



# HO<sub>x</sub> measurements in the summertime upper troposphere over Europe: a comparison of observations to a box model and a 3-D model

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**Abstract.** In situ airborne measurements of OH and HO<sub>2</sub> with the HORUS (HydrOxyl Radical measurement Unit based on fluorescence Spectroscopy) instrument were performed in the summertime upper troposphere across Europe during the HOOVER 2 (HO<sub>x</sub> OVER EuROpe) campaign in July 2007. Complementary measurements of trace gas species and photolysis frequencies were conducted to obtain a broad data set, which has been used to quantify the significant HO<sub>x</sub> sources and sinks. In this study we compare the in situ measurement of OH and HO<sub>2</sub> with simulated mixing ratios from the constrained box model CAABA/MECCA (Chemistry As A Box Model Application/Module Efficiently Calculating the Chemistry of the Atmosphere), and the global circulation model EMAC (ECHAM5/MESSy Atmospheric Chemistry Model). The constrained box model reproduces the observed OH and HO<sub>2</sub> mixing ratios with better agreement (obs/mod median 98 % OH, 96 % HO<sub>2</sub>) than the global model (median 76 % OH, 59 % HO<sub>2</sub>). The observations and the computed HO<sub>x</sub> sources and sinks are used to identify deviations between the models and their impacts on the calculated HO<sub>x</sub> budget.

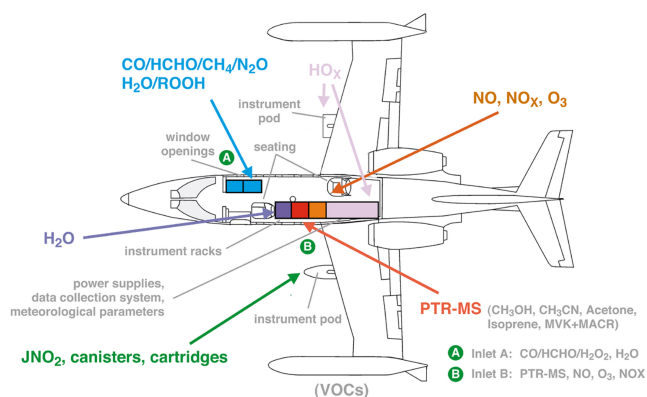
## 1 Introduction

A wide range of chemical compounds emitted by human activities such as industrial processes and traffic, and by vegetation and animals, are oxidised by reactions with the OH radical. Short-lived species are readily oxidised within the planetary boundary layer. Species with a longer lifetime can be mixed and diluted into the free and upper troposphere. There, some gases like HCHO and H<sub>2</sub>O<sub>2</sub> form OH and HO<sub>2</sub> (together called HO<sub>x</sub>) through photolysis; other species are depleted by reactions with OH.

The oxidation capacity of the atmosphere is formed by highly reactive species, which together act as the cleansing agents of the troposphere. The OH radical is the most important oxidising component of the troposphere and has been recognized as the most important cleansing oxidant (Levy, 1971). In the presence of NO<sub>x</sub> or ozone, the OH reaches equilibrium with HO<sub>2</sub> and particularly NO shifts the HO<sub>x</sub> equilibrium towards the OH.

The basic HO<sub>x</sub> chemistry in the (upper) troposphere has previously been summarized in a number of articles (Jaegle et al., 2000; Prather and Jacob, 1997; Ren et al., 2008 and reference therein).

For the scope of this paper a limited number of HO<sub>x</sub> production and loss channels are relevant. The most important tropospheric primary HO<sub>x</sub> source is the photolysis of ozone and the subsequent reaction of O(<sup>1</sup>D) with gas phase water



**Fig. 1.** Payload of the Learjet during HOOVER (© group graphic pool).

(R1 and 2, see Table 1). In the dry air masses of the upper troposphere, where the O(<sup>1</sup>D)-water reaction is relatively ineffective, photolysis of peroxides (R3 and 4) and aldehydes (R5–7) in particular HCHO, lead to substantial OH formation. This can be important in addition to the O(<sup>1</sup>D)/H<sub>2</sub>O-OH source and under some conditions becomes the dominant primary source (Jaegle et al., 2000; Ren et al., 2008; Tan et al., 2001).

HO<sub>x</sub> radicals react with a large range of trace gases. Under low NO<sub>x</sub> conditions in upper tropospheric background air the most important HO<sub>x</sub> sink is the radical-radical reaction of HO<sub>2</sub> either with a second HO<sub>2</sub> radical (R9) or an organic RO<sub>2</sub> radical (R10). The resulting peroxides are in turn precursor species of OH (Klippel et al., 2011), though in the presence of clouds, they are efficiently scavenged (Lelieveld and Crutzen, 1990; Snow et al., 2007). It was reported that, subsequent to the uptake and aqueous phase chemistry, soluble species such as peroxides and HCHO can be released again from cloud droplets through evaporation, as most clouds evaporate rather than precipitate. The microphysical processes under freezing conditions are poorly understood and need further investigation. For example, there is ongoing discussion about H<sub>2</sub>O<sub>2</sub> release from cloud droplets during freezing (Barth et al., 2001; Mari et al., 2000, 2002), which would increase H<sub>2</sub>O<sub>2</sub> concentrations in the upper troposphere. Since the respective H<sub>2</sub>O<sub>2</sub> was released from clouds it was not formed in situ from HO<sub>2</sub> and therefore introduces a potential primary OH source.

A recent study (Klippel et al., 2011) reports that in comparison to field observations the global circulation model EMAC underestimates the H<sub>2</sub>O<sub>2</sub> concentration in the upper troposphere over Europe. Here we investigate the implications of this underestimated H<sub>2</sub>O<sub>2</sub> concentration in the global model on the HO<sub>x</sub> budget.

As shown by Tan et al. (2001) for the tropical Pacific, under upper tropospheric conditions, a substantial fraction of the observed OH and HO<sub>2</sub> radicals is directly formed through



**Fig. 2.** HORUS detection axes are mounted in the wing pod below the wing (© group graphic pool).

**Table 1.** Atmospheric sources (R1–R8), sinks (R9–R11) and cycling (R12–R18) reactions of OH and HO<sub>2</sub> and calibration sources (R19 and R6) of HO<sub>x</sub>.

O <sub>3</sub> + hν	→	O( <sup>1</sup> D) + O <sub>2</sub> (λ ≤ 340 nm)	R1
O( <sup>1</sup> D) + H <sub>2</sub> O	→	2 OH	R2
H <sub>2</sub> O <sub>2</sub> + hν	→	2 OH (λ ≤ 400 nm)	R3
ROOH + hν	→	OH + HO <sub>2</sub> + HCHO	R4
HCHO + hν	→	H + CHO (λ ≤ 335 nm)	R5
H + O <sub>2</sub> + M	→	HO <sub>2</sub> + M	R6
CHO + O <sub>2</sub>	→	HO <sub>2</sub> + CO	R7
HCHO + hν	→	H <sub>2</sub> + CO (λ ≤ 360 nm)	R8
HO <sub>2</sub> + HO <sub>2</sub>	→	H <sub>2</sub> O <sub>2</sub> + O <sub>2</sub>	R9
HO <sub>2</sub> + RO <sub>2</sub>	→	RO <sub>2</sub> H + O <sub>2</sub>	R10
OH + NO <sub>2</sub>	→	HNO <sub>3</sub>	R11
HO <sub>2</sub> + NO	→	OH + NO <sub>2</sub>	R12
HO <sub>2</sub> + O <sub>3</sub>	→	OH + 2 O <sub>2</sub>	R13
OH + CO + O <sub>2</sub>	→	HO <sub>2</sub> + CO <sub>2</sub>	R14
OH + O <sub>3</sub>	→	HO <sub>2</sub> + O <sub>2</sub>	R15
OH + CH <sub>4</sub> (+ O <sub>2</sub> )	→	CH <sub>3</sub> O <sub>2</sub> + H <sub>2</sub> O	R16
CH <sub>3</sub> O <sub>2</sub> + NO	→	CH <sub>3</sub> O + NO <sub>2</sub>	R17
CH <sub>3</sub> O + O <sub>2</sub>	→	HO <sub>2</sub> + HCHO	R18
H <sub>2</sub> O + hν	→	OH + H (λ = 184.9 nm)	R19
H + O <sub>2</sub> + M	→	HO <sub>2</sub> + M	R6

primary production, whereas the dominant fraction is produced by cycling reactions. In general, the HO<sub>x</sub> equilibrium is determined by R12–15. The variability of the HO<sub>x</sub> equilibrium is mainly influenced by the variability of the NO concentrations (R12). Thus, increased or decreased OH mixing ratios can also be due to enhanced or reduced HO<sub>2</sub> cycling to OH. In the following sections, we will analyse the HO<sub>x</sub> source and sink strengths with regard to primary and cycling HO<sub>x</sub> sources in the upper troposphere under conditions observed during the HOOVER summer campaign. A comparison of a simple HO<sub>x</sub> reaction scheme applied as part of a constrained box model and a global 3-D model are used to identify differences, strengths, and weaknesses of the models.

The HOOVER 2 campaign is described in Sect. 2. Technical details of the HO<sub>x</sub> measurement instrument HORUS are given in Sect. 3. Sections 4 and 5 introduce the box model

**Table 2.** Summarized are mean values of observed and simulated (global model) trace gas mixing ratios and a photolysis frequency. The observations displayed are mean values of the four flights presented here and are separated for convectively (convective) affected and unaffected (background) air masses. Tuned model indicates a simulation of the global model with increased photolysis frequencies to match the observation.

	OH [pmol mol <sup>-1</sup> ]	HO <sub>2</sub> [pmol mol <sup>-1</sup> ]	NO [pmol mol <sup>-1</sup> ]	CO [nmol/mol]	O <sub>3</sub> [nmol/mol]	H <sub>2</sub> O <sub>2</sub> [nmol/mol]	J(O <sup>1</sup> D) [s <sup>-1</sup> ]
Background	0.41	15.8	57.5	97.5	81.2	0.69	4.56 10 <sup>-5</sup>
Convective	2.97	14.7	996	107	85.6	1.30	7.57 10 <sup>-5</sup>
Global Model	0.41	8.89	99.3	64.2	102	0.26	3.27 10 <sup>-5</sup>
Tuned Model	0.60	9.39	121	52.3	108	0.21	4.40 10 <sup>-5</sup>

**Table 3.** Techniques employed to measure trace species during HOOVER. Abbreviations are explained in the corresponding text.

