
73 air above. Their study has a long history, with various authors considering different addi-
74 tional terms in the momentum budget to compensate the buoyant downward acceleration.
75 A comprehensive dynamical analysis of this mechanism has been given by Mahrt (1982).
76 In typical conditions, katabatic winds are generally weak and shallow and are considered
77 relevant to explain nocturnal flow down the slope of valley sides. Strong and extensive kata-
78 batic flows, however, have been reported over Antarctica (Parish 1982; Parish and Bromwich
79 1987), where the long polar night, the cold ice-covered surface and the topography combine
80 to produce regions over which “extraordinary” katabatic winds develop (Bromwich 1989;

81 Wendler et al. 1997). They are characterized by near-surface average speeds between 10
82 and 30 m s^{-1} and very high directional persistence. Nocturnal down-slope and down-valley
83 katabatic flows have also been studied in the Rocky Mountains in western North America
84 (e.g. Savage et al. 2008; Whiteman and Zhong 2008). Maximum wind speeds in these cases
85 are generally less than 10 m s^{-1} , with drainage layers that can reach up to $\sim 250 \text{ m AGL}$.

86 The nocturnal down-valley wind mechanism, on the other hand, does not require an in-
87 clined bottom surface at all, as illustrated by the horizontal valley floors used in idealized
88 modeling studies of these flows (e.g. Schmidli et al. 2011). The along-valley pressure gra-
89 dient driving the surface winds in this case is caused by the relative cooling of the valley's
90 atmosphere as compared to a nearby plain (Whiteman (1990); 2000). Crucial in this mech-
91 anism is the geometric shape of the valley in the cross-valley and along-valley directions,
92 as well as the surface energy budget and radiative effects. Besides their thermal forcing,
93 valley winds can also be affected by large-scale winds and pressure gradients. Whiteman
94 and Doran (1993) categorize these influences in terms of the relationship between the di-
95 rections of the valley winds and the upper-air (geostrophic) winds, defining the mechanisms
96 of downward momentum transport (with valley winds following closely the upper-air wind
97 direction), forced channeling (with valley winds aligned to the valley axis in the direction of
98 the upper-air wind) and pressure driven channeling (with valley winds aligned to the val-
99 ley axis in the direction of the upper-air pressure gradient, perpendicular to the upper-air
100 geostrophic wind).

101 Although the mechanisms described above appear well defined and distinct, real cases in
102 complex terrain will generally be affected by a combination of them. In our case, for example,
103 although the three sites studied are located along valley axes, we cannot discard a priori
104 effects of mountain-plain winds associated to the large Andes Cordillera massif existing to
105 the east, nor slope-wind effects due to the inclination of the valley floors. Therefore, following
106 Whiteman (1990), perhaps a proper name for the nocturnal flows documented here would
107 be the mixed term of drainage winds. We will use the term down-valley winds, however,

108 considering that this is their most persistent direction, leaving the issue of the dynamical
109 forcing to a discussion in the final part of this paper, and to subsequent modeling work.

110 2. Site and data

111 *a. Regional topography*

112 Figure 1 shows the topography of the study region and the location of the three 80-m
113 towers analyzed here. All are located along the central axes of wide, gently sloping valleys.
114 Tower SGE is located along Arriero Creek, an east-west oriented dry streambed close to
115 the town of Sierra Gorda. Towers CAN and CAO are located along the Loa River valley,
116 up-valley and down-valley of Calama City ($\sim 140,000$ inhab.), respectively. At the CAO
117 site the valley is also east-west, while at CAN the valley is oriented north to south. Figure
118 2 shows a closer look at the terrain surrounding the towers, including along- and cross-
119 valley topographical profiles. The average slopes along the valley axes are relatively steep,
120 with values ranging between 1.3 and 1.7%, while the cross-sections show some differences
121 between the three sites. CAN, for example, is located where the Loa River valley widens
122 significantly as compared with the up- and down-valley topography, resulting in a local cross-
123 valley transect with a ~ 30 km flat floor. In contrast, transects at CAO and SGE show a
124 well defined U-shape, deeper and wider at CAO as compared to SGE. These variations in
125 relief may explain some of the differences that the down-valley winds exhibit in each site, as
126 shown in Section 3.

127 *b. Available data*

128 The three towers were installed in February 2011. Figure 3 provides an overview of
129 hourly wind speeds at 40 m available at the time of this writing. As the planned measuring
130 period of the towers was limited to 2 years and our focus of analysis will be the cold season

131 between May-October, the datasets can be considered complete for our purposes. Figure 3
132 shows several gaps in data availability at CAN and CAO. Those in 2011 were mostly due
133 to signal interference of the remote data transmission, which was corrected in August 2011,
134 while the lost month of June 2012 at CAO was due to vandalism. For the months from
135 May to October, data availability at CAN, CAO and SGE amounts to 86%, 81%, and 99%,
136 respectively, providing a fair to good representation of winter-spring conditions of years 2011-
137 2012, which presented weak-moderate La Niña conditions, as indicated by the Multivariate
138 ENSO Index of Wolter and Timlin (2011).

139 The main measurements in the towers are wind speed at 5 levels, wind direction at 3
140 levels, and air temperature at 3 levels (heights and sensors provided in Table 1). Each
141 of the sensors used included individual calibration certificates against reference standards,
142 making their observations directly comparable among them. Ancillary meteorological data
143 include near-surface relative humidity, atmospheric pressure, and solar radiation, all data
144 available as 10-minute averages computed from 1 Hz samples. As a means of relating the
145 tower measurements to upper-air conditions, we use 12 UTC (08 LT) vertical profiles of wind
146 and temperature measured at the operational radiosonde station of Antofagasta, located at
147 the coast about 150 km to the WSW of SGE (see Fig. 1). A long-term characterization of
148 the regional lower troposphere based on these upper-air observations can be found in Muñoz
149 et al. (2011).

150 *c. Down-valley flow indices*

151 The observational characterization presented in Section 3 is based upon indices describing
152 the intensity, directional persistence, jet nose height, and stability of the nocturnal flows
153 observed at the three sites. These indices are defined and justified next.

154 At the three sites considered the daytime and nighttime low-level wind regimes are
155 strongly differentiated, as shown by the mean diurnal cycles of 40-m wind speed shown
156 in Fig. 4a. The nighttime phase of the wind regime begins a few hours after sunset, at

157 around 21 LT, when wind speeds start intensifying, attaining maximum mean values by sun-
158 rise, at about 07 LT. The daytime phase, on the other hand, has maximum speeds between
159 15 and 18 LT. Wind directions also show strong contrasts between day and night regimes,
160 as illustrated by the most frequent directions represented with circles in Fig. 4a. At CAN,
161 nighttime (daytime) flow is predominantly from the NNE (SW), while at CAO and SGE
162 the corresponding directions are from the ENE (W). Based on the well-defined and regular
163 diurnal pattern of near-surface winds, the 02-06 LT period is chosen as representative of
164 well-developed nighttime flows.

165 Figures 4b-d show the joint wind speed-wind direction frequency distribution at 40 m
166 AGL for all observations in the 02-06 LT period. The very marked directional preference of
167 the nighttime flow is evident in these figures with 81%, 92%, and 94% of the observations at
168 these hours falling in a 45° sector centered at 10° , 75° , and 80° for stations CAN, CAO, and
169 SGE, respectively. We shall call these sectors as down-valley direction sectors, and define,
170 for each night and site, the index FDV40 as the fraction of 40-m wind direction observations
171 in the 02-06 LT period falling in the corresponding down-valley direction sector. Values
172 of FDV40 are generally high, indicative of the large topographic control of these nocturnal
173 flows. The wind intensity, however, is more variable, as shown by the wind speed ranges in
174 Figs. 4b-d), with 10-minute averages varying generally between 5 m s^{-1} and 20 m s^{-1} . The
175 simple average of 40-m wind speed in the 02-06 LT (WS40) period will be used in Section 3a
176 as intensity index for the nocturnal flows at each site (their variability has been shown in Fig.
177 3). In Sections 3b and 3c, on the other hand, the 40-m down-valley projected wind speed
178 (DVWS40) will be used as an index better suited for comparing down-valley flow intensities
179 among the different sites.

180 The availability of wind speed measurements in the 10-80 m height range has allowed
181 the observation that nighttime flows at these sites frequently take the form of a low-level
182 jet, with wind speeds reaching a local maximum at a certain height above the surface. To
183 describe this, we define the ZX index as the height of the wind speed maximum for each

184 10-minute observation. Figure 5 shows the frequency distribution of ZX and how it varies in
185 the 18-08 LT period. While during daytime and at the beginning of the evening transition
186 ZX corresponds almost always to the highest measurement level, with no indication of a
187 low-level jet, during the night the most common values of ZX are between 20 m and 60
188 m AGL depending on site and hour. In the early evening the maximum wind speeds are
189 most commonly found in the first measurement level (10 m) at all three stations. As time
190 passes, the nose of the jet profile increases in height, being typically at 20 m at SGE and at
191 40 m at CAN and CAO, suggesting that the evening wind reversal is a bottom-up process.
192 The nighttime ZX frequency distributions are less variable at SGE, with more variability
193 observed at the other two sites. At CAO, in particular, a secondary maximum is observed
194 after 02 LT at the 80-m level, indicating that in many cases the nose of the jet is further
195 aloft, or there is no jet structure at all.

196 Finally, the static stability in the tower layer will be described by means of indices
197 computed as 02-06 LT averages of the difference of temperatures measured at the upper
198 (DTU= $T_{80}-T_{40}$) and the lower (DTL= $T_{40}-T_{02}$) levels equipped with temperature sensors.
199 An additional temperature-related index will be the 02-06 LT average of the curvature of
200 the temperature vertical profile, estimated as $CT=T_{80}-2*T_{40}+T_{02}$.

