

1 **Late Oligocene-Miocene proto-Antarctic Circumpolar Current**
2 **dynamics off the Wilkes Land margin, East Antarctica**

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18

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20

21 **ABSTRACT**

22

23 **At present, the Southern Ocean plays an important role in the global climate system and**
24 **in modern Antarctic ice sheet dynamics. Past Southern Ocean configurations are**
25 **however poorly understood. This information is yet important as it may provide**
26 **important insights into the climate system and past ice-sheet behavior under warmer than**
27 **present day climates. Here we study Southern Ocean dynamics during the Oligocene and**

28 Miocene when reconstructed atmospheric CO₂ concentrations were similar to those
29 expected during this century. We reconstruct snapshots of late Oligocene to earliest
30 Miocene (~24.2–23 Ma) paleoceanographic conditions in the East Antarctic Wilkes Land
31 abyssal plain. For this, we combine marine sedimentological, geochemical (X-ray
32 fluorescence, TEX₈₆), palynological and isotopic (ϵ_{Nd}) records from ocean sediments
33 recovered at Deep Sea Drilling Project (DSDP) Site 269. Overall, we find that sediments,
34 delivered to the site by gravity flows and hemipelagic settling during glacial–interglacial
35 cycles, were persistently reworked by a proto-Circumpolar Deep Water (CDW) with
36 varying strengths that result from climatically controlled frontal system migrations. Just
37 prior to 24 Ma, terrigenous input of predominantly fine-grained sediments deposited
38 under weak proto-CDW intensities and poorly ventilated bottom conditions dominates.
39 In comparison, 24 Ma marks the start of episodic events of enhanced proto-CDW current
40 velocities, associated with coarse-grained deposits and better-ventilated bottom
41 conditions. In particular, the dominance of P-cyst and low Calcium (Ca) in the sediments
42 between ~24.2 Ma and 23.6 Ma indicate the presence of an active open ocean upwelling
43 associated with high nutrient conditions. This is supported by TEX₈₆-derived sea surface
44 temperature (SST) data pointing to cool ocean conditions. From ~23.6 to 23.2 Ma, our
45 records reveal an enrichment of Ca in the sediments related to increased calcareous
46 microfossil preservation, high amounts of G-cysts and increasing TEX₈₆-SSTs. This
47 implies warmer water masses reaching the Antarctic margin as the polar front migrated
48 southward. Together with the radiogenic Nd isotope data indicating modern-like CDW
49 values, our records suggest a prominent poleward expansion of proto-CDW over our
50 study site and reduced AABW formation during the latest Oligocene (i.e. ~23.2 Ma ago).
51 Our findings support the notion of a fundamentally different Southern Ocean, with a
52 weaker proto-ACC than present during the late Oligocene and the earliest Miocene.

53

54 **1. Introduction**

55

56 The Antarctic Circumpolar Current (ACC) is the Earth's strongest ocean current (137-
57 162 sverdrup (Sv)) flowing eastward along a 20,000 km pathway around Antarctica (Rintoul
58 et al., 2001; Sokolov and Rintoul, 2009). Owing to the absence of land barriers, the ACC is the
59 only ocean current connecting the Pacific, the Atlantic and the Indian oceans, and consequently
60 influences the entire global ocean circulation (Rintoul, 2018). The ACC pathway is constrained
61 by ocean gateways (i.e. Drake Passage) and the bathymetry of the Southern Ocean. Its strength
62 is mainly controlled by the seafloor topography (Olbers et al., 2004), and the position and
63 intensity of the Southern Westerly Winds (SWW) (Thompson and Solomon, 2002, Aoki et al.,
64 2005; Toggweiler and Russell, 2008; Rignot et al., 2019). At present, the vigorous zonal flow
65 of the ACC prevents the intrusion of warm waters from lower latitudes to penetrate the
66 Antarctic margin and, together with sea-ice presence, contributes to maintain the cold and arid
67 glacial state of Antarctica (e.g., Olbers et al., 2004; Ferrari et al., 2014). The deep layers of the
68 ACC are occupied by a relatively warm and saline water mass, the Circumpolar Deep Water
69 (CDW) (Orsi et al., 1995). Recently, an increasing incursion of CDW into the continental
70 margins has been shown to favor melting and thinning of the Antarctic ice shelves through
71 basal melting (Pritchard et al., 2012; Liu et al., 2015; Nakayama et al., 2018; Rignot et al.,
72 2019). Despite its importance for the Antarctic and the global climate, little is known about the
73 onset and past dynamics of the ACC, as well as its linkages with the Antarctic Ice sheet (AIS)
74 dynamics. This knowledge is especially relevant from past times when climatic conditions
75 were close to the modern and future ones in terms of warmth and atmospheric CO₂
76 concentration.

77 One of these times was the late Oligocene (i.e., ~24.5 Ma), when the reconstructed
78 atmospheric CO₂ concentrations dropped below 600 ppm (600-400 ppm) (Zhang et al., 2013).
79 These values are similar to the modern and projected atmospheric CO₂ concentrations within
80 this century (IPCC, 2013; Meredith et al., 2019). Foster and Rohling (2013) argued that the
81 global ice volume is supposedly less sensitive to CO₂ fluctuations between 600 to 400 ppm. In
82 contrast, benthic foraminiferal oxygen isotope records (e.g., Liebrand et al., 2017) suggest
83 highly fluctuating ice volumes at this time. The drop in CO₂ concentration in the late Oligocene
84 (~24.5 Ma) likely led to climate cooling and ice sheet advance across the Antarctic continental
85 shelves, connecting large areas of marine-based ice with the ocean (Pekar and Christie-Blick,
86 2008; Levy et al., 2019). Ice-proximal geological records (Barrett, 1975; Naish et al., 2008;
87 Kulhanek et al., 2019; Levy et al., 2019) and seismic data (Anderson and Bartek, 1992; Sorlien
88 et al., 2007) provide direct evidence for a major expansion of marine ice sheets across the Ross
89 Sea continental shelf between 24.5 and 24 Ma. However, the oceanographic and climatic
90 conditions leading to the maximum growth of the ice sheet remain poorly known.

91 The global deep-sea benthic $\delta_{18}\text{O}$ records maximum expansion of the AIS between 23.2
92 and 23 Ma (Zachos et al., 2001; Beddow et al., 2016; Liebrand et al., 2017). However, deep-
93 sea benthic $\delta_{18}\text{O}$ records reflect a combination of ice-volume and bottom water temperature
94 and their location (low to mid-latitude versus Antarctic proximal records) determines the
95 different water masses influencing the record and thus masking information from the Antarctic
96 glaciation (e.g., Pekar et al., 2006). Thus, most of the ice volume estimates based on deep-sea
97 benthic $\delta_{18}\text{O}$ records should be taken with caution. Variations in the Southern Ocean circulation
98 and ocean heat transport across the Antarctic continental margin driven by obliquity forcing
99 have been suggested to play a significant role on ice sheet sensitivity during the late Oligocene
100 and Miocene (Salabarnada et al., 2018; Sangiorgi et al., 2018; Levy et al., 2019). This is
101 especially true in times when ice sheets extended into the marine environments (e.g., Jovane et

102 al., 2019; Levy et al., 2019). Sedimentary archives strategically located along latitudinal
103 transects across the main ACC pathway and at the vicinity of the Antarctic ice sheet are
104 however needed to provide direct links between changes in the ocean circulation and ice sheet
105 dynamics (Escutia et al., 2019).

106 Sedimentary records across the Tasman Gateway (Pfuhl and McCave, 2005) and from
107 the South Pacific (Lyle et al., 2007) document a shift to higher velocity bottom currents
108 between 25-23 Ma. This shift has been interpreted to result from the onset of a strong, deep-
109 reaching ACC during the late Oligocene. However, recent comparisons between the dinocysts
110 preserved in sediments from the Integrated Ocean Drilling Program (IODP) Site U1356 off the
111 East Antarctic Wilkes Land margin and strata from Tasmania and south of New Zealand
112 indicate a weaker than present day ACC, at least until the middle Miocene (Bijl et al., 2018b).
113 Oligocene and Miocene paleoceanographic reconstructions off the Wilkes Land margin based
114 on sedimentological data (Salabarnada et al., 2018), dinoflagellate biogeography (Bijl et al.,
115 2018a, b; Sangiorgi et al, 2018) and temperature reconstructions (Hartman et al., 2018) suggest
116 a different oceanographic configuration from that of today in this part of the Southern Ocean.
117 These authors report from multiple lines of evidence warm-temperate sea surface temperatures
118 (SST), limited sea ice expansion and reduced formation of Antarctic bottom waters, linked to
119 a weaker oceanic frontal system, which allowed the intrusion of warmer waters from low
120 latitudes towards the Antarctic margin. These data are consistent with Oligocene numerical
121 simulations, which show weaker global overturning and gyre circulation because of weaker
122 SWW (Herold et al, 2012). In addition, modeling results indicate a limited throughflow of the
123 ACC due to the Australasian paleogeography during the Oligocene (Hill et al., 2013).

124 To decipher the characteristics and dynamics of the ACC and CDW that can then be
125 related to East Antarctic Ice Sheet (EAIS) behavior off Wilkes-Adélie Land during the late
126 Oligocene-Miocene, we report new data from a sediment record recovered by the Deep Sea

127 Drilling Project (DSDP) Leg 28 at Site 269. This site was drilled on the Wilkes Land abyssal
128 plain (Hayes et al., 1975) along the main pathway of the ACC. We focus on the study of the
129 late Oligocene to earliest Miocene (~24.2-23 Ma) record. This record is partly compromised
130 by debris flows at IODP Site U1356, located on the continental rise (Escutia et al., 2011), ~280
131 km landward from Site 269, and is missing in most sedimentary archives around the rest of
132 Antarctica. Because of discontinuous drilling at Site 269, we investigate snapshots of the late
133 Oligocene and early Miocene. Sediment, palynological, geochemical and isotopic data are used
134 to describe and characterise the main changes in sedimentation related to proto-CDW
135 dynamics. The findings at DSDP Site 269 are then compared to results from IODP Site U1356
136 (Escutia et al., 2011, 2014; Salabarnada et al., 2018) (Fig. 1). This latitudinal comparison
137 provides important insights into changes in proto-ACC dynamics that can in turn be related to
138 the evolution of the ice sheet in this region of the East Antarctic margin.

139

140 **2. Site description and oceanographic setting**

141

142 DSDP Leg 28 Site 269 is located on the abyssal plain off the Wilkes Land (61°40.57'S,
143 140°04.21'E, 4282 m water depth) (Hayes et al., 1975) (Fig. 1). Two holes were
144 discontinuously drilled at this site. Our study focuses on Hole 269A, specifically the interval
145 between 655 to 956 meters below sea floor (mbsf) (cores 7R to 13R). Recovered sediments
146 were interpreted shipboard to be mostly turbidites, but evidence of winnowing by bottom
147 currents was also documented (Hayes et al., 1975). Facies were identified within the frame of
148 the DSDP Leg 28 expedition mainly based on the lithology, bioturbation and bed contacts
149 (Piper and Brisco, 1975).

150 At present, Site 269 is located between the modern southern branch of the Polar Front
151 (PF) and the northern branch of the Southern ACC Front (SACCF). In comparison, Site U1356

152 (63°18.6138'S, 135°59.9376'E, 3992 m water depth) lies south of the SACCF, near the
153 Antarctic Divergence (Sokolov and Rintoul, 2002) (Fig. 1). Both sites are today covered by the
154 Adélie Land Bottom Water (ALBW) (Rintoul, 1998), which forms in winter time along the
155 Adélie-Wilkes Land Coast (Orsi et al., 1995). Our reconstruction of the paleoposition of Site
156 269 and Site U1356 is adapted from G-plates geodynamic modeling (<http://www.gplates.org>;
157 Müller et al., 2018), and uses the plate circuit of Seton et al. (2012). It shows that Site 269 has
158 migrated south since the Oligocene but remained located at ~60°S between the late Oligocene-
159 early Miocene, while the paleolatitude of Site U1356 was around 62°S. Geological evidence
160 derived from the analysis of microfossil assemblages from sedimentary records around
161 Antarctica suggest that the proto-Polar Front (PF) was situated close to 60°S between the late
162 Oligocene-early Miocene (Nelson and Cooke, 2001; Cooke et al., 2002).

163 Figure 1 near here.

164 **3. Revising the initial age model**

165

166 We established a new age model based on the integration of new magnetostratigraphic
167 data, dinocyst and calcareous nannofossil biostratigraphy, calibrated using the GTS 2012
168 Astronomic Age Model (Gradstein et al., 2012) (Figs. 2; S1; S2 and Tables S1; S2). The
169 presence of *Operculodinium janduchenei* in the sediment between ± 753-955 (mbsf) (Cores
170 9R 3W to 13R) is assigned to the lower Southern Ocean Dinocyst Zone (SODZ) SODZ8 (Bijl
171 et al., 2018a). This suggests that the bottom of Hole 269A cannot be older than late Oligocene
172 (24.2 Ma). Moreover, the last occurrence of *Operculodinium janduchenei* between 753.27-
173 752.32 mbsf (9R 3W 125-129 cm - 9R 3W 30-33 cm) marks the boundary between dinocysts
174 zones SODZ8 and SODZ9. This boundary is calibrated to 23.6 Ma at U1356 (Bijl et al., 2018a)
175 and is correlated to the reversed polarity found at the bottom of chron C6C (C6Cr). This is
176 corroborated by the presence of *Cyclicargolithus abisectus* at 752.9 mbsf indicating that this

177 interval corresponds to the zone NP 25 (Martini, 1971), which implies a latest Oligocene age.
178 The latter however, must be taken with caution given the scarce amount of calcareous
179 nannofossils, as well as the absence of other conventional markers. The presence of
180 *Impagidinium aculeatum* at 658.82 mbsf (7R 4W 30-34 cm) is assigned to the mid- of SODZ10
181 or older (~23 Ma) (Bijl et al., 2018a). This is correlated with the normal polarity of the
182 geomagnetic chron C6n.2n suggesting also an age of 23 Ma. In addition, based on the
183 biostratigraphic datums we assigned the paleomagnetic reversal at ~ 857.5 mbsf to chron
184 C7n.1n (24 Ma) and at ~703 mbsf to chron C6n.3n (23.23 Ma). Our updated age model
185 suggests an age range for the studied interval from ~24.2 to ~23 Ma (Fig. 2). Initial shipboard
186 data had dated this interval to be of middle Eocene to early Oligocene age (Hayes et al., 1975).
187 We acknowledge that the age model still has large uncertainties due to discontinuous coring,
188 incomplete recovery and the low preservation of microfossils. However, the integration of the
189 well-established Southern Ocean Oligocene-Miocene dinocyst stratigraphy, which was
190 developed at the nearby Site U1356 along with the good correspondence with the
191 magnetostratigraphic data allow us to be relatively confident in our age model.

192 Figure 2 near here.

193 **4. Methods**

194

195 **4.1 Facies analysis**

196

197 A detailed facies analysis was performed on sediment from Hole 269A to determine
198 depositional processes and aid paleoenvironmental reconstructions. We conducted a detailed
199 description of the cores using standard sedimentological techniques (i.e., lithological
200 characterization, contacts, sedimentary structures and textures) in order to produce the
201 lithostratigraphic log in Figure 3 and Supplementary Figure 3. Visual descriptions were aided

202 by high-resolution digital images obtained on the archive halves using a Nikon 60mm camera
203 lens mounted on a custom-built line scanner at the Coast Gulf Repository (CGR), in College
204 Station (Texas, USA). High-resolution images were also used for ichnological analysis (i.e.,
205 type and degree of bioturbation) performed to characterise the paleoenvironmental conditions
206 prevailing in the seafloor. Ichnological analysis was based on the digital treatment of the high-
207 resolution images following previous researches (Dorador and Rodríguez-Tovar, 2014, 2018;
208 Dorador et al., 2014a, 2014b; Rodríguez-Tovar and Dorador, 2015). It was conducted at the
209 Department of Stratigraphy and Paleontology at the University of Granada (Spain). Further
210 characterisation of the biogenic and terrigenous material within each of the lithofacies defined
211 was achieved by: 1) Bulk grain-size analyses that was performed at EPOC (Environnements et
212 Paléoenvironnements Océaniques et Continentaux) (Bordeaux, France). In total 44 sediment
213 samples (dried overnight in an oven at 40°C) were measured in a laser microgranulometer
214 Malvern mastersizer hydro 2000G with automatic samples (0.020 to 2000 µm). 2) Wet sieving
215 and High-Resolution Scanning Electron Microscope (HR-SEM) analysis conducted at the
216 Instituto Andaluz de Ciencias de la Tierra, (CSIC, Spain) and at the Centro de Instrumentación
217 Científica (University of Granada, Spain), respectively. Moreover, continuous measurements
218 of magnetic susceptibility (MS) were taken from the archive half sections of the core. For this,
219 the core sections were left out of the refrigerator overnight to acquire a room temperature before
220 scanning. MS measurements were taken every 2 cm using a Bartington MS2 mounted in a
221 Geotek Multi-Sensor Core Logger (MSCL) at the GCR. Core void measurements were
222 removed from the data set.

