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Cyclic anoxia and organic rich carbonate sediments within a drowned carbonate platform linked to Antarctic ice volume changes: Late Oligocene-early Miocene Maldives

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- Cyclic Anoxia and Organic Rich Carbonate Sediments within a Drowned Carbonate
 Platform linked to Antarctic ice volume changes: Late Oligocene-Early Miocene Maldives
- 3

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46 **ABSTRACT**

This paper reports the newly discovered occurrence of thick sequences (~ 100 m) of 47 Late Oligocene-Early Miocene interbedded organic carbon-rich sediments (sapropels) 48 and pelagic (organic carbon poor) carbonates (~24.5 to 21.5 Ma) at Sites U1466 and 49 U1468 drilled in the Maldives archipelago during the Integrated Ocean Discovery 50 51 Expedition (IODP) 359. This occurrence is unusual in that it is located > 1000 m above the surrounding ocean floor within an inter-atoll basin and not linked to any known 52 global oceanic events. Total organic carbon content reached as high as 35 % in the 53 darker layers, while the interbedded carbonates have concentrations of less than 0.1%. 54 Trace elements characteristic of anoxic waters, such as Mo, V, Cr, U, and Pb, correlate 55 positively with concentrations of organic carbon. Nitrogen isotopic data show no 56 evidence that the intervals of high total organic carbon are related to enhanced 57 productivity driven by upwelling. Instead, high organic carbon is associated with 58 intervals of anoxia. We propose that sea level fluctuations linked to Antarctic ice volume 59 changes restricted exchange with the open ocean causing bottom waters of the inter-60 atoll basin to become anoxic periodically. The architecture of the platform at the end of 61 62 the Oligocene combined with the global sea level highstand set the stage for orbitallydriven sea level changes producing cyclic deposition of sapropels. The proposed 63 mechanism may serve as an analogue for other occurrences of organic carbon-rich 64 65 sediments within carbonate platform settings.

1. **INTRODUCTION**

Organic-rich deposits are abundant in the geological record (Algeo, 2004; Arthur, 1979; 68 69 Jenkyns, 2010; Parrish and Curtis, 1982; Sageman et al., 2003; Schlanger and Jenkyns, 70 1976). Whilst many ancient examples are thought to be related to global anoxic events, there are few examples associated with the development of carbonate platforms (Gardner 71 72 et al., 2013; Koster et al., 1988) and isolated basins (Emeis and Weissert, 2009; Lyons, 1991; Peterson et al., 1991), where restricted ventilation or enhanced productivity caused 73 the preservation of TOC. Here, we report on the newly discovered occurrence of a 74 sequence of alternating organic-rich and -poor sediments from Integrated Ocean 75 Discovery Project (IODP) Sites U1466 and U1468 in the Maldives Inner Sea (Indian 76 Ocean), examining the possible conditions that led to the formation of this unusual deposit 77 in an inner-atoll basin through geochemical and associated with orbitally-driven Antarctic 78 79 ice volume changes, and suggest that the inter-atoll environment of the Maldives may 80 serve as an analogue system for other known organic-rich deposits with previously unclear origins. 81

82

83 **1.1 Background**

This study deals with samples from two cores from the Maldives recovered during Expedition 359 of the Integrated Ocean Discovery Project (IODP) (Betzler et al., 2016a). These sediments are comprised of a ~3 km thick shallow water and hemipelagic carbonate succession (Aubert and Droxler, 1992) that formed on a lower Paleogene (60– 50 Ma) volcanic basement (Duncan and Hargraves, 1990). Although the primary aim of drilling during IODP Expedition 359 was to examine timing of the commencement of the

South Asian Monsoon (Betzler et al., 2018; Betzler et al., 2016b), drilling also revealed,
a ~100 m section of interbedded hemi-pelagic sediments rich in organic carbon at Sites,
U1466 and U1468 (Fig.1), similar to those found in ocean-wide anoxic events.

93

94 2. **METHODS**

Samples were taken from every dark layer and intervening light colored layers between ~ 700 and 800 mbsf at Site U1466 and between ~ 720 and 805 mbsf at U1468. Highresolution sampling was conducted in two core sections, U1466B-56R-1W and U1468A-105X-1W, where 2 cm³ samples were taken at \sim 2 cm intervals (Fig. 2).

99

100 **2.1 Organic** δ^{13} **C and** δ^{15} **N values and Acid-insoluble material**

101 Co-occurring sedimentary organic material was isolated via dissolution in 5 % HCl acid following the method described in Oehlert et al. (2012). The organic carbon residual was 102 combusted in a Costech ECS 4010 (Costech Analytical Technologies, Inc.) Mater was 103 removed and oxygenated species of nitrogen reduced to N₂ using a Cu furnace held at 104 600°C in the standard manner. For C and N isotopic measurements, the CO₂ and N₂ 105 gases produced were transferred to a continuous flow stable isotope ratio mass 106 spectrometer (Thermo Delta V Advantage). The δ^{13} C and δ^{15} N values of the samples are 107 reported relative to the V-PDB (Vienna Pee Dee Belemnite) scale and atmospheric 108 nitrogen respectively. The V-PDB scale for organic carbon is defined by the δ^{13} C value of 109 graphite (USGS24) = -16.05‰ versus V-PDB (Coplen et al., 2006). Within each run 110 (typically 50 samples), a glycine and secondary standard were analyzed every ten 111 samples. The reproducibility of δ^{13} C and δ^{15} N values of the standards was ±0.1 and ±0.3 112

113 % respectively.

114

115 **2.2 Insoluble Residue and Total Organic Carbon**

Weights and percentages of insoluble residue and total organic matter (TOC) were analyzed and calculated following the methods of Oehlert et al. (2012). The standard deviation of these analyses is 0.4 wt% based upon repeated analyses of glycine (n=54).

120 **2.3 Inductively Coupled Plasma Mass Spectrometry (ICP-MS)**

121 Samples from Core Section U1466-56X-1 were analyzed for a range of major and trace elements (Li, Na, Mg, Al, K, Ca, Ti, V, Cr, Mn, Fe, Sr, No, Cd, Pb, Th, and U) using a 122 quadrupole ICP-MS (Thermo iCAP-Q) at Princeton University. The method thoroughly 123 dissolved the carbonate phase and partially dissolved other mineral phases and organic 124 matter so the values analyzed represent a maximum value for the carbonate phase and 125 a minimum of the bulk sample. Approximately 10 to 25 mg samples of bulk sediment 126 were weighed out directly into clean, acid-washed Teflon beakers. The samples were 127 dissolved by adding 5 mL of concentrated (~16 N) distilled nitric acid, capping the 128 129 beakers, and heating them at ca. 75 °C for over 12 hours. The beakers were then opened and dried down, and an additional 2 mL of concentrated nitric acid was added and dried 130 down. The samples were then redissolved in 5 mL of 2% nitric acid. All mineral 131 132 components appeared to be dissolved, although some brown detritus resembling remnant organic matter sometimes remained. After centrifugation, the clear supernatant was 133 diluted up to 200x to yield solutions of approximately 10 ppm Ca. Trace element analyses 134 were performed with matrix-matched standards at approximately 10 ppm Ca and with 135

varying trace element concentrations to bracket those in the samples. Internal errors arebetter than 3% (RSD).

