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Spectrally equivalent time-dependent double wells and unstable anharmonic oscillators

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ABSTRACT: We construct a time-dependent double well potential as an exact spectral equivalent to the explicitly time-dependent negative quartic oscillator with a timedependent mass term. Defining the unstable anharmonic oscillator Hamiltonian on a contour in the lower-half complex plane, the resulting time-dependent non-Hermitian Hamiltonian is first mapped by an exact solution of the time-dependent Dyson equation to a time-dependent Hermitian Hamiltonian defined on the real axis. When unitary transformed, scaled and Fourier transformed we obtain a time-dependent double well potential bounded from below. All tranformations are carried out npn-perturbatively so that all Hamiltonians in this process are spectrally exactly equivalent in the sense that they have identical instantaneous energy eigenvalue spectra.

1. Introduction

Anharmonic oscillators have a wide range of applications in quantum mechanics as they describe for instance delocalization and decoherence of quantum states, e.g. [1]. They also occur naturally in relativistic models, e.g. [2]. From a mathematical point of view their nonlinear nature make them ideal testing grounds for various approximation methods, such as perturbative approaches [3]. Based on a perturbative expansion of the energy eigenvalues it was shown in [4] that the quartic anharmonic oscillator with mass term is spectrally equivalent to a double well potential with linear symmetry breaking. The first hint about the fact that even the unstable quartic anharmonic oscillator posses a well defined bounded real spectrum, despite being unbounded from below on the real axis, was proved in [5, 6], where it was proven that its energy eigenvalues series is Borel summable. The spectral equivalence between an unstable anharmonic oscillator and a complex double well potential was then proven directly by Buslaev and Grecchi in [7].

Subsequently the unstable quartic anharmonic oscillator without mass term was treated in [8] as part of the general series of \mathcal{PT} -symmetric potentials $V(x) = x^2(ix)^{\varepsilon}$, i.e. $\varepsilon = 2$, where it was shown numerically that the Hamiltonians in this series have real and positive spectra for $\varepsilon \geq 2$. Applying the techniques developed in this area of non-Hermitian \mathcal{PT} -symmetric quantum mechanics [9, 10] Jones and Mateo [11] showed that the two Hamiltonians

$$H = p^2 - gx^4$$
, and $h = \frac{p^4}{64g} - \frac{1}{2}p + 16gx^2$, (1.1)

are spectrally equivalent. This was established by first defining H on a suitable contour in the complex plane, $x \to -2i\sqrt{1+ix}$, within the Stoke wedges where the corresponding wavefunctions decay asymptotically. Subsequently the resulting complex Hamiltonian was similarity transformed to a Hermitian Hamiltonian h that is well defined on the real axis.

Here our central aim is to extend the analysis by making the Hamiltonian explicitly time-dependent $H \to H(t)$ through the inclusion of an explicit time-dependence into the coefficients. The similarity transformation acquires then the form

$$h(t) = \eta(t)H(t)\eta^{-1}(t) + i\partial_t \eta(t)\eta^{-1}(t), \qquad (1.2)$$

often referred to as the time-dependent Dyson equation [12, 13, 14, 15, 16, 17, 18, 19, 20], in which $H \neq H^{\dagger}$ is a non-Hermitian explicitly time-dependent Hamiltonian, $h = h^{\dagger}$ a Hermitian explicitly time-dependent Hamiltonian and $\eta(t)$ the time-dependent Dyson map. The latter can be used to define a time-dependent metric $\rho(t)$ via the relation $\rho(t) = \eta^{\dagger}(t)\eta(t)$. Spectral equivalence is then understood on the level of the instantaneous energy eigenvalues for the operators h(t) and the corresponding operator for the non-Hermitian system

$$\tilde{H}(t) = \eta^{-1}(t)h(t)\eta(t) = H(t) + i\eta^{-1}(t)\partial_t\eta(t).$$
(1.3)

Note while H is observable it is not a Hamiltonian governing the time-evolution and satisfying the time-dependent Schrödinger equation. On the other hand the Hamiltonian H(t)is not observable. Besides the aforementioned interest in the unstable anharmonic oscillator itself, there are not many known exact solutions [15, 17, 21, 18, 22, 19, 23, 24, 25, 26, 27, 28, 29, 30] to the time-dependent Dyson equation (1.2), so that any new exact solution provides valuable insights.

