Stratospheric ozone changes from explosive tropical volcanoes: Modelling and ice core constraints

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Key Points:

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12	•	The tropical volcanic eruption in the model shows that the sign of the ozone
13		change is highly sensitive to stratospheric chlorine amounts.
14	•	$\delta^{15} \mathrm{N}(\mathrm{NO}_3^-)$ (a proxy for surface ultra-violet radiation) from the Samalas 1257
15		eruption is obscured by inter-annual variability in the ice core.
16	•	Any $\delta^{15}N(NO_3^-)$ signal from the Samalas eruption will be below 60 per mil. It
17		is unlikely that prolonged complete ozone removal occurred.

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18 Abstract

Major tropical volcanic eruptions have emitted large quantities of stratospheric sul-19 phate and are potential sources of stratospheric chlorine although this is less well 20 constrained by observations. This study combines model and ice core analysis to in-21 vestigate past changes in total column ozone. Historic eruptions are good analogues for 22 future eruptions as stratospheric chlorine levels have been decreasing since the year 23 2000. We perturb the pre-industrial atmosphere of a chemistry-climate model with 24 high and low emissions of sulphate and chlorine. The sign of the resulting Antarctic 25 ozone change is highly sensitive to the background stratospheric chlorine loading. In 26 the first year, the response is dynamical, with ozone increases over Antarctica. In the 27 high HCl (2 Tg emission) experiment, the injected chlorine is slowly transported to the 28 polar regions with subsequent chemical ozone depletion. These model results are then 29 compared to measurements of the stable nitrogen isotopic ratio, $\delta^{15}N(NO_3^-)$, from a 30 low snow accumulation Antarctic ice core from Dronning Maud Land (recovered in 31 2016-17). We expect ozone depletion to lead to increased surface ultraviolet (UV) 32 radiation, enhanced air-snow nitrate photo-chemistry and enrichment in $\delta^{15}N(NO_3^-)$ 33 in the ice core. We focus on the possible ozone depletion event that followed the 34 largest volcanic eruption in the past 1000 years, Samalas in 1257. The characteris-35 tic sulphate signal from this volcano is present in the ice-core but the variability in 36 $\delta^{15} N(NO_3^-)$ dominates any signal arising from changes in UV from ozone depletion. 37 Prolonged complete ozone removal following this eruption is unlikely to have occurred 38 over Antarctica. 30

⁴⁰ Plain Language Summary

Chlorine in the stratosphere destroys ozone that protects the Earth from harmful 41 ultraviolet radiation. Volcanic eruptions in the tropics can emit sulphate and chlorine 42 into the stratosphere. Chlorine levels are currently decreasing and to understand the 43 impact of a volcanic eruption on stratospheric ozone in a future climate, historical 44 eruptions are a useful analogue since the pre-industrial climate also had low chlorine 45 levels. Using a chemistry climate model, we run a set of experiments where we inject 46 different amounts of sulphate and chlorine into the stratosphere over the tropics to 47 simulate different types and strengths of explosive volcanoes and we find that the 48 ozone over Antarctica initially increases over the first year following the eruption. If 49 the volcano emits a large amount of chlorine, ozone then decreases over Antarctica in 50 years two to four following the eruption. We also compare our results to ice-core data 51 around a large historic volcanic eruption, Samalas (1257). 52

⁵³ 1 Introduction

The ozone layer protects life on Earth from ultraviolet (UV) radiation. Explo-54 sive tropical volcanic eruptions can inject volcanic gases into the stratosphere which 55 can disrupt the complex stratospheric chemistry and lead to substantial changes in 56 total column ozone (Solomon, 1999; Robock & Oppenheimer, 2003, for a comprehen-57 sive review). Over the last 1000 years, a number of explosive tropical volcanoes have 58 injected copious volumes of sulphur dioxide (SO₂) and hydrochloric acid (HCl) into 59 the stratosphere. Injection of sulphur dioxide into the stratosphere from an explosive 60 volcanic eruption increases the number of sulphate aerosol particles, providing a larger 61 surface area for heterogeneous reaction to take place on. The impact of this change 62 in stratospheric aerosol loading on ozone is dependent on the stratospheric chlorine 63 loading (e.g., Timmreck, 2012). In a low chlorine atmosphere, N_2O_5 reacts with water vapour on the surfaces of these volcanic aerosols to form HNO₃, effectively seques-65 tering reactive NO_x species into a long-lived reservoir and limiting the availability 66 of NO_x radicals to take part in catalytic reactions which deplete stratospheric ozone 67

(Crutzen, 1970; Johnston, 1971). However, in a high chlorine atmosphere, while the 68 heterogeneous reaction of N_2O_5 on aerosol surfaces has the same effect, halogenated 69 reservoir species also undergo heterogeneous reactions, liberating reactive ClO_x and 70 BrO_x species from long-lived reservoirs (e.g., Solomon, 1999). As a result, a large 71 sulphur dioxide injection is expected to cause polar ozone loss when the chlorine load-72 ing of the stratosphere is high (e.g. Tie & Brasseur, 1995), while in a low chlorine 73 environment, such as a pre-industrial atmosphere or a future atmosphere where the 74 chlorine loading of the stratosphere has declined, it is widely accepted that an injec-75 tion of sulphate from an explosive tropical volcanic eruption will lead to ozone gain 76 over polar regions (Langematz et al., 2018, and references therein). To understand 77 the future atmospheric impact of volcanic eruptions, studying historic eruptions is a 78 useful analog. 79

Estimates of the amount of sulphur dioxide emitted into the stratosphere from 80 eruptions over the past 1000 years are highly variable. For example, sulfate mass 81 concentration records from ice core data give the following estimates for recent tropical 82 eruptions: ~ 10 to 20 Tg SO_2 from Mount Pinatubo in 1991 (Timmreck et al., 2018), \sim 83 60 Tg SO_2 from Mount Tambora in 1815 (Zanchettin et al., 2016) and ~ 100 to 140 Tg 84 SO₂ from the Samalas 1257 series of eruptions (1257, 8.4° S, 116.5°E) (Toohey & Sigl, 85 2017). Samalas is the largest eruption over the last 1000 years and part of a series of 86 4 large eruptions occurring over a period of about 26 years. 87

Some types of explosive volcanoes also emit chlorine and other halogen com-88 pounds. Volcanic stratospheric chlorine emissions are important for ozone destruction 89 reactions (Kutterolf et al., 2013) but are less well constrained, since the highly soluble 90 HCl is scavenged by processes in the volcanic plume (Halmer et al., 2002). In the 91 stratosphere, HCl is the dominant chlorine reservoir species and a source of reactive 92 halogen such as chlorine monoxide, ClO, that destroys ozone. A sophisticated plume 93 model (Textor et al., 2003) suggest that 10% to 20% of the HCl emitted would enter 94 the stratosphere and recent satellite observations have detected HCl injection into the 95 stratosphere from explosive volcanoes (Theys et al., 2014). Geo-chemical evidence by 96 Vidal et al. (2016) suggests that the Samalas eruption could have injected as much as 97 $\sim 230 \,\mathrm{Tg}$ HCl into the atmosphere. In contrast, observations during the 1991 Pinatubo 98 eruption show that the efficiency of the scavenging is highly dependent on atmospheric 99 conditions with barely detectable increases in stratospheric HCl following the eruption 100 (Wallace & Livingston, 1992). Volcanic HCl emissions and the fraction of HCl mass 101 entering the stratosphere are hence highly variable as these depend on the geochem-102 istry of the eruption and the efficiency of the scavenging processes respectively. The 103 type and location of the eruption also play a role. 104