138 2.4 X-Ray Fluorescence

The XRF scans were measured at the Ocean Drilling and Sustainable Earth Science 139 (ODASES) XRF scanning facility at the IODP Gulf Coast Repository (GCR) in College 140 Station, Texas (USA) using a third generation AvaaTech XRF scanner configured to 141 analyze split sediment core halves for elements between Mg and U in the periodic table 142 (Lyle and Backman 2013). The data were acquired with a Canberra X-PIPS Silicon Drift 143 Detector (SDD) with 150 eV X-ray resolution at 5.9 keV and a Canberra Digital Spectrum 144 Analyzer model DAS 1000. The X-Ray source was an Oxford Instruments 100 W Neptune 145 X-ray tube with a rhodium (Rh) target. Raw spectral data were processed using the 146 Canberra WINAXIL software package to produce elemental intensity data. The dual slit 147 system was set to provide down-core spatial resolution of 10 mm and cross-core spatial 148 resolution of 12 mm. The system was set to perform two consecutive runs of the same 149 section, the first one at 9 kV, 0.25 m A and 6 s, and the second one at 30kV, 1.25 mA 150 and 6 s. The first scan provides data for the elements Mg, Al, Ba, Ca, Cl, Cr, Fe, K, Mn, 151 P, S, SI, Rh, and Ti, and the second for Bi, Cd, Cu, Br, Fe, Ga Nb, Ni, Mo, Pb, Rb, Sr, Y, 152 Zn, and Zr. Each core section was removed from refrigeration at least 2 hours before 153 scanning and scraped to clean and smooth the core surface. The split core surface was 154 covered with 4 µm thick Ultralene plastic film to prevent contamination of the X-ray 155 detector. Measurements were taken at 3 cm intervals whenever possible. Because 156 157 measurements cannot be performed if the sediment presents cracks or uneven surface some measurements were skipped or shifted to the nearest suitable area. In addition, 158

Core Sections U1466B-56R-1W and 2W were analyzed at 1 cm resolution. In order to 159 evaluate the reliability of the analysis of certain elements (Fe, Al, Sr, Si K, Ca, Br, Mo, 160 and Ba) a companion study (Kunkelova et al., 2018) analyzed samples of pressed pellets 161 using a fully quantitative conventional XRF method. That study measured 67 samples 162 from Cores 359-U1467C-4H to -6H on a conventional XRF instrument in Edinburgh with 163 164 the same pellets measured on the AvaaTech XRF core scanner College Station in Texas. For these, 1 g aliquots of core material were ground in an agate mortar and pestle and 165 then pressed into 1 cm diameter pellets using a pellet press operating at a pressure of 2 166 167 t/cm2. Pellets were analyzed on a Philips PW2404 XRF spectrometer with a Rh-anode X-ray tube. Corrections for matrix effects on the intensities for major element lines were 168 made using theoretical alpha coefficients calculated using Philips software (Reynolds, 169 170 1963). Intensities of the longer-wavelength trace element lines were corrected for matrix effects using alpha coefficients based on major element concentrations. For other trace 171 elements, matrix corrections were applied using the Rh K alpha Compton scatter line as 172 an internal standard. Line overlap corrections were applied using synthetic standards. 173 The spectrometer was calibrated with 15 USGS and CRPG standards using the values 174 175 given in (Govindaraju, 1994).

176

177 2.5 Inorganic Stable Carbon and Oxygen Isotopes

Carbonate materials were reacted with phosphoric acid using a common acid bath at 90°C and the CO₂ released analyzed using a dual-inlet mass spectrometer (Finnigan-MAT 251 at the University of Miami. The CO₂ gas was calibrated using NBS-19 (National Bureau of Standards) and reported relative to V-PDB using the conventional notation.

182 Replicate analyses yielded a precision < 0.1‰ for both δ^{13} C and δ^{18} O values.

183

184 **2.6 Fatty acids**

In intervals with high organic content, sediment samples from Hole U1466B were 185 186 obtained to analyze fatty acids. Fatty acids were analyzed using an on-line TMAH (tetramethylammonium hydroxide) thermochemolysis. Finely powdered and dried 187 188 sediment sample (ca. 14 mg) was placed in a pyrofoil (foil for pyrolysis) and TMAH 189 reagent (97%, Sigma-Aldrich Co., 25 wt.% methanol; 40 µL) was added. After methanol 190 was evaporated to dryness, the pyrofoil was wrapped and set into a Curie point pyrolyzer (JHP-3: Japan Analytical Industry Co.) and heated at 315°C for 20 s. The resulting TMAH 191 products were introduced into a GC column (DB-5ms, 30m x 0.32mm i.d.; 0.25µm film 192 193 thickness) and analyzed by MS (5975C; Agilent Technologies Co.).

194

195 **2.7 Downhole Logging**

Total and spectral natural gamma radiation (NGR) was measured at 30-cm resolution with a 15-cm sampling interval using standard shipboard methods employed by IODP as described by Betzler et al. (2017).

199

200 **2.8 Statistical Methods**

Factor analysis was performed using Statistica 8.0 and spectral analysis using MATLAB[™] (https://www.mathworks .com). Ages were calculated using the shipboard biostratigraphy (Betzler et al., 2017) and the data were interpolated using a linear method to an interval of 10,000 years (Davis, 1973). Spectral analyses were performed using

Tukey weights and smoothing window of five points. Factor analysis was performed on the ICP-MS data and the results normalized using a Varimax method.

207

208 2.9 Scanning Electron Microscopy

Selected samples were examined from Core U1446B-56R in both the dark and light layers. Small rock chips were mounted on aluminum stubs using epoxy resin, sputter coated with gold-palladium and examined using a Zeiss Ultra Plus Field Emission SEM at the British Natural History Museum.