223

224 **4.2 Major and Trace Element content**

225

226 The elemental composition of sediments derived from X-Ray Fluorescence (XRF) core
227 scanners has been used as an indicator of past climate changes in proximal and distal records
228 from the Antarctic margin. For example, titanium (Ti), aluminium (Al), iron (Fe) versus
229 calcium (Ca) have been interpreted to show changes between terrigenous and biogenic CaCO₃
230 deposition (e.g., Grützner et al., 2005; Hepp et al., 2006; Escutia et al., 2009; Salabarnada et
231 al., 2018). In addition, zirconium-based proxies (e.g. Zr/Ti ratios) have previously been used
232 as an indicator of semiquantitative assessment of bottom current velocities in Oligocene
233 sediments from Site U1356 off Wilkes Land (Salabarnada et al., 2018). This is based on the
234 relative enrichment of heavy minerals such as zircon over less dense minerals (e.g.,
235 aluminosilicates) that result from hydrodynamic winnowing and sorting of heavy minerals
236 (e.g., Bahr et al., 2014). Bromine (Br)/Ti ratios have been used as indicator of organic matter
237 in the sediments (Bahr et al., 2014; Salabarnada et al., 2018). The Br content is associated to
238 total organic carbon (TOC) concentration in the sediments (Seki et al., 2019), and TOC is
239 mainly related to changes in productivity (organic carbon flux in the sediments) and/or changes
240 in the redox state of the sediment (e.g., Jimenez-Espejo et al., 2007). Br is not biased by
241 lithological changes (i.e. mudstones/sandstones), but is affected by organic matter degradation
242 (Bahr et al, 2014).

243 XRF core scanning measurements were conducted at 10kV and 30 kV on the archive
244 split sections from Cores 7R to 13R using an Avaatech XRF core scanner at the GCR. The
245 surface of core sections was cleaned carefully for any gypsum and salts, which might have
246 precipitated and then adjusted manually to form an even surface. A 4µm thick ultralene plastic
247 film was used to cover the core surface in order to avoid contamination while scanning. Due
248 to the presence of cracks in many core sections, spot measurements were taken rather than scan
249 the section continuously.

250 To decipher the different processes influencing the geochemical composition of the
251 sediment, we conducted a Principle Component Analysis (PCA). We used the PAST version
252 2.10 software package (Hammer et al., 2001) following the data pre-treatment in Bahr et al.
253 (2014), including normalization of the data to reduce the signal artefacts related to changes in
254 lithology. For the PCA, we only selected and show elements with a robust signal quality, i.e.,
255 Al, Si, K, Ca, Ti, Fe, Ba, Br, Rb, Sr, and Zr. In this study we report on the following elements
256 and elemental ratios Ca, Ti, Zr/Ti ratios and Br/Ti ratios.

257

258 **4.3 Neodymium isotopes**

259

260 Neodymium (Nd) is delivered to the ocean through the weathering of continental crust,
261 and by exchange with sediment on the continental margin (known as boundary exchange; e.g.,
262 Frank et al., 2002, Wilson et al., 2013). Consequently, water masses forming in different
263 geological basins will be isotopically distinct. Because of its short oceanic residence time in
264 the ocean (~400-1,000 yr) relative to oceanic mixing (1,500 yr) (Tachikawa et al., 2003)
265 records of Nd isotopes allow us to reconstruct past ocean circulation and changes in the
266 weathering inputs. Neodymium isotopes in fossil fish debris are a robust tool to reconstruct
267 changes in ocean circulation by identifying distinct water masses (e.g. Goldstein and
268 Hemming, 2003; Martin and Scher, 2004; van de Flierdt et al., 2016). Neodymium is
269 incorporated into the fish teeth during the fossilization processes at the sediment-water
270 interface and reflects the isotopic composition of seawater in contact with the seafloor at the
271 time of fish tooth deposition, remineralisation and burial (Shaw and Wasserburg, 1985).
272 Neodymium isotope ratios in fossil fish teeth are considered to be resistant to changes in post-
273 burial alteration (Martin and Scher, 2004; Scher et al., 2011). Neodymium isotope ratios
274 ($^{143}\text{Nd}/^{144}\text{Nd}$) are expressed as ϵ_{Nd} , which denotes the deviation of a measured $^{143}\text{Nd}/^{144}\text{Nd}$ ratio

275 from the chondritic uniform reservoir in parts per 10,000 (DePaolo and Wasserburg, 1976;
276 Jacobsen and Wasserburg, 1980).

277 Fish debris was handpicked from the $>63 \mu\text{m}$ sediment fractions that were prepared by
278 wet sieving. Four samples were prepared for fish-tooth and bone debris Nd isotope analyses in
279 the MAGIC laboratories at Imperial College London following the sample preparation detailed
280 in supplementary materials. In addition, to account for a potential influence of the detrital
281 sediment towards the Nd isotope composition of pore waters or overlying bottom water, we
282 also measured the Nd isotope compositions of two detrital sediment samples. Sample
283 preparation for these analyses was conducted in the MAGIC laboratories at Imperial College
284 London as detailed in supplementary materials. The detrital samples were processed using the
285 same ion chromatography as the fish debris.

286 Neodymium isotope ratios for fish debris and detrital sediment samples were
287 determined on a Nu Plasma multiple collector inductively coupled plasma mass spectrometer
288 (MC-ICP-MS) at Imperial College London, operated in static mode. Instrumental mass bias
289 was corrected using the $^{146}\text{Nd}/^{144}\text{Nd}$ ratio of 0.7219. A JNdi-1 isotope standard was run after
290 every sample and all reported $^{143}\text{Nd}/^{144}\text{Nd}$ ratios are corrected to a JNdi value of 0.512115
291 (Tanaka et al., 2000) using bracketing standards. External reproducibility was monitored using
292 the JNdi standards, and accuracy was confirmed by measuring USGS BCR-2 rock standards,
293 which yielded average $^{143}\text{Nd}/^{144}\text{Nd}$ ratios in agreement with the published BCR-2 $^{143}\text{Nd}/^{144}\text{Nd}$
294 ratio of 0.512638 ± 0.000015 (Weis et al., 2006).

295 To correct for the decay of ^{147}Sm to ^{144}Nd within the fish teeth over time we used Sm
296 and Nd concentrations obtained from two samples, which were in good agreement with
297 $^{147}\text{Sm}/^{144}\text{Nd}$ ratios reported from fossil fish teeth in other marine sedimentary records (e.g.,
298 Martin and Scher, 2006; Moiroud et al., 2013; Huck et al., 2017; Wright et al., 2018). When
299 no values available, an average of 0.1286 from the measured samples was applied to calculate

300 $\epsilon_{Nd(t)}$ values. The $^{147}Sm/^{144}Nd$ ratios yielded Nd isotope corrections of 0.19 to 0.21 ϵ_{Nd} units; (t)
301 denotes samples have been corrected for in situ decay of ^{147}Sm .

302 **4.4 TEX₈₆**

303

304 The TetraEtherindex of 86 carbon atoms (TEX₈₆) is a proxy for sea surface temperature
305 (SST). The TEX₈₆ paleothermometer employs the temperature-dependent relative abundance
306 of a suite of thaumarchaeotal membrane lipids; glycerol dibiphytanyl glycerol tetraethers
307 (GDGTs) (Schouten et al., 2002, 2013). In short, this method involves lipid extraction from
308 powdered and freeze-dried sediments using accelerated solvent extraction. Lipid extracts were
309 separated into an apolar, ketone and polar fraction by Al₂O₃ column chromatography using
310 hexane:DCM (9:1), hexane:DCM (1:1) and DCM:MeOH (1:1) as respective eluents. 99 ng of
311 a synthetic C₄₆ (mass-to-charge ratio, $m/z = 744$) GDGT standard was added to the polar
312 fraction, which then was dissolved in hexane:isopropanol (99:1, v/v) to a concentration of ~3
313 mg ml⁻¹ and filtered over a 0.45- μ m polytetrafluoroethylene filter. The dissolved polar
314 fractions were injected and analysed by high-performance liquid chromatography–mass
315 spectrometry (HPLC–MS), using the double column technique for improved separation of co-
316 eluding compounds (Hopmans et al., 2015). GDGT peaks in the HPLC chromatograms were
317 integrated using ChemStation software. Several calibrations have been proposed to translate
318 TEX₈₆ into SST. We employ the TEX₈₆ linear calibration ($SST = 81.5 \times \text{TEX}_{86} - 26.6$ with a
319 calibration error of ± 5.2 °C) by Kim et al. (2010), to enable adequate comparison to existing
320 data, and following elaborate discussions in Hartman et al. (2018). We provide the GDGT peak
321 areas and R-script in the online supplementary materials.

322 Before interpreting the TEX₈₆ results into SST reconstructions samples with overprint,
323 which may affect the reliability of the SST proxy, must be discarded. Terrestrial GDGT input
324 has been reconstructed using the branched and isoprenoid tetraether (BIT) index (Hopmans et

325 al., 2004) as a proxy. Samples with a high BIT index may be biased by soil- and river-derived
326 GDGTs (Hopmans et al., 2004), although the high BIT index may also stem from a decrease
327 in marine crenarchaeol production, as the BIT index is a closed sum between terrestrial GDGTs
328 and the exclusively marine crenarchaeol. Several indices for a potential biased source of GDGTs
329 was investigated and outliers discarded. Namely, methane index (Zhang et al., 2011), flagging
330 overprint by sedimentary methanogenic activity, GDGT-2/GDGT-3 ratio (Taylor et al., 2013),
331 potentially signalling overprint by archaeal communities dwelling deeper into the water
332 column, GDGT-0/Crenarchaeol ratio (Blaga et al., 2009; Damsté et al., 2009; Taylor et al.,
333 2013), flagging overprint by in situ production of isoprenoidal GDGTs in lakes and rivers, and
334 ring index (Zhang et al., 2016), which assesses an overall pelagic character for the different
335 GDGTs within the TEX₈₆ index. High-latitude TEX₈₆ SST reconstructions might be skewed
336 towards summer temperatures (Ho et al., 2014; Schouten et al., 2013; Hartman et al., 2018)
337 and potentially incorporate a subsurface signal (0-200 mbsl) (Hernández-Sánchez et al., 2014;
338 Ho and Lepple, 2015). However, there is a general agreement that TEX₈₆ captures the relative
339 SST trend (Richey and Tierney, 2016).

340

341 **4.5 Palynology – dinoflagellate cyst paleoenvironment reconstruction**

342

343 In total 19 sediment samples (~15g) were processed for palynology and counted at
344 Utrecht University using standard palynological processing and analytical procedures
345 previously described by Bijl et al. (2013; 2018b). Modern dinoflagellate (unicellular planktonic
346 protists) distribution are sensitive to small changes in nutrient availability, SST, salinity,
347 bottom water oxygen, primary productivity and sea-ice cover (Dale, 1996; Prebble et al., 2013;
348 Zonneveld et al., 2013). Approximately 13-16% produce an organic-walled cyst, dinocysts,
349 that can preserve in the fossil record (Head, 1996). Assuming the habitat affinities and feeding

350 strategies of most dinoflagellates can be extrapolated to the fossil assemblages, we can utilize
351 ‘deep-time’ dinocysts assemblages as a paleoceanographic proxy (Bijl et al., 2013; Crouch et
352 al., 2014; Sluijs et al., 2005). In the Southern Ocean, protoperidinioid (P) cysts originate from
353 heterotrophic dinoflagellates and proliferate under increased nutrient conditions, while
354 gonyaulacoid (G) cysts originate from autotrophic or mixotrophic dinoflagellates and reflect
355 oligotrophic conditions (Esper and Zonneveld, 2002). Thus, the relative P/G-cyst ratio can
356 indicate glacial/interglacial variability and ocean frontal movement migrations. Today, the
357 surface sediments at Site 269A consist almost exclusively of P-cysts, specifically the sea-ice
358 affiliated species *Selenopemphix antarctica* (Prebble et al., 2013).

359 **5. RESULTS**

360

361 **5.1. Sedimentation at Hole 269A**

362

363 Sediment from Hole 269A from 655 to 956 mbsf consists of alternations between
364 bioturbated and laminated intervals of terrigenous-rich sediment (Figs. 3a, 3b, 4 and S3).
365 Textural analyses show a low clay content (4 to 20%), a high silt fraction (40 to 80%) and a
366 sand content between <5 and 60 % (Fig. S3). Microfossil preservation is generally low
367 throughout the study interval. Higher preservation of calcareous microfossils was found in the
368 carbonate-cemented beds (Figs. 3a and 5) and within the sediments between 753.5-702.5 mbsf
369 (~23.6 to 23.23 Ma) (Fig. S4). The higher calcareous carbonate preservation is depicted in Ca
370 peaks in Fig. 3d. Higher preservation of diatoms is also observed in the carbonate-cemented
371 facies (Fig. 5) and in Core 7 (660-655.5 mbsf) (~23 Ma).

372 The main observed differences within the bioturbated and laminated intervals are the
373 variations in the silt/sand content, ichnological features, including amount of bioturbation, and
374 carbonate content (carbonate-cemented intervals). Based on this, we differentiate the following

375 sedimentary facies (Figs. 3c and 4): (1) Bioturbated silty claystones to clayey siltstones
376 (Bioturbated mudstone; F1a), bioturbated siltstones to sandy siltstones (F2a), and bioturbated
377 carbonate-cemented facies (F3a). (2) Laminated silty claystones to clayey siltstones
378 (Laminated mudstone; F1b), laminated siltstones to sandy siltstones (F2b), and laminated
379 carbonate-cemented facies (F3b).

380 Figure 3 near here.

381 Bioturbated facies (F1a, F2a and F3a; Figs. 3c, 4 and 5) generally exhibit a structureless
382 and/or mottled texture. Bioturbation varies from low, dominated by *Chondrites* and
383 *Phycosiphon*, to high with abundant and diverse trace fossils, including *Planolites*,
384 *Thalassinoides*, *Nereites*, *Zoophycos* and likely *Scolicia* (Fig. 4). The occurrence of coarser
385 bioturbated facies (F2a) becomes more frequent up-section starting at 858 mbsf, but is also
386 present in the lowermost part of the site at ~ 955.3 mbsf (Fig. 3c).

387 Laminated facies (F1b, F2b and F3b; Figs. 3c, 4 and 5) are characterised by faint and/or
388 distinct sub-mm to mm silty-sandy laminations. Laminated intervals can contain a single or a
389 group of laminae with various sedimentary structures, including continuous and discontinuous
390 planar, wavy, lenticular, ripple and cross laminations (e.g., Fig. 4a). Small soft-sedimentary
391 deformation structures such as convoluted, ball and pillow structures are often observed within
392 the laminations (Fig. 4k). Scarce traces of *Chondrites* may be present.

393 Figure 4 near here.

394 Despite the diagenetic processes associated with the bioturbated and laminated
395 carbonate-cemented facies (F3a, F3b), thin section analyses integrated with HR-SEM images
396 show preservation of microfossils in both facies, including silicified planktonic foraminifers
397 and diatoms (Figs. 5f-h). This observation is noteworthy since the sediments of the study
398 sections were considered almost barren in microfossil (Hayes et al., 1975). The presence of
399 benthic foraminifera with siliceous test suggests diagenetic dissolution and confirms the

400 biogenic origin of these carbonates (Fig. 5). HR-SEM images reveal that the contact between
401 F3b and F3a is sharp and erosive and inverse grading characterizes F3b (Fig. 5e). In total
402 thirteen carbonate-cemented beds were observed, with the thickest ones at 704.2 mbsf and at
403 655.3 mbsf (e.g., Figs. 3c and 3d).

404 Figure 5 near here.

405 The intercalation of bioturbated and laminated facies is characterized by both
406 gradational and sharp contacts. More than 90 sharp contacts were identified (Fig. S3). A
407 common ichnological feature that appears associated with at the sharp contact between the
408 coarse and the fine-grained facies is the presence of pseudo-borings, characterized by well-
409 defined shapes, undeformed, with sharp contacts, infilled with different material to the host
410 sediment (Figs. 4 d-f).

411

412 **5.2 Geochemistry at Hole 269A**

413

414 Down-core variations of Ca content clearly track the 13 carbonate-cemented beds (Fig.
415 3d). In these beds, the dissolution of calcareous microfossils results in carbonate
416 cementation/re-precipitation, as shown by the HR-SEM analyses (Fig. 5h). Additionally, we
417 observe an interval of increased Ca values within the non-carbonate-cemented sediments
418 between 706.8-703.28 mbsf (~23.23 Ma) (Fig. 3d). This increase is also associated with the
419 preservation, even if scarce, of calcareous microfossils (i.e., planktonic, benthic foraminifers
420 and calcareous nannofossils, Fig. S4).

421 In comparison, Ti variations show negative excursions in the carbonate-cemented
422 facies (Fig. 3e). Ti decreases after 858 mbsf (~24 Ma), and more pronounced between 752.5
423 to 655 mbsf (~23.6 to 23 Ma) (Fig. 3e).