2. The time-dependent unstable harmonic oscillator

The Hamiltonian we investigate here is similar to the one in equation (1.1), but with time-dependent coefficient functions and an additional mass term

$$H(z,t) = p^{2} + \frac{m(t)}{4}z^{2} - \frac{g(t)}{16}z^{4}, \qquad m \in \mathbb{R}, g \in \mathbb{R}^{+}.$$
 (2.1)

Defining H(z,t) now on the same contour in the lower-half complex plane $z = -2i\sqrt{1+ix}$ as suggested by Jones and Mateo [11], it is mapped into the non-Hermitian Hamiltonian

$$H(x,t) = p^{2} - \frac{1}{2}p + \frac{i}{2}\{x,p^{2}\} - m(t)(1+ix) + g(t)(x-i)^{2}, \qquad (2.2)$$

with $\{\cdot, \cdot\}$ denoting the anti-commutator. Next we attempt to solve the time-dependent Dyson equation (1.2) to find a Hermitian counterpart h. Making the following general Ansatz for the Dyson map

$$\eta(t) = e^{\alpha(t)x} e^{\beta(t)p^3 + i\gamma(t)p^2 + i\delta(t)p}, \qquad \alpha, \beta, \gamma, \delta \in \mathbb{R},$$
(2.3)

we use the Baker-Campbell-Hausdorff formula to compute the adjoint action of $\eta(t)$ on all terms appearing in H(x,t)

$$\eta x \eta^{-1} = x + \delta + 6\alpha\beta p + 2\gamma p + 3i\alpha^2\beta + 2i\alpha\gamma - 3i\beta p^2, \qquad (2.4)$$

$$\eta p \eta^{-1} = p + i\alpha, \tag{2.5}$$

$$\eta x^{2} \eta^{-1} = x^{2} - 9\beta^{2} p^{4} - 12i\beta (3\alpha\beta + \gamma) p^{3} + (54\alpha^{2}\beta^{2} + 36\alpha\beta\gamma + 4\gamma^{2} - 6i\beta\delta) p^{2} (2.6) + 4(3\alpha\beta + \gamma) [\delta + i\alpha(3\alpha\beta + 2\gamma)] p + 2 (\delta + 3i\alpha^{2}\beta + 2i\alpha\gamma) x + (6\alpha\beta + 2\gamma) \{x, p\} - 3i\beta \{x, p^{2}\} - (3\alpha^{2}\beta + 2\alpha\gamma - i\delta)^{2}, \eta p^{2} \eta^{-1} = p^{2} - \alpha^{2} + 2i\alpha p,$$
(2.7)

$$\eta\{x, p^2\}\eta^{-1} = \{x, p^2\} - 6i\beta p^4 + (24\alpha\beta + 4\gamma)p^3 + (36i\alpha^2\beta + 12i\alpha\gamma + 2\delta)p^2 - 2\alpha^2x \quad (2.8) + 4(i\alpha\delta - 6\alpha^3\beta - 3\alpha^2\gamma)p - 2i\alpha^2(3\alpha^2\beta + 2\alpha\gamma - i\delta) + 4i\alpha\{x, p\}.$$

The gauge like terms in (1.2) and (1.3) are calculated to

$$i\dot{\eta}\eta^{-1} = ix\dot{\alpha} + i\dot{\beta}p^3 - \left(3\dot{\beta}\alpha + \dot{\gamma}\right)p^2 - \left(3i\dot{\beta}\alpha^2 + 2i\dot{\gamma}\alpha + \dot{\delta}\right)p + \dot{\beta}\alpha^3 + \dot{\gamma}\alpha^2 - i\dot{\delta}\alpha, \quad (2.9)$$
$$i\eta^{-1}\dot{\eta} = ix\dot{\alpha} + i\dot{\beta}p^3 - (3\dot{\alpha}\beta + \dot{\gamma})p^2 - (2i\gamma\dot{\alpha} + \dot{\delta})p - i\delta\dot{\alpha}, \quad (2.10)$$

where as commonly used we abbreviate partial derivatives with respect to t by an overdot. Using the expressions in (2.4)-(2.9) for the evaluation of (1.2) and demanding the right hand side to be Hermitian yields the following constraints for the coefficient functions in the Dyson map

$$\alpha = \frac{\dot{g}}{6g}, \quad \beta = \frac{1}{6g}, \quad \gamma = \frac{12g^3 + 6mg^2 + \dot{g}^2 - g\ddot{g}}{4\dot{g}g^2}, \quad \delta = c_1 \frac{g}{\dot{g}} - \frac{g\ln g}{2\dot{g}}, \tag{2.11}$$

with $c_1 \in \mathbb{R}$ being an integration constant. Moreover, the time-dependent coefficient functions in the Hamiltonian (2.1) must be related by the third order differential equation

$$9g^{2}(\ddot{g} - 6g\dot{m}) + 36g\dot{g}(gm - \ddot{g}) + 28\dot{g}^{3} = 0.$$
(2.12)