Species	Technique	Precision	Accuracy	Limit of detection	Time resolution
OH	LIF	0.03 pmol mol <sup>-1</sup>	18 %	0.016 pmol mol <sup>-1</sup>	60 s
HO <sub>2</sub>	LIF	0.42 pmol mol <sup>-1</sup>	18 %	0.33 pmol mol <sup>-1</sup>	60 s
H <sub>2</sub> O <sub>2</sub>	DEF	8.3 %	14 %	24 pmol mol <sup>-1</sup>	1200 s
ROOH	DEF	6.3 %	21 %	< 24 pmol mol <sup>-1</sup>	1200 s
O <sub>3</sub>	CL	±1 ppt/4 %	2 %	2000 pmol mol <sup>-1</sup>	30 s
HCHO	QCL	–	8.6 %	300 pmol mol <sup>-1</sup>	120 s
NO	CL	±8 ppt/7 %	12 %	5 pmol mol <sup>-1</sup>	30 s
CO	QCL	–	1.1 %	200 pmol mol <sup>-1</sup>	2 s
CH <sub>4</sub>	QCL	–	0.57 %	6000 pmol mol <sup>-1</sup>	2 s
J(NO <sub>2</sub> )	FR	1 %	15 %	–	1 s
H <sub>2</sub> O	Humicap	–	–	100 μmol mol <sup>-1</sup>	30 s

and the global model, respectively. Results and discussion of the simulations are presented in the respective model section. The summary and conclusions are part of Sect. 6.

## 2 The HOOVER 2 Campaign

The HOOVER 2 campaign was conceived to study the spatial variability of the oxidation capacity in the summertime troposphere over Europe (see also Klippel et al., 2011).

The most important oxidising species is the OH radical, which has been measured in situ along with its precursors, the hydrogen peroxy radical, ozone, nitrogen monoxide, hydrogen peroxide and total organic peroxides (mean values are summarized in Table 2). To investigate the dominant species believed to impact the OH reactivity, methane, carbon monoxide, and formaldehyde were also measured.

An aircraft (Learjet 35 A) was used to conduct the measurements. A suite of measurement equipment was placed within the cabin and in wing pods which were mounted below the wings. Figure 1 shows a schematic overview of the payload. The wing pod which contained most of the HORUS instrument and infrastructure is shown in Fig. 2.

Since all other instruments used in the study are described elsewhere, only a brief description is given here. Table 3 summarizes technical details of the measurement techniques used.

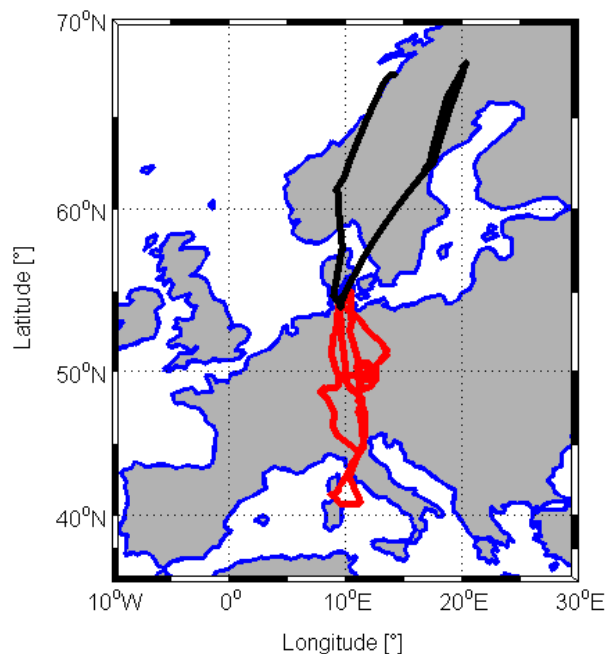
O<sub>3</sub> and NO were measured with a chemiluminescence (CL) detection system. This instrumental setup was already used in a similar configuration during the OOMPH campaign and is described in detail by Beygi et al. (2011).

H<sub>2</sub>O<sub>2</sub> and organic peroxides were observed with a wet chemical system based on derivatisation and fluorescence enzyme (DEF) described in Klippel et al. (2011).

CO, CH<sub>4</sub> and HCHO measurements were performed with a multi-channel infrared quantum cascade laser (QCL) absorption spectrometer (Schiller et al., 2008).

J (NO<sub>2</sub>) data was measured with a set of two filter radiometers (FR) for the downwelling and upwelling fraction. H<sub>2</sub>O was recorded with a Helten Sensor.

Here we report observations obtained during four flights between the Mediterranean (~ 41° N) and sub-polar northern Scandinavia (~ 68° N), performed in the upper troposphere at altitudes higher than 7 km. Flight tracks are depicted in Fig. 3. A typical profile of a HOOVER 2 flight is shown in Fig. 4. During the campaign, the favoured flight level was above 7 km altitude in the upper troposphere and vertical profiles were sampled during takeoff and landing as well as during a midway descent during each flight.



**Fig. 3.** Northern (black) and southern (red) flight tracks of the HOOVER 2 campaign.

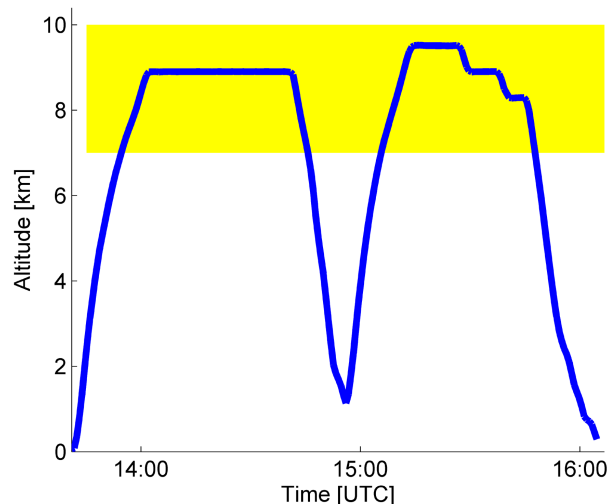
### 3 HORUS instrument

#### 3.1 Description

The OH and HO<sub>2</sub> radical concentrations were measured by using laser-induced fluorescence, using the HORUS instrument, built at the Max Planck Institute for Chemistry in Mainz and redesigned for aircraft campaigns.

The laser system was located within the cabin of the Learjet, while the detection system and most of the infrastructure was mounted in a wing pod below one of the wings. The laser beam was transported into the detection system through 10 m-long optical fibres.

UV light at about 308 nm was used to selectively excite OH radicals ( $A^2\Sigma - X^2\Pi$ ,  $v' = 0 \leftarrow v'' = 0$ ) through the  $Q_1(2)$  transition in a low pressure detection cell (2–8 mbar). The laser pulse repetition frequency was set to 3 kHz during the HOOVER campaigns. To increase the sensitivity of the instrument, the laser light was reflected 32 times using a White cell (White, 1942). Excitation and fluorescence occur at the same wavelength. The fluorescence was measured with a multi-channel photomultiplier perpendicular to the laser beam. The background signal was determined by tuning the laser wavelength off-resonance of the excitation wavelength. The laser was tuned on- and off-resonance with the OH transition every 5 s to determine the fluorescence plus background signals and the background signals, respectively. The off-resonance background fluorescence was then subtracted from the on-resonance OH fluorescence signal. The achieved time resolution was 10 s.



**Fig. 4.** A typical profile of a HOOVER2 flight. Coloured area highlights the upper troposphere.

OH was detected in a first axis. HO<sub>2</sub> was detected in a second axis 16 cm downstream of the first one after addition of NO titration gas to the air stream in order to convert HO<sub>2</sub> to OH. The mean power was 6.2 mW in the OH axis and 0.63 mW in the HO<sub>2</sub> axis throughout the campaign.

The precision was 0.03 pmol mol<sup>-1</sup> (median) for OH and 0.42 pmol mol<sup>-1</sup> for HO<sub>2</sub> at altitudes higher than 7 km. The limit of detection was calculated from the off-resonance measurement and determined to be 0.016 pmol mol<sup>-1</sup> for OH and 0.33 pmol mol<sup>-1</sup> for HO<sub>2</sub> in the upper troposphere for 1 min average data.

#### 3.2 Calibration

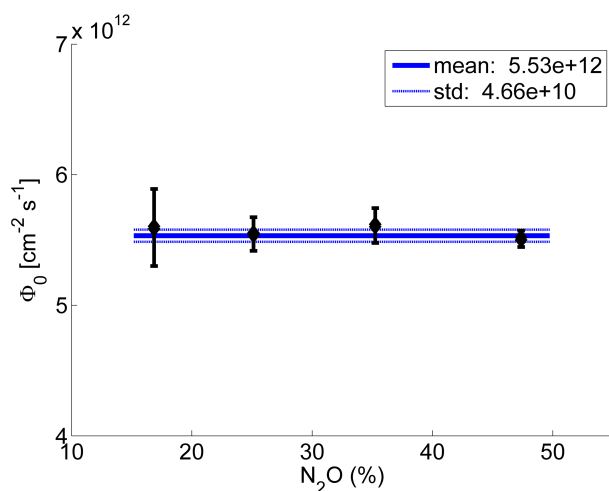
Calibrations were conducted before and after each flight, which started or ended at the airport in Hohn (Germany, 54° N, 9° E). The calibration method of the HORUS instrument is based on the method of Faloon et al. (2004) and described in more detail in Martinez et al. (2010).

As the calibration source, synthetic air (79 % N<sub>2</sub>, 21 % O<sub>2</sub>) was humidified and passed under a Hg Penray lamp. Photolysis of the gaseous water vapour produced OH and HO<sub>2</sub> (R19 and 6, Table 1).

