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1 **Gravimetric Constraints on the Hydrothermal System**
2 **of the Campi Flegrei caldera**

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6 **Key Points:**

- 7 • Hydrothermal plumbing of the central sector of the Campi Flegrei is controlled
8 by volcano-tectonic structures
- 9 • Three shallow-seated (< 1.5 km depth) hydrothermal feeder systems are imaged
10 at Pozzuoli, Solfatara/Pisciarelli volcano and Astroni volcano
- 11 • Low densities of feeder systems are explained by porous caldera-fill material with
12 between 0.38 and 1 vapour volume fraction and between 0 and 0.62 liquid volume
13 fraction in secondary void space

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Abstract

The Campi Flegrei caldera (Italy) has been undergoing unrest over the past five decades including episodes of rapid ground deformation, seismicity and variations in gas emissions. Hydrothermal fluids and gases are released most vigorously in the central sector of the caldera at the fumarolic fields of Solfatara volcano and Pisciarelli. We conducted a high-precision gravity survey coupled with inverse modelling to image the shallow (<2 km depth) structure of the hydrothermal feeder system. Results indicate the presence of three anomalously low density bodies beneath Pozzuoli, Astroni volcano and the Solfatara/Pisciarelli fumarolic fields. The first two are inferred to be sealed hydrothermal systems trapped beneath impermeable cap rock while the latter depicts a plume-like geothermal feeder system reaching the surface via a combination of Solfatara's maar-diatreme structure and the intersection of NW-SE and NE-SW trending regional faults. The density contrasts of the reservoirs from background values are best explained by a multi-phase mixture of caldera-fill containing a secondary and interconnected void volume fraction of between 0.2 and 0.3 that hosts a vapour volume fraction ψ_v of between 0.38 and 1 and a liquid volume fraction ψ_l fraction of between 0 and 0.62. This work highlights the control of volcano-tectonic structures on fluid movement in the shallow crust of hydrothermally active volcanic systems undergoing sustained or periodic unrest.

1 Introduction

Volcanic unrest is often characterised by anomalous seismicity, gas emissions and surface deformation, and is usually attributed to sub-surface magma movement (Sparks, 2003). Volcanic calderas have complex sub-surface structures resulting not at least from the vertical collapse of a pre-existing volcanic edifice and often host both extensive hydrothermal and magmatic reservoirs (Gottsmann & Battaglia, 2008). Hydrothermal systems are a complex interface between magma reservoirs and the surface (Todesco, 2008) and not only produce measurable unrest signals but also modulate geophysical and geochemical signals from underlying magma reservoirs (Chiodini et al., 2002; Gottsmann & Battaglia, 2008; Todesco, 2008; Ingebritsen et al., 2010; Chiodini et al., 2016).

Campi Flegrei caldera (CFC) is a well-documented restless caldera where the separation of the signals from magmatic and hydrothermal sources has not been trivial (Troise et al., 2019). Solfatara volcano and neighbouring Pisciarelli (Fig. 1) host the main surface features of the hydrothermal system at CFC and are located ~ 2.5 kilometres to the

NE of Pozzuoli, the centre of ground deformation over the last 50 years (Di Giuseppe et al., 2015) of unrest. Our current understanding of the structure and dynamics of the hydrothermal system at CFc is informed predominantly by geochemical constraints, geophysical data and resulting models (Caliro et al., 2007; Bruno et al., 2007; Troiano et al., 2019): a multiphase plume of vapour and liquid fuelled by the interaction of magmatic and meteoric fluids at depth rises to feed fumaroles and mud-pools at the surface (Chiodini et al., 2015). While significant sub-surface density variations are expected from this model, gravity data have not been used to contribute to the understanding of the shape and size of the shallow-seated (< 1.5 km) hydrothermal feeder system. Here, we present results from a new gravimetric survey of the central sector of the CFc including a high-resolution gravity survey of Solfatara volcano coupled with data inversion to image the density structure of the upper-most part of the hydrothermal system.

2 Background

2.1 Campi Flegrei Caldera Structure and Recent Unrest History

Campi Flegrei caldera is a ~ 13 -km-wide volcanic caldera in the Campanian Plain near Naples, Italy (Vitale & Isaia, 2014) formed by two major vertical collapses at ~ 40 ka (Giaccio et al., 2017) and ~ 15 ka (Deino et al., 2004). Post-collapse eruptive activity over the last 15 ka generated at least 70 eruptions, mainly concentrated in the central eastern sector of the caldera (Smith et al., 2011; Bevilacqua et al., 2015). More than 20 eruptions occurred in the epoch of activity from 5.8 to 3.8 ka forming landmarks such as Astroni volcano and Agnano caldera (Isaia et al., 2009; Smith et al., 2011), with the latest magmatic eruption in 1538 AD creating Monte Nuovo volcano (Barberi et al., 1984).

The vertical collapses and long-term ground deformation have divided the caldera floor into a block structure (Orsi et al., 1999). The dominant fault trends within the caldera are NW-SE and NE-SW (Di Vito et al., 1999; Florio et al., 1999; Vitale & Isaia, 2014; Isaia et al., 2015) in addition to caldera ring faults (Berrino et al., 2008; Barberi et al., 1991; Gottsmann et al., 2006; Zollo et al., 2003).

The caldera fill is composed of intercalated lava flows, pyroclastic material, and marine and continental sediments (Rosi & Sbrana, 1987; Piochi et al., 2014). Gravity data depict the fill as a broad (~ 6 km wavelength) low density anomaly (Barberi et al., 1991; Capuano et al., 2013). Drilling encountered a zone of thermo-metamorphic rocks below

77 the fill at depths between 2.5 km and 3.1 km, several small igneous intrusions, high ther-
78 mal gradients of 100-170 K km⁻¹ and locally raised isotherms in the central-eastern part
79 (Piochi et al., 2014) of the caldera.

80 Episodes of rapid uplift and subsidence have been recorded at CFc since Roman
81 times (Parascondola, 1947; Bellucci et al., 2006), though subsidence has been the dom-
82 inant long-term trend (Barberi et al., 1984). Episodes of rapid ground uplift occurred
83 in 1969-72 and 1982-84 totalling 3.5 m near Pozzuoli (Barberi et al., 1984). Bucking a
84 trend of slow ground subsidence since 1989, numerous mini-uplift events have occurred
85 producing deformation on the order of centimetres accompanied by seismic activity (Chiodini
86 et al., 2010, 2015) followed by a new episode of sustained uplift beginning in 2005. Fu-
87 marolic flow rate, discharge temperature and seismicity increased at Solfatara volcano
88 and Pisciarelli from 2006 onwards and an increased magmatic contribution was inferred
89 from the composition of the fumarolic gases (Chiodini et al., 2016; Giudicepietro et al.,
90 2019). The cause of the rapid meter-scale uplifts is still controversial with magmatic and
91 hydrothermal sources or a mix of both identified as contributors (Troise et al., 2019). Many
92 authors suggest that the more recent mini-uplift events have an exclusively hydrother-
93 mal origin (Gottsmann et al., 2006; Manconi et al., 2010; D’Auria et al., 2011; Amoruso
94 et al., 2014; Chiodini et al., 2015). NW-SE faults are reactivated during uplift and sub-
95 sidence and may be important pathways for the upward movement of gas and magma
96 to the surface (Vilardo et al., 2010)

97 **2.2 Solfatara volcano and its fumarolic fields**

98 Solfatara volcano is the most thermally active part of the caldera (Di Giuseppe et
99 al., 2015). It releases ten times more thermal energy than the conductive heat flux across
100 the entire caldera floor (Chiodini et al., 2001). Hydrothermal gases and fluids are released
101 most vigorously at Solfatara volcano’s crater floor and its eastern inner wall as well as at
102 the Pisciarelli fumarolic field located on its NE flank (Fig. 1) (Caliro et al., 2007). The
103 100° C isotherm resides only a few hundred metres below the surface of Solfatara’s crater
104 (Piochi et al., 2014).

105 The main hydrothermal features within Solfatara’s crater include a mud-pool named
106 La Fangaia and two main fumaroles named Bocca Nuova and Bocca Grande (see Fig.
107 1). Detailed geological mapping by (Isaia et al., 2015) identifies Solfatara’s structure as
108 a maar-diatreme. High angle normal faults characterise the crater edges and fault planes

109 are predominantly NW-SE. Two lava domes reside in the NE and S crater walls. The
110 crater itself is embedded in the older structure of the Agnano-Monte Spina Complex and
111 fumaroles and hydrothermal activity are concentrated in the fault zones and their in-
112 tersections, where highly fractured rocks act as preferred pathways for fluid ascent.

113 **2.3 The Hydrothermal Plume**

114 The presence of a hydrothermal plume beneath Solfatara volcano was first suggested
115 by Cioni et al. (1984) who, based on geochemical data, proposed that dry steam sepa-
116 rates from a geothermal liquid at 236° C in a highly fractured zone to feed the fuma-
117 role fields. Further compelling evidence was presented by Chiodini et al. (2015) based
118 on fumarole geochemistry, CO₂ flux, water table heights, seismic velocity and InSAR data
119 as well as by thermo-hydro-mechanical modeling (Todesco & Berrino, 2005; Coco et al.,
120 2016). The numerical models require the multi-phase flow of ascending hot fluids (H₂O
121 and CO₂) from depth through a porous medium to reproduce measured fumarole emis-
122 sions, ground deformation and gravity changes. The current conceptual model of the plume
123 suggests that rising magmatic gases flash hydrothermal liquids in a deep 'mixing zone'
124 and form a gas plume which ascends to the surface (Caliro et al., 2007; Chiodini et al.,
125 2015). A summary of relevant geophysical and geochemical surveys of Solfatara volcano
126 and the wider CFc is given in Table 1.

127 **3 Methods**

128 **3.1 Data Acquisition**

129 We performed a new static gravity survey from 8-12 July, 2015 using a Scintrex CG-
130 5 Autograv gravimeter (serial number: 572) in tandem with a TOPCON HiPer Pro Dual-
131 Frequency GNSS base and rover system. The survey area encompassed the highly ur-
132 banised central sector of the CFc (Fig. 2) and contained a total of 85 benchmarks laid
133 out in two different spatial networks.

134 Benchmarks within Solfatara crater were ordered in a dense irregular grid with a
135 minimum spacing of 17 m and included a local gravity control point. The remainder of
136 the benchmarks were spaced more widely along the roads around Solfatara volcano with
137 an average spacing of 1 km and a maximum spacing of 2.5 km. The different spacing per-
138 mitted us to investigate the expression of the hydrothermal plume at Solfatara volcano

139 at a similar scale to several local geophysical studies while also exploring the spatial dis-
 140 tribution of the hydrothermal system across the central sector of the CFc (Fig. 2). Both
 141 the GPS reference receiver and the main gravity base station were located near Monte
 142 Nuovo and all gravity measurements were tied to this reference. The entire network cov-
 143 ered an approximate area of 36 km² and the precision of repeat measurement was $\pm 15 \mu\text{Gal}$
 144 (average of 12 cycles of 30s long readings of 6Hz raw data at each benchmark). Urban
 145 noise led to an average standard error of individual gravity measurements of $\pm 8 \mu\text{Gal}$
 146 which is a factor of between 3 and 5 higher than usually attainable during quiet condi-
 147 tions.

148 We recorded GNSS data for 5-20 min at 1Hz at the survey benchmarks using a roving
 149 receiver/antenna unit. The base receiver/antenna unit recorded continuously at 1
 150 Hz during the survey period. The derived precision of the benchmark locations was gen-
 151 erally under 0.05 m in the vertical and better than 0.04 m in the horizontal after base-
 152 line processing of the benchmark locations against the base station which in turn was
 153 processed against three permanent reference stations of the local INGV Osservatorio Vesu-
 154 viano Permanent Global Navigation Satellite System (GNSS) network and three regional
 155 International GNSS Service (IGS) references (NOT1, MAT1, MEDI).

156 **3.2 Gravity Data Reduction and Correction**

157 The objective of a static gravity survey is to obtain information about the sub-surface
 158 density distribution. The magnitude of gravity at any point is influenced by latitude, el-
 159 evation, topography of the surrounding terrain, Earth and Ocean tides, sub-surface den-
 160 sity variations and instrumental drift (Telford et al., 1991). Raw gravity data are there-
 161 fore composed of several contributions and require careful corrections to obtain the com-
 162 ponent reflecting sub-surface density variations only, known also as the Bouguer anomaly
 163 (BA). Earth tides and instrumental drift are removed first to obtain the observed grav-
 164 ity (g_{obs}) from which the BA can be obtained.

$$BA = g_{obs} - g_n + FAC - BS + TC, \quad (1)$$

165 where g_n is the normal gravity, FAC is the free-air correction, BS is the Bouguer
 166 slab correction and TC is the terrain correction. A detailed description of the data re-
 167 duction is given in the Supplementary Information.

3.3 Data Detrending

The regional Bouguer anomaly is controlled by both shallow and deep-seated density distributions. Long-wavelength features (e.g., spatial variations in deep-seated bedrock thickness) must be removed to reveal the local Bouguer anomaly caused by shallow-seated structures. We calculate a regional gradient of 0.86 mGal/km with a strike of N50°E from the regional Bouguer anomaly data and derive the linearly-detrended residual anomaly (LRA) data for further investigation. Our regional trend compares to a regional gradient of 0.5 mGal/km and a strike of N35°E presented by Cassano and La Torre (1987) who use a much larger and wider-spaced dataset. The Topex gravity data (Sandwell et al., 2013) with a spatial coverage and average station spacing matching more closely with our survey gives a regional gradient of 0.2 mGal/km with a strike of N37° E.

To test the robustness of our results we detrended our data using the regional trend from the Topex data set. While the amplitudes of the resultant anomalies of course change, the presence and location of the main anomalies remain. Therefore, even using the lowest quoted regional gradient, we obtain model results that are reproducible and robust.

Removing a linear trend may not be appropriate in structurally complex areas such as collapse calderas to investigate anomalies associated with a shallow-seated hydrothermal system. Large scale gravity surveys at CFC have consistently shown a negative gravity anomaly associated with low density caldera fill (Nunziata & Rapolla, 1981; Berrino et al., 2008; Capuano et al., 2013). We explore the effect of the fill on our data by constructing a forward model based on the most recent gravity data presented by Capuano et al. (2013) and borehole density data (Barberi et al., 1991; Piochi et al., 2014). The caldera fill is simulated by stacked spheroids within a cylindrical volume of 2 km thickness and 3 km in radius with a density contrast of -300 kg m^{-3} , centered offshore of Pozzuoli (see Supplementary Information Figure S5). Capuano et al. (2013) suggest that the uppermost part of the caldera fill contains remnant high density feeder systems, as well as post-collapse lava flows and domes. We therefore set the top of the model at a depth of 1 km and subtract the simulated caldera-fill anomaly from the regional Bouguer anomaly data. We thus obtain a second local anomaly: the caldera-fill detrended residual anomaly (CRA).

Values for the LRA and CRA anomalies are reported relative to the base station at Monte Nuovo. Secondary anomalies of Solfatara volcano have their values calculated

200 from average background values of Solfatara’s crater floor. Relative values are provided
 201 so that the anomalies are comparable across the two detrended datasets.

202 **3.4 Total Horizontal Gravity Gradiometry**

203 First and second horizontal derivatives of Bouguer gravity data are useful to study
 204 structural controls on gravity anomalies (Cooper & Cowan, 2008). The first derivative
 205 highlights boundaries of buried bodies or faults. The second derivative yields inflection
 206 points of the first gradient and reveals absolute maxima/minima which provide infor-
 207 mation on the shape of buried bodies or inclination of density interfaces. The total hor-
 208 izontal gradients are obtain from

$$\text{THD1} = \sqrt{\left(\frac{\partial g}{\partial x}\right)^2 + \left(\frac{\partial g}{\partial y}\right)^2}, \quad (2)$$

209 where THD1 is the first total horizontal derivative, $\partial g/\partial x$ is the change in grav-
 210 ity in the x direction and $\partial g/\partial y$ is the change in gravity in the y direction (Cooper &
 211 Cowan, 2008). Similarly

$$\text{THD2} = \sqrt{\left(\frac{\partial^2 g}{\partial x^2}\right)^2 + \left(\frac{\partial^2 g}{\partial y^2}\right)^2}, \quad (3)$$

212 where THD2 is the second total horizontal derivative (Fedi, 2002).