424 A principal component analyses (PCA) yielded one major principal component
425 (PC1_{TOT}) that explains 63.8 % of the total variance in the XRF data (Table S4). PC1_{TOT} is
426 characterised by negative loadings for Ca and positive loading for all the other elements (Al,
427 Si, K, Ti, Fe, Br, Ba, Sr, Zr) (Fig. 6a). Prominent negative excursions in PC1_{TOT} are associated
428 with the carbonate-cemented beds. In addition, PC1_{TOT} decreases after 752.5 mbsf (~23.6 Ma)
429 (Fig. S5).

430 Because of the distinct lithological/geochemical characteristics of carbonate-cemented
431 beds compared to the rest of the sediments in our record, we ran another PCA, excluding the
432 carbonate-cemented sediment. The first PC, named PC1 describes the 34.7% of the total
433 variance, with main negative loadings for Zr and Ca and in lesser degree Sr, Br and Si and
434 positive loading for all the other detrital elements (Al, Si, K, Ti, Fe, Ba, Rb) (Table S3; Fig.
435 6b). This pattern suggests accumulation of heavy elements (e.g., Zr) due to sediment sorting
436 likely by bottom currents.

437 Figure 6 near here.

438 Down-core variations of Zr/Ti ratios show high short-term variability (Fig. 3f). In
439 general Zr/Ti show similar patterns with PC1 (Fig. S5) supporting further that both proxies
440 reflect the accumulation of heavy minerals, due to the sorting of bottom currents and/or gravity
441 flows. Zr/Ti ratios show less pronounced variability in fine-grained sediments between 910.5
442 and 901 mbsf. In general, above 858 mbsf (~24 Ma), Zr/Ti ratios show higher values (Fig. 3f).
443 At 856.75 and 857.5 mbsf (~24 Ma) two prominent peaks of high Zr/Ti ratios, are recorded in
444 ripple cross-laminated sandy siltstones, which indicate the strongest episodes of bottom current
445 velocities within our study interval (Fig. S5b). In addition, high Zr/Ti ratios are shown between
446 752.5 to 750.17 mbsf (~23.6 Ma) (Fig. 3f).

447 Br/Ti ratios also increase slightly above 858 mbsf (~24 Ma). It is however between
448 752.5 and 655 mbsf (~23.6 Ma and 23 Ma) that ratios show increased values, which coincide
449 with the higher preservation of microfossils in the non-carbonate cemented sediment (Fig. 3g).

450 The $\epsilon_{Nd(t)}$ data from fish debris range from -9.03 ± 0.25 to -8.07 ± 0.21 (average $\epsilon_{Nd(t)}$
451 values -8.54 ± 0.22) during the late Oligocene (~24 to 23.23 Ma) (Fig. 3h).

452 Late Oligocene detrital sediment samples from Site 269 have $\epsilon_{Nd(t)}$ values of $-12.14 \pm$
453 0.33 and $\epsilon_{Nd(t)} = -13.33 \pm 0.33$, at 854.72 mbsf (~24 Ma) and 751.34 mbsf (~23.6 Ma),
454 respectively, and are within the range of local bedrock composition of proximal areas east of
455 the Metz glacier, within the Wilkes Subglacial Basin (Early Paleozoic granite outcrops; $\epsilon_{Nd} = -$
456 11.2 and -19.8 ; Bertram, et al., 2018; Cook et al., 2013; 2017).

457

458 **5.3 TEX₈₆**

459

460 Of the 15 samples processed, 4 were flagged as outliers with potential for a biased
461 source and thus they were not suitable for TEX₈₆ analysis. The remaining 11 samples had
462 normal values in the indices signaling overprints as mentioned in the methods. SST values
463 were calibrated between 9-14°C, ± 5.2 °C, with bioturbated facies characterized by higher SST
464 values, compared to the laminated facies (Fig. 3i). The relative temperature variability of 3-
465 5°C between laminated and bioturbated sediments, respectively at Hole 269A, is consistent
466 with sedimentological features, glacial-interglacial variability and TEX₈₆-derived SST reported
467 from Site U1356 (Hartman et al., 2018; Salabarnada et al., 2018).

468 **5.4 Palynology**

469

470 All examined samples yielded well to moderately preserved palynomorphs
471 assemblages with *Leiosphaera* (70-80%), in-situ dinocyst (20%) and pollen, while spores and
472 reworked dinocyst are a minor component (Table S2). As is common in high polar latitudes,
473 the absolute abundances of dinocysts remain low throughout the examined succession, and do
474 not exceed 400 cysts/g sediment. Assemblages are dominated by P-cysts, with *Brigantedinium*
475 spp. as most abundant (45-90% of the in situ dinocyst sum), *Lejeunecysta* (<10%) showing no
476 clear trend throughout the section and *Selenopemphix* with its largest abundance at between
477 955-909 mbsf (30%) (Fig. 3k). Even if less abundant (rarely >5%), the samples yielded a G-
478 cyst assemblage that mainly consists of *Batiacasphaera*, *Pyxidinosia*, *Operculodinium* and
479 *Impagidinium*. Similar G-cyst assemblages are present throughout the Oligocene-Miocene
480 record from the U1356 (Bijl et al., 2018b). G-cyst are notably common between 954.3-950.3
481 mbsf (~24.2 Ma) and around 23.6 Ma, with additions of *Spiniferites* spp. (sample 9R 3W 125-
482 129 cm) and *G. inflata* (sample 9R 2W 26-30 cm). Bioturbated facies contain more G-cyst
483 compared to the laminated facies (Fig. 3k).

484

485 **6 DISCUSSION**

486

487 **6.1 Glacial-interglacial sedimentation and short-term polar frontal system dynamics**

488

489 Repeated alternations between laminated and bioturbated facies like those described at
490 Hole 269A (Figs. 3c and 4), are common in deep-water settings around Antarctica and are
491 interpreted to result from changes in sedimentation related to glacial-interglacial cycles,
492 respectively (e.g., Hepp et al., 2006; Lucchi and Rebesco, 2007; Escutia et al., 2009, 2011;
493 Patterson et al., 2014; Salabarnada et al., 2018).

494 Terrigenous laminated deposits at Hole 269A (F1b, F2b, Figs. 3c and 4) are interpreted

495 to form by a complex interplay between sediments delivered by gravity flows (e.g., debris
496 flows and turbidity currents) and bottom currents, during glacial times. Even when laminated
497 deposits preserve turbidite affinity (e.g., sharp/erosional bases, normal grading depicted by
498 visual observations and Zr/Ti variability) (e.g., Figs. S6a and S6c), we find evidence suggesting
499 that these intervals were continuously reworked by bottom currents. This evidence include
500 internal structures such as mud drapes, double mud layers, lenticular laminations, mud-
501 offshoots and a rhythmic character between the intercalation of the muddy/sandy couplets with
502 varying thicknesses (e.g., Figs. 4b, e and k). These structures are typical in contourite deposits
503 and are interpreted to indicate both traction and suspension processes during deposition
504 (Shanmugam et al., 1993; Rebesco et al., 2014). Additionally, the absence or scarcity and low
505 diversity of trace fossils such as *Chondrites* and *Phycosiphon* in the laminated facies can
506 suggest a poorly oxygenated/ventilated environment at seafloor, changes in nutrient
507 availability and likely high sedimentation rates that promote unfavorable conditions for trace
508 markers to thrive (e.g., Lucchi and Rebesco, 2007; Rodríguez-Tovar and Dorador, 2014,
509 Rodríguez-Tovar et al., 2014; 2015a, b, 2019; Hodell et al., 2017). Poorly ventilated conditions
510 at the seafloor where also reported from late Oligocene laminated facies at Site U1356, during
511 glacial times (Salabarnada et al., 2018).

512 In comparison, bioturbated facies (F1a, F2a, Figs 3c and 4) are interpreted to result
513 from mainly hemipelagic sedimentation during interglacial times with continued reworking by
514 bottom currents. The inverse and bi-gradational grading patterns shown by Zr/Ti ratios and MS
515 variations, when resolution is sufficient (Fig. S6) is a strong evidence of bottom current control,
516 suggesting winnowing by bottom currents with fluctuating intensities (e.g., Stow and Faugères,
517 2008). The diverse trace fossil assemblage, indicate more oxygenated/better ventilated
518 conditions, likely higher nutrient content in the seafloor and lower sedimentation favorable to
519 microbenthic trace marker proliferation than during glacial times (e.g., Lucchi and Rebesco,

520 2007; Rodríguez-Tovar et al., 2015a, b). Late Oligocene bioturbated deposits from Site U1356
521 also record an increase in oxygenation at the seafloor when compared with laminated deposits
522 of the same age at the site (Salabarnada et al., 2018).

523 The carbonate-cemented laminated and bioturbated facies (facies: F3a, F3b; Figs 3c
524 and 5) are interpreted to result from sedimentation during the warmest interglacials recorded
525 by our sediments. This is indicated by the high Ca content (Fig. 3d) associated with dissolution
526 of calcareous microfossils, which strongly supports the biogenic origin of these carbonate beds
527 (Figs. 5f, 5g and 5h).

528 Our interpretation of sediments recovered at Site 269 from 956 to 858 mbsf (before 24
529 Ma) is supported by similar late Oligocene (26-25 Ma) bioturbated and laminated sediments
530 recovered at the more proximal IODP Site U1356. Nearly continuous recovery of this interval
531 at Site U1356, allowed for a detailed study of glacial-interglacial cyclicity that is paced by
532 obliquity (Salabarnada et al., 2018). These authors interpreted the changes between laminated
533 and bioturbated facies at Site U1356 to be driven by oceanic frontal migrations forced by
534 glacial-interglacial cycles. In detail, during interglacial times, southward migration of the
535 SWW and the surface oceanic fronts facilitated proto-CDW intrusions closer to the Antarctic
536 margin. This allowed better preservation of calcareous microfossils and enhanced ventilation
537 at the seafloor. The opposite occurred during the glacial times. In addition, even when the
538 interval from 26 and 25 Ma was generally warm as indicated by the prevalence of open water
539 conditions (Bijl et al., 2018), TEX₈₆-derived SST variations (between 1.5 and 3 °C) (Hartman
540 et al., 2018) correlate with the glacial-interglacial cyclicity reported by Salabarnada et al.
541 (2018). In fact, our more distal TEX₈₆-based SST data and dinocyst assemblages, compared
542 to those from Site U1356 (Hartman et al., 2018; Bijl et al., 2018b) support further the oceanic
543 frontal migrations forced by glacial-interglacial cycles. The relative SST variability of 3-5°C
544 at Site 269 may be slightly higher than the glacial-interglacial TEX₈₆-derived SST variability

545 from Site U1356 during the Oligocene-Miocene (Hartman et al., 2018), although a full
546 representation of the glacial-interglacial variability at either site might not have been captured.
547 Additionally, the high abundance of *Brigantedinium spp* (P-cyst) reflect open ocean upwelling
548 and high-nutrient conditions (Harland and Pudsey, 1999; Zonneveld et al., 2013) at Site 269,
549 perhaps even more than at U1356. In addition, the lower amounts of G-cysts compared to Site
550 U1356 indicate that Site 269 was located closer to the upwelling/divergence zone during the
551 glacial times. In contrast, SST increase and higher amounts of G-cysts (lower P/G cyst ratios)
552 similar to modern temperate (interglacial), oligotrophic waters from around Tasmania and
553 southern New Zealand (Prebble et al., 2013) indicate the southward migration of the oceanic
554 fronts during the interglacial times. Our results are consistent with dinocyst assemblages at Site
555 U1356, which show similar changes between oligotrophic, temperate dinocyst assemblages
556 during interglacials, to eutrophic, colder dinocysts during glacial times (Bijl et al., 2018b). The
557 difference in dinocyst assemblages between the two sites may be attributed to a closer
558 proximity of Site 269 to the Antarctic Divergence.

559 Deposition of large stacked debris flow deposits at Site U1356 between ~24.76 Ma and
560 23.23 Ma (Escutia et al., 2011; Passchier et al., 2018) provide limited information regarding
561 ocean configurations during the late Oligocene-early Miocene in this region of the East
562 Antarctic margin. This gap is recorded at Site 269 despite the limitations related to the
563 discontinuous drilling and low-resolution age model. Our sedimentological, geochemical and
564 palynological analyses show that the climate-related ocean polar frontal movement migrations
565 continued throughout the late Oligocene and into the early Miocene. We also note an increase
566 in the frequency of siltstones and sandstones beds between 24 and 23.23 Ma. These sediments
567 could result from local inputs from structural highs depicted nearby in seismic lines that cross
568 the site (De Santis et al., 2003). We note however that Oligocene and the Miocene deposition
569 around these highs is dominated by contourite deposition forming mounded deposits against

570 the highs (De Santis et al., 2003), which could prevent direct delivery of sediment to where
571 Site 269 is located. The correlations previously established between sedimentation at Sites 269
572 and U1356, argues for a regional source of sediments to Site 269 rather than a local one. We
573 therefore argue that these coarse sediments likely correspond to the distal reaches of the debris
574 flow deposits recovered at Site U1356. At this site these deposits have been interpreted to result
575 from ice sheet advances into the continental shelf (Escutia et al., 2011, 2014).

576

577 **6.2 Bottom water signatures at Site 269**

578

579 Geological evidence derived from microfossil assemblages from sedimentary records
580 around Antarctica, including evidence from Site 269, suggest that the proto-Polar Front (PF)
581 was placed close to 60°S between the late Oligocene-early Miocene (Nelson and Cooke, 2001;
582 Cooke et al., 2002). This is further supported by dinocyst assemblage data from Hole 269A
583 indicating that the site was influenced by nutrient-rich upwelling and ice-free waters, with
584 occasionally southward latitudinal transport of waters during the late Oligocene and early
585 Miocene. Fish debris neodymium isotope results from Site 269 (Fig. 3h) are consistent with
586 late Oligocene proto-CDW $\epsilon_{Nd(t)}$ values recorded along the proto-PF on the Kerguelen Plateau
587 (Indian Ocean) (average $\epsilon_{Nd(t)} = -7.8$; Wright et al., 2018) and around Maud Rise (Atlantic
588 Ocean) (average $\epsilon_{Nd(t)} = -8.5$; Scher and Martin, 2004). The combined dataset suggests that a
589 common bottom water mass (proto-CDW) was bathing the South Atlantic and South Indian
590 Ocean along the proto-PF. The pronounced difference between the $\epsilon_{Nd(t)}$ values of the fish
591 debris and the detrital sediment samples from Site 269 ($\epsilon_{Nd(t)} = -12.14 \pm 0.33$, -13.33 ± 0.33
592 (~ 24 Ma and ~ 23.6 Ma, respectively) confirm the water mass signal in the fish debris samples
593 (Fig. S7). Bathymetric reconstructions show that the Southern Indian Ocean basin was already
594 sufficiently deep during the Oligocene and did not contain any large topographic barriers that

595 prevented the flow of proto-CDW from the Kerguelen Plateau to the abyssal plain off the
596 Australian-Antarctic basin (Scotese and Wright, 2018). Based on the above, we conclude that
597 at least during the latest Oligocene (~24 to ~23.23 Ma), Site 269 was covered by a proto-CDW
598 with a Nd signature similar to the present day CDW in the Australian sector of Southern Ocean
599 ($\epsilon_{Nd} = -8.1$ to -9.1 ; Lambelet et al., 2018). This observation contrasts with the present-day
600 bottom water mass configuration at the location of Site 269, which is bathed by Adélie Land
601 Bottom Water (ALBW) (Rintoul, 1998), suggesting a reduced export of ALBW during the late
602 Oligocene. These results are consistent with previous inferences for reduced sea ice in the
603 region (Bijl et al., 2018b; Hartman et al., 2018) diminishing production of ALBW at the
604 Oligocene and Miocene on the Wilkes Land shelf. The less radiogenic $\epsilon_{Nd(t)}$ value around 24
605 Ma ($\epsilon_{Nd(t)} = -9.03 \pm 0.25$) may suggest mixing between proto-CDW and Adélie Coast Bottom
606 Water (ACBW; Oligocene ACBW $\epsilon_{Nd(t)} = -10.6 \pm 0.8$; Huck et al., 2017). However, more data
607 are needed to confirm this hypothesis.

608

609 **6.3 Long-term changes in the proto-ACC dynamics during the late Oligocene to earliest** 610 **Miocene**

611

612 We provide new insights into the ocean configuration during four distinct periods
613 between the late Oligocene-early Miocene from ~24.2 to 24 Ma, at ~24 Ma, from ~23.6 to
614 23.23 Ma, and at ~23 Ma.

615 A remarkable shift in sedimentation occurs at ~24 Ma (858 mbsf) (Fig. 3a). Before 24
616 Ma, predominantly fine-grained sediment deposited under enhanced terrigenous inputs and an
617 overall weak proto-CDW. This interpretation is supported by the high Ti values and low Zr/Ti
618 ratios, respectively, and their low variability, in particular between 910 to 901 mbsf (Figs. 3e
619 and 3f). Furthermore, less frequent bioturbated facies and low Br/Ti ratios suggest a less

620 ventilated environment and low nutrient content at the seafloor (Figs. 3b and 3g). The high
621 abundance of G-cyst assemblages between 954.3 to 950.3 mbsf (Fig. 3k), together with high
622 SST (11.5 to 14°C) (Fig. 3i), indicate that the polar front system was located south of the Site
623 269. However, the concomitant SST drop to 9°C and the absence of G-cysts at 948 mbsf (Figs
624 3i and 3k) indicate northward migration of the polar front, which likely reached Site 269.

