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Magnetic properties of early Pliocene sediments from IODP Site U1467 (Maldives platform) reveal changes in the monsoon system

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Manuscript Details

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Title Magnetic properties of Early Pliocene sediments from IODP Site U1467 (Maldives platform) reveal changes in the monsoon system

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Abstract

We report a study of the magnetic stratigraphy and the anisotropy of isothermal remanent magnetization of Pliocene sediments from IODP Site U1467 drilled in the Maldives platform (Indian Ocean) during Exp. 359. Magnetic stratigraphy gives a precise record of geomagnetic reversals of the early Pliocene from approximately 5.3 Ma to 3.1 Ma providing a detailed age model in an interval, where the biostratigraphic record was scarce. Anisotropy of isothermal magnetization provides data on strength and direction of bottom current during the early Pliocene. The strength of bottom currents recorded by the anisotropy parameter P' , shows a prominent increase at about 4.2 Ma and the currents direction is consistent with that of modern instrumental measurements. Since bottom currents in the Maldives are driven by the monsoon, we speculate that the 4.2 Ma increase of bottom currents could mark the onset of the present-day setting, probably related to the coeval uplift phase of the Himalayan plateau.

Keywords Pliocene magnetic stratigraphy; Anisotropy of isothermal remanent magnetization; Currents strength; Monsoon

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Submission Files Included in this PDF

File Name [File Type]

1467-cover.docx [Cover Letter]

Response to reviewers.docx [Response to Reviewers]

1467-Revised ms with ma changes.docx [Revised Manuscript with Changes Marked]

1467-highlights.docx [Highlights]

Graphical abstract.PDF [Graphical Abstract]

1467-manuscript-revised.docx [Manuscript File]

Fig1.map.PDF [Figure]

Fig2-IRM acquisition.pdf [Figure]

Fig3-z_plot.pdf [Figure]

Fig4-Jelinek plot.pdf [Figure]

Fig5-VGP+dir mod.pdf [Figure]

Fig6-Age model + bio.pdf [Figure]

Fig7-AIRM stereo.pdf [Figure]

Fig8-AIRM_logs.pdf [Figure]

1467-tables.docx [Table]

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Research Data Related to this Submission

Data set

<https://data.mendeley.com/datasets/8383kp7vsb/draft?a=9687bfa8-8328-4ff7-9cc5-4538f2256ff0>

Data for: Magnetic properties of Early Pliocene sediments from IODP Site U1467 (Maldives platform) reveal changes in the monsoon system

Paleomagnetic and rock-magnetic data from IODP Site U1467 from "Magnetic properties of Early Pliocene sediments from IODP Site U1467 (Maldives platform) reveal changes in the monsoon system" by L. Lanci, E. Zanella and Exp. 359 members

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July 10, 2019

Dr. Thierry Correge
Editor of Palaeogeography Palaeoclimatology Palaeoecology

Dear Sir,

I wish to submit the revised version of the manuscript titled "*Magnetic properties of Early Pliocene sediments from IODP Site U1467 (Maldives platform) reveal changes in the monsoon system*" by L. Lanci, et al, to be considered for publication in Palaeogeography Palaeoclimatology Palaeoecology.

I believe we have answered properly to all reviewers' questions and suggestions, which were very useful, as detailed in the file "response to reviewers". Major changes include an amended interpretation of the paleo-currents directions and redrawn figures, that we hope became more attractive. A new co-author has been added, who contributed to the paleoceanographic interpretation.

We hope that the manuscript could be of interest for Paleo3.

Sincerely,

Luca Lanci.

Magnetic properties of early Pliocene sediments from IODP Site U1467 (Maldives platform) reveal changes in the monsoon system, by Lanci et al.

Response to reviewers

Reviewer #1

Q:

Depths throughout the MS:

Since 2011 the depth of IODP cores and samples is expressed as CSF-A, CSF-B, and CCSF (not mbsf or mcd). Please check the guidelines:

<https://www.iodp.org/policies-and-guidelines/142-iodp-depth-scales-terminology-april-2011/file>.

In this regards, the data given in the online Table (U1467_data.xlsx) are not correct and therefore not acceptable (only for practical reasons: every specimen should be traceable). The extended sample code should be in the form of (e.g.) 359U1467B12H1W060/062. This information is given (almost in the complete form) in the NRM direction table BUT NOT in the anisotropy table, where the Hole is not indicated.

The list of samples MUST be associated with the CSF-A value (former mbsf), and this value is missing in both tables.

A: We have changed all depth to meter CSF-A as requested by the reviewers. This apply to figures, tables and to the Excel data sheet.

Q: *Furthermore, there is a major issue regarding the depth in the figures. The magnetostratigraphy presented in figure 6 does not fit with the depth shown in figure 8 (anisotropy data). For example the deepest normal magnetic polarity interval N6 is at about 260–280 m in Fig. 6 but much shallower in Fig. 8, and the scale is the same (mbsf). Even if this should not affect the age of bottom current inception as interpreted from the AIRM data, it must be reviewed.*

A: We acknowledge that there was a mistake in pasting the black and white column into the figure of anisotropy data (former Fig. 6). This mistake have been corrected. Moreover, the new figure (now Fig. 8) has been redrawn in a (hopefully) more attractive format, a smoothed line that shows the data trends was added as suggested by reviewer #1. Moreover, to increase the information shown in the figure we also introduced a color code for the anisotropy shape parameter (T).

Q:

Anisotropy of IRM:

The model put forward in this MS is based on the AIRM and its mathematical expression P' . The background to obtain P' is limited to the laboratory procedure (actually not very well explained, see comment on PDF), and it is

called out in line 313 simply as “anisotropy parameter” without any reference or explanation.

I presume it is the ‘corrected anisotropy degree’ of Jelínek (1981), usually applied to AMS, but I would like to see the equation in the MS, since there is often confusion in the literature about the names of anisotropy parameters. There is no other information regarding the shape of the AIRM tensor, neither related to each sample nor to the mean one. The shape could be easily quantified by the “T” factor of Jelínek (1981)(or maybe, even better with the bootstrap approach of Constable and Tauxe (1990)?). The statistic approach used in the last figure is not indicated, and there is a legend without unit.

A: We have amended the definition of the parameter used in the description of anisotropy. The formula for the corrected anisotropy degree (P') and the shape parameter (T) were provided together with the reference.

A new figure (Fig. 4) with a so-called Jelinek plot, to illustrate the shape of the anisotropy ellipsoids, was added to the manuscript as suggest by reviewer #1 in the annotated ms.

Q: *Furthermore, even if P' is much lower in the “pre-monsoon inception” part of the section, it is important to show a stereographic projection of the eigenvectors, for comparison with the upper part. The distribution of the minimum eigenvector could be elongated NE-SW also before the inception of the monsoon-driven bottom current regime.*

What about also defining a P' minimum value under which the anisotropy is statistically meaningless? It could strengthen the interpretation.

A: We show the stereographic projection of the low P' data as requested in the new Figure 7. Following suggestions we have defined a P' value below which the anisotropy do not have statistically significant orientations and divided the specimens in two sets according to this value. The paleo-currents interpretation was improved by evaluating the AIRM pattern of each sample to infer the current directions that are reported in a circular plot for both sets of samples with high P' and low P' . A statistical test have been applied to ensure that the set with high P' had grouped direction and that the set with lower P' had uniformly distributed directions. Mean directions were calculated only for the set with high P' using Von Mises statistics. Results are shown in the new figure 7.

Other changes following suggestions on the annotated ms.:

As requested in the annotated ms., in section “4.2. Anisotropy of the IRM” we have expanded the paleoenographic explanation of possible monsoon effect on the bottom currents. Admittedly the arguments was very briefly explained even in the cited literature and certainly deserved a better explanation.

Online Table (U1467_data.xlsx):

The table has been reviewed as suggested, in particular sample code for anisotropy were corrected, sample depth was reported in m CSF-A and MDF calculation were added to the table. The complete set of data is available in the excel file.

Figure changes

Figure 2:

Figure 2 was redrawn reporting the specimens code and adding a second panel with the distribution of the median destructive field of the NRM, as suggest by reviewer #1 in the annotated ms.

Figure 3:

We added the labels with demagnetizing fields as requested by reviewer #1 in the annotated ms.

New figure 4:

A new figure (Fig. 4) with a so-called Jelinek plot, to illustrate the shape of the anisotropy ellipsoids, was added to the manuscript as suggest by reviewer #1 in the annotated ms.

New Figure 5:

Figure 5 (former Figure 4) was redrawn reporting the core photographs and the lithostratigraphic units as requested by reviewer #1 in the annotated ms. Moreover, following the suggestions on the annotated ms, were also reported the positions of all measured specimens in order to get a more precise idea of the success rate for magnetostatigraphy.

New figure 7:

Figure 7 was completely redrawn as described above.

New figure 8 (former figure 6):

The new figure 8 have has been redrawn as described above.

Reviewer #2

Q: First of all, given that the journal has a rather wide scope in paleoclimatology, paleogeography, the authors need to explain a bit better the basis and the parameters that are used in magnetic anisotropy. For example, in line 289 they referred to I_{max} and I_{min} , but do not explain what they mean- They do not represent “preferred orientations” (line 289), but the inclination of the maximum and minimum axes of the AIRM matrix.

A: We have changed I_{max} and I_{min} to I_1 and I_3 , which is a “standard” notation for principal axis of AIRM, and defined them as eigenvectors of the IRM anisotropy tensor. We also gave a description of their geological interpretation.

Q: More confusing is the interpretation of Figure 7 (a key part of the paper!): Where is mean direction of 134/00 seen in the stereonet? What do they mean by “elongation axis I_{max}”? (the scatter / grouping of the individual axes?). Also, the same figure contains two types of planes (in blue and red)- what do they mean? And the colored bar on the side?

A: We have amended the interpretation of the paleocurrent data and Figure 7 has been completely redrawn following also suggestions of reviewer #1. We hope the new figure and interpretation answer these questions.

Q: How can they assert that a “the averaged direction of foliation planes” is 137 N? What are the white dots represented? (I assume the axes if minimum ARM, but both the legend and the text fail to explain so). If anything, I can see the NE-SW trend, which would correspond to an azimuth of about 40 deg. So, in general, the paragraph from lines 315 to 319 is rather confusing. The authors do a great job describing possible scenario for the AIRM axes distribution (imbrications, flow-transverse fabrics, flow-aligned fabric, etc.) yet they fail in stating in a clear way which case is what they find and WHY they think so.

So, notice that my main comment is not merely about complementing the legend of such figure, but explaining how the authors go from the basis of the method (lines 294 to 305 to the interpretation of the observed fabric (lines 315 to 322). This part is critical to the paper.

A: As described above, in order to answer this comment and a similar request from reviewer #1, the AIRM interpretation have been improved and the results illustrated in new Figure 7.

Q: Last, but not least, there is a change (stratigraphically upwards) in magnetite concentration, as shown by the increase of the IRM intensity. Such change is accompanied by a slight increase of the anisotropy degree (possible proxy for the strength of the bottom current). The authors interpret such change in anisotropy as the initiation of the monsoon. If that is the case, wouldn't you expect an increase of dust and terrigenous flux from the continent as well? An increase on terrigenous influx, as one would expect, should be conducive to an increase of magnetite as well. So such seemingly contradiction needs to be explained.

A: This question was already explained in the first submission, we have slightly changed it as follow: “From the sedimentological point of view, the decrease of IRM is interpreted as a consequence of changes in the sediment transport mechanism -controlled by wind driven currents- that transferred the sediments and the single-domain magnetite, possibly of biogenic origin, from the shallow platform to the deeper water of Site U1467 (Lüdmann et al., 2013). This process is accelerated by the increased monsoon strength starting at 168 m CSF-A depth.”

1 Magnetic properties of ~~early~~**Early** Pliocene sediments from IODP Site
2 U1467 (Maldives platform) reveal changes in the monsoon system
3

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- 67

68 **Abstract**

69 We report a study of the magnetic stratigraphy and the anisotropy of isothermal
70 remanent magnetization of Pliocene sediments from [International Ocean Discovery Program](#)
71 [\(IODP\) Site U1467](#) drilled in the Maldives platform (Indian Ocean) during Exp. 359. Magnetic
72 stratigraphy gives a precise record of geomagnetic reversals of the early Pliocene from
73 approximately 5.3 Ma to 3.1 Ma providing a detailed age model in an interval, where the
74 biostratigraphic record ~~is~~ scarce. ~~We use the anisotropy~~ [Anisotropy](#) of isothermal
75 [remanent magnetization \(AIRM\) to investigate the statistical orientation of fine magnetic](#)
76 [particles and provide](#) ~~provides~~ data on [the](#) strength and direction of bottom ~~current~~ [current](#)
77 during the early Pliocene. The strength of bottom currents recorded by the ~~AIRM~~ [anisotropy](#)
78 ~~parameter P'~~, shows a prominent increase at [the top of Chron C3n.1n](#) (about 4.2 Ma), and the
79 ~~current~~ [currents](#) direction ([NE - SW](#)) is consistent with that of modern instrumental
80 measurements. Since bottom currents in the Maldives are driven by the monsoon, we
81 speculate that the 4.2 Ma increase of bottom currents could mark the onset of the present-day
82 setting, probably related to the coeval uplift phase of the Himalayan plateau.

83
84 **Keywords:** Paleomagnetism; Pliocene magnetic stratigraphy; Anisotropy of isothermal
85 remanent magnetization; Currents strength; Monsoon

86
87 **1. Introduction**

88 Wind-induced currents are an important factor controlling the sedimentation in
89 ~~the~~ [The](#) Maldives archipelago where they regulate [sediment transport](#) ~~the sediments~~
90 ~~transportation~~ from the atoll to the deeper part of the platform as well as the geometry of the
91 sedimentary bodies (e.g., Betzler et al., 2009; Lüdmann et al., 2013; Betzler et al., [2016b](#)~~2016~~).
92 The onset of these current-driven drift deposits has been related to the increase of monsoon
93 activity starting in the middle Miocene (Betzler et al., [2016b](#)~~2016~~). In this paper we report
94 results from paleo and rock-magnetic analyses from [International Ocean Discovery Program](#)
95 [\(IODP\) Site U1467](#) aiming to improve [the](#) Pliocene age model and investigate the variations of
96 bottom current strength.

97 Site U1467 ($4^{\circ} 51.0155' 0155^{\circ}\text{N}$ and $73^{\circ} 17.0204' 0204^{\circ}\text{E}$) was drilled in the Inner Sea
98 of ~~the~~ [The](#) Maldives ~~archipelago~~ [atoll](#) (Indian Ocean) during IODP Expedition 359 at a water
99 depth of 487.4 m (Betzler et al., 2017) in a distal position compared to the platform margins
100 and moats (Fig. 1). Site U1467 showed nearly horizontally layered seismic reflections of the
101 so-called drift sequences (Betzler et al., 2013, Lüdmann et al., 2013; Wunsch et al., 2017) with

102 no truncations and no indications of mass wasting from the adjacent platform margin.
103 Shipboard analysis suggested that Site U1467 provided a complete and undisturbed
104 succession of all drift sequences from the Late Miocene (Betzler et al., 2017) with good
105 potential for paleoceanographic and paleoclimatic studies. In particular the drift succession,
106 from mid-Miocene to recent, contains several sequences that are potentially related to
107 fluctuations in the monsoon-driven current system. Dating these sequences can yield the ages
108 of changes in the strength and direction of the currents.