213 **3.5 Inverting the Local Bouguer Anomaly Data**

214 We invert the resultant Bouguer gravity data (LRA and CRA) to image causative
 215 density contrasts at depth using GROWTH2.0 (Camacho et al., 2002, 2011). GROWTH2.0
 216 divides the model space into 3D parallelepiped elements and obtains a 3D anomalous
 217 density model using prescribed (a priori) density contrasts. Inherent non-uniqueness in
 218 the inversion is addressed by using a mixed minimisation condition which selects a so-
 219 lution based on least-squared model fitness and model smoothness, or the total anoma-
 220 lous mass. Model inputs include the Bouguer gravity data, cell size, the density contrast
 221 with background density and a balance factor. The balance factor determines the com-
 222 plexity of the model of positive and negative density contrasts with high balance factors
 223 producing simple models. Densities with too high a contrast produce isolated skeletal
 224 bodies and densities with too low a contrast produce inflated and interconnected bod-

225 ies. Both constraints must be explored to find an appropriately complex model with a
 226 plausible density contrast and low auto-correlation (Camacho et al., 2011). As a result,
 227 the minimisation of residuals is insufficient to establish the suitability of the model for
 228 given density contrasts. A mixed minimisation procedure which balances the goodness-
 229 of-fit criterium (model fitness) with the total anomalous mass (model smoothness) min-
 230 imisation condition is applied to select the optimal model. We explored the model space
 231 for each dataset iteratively, searching for suitably complex anomalous bodies and low auto-
 232 correlations for a low cell resolution and repeatedly increasing the resolution and retest-
 233 ing of the density contrast and balance factor at each iteration. A $50 \text{ kg m}^{-3} \text{ km}^{-1}$ in-
 234 crease in background density was implemented to prevent oversizing of anomalous bod-
 235 ies with increasing depth during the inversion.

236 We tested density contrasts in the range of ± 300 to $\pm 600 \text{ kg m}^{-3}$ and balance fac-
 237 tors from 10 to 40 (producing in total 119 model solutions) and selected the model with
 238 the lowest auto-correlation for given model smoothness. This methodology effectively
 239 uses a classic trade-off between model misfit and model simplicity (Gubbins, 2004). The
 240 best solution balances a compromise between adequately fitting the data and produc-
 241 ing a suitably simple model. While the model with the best goodness-of-fit has an auto-
 242 correlation of 0.06 and a balance factor of 20 it yields an array of skeletal bodies of anoma-
 243 lous densities and does not satisfy the mixed minimisation criteria for an optimal solu-
 244 tion. Our optimal model of the CRA has an autocorrelation of 0.13 and a balance fac-
 245 tor of 40 after 58 iterations, while the optimal model of the LRA has an autocorrelation
 246 of 0.14 and a balance factor of 40 after 61 iterations. Details on the inversion procedure
 247 and sensitivity tests are given elsewhere (Camacho et al., 2002, 2011).

248 **4 Results**

249 **4.1 The Bouguer Anomaly**

250 The amplitudes of all anomalies are orders of magnitude above the uncertainties
 251 associated with individual measurements or the terrain correction and are therefore ro-
 252 bust indicators of sub-surface density variations. Figs. 3a-c show the distribution and
 253 amplitudes of the regional and local Bouguer anomalies (LRA and CRA). Linear detrend-
 254 ing (Fig. 3b) reveals a broad ($\sim 4 \text{ km}$ wide) and negative ($\sim -6 \text{ mGal}$) anomaly cen-
 255 tered northwest of Solfatara volcano. It is composed of three distinct lows near Pozzuoli,

256 Solfatara volcano and Astroni volcano. The Pozzuoli anomaly is not present in the CRA
 257 data (Fig. 3c) and the overall negative anomaly is significantly reduced in both wave-
 258 length and amplitude (~ 0.4 km and ~ -4 mGal, respectively) and its centre shifted to-
 259 wards the north. The gravity lows between Solfatara volcano and Astroni volcano per-
 260 sist. Anomalies at the periphery of the survey are poorly constrained and hence ignored.

261 When looking in more detail at the Solfatara area, the patterns of the LRA and
 262 CRA anomalies are similar, but with noticeable differences in the negative amplitudes
 263 (from average values of the crater floor) of the respective gravity lows. The main grav-
 264 ity low in the eastern part of the crater is -1.1 mGal in the LRA from an average value
 265 of -3.15 mGal and -0.76 mGal in the CRA from an average value of 1.42 mGal (Fig. 4a
 266 and b). In both cases, the lows are located close to the fumaroles of Bocca Nuova and
 267 Bocca Grande and extend eastwards towards Pisciarelli. There are gravity highs on the
 268 north-northeastern and southern edges of the crater in both datasets, but the north-northeastern
 269 high is strongest in the LRA data.

270 While there is a small gravity low (~ -0.3 mGal amplitude from background levels
 271 in the crater floor) in the vicinity of La Fangaia in the LRA data (Fig. 4a), this anomaly
 272 is only very weak in the CRA data (Fig. 4b).

273 4.2 Horizontal Derivatives

274 The first horizontal derivative of the LRA (Fig. 5a) reveals strong gradients along
 275 the northeastern crater wall of Solfatara, around the edge of the low gravity region be-
 276 tween La Fangaia and Bocca Grande and more subdued gradients around La Fangaia
 277 and elsewhere in the crater. Prevailing NNE-SSW and NW-SE trends are highlighted
 278 by the gradients (Fig. 5b). The second horizontal derivative suggests similar fault trends
 279 (Fig. 5c and d). The structural trends obtained from total horizontal gravity gradiom-
 280 etry closely match field observations (Fig. 5e and f).

281 4.3 Sub-surface Distribution of Anomalous Mass

282 The optimal LRA and CRA models have a balance factor of 40, an auto-correlation
 283 of 0.14, and an a priori density contrasts of -450 to $+450$ kg m^{-3} (Fig. 6). The models
 284 image three main bodies of negative density contrast beneath Pozzuoli, Astroni volcano
 285 and Solfatara volcano. Although it is difficult to directly relate mathematically derived

286 density contrasts with rock density contrast, the optimal density range matches the 1σ
 287 range in rock densities about an average of 2300 kg/m^3 encountered in boreholes from
 288 Campi Flegrei (Piochi et al., 2014). Despite their inherent non-uniqueness, the models
 289 consistently provide robust results on the density variations at depth for different a pri-
 290 ori density contrasts (Supplementary Figures S6 and S7). The dominant anomalous neg-
 291 ative density bodies persist in all inversions, although as expected they become larger
 292 and more interconnected with decreasing a priori density contrasts.

293 The optimal LRA inversion images the Solfatara/Pisciarelli anomaly as approx-
 294 imately 0.5 km wide and extending from close to the surface to 0.8 km below sea level.
 295 The anomalous body beneath the Pozzuoli area is 1 km in diameter at its widest and ex-
 296 tends from ~ 0.5 km to 1.2 km depth. It is slightly elongated in the NNE-SSW direction.
 297 The Astroni anomaly is elongated E-W, 1.75 km across its widest point and extends from
 298 0.5 km to 1.4 km depth (Fig. 6).

299 The optimal anomalous bodies imaged by the CRA inversion are similar to those
 300 found for the LRA. However, the Pozzuoli anomaly vanishes and the anomalous bod-
 301 ies are imaged at a slightly shallower depth (Fig. 6). The long axis of the Astroni anoma-
 302 lous body is shifted slightly towards the north with respect to the LRA body.

303 Fig. 7 shows the surface traces of the -600 , -450 and -300 kg m^{-3} density isosur-
 304 faces. The inversions of both the LRA (Fig. 7a) and the CRA data (Fig. 7b) consistently
 305 image the Astroni and the Solfatara/Pisciarelli anomalies in the same locations. The Sol-
 306 fatara/Pisciarelli anomaly covers the SE edge of Solfatara crater and extends to Pisci-
 307 arelli in both cases. The Astroni anomaly is centered SW of Astroni crater and covers its
 308 SW wall. The imaging and co-location of the Astroni and Solfatara/Pisciarelli anoma-
 309 lies in both models is an indication of the robustness of the inversion, while the verac-
 310 ity of the Pozzuoli anomaly remains uncertain.

311 5 Discussion

312 5.1 Imaging of distinct reservoirs: Sub-surface controls on fluid distri- 313 bution

314 We present the first gravimetric image of the hydrothermal system at Campi Fle-
 315 grei caldera. Inversions of two differently-detrended data sets (LRA and CRA) provide
 316 robust and reproducible results and image two low-density reservoirs beneath Astroni

317 and Solfatara volcanoes, which we interpret as shallow-seated, fluid-rich hydrothermal
 318 reservoirs. An anomaly beneath Pozzuoli is only imaged by one of the models, which may
 319 be attributed to the lack of offshore gravity data in this survey, potentially preventing
 320 us to properly account for the effect of the caldera fill on the data at Pozzuoli. Both model
 321 results for Pozzuoli are plausible and alternative evidence is required to support the ex-
 322 istence or absence of a low-density reservoir beneath Pozzuoli (see below).

323 The optimal modeled negative density anomalies indicate a $\sim 20\%$ reduction in sub-
 324 surface density from background values. This can be explained by a porous and frac-
 325 tured caldera-fill containing hydrothermal fluids. Borehole data indicate drained host
 326 rock (dominantly volcanic tuff) densities ρ_r between 1600 and 2200 kg m⁻³ in the top
 327 1 km beneath the caldera containing between 5 and 40 vol% inherent void space (Piochi
 328 et al., 2014). To explain the modeled negative density contrasts a reduction in background
 329 bulk host rock density is required. In the hydrothermally active areas imaged in this study
 330 this can, for example, be achieved by the generation of additional (secondary) void space
 331 by fracturing and/or hydrothermal dissolution (scenario 1) or replacing the liquid phase
 332 in undrained porous host rock by a vapour phase (scenario 2). In the former case, the
 333 background bulk densities will be those reported above while, in the latter case undrained
 334 bulk densities of the caldera-fill are in range of 1650 to 2500 kg m⁻³ for given porosi-
 335 ties.

336 We first explore scenario 1 of bulk density reduction from an average background
 337 host rock density ρ_r of 1900 kg m⁻³. Assuming that the reduction in density is primar-
 338 ily driven by the creation of new void space ϕ that is fully connected and can host hy-
 339 drothermal fluids in either vapour (density $\rho_v=1.5$ kg m⁻³) and/or liquid (density $\rho_l=1000$
 340 kg m⁻³) form, the optimal anomalous density contrast $\Delta\rho$ of the reservoirs of ~ -400
 341 ± 25 kg m⁻³ can be explained by a multi-phase mixture of caldera-fill containing an ad-
 342 ditional interconnected void volume fraction of between 0.2 and 0.3 that contains a vapour
 343 volume fraction ψ_v of between 0.38 and 1 and a liquid volume fraction ψ_l fraction 0 and
 344 0.62. The parameter space of conceivable fractions of solids and voids (filled with vapour
 345 and/or liquid) that fit the optimal model for this scenario is shown in Fig. 8 and can be
 346 reproduced by

$$\Delta\rho = (1 - \phi)\rho_r + \phi\psi_l\rho_l + \phi\psi_v\rho_v - \rho_r. \quad (4)$$

347 Given the borehole rock porosity and density ranges, the second end-member sce-
348 nario (vapour replaces liquid in undrained caldera-fill) is only feasible in rocks contain-
349 ing inherent connected void fractions of 0.4 or more in order to explain the optimal den-
350 sity contrast, and is hence if at all only relevant for the top few hundred meters beneath
351 the surface (Piochi et al., 2014). While the most plausible interpretation of the optimal
352 models is a combination of processes associated with both explored scenarios, to explain
353 the modeled density contrast at depths > 250 m scenario 1 must be dominant and from
354 our gravity contrast model alone we favour the creation of additional void space. How-
355 ever, additional constraints are available to help explore the scenarios further.

356 The permeability structure of the central part of the caldera is key to understand-
357 ing the distribution of fluids at the time of the survey. Hot low-density fluids will rise
358 from their source until they attain neutral buoyancy, reach the surface or encounter a
359 barrier to flow, i.e., a zone of reduced permeability. The results suggest the presence of
360 fluid-rich bodies trapped beneath the surface of the CFc at Pozzuoli and Astroni vol-
361 cano while one body discharges freely at Solfatara volcano and Pisciarelli. This implies
362 the presence of an impermeable seal preventing access of fluids to the surface at Pozzuoli
363 and Astroni volcano. Geochemical and electric data indicate the presence of a two-phase
364 hydrothermal plumbing system at the CFc with a gas-dominated regime residing at shal-
365 low (few tens to hundreds of meters) depth beneath the centre of the caldera (Chiodini
366 et al., 2011; Gresse et al., 2017).

367 Permeabilities measured in-situ in boreholes at the CFc vary over 4 orders of mag-
368 nitude ($<10^{-18}$ to $>10^{-14}$) (Piochi et al., 2014). Total Horizontal Gradiometry of Sol-
369 fatara volcano (Fig. 5) shows a correlation of the geometry of low density bodies with
370 the main fault and fracture systems mapped in the field (Isaia et al., 2015). The com-
371 bination of in-situ rock permeability and fracture/fault permeability may explain the dis-
372 tribution of surface expressions of hydrothermal activity in the caldera and their spatio-
373 temporal evolution. Alunitic alteration at the Solfatara and Pisciarelli hydrothermal fields
374 increases rock porosity and permeability and reduces density (Mayer et al., 2016). Crit-
375 ically stressed faults can be hydraulically conductive (Jasim et al., 2015), while miner-
376 alisation can seal previously connected pathways within or around a fault (Sibson, 1994).
377 Faults can thus be both permeable pathways and impermeable inhibitors for fluid flow,
378 depending on the stress regime and degree of alteration. Fluids themselves can modu-
379 late permeability via thermally induced hydraulic fracturing (Knapp & Knight, 1977;

380 Cusano et al., 2008; Saccorotti et al., 2007). Permeability in hydrothermal systems is not
 381 static and changes constantly due to fracturing and cementation from hydrothermal pre-
 382 cipitation, meteoric water invasion or tectonic stresses (Rowland & Sibson, 2004). Faults
 383 at the CFc present fluid pathways on timescales of 1-10 years and 10^2 - 10^3 years (Vilardo
 384 et al., 2010). It is hence conceivable that geophysical surveys conducted over the past
 385 few decades at the CFc provide different snapshots of a constantly evolving hydrother-
 386 mal system. We therefore compare and contrast the Pozzuoli, Astroni and Solfatara/Pisciarelli
 387 anomalies imaged by our study with published results from other investigations (see also
 388 Table 1).