625 Around 24 Ma, our data document the strongest proto-CDW velocities recorded at
626 Hole 269A during the late Oligocene and earliest Miocene, as indicated by the high Zr/Ti peaks
627 associated with cross-sandy laminated intervals that are not preserved elsewhere in the
628 sedimentary record (Figs. 3f and 4e).

629 After 24 Ma, there is an increase in the frequency of deposition of coarser-grained
630 material interbedded with fine-grained sediment (Fig. 3a). Overall Zr/Ti ratios are higher
631 suggesting periods of proto-CDW strengthening, with Zr/Ti peaks associated with Ca
632 enrichment (Figs. 3f and 6b). In addition, more abundant and diverse bioturbation (Figs. 3b
633 and 3c) point to episodes of better ventilated bottom conditions, and likely higher nutrient
634 content at the seafloor, as indicated by a small increase in Br/Ti ratios (Fig. 3g). Between ~24
635 and 23.6 Ma, the frontal system migrated northward, as indicated by the lower SST (9.8-
636 10.8°C) and the dominance of P-cysts, suggesting cool and eutrophic ocean conditions. In
637 addition, the sheer absence of dinocysts may have resulted from low preservation associated to
638 oxidation at sea floor, transport by bottom currents and/or reduced productivity. Given the
639 warmth as reconstructed for the surrounding time intervals (e.g., Bijl et al., 2018b; Hartman et
640 al., 2018), we rule out a permanent ice cover during this time.

641 Figure 7 near here.

642 We infer that the observed shift in depositional environment at 24 Ma was driven by
643 atmospheric and oceanic frontal changes. SWW northward migration due to the major
644 expansion of the AIS between 24.5 and 24 Ma (see Levy et al., 2019 for discussion) is thought

645 to have forced the northward migration of the SWW prior to 24 Ma (Fig. 7a and 7c). This
646 would have resulted in a weak proto-CDW, reduced water mass mixing, and a less ventilated
647 seafloor. In contrast, the increase in proto-CDW intensification after 24 Ma, inferred from the
648 higher Zr/Ti amplitudes, likely indicate times when SWW migrated southwards, aligned with
649 the proto-ACC. This promoted enhanced mixing and better ventilation of the seafloor and
650 likely higher organic matter deposition, in times when the AIS retreated. Our SST and dinocyst
651 data further support the warmer temperatures and dominance of G-cysts, which indicate the
652 stronger influence of warm oligotrophic waters.

653 Between ~23.6 and ~23.23 Ma, sediments record a prominent southward polar frontal
654 system migration, which allowed warmer water to penetrate further southwards (Fig. 7a and
655 7c). This interval is characterized by high Zr/Ti ratios (at ~23.6 Ma, Fig. 3f), high Ca, high
656 SST (11.5-12.9°C) (Fig. 3i) and a reduction in terrigenous input (low Ti values) (Figs. 3d and
657 3e). In addition, there is higher preservation of calcareous microfossils in some intervals (Fig.
658 S2), and dominance of G-cysts (Fig. 3k), similar to those in modern temperate (interglacial),
659 oligotrophic waters from around Tasmania and southern New Zealand (Prebble et al., 2013),
660 high Br/Ti ratios (Fig. 3g) and a thick carbonate cemented bed at ~23.23 Ma (Fig. 3a). Today,
661 calcareous organisms rarely reach the seafloor within and south of the polar front zone. This is
662 because of the presence of the corrosive (CO₂-rich) Antarctic deep waters (Whitehead and
663 Bohaty 2003 and references therein) and strong upwelling (Olbers et al, 2004), which dissolve
664 calcareous rain. A thick carbonate cemented bed is present at Site U1356 dated at 23.23 Ma
665 (Escutia, et al., 2011). The synchronous deposition of these carbonate-cemented beds at Sites
666 U1356 and 269 strongly supports a wide southward expansion of proto-CDW offshore Wilkes
667 Land at ~23.23 Ma. Our ϵ_{Nd} data, despite their low resolution, also suggest a greater influence
668 of proto-CDW during this period compared to ~24 Ma (Fig. 3h). Similar processes occur
669 today, when a reduction in the volume of AABW is compensated by the expansion of the

670 CDW (van Wijk and Rintoul, 2014). Our interpretation is further supported by the absence of
671 sea-ice-related dinocyst species (i.e. *Selenopemphix antarctica*) at both Sites 269 and U1356
672 (Bijl et al., 2018b), which suggests a weaker than modern sea ice season during the Oligocene
673 and Miocene. In addition, the new SST reconstruction at Site 269, similar to those at Site
674 U1356, report warmer surface water conditions than today and argue for a decrease in the
675 potential formation of Antarctic bottom waters (Hartman et al, 2018).

676 At ~23 Ma, high Br/Ti ratios indicate an increase in the organic content in the sediment
677 (Fig. 3), coinciding with the first evidence (scarce) of diatoms at Site 269. A shift from
678 calcareous dominated microfossils to siliceous (e.g., diatomaceous and cherty clay sediments)
679 has also been reported at Site U1356 (Escutia et al., 2011, Escutia et al., 2014; Passchier et al.,
680 2018) arguing for a major regional event likely related to the AIS expansion during the early
681 Miocene, northward expansion of the polar front system and more influence of siliceous
682 productivity. This is supported by the high amount of P-cyst at Site 269 at ~23 Ma, which
683 indicate cool, nutrient-rich upwelling conditions (Fig 3k). After 23 Ma, deposition of the
684 thickest carbonate cemented bed would imply this period was followed by a southward
685 migration of the frontal system (Fig. 3a).

686 In summary, our combined sedimentological, geochemical and palynological data
687 between ~24.2 and ~23 Ma show a dynamic proto-ACC off the Wilkes Land margin during
688 the late Oligocene-Miocene. In addition, the carbonate preservation and low siliceous
689 microfossil preservation at Site 269 (located within the polar front zone) contrasts with modern
690 sedimentation near the Polar Front. This is supported by the presence of oligotrophic, temperate
691 dinocyst assemblages at Sites 269 and U1356 that are similar to the ones found today north of
692 the Polar Front around Tasmania and the Southern New Zealand (Bijl et al., 2018b). Our data
693 therefore suggest a weaker frontal system, characterized by reduced upwelling, which probably
694 allowed southward transport of warm surface waters from lower latitude. This is consistent

695 with numerical modelling results indicating a weaker than present-day proto-ACC (Herold et
696 al., 2012; Hill et al, 2013). These findings argue against the onset of a modern-like ACC during
697 the latest Oligocene (ca 25-23 Ma) based on sedimentary records across the Tasman Gateway
698 (Pfuhl and McCave, 2005) and the South Pacific (Lyle et al., 2007).

699

700 **7. Conclusions**

701

702 Our integrated sedimentological, geochemical, isotopic and palynological data sets from DSDP
703 Site 269 provide new insights into the proto-ACC dynamics during the late Oligocene-early
704 Miocene (~24.2 to 23 Ma) off the eastern Wilkes Land margin. We show that sedimentation at
705 Site 269 is controlled by the persistent reworking of gravity flows and hemipelagic
706 sedimentation by proto-CDW that is characterised by fluctuating current intensities driven by
707 the migration of the frontal system in response to climatic changes. We detail four distinct
708 snapshots (from ~24.2 to 24 Ma, at ~24 Ma, from ~23.6 to ~23.23 Ma, and at ~23 Ma) that we
709 link to changes in the proto-ACC and deep proto-CDW dynamics. Just before 24 Ma, fine-
710 grained sediments were deposited under enhanced terrigenous inputs and weak proto-CDW
711 intensities that resulted in low ventilated bottom conditions and probably low organic content
712 and preservation. At 24 Ma, episodic events of stronger proto-CDW current velocities started,
713 associated with coarser-grained deposits, and better ventilated bottom conditions and slightly
714 higher organic matter content. In addition, TEX₈₆-derived SST data varied between 9 to 13.5°C,
715 while the dominance of P -cysts indicate relatively cool ocean temperatures, upwelling and
716 high-nutrient conditions between ~24 and 23.6 Ma. Together, these evidences suggest that the
717 polar front at this time was located near the site. At ~23.6 Ma, and more pronounced at ~23.23
718 Ma, a prolonged expansion of proto-CDW closer to the Wilkes Land margin is indicated by
719 the higher Ca values, better preservation of calcareous microfossils, higher Br/Ti ratios, high

720 SST (from 11.5 to 12.9°C), high amounts of G-cysts similar to modern temperate and
721 oligotrophic waters, and ϵ_{Nd} data resembling modern-like CDW ϵ_{Nd} signature at Site 269. Given
722 that the record between ~24.9 and 23.23 Ma was masked by debris flow deposition at the more
723 proximal IODP Site U1356, our results provide the first record of ocean configuration for this
724 margin for this time interval. When compared with results from Site U1356 and with numerical
725 modelling, our findings support the notion of a fundamentally different Southern Ocean with a
726 weaker proto-ACC than today during the late Oligocene and earliest Miocene. A weaker frontal
727 system permitted the incursion of warm waters from lower latitude closer to the Antarctic
728 margin and the preservation of carbonate. The synchronous deposition of thick carbonate-
729 cemented beds both at Site 269 and U1356 at 23.23 Ma indicates a regional event of poleward
730 proto-CDW expansion.

731

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747

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1079

1080 **List of Figures**

1081 **Figure 1.** Oceanic frontal system between Antarctica and Australia. Dark Blue dashed lines
1082 schematically represent oceanic fronts today (Sokolov and Rintoul, 2009). Polar water south
1083 of the PF are shaded in darker blue. Map derived from Gplates software (Müller et al., 2018).
1084 Site 269 is marked with yellow symbol and Site U1356 is marked with red symbol. White lines
1085 indicates the continental lithosphere boundary. Black arrows show the pathway of the ACC
1086 today (Rintoul et al., 2001). STF: Subtropical front, SAF: Subantarctic front, PF: Polar Front,
1087 SAACF: Southern ACC front, ACC: Antarctic Circumpolar Current.

1088

1089 **Figure 2.** Age model for Deep Sea Drilling Project Hole 269A based on magnetostratigraphy
1090 constrained by dinocyst and calcareous nannofossil biostratigraphy. **LO:** Last Occurrence. **P:**
1091 Presence. Age model has been calibrated to GTS2012 of Gradstein et al. (2012).

1092

1093 **Figure 3.** Sedimentological, paleontological and geochemical (XRF scanning data and $\epsilon_{Nd(t)}$)
1094 data of DSDP Hole 269A, all plotted versus depth **a:** Graphic lithological log, **b:** sedimentary
1095 structures (see legend), **c:** main facies distribution (see legend), **d:** total Ca counts, **e:** total Ti
1096 counts, **f:** Zr/Ti ratios, **g:** Br/Ti ratios, **h:** ϵ_{Nd} , **i:** TEX₈₆-derived sea surface temperature **k:** %P-
1097 cyst=P-cyst/(P-cyst + G-cyst). Colors of TEX₈₆-derived SST and P-cyst vs G-cyst values reflect
1098 the sedimentary facies. Note core gaps between the cores were removed from the plot (gaps
1099 indicated with //).

1100

1101 **Figure 4.** Representative core photographs from facies assemblages in DSDP Hole 269A,
1102 showing main structures and bed contacts referred to in the text. Facies colour code in the
1103 sidebar to the right of each core image is according to the legend in Figure 3. ft: faint

1104 laminations, pl: planar laminations, wa: wavy laminations, *dm*: double mud layers, rp: ripple,
1105 cr: cross-laminations, mo: mud-offshoots, sd: soft-sedimentary structures, m: mottled, *Ch*:
1106 *Chondrites*, *Pl*: *Planolites*, *Th*: *Thalassinoides*, *Scolicia*: *Sc*. Note single granule in core
1107 photos: g and k. Numbers below core photos show core section and depth (mbsf).

1108

1109 **Figure 5.** Detailed images of carbonate-cemented facies. **a:** A sharp erosive contact (sc-e)
1110 between bioturbated (F3b) and laminated (F3a) carbonate-cemented facies. **b, c, d:** A
1111 carbonate-cemented bioturbated and laminated facies. **e:** Back-scattered electron
1112 photomicrograph showing blow-up example of a sharp erosive contact (sc-e) between
1113 bioturbated and laminated carbonate-cemented facies (red dotted line) and inverse grading
1114 pattern (yellow arrow) above the contact . **f:** Thin section photomicrograph, plain-polarized
1115 light of carbonate-cemented facies showing diatom assemblages within the laminae of facies
1116 F3a. **g:** HR-SEM micrograph and corresponding elemental map of Al, Ca, Si showing diatom
1117 skeletal remains within the carbonate cementation matrix. **h:** foraminifera with siliceous tests.
1118 *Ch*: *Chondrites*, *Pl*: *Planolites*, *Th*: *Thalassinoides*, pl: planar laminations, rp: ripple, cr: cross-
1119 laminations.

1120

1121 **Figure 6.** Principal Component Analysis (PCA) between XRF-scanner data through the study
1122 core section. **a:** PC1_{TOT} with carbonate-cemented facies and **b:** PC1 without carbonate-
1123 cemented facies.

1124

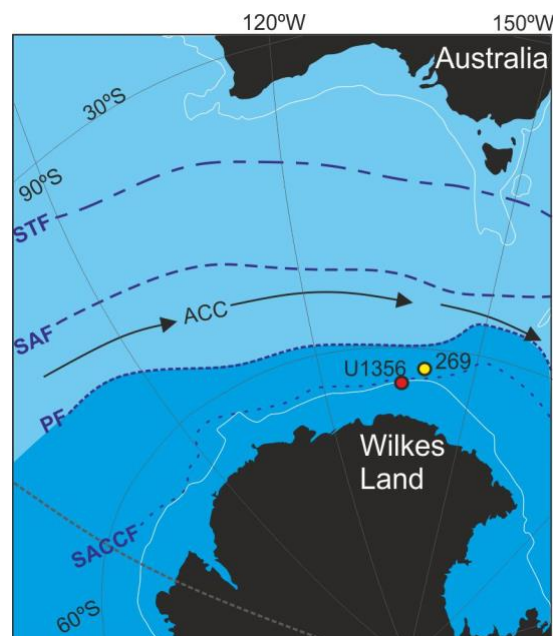
1125

1126 **Figure 7.** Paleoceanographic configuration offshore the Wilkes Land margin around 24-23
1127 Ma. **a:** Reconstructed ocean frontal system between Antarctica and Australia around 24-23 Ma.
1128 Reconstruction of tectonic plates around Tasman Gateway and paleoposition of sites based on

1129 Seton et al., 2012 global plate motion model derived from Gplates software (Müller et al.,
1130 2018). White lines indicate the continental lithosphere boundary. Polar water south of the PF
1131 are shaded in darker blue. Boundary between polar and subantarctic waters during the 23.23
1132 Ma (red dashed line) and during the 24 Ma (yellow dashed line). Frontal constraints west and
1133 east of the Wilkes Land region are from reconstructions of Nelson and Cooke (2001); Cooke
1134 et al. (2002). STF: Subtropical front, SAF Subantarctic front, PF: Polar front. Black arrows
1135 show the pathway of the proto-ACC. Schematic illustrations of proto-CDW dynamics for the
1136 24 Ma (b) and 23.23 Ma (c). **b:** At ~ 24 Ma, Westerlies and PF were located close to Site 269
1137 resulting strong proto-ACC. There is likely enhanced proto-AABW production. **c:** At ~ 23.23
1138 Ma, Westerlies and PF migrated southwards close to Site U1356. Proto-ABBW formation is
1139 reduced. This allowed the relative warmer surface water and proto-CDW to penetrate closer to
1140 Antarctic continent.