201 3. Results

202 a. *Characterization of down-valley jets and stability*

203 For each site we define a down-valley low-level jet (DVJ) case as a night in which the
204 following conditions are met:

- 205 • $FDV_{40} > 85\%$, i.e. more than 85% of the 02-06 LT wind direction observations fall in
206 the down-valley direction sector, and
- 207 • $ZX < 80$ m in more than 85% of the 02-06 LT observations.

208 These conditions are rather strict in demanding persistent down-valley nose-shaped wind
209 profiles, but they still permit the inclusion of a significant number of cases, as illustrated by
210 the time series in Fig. 3 and the occurrence frequency of DVJ cases in Fig. 6a. Site SGE has
211 the maximum percentage of cases along the year (79% annual average), while CAN and CAO
212 have bi-monthly frequencies between 20% and 60%. It must be stressed that cases left out
213 may also be low-level jets, but they have a larger directional variability or have maximum
214 speeds at 80 m or above (especially at CAO), rendering them less suited for characterizing
215 wind and temperature profiles below and above the jet nose.

216 Figure 6b describes the variability of DVJ intensities, as measured by the inter-quartile
217 ranges of WS40 for each site. Typical magnitudes vary in the 5-18 m s⁻¹ range, with an
218 appreciable seasonal modulation showing larger values between May and October (austral
219 winter and spring). Figures 6c-d describe, for DVJ cases, the variability of the stability
220 indices DTL and DTU, respectively. The stability closer to the surface shows little seasonal
221 variation, with temperature differences typically in the 2-4 °C range all year long. The
222 stability in the upper layer, however, has a strong seasonal modulation in phase with the
223 jet intensity (Fig. 6b). This co-variability between wind speed and temperature profiles in
224 DVJs is further analyzed next.

225 In order to, at least partially, filter out the seasonal co-variability of winds and stability,
226 we restrict the subsequent analysis to the months from May to October, leaving 169, 118 and
227 300 DVJ cases for CAN, CAO, and SGE, respectively. Figure 7 displays the mean vertical
228 profiles of wind speed and temperature, distinguishing between cases of weak and strong
229 DVJs (using the median of WS40 as discriminator). A connection between the wind speed
230 and temperature profiles is suggested by Fig. 7, although not as much in terms of the bulk
231 temperature gradient or stability, but rather in terms of the curvature of the temperature
232 profile. Indeed, especially at CAN and SGE, the curvature of the mean temperature profiles
233 changes sign between the weak and strong DVJs. The case-to-case co-variability between
234 temperature and wind speed profiles is examined in Fig. 8. At the three sites there is a

235 marked contrast in the relationship between the DVJ intensity and the stability in the lower
236 and the upper layers. While stronger down-valley flows are generally associated to reduced
237 stability in the lower layer, the opposite association exists with the stability in the upper
238 layer. As a result, a positive correlation between the intensity of the flow and the bulk
239 curvature of the temperature profile is observed, as shown in the rightmost scatter plots of
240 Fig. 8. These relationships are clearer at CAN and SGE in which CT becomes positive in
241 a large number of cases. At CAO, however, only few cases reach a positive curvature in the
242 temperature profile. Reasons for the differences among sites have not been elucidated yet,
243 but topographic differences along and across the valleys may probably play a role (see Fig.
244 2).

245 We interpret the positive correlation between temperature curvature and wind speed as a
246 consequence of the curvature-turbulence relationship proposed by André and Mahrt (1982)
247 (see also Edwards et al. 2006). According to this mechanism, turbulent mixing in the stable
248 boundary layer induces a positive curvature in the temperature profile, opposite to the effect
249 of longwave radiation. As turbulence may be expected to increase with stronger low-level
250 winds, this mixing effect might explain the right hand scatter plots in Fig. 8. Unfortunately,
251 direct measurements of turbulence are not available at these towers. As surrogates, we
252 present in Fig. 9 vertical profiles of turbulence-related variables. The upper panels show
253 distributions of gradient Richardson numbers (Ri) computed based on the temperature and
254 wind profiles. Three layers can be recognized in these figures. Below the jet nose all three
255 sites show very low Ri values, indicative that in this layer wind shear production can sustain
256 turbulence in the presence of buoyant consumption. Largest values of Ri are found in the
257 region of the jet nose due to the low shear existing there. Above the jet, where shear is
258 again significant, Ri values are smaller, but not as much as in the sub-jet layer (especially at
259 CAO and SGE). Shear production in this layer may not be sufficient to sustain turbulence,
260 although turbulence can still exist if transport from the sub-jet layer is important. As an
261 approximation to the turbulence intensity in the different layers, the lower panels in Fig. 9

262 show vertical profiles of the standard deviation of wind speed (σ_{ws}) measured by the propeller
263 anemometers, distinguishing again between weak and strong DVJ cases. Values of σ_{ws} in all
264 levels are comparable, suggesting that turbulence exists throughout the layer sampled by the
265 towers. Largest values of σ_{ws} are found in the lowest measurement levels, consistent with the
266 smaller Ri values found there. The profiles also tend to present a minimum of σ_{ws} at the jet
267 nose level, in accordance again with the small shear production there and the high Ri values.
268 An interesting feature is the change of these mean profiles with DVJ intensity. The stronger
269 down-valley flows are associated to larger values of σ_{ws} in the sub-jet layer, but to smaller
270 values in the layer above the nose. The larger stability existing in this layer associated to the
271 stronger DVJs (see center scatter plots in Fig. 8) may explain this relationship. The overall
272 prevalence of turbulence over the full tower layer and its minimum at the level of the jet
273 suggested by these observations are consistent with those reported by Conangla and Cuxart
274 (2006) for a nocturnal jet of $\sim 10 \text{ m s}^{-1}$ intensity and $\sim 65 \text{ m}$ nose height measured with a
275 100 m instrumented tower in the north Castilian Plateau, Spain.

276 *b. Regional control of down-valley flows and temperatures*

277 1) DOWN-VALLEY WINDS

278 We explore now the co-variability of winds between the three sites and with regional
279 forcings. As flow intensity indices we consider the 40-m down-valley projected wind speed
280 (DVWS40) averaged over hours 02-06 LT for the months of May to October. Upper panels
281 in Fig. 10 show the scatter diagrams of DVWS40 measured at the three towers. A generally
282 good relationship between the down-valley flow strength at all sites is appreciated, especially
283 for the high-values range. The best correlation is found between stations CAO and SGE
284 ($r \sim 0.91$), which may be explained by the fact that they are located in contiguous zonally
285 oriented valleys, in contrast to the meridionally oriented valley of station CAN (see Fig. 2).
286 Therefore, large-scale pressure patterns that modulate down-valley flows in the region may

287 produce a more similar control on CAO and SGE.

288 The Antofagasta radiosondes provide an opportunity for a first-order evaluation of the
289 large-scale controls on the nocturnal down-valley flows. For this purpose, we computed the
290 correlation coefficients between DVWS40 and three indices derived from the Antofagasta
291 soundings (Table 2). In all cases the largest correlations are found for Antofagasta variables
292 in the 2000-3000 m ASL range, coincident with the altitudinal range of the 80-m towers.
293 The first row in Table 2 shows the correlation between the down-valley flows at the three
294 sites and the winds measured by the 12-UTC soundings. More specifically, the correlation
295 is with the component of the upper-air winds coming from the ESE-SE sector, for which the
296 correlation was highest for the three sites. A modest zero-lag correlation, $r \sim 0.5$, is found,
297 with no marked differences among sites.

298 The second and third rows in Table 2 quantify the relationship between the intensity
299 of the down-valley winds and the free-troposphere temperature structure. The down-valley
300 wind mechanism suggests that the relative cooling of the valley air mass relative to the
301 free troposphere at the same altitude should enhance the drainage winds. Indeed, for the
302 three sites the largest correlations found are those between DVWS40 and the difference
303 between the Antofagasta temperature averaged in the 2000-3000 layer (TANTO) and the
304 tower temperatures averaged over the three data levels (TZNIGHT). This suggests that the
305 down-valley wind mechanism may be at work in these flows, especially considering that cor-
306 relations with the temperature difference generally increase as compared to those calculated
307 using the radiosonde temperature alone (comparing second and third row values in Table 2).
308 Nonetheless, all these correlations are relatively modest, and the scatter plots associated to
309 them display a large dispersion (not shown), which may be partially explained by the ~ 200
310 km distance between Antofagasta and the tower sites.

2) TEMPERATURE PATTERNS

Middle and lower panels in Fig. 10 show the scatter diagrams between DTL and DTU values among the three sites. The correlations in these cases are smaller than in the case of the flow strengths, suggesting that the adjustment of the temperature profile to the down-valley flow is influenced by local factors. In this regard, the small range of stabilities measured at CAO in comparison with CAN and SGE could be related to the presence of the city of Calama upwind of CAO. The large roughness associated to the urban land use may preclude the development of intense inversions near the surface.

In terms of the temperature comparison among the different sites, a noticeable feature apparent in Fig. 7 is the higher temperatures in CAO as compared to its up-valley counterpart CAN. While the down-valley mechanism requires colder temperatures upwind at the same altitude, near the surface the temperature usually drops in the down-valley direction, especially if the topography forces the drainage flow to accumulate and form cold pools. In our case, Figs. 7b-f show that the down-valley site CAO has a near-surface temperature about 2.5 °C higher than that measured at CAN. In order to better understand this regional temperature pattern, Fig. 11 shows the average nighttime temperatures measured at the three 80m-towers as a function of altitude. We have also included in the figure corresponding 5-m averaged temperatures of four additional 20-m towers existing in the same valleys (locations shown in Fig. 2). The general trend of warmer temperatures at lower sites is confirmed by all stations, provided the different heights above the ground of the measurements and the different valleys of each station are taken into account. The continuous vertical profile in Fig. 11 shows the corresponding temperature average of the 12 UTC (08 LT) operational radiosonde measurements available at the coast in Antofagasta (see Fig. 1). The well defined subsidence inversion very frequently observed at this location leaves a clear hallmark in the mean temperature profile, characterized by a conspicuous inversion in the 800-1500 m layer (Muñoz et al. 2011). The subsidence inversion, however, rarely exceeds 2000 m ASL leaving the altitude range occupied by the analyzed stations under the influence of the free tropo-

338 sphere above the inversion. Indeed, the mean lapse rate of the radiosonde profile is very close
339 to that of the surface stations. Moreover, while the nighttime near-surface measurements are
340 considerably colder than the radiosonde profile at the same altitude, the 80-m measurements
341 are much closer, suggesting that the surface inversions in these valleys tend to match the
342 free-tropospheric temperature (keeping in mind the differences in the times of the station
343 temperatures and the upper-air temperatures in Fig. 11). These observations suggest that
344 the turbulent mixing associated to the strong down-valley flows is capable of keeping the
345 near-surface temperatures in the valleys coupled to the free-tropospheric thermal profile.

