

Mid- to Late Cambrian docking of the Río de la Plata craton to southwestern Gondwana: age constraints from U–Pb SHRIMP detrital zircon ages from Sierras de Ambato and Velasco (Sierras Pampeanas, Argentina)

SEBASTIÁN O. VERDECCHIA^{1*}, CESAR CASQUET², EDGARDO G. BALDO¹, ROBERT J. PANKHURST³, CARLOS W. RAPELA⁴, MARK FANNING⁵ & CARMEN GALINDO²

¹CICTERRA—Consejo Nacional de Investigaciones Científicas y Técnicas—Universidad Nacional de Córdoba (CONICET—UNC), Av. Vélez Sarsfield 1611, CP X5016CGA, Córdoba, Argentina

²Dpto. de Petrología y Geoquímica, Fac. Ciencias Geológicas, Inst. de Geología Económica (CSIC—Universidad Complutense), 28040 Madrid, Spain

³British Geological Survey, Keyworth, Nottingham NG12 5GG, UK

⁴Centro de Investigaciones Geológicas, Universidad Nacional de la Plata—CONICET, 1900 La Plata, Argentina

⁵Research School of Earth Sciences, The Australian National University, ACT0200 Canberra, Australia

*Corresponding author (e-mail: sverdecchia@gmail.com)

Abstract: The Early Palaeozoic stratigraphy and tectonic history of the Eastern Sierras Pampeanas of central Argentina are complicated by metamorphism and deformation resulting from the Pampean (545–510 Ma) and Famatinian (490–440 Ma) orogenies. We report U–Pb sensitive high-resolution ion microprobe dating of detrital zircons in two metasedimentary successions exposed at Quebrada de La Cébila (*c.* 28°45'S, 66°25'W): the Ambato and the La Cébila metamorphic complexes. The Ambato zircons record age peaks corresponding to Pampean (530 ± 10 Ma), Brasiliano (*c.* 570 and *c.* 640 Ma), Grenville (*c.* 950 to *c.* 1025 Ma) and minor Neoarchaean ages. Similar peaks are also apparent in the La Cébila sample but it additionally contains Palaeoproterozoic zircons (*c.* 2.1 Ga) corresponding to the age of the Río de la Plata craton, from which they are considered to have been sourced. Our interpretation is that the protolith of the Ambato complex was deposited prior to juxtaposition with the craton and is older than the Early Ordovician La Cébila metamorphic complex. We infer that the craton reached its current relative position in the Mid- to Late Cambrian, after the main Pampean tectonothermal event (530–520 Ma) and before deposition of the La Cébila protolith and the Achavil Formation (Sierra de Famatina), which contain comparable detrital zircon populations.

The Palaeoproterozoic Río de la Plata craton of central–eastern Argentina and southern Uruguay has an important role in the tectonic framework of southwestern Gondwana. Historically, this craton was considered the upper plate in collisional models for the Neoproterozoic to Early Cambrian Pampean orogeny (545–510 Ma) during the amalgamation of Gondwana (Escayola *et al.* 2007; Schwartz *et al.* 2008; Ramos *et al.* 2010). However, according to Schwartz & Gromet (2004) and Rapela *et al.* (2007) the pre-Pampean sedimentary Puncoviscana Formation (largely Late Neoproterozoic to Early Cambrian turbidites; see Zimmermann 2005), does not contain the 2.05–2.25 Ga detrital zircons expected in the case of orthogonal collision with the craton. Therefore the present position of the Río de la Plata craton had to be attained during or after the Pampean orogeny. From this evidence a new geotectonic model for the Pampean orogeny was developed by Rapela *et al.* (2007), involving significant right-lateral displacements of continental masses during oblique subduction that preceded collision.

In this paper, we present new U–Pb sensitive high-resolution ion microprobe (SHRIMP) detrital zircon ages from high-grade metasedimentary successions from Quebrada de La Cébila (southern Sierra de Ambato and NE Sierra de Velasco): the Ambato and La Cébila metamorphic complexes (Figs 1 and 2). Provenance patterns, combined with existing geochronological data relevant to Neoproterozoic to Early Palaeozoic metasedi-

mentary successions elsewhere in the Sierras Pampeanas (Rapela *et al.* 2007; Collo *et al.* 2009), constrain the time at which the Río de la Plata craton first became available as a source of sediments for late to post-Pampean basins. This is taken as dating docking of the craton to SW Gondwana.

Geological setting

Orcrops and samples from drill-holes of the Río de la Plata craton suggest that it consists of a mosaic of Palaeoproterozoic igneous and metamorphic terranes of 2260–2020 Ma (e.g. Tandilia belt, Pando belt, Piedra Alta terrane) (Fig. 1; Ramos 1996; Rapela *et al.* 2007; Oyhantçabal *et al.* 2010, and references therein). Its present position was probably reached through large-scale dextral strike-slip movement (present coordinates will be considered throughout the paper) relative to the Puncoviscana Formation that forms the bulk of the sedimentary sequence involved in the Pampean orogen. The latter formation, which overlies an unexposed basement, is thought to have originated on the margin of the Kalahari craton and transferred laterally during oblique subduction of the ephemeral Clymene Ocean (Trindade *et al.* 2006; Rapela *et al.* 2007). In this interpretation, the final stage is represented by the oblique collision in Cambrian times of a large, probably allochthonous, Mesoproterozoic to Palaeoproterozoic terrane in the west that embraced the Western Sierras

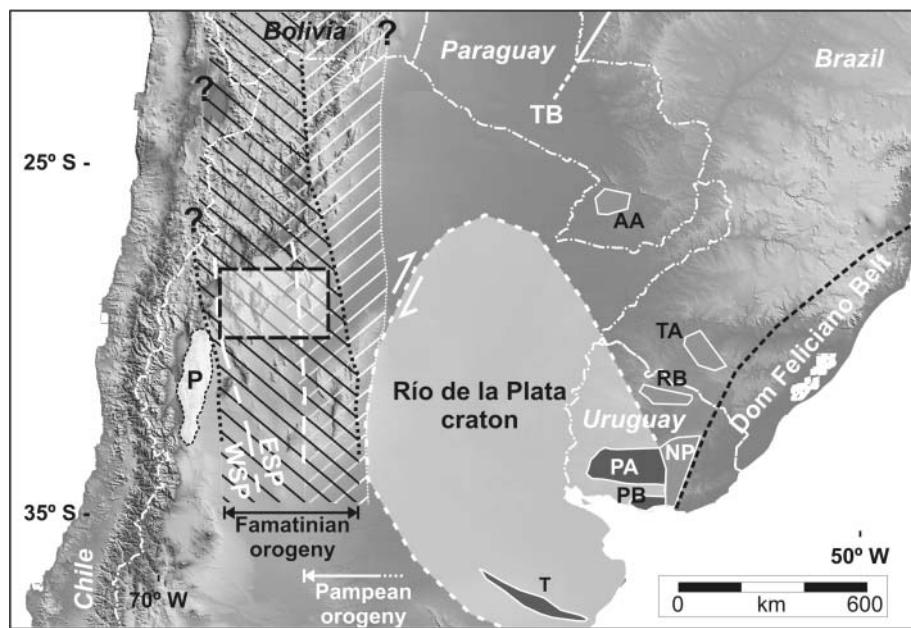


Fig. 1. Digital elevation model (DEM, 90° SRTM type) of central South America showing the Neoproterozoic to Early Palaeozoic tectonic framework and inferred limit of Río de la Plata craton (modified from Cordani *et al.* 2003; Rapela *et al.* 2007; Oyhantçabal *et al.* 2010). TB, Transbrasiliano lineament; T, Tandilia belt; PB, Pando belt; PA, Piedra Alta terrane; NP, Nico Pérez terrane; RB, Rivera block; TA, Tacuarembó block; AA, Asunción arch; ESP, Eastern Sierras Pampeanas; WSP, Western Sierras Pampeanas; P, Precordillera terrane.