The impact of an explosive eruption on stratospheric ozone also depends on dy-105 namical processes. Variability arising from the El Niño-Southern Oscillation (ENSO), 106 the quasi-biennial oscillation (QBO) and the variability in the Brewer-Dobson circula-107 tion are able to affect the ozone response following the eruption (Lehner et al., 2016; 108 Telford et al., 2009). In addition, the radiative heating from the aerosol injection 109 and associated changes to the planetary wave flux from the troposphere are able to 110 alter the stratospheric circulation and hence the transport of aerosols and trace gases 111 (Poberaj et al., 2011). Since the precise time of the year of the historic eruption is 112 often not known, these factors have to be taken into account in the model simulations 113 (Stevenson et al., 2017). 114

Ground-based observations of total column ozone (TCO) commenced in the 1920s and captured the severe decline in the ozone layer resulting from anthropogenic production of long-lived ozone destroying-halocarbons (e.g., Harris et al., 2015, and references therein). However, beyond the relatively short instrumental period, records of total column ozone are non-existent and thus paleo-reconstructions are required to understand how natural phenomena, such as volcanic eruptions, can impact the variability
 of total column ozone.

Recent research has focused on novel Antarctic ice core proxies of surface UV 122 radiation, which can provide constraints on past ozone variability as changes in total 123 column ozone affect the surface UV over Antarctica. The UV proxy is based on 124 the stable isotopic composition of nitrate $(\delta^{15}N(NO_3^-))$ at low accumulation sites in 125 Antarctica (Frey et al., 2009). Theory, laboratory and field experiments have shown 126 that nitrate (NO_3^-) loss from snow and associated isotopic enrichment of $\delta^{15}N(NO_3^-)$ 127 in the NO_3^- fraction remaining in the snow is driven by UV photolysis (Shi et al., 128 2019; Berhanu et al., 2014, 2015; Frey et al., 2009). The presence of the heavier 129 isotope of nitrogen, ¹⁵N, in NO₃⁻ leads to an increase in reduced mass which causes a 130 red shift in the vibrational frequencies and a reduction in zero point energy. The UV 131 absorption peak of the heavier isotope is then narrower and blue shifted resulting in 132 a difference in isotopic fractionation. Further details of this process can be found in 133 Frey et al. (2009). The photolytically-induced fractionation of the $\delta^{15}N(NO_3^-)$ signal 134 is eventually archived in firn and ice. This depends on a number of site-specific factors 135 aside from the UV irradiance including the snow physical properties and the amount 136 and timing of snow accumulation (Erbland et al., 2015, 2013; Noro et al., 2018; Shi 137 et al., 2018). The largest enrichment of $\delta^{15}N(NO_3^-)$ is observed at low accumulation 138 sites on the East Antarctic Plateau, where near surface snow is exposed for more than 139 one summer season to incoming UV radiation (Erbland et al., 2013; Shi et al., 2018). 140

Winton et al. (2019) carried out a comprehensive field and modelling study of the 141 air-snow transfer of NO_3^- at the low snowfall accumulation site at Kohnnen Station 142 in Dronning Maud Land (DML), East Antarctica as part of the ISOL-ICE (ISotopic 143 constraints of past Ozone Layer in polar ICE) project. At the DML site, NO_3^- is 144 recycled three times before it is archived in the snowpack below a depth of 15 cm 145 and within 0.75 years. Sensitivity analysis with a 1D air-snow model, TRANSITS 146 (TRansfer of Atmospheric Nitrate Stable Isotopes To the Snow) (Erbland et al., 2015), 147 of $\delta^{15}N(NO_3^-)$ at DML showed that the dominant factors controlling the archived 148 $\delta^{15} N(NO_3^-)$ signature are the snow accumulation rate and e-folding depth of the surface 149 snowpack for incident UV, with a smaller role from changes in the snowfall timing 150 and TCO. The Winton et al. (2019) study sets the framework for the interpretation 151 of a $\delta^{15}N(NO_3^-)$ record from the new ISOL-ICE ice core drilled in January 2017 at 152 Kohnen Station in Dronning Maud Land, henceforth referred to as the DML site, 153 following the terminology in Winton et al. (2019). The DML region experiences low 154 annual accumulation rates ($< 10 \text{ g cm}^{-2} \text{ yr}^{-1}$) but ice cores from the area still record 155 seasonal, centennial and millennial scale variability in glaciochemistry (Göktas et al., 156 2002; Oerter et al., 2000; Sommer et al., 2000), as well as highly-resolved tropical 157 volcanic eruptions (Hofstede et al., 2004; Severi et al., 2007). This site offers useful 158 potential to investigate the impact of volcanic eruptions on TCO, surface UV radiation 159 and its imprint in the $\delta^{15}N(NO_3^-)$ ice core signature. 160

The aim of this study is to combine modelling studies with ice core evidence 161 to understand the impact on the total column ozone of explosive tropical volcanic 162 eruptions in a low chlorine stratosphere. The first part of this study will explore the 163 sensitivity of ozone over Antarctica to different volcanic emission scenarios using a 164 state-of-the-art chemistry-climate model (UM-UKCA) with additional key heteroge-165 neous and photolysis reactions. The second part of the study examines the $\delta^{15}N(NO_3^-)$ 166 signal for the tropical volcanic eruption, Samalas. Section 2 described the methods 167 used in this paper. We provide a brief overview of the UM-UKCA chemistry-climate 168 model and the additional key heterogeneous and photolysis reactions that were added 169 to improved the representation of stratospheric ozone. A Pinatubo eruption test case 170 is used to validate the response to a present day volcanic eruption. We also provide 171 a brief description of the ice core data and the isotopic analysis. In Section 3.1, we 172

use the model to investigate the response of stratospheric ozone to various volcanic
emission scenarios in a pre-industrial atmosphere. The isotopic constraints offered on
past ozone change from the ice core evidence are presented in Section 3.2. Finally,
Section 4 combines the model results and ice core analysis to discuss the implications
for past and future ozone changes from explosive tropical volcanoes.