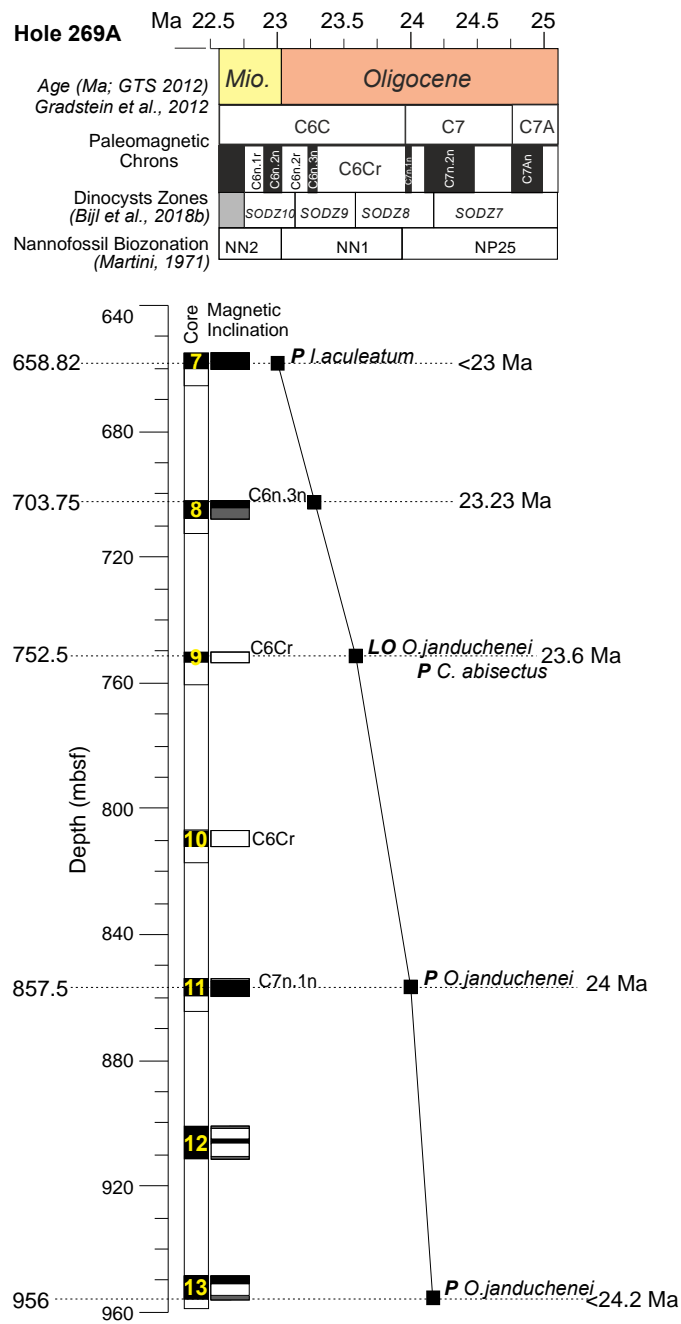
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1143 **Figures**



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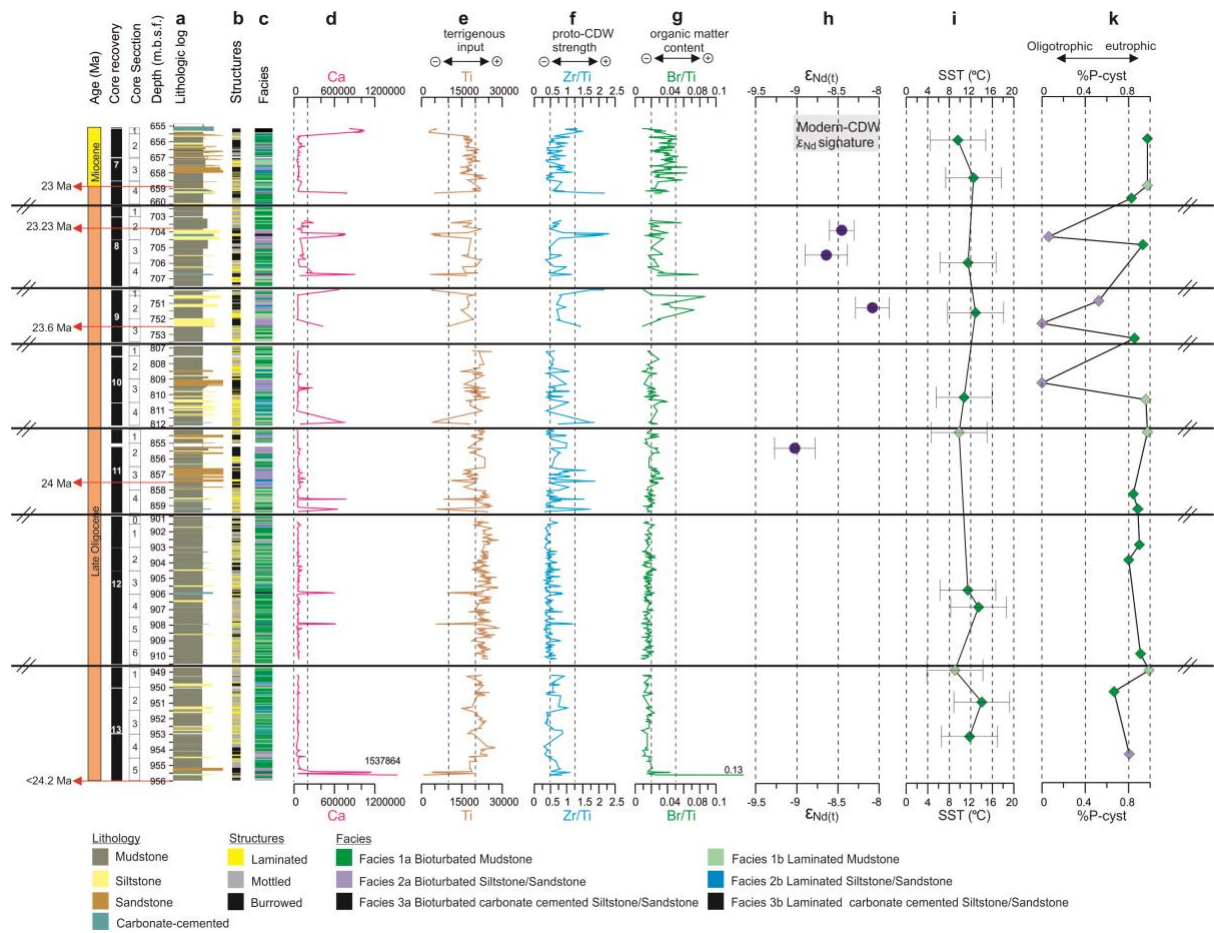
Figure 1- in colour



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Figure 2- in colour



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Figure 3- in colour

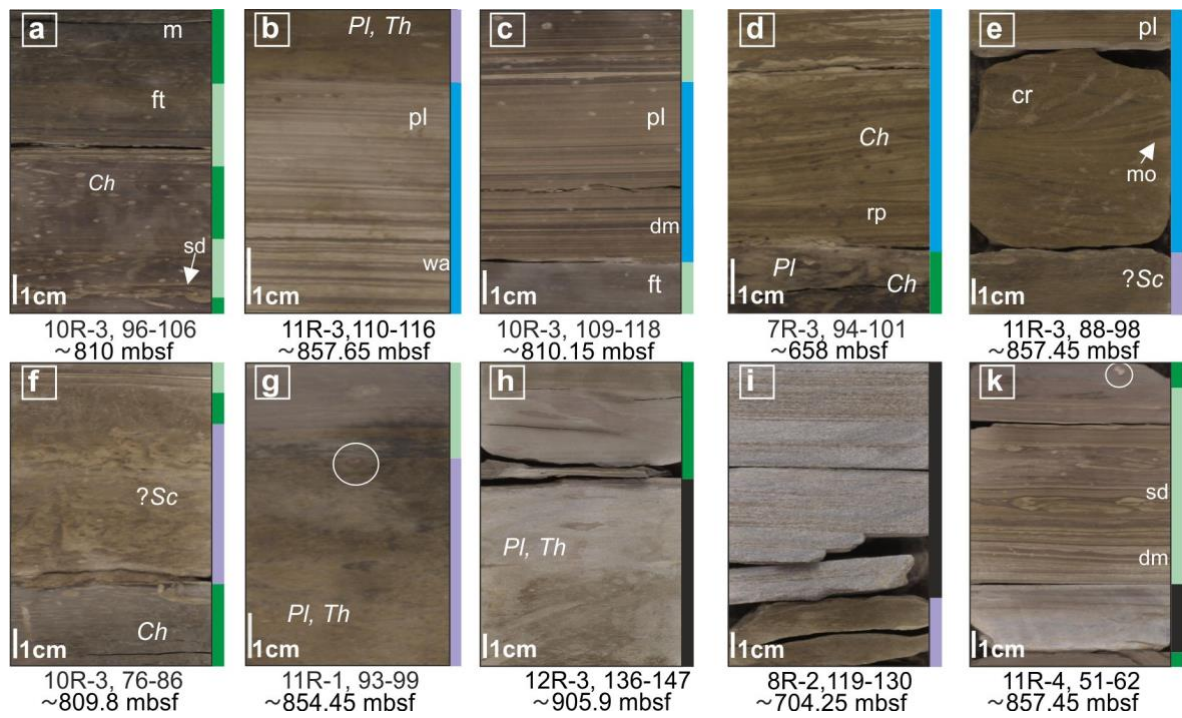


Figure 4- in colour

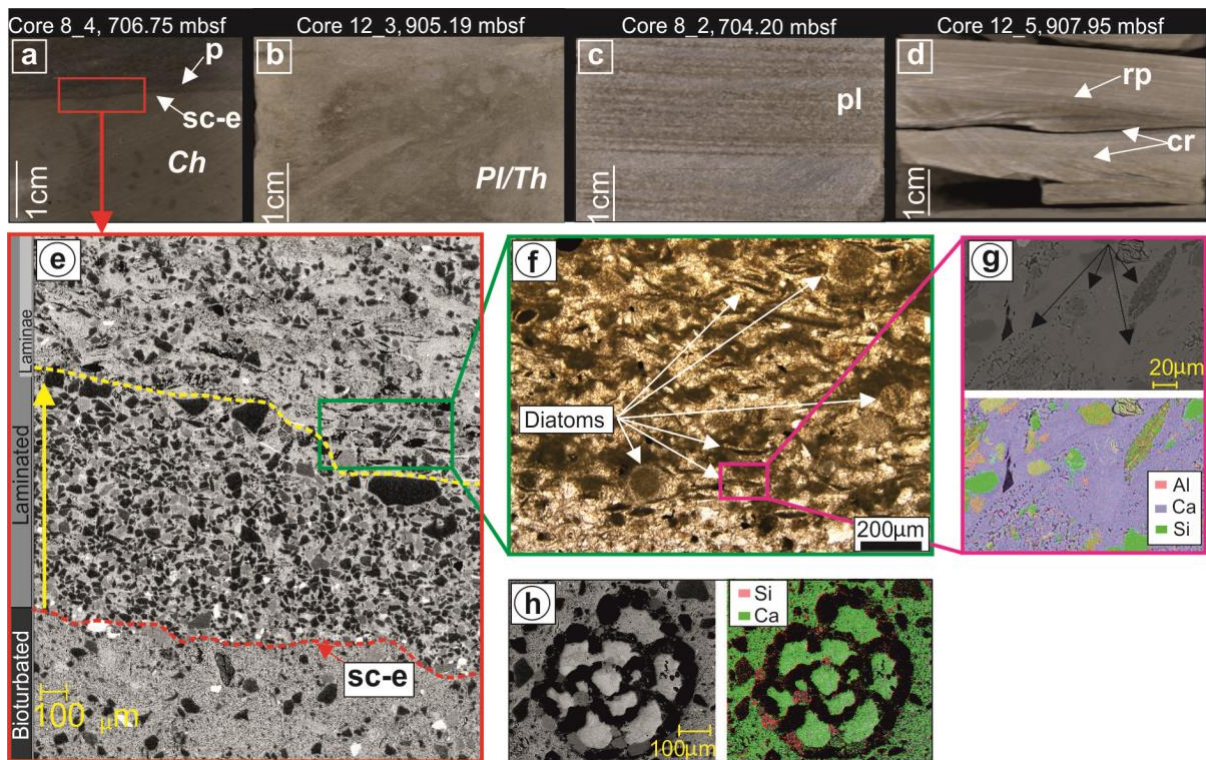
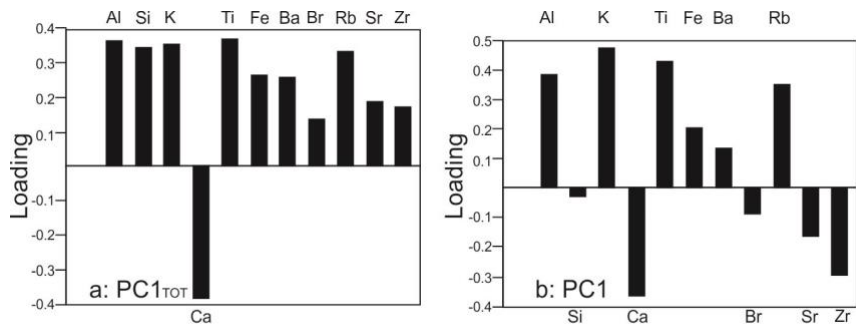


Figure 5-in colour

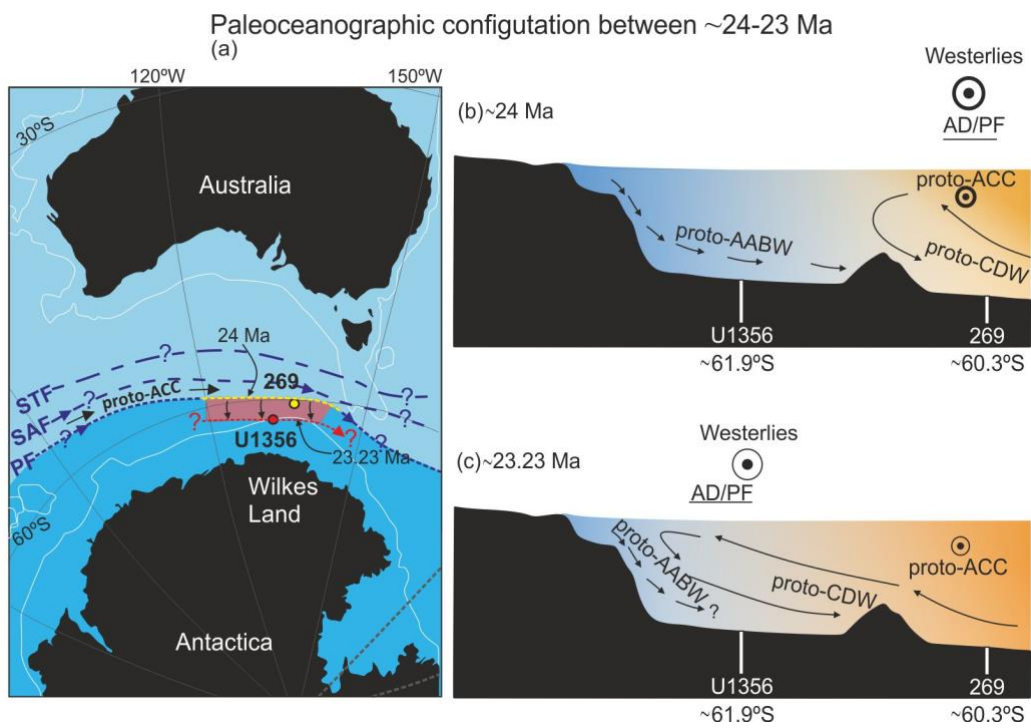


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Figure 6-black and white



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Figure 7-in colour

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1178 **Supplementary Material**

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1180 **Supplementary text**

1181

1182 **Fossil fish tooth sample preparation for Neodymium isotopes**

1183 Fish debris was handpicked from the >63 μm sediment fractions that were prepared by
1184 wet sieving. Four samples were prepared for fish-tooth and bone debris Nd isotope analyses.
1185 All samples were treated with ultraclean 18 M Ω water (MQ water) and methanol-following
1186 Martin and Haley (2000) to remove debris from surfaces and cavities in the MAGIC
1187 laboratories at Imperial College London (see also Huck et al., 2017). Cleaned fish tooth
1188 samples were subsequently transferred into cleaned microcentrifuge tubes and dissolved
1189 overnight in 50 μL of 2M HCl. Dissolved fish debris were loaded on Biorad cation exchange
1190 resin (200-400 μm mesh) to separate the REEs from the sample matrix and Eichrom Ln-Spec
1191 resin (50-100 μm bead size) to separate Nd from the other REEs.

1192

1193 **Bulk sediment sample preparation for Neodymium isotopes**

1194 In order to consider a potential influence of the detrital sediment towards the Nd isotope
1195 composition of pore waters or overlying bottom water, we also measured the Nd isotope
1196 compositions of two detrital sediment samples. Samples were dried and gently homogenised
1197 using mortar and pestle. Approximately 500 mg of homogenised material was subjected to a
1198 carbonate leaching procedure to remove biogenic carbonate using 30 ml of 1.5% buffered
1199 acetic acid (see Biscaye (1965) for more details). Ferromanganese oxides and oxyhydroxides
1200 were leached using a weak reductive solution of 0.005 M Hydroxylamine Hydrochloride
1201 (NH_2OH), followed by a stronger solution of 0.05 M NH_2OH . 50 mg of leached detrital
1202 sediment sample was subsequently digested on a hotplate using a mixture of 1ml of
1203 concentrated HNO_3 , 0.8ml HClO_4 and 2 ml HF. The detrital samples were processed using the
1204 same ion chromatography as the fish debris.

