

Advantages
of a
Grazing Incidence Monochromator
in the
Extreme Ultraviolet

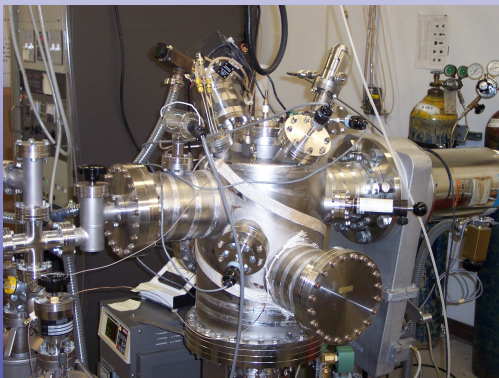
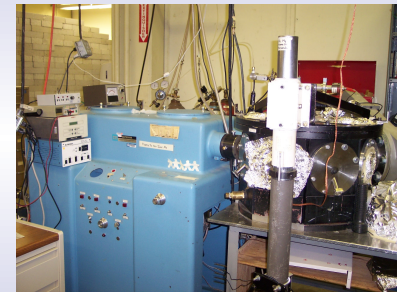
By Sarah Barton

Thin Films

We “grow” films with thicknesses in the range of 200-400 Å



We measure reflectance properties of different films in the extreme ultraviolet



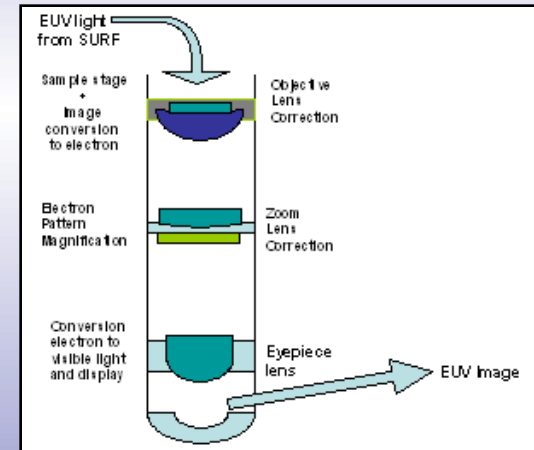
Methods used for creating films include sputtering and evaporation