109 ~~The paleomagnetic~~ Paleomagnetic analysis ~~in~~ of this study ~~provides~~ provide the
110 magnetostratigraphic age of the upper portion of Site U1467 that was sampled with advanced
111 piston coring (APC), and uses the anisotropy of the isothermal remanent magnetization
112 (AIRM) to investigate the statistical orientation of fine magnetic particles aiming to provide a
113 sedimentological record of direction and strength of bottom currents. ~~The drift deposition~~
114 ~~has been related to an increased monsoon strength since the middle Miocene (Betzler et al.,~~
115 ~~2016) and, although there is not a direct effect of winds on currents below the thermocline~~
116 ~~depth, there are ample sedimentological evidence that sea-bottom currents in the Maldives~~
117 ~~are driven by the monsoon (e.g., Betzler et al., 2009, Lüdmann et al., 2013, Betzler et al.,~~
118 ~~2016).~~

119 Shipboard paleomagnetic measurements of Site U1467 (Betzler et al., 2016, Betzler et
120 al., 2017) gave poor results because of a combination of two factors: (i) the very low
121 concentration of magnetic minerals in carbonate platform sediments, which resulted in a very
122 weak natural remanent magnetization (NRM) and (ii) a strong magnetic contamination of the
123 cores due to metallic particles presumably originating from the drilling pipes. The
124 contamination covered the original weak paleomagnetic signal of the sediment preventing
125 any valuable measurements with the shipboard pass-through technique that measures half-
126 cores. Measurements of individual single 7 cm³ box-samples, taken from the inner part of the
127 core, performed aboard, did not show signs of a-significant contamination suggesting that it
128 was restricted to the outer part of the cores. Unfortunately, the NRM natural remanent
129 magnetization (NRM) of these specimens, ranging from ca. 1 x 10⁻⁵ A/m to 1 x 10⁻⁴ A/m, was
130 too weak to be measured reliably using the JR-6 spinner magnetometer available on the Joides
131 Resolution. However, the NRM of the supposedly uncontaminated box-samples is within the
132 range of sensitivity of a DC-SQUIDS cryogenic magnetometer, and this provided the
133 motivation to collect and measure 580 standard sediment specimens with the aim to obtain a
134 reliable magnetic stratigraphy of Site U1467.

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139 2. Material and sampling

140 ~~Standard paleomagnetic~~Paleomagnetic standard specimens (~~Natsuhara-Giken~~
141 ~~sampling~~2x2 plastic cubes, with a volume of 7 cm³) were collected in the upper part of Site
142 U1467 from core sections 359-U1467B-11H to 359-U1467B-~~34H~~33H and from 359-U1467C-
143 10H to 359-U1467C-17H, corresponding to ~~84 m to 302 m core~~ depth ~~from 84 meter~~
144 ~~composite depth (med)~~ below sea floor (~~CSF-A~~), ~~to 330 med~~, at the Gulf Coast Repository
145 (~~GCR~~) at Texas A&M University. Specimens were sampled only from azimuthally-oriented
146 ~~APC~~advanced piston cores, (~~APC~~), since information on core orientation is essential for
147 paleomagnetic studies at equatorial ~~latitudes~~latitude such ~~as for that of~~ Site U1467.

148 According to the shipboard sedimentology, the studied part of Site U1467 was divided
149 into three main lithostratigraphic units (Betzler et al., 2017). The uppermost Unit I was
150 recovered in the top 110 ~~m CSF-A~~med and consists of unlithified, foraminifer-rich wackestone
151 to packstone with a predominance of very fine- to fine-grained wackestone. Unit II extends
152 from ca. 110 ~~m med~~ to 215 ~~m CSF-A~~med; it comprises ~~the~~late Pliocene ~~sediments~~ and is
153 characterized by interlayered unlithified and partially lithified planktonic
154 ~~foraminifera~~foraminifer-rich wackestone and mudstone with pteropods and particulate
155 organic matter. Unit III, which extends from 215 ~~m med~~ to 303 ~~m CSF-A~~med, consists of
156 partially lithified very fine-grained mudstone to wackestone with a dominance of
157 wackestone. The sediment contains abundant planktonic ~~foraminifera~~foraminifers; echinoid
158 spines and sponge spicules are common while benthic ~~foraminifera~~foraminifers are rare.
159 ~~Microfossil~~Microfossils preservation throughout Units III and most of Unit II was generally
160 poor to moderate, ~~and~~ in particular the interval from ca. 150 ~~m med~~ to 300 ~~m med~~ yields no
161 biostratigraphic ages. The interval chosen for paleomagnetic study was also intended to
162 ~~cover~~covers this interval with poor or absent ~~biostratigraphy~~biostratigraphic record.

163

164

165

166 3. Paleomagnetic analysis

167 3.1. Isothermal remanent magnetization

168 Magnetic mineralogy was investigated by acquisition of isothermal remanent
169 magnetization (~~IRM~~) in a set of pilot specimens. ~~IRM was acquired in 12 stepwise increasing~~

170 fields from 0.03 T to 1 T, induced using a ASC pulse magnetizer (Fig. 2a). Results indicate that
171 all measured specimens are characterized by the ~~only~~ presence of only low-coercivity
172 magnetic ~~minerals~~ mineral that saturate in fields between 100 mT and 250 mT. These
173 coercivities are below the maximum coercivity of uniaxial magnetite (e.g. Tauxe , 2002) and
174 suggest (Fig. 2). ~~This suggests~~ a rather homogeneous mineralogy made of ferromagnetic
175 minerals such as magnetite (or maghemite) without any significant presence of diagenetic
176 iron ~~sulphidessulphates~~ that can be distinguished from their higher coercivity (e.g., Tauxe,
177 2002).

178 The saturated isothermal remanent magnetization is relatively weak, as often found in
179 carbonate sediments, because of the low concentration of ferrimagnetic minerals. It ranges~~its~~
180 variations range from 1.0×10^{-3} to 3.7×10^{-2} A/m, suggesting a wide~~comparable~~ variability in
181 the concentration of magnetic minerals.

182

183 *3.2. Natural remanent magnetization*

184 The pass-through shipboard measurements of the NRM showed a huge scope of values
185 with ~~NRM~~-intensity ranging from 5×10^{-6} A/m to 1×10^{-1} A/m along the same core and with a
186 strong downcore decreasing trend (Betzler et al., 2017). This was interpreted as the
187 consequence of steel contamination most likely originating from worn off drill pipes. The
188 bottom part of each APC core, which were the least contaminated, exhibited NRM values
189 ranging from ca. 5×10^{-6} A/m to 1×10^{-4} A/m that were considered reasonable values for
190 carbonate sediments and thus regarded as uncontaminated or only slightly contaminated.
191 Discrete specimens taken from the inner part of the cores showed similar NRM intensity
192 values corroborating the hypothesis that contamination was limited to the outer part of the
193 cores.

194 Based on these remarks, box specimens for the shore-based analysis were collected
195 from the inner part of the core in Site U1467 and were measured using a 2G-enterprise DC-
196 SQUID magnetometers at the CIMaN-ALP laboratory (Cuneo, Italy). Samples were
197 progressively demagnetized in alternating field (AF) up to the maximum field of 100 mT
198 according to a standard paleomagnetic procedure.

199 The directional components of the natural magnetization were calculated using the
200 method of the principal component analysis (Kirschvink, 1980) and the PuffinPlot software
201 (Lurcock and Wilson, 2012). The quality of the measurements and the line fitting was
202 checked by visual inspection of the orthogonal vector plots and was quantified using the
203 maximum angular deviation (MAD).

204 The AF demagnetization technique was effective in demagnetizing the NRM testifying
205 that low-coercivity minerals are the main carriers of the NRM. Vector plots generally show a
206 small viscous overprint removed in fields smaller than 20 mT, while the remaining part of the
207 NRM is demagnetized in the field interval from 20 mT to 100 mT. The component isolated
208 within this coercivity interval was used to calculate the characteristic remanent
209 magnetization (ChRM) when sufficiently linear and well defined. However, the success rate in
210 recovering reliable paleomagnetic directions was generally low and many specimens were
211 discarded because they did not yield ~~to~~ results with acceptable quality. In specimens that gave
212 acceptable results, on average, about 20% of the NRM was removed after AF demagnetization
213 at 20 mT, and ca. 95% of the NRM was removed at field of 60 mT. Moreover, the NRM median
214 destructive field of acceptable specimens (Fig. 2b) has a modal value of 10 mT. The
215 percentage of NRM removed at 60 mT and the values of median destructive field corroborate
216 the results of IRM acquisition suggesting that stable NRM is carried by pseudo-single domain
217 magnetite. In acceptable specimens~~In these specimens~~, the NRM intensity has an average
218 value of~~averaged~~ ca. 2.1×10^{-4} A/m, the average value of MAD obtained from the vector
219 analysis of these specimens was 8.9°. About 10% of specimens have MAD values between 15°
220 and 20°, which although large, ~~are have been~~ considered acceptable ~~anyway~~; most of these
221 specimens are located in the upper part of the investigated interval, between 100 and 180 m
222 depth, or at polarity transitions.~~150 med.~~ Representative orthogonal vector plots for Site
223 U1467 are illustrated in Figure 3.

224 The azimuthal orientation of cores was essential for interpretation of magnetic polarity
225 because the paleomagnetic inclinations of equatorial localities, such as Site U1467 during the
226 Miocene, are very close to zero for both normal and ~~reversed~~~~reversal~~ polarities; hence the
227 geomagnetic polarities are indistinguishable if based only on inclination data. APC cores
228 collected from Site U1467 were oriented using the “tensor tool” that provided a good first-
229 order orientation. The ~~average~~~~averaged~~ declinations from paleomagnetic measurements
230 showed ~~indeed~~ significant ~~departures~~~~departure~~ from ~~the~~ North and discrepancies between
231 cores, suggesting that orientation errors of the tensor tool can be as large as $\pm 30^\circ$. However,
232 although large, these errors did not compromise the polarity of ChRM and there ~~was~~~~were~~ no
233 ambiguity in establishing the magnetic polarity. We did not attempt to remove orientation
234 errors by adjusting the magnetic declination to a mean direction even if this resulted in a
235 reduced precision of the latitude of the virtual geomagnetic pole (VGP).

236

237 *3.3. Anisotropy of isothermal remanent magnetization*

238 In agreement with shipboard measurements, the magnetic susceptibility of box-
 239 samples ~~shows~~ ~~has shown~~ negative (diamagnetic) susceptibility, evidence of the dominating
 240 diamagnetic matrix of CaCO₃ on the ferrimagnetic component. The very weak ~~and negative~~
 241 ~~(diamagnetic)~~ ~~magnetic~~ susceptibility of the carbonate sediments recovered in Site U1467
 242 (Betzler et al., 2017) ~~limit~~ ~~compromises~~ the possibility ~~of using to use~~ the anisotropy of
 243 magnetic susceptibility to investigate the orientation pattern of the magnetic
 244 particles. Therefore, we ~~resorted to using resourced to the anisotropy of isothermal remanent~~
 245 ~~magnetization (AIRM which) that~~ can be measured precisely even in these weakly magnetic
 246 sediments.

247 AIRM measurements were performed in a subset of 75 specimens taken from Core 13
 248 to Core 26 ~~of Hole U1467A.~~ To compute the AIRM an isothermal remanent magnetization
 249 induced with a field of 20 mT was measured and then AF demagnetized, repeating this
 250 procedure along 6 different ~~axes. Each axis was measured twice along opposite directions for~~
 251 ~~a total of 12 AIRM measurements in each specimen directions~~ (e.g., Stephenson et al., 1986;
 252 Jackson, 1991; Potter, 2004). ~~for a total of 12 AIRM measurements in each specimen.~~ The
 253 intensity of isothermal magnetization was measured with a JR-6 spinner magnetometer and
 254 the specimens were demagnetized after each measurement using a tumbling 2G AF-
 255 demagnetizer at a maximum field of 80 mT, before inducing the magnetization in the next
 256 direction. The anisotropy tensor ~~and~~ the ~~directions of the principal IRM axis I_i magnetic~~
 257 ~~lineation~~ (i.e., ~~the eigenvector~~ ~~the direction of the main eigenvector I_{max}~~ ~~of the anisotropy~~
 258 ~~tensor~~) ~~and the foliation plane (i.e., the plane orthogonal to smaller eigenvector I_{min}~~ of the
 259 ~~AIRM anisotropy~~ tensor) were computed from the remanent magnetization using
 260 the AGICO software Anisoft42. ~~The AIRM is therefore represented as a triaxial ellipsoid,~~
 261 ~~whose principal axes correspond to the directions of maximum, intermediate and minimum~~
 262 ~~IRM (I₁ < I₂ < I₃).~~ The anisotropy ellipsoid was described using the corrected anisotropy
 263 ~~degree (P')~~ and the shape parameter ~~(T)~~ computed according to Jelinek (1981)

$$T = (2n_1 - n_2 - n_3) / (n_1 - n_3)$$

$$P' = \sqrt{\exp 2 (a_1^2 + a_2^2 + a_3^2)}$$

267 where $n_i = \ln I_i$, $a_i = \ln \left(\frac{I_i}{I_m} \right)$ and $I_m' = \sqrt[3]{I_1 I_2 I_3}$ with $i = 1, 2, 3$ and are shown in Figure 4.

268 ~~The direction of the largest axis of the anisotropy tensor I₁ represents the magnetic lineation~~
 269 ~~(the preferred orientation of elongated magnetic particles), and the foliation plane is the~~

270 plane that contains the I_1 and I_2 directions, hence is orthogonal to the direction of the smallest
271 axis of the anisotropy tensor I_3 .

272

273 4. Results and discussion

274 4.1. Magnetostratigraphy

275 Declination, inclinations of ChRM and the resulting VGP latitude of Site U1467 are
276 shown in Figure 54 plotted versus depth m CSF-Amed, together with the available
277 biostratigraphic events from shipboard analysis (Betzler et al., 2017). The comparison of
278 measured levels ~~Data quality and success rate are best below 150 med and it is only~~
279 acceptable samples in the upper 50 m of the measured interval. The density of data points in
280 Figure 5 ~~indicates~~ is indicative of the low success rate ~~obtained~~ in finding reliable directions
281 of the ChRM, ~~which was generally low~~. Sometimes, ~~as~~ for instance in specimens taken from
282 Hole U1467C, we could not ~~obtain~~ get any acceptable ~~results~~ result. In general in the upper
283 part of the Site above 110 med, which ~~comprises~~ is comprised in the sedimentological Unit I,
284 the paleomagnetic data yielded poor results probably related to the coarser granulometry of
285 the unlithified wackestone sediments. At ~~depths~~ depth ~~comprised~~ between 100 m and 290
286 med, the quality of the data was ~~sufficient~~ sufficiently good to obtain a reliable record of
287 polarity reversal, although with ~~a~~ variable quality.

