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Citation for published version:

Zhang, X & Curtis, A 2020, 'Seismic Tomography Using Variational Inference Methods', *Journal of Geophysical Research: Solid Earth*, vol. 125, no. 4. <https://doi.org/10.1029/2019JB018589>

Digital Object Identifier (DOI):

[10.1029/2019JB018589](https://doi.org/10.1029/2019JB018589)

Link:

[Link to publication record in Edinburgh Research Explorer](#)

Document Version:

Peer reviewed version

Published In:

Journal of Geophysical Research: Solid Earth

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Seismic tomography using variational inference methods

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Key Points:

- We introduce two variational inference methods: automatic differential variational inference and Stein variational gradient descent.
- We applied the methods to solve synthetic and real-data seismic tomography, producing similar probabilistic results to Monte Carlo methods.
- Variational methods are efficient alternatives to Monte Carlo for generally non-linear Geophysical inverse and inference problems.

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Abstract

13
14 Seismic tomography is a methodology to image the interior of solid or fluid media, and
15 is often used to map properties in the subsurface of the Earth. In order to better inter-
16 pret the resulting images it is important to assess imaging uncertainties. Since tomog-
17 raphy is significantly nonlinear, Monte Carlo sampling methods are often used for this
18 purpose, but they are generally computationally intractable for large datasets and high-
19 dimensional parameter spaces. To extend uncertainty analysis to larger systems we use
20 variational inference methods to conduct seismic tomography. In contrast to Monte Carlo
21 sampling, variational methods solve the Bayesian inference problem as an optimization
22 problem, yet still provide probabilistic results. In this study, we applied two variational
23 methods, automatic differential variational inference (ADVI) and Stein variational gra-
24 dient descent (SVGD), to 2D seismic tomography problems using both synthetic and real
25 data and we compare the results to those from two different Monte Carlo sampling meth-
26 ods. The results show that ADVI provides a biased approximation because of its implicit
27 Gaussian approximation, and cannot be used to find multi-modal posteriors; SVGD can
28 produce more accurate approximations to the results of Monte Carlo sampling methods.
29 Both methods estimate the posterior distribution at significantly lower computational
30 cost, provided that gradients of parameters with respect to data can be calculated ef-
31 ficiently. We expect that the methods can be applied fruitfully to many other types of
32 geophysical inverse problems.

1 Introduction

34 In a variety of geoscientific applications, scientists need to obtain maps of subsur-
35 face properties in order to understand heterogeneity and processes taking place within
36 the Earth. Seismic tomography is a method that is widely used to generate those maps.
37 The maps of interest are usually parameterised in some way, and data are recorded that
38 can be used to constrain the parameters. Tomography is therefore a parameter estima-
39 tion problem, given the data and a physical relationship between data and parameters;
40 since the physical relationships usually predict data given parameter values but not the
41 reverse, seismic tomography involves solving an inverse problem (Curtis & Snieder, 2002).

42 Tomographic problems can be solved using either the full, known physical relation-
43 ships, or by using a linearised procedure which involves creating approximate, linearised
44 physics that is assumed to be accurate close to a particular chosen reference model. In

45 the linearised procedure, one seeks an optimal solution by perturbing the model so as
46 to minimize the misfit between the observed data and the data predicted by the linearised
47 physics. The physics is then re-linearised around this new reference model, and the pro-
48 cess is iterated until the perturbations are sufficiently small. Since most tomography prob-
49 lems are under-determined, some form of regularization must be introduced to solve the
50 system (Aki & Lee, 1976; Dziewonski & Woodhouse, 1987; Iyer & Hirahara, 1993; Taran-
51 tola, 2005). However, regularization is usually chosen using ad hoc criteria which intro-
52 duces poorly understood biases in the results; thus, valuable information can be concealed
53 by regularization (Zhdanov, 2002). Moreover, in nonlinear problems it is almost always
54 impossible to estimate accurate uncertainties in results using linearised methods. There-
55 fore, partially or fully nonlinear tomographic methods have been introduced to geophysics
56 which require no linearisation and which provide accurate estimates of uncertainty us-
57 ing a Bayesian probabilistic formulation of the parameter estimation problem. These in-
58 clude Monte Carlo methods (Mosegaard & Tarantola, 1995; Sambridge, 1999; Malinverno
59 et al., 2000; Malinverno, 2002; Malinverno & Briggs, 2004; Bodin & Sambridge, 2009;
60 Galetti et al., 2015, 2017; Zhang et al., 2018) and methods based on neural networks (Röth
61 & Tarantola, 1994; Devilee et al., 1999; Meier et al., 2007b, 2007a; Shahraeeni & Cur-
62 tis, 2011; Shahraeeni et al., 2012; Käuffl et al., 2013, 2015; Earp & Curtis, 2019).

63 Bayesian methods use Bayes’ theorem to update a *prior* probability distribution
64 function (*pdf* – either a conditional density function or a discrete set of probabilities)
65 with new information from data. The prior pdf describes information available about
66 the parameters of interest prior to the inversion. Bayes’ theorem combines the prior pdf
67 with information derived from the data to produce the total state of information about
68 the parameters post inversion, described by a so-called *posterior* pdf – this process is re-
69 ferred to as Bayesian inference. Thus, in our case Bayesian inference is used to solve the
70 tomographic inverse problem.

71 Monte Carlo methods generate a set (or chain) of samples from the posterior pdf
72 describing the probability distribution of the model given the observed data; thereafter
73 these samples can be used to estimate useful information about that pdf (mean, stan-
74 dard deviation, etc.). The methods are quite general from a theoretical point of view so
75 that in principle they can be applied to any tomographic problems. They have been ex-
76 tended to trans-dimensional inversion using the reversible jump Markov chain Monte Carlo
77 (rj-MCMC) algorithm (Green, 1995), in which the number of parameters (hence the di-

78 dimensionality of parameter space) can vary in the inversion. Consequently the param-
79 eterization itself can be simplified by adapting to the data which improves results on oth-
80 erwise high-dimensional problems (Malinverno et al., 2000; Bodin & Sambridge, 2009;
81 Bodin et al., 2012; Ray et al., 2013; Young et al., 2013; Galetti et al., 2015, 2017; Hawkins
82 & Sambridge, 2015; Piana Agostinetti et al., 2015; Burdick & Lekić, 2017; Galetti & Cur-
83 tis, 2018; Zhang et al., 2018, 2019). Although many applications have been conducted
84 using MCMC sampling methods (previous references, Shen et al., 2012, 2013; Zulfakriza
85 et al., 2014; Zheng et al., 2017; Crowder et al., 2019), they mainly address 1D or 2D to-
86 mography problems due to the high computational expense of Monte Carlo methods. Some
87 studies used MCMC methods for fully 3D tomography using body wave travel time data
88 (Hawkins & Sambridge, 2015; Piana Agostinetti et al., 2015; Burdick & Lekić, 2017) and
89 surface wave dispersion (Zhang et al., 2018, 2019), but the methods demand enormous
90 computational resources. Even in the 1D or 2D case, MCMC methods cannot easily be
91 applied to large datasets which are generally expensive to forward model given a set of
92 parameter values. Moreover, MCMC methods tend to be inefficient at exploring complex,
93 multi-modal probability distributions (Sivia, 1996; Karlin, 2014), which appear to be com-
94 mon in seismic tomography problems.

95 Neural network based methods offer an efficient alternative for certain classes of
96 tomography problems that will be solved many times with new data of the same type.
97 An initial set of Monte Carlo samples is taken from the prior probability distribution over
98 parameter space, and data are computationally forward modelled for each parameter vec-
99 tor. Neural networks are flexible mappings that can be regressed (trained) to emulate
100 the mapping from data to parameter space by fitting the set of examples of that map-
101 ping generated using Monte Carlo (Bishop, 2006). Since for each input data vector the
102 neural network only produces one parameter vector, trade-offs between parameters are
103 not clearly represented in the mapping from data to model parameters. The trained net-
104 work then interpolates the inverse mapping between the examples, and can be applied
105 efficiently to any new, measured data to estimate corresponding parameter values. The
106 first geophysical application of neural network tomography was Röth and Tarantola (1994),
107 but that application did not estimate uncertainties. Forms of networks that estimate to-
108 mographic uncertainties were introduced by Devilee et al. (1999) and Meier et al. (2007b,
109 2007a) and have been applied to surface and body wave tomography in 1D and 2D prob-
110 lems (Meier et al., 2007b, 2007a; Earp & Curtis, 2019). Nevertheless, neural networks

111 still suffer from the computational cost of generating the initial set of training examples.
 112 That set may have to include many more samples than are required for standard Bayesian
 113 MC, because the training set must span the prior pdf whereas standard applications of
 114 MC tomography sample the posterior pdf which is usually more tightly constrained. Neu-
 115 ral networks have the advantage that the training samples need only be calculated once
 116 for any number of data sets whereas MC inversion must perform sampling for every new
 117 data set. However, in high dimensional problems the cost of sampling may be prohibitive
 118 for both MC and NN based methods due to the curse of dimensionality (the exponen-
 119 tial increase in the hypervolume of parameter space as the number of parameters increases
 120 – Curtis & Lomax, 2001).

121 Variational inference provides a different way to solve a Bayesian inference prob-
 122 lem: within a predefined family of probability distributions, one seeks an optimal approx-
 123 imation to a target distribution which in this case is the Bayesian posterior pdf. This
 124 is achieved by minimizing the Kullback-Leibler (KL) divergence (Kullback & Leibler, 1951)
 125 – one possible measure of the difference between two given pdfs (Blatter et al., 2019),
 126 in our case the difference between approximate and target pdfs (Bishop, 2006; Blei et
 127 al., 2017). Since the method casts the inference problem into an optimization problem,
 128 it can be computationally more efficient than either MC sampling or neural network meth-
 129 ods, and provides better scaling to higher dimensional problems. Moreover, it can be used
 130 to take advantage of methods such as stochastic optimization (Robbins & Monro, 1951;
 131 Kubrusly & Gravier, 1973) and distributed optimization by dividing large datasets into
 132 random minibatches – methods which are difficult to apply for MCMC methods since they
 133 may break the reversibility property of Markov chains which is required by most MCMC
 134 methods.

135 In variational inference, the complexity of the approximating family of pdfs deter-
 136 mines the complexity of the optimization. A complex variational family is generally more
 137 difficult to optimize than a simple family. Therefore, many applications are performed
 138 using simple mean-field approximation families (Bishop, 2006; Blei et al., 2017) and struc-
 139 tured families (Saul & Jordan, 1996; Hoffman & Blei, 2015). For example, in Geophysics
 140 the method has been used to invert for the spatial distribution of geological facies given
 141 seismic data using a mean-field approximation (M. A. Nawaz & Curtis, 2018; M. Nawaz
 142 & Curtis, 2019).

143 Even using those simple families, applications of variational inference methods usu-
 144 ally involve tedious derivations and bespoke implementations for each type of problem
 145 which restricts their applicability (Bishop, 2006; Blei et al., 2017; M. A. Nawaz & Cur-
 146 tis, 2018; M. Nawaz & Curtis, 2019). The simplicity of those families also affects the qual-
 147 ity of the approximation to complex distributions. To make variational methods easier
 148 to use, "black box" variational inference methods have been proposed (Kingma & Welling,
 149 2013; Ranganath et al., 2014, 2016). Based on these ideas, Kucukelbir et al. (2017) pro-
 150 posed an automatic variational inference method which can easily be applied to many
 151 Bayesian inference problems. Another set of methods has been proposed based on prob-
 152 ability transformations (Rezende & Mohamed, 2015; Tran et al., 2015; Q. Liu & Wang,
 153 2016; Marzouk et al., 2016); these methods optimise a series of invertible transforms to
 154 approximate the target probability and in this case it is possible to approximate arbi-
 155 trary probability distributions.

156 We apply automatic differential variational inference (ADVI – Kucukelbir et al.,
 157 2017) and Stein variational gradient descent (SVGD – Q. Liu & Wang, 2016) to a 2D
 158 seismic tomography problem. In the following we first describe the basic idea of varia-
 159 tional inference, and then the ADVI and SVGD methods. In section 3 we apply the two
 160 methods to a simple 2D synthetic seismic tomography example and compare their re-
 161 sults with both fixed-dimensional MCMC and rj-MCMC. In section 4 we apply the two
 162 methods to real data from Grane field, North Sea, to study the phase velocity map at
 163 0.9 s and compare the results to those found using rj-MCMC. We thus demonstrate that
 164 variation inference methods can provide efficient alternatives to MCMC methods while
 165 still producing reasonably accurate approximations to Bayesian posterior pdfs. Our aim
 166 is to introduce variational inference methods to the geoscientific community and to en-
 167 courage more research on this topic.

168 **2 Methods**

169 **2.1 Variational inference**

170 Bayesian inference involves calculating or characterising a posterior probability den-
 171 sity function $p(\mathbf{m}|\mathbf{d}_{obs})$ of model parameters \mathbf{m} given the observed data \mathbf{d}_{obs} . Accord-
 172 ing to Bayes' theorem,

$$173 \quad p(\mathbf{m}|\mathbf{d}_{obs}) = \frac{p(\mathbf{d}_{obs}|\mathbf{m})p(\mathbf{m})}{p(\mathbf{d}_{obs})} \quad (1)$$

174 where $p(\mathbf{d}_{obs}|\mathbf{m})$ is called the *likelihood* which is the probability of observing data \mathbf{d}_{obs}
 175 conditional on model \mathbf{m} , $p(\mathbf{m})$ is the *prior* which describes known information about the
 176 model that is independent of the data, and $p(\mathbf{d}_{obs})$ is a normalization factor called the
 177 *evidence* which is constant for a fixed model parameterization. The likelihood is usually
 178 assumed to follow a Gaussian probability density function around the data predicted syn-
 179 thetically from model \mathbf{m} (using the known physical relationships), as this is assumed to
 180 be a reasonable approximation to the pdf of uncertainties or errors in the measured data,
 181 and because noise reduction is performed by stacking, which through the central limit
 182 theorem justifies the use of a Gaussian distribution.

183 Variational inference approximates the above pdf $p(\mathbf{m}|\mathbf{d}_{obs})$ using optimization. First
 184 a family (set) of known distributions $\mathcal{Q} = \{q(\mathbf{m})\}$ is defined. The method then seeks
 185 the best approximation to $p(\mathbf{m}|\mathbf{d}_{obs})$ within that family by minimizing the KL-divergence:

$$186 \quad \text{KL}[q(\mathbf{m})||p(\mathbf{m}|\mathbf{d}_{obs})] = E_q[\log q(\mathbf{m})] - E_q[\log p(\mathbf{m}|\mathbf{d}_{obs})] \quad (2)$$

187 where the expectation is taken with respect to distribution $q(\mathbf{m})$. It can be shown that
 188 $\text{KL}[q||p] \geq 0$ and has zero value if and only if $q(\mathbf{m})$ equals $p(\mathbf{m}|\mathbf{d}_{obs})$ (Kullback & Leibler,
 189 1951). Distribution $q^*(\mathbf{m})$ that minimizes the KL-divergence is therefore the best ap-
 190 proximation to $p(\mathbf{m}|\mathbf{d}_{obs})$ within the family \mathcal{Q} .

191 Combining equations (1) and (2), the KL-divergence becomes:

$$192 \quad \text{KL}[q(\mathbf{m})||p(\mathbf{m}|\mathbf{d}_{obs})] = E_q[\log q(\mathbf{m})] - E_q[\log p(\mathbf{m}, \mathbf{d}_{obs})] + \log p(\mathbf{d}_{obs}) \quad (3)$$

193 The evidence term $\log p(\mathbf{d}_{obs})$ generally cannot be calculated since it involves the eval-
 194 uation of a high dimensional integral which takes exponential time. Instead we calcu-
 195 late the evidence lower bound (ELBO) which is equivalent to the KL-divergence up to
 196 an unknown constant, and is obtained by rearranging equation (3) and using the fact
 197 that $\text{KL}[q||p] \geq 0$:

$$198 \quad \begin{aligned} \text{ELBO}[q] &= E_q[\log p(\mathbf{m}, \mathbf{d}_{obs})] - E_q[\log q(\mathbf{m})] \\ &= \log p(\mathbf{d}_{obs}) - \text{KL}[q(\mathbf{m})||p(\mathbf{m}|\mathbf{d}_{obs})] \end{aligned} \quad (4)$$

200 Thus minimizing the KL-divergence is equivalent to maximizing the ELBO.

201 In variational inference, the choice of the variational family is important because
 202 the flexibility of the variational family determines the power of the approximation. How-
 203 ever, it is usually more difficult to optimize equation (4) over a complex family than a

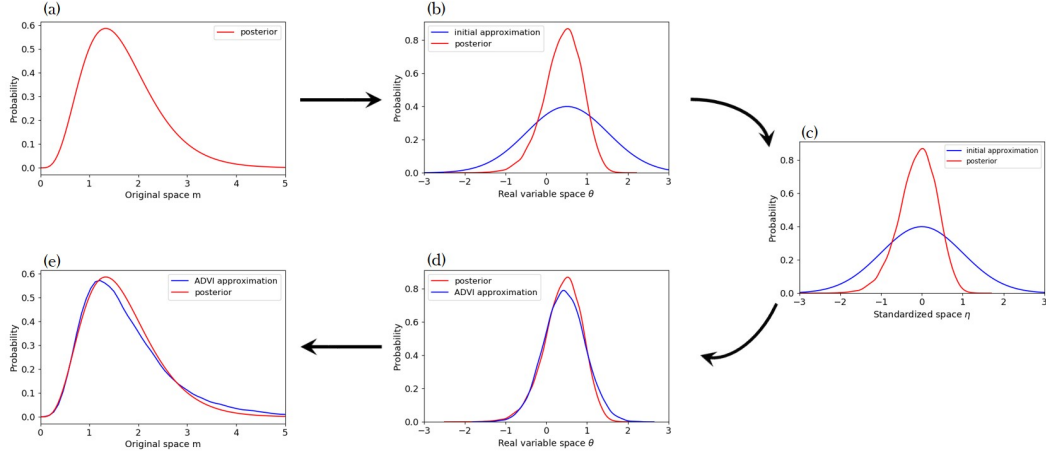


Figure 1. An illustration of the workflow of ADVI. **(a)** An example of a posterior pdf in the original positive half space of parameters \mathbf{m} . **(b)** The posterior pdf in the transformed real variable space θ (red) and an initial Gaussian approximation (blue). **(c)** The posterior pdf (red) and the standard Gaussian distribution (blue) in standardized variable η ; gradients with respect to variational parameters are calculated in this space. **(d)** and **(e)** show the posterior pdf (red) and the approximation obtained using ADVI (blue) in the unconstrained real variable space and the original space, respectively.