389 5.2 The Pozzuoli anomaly

390 Several studies found evidence for a shallow-seated hydrothermal reservoir beneath
 391 Pozzuoli, but its nature is contested:

- 392 • Vanorio et al. (2005) use seismic velocity tomography to delineate a zone of high
 393 V_p/V_s ratios (2.3) beneath Pozzuoli centred at approximately 0.8 km depth and
 394 0.8 km in radius. Below this, they image a low V_p/V_s (1.4) anomaly at 4 km depth.
 395 They interpret these features as a brine caused by steam condensation and a gas
 396 enriched formation, respectively.
- 397 • Chiarabba and Moretti (2006) find a high V_p/V_s region below Pozzuoli at 0-2 km
 398 depth, overlying a low V_p/V_s anomaly at 3 km depth which they interpret as steam
 399 condensation and gas accumulation, respectively.
- 400 • Seismic attenuation tomography by De Siena et al. (2010) in tandem with V_p/V_s
 401 ratios image an anomaly 0-2 km below Pozzuoli, with a ~ 1 km radius. The nature
 402 of this anomaly (liquid or gas dominated), however, remains ambiguous in the study.
- 403 • Chiodini et al. (2015) and Caliro et al. (2007) use geochemical models to predict
 404 the vaporisation of fluids at 2 km depth which then rise to the surface.
- 405 • A stacked (gas-rich pockets beneath liquid dominated systems) arrangement of
 406 fluids is predicted by fault-controlled fluid flow modelling for CFc (Jasim et al.,
 407 2015).

408 In summary, the available evidence is inconclusive regarding the nature of a shallow-
 409 seated reservoir beneath Pozzuoli with indications for either vapour or liquid dominated
 410 regimes. One aspect that needs consideration is the potential for temporal change in the

411 sub-surface phase relationships. Chiarabba and Moretti (2006) and M. Battaglia et al.
 412 (2006) use combined data from 1984 and from 2001, while Vanorio et al. (2005) and De Siena
 413 et al. (2010) used data from 1984, only. Not only do the combined datasets risk mask-
 414 ing of temporal signals, but the system may change over the course of 15 years. Tem-
 415 poral changes in elasticity of the upper crust have been suggested for CFc (Di Luccio
 416 et al., 2015). Cycles of sealing, fracturing and resealing are implied on timescales of decades
 417 to centuries, for example, at Yellowstone caldera where drill cores plugged almost com-
 418 pletely after 25 years (Dobson et al., 2003; Ingebritsen & Sorey, 1988). Our modelling
 419 results are consistent with either a liquid or vapour dominated system with the caveat
 420 that a liquid-dominated regime requires a significantly higher connected porosity com-
 421 pared to a vapour-dominated system to explain the gravity data (Fig. 8).

422 5.3 The Astroni anomaly

423 The Astroni anomaly is unexpected as there are no records of fumaroles or other
 424 geothermal manifestations in the area. The anomaly is located at the convergence of two
 425 crater walls and one might expect high density material here compared to adjacent low
 426 density crater fill. However, there are several geophysical anomalies associated with Ag-
 427 nano caldera (Isaia et al., 2009), in which Astroni volcano is nested:

- 428 • Astroni volcano is seismically active (Chiadini et al., 2017; Saccorotti et al., 2007).
 429 de Lorenzo et al. (2001) and De Siena et al. (2010) image a seismic anomaly be-
 430 neath Agnano at 0-3 km depth which they relate to a high temperature aquifer.
- 431 • Capuano et al. (2013) identify an E-W elongated low-gravity anomaly at Astroni
 432 volcano at 1-2 km depth, which they interpret as a low density gas-rich reservoir
 433 with secondary mineral precipitation as a mechanism for preventing surface ex-
 434 pression. They suggest that hot gases condense near the surface and generate a
 435 water table which in the Agnano plain forms lake Agnano.
- 436 • Troiano et al. (2019) identify a highly resistive (vapour-rich) anomaly correspond-
 437 ing to the south-eastern edge of the Astroni crater, while the Agnano plain is char-
 438 acterised by a mainly conductive anomaly.
- 439 • Water from the Agnano well has high P_{CO_2} values, less negative carbon isotope
 440 signatures than meteoric water and high HCO_3^{-2} concentrations, indicating a con-

441 tribution of heat and hydrothermal fluids from a magmatic system (Venturi et al.,
442 2017).

443 In light of these geochemical and geophysical observations we suggest that the grav-
444 ity anomaly of Astroni volcano is formed by a liquid-dominated highly fractured geother-
445 mal reservoir trapped beneath a less permeable cap rock. Localised seismicity may be
446 due to the movement of fluids along faults and it is plausible that there are pathways
447 linking Solfatara/Pisciarelli and Astroni volcano via a network of mainly NW-SE and
448 NE-SW to NNE-SSW trending faults.

449 **5.4 The Solfatara/Pisciarelli anomaly**

450 The key features of the Bouguer anomalies at Solfatara/Pisciarelli are i) the grav-
451 ity highs on the SW and NE crater walls of Solfatara(>0.5 mGal), ii) the gravity low
452 on the southeastern side of Solfatara’s crater floor (>-1 mGal) and adjacent Pisciarelli
453 and iii) the moderate gravity low near La Fangaia (~-0.3 mGal) (Fig. 4). These com-
454 pare to the findings from an earlier survey of Solfatara (Oliveri del Castillo et al., 1968):
455 i) an elongate ~ 1 mGal gravity high on the NE crater wall, ii) two connected gravity lows
456 reaching from the southern edge of La Fangaia (~-0.3 mGal) to Bocca Grande and Bocca
457 Nuova (~-0.4 mGal), and iii) a small scale ~-0.3 mGal gravity low in the western side
458 of the crater (see Fig. 9). Bruno et al. (2007) demonstrated the spatial correlation of the
459 first three anomalies of the 1968 survey with areas of maximum seismic noise, areas of
460 high CO₂ degassing and elevated temperatures. The horizontal derivative of the 1968
461 gravity data, highlights the role of faults in concentrating these density anomalies (Bruno
462 et al., 2007).

463 We divide the anomalies of our survey into three classes (relative to the average
464 gravity of the crater floor), i) low, ii) moderate low and iii) high to discuss their relation-
465 ship with other recent geophysical observations at Solfatara/Pisciarelli (see Fig. 9 and
466 Table 1).

467 **5.4.1 Low Gravity Class**

468 The low gravity anomaly extends from the eastern side of the crater floor and crosses
469 the crater wall towards Pisciarelli. It coincides with some of the main fumaroles of the
470 geothermal field, which at Solfatara are clustered on its eastern side, where the rocks are

471 intensely fractured (Isaia et al., 2015). We explain the correlation of the gravity low and
472 areas of intense gas emissions in two ways.

473 Firstly, upwelling vapour replaces water in fractures and pore spaces in the rock
474 causing a local decrease in density.

475 Secondly, low density material is expected from intense hydrothermal alteration near
476 the fumaroles where connected porosities can reach values of up to 61 vol% (Mayer et
477 al., 2016) with little variation in hydraulic parameters expected to depths of ~ 500 m (Montanaro
478 et al., 2016). Proximal areas around the fumaroles should therefore experience a sub-
479 stantial reduction in density due to the coupled effects of fluid flux and alteration.

480 At the time of the survey the most vigorous fumarolic activity was located at Bocca
481 Nuova, Bocca Grande and Pisciarelli, with no fumarolic activity noted on top of the crater
482 wall. These fumaroles are also the hottest of those in the Solfatara/Piscarelli area (Chiodini
483 et al., 2001). Isaia et al. (2015) have suggested a link between Bocca Nuova, Bocca Grande
484 and Pisciarelli via faulting and fracturing through the crater wall. This region, there-
485 fore, may be the main pathway for fluids to ascend from depth. Other fumaroles may
486 be fed less voluminously and/or by narrower, subordinate fracture networks, which are
487 below the spatial resolution capability of the gravity survey. The strongest first and sec-
488 ond horizontal gravity gradients (see Fig. 5a and b) are across the crater floor and along
489 the NE crater indicating that the anomalies are strongly influenced by faults and frac-
490 tures.

491 Di Giuseppe et al. (2015) found a high resistivity zone close to Bocca Nuova and
492 Bocca Grande, but little correspondence between resistivity and the other fumaroles. The
493 authors attribute the offset between the resistive body and the 1968 gravity low (see Fig.
494 9) to fluid migration over the time between surveys, but our low gravity anomaly encom-
495 passes both the high resistivity body and the 1968 gravity low. Similar to the 1968 sur-
496 vey, our gravity low increases east-wards from the main fumaroles (Fig. 9) towards Pis-
497 ciarelli.

498 Solfatara undergoes spatio-temporal variations in ground deformation (D’Auria et
499 al., 2012) and has high levels of seismic noise in an arcuate band from the south to the
500 northeast of the crater floor (Bruno et al., 2007). Saccorotti et al. (2007) show maximum
501 likelihood locations of long-period (LP) earthquakes clustered at 500 m depth beneath

502 the SE rim near Bocca Grande. They interpret the LP signals as due to vibrations of
 503 fractures in a buried cavity filled by a water-steam mixture. An interpretation of ultra-
 504 high resolution seismic imaging of the centre of Solfatara divides the first 30 m into a shal-
 505 low zone of aerated tephra underlain by a liquid saturated layer that is deepening in the
 506 direction of La Fangaia and a deeper gas accumulation sloping upwards towards the east-
 507 ern side of the crater (De Landro et al., 2017). Seismic attenuation, a shear-wave veloc-
 508 ity anomaly and low V_p/V_s ratios (Chiarabba & Moretti, 2006; De Siena et al., 2010; Chio-
 509 dini et al., 2015) indicate a shallow gas reservoir around sea level beneath Solfatara.

510 Byrdina et al. (2014) show high resistivity anomalies in both the crater walls and
 511 beneath the crater floor. Bocca Grande is directly above a narrow zone of moderate re-
 512 sistivity and surrounded by a zone of high temperature, high CO₂ and low self-potential.
 513 To the east of Bocca Grande is a high resistivity anomaly, and to the west is a low re-
 514 sistivity zone extending towards La Fangaia. Magnetotelluric (MT) and electrical resis-
 515 tivity (ER) data (Troiano et al., 2014, 2019) depict a moderately resistive anomaly be-
 516 low the eastern crater wall and a high resistivity anomaly beneath the main fumaroles
 517 and Pisciarelli to 2.25 km depth with a radius of ~ 0.15 km. High-resolution ER tomog-
 518 raphy images a gas-dominated reservoir at 60 m depth beneath Solfatara’s crater floor
 519 that feeds Bocca Grande (Gresse et al., 2017).

520 In summary, we propose that the main gravity low is caused by a shallow-seated
 521 (<1000 m depth b.s.l.) accumulation of a two-phase fluid within highly fractured and
 522 porous host rocks (Fig. 8). It is plausible that the eastern crater wall is composed of highly
 523 altered rocks with elevated porosity compared to the rest of the rim, indicating relict and/or
 524 current fluid pathways. The imaged feeder system appears to encompass the most dom-
 525 inant pathway for ascending fluids in the central sector of the caldera through a com-
 526 bination of Solfatara’s maar-diatreme structure (Troiano et al., 2019) and its intersec-
 527 tion with the dominant fault systems (NW-SE and NE-SW to NNE-SSW) of the caldera.

528 **5.4.2 Moderate Low Gravity Class**

529 A moderate gravity low is located near La Fangaia (Fig. 9) and matches the ex-
 530 tent of the 1968 gravity low. The exact location of the shallow-most La Fangaia feeder
 531 system is hard to establish as it appears to change with time. Dried up pits were present
 532 during the survey which must previously have been mudpools. We therefore use the fenced-

533 off area to delineate La Fangaia (brown shape in Fig. 9), although its most active por-
 534 tion was the southern part of the area at the time of the survey. Gentle first and sec-
 535 ond horizontal gravity gradients bound La Fangaia and show the same general direction
 536 as faults mapped in the field (Isaia et al., 2015) (Fig. 5a and b). The moderate gravity
 537 low of La Fangaia is also characterised by high CO₂ flux, low resistivity (with a long lobe
 538 extending westwards beneath the surface), low self-potential, high temperature, elevated
 539 seismic noise and earthquake clustering (Byrdina et al., 2014) (Table 1). These authors
 540 report a positive correlation between CO₂ flux and ground temperature, which are both
 541 anti-correlated with self-potential. The lobate geometry of low resistivity is matched by
 542 the moderate gravity anomaly beyond the boundary of La Fangaia. Seismic noise is high
 543 near the gravity anomaly and has been correlated with anomalous CO₂ degassing (Bruno
 544 et al., 2007).

545 Byrdina et al. (2014) interpret the La Fangaia geophysical anomalies by a liquid
 546 saturated plume with both a downwelling condensing liquid water and an upwelling vapour
 547 and CO₂ mixture. The water table is locally raised at Solfatara, outcropping at La Fan-
 548 gaia (97 m a.s.l) and only 7 m below the surface at the OAK well nearby (see Bruno et
 549 al. (2017)). A two phase (gas and liquid) flow regime feeding La Fangaia is also proposed
 550 by numerical modelling (Rinaldi et al., 2011). This suggests that background densities
 551 for the crater floor are influenced by the presence of liquid water.

552 The moderate gravity low is thus likely formed by a CO₂-bearing hot aquifer con-
 553 tained within altered and high-porosity crater-fill.

554 **5.4.3 High Gravity Class**

555 While the depicted gravity highs are constrained only by a low number of bench-
 556 marks, the anomalies coincide with the location of the Solfatara cryptodome (northeast-
 557 ern high) and the Mount Olibano lava dome (southern high) (Isaia et al., 2015) (see Fig.
 558 9). The northeastern high matches a gravity high detected by the 1968 gravity survey.
 559 Although poor accessibility prevented us from obtaining more measurements on the south-
 560 ern and northeastern rims of the crater, the transition from low density crater fill to the
 561 high density Solfatara cryptodome is well marked by a strong first horizontal gradient
 562 of the Bouguer anomaly. We therefore interpret the gravity highs as remnant domes form-
 563 ing part of the crater rim.

6 Conclusions

The combined use of high-precision gravity and GPS measurements, high-quality DEMs and 3D data inversion has shed light on the shallow-seated hydrothermal system at Campi Flegrei. The results complement a wealth of existing geophysical and geochemical data for the CFC and lend support to a model of a complex sub-surface hydrothermal structure feeding the active fumarolic areas of the central sector of the caldera. We were able to delineate the shallow-seated two-phase hydrothermal plumbing system beneath Solfatara and Pisciarelli and identified a hydrothermal reservoir beneath the Astroni crater. The main gravity anomalies of Solfatara volcano detected in the new survey broadly match those identified by a previous gravity survey conducted in 1968. However, we show that some smaller anomalies may have evolved in size and location over time. This may indicate that within the resolution capabilities of the Bouguer gravity surveys, Solfatara's main hydrothermal feeder system remained broadly unchanged over the past 50 years with the exception of an enlargement towards Pisciarelli, which over the past 15 years has seen a strong increase of hydrothermal activity. Whether or not a separate hydrothermal system resides beneath Pozzuoli cannot be unambiguously answered by our findings, but there are indications for a shallow-seated low-density hydrothermal reservoir during the time of our survey. We encourage additional geophysical and geochemical studies particularly at Astroni volcano and Pozzuoli to test our model results.

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8 Autor contributions

JG conceived the project and received funding. RI facilitated fieldwork and contributed to data collection. NY and JG collected and processed all gravity and GPS data. N.Y. analysed, modeled and interpreted all data as part of her doctorate. JG took the lead in writing the manuscript based on material presented in NY's doctoral thesis with additional data analysis and interpretation. All authors provided critical feedback and helped shape the final version of the manuscript.

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Table 1. Relevant previous surveys of Solfatara/Pisciarelli (marked S) and the wider Campi Flegrei caldera (marked CFc) pertinent to the hydrothermal system of the caldera’s central sector. Question marks (?) indicate undisclosed or unclear survey dates.

