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The Lippmann–Schwinger Formula and One Dimensional Models with Dirac Delta Interactions

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Abstract We show how a proper use of the Lippmann–Schwinger equation $_{5}$ simplifies the calculations to obtain scattering states for one dimensional systems $_{6}$ perturbed by *N* Dirac delta equations. Here, we consider two situations. In the $_{7}$ former, attractive Dirac deltas perturbed the free one dimensional Schrödinger $_{8}$ Hamiltonian. We obtain explicit expressions for scattering and Gamow states. For $_{9}$ completeness, we show that the method to obtain bound states use comparable $_{10}$ formulas, although not based on the Lippmann–Schwinger equation. Then, the $_{11}$ attractive *N* deltas perturbed the one dimensional Salpeter equation. We also obtain $_{12}$ explicit expressions for the scattering wave functions. Here, we need regularisation $_{13}$ techniques that we implement via heat kernel regularisation.

Keywords Scattering states · Schrödinger and Salpeter one dimensional Hamiltonians · Contact perturbations · Gamow wave functions · Lippmann–Schwinger equation

1 Introduction

One of the more used tools in order to understand quantum mechanics are the ¹⁹ solvable models, in particular those which are one dimensional due to their ²⁰ simplicity [1–4]. The more often studied among these models is the free particle ²¹ Schrödinger Hamiltonian decorated with Dirac delta interactions. Relativistic one ²²

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dimensional approaches for the free particle Hamiltonian, such as those named after ²³ Salpeter or Dirac, have also been perturbed with contact interactions of delta type ²⁴ [5–7]. The purpose of the present article is to give a brief review of the recent work ²⁵ by the authors including the perturbation by *N* Dirac deltas of the one dimensional ²⁶ Schrödinger and Salpeter free Hamiltonians [6, 8, 9]. ²⁷

From the physics point of view, point potentials may represent interactions which ²⁸ are very localised in the space and strong and have a vast amount of applications for ²⁹ modelling real physical systems. A well-known model using Dirac delta potentials ³⁰ in non-relativistic quantum mechanics is the so-called Kronig–Penney model [10], ³¹ and it is actually a reference model in describing the band gap structure of metals ³² in solid state physics [11]. In addition, Dirac delta interactions in one or more ³³ dimensions serve as simple pedagogical toy models for the understanding of several ³⁴ quantum non-trivial concepts [12–19]. ³⁵

From the mathematical point of view, contact potentials are the result of the ³⁶ theory of self-adjoint extensions of symmetric operators with equal deficiency ³⁷ indices. In general, there are two methods to obtain these extensions. One is by ³⁸ defining some matching conditions at the nodes (points that support the contact ³⁹ potentials). Other uses the construction of the resolvent operator and often requires ⁴⁰ a renormalisation due to possible divergences in the construction of the resolvent of ⁴¹ the self-adjoint extension. Still a third method relies on a theorem of von Neumann ⁴² that characterises all self-adjoint extensions of a symmetric operator with equal ⁴³ deficiency indices, although this one has been less used. ⁴⁴

We also want to show how the Lippmann–Schwinger formula is useful for 45 this purpose as a simplifying computational tool. Here, we shall use the simplest 46 form of this equation which acquires mathematical sense on Gelfand triplets. The 47 Lippmann–Schwinger formula gives an equation satisfied by the incoming and 48 outgoing plane waves after a scattering process due to a potential *V*. It has the 49 following form: 50

$$|k^{\pm}\rangle = |k\rangle - R_0(E_k \pm i0) V |k^{\pm}\rangle, \qquad (1)$$

where $|k^{\pm}\rangle$ refers to the full scattered incoming (+) and outgoing (-) plane waves, 51 $|k\rangle$ is the free plane wave, V the potential and $R_0(E_k \pm i0)$ is the free resolvent, also 52 called the Green operator. Since it is a function of the complex variable z, R(z), and 53 has a branch cut at the spectrum of the free Hamiltonian (usually $\mathbb{R}^+ \equiv [0, \infty)$), we 54 denote by $R_0(E_k \pm i0)$ the upper and lower limits of R(z) as the imaginary part of 55 z goes to zero. Here, $E_k = (\hbar^2 k^2)/2m$. 56

This paper contains three more sections. In Sect. 2, we briefly discuss the 57 consequences of adding *N* Dirac delta perturbations to the one dimensional free 58 Schrödinger Hamiltonian. In Sect. 3, we do the same with the one dimensional 59 Salpeter Hamiltonian. The analysis of bound states is particularly relevant in both 60 cases. We finish our discussion with the concluding remarks.