213

214

215 3. **RESULTS**

216 **3.1 Age Model**

The ages of the sediments were determined by micropaleontologists on the drilling 217 expedition. The integrated nannofossil and planktonic foraminiferal stratigraphy provides 218 a robust biostratigraphic framework for this carbonate sequence representing drowning 219 of an ancient shallow water carbonate platform. Several reliable biostratigraphic events 220 were placed with a high degree of certainty. These include the first occurrence of 221 Paragloborotalia kugleri, which indicates the Oligocene/Miocene boundary, and the last 222 occurrence of the late Oligocene species Sphenolithus ciperoensis. The description of 223 the sedimentology, age, and various chemical and physical properties can be found in 224 the initial scientific reports (Betzler et al., 2016b) and additional publications (Betzler et 225 226 al., 2016a; Betzler et al., 2018).

227

3.2 Sedimentology

The intervals containing the dark-light cycles (715.34-803.61 mbsf in Site U1466 and 230 728.6-817.55 mbsf in Site U1468) overlie an Oligocene carbonate platform and a 231 succession of shallow-water deposits (Betzler et al., 2016b) (Fig. 1B). These sediments 232 233 consist of wackestone, locally and gradually changing into mudstone or packstone and display an alternation of laminated and poorly bioturbated dark intervals with highly to 234 completely bioturbated light intervals (Fig. 3). Sediments in the dark intervals have 235 236 abundant planktonic foraminifers, fish debris and nannofossils suggesting open marine conditions. The dark layers are finely laminated and locally show scattered discrete small-237 sized burrows of Thalassinoides, Phycosiphon, Palaeophycus, and Planolites. The dark 238 intervals range in thickness from 1 to 25 cm and are intercalated with light wackestone 239 intervals of 5 to 300 cm thick. The sediments in the light intervals are highly bioturbated 240 and are comprised of abundant planktonic foraminifers and nannofossils (Fig. 4). Fi 241 debris are rare to barren. 242

243

244 **3.3 Organic Content**

The darkness of the sediments is related to the TOC of the sediment, which varies between <0.1% within the light layers to as high as 35 % in one of the dark layers. The mean TOC ($\pm 1\sigma$) for the dark layers is 8.78 \pm 5.6% and 5.77 \pm 6.4 % for Sites U1466 and U1468, respectively, compared to the light layers which have TOC of 0.73 \pm 0.78% and 0.15 \pm 0.22%. In both Sites 1466 and 1468, the intervals with the highest TOC, are found towards the base on the dark and light interval (Fig. 5 & 6).

252 3.3.1 Carbon and Nitrogen Isotopic Composition

The mean δ^{13} C values of the TOC (δ^{13} C_{TOC}) are -20.9 ‰ at Site U1466 and -22.7 ‰ at 253 254 Site U1468; there was no statistically significant (SS) variation between the dark and light layers. However, there is a gradual increase in the δ^{13} C values from the base of both 255 Sites U1466 and U1468 to approximately the top of the interval in which the dark and light 256 layers are found. In Site U1466, the increase in δ^{13} C values is from -23 to -17 ‰ while at 257 Site U1468 the δ^{13} C value increases from -24 to -20 ‰. Within the zone of dark and light 258 alternations, the δ^{15} N values of the TOC range from -1 to +2 ‰ with the organic-rich zones 259 generally expressing more negative $\delta^{15}N$ values (Fig. 8). 260

261

262 3.3.2 Fatty Acids

Most of the fatty acids are in the range from C_{12} to C_{32} , with dominance of C_{14} -- C_{18} fatty acids (supplemental material, S1). There is a statistically significant increase in C_{24} -- C_{32} fatty acids below ~770 mbsf and the ratio of C_{24} -- C_{32} fatty acids to C_{14} -- C_{18} fatty acids (long/short ratio) increases towards the base of the dark and light interval.

267

268 3.3.3 CN Ratio

The C/N ratio of the organic material varied between 10 and 20 with no statistically significant differences between the dark and light layers. Site U1468 has higher C/N ratios than U1466 (p<0.05), and the CN ratio increases towards the bottom of the section.

273 **3.4 Carbonate**

274 3.4.1. Oxygen and Carbon Isotopic Composition

Within Core U1466B-56R there is an alternation of ~1‰, in the δ^{18} O values, the increase 275 of which leads the increase in TOC (Fig. 7). The amplitude of these cycles is similar to 276 the early Miocene benthic cycles of δ^{18} O values in foraminifera noted during the same 277 time on Ceara Rise (Zachos et al. 2001). Similar associations were noted in Core 278 U1468A-105X. However, overall the δ^{18} O values of the dark layers (-1.86 ‰ in Site 279 U1466 and -2.29‰ in Site U1468) are slightly more negative than the light layers (-1.60‰ 280 in Site U1466 and -1.96‰ in Site U1468). The δ^{13} C values of the carbonate components 281 are more negative within the dark layers compared to the light layers (+0.84‰ vs. +1.28 282 ‰ in Site U1468; +0.97‰ vs +1.24‰ in Site U1466 (p<0.01)) (Fig. 7). 283

284

285

286 3.4.2. Trace Elemental Analysis

287 X-ray Fluorescence Scanning

The following elements (Ba, Br, Cd, Fe, Cu, Mo, Pb, S, Y, and Zn) showed a strong positive correlation between each other and with the concentration of TOC in the cores (Fig. 8 and 9). Generally, there is a sharp rise in concentration of trace elements at the bottom of each dark interval mirroring changes in the 'darkness' of these intervals.

292 Wet Chemical Analyses

Samples from Section U1466B-56R-1 were analyzed for a range of elements using 293 Inductively Coupled Plasma- Mass Spectrometry (ICP-MS). Whilst some of the elements 294 analyzed were the same as those measured using XRF (Ca, Fe, Mo, Pb, and Sr), several 295 others (Cr, Th, U, and V) were only analyzed using ICP-MS and provide important 296 additional information on the redox state and origin of the sediments. The concentration 297 of certain elements (Ca, Fe, Mo, Pb, Al, and Sr) measured using ICP-MS are generally 298 correlated with intensities measured using XRF, albeit with a lag of ~ 2-4 cm (See 299 supplemental material, S2 for cross plots between XRF and ICP-MS data and 300 301 discussion). For other elements (Cr, Ba, and Mg) no correlation was observed between the two methods (See supplemental material, Analogous to the XRF data, the 302 concentration of a number of elements (AI, Ba, Cd, Cr, Fe, Li, Mo, Pb, Th, V, and U) were 303 positively correlated with TOC in this core (Fig. 7). Principal component analysis shows a 304 preferential clustering of Al, Cd, Cr, Fe, K, Li, Mo, Pb, Th, U, and V accounting for 54% 305 of the variance (supplemental material, Fig. S3-11). 306

307 Comparison to Quantitative XRF Analyses

Correlation coefficients range from about < 0.1 to ~0.9 with K and Si having the highest values of about 0.9. Elements such as Fe, SI, and AI showed a R^2 values > 0.8. A full comparison of the results from both methods are included in the supplemental material (S1).