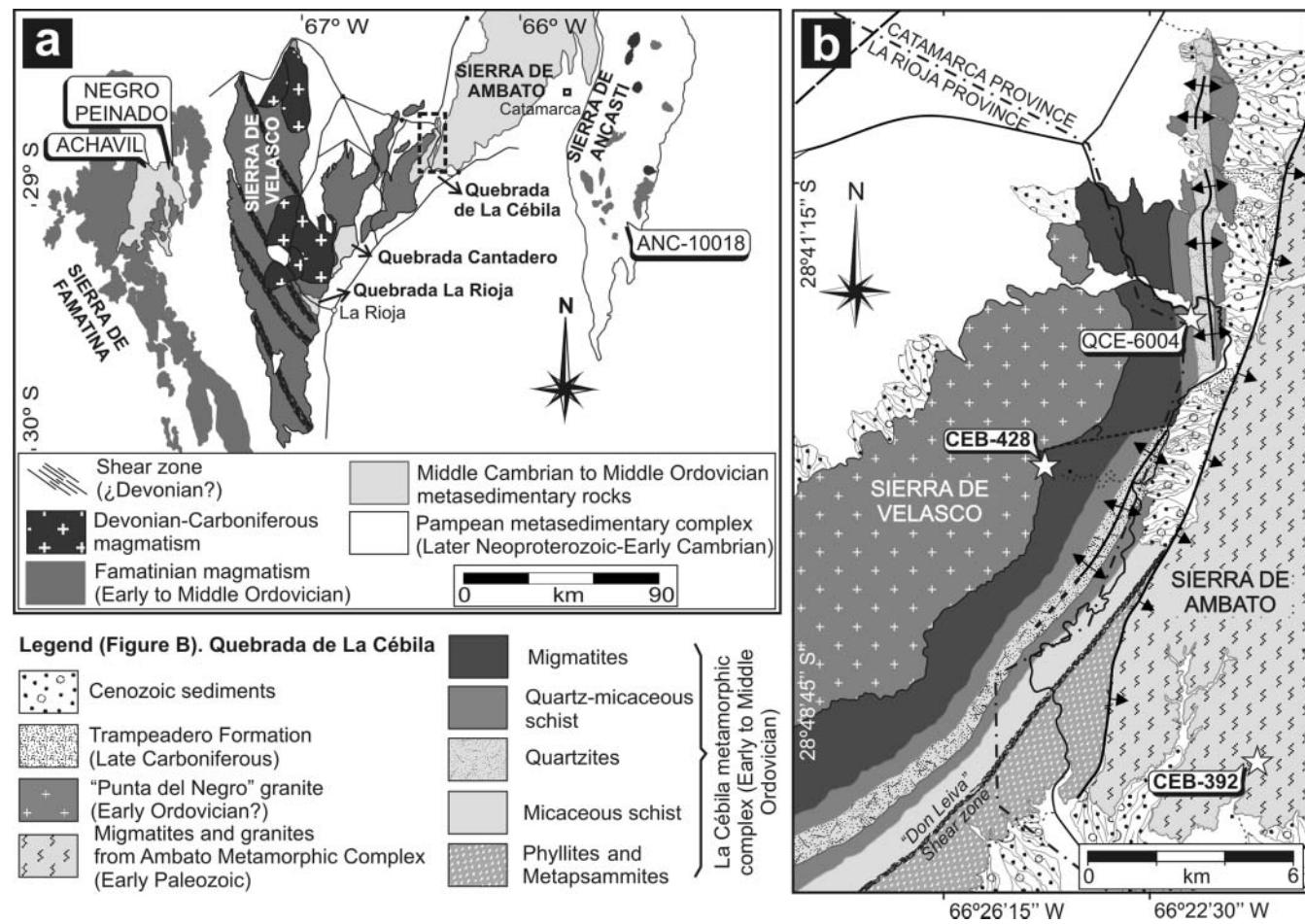


Fig. 2. (a) Schematic geological map of central–western Argentina (after Astini & Dávila 2004; Dahlquist *et al.* 2008; Grosse *et al.* 2008). The main metasedimentary outcrops of the La Cébila metamorphic complex are marked (Quebradas of La Cébila, Cantadero and La Rioja). (b) Geological map of Quebrada de La Cébila. New and previously published sample localities are shown.

Pampeanas (Sierra de Pie de Palo and Sierra de Maz), the Arequipa block (southern Peru; Ramos 2008, and references therein; Casquet *et al.* 2010) and Amazonia among other continental blocks (Rapela *et al.* 2007; Casquet *et al.* 2009). This collision resulted in the Pampean orogenic belt (545–510 Ma, Rapela *et al.* 1998; Schwartz *et al.* 2008) in the Eastern Sierras Pampeanas, juxtaposed with the Río de la Plata craton across a major fault (Fig. 1).

Alternative models for the Pampean orogeny involve either Late Neoproterozoic to Early Cambrian orthogonal collision or ridge subduction against the Río de la Plata craton (Ramos & Vujovich 1993; Escayola *et al.* 2007; Schwartz *et al.* 2008; Ramos *et al.* 2010). Folding and the development of foliation took place in the Early Cambrian and were accompanied by low- to high-grade regional metamorphism under low to medium pressure between 530 and 520 Ma. Intrusion of calc-alkaline I-type plutons and S-type granites (magmatic arc) started at c. 545 Ma (Lira *et al.* 1997; Rapela *et al.* 1998, 2002; Schwartz *et al.* 2008; Martino *et al.* 2009).

The Sierras Pampeanas are blocks of pre-Andean basement tilted during Late Cenozoic flat-slab subduction of the Nazca plate beneath the Central Andean continental margin between 27° and 33°30'S (Jordan & Allmendinger 1986; Ramos *et al.* 2002). Two successive orogenies have long been recognized in the Eastern Sierras Pampeanas: the Pampean orogeny referred to above and the accretionary Famatinian orogeny (Aceñolaza & Toselli 1977; Pankhurst *et al.* 1998; Rapela *et al.* 1998; Dahlquist *et al.* 2008; Fig. 1). The Famatinian belt lies to the west of the Pampean belt and developed mainly between the Late Cambrian and the Early Silurian (490–440 Ma) (e.g. Pankhurst *et al.* 1998, 2000; Astini & Dávila 2004). This orogeny partially reworked the Pampean foreland to the east, and extended well into the Western Sierras Pampeanas on the west (e.g. Rapela *et al.* 1998; Pankhurst *et al.* 2000; Baldo *et al.* 2006; Casquet *et al.* 2008). The Famatinian belt is characterized by Late Cambrian to Early–Middle Ordovician marine and volcanioclastic successions, Early to Mid-Ordovician I- and S-type intrusions (magmatic arcs), minor tonalite–trondjemite–granodiorite suites in the foreland and low- to high-grade, low- to intermediate-pressure metamorphism coeval with foliation development, folding and thrusting (e.g. Pankhurst *et al.* 1998, 2000; Casquet *et al.* 2001, 2008; Rapela *et al.* 2001; Astini 2003, and references therein; Astini & Dávila 2004; Büttner *et al.* 2005; Verdecchia *et al.* 2007; Dahlquist *et al.* 2008; Otamendi *et al.* 2008, 2009; Collo *et al.* 2009). The Western Sierras Pampeanas close to the Andes consist of a Proterozoic basement of Grenville age (c. 1.0–1.3 Ga) that was pervasively reworked by the Famatinian orogeny (Pankhurst & Rapela 1998; Casquet *et al.* 2001, 2008; Varela *et al.* 2004; Vujovich *et al.* 2004; Rapela *et al.* 2010). Evidence for Pampean-age tectonothermal activity in the Western Sierras Pampeanas is provided by U–Pb data, obtained by both thermal ionization mass spectrometry (TIMS) and SHRIMP, and Ar–Ar determinations (e.g. Lucassen & Becchio 2003; Mulcahy *et al.* 2007; Casquet *et al.* 2008).

This work is focused on the southern tip of the Sierra de Ambato and eastern tip of Sierra de Velasco in the Eastern Sierras Pampeanas (Fig. 2), which underwent Famatinian deformation and metamorphism. Two metasedimentary successions are recognized, yielding contrasting detrital zircon ages. The Ambato metamorphic complex (Fig. 2b) mainly consists of high-grade metasedimentary rocks (migmatites and gneisses) and discordant granitic and pegmatitic bodies (Caminos 1979). Larrovere (2009) obtained an Early to Mid-Ordovician metamorphic age on monazite from one migmatite from the central–

northern part of the sierra. However, the ages of igneous rocks and sedimentary protoliths in this region are still unknown. At the southern tip of the Sierra de Ambato, the metamorphic complex overlies the low-grade successions of the La Cébila metamorphic complex across a west-directed Cenozoic reverse fault (Fig. 2b).

The La Cébila metamorphic complex consists of a low- to high-grade metasedimentary succession, peraluminous granites and pegmatitic bodies (Espízúa & Caminos 1979; Verdecchia 2009; Fig. 2b) that crop out discontinuously along the eastern edge of the Sierra de Velasco with the main outcrops along the Quebrada de La Cébila (Fig. 2b). The Sierra de Velasco is a large igneous massif consisting of Ordovician peraluminous to metaluminous granites (Pankhurst *et al.* 2000; Toselli *et al.* 2007), Devonian mylonites (e.g. TIPA shear zone, Höckenreiner *et al.* 2003), and undeformed Carboniferous A-type granitic plutons (Dahlquist *et al.* 2006, 2010; Grosse *et al.* 2008). The La Cébila metamorphic complex consists of phyllites, metapsammites, quartzites, mica- and quartz-schists, gneisses and migmatites with minor calcsilicate rocks, graphite-schist layers and discordant pegmatite bodies (Espízúa & Caminos 1979; Verdecchia 2009). The metamorphic grade increases from very low in the east to high in the west, towards the contact with the Punta del Negro pluton, giving rise to a succession of metamorphic zones (chlorite, biotite, cordierite, andalusite, andalusite–K-feldspar, sillimanite–K-feldspar and cordierite–K-feldspar) roughly parallel to the contact (Verdecchia 2009; Fig. 2b). South of the study area at Quebrada de la Rioja (Fig. 2a) one granitoid that intrudes rocks equivalent to the La Cébila metamorphic complex has yielded an age of 476.4 ± 1.5 Ma (U–Pb isotope dilution (ID)-TIMS on monazite; De los Hoyos *et al.* 2008). Although the age of metamorphism is unknown, Verdecchia (2009) suggested on geological and petrological grounds that Ordovician magmatism was roughly coeval.

An Early Ordovician depositional age for the protoliths of the La Cébila metamorphic complex has been determined from biostratigraphy in quartzites from the sillimanite–K-feldspar zone that preserve a shelly fauna (Verdecchia *et al.* 2007). This age is compatible with U–Pb detrital zircon ages that yielded a maximum sedimentation age of c. 530 Ma (sample QCE-6004, Rapela *et al.* 2007; Fig. 2b). A shallow-water marine siliciclastic platform, in a foreland position relative to the Famatinian magmatic arc to the west, was previously suggested for protoliths of the La Cébila metamorphic complex (Astini *et al.* 2003, 2004; Verdecchia *et al.* 2007; Verdecchia & Baldo 2010).

Samples

One sample from each metamorphic complex referred to above was selected for U–Pb SHRIMP zircon dating.

Sample CEB-392 is a banded migmatite from southern Sierra de Ambato ($28^{\circ}50'29.10''S$, $66^{\circ}20'39.40''W$, see Fig. 2b). The mineral association consists of plagioclase, biotite, quartz, K-feldspar and secondary muscovite and chlorite, with accessory zircon (both in the matrix and as inclusions in biotite), apatite and scarce opaque minerals. The migmatite is a stromatite with alternation of leucosome and melanosome concordant with the foliation. The melanosome is composed of aligned biotite layers (<3–5 mm thick), whereas the leucosome layers (<30 mm thick) have an interlobate mosaic of quartz, plagioclase, K-feldspar and subordinate biotite.

Sample CEB-428 is a paragneiss of the inner La Cébila metamorphic complex collected near the contact with the porphyritic Punta del Negro granite ($28^{\circ}45'46.80''S$,

66°24'29.60"W, Fig. 2b). The mineral association includes cordierite, K-feldspar, biotite, plagioclase, quartz and secondary muscovite, with accessory tourmaline, zircon, monazite, apatite and opaque minerals. Compositional banding is characterized by leucocratic layers <10 mm thick consisting of slightly interlobate aggregates of cordierite, K-feldspar, plagioclase, quartz with minor biotite, and thin biotitic layers (<1 mm thick) with subordinate cordierite, plagioclase, K-feldspar and quartz. In both domains, a foliation is defined by biotite aligned parallel to the banding.

Analytical methods

Zircons were concentrated using standard crushing, washing (to decant slime), heavy liquid, and paramagnetic separation procedures as described by Rapela *et al.* (2007). The zircon-rich heavy mineral concentrates were poured onto double-sided tape, mounted in epoxy together with chips of the Temora reference zircon, sectioned approximately in half, and polished. Cathodoluminescence (CL) images (Fig. 3) were used to decipher the internal structures of the sectioned grains.