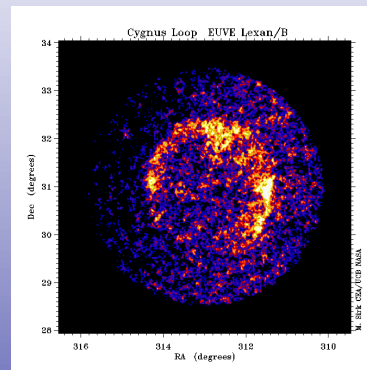
Past Projects and Applications



- ESA Mars Express Probe, Venus Express Probe

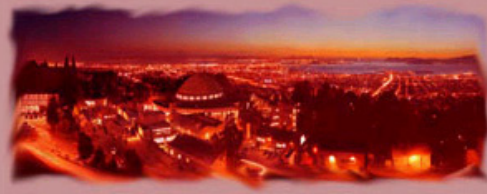


- Astronomy
- Microscopy
- Plasma Diagnostics



The Advanced Light Source at Lawrence Berkeley National Laboratories

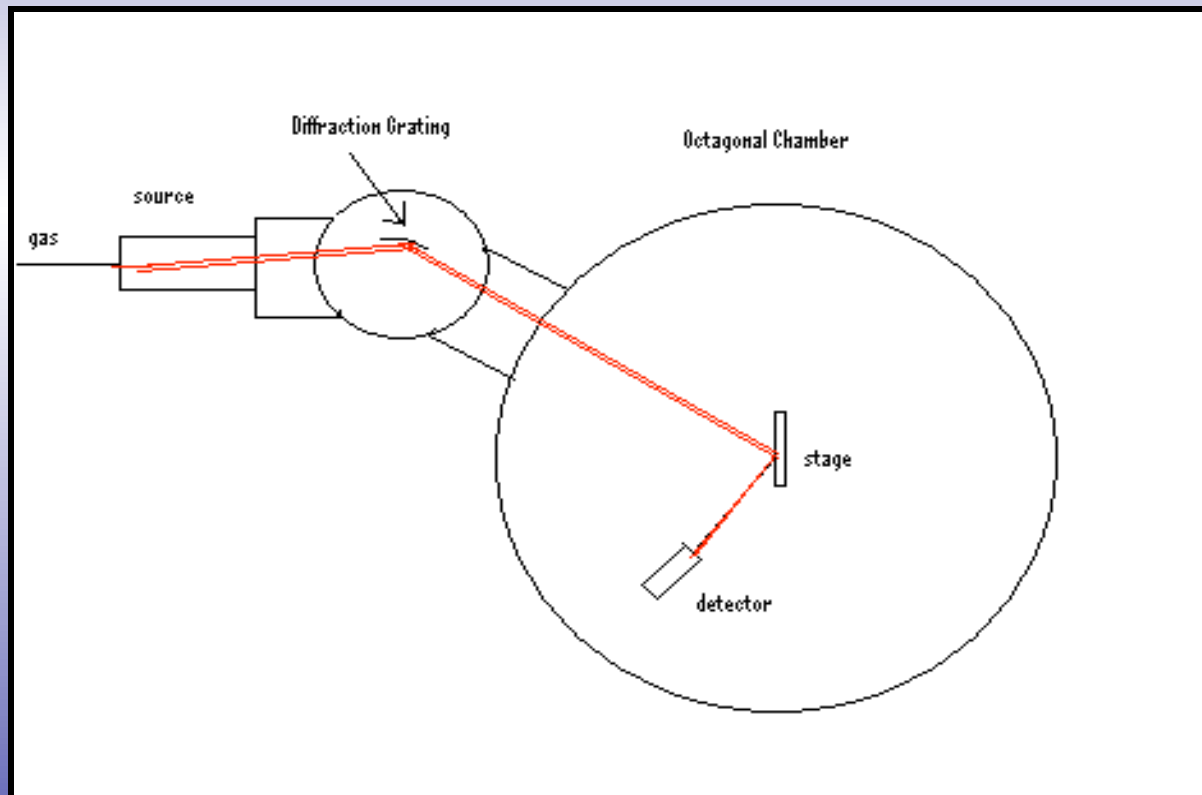
*Center for
X-Ray Optics*



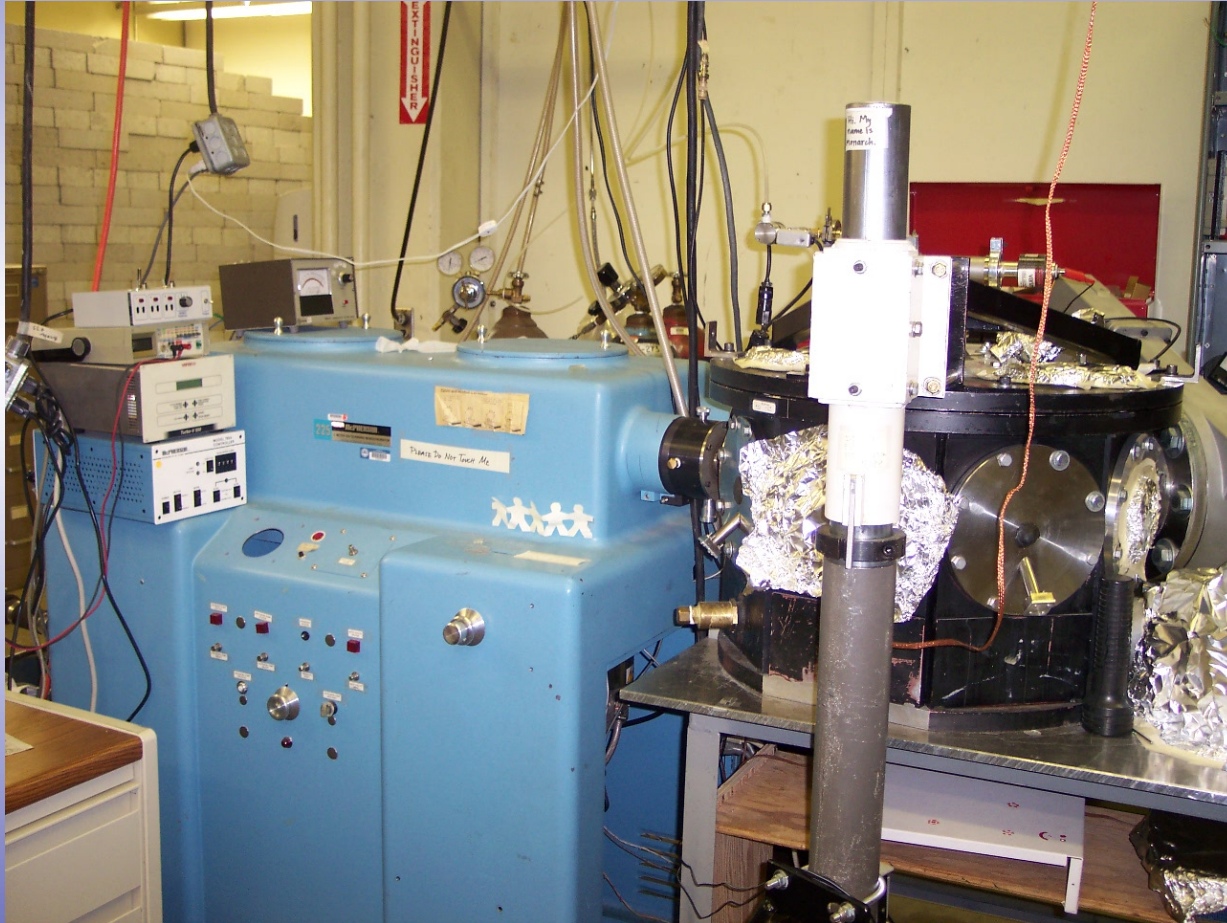
- Synchrotron
- High Intensity
- Shorter Wavelengths
- Continuous Spectrum
- Many Different Gratings
 - Including Grazing Incidence



What is a Grazing Incidence Monochromator Anyway?



