



# Travelling Ionospheric Disturbances studied in the EU H2020 TechTIDE project

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### **Outline**

- Introduction
- TechTIDE objectives and approach
- Case study September 2017
- TechTIDE results
- Summary

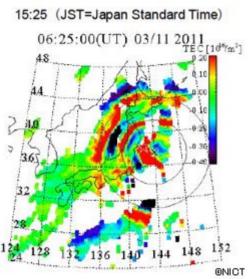






### **Medium Scale TIDs**

- MSTIDs are mostly associated with ionospheric coupling from below through AGW
- Periods: < 60 minutes,</li>
   Wavelength: < 1.000 km</li>



TEC variation about 38 minutes after the earthquake (EQ).

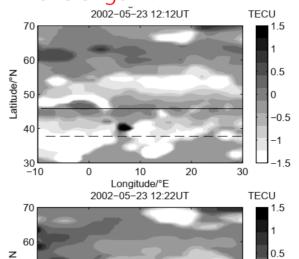
Tsugawa et al. (2011), Earth Planets Space





### **Large Scale TIDs**

- LSTIDs are associated with auroral and geomagnetic activity
- Periods: >= 60 minutes,
   Wavelength: > 1.000 km



10

Longitude/°E

20

30 <del>-</del>

LSTID signatures during 23<sup>rd</sup> May 2002 geomagnetic storm

Borries et al. (2009), Ann. Geophys.

-0.5



## Why do we care about TIDs



Operations at middle and low latitudes depending on predictable ionosphere characteristics are affected by TIDs:

- Systems using the ionosphere as part of their operations such as HF communication, HF geolocation operations
  - direction finding systems
  - > radio communication and broadcasting operations
  - > humanitarian aid organizations and radio amateurs
- Systems for which the ionosphere is a noise
  - > radio-astronomy experiments such as LOFAR and SKA
  - > ground and space-based augmentation systems (EGNOS and GBAS)
  - > enhanced precision positioning systems (such as N-RTK) which are extensively used in agriculture and drilling operations



## **TechTIDE** objectives

The overarching objective of TechTIDE is to design and test new viable TID impact mitigation strategies for the technologies affected by the TIDs and in close collaboration with operators of these technologies, to demonstrate the added value of the proposed mitigation techniques which are based on TechTIDE products.





National Observatory of Athens (NOA), Greece

Deutsches Zentrum für Luft- und Raumfahrt (DLR), Germanv

Ustav Fyziky Atmosfery AV CR (IAP), Czech Republic

Institut Royal Meteorolgique de Belgique (RMI), Belgium

Observatorio del Ebro Fundacion (OE), Spain

Borealis Global Designs Ltd. (BGD), Bulgaria

Leibniz Institute of Atmospheric Physics, Rostock University (L-IAP),

Germany

Universitat Politecnica de Catalunya (UPC), Spain

European Satellite Services Provider (ESSP), France

South Africa National Space Agency (SANSA), South Africa

Watermann Juergen Friedrich Wilhelm (JFWCONSULT), France

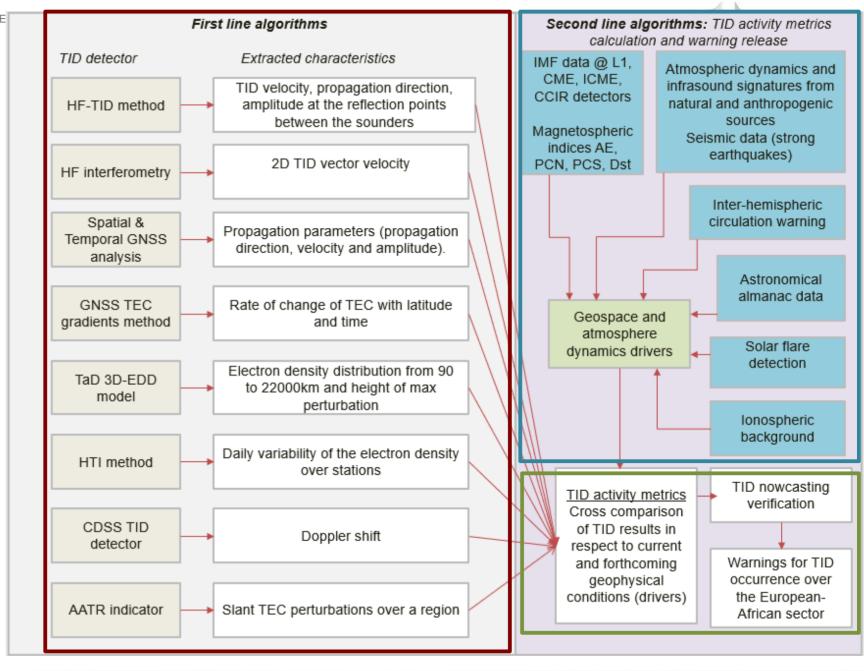
Frederick University (FU), Cyprus

German Federal Police (GFP), Germany



Funded by EC H2020 COMPET Space Weather Call 2017

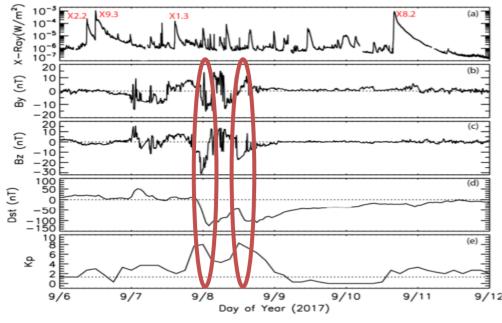
- TID detectors that provide near-real time measurements of TIDs
- Estimation of ionosphere background conditions
- Generation of TID activity metrics and information for non-scientific users