2 One Dimensional Schrödinger Hamiltonian with N Dirac Delta Interactions

The objective of this section is to study the one dimensional Schrödinger Hamil- ⁶⁴ tonian $H_0 = \frac{p^2}{2m}$ perturbed by *N* Dirac deltas located at some points in the real ⁶⁵ axis. This study includes the search for bound states, scattering coefficients and ⁶⁶ resonances provided they exist. As is well known, this perturbed Hamiltonian has ⁶⁷ the form ⁶⁸

$$H := \frac{p^2}{2m} - \sum_{i=1}^N \lambda_i \,\delta(x - a_i) \,, \qquad V := -\sum_{i=1}^N \lambda_i \,\delta(x - a_i) \,, \tag{2}$$

where λ_i and i = 1, 2, ..., N, i = 1, 2, ..., N are *positive* real numbers. The a_i ⁶⁹ show the points supporting the deltas and are called *nodes*. Each of the $-\lambda_i$, with ⁷⁰ $\lambda_i > 0$, is the intensity of the delta located at a_i for all value of i. These coefficients ⁷¹ are chosen to be negative if we want to have bound states. The Schrödinger equation ⁷² produced by (2) is ⁷³

$$-\frac{\hbar^2}{2m}\frac{d^2\psi(x)}{dx^2} - \sum_{i=1}^N \lambda_i \,\delta(x - a_i)\,\psi(x) = E\psi(x)\,. \tag{3}$$

It is interesting to rewrite the interaction V in such a way that the calculations 74 with the aid of the Lippmann–Schwinger equation become easy. For simplicity, let 75 us assume that we have only one first. Then, the potential is $V = \lambda \delta(x - a)$ and the 76 wave function is $\psi(x) = \langle x | \psi \rangle$ [20–24]. In this notation, $(V\psi)(x) = \langle x | V\psi \rangle$ and 77 $\langle x | a \rangle = \delta(x - a)$. Thus, 78

$$(V\psi)(x) = \lambda \,\delta(x-a)\,\psi(a)\,. \tag{4}$$

Next, we note that the potential can be written as $V = \lambda |a\rangle \langle a|$, since then,

$$\langle x|V\psi\rangle = \lambda \,\langle x|a\rangle \langle a|\psi\rangle = \lambda \,\delta(x-a) \,\psi(a) = (V\psi)(x) \,. \tag{5}$$

The generalisation of the expression for the potential V in the case of having N_{80} nodes is the following: 81

$$V = -\sum_{i=1}^{N} \lambda_i |a_i\rangle \langle a_i|.$$
(6)

This is the desired expression. Let us clarify the vectors $|x\rangle$ for any real number ⁸² x are the generalised eigenvalues of the position (multiplication) operator in one ⁸³ dimension with eigenvalue x. As is well known, these vectors do not belong to the ⁸⁴

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Hilbert space on which the multiplication operator acts, but instead to an extension 85 of it endowed with a weak topology. We do not want to enter in these kind of details 86 here, see [21–24]. Vectors $|a_i\rangle$ are precisely of this type with $x = a_i$. 87

The first objective is the search for scattering states. We are introducing the ⁸⁸ procedure in the sequel, although we shall skip some steps in order to reach the ⁸⁹ final result as straightforward as possible. Details may be found in [8, 9]. Let us ⁹⁰ use (6) in the Lippmann–Schwinger equation (1) and multiply the result from the ⁹¹ left by the bra $\langle x |$. We have ⁹²

$$\langle x|k^{\pm}\rangle = \langle x|k\rangle + \sum_{j=1}^{N} \lambda_j \langle x|G_0(E_k \pm i0)|a_j\rangle \langle a_j|k^{\pm}\rangle.$$
(7)