At all sites, spectral natural gamma ray (NGR) logs show that the NGR signal is almost 314 entirely the result of variations in the concentration of U, with K and Th contributing only 315 a minor proportion of the total signal (Betzler et al., 2017). The NGR signal routinely 316 shows a positive correlation with the concentration of TOC (R²=0.40, n=88 for Site U1468) 317 and R²=0.36, n=71 for Site U1466) and can be used to correlate core material with 318 319 downhole logging data, which include in situ total and spectral NGR. Based on changes in the downhole NGR signal, between 30 and 35 dark layers were identified between 24.5 320 and 21.5 Ma, only some of which were recovered during drilling. A spectral analysis of 321 322 these data show peaks at \sim 100 and 50 K years (Fig. 10a).

323 *3.4.4. Resistivity*

The resistivity signal obtained from downhole logging was strongly correlated with the Gamma ray signal. As this tool was located lower on the logging string than the Gamma tool a further 22 m of section could be logged in Hole U1468. A spectral analysis of these data (Fig. 10b) show peaks around 120, 80, and 40K years.

328

329 4. **DISCUSSION**

Previous work has shown that a rimmed platform with a protected basin developed in the Maldives during the Early to Late Oligocene transition when shallow-water carbonate production became restricted to narrow bands at the ocean-ward edges, forming a saucer shape basin perhaps 50-100 m deep (Betzler et al., 2018). The platform drowning at the Oligocene-Miocene transition coincided with a major sea level rise during the early Miocene with an amplitude of approximately 50 m (Miller et al., 2011) and is perhaps

synchronous with platform drowning events recognized in other locations (Mutti et al.,
1997). The drowning in the Maldives is coincident with the development of the sapropels
described in this paper.

339 4.1. Nature of the Dark and Light Layers

While many factors can contribute to the origin of color differences including variations in 340 trace elements such as Fe, it is likely that the major cause of variations in these cores 341 arise from variations in the amount of organic material within the darker layers, identified 342 343 simply on a subjective inspection of color, containing an average of between 5.8 and 9.8 wt % compared to the lighter colored intervals (0.16 to 0.73 wt %). Whilst the presence 344 of abundant oceanic biota suggests that oceanic conditions were prevalent during 345 deposition of both dark and light layers, the darker ones have the characteristics of 346 sediments deposited in anoxic basins including a high TOC (2-39%) and an abundance 347 of redox-sensitive trace elements (Brumsack, 2006; Tribovillard et al., 2006). In contrast 348 to the light layers, the darker intervals have smaller burrows and low bioturbation indices, 349 also clear indicators of reduced oxygenation or enhanced productivity (Reolid and Betzler, 350 2018). 351

The occurrence of such layers is normally thought to be result from either enhanced productivity or through the enhanced preservation of organic material. Reduced organic concentration as a result of sediment dilution which has been proposed in some instances (Bohacs, 2005) to account for variations in the amount of organic material. However, this is not considered to be a realistic option in this case. In reality, most such occurrences arise from a combination of both productivity and preservation, although additional geochemical patterns can indicate the relative importance of these two effects.

360 *4.2. Productivity*

An enhanced productivity origin of the high TOC within the darker layers is supported by 361 slightly higher concentrations of nutrient-related elements such as Cd and to a lesser 362 extent Ba, Cu, and Zn. While the interpretation of variations in these elements is not 363 straight forward (Tribovillard et al., 2006), in some instances higher concentrations of 364 these elements are related to upwelling and hence higher inputs of nutrients such as 365 366 nitrate and phosphate (Boyle, 1981; Lea and Boyle, 1991). Such upwelling might occur along the margins of the platform, influencing productivity within the water over the 367 platform. However, upwelled waters should have been influenced by denitrification and 368 hence the organic components should have had relatively positive $\delta^{15}N$ values (Altabet 369 and Deuser, 1985). This is contrary to is observed with the organic matter in the dark 370 layers having δ^{15} N values close to 0% (Fig. 7). This suggests that nitrogen fixation, rather 371 372 than upwelled nitrate, was the major source of nitrogen for the organic production. Another potential source of nutrients in the darker layers, is atmospheric dust (Figs. 7 -373 374 9). These layers are elevated in AI, K, Li, and Fe. Considering that the Maldives are not significantly influence by riverine source it is likely that atmospheric deposition is the 375 source of these elements. This is supported by work which has shown an increased 376 supply of dust in glacial periods during the past 2 myrs, variation which is clearly related 377 to Milankovitch periodicity (Kunkelova et al., 2018). Nitrogen in dust typically has a δ^{15} N 378 379 value close to zero (Knapp et al., 2010) and therefore the values are consistent with the data measured in this study. Regardless of the source of nutrients, enhanced 380 productivity should have resulted in more positive δ^{13} C values within the carbonate of the 381

dark layers. In fact, the carbonate within the darker layers possessed more negative δ^{13} C values compared to the light layers, suggesting the darker layers cannot be explained by higher productivity. There were no differences in the δ^{13} C_{TOC} values between the dark and light layers. Hence, the δ^{13} C values of both the organic and inorganic components tend to discount the enhanced productivity hypothesis as the major control on valability in TOC within the dark and light layers, although it may have been a contributing factor.

388

389 4.3. Preservation

An enhanced preservation origin of TOC within the dark layers implies that the bottom 390 waters, present during the time that these layers were deposited, periodically became 391 392 anoxic thus inhibiting degradation of organic material. This is supported by the higher concentrations of redox sensitive elements such as Mo, V, Th, Cr, and U within these 393 organic-rich layers (Algeo and Lyons, 2006; Tribovillard et al., 2006). Under reducing 394 conditions Mo, V, Cr, and others are quickly adsorbed by Fe and Mn-oxyhydroxides or 395 incorporated into sulfides. As a consequence, while Modern anoxic basins show a high 396 concentrations of these elements, they are also scavenged from the bottom waters, 397 leading to an eventual depletion of the elements in poorly ventilated basins. Basins which 398 399 are very restricted tend to have lower concentrations, than those basins with moderate restriction. This situation well exemplified by Mo which has been widely used as a 400 paleoredox proxy (Algeo and Maynard, 2004; Piper, 1974; Sageman et al., 2003). 401 402 Variations in the concentration of Mo relative to TOC was studied in a number of Modern anoxic basins, including the Cariaco Basin, the Black Sea, Franvaren Fjord, and Saanich 403