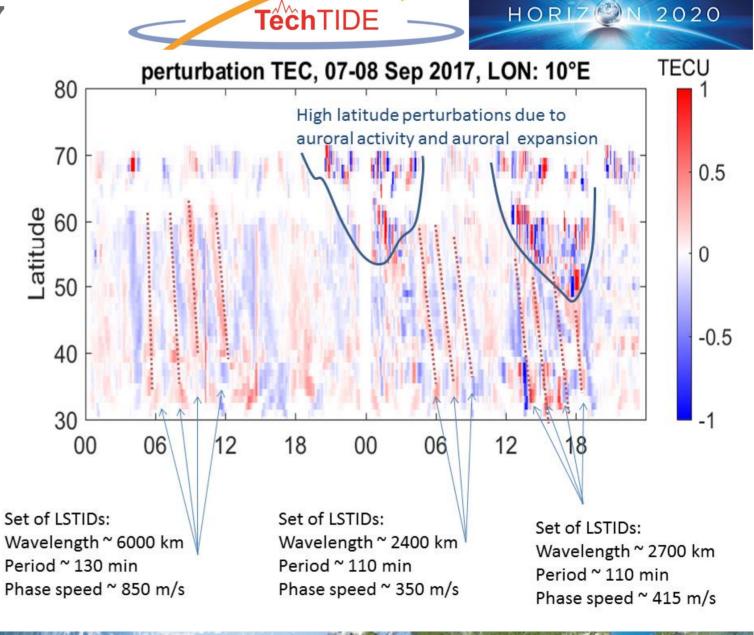




## **Event Analysis September 2017**

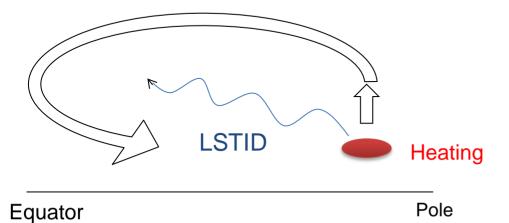
- GNSS based TID detection used for assessment and validation of TechTIDE products
- Many LSTIDs have been observed during 7th and 8th Sept. 2017

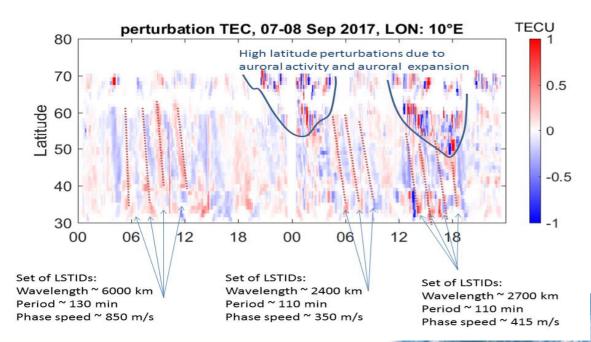






## **Theoretical Background**







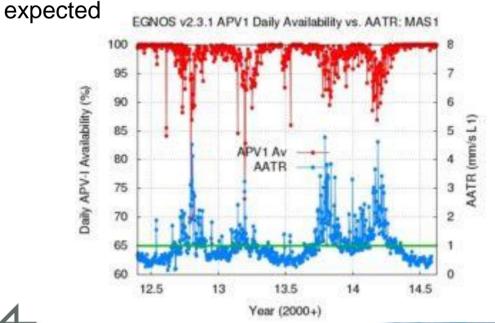
- Very sudden intensive heating generates large scale atmospheric gravity waves in the winds. They are observed in the Total Electron Content (TEC) as **Large Scale Travelling Ionospheric Disturbances** (LSTIDs)
- Source region of LSTIDs is charicterized by strong spatial gradients, depletion and significant fluctuations of the electron density and TEC

## LSTID activity - GNSS precursors

## **Along Arc TEC Rate (AATR)**

 This index based on TEC estimates derived from ground based GNSS measurements

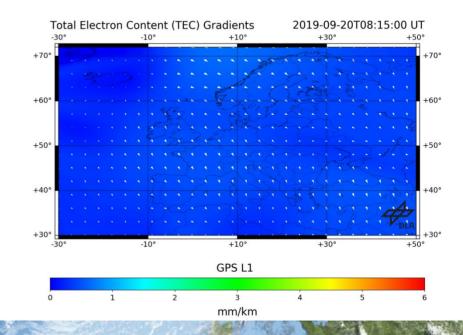
 used to identify the conditions where a degradation in the user performance of the SBAS systems is