204 simple family. Therefore, many applications are performed using the *mean-field* varia-
 205 tional family, which means that the parameters \mathbf{m} are treated as being mutually inde-
 206 pendent (Bishop, 2006; Blei et al., 2017). However, even under that simplifying assump-
 207 tion, traditional variational methods require tedious model-specific derivations and im-
 208 plementations, which restricts their applicability to those problems for which derivations
 209 have been performed (e.g., M. A. Nawaz & Curtis, 2018; M. Nawaz & Curtis, 2019). We
 210 therefore introduce two more general variational methods: the automatic differential vari-
 211 ational inference (ADVI) and the Stein variational gradient descent (SVGD), which can
 212 both be applied to general inverse problems.

213 **2.2 Automatic differential variational inference (ADVI)**

214 Kucukelbir et al. (2017) proposed a general variational method called automatic
 215 differential variational inference (ADVI) based on a Gaussian variational family. In ADVI,
 216 a model with constrained parameters is first transformed to a model with unconstrained
 217 real-valued variables. For example, the velocity model \mathbf{m} that usually has hard bound

218 constraints (such as velocity being greater than zero) can be transformed to an uncon-
 219 strained model $\boldsymbol{\theta} = T(\mathbf{m})$, where T is an invertible and differentiable function (Figure
 220 1a and b). The joint probability $p(\mathbf{m}, \mathbf{d}_{obs})$ then becomes:

$$221 \quad p(\boldsymbol{\theta}, \mathbf{d}_{obs}) = p(\mathbf{m}, \mathbf{d}_{obs}) |det \mathbf{J}_{T^{-1}}(\boldsymbol{\theta})| \quad (5)$$

222 where $\mathbf{J}_{T^{-1}}(\boldsymbol{\theta})$ is the Jacobian matrix of the inverse of T which accounts for the volume
 223 change of the transform, and $|\cdot|$ represents the absolute value. This transform makes
 224 the choice of variational approximations independent of the original model since trans-
 225 formed variables lie in the common unconstrained space of real numbers.

226 In ADVI, we choose a Gaussian variational family (e.g., blue line in Figure 1b),

$$227 \quad q(\boldsymbol{\theta}; \boldsymbol{\phi}) = \mathcal{N}(\boldsymbol{\theta} | \boldsymbol{\mu}, \boldsymbol{\Sigma}) = \mathcal{N}(\boldsymbol{\theta} | \boldsymbol{\mu}, \mathbf{L}\mathbf{L}^T) \quad (6)$$

228 where $\boldsymbol{\phi}$ represents variational parameters $\boldsymbol{\mu}$ and $\boldsymbol{\Sigma}$, $\boldsymbol{\mu}$ is the mean vector and $\boldsymbol{\Sigma}$ is the
 229 covariance matrix. As in Kucukelbir et al. (2017), for computational purposes we use a
 230 Cholesky factorization $\boldsymbol{\Sigma} = \mathbf{L}\mathbf{L}^T$ where \mathbf{L} is a lower-triangular matrix, to re-parameterize
 231 the covariance matrix to ensure that it is positive semidefinite (covariance is positive semidef-
 232 inite by definition). If $\boldsymbol{\Sigma}$ is a diagonal matrix, q reduces to a mean-field approximation
 233 in which the variables are mutually independent; in order to include spatial correlations
 234 in the velocity model we use a full-rank covariance matrix, noting that this incurs a com-
 235 putational cost since it increases the number of variational parameters.

236 In the transformed space, the variational problem is solved by maximizing the ELBO,
 237 written as \mathcal{L} , with respect to variational parameters $\boldsymbol{\phi}$:

$$238 \quad \begin{aligned} \boldsymbol{\phi}^* &= \arg \max_{\boldsymbol{\phi}} \mathcal{L}[q(\boldsymbol{\theta}; \boldsymbol{\phi})] \\ &= \arg \max_{\boldsymbol{\phi}} \mathbb{E}_q [\log p(T^{-1}(\boldsymbol{\theta}), \mathbf{d}_{obs}) + \log |det \mathbf{J}_{T^{-1}}(\boldsymbol{\theta})|] - \mathbb{E}_q [\log q(\boldsymbol{\theta})] \end{aligned} \quad (7)$$

239 This is an optimization problem in an unconstrained space and can be solved using gra-
 240 dient ascent methods without worrying about any constrains on the original variables.

241 However, the gradients of variational parameters are not easy to calculate since the
 242 ELBO involves expectations in a high dimensional space. We therefore transform the
 243 Gaussian distribution $q(\boldsymbol{\theta}; \boldsymbol{\phi})$ into a standard Gaussian $\mathcal{N}(\boldsymbol{\eta} | \mathbf{0}, \mathbf{I})$ (Figure 1c), by $\boldsymbol{\eta} =$

244 $R_\phi(\boldsymbol{\theta}) = \mathbf{L}^{-1}(\boldsymbol{\theta} - \boldsymbol{\mu})$, thereafter the variational problem becomes:

$$\begin{aligned} \phi^* &= \arg \max_{\phi} \mathcal{L}[q(\boldsymbol{\theta}; \phi)] \\ &= \arg \max_{\phi} \mathbb{E}_{\mathcal{N}(\boldsymbol{\eta}|\mathbf{0}, \mathbf{I})} \left[\log p\left(T^{-1}\left(R_\phi^{-1}(\boldsymbol{\eta})\right), \mathbf{d}_{obs}\right) + \log|\det \mathbf{J}_{T^{-1}}\left(R_\phi^{-1}(\boldsymbol{\eta})\right)| \right] - \mathbb{E}_q[\log q(\boldsymbol{\theta})] \end{aligned} \quad (8)$$

245
246 where the first expectation is taken with respect to a standard Gaussian distribution $\mathcal{N}(\boldsymbol{\eta}|\mathbf{0}, \mathbf{I})$.
247 There is no Jacobian term related to this transform since the determinant of the Jaco-
248 bian is equal to one (Kucukelbir et al., 2017). The second expectation $-\mathbb{E}_q[\log q(\boldsymbol{\theta})]$ is
249 not transformed since it has a simple analytic form as does its gradient (Kucukelbir et
250 al., 2017) – see Appendix A.

251 Since the distribution with respect to which the expectation is taken now does not
252 depend on variational parameters, the gradient with respect to variational parameters
253 can be calculated by exchanging the expectation and derivative according to the dom-
254 inated convergence theorem (Çınlar, 2011) and by applying the chain rule – see Appendix
255 B:

$$\nabla_{\boldsymbol{\mu}} \mathcal{L} = \mathbb{E}_{\mathcal{N}(\boldsymbol{\eta}|\mathbf{0}, \mathbf{I})} \left[\nabla_{\mathbf{m}} \log p(\mathbf{m}, \mathbf{d}_{obs}) \nabla_{\boldsymbol{\theta}} T^{-1}(\boldsymbol{\theta}) + \nabla_{\boldsymbol{\theta}} \log|\det \mathbf{J}_{T^{-1}}(\boldsymbol{\theta})| \right] \quad (9)$$

257 The gradient with respect to \mathbf{L} can be obtained similarly,

$$\nabla_{\mathbf{L}} \mathcal{L} = \mathbb{E}_{\mathcal{N}(\boldsymbol{\eta}|\mathbf{0}, \mathbf{I})} \left[(\nabla_{\mathbf{m}} \log p(\mathbf{m}, \mathbf{d}_{obs}) \nabla_{\boldsymbol{\theta}} T^{-1}(\boldsymbol{\theta}) + \nabla_{\boldsymbol{\theta}} \log|\det \mathbf{J}_{T^{-1}}(\boldsymbol{\theta})|) \boldsymbol{\eta}^T \right] + (\mathbf{L}^{-1})^T \quad (10)$$

259 where the expectation is computed with respect to a standard Gaussian distribution, which
260 can be estimated by Monte Carlo (MC) integration. MC integration provides a noisy,
261 unbiased estimation of the expectation and its accuracy increases with the number of
262 samples. Nevertheless, it has been shown that in practice a low number or even a sin-
263 glesample can be sufficient at each iteration since the mean is taken with respect to the
264 standard Gaussian distribution (see discussions and experiments in Kucukelbir et al., 2017).
265 For distributions $p(\mathbf{m}, \mathbf{d}_{obs})$ for which the gradients have analytic forms, the whole pro-
266 cess of computing gradients can be automated (Kucukelbir et al., 2017), hence the name
267 ”automatic differential”. We can then use a gradient ascent method to update the vari-
268 ational parameters and obtain an approximation to the pdf $p(\mathbf{m}|\mathbf{d}_{obs})$ (e.g. Figure 1d).

269 Note that although the method is based on Gaussian variational approximations,
270 the actual shape of the approximation to the posterior $p(\mathbf{m}|\mathbf{d}_{obs})$ over the original pa-
271 rameters \mathbf{m} is determined by the transform T (Figure 1e). It is difficult to determine an
272 optimal transform since that is related to the properties of the unknown posterior (Kucukelbir

273 et al., 2017). In this study we use a commonly-used invertible logarithmic transform (Team
274 et al., 2016),

$$\begin{aligned}
 \theta_i &= T(m_i) = \log(m_i - a_i) - \log(b_i - m_i) \\
 m_i &= T^{-1}(\theta_i) = a_i + \frac{(b_i - a_i)}{1 + \exp(-\theta_i)}
 \end{aligned}
 \tag{11}$$

276 where m_i represents each original constrained parameter, θ_i is the transformed uncon-
277 strained variable, a_i is the original lower bound and b_i the upper bound on m_i . There-
278 fore the quality of the ADVI approximation is limited by the Gaussian approximation
279 in the unconstrained space and by the specific transform T in equation (11).

280 To illustrate the effects of the transform in equation (11), we show an example in
281 Figure 2. The original variable lies in a constrained space between 0.5 and 3.0 (a typ-
282 ical phase velocity range of seismic surface waves). The space is transformed to an un-
283 constrained space using equation (11). If, as in ADVI we assume a standard Gaussian
284 distribution in the transformed space (blue area in Figure 2), the associated probabil-
285 ity distribution in the original space is shown in orange in Figure 2. The actual shape
286 of the distribution in the original space is not Gaussian but is determined by the trans-
287 form T in equation (11). However, under this choice of T it is likely that the probabil-
288 ity distribution in the original space is still unimodal. We thus see that ADVI provides
289 a unimodal approximation of the target posterior pdf around a local optimal parame-
290 ter estimate. This suggests that the method will not be effective for multimodal distri-
291 butions, and the estimated probability distribution depends on the initial value of μ and
292 Σ (Kucukelbir et al., 2017). However, since the maximum a posteriori probability (MAP)
293 estimate has been shown to be effective for parameter estimation in practice, the ADVI
294 method could still be used to provide a good approximation of the distribution around
295 a MAP estimate.

296 **2.3 Stein variational gradient descent (SVGD)**

297 In practice most applications of variational inference use simple families of poste-
298 rior approximations such as a Gaussian approximation (Kucukelbir et al., 2017), mean-
299 field approximations (Blei et al., 2017; M. A. Nawaz & Curtis, 2018; M. Nawaz & Cur-
300 tis, 2019) or other simple structured families (Saul & Jordan, 1996; Hoffman & Blei, 2015).
301 These simple choices significantly restrict the quality of derived posterior approximations.
302 In order to employ a broader family of variational approximations, variational methods
303 based on invertible transforms have been proposed (Rezende & Mohamed, 2015; Tran

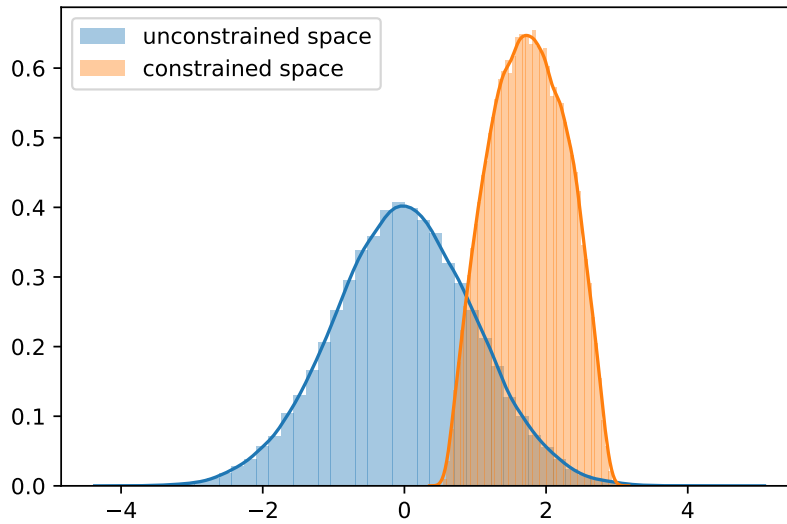


Figure 2. An illustration of the transform in equation (11). The original variable is in a constrained space between 0.5 and 3.0. The blue area shows a standard Gaussian distribution in the transformed unconstrained space and the orange area shows the associated probability distribution in the original space. The probability distributions are estimated using Monte Carlo samples.

304 et al., 2015; Marzouk et al., 2016). In these methods instead of choosing specific forms
 305 for variational approximations, a series of invertible transforms are applied to an initial
 306 distribution, and these transforms are optimized by minimizing the KL-divergence. This
 307 provides a way to approximate arbitrary posterior distributions since a pdf can be trans-
 308 formed to any other pdf as long as the probability measures are absolutely continuous.

309 Stein variational gradient descent (SVGD) is one such algorithm based on an in-
 310 cremental transform (Q. Liu & Wang, 2016). In SVGD, a smooth transform $T(\mathbf{m}) =$
 311 $\mathbf{m} + \epsilon \phi(\mathbf{m})$ is used, where $\mathbf{m} = [m_1, \dots, m_d]$ and m_i is the i^{th} parameter, and $\phi(\mathbf{m}) =$
 312 $[\phi_1, \dots, \phi_d]$ is a smooth vector function that describes the perturbation direction and where
 313 ϵ is the magnitude of the perturbation. It can be shown that when ϵ is sufficiently small,
 314 the transform is invertible since the Jacobian of the transform is close to an identity ma-
 315 trix (Q. Liu & Wang, 2016). Say $q_T(\mathbf{m})$ is the transformed probability distribution of
 316 the initial distribution $q(\mathbf{m})$. Then the gradient of KL-divergence with respect to ϵ can
 317 be computed as (see Appendix C):

$$318 \quad \nabla_{\epsilon} \text{KL}[q_T||p] |_{\epsilon=0} = -\mathbb{E}_q [\text{trace}(\mathcal{A}_p \phi(\mathbf{m}))] \quad (12)$$

319 where \mathcal{A}_p is the Stein operator such that $\mathcal{A}_p \phi(\mathbf{m}) = \nabla_{\mathbf{m}} \log p(\mathbf{m}) \phi(\mathbf{m})^T + \nabla_{\mathbf{m}} \phi(\mathbf{m})$.
 320 This suggests that maximizing the right-hand expectation with respect to $q(\mathbf{m})$ gives the
 321 steepest descent of the KL-divergence, and consequently the KL-divergence can be min-
 322 imized iteratively.

323 It can be shown that the negative gradient of the KL-divergence in equation (12)
 324 can be maximized by using the kernelized Stein discrepancy (Q. Liu et al., 2016). For
 325 two continuous probability densities p and q , the *Stein discrepancy* for a function ϕ in
 326 a function set \mathcal{F} is defined as:

$$327 \quad S[q, p] = \arg \max_{\phi \in \mathcal{F}} \{[\mathbb{E}_q \text{trace}(\mathcal{A}_p \phi(\mathbf{m}))]^2\} \quad (13)$$

328 The Stein discrepancy provides another way to quantify the difference between two dis-
 329 tribution densities (Stein et al., 1972; Gorham & Mackey, 2015). However the Stein dis-
 330 crepancy is not easy to compute for general \mathcal{F} . Therefore, Q. Liu et al. (2016) proposed
 331 a kernelized Stein discrepancy by maximizing equation (13) in the unit ball of a repro-
 332 ducing kernel Hilbert space (RKHS) as follows.