Survey Type	Survey Location	Year of Survey	Author
Bouguer Gravity	S	1968	(Oliveri del Castillo et al., 1968)
	S,CFc	2015	This study
Electrical Resistivity	S	>2008	(Di Giuseppe et al., 2015)
	S	2008-2012	(Byrdina et al., 2014)
	S	2013	(Isaia et al., 2015)
	S	2008-16	(Gresse et al., 2017)
	S	2018	(Troiano et al., 2019)
Magnetotellurics	S	2000?	(Bruno et al., 2007)
	S	2012	(Troiano et al., 2014)
Self Potential	S	2011	(Byrdina et al., 2014)
Gas Measurements	S	2011	(Byrdina et al., 2014)
CO ₂ flux	S	2012-2019	(Tamburello et al., 2019)
Hydrogeological	S	2000?	(Bruno et al., 2007)
InSAR	CFc	1992-2001, 2003-2007	(Vilardo et al., 2010)
	S,CFc	1995-2007	(D’Auria et al., 2012)

Figure 2. Digital Elevation Models (DEMs) and benchmark locations. a) 10 m DEM of Campi Flegrei caldera with the outline of the major collapse structure (dashed line). Survey benchmarks outside the central sector are shown in blue. Pozzuoli (the centre of ground deformation), Solfatara and Pisciarelli are labelled Po, S and Pi, respectively. b) 1 m DEM of the central sector with the benchmarks in blue including the densely spaced benchmark array in the Solfatara crater.

Table 2. Cont'd.

Survey Type	Survey Location	Year of Survey	Author
P/S-wave velocity	S,CFc	2006?	(Zollo et al., 2006)
	S,CFc	2008	(J. Battaglia et al., 2008)
	S,CFc	2005	(Vanorio et al., 2005)
Seismic reflection	CFc	2008?	(Zollo et al., 2008)
Microseismic	S	2000?	(Bruno et al., 2007)
P wave seismic refraction	S	2000?	(Bruno et al., 2007)
Raleigh wave	S	2000?	(Bruno et al., 2007)
Raleigh waves s-wave velocity	S	2001	(Petrosino et al., 2006)
Seismic tomography	S	2009	(Letort et al., 2012)
Seismic attenuation (passive)	S,CFc	1984	(De Siena et al., 2010)
Seismic noise	S	2007	(Petrosino et al., 2012)
S-wave spectra (passive)	S,CFc	2004-2006	(Saccorotti et al., 2007)
Seismic tremor	S,CFc	2000-2019	(Giudicepietro et al., 2019)
Structural geology	S	2007?	(Petrosino et al., 2012)
	S	2013	(Isaia et al., 2015)
Thermal survey	S	2009-2011	(Byrdina et al., 2014)
Volcanological	S	2007?	(Petrosino et al., 2012)
	S	2012-14	(Isaia et al., 2015)

Figure 1. The Campi Flegrei caldera (14.1°E and 40.8°N) and its main fumarolic areas central sector. a) Google Earth image of the caldera indicating the approximate footprint of ground deformation observed since 1982 (dashed line) and the location of Solfatara and Pisciarelli fumarolic fields (black box). b) Details of the hydrothermal expressions of the Solfatara and Pisciarelli fumarolic fields: Solfatara’s fumaroles Bocca Nuova (BN), Bocca Grande (BG) and Le Stufe (LS) and the fence around La Fangaia (LF) are marked by black dashed lines. The main mud-pools and fumaroles of Pisciarelli (Pi) are circled by a dashed yellow line.

Figure 3. Regional and local Bouguer anomaly maps. a) shows the regional Bouguer anomaly, while b) and c) show the local Bouguer anomaly after detrending the data presented in a) for two different components: b) a regional long-wavelength trend to obtain the linearly-detrended residual anomaly (LRA) and c) an additional trend caused by the caldera-fill sediments to obtain the caldera-fill detrended residual anomaly (CRA). The colour bars are scaled to the maximum absolute gravity of the three datasets. Contours are in mGal and benchmarks are shown in white. The data is overlain on the 10 m DEM and Pozzuoli, Solfatara volcano and Astroni volcano are marked by letters P, S and A, respectively.

Figure 4. Local Bouguer anomaly of Solfatara crater. a) LRA, b) CRA. Color bars are not scaled. Contours are in mGal and benchmarks are shown in white. The data is overlain on the 1 m DEM. Black stars mark the locations of La Fangaia (left) and Bocca Grande and Bocca Nuova (right).

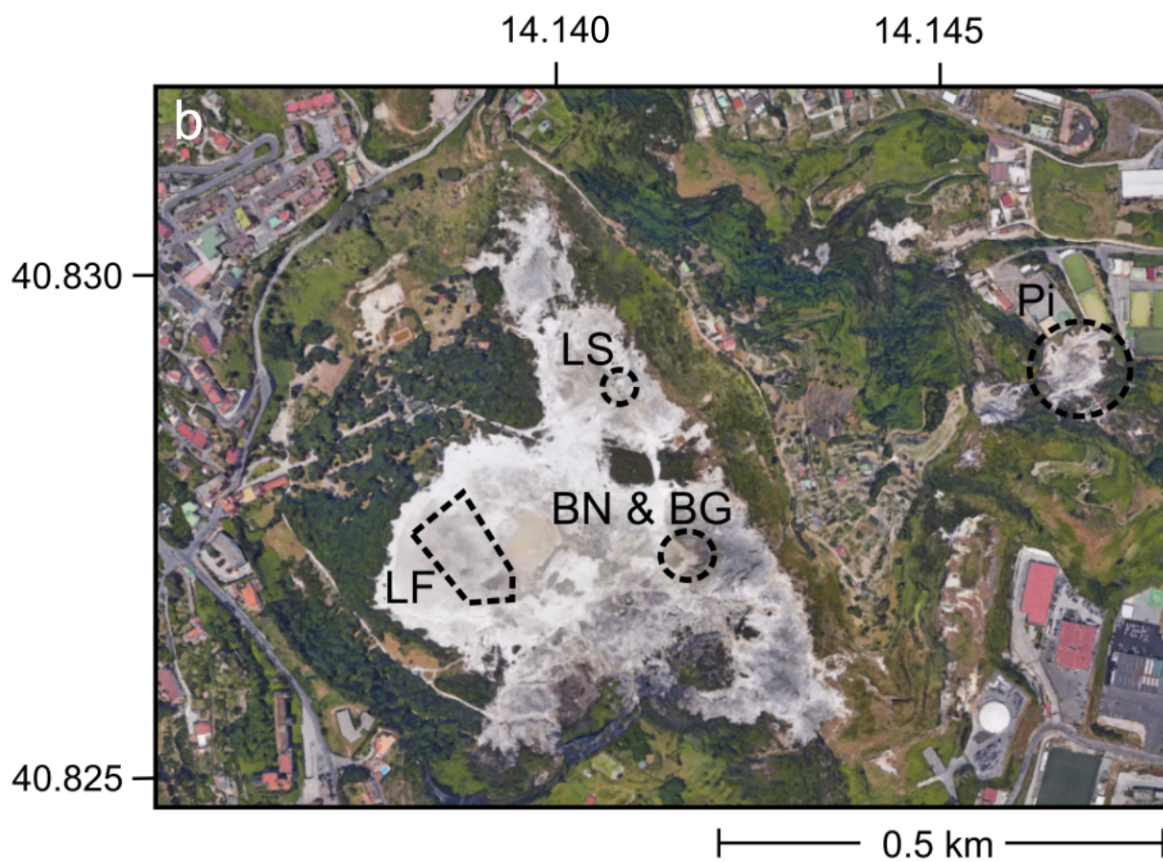
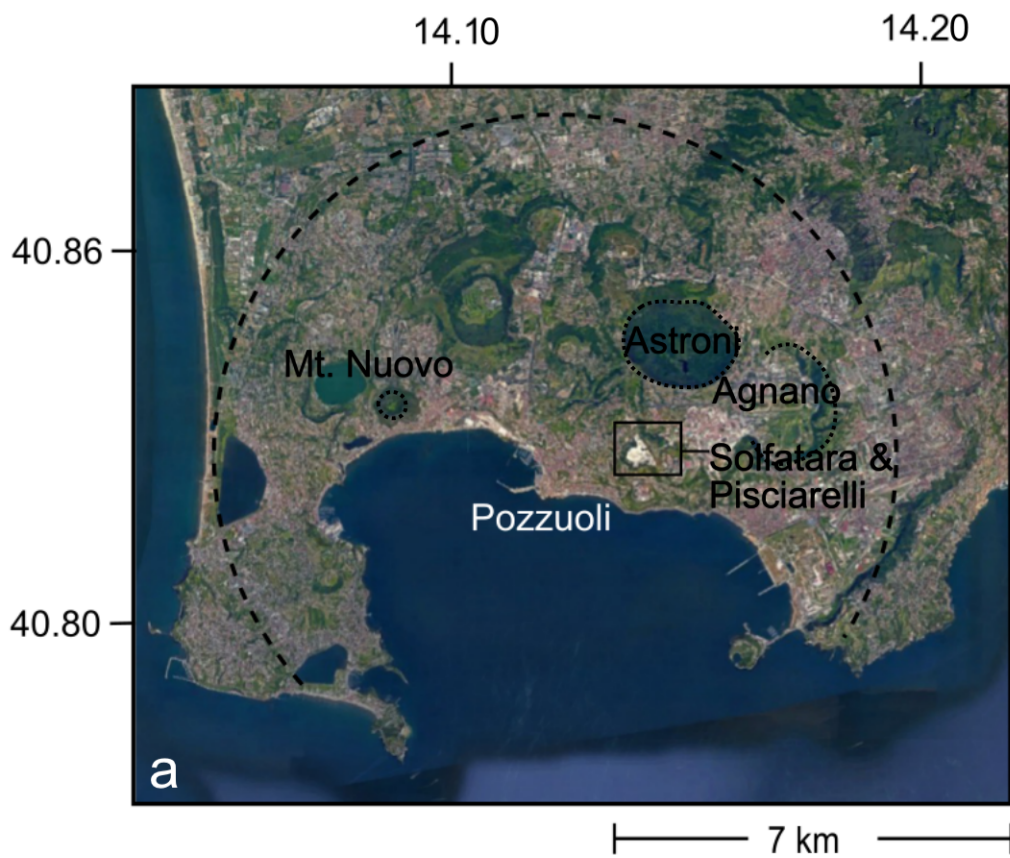
Figure 5. First and second total horizontal derivatives of the LRA anomaly and comparison with field structural data. a) displays the first total horizontal derivative (THD1) with interpreted faults shown as dashed black lines and accompanying rose diagrams in b). c) and d) show the same for the second total horizontal derivative (THD2) with interpreted faults marked by dashed yellow lines. e) displays the DEM of the crater floor and mapped faults (Isaia et al., 2015) as dashed blue lines with f) showing the related structural statistics. Data are binned in 10 degree increments in all rose diagrams (b, d and f). f) shows mapped fault orientations (top, dark blue), fracture orientations (bottom, light blue) and the accompanying dip angles presented in (Isaia et al., 2015). Benchmarks are shown by black crosses. Black stars mark the locations of La Fangaia (left) and Bocca Grande and Bocca Nuova (right).

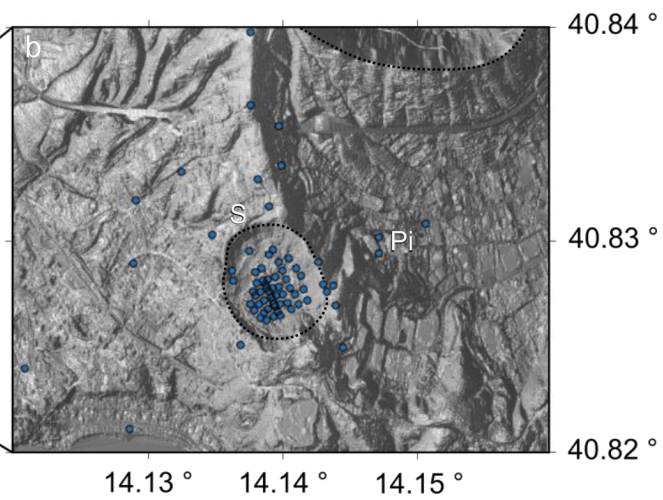
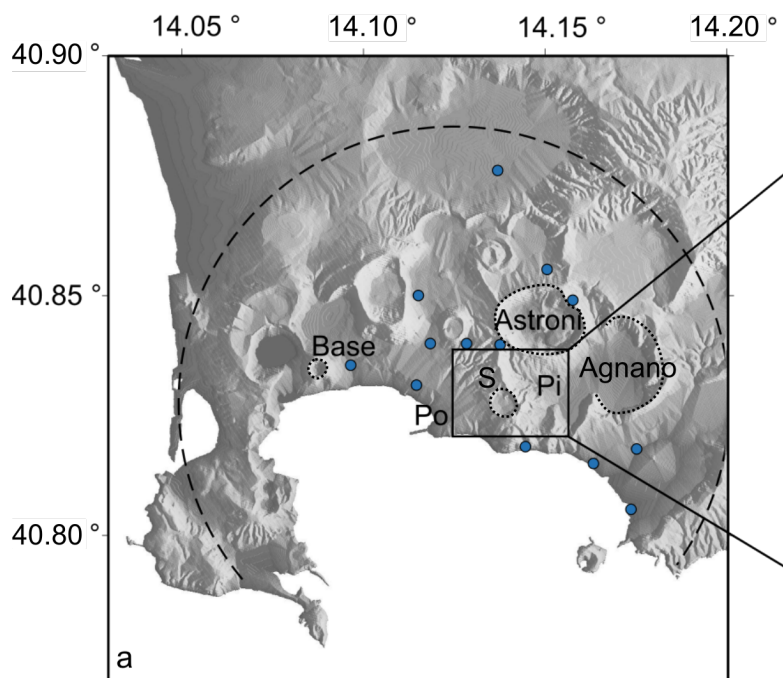
Figure 6. -400 kg m^{-3} density contrast isosurfaces from the optimal inversions for an a priori density contrast of $\pm 450 \text{ kg m}^{-3}$. a) LRA inversion in plan view. b) LRA inversion with a view facing 59° NE. c) CRA inversion in plan view. d) CRA inversion with view facing 59° NE. e) LRA and CRA inversions in plan view. f) LRA and CRA inversions with view facing 59° NE. The main anomalous bodies are located beneath Pozzuoli, Solfatara/Pisciarelli volcano and Astroni volcano and are labelled P, S and A, respectively.