288 In the interval with good data quality (i.e., the central part of the record with smaller
289 MAD), the ChRM inclinations are practically indistinguishable from zero, regardless of the
290 ~~actual~~ polarity, except for transitional directions corresponding to the time elapsed during the
291 reversal of the geomagnetic field. Averaged paleomagnetic directions (Table 1) ~~indicate an~~ 2)
292 ~~indicates a nearly~~ equatorial paleolatitude $\lambda = 0.8^\circ$ ($\lambda_{95}^+ = 4.4^\circ$; $\lambda_{95}^- = -2.8^\circ$) of Site U1467
293 during the ~~early Lower~~ Pliocene within the precision of the paleomagnetic data, ~~(colatitude~~
294 ~~error $dp_{95} = 3.6^\circ$),~~ which is in ~~agreement~~ good accordance with the paleogeographic
295 reconstructions of Besse and Courtillot (2002) and Torsvik et al. (2012).

296 The record of polarity reversals ~~identifies~~ record has identified 6 normal and 6
297 reversed magnetic polarity zones that, based on the biostratigraphic framework, have been
298 interpreted as Chrons C2An.2n to C3n.4r (Fig. 65) (Gradstein et al., 2012). ~~This~~ It has to be
299 ~~noticed that this~~ interpretation is mainly based on the ~~only~~ 4 biostratigraphic events available
300 in the studied section and located in the upper part of the record; ~~however~~. ~~Moreover~~, there
301 is some uncertainty ~~uncertainties~~ even in the biostratigraphic data since the dates based on
302 Last Occurrence ~~lower appearance~~ of foraminifera (*Dentoglobigerina altispira* and
303 *Globorotalia margaritae*) show a relatively large discrepancy with that of calcareous

304 ~~nannofossil~~~~nannoplankton~~ events (LO *Sphenolithus abies* and LO *Reticulofenestra*
305 *pseudoumbilicus*). ~~Even though the~~The few available biostratigraphic markers ~~available and~~
306 ~~their uncertainty~~ leave some room in interpreting which magnetochrons correspond to the
307 measured polarity reversals, ~~we~~. We believe that our interpretation (Fig. 6)~~shown in Figure 5~~
308 is the best compromise between a reduced variability of the sedimentation rate and the
309 available biostratigraphic framework, ~~nonetheless, we realize that this interpretation could~~
310 ~~change if major modification were made to the biostratigraphy data.~~ Within these limitations,
311 the magnetic stratigraphy of Site U1467 ~~provides~~provide a robust age ~~constraint~~constrains in
312 a section ~~where~~were biostratigraphic ~~records are~~record is unavailable.

313 In our interpretation the age of the studied section spans ca. from 5.3 Ma ~~Myr~~ ago to
314 3.1 Ma ~~Myr~~ ago, as reported in ~~detail~~details in Table 21.

315

316 4.2. Anisotropy of the IRM

317 According to a number of studies (e.g., ~~Lüdmann et al., 2013~~; Betzler et al., 2009,
318 ~~2016b~~2016, 2018 and ~~references~~reference therein) bottom currents in ~~the~~The Maldives
319 platform are considered wind-driven and ~~assumed to be are considered~~ a direct consequence
320 of -Asian monsoon. At equatorial latitudes, the link between surface wind and bottom currents
321 extends to a depth of several hundred meters either through Ekman transport or as an
322 undercurrent system and can be seen with modern observations. The present day equatorial
323 Indian Ocean is characterized by seasonally reversing surface currents, known as Wyrтки Jets,
324 driven by zonal winds. Beneath the surface, to a depth of several hundred meters, the flow of
325 the equatorial undercurrent and the equatorial intermediate current has been observed (e.g.,
326 Knox, 1976; Reppin et al., 1999; Schott and McCreary, 2001; Iskandar et al., 2009; Nyadjro et
327 al., 2015). In contrast to other oceans, the Indian Ocean equatorial undercurrent is transient
328 and strongly dependent on winds and pressure gradient variations. Both eastward and
329 westward flows of sub-surface currents have been observed, although modeling studies based
330 on the present day suggest that eastward undercurrents are more likely to occur than
331 westward ones (Schott and McCreary, 2001). Since sub-surface currents develop as
332 consequences of surface wind it is reasonable to assume that stronger surface winds will
333 increase the strength of the undercurrent, and following~~Following~~ this argumentline, we
334 interpret bottom current ~~paleo-direction and~~ strength as proxy of the paleo-monsoon.

335 Magnetic methods are ~~In this case the magnetic method is~~ particularly useful
336 ~~when failing~~ other quick methods for determining paleo flow~~paleoflow~~ from sediment beds
337 such as macroscopic paleocurrent indicators (e.g., cross-stratification and sole marks) are

338 ~~lacking.}~~ In standard analysis of magnetic grain shape fabric, AIRM is considered to be a
339 proxy for the preferred alignment of elongated natural magnetic particles attained in the final
340 stages of transport, with I_1I_{max} and I_3I_{min} representing preferred orientations of the longest
341 and shortest grain axes, respectively (e.g., Hamilton and Rees, 1970, Taira and Scholle, 1979;
342 Novak, 2014, Felletti 2016). The method assumes implicitly that the uniaxial shape-
343 anisotropy of magnetic particles dominates triaxial magnetocrystalline anisotropy, as
344 expected which is a very reasonable assumption for elongated Site U1467 since magnetite
345 particles (e.g., Tauxe, 2002) is the main magnetic mineral.

346 According to theoretical, experimental and field-based fabric studies, two main
347 anisotropic fabric patterns are found (e.g., Harms et al., 1982; Baas et al., 2007): (i) flow-
348 aligned fabric; and (ii) flow-transverse fabric. In flow-aligned fabric the I_1I_{max} axes are
349 oriented parallel to the mean flow direction, while in a flow-transverse fabric, the I_1I_{max} axes
350 are oriented perpendicular to the flow direction. In turbulent flows, grains settling from
351 suspension tend to orient with their I_1I_{max} axes parallel to the flow direction and imbricated
352 upstream (Rusnak, 1957; Allen, 1984). This flow-aligned orientation can be changed into a
353 more stable flow-transverse orientation when the flow becomes strong enough to lift grains
354 and roll them over the surface (e.g. Schwarzacher, 1963; Johansson, 1964; Hendry, 1976,
355 Harms et al., 1982). In both cases the *foliation* planes (i.e., the planes perpendicular to the I_{min}
356 axes) can be imbricated dipping upstream (Harms et al., 1982) and the comparison of their
357 orientation with I_1I_{max} axes can be used to recognize the flow-aligned and flow-transverse
358 fabrics.

359 Deviations from the flow-aligned or the flow-transverse fabrics can occur for a number
360 of reasons (e.g., Baas et al., 2007 and references therein which include) among which spatial
361 changes in current direction, bed surface irregularities, incomplete reorientation of a rolling
362 fabric into a flow-aligned fabric or vice versa, changes in bed roughness and post-depositional
363 modification by bioturbation or soft-sediment deformation. (e.g., Baas et al., 2007 and
364 references therein).

365 We recognise the pattern of each specimenIn Site U1467, we found that the AIRM is
366 significantly large to produce a coherent pattern of orientations only in the upper part on the
367 analysed interval (Fig. 6). The degree of anisotropy (P' parameter) shows a sudden increase
368 from very low values (mean 1.055 ± 0.03) to larger values (mean 1.22 ± 0.17) in the upper ca.
369 168 ± 2 mcd of Site U1467 corresponding to the age of about 4.2 Myr ago in our age model. In
370 the upper part of Site U1467 the directions of the elongation axis I_{max} have a mean direction of
371 $134/00$. Foliation planes are imbricated toward NE and SW (Fig. 7) by comparing the an

372 average angle (θ) between the of about 34° ($a_{95} = 9.4$). The averaged direction of the magnetic
373 lineation I_1 and that of the foliation plunge. If $\theta < 35^\circ$ the pattern is flow-aligned planes of 137°
374 N (angular dispersion 11°) and the flow is taken equal to declination of the I_1 elongation axis
375 in the direction indicate a transverse pattern of the foliation imbrication; if $\theta \geq 55^\circ$ the pattern
376 is flow-transverse and the flow is the declination of $I_1 - 90^\circ$ in the AIRM with a mean flow
377 direction of the foliation imbrication. The intermediate case ($35^\circ < \theta \leq 55^\circ$) is handled by
378 taking directly the imbrication orthogonal to both axis. According to this observation the
379 direction of the foliation plane as the flow direction.

380 In Site U1467, we found that the AIRM is large enough to produce a well-defined
381 pattern of orientations only if the degree of anisotropy $P' \geq 1.1$, which mostly comprises
382 specimens with flow-transverse pattern and located in the upper part on the sediment
383 column. Current directions, foliation planes and I_1 directions are shown in bottom Figure 7 in
384 separated sets for $P' \geq 1.1$ and $P' < 1.1$. In the set with $P' \geq 1.1$, the current directions fall into
385 two distinct groups with nearly opposite modal directions highlighted by the rose diagram
386 (Fig. 7c). Foliation planes also have the opposite plunge and their direction is consistent with
387 the current modes (Fig. 7 a). The mean current directions are computed as a mixture of 2 Von
388 Mises distributions, which is necessary since we have two groups of directions and Von Mises
389 distributions are unimodal. Calculations were performed using the R-package "movMF"
390 (Hornik and Grün, 2014) and returned two independent distributions, the first with mean
391 direction $m=45.3^\circ$ and concentration parameter $k = 6.3$, and a the second with mean
392 direction $m 227.4^\circ$ and concentration parameter $k = 2.2$ (Fig. 7e). The current directions are
393 nearly antipodal as expected for seasonally reversing monsoon-driven currents. In the set
394 with $P' < 1.1$, the flow directions, the foliation planes and I_1 axis appear dispersed, probably
395 because bottom currents were absent or too weak to produce a coherent directional pattern
396 in elongated sediment particles (Fig. 7b and Fig. 7d). A Kuiper test for uniformity accepted
397 the Null hypothesis at the 95% confidence level testifying that these directions do not have a
398 preferential orientation. According to these observations stratigraphic intervals with larger P'
399 indicate the presence of stronger bottom currents that flow flows alternatively toward NE and
400 SW. The N-S components of the observed currents are interpreted as a deflection of
401 equatorial zonal currents in the Inner Sea of the Maldives where bottom currents are forced
402 to follow the sea floor morphology and the directions of the main channels. Inferred current
403 directions are ca. NE and ca. SW and is virtually identical to those that of present-day bottom
404 current data measured by acoustic Doppler profiler by Lüdmann et al. (2013).

405 The presence of bottom currents is not constant throughout the stratigraphic record.
406 In fact the degree of anisotropy P' is generally very small in the lower part of the stratigraphic
407 column (mean 1.05 ± 0.03) and shows larger values (mean 1.22 ± 0.17) in the upper part with
408 a sudden increase at about 168 ± 2 m CSF-A, which corresponds to the top of Chron C3n.1n and
409 an age of about 4.2 Ma (Fig. 8). The increase of anisotropy in the upper ~~ca.~~ 168 m CSF-A is
410 synchronous with a more gradual decrease of IRM intensity, which is indicative of a decreased
411 concentration of magnetic minerals. The decrease of IRM intensity can be interpreted as a
412 superimposed long-term trend with an acceleration starting at the depth of ~ 168 m CSF-A
413 (Fig. 8b). ~~mineral (Fig. 6). Changes in magnetic properties correspond to the depth of seismic~~
414 ~~reflector DS8, which marks the boundary of a sedimentary sequence within the drift deposits~~
415 ~~of the Maldives inner sea (Lüdmann et al., 2018).~~ No changes in the main lithological units
416 were observed at this depth (Betzler et al., 2017), however the decreased concentration of
417 magnetite 2017). ~~The change of magnetic properties at about 168 ± 2 mcd~~ is followed by
418 deteriorated quality of the paleomagnetic measurements and decreased sedimentation rate in
419 the upper part of Site U1467. From the sedimentological point of view, the decrease of IRM is
420 tentatively interpreted as a consequence of changes in the sediment transport mechanism -
421 controlled by wind driven currents- that transferred the sediments and the single-domain
422 magnetite, possibly of biogenic origin, from the shallow platform to the deeper water of Site
423 U1467 (Lüdmann et al., 2013). This process is modified by the increased monsoon strength
424 starting at ~ 168 m CSF-A and the depocenter of drift deposits moving downstream.
425 ~~Regardless of~~ Regardless the reason for the IRM decrease, the increased anisotropy can be
426 associated ~~with to the~~ changes in sedimentation dynamics ~~dynamic~~ that lead to ~~the~~ drift
427 deposition and that has been related to the onset of strong modern monsoon system (Betzler
428 et al., 2016b). 2016). Hence, we assume that stronger bottom currents, indicated by higher
429 anisotropy parameter P' , corresponds to stronger monsoon.

430 Our results suggest that starting from the lower Pliocene (ca. 4.2 Myr ago) the
431 monsoon-related bottom currents became ~~stronger~~ stronger enough to significantly increase the
432 degree of anisotropy and create a mostly transverse pattern in the sediments with significant
433 large AIRM. ~~Increased~~ Increase of monsoon strength could qualitatively be explained with the
434 onset of the intertropical convergence zones (ITCZ) to their present-day position. This implies
435 with a southern shift of the ITCZ south of the Himalayas ~~Himalaya~~ and an increase in the
436 latitudinal separation of the summer and winter ITCZ that moved the winter ITCZ south of
437 ~~the~~ The Maldives (e.g., Allen and Armstrong, 2012 and ~~references~~ reference therein). The
438 Himalayas ~~Himalaya~~ and Tibet have ~~are of~~ primary influences on atmospheric circulation

439 patterns and hence climate of the region. For this reason the surface uplift history of the
440 Himalayan-Tibetan orogen has been suggested to be closely linked to the development of the
441 Asian monsoon (Clift et al., 2008) and in fact, Tibetan plateau and Himalayan uplift is
442 considered necessary for the presence of the strong present day monsoon (Prell and
443 Kutzbach, 1997).

444 During the late Cenozoic the regional uplift may have occurred in two stages, one
445 beginning in the Late Miocene, ~~which that has~~ probably led to the beginning of the drift
446 deposition at 12.9 Ma (Betzler et al., ~~2016b~~,2016) followed by a later Pliocene phase dated
447 approximately from 5 to 2 Myr ago (Harrison et al, 1992; Zheng et al., 2000; An et al., 2001)
448 that could have been recorded in ~~the~~ Site U1467 ~~record~~. Independent evidence supporting a
449 coeval increase of monsoon intensity through enhanced precipitation, occurring at about 4
450 Ma Myr ago, is given by the magnetic susceptibility record from ODP site 758, (Prell and
451 Kutzbach, 1997, An et al., 2001), which is interpreted as the sea-level-mediated fluvial
452 transport from the Ganges and other river systems draining the southern side of the
453 Himalaya-Tibet plateau. Moreover, Zheng et al., (2000) interpret the increase in
454 sedimentation rate and change in depositional facies from redbeds to upward-coarsening
455 conglomerate and debris-flow deposits at the foot of the Kunlun Mountains, as evidence for
456 the uplift of the ~~north-western~~ northwestern Tibetan Plateau ~~began~~ between 3.5 and 4.5 Ma.
457 The timing of increased current strength in ~~the~~ The Maldives platform is compatible with the
458 beginning of the Pliocene uplift stage, and in fact this could mark precisely the beginning of
459 climatic influence of the Pliocene Himalayan uplift at 4.2 Ma Myr ago.