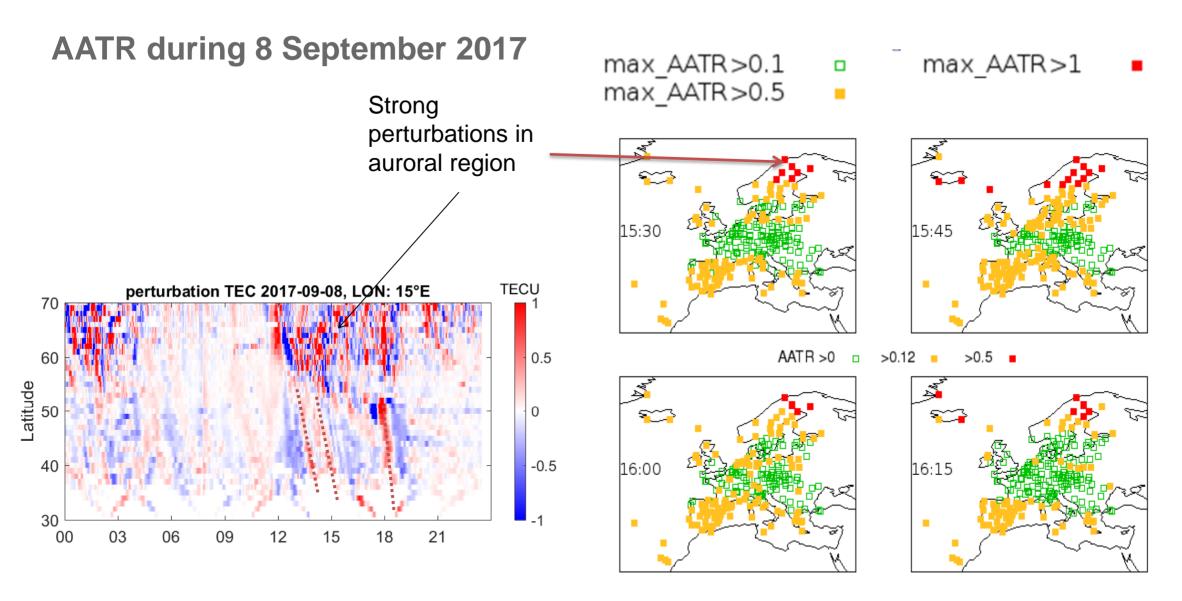


### **TEC Gradients**

- This product is derived from near real-time TEC maps
- Spatial horizontal gradients between the TEC maps (1°x1°) grid points are calculated

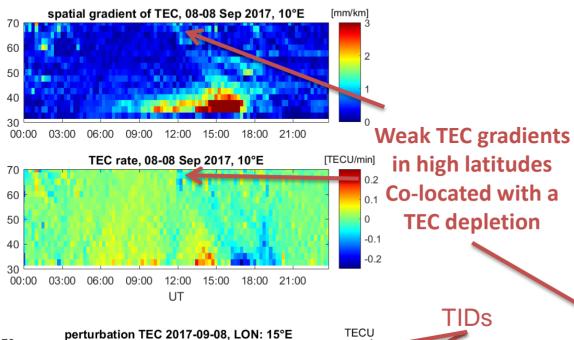


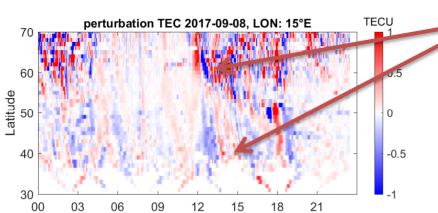






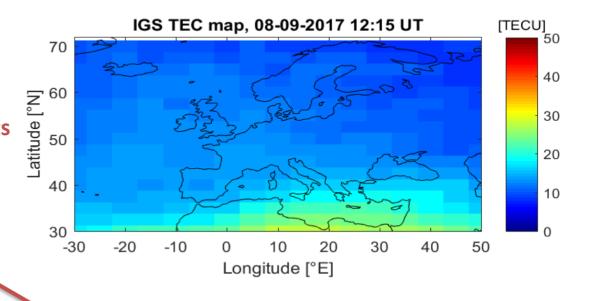
## **TEC gradients during 8 Sept 2017**

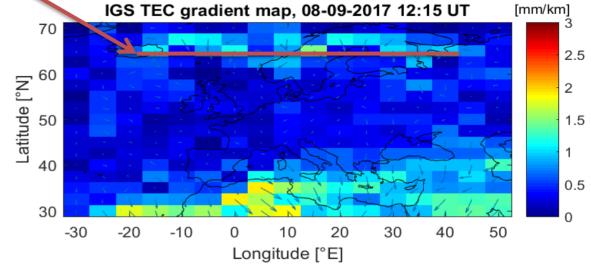














## HF Interferometry method performance

### 7 September 2017

Typical characteristics of LSTIDs whose origin might be auroral.

### Europe

The LSTID activity as observed by the HF-Interferometry reports a dominant disturbance for all the European stations with a periodicity of about 2-h and clear southward propagation occurring in the day-time and in the evening to mid-night hours.