¹⁷⁸ 2 Data and methods

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2.1 Model description, changes

We make use of the coupled chemistry-climate model which consists of the United 180 Kingdom Chemistry and Aerosol (UKCA) module together with the UK Met Office 181 Unified Model (UM) (Walters et al., 2019; Morgenstern et al., 2009; O'Connor et al., 182 2014). The model is free running and with prescribed sea ice and sea surface tempera-183 tures. The original configuration is similar to the Atmospheric Model Intercomparison 184 Project (AMIP) simulation of UK Earth system model (UKESM) submission to the 185 Coupled Model Intercomparison Project Phase 6 (CMIP6) (Eyring et al., 2016). The 186 resolution is 1.875° longitude by 1.25° latitude with 85 vertical levels extending from 187 the surface to 85 km. The UKCA module is run with the combined stratosphere and 188 troposphere chemistry (CheST) option at version 10.9. The model has an internally 189 generated QBO and the dynamics of the stratosphere are well represented (Osprey 190 et al., 2013). The model includes the aerosol scheme, GLOMAP-mode, to simulate 191 the direct and indirect radiative effects (Mann et al., 2010). Aerosol optical proper-192 ties are computed online as the particle size distributions evolve due to micro-physical 193 processes. 194

Stratospheric ozone concentrations are determined by sets of photo-chemical re-195 actions first described by Chapman (1930) plus ozone destroying catalytic cycles in-196 volving chlorine, nitrogen, hydrogen and bromine radical species (Solomon, 1999). The 197 photolysis reactions in the model make use of rates calculated from a combination of 198 the FAST-JX scheme (Wild et al., 2000; Bian & Prather, 2002; Neu et al., 2007) 199 and look-up tables. FAST-JX wavelengths range from 177 to 850 nm over 18 bins 200 and calculates scattering for all bands (Telford et al., 2013). Above about 60 km, a 201 look-up table of photolysis rates (Lary & Pyle, 1991; Morgenstern et al., 2009) is used 202 when wavelengths below 177 nm become important. Heterogeneous reactions are also 203 important for determining stratospheric ozone concentrations in the presence of po-204 lar stratospheric clouds in the polar lower stratosphere or in the presence of sulphate 205 aerosol following explosive volcanic eruptions. Ozone depleting radicals are produced 206 by the photolysis of the products formed from halogen containing compounds react-207 ing on the surface of stratospheric aerosol such as polar stratospheric clouds. These 208 species include hydrochloric acid (HCl), chlorine nitrate ($ClONO_2$), hydrogen bromide 209 (HBr) and bromine nitrate $(BrONO_2)$. Three types of aerosol are considered by the 210 model: ice, nitric acid trihydrate and sulfate aerosol. Above a temperature of about 211 195 K, reactions occur on liquid sulfate aerosol, around 195 K to 188 K, the model 212 forms nitric acid trihydrate particles and below about 188 K, ice particles form. 213

We add 8 new heterogeneous reactions to the model involving chlorine and 214 bromine species in a similar way to Dennison et al. (2019), following the previous 215 work by (Yang et al., 2014), with the main difference being the explicit treatment of 216 the reactions of 4 additional chemical species: Cl_2 , Br_2 , $ClNO_2$ and $BrNO_2$. These 217 species are also photolysed to produce Cl and Br radicals. Reaction rates depend on 218 the probability of a gas molecule colliding irreversibly with the surface of the aerosol 219 and this is given by an uptake coefficient. We update the calculation of the uptake 220 coefficients using the same scheme as Dennison et al. (2019) with the differences listed 221 in Table A1 in the Appendix. The model data used in this paper is archived on the 222 Centre for Environmental Data Analysis (Ming, 2020). 223

224 2.2 Model validation

The changes to the stratospheric chemistry are tested by running the model for 225 30 years in a year 2000 time slice experiment using CMIP6 prescribed trace gases 226 and sea surface temperature forcings. The model is mostly able to reproduce the 227 observed total column ozone and the results are similar to those found by Dennison 228 et al. (2019) in which a more thorough discussion of the changes can be found. The 229 improved match with observed TCO resulting from our model updates is shown in 230 Figure 1(a). The spring ozone hole over Antarctica is deeper than the original model 231 232 with total column ozone values reaching about 175 DU (65 to 90° S average) in October compared to about 200 DU in the original model. These values are closer to those 233 in the ozone values from the satellite ozone dataset from the National Institute of 234 Water and Atmospheric Research – Bodeker Scientific (NIWA-BS) satellite dataset 235 (version 3.4; see http://www.bodekerscientific.com/data/total-column-ozone). 236 The ozone hole minimum in the satellite data reaches about 185 DU although this 237 happens earlier in September. The modified model still under predicts the summer 238 ozone values which take longer to recover compared to observations. This could be due 239 to the vortex breakup being delayed and is a known issue in a number of comprehensive 240 chemistry climate model (Eyring et al., 2010; Butchart et al., 2011; McLandress et al., 241 2012). Overall, our changes to the chemistry scheme lead to an ozone distribution that 242 is very similar to Dennison et al. (2019). 243

To assess the model response to a volcanic perturbation in the present atmo-244 sphere we run an experiment that simulates the eruption of Mount Pinatubo. The 245 model is first spun up using CMIP6 present day forcings, including changing trends in 246 trace gases. We then initialize four ensemble runs using the climate state taken from 247 four different years of the spun up model state. The runs use the CMIP6 trace gas 248 forcings from 1979 to 1994 with the eruption happening in 1991. Although the exact 249 climate state at the time of the Pinatubo eruption is known from observations, the 250 four ensemble runs are done so as to span over the variability arising from the QBO 251 and ENSO. This allows the Pinatubo run to be compared to the pre-industrial volcanic 252 runs in Section 3.1. The timing of historical volcanic eruptions is not well constrained 253 and we do not know the phases of the QBO and ENSO in which the eruptions oc-254 curred. The ensemble is designed to average over this variability. We simulate the 255 Pinatubo eruption as an emission of $10 \text{ Tg } \text{SO}_2$ and 0.02 Tg HCl on 1 June 1991 into 256 the stratosphere as a single vertical plume between 19 and 24 km altitude (the neutral 257 buoyancy height of the plume) at 15.1 °N and 120.2 °E. Mills et al. (2016) discuss the 258 justification for various choices of modelling parameters for Pinatubo. The aim of this 259 experiment is not to reproduce the observations after the Pinatubo eruption exactly 260 but to check that, with the additional chemical reactions and emissions, our model 261 is still able to simulate the broad pattern of the ozone response after a current day 262 explosive volcano. 263

Figure 1(b) shows change in total column ozone from the Pinatubo eruption in 264 the NIWA-Bodeker dataset as the difference between a 1991 to 1994 average and a 265 climatology taken from 1979 to 1990. Similarly, the same change in the model runs 266 is shown in Figure 1(c) but using the average of the four ensemble runs from 1991 267 to 1994 and a climatology taken from 1979 to 1990. A non parametric permutation 268 test is used to determine if the changes seen are larger than the natural variability; 269 changes below the level of the noise is represented by the grey fog which is plotted as 270 overlaid contours at confidence levels of 95, 90, 80, 70 and 60%. The same test is used 271 in all subsequent model plots. The red triangle marks the volcanic eruption in this 272 and subsequent plots. 273

The initial, low latitude, increase in total column ozone south of the volcano in the year following the eruption and the decrease in ozone in Jan 1992 over the North Pole are captured by the model although the changes are shorter lived than in the satellite data. Note that the Antarctic ozone hole is not as prominent a feature in model runs due to the averaging of four ensemble members. Our model ozone changes are qualitatively similar to the Pinatubo case study by Aquila et al. (2012) using a different chemistry-climate model. Aquila et al. (2012) also discuss, in more detail, the possible mechanisms for the stratospheric ozone changes. This experiment demonstrates that our modified model is able to satisfactorily stimulate the ozone changes associated with a present-day volcanic eruption.