333 A Hilbert space is a space \mathcal{H} on which an inner product $\langle, \rangle_{\mathcal{H}}$ is defined. A func-
 334 tion is called a *kernel* if there exists a real Hilbert space and a function φ such that $k(x, y) = \langle$

335 $\varphi(x), \varphi(y) \rangle_{\mathcal{H}}$ (Gretton, 2013). A kernel is said to be positive-definite if the matrix de-
 336 fined by $K_{ij} = k(x_i, x_j)$ is positive definite. Assuming a positive definite kernel $k(\mathbf{m}, \mathbf{m}')$
 337 on $\mathcal{M} \times \mathcal{M}$, its reproducing kernel Hilbert space \mathcal{H} is defined by the closure of the linear
 338 span $\{f : f(\mathbf{m}) = \sum_{i=1}^n a_i k(\mathbf{m}, \mathbf{m}^i), a_i \in \mathcal{R}, n \in \mathcal{N}, \mathbf{m}^i \in \mathcal{M}\}$ with inner products
 339 $\langle f, g \rangle_{\mathcal{H}} = \sum_{i,j} a_i b_j k(\mathbf{m}^i, \mathbf{m}^j)$ for $g(\mathbf{m}) = \sum_i b_i k(\mathbf{m}, \mathbf{m}^i)$. The RKHS has an impor-
 340 tant reproducing property, that is, $f(x) = \langle f(x'), k(x', x) \rangle_{\mathcal{H}}$, such that the evalua-
 341 tion of a function f at x can be represented as an inner product in the Hilbert space.
 342 In a RKHS, the kernelized Stein discrepancy can be defined as (Q. Liu et al., 2016)

$$343 \quad S[q, p] = \arg \max_{\phi \in \mathcal{H}^d} \{E_q [\text{trace}(\mathcal{A}_p \phi(\mathbf{m}))]^2, \quad s.t. \quad \|\phi\|_{\mathcal{H}^d} \leq 1\} \quad (14)$$

344 where \mathcal{H}^d is the RKHS of d -dimensional vector functions. The right side of equation (14)
 345 is found to be equal to,

$$346 \quad \phi^* = \phi_{q,p}^*(\mathbf{m}) / \|\phi_{q,p}^*(\mathbf{m})\|_{\mathcal{H}^d} \quad (15)$$

347 where

$$348 \quad \phi_{q,p}^*(\mathbf{m}) = E_{\{\mathbf{m}' \sim q\}} [\mathcal{A}_p k(\mathbf{m}', \mathbf{m})] \quad (16)$$

349 and for which we have $S[q, p] = \|\phi_{q,p}^*(\mathbf{m})\|_{\mathcal{H}^d}$. Thus the optimal ϕ in equation (12) is
 350 ϕ^* and $\nabla_{\epsilon} \text{KL}[q_T || p] |_{\epsilon=0} = -\sqrt{S[q, p]}$.

351 Given the above solution, the SVGD works as follows: we start from an initial dis-
 352 tribution q_0 , then apply the transform $T_0^*(\mathbf{m}) = \mathbf{m} + \epsilon \phi_{q_0,p}^*(\mathbf{m})$ where we absorb the
 353 normalization term in equation (15) into ϵ ; this updates q_0 to $q_{[T_0]}$ with a decrease in the
 354 KL-divergence of $\epsilon * \sqrt{S[q, p]}$. This process is iterated to obtain an approximation of
 355 the posterior p :

$$356 \quad q_{l+1} = q_{l[T_l^*]}, \quad \text{where } T_l^*(\mathbf{m}) = \mathbf{m} + \epsilon_l \phi_{q_l,p}^*(\mathbf{m}) \quad (17)$$

357 and for sufficiently small $\{\epsilon_l\}$ the process eventually converges to the posterior pdf p . Note
 358 that a large stepsize may lead the Jacobian matrix of transform T to be singular, which
 359 in turn makes the approximation probability fail to converge to the true posterior (Q. Liu,
 360 2017).

361 To calculate the expectation in equation (16) we start from a set of particles (mod-
 362 els) generated using q_0 , and at each step the $\phi_{q,p}^*(\mathbf{m})$ can be estimated by computing
 363 the mean in equation (16) using those particles. Each particle is then updated using the
 364 transform in equation (17), and those particles will form better approximations to the
 365 posterior as the iteration proceeds. This suggests the following algorithm which is schemat-
 366 ically represented in Figure 3:

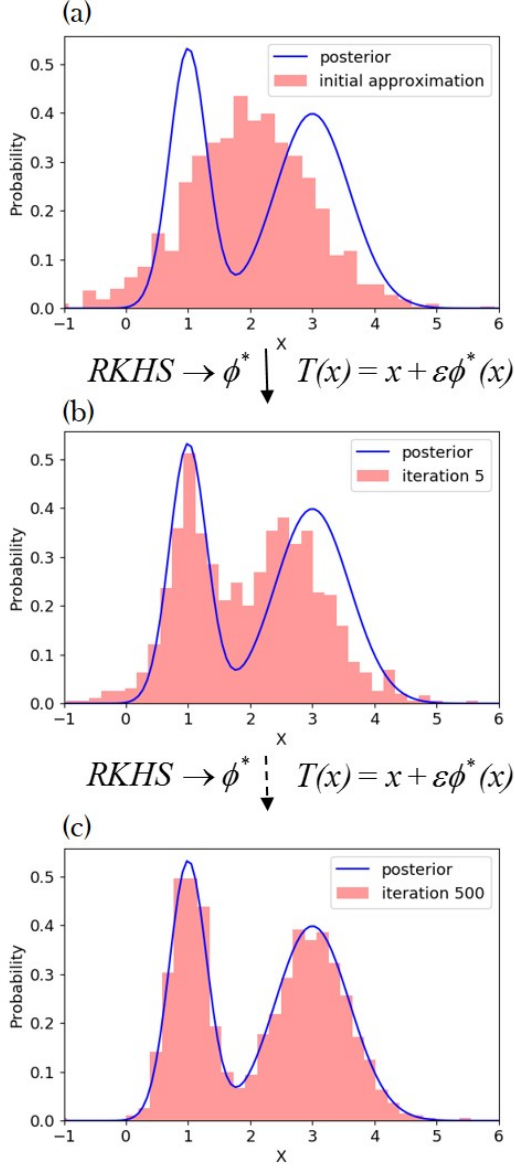


Figure 3. An illustration of the SVGD algorithm. The initial pdf is represented by the density of a set of particles (red histogram) in the top plot. The particles are then updated using a smooth transform $T(x) = x + \epsilon\phi^*(x)$, where ϕ^* is found in a reproducing kernel Hilbert space (RKHS). (a) An example of a posterior pdf (blue line) and an initial distribution (red histogram). (b) The approximating probability distribution after 5 iterations. (c) The approximating probability distribution after 500 iterations.

1. Draw a set of particles $\{\mathbf{m}_i^0\}_{i=1}^n$ from an initial pdf estimate (e.g., the prior).
2. At iteration l , update each particle using:

$$\mathbf{m}_i^{l+1} = \mathbf{m}_i^l + \epsilon_l \phi_{q_l, p}^*(\mathbf{m}_i^l) \quad (18)$$

where

$$\phi_{q_l, p}^*(\mathbf{m}) = \frac{1}{n} \sum_{j=1}^n \left[k(\mathbf{m}_j^l, \mathbf{m}) \nabla_{\mathbf{m}_j^l} \log p(\mathbf{m}_j^l) + \nabla_{\mathbf{m}_j^l} k(\mathbf{m}_j^l, \mathbf{m}) \right] \quad (19)$$

and ϵ_l is the step size at iteration l .

3. Calculate the density of the final set of particles $\{\mathbf{m}_i^*\}_{i=1}^n$ which approximates the posterior probability density function.

For kernel $k(\mathbf{m}, \mathbf{m}')$ we use the radial basis function $k(\mathbf{m}, \mathbf{m}') = \exp(-\frac{1}{h} \|\mathbf{m} - \mathbf{m}'\|^2)$, where h is taken to be $\tilde{d}^2 / \log n$ where \tilde{d} is the median of pairwise distances between all particles. This choice of h is based on the intuition that $\sum_j k(\mathbf{m}_i, \mathbf{m}_j) \approx n \exp(-\frac{1}{h} \tilde{d}^2) = 1$, so that for particle \mathbf{m}_i the two gradient terms in equation (19) are balanced (Q. Liu & Wang, 2016). For the radial basis function kernel the second term in equation (19) becomes $\sum_j \frac{2}{h} (\mathbf{m} - \mathbf{m}_j) k(\mathbf{m}_j, \mathbf{m})$, which drives the particle \mathbf{m} away from neighbouring particles for which the kernel takes large values. Therefore the second term in equation (19) acts as a *repulsive force* preventing particles from collapsing to a single mode, while the first term moves particles towards local high probability areas using the kernel-weighted gradient. If in the kernel $h \rightarrow 0$, the algorithm falls into independent gradient ascent that maximizes $\log p$ for each particle.

Note that since SVGD uses kernelized Stein discrepancy, the choice of kernels may affect the efficiency of the algorithm. In this study we adopted a commonly used kernel: a radial basis function. However, in some cases other kernels may provide a more efficient algorithm, for example, an inverse multiquadric kernel (Gorham & Mackey, 2017), a Hessian kernel (Detommaso et al., 2018) and kernels on a Riemann manifold (C. Liu & Zhu, 2018).

In SVGD, the accuracy of the approximation increases with the number of particles. It has been shown that compared to other particle-based methods, e.g., sequential Monte Carlo methods (Smith, 2013), SVGD requires fewer samples to achieve the same accuracy which makes it a more efficient method (Q. Liu & Wang, 2016). In contrast to sequential Monte Carlo which is a stochastic process, SVGD acts as a deterministic sampling method. If only one particle is used, the second term in equation (19) becomes

398 zero and the method reduces to a typical gradient ascent towards the model with the
 399 maximum a posterior (MAP) pdf value. This suggests that even for a small number of
 400 particles the method could still produce a good parameter estimate since MAP estima-
 401 tion can be an effective method in practice. Thus, in practice one could start from a small
 402 number of particles and gradually increase the number to find an optimal choice.

403 In seismic tomography velocities are usually constrained to lie within a given ve-
 404 locity range. In order to ensure that velocities always lie within the constrains, we first
 405 apply the same transform used in ADVI (equation 11) so that the parameters are in an
 406 unconstrained space. We can then simply use equation (18) to update particles without
 407 explicitly considering the constrains on seismic velocities. The final seismic velocities can
 408 be obtained by transforming particles back to the constrained space.

409 **3 Synthetic tests**

410 We first apply the above methods to a simple 2D synthetic example similar to that
 411 in Galetti et al. (2015). The true model is a homogeneous background with velocity 2
 412 km/s containing a circular low velocity anomaly with a radius of 2 km with velocity 1
 413 km/s . The 16 receivers are evenly distributed around the anomaly approximating a cir-
 414 cular acquisition geometry with radius 4 km (Figure 4a). Each receiver is also treated
 415 as a source to simulate a typical ambient noise interferometry experiment (Campillo &
 416 Paul, 2003; Curtis et al., 2006; Galetti et al., 2015). This produces a total of 120 inter-
 417 receiver travel time data, each of which is computed using a fast marching method of
 418 solving the Eikonal equation over a 100×100 gridded discretisation in space (Rawlinson
 419 & Sambridge, 2004).

420 For variational inversions we use a fixed 21×21 grid of cells to parameterize the
 421 velocity model \mathbf{m} (Figure 4a). The noise level is fixed to be 0.05 s (< 5 percent of travel
 422 times) for all inversions. The prior pdf of the velocity in each cell is set to be a Uniform
 423 distribution between 0.5 km/s and 3.0 km/s to encompass the true model. Travel times
 424 are calculated using the same fast marching method as above over a 100×100 grid, but
 425 using the lower spatial resolution of model properties parameterized in \mathbf{m} . The gradi-
 426 ents for velocity models are calculated by tracing rays backwards from receiver to (vir-
 427 tual) source using the gradient of the travel time field for each receiver pair (Rawlinson
 428 & Sambridge, 2004). For ADVI, the initial mean of the Gaussian distribution in the trans-

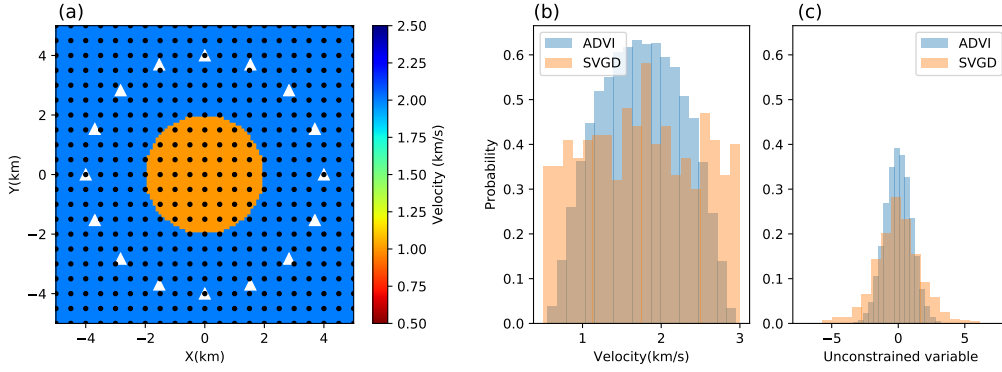


Figure 4. (a) The true velocity model and receivers (white triangle) used in the synthetic test. Sources are at the same locations as receivers to simulate a typical ambient noise experiment. Black dots indicate the locations of grid points used in the inversions. The histograms show the initial distributions of a parameter in the (b) original space (velocity) and (c) transformed unconstrained space for ADVI (blue) and SVGD (orange). In ADVI, the initial distribution is a standard Gaussian in unconstrained space. For simplicity we generated 5000 samples from the standard Gaussian and transformed to the original space to show the initial distribution in the original space. In SVGD the initial distribution is approximated using 800 particles generated from a Uniform distribution in the original space and transformed to the unconstrained space.

429 formed space is chosen to be the value which is the transform of the mean value of the
 430 prior in the original space; the initial covariance matrix is simply set to be an identity
 431 matrix, which turns out to give a standard Gaussian in our case (see blue histogram in
 432 Figure 4c). The shape of the initial distribution in the original space is shown in Fig-
 433 ure 4b (blue histogram). We then used 10,000 iterations to update the variational pa-
 434 rameters (μ and Σ). In order to visualize the results, we generated 5,000 models from
 435 the final approximate posterior probability density in the original space and computed
 436 their mean and standard deviation. For SVGD, we used 800 particles generated from the
 437 prior pdf (orange histogram in Figure 4b) and transformed to an unconstrained space
 438 using equation 11 (orange histogram in Figure 4c). Each particle is then updated using
 439 equation (17) for 500 iterations, then transformed back to seismic velocity. The mean
 440 and standard deviation are then calculated using the values of those particles.

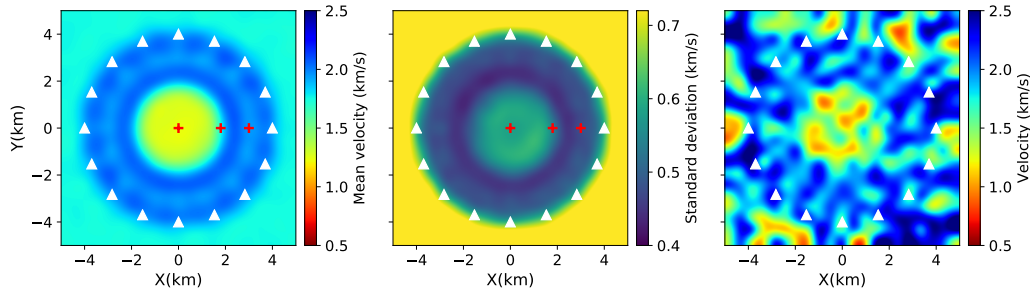


Figure 5. The mean (left), standard deviation (middle) and an individual realization from the approximate posterior distribution (right) obtained using ADVI. The red pluses show locations which are referred to in the main text.