Integrating once and introducing a new parameterization function $\sigma(t)$, we solve this equation by

$$g = \frac{1}{4\sigma^3}$$
, and $m = \frac{4c_2 + \dot{\sigma}^2 - 2\sigma\ddot{\sigma}}{4\sigma^2}$, (2.13)

with $c_2 \in \mathbb{R}$ denoting the integration constant corresponding to the only integration we have carried out. The time-dependent Hermitian Hamiltonian in equation (1.2) then results to

$$h(x,t) = \sigma^3 p^4 + f_{pp}(t)p^2 + f_x(t)x + f_p(t)p + f_{xp}(t)\{x,p\} + f_{xx}(t)x^2 + C(t).$$
(2.14)

with

$$f_{pp} = \frac{\sigma \left\{ \sigma \left[2 \left(\sigma \left(\dot{\sigma}^2 - 4c_2 \right) - 2 \right) \ddot{\sigma} + 16c_2^2 + \dot{\sigma}^4 \right] + 16c_2 \right\} + 4}{4\sigma \dot{\sigma}^2}, \quad f_{xp} = \frac{\left(\sigma \left(\dot{\sigma}^2 - 4c_2 \right) - 2 \right)}{4\sigma^2 \dot{\sigma}}, \\ f_p = \frac{2c_1 \left[\sigma \left(4c_2 + \dot{\sigma}^2 - 2\sigma \ddot{\sigma} \right) + 2 \right] + \ln \left(4\sigma^3 \right)}{12\sigma \dot{\sigma}^2}, \quad f_x = -\frac{2c_1 + \ln \left(4\sigma^3 \right)}{12\sigma^2 \dot{\sigma}}, \quad f_{xx} = \frac{1}{4\sigma^3}, \\ C = \frac{\left(2c_1 + \ln \left(4\sigma^3 \right) \right)^2 + 36\dot{\sigma}^2 \left(4c_2^2 + \ddot{\sigma} \right)}{144\sigma \dot{\sigma}^2} + \frac{1}{8} \left(\dot{\sigma}^2 - 4c_2 \right) \ddot{\sigma} - \frac{\dot{\sigma}^2}{4\sigma^2} \end{cases}$$

We may choose to set $c_1 = c_2 = 0$ and reintroduce the original time-dependent coefficient functions g(t), m(t) so that the Hamiltonian simplifies to

$$h(x,t) = \frac{p^4}{4g} + \left(\frac{18g^2(2g+m)}{\dot{g}^2} + \frac{\dot{g}^2}{72g^3} - \frac{2g+m}{4g}\right)p^2 - \frac{3\left(g^2m+g^3\right)\ln g}{\dot{g}^2}p + \frac{g^2\ln(g)}{\dot{g}}x + \left(\frac{\dot{g}}{12g} - \frac{6g^2}{\dot{g}}\right)\{x,p\} + gx^2 + \frac{1296g^8\ln^2 g + \dot{g}^6 - 36\dot{g}^4g^2(2g+m)}{5184g^5\dot{g}^2} - \frac{m}{2}.$$
 (2.15)

Notice that $\sigma(t)$ can be any function, but the coefficient functions g(t) and m(t) must be related by (2.12) that is (2.13).

The massless case for m(t) = 0 is more restrictive and leads to $\sigma(t)$ being a second order polynomial $\sigma(t) = \kappa_0 + \kappa_1 t + \kappa_2 t^2$ with real constants κ_i . This case is consistently recovered from (2.13) with the choice $c_2 = \kappa_1 \kappa_3 - \kappa_2^2/4$. The solution found for the timeindependent case in [11], would be obtained from (2.3) in the limits $\alpha \to 0$, $\beta \to 1/6g$, $\gamma \to 0$, $\delta \to i$ and $m \to 0$. While this limit obviously exists for α and β , the constraints for γ and δ are different from those reported in (2.11). In fact, setting $\delta(t) \to i\delta(t)$ enforces g to be time-independent and there is no time-dependent solution corresponding to that choice. The energy operator \tilde{H} defined in (1.3) is obtained directly by adding H(x, t) in (2.2) and the gauge-like term in (2.10).