460

461 5. Conclusions

462 Paleomagnetic study of IODP Site U1467 ~~provides~~ provided a magnetic stratigraphy
463 that ~~gives~~ gave an improved age model of the Pliocene portion of Site U1467 compensating for
464 the scarcity of the biostratigraphic data in this time interval. This new age model can
465 potentially be the basis for further astrochronological studies.

466 The analysis of the AIRM has shown evidence of bottom currents with alternating
467 directions ~~similar that correspond~~ to the present-day currents, ~~originating from Asian~~
468 ~~monsoon~~. We found that the strength of the bottom currents inferred from the AIRM-
469 ~~corrected anisotropy degree~~ P' ~~parameter~~ increased suddenly at about 4.2 Myr ago. This is
470 interpreted as, according to our age model. In the formation stratigraphic record, the change
471 in current strength corresponds to the depth of stronger equatorial undercurrents as a
472 consequence of seismic reflector DS8 (Lüdmann et al., 2018) and suggests that the increased

473 monsoon ~~strength-related bottom currents had a direct effect on sediment transportation~~
474 ~~within The Maldives platform.~~

475 A number of studies relate the strength of Asian monsoon to the uplift of the
476 HimalayasHimalaya and Tibetan plateau. We observe that the timing of the increase of bottom
477 currentcurrent (4.2 Ma) coincidescoincide with the increase of fluvial transport to the Bay of
478 Bengal and is compatible with the beginning of the Late Pliocene phase of Himalayan uplift,
479 suggesting that it represents the Maldives record of the Late Pliocene uplift phase. In this case
480 our age model givesgive a precise timing of this event.

481

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487

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- 635

636 **Captions**

637

638 **Figure 1**

639 Location map of IODP Site U1467.

640

641 **Figure 2**

642 Acquisition of isothermal remanent magnetization of representative samples from the
643 investigated site (A) and the estimate of the density distribution of median destructive field of
644 the natural remanent magnetization (B). Isothermal remanent magnetization acquisition
645 shows that all-measured samples are saturated at fields higher thanabove 100-150 mT
646 indicating ~~that~~ the presence of only magnetic mineral has low -coercivity minerals. The low-
647 Low coercivity rulesrule out the presence of relevant amountsamount of hematite or
648 diagenetic iron-sulphides and suggests that magnetite (or maghemite) is the main magnetic
649 mineral in the sediments. The histogram and the density distribution of the median
650 destructive field has a mode of about 10 mT confirming that the natural remanent
651 magnetization is carried by low-coercivity minerals.

652

653 **Figure 3**

654 Representative examples of vector plots of alternating field demagnetization of natural
655 remanent magnetization.NRM. Stepwise demagnetization of natural remanent
656 magnetizationNRM of sediments from Site U1467 shows generally a very small overprint,
657 which is removed at a maximum field of 10-20 mT, followed by a linear path toward the origin
658 that ~~iswas~~ interpreted as the characteristic remanent magnetization.ChRM. Blue segments
659 represent the direction of the characteristic remanent magnetizationChRM computed as the
660 best-fit line of the selected demagnetization steps (shown in red).

661

662 **Figure 4**

663 Jelinek plot (Jelinek, 1981) illustrating the shape of anisotropy tensor (T) and corrected
664 degree of anisotropy (P'). Symbol size is proportional to the intensity of isothermal remanent
665 magnetization.

666

667 **Figure 54**

668 ChRM directions (Declination and Inclination) , maximum angular deviation and virtual
669 geomagnetic pole latitude plotted against core depth (m CFS-A)-. The latitude of the virtual

670 geomagnetic pole ~~is (VGP latitude) in~~ computed from the declination and inclination ~~in~~
671 order to better interpret the geomagnetic polarities, which are reported in the left column as
672 black and white intervals for normal and reversed polarity, respectively. The horizontal
673 dashed lines indicate cores breaks and the small symbols in the left side of the VGP Latitude
674 panel indicates the measured levels. The biostratigraphic events, core photographs and
675 sedimentary units from Betzler et al. (2017) are also reported.
676 Notice that the paleomagnetic inclinations are not significantly different from zero except for
677 transitional directions, indicating an equatorial paleo-latitude of the site.

678
679

680 Figure ~~65~~

681 Paleomagnetic interpretation and age model of the studied portion of Site U1467.
682 Shipboard biostratigraphic events are reported to provide the general age frame. The reversal
683 polarity sequence, N1 to N6, from Site U1467 is shown in the right-vertical axis. ~~The, the~~ open
684 circles connected by the red line represent the correlation of this polarity reversal sequence
685 to the reference geomagnetic polarity scale on the horizontal upper axis, ~~according to our~~
686 interpretation.

687
688

689 Figure ~~7~~

690 A and B) Equal area projection of the main anisotropy axis I_1 and foliation planes for the
691 specimens sets with $P' \geq 1.1$ and $P' < 1.1$, respectively. I_1 axis are shown in different colours
692 depending on their flow pattern. The set with $P' \geq 1.1$, mostly taken above 168 ± 2 m CSF-A,
693 shows foliation planes imbricated along the current direction, in this case imbrications
694 approximately toward NE and SW indicates currents flowing alternatively in these opposite
695 directions. C and D) Current directions shows in the circular plots (dots) together with their
696 rose diagram. The set with $P' \geq 1.1$ shows two distinct modal values while the set with $P' < 1.1$
697 have uniformly distributed directions. E) Von Mises distributions and mean values (red
698 arrows) for the set of current directions with $P' \geq 1.1$.

699

700 Figure ~~86~~

701 Summary of anisotropy of isothermal remanent magnetization AIRM data versus depth. ~~The- P'~~
702 ~~parameter~~ indicates the corrected anisotropy degree ~~of anisotropy~~, the shape parameter T is
703 illustrated with a colour code. The IRM is indicative of concentration of magnetic minerals.

704 Data have been smoothed using the locally weighted regression method (Cleveland 1979,
705 Cleveland et al., 1992) to illustrate the main trend. The 95% confidence level is shown by the
706 grey band. The reversal polarity column provides a time frame and ties the age of the
707 greengrey band marking the shift toward higher anisotropy and lower IRM intensity to the
708 top of chron C3n.1n.

709

710

711 ~~Figure 7~~

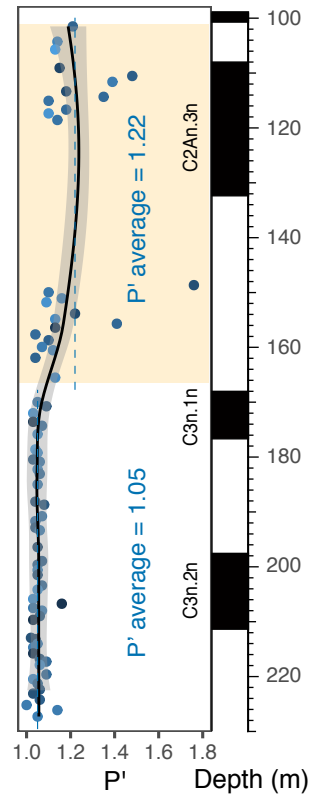
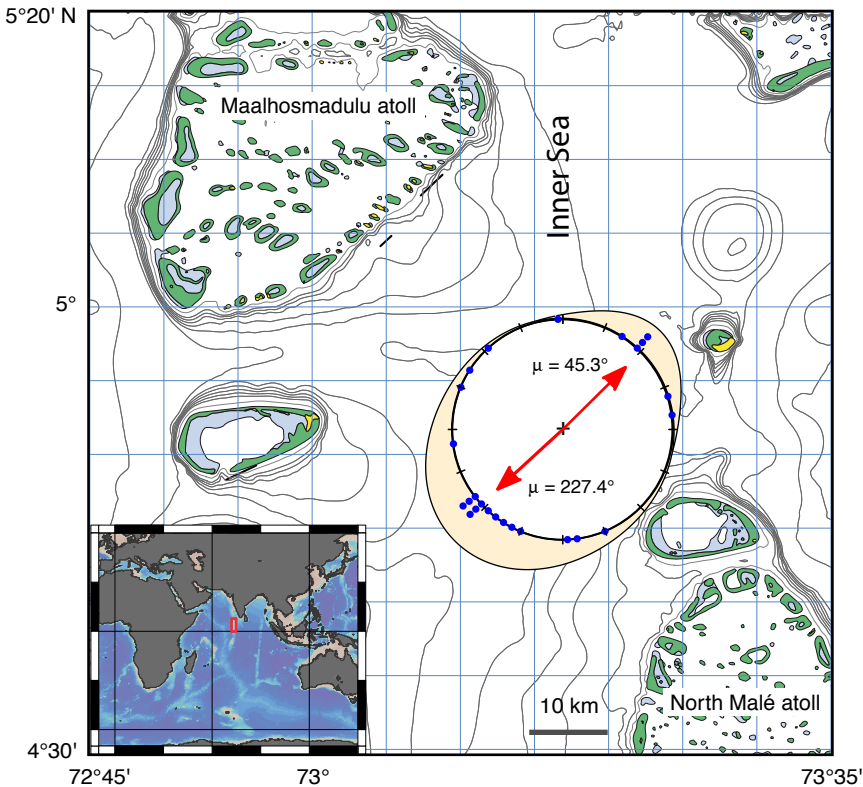
712

713

714 ~~Equal area projection of the I_{\min} -anisotropy axis and foliation planes restricted to~~
715 ~~samples taken above 168 ± 2 mcd, which correspond to the Site section with higher P' .~~
716 ~~Foliation planes are imbricated along the current direction, in this case imbrications~~
717 ~~approximately toward NE and SW indicates currents flowing alternatively in these opposite~~
718 ~~directions. The planes show mostly imbrications with moderate inclinations; planes in red are~~
719 ~~somewhat anomalous because of the high inclination.~~

Highlights

- Magneto-stratigraphic age model of Pliocene sediments from Site U1467
- Record of bottom currents from anisotropy of isothermal remanent magnetization
- Monsoon-related bottom currents increase at about 4.2 Ma
- Bottom currents increase is coeval with Late Pliocene phase of Himalayan uplift



1 Magnetic properties of early Pliocene sediments from IODP Site
2 U1467 (Maldives platform) reveal changes in the monsoon system

3
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- 66

67 **Abstract**

68 We report a study of the magnetic stratigraphy and the anisotropy of isothermal
69 remanent magnetization of Pliocene sediments from International Ocean Discovery Program
70 (IODP) Site U1467 drilled in the Maldives platform (Indian Ocean) during Exp. 359. Magnetic
71 stratigraphy gives a precise record of geomagnetic reversals of the early Pliocene from
72 approximately 5.3 Ma to 3.1 Ma providing a detailed age model in an interval where the
73 biostratigraphic record is scarce. We use the anisotropy of isothermal remanent
74 magnetization (AIRM) to investigate the statistical orientation of fine magnetic particles and
75 provide data on the strength and direction of bottom currents during the early Pliocene. The
76 strength of bottom currents recorded by the AIRM, shows a prominent increase at the top of
77 Chron C3n.1n (about 4.2 Ma), and the current direction (NE - SW) is consistent with that of
78 modern instrumental measurements. Since bottom currents in the Maldives are driven by the
79 monsoon, we speculate that the 4.2 Ma increase of bottom currents could mark the onset of
80 the present-day setting, probably related to the coeval uplift phase of the Himalayan plateau.

81

82 **Keywords:** Paleomagnetism; Pliocene magnetic stratigraphy; Anisotropy of isothermal
83 remanent magnetization; Currents strength; Monsoon

84

85 **1. Introduction**

86 Wind-induced currents are an important factor controlling the sedimentation in the
87 Maldives archipelago where they regulate sediment transport from the atoll to the deeper
88 part of the platform as well as the geometry of the sedimentary bodies (e.g., Betzler et al.,
89 2009; Lüdmann et al., 2013; Betzler et al., 2016b). The onset of these current-driven drift
90 deposits has been related to the increase of monsoon activity starting in the middle Miocene
91 (Betzler et al., 2016b). In this paper we report results from paleo and rock-magnetic analyses
92 from International Ocean Discovery Program (IODP) Site U1467 aiming to improve the
93 Pliocene age model and investigate the variations of bottom current strength.

94 Site U1467 (4° 51.0155' N and 73° 17.0204' E) was drilled in the Inner Sea of the
95 Maldives archipelago (Indian Ocean) during IODP Expedition 359 at a water depth of 487.4 m
96 (Betzler et al., 2017) in a distal position compared to the platform margins and moats (Fig. 1).
97 Site U1467 showed nearly horizontally layered seismic reflections of the so-called drift
98 sequences (Betzler et al., 2013, Lüdmann et al., 2013; Wunsch et al., 2017) with no
99 truncations and no indications of mass wasting from the adjacent platform margin. Shipboard
100 analysis suggested that Site U1467 provided a complete and undisturbed succession of all

101 drift sequences from the Late Miocene (Betzler et al., 2017) with good potential for
102 paleoceanographic and paleoclimatic studies. In particular the drift succession, from mid-
103 Miocene to recent, contains several sequences that are potentially related to fluctuations in
104 the monsoon-driven current system. Dating these sequences can yield the ages of changes in
105 the strength and direction of the currents.

106 The paleomagnetic analysis in this study provides the magnetostratigraphic age of the
107 upper portion of Site U1467 that was sampled with advanced piston coring (APC), and uses
108 the anisotropy of the isothermal remanent magnetization (AIRM) to investigate the statistical
109 orientation of fine magnetic particles aiming to provide a sedimentological record of direction
110 and strength of bottom currents.

111 Shipboard paleomagnetic measurements of Site U1467 (Betzler et al., 2016, Betzler et
112 al., 2017) gave poor results because of a combination of two factors: (i) the very low
113 concentration of magnetic minerals in carbonate platform sediments, which resulted in a very
114 weak natural remanent magnetization (NRM) and (ii) a strong magnetic contamination of the
115 cores due to metallic particles presumably originating from the drilling pipes. The
116 contamination covered the original weak paleomagnetic signal of the sediment preventing
117 any valuable measurements with the shipboard pass-through technique that measures half-
118 cores. Measurements of individual 7 cm³ box-samples, taken from the inner part of the core,
119 performed aboard, did not show signs of significant contamination suggesting that it was
120 restricted to the outer part of the cores. Unfortunately, the NRM of these specimens, ranging
121 from ca. 1×10^{-5} A/m to 1×10^{-4} A/m, was too weak to be measured reliably using the JR-6
122 spinner magnetometer available on the Joides Resolution. However, the NRM of the
123 supposedly uncontaminated box-samples is within the range of sensitivity of a DC-SQUIDS
124 cryogenic magnetometer, and this provided the motivation to collect and measure 580
125 standard sediment specimens with the aim to obtain a reliable magnetic stratigraphy of Site
126 U1467.