The U–Th–Pb analyses were made using SHRIMP RG at the Research School of Earth Sciences, The Australian National University, Canberra, Australia as described by Williams (1998, and references therein). Each analysis consisted of four scans through the mass range, with the reference zircon analysed once for every five unknowns. Data were reduced using the SQUID Excel macro of Ludwig (2001). Because young zircons with normal U contents have low $^{207}\text{Pb}/^{235}\text{U}$ ratios and statistically highly imprecise $^{204}\text{Pb}/^{206}\text{Pb}$ ratios, common-Pb correction was made using the measured ^{204}Pb measurements only for ages older than c. 1100 Ma, and ^{207}Pb measurement for younger ages (see Williams 1998); in the latter case there are no common-Pb corrected $^{207}\text{Pb}/^{235}\text{U}$ ratios reported in Table 1.

Tera–Wasserburg concordia plots (Fig. 4) and probability density plots with stacked histograms (Fig. 5) were constructed, and weighted-mean $^{206}\text{Pb}/^{238}\text{U}$ age calculations were carried out

using ISOPLOT/Ex (Ludwig 2003). Uncertainties on all calculated ages are reported as 95% confidence limits.

U–Pb results

CEB-392 (Ambato metamorphic complex)

Zircon grains are up to 200 μm long, rounded, anhedral, but with a minority of euhedral prismatic crystals, some of the latter showing bi-pyramidal terminations (Fig. 3a–c). CL images show irregular detrital cores and discordant low-luminescence overgrowths, some showing faint oscillatory zoning that is interpreted as metamorphic in origin. Sixty-one grains were analysed (Table 1). The zircons have moderate Th/U ratios, mostly ≤ 0.5 , and their $^{206}\text{Pb}/^{238}\text{U}$ ages range from c. 460 to c. 2650 Ma, with only three analyses being more than 10% discordant. The inheritance pattern shows a bimodal distribution with 54 ages in the range 460–1200 Ma and five Archaean ages at c. 2600 Ma (for which $^{207}\text{Pb}/^{206}\text{Pb}$ ages are considered more reliable than $^{238}\text{U}/^{206}\text{Pb}$ ages). There is a notable absence of Palaeoproterozoic ages. The two youngest ages (459 ± 5 and 471 ± 5 Ma) are from rims with high U (>1000 ppm) and low Th/U (<0.05), which suggests a metamorphic origin during the Famatinian orogeny and provides a minimum age for sedimentation, although the precise age of metamorphism cannot be determined with only two results. There is a major peak at c. 530 ± 10 Ma, defined by six analyses of grains showing oscillatory zoning characteristic of igneous crystallization, and hence it is assumed that these are of detrital origin, constraining a maximum possible age for deposition. Other significant post-Archaean age peaks occur at c. 570 Ma, c. 640 Ma, c. 900 Ma, 950–1025 Ma (concentrated at c. 1015 Ma) and, perhaps, c. 1165 Ma.

CEB-428 (La Cébila metamorphic complex)

Zircon grains are up to 100 μm in size and show a variety of rounded to euhedral prismatic shapes, some with bi-pyramidal

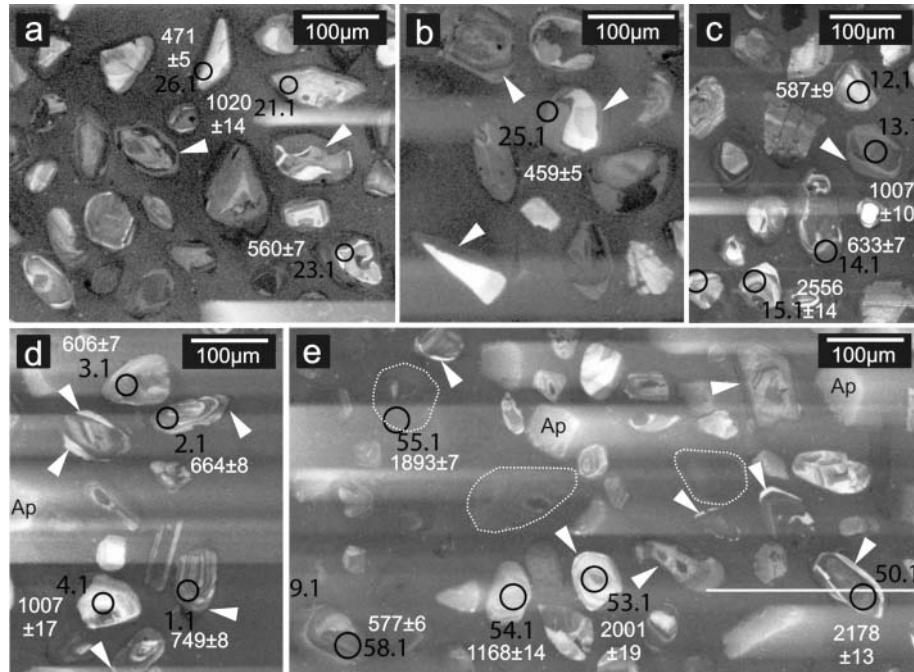


Fig. 3. CL images showing examples of the zircon grains analysed: (a–c) migmatite (CEB-392) from the Ambato metamorphic complex; (d, e) paragneiss (CEB-428) from the La Cébila metamorphic complex. Rims are indicated with white arrowheads.

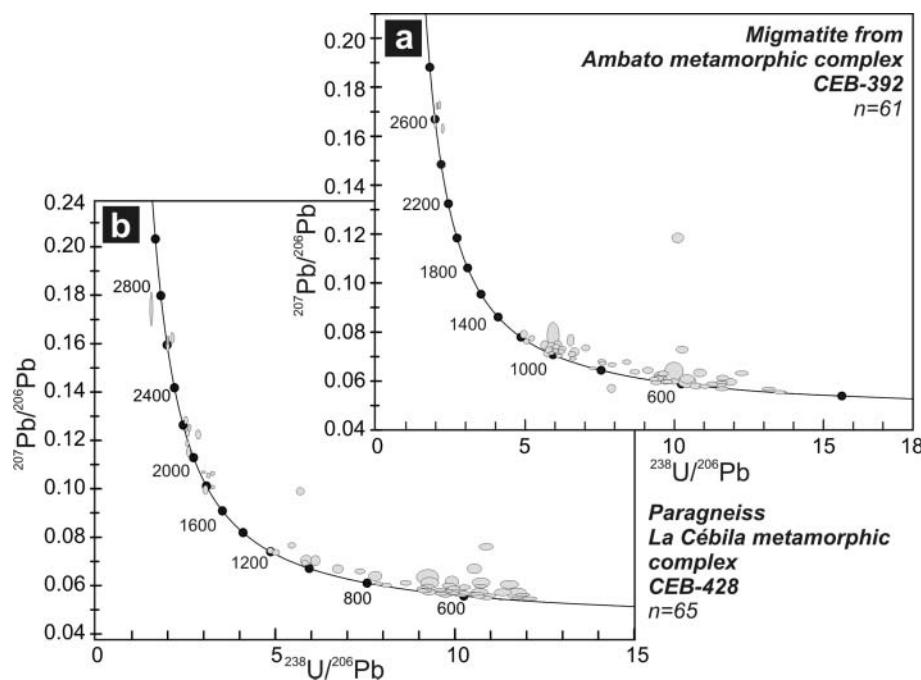


Fig. 4. Tera–Wasserburg plots for U–Pb SHRIMP data for (a) migmatite sample CEB-392 from the Ambato metamorphic complex and (b) paragneiss sample CEB-428 from the La Cébila metamorphic complex.

terminations (Fig. 3d and e). Many grains are fragments. Strong oscillatory zoning is evident in the CL images. Some grains show zoned cores overgrown by thin rims of low luminescence (Fig. 3d and e). On the other hand, some of the anhedral zircon grains show little or no zoning. Sixty-five grains were analysed, all of which yielded ages that are mostly less than 10% discordant. The majority of the grains exhibit Th/U ratios in the range 0.2–1.0, which is normal for igneous zircon. The inheritance pattern shows a similar but more continuous spread of ages than that for sample CEB-392. Five youngest ages form a coherent group with a mean age of 520 ± 10 Ma, but more prominent peaks are defined at c. 570, c. 610 and c. 660 Ma. Early Neoproterozoic ages in the range 950–1150 Ma (with a small peak at c. 1015 Ma) are less common than in CEB-392, but there is a significant grouping of Palaeoproterozoic $^{207}\text{Pb}/^{206}\text{Pb}$ ages (1750–2200 Ma, with possible minor peaks at c. 1890, c. 2050 and 2150 Ma), as well as a few Archaean ages of c. 2600 Ma.

Discussion

Neoproterozoic to Early Cambrian sedimentary rocks that were involved in the Pampean orogeny are characterized by detrital zircon provenance patterns with well-developed Brasiliano age peaks between 680 and 570 Ma and Grenvillian age peaks between 950 and 1100 Ma, but lack Pampean magmatic and metamorphic zircons with ages in the range 545–520 Ma. A minor group of Palaeoproterozoic (1.7–2.0 Ga) and Archaean (c. 2.6 Ga) grains is also present (see Sims *et al.* 1998; Pankhurst *et al.* 2000; Schwartz & Gromet 2004; Escayola *et al.* 2007; Rapela *et al.* 2007; Drobe *et al.* 2009; Adams *et al.* 2011). In contrast, the migmatite of the Ambato metamorphic complex records a maximum sedimentation age of 530 ± 10 Ma (CEB-392, Fig. 5). Age peaks corresponding to Brasiliano ages (c. 570 and c. 640 Ma) and Grenville ages (c. 950 to c. 1025 Ma) are present, whereas Palaeoproterozoic ages characteristic of the Río de la Plata craton (2260–2020 Ma; Rapela *et al.* 2007, and references

therein) are absent. Two ages of 471 ± 5 and 459 ± 5 Ma (1 σ uncertainties) determined for overgrowths on zircon grains from the Ambato complex sample are interpreted as closely approximating the age of the Famatinian metamorphic overprint and are compatible with an Early to Mid-Ordovician U–Pb monazite age (c. 470 Ma) determined by Larrovere (2009). The protolith of the complex was thus deposited between Early Cambrian and Mid-Ordovician time (based on the ICS Stratigraphic Chart of 2009).