Inlet (Algeo and Lyons, 2006). These basins have sedimentary Mo/TOC ratios ranging 404 from ~ 5 to 50 (x 10^{-4}). By comparison samples from the darker organic-rich layers of 405 Core U1466-56R have Mo/TOC ratios of approximately 30, similar to those found in the 406 Cariaco Basin and near the highest ratios found in Modern anoxic basins (Fig. 11). Using 407 the correlation between Mo/TOC ratios presented by Algeo and Lyons (2006) would 408 suggest a ventilation age of ancient inner sea of between 10 and 100 years, intermediary 409 between basins such as the Black Sea (ventilation age> 500 years) and the Saanich Inlet 410 (ventilation age < 2 years) (Algeo and Lyons, 2006). 411

412

Based on the variations in the concentrations of various trace elements, there appears to 413 be evidence for enhanced productivity, driven probably by input of dust, as well as anoxia 414 during the deposition of the dark layers. Such anoxic conditions could have arisen in 415 response to sea-level changes, allowing the bottom waters in the basin to become 416 isolated by virtue of a relatively shallow sill depth connecting the enclosed basin to the 417 surrounding ocean. Similar control has been postulated for the organic-rich sediments in 418 the Cariaco Basin (Peterson et al., 1991), and Black Sea (Lyons, 1991) and it is probably 419 more than coincidental that the Mo/TOC ratios within the dark layers from core U1466-420 56R are similar to those found in the Cariaco Basin. A potential model of this mechanism 421 in shown in Fig. 12. Support for this mechanism of sea-level variation driving restriction 422 and enhanced organic matter preservation comes from other paleoceanographic 423 indicators from this time. During the late Oligocene-Early Miocene changes in the volume 424 425 of the Antarctic ice sheet produced variations of approximately 50m in sea level, well documented in the benthic (*Cibicides* sp.) oxygen isotope record (Liebrand et al., 2017; 426

Zachos et al., 2001). Such variation is also present in the δ^{18} O values of the carbonate 427 fraction in Core U1466B-56R, which shows oscillations of about 1‰ between the dark 428 and light layers. The changes in TOC tend to lag changes in the δ^{18} O values, as would 429 430 be expected if sea level were a driving force in the preservation of TOC (Fig. 8). Sea level control of anoxia is also supported by the spectral analysis of the gamma ray and 431 resistivity signals obtained from the logging. These reveal the presence of signals 432 approximately coincident with tilt (40K) and eccentricity (100-120K) frequencies (Figs. 11 433 A & B). Higher gamma ray signals are predominantly a result of increases in the 434 concentration of U which correspond to periods of bottom-water anoxia which formed 435 during sea-level low stands. Similar spectral frequencies were noted in the Zachos et al. 436 (2001) and Liebrand et al. (2017) datasets. The 20K frequency was not observed in this 437 438 study because of the low spatial resolution of the logging tool relative to the rate of sediment deposition. We do not see a mechanism such as proposed for the production 439 of sapropels in Mediterranean in which increased freshwater input was responsible for 440 the development of anoxia (Emeis and Weissert, 2009; Rossignol-Strick, 1985). In the 441 case of the Maldives there is no riverine source, such as in the Mediterranean, and 442 therefore all the freshwater would have to be derived from rainwater. Not only would the 443 volumes of water necessary be unrealistic, but this freshwater would cause the carbonate 444 component to have more negative δ^{18} O values during the periods of dark layer formation, 445 opposite to what is observed. 446

447

448 4.4 Changes during the deposition of the 'Dark and Light' Interval

Several observations shed light on the evolution of the light and dark sequences. First, 449 the frequency of the alternations and variations in TOC were much greater in the older 450 portion of the sequence suggesting that the isolation of the basin was much more effective 451 during this time. Hence, later intervals did not experience the same degree of anoxia as 452 evidenced by lighter shade of 'dark' layers (lower TOC) and the presence of burrows. At 453 the same time the $\delta^{13}C_{TOC}$ values, while not showing significant changes between the 454 455 dark and light layers, shows an increase from the bottom to the top of the dark and light 456 interval (23.1 to 20 Ma) (Fig. 8 & 9). These more negative values may have originated 457 from higher input of terrestrial material at bottom of the sapropel-bearing interval. Such an interpretation is supported by the relative distribution of short (C₁₄–C₁₈) and long (C₂₄– 458 459 C₃₂) fatty acids (See supplemental material, Fig. S1-1 and S1-2) and changes in the C/N ratio of the TOC. The short chain fatty acids are mainly derived from aquatic organisms 460 such as phytoplankton, while long ones are derived from higher plants (Cranwell, 1974; 461 Ishiwatari et al., 2006; Meyers and Eadie, 1993). The trend within the dark and light 462 intervals at Site U1466 shows a pronounced decrease from the bottom to the top of the 463 interval. This may be a result of an increase in water depth from the late Oligocene to the 464 Early Miocene meaning that the changes in sea level would be less effective in producing 465 anoxia or alternatively a reduction in the magnitude of sea-level changes as suggested 466 by the data of Liebrand et al. (2017). At Site U1468 there is a statistically significant 467 decrease in the CN ratio in this interval, also interpretable as a reduction in the influence 468 of terrestrial plants (Meyers and Eadie, 1993). At Site U1466, the C/N ratio is already 469 close to the value expected for marine TOC throughout, and more positive $\delta^{13}C_{TOC}$ values 470 compared to Site U1468 suggest that the site received less terrestrial TOC. 471

473 4.6. Impact of Dark and Light Layers on Global Carbon Cycle

Numerous studies have noted that there was globally an approximate 0.5 ‰ increase in 474 the δ^{13} C values of certain foraminifera coincident with the Oligocene-Miocene boundary 475 (Hodell and Woodruff, 1994; Liebrand et al., 2017; Woodruff and Savin, 1989; Wright and 476 Miller, 1992), although no specific carbon burial event has been identified that coincides 477 478 with this change. Whilst the direction and the timing of the change is consistent with the increased burial of organic carbon noted in this paper, the amount of carbon theoretically 479 deposited in the Maldives during the time interval is insufficient to induce a global change 480 in the δ^{13} C value of seawater DIC of 0.5‰. However, it is possible that the phenomenon 481 of dark and light layers observed in the Maldives during this time was not unique and that 482 there were other systems, yet undocumented, where global sea level changes drove 483 similar changes in organic matter burial. For example, there is global wide increase of ~ 484 0.3 ‰ in the δ^{13} C value of oceanic carbonates during the time represented by the 485 sapropels (Hodell and Woodruff, 1994; Woodruff and Savin, 1989). However, as an 486 increase of 0.3 ‰ in the δ^{13} C value of the oceans would require about a 4% increase in 487 488 the global burial rate of organic carbon it seems unlikely that the Maldives alone could not account for the observed global change in δ^{13} C values. For example, if similar sapropels 489 were present in the Maldives at the same time, then the potential area available for these 490 to form would be ~45,000 km². Assuming that 50% of these strata were similar to the 491 dark layers found in this study and that they contained on average 10% organic carbon, 492 then the amount of buried organic carbon, based on a Berner type model (Berner et al., 493 1983), would be approximately 100 times less than needed to account for the observed 494

change in the global δ^{13} C value and other therefore to account for the 0.3 ‰ change in δ^{13} C values other sites of organic accumulation must have been active at this time.