346 *c. Slope-cooling forcing of down-valley flows*

347 In the previous sub-section we assessed the down-valley wind mechanism as a factor of
348 the observed nocturnal flows, by relating them to the regional wind and thermal structure.
349 While some support for the down-valley mechanism was found, this observational work would
350 not be complete without an estimate of the slope-wind mechanism as well. As pointed out
351 in the Introduction, real complex-terrain winds are prone to be affected by various factors
352 in different degrees, and considering the significant slope of these valley axes ($\sim 1.5\%$, see
353 Section 2), the cooling of the sloping surface air might play a role as well.

354 Figures 12a-b show the mean evolution of the down-valley component of the wind and
355 the near surface stability at the three sites. The existence of two distinct phases during
356 nighttime is suggested by these plots. A transition phase occurs between 18 LT and 02 LT,
357 in which the down-valley flow intensity increases steadily (Fig. 12a). Afterwards, between 02
358 and 07 LT, a quasi-stationary regime of flow and stability is observed. The 40m-02m stability
359 evolution during the transition phase shows an interesting behavior (Fig. 12b). It changes
360 sign at sunset (18 LT), just before the time when the down-valley flows begin to increase,
361 then it reaches a maximum in the middle of the transition and it finally adjusts to a lower
362 value in the quasi-stationary phase. Although the details of the transitions are different in
363 each station, the initial and final states and the shape of the transition are similar for the

364 three of them. The joint evolution of down-valley flow and stability during the transition
 365 suggests a mutual adjustment. Development of the surface inversion appears to accelerate
 366 the flow in the down-valley direction, with the flow at some time becoming so intense that its
 367 accompanying turbulent mixing reduces the near-surface stability, and a sort of equilibrium
 368 between stability and flow strength is finally established.

369 Some quantification of the slope mechanism is possible based on the available observa-
 370 tions. To do so, we compute the down-valley pressure gradient driven by a surface temper-
 371 ature inversion over a sloped surface. Let us model the temperature inversion as

$$372 \quad T(z) = T_\infty - \Delta_o \exp(-z/H), \quad (1)$$

373 where T_∞ is the upper-air temperature, Δ_o represents the intensity of the surface cooling,
 374 H is the height scale of the inversion, and z is the local vertical coordinate. The hydrostatic
 375 equation applied to this temperature structure results in

$$376 \quad \frac{dp}{dz} = -\frac{gp}{RT_\infty} \left(1 - \frac{\Delta_o}{T_\infty} \exp(-z/H) \right)^{-1}, \quad (2)$$

377 where g is the acceleration of gravity and R is the ideal gas constant for air. The three
 378 parameters T_∞ , Δ_o , and H are assumed uniform in space, so that horizontal pressure gra-
 379 dients arise solely due to the along-slope change of altitude, Z , for a constant value of z , as
 380 given by $Z = Z_T(x) + z$, where Z_T is the surface altitude and x is the distance coordinate in
 381 the up-valley direction. Assuming a small slope, $s = dZ_T/dx$, and small perturbations with
 382 respect to density ρ_o , and temperature T_o , the horizontal pressure gradient at the surface in
 383 the up-valley direction is approximately

$$384 \quad \frac{1}{\rho_o} \frac{\partial p}{\partial x} \sim sg \frac{\Delta_o}{T_o}. \quad (3)$$

385 If we estimate Δ_o based on the T40-T02 difference shown in Fig. 12b, then its linear time
 386 increase during the transition can be approximated by $\Delta_o \sim (5^\circ\text{C}/4\text{hr})t$, where t is time
 387 after sunset. Replacing this expression in (3) and integrating in time for a 4-hr transition,
 388 we obtain that the pressure gradient can support an increase of 18 m s^{-1} in the downslope

389 flow in the four hours (we have used $s \sim 0.015$, and $T_o \sim 300$ K). This value is quite large
 390 but of the same order as the ~ 12 m s⁻¹ observed increase in downslope flow during the
 391 transition (Fig. 12a), the overestimation suggesting that turbulent friction plays a role as
 392 well. Indeed, the role of turbulence is presumably important after the stability and flow
 393 have mutually adjusted. If we assume that in the quasi-stationary regime a balance exists
 394 between the turbulence and pressure gradient forces, then

$$395 \quad \frac{\partial \overline{u'w'}}{\partial z} \sim -sg \frac{\Delta_o}{T_o}, \quad (4)$$

396 where $\overline{u'w'}$ is the turbulent momentum flux. The left hand side of (4) can be estimated as

$$397 \quad \frac{\partial \overline{u'w'}}{\partial z} \sim -\frac{u_*^2}{H_n}, \quad (5)$$

398 where u_* is the surface friction velocity and H_n is a height scale at which $\overline{u'w'} \sim 0$. Combining
 399 (4) and (5), an estimate for H_n is

$$400 \quad H_n \sim u_*^2 \left(sg \frac{\Delta_o}{T_o} \right)^{-1}. \quad (6)$$

401 The friction velocity can in principle be related to the wind speed at 10 m AGL using
 402 atmospheric surface layer relationships. During the quasi-stationary regime, wind speed at
 403 10 m AGL is ~ 10 m s⁻¹ (Fig. 7) and assuming near-neutral conditions in the surface layer,
 404 we can estimate u_* by using a logarithmic wind profile and a surface roughness appropriate
 405 for a desert ($z_o \sim 10^{-3}$ m), obtaining $u_* \sim 0.4$ m s⁻¹. Figure 12b shows that in the
 406 equilibrium $\Delta_o \sim 5$ °C (we include also the temperature difference between 80 m and 40 m
 407 which in the equilibrium regime is significant), so that the scale height $H_n \sim 65$ m. This
 408 scale is higher but of the same order as the observed height of the jet nose (between 20 m
 409 and 60 m AGL according to Fig. 5), at which, by symmetry arguments, $\overline{u'w'}$ can indeed be
 410 argued to be close to zero, consistent also with the turbulence modeling of nocturnal jets by
 411 Conangla and Cuxart (2006) and Cuxart and Jiménez (2007). Two reasons can be thought
 412 of to explain the overestimation of H_n with respect to the observed nose jet heights: 1)
 413 the neutral assumption in the estimation of u_* produces probably an overestimation of u_* ,

414 considering that near-surface conditions are more stable than neutral; 2) if the actual H_n
415 is in the 20-40 m AGL range, then wind speed measurements at 10 m AGL fall probably
416 outside the atmospheric surface layer (conventionally defined as the layer where turbulent
417 fluxes change less than 10% from their surface values), so that the u_* estimate based on
418 the 10 m wind speed is again questionable. Neither of these two considerations, however,
419 invalidates the interpretation of (6) as an estimate for the nose height scale of an equilibrium
420 turbulent down-slope jet. Application of (6) to explain the observed case-to-case variability
421 of ZX produces the scatter plots in Figs. 12c-e. While the overestimation of nose height
422 values persists, a positive correlation between H_n and ZX is found at the three sites, being
423 largest at CAN, where about 70% of the observed ZX variance is explained by (6).

424 Summarizing this sub-section, we conclude that during the transient phase and during the
425 quasi-equilibrium regime of these down-valley flows, the joint evolution of the stability and
426 the flow intensity suggests that the down-slope pressure gradient due to the surface cooling
427 and the turbulent wind stress are both important terms in their momentum budgets.

428 4. Summary and Conclusions

429 To the authors's knowledge, this communication provides the first extended (~ 2 years,
430 0-80 m AGL, 3 sites) documentation of the surface wind flow over the Atacama Desert region
431 of Northern Chile. The most conspicuous feature of the observed wind regime is the existence
432 of a strong nocturnal drainage flow with hourly averaged down-valley speeds reaching up to
433 about 20 m s^{-1} during the cold season. The multi-level measurements reveal nose-shaped
434 wind profiles with maximum values occurring at about 20-60 m AGL. The frequency of
435 occurrence of persistent down-valley nocturnal jets with a well-defined nose-shape profile in
436 the 0-80 m layer shows little seasonal preference. Their intensity, however, shows a marked
437 seasonality with stronger flows occurring between May and October. Stronger down-valley
438 jets are associated to a more positive curvature of the temperature profile, probably due

439 to the effect of turbulent mixing. Standard deviation of wind speed suggests that strong
440 turbulence prevails in the full tower layer, despite gradient Richardson numbers being large
441 at the jet nose and above it.

442 Turbulent mixing also appears to be responsible for the warmer temperatures observed
443 at the lower altitude sites down the valleys. In fact, the mean altitudinal lapse rate of
444 near-surface temperature matches closely the free-tropospheric lapse rate measured at an
445 aerological station located at the coast. Regressions between these upper-air measurements
446 and the tower observations show that there exists a moderate control of valley nocturnal
447 flows by large scale factors. The highest correlations found are those between the down-
448 valley flow intensity and the temperature difference between the free troposphere and the
449 valley, providing some support for the down-valley wind mechanism. Nevertheless, estimates
450 of terms in the momentum budget for the development and the quasi-stationary phases of
451 the down-valley flows suggest that the pressure gradient force due to the near-surface cooling
452 along the sloping valley axes also plays an important role in these drainage flows. A scale for
453 the height of the wind speed maximum in equilibrium turbulent down-slope jets is proposed
454 based on the surface friction velocity and the surface inversion intensity. At site CAN
455 this scale explains about 70% of the case-to-case variance of the observed jet nose height,
456 although a significant overestimation is found, probably due to an overestimation of the
457 friction velocity.