For convenience, we shall use the notation $G_0(x, y; E_k \pm i0) := \langle x | G_0(E_k \pm 93) | y \rangle$ in the sequel. Also, we recall that $\langle x | k \rangle$ is the free plane wave and $94 \psi_k^{\pm}(x) := \langle x | k^{\pm} \rangle$ the perturbed plane wave in the coordinate representation. In 95 consequence, (7) can be written as (Henceforth we shall consider the sign plus in (7) 96 only, for simplicity. Similar results would be obtained with the other choice.) 97

$$\psi_k^+(x) = e^{ikx} + \sum_{j=1}^N \lambda_j \, G_0(x, a_j; E_k + i0) \, \psi_k^+(a_j) \,, \tag{8}$$

The goal is now to obtain the explicit form of $\psi^+(x)$, for which we have to find 98 the explicit form of the terms under the sum in (8). First, let as choose as values 99 of x in (8) the $\{a_j\}$. We obtain the following linear system of N equations for N 100 indeterminates: 101

$$e^{ika_i} = \psi^+(a_i) \left[1 - \lambda_i \, G_0(a_i, a_i; E_k + i0)\right]$$

$$-\sum_{j \neq i}^N \lambda_j \, G_0(a_i, a_j; E_k + i0) \, \psi^+(a_j) \,, \qquad i = 1, 2, \dots, N \,.$$
(9)

This system can be rewritten in matrix form. If $\Phi \equiv \{\Phi_{ij}\}$ is the $N \times N$ matrix 102 with matrix elements 103

$$\Phi_{ij}(E_k + i0) = \begin{cases} 1 - \lambda_i G_0(a_i, a_i; E_k + i0) \text{ if } i = j, \\ \lambda_j G_0(a_i, a_j; E_k + i0) \text{ if } i \neq j. \end{cases}$$
(10)

Then, Eqs. (9) take the form,

$$\sum_{j=1}^{N} \Phi_{ij}(E_k + i0) \psi_k^+(a_j) = e^{ika_j}, \qquad j = 1, 2, \dots, N,$$
(11)

with solution,

$$\psi_k^+(a_j) = \sum_{j=1}^N \left[\Phi^{-1}(E_k + i0) \right]_{ij} e^{ika_j} , \qquad (12)$$

where Φ^{-1} is the inverse of the matrix Φ . In consequence, the final form of (8) is 106

$$\psi_k^+(x) = e^{ikx} + \sum_{j=1}^N \lambda_j G_0(x, a_j; E_k + i0) \left[\Phi^{-1}(E_k + i0) \right]_{ij} e^{ika_j}.$$
 (13)

Then, we have to find the Green function $G_0(x, a_j; E_k + i0)$. We do not intend 107 to describe the procedure here, which is explained in detail in [9]. Once we have 108 obtained this Green function, using (10), we finally get all matrix elements of Φ . 109 The final results are 110

$$G_0(x, a_j; E_k + i0) = \frac{im}{\hbar^2 k} e^{k|x-a_i|}$$
(14)

and

$$\Phi_{ij}(E_k + i0) = \begin{cases} 1 - \frac{im\lambda_i}{\hbar^2 k} & \text{if } i = j, \\ -\sqrt{\lambda_i \lambda_j} & \frac{im}{\hbar^2 k} e^{ik|a_i - a_j|} & \text{if } i \neq j. \end{cases}$$
(15)

Then, we have determined all the perturbed plane waves $\psi_k^+(x)$. For $\psi_k^-(x)$, we 112 follow a similar procedure. Always recall that $E_k = (\hbar^2 k^2)/2m$. 113

2.1 Search for Bound States

So far, we have found the scattering states corresponding to the total (or perturbed) ¹¹⁵ Hamiltonian, for which we have used the Lippmann–Schwinger equation as main ¹¹⁶ tool. Next, we search for the possible existence of bound states, where the search ¹¹⁷ could be carried out with similar tools to those used in the precedent discussion. ¹¹⁸

We proceed as follows: Let us use the simplified notation $|f_i\rangle := \sqrt{\lambda_i} |a_i\rangle$, so 119 that the total Hamiltonian (2) may be written as 120

$$H = \frac{p^2}{2m} - \sum_{i=1}^{N} |f_i\rangle \langle f_i|.$$
 (16)

The corresponding Schrödinger equation reads

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$$\left\langle x \left| \frac{p^2}{2m} \right| \psi \right\rangle - \sum_{i=1}^{N} \left\langle x \right| f_i \right\rangle \left\langle f_i \right| \psi \right\rangle = E \left\langle x \right| \psi \right\rangle .$$
(17)

Bound states correspond to solutions of (17) with negative *E* and square integrable the wave function $\psi(x) \equiv \langle x | \psi \rangle$.