The theoretically produced concentrations of OH and HO<sub>2</sub> were calculated according to Eq. (1):

$$[\text{OH}] = [\text{HO}_2] = \Phi_0 \sigma_{\text{H}_2\text{O}} [\text{H}_2\text{O}] t f_{\text{O}_2}, \quad (1)$$

where  $\sigma$  is the absorption cross section of H<sub>2</sub>O at 184.9 nm, [H<sub>2</sub>O] is the water vapour concentration,  $t$  is irradiation time,  $f$  is a correction factor for the flux reduction due to absorption by O<sub>2</sub> throughout the height of the photolysis chamber, and  $\Phi_0$  is the photon flux of the Penray lamp which was determined by N<sub>2</sub>O actinometry (Fig. 5). This method is described in more detail in Martinez et al. (2010).



**Fig. 5.** Photon flux  $\Phi_0$  calculated from NO measurement for different N<sub>2</sub>O mixing ratios.

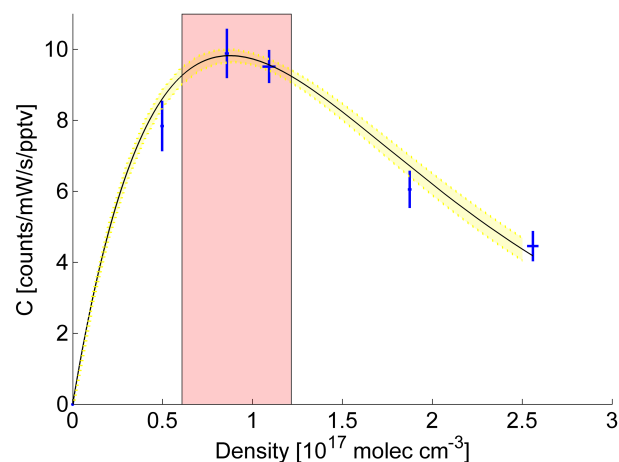
The calculated OH and HO<sub>2</sub> concentrations were corrected for wall losses within the calibration system (6 % for OH and 2 % for HO<sub>2</sub>) and then correlated with the measured fluorescence signal to obtain the instrument sensitivity, while potential OH and HO<sub>2</sub> loss at the wall surface within the instrument is considered through the calibration itself.

The detected fluorescence signal of the excited OH radicals at a given OH concentration depends mainly on laser power and internal pressure. The pressure determines the density and therefore the quantity of wall loss and the quenching efficiency of collisions between excited OH and N<sub>2</sub>, O<sub>2</sub> and H<sub>2</sub>O molecules, respectively. The internal density depends on the nozzle diameter at the inlet and on the ambient pressure, which changes with flight altitude.

Therefore, the ground-based calibration took into account this density dependency of the OH fluorescence signal. The internal pressure was changed by using different pinhole diameters during calibration. All ground-based calibrations were used to calculate a global set of fit parameters. This set has been applied to fit the individual calibrations before and after each flight to determine the sensitivity for each pressure within the range of internal pressures experienced during the flight.

Figure 6 shows the density dependence of the mean sensitivity,  $C$ , observed during the calibration of the HOOVER summer campaign. The variability of the sensitivity with increasing pressure is a function of increasing density (due to an increasing OH concentration), decreasing wall loss and enhanced quenching (Faloona et al., 2004). The highlighted red area indicates the optimal internal pressure range of 2.5 to 5 mbar to be reached in the upper troposphere during the HOOVER flights.

In contrast to Martinez et al. (2010), a significant water dependency other than quenching was not observed during the calibrations. The OH concentration is calculated with the re-



**Fig. 6.** Pressure dependency of the instrument sensitivity. Highest sensitivity is found between 2.5 and 5 mbar in the highlighted red area. Points shown are mean values of individual calibrations; the error bars indicate the variability between the individual calibrations. The black line represents a chi-square fit and the yellow area indicates the uncertainty.

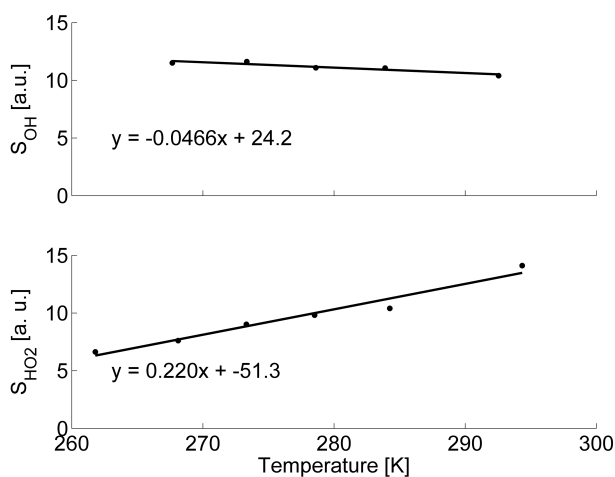
spective sensitivity,  $C$ , and the measured signal,  $S$ , as shown in Eq. (2):

$$[\text{OH}] = \frac{S(\text{OH})}{C(P, \text{H}_2\text{O}, T, \rho)}. \quad (2)$$

### 3.3 Temperature-dependent calibration

Due to its high reactivity, HO<sub>x</sub> radicals could be lost within the detection system by wall contact. The design of the detection system consists of an inlet tube, a detection chamber to record the OH fluorescence signal, a spacer tube with an NO injector and a second detection chamber to record the “HO<sub>2</sub>” fluorescence signal. The sample is drawn through a nozzle into a reduced pressure detection system. Calculations of the velocity field show a compact jet between the inlet nozzle and the detection chamber, indicating no wall contact of the air sample while transiting the inlet tube. However, the jet broadens and makes wall contact downstream of the first detection axis. Thus, OH and HO<sub>2</sub> losses can occur downstream of the first detection axis during measurement and calibration and are parameterized within the calibrated instrument sensitivity, with respect to pressure, density, and humidity variability.

In order to characterize a potential temperature-dependent sensitivity of HO<sub>2</sub> at low temperatures the inlet tube was wrapped in a cooling coil to simulate observed temperature profiles under laboratory conditions at constant OH and HO<sub>2</sub> mixing ratios. While the observed ambient temperatures had been between 256 and 223 K in the upper troposphere the internal temperature within the wing pod had not been lower than 253 K. During the temperature-dependent calibration



**Fig. 7.** A calibration at different temperatures was conducted to study the effect of temperature-dependent surface losses of HO<sub>2</sub> within the inlet and detection system of HORUS under upper tropospheric conditions. Here the signal  $S$ , which is normalized to laser power and time, is shown as a function of temperature.

this observed internal temperature range and its temperature gradient were simulated.

Therefore, the distance between the winding of the cooling lines had been adapted to establish a temperature gradient similar to the one observed during flight. Under these conditions, a slightly higher signal (Fig. 7) appeared at reduced temperatures in the OH axis. In contrast to this, the fluorescence signal of HO<sub>2</sub> significantly decreased in the second axis with decreasing temperature.

Since there is no evidence for a changing OH loss in the inlet tube the change of the OH signal is likely to be caused from cooling down the calibration unit or nearby electronics. Therefore, we corrected the elevated OH signal at reduced temperatures for both fluorescence signals of OH and HO<sub>2</sub>. Additionally, the HO<sub>2</sub> measurement was corrected accordingly to the lower fluorescence signal observed at reduced temperatures.

### 3.4 Interferences

Species which fluoresce at similar wavelengths as OH could interfere with the OH fluorescence measurement. By measuring the off-resonance signal at slightly larger and smaller wavelengths, the broad fluorescence signal of any potential interfering compounds or scatter from detection cell surfaces can be taken into account. The observed off-resonance signal was subtracted from the OH signal.

Possible interferences for our instrument have been studied e.g. during the HO<sub>x</sub>Comp campaign, a formal blind inter-comparison of six HO<sub>x</sub> instruments (4 LIF, 1 DOAS (Differential Optical Absorption Spectroscopy) and 1 CIMS (Chemical Ionisation Mass Spectroscopy)). Measurements were conducted in moderately polluted ambient air masses

and under different conditions in the atmosphere simulation chamber SAPHIR in Jülich, Germany. The inter-comparison of the OH measurements performed with HORUS showed a linear correlation to the measurements of all five other instruments under daylight conditions (Schlosser et al., 2009). No unknown OH-interference was found for H<sub>2</sub>O, O<sub>3</sub>, NO<sub>x</sub>, RO<sub>x</sub> and several VOCs.

Mao et al. (2012) report experimental evidence that indicates an OH interference possibly from oxidation of VOC observed in a California forest within the planetary boundary layer. Also measurements using the HORUS system showed an OH interference during recent ground-based field campaigns.

In contrast to observations within the planetary boundary layer, airborne side-by-side measurement of OH and HO<sub>2</sub> in the troposphere comparing a LIF-based instrument (ATHOS) and CIMS-technique shows in general reasonable agreement between the techniques (Ren et al., 2012). The combined measurement uncertainties are generally covering the differences indicating no additional OH interference.

In the paper of Baker et al. (2010) NMHCs observed in spring (Table 4) partly observed over Europe are highlighted. The observed levels of NMHCs would not introduce a significant OH reactivity or interference. Summertime observations obtained in 2007 from the CARIBIC project showed comparable NMHC mixing ratios in the upper troposphere over Europe compared to the published spring mixing ratios (A. Baker, personal communication, 2012) and compare to published VOC mixing ratios obtained in northern mid-latitudes over continents (Jaegle et al., 2000).

Therefore, no OH interference in the upper troposphere over Europe is expected to affect the OH measurement significantly, even though the uncertainty of the OH measurement might be somewhat higher in convectively influenced air masses.