458 The flow system reported here has immediate importance in the characterization of the
459 wind energy potential of the Atacama Desert. As mentioned in the Introduction, these
460 measurements were undertaken by the Chilean government in order to explore the possibility
461 of wind farm development in the Atacama region, where the current demand for energy is
462 particularly high. Our results show that, while there is possibly a good wind resource at all
463 measurement sites, the shape and the relationship with stability of the vertical profiles of
464 wind speed sustained by these flows can be quite different to what is commonly found in wind
465 farm sites over flat terrain (e.g. van den Berg 2008), a fact that must be properly taken into

466 account in the evaluation and design of future wind energy projects in this area. Beyond this
467 applied aspect, further studies should better define the spatial structure and organization of
468 the nocturnal circulation, in order to assess how representative these measurements are of
469 the general low-level wind-regime of the Atacama Desert. Additional investigations should
470 better characterize the evening and morning transitions of these winds and their turbulent
471 structure, as well as elucidate their possible interaction with the coastal boundary layer
472 and subsidence inversion existing to the west and quantify their potential role in the zonal
473 transport of natural and anthropogenic aerosols from the continent to the SEP stratocumulus
474 layer existing offshore. Studies on the numerical modeling and the synoptic forcing of these
475 flows are currently underway in order to address some of these issues.

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TABLE 1. Measurement heights and main NRG Systems Inc. sensors available in 80 m towers.

variable	heights (m AGL)	sensors	main characteristics
wind speed	10, 20, 40, 60, 80	NRG#40C	3-cup anemometer; 0.78 m s ⁻¹ threshold; 3.0 m distance constant.
wind direction	10, 40, 80	NRG#200P	continuous rotation potentiometer; 1.0 m s ⁻¹ threshold.
temperature	2, 40, 80	NRG#110S	integrated circuit temperature sensor; 10 min thermal time constant.

TABLE 2. Correlation coefficients between tower measurements and Antofagasta radiosonde data in the 2000 - 3000 m ASL layer (May-October).

tower variable	Antofagasta-related variable	CAN	CAO	SGE
DVWS40	ESE-SE Wind	0.41	0.53	0.49
DVWS40	TANTO	0.54	0.54	0.62
DVWS40	TANTO-TZNIGHT	0.60	0.71	0.81

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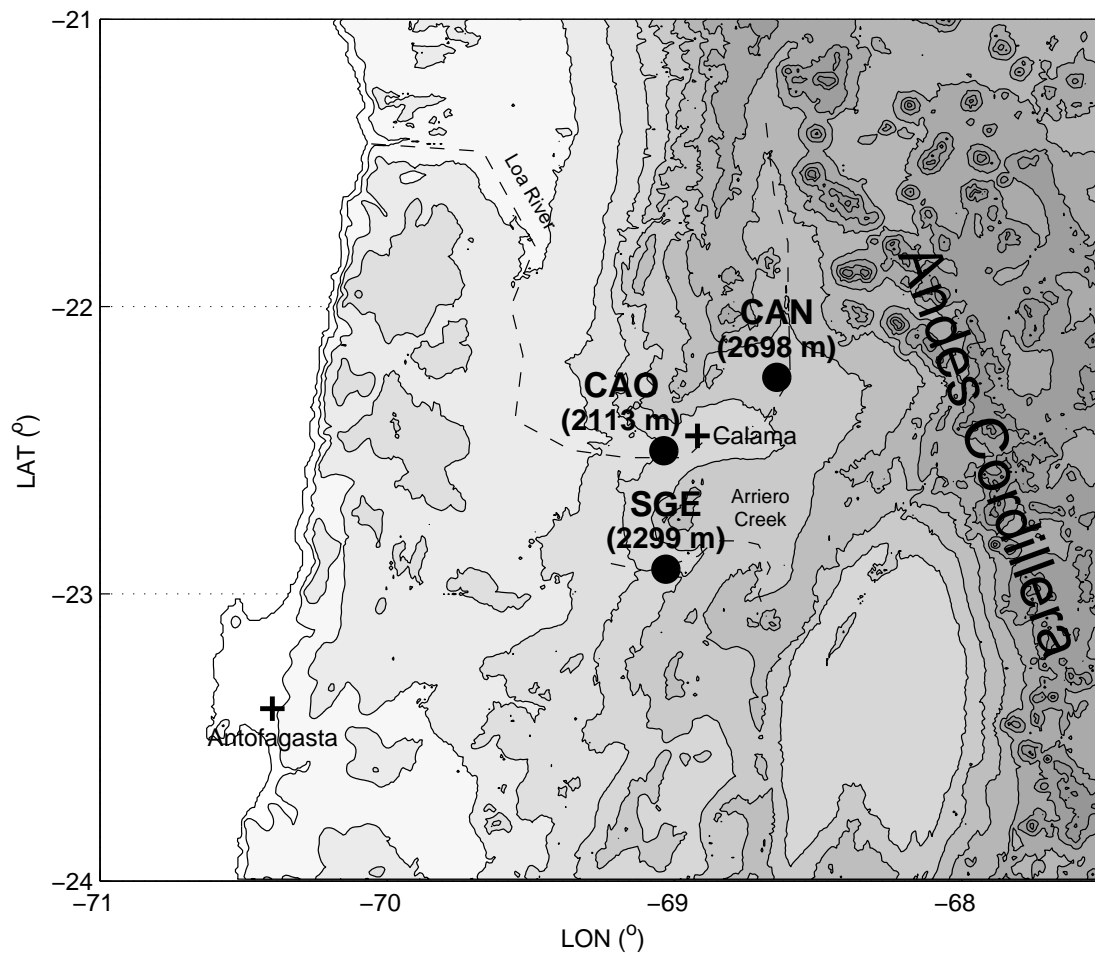


FIG. 1. Topography of the study region and location of 80-m towers (altitude of the stations shown in parenthesis). Terrain heights are contoured from 0 to 6000 m ASL every 500 m. Pluses (+) indicate radiosonde station (Antofagasta) and Calama city. Dashed lines mark Loa River and Arriero Creek.