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131 **2. Material and sampling**

132 Standard paleomagnetic specimens (Natsuhara-Giken sampling cubes, with a volume
133 of 7 cm³) were collected in the upper part of Site U1467 from core sections 359-U1467B-11H
134 to 359-U1467B-34H and from 359-U1467C-10H to 359-U1467C-17H, corresponding to 84 m

135 to 302 m core depth below sea floor (CSF-A), at the Gulf Coast Repository at Texas A&M
136 University. Specimens were sampled only from azimuthally-oriented APC cores, since
137 information on core orientation is essential for paleomagnetic studies at equatorial latitudes
138 such as for Site U1467.

139 According to the shipboard sedimentology, the studied part of Site U1467 was divided
140 into three main lithostratigraphic units (Betzler et al., 2017). The uppermost Unit I was
141 recovered in the top 110 m CSF-A and consists of unlithified, foraminifer-rich wackestone to
142 packstone with a predominance of very fine- to fine-grained wackestone. Unit II extends from
143 ca. 110 m to 215 m CSF-A; it comprises late Pliocene sediments and is characterized by
144 interlayered unlithified and partially lithified planktonic foraminifera-rich wackestone and
145 mudstone with pteropods and particulate organic matter. Unit III, which extends from 215 m
146 to 303 m CSF-A, consists of partially lithified very fine-grained mudstone to wackestone with
147 a dominance of wackestone. The sediment contains abundant planktonic foraminifera;
148 echinoid spines and sponge spicules are common while benthic foraminifera are rare.
149 Microfossil preservation throughout Units III and most of Unit II was generally poor to
150 moderate, and in particular the interval from ca. 150 m to 300 m yields no biostratigraphic
151 ages. The interval chosen for paleomagnetic study was also intended to cover this interval
152 with poor or absent biostratigraphy.

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156 **3. Paleomagnetic analysis**

157 *3.1. Isothermal remanent magnetization*

158 Magnetic mineralogy was investigated by acquisition of isothermal remanent
159 magnetization (IRM) in a set of pilot specimens. IRM was acquired in 12 stepwise increasing
160 fields from 0.03 T to 1 T, induced using a ASC pulse magnetizer (Fig. 2a). Results indicate that
161 all measured specimens are characterized by the presence of only low-coercivity magnetic
162 minerals that saturate in fields between 100 mT and 250 mT. These coercivities are below the
163 maximum coercivity of uniaxial magnetite (e.g. Tauxe , 2002) and suggest a rather
164 homogeneous mineralogy made of ferromagnetic minerals such as magnetite (or maghemite)
165 without any significant presence of diagenetic iron sulphides that can be distinguished from
166 their higher coercivity (e.g., Tauxe, 2002).

167 The saturated isothermal remanent magnetization is relatively weak, as often found in
168 carbonate sediments, because of the low concentration of ferrimagnetic minerals. It ranges

169 from 1.0×10^{-3} to 3.7×10^{-2} A/m, suggesting a wide variability in the concentration of magnetic
170 minerals.

171

172 *3.2. Natural remanent magnetization*

173 The pass-through shipboard measurements of the NRM showed a huge scope of values
174 with intensity ranging from 5×10^{-6} A/m to 1×10^{-1} A/m along the same core and with a
175 strong downcore decreasing trend (Betzler et al., 2017). This was interpreted as the
176 consequence of steel contamination most likely originating from worn off drill pipes. The
177 bottom part of each APC core, which were the least contaminated, exhibited NRM values
178 ranging from ca. 5×10^{-6} A/m to 1×10^{-4} A/m that were considered reasonable values for
179 carbonate sediments and thus regarded as uncontaminated or only slightly contaminated.
180 Discrete specimens taken from the inner part of the cores showed similar NRM intensity
181 values corroborating the hypothesis that contamination was limited to the outer part of the
182 cores.

183 Based on these remarks, box specimens for the shore-based analysis were collected
184 from the inner part of the core in Site U1467 and were measured using a 2G-enterprise DC-
185 SQUID magnetometers at the CIMaN-ALP laboratory (Cuneo, Italy). Samples were
186 progressively demagnetized in alternating field (AF) up to the maximum field of 100 mT
187 according to a standard paleomagnetic procedure.

188 The directional components of the natural magnetization were calculated using the
189 method of the principal component analysis (Kirschvink, 1980) and the PuffinPlot software
190 (Lurcock and Wilson, 2012). The quality of the measurements and the line fitting was checked
191 by visual inspection of the orthogonal vector plots and was quantified using the maximum
192 angular deviation (MAD).

193 The AF demagnetization technique was effective in demagnetizing the NRM testifying
194 that low-coercivity minerals are the main carriers of the NRM. Vector plots generally show a
195 small viscous overprint removed in fields smaller than 20 mT, while the remaining part of the
196 NRM is demagnetized in the field interval from 20 mT to 100 mT. The component isolated
197 within this coercivity interval was used to calculate the characteristic remanent
198 magnetization (ChRM) when sufficiently linear and well defined. However, the success rate in
199 recovering reliable paleomagnetic directions was generally low and many specimens were
200 discarded because they did not yield results with acceptable quality. In specimens that gave
201 acceptable results, on average, about 20% of the NRM was removed after AF demagnetization
202 at 20 mT, and ca. 95% of the NRM was removed at field of 60 mT. Moreover, the NRM median

203 destructive field of acceptable specimens (Fig. 2b) has a modal value of 10 mT. The
204 percentage of NRM removed at 60 mT and the values of median destructive field corroborate
205 the results of IRM acquisition suggesting that stable NRM is carried by pseudo-single domain
206 magnetite. In acceptable specimens, the NRM intensity has an average value of ca. 2.1×10^{-4}
207 A/m, the average value of MAD obtained from the vector analysis of these specimens was 8.9° .
208 About 10% of specimens have MAD values between 15° and 20° , which although large, are
209 considered acceptable; most of these specimens are located in the upper part of the
210 investigated interval, between 100 and 180 m depth, or at polarity transitions. Representative
211 orthogonal vector plots for Site U1467 are illustrated in Figure 3.

212 The azimuthal orientation of cores was essential for interpretation of magnetic polarity
213 because the paleomagnetic inclinations of equatorial localities, such as Site U1467 during the
214 Miocene, are very close to zero for both normal and reversed polarities; hence the
215 geomagnetic polarities are indistinguishable if based only on inclination data. APC cores
216 collected from Site U1467 were oriented using the “tensor tool” that provided a good first-
217 order orientation. The average declinations from paleomagnetic measurements showed
218 significant departures from North and discrepancies between cores, suggesting that
219 orientation errors of the tensor tool can be as large as $\pm 30^\circ$. However, although large, these
220 errors did not compromise the polarity of ChRM and there was no ambiguity in establishing
221 the magnetic polarity. We did not attempt to remove orientation errors by adjusting the
222 magnetic declination to a mean direction even if this resulted in a reduced precision of the
223 latitude of the virtual geomagnetic pole (VGP).

224

225 *3.3. Anisotropy of isothermal remanent magnetization*

226 In agreement with shipboard measurements, the magnetic susceptibility of box-
227 samples shows negative (diamagnetic) susceptibility, evidence of the dominating diamagnetic
228 matrix of CaCO_3 on the ferrimagnetic component. The very weak diamagnetic susceptibility of
229 the carbonate sediments recovered in Site U1467 (Betzler et al., 2017) limits the possibility of
230 using the anisotropy of magnetic susceptibility to investigate the orientation pattern of the
231 magnetic particles. Therefore, we resorted to using AIRM which can be measured precisely
232 even in these weakly magnetic sediments.

233 AIRM measurements were performed in a subset of 75 specimens taken from Core 13
234 to Core 26 of Hole U1467A. To compute the AIRM an isothermal remanent magnetization
235 induced with a field of 20 mT was measured and then AF demagnetized, repeating this
236 procedure along 6 different axes. Each axis was measured twice along opposite directions for

237 a total of 12 AIRM measurements in each specimen (e.g., Stephenson et al., 1986; Jackson,
238 1991; Potter, 2004). The intensity of isothermal magnetization was measured with a JR-6
239 spinner magnetometer and the specimens were demagnetized after each measurement using
240 a tumbling 2G AF-demagnetizer at a maximum field of 80 mT, before inducing the
241 magnetization in the next direction. The anisotropy tensor and the directions of the principal
242 IRM axis I_i (i.e., the eigenvectors of the AIRM tensor) were computed from the remanent
243 magnetization using the AGICO software Anisoft42. The AIRM is therefore represented as a
244 triaxial ellipsoid, whose principal axes correspond to the directions of maximum,
245 intermediate and minimum IRM ($I_1 < I_2 < I_3$). The anisotropy ellipsoid was described using
246 the corrected anisotropy degree (P') and the shape parameter (T) computed according to
247 Jelinek (1981)

$$T = (2n_1 - n_2 - n_3)/(n_1 - n_3)$$
$$P' = \sqrt{\exp 2 (a_1^2 + a_2^2 + a_3^2)}$$

249 where $n_i = \ln I_i$, $a_i = \ln \left(\frac{I_i}{I_m} \right)$ and $I_m = \sqrt[3]{I_1 I_2 I_3}$ with $i = 1, 2, 3$ and are shown in Figure 4.

251 The direction of the largest axis of the anisotropy tensor I_1 represents the magnetic lineation
252 (the preferred orientation of elongated magnetic particles), and the foliation plane is the
253 plane that contains the I_1 and I_2 directions, hence is orthogonal to the direction of the smallest
254 axis of the anisotropy tensor I_3 .

255

256 4. Results and discussion

257 4.1. Magnetostratigraphy

258 Declination, inclinations of ChRM and the resulting VGP latitude of Site U1467 are
259 shown in Figure 5 plotted versus depth m CSF-A, together with the available biostratigraphic
260 events from shipboard analysis (Betzler et al., 2017). The comparison of measured levels and
261 acceptable samples in Figure 5 indicates the low success rate in finding reliable directions of
262 the ChRM. Sometimes, for instance in specimens taken from Hole U1467C, we could not
263 obtain any acceptable results. In general in the upper part of the Site above 110 m, which
264 comprises the sedimentological Unit I, the paleomagnetic data yielded poor results probably
265 related to the coarser granulometry of the unlithified wackestone sediments. At depths
266 between 100 m and 290 m, the quality of the data was sufficient to obtain a reliable record of
267 polarity reversal, although with variable quality.

268 In the interval with good data quality (i.e., the central part of the record with smaller
269 MAD), the ChRM inclinations are practically indistinguishable from zero, regardless of the

270 polarity, except for transitional directions corresponding to the time elapsed during the
271 reversal of the geomagnetic field. Averaged paleomagnetic directions (Table 1) indicate an
272 equatorial paleolatitude $\lambda = 0.8^\circ$ ($\lambda_{95}^+ = 4.4^\circ$; $\lambda_{95}^- = -2.8^\circ$) of Site U1467 during the early
273 Pliocene within the precision of the paleomagnetic data, which is in agreement with the
274 paleogeographic reconstructions of Besse and Courtillot (2002) and Torsvik et al. (2012).

275 The record of polarity reversals identifies 6 normal and 6 reversed magnetic polarity
276 zones that, based on the biostratigraphic framework, have been interpreted as Chrons
277 C2An.2n to C3n.4r (Fig. 6) (Gradstein et al., 2012). This interpretation is mainly based on the 4
278 biostratigraphic events available in the studied section and located in the upper part of the
279 record; however, there is some uncertainty even in the biostratigraphic data since the dates
280 based on Last Occurrence of foraminifera (*Dentoglobigerina altispira* and *Globorotalia*
281 *margaritae*) show a relatively large discrepancy with that of calcareous nannofossil events
282 (LO *Sphenolithus abies* and LO *Reticulofenestra pseudoumbilicus*). Even though the few
283 available biostratigraphic markers leave some room in interpreting which magnetochrons
284 correspond to the measured polarity reversals, we believe that our interpretation (Fig. 6) is
285 the best compromise between a reduced variability of the sedimentation rate and the
286 available biostratigraphic framework. Within these limitations, the magnetic stratigraphy of
287 Site U1467 provides a robust age constraint in a section where biostratigraphic records are
288 unavailable. In our interpretation the age of the studied section spans ca. from 5.3 Ma to 3.1
289 Ma, as reported in detail in Table 2.

290

291 4.2. Anisotropy of the IRM

292 According to a number of studies (e.g., Betzler et al., 2009, 2016b, 2018 and references
293 therein) bottom currents in the Maldives platform are considered wind-driven and assumed
294 to be a direct consequence of Asian monsoon. At equatorial latitudes, the link between
295 surface wind and bottom currents extends to a depth of several hundred meters either
296 through Ekman transport or as an undercurrent system and can be seen with modern
297 observations. The present day equatorial Indian Ocean is characterized by seasonally
298 reversing surface currents, known as Wyrтки Jets, driven by zonal winds. Beneath the surface,
299 to a depth of several hundred meters, the flow of the equatorial undercurrent and the
300 equatorial intermediate current has been observed (e.g., Knox, 1976; Reppin et al., 1999;
301 Schott and McCreary, 2001; Iskandar et al., 2009; Nyadjro et al., 2015). In contrast to other
302 oceans, the Indian Ocean equatorial undercurrent is transient and strongly dependent on
303 winds and pressure gradient variations. Both eastward and westward flows of sub-surface

304 currents have been observed, although modeling studies based on the present day suggest
305 that eastward undercurrents are more likely to occur than westward ones (Schott and
306 McCreary, 2001). Since sub-surface currents develop as consequences of surface wind it is
307 reasonable to assume that stronger surface winds will increase the strength of the
308 undercurrent, and following this argument, we interpret bottom current strength as proxy of
309 the paleo-monsoon.

310 Magnetic methods are particularly useful when other quick methods for determining
311 paleo flow from sediment beds such as macroscopic paleocurrent indicators (e.g., cross-
312 stratification and sole marks) are lacking. In standard analysis of magnetic grain shape fabric,
313 AIRM is considered to be a proxy for the preferred alignment of elongated natural magnetic
314 particles attained in the final stages of transport, with I_1 and I_3 representing preferred
315 orientations of the longest and shortest grain axes, respectively (e.g., Hamilton and Rees,
316 1970, Taira and Scholle, 1979; Novak, 2014, Felletti 2016). The method assumes implicitly
317 that the uniaxial shape-anisotropy of magnetic particles dominates triaxial
318 magnetocrystalline anisotropy, as expected for elongated magnetite particles (e.g., Tauxe,
319 2002).