Fuchs et al. (2011) reports an HO<sub>2</sub> interference with RO<sub>2</sub> radicals for LIF instruments. This is stated to occur due to the NO injection into the air flow, which requires high NO mixing ratios up to more than 1000 μmol mol<sup>-1</sup>, in order to quantitatively convert HO<sub>2</sub> into OH (R12). Simultaneous production of HO<sub>2</sub> from RO<sub>2</sub> can lead to additional OH production via RO<sub>2</sub> (R17 and 18). It was reported that the HO<sub>2</sub> yield is highest when the RO<sub>2</sub> is formed by reaction of OH with unsaturated organic compounds. For small saturated hydrocarbons, the OH forming reactions of CH<sub>3</sub>O<sub>2</sub> (< 5 %, Holland et al., 2003) and C<sub>2</sub>H<sub>5</sub>O<sub>2</sub> are negligible in the short reaction time between NO injection and OH detection.

The HOOVER measurements presented here were conducted in the upper troposphere. Thus, a considerable HO<sub>2</sub> interference impacting the measurement results is not likely, since ambient isoprene mixing ratios were lower than the limit of detection. Even in the air masses probed in the outflow region of the convective element no increased isoprene mixing ratios were observed. Assuming that oxidation of NMHCs (Non-Methane HydroCarbons) during



the convective transport formed RO<sub>2</sub> radicals, the uncertainty of the observed HO<sub>2</sub> mixing ratios can be somewhat higher due to a potential RO<sub>2</sub> interference.

In agreement with this the global model mixing ratios of NMHCs were negligible; e.g. the median upper tropospheric mixing ratio of isoprene was  $1.86 \times 10^{-16}$  mol mol<sup>-1</sup> along the HOOVER flight trajectory. However, no NMHCs have been applied to the box model either. The results of the constrained box model simulations reproduce the OH and HO<sub>2</sub> concentrations indicating that the important HO<sub>x</sub> chemistry is covered. Therefore, differences between the model results and their comparison to the observations are apparently not likely to be caused by the negligible NMHC concentrations within the global model.

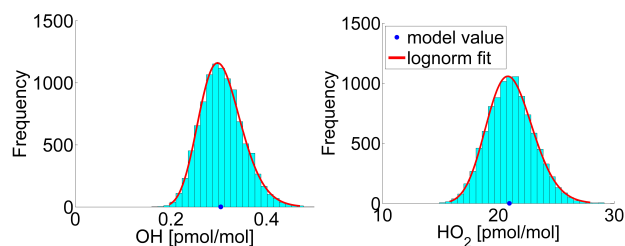
## 4 Box model CAABA/MECCA

### 4.1 Description

We used the atmospheric chemistry box model CAABA/MECCA-2.7b (Chemistry As A Boxmodel Application/Module Efficiently Calculating the Chemistry of the Atmosphere). The parts of the model that were used for this study do not differ substantially from version 3.0 described by Sander et al. (2011a). To connect the box model CAABA to the MECCA chemistry module and to physical processes, we used the MESSy (Modular Earth Submodel System) interface by Jöckel et al. (2005). MESSy formally describes a way to connect different process and diagnostic submodels to a more comprehensive model system. The key ideas are the modular, object-oriented approach, and the strict separation between model infrastructure (such as data exchange, input/output, memory management, etc.) from the actual scientific content (e.g. the parameterizations, numerical solvers, etc.) into 4 different software layers. Information about the implementation of the recent (2nd generation) MESSy infrastructure have been documented by Jöckel et al. (2010). The model simulations were executed in multi-simulation and steady-state mode, see Sander et al. (2011a) for details. Thus, multiple model simulations were performed, and each model simulation continued until the relative changes of OH and HO<sub>2</sub> according to Eq. (4) were less than  $10^{-6}$  within a 15 min time step:

$$\left| \frac{\Delta[\text{OH}]}{[\text{OH}]} \right| = \frac{[\text{OH}(t + \Delta t)] - [\text{OH}(t)]}{[\text{OH}(t)]} \leq 10^{-6}. \quad (3)$$

Measured values were split into data sets for 60 s time intervals, and one model simulation was performed for each data set. The model simulations were constrained by fixing the measured species (NO, CO, O<sub>3</sub>, H<sub>2</sub>O, H<sub>2</sub>O<sub>2</sub>, sum of ROOH as CH<sub>3</sub>OOH, and CH<sub>4</sub>) to their observed values. H<sub>2</sub> was fixed to  $0.6 \mu\text{mol}/\text{mol}^{-1}$ , and CH<sub>4</sub> to  $1.8 \mu\text{mol}/\text{mol}^{-1}$  when not measured.



**Fig. 8.** Distribution of the OH- and HO<sub>2</sub>-mixing ratio derived from Monte-Carlo simulation of the data set observed in the upper troposphere over Europe, see Sect. 4 for details. The model value indicates the simulated mixing ratio resulting from the model run, when no Monte-Carlo variation was applied.

From the comprehensive chemical reaction mechanism (Sander et al., 2011a), we have selected the basic O<sub>3</sub>, NO<sub>x</sub>, HO<sub>x</sub>, and CH<sub>4</sub> chemistry. Reactions of higher hydrocarbons ( $\geq \text{C}_2$ ), halogens, and sulphur compounds were switched off for our model simulation. The simple HO<sub>x</sub> budget scheme, which is expected to dominate HO<sub>x</sub> production and loss in the constrained box model under upper tropospheric conditions, is summarized in Table 1. As mentioned previously, rather low mixing ratios of NMHCs were computed by the global model. In comparison to global model simulations, neglecting the OH reactivity due to reactions with NMHCs should not influence the box model results significantly.

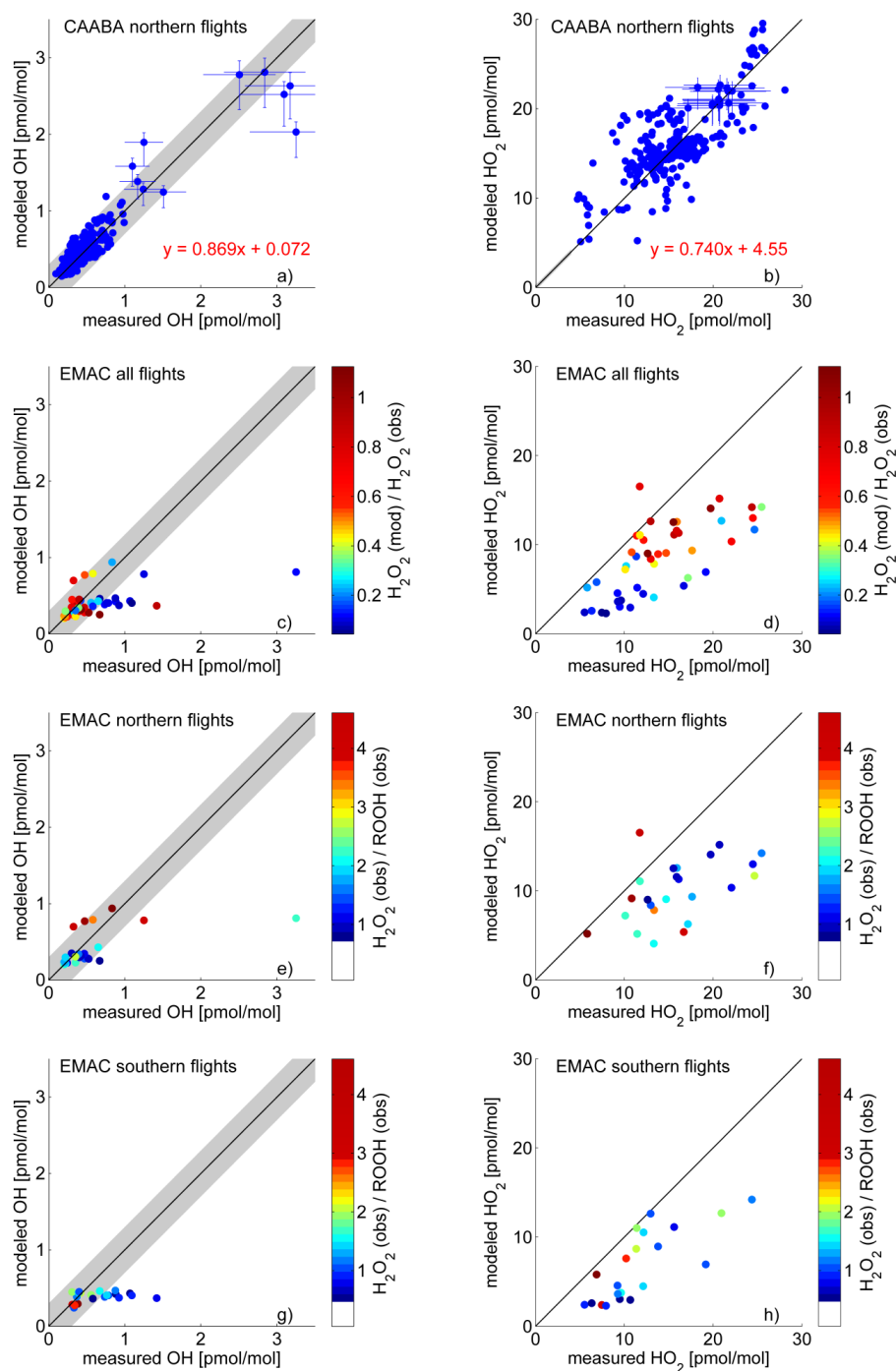
Photolysis frequencies, which were needed as model input, were calculated with the radiative transfer model TUV (Tropospheric Ultraviolet-Visible model V 4.1) (Madronich and Flocke, 1998) along the flight trajectories. As Palancar et al. (2011) stated, the TUV model is able to reproduce an observed actinic flux under cloud-free conditions ( $1.01 \pm 0.04$ ). A comparison of all obtained observations to cloud-free model results showed an enlarged uncertainty of the model results ( $1.1 \pm 0.3$ ).

To minimize cloud effects, we used measured  $J(\text{NO}_2)$  to scale the calculated values for cloud effects as described in Stickler et al. (2006).

CAABA simulations were performed for the northbound flights from southern Germany (48° N) to northern Scandinavia (68° N) only since NO and CO measurements were not available during the southbound flights.