The Ambato migmatite detrital zircon age pattern resembles that of the Middle Cambrian Negro Peinado Formation of the Sierra de Famatina (Collo *et al.* 2009), which yielded a youngest detrital zircon age of 505 ± 13 Ma (and a more significant peak at 522 ± 8 Ma) but also lacks zircons of Río de la Plata craton age (see Fig. 5). Brasiliano age peaks are present in both samples (stronger in CEB-392) and are consistent with provenance from reworking of Late Neoproterozoic sedimentary successions such as the Puncoviscana Formation (e.g. Drobe *et al.* 2009; Adams *et al.* 2011; Hauser *et al.* 2011) and the equivalent Ancasti metamorphic complex of the easternmost Sierras Pampeanas (Rapela *et al.* 2007; Murra *et al.* 2011; see Fig. 5). Provenance of the Puncoviscana Formation has been related in part to sources in the Kalahari craton (see Schwartz & Gromet 2004; Rapela *et al.* 2007). On the other hand, 1165–950 Ma zircon ages in the Ambato metamorphic complex and the Negro Peinado Formation suggest derivation from a Grenville-age basement similar to that recognized in the Western Sierras Pampeanas, probably the tip of a much larger terrane that collided during the Pampean orogeny embracing the Arequipa block and Amazonia (see Rapela *et al.* 2007; Casquet *et al.* 2008, 2010). Alternatively, the Grenvillian-age grains might be derived from reworking of the underlying Puncoviscana Formation. Collo *et al.* (2009) interpreted the Negro Peinado Formation as deposited in a Mid-Cambrian foreland basin related to the Pampean orogen exposed in the east.

Protoliths of the La Cébila metamorphic complex are Early Ordovician according to fossil remains in quartzitic layers (Verdecchia *et al.* 2007). The provenance pattern of the para-

Table 1. Summary of SHRIMP U–Pb zircon results for Ambato and La Cebila metamorphic complexes

Grain and spot	U (ppm)	Th (ppm)	Th/U	$^{206}\text{Pb} / ^{204}\text{Pb}$		F206 (%)	Total ratios	Radiogenic ratios			P	$^{206}\text{Pb} / ^{238}\text{U}$	\pm	$^{207}\text{Pb} / ^{206}\text{Pb}$	\pm	% Disc.	
				$^{238}\text{U} / ^{206}\text{Pb}$	\pm			$^{206}\text{Pb} / ^{235}\text{U}$	\pm	$^{207}\text{Pb} / ^{206}\text{Pb}$							
<i>CFB-392 Ambato</i>																	
1.1	65	20	0.31	10	0.000997	1.70	5.689	0.091	0.0765	0.0012	0.1758	0.0029	1.854	0.067	0.736	1044	16
2.1	2366	19	0.01	279	0.000014	0.02	7.292	0.075	0.0664	0.0004	0.1371	0.0014	1.254	0.015	0.0663	828	8
3.1	527	28	0.05	41	0.000046	<0.01	11.048	0.127	0.0586	0.0006	0.0905	0.0011	0.1998	0.0031	0.0792	559	6
4.1	107	48	0.45	18	0.000141	0.24	4.992	0.076	0.0812	0.0011	0.1977	0.0011	0.1272	0.0017	0.0604	1174	16
5.1	609	356	0.58	66	0.000101	0.18	7.922	0.093	0.0577	0.0011	0.1272	0.0017	1.120	0.055	0.0639	772	9
6.1	44	75	1.72	6	0.000877	1.50	5.972	0.129	0.0806	0.0037	0.1649	0.0037	1.547	0.144	0.0680	984	20
7.1	459	126	0.27	42	0.000230	<0.01	9.429	0.114	0.0603	0.0007	0.1062	0.0013	1.062	0.021	0.0061	651	8
8.1	464	6	0.01	35	0.000151	0.13	11.301	0.136	0.0595	0.0007	0.0884	0.0011	1.169	0.116	0.0006	1109	52
9.1	515	119	0.23	38	–	0.15	11.551	0.137	0.0593	0.0007	0.0864	0.0010	1.176	0.117	0.0006	816	12
10.1	526	429	0.82	227	0.000013	0.02	4.988	0.022	0.1749	0.0013	0.0529	0.0017	1.140	0.055	0.0680	1176	40
11.1	582	145	0.25	50	0.000053	0.05	10.012	0.117	0.0607	0.0007	0.0998	0.0012	1.163	0.144	0.0680	737	5
12.1	158	60	0.38	13	0.000515	0.26	10.460	0.160	0.0617	0.0012	0.0954	0.0015	1.144	0.144	0.0680	887	-13
13.1	735	269	0.37	107	0.000058	0.10	5.912	0.066	0.0729	0.0005	0.1690	0.0019	1.680	0.023	0.0006	989	16
14.1	445	21	0.05	40	0.000165	0.40	9.651	0.117	0.0641	0.0007	0.1032	0.0013	1.254	0.015	0.0006	633	7
15.1	259	107	0.41	98	0.000036	0.05	2.261	0.028	0.1703	0.0014	0.4421	0.0055	1.0354	0.015	0.0006	2360	24
16.1	39	95	0.27	33	0.000156	0.27	6.271	0.082	0.0693	0.0009	0.1590	0.0014	1.591	0.037	0.0006	546	6
17.1	515	97	0.19	58	–	<0.01	7.594	0.089	0.0693	0.0009	0.1317	0.0016	1.260	0.019	0.0006	954	12
18.1	305	24	0.08	40	0.000058	0.10	6.593	0.084	0.0724	0.0008	0.1514	0.0020	1.481	0.029	0.0009	955	27
19.1	226	250	1.11	93	0.000034	0.05	2.079	0.027	0.1800	0.0009	0.4807	0.0061	1.1905	0.167	0.0006	2649	5
20.1	1719	2	0.00	127	0.000156	0.29	11.616	0.124	0.0604	0.0004	0.0858	0.0009	1.250	0.027	0.0006	531	6
21.1	241	144	0.60	39	0.000193	0.32	5.252	0.070	0.0794	0.0009	0.1898	0.0025	2.006	0.044	0.0006	1120	14
22.1	833	99	0.12	73	0.000059	<0.01	9.802	0.110	0.0605	0.0008	0.1020	0.0012	1.020	0.020	0.0006	626	7
23.1	307	26	0.08	26	0.000220	0.26	7.86	0.153	0.1229	0.0014	0.0908	0.0019	1.029	0.026	0.0006	560	7
24.1	238	61	0.26	35	0.000346	0.59	5.828	0.079	0.0743	0.0009	0.1710	0.0024	1.685	0.049	0.0015	1018	13
25.1	1357	6	0.00	86	0.000071	<0.01	13.563	0.149	0.0574	0.0008	0.0737	0.0013	1.0574	0.023	0.0006	459	5
26.1	1201	47	0.04	78	0.000121	0.11	13.187	0.147	0.0574	0.0008	0.0757	0.0013	1.0574	0.023	0.0006	471	5
27.1	525	101	0.19	47	0.000244	0.20	9.547	0.116	0.0627	0.0007	0.1045	0.0013	1.154	0.023	0.0006	641	8
28.1	289	202	0.70	49	–	<0.01	5.073	0.065	0.0765	0.0008	0.1973	0.0025	2.143	0.038	0.0006	1161	14
29.1	549	226	0.41	76	0.000111	0.19	6.169	0.102	0.0708	0.0008	0.1618	0.0019	1.566	0.025	0.0006	967	11
30.1	239	76	0.32	22	0.000288	0.26	9.540	0.136	0.0632	0.0011	0.1045	0.0015	1.045	0.024	0.0006	935	22
31.1	854	41	0.05	63	0.000084	<0.01	11.629	0.133	0.0575	0.0006	0.0861	0.0012	1.0574	0.023	0.0006	532	6
32.1	351	49	0.14	25	0.000050	0.34	11.864	0.160	0.0605	0.0010	0.0840	0.0012	1.0574	0.023	0.0006	520	7
33.1	453	183	0.40	64	0.000030	0.05	6.065	0.073	0.0575	0.0007	0.1648	0.0020	1.710	0.026	0.0006	983	11
34.1	664	111	0.17	58	0.000159	<0.01	9.748	0.114	0.0627	0.0008	0.1026	0.0012	1.1517	0.0165	0.0006	1076	19
35.1	71	26	0.36	6	0.000061	0.56	10.015	0.209	0.0648	0.0008	0.0993	0.0021	1.204	0.020	0.0006	630	7
36.1	314	58	0.19	26	0.000608	1.82	10.291	0.135	0.0744	0.0010	0.0954	0.0013	1.0988	0.0115	0.0006	610	13
37.1	734	44	0.06	75	0.000035	0.70	8.428	0.096	0.0691	0.0006	0.1178	0.0012	1.443	0.047	0.0006	718	8
38.1	178	49	0.27	23	0.000023	0.41	6.684	0.098	0.0736	0.0011	0.1490	0.0022	1.443	0.047	0.0006	935	12
39.1	269	90	0.33	107	0.000002	0.00	2.161	0.028	0.1805	0.0011	0.4628	0.0060	1.1517	0.165	0.0006	2657	10
40.1	743	76	1.10	110	0.000126	0.22	7.623	0.086	0.0604	0.0006	0.1309	0.0015	1.204	0.020	0.0006	793	8
41.1	597	79	0.13	59	0.000139	0.24	8.708	0.103	0.0647	0.0007	0.1146	0.0014	1.0988	0.0115	0.0006	699	8
42.1	356	24	0.07	33	0.000173	0.46	9.146	0.119	0.0656	0.0009	0.1088	0.0015	1.069	0.021	0.0006	666	8
43.1	321	18	0.06	25	0.000094	0.67	10.886	0.145	0.0644	0.0011	0.0912	0.0012	1.0844	0.022	0.0006	563	7
44.1	839	401	0.48	109	0.000045	0.08	6.627	0.074	0.0706	0.0008	0.1508	0.0017	1.454	0.020	0.0006	905	9
45.1	432	94	0.22	53	0.000389	0.67	7.061	0.090	0.0752	0.0008	0.1403	0.0019	1.304	0.037	0.0006	846	11
46.1	626	95	0.15	68	0.000022	0.40	7.939	0.093	0.0679	0.0006	0.1255	0.0015	1.0720	0.0196	0.0006	829	25
47.1	91	19	0.21	7	0.0000453	0.13	10.448	0.202	0.0607	0.0017	0.0956	0.0019	1.0602	0.026	0.0006	588	11
48.1	1196	252	0.21	110	0.000199	0.03	9.356	0.104	0.0617	0.0005	0.1069	0.0012	1.0623	0.019	0.0006	654	7
49.1	871	17	0.02	64	0.000242	0.53	11.652	0.140	0.0623	0.0007	0.0854	0.0010	1.0623	0.025	0.0006	528	6
50.1	186	65	0.35	26	0.000602	1.03	6.129	0.087	0.0768	0.0010	0.1628	0.0025	1.687	0.061	0.0006	973	14
51.1	1812	20	0.01	127	0.000554	0.87	12.293	0.141	0.0643	0.0007	0.0806	0.0009	1.0628	0.026	0.0006	500	6
52.1	495	6	0.01	65	0.000124	0.06	6.543	0.078	0.0785	0.0017	0.1527	0.0017	1.643	0.041	0.0006	916	10
53.1	294	167	0.57	126	0.000003	<0.01	2.013	0.025	0.1720	0.0008	0.4966	0.0063	1.1773	0.159	0.0006	2599	27
54.1	569	78	0.14	78	0.000073	0.13	6.235	0.073	0.0736	0.0006	0.1602	0.0019	1.0602	0.026	0.0006	1001	23
55.1	167	44	0.26	24	0.000375	0.64	5.953	0.089	0.0740	0.0011	0.1681	0.0026	1.733	0.054	0.0006	1072	54
56.1	614	162	0.26	87	0.000050	0.09	6.049	0.070	0.0740	0.0011	0.1652	0.0019	1.669	0.026	0.0006	986	11
57.1																	