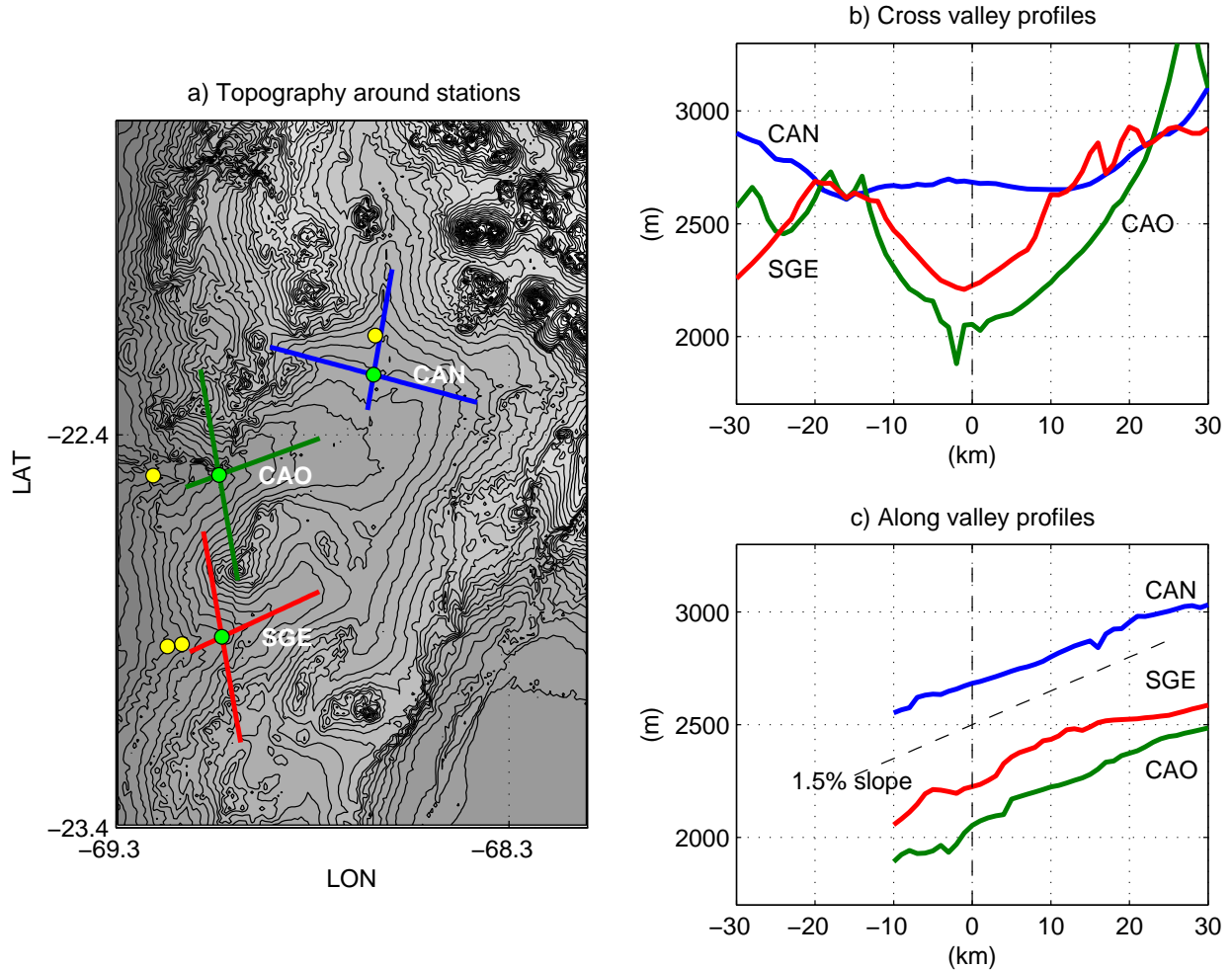


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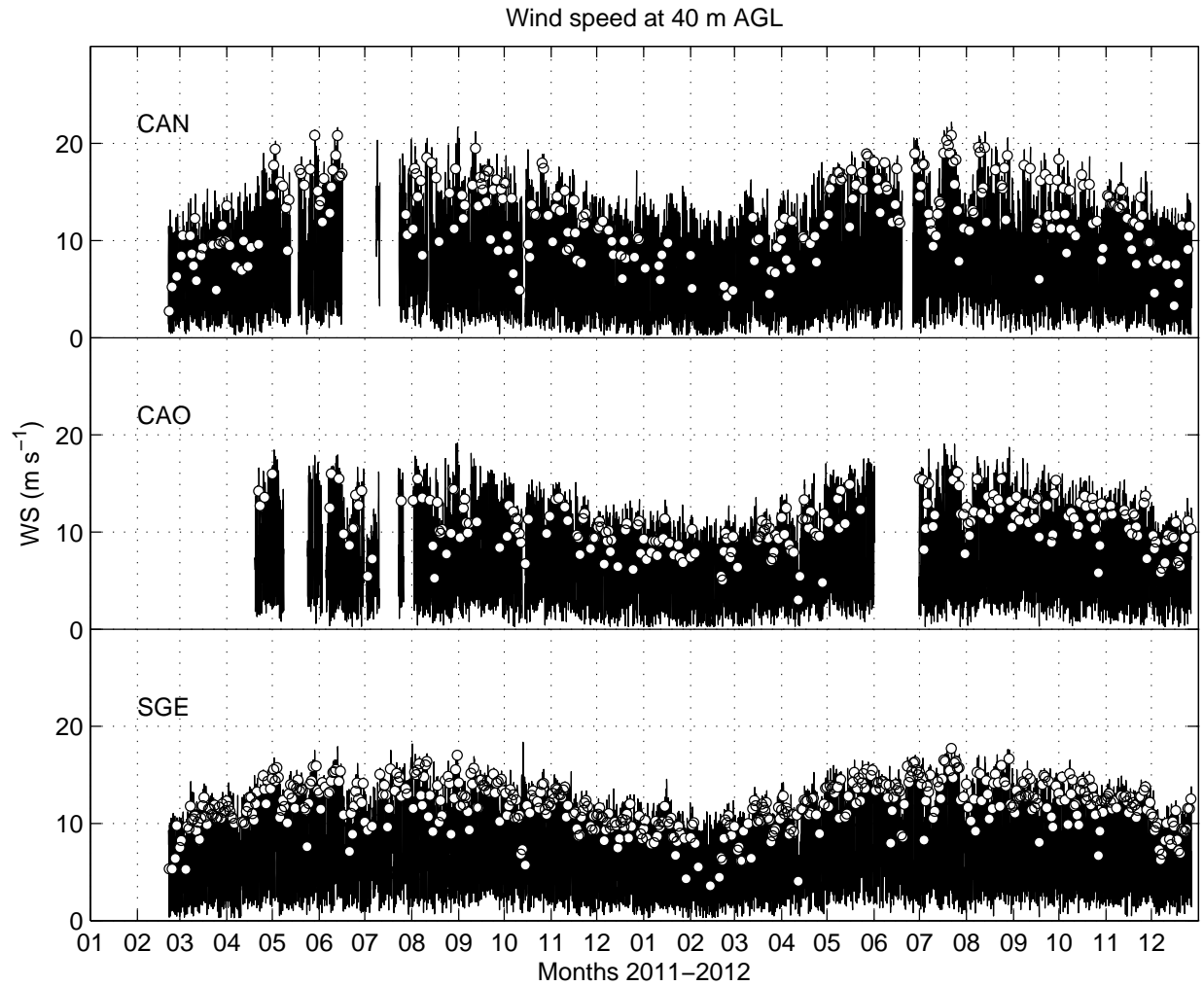


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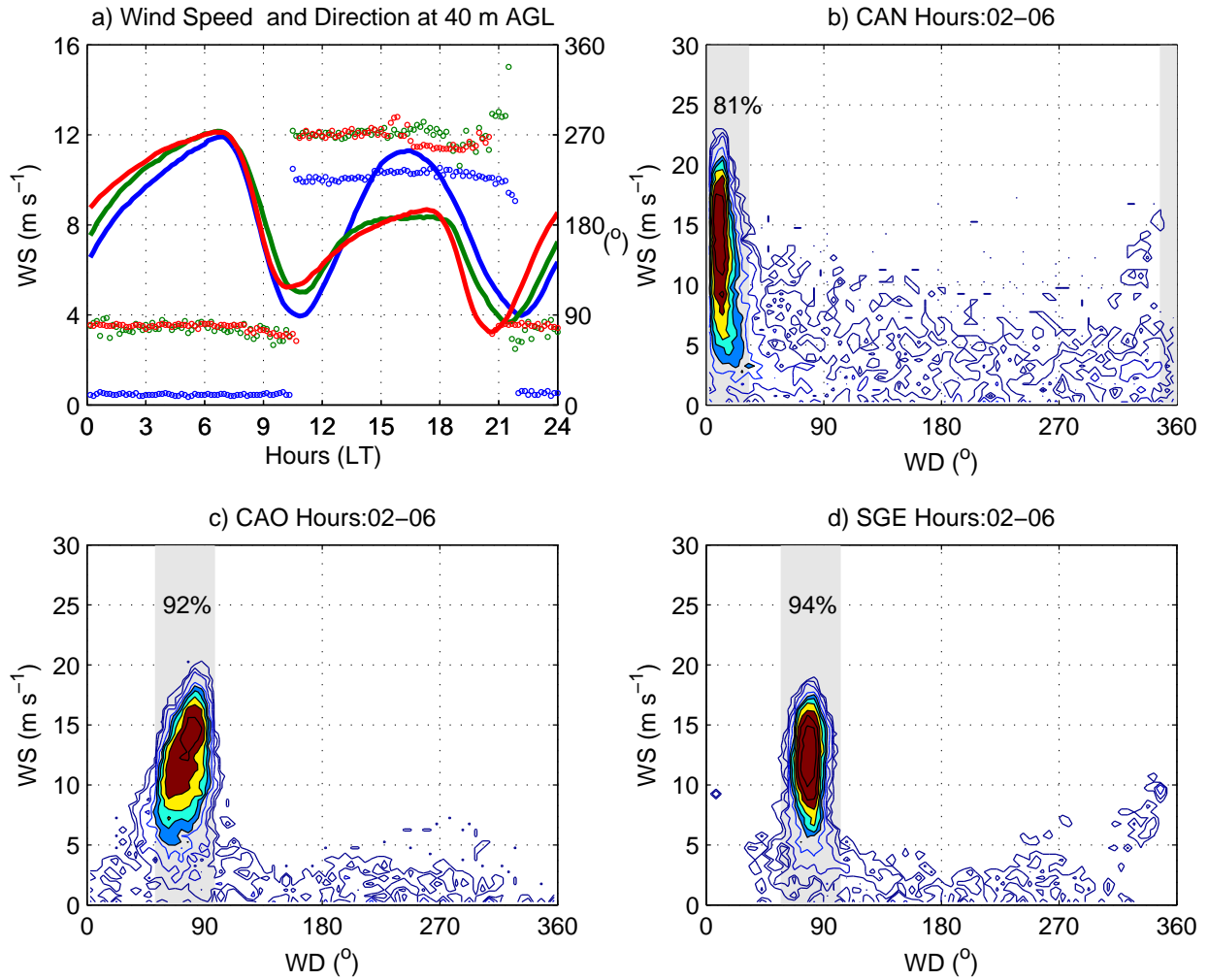


FIG. 4. a) Diurnal cycles of mean wind speed (lines, left scale) and most frequent wind direction (circles, right scale) measured at 40 m AGL for the full period of record at CAN (blue), CAO (green) and SGE (red). b) CAN wind direction-wind speed joint frequency distribution at 40 m AGL for 10-minute averages in the 02-06 LT period. Frequency contours are logarithmically spaced in order to display infrequent occurrences (line contours) and more common occurrences (colored contours). c) as b) but for CAO. d) as b) but for SGE. Gray-shaded regions in b)-d) mark the down-valley direction sectors defined for each station. The percentage of 02-06 LT wind direction observations falling in this sector is annotated for each site.