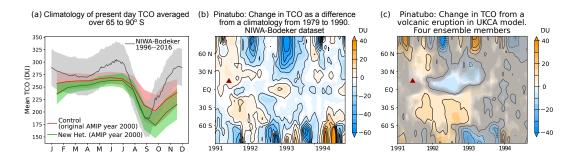


Figure 1. (a) Climatology of total column ozone (TCO) (DU) for the present climate from the NIWA-Bodeker satellite dataset (1996–2016) in black, a 30 year timeslice run of the year 2000 from the original AMIP model setup in red and the corresponding timeslice with the modified model with new heterogeneous reactions and emission files in green. Shaded bands show ± 1 standard deviation. (b) Change in Bodeker ozone following the Pinatubo eruption (red triangle) as a difference from a climatology taken from years 1979 to 1990. The QBO signal has been filtered out. (c) Change in TCO (DU) following the Pinatubo eruption (10 Tg SO₂, 0.02 Tg HCl) in the model. The plot shows the difference from a climatology taken from 1979 to 1990 and is the average of four ensemble members. The grey fog illustrates regions where the signal is below the level of the noise (see the main text for further details). The red triangle marks the volcanic eruption. Note the different colour scales between (b) and (c).

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2.3 Ice core analysis

The first high-resolution record of $\delta^{15} N(NO_3^-)$ was obtained for the last 1.3 kyr 285 from the 120 m ISOL-ICE ice core. The core was drilled in the clean air sector at 286 Kohnen Station, DML on the high-elevation East Antarctic Plateau (2892 m above 287 sea-level; 74.9961° S, 0.094717° E) in January 2017. A full description of the methods 288 for the ISOL-ICE ice core can be found in Winton et al. (2019) and only a brief 289 summary is given here. The core was analysed for i) continuous flow analysis (CFA) of 290 nitrate (NO_3) , sodium (Na) and magnesium (Mg) mass concentrations and electrolytic 291 meltwater conductivity at the British Antarctic Survey (BAS), Cambridge, and ii) 292 discrete sections for the $\delta^{15}N(NO_3^-)$ composition at the Institute of Environmental 293 Geosciences (IGE), University of Grenoble. Here we report the dated section of the 294 ice core from 1227 to 1350 AD (69.8 to 79.4 m) covering the Samalas eruption in 1257. 295 Dating was achieved by annual layer counting of measured concentrations of Na and 296 Mg following previous studies at DML (Göktas et al., 2002; Weller & Wagenbach, 297 2007; Weller et al., 2008) constrained by well-dated volcanic horizons (further details 298 can be found in Table B1). An age uncertainty of ± 3 years is estimated at the base 299 of the ice core. High-resolution sampling for $\delta^{15}N(NO_3^-)$ analysis was carried out 300 i) across volcanic horizons with a sample resolution of 5 to 30 cm, and ii) in 10 cm 301 resolution baseline samples 1 m either side of the volcanic peak. A total of 119 discrete 302 measurements of $\delta^{15}N(NO_3^-)$ are reported here. Discrete $\delta^{15}N(NO_3^-)$ samples were pre-303 concentrated and analysed using the denitrifier method following Frey et al. (2009) 304

and Morin et al. (2009). The nitrogen isotopic ratio was referenced against N₂-Air (Mariotti, 1983). We report ¹⁵N/¹⁴N of NO₃⁻ (δ^{15} N(NO₃⁻)) as δ -values: δ^{15} N(NO₃⁻) = $\begin{pmatrix} \frac{R_{\text{sample}}}{R_{\text{standard}}} - 1 \end{pmatrix}$ where R is the elemental isotopic ratio in the sample and standard respectively. The overall accuracy of the method for δ^{15} N(NO₃⁻) is 3 per mil.

309 **3 Results**

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3.1 Volcanic perturbations in model

Using the CMIP6 pre-industrial forcings, a year 1850 control run is produced. 311 The control run is 60 years long excluding 10 years of spin up which are discarded. The 312 effect from explosive volcanoes on the stratosphere is investigated by running a series of 313 four volcanic perturbation runs spun off from four different years of the control run to 314 represent the variability arising from different ENSO and QBO states in a similar way 315 to the Pinatubo case study in Section 2.2. The volcanic emissions are prescribed in a 316 similar way to the Pinatubo eruption on 1 September of the first year of the run. Since 317 historical volcanic eruptions are variable and HCl emissions are less well constrained, 318 we develop a matrix of simulations that spans the uncertainty in emissions. The six 319 sets of experiments have one of low SO_2 (10 Tg) or high SO_2 (100 Tg) paired with no 320 HCl, low HCl (0.02 Tg) and high HCl (2 Tg). Changes are plotted as the difference 321 between the average of the four perturbation runs and a climatology derived from the 322 control run. 323

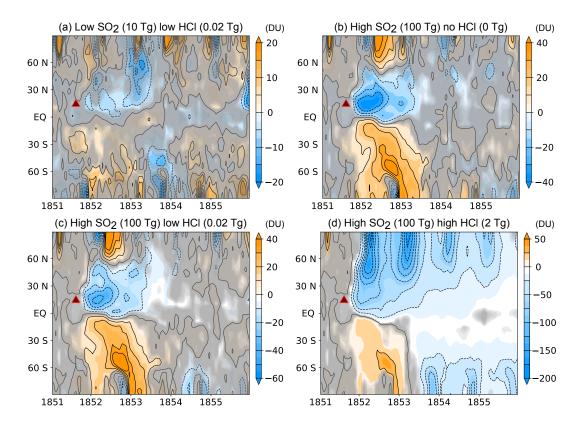


Figure 2. Change in total column ozone (DU) for the pre-industrial volcanic perturbation experiments. The plots show the difference between the average of four ensemble members and a single climatology drawn from a 60 year run. The emission scenarios shown are (a) low SO_2 , low HCl case (b) high SO_2 , no HCl (c) high SO_2 , low HCl and (d) high SO_2 , high HCl. The red triangle denotes the location of the injection. Note the different colour scales.