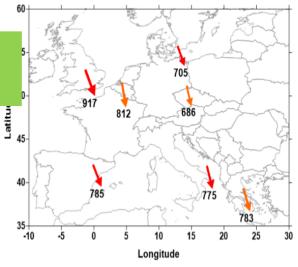
### **South Africa**

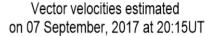
Stations in South Africa observe a dominant disturbance with a periodicity close to 1.5-h and clear northward propagation occurring in the evening.

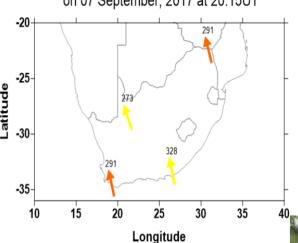




Vector velocities estimated on 07 September, 2017 at 12:00U7









## **HF-TID** method





**TéchTIDE** 

Strong TID activity in the evening hours.

Propagation: north-tosouth, velocity: 300 –

1000 m/s

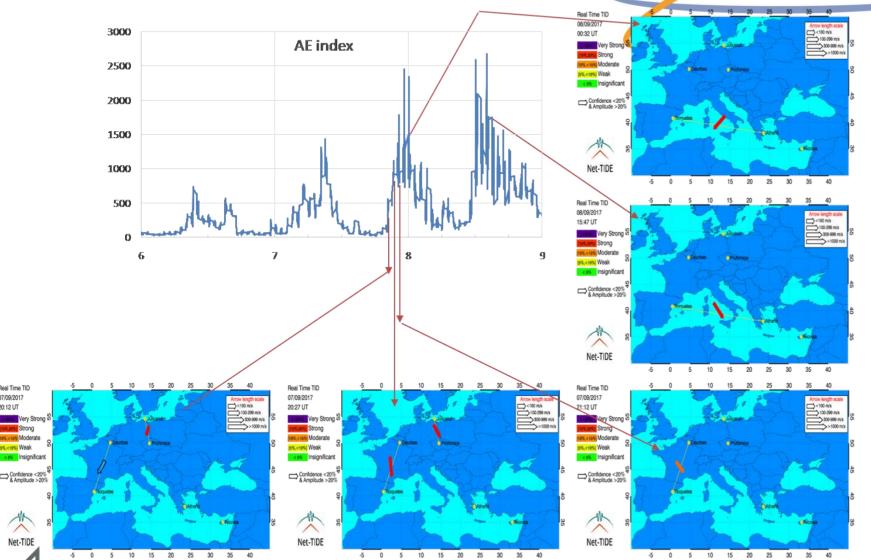
### **September 8:**

Strong TID activity in the early morning and daytime hours.