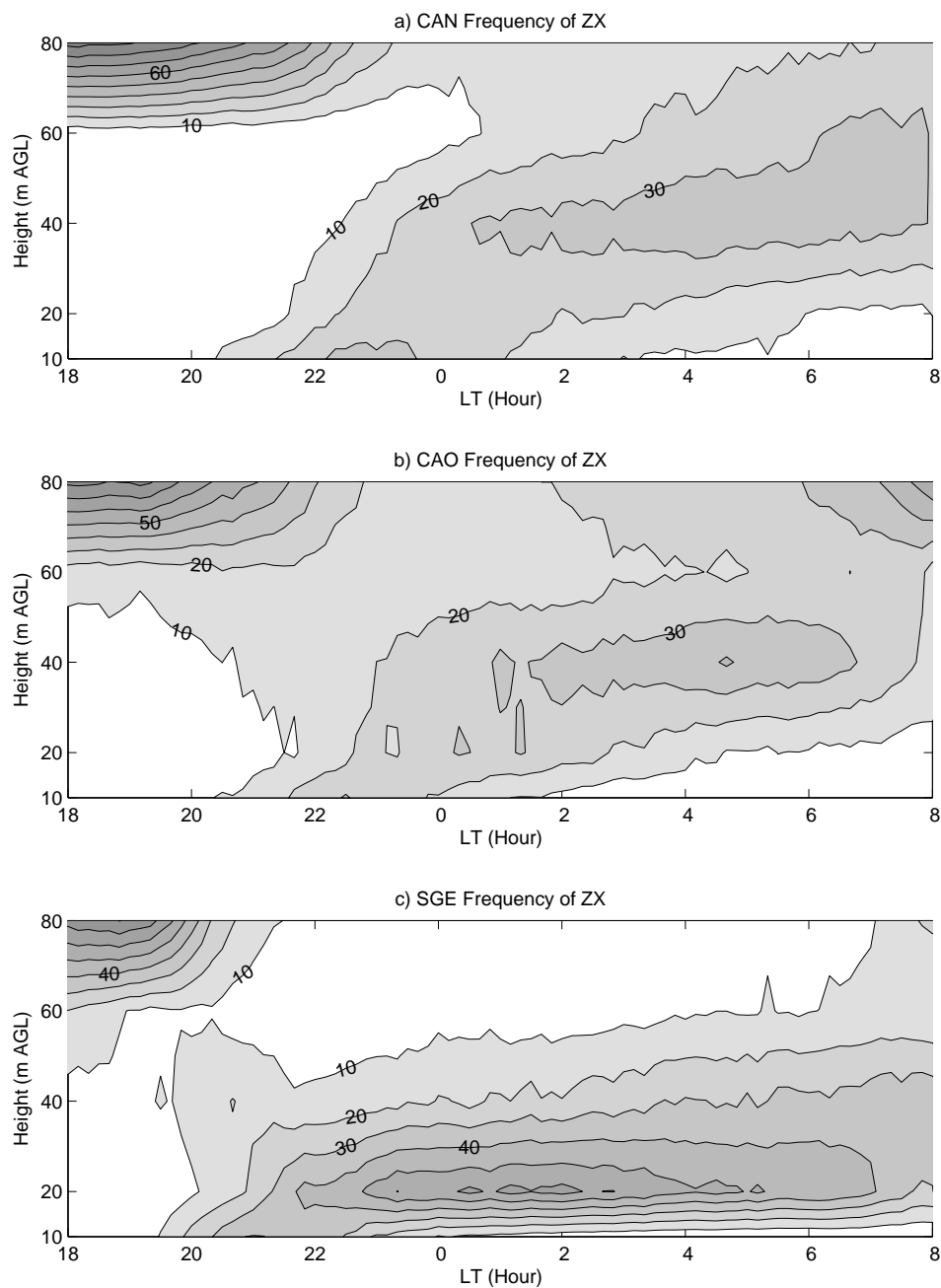


FIG. 5. Time evolution of the frequency distribution of the height of maximum wind speed (ZX) measured at a) CAN, b) CAO, c) SGE. Frequency contours drawn from 10% to 90% every 10%. Only the period from 18 to 08 LT is shown with 10-minute time resolution. Wind speed data is available only at measuring heights of 10 m, 20 m, 40 m, 60 m, and 80 m AGL. The full period of record at each station has been considered.

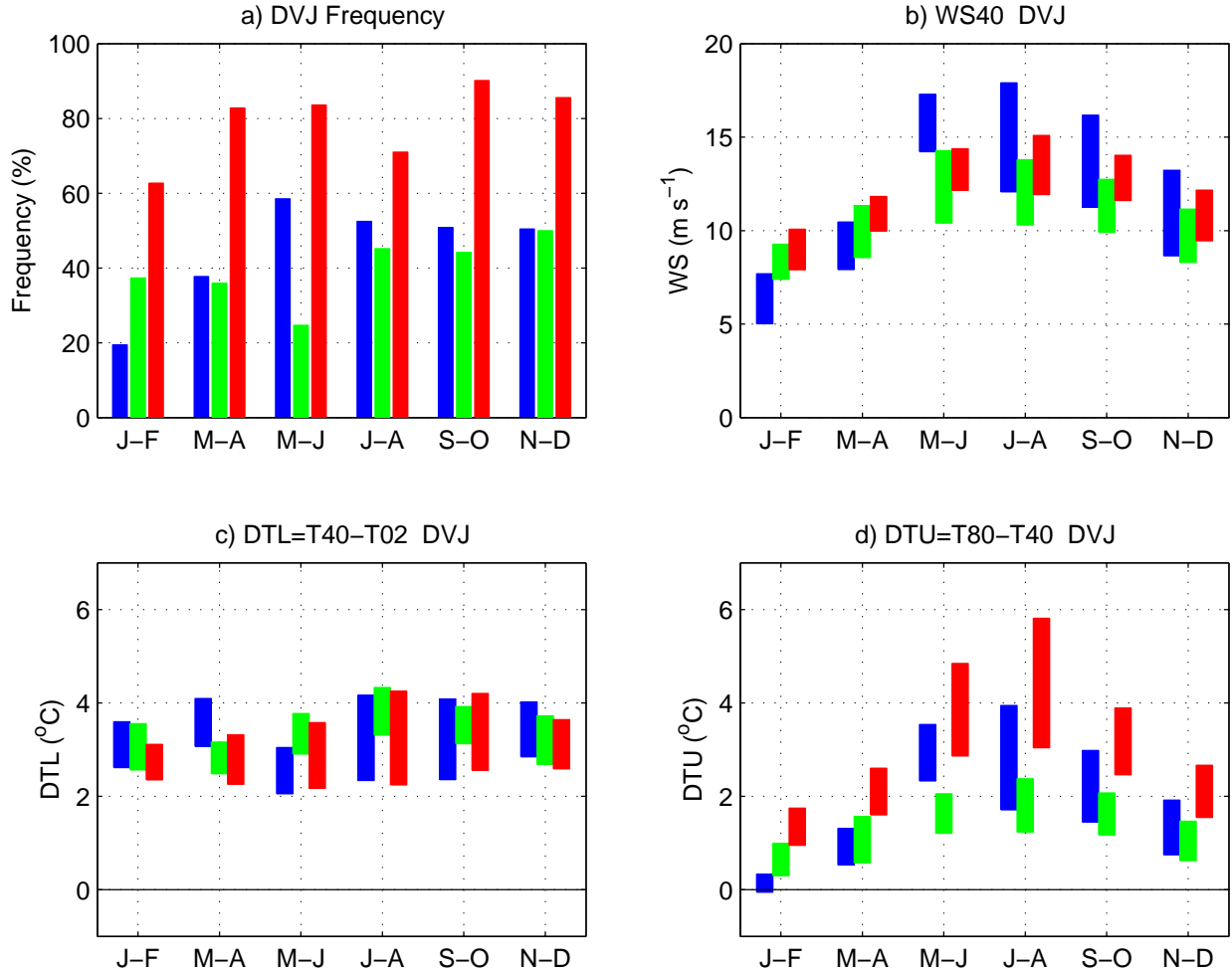


FIG. 6. a) Bi-monthly occurrence frequency of down-valley jets (DVJs) as defined in the text. b) Bi-monthly variation of inter-quartile ranges of DVJ intensities at 40 m AGL averaged in the 02-06 LT period. c) As b) but for 40 m-02 m temperature difference during DVJ events. d) As c) but for 80 m-40 m temperature difference. In all panels colors blue, green and red refer to stations CAN, CAO and SGE, respectively.

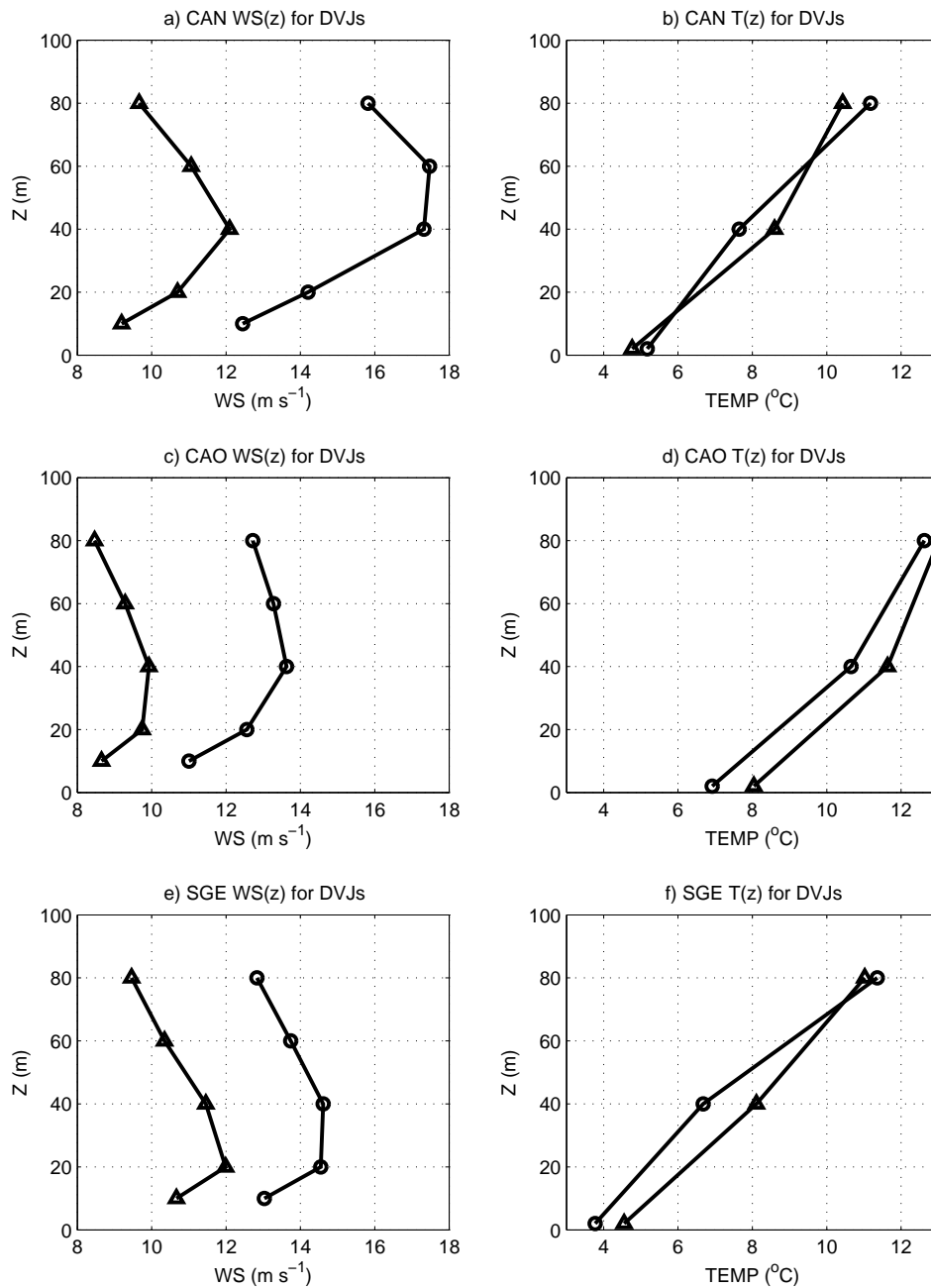


FIG. 7. Vertical profiles of 02-06 LT-averaged wind speed (left panels) and temperature (right panels) averaged over strong (circles) and weak (triangles) DVJs for the months of May to October at stations CAN (upper), CAO (middle) and SGE (bottom). Threshold between weak and strong DVJs is based on the median of WS40 at each site (15.4 m s^{-1} , 11.8 m s^{-1} , and 13.3 m s^{-1} for CAN, CAO, and SGE, respectively).