Next, insert the completeness relation $1 = \frac{1}{2\pi\hbar} \int |p\rangle \langle p| dp$ in front of $|\psi\rangle$ and 124 $|f_i\rangle$. Define $\tilde{\psi}(p) := \langle p|\psi\rangle$, which is indeed the Fourier transform of $\langle x|\psi\rangle$, and 125 write $\phi(a_i) := \langle f_i|\psi\rangle = \sqrt{\lambda_i} \langle a_i|\psi\rangle = \sqrt{\lambda_i} \psi(a_i)$. Recall that $\langle x|p\rangle = e^{\frac{i}{\hbar}px}$. 126 Then, (17) becomes 127

$$\int_{-\infty}^{\infty} \frac{dp}{2\pi\hbar} e^{\frac{i}{\hbar}px} \,\widetilde{\psi}(p) \left(\frac{p^2}{2m} - E\right) = \sum_{i=1}^{N} \sqrt{\lambda_i} \int_{-\infty}^{\infty} \frac{dp}{2\pi\hbar} e^{\frac{i}{\hbar}p(x-a_i)} \,\phi(a_i) \,.$$
(18)

From (18) and the properties of the Fourier transform, we have that

$$\widetilde{\psi}(p) = \sum_{i=1}^{N} \sqrt{\lambda_i} \, \frac{e^{-\frac{i}{\hbar} p a_i}}{\frac{p^2}{2m} - E} \phi(a_i) \,. \tag{19}$$

But $\tilde{\psi}(p)$ is the Fourier transform of the solution $\psi(x)$ of the Schrödinger 129 equation (17). Let us use this idea to conclude that (take $x = a_i$) 130

$$\psi(a_i) = \sum_{i=1}^N \sqrt{\lambda_i} \int_{-\infty}^\infty \frac{dp}{2\pi\hbar} \frac{e^{-\frac{i}{\hbar}pa_i}}{\frac{p^2}{2m} - E} \phi(a_i) \,. \tag{20}$$

Multiply both sides in (20) by $\sqrt{\lambda_i}$ and recalling that $\phi(a_i) = \sqrt{\lambda_i} \psi(a_i)$, we 131 arrive to an equation of the form: 132

$$\sum_{j=1}^{N} \Phi_{ij}(E) \phi(a_j) = 0.$$
(21)

Find details in [8]. It is beyond a mere coincidence that the matrix elements $\Phi \equiv 133$ { $\Phi_{ij}(E)$ } are identical to those of (15) with the replacement $k = \sqrt{2m|E|}$, so that [8] 134

$$\Phi_{ij}(E) = \begin{cases} 1 - \frac{m\lambda_i}{\hbar\sqrt{2m|E|}} & \text{if } i = j, \\ -\frac{m\sqrt{\lambda_i\lambda_j}}{\hbar\sqrt{2m|E|}} \exp\left(-\sqrt{2m|E|} |a_i - a_j|/\hbar\right) & \text{if } i \neq j. \end{cases}$$
(22)

Since Eq. (21) has come directly from (17), it is a necessary condition for the ¹³⁵ existence of solutions of (17) with the desired properties. This equation has non- ¹³⁶

trivial solutions $\{\phi(a_j)\}\$ if and only if det $\Phi(E) = 0$. Therefore, the bound states 137 energies are solutions of the transcendental equation det $\Phi(E) = 0.^1$ 138

For a systematic calculation of the bound states, let us consider the following 139 eigenvalue problem: 140

$$\Phi(E)A(E) = \omega(E)A(E), \qquad (23)$$

where $\omega(E)$ are the eigenvalues of the $N \times N$ matrix $\Phi(E)$ and A(E) their 141 corresponding eigenvectors. Equations (21) and (23) coincide if and only if $\omega(E) = 142$ 0 and then, the bound states energies have to be the solutions of the transcendental 143 equation $\omega(E) = 0$ its eigenvectors being those with components equal to $\phi(a_j)$. 144 If we assume no degeneracy, the wave function corresponding to the energy value 145 E_i with eigenvector $A(E) \equiv (\phi(a_1), \dots, \phi(a_N))$ takes the form (19) with $E = E_i$. 146 In the coordinate representation, the wave function is just its Fourier transform. For 147 further comments, see [6, 8].