441 To demonstrate the variational methods we compare the results with the fixed-dimensional
 442 Metropolis-Hastings MCMC (MH-McMC) method (Metropolis & Ulam, 1949; Hastings,
 443 1970; Mosegaard & Tarantola, 1995; Malinverno et al., 2000) and the rj-McMC method
 444 (Green, 1995; Bodin & Sambridge, 2009; Galetti et al., 2015; Zhang et al., 2018). For
 445 MH-McMC inversion we used the same parameterization as for the variational methods
 446 (a 21×21 grid). A Gaussian perturbation is used as the proposal distribution used to
 447 generate potential MCMC samples, for which the step length is chosen by trial and er-
 448 ror to give an acceptance ratio between 20 and 50 percent. We used a total of 6 chains,
 449 each of which used 2,000,000 iterations with a burn-in period of 1,000,000 iterations. To
 450 reduce the correlation between samples we only retain every 50^{th} sample in each chain
 451 after the burn-in period. The mean and standard deviation are then calculated using those
 452 samples. For rj-McMC inversion we use Voronoi cells to parameterize the model (Bodin
 453 & Sambridge, 2009), for which the prior pdf of the number of cells is set to be a Uni-
 454 form distribution between 4 and 100. The proposal distribution for fixed-dimensional steps
 455 (changing the velocity of a cell or moving a cell) is chosen in a similar way as in MH-
 456 McMC. For trans-dimensional steps (adding or deleting a cell) the proposal distribution
 457 is chosen as the prior pdf (Zhang et al., 2018). We used a total of 6 chains, each of which
 458 contained 500,000 iterations with a burn-in period of 300,000. Similarly to the fixed-dimensional
 459 inversion the chain was thinned by a factor of 50 post burn-in.

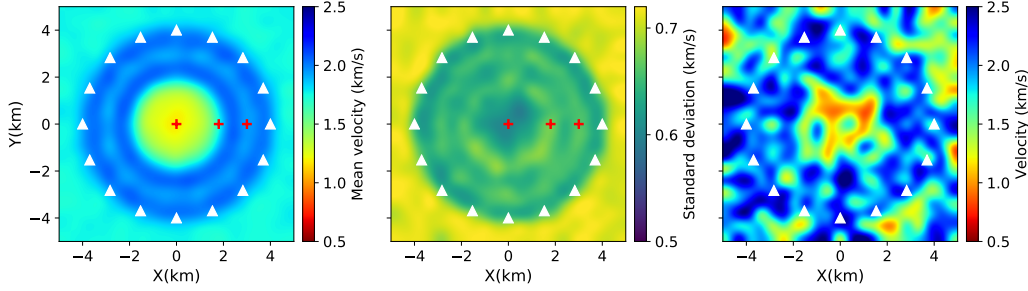


Figure 6. The mean (left), standard deviation (middle) and an individual realization from the approximate posterior distribution (right) obtained using SVGD. The red pluses show locations which are referred to in the main text.

3.1 Results

Figure 5 shows the mean, standard deviation and an individual realization from the approximate posterior distribution calculated using ADVI. The mean model successfully recovers the low velocity anomaly within the receiver array except that the velocity value is slightly higher ($\sim 1.2 \text{ km/s}$) than the true value (1.0 km/s). Between the location of the central anomaly and that of the receiver array there is a slightly lower velocity loop. The standard deviation map shows standard deviations similar to that of the prior (0.72 km/s) outside of the array, and clearly higher uncertainties at the location of the central anomaly. The standard deviations around the central anomaly are slightly higher than those at the center. Figure 6 shows the results from SVGD. Similarly, the velocity of the low velocity anomaly ($\sim 1.2 \text{ km/s}$) is slightly higher than the true value and a slightly lower velocity loop is also observed between the central anomaly and the receiver array. There is a clear higher uncertainty loop around the central anomaly; this has been observed previously and represent uncertainty due to the trade-off between the velocity of the anomaly and its shape (Galetti et al., 2015; Zhang et al., 2018). There is also another higher uncertainty loop associated with the lower velocity loop between the central anomaly and the receiver array. In contrast to this result, the loop cannot be observed in the results of ADVI.

To validate and better understand these results, Figure 7 shows the results from MH-McMC. The mean velocity model is very similar to the results from ADVI and SVGD. For example, the velocity value of the low velocity anomaly is higher than the true value,

481 which suggests that the mean value of the posterior under the specified parameteriza-
 482 tion is genuinely biased towards higher values than the true value. A lower velocity loop
 483 is also observed between the circular anomaly and the receiver array. The standard de-
 484 viation map shows similar results to those from SVGD: there is a higher uncertainty loop
 485 around the central anomaly and another one associated with the lower velocity loop be-
 486 tween the circular anomaly and the receiver array. The latter loop suggests that this area
 487 is not well constrained by the data, and therefore the mean velocity tends towards the
 488 mean value of the prior which is lower than the true value. We do not observe the clear
 489 higher uncertainty loops in the result of ADVI which may be due to the Gaussian ap-
 490 proximation which is used to fit a non-Gaussian posterior. In Figure 8 we show the re-
 491 sults from rj-McMC. Compared to the results from the fixed-parameterization inversions,
 492 the mean velocity is a more accurate estimate of the true model and uncertainty across
 493 the model is also lower. For example, the middle low velocity anomaly has almost the
 494 same value as the true model and has standard deviation of only $\sim 0.3 \text{ km/s}$ compared
 495 to values significantly greater than 0.3 km/s for all other methods. Between the mid-
 496 dle anomaly and the receivers, the model is determined better than in the fixed-paramterization
 497 inversions (with a standard deviation smaller than 0.1 km/s). This is because in rj-McMC
 498 the model parameterization adapts to the data which usually results in a lower-dimensional
 499 parameter space due to the natural parsimony of the method. For example, the aver-
 500 age dimensionality of the parameter space in the rj-McMC inversion is around 10; for
 501 comparison the fixed-parameterization inversions all have dimensionality fixed to be 441.
 502 The standard deviation map from the rj-McMC also shows a clear higher uncertainty
 503 loop within the array around the low velocity anomaly, and high uncertainties outside
 504 of the array where there is no data coverage.

505 Note that individual models from fixed-parameterization inversions (ADVI, SVGD
 506 and MH-McMC) show complex structures because of their higher dimensionality and the
 507 simple Uniform prior distribution that we adopted (right panels in Figure 5, 6 and 7).
 508 This might not be appropriate since the real Earth may have a smoother structure (de
 509 Pasquale & Linde, 2016; Ray & Myer, 2019). In that case, more informative prior in-
 510 formation including some form of regularization might be used to produce smoother in-
 511 dividual models (MacKay, 2003).

512 The results in Figure 8 do not show the double-loop uncertainty structure that is
 513 observed in the SVGD and MH-McMC results. The rj-McMC method contains an im-

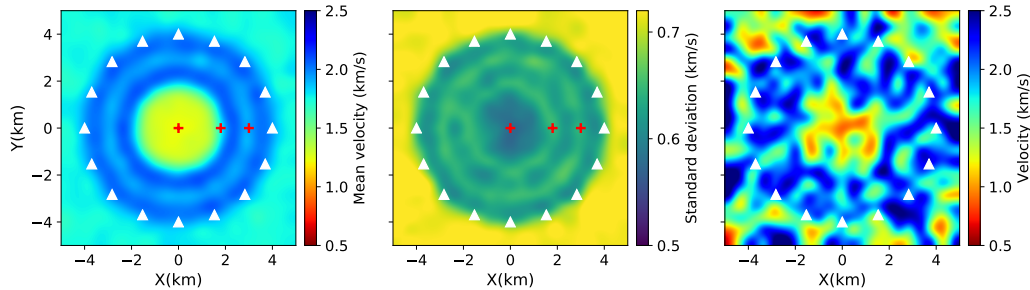


Figure 7. The mean (left), standard deviation (middle) and an individual realization from the approximate posterior distribution (right) obtained using MH-McMC. The red pluses show the point location which are referred to in the text.

514 plicit natural parsimony – the method tends to use fewer rather than more cells when-
 515 ever possible. While this may be useful in order to reduce the dimensionality of param-
 516 eter space, it is also possible that it causes some detailed features of the velocity or un-
 517 certainty structure to be omitted, much like a smoothing regularization condition in other
 518 tomographic methods. Since the double-loop structure appears to be a robust feature
 519 of the image uncertainty, we assume that the parsimony has indeed regularised some of
 520 the image structure out of the rj-McMC results.

521 Note that the result from rj-McMC is fundamentally different from results obtained
 522 using the fixed-parameterization inversions (ADVI, SVGD and MH-McMC) because of
 523 its entirely different parameterization. While the other inversion results are parameter-
 524 ized over a regular grid and can themselves be regarded as pixelated images, rj-McMC
 525 produces a set of models that are vectors containing positions and velocities of Voronoi
 526 cells, which can be transformed to an image on a regular grid (right panel in Figure 8).
 527 However, the Voronoi parametrization imposes prior restrictions on the pixelated form
 528 of models, for example all pixels within each Voronoi cell have identical velocities. As a
 529 result rj-McMC produces very different results to those obtained using the other meth-
 530 ods. In fact the choice of parameterization in rj-McMC can impose a variety of restric-
 531 tions on models, and different parameterizations can produce very different standard de-
 532 viation structures (Hawkins et al., 2019). Thus the results of rj-McMC must always be
 533 interpreted in the light of the specific prior information imposed by the parameteriza-
 534 tion deployed.

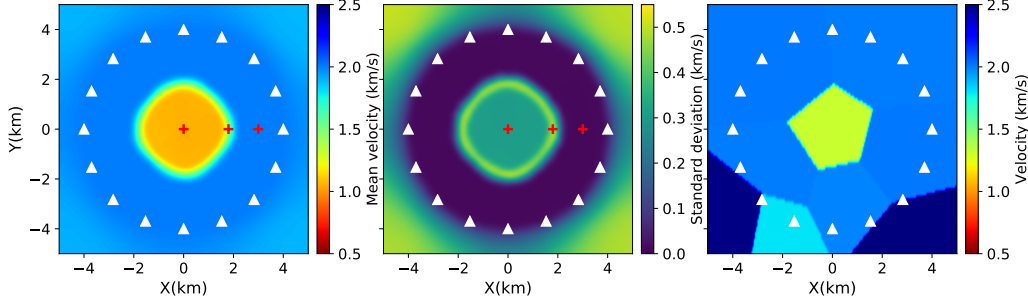


Figure 8. The mean (left), standard deviation (middle) and an individual realization from the approximate posterior distribution (right) obtained using trans-dimensional rj-McMC. The red pluses show the point location which are referred to in the text.

535 To further analyse the results, in Figure 9 we show marginal probability distribu-
 536 tions from the different inversion methods at three points (plus signs in Figure 5, 6, 7,
 537 and 8): point (0, 0) at the middle of the model, point (1.8, 0) at the boundary of the low
 538 velocity anomaly which has higher uncertainties, and point (3, 0) which also has higher
 539 uncertainties in the results from SVGD and MH-McMC. Due to symmetries of the model,
 540 marginal distributions at these three points are sufficient to reflect much of the entire
 541 set of single-parameter marginal probability distributions. At point (0, 0), the three fixed-
 542 parameterization methods produce similar marginal probability distributions. However,
 543 the marginal distribution from rj-McMC is narrower and concentrates around the true
 544 solution (1.0 km/s). This is likely due to the fact that in rj-McMC we have a much smaller
 545 parameter space than in the fixed-parameterization inversions. To assess the convergence
 546 we show the marginal distributions obtained by doubling the number of iterations in ADVI
 547 and SVGD with an red line in Figure 9a and b. The results show that increasing iter-
 548 ations only slightly improves the marginal distributions, suggesting that they have nearly
 549 converged. The black line in Figure 9b shows the marginal distribution obtained using
 550 more particles (1,600) with the same number of iterations (500). The result is almost
 551 the same as the result obtained using the original set of particles which suggests that 800
 552 particles are sufficient in this case. At point (1.8, 0), the marginal distributions from the
 553 three fixed-parameterization inversions become broader which explains the higher un-
 554 certainty loops observed in the standard deviation maps. The distribution from ADVI
 555 is more centrally focussed than the other two, which is again suggestive of the limita-

556 tions of that method caused by the Gaussian approximation. The distributions from SVGD
 557 and MH-McMC are more similar to each other and are close to the prior – a Uniform
 558 distribution – which suggests that the area is not well constrained by the data. By con-
 559 trast, the result from rj-McMC shows a clearly multimodal distribution with one mode
 560 centred around the velocity of the anomaly (1 *km/s*) and the other around the background
 561 velocity (2 *km/s*) as discussed in Galetti et al. (2015). This multimodal distribution re-
 562 flects the fact that it is not clear whether this point is inside or outside of the anomaly
 563 which produces the higher uncertainty loop in the standard deviation map. This sug-
 564 gests that there are different causes of the higher uncertainty loops in the different mod-
 565 els. In the fixed-parameterization inversions (ADVI, SVGD and MH-McMC) the higher
 566 uncertainty loops are mainly caused by the low resolution of the data at the boundary
 567 of the low velocity anomaly which produces broader marginal distributions. In the rj-
 568 McMC inversion, the higher uncertainty loops are mainly caused by multimodality in
 569 the posterior pdf. At point (3.0, 0) similarly to the point (0, 0), the marginal distribu-
 570 tions from the three fixed-parameterization inversions have similar shape and are much
 571 broader than the result from rj-McMC. Compared to the results from SVGD and MH-
 572 McMC, the result from ADVI again shows a more centrally-focussed distribution rem-
 573 iniscent of the Gaussian limitation implicit in ADVI. In the result of rj-McMC the marginal
 574 distribution concentrates to a very narrow distribution around the true value. Overall
 575 the marginal distributions from the fixed-parameterization inversions are broader than
 576 the result from rj-McMC due to their far larger parameter space. Note that although
 577 the marginal distributions from SVGD and MH-McMC have slightly different shape which
 578 causes differences in the magnitudes of their standard deviation maps, the maps are es-
 579 sentially similar from these quite different methods which suggests that the results are
 580 (approximately) correct.

581 **3.2 Computational cost**

582 Table 1 summarises the computational cost of the different methods. ADVI involves
 583 10,000 forward simulations which takes 0.45 CPU hours. However, note that in ADVI
 584 we used the full-rank covariance matrix which becomes huge in high dimensional param-
 585 eter spaces which could makes the method inefficient. SVGD involves 400,000 forward
 586 simulations which takes 8.53 CPU hours. This appears to make it less efficient than ADVI,
 587 however SVGD can produce a more accurate approximation to the posterior pdf than

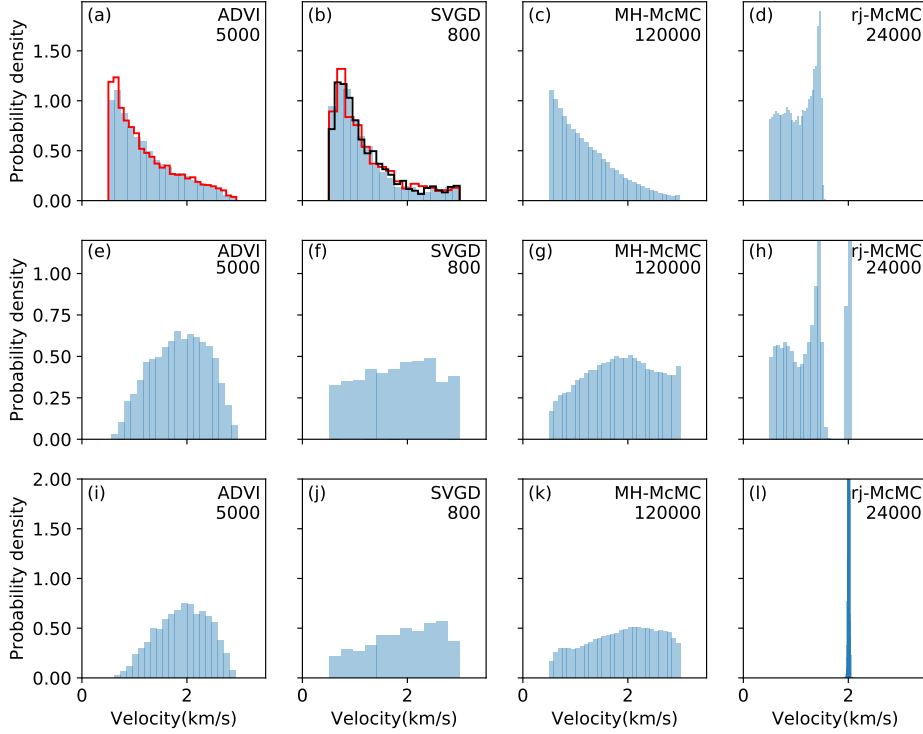


Figure 9. The marginal posterior pdfs of velocity at three points (pluses in Figure 3,4,5,6) derived using different methods. **(a)**, **(b)**, **(c)** and **(d)** show the marginal posterior distributions of velocity at the point (0,0) from ADVI, SVGD, MH-McMC and rj-McMC respectively. **(e)**, **(f)**, **(g)** and **(h)** show the marginal distributions at the point (1.8,0) from the four methods respectively, and **(i)**, **(j)**, **(k)** and **(l)** show the marginal distributions at the point (3,0) from the four methods respectively. The red lines in **(a)** and **(b)** are marginal distributions obtained by doubling the number of iterations and the black line in **(b)** shows the marginal distribution obtained using 1,600 particles. The number at the top-right of each figure shows the number of Monte Carlo samples.