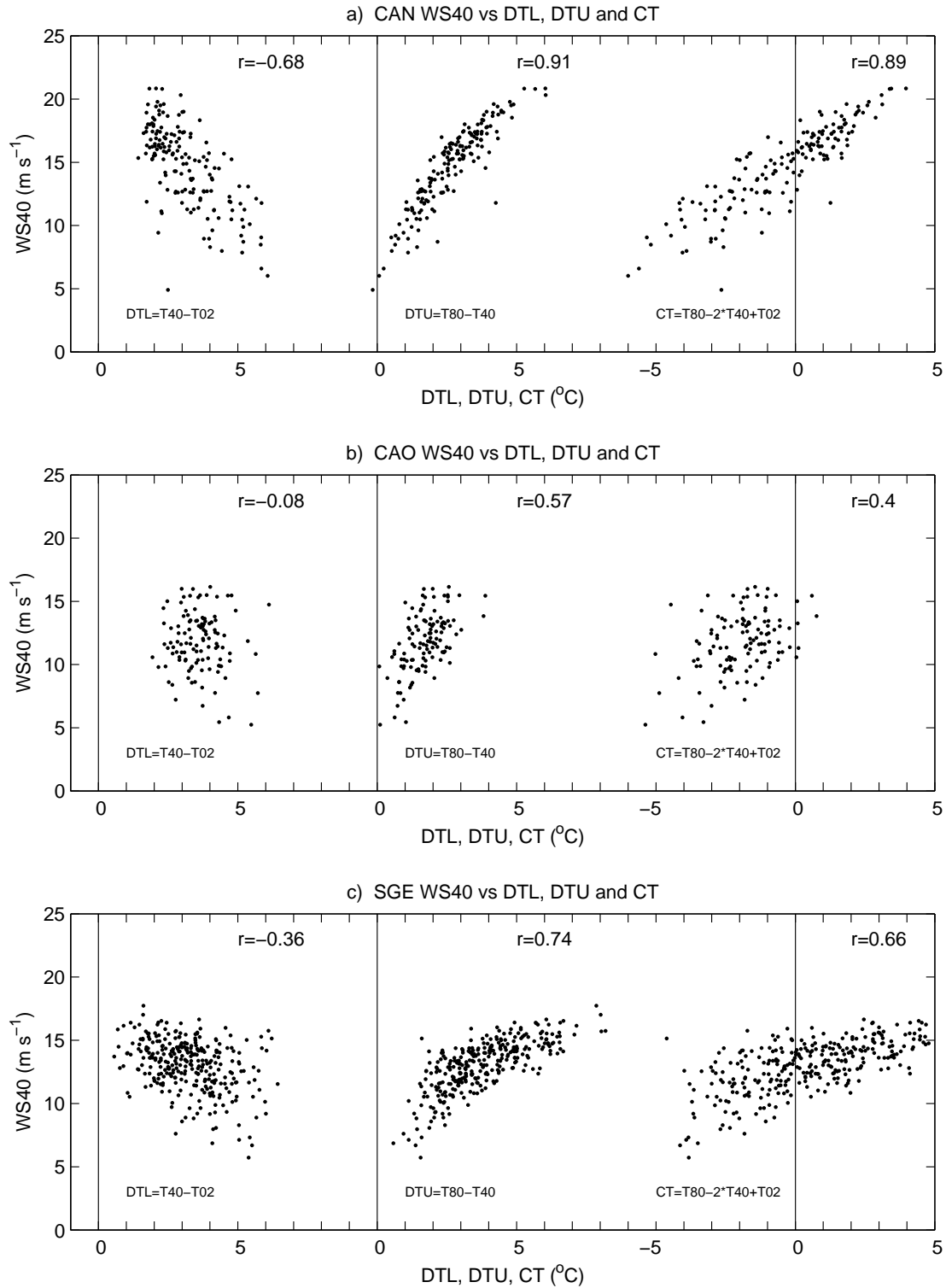


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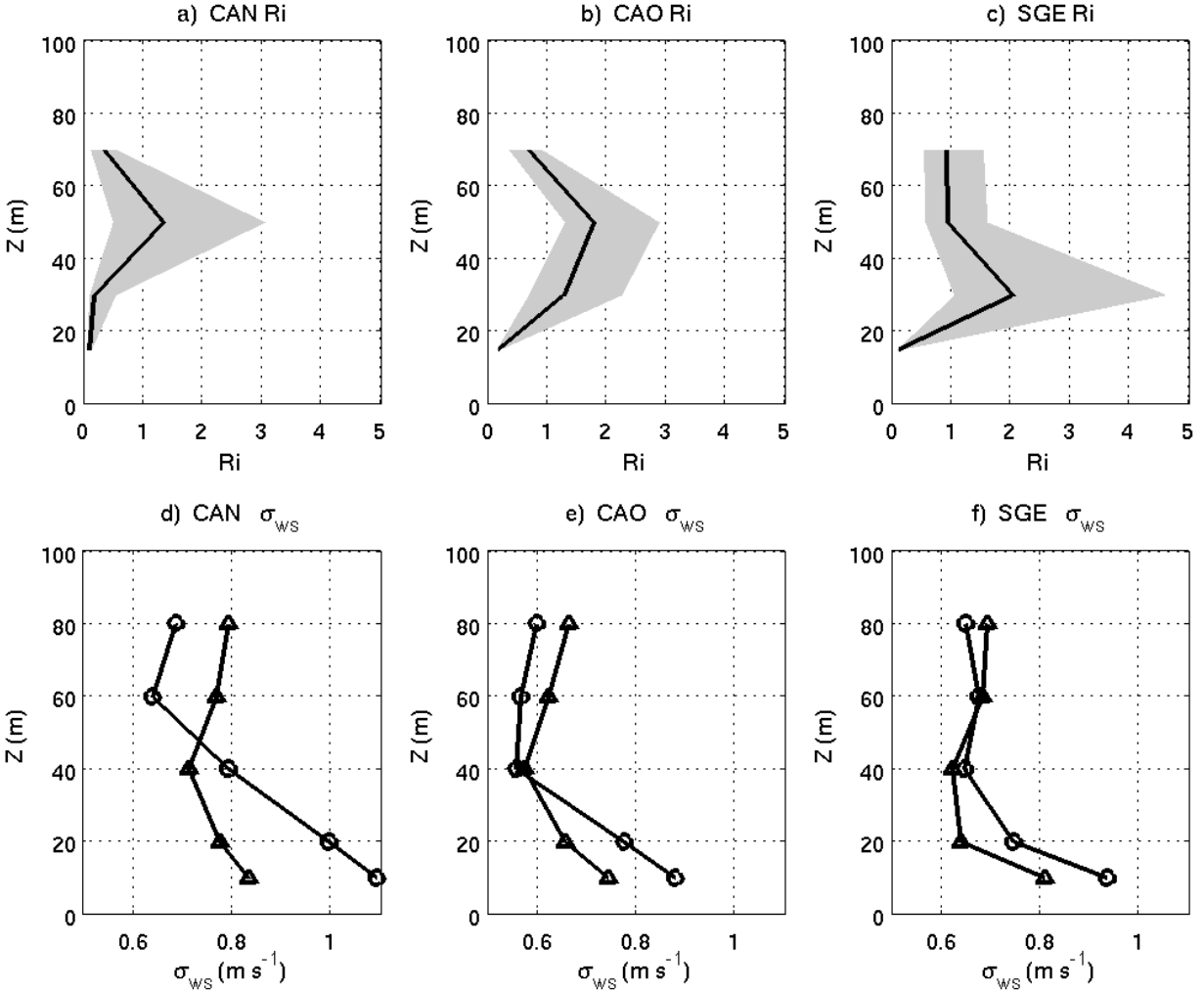


FIG. 9. Upper panels: Vertical profiles of Richardson numbers during DVJ events (02-06 LT averages). Shading marks the inter-quartile range and continuous line marks the median values. Lower panels: Vertical profiles of wind speed standard deviation averaged over strong (circles) and weak (triangles) DVJs. Left, middle and right panels correspond to stations CAN, CAO and SGE, respectively. Analysis period includes the months from May to October.