2.2 Resonances and Gamow States

The Lippmann–Schwinger equation is also useful for the construction of Gamow 150 states, which are vector states for resonances. In a resonant scattering process [25] 151 produced by a Hamiltonian pair, say $\{H_0, H\}$, where H_0 is a *free* Hamiltonian and 152 $H = H_0 + V$, where V is the interaction, the Gamow vectors, ψ^{\pm} , for a resonance 153 with energy E_R and inverse of the mean life given by Γ are two eigenvectors of H 154 with respective eigenvalues $E_R \pm \Gamma/2$, i.e., $H\psi^{\pm} = (E_R \mp \Gamma/2) \psi^{\pm}$ [25]. This 155 property shows that the Gamow vector ψ^+ decays exponentially as $t \mapsto \infty$ (and 156 ψ^- decays exponentially as $t \mapsto -\infty$, they are time reversal of each other). This 157 situation produces two problems, one from the point of view of physics and the other 158 from the point of view of mathematics. 159

Although exponential decay for simple quantum unstable systems has been 160 detected for essentially for all values of time, deviations for these exponential law 161 have been detected for very short or very large times [26, 27]. Since these deviations 162 certainly occur under these conditions only, they are very difficult to be detected. For 163 most values of time, exponential decay serves as an excellent approximation. This 164 is why Gamow vectors are useful as good approximations of decaying states. 165

A self-adjoint operator on Hilbert space, as is the case of the Hamiltonian H, 166 cannot have complex eigenvalues with corresponding eigenvectors in this Hilbert 167 space. Thus, Gamow vectors are well-defined objects on some extensions of Hilbert 168 spaces called rigged Hilbert spaces [25, 28–30]. 169

¹As a matter of fact, this also follows because $\Phi(E)$ appears in the denominator of the resolvent of the total Hamiltonian *H*.

Let us briefly sketch the use of (1) to obtain an explicit expression of the Gamow 170 vectors as eigenvectors of H with eigenvalue $E_R \pm \Gamma/2$. Details may be found 171 in [9, 31]. If we multiply Eq. (1) to the right by the bra $\langle \psi |$, we obtain a complex 172 function on the variable k. With adequate choices of the space of bras, this results on 173 meromorphic functions of complex variable defined at least on a half plane [29, 30]. 174 Let us assume that this is the case and omit the bra $\langle \psi |$. Then, if we define k_R as 175

$$z_R := E_R - \Gamma/2 = \frac{k_R^2 \hbar^2}{2m},$$
(24)

we may consider the analytic extension of (1) to the value of k given by k_R ,

$$|k_R^+\rangle = |k_R\rangle - G_0(z_R) V |k_R^+\rangle.$$
⁽²⁵⁾

It is important to remark that z_R is a pole of the Green function corresponding to the total Hamiltonian H, but not of the free Hamiltonian H_0 , just by the characterisation of resonances using the resolvent [32]. Then, $G_0(z_R)$ is well defined and so is $|k_R^+\rangle$, 179 which has the property [9, 31, 33] 180

$$H|k_R^+\rangle = z_R |k_R^+\rangle.$$
⁽²⁶⁾

Thus, $|k_R^+\rangle$ is one of the Gamow vectors with resonance pole z_R (the other can be 181 obtained exactly in the same way, just replacing z_R by its complex conjugate z_R^* and 182 taking the minus sign in (1). This Gamow vector in the coordinate representation is 183 $\psi_R^+(x) := \langle x | k_R^+ \rangle$, so that 184

$$(H\psi_R^+)(x) = \langle x|H|k_R^+\rangle = z_R \langle x|k_R^+\rangle = z_R \psi_R^+(x).$$
⁽²⁷⁾