### 4.2 Monte-Carlo simulations

The Monte-Carlo method was used to estimate the uncertainty of simulated OH and HO<sub>2</sub> values as a result of uncertainties in the values of the chemical rate coefficients. Each Monte-Carlo simulation encompasses a set of 9999 individual model simulations. For each of these model simulations, all rate coefficients were multiplied with randomly chosen factors. These factors are scaled with the



**Fig. 9.** Comparison of model calculations to observed OH and HO<sub>2</sub>. CAABA reproduces the observed OH mixing ratios within  $\pm 0.3 \text{ pmol mol}^{-1}$  (shaded area) in the upper troposphere. The 1 : 1 line is given in black, the equation given in red indicates the result of a linear fit. In subplots (a) and (b) a limited number of error bars is shown to improve readability.

uncertainties of the rate coefficients (taken from the IUPAC and JPL recommendations, Atkinson et al., 2007; Sander et al., 2011b) and are centered around 1.

Results of one Monte-Carlo simulation of OH and HO<sub>2</sub> are given in Fig. 8. The calculated median and percentiles of OH

and HO<sub>2</sub> were  $0.304^{+0.02}_{-0.05}$  and  $21.0^{+1.0}_{-2.3} \text{ pmol mol}^{-1}$ , respectively. The percentiles at 31.7 and 68.3 % of the distribution were used as an estimate of the  $1\sigma$ -uncertainty.



### 4.3 Results

CAABA reproduces the observed OH and HO<sub>2</sub> mixing ratios reasonably well. Figure 9a and b shows that more than 98 % of the calculated OH mixing ratios are scattered within a range of  $\pm 0.3$  pmol mol<sup>-1</sup> around the observed values. The scatter of the calculated HO<sub>2</sub> mixing ratios is somewhat higher. The error bars shown in Fig. 9a and b indicate the uncertainties of the measured values (horizontal) and the Monte-Carlo-simulations (vertical), the latter being due to the uncertainty of the rate coefficients.

The number distribution of the calculated quotients of simulated and observed OH mixing ratios result in an asymmetric (skew) distribution with an arithmetic mean value around 1. Since there are no negative values allowed for both of the number distributions half of the calculated distribution is projected to the interval [0, 1], whereas the remaining part is projected to [1 + inf] causing the log-normal distribution.

The box model underestimates the observed OH and HO<sub>2</sub> mixing ratios marginally, since the slope of the fit in Fig. 9a and b, the maximum of the probability distribution and the maximum of the fitted log-normal distribution are all slightly smaller than 1 for OH and HO<sub>2</sub>. The median underestimation is 1.8 % for OH and 3.7 % for HO<sub>2</sub>, typically within the uncertainty of the measurements and of the rate constants used.

### 4.4 Comparison to previous campaigns

Ren et al. (2008) compared observed and (box model) simulated OH and HO<sub>2</sub> mixing ratios (altitude bins) of three airborne campaigns. Two of these campaigns were performed in mid-latitudes. In summer 2004 INTEX-(N)A was carried out over Northern America and the western Atlantic Ocean (Singh et al., 2006). In spring 2001 TRACE-P was conducted over the northern Pacific region (Jacob et al., 2003). PEM-(T)B was performed in spring 1999 in the tropical Pacific area (Browell et al., 2001).

The study of Ren et al. (2008) reports that a constrained box model reproduces the observed OH mixing ratios with good agreement for all three campaigns. Also the observed HO<sub>2</sub> mixing ratios were comparable to the model calculations of the TRACE-P and PEM-(T)B campaign. During the INTEX-(N)A campaign a higher load of pollutants was observed in the air masses of the upper troposphere compared to TRACE-P and PEM-(T)B.

#### 4.4.1 Comparison to INTEX-(N)A

For INTEX-(N)A 0.3 pmol mol<sup>-1</sup> (median value) OH at 7 km and 0.7 pmol mol<sup>-1</sup> OH at 11 km were reported. During the HOOVER campaign median OH mixing ratios were between 0.3 pmol mol<sup>-1</sup> at 7 km and 1 pmol mol<sup>-1</sup> at 11 km. While Ren et al. (2008) report OH mixing ratios between 0.1 and 0.8 pmol mol<sup>-1</sup>, a larger range of OH mixing ratios of up to more than 3 pmol mol<sup>-1</sup> was observed in the upper tro-

posphere during the HOOVER campaign. A similar trend of increasing OH mixing ratios with altitude was observed during both campaigns.

The observed-to-calculated comparison for the INTEX-(N)A campaign showed that the box model used to reproduce the observed OH mixing ratios in the upper troposphere produced good agreement. The same result is found for the comparison between the HOOVER OH observation and box model simulation.

The observed HO<sub>2</sub> mixing ratios during INTEX-(N)A were roughly between 10 and 30 pmol mol<sup>-1</sup> and are comparable to the HOOVER observations of HO<sub>2</sub> mixing ratios of about 5–30 pmol mol<sup>-1</sup>. With increasing altitude a decreasing trend of HO<sub>2</sub> mixing ratios was found for both campaigns.

Under upper tropospheric conditions as observed during INTEX-(N)A the model tended to underestimate HO<sub>2</sub> up to a factor of more than 3 (median values of altitude bins) The planetary boundary layer and mid-troposphere HO<sub>2</sub> mixing ratios are generally satisfactorily reproduced by the box model, while starting at an altitude of 8 km HO<sub>2</sub> is underestimated, with an increasing trend until 11 km height. In contrast to Ren et al. (2008), we did not observe an offset between the measured and the box model simulated HO<sub>2</sub> mixing ratios. We find a good agreement between observations and calculations, indicating no major lack of understanding of the chemistry under the probed background conditions.

#### 4.4.2 Comparison to TRACE-P

During TRACE-P flight 15 air masses in the outflow region of a convective element and convectively unaffected air masses in the upper troposphere were probed. OH mixing ratios between 0.1 and 1 pmol mol<sup>-1</sup> and HO<sub>2</sub> mixing ratios between 6 and 30 pmol mol<sup>-1</sup> were observed. These observations compare well to the HOOVER observations, even though the high OH mixing ratios of more than 3 pmol mol<sup>-1</sup> observed in the outflow region over Germany are not reached.

During both campaigns a similar trend of increasing OH mixing ratios with altitude was observed in the upper troposphere. During TRACE-P flight 15 HO<sub>2</sub> slightly increases with altitude, while the HOOVER observation shows a decreasing trend of HO<sub>2</sub> mixing ratios.

The model used reproduces OH reasonably (Ren et al., 2008) and likewise CAABA/MECCA reproduces the OH observed during HOOVER. However HO<sub>2</sub> mixing ratios are well reproduced by the CAABA/MECCA model while HO<sub>2</sub> mixing ratios observed during TRACE-P are underpredicted by 33–50 %, showing an increasing trend with altitude.

However, upper tropospheric OH and HO<sub>2</sub> mixing ratios obtained outside of directly convectively influenced air masses were not higher than roughly 0.3 pmol mol<sup>-1</sup> for OH and mostly lower than 20 pmol mol<sup>-1</sup> for HO<sub>2</sub>,

respectively. Therefore significantly lower OH mixing ratios were observed during the TRACE-P campaign if considering the whole campaign. HO<sub>2</sub> mixing ratios are comparable in general, even though higher HO<sub>2</sub> mixing ratios of up to 30 pmol mol<sup>-1</sup> were observed during the HOOVER campaigns.

#### 4.4.3 Comparison to PEM-(T)B

OH and HO<sub>2</sub> observations obtained during PEM-(T)B (Tan et al., 2001) were 0.1 and 0.3 pmol mol<sup>-1</sup> (median values), respectively, and thus much lower than the 1 pmol mol<sup>-1</sup> (median value) reached during HOOVER. HO<sub>2</sub> mixing ratios are between 3 and 12 pmol mol<sup>-1</sup> (median values) and show a higher variability than the HOOVER observation of 10–20 pmol mol<sup>-1</sup> (median values).

An increasing trend of OH mixing ratios and a mostly decreasing trend of HO<sub>2</sub> mixing ratios as a function of altitude was found during PEM-(T)B and is generally similar to the trends observed during HOOVER.

While the model used to reproduce the OH mixing ratios well for the INTEX-(N)A and TRACE-P campaigns, OH is underestimated by up to 50 % for the PEM-(T)B campaign. In contrast the HO<sub>2</sub> is well reproduced for the PEM-(T) conditions, while HO<sub>2</sub> was underpredicted for the INTEX-(N)A and TRACE-P campaigns.

During HOOVER 2, in the upper troposphere observed OH mixing ratios are on average comparable to those observed during INTEX-(N)A and higher than the observed OH mixing ratios of the PEM-(T)B, and TRACE-P campaigns. However, OH mixing ratios observed under NO<sub>x</sub> levels higher than 50 pmol mol<sup>-1</sup> exceed the observed OH mixing ratios of the respective campaigns. Observed HO<sub>2</sub> mixing ratios compare to those found during the INTEX-(N)A campaign and also compare to the HO<sub>2</sub> mixing ratios observed during the TRACE-P campaign, even though higher HO<sub>2</sub> mixing ratios were found during the HOOVER 2 campaign.

With regard to CO, NO and ozone, the conditions observed during HOOVER 2 are more comparable to INTEX-(N)A, while the higher OH mixing ratios found are likely due to a higher HO<sub>2</sub> conversion rate caused by enhanced NO mixing ratios observed in convectively affected air masses in the upper troposphere over Europe (Olson et al., 2006, 2004; Ren et al., 2008; Tan et al., 2001).

## 5 General circulation model EMAC

### 5.1 Description

EMAC (ECHAM5 version 5.3.02/MESSy version 2.41 Atmospheric Chemistry) is based on the general circulation model ECHAM5 (ECMWF, Hamburg, Version 5, Roeckner et al., 2006), which is coupled to the Modular Earth Submodel System (MESSy, Jöckel et al., 2010). The chemistry

scheme is MECCA as also used in CAABA/MECCA box model simulations. However, in the 3D simulation analysed here, the initialized number of trace gas species is much larger and includes NMHC species compared to the box model and the species concentrations are not constrained by measurements. A description of the applied emission inventory is given in Jöckel et al. (2010).