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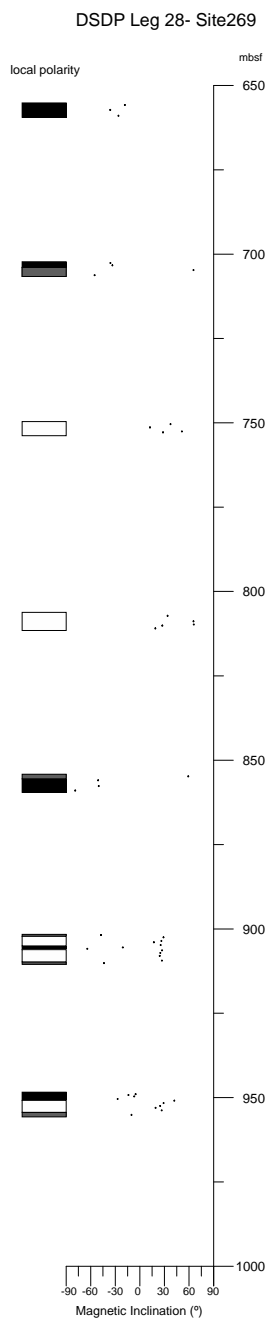
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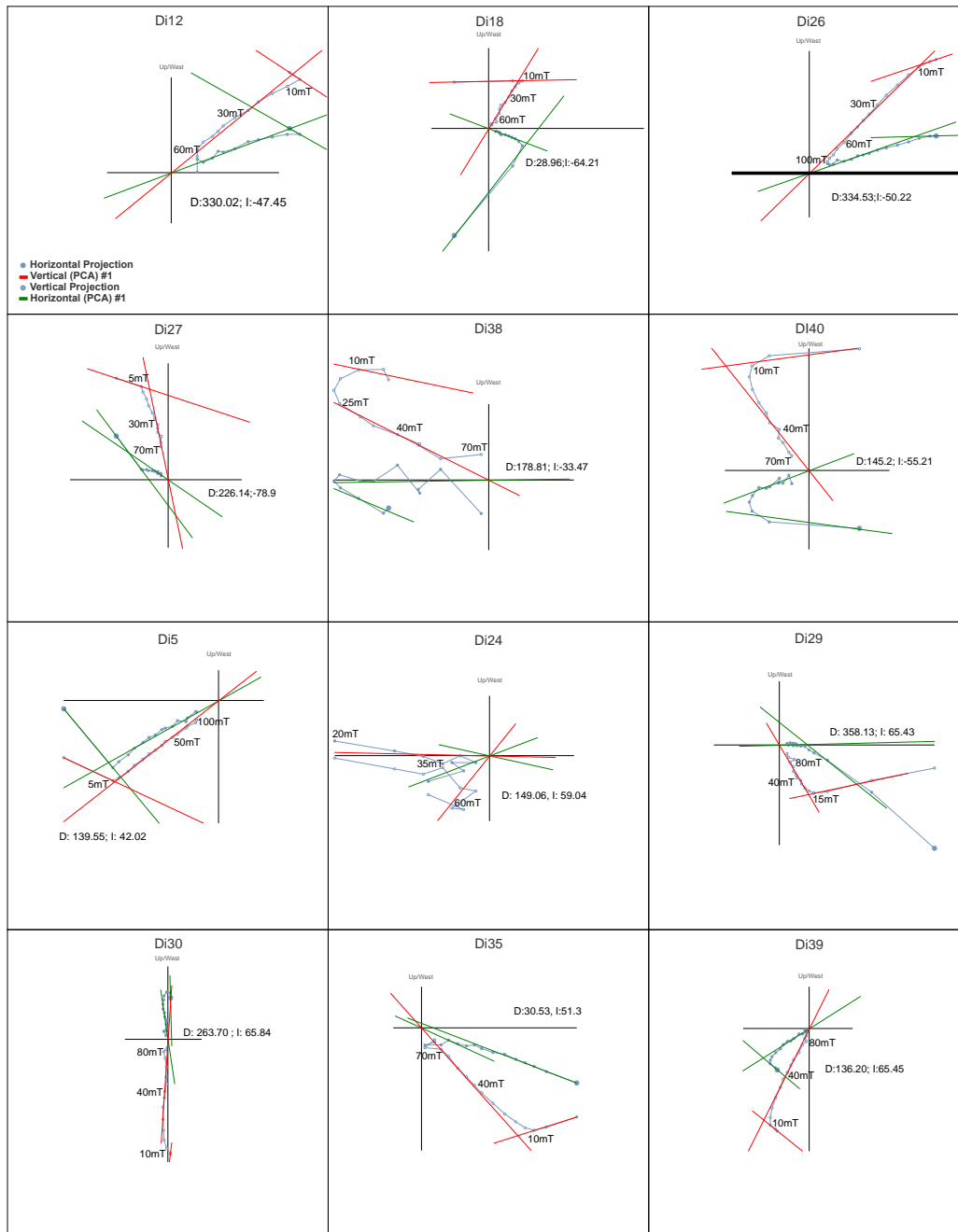
1211 **Supplementary figures**

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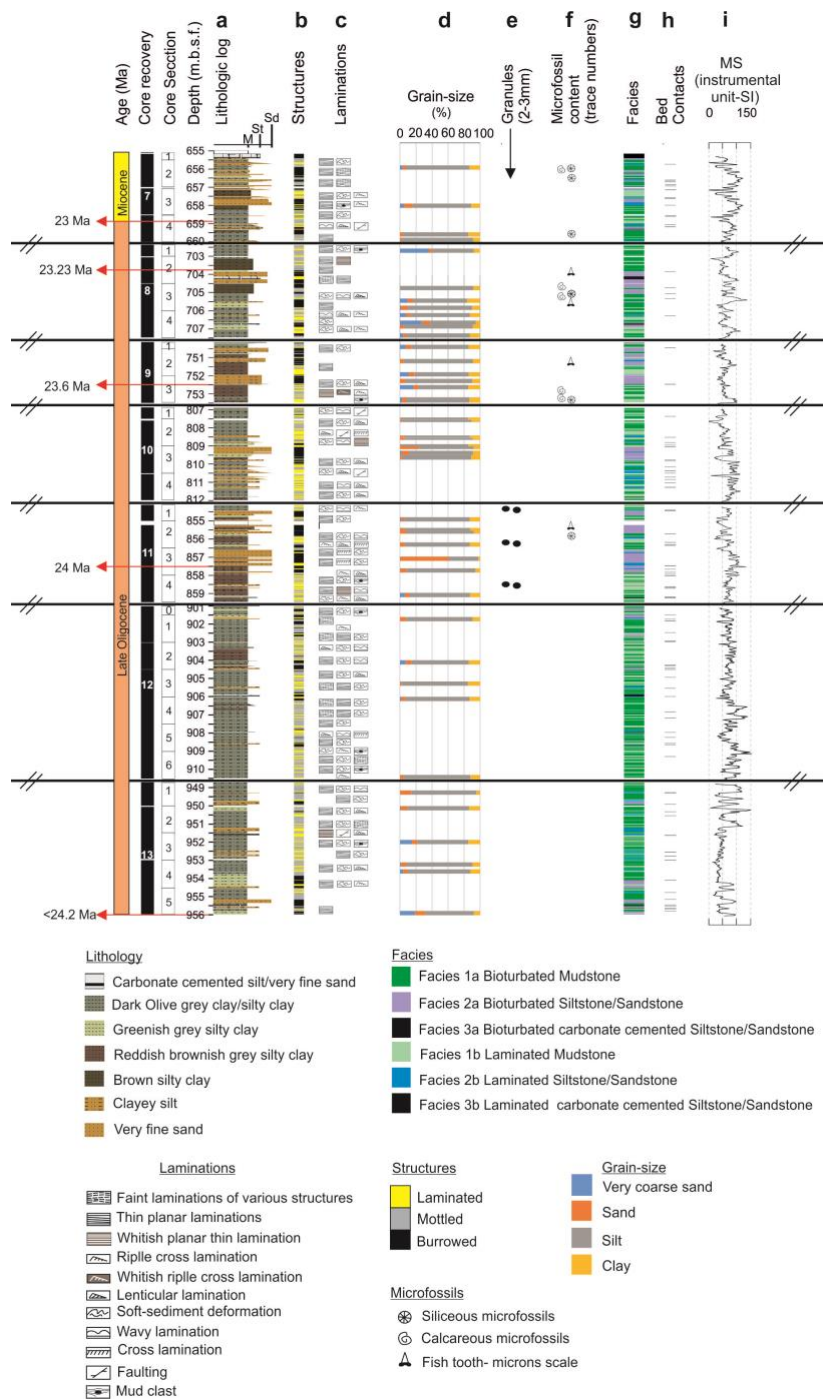
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1214 **Figure S1.** Stratigraphic location of the calculated Inclinations for the Site 269. A local
1215 polarity, where negative (positive) inclinations indicate normal (reversed) polarity in black
1216 (white), is also provided. Magnetozones are defined by more than one sample, otherwise they
1217 are shown in grey.



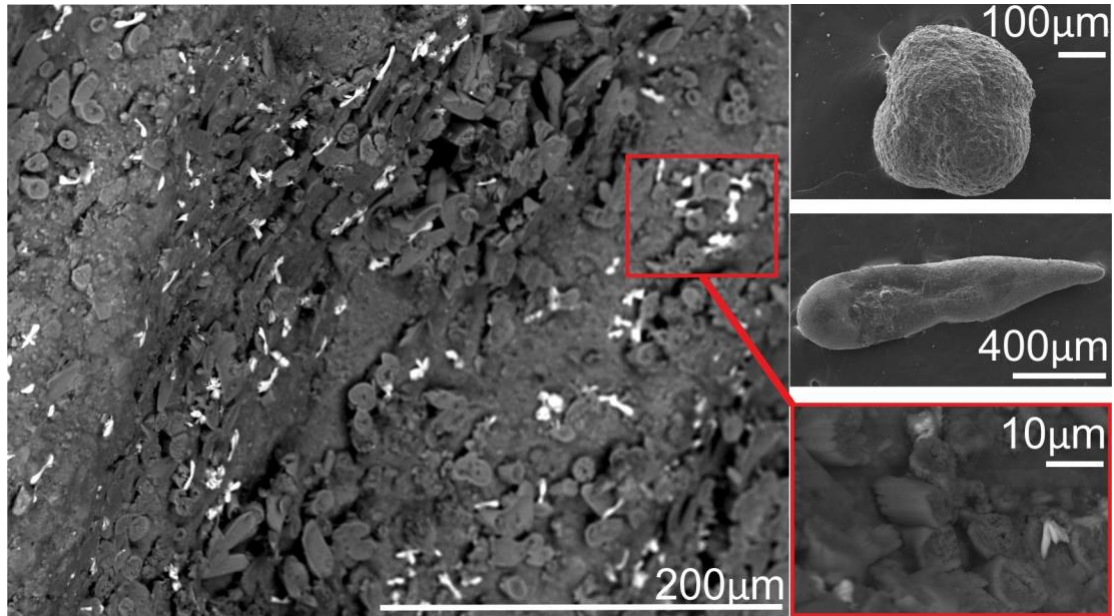
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1219 **Figure S2.** Orthogonal plots of the representative samples, showing two distinctive directions,
 1220 both in normal samples and in reversed samples. Inclination values are also indicated. Open
 1221 plots indicate inclinations (vertical projection). All calculated directions are available in Table
 1222 S1. Samples were calculated by means of the Paldir and paleomagnetism.org (Koymans et al.,
 1223 2016) programs.



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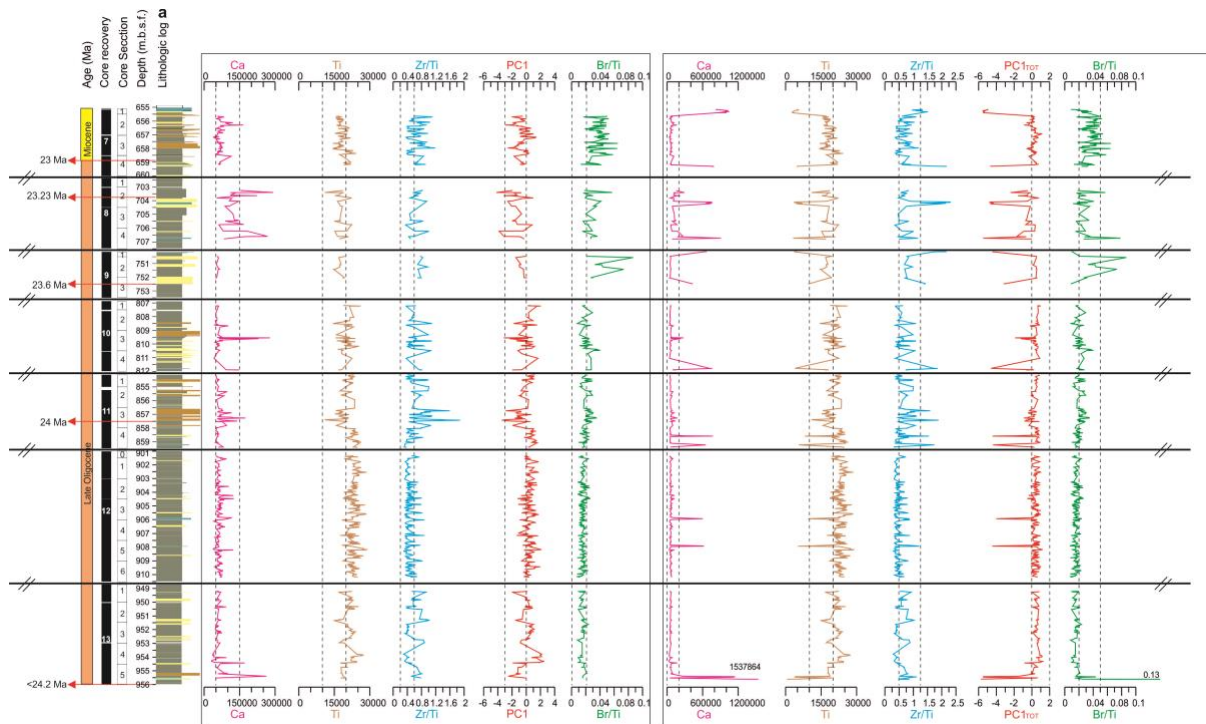
1227 **Figure S3. a:** Detailed lithological log, **b:** structures, **c:** types of laminations, **d:** grain-size
1228 data, **e:** presence of granules 2-3mm, **f:** visually observed microfossils (trace amounts), **g:** main
1229 facies distribution, **h:** sharp contacts between facies, **i:** magnetic susceptibility (MS). Note core
1230 gaps between different cores due to discontinuous coring (//). M: Mudstone, St: Siltstone: Sd
1231 Sandstone.



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1234 **Figure S4.** HR-SEM image showing calcareous microfossils associated with barite and
1235 representative species of foraminifers found in cores 8R and 9R.