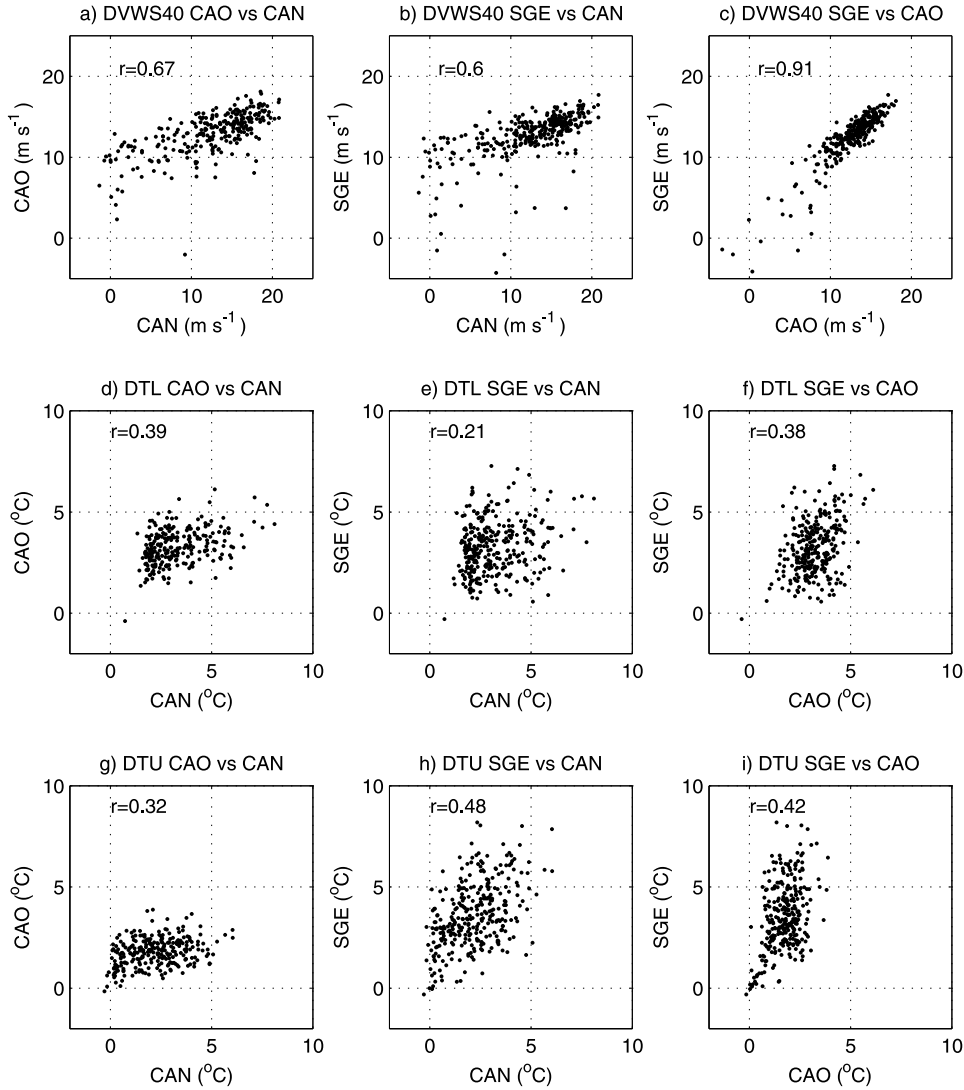


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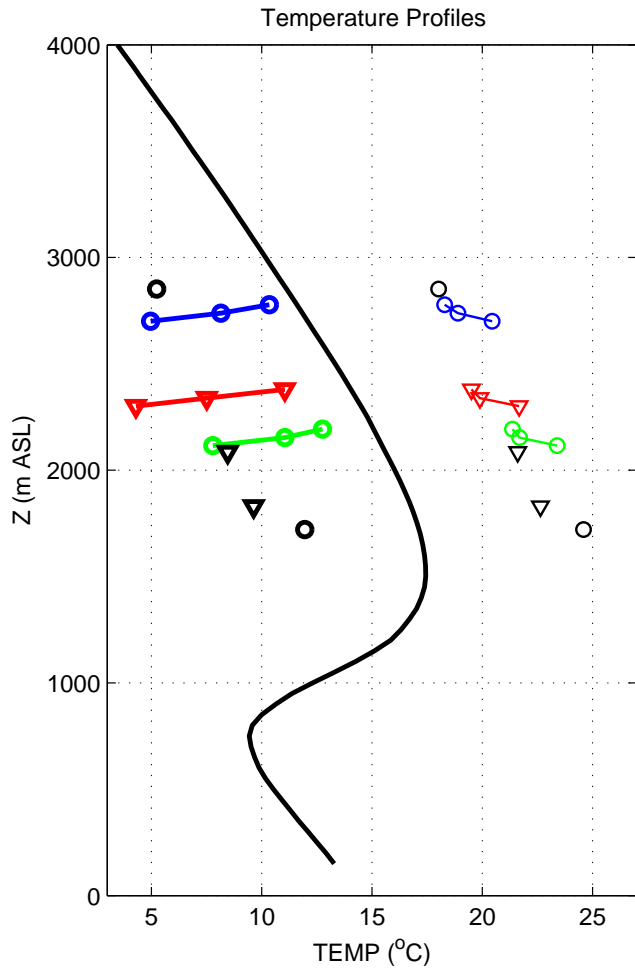


FIG. 11. Temperatures averaged over the months from May to October plotted against the altitude above sea level of the sensor. Black continuous line: 12 UTC (08 LT) Antofagasta radiosondes. Colored profiles for 02-06 LT (bold) and 14-18 LT (fine) averages at CAN (blue), CAO (green) and SGE (red). Black symbols: 02-06 LT (bold) and 14-18 LT (fine) averages of 5-m temperatures measured at 20-m meteorological masts (locations shown in Fig. 2). Triangles and circles indicate stations located along Arriero Creek and Loa River valleys, respectively.

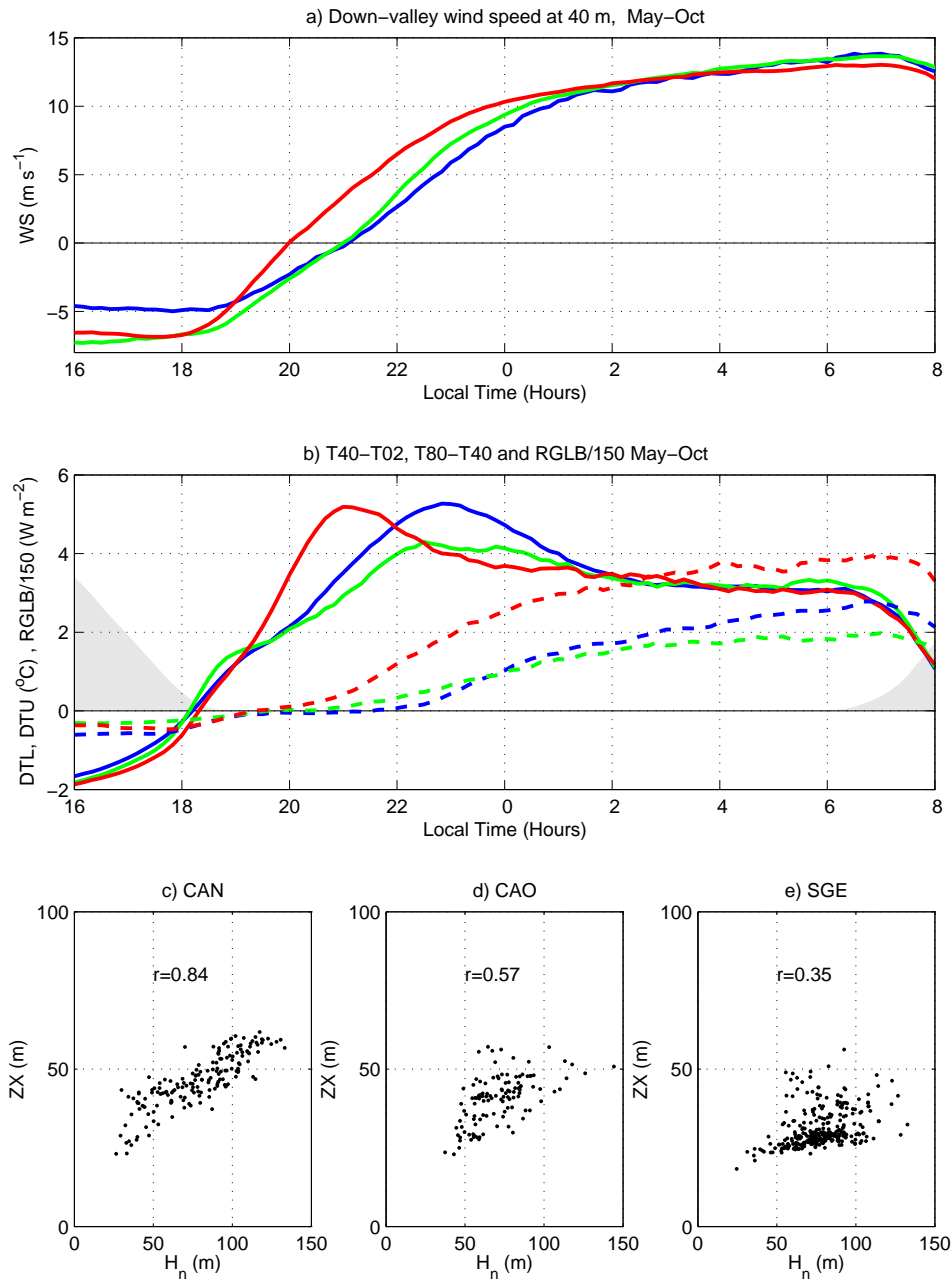


FIG. 12. a) May-October average of down-valley projected wind speed at 40 m AGL (continuous) for stations CAN (blue), CAO (green) and SGE (red) for period between 16 and 08 LT. b) As a), but for 40 m - 02 m temperature difference (continuous), 80 m - 40 m temperature difference (dashed), and mean surface solar radiation scaled by 150 (shaded). c) Scatter plot between observed jet nose heights (ZX) and estimates based on equation (6) for station CAN. d) as c) but for CAO. e) as c) but for SGE. In order to have a continuous estimate of ZX, a second-order polynomial was fitted to the three data levels around the observed wind speed maximum.