588 ADVI which is limited by the Gaussian approximation. Note that SVGD can easily be
 589 parallelized by computing the gradients in equation (19) in parallel, making the method
 590 more time-efficient. For example, the above example takes 0.97 hours when parallelized
 591 using 10 cores. In comparison, MH-McMC requires 2,000,000 simulations for one chain
 592 which takes about 80.05 CPU hours, so for all 6 chains it requires 480.3 CPU hours in
 593 total. The rj-McMC run involved 500,000 simulations for one chain which takes about
 594 17.1 CPU hours, so 102.6 CPU hours in total for 6 chains. The Monte Carlo methods
 595 use evaluations of the likelihood and prior distribution at each sample whereas both vari-
 596 ational methods also deploy the information in the various gradients in equations 9, 10
 597 and 19. The number of simulations is therefore not a good metric to compare the four
 598 methods, since the gradients in this case are calculated by ray tracing which require more
 599 calculations per simulation in Table 1 compared to MC. CPU hours is a fairer metric for
 600 comparison, but of course this depends on the mechanism by which gradients are obtained:
 601 in other forward or inverse problems it is even possible that the variational methods take
 602 longer than Monte Carlo if estimating gradients requires extensive computation.

603 In the comparison in Table 1, rj-McMC is more efficient than MH-McMC due to
 604 the fact that rj-McMC explores a much smaller parameter space than the fixed param-
 605 eterization in MH-McMC. However, note that this might not always be true since trans-
 606 dimensional steps in rj-McMC usually have a very low probability of being accepted (Bodin
 607 & Sambridge, 2009; Zhang et al., 2018) and the method is generally significantly more
 608 difficult to tune (Green & Hastie, 2009). Overall, obtaining solutions from variational
 609 methods (ADVI, SVGD) is more efficient than Monte Carlo methods since they turn the
 610 Bayesian inference problem into an optimization problem. This also makes variational
 611 inference methods applicable to larger-datasets, and offers the advantage that very large
 612 datasets can be divided into random minibatches and inverted using stochastic optimiza-
 613 tion (Robbins & Monro, 1951; Kubrusly & Gravier, 1973) together with distributed com-
 614 putation. Monte Carlo methods are very computationally expensive for large datasets.
 615 Of course, the above comparison depends on the methods used to assess convergence for
 616 each method, which introduces some subjectivity in the comparison so that the abso-
 617 lute time required by each method may not be entirely accurate. Nevertheless, from all
 618 tests that we have conducted it is clear that variational methods produce solutions far
 619 more efficiently than Metropolis-Hastings and rj-McMC methods. Note that some other
 620 Monte Carlo sampling methods, e.g. Hamiltonian Monte Carlo, also use gradient infor-

Table 1. The comparison of computational cost for all 4 methods

Method	Number of simulations	CPU hours
ADVI	10,000	0.45
SVGD	400,000	8.53
MH-McMC	12,000,000	480.3
rj-McMC	3,000,000	102.6

621 mation and may be more efficient than Metropolis-Hastings methods (Neal et al., 2011;
 622 Sen & Biswas, 2017; Fichtner et al., 2018).

623 4 Application to Grane field

624 The Grane field is situated in the North sea, and contains a permanent monitor-
 625 ing system composed of 3458 four-component sensors measuring 3 orthogonal compo-
 626 nents of particle velocity and water pressure variations due to passing seismic waves. Zhang
 627 et al. (2019) used beamforming to show that the noise sources measured in the Grane
 628 field are nearly omnidirectional, which allows us to use ambient seismic noise tomogra-
 629 phy to study the subsurface of the field. To reduce the computational cost, in this study
 630 we down-sampled the number of receivers by a factor of 10 which results in 346 receivers,
 631 and we only used 35 receivers as virtual sources (Figure 10a). Cross-correlations are com-
 632 puted between vertical component recordings at pairs consisting of a virtual source and
 633 a receiver using half-hour time segments, and the set of correlations for each pair were
 634 stacked over 6.5 hours. This process produces approximate virtual-source seismograms
 635 of Rayleigh-type Scholte waves (Campillo & Paul, 2003; Shapiro et al., 2005; Curtis et
 636 al., 2006). Phase velocity dispersion curves for each (virtual) source-receiver pair are then
 637 automatically picked using an image transformation technique: for all processing details
 638 see Zhang et al. (2019) which presents a complete ambient noise analysis of the field and
 639 presents tomographic phase velocity maps at various frequencies as well as estimated shear-
 640 velocity structure of the near seabed subsurface. Here we use the recording phase veloc-
 641 ity data at 0.9 s period.

642 We apply the variational inference methods ADVI and SVGD, and rj-McMC to the
 643 data to obtain phase velocity maps at 0.9 s and compare the results. For variational meth-

644 ods, the field is parametrized using a regular 26×71 grid with a spacing of 0.2 km at
 645 both x and y directions giving a velocity model dimensionality of 1846. Due to its com-
 646 putational cost in high dimensional spaces we do not apply MH-McMC. The data noise
 647 level is set to be 0.05 s, which is an average value estimated by the hierarchical Bayesian
 648 Monte Carlo inversion of Zhang et al. (2019). The prior pdf of phase velocity in each model
 649 cell is set to be a Uniform distribution between 0.35 km/s and 0.55 km/s , which is se-
 650 lected to be wider than the minimum (0.4 km/s) and maximum (0.5 km/s) phase veloc-
 651 ity picked from cross-correlations. The initial probability distribution for ADVI is cho-
 652 sen similarly to that in the synthetic tests: a standard Gaussian distribution in the un-
 653 constrained space (blue histogram in Figure 10c), and its shape in the original space is
 654 shown in Figure 10b (blue histogram). For SVGD, the initial distribution is approximated
 655 using 1000 particles generated from the prior in the original space (orange histogram in
 656 Figure 10b) and transformed to the unconstrained space (orange histogram in Figure 10c).
 657 We then applied 10,000 iterations for ADVI and 500 iterations for SVGD. Similarly to
 658 the synthetic test above for rj-McMC we use Voronoi cells to parameterize the model.
 659 The prior pdf of the number of cells is set to be a discrete Uniform distribution between
 660 30 and 200, and the data noise level is estimated hierarchically during the inversion (Zhang
 661 et al., 2018). Proposal distributions are the same as in the synthetic test above. We used
 662 a total of 16 chains, each of which contains 800,000 iterations including a burn-in period
 663 of 400,000. To reduce the correlation between samples we only retain every 50th sam-
 664 ple post burn-in for our final ensemble.

665 Figure 11 shows the mean and standard deviation maps from ADVI. The mean phase
 666 velocity map shows a clear low velocity anomaly around the centre of the field from $Y=6$
 667 km to $Y=10$ km and another at the western edge between $Y=8$ km and $Y=10$ km. These
 668 were also observed by (Zhang et al., 2019) using Eikonal tomography, who showed that
 669 they are correlated with areas of higher density of pockmarks on the seabed, suggest-
 670 ing that they are caused by near surface fluid flow effects. At the western edge between
 671 $Y=6$ km and $Y=8$ km and at the northwestern edge there are high velocity anomalies
 672 which were also observed in the results of Zhang et al. (2019). In the north between $Y=11$
 673 km and $Y=12$ km and along the eastern edge between $Y=7$ km and $Y=10$ km the model
 674 shows some low velocity anomalies. Moreover, there are some small anomalies distributed
 675 across the field. For example, to the south of the central low velocity anomaly around
 676 $Y=6$ km there are several other low velocity anomalies. Similarly there is a small low

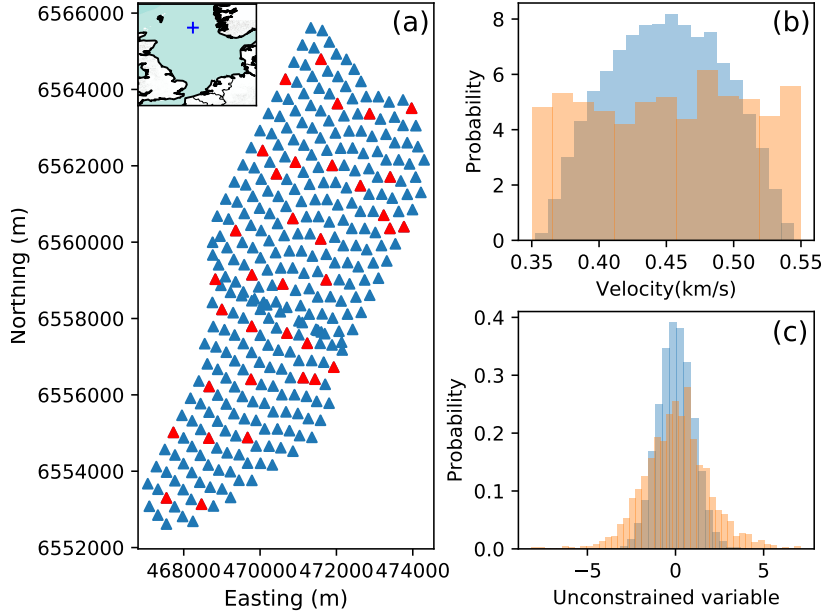


Figure 10. (a) The distribution of receiver (blue and red triangles) across the Grane field used in this study. Red triangles show the receivers that were used as virtual sources. The blue plus in the inset map shows the location of Grane field. The histograms show the initial distributions of a parameter in the (b) original (velocity) space and (c) transformed unconstrained space for ADVI (blue) and for SVGD (orange). Similar to Figure 4, we used 5000 Monte Carlo samples to show probability distributions in both the original and the unconstrained space for ADVI. The initial distribution for SVGD is approximated using 1000 particles generated from the prior (a Uniform distribution) in the original space and transformed to the unconstrained space.

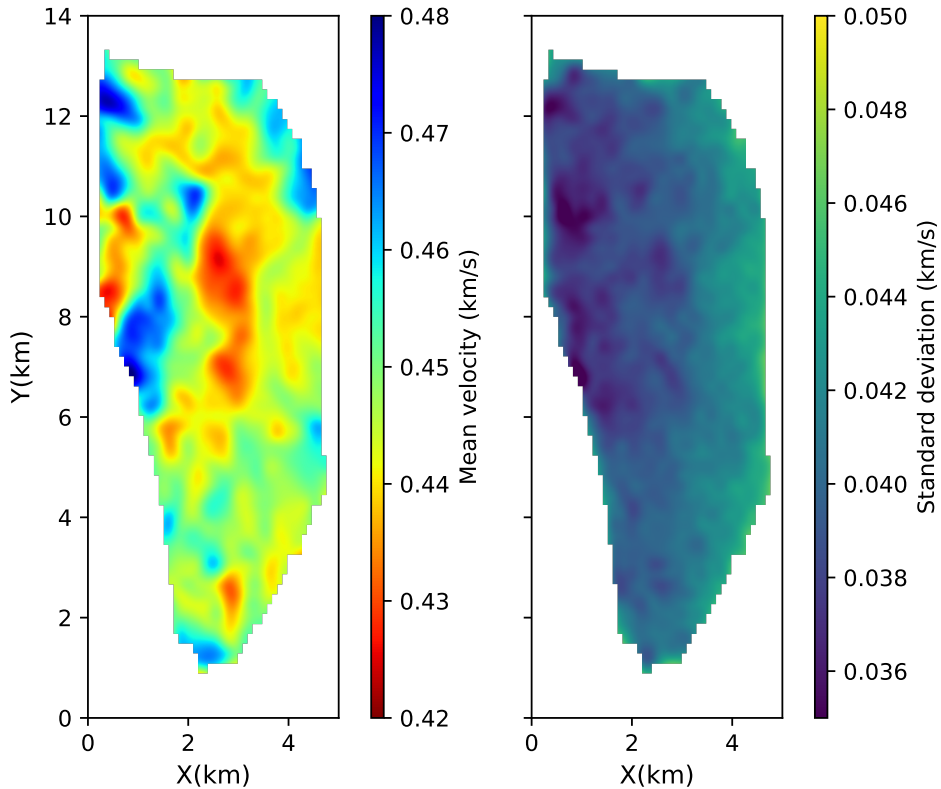


Figure 11. The mean (left) and standard deviation map (right) from ADVI.

677 velocity anomaly and a small high velocity anomaly in the south of the field around $Y=2.5$
 678 km, and a small high velocity anomaly in the north around $Y=10.5$ km.

679 Overall the standard deviation map shows that uncertainty in the west is lower than
 680 in the east. At the western edge there are some low uncertainty areas which are asso-
 681 ciated with velocity anomalies. For example, the low uncertainty area between $Y=6$ km
 682 and $Y=8$ km is associated with the high velocity anomaly at the same location. Sim-
 683 ilarly the high velocity anomaly at the northwestern edge around $Y=12$ km shows a lower
 684 uncertainty, and the middle low velocity anomaly also shows slightly lower uncertain-
 685 ties. This might suggest that these velocity structures are well-constrained by the data.
 686 However, in the synthetic tests we noticed that the ADVI can produce biased standard

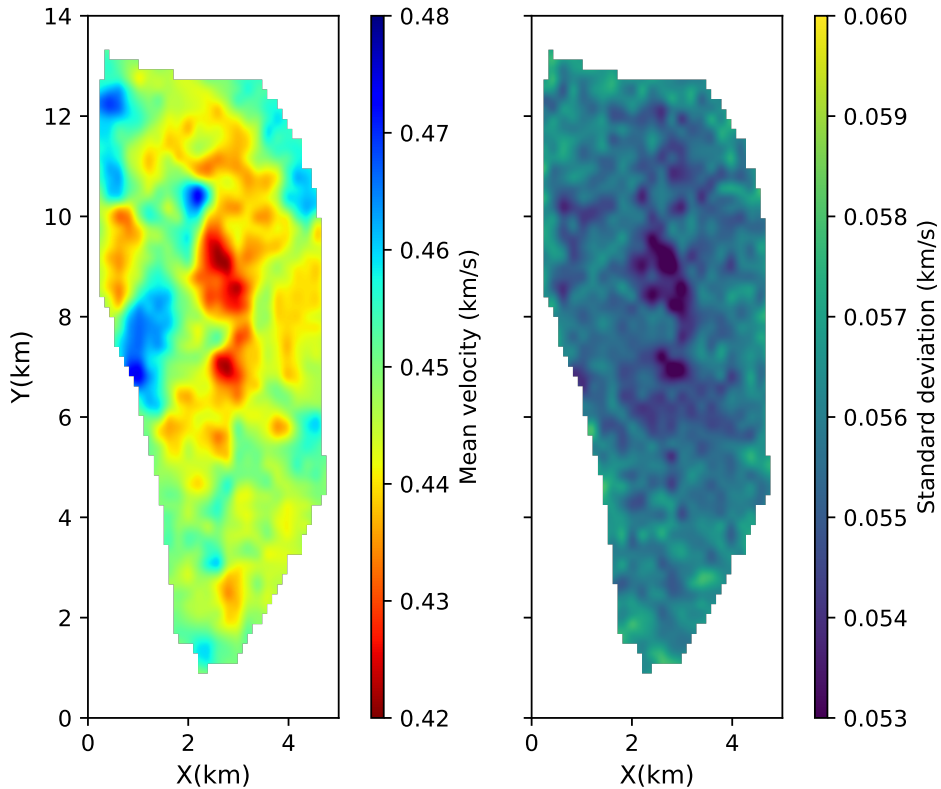


Figure 12. The mean (left) and standard deviation map (right) from SVGD.

687 deviation maps due to the Gaussian approximation, so these uncertainty properties may
 688 not be robust.

689 We show the mean and standard deviation maps obtained using SVGD in Figure
 690 12. The mean velocity map shows very similar structures to the result from ADVI, ex-
 691 cept that the velocity magnitudes are slightly different. For example, we observe the cen-
 692 tral low velocity anomaly and one at the western edge which appeared in the mean ve-
 693 locity map from ADVI and are related to the density distribution of pockmarks. Sim-
 694 ilarly there are high velocity anomalies at the western edge and a low velocity anomaly
 695 at the eastern edge. Even for more detailed structure, e.g., the low velocity anomalies
 696 at the north ($Y \gtrsim 10$ km), the low velocity anomalies around $Y=6$ km and the small ve-
 697 locity anomalies around $Y=2.5$ km, the two results show highly consistent properties be-

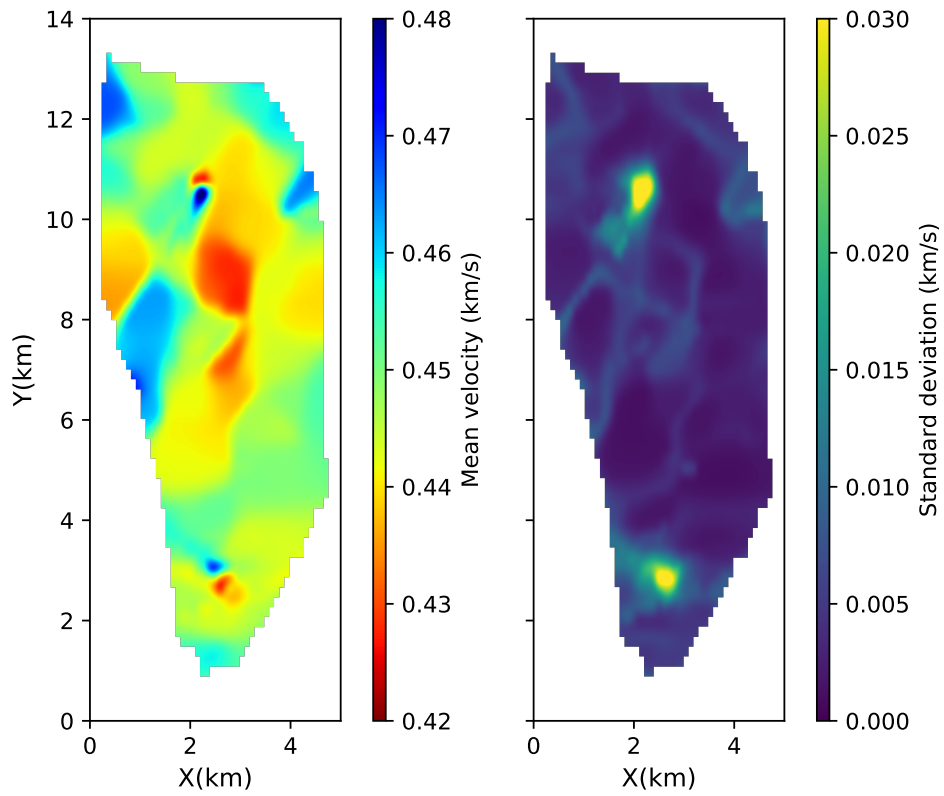


Figure 13. The mean (left) and standard deviation map (right) from rj-McMC.