Propagation: north-to-

south, velocity: 300 -

1000 m/s

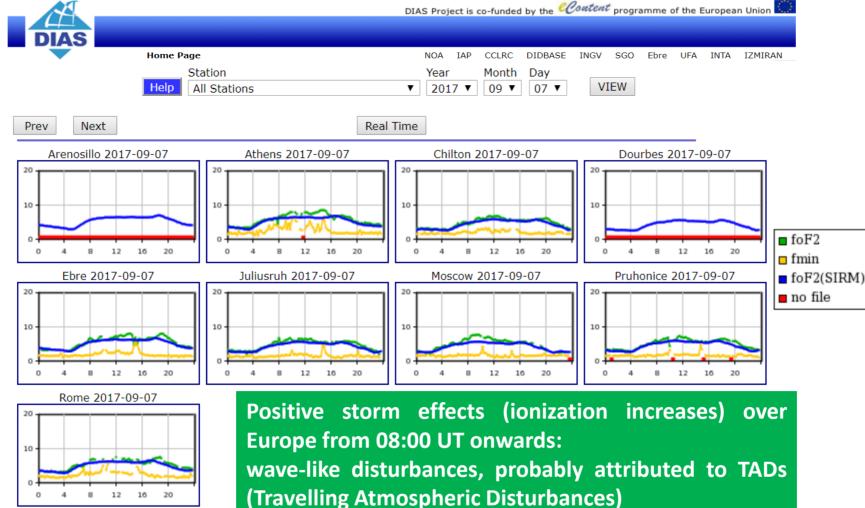


## **Background ionosphere**









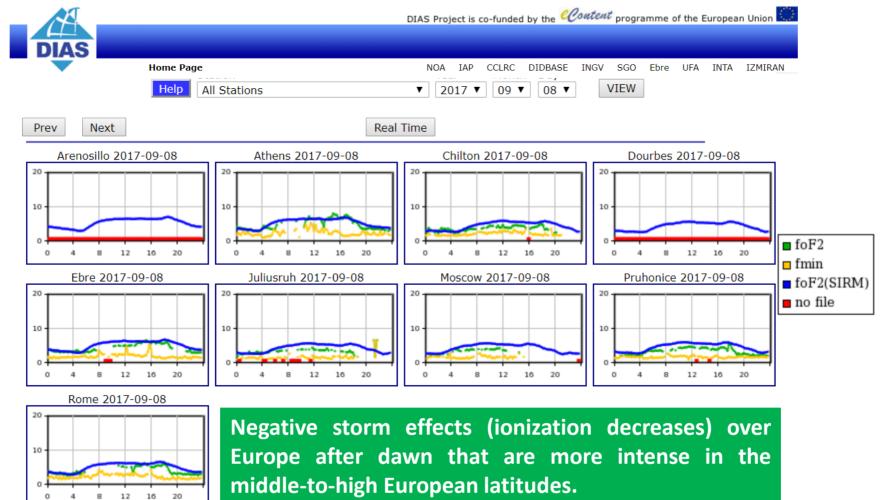


## **Background ionosphere**











## 1st Release of TechTIDE system







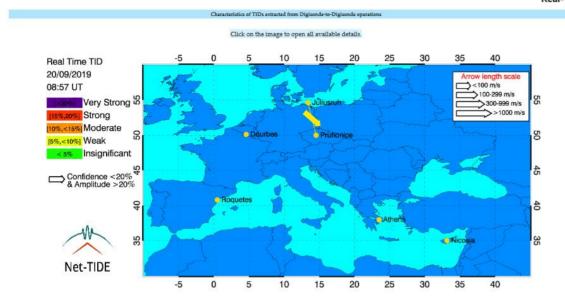
www.tech-tide.eu

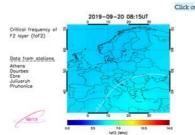


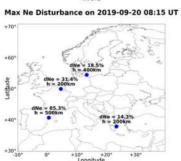
# TechTIDE Project New! Second TechTIDE Users' Workshop, 9 October 2019, Prague, Czech Republic, Click-Here!

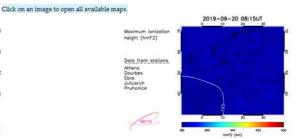
Real-Time Conditions

#### Real-Time Conditions









Max Ne Disturbance on 2019-09-20 08:20 UT



Electron density perturbation calculated with the TaD model

## **TechTIDE** activity report

- All results of TID detection are collected in an overview table
- Differences do result from different affected regions/altitudes in the ionosphere
- It illustrates the complexity of TID propagation in the ionosphere





Activity	Method/Indicators	Time interval		
		September, 6	September, 7	September, 8
Overall	<ul> <li>Solar flare indicators</li> <li>Geomagnetic activity (Kp, Dst)</li> <li>Magnetospheric activity (AE)</li> <li>AATR</li> <li>Ionospheric characteristics (SNR, foF2, TEC)</li> </ul>	Solar flare activity	<ul> <li>Enhanced geomagnetic and magnetospheric activity</li> <li>Enhanced ionospheric activity: Positive storm effects over Europe from 08:00 UT onwards</li> </ul>	<ul> <li>Enhanced geomagnetic and magnetospheric activity</li> <li>Enhanced ionospheric activity: Negative storm effects over Europe after dawn</li> </ul>
TIDs	AATR	-	-	Weak LSTIDs that propagate southwards from the auroral region (15:30 - 16:00 UT).
	HF-Interferometry	-	<ul> <li>Europe: LSTIDs with a periodicity of about 2-h and southward propagation (600-800 m/s) occurring in the day-time and in the evening to midnight hours</li> <li>South Africa: LSTIDs with a periodicity close to 1.5-h and northward propagation occurring in the evening (200-1000 m/s).</li> </ul>	<ul> <li>Europe: LSTIDs activity in early hours (until 2:00 UT) - azimuth of 180º (from true north) and a velocity of about 400m/s.</li> <li>South Africa: LSTIDs with a periodicity of about 1.5-h and northward propagation occurring in the day-time hours. The disturbance propagated with an azimuth of 0º (from true north) and a velocity of about 450m/s.</li> </ul>
	HF-TID	-	Europe: Strong TID activity in the evening hours.     Propagation: north-to-south, velocity: 300 – 1000 m/s	Europe: Strong TID activity in the early morning and daytime hours.     Propagation: north-to-south, velocity: 300 – 1000 m/s
	TaD			Disturbances in the daytime hours that are maximum in lower altitudes



## **Summary**

- TechTIDE is a project which aims to help users in the GNSS and HF application domains to mitigate the impact of TIDs
- Many different techniques for detection of TIDs have been implemented. They use different instruments and complement eachother in the detection of different types of TIDs
- The first release of TechTIDE near real-time services has been issued in May 2019
- Second and third release are planned for October 2019 and April 2020. We will add:
  - Activity reports
  - Alarms
  - User friendly navigation
  - ...