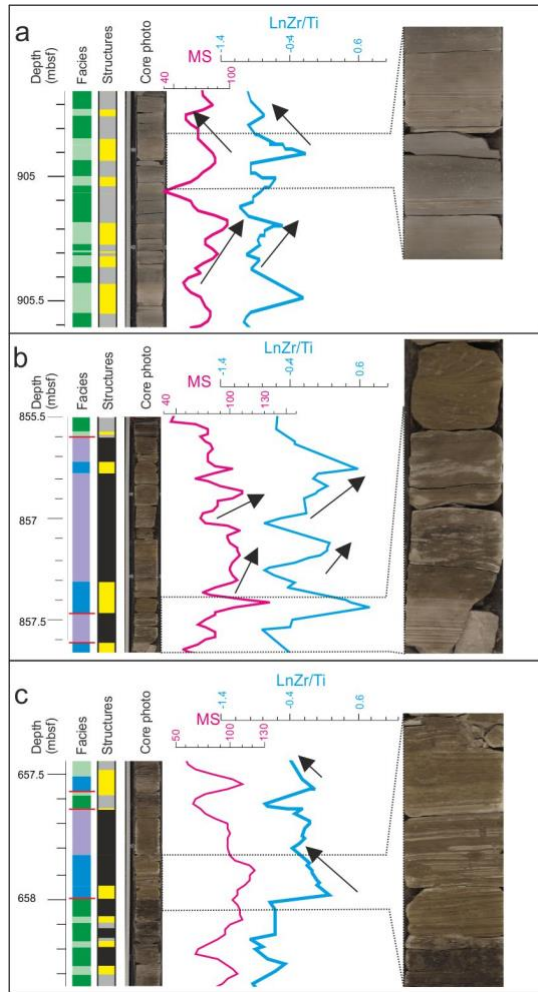
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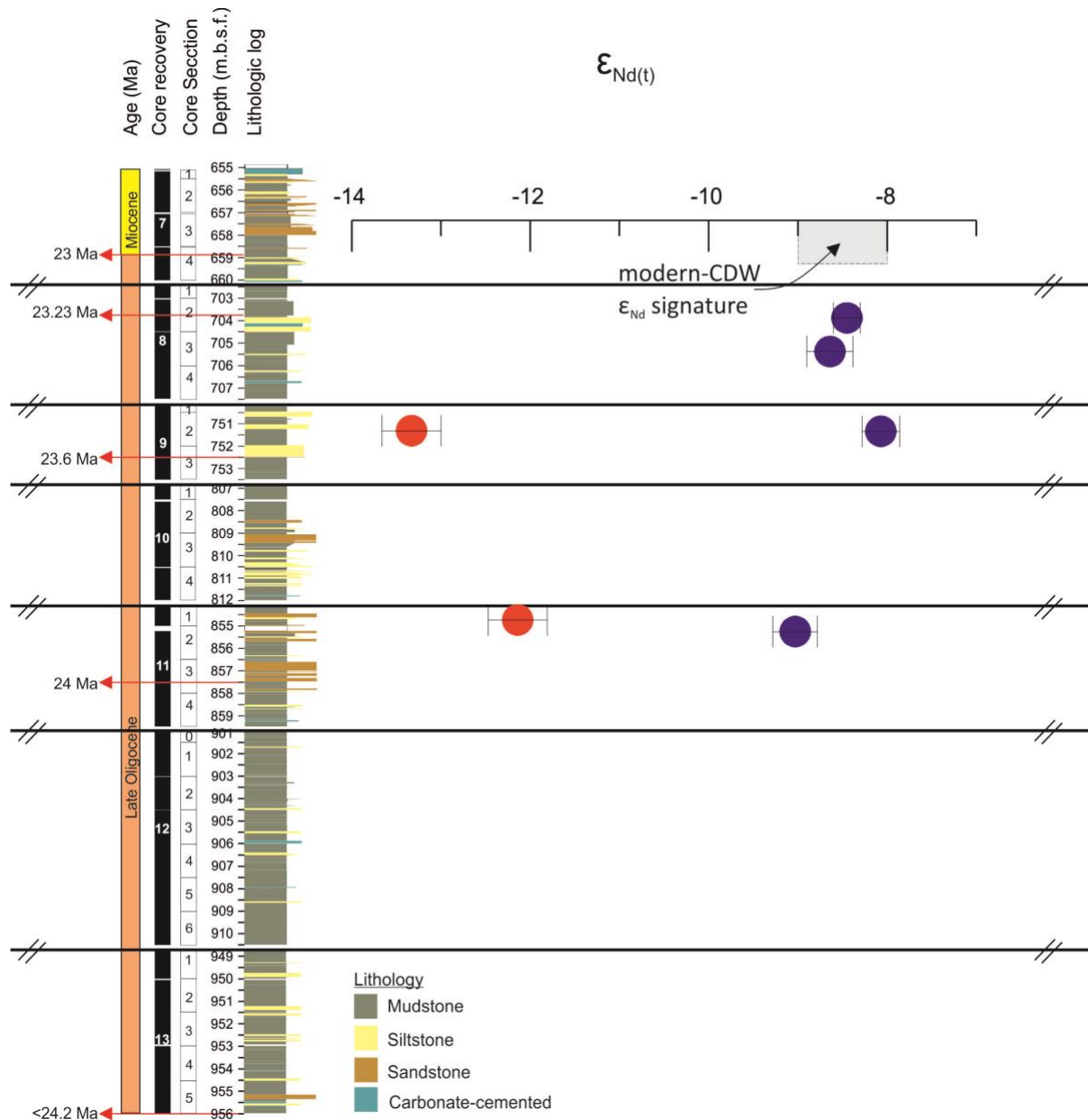
1239 **Figure S5.** Lithologic log (a) and down-core variations of (i) selected XRF-scanning data (in
1240 total counts), elemental ratios and PC1 without the carbonate-cemented facies (left) and (ii)
1241 selected XRF-scanning data, elemental ratios and PC1_{tot} with the carbonate-cemented facies
1242 (right). Note core gaps between different cores due to discontinuous coring (//).

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Figure S6. Representative examples of facies associations, including structures, magnetic susceptibility (MS) and Zr/Ti data. **a:** Fine-grained facies associations. **b** and **c:** Coarse-grained facies associations. Arrows indicate inverse and normal grading patterns. Red lines between facies indicate sharp bed contacts. Facies colour and structure codes are given in Figure 3.



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1255 **Figure S7.** Neodymium isotopic composition of fish debris (blue) and bulk detrital sediment
1256 ϵ_{Nd} values (red) from Hole 269A

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Table S1. Paleomagnetic dataset at Hole 269A Cores 7R-13R

Leg	Site	Hole	Core	Section	Top (cm)	Bot (cm)	Depth (mbsf)	Sample Name	Declination	Inclination	Intensity (A/m ²)	Strat	MAD	Forced	Type	Coordinates	Bedding Strike	Bedding Dip	Num Step	Min Step	Max Step
28	269	A	13	1	53	55	949.03	d11	53.76352373	-5.199547693	12816.60013	7	4.408449078	TRUE	dir	Geographic Coordinates	90	0	5	15	35
28	269	A	13	1	76	78	949.26	d12	187.5173705	-13.90078791	6251.883574	6	3.301610054	TRUE	dir	Geographic Coordinates	90	0	11	25	100
28	269	A	13	1	114	116	949.64	d13	140.0669862	-6.894508896	6266.399988	5	1.729861671	TRUE	dir	Geographic Coordinates	90	0	12	15	90
28	269	A	13	2	44	46	950.44	d14	62.01754422	-27.06342913	16272.10342	4	1.777811659	TRUE	dir	Geographic Coordinates	90	0	12	20	100
28	269	A	13	2	95	98	950.95	d15	139.5556692	42.02077811	14790.02487	3	1.980377917	TRUE	dir	Geographic Coordinates	90	0	14	10	100
28	269	A	13	3	13	15	951.63	d16	99.02692451	28.921119504	9137.591091	2	2.41338962	TRUE	dir	Geographic Coordinates	90	0	8	25	70
28	269	A	13	3	104	106	952.54	d17	182.0068096	24.78570975	11344.96448	1	1.967135126	TRUE	dir	Geographic Coordinates	90	0	12	15	90
28	269	A	13	4	3	5	953.03	d18	222.6946459	19.20222119	5656.079284	8	2.10642605	TRUE	dir	Geographic Coordinates	90	0	8	30	90
28	269	A	13	4	82	84	953.82	d19	35.35353282	26.65869667	4713.510042	7	4.291701509	TRUE	dir	Geographic Coordinates	90	0	8	30	80
28	269	A	13	5	67	69	955.17	d10	43.23591546	-10.24051787	1349.560528	6	13.32826152	TRUE	dir	Geographic Coordinates	90	0	6	5	30
28	269	A	12	cc	108	110		d11	98.56922818	42.67535279	23715.78417	5	2.340095871	TRUE	dir	Geographic Coordinates	90	0	9	20	70
28	269	A	12	1	31	33	901.81	d12	330.0166967	-47.4554181	9962.928043	4	2.367538165	TRUE	dir	Geographic Coordinates	90	0	10	10	60
28	269	A	12	1	100	105	902.5	d13	308.9936867	28.80580856	4401.981022	3	22.37747839	TRUE	dir	Geographic Coordinates	90	0	7	30	70
28	269	A	12	2	55	58	903.55	d14	255.5246598	26.2429681	21456.26987	2	3.458319655	TRUE	dir	Geographic Coordinates	90	0	10	25	70
28	269	A	12	2	99	102	903.99	d15	43.87058825	17.08826976	33370.27581	1	1.930333287	TRUE	dir	Geographic Coordinates	90	0	8	15	90
28	269	A	12	3	31	37	904.81	d16	311.4540766	25.38271471	6261.890212	8	7.110635407	TRUE	dir	Geographic Coordinates	90	0	3	70	100
28	269	A	12	3	100	104	905.5	d17	112.8863995	-20.79215717	21881.08232	7	2.435042272	TRUE	dir	Geographic Coordinates	90	0	9	20	70
28	269	A	12	3	139	143	905.89	d18	28.96001016	-64.21128488	18086.04801	6	2.041162135	TRUE	dir	Geographic Coordinates	90	0	9	20	70
28	269	A	12	4	35	38	906.35	d19	141.8096984	27.06084292	19862.01544	5	3.18722628	TRUE	dir	Geographic Coordinates	90	0	8	25	70
28	269	A	12	4	114	114	907.14	d20	344.1505266	24.98480412	7524.080645	4	5.534615148	TRUE	dir	Geographic Coordinates	90	0	7	15	50
28	269	A	12	5	52	54	908.02	d21	69.08596658	24.27440215	17408.94553	2	2.165969334	TRUE	dir	Geographic Coordinates	90	0	9	25	80
28	269	A	12	6	43	46	909.43	d22	281.6904156	26.9239216	14765.06822	2	2.552548953	TRUE	dir	Geographic Coordinates	90	0	5	40	70
28	269	A	12	6	119	120	910.19	d23	153.9393002	-43.82377169	17773.54343	1	3.971015509	TRUE	dir	Geographic Coordinates	90	0	8	25	70
28	269	A	11	1	128	131	854.78	d24	149.0593065	59.0440706	6141721.081	8	13.68602895	TRUE	dir	Geographic Coordinates	90	0	4	50	80
28	269	A	11	2	100	102	856	d25	326.9019884	-50.97461873	289344012	7	3.08101949	TRUE	dir	Geographic Coordinates	90	0	9	30	90
28	269	A	11	3	116	119	857.66	d26	334.5364202	-50.21976757	388489949.3	6	2.845164352	TRUE	dir	Geographic Coordinates	90	0	11	25	100
28	269	A	11	4	102	104	859.02	d27	226.1544957	-78.90314623	98982659.64	5	1.643774278	TRUE	dir	Geographic Coordinates	90	0	7	30	70
28	269	A	10	1	118	120	807.18	d28	127.4962079	33.84819793	125506107.5	4	3.888959636	TRUE	dir	Geographic Coordinates	90	0	8	25	70
28	269	A	10	2	135	137	808.85	d29	358.1326336	65.43141803	34389970.9	3	4.194877001	TRUE	dir	Geographic Coordinates	90	0	9	30	90
28	269	A	10	3	79	81	809.79	d30	263.7014199	65.84537641	218693976.6	2	1.907142101	TRUE	dir	Geographic Coordinates	90	0	8	30	80
28	269	A	10	3	113	115	810.13	d31	312.9883864	27.39953287	639061206.2	1	0.796159944	TRUE	dir	Geographic Coordinates	90	0	9	30	90
28	269	A	10	4	47	49	810.97	d32	281.4697077	18.89465552	1001907661	8	1.089722914	TRUE	dir	Geographic Coordinates	90	0	9	30	90
28	269	A	9	1	140	141	750.4	d33	6.985047386	37.39524013	71105293.91	7	7.206328365	TRUE	dir	Geographic Coordinates	90	0	6	35	70
28	269	A	9	2	91	92	751.41	d34	142.4203444	12.33232116	310016716.7	6	2.480805913	TRUE	dir	Geographic Coordinates	90	0	9	35	100
28	269	A	9	3	56	57	752.56	d35	30.53633776	51.30792985	222359261.8	5	9.025097405	TRUE	dir	Geographic Coordinates	90	0	7	40	90
28	269	A	9	3	82	84	752.82	d36	148.2317044	28.23487618	320725804.9	4	2.837413	TRUE	dir	Geographic Coordinates	90	0	9	30	90
28	269	A	8	1	112	114	702.62	d37	296.5709821	-35.97222585	178226512.9	3	2.017011139	TRUE	dir	Geographic Coordinates	90	0	6	30	60
28	269	A	8	2	32	34	703.32	d38	178.8055694	-33.57064374	31276157.15	2	15.17208249	TRUE	dir	Geographic Coordinates	90	0	6	35	70
28	269	A	8	3	21	23	704.71	d39	136.2052788	65.45985345	254911417.1	1	1.883407029	TRUE	dir	Geographic Coordinates	90	0	8	30	80
28	269	A	8	4	24	26	706.24	d40	154.1991718	-55.25178141	12870874.8	7	5.107651031	TRUE	dir	Geographic Coordinates	90	0	8	25	70
28	269	A	7	2	20	20	655.78	d42	199.0415962	-18.28783989	71599055.3	8	6.586390491	TRUE	dir	Geographic Coordinates	90	0	3	15	25
28	269	A	7	3	21	23	657.21	d43	193.5689914	-36.21456441	77389858.44	6	1.903325293	TRUE	dir	Geographic Coordinates	90	0	6	35	70
28	269	A	7	4	51	53	659.01	d44	168.0187196	-26.12330831	72845114.34	5	6.719835872	TRUE	dir	Geographic Coordinates	90	0	4	40	60

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Table S2. Palynomorphs assemblages at Hole 269A; Cores 7R-13R.

Sample Name	Core Section	Interval (cm)	Depth core (mbsf)	Lycopodium	Lycopodium counted	Weight (g)	Brigantedinium spp.	Brigantedinium cavate	Brigantedinium pynei	Selenopemphix spp.	Selenopemphix antarctica	Selenopemphix Nephroides	Selenopemphix Undulata	Lejeunecysta spp.	Lejeunecysta acuminata	Protoperidinium	Malvinia escutiana	Vozzhennikovia spp.	Spinidinium spp.	Spinidinium macmurdoense	Wetzellia spp.	Rhombodinium draco	Deflandrea spp.	Deflandrea antarctica	Phthanoperidinium spp.	Total protoperidinioids
DSDP Leg 28 Site 269A	006R 3W	100-104	610.52	20848	2549	19.4551	9																			10
DSDP Leg 28 Site 269A	007R 2W	30-34	655.82	20848	648	18.9571	174	1	1	2		3		4	1											195
DSDP Leg 28 Site 269A	007R 4W	30-34	658.82	20848	1538	20.0307	127	14						2.5		9		10	1	2						156.5
DSDP Leg 28 Site 269A	007R 4W	113-117	659.65	20848	2533	19.4604	49	7		1						2		3.5								62.5
DSDP Leg 28 Site 269A	008R 2W	131-135	704.33	20848	2283	20.1364	1											1								2
DSDP Leg 28 Site 269A	008R 3W	28-32	704.8	20848	2243	19.3753	124	15		7				13.5				5	3							167.5
DSDP Leg 28 Site 269A	009R 2W	26-30	750.78	20848	3505	16.3358	45	4											2	1						59
DSDP Leg 28 Site 269A	009R 3W	30-34	752.32	20848	1863	19.8009																				0
DSDP Leg 28 Site 269A	009R 3W	125-129	753.27	20848	2285	18.2004	86	2		3				10			3	62.5	2	2						171.5
DSDP Leg 28 Site 269A	010R 3W	25-29	809.27	20848	2017	20.1234																				0
DSDP Leg 28 Site 269A	010R 4W	25-29	810.77	20848	1999	19.1668	65			3				5.5												73.5
DSDP Leg 28 Site 269A	011R 1W	75-79	854.27	20848	1766	14.4943	187			7	1			2	1											198
DSDP Leg 28 Site 269A	011R 4W	22-26	858.24	20848	2128	20.0815	82	8		7		0.5		6				4.5								108
DSDP Leg 28 Site 269A	011R 4W	125-129	859.27	20848	2100	20.0162	150	25		4								1	1							181
DSDP Leg 28 Site 269A	012R 2W	35-39	902.87	20848	1865	19.4047	88											2								90
DSDP Leg 28 Site 269A	012R 2W	125-129	903.77	20848	2018	20.0258	123	5		6				0.5				14	3	1						156.5
DSDP Leg 28 Site 269A	012R 6W	125-129	909.77	20848	2069	20.0322	169	2		20.5				4				2								215.5
DSDP Leg 28 Site 269A	013R 1W	33-37	948.85	20848	840	19.2824	134			19				1				1								200
DSDP Leg 28 Site 269A	013R 2W	26-30	950.28	20848	1642	20.0555	159.5			35				1	1											226
DSDP Leg 28 Site 269A	013R 4W	125-129	954.27	20848	839	18.1817	148			26	5	43.5		5					1							228.5

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1288 **Table S4.** Concentrations of GDGTs at Hole 269A; Cores 7R-13R. All samples and
1289 corresponding depths, GDGT peak area values, TEX86 (Schouten et al., 2002) and BIT index
1290 values (Hopmans et al., 2004), Methane Index (Methzhang) values (Zhang et al., 2011),
1291 GDGT2/Crenarchaeol ratios (Weijers et al., 2011), GDGT-0/Crenarchaeol ratios (Blaga et al.,
1292 2009) and GDGT-2/GDGT-3 ratios (Taylor et al., 2013), and RING index (Sinninghe Damsté,
1293 2016). Discarded samples (OUTLIER=TRUE) with outlier values are based on BIT > 0.4,
1294 GDGT2/GDGT3` > 5, `GDGT0/cren` > 2 and `Methzhang` > 0.3.