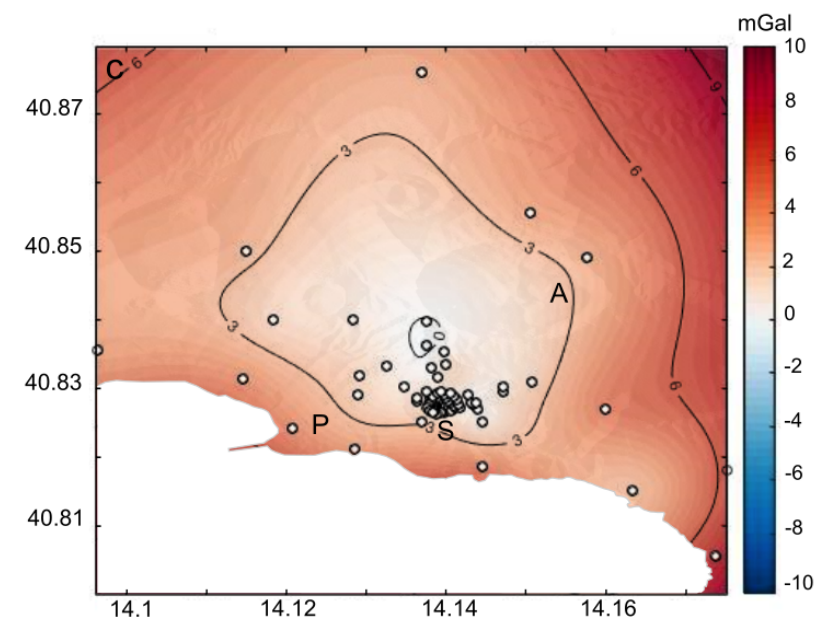
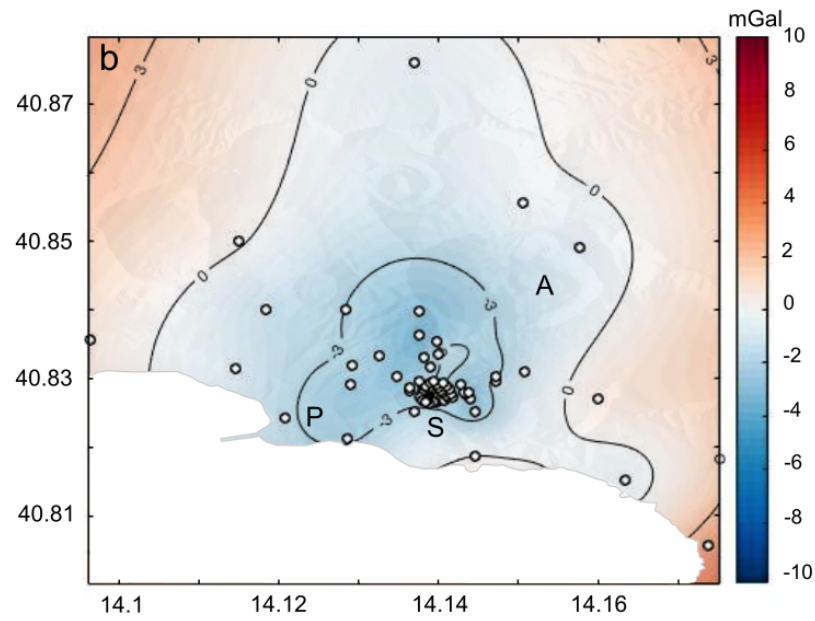
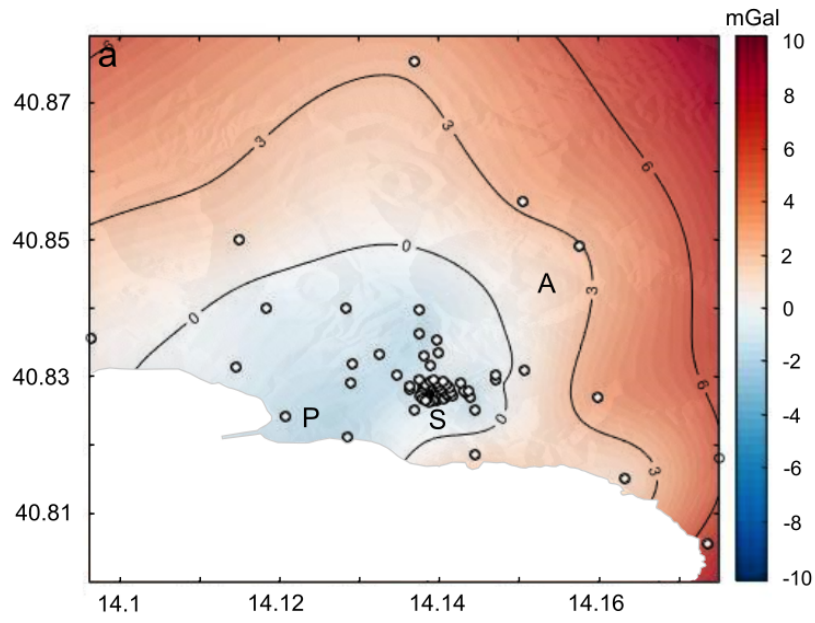
Figure 7. Surface traces of anomalous bodies for density contrasts of -600 , -450 and -300 kg m^{-3} represented by dotted, solid and dashed lines, respectively. a) LRA model. b) CRA model. Location labelings as in Fig. 6.

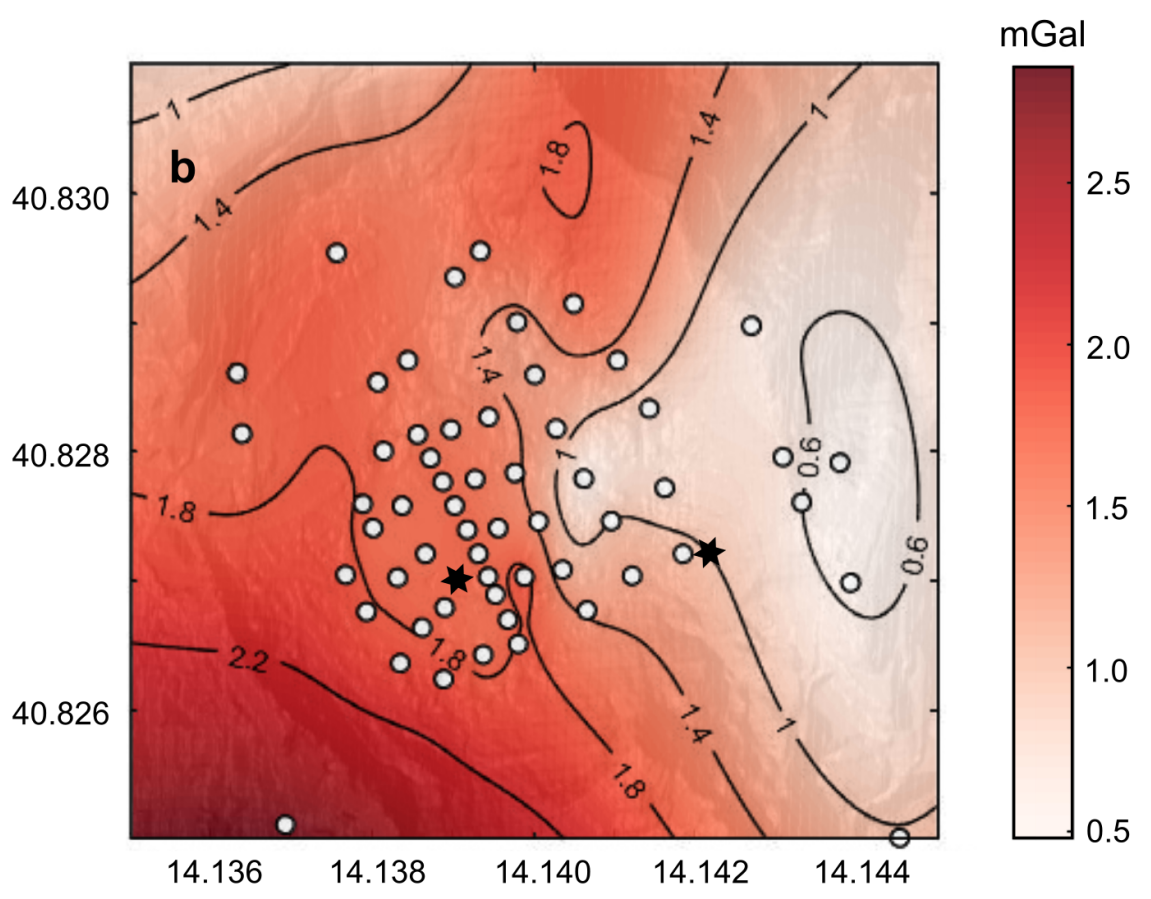
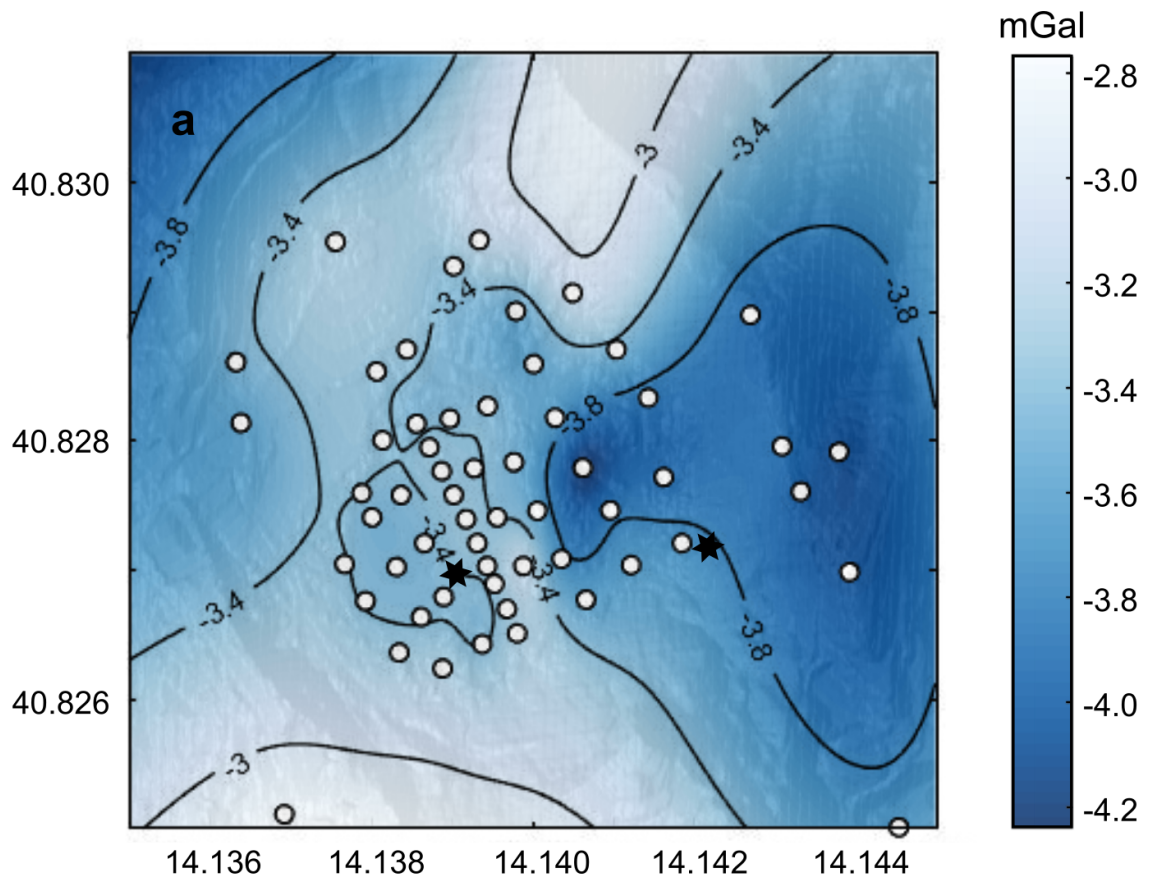
Figure 8. Diagram of the bulk density contrast parameter space of the proposed hydrothermal reservoirs for scenario 1 (see text for further explanation). A range of phase (solid, liquid, vapour) volume fraction combinations can explain the optimal density contrast of $\sim -400 \pm 25 \text{ kg m}^{-3}$ and include porous and fractured caldera-fill (host) rocks with an average bulk density of 1900 kg m^{-3} for depths between the surface and 1 km (Piochi et al., 2014)) containing hydrothermal vapour (1.5 kg m^{-3}) and/or liquid (1000 kg m^{-3}) in secondary void space. The optimal anomalous density contrast is highlighted by the gray-colored area and can be explained by a multi-phase mixture of host rock containing a volume fraction of voids between 0.2 and 0.3 that contains a vapour volume fraction of between 0.38 and 1 and a liquid volume fraction of between 0 and 0.62. The red circle marks one of the possible permutation of volume fractions for illustration: solids (0.76) and voids (0.24) with the latter containing vapour (0.8) and liquid (0.2). The colour bar shows the bulk density contrast in kg m^{-3} .

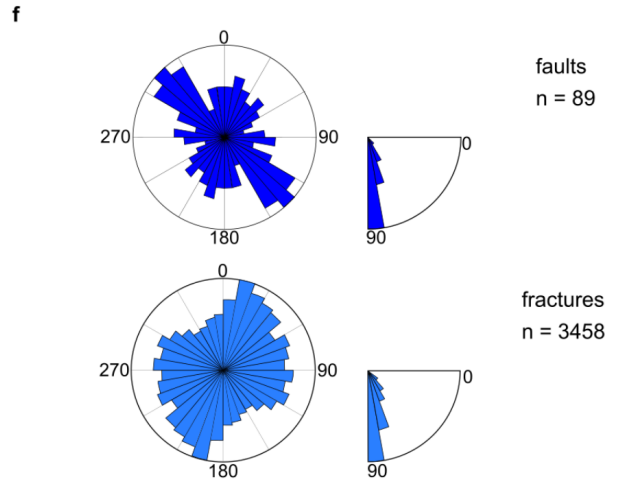
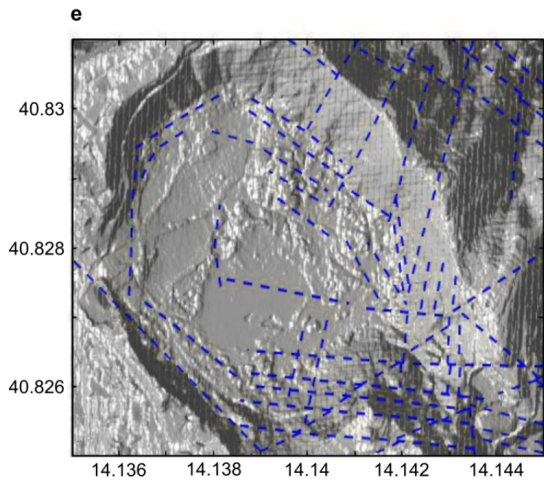
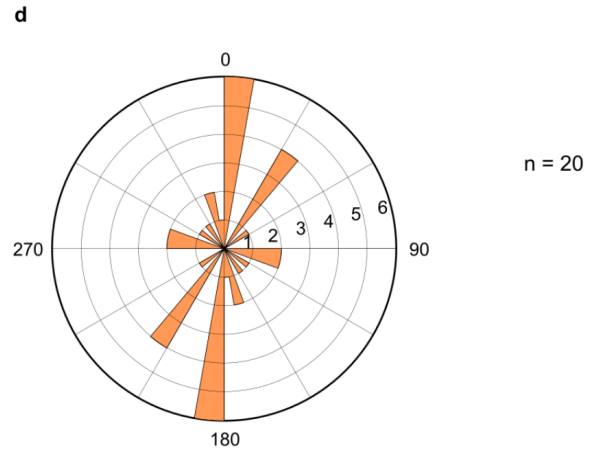
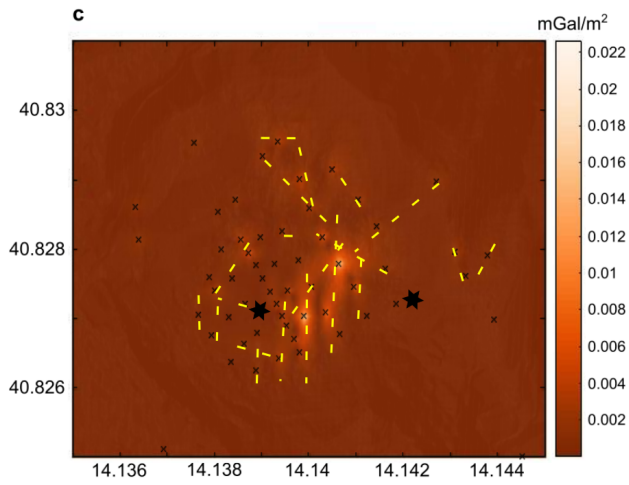
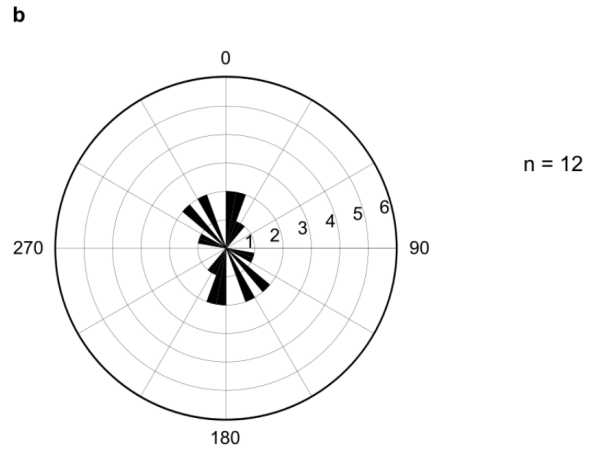
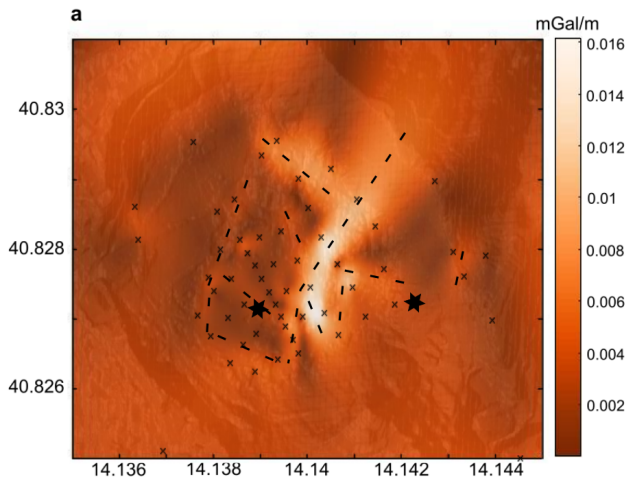
Figure 9. Comparison sketch of previous electrical resistivity (2015) and gravity (1968) anomalies of the Solfatara crater with the LRA data from this study. Dotted lines denote low amplitude anomalies, dashed lines denote moderately low amplitude anomalies and solid lines denote high amplitude anomalies. The crater outline, Pisciarelli, La Fangaia, fumaroles and lava domes are shown. Colour scheme for the LRA gravity anomalies is given in the caption.