Figure 2 shows the change in total column ozone in the (a) low SO_2 + low HCl, 324 (b) high SO_2 + no HCl, (c) high SO_2 + low HCl and (d) high SO_2 + high HCl cases. 325 The low SO_2 + no HCl case and high SO_2 + no HCl cases are found to be qualitatively 326 similar to two further experiments (not shown): the low SO_2 + low HCl and high SO_2 327 + low HCl cases, respectively. This is expected since the stratospheric chlorine loading 328 is low (< 0.4 ppbv of HCl over the polar region averaged between 65 to 90° S), similar 329 to a pre-industrial atmosphere and we do not observe large depletion of ozone depletion 330 events by chlorine radicals during spring to form ozone holes. 331

332 The low SO_2 + low HCl case in Figure 2(a) represents the ozone response to a Pinatubo-like explosive volcano in a pre-industrial atmosphere. It shows that the 333 changes in TCO are small and dominated by internal variability in most regions. This 334 should be contrasted with the Pinatubo case study shown previously in Figure 1(c) 335 where an eruption of an equivalent magnitude was able to cause significant ozone 336 changes, including an ozone depletion of about 20 DU in the year following the eruption 337 over Antarctica. In contrast, under scenarios of low or no HCl but when the SO_2 338 emitted is high (Figures 2(b) and (c)), substantial changes in total column ozone are 339 seen for 1.5 years following the eruption. These two cases (high SO_2 and no HCl 340 case, high SO_2 and low HCl) are qualitatively similar suggesting that transport effects 341 still dominate when the amount of HCl is low in a pre-industrial atmosphere and 342 the volcanic chlorine injection is not sufficient to make a significant change to the 343 background stratospheric chlorine loading. The primary impact of a large injection of 344 SO₂ is to locally decrease TCO in the tropics and increase TCO at high latitudes, via 345 the mechanisms described below. 346

Since chemical, dynamical and radiative processes are coupled in the model, 347 it is difficult to quantify the contribution from individual processes but the results 348 suggest that the main driver of the ozone changes is dynamical in the year following 349 the eruption. The SO_2 aerosol leads to both longwave and shortwave heating in the 350 lower stratosphere (Robock, 2000) resulting in a change in the meridional circulation 351 as shown in Figure 3(a). The increased upwelling brings more ozone-poor tropospheric 352 air into the lower stratosphere leading to lower total column ozone. In contrast, the 353 decreases in upwelling outside the initial SO_2 cloud results in an increase in ozone in 354 the regions polewards of the SO_2 cloud in both hemispheres. Compared to the changes 355 in transport, the changes to the partitioning between radicals and reservoir species for 356 ClO_x , HO_x and NO_x appear to be a second order effect (not shown). The warming 357 in the lower stratosphere results in a warming of the cold point region. This region 358 controls the freeze-drying of water vapour entering the lower stratosphere and warmer 359 temperatures will result in a moistening of the stratosphere and subsequent changes 360 to HO_x chemistry. Changes in SO_2 aerosol also change the partitioning between NO_y 361 and N_2O_5 in the polar regions which can result in ozone changes but these have not 362 been quantified in this study. 363

In contrast, when a substantial amount of HCl together with SO_2 is injected 364 into the stratosphere (high SO_2 and high HCl case, Figure 2(d)), large, chemical 365 ozone depletion occurs over the polar regions during spring time in the year two to 366 four following the eruption. The initial, low latitude, response still appears to be 367 dynamical but when the injected chlorine reaches polar regions (Figure 3(b)), catalytic 368 destruction of ozone occurs in the polar vortex during spring. The mixing ratio of HCl 369 reaches values of up to 4 ppbv and 1.3 ppbv at 20 km over the North and South poles 370 respectively. These values are comparable to the present day (year 2000) values of 371 the equivalent effective stratospheric chlorine of $\sim 3 \text{ ppbv}$. The high SO₂ and high 372 HCl scenario is the one experiment where we observed prolonged ozone destruction 373 occurring over a number of years over Antarctica with a maximum decrease in total 374 column ozone of $\sim 90 \,\mathrm{DU}$ in spring of the second year after the eruption. Over the 375

North pole, stratospheric ozone is nearly completely removed in the spring for at least
 four years following the eruption.

The results are sensitive to the date, location and height of the injection in 378 the tropics. A discussion of the sensitivity of eruption source parameters on volcanic 379 radiative forcing can be found in Marshall et al. (2019). In our experiment, the lower 380 branch of the Brewer-Dobson circulation is stronger in the Northern Hemisphere in 381 September and hence the injected chlorine is primarily advected to the North pole 382 in the months following the eruption. It takes about 1.5 years for chlorine to be 383 transported to the South pole. The duration of the response to a volcanic eruption 384 is controlled by stratospheric dynamics and the material that is injected in the lower 385 stratosphere is transported to the troposphere and removed within 2 to 5 years. The 386 injected chlorine will eventually be removed from the polar stratosphere. 387

In summary, in a pre-industrial atmosphere with low chlorine levels in the strato-388 sphere, we do not detect a significant ozone response to a Pinatubo strength eruption 389 in the model. A large explosive volcano, of similar magnitude to Samalas with no 390 or low HCl produces an increase in total column ozone over Antarctica. The change 391 is short-lived (~ 2 years) and primarily driven by transport changes. In contrast, if 392 a volcanic injection of HCl (2 Tg in our experiments) is able to raise stratospheric 393 chlorine concentrations closer to present day levels, ozone depleting chemical reactions 394 will occur to produce Antarctic ozone depletion in spring as long as sufficient HCl is 395 present. The stratospheric lifetime of chlorine is determined by the age of air and the 396 strength of the stratospheric circulation. When the chlorine reaches the troposphere, it 397 is washed out, giving a lifetime of about 5 years for HCl entering in the shallow branch 398 of the Brewer-Dobson circulation. The increase in surface UV, resulting from ozone 399 depletion, will affect the $\delta^{15}N(NO_3^-)$ ratio in the snow pack. The timing of the change 400 in surface UV could lag, by a number of years, behind that of the volcanic sulphate 401 signal in ice cores, since sulphate arrives via tropospheric and stratospheric transport 402 whilst the UV signal is dependent on stratospheric ozone depletion. An additional 403 difficulty is that the timing of the arrival of the signal depends on the season of the 404 eruption; a quantity that is unknown for most volcanoes over the past 1000 years. 405

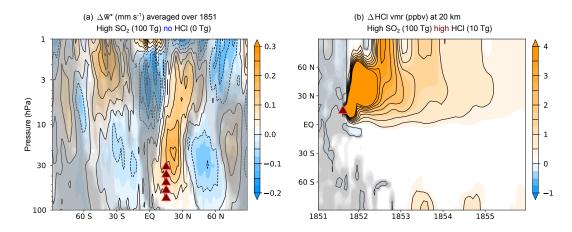


Figure 3. (a) Change in the mean residual vertical velocity for the high SO_2 and no HCl case calculated as the difference between the average over the first year following the eruption of four ensemble members and the mean of the control run to show the dynamical changes. The red triangles represent the location and vertical extent of the volcanic eruption. (b) Change in HCl volume mixing ratio (ppbv) at 20 km for the high SO_2 and high HCl case from the control run to show chemical changes.