The submodels ONLEM, OFFLEM, TNUDGE and DRY-DEP were applied to calculate primary emissions and dry deposition of trace gases and aerosols (Kerkweg et al., 2006a, b). Wet scavenging on cloud particles and aqueous-phase chemistry in cloud droplets were simulated with the submodel SCAV (Tost et al., 2006).

EMAC was applied in the T42L90MA-resolution, i.e. with a spherical truncation of T42 (corresponding to a quadratic Gaussian grid of  $\approx 2.8$  by  $2.8$  degrees in latitude and longitude) with 90 vertical hybrid pressure levels up to 0.01 hPa.

In order to fly “virtually” through the grid boxes of the global numerical chemistry and climate simulation system the S4D submodel was used (Jöckel et al., 2010). It performs a bi-linear interpolation of the model results in space and time along the flight track positions.

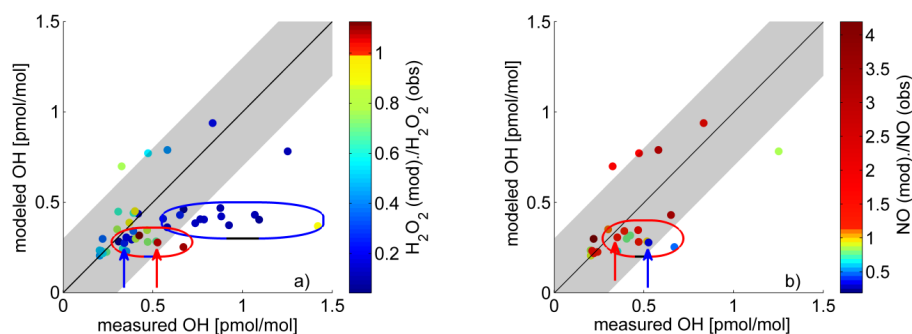
$$\text{FTT} \subset \text{MT} \pm \frac{\Delta t}{2} \quad (4)$$

For each time interval, S4D searches for the corresponding measurement point along the flight track according to Eq. (4). Whenever a flight track time point (FTT) is within half of the model time step ( $\Delta t/2$ ) around the model time (MT) the submodel performs two horizontal interpolations using the 4 closest grid points at two model levels (above and below the flight track), and then a linear interpolation in the vertical to the flight track. The result is written into an output file.

This method provides a model-calculated value of all tracers and reaction rates along the flight track with a time resolution of the model time step (12 min in this analysed simulation).

### 5.2 Results

In general, EMAC significantly underestimates the observed OH and HO<sub>2</sub> mixing ratios (Fig. 9c, d) in the upper troposphere along the flight tracks. Most of the simulated OH values are scattered around the observed ones within a range of  $\pm 0.3$  pmol mol<sup>-1</sup>, as also perceived from the CAABA simulation. However, the fraction (75 %) of simulated values within the range is substantially smaller compared to the CAABA calculations (98 %). Furthermore, the EMAC model systematically underestimates OH and HO<sub>2</sub> (Fig. 9c, d). Overall 67 % of the calculated OH values and 94 % of the calculated HO<sub>2</sub> values are smaller than the observed. The median underestimation of OH and HO<sub>2</sub> compared to the observations is 24 % and 41 %, respectively.



**Fig. 10.** Highlighted are underestimated OH mixing ratios at underestimated H<sub>2</sub>O<sub>2</sub> mixing ratios (blue ellipse). Within the red ellipse some data points show slightly underestimated OH at relatively well reproduced H<sub>2</sub>O<sub>2</sub> but severely underestimated NO, while at other times OH was reasonably reproduced even though H<sub>2</sub>O<sub>2</sub> mixing ratios in the model were much too low and NO was overestimated. Figure 10a is a cut-out of Fig. 9c. To highlight the statement the colour coding was adapted.

The comparison of EMAC simulated and observed OH and HO<sub>2</sub> mixing ratios also show a log-normal like distribution. The log-normal like distribution is found because there are no negative values allowed for both of the number distributions. As previously described for CAABA, half of the calculated distribution is projected to the interval [0, 1], whereas the remaining part is projected to [1 + inf]. However, in contrast to the CAABA results this number distribution is shifted towards a large underestimation of OH and HO<sub>2</sub>.

Since both models utilize a similar chemical mechanism the underestimation of both species by the global model indicates that (1) precursor mixing ratios and (2) possibly reaction and photolysis rates are less well reproduced within the unconstrained 3-D model, because the global model is not constrained to changes of meteorological parameters such as temperature, and pressure. Enhanced photolysis frequencies were observed closely above clouds and can be underestimated due to a different distribution of clouds in the model simulation. Furthermore, the 3-D model simulates additional processes not represented by the box model, such as large-scale, convective, and diffusive tracer transport and wet-scavenging of soluble constituents. Small-scale processes like convective transport cannot be resolved by large-scale models leading to over- and underestimations of trace gases in the plumes.

A comparison of the observed and simulated OH and HO<sub>2</sub> mixing ratios (Figs. 9–11) and the computed HO<sub>x</sub> formation (Fig. 12) highlights several differences in the OH budget, which directly affect the calculated OH mixing ratios. In the following comparison between observation and the global 3-D-model output, we focus on the impact of hydrogen peroxide as a primary OH source, the impact of CO and NO on the HO<sub>x</sub> equilibrium and the influence of convective transport on the HO<sub>x</sub> budget comparison.

### 5.2.1 Influence of H<sub>2</sub>O<sub>2</sub>

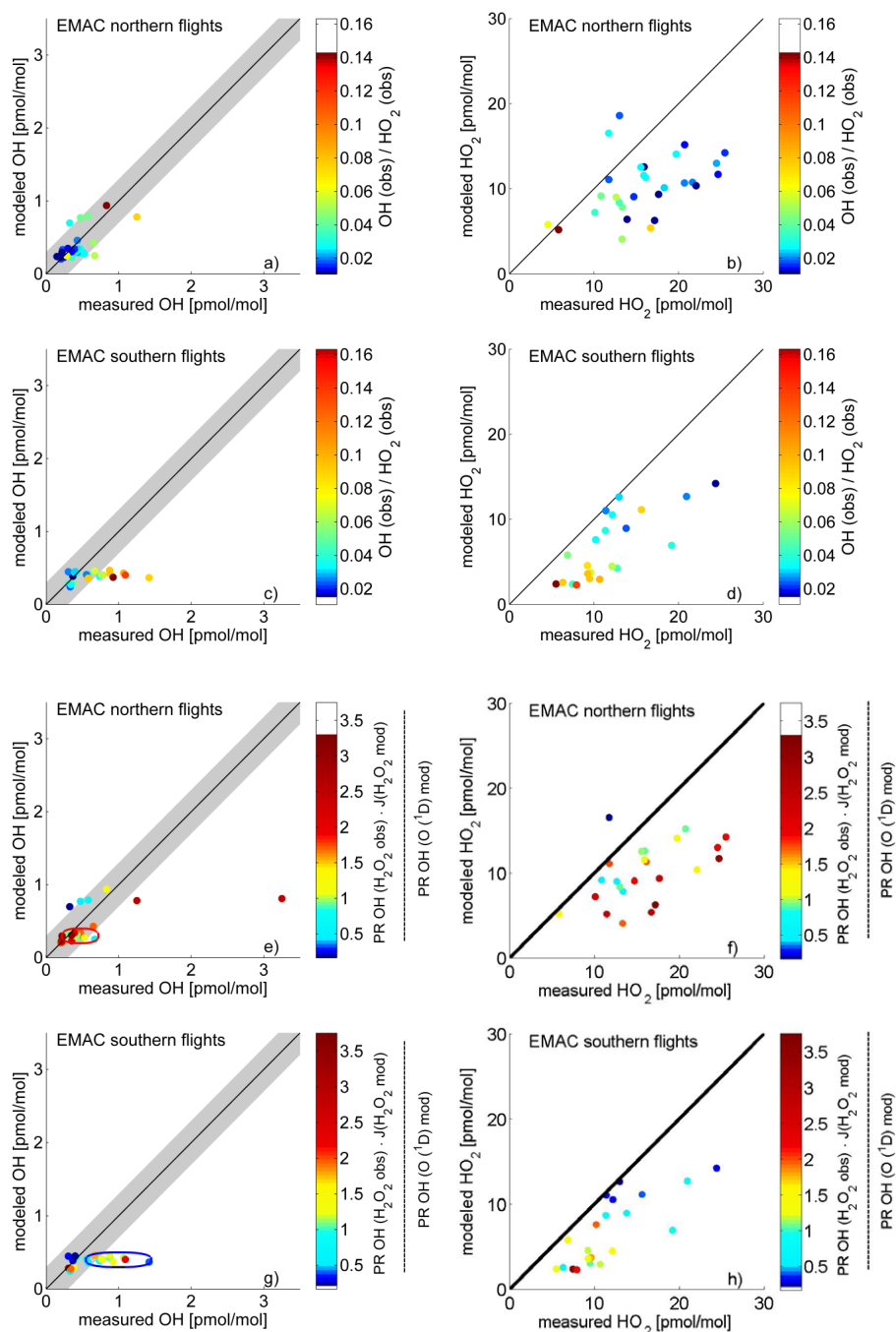
In Fig. 10a the blue ellipse highlights that the EMAC model tends to underestimate the observed OH mixing ratios most strongly when the H<sub>2</sub>O<sub>2</sub> mixing ratio is strongly underestimated compared to the observations.

This underprediction of OH is not always directly linked to the quality of the H<sub>2</sub>O<sub>2</sub> reproduction within the global model. The red ellipse in Fig. 10a highlights data when the model reproduces the OH even though the H<sub>2</sub>O<sub>2</sub> is clearly underestimated (blue arrow). In contrast, for other points for which H<sub>2</sub>O<sub>2</sub> mixing ratios are well reproduced by the model, the OH is underestimated (red arrow).