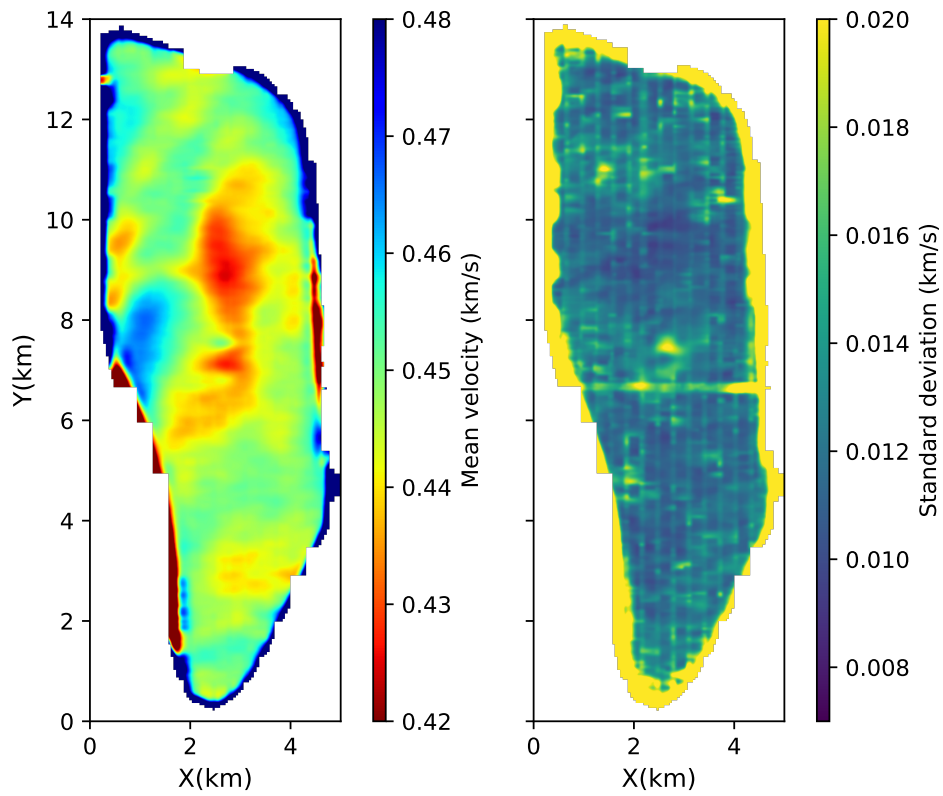


Figure 14. The mean (left) and standard deviation map (right) obtained using Eikonal tomography by Zhang et al. (2019).

698 tween the two methods. This suggests that we have obtained accurate mean phase ve-
699 locity maps given the fixed, gridded model parameterization and the observed data.

700 Despite the similarity in the mean results, the standard deviation map from SVGD
701 is quite different from the results from ADVI, which is consistent with similar variations
702 that we observed in the synthetic tests. For example, there is no clear magnitude dif-
703 ference between the west and the east as appeared in the result from ADVI. There is a
704 clear low uncertainty area associated with the central low velocity anomaly, which is slightly
705 lower in magnitude than the result from ADVI. Similarly there is a slightly lower un-
706 certainty area at the western edge associated with the low velocity anomaly at the same
707 location. The south-central low velocity anomaly around $Y=6$ km also exhibits relatively
708 lower uncertainties, which suggests that those small low velocity anomalies in this area
709 may reflect true properties of the subsurface. Similarly there are some low uncertainty
710 structures at the north around $Y=11$ km which are associated with low velocity anoma-
711 lies. Note that due to the Gaussian approximation in ADVI, the standard deviation re-
712 sults from SVGD show different magnitudes as we saw in the synthetic tests.

713 Figure 13 shows the mean and standard deviation maps obtained from rj-McMC.
714 The mean velocity map shows broadly similar structures to the results from ADVI and
715 SVGD. For example, we also observed the middle low velocity anomaly, the low veloc-
716 ity anomalies at the western and eastern edges and the high velocity anomalies at the
717 western edge. However, compared to the previous results these structures are smoother
718 which is probably caused by the natural parsimony that is implicit within the rj-McMC
719 inversion method (Green, 1995; Bodin & Sambridge, 2009) similarly to the synthetic tests
720 above. The small velocity anomalies in the previous results disappear in the result from
721 rj-McMC; this may also be caused by the natural parsimony of rj-McMC, or by overfit-
722 ting of data in the variational methods due to the fixed parameterization. However, the
723 small high and low velocity anomalies around $Y=2.5$ km and around $Y=10.5$ km still
724 exist, which suggests that these detailed velocity structures may represent real proper-
725 ties of the subsurface (or are caused by a consistent bias in the data).

726 Similarly to the synthetic tests, the standard deviation map from rj-McMC shows
727 significantly smaller uncertainties (< 0.01 km/s) than the results from ADVI (~ 0.04
728 km/s) and SVGD (~ 0.055 km/s), which is probably caused by a lower dimensionality
729 of parameter space used in rj-McMC (around 60 Voronoi cells were used) than in vari-

730 ational methods (1846), resulting in fewer trade-offs between parameters. However, there
731 are higher uncertainties at the location of the small velocity anomalies at $Y=2.5$ km and
732 at $Y=10.5$ km, which is probably due to the fact that not all chains found these small
733 structures.

734 To compare our results with traditional methods, Figure 14 shows the mean and
735 standard deviation maps obtained using Eikonal tomography by Zhang et al. (2019) us-
736 ing all of the available data (3458 virtual sources and 3458 receivers). The mean veloc-
737 ity model shows similar but slightly smoother structures compared to those obtained us-
738 ing ADVI and SVGD. This may be because the larger quantity of data used in Eikonal
739 tomography reduces the noise and stabilizes the results, or because the interpolation used
740 in Eikonal tomography regularizes (smooths) small scale structure. The standard devi-
741 ation map shows lower uncertainties at the location of the middle low velocity anomaly
742 which is similar to that obtained using SVGD. This again suggests that SVGD can pro-
743 duce a more accurate standard deviation estimate than ADVI. The mean velocity model
744 from rj-McMC shows smoother structures than that from Eikonal tomography, which
745 may suggest that rj-McMC omits small scale structure due to its implicit parsimony. The
746 standard deviation map from rj-McMC also does not show similar structures to those
747 obtained using Eikonal tomography or SVGD due to the completely different parame-
748 terizations employed.

749 In the inversion, ADVI involved 10,000 forward simulations which took 5.1 CPU
750 hours and SVGD involved 500,000 forward simulations which required 141.8 CPU hours.
751 By contrast the rj-McMC involved 12,800,000 forward simulations to obtain an accept-
752 able result which required 1,866.1 CPU hours. In real time, SVGD was in fact parallelised
753 using 12 cores which took 12.1 hours to run, while rj-McMC was parallelised using 16
754 cores which therefore took about 5 days. We conclude that, although the variational meth-
755 ods produce higher uncertainty estimates, they can produce similar parameter estimates
756 (mean velocity) at hugely reduced computational cost, and indeed our synthetic tests
757 suggest that the variational SVGD image uncertainty results may in fact be more cor-
758 rect.

5 Discussion

We have shown that variational methods (ADVI and SVGD) can be applied to seismic tomography problems and provide efficient alternatives to MCMC. ADVI produces biased posterior pdfs because of its implicit Gaussian approximation, and cannot be applied to problems with multi-modal posteriors. However, it still generates an accurate estimate of the mean model. Given that it is very efficient (only requiring 10,000 forward simulations) the method could be useful in scenarios where efficiency is important and a Gaussian approximation is sufficient for uncertainty analysis. Alternatively a mixture of Gaussians approximation might be used to improve the accuracy of the algorithm (Zobay et al., 2014; Arenz et al., 2018). In a very high dimensional case, ADVI could become less efficient because of the increased size of the Gaussian covariance matrix. In that case one could use a mean-field approximation (setting model covariances to zero), or use a sparse covariance matrix to reduce computational cost since seismic velocity in any cell is often most strongly correlated with that in neighbouring cells.

SVGD can produce a good approximation to posterior pdfs. However, since it is based on a number of particles, the method is more computationally costly than ADVI. In this study we parallelized the computation of gradients to improve the efficiency, and for large datasets further improvements can be obtained by using random minibatches to perform the inversion (Q. Liu & Wang, 2016). Such a strategy can be applied to any variational inference method (e.g. also ADVI) since variational methods solve an optimization rather than a stochastic sampling problem. In comparison, this strategy cannot easily be used in MCMC based methods since it may break the detailed balance requirement of MCMC (Blei et al., 2017). Though it has been shown that SVGD requires fewer particles than particle-based sampling methods (e.g., sequential Monte Carlo) in the sense that it reduces to finding the MAP model if only one particle is used, the optimal choice of the number of particles remains unclear, especially for very high dimensional spaces. In the case of very high dimensionality another possibility is to use normalizing flows – a variational method based on a series of specific invertible transforms (Rezende & Mohamed, 2015).

Monte Carlo and variational inference are different types of methods that solve the same problem. Monte Carlo simulates a set of Markov chains and uses samples of those chains to approximate the posterior pdf, while variational inference solves an optimiza-

791 tion problem to find the closest pdf to the posterior within a given family of probabil-
792 ity distributions. Monte Carlo methods provide guarantees that samples are asymptot-
793 ically distributed according to the posterior pdf as the number of samples tends to in-
794 finity (Robert & Casella, 2013), while the statistical properties of variational inference
795 algorithms are still unknown (Blei et al., 2017). It is possible to combine the two meth-
796 ods to capitalise on the merits of both. For example, the approximate posterior pdf from
797 an efficient variational method (e.g. ADVI) can be used as a proposal distribution for
798 Metropolis-Hastings (De Freitas et al., 2001) to improve the efficiency of McMC, or McMC
799 steps can be integrated to the variational approximation to improve the accuracy of vari-
800 ational methods (Salimans et al., 2015).

801 We used a fixed regular grid of cells to parameterize the tomographic model in the
802 variational methods, which might introduce overfitting of the data. For example, the mean
803 velocity models in the synthetic tests show a slightly lower velocity loop between the low
804 velocity anomaly and the receivers, and the uncertainties obtained from fixed-parameterization
805 inversions are significantly higher than the results from rj-McMC. However, although rj-
806 McMC produces lower uncertainty estimates, small scale structures can be omitted in
807 the results of rj-McMC due to their implicitly imposed parsimony. For example, in our
808 real data example, small scale structures in the results of variational inference methods
809 and Eikonal tomography are smoothed out in the results of rj-McMC. Indeed the param-
810 eterization used in rj-McMC imposes restrictions on models, and different parameter-
811 izations can produce different uncertainties (Hawkins et al., 2019). This makes the in-
812 terpretation and use of uncertainties from rj-McMC difficult.

813 It is not easy to determine an optimal grid in variational inference methods since
814 this introduces a trade off between resolution of the model and overfitting of the data.
815 Therefore, it might be necessary to use a more flexible parameterization, e.g., Voronoi
816 cells (Bodin & Sambridge, 2009; Zhang et al., 2018) or wavelet parameterization (Fang
817 & Zhang, 2014; Hawkins & Sambridge, 2015; Zhang & Zhang, 2015). It may also be pos-
818 sible to apply a series of different parameterizations and select the best one using model
819 selection theory (Walter & Pronzato, 1997; Curtis & Snieder, 1997; Arnold & Curtis, 2018).
820 Note that it would make the methods less computationally efficient to find an optimal
821 parameterization because we may need to run a series of optimization problems with dif-
822 ferent parameterizations. However, in cases with very large datasets which may more
823 suitably be solved by variational inference methods, it might instead be sufficient to use

824 a parameterization with the highest resolution that the frequency of the data could re-
825 solve. Instead some more informative prior or regularization may be used to reduce the
826 magnitude of uncertainty estimates and to better constrain the model (MacKay, 2003;
827 Ray & Myer, 2019).

828 In our experiments the results from *rj-McMC* are significantly different from the
829 results obtained using variational methods or *MH-McMC*. This is essentially caused by
830 different parameterizations. In *ADVI*, *SVGD* and *MH-McMC* we invert for a pixelated
831 image, while in *rj-McMC* we invert for a distribution of parameters that represent lo-
832 cations and shapes of cells and their constant velocities, the pointwise spatial mean of
833 which is visualized as an image. Therefore even though we visualized them in the same
834 way, the results are essentially not directly comparable. Nevertheless, the comparison
835 with *rj-McMC* is interesting because until now a quite different alternative probabilis-
836 tic method was never used to estimate the posterior of images from the same realistic
837 tomography problem. The results here demonstrate that the *rj-McMC* method as ap-
838 plied in most tomography papers gives significantly different solutions than we might pre-
839 viously have thought; specifically, it does not produce the posterior distribution of the
840 pixelated image that is usually shown in scientific papers (e.g., Bodin & Sambridge, 2009;
841 Zulfakriza et al., 2014; Galetti et al., 2015; Crowder et al., 2019). Rather, it samples a
842 probability distribution in a particular irregular and variably parametrized model space
843 and results should be interpreted as such. Note that some other methods, e.g. *rj-McMC*
844 with Gaussian processes, may provide results that can be compared between all sampling
845 methods, and provide a means of injecting prior information with adaptable complex-
846 ity into the sampling scheme (Ray & Myer, 2019).

847 In this study we used a fixed data noise level in the variational methods. It has been
848 shown that an improper noise level can introduce biases in tomographic results (Bodin
849 & Sambridge, 2009; Zhang et al., 2019), so in our example we used the noise level esti-
850 mated by hierarchical *McMC*. It can also be estimated by a variety of other methods (Bensen
851 et al., 2009; Yao & Van Der Hilst, 2009; Weaver et al., 2011; Nicolson et al., 2012, 2014),
852 and maximum likelihood methods (Sambridge, 2013; Ray et al., 2016; Ray & Myer, 2019).
853 In future it might also be possible to include the noise parameters in variational meth-
854 ods in a hierarchical way.

855 In this study we applied variational inference methods to simple 2D tomography
856 problems, but it is straightforward to apply the methods to any geophysical inverse prob-
857 lems whose gradients with respect to the model can be computed efficiently. For exam-
858 ple, variational methods can be applied to 3D seismic tomography problems to provide
859 efficient approximation, which generally demands enormous computational resources us-
860 ing McMC methods (Hawkins & Sambridge, 2015; Zhang et al., 2018, 2019). The meth-
861 ods also provide possibilities to perform Bayesian inference for full waveform inversion,
862 which is generally very expensive for McMC (Ray et al., 2017) and suffers from noto-
863 rious multimodality in the likelihoods. SVGD provides a possible way to approximate
864 these complex distributions given that theoretically it can approximate arbitrary distri-
865 butions.

866 **6 Conclusion**

867 We introduced two variational inference methods to geophysical tomography – au-
868 tomatic differential variational inference (ADVI) and Stein variational gradient descent
869 (SVGD), and applied them to 2D seismic tomography problems using both synthetic and
870 real data. Compared to the Markov chain Monte Carlo (McMC) method, ADVI provides
871 an efficient but biased approximation to Bayesian posterior probability density functions,
872 and cannot be applied to find multi-modal posteriors because of its implicit Gaussian
873 assumption. In contrast, SVGD is slightly slower than ADVI but produces a more ac-
874 curate approximation. The real data example shows that ADVI and SVGD produce very
875 similar mean velocity models, even though their uncertainty estimates are different . The
876 mean velocity models are very similar to those produced by reversible jump McMC (rj-
877 McMC), except that the mean model from rj-McMC is smoother because of the much
878 lower dimensionality of its parameter space. Variational methods thus can provide ef-
879 ficient approximate alternatives to McMC methods, and can be applied to many geo-
880 physical inverse problems.