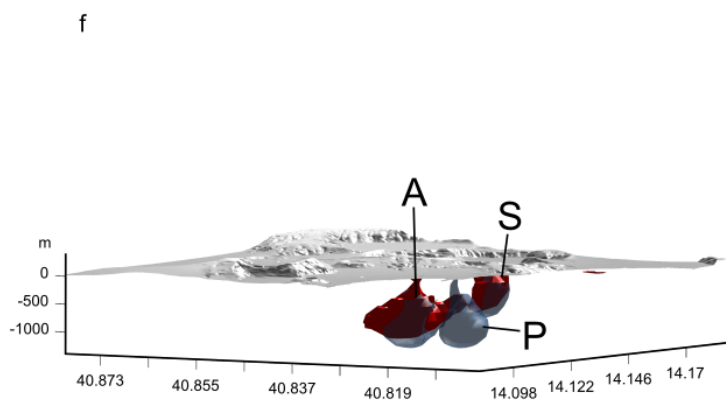
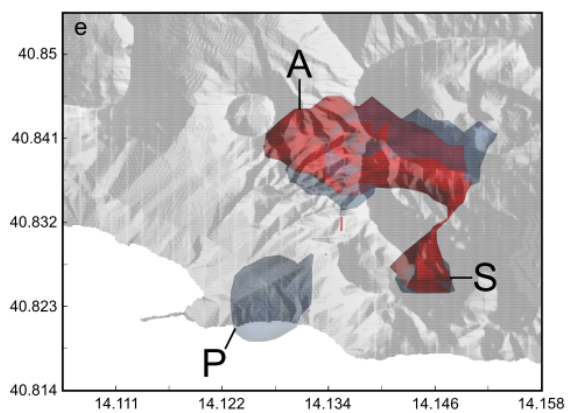
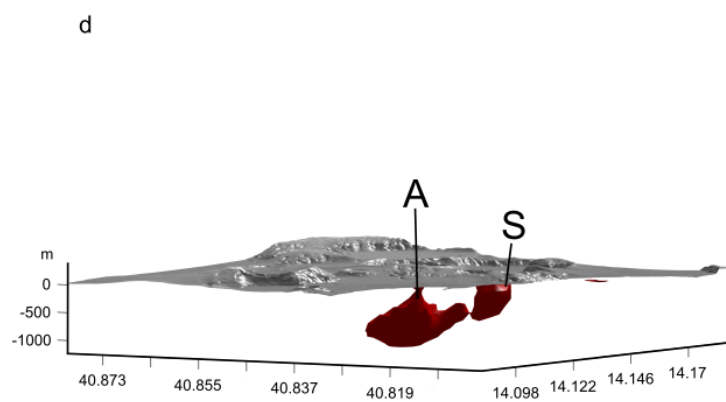
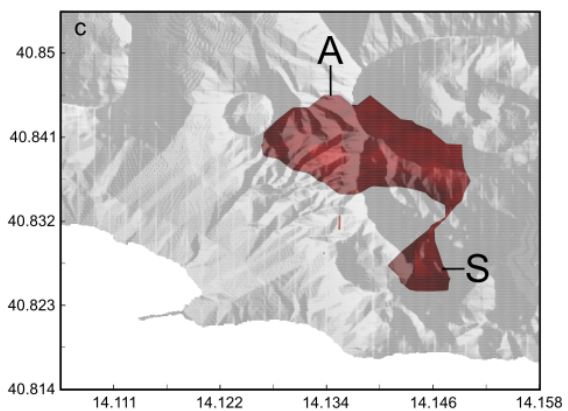
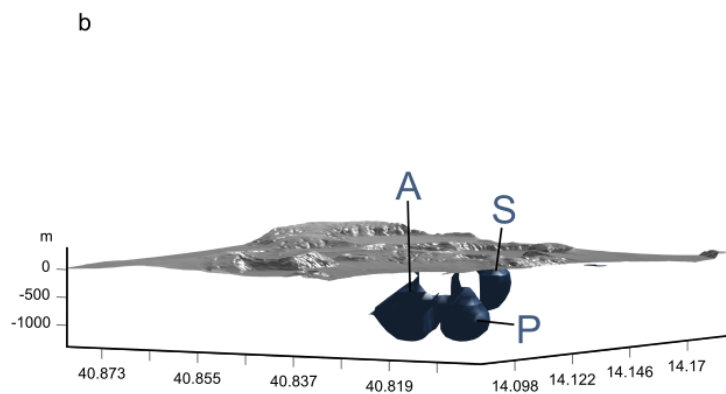
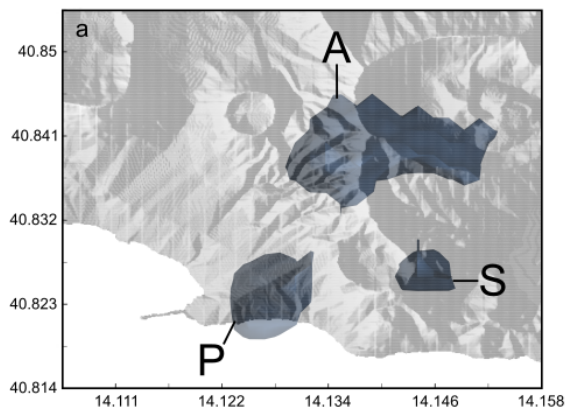


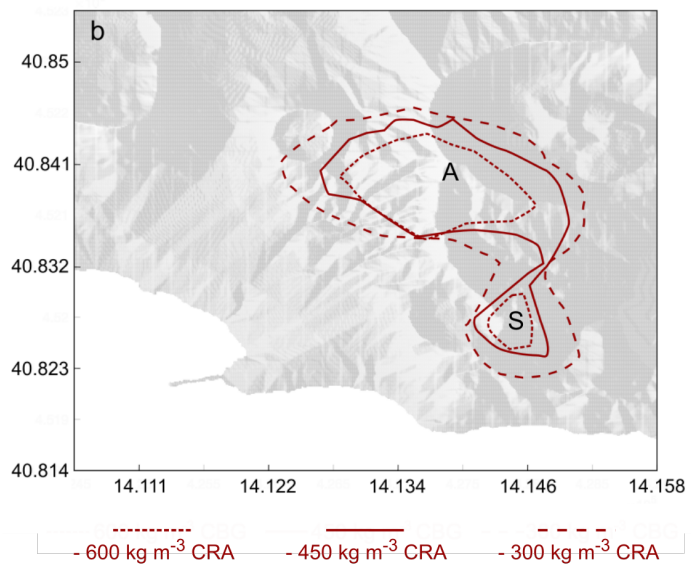
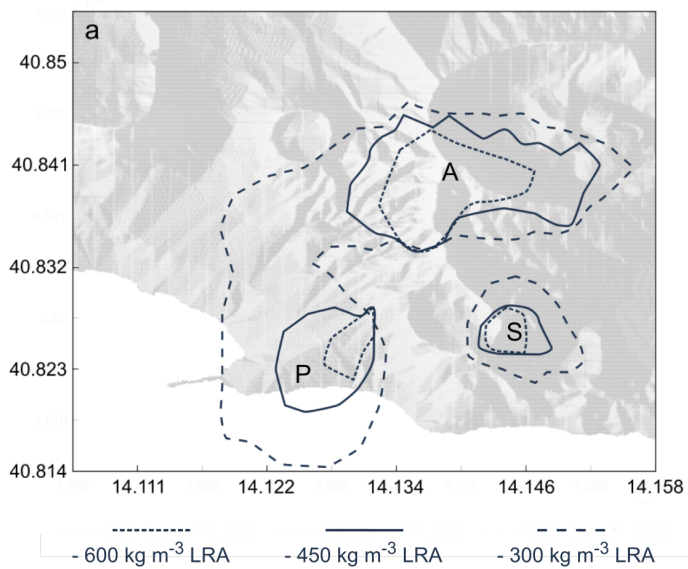


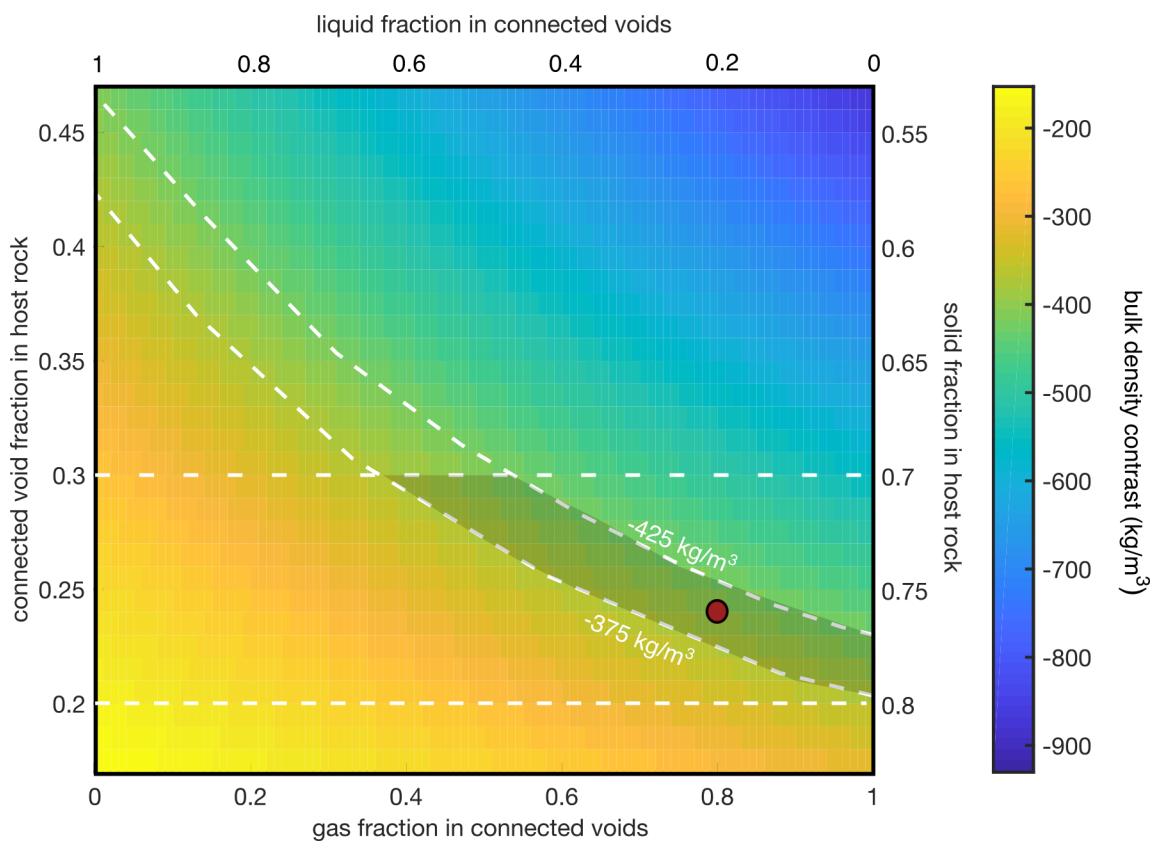


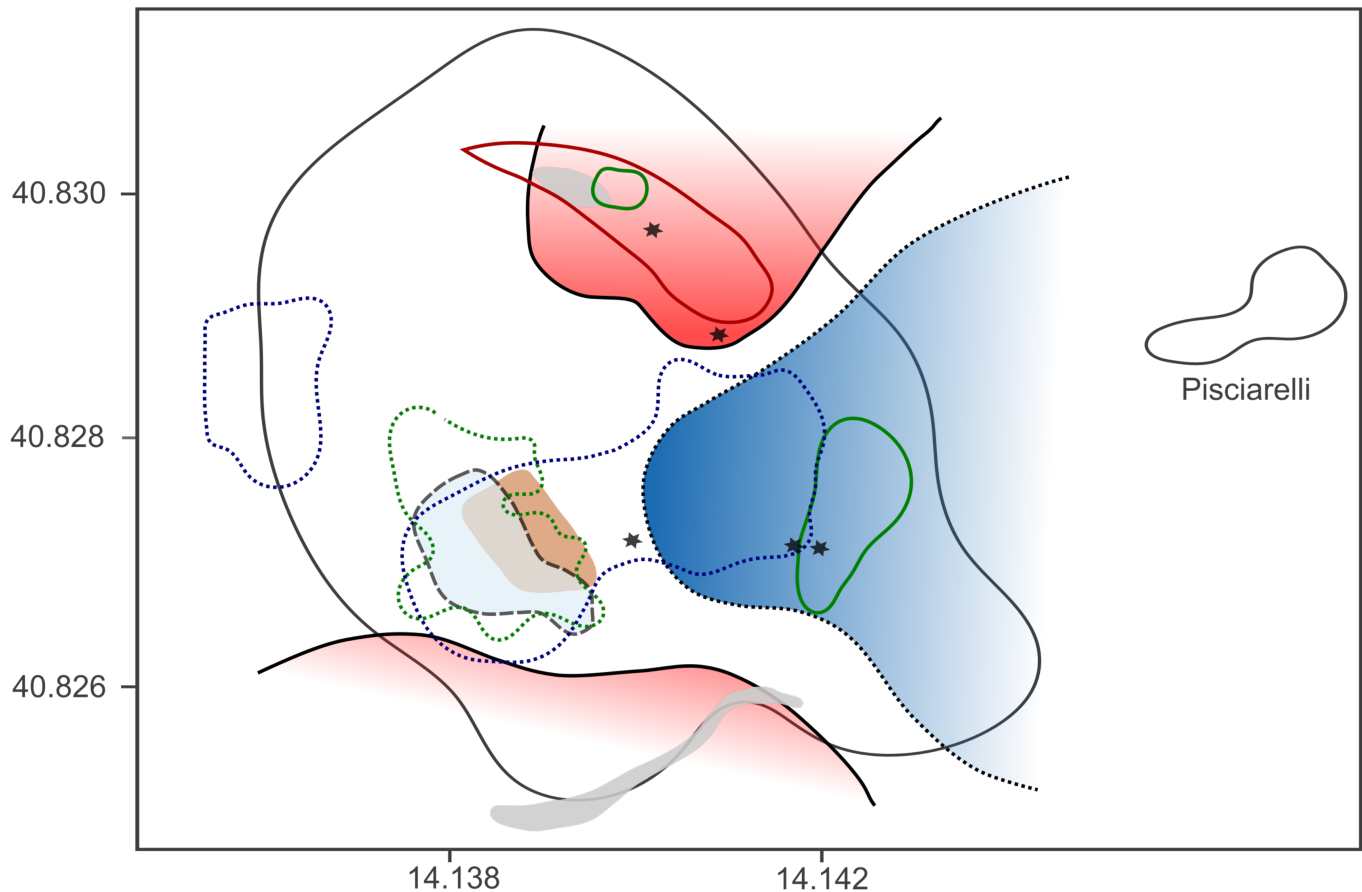












- ★ Fumaroles
- Crater rim
- Lava Dome
- La Fangaia
- High Resistivity (Di Giuseppe et. al 2015)
- ⋯ Low Resistivity (Di Giuseppe et. al 2015)
- High Gravity (Oliveri del Castillo 1968)
- ⋯ Low Gravity (Oliveri del Castillo 1968)
- Low Gravity (This Study)
- ⋯ Moderate Gravity (This Study)
- High Gravity (This Study)

Supporting Information for ”Gravimetric Constraints on the Hydrothermal System of the Campi Flegrei caldera”

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1. Gravity Data Reduction and Correction
2. Information on Caldera Fill Modelling
3. Additional Inversion Models
4. Figures S1 to S7

Introduction This document provides supporting information on the gravity data reduction and inverse modelling as part of the gravity study of the central sector of the Campi Flegrei caldera (Italy).