Figure 10b shows that when the OH is well reproduced despite underestimated H<sub>2</sub>O<sub>2</sub> mixing ratios, NO mixing ratios are simultaneously overestimated (red arrow in subplot b). Apparently, the underrated OH source from H<sub>2</sub>O<sub>2</sub> photolysis is compensated in the model by a stronger OH source from reactions of NO. When OH is underestimated despite well reproduced H<sub>2</sub>O<sub>2</sub>, NO mixing ratios are underestimated (blue arrow) leading to a too low OH formation from this source.

HO<sub>2</sub> is almost always underestimated, even if the H<sub>2</sub>O<sub>2</sub> is well reproduced by the model, as seen in Fig. 9d. A clear correlation links the lowest and the most underestimated HO<sub>2</sub> mixing ratio to the level of H<sub>2</sub>O<sub>2</sub> underestimation. Coupled convective transport of H<sub>2</sub>O<sub>2</sub> and NO<sub>x</sub> and local NO<sub>x</sub> formation by lightning might lead to an underestimation of H<sub>2</sub>O<sub>2</sub> and NO<sub>x</sub> in the convectively influenced air masses within the upper troposphere. Therefore, an underestimation of simulated OH can be promoted from an underestimated HO<sub>2</sub> conversion into OH.

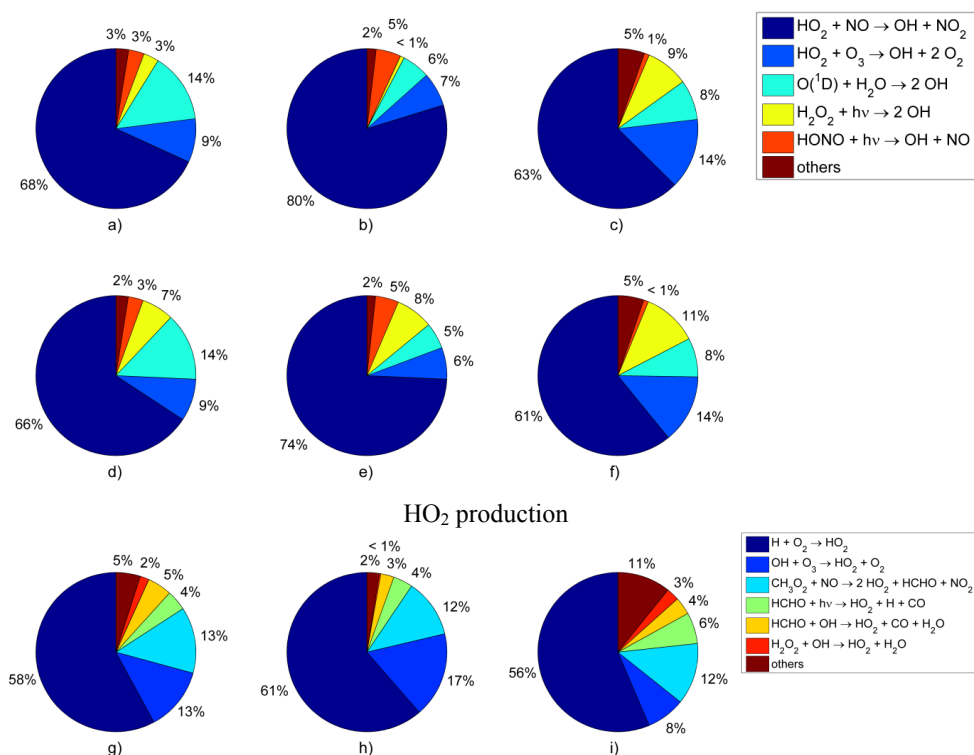
The OH production through H<sub>2</sub>O<sub>2</sub> photolysis varies between less than 1 and 9 % when the simulated H<sub>2</sub>O<sub>2</sub> is not constrained to measurements (Fig. 12). The most important primary HO<sub>2</sub> source (HCHO photolysis) is more stable and contributes 4 % to the total HO<sub>2</sub> production under all separated conditions (Fig. 12). This is consistent with a study of Klippel et al. (2011), which reports that HCHO



**Fig. 11.** Comparison of model simulation to observed OH and HO<sub>2</sub> mixing ratios. Colour coding highlights the conversion ratio and the ratio of primary production rates. The 1 : 1 lines are given in black.

is almost homogeneously distributed over central Europe. Since EMAC is realistically reproducing altitudinal and latitudinal gradients of the HCHO distribution and also capturing the elevated HCHO mixing ratios observed over the Mediterranean, EMAC captures the dominant primary HO<sub>2</sub> source through HCHO photolysis.

Sequentially, a sensitivity study (Klippel et al., 2011) showed that parameterization of clouds and precipitation leads to overestimated wet scavenging of H<sub>2</sub>O<sub>2</sub>. Switching off wet scavenging had no major impact on HCHO but improved the reproduction of H<sub>2</sub>O<sub>2</sub> significantly and therefore increased the OH formation through H<sub>2</sub>O<sub>2</sub> photolysis.



**Fig. 12.** Shown is the relative contribution of the main OH and HO<sub>2</sub> forming reactions to the total OH and HO<sub>2</sub> production during northern flight tracks, southern flight tracks and for the conditions the global model tends to underestimate the OH mixing ratio. Subplots (d), (e) and (f) show the OH formation with simulated H<sub>2</sub>O<sub>2</sub> corrected to observed values.

### 5.2.2 OH production rate by H<sub>2</sub>O<sub>2</sub> photolysis

Since the OH forming reaction of O(<sup>1</sup>D) weakens with altitude, due to lower water content of the air masses, other OH-producing pathways, like photolysis of H<sub>2</sub>O<sub>2</sub> and HCHO, become more important (Tan et al., 2001). Figure 12a–c shows the relative contribution of H<sub>2</sub>O<sub>2</sub> photolysis to the total OH production as the model simulates it. Compared to the O(<sup>1</sup>D)-source of OH (14 % of the total OH production), H<sub>2</sub>O<sub>2</sub> photolysis is relatively weak (3 %) on southern flight tracks. On northern flight tracks the H<sub>2</sub>O<sub>2</sub> photolysis (9 %) exceeds the weaker O(<sup>1</sup>D) contribution of 8 %.

It was shown above in Fig. 9c and d that H<sub>2</sub>O<sub>2</sub> is underestimated by the model. This leads to a too small primary OH source through H<sub>2</sub>O<sub>2</sub> photolysis. The colour coding of Fig. 11e–h gives the relative importance of the H<sub>2</sub>O<sub>2</sub>-dependent OH source in relation to the strength of the O(<sup>1</sup>D)-dependent OH source. Thus, an OH production budget corrected to observed H<sub>2</sub>O<sub>2</sub> mixing ratios (Fig. 12d–f) underlines the higher contribution of the H<sub>2</sub>O<sub>2</sub> photolysis to the (primary) OH production for all three cases (southern and northern flight tracks and for underestimated OH). A comparison of the upper to the middle panel of Figure 12 indicates that the contribution of H<sub>2</sub>O<sub>2</sub> photolysis to the OH production is underestimated most for the outlying dots of

OH mixing ratios on the southern leg (see blue rectangle in Fig. 11g). Thus a correction of the simulated H<sub>2</sub>O<sub>2</sub> to observed H<sub>2</sub>O<sub>2</sub> mixing ratios extend the H<sub>2</sub>O<sub>2</sub> contribution from 0.79–7.60 % for the total OH production and from 6.87–43.3 % for the primary OH production.

The H<sub>2</sub>O<sub>2</sub> photolysis is further underestimated since the model tends to underestimate photolysis rates. Table 2 highlights mean values of J(O<sup>1</sup>D) indicating an underestimation of about 30 % between simulated and J(O<sup>1</sup>D) observed under background conditions. Photolysis frequencies are even more underestimated when convectively influenced air masses have been probed. Photolysis frequencies were most likely enhanced because the flight track was close above and around clouds, causing high scattering of sunlight. Such clouds are likely not resolved, in particular at the observed position, by the global model.

An additional model run with scaled (by a factor of 1.4) photolysis frequencies to match the observation was performed in order to investigate changes of the HO<sub>x</sub> mixing ratios. Slightly reduced H<sub>2</sub>O<sub>2</sub> and HCHO and enhanced ozone introduce an enhanced primary HO<sub>x</sub> production rate due to an increase of the photolysis frequencies. Therefore, higher OH (50 %) and HO<sub>2</sub> (5.6 %) mixing ratios were simulated by the global model (Table 2) due to enhanced photolysis frequencies. However, HO<sub>2</sub> and thus HO<sub>x</sub> mixing ratios are still

underestimated ( $\sim 40\%$  compared to observed background conditions, see Table 2). Increased NO mixing ratios enhance the HO<sub>2</sub> conversion into OH and lower CO mixing ratios weaken the back-cycling into HO<sub>2</sub>. Therefore, the model overestimates OH mixing ratios (46% to background conditions) in average but still underestimates the OH mixing ratios when H<sub>2</sub>O<sub>2</sub> is underestimated.

A more complex situation is found in the highlighted area shown in Fig. 11e. Here, the slight underestimation of OH might be explained by the H<sub>2</sub>O<sub>2</sub> mixing ratio, since the H<sub>2</sub>O<sub>2</sub>-dependent OH production rate is an important contributor to the overall primary OH production. However, no correlation to the degree of H<sub>2</sub>O<sub>2</sub> underestimation could be found (Fig. 10a). Figure 10b indicates in the same area a correlation between the degree of OH underestimation and the underestimation of NO, since OH is better reproduced at higher NO ratios (modelled to observed) and lower H<sub>2</sub>O<sub>2</sub> ratios. In contrast to the data shown in the blue ellipse (same Figures), here OH is underestimated when the model reproduces the H<sub>2</sub>O<sub>2</sub>. Thus, a missing HO<sub>2</sub> conversion rate is likely responsible for the OH underestimation.

The comparison of observed and simulated HO<sub>2</sub> shows a correlation between the degree of HO<sub>2</sub> underestimation and the missing H<sub>2</sub>O<sub>2</sub> (Fig. 9d) as well as the importance of the H<sub>2</sub>O<sub>2</sub>-dependent OH source (Fig. 11f, h) for the OH.