881 **Acknowledgments**

882 The authors would like to thank the Grane license partners Equinor ASA, Petoro AS,
883 ExxonMobil E&P Norway AS, and ConocoPhillips Skandinavia AS for allowing us to
884 publish this work. The views and opinions expressed in this paper are those of the au-
885 thors and are not necessarily shared by the license partners. The authors thank the Ed-

886 Edinburgh Interferometry Project sponsors (Schlumberger, Equinor and Total) for support-
887 ing this research. This work used the Cirrus UK National Tier-2 HPC Service at EPCC
888 (<http://www.cirrus.ac.uk>). The data used in this study are available at Edinburgh DataShare
889 (<https://doi.org/10.7488/ds/2607>).

890 References

- 891 Aki, K., & Lee, W. (1976). Determination of three-dimensional velocity anomalies
892 under a seismic array using first P arrival times from local earthquakes: 1. a
893 homogeneous initial model. *Journal of Geophysical research*, *81*(23), 4381–
894 4399.
- 895 Arenz, O., Zhong, M., & Neumann, G. (2018). Efficient gradient-free variational
896 inference using policy search. In *International conference on machine learning*
897 (pp. 234–243).
- 898 Arnold, R., & Curtis, A. (2018). Interrogation theory. *Geophysical Journal Interna-*
899 *tional*, *214*(3), 1830–1846.
- 900 Bensen, G., Ritzwoller, M., & Yang, Y. (2009). A 3-D shear velocity model of the
901 crust and uppermost mantle beneath the United States from ambient seismic
902 noise. *Geophysical Journal International*, *177*(3), 1177–1196.
- 903 Bishop, C. M. (2006). *Pattern recognition and machine learning*. springer.
- 904 Blatter, D., Key, K., Ray, A., Gustafson, C., & Evans, R. (2019). Bayesian joint
905 inversion of controlled source electromagnetic and magnetotelluric data to im-
906 age freshwater aquifer offshore new jersey. *Geophysical Journal International*,
907 *218*(3), 1822–1837.
- 908 Blei, D. M., Kucukelbir, A., & McAuliffe, J. D. (2017). Variational inference:
909 A review for statisticians. *Journal of the American Statistical Association*,
910 *112*(518), 859–877.
- 911 Bodin, T., & Sambridge, M. (2009). Seismic tomography with the reversible jump
912 algorithm. *Geophysical Journal International*, *178*(3), 1411–1436.
- 913 Bodin, T., Sambridge, M., Tkalčić, H., Arroucau, P., Gallagher, K., & Rawlinson,
914 N. (2012). Transdimensional inversion of receiver functions and surface wave
915 dispersion. *Journal of Geophysical Research: Solid Earth*, *117*(B2).
- 916 Burdick, S., & Lekić, V. (2017). Velocity variations and uncertainty from transdi-
917 mensional P-wave tomography of North America. *Geophysical Journal Interna-*

- 918 *tional*, 209(2), 1337–1351.
- 919 Campillo, M., & Paul, A. (2003). Long-range correlations in the diffuse seismic coda.
920 *Science*, 299(5606), 547–549.
- 921 Çınlar, E. (2011). *Probability and stochastics* (Vol. 261). Springer Science & Busi-
922 ness Media.
- 923 Crowder, E., Rawlinson, N., Pilia, S., Cornwell, D., & Reading, A. (2019). Trans-
924 dimensional ambient noise tomography of Bass Strait, southeast Australia,
925 reveals the sedimentary basin and deep crustal structure beneath a failed
926 continental rift. *Geophysical Journal International*, 217(2), 970–987.
- 927 Curtis, A., Gerstoft, P., Sato, H., Snieder, R., & Wapenaar, K. (2006). Seismic inter-
928 ferometry – turning noise into signal. *The Leading Edge*, 25(9), 1082–1092.
- 929 Curtis, A., & Lomax, A. (2001). Prior information, sampling distributions, and the
930 curse of dimensionality. *Geophysics*, 66(2), 372–378.
- 931 Curtis, A., & Snieder, R. (1997). Reconditioning inverse problems using the genetic
932 algorithm and revised parameterization. *Geophysics*, 62(5), 1524–1532.
- 933 Curtis, A., & Snieder, R. (2002). Probing the earth’s interior with seismic tomogra-
934 phy. *International Geophysics Series*, 81(A), 861–874.
- 935 De Freitas, N., Højen-Sørensen, P., Jordan, M. I., & Russell, S. (2001). Variational
936 MCMC. In *Proceedings of the seventeenth conference on uncertainty in artificial*
937 *intelligence* (pp. 120–127).
- 938 de Pasquale, G., & Linde, N. (2016). On structure-based priors in bayesian geophys-
939 ical inversion. *Geophysical Journal International*, 208(3), 1342–1358.
- 940 Detommaso, G., Cui, T., Marzouk, Y., Spantini, A., & Scheichl, R. (2018). A stein
941 variational newton method. In *Advances in neural information processing sys-*
942 *tems* (pp. 9169–9179).
- 943 Devilee, R., Curtis, A., & Roy-Chowdhury, K. (1999). An efficient, probabilistic
944 neural network approach to solving inverse problems: Inverting surface wave
945 velocities for Eurasian crustal thickness. *Journal of Geophysical Research:*
946 *Solid Earth*, 104(B12), 28841–28857.
- 947 Dziewonski, A. M., & Woodhouse, J. H. (1987). Global images of the Earth’s inte-
948 rior. *Science*, 236(4797), 37–48.
- 949 Earp, S., & Curtis, A. (2019). Probabilistic neural-network based 2D travel time to-
950 mography. *arXiv preprint arXiv:1907.00541*.

- 951 Fang, H., & Zhang, H. (2014). Wavelet-based double-difference seismic tomography
 952 with sparsity regularization. *Geophysical Journal International*, *199*(2), 944–
 953 955.
- 954 Fichtner, A., Zunino, A., & Gebraad, L. (2018). Hamiltonian monte carlo solution
 955 of tomographic inverse problems. *Geophysical Journal International*, *216*(2),
 956 1344–1363.
- 957 Galetti, E., & Curtis, A. (2018). Transdimensional electrical resistivity tomography.
 958 *Journal of Geophysical Research: Solid Earth*, *123*(8), 6347–6377.
- 959 Galetti, E., Curtis, A., Baptie, B., Jenkins, D., & Nicolson, H. (2017). Transdimen-
 960 sional love-wave tomography of the British Isles and shear-velocity structure
 961 of the east Irish Sea Basin from ambient-noise interferometry. *Geophysical*
 962 *Journal International*, *208*(1), 36–58.
- 963 Galetti, E., Curtis, A., Meles, G. A., & Baptie, B. (2015). Uncertainty loops in
 964 travel-time tomography from nonlinear wave physics. *Physical review letters*,
 965 *114*(14), 148501.
- 966 Gorham, J., & Mackey, L. (2015). Measuring sample quality with Stein’s method. In
 967 *Advances in neural information processing systems* (pp. 226–234).
- 968 Gorham, J., & Mackey, L. (2017). Measuring sample quality with kernels. In *Pro-*
 969 *ceedings of the 34th international conference on machine learning-volume 70*
 970 (pp. 1292–1301).
- 971 Green, P. J. (1995). Reversible jump Markov chain Monte Carlo computation and
 972 Byesian model determination. *Biometrika*, 711–732.
- 973 Green, P. J., & Hastie, D. I. (2009). Reversible jump MCMC. *Genetics*, *155*(3),
 974 1391–1403.
- 975 Gretton, A. (2013). Introduction to RKHS, and some simple kernel algorithms.
- 976 Hastings, W. K. (1970). Monte Carlo sampling methods using Markov chains and
 977 their applications. *Biometrika*, *57*(1), 97–109.
- 978 Hawkins, R., Bodin, T., Sambridge, M., Choblet, G., & Husson, L. (2019). Trans-
 979 dimensional surface reconstruction with different classes of parameterization.
 980 *Geochemistry, Geophysics, Geosystems*, *20*(1), 505–529.
- 981 Hawkins, R., & Sambridge, M. (2015). Geophysical imaging using trans-dimensional
 982 trees. *Geophysical Journal International*, *203*(2), 972–1000.
- 983 Hoffman, M. D., & Blei, D. M. (2015). Structured stochastic variational inference.

- 984 In *Artificial intelligence and statistics*.
- 985 Iyer, H., & Hirahara, K. (1993). *Seismic tomography: Theory and practice*. Springer
986 Science & Business Media.
- 987 Karlin, S. (2014). *A first course in stochastic processes*. Academic press.
- 988 Käüfl, P., Valentine, A., de Wit, R., & Trampert, J. (2015). Robust and fast prob-
989 abilistic source parameter estimation from near-field displacement waveforms
990 using pattern recognition. *Bulletin of the Seismological Society of America*,
991 *105*(4), 2299–2312.
- 992 Käüfl, P., Valentine, A. P., O’Toole, T. B., & Trampert, J. (2013). A framework for
993 fast probabilistic centroid-moment-tensor determination – inversion of regional
994 static displacement measurements. *Geophysical Journal International*, *196*(3),
995 1676–1693.
- 996 Kingma, D. P., & Welling, M. (2013). Auto-encoding variational Byes. *arXiv*
997 *preprint arXiv:1312.6114*.
- 998 Kubrusly, C., & Gravier, J. (1973). Stochastic approximation algorithms and ap-
999 plications. In *1973 ieee conference on decision and control including the 12th*
1000 *symposium on adaptive processes* (pp. 763–766).
- 1001 Kucukelbir, A., Tran, D., Ranganath, R., Gelman, A., & Blei, D. M. (2017). Au-
1002 tomatic differentiation variational inference. *The Journal of Machine Learning*
1003 *Research*, *18*(1), 430–474.
- 1004 Kullback, S., & Leibler, R. A. (1951). On information and sufficiency. *The annals of*
1005 *mathematical statistics*, *22*(1), 79–86.
- 1006 Liu, C., & Zhu, J. (2018). Riemannian stein variational gradient descent for bayesian
1007 inference. In *Thirty-second aaii conference on artificial intelligence*.
- 1008 Liu, Q. (2017). Stein variational gradient descent as gradient flow. In *Advances in*
1009 *neural information processing systems* (pp. 3115–3123).
- 1010 Liu, Q., Lee, J., & Jordan, M. (2016). A kernelized Stein discrepancy for goodness-
1011 of-fit tests. In *International conference on machine learning* (pp. 276–284).
- 1012 Liu, Q., & Wang, D. (2016). Stein variational gradient descent: A general purpose
1013 Byesian inference algorithm. In *Advances in neural information processing sys-*
1014 *tems* (pp. 2378–2386).
- 1015 MacKay, D. J. (2003). *Information theory, inference and learning algorithms*. Cam-
1016 bridge university press.

- 1017 Malinverno, A. (2002). Parsimonious Bayesian Markov chain Monte Carlo inversion
 1018 in a nonlinear geophysical problem. *Geophysical Journal International*, *151*(3),
 1019 675–688.
- 1020 Malinverno, A., & Briggs, V. A. (2004). Expanded uncertainty quantification in
 1021 inverse problems: Hierarchical Bayes and empirical Bayes. *Geophysics*, *69*(4),
 1022 1005–1016.
- 1023 Malinverno, A., Leaney, S., et al. (2000). A Monte Carlo method to quantify uncer-
 1024 tainty in the inversion of zero-offset VSP data. In *2000 seg annual meeting*.
- 1025 Marzouk, Y., Moselhy, T., Parno, M., & Spantini, A. (2016). An introduction to
 1026 sampling via measure transport. *arXiv preprint arXiv:1602.05023*.
- 1027 Meier, U., Curtis, A., & Trampert, J. (2007a). A global crustal model constrained
 1028 by nonlinearised inversion of fundamental mode surface waves. *Geophysical Re-
 1029 search Letters*, *34*, L16304.
- 1030 Meier, U., Curtis, A., & Trampert, J. (2007b). Global crustal thickness from neu-
 1031 ral network inversion of surface wave data. *Geophysical Journal International*,
 1032 *169*(2), 706–722.
- 1033 Metropolis, N., & Ulam, S. (1949). The Monte Carlo method. *Journal of the Ameri-
 1034 can statistical association*, *44*(247), 335–341.
- 1035 Mosegaard, K., & Tarantola, A. (1995). Monte Carlo sampling of solutions to
 1036 inverse problems. *Journal of Geophysical Research: Solid Earth*, *100*(B7),
 1037 12431–12447.
- 1038 Nawaz, M., & Curtis, A. (2019). Rapid discriminative variational Bayesian inversion
 1039 of geophysical data for the spatial distribution of geological properties. *Journal
 1040 of Geophysical Research: Solid Earth*.
- 1041 Nawaz, M. A., & Curtis, A. (2018). Variational Bayesian inversion (VBI) of quasi-
 1042 localized seismic attributes for the spatial distribution of geological facies. *Geo-
 1043 physical Journal International*, *214*(2), 845–875.
- 1044 Neal, R. M., et al. (2011). Mcmc using hamiltonian dynamics. *Handbook of markov
 1045 chain monte carlo*, *2*(11), 2.
- 1046 Nicolson, H., Curtis, A., & Baptie, B. (2014). Rayleigh wave tomography of the
 1047 British Isles from ambient seismic noise. *Geophysical Journal International*,
 1048 *198*(2), 637–655.
- 1049 Nicolson, H., Curtis, A., Baptie, B., & Galetti, E. (2012). Seismic interferometry

- 1050 and ambient noise tomography in the British Isles. *Proceedings of the Geolo-*
 1051 *gists' Association*, 123(1), 74–86.
- 1052 Piana Agostinetti, N., Giacomuzzi, G., & Malinverno, A. (2015). Local three-
 1053 dimensional earthquake tomography by trans-dimensional Monte Carlo sam-
 1054 pling. *Geophysical Journal International*, 201(3), 1598–1617.
- 1055 Ranganath, R., Gerrish, S., & Blei, D. (2014). Black box variational inference. In
 1056 *Artificial intelligence and statistics* (pp. 814–822).
- 1057 Ranganath, R., Tran, D., & Blei, D. (2016). Hierarchical variational models. In *In-*
 1058 *ternational conference on machine learning* (pp. 324–333).
- 1059 Rawlinson, N., & Sambridge, M. (2004). Multiple reflection and transmission
 1060 phases in complex layered media using a multistage fast marching method.
 1061 *Geophysics*, 69(5), 1338–1350.
- 1062 Ray, A., Alumbaugh, D. L., Hoversten, G. M., & Key, K. (2013). Robust and ac-
 1063 celerated Bayesian inversion of marine controlled-source electromagnetic data
 1064 using parallel tempering. *Geophysics*, 78(6), E271–E280.
- 1065 Ray, A., Kaplan, S., Washbourne, J., & Albertin, U. (2017). Low frequency full
 1066 waveform seismic inversion within a tree based Bayesian framework. *Geophysi-*
 1067 *cal Journal International*, 212(1), 522–542.
- 1068 Ray, A., & Myer, D. (2019). Bayesian geophysical inversion with trans-dimensional
 1069 gaussian process machine learning. *Geophysical Journal International*, 217(3),
 1070 1706–1726.
- 1071 Ray, A., Sekar, A., Hoversten, G. M., & Albertin, U. (2016). Frequency domain
 1072 full waveform elastic inversion of marine seismic data from the alba field using
 1073 a bayesian trans-dimensional algorithm. *Geophysical Journal International*,
 1074 205(2), 915–937.
- 1075 Rezende, D. J., & Mohamed, S. (2015). Variational inference with normalizing flows.
 1076 *arXiv preprint arXiv:1505.05770*.
- 1077 Robbins, H., & Monro, S. (1951). A stochastic approximation method. *The annals*
 1078 *of mathematical statistics*, 400–407.
- 1079 Robert, C., & Casella, G. (2013). *Monte Carlo statistical methods*. Springer Science
 1080 & Business Media.
- 1081 Röth, G., & Tarantola, A. (1994). Neural networks and inversion of seismic data.
 1082 *Journal of Geophysical Research: Solid Earth*, 99(B4), 6753–6768.