320 According to theoretical, experimental and field-based fabric studies, two main
321 anisotropic fabric patterns are found (e.g., Harms et al., 1982; Baas et al., 2007): (i) flow-
322 aligned fabric; and (ii) flow-transverse fabric. In flow-aligned fabric the I_1 axes are oriented
323 parallel to the mean flow direction, while in a flow-transverse fabric the I_1 axes are oriented
324 perpendicular to the flow direction. In turbulent flows, grains settling from suspension tend
325 to orient with their I_1 axes parallel to the flow direction and imbricated upstream (Rusnak,
326 1957; Allen, 1984). This flow-aligned orientation can be changed into a more stable flow-
327 transverse orientation when the flow becomes strong enough to lift grains and roll them over
328 the surface (e.g. Schwarzacher, 1963; Johansson, 1964; Hendry, 1976, Harms et al., 1982). In
329 both cases the *foliation* planes can be imbricated dipping upstream (Harms et al., 1982) and
330 the comparison of their orientation with I_1 axes can be used to recognize the flow-aligned and
331 flow-transverse fabrics. Deviations from the flow-aligned or the flow-transverse fabrics can
332 occur for a number of reasons which include spatial changes in current direction, bed surface
333 irregularities, incomplete reorientation of a rolling fabric into a flow-aligned fabric or vice
334 versa, changes in bed roughness and post-depositional modification by bioturbation or soft-
335 sediment deformation (e.g., Baas et al., 2007 and references therein).

336 We recognise the pattern of each specimen by comparing the angle (θ) between the
337 direction of the magnetic lineation I_1 and that of the foliation plunge. If $\theta < 35^\circ$ the pattern is

338 flow-aligned and the flow is taken equal to declination of the I_1 axis in the direction of the
339 foliation imbrication; if $\theta \geq 55^\circ$ the pattern is flow-transverse and the flow is the declination of
340 $I_1 - 90^\circ$ in the direction of the foliation imbrication. The intermediate case ($35^\circ < \theta \leq 55^\circ$) is
341 handled by taking directly the imbrication direction of the foliation plane as the flow
342 direction.

343 In Site U1467, we found that the AIRM is large enough to produce a well-defined
344 pattern of orientations only if the degree of anisotropy $P' \geq 1.1$, which mostly comprises
345 specimens with flow-transverse pattern and located in the upper part on the sediment
346 column. Current directions, foliation planes and I_1 directions are shown in Figure 7 in
347 separated sets for $P' \geq 1.1$ and $P' < 1.1$. In the set with $P' \geq 1.1$, the current directions fall into
348 two distinct groups with nearly opposite modal directions highlighted by the rose diagram
349 (Fig. 7c). Foliation planes also have the opposite plunge and their direction is consistent with
350 the current modes (Fig. 7 a). The mean current directions are computed as a mixture of 2 Von
351 Mises distributions, which is necessary since we have two groups of directions and Von Mises
352 distributions are unimodal. Calculations were performed using the R-package "movMF"
353 (Hornik and Grün, 2014) and returned two independent distributions, the first with mean
354 direction $m=45.3^\circ$ and concentration parameter $k = 6.3$, and a the second with mean
355 direction $m 227.4^\circ$ and concentration parameter $k = 2.2$ (Fig. 7e). The current directions are
356 nearly antipodal as expected for seasonally reversing monsoon-driven currents. In the set
357 with $P' < 1.1$, the flow directions, the foliation planes and I_1 axis appear dispersed, probably
358 because bottom currents were absent or too weak to produce a coherent directional pattern
359 in elongated sediment particles (Fig. 7b and Fig. 7d). A Kuiper test for uniformity accepted
360 the Null hypothesis at the 95% confidence level testifying that these directions do not have a
361 preferential orientation. According to these observations stratigraphic intervals with larger P'
362 indicate the presence of stronger bottom currents that flow alternatively toward NE and SW.
363 The N-S components of the observed currents are interpreted as a deflection of equatorial
364 zonal currents in the Inner Sea of the Maldives where bottom currents are forced to follow the
365 sea floor morphology and the directions of the main channels. Inferred current directions are
366 virtually identical to those of present-day bottom current data measured by acoustic Doppler
367 profiler by Lüdmann et al. (2013).

368 The presence of bottom currents is not constant throughout the stratigraphic record.
369 In fact the degree of anisotropy P' is generally very small in the lower part of the stratigraphic
370 column (mean 1.05 ± 0.03) and shows larger values (mean 1.22 ± 0.17) in the upper part with
371 a sudden increase at about 168 ± 2 m CSF-A, which corresponds to the top of Chron C3n.1n and

372 an age of about 4.2 Ma (Fig. 8). The increase of anisotropy in the upper 168 m CSF-A is
373 synchronous with a more gradual decrease of IRM intensity, which is indicative of a decreased
374 concentration of magnetic minerals. The decrease of IRM intensity can be interpreted as a
375 superimposed long-term trend with an acceleration starting at the depth of ~168 m CSF-A
376 (Fig. 8b). No changes in the main lithological units were observed at this depth (Betzler et al.,
377 2017), however the decreased concentration of magnetite is followed by deteriorated quality
378 of the paleomagnetic measurements and decreased sedimentation rate in the upper part of
379 Site U1467. From the sedimentological point of view, the decrease of IRM is interpreted as a
380 consequence of changes in the sediment transport mechanism -controlled by wind driven
381 currents- that transferred the sediments and the single-domain magnetite, possibly of
382 biogenic origin, from the shallow platform to the deeper water of Site U1467 (Lüdmann et al.,
383 2013). This process is modified by the increased monsoon strength starting at ~168 m CSF-A
384 and the depocenter of drift deposits moving downstream. Regardless of the reason for the
385 IRM decrease, the increased anisotropy can be associated with changes in sedimentation
386 dynamics that lead to drift deposition and that has been related to the onset of strong modern
387 monsoon system (Betzler et al., 2016b).

388 Our results suggest that starting from the lower Pliocene (ca. 4.2 Myr ago) the
389 monsoon-related bottom currents became strong enough to significantly increase the degree
390 of anisotropy and create a mostly transverse pattern in the sediments with large AIRM.
391 Increased monsoon strength could qualitatively be explained with the onset of the
392 intertropical convergence zones (ITCZ) to their present-day position. This implies a southern
393 shift of the ITCZ south of the Himalayas and an increase in the latitudinal separation of the
394 summer and winter ITCZ that moved the winter ITCZ south of the Maldives (e.g., Allen and
395 Armstrong, 2012 and references therein). The Himalayas and Tibet have a primary influences
396 on atmospheric circulation patterns and hence climate of the region. For this reason the
397 surface uplift history of the Himalayan-Tibetan orogen has been suggested to be closely linked
398 to the development of the Asian monsoon (Clift et al., 2008) and in fact, Tibetan plateau and
399 Himalayan uplift is considered necessary for the presence of the strong present day monsoon
400 (Prell and Kutzbach, 1997).

401 During the late Cenozoic the regional uplift may have occurred in two stages, one
402 beginning in the Late Miocene, which probably led to the beginning of the drift deposition at
403 12.9 Ma (Betzler et al., 2016b), followed by a later Pliocene phase dated approximately from 5
404 to 2 Myr ago (Harrison et al, 1992; Zheng et al., 2000; An et al., 2001) that could have been
405 recorded in Site U1467. Independent evidence supporting a coeval increase of monsoon

406 intensity through enhanced precipitation, occurring at about 4 Ma, is given by the magnetic
407 susceptibility record from ODP site 758, (Prell and Kutzbach, 1997, An et al., 2001), which is
408 interpreted as the sea-level-mediated fluvial transport from the Ganges and other river
409 systems draining the southern side of the Himalaya-Tibet plateau. Moreover, Zheng et al.,
410 (2000) interpret the increase in sedimentation rate and change in depositional facies from
411 redbeds to upward-coarsening conglomerate and debris-flow deposits at the foot of the
412 Kunlun Mountains as evidence for the uplift of the north-western Tibetan Plateau between 3.5
413 and 4.5 Ma. The timing of increased current strength in the Maldives platform is compatible
414 with the beginning of the Pliocene uplift stage, and in fact this could mark precisely the
415 beginning of climatic influence of the Pliocene Himalayan uplift at 4.2 Ma.

416

417 **5. Conclusions**

418 Paleomagnetic study of IODP Site U1467 provides a magnetic stratigraphy that gives
419 an improved age model of the Pliocene portion of Site U1467 compensating for the scarcity of
420 the biostratigraphic data in this time interval. This new age model can potentially be the basis
421 for further astrochronological studies.

422 The analysis of the AIRM has shown evidence of bottom currents with alternating
423 directions similar to the present-day currents. We found that the strength of the bottom
424 currents inferred from the AIRM-corrected anisotropy degree P' increased suddenly at about
425 4.2 Myr ago. This is interpreted as the formation of stronger equatorial undercurrents as a
426 consequence of increased monsoon strength.

427 A number of studies relate the strength of Asian monsoon to the uplift of the Himalayas
428 and Tibetan plateau. We observe that the timing of the increase of bottom currents (4.2 Ma)
429 coincides with the increase of fluvial transport to the Bay of Bengal and is compatible with the
430 beginning of the Late Pliocene phase of Himalayan uplift, suggesting that it represents the
431 Maldives record of the Late Pliocene uplift phase. In this case our age model gives a precise
432 timing of this event.

433

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439

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582 **Captions**

583

584 **Figure 1**

585 Location map of IODP Site U1467.

586

587 **Figure 2**

588 Acquisition of isothermal remanent magnetization of representative samples from the
589 investigated site (A) and the estimate of the density distribution of median destructive field of
590 the natural remanent magnetization (B). Isothermal remanent magnetization acquisition
591 shows that all samples are saturated at fields higher than 100-150 mT indicating the presence
592 of low coercivity minerals. The low coercivity rules out the presence of relevant amounts of
593 hematite or diagenetic iron-sulphides and suggests that magnetite (or maghemite) is the main
594 magnetic mineral in the sediments. The histogram and the density distribution of the median
595 destructive field has a mode of about 10 mT confirming that the natural remanent
596 magnetization is carried by low-coercivity minerals.

597

598 **Figure 3**

599 Representative examples of vector plots of alternating field demagnetization of natural
600 remanent magnetization. Stepwise demagnetization of natural remanent magnetization of
601 sediments from Site U1467 shows generally a very small overprint, which is removed at a
602 maximum field of 10-20 mT, followed by a linear path toward the origin that is interpreted as
603 the characteristic remanent magnetization. Blue segments represent the direction of the
604 characteristic remanent magnetization computed as the best-fit line of the selected
605 demagnetization steps (shown in red).

606

607 **Figure 4**

608 Jelinek plot (Jelinek, 1981) illustrating the shape of anisotropy tensor (T) and corrected
609 degree of anisotropy (P'). Symbol size is proportional to the intensity of isothermal remanent
610 magnetization.

611

612 **Figure 5**

613 ChRM directions (Declination and Inclination) , maximum angular deviation and virtual
614 geomagnetic pole latitude plotted against core depth (m CFS-A). The latitude of the virtual
615 geomagnetic pole is computed from the declination and inclination in order to better

616 interpret the geomagnetic polarities, which are reported in the left column as black and white
617 intervals for normal and reversed polarity, respectively. The horizontal dashed lines indicate
618 cores breaks and the small symbols in the left side of the VGP Latitude panel indicates the
619 measured levels. The biostratigraphic events, core photographs and sedimentary units from
620 Betzler et al. (2017) are also reported. Notice that the paleomagnetic inclinations are not
621 significantly different from zero except for transitional directions, indicating an equatorial
622 paleo-latitude of the site.

623

624 Figure 6

625 Paleomagnetic interpretation and age model of the studied portion of Site U1467. Shipboard
626 biostratigraphic events are reported to provide the general age frame. The reversal polarity
627 sequence, N1 to N6, from Site U1467 is shown in the right-vertical axis. The open circles
628 connected by the red line represent the correlation of this polarity reversal sequence to the
629 reference geomagnetic polarity scale on the horizontal upper axis.

630

631 Figure 7

632 A and B) Equal area projection of the main anisotropy axis I_1 and foliation planes for the
633 specimens sets with $P' \geq 1.1$ and $P' < 1.1$, respectively. I_1 axis are shown in different colours
634 depending on their flow pattern. The set with $P' \geq 1.1$, mostly taken above 168 ± 2 m CSF-A,
635 shows foliation planes imbricated along the current direction, in this case imbrications
636 approximately toward NE and SW indicates currents flowing alternatively in these opposite
637 directions. C and D) Current directions shows in the circular plots (dots) together with their
638 rose diagram. The set with $P' \geq 1.1$ shows two distinct modal values while the set with $P' < 1.1$
639 have uniformly distributed directions. E) Von Mises distributions and mean values (red
640 arrows) for the set of current directions with $P' \geq 1.1$.

641

642 Figure 8

643 Summary of anisotropy of isothermal remanent magnetization data versus depth. P' indicates
644 the corrected anisotropy degree, the shape parameter T is illustrated with a colour code. The
645 IRM is indicative of concentration of magnetic minerals. Data have been smoothed using the
646 locally weighted regression method (Cleveland 1979, Cleveland et al., 1992) to illustrate the
647 main trend. The 95% confidence level is shown by the grey band. The reversal polarity
648 column provides a time frame and ties the age of the green band marking the shift toward
649 higher anisotropy and lower IRM intensity to the top of chron C3n.1n.