CEB-428 La Cebolla											
1.1	409	205	0.50	43	0.000046	0.03	8.118	0.094	0.1231	0.0006	0.0015
2.1	378	75	0.20	35	0.000162	0.06	9.206	0.110	0.0622	0.0007	0.0013
3.1	275	194	0.71	23	0.000168	0.20	10.121	0.127	0.0617	0.0009	0.0013
4.1	62	57	0.91	9	0.000351	0.60	5.877	0.104	0.0759	0.0016	0.0031
5.1	260	86	0.33	38	0.000087	0.15	5.857	0.072	0.0745	0.0008	0.0021
6.1	306	171	0.56	83	0.000025	0.04	3.157	0.037	0.1146	0.0007	0.0037
7.1	215	201	0.94	20	-	<0.01	9.311	0.123	0.0614	0.0009	0.0015
8.1	532	233	0.44	139	0.000025	0.04	3.286	0.036	0.1094	0.0005	0.0042
9.1	108	86	0.80	46	0.000021	0.03	2.029	0.029	0.1747	0.0023	0.0023
10.1	143	77	0.54	13	-0.000085	0.02	9.749	0.146	0.0609	0.0016	0.0125
11.1	109	109	0.55	66	0.000024	0.04	2.576	0.032	0.1908	0.0008	0.0012
12.1	70	46	0.65	39	-	<0.01	1.566	0.024	0.0617	0.0051	0.0051
13.1	111	78	0.70	12	0.000369	0.44	7.803	0.121	0.0686	0.0014	0.0020
14.1	200	73	0.37	66	0.000057	0.09	2.606	0.032	0.1255	0.0016	0.0048
15.1	178	142	0.80	15	0.000312	0.14	9.992	0.141	0.0614	0.0011	0.0099
16.1	619	466	0.75	53	0.000122	0.36	9.976	0.113	0.0632	0.0006	0.0099
17.1	73	108	1.47	20	0.000053	0.08	3.083	0.051	0.1082	0.0014	0.0045
18.1	185	70	0.38	15	0.00004	0.25	10.699	0.153	0.0612	0.0012	0.0032
19.1	367	42	0.11	35	0.000086	0.11	9.132	0.111	0.0628	0.0008	0.0014
20.1	304	145	0.48	129	0.000017	0.02	2.029	0.024	0.1766	0.0013	0.0057
21.1	144	114	0.79	12	0.000811	0.82	10.751	0.164	0.0657	0.0013	0.0092
22.1	828	288	0.35	60	0.000054	0.19	11.815	0.131	0.0594	0.0005	0.0045
23.1	190	155	0.82	14	0.000041	0.84	11.536	0.165	0.0649	0.0012	0.0013
24.1	246	69	0.28	39	0.000075	0.13	5.485	0.068	0.0827	0.0008	0.0024
25.1	1104	194	0.18	95	0.000061	0.07	9.993	0.107	0.0609	0.0004	0.0011
26.1	714	148	0.21	57	0.000426	0.61	10.742	0.120	0.0640	0.0006	0.0025
27.1	325	112	0.34	38	0.000121	0.21	7.376	0.090	0.0709	0.0008	0.0153
28.1	165	51	0.31	12	0.000003	0.33	11.740	0.178	0.0606	0.0017	0.0049
29.1	291	59	0.29	30	0.000075	0.13	5.713	0.075	0.0649	0.0011	0.0024
30.1	325	422	1.30	28	0.000175	0.46	9.860	0.122	0.0642	0.0008	0.0110
31.1	219	110	0.50	17	0.000283	0.35	11.356	0.155	0.0612	0.0011	0.0077
32.1	84	346	0.72	45	0.000024	0.01	9.139	0.106	0.0620	0.0006	0.0194
33.1	30	35	0.35	25	-	<0.01	2.876	0.044	0.1334	0.0014	0.0379
34.1	453	398	0.88	37	0.000860	1.56	10.562	0.127	0.0720	0.0014	0.0932
35.1	406	105	0.26	33	0.000122	0.10	10.591	0.128	0.0602	0.0008	0.0043
36.1	1282	65	0.05	93	0.000035	0.30	11.905	0.128	0.0601	0.0004	0.0037
37.1	203	97	0.48	36	0.000110	0.18	4.885	0.062	0.0803	0.0008	0.0204
38.1	1226	347	0.28	133	0.000087	0.04	7.892	0.084	0.1267	0.0014	0.0164
39.1	424	239	0.56	42	0.000117	0.32	8.647	0.102	0.0655	0.0007	0.1153
40.1	179	96	0.53	61	0.000127	0.19	2.531	0.032	0.1399	0.0012	0.3943
41.1	859	313	0.36	67	0.000026	0.10	9.986	0.121	0.0579	0.0005	0.0051
42.1	118	102	0.86	47	0.000024	0.03	2.159	0.032	0.1776	0.0015	0.0070
43.1	360	77	0.21	31	0.000045	0.71	9.938	0.125	0.0661	0.0017	0.0099
44.1	579	155	0.27	48	0.000392	0.46	10.272	0.119	0.0635	0.0007	0.0069
45.1	332	109	0.33	87	0.000050	0.08	3.274	0.038	0.1156	0.0007	0.0352
46.1	1870	120	0.06	158	0.000005	<0.01	10.192	0.107	0.0599	0.0004	0.0981
47.1	115	51	0.44	15	0.000326	0.57	6.759	0.103	0.0719	0.0012	0.0233
48.1	403	139	0.34	37	0.000061	0.28	9.261	0.112	0.0639	0.0008	0.0177
49.1	874	490	0.56	70	0.000124	<0.01	10.685	0.118	0.0792	0.0013	0.0041
50.1	583	57	0.10	190	0.000023	0.03	2.633	0.029	0.1364	0.0010	0.3796
51.1	229	82	0.36	32	0.000042	0.07	6.136	0.079	0.0757	0.0012	0.1629
52.1	874	485	0.85	79	0.000075	0.16	9.597	0.105	0.0623	0.0005	0.1040
53.1	591	151	0.63	49	0.000134	0.20	6.265	0.035	0.1249	0.0012	0.3633
54.1	159	52	0.32	27	0.000081	0.14	5.030	0.067	0.0794	0.0009	0.3759
55.1	743	1628	2.19	211	0.000021	0.03	3.031	0.032	0.1161	0.0004	0.3299
56.1	987	667	0.68	70	0.000073	0.13	12.093	0.132	0.0584	0.0006	0.0826
57.1	987	792	0.80	88	0.000022	<0.01	9.634	0.106	0.0606	0.0005	0.1038
58.1	568	375	0.66	46	0.000097	0.11	10.659	0.122	0.0602	0.0008	0.0937
59.1	964	444	0.67	76	0.000010	<0.01	10.907	0.120	0.0586	0.0005	0.0917
60.1	82	46	0.57	8	0.000162	0.51	9.317	0.161	0.0656	0.0016	0.1068
61.1	147	98	0.66	49	0.000093	0.14	2.583	0.034	0.1341	0.0010	0.3866
62.1	612	96	0.16	48	0.001856	0.89	10.887	0.132	0.0821	0.0004	0.0892
63.1	43	16	0.37	4	0.000037	0.82	9.263	0.204	0.0682	0.0023	0.1071
64.1	318	489	1.54	27	0.000093	0.09	10.236	0.128	0.0606	0.0008	0.0976
65.1	6252	669	0.11	470	0.000015	<0.01	11.427	0.116	0.0579	0.0002	0.0876

For CEB-392, error in Temora reference zircon calibration was 0.74% for the analytical session. For CEB-428, error in Temora reference zircon calibration was 0.79% for the analytical session. Uncertainties are given at the 1 σ level. f206 (%) denotes the percentage of ^{206}Pb that is common Pb. For areas older than c. 1100 Ma, correction for common Pb made using the measured $^{38}\text{U} / ^{206}\text{Pb}$ and $^{207}\text{Pb} / ^{206}\text{Pb}$ ratio. For areas younger than c. 1100 Ma, correction for common Pb made using $^{207}\text{Pb} / ^{206}\text{Pb}$. Pb* is radiogenic ^{206}Pb .