#### Overall coordinator:

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Twitter: @Tech\_TIDE <a href="https://twitter.com/Tech\_TIDE">https://twitter.com/Tech\_TIDE</a>

ResearchGate:

https://www.researchgate.net/project/TechTIDE-2



## Thank you for your attention



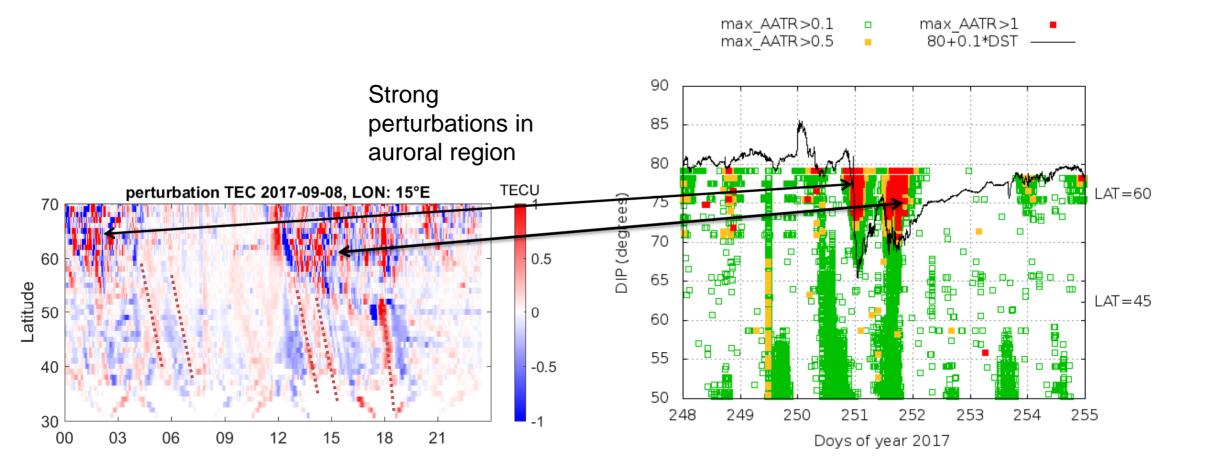


## **Backup slides**



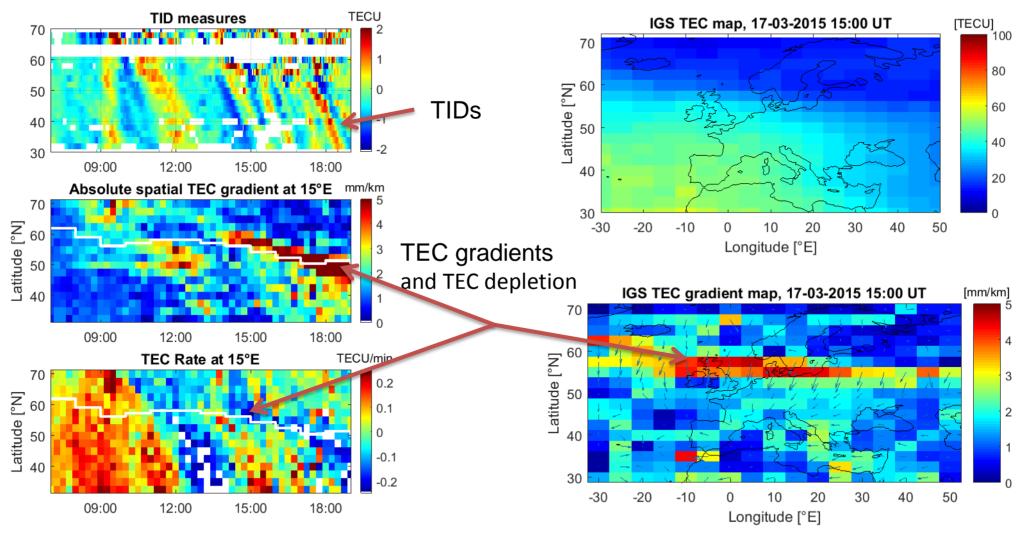


## **AATR during 8 September 2017**

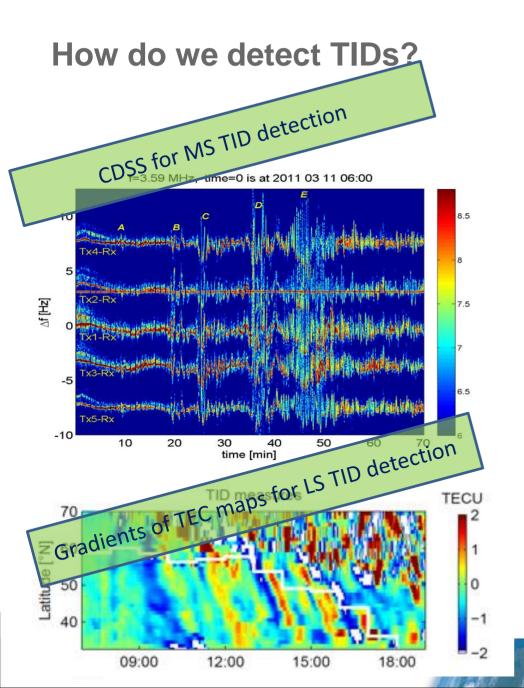




### **Event 17 Mar 2015**

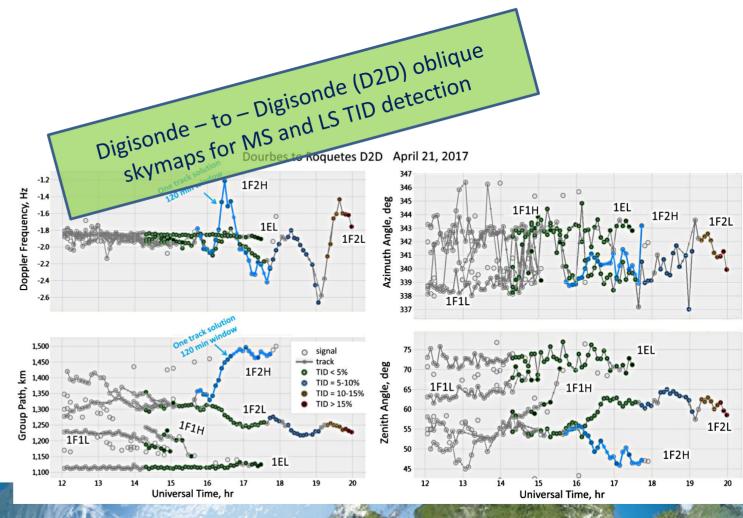


Borries, C., A. M. Mahrous, N. M. Ellahouny, and R. Badeke (2016), Multiple ionospheric perturbations during the Saint Patrick's Day storm 2015 in the European-African sector, J. Geophys. Res. Space Physics, 121, doi:10.1002/2016JA023178.









## HF interferometry method

### 8 September 2017

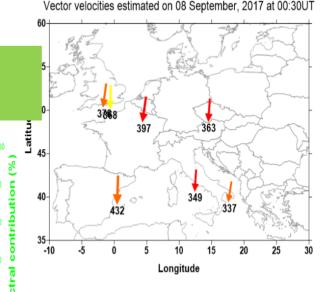
Typical characteristics of LSTIDs whose origin might be auroral.

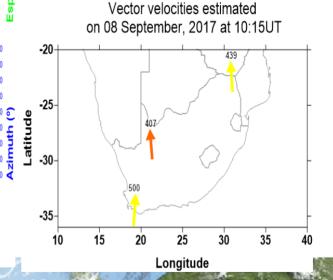
### **Europe**

LSTID activity was also observed in early hours of 8 September (until 2:00 UT) for all the European stations, as continuation of the event reported for 7 of September - azimuth of 180° (from true north) and a velocity of about 400m/s.

### **South Africa**

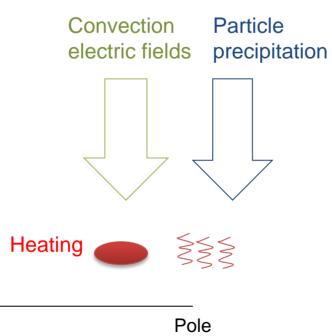