650

651

652

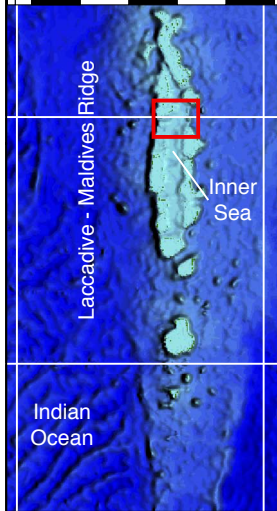
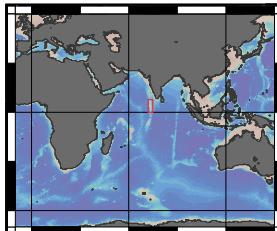
653

72°45'

73°

73°35'

5°20' N



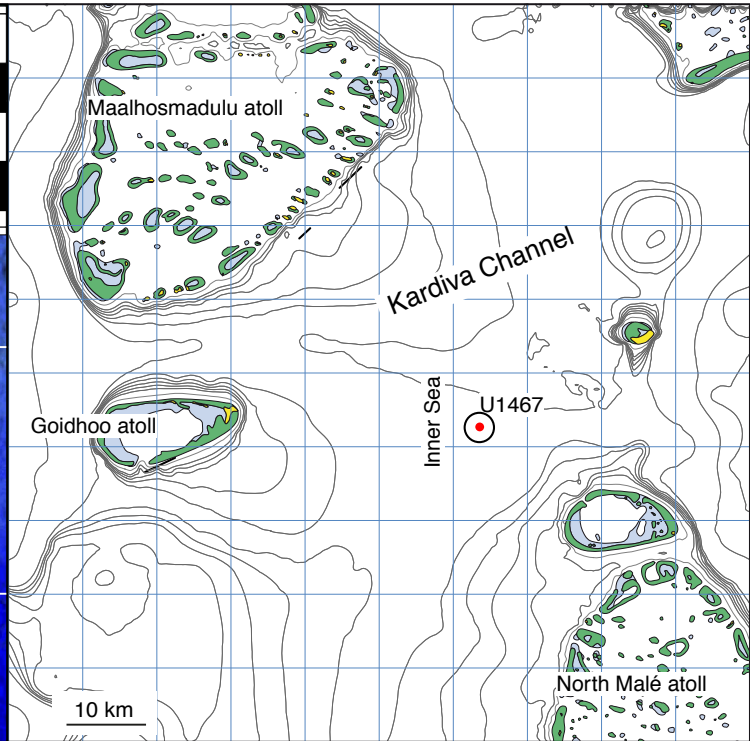
5° N

0°

Indian Ocean

70°E

75°



Maalhosmadulu atoll

Kardiva Channel

Inner Sea

U1467

Goidhoo atoll

North Malé atoll

10 km

5°

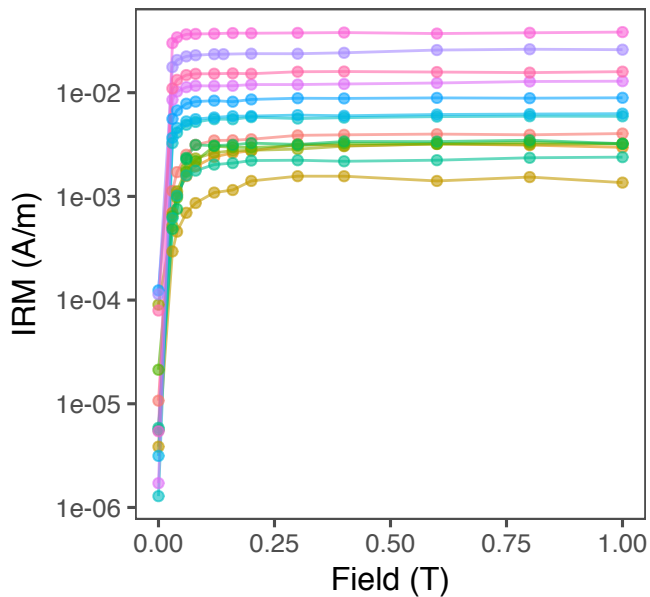
4°30'

Sample ID

- B13H5W129
- B14H5W090
- B15H3W100
- B16H2W050
- B17H1W134
- B18H3W110
- B19H2W004
- B20H5W095
- B21H5W014
- B22H5W128
- B23H3W089
- B24H3W145
- B25H5W127
- B26H1W073

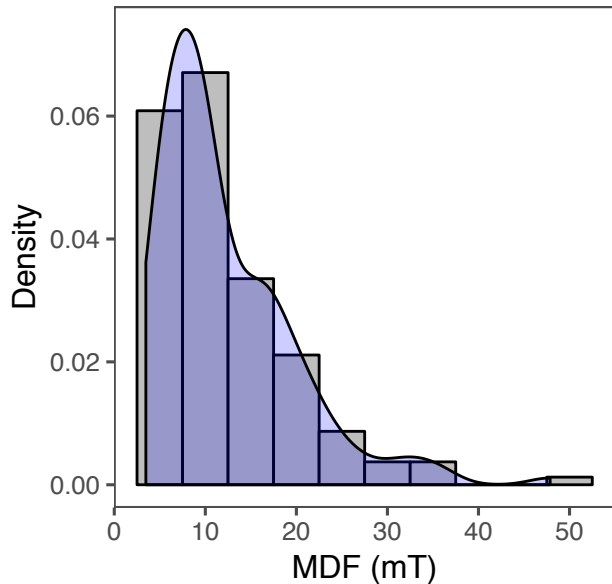
A

IRM acquisition

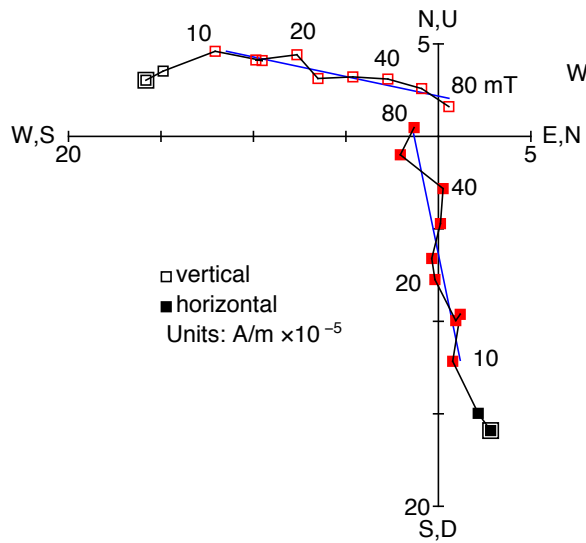


B

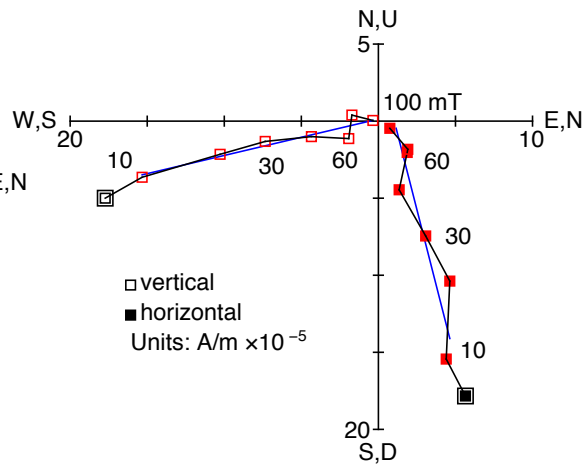
MDF distribution



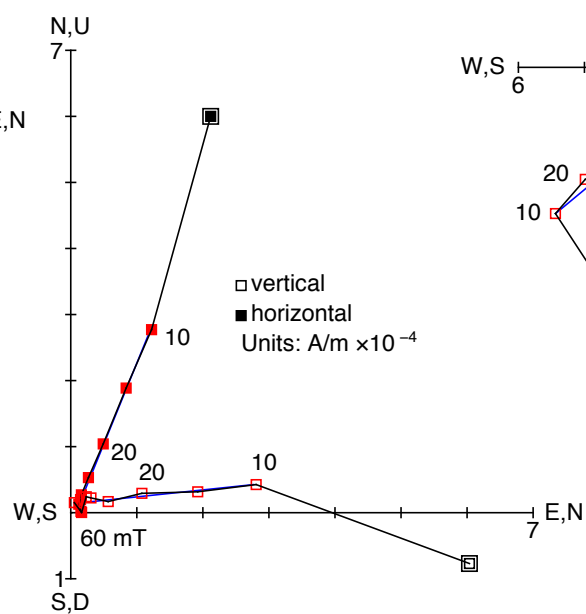
Sample: B13H3W098



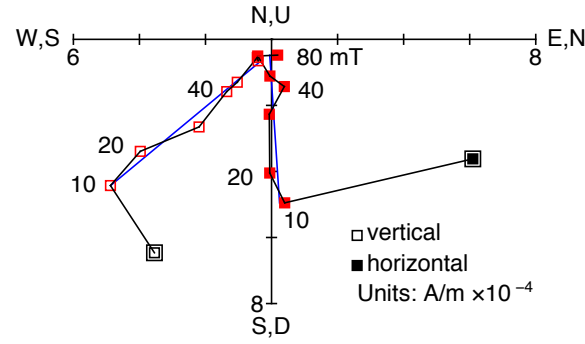
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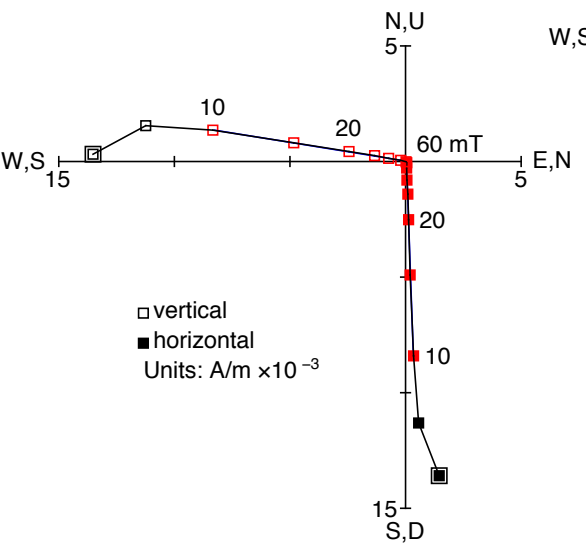
Sample: B20H6W017



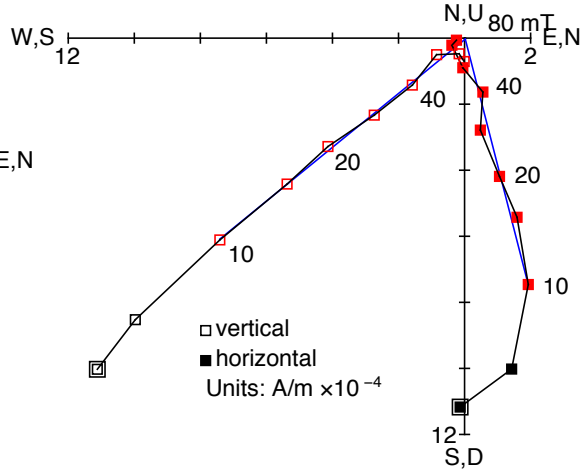
Sample: B21H2W123



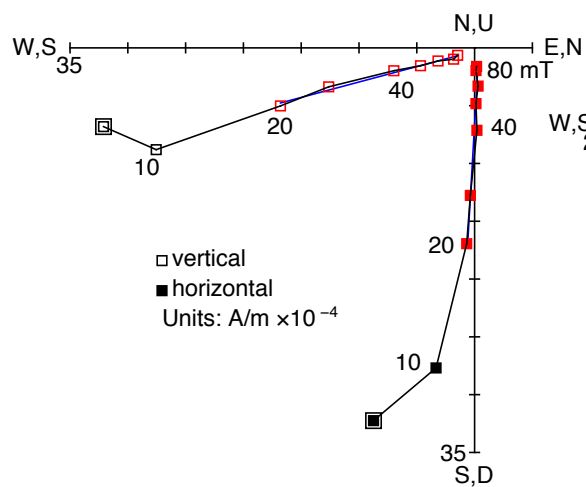
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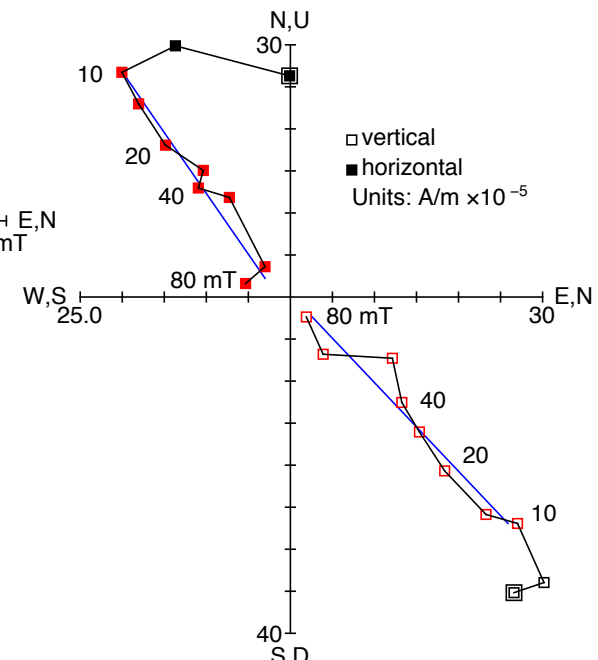
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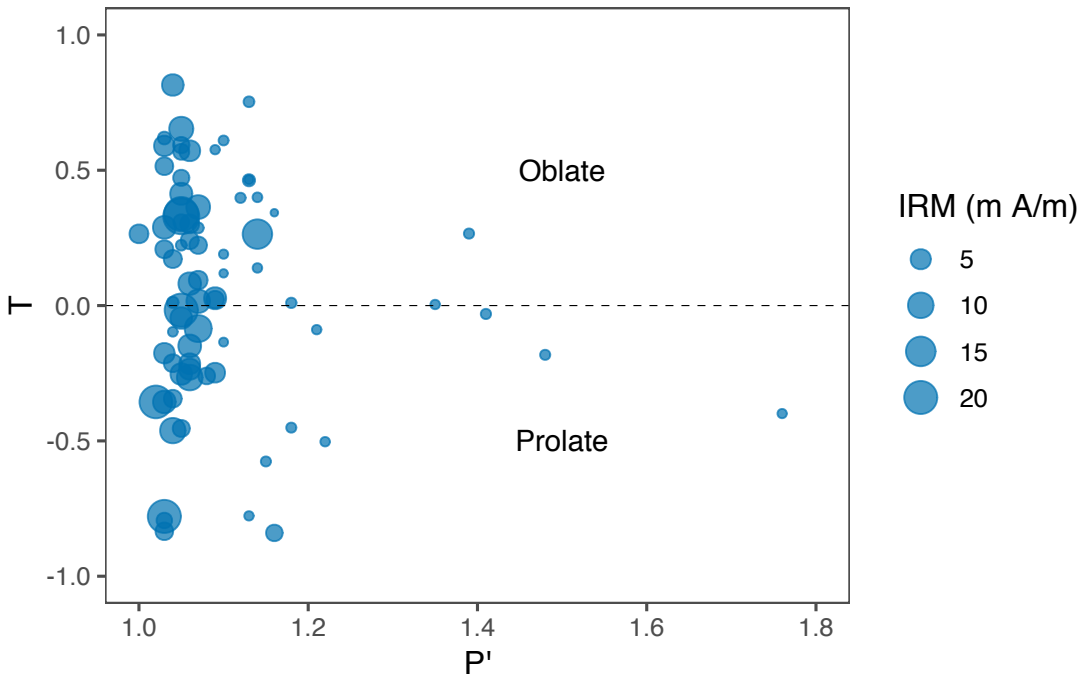
Sample: C15H3W147