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1. Gravity Data Reduction and Correction

1.1. Instrumental Drift

Instrumental drift refers to changes in the instrument with time which affect the reproducibility of measurements. We corrected for drift using standard protocols of conducting single and multiple measurement loops along a series of benchmarks starting and ending at the same reference or control points.

While the drift in the data follows an average linear trend of ~ -0.675 mGal/day over the course of the survey (Fig. S1), the dataset for each day was corrected using the drift obtained by repeat readings at the base station as well as at control points. For the survey within the Solfatara crater, a control point located just outside the dense survey grid was measured multiple times per day and was tied to the base station twice a day.

1.2. Earth Tides

We used the CG-5 Autograv’s inbuilt Earth tide removal algorithm. Ocean loading effects are negligible in the Mediterranean (Ray, 1999) and have not been considered in our survey.

1.3. Normal Gravity

We use the International Gravity Formula (IGF) 1980 (Moritz, 1992) to calculate the normal gravity (g_n) for each gravity benchmark and the gravity reference using:

$$g_n = 978032677.14 \frac{1 + 0.0019319 \sin^2\phi}{(0.0066944 \sin^2\phi)^2}, \quad (1)$$

where ϕ is the latitude of the station.

Normal gravity at the main gravity base station was calculated at 980244.4306 mGal which all benchmarks were tied to accordingly to account for the effect of latitude on the data.

1.4. Free-Air Correction

The free-air correction (FAC) reduces the gravity data to the same equipotential elevation:

$$FAC(mGal) = -0.3086 h , \quad (2)$$

where h is the height difference in metres from the reference surface. This correction is added to gravity measurements made above the reference surface and subtracted below the reference surface (Telford et al., 1991). -0.3086 mGal/m is the theoretical free-air gradient. Berrino, Corrado, Luongo, and Toro (1984) measured a value of -0.290 mGal/m, and we calculate a value of -0.289 mGal/m from the published Bouguer anomaly data. However, we use the theoretical free-air gradient for the corrections noting that it is a value used more broadly for Bouguer data analysis. Using the theoretical vs the calculated or measured value adds an uncertainty of a few μ Gal to the Bouguer anomaly given uncertainties in benchmark locations. This error is negligible compared to the uncertainty due to terrain effects (see below). The average error of GPS-derived benchmark elevation of ± 0.024 m contributes an uncertainty of ± 7 μ Gal to the data based on the theoretical free-air gradient.

The resulting free-air anomaly map is shown in Fig. S2.

1.5. Bouguer Slab and Terrain Correction

The contributions to gravity from rocks located between the elevation of the gravity benchmarks and the elevation of the reference surface is accounted for by the Bouguer slab correction (BS) and the terrain correction (TC). The Bouguer slab correction (Bullard A) approximates this material as a slab of infinite horizontal extent, finite thickness and constant density (Robinson, 1988)

$$BS(mGal) = -0.0419 \rho h, \quad (3)$$

where ρ is the density of the slab (described below) and h is the height difference (m) from the reference datum.

The uncertainty of benchmark elevations and the tested density increments (100 kg m^{-3}) generates an uncertainty of $4 \mu\text{Gal}$ in the BS correction.

The approximation of the Bouguer slab only holds if the nearby topography is extremely subdued (Robinson, 1988). This is not the case at Campi Flegrei and one must consider an additional terrain correction, which we perform in MATLAB using high resolution onshore Digital Elevation Models (DEMs). Within and nearby the steep sided Solfatara crater, we used a 1 m DEM. We took care to place benchmarks in locations free from significant nearby (within the first Hammer zone) topographic changes to mitigate the effects of near-field topography. A 10 m DEM (Tarquini et al., 2012) was appropriate for the more distal region surrounding Solfatara. The RMS error between the GPS benchmark heights and heights at the same locations in the DEMs was $\sim 2 \text{ m}$ and $\sim 0.4 \text{ m}$ for the 10 m and the 1 m DEMs, respectively. We accounted for these offsets in the processing. We also incorporated bathymetric data (Ryan et al., 2009) to calculate terrain effects induced by

offshore topography. As the density of sea water (1024 kg m^{-3}) is much lower than that of rock this must be accounted for separately in the terrain correction.

We constructed separate DEMs for the onshore and offshore portions of the terrain correction. To test the uncertainty of the terrain correction by offsets between DEM heights and GPS derived heights we generated two normally distributed random topographies, ranging from 0 m in height to the RMS error between the GPS benchmark heights and 10 m and 1 m DEM heights (2 m and 0.4 m, respectively). We calculated the terrain correction for two benchmarks located 1 km apart (for the 10 m DEM) and 50 m apart (for the 1 m DEM) to establish the difference between the two terrain corrections. This was repeated 100 times and the 1σ error of each distribution was found to be $130 \mu\text{Gal}$ and $13 \mu\text{Gal}$ for the 10 m and 1 m DEM, respectively.

To calculate the cumulative terrain effect we followed the approach of Hammer (1939), but calculated the terrain correction at each DEM data point rather than for each Hammer chart compartment. The distance from each benchmark to every DEM data point is calculated and used to weigh the influence of the height difference between each DEM data point and the benchmark;

$$TC_1 = \left(\frac{1}{r} - \frac{1}{r^2 + \Delta z^2} \right) \Delta x^2, \quad (4)$$

where r is the radial distance from the benchmark to each DEM data point in metres, Δz is the height difference between the benchmark and the DEM data point and Δx is the DEM spacing. The total terrain correction for each benchmark is then calculated;

$$TC_2 = \rho G \sum TC_1, \quad (5)$$

where G is the universal gravitational constant. The terrain correction was calculated for the onshore portion of the survey using the onshore DEM and an appropriate density (described below). The offshore portion required a density equal to that of the onshore density minus the density of seawater (1024 kg m^{-3})

The density for both the Bouguer correction and the terrain correction should be close to the average sub-surface density (Robinson, 1988). If density data is not available it is possible to estimate it using the Bouguer anomaly and elevation data. Traditionally this is done using the Nettleton method (Nettleton, 1976). This method involves plotting profiles of Bouguer gravity (for a range of different terrain densities) against topography and calculating the correlations between the Bouguer anomalies for a range of different terrain densities and topography. The terrain density giving the least correlation between Bouguer anomaly and topography is selected as the best estimate of terrain density. (Fig. S3) However, profiles over structurally controlled features may not be appropriate as density might change with elevation (Nettleton, 1976). While the profiles displayed in Fig. S3 show some correlation with topography, the least correlated profile is between 1800 and 1900 kg m^{-3} . The interpolation of the Bouguer gravity data has a smoothing effect particularly in areas of sparse data and makes correlating the elevation and Bouguer gravity profiles qualitative rather than quantitative.

To explore a quantitative approach we correlate the Bouguer anomaly for different terrain densities with elevation at each benchmark (Eshaghzadeh et al., 2015). This correlation is plotted against the tested densities and a least-squares straight line is fitted to the resultant data. Once this line is subtracted from the data, the ‘correlation difference’ is plotted against the density and returns an unique optimal terrain density. The resultant value of 1900 kg m^{-3} (Fig. S4) is in agreement with the value derived by the Nettleton method and compares with densities of material recovered from boreholes at Campi Flegrei in the range from $\sim 1600 \text{ kg m}^{-3}$ (tuff) to $\sim 2800 \text{ kg m}^{-3}$ (thermometamorphic rocks) (Barberi et al., 1991; Piochi et al., 2014). The average density of the first 500 m of all borehole data is approximately 1800 kg m^{-3} and the average density of all the borehole data is 2300 kg m^{-3} (Barberi et al., 1991). We hence regard the value of 1900 kg m^{-3} for both the terrain and Bouguer slab densities as a mathematically robust and geologically plausible estimate.

2. Caldera Fill Modelling

The effect of the caldera fill on the regional Bouguer anomaly is shown in Fig. S5.

3. Inversion Models for ± 600 and $\pm 300 \text{ kg m}^{-3}$ A Priori Density Contrasts

Results are shown in Figs. S6 and S7.

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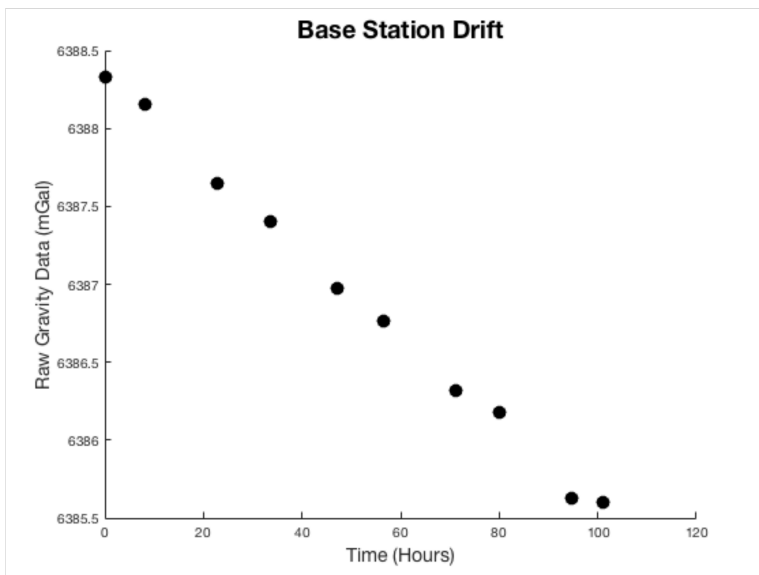


Figure S1. Instrument drift of ~ -0.675 mGal/day established by gravity readings at the base station over the course of the survey.

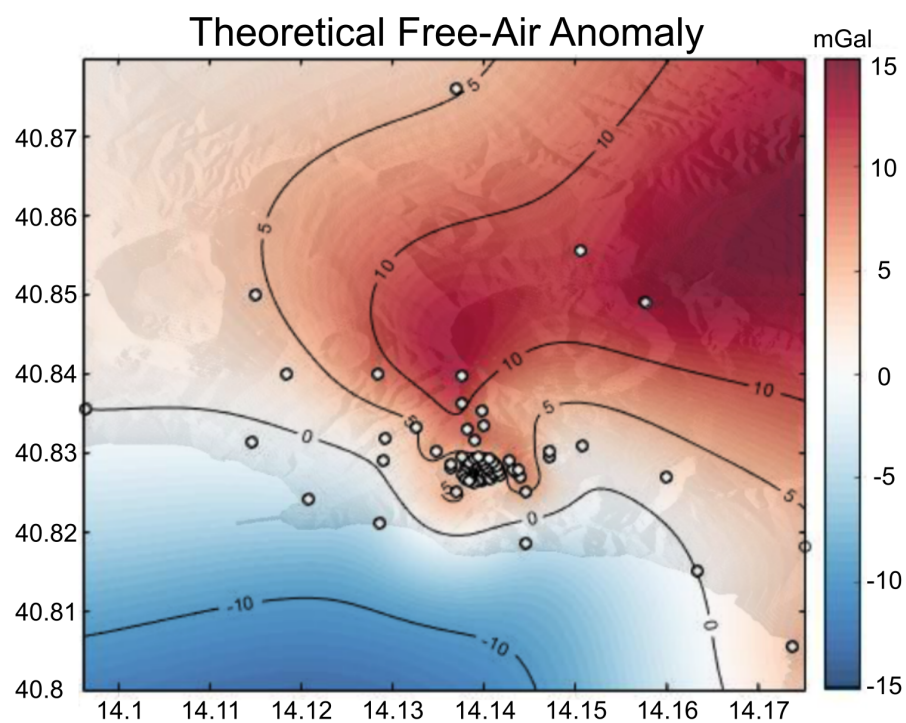


Figure S2. Free-air anomaly map superimposed on the 10 m DEM with benchmarks in white and contours in mGal.

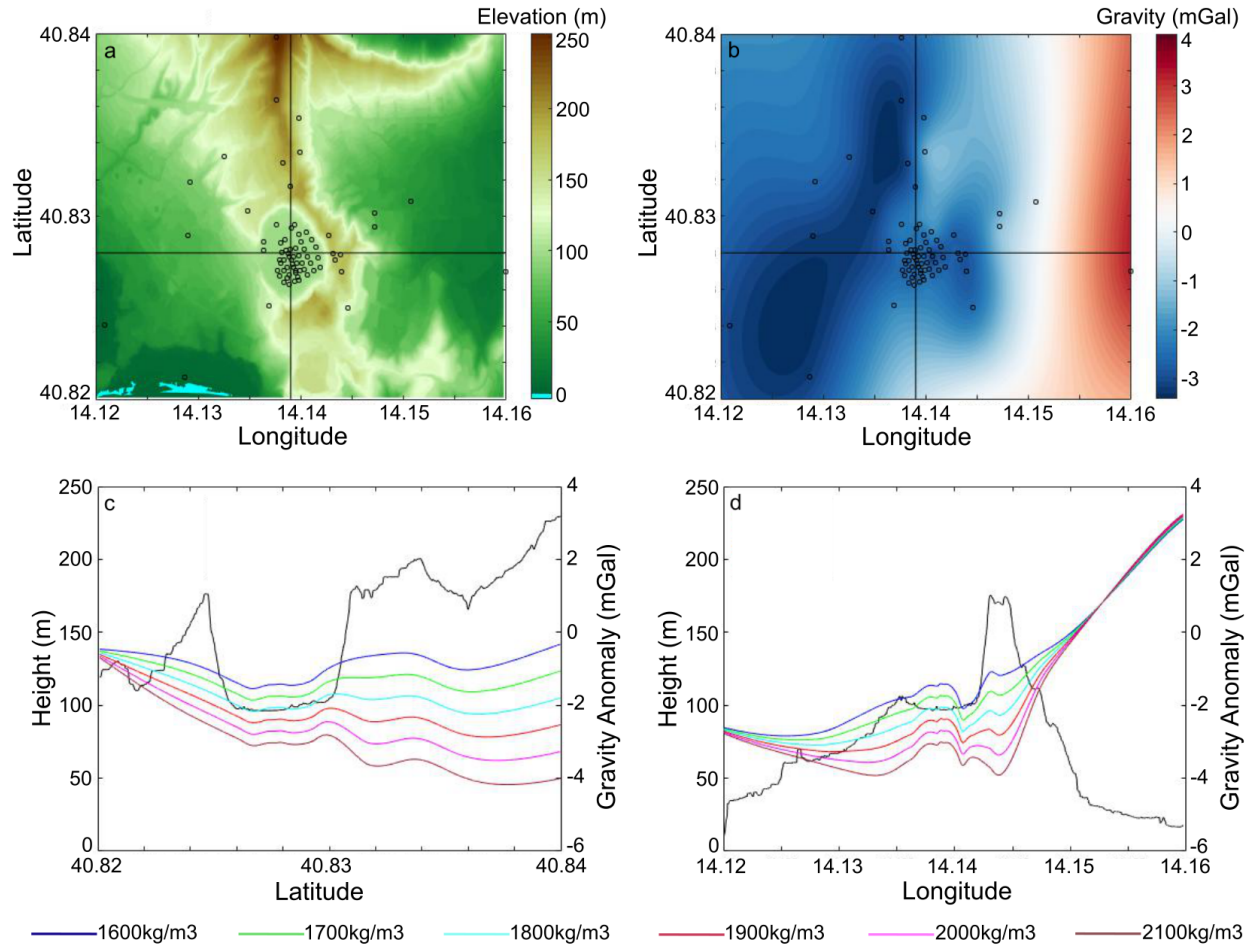


Figure S3. Derivation of the terrain density using Nettleton's method. a) The 1m DEM with EW and NS profiles at 40.828°N and 14.139°E, respectively shown as black lines. b) Bouguer anomaly map for a 1900 kg m⁻³ Bouguer slab and terrain density with NS and EW profiles shown as black lines. c) Comparison of Bouguer anomaly data (for slab and terrain densities as per the legend) with topography (black line) along the N-S profile. d) same as c) for the EW profile.