A high OH mixing ratio of more than 3 pmol mol<sup>-1</sup> was observed in the outflow region of a convective system over Germany. The upper tropospheric composition was in this case affected by fresh convective outflow, not resolved by the global model, which parameterizes these small-scale processes. Nevertheless, these observations give an indication of how the mixing ratios of different tracers in the EMAC model can be related to air masses within the outflow region of a small-scale convective regime.

### 5.2.3 Influence of convective transport

For the southern flight tracks NO and CO measurements were not available. Therefore, the ratio between H<sub>2</sub>O<sub>2</sub> and organic peroxides (referred as ROOH) was taken as a proxy for convective transport. Even though identifying convectively influenced air masses by the peroxide ratio has a larger uncertainty than identifying convective influence through other trace gas mixing ratios convectively influenced air masses were identified using the ratio of H<sub>2</sub>O<sub>2</sub> to CH<sub>3</sub>OOH (based on (Snow et al., 2007) and references therein). In our study, enhanced H<sub>2</sub>O<sub>2</sub> mixing ratios were typically found in convectively affected air masses. Nevertheless, the H<sub>2</sub>O<sub>2</sub>/ROOH ratio was reduced, since ROOH was even more strongly enhanced and less scavenged through the aqueous phase removal and rainout. A more detailed discussion will be presented in Bozem et al. (2013).

Small peroxide ratios indicate influences of convection on the trace gas distribution whenever the OH and HO<sub>2</sub> underestimation is prominent on the southern flight tracks (Fig. 9g,

h). As indicated by the OH/HO<sub>2</sub> ratio (Fig. 11c), along with peroxides NO<sub>x</sub> was convectively injected into the upper troposphere, shifting the HO<sub>x</sub> equilibrium towards the OH. Underprediction of HO<sub>2</sub> then results in a promoted OH underestimation, since a primary OH source (photolysis of H<sub>2</sub>O<sub>2</sub>) and HO<sub>2</sub> cycling (with NO) into OH is underestimated.

The observed OH and HO<sub>2</sub> mixing ratios of the northern flight tracks do not appear to be much affected by convective transport.

### 5.2.4 HO<sub>x</sub> equilibrium

The HO<sub>x</sub> equilibrium is mainly influenced by NO and CO (R12 and 14), which inter-convert the two species (Jaegle et al., 2001). The HO<sub>x</sub> sum is virtually always underestimated by the global model. An underestimation of CO and a mostly overestimated NO reproduces the OH relatively well, while HO<sub>2</sub> is strongly underestimated (see Fig. 10b, red ellipse). HO<sub>2</sub> is reproduced best when the CO-ratio (simulated/measured not shown) is close to 1 and calculated OH cycling rates into HO<sub>2</sub> are close to realistic values.

However, the model tends to underestimate the photolysis frequencies (see J(O<sup>1</sup>D) in Table 2) along the flight tracks. Since enhanced photolysis frequencies introduce a higher primary HO<sub>x</sub> production within the model mechanism, HO<sub>x</sub> mixing ratios would increase and HO<sub>2</sub> mixing ratios match the observations somewhat better. However, HO<sub>2</sub> mixing ratios still remain underestimated while the mean OH mixing ratio is overestimated; both indicating that the HO<sub>x</sub> budget is not reproduced by the simulation. OH mixing ratios observed along with underestimated H<sub>2</sub>O<sub>2</sub> mixing ratios (Fig. 10a, blue ellipse) remain underestimated since H<sub>2</sub>O<sub>2</sub> mixing ratios are heavily underestimated while the photolysis frequencies are moderately only underestimated.

### 5.2.5 OH/HO<sub>2</sub> ratio

The variability of the OH/HO<sub>2</sub> ratio is mainly influenced by the NO variability (Ren et al., 2008), which is usually higher than the CO variability. During background conditions as observed during the HOOVER summer flights shown here, the mean CO mixing ratio was 97.5 pmol mol<sup>-1</sup>, with a standard deviation of 15% and the NO mean value was 57.5 pmol mol<sup>-1</sup> with a standard deviation of 99%. Other studies (Jaegle et al., 2001; Martinez et al., 2003; Ren et al., 2008) report the same NO related OH/HO<sub>2</sub> behaviour. Accordingly, the OH/HO<sub>2</sub> ratio can be used to display relative changes of the HO<sub>2</sub> conversion rate due to changes in NO.

The colour coding in Fig. 11a–d displays increasing OH mixing ratios with an increasing OH/HO<sub>2</sub> ratio along the northern flight leg and an increasing OH underestimation along the southbound flight leg. In both regions, smaller HO<sub>2</sub> mixing ratios are found during high OH/HO<sub>2</sub> ratios and no correlation to the level of underestimation of HO<sub>2</sub> is observed.



### 5.2.6 Comparison to previous model studies

HO<sub>x</sub> observations from the ATHOS instrument were compared to results of GEOS-Chem simulations indicating that the HO<sub>x</sub> budget is not entirely resolved by the global model. For the INTEX-(N)A campaign an OH and HO<sub>2</sub> overestimation of 60 and 30 %, respectively, were reported for higher altitudes (Hudman, 2007). A later recalibration of the ATHOS instrument revised OH and HO<sub>2</sub> mixing ratios to be a factor of 1.64 higher (Ren et al., 2008). Since the global model then tends to underestimate HO<sub>2</sub>, the HO<sub>2</sub> and HO<sub>x</sub> budget remains uncovered. Zhang et al. (2008) report that the same global model reproduces the HO<sub>2</sub> observations obtained during the INTEX-B campaign but overestimates OH by 27 %. A comparison of observed HO<sub>x</sub> mixing ratios observed during the ARCTAS campaign to the global model found overestimated HO<sub>2</sub>, speculatively caused by underestimated uptake of HO<sub>2</sub> on aerosols (Mao, 2010).

## 6 Summary and conclusions

The HO<sub>x</sub> measurements performed during the HOOVER summer campaign show unexpectedly high OH mixing ratios up to more than 3 pmol mol<sup>-1</sup> and HO<sub>2</sub> mixing ratios up to more than 25 pmol mol<sup>-1</sup> in the upper troposphere over central and northern Europe. The comparison of the measurements to the box model CAABA/MECCA calculated mixing ratios reveal excellent agreement between calculations and observations. These calculations were done with a simple chemical mechanism, constrained by the measurements, indicating no major lack of understanding of the chemistry under the probed background conditions. The comparison to the global 3-D circulation model EMAC (executed with a similar chemical mechanism as used for the constrained CAABA calculations) shows a substantial and systematic underestimation of the observed OH and HO<sub>2</sub> mixing ratios. This underestimation of the HO<sub>x</sub> mixing ratios is caused by underestimated HO<sub>x</sub> precursor mixing ratios, in particular H<sub>2</sub>O<sub>2</sub>, related to the treatment of this gas in the transport and deposition routines. The study of Klippel et al. (2011) indicates that the global model fails to reproduce H<sub>2</sub>O<sub>2</sub> since microphysical processes of wet scavenging of soluble trace gas species is neither fully understood nor well parameterized within the global model. Since H<sub>2</sub>O<sub>2</sub> can be an important upper tropospheric primary source of OH, the global model consistently underestimates the OH mixing ratio when underestimating the H<sub>2</sub>O<sub>2</sub>. However, the model tends to underestimate the photolysis frequencies (see J(O<sup>1</sup>D) in Table 2). An analysis of the OH formation budget shows that a correction of simulated H<sub>2</sub>O<sub>2</sub> mixing ratios to observed mixing ratios increases the OH formation through H<sub>2</sub>O<sub>2</sub> photolysis by up to a factor of 10.

The model tends to underestimate not only H<sub>2</sub>O<sub>2</sub> but also NO mixing ratios in convectively transported air masses. Fur-

thermore, photolysis frequencies are underestimated by the global model along the flight tracks. Scaling the photolysis frequencies to match the observation results in overestimated OH mixing ratios and slightly less underestimated HO<sub>2</sub> mixing ratios. The HO<sub>2</sub> and thus HO<sub>x</sub> mixing ratios remain underestimated. Increased photolysis rates result in enhanced photolysis of HCHO and an increase in all photolytic HO<sub>x</sub> sources. Simulated HCHO and H<sub>2</sub>O<sub>2</sub> mixing ratios would be slightly lower due to this higher photolytic loss of the two tracers.

Since the model reproduces the observed gradients of HCHO realistically (Klippel et al., 2011), no further HO<sub>2</sub> formation is indicated due to underestimated HCHO.

For the southern flight tracks NO and CO measurements were not available. As indicated by the OH/HO<sub>2</sub> ratio, along with peroxides NO<sub>x</sub> was convectively injected into the upper troposphere shifting the HO<sub>x</sub> equilibrium towards OH. Underprediction of HO<sub>2</sub> than results in a stronger OH underestimation, since a primary OH source (photolysis of H<sub>2</sub>O<sub>2</sub>) as well as HO<sub>2</sub> cycling (with NO) into OH is underestimated.

Apparently, in the case of a NO overestimation the too high HO<sub>2</sub> conversion rates can balance the underestimated H<sub>2</sub>O<sub>2</sub> photolysis rates and the simulated OH is in agreement with the observations. In this case, the HO<sub>x</sub> budget is not represented within the model and OH is reproduced through incorrect source strengths. However, further observations and comparisons are needed to investigate the significance of this finding.

As other studies have shown, it is a challenge for global models to calculate the transport of highly soluble trace gases like H<sub>2</sub>O<sub>2</sub> and HCHO (Klippel et al., 2011; Tost et al., 2006). Models refined to solve small-scale processes are needed to investigate convective transport processes and scavenging of soluble trace species, as well as in cloud chemistry and the rejection of peroxides during freezing (Lelieveld and Crutzen, 1990; Mari et al., 2000).

We have shown that the uncertainties related to these processes lead to substantial ambiguity in the model calculated oxidation capacity of the upper troposphere.

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