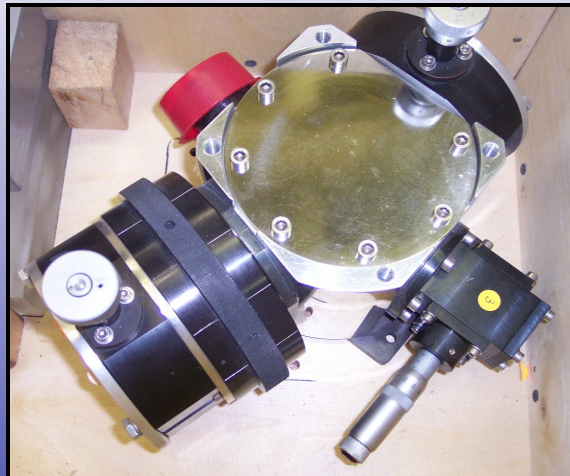
Set-Up



-- Insert
GIMS
here

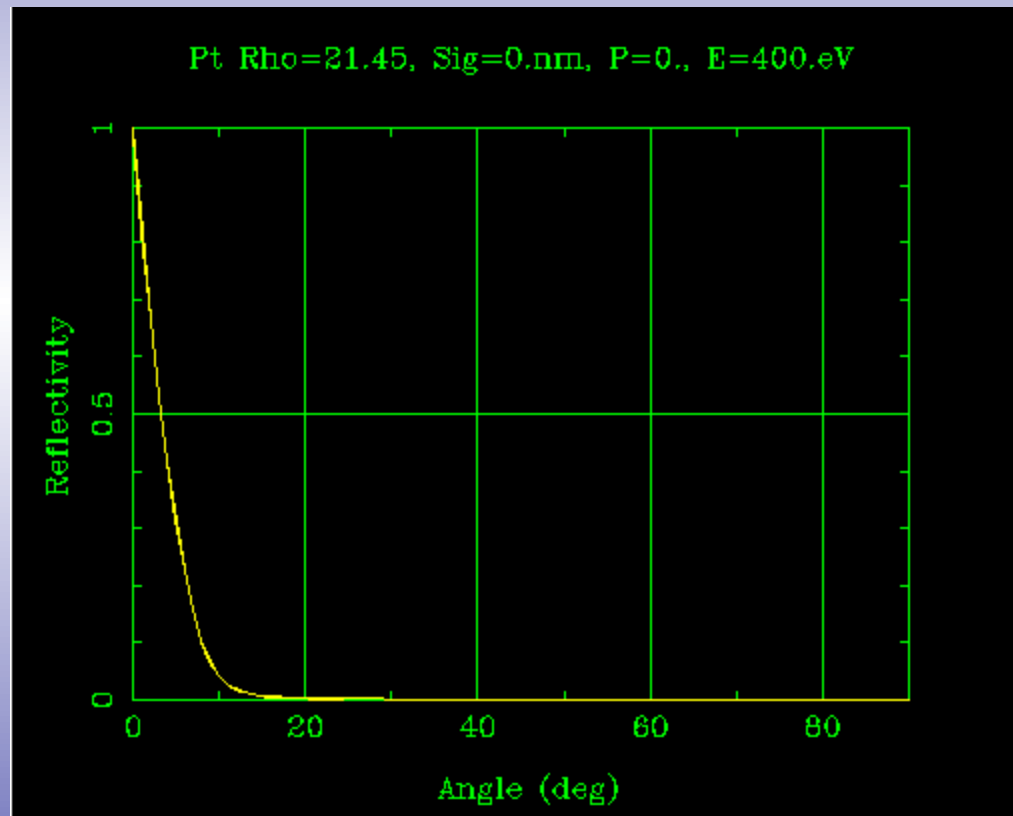
Differences

- Near-Grazing vs. Near-Normal Angles
- More Reflective for Higher Wavelengths
 - Possible Higher Intensity
 - Smaller Size