1 Discordant point excluded from Figure 5.

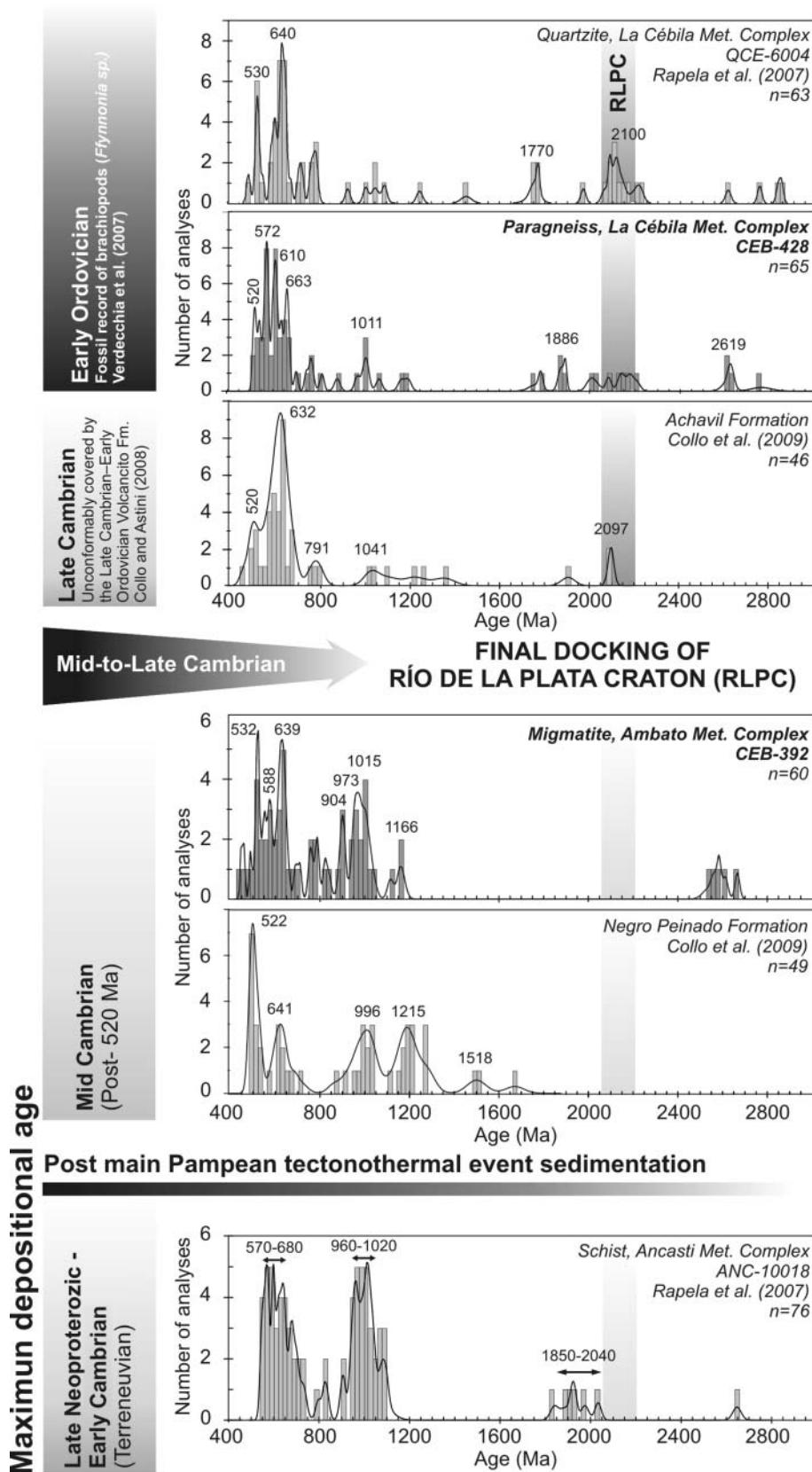


Fig. 5. Detrital zircon U-Pb age patterns for metasedimentary samples from the La Cébila metamorphic complex (CEB-428, this work; quartzite QCE-6004, Rapela *et al.* 2007), the Ambato metamorphic complex (CEB-392, this work, excluding one discordant point from Table 1), the Ancasti metamorphic complex (schist ANC-10018, Rapela *et al.* 2007), and the Negro Peinado and Achavil formations (Collo *et al.* 2009). The position of final docking of Río de la Plata craton (RLPC) is indicated as Mid- to Late Cambrian.

gneiss studied here is characterized by Pampean (*c.* 520 to *c.* 540 Ma), Brasiliano (*c.* 570 to *c.* 660 Ma) and Grenville (*c.* 1015 Ma) age peaks, but significantly includes Palaeoproterozoic zircons with ages that match those of the Río de la Plata craton

(2.02–2.26 Ga; Hartmann *et al.* 2002; Rapela *et al.* 2007) and a few older grains. Similar results were obtained on a quartzite from this same complex by Rapela *et al.* (2007) (Fig. 5), which strengthens the view that by the Early Ordovician the Rio de la

Plata craton was in its present position, providing clastic detritus to wide sedimentary basins in the west that joined those resulting from the reworking of the underlying metasedimentary rocks.

Detrital zircons with Río de la Plata craton ages of c. 2100 Ma were also recorded by Collo *et al.* (2009) in the Achavil Formation, as well as Pampean (c. 520 Ma), Brasiliano (c. 630 Ma) and Grenville (c. 1040–1120 Ma) age peaks, much like the La Cébila metamorphic complex provenance. The maximum age of the Achavil Formation is poorly constrained as the youngest detrital zircon age peak is 519 ± 23 Ma (1σ). The minimum sedimentation age, however, is Late Cambrian, as the base of the overlying Volcancito Formation contains fossils of that age (Astini 2003; Albanezi *et al.* 2005). Collo & Astini (2008) and Collo *et al.* (2009) suggested that the Achavil Formation had to be late Middle to Late Cambrian and younger than the Negro Peinado Formation on the basis of stratigraphical evidence and detrital zircon age patterns. Thus the record of Rio de la Plata craton influence extends back into the Late Cambrian.

From these results we suggest here that the Ambato metamorphic complex is older than both the La Cébila metamorphic complex and the Achavil Formation, and is probably equivalent to the Middle Cambrian Negro Peinado Formation of Sierra de Famatina. Both the Ambato metamorphic complex and the Negro Peinado Formation lack detrital zircons of Rio de la Plata age (Fig. 5). On the other hand, the Achavil Formation and the La Cébila metamorphic complex, although not equivalent stratigraphically, both contain detrital zircon populations with age peaks in part similar to those of the Ambato and Negro Peinado formations, but additionally include Palaeoproterozoic zircons that can be related to the Rio de la Plata craton. The latter implies that by the Mid- or Late Cambrian the Rio de la Plata craton had reached a palaeogeographical position close to the present position. Moreover, the Rio de la Plata craton does not show evidence of deformation and metamorphism of Pampean age (Rapela *et al.* 2007), which implies that it was the source of zircons for the Achavil sedimentary basin after the main Pampean tectonothermal events (i.e. between 545 and 520 Ma). Orographic barriers such as the rising Pampean orogen played a transient role sometime between 520 and 510 Ma. By the time the Achavil Formation was deposited, such a barrier did not exist and Palaeoproterozoic zircons could easily reach sedimentary realms in the west. We conclude that the Rio de la Plata craton reached a position close to the present one sometime between the end of the Pampean tectonothermal event (c. 520 Ma) and the deposition of the Achavil Formation in the Mid- to Late Cambrian.

One alternative source for zircons with Palaeoproterozoic ages similar to the Rio de la Plata craton could have been in the Amazonia craton, particularly in the Maroni–Itacaiunas province of northern Amazonia (e.g. Cordani & Teixeira 2007). If a static palaeogeographical model is implied (i.e. that Amazonia were in a position similar to its present position relative to other cratons, basins and orogenic belts of southern South America), then zircons would have had to travel between 3000 and 4000 km from the Maroni–Itacaiunas province (all the way round the craton itself) to reach the La Cébila basin. This enormous distance might be feasible if long-range energetic submarine currents were involved, but we can also add the fact that to our knowledge there is no evidence for the anticipated string of basins similar to La Cébila (Early Ordovician) that would connect the Sierras Pampeanas with the Maroni–Itacaiunas region of Amazonia. The possibility of the Palaeoproterozoic Arequipa block of southern Peru as another alternative source to the Rio de la Plata craton is more difficult to justify: Trans-

Amazonian or Eburnian ages between c. 1.8 and 2.1 Ga have been recorded from this block (Loewy *et al.* 2004; Casquet *et al.* 2010), but also sedimentary rocks with detrital zircons with ages between c. 1.2 and 1.6 Ga that are not recorded in the La Cébila basin.

Conclusions

The metasedimentary Ambato metamorphic complex and the La Cébila metamorphic complex that crop out at the southern end of the Sierra de Ambato and NE tip of Sierra de Velasco are Middle Cambrian and Early Ordovician respectively. The protoliths for these complexes were deposited before the start of the Famatinian magmatism in the late Early Ordovician, and after the Pampean orogeny. Sources for most of the detrital zircons in both formations can be found in neighbouring regions, particularly in the Pampean orogen and the Proterozoic Western Sierras Pampeanas. However, Palaeoproterozoic zircons are recorded only in the La Cébila metamorphic complex. The source for these zircons probably lay in the Rio de la Plata craton to the east. The Late Cambrian Achavil Formation also contains Palaeoproterozoic zircons.

The Rio de la Plata craton was juxtaposed obliquely to the Pampean orogen, reaching a position close to its present one sometime between the end of the main Pampean tectonothermal event (c. 520 Ma) and when it became the source for Palaeoproterozoic detrital zircons in the Achavil Formation (i.e. during the Mid- to Late Cambrian interval). The situation persisted in the earliest Ordovician when the La Cébila sedimentary succession was deposited.

Financial support for this paper was provided by Argentine Grants PIP-CONICET 5719 and FONCYT BID 1728/0C AR PICT-1009 and Spanish MEC Grants CGL2005-02065/BTE, CGL2009-07984 and UCM-Santander GR58/08. We acknowledge comments by V. Ramos and one anonymous reviewer.