3.2 Ice core results

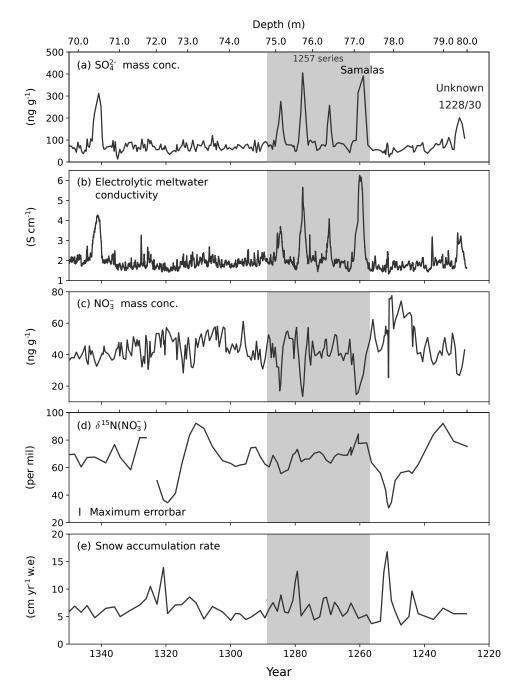


Figure 4. 1227 to 1350 AD section of the ISOL-ICE ice core data from DML, Antarctica. Age is plotted along the bottom and the corresponding ice depth along the top. The vertical grey region marks the dates around the 1257 series of volcanoes. (a) Sulphate mass concentrations. (b) Electrolytic melt water conductivity. (c) Nitrate mass concentration (d) Isotopic ratio of ${}^{15}N/{}^{14}N$ of NO₃⁻ ($\delta^{15}N(NO_3^-)$) given as δ -values. (e) Snow accumulation rate in (cm yr⁻¹ water equivalent (w.e)). Note that the various quantities are available at different time resolutions depending on the analysis method used.

We expect a tropical volcanic eruption to lead to a sulphate signal in the ice record. The previous modelling studies show that high SO₂ and high HCl eruptions can cause a decrease in TCO which would increase the UV dose reaching the surface at the ice core site. As a result, stronger photolysis would enhance NO₃⁻ loss, redistribution and recycling from snowpack, decreasing the NO₃⁻ mass concentration and enriching the $\delta^{15}N(NO_3^-)$ signature.

The ISOL-ICE ice core data from 1227 to 1350 AD is illustrated in Figure 4. The 413 ice core captures a clear signal of the 1257 Samalas series of four volcanic eruptions as 414 indicated by elevated sulphate mass concentrations and electrolytic meltwater conduc-415 tivity levels above the background in the ice core (Figures 4(a) and (b)). This pattern 416 is consistently observed in ice cores across DML and across the wider Antarctic region 417 (e.g., Hofstede et al. (2004); Göktas et al. (2002)), where sulphate originated from the 418 1257 series of eruptions, was transported via the stratosphere to Antarctica (Baroni 419 et al., 2008). Nitrate mass concentrations in the ice core decrease coincident with 420 the four large volcanic eruptions (Figure 4(c)). This observation has been reported 421 for other volcanic eruptions in Antarctic and Greenland ice cores, and is thought to 422 occur from the displacement of NO_3^- away from the highly acidic (sulphuric acid) 423 volcanic layers (Wolff, 1995; Laj et al., 1993; Legrand & Kirchner, 1990). This post-424 depositional process, unrelated to photolysis, leads to the anti-correlation between 425 the sulfate peaks and NO_3^- during the volcanic eruptions. Based on other records of 426 NO_3^- in Antarctica (Pasteris et al., 2014), we expect the NO_3^- mass concentration to 427 be correlated to the accumulation rate outside of the volcanic eruptions. We do not 428 see this positive correlation in the background variability in the ISOL-ICE ice core 429 (Figure 4(c) and (e); $R^2 = 0.04, p < 10^{-3}$ with data from five years either side of 430 the volcanic eruptions is not used). The $\delta^{15}N(NO_3^-)$ is weakly anti-correlated to the 431 accumulation rate (Figure 4(d) and (e); $R^2 = 0.2, p < 10^{-4}$ again with five years either 432 side of the volcanic eruptions removed) as would be expected from spatial transects 433 across Antarctica (Erbland et al., 2015, 2013; Noro et al., 2018; Shi et al., 2018), and 434 sensitivity tests of variable accumulation rate on the $\delta^{15}N(NO_3^-)$ signal at the DML 435 site (Winton et al., 2019). 436

The accumulation rate is variable at the DML site (2.5 to 11 cm yr^{-1} water 437 equivalent) (Oerter et al., 2000; Sommer et al., 2000) and there is no trend over the 438 last 1000 years. We speculate that changes in the accumulation rate will lead to 439 changes in e-folding depth over time which can account for part of the variability of 440 the $\delta^{15}N(NO_3^-)$ signal (Winton et al., 2019), with a smaller contribution from extreme 441 precipitation events (Turner et al., 2019). The e-folding depth of the local snowpack 442 depends on snow physical properties and contributes to the $\delta^{15}N(NO_3^-)$ signal eventu-443 ally preserved in local firm and ice (Winton et al., 2019). Unfortunately, the variability 444 of e-folding depth in the past is not known and may be a source of additional noise in 445 the $\delta^{15}N(NO_3^-)$ signal. 446

We see no enrichment of the $\delta^{15}N(NO_3^-)$ signal above the background variability 447 during the four volcanic eruptions Figure 4(d). We speculate on possible reasons for 448 thus lack of enrichment after the 1257 series. Firstly, the $\delta^{15}N(NO_3^-)$ UV proxy is 449 not sensitive enough to record TCO and concurrent surface UV changes lasting only 450 a few years. Winton et al. (2019) assessed the sensitivity of the $\delta^{15}N(NO_3^-)$ UV proxy 451 to changes in total column ozone using the TRANSITS model (Erbland et al., 2015). 452 We expect that a decrease in the total column ozone of 100 DU, estimated for a large 453 eruption on the magnitude of Samalas (assuming an eruption in September), would result in a 25 per mil increase in $\delta^{15}N(NO_3^-)$ at DML. However, this is below the inter-455 annual $\delta^{15}N(NO_3^-)$ variability of 30 to 90 per mil at DML (over the period 1227 to 1350 456 AD), and thus the development of a volcanic induced-large ozone depletion in spring 457 is unlikely to be observed above the natural background $\delta^{15}N(NO_3^-)$ variability. Note 458 that the inter-annual variability of $\delta^{15}N(NO_3^-)$ is larger than the seasonal variability of 459

about ± 25 per mil of $\delta^{15}N(NO_3^-)$ seen at the bottom of the snow pits in Winton et al. (2019). Despite DML having a relatively low snow accumulation rate, the sensitivity of the $\delta^{15}N(NO_3^-)$ UV proxy is low at this site. Secondly, although we observe a significant decrease in the NO_3^- concentration during the volcanic eruptions, we cannot rule out the possibility that the lower NO_3^- concentrations are due to migration of $NO_3^$ in acidic layers. Lastly, the impact of acidic volcanic layers on the $\delta^{15}N(NO_3^-)$ has yet to be quantified.