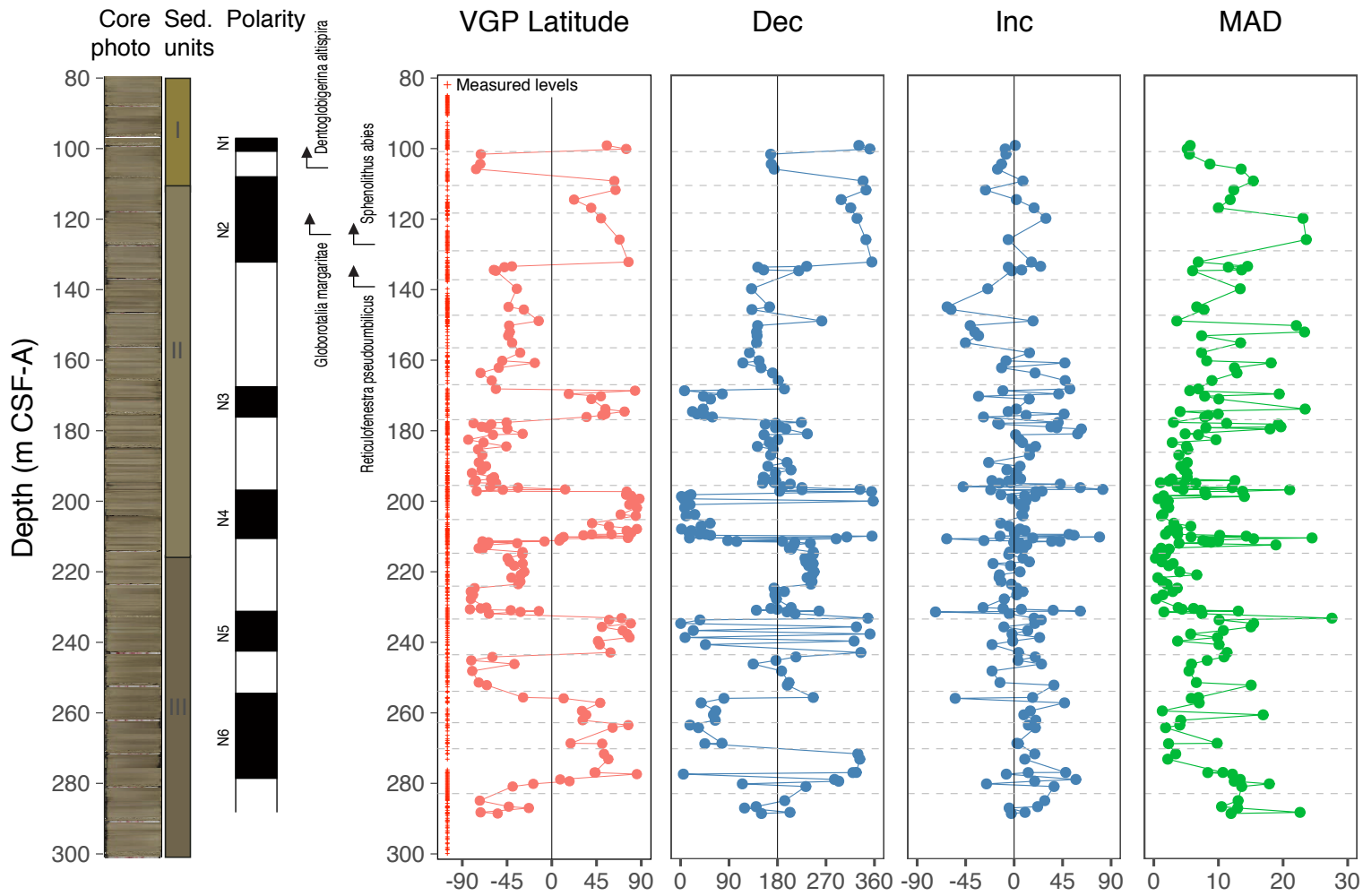


Sample: C16H4W048

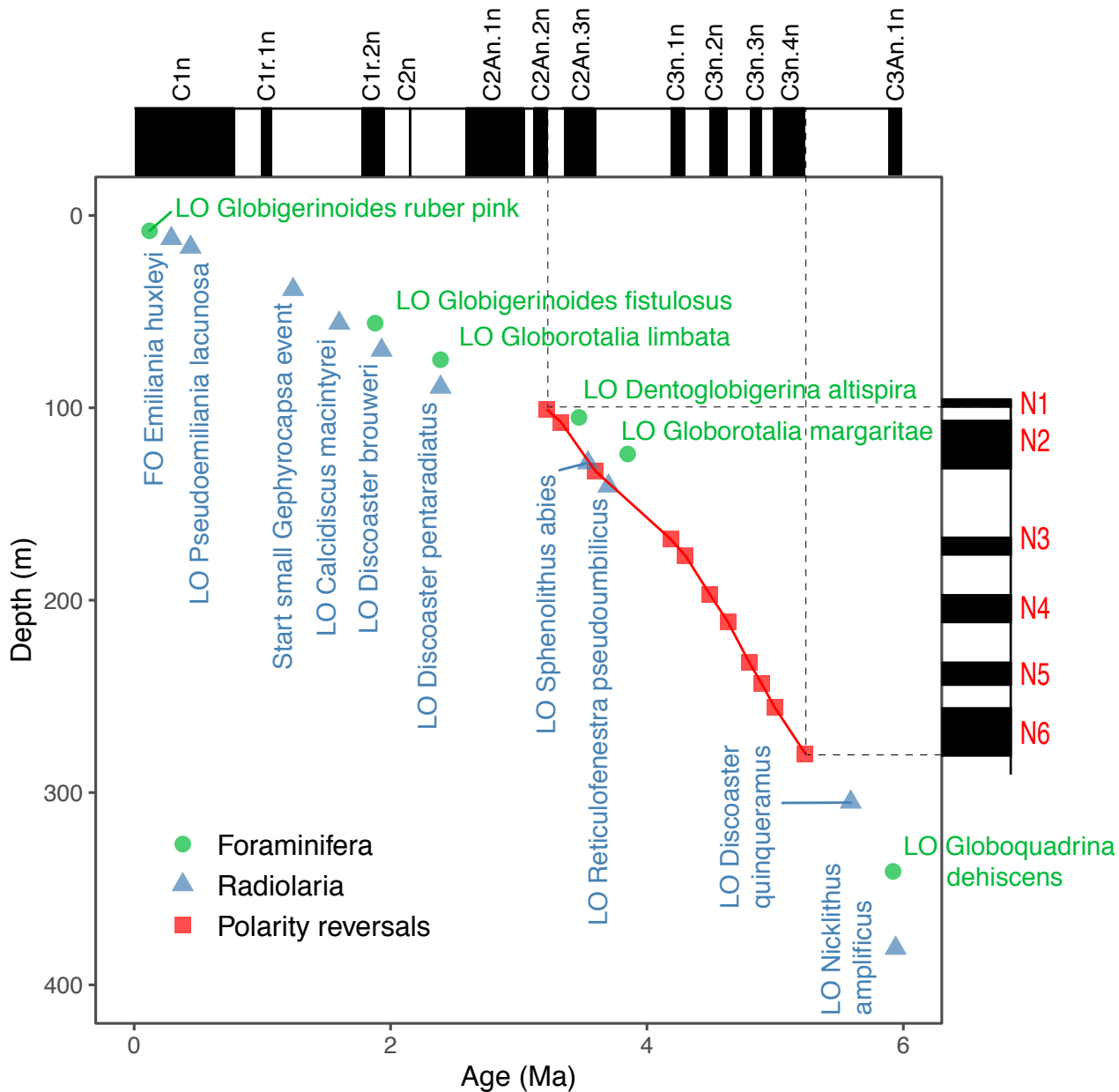


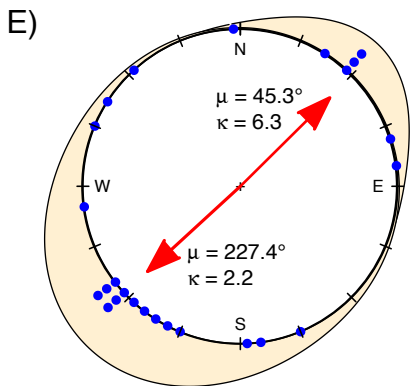
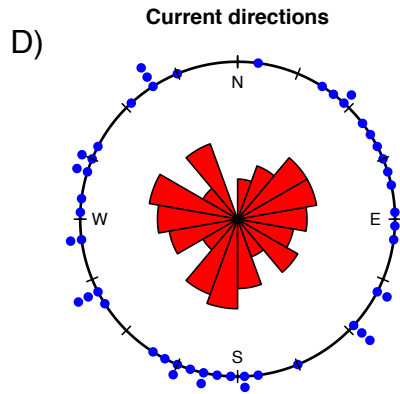
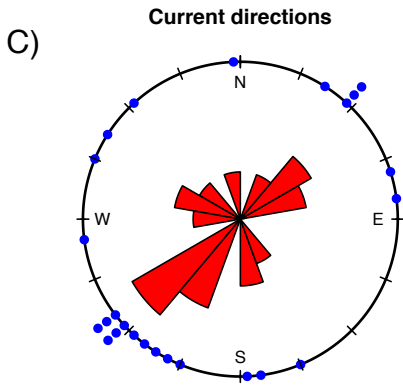
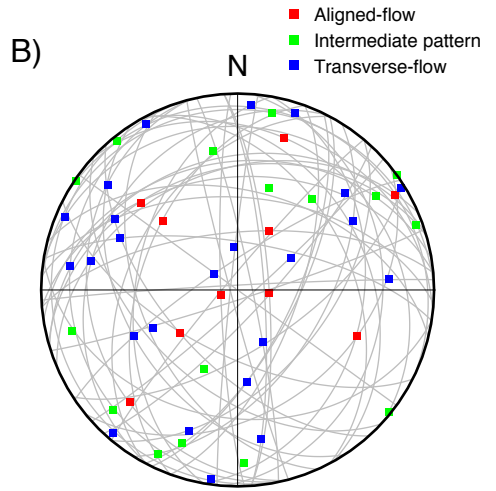
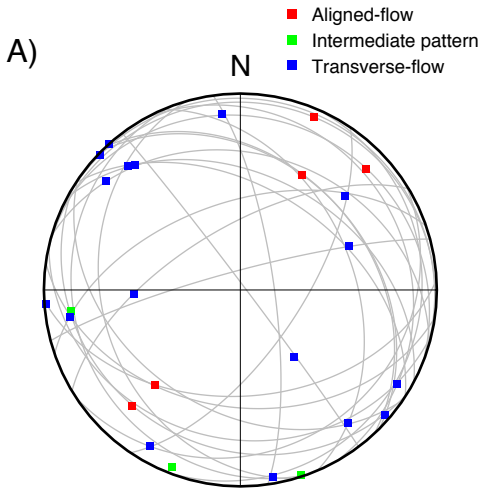
Jelinek plot





Age model



$P' \geq 1.1$ $P' < 1.1$ 

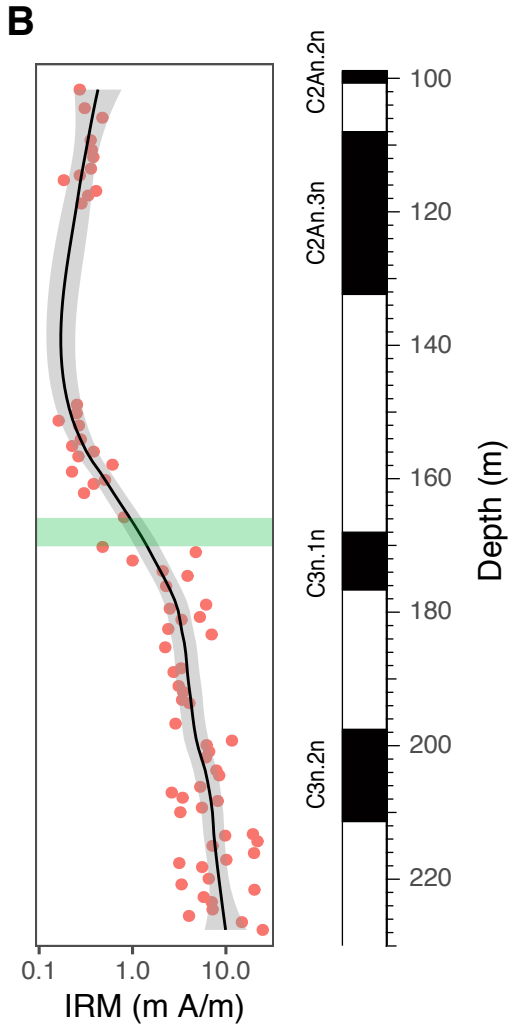
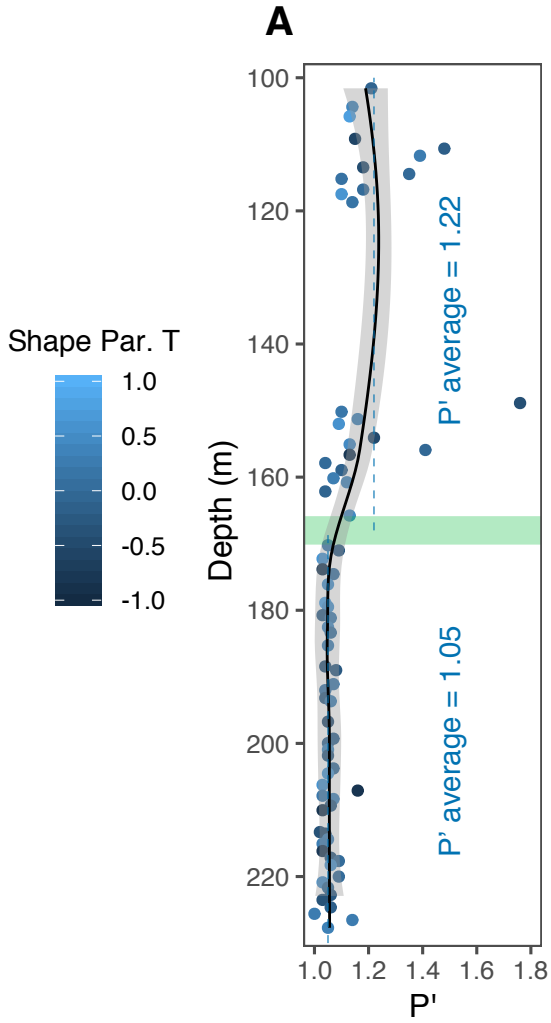


Table 1; Mean direction, Virtual Geomagnetic Pole position and paleo latitude of Site U1467 (the relatively low precision of the data is mostly a consequence of the poor azimuthal orientation of the cores).

Fisher Statistics	Dec = 5.7, Inc = 1.6, R = 121.91, k = 3.12, $a_{95} = 7.2$, N = 179
VGP	Lat = 82.9, Long = 198.5, $dm_{95} = 7.2$ $dp_{95} = 3.6$
Paleo Latitude	$l = 0.8^\circ$, $l_{+95} = 4.4^\circ$, $l_{-95} = -2.8^\circ$

Table 2: Magnetostratigraphic reversals

Chron	Age (Ma)	Depth Top (m CSF-A)	Depth Bottom (m CSF-A)
C2An.2n Bottom	3.220	100.10	101.59
C2An.3n Top	3.330	105.81	109.18
C2An.3n Bottom	3.596	132.15	133.38
C3n.1.n Top	4.187	168.14	168.65
C3n.1n Bottom	4.300	176.11	177.64
C3n.2n Top	4.493	197.15	197.26
C3n.2n Bottom	4.631	211.15	211.36
C3n.3n Top	4.799	231.89	233.15
C3n.3n Bottom	4.896	242.89	244.14
C3n.4n Top	4.997	255.63	255.89
C3n.4n Bottom	5.235	279.39	280.14