References

- ACEÑOLAZA, F. & TOSSELLI, A. 1977. Observaciones geológicas y paleontológicas sobre el Ordovícico de la zona de Chaschuil, Provincia de Catamarca. *Acta Geológica Lilloana, Tucumán*, **XIV**, 233–259.
- ADAMS, C.J., MILLER, H., ACEÑOLAZA, F.G., TOSSELLI, A.J. & GRIFFIN, W.L. 2011. The Pacific Gondwana margin in the late Neoproterozoic–early Paleozoic: Detrital zircon U–Pb ages from metasediments in northwest Argentina reveal their maximum age, provenance and tectonic setting. *Gondwana Research*, **19**, 71–83.
- ALBANEZI, G.L., ESTEBAN, S.B., ORTEGA, G., HÜNICKEN, M.A. & BARNES, C.R. 2005. Bioestratigrafía y ambientes sedimentarios de las Formaciones Volcancito y Bordo Atravesado (Cámbrico Superior–Ordovícico Inferior), Sistema de Famatina, provincia de La Rioja. In: DAHLQUIST, J.A., BALDO, E.G. & ALASINO, P.H. (eds) *Geología de la provincia de La Rioja, Precámbrico–Paleozoico Inferior*. Asociación Geológica Argentina, Serie D, **8**, 42–64.
- ASTINI, R.A. 2003. The Ordovician proto-Andean basins. In: BENEDETTO, J.L. (ed.) *Ordovician Fossils of Argentina*. Secretaría de Ciencia y Tecnología, Universidad Nacional de Córdoba, 1–74.
- ASTINI, R.A. & DÁVILA, F.M. 2004. Ordovician back arc foreland and Ocloyic thrust belt development on the western Gondwana margin as a response to Precordillera terrane accretion. *Tectonics*, **23**, TC4008, doi:10.1029/2003TC001620.
- ASTINI, R.A., DÁVILA, F.M., RAPELA, C.W., PANKHURST, R.J. & FANNING, C.M. 2003. Ordovician back-arc clastic wedge in the Famatina Ranges: New ages and implications for reconstruction of the Proto-Andean Gondwana margin. In: ALBANEZI, G.L., BERESI, M.S. & PERALTA, S.H. (eds) *Ordovician from the Andes*. Serie de Correlación Geológica, San Miguel de Tucumán, Argentina, **17**, 375–380.
- ASTINI, R.A., DÁVILA, F.M., COLLO, G. & MARTINA, F. 2004. La Formación La Aguadita (Ordovícico medio-superior): Su implicancia en la evolución temprana del Famatina y como parte del orógeno ocloyico en el noroeste argentino. In: DAHLQUIST, J.A., BALDO, E.G. & ALASINO, P.H. (eds) *Geología*

- de la provincia de La Rioja, Precámbrico–Paleozoico Inferior.* Asociación Geológica Argentina, Serie D, **8**, 67–84.
- BALDO, E., CASQUET, C., ET AL. 2006. Neoproterozoic A-type magmatism in the Western Sierras Pampeanas (Argentina): evidence for Rodinia break-up along a proto-Iapetus rift? *Terra Nova*, **18**, 388–394.
- BÜTTNER, S.H., GLODNY, J., LUCASSEN, F., WEMMERD, K., ERDMANN, S., HANDLER, R. & FRANZ, G. 2005. Ordovician metamorphism and plutonism in the Sierra de Quilmes metamorphic complex: Implications for the tectonic setting of the northern Sierras Pampeanas (NW Argentina). *Lithos*, **83**, 143–181.
- CAMINOS, R. 1979. Sierras Pampeanas Noroccidentales. Salta, Tucumán, Catamarca, La Rioja y San Juan. In: TURNER, J.C.M. (ed.) *Segundo Simposio de Geología Regional Argentina*. Academia Nacional de Ciencias, Córdoba, 225–291.
- CASQUET, C., BALDO, E., PANKHURST, R.J., RAPELA, C.W., GALINDO, C., FANNING, C.M. & SAAVEDRA, J. 2001. Involvement of the Argentine Precordillera Terrane in the Famatinian mobile belt: Geochronological (U–Pb SHRIMP) and metamorphic evidence from the Sierra de Pie de Palo. *Geology*, **29**, 703–706.
- CASQUET, C., PANKHURST, R.J., ET AL. 2008. The Mesoproterozoic Maz terrane in the Western Sierras Pampeanas, Argentina, equivalent to the Arequipa–Antofalla block of southern Peru? Implications for West Gondwana margin evolution. *Gondwana Research*, **13**, 163–175.
- CASQUET, C., RAPELA, C.W., PANKHURST, R.J., BALDO, E., GALINDO, C., FANNING, M. & SAAVEDRA, J. 2009. Proterozoic terranes in southern South America: Accretion to Amazonia, involvement in Rodina formation and further west Gondwana accretion (Abstract). In: *Rodinia: Supercontinents, Superplumes and Scotland. Fermor Meeting, Edinburgh, 6–13 September 2009*. School of Geosciences, University of Edinburgh/Geological Society of London, 61.
- CASQUET, C., FANNING, C.M., GALINDO, C., PANKHURST, R.J., RAPELA, C.W. & TORRES, P. 2010. The Arequipa Massif of Peru: New SHRIMP and isotope constraints on a Paleoproterozoic infier in the Grenvillian orogen. *Journal of South American Earth Sciences*, **29**, 128–142.
- COLLO, G. & ASTINI, R.A. 2008. La Formación Achaval: una unidad diferenciable dentro del basamento metamórfico de bajo grado del Famatina en la región pampeana de los Andes Centrales. *Revista de la Asociación Geológica Argentina*, **63**, 344–362.
- COLLO, G., ASTINI, R.A., CAWOOD, P.A., BUCHAN, C. & PIMENTEL, M. 2009. U–Pb detrital zircon ages and Sm–Nd isotopic features in low-grade metasedimentary rocks of the Famatina belt: implications for late Neoproterozoic–early Palaeozoic evolution of the proto-Andean margin of Gondwana. *Journal of the Geological Society, London*, **116**, 1–17.
- CORDANI, U.G. & TEIXEIRA, W. 2007. Proterozoic accretionary belts in the Amazonian Craton. In: HATCHER, R.D., JR., CARLSON, M.P., McBRIDE, J.H. & MARTINEZ CATALÁN, J.R. (eds) *4-D Framework of Continental Crust*. Geological Society of America, Memoirs, **200**, 297–320.
- CORDANI, U.G., D'AGRELLA-FILHO, M.S., BRITO-NEVES, B.B. & TRINDADE, R.I.F. 2003. Tearing up Rodinia: the Neoproterozoic palaeogeography of South American cratonic fragments. *Terra Nova*, **15**, 350–359.
- DAHLQUIST, J.A., PANKHURST, R.J., ET AL. 2006. The San Blas Pluton: An example of Carboniferous plutonism in the Sierras Pampeanas, Argentina. *Journal of South American Earth Sciences*, **20**, 341–350.
- DAHLQUIST, J.A., PANKHURST, R.J., ET AL. 2008. New SHRIMP U–Pb data from the Famatina Complex: constraining Early–Mid Ordovician Famatinian magmatism in the Sierras Pampeanas, Argentina. *Geologica Acta*, **6**, 319–333.
- DAHLQUIST, J.A., ALASINO, P.H., EBY, N., GALINDO, C. & CASQUET, C. 2010. Fault controlled Carboniferous A-type magmatism in the proto-Andean foreland (Sierras Pampeanas, Argentina): Geochemical constraints and petrogenesis. *Lithos*, **115**, 65–81.
- DE LOS HOYOS, C.R., BASEI, M.A., ROSSI, J.N. & TOSELLI, A.J. 2008. Four new ID-TIMS U–Pb monazite ages for deformed and undeformed granitoids in the eastern sector of the Velasco range, Sierras Pampeanas, Argentina. In: *VI South American Symposium on Isotope Geology, Extended Abstracts Volume (CD-ROM)*.
- DROBE, M., LÓPEZ DE LUCHI, M.G., STEENKEN, A., FREI, R., NAUMANN, R., SIEGESMUND, S. & WEMMER, K. 2009. Provenance of the late Proterozoic to early Cambrian metaclastic sediments of the Sierra de San Luis (Eastern Sierras Pampeanas) and Cordillera Oriental, Argentina. *Journal of South American Earth Sciences*, **28**, 239–262.
- ESCAVOLA, M.P., PIMENTEL, M. & ARMSTRONG, R. 2007. Neoproterozoic backarc basin: Sensitive high-resolution ion microprobe U–Pb and Sm–Nd isotopic evidence from the Eastern Pampean Ranges, Argentina. *Geology*, **35**, 495–498.
- ESPIZÚA, L. & CAMINOS, R. 1979. Las rocas metamórficas de la Formación La Cébila, Sierra de Ambato, provincias de Catamarca y La Rioja. *Boletín de la Academia Nacional de Ciencias de Córdoba, Argentina*, **53**, 125–142.
- GROSSE, P., SÖLLNER, F., BÁEZ, M., TOSELLI, A.J., ROSSI, J.N. & DE LA ROSA, J. 2008. Lower Carboniferous post-orogenic granites in central–eastern Sierra de Velasco, Sierras Pampeanas, Argentina: U–Pb monazite geochronology, geochemistry and Sr–Nd isotopes. *International Journal of Earth Sciences*, **98**, 1001–1025.
- HARTMANN, L.A., SANTOS, J.O.S., CINGOLANI, C.A. & MCNAUGHTON, N.J. 2002. Two Palaeoproterozoic orogenies in the evolution of the Tandilia Belt, Buenos Aires, as evidenced by zircon U–Pb SHRIMP geochronology. *International Geology Review*, **44**, 528–543.
- HAUSER, N., MATTEINI, M., OMARINI, R.H. & PIMENTEL, M.M. 2011. Combined U–Pb and Lu–Hf isotope data on turbidites of the Paleozoic basement of NW Argentina and petrology of associated igneous rocks: Implications for the tectonic evolution of western Gondwana between 560 and 460 Ma. *Gondwana Research*, **19**, 100–127.
- HÖCKENREINER, M., SÖLLNER, F. & MILLER, H. 2003. Dating the TIPA shear zone: an Early Devonian terrane boundary between Famatinian and Pampean systems (NW-Argentina). *Journal of South American Earth Sciences*, **16**, 45–66.
- JORDAN, T. & ALLMENDINGER, R. 1986. The Sierras Pampeanas of Argentina: a modern analogue of Laramide deformation. *American Journal of Science*, **286**, 737–764.
- LARROVERE, M. 2009. *Petrología de la faja migmatítica entre el flanco noroccidental de la sierra de Ancasti, su continuación en la sierra de Aconquija y el flanco nororiental de la sierra de Ambato*. PhD thesis, Universidad Nacional de Córdoba.
- LIRA, R.R., MILLONE, H.A., KIRSCHBAUM, A.M. & MORENO, R.S. 1997. Calc-alkaline arc granitoid activity in the Sierra Norte–Ambargasta Ranges, central Argentina. *Journal of South American Earth Sciences*, **10**, 157–177.
- LOEWY, S., CONNELLY, J.N. & DALZIEL, I.W.D. 2004. An orphaned block: The Arequipa–Antofalla Basement of central Andean margin of South America. *Geological Society of America Bulletin*, **116**, 171–187.
- LUCASSEN, F. & BECCIO, R. 2003. Timing of high-grade metamorphism: early Palaeozoic U–Pb formation ages of titanite indicate long-standing high-T conditions at the western margin of Gondwana (Argentina, 26–29°S). *Journal of Metamorphic Geology*, **21**, 649–662.
- LUDWIG, K.R. 2001. *SQUID 1.02, a User's Manual*. Berkeley Geochronology Center, Special Publication, 2.
- LUDWIG, K.R. 2003. *Isoplot/ExVersion 3.0, a Geochronological Toolkit for Microsoft Excel*. Berkeley Geochronology Center, Special Publication, 4.
- MARTINO, R.D., GUERESCHI, A.B. & SFRAGULLA, J.A. 2009. Petrography, structure and tectonic significance of the Tuclame banded schists in the Sierras Pampeanas de Córdoba and their relationship with the Argentinian North-western metamorphic basement. *Journal of South American Earth Sciences*, **27**, 280–298.
- MULCAHY, S.R., ROESKE, S.M., MCCLELLAND, W.C., NOMADE, S. & RENNE, P.R. 2007. Cambrian initiation of the Las Pirquitas thrust of the western Sierras Pampeanas, Argentina: implications for the tectonic evolution of the proto-Andean margin of South America. *Geology*, **35**, 443–446.
- MURRA, J., BALDO, E., GALINDO, C., CASQUET, C., PANKHURST, R., RAPELA, C. & DAHLQUIST, J. 2011. Sr, C and O isotope composition of marbles from the Sierra de Ancasti, Eastern Sierras Pampeanas, Argentina: age and constraints for the Neoproterozoic–Lower Paleozoic evolution of the proto-Gondwana margin. *Geologica Acta*, **9**, 1–23.
- OTAMENDI, J.E., TIBALDI, A.M., VUJOVICH, G.I. & VIÑAO, G.A. 2008. Metamorphic evolution of migmatites from the deep Famatinian arc crust exposed in Sierras Valle Fértil–La Huerta, San Juan, Argentina. *Journal of South American Earth Sciences*, **25**, 313–335.
- OTAMENDI, J.E., VUJOVICH, G.I., DE LA ROSA, J.D., TIBALDI, A.M., CASTRO, A., MARTINO, R.D. & PINOTTI, L.P. 2009. Geology and petrology of a deep crustal zone from the Famatinian paleo-arc, Sierras Valle Fértil–La Huerta, San Juan, Argentina. *Journal of South American Earth Sciences*, **27**, 258–279.
- OYHANTÇABAL, P., SIEGESMUND, S. & WEMMER, K. 2010. The Río de la Plata Craton: a review of units, boundaries, ages and isotopic signature. *International Journal of Earth Sciences*, doi:10.1007/s00531-010-0580-8.
- PANKHURST, R.J. & RAPELA, C.W. 1998. The Proto-Andean margin of Gondwana: an introduction. In: PANKHURST, R.J. & RAPELA, C.W. (eds) *The Proto-Andean Margin of Gondwana*. Geological Society, London, Special Publications, **142**, 1–9.
- PANKHURST, R., RAPELA, C., SAAVEDRA, J., BALDO, E., DAHLQUIST, J., PASCUA, I. & FANNING, C. 1998. The Famatinian magmatic arc in the central Sierras Pampeanas: an Early to Mid-Ordovician continental arc on the Gondwana margin. In: PANKHURST, R. & RAPELA, C. (eds) *The Proto-Andean Margin of Gondwana*. Geological Society, London, Special Publications, **142**, 343–397.
- PANKHURST, R., RAPELA, C. & FANNING, C. 2000. Age and origin of coeval TTG, I- and S-type granites in the Famatinian belt of NW Argentina. *Transactions of the Royal Society of Edinburgh: Earth Sciences*, **91**, 151–168.