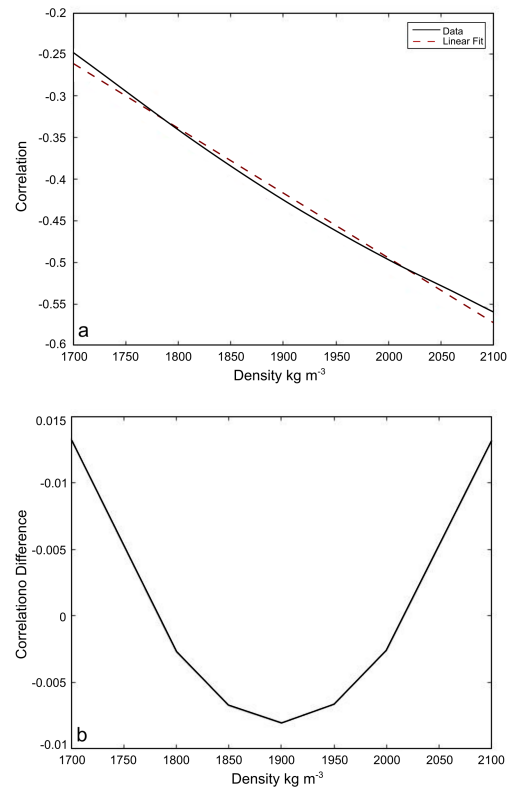


Figure S4. a) Correlation between topography and terrain density with the best linear fit (dashed red line). b) The residuals between the correlation curve and its linear trend plotted against density yields an unique minimum value of the terrain density which minimises the topographic effects on the Bouguer anomaly (Eshaghzadeh et al., 2015).

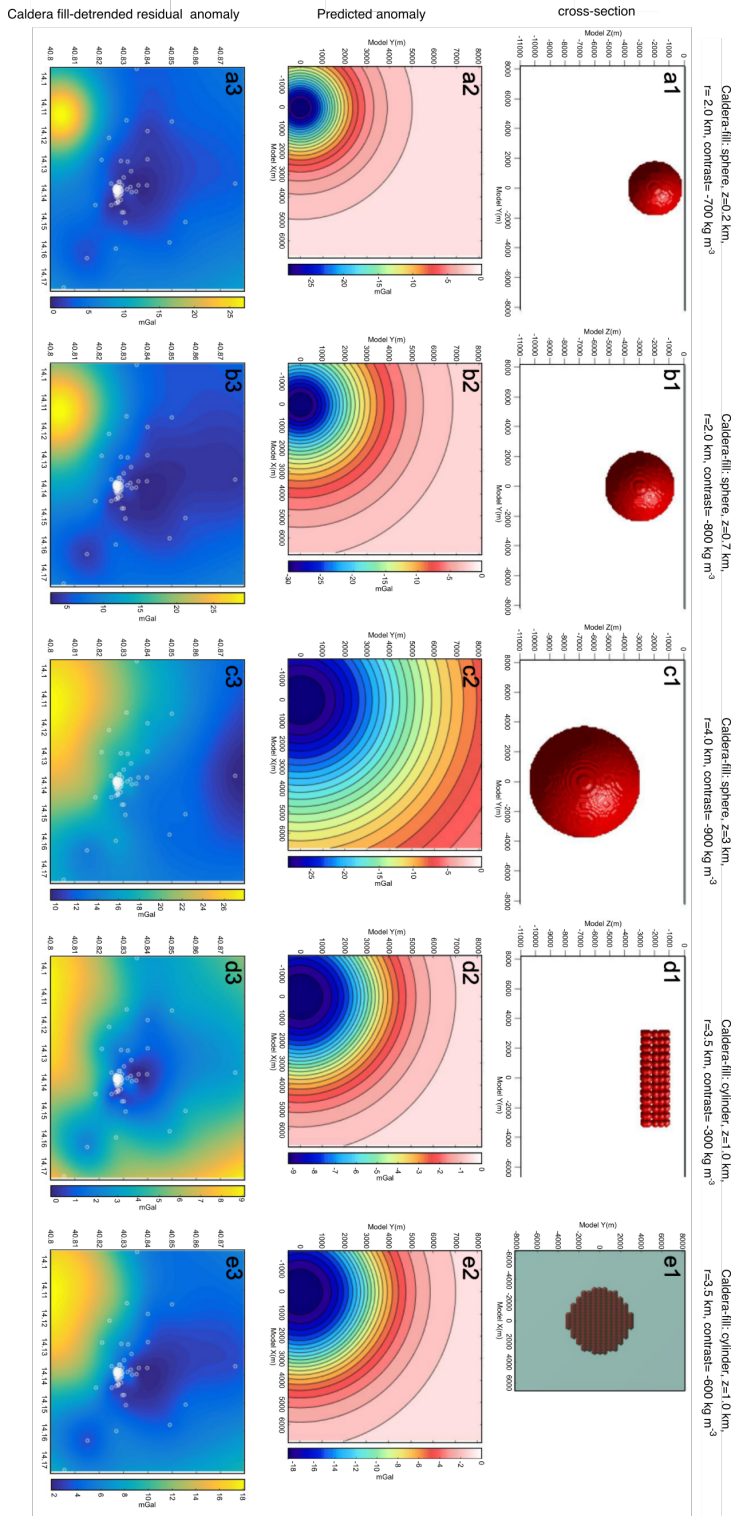


Figure S5. The effect of the caldera fill on the regional Bouguer anomaly for spherical (panel letters a-c) and cylindrical (panel letters c and d) representations of the fill. Panels numbers (1) show model cross sections for given shape, depth to top of body (z) and radius (r) and negative density contrast, except for e1 which illustrates the model shown in d1 in plan view. Panels numbers (2) give the predicted gravity anomaly at the surface. Panel numbers (3) show the resultant local Bouguer gravity anomaly once the caldera fill anomaly has been subtracted from the regional Bouguer data set to obtain the caldera-fill detrended residual anomaly (CRA) (see main text for details). The single sphere models (panels a-c) were rejected for lack of fitting with results from gravity modeling presented by (Capuano et al., 2013), while the cylinder models (panels d-e) provided a much better match with d3 providing the best. The resultant CRA shown in d3 is identical to the CRA presented in Fig. 3c of the main paper and is used for the inversions.

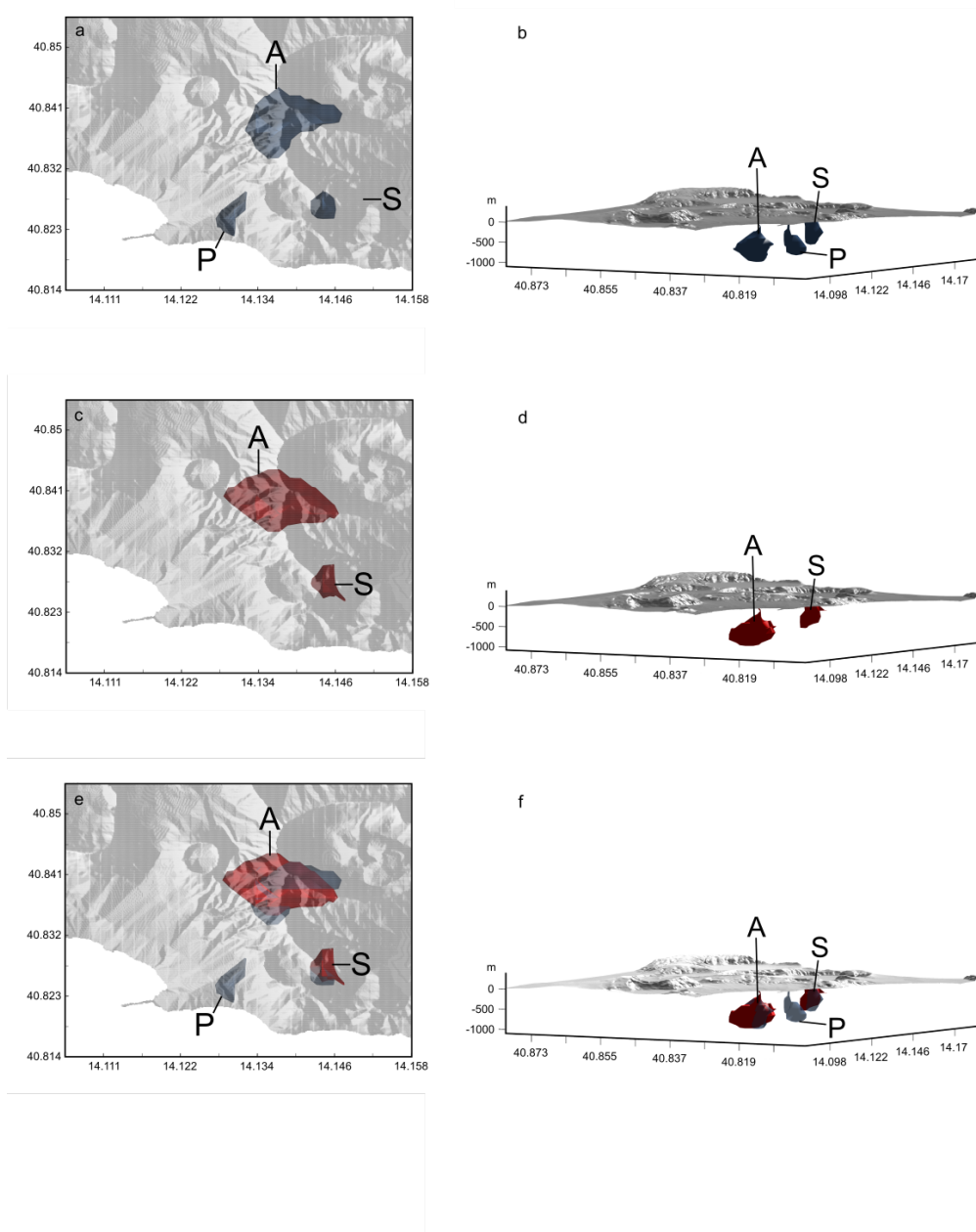


Figure S6. -550 kg m^{-3} density contrast isosurfaces from inversions with an a priori density contrast of $\pm 600 \text{ kg m}^{-3}$. a) LRA inversion in plan view. b) LRA inversion with a view facing 59° NE. c) CRA inversion in plan view. d) CRA inversion with view facing 59° NE. e) LRA and CRA inversions in plan view. f) LRA and CRA inversions with view facing 59° NE. The main anomalous bodies are located beneath Pozzuoli, Solfatara/Pisciarelli volcano and Astroni volcano and are labelled P, S and A, respectively.

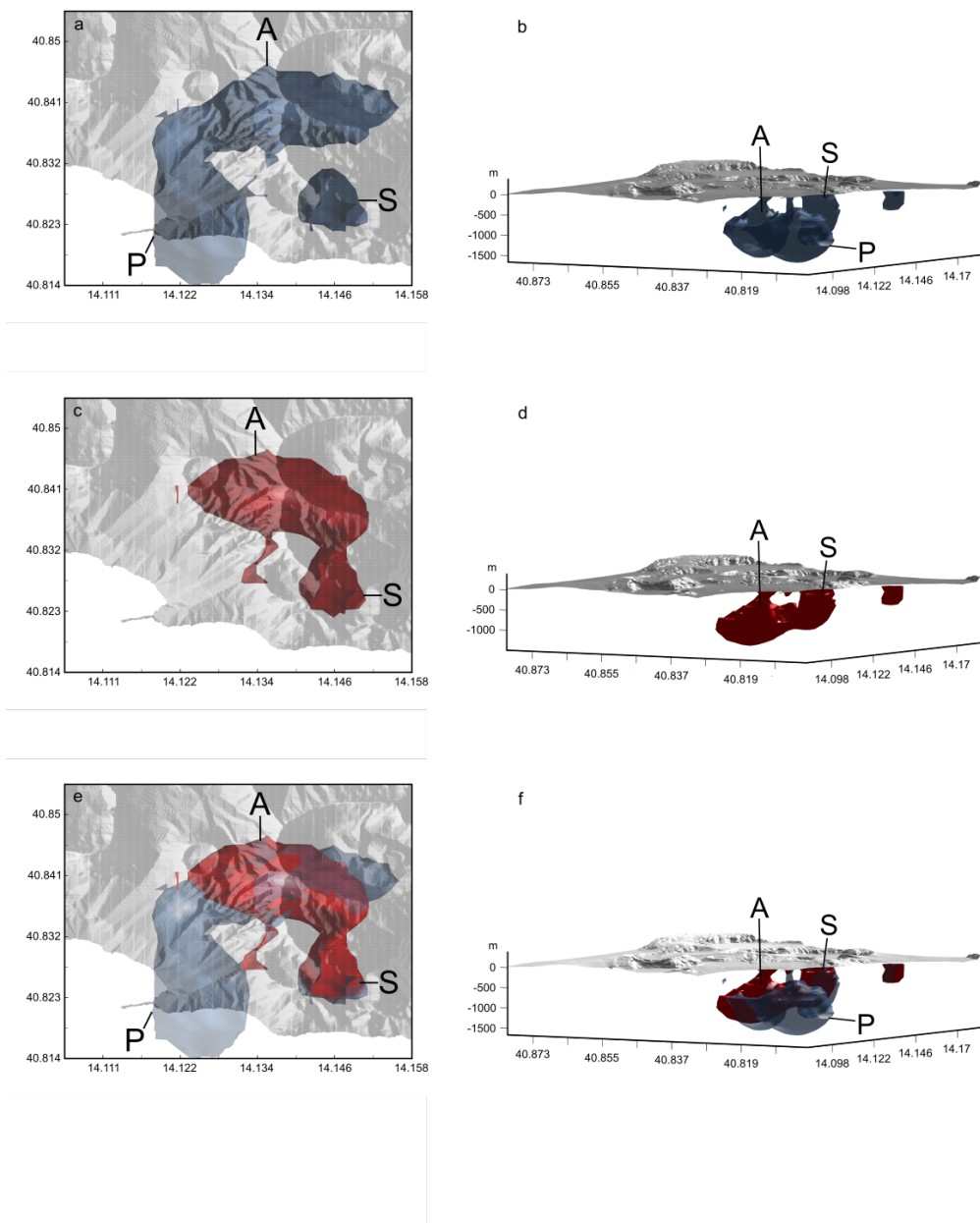


Figure S7. Same as Figure S6 but showing the -250 kg m^{-3} density contrast isosurfaces from inversions with an a priori density contrast of $\pm 300 \text{ kg m}^{-3}$.