- 1083 Salimans, T., Kingma, D., & Welling, M. (2015). Markov chain Monte Carlo and
 1084 variational inference: Bridging the gap. In *International conference on machine*
 1085 *learning* (pp. 1218–1226).
- 1086 Sambridge, M. (1999). Geophysical inversion with a neighbourhood algorithm – i.
 1087 searching a parameter space. *Geophysical journal international*, *138*(2), 479–
 1088 494.
- 1089 Sambridge, M. (2013). A parallel tempering algorithm for probabilistic sampling and
 1090 multimodal optimization. *Geophysical Journal International*, ggt342.
- 1091 Saul, L. K., & Jordan, M. I. (1996). Exploiting tractable substructures in intractable
 1092 networks. In *Advances in neural information processing systems* (pp. 486–
 1093 492).
- 1094 Sen, M. K., & Biswas, R. (2017). Transdimensional seismic inversion using the
 1095 reversible jump hamiltonian monte carlo algorithm. *Geophysics*, *82*(3), R119–
 1096 R134.
- 1097 Shahraeeni, M. S., & Curtis, A. (2011). Fast probabilistic nonlinear petrophysical in-
 1098 version. *Geophysics*, *76*(2), E45–E58.
- 1099 Shahraeeni, M. S., Curtis, A., & Chao, G. (2012). Fast probabilistic petrophysical
 1100 mapping of reservoirs from 3D seismic data. *Geophysics*, *77*(3), O1–O19.
- 1101 Shapiro, N. M., Campillo, M., Stehly, L., & Ritzwoller, M. H. (2005). High-
 1102 resolution surface-wave tomography from ambient seismic noise. *Science*,
 1103 *307*(5715), 1615–1618.
- 1104 Shen, W., Ritzwoller, M. H., & Schulte-Pelkum, V. (2013). A 3-D model of the crust
 1105 and uppermost mantle beneath the central and western US by joint inver-
 1106 sion of receiver functions and surface wave dispersion. *Journal of Geophysical*
 1107 *Research: Solid Earth*, *118*(1), 262–276.
- 1108 Shen, W., Ritzwoller, M. H., Schulte-Pelkum, V., & Lin, F.-C. (2012). Joint inver-
 1109 sion of surface wave dispersion and receiver functions: a Bayesian Monte-Carlo
 1110 approach. *Geophysical Journal International*, *192*(2), 807–836.
- 1111 Sivia, D. (1996). *Data analysis: A Bayesian tutorial* (oxford science publications).
- 1112 Smith, A. (2013). *Sequential Monte Carlo methods in practice*. Springer Science &
 1113 Business Media.
- 1114 Stein, C., et al. (1972). A bound for the error in the normal approximation to the
 1115 distribution of a sum of dependent random variables. In *Proceedings of the*

- 1116 *sixth berkeley symposium on mathematical statistics and probability, volume 2:*
 1117 *Probability theory.*
- 1118 Tarantola, A. (2005). *Inverse problem theory and methods for model parameter esti-*
 1119 *mation* (Vol. 89). SIAM.
- 1120 Team, S. D., et al. (2016). Stan modeling language users guide and reference man-
 1121 ual. *Technical report.*
- 1122 Tran, D., Ranganath, R., & Blei, D. M. (2015). The variational Gaussian process.
 1123 *arXiv preprint arXiv:1511.06499.*
- 1124 Walter, E., & Pronzato, L. (1997). *Identification of parametric models from experi-*
 1125 *mental data.* Springer Verlag.
- 1126 Weaver, R. L., Hadziioannou, C., Larose, E., & Campillo, M. (2011). On the pre-
 1127 cision of noise correlation interferometry. *Geophysical Journal International,*
 1128 *185(3), 1384–1392.*
- 1129 Yao, H., & Van Der Hilst, R. D. (2009). Analysis of ambient noise energy distribu-
 1130 tion and phase velocity bias in ambient noise tomography, with application to
 1131 SE tibet. *Geophysical Journal International, 179(2), 1113–1132.*
- 1132 Young, M. K., Rawlinson, N., & Bodin, T. (2013). Transdimensional inversion of
 1133 ambient seismic noise for 3D shear velocity structure of the Tasmanian crust.
 1134 *Geophysics, 78(3), WB49–WB62.*
- 1135 Zhang, X., Curtis, A., Galetti, E., & de Ridder, S. (2018). 3-D Monte Carlo surface
 1136 wave tomography. *Geophysical Journal International, 215(3), 1644–1658.*
- 1137 Zhang, X., Hansteen, F., & Curtis, A. (2019). Fully 3D Monte Carlo ambient noise
 1138 tomography over Grane field. In *81st eage conference and exhibition 2019.*
- 1139 Zhang, X., & Zhang, H. (2015). Wavelet-based time-dependent travel time tomog-
 1140 raphy method and its application in imaging the Etna volcano in Italy. *Journal*
 1141 *of Geophysical Research: Solid Earth, 120(10), 7068–7084.*
- 1142 Zhdanov, M. S. (2002). *Geophysical inverse theory and regularization problems*
 1143 (Vol. 36). Elsevier.
- 1144 Zheng, D., Saygin, E., Cummins, P., Ge, Z., Min, Z., Cipta, A., & Yang, R. (2017).
 1145 Transdimensional Bayesian seismic ambient noise tomography across SE tibet.
 1146 *Journal of Asian Earth Sciences, 134, 86–93.*
- 1147 Zobay, O., et al. (2014). Variational bayesian inference with gaussian-mixture ap-
 1148 proximations. *Electronic Journal of Statistics, 8(1), 355–389.*

1149 Zulfakriza, Z., Saygin, E., Cummins, P., Widiyantoro, S., Nugraha, A. D., Lühr,
 1150 B.-G., & Bodin, T. (2014). Upper crustal structure of central Java, Indonesia,
 1151 from transdimensional seismic ambient noise tomography. *Geophysical Journal*
 1152 *International*, 197(1), 630–635.

1153 Appendix A The entropy of a Gaussian distribution

1154 The entropy $H[q(\boldsymbol{\theta}; \boldsymbol{\phi})]$ of a Gaussian distribution $\mathcal{N}(\boldsymbol{\theta}|\boldsymbol{\mu}, \mathbf{L}\mathbf{L}^T)$ is:

$$\begin{aligned}
 1155 \quad H[q(\boldsymbol{\theta}; \boldsymbol{\phi})] &= -\mathbb{E}_q[\log q(\boldsymbol{\theta})] \\
 1156 &= -\int \mathcal{N}(\boldsymbol{\theta}|\boldsymbol{\mu}, \mathbf{L}\mathbf{L}^T) \log \mathcal{N}(\boldsymbol{\theta}|\boldsymbol{\mu}, \mathbf{L}\mathbf{L}^T) d\boldsymbol{\theta} \\
 1157 &= \frac{k}{2} + \frac{k}{2} \log(2\pi) + \frac{1}{2} \log|\det(\mathbf{L}\mathbf{L}^T)|
 \end{aligned}$$

1158 where k is the dimension of vector $\boldsymbol{\theta}$. The gradients with respect to $\boldsymbol{\mu}$ and \mathbf{L} can be eas-
 1159 ily calculated (see Appendix B).

1160 Appendix B Gradients of the ELBO in ADVI

1161 We first describe the dominated convergence theorem (DCT) (Çınlar, 2011):

1162 **Theorem** Assume $X \in \mathcal{X}$ is a random variable and $f : \mathbb{R} \times \mathcal{X} \rightarrow \mathbb{R}$ is a function
 1163 such that $f(t, X)$ is integrable for all t and $\frac{\partial f(t, X)}{\partial t}$ exists for each t . Assume that there
 1164 is a random variable Z such that $|\frac{\partial f(t, X)}{\partial t}| \leq Z$ for all t and $\mathbb{E}(Z) < \infty$. Then

$$1165 \quad \frac{\partial}{\partial t} \mathbb{E}(f(t, X)) = \mathbb{E}\left(\frac{\partial}{\partial t} f(t, X)\right)$$

1166 The proof of this theorem is given in Çınlar (2011).

1167 We then calculate the gradients in equation (9) and (10) based on Kucukelbir et
 1168 al. (2017). The ELBO \mathcal{L} is:

$$1169 \quad \mathcal{L} = \mathbb{E}_{\mathcal{N}(\boldsymbol{\eta}|\mathbf{0}, \mathbf{I})} \left[\log p(T^{-1}(R_{\boldsymbol{\phi}}^{-1}(\boldsymbol{\eta})), \mathbf{d}_{obs}) + \log|\det \mathbf{J}_{T^{-1}}(R_{\boldsymbol{\phi}}^{-1}(\boldsymbol{\eta}))| \right] + H[q(\boldsymbol{\theta}; \boldsymbol{\phi})]$$

1170 where $H[q(\boldsymbol{\theta}; \boldsymbol{\phi})] = \mathbb{E}_q[\log q(\boldsymbol{\theta})]$ is the entropy of distribution q . Assume $\frac{\partial}{\partial \boldsymbol{\phi}} \log p$ is bounded
 1171 where $\boldsymbol{\phi}$ represents variational parameters $\boldsymbol{\mu}$ and \mathbf{L} , then the gradients can be computed
 1172 by exchanging the derivative and the expectation using the dominated convergence the-
 1173 orom (DCT) and applying the chain rule:

$$1174 \quad \nabla_{\boldsymbol{\mu}} \mathcal{L} = \nabla_{\boldsymbol{\mu}} \left\{ \mathbb{E}_{\mathcal{N}(\boldsymbol{\eta}|\mathbf{0}, \mathbf{I})} \left[\log p(T^{-1}(R_{\boldsymbol{\phi}}^{-1}(\boldsymbol{\eta})), \mathbf{d}_{obs}) + \log|\det \mathbf{J}_{T^{-1}}(R_{\boldsymbol{\phi}}^{-1}(\boldsymbol{\eta}))| \right] + H[q(\boldsymbol{\theta}; \boldsymbol{\phi})] \right\}$$

1175 Applying the DCT and since \mathbf{H} does not depend on $\boldsymbol{\mu}$,

$$1176 \quad \nabla_{\boldsymbol{\mu}} \mathcal{L} = \mathbb{E}_{\mathcal{N}(\boldsymbol{\eta}|\mathbf{0}, \mathbf{I})} \left[\nabla_{\boldsymbol{\mu}} \left\{ \log p(T^{-1}(R_{\boldsymbol{\phi}}^{-1}(\boldsymbol{\eta})), \mathbf{d}_{obs}) \right\} + \nabla_{\boldsymbol{\mu}} \left(\log |\det \mathbf{J}_{T^{-1}}(R_{\boldsymbol{\phi}}^{-1}(\boldsymbol{\eta}))| \right) \right]$$

1177 Applying the chain rule,

$$1178 \quad \begin{aligned} \nabla_{\boldsymbol{\mu}} \mathcal{L} &= \mathbb{E}_{\mathcal{N}(\boldsymbol{\eta}|\mathbf{0}, \mathbf{I})} \left[\nabla_{\mathbf{m}} \log p(\mathbf{m}, \mathbf{d}_{obs}) \nabla_{\boldsymbol{\theta}} T^{-1}(\boldsymbol{\theta}) \nabla_{\boldsymbol{\mu}} R_{\boldsymbol{\phi}}^{-1}(\boldsymbol{\eta}) + \nabla_{\boldsymbol{\theta}} \log |\det \mathbf{J}_{T^{-1}}(\boldsymbol{\theta})| \nabla_{\boldsymbol{\mu}} R_{\boldsymbol{\phi}}^{-1}(\boldsymbol{\eta}) \right] \\ 1179 \quad &= \mathbb{E}_{\mathcal{N}(\boldsymbol{\eta}|\mathbf{0}, \mathbf{I})} \left[\nabla_{\mathbf{m}} \log p(\mathbf{m}, \mathbf{d}_{obs}) \nabla_{\boldsymbol{\theta}} T^{-1}(\boldsymbol{\theta}) + \nabla_{\boldsymbol{\theta}} \log |\det \mathbf{J}_{T^{-1}}(\boldsymbol{\theta})| \right] \end{aligned}$$

1180 The gradient with respect to \mathbf{L} can be obtained similarly,

$$1181 \quad \begin{aligned} \nabla_{\mathbf{L}} \mathcal{L} &= \nabla_{\mathbf{L}} \left\{ \mathbb{E}_{\mathcal{N}(\boldsymbol{\eta}|\mathbf{0}, \mathbf{I})} \left[\log p(T^{-1}(R_{\boldsymbol{\phi}}^{-1}(\boldsymbol{\eta})), \mathbf{d}_{obs}) + \log |\det \mathbf{J}_{T^{-1}}(R_{\boldsymbol{\phi}}^{-1}(\boldsymbol{\eta}))| \right] \right. \\ 1182 \quad &\quad \left. + \frac{k}{2} + \frac{k}{2} \log(2\pi) + \frac{1}{2} \log |\det(\mathbf{L}\mathbf{L}^T)| \right\} \end{aligned}$$

1183 Applying the DCT

$$1184 \quad \begin{aligned} \nabla_{\mathbf{L}} \mathcal{L} &= \mathbb{E}_{\mathcal{N}(\boldsymbol{\eta}|\mathbf{0}, \mathbf{I})} \left[\nabla_{\mathbf{L}} \left\{ \log p(T^{-1}(R_{\boldsymbol{\phi}}^{-1}(\boldsymbol{\eta})), \mathbf{d}_{obs}) \right\} + \nabla_{\mathbf{L}} \left(\log |\det \mathbf{J}_{T^{-1}}(R_{\boldsymbol{\phi}}^{-1}(\boldsymbol{\eta}))| \right) \right] \\ 1185 \quad &\quad + \nabla_{\mathbf{L}} \frac{1}{2} \log |\det(\mathbf{L}\mathbf{L}^T)| \end{aligned}$$

1186 and applying the chain rule we obtain

$$1187 \quad \begin{aligned} \nabla_{\mathbf{L}} \mathcal{L} &= \mathbb{E}_{\mathcal{N}(\boldsymbol{\eta}|\mathbf{0}, \mathbf{I})} \left[\nabla_{\mathbf{m}} \log p(\mathbf{m}, \mathbf{d}_{obs}) \nabla_{\boldsymbol{\theta}} T^{-1}(\boldsymbol{\theta}) \nabla_{\mathbf{L}} R_{\boldsymbol{\phi}}^{-1}(\boldsymbol{\eta}) + \nabla_{\boldsymbol{\theta}} \log |\det \mathbf{J}_{T^{-1}}(\boldsymbol{\theta})| \nabla_{\mathbf{L}} R_{\boldsymbol{\phi}}^{-1}(\boldsymbol{\eta}) \right] + (\mathbf{L}^{-1})^T \\ 1188 \quad &= \mathbb{E}_{\mathcal{N}(\boldsymbol{\eta}|\mathbf{0}, \mathbf{I})} \left[(\nabla_{\mathbf{m}} \log p(\mathbf{m}, \mathbf{d}_{obs}) \nabla_{\boldsymbol{\theta}} T^{-1}(\boldsymbol{\theta}) + \nabla_{\boldsymbol{\theta}} \log |\det \mathbf{J}_{T^{-1}}(\boldsymbol{\theta})|) \boldsymbol{\eta}^T \right] + (\mathbf{L}^{-1})^T \end{aligned}$$

1189 Appendix C Gradients of KL-divergence in SVGD

1190 We calculate the gradient in equation (12) following Q. Liu and Wang (2016). De-
1191 note T^{-1} as the inverse transform of T . Then by changing the variable,

$$1192 \quad \text{KL}[q_T || p] = \text{KL}[q || p_{T^{-1}}]$$

1193 and hence

$$1194 \quad \begin{aligned} \nabla_{\epsilon} \text{KL}[q_T || p] |_{\epsilon=0} &= \nabla_{\epsilon} \text{KL}[q || p_{T^{-1}}] |_{\epsilon=0} \\ 1195 \quad &= \nabla_{\epsilon} [\mathbb{E}_q \log q(\mathbf{m}) - \mathbb{E}_q \log p_{T^{-1}}(\mathbf{m})] \end{aligned}$$

1196 and since $q(\mathbf{m})$ does not depend on ϵ

$$1197 \quad \nabla_{\epsilon} \text{KL}[q_T || p] |_{\epsilon=0} = -\mathbb{E}_q [\nabla_{\epsilon} \log p_{T^{-1}}(\mathbf{m})]$$

1198 where $p_{T^{-1}}(\mathbf{m}) = p(T(\mathbf{m})) \cdot |\det(\nabla_{\mathbf{m}} T(\mathbf{m}))|$. Therefore

$$1199 \quad \nabla_{\epsilon} \log p_{T^{-1}}(\mathbf{m}) = (\nabla_{\mathbf{m}} \log(p(\mathbf{m})))^T \nabla_{\epsilon} T(\mathbf{m}) + \text{trace} \left((\nabla_{\mathbf{m}} T(\mathbf{m}))^{-1} \cdot \nabla_{\epsilon} \nabla_{\mathbf{m}} T(\mathbf{m}) \right)$$

1200 where $T(\mathbf{m}) = \mathbf{m} + \epsilon\phi(\mathbf{m})$, $\nabla_{\epsilon}T(\mathbf{m} = \phi(\mathbf{m}))$ and $\nabla_{\mathbf{m}}T(\mathbf{m})|_{\epsilon=0} = \mathbf{I}$, and so

$$\begin{aligned}
 1201 \quad \nabla_{\epsilon}\text{KL}[q_T||p]|_{\epsilon=0} &= -\mathbf{E}_q \left[(\nabla_{\mathbf{m}}\log(p(\mathbf{m})))^T \phi(\mathbf{m}) + \text{trace}(\nabla_{\mathbf{m}}\phi(\mathbf{m})) \right] \\
 1202 &= -\mathbf{E}_q \left[\text{trace}(\nabla_{\mathbf{m}}\log(p(\mathbf{m}))\phi(\mathbf{m})^T) + \text{trace}(\nabla_{\mathbf{m}}\phi(\mathbf{m})) \right] \\
 1203 &= -\mathbf{E}_q [\text{trace}(\mathcal{A}_p\phi(\mathbf{m}))]
 \end{aligned}$$

1204 where $\mathcal{A}_p\phi(\mathbf{m}) = \nabla_{\mathbf{m}}\log p(\mathbf{m})\phi(\mathbf{m})^T + \nabla_{\mathbf{m}}\phi(\mathbf{m})$ is the Stein operator.