- RAMOS, V.A. 1996. Evolución tectónica de la Plataforma Continental. In: RAMOS, V.A. & TURIC, M.A. (eds) *Geología y Recursos Naturales de la Plataforma Continental Argentina*. Asociación Geológica Argentina–Instituto Argentino del Petróleo, Buenos Aires, 385–404.
- RAMOS, V.A. 2008. The basement of the Central Andes: The Arequipa and related terranes. *Annual Review of Earth and Planetary Sciences*, **36**, 289–324.
- RAMOS, V.A. & VUJOVICH, G.I. 1993. The Pampia Craton within Western Gondwanaland. In: ORTEGA-GUTIÉRREZ, F., CONEY, P., CENTENO-GARCÍA, E. & GÓMEZ-CABALLERO, A. (eds) *Proceedings of the First Circum-Pacific and Circum-Atlantic Terrane Conference, Mexico*. Universidad Nacional Autónoma de México, Instituto de Geología, Guanajuato, México, 113–116.
- RAMOS, V.A., CRISTALLINI, E.O. & PÉREZ, D.J. 2002. The Pampean flat slab of the Central Andes. *Journal of South American Earth Sciences*, **15**, 59–78.
- RAMOS, V.A., VUJOVICH, G., MARTINO, R. & OTAMENDI, J. 2010. Pampia: a large cratonic block missing in the Rodinia supercontinent. *Journal of Geodynamics*, **50**, 243–255.
- RAPELA, C.W., PANKHURST, R.J., CASQUET, C., BALDO, E., SAAVEDRA, J., GALINDO, C. & FANNING, C.M. 1998. The Pampean Orogeny of the southern proto-Andes: evidence for Cambrian continental collision in the Sierras de Córdoba. In: PANKHURST, R.J. & RAPELA, C.W. (eds) *The Proto-Andean Margin of Gondwana*. Geological Society, London, Special Publications, **142**, 181–217.
- RAPELA, C.W., PANKHURST, R.J., BALDO, E., CASQUET, C., GALINDO, C., FANNING, C.M. & SAAVEDRA, J. 2001. Ordovician metamorphism in the Sierras Pampeanas: New U–Pb SHRIMP ages in Central–East Valle Fértil and the Velasco Batholith. In: *III Simposio Sudamericano de Geología Isotópica (III SSAGI), Pucón (CD-ROM)*, Article 616.
- RAPELA, C.W., BALDO, E.G., PANKHURST, R.J. & SAAVEDRA, J. 2002. Cordierite and leucogranite formation during emplacement of highly peraluminous magma: the El Pilón Granite Complex (Sierras Pampeanas, Argentina). *Journal of Petrology*, **43**, 1003–1028.
- RAPELA, C., PANKHURST, R., ET AL. 2007. The Río de la Plata craton and the assembly of SW Gondwana. *Earth-Science Reviews*, **83**, 49–82.
- RAPELA, C.W., PANKHURST, R.J., CASQUET, A., BALDO, E., GALINDO, C., FANNING, C.M. & DAHLQUIST, J.M. 2010. The Western Sierras Pampeanas: Protracted Grenville-age history (1330–1030 Ma) of intra-oceanic arcs, subduction–accretion at continental-edge and AMCG intraplate magmatism. *Journal of South American Earth Sciences*, **29**, 105–127.
- SCHWARTZ, J.J. & GROMET, L.P. 2004. Provenance of a late Proterozoic–early Cambrian basin, Sierras de Córdoba, Argentina. *Precambrian Research*, **129**, 1–21.
- SCHWARTZ, J.J., GROMET, L.P. & MIRÓ, R. 2008. Timing and duration of the calcalkaline arc of the Pampean Orogeny: implications for the Late Neoproterozoic to Cambrian evolution of Western Gondwana. *Journal of Geology*, **116**, 39–61.
- SIMS, J.P., IRELAND, T.R., ET AL. 1998. U–Pb, Th–Pb and Ar–Ar geochronology form the southern Sierras Pampeanas: implication for the Palaeozoic tectonic evolution of the western Gondwana margin. In: PANKHURST, R.J. & RAPELA, C.W. (eds) *The Proto-Andean Margin of Gondwana*. Geological Society, London, Special Publications, **142**, 259–281.
- TOSELLI, A.J., MILLER, H., ACEÑOLAZA, F.G., ROSSI, J.N. & SÖLLNER, F. 2007. The Sierra de Velasco (northwestern Argentina)—an example for polyphase magmatism at the margin of Gondwana. *Neues Jahrbuch für Geologie und Paläontologie, Abhandlungen*, **246**, 325–345.
- TRINDADE, R.I.F., D'AGRELLA-FILHO, M.S., EPOF, I. & BRITO NEVES, B.B. 2006. Paleomagnetism of Early Cambrian Itabaiana mafic dikes (NE Brazil) and the final assembly of Gondwana. *Earth and Planetary Science Letters*, **244**, 361–377.
- VARELA, R., SATO, A., BASEI, M.A.S. & SIGA, O., JR. 2004. Proterozoico medio y Paleozoico inferior de la Sierra de Umango, antepais andino (29°S), Argentina: edades U–Pb y caracterizaciones isotópicas. *Revista Geológica de Chile*, **30**, 265–284.
- VERDECCHIA, S.O. 2009. *Las metamorfitas de baja presión vinculadas al arco magmático famatiniano: las unidades metamórficas de la Quebrada de La Cébila y el borde oriental del Velasco, Provincia de La Rioja—Argentina*. PhD thesis, Universidad Nacional de Córdoba.
- VERDECCHIA, S.O. & BALDO, E.G. 2010. Geoquímica y procedencia de los metasedimentos ordovícicos del complejo metamórfico La Cébila, provincia de La Rioja, Argentina. *Revista Mexicana de Ciencias Geológicas*, **27**, 97–111.
- VERDECCHIA, S.O., BALDO, E.G., BENEDETTO, J.L. & BORGHI, P.A. 2007. The first shelly faunas from metamorphic rocks of the Sierras Pampeanas (La Cébila Formation, Sierra de Ambato, Argentina): age and paleogeographic implications. *Ameghiniana*, **44**, 493–498.
- VUJOVICH, G.I., VAN STAAL, C.R. & DAVIS, W. 2004. Age constraints and the tectonic evolution and provenance of the Pie de Palo Complex, Cuyania composite terrane, and the Famatinian orogeny in the Sierra de Pie de Palo, San Juan, Argentina. *Gondwana Research*, **7**, 1041–1056.
- WILLIAMS, I.S. 1998. U–Th–Pb geochronology by ion microprobe. In: MCKIBBEN, M.A., SHANKS, W.C., III & RIDLEY, W.I. (eds) *Applications of Microanalytical Techniques to Understanding Mineralizing Processes*. Reviews of Economic Geology, **7**, 1–35.
- ZIMMERMANN, U. 2005. Provenance studies of very low- to low-grade metasedimentary rocks of the Puncoviscana complex, northwest Argentina. In: VAUGHAN, A.P.M., LEAT, P.T. & PANKHURST, R.J. (eds) *Terrane Processes at the Margins of Gondwana*. Geological Society, London, Special Publications, **246**, 381–416.

Received 16 September 2010; revised typescript accepted 21 February 2011.

Scientific editing by Quentin Crowley.