467 4 Discussion and conclusions

The aim of this paper is to understand the impact on the total column ozone of 468 explosive tropical volcanic eruptions in a low chlorine stratosphere and to search for 469 evidence of these changes in the ice core record over the last 1000 years. We made use 470 of the UM-UKCA chemistry-climate model, with improved heterogeneous reactions 471 and emissions, to model the evolution of ozone after different injections scenarios of 472 SO_2 and HCl representing possible past volcanic eruptions. We then compare the 473 model results to the $\delta^{15}N(NO_3^-)$ isotopic ratio from the recently obtained ISOL-ICE 474 ice core. Winton et al. (2019) and earlier work (Berhanu et al., 2015; Erbland et al., 475 2015) suggest that it may be possible to use $\delta^{15} N(NO_3^-)$ as a UV proxy for Antarctic 476 ozone changes, after accounting for accumulation rate changes. A decrease in ozone 477 leads to increased surface UV which increases the fractionation of $\delta^{15}N(NO_3^-)$ in the 478 photolytically active region of the snowpack. The resulting $\delta^{15}N(NO_3^-)$ isotopic signal, 479 which integrates the UV signal seen over several years, is then buried. We analyse the 480 $\delta^{15}N(NO_3^-)$ ice core signature around the period of the Samalas eruption to reconstruct 481 past UV changes. 482

The model experiments show that a "Pinatubo-like" eruption (low SO₂, 10 Tg 483 and low HCl, 0.02 Tg) in a pre-industrial atmosphere does not produce a significant 484 response in ozone over Antarctica (Figure 2(c)) whilst the high SO₂ (100 Tg) volcances 485 (with no or low HCl) both produce increases in ozone over Antarctica that are short-486 lived, lasting about 1.5 years (Figure 2(b) and (c)). The pattern of ozone changes 187 for the latter are primarily caused by transport changes arising from changes to the Brewer-Dobson circulation (Figure 3(a)). In contrast, when the amounts of SO₂ and 489 HCl emitted are both high (high SO₂, 100 Tg and high HCl, 2 Tg) and the HCl loading 490 over the polar regions becomes comparable to present day stratospheric values, our 491 model results show significant ozone depletion over both poles (Figure 2(d)) for at 492 least four years following the eruption. Note that the chemical reactions that destroy 493 ozone are only able to occur when HCl in the stratosphere reaches the polar regions 494 and hence the timing of the springtime ozone depletion depends strongly on the date 495 of the eruption. Since we model the eruption as occurring on 1 September, we find 496 that it takes about 1 year for the injected HCl from the volcano to reach Antarctica 497 (Figure 3(b)). Before the HCl reaches Antarctica, the increase in ozone over the 498 Southern Hemisphere is caused by the same dynamical changes as in the low or no 499 HCl model experiments. 500

Our results for the ozone response for the high SO_2 and high HCl experiments 501 are quantitatively similar to the experiments by Wade (2018) and Brenna et al. (2019)502 in terms of the ozone depletion seen over Antarctica. The experiments by Wade 503 (2018) use a coupled atmosphere and ocean with interactive chemistry (HadGEM3-ES 504 model). They explore the range of uncertainty in the emission parameters for the 505 Samalas eruption and vary SO_2 from 94.8 to 142.2 Tg, HCl from 0 to 46.68 Tg and 506 HBr from 0 to 0.263 Tg with a maximum injection height of 34 km. In their highest 507 emission scenario, they observe a decrease of 75% in the global stratospheric ozone 508 and Antarctic ozone depletion $(> 240 \,\mathrm{DU})$ for six years following the eruption. In the 509 work by Brenna et al. (2019), they impose a Central American explosive volcano in 510 a chemistry climate model (CESM1) in which the effect of sulphuric acid aerosols are 511

imposed as a modified El Chichòn surface area density forcing equivalent to 30 Tg SO₂.
The results from their experiment with 2.93 Tg Cl, 9.5 Mt Br at 14°N, 89°W with an
injection height of 29.7 hPa on January 1 (their Figure 3(c)) are qualitative similar to
our results in Figure 2(d). Brenna et al. (2019) show that the average ozone decreases
by more than 120 DU over the polar cap and observe a similar ozone increase over
Antarctica in the year after that eruption which is followed by a series of four years
with large spring-time ozone depletion.

The results above and the model experiments in this work suggest that if a trop-519 ical volcano emits a substantial amount of SO_2 and HCl (high SO_2 , 100 Tg and high 520 HCl, 2 Tg in our case), prolonged ozone depletion, lasting more than four years, will 521 occur over Antarctica. We choose to focus on the ice core record around the Samalas 522 eruption (part of the 1257 series of four volcanoes) since ice core and geochemical 523 evidence suggests that this volcano was the largest in the past 1000 years in terms of 524 SO_2 and HCl emissions although there is significant uncertainty in the amount of HCl 525 that was able to reach the stratosphere from this eruption (Halmer et al., 2002). The 526 1257 series of volcanoes that includes Samalas consists of four eruptions that occur at 527 intervals of 10, 8 and 8 years. If all four eruptions caused ozone depletion, we expect 528 to see a prolonged period of increase in $\delta^{15}N(NO_3^-)$ in the ice core. 529