Dominant disturbance with a periodicity of about 1.5-h and northward propagation occurring in the day-time hours. The disturbance propagated with an azimuth of  $0^{\circ}$  (from true north) and a velocity of about 450 m/s.







## **Theoretical Background**



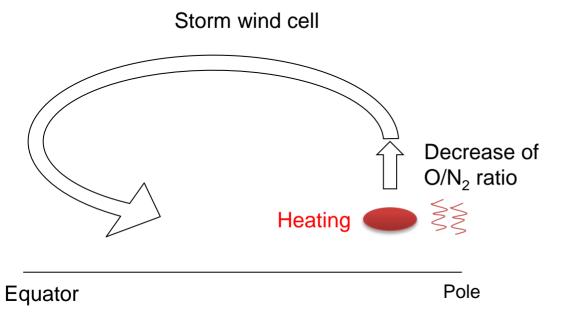


- During geomagnetic storms, large amounts of energy is transferred from the solar wind into the Earth system
- The energy is deposited in high latitudes, via convection electric fields and particle precipitation in the auroral region
- The electric fields drive currents in the auroral region
- Dissipation of currents and precipitation result in significant heating in the auroral region
- Electrojet heating has very large horizontal extension while
- Heating through precipitation is very localized



Equator

## **Theoretical Background**





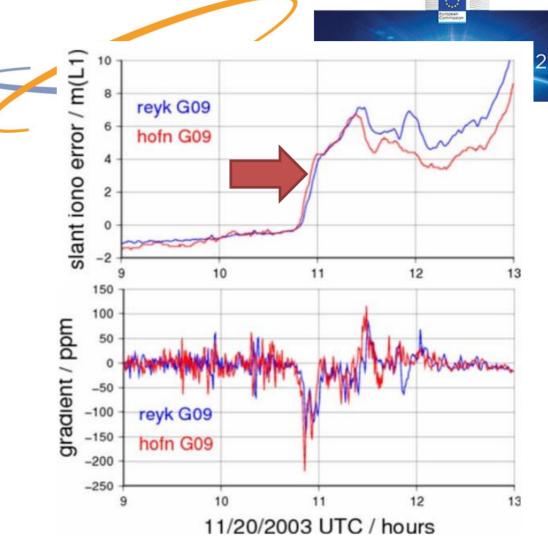
- Both heating effects have impact on thermosphereionosphere system
- Precipitation causes strong localized fluctuations in the thermosphere composition, winds and hence strong flucutuations in the electron density. It is associated with scintillation effects
- Current heating generates large scale wind systems and cause the expansion of the thermosphere.
  - The electron density in the auroral region decreases
  - The electron density in mid-latitudes increases
  - There are significant gradients in the electron density at the border of the auroral region