Now, let us go back to the *N* Dirac deltas interaction and, consequently, take 185 in (25) the form of the potential given by $V = -\sum_{i=1}^{N} \lambda_i |a_i\rangle \langle a_i|$. Multiply the 186 result of this operation to the right by the bra $\langle x|$ and divide k_R into real and 187 imaginary parts, $k_R = k_r - ik_I$. We have that $\langle x|k_R \rangle = e^{ik_R x} = e^{ik_r x} e^{-ik_I x}$ and 188

$$\psi_{k}^{+}(x) = \langle x | k_{R}^{+} \rangle = \langle x | k_{R} \rangle + \sum_{i=1}^{N} \lambda_{i} \langle x | G_{0}(z_{R}) | a_{i} \rangle \langle a_{i} | k_{R}^{+} \rangle$$

$$= e^{ik_{r}x} e^{k_{I}x} + \sum_{i=1}^{N} \lambda_{i} G_{0}(x, a_{i}; z_{R}) \psi_{R}^{+}(a_{i}) = e^{ik_{r}x} e^{k_{I}x} \qquad (28)$$

$$+ \sum_{i=1}^{N} \lambda_{i} \sum_{j=1}^{N} \frac{im \sqrt{\lambda_{i} \lambda_{j}}}{\hbar^{2}(k_{r} - ik_{I})} \left[e^{i(k_{r} - ik_{I}) | x - a_{i} |} \Phi^{-1}(z_{R}) \right]_{ij} e^{i(k_{r} - ik_{I})a_{j}}.$$

A similar result can be obtained for the Gamow wave function $\psi^{-}(x)$. In 189 principle, both Gamow functions will be equally suitable to play the role of wave 190 function for the resonance state. The only technical difference is that one represents 191 the time reversal of the other [30]. Observe that $\psi_{k}^{+}(x) \mapsto \infty$ as $x \mapsto \infty$. 192 Gamow wave functions cannot be normalised in the usual sense of square integrable 193 normalisation, but in sharp contrast with the plane waves (Dirac kets) which are not 194 normalisable although bounded, Gamow functions show an exponential growing at 195 the spatial infinite. This behaviour has been often called the exponential catastrophe. 196 This is not such a problem with a proper interpretation of the Gamow wave function 197 in terms of generalised functions in a suitable rigged Hilbert space. Still, this exponential behaviour creates some particular problems such as the difficulties arisen in 199 order to fix a proper definition of averages of observables in Gamow states [34, 35]. 200

3 One Dimensional Salpeter Hamiltonian with N Deltas

The one dimensional Salpeter Hamiltonian decorated with N Dirac deltas has the $_{202}$ following form (c = 1): $_{203}$

$$H := \sqrt{p^2 + m^2} - \sum_{i=1}^N \lambda_i \,\delta(x - a_i) \,, \qquad H_0 := \sqrt{p^2 + m^2} \,. \tag{29}$$

Here, H_0 is the free Salpeter Hamiltonian. The definition of a self-adjoint version 204 for H in (29) is not as simple as is in the Schrödinger case, where it is sufficient to 205 impose correct matching conditions at the nodes. This self-adjoint version is usually 206 determined by a proper choice of the resolvent operator of H, which should be 207 obtained from the resolvent operator of H_0 by the Krein formula. However, this 208 procedure leads to divergences in our case, so that a regularisation procedure is 209 in order here [5, 6]. We have chosen heat kernel regularisation for several reasons 210 discussed in [6]. Let us sketch briefly the procedure. First of all, we write the 211 Hamiltonian H as in (29) as 212

$$H = \sqrt{p^2 + m^2} - \sum_{i=1}^{N} \lambda_i |a_i\rangle \langle a_i|, \qquad (30)$$

exactly as we did for the cases studied in the previous section. The next step is to 213 write an ϵ -regularised version of (30) as 214

$$H_{\epsilon} = \sqrt{p^2 + m^2} - \sum_{i=1}^{N} \lambda_i(\epsilon) |a_i^{\epsilon}\rangle \langle a_i^{\epsilon}|, \qquad (31)$$

where the new kets $|a_i^{\epsilon}\rangle$ are defined in such a way that $\langle x|a_i^{\epsilon}\rangle := K_{\epsilon/2}(x, a_i)$, where 215 the function $K_t(x, y)$ is the so-called heat kernel, which is the fundamental solution 216 of the heat equation of the form: 217

$$\sqrt{p^2 + m^2} K_t(x, y) = -\frac{\partial K_t(x, y)}{\partial t}, \qquad (32)$$

and the weights $\lambda(\epsilon)$ are also chosen as functions of the parameter ϵ , such that 218 $\lim_{\epsilon \to 0^+} \lambda_i(\epsilon) \mapsto \lambda_i, i = 1, 2, ..., N$. The interest of this choice for $\langle x | a_i^{\epsilon} \rangle$ comes 219 after the limiting property $\langle x | a_i^{\epsilon} \rangle \mapsto \langle x | a_i \rangle = \delta(x - a_i)$ as $\epsilon \mapsto 0^+$. 220