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Sample code	Depth(mbsf)	GDGT-0	GDGT-1	GDGT-2	GDGT-3	GDGT-4	GDGT-4'	GDGT-IIIa	GDGT-IIIa'	SUM GDGT-IIIa	GDGT-IIIa	GDGT-IIIa'	SUM GDGT-IIIa	GDGT-IIb	
28-269A-7R-2W 30-34cm	655.8	2942750.0	159364.0	0.0	62727.9	18918.3	2706500.0	45714.4	47578.7	26296.1	73874.8	18699.4	13806.1	32505.5	6965.0
28-269A-7R-3W 136-136cm	658.4	12613497.0	717051.9	119649.4	385648.6	77529.1	12553429.0	198797.5	314790.5	156020.8	470811.3	67665.6	73249.0	140914.6	52635.8
28-269A-7R-4W 127-131	659.8	28695870.0	829447.3	248715.4	461230.4	57143.8	8250699.5	231671.9	139414.0	94517.9	233931.9	51122.1	38731.7	89853.8	32993.8
28-269A-8R-3W 142-146	705.9	4760829.0	204032.7	38097.8	92975.4	29892.9	4139520.5	56135.2	167024.0	105912.6	272936.6	90077.1	50242.6	140319.7	54628.9
28-269A-9R-2W 106-110cm	753.1	17138900.0	576784.9	167579.6	251321.1	1177106.8	16933678.0	174104.2	230212.3	147753.3	377965.6	70443.5	65665.1	136108.6	69136.6
28-269A-10R-3W 25-28cm	809.3	28890.9	2495.0	7904.5	2223.6	2104.2	4668.5	2071.6	5271.6	2961.4	8233.0	3065.6	5508.6	8574.2	0.0
28-269A-10R-3W 116-120cm	810.2	6003169.0	254334.4	56686.4	86115.8	45950.0	6143775.5	83228.2	151393.5	86406.7	237800.2	52180.1	44458.8	96638.9	42485.1
28-269A-11R-1W 75-79cm	854.3	795513.0	42238.5	0.0	15689.7	5634.5	861324.0	12810.7	19932.8	10570.1	30502.9	8328.0	4752.0	13079.9	5463.0
28-269A-11R-3W 23-27cm	856.8	33051.7	2651.7	0.0	2004.7	2018.5	22133.5	2057.4	9372.8	7193.4	16566.2	3557.1	4061.8	7618.9	0.0
28-269A-12R-2W 40-43cm	902.9	194355.7	10943.2	0.0	3960.4	2268.3	183369.7	4365.3	7947.7	9103.8	17051.5	5069.7	6338.1	11407.8	2383.1
28-269A-12R-4W 25-28cm	905.8	3728442.0	168857.0	33074.1	70626.6	29034.4	3919071.8	48228.7	144021.0	79521.6	223542.6	68037.8	46249.2	114287.0	31120.5
28-269A-12R-4W 125-128cm	906.8	3072708.0	130516.6	16421.6	59769.3	22098.5	3129816.3	44337.4	76395.4	44885.9	121281.3	39721.6	24584.5	64306.1	21236.2
28-269A-13R-1W 33-37cm	948.9	459448.0	21973.4	0.0	7378.4	3209.4	486734.0	6502.2	25618.2	14374.4	39992.6	14912.4	9208.5	24120.9	6242.0
28-269A-13R-2W 87-91cm	950.9	2347130.0	105004.3	0.0	51535.6	19066.5	2323012.3	33990.5	168605.3	101849.8	270455.1	92141.9	50565.0	142706.9	41430.6
28-269A-13R-4W 13-17cm	953.2	2251815.8	122400.4	21870.6	52940.5	18491.2	2273169.8	37531.8	98155.3	57310.7	155466.0	54374.1	31827.6	86201.7	28874.5

GDGT-11b ¹	SUM GDGT-11b	OH-1034	GDGT-11c	GDGT-11c ¹	SUM GDGT-11c	GDGT-1a	GDGT-1b	GDGT-1c	TEX86	SST	BIT	SUM_ISO	FA1302:ISO	FA1300:ISO	FA1298:ISO				
1034-tot	OH-1034	1032-1	1032-2	1032-tot	1022	1020	OH-1020a	OH-1020b	OH-1020c	OH-1020d	1018								
14544.7	21509.7	0.0	3749.1	9198.9	12948.0	40162.4	14362.6	0.0	0.0	0.0	0.0	12066.1	0.44	9.60	0.05	5935974.60	0.496	0.027	0.011
80306.2	132942.0	104047.4	13712.3	24680.3	38392.6	259286.0	48214.7	9688.9	309723.9	389706.7	31440.2	44094.6	0.48	12.52	0.06	26545953.10	0.475	0.027	0.015
35129.8	68123.6	37714.4	9001.3	13210.8	22212.1	115207.8	21693.6	12370.7	161634.8	228429.3	19674.6	17263.1	0.47	12.10	0.05	38526062.90	0.745	0.022	0.012
37768.2	92397.1	48924.1	13229.0	14742.5	27971.5	195084.8	29822.2	6062.6	214955.1	307641.0	14341.2	26600.7	0.47	11.49	0.13	9283385.70	0.513	0.022	0.010
69015.9	138152.5	55351.3	22621.5	25153.8	47775.3	379003.8	51333.5	18512.8	350091.6	461749.2	22602.7	47889.6	0.48	12.90	0.05	35191895.00	0.487	0.016	0.007
0.0	0.0	0.0	0.0	0.0	0.0	5851.6	2279.7	2355.4	31366.7	41509.1	9545.0	2403.8	0.72	32.04	0.83	42453.80	0.681	0.059	0.052
34092.4	76577.5	47095.5	10714.1	14235.1	24949.2	167211.4	23949.2	3173.1	198778.5	257938.0	11349.6	19038.2	0.46	10.76	0.08	12616572.90	0.476	0.020	0.007
4166.0	9629.0	0.0	1101.0	1475.0	2576.0	15798.4	2882.8	0.0	0.0	0.0	0.0	1790.2	0.45	9.83	0.06	1733210.38	0.459	0.024	0.009
0.0	0.0	0.0	0.0	0.0	0.0	3545.1	0.0	2460.4	41030.8	52970.5	19288.3	0.0	0.70	30.15	0.56	63917.50	0.517	0.041	0.031
4796.7	7179.8	0.0	2537.9	0.0	2537.9	0.0	3250.4	2962.9	48828.5	58378.8	5677.2	3067.2	0.49	13.49	NA	399262.60	0.487	0.027	0.010
25491.3	56611.8	58382.1	7456.9	9845.0	17301.9	160392.1	17823.1	4769.6	113524.4	137038.3	8631.3	11332.4	0.47	11.45	0.11	7964260.50	0.468	0.021	0.009
17262.9	38499.1	21836.9	7235.3	6543.2	13778.5	102277.7	15114.2	3591.8	77430.9	100388.7	8322.0	10313.6	0.49	13.47	0.08	6459246.10	0.476	0.020	0.009
5069.5	11311.4	0.0	1511.2	1574.5	3085.7	25200.1	3997.7	0.0	0.0	0.0	0.0	1824.3	0.44	9.06	0.16	985245.36	0.466	0.022	0.007
27951.3	69381.9	0.0	10227.7	11422.2	21649.9	151594.1	26127.8	3518.7	56499.7	79548.7	7004.6	11069.5	0.50	14.07	0.20	4879739.20	0.481	0.022	0.011
21519.9	50394.4	0.0	5752.5	7663.4	13415.9	111762.2	16782.5	4996.4	84776.7	101662.3	12490.3	8417.6	0.47	11.78	0.13	4756349.50	0.473	0.026	0.011

FA1296.ISO	FA1292.ISO	FA1292.ISO	GDGT2/GDGT3	GDGT2/cren	GDGT0/cren	Methhang	RING	SUM.ALL	SUM.BR	GBT	pH	MBT	MAT	outlierBIT	outlierGDGT/3	outlierGDGT0/cren	outlierMethhang	outlierRING	OUTLIER
0.003	0.456	0.008	3.316	0.023	1.087	0.081	1.912	5902393.59	207429.19	0.116	8.350	0.321	10.104	FALSE	FALSE	FALSE	FALSE	FALSE	FALSE
0.003	0.473	0.007	4.974	0.031	1.005	0.085	1.986	26500379.30	1134655.80	0.019	8.603	0.310	10.310	FALSE	FALSE	FALSE	FALSE	FALSE	FALSE
0.001	0.214	0.006	8.071	0.056	3.478	0.137	0.931	37746527.30	568285.90	0.094	8.407	0.271	8.686	FALSE	TRUE	FALSE	FALSE	FALSE	TRUE
0.003	0.446	0.006	3.110	0.022	1.150	0.072	1.859	9741617.30	785132.60	0.144	8.278	0.320	9.926	FALSE	FALSE	FALSE	FALSE	FALSE	FALSE
0.003	0.481	0.005	2.146	0.015	1.012	0.052	1.985	35424911.10	1178228.90	-0.005	8.664	0.406	13.419	FALSE	FALSE	FALSE	FALSE	FALSE	FALSE
0.050	0.110	0.049	1.057	0.476	6.188	0.503	0.947	62973.30	27342.30	0.678	6.890	0.385	8.912	TRUE	FALSE	FALSE	FALSE	TRUE	TRUE
0.004	0.487	0.007	1.874	0.014	0.977	0.058	2.019	12876337.30	646164.60	0.079	8.446	0.325	10.446	FALSE	FALSE	FALSE	FALSE	FALSE	FALSE
0.003	0.497	0.007	2.785	0.018	0.924	0.068	2.070	1745906.93	76259.23	0.106	8.377	0.268	8.532	FALSE	FALSE	FALSE	FALSE	FALSE	FALSE
0.032	0.346	0.032	0.993	0.091	1.493	0.216	1.713	NA	NA	NA	NA	NA	NA	TRUE	FALSE	FALSE	FALSE	TRUE	TRUE
0.006	0.459	0.011	1.746	0.022	1.060	0.084	1.945	NA	NA	0.148	8.268	NA	NA	FALSE	FALSE	FALSE	FALSE	FALSE	TRUE
0.004	0.492	0.006	2.433	0.018	0.951	0.063	2.042	8297033.40	601290.90	0.249	8.004	0.315	9.170	FALSE	FALSE	FALSE	FALSE	FALSE	FALSE
0.003	0.485	0.007	2.705	0.019	0.982	0.063	2.015	6612432.20	365570.50	0.171	8.208	0.349	10.672	FALSE	FALSE	FALSE	FALSE	FALSE	FALSE
0.003	0.494	0.007	2.299	0.015	0.944	0.062	2.050	1062216.84	109532.69	0.264	7.965	0.283	8.093	FALSE	FALSE	FALSE	FALSE	FALSE	FALSE
0.004	0.476	0.007	2.703	0.022	1.010	0.069	1.986	5397118.00	692985.20	0.247	8.009	0.272	7.853	FALSE	FALSE	FALSE	FALSE	FALSE	FALSE
0.004	0.478	0.008	2.863	0.023	0.991	0.077	2.003	5004957.70	442440.30	0.186	8.169	0.310	9.354	FALSE	FALSE	FALSE	FALSE	FALSE	FALSE

1303

1304 **R-script: GDGT Peak area calculation**

1305 **required packages**

1306 `library(caTools)`

1307 `library(ggplot2)`

1308 `library(dplyr)`

1309 `##`

1310 `## Attaching package: 'dplyr'`

1311 `## The following objects are masked from 'package:stats':`

1312 `##`

1313 `## filter, lag`

1314 `## The following objects are masked from 'package:base':`

1315 `##`

1316 `## intersect, setdiff, setequal, union`

1317 `library(readr)`

1318 `library(readxl)`

1319 `library(tidyverse)`

1320 `## — Attaching packages ————— — tidyverse 1.`

1321 `2.1 —`

1322 `## ✓ tibble 2.1.3 ✓ purrr 0.3.2`

1323 `## ✓ tidyr 0.8.3 ✓ stringr 1.3.1`

1324 `## ✓ tibble 2.1.3 ✓ forcats 0.3.0`

1325 `## — Conflicts ————— tidyverse_conflict`

1326 `s() —`

1327 `## ✗ dplyr::filter() masks stats::filter()`

1328 `## ✗ dplyr::lag() masks stats::lag()`

1329 `library(plotly)`

1330 `##`

1331 `## Attaching package: 'plotly'`

1332 `## The following object is masked from 'package:ggplot2':`

1333 `##`

1334 `## last_plot`

1335 `## The following object is masked from 'package:stats':`

1336 `##`

1337 `## filter`

1338 `## The following object is masked from 'package:graphics':`

1339 `##`

1340 `## layout`

1341 `library(ggtern)`

1342 `## --`

1343 `## Remember to cite, run citation(package = 'ggtern') for further info.`

1344 `## --`

1345 `##`

1346 `## Attaching package: 'ggtern'`

1347 `## The following objects are masked from 'package:ggplot2':`

1348 `##`

1349 `## %+%, aes, annotate, calc_element, ggplot, ggplot_build,`

1350 `## ggplot_gtable, ggplotGrob, ggsave, layer_data, theme,`

1351 `## theme_bw, theme_classic, theme_dark, theme_gray, theme_light,`

1352 `## theme_linedraw, theme_minimal, theme_void`

1353 `theme = theme_set(theme_classic())`

1354 `library(readxl)`

1355 `setwd("~/Documents/Projects/Site 269")`

```

1356 OGdata <- read_excel("269 OG.xlsx")
1357 #shows data tables below
1358 OGdata
1359 adds collumns with calculations from GDGT peak areas
1360 #adds collumns with calculations from GDGT peak areas
1361 OGdata <- OGdata %>%
1362 mutate(
1363   #TEX86 ratio from Schouten et al. 2002
1364   TEX86 = (`1298` + `1296` + `1292`)/(`1300` + `1298` + `1296` + `1292`
1365   `),
1366   #SST calibrations from Kim et al., 2010; TEX86 linear calibration model (0 m)
1367   SST = 81.5*TEX86 -26.6,
1368   #Hopmans et al. 2004
1369   BIT = (`1050-tot` + `1036-tot` + `1022`)/(`1292` + `1050-tot` + `1036`
1370   -tot` + `1022`),
1371   #fractional abundances needed for the RING index
1372   SUM.ISO = `1302` + `1300` + `1298` + `1296` + `1292` + `1292`,
1373   FA1302.ISO = `1302` / SUM.ISO,
1374   FA1300.ISO = `1300` / SUM.ISO,
1375   FA1298.ISO = `1298` / SUM.ISO,
1376   FA1296.ISO = `1296` / SUM.ISO,
1377   FA1292.ISO = `1292` / SUM.ISO,
1378   `FA1292'.ISO` = `1292` / SUM.ISO,
1379   #depth contribution (Taylor et al., 2013) AOM (Weijers et al., 2011), lake in
1380   situ production (Blaga et al., 2009), methanogenic (Zhang et al.) indices and
1381   the RING index
1382   `GDGT2/GDGT3` = `1298` / `1296`,
1383   `GDGT2/cren` = `1298` / `1292`,
1384   `GDGT0/cren` = `1302` / `1292`,
1385   `Methzhang` = (`1300` + `1298` + `1296`) / (`1300` + `1298` + `1296`
1386   + `1292` + `1292`),
1387   `RING` = FA1300.ISO + 2*FA1298.ISO + 3*FA1296.ISO + 4*FA1292.ISO + 4*
1388   `FA1292'.ISO`,
1389   #Summarizing
1390   SUM.ALL = `1302` + `1292` + `1292` + `1022` + `1020` + `1018` + `103`
1391   6-tot` + `1034-tot` + `1032-tot` + `1050-tot`,
1392   SUM.BR = `1022` + `1020` + `1018` + `1036-tot` + `1034-tot` + `1032-t`
1393   ot` + `1050-tot`,
1394   # weijers et al., 2007
1395   CBT = -log10((`1020` + `1034-tot`) / (`1020` + `1036-tot`)),
1396   pH = (3.3283-CBT)/ 0.3847,
1397   MBT = (`1022` + `1020` + `1018`) / SUM.BR,
1398   #Peterse et al., 2012 (MBT' = MBT, because 1048 and 1046 were not integrated)
1399   `MAT` = 0.81 - 5.67*CBT + 31*MBT,
1400   )
1401 CAPDELRLIL <- function(TEX86) -1.15*TEX86 + 2.98*TEX86^2 + 1.49
1402 CAPDELRLIU <- function(TEX86) -0.39*TEX86 + 3.66*TEX86^2 + 1.69
1403 Outlier rules for TEX86
1404 #If:
1405 #RING outside CAPDELRI curves,
1406 #GDGT2/GDGT3 > 5 OR,
1407 #GDGT2/cren > 0.1 OR,
1408 #GDGT0/cren > 2 OR,
1409 #BIT > 0.4 OR,
1410 #Methzhang 1300+1298+1296/1300+1298+1296+1292+1292' > 0.3

```



```

1411 #-->
1412 #rejectTEX86
1413 #this adds columns to your dataframe with a logical value (TRUE or FALSE) for
1414 each outlier rule. The final OUTLIER column is added with a logical value for
1415 whether any of the other outlier rule is true
1416 OGdata <- OGdata %>%
1417 mutate(
1418   `outlierBIT` = BIT > 0.4,
1419   `outlierGDGT2/3` = `GDGT2/GDGT3` > 5,
1420   `outlierGDGT0/cren` = `GDGT0/cren` > 2,
1421   `outlierMethzhang` = `Methzhang` > 0.3,
1422   outlierRING = RING<CAPDELRI.L(TEX86) | RING>CAPDELRI.U(TEX86)
1423 ,
1424   OUTLIER = `outlierBIT` | `outlierGDGT2/3` | `outlierGDGT0/cre
1425 n` | `outlierMethzhang` | outlierRING
1426 )
1427 #shows new

```

1428

1429 **Supplementary References**

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