## Validation of method with published events

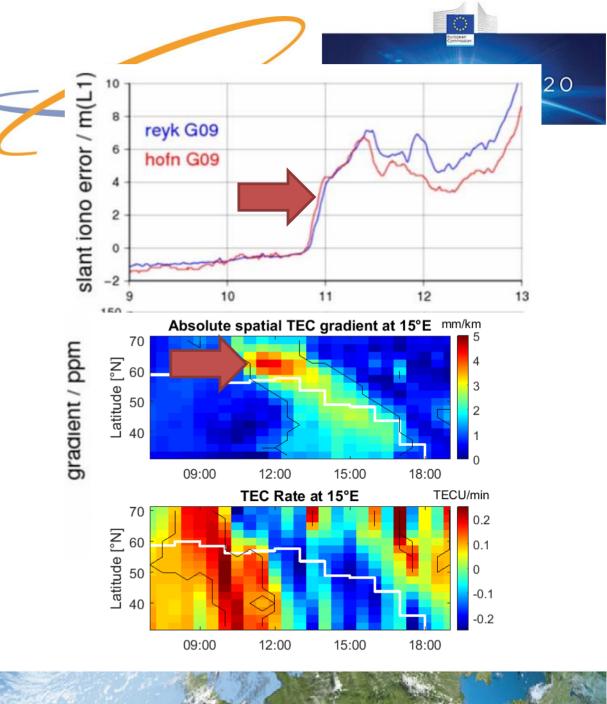
- Extreme TEC gradient published in Mayer et al. (2008)
- Max. gradient at around 10:55 UT at 63°N (Iceland)
- Max amplitude: 2.5 TECU/min
- Speed 117 m/s in North-South direction

Mayer, C. and Jakowski, N. and Borries, C. and Pannowitsch, T. and Belabbas, B. (2008) Extreme ionospheric conditions over Europe observed during the last solar cycle, Proceedings of the 4th ESA Workshop on Satellite Navigation User Equipment Technologies





- Extreme TEC gradient published in Mayer et al. (2008)
- Max. gradient at around 10:55 UT at 63°N (Iceland)
- Max amplitude: 2.5 TECU/min
- Speed 117 m/s in North-South direction
- TEC maps can identify significant gradient
- · Gradient moves equatorward
- Speed ~ 154 m/s → roughly same magnitude
- Max amplitude ~ 0.2 TECU/min → one magnitude less
- → Method applies to identify gradients correctly
- → However, the gradients are a lot sharper/ steeper than what can be identified in the TEC maps

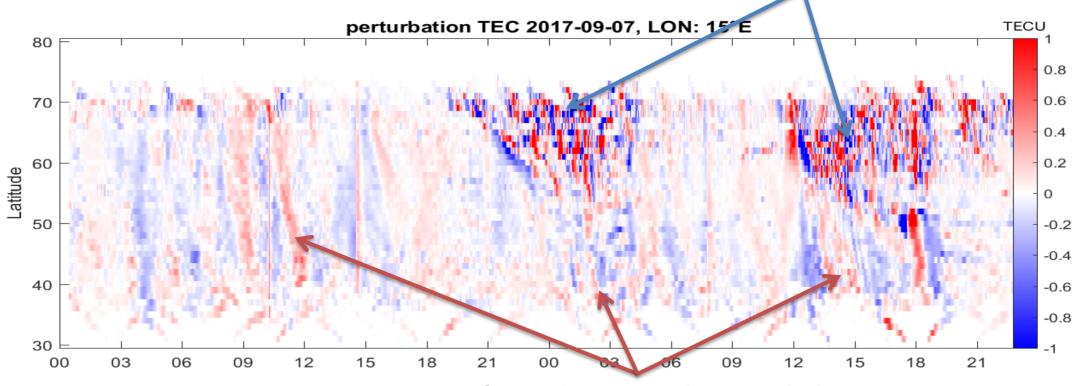


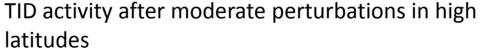


## **TID** activity observed in **TEC**



High latitude perturbations due to precipitation and current activity







## **Time-latitude plots of TEC gradients**





More perturbations in high latitudes, but no significant TEC gradients.

#### Disturbed Quiet spatial gradient of TEC, 07-08 Sep 2017, 10°E [mm/km] spatial gradient of TEC, 05-06 Sep 2017, 10°E [mm/km] 70 60 60 50 50 40 40 30 00:00 06:00 06:00 18:00 12:00 18:00 00:00 06:00 12:00 18:00 TEC rate, 07-08 Sep 2517, 10°E TEC rate, 05-06 Sep 2017, 10°E [TECU/min] [TECU/min] 70 0.2 0.2 60 60 0.1 0.1 50 50 -0.1-0.1 40 40 -0.2 -0.230 00:00 06:00 18:00 00:00 12:00 06:00 12:00 18:00 06:00 12:00 06:00 00:00 18:00 00:00 18:00 UT UT