497

498 **5.0 CONCLUSIONS**

499 We have identified the presence of ~100 m thick sequences of organic-rich sediments (sapropels) alternating with organic poor oceanic carbonates in late-Oligocene and Early-500 Miocene strata in the Maldives Archipelago. The organic-rich components contain up to 501 30% organic carbon and have trace elements signatures characteristic of formation within 502 anoxic bottom waters. These alternations occur within an atoll-like setting 1000 of meters 503 above the surrounding ocean basin. Based on stable O and C isotopic variations within 504 the carbonate components as well as spectral analysis of the NGR component of the 505 logging data, we propose that these sapropels formed in response to the variations in 506 507 orbitally induced sea-level fluctuations linked to Antarctic ice volume changes superimposed on a major transgression, which flooded a Late-Oligocene shallow-water 508 509 carbonate platform. During the orbitally induced changes, the bottom waters within the 510 atoll lagoon became anoxic and conducive to the preservation of organic material. As the water depth within the atoll became greater towards the end of the sapropels bearing 511 512 interval, the isolation of the bottom water became less effective and the intensity of the sapropels diminished. While sea level induced ventilation has been proposed for the 513 514 development of anoxia in closed basins such as the Cariaco Basin (Peterson et al., 1991) and Mediterranean (Emeis and Weissert, 2009) leading to the formation of sapropels, 515 516 subtle changes in sea level might also be responsible for the formation of organic-rich

intervals within carbonate platform settings (Koster et al., 1988) as demonstrated in this
study.

519

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646 **FIGURE CAPTIONS**

647

Figure 1: A) Map showing the location of Sites U1466 and U1468 drilled during IODP
Expedition 359. B) Seismic line showing the location and depth of penetration of Sites
U1466 and U1468.

651

Figure 2: Core photograph of dark and light layers from A) U1466B-56R-1 & 2, and B)U1468A-105X-1.

654

Figure 3: Facies and components of the dark and light layers. A) Close-up and B) photomicrograph of a dark layer at 790.8 mbsf at Site U1466 (U1466B-56R-1W-117) showing a wackestone with trace fossils (*Phycosiphon*-Ps, *Planolites*-Pl, and *Thalassinoides*-Th) and allochems (fish bones-FB, planktonic foraminifera-PF, and organic matter-TOC). C) Close-up and D) photomicrograph of a light interval at 806.960 mbsf at Site U1468 (U1468A-105X-1W-76) displaying a wackestone with a burrow of *Palaeophycus* (Pa), and common planktonic foraminifera (PF).

662

Figure 4: Scanning electron micrographs of fracture surfaces of sediment samples froma pair of light-dark levels with particularly good nannofossil preservation.

A, B) Dark layer, sample U1466B-56R-1W 45cm. The dominant coccolith
 is *Cyclagelosphaera floridana*, including several coccospheres, *Sphenolithus disbelemnos* is also common (conical nannoliths with honeycomb-like fabric).

668

C, D) Light layer, sample U1466B-56R-1W 35cm. The dominant coccoliths
are *Cyclagelosphaera floridana* (larger circular coccoliths) and *Umbilicosphaera jafari* (smaller circular coccoliths), note also strongly overgrown specimens of *Discoaster deflandrei* (lower right in D).

673

Figure 5: Changes in the percent organic content, $\delta^{13}C_{OM}$, $\delta^{15}N$ and C/N ratio between 700 and 800 mbsf from U1466 plotted against age. Also shown are the compilation of benthic $\delta^{18}O$ values for the same time period from Zachos et al. (2001).

677

Figure 6: Changes in the percent organic content, δ^{13} Com, δ^{15} N, C/N ratio and

NGR (logging data) between 700 and 800 mbsf from U1468.

680

Figure 7: Changes in the concentration of nutrients like elements (Cd and Ba), detrital elements (Fe and Al), and Redox elements (Mo, V, Th, and U)determined by ICP-MS compared to the percentage of organic carbon, the δ^{18} O and δ^{13} C values of the carbonate and the δ^{15} N values measured in the TOC in U1466B-56R-1. The solid line in each graph represents a three-point moving average. All metal concentrations in ppm.

Figure 8: XRF data from U1466B-56R-1 showing the variation in redox, dust, nutrient,
and organic related elements relative to the TOC and the dark and light layers.
Concentration units for all elements are in counts per second.

690

Figure 9: XRF data from U1468A-105x-1 showing the variation in redox and dust related
elements associated with the dark and light layers. Concentration units for all elements
are in counts per second.

694

Figure 10: A) Spectral analysis of natural gamma data obtained from the logging run at Site U1468 between a depth of 700 and 800 mbsf. B) Spectral analysis of resistivity data from the logging run at Site U1468 between a depth of 700 and 800 mbsf. For both gamma and resistivity age assignments have been made using shipboard biostratigraphy and have been interpolated to a time interval of 10,000 years.

700

Figure 11: Data from Core U146656R show the relationship between the concentration of Mo and the TOC (red symbols) relative to similar data published by Algeo and Lyons (2006) from the Black Sea and Saanich Inlet. Data for the dark intervals are highlighted in the oval while the light layers have values of Mo and TOC close to zero.

705

Figure 12: A cartoon demonstrating how sea level changes could lead to the development
of anoxia. In the upper panel the inter-atoll basin is connected to the surrounding ocean
and water flows freely into the basin. In the lower panel, sea level has fallen and access
to the basin become restricted leading to the development of anoxia.







712 Figure 2



714 Figure 3



716 A and B) U1466A-56x-1-45 (Black Layer)

C and D) U1466A-56x-1-35 (White Layer)

717 Figure 4



720 Figure 5





723 Figure 6



727 Figure 7



730 Figure 8



734 Figure 9



737 Figure 10





Figure 11





Figure S1-1: Representative mass fragmentograms (m/z 74) of fatty acids in the Maldives sediment (U1466) between 702.9 and 803.5 mbsf. Circles and each number denote fatty acids and carbon atom number, respectively.