Our record of the isotopic ratio of $\delta^{15}N(NO_3^-)$ in the ice core around the 1257 530 series eruptions shows that using $\delta^{15} N(NO_3^-)$ as a proxy for ozone changes is not able 531 to detect significant prolonged ozone depletion above the level of inter-annual vari-532 ability. Spatial transects across Antarctica (Noro et al., 2018, and references therein), 533 supported by air snow-photochemistry modelling (TRANSITS) (Winton et al., 2019; 534 Erbland et al., 2015) show a strong non linear dependence of $\delta^{15}N(NO_3^-)$ on snow ac-535 cumulation rate. Deeper ice core records in Greenland have observed a dependence of 536 $\delta^{15}N(NO_2^{-})$ and accumulation rate on glacial-interglacial transition timescales (Frever 537 et al., 1996). In this paper, we present the highest resolution record in ice cores 538 and find that both $\delta^{15}N(NO_3^-)$ and the accumulation rate show a substantial inter-539 annual variability (about 30 to 90 per mil and 4 to 18 cm yr⁻¹ water equivalent (w.e) 540 respectively in the ice record at 60 - 70 m depth). However, we do not observe a 541 clear relationship between the two on centennial timescales making it challenging to 542 disentangle the $\delta^{15}N(NO_3^-)$ signal from other changes such as those in accumulation 543 rate. Winton et al. (2019) show that for a 100 DU change in total column ozone (Fig-544 ure 2(d)), we expect to see a change of about 25 per mil in $\delta^{15}N(NO_3^-)$ with a roughly 545 linear relationship. This is below the level of inter-annual variability in $\delta^{15}N(NO_3^-)$ 546 seen in the ice core of about 60 per mil (the maximum uncertainty in our samples is 547 less than ± 3 per mil over this time period). The snow pack also integrates UV changes 548 over a couple of years and smooths out seasonal variability. For a $\delta^{15}N(NO_3^-)$ signal 549 to have been detected at the DML site from the 1257 eruptions, we suggest that it 550 would require a prolonged period (several years) of near complete ozone destruction, 551 for instance, during the series of seven stratospheric volcanic eruptions that occurred 552 over a deglaciation $\sim 17.7 \,\mathrm{ka}$ (McConnell et al., 2017). With the additional caveat 553 that the timing and magnitude of ozone changes depends on the season of the erup-554 tion and assuming the response scales linearly with forcing, these results suggest that 555 for a signal to be seen above the inter-annual variability, we would need prolonged near 556 complete ozone destruction (over 300 DU). Since we do not see a $\delta^{15}N(NO_3^-)$ signal 557 of this magnitude in the ice core, this provides a constraint on the magnitude of past 558 ozone changes caused by the 1257 eruptions. We conclude that that it is unlikely that 559 the Samalas 1257 series injected enough HCl to cause complete ozone removal over 560 Antarctica. 561

In summary, we have evaluated the impact of various explosive tropical volcanic emission scenarios on stratospheric ozone changes in a pre-industrial atmosphere and found that the sign of the ozone change over the polar regions depends on the amount of

Reaction	Uptake coefficient		
	Ice	Nitric acid trihydrate	Sulphate aerosol
$\hline ClONO_2 + HCl \rightarrow Cl_2 + HNO_3 \qquad \qquad$	0.3	0.3	f
$CIONO_2 + H_2O \rightarrow HOCl + HNO_3$	0.3	0.006	f
$HOCl + HCl \rightarrow Cl_2 + H_2O$	0.3	0.3	f
$N_2O_5 + H_2O \rightarrow 2 HNO_3$	0.03	0.006	0.1
$N_2O_5 + HCl \rightarrow ClNO_2 + HNO_3$	0.03	0.006	0.02
$HOBr + HCl \rightarrow BrCl + H_2O$	0.25	0.25	0.1
$BrONO_2 + HCl \rightarrow BrCl + HNO_3$	0.3	0.3	0.01
$BrONO_2 + H_2O \rightarrow HOBr + HNO_3$	0.3	0.001	0.01
$\mathrm{HOBr} + \mathrm{HBr} \rightarrow \mathrm{Br}_2 + \mathrm{H}_2\mathrm{O}$	0.25	0.25	0.1
$HOCl + HBr \rightarrow BrCl + H_2O$	0.25	0.25	0.02
$ClONO_2 + HBr \rightarrow BrCl + HNO_3$	0.56	0.56	0.02
$BrONO_2 + HBr \rightarrow Br_2 + HNO_3$	0.3	0.3	0.01
$N_2O_5 + HBr \rightarrow BrNO_2 + HNO_3$	0.05	0.001	0.02

f denotes uptake coefficients calculated using the method in Shi et al. (2001).

 Table A1.
 New heterogeneous reactions added to the UKCA module together with the uptake coefficients.

HCl injected by the eruption. $\delta^{15}N(NO_3^-)$ can theoretically be used as a proxy for UV 565 and thus has the potential to indicate changes in past TCO. Changes in $\delta^{15}N(NO_3^-)$ 566 could be positive or negative (indicating either increases or decreased in TCO) de-567 pending on the type of volcanic eruption and they are unlikely to be synchronous with 568 sulphate peaks because of different transport pathways and the different timings of the 569 ozone changes. Using a novel high resolution $\delta^{15}N(NO_3^-)$ ice core record, we are un-570 able to detect a signal from the largest volcanic eruption (1257 series) in the past 1000 571 years as there is a large inter-annual variability in the $\delta^{15} N(NO_3^-)$ record. We recom-572 mend that future studies of this nature should first understand why the $\delta^{15}N(NO_3^-)$ 573 record has a large variability at DML site if $\delta^{15}N(NO_3^-)$ is to be used to constrain the 574 ozone change. A site with lower variability than 25 per mil in $\delta^{15}N(NO_3^-)$ could be 575 considered although increasing the sensitivity of the UV proxy by choosing a site with 576 lower annual accumulation comes at the expense of reduced time resolution making it 577 less likely to resolve volcanic eruptions. 578

579 Appendix A Model improvements

A1 Heterogeneous and photolysis reactions

Table A1 lists the new heterogeneous reactions added to the UKCA module together with the uptake coefficients on ice, nitric acid trihydrate and sulfate aerosol. This can be compared to Table 1 in Dennison et al. (2019). We use the method in Shi et al. (2001) to calculate the values of the uptake coefficients that are not constant and are denoted by f in Table A1.

586 A2 Bromocarbon emissions

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The emission flux datasets of the five very short lived bromocarbon species (CH₃Br, CH₂BrCl, CH₂Br₂, CHBr₂Cl, CHBrCl₂) are explicitly included as emission files. These are similar to the ones used in Yang et al. (2014), which are based on

Volcano	Eruption date	Arrival date	Peak Depth (m)	Start Depth (m)
Kuwae ^a	1450	1454	61.01	61.13
1285^{b}	1285	1285	75.12	75.22
1277^{b}	1277	1277	75.77	75.9
1269^{b}	1269	1269	76.41	77.46
Samalas $1257^{\rm b}$	1257	1259	77.12	77.23
Unknown $1228/30^{\rm b}$	1229	1229	79.33	79.43

 $^{\rm a}$ Zielinski et al. (1994) $^{\rm b}$ Langway Jr. et al. (1995)

Table B1. Volcanic horizons identified from the sulfate and electrical meltwater conductivity records. Eruption date of the volcano and arrival dates of the sulfate in the ice core are obtained from Zielinski et al. (1994) and Langway Jr. et al. (1995) except for the Unknown 1228/30 volcano where the precise eruption date is not known. We choose 1229 as the eruption and arrival date for dating purposes.

the original work (scenario 5) of Warwick et al. (2006), except for the emissions of CH_2Br_2 . The emissions of CH_2Br_2 were scaled to give a total emission of $57 \,\mathrm{Gg\,yr^{-1}}$, corresponding to 50% of the original flux and in better agreement with Liang et al. (2010) and Ordóñez et al. (2012). The combined effect of the bromocarbons is to provide ~ 5pptv of inorganic bromine to the stratosphere (Yang et al., 2014) in a pre-industrial atmosphere.

596 Appendix B Ice core analysis

Table B1 shows the volcanic horizons identified from the sulfate and electrical meltwater conductivity records in the ISOL-ICE ice core.

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