Now, we go back to the Lippmann–Schwinger equation (1), where in the present 221 case $E_k = \sqrt{p^2 + m^2}$ and V is as in (31). This gives 222

$$|k^{\pm}(\epsilon)\rangle = |k\rangle + \sum_{j=1}^{N} \lambda_j(\epsilon) R_0(E_k \pm i0) |a_j^{\epsilon}\rangle \langle a_j^{\epsilon}|k^{\pm}\rangle.$$
(33)

Let us choose the plus sign in (33) and use for brevity the following notation: 223 $|f_i^{\epsilon}\rangle := \sqrt{\lambda_i(\epsilon)} |a_i^{\epsilon}\rangle$. Then, we choose one subindex *i* and isolate the corresponding 224 term in (33): 225

$$|k^{+}(\epsilon)\rangle = |k\rangle + R_{0}(E_{k} + i0) |f_{i}^{\epsilon}\rangle \langle f_{i}^{\epsilon}|k^{+}(\epsilon)\rangle + \sum_{j\neq i}^{N} R_{0}(E_{k} + i0) |f_{j}^{\epsilon}\rangle \langle f_{j}^{\epsilon}|k^{+}(\epsilon)\rangle, \qquad (34)$$

before multiplying (34) to the left by the ket $\langle f_i^{\epsilon} |$. This gives

$$\begin{bmatrix} 1 - \langle f_i^{\epsilon} | R_0(E_k + i0) | f_i^{\epsilon} \rangle \end{bmatrix} \langle f_i^{\epsilon} | k^+(\epsilon) \rangle$$
$$- \sum_{i \neq j}^{N} \left[\langle f_i^{\epsilon} | R_0(E_k + i0) | f_i^{\epsilon} \rangle \right] \langle f_i^{\epsilon} | k^+(\epsilon) \rangle = \langle f_i^{\epsilon} | k \rangle , \qquad (35)$$

expression valid for i = 1, 2, ..., N. This may be written in the matrix form as 227

$$\sum_{j=1}^{N} T_{ij}(\epsilon, E_k + i0) \langle f_j^{\epsilon} | k^+(\epsilon) \rangle = \langle f_i^{\epsilon} | k \rangle, \qquad j = 1, 2, \dots, N,$$
(36)

with

$$T_{ij}(\epsilon, E_k + i0) = \begin{cases} 1 - \langle f_i^{\epsilon} | R_0(E_k + i0) | f_i^{\epsilon} \rangle & \text{if } i = j, \\ - \langle f_i^{\epsilon} | R_0(E_k + i0) | f_j^{\epsilon} \rangle & \text{if } i \neq j. \end{cases}$$
(37)

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Therefore, we may write the solution of (36) as

$$\left\langle f_{i}^{\epsilon} | k^{+}(\epsilon) \right\rangle = \sum_{j=1}^{N} \left[T^{-1} \left(\epsilon, E_{k} + i0 \right) \right]_{ij} \left\langle f_{j}^{\epsilon} | k \right\rangle.$$
(38)

We use (38) in (35) and, then, multiply the result to the left by the bra $\langle x |$. This gives 230

$$\psi^{+}(\epsilon, x) := \langle x|k^{+}(\epsilon) \rangle$$

$$= \langle x|k \rangle + \sum_{i,j=1}^{N} \langle x| R_{0}(E_{k} + i0) | f_{i}^{\epsilon} \rangle [T^{-1}(\epsilon, E_{k} + i0)]_{ij} \langle f_{j}^{\epsilon} | k \rangle$$

$$= e^{ikx} + \sum_{i,j=1}^{N} \langle x| R_{0}(E_{k} + i0) | a_{i}^{\epsilon} \rangle [\Phi^{-1}(\epsilon, E_{k} + i0)]_{ij} \langle a_{j}^{\epsilon} | k \rangle, \quad (39)$$

with

$$\Phi_{ij}(\epsilon, E_k + i0) = \begin{cases} \frac{1}{\lambda_i(\epsilon)} - \langle a_i^{\epsilon} | R_0(E_k + i0) | a_i^{\epsilon} \rangle & \text{if } i = j, \\ -\langle a_i^{\epsilon} | R_0(E_k + i0) | a_j^{\epsilon} \rangle & \text{if } i \neq j. \end{cases}$$
(40)