Figure S1-2: Vertical profiles of n-C₂₂–C₃₂ fatty acids to n-C₁₄–C₃₂ fatty acids (long/short) ratio from U1466 between 702.9 and 803.5 mbsf. Long/short ratio; $(C_{24}+C_{26}+C_{28}+C_{30}+C_{32})/(C_{14}+C_{16}+C_{18})$.

Supplemental Material S2: Correlation between XRF and ICP-MS data

Samples for the ICP-MS were taken every 2 cms from Core 1466B-56R-1A and analyzed using the methods described in the accompanying paper. The XRF measurement were made on the archive half of the core at an interval of 1 cm. There were nine elements measured using both methods. Note that elements which gave negative counts using the scanning XRF were not compared. For comparison of the intensity data from the XRF with the ICP-MS data, the XRF data were interpolated to a sample interval of 2cm using a rectangular interpolation method adapted from Davis (1973). The data were then compared using a Pearson regression coefficient. As a result of the fact that the cores were not completed filled, there was the possibility that the depth in the archive and working portions of the core were different. For this reason the correlation between the XRF and ICP-MS data was subject to a lag analysis to determine the optimum correlation. The correlation coefficients are shown in Table S1. With the exception of Cr all showed positive covariance between the XRF and ICP-MS measurements. Generally there was about a 2-4 cm mismatch between the working and archive halves of the cores.

	0	2 cm	4 cm				
Fe	0.17	0.36	0.67				
Sr	0.26	0.36	0.28				
Мо	0.40	0.74	0.51				
Pb	0.05	0.37	0.28				
Cr	-ve	-ve	-ve				
Ва	0.02	0.16	0.13				
Al	0.06	0.27	0.47				
Mg	0.06	0.15	0.21				
Cd	0.21	0.43	0.06				

Table S1-1: R² values between XRF and ICPMS analyses as a function of lags. Those in red are statistically significant at the 99% confidence limits.

Further discrepancy between the two measurements arises as a result of nature of the analysis. The XRF technique targets the entire sample, including acid soluble and insoluble components, while the ICP-MS only analyzes the acid soluble portion plus any elements leached from the acid insoluble component. In addition the core is often heterogeneous in composition and the ICP-MS only analyses a small portion of the core while the XRF integrates signals from a much larger region. Consequently the concentrations measured by ICP-MS can be considered a minimum value. This differential leaching probably accounts for the discrepancies in the lag analysis which suggests different components were leached during the ICPMS analyses in different intervals. The absence of a correlation between Ba, Cr Mg and Cd probably results from low detection limits of the XRF.

References

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Figure S2-1: Comparison of concentration of Fe measured using ICPMS with the intensity derived from the XRF at a lag of 2 cm between working and archive halves of the core.



Figure S2-2: Comparison of concentration of Sr measured using ICPMS with the intensity derived from the XRF at a lag of 2 cm between working and archive halves of the core.



Figure S2-3: Comparison of concentration of Mo measured using ICPMS with the intensity derived from the XRF at a lag of 2 cm between working and archive halves of the core.



Figure S2-4: Comparison of concentration of Pb measured using ICPMS with the intensity derived from the XRF at a lag of 2 cm between working and archive halves of the core.



Figure S2-5: Comparison of concentration of Al measured using ICPMS with the intensity derived from the XRF at a lag of 2 cm between working and archive halves of the core.



Figure S2-6: Comparison of concentration of Mg measured using ICPMS with the intensity derived from the XRF at a lag of 2 cm between working and archive halves of the core.



Figure S2-7: Comparison of concentration of Cd measured using ICPMS with the intensity derived from the XRF at a lag of 2 cm between working and archive halves of the core.

Comparison between Scanning XRF and Quantitative XRF

In order to evaluate the reliability of the analysis of certain elements (Fe, Al, Sr, Si K, Ca, Br, Mo, and Ba), a companion study (Kunkelova et al., 2018) analyzed samples of pressed pellets using a fully quantitative conventional XRF method. Here we report on the correlation between the R² values between the intensities and the concentrations are shown in Table S1-2. These show a range of between < 0.01 and >0.9. In most instances the low correlation coefficients arise as a result low concentration of the element (i.e. Mg, Pb, Cr, Ba, Zn, Ti, and Ni) leading to low sensitivity. In other instances (i.e. Ca) it could result of the relatively invariant concentration within the samples analyzed. The fact that some elements (Fe, Sr, Al, Si, K, and Br) all show high correlation coefficients provides confidence that variations in the intensity of the signal derived from the scanning XRF reflect true variations within the measured section.

Fe	0.79
Sr	0.77
Mo	0.01
Pb	
Cr	0.01
Ва	0.00
AL	0.84
Mg	0.12
Cd	
Si	0.89
К	0.90
Са	0.02
Br	0.48
Zn	0.02
Ni	0.04
Ti	0.18

Table S1-2: Correlation between intensities made using the scanning XRF and the quantitative XRF on the identical samples. Samples highlighted in red are statistically significant (p<0.05).

Supplemental Material S3: Correlation between TOC and ICP-MS Data

The following section contains graphs of the correlation between the percentile of organic material and elements measuring using ICP-MS from Section U1466B-56R-1 between 789.65 and 790.8 mbsf.

Correlation coefficients between the concentration of organic carbon, organic nitrogen, CN ratio, $\delta^{13}C_{inoganic}$, $\delta^{13}O_{inorganic}$, $\delta^{13}C_{organic}$, $\delta^{13}N_{organic}$, and all the major and trace elements are included in Table S2-1.

Factor analysis of the data are shown in Figure S2-12 and Table S2-2.



Figure S3-1: Correlation between the percentile organic carbon and concentration of V using ICP-MS.



Figure S3-2: Correlation between the percentile organic carbon and concentration of Fe using ICP-MS.



Figure S3-3: Correlation between the percentile organic carbon and concentration of Mo using ICP-MS.



Figure S3-4: Correlation between the percentile organic carbon and concentration of Th using ICP-MS.



Figure S3-5: Correlation between the percentile organic carbon and concentration of Cd using ICP-MS.



Figure S3-6: Correlation between the percentile organic carbon and concentration of Pb using ICP-MS.



Figure S3-7: Correlation between the percentile organic carbon and concentration of U using ICP-MS.



Figure S3-8: Correlation between the percentile organic carbon and concentration of Cr using ICP-MS.



Figure S3-9: Correlation between the percentile organic carbon and concentration of Li using ICP-MS.



Figure S3-10: Correlation between the percentile organic carbon and concentration of K using ICP-MS.



Figure S3-11: Principal component analysis of ICPMS data identifying three components which account for 73% of the variance. Above shows the correlation between Factor 1 and Factor 2. Data within the ovals show loading > 0.7.