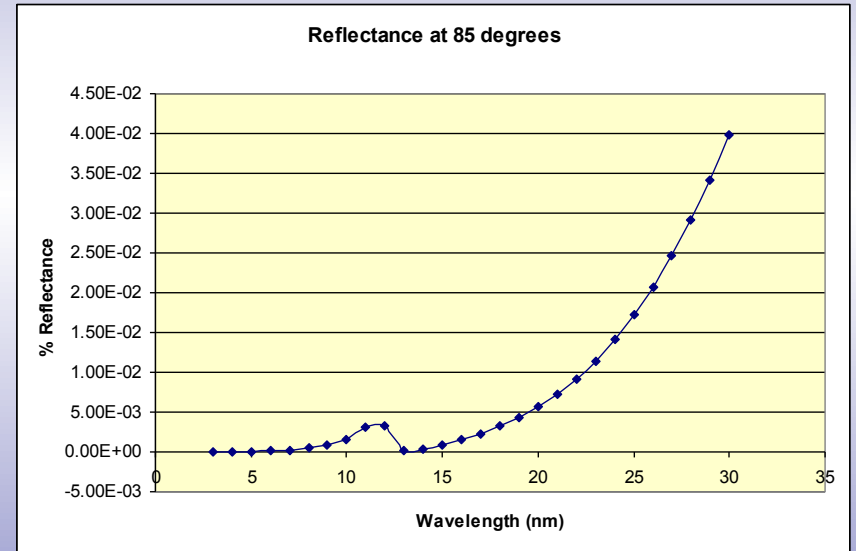
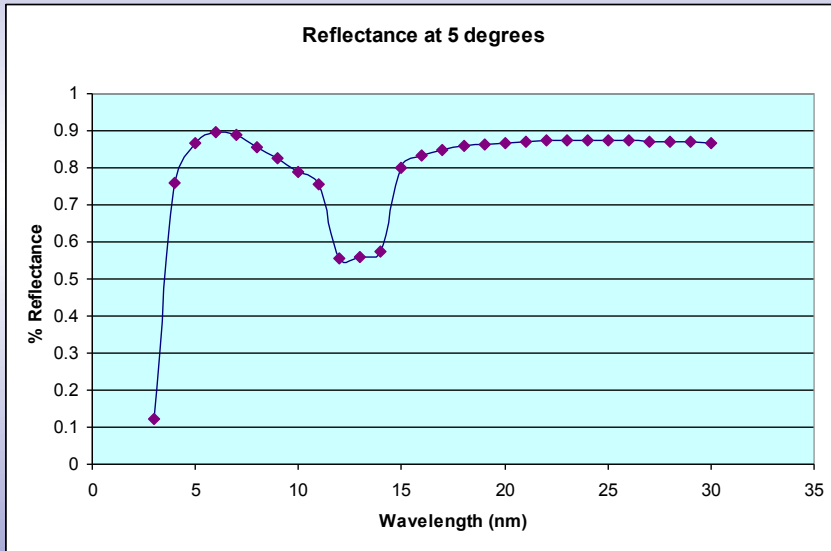


Reflection and Absorption

- 30 Å
- Pt



Why We Care



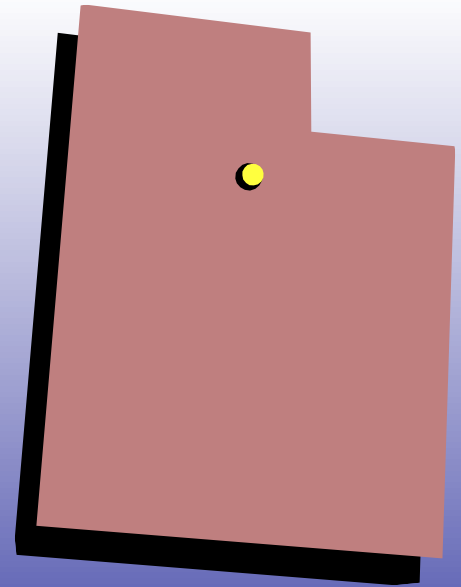
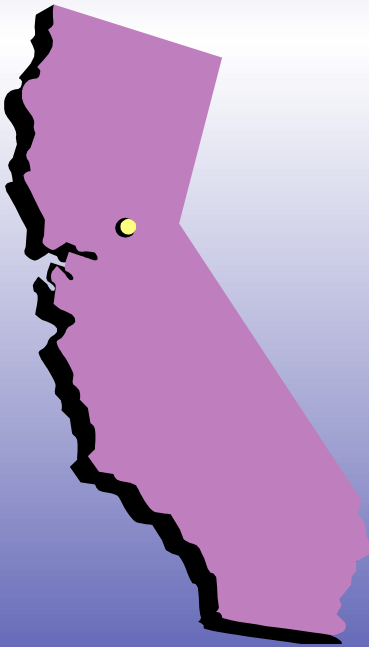
Benefits

- Less Absorption
- Higher Wavelengths
- 2 Sources

- And...

Convenience!

The Under Ground Lab
is a lot closer than
California!



Acknowledgements

- Dr. Alexander Shevelko
 - Dr. Turley
 - Dr. Allred
- BYU Thin Films Group
- NASA Rocky Mountain Consortium
 - John Ellsworth