The next step is to take the limit $\epsilon \mapsto 0$, for which we need a determination of 232 the functions $\lambda_i(\epsilon)$ for all values of i = 1, 2, ..., N. This has been motivated and 233 determined in Section II in [6] and is 234

$$\frac{1}{\lambda_i(\epsilon)} = \frac{1}{\lambda_i(M_i)} + \int_0^\infty dt \, K_{t+\epsilon}(a_i, a_i) \, e^{tM_i} \,, \tag{41}$$

where $K_t(x, y)$ is the heat kernel and M_i is an unphysical renormalisation scale that ²³⁵ is chosen to be the energy of the bound state E_B^i corresponding to the bound state ²³⁶ of the *i*-th delta [6]. This gives in the limit $\epsilon \mapsto 0$, ²³⁷

$$\psi_k^+(x) = e^{ikx} + \sum_{i,j=1}^N \langle x | R_0(E_k + i0) | a_i \rangle \left[\Phi^{-1}(E_k + i0) \right]_{ij} e^{ika_j} .$$
(42)

Here,

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$$\langle x | R_0(E_k + i0) | a_i \rangle = \frac{i\sqrt{k^2 + m^2}}{k} e^{ik|x - a_i|} + \frac{1}{\pi} \int_m^\infty d\mu \, e^{-\mu|x - a_i|} \, \frac{\sqrt{\mu^2 - m^2}}{\mu^2 + k^2}$$

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2	3	8

and

.

$$\begin{aligned}
\varphi_{ij}(E_k + i0) &= \\
&= \begin{cases} -\frac{1}{\lambda(E_{\lambda}, E_B^i)} - \frac{iE_k}{\sqrt{E_k^2 - m^2}} & \text{if } i = j, \\
-\frac{iE_k}{\sqrt{E_k^2 - m^2}} e^{i\sqrt{E_k^2 - m^2} |x - a_j|} - \frac{1}{\pi} \int_m^\infty d\mu \, e^{-\mu |x - a_i|} \frac{\sqrt{\mu^2 - m^2}}{\mu^2 + E_k^2 - m^2} & \text{if } i \neq j,
\end{aligned}$$

where

$$\frac{1}{\lambda \left(E_{\lambda}, E_{B}^{i}\right)} = \frac{-E_{k}}{\pi \sqrt{E_{k}^{2} - m^{2}}} \operatorname{arctanh}\left(\frac{\sqrt{E_{k}^{2} - m^{2}}}{E_{k}}\right) - \frac{E_{B}^{i}}{\pi \sqrt{m^{2} - \left(E_{B}^{i}\right)^{2}}} \left(\frac{\pi}{2} + \arcsin\frac{E_{B}^{i}}{m}\right),$$

where E_B^i has been defined before and $\mu := \min_i E_B^i$. The conclusion is that the ²⁴² Lippmann–Schwinger equation gives in a rather straightforward manner the exact ²⁴³ form of the scattering states in a rather cumbersome situation as the one discussed ²⁴⁴ along the present section. Explicit expressions for transmission and reflection ²⁴⁵ coefficients can be also derived from the above expressions. ²⁴⁶

4 Concluding Remarks

The Lippmann–Schwinger equation is a useful tool that permits to obtain explicit 248 forms for the scattering states produced by some potential. When this potential 249 is a finite set of Dirac delta interactions, one may find explicit expressions for 250 these scattering states. We have shown that this is the case when perturbing the 251 free Schrödinger one dimensional and the Salpeter Hamiltonians with *N* attractive 252 deltas. In the first case, we have also shown that the Lippmann–Schwinger equation 253 gives explicit expressions for Gamow wave functions which are the wave function 254 for the purely exponential decay part of resonance states. The discussion on the 255 search for bound states for the Schrödinger case includes similar methods. 256

The one dimensional Salpeter Hamiltonian with *N* attractive deltas is much more 257 complicated as it requires of a regularisation procedure that we implement with the 258 use of the heat kernel for the pseudo-differential operator $\sqrt{-d^2/dx^2 + m^2}$. In this 259 case, we also obtain the exact form of the scattering states. 260

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