Let us now eliminate the terms in h(x, t) proportionate to x and $\{x, p\}$ by means of a unitary transformation

$$U = e^{-i\frac{f_{xp}}{2f_{xx}}p^2 - i\frac{f_x}{2f_{xx}}p},$$
(2.16)

which leads to the unitary transformed Hamiltonian

$$\hat{h}(x,t) = \sigma^3 p^4 + \left(f_{pp} - \frac{f_{xp}^2}{f_{xx}} \right) p^2 + \left(f_p - \frac{f_x f_{xp}}{f_{xx}} \right) p + f_{xx} x^2 + C - \frac{f_x^2}{4f_{xx}}.$$
(2.17)

Similarly as in the time-independent case [11], we may scale this Hamiltonian, albeit now with a time-dependent function, $x \to (f_{xx})^{-1/2}x$. Subsequently we Fourier transform $\hat{h}(x,t)$ so that it is viewed in momentum space. In this way we obtain a spectrally equivalent Hamiltonian with a time-dependent potential

$$\tilde{h}(y,t) = p_y^2 + \sigma^3 f_{xx}^2 y^4 + \left(f_{xx} f_{pp} - f_{xp}^2\right) y^2 + \left(\sqrt{f_{xx}} f_p - \frac{f_x f_{xp}}{\sqrt{f_{xx}}}\right) y + C - \frac{f_x^2}{4f_{xx}}, (2.18)$$

$$= \frac{g}{4} y^2 \left(y^2 + \frac{\dot{g}^2}{36g^3} + \frac{72g^2m}{\dot{g}^2} - \frac{m}{g} + 2\right) + \frac{\left(36g^2m + \dot{g}^2\right)\sqrt{g}\ln g}{12\dot{g}^2} y \qquad (2.19)$$

$$+ \frac{\dot{g}^4}{5184g^5} - \frac{\dot{g}^2m}{144g^3} - \frac{\dot{g}^2}{72g^2} - \frac{m}{2},$$

where for simplicity we have set $c_1 = c_2 = 0$ in (2.19). The potential in h(y,t) is a double well that is bounded from below. We illustrate this for a specific choice of $\sigma(t)$, that is g(t) and m(t), in figure 1.

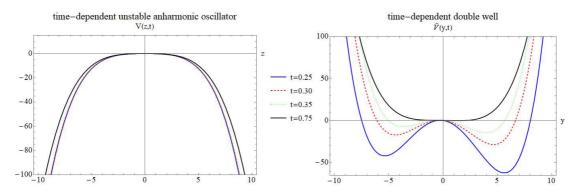


Figure 1: Spectrally equivalent time-dependent anharmonic oscillator potential V(z,t) in (2.1) and time-dependent double well potential $\tilde{V}(y,t)$ in (2.19) for $\sigma(t) = \cosh t$, $g(t) = 1/4 \cosh^3 t$, $m(t) = (\tanh^2 t - 2)/4$ at different values of time.

3. Conclusions

We have proven the remarkable fact that the time-dependent unstable anharmonic oscillator is spectrally equivalent to a time-dependent double well potential that is bounded from below. The transformations we carried out are summarized as follows:

$$H(z,t) \stackrel{z \to x}{\to} H(x,t) \stackrel{\text{Dyson}}{\to} h(x,t) \stackrel{\text{unitary transform}}{\to} \hat{h}(x,t) \stackrel{\text{Fourier}}{\to} \tilde{h}(y,t).$$

We have first transformed the time-dependent anharmonic oscillator H(z,t) from a complex contour in a Stokes wedge to the real axis H(x,t). The resulting non-Hermitian Hamiltonian H(x,t) was then mapped by mean of a time-dependent Dyson map $\eta(t)$ to a time-dependent Hermitian Hamiltonian h(x,t). It turned out that the Dyson map can not be obtained by simply introducing time-dependence into the known solution for the time-independent case [11], but it required to complexify one of the constants and the inclusion of two additional factors. In order to obtain a potential Hamiltonian we have unitary transformed h(x,t) into a spectrally equivalent Hamiltonian $\hat{h}(x,t)$, which when Fourier transformed leads to $\tilde{h}(y,t)$ that involved a time-dependent double well potential.

A detailed analysis of the spectra and eigenfunctions using approximation methods for time-dependent potential [31] is left for future investigations. Moreover, it is well known that Dyson maps are not unique, in the time-dependent as well as time-independent case, and it might therefore be interesting to explore whether additional spectrally equivalent Hamiltonians to H(z,t) can be found in the same fashion for new type of maps.

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