	Means	Std.Dev.	Depth	d15N	d13C	N%	C%	CN	d13Cin	d18Oin	Li	Na	Mg	Al	К	Ca	Ti	V	Cr	Mn	Fe	Sr	Мо	Cd	Pb	Th
Depth	790.2	0.22																								
d15N	0.0	0.97	0.25																							
d13C	-20.3	0.42	-0.25	0.09	1.00																					
N%	0.0	0.00	-0.15	-0.43	-0.14																					
C%	0.0	0.06	0.08	-0.69	-0.10	0.22																				
CN	17.3	3.19	-0.66	-0.38	0.17	0.38	0.25																			
d13Cin	1.1	0.44	-0.15	0.67	-0.05	-0.22	-0.73	-0.14																		
d18Oin	-1.9	0.31	-0.07	-0.31	-0.16	0.14	0.05	0.01	-0.04																	
Li	1.6	1.28	0.06	-0.59	-0.26	0.20	0.79	0.28	-0.70	0.08																
Na	4608.9	1697.91	-0.30	-0.05	0.21	0.00	-0.07	0.10	0.03	-0.20	-0.01															
Mg	2942.4	993.60	0.51	0.04	-0.16	-0.26	0.35	-0.14	-0.20	-0.18	0.51	0.04														
AI	1013.5	1149.32	0.12	-0.55	-0.34	0.21	0.73	0.23	-0.64	0.09	0.97	-0.05	0.55													
K	499.1	488.41	0.06	-0.58	-0.31	0.23	0.71	0.27	-0.62	0.12	0.96	0.04	0.53	0.99												
Ca	336388.1	64849.11	-0.07	0.52	0.10	-0.21	-0.54	0.00	0.60	-0.12	-0.47	0.32	0.12	-0.50	-0.46											
Ti	294.0	70.61	-0.04	0.41	0.15	-0.21	-0.46	0.15	0.53	-0.04	-0.32	0.10	0.22	-0.35	-0.32	0.87										
V	53.3	60.50	0.03	-0.73	-0.04	0.22	0.93	0.30	-0.82	0.13	0.86	-0.05	0.33	0.80	0.79	-0.58	-0.46									
Cr	18.3	19.46	0.05	-0.69	-0.01	0.23	0.93	0.31	-0.78	0.03	0.79	-0.07	0.30	0.76	0.75	-0.57	-0.44	0.96								
Mn	44.8	25.31	0.55	0.55	-0.01	-0.48	-0.27	-0.45	0.29	-0.23	-0.16	-0.01	0.72	-0.13	-0.16	0.51	0.52	-0.32	-0.33							
Fe	1609.1	1067.44	0.22	-0.51	-0.40	0.24	0.68	0.19	-0.56	0.08	0.93	-0.04	0.63	0.98	0.97	-0.42	-0.26	0.74	0.71	-0.02						
Sr	1677.3	363.81	-0.61	-0.14	-0.03	0.35	-0.11	0.59	0.22	0.10	0.02	0.41	-0.28	-0.05	0.02	0.52	0.47	-0.08	-0.08	-0.35	-0.06					
Мо	44.6	65.04	0.02	-0.70	0.00	0.21	0.89	0.27	-0.71	0.14	0.72	-0.05	0.22	0.63	0.60	-0.55	-0.46	0.90	0.85	-0.30	0.58	-0.09				
Cd	2.6	3.32	0.11	-0.65	-0.23	0.19	0.89	0.24	-0.63	0.06	0.73	-0.10	0.36	0.71	0.67	-0.59	-0.48	0.84	0.84	-0.21	0.71	-0.19	0.88			
Pb	0.8	0.67	0.06	-0.44	-0.36	0.31	0.68	0.28	-0.44	-0.04	0.84	-0.01	0.52	0.89	0.88	-0.38	-0.24	0.68	0.72	-0.11	0.90	0.04	0.51	0.69		
Th	0.3	0.27	-0.01	-0.32	-0.35	0.25	0.58	0.24	-0.44	-0.03	0.86	-0.05	0.37	0.87	0.83	-0.41	-0.28	0.63	0.62	-0.18	0.82	0.08	0.49	0.58	0.90	
U	17.8	18.20	0.16	-0.69	-0.18	0.24	0.93	0.25	-0.73	0.09	0.80	-0.09	0.38	0.78	0.75	-0.59	-0.46	0.93	0.93	-0.24	0.77	-0.16	0.90	0.96	0.74	0.62

Table S3-1: Correlation matrix of ICPMS, percent C and N, and isotopic data. Values in red show significant correlation at the 95% confidence limits.

	Factor - 1	Factor - 2	Factor - 3					
Depth	-0.07333	0.804629	-0.279774					
dN	0.72641	0.438625	0.103196					
dC	0.23302	-0.242078	-0.113724					
ON	-0.30728	-0.491290	0.077453					
00	-0.91027	-0.018920	-0.089858					
CN	-0.28911	-0.640059	0.445096					
dCin	0.77712	0.046759	0.266529					
dOin	-0.10928	-0.223983	-0.125105					
Li	-0.93069	0.102555	0.231112					
Na	0.07538	-0.208578	0.421388					
Mg	-0.38372	0.773377	0.393540					
AI	-0.91814	0.176567	0.224815					
K	-0.90013	0.118432	0.281959					
Ca	0.64484	0.103922	0.643812					
Ti	0.50959	0.141620	0.681751					
V	-0.94580	-0.068646	-0.062370					
Cr	-0.92109	-0.067770	-0.069652					
Mn	0.30877	0.842163	0.256335					
Fe	-0.87952	0.264845	0.285345					
Sr	0.09616	-0.618233	0.707159					
Мо	-0.84380	-0.115370	-0.169176					
Cd	-0.88552	0.042328	-0.115011					
Pb	-0.83442	0.150592	0.350650					
Th	-0.77644	0.093357	0.311428					
U	-0.94297	0.038720	-0.088403					
Expl.Var	11.86319	3.549261	2.721456					
Prp.Totl	0.47453	0.141970	0.108858					

Table S3-2: Factor loadings of the geochemical parameters. Factors >0.7 are marked.

Supplemental S-4: Comparison with benthic oxygen isotope data



Comparison of record of benthic $\delta^{\rm 18}{\rm O}$ values with data from U1466 and U1468.

Figure S4-1: Reproduction of Figure 5 with δ^{18} O values from Zachos et al (2001). The d18O values show a gradual increase consistent with an increase in sea level during this period.



Figure S4-2: Reproduction of Figure 6 with δ^{18} O values from Zachos et al (2001).